Rotational spectroscopy of n-propanol: Aa and Ag conformers

O. Zingsheim, J. Maßen, H. S. P. Müller, B. Heyne, M. Fatima, L. Bonah, A. Belloche, F. Lewen, and S. Schlemmer

Context. The primary alcohol n-propanol (i.e., normal-propanol or propan-1-ol; \(\text{C}_3\text{H}_7\text{OH}\)) occurs in five different conformers: \(\text{Ga, Gg, Gf, Aa, and Ag}\). All rotational spectra of the three conformers of the \(G\) family are well described, making astronomical search of their spectroscopic signatures possible, as opposed to those of the \(Aa\) and \(Ag\) conformers.

Aims. Our goal is to facilitate the astronomical detection of \(Aa\) and \(Ag\) conformers of n-propanol by characterizing their rotational spectra.

Methods. We recorded the rotational spectra of n-propanol in the frequency domain of 18–505 GHz. Additional double-modulation double-resonance (DM-DR) measurements were performed, more specifically with the goal to unambiguously assign weak transitions of the \(Aa\) conformer and to verify assignments of the \(Ag\) conformer.

Results. We derived a spectroscopic quantum mechanical model with experimental accuracy (with \(J_{\text{max}} = 70\) and \(K_{\text{e,\text{max}}} = 6\)) for \(Aa\) n-propanol. Furthermore, we unambiguously assigned transitions (with \(J_{\text{max}} = 69\) and \(K_{\text{e,\text{max}}} = 9\)) of \(Ag\) n-propanol; in doing so, we prove the existence of two tunneling states, \(A_{\text{g}}^+\) and \(A_{\text{g}}^−\).

Conclusions. The astronomical search of all five conformers of n-propanol is now possible via their rotational signatures. These are applied in a companion article on the detection of n-propanol toward the hot molecular core Sgr B2(N2).

Key words. molecular data – methods: laboratory: molecular – techniques: spectroscopic – ISM: molecules – radio lines: ISM – submillimeter: ISM

1. Introduction

To date, more than 260 molecules have been detected in space\(^1\). Many rotational molecular spectra are collected in databases such as the Cologne Database for Molecular Spectroscopy (CDMS; Müller et al. 2001, 2005; Endres et al. 2016), allowing astronomers to search for their spectroscopic signatures in space. Several complex organic molecules (COMs) have been found thus far, for instance, in the warm gas around protostars in star forming regions, such as formamide (Rubin et al. 1971), ethyl cyanide (Johnson et al. 1977), and ethanediol (Zuckerman et al. 1975; Pearson et al. 1997) have been detected. Ethanol occurs as two conformers, so-called anti-ethanol (Zuckerman et al. 1975) and gauche-ethanol (Pearson et al. 1997), and both have been referenced based on their first detection. The next heavier alcohol, propanol (\(\text{C}_3\text{H}_7\text{OH}\)), is thus another candidate for a detection in space, particularly in environments with large amounts of ethanol. The possible detection of propanol may be further supported by the detection of its isomer, ethyl methyl ether (or methoxyethane, \(\text{CH}_3\text{CH}_2\text{OCH}_3\)) in space (Tercero et al. 2015, 2018; Manigand et al. 2020).

Propanol itself occurs in two isomers, as a primary alcohol n-propanol (normal-propanol, propan-1-ol, normal propyl alcohol; \(\text{CH}_3\text{CH}_2\text{CH}_2\text{OH}\), where the hydroxyl group (–OH) is attached to the outer carbon atom, and as a secondary alcohol t-propanol (iso-propanol, propan-2-ol, or iso-propyl alcohol; \(\text{CH}_3\text{CH(OH)}\text{CH}_3\)), where the hydroxyl group is attached to the central carbon atom. n-Propanol, the subject of this study, occurs in five different conformers: \(\text{Ga, Gg, Gf, Aa, and Ag}\), as described by Abdurakhmanov et al. (1968, 1970, 1976); Dreizler & Scappini (1981); Lotta et al. (1984); Maeda et al. (2006); Kisiel et al. (2010) and references therein. The capital letters \(G\) and \(A\) refer to the rotation of the heavy nuclear plane \(C-C-C\) compared to \(C-C-O\). An “anti” (A) configuration\(^2\) describes a rotation by 180° and “gauche” (G) by about 60°. The small letters refer to the rotation of the dihedral angle of the hydroxy group (–OH).

The G family of conformers (\(Ga, Gg, Gf\)) of n-propanol has been rather extensively studied in rotational spectroscopy, as in, for instance, Abdurakhmanov et al. (1970); Maeda et al. (2006); Kisiel et al. (2010) and references therein. A single state analysis of \(Ga\) was able to fit 2865 lines up to 375 GHz, but systematic deviations occurred for some transitions with higher \(J\) quantum numbers (Maeda et al. 2006). Based on this observation, a combined analysis of the three \(G\) conformers was

\(^{1}\) The supplementary material, our predictions of the rotational spectra (*cat files), are only available at the CDS via anonymous ftp to cdsarc.u-strasbg.fr (130.79.128.5) or via http://cdsarc.u-strasbg.fr/viz-bin/cat/?/A+/662/A111

\(^{2}\) The term “trans” (T) was often used instead of “anti” historically, although it should only be used to refer to planar molecules.

Astronomy & Astrophysics

Open Access article, published by EDP Sciences, under the terms of the Creative Commons Attribution License (https://creativecommons.org/licenses/by/4.0/), which permits unrestricted use, distribution, and reproduction in any medium, provided the original work is properly cited.

This article is published under the Subscribe-to-Open model. Subscribe to A&A to support open access publication.
carried out by taking into account Coriolis interaction and Fermi resonances (Kisiel et al. 2010). Thanks to the quantum mechanical models of rotational spectra with experimental accuracy (Kisiel et al. 2010), conformers of the G family can be found in space. Furthermore, this distinctive analysis gives confidence that the G and A conformers can be treated separately as within the G family, the conformers seem to be unaffected by conformers of the A family (Aa and Ag).

The ab initio calculations have shown some scatter for the relative energies of n-propanol conformers, as, for instance, in Lotta et al. (1984); Kahn & Bruce (2005) and references therein (a detailed overview is given in Maeda et al. 2006, but they agree in the fact that all conformers, in particular from both G and A families, are rather close in energy). Considering only Boltzmann statistics, the Aa and Ag conformers are expected to have about the same chance of being detected in the warm interstellar medium as the G family conformers.

The most advanced rotational analysis of the Aa n-propanol was derived by using a Fourier transform microwave (FTMW) spectrometer (8–18 GHz) and a Stark modulated microwave spectrometer (up to 40 GHz) (Dreizler & Scappini 1981). Four conformers of n-propanol occur as mirror configurations (two-fold degeneracy due to “left” and “right-handed” versions, C₁ symmetry), except for the Aa conformer, which has a plane of symmetry and one-fold degeneracy (Cᵥ symmetry), as shown in Fig. 1. The Ag conformer occurs in mirror images (see Fig. 1), however, the barrier for hydrogen tunneling between the two equivalent Ag forms is low enough to distinguish symmetric Ag⁺ and antisymmetric Ag⁻ tunneling states (Abdurakhmanov et al. 1976; Kisiel et al. 2010). The tunneling transition state has Cᵥ symmetry and tunneling-rotation interactions need to be taken into account, as has been done, for instance, for gauche-ethanol (Pearson et al. 2008) and gauche-propanal more recently (Pickett & Scroggin 1974; Zingsheim et al. 2017, 2022). So far, only eleven doublet features up to 30 GHz were tentatively assigned to Ag⁺ and Ag⁻ tunneling states with J_max = 4 (Abdurakhmanov et al. 1976).

The main aim of the present study is to facilitate the astronomical detection of Aa n-propanol and Ag n-propanol. For this purpose, the rotational spectra are measured in selected frequency ranges from 18 to 505 GHz. The employed experimental setups are presented in Sect. 2. The experimental data are analyzed in Sect. 3 with a focus on secure assignments of the perturbed systems. Next, the spectroscopic results and the astrochemical search of n-propanol are discussed in Sect. 4. Finally, our main conclusions are given in Sect. 5.

2. Experimental techniques

We recorded rotational spectra of n-propanol (i) in the 17.5–26.5 GHz region using chirped-pulse Fourier transform microwave (CP–FTMW) jet spectroscopy and (ii) in the regions of 33–67 GHz, 70–129 GHz, and 160–505 GHz using conventional 2f demodulation absorption techniques. Additional (iii) double-modulation double-resonance (DM–DR) measurements support assignments of weaker transitions to the Aa conformer and, in particular, secure the identification of various ⁴R series of the Ag conformer. The (i) CP–FTMW, (ii) conventional (sub-)millimeter wave (MMW) absorption spectrometers, and (iii) the DM–DR setup are introduced in detail in (i) Hermanns et al. (2019), (ii) Ordu et al. (2019); Martin-Drumel et al. (2015), and (iii) Zingsheim et al. (2021), respectively. Thus, their main characteristics are only briefly summarized in the following.

In terms of (i) the 17.5–26.5 GHz range was covered in steps of 1 GHz bandwidth using a CP–FTMW spectrometer. In contrast to the above-described setup (Hermanns et al. 2019), mixing up the CP and mixing down the molecular signal is no longer required as the CP is generated directly in the target frequency region using an arbitrary waveform generator (Keysight M8195A 65 GS/s). In the present setup, each 1 GHz broad CP spanned 2 µs and was amplified with a 4 W amplifier. These amplified pulses were then broadcast via a horn antenna to excite the molecular ensemble. After a pulse was stopped, the resulting free induction decay (FID) of the molecules was collected with another horn antenna, amplified using a low noise amplifier, and recorded for 10 µs on a 100 GS/s oscilloscope. Each measurement consists of 65 000–100 000 acquisitions, which were performed with a 10 Hz repetition rate. The averaged FIDs were then Fourier transformed using a Hanning window, which resulted in a frequency resolution of about 290 kHz, considering the full width at half maximum (FWHM). The sample was prepared by mixing 1 mL of liquid n-propanol (Sigma-Aldrich, 97.7% purity) with 7 bar neon in a 10 L container. The mixture was supersonically expanded with a relative stagnation pressure of 1 bar and an opening duration of 500 µs.

For (ii) either an Agilent E8257D or a Rohde & Schwarz SMF100A synthesizer generated electromagnetic waves. Electromagnetic waves were directly coupled to a horn antenna to record single spectra in the 33–67 GHz region. We used in-house developed electronics for operating an amplifier tripler chain in full saturation mode to fully cover the 70–129 GHz region and used commercially available cascaded doublers and triplers from Virginia Diodes Inc. (VDI) to fully cover the 160–505 GHz region. Frequency modulation (FM) increases the signal-to-noise ratio (S/N) and the 2f demodulation creates absorption features that are close to the second derivative of a Voigt profile.

---

Footnote:

⁴ Transitions with identical selection rules are called a series, here abbreviated as ΔK_a,ΔJ, with ΔJ represented by capital letters P, Q, and R for ΔJ = −1,±0 or +1, respectively, and ΔK_a = −1,±0, or +1 by superscripts p, q, or r, respectively. For instance, a-type transitions (ΔK_a = 0) with ΔJ = +1 are called ⁴R series transitions. A subscript may be used additionally to highlight the respective K_a quantum number (⁴R_{K_a}, series).

A111, page 2 of 17
We used a modulation frequency of $f \approx 47$ kHz and its amplitude is on the order of the FWHM of the absorption profiles. The absorption path was 14 m in a single pass configuration for measurements up to 130 GHz (Ordu et al. 2019) and 10 m in double pass configuration otherwise (Martin-Drumel et al. 2015). Here, $n$-propanol was vaporized into the cell to pressures in the range of 10–40 µbar.

For (iii) we used both synthesizers, namely, Agilent as the so-called probe and Rohde & Schwarz as the pump, simultaneously for DM–DR measurements; a modulated pump radiation was applied in addition to the frequency modulated probe radiation. The probe signal was usually frequency tripled, and in-house developed electronics used for operation in full saturation mode ensured a stable output power of about 1 mW (0 dBm). The pump source, an active frequency multiplier (AFM6 70–110 + 14 kHz from RPG Radiometer Physics), delivered output powers of up to 60 mW (17.8 dBm). The final pump and probe frequencies were both located in the W-band. We performed some additional measurements with the probe frequency being around 200 GHz using commercial components by VDI. It is only the probe transitions that share an energy level with the pump transition that are affected by the additional source. An additional 1/f demodulation of the pumped signal is experimentally realizing a difference spectrum of the on- and off-resonant measurements, or, in other words, a subtraction of a conventional and a DR measurement. In this way, it is only the lines of interest remaining and the assignment of even weak or blended features is facilitated (Zingsheim et al. 2021).

3. Results: Spectroscopic signatures

The measurement results for $Aa$ and $Ag n$-propanol are presented in this section. The central focus of this chapter is the analyses and the resulting quantum mechanical models of rotational spectra, as well as the verification of assignments, since both conformers show perturbed patterns. Our assignment procedures benefited, in terms of speed and unambiguity, from the LLWP software, which is based on Loomis-Wood plots5 (Bonah et al. 2022). Adjacent transitions of certain series are plotted around their predicted frequency in Loomis-Wood plots to increase the confidence of correct assignments or to visualize systematic deviations between observed and predicted transition frequencies.

3.1. $Aa n$-propanol

The $Aa$ conformer is an asymmetric rotor quite close to the prolate limit with $\kappa = -0.978$ ($\kappa = [2B - A - C]/[A - C]$). The dipole moments were determined with Stark effect measurements to be $\mu_a = 0.21(7)$ D and $\mu_b = 1.54(2)$ D (Abdurakhmanov et al. 1970). In total, 73 lines (118 transitions) were assigned and fit up to 40 GHz in a previous study (Dreizler & Scappini 1981). Therein, all lines from Abdurakhmanov et al. (1968) were remeasured; $a$- and $b$-type transitions were assigned and fit and in particular the assignments of weak $a$-type transitions were secured with the help of MW–MW DR measurements (Dreizler & Scappini 1981). A splitting due to the internal rotation of the methyl group (–CH₃) could sometimes be resolved (14 A + 14 E transitions) and together with 45 blended doublet features, the transition frequencies were reproduced with a standard deviation of 195 kHz (Dreizler & Scappini 1981). Thereby, the potential barrier of the methyl group internal rotation was determined to be $V_3 = 955(21)$ cm⁻¹ (or 2730(60) cal mol⁻¹) and the angle of the internal rotation axis $i$ to the principal moment of inertia axis $a$ to be $\langle i,a \rangle = (29 \pm 1)^\circ$ (Dreizler & Scappini 1981).

We used the available parameters and transition frequencies (Dreizler & Scappini 1981) to set up an initial spectroscopic model of an asymmetric rotor, including three-fold internal rotation in ERHAM (Groner 1992, 1997, 2012). Thanks to the initial prediction and with the application of an iterative assignment and fit strategy, we could assign strong and also weak $b$-type transitions in a straightforward fashion (see Fig. 2). Transitions from Dreizler & Scappini (1981) were remeasured if possible6. Furthermore, we secured assignments of weak $a$-type transitions in the $W$-band region with the DM–DR technique. Additionally, transitions of the ‘$R$’ series ($J_{K_a,K_c} = J_{J,J-3} \leftarrow (J - 1)_{J,J+4}$) and the ‘$Q$’ series transitions in the $W$-band region with the DM–DR technique. Additionally, transitions of the ‘$R$’ series ($J_{K_a,K_c} = J_{J,J-3} \leftarrow (J - 1)_{J,J+4}$) and the ‘$Q$’ series transitions in the $W$-band region with the DM–DR technique. Additionally, transitions of the ‘$R$’ series ($J_{K_a,K_c} = J_{J,J-3} \leftarrow (J - 1)_{J,J+4}$) and the ‘$Q$’ series transitions in the $W$-band region with the DM–DR technique. Additionally, transitions of the ‘$R$’ series ($J_{K_a,K_c} = J_{J,J-3} \leftarrow (J - 1)_{J,J+4}$) and the ‘$Q$’ series transitions in the $W$-band region with the DM–DR technique. Additionally, transitions of the ‘$R$’ series ($J_{K_a,K_c} = J_{J,J-3} \leftarrow (J - 1)_{J,J+4}$) and the ‘$Q$’ series transitions in the $W$-band region with the DM–DR technique.

5 Download at https://llwp.astro.uni-koeln.de/

6 Two $Q$-branch $a$-type lines (or four transitions) were not included in the analysis of Dreizler & Scappini (1981), but were secured by them with MW–MW DR spectroscopy; These lines are successfully incorporated to the fit in this study. However, six other transitions have been unweighted in this study due to conspicuously large deviations.

Fig. 2. Loomis–Wood plot of $Aa n$-propanol – screenshot of the LLWP software (Bonah et al. 2022). Shown here is a doublet series ($A$ and $E$ components of the methyl group internal rotation) of the $J_{K_a,K_c} = J_{J,J-3} \leftarrow J_{J,J-2}$ $b$-type transitions from $J = 6$ (bottom panel) to $J = 36$ (top panel). Predictions of the $A$ components are set as center frequencies and are depicted by green sticks. The $E$ components are depicted by red ones. This $Q$ series with $K_a + K_c = J$ and $\Delta K_a, \Delta K_c = +1,-1$ is well reproduced by our spectroscopic model described in Table 2 up to $J = 22$, however, transitions up to $J = 41 (\Delta v \approx 90$ MHz) are straightforwardly assignable.
choice of fit transitions is somewhat ambiguous: the quantum number coverage is visualized in Fig. C.1. Larger systematic deviations, \( \Delta \nu = \nu_{\text{obs}} - \nu_{\text{calc}} \), occur from our final predictions for various series (cf. Fig. 2). These deviations start at similar \( J \) values for series involving identical quantum numbers, see Table 1. However, the assignment of these transitions is unambiguous as demonstrated in Fig. 2 and the assignment of internal rotation components is also secured (cf. Fig. 3). Finally, we could assign 1255 transitions with frequencies up to 505 GHz and \( J_{\text{max}} = 70 \), thereof, 928 transitions are well reproduced by the prediction. The determined spectroscopic parameters can be found in Table 2.

### 3.2. Ag \( n \)-propanol

The three members of the \( G \) family were extensively described in the literature plus the \( Aa \) conformer to some extent. The situation is different for the fifth conformer of \( n \)-propanol: \( Ag \). There is no spectroscopic model of rotational spectra with experimental accuracy available, and only 22 \( J_{K_g,K_c} \leftarrow (J-1)K_g,K_c-1 \) \( QR \) series transitions (1, 3, 5, and 2 transitions for \( J = 1, 2, 3, \) and 4, respectively, for each of the two tunneling states), were tentatively assigned in the MW region (Abdurakhmanov et al. 1976). \( \)In the upper panel. \( E \) transitions, which are forbidden in rigid rotors (Herschbach & Swalen 1958), secure the assignment of internal rotation components. We note that slightly off-resonant pumping leads to small, asymmetric features of \( A \) and \( E \) components in the DM–DR spectrum of the other one, as pumped \( A \) and \( E \) components are rather close in frequency, cf. Zingsheim et al. (2021).

The situation is different for the fifth conformer of \( n \)-propanol: \( Ag \). There is no spectroscopic model of rotational spectra with experimental accuracy available, and only 22 \( J_{K_g,K_c} \leftarrow (J-1)K_g,K_c-1 \) \( QR \) series transitions (1, 3, 5, and 2 transitions for \( J = 1, 2, 3, \) and 4, respectively, for each of the two tunneling states), were tentatively assigned in the MW region (Abdurakhmanov et al. 1976). \( \)In the upper panel. \( E \) transitions, which are forbidden in rigid rotors (Herschbach & Swalen 1958), secure the assignment of internal rotation components. We note that slightly off-resonant pumping leads to small, asymmetric features of \( A \) and \( E \) components in the DM–DR spectrum of the other one, as pumped \( A \) and \( E \) components are rather close in frequency, cf. Zingsheim et al. (2021).

### Table 1. Quantum number overview of the final analysis of \( Aa \) \( n \)-propanol.

| \( K_a \) \( (a) \) | \( K_a'' + K_c'' \) \( a \)-type \( (b) \) | \( b \)-type \( (b) \) |
|---|---|---|
| \( \pm 0, -1 \) | \( \pm 1, -3 \) | \( \pm 1, -1 \) |
| \( \pm 0, -1 \) | \( \pm 1, -1 \) | \( \pm 1, -3 \) |
| 0 | \( J = 16(16) \) | \( \rightarrow \) | \( 32(32) \) | \( 69(70) \) | \( 68(70) \) |
| 1 | \( J + 1 \) \( 17(17) \) | \( 7(7) \) | \( 28(28) \) | \( 20(20) \) | \( 29(29) \) |
| 1 | \( J = 16(16) \) | \( 7(7) \) | \( 31(41) \) | \( 31(53) \) | \( 41(53) \) |
| 2 | \( J + 1 \) \( 16(16) \) | \( 13(13) \) | \( 26(26) \) | \( 23(32) \) | \( 22(26) \) |
| 2 | \( J = 16(16) \) | \( 13(13) \) | \( 22(41) \) | \( 20(37) \) | \( 22(26) \) |
| 3 | \( J + 1 \) \( 16(16) \) | \( 9(25) \) | \( 9(12) \) | \( \rightarrow \) | \( \rightarrow \) |
| 3 | \( J = 16(16) \) | \( 9(11) \) | \( 9(13) \) | \( \rightarrow \) | \( \rightarrow \) |
| 4 | \( J + 1 \) \( 11(11) \) | \( \rightarrow \) | \( \rightarrow \) | \( \rightarrow \) | \( \rightarrow \) |
| 4 | \( J = 11(11) \) | \( \rightarrow \) | \( \rightarrow \) | \( \rightarrow \) | \( \rightarrow \) |

**Notes.** Values represent \( J_{\text{max}} \), the highest \( J \) quantum number of well reproduced transitions of a given series in our final spectroscopic model (Values in parentheses represent \( J_{\text{max}} \) of assigned transitions); \( Q \) transitions are neglected here. \( \)The given \( K_c \) is always the lower one of \( K_c' \) and \( K_c'' \). Selection rules for \( K_2 \) and \( K_c \) quantum numbers: \( \Delta K_2 = K_2' - K_2'' \), \( \Delta K_c = K_c' - K_c'' \). Upper values are for asymmetry side \( K_c'' + K_2'' = J'' + 1 \) and lower ones for \( K_c' + K_2' = J' + 1 \). Additionally, \( \Delta K_2 \) with \( -1, 0, +1 \) is also represented by the superscripts \( p, q, r \), respectively. The capital letters \( P, Q, R \), represent \( \Delta J \) with \( -1, 0, +1 \) respectively.

7 Allowed \( a \)- and \( b \)-type transitions occur within a tunneling state: \( Ag^+ \leftrightarrow Ag^+ \) and \( Ag^- \leftrightarrow Ag^- \), whereas \( c \)-type transitions occur between the tunneling states: \( Ag^+ \leftrightarrow Ag^- \).
doublet series for each $^4R_A$, $^4R_5$, $^4R_6$, $^4R_7$, $^4R_8$, and $^4R_9$ series ($4 \leq K_a \leq 9$) proving the existence of $0^+$ and $0^-$ tunneling states (see Fig. 4). The nomenclature $0^+$ and $0^-$ is used to highlight that the vibrational ground state $v = 0$ is observed. We assigned the tunneling state to transitions with $4 \leq K_a \leq 9$ in the way that we assumed $0^+$ transitions to be always lower in frequency than ones of $0^-$, which is usually the case if identical transitions are compared (transitions with same $J$'s, $K_a$'s, and $K_c$'s).

A further confirmation of unambiguous assignments is the observed asymmetry splitting for $^4R_A$, $^4R_5$, $^4R_6$, and $^4R_7$ series. The two asymmetry components, transitions with identical $J$ and $K_a$ but either $K_c = J - K_a$ or $K_c = J + K_a + 1$, are frequently blended for low $J$ and higher $K_a$ values (prolate pairing), but this degeneracy is lifted for higher $J$ (cf. colored lines in Fig. 4 which are split into two lines at higher $J$). The asymmetry splitting is the frequency difference of the two $\alpha$-type transitions with $J_{K_a,J-K_a}$ and $J_{K_a,J+K_c+1}$. The predicted asymmetry splittings, from the ab initio calculations of Kisiel et al. (2010), show the right order of magnitude when compared to the observed ones of both tunneling states; more importantly, they show similar systematic deviations for all $K_a$ numbers (see Fig. 5).

At this point, we tried to reproduce the determined frequencies of assigned transitions with $4 \leq K_a \leq 9$ by applying a common asymmetric rotor Hamiltonian, which is comparable to what has been done for the $Aa$ conformer (cf. Table 2); This time the SPFIT program was used (Pickett 1991). Many series seem to deviate from asymmetric rotor spectra. We expect identical $^4R_{K_c}$ series (same $K_a$ and same asymmetry side) of the two tunneling states to appear as nearly parallel lines in Fig. 4, whose spacing should decrease for increasing $K_a$ quantum numbers. This is not the case, for instance, the two series with $K_a = 8$ (purple lines) have quite different slopes and the tunneling splitting of series with $K_a = 6$ (green lines) is larger than that of the series with $K_a = 5$ (red lines). We tried to incorporate even order Coriolis-type parameters ($F_{ab}$, $F_{ac}$, and $F_{bd}$) to account for possible tunneling-rotation interaction between the two tunneling states, $Ag^+$ and $Ag^-$, as has been done successfully several times for similar molecular systems in the past; for instance, in Pickett (1972); Christen & Müller (2003); Pearson et al. (1997); Zingsheim et al. (2022). However, for the case of $Ag$ n-propanol, still no experimental accuracy for our predictions has been reached.

We continued assigning $J$, $K_a$, and $K_c$ quantum numbers to transitions with $K_a < 3$. However, transitions from one of the two series, $Ag^+$ and $Ag^-$, are not strictly lower in frequency than the other, in contrast to series with $4 \leq K_a \leq 9$. Crossing series are observed in Loomis-Wood plots (Figs. B.2–B.5). Therefore, the tunneling state assignments remain speculative for transitions with $K_a < 3$. The two tunneling series ($^4R$ series with identical $K_a$ and $K_c$ quantum numbers) with $K_a < 3$ are even less parallel than observed ones with $4 \leq K_a \leq 9$. We expect states with $K_a < 3$ to be strongly influenced by tunneling-rotation interaction between $Ag^+$ and $Ag^-$ tunneling states, cf. Fig. C.2.
4. Discussion

In this work, we present an extended spectroscopic quantum mechanical model based on additional assignments in the MMW region for Aa n-propanol and secured assignments of both tunneling states, Aa+ and Aa−, for Ag n-propanol. In the following, we discuss the spectroscopic results of the Aa and Ag conformers (Sects. 4.1 and 4.2). Subsequently, we offer some ideas to overcome certain limitations, specifically, thanks to a global analysis of both conformers (Sect. 4.3). Next, spectroscopic results of n-propanol are judged in regard of astronomical searches (Sect. 4.4). The rotational data are used in a companion article for its astronomical detection (Sect. 4.5). Moreover, we showed that DM–DR measurements can straightforwardly be extended to frequencies higher than the W-band, see Fig. 3, as was finally, we could not securely assign 9Rj transitions as too many candidate series are close in frequency (Fig. B.6).

In summary, we could prove the existence of the two tunneling states Ag+ and Ag− and assign lines with Jmax = 69 and Kmax = 9. A quantum number overview of assigned transitions is visualized in Fig. C.1. Deriving a spectroscopic quantum mechanical model with experimental accuracy is beyond the scope of this work because of the multiple perturbations whose origins are unknown at present.

Table 2. Spectroscopic parameters of Aa n-propanol.

| Parameter | Theory Ref. K10 | Experiment Ref. DS81 | This work |
|-----------|----------------|----------------------|-----------|
| A         | 26 579         | 26 401.671(50)       | 26 401.4395(40) |
| B         | 3784           | 3802.154(11)         | 3802.160 25(39) |
| C         | 3531           | 3549.543(20)         | 3549.454 27(33) |

| Parameter | Theory Ref. K10 | Experiment Ref. DS81 | This work |
|-----------|----------------|----------------------|-----------|
| ΔJ        | 5.579          | 141.5(47)            | 149.31(55) |
| ΔJK       | 3.25           | 9.579(99)            | -0.3608(295) |
| ΔJ        | 0.828          | 1.672(73)            | 0.9907(48) |
| δK        | 0.361          | -29.6(69)            | -3.250(164) |
| δJ        | 0.0737         | 0.0943(45)           | 0.129324(170) |

ΦK = 10^6, ΦJK = 10^6, ΦJ = 10^6, Φq = 10^6, φq = 10^6, φJ = 10^6

| Parameter | Theory Ref. K10 | Experiment Ref. DS81 | This work |
|-----------|----------------|----------------------|-----------|
| J         | 113            | 928                  |           |
| LJK       | 73             | 689                  |           |
| rms J/kHz | 195            | 48                   |           |
| urms (a)  | 1.07           |                      |           |

Notes. Watson’s A reduction Hamiltonian of asymmetric rotors is used with ERHAM. A, B, and C are rotational constants; Δ’s, Φ’s, and L’s are quartic, sextic, and octic centrifugal distortion constants, respectively. The angle <i,a>, the unitless ρ value, and ε10 describe splittings due to the internal rotation of the –CH₃ group. 9 We remeasured 34 transitions of Dreizler & Scappini (1981) and unweighted 6 Q-branch a-type transitions. In total, we assigned 1255 transitions.

References. K10 (Kisiel et al. 2010); DS81 (Dreizler & Scappini 1981).
demonstrated in the original DM–DR manuscript (Zingsheim et al. 2021).

4.1. Aa n-propanol

We extended spectroscopic assignments of Aa n-propanol up to 505 GHz with \( J_{\text{max}} = 70 \) and \( K_{\text{v, max}} = 6 \). Finally, 928 transitions are well reproduced by the spectroscopic parameters in Table 2. Larger systematic deviations \( \Delta \nu \) from our predictions occur for certain series (cf. also Fig. 2); deviations start from transitions with similar \( J \) when series with identical \( K_v \) values are compared. The deviations occur at rather low \( K_v \) values and the origins of these perturbations are unknown so far. Transitions involving \( K_v = 0 \) and 1 seem to be unperturbed as \( b \)-type transitions involving \( K_v = 1 \leftrightarrow 0 \) are well reproduced in our final spectroscopic model. On the other hand, \( b \)-type transitions with \( K_v = 2 \leftrightarrow 1, K_v = 3 \leftrightarrow 2, \) and \( K_v = 4 \leftrightarrow 3 \) are only well reproduced up to around \( J = 30, J = 20 \), and \( J = 10 \), respectively (see Table 1). This case is similar to that of \( a \)-ethanol, as discussed for the \( ^{13} \)C isotopomers (Bouchez et al. 2012).

If different states are interacting, consecutive transitions of one series are frequently perturbed. In particular, systematic deviations occur around the strongest perturbed energy levels. So-called avoided crossing patterns may be observed; in this case, transitions belonging to the same series with somewhat smaller and larger quantum numbers than heavily perturbed ones can often be assigned with confidence and, more importantly, reproduced to experimental accuracy without treating the interaction. In this way, perturbed energy levels can be located and transitions from these levels can be unweighted in the analysis if the origin of the interaction is unknown or indescribable. For instance, at the start of the analysis of gauche-propanol, certain \( \nu = 0 \) series had to be unweighted (Zingsheim et al. 2017), but they could be properly described by including Fermi resonances and Coriolis interaction in a global analysis with \( \nu_{24} = 1 \) (Zingsheim et al. 2022). In the case of Aa n-propanol, no transitions with higher \( J \) values than the ones that deviate from the prediction could be fit to experimental accuracy; the strongest perturbed energy levels and therefore the center of the perturbation, or the strongest perturbed levels and their exact energies, are unknown to this point. Therefore, we expect that our final spectroscopic model is effective since already perturbed transitions may be included in the line list.

4.2. Ag n-propanol

We thoroughly examined assignments of Ag n-propanol in the MW and MMW regions and proved the existence of the two tunneling states \( Ag^+ \) and \( Ag^- \). First, DM–DR measurements linked various \( ^1R \) series transitions and assignment of these series was undertaken with the support of the LLWPF software. We could not confirm doublet patterns from the literature (cf. Fig. 2). The ab initio calculations (Kisiel et al. 2010) allowed us to pinpoint patterns and, in particular, the verification of assignments by studying the asymmetry splitting, cf. Fig. 5. In this way, assignments up to \( J_{\text{max}} = 69 \) and \( K_{\text{v, max}} = 9 \) with frequencies up to 505 GHz have been achieved. A derivation of a spectroscopic quantum mechanical model to fit the experimental data accurately is beyond the scope of this work.

4.3. Global analysis: Aa, Ag+, and Ag− conformers

In the future, Aa n-propanol should be described in a more sophisticated spectroscopic quantum mechanical model and an initial description of Ag n-propanol with experimental accuracy should be derived. This may be possible only through a global analysis of Aa, Ag+, and Ag− conformers. The rotational energy levels of the \( G \) family seem to be isolated from \( A \) conformers ones (Kisiel et al. 2010), therefore, an additional incorporation of \( G \) conformers in a global analysis is expected to be unnecessary.

A global analysis of the \( A \) family is expected to be successful only if Coriolis interaction and Fermi resonance between the \( Aa \) conformer and \( Ag^+ \) or \( Ag^- \) tunneling states and, very importantly, also tunneling-rotation interactions between the \( Ag^+ \) and \( Ag^- \) tunneling states are considered. The first interaction term which should be investigated are the ones used to describe the system anti, gauche−, and gauche+ ethanol (Pearson et al. 2008), as the group theoretical considerations and probably also the energy splittings are similar. A quantum mechanical treatment taking interactions into account would also allow to determine relative energies as has been done for the \( G \) family (Kisiel et al. 2010). Furthermore, the potential of the –OH group rotation may be studied in detail by taking also higher vibrationally excited states into account.

4.4. Astronomical search of n-propanol

The partition function used for creating the predictions of all five conformers, given in the supplementary material, is \( Q_{GS} = 513984.0861 \) (log \( Q = 5.7109 \)) at 300 K. This value is derived from the extensive parameter set from Kisiel et al. (2010) and takes the three \( G \) conformers and their relative energies into account. Thereby, the degeneracy factor was set to 4 (2 for the internal rotation of the –CH\(_3\) group plus 2 when taking the mirror images into account). Conformational correction factors taking into account the existence of the \( A \) conformers as well as vibrational correction factors, can be found in Table A1.

Predictions of \( Ga, Gg, \) and \( Gg' \) are based on the parameter set of Kisiel et al. (2010). We note that the relative energy of \( Ga \) was set to \( E_{Ga} = 0 \) cm\(^{-1}\) (consequently, \( E_{Gg} = 47.8 \) cm\(^{-1}\) and \( E_{Gg'} = 50.8 \) cm\(^{-1}\)), instead of using \( Gg \) as reference point, as done in the original data. This adaption is inevitable to derive meaningful intensities and lower state energies in the predictions. Predictions in Pickett’s SPCAT format (*.cat files; Pickett 1991) can be found in the supplementary material and will be added to the CDMS as well.

The predictions of Aa n-propanol are limited to transitions with \( J \) up to 70, 30, 20, and 10 for \( K_a = 1, 2, 3, \) and 4, respectively. All assigned transitions with \( K_a \leq 5 \) can be used in astronomical observations, but measured transition frequencies should replace predicted ones. Transitions with \( K_a > 5 \) should not be used in astronomical observations due to the limitations of our spectroscopic quantum mechanical model. The intensities and lower state energies in our predictions were corrected by an assumed relative energy of \( E_{Aa} = 30 \) cm\(^{-1}\). Overall, the number of unambiguously assigned transitions up to 505 GHz

---

8 Future analyses may also benefit by taking into account preliminary results of Kisiel et al. (2010), presented at the International Symposium on Molecular Spectroscopy (ISMS) 2006. We were made aware of this after submission; assignments are not publicly available.

9 Caution is advised concerning the number of interaction parameters; Using too many leads to high correlations and convergence problems. Using too few may lead to a more stable fit, but may not account for all perturbations, or, if it does, at the expense of unphysical parameters.

10 Our predictions (*.cat files) will be available together with further information; cf. Appendix A, in the CDMS soon.
and the quantum number coverage for low \(K_a\) values is sufficient to search for rotational transitions of the \(Aa\) conformer in space.

An astronomical detection of \(Ag\) \(n\)-propanol is facilitated thanks to many unambiguously assigned transitions. For creating meaningful predictions, despite the fact that a quantum mechanical model with experimental accuracy is missing, we used, in a first step, ab initio calculations that are based on Kisiel et al. (2010) to derive theoretical predictions. Even though predicted frequencies do not match experimental accuracy at all, meaningful physical parameters, such as the intensities, lower-state energy, and degeneracies of transitions are obtained by assuming a relative energy of \(E_{Ag} = 45.5\) cm\(^{-1}\). Finally, we replaced the predicted frequencies by the measured ones and neglected all unassigned transitions. In doing so, an astronomical detection of \(Ag\) \(n\)-propanol is possible, but deriving a proper spectroscopic quantum mechanical model should be done in particular to derive relative energies and to enlarge the number of usable transitions.

4.5. Astronomical detection of \(n\)-propanol

Some members of our team searched for distinctive rotational signatures of all five conformers of \(n\)-propanol in a companion astronomical study that uses the survey called Re-exploring Molecular Complexity with ALMA (ReMoCa; details given in Belloche et al. 2019) toward Sagittarius (Sgr) B2(N) (Belloche et al. 2022). This astronomical study reports the identification of \(n\)-propanol toward the secondary hot molecular core, Sgr B2(N2), thanks to the spectroscopic predictions from this work. The detection relies on eight and five lines of the \(Gg^j\) and \(Ag\) conformers, respectively. The catalogs produced in this work were instrumental in securing this first astronomical detection of \(n\)-propanol in a hot core.

5. Conclusion

We recorded spectra of \(n\)-propanol in the region of 18–505 GHz, which allowed us to analyze the rotational signatures of its \(Aa\) and \(Ag\) conformers, in addition to the already extensively studied \(Ga\), \(Gg\), and \(Gg^j\) conformers (Kisiel et al. 2010). Therefore, rotational signatures of all five conformers of \(n\)-propanol are available at present and the main conclusions of this work are as follows:

1. An extended quantum mechanical model of rotational spectra of \(Aa\) \(n\)-propanol has been made available.
2. The existence of the two tunneling states, \(Ag^+\) and \(Ag^-\), is proven for the first time by unambiguously assigning transitions of \(Ag\) \(n\)-propanol.
3. A global analysis of \(Aa\) and \(Ag\) conformers is recommended if the spectroscopic description of the \(Aa\) conformer needs to be extended or if a quantum mechanical model with experimental accuracy of the \(Ag\) conformer is the target.
4. Predictions of the rotational spectra of all five conformers of \(n\)-propanol (*.cat files in the supplementary material) enable their astronomical search.

In addition, the spectroscopic analysis presented in this work allowed for the astronomical detection of \(n\)-propanol toward a hot core in a companion article (Belloche et al. 2022).

Acknowledgements. We would like to thank Dr. Matthias Justen and Dr. Leonid Surin for their help in translating Russian manuscripts and providing literature data from the VINITI database. B.H. acknowledges financial support by the Deutsche Forschungsgemeinschaft (DFG; project ID WE 5874/1-1). This work has been supported via Collaborative Research Center 956, sub-projects B3 and B4, and the “Cologne Center for Terahertz Spectroscopy”, both funded by the DFG (project IDs 18408867 and SCHL 341/15-1, respectively).

References

Abdarakhmanov, A., Rahimova, R., & Imanov, L. 1968, Izv. Akad. Nauk. Azerb. SSR, 3, 15
Abdarakhmanov, A., Rahimova, R., & Imanov, L. 1970, Phys. Lett. A, 32, 123
Abdarakhmanov, A., Ragimova, R., Veliyulin, E., & Imanov, L. 1976, VINITI (a Russian database), 5835-76, 6
Ball, J. A., Gottlieb, C. A., Lilley, A. E., & Radford, H. E. 1970, Astrophys. J. Lett., 162, L203
Belloche, A., Garrod, R. T., Müller, H. S. P., et al. 2019, A&A, 628, A10
Belloche, A., Garrod, R. T., Zingsheim, O., Müller, H. S. P., & Menten, K. M. 2022, A&A, 662, A110
Bonah, L., Zingsheim, O., Müller, H. S., Guillemin, J.-C., & Schlemmer, S. 2022, J. Mol. Spectrosc., submitted
Bouchez, A., Walters, A., Müller, H. S., et al. 2012, J. Quant. Spectr. Rad. Transf., 113, 1148
Christen, D., & Müller, H. S. P. 2003, Phys. Chem. Chem. Phys., 5, 3600
Dreizler, H., & Scappini, F. 1981, Z. Naturforsch. A, 36, 1187
Endres, C. P., Schlemmer, S., Schilke, P., Stutzki, J., & Müller, H. S. 2016, J. Mol. Spectr., 327, 95
Fukushima, K., & Zvolinski, B. J. 1968, J. Mol. Spectr., 26, 368
Groner, P. 1992, J. Mol. Spectr., 156, 164
Groner, P. 1997, J. Chem. Phys., 107, 4483
Groner, P. 2012, J. Mol. Spectr., 278, 52
Herrmanns, M., Wehres, N., Lewen, F., Müller, H., & Schlemmer, S. 2019, J. Mol. Spectr., 358, 25
Herschbach, D. R., & Swalen, J. D. 1958, J. Chem. Phys., 29, 761
Hollis, J. M., Lovas, F. J., Jewell, P. R., & Coudert, L. H. 2002, ApJL, 571, L59
Johnson, D. R., Lovas, F. J., Gottlieb, C. A., et al. 1977, ApJ, 218, 370
Jørgensen, J. K., van der Wiel, M. H. D., Coutens, A., et al. 2016, A&A, 595, A117
Kahn, K., & Bruce, T. C. 2005, ChemPhysChem, 6, 487
Kisiel, Z., Dorosch, O., Maeda, A., et al. 2010, Phys. Chem. Chem. Phys., 12, 8329
Lotta, T., Murto, J., Räisänen, M., & Aspiaula, A. 1984, Chem. Phys., 86, 105
Maeda, A., Lucia, F. C. D., Herbst, E., et al. 2006, ApJS, 162, 428
Manigand, S., Jörgensen, J. K., Calculat, H., et al. 2020, A&A, 635, A48
Martin-Drumel, M.-A., van Wijngaarden, J., Zingsheim, O., et al. 2015, J. Mol. Spectr., submitted
Müller, H. S., Thornton, S., Roth, D., & Winnewisser, G. 2005, J. Mol. Struct., 742, 215
Ordu, M. H., Zingsheim, O., Belloche, A., et al. 2019, Astron. Astrophys., 629, A72
Pearson, J. C., Sastry, K. V. L. N., Herbst, E., & Lucia, F. C. D. 1997, ApJ, 480, 420
Pearson, J. C., Brauer, C. S., & Drouin, B. J. 2008, J. Mol. Spectr., 251, 394
Pickett, H. M. 1972, J. Chem. Phys., 56, 1715
Pickett, H. M. 1991, J. Mol. Spectr., 148, 371
Pickett, H. M., & Scroggin, D. G. 1974, J. Chem. Phys., 61, 3954
Rubin, R. H., Swenson, G. W., J., Benson, R. C., Tigelaar, H. L., & Flygare, W. H. 1971, ApJL, 169, L99
Tercero, B., Cernicharo, J., López, A., et al. 2015, A&A, 582, L1
Tercero, B., Cuadrado, S., López, A., et al. 2015, A&A, 582, L1
Tercero, B., Cernicharo, J., López, A., et al. 2018, A&A, 620, L6
Zingsheim, O., Müller, H. S. P., Lewen, F., Jörgensen, J. K., & Schlemmer, S. 2017, J. Mol. Spectr., 342, 125
Zingsheim, O., Bonah, L., Lewen, F., et al. 2021, J. Mol. Spectr., 381, 111519
Zingsheim, O., Müller, H. S., Bonah, L., et al. 2022, J. Mol. Spectr., 384, 111565
Zuckerman, B., Turner, B., Johnson, D., et al. 1975, ApJ, 196, L99
Appendix A: Partition function $Q$ of n-propanol

We calculated the partition function of the three G conformers $Q_{Gx}$ with SP-CAT to be $513984.0861$ at 300 K, which takes into account the $G_a$, $G_g$, and $G_{g'}$ conformers and a degeneracy factor of 4 (a factor of 2 for the internal rotation of the $-\text{CH}_3$ group and another factor of 2 for mirror images of the respective conformers; see next paragraph for more details). This value is based on the quantum mechanical description from Kiesiel et al. (2010).

The spectroscopic quantum mechanical models of $Aa$ and $Ag$ conformers are not as accurate as the ones of the $G$ family and the relative energies of $E_{Aa}$ and $E_{Ag}$ to $E_{Gx}$ are somewhat uncertain, thus, we used the partition function derived from the $G$ family $Q_{Gx}$ for creating predictions (*.cat files) of all five conformers. For the predictions of $Aa$ and $Ag$ conformers, the degeneracy factor is 2 and 4, respectively. Both conformers have an internal rotor (A and E transition, whereby both components are blended for the Agd conformer), but the $Aa$ conformer does not occur in a mirror configuration as the other four conformers do. In this way, we derived correct relative intensities of all five conformers.

For predicting correct absolute intensities, which are fundamental for deriving meaningful column densities of molecules in space, the existence of both A conformers and all vibrational states need to be taken into account in order to derive $Q_{Ta}$. Therefore, we present conformational and vibrational correction factors in Table A.1. The conformational correction factors are derived as follows:

$$F^T_{\text{conf}} = \frac{Q^T_{Ga(1)} \cdot f_{Aa} \cdot f_{Aa} \cdot BF^T_{Ar} + Q^T_{Ga(1)} \cdot f_{Ag} \cdot f_{Ag} \cdot BF^T_{Ag} + Q^T_{Gx}}{Q^T_{Gx}}$$

where the partition function $Q^T_{Ga(1)}$ is calculated with degeneracy factor 1, for instance, to be $Q_{Ga(1)} = 49318.3481$ at 300 K. The first factor $f = 2$ takes the internal rotation into account and the second factor, namely, $f = 1$ or 2, the conformer multiplicity of Aa and Ag, respectively. The Boltzmann factor $BF^T_{\text{conf}} = \exp(-E_{\text{conf}}/k_BT)$ of both conformers is derived by assuming $E_{Aa} = 30.0 \text{ cm}^{-1}$ and $E_{Ag} = 45.5 \text{ cm}^{-1}$. The vibrational correction factors were calculated using fundamental vibrational frequencies (with $\nu \leq 1200 \text{ cm}^{-1}$) of the $Aa$ conformer (Fukushima & Zwolinski 1968), that is, $\nu_{30} = 131 \text{ cm}^{-1}$, $\nu_{29} = 240 \text{ cm}^{-1}$, $\nu_{28} = 286 \text{ cm}^{-1}$, $\nu_{27} = 332 \text{ cm}^{-1}$, $\nu_{26} = 463 \text{ cm}^{-1}$, $\nu_{25} = 758 \text{ cm}^{-1}$, $\nu_{24} = 858 \text{ cm}^{-1}$, $\nu_{23} = 898 \text{ cm}^{-1}$, $\nu_{22} = 971 \text{ cm}^{-1}$, $\nu_{21} = 1013 \text{ cm}^{-1}$, $\nu_{20} = 1047 \text{ cm}^{-1}$, and $\nu_{19} = 1066 \text{ cm}^{-1}$. Applying the correction factors from Table A.1 allows to derive accurate column densities of n-propanol and derive relative abundances to other molecular species in astronomical sources.

Appendix B: Rotational signatures of Ag n-propanol

First, a part of the MW spectrum of n-propanol and propanal, depicting the $3_{K_a,K_a} \leftrightarrow 2_{K_a,K_a-1}$ a-type transitions, recorded with the CP-FTMW spectrometer, are shown in Fig. B.1. Rotational signatures are shown in the frequency regions 21.5--22.5 GHz and 25.0--26.0 GHz, respectively. Both measurements were performed under similar conditions. The rotational spectrum of gauche-propanol is well-understood (Zingsheim et al. 2022), in particular doublet structures due to two tunneling states are observed: Two $K_a = 1$ transitions which are close in frequency and have similar intensities. We note that tunneling-rotation interaction mixes the wave functions of involved rotational energy levels, in this case intensity borrowing may play a role and nominally forbidden transitions can be observed (cf. transitions marked with $\nu = 0' \leftrightarrow 0$'). The interpretation of the rotational spectrum of Ag n-propanol is more complicated. Assigned doublet transitions for Ag n-propanol from Abdurakhmanov et al. (1976) are marked in red, but cannot be verified in our CP-FTMW jet experiment. However, two transitions with similar intensities close in frequency are found twice. Comparison of the rotational spectra of both molecules leads to tentative assignments of $K_a = 1$ transitions of Ag n-propanol. The appearance of two transitions close in frequency were first hints of the existence of two tunneling states ($Ag^{+}$ and $Ag^{−}$), similar to what is observed for gauche-propanol. The different tunneling splittings of transitions with $K_a = 1$ for Ag n-propanol suggest that involved rotational energy levels are perturbed due to so far untreated interactions. We note that the transition marked with an asterisk (*) is another possible candidate for a transition with $K_a = 1$, if $^3R$ series are tracked, cf. Fig. B.4. Unambiguously assigning these transitions can be done if a spectroscopic quantum mechanical model with experimental accuracy is derived.

Second, typical spectroscopic signatures of $^3R_{K_a}$ series transitions allow for unambiguous assignments of the $Ag$ conformer, despite a missing quantum mechanical description. A broader Fortrat diagram than shown in Fig. 4 is presented in Fig. B.2. Then, Loomis-Wood plots are shown for series with $0 \leq K_a \leq 3$ to further document these assignments (see Figs. B.3--B.6). Series with $4 \leq K_a \leq 9$, which have been assigned in this work as well, are already well represented by Fig. 4 in the main article. The Loomis-Wood plots are all screenshots of the LLWP software used for assigning transitions within this work (Bonah et al. 2022).

Appendix C: Overview of the rotational energy levels of Aa and Ag n-propanol

We present a quantum number overview of our analysis for $Aa$ and $Ag$ conformers of n-propanol here as we could not reproduce all assigned transitions to experimental uncertainty. However, all assigned transitions can be used for astronomical searches of the rotational spectra of both conformers, therefore, quantum number overviews are given for clarification (see Fig. C.1). Furthermore, an estimated reduced energy diagram of Ag n-propanol is shown, which illustrates the location of tunneling-rotation interactions (see Fig. C.2). Of course, interactions between rotational energy levels of Ag n-propanol and those of Aa n-propanol should also be considered if a global analysis is performed, but these interactions are not illustrated here.

---

11 Reducing the total degeneracy factor of 4 is not possible if only one column density for all five conformers of n-propanol are to be derived, e.g., in local thermal equilibrium conditions in space, as the $Aa$ conformer does not exist in a mirror image form and A and E internal rotation components are resolved.
Table A.1: Partition function $Q_G$ of all three $G$ conformers (with spin weight 4), $Q_{Ga(1)}$ of $Ga$ (with spin weight 1), together with conformational correction factors $F_{\text{Conf}}$, vibrational correction factors $F_{\text{Vib}}$, and their product $F_{\text{Tot}}$ for various temperatures of interest for an astronomical detection.

| $T$ (K) | $Q_G$ | log $Q_G$ | $Q_{Ga(1)}$ | $F_{\text{Conf}}$ | $F_{\text{Vib}}$ | $F_{\text{Tot}}$ |
|---------|-------|-----------|-------------|-------------------|----------------|----------------|
| 300     | 513984.0861 | 5.7109 | 49388.3481 | 1.4748 | 6.4531 | 9.5167 |
| 225     | 318241.7669 | 5.5028 | 32031.6352 | 1.4671 | 3.2754 | 4.8055 |
| 200     | 260495.0562 | 5.4158 | 26839.0622 | 1.4631 | 2.6370 | 3.8583 |
| 180     | 217323.4156 | 5.3371 | 22911.7137 | 1.4590 | 2.2291 | 3.2523 |
| 160     | 177034.0151 | 5.2481 | 19197.9679 | 1.4537 | 1.8966 | 2.7571 |
| 150     | 158035.3077 | 5.1988 | 17425.1609 | 1.4504 | 1.7549 | 2.5454 |
| 140     | 139840.8805 | 5.1456 | 15710.7948 | 1.4466 | 1.6280 | 2.3551 |
| 120     | 106007.9447 | 5.0253 | 12465.7151 | 1.4367 | 1.4146 | 2.0325 |
| 100     | 75871.4993 | 4.8801 | 9481.9699  | 1.4221 | 1.2499 | 1.7775 |
| 75      | 44088.0429 | 4.6443 | 6158.4144  | 1.3905 | 1.1058 | 1.5376 |
| 37.5    | 11387.5229 | 4.0564 | 2178.3969  | 1.2546 | 1.0067 | 1.2630 |
| 18.75   | 3228.9908 | 3.5091 | 771.4138   | 1.0769 | 1.0000 | 1.0770 |
| 9.375   | 1095.9636 | 3.0398 | 273.6960   | 1.0059 | 1.0000 | 1.0059 |
Fig. B.1: Microwave spectra of $n$-propanol and propanal. Assignments of $a$-type transitions with $3_{K_a,K_c} \leftarrow 2_{K_a,K_c-1}$ are shown for both molecules ($K_a$’s are given in the figure). **Top panel:** Rotational spectrum of gauche-propanal with assignments from Zingsheim et al. (2022) marked in blue. **Bottom panel:** Assigned doublet transitions for $Ag$ $n$-propanol from Abdurakhmanov et al. (1976) are marked in red, but cannot be verified in our CP-FTMW jet experiment. Our tentative assignments of $K_a = 1$ transitions of $Ag$ $n$-propanol are marked in blue. We note that the transition marked with an asterisk (*) is another possible candidate for a transition with $K_a = 1$, more information can be found in the text.
Fig. B.2: Fortrat diagram of \(n\)-propanol. The colored lines in the Fortrat diagram mark linkages of \(^4R_k\) series transitions of \(Ag\) \(n\)-propanol, \(J_{K_aK_c} \leftarrow (J-1)_{K_aK_c-1}\). DM-DR spectroscopy secured assignments in the W-band region (from \(J = 10\) to \(J = 16\)). Tentative assignments of \(K_a = 1\) transitions derived from the CP-FTMW measurement (Fig. B.1) are marked by the blue arrows. Considering found \(^4R_1\) series, another candidate transition is marked by the blue asterisk. A quantum mechanical model with experimental accuracy can solve the ambiguity of these tentative assignments. The gray area depicts the region of the Fortrat diagram in Fig. 4.

Fig. B.3: Loomis-Wood plot around the assigned \(^4R_0\) series transitions in the W-band region of the \(Ag\) conformer, which are shown by red sticks.
Fig. B.4: Loomis-Wood plot around the assigned \(^9R_1\) series transitions in the W-band region of the \(A_g\) conformer, which are shown by red sticks. The blue asterisk (*) marks the same transition as is marked in Fig. B.1.
Fig. B.5: Loomis-Wood plot around the assigned $^4R_2$ series transitions in the W-band region of the $Ag$ conformer, which are shown by red sticks. The series higher in frequency with $K_a + K_c = J + 1$ is probably strongly interacting with a so far unknown state around $J = 12$. 

(a) $K_a + K_c = J + 1$

(b) $K_a + K_c = J$
Fig. B.6: Loomis-Wood plot around assumed \( {^9}R_3 \) series transitions in the W-band region. No assignments with \( K_a = 3 \) are secured yet, but the existence of such can be seen by the many unassigned transitions in the shown region. Four series close in frequency are expected; Two asymmetry sides for both tunneling states.
A111, page 16 of 17

Fig. C.1: Quantum number overview of the analysis of (a) Aa n-propanol and (b) Ag n-propanol. All rotational energy levels of assigned transitions are visualized with the $J$ quantum numbers on the x-axis, $K_a$ quantum numbers on the y-axis, and $K_c$ quantum numbers are either depicted by squares or circles for $K_c = J - K_a$ or $K_c = J - K_a + 1$, respectively. Filled markers depict quantum numbers of assigned transitions whose frequencies are fit to experimental accuracy using our final spectroscopic model in Table 2. The empty circles mark assigned transitions which are not included in the final analysis, but can be used for astronomical searches. For Aa n-propanol, we note that we do not differentiate between A and E components of each rotational energy level as in overwhelming majority of cases both components are assigned per transition. For Ag n-propanol, the rotational energy levels of the $0^+$ and $0^-$ tunneling states are given in blue and red, respectively. Triangles are used for transitions whose tunneling state assignments have only been guessed so far.
Interactions occur in this region and determining an accurate energy difference is essential to properly take them into account. Even though the exact energy difference between the two tunneling states is unknown so far, many tunneling-rotation interactions occur in this region and determining an accurate energy difference is essential to properly take them into account.

The energies of the rotational levels of the two tunneling states will differ. We set the energy difference of the two tunneling states to 3 cm\(^{-1}\) (a) Reduced energy diagram for rotational energy levels with \(K_a = \leq 9\).

The reduced energy diagram for rotational energy levels with \(K_a = \leq 3\).

Fig. C.2: Reduced energy diagram of \(Ag\) n-propanol. The reduced energy is \(E_{\text{red}} = E_r + E_{\text{rot}} - J \cdot (J+1) \cdot (B+C)/2\). Thereby, the energies of the rotational levels \(E_{\text{rot}}\) are calculated based on ab initio calculations of Kisiel et al. (2010). We used the same parameters for both tunneling states as ab initio calculations are presented for a single state only. We note that actual parameters of the two tunneling states will differ. We set the energy difference of the two tunneling states to 3 cm\(^{-1}\), i.e. \(E_T^{0+} = 0 \text{ cm}^{-1}\) and \(E_T^{0-} = 3 \text{ cm}^{-1}\), on the assumption that the splitting is similar to what has been observed for gauche-ethanol (Pearson et al. 1997). (a) The reduced energy diagram for rotational energy levels with \(K_a < 9\) shows that series with \(K_a > 5\) may not cross each other. We note that it may happen if different rotational parameters for the two tunneling states are taken into account. (b) The reduced energy diagram for rotational energy levels with \(K_a \leq 4\) is a zoom of the y-axis of (a). We note that many series cross each other and even though the exact energy difference between the two tunneling states \(Ag^+\) and \(Ag^-\) is unknown so far, many tunneling-rotation interactions occur in this region and determining an accurate energy difference is essential to properly take them into account.