Influence Functions in Deep Learning Are Fragile

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Abstract

Influence functions approximate the effect of training samples in test-time predictions and have a wide variety of applications in machine learning interpretability and uncertainty estimation. A commonly-used (first-order) influence function can be implemented efficiently as a post-hoc method requiring access only to the gradients and Hessian of the model. For linear models, influence functions are well-defined due to the convexity of the underlying loss function and are generally accurate even across difficult settings where model changes are fairly large such as estimating group influences. Influence functions, however, are not well-understood in the context of deep learning with non-convex loss functions. In this paper, we provide a comprehensive and large-scale empirical study of successes and failures of influence functions in neural network models trained on datasets such as Iris, MNIST, CIFAR-10 and ImageNet. Through our extensive experiments, we show that the network architecture, its depth and width, as well as the extent of model parameterization and regularization techniques have strong effects in the accuracy of influence functions. In particular, we find that (i) influence estimates are fairly accurate for shallow networks, while for deeper networks the estimates are often erroneous; (ii) for certain network architectures and datasets, training with weight-decay regularization is important to get high-quality influence estimates; and (iii) the accuracy of influence estimates can vary significantly depending on the examined test points. These results suggest that in general influence functions in deep learning are fragile and call for developing improved influence estimation methods to mitigate these issues in non-convex setups.

1 Introduction

In machine learning, influence functions [1,2] can be used to estimate the change in model parameters when the empirical weight distribution of the training samples is perturbed by an infinitesimal amount. This approximation is cheaper to compute compared to the expensive process of repeatedly re-training the model to retrieve the exact parameter changes. Influence functions could thus be used to understand the effect of removing an individual training point (or, groups of training samples) on the model predictions at the test-time. Leveraging a first-order Taylor’s approximation of the loss function, [3] has shown that a (first-order) influence function, computed using the gradient and the Hessian of the loss function, can be useful to interpret machine learning models, fix mislabelled training samples and create data poisoning attacks.

Influence functions are in general well-defined and studied for models such as logistic regression or smooth SVM [3], where the underlying loss-function is convex. For convex loss functions, influence functions are also accurate even when the model perturbations are fairly large (e.g. in the group influence case [4,5]). However, when the convexity assumption of the underlying loss function is violated, which is the case in deep learning, the behaviour of influence functions is not well understood and is still an open area of research. With recent advances in computer vision [6], natural...
language processing \cite{7}, high-stakes applications such as medicine \cite{8} and finance \cite{9}, it has become particularly important to interpret deep model predictions. This makes it critical to understand influence functions in the context of deep learning, which is the main focus of our paper.

Although \cite{3} has developed an influence function in the convex case, it has provided some successful empirical results of the application of the influence function in a non-convex case as well. In particular, \cite{3} has shown that for a small CNN with around 2,600 parameters trained on 10\% of MNIST, influence functions are fairly accurate. However, it has remained an open question if a successful application of influence functions can be generalized to different deep model architectures and practical input datasets such as full MNIST, CIFAR-10 and ImageNet. In this paper, we aim to shed some light on this question.

In the case of deep networks, several factors might have an impact on influence estimates: (i) due to non-convexity of the loss function, different initializations of the perturbed model can lead to significantly different model parameters (with approximately similar loss values); (ii) even if the initialization of the model is fixed, the curvature values of the network (i.e. eigenvalues of the Hessian matrix) at optimal model parameters might be very large in very deep networks, leading to a significant Taylor’s approximation error of the loss function and thus resulting in poor influence estimates; (iii) for large neural networks, computing the exact inverse-Hessian Vector product, required in computation of influence estimates, can be computationally very expensive. Thus, one needs to use approximate inverse-Hessian Vector product techniques which might be erroneous; resulting in low quality influence estimates; and finally (iv) different architectures can have different loss landscape geometries near the optimal model parameters, leading to varying influence estimates.

In this paper, we study aforementioned issues of using influence functions in deep learning through an extensive experimental study on progressively-growing complex models and datasets. We first start our analysis with a case study of a small neural network for the Iris dataset where the exact Hessian matrix can be computed. We then progressively increase the complexity of the network and analyse a CNN architecture (depth of 6) trained on 10\% of MNIST dataset, similar to \cite{3}. Next, we evaluate the accuracy of influence estimates for more complex deep architectures (e.g. ResNets) trained on MNIST and CIFAR-10. Finally, we compute influence estimates on the ImageNet dataset using ResNet-50.

We make the following observations through our analysis:

- We find that the network depth and width have a strong impact on influence estimates. In particular, we show that influence estimates are fairly accurate when the network is shallow, while for deeper models, influence estimates are often erroneous. We attribute this partially to the increasing curvature values of the network as the depth increases.
- We observe that the weight decay regularization is important to obtain high quality influence estimates in certain architectures and datasets.
- We show that the inverse-Hessian Vector product approximation techniques such as stochastic estimation \cite{10} are erroneous, especially when the network is deep. This can contribute to the low quality of influence estimates in deep models.
- We observe that the choice of test-point has a significant impact on the quality of influence estimates, across different datasets and architectures.
- In very large-scale datasets such as ImageNet, we have found that even ground-truth influence estimates (obtained by leave-one-out re-training) can be inaccurate and noisy partially due to the model’s training and convergence.

These results highlight sensitivity of current influence functions in deep learning and call for developing robust influence estimators to be used in large-scale machine learning applications.

2 Related Works

Influence functions are primarily used to identify important training samples for test-time predictions and debug machine learning models \cite{3}. In recent times, there is an increase in the applications of influence functions for tasks other than interpretability. For e.g. \cite{11} has used influence functions to audit the reliability of predictions by producing confidence intervals for test-time predictions. In NLP, influence functions have been used to detect biases in word-embeddings \cite{12} whereas in the domain of
3 Basics of Influence Function

Consider \( h \) to be a function parameterized by \( \theta \) which maps from an input feature space \( \mathcal{X} \) to an output space denoted by \( \mathcal{Y} \). The training samples are denoted by the set \( \mathcal{S} = \{ z_i : (x_i, y_i) \}_{i=1}^{n} \), while the loss function is represented by \( \ell(h_{\theta}(z)) \) for a particular training example \( z \). The standard empirical risk minimization solves the following optimization problem:

\[
\theta^* = \arg \min_{\theta} \frac{1}{n} \sum_{i=1}^{n} \ell(h_{\theta}(z_i)).
\]  

(1)

Up-weighting a training example \( z \) by an infinitesimal amount \( \epsilon \) leads to a new set of model parameters denoted by \( \theta^\epsilon(z) \). This set of new model parameters \( \theta^\epsilon(z) \) is obtained by solving:

\[
\theta^\epsilon(z) = \arg \min_{\theta} \frac{1}{n} \sum_{i=1}^{n} \ell(h_{\theta}(z_i)) + \epsilon \ell(h_{\theta}(z)).
\]  

(2)

Removing a training point \( z \) is similar to up-weighting its corresponding weight by \( \epsilon = -1/n \) in Equation (1). The main idea used by \( 3, 4, 11 \) is to approximate \( \theta^\epsilon(z) \) by the first-order Taylor series expansion around the optimal model parameters represented by \( \theta^* \), which leads to:

\[
\theta^\epsilon(z) \approx \theta^* - \epsilon H_{\theta^*}^{-1} \nabla_{\theta} \ell(h_{\theta^*}(z)),
\]  

(3)

where \( H_{\theta^*} \) represents the Hessian with respect to model parameters \( \theta^* \). Following the classical result of \( 11 \), the change in the model parameters \( (\Delta \theta = \theta^\epsilon(z) - \theta^*) \) on up-weighting the training example \( z \) can be approximated by the influence function \( 3 \) and is denoted by \( \mathcal{I}(z) \) as follows:

\[
\mathcal{I}(z) = \left. \frac{d\theta^\epsilon(z)}{d\epsilon} \right|_{\epsilon=0} = -H_{\theta^*}^{-1} \nabla_{\theta} \ell(h_{\theta^*}(z)).
\]  

(4)

The change in the loss value for a particular test point \( z_t \) when a training point \( z \) is up-weighted can be approximated as a closed form expression by the chain rule \( 3 \):

\[
\mathcal{I}(z, z_t) = -\nabla \ell(h_{\theta^*}(z_t))^T H_{\theta^*}^{-1} \nabla \ell(h_{\theta^*}(z)).
\]  

(5)

\( \mathcal{I}(z, z_t)/n \) is approximately the change in the loss for the test-sample \( z_t \) when a training sample \( z \) is removed from the training set. This result is, however, based on the assumption that the underlying loss function is strictly convex in the model parameters \( \theta \) and the Hessian \( H_{\theta^*} \) is a positive-definite matrix \( 3 \). For large models, computing and inverting the exact Hessian \( H_{\theta^*} \) is expensive. In such cases, the inverse-Hessian Vector product can be computed efficiently with a combination of Hessian-vector product rule \( 18 \) and optimization techniques (see Section 4 for details).

4 Computing Inverse-Hessian Vector Product

In large over-parameterized deep networks, computing and inverting the exact Hessian \( H_{\theta^*} \) is expensive. In such cases, the Hessian-vector product rule \( 18 \) or stochastic estimation \( 10 \) to compute the approximate inverse-Hessian Vector product. More specifically, to compute \( t = H_{\theta^*}^{-1} v \), we solve the following optimization problem using conjugate-gradient: \( t^* = \arg \min_t \{ \frac{1}{2} t^T H_{\theta^*} t - v^T t \} \), where \( v = \nabla_{\theta} \ell(h_{\theta^*}(z_t)) \). This optimization, however,
Figure 1: Iris dataset experimental results - (a,b) Comparison of norm of parameter changes computed with influence function vs re-training; (a) trained with weight-decay; (b) trained without weight-decay. (c) Spearman correlation vs. network depth. (d) Spearman correlation vs. network width.

requires the Hessian $H_{\theta^*}$ to be a positive definite matrix, which is not true in case of deep networks due to the presence of negative eigenvalues. In practice, the Hessian can be regularized by adding a damping factor of $\lambda$ to its eigenvalues (i.e. $H_{\theta^*} + \lambda I$) to make it positive definite. In deep models, with a large number of parameters and large training set, conjugate-gradient is often expensive as it requires computing the Hessian-vector product \[18\] for every data sample in the training set. In those cases, stochastic estimation techniques \[10\] have been used which are fast as they do not require going through all the training samples. In stochastic estimation, the inverse Hessian is computed using a recursive reformulation of the Taylor expansion: $H_j^{-1} = I + (I - H)H_{j-1}^{-1}$ where $j$ is the recursion depth hyperparameter. A training example $z_i$ is uniformly sampled and $\nabla^2 \ell(h_{\theta^*}(z_i))$ is used as an estimator for computing $H$. This technique also requires tuning a scaling hyperparameter $\gamma$ and a damping hyperparameter $\beta$. In our experiments with large deep models, we use the stochastic estimation method to compute the inverse-Hessian Vector product.

5 Experiments

Datasets: We first study the behaviour of influence functions in a small Iris dataset \[20\], where the exact Hessian can be computed. Further, we progressively increase the complexity of the model and datasets: we use small MNIST \[3\] to evaluate the accuracy of influence functions in a small CNN architecture with a depth of 6. Next, we study influence functions on modern deep architectures trained on the standard MNIST \[21\] and CIFAR-10 \[22\] datasets. Finally, to understand how influence functions scale to large datasets, we use ImageNet \[23\] to compute the influence estimates.

Evaluation Metrics: We evaluate the accuracy of influence estimates at a given test point $z_t$ using both Pearson \[24\] and Spearman rank-order correlation \[25\] with the ground-truth (obtained by re-training the model) across a set of training points. Most of the existing interpretability methods desire that influential examples are ranked in the correct order of their importance \[26\]. Therefore, to evaluate the accuracy of influence estimates, Spearman correlation is often a better choice. In our experiments in particular, we evaluate the correlations with the top influential training points.

5.1 Understanding Influence Functions when the Exact Hessian Can be Computed

Setup: Computing influence estimates with the exact Hessian has certain advantages in our study: a) it bypasses inverse-Hessian Vector product approximation techniques which induce errors in computing influence estimates. Thus, we can compare influence estimates computed with exact vs. approximate inverse-Hessian Vector products to quantify this type of error; b) The deviation of the parameters computed with the influence function from the exact parameters can be computed exactly. This information can be useful to further quantify the error incurred by (first-order) influence estimates in the non-convex setup. However, computations of the exact Hessian matrix and its inverse are only computationally feasible for models with small number of parameters. Thus, we use the Iris dataset along with a small feed-forward neural network to analyse the behaviour of influence function computed with the exact Hessian in a non-convex setup. We train models to convergence for 60k iterations with full-batch gradient descent. To obtain the ground-truth estimates, we re-train

\[\text{It is assumed that } \forall i, \nabla^2 \ell(h_{\theta^*}(z_i)) \succeq 0; \[3\] notes that if this is not true, the loss can be scaled down without affecting the parameters. The scaling factor is a hyperparameter which helps the convergence of the Taylor series. The damping coefficient is added to the diagonal of the Hessian matrix to make it invertible.\]
the models for 7.5k steps, starting from the optimal model parameters. For our analysis, we choose the test-point with the maximum loss and evaluate the accuracy of influence estimates with the ground-truth amongst of the top 16.6% of the training points. Through our experiments with the exact Hessian, we answer some relevant questions related to how properties of the network such as depth, width and regularizers (e.g. weight-decay) affect the influence estimates.

**The Effect of Weight-Decay**: One of the simpler and common regularization techniques used to train neural networks is weight-decay regularization. In particular, a term $\lambda \|\theta\|_2^2$, penalizing the scaled norm of the model parameters is added to the objective function, during training, where $\lambda$ is a hyperparameter which needs to be tuned. We train a simple feed-forward network\(^2\) with and without weight-decay regularization. For the network trained with a weight-decay regularization, we observe a Spearman correlation of 0.97 between the influence estimates and the ground-truth estimates. In comparison, for the network trained without a weight-decay regularization, the Spearman correlation estimates decrease to 0.508. In this case, we notice that the Hessian matrix $H_{\theta^*}$ is singular, thus a damping factor of 0.001 is added to the Hessian matrix, to make it invertible. To further understand the reason for this decrease in the quality of influence estimates, we compare the following metric across all training examples: a) Norm of the model parameter changes computed by re-training; b) Norm of the model parameter changes computed using the influence function (i.e. $\|H_{\theta^*}^{-1}\nabla \ell(z_i)\|_2 \; \forall i \in [1, n]$) (Fig. 1-(a,b)). We observe that when the network is trained without weight-decay, changes in model parameters computed with the influence function have a significantly larger deviation from those computed using re-training. This essentially suggests that the gap in Taylor expansion, using (first-order) influence estimates is large, when the model is trained without weight-decay. We observe similar results with smooth activation functions such as tanh (see the Appendix for details).

**The Effect Of Network Depth**: From Fig. 1(c), we see that network depth has a dramatic effect on the quality of influence estimates. For example, when the depth of the network is increased to 8, we notice a significant decrease in the Spearman correlation estimates. To further our understanding about the decrease in the quality of influence estimates when the network is deeper, we compute the gap in the approximation between the ground-truth parameter changes (computed by re-training) and the approximate parameter changes (computed using the influence function). To quantify the error gap, we compute the Spearman correlation estimates between the norm of true and approximate parameter changes across the top 16.6% of the influential examples. We find that with increasing depth, the Spearman correlation estimates between the norm of the true and approximate parameter changes decrease. From Fig. 2(c), we see that the approximation error gap is particularly large when the depth of the network is more than 5. We also notice a consistent increase in the curvature of the loss function (Fig. 2(b)), as the network becomes deeper. This possibly suggests that the curvature information of the network can be an upper bound in the approximation error gap between the true parameters and the ones computed using the influence function. Even in case of non-smooth activation functions like ReLU, we have a similar observation. (see the Appendix for more details).

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\(^2\) With width of 5, depth of 1 and ReLU activations
Figure 3: Experiments on small MNIST using a CNN architecture. (a) Estimation of influence function with and without weight decay on (a) the top influential points, (b) training points at 30th percentile of influence score distribution. (c) Correlation vs the weight decay factor (evaluated on the top influential points).

The Effect Of Network Width: To see the effect of the network width on the quality of influence estimates, we evaluate the influence estimates for a feed-forward network of constant depth, by progressively increasing its width. From Fig. 1-(d), we observe that with an increase in network width, the Spearman correlation decreases consistently. For example, we find that the Spearman correlation decreases from 0.82 to 0.56, when the width of the network is increased from 8 to 50. This observation suggests that over-parameterizing a network by increasing its width has a strong impact in the quality of influence estimates.

The Effect of Stochastic Estimation on inverse-Hessian Vector Product: For large deep networks, the inverse-Hessian Vector product is computed using stochastic estimation[10], as the exact Hessian matrix cannot be computed and inverted. To understand the effectiveness of stochastic approximation, we compute the influence estimates with both the exact Hessian and stochastic estimation. We observe that across different network depths, the influence estimates computed with stochastic estimation have a lower Spearman correlation when compared to the ones computed with the exact Hessian. From Fig. 2-(a), we find that the error in the approximation is more, when the network is deeper.

5.2 Understanding Influence Functions in Shallow CNN Architectures

Setup: In this section, we perform a case study using a CNN architecture[3] with around 2,600 parameters on the small MNIST dataset (i.e. 10% of MNIST); a similar setup used in [3]. To assess the accuracy of influence estimates, we select a set of test-points with high test-losses computed at the optimal model parameters. For each of the test points, we select 100 training samples with the highest influence scores and compute the ground-truth influence by re-training the model. We also select 100 training points with influence scores at the 30th percentile of the entire influence score distribution. These training points have low influence scores and a lower variance in their scores when compared to the top influential points. The model is trained with and without weight-decay regularization.

When trained with a weight-decay regularization and evaluated based on the top influential points, we find that the correlation influence estimates are consistently significant (Fig. 3(a)). This is consistent with the results reported in [3]. However, when the evaluation is done with the set of training samples at the 30th percentile of the influence score distribution, the correlation estimates decrease significantly (Fig. 3(b)). This shows that influence estimates of only the top influential points are precise when compared to ground-truth re-trainings. Furthermore, without the weight-decay regularization, influence estimates in both cases are poor across all the test-points (Fig. 3(a,b)).

To further understand the impact of weight-decay on influence estimates, we train the network with different weight-decay regularization factors. From Fig. 3(c), we see that the selection of weight-decay factor is important in getting high-quality influence estimates. For this specific CNN architecture, we notice that the correlations start decreasing when the weight-decay factor is greater than 0.01. Moreover, from Fig. 3(a,b), we find that the selection of test-point also has a strong impact

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3The model is trained for 500k iterations to reach convergence with the optimal model parameters \( \theta^* \). The ground-truth estimates are obtained by re-training the models from the optimal parameter set \( \theta^* \) for 30k iterations. When trained with a weight-decay, a regularization factor of 0.001 is used.
on the quality of influence estimates. For example, when the network is trained with weight-decay and the influence estimates are computed for top influential training points, we notice that the Spearman correlation estimates range from 0.92 to 0.38 across different test-points and have a high variance.

These results show that despite some successful applications of influence functions in this non-convex setup, as reported in [3], their performances are very sensitive to hyper-parameters of the experiment as well as to the training procedure. In the next two sections, we assess the quality of influence estimates on more complex architectures and datasets including MNIST, CIFAR-10 and ImageNet. In particular, we desire to understand, if the insights gained from experiments on smaller networks can be generalized to more complex networks and datasets.

### 5.3 Understanding Influence Functions in Deep Architectures

**Setup:** In this section, we evaluate the accuracy of influence estimates using MNIST [21] and CIFAR-10 [22] datasets across different network architectures including small CNN[3], LeNet [27], ResNet-18, ResNet-50 [28], VGG-13 and VGG-14 [29]. To compute influence estimates, we choose two test points for each architecture: a) the test-point with the highest loss, and b) the test-point at the 50th percentile of the losses of all test points. For each of these two test points, we select the top 40 influential training samples and compute the correlation of their influence estimates with the ground-truth estimates. To compute the ground-truth influence estimates, we follow the strategy of [3], where we re-train the models from optimal parameters for 6% of the steps used for training the optimal model. When the networks are trained with a weight-decay regularization, we use a constant weight-decay factor of 0.001 across all the architectures (see Appendix for more details).

**Results On MNIST:** From Table 1 we observe that for the test-point with the highest loss, the influence estimates in the small CNN and LeNet architectures (trained with the weight-decay regularization) have high qualities. These networks have 2.6k and 44k parameters, respectively, and are relatively smaller and less deep than the other networks used in our experimental setup. As the depth of the network increases, we observe a consistent decrease in quality of influence estimates. For the test-point evaluated at the 50th percentile of the loss, the correlations decrease marginally across most of the architectures. We find that on CIFAR-10, even architectures trained without weight-decay regularization have highly significant correlation estimates when evaluated with the test-point which incurs the highest loss.

**Results On CIFAR-10:** For CIFAR-10, across all architectures trained with the weight-decay regularization, we observe that the correlation estimates for the test-point with the highest loss are highly significant. For example, the correlation estimates are above 0.6 for a majority of the network architectures. However, for the test-point evaluated at the 50th percentile of the loss, the correlations decrease marginally across most of the architectures. We find that on CIFAR-10, even architectures trained without weight-decay regularization have highly significant correlation estimates when evaluated with the test-point which incurs the highest loss.

In case of MNIST, we have found that in shallow networks, the influence estimates are fairly accurate while for deeper networks, the quality of influence estimates decrease. For CIFAR-10, although the influence estimates are significant, we found that the correlations are marginally lower in deeper networks such as ResNet-50. The improved quality of influence estimates in CIFAR-10 can be attributed to the fact that for a similar depth, architectures trained on CIFAR-10 are less over-

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Table 1: Correlation estimates on MNIST And CIFAR-10 ; A=Test-point with highest loss; B=Test-point at the 50th percentile of test-loss spectrum; P=Pearson correlation; S=Spearman correlation

| Dataset | MNIST | CIFAR-10 |
|---------|-------|----------|
| Architecture | A (With Decay) | B (With Decay) | A (Without Decay) | A (With Decay) | B (With Decay) | A (Without Decay) |
| Small CNN | 0.95 | 0.32 | 0.41 | 0.61 | 0.36 | 0.47 |
| LeNet | 0.83 | 0.34 | 0.18 | 0.59 | 0.26 | 0.38 |
| VGG13 | 0.92 | 0.32 | 0.35 | 0.59 | 0.26 | 0.38 |
| VGG14 | 0.81 | 0.28 | 0.29 | 0.42 | 0.25 | 0.36 |
| ResNet18 | 0.41 | 0.28 | 0.14 | 0.42 | 0.25 | 0.36 |
| ResNet50 | 0.24 | 0.28 | 0.08 | 0.46 | 0.24 | 0.46 |

For CIFAR-10, evaluations on small CNN have not been performed due to the poor test accuracy.
parameterized compared to architectures trained on MNIST. Note that, in Section 5.1, where the exact Hessian matrix can be computed, we observed that over-parameterization decreases the quality of influence estimates. From Table(1), we also observed that the selection of test-point has a significant impact on the quality of influence estimates. Furthermore, we noticed large variations in the quality of influence estimates across different architectures. In general we found that influence estimates for small CNN and LeNet are reasonably accurate, while for ResNet-50, the quality of estimates decrease across both MNIST and CIFAR-10. Precise reasons for these variations are difficult to establish. We hypothesize that it can be due to the following factors: (i) Different architectures trained on different datasets have contrasting characteristics of loss landscapes at the optimal parameters which can have an impact on influence estimates. (ii) The weight-decay factor may need to be set differently in various architectures, to obtain high quality influence estimates.

5.4 Is Scaling Influence Estimates To ImageNet Possible?

To illustrate this, we conduct an additional influence estimation on ImageNet. After training an initial model to 92.302% top5 test accuracy, we select two test points at random, calculate influence over the entire training set, and then select the top 50 points by their influences as candidates for re-training. We then use the re-training procedure suggested by [3], which starts leave-one-out re-training from the parameter set obtained after the initial training. We re-train for an additional 2 epochs, approximately 5% of the original training time, and calculate the correlations. We observe that for both test points, both Pearson and Spearman correlations are very low (less than 0.15, see details in the appendix).

In our experiments, we observe high variability among ground-truth estimates obtained by re-training the model (see the appendix for details). We conjecture that this may be partially due to the fact that the original model has not be fully converged. To study this, we train the original model with all training points for an additional 2 epochs and measure the change in the test loss. We find that the overall top5 test accuracy has improved slightly to 92.336 % (+0.034) and the loss for one of the considered test points has decreased by relatively a significant amount of 0.679. However, the loss for the other point has increased slightly by 0.066. Such changes in loss values can therefore out-power the effect of leave-one-out retraining procedure. Second, we calculate the 2-norm of the weight gradients, which should be close to zero near an optimal point, and compare it to a standard pre-trained ImageNet ResNet-50 model as a baseline. We find these norms to be 20.18 and 15.89, respectively, showing our model has similar weight gradient norm to the baseline. Although these norms are relatively small given that there are 25.5M parameters, further re-training the model still changes loss values for some samples significantly, making the ground-truth estimates noisy. We suggest that one way to obtain reliable ground truth estimates may be obtained through leave-one-out re-training at this scale.

6 Conclusion

In this paper, we present a comprehensive analysis of the successes and failures of influence functions in deep learning. Through our experiments on datasets including Iris, MNIST, CIFAR-10, ImageNet and architectures including LeNet, VGGNets, ResNets, we have demonstrated that influence functions in deep learning are fragile in general. We have shown that several factors such as the weight-decay, depth and width of the network, the network architecture, stochastic approximation and the selection of test points, all have strong effects in the quality of influence estimates. In general, we have observed that influence estimates are fairly accurate in shallow architectures such as small CNN and LeNet, while in very deep and wide architectures such as ResNet-50, the estimates are often erroneous. Additionally, we have scaled up influence computations to the ImageNet scale, where
we have observed influence estimates are highly imprecise when compared to the ground-truth re-trainings. These results call for developing robust influence estimators in the non-convex setups of deep learning.

7 Broader Impact

In recent times, there has been a significant increase in the usage of deep learning models in applications such as NLP [7], computer vision [6] and healthcare [30]. In NLP, deep learning models are often used for machine comprehension [31], summarization [32] and question-answering [33] tasks, amongst others. In vision, deep models are frequently used in applications such as object detection [34], image segmentation [35] and inpainting [36]. Deep learning is also used in sensitive applications such as medical vision [8] and cancer detection/analysis [37]. In view of these applications and the black-box nature of decision making by deep learning systems, it has become particularly important to interpret model predictions [38]. For example, an individual using a deep learning model might want to understand the reasoning behind the decision made by the model. Model interpretability is thus useful in enabling humans/end-users to trust black-box deep learning systems. One way to interpret test-time predictions is by detecting influential training samples responsible for particular predictions. Influence functions [3, 4, 1, 2], coming from robust statistics, is one of the methods to identify such important training examples (or influential training groups). However, methods based on influence functions are not well-understood in deep learning settings. Providing an understanding about the performance of influence functions in deep learning has been the primary focus of our paper. We believe our work can help accelerate research in influence functions and model interpretability in general, with a specific focus on deep learning. To the best of our knowledge, our work will facilitate research in explainable machine learning and does not lead to any negative or unethical societal impact.

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Appendix

A Additional Experimental Results on Iris Dataset

In this section, we provide additional experimental results to understand the effect of network depth on the correlation estimates for ReLU networks. From Fig. 4, we observe that even in case of architectures trained with non-smooth activation functions such as ReLU, the correlation estimates consistently decrease with depth. Similar to our findings in case of networks trained with tanh activation (as shown in the main text), we observe that the top eigenvalue of the Hessian matrix and the Taylor’s approximation gap increases with depth.

![Figure 4: Additional Iris experimental results for ReLU networks: (a) Spearman correlation vs. network depth; (b) Top eigenvalue of the Hessian vs. network depth; (c) Spearman correlation between the norm of parameter changes computed with influence function vs. re-training.](image1)

In the main text, we reported that when a network with ReLU activation is trained with a weight-decay regularization, the correlation estimates are significant and the Taylor’s approximation gap is less. We find a similar result even with smoother activation functions such as tanh. From Fig. 5, we observe that when a network with tanh activation is trained with a weight-decay regularization, the Taylor’s approximation gap is less. However when the network is trained without a weight-decay regularization, the Taylor’s expansion gap is large resulting in poor quality of influence estimates.

![Figure 5: Additional Iris experimental results for tanh networks; (a) When trained with weight-decay, the Taylor’s approximation gap is small; (b) When trained without weight-decay, the Taylor’s expansion gap is large. These results are similar to our findings for ReLU networks which are reported in the main text.](image2)
B Visualisation of Top Influential Points

In this section, we visualise the top influential training samples corresponding to a given test-point. In the main text, we noted that the selection of test-points has a strong impact on the quality of influence estimates. Additionally, we also observe that the selection of test-points has an impact on the semantic-level similarities between inferred influential training points and the test-points being evaluated. For example, in Fig. 6, we observe that 2 out of the top 5 influential points are not from the same class as the test-point with index 1479. However in Fig. 7, we observe that all the top 5 influential training samples are semantically similar and from the same class as the evaluated test-point with index 7196.

Figure 6: Top 5 influential points for the test point: 1479 (CIFAR-10). The model is a ResNet-18 trained with a weight-decay regularization; Only 3 out of the 5 points are semantically similar to the test-point with class "Bird".

Figure 7: Top 5 influential points for the test point: 7196 (CIFAR-10). The model is a ResNet-18 trained with a weight-decay regularization; All the 5 training points are semantically similar to the test-point from the class "Airplane".

C Running Times

In this section, we provide computational running times for (first-order) influence function estimations. We note that in models with a large number of parameters, the influence computation is relatively slow. However, even in large deep models, it is still faster than re-training the model for every training example. In our implementation, for a given test-point $z_{\text{test}}$, we first compute $c = H_{\theta^*}^{-1} \nabla \ell(h_{\theta^*}(z_{\text{test}}))$ once which is the most computationally expensive step. We then compute a vector dot product i.e. $c^T \nabla \ell(h_{\theta^*}(z_i)) \forall i \in [1,n]$. In Table 2, we provide the computational running times for estimating influence functions in different network architectures.

D Additional Experimental Details on ImageNet Influence Calculations

In this section we give further details on the influence estimation on ImageNet. To help address the high computational cost of training and re-training, we utilize highly optimized ImageNet training schemes such as those submitted to the DAWNBench competition [39]. In particular we use the scheme published from [40], for the ResNet-50 architecture which uses several training tricks including progressive image resizing, weight decay tuning, dynamic batch sizes [41], learning rates [42], and half-precision floats. Although these techniques are unorthodox, they are sufficient for our purposes since we need only to compare between the fully trained and re-trained models. We replicate this scheme and obtain a top-5 validation accuracy of 92.302%.

5https://www.fast.ai/2018/08/10/fastai-diu-imagenet/
| Architecture | Influence Computation Time (MNIST) | Influence Computation Time (CIFAR-10) |
|--------------|-----------------------------------|--------------------------------------|
| Small CNN    | 141.13 ± 0.51                     | N/A                                  |
| LeNet        | 162.6 ± 2.20                      | 136.39 ± 3.16                       |
| VGG13        | 3886.23 ± 3.45                    | 4416.54 ± 2.01                      |
| VGG14        | 4619.11 ± 5.08                    | 4620.69 ± 6.11                      |
| ResNet-18    | 960.08 ± 4.67                     | 910.58 ± 8.49                       |
| ResNet-50    | 4323.13 ± 8.26                    | 3857.66 ± 21.6                      |

Table 2: Computational running times for influence function across different architectures

We now give further details on the test points selected. The first has a test loss at the 83rd percentile (loss=2.634, index = 13,923, class=kit fox), the second has the test loss at the 37th percentile (loss=0.081, index = 2,257, class=gila monster), where the indices refer to where they appear in test_loader.loader.dataset. We visualize these test points in Figure 9.

Next, for each of these test points, we compute influence across the entire dataset and select the top 50 training points by influence scores. We visualize 25 of these points in Figures 10 and 11. We observe that there is qualitative similarity between the test points and some of their respective most influential training points, but not others. Although there is qualitative similarity in some cases, the results are still overall weak quantitatively.

We plot the obtained correlations in Figure 8.

For computing the weight gradient norm, we take the mean norm in batches of size 128 over the entire dataset for both our model and a standard PyTorch pretrained model as a baseline, both of which are ResNet-50 models with around 25.5M parameters.

![ImageNet Influence Correlations](image)

Figure 8: ImageNet influence estimation results for the selected test points 13,923 (left) and 2,257 (right). X-axis is change in test loss after removal of a training point and retraining as described in the text. Y-axis is the change in test loss estimated with influence function. Pearson and Spearman correlations are shown in the caption. Correlations are low, showing the weakness of this influence estimation.
Figure 9: Selected test points for influence estimation.
Figure 10: Top 25 ImageNet training points by influence for test point 13,293, kit fox. Many of the identified classes are furred mammals, e.g. red wolf, basenji, and dingo, which have visual similarity to the test point. Other examples are questionable, e.g. the common iguana, and African elephant. Although there is qualitative similarity in some cases, the results are still overall weak quantitatively.
Figure 11: Top 25 ImageNet training points by influence for test point 2,257, gila monster. Many of the identified
classes are spotted lizards, e.g. banded gecko ad European fire salamander, which have visual similarity to
the test point. Other examples are questionable, e.g. the stingray, coral fungus, and barrow. Although there is
qualitative similarity in some cases, the results are still overall weak quantitatively.