DISSIPATION FLOW-FRAMES: PARTICLE, ENERGY, THERMOMETER

P. Ván\textsuperscript{1,2,3} and T.S. Biro\textsuperscript{1}

\textsuperscript{1}Dept. of Theoretical Physics, Wigner Research Centre for Physics, Institute for Particle and Nuclear Physics, H-1525 Budapest, Konkoly Thege Miklós út 29-33, Hungary;  
\textsuperscript{2}Dept. of Energy Engineering, Budapest Univ. of Technology and Economics, H-1111, Budapest, Műegyetem rkp. 3-9, Hungary;  
\textsuperscript{3}Montavid Thermodynamic Research Group  
Email: van.peter@wigner.mta.hu

ABSTRACT

We associate the following physical co-mover conditions to different frame choices: i) Eckart: particle flow, ii) Landau-Lifshitz: energy flow, iii) Jüttner: moving thermometer frame. The role of fixing a flow-frame is analysed with respect to local equilibrium concentrating on dissipative currents and forces in single component relativistic fluids. The special role of a "Jüttner frame" is explored and contrasted to the more common Eckart and Landau-Lifshitz choices.

1 Introduction

In dissipative theories of relativistic fluids we deal with four fundamental questions.

The first considers causality. Only divergence type theories are, in general, causal because there the symmetric hyperbolicity of the system of nonlinear evolution equations is established by construction\cite{1; 2; 3; 4; 5; 6}. The weaker version of causality requires for the symmetric hyperbolicity only for the linearized equations, and allows for characteristic speeds less than the speed of light\cite{7}. This weak causality was studied in the Israel-Stewart theory; numerous resulting inequalities are given in \cite{8}. From a physical point of view the causality of theories with parabolic differential equations should also be possible. In this case the validity of the continuum description is restricted by the characteristic maximal speeds\cite{9; 10; 11; 12}. A necessary condition for this type of restrictions requires the damping of the perturbations, equivalent to the linear stability of the theory\cite{13}.

The second question deals with generic stability. Generic stability is the linear stability of the homogeneous equilibrium solutions. The simplest relativistic generalization of the nonrelativistic Fourier-Navier-Stokes equations was proved to be unstable by Hiscock and Lindblom\cite{14}. In the sequel they formulated mathematical conditions of generic stability of the Israel-Stewart theory\cite{8} specified to the Eckart frame. However, due to the overwhelming complexity of these conditions they are not connected to reasonable properties of equations of state or transport coefficients. Since then several different propositions arose suggesting a first order theory, mostly motivated by the restoration of the generic stability\cite{15; 16; 12; 17; 18; 19; 20; 21}.

The third question is the correct distinction between ideal and dissipative fluids, especially in a relativistic context. It is customary to assume that perfect, nondissipative fluids are characterized by a special form of the energy-momentum tensor and the particle current density vector\cite{22; 23}. On the other hand physical dissipation is accompanied by nonzero entropy production. From this point of view there is a more extended family of perfect fluids beyond the customarily treated ones\cite{24}. These distinctions are technically addressed by the so called matching conditions\cite{25; 20; 21; 26}.

Finally the proper choice of flow-frames continues to be an unsettled question\cite{16}. One generally believes that in relativistic fluids the flow field $u^a$ can be chosen arbitrarily, since it is a somewhat vaguely defined physical property, belonging to the flow of volatile quantities, once the energy, once the conserved charge. In this situation it is customary to fix the flow either to the motion of particles (Eckart frame)\cite{27}, or that of the energy density (Landau-Lifshitz frame)\cite{28}. The flow fixing determines a continuous set of local rest frames in the fluid: we shall refer to the different choices of fixing the flow as flow-frames or frames. Contrary to the belief in a free choice of the flow-frame we point out that this may not be completely arbitrary, as one associates a given physical content of the dissipation to each. Further choices than the two classical ones should be preferred by demanding a given form of local Gibbs relations.

In this paper we present the general flow-frame, the separation of perfect and dissipative parts of energy-momentum and particle number current density and their relation to generic stability. The key theoretical aspect connecting these problems is relativistic thermodynamics. Our most important observation is that the usual assumption of kinetic equilibrium by introducing the velocity field parallel to the local thermometer and Lagrange multiplier field $\beta^a$ also appearing in the collision invariant $\psi = \alpha + \beta^a k_a$, already acts as a flow-frame fixing. This choice we tag as thermometer frame or Jüttner frame, distinguishing from the Eckart, Landau-Lifshitz and other conventions.

2 General one component dissipative relativistic fluids

In this paper the Lorentz form is given as $g^{ab} = \text{diag}(1, 1, 1, 1)$ and all indexes $a, b, c, \ldots$ run over 0, 1, 2, 3. We use natural units, $\hbar = k = c = 1$.

A single component fluid is characterized by the particle number density $n^a$, the particle current density $j^{ac}$, and the energy-momentum tensor $\epsilon^{abc}$.

The evolution equations for these quantities are obtained from the following action

\begin{align*}
S = \int d^4x \left\{ \frac{1}{2} \sqrt{\gamma} \left( g^{ab} \partial_a \phi \partial_b \phi + \epsilon^{abc} \partial_a j^{bc} \right) + \frac{1}{2} \sqrt{\gamma} \left( \gamma^{abcd} \partial_a \phi \partial_b \phi \right) \right\} + \int d^3x \left\{ \frac{1}{2} \sqrt{\gamma} \left( \gamma^{abc} \partial_a \phi \partial_b \phi \right) \right\}
\end{align*}

where $\phi$ is the fluid field, $\gamma$ is the determinant of the induced metric, and the dominant energy-momentum tensor is given by

\begin{align*}
\epsilon^{abcd} = n^a \epsilon_{abc} + j^{ac} j^{bd} - j^{ab} j^{cd} - \frac{1}{2} n^a \epsilon_{abc} \epsilon^{bd}
\end{align*}

The first term in the integrand is the relativistic Euler equation for the fluid field $\phi$, which is defined as the difference between the energy-momentum density and the volume form $\sqrt{\gamma}$. The second term is the relativistic equation of state, which is a function of the fluid field $\phi$ and the energy-momentum tensor $\epsilon^{abc}$.

The evolution equations for the fluid field $\phi$ are obtained from the following action

\begin{align*}
S = \int d^4x \left\{ \frac{1}{2} \sqrt{\gamma} \left( g^{ab} \partial_a \phi \partial_b \phi + \epsilon^{abc} \partial_a j^{bc} \right) + \frac{1}{2} \sqrt{\gamma} \left( \gamma^{abcd} \partial_a \phi \partial_b \phi \right) \right\} + \int d^3x \left\{ \frac{1}{2} \sqrt{\gamma} \left( \gamma^{abc} \partial_a \phi \partial_b \phi \right) \right\}
\end{align*}

where $\phi$ is the fluid field, $\gamma$ is the determinant of the induced metric, and the dominant energy-momentum tensor is given by

\begin{align*}
\epsilon^{abcd} = n^a \epsilon_{abc} + j^{ac} j^{bd} - j^{ab} j^{cd} - \frac{1}{2} n^a \epsilon_{abc} \epsilon^{bd}
\end{align*}

The first term in the integrand is the relativistic Euler equation for the fluid field $\phi$, which is defined as the difference between the energy-momentum density and the volume form $\sqrt{\gamma}$. The second term is the relativistic equation of state, which is a function of the fluid field $\phi$ and the energy-momentum tensor $\epsilon^{abc}$.
number four-vector \( N^a \) and the symmetric energy-momentum density tensor \( T^{ab} \). The velocity field of the fluid, the flow-frame \( u^a \), introduces a local rest frame and the basic fields \( N^a \) and \( T^{ab} \) can be expanded by their local rest frame components parallel and perpendicular to the flow:

\[
N^a = nu^a + j^a, \quad T^{ab} = eu^a u^b + q^{ab} + P^{ab}.
\]

Here \( n \) is the flow-frame particle number density, \( j^a \) is in this frame the non-convective particle number current density, \( e \) is the energy density, \( q^{ab} \) is the momentum density and \( P^{ab} \) is the pressure tensor. These components are flow-frame dependent, in particular \( j^a u_a = 0 \), \( q^{ab} u_{a} = 0 \) and \( P^{ab} u_b = 0 \). Introducing the substantial time derivative \( \frac{\partial}{\partial t} := u^a \partial_a \) denoted by and over-dot, the balances of the particle current density and energy-momentum are expressed by the local rest frame quantities:

\[
\partial_a N^a = \dot{n} + n \dot{u}^a u_a + \dot{\alpha}_a q^{ab} = 0, \quad \partial_b T^{ab} = \dot{e} u^a + e \dot{u}^a + e u^a \partial_a \dot{u}^b + \dot{\psi} + q^{ab} \partial_b u^a + u^a \partial_b q^{ab} = 0. \quad (3)
\]

The energy and momentum balances are the time and space-like components of the energy-momentum balance projected in the flow-frame:

\[
u_a \partial_b T^{ab} = \dot{e} + e \partial_b \dot{u}^b + u_a \dot{q}^{ab} + \dot{\alpha}_a q^{ab} - P^{ab} \partial_b u_a = 0, \quad \Delta_c \partial_b T^{cb} = \dot{e} u^a + \Delta_c \dot{q}^{ab} + q^{ab} \partial_b u^a + \dot{\psi} + q^{ab} \partial_b u^a + \Delta_c \partial_b P^{ab} = 0. \quad (5)
\]

The frame related quantities are important in the separation of the ideal and dissipative parts of the basic fields. This separation is best performed by analyzing the thermodynamical relations. In order to achieve this one postulates the existence of an additional vector field, the entropy current as a function of the local equilibrium values. There also may exist non-dissipative currents (presumably driven by non-dissipating forces, like the Lorentz force in magnetic fields). The thermodynamic approach aims at the separation of dissipative and non-dissipative local currents, in order to ensure the positivity of the expression (7). Physical freedom in the choice of a flow-frame should be restricted to different handlings of non-dissipative currents.

It is natural to introduce the Jüttner frame \( u^a_j \) defined by the direction of \( \dot{\beta}^a \) (thermometer motion):

\[
\psi^a_j = \frac{\dot{\beta}^a}{\sqrt{1 + \beta^a P_\beta}}
\]

In that frame the equilibrium fields are decomposed as:

\[
N^a = n_j \psi^a_j, \quad T^{ab} = e_j \psi^a_j \psi^b_j - P \Delta^{ab}, \quad S^a_0 = (\beta \psi_j - \alpha n_j) \psi^a_j.
\]

3 Thermodynamics of relativistic fluids – equilibrium

The concept of perfect fluids deals with the absence of dissipation, the entropy production vanishes:

\[
\Sigma_0 = \partial_a S^a_0 + \alpha \partial_a N^a - \beta_b \partial_a T^{ba}_0 = 0. \quad (8)
\]

The equilibrium entropy density \( S^a_0 \) is connected to the equilibrium particle number density \( N^a_0 \) and equilibrium energy-momentum density \( T^{ab}_0 \) by the following definition of the isotropic pressure:

\[
p_0 \dot{\beta}^a = S^a_0 + \alpha N^a_0 - \beta_b T^{ab}_0. \quad (9)
\]

Standard kinetic theory definitions and calculations satisfy the above expressions. Then \( \alpha \) and \( \dot{\beta}^a \) are coefficients in the collision invariant of the equilibrium distribution function, \( \psi = \alpha + \beta^a u^a \), and the pressure is that of an ideal gas \( p_0 = n_0 T \).

Kinetic theory describes a perfect fluid by the detailed balance requirement. Out of equilibrium dissipation can occur. In a dissipative fluid all physical quantities in principle deviate from their local equilibrium values. There also may exist non-dissipative currents (presumably driven by non-dissipating forces, like the Lorentz force in magnetic fields). The thermodynamic approach aims at the separation of dissipative and non-dissipative local currents, in order to ensure the positivity of the expression (7). Physical freedom in the choice of a flow-frame should be restricted to different handlings of non-dissipative currents.

where \( h_j = e_j + p_0 \) is the equilibrium enthalpy density in the Jüttner frame, and \( \beta_j = \beta^a u^a_j = 1/T \) is the reciprocal Jüttner temperature. \( \alpha, \beta_a \) and \( p_0 \) do not carry a frame index, because they are introduced before specifying the flow-frame. On the other hand the representations (11)-(13) are frame dependent. In a general flow-frame \( u^a \), that is not parallel to \( \dot{\beta}^a \), one can characterize this difference by introducing \( w^a = \dot{\beta}^a / (\dot{\beta}^b u_b) - u^a \). Then \( w^a \) is orthogonal to \( u^a \) (\( w^a u_a = 0 \)) and spacelike (\( w^a w_a = -w^2 \)). The Lagrange multiplier four-vector, \( \beta^a \), can be split as:

\[
\beta^a = \beta_j \psi^a_j = \beta (u^a + w^a), \quad (14)
\]

where \( \beta = \beta^a u_a \) is the reciprocal temperature in a general frame.
defined by \( u^a \). The equilibrium fields in this frame are given as

\[
N^a_0 = n_0 u^a + j^a_0, \quad (15)
\]

\[
T^{ab}_0 = e_0 u^a u^b + q^{ab}_0 u^b + \frac{q^{ab}_0}{h_0} - p \Delta^{ab} + \frac{q^{ab}_0}{h_0}, \quad (16)
\]

\[
S^a_0 = (\beta h_0 + \beta w_a q^a_0 - \alpha(\theta_0)) u^a + \beta q^a_0 - \alpha f^a_0. \quad (17)
\]

Here \( \beta = \beta_j / \sqrt{1 - w^2}, \quad n_0 = n_j / \sqrt{1 - w^2}, \quad e_0 = (e_j + pw^2) / (1 - w^2), \quad \alpha \) and \( p_0 \) does not change, \( \beta_0^a = n_0 w^a \) [24], (15)-(17) and (11)-(13) are the forms of the same equilibrium fields in the Jüttner and in the general frames respectively. In the specific equilibrium the Jüttner, Eckart and Landau-Lifshitz frames coincide, the different choices lead to the same condition: \( w_a = 0 \).

### 4 Thermodynamics of relativistic fluids – out of equilibrium

In classical non-equilibrium thermodynamics, without internal variables, one assumes that the gradients of the equilibrium fields characterize the deviation from local kinetic equilibrium. In that case the concept of local equilibrium is not modified. The internal variable theories, like the Israel-Stewart theory [33; 34; 25; 35; 36] or GENERIC [37; 38], choose a different characterization: local equilibrium is modified, some formerly dissipative currents appear among the state variables and as a consequence their contribution may reduce the entropy production. The relativistic theories revealed that the flow-frame fixing plays a special role in the specification of local equilibrium. It has been an observation of Planck and Einstein, that the momentum density (energy current density) is not purely dissipative and therefore in relativistic theories it has to be taken into account even in local equilibrium [39; 40].

Our starting point is the fundamental inequality of the second law (7). We introduce the following relation of the fields out of equilibrium, as a generalization of (9):

\[
S^a + \alpha \gamma^a - \beta_0 T^{ab} = \Phi^a. \quad (18)
\]

With a general \( \Phi^a \) this relation is valid without any restriction. In a general flow-frame, \( u^a \), we define the thermostatic pressure as:

\[
p = \frac{u_a \Phi^a}{\beta}. \quad (19)
\]

Therefore the general form of the potential \( \Phi^a \) can be written as

\[
\Phi^a = \frac{\beta p(u^a + \gamma^a)}{\beta}, \quad \text{where } u^a \gamma_a = 0. \quad (20)
\]

The parallel and perpendicular components of (18) to the flow \( u^a \) are

\[
s + \alpha \gamma - \beta(h + w_0 \gamma^b) = 0, \quad (21)
\]

\[
\alpha f^a - \beta(q^a + w_b \Pi^{ab}) + \beta w_a \gamma_a = 0, \quad (22)
\]

where \( h = e + p \) and \( \Pi^{ab} = P^{ab} + p \Delta^{ab} \) is the viscous pressure. Then we rewrite the entropy production (7) with flow related quantities:

\[
\Sigma = \partial_a S^a + \alpha \partial_a \gamma^a - \beta_0 \partial_a T^{ab} = -N^a \partial_a \alpha + T^{ab} \partial_a \beta_b + \partial_a \Phi^a = -n \alpha + \beta \gamma^a + \beta p + \Pi^{ab} \partial_a \beta_b - j^a \partial_a \alpha + \frac{n \alpha}{\beta} w_0 \alpha^b + \frac{n}{\beta} w_b \beta^a + \beta w_0 (q^b - h w^b) + \beta w_0 \gamma_a \partial_a w_b + \beta \partial a \gamma^b (\gamma^a - w_a) + \gamma^a \beta a \gamma^b. \quad (23)
\]

Thermodynamics is taken into account by the following two postulates.

1) The underlined part in the above expression with proper time derivatives (total differentials) is zero.

\[
\beta \frac{d}{dt} p = n \frac{d}{dt} \alpha - h \frac{d}{dt} \beta - q^a \frac{d}{dt} \gamma^a . \quad (24)
\]

This is the relativistic Gibbs-Duhem relation. Considering this together with the vanishing differential of (21), we obtain the Gibbs relation [41]:

\[
\beta (de + w_a \gamma^a) = ds + \alpha dn. \quad (25)
\]

Based on this result we conclude that the entropy has to be given by a functional relationship between the local densities (but certainly including the momentum density \( q^a \)), i.e. the proper relativistic and local equation of state is a function \( s(e, q^a, n) \). It has the following partial derivatives:

\[
\frac{\partial s}{\partial e} |_{q^a, n} = \beta, \quad \frac{\partial s}{\partial n} |_{e, q^a} = -\alpha, \quad \frac{\partial s}{\partial q^a} |_{e, n} = \beta w_a. \quad (26)
\]

identifying the thermodynamical entropic intensive parameters as being \( \beta, \alpha = \beta \mu \) and \( \beta \omega^a \). The four-vector \( w^a \) is constrained by its orthogonality to the local flow, so it contains independent information on a spatial three-vector only. In isotropic media this degree of freedom is reduced to the length of this vector, \( w_a \). In cases containing radiation it appears as a velocity parameter of the Doppler shift [41].

By utilizing the above functional form of the equation of state one derives that the pressure, the intensive parameter associated to mechanical work, satisfies the following four-vector generalized Gibbs-Duhem relation, now written by the traditional differentials:

\[
\beta \partial_a p = n \partial_a \alpha - h \partial_a \beta - q^a \partial_a (\beta w_b). \quad (27)
\]

2) Our second postulate is \( g^a = w^a \). By doing so we spell out the fundamental compatibility of non-equilibrium (18) with the equilibrium (9) definitions of pressure. In this way we treat the non-dissipative part of the thermodynamical potential, and with that the influence of the pressure gradient on the entropy production rate possibly closest to the ideal gas behavior. This is a special matching condition known from kinetic theory (\( \delta n = 0, \delta e = 0 \)): in this case the pressure four-vector \( \Phi^a \) is parallel to the thermoster vector \( \beta^a \).

Now a short calculation reduces (23) to a form collecting terms according to the gradients of intensives. A chemical diffusion part is associated to \( \partial_a \alpha \), a heat diffusion (Fourier-) part to \( \beta p \), and a viscous diffusion part to \( \Pi^{ab} \partial_a \beta_b \).
to the gradient of \( \beta \), and finally a viscosity term with the symmetric gradient tensor of the full four-vector \( \partial_{a} \beta_{b} \). We also gain one further term containing the gradient of the difference between \( u^{a} \) and \( w^{a} \). The antisymmetry of the multiplier enforces the antisymmetry of this velocity related gradient, therefore this term we tag as \(^{\prime\prime}\text{vorticity}^{\prime\prime}\). We arrive at the following expression:

\[
\Sigma = (nw^{a} - f^{a}) \partial_{a} \alpha + (q^{a} - hw^{a}) (\partial_{a} \beta + \beta \partial_{a} u_{b}) + (\Pi^{ab} - q^{a} w^{b}) \partial_{a} \beta_{b} + q^{[a} w^{b]} \partial_{a} (\beta (u_{b} - w_{b})) \geq 0. \tag{28}
\]

Here \( q^{[a} w^{b]} \) and \( q^{a} w^{b} \) denotes the symmetric and antisymmetric parts of \( q^{a} w^{b} \) respectively. (28) is the entropy production rate without fixing the flow-frame. For a perfect fluid, characterized by (15)-(16), the local entropy production is zero. Now it is straightforward to identify thermodynamic fluxes and forces and establish functional relationships, that are strictly linear in the first approximation:\(^2\):

| Fluxes  | Diffusive | Thermal | Viscous | Vortical |
|---------|-----------|---------|---------|----------|
| \( nw^{a} - f^{a} \) | \( q^{a} - hw^{a} \) | \( \Pi^{ab} - q^{a} w^{b} \) | \( q^{[a} w^{b]} \) |

Table 1. Thermodynamic fluxes and forces in a general flow frame.

Here \( \nabla_{a} = \Delta^{b}_{a} \partial_{b} \). The corresponding linear response relations for isotropic continua are:

\[
nw^{a} - f^{a} = D \nabla^{a} \alpha + \sigma (\nabla^{a} \beta + \beta \partial_{a} u_{b}), \tag{29}
\]

\[
q^{a} - hw^{a} = \sigma \nabla^{a} \alpha + \lambda (\nabla^{a} \beta + \beta \partial_{a} u_{b}), \tag{30}
\]

\[
\Pi^{ab} - q^{a} w^{b} = \zeta \Delta^{ab} \nabla^{c} \beta_{c} + 2 \eta \Delta^{[bc} \nabla^{d]} \beta_{c}, \tag{31}
\]

\[
q^{[a} w^{b]} = \chi \Delta^{[bc} \nabla^{d]} (\beta (u_{c} - w_{c})). \tag{32}
\]

Here \( \{ \} \) denotes the symmetric traceless part in the bracketed indices, \( \lambda \) is the heat conduction coefficient, \( D \) is the diffusion coefficient, \( \sigma \) is the Soret-Dufour coefficient of thermal diffusion. \( \zeta \) is the bulk viscosity, \( \eta \) is the shear viscosity, and \( \chi \) is the vortical viscosity coefficient. Because of the nonnegative entropy production (28) the linear transport coefficients must fulfill the following inequalities:

\[
D \geq 0, \quad \lambda \geq 0, \quad \lambda D - \sigma^{2} \geq 0, \quad \zeta \geq 0, \quad \eta \geq 0, \quad \chi > 0. \tag{33}
\]

Here the first three inequalities are coupled channel conditions for stability, while the last three are independent ones.

The procedure described here ensures the existence of a homogeneous flow field as a time independent solution of the equations of motion of the fluid. That is why deviation from local equilibrium is best characterized by gradients of the basic fields in the first approximation.

In the following we study some important particular choices for the flow-frame.

5 Thermometer frame

The thermometer or Jüttner frame is the natural choice in kinetic theory calculations. In this case the direction of \( \beta^{a} \) defines the flow-frame similarly to the natural frame in perfect fluids:

\[
\beta = \sqrt{\|b \beta^{b} \|^2} \quad \text{and} \quad u^{a} = \beta^{a} / \beta. \quad \text{In this section we apply this definition of the flow-frame. Then the local equilibrium relations are:}
\]

\[
s + \alpha n - \beta h = 0, \quad ds + \alpha dn - \beta de = 0. \tag{34}
\]

The entropy current density, \( J^{a} \) satisfies

\[
J^{a} + \alpha f^{a} - \beta q^{a} = 0^{a}, \tag{35}
\]

and the entropy production rate fulfills the inequality:

\[
\Sigma = - (\epsilon - f^{a} \partial_{a} \alpha + q^{a} (\partial_{a} \beta + \beta \partial_{a} u_{b}) + \beta \Pi^{ab} \partial_{a} u_{b} \geq 0. \tag{36}
\]

This form of the entropy production was derived originally by Eckart restricting to the case \( q^{a} = w^{a} = 0 \). Eckart identified the following thermodynamic fluxes and forces

| Fluxes  | Diffusive | Thermal | Mechanical |
|---------|-----------|---------|------------|
| \( nw^{a} - f^{a} \) | \( q^{a} - hw^{a} \) | \( \Pi^{ab} - q^{a} w^{b} \) | \( \Pi^{ab} \) |

Table 2. Thermodynamic fluxes and forces by Eckart.

Unfortunately in this case a generic instability occurs, the linear instability of the homogeneous equilibrium, as it was proved by Hiscock and Lindblom in [14]. Nonnegative entropy production is established only if considering the basic balance equations (for energy, momentum and further conserved Noether-charges) as constraints. However, by deriving (36) the balance of momentum does not enter the calculations. Therefore the linear relation between the thermal part of the fluxes and forces with the acceleration term, \( \beta \delta \partial^{a} \beta \), connects changes in these quantities irrespective to the momentum balance equation (6). A correct treatment of thermodynamic forces and fluxes on the other hand should introduce the momentum balance into the above entropy production formula. A short calculation leads to:

\[
\Sigma = \left( \frac{n}{h} q^{a} - f^{a} \right) \partial_{a} \alpha - \beta \partial_{a} \left( q_{a} + q_{a} w_{b} + q_{b} \partial_{b} u_{a} + \partial_{a} \Pi_{b}^{b} \right) + \beta \Pi^{ab} \partial_{a} u_{b} \geq 0. \tag{37}
\]

This step makes an important difference with respect to stability properties of the homogeneous equilibrium of a fluid. The corresponding thermodynamic fluxes and forces in the Jüttner frame are

| Fluxes  | Diffusive | Thermal | Mechanical |
|---------|-----------|---------|------------|
| \( nw^{a} - f^{a} \) | \( q^{a} - hw^{a} \) | \( \Pi^{ab} - q^{a} w^{b} \) | \( \Pi^{ab} \) |

Table 3. Thermodynamic fluxes and forces in Jüttner frame providing generic stability.

Here \( X^{a} \) is a convenient abbreviation for the thermal force, the thermodynamical force associated to the dissipative current of the heat. Linear transport relations for isotropic continua in the Jüttner frame can now be easily established:

\[
\frac{n}{h} q^{a} - f^{a} = D \nabla^{a} \alpha + \sigma X^{a}, \tag{38}
\]

\[
- \frac{\beta}{h} q^{a} = \sigma \nabla^{a} \alpha + \lambda X^{a}, \tag{39}
\]

\[
\beta \Pi^{ab} = \tau \Delta^{bc} \partial_{c} u^{f} + 2 \eta \Delta^{[bc} \nabla^{d]} u^{f}. \tag{40}
\]
With this modification the generic stability of the theory in Jüttner frame is established: the heat transfer vector $q^a$ receives a positive relaxation factor, $\beta/\hbar > 0$. It is easy to realize that by ignoring viscosity, component diffusion and cross effects, in homogeneous equilibrium, where all spacelike projected gradients of the velocity field vanish, the only surviving term in the thermal force is that with the total time derivative of the heat vector:

$$\lambda X^a = \lambda \Delta^{ab} q_b = -\frac{\beta}{\hbar} q^a. \quad (41)$$

Multiplied by $q^a$ this leads to a relaxation equation for the length of the vector, $Q = -q_a q^a$ as follows

$$\dot{Q} = -2\frac{\beta}{\hbar}\lambda Q. \quad (42)$$

This means a relaxation towards the $q^a = 0$ value of the energy current density.

An important property of these equations is the expected generic stability of the homogeneous equilibrium. Without the detailed calculations (to be shown elsewhere) we want to emphasize that the conditions of generic stability are purely thermodynamic. Namely, it is satisfied whenever the transport coefficients $\lambda, \eta$ are nonnegative and the following inequalities for thermodynamic stability i.e. the concavity of the entropy $s(e, n, q^a)$ are satisfied:

$$\partial_a T > 0, \quad \partial_a \frac{\beta}{T} > 0 \quad \partial_a T \partial_a \frac{\beta}{T} - \left( \frac{\partial_a T}{T} \right)^2 \geq 0. \quad (43)$$

### 6 Other flow-frames

The other flow-frames can be conveniently defined in our general framework.

The Eckart frame is defined by the direction of the particle current density vector $u^a = N^a / \sqrt{N_p N_a}$. One realizes that in case of dissipative fluids the Jüttner and Eckart frames do not coincide.

In case of a Landau-Lifshitz frame the flow field is defined by the direction of the momentum density vector $u^a = u^a T^b_{\mu}/||u^a T^b||$, therefore $q^a = 0^a$. In case of dissipative fluids the Jüttner and Landau-Lifshitz frames also do not coincide. However, in the absence of $q^a$, the thermodynamic relations are similar to the ones in a Jüttner frame

$$s + \alpha n - \beta h = 0, \quad \delta s + \alpha \delta n - \beta \delta e = 0, \quad \delta n \alpha - \delta h \beta - \delta p = 0. \quad (44)$$

In principle there are several further possibilities of frame fixing. One of them introduces $w^a = \beta^a / \hbar$. This choice fixes the velocity field compatible to some kinetic theory calculations [24; 18].

Once a choice of the linear response has been made, one can transform the description in one frame to the other. The different transport coefficients are not equivalent, a constant invariant coefficient may become flow-frame dependent in other frames. Wether the primary flow-frame independent choice is preferred or not requires further investigation.

### 7 Summary

Thermodynamic relations in relativistic fluids adhere to flow-frames, while dividing spacial homogeneous changes from the forces enforcing this homogeneity. It is made transparent in the train of thoughts from (18) to (23), where we calculated entropy production separating comoving time derivatives and spacial gradients. We have seen, that $\alpha, \beta^a = \beta (u^a + w^a)$, and $p$ are flow-frame independent. Then local equilibrium was postulated by the thermodynamic relation (25), containing homogeneous thermodynamics. In [41], presenting a different reasoning, we have shown that the different transformation formulas of the relativistic temperature, due to Planck-Einstein, Blanuša-Ott, Landsberg and Doppler, can be unified and reasonably explained in exactly this thermodynamic framework.

We propose that the thermometer frame, defined in (10), should be a preferred choice. In general $\beta^a$ can be divided into parts orthogonal and parallel to the flow $u$: $\beta^a = \beta (u^a + w^a)$, where $w^a w_a = 0$. We have revealed how far this choice differs from the Eckart and Landau-Lifshitz frames. There are arguments, that the widely used Landau-Lifshitz frame should be preferred [10; 42]. However, these studies do not distinguish the thermometer frame.

The entropy production in a general frame (28) helps to recognize

- that viscous pressure is damps the inhomogeneities in $\beta^a$,
- that there are perfect fluids with zero entropy production but $q^a/h = w^a \neq 0$ and $\Pi_{ab} = \hbar u^a w^b \neq 0$,
- there is a vorticity related dissipative term.

Furthermore we have mentioned, that generic stability is properly derived if the momentum balance constraint is also considered in the calculation of the entropy production (36).

In our previous works we have shown further examples of flow-frames. In [12; 17; 43; 41] the $w^a = q^a/h$ case was explored and in [24] and [18] we have analyzed the kinetic theory compatibility and thermodynamics when $w^a = q^a/h$. Thermodynamic considerations show, that the coupling of the momentum balance to the entropy production cannot be avoided [43]. It was proven independently of the Eckart and Landau-Lifshitz frame for $w^a = q^a/h$ in [17], for the $w^a = q^a/e$ case a partial proof was given in [18].

### 8 Acknowledgement

The work was supported by the grants Otka K81161, K104260 and TT 10-1-2011-0061/ZA-15-2009. The authors thank Etele Molnár for valuable discussions.

### REFERENCES

[1] R. Geroch. On hyperbolic "theories" of relativistic dissipative fluids. 2001. arXiv:gr-qc/0103112.

[2] R. Geroch. Relativistic theories of dissipative fluids. Journal of Mathematical Physics, 36(8):4226–4241, 1995.

[3] R. Geroch and L. Lindblom. Dissipative relativistic fluid theories of divergence type. Physical Review D, 41:1855–1861, 1990.

[4] R. Geroch and L. Lindblom. Causal theories of dissipative relativistic fluids. Annals of Physics, 207:394–416, 1991.

[5] I-S. Liu, I. Müller, and T. Ruggeri. Relativistic thermodynamics of gases. Annals of Physics, 169:191–219, 1986.
