The radiative lifetime of metastable CO (a^3Π, v=0)

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(Received 11 October 2007; accepted 26 October 2007; published online 13 December 2007)

We present a combined experimental and theoretical study on the radiative lifetime of CO in the a^3Π1,2, v=0 state. CO molecules in a beam are prepared in selected rotational levels of this metastable state, Stark-decelerated, and electrostatically trapped. From the phosphorescence decay in the trap, the radiative lifetime is measured to be 2.63±0.03 ms for the a^3Π1, v=0, J=1 level. From the spin-orbit coupling between the a^3Π1 and the A^1Π states a 20% longer radiative lifetime of 3.16 ms is calculated for this level. It is concluded that coupling to other ^1Π states contributes to the observed phosphorescence rate of metastable CO. © 2007 American Institute of Physics. [DOI: 10.1063/1.2813888]

Triggered by the observation of CO in the upper atmosphere of Mars with a UV spectrometer on board the Mariner 6 spacecraft in 1969,1 researchers started studying the UV sphere of Mars with a UV spectrometer on board the Mariner spacecraft. The main mechanism for the short lifetime of CO in the upper atmosphere of Mars was believed to be photodissociation on the H20849-X2Π ground state. However, the detection of CO in an upper atmospheric region with a sufficiently high temperature to allow efficient photodissociation of CO was not straightforward, as CO was found at a lower altitude than expected. The detection of CO in the UV sphere of Mars was later confirmed by observations with the Mariner 9 spacecraft in 1971. The observed UV spectrum was consistent with the absorption line-shape and the spectral profile of the light source to be accurately known. It is more problematic that the Franck-Condon factors for the vibrational bands of this transition also have to be known to be able to deduce a radiative lifetime. For phosphorescence decay measurements, on the other hand, the rovibrational state distribution of the metastable molecules needs to be known, and, on the timescale of the phosphorescence, collisions need to be avoided. To achieve this, laser excitation of CO to single rovibrational levels of the metastable state in the collision-free environment of a molecular beam has been used.8–10 An intrinsic problem in these molecular beam experiments is, however, that the molecules move with a high speed, typically limiting the time during which the phosphorescence can be detected to a fraction of a millisecond.

The Stark deceleration of a beam of polar molecules followed by electrostatic trapping enables the observation of state-selected molecules for several seconds.11 Recently, the radiative lifetime of vibrationally excited OH radicals was determined by recording the temporal decay of their population in the trap.12 Here we report the Stark deceleration and electrostatic trapping of metastable CO molecules that are laser prepared in either the (1,0,1) or in the (2,0,2) level. The radiative lifetimes are measured by monitoring the phosphorescence decay of the trapped molecules. Calculations to rationalize the observed lifetimes are presented.

The Stark deceleration of metastable CO molecules has been described in detail before.13 In the present experiment, we use a Stark decelerator with 108 electric field stages to load the metastable CO molecules in an electrostatic trap. The experimental setup is schematically shown in Fig. 1. A pulsed beam of CO molecules with a mean velocity of 320 m/s is produced by expanding a mixture of 20% CO in Xe from a cooled pulsed valve (T=203 K). A packet of CO molecules in the upper (low-field seeking) A-doublet component of the (1,0,1) or (2,0,2) level in the metastable a^3Π

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References

1. James, J. J. Phys. Chem. 1997, 101, 11152–11156.

2. Gilijamse, J. J.; Hoekstra, S.; Meek, S. A.; Metsälä, M.; van de Meerakker, S. Y. T.; Meijer, G. Chem. Phys. Lett. 2005, 401, 157–161.

3. Gilijamse, J. J.; Groenenboom, G. C.; Meek, S. A.; Metsälä, M.; van de Meerakker, S. Y. T.; Meijer, G. J. Chem. Phys. 2007, 127, 221102-1–221102-4.

4. Gilijamse, J. J.; Hoekstra, S.; Meek, S. A.; Metsälä, M.; van de Meerakker, S. Y. T.; Meijer, G. J. Chem. Phys. 2005, 123, 125801.

5. Gilijamse, J. J.; Hoekstra, S.; Meek, S. A.; Metsälä, M.; van de Meerakker, S. Y. T.; Meijer, G. J. Chem. Phys. 2006, 124, 134303.

6. Gilijamse, J. J.; Hoekstra, S.; Meek, S. A.; Metsälä, M.; van de Meerakker, S. Y. T.; Meijer, G. J. Chem. Phys. 2006, 124, 134303.

7. Gilijamse, J. J.; Hoekstra, S.; Meek, S. A.; Metsälä, M.; van de Meerakker, S. Y. T.; Meijer, G. J. Chem. Phys. 2006, 124, 134303.

8. Gilijamse, J. J.; Hoekstra, S.; Meek, S. A.; Metsälä, M.; van de Meerakker, S. Y. T.; Meijer, G. J. Chem. Phys. 2006, 124, 134303.

9. Gilijamse, J. J.; Hoekstra, S.; Meek, S. A.; Metsälä, M.; van de Meerakker, S. Y. T.; Meijer, G. J. Chem. Phys. 2006, 124, 134303.

10. Gilijamse, J. J.; Hoekstra, S.; Meek, S. A.; Metsälä, M.; van de Meerakker, S. Y. T.; Meijer, G. J. Chem. Phys. 2006, 124, 134303.

11. Gilijamse, J. J.; Hoekstra, S.; Meek, S. A.; Metsälä, M.; van de Meerakker, S. Y. T.; Meijer, G. J. Chem. Phys. 2006, 124, 134303.

12. Gilijamse, J. J.; Hoekstra, S.; Meek, S. A.; Metsälä, M.; van de Meerakker, S. Y. T.; Meijer, G. J. Chem. Phys. 2006, 124, 134303.

13. Gilijamse, J. J.; Hoekstra, S.; Meek, S. A.; Metsälä, M.; van de Meerakker, S. Y. T.; Meijer, G. J. Chem. Phys. 2006, 124, 134303.
state is prepared by direct laser excitation from the $X^1\Sigma^+$, $\nu=0$ ground state. After passing through a skimmer, the packet of molecules is slowed down in the Stark decelerator and subsequently loaded and confined in an electrostatic quadrupole trap. When in the (1,0,1) level, about $10^3$ CO molecules are trapped at a density of $10^8$/cm$^3$ and at a temperature of around 20 mK; in the (2,0,2) level an order of magnitude less molecules are trapped at a somewhat higher temperature. A detailed description of the Stark decelerator and of the electrostatic trap that have been used can be found elsewhere.

The ultraviolet phosphorescence back to the electronic ground state that escapes through a hole in one of the trap electrodes is imaged with a lens onto a photomultiplier tube (PMT). In the upper part of Fig. 2 a semilogarithmic plot of the phosphorescence signal is shown as a function of time. The CO molecules are prepared in the (1,0,1) state at $t=0$ ms, and the strong phosphorescence peak around 4–5 ms results from molecules that pass through the trap with (more or less) the initial beam velocity. At 9.3 ms, the decelerated packet of molecules arrives in the trap center and the trap is switched on. After some initial oscillations, caused by the collective motion of the molecules in the trap, an exponentially decaying phosphorescence signal is observed. This signal is shown on a linear scale in panel (b) of Fig. 2. A weighted least-squares single exponential fit to this data in the time-interval from 13 to 38 ms gives a decay time of 2.63±0.03 ms, where the error is an estimate accounting for the systematic effect of the initial oscillations. The phosphorescence of the trapped molecules is measured in the presence of the trapping fields, i.e., in an electric field ranging from 0 to 10 kV/cm. Under these conditions the opposite parity states of the $\Lambda$-doublet are mixed.

For the measurement of the phosphorescence decay of CO molecules in the (2,0,2) level, shown in panel (c) of Fig. 2, the repetition frequency of the molecular beam experiment is reduced from the normally used 10 Hz to 2 Hz to allow for a ≈500 ms observation time. To avoid detection of straylight from the excitation laser, that keeps on running at 10 Hz, no phosphorescence signal is recorded during a 6 ms interval around the times that this laser is fired. A single exponential fit to the data in the 50–450 ms interval yields a decay time of 140±4 ms.

Apart from the phosphorescence to the ground state, optical pumping by blackbody radiation and collisions with background gas can lead to a decay of the signal of trapped molecules; at the present densities, collisions between trapped molecules do not play a role. Optical pumping due to room temperature blackbody radiation is calculated to proceed at a rate of around 0.014 s$^{-1}$ for CO in the $a^2\Pi$, $\nu=0$ state. The contribution of optical pumping to the overall trap loss rate can therefore safely be neglected. The loss rate due to background collisions was measured to be 0.17 s$^{-1}$ for trapped ground-state OH and OD radicals in this apparatus. As these measurements were done under otherwise identical conditions, we assume the same collisional trap loss rate for metastable CO. For CO molecules in the (1,0,1) level the loss due to collisions can then also be neglected; for molecules in the (2,0,2) level collisions slightly contribute to the observed trap loss rate. After correction for this, a radiative lifetime for CO molecules in the (1,0,1) level of 2.63±0.03 ms is found, whereas the radiative lifetime for molecules in the (2,0,2) level is determined as 143±4 ms. The ratio of these lifetimes is 54.4±1.6, which is in good agreement with previous measurements obtained in a metastable CO beam.
agreement with the ratio of 54.66±0.01 that is expected from the ratio of the Ω=1 character in the respective wavefunctions.\textsuperscript{15}

To calculate the radiative lifetime of the (1,0,1) level, we use the model as originally described by James.\textsuperscript{3} Essential in this model are the matrix elements $\langle v_a | h_{A,a}(r) | v_a \rangle$ of the $r$-dependent spin-orbit coupling $h_{A,a}(r)$ and the matrix elements $\langle v_a | \mu_{A,X}(r) | v_b \rangle$ of the $r$-dependent electronic $A \ ^3\Pi \rightarrow X ^1\Sigma^+$ transition dipole moment $\mu_{A,X}(r)$, together with the energy separation of the spin-orbit coupled states. We have computed the wave functions $| v_a \rangle$, $| v_b \rangle$, and $| v_c \rangle$, for the vibrational levels of the $X ^1\Sigma^+$, the $A ^3\Pi$, and the $a ^3\Pi$ states, respectively, with the sinc-function-dependent variable representation (DVR) method.\textsuperscript{16} The diatomic potentials are computed with the Rydberg-Klein-Rees (RKR) program of Ref. 17, taking the required spectroscopic data for the $X$-state from Ref. 18 and for the $A$ and $a$ states from Ref. 19.

James approximated the offdiagonal spin-orbit coupling $h_{A,a}(r)$ as minus the diagonal spin-orbit coupling constant of the $a ^3\Pi$ state, $-h_a(r)$.\textsuperscript{3} In the present work we used the full Breit-Pauli spin-orbit operator and we computed the electronic wave functions at the internally contracted multireference single and double excitation (MRCI) level employing molecular orbitals from a state averaged complete active space multiconfigurational self-consistent field (CASSCF) calculation with the MOLPRO computer program. The active space consisted of the full valence space extended with a $\sigma$, a $\pi_x$, a $\pi_y$, and a $\delta_{\sigma}$ orbital (in all calculations $C_{2\pi}$ point group symmetry was used). The $1s$ orbitals were optimized but not correlated. The one-electron basis consists of the core-valence correlation consistent quintuple-zeta (cc-pCV5Z) basis set augmented with tight $p$ functions with exponents 462.5 and 950.4 for carbon and oxygen, respectively. The exponents were energy-optimized in atomic calculations. The accuracy of such an approach was demonstrated by Nicklass et al.\textsuperscript{20} The transition dipole moment $\mu_{A,X}(r)$ was also computed at the CASSCF+MRCI level. The same extended active space was used, but the orbitals for the $X$ and $A$ states were optimized independently and the cc-pV6Z basis was used. The results are presented in Fig. 3.

Using these results, we computed a radiative lifetime for CO molecules in the (1,0,1) state of 3.16 ms. In order to understand the 20\% discrepancy with the experimental value, we performed an extensive analysis on the quality of the potentials and on that of the calculations of the spin-orbit coupling and transition dipole moment functions.

James arrived at a calculated value of 2.93 ms for the (1,0,1) level.\textsuperscript{3} We repeated the calculation of James by employing RKR potentials based on spectroscopic data available in 1971 and found a value of 2.94 ms. With our new RKR potentials and James’ spin-orbit coupling and transition dipole function we obtain 2.93 ms. We therefore do not find the extreme sensitivity to the potential reported by Sykora and Vidal, and cannot reproduce their value of 3.41 ms.\textsuperscript{10}

We computed the spin-orbit constants of the vibrational levels in the $a ^3\Pi$ state, $\langle v_a | h_a(r) | v_a \rangle$, because these can be compared to experimental values. The computed values for the levels $v_a=0$–8 are all between 1.4\% and 1.5\% lower than the spectroscopic values.\textsuperscript{19} We checked the dependence of the spin-orbit coupling on the one-electron basis set by computing $h_{A,a}$ and $h_a$ at $r=2.3a_0$ with a cc-pCVQZ+p and a cc-pCVTZ+p basis set. This gives couplings that are about 0.5\% and 1.8\% higher, respectively, than the value obtained in the cc-pCVSZ+p basis. We determined the effect of core-valence correlation on the spin-orbit coupling in a calculation with a smaller active space and found that it may increase the couplings by about 0.5\%.

Replacing our $\mu_{A,X}(r)$ with the one computed by Spielfiedel et al.\textsuperscript{21} gives a value of 3.35 ms for the lifetime of the (1,0,1) level. The \textit{ab initio} method employed in Ref. 21 is similar to ours. The difference between their $\mu_{A,X}(r)$ and ours increases at shorter distances as shown in Fig. 3(a). The main reason for the difference is that we optimize the orbitals for the $A$ and the $X$ states separately, whereas they employed orbitals from a state averaged CASSCF calculation.

We compute a total oscillator strength for the $A ^3\Pi$, $v \equiv 12-X ^1\Sigma^+$, $v=0$ transitions of 0.1764. This is 2.4\% lower than the value of 0.1807 obtained from high resolution electron energy loss spectroscopy.\textsuperscript{22} With $\mu_{A,X}$ from Ref. 21 we find a total oscillator strength of 0.1599, i.e., 11.5\% below the experimental value.

Another test of $\mu_{A,X}$ is provided by the radiative lifetime measurements of the $A ^3\Pi$ state by Field et al.\textsuperscript{23} Their deperturbed values for $v=0$–7 are all of the same quality and decrease smoothly from $\tau(v_a=0)=9.9±0.1$ ns to $\tau(v_a=7) = 8.95±0.1$ ns. In our calculations, however, the lifetimes in-
TABLE I. The $n^3\Pi-\chi^1\Sigma^+$ transition dipole moment $\mu_{\alpha,X}$ and the $n^3\Pi-a^3\Pi$ spin-orbit coupling $b_{s\alpha}$ at $r=2.3a_0$, computed at the CASSCF +MRCI level (valence CAS extended with two $\pi_4$ and two $\pi_6$ orbitals) in an aug-cc-pVTZ basis, the energy separation $n^3\Pi-a^3\Pi$ ($E_n-E_{a\alpha}$, with Davidson correction), and the scaled lifetime $\tau_i$ of the (1,0,1) level when the states $1, \ldots, n$ are included.

| $n$ | $E_n-E_{a\alpha}$ (cm$^{-1}$) | $\mu_{\alpha,X}$ (e\(a_0\)) | $b_{s\alpha}$ (cm$^{-1}$) | $\tau_i$ (ms) |
|-----|-------------------------------|-------------------------------|--------------------------|---------------|
| 1   | 16 058                        | 0.542                         | -36.3                    | 3.16          |
| 2   | 48 577                        | -0.513                        | 0.1                      | 3.15          |
| 3   | 54 284                        | 0.369                         | -12.9                    | 2.74          |
| 4   | 67 572                        | 0.016                         | 17.7                     | 2.76          |
| 5   | 70 981                        | 0.000                         | 0.2                      | 2.76          |
| 6   | 76 952                        | 0.303                         | 8.9                      | 2.92          |
| 7   | 81 025                        | -0.105                        | 20.8                     | 2.80          |

These two measurements are mutually consistent and yield an accurate value of the radiative lifetime of the a $^3\Pi_1$, $v=0$, $J=1$ level of 2.63±0.03 ms. Now that the radiative lifetime of this level is known with unprecedented precision, discrepancies with earlier calculated values have become apparent. This prompted us to perform more detailed calculations for this level. When only spin-orbit coupling of the a $^3\Pi_1$ state with the A $^1\Pi_1$ state is included, we compute a lifetime of 3.16 ms. We show that obtaining agreement between theory and experiment will require the calculation of the contribution of higher $^1\Pi$ states.

We acknowledge helpful discussions with A. van der Avoird and R. W. Field. One of the authors (M.M.) is grateful to the Academy of Finland for financial support.

1. C. A. Barth, W. G. Fastie, C. W. Hord, J. B. Pearce, K. K. Kelly, A. I. Stewart, G. E. Thomas, G. P. Anderson, and O. F. Raper, Science 165, 1004 (1969).
2. T. C. James, J. Mol. Spectrosc. 40, 545 (1971).
3. T. C. James, J. Chem. Phys. 55, 4118 (1971).
4. W. L. Borst and E. C. Zipf, Phys. Rev. A 3, 979 (1971).
5. G. M. Lawrence, Chem. Phys. Lett. 9, 575 (1971).
6. T. G. Slanger and G. Black, J. Chem. Phys. 55, 2164 (1971).
7. C. E. Johnson and R. S. van Dyck, Jr., J. Chem. Phys. 56, 1506 (1972).
8. R. T. Jongma, G. Berden, and G. Meijer, J. Chem. Phys. 107, 7034 (1997).
9. T. Sykora and C. R. Vidal, J. Chem. Phys. 110, 6319 (1999).
10. T. Sykora and C. R. Vidal, J. Chem. Phys. 112, 5320 (2000).
11. S. Hoekstra et al., Phys. Rev. Lett. 98, 133001 (2007).
12. S. Y. T. van de Meerakker, N. Vanhaecke, M. P. J. van der Loo, G. C. Groenenboom, and G. Meijer, Phys. Rev. Lett. 95, 013003 (2005).
13. H. L. Bethlem and G. Meijer, Int. Rev. Phys. Chem. 22, 73 (2003).
14. S. Y. T. van de Meerakker, N. Vanhaecke, and G. Meijer, Annu. Rev. Phys. Chem. 57, 159 (2006).
15. A. Wada and H. Kanamori, J. Mol. Spectrosc. 200, 196 (2000).
16. G. C. Groenenboom and D. T. Colbert, J. Chem. Phys. 99, 9681 (1993).
17. R. J. Le Roy, University of Waterloo, Chemical Physics Research Report No. CP-657R, 2004, http://leroy.uwaterloo.ca/programs
18. A. Le Floch, Mol. Phys. 72, 133 (1991).
19. R. W. Field, S. G. Tilford, R. A. Howard, and J. D. Simmons, J. Mol. Spectrosc. 44, 347 (1972).
20. A. Nicklass, K. A. Peterson, A. Berning, H.-J. Werner, and P. J. Knowles, J. Chem. Phys. 112, 5624 (2000).
21. A. Spielfiedel, W.-U. Tchang-Brillet, F. Dayou, and N. Feautrier, Astron. Astrophys. 346, 699 (1999).
22. W. F. Chan, G. Cooper, and C. E. Brion, Chem. Phys. 170, 123 (1993).
23. R. W. Field, O. Benoist d’Azy, M. Lavollée, R. Lopez-Delgado, and A. Tramer, J. Chem. Phys. 78, 2838 (1983).