Pressure-temperature phase diagram of charge ordering in Nd$_{1/2}$Sr$_{1/2}$MnO$_3$

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Abstract — We observe how the charge-ordering (CO) temperature $T_{CO}$ of Nd$_{1/2}$Sr$_{1/2}$MnO$_3$ decreases with the external pressure $p$ from 160 K at $p = 0$ down to 30 K at $p \approx 4.5$ GPa, by measuring the values $p$, $T$ where the far-infrared spectral weight of the metallic phase is fully recovered. We thus determine the $(p, T)$ phase diagram of CO in that manganite. We also find that the parameter $d(\ln T_{CO})/dp$ which describes this metallization from the CO phase is equal and opposite to the quantity $d(\ln T_c)/dp$ which governs the metallization of the paramagnetic state at comparable Curie temperatures $T_c$, in similar manganites at half-doping.

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Introduction. — The $(x, T)$ phase diagram of the manganites $A_1-xB_xMnO_3$, where $A$ is a lanthanoid and $B$ an alkali earth, includes charge-ordered phases for $0.5 \leq x \leq 0.85$ and $T < T_{CO}$. $T_{CO}$, the charge-ordering (CO) temperature, in the absence of an external field may vary from 160 K for Nd$_0.5$Sr$_0.5$MnO$_3$ (NSMO) [1] to more than 500 K for Bi$_{0.5}$Sr$_{0.5}$MnO$_3$ [2]. The CO is usually associated with both orbital ordering and antiferromagnetism (AFM) [3], so that the ordered phase may collapse under magnetic fields $B$ which span from 10 T for NSMO [4] to more than 50 T for Bi$_{0.5}$Sr$_{0.5}$MnO$_3$ [2]. For a long time, CO in manganites has been discussed in terms of arrays of Jahn-Teller small polarons [5]. However, in manganites with $T_{CO}$ below room temperature and with moderate critical fields, there are strong indications [6-9] that the CO phase is better described in terms of Charge Density Waves (CDW) [10]. This instability of the CO liquid takes place in a regime of moderately weak electron-phonon coupling and is associated with a perfect (in the ideal one-dimensional CDW) or partial nesting of the Fermi surface.

In the optical conductivity $\sigma(\omega)$, the transition to a CO phase shows up in the appearance of an insulating gap $E_g$ in the mid infrared [11] and of absorption peaks in the Terahertz range [8,12]. Conversely, the infrared-gap filling —with partial shielding of the phonon peaks by a free-carrier absorption— provides evidence for an insulator-to-metal transition (IMT) due to the CO collapse. Besides $T$ and $B$, an external pressure $p$ [13] can destabilize the charge order. Pressures in the GPa range have been shown to destroy the CO ground state of CeTe$_3$ at room temperature [13]. In the manganites, no infrared measurements at high pressure have been reported in a CO state to our knowledge. However, transitions from the paramagnetic (PM) phase to the ferromagnetic (FM), metallic phase, induced by pressure, were observed by monitoring either the mid [14] or the far infrared [15] $\sigma(\omega)$. In the present experiment we have investigated the effect of $p$ on the CO phase of a manganite at low temperature, by measuring its $\sigma(\omega)$ in the far infrared at different pressures and temperatures. The selected system is Nd$_{1/2}$Sr$_{1/2}$MnO$_3$ (NSMO), a manganite which exhibits both a homogeneous, commensurate CO with periodicity $2a$ ($a = $ lattice parameter), and AFM (of charge-exchange, CE, type) below a $T_{CDW} \approx 160$ K (onset) [1] at ambient pressure. This transition is clearly detected in the infrared through a CO gap at $E_g \sim 0.1$ eV [8], which opens in the Drude term of the ferromagnetic (FM) metallic phase. This is stabilized by the double-exchange mechanism [3] from $T_{CO}$ to a Curie temperature $T_c \approx 255$ K [16]. Here we detect the IMT induced by pressure at $T < T_{CO}$ by

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measuring the spectral weight
\[ W(p, T) = \int_{\Omega_1}^{\Omega_2} \sigma(\omega, T) \ d\omega \] (1)
at zero pressure in the metal at 200 K and observing how it is recovered in the low-\( T \) AFM phase due to closure of the CO gap. In our case, \( \Omega_1 = 120 \text{ cm}^{-1} \) and \( \Omega_2 = 1000 \text{ cm}^{-1} \).

**Experimental.** – The optical conductivity of NSMO was extracted from reflectivity measurements performed down to 30 K in a diamond anvil cell (DAC). In order to obtain a reasonable photon flux through the small aperture of the DAC [17], we used instead of conventional sources the InfraRed Synchrotron Radiation (IRSR) available at the beam line BL43IR of the storage ring SPring-8 [18].

The sample, with a reflecting surface of about \( 100 \times 100 \mu\text{m} \) corresponding to the \( ab \)-plane, was cut from a single crystal grown by the floating-zone method [6]. It was inserted in a DAC, together with a parallel gold mirror for reference, and with a small ruby. The frequency of its fluorescence peak provided at any \( T \), after long thermalization, the pressure \( p \) at \( T \). The pressure medium was Daphne7474 [19] and the cell was inserted in a liquid-helium cryostat. Due to the surprising low pressures at which the AFM-FM transition occurs, the contraction of the cell upon cooling played a major role. For this reason, one had often to select different \( p \) values for the measurements at different \( T \).

The intensity \( I_s \) reflected by the sample was divided by the \( I_r \) reflected in the same conditions by the mirror, to obtain the reflectivity \( R(\omega) \) at the interface between the manganite and the diamond window. This quantity is shown in fig. 1 at different temperatures and pressures. Even these raw data clearly show the effect of pressure on the optical response. In (a) at 200 K (in the FM phase) \( R(\omega) \) increases slightly at high \( \omega \) and then decreases in the whole range for increasing \( p \). In (c) at 30 K, deeply in the CO ground state, both the overall reflectivity and the phonon shielding increase with pressure. At 4.6 GPa the shape of \( R(\omega) \) is again similar to that observed in panel (a) at \( p \sim 0 \). At 160 K (b), the metallic \( R(\omega) \) is reached at 2.7 GPa. Due to the limited frequency range accessible to the experiment, standard Kramers-Kronig transformations could not be used to extract \( \sigma(\omega) \) from \( R(\omega) \). This quantity was instead obtained by fitting, to the data of fig. 1, the formula
\[ R(\omega) = \frac{\sqrt{\epsilon(\omega) - n_d^2}}{\sqrt{\epsilon(\omega) + n_d^2}}, \] (2)
where the refractive index of diamond \( n_d \) was kept real and equal to 2.37, a value obtained by averaging \( n_d(\omega) \) [20] over the whole measuring range. An example of such a fit is shown in fig. 1(c).

For the dielectric function \( \epsilon(\omega) \) we assumed the usual expansion in terms of Drude and Lorentz oscillators from free charges and phonons, respectively. The fit also requires a polaron-like excitation [21] centered around \( 500 \text{ cm}^{-1} \), and the inclusion of the tails of the higher-frequency bands previously measured in NSMO [22]. An example of such decomposition is shown at 200 K in panel (a) of fig. 2, where \( \sigma(\omega) = \omega\epsilon_2(\omega)/4\pi \) is reported for the same temperatures and pressures as in fig. 1. In the FM phase at 200 K (fig. 2(a)) one sees again that at the highest \( p \) the sample has partly lost its metallic character. In the CO-AFM phase (panels (b) and (c)), the effect of pressure is opposite. It promotes a metallization process which shows up clearly both at 30 K (fig. 2b) and at 160 K (c), through an increase of the overall conductivity and a partial shielding of the phonon peaks.

**Results and discussion.** – In order to describe quantitatively the metallization process, we have calculated \( W \) and normalized its value to the corresponding \( W_0 \) in the metallic FM phase at the lowest pressure (\( T = 200 \text{ K}, P = 0.05 \text{ GPa} \)). The results are reported as a function of pressure in fig. 3. For sake of clarity, data taken above 130 K are shown in panel (a), those below this temperature in panel (b). In the FM phase at \( T = 200 \text{ K} \) (fig. 3(a)), the behavior of \( W \) indicates a significant decrease in the metallic character of NSMO around \( p^* \simeq 1.5 \text{ GPa} \). This effect, similar to that reported in \( \text{La}_{0.75}\text{Ca}_{0.25}\text{MnO}_3 \) [23],

![Fig. 1: (Color online) Reflectivity \( R(\omega) \) of \( \text{Nd}_{1/2}\text{Sr}_{1/2}\text{MnO}_3 \) at three different temperatures and under increasing pressures, as measured at the interface with the diamond window. The dotted line in (c) is an example of the Drude-Lorentz fit to data, as explained in the text.](image-url)
points out the increase of the super-exchange interaction, which tends to establish an AFM coupling between Mn nearest neighbors. In fig. 3(a), the drop in $W$ occurs at a much lower pressure than in La$_{0.75}$Ca$_{0.25}$MnO$_3$ (23 GPa), but in this compound the ground state is FM, while in NSMO the FM phase is unstable with respect to the AFM ground state. The decrease of $W$ at 200 K under pressure may then be attributed to the appearance of AFM zones, in a phase-separation scenario often invoked for the manganites at half-doping [3].

In the four sets of data at or below 160 K in fig. 3(a), $W$ increases monotonically with $p$. In the end, the metallic spectral weight $W_0$ is recovered at a pressure $p_M$ which increases for decreasing temperature. One thus can associate with each $p_M$ a $T_{CO}$ which marks the collapse of the CO phase. Those pairs of values will be used below to plot a boundary line between the CO and the metallic phase of NSMO. Below 130 K (fig. 3(b)), all data show a similar behavior and therefore they are grouped around a single line. At the lowest $T$ $W \simeq W_0$ at $p_M \simeq 4.5$ GPa, which therefore is the CO critical pressure. This value is in excellent agreement with the dc resistivity which, above 4.5 GPa, does not show any CO transition [24] and is rather low. This confirms that the interactions which stabilize the CO-AFM phase in manganites like NSMO are not so strong as to prevent the application of the CDW model [8].

Let us now discuss what is implied in the above results. Pressure promotes the IMT by reducing either the average Mn-O distance $d$ and the average angle $\beta$ of bonds Mn-O-Mn in the $ab$-plane. Then, $\pi - \beta$ measures the distortion of the MnO$_6$ octahedra [25]. Both mechanisms affect directly the spectral weight $W$ through the electron bandwidth, and in a tight-binding approximation [26]

$$W(p, T < T_{CO}) \approx \frac{\cos[|\pi - \beta|/2]}{d^{3.5}}.$$  \hspace{1cm} (3)

However, $d$ decreases by a bare 0.7% between 0 and 4.5 GPa, while in the ground CE state $\beta = 157^\circ$ [27]. Even assuming $\beta = \pi$ at $p_M$, according to eq. (3) the isotropic octahedra increase $W$ by just a 4%. In fig. 3(b), $W$ grows instead by a factor of 3. Therefore, the small pressure-induced changes in eq. (3) are hugely amplified by non-linear effects. The most important one is obviously the transition from the AFM super-exchange to the FM double-exchange between the Mn$^{3+}$ and Mn$^{4+}$ ions. However, also the charge-lattice interaction is expected to play a role via the Jahn-Teller effect [25].

The change of slope at $p^*$ in fig. 3(b) may be related to an observation [28] in the similar compound Nd$_{1/2}$Ca$_{1/2}$MnO$_3$. Therein, for increasing $p$, while the lattice parameter $a$ decreases monotonically, the
Jahn-Teller distortion was found to decrease up to 7 GPa, and then to increase again. Its “V”-shaped behavior vs. \( p \) was attributed to a subtle shear strain of the octahedra. In NSMO this should happen at pressures lower than 7 GPa, due to the larger ionic volume of Sr with respect to Ca, and may explain the observation at \( p^* \). However, the origin of the crossover in fig. 3 at the same \( p^* \) in both phases FM and CO, is not clear yet.

The pairs of values \( [p_M(T), T_{\text{CO}}] \) obtained from fig. 3 at \( W/W_0 \approx 1 \) allow one to build up the boundary line between the metallic and the CO phase reported in fig. 4, which is obtained by a power-law fit to data (dots). The transition pressure increases steeply below \( T_{\text{CO}} \), to saturate at about 4.5 GPa at low temperature. This behavior is well illustrated by the variation of the parameter \[29,30]\]

\[
\Psi = \frac{1}{T_M} dT_M/dp_M = d(\ln T_M)/dp_M
\]

which was originally reported for the PM-FM transition at the Curie temperature \( T_c \) and which here can be calculated from the derivative of the boundary line in fig. 4 (see the scheme in the inset of fig. 5). Therefore, in eq. (4), \( T_M = T_c \) for the PM-FM metallization (\( \Psi > 0 \)) and \( T_M = T_{\text{CO}} \) for the AFM-FM transition (\( \Psi < 0 \)). In fig. 5 \( \Psi \) is plotted as extracted from the present data (red solid line) and compared with the (dashed) line reported in ref. [30]. This was traced through three values (empty symbols), having opposite sign with respect to the present ones, and taken around \( T_c \) in other manganites at half-doping: SmNSM0 = \( \text{Sm}_{0.875}\text{Nd}_{0.125}\text{Sr}_{0.5}\text{MnO}_{3} \) \( (T_c = 115 \text{ K}) \); NSPMO = \( \text{Nd}_{0.5}\text{Sr}_{0.36}\text{Pb}_{0.14}\text{MnO}_{3} \) \( (T_c = 204 \text{ K}) \); LSMO = \( \text{La}_{0.5}\text{Sr}_{0.5}\text{MnO}_{3} \) \( (T_c = 360 \text{ K}) \). The agreement both in the slope and in the absolute values of \( \Psi \) in the common range of transition temperatures is impressive, but can be explained by the following argument. Indeed \[29\], \( T_c \propto D \) and \( T_{\text{CO}} \propto U/D \) where \( D \) is the carrier bandwidth renormalized by polaronic effects \[31\] and \( U \) is the correlation energy (assumed to be independent of \( p \)). Therefore, at the same \( T \),

\[
d(\ln T_c)/dp_M = -d(\ln T_{\text{CO}})/dp_M.
\]

**Conclusion.** We have here reported the optical observation of the collapse, induced by pressure, of the charge-ordered phase in NSMO at low temperature. The process slows down at about the same pressure \( p^* \) where the metal shows an instability above \( T_{\text{CO}} \). Through the analysis of the spectral weight we obtain the \((p,T)\) phase diagram of the charge order and the parameter \( \Psi \) of the AFM-FM transition. This is equal and opposite to that measured for the PM-FM transition at comparable Curie temperatures in other manganites at half-doping. This confirms experimentally that the metallization induced by pressure in manganites is governed by the same physics, either if one starts from a paramagnetic insulator, or from a charge-ordered antiferromagnet.

We thank X. Blasco for providing the NSMO crystal which the present sample was cut from.

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![Graph of Pressure-temperature diagram of the CO phase in Nd\(_{1/2}\)Sr\(_{1/2}\)MnO\(_3\).](image)

![Graph showing Absolute values of the \( \Psi \) defined in eq. (4).](image)
REFERENCES

[1] Caignaert V. et al., Solid State Commun., 99 (1996) 173.
[2] Frontera C. et al., J. Phys.: Condens. Matter, 13 (2001) 1071.
[3] Dagotto E., Hotta T. and Moreo A., Phys. Rep., 344 (2001) 1.
[4] Popov Yu. F. et al., Phys. Solid State, 45 (2003) 1221.
[5] Goodenough J., Phys. Rev., 100 (1955) 564.
[6] Herrero-Martin J. et al., Phys. Rev. B, 70 (2004) 024408.
[7] Loudon J. C. et al., Phys. Rev. Lett., 94 (2005) 097202.
[8] Nucara A. et al., Phys. Rev. Lett., 101 (2008) 066407.
[9] Milward G. C. et al., Solid State Commun., 14 (1974) 703.
[10] Grünner G., Rev. Mod. Phys., 60 (1988) 1129.
[11] Calvani P. et al., Phys. Rev. Lett., 81 (1998) 4504.
[12] Zhukova E. et al., EPL, 90 (2010) 17005.
[13] Sacchetti A. et al., Phys. Rev. Lett., 98 (2007) 026401.
[14] Congeduti A. et al., Phys. Rev. B, 63 (2001) 184410.
[15] Sacchetti A. et al., Phys. Rev. Lett., 96 (2006) 035503.
[16] Tokura Y. et al., Phys. Rev. Lett., 76 (1996) 3184.
[17] Cestelli Guidi M. et al., J. Opt. Soc. Am. A, 22 (2005) 2810.
[18] Ikeimoto Y. et al., Infrared Phys. Technol., 45 (2004) 369.
[19] Murata K. et al., Rev. Sci. Instrum., 79 (2008) 085101.
[20] Dore P. et al., Appl. Opt., 37 (1998) 5731.
[21] Calvani P., Riv. Nuovo Cimento, 24 (2001) 1.
[22] Nucara A. et al., Phys. Rev. B, 77 (2008) 064431.
[23] Ding Y. et al., Phys. Rev. Lett., 102 (2009) 237201.
[24] Yu R. C. et al., J. Appl. Phys., 97 (2005) 083910.
[25] Laukhin V., Fontcuberta J., García-Muñoz J. L. and Obradors X., Phys. Rev. B, 56 (1997) R10009.
[26] Radaelli P. G. et al., Phys. Rev. B, 56 (1997) 8265.
[27] Kajimoto R. et al., Phys. Rev. B, 60 (1999) 9506.
[28] Arulraj A. et al., Phys. Rev. Lett., 94 (2005) 165504.
[29] Fontcuberta J., Laukhin V. and Obradors X., Phys. Rev. B, 60 (1999) 6266.
[30] Zhao G. M., Hunt M. B. and Keller H., Phys. Rev. Lett., 78 (1997) 955.
[31] Millis A. J., Shraiman B. I. and Mueller R., Phys. Rev. Lett., 77 (1996) 175.