ERRATUM: “SEARCHING FOR PLANETS IN THE HYADES. IV. DIFFERENTIAL ABUNDANCE ANALYSIS OF HYADES DWARFS” (AJ, 125, 3185 [2003])

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In the bottom panel of Figure 5, the data from R. P. Kraft (ApJ, 142, 681 [1965]) (open squares) were incorrectly plotted. The corrected figure is included in this erratum. Because the fits shown in both the top and bottom panels of Figure 5 were to our data alone and not to the Kraft data, the adjustment of these data points does not affect the results of this paper. However, we direct the reader to D. R. Soderblom et al. (ApJS, 85, 315 [1993]) for a more in-depth analysis of the rotational velocity of the higher mass objects in the Hyades. We wish to thank D. Soderblom for pointing out this error to us.

![Figure 5](image-url)

**Fig. 5.—** Plot of $v \sin i$ of observed Hyades stars (including the same possible nonmembers, as discussed in Fig. 3). The symbols are the same as in Fig. 3. The solid curve is a polynomial fit to the upper envelope of our data (filled circles) as discussed in the text, the dashed curve is a power-law fit to our data only, and the dotted curve is an exponential fit to our data, as well. The top panel shows only our data, while the bottom panel adds in $v \sin i$ measurements from E. Böhm-Vitense et al. (ApJ, 569, 941 [2002]) (open triangles), W. Benz et al. (A&A, 138, 93 [1984]) (open stars), and selected dwarfs from R. P. Kraft (ApJ, 142, 681 [1965]) (open squares).