Interaction of charged impurities and Rydberg excitons in cuprous oxide

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We investigate the influence of a static, uncorrelated distribution of charged impurities on the spectrum of bound excitons in the copper oxide Cu$_2$O. We show that the statistical distribution of Stark shifts and ionisation rates leads to the vanishing of Rydberg resonances into an apparent continuum. The appearance of additional absorption lines due to the broken rotational symmetry, together with spatially inhomogeneous Stark shifts, leads to a modification of the observed line shapes that agree qualitatively with the changes observed in the experiment.

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I. INTRODUCTION

Semiconductor Wannier excitons are quasi-particles comprised of an electron and a hole bound by their mutual Coulomb interaction [1]. These states have first been observed in the 1950s in Cu$_2$O [2, 3], where they appear as a series of resonances below the band gap and show remarkable resemblance to the hydrogenic Rydberg series. Recently, excitons with large principal quantum numbers of up to $n = 25$, termed Rydberg excitons, have been observed in Cu$_2$O [4]. These Rydberg excitons are very sensitive to perturbations of their surroundings, just like their atomic counterparts. For example, an intensity-dependent bleaching of the resonances has been observed [4], which has been interpreted as an excitonic Rydberg blockade. Furthermore, the deviations of the spectrum of Rydberg excitons from a purely hydrogenic series can be combined into a (phenomenological) quantum defect $\delta_{n,\ell}$ that is induced by the nonparabolic hole dispersion and other central-cell corrections [5-7].

Since their first observation, the influence of electric and magnetic fields on Rydberg excitons [8-11], the mutual dipole-dipole interaction between them [12], their fluorescence [13] and inter-excitonic transitions [14] have been studied. Additionally, proposals have been put forward to use them for the implementation of masers [15, 16] as well as the realisation of topological spin phases in lattice potentials [17].

Another effect that has sparked substantial interest is the perturbation by free carriers, i.e. the electron-hole plasma [18-19]. It has been observed that the introduction of an electron-hole plasma by pumping above the band gap leads to a Mott transition for the Rydberg excitons. There, the band gap is lowered but the positions of the excitonic resonances are almost unaffected until the band gap crosses them, and the resonances vanish into the ionisation continuum [18]. The apparent suppression of the highest exciton resonances follows a similar phenomenology as the Rydberg blockade mechanism. In addition to this plasma-induced shift of the band gap, the experiments revealed a static shift that is already present without the introduction of free charge carriers.

It has long been proposed that disorder introduced, e.g., by charged impurities might lead to the appearance of an exponential decay of the absorption coefficient below the band gap, as well as a shifted band gap [20]. These charged impurities introduce a static electric field in which the exciton resonances may shift or ionise. This can result in a downward shift of the edge of the absorption continuum which can be interpreted as a reduction of the band gap. The influence of the static charged impurities can, at least for low densities of the impurities, be modeled by methods originally derived for atomic systems in ionic plasmas, the micro-field distributions. These describe the statistical distribution of local electric fields which in turn induce statistically distributed Stark shifts and ionisation broadening for the excitonic states. In the absence of a screening plasma, the suitable micro-field distribution is the one derived by Holtsmark in 1919 [21] as all assumptions (static, uncorrelated and (locally) homogeneous charge distribution) should be fulfilled. We will therefore use it in this work to assess the influence of charged impurities on the absorption spectra of the Rydberg excitons in Cu$_2$O.

The article is structured as follows: In Sec. II we describe the modeling Hamiltonian used as well as the Holtsmark micro-field distribution and discuss the central assumptions. In Sec. III we will present the numerical spectra, analyse their line parameters and compare them to experimental spectra. Finally, we will provide a discussion of our results and an outlook in Sec. IV.

II. THEORY OF STARK-SHIFTED EXCITONS

In a perfect semiconductor crystal, the unexcited crystal ground state does not contain any conduction-band electrons or valence-band holes. In the presence of impurities, however, some charge carriers may be trapped by the potentials induced by these impurities. This can prevent electron-hole pairs from recombining and leads to a persistent population of charged impurities, even at crystal temperatures near the absolute temperature minimum.
The real-space Wannier equation for an exciton perturbed by external, static charges has the form

$$\mathcal{H}_0 + \frac{e^2}{4\pi \varepsilon} \sum_i s_i \left( \frac{1}{|r_i - R_i|} - \frac{1}{|r_i - R_h|} \right) \phi(r_e, r_h)$$

(1)

$$= E \phi(r_e, r_h),$$

where $\mathcal{H}_0$ is the unperturbed excitonic Hamiltonian, $s_i = \pm 1$ is the sign of the perturbing charge and $\varepsilon = \varepsilon_0 \varepsilon_r$ is the crystal permittivity with $\varepsilon_r = 7.5$ [22]. Furthermore, $r_{e/h}$ denote the coordinates of the electron and hole that form the exciton, and the $R_i$ are the coordinates of the charged impurities. Focussing on only one of the charges, introducing center-of-mass and relative coordinates $R$ and $r$, respectively, as well as $\varrho_i = R - R_i$ gives for the interaction Hamiltonian $\mathcal{H}_i$

$$\mathcal{H}_i = \frac{e^2}{4\pi \varepsilon} s_i \left( \frac{1}{|\varrho_i - \alpha r|} - \frac{1}{|\varrho_i + \beta r|} \right)$$

(2)

where $\alpha = m_e/(m_e + m_h)$ and $\beta = m_h/(m_e + m_h)$ are the relative electron and hole masses, respectively. A Taylor expansion around $r = 0$ up to first order in $r$ yields

$$\mathcal{H}_i \approx \frac{e^2}{4\pi \varepsilon} s_i \frac{\varrho_i \cdot r}{|\varrho_i|^3}$$

(3)

and thus

$$\left[ \mathcal{H}_0 + \frac{e^2}{4\pi \varepsilon} \sum_i s_i \frac{\varrho_i \cdot r}{|\varrho_i|^3} \right] \phi(r)$$

(4)

$$= \left[ \mathcal{H}_0 + e F \cdot r \right] \phi(r) = E \phi(r)$$

where

$$F = \sum_i F_i = \frac{e}{4\pi \varepsilon} \sum_i s_i \frac{\varrho_i}{|\varrho_i|^3}$$

(5)

is the total electric field of all charged impurities. Here, the implicit assumption has been made that the length scale on which $F$ varies is large compared to the extension of excitonic states. In this case, the center-of-mass and relative coordinates can be separated if $\mathcal{H}_0$ also permits such a separation, and the truncation of the Taylor expansion after the first non-vanishing term is justified.

Under the assumption of a static, uncorrelated and homogeneous distribution of perturbing charges, the microfield distribution can be derived from Eq. (5), yielding the Holtzmark distribution [21]

$$P(\xi) d\xi = \frac{2}{\pi} \xi d\xi \int_0^\infty dx \ x^2 e^{-x^2/2} \sin(\xi x)$$

(6)

of the normalised electric field $\xi = |F|/F_0$. The normalisation factor $F_0$ corresponds closely to the field induced by a single impurity at a distance of $R_0 = \sqrt{3/(4\pi \rho_e)}$

$$F_0 = \frac{e}{2\varepsilon} \left[ 4\rho_e \right]^{2/3} R_0^{1/3} \approx \frac{e}{4\pi \varepsilon R_0^2}$$

(7)

where $\rho_e$ denotes the density of charged impurities. $R_0$ coincides roughly with the average distance of the nearest-neighbour impurity at any point. The signs of the perturbing charges $s_i$ are irrelevant, as long as the Taylor expansion in Eq. (3) is limited to the term linear in $r$. Micro-field distributions for more involved scenarios have been derived including, e.g., a screening plasma and charge-carrier correlations [23, 24].

We will focus on the simplest scenario of unscreened charges interacting with hydrogen-like excitons fulfilling the nonparabolic Wannier equation

$$\mathcal{H}_0 \phi(r) = \left[ \frac{p^2}{2\mu} + \Delta T_h(p^2) - \frac{e^2}{4\pi \varepsilon r} \right] \phi(r) = E \phi(r)$$

(8)

where $\Delta T_h(p^2)$ is the nonparabolic part of the hole dispersion which is responsible for the excitonic quantum defects. The approach to solve this equation based on reformulating it as a Sturmian Coulomb problem [25] has been outlined in Ref. [2]. The relative absorption coefficients of the Stark spectra $\kappa_0(\omega, F)$ are then derived by diagonalisation of the Wannier equation [1] in the basis of the eigenstates of $\mathcal{H}_0$.

If the Hamiltonian of the unperturbed exciton has $O(3)$ symmetry as in our model, the quantisation axis can be chosen parallel to $F$. In this case, the excitonic quantisation axes are distributed statistically, which can be taken into account by regarding the exciting light field as unpolarised with respect to the quantisation axis and leads to spectra independent of the field direction $\kappa_0(\omega, F) = \kappa_0(\omega, F) = \sum_i L_i(\omega, F)$.

The resonance line shapes are modeled by asymmetric Lorentzians [26]

$$L_i(\omega) = \frac{f_i \Gamma_i}{\pi} \frac{\Gamma_i}{\omega - \omega_i^2 + \Gamma_i^2/4}$$

(9)

where $f_i$ is the oscillator strength, $\Gamma_i$ the FWHM linewidth, $\omega_i$ the frequency and $q_i$ the asymmetry parameter of the $i$-th resonance. For isolated $P$-excitons, the asymmetry can be linked to the frequency dependence of the phononic scattering [26]. In the following, however, the $q_i$ have to be interpreted as empirical parameters used to describe and compare the shapes of the absorption lines in the experimental and numerical spectra as the apparent change in the asymmetry is induced by the superposition of multiple lines. In our model, all of these parameters except for the $q_i$ depend on the electric field $F$. The function $g(x)$ has been chosen as

$$g(x) = \begin{cases} 1 & \text{if } |x| \leq 4 \\ e^{-(|x|/4)^2} & \text{else} \end{cases}$$

(10)

The corresponding line shape resembles an asymmetric Lorentzian with constant asymmetry parameter in the vicinity of the resonance, and a symmetric Lorentzian.
far away from it. This line shape has been chosen as the use of asymmetric Lorentzians with constant asymmetry \( g(x) = 1 \) leads to a linearly decreasing absorption at the band gap due to the long range decay \( \propto -{(\omega - \omega_i)^{-1}} \) of all resonances below it. This behaviour is not observed in the experiment, where the absorption increases linearly at the band gap as predicted by Elliott [27] and reproduced by symmetric Lorentzians with their long range decay \( \propto (\omega - \omega_i)^{-2} \).

The optical transition matrix elements \( o_{n,\ell,m} \) from the crystal vacuum to the eigenstates of \( H_0 \) are proportional to [27]

\[
o_{n,\ell,m} \propto \begin{cases} 
\frac{\partial}{\partial r} R_{n\ell}(r) \big|_{r=0} & \text{if } \ell = 1 \\
0 & \text{else ,} 
\end{cases}
\]

for unpolarised light, where \( R_{n\ell}(r) \) denotes the radial part of the real-space envelope function. The relative oscillator strengths of the Stark excitons can then be calculated via

\[
f_i \propto |c_i \cdot o|^2
\]

where \( c_i \) is the \( i \)-th algebraic eigenvector of the Hamiltonian in Eq. (4) and \( o \) is the vector of the relative transition matrix elements in Eq. (11), expressed in the chosen basis.

The linewidths were calculated as laid out in Ref. [28] for the unperturbed \( P \) and \( F \) excitons (see Tab. I). They contain all relevant phononic scattering paths into the yellow 1S and 2S excitons, namely the scattering by LO phonons via the Fröhlich mechanism and the deformation potential scattering by the \( \Gamma_{3/5} \) phonons. For the \( S \) and \( D \) excitons, the experimental linewidths are not well described by this theory. We did therefore use extrapolated experimental results from SHG spectra [29] for the 3S and 3D state, respectively, giving \( \Gamma_{nS} = 2 \text{meV} \; n^{-3} \) and \( \Gamma_{nD} = 3 \text{meV} \; n^{-3} \). In addition, the complete experimental spectra seem to be broadened by \( 5 - 6 \text{meV} \). Figure 1 shows the experimental linewidths of one particular absorption spectrum. The deviation from the theoretically expected scaling \( \propto (n^2 - 1)/n^5 \) already observed in Ref. [4] could be explained by the convolution of the spectrum with a broadening Lorentzian, whose origin is unknown to us. We modeled it by adding \( 5.55 \text{meV} \) to all input linewidths.

The asymmetry of the lines was taken to be \( g_i = -0.24 \) for all lines, derived from fits to experimental spectra. Clearly, the model Hamiltonian in Eq. (3) evaluated in a basis of bound excitonic states can only be a reasonable description for Stark-excitons that are themselves bound. This problem could be addressed by complex scaling techniques [11] or the introduction of a complex absorbing potential [30]. Fortunately, the states above the classical ionisation threshold [31]

\[
E_{\text{ion}}(F) = -\sqrt{\frac{eVF}{\pi \varepsilon}}
\]

tend to be broadened by the ionisation as well as the averaging over the micro-field distribution as they are very sensitive to variations of the electric field. Our assumption is thus, that the ionised states only contribute to a continuous background to the absorption spectrum but do not account for prominent absorption lines. Once the Stark spectra are calculated, the Holtsmark spectra can be derived via

\[
\kappa(\omega, \rho_{\alpha}) = \frac{1}{F_0(\rho_{\alpha})} \int_0^\infty dF P \left( \frac{F}{F_0(\rho_{\alpha})} \right) \kappa_0(\omega, F).
\]

The integration was performed on a logarithmic grid in \( F \) with \( F_n/F_{n-1} = 1.001 \) for field strengths from \( 1 \text{mV} \; \text{m}^{-1} \) to \( 100 \text{kV} \; \text{m}^{-1} \) via the finite difference summation \( \int dF \; f(F) \approx \sum_n \Delta F_n f(F_n) \) where \( \Delta F_n = (F_{n+1} - nD = 3 \text{meV} \) taken from Ref. [28].

| \( n \) | \( P \) | \( F \) | \( n \) | \( P \) | \( F \) |
|-----|-----|-----|-----|-----|-----|
| 2   | 1896 | 3  | 6484.4 | 3  | 6484.4 |
| 4   | 2685 | 8 | 162  | 13 | 1939 | 7  | 489 |
| 6   | 8168 | 5 | 024  | 7  | 5174 | 3  | 75 |
| 8   | 3749 | 2 | 974  | 9  | 245  | 2  | 369 |
| 10  | 1789 | 1 | 89   | 11 | 1346 | 1  | 516 |
| 12  | 1038 | 1 | 226  | 13 | 817  | 1  | 001 |
| 14  | 6545 | 0 | 825  | 15 | 5325 | 0  | 687 |
| 16  | 4399 | 0 | 577  | 17 | 366  | 0  | 488 |
| 18  | 3104 | 0 | 416  | 19 | 2623 | 0  | 358 |
| 20  | 2249 | 0 | 31   | 21 | 1943 | 0  | 27 |
| 22  | 1691 | 0 | 236  | 23 | 148  | 0  | 208 |
| 24  | 1303 | 0 | 184  | 25 | 108  | 0  | 184 |

TABLE I: Input FWHM linewidths in \( \text{meV} \) taken from Ref. [28].
To summarise this section, the central assumptions of our model are:

1. The charged impurities are static and their distribution is homogeneous and uncorrelated.
2. The electric field induced by the impurities varies on length scales considerably larger than the extension of the excitonic states of interest.
3. The spectral structure is dominated by bound excitons below the classical ionisation threshold.

III. NUMERICAL RESULTS AND COMPARISON TO EXPERIMENTAL DATA

We will now apply our numerical method to the resonance spectrum of Rydberg excitons and compare with two experimental absorption spectra with maximum observable principal quantum numbers of $n_{\text{max}} \approx 25$ (hereafter $\text{S1}$) and $n_{\text{max}} \approx 13$ (hereafter $\text{S2}$). The spectrum $\text{S1}$ measured at $1.2\,\text{K}$ is the one used in the Ref. [4] and $\text{S2}$ was measured at $1.3\,\text{K}$. The quantity $n_{\text{max}}$ denotes the principal quantum number above which the resonances form an apparent absorption continuum and no individual lines can be resolved. There is, of course, some uncertainty in the definition of the highest observable principal quantum number $n_{\text{max}}(\rho_c)$. In our analysis, a resonance was considered to have vanished as soon as its spectral range could not be reliably fitted with the line shape in Eq. [9]. Figure 2 (a) compares $\text{S1}$ with a numerical spectrum derived for $\rho_c = 1.2 \times 10^9\,\text{cm}^{-3}$ which was chosen to reproduce $n_{\text{max}}$ while (b) compares $\text{S2}$ to a numerical spectrum for $\rho_c = 10^{11}\,\text{cm}^{-3}$. The numerical spectrum in Fig. 2 (b) shows additional lines corresponding predominantly to the $S$, $D$- and $F$-excitons, which become dipole allowed due to the broken rotational symmetry.

For the numerical computation, we took into account all states with $\ell \leq 25$, $n_r = n - \ell - 1 \leq 100$ as well as $m = 0, \pm 1$. The calculation can be restricted to these three magnetic quantum numbers as the Stark Hamiltonian has cylindrical symmetry which ensures that $m$ remains a good quantum number (if the quantisation axis is chosen as $z \parallel \mathbf{F}$) and the optically active $P$-states can only be mixed into other states with $m = 0, \pm 1$.

One observes that:

1. Excitons with large principal quantum numbers smear out and form an absorption continuum while the total oscillator strength is conserved.
2. The transition from negative asymmetry parameters $q_n$ for low principal quantum numbers to positive ones for the highest $n$, which have been observed in experimental spectra, is reproduced.
3. Due to the breaking of the rotational symmetries by the Stark effect, additional absorption lines – corresponding to initially dark states – appear in the numerical spectra for high impurity densities. The strongest additional lines correspond to the $S$-, $D$- and $F$-states.

Figure 2 shows the line parameters derived by fits to the numerical spectra under the assumption that the background below each line is constant over its width. The oscillator strength $f$ [Fig. 3 (a)] drops off steeply before the lines vanish starting at $n \approx 2n_{\text{max}}(\rho_c)/3$, an observation that could be explained by neither plasma nor phonon interactions [19]. Compared to the experiment, however, the oscillator strength follows the $(n^2 - 1)/n^5$ scaling for longer and drops off more steeply for large $n$. Note, that the experimental oscillator strengths in Fig. 3 (a) have been normalised to the numerical ones at $n = 5$ as their absolute values cannot be compared.

The FWHM linewidths [Fig. 3 (b)] start to deviate from the linewidths of the unperturbed resonances $n \approx n_{\text{max}}(\rho_c)/2$ and drop off shortly before $n_{\text{max}}(\rho_c)$. Addi-
tional inhomogeneous broadening could be introduced by the higher orders of the Taylor expansion, Eq. (3), which will become relevant when the electric field varies on the length scale of the exciton diameter, the crystal is inhomogeneously strained, or states below the classical ionisation threshold are ionised via tunneling processes [10]. Furthermore, there could be additional sources of microfields with different micro-field distributions, for example optical phonons [20] or surface charges. The asymmetry parameter $q$ [Fig. 3 (c)] deviates from the value for the unperturbed lines for large principal quantum numbers and changes sign for $n \approx 4n_{\text{max}}(\rho_{ci})/5$.

From the numerical spectra, we can extract the maximally observable principal quantum number as well as the shift of the band gap. Figure 4 shows the dependence of this band-gap shift on the density $\rho_{ci}$ of charged impurities. To a good approximation, it follows a power law $\Delta E_g(\rho_{ci}) = -(0.71 \pm 0.17) \text{meV} (\rho_{ci}/\text{cm}^{-3})^{0.254 \pm 0.011} \propto \rho_{ci}^{1/4}$. This scaling agrees with the dependence of the band-gap shift on the plasma density derived from many-body theory [19].

\[ \Delta E_g(\rho_{ci}) = -(0.71 \pm 0.17) \text{meV} (\rho_{ci}/\text{cm}^{-3})^{0.254 \pm 0.011} \propto \rho_{ci}^{1/4} \]

IV. DISCUSSION AND OUTLOOK

In this work, we have numerically investigated the influence of charged impurities on the spectrum of (Rydberg) excitons in Cu$_2$O. Our calculations reproduce experimentally observed phenomena such as the vanishing of the resonances with high principal quantum numbers into an apparent absorption continuum, accompanied by a drop of the oscillator strength of the discernible lines, a broadening as well as a change of the line shape towards positive asymmetry parameters $q$.

The breaking of the rotational symmetries inherent in our model leads to the redistribution of oscillator strength to initially dark states and the corresponding appearance of weak additional absorption lines in the spectra for impurity densities greater $\sim 10^{10}$ cm$^{-3}$. In the experimental spectrum S2 (Fig. 2 (b)) there are indeed some indications for such peaks, however, the signal/noise ratio of the present spectrum does not allow for a conclusive analysis. In every case, the non-appearance of such peaks may be used to establish an upper bound on the charged impurity density of given crystal samples.
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