Parameterisation space for Cornell potential in a QCD potential model

Krishna Kingkar Pathak1, Satyadeep Bhattacharya2, Tapashi Das3,a
1 Department of Physics, Arya Vidyapeeth College, Guwahati 781016, India
2 Impact Hub Institute for Competitive Examination, Dimapur, Nagaland 797112, India
3 Department of Physics, Madhab Choudhury College, Barpeta 781301, India

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Abstract We make a critical analysis on the free parameters of the Cornell potential \(-\frac{4\alpha_s}{r^3} + br + c\) and provide a parameterisation space for the strong coupling constant \(\alpha_s\) and the constant shift ‘c’ for choosing linear part as perturbation in the potential model. In the analysis of heavy-light mesons (D, Ds, B, Bs and Bc), we have found a wide range of values for the coupling constant i.e \(0.20 \leq \alpha_s \leq 0.64\) with \(-1.2 \leq c \leq -0.66\) which can be used to treat the confining part as perturbation.

1 Introduction

QCD potential between a quark and anti quark has been the first important ingredient of phenomenological model to study Hadron Physics. The potential model is found to be successful in providing both qualitative and quantitative description of hadron spectrum and its decay modes. To deal with a potential model, the choice of the correct QCD potential is the most important ingredient for its success. There are several acceptable potentials in QCD, which is based upon the two important facts of QCD i.e. the confinement and asymptotic freedom of quarks. To study quark-anti quark bound states, some of the accepted and commonly used potentials are, Cornell potential [1]: \(-\frac{4\alpha_s}{r^3} + Br + c\), Power law potential [2,3]: \(-Ar^\alpha + Dr^\beta + c\), Logarithmic potential [4]: \(A + B ln r\), Richardson potential [5]:\(Ar - \frac{B}{rln \frac{1}{r}}\) etc.

However, the potential parameters in the different potentials as well as in the different models are found to vary within a noticeable range. For example, the value of ‘c’ is found to vary from model to model. In the work of relativistic quark model, Faustov et al. has used \(c = -0.3\) GeV [6], Eichten et al. in Ref. [7] has taken it to be \(c = 0.50805\) GeV, Mao Zhi Yang has taken \(c = -0.19\) GeV [8], Scora and Isgur considered \(c = -0.81\) GeV in Ref. [9], whereas Grant and Rosner [10] considered a large negative value for \(c = -1.305\) GeV in a power law potential. The value of the strong running coupling constant \(\alpha_s\) appearing in the QCD potential has also a wide range from 0.22 to 0.64 in different theoretical works as well as in the different schemes like \(\overline{MS}, V\)-scheme etc. [11–14].

In this work, we review our previous work [15,16] put forward some comments on the perturbation theory of choosing the linear part of the Cornell potential as perturbation and provide a parameterisation space for the strong coupling constant \(\alpha_s\) and the constant shift c. While taking the Coulomb part as parent and linear as perturbation one important point is to be noted that for this choice, the perturbation is possible only for very small value of ‘b’. In infinite mass limit, the value of ‘b’ is \(\leq 0.03\) GeV2 [15] which are much smaller than the value of ‘b’ in charmonium spectroscopy (\(\approx 0.183\) GeV2).

The paper is organized as follows, in the Sect. 2, we describe the Cornell potential and its parameters. In Sect. 3, we describe the QCD potential model with linear part as perturbation. In Sect. 4, we discuss the constraints on the strong coupling constant \(\alpha_s\) and \(c\). Section 5 contains the conclusions.

2 The Cornell potential and perturbation

For mesons, the one gluon exchange contribution between quark and antiquark is given by the coulombic potential,

\[ V(r) = -\frac{4\alpha_s}{3r}. \]  

(1)

Here, \(-\frac{4}{3}\) is due to the SU(3) color factor and \(\alpha_s\) is the strong coupling constant.
For large distance, we have confinement of quarks. The confining potential is given by
\[ V(r) = br, \]  
(2)
where \( b \) is known as confinement parameter and phenomenologically \( b = 0.183 \text{ GeV}^2 \) [16].

Considering the two effects, Cornell potential is written as:
\[ V(r) = -\frac{4\alpha_s}{3r} + br + c \]  
(3)
which is the sum of coulombic and linear potential with a scale factor ‘c’. Both the potentials play decisive role in the quark dynamics and their separation is not possible. Besides there is no appropriate small parameter so that one of the potential within a perturbation theory can be made perturbative.

With the Cornell Potential, one cannot solve the Schrodinger equation in quantum mechanics except for some simple models. Therefore, physicists opt for developing efficient approximate methods. Perturbation theory is one of the helpful tools to get an approximate wavefunction with the Cornell potential. In fact, perturbation theory is considered to be one of the approximate methods which most appeals to intuition.

However, in perturbation theory, one has to check the convergence of the series which appears in the procedure. If the rate of convergence of the perturbation series is sufficiently high then we may expect accurate results.

The advantage of taking Cornell potential for study is that it leads naturally two choices of “parent” Hamiltonian, one based on the Coulomb part and the other on the linear term, which can be usefully compared. It is expected that, in choosing the perturbative part of the potential a dominant role is played by critical \( r_0 \) where the potential \( V(r) = 0 \). Aitchison and Dudek in Ref. [17] put an argument that if the size of a state measured by \( r < r_0 \), then the Coulomb part as the “Parent” will perform better and if not so the linear part as “parent” will perform better. The Aitchison’s work also showed the results that with Coulombic part as perturbation(VIPT), bottomonium spectra are well explained than Charmonium where as Charmonium states are well explained with linear part as parent. It becomes noteworthy in this context that the critical distance \( r_0 \) is not a constant and can be enhanced by reducing \( b \) and \( c \) or by increasing \( \alpha_s \).

3 The QCD potential model

In this work, a specific potential model with linear part as perturbation is taken into consideration. For completeness and proper reference we put the last modified version of the wavefunction with coulombic part as parent as in Refs. [16, 18, 19].

The non-relativistic predictions of potential models with a non-relativistic Hamiltonian for the heavy-light and heavy-heavy mesons are found to be in fair agreements with the updated theoretical, experimental and lattice results. Hence, we start with the ground state \((l = 0)\) spin independent non-relativistic Fermi–Breit Hamiltonian (without the contact term),
\[ H = -\frac{\nabla^2}{2\mu} - \frac{4\alpha_s}{3r} + br + c. \]  
(4)

With the linear term ‘\( br + c \)’ as perturbation and using Dalgarno method, the wave function in the model is obtained as [16, 18–20]
\[ \psi_{rel+conf}(r) = \frac{N'}{\sqrt{\pi a_0^3}} e^{-\frac{r}{a_0}} \left(C' - \frac{\mu ba_0 r^2}{2}\right) \left(\frac{r}{a_0}\right)^{-\epsilon} , \]  
(5)
where
\[ N' = \frac{2^{2\frac{1}{2}}}{\sqrt{(2^{2\frac{1}{2}}\Gamma(3-2\epsilon)\Gamma^2 - \frac{4}{5^{\frac{1}{2}}}\mu b a_0^2 \Gamma(5-2\epsilon)\Gamma^2 + \frac{1}{3^\frac{1}{2}}\mu^2 b^2 a_0^2 \Gamma(7-2\epsilon)}}} \]  
(6)
\[ C' = 1 + c A_0 \sqrt{\pi a_0^3}, \]  
(7)
\[ \mu = \frac{m_i m_j}{m_i + m_j}, \]  
(8)
\[ a_0 = \left(\frac{\frac{4}{3} \mu \alpha_s}{\mu}\right)^{-1}, \]  
(9)
\[ \epsilon = 1 - \sqrt{\left(\frac{4}{\frac{5}{3} \alpha_s}\right)} ^2 . \]  
(10)

The QCD potential is taken as
\[ V(r) = -\frac{4}{3r} \alpha_s + br + c. \]  
(11)

Here \( A_0 \) is the undetermined factor appearing in the series solution of the Schrödinger equation. The term \( \left(\frac{r}{a_0}\right)^{-\epsilon} \) in Eq. (5) is the Dirac factor and was introduced to incorporate relativistic effect [19, 21, 22]. It is to be noted that the factor \( \left(\frac{r}{a_0}\right)^{-\epsilon} \) to the wave function (5) develops a singularity at \( r \rightarrow 0 \) and the masses of scalar mesons become negative, therefore the term can be omitted from the wave function as discussed in Ref. [12].

Similarly, the wave function with coulombic part ‘\(-\frac{4\alpha_s}{3r} + c\)’ as perturbation and linear part as parent (up to O(4)) is given by [23],
\[ \psi_{rel+lin}(r) = \frac{N''}{r} \left[1 + A_1(r)r + A_2(r)r^2 + A_3(r)r^3 \right] 
+ A_4(r)r^4 \right] A_1[r \rho_1 + \rho_0] \left(\frac{r}{a_0}\right)^{-\epsilon} \]  
(12)
where \( A_i[r] \) is the Airy function [24] and \( N'' \) is the normalization constant

\[
N'' = \frac{1}{\left[ \int_0^\infty 4\pi \left[ 1 + A_1(r) + A_2(r)r^2 + A_3(r)r^3 + A_4(r)r^4 \right]^2 (A_i[\rho_1 + \rho_0])^2 \left( \frac{r}{m} \right)^{-2e} dr \right]^{1/2}}.
\] (13)

The co-efficients of the series solution as occurred in Dal-garno’s method of perturbation, are the functions of \( \alpha_s, \mu, b \) and \( c \):

\[
A_1 = \frac{-2\mu^2 A_i}{2\rho_1 k_1 + \rho_1^2 k_2}
\]
(14)

\[
A_2 = \frac{-2\mu (W^1 - c)}{2 + 4\rho_1 k_1 + \rho_1^2 k_2}
\]
(15)

\[
A_3 = \frac{-2\mu W^0 A_i}{6 + 6\rho_1 k_1 + \rho_1^2 k_2}
\]
(16)

\[
A_4 = \frac{-2\mu W^0 A_2 + 2\mu b A_1}{12 + 8\rho_1 k_1 + \rho_1^2 k_2}
\]
(17)

The different parameters are given by,

\[
\rho_1 = (2\mu b)^{1/2}
\]
(18)

\[
\rho_0 = -\left[ \frac{3\pi(4n - 1)}{8} \right]^{1/2}.
\]
(19)

(In our case \( n = 1 \) for ground state)

\[
k_1 = 1 + \frac{k}{r}
\]
(20)

\[
k = \frac{0.3550281 - (0.2588194)\rho_0}{(0.2588194)\rho_1}
\]
(21)

\[
k_2 = \frac{k^2}{r^2}
\]
(22)

\[
W^1 = \int \psi^{(0)\ast} H^\ast \psi^{(0)} d\tau
\]
(23)

\[
W^0 = \int \psi^{(0)\ast} H_0 \psi^{(0)} d\tau.
\]
(24)

Let us now discuss the constraints on the parameters \( \alpha_s \) and \( c \) in the model.

4 Constraints from two points of view

To evaluate a narrow range of the free parameter in the model, we consider the two constraints; one from the expectation value and the other from the convergence point of view in the model.

4.1 From the condition of expectation value

Considering the argument of Aitchison and Dudek [17] \( \langle r \rangle < r_0 \) to treat the linear part as perturbation, we should get the size of a state measured by \( \langle r \rangle < r_0 \), where \( r_0 \) is the critical distance at which \( V(r_0) = 0 \). Now,

\[
\langle r \rangle_{coul} = \int \psi^\ast r \psi dr = \frac{3\alpha_0}{2} = r_1 (\text{say})
\]

and the critical distance \( r_0 \) at which \( V(r_0) = 0 \) can be obtained by the relation

\[
br_0^2 + cr_0 - \frac{4\alpha_s}{3} = 0.
\]

The variation of \( r_1 \) and \( r_0 \) with the model parameters can be easily studied from the above relations and the results are tabulated in Table 1. From the results it is clear that to treat linear part as perturbation, the value of upper scale of \( \alpha_s \) is different for different mesons. The maximum value for \( B \) meson is found to be \( \alpha_s = 0.64 \) beyond which the condition \( \langle r \rangle < r_0 \) invalid.

4.2 From the convergence point of view

From the momentum transform , we see that for a lower cut-off value of \( Q_0^2 \), either one has to consider a very small value of \( b \) or to increase the value of \( \alpha_s \), which is obvious, since in both the cases coulombic part will be more dominant. However, reality condition of \( \epsilon \) from Eq. (10) demands that \( \alpha_s \leq \frac{3}{4} \) and hence one can not go beyond \( \alpha_s = 0.75 \) in this approach.

From the convergence point of view, the perturbative condition demands [20]

\[
\frac{(4 - \epsilon)(3 - \epsilon)\mu b a_0^3}{2(1 + a_0^2 Q^2)} \ll C'.
\]
(25)

For a positive cut off \( Q_0^2 \), we can write

\[
\frac{(4 - \epsilon)(3 - \epsilon)\mu b a_0^3}{2(1 + a_0^2 Q_0^2)} = C'.
\]
(26)

The values of \( Q_0^2 \) with \( b = 0.183 \) GeV\(^2\) for \( B \) and\( D \) mesons are shown in Table 2.

Thus to incorporate lower value of \( Q^2 \) (\( Q^2 \leq \Lambda^2_{QCD} \)), with linear part as perturbation, one expects a bound of \( \alpha_s \leq 0.64 \).
Table 1  Values of $r_1$ and $r_0$ for different mesons with $m_{u/d} = 0.33$ GeV, $m_s = 0.483$ GeV, $m_c = 1.55$ GeV, $m_b = 4.97$ GeV, $b = 0.183$ GeV$^2$ and $cA_0 = 1$ GeV

| Mesons | $\alpha_s = 0.20$ | $\alpha_s = 0.36$ | $\alpha_s = 0.45$ | $\alpha_s = 0.58$ | $\alpha_s = 0.60$ | $\alpha_s = 0.64$ |
|--------|-----------------|-----------------|-----------------|-----------------|-----------------|-----------------|
|        | $r_1$ | $r_0$ | $r_1$ | $r_0$ | $r_1$ | $r_0$ | $r_1$ | $r_0$ | $r_1$ | $r_0$ | $r_1$ | $r_0$ |
| $D(c\bar{u}/c\bar{d})$ | – | – | – | – | 6.990 | 7.148 | 6.757 | 7.167 | 6.334 | 7.204 |
| $D(c\bar{s})$ | – | – | – | – | 5.422 | 6.151 | 5.241 | 6.172 | – | – |
| $B(\bar{b}u/\bar{b}d)$ | – | – | – | – | 6.143 | 6.151 | 5.938 | 6.172 | – | – |
| $B_s(\bar{b}s)$ | – | – | – | – | 5.898 | 6.010 | 4.576 | 6.151 | – | – |
| $B_c(\bar{b}c)$ | 4.803 | 5.719 | 2.668 | 5.908 | – | – | – | – | – | – |

Table 2  Allowed range of $\alpha_s$ from the limit of $Q^2_0$ in the Model

| Mesons | $\alpha_s = 0.20$ | $\alpha_s = 0.36$ | $\alpha_s = 0.45$ | $\alpha_s = 0.58$ | $\alpha_s = 0.60$ | $\alpha_s = 0.64$ |
|--------|-----------------|-----------------|-----------------|-----------------|-----------------|-----------------|
|        | $r_1$ | $r_0$ | $r_1$ | $r_0$ | $r_1$ | $r_0$ | $r_1$ | $r_0$ | $r_1$ | $r_0$ |
| $D(\mu = 0.2774$ GeV) | – | – | – | – | 0.01409 | 0.01037 | 0.00204 |
| $D_s(\mu = 0.3576$ GeV) | – | – | – | – | 0.00939 | 0.00322 | – |
| $B(\mu = 0.3157$ GeV) | – | – | – | – | 0.0125 | 0.00775 | – |
| $B_s(\mu = 0.4238$ GeV) | – | – | – | – | 0.04480 | 0.00098 | – |
| $B_c(\mu = 1.171$ GeV) | 0.26378 | 0.092543 | – | – | – | – |

Table 3  Allowed range of $\alpha_s$ and $c$ for different mesons under the constraints

| Mesons | $\alpha_s$ | $c$ |
|--------|------------|-----|
|        | $D(\mu = 0.2774$ GeV) | 0.570–0.640 | $\leq -1.2$ |
|        | $D_s(\mu = 0.3576$ GeV) | 0.575–0.610 | -0.860 to -0.785 |
|        | $B(\mu = 0.3157$ GeV) | 0.580–0.629 | -0.998 to -0.887 |
|        | $B_s(\mu = 0.4238$ GeV) | 0.450–0.582 | -0.997 to -0.663 |
|        | $B_c(\mu = 1.171$ GeV) | 0.200–0.409 | -0.994 to -0.197 |

Fig. 1  Range of $\alpha_s$ and $c$ for $D, D_s, B, B_S$ and $B_c$
4.3 Constraints on $\alpha_s$

In the analysis, we further see that the value of $\alpha_s$ as well as the model parameter ‘c’ also play a crucial role in choosing the parent and perturbative terms. From the above two constraints we make an individual range of $\alpha_s$ for different heavy light mesons and tabulate in Table 3. The range of ‘$\alpha_s$’ and ‘c’ within the two constraints are further specified in Fig. 1.

5 Conclusions

In this work, we mainly focus in finding the analytical conditions to treat the linear part of the Cornell potential as perturbation. In the analysis we consider two constraints and evaluate a parameterisation space for $\alpha_s$ and c in the range of $\alpha_s$ is found to be $0.20 \leq \alpha_s \leq 0.64$ with $-1.2 \leq c \leq -0.66$. We further note that the positive value of $c$ as is used in [7] is excluded in the model and other values of $\alpha_s$ and ‘c’ as is used in different phenomenological works [6–8,18–20] are found to be valid within the parameterisation space provided in Fig. 1. Further it is to be noted that the allowed space for $B_s$ meson is too large in comparison to other heavy-light mesons $D, D_s, B, B_s$ which may be due to the constituents of two heavy quarks and needs a detailed study to be carried out.

However with linear part as perturbation, if the value of $\alpha_s$ in the above range is taken to be granted, then with the same potential another possibility of considering the coulombic part as perturbation also arises for a value of $\alpha_s \leq 0.20$.

Data Availability Statement This manuscript has no associated data or the data will not be deposited. [Authors’ comment: The data can be deposited.]

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