Weak and strong interactions between dark solitons and dispersive waves

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The effect of mutual interaction between dark solitons and dispersive waves is investigated numerically and analytically. The condition of the resonant scattering of dispersive waves on dark solitons is derived and compared against the results of numerical simulations. It is shown that the interaction with intense dispersive waves affects the dynamics of the solitons strongly changing their frequencies and accelerating or decelerating the solitons. It is also demonstrated that two dark solitons can form a cavity for dispersive waves bouncing between the two dark solitons. The differences of the resonant scattering of the dispersive waves on the dark and bright solitons are discussed. In particular we demonstrate that two dark solitons and dispersive wave bouncing in between them create solitonic cavity with convex “mirrors” unlike the concave “mirror” in case of the bright solitons.

INTRODUCTION

Optical dark solitons were predicted [1] and experimentally realized [2] several decades ago and since then attracted considerable attention [3–9] in optical community. The interest to the dark solitons is motivated by their fascinating fundamental properties and by possible applications for telecommunications [3–4]. Dark solitons exist in the normal dispersion regime and they are characterized by the dip on top of the flat intensity background. Vectorial dark solitons [10] as well as coupled bright and dark solitons were considered [11–12]. Also the dynamics of dark solitons in presence of higher order dispersive effects was well covered [3–13]. Several studies have also addressed emission of dispersive waves (DWs) by dark solitons [13–15] as well as emission of radiation by shock waves [16] propagating in normal dispersion region.

Discussing the interaction of dispersive waves and bright solitons it is relevant to remind that extensive studies were performed on scattering of DWs on bright solitons, with solitons properties remaining invariant [18–22]. In particular it was also demonstrated that strong resonant scattering of DWs on solitons can accelerate or decelerate the solitons with consequent frequency up- or downshifts [23, 24]. This mechanism can be viewed as a method of all-optical switching [24–26] or alternative technique for generation of broad and coherent supercontinuum [25]. A very interesting scenario occurs when DWs emitted by one soliton interacts with a neighboring solitons through the dispersive waves trapped between the solitons. The multiple scattering of DW on solitons results in the mutual attraction of the solitons and eventually to their collision [27]. Furthermore, these collisions can cause fusion of two solitons [28]. This phenomenon was proven to be responsible for the appearance of multiple soliton knot patterns [23–25] during the complex supercontinuum generation process. Very recently the effect of strong variations of the solitons trajectories under the action of the DW was experimentally demonstrated [26]. An experimental observation of the solitonic cavities confining DW was very recently reported in [27].

This work aims to address the interaction of dark solitons with external DWs and to explore the possibility of manipulation of dark solitons by DWs. We will also demonstrate a possibility to create solitonic cavities made from dark solitons and DWs oscillating in between them.

INTERACTION OF WEAK DISPERSIVE WAVES WITH DARK SOLITONS

For a weak DW interacting with the bright soliton it is possible to construct an analytic theory predicting the location of the resonance frequencies [18–19]. Here we develop such a theoretical approach based on the perturbation theory for the case of dark solitons interacting with weak DWs.

Dynamics of the dimensionless amplitude $u(z,t)$ in normal dispersion regime with significant third-order dispersion can be described by a generalized nonlinear Schrödinger equation

$$i\partial_z u - \frac{1}{2} \partial^2_t u - \frac{i}{6} \beta_3 \partial^3_t u + |u|^2 u = 0$$

We seek for a solution in the form of a dark soliton perturbed by a sum of incident DW and scattered radiation, generated by interaction between DW and the soliton

$$u(z,t) = [\psi_{sol}(z,t) + \psi_{inc}(z,t) + \psi_{sc}(z,t)] e^{i\omega_0 z}$$

where $\psi_{sol} = u_0 \tanh(u_0 z)$. Assuming that both DW and scattered radiation are small compared to the soliton

$$\psi_{inc}(z,t) = \sum_{n=1}^{\infty} a_n e^{i\theta_n z}$$

The conditions for resonant scattering of DWs on dark solitons can be derived by considering small $\psi_{inc}(z,t)$ and $|\psi_{sol}|^2 u$ expansion.
background \( u_0 \) and retaining only the linear terms we can write the following equations for \( \psi_{inc} \) and \( \psi_{sc} \)

\[
i\partial_z \psi_{inc} - i \frac{1}{2} \beta_3^2 \psi_{inc} - i \frac{1}{6} \beta_3^2 \partial_t^2 \psi_{inc} + u_0^2 (\psi_{inc} + \psi_{inc}^+) = 0
\]

\( \text{(2)} \)

\[
i\partial_z \psi_{sc} - i \frac{1}{2} \beta_3^2 \psi_{sc} - i \frac{1}{6} \beta_3^2 \partial_t^2 \psi_{sc} + (2 \psi_{sol}^2 - u_0^2) \psi_{sc} + \psi_{sol}^* \psi_{sc} = i \frac{1}{6} \beta_3 \partial_t \psi_{sol} - 2 (\psi_{sc}^2 - u_0^2) \psi_{inc} - (\psi_{sc}^2 - u_0^2) \psi_{inc}^+
\]

\( \text{(3)} \)

Right hand’s side of equation includes three driving terms: first one generates Cherenkov radiation of the soliton (which has been already analyzed in [13]), while the other two correspond to FWM process between DW and the soliton. Seeking for a plane wave solution for (2) by setting

\[
\psi_{inc} = \psi_+ e^{ik(\omega_{inc})z - i\omega_{inc}t} + \psi_- e^{-ik(\omega_{inc})z + i\omega_{inc}t}
\]

we arrive at the dispersion relation

\[
k(\omega_{inc}) = \frac{1}{6} \beta_3 \omega_{inc}^3 + \omega_{inc} \sqrt{\frac{1}{4} \omega_{inc}^2 + u_0^2}
\]

\( \text{(4)} \)

The condition of the resonance excitation of the delocalized eigenmodes of the medium can be obtained by the analysis of the asymptotic of the eigenfunctions when where \( \psi_{sol}^2 \) approaches \( u_0^2 \). The equation then turns into a version of (2) with additional forcing terms and for the FWM resonance frequencies \( \omega_{sc} \) we can write

\[
\frac{1}{6} \beta_3 \omega_{sc}^3 \pm \omega_{sc} \sqrt{\frac{1}{4} \omega_{sc}^2 + u_0^2} = \frac{1}{6} \beta_3 \omega_{inc}^3 \pm \omega_{inc} \sqrt{\frac{1}{4} \omega_{inc}^2 + u_0^2}
\]

\( \text{(5)} \)

To confirm the analysis above we have conducted a series of numerical simulations. We solved using split-step Fourier method with the initial condition

\[
u(0, t) = u_0 \tanh(u_0 t) + A \exp(-(t-t_1)^2/\tau^2) \exp(-i\omega_{inc} t)
\]

where \( A \) is the amplitude, \( t_1 \) is the initial position of the center and \( \tau \) is the width of DW. We started with a relatively weak dispersive wave with amplitude of \( A = 0.1 \) reflecting off the soliton with \( u_0 = 1 \). A typical result of simulation is presented on Fig. 1. As it is shown on Fig. 1 (b) around the collision point \( (z \approx 35) \) the frequency of the dispersive radiation changes. The new frequency of the DW is very well predicted by the resonance condition \( \text{(5)} \), see Fig. 1 (c) showing the graphic solution of the resonance condition. To further test \( \text{(5)} \) we have varied the frequency \( \omega_{inc} \) of the incident pulse of DW and measuring the scattering frequency \( \omega_{sc} \) at the end of simulation distance. The numerically measured frequencies are shown in Fig. 2 by red squares alongside with the prediction of the resonance condition shown by the black curve. As one can see, the resonance condition predicts the frequency of the scattered DW \( \omega_{sc} \) very well.

It should be noted here that the resonance condition predicts two resonant frequencies even if the dispersion characteristics of the DW has only one branch. In our case the dispersion characteristics of the small excitation on the background of the dark soliton has two branches and the resonant scattering between the DW belonging to different branches of the the dispersion characteristics is also possible. However only one of the resonant scattering channels is efficient and in this paper we discuss only this kind of scattering.

![Figure 1: A relatively weak DW (\( A = 0.10 \)) scattering on a dark soliton. (a) is the field intensity minus the background \( |u(z, t)|^2 - u_0^2 \), (b) is the spectral density \( |u(z, \omega)| \) in the relevant region of frequency domain, (c) displays a graphical solution to \( \text{(5)} \) where the blue curve and the red line are, respectively, the left and the right hand’s sides of the equation.](image)

![Figure 2: Scattered frequency \( \omega_{sc} \) as a function of incident frequency \( \omega_{inc} \). Solid line is the theoretical prediction according to \( \text{(5)} \), markers are found by numerical modeling.](image)
INTERACTION OF STRONG DISPERSIVE WAVES WITH DARK SOLITONS AND DARK SOLITONIC CAVITIES

The interaction of an intensive DW with a soliton is a more complicated case with richer dynamics, which, unfortunately, cannot be adequately described by the analytics developed in the previous section. The main drawback of the simple linearized theory is that we have assumed that the soliton does not change while interacting with DW, which is, strictly speaking, not entirely true even for a weak DW. Collision with more intensive DWs seems to have a substantial effect on the solitons and, as a consequence, on the spectral structure of the DW itself. For example, twofold increase in the amplitude of DW \((A = 0.20)\) leads to DW pushing the soliton from its original trajectory as it is demonstrated on Fig. 3. Importantly, the steering of the trajectory in case of the dark soliton takes place in the direction opposite to that of the incident DW. This is completely different from the case of the bright solitons that is steered toward the incident DW. Looking at the process in the frequency domain (see Fig. 3b)) one can notice that scattered wave is considerably broadened spectrally. This can be explained by the fact that the soliton frequency is changing under the action of the incident DW and therefore different parts of the incident pulse scatter into different frequencies enriching the spectrum of the DW.

Further increase of the intensity of DW leads to a decomposition of the black soliton that manifest itself by a frequency continuum in the spectral domain. Fig. 4 demonstrates an outcome of the simulation in which an intensive DW with the amplitude \((A = 0.30)\) collides with the soliton causing disintegration of the soliton.

A very interesting configuration arises when we launch a relatively weak DW in a free space bound by two dark solitons. Initial condition in this case is changed to

\[
\begin{align*}
    u(0, t) &= u_{\text{cut}}(t) + A \exp(-t^2/r^2) \exp(-i\omega_{\text{inc}} t) \\
    u_{\text{cut}}(t) &= \begin{cases} 
    -u_0 \tanh(u_0(t + t_1)), & t < 0 \\
    +u_0 \tanh(u_0(t - t_1)), & t \geq 0
    \end{cases}
\end{align*}
\]

where \(t_1\) is the soliton launch position. In this case two solitons act as boundaries trapping DW inside a soliton cavity. DW periodically bounces off the walls of the cavity while changing its carrier frequency in a periodical fashion, as it is illustrated in Figs. 5, 6. Since we are dealing with multiple scatterings, soliton trajectory can be significantly perturbed even by a relatively weak DW with amplitude \(A = 0.1\) (see Fig. 6). There is a major difference from the case of two bright solitons creating a solitonic cavity [27, 29, 31]. The two dark solitons create a “convex mirror” like cavity when reflecting the DW and not “concave mirror” cavity as produced by the bright solitons.

CONCLUSIONS

The resonant scattering of weak and strong dispersive waves on dark solitons is studied analytically and numerically. An analytical formula predicting the frequency of the scattered wave was derived for the case of weak dispersive waves. Numerical simulations show that the resonant condition predicts the frequency of the scattered waves with very high accuracy. The interaction between the solitons and relatively intense DW was studied by means of pure numerical simulations. It is demonstrated that strong dispersive waves can affect dark soliton’s trajectory strongly which opens a way to control dark solitons by dispersive waves. An important observation is that the bright soliton trajectories bends against the di-
FIG. 5: A very weak DW \((A = 0.05)\) trapped in a two soliton cavity. (a) is the field intensity minus the background \(|u(z, t)|^2 - u_0^2\), (b) is the spectral density \(|u(z, \omega)|\) in the relevant region of frequency domain.

FIG. 6: A weak DW \((A = 0.1)\) trapped in a two soliton cavity. (a) is the field intensity minus the background \(|u(z, t)|^2 - u_0^2\), (b) is the spectral density \(|u(z, \omega)|\) in the relevant region of frequency domain.

Direction of the incident DW and the dark solitons trajectories bend in the direction of the incident DW. It was also demonstrated that two dark solitons allow creation of a solitonic cavity being able to trap dispersive waves making them bouncing between the solitons behaving like a couple of convex mirrors. The radiation trapped between the solitons results in the dispersive-waves mediated repulsion of the solitons at the distances much larger than the characteristic size of the solitons.

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