The cross-over from tunnelling to multiphoton ionization of atoms

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We present a theory illuminating the cross-over from strong-field tunnelling ionization to weak-field multiphoton ionization in the interaction of a classical laser field with a hydrogen atom. A simple formula is derived in which the ionization amplitude appears as a product of two separate amplitudes. The first describes the initial polarization of the atom by virtual multiphoton absorption and the second the subsequent tunnelling out of the polarized atom. Tunnelling directly from the ground state and multiphoton absorption without tunnelling appear naturally as the limits of the theory.

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I. INTRODUCTION

Traditionally the interaction of a strong laser field with an atom which leads to ionization is considered to occur by two contrasting mechanisms, as limiting cases of the so-called Keldysh parameter \( \gamma = \sqrt{I_p/(2U_p)} \) where \( I_p \) is the atom ionization potential and \( U_p \) is the laser ponderomotive potential. The Keldysh parameter can be represented in atomic units also as a Lorentzian potential. The Keldysh parameter can be estimated in atomic units also as \( \gamma = \omega(2I_p)^{1/2}/E_0 \), where the classical laser electric field has frequency \( \omega \) and strength \( E_0 \). For \( \gamma < 1 \) or low frequency and high intensity, ionization proceeds by tunnelling out of the atomic potential under the influence of the potential supplied by the field. When \( \gamma > 1 \), for high frequency and relatively low intensity, then direct absorption of several photons occurs giving rise to ionization and possibly "above-threshold" ionization (ATI) due to absorption of further photons by electrons already in the continuum.

The earliest theoretical treatments were based on the "strong-field approximation" (SFA) by Keldysh \[1\], Faisal \[2\] and Reiss \[3\], known as the KFR approach. More recent developments e.g.\[4\],[5], \[6\] contain extensions of the SFA and in certain cases fully numerical calculations are now available \[7\],[8], \[9\].

The SFA and its extensions have been very successful in describing the ionization of atoms by laser fields. It involves the approximation of the exact \( T \)-matrix element for the ionizing transition by a matrix element of the form

\[
f = \int_{-\infty}^{\infty} \langle \phi_f(t') | V_F(t') | \phi_i(t') \rangle \, dt',
\]

where \( | \phi_i \rangle \) is the initial eigenstate of the atom alone, \( | \phi_f \rangle \) is the continuum eigenstate of the laser field alone and \( V_F \) is the interaction of atomic electron with laser field. From the outset it was shown that both limits are contained in this theory. This is plausible if one considers that in a Floquet picture \( | \phi_f \rangle \) contains the electron coupled to any number of photons, so that multiphoton ionization is described. Additionally, in the opposite limit of \( \gamma < 1 \), a semi-classical description of \( | \phi_f \rangle \) in the classical laser field corresponds to a tunnelling interpretation. Nevertheless, a clear physical picture of the intermediate region and the transition between the two limits does not emerge.

Recently there has been renewed interest in developing the theory to explain more physically how the transition from direct tunnelling out of the ground state to the seemingly quite different mechanism of multi-photon absorption occurs. A first step in this direction was made by Klaiber et.al. \[14\] who recognized that during tunnelling the electron may absorb energy from the field to emerge at the tunnel exit with more energy than when it started to tunnel. This "non-adiabatic" energy absorption was treated by classical mechanics. Clearly however, as will be explained in more detail below, the energy gain can also be thought of as the absorption of photons during the tunnelling process.

In a recent letter \[15\] a numerical study of ionization in extremely strong laser fields was reported. The results were interpreted by pointing out that tunnelling may not only occur directly out of the atomic ground state (the hydrogen atom was used as example). Rather, excitation of the atom to higher bound states may occur followed by tunnelling. Indeed, since the binding energy is then much reduced, as in the non-adiabatic case, tunnelling is more efficacious from such excited states. In fact the picture presented corresponds to "over-the-barrier" ionization out of excited states without the need for tunnelling. Although plausible, we feel that such a picture is not quite correct. The authors considered that the real time-independent hydrogen eigenstates must first be populated by multi-photon absorption. Since for the case of hydrogen the first excited \( N = 2 \) manifold lies 10.4\,eV above the ground state, a significant multiphoton tran-
sition is required to populate these states. Higher mani-
folds were not considered.

Here we put forward an alternative picture in which
the dominant role is played not by real eigenstates but by
virtual "off the energy shell" states of the hydrogen atom.
For infra-red and visible photons these energy states lie
far below the $N = 2$ excited manifold but nevertheless
tunnelling can occur from them. We will show that the
picture of virtual absorption allows a simple description
of the smooth transition from tunnelling to multiphoton
ionization to be given. We make clear from the outset
that our aim is not to develop a theory with which to
confront specific experimental data. There one cannot
compete with fully numerical methods. Rather it is to
expose the physical mechanism of laser-atom interaction
and to explain in a simple picture how the cross-over from
tunnelling to multiphoton ionization arises.

When an atom is subject to an electric field whose fre-
cequency is not resonant with transition to an eigenstate,
there is an interaction and distortion which is usually re-
ferred to classically as polarization of the electron cloud.
In the quantized photon picture this is ascribed to the
virtual absorption of photons. After each photon absorp-
tion there is a changed wavefunction and since the energy
is higher, this wavefunction usually extends to larger
distance. That is, as each photon is absorbed virtually, the
atom "swells" in extent. Clearly, ionization by real pho-
ton absorption or by tunnelling, can occur readily from
such extended and weakly-bound virtual states. A calcu-
lation of this process of virtual excitation followed by
tunnelling is the subject of this paper.

Already in 1988 the virtual multiphoton off-shell
atomic states were used to provide the first explana-
tion of electron angular distributions in ATI of the hydrogen
atom [10] by comparison with experiments of Feldmann
et.al. [11]. Unfortunately, although studied extensively
in unpublished work of Kracke [12] and used in ion-atom
collisions [13], no further discussion of the nature and
properties of such wavefunctions seems to have been pub-
lished.

The plan of the paper is as follows. In section II we
present a critique of the standard scattering theory used
to calculate the transition amplitude (T-matrix element)
to continuum states. We show that the transition am-
plitude to a continuum state can be represented in an
intuitively appealing way in that it appears as a direct
product of the amplitude for virtual $n$-photon absorp-
tion and the probability for subsequent tunnelling from
this virtual excited state. The results of calculations for
the realistic case of the three-dimensional hydrogen atom
are presented in section III. The virtual absorption wave-
function is calculated numerically by iteration of the in-
homogeneous Schr"odinger equation and for the tunnelling
wavefunction the quasi-static approximation is employed
using a separation in parabolic coordinates. The results
indeed exhibit a smooth and continuous transition from
optimum tunnelling directly from the ground state, to
dominant multiphoton ionization, as the Keldysh param-
eter is varied. Throughout we use atomic units in which
the electron charge, the electron mass and $\hbar$ are equal to
unity.

II. THE BASIC EQUATIONS

A. The strong-field approximation

The dynamics of ionization of an initially-bound elec-
tron in a strong laser field is essentially decided by the
competition between two electric fields; that of the par-
etent nucleus and that of the external laser. As such there
is great similarity with the theory of electron exchange
in ion-atom collisions where the two competing fields are
those of the two nuclei and involve two frames of refer-
cence, the laboratory frame of the parent nucleus and the
moving frame of the incident nucleus. Indeed "over the
barrier" ionization was first formulated for the ion-atom
problem. This analogy will emerge also in the formulas
presented here and perhaps casts a new light on the SFA.

We consider a total Hamiltonian

$$H(t) = H_i + V_F(t) = K + V + V_F(t) = H_F + V,$$  \hspace{1cm} (2)

where $K$ is the electron kinetic energy operator, $V$ is the
nuclear potential and $V_F(t)$ is the interaction between
electron and laser field (considered to be a classical field).
The electron wavefunction at time $t$ is given by solution of the equation

$$H(t)\Psi(t) = i\frac{\partial \Psi(t)}{\partial t}.$$ \hspace{1cm} (3)

The transition probability amplitude from an initial to a
final state at time $t$ can be expressed in two equivalent
post and prior forms, i.e.

$$f(t) = \langle \phi_f(t) | \Psi_i^+(t) \rangle = \langle \Psi_f^+(t) | \phi_i(t) \rangle.$$ \hspace{1cm} (4)

The two exact wavefunctions propagate forward in time
with $\Psi_f^+(t) \rightarrow \phi_i$ as $t \rightarrow -\infty$ and backwards in time with
$\Psi_f^+(t) \rightarrow \phi_f$ as $t \rightarrow \infty$ respectively. If one considers that
$\phi_i$ is an eigenstate of $H_i$ and $\phi_f$ is an eigenstate of $H_f$, then from the Schr"odinger equation one can show that

$$f(t) = \int_{-\infty}^{t} \langle \phi_f(t') | V | \Psi_i^+(t') \rangle \, dt'$$ \hspace{1cm} (5)

for the post form or

$$f(t) = \int_{t}^{\infty} \langle \Psi_f^+(t') | V_F(t') | \phi_i(t') \rangle \, dt'$$ \hspace{1cm} (6)

for the prior form. These two expressions are exact.

An approximation that has received much attention for
ionization is the SFA of Eq. (1). In the formalism of re-
arrangement given here, one notes that $\phi_f(t)$ is an eigen-
state of the (electron + field) Hamiltonian $H_f$. Hence
this is a Volkov state and the SFA is made simply by
replacing $\Psi_f^+$ in Eq. [5] by the initial state $\phi_i$. Interestingly, although often termed "non-perturbative" now the SFA appears as the first Born term for re-arrangement of the electron between eigenstates of the two potentials

$$f^{SFA} = \int_{-\infty}^{\infty} \langle \phi_f(t') | V | \phi_i(t') \rangle \, dt'.$$  \hfill (7)

As in general re-arrangement scattering [10], one can show that the equivalent first Born approximation putting $\Psi_f^+ \approx \phi_f$ in the prior form Eq. [6] is identically equal i.e.

$$f^{SFA} = \int_{-\infty}^{\infty} \langle \phi_f(t') | V_F(t') | \phi_i(t') \rangle \, dt',$$\hfill (8)

so that one can use either potential in the first Born SFA re-arrangement matrix element.

In the length gauge the Volkov state reads (in units with $c = \hbar = m = 1$ and $c = 137$)

$$\phi_f^v(r,t) = \exp[i(p + A(t)/c).r - i \int_0^t dt'(p + A(t'/c)^2)/2]$$ \hfill (9)

where $A$ is the vector potential. This is simply the Kramers-Henneberg space-translation plane wave [2] describing the electron stationary in the moving field. The additional exponential energy and momentum factors involving $A$ are identical to the "electron translation factors" appearing on final state wavefunctions in ion-atom electron capture, where the electron is also stationary in the moving field of the ion [23]. This justifies our view of the SFA as a collisional re-arrangement process in first Born approximation.

In approximate evaluations in collision theory, the time integrated forms Eq. [5], Eq. [6] of the transition amplitude are usually preferred to the direct projection forms Eq. [4]. Basically this is because, if $\phi_i$ and $\phi_f$ are orthogonal as is usually the case, the forms Eq. [1] give zero for the first-order amplitude whereas the integral forms Eq. [7] and Eq. [8] give a finite result.

By contrast, in numerically accurate propagations of the time-dependent wavefunction it is more direct to use the projection form Eq. [1]. This will be the strategy adopted in this paper and it removes a certain ambiguity in the physical interpretation of the SFA when the two equivalent forms Eq. [7] and Eq. [8] are used. In the form Eq. [7] one would say that ionization out of the initial state occurs by the electron scattering from its parent nucleus and then accessing the Volkov state, describing either tunnelling or absorption of photons depending upon the value of $\gamma$. However, the form Eq. [8] would be interpreted as an initial absorption of a single photon via $V_F$, followed by overlap on the same Volkov state. Which physical picture is correct ?

In the following we describe ionization in a unified way in that we approximate $\Psi_f^+$ in Eq. [5] essentially by a product of a state which initially has absorbed virtually a certain number of photons and a semi-classical state describing subsequent tunnelling in the full potential of the nuclear and laser electric fields. This describes a continuous transition from tunnelling to multiphoton regime according as to which element of the product states is dominant.

B. The approximate transition matrix element

We begin, not with the standard form Eq. [5] of the transition amplitude, but with the direct time propagation of Eq. [4]

$$f(t) = \langle \phi_f(t) | \Psi_i^+(t) \rangle = \langle p_f | U(t, -\infty) | \phi_i \rangle$$ \hfill (10)

where $|p_f\rangle$ is the final momentum state of the continuum electron, $|\phi_i\rangle$ is the initial atomic state and

$$\left(H(t) - i \frac{\partial}{\partial t}\right) U(t, t') = 0$$ \hfill (11)

is the full time propagator. Our approach will be to approximate the time development as occurring initially, up to a time $t_i$ with the laser field as a perturbation, followed by a propagation in the static field of the laser plus atomic Coulomb potential. That is, we write the full time-development operator as a product, 

$$f(t) = \langle p_f | U(t, t_i) U(t_i, -\infty) | \phi_i \rangle$$

$$= \langle p_f | U(t, t_i) | \psi(t_i) \rangle$$ \hfill (12)

where $|\psi(t_i)\rangle$ is an off-shell atomic state with photons absorbed virtually. This describes an initial polarization of the atom by the laser field. In the next section this state will be expanded in states in which a given number $n$ of photons has been absorbed virtually. The operator $U(t, t_i)$ then describes the subsequent tunnelling transition of the electron to a final ionized state. Then we write

$$\langle p_f | U(t, t_i) | \psi(t_i) \rangle$$

$$= \int d r_f \langle p_f | r_f \rangle \langle r_f | U(t, t_i) | \psi(t_i) \rangle.$$ \hfill (13)

Since the propagation through the tunnelling region and beyond as a continuum electron will be described subsequently by a semi-classical wavefunction, we will define ionization probability as given by the probability density $|\langle r_f | U(t, t_i) | \psi(t_i) \rangle|^2$ at a point $r_f$, corresponding to the exit from the tunnelling region. In appendix B it is shown that, as a result of the imaging theorem [20], this is equal to the transition probability $|\langle p_f | U(t, t_i) | \psi(t_i) \rangle|^2$.

The matrix element to be calculated can be written as an integral of the product of the transition amplitude to a virtual state multiplied by the tunnelling amplitude i.e.

$$\Psi^+(r_f t_f) \equiv \langle r_f | U(t_f, t_i) | \psi(t_i) \rangle$$

$$= \int \langle r_f | U(t_f, t_i) | r_i \rangle \langle r_i | \psi(t_i) \rangle \, dr_i$$ \hfill (14)

where we have introduced the kernel $K(r_f t_f, r_i t_i)$. 


III. THE OFF-SHELL WAVEFUNCTIONS

We consider ionization of a hydrogen atom i.e. we take $H_t \equiv H_0 = K + V$ where $V(r) = -\kappa/r$, the ionization potential $I_p = \kappa^2 / 2$ and nuclear charge $\kappa = 1$. The initial bound 1s ground state is

$$\psi_0(r) = \sqrt{\frac{\kappa^3}{\pi}} e^{-\kappa r}. \quad (15)$$

The first task is to calculate the virtual state $|\psi(t_i)\rangle$. We approximate the exact state by its lowest-order perturbation result. Hence, for this part of the ionization process the laser field is taken to be effectively a c.w. pulse. Then, the hydrogen atom is driven by a periodic circularly polarized laser with electric field $F(t) = E_0\hat{x}\cos[\omega t] + E_0\hat{y}\sin[\omega t]$. The time-dependent Schrödinger-equation then reads

$$H_0|\psi(t)\rangle - r \cdot \mathbf{F}(t)|\psi(t)\rangle - i \frac{\partial |\psi(t)\rangle}{\partial t} = 0 \quad (16)$$

Since the laser field is periodic the state vector can be expanded in a Floquet Fourier series

$$|\psi(t)\rangle = \exp[iI_p t] \sum_j |\psi_j\rangle \exp[-ij\omega t] \quad (17)$$

Inserting this expression into the Schrödinger equation, multiplying by $\exp[i\omega t]$ and integrating over all time yields

$$[H_0 + (I_p - n\omega)]|\psi_n\rangle = \frac{E_0}{2} \left( (x + iy)|\psi_{n-1}\rangle + (x - iy)|\psi_{n+1}\rangle \right) \quad (18)$$

This equation describes the population of the state $|\psi_n\rangle$ by absorption or emission of a photon from neighboring states. Since we consider the initial state as the ground state, in accordance with perturbation theory we retain only the lower state in the inhomogeneous term to give

$$[H_0 + (I_p - n\omega)]|\psi_n\rangle = \frac{E_0}{2} (x + iy)|\psi_{n-1}\rangle. \quad (19)$$

This is the inhomogeneous equation for the off-shell Coulomb wavefunctions. In [10] it was solved iteratively in a numerical procedure for the absorption of up to nine photons to calculate angular distribution of ATI continuum electrons. Here we restrict discussion to virtual states which are still bound. In appendix A we show how the inhomogeneous equation for the radial wavefunction is derived and solved. The method goes back to Dalgarno and Lewis in 1955 [21] and has been used often in early work on multiphoton ionization e.g. in [22, 23, 24]. The results are shown on Fig.1 for the radial density $|\mathcal{R}_n(r)|^2$ as a function of distance $r$ from the nucleus. Since the absolute magnitude of the virtual wavefunctions decrease strongly with $n$, we have normalized each magnitude to unity by defining $\mathcal{R}_n(r) \equiv R_n(r)/R_0(r_n)$, where $r_n$ is the position of the wavefunction maximum.

For circular polarization the orbital angular momentum quantum numbers $(l, m)$ are simply $l = m = n$.

The absorption of multiple photons is usually depicted as a vertical process in the atomic potential but the main feature of the off-shell wavefunctions shown in Fig.[1] is that, as the energy and angular momentum of the electron increases, the wavefunction has its maximum at larger and larger $r$ values. In Fig.1 to illustrate clearly the shift of the wavefunction from the nucleus, the modulus squared of each wavefunction for successive photon absorption has been normalized by dividing by its maximum value $\cdot$. The actual magnitude of the wavefunction decreases with each iteration due to the $E_0$ factor. Of course, the shift of the wavefunction to larger distances as binding energy decreases is also a feature of the on-shell eigenstates of the hydrogen atom. The effect is amplified here by the dipole operator in the inhomogenous term. The important point for subsequent tunnelling is that this virtual wavefunction has significant amplitude in the tunnelling region of the combined atomic and laser electric field potential. As shown in the next section, this leads to enhanced tunnelling out of virtually-excited states compared to that from the ground state.

It is interesting to compare the energy gain by virtual photon absorption treated here, with the energy gain calculated in the “non-adiabatic” tunnelling picture of [14]. Since the former is calculated in quantum picture and the latter in classical mechanics, the quantities to be compared are somewhat arbitrary. However, to be precise, we plot the effective total radial energy of the quantum case against the total energy of the classical case, both evaluated in the tunnelling direction $x$. That is, we plot $-(\kappa n^2 + n(n + 1)/r^2)/2$ (see Eq. (A3)) at the maximum value of the wavefunction against the energy $(i^2/2 - rE_0 - \kappa/r)$ for the classical energy gain [14]. This comparison is shown in Fig.2.

In the cases $\gamma = 0.66$ and $\gamma = 2.0$, Fig.2 a and b, with low frequency $\omega = 0.025$ a.u. there is reasonable quantitative agreement but good qualitative agreement. For $\gamma = 0.66$, the frequency implies that the energy increase,
due to five photons absorbed, is small on the energy scale shown. However, the quantum calculation shows the wavefunction penetrating into the tunnelling region as photons are absorbed. The same is true for the multi-photon ionization regime $\gamma = 6.0$ shown in Fig. 2c. Here one observes 16 photons absorbed virtually with the energy increasing as a function of position in almost exactly the same way as in the classical calculation. Of course many photons corresponds to the classical limit but the close agreement of the two estimates of energy versus position is quite noteworthy.

IV. THE TUNNELLING WAVEFUNCTION

The second task is the calculation of the tunnelling probabilities for different starting values of $r_i$ in Eq. (20) and integration over $r_i$ to obtain total ionization probability. Eq. (20) we calculate the wavefunction at tunnel exit

$$\Psi^+(r_f) = \int K(r_f, r_i) \psi_n(r_i) \, dr_i$$  \hspace{1cm} (20)$$

The tunnelling wavefunction i.e. the kernel as a function of $r_i$, with the (static) field of strength $E_0$ in the $x$ direction, satisfies the Schrödinger equation

$$\left( -\frac{\Delta}{2} - \frac{1}{r} - xE_0 \right) K = -\frac{\kappa_n^2}{2} K$$  \hspace{1cm} (21)$$

where $\kappa_n^2 \equiv 2(I_p - n\omega)$. Following [17] and [18] the tunnelling wavefunction is calculated in parabolic coordinates $(\eta, \xi, \phi)$ with $x = (\eta - \xi)/2$, $y = \sqrt{\eta} \cos \phi$, $z = \sqrt{\eta} \sin \phi$ since the equation separates in these coordinates. The tunnelling is described by the $\eta$ equation and, as shown in detail in [19], can taken to the lowest order in $h$ i.e. the WKB solution. This gives the semi-classical tunnelling wavefunction. Of course the virtual wavefunction $\psi_n(r_i)$ in Eq. (20) is calculated in spherical coordinates with the $z$ axis perpendicular to the plane of polarization. Then, for circular polarization, the state with angular momentum $l = n$ is populated and with the highest $m = n$ value. The electron density in the excited state is aligned in the $xy$ plane and correspondingly we take $m = 0$ only with respect to the parabolic $\phi$ dependence, which gives optimum tunnelling [19]. With these approximations the function $K(r_f, r_i)$ can be calculated and the integral over $r_i$ in Eq. (20) performed numerically. Note that, in parabolic coordinates, the integrand is exactly of the form considered by Landau and Lifshitz [17] for ionization from the ground state. In their calculation they simply assumed a particular starting point $r_i$. Here we have performed the integral over all $r_i$.

The final ionization probability is a product of the two competing processes of multi-photon absorption and under-the-barrier tunnelling. In perturbation theory, the $n$-photon wavefunction $\psi_n(r_i)$ contains a time-dependent phase factor $\exp[i(I_p - n\omega)t_i]$. Hence we will treat each $n$-photon state separately corresponding to a different final energy. Also, to make the calculation of tunnelling probability tractable, as in [14], we will describe tunnelling in the static electric field at a time corresponding to the maximum of the field strength. Then the differential ionization probability out of a state with $n$ photons absorbed virtually is time independent

$$\frac{dP_n}{d\theta_f} = \left| \Psi^+_n(r_f) \right|^2 = \left| \int K(r_f, r_i) \psi_n(r_i) \, dr_i \right|^2.$$  \hspace{1cm} (22)$$

FIG. 2. The total potential in the $x$ direction. The dots indicate the position of the wavefunction maximum as a function of energy for increasing number of photons. The dashed line is the classical energy gain.
The $n$-photon absorption probability decreases as $E_n^{2n}$ whereas the tunnelling probability increases exponentially in $(I_p - n\omega)$. Below, we consider a fixed frequency, low enough to justify the quasi-static tunnelling approximation but requiring many photons to be absorbed to reach the ionization threshold. Varying $\gamma$ then corresponds to varying field strength. The maximum value of field strength determines the height of the potential barrier for tunnelling and so has a decisive effect on the tunnelling probability. The competition between the probability to access a state by photon absorption and the probability to tunnel out from that state decides the dominant mode of photo-ionization. The results illustrate this influence of the laser field strength on the ionization mechanism and are presented in Fig.3.

V. THE IONIZATION PROBABILITIES

In Fig.3 we show the ionization probabilities as a function of the number of virtually-absorbed photons, up to an energy corresponding to the top of the potential barrier. To illustrate the relative probabilities for ionization from each virtual state, we plot the quantity

$$P_n^{rel} = \frac{|\Psi_n^+(r_f)|^2}{|\Psi_{max}^+(r_f)|^2}$$

(23)

for each $n$, where $|\Psi_{max}^+|^2$ corresponds to the $n$ value giving maximum ionization probability and $r_f$ is the tunnel exit. In all cases, $\omega$ is fixed at a value 0.025 a.u. Direct photo-ionization corresponds to absorption of 20 photons. For small $\gamma$ equal to 0.66, shown in Fig.3a and corresponding to a field strength of 0.04 a.u., ionization occurs most probably directly out of the ground state. Principally this is because the height and width of the potential barrier falls with increasing maximum field strength and tunnelling probability depends upon it exponentially. Here the barrier is such that tunnelling can take place from the ground state. Although tunnelling from higher energy states is even more probable, this is more than offset by the reduced probability of photon absorption, leading to a monotonic decrease of ionization probability as a function of the number of photons absorbed.

By contrast, at high value 6.0 of $\gamma$, Fig.3c, the picture is quite different. In this case the field strength is only 0.003 a.u. and the height of the barrier suppresses tunnelling strongly from the lower-energy states. Here one sees a monotonic increase of probability with photons absorbed corresponding to enhanced tunnelling out of successively higher-energy states. Clearly this limit corresponds to direct multiphoton ionization, as in the ATI calculations of [10]. Paradoxically, it is the increasing tunnelling rate that leads to the increase of ionization probability with photon number, what is normally referred to as the multiphoton ionization limit. As field strength decreases this leads in turn to the most probable transition being due to no tunnelling at all, i.e. over-the-barrier release of electrons.

Note that we are comparing always relative probabilities as a function of photons absorbed. Since the field strength is ten times lower, the absolute probabilities are lower for $\gamma = 6.0$ than for $\gamma = 0.66$ due to the lower field strength for fixed photon frequency.

As one might expect, intermediate values of $\gamma$ interpolate between these two limits and exhibit preferential tunnelling from virtual excited states which are still below the top of the barrier. For the higher-energy states the probability falls again even though tunnelling is enhanced. The example $\gamma = 2.0$, corresponding to field strength 0.01 a.u. is shown in Fig.3b. Here the calculation predicts that a maximum ionization probability occurs for three photon absorption, followed by tunnelling. The three panels of Fig.3 illustrate nicely the cross-over from ground-state tunnelling to multiphoton ionization.

![Fig. 3. The relative probabilities of ionization as a function of increasing numbers n of virtually-absorbed photons of frequency 0.025 a.u.](image-url)
without tunnelling.

The same pattern emerges for different values of $\omega$, although for higher frequencies, since for hydrogen $\gamma = \omega/E_0$, the cross-over to predominant ground-state tunnelling (and even over-the-barrier escape) occurs at values of $\gamma$ exceeding unity. This is illustrated in the subsequent figures where we consider the two cases of $\omega = 0.05$ a.u. and 0.1 a.u.

In Fig 4a, we show the relative probabilities for $\omega = 0.05$ a.u. and $\gamma = 1.0$ giving $E_0 = 0.05$ a.u.. Already for this field strength there is most probable tunnelling out of the ground state and the multiphoton excitation probability falls off rapidly for $n = 1$. At intermediate $\gamma = 2.0$, shown in Fig 4b, the most probable tunnelling has shifted to $n = 1$. At higher $\gamma = 8.0$, Fig 4c, corresponding to field strength $E_0 = 0.006$, the transition to preferential multiphoton ionization has been made completely. The same is true for $\omega = 0.10$ and $\gamma = 8.0$

shown in Fig 5. However, for this frequency, the field is $E_0 = 0.012$ and only the virtual absorption of two photons is necessary to reach the top of the barrier. In the case of $\omega = 0.10$ (not shown), already at $\gamma = 2.0$, the field $E_0 = 0.05$ is such that the barrier is so low that ionization occurs over the barrier after the absorption of just one photon and there is essentially no tunnelling.

VI. CONCLUSIONS

We have derived a simple intuitive expression, Eq. (14) describing the ionization of the hydrogen atom by a classical laser field as consisting of two steps. The ionization amplitude then factors into a product of the separate amplitudes of the two steps occurring. The first step is a polarization of the atom and energy increase of the electron due to the virtual absorption of photons. The second step is the tunnelling of the virtually-excited electron out of the total (atom + field) static potential leading to ionization.

The virtual absorption of photons leads to the electron gaining energy as it recedes from the nucleus and this mechanism supports the supposition of Klaiber et.al. [14] who described the “non-adiabatic” energy gain by classical mechanics. Indeed, there is close agreement, for large number of photons absorbed, between the energy gain predicted in our quantum perturbation theory and that ascribed to classical motion.

The results for the relative ionization probabilities as a function of the number of virtually absorbed photons are presented for fixed laser frequencies but low enough as to require many photons to be absorbed to reach the ionization threshold. The Keldysh parameter $\gamma$ then is inversely proportional to the peak field strength and this decides the position and value of the peak of the tunnelling potential that is decisive for the tunnelling probability. The results demonstrate a continuous smooth transition between the two limits in which the maximum probability is associated with direct ground-state tunnelling for the higher strengths and complete multiphoton
field ionization can set in for \( \gamma \)
the tunnelling region and indeed direct over-the-barrier
manuscript.

The transition is indicated schematically in Fig. 6 which
emphasizes that, contrary to the usual depiction of a
vertical transition in space, as the electron absorbs en-
ergies by virtual photon absorption, the atomic wavefunc-
tion swells in size. Schematically and following tradition,
the atomic wavefunction is given in Eq. (19)

\[
\Psi_{\text{off}}(r) = \frac{1}{\sqrt{2 \kappa \gamma}} \left( e^{i \phi} \psi_n - \kappa \gamma \psi_{n-1} \right)
\]

However, as we have seen in the example of \( \omega = 0.10 \) a.u.,
the tunnelling region and indeed direct over-the-barrier
field ionization can set in for \( \gamma \) values greater than unity.

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Appendix A: Off-shell Coulomb wavefunctions

The inhomogeneous equation for the off-shell Coulomb
wavefunction is given in Eq. (19)

\[
(H_0 + (I_p - n \omega)) | \psi_n \rangle = \frac{E_0}{2} (x + iy) | \psi_{n-1} \rangle.
\]

In spherical coordinates this is

\[
\frac{1}{r^2} \nabla^2 \psi_n + \left( \frac{I_p}{r} - n \omega \right) \psi_n = \frac{E_0}{r^2} \sin(\theta) e^{i \phi} | \psi_{n-1} \rangle.
\]

Projecting on spherical harmonics yields an iterative
equation for the radial part of the excited states \( \psi_n \). With
\( \psi_0(r) = R_0(r) Y_{0,0}(\theta, \phi)/r \) and \( R_0(r) = 2 \sqrt{3} r \exp(-r \kappa) \)
the first order equation reads:

\[
R_1''(r) - \frac{2 R_1(r)}{r^2} + \left( \frac{I_p + \omega + \kappa}{r} \right) R_1(r) = -\frac{2}{3} \sqrt{\frac{2}{3}} \frac{E_0}{r} R_0(r).
\]

The nth order equation to be solved iteratively is

\[
R_n''(r) + \left( -\kappa_n^2 - \frac{n(n+1)}{r^2} + \frac{2 \kappa}{r} \right) R_n(r) = -\sqrt{\frac{2n}{2n+1}} r E_0 R_{n-1}(r),
\]

where \( \kappa^2_n \equiv 2(I_p - n \omega) \).

The homogeneous equation has two solution functions
that are:

\[
y_1(r) = W_{\frac{\kappa}{\kappa_n}, n + \frac{1}{2}} (2r \kappa_n) \]
\[
y_2(r) = M_{\frac{\kappa}{\kappa_n}, n + \frac{1}{2}} (2r \kappa_n),
\]

where \( M \) and \( W \) are the Whittaker-functions. Taking
into account the asymptotic behavior of these functions
the nth off-shell wavefunction in the 3D-Coulomb potential
will be given via the expression for the radial functions

\[
R_n(r) = y_1(r) \int_0^r dz f(z) y_2(z) \frac{W(z)}{W(z)}
\]
\[
+y_2(r) \int_r^\infty dz f(z) y_1(z) \frac{W(z)}{W(z)},
\]

with the Wronskian
\( W = y_1 y_2' - y_2 y_1' \).

Appendix B: Imaging Theorem

With a final measured momentum state \( |p\rangle \) the prob-
ability amplitude \( f(t) \), the projection on the exact time-
propagating state, can be written \( f(t) = \langle p \mid \Psi^+(t) \rangle \equiv \Psi^+(p, t) \). One notes that for a free electron this is just
the Fourier transform of the exact spatial wavefunction. However, it is defined more generally e.g. in an asympto-
tic Coulomb potential. The imaging theorem (IT) \( [20] \)
shows that in the region where the semi-classical wave-

function is valid, the amplitude \( \Psi^+(p, t) \) can be related to the position wavefunction \( \Psi^+(rt) \) of Eq. (14),
which is the quantity we calculate. Specifically, the IT equates
the probabilities,

\[
\left| \Psi^+(p, t) \right|^2 \, dp = \left| \Psi^+(r, t) \right|^2 \, dr
\]

at all points connecting the momentum \( p \) with position \( r \)
and vice versa along a classical trajectory. We have put
\( r = rf \), the position corresponding to the barrier exit for
each virtual state energy. Thereby we equate the proba-
bility density in position space with the probability den-
sity for projection onto a momentum state. This means
that we have assumed that semi-classics is valid immedi-
ately following the electron’s transition to a continuum
state. This is an approximation but is compatible with our
use of semi-classics to describe the under-the-barrier motion. The IT also lends credence to the strategy of
Ni et.al. \( [9] \) who use classical mechanics to propagate
numerically-calculated probability densities backwards in
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