Inverted crossover resonance aiding laser cooling of $^{171}$Yb

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We observe an inverted crossover resonance in $\pi$-driven four-level systems, where $F' - F = 0, +1$. The signal is observed through saturated absorption spectroscopy of the $(6s^2)\, ^1S_0 - (6s6p)\, ^3P_1$ transition in $^{171}$Yb, where the nuclear spin $I = 1/2$. The enhanced absorption signal is used to generate a dispersive curve for 556 nm laser frequency stabilisation and the stabilised light cools $^{171}$Yb atoms in a two-stage magneto-optical trap, achieving temperatures of 20 $\mu$K. The Doppler-free spectroscopy scheme is further used to measure isotopic frequency shifts and hyperfine separations for the intercombination line in ytterbium.

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1. Introduction

Neutral ytterbium finds a niche in many ultra-cold atom explorations; examples include degenerate Fermi gases \cite{1,2}, artificial gauge potentials \cite{3,4}, quantum many body simulations \cite{5,6} and atomic clocks \cite{7,8}, along with the potential application in quantum computation \cite{9}. In the majority of instances the intercombination line ($^1S_0 - ^3P_1$) is used for laser cooling, typically in a magneto-optical trap. With a Doppler cooling limit of 4.5 $\mu$K the low temperatures are conducive to a range of applications; for example, loading optical lattice traps and from there producing all optical Bose-Einstein condensates \cite{10} or degenerate Fermi gases \cite{11,12}. With a natural line-width of 184 kHz the full potential of the laser cooling is only attainable if the frequency noise of the 556 nm light is sufficiently low. One solution is to use an optical cavity with a reasonably high finesse and low frequency drift to stabilise the 556 nm light frequency \cite{13,14}. An alternative is to lock the green light to an element of a well stabilised frequency comb \cite{15}. A further means is to lock to the center of a Doppler profile produced with a thermal atomic beam \cite{16,17}. This helps maintain long term stability of the green light; however, keeping the frequency excursions close to 180 kHz is challenging \cite{18}. Here we demonstrate a means of conferring short-to-long term frequency stability on 556 nm light without the need for an optical cavity or frequency comb lock. We use frequency-modulated spectroscopy where the 556 nm beam intersects a Yb atomic beam at normal incidence, and reflect the beam with a lens-mirror combination to create a pump-probe scheme for fluorescence spectroscopy. Unexpectedly, we observe enhanced absorption (rather than saturation) behaviour in saturated absorption spectroscopy (SAS) for both of the hyperfine lines in the $^1S_0 - ^3P_1$ transition of $^{171}$Yb. The effect occurs when a small DC magnetic field, of the order of a few hundredths of a millitesla, is applied across the interaction zone. The field strength is below that required to resolve the Zeeman sub-levels; though, is still apparent when the Zeeman sub-levels are separated. We have identified the enhanced absorption as an inverted crossover resonance between two principal transitions \cite{22,23}, where the principal transitions arise due to the Zeeman splitting of magnetic sub-states. The signal we observe is similar to predictions made by models of electromagnetically induced absorption (EIA) such as that by Taichenachev et al. \cite{24} and Goren et al. \cite{25}. However, the pump-probe scheme with linearly polarised light that is used in the experiment here does not permit a transfer of coherence as required for EIA.

Doppler-free spectroscopy in Yb for frequency stabilisation of 556 nm light has been reported previously, but without reference to detail \cite{3}. We also note that there is recent discussion about coherent population trapping (CPT) with the intercombination line in $^{174}$Yb \cite{26,27}. A characteristic that we share with these CPT investigations is that there is saturation of the transition by both pump and probe fields. Our observation of an inverted crossover resonance may be the first for a group II atom.

The instability of the 556 nm light frequency, locked to the crossover resonance, is assessed by use of an optical frequency comb whose mode spacing is referenced to a hydrogen maser. Further validation of the the green light’s frequency instability is performed by using the 556 nm light in a two-stage magneto-optical trap of $^{171}$Yb ($F' = 3/2$), producing temperatures below 20 $\mu$K with a standard deviation of 2 $\mu$K across nearly sixty sequential measurements ($F'$ represents the total angular momentum of the exited state).

Our setup also enables a measure of frequency differences between isotopic lines and hyperfine separations with statistical uncertainties of 60 kHz or less for most of the frequency intervals. These are determined via absolute optical frequency measurements \cite{28}. We provide a...
summary of these measurements and compare with previous reports. Relatively good agreement is found, though a number of frequency separations differ by more than two standard deviations. High-resolution spectroscopy on the Yb intercombination line has previously been performed via $V$-type double-resonance spectroscopy, but without quantitative frequency measurements [30].

This paper is organised as follows. Section 1 describes the experimental details for the saturated absorption spectroscopy (SAS) and frequency measurements. Sect. 2 presents the spectra generated from the saturated absorption spectroscopy and gives information pertaining to the enhanced absorption. Sect. 3 describes the two-stage laser cooling scheme of $^{171}\text{Yb}$ and presents results related to the stabilisation of the 556 nm light, and Sect. 4 summarises the isotopic frequency shift measurements carried out with the saturated absorption scheme.

2. Saturated absorption spectroscopy

The saturated absorption scheme comprises a retro-reflected 556 nm beam at right angles to the atomic beam and the 556 nm fluorescence is detected by a photomultiplier tube (Hamamatsu H10492-003). We take advantage of the Yb atomic beam leading to a magneto-optical trap (MOT) used for clock line spectroscopy [29]. Frequency modulation at 33 kHz is applied to the light via an acousto-optic modulator (AOM) allowing for third harmonic detection with a lock-in amplifier. Helmholtz coils in close vicinity of the saturated absorption zone are used to set the magnetic field in the vertical direction. For isotopic frequency shift measurements, the magnetic field is set to cancel the vertical component of the Earth’s $\mathbf{B}$-field (0.059 mT upward), and the polarisation of the 556 nm light is set vertically (i.e. parallel to that of $\mathbf{B}$ when it is non-zero). The zero $B$-field condition has been determined by several means; one of which is to rotate the light polarisation by 90° and observe the Zeeman splitting, which is stronger for the $\sigma$ transitions. The intensity profile of the 556 nm beam is elliptical with the major axis aligned with the atomic beam.

The temperature of the oven (Createc, GmbH) is maintained at 430°C producing an effusion of Yb atoms through a trio of stainless steel tubes with a diameter of 1 mm and length 20 mm. The green fluorescence passes through a spatial filtering scheme to minimise stray light.

We estimate the flow rate of Yb atoms through the 556 nm beams to be $\sim(3)\times10^{13}$ s$^{-1}$ based on molecular flow rates through a cylindrical tube [31] (an estimate based on fluorescence detection is an order of magnitude less). Based on a mean velocity of 320 m s$^{-1}$ and an intensity well above saturation, the number of absorption/emission cycles is about 15 per atom as the atoms traverse the pump and probe beams. The 556 nm light is produced via resonant frequency doubling of 1112 nm light, which is generated from a commercial fiber laser that injection-locks a diode laser [32]. A bow-tie configuration doubling cavity uses an 18 mm length periodically poled potassium titanyl phosphate crystal for the non-linear conversion and 75 mm radius of curvature focusing mirrors. The line-width of the 556 nm radiation has been found to be no greater than 410 kHz [32]. Based on the manufacturer specifications for the fiber laser, the line-width of the green light should be below 240 kHz — slightly larger than the natural line-width of the intercombination line.

For generating the spectra the 556 nm frequency is slowly swept via a piezo transducer in the fiber laser. We calibrate the frequency span by use of a frequency comb. There is 0.8 mW of 1112 nm light combined with
the nearest comb element on an avalanche photodiode to produce an RF beat signal with 29 dB signal-to-noise ratio (SNR) with 500 kHz resolution bandwidth. The beat signal is filtered with a 10 MHz wide (−3 dB) band pass filter centered at 30 MHz. The frequency comb, referenced to a hydrogen maser, is used to make frequency measurements for all of the Yb isotopic lines, except for $^{168}$Yb with 0.13 \% relative abundance and $^{173}$Yb ($F' = 3/2$) (Sect. 4).

$$g_F = g_I = - \frac{1}{I} \left( \frac{\mu_N}{\mu_B} \right) \left( \frac{\mu_I}{\mu_N} \right)$$

Fig. 2. (a) Third harmonic dispersive curves for saturated absorption spectroscopy of the $^1S_0 - ^3P_1$ line of $^{171}$Yb ($F' = 1/2$) for two values of magnetic field: (i) near zero field and (ii) $B = 0.39$ mT. A vertical offset is applied for clarity. The result in trace (ii) occurs for both directions of the applied $B$-field. (b) The integrated dispersive profiles of (a). The averaging time per data point was 0.2 s.

A comparison of the dispersive spectra for $^{171}$Yb ($F' = 1/2$) at two different magnetic field strengths is shown in Fig. 2. Trace (a) with zero $B$-field shows the usual saturated absorption feature (or Lamb dip), while at $\sim 0.39$ mT there is a strong inversion of the dip showing enhanced absorption. The enhancement is seen for both magnetic field directions and for both $^{171}$Yb hyperfine lines, however the effect is more pronounced with $^{171}$Yb ($F' = 1/2$). The strength of the enhanced absorption is reduced by about 30 \% in the case of $^{171}$Yb ($F' = 3/2$). The difference between the two traces shows that there is a substantial change in behaviour with the simple addition of a small DC magnetic field.

The branching ratios for the de-excitation are readily calculated for the $^{171}$Yb ($F' = 1/2$) and $^{171}$Yb ($F' = 3/2$) systems [33]. We illustrate these in Fig. 3. For $^{171}$Yb ($F' = 1/2$) the probability of the downward $\sigma$ transitions is 2/3 while for $^{171}$Yb ($F' = 3/2$) it is 1/3. This correlates with the effect being more pronounced for $F' = 1/2$. The factors relevant here are the natural line-width of the transition (Γ$^{556}$ = 184 kHz), the Zeeman shift of the $^3P_1$ m$_F$ levels (±14 MHz·m$^{-1}$) and the Zeeman shift of the ground state m$_F$ levels (±3.76 kHz·m$^{-1}$). The latter is evaluated with $\Delta \omega = g_F \mu_B m_F B/h$, where $g_F$ is the Landé $g$-factor and $\mu_B$ is the Bohr magneton. For the case of hyperfine states in an external magnetic field with $J = 0$ and $F = I$ the Landé $g$-factor simplifies to become

Enhanced absorption (or an inverted Lamb dip) has been observed previously in SAS; for example, in $^{133}$Cs [37] and $^{40}$Ca [38]. In the former case the enhanced absorption occurs with a change in $F$ or in crossover resonances. For example, the saturation behaviour depends strongly on the magnetic field for the $F = 2 \rightarrow F' = 2, 3$ crossover resonance. The level system in $^{171}$Yb is simpler with only four Zeeman sub-states involved. We may be reporting the first example of a crossover resonance where

\[ \text{where the first quotient is the ratio of nuclear and Bohr magnetons (hence, equal to the ratio of electron and proton masses) and the second quotient is the nuclear magnetic moment expressed in nuclear magnetons} \]

\[ \text{For} \mu_I/\mu_N \text{we take the value of 0.49367(1) from published tables} \]

\[ \text{The ground state Zeeman splitting is} 2.7 \times 10^{-4} \text{times weaker than the upper state splitting. This degree of splitting may not be relevant here, but is included as a matter of interest. The Rabi frequency for the pump is} \]

\[ \text{The enhanced absorption signal is best described as an inverted crossover resonance, where the principal transitions are those depicted in Fig. 3 and are driven with} \]

\[ \text{The modelling carried out by Vafafard et al.} \]
For $^{40}\text{Ca}$ the inverted dip occurs for a particular range of intensity levels and forms a triplet at high intensity. In our case the enhanced absorption is manifest for the 556 nm beam intensity range of 18 $I_S^{(556)}$ to 910 $I_S^{(556)}$, where $I_S^{(556)} = 1.4 \text{ W m}^{-2}$ is the saturation intensity for the $^1S_0 - ^3P_1$ transition (we have stated the combined intensity of the two beams). The only changes that occur to the line are an enhanced SNR with some line broadening, though both appear to become constant above about 150 $I_S^{(556)}$. Below 18 $I_S^{(556)}$ the signal is too weak to discern, so it is not clear what happens near saturation.

For the $^{171}\text{Yb}$ ($F' = 3/2$) hyperfine transition with a nullued magnetic field, a very small perturbation to the line is seen that we attribute to the $^{173}\text{Yb}$ ($F' = 3/2$) line. It is consistent with previous reports that the latter is within $3 \text{ MHz}$ of the $^{171}\text{Yb}$ ($F' = 3/2$) line and higher in frequency [39,40].

![Fig. 3. Branching ratios from the upper state Zeeman sub-levels in $^{171}\text{Yb}$ for the $F = \frac{1}{2} - F' = \frac{1}{2}$ and $F = \frac{1}{2} - F' = \frac{3}{2}$ transitions. For the excitation phase the 556 nm light polarisation is parallel to the magnetic field, implying $\Delta m_F = 0$. The separation of the upper and lower states is not drawn to scale. The sub-level separation rates are: 14 MHz mT$^{-1}$ for $F = \frac{1}{2} - F' = \frac{1}{2}$, 7 MHz mT$^{-1}$ for $F = \frac{1}{2} - F' = \frac{3}{2}$ and 3.76 kHz mT$^{-1}$ for the ground state.](image)

An inverted crossover resonance is again observed for $^{173}\text{Yb}$ ($F = \frac{5}{2} \rightarrow \frac{3}{2}$), however the contrast is significantly weaker than for $^{171}\text{Yb}$. The spin-$\frac{5}{2}$ system has a weaker Zeeman shift dependence (1.2 MHz mT$^{-1}$). Observations have been made for magnetic fields up to 1.2 mT without a high contrast resonance occurring. The higher spin system generates an abundance of transition pathways, which likely prevents the observation of a clear crossover resonance since the multi-level splitting encompasses a broader atomic velocity range.

A crossover resonance is also observed for $^{172}\text{Yb}$ (and by inference all the bosonic isotopes) when $E$ and $B$ are perpendicular, thus driving $\Delta m_F = \pm 1$ transitions. In this case the resonance exhibits reduced absorption (i.e., it is a normal Lamb dip). Here, the two atomic velocity groups share only two transition pathways.

The enhanced absorption signal of $^{171}\text{Yb}$ ($F' = 3/2$) is used to stabilise the frequency of the 556 nm light by locking to the center of the dispersion curve. The correction signal is sent to the piezo transducer in the 1112 nm master fiber laser with a servo bandwidth of a few Hertz. The instability of the 556 nm light frequency was assessed by counting the beat signal frequency generated by the heterodyning of 1112 nm light with an adjacent element of the frequency comb. As mentioned, the mode spacing (and offset frequency) of the comb were referenced to a hydrogen maser. The instability at 1 s integration time is approximately 70 kHz (for the carrier frequency corresponding to 556 nm) and integrates down as $\tau^{-1/2}$, reducing to 6 kHz after 100 s of integration time.

### 3. Laser Cooling

This stabilised 556 nm light is used in a two-stage laser cooling scheme, where atoms are first trapped with 399 nm light and a quadrupole magnetic field; then for 50 ms the violet light is switched off and the atoms are further cooled with the 556 nm light. The dual-MOT scheme is depicted in Fig. 1b. The 399 nm and 556 nm laser beams are combined with long-wavelength pass filters for each of the MOT beams, and achromatic quarter wave plates are used to set the handedness of the circular polarization (simultaneously for both wavelengths). Since the 399 nm light is responsible for the initial capturing of the Yb atoms, the 556 nm beam diameter in the MOT can be set small, so posing less need for high power (through, at the expense of more stringent alignment conditions). The 556 nm beam radius ($e^{-2}$) is 2.0 mm and approximately 9 mW in total from the six beams reaches the atoms.

The difference frequency of RF signals driving two acousto-optic modulators is used to set the frequency detuning of the 556 nm MOT light, where one AOM shifts the frequency of the light sent to the saturation absorption scheme and the other the MOT light. The transfer efficiency is found to be highest for a frequency detuning of $-6.1 \text{ MHz}$ or about $-33 \Gamma^{(556)}$.

Temperature measurements were performed by recording the free expansion of the atom cloud on a CCD; each successive image is generated through the reloading and release of the MOT [41] (the initial release time was 1 ms and incremented in 0.75 ms steps for a series of five frames, the CCD exposure time was 1 ms). The optical imaging and filtering scheme produces a magni-
fication of unity. The event sequence for the dual-colour MOT is summarised in Fig. 5 and Fig. 6 shows a succession of temperature and cloud radius measurements. Here the 399 nm light was switched off for 50 ms before the release of the remaining electromagnetic fields. Apart from the free expansion period, the magnetic quadrupole field was held fixed at 0.42 T·m⁻¹ (axial direction). This value was calculated based on the design of the MOT coils and current. We found that a reduction in B-field gradient for the 556 nm MOT loading did not aid in the transfer (further discussion below). The frequency detuning, δ, of the 556 nm light was at a −33 Γ(556) for the transfer and then swept to −17 Γ(556) in concurrence with a reduction in intensity to ∼200 I(556) S, from a maximum of 10³ I(556) S. Ramping down the intensity of the 556 nm light is found to have influence only when used in combination with a reduction in δ. The uncertainties given by the error bars are generated from the line fitting of r²rms versus t² produced by the free expansion of the atom cloud. The weighted mean temperature is 18.2 µK and the fluctuation in temperature, as represented by the standard deviation of the data set, is 2.0 µK. The low level of temperature variation from shot to shot (or cycle to cycle) is a good indication of the stability of the 556 nm laser frequency and hence the merit of locking to the inverted crossover resonance. Reducing the green light intensity closer to I(556) S has yet to produce lower atomic cloud temperatures (where the trapping force is still about ten times greater than that due to gravity). The 18 µK temperature measured here is comparable to that achieved in other Yb MOTs that use the intercombination line (e.g. Ref. [13] reports 15 µK). For reference, the effect of releasing the atoms from the light fields only and not the quadrupole B-field is indicated by the two triangular data points in Fig. 6.

We note that the 556 nm light is shut off via switching of the AOM’s RF drive. This leaves a residual 0.05 I(556) S of light intensity at the MOT. We confirmed that this did not influence the temperature measurements by use of a physical shutter (Cambridge Technology) to switch the 556 nm beams on and off [42]. No change in the temperature recordings was evident. The limitation preventing lower temperatures is still to be determined (as previously noted, the limit set by 1D Doppler cooling theory is 4.5 µK). A likely influence is the Zeeman slower beam, which remains on continuously and is only detuned by −7 Γ(399) from the center of the 1S₀ − 1P₁ transition. However, variations of the Zeeman beam intensity by 50 % (from ∼2.5 I(399) S to ∼5 I(399) S) show no discernible influence on the atomic cloud temperature. Based on ion pump current readings the chamber pressure was approximately 5×10⁻⁸ Pa (or 5×10⁻¹⁰ mbar). Remaining higher frequency noise on the green light is a possible reason given that we rely on the intrinsic linewidth of the 1112 nm master laser and servo bandwidth of the resonant frequency doubling cavity. The 556 nm MOT beams are also prone to very slight direction changes as the frequency detuning is reduced. We have also studied the temperature of the composite bosonic isotope ¹⁷⁴Yb with the same event sequence. Here there is no enhanced absorption signal in the SAS, so the laser stabilising is with respect to the more usual Lamb-dip signal. The combination of reducing δ and the 556 nm intensity before the free expansion is very pronounced. We observe a temperature reduction from 110 µK to 45 µK through this additional cooling phase, which is carried out over 15 ms.

Fig. 5. Event sequence for temperature measurements with the two-stage magneto-optical trap. I, intensity; D, detection period; ∆tᵣ, free expansion duration; Iₛ = 595 W m⁻² and 1.4 W m⁻² for the 1S₀ − 1P₁ and 1S₀ − 3P₁ transitions, respectively; δ represents the frequency detuning from line-center. Γ(399) = 2π × 28.9 × 10⁶ rad s⁻¹ and Γ(556) = 2π × 184 × 10⁵ rad s⁻¹. The additional cooling phase produced by the 556 nm light occurs over the period ∆tᵣ, set to 50 ms. ∆tᵣ and ∆tₜ are not drawn to scale.

Fig. 6. Temperature and MOT cloud radius recorded for ¹⁷¹Yb (F' = ½) using the inverted crossover resonance to stabilise the 556 nm light. The weighted mean temperature for this sequence is 18.2 ± 2.0 µK. The two triangular data points show the effect of not releasing the atoms from the magnetic quadrupole field for the free expansion period.
The cloud radius (root-mean squared value) in Fig. 6 gives a mean value of about ∼0.18 mm. From the fluorescence received by a photo-multiplier tube (or with the CCD, calibrated) we can infer the number of atoms in the MOT. The collection efficiency is about 0.5%, and the PMT sensitivity is 0.085 V nW⁻¹ for the choice of gain setting. Taking into account the frequency detuning and light intensity we estimate the number of $^{172}$Yb atoms to be $2 \times 10^9$, corresponding to a (peak) density of $3 \times 10^{10}$ cm⁻³. This number reduces by 40% in the transfer from the violet to green MOT and the number of $^{171}$Yb atoms trapped is ∼40% that of $^{172}$Yb. The low atom number is attributed to the low atomic flux rate from the oven, due to the conservative nozzle design. As the need arises this nozzle will be replaced such that the flow rate is enhanced. A higher flux rate is not anticipated to influence the atomic cloud temperature significantly (the present flow rate of remaining thermal atoms through the MOT cloud is roughly $1 \times 10^9$ s⁻¹).

In previous work the transfer of atoms to the 556 nm MOT is sometimes carried out with a reduction in the magnetic field gradient, $dB/dz$. In our case a reduction in $dB/dz$ reduces the transfer efficiency. The magnetic field gradient of 0.42 T m⁻¹ for the green MOT, at first, appears surprisingly high. We can estimate the maximum $B$-gradient by considering the adiabatic condition that $dB/dz < \hbar k T / 2m \mu$, where $k$ is the wave number, $m$ the mass of the atom, $v$ the velocity and $\mu$ is the Zeeman shift coefficient. For the $^1S_0 - ^3P_1$ transition in $^{171}$Yb (F' = 3/2), $\mu = 2\pi \times 0.7 \times 10^{10}$ Hz T⁻¹. With a first stage temperature of ∼300 μK [41] all the atoms can in principle be transferred with a capture velocity of about 0.6 m s⁻¹. The corresponding maximum $B$-field gradient is 1.0 T m⁻¹, safely above what we find in the experiment.

### 4. Isotopic shifts and hyperfine separations

The saturated absorption scheme provides a convenient means of measuring isotopic frequency shifts between Yb isotopes and the hyperfine separations within the fermionic isotopes. To our knowledge, sub-Doppler spectroscopy with quantitative measurements has not been carried out previously on the $^1S_0 - ^3P_1$ Yb lines. The frequency comb (referred to above) is used to measure absolute optical frequencies with respect to a hydrogen maser for all the lines. For each transition [apart from $^{173}$Yb (F = 5/2 → 3/2) and $^{168}$Yb] the 556 nm light is locked to the center of the saturated absorption dip and the frequency of the optical comb beat is recorded with a frequency counter, set to a gate time of 0.4 s, over a mean duration of 200 s. The Allan deviation typically reduces to ∼10 kHz with the accumulated data. Our reported values in Table 1 are the mean from six separate measurement runs taken over the course of two months. During a single run the absolute frequencies are measured within a two hour period, hence factors that could affect the absolute frequencies, such as slight beam direction changes, are made common to all the lines. We evaluate the frequency separations with respect to the $^{176}$Yb isotopic line to aid comparison with previous work. For convenience, in the right most column we list the frequency separations with respect to the lowest frequency line: $^{172}$Yb ($F = \frac{5}{2} \rightarrow \frac{3}{2}$).

These measurements were performed with a nulled magnetic field avoiding separation of the Zeeman components and the optical pumping behaviour described above. We note that for the $^{171}$Yb (F' = 3/2) line there is a ∼200 kHz shift between the center of the enhanced absorption line and the zero-$B$-field line, while for the $^{171}$Yb (F' = 1/2) line this shift is below 40 kHz.

Our quoted uncertainties are the statistical uncertainties ($\sigma / \sqrt{n - 1}$) produced over the $n = 6$ separate measurement trials, rather than an evaluation of systematic shifts. This is done because the largest uncertainty arises from identifying the line-center for the lock point, rather than from systematic shifts. Note, even in the case of line-center offsets these will be largely common between the different lines as the lock point on the discriminator is not changed during a single measurement run (though, prior to the frequency measurement, a scan of the line was in most cases carried out to confirm symmetry about the lock point). There will, however, be an associated shift due to the different line strengths, which we estimate to be below 30 kHz when considering the strongest and weakest lines. For the most abundant isotopes the gain of the PMT was reduced to help minimise this shift. A statistical uncertainty over a number of measurements should incorporate errors associated with identifying the line-center. Shifts due to the Stark effect or frequency counting errors are well below the uncertainties reported here (and would also be common to all the lines). In support, a 556 nm beam intensity increase by a factor of ten showed no measurable change to the (absolute) frequencies.

Some of the frequency separations differ by more than two standard deviations in comparison with previous measurements; notably the separations to the $^{173}$Yb (F' = 7/2), $^{171}$Yb (F' = 1/2) and $^{172}$Yb lines. In the case of $^{171}$Yb (F' = 3/2) we have better agreement with Ref. [43]. We do not have an explanation for these differences. As noted earlier, the $^{171}$Yb (F' = 3/2) line is neighboured with the $^{173}$Yb (F' = 3/2) line. The latter is only marginally visible with a +2.5 (3) MHz separation from the former.

### 5. Conclusions

We have identified an inverted crossover resonance in a di-valent atom by use of saturated absorption spectroscopy where the transitions are driven with linearly polarised retro-reflected beams. It occurs with the two hyperfine lines in neutral $^{171}$Yb when a small magnetic bias field is applied parallel to the light’s $E$-polarisation. The forward and reverse beams are overlapped and orthogonal to a thermal atomic beam. The quality factor of the crossover resonance is ∼2×10⁸. A small magnetic bias field cre-
ates a simple optical pumping scheme that produces enhanced absorption via Zeeman splitting of the $J = 0$ ground state. Branching to a ground state doublet produces multiple pathways for the atoms through the four states, creating a scheme free of saturation. The photon absorption only becomes important by the interaction time.

The $F = 1/2$ energy level structure of the $^{171}$Yb intercombination line is not common; two other examples are $^{199}$Hg and $^{111,113}$Cd, partly explaining why an inverted crossover resonance has rarely (if at all) been observed in divalent atoms. An inverted crossover resonance is also observed with $^{173}$Yb, where $I = 5/2$, but with a much poorer contrast and broader line shape. This is most likely because the multi-level splitting encompasses a broader velocity range, thus preventing the formation of a narrow resonance. Crossover resonances are further observed for the bosonic isotopes when the light’s $E$-polarisation is perpendicular to the magnetic field. In this case the resonance forms a normal saturation feature.

A dispersive signal derived from the enhanced absorption resonance is used to frequency stabilise 556 nm light and in so doing enable laser cooling of $^{171}$Yb atoms in a dual-stage magneto-optical trap. Atom-cloud temperatures of $\sim 20 \mu$K are produced with cycle-to-cycle variations of $2 \mu$K. This is achieved without the use of an optical reference cavity for high-frequency noise suppression. The saturated absorption scheme has also provided the highest resolution measurements to date of the isotopic shifts and hyperfine separations of the $(6s^2) \ ^1S_0 - (6s6\sigma)^3P_1$ transition in ytterbium.

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Table 1. Summary of the measured $^1S_0 - ^3P_1$ isotopic shifts and hyperfine splittings in ytterbium. Some previously reported values are included for comparison.

| Transition | This work | Ref. [40] | Ref. [43] | This work |
|------------|-----------|-----------|-----------|-----------|
| $^{173}$Yb ($^5_2 \rightarrow ^7_2$) | -1431.363 (37) | -1431.872 (60) | -1432.6 (12) | 0 |
| $^{171}$Yb ($^5_2 \rightarrow ^7_2$) | -1177.725 (33) | -1177.231 (60) | -1177.3 (11) | 254.637 (10) |
| $^{176}$Yb | 0 | 0 | 0 | 1431.363 (37) |
| $^{174}$Yb | 954.628 (26) | 954.832 (60) | 954.628 (9) | 2386.011 (39) |
| $^{172}$Yb | 1955.444 (24) | 1954.852 (60) | 1954.8 (16) | 3386.827 (38) |
| $^{170}$Yb | 3241.229 (58) | 3241.175 (44) | 3241.5 (28) | 4672.612 (35) |
| $^{173}$Yb ($^5_2 \rightarrow ^7_2$) | 3266.554 (61) | 3266.243 (60) | 3267.1 (28) | 4697.920 (35) |
| $^{168}$Yb | -4609.960 (80) | -4611.9 (44) | 4611.9 (44) | - |
| $^{173}$Yb ($^{1}_{2} \rightarrow ^{3}_{2}$) | 4761.175 (44) | 4759.44 (80) | 4761.8 (37) | 6192.416 (56) |
| $^{173}$Yb ($^{1}_{2} \rightarrow ^{3}_{2}$) | 4763.680 (300) | 4762.11 (120) | 4761.8 (37) | 6194.920 (300) |
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