Hard Single Diffraction in $\bar{p}p$ Collisions at $\sqrt{s} = 630$ and 1800 GeV

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Abstract

Using the DØ detector, we have studied events produced in \( \bar{p}p \) collisions that contain large forward regions with very little energy deposition (“rapidity gaps”) and concurrent jet production at center-of-mass energies of \( \sqrt{s} = 630 \) and 1800 GeV. The fractions of forward and central jet events associated with such rapidity gaps are measured and compared to predictions from Monte Carlo models. For hard diffractive candidate events, we use the calorimeter to extract the fractional momentum loss of the scattered protons.
Inelastic diffractive collisions are responsible for 10–15% of the $pp$ total cross section and have been described by Regge theory through the exchange of a pomeron [1]. Diffractive events are characterized by the absence of significant hadronic particle activity over a large region of rapidity or pseudorapidity ($\eta = -\ln[\tan(\theta/2)]$, where $\theta$ is the polar angle relative to the beam). This empty region is called a rapidity gap and can be used as an experimental signature for diffraction. Recent interest in diffraction has centered on the possible partonic nature of the pomeron in the framework of quantum chromodynamics (QCD), as suggested by Ingelman and Schlein [2]. Hard single diffraction (HSD), which combines diffraction and a hard scatter (such as jet or $W$-boson production), can be used to study the properties of the pomeron.

The partonic nature of the pomeron was first inferred by the UA8 experiment [3] at the CERN $Sp$S collider at $\sqrt{s} = 630$ GeV from studies of diffractive jet events. Recent analyses of diffractive jet production [4] and diffractive $W$-boson production [7] are consistent with a predominantly hard gluonic pomeron, but measured rates at the Fermilab Tevatron are several times lower than predictions based on data from the DESY $ep$ collider HERA [8]. In this Letter we present new measurements of the characteristics of diffractive jet events, and of the fraction of central and forward jet events that contain forward rapidity gaps (“gap fraction”) at center-of-mass energies $\sqrt{s} = 630$ and 1800 GeV. These measurements augment previous results from the CDF collaboration on the gap fraction for forward jets at $\sqrt{s} = 1800$ GeV [4] and place further constraints on diffractive models.

In the DØ detector [9], jets are measured using the uranium/liquid-argon calorimeters with an electromagnetic section extending to $|\eta| < 4.1$ and coverage for hadrons to $|\eta| < 5.2$. Jets are reconstructed using a fixed-cone algorithm with radius $R = \sqrt{\Delta\eta^2 + \Delta\phi^2} = 0.7$ ($\phi$ is the azimuthal angle). The jets are corrected using standard DØ routines for jet-energy scale [10], except that there is no subtraction of energy from spectator parton interactions, since these are unlikely for diffractive events.

To identify rapidity gaps, we measure the number of tiles containing a signal in the LO forward scintillator arrays ($n_{L\text{LO}}$), and towers ($\Delta\eta \times \Delta\phi = 0.1 \times 0.1$) above threshold in the calorimeters ($n_{\text{CAL}}$). The LO arrays provide partial coverage in the region $2.3 < |\eta| < 4.3$. A portion of the two forward calorimeters ($3.0 < |\eta| < 5.2$) is used to measure the calorimeter multiplicity, with a particle tagged by the deposition of more than 150 (500) MeV of energy in an electromagnetic (hadronic) calorimeter tower. The thresholds are set to give negligible noise from uranium decays, while maximizing sensitivity to energetic particles [11].

For $\sqrt{s} = 630$ and 1800 GeV, we use triggers which required at least two jets with transverse energy $E_T > 12$ or 15 GeV (see Table I) to study the dependence of the gap fraction on jet location. The forward jet triggers required the two leading jets to both have $\eta > 1.6$ (or $\eta < -1.6$), while the central jet triggers had an offline requirement of $|\eta| < 1.0$. These data were obtained during special low luminosity runs, with typical instantaneous luminosities much less than $1 \times 10^{30} \text{cm}^{-2}\text{s}^{-1}$. At each $\sqrt{s}$, we also implemented the so-called single veto trigger (SV), a dijet trigger that required a rapidity gap on one side (using the LO detector). The SV trigger was used to obtain large samples of single diffractive candidate events. The events in the final data samples all have a single $\overline{p}p$ interaction requirement, a vertex position within 50 cm of the center of the interaction region, and two leading jets that satisfy standard quality criteria [12]. The number of events in each of the final data samples and the integrated luminosities ($\mathcal{L}$) are given in Table I.
The $n_{LØ}$ versus $n_{CAL}$ distributions for central and forward jet events at $\sqrt{s} = 630$ and 1800 GeV are shown in Fig. 1. For forward jet events, these quantities are defined by the $\eta$ region on the side opposite the two leading jets, while for central jet events they are defined by the forward $\eta$ interval that has the lower multiplicity. The distributions display a peak at zero multiplicity ($n_{CAL} = n_{LØ} = 0$), in qualitative agreement with expectations for a diffractive component in the data.

![Graphs showing multiplicity distributions](image)

FIG. 1. Multiplicity distributions at $\sqrt{s} = 1800$ GeV for (a) forward and (b) central jet events, and at $\sqrt{s} = 630$ GeV for (c) forward and (d) central jet events.

The gap fraction is extracted from a two-dimensional fit to the lego plot of $n_{LØ}$ versus $n_{CAL}$. The non-diffractive (high multiplicity) background is fitted in the signal region using a four-parameter polynomial, and the signal is fitted with a falling exponential, as suggested
by Monte Carlo [1]. Figure 2 shows the multiplicity distribution from Fig. 1(a), and the resulting fitted signal, fitted background, and normalized distribution of pulls ((data-fit)/√N). All distributions have adequate fits, with χ²/dof < 1.2.

Table I shows the gap fractions obtained for the four event samples. The values range from (0.22 ± 0.05)% for central jets at √s = 1800 GeV, to (1.19 ± 0.08)% for forward jets at √s = 630 GeV. Uncertainties are dominated by those on the fit parameters. Additional small uncertainties from the dependence on the range of multiplicity used in the fits were added in quadrature. Potential sources of systematic error, such as the number of fit parameters, jet energy scale, trigger turn-on, tower threshold, luminosity, residual noise, and jet quality, yield only negligible variations in the gap fractions [11].

Table I shows that the gap fractions at √s = 630 GeV are larger than gap fractions at √s = 1800 GeV and that gap fractions for forward jets are larger than for central jets. Table I also lists predicted gap fractions for several possible pomeron structure functions (discussed below).

![FIG. 2. The (a) data from Fig. 1(a), and corresponding (b) fitted signal, (c) fitted background, and (d) normalized pull distributions.](image)

We compare the data to Monte Carlo (MC) simulations using the hard diffractive event generator POMPYT [13], which is based on the non-diffractive PYTHIA [14] program. In POMPYT, a pomeron is emitted from the proton with a certain probability (called the flux factor [2]), and has a structure functions s(β), where β is the fractional momentum of the pomeron carried by the hard parton. We used the standard Donnachie-Lanshoff flux factor [12] in this analysis and compare our data to four structure functions: (i) “hard gluon,” a pomeron consisting of two gluons, s(β) ∝ β(1 − β); (ii) “flat gluon,” s(β) ∝ constant; (iii) “soft gluon,” s(β) ∝ (1 − β)^5; and (iv) “quark,” the two-quark analog of (i). In each case, the gap fraction is defined as the cross section for jet events with a rapidity gap based on POMPYT divided by the jet cross section from PYTHIA Many uncertainties, such as the
choice of proton parton densities, cancel in the ratio. An MC version of the fitting method is applied to correct for diffractive events that fail the gap selection criteria. By applying the appropriate correction factor (which ranges from a few per cent for soft gluon central jets to about 80% for hard gluon forward jets) to each MC prediction and comparing to the data, we make no assumptions about which model (if any) is correct \[16\].

Monte Carlo gap fractions are shown in Table \[\text{I}\]. The systematic uncertainties are typically dominated by the difference in energy scale between data and Monte Carlo, but also include uncertainties from the fitting procedure. We observe that rates for harder gluon structures are far higher than supported by data, while the quark structure is in reasonable agreement with the data. The quark structure, however, has previously been shown to predict an excessive rate of diffractive $W$-Bosons \[9\].

A hard gluonic pomeron is capable of describing previous measurements \[4, 7\], if combined with a flux factor that decreases with increasing $\sqrt{s}$ \[\text{I}\]. The ratios of gap fractions shown in the lower half of Table \[\text{I}\] provide new information, since the flux factor cancels for the same $\sqrt{s}$, and dependence on the flux factor is reduced for different $\sqrt{s}$. The ratios for jets with $|\eta| > 1.6$ to jets with $|\eta| < 1.0$ show clear disagreement between the data and predictions for a hard-gluon pomeron structure, despite this cancellation. A gluon-dominated pomeron containing both soft and hard components, combined with a reduced flux factor, could describe all the data samples.

**TABLE II.** The measured and predicted gap fractions and their ratios.

| Sample  | $|\eta| > 1.6$ | $|\eta| < 1.0$ | $|\eta| > 1.6$ | $|\eta| < 1.0$ |
|---------|---------------|---------------|---------------|---------------|
| 1800 GeV | \(0.65 \pm 0.04\)% | \(2.2 \pm 0.3\)% | \(2.2 \pm 0.3\)% | \(1.4 \pm 0.2\)% |
| 1800 GeV | \(0.22 \pm 0.05\)% | \(2.5 \pm 0.4\)% | \(3.5 \pm 0.5\)% | \(0.05 \pm 0.01\)% |
| 630 GeV | \(1.19 \pm 0.08\)% | \(3.9 \pm 0.9\)% | \(3.1 \pm 0.8\)% | \(1.9 \pm 0.4\)% |
| 630 GeV | \(0.90 \pm 0.06\)% | \(5.2 \pm 0.7\)% | \(6.3 \pm 0.9\)% | \(0.14 \pm 0.04\)% |

Ratios of Gap Fractions

| Sample  | $|\eta| > 1.6$ | $|\eta| < 1.0$ | $|\eta| > 1.6$ | $|\eta| < 1.0$ |
|---------|---------------|---------------|---------------|---------------|
| 630/1800 | 1.8 \pm 0.2 | 1.7 \pm 0.4 | 1.4 \pm 0.3 | 1.4 \pm 0.3 |
| 630/1800 | 4.1 \pm 0.9 | 2.1 \pm 0.4 | 1.8 \pm 0.3 | 3.1 \pm 1.1 |
| 1800 | 3.0 \pm 0.7 | 0.88 \pm 0.18 | 0.64 \pm 0.12 | 30. \pm 8. |
| 630 | 1.3 \pm 0.1 | 0.75 \pm 0.16 | 0.48 \pm 0.12 | 13. \pm 4. |

The characteristics of the HSD events were examined using the high statistics SV trigger. We plot in Fig. 3 the distributions of the number of jets, the $E_T$-weighted rms jet widths, the $\Delta \phi$ between the two leading jets, and the relative ratio of diffractive to non-diffractive events as a function of the average $E_T$ of the two leading jets, for central jets at $\sqrt{s} = 1800$ GeV. The solid lines in Fig. 3(a)–(c) correspond to the distributions for HSD candidate events ($n_{\text{CAL}} = n_{\text{LO}} = 0$), and the dashed lines show the distributions for non-diffractive events ($n_{\text{CAL}} > 0$ and $n_{\text{LO}} > 0$). These plots show that the diffractive events appear to have less overall radiation. Figure 3(d) indicates that there is little dependence of the gap fraction on average jet $E_T$. The MC samples (not shown) have characteristics similar to the data.
Finally, we measure the fractional momentum loss of the proton $\xi$, defined as [8]:

$$\xi \approx \frac{1}{\sqrt{s}} \sum_i E_T e^{\eta_i}$$  \hspace{1cm} (0.1)$$

where the summation is over all observed particles. The outgoing scattered proton or antiproton (and the rapidity gap) is defined to be at positive $\eta$. Equation (1) weights heavily the well-measured central region near the rapidity gap, while particles that escape down the beam pipe at negative $\eta$ give a negligible contribution. Using POMPYT events, where $\xi$ can be determined from the momentum of the scattered proton, we have verified that Eq. (1) is reliable at both values of $\sqrt{s}$ and for different pomeron structures. A scale factor $(2.2 \pm 0.3)$ derived from Monte Carlo is used to convert $\xi$ measured from all particles to that from just electromagnetic calorimetric energy depositions [11]. The $\xi$ distributions for forward and central jets at $\sqrt{s} = 630$ and 1800 GeV are displayed in Fig. 4 with the shaded region showing the variance in the distribution due to energy scale uncertainties. Energy-scale uncertainties result in a shift in $\xi$ such that if the true distribution were below the histogram at small $\xi$, it would be above the histogram at large $\xi$.

The $\xi$ distributions show the expected kinematic behavior of diffraction ($M = \sqrt{\xi s}$, where $M$ is the mass of the diffractive system), peaking at larger $\xi$ for central jets than for forward jets. Forward and central jets at $\sqrt{s} = 630$ GeV also peak at larger $\xi$ values with respect to the corresponding distributions at $\sqrt{s} = 1800$ GeV, since for fixed diffractive mass, smaller $\sqrt{s}$ implies larger $\xi$. Even though pomeron exchange is thought to dominate only
FIG. 4. The $\xi$ distributions for $\sqrt{s} = 1800$ GeV (a) forward and (b) central jets and for $\sqrt{s} = 630$ GeV (c) forward and (d) central jets, using the SV trigger with $n_{\text{CAL}} = n_{\text{LLO}} = 0$. The shaded region shows the variance in the distribution due to energy scale uncertainties (see text).

for $\xi < 0.05$, the trends of the $\xi$ distributions can be reproduced by POMPYT. Without the observation of the scattered proton, the interpretation of these large $\xi$ rapidity gap events is uncertain.

We have measured properties of hard single diffraction at $\sqrt{s} = 630$ and 1800 GeV with jets at forward and central rapidities. The gap fractions have been measured without applying model-dependent corrections. Within the Ingelman-Schlein model, our data can be reasonably described by a pomeron composed dominantly of quarks. For the model to describe our data as well as previous measurements, a reduced flux factor convoluted with a gluonic pomeron containing significant soft and hard components is required. We have also measured the fractional momentum lost by the scattered proton and found it greater than typically expected for pomeron exchange.

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