Performance of the collective three-level quantum thermal engine

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We investigate the performance of a microscopic quantum heat engine consisting of $V$– or $\Lambda$–type emitters interacting collectively or independently when being in contact with environmental thermal reservoirs. Though the efficiency of a Carnot’s cycle is always higher than those associated with these setups, we have found that the performance of the cooperative $\Lambda$–type heat engine may be larger than that of the $V$–type under similar conditions. Cooperativity among the emitters plays an important role for the $\Lambda$–type setup, significantly improving its performance, while is less relevant for a $V$–type thermal engine. This is because the population inversion on the working atomic transition as well as its off-diagonal elements behave differently for these two atomic ensembles.

I. INTRODUCTION

Quantum heat engines, converting thermal energy into mechanical work, have attracted considerably attention since already many years [1, 2], also, in context of laws of quantum thermodynamics applied to small systems [3–7]. Based on those laws, it was demonstrated that the efficiencies of few-level thermal quantum engines are limited by the Carnot efficiency [1–8], though, one may go beyond the Carnot limit via squeezed thermal reservoirs [9]. A thorough analysis of a three-level system as amplifiers or attenuators was given in [10], while the electromagnetically induced transparency may be used to construct a quantum heat engines [11]. The quantum statistics of a single-atom heat engine was investigated too, in Ref. [12]. Substantial work was performed with respect to an experimental realization of a quantum thermal engine [13–17]. Particularly, it was demonstrated that quantum effects are responsible for enhancing the output power of a quantum microscopic heat engine compared to that of any classical one using the same resources [16, 17]. Earlier, it was emphasized the relevance of quantum effects in extracting work from a single thermal bath [18, 19].

The quantum behaviors of a microscopic thermal engine may change considerably if collective phenomena among its elements are taken into account [20, 21]. From this point of view, the role of entanglement in a small self-contained quantum refrigerator was investigated in [22], whereas the performances of quantum heat engines can be enhanced via collective interactions among many few-level emitters, used as working substance [23–27]. Furthermore, the output work may scale quadratically with the number of elements constituting the cooperative thermal engine [23, 24]. Collective effects greatly enhance the charging power of quantum batteries [28], while the quantum thermometer, that is, the precision of the temperature estimation, improves for larger spin ensembles too [29]. An ensemble of indistinguishable quantum machines can give rise as well to a genuine quantum enhancement of the collective thermodynamic performance [30]. Moreover, a quantum Otto cycle in which the medium, an interacting ultracold gas, is driven between a superfluid and an insulating phase can outperform similar single particle cycles [31].

Thus, motivated by the recent substantial progress towards this issue, we investigate here the quantum performance of a microscopic thermal engine composed generally from $N$ three-level $V$– or $\Lambda$–type emitters. More precisely, the working subsystem may consist of single or multiple atoms interacting independently or collectively, in the Dicke’s sense [32–37], via the surrounding thermal reservoirs. Particularly, the hot bath acts on the $|3\rangle \leftrightarrow |1\rangle$ transition, while the cold one on the $|2\rangle \leftrightarrow |1\rangle$ atomic transition, respectively, as it is depicted in Fig. 1, in analogy with the two heat reservoirs forming the part of a macroscopic classical motor operating between two thermal baths. To close the cycle, a weak and coherent electromagnetic field is applied on the $|3\rangle \leftrightarrow |2\rangle$ transition, converting the incoherent thermal energy of the heat reservoirs to an output work in an inverted populated atomic medium. Therefore, we are interested in the performance of this process in a collective heat engine. Also, we have assumed that the external applied coherent field modifies insignificantly the sample’s steady-state achieved due to the environmental thermal reservoirs alone. This fact allows us to find the steady-state solutions of the corresponding master equations describing the collective three-level samples and, actually, in this way one can investigate the population inversion on the working transition and the performance of such a cooperative quantum thermal engine. We have found that generally the performance of a heat engine formed from an independent $V$–type atomic ensemble is larger than that when cooperativity among the emitters would be relevant. On the other hand, the corresponding performance for collectively interacting $\Lambda$–type emitters is always larger than that for independent atoms. Furthermore, in similar conditions, the quantum performance for a $\Lambda$–type ensemble is larger than the one for a $V$–type atomic sample. However, their efficiencies are smaller than the efficiency of the Carnot cycle.

The article is organized as follows. In Sec. II we describe the analytical approach and the collective population dynamics of the system of interest, while in Sec. III

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we analyze the corresponding quantum performance of the cooperative three-level microscopic heat engine. The summary is given in Sec. IV.

II. COLLECTIVE POPULATION DYNAMICS OF THE MICROSCOPIC THREE-LEVEL THERMAL ENGINE

The master equation describing a collection of $N$ three-level $V$-type ensembles, see Fig. 1(a) and Appendix A, interacting with a weak external coherent electromagnetic field on the transition $|3\rangle \leftrightarrow |2\rangle$, with the hot thermal bath on the $|3\rangle \leftrightarrow |1\rangle$ transition, while the cold one is acting on the $|2\rangle \leftrightarrow |1\rangle$ atomic transition, respectively, in the Born-Markov approximations \[\text{[33–37]}, \text{is:}\]

$$\frac{d}{dt} \rho(t) + i\Omega[S_{32} + S_{23}, \rho] = -\sum_{\alpha \in \{1, 2, 3\}} \frac{\gamma_\alpha}{2} (1 + \bar{n}_\alpha) \{[S_{1\alpha}, S_{1\alpha}^\dagger] + H.c.\} - \sum_{\alpha \in \{1, 2, 3\}} \frac{\gamma_\alpha n_\alpha}{2} \{[S_{1\alpha}, S_{1\alpha}^\dagger] + H.c.\}. \tag{1}$$

Here the collective operators $S_{\alpha\beta} = \sum_{j=1}^{N} S^{(j)}_{\alpha\beta}$, with $\{\alpha, \beta \in 1, 2, 3\}$ and $S^{(j)}_{\alpha\beta} = |\alpha\rangle_j \langle \beta|$], satisfy the commutation relations $[S_{\alpha\beta}, S_{\gamma\delta}^\dagger] = \delta_{\beta\gamma} S_{\alpha\delta} - \delta_{\alpha\gamma} S_{\beta\delta}$. The single-atom spontaneous decay rate on transition $|\alpha\rangle \rightarrow |1\rangle$ is given by $\gamma_\alpha$, $\{\alpha \in 2, 3\}$, whereas the corresponding mean thermal photon number due to the environmental thermal reservoirs, at temperature $T_\alpha$, is given by the following expression: $\bar{n}_\alpha = \exp(\frac{\hbar \omega_\alpha}{k_B T_\alpha}) - 1^{-1}$, with $k_B$ being the Boltzmann’s constant and $\omega_\alpha = \omega_\beta - \omega_\gamma \Omega$, considered real, is the corresponding Rabi frequency.

Since the external coherent electromagnetic field applied on the working transition $|3\rangle \leftrightarrow |2\rangle$ is considered weak, that is $\Omega \ll \{N_2^3, N_3^3\}$, the final steady state is determined mainly by the heat reservoirs. Therefore, the steady-state solution of the master equation 1, when $\Omega = 0$, is given by the following expression:

$$\rho_s = Z^{-1} e^{-\xi_3 S_{32}} e^{-\xi_2 S_{23}}, \tag{2}$$

where $Z$ is determined by the requirement $Tr\{\rho_s\} = 1$. The substitution of the steady-state solution 2 in the Eq. 1 results in

$$\xi_\alpha = \ln\left(\frac{1 + \bar{n}_\alpha}{n_\alpha}\right) \equiv -\ln(\eta_\alpha), \tag{3}$$

where $\eta_\alpha = n_\alpha / (1 + n_\alpha) < 1, \{\alpha \in 2, 3\}$. Notice that the master equation describing a collection of $\Lambda$-type three-level emitters, see Fig. 1(b), can be obtained from Eq. 1 by swapping the two indices of each transition operator having $\alpha$ or $\beta$ as one of the indices, e.g., $S_{\alpha\beta} \leftrightarrow S_{\beta\alpha}$. The corresponding steady-state solution is given by Eq. 2 with, however, $\xi_\alpha = \ln(\eta_\alpha), \{\alpha \in 2, 3\} \[38, 39\].

The expectation values of any diagonal statistical moments are obtained by introducing the coherent atomic states $|N, n, m\rangle$ corresponding to the $su(3)$ algebra of the operators $S_{\alpha\beta} \[38, 39\]$. The state $|N, n, m\rangle$ denotes a symmetric collective state of $N$ particles with $n$ atoms in bare state $|1\rangle, (m-n)$ in bare state $|2\rangle$ and $(N-m)$ atoms in bare state $|3\rangle$. For a given $N$, the admissible values of $\{n, m\}$ are $n = 0, 1, 2, \cdots, N$ while $m = n, n + 1, \cdots, N$. Thus, the expectation values of the collective population operators, $\langle S_{\alpha\alpha} \rangle = -Z^{-1} (\partial/\partial \xi_\alpha) Z, \{\alpha \in 2, 3\}$, with

$$Z = \sum_{n=0}^{N} \sum_{m=n}^{N} e^{-\xi_2 (m-n)} e^{-\xi_3 (N-m)},$$

needed in the subsequent discussion of the sample’s efficiency for a $V$-type three-level ensemble, can be obtained explicitly as:

$$\langle S_{11} \rangle = \frac{1}{\eta_3 - 1} \left(\frac{(n_2 - n_3)}{(\eta_2 - \eta_3)}(1 + N - n_2(N + 2) + n_2^{N+2}) \right) + \frac{(n_2 - n_3)}{(\eta_2 - \eta_3)}(1 + N - n_2(N + 2) + n_2^{N+2}) \right), \tag{4}$$

and $\langle S_{11} \rangle + \langle S_{22} \rangle + \langle S_{33} \rangle = N$. From the above expressions follow that if $n_2 = n_3$ then $\langle S_{33} \rangle = \langle S_{22} \rangle$. Importantly for the incoming discussions,

$$\langle S_{33} \rangle > \langle S_{22} \rangle, \text{ only when } \eta_3 > \eta_2. \tag{5}$$

If $\eta_3 = 0$, then $\langle S_{33} \rangle = 0$, while

$$\langle S_{22} \rangle = \frac{n_2(1 - n_2^N) - N(1 - n_2)n_2^{N+1}}{(1 - n_2)(1 - n_2^{N+1})}, \tag{6}$$

that is, we have recovered the corresponding result for a collection of two-level atoms, $|2\rangle \leftrightarrow |1\rangle$, in a thermal reservoir \[40\].

![Figure 1](image-url)
The corresponding collective steady-state population dynamics for a $\Lambda$-type three-level ensemble is given as follows:

\[
\begin{align*}
\langle S_{11} \rangle &= \frac{\eta_2 \eta_3 (\eta_1^N A_N(\eta_2, \eta_3) - (1 - \eta_3)^2 \eta_1^N)}{(1 - \eta_2)(1 - \eta_3)D_N(\eta_2, \eta_3)}, \\
\langle S_{22} \rangle &= \frac{\eta_3 (\eta_1^N B_N(\eta_2, \eta_3) + (1 - \eta_2)^2 \eta_1^{N+1})}{(1 - \eta_2)(\eta_3 - \eta_2)D_N(\eta_2, \eta_3)},
\end{align*}
\]

where $\langle S_{33} \rangle = N - \langle S_{22} \rangle - \langle S_{11} \rangle$. Here, $A_N(\eta_2, \eta_3) = (1 - \eta_3)^2 + \eta_3^2(2 - \eta_2 + N(1 - \eta_2)(1 - \eta_3) - \eta_2(\eta_3 - \eta_2))$, $B_N(\eta_2, \eta_3) = (1 - \eta_3)((2 + N)\eta_2^2 + N(1 - \eta_2)(1 + N)(1 + \eta_3)) - (\eta_2 - \eta_3)^2 \eta_3^{N+1}$, and $D_N(\eta_2, \eta_3) = (1 - \eta_3)\eta_3^{N+1} + (\eta_2 - 1)\eta_2^{N+1} + (\eta_3 - \eta_2)(\eta_2\eta_3)^{N+1}$. Some limiting cases of the steady-state population dynamics, given by (7), are as follows: when $\eta_2 \to 0$ or $\eta_3 \to 0$ then $\langle S_{11} \rangle \to 0$. If $\eta_2 \to 0$ while $\eta_3 \neq 0$, then $\langle S_{22} \rangle \to N$. Conversely, if $\eta_3 \to 0$ while $\eta_2 \neq 0$, then $\langle S_{33} \rangle \to N$. Again here, if $\eta_2 = \eta_3$ then $\langle S_{22} \rangle = \langle S_{33} \rangle$, whereas

\[
\langle S_{22} \rangle > \langle S_{33} \rangle \text{ only if } \eta_3 > \eta_2.
\]

The cooperative population dynamics given by the expressions (4) thru (8) will be analyzed, in the next Section, in the context of the performance of the microscopic thermal engine composed of individual or $N$ three-level emitters, interacting independently or collectively.

**III. THE PERFORMANCE OF THE COLLECTIVE QUANTUM THERMAL ENGINE**

The maximal quantum efficiency of the microscopic collective thermal engine for both samples (see Fig. 1), described in the previous Section, is given by the ratio of the energy emitted as work to the energy absorbed from the hot reservoir $\eta_{\omega} \beta_0$, namely,

\[
\varepsilon = \frac{\omega_{32}}{\omega_{31}}.
\]

This efficiency realizes for the transition cycle $[1] \to [3] \to [2] \to [1]$, for $V$-type ensembles, and $[3] \to [1] \to [2] \to [3]$ for $\Lambda$-type emitters, respectively, see Fig. 1. However, in the steady-state, the performance will be governed by the mean values of the populations in these states as well as the induced coherences. Therefore, the output generated work (power output) will be proportional to the polarization induced by the applied weak electromagnetic field on the working transition $[3] \leftrightarrow [2]$. Particularly, for a $V$-type ensemble the output work is $P_a = i\hbar \Omega(\omega_{31}\gamma_{31} + \omega_{21}\gamma_{22} + \omega_{23}\gamma_{23}) = i\hbar \Omega(\omega_{31} - \omega_{21})(\langle S_{31} \rangle - \langle S_{21} \rangle)$, whereas for $\Lambda$-type emitters is $P_a = i\hbar \Omega(\omega_{31}\gamma_{31} + \omega_{21}\gamma_{22} + \omega_{23}\gamma_{23}) = i\hbar \Omega(\omega_{31} - \omega_{21})(\langle S_{31} \rangle - \langle S_{21} \rangle)$, see e.g. (10). In the following, we shall relate it with the population inversion between working levels, $\{[3], [2]\}$, and investigate the performance of the microscopic three-level heat engine for individual atoms as well as for independently or collectively interacting emitters, respectively.

Thus, for a $V$-type atomic ensemble, we have

\[
\frac{d}{dt} \langle S_{23} \rangle = i\Omega(\langle S_{33} \rangle - \langle S_{22} \rangle) - \gamma_a V \langle S_{23} \rangle / 2
\]

\[
- (\gamma_2 + \gamma_3)(\langle S_{13} \rangle \langle S_{21} \rangle / 2) / 2,
\]

where $\gamma_a V = \gamma_2(1 + \bar{n}_2) + \gamma_3(1 + \bar{n}_3)$, while for the $\Lambda$-type emitters one has:

\[
\frac{d}{dt} \langle S_{32} \rangle = i\Omega(\langle S_{22} \rangle - \langle S_{33} \rangle) - \gamma_{3\Lambda} \langle S_{32} \rangle / 2
\]

\[
+ (\gamma_2 + \gamma_3)(\langle S_{13} \rangle \langle S_{21} \rangle / 2),
\]

with $\gamma_{3\Lambda} = \gamma_2 \bar{n}_2 + \gamma_3 \bar{n}_3$. One can observe that the last terms of Eqs. (10,11) account for the collective effects among the three-level emitters. For larger atomic ensembles, i.e. $N \gg 1$, one can decouple the collective correlator $\langle S_{13} \rangle \langle S_{31} \rangle$, $\alpha \neq \beta$, as follows $\langle S_{13} \rangle \approx \langle S_{11} \rangle \langle S_{33} \rangle$. The decoupling is valid as long as the fluctuations of the population in the $[1]$ state, i.e. $\sqrt{\Delta S_{11}} = \sqrt{\langle S_{11}^2 \rangle - \langle S_{11} \rangle^2}/N^2$, which may scale as $\sqrt{\Delta S_{11}} \sim N^{-1/2}$, are negligible. This will allow us to obtain the following steady-state expressions for the off-diagonal elements:

\[
\langle S_{23} \rangle = \frac{i\Omega}{\Gamma_V} (\langle S_{33} \rangle - \langle S_{22} \rangle),
\]

(12)

for a $V$-type ensemble, and

\[
\langle S_{32} \rangle = \frac{i\Omega}{\Gamma_\Lambda} (\langle S_{22} \rangle - \langle S_{33} \rangle),
\]

(13)

for $\Lambda$-type emitters, respectively. Here,

\[
\Gamma_V = (\gamma_a V + \gamma_c) / 2,
\]

(14)

\[
\Gamma_\Lambda = (\gamma_{3\Lambda} - \gamma_c) / 2,
\]

with $\gamma_c = (\gamma_2 + \gamma_3)\langle S_{11} \rangle$ describing the collective contribution. A relevant aspect here is that both the population differences as well as $\gamma_{3\Lambda}$ depend on the number of emitters when they interact collectively. For single- or many-independent emitters, $\gamma_c = 0$. Therefore, in order to clarify the role played by the cooperativity among the three-level atoms with respect to the output generated work of a many-particle quantum heat engine or its efficiency, in the following we investigate the corresponding steady-state population quantum dynamics entering in Exps. (12,13).

For the sake of comparison, firstly, we shall discuss the efficiency of the quantum three-level engine for a single-emitter or an independent atomic ensemble [1] and, then, for collectively interacting emitters. As we have mentioned in the previous Section, the population quantum dynamics in the steady-state is mainly due to the environmental thermal reservoirs, because of the weakness of the applied external electromagnetic field, i.e. $\Omega \ll \{N\gamma_2, N\gamma_3\}$. For a single $V$-type atomic system, from Eqs. (14) when $N = 1$, one obtains:

\[
\langle S_{33} \rangle = \frac{\eta_3}{1 + \eta_2 + \eta_3}, \quad \text{and} \quad \langle S_{22} \rangle = \frac{\eta_2}{1 + \eta_2 + \eta_3},
\]

(15)
FIG. 2: The scaled steady-state inversion operator \((\langle S_{33} \rangle - \langle S_{22} \rangle)/N\), for \(V\)-type atomic ensembles, as a function of \(\eta_2\) when (a) \(\eta_3 = 0.8\) and (b) \(\eta_3 = 0.9\), respectively. The dashed curves describe the situation of a single-atom or independent atomic ensembles, whereas the solid ones depict the case of collectively interacting emitters with \(N = 100\). Notice that \(0 \leq \eta_2 \leq 0.99\) here.

while their ratio is, respectively,

\[
\frac{\langle S_{33} \rangle}{\langle S_{22} \rangle} = \frac{\eta_3}{\eta_2}
\]

(16)

Since \(\eta_\alpha = \exp\left[-\hbar \omega_{\alpha 1}/(k_B T_\alpha)\right]\), \(\{\alpha = 3, 2\}\), one then has that:

\[
\frac{\langle S_{33} \rangle}{\langle S_{22} \rangle} = \exp\left[\frac{\hbar \omega_{32}}{k_B T_2}(\varepsilon_C/\varepsilon - 1)\right],
\]

where

\[
\varepsilon_C = 1 - \frac{T_2}{T_3},
\]

is the Carnot efficiency. Population inversion, i.e. \(\langle S_{33} \rangle > \langle S_{22} \rangle\), occurs only if \(\eta_3 > \eta_2\) meaning that \(\omega_{21}/2T_2 > \omega_{31}/2T_3\) or, equivalently (see the population ratio (17)),

\[
\varepsilon < \varepsilon_C.
\]

(19)

Correspondingly, for a single \(\Lambda\)-type atom, see Fig. 1(b), from Eqs. (7) when \(N = 1\), we have:

\[
\langle S_{22} \rangle = \frac{\eta_3}{\eta_2 + \eta_3 + \eta_2 \eta_3}, \quad \text{and} \quad \langle S_{33} \rangle = \frac{\eta_2}{\eta_2 + \eta_3 + \eta_2 \eta_3},
\]

(20)

FIG. 3: The scaled output work, proportional to the imaginary part of \((\langle S_{23} \rangle)/N\) [in units of \(\Omega/\gamma\)], given by Eq. (12) for \(V\)-type atomic ensembles, as a function of \(\eta_2\) when (a) \(\eta_3 = 0.8\) and (b) \(\eta_3 = 0.9\), respectively. The dashed curves describe the situation of independent atomic ensembles, whereas the solid ones depict the case of collectively interacting emitters. The solid lines were plotted by multiplying with a factor equal to 20. Here, \(\gamma_2 = \gamma_3 \equiv \gamma\), \(N = 100\) while \(\eta_2 \leq 0.99\).

FIG. 4: The scaled steady-state inversion operator \((\langle S_{22} \rangle - \langle S_{33} \rangle)/N\), for \(\Lambda\)-type atomic ensembles, as a function of \(\eta_2\) when (a) \(\eta_3 = 0.8\) and (b) \(\eta_3 = 0.9\), respectively. Other parameters are as in Fig 2.

with, however,

\[
\frac{\langle S_{22} \rangle}{\langle S_{33} \rangle} = \frac{\eta_3}{\eta_2}
\]

(21)

or [1]

\[
\frac{\langle S_{22} \rangle}{\langle S_{33} \rangle} = \exp\left[\frac{\hbar \omega_{23}}{k_B T_2}(\varepsilon_C/\varepsilon - 1)\right],
\]

(22)

Here, the population inversion, i.e. \(\langle S_{22} \rangle > \langle S_{33} \rangle\), happens only if \(\eta_3 > \eta_2\) or \(\omega_{12}/T_2 > \omega_{13}/T_3\) which, again, implies that \(\varepsilon < \varepsilon_C\), see the ratio (22). Notice that for an independent \(V\)- or \(\Lambda\)-type atomic ensemble one should multiply the expressions (15) and (20) by \(N\), with \(\sum_\alpha \langle S_{\alpha \alpha} \rangle = N\), \(\{\alpha = 1, 2, 3\}\). Thus, concluding this part, the efficiency of such a \(V\)- or \(\Lambda\)-type heat engine consisting of individual atoms or an ensemble of independent emitters will be always smaller than that given by the Carnot cycle, i.e., \(\varepsilon_{ind} < \varepsilon_C\). Furthermore, the output work of the heat engine for \(\Lambda\)-type single or independent emitters is generally higher than the corresponding work for \(V\)-type emitters, in similar conditions (follow the incoming discussions). This is because the population inversion for \(\Lambda\) atoms is larger than for \(V\) emitters, compare the dashed curves of Fig. 2(a) and Fig. 2(a) as well as of Fig. 2(b) and Fig. 3(b).

In what follows, we shall focus on a collectively interacting \(V\)- or \(\Lambda\)-type three-level ensemble. According to

FIG. 5: The scaled output work, proportional to the imaginary part of \((\langle S_{23} \rangle)/N\) [in units of \(\Omega/\gamma\)], given by Eq. (13) for \(\Lambda\)-type atomic ensembles, as a function of \(\eta_2\) when (a) \(\eta_3 = 0.8\) and (b) \(\eta_3 = 0.9\), respectively. The dashed lines describe the situation of independent atomic ensembles, whereas the solid ones depict the case of collectively interacting emitters. Here, \(\gamma_2 = \gamma_3 \equiv \gamma\), \(N = 100\) while \(\eta_2 \leq 0.99\).
Eqs. (17), the population dynamics of a collection of $N$ three-level emitters depends on $\{n_2, n_3\}$ in a sophisticated way, but not on their ratio. Therefore, simple expressions similar to those given by (12) would not be possible for collectively interacting emitters. On the other side, certainly, $\langle S_{33}\rangle > \langle S_{22}\rangle$, for collectively interacting emitters in $V$–type ensembles, while $\langle S_{22}\rangle > \langle S_{33}\rangle$, for collectively interacting atoms in $\Lambda$–type samples, only if $n_3 > n_2$ as it was emphasized in the previous Section. Since the ratio $n_3/n_2$ is given by Exps. (17,22), one can conjecture then that the efficiency of the collective three-level thermal engine, $\varepsilon_{\text{col}}$, is smaller than the efficiency of the Carnot cycle, i.e., $\varepsilon_{\text{col}} < \varepsilon_C$. While the efficiency of the three-level $V$– or $\Lambda$–type thermal engine is smaller than that of the Carnot’s cycle, regardless of the interaction nature among the emitters, the performances of these samples differ if atoms are considered independent or collectively interacting. To elucidate this issue, in Figure 2, for a $V$–type atomic sample, we plot the scaled inversion operator $\langle (S_{33}) - (S_{22}) \rangle / N$ as a function of $n_2$, while $n_3$ being fixed. The dashed lines depict the case of independent atoms whereas the solid ones are for collectively interacting emitters. Depending on the strength of the hot bath, i.e. $n_3$, the inversion can be larger for independent or collectively interacting atoms, compare Fig. 2(a) and Fig. 2(b). However, for collectively interacting emitters, the output work is inversely proportional to the population into the state $|1\rangle$, see Exps. (12,14). Therefore, in Figure 3, we plot the scaled output work, proportional to the imaginary positive part of $\langle S_{23}\rangle / N$ given by Exp. (12), for independently or collectively interacting $V$–type emitters. By inspecting this figure, one can observe that generally the output work of a $V$–type heat engine is larger for an ensemble of independently interacting emitters rather than if would interact collectively, and it linearly depends on the number of emitters (note that the collective curves were obtained by multiplying with 20). Also, in order to focus on the influence of the thermal environmental reservoirs only, we have considered that $\gamma_2 = \gamma_3 = \gamma$. On the other hand, Figure 4 depicts the steady-state behaviors of the scaled inversion operator corresponding to the $\Lambda$–type ensemble, i.e. $\langle (S_{22}) - (S_{33}) \rangle / N$, as a function of the strength of the cold bath $n_2$ and for different values of $n_3$. In contrast to the $V$–type three-level ensembles, for $\Lambda$–types ones the positive inversion is always higher for collectively interacting emitters, compare Fig. 2 and Fig. 4, as well as the dashed and solid lines of Fig. 4, respectively. Again here, the output work is inversely proportional to the population in the state $|1\rangle$, see Exps. (13,14). Actually, for collectively interacting $\Lambda$–type emitters this state is less populated for $n_3 < n_2 < 1$ and $N \gg 1$. Remarkable here, $\Gamma_\Lambda$ is smaller for collectively interacting ensembles, see Exp. (14), meaning that we have an increase in the output work from this reason as well as because the inversion enhances too, see Exp. (13) and Fig. (4). So, Fig. 5 shows the scaled output work, proportional to the imaginary and positive part of $\langle S_{32}\rangle / N$ given by Eq. (13) for independently or collectively interacting $\Lambda$–type atomic ensembles. We observe here a significant enhance of the output work generated for collectively interacting emitters, compared to the independent atoms case. Thus, the performance of a thermal engine based on the collectively interacting $\Lambda$–type atoms is larger than for a similar ensemble consisting of independently interacting emitters. Moreover, in a $\Lambda$–type ensemble one can almost completely transfer the population to the state $|2\rangle$, see Fig. (4) when $n_2 < n_3$, meaning that in similar conditions the performance of a thermal quantum heat engine consisting of $\Lambda$–type three-level emitters would be higher than the performance of a thermal engine formed from $V$–type three-level emitters, compare Fig. (4) and Fig. 5, correspondingly.

Thus, generalizing here, a microscopic $\Lambda$–type three-level quantum heat engine may have an advantage over a similar one formed, respectively, of an ensemble of $V$–type three-level emitters. Furthermore, the output work of a cooperative $\Lambda$–type thermal quantum engine, i.e. its performance, is larger than that of a heat engine consisting from single or independently interacting $\Lambda$–atoms. Actually, it is greater than the one for single, independently or collectively interacting $V$–type three-level emitters, in similar conditions.

IV. SUMMARY

Summarizing, we have investigated the efficiency and performance of a microscopic three-level quantum heat engine and have elucidated the role the collectivity among the emitters plays with respect to this issue. Since in a cooperative $\Lambda$–type ensemble one can more efficiently create population inversion on the involved working atomic transition, the energy conversion of the incoherent thermal reservoirs towards the coherently applied electromagnetic field probing that transition is highly improved compared to an independent atomic ensemble or to an independently or collectively interacting $V$–type emitters of a microscopic heat engine, under identical conditions. Furthermore, the quantum efficiency of a Carnot cycle is always better than that characterizing these setups, regardless of the cooperativity among emitters.

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Appendix A: Master equation for an ensemble of \(V\)-type three-level emitters

The Master equation describing an arbitrary collection of \(N\) three-level \(V\)-type emitters interacting with their environmental reservoirs as well as with a coherent electromagnetic field is, see Fig. 1(a):

\[
\frac{d}{dt}\rho(t) + i \sum_{j=1}^{N} \Omega_j \left[ S_{j2}^{(j)} e^{i \mathbf{k}_L \cdot \mathbf{r}_j} + S_{j3}^{(j)} e^{-i \mathbf{k}_L \cdot \mathbf{r}_j}, \rho \right] + i \sum_{j \neq l=1}^{N} \sum_{\alpha \in \{2, 3\}} \frac{\gamma_\alpha \Omega_{jl}^{(\alpha)}}{2} \left[ S_{\alpha 1}^{(j)} S_{l1}^{(l)} \rho \right] = - \sum_{j,l=1}^{N} \sum_{\alpha \in \{2, 3\}} \frac{\gamma_\alpha \lambda_{jl}^{(\alpha)}}{2} \left( 1 + \tilde{n}_\alpha \right) \left[ S_{\alpha 1}^{(j)} S_{l1}^{(l)} \rho \right] + \tilde{n}_\alpha \left[ S_{l1}^{(j)} S_{\alpha 1}^{(l)} \rho \right] + H.c.. \tag{A1}
\]

For dipole-allowed transitions one has that \(\lambda_{jl}^{(\alpha)} = \sin(2\pi r_{jl}/\lambda_{\alpha 1})/(2\pi r_{jl}/\lambda_{\alpha 1})\) and \(\Omega_{jl}^{(\alpha)} = \cos(2\pi r_{jl}/\lambda_{\alpha 1})/(2\pi r_{jl}/\lambda_{\alpha 1})\), \(\alpha \in \{2, 3\}\), where we have averaged over all dipole orientations, whereas \(r_{jl} = |\mathbf{r}_j - \mathbf{r}_l|\) are the inter-particle intervals between the \(j\)th and the \(l\)th emitters, respectively. \(\lambda_{\alpha 1}\) is the wavelength of the photon emitted on the \(|\alpha\rangle \rightarrow |1\rangle\) transition, respectively. Further, \(S_{\alpha j}^{(j)} = |\alpha\rangle_{j}\langle \beta|\), with \((\alpha, \beta) \in \{1, 2, 3\}\), represents the population of the state \(|\alpha\rangle\) in the \(j\)-th atom, if \(\alpha = \beta\), or the transition operator from \(|\beta\rangle\) to \(|\alpha\rangle\) of the \(j\)-th atom when \(\alpha \neq \beta\). The atomic operators obey the commutation relations:

\[
\left[ S_{\alpha \beta}^{(j)}, S_{\gamma \alpha}^{(l)} \right] = \delta_{\beta \gamma} \hat{S}_{\alpha \alpha}^{(j)} - \delta_{\alpha \gamma} \hat{S}_{\beta \beta}^{(j)}. \tag{A2}
\]

The dipole-dipole interaction term, proportional to \(\Omega_{ddj}^{(\alpha)}\), was not included in the master equation since we anticipated that it will not influence the established steady state, i.e. the steady-state solution commutes with \(\hat{S}_{\alpha 1} S_{1\alpha}\), \(\alpha \in \{2, 3\}\). This also means that \(\Omega_{ddj}^{(\alpha)} < N\gamma_\alpha\), \(\alpha \in \{2, 3\}\), that is, the dipole-dipole caused shift among multiple multi-particle states should be smaller than the collective decay rates. As well, in the coherent pumping term the exponent factor \(e^{\pm i \mathbf{k}_L \cdot \mathbf{r}_j} \rightarrow 1\) when the wavelength \(\lambda_L\) of the external coherent pumping source is bigger than the ensemble's size, i.e., the emitters are in an equivalent position with respect to the coherent driving. Finally, if \(j = l = 1\), one obtains from Eq. (A1) the corresponding master equation for a single \(V\)-type atom. The master equation for independent emitters, i.e. if \(r_{jl}/\lambda_{\alpha 1} \gg 1\), can be obtained from (A1) when \(j = l\) and \(\lambda_{jj}^{(\alpha)} \rightarrow 1\), while \(\Omega_{jj}^{(\alpha)} \lambda_{jj}^{(\alpha)} \rightarrow 0\), \(\alpha \in \{2, 3\}\).

In the same way can be analyzed the respective master equation describing \(\Lambda\)-type emitters, see Fig. 1(b).

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