Lepton-flavour violation
in the light of leptogenesis and muon $g - 2$

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**Abstract**

A discrepancy between the experimental and theoretical results was reported for the anomalous magnetic moment of the muon. In leptogenesis scenarios which are due to a decay of right-handed neutrinos, this anomaly leads to large violations of lepton flavour within the framework of supersymmetric see-saw models. It is shown that for a hierarchical right-handed neutrino mass spectrum, we generically expect to observe $\mu \rightarrow e\gamma$ in the near future experiment.
I. INTRODUCTION

The experiments have measured the neutrino oscillations and established a mass of neutrinos. This cannot be explained in the Standard Model (SM) framework and thus is an evidence of physics beyond SM. From the cosmological side, one of the most challenging mysteries of the universe is the source of the asymmetry of the baryon number density. Since it is quite difficult to explain the asymmetry in SM, we are required to introduce new mechanism to generate it.

Recently, the anomalous magnetic moment \((g - 2)\) of the muon suggested a contribution of the new physics. The E821 experiment at Brookhaven measured it at the extremely precise level\([1]\), and the results have been compared with the SM prediction. According to the latest update of the hadronic contribution to the SM value\([2]\), which is based on the \(e^+ e^-\) collision data, a difference was reported between these two results as (see \([3]\) and references therein for details)

\[
a_\mu (\text{exp.}) - a_\mu (\text{SM}) = 302(88) \times 10^{-11},
\]

which means \(3.4\sigma\) deviation. One may expect that this discrepancy is explained by hadronic uncertainties in the SM prediction. However, too large corrections are then required to fix \([\text{1}]\) in the hadron sector\([3]\). Instead of pursuing this idea, we consider new contributions from physics beyond SM.

These features are easily achieved by the supersymmetric (SUSY) see-saw scenarios. In the SUSY models, all the SM particles are accompanied by their superpartners, and these new particles affect low-scale phenomena through radiative corrections. In fact, the discrepancy of the muon \(g - 2\) can be saturated due to a \(\tan \beta\) enhancement\([4]\). On the other hand, the neutrino oscillations are realized by the Type I see-saw mechanism, introducing the right-handed sector\([5]\). Furthermore, against the mystery of the baryon asymmetry, the see-saw mechanism provides an elegant solution. By producing the right-handed neutrinos at an early stage of the universe, the lepton asymmetry would be generated at the decay of those neutrinos, and then leads to the baryon asymmetry of Universe (BAU) via the sphaleron effect.

As a prediction of the SUSY see-saw scenarios, the lepton flavours are violated at the weak scale. These violations are tightly correlated with the leptogenesis and muon \(g - 2\).
The measurements of the neutrino oscillations indicate flavour-violations in the neutrino sector. Such violations are naturally transmitted into the charged-lepton sector via the renormalization group evolutions. At the weak scale, the lepton-flavour violating processes are induced by superparticles. Those superparticles contribute as well to the prediction of the muon $g-2$. Thus, they cannot be arbitrary heavy in order to explain the muon $g-2$ anomaly \(^1\). On the other hand, the right-handed neutrino mass scale is required to be large for the leptogenesis to work successfully. Such a large scale enhances the neutrino Yukawa coupling including the flavour-violating effects. Consequently, in the leptogenesis scenarios, it is considered that the lepton-flavour violating processes tend to be sizable in the light of the anomaly of the muon $g-2$. In this letter, we will show that for a hierarchical right-handed neutrino mass spectrum, $\mu \rightarrow e\gamma$ is generically expected to be observed in the near future experiment.

II. CORRELATIONS AMONG OBSERVABLES

We first briefly review the muon $g-2$. The SUSY contributions consist of the two diagrams; those mediated by the charginos and neutralinos. Considering the coupling strength, the former usually dominates the SUSY contributions. The result is approximately obtained as \(^4\)

$$\delta a_{\mu} \simeq \frac{5\alpha_2}{48\pi} \frac{m_{\mu}^2}{m_{\text{SUSY}}^2} \text{sign}(M_2\mu_H) \tan \beta,$$

(2)

where $M_2$ is a wino mass, $\mu_H$ is a Higgsino mass and $\tan \beta = \langle H_u \rangle / \langle H_d \rangle$ is a ratio of two vacuum expectation values of Higgses. In this expression, we have set all soft masses to be equal for simplicity. In the letter, we choose a sign of $(M_2\mu_H)$ to be positive to explain \(^1\), which is irrelevant for the following two observables.

The second observable is $\text{Br}(\mu \rightarrow e\gamma)$. The SM contribution is known to be highly suppressed, and the SUSY effects determines the branching ratio. The SUSY diagrams are the same as those of the muon $g-2$ except for the flavour dependence. Let us consider that the soft breaking slepton masses are flavour universal at the cutoff scale, $\mu_X$, which is assumed to be larger than the right-handed neutrino mass scale, $\mu_R$. Even with this flavour-universal condition, in SUSY see-saw models, neutrino Yukawa couplings induce the left-handed slepton mixing through the renormalization running from $\mu_X$ down to $\mu_R$. Thus
we expect significant SUSY contributions to Br($\mu \rightarrow e\gamma$). With the universal slepton mass boundary condition, the branching ratio is represented as

$$\text{Br}(\mu \rightarrow e\gamma) \sim \frac{\alpha^3}{G_F^2} \frac{|(Y_N^\dagger Y_N)_{12} \ln (\mu_R/\mu_X)|^2}{m_{\text{SUSY}}^2} \tan^2 \beta ,$$

(3)

up to a numerical factor, where $Y_N$ is a neutrino Yukawa coupling matrix in the basis of diagonal charge lepton Yukawa couplings and right-handed neutrino mass matrix. Here, all the soft SUSY breaking parameters including those during the renormalization running were simply set to be equal. Comparing with Eq. (2), a naive relation between $\delta a_\mu$ and Br($\mu \rightarrow e\gamma$) is found as $\text{Br}(\mu \rightarrow e\gamma) \propto (\delta a_\mu)^2$ for fixed $Y_N$.

It should be mentioned that we showed Eqs. (2) and (3) to clarify the dependence on the model parameters, while for the following numerical analysis, we do not rely on these simplifications. Namely, we will evaluate $\delta a_\mu$ and Br($\mu \rightarrow e\gamma$) completely up to the one-loop level after solving a set of renormalization group equations and diagonalizing the mass matrix in the next section.

The production of the right-handed neutrino leads to the generation of the lepton asymmetry. In leptogenesis scenarios such as the thermal and non-thermal ones, the lepton asymmetry is proportional to CP asymmetry in a decay of a lightest right-handed neutrino, $N_1$, into lepton doublet, $L$, and Higgs, $H_u$. The CP asymmetry parameter $\epsilon_1$ is (see [8] for a review)

$$\epsilon_1 \equiv \frac{\Gamma(N_1 \rightarrow L + H_u) - \Gamma(N_1 \rightarrow L^c + H_u^c)}{\Gamma(N_1 \rightarrow L + H_u) + \Gamma(N_1 \rightarrow L^c + H_u^c)} = \frac{1}{8\pi (Y_N Y_N^\dagger)_{11}} \sum_{i \neq 1} \text{Im} \left[ (Y_N Y_N^\dagger)_{i1}^2 f(M_i^2/M_1^2) \right] ,$$

(4)

where $M_i$’s are masses of the right-handed neutrinos and a loop function $f(x)$ is given as

$$f(x) = \sqrt{x} \ln \left( 1 + \frac{1}{x} \right) - \frac{2\sqrt{x}}{x - 1} .$$

(5)

Here and in the following analysis, we ignore flavour-dependent effects on the leptogenesis because they are generically negligible. The above expression is valid for a hierarchical right-handed neutrino mass spectrum, i.e., $|M_{2,3} - M_1| \gg \Gamma_{2,3} + \Gamma_1$, where $\Gamma_i$ is the decay width of the $i$th right-handed neutrino and estimated at the tree level as

$$\Gamma_i = \frac{(Y_N Y_N^\dagger)_{ii}}{8\pi} M_i .$$

(6)
When $M_2$ becomes very close to $M_1$, the resonant effects contribute to the lepton asymmetry. This topic will be discussed in the final section. In the following analysis, we will focus on the hierarchical (off-resonant) case, $M_1 \ll M_2 \lesssim M_3$.

The neutrino Yukawa coupling matrix and the right-handed neutrino mass matrix are connected to the light neutrino mass matrix, $m_\nu$ through the see-saw relation,

$$(Y_N)_{ki} \frac{1}{M_k} (Y_N)_{kj} \langle H_u \rangle^2 = (m_\nu)_{ij} = U_{ik} m_k U_{jk}^* ,$$

where $m_i$ is mass eigenvalues of the light neutrinos and $U$ is Pontecorvo–Maki–Nakagawa–Sakata (PMNS) matrix\cite{11}. From data of neutrino oscillation experiments, parameters in $U$ and the squared mass differences are determined as\cite{12}

$$|\Delta m^2_{\text{atm}}| \equiv |m_3^2 - m_1^2| \sim 2.5 \times 10^{-3} \text{eV}^2 , \quad \sin^2 2\theta_{\text{atm}} = 4|U_{23}^2|(1 - |U_{23}^2|) \sim 1.0 ,$$

$$\Delta m^2_{\odot} \equiv m_2^2 - m_1^2 \sim 8.0 \times 10^{-5} \text{eV}^2 , \quad \tan^2 \theta_{\odot} \simeq \frac{|U_{12}^2|}{|U_{11}^2|} \sim 0.4 ,$$

$$|U_{13}| \lesssim 0.2 .$$

In order to incorporate the relation Eq. (7) in the analysis, it is useful to use the following parametrization\cite{13},

$$(Y_N)_{ij} = \frac{1}{\langle H_u \rangle} \sqrt{M_i} R_{ik} \sqrt{m_k} U_{jk}^* ,$$

with $R_{ij}$ satisfying $\sum_k R_{ik} R_{jk} = \delta_{ij}$. $R_{ij}$ has six real parameters.

In a lot of leptogenesis scenarios, the lepton asymmetry is favored to be as large as possible. With the above parametrization, it is easily shown that a size of the CP asymmetry parameter $\epsilon_1$ has an upper bound which is much less than one. With $M_1/M_2,3 \ll 1$ and

$$\max(|R_{ij}|) < \mathcal{O}(M_{2,3}/M_1)$$

(see\cite{14}), one gets\cite{14}

$$|\epsilon_1| \simeq \frac{1}{8\pi \langle Y_N Y_N \rangle_{11}} \left| \sum_{i \neq 1} \text{Im} \left( (Y_N Y_N^\dagger)_{1i}^2 \frac{M_i}{M_1} \right) \right| = \frac{M_1}{8\pi \langle H_u \rangle^2} \left| \sum_i \text{Im}(m_i^2 R_{i1}) \right| \sim \frac{M_1}{8\pi \langle H_u \rangle^2 m_1 + m_3} \Delta m^2_{\atm} .$$

One can find that $|\epsilon_1|$ is maximized when $R_{12} = 0$ and $|\text{Re}(R_{13})| = |\text{Im}(R_{13})|$ are satisfied. For a given maximal $\epsilon_1 \equiv \epsilon_1^{\text{max}}$, the right-handed neutrino mass is estimated as

$$M_1 \simeq 1.5 \times 10^{10} \text{GeV} \left( \frac{\epsilon_1^{\text{max}}}{10^{-6}} \right) \left( \frac{\langle H_u \rangle}{174 \text{GeV}} \right)^2 \left( \frac{\Delta m^2_{\atm}}{2.5 \times 10^{-3} \text{eV}^2} \right)^{-1} \left( \frac{m_1 + m_3}{0.05 \text{eV}} \right) .$$

We expect that a large asymmetry of the lepton number density leads to a sizable branching ratio of $\mu \to e\gamma$\cite{15, 16}. From Eq. (10), $|\epsilon_1|$ is proportional to $M_1$, and thus the CP
asymmetry, i.e., the lepton asymmetry, is enhanced for heavier $M_1$. Then, by satisfying the relation $M_1 \ll M_2 \lesssim M_3$, heavier $M_1$ leads to larger elements of $Y_N$. As was shown in Eq. (3), large $Y_N$ enhances $\text{Br}(\mu \rightarrow e\gamma)$. Thus, for leptogenesis scenarios to successfully work, $\text{Br}(\mu \rightarrow e\gamma)$ tends to be large.

One may consider that the SUSY contributions to the $\mu \rightarrow e\gamma$ amplitude are suppressed by heavy superparticles. However, the anomaly of the muon $g - 2$, Eq. (1), prohibits the particles to be decoupled as long as it is explained by the SUSY contributions. Especially since the branching ratio of $\mu \rightarrow e\gamma$, Eq. (3), is tightly correlated with the SUSY contribution to the muon $g - 2$, Eq. (2), it is very hard to suppress $\text{Br}(\mu \rightarrow e\gamma)$ with keeping both $\delta a_\mu$ and $\epsilon_1$ large.

We should mention specific cases in which $\text{Br}(\mu \rightarrow e\gamma)$ is suppressed. First of all, the flavour structure of the Yukawa coupling potentially causes a cancellation in the decay amplitude. Actually, the amplitude is proportional to $(Y_N^\dagger Y_N)_{12} = (Y_N)_{11}^* (Y_N)_{12} + (Y_N)_{21}^* (Y_N)_{22} + (Y_N)_{31}^* (Y_N)_{32}$. Thus, an accidental cancellation may happen among these complex numbers with $\delta a_\mu$ and $\epsilon_1$ fixed. We might also obtain cancellations between the chargino and neutralino contributions, or by taking into account the initial flavour-changing components of the soft SUSY breaking parameters. These cases will be commented in the final section. In any case, the cancellations are considered to be accidental in general and thus regarded as a fine-tuning. As the second case, let us consider texture structures of the right-handed Yukawa coupling. Assigning texture zeros properly for the Yukawa coupling, $\text{Br}(\mu \rightarrow e\gamma)$ can be suppressed[16]. Then, instead of $\mu \rightarrow e\gamma$, it is likely to observe other lepton flavour violating processes, e.g., $\tau \rightarrow \mu(e)\gamma$. Note that, since the light-neutrino mass spectrum is predicted to be specific in this framework, it is expected to identify such a case in future. In the following study, we will discuss a lower bound of $\text{Br}(\mu \rightarrow e\gamma)$ in generic conditions. Namely, we will assume no fine-tunings in the neutrino sector, and not include the specific texture setup.
Lower bounds of the predicted $\text{Br}(\mu \to e\gamma)$ are displayed in Fig. 1 as a function of $\epsilon_1$ for fixed $\delta a_\mu$. Here, we scanned the SUSY parameters and the neutrino Yukawa couplings satisfying the experimental data. In the analysis, we set $\mu_X$ to be the grand-unification (GUT) scale and imposed the following assumptions; neglecting resonant effects of the leptogenesis and generic structure for the lepton sector. In particular, as for the latter assumption, we assumed no fine-tunings in the right-handed neutrino Yukawa coupling and sought the parameter points which minimize $|(Y_N)_{11}^*(Y_N)_{12}| + |(Y_N)_{21}^*(Y_N)_{22}| + |(Y_N)_{31}^*(Y_N)_{32}|$ for fixed $\epsilon_1$. This means that $|(Y_N^\dagger Y_N)_{12}|$ is minimized without an accidental cancellation among three different terms in the summation on the lines.

It can be said that the lower bounds in Fig. 1 are conservative under the above assumptions. We can check that on the lines, $m_1 = 0$ and $U_{13} = 0$ are satisfied, and $|\epsilon_1|$ is maximized by satisfying $R_{12} = 0$ and $\text{arg}(R_{13}) = \pi/4$. When we increase the lightest neutrino mass, according to Eq. (10) the maximal value of $\epsilon_1$ is suppressed by $\Delta m_{\text{atm}}^2/(m_1 + m_3)$. Namely, with $\epsilon_1$ fixed, $M_1$ increases for larger $m_1$, and thus, the lower bound of $\text{Br}(\mu \to e\gamma)$ goes up. In addition, $\text{Br}(\mu \to e\gamma)$ is minimized when both $m_1 = 0$ and $U_{13} = 0$ are satisfied. This is because $(Y_N)_{11}$ becomes naturally small. From (9), if either $m_1$ or $U_{13}$ is finite, the branching ratio of $\mu \to e\gamma$ receives an additional contribution from $(Y_N)_{11}^*(Y_N)_{12}$. Also, the CP violation phases in the PMNS matrix are taken to be zero, otherwise $\text{Br}(\mu \to e\gamma)$ becomes larger. Although we assumed that the slepton mass matrix is universal at the cutoff scale, introducing the off-diagonal components is just additive to the branching ratio. In conclusion, the lower bounds in Fig. 1 are conservative.

It is stressed that the lower bound of $\text{Br}(\mu \to e\gamma)$ depends on $M_2/M_1$ but is independent of $M_3$. It is because on the lines of the lower bound, $R_{ij}$ still has enough degrees of freedom to give the same minimal value of $\text{Br}(\mu \to e\gamma)$ for different values of $M_3/M_2$. In other words, the minimum of $\text{Br}(\mu \to e\gamma)$ has a flat direction in the parameter space. It is obtained that as long as $M_2/M_1$ is larger than 10, the lower bound is proportional to $M_2/M_1$. While for smaller $M_2/M_1$, the lower bound of $\text{Br}(\mu \to e\gamma)$ is not a linear function of $M_2/M_1$, because contributions from the next leading order in loop function $f(M_2^2/M_1^2)$ which is a term of

\footnote{In the analysis, we set the soft mass of the selectron to be degenerate with that of the smuon. When the selectron is much heavier than the smuon, $\text{Br}(\mu \to e\gamma)$ can be suppressed for fixed $\delta a_\mu$.}
order $M_1^2/M_2^2$ are non-negligible. On the other hand, the lower bound is almost independent of $\tan \beta$. This is simply because $\text{Br}(\mu \to e\gamma)$ is proportional to $(\delta a_\mu)^2$. Thus, the branching ratio remains the same for fixed $\delta a_\mu$.

From Fig. 11 in order to obtain $\delta a_\mu > 2.1 \times 10^{-9}$ (1σ), the leptogenesis scenarios with $\epsilon_1 > 10^{-5}$ provide the generic lower bound of $\text{Br}(\mu \to e\gamma) > 10^{-13}$ for the hierarchy $M_2/M_1 > 10$ with $M_1 \gtrsim 1.5 \times 10^{11}\text{GeV}$. This is the sensitivity at which we expect to observe $\mu \to e\gamma$ in near future such as in the MEG experiment [18]. For smaller $\epsilon_1 = 10^{-6}$ with $M_2/M_1 = 10$, there is the parameter region where we will not detect $\mu \to e\gamma$, while for larger $M_2/M_1$ such as $\approx 100$, the lower bound exceeds the experimental sensitivity with $\delta a_\mu > 2.1 \times 10^{-9}$. On the other hand, for a smaller hierarchy case, $M_2/M_1 = 3$, the lower bound of $\text{Br}(\mu \to e\gamma)$ reaches $O(10^{-13})$ when $\epsilon_1$ is larger than $10^{-4}$, which corresponds to $M_1 \gtrsim 10^{12}\text{GeV}$. As a result, since a lot of leptogenesis models practically require large CP asymmetry, we expect to observe the lepton-flavour violating muon decay in near future in the light of the muon $g - 2$ anomaly when $M_2/M_1$ is hierarchical.

IV. CONCLUSIONS AND DISCUSSION

The reported discrepancy of the muon $g - 2$ has impacts on low-energy phenomena. In this letter, we focused on the lepton flavour violation. We showed that in the light of this anomaly, successful leptogenesis scenarios prefer a sizable branching ratio of the lepton-flavour violating muon decay. Since the muon $g - 2$ anomaly favors superparticles to stay in rather low-energy regime, e.g., $\lesssim 1\text{TeV}$, the SUSY contributions to the muon $g - 2$ can be checked directly by probing those particles in the forthcoming Large Hadron Collider. When the SUSY contributions will be confirmed in the experiment, neglecting resonant leptogenesis and setting $\mu_X > \mu_R$, we found that leptogenesis scenarios generically predict $\text{Br}(\mu \to e\gamma)$ to become larger than $O(10^{-13})$ in wide parameter regions. Thus, we expect to detect $\mu \to e\gamma$ in the near future experiment.

In the analysis, we have focused on the CP asymmetry parameter, $\epsilon_1$, to discuss leptogenesis. The lepton asymmetry depends on the thermal history of the universe as well, particularly on the production channels of the right-handed neutrinos. As a natural channel, they are produced in the thermal bath effectively when the reheating temperature exceeds the right-handed neutrino scale. In this thermal leptogenesis scenarios, the lepton asymme-
FIG. 1: Generic lower bound on the prediction of Br(μ → eγ) for fixed δa_μ and ε_1. Here, μ_X is set to be the GUT scale. In the analysis, we assumed no fine-tunings and did not take specific texture structures so that we avoid accidental suppression of Br(μ → eγ). We also neglect the resonant effects for the leptogenesis, which become significant when M_2 is very close to M_1. In the graphs, the ratio M_2/M_1 is taken as (a) M_2/M_1 = 3, (b) = 10, (c) = 30, and = 100 with M_1 varied. The horizontal line at Br(μ → eγ) = 10^{-13} represents the sensitivity of the near future experiment.

In the absence of the initial right-handed neutrinos, κ can be as large as O(10^{-1}), and thus ε_1 ≳ 10^{-6} is needed to explain the present measurements of BAU. This maximal κ is achieved when a washout mass parameter, \tilde{m}_1 = (Y_N)_{1i}(M_i)^{-1}(Y^*_N)^{1i}\langle H_u \rangle^2, is tuned to be about 0.001eV. We checked that this condition can be realized on the lines of the lower bound of Br(μ → eγ). On the other hand, since the efficiency factor decreases very quickly...
when \( \tilde{m}_1 \) differs from the maximal efficiency value, the efficiency factor easily takes smaller value. In practice, we need \( \epsilon_1 \gtrsim 2 \times 10^{-6} \)\(^8\), and then the lower bound of \( \text{Br}(\mu \to e\gamma) \) is obtained to be larger than \( O(10^{-13}) \) for \( M_2/M_1 > 30 \) with \( \delta a_\mu > 2.1 \times 10^{-9} \)\(^b\). In this parameter region, we expect to measure the flavour-changing decay of the muon in near future. In another scenario, the right-handed neutrino may be produced non-thermally. In this case, the resultant lepton asymmetry strongly depends on physics in high-energy scale. Even in this case, the lepton asymmetry is proportional to the CP asymmetry parameter, \( \epsilon_1 \), and we obtain the lower bound of \( \text{Br}(\mu \to e\gamma) \).

If the lepton-flavour violations will not be observed in the future experiments, the absence of signals does not always exclude the leptogenesis scenarios as a source of BAU. In this letter, we imposed several conditions to obtain the lower bound. First of all, we assumed no accidental cancellations in the decay amplitude of the lepton-flavour violation. However, it is possible to suppress the process in some specific cases. Actually, \( \text{Br}(\mu \to e\gamma) \) can be lower than the bounds in Fig. \( \square \) when the right-handed neutrinos have a special flavour structure which suppresses \( |(Y_{N}^{\dagger}Y_{N})_{12}| \). Another cancellation may happen between the chargino and neutralino contributions, at some spots in the parameter space. Also, taking into account initial flavour-changing components of the soft SUSY breaking parameters at the mediation scale, cancellations may also happen because their contribution can destructively interfere with those from the right-handed neutrino Yukawa coupling. The second possibility is obtained by concerning the assumption of \( \mu_X \gtrsim \mu_R \). The charged-lepton sector receives the flavour-changing corrections through the renormalization evolutions only when the soft SUSY breaking effects are mediated before the right-handed neutrinos decouple. We have assumed that the mediation takes place at the GUT scale, while the gauge-mediated SUSY breaking scenarios generally have lower \( \mu_X \). The scale dependence is just logarithmic and very weak when the messenger scale is far from the right-handed neutrino mass scale. However, if the messenger scale approaches very close to the right-handed neutrino scale, the lepton-flavour violating processes become suppressed\(^{19}\). In those cases, the lepton-flavour violations will not be measured in near future even in leptogenesis scenarios with the muon \( g - 2 \) anomaly.

\(^b\) Since larger \( \epsilon_1 \) corresponds to heavier \( M_1 \), a higher reheating temperature is required, and thus the cosmological problem of the thermal gravitino production tends to be severer.
The lepton asymmetry may be enhanced compared with the evaluation in this letter. Actually, the right-handed neutrinos can have a (quite) degenerate mass spectrum. In the case of the hierarchical spectrum, the lepton asymmetry is obtained only from the decay of the lightest right-handed neutrino, while when the heavier right-handed neutrino mass becomes close to the lightest one, i.e., \( M_2 \to M_1 \), CP asymmetric decay of \( N_2 \) simultaneously contributes to the lepton asymmetry\(^9, 10\). Actually, Eq. (4) diverges for \( M_2 \to M_1 \), and taking into account \( \epsilon_2 \) in addition to \( \epsilon_1 \), the CP asymmetry is enhanced and becomes maximized when \( |M_2 - M_1| \sim \Gamma_2,3 + \Gamma_1 \)\(^9\). As another possibility, the flavour effect might affect the leptogenesis. This effect can modify the estimation of the CP asymmetry of the right-handed neutrino decay. For instance, even if the total CP asymmetry, \( \epsilon_1 \), is canceled, sufficient CP asymmetry is potentially produced by flavour effects\(^{21}\). Anyway, all of these possibilities restrict the right-handed neutrino structure. Thus, we may observe distinct signatures, e.g., \( \tau \to \mu \gamma \) and \( \tau \to e \gamma \), in future experiments in the light of the muon \( g - 2 \) anomaly. We will discuss these contents in future works.

Acknowledgment

The authors thank to Wilfried Buchmüller and Alejandro Ibarra for useful discussions and comments.

[1] G. W. Bennett et al. [Muon G-2 Collaboration], Phys. Rev. D 73, 072003 (2006).
[2] K. Hagiwara, A. D. Martin, D. Nomura and T. Teubner, Phys. Lett. B 649, 173 (2007).
[3] M. Passera, W. J. Marciano and A. Sirlin, arXiv:0804.1142 [hep-ph].
[4] J.L. Lopez, D.V. Nanopoulos and X. Wang, Phys. Rev. D 49, 366 (1994); U. Chattopadhyay and P. Nath, Phys. Rev. D 53, 1648 (1996); T. Moroi, Phys. Rev. D 53, 6565 (1996); Erratum-ibid, D 56, 4424 (1997).
[5] P. Minkowski, Phys. Lett. B 67 (1977) 421; M. Gell-Mann, P. Ramond and R. Slansky, Proceedings of the Supergravity Stony Brook Workshop, New York 1979, eds. P. Van Nieuwenhuizen and D. Freedman; T. Yanagida, Proceedings of the Workshop on Unified Theories and Baryon Number in the Universe, Tsukuba, Japan 1979, eds A. Sawada and A. Sugamoto.
[6] F. Borzumati and A. Masiero, Phys. Rev. Lett. 57, 961 (1986); J. Hisano, T. Moroi, K. Tobe, M. Yamaguchi and T. Yanagida, Phys. Lett. B 357, 579 (1995); J. Hisano, T. Moroi, K. Tobe and M. Yamaguchi, Phys. Rev. D 53, 2442 (1996); S. T. Petcov, S. Profumo, Y. Takanishi and C. E. Yaguna, Nucl. Phys. B 676, 453 (2004).

[7] J. Hisano and K. Tobe, Phys. Lett. B 510, 197 (2001).

[8] W. Buchmuller, P. Di Bari and M. Plumacher, Annals Phys. 315 305 (2005).

[9] A. Pilaftsis, Int. J. Mod. Phys. A 14, 1811 (1999).

[10] A. Pilaftsis and T. E. J. Underwood, Nucl. Phys. B 692, 303 (2004); A. Pilaftsis and T. E. J. Underwood, Phys. Rev. D 72, 113001 (2005); T. Hambye, Nucl. Phys. B 633, 171 (2002); T. Hambye, J. March-Russell and S. M. West, JHEP 0407, 070 (2004); Z. z. Xing and S. Zhou, Phys. Lett. B 653, 278 (2007); G. C. Branco, A. J. Buras, S. Jager, S. Uhlig and A. Weiler, JHEP 0709, 004 (2007); E. J. Chun and K. Turzynski, Phys. Rev. D 76, 053008 (2007); T. Kitabayashi, Phys. Rev. D 76, 033002 (2007).

[11] B. Pontecorvo, Zh. Eksp. Teor. Fiz. 33 549 (1957), 34 (1958) 247 and 53 1717 (1967); Z. Maki, M. Nakagawa and S. Sakata, Prog. Theor. Phys. 28 870 (1962).

[12] A. Bandyopadhyay, S. Choubey, S. Goswami, S. T. Petcov and D. P. Roy, Phys. Lett. B 608, 115 (2005); Phys. Rev. D 72, 033013 (2005); T. Schwetz, Phys. Scripta T127 1 (2006); G. L. Fogli, E. Lisi, A. Marrone, A. Palazzo and A. M. Rotunno, Prog. Part. Nucl. Phys. 57, 71 (2006).

[13] J.A. Casas and A. Ibarra, Nucl. Phys. B618 171 (2001).

[14] S. Davidson and A. Ibarra, Phys. Lett. B 535, 25 (2002).

[15] S. Pascoli, S. T. Petcov and W. Rodejohann, Phys. Rev. D 68 093007 (2003); S. T. Petcov, W. Rodejohann, T. Shindou and Y. Takanishi, Nucl. Phys. B 739 208 (2006).

[16] A. Ibarra and G. G. Ross, Phys. Lett. B 591 285 (2004).

[17] M. Raidal, A. Strumia and K. Turzynski, Phys. Lett. B 609, 351 (2005); Erratum-ibid. B 632, 752 (2006).

[18] S. Ritt [MEG Collaboration], Nucl. Phys. Proc. Suppl. 162, 279 (2006); M. Grassi [MEG Collaboration], Nucl. Phys. Proc. Suppl. 149, 369 (2005).

[19] K. Tobe, J. D. Wells and T. Yanagida, Phys. Rev. D 69, 035010 (2004).

[20] R. Barbieri, P. Creminelli, A. Strumia and N. Tetradis, Nucl. Phys. B 575, 61 (2000); H. B. Nielsen and Y. Takanishi, Nucl. Phys. B 636, 305 (2002); O. Vives, Phys. Rev. D
73, 073006 (2006); A. Abada, S. Davidson, F. X. Josse-Michaux, M. Losada and A. Riotto, JCAP 0604, 004 (2006); A. Abada, S. Davidson, A. Ibarra, F. X. Josse-Michaux, M. Losada and A. Riotto, JHEP 0609, 010 (2006); E. Nardi, Y. Nir, E. Roulet and J. Racker, JHEP 0601, 164 (2006).

[21] S. Pascoli, S. T. Petcov and A. Riotto, Phys. Rev. D 75, 083511 (2007); Nucl. Phys. B 774, 1 (2007).