Extension of Bertrand’s theorem and factorization of the radial Schrödinger equation

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(September 13, 2018)

I. INTRODUCTION

The Bertrand’s theorem is extended, i.e. closed orbits still may exist for central potentials other than the power law Coulomb potential and isotropic harmonic oscillator. It is shown that for the combined potential \( V(r) = W(r) + b/r^2 \) (\( W(r) = ar^n \)), when (and only when) \( W(r) \) is the Coulomb potential or isotropic harmonic oscillator, closed orbits still exist for suitable angular momentum. The correspondence between the closeness of classical orbits and the existence of raising and lowering operators derived from the factorization of the radial Schrödinger equation is investigated.

PACS number(s): 03.65. -w; 03.65. Ge

II. EXTENSION OF BERTRAND’S THEOREM

It will be shown that if the restriction of a pure power law central potential is relaxed, the Bertrand’s theorem may be extended. In particular, when an additional central potential \( b/r^2 \) is added to \( W(r) = ar^n \),

\[
V(r) = W(r) + b/r^2,
\]

we will show that when and only when \( W(r) \) is the Coulomb potential or isotropic harmonic potential, there still exist closed orbits for some suitable angular momenta.

The equations of motion of a particle in the central potential (4) is expressed as (\( \mu = 1 \))

\[
\begin{align*}
\dot{r} &= p_r/r^2, \\
p_r &= L^2/r^3 + f(r), \\
\dot{L} &= 0,
\end{align*}
\]

where \( L \) is the angular momentum and \( f(r) = -dV/dr = g(r) + 2b/r^3 \), \( g(r) = -dW/dr = -ar^{n-1} \). It was shown (3) that for a circular orbit at radius \( r_0 \) to be possible, the force must be attractive and

\[
E = L^2/2r_0^2 + V(r_0), \quad f(r_0) = -L^2/r_0^3.
\]

To guarantee the stability of a circular orbit \( r = r_0 \), it is required (2) that the second derivative of the effective potential \( U(r) = V(r) + L^2/2r^2 \) must be positive, i.e.

\[
f''(r_0) - 3f(r_0)/r_0 > 0.
\]

It is noted that the inclusion of the additional force \( 2b/r^3 \) does not alter the stability condition (2), so we have

\[
-g'(r_0) - 3g(r_0)/r_0 > 0, \quad \text{i.e.} \quad a(n+2) > 0.
\]
Thus $\nu < -2$ or $\nu > 0$ for $a > 0$, and $-2 < \nu < 0$ for $a < 0$. Since we are interested in the bounded motion, the case $\nu < -2$ for $a > 0$ is excluded. Thus, we obtain the stability condition for circular orbits: $av > 0$, i.e., $g(r) < 0$.

To seek the condition for closeness of an orbit, let us consider the orbit equation (3) for potential (1):

$$d^2u \over d\theta^2 + (1 + 2b/L^2)u = -{1 \over L^2} \overline{d} W(\overline{u})_0, \quad u = {1 \over r}.$$  
\hspace{1cm} (6)

Introducing $\kappa = \sqrt{1 + 2b/L^2}$, $\xi = \kappa u$ and $\phi = \kappa \theta$, we get

$$d^2\xi \over d\phi^2 + \xi = J(\xi), \quad J(\xi) = -{1 \over L^2} \overline{d} W(\overline{\kappa \xi}).$$  
\hspace{1cm} (7)

Using (3), it can be shown that at $\xi = \xi_0 = \kappa/r_0$, $J(\xi_0) = 0$. In addition, using (3) and $g(r) < 0$, one get $\beta^2 = 1 - J'(\xi_0) = 3 + \xi_0 \overline{d}\overline{d}\vert_{r=r_0} > 0$, then $\beta^2 = \nu + 2$.

Thus, the stability condition for circular orbits also may be expressed as: $\beta^2 > 2$ for $a > 0$, and $2 > \beta^2 > 0$ for $a < 0$.

Using the Taylor expansion of $J(\xi)$ at $\xi_0$, a deviation of an orbit from a circle $\delta \xi = \xi - \xi_0$, is governed by

$$d^2\delta \xi \over d\phi^2 + \delta \xi = J'(\xi_0)\delta \xi + {1 \over 6} J''(\xi_0)(\delta \xi)^2 + O((\delta \xi)^4).$$  
\hspace{1cm} (8)

For a small deviation of an orbit from circularity, to the first order of $\delta \xi$, (8) is reduced to

$$d^2\delta \xi \over d\phi^2 + \beta^2 \delta \xi = 0.$$  
\hspace{1cm} (9)

The solution to (9) can be expressed as:

$$\delta \xi = a_1 \cos(\beta \phi), \quad \kappa u = {a_1 \over \kappa} \cos(\beta \kappa \theta).$$  
\hspace{1cm} (10)

It can be seen that for $\beta \kappa = q/p$ (a rational number, $q$ and $p$ are integers), when $\theta = \theta_0 + 2\pi p$, $\delta u = \delta u_0$, i.e., the orbit remains closed. Thus, we get the condition for closeness of orbits against a small deviation from circularity: $\beta \kappa$ must be a rational number.

To seek the condition for a closed orbit when the deviation from circularity is considerable, $\delta \xi$ can be expressed as the Fourier expansion, i.e., in addition to the fundamental term in (9), more terms in the Fourier expansion should be included in $\delta \xi$. Substituting $\delta \xi$ in (9) and using the arguments similar to (8), one get $\beta^2=0$, 1 and 4 for closed orbits. Thus, we obtain the conclusion: when (and only when) $\beta^2 = 1$ ($\nu = -1$) for $a < 0$, and $\beta^2 = 4$ ($\nu = 2$) for $a > 0$, the orbit is closed. Then, $\kappa$ must be a rational number. Therefore, for the potential (1), when (and only when) $W(r)$ is the Coulomb potential or isotropic harmonic oscillator, closed orbits still exist for suitable angular momentum $L = \sqrt{2b/(\kappa^2 - 1)}$, where $\kappa$ is a rational number.

Example: For the valence electron in an alkalis atom $V(r)$ may be approximately expressed as ($e = \mu = \hbar = 1$)

$$V(r) = -\frac{1}{r} - \frac{\lambda}{r^2}, \quad (0 < \lambda \ll 1).$$  
\hspace{1cm} (11)

In this case, the orbit equation (3) is reduced to

$$d\theta = -du/\sqrt{2/\lambda L^2 + 2u/L^2 - \kappa^2 u^2},$$  
\hspace{1cm} (12)

where $\kappa = \sqrt{1 - 2\lambda/L^2}$. Integrating (12), we get

$$u = \frac{1}{r} \left[1 + \frac{1}{L^2 \kappa^2} \left(1 + 2E/\lambda \right) \cos(\kappa \theta - \theta_0)\right].$$  
\hspace{1cm} (13)

where $\sqrt{1 + 2E/\lambda \kappa^2} = \sqrt{1 + 2E/(L^2 - 2\lambda)} \geq 0$ to make the solution (13) having sense. For $1 + 2E/\lambda \kappa^2 = 0$, the orbits are circular. Because $\kappa < 1$ for $\lambda > 0$, no elliptic orbits exist. However, when $\kappa$ is a rational number, closed orbits (other than ellipses) still exist. Some illustrative examples are displayed in Fig. 1.

FIG. 1. The closed orbits of the valence electron in an alkalis atom ($\lambda = 0.2$).

III. FACTORIZATION OF THE RADIAL SCHRÖDINGER EQUATION

For a particle in the central field $V(r)$, the energy eigenfunction may be chosen as the simultaneous eigenstate of the complete set of conserved observables ($H, \mathbf{L}^2, \mathbf{L}_z$), and the radial wave function $R_{n,l}(r) = \chi_{n,l}(r)/r$, ($l$ are nonnegative integers) satisfies the radial Schrödinger equation ($\hbar = \mu = 1$)

$$H \chi_{n,l}(r) = E_n \chi_{n,l}(r),$$  
\hspace{1cm} (14)

For the potential (1), (14) is reduced to

$$[d^2/dr^2 - {l(l+1) \over r^2} - 2W(r) - \frac{2b}{L^2}] \chi_{n,l}(r) = -2E_n \chi_{n,l}(r).$$  
\hspace{1cm} (15)

Let $l' = (l + 1) + 2b$, i.e., $l' = -\frac{1}{2} + (l + \frac{1}{2})\sqrt{1 + 2b/(l + 1/2)^2}$, (15) can be rewritten as

$$D_{l'}(r)\chi_{n,l'}(r) = -2E_n \chi_{n,l'}(r),$$  
\hspace{1cm} (16)

which has the same form as the usual radial Schrödinger equation with the central potential $W(r)$, but the nonnegative integer quantum number $l$ is replaced by a non-integer $l'$. It has been shown (15) that if the power
law central potential $W(r)$ is assumed, only for the Coulomb potential or isotropic harmonic oscillator the radial Schrödinger equation can be factorized to derive the raising and lowering operators of $l'$ ($\Delta l' = \pm 1$). However, for $\Delta l' = \pm 1$, the shift in $l$ is by no means $\pm 1$. Therefore, though the radial Schrödinger equation can be factorized for the central potential (1), no angular momentum raising and lowering operators ($\Delta l = \pm 1$) can be constructed, which is intimately connected with the breaking of the dynamical symmetry of Coulomb potential and isotropic harmonic oscillator due to the added term $b/r^2$ in (1).

For the Coulomb potential and isotropic harmonic oscillator, whose exact eigenvalues are well known, the energy raising and lowering operators have been derived by using different approaches in (11). For a particle in a more general central potential, the exact energy eigenvalues are usually unknown. To derive the energy raising and lowering operators, we may use the WKB approximation. According to the WKB approximation, the eigenvalues of bounded states of (10) take the form (14) (note, here the usual angular momentum $l$ has been replaced by $l'$):

$$ E_n = \alpha_\nu n^{2\nu - 1}. $$

where $\alpha_\nu = a^{2\nu - 1/2} |\nu \sqrt{\pi} r^{\nu + 1/2}|^{2\nu} > 0$, $n = n_r + l'/2 + \frac{3}{4}$ for $a > 0$, and $\alpha_\nu = -(a)^{2\nu - 1/2} |\nu \sqrt{\pi} r^{\nu + 1/2}|^{2\nu} < 0$, $n = n_r + \frac{3}{2} + l'/2 - 2/\nu + 1$ for $a < 0$. Thus, (16) may be recast into

$$ D_n(r)\chi_n,l,v(r) = \nu (l'+1)\chi_n,l,v(r), $$

$$ D_n(r) = r^2 d^2/dr^2 - 2ar^{\nu+2} + 2\alpha_\nu n^{2\nu/(\nu+2)^2} r^2. $$

(1) For $a > 0$, the factorizability of (18) requires

$$ (rd/dr - \sqrt{2ar^{\nu+1}} + A)(rd/dr + \sqrt{2ar^{\nu+1}} + B)\chi_n,l,v(r) = D_n(r)\chi_n,l,v(r) + AB\chi_n,l,v(r), $$

then we have

$$ A + B + 1 = 0, $$

$$ \sqrt{2ar^{\nu+1}}(A - B + \nu/2 + 1) = 2\alpha_\nu n^{2\nu/(\nu+2)^2} r^2. $$

(20)

The solution to (20) is $\nu = 2$, $\alpha_\nu = 2\sqrt{2a}$, $A = 2n - 3/2$, and $B = -(2n - 1/2)$. Thus, (19) is reduced to

$$ (rd/dr - \sqrt{2ar^{2}} + 2n - 3/2)(rd/dr + \sqrt{2ar^{2}} - 2n + 3/2)\chi_n,l,v(r) = D_n(r)\chi_n,l,v(r) $$

$$ -(2n - 3/2)(2n - 1/2)\chi_n,l,v(r). $$

(21)

Similarly, (18) also can be factorized as

$$ (rd/dr + \sqrt{2ar^{2}} - (2n + 3/2)(rd/dr - \sqrt{2ar^{2}} + (2n + 1/2)\chi_n,l,v(r) $$

$$ -(2n + 1/2)(2n + 3/2)\chi_n,l,v(r). $$

Using (21) and (22), one may obtain the energy raising and lowering operators

$$ \chi_{n+2,l,v}(r) \sim [rd/dr - \sqrt{2ar^{2}} + (2n + 1/2)]\chi_{n,l,v}(r), $$

$$ \chi_{n-2,l,v}(r) \sim [rd/dr + \sqrt{2ar^{2}} - (2n - 1/2)]\chi_{n,l,v}(r). $$

(23)

For $a < 0$, the factorizability of (18) requires

$$ (rd/dr - \sqrt{-2\alpha_\nu n^{2\nu/(\nu+2)^2}} + A)(rd/dr + \sqrt{-2\alpha_\nu n^{2\nu/(\nu+2)^2}} + B)\chi_n,l,v(r) = D_n(r)\chi_n,l,v(r) $$

$$ + AB\chi_n,l,v(r), $$

then we have

$$ A + B + 1 = 0, $$

$$ \sqrt{-2\alpha_\nu n^{2\nu/(\nu+2)^2}}(A - B + \nu/2 + 1) = -2\alpha_\nu n^{2\nu/(\nu+2)^2} r^2. $$

(25)

The solution to (25) is $\nu = -1$, $\alpha_\nu = -a^2/2$, $A = n - 1$, and $B = -n$. Thus, (24) is reduced to

$$ [rd/dr + ar/n + n - 1][rd/dr - ar/n] $$

$$ -n\chi_n,l,v(r) = D_n(r)\chi_n,l,v(r) - (n-1)n\chi_n,l,v(r). $$

(26)

Similarly, (18) also can be factorized as

$$ [rd/dr - ar/n - (n+1)][rd/dr + ar/n + n] $$

$$ +n\chi_n,l,v(r) = D_n(r)\chi_n,l,v(r) - (n+1)n\chi_n,l,v(r). $$

(27)

Using (26) and (27) and introducing the scaling operator $M(k)$ defined by $M(k)f(r) = f(kr)$, one obtains the energy raising and lowering operators

$$ \chi_{n+1,l,v}(r) \sim M(k)(n-1)(rd/dr + ar/n + n)\chi_n(l,v)(r) $$

$$ = [rd/dr + ar/(n+1) + n]M(k-1)(n-1)\chi_n(l,v)(r), $$

$$ \chi_{n-1,l,v}(r) \sim M(k)(n+1)(rd/dr - ar/n - n)\chi_n(l,v)(r) $$

$$ = [rd/dr - ar/(n-1) - n]M(k-1)(n+1)\chi_n(l,v)(r). $$

(28)

Thus, we arrive at the conclusion that, for the potential (1), only when $W(r)$ is the Coulomb potential or isotropic harmonic oscillator, the radial Schrödinger equation can be factorized to derive the energy raising and lowering operators.

For a classical particle in the Coulomb potential $W(r) = -1/r$, the orbit is always closed for any negative energy ($E < 0$) and positive angular momentum $L$, i.e., an ellipse, of which the length of semi-major axis is $a = 1/2|E|$ and the eccentricity $e = \sqrt{1 - 2|E|L^2}$. This is guaranteed by the existence of an additional conserved quantity—the Runge-Lenz vector $R = p \times L - r/v$, which determines the direction of semi-major axis and the eccentricity $e = |R|$. In quantum mechanics, the angular momentum and bound energy are quantized. Therefore, it is not surprising that for the Coulomb potential the radial Schrödinger equation can be factorized and there exist four kinds of raising and lowering operators connecting various eigenstates of energy and angular momentum (11).

However, for the combined central potential (11), the dynamical symmetry $O_4$ is broken and $R$ no longer remains constant, and the orbit, in general, is not closed.
It has been shown in sect. II that for suitable angular momenta $L = \sqrt{2\lambda/(1 - \kappa^2)}$ ($\kappa$ is a rational number), the orbits are still closed (but not ellipses). Therefore, it is understandable why no angular momentum raising and lowering operators $(\Delta l = \pm 1)$ exist, but energy raising and lowering operators for a fixed $l'$ (see (28)) still can be constructed. In fact, for $\lambda = 0$, (28) is just the raising and lowering operators $B(l, n \uparrow)$ and $B(l, n \downarrow)$ given in eq. (33) of [6]. It is noted that the angular momentum raising and lowering operators $A$, $C$ and $D$ in [6] no longer exist for the combined potential (11). For the isotropic harmonic oscillator, the situation is similar.

IV. SUMMARY

The Bertrand’s theorem is extended in this paper. For a classical particle in a power law central potential $W(r)$, the Bertrand’s theorem does hold, i.e., only for the Coulomb potential or isotropic harmonic oscillator, closed orbits always exist for continuous energy and angular momentum. Accordingly, in quantum mechanics there exist both energy and angular momentum raising and lowering operators connecting neighboring discrete energy and angular momentum eigenstates. For the combined potential (1), it is shown that when $W(r)$ is the Coulomb potential or isotropic harmonic oscillator, classical closed orbits (other than ellipse) still exist for continuous energy, but only for suitable angular momenta. From this one can understand why in this case there exist only energy raising and lowering operators in quantum mechanics, but no angular momentum raising and lowering operators.

This work was supported in part by the Post-Doctoral Foundation of China (ZBW) and the National Natural Science Foundation of China (JYZ).

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