On extreme field limits in high power laser matter interactions:
radiation dominant regimes in high intensity electromagnetic wave
interaction with electrons

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Abstract We discuss the key important regimes of electromagnetic field interaction with charged particles. Main attention is paid to the nonlinear Thomson/Compton scattering regime with the radiation friction and quantum electrodynamics effects taken into account. This process opens a channel of high efficiency electromagnetic energy conversion into hard electromagnetic radiation in the form of ultra short high power gamma ray flashes.

I. INTRODUCTION

Radiation of present day lasers approaches the intensity regimes where in the electromagnetic (EM) wave interaction with matter the radiation friction effects on the charged particle dynamics become dominant \cite{1, 2}. At these limits the electron dynamics become dissipative \cite{3, 4} with fast conversion of the EM wave energy to hard EM radiation, which for typical laser parameters is in the gamma-ray range \cite{5, 6}. For laser radiation with 1 \textmu m wavelength the radiation friction force changes the scenario of the EM wave interaction with matter at the intensity of about $I_R \approx 10^{23}$W/cm$^2$. For the laser intensity close to $I_R$ also novel physics of abundant electron-positron pair creation comes into play \cite{7, 8} (see \cite{9, 10} and \cite{11, 12}). In this regime, the electron (positron) interaction with the EM field is principally determined by a counterplay between the radiation friction and quantum effects. The quantum electrodynamics (QED) effects weaken the EM emission by the relativistic electron resulting in the lowering of the radiation friction \cite{17, 18}. In an extremely high intensity EM field vacuum looses its property to be a empty substance showing such various nonlinear QED processes as vacuum polarization, electron-positron pair plasma creation, and other properties depending on the EM field amplitude.

In Fig. 1 we present a schematic of nonlinear QED processes which realization depends not only on the EM radiation intensity, but also on the photon and charged particle energy and on the EM field configuration \cite{13, 14, 15}. Fig. 1 illustrates a transition from the relativistic interaction regime (Rel) to dominant radiation friction (RF) and nonlinear Thomson scattering (NTS) through the limit when the quantum electrodynamics comes into play (QED) and to electron-positron avalanche/cascade development (A/C), towards nonlinear vacuum with electron-positron creation ($E_S$) and nonlinear vacuum polarization (VP) while the EM field intensity grows.
FIG. 1. Schematic of nonlinear QED processes. Laser-matter interaction transition from the relativistic regime (Rel) to dominant radiation friction (RF) and nonlinear Thomson scattering (NTS) through the limit when quantum electrodynamics comes into play (QED) and to electron-positron avalanche/cascade development (A/C), towards nonlinear vacuum with electron-positron creation (E\textsubscript{S}) and nonlinear vacuum polarization (VP) while the EM field intensity increases.

The probabilities of the processes involving extremely high intensity EM field interaction with electrons, positrons and photons are determined by several dimensionless parameters.

When the normalized dimensionless EM wave amplitude

\[ a = \frac{eE\lambda}{2\pi me^2} \]

exceeds unity, \( a > 1 \), the energy of the electron quivering in the field of the wave becomes relativistic. Here \( \lambda = 2\pi c/\omega \) with \( \omega \) being the EM wave frequency. The EM wave intensity is expressed via the normalized amplitude, \( a \), as

\[ I_{rel} = \frac{m_e^2 c^3 a^2}{4\pi e^2} \]

\[ = 1.37 \times 10^{18} a^2 \left( \frac{1\mu m}{\lambda} \right)^2 \text{ W/cm}^2. \]

In a plane EM wave the parameter \( a \) is related to the Lorentz invariant, which being expressed via the 4-potential of the electromagnetic field, \( A^\mu \), is equal to \( a = e\sqrt{A^\mu A_\mu}/m_e c^2 \). Here and below a summation over repeating indices \( \mu = 0, 1, 2, 3 \) is assumed.

A relativistic electron interacting with an EM wave emits high energy photons. Here and below for the sake of simplicity we analyse the dynamics of a radiating electron in a homogeneous rotating electric field, which corresponds to the nodes of two colliding EM waves, where the wave magnetic field vanishes, and/or to the electron interaction with an EM wave in near-critical plasmas in the frame of reference moving with the group velocity of the wave [19].

In the regime of Nonlinear Thomson Scattering (NTS) the power emitted is proportional to the fourth power of its energy, \( m_e c^2 \gamma_e \),

\[ P_\gamma \approx \varepsilon_{rad} m_e c^2 \omega \gamma_e^4. \]

The dimensionless parameter,

\[ \varepsilon_{rad} = \frac{4\pi r_e}{3\lambda} = 1.17 \times 10^{-8} \left( \frac{1\mu m}{\lambda} \right), \]

proportional to the ratio of the classical electron radius \( r_e = e^2/m_e c^2 = 2.8 \times 10^{-13} \) cm and the EM wave wavelength \( \lambda \) characterizes the role of radiation losses. The maximal rate at which an electron can acquire the energy from the EM field is approximately equal to \( m_e c^2 \omega a \). The condition of the balance between the acquired and lost energy for the electron Lorentz factor equal to \( \gamma_e = a \) shows that the radiation effects become dominant at \( a > a_{rad} = \varepsilon_{rad}^{-1/3} \), i.e. at the EM wave intensity above

\[ I_R = \left( \frac{3}{2} \right)^{2/3} \frac{m_e^{8/3} c^5 \omega^{4/3}}{4\pi e^{10/3}} \]

\[ = 2.65 \times 10^{23} \left( \frac{1\mu m}{\lambda} \right)^{4/3} \text{ W/cm}^2. \]

The characteristic frequency of the emitted radiation is proportional to the cube of the electron energy, \( \omega \approx 0.3 \omega \gamma_e^3 \). QED effects become important, when the energy of the photon generated by Thomson (Compton) scattering is of the order of the electron energy, i.e. \( h\omega_m \approx m_e c^2 \gamma_e \). If \( \gamma_e = a \) this yields the QED limit on the EM field amplitude, \( a^2/a_{S} > 1 \).
Here the dimensionless parameter
\[
a_S = \frac{eE_S \lambda}{2\pi m_e c^2} = \frac{\lambda}{\hbar \omega} = 4.2 \times 10^5 \left( \frac{\lambda}{\mu \text{m}} \right) \tag{6}
\]
is the normalized critical electric field of quantum electrodynamics, \(E_S = m_e^3 e^3 / e c) [21] \), with \(\lambda_C = 2\pi\hbar / m_e c = 2.42 \times 10^{-10} \text{cm} \) being the Compton wavelength. The EM radiation intensity for the wave with the amplitude \(a_S\) is
\[
I_S = \frac{m_e^3 e^5 \omega}{4\pi^2 \hbar^2} = 2.36 \times 10^{28} \text{ W/cm}^2. \tag{7}
\]
The above obtained QED limit, \(a^2 / a_S > 1\), corresponds to the condition \(\chi_e > 1\), where relativistic and gauge invariant parameter \(\chi_e\)
\[
\chi_e = \sqrt{\frac{(F^\mu \nu k_\nu)^2}{E_S m_e c}} \approx 2 \frac{a}{a_S} \gamma e \tag{8}
\]
characterizes the probability of the gamma-photon emission by the electron with 4-momentum \(p_\nu\) in the field of the EM wave. The 4-tensor of the EM field is defined as \(F^\mu_\nu = \partial^\mu A_\nu - \partial^\nu A_\mu\). The QED limit is reached for the EM wave intensity of the order of
\[
I_Q = \frac{m_e^3 e^5 \omega}{8\pi^2 \hbar^2} = 5.75 \times 10^{23} \left( \frac{1 \mu \text{m}}{\lambda} \right) \text{ W/cm}^2. \tag{9}
\]
The dimensionless parameter
\[
\chi_\gamma = \frac{\hbar \sqrt{(F^\mu \nu k_\nu)^2}}{E_S m_e c} \approx 2 \frac{a}{a_S} \frac{\omega_\gamma}{\omega} \tag{10}
\]
determines the probability of the electron-positron pair creation by the photon with the energy \(\hbar \omega_\gamma\) in the EM field via the Breit-Wheeler process [22, 23].

Nonlinear vacuum properties are determined by the Poincare invariants,
\[
F = \frac{(B^2 - E^2)}{2} \tag{11}
\]
and
\[
G = (E \cdot B). \tag{12}
\]
\[\text{FIG. 2. Regimes of EM field interaction in the space of parameters } \frac{E}{E_S}, \chi_\gamma, \text{ and } \chi_e. \text{ The limit } E/E_S \to 1 \text{ corresponds to the nonlinear QED vacuum regimes with the electron-positron pair creation from vacuum and photon-photon interaction. When the parameter } \chi_e \text{ becomes large, multi-photon Compton scattering results in the high energy photon generation. For } \chi_\gamma > 1 \text{ the multiphoton Breit-Wheeler process results in electron-positron pair generation via the gamma-photon interaction with a strong EM field. The nonlinear Thomson scattering regime is realized for } a \gg 1 \text{ with the radiation friction effects coming in play at } a > a_{rad}. \]

The probability of the electron-positron pair creation depends on
\[
E = \sqrt{F^2 + G^2 - F} \tag{13}
\]
and
\[
B = \sqrt{F^2 + G^2 + F}, \tag{14}
\]
which are the electric and magnetic fields in the frame of reference where they are parallel.

Figure 2 illustrates different regimes of the EM interaction in the space of parameters \(E/E_S, \chi_\gamma, \text{ and } \chi_e\). In the limit \(E/E_S \to 1\) the EM waves can create electron-positron pairs from vacuum [2, 25], the EM wave can interact via photon-photon collisions [26] and vacuum polarization [27]. When the parameter \(\chi_e\) becomes large the multiphoton Compton scattering results in the high energy photon generation. For \(\chi_\gamma > 1\) the multiphoton Breit-Wheeler
process results in electron-positron pair generation via the gamma-photon interaction with the strong EM field. The nonlinear Thomson scattering (NTS) regime is realized for $a > a_{rad} = \varepsilon_{rad}^{-1/3}$ with the scattering cross section depending on the EM field amplitude. If $a > a_{rad} = \varepsilon_{rad}^{-1/3}$, radiation friction effects play a key role.

The comparsion of expressions given by Eqs. 5 and 9 shows that for the EM wave length equal to $\lambda \approx 0.8 \mu m$ the intensities $I_R$ and $I_Q$ are of the same order of magnitude as has been noted in Ref. 28. The curves $I_R(\omega)$ and $I_Q(\omega)$ intersect to each other at the frequency equal to

$$\omega = \frac{\hbar c}{m e^{2} \lambda}$$

(15)
corresponding to the wavelength of $\lambda_1 = 0.821 \mu m$ and the photon energy of the order of 1.5 $eV$.

It is convenient to rewrite expressions for $I_R(\omega)$ and $I_Q(\omega)$ in terms of the frequency $\omega_1$ as

$$I_R = \frac{m_1 c^2 e^2}{144 \pi \hbar^4} \left( \frac{\omega}{\omega_1} \right)^{4/3}$$

$$= 3.8 \times 10^{23} \left( \frac{821 nm}{\lambda} \right)^{4/3} \frac{W}{cm^2}$$

(16)

and

$$I_Q = \frac{m_1 c^2 e^2}{144 \pi \hbar^4} \left( \frac{\omega}{\omega_1} \right)$$

$$= 3.8 \times 10^{23} \left( \frac{821 nm}{\lambda} \right) \frac{W}{cm^2}.$$

(17)

In the present paper we mainly pay attention to the nonlinear Thomson/Compton scattering regime when the radiation friction and quantum electrodynamics effects play comparably important roles.

II. RADIATION FRICTION EFFECTS ON CHARGED PARTICLE MOTION

In order to describe the relativistic electron dynamics in the EM field we shall use the equations of electron motion with the radiation friction force in the Landau-Lifshitz form with a form-factor taking into account the QED weakening of the radiation friction. The EM wave is modeled by rotating electric field, which as noted above corresponds to the transformation to the boosted frame of reference moving with the group velocity of the wave. In this frame of reference the electron equations of motion can be written as

$$\frac{dq}{d\tau} = -a e \frac{\varepsilon_{rad}}{\gamma} \left( \chi_e + \frac{a}{(q \cdot a)^2} \right) \left( \frac{\hbar c}{m e^2} \right)^2$$

(18)

where $\gamma = \sqrt{1 + \frac{q^2}{c^2} + \frac{q^2}{c^2} + \frac{q^2}{c^2}}$ is the electron Lorentz factor. The form factor $G_e(\chi_e)$ with $\chi_e$ defined by Eq. (8) describes the radiation friction reduction due to quantum effects. In 3D notation the parameter $\chi_e$ given by Eq. (8) reads

$$\chi_e = \gamma_e \frac{c}{E_S} \sqrt{(E + \frac{1}{c} \mathbf{v} \times \mathbf{B})^2 - (\mathbf{v} \cdot \mathbf{E})^2}.$$

(19)

We introduce the normalized variables,

$$\tau = \omega t, \quad q = \frac{p}{mc}, \quad a = \frac{cE}{m c \omega c}.$$

(20)

The dimensionless parameter defined by Eq. (1) determines the role of the radiation friction.

As well known, QED effects weaken the radiation friction because in the quantum regime the charged particle emits less radiation. In the QED regime the recoil due to photon emission becomes important. According to Refs. 17 the total radiated intensity is reduced by a factor depending on the quantum parameter $\chi_e$.

In addition, in Eq. (18) following an approach used in Refs. 7, 9, 29, 30 we take into account the QED effects by using the form-factor, $G_e(\chi_e)$, equal to the ratio of the full radiation intensity, $I$, to the intensity emitted by a classical electron, $I_{cl}$. Using the results of Refs. 21, 31, 32 we can write the form-factor $G_e(\chi_e)$ as

$$G_e(\chi_e) =$$
\[
\int_0^\infty \frac{12 + 15\chi_e x^{3/2} + 12\chi_e^2 x^3}{4 (1 + \chi_e x^{3/2})^4} \Phi'(x) dx, \quad (21)
\]
where \(\Phi(x)\) is the Airy function \([33]\). Here, we neglect the effects of the discret nature of the photon emission in the QED (see \([28, 30, 34]\)).

In the limit \(\chi_e \ll 1\) the form-factor \(G(\chi_e)\) tends to unity as
\[
G_e(\chi_e) = 1 - \frac{55\sqrt{3}}{16} \chi_e + 48\chi_e^2 - ... \approx 1 - 5.95\chi_e + 48\chi_e^2 - ... \quad (22)
\]
For \(\chi_e \gg 1\) it tends to zero as
\[
G_e(\chi_e) = \frac{32\pi}{273^{5/6}\Gamma(1/3)}\chi_e^{4/3} - \frac{1}{\chi_e^2} + \frac{110\pi}{81\Gamma(2/3)}\chi_e^{8/3} - \frac{113^{1/2}}{5\chi_e^3} + ... \approx \frac{0.5564}{\chi_e^{4/3}} - \frac{1}{\chi_e^2} + \frac{2.6}{\chi_e^{8/3}} - \frac{3.81}{\chi_e^3} + ... \quad (23)
\]
However expression (21) and asymptotical dependences (22) and (23) are not convenient for implementing them in computer codes. For the sake of calculation simplicity the approximation
\[
G_{BKS}(\chi_e) \approx \frac{1}{[1 + 4.8(1 + \chi_e)\ln(1 + 1.7\chi_e) + 2.44\chi_e^{2/3}]/1.5, \quad (24)}
\]
is usually used \([29]\). This expression can be further simplified. The function
\[
G_R(\chi_e) \approx \frac{1}{(1 + 8.93\chi_e + 2.41\chi_e^2)^{2/3}}, \quad (25)
\]
which we shall use below, within the interval \(0 < \chi_e < 10\) has an accuracy of approximation better than 1%. In Fig. 3 we plot the functions \(G_e(\chi_e)\) and \(G_R(\chi_e)\) given by Eqs. (21) and (25), respectively. We see that that their difference is negligebly small.

We consider a rotating electric field given by
\[
a = -a \left( e_2 \cos \tau + e_3 \sin \tau \right), \quad (26)
\]
where \(e_2\) and \(e_3\) are unit vectors along the co-ordinate axis in the plane perpendicular to the direction of the EM wave propagation. We represent the electron momentum as (see also Ref. \([35]\))
\[
\begin{pmatrix}
q_1 \\
q_2 \\
q_3
\end{pmatrix}
= \begin{pmatrix}
1 & 0 & 0 \\
0 & \cos \tau & \sin \tau \\
0 & \sin \tau & -\cos \tau
\end{pmatrix}
\begin{pmatrix}
q_{1||} \\
q_{2\perp} \\
q_{3\perp}
\end{pmatrix}, \quad (27)
\]
In the case of a rotating electric field Eq. (19) yields for the QED parameter
\[
\chi_e = \frac{a^2}{a_S} \frac{1 + q_{1||}^2 + q_{2\perp}^2}{\gamma_e}, \quad (28)
\]
with \(\gamma_e = \sqrt{1 + q_1^2 + q_{2\perp}^2 + q_{3\perp}^2}\).

Substituting expressions (27) into the equations of the electron motion (18), we obtain
\[
\frac{dq_{1||}}{d\tau} + q_{2\perp} = a - \varepsilon_{rad} G_e(\chi_e)a q_{1||} q_{2\perp}^2 \gamma_e, \quad (29)
\]
\[
\frac{dq_{2\perp}}{d\tau} - q_{1||} =
\]
\[
\frac{dq_{3\perp}}{d\tau} - q_{1||} =
\]

FIG. 3. Functions \(G_e\) (lower curve) and \(G_R\) (upper curve) given by Eqs. (21) and (25), respectively, versus \(\chi_e\).
Here and below we do not consider ultrarelativistic electron beam interaction with the laser radiation. We assume that the longitudinal component of the electron momentum, \( q_1 \), is much less than the transverse momentum component \( \sqrt{q_2^2 + q_3^2} \), which corresponds to the case of the laser pulse interacting with a near-critical density plasma, when \( n \approx m_e \omega^2 / 4 \pi e^2 \) and \( q_1^{(0)} \approx 1 \), the change of the momentum component is negligibly provided that the laser pulse duration is small enough.

Multiplying equation (29) by \( q_\parallel / \gamma_e \) and equation (30) by \( q_\perp / \gamma_e \) and taking the sum we find

\[
\frac{d \gamma_e}{d \tau} = \frac{a_\parallel}{\gamma_e} - \varepsilon_{rad} G_e(\chi_e) \left( a q_\perp + a^2 q_\perp^2 \right),
\]

which shows how the electron acquires energy from the electromagnetic wave and loses it due to radiation friction.

Typical solutions of this system of equations are presented in Fig. 4, where we show dependences of the \( q_2 \) and \( q_3 \) components of the electron momentum with respect to time (l.h.s column, Fig. 4a,c), and the \( q_\parallel \) and \( q_\perp \) components of the electron momentum (central column, Fig. 4b,d). As we see, for \( a \ll \varepsilon_{rad}^{-1/3} \) the electron oscillations in the \((q_\parallel,q_\perp)\) plane decay slowly while for the laser pulse amplitude values equal to or above \( \varepsilon_{rad}^{-1/3} \), the electron oscillations in the rotating coordinate system decay during a time of the order of or less than the wave period. The dependence on time of the QED parameters \( \chi_e(t) \) and \( G(t) \) is plotted in Fig. 4e,f. We see that for \( a = 0.25 \varepsilon_{rad}^{-1/3} \), i.e. at the laser intensity of the order of \( 10^{21} \) W/cm\(^2\) for 1\(\mu\)m wavelength laser radiation the quantum reducing of the radiation friction force is negligibly small. Asymptotically at \( t \to \infty \) we have \( \chi_e \approx 0.1 \) and \( G_{e^{1/3}} \approx 0.95 \). When \( a = 0.75 \varepsilon_{rad}^{-1/3} \), the quantum correction of the the radiation friction force becomes more significant with \( \chi_e \approx 0.3 \) and \( G_{e^{1/3}} \approx 0.75 \) in the limit \( t \to \infty \).

Using the above obtained results we can find border lines between the domain where the dominant radiation friction regime takes place (this is the domain II in Fig. 3) and the domain IV in Fig. 3 where QED effects must be taken into account. They are given by different dependencies of the EM radiation intensity on the wave frequency. As was shown in Ref. 35, in the limit \( a > a_{rad} \) the transverse component of the electron momentum \( q_\perp \approx (a_\parallel \varepsilon_{rad})^{1/4} \) is substantially less than the component parallel to the electric field \( q_\parallel = (a \varepsilon_{rad})^{1/4} \). Quantum effects become important, when the value of the QED parameter \( \chi_e \), which becomes equal to \( \chi_e \approx q_\parallel / a_s \), approaches unity. This yields in the limit \( \omega \ll \omega_1 \) for the EM wave intensity

\[
I_{R-Q} = \frac{m_e^4 c^2 e^4}{9 \pi \hbar^4} = 5.6 \times 10^{21} \frac{W}{cm^2}.
\]

The border line between domains III and IV for \( \omega \gg \omega_1 \) in Fig. 3 corresponds to the limit \( \chi_e \gg 1 \), when the asymptotic expression for the form factor \( G_e(\chi_2) \) is given by Eq. (23). From the equations of the electron motion (29) and (30) we obtain, that the electron normalized energy depends on the EM field amplitude as \( \gamma_e \approx (a_\parallel / \varepsilon_{rad} a_S^{1/3})^{1/8} \). This yields for the EM wave amplitude at the border line \( a_{Q-R} \approx 96 a_{3}^{4} \varepsilon_{rad}^{-3} \approx 324 (\lambda_c^4 / a_s^{3/4}) \). Using this expression for \( a_{Q-R} \) we obtain a dependence of \( I_{Q-R}(\omega) \) in the form

\[
I_{Q-R} \approx 8 \sqrt{c^5 \hbar^8 \omega^4 \varepsilon_{rad}^{-3}} = 8 \times 10^{21} \left( \frac{\omega}{\omega_1} \right)^4 \frac{W}{cm^2}.
\]

In Fig. 5 we plot the border curves \( I_R, I_{R-Q}, I_Q \) and \( I_{Q-R} \) versus the EM wave frequency. These curves subdivide the plane \( I, \omega \) to four domains. In the first domain, (I) neither radiation friction nor QED effects are important for the relativistic interaction of an electron with the EM field. In the second domain, (II) the electron – EM wave interaction is dominated by the radiation friction force effects. In the third domain, (III) the QED effects come into play with insignificant radiation friction force effects. In the high intensity limit, (IV) both
the QED and radiation friction force effects determine the radiating charged particle dynamics in the EM wave. The dashed lines show the dependences \( I_R(\omega) \) at \( 0 < \omega < \omega_1 \) and \( I_Q(\omega) \) at \( \omega > \omega_1 \).

The quantum nature of the photon emission process results in the recoil effect which causes the electron trajectory broadening (see review article [36] and literature cited in). Following this paper we can write equation describing dependence on time of the mean square of the electron trajectory deflection

\[
\frac{d\Delta r^2}{dt} = \frac{55}{48\sqrt{3}} \frac{r_e c \lambda C}{\lambda} \gamma_e^5.
\] (34)

A condition of the electron deflection during the laser wave period \( 2\pi/\omega \) being of the order of the laser wavelength, \( \Delta r \approx \lambda \), shows that the quantum diffusion becomes important at the electron quiver energy corresponding to

\[
\gamma_D = \left( \frac{\sqrt{3}}{55\pi} \frac{m_e^2 c^5}{\hbar \omega^2 \hbar e^2} \right)^{1/5} = \frac{2.4}{\alpha^2} \left( \frac{\omega_1}{\omega} \right)^{2/5},
\] (35)
i.e. above approximately 25 GeV. Corresponding intensity is

\[
I_D = 7.2 \times 10^{26} \left( \frac{\omega_1}{\omega} \right)^{4/5} \text{ W cm}^2. \quad (36)
\]

III. INTEGRAL SCATTERING CROSS SECTION

According to Eq. [31], the energy flux reemitted by the electron is equal to

\[
e(\mathbf{v} \cdot \mathbf{E}) \approx \varepsilon_{\text{rad}} G_e(\chi_e) m_e^2 c^2 \omega_e (a q_\perp + a^2 q_\perp^2).
\] (37)
For $\chi_e \ll 1$, i.e.

$$1 \ll \varepsilon_{rad}^{-1/3} \ll a \ll \varepsilon_{rad} a^2,$$  \hspace{1cm} (41)

it reaches a maximum of $\sigma \approx 0.53 \sigma_T \varepsilon_{rad}^{-2/3}$ at $a = 1.1 \varepsilon_{rad}^{-1/3}$. These dependences correspond to the domains (I) and (III) in Fig. 6, where the radiation friction effects are weak. Then for $a \gg \varepsilon_{rad}^{-1/3}$ the scattering cross section decreases according to

$$\sigma \approx \sigma_T / a \varepsilon_{rad},$$  \hspace{1cm} (42)

as seen in Fig. 6 as a consequence of the fact that the maximal power reemitted by the electron cannot exceed $eE/\gamma$. This gives a constraint on the scattering cross section: $4\pi e^2/E$. In the limit $\varepsilon_{rad}^{-1/3} \ll a \ll \varepsilon_{rad} a^2$, the electron momentum components are $q_{\perp} \approx (a/\varepsilon_{rad})^{-1}$ and $q_{\parallel} \approx (a/\varepsilon_{rad})^{1/4}$ with $q_{\perp} \ll q_{\parallel}$ \cite{55}, which corresponds to domain (II) in Fig. 6, where the radiation friction is dominant.

Dependences of the scattering cross section normalized by $\sigma_T$ and the electron energy $\gamma_e$ on the laser pulse amplitude $a$ are shown in Fig. 6. The parameters $\varepsilon_{rad} = 1.75 \times 10^{-8}$ and $a_S = 2.8 \times 10^{-3}$ (curves 2) correspond to the EM frequency equal to $\omega_1$ given by Eq. (15).

For $\omega = \omega_1 / 12.5$ (curves 1) the EM wave length is about $\lambda = 10 \mu m$, which is typical for the CO$_2$ laser wavelength range. The case 3 with $\omega = 12.5 \omega_1$ corresponds to the parameters in the domain (IV) in Fig. 6 with the radiation friction lowered by quantum effects.

**IV. COMPUTER SIMULATIONS OF THE RADIATION FRICTION AND QED EFFECTS ON THE LASER PULSE INTERACTION WITH INHOMOGENEOUS PLASMA**

The electron quivering in the laser field emits photons whose energy is proportional to the cube of the electron Lorentz factor: $\hbar \omega_\gamma \approx 0.3 \hbar \omega_\gamma 3$. For $a \ll \varepsilon_{rad}^{-1/3}$ a typical value of the photon frequency is proportional to $a^3$. In the
FIG. 7. The concept of high power gamma-ray flash generation in the laser-matter interaction.

If the limit of high laser intensity $a \gg \varepsilon_{\text{rad}}^{-1/3}$, the frequency scales as $\omega_2 = \omega (a/\varepsilon_{\text{rad}})^{3/4}$. For multi-petawatt laser radiation the emitted photon energy is in the gamma ray energy range.

The gamma-ray pulse energy, duration and divergence are determined by the laser pulse amplitude and by the plasma target density scale length. We analyse the density scale length effect on the parameters of the emitted gamma-flashes. Fig. 7 illustrates the concept of the high power gamma-ray flash generation in the laser-matter interaction.

During interaction of super-high-power laser light with matter the laser pulse is subject to various instabilities. Among them the most important is the relativistic self-focusing resulting in the laser pulse modulation and channeling. It leads to the increase of the laser pulse amplitude and to the decrease of the electron density in the interaction region, which change the laser energy depletion length and the parameters of the gamma-rays emitted. Thorough studying of these effects and of the effects of the plasma inhomogeneity require computer simulations.

We performed studies of the laser pulse interaction with high density targets using a two-dimensional (2D) particle-in-cell (PIC) code where the radiation friction force has been incorporated in the Landau-Lifshitz form as has also been done in Refs. [37]. In addition, the QED radiation friction weakening is taken into account by multiplying the radiation force by a form-factor $G_e(\chi_e)$ given by Eq. (25). We note that results of detailed computer simulations of the high power gamma ray flash emission by a multi-petawatt pulse laser have been presented in Ref. [8], where the plasma target inhomogeneity has been suggested for optimization of the EM wave energy conversion to the gamma ray energy flash.

In the simulations, the laser pulse has the power $P_{\text{las}}$ equal to 100PW with the pulse duration of $T_{\text{las}} = 30$ fs.

In the tailored plasma target the density changes exponentially, $n \propto \exp(x/L)$, with the plasma inhomogeneity scale length, $L$, from 0.1 ncr to 350 ncr in the interval $\approx 20 \mu m$, and then becomes constant having a thickness of 10 $\mu m$.

The simulation box has a width equal to 80 $\mu m$ and a length varying from 50 $\mu m$ to 210 $\mu m$. The mesh has a spatial resolution of $\Delta x = \Delta y$ varying from $1/40 \mu m$ to 1/200 $\mu m$ with a temporal resolution of $\Delta t = 0.0025$ fs. The plasma is comprised of electrons and ions with a mass to charge ratio corresponding to $A/Z = 2$, corresponding to fully ionized Carbon. The average number of quasiparticles is
Simulation results for the parameters of interest are shown in Figs. 7 and 8. The 100 PW laser pulse with the normalized amplitude $a = 474$ interacts with the plasma target, whose density inhomogeneity is characterized by a scale length equal to $L = 2.5 \mu m$.

Fig. 8a) shows the dependence on time of the electron $E_e$, ion $E_i$ and total energy $E_{\text{tot}}$ when the radiation friction force and the QED effects are neglected. We see that the laser pulse energy at first is converted to the electron energy and then to the ion energy. In Fig. 8b) we plot the same dependences of the electron, ion and total energy and the energy of emitted gammarays in time for the case with the radiation friction force taken into account but without QED effects. In this case a major part of the laser radiation becomes lower as it is seen in Figs. 7c) and e). The maximum gamma ray pulse power is equal to 10 PW. If the radiation friction force and the QED effects are incorporated into the model and the QED effects are neglected (9b) and e)), and when the radiation friction force and QED effects are neglected (9a), where the electron density distribution in the $(x,y)$ plane is shown at $t = 140$fs and d) with the laser electric field distribution presented for the same time), the conversion of the laser pulse energy to the gamma-ray flash, whose density inhomogeneity is characterized by a scale length equal to $L = 2.5 \mu m$.

The maximal gamma ray pulse power is equal to 45 PW. If the radiation friction force and the QED effects are taken into account (see 9a), where the electron density distribution in the $(x,y)$ plane is shown at $t = 140$fs and d) with the laser electric field distribution presented for the same time), the conversion of the laser pulse energy to the gamma-ray flash, whose density inhomogeneity is characterized by a scale length equal to $L = 2.5 \mu m$.

V. SPECTRUM OF THE RADIATION EMITTED BY AN ENSEMBLE OF ULTRARELATIVISTIC ELECTRONS

The frequency spectrum of a relativistic electron rotating along the trajectory with the radius of curvature, $R$, is given by the expression,

$$\frac{dI_{\gamma}(\omega, \mathcal{E})}{d\omega} = \frac{\sqrt{3} e^2}{2\pi c m_e c^2} \int_{u(\omega, \mathcal{E})}^{\infty} K_{5/3}(x) dx. \quad (43)$$

which is well known in the classical electrodynamics [20]. Here $K_{\nu}(z)$ is the modified Bessel function [22]. The electron energy is assumed to be $\mathcal{E} \gg m_c^2$, $m_e$ and $c$ are the electron mass and the speed of light in vacuum, respectively. The function of the electron energy, $u(\omega, \mathcal{E}) = \omega/\omega_{c}$, is the ratio of the ratio of the emitted photon frequency $\omega_{c}$ to the critical frequency equal to

$$\omega_{c} = \frac{3c}{2R} \left( \frac{\mathcal{E}}{m_e c^2} \right)^3. \quad (44)$$

In the limit $u(\omega, \mathcal{E}) \ll 1$ expression (43) yields $dI/d\omega_{\gamma} \approx (e^2/c)(\omega_{c} R/c)^{1/3}$, while for $u(\omega, \mathcal{E}) \gg 1$ we have

$$\frac{dI}{d\omega_{\gamma}} \approx \sqrt{\frac{3\pi}{2}} \frac{e^2}{c m_e c^2} u(\omega_{\gamma}, \mathcal{E})^2 e^{-u(\omega_{\gamma}, \mathcal{E})}. \quad (45)$$
FIG. 8. Computer simulation of a 100 PW laser pulse interaction with a tailored target. a) Time evolution of the electron $E_e$, ion $E_i$ and total energy $E_{\text{tot}}$ with no radiation friction force and QED effects. b) $E_e$, ion $E_i$, gamma-ray $E_{\gamma}$ and total energy $E_{\text{tot}}$ with the radiation friction force but without QED effects. c) $E_e$, ion $E_i$, gamma-ray $E_{\gamma}$ and total energy $E_{\text{tot}}$ with the radiation friction force and with the QED effects. d) The gamma ray pulse power $P_{\gamma}$ and energy $E_{\gamma}$ vs time with the radiation friction force but without QED effects. e) $P_{\gamma}$ and $E_{\gamma}$ vs time with the radiation friction force and QED effects. Time is measured in fs. Units of the energy and power are J and PW, respectively.

FIG. 9. Computer simulation of the hole boring by a 100 PW laser in the tailored target. a,b,c) The electron density distribution and d,e,f) and the laser electric field in the $(x,y)$ plane at $t = 140$fs. For a) and d) no radiation friction force and QED effects are taken into account. b) and e) show the hole boring, when the radiation friction force is present but without the QED effects. c) and f) correspond to the case of the nonvanishing radiation friction force and QED effects.
FIG. 10. Comparison of the conversion efficiency of the laser energy to the energy of the gamma-rays for a 30 fs laser pulse interacting with the tailored target when the quantum correction of the radiation friction is taken into account (dots) and when it is neglected (solid curve).

Electrons rotating along the circles with the same radius emit electromagnetic radiation with the same frequency spectrum independently of the particular radiation mechanism.

In the case of synchrotron radiation when the ultrarelativistic electron with the energy $E \gg m_e c^2$ rotates in a homogeneous magnetic field the Larmor radius in the limit $E \gg m_e c^2$ is equal to $r_B = E/eB$. The characteristic photon frequency of the synchrotron radiation according to Eq. (43) is given by $\omega_B = eB/m_e c$ is the Larmor frequency.

The characteristic energy of a photon emitted via nonlinear Thomson scattering of a circularly polarized electromagnetic wave scales with the electron quiver energy as $\omega_c = (3/2)\omega_0(E/m_e c^2)^3$, where $\omega_0 = eB/m_e c$ is the laser frequency. The energy of the electron quivering in plasma under the action of an electromagnetic wave with an amplitude of $a = eE/m_\gamma c \gg 1$ is of the order of $E = am_\gamma c^2$. For a laser frequency of the order of $10^{15}$ s$^{-1}$ the emitted photon energy is in the gamma-ray range if $a > 10^2$ which corresponds to a laser intensity higher than $10^{22}$ W/cm$^2$.

Computer simulations of a petawatt power short laser pulse interaction with a near-critical plasma [8] show that the electron energy spectra in the interval $E < E_m$ can be approximated by a power-law dependence. For the energy $E < E_m$ the spectrum has exponential form. Under the conditions of the simulations presented in Ref. [8] the energy, $E_m$, is of the order of 50 MeV. The electron energy distribution can be described by the dependence

$$\frac{dN(E)}{dE} = K E^{-\kappa} \exp\left(-\frac{E}{E_m}\right). \quad (46)$$

For the parameters corresponding to the results presented in [8] the power index equals $\kappa \approx 0.8$.

The averaged spectrum of the emitted photons is given by the integral

$$J(\omega_\gamma) = \int_0^\infty \frac{dI_\gamma(\omega_\gamma, E) \, dN(E)}{d\omega_\gamma \, dE} \, d\omega_\gamma, \quad (47)$$

where the functions $dI_\gamma/d\omega_\gamma$ and $dN(E)/dE$ are determined by Eqs. (43) and (46).

We analyze the spectrum of the radiation emitted in the process of nonlinear Thomson scattering of a circularly polarized electromagnetic wave, when the function $u(\omega_\gamma, E)$ is

$$u_C(\omega_\gamma, E) = \frac{2}{3} \omega_\gamma \left(\frac{m_e c^2}{E}\right)^3. \quad (48)$$

Substituting the functions (43) and (46) to (47) and performing integration over $E$ we obtain the averaged spectrum of the emitted photons

$$J_C(\omega_\gamma) = Q_C \omega_\gamma^{2/3} \, F_C(\epsilon_m, \kappa), \quad (49)$$

where

$$Q_C = \frac{\sqrt{3} K e^2}{2\pi c} \left(\frac{2}{3\omega_0}\right)^{2/3} (m_e c^2)^{1-\kappa}, \quad (50)$$

$$\epsilon_m = \frac{E_m}{m_e c^2} \left(\frac{3\omega_0}{2\omega_\gamma}\right)^{1/2}, \quad (51)$$

and

$$F_C(\epsilon_m, \kappa) =$$
\[ \epsilon_m^{-(1+\kappa)} \int_0^\infty K_{5/3}(x) \Gamma \left( -1 - \kappa, \frac{1}{x^\kappa \epsilon_m} \right) dx. \] (52)

Here \( \Gamma(a, z) \) is the incomplete gamma function \[33\]. The function \( F_C(\epsilon_m, \kappa) \) can be expressed in terms of hypergeometric functions.

In the low frequency range, \( \omega_\gamma \ll (3/2)\omega_0(\mathcal{E}_m/m_e c^2)^3 \), using the asymptotic representation of the incomplete gamma function at \( x^{-\frac{1}{2}} \epsilon_m^{-1} \to 0 \),

\[ \Gamma \left( -1 - \kappa, \frac{1}{x^\kappa \epsilon_m} \right) = \Gamma(-1 - \kappa) + \epsilon_m^{-(1+\kappa)} \frac{x^{-\frac{1}{2}}}{1+\kappa} - \frac{\epsilon_m^{1-\frac{3}{2}}}{\kappa} + \ldots, \] (53)

we obtain

\[ F_C(\infty, \kappa) = \frac{2^{\frac{\kappa-2}{3}}}{1+\kappa} \Gamma \left( \frac{\kappa+9}{6} \right) \Gamma \left( \frac{\kappa+1}{6} \right), \] (54)

where \( \Gamma(x) \) is the gamma function \[33\]. This expression is valid provided \( \kappa > 1 \). We see that the spectrum has the power form,

\[ J_C(\omega_\gamma) \sim \omega_\gamma^{-\frac{2\kappa}{3}}. \] (55)

If the power index is greater than 2, \( \kappa > 2 \), the dependence \( J_C(\omega_\gamma) \) is monotonically decreasing as \( \omega_\gamma/\omega_0 \to \infty \). For \( 1 < \kappa < 2 \) the function \( J_C(\omega_\gamma) \) grows with \( \omega_\gamma \). In this case as well, if the power index is less than unity, \( \kappa < 1 \), we should take into account that the electron energy spectrum is truncated at the high energy end.

Analyzing the asymptotic behaviour of the function \( F_C(\epsilon_m, \kappa) \) given by Eq. (52) as \( \epsilon_m \to \infty \) for \( \kappa < 1 \) we find the low frequency scaling \( \omega_\gamma \ll (3/2)\omega_0(\mathcal{E}_m/m_e c^2)^3, \kappa < 1 \)

\[ J_C(\omega_\gamma) \approx \sqrt{3\epsilon_c^2} \left( \frac{\mathcal{E}_m}{m_e c^2} \right)^{(1-\kappa)} \Gamma \left( \frac{2}{3} \right) \Gamma(1 - \kappa) \left( \frac{\omega_\gamma}{3\omega_0} \right)^{1/3}. \] (56)

Considering the limit \( \epsilon_m \to 0 \), which corresponds to the high frequency range, \( \omega_\gamma \gg (3/2)\omega_0(\mathcal{E}_m/m_e c^2)^3 \), we obtain

\[ F_C(\epsilon_m, \kappa) \approx \sqrt{\frac{\pi}{2}} \epsilon_m^{-\kappa} \int_0^\infty dx \frac{1}{x^{1+\kappa}} \exp \left( -x - \frac{1}{x^{1/3} \epsilon_m} \right) \] \( \propto \exp \left( -\frac{4}{3^{1/3} \epsilon_m^{3/4}} \right) \), (57)

i.e., the spectrum decreases exponentially with photon frequency,

\[ J_C(\omega_\gamma) \sim \exp \left[ -4 \left( \frac{2\omega_\gamma m_e^3 c^2}{51\omega_0^3 \mathcal{E}_m^3} \right)^{1/4} \right]. \] (58)

As a result we see that the frequency spectrum has a maximum at

\[ \omega_{\gamma,m} \approx 14.2 \times \omega_0 \left( \frac{\mathcal{E}_m}{m_e c^2} \right)^3. \] (59)

For the parameters of the simulations presented in \[8\] the characteristic energy \( \mathcal{E}_m \) is of the order of 38 MeV. This yields for the photon energy, \( \varepsilon_{\gamma,m} = h\omega_{\gamma,m} \), corresponding to the spectrum maximum \( \varepsilon_{\gamma,m} \approx 6 \) MeV.

**VI. CONCLUSION**

With Kilo-Joule lasers high field science will enter novel regimes of electromagnetic radiation interaction with matter when radiation friction force effects result in the high efficiency conversion of the energy of the EM wave into the energy of hard EM radiation in the form of ultra short high power gamma ray flashes. The energy spectrum of the gamma-ray flash emitted by the relativistic electrons has a typical form with a maximum which dependence on the parameters is given by Eq. (59).
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[1] G. A. Mourou, T. Tajima, S. V. Bulanov, Rev. Mod. Phys. 78, 309 (2006).
[2] M. Marklund and P. Shukla, Rev. Mod. Phys. 78, 591 (2006); A. Di Piazza, C. Muller, K. Z. Hatsagortsyan, C. H. Keitel, Rev. Mod. Phys. 84, 1177 (2012).
[3] Ya. B. Zel’dovich, Sov. Phys. Usp. 18, 97 (1975).
[4] A. Zhidkov, J. Koga, A. Sasaki, M. Uesaka, Phys. Rev. Lett. 88, 185002 (2002).
[5] J. K. Koga, T. Zh. Esirkepov, S. V. Bulanov, Phys. Plasmas 12, 093106 (2005).
[6] A. Thomas, C. Ridgers, S. S. Bulanov, B. J. Griffin, S. Mangles, Phys. Rev. 2, 041004 (2012).
[7] C. P. Ridgers, C. S. Brady, R. Duclous, J. G. Kirk, K. Bennett, T. D. Arber, A. P. L. Robinson, A. R. Bell, Phys. Rev. Lett. 108, 165006 (2012).
[8] T. Nakamura, J. K. Koga, T. Zh. Esirkepov, M. Kando, G. Korn, S. V. Bulanov, Phys. Rev. Lett. 108, 195001 (2012).
[9] A. R. Bell and J. G. Kirk, Phys. Rev. Lett. 101, 200403 (2008).
[10] J. G. Kirk, A. R. Bell, I. Arka, Plasma Phys. Contr. Fusion 51, 085008 (2009).
[11] A. M. Fedotov, N. B. Narozhny, G. Mourou, G. Korn, Phys. Rev. Lett. 105, 080402 (2010).
[12] E. N. Nerush, I. Yu. Kostyukov, M. V. Legkov, A. M. Fedotov, N. B. Narozhny, N. V. Elkina, H. Ruhl, Phys. Rev. Lett. 106, 035001 (2011).
[13] S. S. Bulanov, T. Esirkepov, J. Koga, A. Thomas, S. V. Bulanov, Phys. Rev. Lett. 105, 220407 (2010).
[14] S. S. Bulanov, V. D. Mur, N. B. Narozhny, J. Nee, V. S. Popov, Phys. Rev. Lett. 104, 220404 (2010).
[15] R. Schtzhold, H. Gies, G. Dunne, Phys. Rev. Lett. 101, 130404 (2008).
[16] S. V. Bulanov, T. Zh. Esirkepov, Y. Hayashi, M. Kando, H. Kiriyama, J. K. Koga, K. Kondo, H. Kotaki, A. S. Pirozhkov, S. S. Bulanov, A. G. Zhidkov, P. Chen, D. Neely, Y. Kato, N. B. Narozhny, G. Korn, Nuclear Instr. Meth. Phys. Res. A 660, 31 (2011).
[17] J. Schwinger, Proc. Natl. Acad. Sci. U.S.A. 40, 132 (1954).
[18] A. A. Sokolov, N. P. Klepikov, I. M. Ternov, Sov. Phys. JETP 24, 249 (1964).
[19] A. C.-L. Chian, Phys. Rev. A 24, 2773 (1981); S. S. Bulanov, Phys. Rev. E 69, 036408 (2004).
[20] A. B. Berestetskii, E. M. Lifshitz, The Classical Theory of Fields (Pergamon, Oxford, 1975).
[21] V. B. Berestetskii, E. M. Lifshitz, and L. P. Pitaevskii, Quantum Electrodynamics (Pergamon, Oxford, 1982).
[22] G. Breit and J. A. Wheeler, Phys. Rev. 46, 1087 (1934).
[23] A. I. Nikishov and V. I. Ritus, Sov. Phys. Usp. 13, 303 (1970); T. Heinzl, A. Ilderton, M. Marklund, Phys. Rev. D 81, 051902(R) (2010).
[24] V. S. Popov, Phys. Usp. 47, 855 (2004).
[25] R. Ruffini, G. Vereshchagin, S. S. Xue, Physics Reports 487, 1 (2010).
[26] J. K. Koga, S. V. Bulanov, T. Esirkepov, A. Pirozhkov, M. Kando, Phys. Rev. A 86, 053823 (2012).
[27] N. Rosanov, JETP Lett. 88, 501 (2008).
[28] C. S. Brady, C. P. Ridgers, T. D. Arber, A. R. Bell, J. G. Kirk, Phys. Rev. Lett. 109, 245006 (2012).
[29] V. Baier, V. Katkov, V. Strakhovenko, Electromagnetic Processes at High Energies in Oriented Single Crystals (World Scientific, 1998).
[30] S. S. Bulanov, E. Esarey, C. B. Schroeder, and W. P. Leemans, in preparation.
[31] V. I. Ritus, in Issues in Intense-Field Quantum Electrodynamics (Nova Science, Commack, 1987).
[32] I. V. Sokolov, J. A. Nee, V. P. Yanovsky, N. M. Naumova, G. A. Mourou, Phys. Rev. E 81, 036412 (2010).
[33] M. Abramowitz and I. A. Stegun, Handbook of Mathematical Functions with Formulas, Graphs, and Mathematical Tables (Dover, New York, 1965).
[34] R. Duclous, J. G. Kirk, A. R. Bell, Plasma Phys. Contr. Fusion 53, 015009 (2011).
[35] S. V. Bulanov, T. Esirkepov, M. Kando, J. Koga, S. S. Bulanov, Phys. Rev. E 84, 056605 (2011).

[36] I. M. Ternov, Phys. Uspekhi 165, 429 (1995).

[37] N. Naumova, T. Schlegel, V. T. Tikhonchuk, C. Labaune, I. V. Sokolov, G. Mourou, Phys. Rev. Lett. 102, 025002 (2009); M. Tamburini, F. Pegoraro, A. Di Piazza, C. H. Keitel, T. V. Liseykina, A. Macchi, New J. Phys. 12, 123005 (2010).