Effective Low–Energy Potential for Slow Dirac Fermions in Einstein–Cartan Gravity with Torsion and Chameleon∗

A. N. Ivanov† and M. Wellenzohn† ⊕

1Atominstitut, Technische Universität Wien, Stadionallee 2, A-1020 Wien, Austria
2FH Campus Wien, University of Applied Sciences, Favoritenstraße 226, 1100 Wien, Austria

(Dated: November 30, 2015)

We derive the most general effective low–energy potential to order $O(1/m)$ for slow Dirac fermions with mass $m$, coupled to gravitational, chameleon and torsion fields in the Einstein–Cartan gravity. The obtained results can be applied to the experimental analysis of gravitational, chameleon and torsion interactions in terrestrial laboratories. We discuss the use of rotating coordinate systems, caused by rotations of devices, for measurements of the torsion vector and tensor components, caused by minimal torsion–fermion couplings (Ivanov and Wellenzohn, Phys. Rev. D 92, 065006 (2015)). Using the most general form of a metric tensor of curved spacetimes in rotating coordinate systems, proposed by Obukhov, Silenko, and Teryaev (Phys. Rev. D 84, 024025 (2011)), we extend this metric by the inclusion of the chameleon field and calculate the set of vierbein fields, in terms of which Dirac fermions couple to torsion vector and tensor components through minimal torsion–fermion couplings. For such a set of vierbein fields we discuss a part of the effective low–energy potential for slow Dirac fermions, coupled to gravitational, chameleon and torsion fields to order $O(1)$ in the large fermion mass expansion.

PACS numbers: 03.65.Pm, 04.25.-g, 04.25.Nx, 14.80.Va

I. INTRODUCTION

In terrestrial laboratories [1]–[7] gravitational and chameleon interactions are being investigated in terms of cold and ultracold neutrons through some effective low–energy potentials [8]–[11]. The low–energy torsion–fermion interactions of the pseudoscalar and axial–vector components of torsion field have been derived and estimated by Lämmerzahl [12] and Obukhov, Silenko, and Teryaev [13]. The most general torsion–fermion interactions of constant torsion fields have been proposed and estimated by Kostelecky, Russell, and Tasson [14]. The results, obtained by Lämmerzahl [12], Obukhov et al. [13] and Kostelecky et al. [14], have been discussed in [15]. An attempt of a direct measurement of torsion–fermion interactions with constant torsion fields, proposed by Kostelecky et al. [14], has been undertaken by Lehnert, Snow and Yan [16].

In this paper we derive the most general effective low–energy potential to order $1/m$ for slow Dirac fermions with mass $m$, coupled to gravitational, chameleon and torsion fields in the Einstein–Cartan gravity.

The chameleon part of such an effective low–energy potential contains new chameleon–fermion interactions with respect to those calculated in [10, 15]. These new chameleon–fermion interactions can be used for more detailed experimental analysis of the properties of the chameleon field [17, 18]. The chameleon field, changing its mass in dependence of a mass density of environment, has been invented to avoid the problem of violation of the equivalence principle [19]. In addition the chameleon field can be also identified with a quintessence (canonical scalar field) [20, 21], which has been postulated for an explanation of the late–time acceleration of the Universe expansion [22, 23]. The laboratory probes of the chameleon field, coupled to a matter in conformal way [17, 18], may also shed light on dark energy dynamics [25, 26].

Torsion is an additional to a metric tensor natural geometrical quantity characterizing spacetime geometry through spin–matter interactions [24, 25]. It allows to probe rotational degrees of freedom of spacetime in terrestrial laboratories. Torsion is described by a third–order tensor $T_{\sigma \mu \nu}$, antisymmetric with respect to indices $\mu$ and $\nu$, i.e. $T_{\sigma \mu \nu} = -T_{\sigma \nu \mu}$. It possesses 24 independent components, which can be decomposed into 4 axial–vector $B_\mu$, 4 vector $E_\mu$ and 16 tensor $M_{\sigma \mu \nu}$ components [14] (see also Eq. (5) and Eq. (4)). The torsion part of the effective low–energy potential, derived in our paper, is caused by torsion–fermion minimal couplings for all torsion components only. An importance of this part of the effective low–energy potential is related to a possible solution of the following

∗ We dedicate this paper to 200 Jubilee of Vienna University of Technology
†Electronic address: ivanov@kph.tuwien.ac.at
⊕Electronic address: max.wellenzohn@gmail.com
problem. As has been shown in \[15,\] only torsion axial–vector $B_\mu$ components are present in the torsion–fermion minimal couplings in the curved spacetimes with metrics, providing vanishing time–space (space–time) components of the vierbein fields. Since these are usual metrics of spacetimes in terrestrial laboratories, in such spacetimes torsion vector $E_\mu$ and tensor $M_{\sigma\mu}$ components, coupled to Dirac fermions, can be introduced only through non–minimal torsion–fermion couplings with phenomenological coupling constants \[14\] (see also \[15\]). The presence of phenomenological coupling constants screens real values of torsion vector $E_\mu$ and tensor $M_{\sigma\mu}$ components. Thus, a search for possible ways of measurements of torsion vector $E_\mu$ and tensor $M_{\sigma\mu}$ components through torsion–fermion minimal couplings is of great deal of importance for understanding of correct values of torsion.

We show that these measurements can be in principle possible in curved spacetimes, described by metric tensors with non–diagonal components. These metric tensors define non–vanishing time–space (space–time) components of the vierbein fields, in terms of which slow fermions couple to torsion vector $E_\mu$ and tensor $M_{\sigma\mu}$ components through minimal torsion–fermion couplings.

It is well–known \[31\] that in rotating coordinate systems spacetimes are described by non–diagonal metric tensors. In terrestrial laboratories spacetimes with non–diagonal metric tensors can be in principle realized by means of rotating devices (neutron interferometers) \[32, 33\] (see also a book by Rauh and Werner \[34\] for a necessary information on neutron interferometry. Thus, we propose to measure torsion vector $E_\mu$ and tensor $M_{\sigma\mu}$ components through minimal torsion–fermion couplings in rotating coordinate systems, caused by rotating devices. In our analysis of curved spacetimes in rotating coordinate systems we follow the papers by Hehl and Ni \[35\] and Obukhov, Silenko, and Teryaev \[36, 37\].

The paper is organized as follows. In section II we derive the most general Hamilton operator for the Dirac fermions in the Einstein–Cartan gravity with chameleon and torsion. We adduce the Schrödinger–Pauli equation for slow Dirac fermions, coupled to the effective low–energy potential, caused by gravitational, chameleon and torsion fields. In section III we calculate the vierbein fields, related to the most general metric tensors of the curved spacetimes in the Einstein–Cartan gravity. We demonstrate that these metric tensors provide the Schr"odinger–Pauli equation for slow Dirac fermions, coupled to gravitational, chameleon and torsion fields in the Einstein–Cartan gravity.

### II. SLOW DIRAC FERMIONS IN THE EINSTEIN–CARTAN GRAVITY WITH TORSION AND CHAMELEON

In gravitational theories with chameleon field fermions couple to the chameleon field $\phi(x)$ through the metric $\tilde{g}_{\mu\nu}(x)$ in the Jordan frame related to the metric $g_{\mu\nu}(x)$ in the Einstein frame by $\tilde{g}_{\mu\nu}(x) = f^2(x)g_{\mu\nu}(x)$, where $f(x) = e^{\beta\phi(x)/M_{Pl}}$ is the conformal factor \[17, 18\] (see also \[11\]), $\beta$ is the chameleon–matter coupling constant and $M_{Pl} = 1/\sqrt{8\pi G_N} = 2.435 \times 10^{18}\text{ eV}$ is the reduced Planck mass and $G_N$ is the gravitational coupling \[38\].

For the derivation of the effective low–energy potential we start with the analysis of the Dirac equation for fermions with mass $m$, coupled to the chameleon field in the spacetime with torsion and the metric $d\tilde{s}^2 = \tilde{g}_{\mu\nu}(x)dx^\mu dx^\nu$ (the Jordan–frame metric). The Dirac–fermion action we take in the following form \[39\] (see also \[15\])

$$S_\psi = \int d^4x \sqrt{\tilde{g}} \left( \frac{i}{2} \tilde{\gamma}^\alpha(x) \tilde{\psi}(x) \gamma^\lambda D_\mu \tilde{\psi}(x) - m \tilde{\psi}(x) \tilde{\psi}(x) \right),$$

where $\tilde{\gamma}^\alpha(x)$ and $\gamma^\lambda$ are the vierbein fields, mapping the curved spacetime onto the Minkowski spacetime \[39\] (see also \[11, 15\]), and the Dirac matrices in the Minkowski spacetime \[40\], respectively. The first term in the brackets of Eq. (1) takes the form $\tilde{\varepsilon}^\alpha_\mu(x) \tilde{\psi}(x) \gamma^\lambda \tilde{D}_\mu \tilde{\psi}(x) = \tilde{\varepsilon}^\mu_\alpha(x) (\tilde{\psi}(x) \gamma^\lambda D_\mu \tilde{\psi}(x) - (\tilde{\psi}(x) \tilde{D}_\mu) \gamma^\lambda \tilde{\psi}(x) \tilde{\psi}(x))$, where $D_\mu \tilde{\psi}(x)$ and $(\tilde{\psi}(x) \tilde{D}_\mu)$ are the covariant derivatives defined by \[11, 15\]

$$D_\mu \tilde{\psi}(x) = \partial_\mu \tilde{\psi}(x) - \tilde{\Gamma}_\mu \tilde{\psi}(x), \quad (\tilde{\psi}(x) \tilde{D}_\mu) = \partial_\mu \tilde{\psi}(x) - \gamma^0 \tilde{\Gamma}_\mu^0(x) \gamma^0.$$

The spin affine connection $\tilde{\Gamma}_\mu^\alpha(x)$ is given by \[17\]

$$\tilde{\Gamma}_\mu^\alpha(x) = \frac{i}{4} \tilde{\omega}_{\mu\alpha\beta}(x) \sigma^{\alpha\beta},$$

where $\sigma^{\alpha\beta} = (i/2)(\gamma^\alpha \gamma^\beta - \gamma^\beta \gamma^\alpha)$ are the Dirac matrices in the Minkowski spacetime \[40\] and the spin connection
The derivation of the Dirac equation in the curved spacetime with the metric tensor $\tilde{g}_{\mu\nu}(x)$ can be represented in the following irreducible form \cite{11}

\[
\tilde{T}_{\sigma\mu\nu}(x) = \frac{1}{3} \left( \tilde{g}_{\sigma\mu}(x) \tilde{E}_{\nu}(x) - \tilde{g}_{\sigma\nu}(x) \tilde{E}_{\mu}(x) \right) + \frac{1}{3} \tilde{\varepsilon}_{\sigma\mu\nu\alpha}(x) \tilde{B}^\alpha(x) + \tilde{M}_{\sigma\mu\nu}(x),
\]

where the 4–vector $\tilde{E}_{\nu}(x)$ and axial 4–vector $\tilde{B}^\alpha(x)$ fields, possessing 4 independent components each, are defined by

\[
\tilde{E}_{\nu}(x) = \tilde{g}^{\sigma\mu}(x) \tilde{T}_{\sigma\mu\nu}(x), \quad \tilde{B}^\alpha(x) = \frac{1}{2} \varepsilon^{\sigma\mu\nu}(x) \tilde{T}_{\sigma\mu\nu}(x).
\]

Here $\tilde{\varepsilon}_{\sigma\mu\nu\alpha}(x) = \sqrt{-\tilde{g}(x)} \varepsilon_{\mu\nu\sigma\alpha}$ and $\tilde{\varepsilon}^{\sigma\mu\nu}(x) = \varepsilon^{\sigma\mu\nu}/\sqrt{-\tilde{g}(x)}$ are covariant Levi–Civita tensors in the curved spacetime with the Jordan metric $\tilde{g}_{\mu\nu}(x)$ and the definition $\epsilon^{0123} = -\epsilon_{0123} = +1$ \cite{31}. For the derivation of the axial–vector field $\tilde{B}^\alpha(x)$ in terms of the torsion tensor field $\tilde{T}_{\sigma\mu\nu}(x)$ we have used the relation $\epsilon^{\sigma\mu\nu} \varepsilon_{\mu\nu\beta} = -6 \delta^\alpha_\beta$ \cite{40}.

The residual 16 independent components of the torsion field $\tilde{T}_{\sigma\mu\nu}(x)$ can be attributed to the tensor field $\tilde{M}_{\sigma\mu\nu}(x)$, which obeys the constraints $\tilde{g}^{\mu\nu}(x)\tilde{M}_{\sigma\mu\nu}(x) = \varepsilon^{\sigma\mu\nu}(x)\tilde{M}_{\sigma\mu\nu}(x) = 0$ \cite{14}.

The derivation of the Dirac equation in the curved spacetime with the metric tensor $\tilde{g}_{\mu\nu}(x)$ and torsion we have carried out in the Appendix \cite{12}. The result is

\[
\left( i \tilde{e}_\lambda^\mu(x) \gamma^\lambda D_\mu - \frac{1}{2} i \tilde{\varpi}_{\mu\alpha\beta}(x) \tilde{e}_\lambda^\mu(x) \gamma^\lambda - \frac{1}{2} i \dot{\tilde{\varpi}}_{\mu\alpha\beta}(x) \tilde{e}_\lambda^\mu(x) \left( \gamma^\beta \gamma^\alpha + \frac{i}{4} [\sigma^{\alpha\beta}, \gamma^\lambda] \right) - m \right) \psi(x) = 0,
\]

where $[\sigma^{\alpha\beta}, \gamma^\lambda] = \sigma^{\beta\lambda} \gamma^\alpha - \gamma^\beta \sigma^{\alpha\lambda} = 2i (\gamma^\beta \gamma^\alpha - \gamma^\alpha \gamma^\beta)$ \cite{12}. The Dirac equation Eq.\[8] agrees well with that derived by Kostelecky (see Eq.(18) of Ref.\[39]).

The vierbein fields $\tilde{e}_\mu^\nu(x)$ and $\tilde{e}_\nu^\mu(x)$ in the Jordan frame are related to the vierbein fields $e_\mu^\nu(x)$ and $e_\nu^\mu(x)$ in the Einstein frame by

\[
\tilde{e}_\mu^\nu(x) = f(x) e_\mu^\nu(x), \quad \tilde{e}_\nu^\mu(x) = e_\nu^\mu(x)/f(x).
\]

The Dirac equation Eq.\[8] in its standard form reads

\[
i \frac{\partial \psi(t, \vec{r})}{\partial t} = H \psi(t, \vec{r}),
\]

where $\psi(t, \vec{r})$ is the Dirac wave function of the fermion with mass $m$. The Hamiltonian operator $H$ in its non–perturbative form is given by

\[
H = \tilde{E}_0^\mu(x) \gamma^\mu m - \tilde{E}_0^\mu(x) \tilde{e}_\mu^\nu(x) i \gamma^0 \gamma^\lambda \frac{\partial}{\partial \vec{x}^\nu} - \tilde{E}_0^0(x) \tilde{e}_0^\lambda(x) i \gamma^0 \gamma^\lambda \frac{\partial}{\partial t} + \frac{1}{2} i \tilde{E}_0^\mu(x) \gamma^\mu \left( \tilde{\varpi}_{\sigma\mu\nu}(x) \tilde{e}_\sigma^\lambda(x) \gamma^\lambda + \dot{\tilde{\varpi}}_{\mu\alpha\beta}(x) \tilde{e}_\lambda^\mu(x) \eta^{\alpha\beta} \gamma^\lambda \right) + \frac{1}{2} \tilde{E}_0^0(x) \dot{\tilde{\varpi}}_{\mu\alpha\beta}(x) \gamma^\mu \sigma^{\alpha\beta} \eta^{\gamma\delta} \gamma^\lambda,
\]

where the vierbein field $\tilde{e}_0^\mu(x)$ is defined by $\tilde{E}_0^0(x) = e_0^\mu(x)/(1 - e_0^\mu(x) \tilde{e}_0^\mu(x)) = \tilde{e}_0^\mu(x)/(1 - e_0^\mu(x) \tilde{e}_0^\mu(x)) = 1/\tilde{e}_0^\mu(x)$. The definition for the vierbein field $\tilde{E}_0^0(x)$ follows from the relations $\tilde{e}_0^\mu(x) \tilde{e}_0^\lambda(x) = \delta^\mu_\lambda$ and $\tilde{e}_0^\mu(x) \tilde{e}_0^\mu(x) = \delta^\mu_\mu$. In addition
we have used that \( \{ \sigma^\alpha, \gamma^\lambda \} = \sigma^\alpha \gamma^\lambda + \gamma^\lambda \sigma^\alpha = -2 \epsilon^{\lambda \delta \beta} \gamma^\gamma \) \([13]\) and \( \gamma^\gamma = i \gamma^0 \gamma^1 \gamma^2 \gamma^3 \) \([10]\). For the derivation of the effective low-energy potential it is convenient to transcribe the Hamilton operator Eq. \((14)\) into the form

\[
H = \tilde{E}_0^0(x) \gamma^0 m + \frac{1}{2} i \tilde{E}_0^0(x) \left( \hat{T}^\alpha \rho_{\alpha \mu}(x) \gamma^\mu(x) + \tilde{\omega}_{\mu \beta}(x) \gamma^\mu(x) \gamma^\beta \right) + \frac{1}{4} \tilde{E}_0^0(x) \left( \tilde{\omega}_{\mu \beta}(x) \gamma^\mu(x) + \tilde{\omega}_{\mu \beta}(x) \gamma^\mu(x) \right) \epsilon^{jk} \Sigma_k - \tilde{E}_0^0(x) \gamma^\mu(x) i \frac{\partial}{\partial x^j} - \tilde{E}_0^0(x) \gamma^\mu(x) i \gamma^\mu \gamma^\beta \frac{\partial}{\partial t} + \frac{1}{2} \tilde{E}_0^0(x) \left( \hat{T}^\alpha \rho_{\alpha \mu}(x) \gamma^\mu(x) + \tilde{\omega}_{\mu \beta}(x) \gamma^\mu(x) \gamma^\beta \right) \epsilon^{jk} \gamma^5, \tag{12}
\]

where we have denoted \( \tilde{\omega}_{\mu \beta}(x) = \tilde{\omega}_{\mu \beta}(x) - \tilde{\omega}_{\mu \beta}(x) \) and used \( \epsilon^{jk} \gamma^5 = -\epsilon^{jk} \) \([10]\). It is well-known \([13]\) that the Hamilton operator Eq. \((12)\) is not hermitian. This is because of the factor \( \sqrt{-g} \) in the definition of the 4-dimensional covariant volume element \( d^4x \sqrt{-g} \) in the curved spacetime. In order to deal with the hermitian Hamilton operator we have to make the following transformation of the fermion wave function and the Hamilton operator \([13]\):

\[
\psi(x) = \left( \sqrt{-g(x)} \tilde{e}_0^0(x) \right)^{-1/2} \psi'(x),
\]

\[
H' = \left( \sqrt{-g(x)} \tilde{e}_0^0(x) \right)^{1/2} H \left( \sqrt{-g(x)} \tilde{e}_0^0(x) \right)^{-1/2} - i \left( \sqrt{-g(x)} \tilde{e}_0^0(x) \right)^{1/2} \frac{\partial}{\partial t} \left( \sqrt{-g(x)} \tilde{e}_0^0(x) \right)^{-1/2}. \tag{13}
\]

We would like to note that the last term in the Hamilton operator \(H'\) acts on the fermion wave function as a multiplication operator and does not differentiate it with respect to time. This means that we do not need to add in the definition of the Hamilton operator \(H'\) a time derivative operator \(i \partial / \partial t\) \([51, 52]\). The Dirac equation for the wave function \(\psi'(t, \vec{r})\) retains its standard form

\[
\frac{i}{\sqrt{-g}} \frac{\partial \psi'(t, \vec{r})}{\partial t} = H' \psi'(t, \vec{r}), \tag{14}
\]

where the hermitian Hamilton operator \(H'\) is given by

\[
H' = \tilde{E}_0^0(x) \gamma^0 m + \frac{1}{2} i \tilde{E}_0^0(x) \left( \hat{T}^\alpha \rho_{\alpha \mu}(x) \gamma^\mu(x) + \tilde{\omega}_{\mu \beta}(x) \gamma^\mu(x) \gamma^\beta \right) + \frac{1}{2} \tilde{E}_0^0(x) \left( \hat{T}^\alpha \rho_{\alpha \mu}(x) \gamma^\mu(x) + \tilde{\omega}_{\mu \beta}(x) \gamma^\mu(x) \right) \epsilon^{jk} \Sigma_k - \tilde{E}_0^0(x) \gamma^\mu(x) i \frac{\partial}{\partial x^j} - \tilde{E}_0^0(x) \gamma^\mu(x) i \gamma^\mu \gamma^\beta \frac{\partial}{\partial t} + \frac{1}{2} \tilde{E}_0^0(x) \left( \hat{T}^\alpha \rho_{\alpha \mu}(x) \gamma^\mu(x) + \tilde{\omega}_{\mu \beta}(x) \gamma^\mu(x) \gamma^\beta \right) \epsilon^{jk} \gamma^5. \tag{15}
\]

Using the relation (see Eq.(A-9) of Ref.\([13]\))

\[
\frac{1}{2} \sqrt{-g} \frac{\partial}{\partial t} \left( \sqrt{-g} \tilde{e}_0^0(x) \right) = -\frac{1}{2} \tilde{T}^\alpha \rho_{\alpha \mu}(x) \gamma^\mu(x) - \frac{1}{2} i \tilde{\omega}_{\mu \beta}(x) \gamma^\mu(x) \gamma^\beta - \frac{1}{2} i \sqrt{-g} \frac{\partial}{\partial x^j} \left( \sqrt{-g} \tilde{e}_0^0(x) \right), \tag{16}
\]

we transcribe the Hamilton operator Eq.\((15)\) into the form

\[
H' = \tilde{E}_0^0(x) \gamma^0 m - \frac{1}{2} i \tilde{E}_0^0(x) \left( \hat{T}^\alpha \rho_{\alpha \mu}(x) \gamma^\mu(x) + \tilde{\omega}_{\mu \beta}(x) \gamma^\mu(x) \gamma^\beta \right) + \frac{1}{4} \tilde{E}_0^0(x) \left( \hat{T}^\alpha \rho_{\alpha \mu}(x) \gamma^\mu(x) + \tilde{\omega}_{\mu \beta}(x) \gamma^\mu(x) \gamma^\beta \right) \epsilon^{jk} \Sigma_k - \tilde{E}_0^0(x) \gamma^\mu(x) i \frac{\partial}{\partial x^j} - \tilde{E}_0^0(x) \gamma^\mu(x) i \gamma^\mu \gamma^\beta \frac{\partial}{\partial t} + \frac{1}{2} \tilde{E}_0^0(x) \left( \hat{T}^\alpha \rho_{\alpha \mu}(x) \gamma^\mu(x) + \tilde{\omega}_{\mu \beta}(x) \gamma^\mu(x) \gamma^\beta \right) \epsilon^{jk} \gamma^5. \tag{17}
\]
According to the Foldy–Wouthuysen classification \[11\], the operators in the first two lines of Eq. (17) are even, whereas all other operators are odd. For a derivation of a low–energy effective Hamilton operator of slow fermions all odd operators should be removed by some unitary transformations \[11\]. Skipping standard intermediate Foldy–Wouthuysen calculations, which are given in the Appendix, we arrive at the Schrödinger–Pauli equation

\[
i \frac{\partial \Psi(t, \vec{r})}{\partial t} = \left( -\frac{1}{2m} \Delta + \Phi_{\text{eff}}(t, \vec{r}, \vec{\sigma}) \right) \Psi(t, \vec{r}),
\]

where \( \Psi(t, \vec{r}) \) is the large component of the Dirac wave function of slow fermions with mass \( m \) and \( \Phi_{\text{eff}}(t, \vec{r}, \vec{\sigma}) \) is the effective low–energy potential for slow fermions, coupled to gravitational, chameleon and torsion fields, and \( \vec{\sigma} \) are the \( 2 \times 2 \) Pauli matrices \[40\]. The exact expression of the potential \( \Phi_{\text{eff}}(t, \vec{r}, \vec{\sigma}) \), calculated to order \( O(1/m) \), is given by Eq. (A-15) of the Appendix.

### III. CONCLUSIVE DISCUSSION

We have analysed the low–energy approximation of the Dirac equation for fermions with mass \( m \) in the Einstein–Cartan gravity with torsion and chameleon. Using the Foldy–Wouthuysen transformations we have derived the most general low–energy potential to order \( 1/m \) for slow Dirac fermions, coupled to gravitational, torsion and chameleon fields. The aim of the derivation of this effective low–energy potential is addressed to the investigation of spacetimes in which torsion vector \( \vec{T} \) and tensor components, coupled minimally to slow Dirac fermions, can be in principle observable. As has been shown in \[15\] for metric tensors, yielding vanishing non–diagonal time–space (space–time) components of the vierbein fields, in the perturbative regime for gravitational, torsion and chameleon fields only torsion axial–vector components survive in the low–energy approximation of the minimal torsion–fermion couplings.

In the Appendix for the derivation of the effective low–energy potential \( \Phi_{\text{eff}}(t, \vec{r}, \vec{\sigma}) \) (see Eq. (A-15)) we have introduced the operators \( A, B, C, D_j, F_j, G_j, K \) and \( L^j \), which are defined in Eq. (A-2). In the approximation \[15\] applied to our approach these operators behave as follows

\[
A = 1 + O(g, \phi), \quad B = 0, \quad C^j = \frac{1}{4} B^j = O(g, \phi), \quad D_j = \tilde{\delta}^j = O(g, \phi), \quad F_j = 0, \quad G_j = O(g, \phi), \quad K = -\frac{1}{4} \hat{K} + O(g, \phi), \quad L^j = 0,
\]

where \( B^j = \frac{1}{2} \tilde{\epsilon}^{jkl}(T_{jk0} + T_{kj0} + T_{0jk}) \) and \( \hat{K} = \frac{1}{2} \tilde{\epsilon}^{jkl}T_{ijk} \) are the torsion axial–vector and pseudoscalar components \[15\], and \( O(g, \phi) \) are the linear order contributions of gravitational and chameleon fields. The non–trivial linear order contributions of the torsion vector and tensor components can appear only in spacetimes with non–diagonal metric tensors, yielding non–vanishing non–diagonal time–space (space–time) components of the vierbein fields. It is well–known that in the rotating coordinate system spacetime is described by a non–diagonal metric tensor \[31\] with the non–vanishing time–space (space–time) components \( g_{0j}(x) \) proportional to the angular velocity (see also \[35–37\]).

The phase–shift induced by a rotational motion of an optical interferometer was first proposed by Sagnac \[42\] and observed by Michelson, Gale, and Pearson \[43\]. In spite of the fact that the inertial properties of photons and neutrons are different, the analogous effect for the phase–shift of slow neutrons was predicted by Page \[44\] and measured by Werner et al. \[45\], Atwood et al. \[32\] and Mashhoon \[33\]. For the measurement of such a phase–shift Atwood et al. \[32\] and Mashhoon \[33\] used the rotating two–crystal neutron interferometer and the neutron interferometer in the rotating reference frame, respectively. According to an equivalence between a rotating coordinate system and a gravitational field or a curved spacetime with a corresponding metric tensor \[31\], the experimental setup of the experiments by Atwood et al. \[32\] and Mashhoon \[33\] should determine metric tensors of curved spacetimes, created by rotating neutron interferometers in the gravitational field of the Earth.

Following such an equivalence, for the experimental analysis of fermion–torsion interactions, described by the effective low–energy potential \( \Phi_{\text{eff}}(t, \vec{r}, \vec{\sigma}) \) given by Eq. (A-15), we have to determine the metric tensor of the curved spacetime and calculate the vierbein fields, caused by the experimental setup of possible experiments. The line element in spacetimes, created by rotating devices in an arbitrary gravitational field, we take in the most general form, proposed by Obukhov, Silenko, and Teryaev \[37\]:

\[
d\vec{s}^2 = \tilde{V}^2(x) dt^2 + \eta_{ij} \tilde{W}^i_j(x) \tilde{W}^j_i(x) \left[ dx^j - K^j(x) dt \right] \left[ dx^i - K^i(x) dt \right],
\]

where the functions \( \tilde{V}^2(x) \) and \( \tilde{W}^i_j(x) \) are defined by an arbitrary gravitational field. In comparison with Obukhov et al. \[37\] they are modified by the chameleon field. In turn, the functions \( K^j(x) \) are caused by rotations. The
components of the metric tensor $\bar{g}_{\mu\nu}(x)$ are equal to

$$
\bar{g}_{00}(x) = \bar{V}^2(x) + (1 - \xi) \eta_{ij} \bar{W}_i(x)\bar{W}_j(x)K^i(x)K^j(x), \quad \bar{g}_{0j}(x) = -\eta_{ij} \bar{W}_i(x)\bar{W}_j(x)K^i(x),
$$

$$
\bar{g}_{ij}(x) = \eta_{ij} \bar{W}_i(x)\bar{W}_j(x).
$$

(21)

The vierbein fields $\bar{e}_\mu^j(x)$ are defined by the relation $\bar{g}_{\mu\nu} = \eta_{\alpha\beta}\bar{e}_\mu^\alpha(x)\bar{e}_\nu^\beta(x)$. Solving Eq. (22) for $\bar{g}_{00}(x)$ and $\bar{g}_{ij}(x)$ we get

$$
\bar{e}_0^0(x) = \sqrt{\bar{V}^2(x)} + (1 - \xi) \eta_{ij} \bar{W}_i(x)\bar{W}_j(x)K^i(x)K^j(x), \quad \bar{e}_0^j(x) = -\sqrt{\xi} \bar{W}_j(x)K^j(x),
$$

$$
\bar{e}_i^0(x) = 0, \quad \bar{e}_i^j(x) = \bar{W}_j(x),
$$

(23)

where $\xi$ is a parameter, which can be fixed from Eq. (21) for $\bar{g}_{0j}(x)$. Indeed, using the vierbein fields Eq. (23) we obtain

$$
\bar{g}_{0j}(x) = -\sqrt{\xi} \eta_{ij} \bar{W}_i(x)\bar{W}_j(x)K^i(x).
$$

(24)

From the comparison of Eq. (21) with Eq. (24) we obtain $\xi = 1$. This gives the vierbein fields, given by Eq. (23), equal to

$$
\bar{e}_0^0(x) = \bar{V}(x), \quad \bar{e}_0^j(x) = -\bar{W}_j(x)K^j(x), \quad \bar{e}_i^0(x) = 0, \quad \bar{e}_i^j(x) = \bar{W}_j(x).
$$

(25)

For the calculation of the vierbein fields $\bar{e}_\mu^j(x)$ we use the relations $\bar{e}_\mu^j(x) = \delta_\mu^j\bar{e}_\nu^0(x)\bar{e}_\nu^\nu(x)$.

(26)

Skipping intermediate calculations we obtain

$$
\bar{e}_0^0(x) = \frac{1}{\bar{V}(x)}, \quad \bar{e}_0^j(x) = 0, \quad \bar{e}_i^0(x) = \frac{K^i(x)}{\bar{V}(x)}, \quad \bar{e}_i^j(x) = \bar{W}_j(x).
$$

(27)

The vierbein fields in Eq. (26) and Eq. (27) have been calculated at the assumption that the functions $W_j^j(x)$ and $W_j^i(x)$ obey the orthogonality relations

$$
W_j^j(x)W_i^j(x) = \delta_i^j, \quad W_j^i(x)W_i^j(x) = \delta_i^j,
$$

(28)

which are fulfilled for the Schwarzschild metric in the weak gravitational field of the Earth approximation. For the verification of the correctness of the obtained vierbein fields we construct the metric tensor $\bar{g}_{\mu\nu}(x)$. In terms of the vierbein fields $\bar{e}_\mu^j(x)$ it is determined by

$$
\bar{g}^{\mu\nu}(x) = \eta^{\alpha\beta}\bar{e}_\mu^\alpha(x)\bar{e}_\nu^\beta(x).
$$

(29)

Using the vierbein fields Eq. (24) for the components of the metric tensor $\bar{g}^{\mu\nu}(x)$ we obtain the following expressions

$$
\bar{g}^{00}(x) = \frac{1}{\bar{V}^2(x)}, \quad \bar{g}^{0j}(x) = \frac{K_j(x)}{\bar{V}(x)},
$$

$$
\bar{g}^{ij}(x) = \frac{K_i(x)K_j(x)}{\bar{V}^2(x)} + \eta^{ij}\bar{W}_k(x)\bar{W}_k(x).
$$

(30)

One may show that the metric tensors $\bar{g}_{\mu\nu}(x)$ and $\bar{g}^{\mu\nu}(x)$, given by Eq. (21) and Eq. (30), respectively, obey the relation $\bar{g}^{\alpha\beta}(x)\bar{g}_{\alpha\beta}(x) = \delta^{\mu\nu}$. Then, because of $\bar{e}_i^0(x) = \bar{e}_i^j(x) = 0$ for the vierbein fields Eq. (24) and Eq. (27) we get $\bar{E}_0^0(x) = \bar{e}_0^0(x)$. 

For the vierbein fields Eq. (25) and Eq. (27) torsion–fermion interactions are yielded by the operators $C_\ell$, $G_j$ and $K$ only. Since the contributions of the torsion–fermion interactions, caused by the operators $G_j$ and $K$, are suppressed by a factor of $1/m$, below we analyse the contributions of the torsion–fermion interactions, caused by the operator $C_\ell$, which appear to order $O(1)$ in the large fermion mass expansion. We take also into account the contributions of the operators $A$, $B$ and $L^j$ in order to derive a complete set of gravitational, chameleon and torsion interactions with slow fermions to order $O(1)$ in the large fermion mass expansion. The analysis of contributions of the operators $G_j$ and $K$ goes beyond the scope of this paper. We are planning to perform such an analysis in our forthcoming publication.

For the vierbein fields Eq. (25) and Eq. (27) the operators Eq. (A-2) are given by

\[
A = e^0_0(x),
B = -\frac{1}{2} i e^0_0(x) \frac{1}{\sqrt{-g}} \frac{\partial}{\partial x^i} \left( \sqrt{-\tilde{g}} \tilde{e}^i_0(x) \right) + \frac{1}{2} i \left( e^0_0(x) \right)^2 \tilde{e}^i_0(x) \frac{1}{\sqrt{-\tilde{g}(x)}} \frac{\partial}{\partial x^i} \left( \sqrt{-\tilde{g}(x)} \tilde{e}^j_0(x) \right),
C^i = \frac{1}{4} e^0_0(x) \left( \tilde{\omega}_{0jk}(x) \tilde{e}^0_0(x) + \tilde{\omega}_{0jk}(x) \tilde{e}^0_0(x) + \tilde{\omega}_{0jk}(x) \tilde{e}^0_0(x) + \tilde{\omega}_{0jk}(x) \tilde{e}^0_0(x) \eta^{jk} \right)
\]

\[
+ \frac{1}{2} \left( e^0_0(x) \right)^2 \tilde{e}^j_0(x) \frac{1}{\sqrt{-\tilde{g}(x)}} \frac{\partial}{\partial x^j} \left( \sqrt{-\tilde{g}(x)} e^j_0(x) \right),
K = -\frac{1}{4} \tilde{\omega}_{0jk}(x) e^0_0(x) \tilde{e}^j_0(x) \eta^{jk},
L^j = -e^0_0(x) \tilde{e}^j_0(x).
\] (31)

For the analysis of interactions of slow Dirac fermions with gravitational, chameleon and torsion fields, caused by non–diagonal space–time components $\tilde{e}^j_0(x)$ of the vierbein fields we assume a motion of Dirac fermions with mass $m$ in the curved spacetime with the Schwarzschild metric, taken in the weak gravitational field of the Earth approximation and modified by the contributions of the chameleon field and rotation. The line element of such a spacetime is given by

\[
d\mathbf{s}^2 = (1 + 2U_+) \, dt^2 + 2 \left( 1 - 2U_- \right) \mathbf{K} \cdot d\mathbf{r} \, dt - (1 - 2U_-) \, dr^2,
\] (32)

where we have neglected the contribution of the terms of order $\mathbf{K}^2$ that is well justified in the terrestrial laboratories and kept the contributions of the chameleon field to linear order $O(1)$. The potentials $U_\pm$ are equal to $15$

\[
U_\pm = U_E \pm \frac{\beta}{M_{\text{Pl}}} \phi(x),
\] (33)

where $U_E = \tilde{g} \cdot \tilde{r}$ is the Newtonian gravitational potential of the Earth and $\tilde{g}$ is the gravitational acceleration $15$. To linear order contributions of the gravitational and chameleon field the vierbein fields Eq. (25) and Eq. (27) read

\[
e^0_0(x) = 1 + U_+, \quad \tilde{e}^0_0(x) = - \left( 1 - U_- \right) K^j(x), \quad \tilde{e}^0_0(x) = 0, \quad \tilde{e}^j_0(x) = 1 - U_+, \quad \tilde{e}^j_0(x) = + \left( 1 - U_- \right) K_j(x), \quad \tilde{e}^j_0(x) = 0, \quad \tilde{e}^j_0(x) = 1 - U_+, \quad \tilde{e}^j_0(x) = 0.
\] (34)

The diagonal components of the vierbein fields agree well with those, calculated in $15$. In such a spacetime the operators $A$, $B$, $C_\ell$ and $L^j$ are equal to

\[
A = 1 + U_+, 
B = -\frac{1}{2} i \text{div} \mathbf{K}, 
C_\ell = -\frac{1}{4} (\text{rot} \mathbf{K})^i + \frac{1}{4} B^i + \frac{1}{4} \ell^{ijk} K_j T_{0ik} + \frac{1}{4} \ell^{ijk} T_{jka} K^a = 
-\frac{1}{4} (\text{rot} \mathbf{K})^i + \frac{1}{4} B^i + \frac{1}{6} K \ell^i + \frac{1}{4} \ell^{ijk} K_j M_{0ik} + \frac{1}{4} \ell^{ijk} M_{jka} K^a, 
L^j = -K^j,
\] (35)
where to linear order approximation $\tilde{T}_{\mu \nu} = T_{\mu \nu}$, $B^\ell = \frac{1}{2} \epsilon^{\ell j k} (T_{j k \theta} + T_{k \theta j} + T_{\theta j k})$ and $K = \frac{1}{2} \epsilon^{a b c} T_{a b c}$. Then, we have used Eq. (3) and $\sqrt{-g} = 1 + U_+ - 3U_-$, calculated to linear order of the gravitational and chameleon field and at the neglect the contribution of order $O(\tilde{K}^2)$. For the calculation of the operators in Eq. (35) we have not distinguished indices in the Minkowski and curved spacetime. This is correct, since the operators Eq. (35) are defined in the perturbative regime for gravitational, chameleon and torsion fields and describe corresponding interactions of slow Dirac fermions in the Minkowski spacetime.

For curved spacetimes with the metric Eq. (32) the contribution of the operators Eq. (35) to the effective low-energy potential $\Phi_{\text{eff}}(t, \vec{r}, \vec{\sigma})$, calculated to order $O(1)$ in the large fermion mass expansion, is given by

$$
\Phi_{\text{eff}}(t, \vec{r}, \vec{\sigma}) = m U_+ - \frac{1}{2} i \text{div}\tilde{K} - \tilde{K} \cdot \vec{\nabla} + \frac{1}{4} \vec{\sigma} \cdot \vec{\omega} - \frac{1}{4} \vec{\sigma} \cdot \vec{B} - \frac{1}{6} K \vec{\sigma} \cdot \vec{K} + \frac{1}{4} \sigma_\ell \epsilon^{\ell j k} K_j \mathcal{M}_{00k} + \frac{1}{4} \sigma_\ell \epsilon^{\ell j k} \mathcal{M}_{jka} K^a,
$$

(36)

where $\sigma_j = (-\vec{\sigma})_j$ and we have used that $\epsilon_{j k a 0} = \epsilon_{j k a}$ with $\epsilon_{123} = +1$. It is important to note that torsion vector $\vec{E}$ components do no couple to slow Dirac fermions to order $O(1)$ in the large fermion mass expansion and to linear order approximation of the torsion field.

For rotating coordinate systems with an angular velocity $\vec{\omega}$ the vector functions $K^i$ are equal to $K^i = - (\vec{\omega} \times \vec{r})^i$, with $\vec{r}$ the orbital momentum operator of slow fermions, and the third term $-\vec{\omega} \cdot \vec{S}$ is the spin operator of slow fermions, agree well with the results, obtained by Hehl and Ni [35]. The interactions $-\vec{\omega} \cdot \vec{L}$ and $-\vec{\omega} \cdot \vec{S}$ were analysed in the experiments by Werner, Staudenmann, and Colella [36] and by Mashhoon [33]. The fourth term, describing torsion–spin–matter interaction of the torsion axial–vector components, was derived by Lämmerzahl and Obukhov, Silenko, and Teryaev [13] (see also [15]). The other terms in the effective low-energy potential Eq. (35) are new. The fifth term $(1/6)K \vec{\sigma} \cdot (\vec{\omega} \times \vec{r})$ is a new low–energy interaction of the torsion pseudoscalar component $K$ with slow Dirac fermions. In turn, the last two terms in Eq. (36) describe new low–energy interactions
d

d of torsion tensor $\mathcal{M}_{00k}$ and $\mathcal{M}_{jka}$ components with slow Dirac fermions, caused by minimal torsion–fermion couplings without phenomenological coupling constants. According to estimates by Kostelecky et al. [11], the constant torsion tensor components $\mathcal{M}_{00k}$ and $\mathcal{M}_{jka}$, multiplied by a phenomenological coupling constant $\xi_5^{(5)}$, are restricted by $|\xi_5^{(5)}| \mathcal{M}_{00k} < 10^{-27}$ and $|\xi_5^{(5)}| \mathcal{M}_{jka} < 10^{-26}$, respectively. Recently the non–minimal torsion–matter couplings have been also discussed by Puetzfeld and Obukhov [10]. As regards torsion vector components $\vec{E}$, we have found that slow Dirac fermions do not couple to them to order $O(1)$ in the large fermion mass expansion and to linear order of the torsion field approximation.

The upper bound of the linear superposition of the constant torsion vector and axial–vector components $|\gamma| < 9.1 \times 10^{-23}$ GeV, measured by Lehnert, Snow and Yan [16] by means of an investigation of a spin rotation of cold neutrons in the liquid $^4$He, is by a factor $10^3$ larger compared with the estimate $|\gamma| < 10^{-27}$ GeV, obtained in [15]. Thus, we have shown that to linear order of the torsion field approximation in spacetimes of rotating coordinate systems the contributions of only torsion pseudoscalar $K$ and tensor $\mathcal{M}_{00k}$ and $\mathcal{M}_{jka}$ components, caused by minimal torsion–fermion couplings, appear to order $O(1)$ in the large fermion mass expansion. The certain steps in the realization of curved spacetimes in terrestrial laboratories by using rotating devices (neutron interferometers) were made by Atwood et al. [52] and Mashhoon [53]. The measurements of the transition frequencies between quantum gravitational states of ultracold neutrons in the qBounce experiments [1] as functions of an angular velocity $\vec{\omega}$ of a rotating mirror should provide a new level of highly precise probes of the properties of the Einstein–Cartan gravity, dark energy and evolution of the Universe. Of course, the measurements of new gravitational, chameleon and torsion interactions in Eq. (37) as well as other interactions in the effective low–energy potential Eq. (4–15) can be carried out by using rotating neutron interferometers [52, 54].

Now we would like to discuss shortly the Foldy–Wouthuysen method [41], which we use in this paper for the derivation of the effective low-energy potential for slow Dirac fermions, coupled to gravitational, chameleon and torsion fields. Mainly this discussion concerns uniqueness and accuracy of the Foldy–Wouthuysen representation of
the Dirac Hamilton operator, obtained by the Foldy–Wouthuysen transformation. It is well-known that the Foldy–Wouthuysen method of a transformation of a Dirac Hamilton operator to a form, containing only even (diagonal) operators (as regards the definition of odd and even operators a reader might consult Ref. [11] or look up in the Appendix to this paper), is not unique and there are some other methods of transformation of a Dirac Hamilton operator to a diagonal form. A very nice survey of possible methods of transformation of a Dirac Hamilton operator for fermions with mass $m$ to a diagonal form, containing only even operators, one can find in the paper by Vries [48].

The Foldy–Wouthuysen method, removing odd operators from a Dirac Hamilton operator for fermions with mass $m$ by Foldy–Wouthuysen unitary transformations, allows to reduce a Dirac Hamilton operator to a non-relativistic form in the approximation of a large fermion mass expansion by a set of unitary transformations or by the iterative Foldy–Wouthuysen method. The obtained non-relativistic Hamilton operator is given by an infinite series of even operators in powers of $1/m$, which does not seem to give hope for a closed-form operator. A problem of a closed form of a transformed Dirac Hamilton operator, expressed in terms of only even operators, was investigated by Eriksen [17]. Eriksen showed that the unitary transformation $e^{iS} = \sqrt{\gamma^0 H/\sqrt{\lambda}}$, where $\gamma^0$ and $H$ are the Dirac matrix and a Dirac Hamilton operator, allows to transform a Dirac Hamilton operator $H$ to a square root of an even operator. However, Eriksen’s unitary operator $e^{i\overline{S}} = \sqrt{\gamma^0 H/\sqrt{\lambda}}$, leading to a closed-form of a transformed Dirac Hamilton operators, suffers from ambiguous definition. In order to define the operator $e^{iS} = \sqrt{\gamma^0 H/\sqrt{\lambda}}$ unambiguously one has to assume that the square root of a unit operator is a unit operator. For recent discussion of a square root operator definition and analyses of the Dirac Hamilton operators by means of the Eriksen method we propose a reader the papers by Silenko [49], Neznamov and Silenko [50] and Silenko [51–53]. According to Eriksen [17], the unitary operator $e^{iS} = \sqrt{\gamma^0 H/\sqrt{\lambda}}$, providing an exact diagonalization of the Dirac Hamilton operator $H$, can be defined by $e^{iS} = \sqrt{\gamma^0 \lambda} = (1 + \gamma^0 \lambda)/\sqrt{1 + \gamma^0 \lambda} + \gamma^0 \lambda$, where $\lambda = H/\sqrt{\lambda}$. Another problem of the Foldy–Wouthuysen method concerns an accuracy of the Foldy–Wouthuysen representation of a Dirac Hamilton operator in comparison with a large fermion mass expansion of an exact form of a transformed Dirac Hamilton operator. For the first time such a problem was discussed by Eriksen and Kolsrud [54]. Recently this problem has been investigated by Neznamov and Silenko [50] and Silenko [51–53]. According to [54] and [50–53], the Foldy–Wouthuysen representation of a Dirac Hamilton operator, obtained by a set of unitary transformations, can but not coincide with a large fermion mass expansion of an exact transformed Dirac Hamilton operator, diagonalized by means of only one unitary transformation (e.g. the Eriksen transformation). As has been pointed out by Neznamov and Silenko [50] and Silenko [51–53], such a disagreement can be explained by a non-commutativity of Foldy–Wouthuysen unitary transformations in the iterative Foldy–Wouthuysen method. For example, in our case we have diagonalized the Dirac Hamilton operator Eq. (17) by three unitary transformations of the Dirac wave functions $e^{iS_1}$, $e^{iS_2}$ and $e^{iS_3}$ (see the Appendix), respectively. A resulting unitary transformation is equal to $e^{iS} = e^{iS_3}e^{iS_2}e^{iS_1}$. According to [50] and [51–53], a coincidence of our result for the effective low-energy potential Eq. (A-15) with a large fermion mass expansion of an exact transformed Dirac Hamilton operator, obtained by means of only one unitary transformation (e.g. the Eriksen transformation), can be expected only for the validity of the relation $e^{iS_3}e^{iS_2}e^{iS_1} = e^{i(S_1+S_2+S_3)}$, which demands a commutativity of the operators $[S_i, S_j] = 0$ for $i \neq j = 1, 2, 3$. Since the unitary operators do not commute $[S_i, S_j] \neq 0$ for $i \neq j = 1, 2, 3$ (see Appendix), such the relation $e^{iS_3}e^{iS_2}e^{iS_1} = e^{i(S_1+S_2+S_3)}$ is not valid and one may expect some deviations of the effective low-energy potential Eq. (A-15) from that derived by a large fermions mass expansion of an exact diagonalized Dirac Hamilton operator. However, one may show that any deviations can appear only to order $O(1/m^2)$. The later can be justified by the observation that $[S_2, S_1] = O(1/m^2)$ and $[S_3, S_1] = O(1/m^3)$ for $j = 1, 2$. Hence, to order $O(1/m)$, which we have kept for the derivation of the effective low-energy potential $\Phi_{\text{eff}}$ in Eq. (A-15), these two fermion mass expansions should coincide. A method of the calculation of the corrections to the Foldy–Wouthuysen representation of a Dirac Hamilton operator has been discussed in detail by Silenko [53]. For example, suppose that two Foldy–Wouthuysen unitary transformations with operators $e^{iS_1}$ and $e^{iS_2}$, performed one after another, diagonalize a Dirac Hamilton operator, i.e. $H \rightarrow H_{\text{FW}}$, where $H_{\text{FW}}$ is a Dirac Hamilton operator in the Foldy–Wouthuysen representation. According to Silenko [53], an additional Foldy–Wouthuysen unitary transformation $U_{\text{corr}} = \exp(-\frac{i}{2} [S_1, S_2])$ should allow to cancel an error of the iterative Foldy–Wouthuysen method in the leading order. Such a correction is valid if the commutators $[S_1, [S_1, S_2]]$ and $[S_2, [S_1, S_2]]$ and commutators of higher orders can be neglected with respect to the commutator $[S_1, S_2]$. Since in our case this constraint is fulfilled, the correction to the effective low-energy potential Eq. (A-15) can be calculated by means of the unitary transformation $U_{\text{corr}} = \exp(-\frac{i}{2} [S_1, S_2])$. However, in our case $[S_1, S_2] = O(1/m^3)$ and the effective low-energy potential Eq. (A-15) is calculated to order $O(1/m)$. This might imply that the effective low-energy potential Eq. (A-15) should in principle coincide with a large fermion mass expansion to order $O(1/m)$ of an exact diagonalized Dirac Hamilton operator by, for example, the Eriksen method [17, 50–53].
IV. ACKNOWLEDGEMENTS

We are grateful to Hartmut Abele for stimulating discussions. This work was supported by the Austrian “Fonds zur Förderung der Wissenschaftlichen Forschung” (FWF) under the contract I689-N16.

V. APPENDIX A: DERIVATION OF THE EFFECTIVE LOW-ENERGY POTENTIAL $\Phi_{\text{eff}}(t, \vec{r}, \vec{\sigma})$ IN EQ. (18)

For the derivation of the low-energy potential $\Phi_{\text{eff}}(t, \vec{r}, \vec{\sigma})$ for slow Dirac fermions with mass $m$ in the Schrödinger–Pauli equation Eq. (18) we define the Hamilton operator Eq. (17) as follows

$$H' = A i \gamma_0 m + B + C i \Sigma_i + D^i_j i \gamma_0 \gamma^j \frac{\partial}{\partial x^j} + F_j i \gamma_0 \gamma^j \frac{\partial}{\partial t} + G_j i \gamma_0 \gamma^j + K \gamma^5 + L^j i \frac{\partial}{\partial x^j}$$  (A-1)

where we have defined

\begin{align*}
A &= \bar{E}_0^0(x), \\
B &= -\frac{1}{2i} \bar{E}_0^0(x) \frac{1}{\sqrt{-\bar{g}}} \frac{\partial}{\partial x^j} \left( \sqrt{-\bar{g}} \bar{\epsilon}_0^j(x) \right) + \frac{1}{2i} \left( \bar{E}_0^0(x) \right)^2 \frac{1}{\sqrt{-\bar{g}}} \frac{\partial}{\partial x^j} \left( \sqrt{-\bar{g}} \bar{\epsilon}_0^j(x) \right), \\
C^j &= \frac{1}{4} \bar{E}_0^0(x) \left( \tilde{\omega}_{\mu j}^k(x) \bar{\epsilon}_0^k(x) + \tilde{\omega}_{\mu [j]}^k(x) \bar{\epsilon}_0^k(x) \right) e^{ik}, \\
D^j_i &= -\bar{E}_0^0(x) \bar{\epsilon}_0^j(x), \\
F_j &= -\bar{E}_0^0(x) \bar{\epsilon}_0^j(x), \\
G_j &= \frac{1}{2} \bar{E}_0^0(x) \left( \bar{T}_{\mu j}^\alpha \bar{\epsilon}_0^\alpha(x) + \tilde{\omega}_{\mu j}^\beta(x) \bar{\epsilon}_0^\beta(x) \gamma^{\beta} \right) - \frac{1}{2} \left( \bar{E}_0^0(x) \right)^2 \frac{1}{\sqrt{-\bar{g}}} \frac{\partial}{\partial x^j} \left( \sqrt{-\bar{g}} \bar{\epsilon}_0^j(x) \right) - \frac{1}{2} \bar{E}_0^0(x) \bar{\epsilon}_0^j(x) \frac{1}{\sqrt{-\bar{g}}} \frac{\partial}{\partial x^j} \left( \sqrt{-\bar{g}} \bar{\epsilon}_0^j(x) \right), \\
K &= \frac{1}{4} \bar{E}_0^0(x) \tilde{\omega}_{\mu j}^k(x) \bar{\epsilon}_0^k(x) e^{ik}, \\
L^j &= -\bar{E}_0^0(x) \bar{\epsilon}_0^j(x).  \quad (A-2)
\end{align*}

For the elimination of the odd operators we perform the Foldy–Wouthuysen unitary transformation of the wave function $\psi'(x) = e^{-iS_1} \psi_1(x)$ and the Hamilton operator \[ 11 \]

$$H_1 = e^{+iS_1} H' e^{-iS_1} - i e^{iS_1} \frac{\partial}{\partial t} e^{-iS_1} = H' - \frac{\partial S_1}{\partial t} + i \left[ S_1, H' \right] - \frac{1}{2} \left[ S_1, \left[ S_1, H' \right] - \frac{1}{3} \frac{\partial S_1}{\partial t} \right] + \ldots \quad (A-3)$$

The time derivative appears because of a time dependence of the chameleon and torsion fields. Then, following \[ 11 \]

we take the operator $S_1$ in the form

$$S_1 = -\frac{i}{2mA} \gamma^0 \left( D^j_i i \gamma^0 \gamma^j \frac{\partial}{\partial x^j} + F_j i \gamma^0 \gamma^j \frac{\partial}{\partial t} + G_j i \gamma^0 \gamma^j + K \gamma^5 \right).$$  \quad (A-4)

The time derivative of $S_1$ and the commutators in Eq. (A-3) are equal to

$$\frac{\partial S_1}{\partial t} = \frac{1}{2m} \frac{\partial}{\partial t} \left( \frac{D^j_i}{A} \right) \gamma^j \frac{\partial}{\partial x^j} + \frac{1}{2m} \frac{\partial}{\partial x^j} \left( \frac{F_j}{A} \right) \gamma^j \frac{\partial}{\partial t} + \frac{1}{2m} \frac{\partial}{\partial x^j} \left( \frac{G_j}{A} \right) \gamma^j - \frac{1}{2m} \frac{\partial}{\partial t} \left( \frac{K}{A} \right) i \gamma^0 \gamma^5$$  \quad (A-5)

and

$$i \left[ S_1, H' - \frac{1}{2} \frac{\partial S_1}{\partial t} \right] =$$

$$= -\frac{1}{2m} i \gamma^0 \gamma^j \frac{\partial}{\partial x^j} - F_j i \gamma^0 \gamma^j \frac{\partial}{\partial t} - G_j i \gamma^0 \gamma^j - K \gamma^5 - \frac{1}{2} \frac{i}{A} \gamma^0 \gamma^j \frac{D^j_i}{A} \frac{\partial A}{\partial x^j} - \frac{1}{2} \frac{i}{A} \gamma^0 \gamma^j \frac{F_j}{A} \frac{\partial A}{\partial x^j} + \frac{1}{2m} i \gamma^0 \gamma^j \frac{B}{A} \frac{\partial}{\partial x^j} + \frac{1}{2m} i \gamma^0 \gamma^j \frac{F_j}{A} \frac{\partial B}{\partial x^j} - \frac{1}{2m} \frac{\partial}{\partial x^j} \left( \frac{G_j}{A} \right) \gamma^j - \frac{1}{2m} \frac{\partial}{\partial t} \left( \frac{K}{A} \right) i \gamma^0 \gamma^5 + \frac{1}{2m} i \gamma^0 \gamma^j \frac{D^j_i}{A} \frac{\partial C_i}{\partial x^j} + \frac{1}{2m} i \gamma^0 \gamma^j \frac{F_j}{A} \frac{\partial C_i}{\partial x^j} i \gamma^5 + \frac{1}{2m} i \gamma^0 \gamma^j \frac{C_i}{A} \frac{\partial}{\partial x^j}$$

$$+ \frac{1}{m} \frac{i}{A} \gamma^5 \frac{\partial}{\partial x^j} \left( \frac{D^j_i}{A} \right) \gamma^j - \frac{1}{2m} \frac{i}{A} \gamma^0 \gamma^j \frac{F_j}{A} \frac{\partial C_i}{\partial x^j} i \gamma^0 \gamma^5 + \frac{1}{2m} i \gamma^0 \gamma^j \frac{C_i}{A} \frac{\partial}{\partial x^j}$$

$$+ \frac{1}{m} \frac{i}{A} \gamma^5 \frac{\partial}{\partial x^j} \left( \frac{D^j_i}{A} \right) \gamma^j - \frac{1}{2m} \frac{i}{A} \gamma^0 \gamma^j \frac{F_j}{A} \frac{\partial C_i}{\partial x^j} i \gamma^0 \gamma^5 + \frac{1}{2m} i \gamma^0 \gamma^j \frac{C_i}{A} \frac{\partial}{\partial x^j}$$
\[
\begin{align*}
&+ \frac{1}{mA} \partial_{\mu} \gamma_{k} F_{j} C_{i} \frac{\partial}{\partial t} + \frac{1}{mA} \partial_{\mu} \gamma_{k} G_{j} C_{i} + \frac{1}{mA} \gamma_{0} \eta^{j \bar{k}} D_{j}^{i} D_{k}^{\bar{k}} \frac{\partial^{2}}{\partial x^{j} \partial x^{k}} \\
&+ \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} D_{j}^{i} \frac{\partial D_{k}^{\bar{k}}}{\partial x^{j}} \frac{\partial}{\partial x^{k}} + \frac{1}{2mA} \gamma^{0} \eta^{j \bar{k}} D_{k}^{j} \frac{\partial D_{j}^{\bar{k}}}{\partial x^{k}} \frac{\partial}{\partial x^{j}} + \frac{1}{2mA} \gamma_{0} i \epsilon^{j \bar{k} \ell} \Sigma_{j} D_{j}^{i} \frac{\partial D_{k}^{\bar{k}}}{\partial x^{\ell}} \frac{\partial}{\partial x^{k}} \\
&- \frac{1}{2m} \gamma_{0} i \epsilon^{j \bar{k} \ell} \Sigma_{j} D_{k}^{\ell} \frac{\partial}{\partial x^{\ell}} \left( \frac{F_{j}^{i}}{A} \right) \frac{\partial}{\partial x^{j}} + \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} G_{j} D_{k}^{j} \frac{\partial}{\partial x^{k}} + \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} D_{k}^{j} \frac{\partial}{\partial x^{k}} \left( \frac{G_{j}^{i}}{A} \right) \\
&+ \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} D_{k}^{j} F_{k} \frac{\partial}{\partial x^{k}} + \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} D_{k}^{j} \frac{\partial F_{k}}{\partial x^{k}} \frac{\partial}{\partial x^{j}} + \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} F_{k} \frac{\partial}{\partial x^{k}} \frac{\partial}{\partial x^{j}} \\
&+ \frac{1}{2mA} \gamma_{0} i \epsilon^{j \bar{k} \ell} \Sigma_{j} F_{k} \frac{\partial}{\partial x^{k}} \frac{\partial}{\partial x^{j}} + \frac{1}{2mA} \gamma_{0} i \epsilon^{j \bar{k} \ell} \Sigma_{j} F_{k} \frac{\partial}{\partial x^{k}} \frac{\partial}{\partial x^{j}} + \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} R_{j} \frac{\partial}{\partial x^{k}} \frac{\partial}{\partial x^{j}} \\
&- \frac{1}{2m} \gamma_{0} i \epsilon^{j \bar{k} \ell} \Sigma_{j} F_{k} \frac{\partial}{\partial x^{k}} \frac{\partial}{\partial x^{j}} \left( \frac{G_{j}^{i}}{A} \right) + \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} K D_{k}^{j} \frac{\partial}{\partial x^{k}} + \frac{1}{2mA} i \gamma_{0} \Sigma_{k} D_{k}^{j} \frac{\partial}{\partial x^{j}} \left( \frac{K^{i}}{A} \right) \\
&+ \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} D_{j}^{i} F_{k} \frac{\partial}{\partial x^{k}} + \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} D_{k}^{j} F_{k} \frac{\partial}{\partial x^{j}} + \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} F_{k} \frac{\partial}{\partial x^{j}} \frac{\partial}{\partial x^{k}} \\
&+ \frac{1}{2mA} \gamma_{0} i \epsilon^{j \bar{k} \ell} \Sigma_{j} F_{k} \frac{\partial}{\partial x^{k}} \frac{\partial}{\partial x^{j}} + \frac{1}{2mA} \gamma_{0} i \epsilon^{j \bar{k} \ell} \Sigma_{j} F_{k} \frac{\partial}{\partial x^{k}} \frac{\partial}{\partial x^{j}} + \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} R_{j} \frac{\partial}{\partial x^{k}} \frac{\partial}{\partial x^{j}} \\
&- \frac{1}{2m} \gamma_{0} i \epsilon^{j \bar{k} \ell} \Sigma_{j} F_{k} \frac{\partial}{\partial x^{k}} \frac{\partial}{\partial x^{j}} \left( \frac{G_{j}^{i}}{A} \right) + \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} K D_{k}^{j} \frac{\partial}{\partial x^{k}} + \frac{1}{2mA} i \gamma_{0} \Sigma_{k} D_{k}^{j} \frac{\partial}{\partial x^{j}} \left( \frac{K^{i}}{A} \right) \\
&+ \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} D_{j}^{i} F_{k} \frac{\partial}{\partial x^{k}} + \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} D_{k}^{j} F_{k} \frac{\partial}{\partial x^{j}} + \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} F_{k} \frac{\partial}{\partial x^{j}} \frac{\partial}{\partial x^{k}} \\
&+ \frac{1}{2mA} \gamma_{0} i \epsilon^{j \bar{k} \ell} \Sigma_{j} F_{k} \frac{\partial}{\partial x^{k}} \frac{\partial}{\partial x^{j}} + \frac{1}{2mA} \gamma_{0} i \epsilon^{j \bar{k} \ell} \Sigma_{j} F_{k} \frac{\partial}{\partial x^{k}} \frac{\partial}{\partial x^{j}} + \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} R_{j} \frac{\partial}{\partial x^{k}} \frac{\partial}{\partial x^{j}} \\
&- \frac{1}{2m} \gamma_{0} i \epsilon^{j \bar{k} \ell} \Sigma_{j} F_{k} \frac{\partial}{\partial x^{k}} \frac{\partial}{\partial x^{j}} \left( \frac{G_{j}^{i}}{A} \right) + \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} K D_{k}^{j} \frac{\partial}{\partial x^{k}} + \frac{1}{2mA} i \gamma_{0} \Sigma_{k} D_{k}^{j} \frac{\partial}{\partial x^{j}} \left( \frac{K^{i}}{A} \right)
\end{align*}
\]

\[\text{and}\]

\[\frac{i^{2}}{2} \left[ S_{1}, [S_{1}, W - \frac{1}{3} \partial S_{1} / \partial t] \right] = \]

\[= - \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} D_{j}^{i} D_{k}^{\bar{k}} \frac{\partial^{2}}{\partial x^{j} \partial x^{k}} - \frac{1}{4mA} \gamma_{0} \eta^{j \bar{k}} D_{j}^{i} \frac{\partial D_{k}^{\bar{k}}}{\partial x^{j}} \frac{\partial}{\partial x^{k}} - \frac{1}{4mA} \gamma_{0} \eta^{j \bar{k}} D_{k}^{j} \frac{\partial D_{j}^{\bar{k}}}{\partial x^{k}} \frac{\partial}{\partial x^{j}} - \frac{1}{2mA} \gamma_{0} \eta^{j \bar{k}} D_{j}^{i} \frac{\partial^{2}}{\partial x^{j} \partial x^{k}}
\]
\[ -\frac{1}{2mA} \gamma^0 \eta^j D_j F_k \frac{\partial^2}{\partial \mathbf{x}^2} - \frac{1}{4mA} \gamma^0 \eta^j D_j \frac{\partial F_k}{\partial \mathbf{x}} \frac{\partial}{\partial t} - \frac{1}{4mA} \gamma^0 \eta^j F_k \frac{\partial (D_j^A)}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} - \frac{1}{4mA} \gamma^0 i \epsilon^{jkl} \Sigma_i D_j \frac{\partial F_k}{\partial \mathbf{x}} \frac{\partial}{\partial t} \]

\[ + \frac{1}{4mA} \gamma^0 \eta^j D_j \frac{\partial}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} - \frac{1}{2mA} \gamma^0 \eta^j F_j F_k \frac{\partial^2}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} - \frac{1}{4mA} \gamma^0 \eta^j F_j \frac{\partial D_k}{\partial \mathbf{x}} \frac{\partial}{\partial t} + \frac{1}{4mA} \gamma^0 i \epsilon^{jkl} \Sigma_i D_j \frac{\partial F_k}{\partial \mathbf{x}} \frac{\partial}{\partial t} - \frac{1}{2mA} \gamma^0 \eta^j G_j \frac{\partial}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} \]

\[ - \frac{1}{4mA} \gamma^0 \eta^j F_k \frac{\partial}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} - \frac{1}{2mA} \gamma^0 \eta^j D_j \frac{\partial G_k}{\partial \mathbf{x}} \frac{\partial}{\partial t} - \frac{1}{4mA} \gamma^0 i \epsilon^{jkl} \Sigma_j D_j \frac{\partial G_k}{\partial \mathbf{x}} \frac{\partial}{\partial t} - \frac{1}{2mA} \gamma^0 \eta^j F_j \frac{\partial D_k}{\partial \mathbf{x}} \frac{\partial}{\partial t} - \frac{1}{2mA} \gamma^0 \eta^j G_j \frac{\partial}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} \]

\[ - \frac{1}{4mA} \gamma^0 \eta^j F_k \frac{\partial G_k}{\partial \mathbf{x}} \frac{\partial}{\partial t} - \frac{1}{2mA} \gamma^0 \eta^j D_j \frac{\partial G_k}{\partial \mathbf{x}} \frac{\partial}{\partial t} - \frac{1}{4mA} \gamma^0 i \epsilon^{jkl} \Sigma_j D_j \frac{\partial G_k}{\partial \mathbf{x}} \frac{\partial}{\partial t} - \frac{1}{2mA} \gamma^0 \eta^j F_j \frac{\partial D_k}{\partial \mathbf{x}} \frac{\partial}{\partial t} - \frac{1}{2mA} \gamma^0 \eta^j G_j \frac{\partial}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} \]

\[ - \frac{1}{8mA} \gamma^0 i \epsilon^{jkl} \Sigma_i D_j \frac{\partial G_k}{\partial \mathbf{x}} \frac{\partial}{\partial t} - \frac{1}{4mA} \gamma^0 \eta^j F_j \frac{\partial D_k}{\partial \mathbf{x}} \frac{\partial}{\partial t} - \frac{1}{8mA} \gamma^0 i \epsilon^{jkl} \Sigma_i D_j \frac{\partial G_k}{\partial \mathbf{x}} \frac{\partial}{\partial t} - \frac{1}{8mA} \gamma^0 i \epsilon^{jkl} \Sigma_i D_j \frac{\partial G_k}{\partial \mathbf{x}} \frac{\partial}{\partial t} - \frac{1}{8mA} \gamma^0 i \epsilon^{jkl} \Sigma_i D_j \frac{\partial G_k}{\partial \mathbf{x}} \frac{\partial}{\partial t} \]

\[ - \frac{1}{4mA} \gamma^0 \eta^j F_k \frac{\partial A}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} + \frac{1}{4mA} \gamma^0 \eta^j F_k \frac{\partial A}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} + \frac{1}{4mA} \gamma^0 \eta^j F_k \frac{\partial A}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} \]

\[ \text{(A-7)} \]

The Hamilton operator $H_1$ decomposes into two parts $H_1 = H_{1-even} + H_{1-odd}$, where the operators $H_{1-even}$ and $H_{1-odd}$ are equal to

\[ H_{1-even} = A \gamma^0 m + B + C^j \Sigma_j + D^j i \frac{\partial}{\partial \mathbf{x}^2} + \frac{1}{2mA} \gamma^0 \eta^j D_j \frac{\partial^2}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} \]

\[ + \frac{1}{4mA} \gamma^0 \eta^j D_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} + \frac{1}{4mA} \gamma^0 \eta^j D_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} + \frac{1}{2mA} \gamma^0 \eta^j F_j \frac{\partial^2}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} + \frac{1}{4mA} \gamma^0 i \epsilon^{jkl} \Sigma_j D_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} \]

\[ - \frac{1}{4mA} \gamma^0 i \epsilon^{jkl} \Sigma_j D_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} + \frac{1}{2mA} \gamma^0 \eta^j F_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} + \frac{1}{4mA} \gamma^0 \eta^j F_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} \]

\[ + \frac{1}{4mA} \gamma^0 \eta^j D_j \frac{\partial}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} + \frac{1}{2mA} \gamma^0 \eta^j D_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} + \frac{1}{4mA} \gamma^0 \eta^j F_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} \]

\[ + \frac{1}{4mA} \gamma^0 \eta^j F_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} + \frac{1}{4mA} \gamma^0 \eta^j D_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} + \frac{1}{4mA} \gamma^0 \eta^j F_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} \]

\[ - \frac{1}{4mA} \gamma^0 i \epsilon^{jkl} \Sigma_j D_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} + \frac{1}{4mA} \gamma^0 i \epsilon^{jkl} \Sigma_j D_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} \]

\[ - \frac{1}{4mA} \gamma^0 i \epsilon^{jkl} \Sigma_j D_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} + \frac{1}{4mA} \gamma^0 i \epsilon^{jkl} \Sigma_j D_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} + \frac{1}{4mA} \gamma^0 i \epsilon^{jkl} \Sigma_j D_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} + \frac{1}{4mA} \gamma^0 i \epsilon^{jkl} \Sigma_j D_j \frac{\partial D_k}{\partial \mathbf{x}^2} \frac{\partial}{\partial t} \]
For the calculation of the Hamilton operator $H$ where the operator 
\[
\hat{H}(1/m) = e_{\text{odd}}.
\]
In order to remove the odd contributions of order $O(1/m^2)$ we arrive at the Hamilton operator $H_2 = H_{2\text{even}} + H_{2\text{odd}}$ where
\[
\hat{H}_{2\text{even}} = \hat{H}_{1\text{even}} + \frac{1}{4mA^2} \gamma_{ij} D^j_i \frac{\partial A}{\partial x^j} - \frac{1}{4mA^2} \gamma_{lj} \frac{\partial L}{\partial x^l} - \frac{1}{2mA} \gamma_{lj} \frac{\partial L}{\partial x^j} - \frac{1}{2mA} \gamma_{lj} \frac{\partial L}{\partial x^j}.
\]
The contribution of the Hamilton operator $H_{2odd}$ we remove by the third unitary transformation of the wave function $\psi_2(x) = e^{-iS_3} \psi_3(x)$ and the Hamilton operator $H_3$:

$$H_3 = e^{+iS_3} H_2 e^{-iS_3} - i e^{iS_3} \frac{\partial}{\partial t} e^{-iS_3} = H_2 - \frac{\partial S_3}{\partial t} + i \left[ S_3, H_2 - \frac{\partial S_3}{\partial t} \right] + \ldots ,$$

(A-12)

where the operator $S_3$ is given by $S_3 = -(i/2m) \gamma_0 H_{2odd} = O(1/m^2)$. Neglecting the contributions of order $O(1/m^2)$ the low–energy approximation of the Dirac Hamilton operator for slow fermions is equal to $H_3 = H_{2even}$, which we denote $H_3 = H_{2even} = H_{FW}$. Thus, the Dirac equation in the low–energy approximation takes the form

$$i \frac{\partial \psi_3(x)}{\partial t} = H_{FW} \psi_3(x),$$

(A-13)

where $\psi_3(x) = e^{iS_4} e^{iS_3} \psi(x)$. Following the standard procedure [11] and multiplying the both sides of Eq.(A-13) by the matrix $(1 + \gamma_0)/2$ we arrive at the Schrödinger–Pauli equation

$$i \frac{\partial \Psi(t, \vec{r}, \vec{\sigma})}{\partial t} = \left( -\frac{1}{2m} \Delta + \Phi_{eff}(t, \vec{r}, \vec{\sigma}) \right) \Psi(t, \vec{r}),$$

(A-14)

where $\Psi(t, \vec{r}) = \frac{1 + \gamma_0}{2} \psi_3(x)$ is the large component of the slow Dirac fermion wave function and $\Phi_{eff}(t, \vec{r}, \vec{\sigma})$ is the effective low–energy potential.
where $\sigma_a = (-\sigma^\dagger)_a$ and $\sigma_a = -\sigma^a$. The effective potential Eq. (A-15) is the most general effective low-energy potential for slow Dirac fermions in the Einstein–Cartan gravity with torsion and chameleon, calculated to order $O(1/m)$.

We would like to note that the effective low-energy potential Eq. (A-15) is derived by using the Foldy–Wouthuysen method or the Foldy–Wouthuysen unitary transformations of wave functions of Dirac fermions with mass $m$, leading to a non–relativistic Hamilton operator, expressed in terms of even operators only in a form of the large fermion mass expansion in powers of $1/m$. It is known that such a method of reduction of the Dirac Hamilton operator to a form, containing only even operators, is not unique and there are some other methods of unitary transformations such as the Eriksen method [47] and others, which were well discussed by de Vries [48]. In section III we give a short comparison of the Foldy–Wouthuysen with the Eriksen one only, since other methods of unitary transformations of a Dirac Hamilton operator to a form, containing only even operators, seem to be cumbersome when compared to the Eriksen method [48]. We discuss also an accuracy of the Foldy–Wouthuysen representation of a Dirac Hamilton operator.

[1] H. Abele, T. Jenke, H. Leeb, and J. Schmiedmayer, Phys. Rev. D 81, 065019 (2010).
[2] T. Jenke, P. Geltenbort, H. Lemmel, H. Abele, Nature Physics 7, 468 (2011).
[3] H. Abele, T. Jenke, D. Studer, and P. Geltenbort, Nucl. Phys. A 827, 593c (2009).
[4] T. Jenke, D. Studler, H. Abele, and P. Geltenbort, Nucl. Instr. and Meth. in Physics Res. A 611, 318 (2009).
[5] H. Abele and H. Leeb, New J. Phys. 14, 055010 (2012).
[6] T. Jenke, G. Cronenberg, J. Bürdinger, L. A. Chizhova, P. Geltenbort, A. N. Ivanov, T. Lauer, T. Lins, S. Rotter, H. Saul, U. Schmidt, and H. Abele, Phys. Rev. Lett. 112, 115105 (2014).
[7] H. Lemmel, Ph. Brax, A. N. Ivanov, T. Jenke, G. Pignon, M. Pitschmann, T. Potocar, M. Wellenzohn, M. Zawisky, and H. Abele, Phys. Lett. B 743, 310 (2015).
[8] Ph. Brax and G. Pignon, Phys. Rev. Lett. 107, 111301 (2011).
[9] A. N. Ivanov, R. Höllwieser, T. Jenke, M. Wellenzohn, and H. Abele, Phys. Rev. D 87, 105013 (2013).
[10] A. N. Ivanov and M. Pitschmann, Phys. Rev. D 90, 045040 (2014) and references therein.
[11] A. N. Ivanov and M. Wellenzohn, Phys. Rev. D 91, 085025 (2015).
[12] C. Lämmerzahl, Phys. Lett. A 228, 223 (1997).
[13] Yu. N. Obukhov, A. J. Silenko, and O. V. Teryaev, Phys. Rev. D 90, 124068 (2014) and references therein.
[14] V. A. Kostelecky, N. Russell, and J. D. Tasson, Phys. Rev. Lett. 100, 111102 (2008).
[15] A. N. Ivanov and M. Wellenzohn, Phys. Rev. D 92, 065006 (2015); arXiv: 1509.04014 [gr-qc].
[16] R. Lehnert, W. M. Snow, and H. Yan, Phys. Lett. B 730, 353 (2014); Erratum Phys. Lett. B 744, 415 (2015).
[17] J. Khoury and A. Weltman, Phys. Rev. Lett. 93, 171104 (2004); Phys. Rev. D 69, 044026 (2004).
[18] D. F. Mota and D. J. Shaw, Phys. Rev. D 75, 063501 (2007); Phys. Rev. Lett. 97, 151102 (2007).
[19] C. M. Will, in THEORY AND EXPERIMENT IN GRAVITATIONAL PHYSICS, Cambridge University Press, Cambridge 1993.
[20] I. Zlatev, L. Wang, and P. J. Steinhardt, Phys. Rev. Lett. 82, 896 (1999); P. J. Steinhardt, L. Wang, and I. Zlatev, Phys. Rev. D 59, 123504 (1999).
[21] S. Tsujikawa, Class. Quantum Grav. 30, 214003 (2013).
[22] S. Perlmutter et al., Bull. Am. Astron. Soc. 29, 1351 (1997).
[23] A. G. Riess et al., Astron. J. 116, 1009 (1998).
[24] S. Perlmutter et al., Astron. J. 517, 565 (1999).
[25] E. J. Copeland, M. Sami, S. Tsujikawa, Int. J. Mod. Phys. D 15, 1753 (2006).
[26] J. A. Frieman, M. S. Turner, and D. Huterer, Annu. Rev. Astro. Astrophys. 46, 385 (2008).
[27] Ph. Brax and A.-C. Davis, Phys. Rev. D 91, 063503 (2015) and references therein.
[28] G. Pignol, Int. J. Mod. Phys. A 30, 1530048 (2015).
[29] F. W. Hehl, J. D. McRea, E. W. Mielke, and Y. Ne’eman, Phys. Rep. 258, 1 (1995) and references therein.
[30] R. T. Hammond, Rep. Prog. Phys. 65, 599 (2002) and references therein.
[31] L. D. Landau and E. M. Lifshitz, in LEHRBUCH DER THEORETISCHEN PHYSIK, Band II: KLASSISCHE FELDTHEORIE, Verlag Harri Deutsch, Thun und Frankfurt am Main, 2008.
[32] D. K. Atwood, M. A. Horne, C. G. Shull, and J. Arthur, Phys. Rev. Lett. 52, 1673 (1984).
[33] B. Mashhoun, Phys. Rev. Lett. 61, 2639 (1988).
[34] H. Rauch and S. A. Werner, in NEUTRON INTERFEROMETRY: Lessons in Experimental Quantum Mechanics, Second Edition, Oxford University Press, Oxford, 2015.
[35] F. W. Hehl and W.-T. Ni, Phys. Rev. D 42, 2045 (1990) and references therein.
[36] Yu. N. Obukhov, A. J. Silenko, and O. V. Teryaev, Phys. Rev. D 80, 064044 (2009).
[37] Yu. N. Obukhov, A. J. Silenko, and O. V. Teryaev, Phys. Rev. D 84, 024045 (2011).
[38] K. A. Olive et al. (Particle Data Group), Chin. Phys. A 38, 090001 (2014).
[39] V. A. Kostelecky, Phys. Rev. D 69, 105009 (2004).
[40] C. Itzykson and J. Zuber, in Quantum Field Theory, McGraw–Hill, New York 1980.
[41] L. L. Foldy and S. A. Wouthuysen, Phys. Rev. 78, 29 (1950).
[42] M. G. Sagnac, C. R. Acad. Sci. 157, 708, 1410 (1913).
[43] A. A. Michelson, H. G. Gale, and F. Pearson, Astrophys. J. 61, 140 (1925).
[44] L. A. Page, Phys. Rev. Lett. 35, 543 (1975).
[45] S. A. Werner, J.–L. Staudenmann, and R. Colella, Phys. Rev. Lett. 42, 1103 (1979).
[46] D. Puetzfeld and Yu. N. Obukhov, Int. J. Mod. Phys. D 23, 1442004 (2014).
[47] E. Eriksen, Phys. Rev. 111, 1011 (1958).
[48] E. de Vries, Forschr. Phys. 18, 149 (1970).
[49] A. J. Silenko, J. Math. Phys. 44, 2952 (2003).
[50] V. P. Neznamov and A. J. Silenko, J. of Math. Phys. 50, 122302 (2009).
[51] A. J. Silenko, Theor. Math. Phys. 176, 987 (2013).
[52] A. J. Silenko, Phys. Rev. A 91, 012111 (2015).
[53] A. J. Silenko, Phys. Rev. A 91, 022103 (2015).
[54] E. Eriksen and M. Kolsrud, Nuovo Cim. Suppl. 18, 1 (1960).