On the Derivative Expansion for the Electromagnetic Casimir 
Free Energy at High Temperatures

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Abstract

We study the contribution of the thermal zero modes to the Casimir free energy, in the case of a fluctuating electromagnetic (EM) field in the presence of real materials described by frequency-dependent, local and isotropic permittivity ($\epsilon$) and permeability ($\mu$) functions.

Those zero modes, present at any finite temperature, become dominant at high temperatures, since the theory is dimensionally reduced. Our work, within the context of the Derivative Expansion (DE) approach, focusses on the emergence of non analyticities in that dimensionally reduced theory. We conclude that the DE is well defined whenever the function $\Omega(\omega)$, defined by $[\Omega(\omega)]^2 \equiv \omega^2 \epsilon(\omega)$, vanishes in the zero-frequency limit, for at least one of the two material media involved.

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I. INTRODUCTION

Casimir forces are one of the most remarkable macroscopic manifestations of the vacuum fluctuations of the EM field [1]. The precision achieved in recent experiments to measure those forces has encouraged further work to obtain more detailed theoretical predictions for them; in particular, it has become increasingly clear that the geometry and electromagnetic properties of the materials are the two most important aspects that must be taken into account for those predictions to be accurate.

In spite of the intense activity in this field of research, comparison between theory and experiment is not yet entirely satisfying, what leaves some room for debate [2]. Among the most widely used tools to tackle these issues, Lifshitz formula [3] occupies a prominent place. Indeed, on the one hand it can be used as the starting point for computing the Casimir force away from the idealized, perfect conductivity case. On the other, it can also be applied to incorporate non-trivial geometrical effects, at least when the surfaces involved are close to each other and smooth. Indeed, Lifshitz formula, originally meant just for flat and parallel slabs may also be thought of as the leading term, the so called Proximity Force Approximation [4] (PFA), in a Derivative Expansion (DE) of the free energy, regarded as a functional of the shape of the function(s) defining the surfaces involved [5, 6]. In the language of Quantum Field Theory, the DE may be consistently viewed as a low-momentum expansion of the vertex functions, with the PFA playing the role of leading term, involving the summation of an infinite number of vertex functions at zero-momentum. The next to leading order term, in turn, requires the knowledge of vertex functions at the second order in momentum. A similar approach has also been applied in [7] to find an interesting approximation for the Casimir-Polder interaction force, assuming that the surfaces are gently curved.

Regarding the EM properties of the media involved, it is usually sufficient to use a version of Lifshitz formula which describes them by means of an isotropic, frequency dependent permittivity $\varepsilon(\omega)$ and, when magnetic properties are relevant, also its permeability $\mu(\omega)$. Note, however, that even though Lifshitz formula has been derived using a variety of approaches [8], its applicability to dissipative systems is still debated [2, 9, 10].

It has been shown that, in some particular situations, non-analyticities in the expansion of the vertex functions could produce non-local NTLO corrections to the PFA within the
context of the DE. This has been explicitly shown for quantum scalar fields at non-zero temperature in the presence of perfect mirrors described by Neumann boundary conditions) [11]. This also holds true for a real scalar field with Neumann conditions in 2+1 dimensions, albeit it can be shown that the non-analyticity can be cured (in a concrete model) by introducing a small departure from perfect Neumann conditions [12].

In this article, we analyse the emergence of non-analyticities in the DE for real materials which, based on the insight from our previous work [12], should come from contributions due to the dimensionally reduced massless modes, which appear as the zero frequency terms in a Matsubara expansion of the fields. To carry out such an analysis, it is convenient to have a general formula for the DE, corresponding to the free energy for the EM field in the presence of media with realistic properties. This kind of formula has been used (although not explicitly displayed) by other authors in a previous work [6]. There, the Casimir free energy has been evaluated using the scattering matrix approach, based on earlier results about the $S$-matrix elements for the scattering of electromagnetic waves, computed in a perturbative expansion in the departure from the planar surface case [13].

It is our aim in this work to carry out an analysis of the validity of the DE using a different approach, for a closely related physical system: we incorporate the possibility for the media to have non-trivial magnetic properties. On the other hand, our approach to the construction of the DE proceeds along different lines: we use a functional integral approach and the Matsubara formalism to single out the zero mode contributions from the very beginning. We then apply the DE approach, performing an independent perturbative calculation in that dimensionally reduced theory. We find conditions for the DE to be well-defined, and single out models where those conditions are not met.

In the first principles functional approach that we follow, the derivation dwells on a subtle point related to the choice of the temporal gauge for the EM field at a non-zero temperature [14], which allows for a clean isolation of the problematic zero mode, the origin of the “plasma-Drude” controversy in the calculation of the Casimir force at a non-vanishing temperature [2, 9, 15]. Remarkably, it is precisely this zero mode the same which, depending on the case considered, may or not exhibit a non-analytic behavior as a function of the external momenta of the generalized vertices.

The structure of this article is as follows: in Sect. II we present the system that we want to study and the main assumptions made in order to calculate its Casimir free energy,
having in mind its treatment within the context of the DE approach. Then, in Sect. we present explicit results on the zero-mode contributions to the free energy, for media described by non trivial permittivities or permeabilities. In Sect. we study the eventual emergence of non-analyticities in the NTLO correction of the PFA. Finally, in Sect. we present our conclusions. Some technical but nevertheless relevant details of the calculations are presented in Appendix A.

II. THE SYSTEM

A. Geometrical set-up, definitions and conventions

The geometry of the systems that we consider can be determined by specifying just two surfaces, denoted by $L$ and $R$. Our construction will begin with cases where they are the boundaries between two spatial regions, filled by material media.

The surfaces $L$ and $R$ are, respectively, defined by $x_3 = 0$ and $x_3 = \psi(x_1, x_2)$, where $\psi$ is a smooth function of two Cartesian coordinates on an $x_3 = \text{constant}$ plane, for which we will adopt the shorthand notation $\mathbf{x}_\parallel \equiv (x_1, x_2)$.

A real medium will be described by introducing isotropic, spatially local permittivities and permeabilities (in other words, they are defined by scalar functions, since the corresponding tensors are proportional to the identity matrix).

The perfect-conductor case will be obtained as a particular limit.

B. Free energy and partition function

Since we deal with a quantum field at a finite temperature, it is convenient to introduce its free energy $F(\psi)$, a function of the inverse temperature $\beta = T^{-1}$ (in our conventions, Boltzmann’s constant $k_B \equiv 1$) and a real functional, i.e., a real-valued function, of $\psi(\mathbf{x}_\parallel)$. For the sake of notational clarity, the dependence on $\beta$ is not explicitly shown. $F$ may be written in terms of the partition function, $Z(\psi)$, as follows:

$$F(\psi) = -\frac{1}{\beta} \log \left[ \frac{Z(\psi)}{Z_0} \right],$$

where the denominator, $Z_0$, denotes the partition function for the free EM field, i.e., in the absence of media. The effect of that denominator is to subtract the free energy of a free
 Bose gas of photons in the absence of the mirrors, which does not contribute to the force between them. There are other contributions to the free energy which are independent of the distance between the two media, namely, that are invariant under \( \psi(x_\parallel) \rightarrow \psi(x_\parallel) + b \) (\( b \equiv \text{constant} \)). They can, usually in a rather straightforward way, be identified with self-energies of the mirrors. That is, contributions which measure the energy of the distorted vacuum corresponding to the EM field in the presence of just one medium (taking the zero-point energy of the EM field as reference). Since we are ultimately interested in the calculation of the part of the free energy which is responsible for the (normal) force between the two media, those terms will be discarded.

In the Matsubara (imaginary-time) formalism, a functional integral expression for the partition function \( Z(\psi) \) can be constructed by integrating over field configurations depending on the spatial coordinates \( x \) and the imaginary time \( x_0 \equiv \tau \). The fields are periodic, with period \( \beta \), in the imaginary time. Denoting by \( A = (A_\mu) \), \( (\mu = 0,1,2,3) \) the 4-potential in Euclidean spacetime, the form of \( Z(\psi) \) is as follows:

\[
Z(\psi) = \int [DA] e^{-S_{\text{inv}}(A)}
\]

where \( S_{\text{inv}}(A) \) is the gauge-invariant action for \( A \), with gauge transformations given by \( A_\mu(x) \rightarrow A_\mu(x) + \delta_\lambda A_\mu(x) \), \( \delta_\lambda A_\mu(x) = \partial_\mu \lambda(x) \). \( DA \) is the unconstrained functional-integration measure, while \([DA]\) is used to denote that measure after gauge fixing.

In terms of the components of the field strength tensor \( F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu \), the form of the gauge-invariant action is:

\[
S_{\text{inv}}(A) = \int_0^\beta d\tau \int_0^\beta d\tau' \int d^3x \left[ \frac{1}{2} F_{0j}(\tau, x) \epsilon(\tau - \tau', x) F_{0j}(\tau', x) \\
+ \frac{1}{4} F_{ij}(\tau, x) \mu^{-1}(\tau - \tau', x) F_{ij}(\tau', x) \right],
\]

where indices from the middle of the Roman alphabet run over spatial indices (Einstein summation convention has been adopted), and \( \epsilon(\tau - \tau', x) \) and \( \mu(\tau - \tau', x) \) denote the Euclidean versions of the permittivity and permeability, respectively (\( \mu^{-1} \) is the inverse integral kernel of \( \mu \), with respect to its time-like arguments). Space locality of those response functions has been assumed implicitly.

On account of our assumptions on the geometry of the system, we have that:

\[
\epsilon(\tau - \tau', x) = \theta(-x_3) \epsilon_L(\tau - \tau') + \theta(x_3) \theta(\psi(x_\parallel) - x_3) + \theta(x_3 - \psi(x_\parallel)) \epsilon_R(\tau - \tau') \]

\[
\mu(\tau - \tau', x) = \theta(-x_3) \mu_L(\tau - \tau') + \theta(x_3) \theta(\psi(x_\parallel) - x_3) + \theta(x_3 - \psi(x_\parallel)) \mu_R(\tau - \tau'),
\]
where $\epsilon_{L,R}(\tau - \tau')$ and $\mu_{L,R}(\tau - \tau')$ characterize the permittivity and permeability of the respective mirror.

Note that the free (vacuum) form of the action is

$$S_{\text{inv}}^{(\text{vac})}(A) = \frac{1}{4} \int d^4x \, F_{\mu\nu}F_{\mu\nu},$$

where $d^4x \equiv d\tau d\mathbf{x}$, the integral over $\tau$ goes from 0 to $\beta$ (the fields being periodic, any interval of length $\beta$ can be used as the extent of the imaginary time coordinate to define the action).

C. Matsubara modes and gauge fixing

The action is invariant under translations in the imaginary time: $\tau \to \tau + \text{constant}$. This suggests the use of mixed Fourier transformations for the fields, as well as for the response functions:

$$A_\mu(\tau, \mathbf{x}) = \frac{1}{\beta} \sum_{n=-\infty}^{+\infty} \tilde{A}_\mu^{(n)}(\mathbf{x}) e^{i\omega_n \tau}$$

$$\epsilon(\tau - \tau', \mathbf{x}) = \frac{1}{\beta} \sum_{n=-\infty}^{+\infty} \tilde{\epsilon}^{(n)}(\mathbf{x}) e^{i\omega_n (\tau - \tau')}$$

$$\mu(\tau - \tau', \mathbf{x}) = \frac{1}{\beta} \sum_{n=-\infty}^{+\infty} \tilde{\mu}^{(n)}(\mathbf{x}) e^{i\omega_n (\tau - \tau')}$$

where $\omega_n \equiv \frac{2\pi n}{\beta}$ ($n \in \mathbb{Z}$) are the Matsubara frequencies. The gauge transformations can be represented in Fourier space; to that end, note that, since the field $A_\mu(\tau, \mathbf{x})$ is periodic, so must be $\delta_\lambda A_\mu(\tau, \mathbf{x})$. Then we can expand this object in terms of a Fourier series,

$$\delta_\lambda A_\mu(\tau, \mathbf{x}) = \frac{1}{\beta} \sum_{n=-\infty}^{+\infty} c_{\mu}^{(n)}(\mathbf{x}) e^{i\omega_n \tau}.$$  

Note, however, that $\lambda(\tau, \mathbf{x})$ needs not to be periodic. It may be seen that its more general form is as follows:

$$\lambda(\tau, \mathbf{x}) = C \tau + D + \lambda_P(\tau, \mathbf{x})$$

where $C$ and $D$ are constants, and $\lambda_P(\tau, \mathbf{x})$ is periodic:

$$\lambda_P(\tau, \mathbf{x}) = \frac{1}{\beta} \sum_{n=-\infty}^{+\infty} \tilde{\lambda}_P^{(n)}(\mathbf{x}) e^{i\omega_n \tau}.$$
Hence, the full set of gauge transformations for the gauge field may be written, using a Fourier transform for the time arguments, as follows:

\[
\delta_{\lambda} \tilde{A}^{(n)}(x) = \partial_j \tilde{A}^{(n)}_\mu(x)
\]

\[
\delta_{\lambda} \tilde{A}^{(n)}_0(x) = \begin{cases} 
   i\omega_n \tilde{A}^{(n)}_\mu(x) & \text{if } n \neq 0 \\
   C & \text{if } n = 0
\end{cases}.
\] (10)

From the explicit form above for the gauge transformations in the same ‘mixed’ Fourier representation, we see that one possible gauge fixing condition, which we will adopt throughout this article, is the temporal gauge, which in this finite-temperature setup is given by:

\[
\tilde{A}^{(n)}_0(x) = 0, \quad \forall n \neq 0.
\] (11)

We then write the thermal partition function using the mixed Fourier representation. Using the properties: \( \tilde{A}^{(n)*}_\mu = \tilde{A}^{(-n)}_\mu \), \( \epsilon^{(-n)} = \epsilon^{(n)} \) and \( \bar{\mu}^{(-n)} = \bar{\mu}^{(n)} \), it is rather straightforward to see that it may be written as an infinite product of (decoupled) integrals, involving the zero mode, and just the positive (or, alternatively, negative) modes:

\[
Z(\psi) = \prod_{n=0}^{\infty} Z^{(n)}(\psi)
\] (12)

where

\[
Z^{(0)}(\psi) = \int [D\tilde{A}^{(0)}_0 D\tilde{A}^{(0)}_j] e^{-S^{(0)}(\tilde{A}^{(0)}_0, \tilde{A}^{(0)}_j)}
\] (13)

and, for \( n \geq 1 \),

\[
Z^{(n)}(\psi) = \int [D\tilde{A}^{(n)}_j D\tilde{A}^{(n)*}_j] e^{-S^{(n)}(\tilde{A}^{(n)}_j, \tilde{A}^{(n)*}_j)}
\] (14)

where \( S^{(0)} \) involves two real fields: one of them is the zero-frequency component of \( \tilde{A}_0 \), which behaves as a scalar in the Euclidean 2 + 1 dimensional space. The other, corresponding to the zero-mode for \( \tilde{A}_j \), is a real vector field:

\[
S^{(0)}(\tilde{A}^{(0)}_0, \tilde{A}^{(0)}_j) = \frac{1}{\beta} \int d^3 x \left[ \frac{1}{2} \tilde{\epsilon}^{(0)}(x)(\partial_j \tilde{A}^{(0)}_0)^2 + \frac{1}{4 \bar{\mu}^{(0)}(x)}(\tilde{F}^{(0)}_{jk})^2 + \frac{1}{2} \Omega_0^2(x)(\tilde{A}^{(0)}_j)^2 \right]
\] (15)

where we have introduced:

\[
\Omega_0^2(x) \equiv \lim_{n \to 0} \left[ \omega_n^2 \tilde{\epsilon}^{(n)}(x) \right].
\] (16)

Note that \( \Omega_0 \) vanishes for a dielectric and also for a metal described by the Drude model. On the other hand, it equals the plasma frequency for a metal described by the plasma model.
Since, by assumption, there is vacuum between the two mirrors, $\tilde{\epsilon}^{(0)} = \tilde{\mu}^{(0)} = 1$, and $\Omega_0^2 = 0$ in that region. Thus the scalar and vector fields behave there as a free massless scalar and a free gauge field, respectively. In each one of the regions occupied by a mirror, the scalar field Lagrangian is multiplied by a constant (in a wave-function renormalization fashion), while the vector field also has a constant Proca-like mass (each constant is determined by the properties of the medium on the mirror considered). Besides the latter also has a factor similar to the scalar field one, but determined by the permeability.

As a consequence of the temporal gauge choice, there is no scalar field mode in the $S_{n>0}$ terms. They always involve just a complex vector field:

$$S^{(n)}(\tilde{A}_j^{(n)}, \tilde{A}_j^{(n)*}) = \frac{1}{\beta} \int d^3x \left( \frac{1}{2} |\tilde{F}_{jk}^{(n)}|^2 + \Omega_n^2(x) |\tilde{A}_j^{(0)}|^2 \right),$$

(17)

with a Proca-like mass $\Omega_n^2(x) = \tilde{\epsilon}^{(n)}(x) \omega_n^2$ which, contrary to what happened for the zero mode, is non-vanishing between the mirrors. Note that, except for the $n = 0$ mode, each term corresponds to combining two different Matsubara modes, namely, $n$ and $-n$ into the action for a complex field. This is possible because of the reality of the permittivity in Fourier space, which is in turn a reflection of the parity in its time argument.

In any case, the free energy will be obtained as a sum of infinite terms, each one corresponding to a given value of the index $n$,

$$F(\psi) = \sum_{n=0}^{\infty} F^{(n)}(\psi),$$

(18)

where

$$F^{(n)}(\psi) = -\frac{1}{\beta} \log \left[ \frac{Z^{(n)}(\psi)}{Z_0^{(n)}} \right].$$

(19)

In the special case of $n = 0$, we may split the free energy according to its origin being the scalar ($s$) or vector ($v$) fields:

$$F^{(0)} = F_s^{(0)}(\psi) + F_v^{(0)}(\psi),$$

(20)

while for $n \geq 1$, we only have contributions originated in a vector field $F^{(n)} = F_v^{(n)}$, $\forall n \geq 1$.

III. RESULTS FOR THE ZERO MODE FREE ENERGIES

This contribution to the free energy is composed of two independent terms, $F^{(0)} = F_s^{(0)} + F_v^{(0)}$, each one of them can be obtained from the calculation of a path integral over an
(unconstrained) real field, namely:

\[
e^{-\beta F_s(\varphi)} = \int \mathcal{D}\tilde{A}_0^{(0)} e^{-\frac{1}{\beta} \int d^3x \tilde{A}_0^{(0)}(x)(\partial_j \tilde{A}_0^{(0)})^2}
\]  

(21) and

\[
e^{-\beta F_v^{(0)}(\varphi)} = \int \mathcal{D}\tilde{A}_0^{(0)} e^{-\frac{1}{\beta} \int d^3x [\frac{1}{4\tilde{\mu}_0^{(0)}(x)}(\tilde{F}_0^{(0)})^2 + \frac{1}{2} \Omega_0^2(\varphi)(\tilde{A}_0^{(0)})^2]}
\]  

(22)

The functions \(\tilde{\epsilon}^{(0)}\), \(\tilde{\mu}^{(0)}\) and \(\Omega_0\) are model-dependent. Nevertheless, since by assumption we have a vacuum between the plates, we can write:

\[
\tilde{\epsilon}^{(0)}(x) = \epsilon_L \theta(-x_3) + \theta(x_3) \theta(\varphi(x) - x_3) + \epsilon_R \theta(x_3 - \varphi(x))
\]  

(23) and

\[
\tilde{\mu}^{(0)}(x) = \mu_L \theta(-x_3) + \theta(x_3) \theta(\varphi(x) - x_3) + \mu_R \theta(x_3 - \varphi(x))
\]  

(24)

and

\[
\Omega_0^2(x) = \Omega_L^2 \theta(-x_3) + \Omega_R^2 \theta(x_3 - \varphi(x))
\]  

(25)

where we have introduced the constants:

\[
\epsilon_{L,R} = \lim_{n \to 0} \epsilon_{L,R}^{(n)}
\]

\[
\Omega_{L,R}^2 = \lim_{n \to 0} [\omega_{L,R}^2(\tilde{\epsilon}_L^{(n)}, \tilde{\mu}_R^{(n)})]
\]  

(26)

which are entirely determined by the parameters of the model used for the permittivity.

We take advantage of the fact that the two contributions above are decoupled, to consider and present the corresponding results separately below. Before doing this, it is instructive to discuss the limit of perfectly conducting materials for the scalar and vector contributions. The \(\tilde{A}_0^{(0)}\) field behaves as a scalar, and when the infinite permittivity limit is taken, its gradient inside the region occupied by the mirror vanishes. Thus the field is constant in that region; assuming that the conductors are grounded, that constant vanishes, so that the field itself is zero. That is, this field becomes a Dirichlet mode, corresponding to the transverse magnetic (TM) EM mode.

The vector zero mode, on the other hand, behaves as an EM field in 2 + 1 dimensions. If \(\Omega_L\), say, tends to infinity, then the that EM field will vanish identically on the region occupied by that mirror. It then has perfect conductor boundary conditions at \(x_3 = 0\). (Of course, the same will happen on the other surface if the corresponding constant tends to infinity.) But we have shown this to be equivalent to a real scalar field with Neumann conditions \([12]\). Therefore, this is a Neumann mode, corresponding to the transverse electric (TE) EM mode.
A. The scalar zero-mode contribution $F_s(\psi)$

We first note that a formal result of the functional integral for this mode can be written in terms of the determinant of an operator $K_s$:

$$e^{-\beta F_s(\psi)} = [\det K_s]^{-\frac{1}{2}}$$

$$K_s = -\partial_j \tilde{\epsilon}(0)(x) \partial_j,$$  \hspace{1cm} (27)

where we have neglected a global constant which does not contribute to the interaction energy between the mirrors.

Within the DE approach to the second order, $F_s$ is given by an expression with the form:

$$[F_s(\psi)]_{DE} = \int d^2x_\parallel \left[ V_s(\psi) + Z_s(\psi)\partial_a \psi(x) \partial_a \psi(x) \right].$$  \hspace{1cm} (28)

The functions $V_s$ and $Z_s$ may be determined from the first and third terms in the expansion of $F_s(a + \eta)$ in powers of $\eta$: $F_s(a + \eta) = F_s,0(a) + F_s,1(a, \eta) + F_s,2(a, \eta) + \ldots$, as described in our previous works [5].

1. Zeroth order term

The first term, which contains no derivatives, allows us to construct the corresponding ‘potential’ in the DE,

$$V_s(a) = \frac{1}{L^2} F_s,0(a),$$  \hspace{1cm} (29)

where $L^2$ denotes the area of each surface ($L \to \infty$). From (27), this implies

$$V_s(a) = \frac{1}{2\beta} \int \frac{d^2k_\parallel}{(2\pi)^2} \log \det [K_s|\psi(x_\parallel=a)] .$$  \hspace{1cm} (30)

This contribution can be evaluated exactly, taking advantage of the fact that it corresponds to a system where $\tilde{\epsilon}(0)(x)$ depends only on $x_3$, by using, for example, the Gelfand-Yaglom theorem approach [16]. Since the calculation is already a standard one, we omit the details. The result is:

$$V_s(a) = \frac{1}{2\beta} \int \frac{d^2k_\parallel}{(2\pi)^2} \log \left(1 - r_{L,R} e^{-2|k_1|a}\right) ,$$  \hspace{1cm} (31)

where

$$r_{L,R} \equiv \frac{1 - \epsilon_{L,R}}{1 + \epsilon_{L,R}} .$$  \hspace{1cm} (32)

As expected, in the limit $\epsilon_{L,R} \to \infty$ one obtains the result corresponding to a scalar field in $2 + 1$ dimensions satisfying Dirichlet boundary conditions times $1/\beta$ [11].
2. First order term

The first order term, \( F_{s,1} \), is not required in order to determine the DE; however, we calculate it in order to have a consistency check for some of the ingredients we use in our work. We first note that \( F_s(\psi) \) can be conveniently written in an equivalent form where the zeroth order term is extracted explicitly,

\[
e^{-\beta F_s} = e^{-\beta F_{s,0}} e^{-\beta F_{s,I}}
\]  

where \( F_{s,I} \) contains all the perturbative corrections:

\[
F_{s,I} = -\frac{1}{\beta} \log \langle e^{-S_{s,I}} \rangle .
\]

The average symbol is given by a functional integral with the action corresponding to \( \eta \equiv 0 \):

\[
\langle \ldots \rangle = \frac{\int \mathcal{D}\tilde{A}^{(0)}_0 \ldots e^{-S_{s,0}(\tilde{A}^{(0)}_0)}}{\int \mathcal{D}\tilde{A}^{(0)}_0 e^{-S_{s,0}(\tilde{A}^{(0)}_0)}} .
\]

Here,

\[
S_s = S_{s,0} + S_{s,I}
\]

and

\[
S_{s,0} \equiv S_s \big|_{\psi(x_3) = a} .
\]

Then, the first order term becomes:

\[
F_{s,1} = \frac{1}{\beta} \langle S_{s,1} \rangle ,
\]

where the first-order term in the action, \( S_{s,1} \), is given by

\[
S_{s,1} = \frac{1 - \epsilon R}{2\beta} \int d^3x \delta(x_3 - a) |\nabla A^{(0)}_0(x)|^2 .
\]

Thus,

\[
F_{s,1} = \frac{1 - \epsilon R}{2\beta^2} \int d^3x \delta(x_3 - a) \Delta_{jl}(x_\parallel, x_\parallel)
\]

where we have introduced the objects:

\[
\Delta_{jl}(x_\parallel, y_\parallel) \equiv \Delta_{jl}(x_\parallel - y_\parallel) = \langle \partial_j \tilde{A}^{(0)}_0(x) \partial_l \tilde{A}^{(0)}_0(y) \rangle_{x_3,y_3 \to a} .
\]

On the other hand, the correlation function which coincidence limit appears above, is ill-defined when one of the indices \( j, l \) equals 3. Indeed, the derivatives of the field have a
discontinuity whenever the permittivity has a jump. We then introduce in $S_{s,1}$, the point split action, before evaluating the average:

$$S_{s,1} = \frac{1 - \epsilon_R}{2\beta} \int d^3x \partial_j \tilde{A}_0^{(0)}(x_a + \eta) \partial_j \tilde{A}_0^{(0)}(x_a - \eta).$$

And obtain the result for the free energy by taking the $\eta \to 0$ limit after evaluating the average.

Thus, we see that the first-order term is given by:

$$F_{s,1} = \frac{1 - \epsilon_R}{2\beta^2} \int d^2x || \Delta^{+-}_{aa}(x_a, x_a) + \Delta_{33}^{+-}(x_a, x_a) \| \eta(x_a),$$

where

$$\Delta^{+-}_{jl}(x_a - y_b) = \left[ \left. (\partial_j \tilde{A}_0^{(0)}(x) \partial_l \tilde{A}_0^{(0)}(y) \right) \right]_{x_3 \to a, y_3 \to a} - \int \frac{d^2k||}{(2\pi)^2} \tilde{\Delta}^{+-}_{jl}(k) e^{ik|| \cdot (x_a - y_b)}.$$ 

We note that the $\pm$ in $\Delta_{ab}$ ($a \neq 3, b \neq 3$) may be omitted, since that object is well defined, continuous at $x_3 = y_3 = a$.

A rather straightforward calculation shows that, when both indices are different from 3:

$$\tilde{\Delta}_{ab}(k||) = \beta \tilde{\Delta}(k||) k_a k_b,$$

with:

$$\tilde{\Delta}(k||) = \frac{1}{2||k||} \left\{ \frac{1}{\epsilon_R} + \frac{1}{1 - r_L r_R e^{-2||k||} a} \left\{ r_L \left( \frac{2}{1 + \epsilon_R} + \frac{r_R}{\epsilon_R} e^{-2||k||} a - \frac{r_R}{\epsilon_R} \right) \right\} \right\}. $$

On the other hand,

$$\tilde{\Delta}_{33}^{+-}(k||) = \tilde{\Delta}_{33}^{-+}(k||) = -\beta \epsilon_R ||k|| \left[ \frac{1}{\epsilon_R} - \tilde{\Delta}(k||) ||k|| \right] $$

with the same $\tilde{\Delta}$ as in the previous equation.

Thus, at the first order:

$$F_{s,1} = \frac{1 - \epsilon_R}{2\beta^2} \int d^2x || \eta(x_a) \left( \int \frac{d^2k||}{(2\pi)^2} \left[ \tilde{\Delta}_{aa}(k||) + \tilde{\Delta}_{33}^{+-}(k||) \right] \right) \right.$$ 

Discarding $a$-independent terms, we find that the free energy per unit area is:

$$F_{s,1} = \frac{F_{s,1}}{L^2} = \eta_0 \frac{r_L r_R}{\beta} \int \frac{d^2k||}{(2\pi)^2} \frac{||k||}{e^{2||k||} a - r_L r_R},$$

with $\eta_0 \equiv \int d^2x || \eta(x_a) ||$. The consistency check is completed by noting that this term agrees with $\eta_0 \frac{\partial}{\partial a} V_s(a)$. 

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3. Second order term

The second order term is obtained by collecting the second-order terms in the expansion in powers of \( \eta \). Discarding self-energy like contributions, and with the same notations introduced above, we see that it is, in principle, given by:

\[
F_{s,2} = -\frac{1}{4\beta^3} (1 - \epsilon_R)^2 \int d^2x_\parallel \int d^2y_\parallel \left[ (\partial_j \widetilde{A}_0^{(0)}(x) \partial_k \widetilde{A}_0^{(0)}(y)) \right]_{x_3 = y_3 = 0}^2 \eta(x_\parallel) \eta(y_\parallel). \tag{50}
\]

The expression is, however, ill-defined. Using the point splitting procedure, in the same form as in the first-order calculation, to make sense of the ill-defined vertices, we obtain:

\[
F_{s,2} = -\frac{1}{8\beta^3} (1 - \epsilon_R)^2 \int d^2x_\parallel \int d^2y_\parallel \left[ 2\Delta_{ab}(x_\parallel - y_\parallel) \Delta_{ab}(x_\parallel - y_\parallel) \\
+ \Delta_{33}^{+}(x_\parallel - y_\parallel) \Delta_{33}^{--}(x_\parallel - y_\parallel) + \Delta_{33}^{++}(x_\parallel - y_\parallel) \Delta_{33}^{--}(x_\parallel - y_\parallel) \\
+ \Delta_{3a}^{++}(x_\parallel - y_\parallel) \Delta_{3a}^{--}(x_\parallel - y_\parallel) + \Delta_{3a}^{--}(x_\parallel - y_\parallel) \Delta_{3a}^{--}(x_\parallel - y_\parallel) \right] \eta(x_\parallel) \eta(y_\parallel). \tag{51}
\]

This term, being quadratic in \( \eta \), can be represented in (parallel) Fourier space in terms of a kernel \( f_s^{(2)} \):

\[
F_s^{(2)} = \frac{1}{2} \int \frac{d^2k_\parallel}{(2\pi)^2} f_s^{(2)}(k_\parallel, a) |\tilde{\eta}(k_\parallel)|^2,
\]

where \( k_\parallel = |k_\parallel| \) The function \( Z_s(\psi) \) in Eq. (28) can be obtained from

\[
Z_s(\psi) = \frac{1}{2} \frac{\partial f_s^{(2)}}{\partial k_\parallel^2}(0, \psi). \tag{53}
\]

Note that, as \( \epsilon_{R,L} \) are dimensionless, dimensional analysis implies that \( Z_s \) is \( 1/(\beta \psi^4) \) times a function of \( \epsilon_{R,L} \).

The explicit form for \( f_s^{(2)} \) becomes:

\[
f_s^{(2)}(k_\parallel, a) = -\frac{1}{2\beta} (1 - \epsilon_R)^2 \int \frac{d^2p_\parallel}{(2\pi)^2} \left\{ \Delta(p_\parallel) \tilde{\Delta}(p_\parallel + k_\parallel) \left[ (p_\parallel \cdot (p_\parallel + k_\parallel))^2 \\
- 2\epsilon_R |p_\parallel| |p_\parallel + k_\parallel| (p_\parallel \cdot (p_\parallel + k_\parallel)) \right] \\
+ \epsilon_R^2 \left( \frac{1}{\epsilon_R} |p_\parallel| + \tilde{\Delta}(p_\parallel) |p_\parallel|^2 \right) \left( \frac{1}{\epsilon_R} |p_\parallel + k_\parallel| + \tilde{\Delta}(p_\parallel + k_\parallel) |p_\parallel + k_\parallel|^2 \right) \right\}. \tag{54}
\]

Subtracting from \( f_s^{(2)} \) its value at \( a \to \infty \) we obtain, in the particular case \( \epsilon_R = \epsilon_L = \epsilon \)

\[
f_s^{(2)}(k_\parallel, a) = -\frac{1}{2\beta} (1 - \epsilon)^2 \int \frac{d^2p_\parallel}{(2\pi)^2} \left\{ \frac{1}{|p_\parallel| |p_\parallel + k_\parallel|} \left( \frac{A_-(p_\parallel)}{|A_-(p_\parallel)|} A_-(p_\parallel + k_\parallel) - \frac{1}{(1 + \epsilon)^2} \right) \\
\times \left[ (p_\parallel \cdot (p_\parallel + k_\parallel))^2 - 2\epsilon |p_\parallel||p_\parallel + k_\parallel| (p_\parallel \cdot (p_\parallel + k_\parallel)) \right] \\
+ |p_\parallel||p_\parallel + k_\parallel| (\epsilon^2 A_+(p_\parallel) A_+(p_\parallel + k_\parallel) - \frac{1}{(1 + \epsilon)^2}) \right\}, \tag{55}
\]

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where

\[ A_{\pm}(p) = \frac{1}{2\epsilon} \left( 1 \pm \frac{r(1 - e^{-2|p||a}}{1 - r^2 e^{-2|p||a}} \right). \]  

(56)

In the limit \( \epsilon \to \infty \), and omitting a term independent of \( k || a \), it reduces to

\[ f_s^{(2)}(k || a) = -\frac{2}{\beta} \int \frac{d^2p ||}{(2\pi)^2} \frac{|p || |p || + k ||}{(1 - e^{-2|p||a})(e^{2|p||a} + k || a - 1)}. \]  

(57)

As expected, this result coincides with \( 1/\beta \) times the one obtained for a Dirichlet scalar field in 2 + 1 dimensions [11].

It is interesting to remark that Eq. (55) can be obtained with a rather different approach based on the scattering formula for the Casimir free energy [9], using standard perturbation theory for the analysis of the incidence of classical electromagnetic waves on rough surfaces of small slope [13] (see Appendix A). The derivation presented here sheds light on the gauge fixing procedure in the functional integral: the temporal gauge does not fix the zero-frequency component of \( \tilde{A}_0 \), giving rise to the scalar contribution to the Casimir free energy.

B. The vector zero-mode contribution \( F_v(\psi) \)

Again, within the DE approach to the second order, we will have for \( F_s \) an expression with the form:

\[ [F_v(\psi)]_{DE} = \int d^2x || V_v(\psi) + Z_v(\psi) \partial_a \psi(\mathbf{x}) \partial_a \psi(\mathbf{x})]. \]  

(58)

The functions \( V_v \) and \( Z_v \) may be determined from the first and third terms in the expansion of \( F_v(a + \eta) \) in powers of \( \eta \): \( F_v(a + \eta) = F_v,0(a) + F_v,1(a, \eta) + F_v,2(a, \eta) + \ldots \).

In what follows we will consider separately the case of a magnetic materials with a permittivity such that \( \Omega_0(\mathbf{x}) \) vanishes identically (for example, if the permittivity has a regular zero-frequency limit) and the case of a non-magnetic materials with \( \tilde{\mu}^{(0)} = 1 \) and \( \Omega_0(\mathbf{x}) \neq 0 \).

1. The case \( \tilde{\mu}^{(0)} \neq 1 \) and \( \Omega_0 = 0 \).

In this situation, the Casimir free energy is given by:

\[ e^{-\beta F_v^{(0)}(\psi)} = \int \mathcal{D}A_j^{(0)} e^{-\frac{1}{4\beta} \int d^2x \frac{1}{\tilde{\mu}^{(0)}(\mathbf{x})} \langle \tilde{F}_j^{(0)} \rangle^2}. \]  

(59)
In order to evaluate the functional integral, we apply a result we obtained in a previous paper \[12\], whereby we have shown that the calculation of the effective action for a gauge field in the presence of imperfect mirrors could be mapped to a scalar field model. Indeed, using the duality

$$\partial_j \phi \leftrightarrow \epsilon_{ijk} \partial_i A_k,$$

the free energy can be written in terms of the scalar field as

$$e^{-\beta F_0^{(0)}(\phi)} = \int \mathcal{D} \phi e^{-\frac{1}{2\beta} \int d^4x \frac{1}{\mu^{(0)}(x)} (\partial_j \phi)^2}.$$ 

(60)

Therefore, the result for the previous integral can be borrowed from the ones we have already obtained for the scalar zero mode, by extending the results to $0 < \epsilon < 1$ and replacing $\epsilon_{L,R} \rightarrow \frac{1}{\mu_{L,R}}$.

One can readily check that in the limit $\mu_{L,R} \rightarrow \infty$ one recovers the TE mode of the perfect conductor case. Indeed, taking the corresponding limit $\epsilon_{L,R} \rightarrow 0$ in Eq.(31) and (55) one obtains

$$f^{(2)}_v(k_{||},a) = -\frac{2}{\beta} \int \frac{d^2p_{||}}{(2\pi)^2} \log \left(1 - \rho L \rho R e^{-2|k_{||}|^a}\right),$$

(62)

which is $1/\beta$ times the result for a Neumann scalar field in $2 + 1$ dimensions.

2. The case $\bar{\mu}^{(0)} = 1$ and $\Omega_0 \neq 0$.

The ‘potential’ term in the DE can again be evaluated exactly, since it corresponds to a system where $\bar{\epsilon}^{(0)}(x)$ depends only on $x_3$. An application of the Gelfand-Yaglom theorem approach yields:

$$V_v(a) = \frac{1}{2\beta} \int \frac{d^2k_{||}}{(2\pi)^2} \log \left(1 - \rho L \rho R e^{-2|k_{||}|^a}\right),$$

(63)

where we have introduced:

$$\rho_{L,R}(k_{||}) = \frac{\Omega_{L,R}^2}{\left(|k_{||}| + \sqrt{|k_{||}|^2 + \Omega_{L,R}^2}\right)^2} = \frac{\sqrt{|k_{||}|^2 + \Omega_{L,R}^2} - |k_{||}|}{\sqrt{|k_{||}|^2 + \Omega_{L,R}^2} + |k_{||}|}.$$  

(64)

Having already illustrated the computation of the Casimir free energy using the functional approach, and in order to avoid a rather lengthy calculation, for the second order term we
will use the known results based on the scattering matrix approach. As shown in Appendix A, the result for the second order kernel is, in the particular case $\Omega_L = \Omega_R = \Omega$:

$$f_s^{(2)}(k_\parallel) = -\frac{2}{\beta} \int \frac{d^2 p_\parallel}{(2\pi)^2} \frac{|\rho^2(p_\parallel)|}{g(p_\parallel)} e^{-2|p_\parallel|a} \left\{ \sqrt{\Omega^2 + |p_\parallel|^2} - \sqrt{\Omega^2 + |p_\parallel + k_\parallel|^2} + \frac{(p_\parallel \cdot (p_\parallel + k_\parallel))^2}{|p_\parallel|^2 |p_\parallel + k_\parallel| g(p_\parallel + k_\parallel)} \right\}, \quad (65)$$

where we introduced the notation

$$g(k_\parallel) = 1 - \rho^2(k_\parallel) e^{-2|k_\parallel|^a}. \quad (66)$$

The perfect conductor limit is obtained as $\Omega \to \infty$. In this limit, the difference of square roots in Eq.(65) vanishes, the reflection coefficient $\rho \to 1$, and the final result corresponds to a Neumann scalar field in $2 + 1$ dimensions (see Eq.(62)).

**IV. ANALYTICITY OF THE DERIVATIVE EXPANSION**

In this section we analyze the structure of the NTLO correction to the PFA. As repeatedly emphasized in our previous works [5, 11, 12], the DE is based on the low momentum expansion of the kernels that appear in the perturbative evaluation of the Casimir free energy, when the shape of the interface between different media is $\psi(x_\parallel) = a + \eta(x_\parallel)$ with $a \ll \eta$ (in the present work, these kernels are $f_s^{(2)}$ and $f_\psi^{(2)}$).

If the kernels admit an expansion in powers of $k_\parallel^2$, the NTLO correction to the PFA can be written locally in terms of derivatives of $\psi$. In the opposite case, when the low momentum behavior is non-analytic in $k_\parallel^2$, the NTLO correction is non-local. For perfect conductors, we have shown that at finite temperature the TE contribution contains a non-analytic contribution [11]. The origin of the non-analyticity is in the zero-frequency mode, which behaves as a scalar field with Neumann boundary conditions. Imperfect boundary conditions may restore the analyticity of the kernels, at least in the toy model considered in Ref. [12]. A question that naturally suggests itself is whether the departure from perfect conditions will always restore the analyticity. To that end, the question we address here is whether the boundary conditions for the EM field in the presence of real materials at finite temperature make the kernels analytic or not. In other words, if the resulting NTLO correction to PFA, for realistic conditions at finite temperature, is spatially local or not.
In order to analyze the low momentum behavior of \(f_s^{(2)}\) and \(f_v^{(2)}\), we can proceed as follows: we first expand the integrand defining both kernels (see Eqs. (55) and (65)) in powers of \(k^2_{\parallel}\), and then look for eventual infrared divergences in the term proportional to the first power of \(k^2_{\parallel}\), which, when it is well defined, constitutes the NTLO correction to PFA in the DE approach.

The expansion of \(|p_{\parallel} + k_{\parallel}|\) in powers of \(k_{\parallel}\) generates inverse powers of \(p_{\parallel}\): assuming that \(k_{\parallel}\) points in the \(x\)-direction we have

\[
|p_{\parallel} + k_{\parallel}| = \sqrt{|p_{\parallel}|^2 + |k_{\parallel}|^2 + 2|p_{\parallel}||k_{\parallel}| \cos \theta \simeq |p_{\parallel}| + \cos \theta |k_{\parallel}| + \frac{1}{2|p_{\parallel}|} \sin^2 \theta |k_{\parallel}|^2 + \ldots , \tag{67}
\]

so the term proportional to \(k^2_{\parallel}\) is inversely proportional to \(p_{\parallel}\). This is one source of potential infrared divergences. Additional inverse powers of \(p_{\parallel}\) may appear when the reflection coefficients tend to one (see Eq. (66)).

For the scalar kernel (Eq. (55)), the functions \(A_{\pm}\) and their derivatives contain at most one inverse power of \(p_{\parallel}\). Therefore, by power counting one can check that for any value \(0 < \epsilon < \infty\) there are no infrared divergences and the NTLO correction to PFA is local (the consideration of values \(0 < \epsilon < 1\) is useful, since to analyze the case of magnetic materials, we apply the duality \(\mu \equiv \frac{1}{\epsilon}\)). In the limit \(\epsilon \to \infty\) one obtains the Dirichlet kernel, that does not contain infrared divergences. In the opposite limit \(\epsilon \to 0\) the result corresponds to the Neumann kernel, which has an infrared divergence when expanded in powers of \(k_{\parallel}\).

Therefore, we conclude that the scalar contribution generates a well defined DE, and that magnetic materials regulate the non-analyticity of the TE mode when considering very large (but not infinite) values of \(\mu\).

We have confirmed these results with numerical evaluations. In Fig. 1 we plot \(Z_s\), defined in Eq. (53), as a function of \(\epsilon\). \(Z_s = 0\) for \(\epsilon = 1\) and tends to the 2+1 Dirichlet value

\[
Z_s^D = -\frac{\Gamma(3/2)[1 + 6\zeta(3)]}{12(4\pi)^{3/2}} \simeq -0.0136 \tag{68}
\]

for large values of \(\epsilon\).

In Fig. 2 we plot \(Z_v\) in the case of magnetic materials as a function of the permeability \(\mu\). As described in Section III B 1, \(Z_v\) formally coincides with \(Z_s\) after the replacement \(\epsilon \to 1/\mu\), so the numerical evaluation is similar to the previous one. \(Z_v\) vanishes as \(\mu \to 1\) and diverges for \(\mu \gg 1\), as expected for a Neumann mode in 2+1 dimensions.

Let us now consider the vector zero-mode contribution in Eq. (65). The term proportional to \(\sqrt{\Omega^2 + |p_{\parallel} + k_{\parallel}|^2}\) does not generate inverse powers of \(p_{\parallel}\) when expanded in powers of \(k_{\parallel}\)
FIG. 1. Numerical evaluation of $Z_s(a)$ (in units of $\beta a^4$) given in Eq.(53) as a function of $\epsilon$. $Z_s$ vanishes when $\epsilon \to 1$ and tends to the value for a $2 + 1$ Dirichlet mode (-0.0136) when $\epsilon \to \infty$.

FIG. 2. Numerical evaluation of $Z_v(a)$ (in units of $\beta a^4$) for the magnetic case as a function of $\mu$. $Z_v = 0$ for $\mu = 1$ and diverges when $\mu \gg 1$, as expected for a 2+1 Neumann mode.

(the plasma frequency acts as an infrared regulator). To analyze the last term, it is useful to consider the identity

$$
\frac{(p_\parallel \cdot (p_\parallel + k_\parallel))^2}{|p_\parallel||p_\parallel + k_\parallel|} = |p_\parallel||p_\parallel + k_\parallel| - \frac{|p_\parallel||\sin^2 \theta|}{|p_\parallel + k_\parallel|}|k_\parallel|^2.
$$

(69)

The second term in the above equation generates an infrared divergence in the vector kernel. Indeed, when inserting this identity in Eq.(65), the term proportional to $\sin^2 \theta$ reads, up to order $|k_\parallel|^2$,

$$
-\frac{2|k_\parallel|^2}{\beta} \int \frac{d^2p_\parallel}{(2\pi)^2} \frac{\rho^2(p_\parallel)}{g^2(p_\parallel)} e^{-2|p_\parallel|\alpha \sin^2 \theta}.
$$

(70)
For small values of $p_\parallel$ we have $\rho^2 \simeq 1$ and $g \simeq p_\parallel$, and therefore the integral above is logarithmically divergent in the infrared. This result shows that the NTLO correction to PFA for the zero-frequency vector mode is non-local as long as $\Omega$ is different from zero. Unlike for magnetic materials, a finite value of $\Omega$ does not regulate the infrared divergence. This can be traced back to the fact that the reflection coefficient $\rho^2 \to 1$ as $p_\parallel \to 0$, for any value of $\Omega$.

V. CONCLUSIONS

In this paper we have analysed in detail the zero-frequency contribution to the Casimir free energy in the presence of real materials. We used a functional approach, and clarified the use of the temporal gauge in the context of the Casimir effect, since it does not fix the zero frequency mode of $\tilde{A}_0$. This mode does in fact generate a TM contribution to the free energy.

We have also computed the kernels that are necessary to obtain the DE of the Casimir free energy, in order to discuss the validity of the DE for real materials. We have shown that the TM contribution does always produce a NTLO correction to the PFA, which is local in derivatives of the function $\psi$ that defines the shape of the curved interface. The same happens for the TE contribution in the case of magnetic materials. There is only one situation where the DE would fail, that is, when $\omega^2 \epsilon(\omega) \to \Omega^2 \neq 0$ as $\omega \to 0$ for both mirrors \[17\]. In terms of the models usually considered in the Casimir literature to describe real materials, this condition corresponds to the plasma model.

In summary, the non-analyticities we observed for perfect conductors in our previous work \[11\], survive only under the assumption of perfectly lossless materials. Related with this, in recent works \[10, 18\] it has been claimed that the Lifshitz-Matsubara formula does not apply for the plasma model, and that it should be understood as the lossless limit of the Drude model (note however that this claim has been contested in Ref. \[19\]). If this were the case, then $\Omega^2$ would vanish, and there would be no vector (or TE) contribution to the zero mode. Consequently, the NTLO correction to PFA would be local even in this lossless limit.
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Appendix A: The scattering approach

In this Appendix we outline the calculation of \( f^{(2)} = f_s^{(2)} + f_v^{(2)} \) using the scattering approach. In Ref.[6] it is shown that

\[
f^{(2)}(k_\parallel, a) = -\frac{1}{\beta} \int \frac{d^2p_\parallel}{(2\pi)^2} \sum_Q \frac{|p_\parallel r_Q(p_\parallel)|}{g_Q(p_\parallel)} e^{-2|p_\parallel|a} [(B_2)_{QQ}(p_\parallel, p_\parallel, p_\parallel + k_\parallel) \nonumber \\
+ 2 \sum_{Q'} \frac{|p_\parallel + k_\parallel r_{Q'}(p_\parallel + k_\parallel)|}{g_{Q'}(p_\parallel + k_\parallel)} e^{-2|p_\parallel + k_\parallel|a} B_{QQ'}(p_\parallel, p_\parallel + k_\parallel) B_{QQ'}(p_\parallel + k_\parallel, p_\parallel)] , \tag{A1}
\]

where the indices \( Q, Q' \) denote the two polarizations, \( r_Q \) are the Fresnel reflection coefficients and

\[
g_Q(p_\parallel) = 1 - r_Q^2 e^{-2|p_\parallel|a} . \tag{A2}
\]

We are assuming here that both media have the same EM properties.

The crucial ingredients in the above formula are the coefficients \( B_{QQ'} \) and \((B_2)_{QQ}\), which describe the scattering of a classical EM wave at the rough interface between two homogeneous half-spaces. These coefficients have been computed in Ref.[13] using small-slope perturbation theory. The notation there is slightly different: \( B_{QQ'} \rightarrow B_{22}^{\alpha\alpha_0} \) and \((B_2)_{QQ} \rightarrow i(B_2)_{22}^{\alpha\alpha_0} \). The situation where both media have the same permittivity \( \epsilon \) is obtained by setting \( \epsilon_1 = \epsilon \) and \( \epsilon_2 = 1 \) in the results of Ref.[13].

As the general formula for \( f^{(2)} \) has not been explicitly displayed in previous works, and is needed for the analysis of the emergence of non-analyticities of the DE, we will work it out here. The coefficients \( B_{QQ'} \) are given in Eq.(4.15) of Ref.[13]. One can check that, as we are considering the zero-frequency mode, they vanish for \( Q \neq Q' \). Therefore, \( f^{(2)} \) in Eq.(A1) has two separate contributions: \( Q = Q' = 1 \) and \( Q = Q' = 2 \), which correspond to the scalar \((f_s^{(2)})\) and vector \((f_v^{(2)})\) contributions, respectively. The coefficients \((B_2)_{QQ}\) are given in the Appendix D of Ref.[13].
1. Scalar contribution

The Fresnel reflection coefficient reads in this case

\[ r_1 = \frac{\epsilon - 1}{\epsilon + 1}, \tag{A3} \]

and the scattering coefficients

\[ B_{11}(p_\parallel, p'_\parallel) = -\frac{(\epsilon - 1)}{(\epsilon + 1)^2} \left[ \epsilon + \frac{p_\parallel \cdot p'_\parallel}{|p_\parallel||p'_\parallel|} \right] \tag{A4} \]

\[ (B_2)_{11}(p_\parallel, p_\parallel, p'_\parallel) = \frac{2(\epsilon - 1)}{(\epsilon + 1)^2|p_\parallel|^2} \left[ -\frac{(\epsilon - 1)}{(\epsilon + 1)|p'_\parallel|^2} (|p_\parallel|^2|p'_\parallel|^2 - (p_\parallel \cdot p'_\parallel)^2) + \frac{4\epsilon |p_\parallel|}{(\epsilon + 1)} p_\parallel \cdot p'_\parallel \right]. \]

Inserting Eq. (A4) into Eq. (A1) one obtains the scalar contribution to \( f^{(2)} \). After a rather long calculation one can show that the result coincides with Eq. (55).

2. Vector contribution

For \( Q = 2 \), the Fresnel reflection coefficient reads

\[ r_2 = \frac{|k_\parallel| - \sqrt{|k_\parallel|^2 + \Omega^2}}{|k_\parallel| + \sqrt{|k_\parallel|^2 + \Omega^2}}. \tag{A5} \]

Note that it coincides with the coefficient \(-\rho\) defined in Eq. (64), in the particular case \( \Omega_L = \Omega_R = \Omega \).

On the other hand we have

\[ B_{22}(p_\parallel, p'_\parallel) = \frac{\Omega^2}{(\sqrt{\Omega^2 + |p_\parallel|^2 + |p'_\parallel|^2})(\sqrt{\Omega^2 + |p_\parallel|^2 + |p'_\parallel|^2})} \frac{p_\parallel \cdot p'_\parallel}{|p_\parallel||p'_\parallel|} \]

\[ (B_2)_{22}(p_\parallel, p_\parallel, p'_\parallel) = -2\rho(p_\parallel) \left[ \left( \frac{p_\parallel \cdot p'_\parallel}{|p_\parallel|^2 |p'_\parallel|^2} + \sqrt{\Omega^2 + |p_\parallel|^2} - \sqrt{\Omega^2 + |p'_\parallel|^2} \right) \right]. \tag{A6} \]

Inserting Eq. (A6) into Eq. (A1) one obtains the vector contribution to \( f^{(2)} \), which is given in Eq. (65).

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