Einstein’s Special Relativity:
The Hyperbolic Geometric Viewpoint

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May 2, 2009

Conference on Mathematics, Physics and Philosophy
on the Interpretations of Relativity, II,
Budapest, 4-6 September, 2009

Published in:
PIRT Conference Proc., 4-6 Sept. 2009, Budapest, pages 1–35.

ABSTRACT  The analytic hyperbolic geometric viewpoint of
Einstein’s special theory of relativity is presented. Owing to the
introduction of vectors into hyperbolic geometry, where they are
called *gyrovectors*, the use of analytic hyperbolic geometry extends
Einstein’s unfinished symphony significantly, elevating it to the sta-
tus of a mathematical theory that could be emulated to the benefit
of the entire mathematical and physical community. The resulting
theory involves a gyrovector space approach to hyperbolic geo-
metry and relativistic mechanics, and could be studied with profit by
anyone with a sufficient background in the common vector space
approach to Euclidean geometry and classical mechanics. Einstein
noted in his 1905 paper that founded the special theory of relativity
that his velocity addition law satisfies the law of velocity par-
allelogram only to a first approximation. Within our hyperbolic
geometric viewpoint of special relativity it becomes clear that Ein-
stein’s velocity addition law leads to a hyperbolic parallelogram
addition law of Einsteinian velocities, which is supported exper-
imentally by the cosmological effect known as stellar aberration
and its relativistic interpretation. The latter, in turn, is supported
experimentally by the “GP-B” gyroscope experiment developed
by NASA and Stanford University. Furthermore, the hyperbolic
viewpoint of special relativity meshes extraordinarily well with the
Minkowskian four-vector formalism of special relativity, revealing
that the seemingly notorious relativistic mass meshes up with the
four-vector formalism as well, owing to the natural emergence of
dark matter. It is therefore hoped that both special relativity and
its underlying analytic hyperbolic geometry will become part of the
lore learned by all undergraduate and graduate mathematics and
physics students.

Key words: Special Relativity, Hyperbolic Geometry, Einstein’s Velocity Addition
Law, Stellar Aberration, Dark Matter, Gyrogroups, Gyrovectors.
1. Introduction

It is a pleasure for me to be given this opportunity to participate in this Conference on Mathematics, Physics and Philosophy in the Interpretations of relativity, II, in order to present the novel hyperbolic geometric viewpoint of Einstein’s Special Theory of Relativity.

The hyperbolic geometry of Bolyai and Lobachevsky underlies relativistic physics just as Euclidean geometry underlies classical physics. As such, it suggests a rather radical break of the traditional study of Einstein’s special theory of relativity, required for the restoration of the glory and harmony of special relativity in the Twenty-First Century. The mathematical structure that Einstein’s velocity addition law encodes emerges in all its splendor, leading to a no-commutative-nonassociative algebraic setting for the Beltrami-Klein ball model of hyperbolic geometry.

Einstein’s velocity addition law is non-commutative (and non-associative, as well). Indeed, Einstein’s exposé of velocity composition for two inertial systems emphasizes the lack of symmetry in the formula for the direction of the relative composite velocity vector [8, pp. 905 – 906] [54, p. 117]. Émile Borel’s attempt to “repair” the seemingly “defective” Einstein velocity addition in the years following 1912 is described by Walter in [54, p. 117]: “Borel could construct a tetrahedron in kinematic space, and determined thereby both the direction and magnitude of relative [composite] velocity in a symmetric manner.”

The goal of this lecture is to demonstrate that Einstein’s velocity addition law possesses rich structure and, hence, should be placed centrally in special relativity theory along with the algebraic structure that it encodes. We will find that Einstein’s velocity addition law encodes a vector space-like structure, called an Einstein gyrovector space, which forms the setting for the Beltrami-Klein ball model of hyperbolic geometry just as vector spaces form the setting for the standard model of Euclidean geometry.

Accordingly, placing Einstein’s velocity addition law centrally in the basis of Einstein’s special theory of relativity amounts to the study of the theory by means of its underlying analytic hyperbolic geometry. The resulting hyperbolic geometric viewpoint of Einstein’s special theory of relativity is rewarding. It suggests, for instance, the introduction of vectors into hyperbolic geometry, where they are called gyrovectors, and the study of Einsteinian velocities as gyrovectors that add according to the gyroparallelogram addition law.

Indeed, in the years 1908 – 1914, the period which experienced a dramatic flowering of creativity in the special theory of relativity, the Croatian physicist and mathematician Vladimir Varičak (1865 – 1942), professor and rector of Zagreb University, showed that this theory has a natural interpretation in hyperbolic geometry [51, 2]. However, much to his chagrin, he had to admit in 1924 [52, p. 80] that the adaption of vector algebra for use in hyperbolic geometry was just not feasible, as Scott Walter notes in [54, p. 121]. Vladimir Varičak’s hyperbolic geometry program, cited by Pauli [31, p. 74], is described by Walter in [54, p. 112–115].

We will find in this lecture that Borel’s attempt to “repair” Einstein’s velocity addition law, and Varičak’s concern about the lack of vector algebra in hyperbolic geometry are both unjustified. Indeed, we will find that Einstein’s velocity addition law admits a gyrovector space structure, in which Einsteinian velocities are
gyrovectors that add commutatively according to the gyroparallelogram addition law. The prefix “gyro” that we use extensively to capture analogies stems from Thomas gyration, which is the mathematical abstraction of the special relativistic effect known as Thomas precession. In fact, it is the mere introduction of Thomas gyration that turns Euclidean geometry into hyperbolic geometry, and some results of classical mechanics into corresponding results in relativistic mechanics.

2. Einstein Velocity Addition

Let $c$ be any positive constant, let $(V, +, \cdot)$ be any real inner product space, and let

$$V_c = \{ v \in V : \| v \| < c \}$$

be the $c$-ball of all relativistically admissible velocities of material particles. It is the open ball of radius $c$, centered at the origin of the real inner product space $V$, consisting of all vectors $v$ in $V$ with magnitude $\| v \|$ smaller than $c$.

Einstein velocity addition in the $c$-ball of all relativistically admissible velocities is given by the equation [33, Eq. 2.9.2], [29, p. 55], [13], [44],

$$u \oplus v = \frac{1}{1 + \frac{u \cdot v}{c^2}} \left\{ u + \frac{1}{\gamma_u} v + \frac{\gamma_u}{1 + \gamma_u} (u \cdot v) u \right\}$$

satisfying the gamma identity

$$\gamma_{u \oplus v} = \gamma_u \gamma_v \left( 1 + \frac{u \cdot v}{c^2} \right)$$

for all $u, v \in V_c$, where $\gamma_u$ is the gamma factor

$$\gamma_v = \frac{1}{\sqrt{1 - \frac{\| v \|^2}{c^2}}}$$

in the $c$-ball $V_c$. Here $u \cdot v$ and $\| v \|$ represent the inner product and the norm that the ball $V_c$ inherits from its space $V$.

A frequently used identity that follows from (4) is

$$\frac{v^2}{c^2} = \frac{\gamma_v^2 - 1}{\gamma_v^2}$$

where we use the notation $v^2 = v \cdot v = \| v \|^2$.

In physical applications, $V = \mathbb{R}^3$ is the Euclidean 3-space, which is the space of all classical, Newtonian velocities, and $V_c = \mathbb{R}_c^3 \subset \mathbb{R}^3$ is the $c$-ball of $\mathbb{R}^3$ of all relativistically admissible, Einsteinian velocities. Furthermore, the constant $c$ represents in physical applications the vacuum speed of light.

Einstein addition (2) of relativistically admissible velocities was introduced by Einstein in his 1905 paper [8] [9, p. 141] that founded the special theory of relativity, where the magnitudes of the two sides of Einstein addition (2) are presented. One has to remember here that the Euclidean 3-vector algebra was not so widely known in 1905 and, consequently, was not used by Einstein. Einstein calculated in [8] the behavior of the velocity components parallel and orthogonal to the relative velocity between inertial systems, which is as close as one can get without vectors to the vectorial version (2).
We naturally use the abbreviation \( u \ominus v = u \oplus (-v) \) for Einstein subtraction, so that, for instance, \( v \ominus v = 0 \), \( \ominus v = 0 \ominus v = -v \) and, in particular,

\[
\ominus (u \oplus v) = \ominus u \ominus v
\]

and

\[
\ominus u \ominus (u \ominus v) = v
\]

for all \( u, v \) in the ball, in full analogy with vector addition and subtraction. Identity (6) is known as the *automorphic inverse property*, and Identity (7) is known as the *left cancellation law* of Einstein addition \([45, 47, 50]\). We may note that Einstein addition does not obey the immediate right counterpart of the left cancellation law (7) since, in general,

\[
(u \oplus v) \ominus v \neq u
\]

However, this seemingly lack of a right cancellation law will be repaired in (26), following the suggestive introduction of a second gyrogroup binary operation in Def. 3 below, which captures analogies.

In the Newtonian limit of large \( c \), \( c \to \infty \), the ball \( V_c \) expands to the whole of its space \( V \), as we see from (1), and Einstein addition \( \oplus \) in \( V_c \) reduces to the ordinary vector addition \( + \) in \( V \), as we see from (2) and (4).

Einstein addition is noncommutative \([4]\). While \( \|u \oplus v\| = \|v \oplus u\| \), we have, in general,

\[
u \oplus v \neq v \oplus u
\]

Moreover, Einstein addition is also nonassociative since, in general,

\[
(u \oplus v) \oplus w \neq u \oplus (v \oplus w)
\]

It seems that following the breakdown of commutativity and associativity in Einstein addition some mathematical regularity has been lost in the transition from Newton’s velocity vector addition in \( V \) to Einstein’s velocity addition \( \oplus \) in \( V_c \). This is, however, not the case since, as we will see in Sec. 3, Thomas gyration comes to the rescue \([14, 15, 17, 50, 55, 32]\). Indeed, we will find in Sec. 3 that the mere introduction of gyration endows the Einstein groupoid \( (V_c, \oplus) \) with a grouplike rich structure \([42]\) that we call a *gyrogroup*.

When the nonzero vectors \( u, v \in V_c \subset V \) are parallel in \( V \), \( u \parallel v \), that is, \( u = \lambda v \) for some \( 0 \neq \lambda \in \mathbb{R} \), Einstein addition reduces to the Einstein addition of parallel velocities \([56, p. 50]\),

\[
\|u \oplus v\| = \frac{1 + \frac{1}{c^2} \|u\| \|v\|}{1 + \frac{1}{c^2} \|u\| \|v\|} \|u\| \|v\|
\]

which was confirmed experimentally by the Fizeau’s 1851 experiment \([28]\). Owing to its simplicity, some books on special relativity present Einstein velocity addition in its restricted form (11) rather than its general form (2).

The restricted Einstein addition (11) is both commutative and associative. Accordingly, the restricted Einstein addition is a group operation, as Einstein noted in \([5]\); see \([3\ p. 142]\). In contrast, Einstein made no remark about group properties of his addition of velocities that need not be parallel. Indeed, the general Einstein
addition \([\mathrm{2}]\) is not a group operation but, rather, a gyrocommutative gyrogroup operation, a structure that was discovered more than 80 years later, in 1988 \([\mathrm{41}]\), and is presented in Def. \([\mathrm{1}]\) in Sec. \([\mathrm{3}]\).

3. **Thomas Gyration and Einstein Gyrogroups**

For any \(u, v \in \mathbb{V}_c\), let \(\text{gyr}[u, v] : \mathbb{V}_c \to \mathbb{V}_c\) be the self-map of \(\mathbb{V}_c\) given in terms of Einstein addition \(\oplus\) by the equation \([\mathrm{41}]\)

\[
\text{gyr}[u, v]w = \ominus(u \oplus v) \ominus \{u \oplus (v \oplus w)\}
\]

where \(\ominus v = -v\), for all \(w \in \mathbb{V}_c\). The self-map \(\text{gyr}[u, v]\) of \(\mathbb{V}_c\), which takes \(w \in \mathbb{V}_c\) into \(\ominus(u \oplus v) \ominus \{u \oplus (v \oplus w)\} \in \mathbb{V}_c\), is called the **Thomas gyration** generated by \(u \) and \(v\). Thomas gyration is the mathematical abstraction of the relativistic effect known as Thomas precession \([\mathrm{44}, \text{Chap. 1}], [\mathrm{47}, \text{Sec. 10.3}]\), and it has an interpretation in hyperbolic geometry \([\mathrm{53}]\) as the negative hyperbolic triangle defect \([\mathrm{47}, \text{Theorem 8.55}]\).

In the Newtonian limit, \(c \to \infty\), Einstein addition \(\oplus\) in \(\mathbb{V}_s\) reduces to the common vector addition \(+\) in \(\mathbb{V}\), which is associative. Accordingly, in this limit the gyration \(\text{gyr}[u, v]\) in \([\mathrm{12}]\) reduces to the identity map of \(\mathbb{V}\). Hence, as expected, Thomas gyration \(\text{gyr}[u, v]\), \(u, v \in \mathbb{V}_c\), vanish (that is, they become trivial) in the Newtonian limit.

It is clear from the gyration equation \([\mathrm{12}]\) that gyration measures the extent to which Einstein addition is nonassociative, where associativity corresponds to trivial gyration.

The gyration equation \([\mathrm{12}]\) can be manipulated (with the help of computer algebra) into the equation

\[
\text{gyr}[u, v]w = w + \frac{Au + Bv}{D}
\]

where

\[
A = -\frac{1}{c^2} \frac{\gamma_u^2}{(\gamma_u + 1)}(\gamma_v - 1)(u \cdot w) + \frac{1}{c^2} \gamma_u \gamma_v (v \cdot w)
\]

\[
+ \frac{2}{c^4} \frac{\gamma^2_{u\gamma_v}}{(\gamma_u + 1)(\gamma_v + 1)}(u \cdot v)(v \cdot w)
\]

\[
B = -\frac{1}{c^2} \frac{\gamma_v}{\gamma_v + 1}(\gamma_u (\gamma_v + 1)(u \cdot w) + (\gamma_u - 1) \gamma_v (v \cdot w))
\]

\[
D = \gamma_u \gamma_v (1 + \frac{u \cdot v}{c^2}) + 1 = \gamma_{u\gamma_v} + 1 > 1
\]

for all \(u, v, w \in \mathbb{V}_c\). Allowing \(w \in \mathbb{V} \supset \mathbb{V}_c\) in \([\mathrm{13}]\) \(- [\mathrm{14}]\), gyration \(\text{gyr}[u, v]\) are expendable to linear maps of \(\mathbb{V}\) for all \(u, v \in \mathbb{V}_c\).
In each of the three special cases when (i) \( u = 0 \), or (ii) \( v = 0 \), or (iii) \( u \) and \( v \) are parallel in \( V \), \( u \parallel v \), we have \( Au + Bv = 0 \) so that \( \text{gyr}[u, v] \) is trivial,

\[
\begin{align*}
\text{gyr}[0, v]w &= w \\
\text{gyr}[u, 0]w &= w \\
\text{gyr}[u, v]w &= w, \quad u \parallel v
\end{align*}
\]

for all \( u, v \in V_c \), and all \( w \in V \).

It follows from (13) that

\[
\text{gyr}[v, u](\text{gyr}[u, v]w) = w
\]

for all \( u, v \in V_c \), \( w \in V \), so that gyrations are invertible linear maps of \( V \), the inverse of \( \text{gyr}[u, v] \) being \( \text{gyr}[v, u] \) for all \( u, v \in V_c \).

Gyrations keep the inner product of elements of the ball \( V_c \) invariant, that is,

\[
\text{gyr}[u, v]a - \text{gyr}[u, v]b = a - b
\]

for all \( a, b, u, v \in V_c \). Hence, \( \text{gyr}[u, v] \) is an isometry of \( V_c \), keeping the norm of elements of the ball \( V_c \) invariant,

\[
\|\text{gyr}[u, v]w\| = \|w\|
\]

Accordingly, \( \text{gyr}[u, v] \) represents a rotation of the ball \( V_c \) about its origin for any \( u, v \in V_c \).

The invertible self-map \( \text{gyr}[u, v] \) of \( V_c \) respects Einstein addition in \( V_c \),

\[
\text{gyr}[u, v](a \oplus b) = \text{gyr}[u, v]a \oplus \text{gyr}[u, v]b
\]

for all \( a, b, u, v \in V_c \), so that \( \text{gyr}[u, v] \) is an automorphism of the Einstein groupoid \((V_c, \oplus)\). We recall that an automorphism of a groupoid \((V_c, \oplus)\) is a bijective self-map of the groupoid \( V_c \) that respects its binary operation, that is, it satisfies (19).

Under bijection composition the automorphisms of a groupoid \((V_c, \oplus)\) form a group known as the automorphism group, and denoted \( \text{Aut}(V_c, \oplus) \). Being special automorphisms, Thomas gyrations \( \text{gyr}[u, v] \in \text{Aut}(V_c, \oplus) \), \( u, v \in V_c \), are also called gyroautomorphisms, \( \text{gyr} \) being the gyroautomorphism generator called the gyrator.

The gyroautomorphisms \( \text{gyr}[u, v] \) regulate Einstein addition in the ball \( V_c \), giving rise to the following nonassociative algebraic laws that “repair” the breakdown of commutativity and associativity in Einstein addition:

\[
\begin{align*}
\text{u} \oplus \text{v} &= \text{gyr}[\text{u}, \text{v}](\text{v} \ominus \text{u}) & \text{Gyrocommutativity} \\
\text{u} \ominus (\text{v} \ominus \text{w}) &= (\text{u} \ominus \text{v}) \ominus \text{gyr}[\text{u}, \text{v}]\text{w} & \text{Left Gyroassociativity} \\
(\text{u} \ominus \text{v}) \ominus \text{w} &= \text{u} \ominus (\text{v} \ominus \text{gyr}[\text{v}, \text{u}]\text{w}) & \text{Right Gyroassociativity}
\end{align*}
\]

for all \( u, v, w \in V_c \).

Owing to the gyrocommutative law in (20), Thomas gyration is recognized as the familiar Thomas precession. The gyrocommutative law was already known to Silberstein in 1914 [36] in the following sense. The Thomas precession generated by \( u, v \in \mathbb{R}^3 \) is the unique rotation that takes \( v \ominus u \) into \( u \ominus v \) about an axis perpendicular to the plane of \( u \) and \( v \) through an angle \( < \pi \) in \( V \), thus giving rise to the gyrocommutative law. Obviously, Silberstein did not use the terms “Thomas precession” and “gyrocommutative law” since these terms have been coined later,
respectively, following Thomas’ 1926 paper [40], and in 1991 [42, 43]. We may remark that Thomas precession has purely kinematical origin, as emphasized in [46]. Accordingly, the presence of Thomas precession is not connected with the action of any force.

Contrasting the discovery before 1914 of what we presently call the gyrocommutative law, the gyroassociative laws, left and right, were discovered about 75 years later, in 1988 [41].

A most important property of Thomas gyration is the so called loop property (left and right),

\[
\begin{align*}
\text{(21)} & \quad \text{gyr}[u \oplus v, v] = \text{gyr}[u, v] \quad \text{Left Loop Property} \\
\text{gyr}[u, v \oplus u] = \text{gyr}[u, v] \quad \text{Right Loop Property}
\end{align*}
\]

for all \(u, v \in V_c\). The left loop property will prove useful in (25) in solving a basic gyrogroup equation.

The grouplike groupoid \(V_c, \oplus\) that regulates Einstein addition, \(\oplus\), in the ball \(V_c\) of the Euclidean 3-space \(V\) is a gyrocommutative gyrogroup called an Einstein gyrogroup. Einstein gyrogroups and gyrovector spaces are studied in [44, 45, 47, 50]. Gyrogroups are not peculiar to Einstein addition [48]. Rather, they are abound in the theory of groups [14, 15, 12], loops [18], quasigroup [19, 23], and Lie groups [20, 21, 22].

Taking the key features of Einstein velocity addition law, and guided by analogies with groups, we are led to the following formal definition of abstract gyrogroups:

**Definition 1. (Gyrogroups).** A groupoid is a non-empty set with a binary operation. A groupoid \((G, \oplus)\) is a gyrogroup if its binary operation satisfies the following axioms. In \(G\) there is at least one element, 0, called a left identity, satisfying

\((G1)\) \quad 0 \oplus a = a

for all \(a \in G\). There is an element \(0 \in G\) satisfying axiom \((G1)\) such that for each \(a \in G\) there is an element \(\ominus a \in G\), called a left inverse of \(a\), satisfying

\((G2)\) \quad \ominus a \oplus a = 0 .

Moreover, for any \(a, b, c \in G\) there exists a unique element \(\text{gyr}[a, b]c \in G\) such that the binary operation obeys the left gyroassociative law

\((G3)\) \quad a \oplus (b \oplus c) = (a \oplus b) \oplus \text{gyr}[a, b]c .

The map \(\text{gyr}[a, b] : G \to G\) given by \(c \mapsto \text{gyr}[a, b]c\) is an automorphism of the groupoid \((G, \oplus)\), that is,

\((G4)\) \quad \text{gyr}[a, b] \in \text{Aut}(G, \oplus) ,

and the automorphism \(\text{gyr}[a, b]\) of \(G\) is called the gyroautomorphism, or the gyration, of \(G\) generated by \(a, b \in G\). The operator \(\text{gyr} : G \times G \to \text{Aut}(G, \oplus)\) is called the gyrator of \(G\). Finally, the gyroautomorphism \(\text{gyr}[a, b]\) generated by any \(a, b \in G\) possesses the left loop property

\((G5)\) \quad \text{gyr}[a, b] = \text{gyr}[a \oplus b, b] .

The first pair of the gyrogroup axioms are like the group axioms. The last pair present the gyrator axioms and the middle axiom links the two pairs.

As in group theory, we use the notation \(a \ominus b = a \oplus (\ominus b)\) in gyrogroup theory as well.

Some groups are commutative. In full analogy, some gyrogroups are gyrocommutative.
Definition 2. (Gyrocommutative Gyrogroups). A gyrogroup \((G, \oplus)\) is gyrocommutative if its binary operation obeys the gyrocommutative law \((G6)\)

\[ a \oplus b = \text{gyr}[a, b](b \oplus a) \]

for all \(a, b \in G\).

We thus see that it is Thomas gyration that regulates Einstein addition, endowing it with the rich structure of a gyrocommutative gyrogroup.

The gyrogroup operation (or, addition) of any gyrogroup has an associated dual operation, called the gyrogroup cooperation (or, coaddition), the definition of which follows:

Definition 3. (The Gyrogroup Cooperation (Coaddition)). Let \((G, \oplus)\) be a gyrogroup with gyrogroup operation (or, addition) \(\oplus\). The gyrogroup cooperation (or, coaddition) \(\ominus\) is a second binary operation in \(G\) given by the equation

\[ a \ominus b = a \oplus \text{gyr}[a, b]b \]

for all \(a, b \in G\).

Replacing \(b\) by \(\ominus b\) in \((22)\) we have the cosubtraction identity

\[ a \ominus b := a \oplus (\ominus b) = a \ominus \text{gyr}[a, b]b \]

for all \(a, b \in G\).

To motivate the introduction of the gyrogroup cooperation, let us solve the equation

\[ x \oplus a = b \]

for the unknown \(x\) in a gyrogroup \((G, \oplus)\).

Assuming that a solution \(x\) exists, we have the following chain of equations

\[
\begin{align*}
x &= x \oplus 0 \\
&= x \oplus (a \ominus a) \\
&= (x \oplus a) \oplus \text{gyr}[x, a](\ominus a) \\
&= (x \oplus a) \ominus \text{gyr}[x, a]a \\
&= (x \ominus a) \ominus \text{gyr}[x \ominus a, a]a \\
&= b \ominus \text{gyr}[b, a]a \\
&= b \ominus a
\end{align*}
\]

where the gyrogroup cosubtraction, \((23)\), which captures here an obvious analogy, comes into play. Hence, if a solution \(x\) to the gyrogroup equation \((24)\) exists, it must be given uniquely by \((25)\). One can finally show that the latter is indeed a solution \([47, \text{Sec. 2.4}]\).

The gyrogroup cooperation is introduced into gyrogroups in order to capture useful analogies between gyrogroups and groups, and to uncover duality symmetries with the gyrogroup operation. Thus, for instance, the gyrogroup cooperation recovers the seemingly missing right counterpart of the left cancellation law \((7)\), giving rise to the following right cancellation law:

\[ (v \oplus u) \ominus u = v \]
for all \( u, v \) in the ball. Furthermore, the right cancellation law (26) can be dualized, giving rise to the dual right cancellation law

\[(v \oplus u) \ominus u = v\]

As an example, and for later reference, we note that it follows from (26) that

\[(28) \quad d = (b \oplus c) \ominus a \quad \implies \quad b \oplus c = d \oplus a\]

in any gyrocommutative gyrogroup.

A gyrogroup cooperation is commutative if and only if the gyrogroup is gyrocommutative [45, Theorem 3.4] [47, Theorem 3.4]. Hence, in particular, Einstein coaddition is commutative. Indeed, Einstein coaddition, \( \oplus \), in an Einstein gyrogroup \((V_s, \oplus, \ominus)\) is given by the equation [47, Eq. 3.195]

\[(29) \quad u \oplus v = 2 \otimes \frac{\gamma_u u + \gamma_v v}{\gamma_u + \gamma_v}\]

which is commutative, as expected. Hence, for instance, \( d \ominus a = a \ominus d \) in (28). The symbol \( \otimes \) in (29) represents scalar multiplication so that, for instance, \( 2 \otimes v = v \oplus v \), for all \( v \) in a gyrogroup \((G, \oplus)\), as explained in Sec. 4.

4. EINSTEIN GYROVECTOR SPACES

Einstein addition in the ball admits scalar multiplication, giving rise to the following definition [43].

**Definition 4.** An Einstein gyrovector space \((V_s, \oplus, \otimes)\) is an Einstein gyrogroup \((V_s, \oplus)\), \(V_s \subseteq V\), with scalar multiplication \( \otimes \) given by the equation

\[(30) \quad r \otimes v = s \left( \left(1 + \frac{\|v\|}{s}\right)^r - \left(1 - \frac{\|v\|}{s}\right)^r \right) \frac{v}{\|v\|}\]

\[= s \tanh(r \tanh^{-1} \frac{\|v\|}{s}) \frac{v}{\|v\|}\]

where \( r \) is any real number, \( r \in \mathbb{R} \), \( v \in V_s \), \( v \neq 0 \), and \( r \otimes 0 = 0 \), and with which we use the notation \( v \otimes r = r \otimes v \).

Einstein gyrovector spaces are studied in [47, Sec. 6.18] and [50]. Einstein scalar multiplication does not distribute over Einstein addition, but it possesses other properties of vector spaces. For any positive integer \( n \), and for all real numbers \( r_1, r_2 \in \mathbb{R} \), and \( v \in V_s \), we have

\[
n \otimes v = v \oplus \ldots \oplus v \quad n \text{ terms}
\]

\[
(r_1 + r_2) \otimes v = r_1 \otimes v \oplus r_2 \otimes v
\]

Scalar Distributive Law

\[
(r_1 r_2) \otimes v = r_1 \otimes (r_2 \otimes v)
\]

Scalar Associative Law

\[
r \otimes (r_1 \otimes v) = r \otimes (r_1 \otimes v) \oplus r \otimes (r_2 \otimes v)
\]

Monodistributive Law

in any Einstein gyrovector space \((V_s, \oplus, \otimes)\).
Any Einstein gyrovector space \((V_s, \oplus, \otimes)\) inherits an inner product and a norm from its vector space \(V\). These turn out to be invariant under gyrations, that is,
\[
\text{gyr}[a, b]u \cdot \text{gyr}[a, b]v = u \cdot v
\]
\[
\|\text{gyr}[a, b]v\| = \|v\|
\]
for all \(a, b, u, v \in V_s\).

Unlike vector spaces, Einstein gyrovector spaces \((V_s, \oplus, \otimes)\) do not possess the distributive law since, in general,
\[
r \otimes (u \oplus v) \neq r \otimes u \oplus r \otimes v
\]
for \(r \in \mathbb{R}\) and \(u, v \in V_s\). One might suppose that there is a price to pay in mathematical regularity when replacing ordinary vector addition with Einstein addition, but this is not the case as demonstrated in [44, 45, 47], and as noted by S. Walter in [55].

In full analogy with the common Euclidean distance function, Einstein addition gives rise to the gyrodistance function
\[
d_\oplus(a, b) = \|\oplus a \oplus b\|
\]
that obeys the gyrotriangle inequality [47, Theorem 6.9] for any \(a, b, p \in V_s\) in an Einstein gyrovector space \((V_s, \oplus, \otimes)\). The gyrodistance function is invariant under the group of motions of any Einstein gyrovector space, that is, under left gyrotranslations and rotations of the space [47, Sec. 4]. The gyrotriangle inequality (33) reduces to a corresponding gyrotriangle equality,
\[
d_\oplus(a, b) = d_\oplus(a, p) \oplus d_\oplus(p, b)
\]
if and only if point \(p\) lies between points \(a\) and \(b\), that is, point \(p\) lies on the gyrosegment \(ab\), as shown in Fig. 2 for points \(A, B,\) and \(P\). Accordingly, the gyrodistance function is gyroadditive on gyrolines, as formulated in [55], and illustrated graphically in Fig. 2.

Furthermore, the Einstein gyrodistance function (33) in any \(n\)-dimensional Einstein gyrovector space \((\mathbb{R}_c^n, \oplus, \otimes)\), \(\mathbb{R}_c^n \subset \mathbb{R}^n\) being the \(n\)-dimensional \(c\)-ball, possesses a familiar Riemannian line element. It gives rise to the Riemannian line element \(ds_c^2\) of the Einstein gyrovector space with its gyrometric (33),
\[
d_\oplus^2 = \|
\]
\[
= \frac{c^2}{c^2 - x^2} dx^2 + \left(\frac{c^2}{c^2 - x^2}\right)^2 (x \cdot dx)^2
\]
where \(dx^2 = dx \cdot dx\), as shown in [47, Theorem 7.6].

Remarkably, the Riemannian line element \(ds_c^2\) in (36) turns out to be the well-known Riemannian line element that the Italian mathematician Eugenio Beltrami introduced in 1868 in order to study hyperbolic geometry by a Euclidean disc model, now known as the Beltrami-Klein disc [27, p. 220, 3]. An English translation of his historically significant 1868 essay on the interpretation of non-Euclidean geometry is found in [38]. The significance of Beltrami’s 1868 essay rests on the generally known fact that it was the first to offer a concrete interpretation of hyperbolic geometry.
Figure 1. The unique gyroline $L_{AB}$ in an Einstein gyrovector space $(\mathbb{R}^n_s, \oplus, \otimes)$ through two given points $A$ and $B$. The case of the Einstein gyrovector plane, when $\mathbb{R}^n_s = \mathbb{R}^2_s = 1$ is the real open unit disc, is shown.

Figure 2. The gyrosegment $AB$ that links the points $A$ and $B$ in $(\mathbb{R}^n_s, \oplus, \otimes)$, with one of its generic points $P$ and its gyromidpoint $M_{AB}$. The point $P$ lies between $A$ and $B$ and, hence, obeys the gyrotriangle equality.

by interpreting “straight lines” as geodesics on a surface of a constant negative curvature. Beltrami, thus, constructed a Euclidean disc model of the hyperbolic plane [27] [38], which now bears his name along with the name of Klein.

We have thus found that the Beltrami-Klein ball model of hyperbolic geometry is regulated algebraically by Einstein gyrovector spaces with their gyrodistance function [33] and Riemannian line element [36], just as the standard model of Euclidean geometry is regulated algebraically by vector spaces with their Euclidean distance function and Riemannian line element $ds^2 = dx^2$.

In full analogy with Euclidean geometry, the unique Einstein gyroline $L_{AB}$, Fig. 1, that passes through two given points $A$ and $B$ in an Einstein gyrovector space $V_s = (V_s, \oplus, \otimes)$ is given by the equation

$$L_{AB} = A \oplus (\oplus A \oplus B) \otimes t$$

Einstein gyrolines are chords of the ball, which turn out to be the familiar geodesics of the Beltrami-Klein ball model of hyperbolic geometry [27]. Accordingly, Einstein gyrosegments are Euclidean segments, as shown in Fig. 2.

The gyromidpoint $M_{AB}$ of gyrosegment $AB$ is the unique point of the gyrosegment that satisfies the equation $d_\otimes(M_{AB}, A) = d_\otimes(M_{AB}, B)$. It is given by each of the following equations [50, Theorem 3.33], Fig. 2

$$M_{AB} = A \oplus (\oplus A \oplus B) \otimes \frac{1}{2} = \frac{\gamma_A A + \gamma_B B}{\gamma_A + \gamma_B} = \frac{1}{2} \otimes (A \boxplus B)$$
The Gyroparallelogram Condition: $D = (B \oplus C) \ominus A$

\[
\begin{align*}
M_{AD} &= \frac{\gamma_A^B + \gamma_A^D}{\gamma_A^B + \gamma_A^D} = \frac{1}{2}(A \oplus D) \\
M_{BC} &= \frac{\gamma_B^C + \gamma_B^C}{\gamma_B^C + \gamma_B^C} = \frac{1}{2}(A \oplus C) \\
M_{ABDC} &= \frac{\gamma_A^B + \gamma_B^C + \gamma_C^D}{\gamma_A^B + \gamma_B^C + \gamma_C^D} = \frac{1}{2}(A \oplus C) \\
\end{align*}
\]

\[
\mathbf{u} \oplus \mathbf{v} = \mathbf{w}
\]

Figure 3. The Einstein gyroparallelogram and its addition law. The gyrodiagonals $AD$ and $BC$ of gyroparallelogram $ABDC$ intersect each other at their gyromidpoints. Detailed studies of the gyroparallelogram and its extension to higher dimensional gyroparallelepipeds are presented in [45, 47]. The gyroparallelogram addition law plays an important role in our gyrovector space approach to hyperbolic geometry, studied in [47, 50].

in full analogy with Euclidean midpoints.

In Euclidean geometry a parallelogram is a quadrilateral the two diagonals of which intersect at their midpoints. In full analogy, in hyperbolic geometry a gyroparallelogram is a gyroquadrilateral the two gyrodiagonals of which intersect at their gyromidpoints. Accordingly, if $A$, $B$ and $C$ are any three non-gyrocollinear points (that is, they do not lie on a gyroline) in an Einstein gyrovector space, and if a fourth point $D$ is given by the gyroparallelogram condition

\[
D = (B \oplus C) \ominus A
\]

then the gyroquadrilateral $ABDC$ is a gyroparallelogram, shown in Fig. 3.

Indeed, the two gyrodiagonals of gyroquadrilateral $ABDC$ are the gyrosegments $AD$ and $BC$, shown in Fig. 3 the gyromidpoints of which coincide, that is,

\[
\frac{1}{2}(A \oplus D) = \frac{1}{2}(B \oplus C)
\]

The result in (40) follows from the gyroparallelogram condition (39), the implication in (28) and a gyromidpoint identity in (38).

In his 1905 paper that founded the special theory of relativity [8], Einstein noted about his addition law of relativistically admissible velocities:

“Das Gesetz vom Parallelogramm der Geschwindigkeiten gilt also nach unserer Theorie nur in erster Annäherung.”

A. Einstein [8]
Figure 4. Two equivalent vectors in a Euclidean vector plane \((\mathbb{R}^2, +, \cdot)\). The two vectors are parallel and have equal values and, hence, equal lengths.

Figure 5. Two equivalent gyrovectors in an Einstein gyrovector plane \((\mathbb{R}^2_c, \oplus, \otimes)\). The two gyrovectors have equal values and, hence, equal gyrolengths.

[English translation: Thus the law of velocity parallelogram is valid according to our theory only to a first approximation.]

Indeed, Einstein velocity addition, \(\oplus\), is noncommutative and does not give rise to an exact “velocity parallelogram” in Euclidean geometry. However, as we see in Fig. 3, Einstein velocity coaddition, \(\boxplus\), which is commutative, does give rise to an exact “velocity gyroparallelogram” in hyperbolic geometry.

The breakdown of commutativity in Einstein velocity addition law seemed undesirable to the famous mathematician Émile Borel. Borel’s resulting attempt to “repair” the seemingly “defective” Einstein velocity addition in the years following 1912 is described by Walter in [54, p. 117]. Here, however, we see that there is no need to repair Einstein velocity addition law for being noncommutative since, despite of being noncommutative, it gives rise to the gyroparallelogram law of gyrovector addition, which turns out to be commutative. The compatibility of our gyroparallelogram addition law of Einsteinian velocities with cosmological observations of stellar aberration will be discussed in Sec. 8. The extension of the gyroparallelogram addition law of \(n = 2\) summands to a corresponding gyroparallelepiped addition law of \(n > 2\) summands is presented in [47, Theorem 10.6].

5. VECTORS AND GYROVECTORS

Elements of a real inner product space \(\mathbb{V} = (\mathbb{V}, +, \cdot)\), called points and denoted by capital italic letters, \(A, B, P, Q, \) etc, give rise to vectors in \(\mathbb{V}\), denoted by bold roman lowercase letters \(\mathbf{u}, \mathbf{v}, \) etc. Any two ordered points \(P, Q \in \mathbb{V}\) give rise to a unique rooted vector \(\mathbf{v} \in \mathbb{V}\), rooted at the point \(P\). It has a tail at the point \(P\) and a head at the point \(Q\), and it has the value \(-P + Q\),

\[
\mathbf{v} = -P + Q
\]
The length of the rooted vector \( v = -P + Q \) is the distance between the points \( P \) and \( Q \), given by the equation

\[
\|v\| = \| -P + Q \|
\]

Two rooted vectors \( -P + Q \) and \( -R + S \) are equivalent if they have the same value, that is,

\[
-P + Q \sim -R + S \quad \text{if and only if} \quad -P + Q = -R + S
\]

The relation \( \sim \) in (43) between rooted vectors is reflexive, symmetric and transitive, so that it is an equivalence relations that gives rise to equivalence classes of rooted vectors. Two equivalent rooted vectors in a Euclidean vector plane are shown in Fig. 4. To liberate rooted vectors from their roots we define a vector to be an equivalence class of rooted vectors. The vector \( -P + Q \) is thus a representative of all rooted vectors with value \( -P + Q \).

A point \( P \in \mathbb{V} \) is identified with the vector \( -O + P \), \( O \) being the arbitrarily selected origin of the space \( \mathbb{V} \). Hence, the algebra of vectors can be applied to points as well. Naturally, geometric and physical properties regulated by a vector space are independent of the choice of the origin.

Let \( A, B, C \in \mathbb{V} \) be three non-collinear points, and let

\[
\begin{align*}
\mathbf{u} &= -A + B \\
\mathbf{v} &= -A + C
\end{align*}
\]

be two vectors in \( \mathbb{V} \) that, without loss of generality, possess the same tail, \( A \). Furthermore, let \( D \) be a point of \( \mathbb{V} \) given by the parallelogram condition

\[
D = B + C - A
\]

Then, the quadrilateral \( ABDC \) is a parallelogram in Euclidean geometry in the sense that its two diagonals, \( AD \) and \( BC \), intersect at their midpoints, that is,

\[
\frac{1}{2}(A + D) = \frac{1}{2}(B + C)
\]

Then, the vector addition of the vectors \( \mathbf{u} \) and \( \mathbf{v} \) that generate the parallelogram \( ABDC \) is the vector \( \mathbf{w} \) given by the well-known parallelogram addition law,

\[
\mathbf{w} = -A + D
\]

where \( \mathbf{w} \) is the vector formed by the diagonal \( AD \) of the parallelogram \( ABDC \).

Vectors in the space \( \mathbb{V} \) are, thus, equivalence classes of ordered pairs of points that add according to the parallelogram law.

Gyrovectors emerge in Einstein gyrovector spaces \( \mathbb{V}_s = (\mathbb{V}_s, \oplus, \otimes) \) in a way fully analogous to the way vectors emerge in the spaces \( \mathbb{V} \). Elements of \( \mathbb{V}_s \), called points and denoted by capital italic letters, \( A, B, P, Q, \) etc, give rise to gyrovectors in \( \mathbb{V}_s \), denoted by bold roman lowercase letters \( \mathbf{u}, \mathbf{v}, \) etc. Any two ordered points \( P, Q \in \mathbb{V}_s \) give rise to a unique rooted gyrovector \( \mathbf{v} \in \mathbb{V}_s \), rooted at the point \( P \). It has a tail at the point \( P \) and a head at the point \( Q \), and it has the value \( \oplus P \oplus Q \),

\[
\mathbf{v} = \oplus P \oplus Q
\]

The gyrolength of the rooted gyrovector \( \mathbf{v} = \oplus P \oplus Q \) is the gyrodistance between the points \( P \) and \( Q \), given by the equation

\[
\|\mathbf{v}\| = \|\oplus P \oplus Q\|
\]
Two rooted gyrovectors $\oplus P \oplus Q$ and $\oplus R \oplus S$ are equivalent if they have the same value, that is,
\begin{equation}
\oplus P \oplus Q \sim \oplus R \oplus S \quad \text{if and only if} \quad \oplus P \oplus Q = \oplus R \oplus S
\end{equation}
The relation $\sim$ in (50) between rooted gyrovectors is reflexive, symmetric and transitive, so that it is an equivalence relation that gives rise to equivalence classes of rooted gyrovectors. Two equivalent rooted gyrovectors in an Einstein gyrovector plane are shown in Fig. 5. To liberate rooted gyrovectors from their roots we define a gyrovector to be an equivalence class of rooted gyrovectors. The gyrovector $\oplus P \oplus Q$ is thus a representative of all rooted gyrovectors with value $\oplus P \oplus Q$.

A point $P$ of a gyrovector space $(V_s, \oplus, \otimes)$ is identified with the gyrovector $\ominus O \oplus P$, $O$ being the arbitrarily selected origin of the space $V_s$. Hence, the algebra of gyrovectors can be applied to points as well. Naturally, geometric and physical properties regulated by a gyrovector space are independent of the choice of the origin.

Let $A, B, C \in V_s$ be three non-gyrocollinear points of an Einstein gyrovector space $(V_s, \oplus, \otimes)$, and let
\begin{equation}
\mathbf{u} = \ominus A \oplus B \quad \mathbf{v} = \ominus A \oplus C
\end{equation}
be two gyrovectors in $V$ that, without loss of generality, possess the same tail, $A$. Furthermore, let $D$ be a point of $V_s$ given by the gyroparallelogram condition
\begin{equation}
D = (B \boxplus C) \ominus A
\end{equation}
Then, the gyroquadrilateral $ABDC$ is a gyroparallelogram in the Beltrami-Klein ball model of hyperbolic geometry in the sense that its two gyrodiagonals, $AD$ and $BC$, intersect at their gyromidpoints, that is,
\begin{equation}
\frac{1}{2}(A \oplus D) = \frac{1}{2}(B \oplus C)
\end{equation}
as explained in Sec.[4] and illustrated in Fig.[3] Then, the gyrovector addition of the gyrovectors $\mathbf{u}$ and $\mathbf{v}$ that generate the gyroparallelogram $ABDC$ is the gyrovector $\mathbf{w}$ given by the gyroparallelogram addition law, Fig.[3]
\begin{equation}
\mathbf{w} = \ominus A \oplus D = \mathbf{u} \oplus \mathbf{v}
\end{equation}
where $\mathbf{w}$ is the gyrovector formed by the gyrodiagonal $AD$ of the gyroparallelogram $ABDC$.

Gyrovectors in the ball $V_s$ are, thus, equivalence classes of ordered pairs of points that add according to the gyroparallelogram law.

6. GYROTRIGONOMETRY IN EINSTEIN GYROVECTOR SPACES

Let $A, B$ and $C$ be three non-gyrocollinear points in an Einstein gyrovector space $(V_s, \oplus, \otimes)$, and let $\ominus A \oplus B$ and $\ominus A \oplus C$ be the resulting two nonzero gyrovectors with the common tail $A$ and included gyroangle $\alpha$, as shown in Fig.[6] The gyrolength of gyrovector $\ominus A \oplus B$, $\| \ominus A \oplus B \|$, is nonzero, and its associated gyrovector
\begin{equation}
\frac{\ominus A \oplus B}{\| \ominus A \oplus B \|}
\end{equation}
is called a unit gyrovector. Guided by analogies with Euclidean trigonometry, the gyrocosine, \( \cos \), of the gyroangle \( \alpha \) included by the two gyrovectors that emanate from point \( A \) is given by the equation

\[
(56) \quad \cos \alpha = \frac{∇B∇A}{\|∇B∇A\|} \frac{∇C∇A}{\|∇C∇A\|} \frac{∇A∇B}{\|∇A∇B\|} \frac{∇A∇C}{\|∇A∇C\|}
\]

as illustrated in Fig. 6 for gyrotriangle gyroangles. The gyroangle is invariant under the group of motions of any Einstein gyrovector space, that is, under left gyrotranslations and rotations of the space [47, Theorem 8.6].

Applying the gamma identity [3] and the definition of the gyrocosine function in [56] to the gyrovector sides of a gyrotriangle \( ABC \), using the standard gyrotriangle notation in Fig. 6 we obtain the following law of gyrocosines [50, Sec. 4.5], [47, Sec. 12.2],

\[
\begin{align*}
\gamma_a &= \gamma_b \gamma_c (1 - b_s c_s \cos \alpha) \\
\gamma_b &= \gamma_a \gamma_c (1 - a_s c_s \cos \beta) \\
\gamma_c &= \gamma_a \gamma_b (1 - a_s b_s \cos \gamma)
\end{align*}
\]

\[
\delta = \pi - (\alpha + \beta + \gamma) > 0
\]
Like Euclidean triangles, the gyroangles of a gyrotriangle are uniquely determined by its sides. Solving the system (57) of three identities for the three unknowns \( \cos \alpha \), \( \cos \beta \) and \( \cos \gamma \), and employing (5), we obtain the following theorem.

**Theorem 5. (The Law of Gyrocosines; The SSS to AAA Conversion Law).** Let \( \triangle ABC \) be a gyrotriangle in an Einstein gyrovector space \( (V_s, \oplus, \otimes) \). Then, in the gyrotriangle notation in Fig. 6,

\[
\begin{align*}
\cos \alpha &= \frac{-\gamma_a + \gamma_b \gamma_c}{\gamma_b \gamma_c b_a c_a} = \frac{-\gamma_a + \gamma_b \gamma_c}{\sqrt{\gamma_b^2 - 1} \sqrt{\gamma_c^2 - 1}} \\
\cos \beta &= \frac{-\gamma_b + \gamma_a \gamma_c}{\gamma_a \gamma_c a_b c_b} = \frac{-\gamma_b + \gamma_a \gamma_c}{\sqrt{\gamma_a^2 - 1} \sqrt{\gamma_c^2 - 1}} \\
\cos \gamma &= \frac{-\gamma_c + \gamma_a \gamma_b}{\gamma_a \gamma_b a_c b_a} = \frac{-\gamma_c + \gamma_a \gamma_b}{\sqrt{\gamma_a^2 - 1} \sqrt{\gamma_b^2 - 1}}
\end{align*}
\]

The identities in (58) form the SSS (Side-Side-Side) to AAA (gyroAngle-gyroAngle-gyroAngle) conversion law in Einstein gyrovector spaces. This law is useful for calculating the gyroangles of a gyrotriangle in an Einstein gyrovector space when its sides (that is, its side-gyrolengths) are known.

In full analogy with the trigonometry of triangles, the **gyrosine** of a gyrotriangle gyroangle \( \alpha \) is nonnegative, given by the equation

\[\sin \alpha = \sqrt{1 - \cos^2 \alpha}\]  

(59)

Hence, it follows from Theorem 5 that the gyrosine of the gyrotriangle gyroangles in that Theorem are given by

\[
\begin{align*}
\sin \alpha &= \sqrt{1 + 2 \gamma_a \gamma_b \gamma_c - \gamma_a^2 - \gamma_b^2 - \gamma_c^2} \\
\sqrt{\gamma_a^2 - 1} \sqrt{\gamma_b^2 - 1} \\
\sin \beta &= \sqrt{1 + 2 \gamma_a \gamma_b \gamma_c - \gamma_a^2 - \gamma_b^2 - \gamma_c^2} \\
\sqrt{\gamma_a^2 - 1} \sqrt{\gamma_b^2 - 1} \\
\sin \gamma &= \sqrt{1 + 2 \gamma_a \gamma_b \gamma_c - \gamma_a^2 - \gamma_b^2 - \gamma_c^2} \\
\sqrt{\gamma_a^2 - 1} \sqrt{\gamma_b^2 - 1}
\end{align*}
\]

(60)

Identities (60) immediately give rise to the identities

\[
\frac{\sin \alpha}{\sqrt{\gamma_a^2 - 1}} = \frac{\sin \beta}{\sqrt{\gamma_b^2 - 1}} = \frac{\sin \gamma}{\sqrt{\gamma_c^2 - 1}}
\]

(61)

that form the law of gyrosines.

Unlike Euclidean triangles, the side gyrolengths of a gyrotriangle are uniquely determined by its gyroangles, as the following theorem demonstrates.

**Theorem 6. (The AAA to SSS Conversion Law).** Let \( \triangle ABC \) be a gyrotriangle in an Einstein gyrovector space \( (V_s, \oplus, \otimes) \). Then, in the gyrotriangle notation in
\[
\begin{align*}
\sin \alpha &= \frac{\gamma_a}{\gamma_c}, \\
\cos \alpha &= \frac{b}{c}, \\
\sin^2 \alpha + \cos^2 \alpha &= 1
\end{align*}
\]

\[
\begin{align*}
\sin \beta &= \frac{\gamma_b}{\gamma_c}, \\
\cos \beta &= \frac{a}{c}, \\
\sin^2 \beta + \cos^2 \beta &= 1
\end{align*}
\]

Figure 7. Gyrotrigonometry in an Einstein gyrovector plane \((\mathbb{R}_s^2, \oplus, \otimes)\).

**Proof.** Let \(ABC\) be a gyrotriangle in an Einstein gyrovector space \((V_s, \oplus, \otimes)\) with its standard notation in Fig. 6. It follows straightforwardly from the SSS to AAA conversion law (58) that

\[
\left(\frac{\cos \alpha + \cos \beta \cos \gamma}{\sin \beta \sin \gamma}\right)^2 = \frac{(\cos \alpha + \cos \beta \cos \gamma)^2}{(1 - \cos^2 \beta)(1 - \cos^2 \gamma)} = \gamma_a^2
\]

implying the first identity in (62). The remaining two identities in (62) are obtained from (58) by permutation of vertices. \(\square\)

The identities in (62) form the AAA to SSS conversion law. This law is useful for calculating the sides (that is, the side-gyrolengths) of a gyrotriangle in an Einstein gyrovector space when its gyroangles are known. Thus, for instance, \(\gamma_a\) is obtained from the first identity in (62), and \(a\) is obtained from \(\gamma_a\) by Identity (5).

Let \(ABC\) be a right gyrotriangle in an Einstein gyrovector space \((V_s, \oplus, \otimes)\) with the right gyroangle \(\gamma = \pi/2\), as shown in Fig. 7 for \(V_s = \mathbb{R}_s^2\). It follows from (62) with \(\gamma = \pi/2\) that the sides \(a, b\) and \(c\) of gyrotriangle \(ABC\) in Fig. 7 are related to
the acute gyroangles $\alpha$ and $\beta$ of the gyrotriangle by the equations

\[
\begin{align*}
\gamma_a &= \frac{\cos \alpha}{\sin \beta} \\
\gamma_b &= \frac{\cos \beta}{\sin \alpha} \\
\gamma_c &= \frac{\cos \alpha \cos \beta}{\sin \alpha \sin \beta}
\end{align*}
\]

(64)

The identities in (64) imply the Einstein-Pythagoras Identity

(65)

\[
\gamma_a \gamma_b = \gamma_c
\]

for a right gyrotriangle $ABC$ with hypotenuse $c$ and legs $a$ and $b$ in an Einstein gyrovector space, Fig. 7.

Let $a$, $b$ and $c$ be the respective gyrolengths of the two legs $a$, $b$ and the hypotenuse $c$ of a right gyrotriangle $ABC$ in an Einstein gyrovector space $(\mathbb{V}_s, \oplus, \otimes)$, Fig. 7. By (5), p. 3, and (64) we have

\[
\begin{align*}
\left(\frac{a}{c}\right)^2 &= \frac{(\gamma_a^2 - 1)/\gamma_a^2}{(\gamma_c^2 - 1)/\gamma_c^2} = \cos^2 \beta \\
\left(\frac{b}{c}\right)^2 &= \frac{(\gamma_b^2 - 1)/\gamma_b^2}{(\gamma_c^2 - 1)/\gamma_c^2} = \cos^2 \alpha
\end{align*}
\]

(66)

where $\gamma_a$, $\gamma_b$ and $\gamma_c$ are related by (65).

Similarly, by (5), and (64) we also have

\[
\begin{align*}
\left(\frac{\gamma_a a}{\gamma_c c}\right)^2 &= \frac{\gamma_a^2 - 1}{\gamma_c^2 - 1} = \sin^2 \alpha \\
\left(\frac{\gamma_b b}{\gamma_c c}\right)^2 &= \frac{\gamma_b^2 - 1}{\gamma_c^2 - 1} = \sin^2 \beta
\end{align*}
\]

(67)

Identities (66) and (67) imply

\[
\begin{align*}
\left(\frac{a}{c}\right)^2 + \left(\frac{\gamma_b b}{\gamma_c c}\right)^2 &= 1 \\
\left(\frac{\gamma_a a}{\gamma_c c}\right)^2 + \left(\frac{b}{c}\right)^2 &= 1
\end{align*}
\]

(68)

and, as shown in Fig. 7

\[
\begin{align*}
\cos \alpha &= \frac{b}{c} \\
\cos \beta &= \frac{a}{c}
\end{align*}
\]

(69)
and
\[ \sin \alpha = \frac{\gamma a}{\gamma c} \]
\[ \sin \beta = \frac{\gamma b}{\gamma c} \]

(70)

Interestingly, we see from (69) – (70) that the gyrocosine function of an acute gyroangle of a right gyrotriangle in an Einstein gyrovector space has the same form as its Euclidean counterpart, the cosine function. In contrast, it is only modulo gamma factors that the gysine function has the same form as its Euclidean counterpart, the sine function.

Identities (68) give rise to the following two distinct Einsteinian-Pythagorean identities,
\[ a^2 + \left( \frac{\gamma b}{\gamma c} \right)^2 b^2 = c^2 \]
\[ \left( \frac{\gamma a}{\gamma c} \right)^2 a^2 + b^2 = c^2 \]

for a right gyrotriangle with hypotenuse \( c \) and legs \( a \) and \( b \) in an Einstein gyrovector space. The two distinct Einsteinian-Pythagorean identities in (71) that each Einsteinian right gyrotriangle possesses converge in the Newtonian-Euclidean limit of large \( s, s \to \infty \), to the single Pythagorean identity
\[ a^2 + b^2 = c^2 \]

(72)

that each Euclidean right-angled triangle possesses.

7. The Interpretation of Classical Stellar Aberration by Vectors and Trigonometry AND The Interpretation of Relativistic Stellar Aberration by Gyrovectors and Gyrotrigonometry

The universe is our laboratory, we are the experimenters asking nature whether, in the limit of negligible force, relativistic velocities add

(1) according to Einstein’s 1905 velocity addition law (2), or
(2) according to the Einstein gyroparallelogram addition law (54).

Fortunately, the cosmological phenomenon of stellar aberration comes to the rescue, as we will see in this section. We will find that owing to the validity of well-known relativistic stellar aberration formulas,

(i) Einsteinian, relativistic velocities are gyrovectors that add according to the gyroparallelogram addition law (54), Fig. 3, which is commutative, just as
(ii) Newtonian, classical velocities are vectors that add according to the common parallelogram addition law.
To set the stage for our study of the relativistic stellar aberration, we begin with the study of the classical particle aberration, Fig. 8, which will be extended to the study of the relativistic particle aberration, Fig. 9, by gyro-analogies.

Figs. 8 and 9 present, respectively, the Newtonian velocity space $\mathbb{R}^2$, and the Einsteinian velocity space $\mathbb{R}^2_{c=1}$, along with several of their points, where only two dimensions are shown for clarity. The origin, $O$, of each of these two velocity spaces, not shown in Figs. 8–9, is arbitrarily selected, representing an arbitrarily selected inertial rest frame $\Sigma_0$.

Points of a velocity space represent uniform velocities relative to the rest frame $\Sigma_0$. In particular, the points $E$, $S$ and $P$ in Figs. 8–9 represent, respectively, the velocity of the Earth, the Sun, and a Particle (emitted, for instance, from a star) relative to the rest frame $\Sigma_0$.

Accordingly, the Newtonian velocity vector $\mathbf{v}$ of the Sun relative to the Earth is
\begin{equation}
\mathbf{v} = -E + S
\end{equation}
and the Einsteinian velocity gyrovector $\mathbf{v}$ of the Sun relative to the Earth is
\begin{equation}
\mathbf{v} = \ominus E \oplus S
\end{equation}
as shown, respectively, in Figs. 8–9.

Similarly, the particle \( P \) moves uniformly relative to the Earth and relative to the Sun with respective Newtonian velocities, Fig. 8.

\[
P_e = -E + P \\
P_s = -S + P
\]

(74a)

In full analogy with (74a), the Einsteinian velocities of the particle \( P \) relative to the Earth and relative to the Sun are, Fig. 9.

\[
P_e = \odot E \oplus P \\
P_s = \odot S \oplus P
\]

(74b)

The Newtonian velocities (74a) of the particle make angles \( \theta_e \) and \( \theta_s \), respectively, with the Newtonian velocity \( v \) in (73a), as shown in Fig. 8.

Similarly, the Einsteinian velocities (74b) of the particle make gyroangles \( \theta_e \) and \( \theta_s \), respectively, with the Einsteinian velocity \( v \) in (73b), as shown in Fig. 9.

Following Fig. 8, classical particle aberration is the angular change \( \theta_s - \theta_e \) in the apparent direction of a moving particle caused by the motion with Newtonian relative velocity \( v \), (73a), between \( E \) and \( S \). A relationship between the angles \( \theta_s \) and \( \theta_e \) is called a classical particle aberration formula.

Similarly, following Fig. 9, relativistic particle aberration is the gyroangular change \( \theta_s - \theta_e \) in the apparent direction of a moving particle caused by the motion with Einsteinian relative velocity \( v \), (73b), between \( E \) and \( S \). A relationship between the gyroangles \( \theta_s \) and \( \theta_e \) is called a relativistic particle aberration formula.

In order to uncover classical particle aberration formulas we draw the altitude \( PQ \) from vertex \( P \) to side \( ES \) (extended if necessary) obtaining the right-angled triangle \( EQP \) in Fig. 8. The latter, in turn, enables the parallelogram law and the triangle law of Newtonian velocity addition, and the triangle equality and trigonometry, to be applied, obtaining the following two, mutually equivalent, classical particle aberration formulas,

\[
cot \theta_e = \cot \theta_s + \frac{v}{p_s \sin \theta_s}
\]

(75a)

\[
cot \theta_s = \cot \theta_e - \frac{v}{p_e \sin \theta_e}
\]

The resulting classical particle aberration formulas (75a) are in full agreement with formulas available in the literature; see, for instance, [39, Eq. (134), p. 147].

The details of obtaining the classical particle aberration formulas (75a), illustrated in Fig. 8 are presented in [14, Chap. 13] and, hence, will not be presented here.

In full analogy, in order to uncover relativistic particle aberration formulas we draw the altitude \( PQ \) from vertex \( P \) to side \( ES \) (extended if necessary) obtaining the right gyrotriangle \( EQP \) in Fig. 9. The latter, in turn, enables the gyrotriangle law and the gyroparallelogram law of Einsteinian velocity addition, and the gyrotriangle equality and gyrotrigonometry, to be applied, obtaining the following two, mutually
equivalent, relativistic particle aberration formulas,

\[
\cot \theta_c = \gamma_v (\cot \theta_s + \frac{v}{p_s \sin \theta_s}) \\
\cot \theta_s = \gamma_v (\cot \theta_c - \frac{v}{p_c \sin \theta_c})
\] (75b)

The resulting relativistic particle aberration formulas (75b) are in full agreement with formulas available in the literature; see, for instance, [33, p. 53], [34, p. 86] and [24, pp. 12–14].

The details of obtaining the relativistic particle aberration formulas (75b), illustrated in Fig. 9, are presented in [47, Chap. 13] and, hence, will not be presented here.

In Euclidean geometry the triangle law and the parallelogram law of vector addition are equivalent. In full analogy, their gyro-counterparts are equivalent as well, as explained in [50, Sec. 4.3] in detail.

The equivalence of the two equations in (75a) implies \( p_s \sin \theta_s = p_e \sin \theta_e \), thus recovering the law of sines for the Euclidean triangle ESP in Fig. 8, noting that \( \sin \theta_s = \sin(\pi - \theta_s) \).

In full analogy, the equivalence of the two equations in (75b) implies the relativistic law of gyrosines [47, Theorem 12.5],

\[
\frac{\gamma_p s}{\sin \theta_e} = \frac{\gamma_p e}{\sin \theta_s} 
\] (76b)

for the gyrotriangle ESP in Fig. 9, noting that \( \sin \theta_s = \sin(\pi - \theta_s) \).

In this section we have described the way to recover the well-known classical particle aberration formulas (75a) by employing trigonometry, the triangle equality, the triangle addition law, and the parallelogram addition law of Newtonian velocities.

In full analogy, we have described in this section the way to recover the well-known relativistic particle aberration formulas (75b) by employing gyrotrigonometry, the gyrotriangle equality, the gyrotriangle addition law, and the gyroparallelogram addition law of Einsteinian velocities.

In contrast, the well-known relativistic particle aberration formulas (75b) are obtained in the literature by employing the Lorentz transformation group of special relativity.

What is remarkable here is that the relativistic particle aberration formulas (75b), which are commonly obtained in the literature by Lorentz transformation considerations, are recovered here by gyrotrigonometry and the gyroparallelogram addition law of Einsteinian velocities, in full analogy with the recovery of their classical counterparts. This remarkable way of recovering the particle aberration formulas (76) demonstrates that since special relativity is governed by the Lorentz transformation group, Einsteinian velocities in special relativity add according to the gyroparallelogram addition law, just as Newtonian velocities add according to the parallelogram addition law.
Hence, any experiment that confirms the validity of the relativistic particle aberration formulas (75b), amounts to an experiment that confirms the validity of the gyroparallelogram addition law of Einsteinian velocities.

In the special case when the particle $P$ in Fig. 9 is a photon emitted from a star, the Einsteinian speeds of the photon relative to both $E$ and $S$ is $p_e = p_s = c$, and the relativistic particle aberration formulas (75b) reduce to the corresponding stellar aberration formulas,

$$\cot \theta_e = \gamma_v \frac{\cos \theta_s + v/c}{\sin \theta_s}$$

$$\cot \theta_s = \gamma_v \frac{\cos \theta_e - v/c}{\sin \theta_e}$$

The discovery of stellar aberration, which results from the velocity of the Earth in its annual orbit about the Sun, by the English astronomer James Bradley in the 1720s, is described, for instance, in [37].

A high precision test of the validity of the stellar aberration formulas (77) in special relativity has recently been obtained as a byproduct of the “GP-B” gyroscope experiment. Indeed, the validity of the stellar aberration formulas (77) is central for the success of the “GP-B” gyroscope experiment developed by NASA and Stanford University [11] to test two unverified predictions of Einstein’s general theory of relativity [10, 16].

The GP-B space gyroscopes encountered two kinds of stellar aberration. Orbital aberration with 97.5-minute period of $\pm 5.1856$ arc-seconds that results from the motion of the gyroscopes around the earth, and annual aberration with one year period of about $\pm 20.4958$ arc-seconds that results from the motion of the earth (and the gyroscopes) around the sun. These aberrations, calculated by methods of special relativity, were used to calibrate the gyroscopes and their accompanying instruments.

If the “GP-B” gyroscope experiment proves successful, it could be considered as an experimental evidence of the validity of the stellar aberration formulas (77) and, hence, the validity of the relativistic particle aberration formulas (75b) as well. The latter, in turn, could be considered as an experimental evidence of the validity of the gyroparallelogram addition law of Einsteinian velocities. Indeed, the preliminary analysis of data has confirmed the theoretical prediction of the “GP-B” gyroscope experiment [17], so that the experiment seems to prove successful.

8. THE RELATIVISTIC MASS AND DARK MATTER

Let

$$S = S(m_k, v_k, \Sigma_0, N)$$

be an isolated system of $N$ noninteracting material particles the $k$-th particle of which has invariant mass $m_k > 0$ and relativistically admissible velocity $v_k \in \mathbb{V}_c$ relative to an inertial frame $\Sigma_0$, $k = 1, \ldots, N$.

Assuming that the four-momentum is additive, the sum of the four-momenta of the $N$ particles of the system $S$ gives the four-momentum $(m_0 \gamma_{v_0}, m_0 \gamma_{v_0} v_0)^t$ of $S$. 
Accordingly,

\[
\sum_{k=1}^{N} m_k \gamma_{v_k} = m_0 \gamma_{v_0}
\]  

(79)

where the invariant masses \(m_k > 0\) and the velocities \(v_k, k = 1, \ldots, N\), relative to \(\Sigma_0\) of the constituent particles of \(S\) are given, while the unique invariant mass \(m_0\) of \(S\) and the unique velocity \(v_0\) of the CM frame of \(S\) relative to \(\Sigma_0\) are to be determined.

In general, if \(m_0\) in (79) exists, it satisfies the inequality

\[
m_0 \neq \sum_{k=1}^{N} m_k
\]

(80)

leading to the conclusion that in special relativity mass, \(m_k\), is not additive [26], and relativistic mass, \(m_k \gamma_{v_k}\), does not mesh up with the Minkowskian four-vector formalism of special relativity [30, 5, 1].

It follows immediately from (79) that if \(m_0 \neq 0\) exists, then \(v_0\) is given by the equation

\[
v_0 = \frac{\sum_{k=1}^{N} m_k \gamma_{v_k} v_k}{\sum_{k=1}^{N} m_k \gamma_{v_k}}
\]

(81)

Employing the gyrocommutative gyrogroup structure of Einstein’s velocity addition law, it is found in [47, Chap. 11] that the unique solution of (79) for the unknown \(m_0 > 0\) is given by the equation [49, ?]

\[
m_0 = \sqrt{\left(\sum_{k=1}^{N} m_k \right)^2 + 2 \sum_{j,k=1}^{N} m_j m_k (\gamma_{v_j} \gamma_{v_k} - 1)}
\]

(82)

Hence, if the four-momentum is additive then, by (79), the velocity \(v_0\) of the CM frame of \(S\) relative to \(\Sigma_0\) is given by (81), and the invariant mass \(m_0\) of \(S\) is given by (82).

Furthermore, it follows from (79) that the relativistic mass \(m_0 \gamma_{v_0}\) is additive, that is,

\[
m_0 \gamma_{v_0} = \sum_{k=1}^{N} m_k \gamma_{v_k}
\]

(83)

We thus see that owing to the introduction of the invariant mass \(m_0\) of a system of particles, given by (82), the relativistic mass is additive, and it meshes extraordinarily well with the Minkowskian four-vector formalism of special relativity.

Suggestively, we define the Newtonian mass, \(m_{\text{newton}}\), of the system \(S\) by the equation

\[
m_{\text{newton}} := \sum_{k=1}^{N} m_k
\]

(84)
and the dark mass, \( m_{\text{dark}} \), of the system \( S \) by the equation

\[
(85) \quad m_{\text{dark}} := \sqrt{\sum_{j,k=1}^{N} m_j m_k \left( \gamma_{\text{\textcircled{i}}j} \text{\textcircled{j}} k \right) - 1}
\]

so that (82) can be written as

\[
(86) \quad m_0 = \sqrt{m_{\text{newton}}^2 + m_{\text{dark}}^2}
\]

The dark mass in (85) measures the extent to which the system \( S \) deviates away from rigidity. Gravitationally, dark mass behaves just like ordinary mass, as postulated in cosmology [6, p. 37]. However, it is undetectable by all means other than gravity since it is fictitious, or virtual, in the sense that it is generated solely by relative motion between constituent objects of the system.

Dark matter was introduced into cosmology as an ad hoc postulate, hypothesized to provide observed missing gravitational force [7]. In contrast, dark mass emerges here as a consequence of the covariance of Einstein’s special theory of relativity, and it stems from relative motion between constituent objects of a system. All relative velocities between the constituent particles of a rigid system vanish, so that the dark mass of a rigid system vanishes as well.

Under special circumstances dark matter may appear or disappear in galaxies, a fact that may increase or decrease the total mass of galaxies which may, in turn, decelerate or accelerate the expansion of the Universe. These special circumstances are characterized by supernovae and star formation.

Each stellar explosion, a supernova, creates relative speeds between objects that were at rest relative to each other prior to the explosion. The resulting generated relative speeds increase the dark mass of the region of the supernova.

Conversely, relative speeds of objects that converge into a star vanish in the process of star formation, resulting in the decrease of the dark mass of a star formation region.

Dark mass is observed in particle physics as well. Let us consider two particles with rest masses \( m_1 \) and \( m_2 \), and velocities \( v_1 \) and \( v_2 \) relative to an inertial rest frame \( \Sigma_0 \), respectively. If these particles were to collide and stick, the rest mass \( m_0 \) and the velocity \( v_0 \) relative to \( \Sigma_0 \) of the resulting composite particle would satisfy the four-momentum conservation law (79), that is

\[
(87) \quad m_0 \begin{pmatrix} \gamma v_0 \\ \gamma v_0 v_0 \end{pmatrix} = m_1 \begin{pmatrix} \gamma v_1 \\ \gamma v_1 v_1 \end{pmatrix} + m_2 \begin{pmatrix} \gamma v_2 \\ \gamma v_2 v_2 \end{pmatrix}
\]

Hence, by (82) and (86),

\[
(88) \quad m_0 = \sqrt{(m_1 + m_2)^2 + 2m_1 m_2 \left( \gamma_{\text{\textcircled{i}v_1}} \text{\textcircled{j}v_2} - 1 \right)} = \sqrt{m_{\text{newton}}^2 + m_{\text{dark}}^2}
\]

where

\[
(89) \quad m_{\text{newton}} = m_1 + m_2 \\
m_{\text{dark}} = 2m_1 m_2 \left( \gamma_{\text{\textcircled{i}v_1}} \text{\textcircled{j}v_2} - 1 \right) > 0
\]
and, by (81),

\[ v_0 = \frac{m_1 \gamma v_1 + m_2 \gamma v_2}{m_1 \gamma v_1 + m_2 \gamma v_2} \]  

Hence, the relativistic mass of the composite particle is \( m_0 \gamma v_0 \), where \( m_0 \) is given by (88), and \( v_0 \) is given by (90), satisfying by (83),

\[ m_0 \gamma v_0 = m_1 \gamma v_1 + m_2 \gamma v_2 \]

It is clear from (88) – (89) that the Newtonian mass, \( m_{\text{newton}} \), is conserved during the collision. It is only the total invariant mass, \( m_0 \), which is increased following the collision owing to the emergence of the dark mass \( m_{\text{dark}} \).

Examples of particles that collide and stick, as described in (87) – (91), are observed in experimental searches for new particles in high-energy particle colliders.

We thus see that our study of the dark mass may be applied on the subatomic scale as well as on the scale of the cosmos.

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