Investigation of POWER8 processors for astronomical adaptive optics real-time control

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ABSTRACT

The forthcoming Extremely Large Telescopes (ELTs) all require adaptive optics systems for their successful operation. The real-time control for these systems becomes computationally challenging, in part limited by the memory bandwidths required for wavefront reconstruction. We investigate new POWER8 processor technologies applied to the problem of real-time control for adaptive optics. These processors have a large memory bandwidth, and we show that they are suitable for operation of first-light ELT instrumentation, and propose some potential real-time control system designs. A central processing unit (CPU)-based real-time control system significantly reduces complexity, improves maintainability, and leads to increased longevity for the real-time control system.

Key words: instrumentation: adaptive optics – instrumentation: miscellaneous – methods: numerical.

1 INTRODUCTION

The forthcoming Extremely Large Telescopes (ELTs; Johns 2008; Nelson & Sanders 2008; Spyromilio et al. 2008) will all rely on adaptive optics (AO) systems (Babcock 1953) for their successful operation, allowing the degrading effects of atmospheric turbulence to be greatly reduced. An AO system actively measures wavefront perturbations introduced by the Earth’s atmosphere, and attempts to mitigate these in real-time (on millisecond time-scales) using one or more deformable mirrors (DMs). This is a computationally demanding task, and requires a dedicated real-time control system (RTCS). Computational requirements scale with the forth power of telescope diameter when considering traditional RTCS algorithms: for a given level of AO correction, the DM pitch must remain constant, and so the number of subapertures across the telescope pupil scales with telescope diameter, $d$. The total number of subapertures and actuators therefore each scale as $O(d^2)$, and therefore the number of operations required for wavefront reconstruction (a matrix–vector multiplication) scales as $O(d^4)$. Due to this rapid scaling of computational complexity, careful design considerations must be made when designing real-time control systems for the ELTs.

These RTCSs must be designed with long lifetimes, since the AO instruments on these telescopes are expected to be operational for at least thirty years (Vernet et al. 2012). Therefore maintenance, of both software and hardware is key to success. An RTCS design which is hardware ambiguous, i.e. does not require a particular hardware set to operate, is clearly advantageous. Previous system designs have frequently relied on specific hardware, typically digital signal processors (DSPs) and field programmable gate arrays (FPGAs; for example the ESO SPARTA system, Fedrigo et al. 2006), which, due to long periods spent in design, are often close to obsolescence even during commissioning, with availability of spare parts becoming problematic, and specific programming knowledge required. Hardware failure of these systems then poses the risk that an entire new system will require designing, with the original software not being portable to new hardware.

In recent years, there has been much success with hardware agnostic AO RTCSs which operate on conventional PC hardware, including the Durham AO real-time controller (DARC; Basden et al. 2010; Basden & Myers 2012), which is a generic system, used by the CANARY AO on-sky demonstrator instrument (Myers et al. 2008), and the real-time control system for the Gemini South telescope GeMS AO system (Rigaut et al. 2012). In theory, such systems simply require a recompilation of the source code to be ported to other (similar) hardware platforms, and are easy to move on to upgraded hardware. In practice, the advent of binary driver code, e.g. for wavefront sensors (WFSs) and DMs, means that porting is not always possible. Although porting to new hardware is typically limited to other PC-like systems that have an operating system running on a central processing unit (CPU), this is not always the case. In particular, DARC system has a modular design which allows parts of the real-time pipeline to be placed in alternative hardware, including for example:

(i) module and slope calculation in FPGA using a customized version of the SPARTA system (Fedrigo et al. 2006)
(ii) wavefront reconstruction using graphics processing units (GPUs; Basden et al. 2010)
(iii) a full GPU pipeline, from raw WFS images to DM demands.
However, this system still requires a CPU based core to oversee control of the hardware accelerators.

For ELT-scale AO systems, the largest computational requirements come from wavefront reconstruction algorithms, which typically use a matrix–vector multiplication (MVM) to obtain DM surface shape from WFS slope measurements. On conventional PC hardware, this algorithm is memory-bound, rather than compute-bound, and so for low latency operation, systems with large memory bandwidth are required. For this reason, accelerator cards (such as GPUs) are considered in designs for ELT-scale RTCSs to provide the necessary memory bandwidths for these algorithms. However, this in itself raises new problems in moving data into and out of the accelerator for processing, which adds time and hence latency to the RTCS pipeline. Designs that minimize this latency are key.

1.1 The POWER8 processor

The specification and road-map of the IBM POWER8 processor (Sinharoy et al. 2015) seems promising for AO RTCSs, with two key relevant features: A memory bandwidth approaching that of GPUs (up to 230 GB/s), and support for a novel interconnect technology (NVLink; Foley 2014) due for release in 2017 that will provide an order of magnitude increase in data bandwidth between processor and GPU. Additionally, the OpenPower foundation has the potential for providing novel hardware acceleration architectures tightly coupled with POWER8 processors via the Coherent Accelerator Processor Interface (CAPI; Stuecheli et al. 2015), including a currently available offering from the company Nallatech. The memory bandwidth of these processors is significantly larger than other available CPUs, hence the interest for AO real-time control, and a concise overview of the memory subsystems is given by Starke et al. (2015).

Here, we provide details of initial performance testing of the DARC RTCS on a POWER8 system. In Section 2, we discuss the system configuration, RTCS installation process and the tests that we perform. In Section 3, we present our findings, and we conclude in Section 4.

2 THE DARC REAL-TIME CONTROLLER ON A POWER8 SYSTEM

Most of the results that we will present here are performed on a low-end Tyan OpenPower Customer Reference system, model GN70-BP010, hosted at Durham. This system has a single four-core POWER8 processor clocked at 3 GHz. Each core has eight-way symmetric multithreading, providing a total of 32 hardware threads. The system has 16 GB DDR3 (1.6 GHz) RAM, controlled by a single Centaur memory controller. The total theoretical memory bandwidth for this system is 28.8 GB s⁻¹ between CPU and main memory (19.2 GB s⁻¹ read, 9.6 GB s⁻¹ write).

We have also had limited cloud access to a more powerful S824 POWER8 system with two 12-core processors (to which our machine instance had access to 22 cores), each eight-way threaded, providing a total of 176 hardware threads. Half of the memory banks of this machine are populated, and thus a total memory bandwidth of about 59 GB s⁻¹ for read operations, and 29.5 GB s⁻¹ for write operations is available. The operating system of this machine was run behind a hypervisor. Both of these systems run the Ubuntu operating system (14.10). Results presented here are from our low-end system unless stated otherwise.

2.1 Real-time control system installation

We use the publicly available DARC AO RTCS system, with source code downloaded from the sourceforge hosting site. Installation on a POWER8 system was trivial: we simply had to remove three unsupported compiler options from the Makefile (-msse2 -mfpmath=ssse3 -march=native) and then compile and install in the usual way. All of the required library dependencies were available from the Ubuntu repositories, and downloaded automatically as part of the DARC installation process. We did not attempt to optimize DARC using compiler flags specific to the POWER8 processor, and we used the freely available gcc compiler, for which source code is available (important for lifetime considerations).

We investigated the use of GigE Vision cameras for WFSs, using the open-source Aravis library, with modifications specifically to allow access to the camera pixel stream, rather than full-frame access (to reduce RTCS latency). Because this library is entirely open-source, and does not require any hardware drivers, there were no issues with binary drivers. This library provides access to a number of WFSs that have been used on-sky with the CANARY AO system, including an Imperx Bobcat camera, an Emergent Vision Technologies HS2000 10 GBit camera and a First-Light OCAM2S camera. During operation, as soon as sufficient pixels have arrived at the computer to complete a given subaperture, this subaperture is processed by a thread (calibration, slope calculation and partial reconstruction). The thread then returns to compute the next available subaperture, in a round-robin fashion. Once all subapertures for a given frame have been processed, each thread will have a partial DM vector, and these are then combined in a reduction step to yield the final DM command.

To further demonstrate the proof of concept of a complete AO system, we selected an Alpao 241 actuator DM with an Ethernet interface. It was necessary to develop our own library interface for this DM since source code for the Software Developers Kit was not available, and the binary libraries were for X86 architectures. However, control of this DM involves sending a User Datagram Protocol packet, and so was trivial to implement. A closed-loop AO system driven by a POWER8 server is therefore feasible using an existing RTCS.

2.2 Testing real-time performance

We investigate the performance of DARC on POWER8 by configuring the system as would be used in a number of different AO cases. These are

(i) a 40 × 40 subaperture single conjugate AO (SCAO) system,
(ii) a 80 × 80 subaperture SCAO system,
(iii) a 80 × 80 subaperture system with increased actuator counts.

For each of these cases, we investigate performance for different sized subapertures, i.e. different numbers of pixels per subaperture.

The third case can be viewed as a single WFS of the proposed European Extremely Large Telescope (E-ELT) multi-conjugate adaptive optics (MCAO) instrument (Foppiani et al. 2010) with computation of a full set of partial DM demands. A full MCAO real-time control system could then be comprised of one compute node per WFS, with combination of partial DM demands being computed as a (low operation count) final processing step to give the demands to be sent to the DMs. We discuss this further in Section 3.4.

Our tests presented here do not include a physical WFS camera or DM, since we do not have suitable equipment available (specifically, cameras with sufficient pixels and frame rates, and...
a DM with enough actuators). Rather, we concentrate on the core computational pipeline. Our previous experience has shown that introducing a physical camera to a system has little impact on overall performance (maximum achievable frame rate), provided the camera itself is capable of reaching these frame rates. Because the DARC RTCS can process pixels as they arrive at the computer, then once the last pixel for a given frame arrives, most of the computation has typically already completed. The RTCS is used without frame piping here, i.e. there are never two frames being processed at once, so that the frame rate represents the computation time of a given frame. We note that with a real camera, expected readout time and data transfer time will depend very much on camera model, and in astronomical AO the readout time is often the limiting factor in achievable frame rate (likely to be the case for the forthcoming ELTs), and for true latency considerations, this should be taken into account. For example, for a camera with a maximum frame rate of 500 Hz, the readout time (and exposure time) will be 2 ms. Assuming that data is transferred as it is read out (rather than buffered), this means there will be a delay of 4 ms from start of exposure to last pixel arriving at the computer (by which time, most of the computation will have completed). However, an investigation of camera latency is beyond the scope of this paper.

Of key importance in the approach that we take is that we are using a fully configured AO RTCS, which has been proven on-sky. When benchmarking hardware performance, it can be tempting to write simple benchmarking code which investigates the key algorithms under consideration, i.e. image calibration (vector operations), slope computation (vector and reduction operations), and wavefront reconstruction (MVM). However, this leads to optimistic estimates, since the benchmark is grossly simplified and bears little resemblance to actual code that would be usable on-sky at a telescope.

### 2.2.1 The performance metric

We define the performance of the RTCS by measuring the time taken to perform the computation for each AO system frame. In the default DARC configuration, which we use here, the computation of each frame must be completed before the next frame is started. This therefore means that the inverse of the frame computation time gives the maximum achievable frame rate for the AO system. This behaviour is critical for optimizing AO system latency on a given hardware set.

The DARC RTCS uses a horizontal processing strategy (Basden et al. 2010) with each thread operating on WFS data from start to finish, rather than having different threads performing individual tasks (e.g. a set of threads for image calibration, a set for slope computation, and a set for wavefront reconstruction). This strategy allows automatic load balancing by the operating system, and simplifies performance optimization: the main parameter to be optimized is the number of processing threads, rather than balancing the number of threads per algorithm which can become a complex optimization problem. Of further consideration is the number of subapertures that each thread should process at once, influencing the order of memory operations and the size of the partial MVMs. If this is too small, then many inefficient small MVM operations will reduce the performance, while if too large, a small number of large MVM operations will lead to a saturation of memory bandwidth, resulting in threads being work-starved.

### 3 RTCS PERFORMANCE ON POWER8

We now consider the achievable performance on the POWER8 systems under investigation, and consider the application for future RTCS designs. For each case, we investigate changing the number of threads used by DARC, and the processing block size used, i.e. the number of subapertures processed together as a block.

#### 3.1 An 8 m XAO system

We investigate the case of a eXtreme AO (XAO) system on an 8 m telescope with 20 cm subapertures (40 x 40), and results are shown in Fig. 1. Here, it can be seen that with the low-end system a maximum frame rate of nearly 2 kHz is achieved. In this case, the control matrix size is 1304 x 2480, requiring a memory bandwidth of 23.4 GB s⁻¹ to read this from main memory every RTCS iteration at this frame rate. This is larger than the available memory bandwidth (19.2 GB s⁻¹) and therefore, the control matrix (12 MB) is being stored in the large L3 cache (32 MB).

RTCS processing tasks are divided among a selected number of threads, and we see that using 31 threads provides best performance. The processor has four cores, each with eight-way simultaneous multithreading capability (i.e. 32 virtual cores). Of particular note is the linearity of these curves between eight threads and the peak: the RTCS pipeline is seen to be highly parallelizable with performance scaling almost directly with the number of cores available.

We also consider the case when this system has a larger number of actuators to control, e.g. for a woofer-tweeter system. This is of
POWER8 processors for AO real-time control

Figure 1. Achievable RTCS frame rate as a function of number of processing threads used. The individual lines represent the number of times (given by the legend) threads are reused each frame (affecting the number of partial matrix–vector products that are implemented).

Figure 2. Maximum achievable RTCS frame rate as a function of number of actuators controlled for a 40 × 40 subaperture system. Inset is shown the corresponding memory bandwidth required by the MVM to achieve this frame rate.

Figure 3. Maximum AO frame-rate as a function of number of pixels per subaperture (with 80 × 80 square subapertures).

maximum frame rate is 100.2 Hz on our low-end system, requiring a memory bandwidth of 18.9 GB s⁻¹ to read the matrix from memory each iteration (it is too large to fit in cache), in addition to reading calibration image and other memory operations. This is very close to the theoretical maximum memory bandwidth, and so we conclude that the POWER8 architecture is optimized and pipelined in such a way as to achieve peak performance for mixed processing tasks.

3.2 A single ELT WFS

We investigate the case of an E-ELT SCAO system, with a single WFS with 80 × 80 subapertures (with 6 × 6 pixels per subaperture), and a control matrix of size 5160 × 9824 (193 MB). In this case, the
3.2.1 Thread counts

We investigate how the number of processing threads affects the achievable AO frame rate. Fig. 4 shows that using close to, but less than, the number of hardware threads (32) provides best performance. Of particular note here is that (in comparison with Fig. 1) performance no longer scales directly with the number of processing cores. This is because this larger problem size is memory bandwidth limited, rather than compute limited.

3.2.2 Amdahl’s law

Amdahl’s law (Amdahl 1967) states that the performance gain in a system through parallelization (or other) techniques is limited by the fraction of time spent within the parts of the system benefiting from those improvements.

In the case of a high-order AO RTCS, the limiting performance factor is memory bandwidth, required for wavefront reconstruction. Increasing available memory bandwidth will only continue to significantly improve performance while other parts of the computational pipeline (namely image calibration and slope calculation) do not begin to dominate the computation time. Therefore, to be able to make scaled performance predictions, we need to be able to determine the time taken for these operations which are compute limited rather than memory bandwidth limited.

We therefore investigate performance with and without wavefront reconstruction. For the case without wavefront reconstruction, we are interested in how well the POWER8 system can process pixel information and produce wavefront slopes, and assume that the reconstruction could be performed elsewhere (i.e. in a GPU, using NVLink), though of course this may introduce additional latency.

Fig. 5 shows maximum achievable frame rates for the AO RTCS processing pipeline when the large MVM for wavefront reconstruction is removed, and thus places an approximate limit on achievable performance for these processors when unlimited memory bandwidth is available. Therefore, we can see that when using a POWER8 system with greater memory bandwidth (up to $256 \text{ GB s}^{-1}$ read bandwidth for a dual-processor server), frame rates of nearly $1.3 \text{ kHz}$ should be available for this system, limited by the memory bandwidth for wavefront reconstruction, since we know that other aspects of the real-time pipeline can be performed faster than this ($1.6 \text{ kHz}$ on our low-end system, and faster on a high end 24-core server).

3.3 A multiple mirror ELT SCAO system

To investigate the performance of this ELT-scale SCAO system further, we consider the case of multiple mirror SCAO systems, i.e. with an increased number of actuators. This increases the control matrix size, and thus allows us to investigate performance limiting factors for different AO system configurations. We also investigate performance with different subaperture sizes (pixels per subaperture), so that we can separate compute intensive and memory intensive tasks.

Fig. 6 shows maximum AO frame rate on our low-end POWER8 hardware as a function of control matrix size.
slightly reduced. We believe that this is due to less of the larger matrix being cached, i.e. when there is a larger matrix to read, cache prediction is not so good. However, the system is still able to achieve nearly 90 per cent of theoretical memory bandwidth during the AO system loop.

3.3.1 Operation at necessary frame rates

The maximum frame rates reported so far have not been sufficient for an on-sky ELT AO system. However, we have only been able to perform benchmarking on a low-end system. Due to the high utilization of available memory bandwidth (close to 100 per cent), we can make predictions as to maximum achievable frame rates for currently available higher end systems. A POWER8 S824 system contains two processors, each with up to 128 GB s\(^{-1}\) memory bandwidth for read operations, a combined factor of 13.3 times greater than our system. If memory bandwidth is the limiting factor, we could expect an AO frame rate of greater than 1.2 kHz for an ELT-scale SCAO system using an S824 system. It is likely that other parts of the computational pipeline would start to limit performance so that this frame rate would not be achieved. In Section 3.2.2, we have investigated performance on our low-end system with the MVM removed, to demonstrate that pixel processing and slope computation at higher frame rates is achievable. Therefore, with sufficient memory bandwidth, ELT frame rates are easily available on an existing POWER8 server.

3.4 An ELT MCAO system

We have considered the performance case for an ELT-scale SCAO system, and we now use this information to consider MCAO system design. The E-ELT MCAO instrument, MAORY (Foppiani et al. 2010), is likely to have 4–6 laser guide stars (LGSs) and up to three natural guide star (NGS) low-order WFSs, with a total of two or three DMs (including the telescope M4 DM), operating up to 10 000 actuators with a 500 Hz frame rate.

Processing of WFS images to yield wavefront gradients is independent, i.e. slopes obtained by processing one WFS do not depend on the processing of other WFSs. Similarly, when using conventional MVM wavefront reconstruction methods (we discuss other methods in Section 3.7), the slopes from each WFS can be used independently of other WFSs to compute a partial set of DM commands. The partial DM commands from each WFS can then be summed, yielding the final DM demands to be applied to the mirror, in a low count vector addition operation.

We therefore now consider a MCAO control solution which has a separate POWER8 server for each LGS WFS (directly connected), and an additional POWER8 server for the three NGS, with partial DM demands being sent to one server for summation to yield the final DM demands, as shown in Fig. 7. We note that since the NGS are likely to be of lower order (resulting in a smaller MVM), it would be possible to process all NGS in a single server, reducing cost and complexity. This server is then also used to collate the partial DM demands, which will arrive over more than one 10 G Ethernet link to reduce latency.

With this control solution, each server therefore has to process a single WFS, and between 8000–10 000 actuators, and so we can directly estimate expected performance using Fig. 3, which by scaling to the memory bandwidth available in a S824 system, will yield frame rates above 500 Hz, the MAORY design goal. Further processor improvements over the next few years (for example the

![Figure 7. A schematic design showing components for a ELT MCAO real-time control system, and the links between them. WFSs are connected individually to a POWER8 server, which computes partial DM demands. These are then summed before being sent to the DM.](http://mnras.oxfordjournals.org/)

POWER9 processor in 2017) will improve performance further, and be available within the time frame of MAORY system development.

3.5 An ELT MOAO system

We now consider requirements for an ELT-scale multi-object AO (MOAO) system. The E-ELT MOAO instrument is likely to be MOSAIC (Hammer et al. 2014), and will use six LGSs and up to five NGS. Up to 20 MOAO channels are proposed, each with a DM, in addition to the main telescope M4 DM.

Fig. 8 shows a possible schematic design for the MOAO real-time control system. In summary, 21 servers are required, one for each DM, including the M4 mirror. Each server receives images from three or four WFSs and processes these to provide wavefront slope information. These wavefront slopes are then shared with two other servers, which in return also share the wavefront slope information computed from their WFSs. Therefore, each server will have access to the 11 WFS slope vectors. Each server then performs a tomographic wavefront reconstruction, projected along a given line of sight, and sends the DM demands to the relevant DM.

With this design, each server is responsible for processing four WFS images, and performing a MVM with a matrix size of about 100 000 × 5000. At the desired frame rate of 250 Hz, this represents a required memory bandwidth of about 470 GB s\(^{-1}\), which is achievable using a four-socket POWER8 server (e.g. the S850 system, which has a read memory bandwidth of 512 GB s\(^{-1}\)), though is above that obtainable in a single dual socket server. It is likely that within the next decade (the time-frame for ELT MOAO instrument

![Figure 8. A schematic design showing components for a ELT MOAO real-time control system, and the links between them. Four WFSs are connected to a server, which computes slope measurements, and shares these with two other servers. Each server then has access to all WFS slope measurements, and computes DM demands for a single DM.](http://mnras.oxfordjournals.org/)
development), significant improvements in memory bandwidth will be realized, enabling this performance goal to be met with even greater overhead, reducing latency. Additionally, the inclusion of one or two GPUs to the system (taking advantage of the forthcoming high performance NVLINK interconnect; Foley 2014) specifically to perform MVM would further reduce latency. We discuss this further in Section 3.7.

It should be noted that with this design, the wavefront reconstruction for each DM is independent, allowing different algorithms to be trialled with performance comparisons made while the system is in operation. This capability will be key to maximizing MOAO performance.

### 3.6 Variation in latency

The variation of AO system latency, or jitter, is a key parameter when developing a real-time control system. If this jitter is large, then there will be frequent delays in the AO processing pipeline, leading to reduced AO performance. This is particularly critical for higher order AO systems. Fig. 9 shows the variation in latency measured over 100 000 frames on the POWER8 server for both the 40 × 40 and 80 × 80 subaperture systems. For the higher order case, the variation in latency follows a Gaussian distribution, with a FWHM of 1.4 ms, 5 per cent of the mean frame time. No frames take more than twice the mean frame time, and 99 per cent of frames take less than 8 per cent longer than the mean time.

For the low-order case, the variation in latency is no longer Gaussian, showing an extended tail, and additional features that may be related to the granularity of the timer. The rms jitter is 62 µs. Here, less than 0.01 per cent of frames take longer than twice the mean frame time to complete, and 99 per cent of frames take less than 38 per cent longer than the mean frame time to complete.

We are currently using a stock Ubuntu kernel (3.16.0-23). The use of a real-time kernel would further improve this jitter, though we do not investigate here as this is not yet available.

### 3.7 Further considerations

We have so far only considered the basic AO RTCS pipeline operations, including wavefront reconstruction using a MVM algorithm, image calibration and slope computation. However, for an ELT, this is unlikely to be sufficient, as further algorithms will be necessary, for example the linear-quadratic-Gaussian (LQG) control as demonstrated by CANARY, for vibration mitigation (Sivo et al. 2013), which involves several MVM operations.

Current implementations of LQG demonstrated on-sky have required significantly more computational power and memory bandwidth than a conventional MVM algorithm, and so the hardware that we are investigating here may not be suitable for these algorithms. There are two alternatives: LQG is an active area of research, and efficient implementations are being developed (Gray & Le Roux 2012). Alternatively, hardware acceleration techniques can be considered.

A requirement for additional hardware acceleration will benefit significantly from the proposed high-speed NVLINK and CAPI interconnects under development for future POWER8 processors and hardware accelerators. Specific hardware, for example GPUs or FPGAs, can be used to provide acceleration of given algorithms, in this case, the wavefront reconstruction problem. A high-speed, low latency link is key to enabling this, as it will maintain low system latency: improved algorithmic behaviour will only improve AO system performance if the algorithms do not lead to significant increases in AO system latency. A key feature of the CAPI interface is that it enables abstracted code to be developed with accelerators sharing the same memory address space as the CPU, allowing code to be developed independently of the physical hardware acceleration used.

A high bandwidth, low latency accelerator interconnect is also essential for future designs of ELT-scale XAO real-time systems. For these systems, low latency is critical.

#### 3.7.1 Future-proofing AO real-time control

We have demonstrated that an existing AO RTCS can be ported to an alternative processor technology with very little effort, and that this technology has the potential to enable AO real-time control for first-light ELT AO instruments without the requirement for additional hardware acceleration. This greatly simplifies RTCS design, and provides greater confidence that the RTCS software will be able to operate for the foreseeable future, independent of underlying hardware changes (provided a C compiler exists). No proprietary libraries are necessary, and full source code for this system is available.

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**Figure 9.** (a) A histogram of frame computation times for an 80 × 80 SCAO system. (b) A histogram of frame computation times for a 40 × 40 SCAO system.
Of key importance here is that an ELT-scale AO real-time control system can be developed in the widely used C programming language, and does not require any custom hardware, or any niche untransferable skills. Transferability of this system to other processor types give a significant degree of confidence that a system developed in this way will remain operable, configurable, upgradable and hardware independent for the foreseeable future. This is a key advantage for telescopes with expected operational lifetimes approaching a century.

4 CONCLUSION

We have investigated the use of a freely available, open source, AO RTCS on new POWER8 hardware. We find that installation on this hardware was trivial, demonstrated the use of WFSs and a DM, and find that computational performance is in line with expectations, with ELT-scale AO RTCS performance being limited by available memory bandwidth, of which our RTCS typically reaches above 90 per cent of the theoretical maximum. The large potential memory bandwidth of the POWER8 CPU, along with forthcoming innovations enabling high bandwidth communication between the CPU and other hardware (including GPUs, with NVLink), means that POWER8 systems are a prime contender for use with ELT-scale AO RTCSs, and that using conventional computer server technology is highly attractive to maintain longevity, upgradability and comprehension of these systems.

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