Broadband Spectral Analysis Algorithm with High-Frequency Resolution for Elimination of Overlap Interference between Adjacent Channels

Xianhai Wang 1,2, Teng Wang 1,*, Chuan Yin 3, Jun Han 4, Qiao Meng 5 and Chen Wang 4

1 School of Electronic Engineering, Changzhou College of Information Technology, Changzhou 213164, China; wangxianhai@czcit.edu.cn
2 Changzhou Industrial Internet Research Institute Co., Ltd., Changzhou 213164, China
3 College of Electronics and Information, Hangzhou Dianzi University, Hangzhou 310018, China; yinc@hdu.edu.cn
4 The National Astronomical Observatories, Chinese Academy of Sciences, Beijing 100101, China; hanjun@nao.cas.cn (J.H.); wangchen@nao.cas.cn (C.W.)
5 Institute of RF-OE-ICs, Southeast University, Nanjing 210096, China; mengqiao@seu.edu.cn
* Correspondence: wangteng@czcit.edu.cn

Abstract: Spectral lines can be analysed to determine the physical properties of molecular clouds and the astrochemical processes in the formation area of massive stars. To improve the observation technology of radio astronomy, this paper proposes and compares two spectral analysis algorithms (improved weighted overlap-add (IWOLA) + FFT and IWOLA + weighted overlap-add (WOLA)). The proposed algorithms can obtain an ultra-high-frequency resolution for real-valued wideband signals, eliminate the signal overlapping interference between adjacent channels, substantially decrease the required hardware resources, especially multipliers, adders, and memory resources, and reduce the system design complexity. The IWOLA + FFT algorithm consists of an improved weighted overlap-add (IWOLA) filter bank, fast Fourier transform (FFT), a specific decimation for the output data from the IWOLA filter bank, and a selection for part of the output data from the FFT. The IWOLA + WOLA algorithm consists of the same modules as the IWOLA + FFT algorithm, with the second-stage FFT replaced by the modules of the weighted overlap-add (WOLA) filter bank and the accumulation for each sub-band. Based on an analysis of the underlying principles and characteristics of both algorithms, the IWOLA + FFT algorithm demonstrated a spectrum with a high frequency resolution and a comparable performance to an ultra-large-scale FFT, based on two smaller FFTs and a flexible architecture instead of a ultra-large-scale FFT. The IWOLA + WOLA algorithm contains the same function as the IWOLA + FFT algorithm and demonstrates a higher performance. The proposed algorithms eliminated the interference between the adjacent channels within the entire Nyquist frequency bandwidth. The simulation results verify the accuracy and spectral analysis performances of the proposed algorithms.

Keywords: channelisation; polyphase filter bank; WOLA filter bank; IWOLA filter bank

1. Introduction

The observation of spectral lines can be used to accurately study the physical properties of molecular clouds in the formation area of massive stars, e.g., the temperature, density, radiation field, and astrochemical processes [1–5]. Therefore, the observation of spectral lines is critical in radio astronomy. During radio-astronomical observations, digital backends are generally required to have a wide bandwidth and high-frequency resolution [6–9]. To trace smaller-scale structures and more detailed dynamic characteristics, spectral line observations with high-frequency resolutions are conducted to extract information from the target celestial bodies. A frequency resolution of several hertz or 1 Hz is required in numerous high-precision spectral line observations. For example, during the
Zeeman effect observation, if a neutral HI is observed, the frequency shift generated by each micro-Gauss magnetic field is 2.8 Hz. This requires a digital backend with a frequency resolution that should be higher than 2.8 Hz [10]. In addition, when observing the spectral lines, especially in the high-frequency band, the broadband receivers can simultaneously receive more information to improve the observation efficiency.

The requirements of the wideband and high-frequency resolution make it difficult to obtain the digital backends for radio-astronomical observations. At present, most signal processing techniques in the digital backends, such as spectral analysis, are implemented with a field programmable gate array (FPGA). Alternatively, according to the task requirements, signal pre-processing is first achieved in the FPGA, and further processing is performed by a computer with graphics processing units [11–13]. The maximum number of points of the available fast Fourier transform (FFT) intellectual property (IP) cores in the FPGAs used for the spectral analysis is 65536. If a resolution of 1 Hz is required for a signal with a 1 MHz bandwidth, a 106-point plain FFT is required, which necessitates a specialized design with a significant workload. The operating frequency of an actual FPGA is generally several hundred MHz. This indicates that a signal with an ultra-wide bandwidth greater than several hundred MHz cannot be directly processed by a simple FFT implemented on FPGA. Therefore, the direct application of a large FFT IP core in an FPGA to the ultra-large-scale FFT processing of wideband signals cannot be achieved to obtain an ultra-high-frequency resolution. Two methods are mainly used in radio astronomy to achieve this objective. A potential method involves the use of two-stage FFTs [14]. A commonly used method involves the combination of a polyphase filter bank (PFB) (sometimes called a WOLA filter bank) and an FFT or cascaded PFB structure [15–20]. For example, the combination regime is used as a spectrum analyser for the search for extra-terrestrial intelligence (SETI). It has a frequency resolution of 91 kHz when the processing bandwidth is 1.5 GHz, and a maximum frequency of 1 Hz when the processing bandwidth is 150 MHz [7,19,21]. In comparison with the two-stage FFT method, the aforementioned combined structure improves the frequency response characteristics for the first-stage channelised processing and the computational efficiency. This paper proposes a modified combination of the PFB + FFT to obtain improved flexibility with a more simplified design, and to conserve hardware resources in the FPGA. In addition, this paper proposes an IWOLA + WOLA algorithm to demonstrate an improved performance.

In this paper, a spectral analysis algorithm with an ultra-high-frequency resolution is first proposed for broadband real-valued signals. This algorithm uses an improved WOLA (IWOLA) [22] channelised filter bank to channelise the broadband signals preliminarily, and decimates the output data from the IWOLA filter bank appropriately. The IWOLA algorithm applies an equivalent Hilbert transform to the normal WOLA filter bank by shifting the centre frequency of every sub-band by half of the frequency bin, such that the IWOLA filter bank provides K/2 independently output complex sub-bands for real-valued signals instead of the general K/2 + 1 sub-bands, thus reducing the subsequent processing units by one set. After the decimation, FFT processing for the decimated data is carried out to perform spectral analysis with an ultra-high-frequency resolution. Furthermore, by selecting only a portion of the output data from the FFT, the interference between the adjacent channels can be eliminated within the entire Nyquist frequency bandwidth. To obtain a higher channelised performance, an optimized algorithm is finally proposed, which uses the IWOLA filter bank and the second-stage WOLA filter bank, instead of the second-stage FFT.

The remainder of this paper is organised as follows. In Section 2, the underlying principle and architecture of the proposed IWOLA + FFT algorithm are presented. In Section 3, the characteristics of the IWOLA filter bank and IWOLA + FFT algorithms are detailed, including the channelisation characteristics, spectra performance, and elimination of the overlapping interference between the adjacent channels. To improve performance of the IWOLA + FFT algorithm, the IWOLA + WOLA algorithm is proposed, and its performance analysis results are presented in Section 4. The simulation of the proposed
algorithms and analysis results are presented in Section 5, followed by the conclusions in Section 6.

2. Basic Algorithm Principle

The WOLA algorithm was originally used in the short-time Fourier transform (STFT) [23], and the WOLA filter bank has been widely used in many fields such as sub-band coding, audio processing, and radio astronomy [24–27]. The output of the WOLA filter bank involves real- and complex-valued signals, and the number of the independent output sub-channels is $K/2 + 1$ (where $K$ is the number of sub-bands); thus, the processing complexity and consumption of hardware resources is increased. To simplify the hardware implementation of the algorithm, the IWOLA algorithm [22] was proposed, which shifts the centre frequencies of every sub-band of the WOLA filter bank by $\pi/K$ along the positive frequency axis. In particular, the expression of the IWOLA filter bank is as follows:

$$y(\nu, k) = e^{-j\pi(2k + 1)\nu D/K} \sum_{p=0}^{K-1} F_{\nu}(p)e^{-j2\pi kp/K}, \tag{1}$$

where

$$R_{\nu}(p) = \sum_{l=0}^{D-1} (-1)^l x(\nu D + lK + p)w(-lK - p), \tag{2}$$

$$F_{\nu}(p) = R_{\nu}(p)e^{-j\pi p/K}, \tag{3}$$

where $l$ is an arbitrary integer that ranges from 0–$D$-1; $k$ denotes the channel number; $D$ denotes the decimation factor; $p = 0, 1, 2, \ldots, K-1$; and $w$ is a window function with coefficients that can be obtained by the design command fir1 in MATLAB for a finite impulse response (FIR) low-pass filter. The cut-off frequency of the FIR low-pass filter is $f_s/2K$.

This process of shifting frequency is equivalent to a Hilbert transform applied to the WOLA filter bank, thus causing all independent output sub-channels to exhibit identical features. The identities of the sub-bands of the IWOLA filter bank provide simplification for the hardware implementation of the processing. Furthermore, for a real-valued input signal, the output data from the $K/2 + 1$ sub-channels of the WOLA filter bank should be stored and processed further, whereas the output data from the $K/2$ sub-channels of the IWOLA filter bank should be stored and processed further. This reduces the amount of data processing and conserves off-chip memory resources, especially for a large second-stage FFT.

The overall framework of the spectral analysis algorithm based on the combined IWOLA + FFT structure is shown in Figure 1.
The output data of Sub-bands 0–K/2–1 of the IWOLA filter bank are transposed and obtained in-turn as outputs. If \( K = D \), a maximally decimated filter bank (MDFB) \[28\] is produced, which has a relatively simple design and consumes less hardware resources. However, it leads to an aliasing effect (overlapping interference), as shown in Figure 2a. To avoid the aliasing effect, \( D \leq K/2 \) is required, which results in a non-maximally decimated filter bank (NMDFB), as shown in Figure 2b. Generally, for a moderately large value of \( K \), \( D \) is set as \( K/2 \). However, when \( K \) is significantly large, e.g., \( K = 2^{20}, D = 2^{19} \). This leads to a low-pass filter prototype \( w \) with an order of magnitude of \( 2^{39} \), which consumes a significantly large amount of hardware resources. In this case, assuming that \( D < K/2 \) (e.g., \( K = 2^{20} \) and \( D = 2^{18} \) or smaller), the maximum order number of the low-pass filter prototype is \( 2^{38} \), which conserves half of the hardware resources, especially the multipliers and adders, at the expense of a slight performance reduction (slight decrease in the stopband attenuation). Under the condition of \( D < K/2 \), the data rate of each sub-band of the IWOLA filter bank is larger than the bandwidth value of each sub-band. Thus, if an FFT is implemented directly after the IWOLA filter bank, a large amount of computing resources are wasted. Therefore, before the FFT process, the data of each sub-band of the IWOLA filter bank should be decimated by a factor \( D_1 \), as shown in Figure 1. The decimation factor \( D_1 \) can be expressed as follows:

\[
D_1 = \frac{K}{2D}
\]

To make the decimating factor \( D_1 \) a positive integer, \( D < K/2 \) and \( K \) is divisible by \( D \). After the decimation by \( D_1 \), the data rate of each sub-band decreases to a value twice its bandwidth, which is critical for solving the problem of interference from the neighbouring sub-channels. This leads to the maximal conservation of the computing resources, given that two moderate FFTs are used instead of an ultra-large FFT. The overlapping interferences between the adjacent sub-bands can be eliminated by appropriately selecting part of the output data from the \( N_1 \)-point FFT.

After the application of \( N_1 \)-point FFT to the decimated data of Sub-bands 0–K/2, the resulting frequency resolution \( df \) for each sub-band can be expressed as follows:

\[
df = \frac{f_s}{D/D_1/N_1} = \frac{2f_s}{K/N_1}
\]

To eliminate the interference between the adjacent sub-bands, a part of the FFT bins of each sub-band is selected for the bin combination. The method of bin selection and
the combination is shown in Table 1. The number of selected FFT bins $N_2$ for Sub-bands $0-K/2-1$ can be expressed as follows:

$$N_2 = f_s / K / (f_s / D / D_1 / N_1) = D \times D_1 \times N_1 / K$$

(6)

Table 1. Bin selection and sorting diagram. It should be noted that $f_s$ is a sampling frequency. A signal within the bandwidth of $f_s/2$ is first split into $K$ sub-bands. Each of the first $K/2$ sub-bands are divided into $N_1$ FFT bins. Sub-bands $0-K/2-1$ represent complex signals. Consequently, the number of selected FFT bins for Sub-bands $0-K/2-1$ are all $N_2$. By selecting the appropriate bins from the $N_1$ FFT bins for each sub-band, the frequency-domain data for the real-valued signals within $f_s/2$ are obtained.

| Bandwidth = $f_s/2$ |
|---------------------|
| Sub-Bands 0–K/2–1 (Complex Signal) |
| 0 | 1 | … | $N_2/2$–1 | $N_1$–1 | $N_1$–2 | … | $N_1$–$N_2/2$ |

It is necessary for $N_2$ to be an integer, which can be readily obtained. Sub-bands $0-K/2-1$ are all complex signals, which represent the entire frequency range of real-valued signals.

As shown in Table 1, the bins from $0-N_2/2-1$ and $N_1$–$N_2/2$ to $N_1$–1 are selected separately from the $N_1$ FFT bins for each of the sub-bands from $0-K/2-1$. The selected bins $N_1$–$N_2/2$ to $N_1$–1 from each of the sub-bands from $0-K/2-1$, respectively, are reversed and re-ordered. The selected bins in Sub-bands $0-K/2-1$ are combined into one group along the order of the channel numbering. This represents the frequency-domain data for the real-valued signals within $f_s/2$, and the signal power can then be calculated.

Based on the analysis presented above, the algorithm splits an ultra-large-scale FFT into a two-stage process, in which two smaller-scale FFTs are required. The $N$-point FFT requires a magnitude order of $O(N \log_2 N)$ operations. Therefore, if spectral processing with an ultra-high-frequency resolution ($N$ is extremely large) is implemented in the FPGA, the algorithm reduces the operations to the magnitude order of $O(K \log_2 K) + O(N_1 \log_2 N_1)$ ($K$ and $N_1 << N$). This conserves hardware resources and compensates for the lack of a commercially available enormous point FFT IP core.

As mentioned before, the WOLA filter bank is equivalent to the PFB. They are both efficient implementation of the channelization processing. Therefore, their computational and hardware complexity is equivalent. In comparison with the expressions of the IWOLA filter bank and the WOLA filter bank [22], the expression of the IWOLA filter bank involves two terms more: $(-1)^l$ and $e^{-j \pi p / K}$. It is obvious that the first term does not increase the computational and hardware complexity; the second term only necessitates one complex multiplication and hardware complexity; two block RAMs (BRAMs) in the FPGA. The capacity of the two extra BRAMs is $K \times 16$ bit, given that each data of the second term is expressed with 16 bits in the FPGA. The function of the data decimation is to select or discard data, which does not increase the computational and hardware complexity.

3. Analysis of the IWOLA + FFT Algorithm Characteristics

3.1. Subsection

The channel characteristics of the IWOLA filter bank can be designed based on the task requirements. By selecting different values of the decimating factor $D$ of the IWOLA filter bank, the channel frequency responses of the IWOLA filter bank with different performances can be obtained, as shown in Figure 3. To simplify the description, $K$ and $D$ were set as small values. When $D = 16$, the overlap between the adjacent channels was significantly less than $D = 8$, and the pass-band gain in the sub-bands was flatter than that in the case of $D = 8$. Therefore, the performance of the IWOLA filter bank was improved
with an increase in $D$. However, it required more hardware resources. During the actual applications, the designs can be traded off according to specific needs.

![Frequency response of channels obtained from IWOLA filter bank (K=32, D=8)](image)

![Frequency response of channels obtained from IWOLA filter bank (K=32, D=16)](image)

**Figure 3.** Frequency response of the channels in the IWOLA filter bank (where $D$ denotes a decimating factor for the IWOLA filter bank).

Under the same conditions as the IWOLA filter bank, the frequency responses of the WOLA filter bank are presented in Figure 4. The frequency response for $D = 16$ was superior to that for $D = 8$, given that the former exhibited a flatter channel gain and narrower overlaps between neighbouring sub-channels. However, more hardware resources were consumed. In comparison, the frequency response characteristic of the IWOLA filter bank was identical to that of the WOLA filter bank, with the exception of shift in the centre frequency of each sub-band by $\pi/K$. 

![Image of frequency response for WOLA filter bank](image)
3.2. Elimination of the Overlap Interference between the Adjacent Subchannels

After the decimation by $D_1$, the output data rate for each subchannel in the IWOLA filter bank was twice the value of the subchannel bandwidth. Therefore, after the $N_1$-point FFT, only the FFT bins that correspond to the actual frequency of each sub-band should be selected. Figure 5 presents the actual bandwidth and sampling frequency of each sub-band after the decimation. The actual sampling frequency of each sub-channel $f_{s1}$ was equal to $f_s / (D \times D_1)$. The shaded parts represent the actual bandwidth $B$ of each sub-band, $f_2$ is the cut-off frequency of the frequency response characteristic of the aforementioned window function, and $f_1$ is a mirror frequency of $f_2$ relative to $f_{s1}/2$. The number of selected bins for Sub-bands $0$–$K/2$–$1$ $N_2$ was obtained from (6), $B = 2 \times f_2 = N_2 \times df$, instead of $f_{s1}$, and $f_1 = f_{s1} - f_2$. For Sub-bands $0$–$K/2$–$1$, their signals were complex. Therefore, it was necessary to select $N_2 N_1$-FFT bins, in which $N_2/2 N_1$-FFT bins, which were numbered from zero to the point corresponding to $f_2$, were sequentially selected; and the other $N_2/2$ bins were sequentially selected from the point corresponding to $f_1$ to the point corresponding to $f_{s1}$. Thereafter, a combination of the two parts of the bins can only cover the bandwidth of the complex signals in Sub-bands $0$–$K/2$–$1$. 

**Figure 4.** Frequency response of the channels in the WOLA filter bank (where D denotes a decimating factor for the WOLA filter bank).
Throughout the process of the proposed algorithm, the part of the $N_1$-point FFT bins that does not correspond to the actual bandwidth of each sub-band is discarded. This eliminates the signal overlap interference between the adjacent sub-bands. Moreover, the decimation of $D_1$ reduces the operating frequency for further processing; thus ensuring the utilization of the computing resources.

### 3.3. Performance Comparison of the IWOLA + FFT Algorithm with a Direct FFT

An assessment of the scalloping losses of the IWOLA filter bank, proposed algorithm, and direct FFT is presented below. First, we compared the scalloping loss of the IWOLA filter bank with that of the direct FFT by sweeping the frequency domain with 11 sinusoidal signals from one bin centre (e.g., the first bin) to an adjacent bin centre (the second bin). Thereafter, the maximum magnitude was plotted as shown in Figure 6. The analysis results revealed that the IWOLA filter bank exhibited a gain depression of approximately 0.36 near the border of the adjacent sub-channels. The adverse influence on the IWOLA + FFT algorithm is detailed further in this manuscript, and can be minimized by changing the characteristics of the window function for the IWOLA filter bank. We scanned the frequency characteristics from one bin centre to an adjacent bin centre using 11 sinusoidal signals. The improved result is plotted in Figure 7, which indicates a flatter gain curve of the IWOLA filter bank. This implies that the gain loss problem caused by channelisation was solved, and the scalloping loss of the IWOLA + FFT structure only depends on the succeeding FFT. Furthermore, under the assumption that the frequency resolution of the IWOLA + FFT algorithm is the same as that of the direct FFT, we analysed the scalloping losses using a similar approach to Figure 6. We then plotted the normalized magnitude, as shown in Figure 8. As can be seen from the figure, the proposed algorithm demonstrated the same scalloping loss characteristics as the direct FFT.
Figure 7. Scalloping loss of the IWOLA filter bank and the direct FFT. The results revealed that the IWOLA filter bank exhibited a flatter gain feature than the FFT.

Figure 8. Scalloping loss of the IWOLA + FFT algorithm and the direct FFT.

As can be observed from Figure 3, for the channelisation characteristics of the IWOLA filter bank, there was a gain depression near the border between the two adjacent channels due to the channelisation process. Therefore, it had an adverse influence on the characteristics of the proposed algorithm. We obtained the channelised gain flatness characteristics, as shown in Figure 9, by sweeping the frequency domain with 31 sinusoidal signals from the centre of one sub-channel to that of the neighbouring sub-channel and plotting the normalized magnitude. A lower amplitude gain was caused by the IWOLA filter bank near the border of the adjacent sub-channels. With the proposed algorithm, this problem was solved by changing the characteristics of the window function for the IWOLA filter bank, which did not influence the adjacent sub-bands due to the NMDFB. The characteristics of the window function can be improved by increasing its cut-off frequency by 1.4 times. We analysed the frequency channel gain flatness using a similar approach to Figure 9. We then plotted the normalized magnitude, as shown in Figure 10, which indicates that Gain depression near the border of the adjacent sub-channels was improved by changing the characteristics of the window function for the IWOLA filter bank. The improved frequency channel gain flatness characteristics were the same as those of the direct FFT. There was a gain fluctuation of approximately 0.32.
The flat channel characteristics of the IWOLA filter bank ensure that the IWOLA + FFT algorithm features the same spectral leakage as the direct FFT algorithm. A sinusoidal signal with a frequency of 45 MHz was inputted to the two algorithms, where \( K = 1024, \ N_1 = 2048, \ f_s = 800 \text{ MHz}, \ D = 8, \) and the frequency resolution was approximately 763 Hz. We then plotted the spectral leakage characteristics, as shown in Figure 11.

Under the assumption of streaming FFT processing, it is common knowledge that direct FFT demonstrates an excellent transient response. However, the IWOLA filter bank features a poor transient response, given that a long time period is required for the window function module to fill up data. The transient response of the IWOLA filter bank based on the input of a sinusoidal signal is plotted in Figure 12. As can be observed from the figure, the IWOLA filter bank output reached stability after the sixteenth frame (approximately greater than the decimation factor by a factor of two). This resulted in a decrease in resolution in the time-domain. However, the transient response time is not an issue for many radio astronomy applications, since a general accumulation process can be applied in radio astronomy to increase the signal-to-noise ratio (SNR) at the expense of a decrease in the time resolution.
The proposed algorithm uses the small-scale IWOLA filter bank and FFT instead of the ultra-large FFT calculation to perform spectral analysis on the broadband signals with an ultra-high-frequency resolution. As observed from the analysis results presented in the previous sections, the designed IWOLA filter bank in the IWOLA + FFT algorithm demonstrates flat channel characteristics. In addition, the proposed algorithm enhances the gain near the border of the two neighbouring sub-bands, with a channel characteristic that is identical to that of the second-stage direct FFT. To improve the channel characteristics, the second-stage FFT can be substituted with a WOLA filter bank, which is readily implementable.

4. Optimization of the IWOLA + FFT Algorithm
4.1. Architecture of the IWOLA + WOLA Algorithm

Based on the analysis results, the IWOLA + FFT algorithm can be used to perform ultra-large FFT calculations, with a nearly identical channel performance with respect to the ultra-large FFT, thus achieving spectral analysis with an ultra-fine frequency resolution. For improved channel characteristics, this paper proposes an IWOLA + WOLA algorithm by substituting the second-stage FFT with the WOLA filter bank to optimize the IWOLA + FFT algorithm. The overall framework of the IWOLA + WOLA algorithm is shown in Figure 13.
The three upper modules of the IWOLA + WOLA algorithm in Figure 13 are the same as those of the IWOLA + FFT algorithm. The WOLA filter bank outputs are the time-frequency spectrum for each subchannel ranging from 0–K/2. Therefore, the output data of the WOLA filter bank should be accumulated for linear frequency modulation (LFM) signals if the LFM signal is modulated within a long time-period. Otherwise the bandwidth of the LFM signals may be reduced after processing by the IWOLA + WOLA algorithm. In addition, the data accumulation for each sub-channel can improve the SNR. Before the output, the data of each sub-channel from the module of the power calculation and accumulation should then be re-ordered, given that the input data into the succeeding WOLA filter bank are complex. The re-ordering method is presented in Table 2. Furthermore, the re-ordered data should be partially selected to represent the entire sub-channel due to the NMDFB, as shown in Table 3. The number of selected data is K1/2. The selected data can then be combined into one group according to the order of the channel numbering.

Table 2. Bin reordering diagram for each sub-band (where \( K_1 \) is the bin number of the second-stage WOLA filter bank, and all data feeding into the WOLA filter bank are complex).

| Bandwidth = \( f_s/(D_k \times D_1) \) | Each Sub-Band (Complex Signal) |
|--------------------------------------|-------------------------------|
| Before re-order 0 1 ... K1/2 − 1 K1/2 K1/2 + 1 ... K1 − 1 | After re-order 0 K1 − 1 ... K1/2 + 1 K1/2 K1/2 − 1 ... 1 |

Table 3. Bin selection diagram for each sub-band (where \( K_1 \) is the bin number of the second-stage WOLA filter bank, and all data feeding into the WOLA filter bank are complex. The number of selected data for each sub-channel is K1/2).

| Bandwidth = \( f_s/(D_k \times D_1) \) | Each Sub-Band (Complex Signal) |
|--------------------------------------|-------------------------------|
| Before data selection 0 1 ... K1 − 1 | After data selection K1/2 − K1/4 K1/2 − K1/4 + 1 ... K1/2 K1/2 + 1 ... K1/2 + K1/4 − 1 |

4.2. Performance Comparison between the IWOLA + WOLA and IWOLA + FFT Algorithms

To demonstrate the advantages of the IWOLA+WOLA algorithm, a signal with a frequency of 45 MHz was inputted into the IWOLA + WOLA structure, IWOLA + FFT structure, where \( K = 1024, K_1 = 2048, f_s = 800 \) MHz, \( D = 8 \), and the frequency resolution was approximately 763 Hz. The spectral leakage characteristics were then plotted, as shown in Figure 14. As can be observed from the figure, the IWOLA + WOLA algorithm
featured a spectral leakage less than that of the IWOLA + FFT algorithm by a factor of approximately 2.1.

Figure 14. Comparison of the spectral leakage.

To analyse the SNRs of the IWOLA + WOLA and IWOLA + FFT algorithms, an input LFM signal with a frequency of 160 MHz, bandwidth of 80 MHz, and SNR of 30 dB was inputted to both algorithms, where $K = 1024$, $K_1 = 2048$, $f_s = 800$ MHz, $D = 8$, and the frequency resolution was approximately 763 Hz. The SNR performance was plotted as shown in Figure 15. As can be observed from the figure, the IWOLA + WOLA structure demonstrated an SNR of 31.27 dB, which was superior to that of the IWOLA + FFT structure (24.65 dB).

Figure 15. The SNR comparison. The SNR for the IWOLA + WOLA algorithm was 6.62 dB higher than that of the IWOLA + FFT algorithm.

Furthermore, we analysed the channelised gain flatness characteristics of the IWOLA + WOLA and the IWOLA + FFT algorithms by scanning the spectrum from one sub-channel centre to an adjacent sub-channel centre using 31 sinusoidal signals. The channelised gain flatness characteristics are shown in Figure 16. The IWOLA + WOLA algorithm demonstrated a fluctuation range of approximately 0.12, whereas the IWOLA + FFT algorithm demonstrated a fluctuation range of approximately 0.32. Thus, the IWOLA + WOLA algorithm featured a fluctuation range that was smaller than that of the IWOLA + FFT algorithm.
Based on the analysis results above, although the IWOLA + WOLA algorithm uses a smaller two-stage FFT with high computation efficiency instead of an ultra-large FFT, which is similar to the IWOLA + FFT algorithm; the IWOLA + WOLA demonstrated a superior performance to the IWOLA + FFT algorithm and direct FFT algorithm.

5. Algorithm Simulation Analysis

To verify the accuracy and performance of the proposed broadband spectral analysis algorithms with a high-frequency resolution, the following simulation analyses were carried out in MATLAB (the version: R2017a). The following parameters were set according to Equations (4)–(6): $K = 1024$, $D = 8$, $N_1 = 2048$, $K_1 = 2048$, $D_2 = 8$; and the sampling frequency of the real-valued signal $f_s = 800$ MHz, $D_1 = 64$, $df = 763$ Hz, and $N_2 = 1024$. To analyse the algorithm performances, three types of methods were used to calculate the spectrum for the simulated signals. The first involved the direct performance of an ultra-large FFT (direct FFT) to obtain a spectral resolution of 763 Hz. The remaining two methods involved the use of the IWOLA + FFT and IWOLA + WOLA algorithms to achieve the same spectral resolution.

First, a real sinusoidal signal with a frequency of 40.00 MHz was inputted to the three algorithms, to evaluate their gains and SNRs, as shown in Figure 17. As can be observed from the figure, the IWOLA + FFT structure exhibited a gain increase of approximately 3 dB when compared with the direct FFT, and a nearly identical SNR. The IWOLA + WOLA structure exhibited a gain decrease of approximately 5 dB when compared with the IWOLA + FFT structure. However, this may not be significant, given that the SNR of the IWOLA + WOLA structure was 6 dB higher than that of the IWOLA + FFT structure. Thus, the gain of the IWOLA + WOLA can be any value, which can then be multiplied by a constant to meet the design requirements only if its SNR is equal or superior to that of the IWOLA + FFT algorithm. Furthermore, to further verify the accuracy of the proposed algorithms, we simultaneously inputted four signals into the three algorithms. The input signal consisted of two real sinusoidal signals with frequencies of 40.00 MHz and 40.000763 MHz, respectively, and two chirp signals, in addition to Gaussian white noise. The starting frequencies of the two chirp signals were 80 MHz and 160 MHz, respectively. Moreover, their bandwidths were 0.78125 MHz, which is equivalent to the subchannel bandwidth of the IWOLA filter bank, and 80 MHz, respectively. The spectral data obtained by the three methods are shown in Figure 18, where (a) presents a spectrum obtained by the direct FFT algorithm, (b) presents a spectrum obtained by the IWOLA + FFT algorithm, and (c) presents a spectrum obtained by the IWOLA + WOLA algorithm. There were two real sinusoidal signals at A, and one chirp signal at B, which are detailed further in this manuscript. As can be observed from the figure, the proposed algorithms demonstrated flat
gain characteristics. The bandwidth of the second chirp signal occupies multiple sub-bands, but its gain between sub-bands are flat and identical.

Figure 17. Spectrum gain and SNR for the three algorithms (for a sinusoidal signal with a frequency of 40 MHz).

Figure 18. Comparison of the simulation analysis.
There were two real sinusoidal signals at A, with frequencies of 40 MHz and 40.000763 MHz. The magnified views for Point A in Figure 18 are shown in Figure 19. When comparing Figure 19a–c, the resultant spectra obtained by the direct FFT and IWOLA + FFT algorithms for the signal at Point A were the same, whereas the IWOLA + WOLA algorithm demonstrated a superior spectrum. The two proposed algorithms demonstrated the same frequency resolution as the direct FFT algorithm, namely, 763 Hz. The magnified views for Point B in Figure 18 are shown in Figure 20. The chirp signal frequency started at 80 MHz with a bandwidth of 0.78125 MHz, which was equal to each sub-band width of the IWOLA filter bank.

In addition, a coherent dispersed pulse was generated in MATLAB as a simulated pulsar signal, with a dispersion measure (DM) of 26.776 pc cm$^{-3}$. The frequency ranged from 1200–1600 MHz, and its sampling frequency and observation time were 800 MHz and 35 ms, respectively. The simulated pulsar signal with a white Gaussian noise (SNR = 0 dB) is shown in Figure 21. By setting the decimation factor $D = 8$, the number of channels was set as $K = 32$, and the bin number $K_1$ and the decimation factor $D_2$ of the second-stage WOLA filter bank were 2048 and 8, respectively. The simulated pulsar signal was inputted to the proposed IWOLA + WOLA algorithm. This yielded the spectrum shown in Figure 22 with a frequency resolution of approximately 24.4 kHz. As can be seen from the figure that Before dedispersion, there was dispersion smearing at the observed frequencies within an observation duration of 35 ms. The pulsar was dispersed within a range of 1200–1600 MHz. After a direct dedispersion algorithm, the dispersed signal was aligned to the arrival time, which corresponded to 1200 MHz. The dedispersed signal is displayed in Figure 23. It was subjected to a succeeding overlap-adding processing, which yielded the dedispersed time series of the simulated pulsar, as shown in Figure 24. As can be observed from the figure, the SNR of the pulsar signal output from the IWOLA + WOLA algorithm was 24.07 dB.
Figure 20. Comparison of the simulation analysis results (magnification of Point B).

(a) Spectrum of the direct FFT algorithm (magnification of Point B)

(b) Spectrum of using the IWOLA + FFT algorithm (magnification of Point B)

(c) Spectrum of the IWOLA + WOLA algorithm (magnification of Point B)

Figure 21. Coherent dispersed pulses with noise.
Figure 22. Spectrogram of the simulated pulsar obtained by the proposed algorithm (before dedispersion).

Figure 23. Spectrogram of the simulated pulsar obtained by the proposed algorithm (after dedispersion).

Figure 24. Dedispersed time series of the pulsar after overlap-adding by the channels (overlap-adding for all the 16,384 channels).

The simulation results revealed that the proposed IWOLA + WOLA algorithm demonstrated a superior accuracy and performance to the direct FFT and the IWOLA + FFT
algorithms. The proposed IWOLA + FFT and IWOLA + WOLA algorithms solved the overlapping interference problem of the adjacent subchannels caused by the first-stage channelisation that is realised by an IWOLA filter bank. By selecting the appropriate $K$, $D$, $D_2$, and $K_1$ values, in addition to the operating frequency for hardware such as FPGAs, a broadband spectral analysis system can be designed with a satisfactory performance that consumes less resources.

6. Conclusions

Two broadband spectral analysis algorithms with an ultra-high-frequency resolution, namely, the IWOLA + FFT algorithm and the IWOLA + WOLA algorithm, are proposed in this paper. The IWOLA + FFT algorithm was constructed by combining an IWOLA filter bank with FFT, in addition to a specific decimation for the output data from the IWOLA filter bank and the data selection for the output data from the FFT. The algorithm disassembles an ultra-large-scale FFT process into two-stage processes. Hence, two small-scale FFTs are required, which can achieve flexibility for the system implementation. This implies that if the algorithm is implemented with hardware such as an FPGA, it is not necessary to design an ultra-large point FFT. The algorithm reduces the subsequent processing units by one set for real-valued signals by shifting the centre frequency of every sub-band of the WOLA filter bank by half of the frequency bin, conserves hardware resources, and decreases the implementation complexity according to the identity of each sub-band of the IWOLA filter bank. The IWOLA + FFT algorithm demonstrates a comparable spectral analysis performance to that of the direct FFT algorithm and eliminates the overlap interference between the adjacent channels within the entire Nyquist frequency bandwidth. To improve the channelised performance, the IWOLA + WOLA algorithm was further proposed and analysed, which includes the second-stage WOLA filter bank and accumulation module instead of the second-stage FFT. The IWOLA + WOLA algorithm demonstrates the same function as the IWOLA + FFT algorithm, with a superior spectral analysis performance. By changing the algorithm parameters $K$, $D$, $D_2$, and $K_1$, the spectral analysis can be readily performed with an ultra-high-frequency resolution and satisfactory performance. Future research will be focused on design implementation and optimization in a FPGA, and the deployment of the proposed schemes at an observatory.

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