A Numberable Set of Exact Solutions for the Macroscopic Approach to Extended Thermodynamics of Polyatomic Gases with Many Moments

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A new model for Polyatomic Gases with an arbitrary but fixed number of moments has been recently proposed and investigated in the framework of Extended Thermodynamics; the arbitrariness of the number of moments is linked to a number \( N \) and the resulting model is called an \( N \)-Model. This model has been elaborated in order to take into account the entropy principle, the Galilean relativity principle, and some symmetry conditions. It has been proved that the solution for all these conditions exists, but it has not been written explicitly because hard notation is necessary; it has only been shown how the theory is self-generating in the sense that if we know the closure of the \( N \)-Model, then we will be able to find that of \((N+1)\)-Model. Up to now only a single particular solution has been found in this regard. Instead of this, we find here a numberable set of exact solutions which hold for every fixed number \( N \).

1. Introduction

The model for Polyatomic Gases considers for every number \( N \) the following balance equations:

\[
\begin{align*}
\partial_t F^{i_1\cdots i_n} + \partial_{x_k} F^{kj_1\cdots j_n} &= P^{i_1\cdots i_n} \quad \text{for } n = 0, \ldots, N + 2, \\
\partial_t G^{i_1\cdots i_m} + \partial_{x_k} G^{kj_1\cdots j_m} &= Q^{i_1\cdots i_m} \quad \text{for } m = 0, \ldots, N
\end{align*}
\]  

(1)

and the resulting model is called an \( N \)-Model.

Here \( F^{i_1\cdots i_n} \) represents the momentum-like hierarchy, while \( G^{i_1\cdots i_m} \) represents the energy-like hierarchy; the quantities with an index \( k \) represent their fluxes, while \( P^{i_1\cdots i_n} \) and \( Q^{i_1\cdots i_m} \) represent the production terms; for the conservation laws of mass and momentum we have \( P = 0 \) and \( F = 0 \), while for the conservation of energy we have \( Q = 0 \).

This model represents a novelty with respect to the earlier versions of Extended Thermodynamics (ET) which used only \( (1)_1 \) obtaining the important result to have a symmetric hyperbolic system of partial differential equations, with finite speed of propagation of shock waves and other important properties. A very short list of significative articles in this regard is given in [1–7]. These results constituted an important improvement of the models of Ordinary Thermodynamics. A very significative article in this previous framework is given in [8]. But ET was still not completely satisfactory because it is too much restrictive also on the state function relating the pressure \( p \) to the mass density \( \rho \) and energy density \( \epsilon \). This problem has been overcome in Extended Thermodynamics of Polyatomic Gases by considering all the balance equations (1) of which the first one is called the “Mass-Block” of equations and the second is called the “Energy-Block” of equations. See, for example, the articles [9–16].

It has to be noted that in the literature here cited the balance equations are not included in the \( N \)-Model, but they can be easily recovered with the theory of subsystems (see [17]) by simply putting the Lagrange multiplier of the last equation in the Mass-Block equal to zero. In \([11, 12]\) it has been shown that the wave speeds of the \( N \)-Model are the same as those of monatomic gases; but this proof has been
given with the kinetic equations which are only a subcase of those of the macroscopic approach (in effect, the whole closure here found is zero in the kinetic approach and only parts (5) and (6) remains). In any case there is no problem for what we have said at the beginning of this note. Moreover, a priori the Bhatnagar–Gross–Krook equation (see the equation before (14) in [10]) can be replaced by an infinite hierarchy of balance equations for moments and only for practical reasons some of them are eliminated; consequently, there is no inconsistence if another of these balance equations is given.

In (1) the independent variables are \( F^{i-1}_{i} \) and \( G^{i-1}_{i} \); the quantities \( F^{ki}_{i} \) and \( G^{ki}_{i} \) are their corresponding fluxes. The problem will be closed when we know the expressions of \( F^{ki}_{i} \) and \( G^{ki}_{i} \) as functions of the independent variables. Restrictions on their generality are obtained by imposing the entropy principle, the Galilean relativity principle, and the symmetry conditions. These last ones express the fact that each flux in (1) is equal to the independent variable of the subsequent equation, except for the flux in the last equation of the Mass-Block and for that in the last equation of the Energy-Block; for these last fluxes we know only that they are symmetric tensors.

The Entropy Principle, exploited through Liu’s Theorem [18] and a bright idea conceived by Ruggeri and Strumia [19], becomes equivalent to assuming the existence of Lagrange Multipliers \( \mu_A \) and \( \lambda_B \) which can be taken as independent variables and, after that, we have

\[
\begin{align*}
F^{i-1}_{i} &= \frac{\partial h'}{\partial \mu_{i-1}} , \\
G^{i-1}_{i} &= \frac{\partial h'}{\partial \lambda_{i-1}} , \\
F^{ki}_{i} &= \frac{\partial h_k}{\partial \mu_{i-1}} , \\
G^{ki}_{i} &= \frac{\partial h_k}{\partial \lambda_{i-1}},
\end{align*}
\]

which expresses all the moments in terms of only two unknown functions, the 4-potential \( h' \) and \( h^k \). Their relation with the entropy \( h \) and its flux \( h^k \) is given by

\[
\begin{align*}
h &= -h' + F^{i-1}_{i} \mu_{i-1} + G^{i-1}_{i} \lambda_{i-1} , \\
h^k &= -h^{ik} + F^{ki}_{i} \mu_{i-1} + G^{ki}_{i} \lambda_{i-1}.
\end{align*}
\]

A nice consequence of (2) is that the field equations assume the symmetric form. Another consequence of (2) is that the above-mentioned symmetry conditions and the Galilean relativity principle can be expressed as

\[
\begin{align*}
\frac{\partial h^k}{\partial \mu_{i-1}} &= \frac{\partial h'}{\partial \mu_{i-1}} & \text{for } n = 0, \ldots, N + 1; \\
\frac{\partial h^k}{\partial \lambda_{i-1}} &= 0 , \\
\frac{\partial h^k}{\partial \lambda_{i-1}} &= \frac{\partial h'}{\partial \lambda_{i-1}} & \text{for } m = 0, \ldots, N - 1; \\
\frac{\partial h^k}{\partial \lambda_{i-1}} &= 0 ,
\end{align*}
\]

The equivalence of (4), to the Galilean relativity principle is proved in literature such as [13, 20, 21] and on Representation Theorems such as that of [22].

In (5), (6), and (7) we report an already known particular solution (see [23]) for the conditions (4). But it is not the general solution; in fact, in the present article we will exhibit a significative set of other solutions depending on a numberable family of arbitrary single variable functions \( g_b(\lambda) \) and we report it in the next section, (8) and (9). It is expressed through Taylor’s expansion around equilibrium which is defined as the state where \( \mu_{i-1} = 0 \), for \( n = 1, \ldots, N + 2 \), and \( \lambda_{i-1} = 0 \), for \( m = 1, \ldots, N \), so that the only variables which are not zero at equilibrium are \( \mu \) and \( \lambda \). The first of these is the chemical potential, while \( \lambda = 1/2T \) with \( T \) absolute temperature.

Obviously, also the velocity is nonvanishing at equilibrium; but if we use a decomposition in nonconvective parts and in velocity dependent parts, we obtain the Lagrange multiplier \( \mu \) of the momentum conservation law in terms of the other Lagrange multipliers and we see that its velocity independent part is zero at equilibrium. Moreover, from the fact that the velocity independent parts of \( \mu \) and \( \lambda \) are zero at equilibrium, it follows that they are zero also as a whole, without decomposition; the only exceptions are for the Lagrange multipliers of the mass and of the energy conservation laws (see equation (25) of [21], e.g.).

The proof of the fact that (8) and (9) give a solution for (4) is here split into two parts: the first one is reported in Section 3 and the second and final part is described in Section 4.

In conclusion, the sum of the solution reported in (5), (6), and (7) and the present one reported in (8) and (9) is a more general solution for our conditions.
2. A Set of Exact Solutions for Conditions (4)

Before writing our solution we need an apparently complicated notation.

To do the derivatives with respect to \( \mu_{i,j} \) a number of times, we use the compact form of our conditions and it is

\[
\frac{\partial \psi_{n+1}}{\partial \mu} = \psi_n.
\]

But the above one is not the unique solution; in fact, by using the same notation, we can now say that a more general solution can be obtained by adding to (5) and (6) the following expressions (8) and (9); that is,

\[
\frac{1}{r_N!} \sum_{\sum_{i=1}^{N} ip_i + \sum_{i=1}^{N} ir_i \text{ even}} \frac{1}{p_0!} \frac{1}{p_1!} \ldots \frac{1}{p_{N+2}!} \frac{1}{r_N!} \left[ \frac{1 + \frac{N+2}{2} \sum_{i=1}^{N} ip_i + \sum_{i=1}^{N} ir_i}{1 + \frac{N+2}{2} \sum_{i=1}^{N} ip_i + \sum_{i=1}^{N} ir_i} \right]!!
\]

\[
\frac{\partial \psi_{n+1}}{\partial \mu} = \psi_n.
\]

where \( \psi_n \) is a family of functions depending only on \( \mu \) and \( \lambda \) and constrained only by

\[
\frac{1}{r_N!} \sum_{\sum_{i=1}^{N} ip_i + \sum_{i=1}^{N} ir_i \text{ odd}} \frac{1}{p_0!} \frac{1}{p_1!} \ldots \frac{1}{p_{N+2}!} \frac{1}{r_N!} \left[ \frac{1 + \frac{N+2}{2} \sum_{i=1}^{N} ip_i + \sum_{i=1}^{N} ir_i}{1 + \frac{N+2}{2} \sum_{i=1}^{N} ip_i + \sum_{i=1}^{N} ir_i} \right]!!
\]

\[
\frac{\partial \psi_{n+1}}{\partial \mu} = \psi_n.
\]
a shortened symbol denoting the product of some $\delta^-$ each one with 2 indexes and with a final symmetrization over all these indexes.

We note the following:

(i) If $N = 0$, the function $g$ in the expression of $h_k^r$ has index $-2p_0 - p_1$ and consequently this $g$ may be different from zero only if $p_0 = 0$ and $p_1 = 0$; but in this case $p_1 + 2p_2$ is not odd. This fact implies that $h_k^r = 0$. Similarly, the function $g$ in the expression of $h_l^r$ has index $-2p_0 - p_1 - 1$ so that $g = 0$ and $h_l^r = 0$. This confirms the results of [24] for this case $N = 0$, because in the sum of (8) and (9) and the corresponding results (5) and (6) of [23] only these last ones remain in the case $N = 0$.

(ii) If $N = 1$, the function $g$ in the expression of $h_k^r$ has index $-2p_0 - p_1 + p_2 + r_1$ and this is an odd number because $p_1 + 2p_2 + 3p_3 + r_1$ must be even. Consequently, the index $p_0$ belongs to the interval $[0, (−p_1 + p_2 + r_1 - 1)/2]$ if this is not the empty set. By taking into account this fact we see that the sum of the solution reported in (6) and the expression (9) for the case $N = 1$ is the same as that in [24]. Similarly, the function $g$ in the expression of $h_l^r$ has index $−2p_0 − p_1 + p_2 + r_1 − 1$ and this is an odd number because $p_1 + 2p_2 + 3p_3 + r_1$ must be even. Consequently, the index $p_0$ belongs to the interval $[0, (−p_1 + p_2 + r_1 − 2)/2]$ if this is not the empty set. By taking into account this fact we see that the sum of our expression (8) and that in (5) for the case $N = 1$ is the same as that in [24].

(iii) If $N \geq 2$ our solution is completely new.

3. The New Solution (8) and (9) Satisfies (4)\textsubscript{1–4}

In this section we prove that (8) and (9) are a solution for (4)\textsubscript{1–4}, putting off (4)\textsubscript{5} to Section 4:

(i) Let us substitute firstly (8) and (9) in (4). We obtain for both sides the following expression:

\[
\sum_{p_0 \cdots p_N, \sum r_i = m} \frac{1}{p_0!} \frac{1}{p_{N+2}!} \frac{1}{r_1!} \frac{1}{r_2!} \cdots \frac{1}{r_N!} \sum_{i=1}^{N+2} \frac{1}{p_i!} \frac{1}{p_{N+2}!} \frac{1}{r_1!} \frac{1}{r_2!} \cdots \frac{1}{r_N!} \left[ m \sum_{i=1}^{N+2} \frac{1}{p_i!} \frac{1}{p_{N+2}!} \frac{1}{r_1!} \frac{1}{r_2!} \cdots \frac{1}{r_N!} \left( \sum_{i=1}^{N+2} \frac{1}{p_i!} \frac{1}{p_{N+2}!} \frac{1}{r_1!} \frac{1}{r_2!} \cdots \frac{1}{r_N!} \right) \right]!!
\]

where we have taken into account that $\mu_{i_1 \cdots i_m}$ here is denoted with $\mu_{A_{i_1}}$ or with $\mu_{A_{i_1}}$ and so on up to $\mu_{A_{i_m}}$. Similarly, $\mu_{k_1 \cdots k_m}$ is denoted with $\mu_{A_{k_1}}$, or with $\mu_{A_{k_1}}$ and so on up to $\mu_{A_{k_m}}$. Consequently, the left hand side of (4)\textsubscript{3} causes, with respect to expression (9), a rise of one unity of the index $\rho_2$; similarly, the right hand side of (4)\textsubscript{3} causes, with respect to expression (8), a rise of one unity of the index $\rho_{m+1}$.

(ii) It is easy to verify that (9) satisfies (4)\textsubscript{1, 2, 4}.

(iii) By substituting (8) and (9) in (4)\textsubscript{3} in the case $m \geq 1$, we obtain for both sides the following expression:

\[
\sum_{p_0 \cdots p_N, \sum r_i = m} \frac{1}{p_0!} \frac{1}{p_{N+2}!} \frac{1}{r_1!} \frac{1}{r_2!} \cdots \frac{1}{r_N!} \left[ m \sum_{i=1}^{N+2} \frac{1}{p_i!} \frac{1}{p_{N+2}!} \frac{1}{r_1!} \frac{1}{r_2!} \cdots \frac{1}{r_N!} \left( \sum_{i=1}^{N+2} \frac{1}{p_i!} \frac{1}{p_{N+2}!} \frac{1}{r_1!} \frac{1}{r_2!} \cdots \frac{1}{r_N!} \right) \right]!!
\]

where we have taken into account that $\lambda_{i_1 \cdots i_m}$ here is denoted with $\lambda_{B_{i_1}}$ or with $\lambda_{B_{i_1}}$ and so on up to $\lambda_{B_{i_m}}$. Similarly, $\lambda_{k_1 \cdots k_m}$ is denoted with $\lambda_{B_{k_1}}$, or with $\lambda_{B_{k_1}}$ and so on up to $\lambda_{B_{k_m}}$. Consequently, the left hand side of (4)\textsubscript{3}, with $m \geq 1$ causes, with respect to expression (9), a rise of one unity of the index $r_m$; similarly, the right hand side of (4)\textsubscript{3}, with $m \geq 1$ causes, with respect to expression (8), a rise of one unity of the index $r_{m+1}$.

(iv) Let us conclude substituting (8) and (9) into (4)\textsubscript{3} in the case $m = 0$.

We have the same situation of the above case for its right hand side, while for its left hand side we have to take simply the derivative of (9) with respect to $\lambda$. More precisely, we obtain for both sides the following expression:

\[
\sum_{p_0 \cdots p_N, \sum r_i = m} \frac{1}{p_0!} \frac{1}{p_{N+2}!} \frac{1}{r_1!} \frac{1}{r_2!} \cdots \frac{1}{r_N!} \left[ m \sum_{i=1}^{N+2} \frac{1}{p_i!} \frac{1}{p_{N+2}!} \frac{1}{r_1!} \frac{1}{r_2!} \cdots \frac{1}{r_N!} \left( \sum_{i=1}^{N+2} \frac{1}{p_i!} \frac{1}{p_{N+2}!} \frac{1}{r_1!} \frac{1}{r_2!} \cdots \frac{1}{r_N!} \right) \right]!!
\]
We can insert a symmetrization over all the indexes of $\delta$; moreover, we have also substituted $\mu_{ij_{j_{n+1}}}$ with $\delta_{j_{n+1},j_{n+1}}$.

In the resulting expression we can insert a symmetrization over all the indexes of the set $j_{1} \cdots j_{n+1}$. For the product of the other cases, let us consider the shortened expression $\delta_{j_{1} \cdots j_{n+1}}(\delta_{k_{1} \cdots k_{n+1}}) \mu_{\mu_{1} \cdots \mu_{n+1}}$; here we can exchange the names of the indexes $j$ with those of $k$, so that the above shortened expression becomes $\delta_{k_{1} \cdots k_{n+1}}(\delta_{k_{1} \cdots k_{n+1}}) \mu_{k_{1} \cdots k_{n+1}}$.

Now we can exchange the indexes $k_{1} \cdots k_{n}$ with the indexes $\delta_{j_{1} \cdots j_{n+1}}$. Comparing this result with the expression which we started from, it is the same if we had exchanged the indexes $j_{n+1}$ and $k_{n+1}$. This completes the proof that the expression (14) remains the same if we exchange two indexes of the set $j_{1} \cdots j_{n+1}$. If we replace $\delta_{j_{1} \cdots j_{n+1}}$ with $\delta_{j_{1} \cdots j_{n+1}}$; so we can insert there a symmetrization over those indexes and (14) becomes
where underlined indexes denote symmetrization over these indexes. Now we observe that $(n + 1)p_{m+1}$ is exactly the number of the indexes of the set $j_1 \cdots j_{n+1} A_{n+1,1, \cdots A_{n+1,p_{m+1}}}$ and that, thanks to the summation $\sum_{m=0}^{N+1}$, the index near $i$ in $\delta A_{m+1}$ can be every index of the set $A_{1,1} \cdots A_{1,p} \cdots A_{m+1,1,p_{m+1}} \cdots A_{N+2,1} \cdots A_{N+2,p_{N+2}}$.

These facts allow rewriting (15) as

$$\sum_{n=0}^{N+1} (n + 1) \mu_{j_1 \cdots j_n} \frac{\partial h'}{\partial H_{j_1 \cdots j_n}} = \sum_{m=0}^{N+1} (n + 1) p_{m+1} \cdot \frac{1}{p_0!} \cdot \frac{1}{p_{N+1}!} \cdot \frac{1}{r_1!} \cdot \frac{1}{r_2!} \cdots \frac{1}{r_N!} \cdot \frac{1}{1 + \sum_{i=1}^{N+2} p_i + \sum_{j=1}^{N+1} i r_j} \cdot g_{\sum_{m=0}^{N+1} (i-2)p_i + \sum_{i=1}^{N+1} i r_i - 1} \left[ \frac{-1}{2\lambda} \right]^{\sum_{m=0}^{N+2} (j-1)p_j - \sum_{i=1}^{N+1} i r_i}$$

(ii) For the fourth term of (13) we can do similar passages (the difference is that we have $\lambda_-$ instead of $\mu_-$, $N \rightarrow 2$ instead of $N$, and $s$ instead of $n$); in this way that term becomes

$$\sum_{s=0}^{N-1} (s + 1) \lambda_{\delta h - h_s} \frac{\partial h'}{\partial \lambda_{h_{s-1} - h_s}} = \sum_{s=0}^{N-1} \frac{1}{(s + 1) r_{s+1}} \cdot \frac{1}{p_0!} \cdot \frac{1}{p_{N+1}!} \cdot \frac{1}{r_1!} \cdot \frac{1}{r_2!} \cdots \frac{1}{r_N!} \cdot \frac{1}{1 + \sum_{i=1}^{N+2} p_i + \sum_{j=1}^{N+1} i r_j} \cdot g_{\sum_{s=0}^{N-1} (i-2)p_i + \sum_{i=1}^{N+1} i r_i - 1} \left[ \frac{-1}{2\lambda} \right]^{\sum_{s=0}^{N+2} (j-1)p_j - \sum_{i=1}^{N+1} i r_i}$$

(iii) If we look at the last term of (13), we see that it can be written together with (16) and (17) and they become

$$\sum_{n=0}^{N+1} \frac{\partial h'}{\partial H_{j_1 \cdots j_n}} (n + 1) \mu_{j_1 \cdots j_n} + \sum_{s=0}^{N-1} \frac{\partial h'}{\partial \lambda_{h_{s-1} - h_s}} (s + 1) \lambda_{j_1 \cdots j_s} + H^s \delta = \left[ 1 + \sum_{n=0}^{N+1} (n + 1) p_{n+1} \cdot \frac{1}{p_0!} \cdot \frac{1}{p_{N+1}!} \cdot \frac{1}{r_1!} \cdot \frac{1}{r_2!} \cdots \frac{1}{r_N!} \cdot \frac{1}{1 + \sum_{i=1}^{N+2} p_i + \sum_{j=1}^{N+1} i r_j} \cdot g_{\sum_{m=0}^{N+1} (i-2)p_i + \sum_{i=1}^{N+1} i r_i - 1} \left[ \frac{-1}{2\lambda} \right]^{\sum_{m=0}^{N+2} (j-1)p_j - \sum_{i=1}^{N+1} i r_i} \right]$$

where also the index $k$ has been put under the symmetrization and we recall that a property of symmetrization is $\delta k \delta s = \delta n $ $\delta k = \delta (k-)$ and that, thanks to the square bracket at the beginning of the right hand side of (18) has become equal to the denominator of the half factorial of that right hand side.

(iv) For the second term of (13), we can use (8) and the derivation causes someway the rising of one unity of the index $p_2$; the result is

$$\frac{\partial^2 H}{\partial r_1 \partial r_{s+1}} \frac{\partial h'}{\partial h_{s-1} - h_s} (n + 1) \mu_{j_1 \cdots j_n} + \sum_{s=0}^{N-1} \frac{\partial h}{\partial \lambda_{h_{s-1} - h_s}} (s + 1) \lambda_{j_1 \cdots j_s} + H^s \delta = \left[ 1 + \sum_{n=0}^{N+1} (n + 1) p_{n+1} \cdot \frac{1}{p_0!} \cdot \frac{1}{p_{N+1}!} \cdot \frac{1}{r_1!} \cdot \frac{1}{r_2!} \cdots \frac{1}{r_N!} \cdot \frac{1}{1 + \sum_{i=1}^{N+2} p_i + \sum_{j=1}^{N+1} i r_j} \cdot g_{\sum_{m=0}^{N+1} (i-2)p_i + \sum_{i=1}^{N+1} i r_i - 1} \left[ \frac{-1}{2\lambda} \right]^{\sum_{m=0}^{N+2} (j-1)p_j - \sum_{i=1}^{N+1} i r_i} \right]$$
\[
\frac{\partial^2 \lambda}{\partial \eta_{k}^2} = 2 \lambda \sum_{p_0, p_1, \ldots, r_N} \frac{1}{p_0!} p_1! \ldots \frac{1}{p_{N+2}!} r_1! r_2! \\
\cdots \frac{1}{r_N!} \left[ \left( \sum_{i=1}^{N+2} i p_i \right) + \sum_{i=1}^{N+2} i r_i \right]!! \]

\[
\frac{\partial^2 \lambda}{\partial \eta_{k}^2} \sum_{2 \lambda} \left( \frac{-1}{2\lambda} \sum_{i=1}^{N+2} (3-\lambda) p_i - \sum_{i=1}^{N+2} i r_i \right) \]

(19)

It is true that with change of index we have in the summation the extra term with \( r_{s+2} = 0 \); but it does not effect the result for the presence of the coefficient \( r_{s+1} \).

If we look at the expressions (18), (19), and (21) we conclude that to prove (13) it is sufficient that the following relation holds:

\[
\frac{\partial^2 \lambda}{\partial \eta_{k}^2} \left( \frac{-1}{2\lambda} \sum_{i=1}^{N+2} (3-\lambda) p_i - \sum_{i=1}^{N+2} i r_i \right) = 0,
\]

and to prove this relation it is sufficient to prove that

\[
\frac{\partial}{\partial \lambda} \lambda \left( -\frac{1}{2\lambda} \right) + 2 \lambda \frac{\partial}{\partial \lambda} \left( -\frac{1}{2\lambda} \right) + 2 r_{s+1} \frac{\partial}{\partial r_{s+1}} \left( -\frac{1}{2\lambda} \right) = 0,
\]

(23)

where we have put \( r = r_1 + r_2 + \cdots + r_N \) and \( \delta = \frac{\partial}{\partial \lambda} \left( -\frac{1}{2\lambda} \right) g_{s+1}^{s+2} \). Obviously, (23) is an identity because we have

\[
\frac{\partial^2 \lambda}{\partial \eta_{k}^2} \left( -\frac{1}{2\lambda} \right) = -2 \lambda \frac{\partial}{\partial \lambda} \left( -\frac{1}{2\lambda} \right) - 2 r_{s+1} \frac{\partial}{\partial r_{s+1}} \left( -\frac{1}{2\lambda} \right).
\]

This concludes all the arguments which we had to prove.
5. Conclusions

We conclude that, in the particular case $N = 0$, we obtain the 11 moments' model and the solution here indicated is exactly that already obtained in literature with the macroscopic approach. In the case $N = 1$ we have that the present one is a particular solution for that already known in literature. So we have seen how it transfers itself also to higher values of $N$. In this way we are approaching the goal to obtain the most general solution and this will be the object of a future research.

Competing Interests

The authors declare that they have no competing interests.

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