The climate impact of the Drake Passage opening from a fluid dynamics point of view: the role of Antarctic glaciation

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ABSTRACT

Pronounced global cooling around the Eocene-Oligocene transition (EOT) was a pivotal event in Earth’s climate history, controversially associated with the opening of the Drake Passage. Using a physical laboratory model we revisit the fluid dynamics of this marked reorganization of ocean circulation. Here we show, seemingly contradicting paleoclimate records, that in our experiments opening the pathway yields higher values of mean water surface temperature than the “closed” configuration. This mismatch points to the importance of the role ice albedo feedback plays in the investigated EOT-like transition, a component that is not captured in the laboratory model. Our conclusion is supported by numerical simulations performed in a global climate model (GCM) of intermediate complexity, where both “closed” and “open” configurations were explored, with and without active sea-ice dynamics. The GCM results indicate that sea surface temperatures would change in the opposite direction following an opening event in the two ice dynamics settings, and the results are therefore consistent both with the laboratory experiment (slight warming after opening) and the paleoclimatic data (pronounced cooling after opening). It follows that in the hypothetical case of an initially ice-free Antarctica the continent could have become even warmer after the opening, a scenario not indicated by paleotemperature reconstruction. These results provide circumstantial evidence supporting a particular EOT scenario in which Antarctica has already been – at least partially – covered with ice when the Drake Passage fully opened.

Introduction

The opening of major gateways in the Southern Ocean, namely the Drake Passage and the Tasmanian Seaway around the Eocene-Oligocene boundary ca. 34 Ma, is widely regarded as a key contributor to the thermal isolation of Antarctica (that has reached its polar position already in the Cretaceous period) and the glaciation of the continent1,2. The Eocene-Oligocene transition (EOT) is marked by a major climatic change on the global scale, with paleotemperature records indicating a rapid drop of global mean temperature. The EOT also coincides with the largest extinction event in the Cenozoic3, and extraterrestrial impact events are also known from this interval4.

The opening of the southern seaways enabled the development of the Antarctic Circumpolar Current (ACC) which inhibits
warm surface currents to transport heat from the Tropics to the Antarctic continent. The details of the role that the onset of ACC played in the growth of continental-scale Antarctic ice sheet (AIS) are still far from being fully resolved. In their seminal paper DeConto and Pollard argued that the opening of gateways was merely a secondary factor for AIS development and falling atmospheric CO₂ level was the main driver of EOT. However, significant gaps and contradictions exist in the available CO₂ proxy data in the most relevant time interval between 40 and 30 Ma, thus setting a precise timeline of events (i.e. CO₂ decrease, temperature drop, Drake Passage opening), and hence to establish their causal links remains controversial. Here we revisit this problem with a focus on the climate impact of Drake Passage opening, using a novel combination of a laboratory experiment and numerical simulations.

The present-day ACC is dominantly driven by wind stress and the meridional buoyancy contrast (yielding quasi-geostrophic flow) of the ocean itself, whose sea surface temperature (SST) field exhibits a strong and practically zonally invariant temperature difference of ≈ 10°C between the latitudes of 70°S and 35°S. The enormous material and momentum transport associated with ACC – estimated at 100–150 Sv making it the largest ocean current – is facilitated by the unique geographical boundary conditions of the Southern Ocean, the only ocean basin on the planet that is not confined by meridional continental boundaries. In the bulk, the “thermal wind balance” facilitated by this flow maintains tilted density interfaces (isopycnals) sloping northward, yielding baroclinic instability in the ACC. Blocking this circumpolar zonal pathway with a continental barrier deviate the flow to form meridional boundary currents which in turn lead to the buildup of a zonal (east-west) gradient in sea surface temperature and a significantly larger meridional heat transport between the equatorial and polar regions than in the “open” configuration (i.e. without the barrier).

Inspired by the Drake Passage opening – but not aspiring to accurately model the actual paleoceanographic processes in their vast complexity – the present study investigates the underlying physical basics of the heat exchange dynamics in a topologically similar, but radically simplified system. The widespread claim which we intend to critically address is that an opening event resulting in the formation of a circumpolar pathway could, in itself, lead to the thermal isolation (and glaciation) of the polar region and, eventually, the decrease of global temperature, via blocking meridional heatflow. If so, this reasoning would suggest that Antarctica may not have necessarily been covered with land ice before the formation of the circumpolar pathway. It is also possible, however, that due to other reasons the glaciation preceded the Drake Passage opening, and the latter merely accelerated the cooling. Thus, regarding the physical basics of the problem, the key question we are focusing on is whether or not the topological reorganization of the ocean currents, facilitated by an opening event, is sufficient to yield cooling per se. Our experimental and numerical findings suggest that if we do not incorporate ice cover in our models, representing the case of an initially ice-free Antarctica, then the expected global cooling does not occur, somewhat contradicting the traditional reasoning. However, if land ice is already present on the continent, then the opening can indeed enhance further glaciation and cooling. These findings, although admittedly conceptual in nature, may contribute to the better understanding of the timing of the various EOT-related events.

**Laboratory experiments**

Due to hydrodynamic similarity certain key aspects of the complexity of such planet-scale flow systems can be modeled rather accurately in relatively simple tabletop-size experiments. The apparatus applied in the present work is one version of a laboratory set-up widely used in atmospheric and ocean dynamics; the so-called differentially heated rotating annulus, introduced by Raymond Hide in the early 1950s. The setting captures the two most relevant factors in the formation of ACC-like geostrophic jets and baroclinic surface vortices, namely the meridional temperature gradient and planetary rotation (Fig. 1). The applicability of the setting to investigate large-scale ocean circulation in the “closed” Drake Passage configuration has been demonstrated in a recent pilot study.

The layout of the apparatus is sketched in Fig. 1. The tank is mounted on a turntable so that its axis of symmetry coincides with the axis of rotation. The vertical sidewalls of the tank consist of two coaxial cylinders: the cooled inner cylinder (modeling Antarctica) and the heated outer rim. The working fluid (water), representing the Southern Ocean, occupies the annular gap between the two cylinders, and therefore experiences a certain “meridional” temperature difference ΔT. The top of the tank is not covered to enable the observation of the water surface temperature (WST) field with infrared thermography. The geometrical and dynamical parameters of the setup had to be set to ensure approximate dynamic similarity with the actual Southern Ocean are discussed in Methods. A thin removable vertical insulating barrier reaching through the full water depth represents the Antarctica-South America land bridge (light brown wall in Fig. 1).

Each experimental run was conducted as follows. After setting the temperature of the inner and outer sidewalls and the rotation rate to their prescribed values, the experiment was kept running for over ~3500 revolutions – a time interval sufficient for quasi-stationary flow patterns to develop – with the barrier in place (“closed” configuration) of which the last ~1700 revolution-long leg was later evaluated. Then the barrier was removed by pulling it upward instantaneously, while the rotation rate and the thermal control remained unchanged. In the resulting “open” configuration the data acquisition continued for an additional ~1700 revolutions. (More details about the measurement techniques are given in the Methods section.)
Figure 1. The sketch of the experimental set-up. The water body is located in the annular domain between the two coaxial cylinders. The entire tank is rotating around its axis in the clockwise direction – as indicated by the black arrow – to imitate the Southern Hemisphere configuration. The inner and outer sidewalls are kept at constant cold and warm temperature, respectively, maintaining a radial (“meridional”) temperature difference forcing, $\Delta T_r$. The removable vertical obstacle is sketched with light brown. The mount of the infrared sensor is standing in the laboratory near the turntable (i.e. not co-rotating), and thus scans the surface temperature field of the water. The sensor’s field of view is the domain marked with yellow.

Figure 2. Water surface temperature patterns. a, Infrared thermographic image of the water surface temperature (WST) anomalies in a “closed” laboratory experiment ($\Omega = 2.0$ rad/s, $\Delta T_r = 11\,^\circ C$, $H = 5$ cm). b, Annual average of the sea surface temperature (SST) anomalies in a “closed” numerical simulation in the PlaSim GCM (“no ice” configuration). The zonal averages are subtracted in each grid cell for better visibility.

Fig.2a shows a typical WST map from a “closed” experiment with a clearly visible zonal temperature gradient and a pronounced cold anomaly (blue) at the “eastern” side of the barrier (as reported in\textsuperscript{12}). The abrupt reorganization of the flow following the removal of the barrier is reflected in the exemplary time series in Figs.3a and b. Panel a) shows the “meridionally” (i.e. radially) averaged, zonally scanned exemplary temperature time series smoothed by a ca. 10-revolution (491-point) running mean. The record, hence, accounts for the temporal development of the spatially averaged WST field. The blue section of the curve corresponds to the “closed” state, whereas the “open” leg of the experiment, exhibiting significantly higher mean temperatures (after a roughly 200-revolution transient phase) is marked with red. The rapid transition is also indicated by the abrupt drop of the running standard deviation of the temperature field (using the same window length) (Fig.3b). The histograms of Fig.3c represent the spatial and temporal variability of the WST distribution of the entire water surface before and after the opening event (with the same color coding as in panels a) and b)). Apparently, the opening caused a significant shift of the distribution towards higher temperatures, the disappearance of the cold anomaly (cf. Fig.2a) and, as a consequence, of the fat left tail in the “closed” histogram in panel c).

The histograms of Fig.3d and e are obtained in a control experiment where the set-up was equipped with co-rotating temperature sensors (thermocouples) measuring near-surface and near-bottom water temperature time series at mid-radius, at an azimuthal position angle (“longitude”) of ca. $90^\circ$ from the barrier. Panels d) and e) show the measured temperature distributions before and after the opening, respectively. Clearly, the surface (orange) and bottom (green) temperature distributions of the
Figure 3. Water surface temperature (WST) distributions before and after the opening. a, The time series of the spatially averaged WST field at each time instant, following a 10-revolution running averaging for smoothing. The change of color (from blue to red) marks the removal of the barrier ($\Omega = 2.0$ rad/s, $\Delta T_r = 18.7^\circ$C). b, The spatio-temporal standard deviation of the fluctuations of the WST field, as calculated with a 10-revolution moving window in the same experiment as in panel a). c, Histograms of the WST field of the entire surface acquired from the data of $\sim$1700 revolutions in both the “closed” (blue) and “open” (red) legs of the same experiment as of panels a and b. d, e, Histograms of temperature time series obtained at the same horizontal location via co-rotating temperature sensors placed underneath each other, one close to the surface, at a depth of 6 mm and the other above the bottom by 5 mm from a control experiment (with a total water height of $H = 5$ cm). The near-surface histograms are shown with orange and the near-bottom ones are green. Panel d represents the “closed” and panel e the “open” leg of the same experiment ($\Omega = 2.0$ rad/s, $\Delta T_r = 11^\circ$C).

“closed” case largely overlap with each other, whereas in the “open” configuration the surface and bottom histograms shifted toward higher and lower temperatures, respectively, indicating stronger thermal isolation between the upper and lower layers.

Figure 4. Temperature statistics before and after the opening. a, Zonal and temporal mean WST after vs. before the opening, from 7 experiments (marked with different colors in panels a,b, and c, and from 12 zonal contours (“latitudes”). b, c, Zonal and temporal standard deviation (b) and skewness (c) of the same signals, after vs. before the opening.
The comparison of the mean, standard deviation (StDev) and skewness (Sk) of the first (closed) and second (open) leg of the WST signals from the 12 infrared channels (“latitudes”) in 7 experiments is presented in Fig.4a, b, and c, respectively. The runs differed in the preset values of the “meridional” temperature contrast \( \Delta T \) (see Methods section). Each symbol represents one channel (“latitude”) of a given experiment, and the different runs are distinguished by the coloring of the data points. Thus, in panel a) one data point marks a time-averaged “zonal mean” temperature \( \langle T \rangle \) for one “latitude” and one experimental run. The point’s horizontal coordinate represents its value in the closed leg and the vertical coordinate gives its value in the open leg of the experiment. Apparently, all values scatter above the dashed \( y = x \) line. Therefore, the zonal and temporal average WST has increased in each run and at all “latitudes” after the removal of the barrier, in concert with the vanishing of the cold surface anomaly. This change is accompanied with a clear decrease of the signals’ standard deviations (Fig.4b) and – as the histogram of Fig.3b also demonstrates – the transition from typically left-skewed distributions towards more symmetrical ones (Fig.4c).

Our fluid dynamical interpretation of the above observations is the following: in the “closed” configuration the barrier inhibits the buildup of a geostrophic zonal flow and facilitates full-depth “meridional” overturning that yields substantial mixing between the near-surface and deeper domains of working fluid. In other words, there are less coherent baroclinic waves or eddies and a stronger turbulence and mixing which homogenizes the temperature at the upper and lower levels. Hence, in this case, the average surface temperature can be considered as a representative surrogate of the volumetric average of the whole water body. In the “open” setting, however, heat transport in the “meridional” direction practically occurs by the means of baroclinic eddy (turbulent or wave) transport that is stronger close to the surface due to the free surface condition. These surface coherent structures are unable to mix and exchange heat effectively with the colder bulk (cf. Fig.3c) and therefore conserve a generally warmer WST field.

To summarize, the sign of the surface temperature shift observed in the experiments is the opposite of the one associated with the actual EOT climate change. Obviously, we can conclude that the laboratory experiment is not a sufficient representation of the Southern Ocean in this respect. Although the laboratory model captures the basic overturning dynamics of the actual ocean, it obviously lacks important feedback mechanisms, most notably atmospheric and ice dynamics. In order to study the combined effect of these components of the climate system and to disentangle the competing feedback processes, we turned to numerical simulations.

**Simulations in a GCM**

Our numerical investigations were conducted with a general circulation model or global climate model (GCM) of intermediate complexity, the so-called Planet Simulator (PlaSim), developed at the University of Hamburg\(^{18,19}\), the technical details of which are described in the Methods section. The model incorporates atmospheric and – when coupled with a large-scale geostrophic (LSG) ocean module\(^{20}\) – full-depth ocean dynamics.

Instead of implementing the realistic paleogeographic shorelines and the (mostly unknown) paleobathymetry of the Eocene epoch, the “closed” configuration was modeled simply by a full-depth meridional barrier (“dam”) in the Drake Passage connecting the tip of South America and the Antarctic peninsula, while otherwise keeping the present-day continental arrangement unchanged, as sketched in Fig.2b. In the zonal direction the barrier spanned two grid cells (i.e. a domain of 11.2°). Each simulation covered a 1000-year time frame, of which the last, already quasi-stationary 200-year interval was evaluated in terms of annually averaged surface temperature fields both in the “closed” and “open” configurations (no barrier, modern topography). The model runs were conducted using the widely accepted estimate\(^{21}\) for the Late Eocene-Early Oligocene CO\(_2\) level, i.e. 750 ppm.

As an adjustable setting, the so-called ice module of PlaSim could be switched on or off. When active, the module computes the extent and thickness of sea ice dynamically as the temperature field changes. When turned off, however, even when the water temperature decreases below freezing point in a given grid cell, the formation of ice shelves is inhibited and, hence, so are the resulting effects on the system (most notably, the ice-albedo feedback). The latter setting therefore approximates the laboratory experiment where ice dynamics is also absent. Fig.2b shows the sea surface temperature anomaly field from a “closed” simulation (“no ice” configuration), computed by averaging over the aforementioned 200-year period and subtracting the zonal means in each grid cell. The pattern indeed resembles the one seen in the experimental model (Fig.2a) with pronounced warm and cold anomalies at the Pacific and Atlantic sides of the barrier, respectively.

Panels a) and b) of Fig.5 show histograms of annually averaged surface temperature fields in the domain of the Southern Ocean (i.e. between latitude 55° S and the Antarctic circle) as acquired from the aforementioned 200-year-long series of quasi-stationary data. The blue and red graphs represent the distributions from the “closed” and the “open” configurations, respectively. The temperature scale was adjusted so that the mean of the “closed” record is subtracted from all measured values, thus the “closed” histograms are centered to zero in both panels. In panel a) the GCM’s ice module was turned off, whereas in panel b) it was active. The most striking feature of the simulations with inactive ice module (Fig.5a) is that the histograms remain practically identical in the “open” and “closed” cases. However, when the ice module is active in the model (Fig.5b),
Figure 5. Surface temperature anomalies in the GCM. a, b. Histograms of the sea surface temperature anomaly field of the domain of the Southern Ocean (between latitude 55° S and the Antarctic circle) acquired from the annually averaged temperature fields of 200 subsequent years in both the “closed” (blue) and the “open” (red) configurations. The anomalies here are given as the deviation from the mean of the “closed” distributions (hence both “closed” histograms are exactly zero-centered). Panels a and b correspond to the inactive and active ice dynamics presets, respectively. c Zonally and temporally averaged meridional profiles of surface temperature change. The differences between the zonal and temporal mean temperatures of the two barrier configurations $⟨T_{\text{open}}⟩ - ⟨T_{\text{closed}}⟩$ subjected to otherwise identical conditions.

The “open” configuration yields markedly lower surface temperatures in the Southern Ocean, in qualitative agreement with the EOT paleotemperature records. The latitude-dependence of the (imaginary) opening event’s impact on the zonal and temporal mean surface temperatures of the Southern Ocean and Antarctica is demonstrated in Fig. 5c. The difference between the average values from the “open” and “closed” configuration $⟨T_{\text{open}}⟩ - ⟨T_{\text{closed}}⟩$ is plotted here against the latitude. In the absence of ice dynamics (inactive ice module, black curve) the only region where the opening yields detectable cooling is the close vicinity of the pole. The rest of Antarctica, i.e. the domain up to ca. 70°S exhibits markedly higher temperatures in the “open” configuration than in the “closed” one. In the fully ocean-covered latitudes (North of the 70°S circle) the difference between the “closed” and “open” cases is negligible (cf. Fig. 5a). Averaging over the investigated domain (or the entire globe), therefore a slight temperature increase can be observed in the “open” configuration relative to the “closed” one. With active ice module, however, the “open” configuration (orange curve) clearly has lower surface temperatures at all latitudes of the Southern Ocean and Antarctica, consistently with the EOT records. Apparently, the magnitude of the “open-closed” temperature difference profile increases monotonically toward the South Pole.

To summarize, ice dynamics – most notably ice-albedo feedback – appears to play a key role in the chain of mechanisms that connect the opening of the Drake Passage and the overall decrease of temperature. Without ice dynamics the opening per se does not cause such cooling, in agreement with the findings from the laboratory experiment.

Discussion

Our results indicate that the classic interpretation of the EOT global cooling, and the narrative of Antarctic glaciation in particular, needs to be reconsidered. The findings from a conceptual laboratory model as well as numerical simulations in a GCM suggest that under constant thermal forcing conditions the opening of a meridional barrier does not necessarily yield an overall decrease of surface temperatures. Moreover, it may even have the opposite effect – a slight warming in the investigated region – if ice formation is inhibited. If, however, the conditions are such that they allow for ice coverage to be present in the “closed” and “open” configurations alike, then indeed – in line with the paleoclimate record – the South polar region gets markedly colder in the “open” case than in the “closed” one.

Earlier numerical simulations also indicated that closing the present-day Drake Passage alone would not necessarily affect the mean SST in the Southern Ocean significantly. Instead, the observed differences between the closed and open configurations were found to be very sensitive to the CO$_2$ concentration and, hence, the sea ice cover in the region. Moreover, another work using a fully coupled atmosphere-ocean model with a mid-Oligocene geography suggests that the development of ACC appears to be more of a consequence than a driver of the global cooling, and as such it rather acted as a feedback mechanism in the EOT. More recent numerical studies also emphasized the amplifying role of the ice-albedo feedback in the process, in agreement with our findings.

The question arises, however, of why we observe increased water surface temperatures in the laboratory experiment after the
opening, whereas the Southern Ocean’s mean surface temperature remains practically unaffected by changing the configuration of the Drake Passage in the ice-free GCM simulations (Fig.5). First, it is important to realize that the thermal boundary conditions are somewhat different in the two settings. In the laboratory, the temperature of the middle cylinder (representing Antarctica) is kept at a fixed, prescribed value, whereas in the GCM dynamic heat exchange takes place between Antarctica and the ocean. As the upper curve of Fig.5c shows, certain fully or partially land-covered latitudes (i.e. approximately those south of 70°S) indeed exhibit higher temperatures in the open configuration, which may well be the consequence of heat release from the sea surface to the atmosphere. In the open ice-free GCM run the meridional temperature contrast is apparently below the one found in the closed ice-free configuration. Smaller meridional temperature gradient is typically associated with elevated global mean temperatures, as observed in the case of the present-day climate change as well. Therefore, the experimental and numerical results, despite their obvious differences, reassuringly point to the same direction in the sense that no significant decrease of water (or land) temperatures could be detected in the absence of ice dynamics. To conclude, it appears that the opening of the circumpolar pathway per se is not even qualitatively sufficient to account for glaciation and global cooling if Antarctica is fully ice-free.

Our laboratory experiment is a radically simplified, abstract representation of the Southern Ocean. The setting does not capture wind stress or beta effect, just to name two crucial factors that propel and shape the present-day Antarctic Circumpolar Current (ACC). The beta effect would also play a key role in a more realistic closed configuration, as the northward flow at the closed passage would likely form a western boundary current. Nor does our laboratory model involve the effects of other continental boundaries and heat and material exchange with the ocean gyres. Nevertheless, the lessons learned here on the heat transfer of quasi-geostrophic flows in the two most basic topological configurations are relevant and insightful and we believe that they would hold up in more complex (realistic) cases as well. Removing the barrier allows the formation of a meandering circumpolar jet, which blocks full-depth overturning and thus hinders meridional heat transfer. In the lack of overturning, the mixing between cold deep water and warmer surface layers becomes markedly reduced, hence, surface temperatures increase.

This finding, even though it may seem counterintuitive, provides circumstantial evidence in favor of the assumption that Antarctica may already have been, at least partially, ice-covered before the opening of the Drake Passage⁵. With land ice already present on the continent, the opening could indeed catalyze further ice formation (presumably via ice-albedo feedback) and could eventually lead to the thermal isolation of Antarctica and the strengthening of the ACC. However, in the hypothetical case of an initially ice-free Antarctica the continent could have become even warmer following the opening, a scenario not indicated by paleotemperature reconstructions. The proposed story line is consistent with the interpretation of DeConto and Pollard⁶, who argued that the falling atmospheric CO₂ level was the primary cause for the glaciation of Antarctica, and consequently, this event could possibly predate the opening of the gateway. However, in light of our results, once a permanent ice cap was already present, the opening indeed contributed to the further increase of meridional temperature gradients and to the observed rapid drop of the global mean temperature ca. 34 million years ago.

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Methods

Laboratory experiments

The desired hydrodynamical similarity of the flow in the experiment to the actual ocean circulation requires certain nondimensional numbers (similarity parameters) to match, at least to the order-of-magnitude. In case of thermally driven rotating flows that are studied here, the most important similarity parameter is the thermal Rossby number (also known as Hide number) $Ro_T$, which has the form

$$Ro_T = \frac{\alpha g H \Delta T_r}{\Omega^2 L^2},$$

(1)

where $g = 9.81 \text{ m/s}^2$ is the gravitational acceleration, $\alpha = 2.07 \times 10^{-4} \ 1/\degree C$ represents the volumetric thermal expansion coefficient of the fluid, $L$ is the characteristic horizontal length-scale (e.g. basin size), $H$ denotes the fluid depth, $\Omega$ is the rotation rate (angular velocity) of the system and $\Delta T_r$ is the lateral (‘meridional’) temperature difference. The other nondimensional parameter of key relevance for the present configuration is the Taylor number $Ta$ which quantifies the relative importance of viscous effects:

$$Ta = \frac{4 \Omega^2 L^5}{\nu^2 H},$$

(2)
where $\nu = 1.004 \times 10^{-6} \text{ m}^2/\text{s}$ denotes the kinematic viscosity of the fluid. The values of these parameters in large-scale ocean currents – with a horizontal length scale comparable to the radius of Earth – are typically of the orders of $\mathcal{O}(Ro_T) \leq 10^{-2}$ and $\mathcal{O}(Ta) \geq 10^8$.

For our experiments two differentially heated rotating annuli were used, one located at the von Karman Laboratory of the Eötvös University (Budapest, Hungary) and another at the Fluid Centre of the Brandenburg University of Technology (Cottbus, Germany). The geometrical parameters of the Budapest tank are as follows: the radius of the inner (cooled) cylinder is $R_1 = 4.5 \text{ cm}$, the radius of the outer (heated) cylindrical sidewall is $R_2 = 15 \text{ cm}$ and the applied water depth was $H = 5 \text{ cm}$. The values of the same parameters in the Cottbus set-up are $R_1 = 4.5 \text{ cm}$, $R_2 = 12 \text{ cm}$, $H = 5 \text{ cm}$. The rotation rate of the experimental tank was set to a constant angular velocity $\Omega = 2.0 \text{ rad/s}$ in all cases. The blocking barriers were made of wood and acrylic and had a width of $R_2 - R_1$, a thickness of $d = 0.5 \text{ cm}$ and blocked the flow in the full depth. The temperature difference $\Delta T_r$ between the sidewalls that is to be kept constant throughout the experiments is regulated by means of Laude ProLine heating and cooling thermostats.

Seven of the experiment runs were conducted in the Budapest tank, which is equipped with a $4 \times 16$ pixel Melexis thermal imaging sensor, mounted above the tank as sketched in Fig.1. The sensor is characterized by a narrow radial “footprint” with an effective field of view of $16^\circ \times 60^\circ$ spanning from the inner to the outer cylindrical sidewall. 12 pixels in the radial (“meridional”) domain of the footprint is occupied by the free water surface of the annular gap. As the tank rotates underneath, the sensor thus scans the water surface temperature (WST) field with a sampling rate of 10 Hz. It is to be noted, that the penetration depth of the applied wavelength range (i.e. $7.5 - 14\mu\text{m}$) into water is less than a millimeter. The thermographic image of Fig.2a was obtained by a co-rotating InfraTec VarioCam infrared camera mounted above the set-up, operating in the same spectral range as the Melexis sensor.

The temperature setpoints and the corresponding values of $Ro_T$ and $Ta$ – calculated with $L = R_2 - R_1$ as the horizontal scale – are listed in Table 1. All the experiment runs started with a closed barrier and after at least 1800 revolutions (1.5 hours) the barrier was removed instantaneously by pulling it upward manually. Instead of running separate experiments with “open” and “closed” configurations, we have chosen to apply this dynamic procedure in order to ensure that the external conditions, affected by uncontrollable variations of the laboratory environment, stay precisely the same in the “open” and “closed” cases.

| run ID | $\langle \Delta T_r \rangle \ [^\circ \text{C}]$ | $Ro_T$ |
|--------|---------------------------------|--------|
| 1B     | 18.7                            | 0.043  |
| 2B     | 22.7                            | 0.052  |
| 3B     | 25.6                            | 0.059  |
| 4B     | 14.6                            | 0.034  |
| 5B     | 17.8                            | 0.041  |
| 6B     | 21.9                            | 0.050  |
| 7B     | 14.6                            | 0.034  |
| 1C     | 11.0                            | 0.045  |
| 2C     | 11.0                            | 0.045  |

Table 1. The list of experiments. The values of $\langle \Delta T_r \rangle$ represent the temporal average temperature differences between the inner and outer walls (transients are omitted). Experiment IDs ending with B and C denote runs conducted in the Budapest and Cottbus tanks, respectively. The value of the Taylor number for the Budapest tank is $Ta_B = 4.05 \times 10^9$ and for the Cottbus tank $Ta_C = 8.26 \times 10^9$ in all experiments.

The studied parameter range falls in the same geostrophic turbulent dynamical regime as the aforementioned values ($\mathcal{O}(Ro_T) \leq 10^{-2}$ and $\mathcal{O}(Ta) \geq 10^8$) representing ACC, see e.g. the regime diagram in Ref.\textsuperscript{1}, therefore the fact that the nondimensional numbers do not match precisely are not relevant from a qualitative point of view.

The control experiment to compare the temperature distributions in the bulk – inaccessible for the infrared sensors – and near the surface (see Fig.3c,d) was conducted in the Cottbus tank, using Ahlborn ALMEMO NiCr sensors with a relative resolution of $0.05^\circ \text{C}$ and a sampling rate of 1 Hz. The sensors for the near-surface and near-bottom temperatures were both fixed onto the same co-rotating mast above the free surface of the rotating annulus, and penetrated 6 mm-deep into the water surface and 5 mm from the bottom, respectively. The data were streamed in real-time through the co-rotating UHF module ALMEMO 8590-9.

The Planet Simulator (PlaSim) GCM

In the present study the climate model, Planet Simulator (PlaSim)\textsuperscript{2} is used to simulate Earth’s climate. PlaSim was developed to understand the main physical processes in climate dynamics. In previous studies (see, e.g., \textsuperscript{3–5}) it proved to be an appropriate
numerical tool to investigate possible behaviour of the climate system on global scale.

For our purposes, we make use of the same atmospheric setup as in Ref. 3, i.e., the horizontal resolution of the simulations is T21, which yields a grid of approximately 5.6° × 5.6°. As a non-standard aspect of our model setup, the PlaSim atmosphere is coupled to a large scale geostrophic (LSG) ocean (original name: The Hamburg Large Scale Geostrophic Ocean General Circulation Model (Cycle 1)7. The atmospheric dynamics are described by primitive equations that represent conservation laws, thermodynamics and the hydrostatic approximation. Through parameterization, the model accounts for numerous unresolved processes, including sea ice formation, which is a key feature in our study. The so-called sea ice module is based on the zero layer model of Ref. 9. This model computes the thickness of the sea ice from the thermodynamic balances at the top and the bottom of the sea ice. The zero layer model assumes the temperature gradient in the ice to be linear and eliminates the capacity of the ice to store heat. For each marine cell on the grid, sea ice is allowed to form when the surface temperature drops below 271.25 K (−1.90 °C). If a grid cell is covered by sea ice, snowfall is accumulated on top of the ice. Snow is converted to sea ice if there is sufficient snow to suppress the ice/snow interface below the sea level. The typical sea ice thickness for the fully ice-covered sea is around 10 meters.

The LSG ocean model is based on the observations that for large scale ocean circulation models designed for climate studies, the relevant characteristic spatial scales are large compared to the internal Rossby radius throughout most of the ocean. At the same time the characteristic time scales are large compared with the periods of gravity modes and barotropic Rossby wave modes. The LSG ocean model was developed by Maier-Reimer and Mikolajewicz in early 1990s7. This LSG ocean model was originally proposed by8, and is described more fully by9, and it has been used in a number of climate and paleoclimate studies (see, e.g.,9–13). In Maier-Reimer et al. 1993 the LSG ocean model was investigated in details. It has been showed that the simulated mean ocean circulation for appropriately chosen surface forcing fields adequately and realistically reproduces the principal water mass properties, residence times, and large-scale transport properties of the observed ocean circulation within the constraints of the model resolution. We used the default resolution of 3.5 x 3.5 degrees and 22 non-equidistant vertical layers along with a realistic present-day bathymetry. The typical maximum basin depth is 5500 meters. We mention that the LSG model was also applied previously to investigate the role of the closed Drake Passage in ocean dynamics12. In our model setup the original LSG model is modified, the "closed" configuration was modelled simply by a full-depth meridional barrier (“dam”) in the Drake Passage connecting the tip of South America and the Antarctic Peninsula, otherwise keeping the present-day continental arrangement unchanged. The initial state and other properties of the LSG model ocean are the default ones7. The LSG ocean is spun up with present geography and bottom topography for a period of 10000 years to reach steady state. After the spin-up we compute a 1000 year long period to investigate the impact of the closed/open Drake Passage.

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Author contributions statement

M.V., T.B. and M.H. led the interpretation of results and the writing of the manuscript. A.H., I.D.B., C.R., U.H., M.V. and T.B. prepared and oversaw the experimental set-up and conducted the laboratory experiments. M.H. implemented and carried out the numerical simulations in PlaSim. A.P. developed the infrared sensor system. A.N. and J.P. provided the essential paleoclimatological and geological input for the research work and helped the interpretation of the results. All authors reviewed the manuscript.

Additional information

Competing interests The authors declare that they have no competing interests.