Twisted K-Theory as a BRST Cohomology

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Abstract

We use the BRST formalism to classify the gauge orbits of type II string theory’s Ramond-Ramond (RR) field strengths under large RR gauge transformations of the RR gauge potentials. We find that this construction is identical to the Atiyah-Hirzebruch spectral sequence construction of twisted K-theory, where the Atiyah-Hirzebruch differentials are the BRST operators. The actions of the large gauge transformations on the field strengths that lie in an integral lattice of de Rham cohomology are found using supergravity, while the action on $\mathbb{Z}_2$ torsion classes is found using the Freed-Witten anomaly. We speculate that an S-duality covariant classification may be obtained by including NSNS gauge transformations and using the BV formalism. An example of a $\mathbb{Z}_3$ torsion generalization of the Freed-Witten anomaly is provided.

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1 Introduction

It was once believed that, in the absence of D-branes, Ramond-Ramond field strengths were classified by twisted K-theory \cite{1}. Evidence for this conjecture came from an analysis of symmetric boundary conditions in the worldvolume theories of open strings \cite{2}, from the analysis of the Chan-Paton bundles on various unstable D-branes \cite{3} and from the conditions imposed on D-brane embeddings by global worldsheet anomaly cancellation \cite{4}. However within a few years, Diaconescu, Moore and Witten realized \cite{5} that, in type IIB string theory, the twisted K-theory classification of fluxes is inconsistent with S-duality and so is incorrect.

With hindsight this was not surprising, as none of the derivations mentioned above is covariant with respect to S-duality. For example, searching for D-branes as boundary conditions in the conformal field theories will not yield an S-duality covariant classification of branes unless one also includes NS branes and examines the boundary conditions of \((p,q)\)-strings, which would be difficult as the fundamental string coupling diverges near an NS5-brane. Witten’s construction of untwisted K-theory from tachyon condensation \cite{6} of stacks of D9 and anti D9-branes also could not be S-dualized as D9-branes are already poorly understood, and their S-duals are unknown if they exist at all. Freed and Witten’s worldsheet anomaly was easily S-duality covariantized \cite{7}, however except in the case of 3-form field strengths, it was not known how to use it to find a consistency condition for the field strengths.

Some authors have suggested that this inconsistency should be resolved by replacing twisted K-theory with an entirely different generalized cohomology theory \cite{10}. Instead, in this note we suggest a new approach to this old problem, using the BRST formalism. In particular, we identify a set of large Ramond-Ramond gauge transformations of the type II supergravity action, essentially those which keep the Wilson loops invariant, and we show that they act nontrivially on the field strengths. These symmetries act on the integral lattice of de Rham cohomology which satisfies the Dirac quantization condition. This lattice is often smaller than the full integral cohomology, in which the various field strengths are believed to be valued. We use the Freed-Witten anomaly to extend the gauge transformations to the full integral cohomology.
We then make a simple observation, which is the central result of this paper. The BRST cohomology of the RR gauge transformations is isomorphic, as a set, to twisted K-theory. In particular Atiyah and Hirzebruch have shown that K-theory may be constructed from integral cohomology by taking its cohomology with respect to a series of differential operators. In other words, a quotient of a subset of cohomology approximates K-theory, and a quotient of a subset of that is a better approximation, and so on. We argue that Atiyah and Hirzebruch’s differential operators are precisely the BRST charges corresponding to the large RR gauge transformations. In addition the extra, wrong degree, cohomology classes that one adds to the sequence are the ghosts and antighosts. This is an unusual application of the BRST formalism because the gauge symmetry that we consider is discrete, and so in particular the ghosts and antighosts do not enjoy propagating degrees of freedom.

This all suggests that to find the S-duality covariant classification of RR and NSNS field strengths in IIB, and also the correct classification in IIA, one needs to take the BRST cohomology not only with respect to the RR gauge transformations but also with respect to the NSNS gauge transformations. This is not easy, as the NS 3-form field strength is itself used in the RR gauge transformations, and as a result these two symmetries cannot be disentangled. The result appears to be that one does not find that the collection of NSNS and RR field strengths together is a cohomology or even an additive group. This is no surprise, as the field strengths are solutions to supergravity equations of motion which are nonlinear. Quotienting by the NSNS transformations one loses more than just the addition however, the S-duality covariant classification in general often has a different cardinality than the twisted K-theory, as was shown in the case of the Klebanov-Strassler geometry in [11].

We hope that the BRST construction of twisted K-theory will have other applications. For example, Witten has speculated [12] that there may be a twisted K-theory based formulation of type II string theories. However efforts at constructing such a formulation have been complicated by the fact that twisted K-theory classes are not easily expressed as fields, and so not easily treated using the standard tools of quantum field theories. However using the current construction one need only start with fields that are ordinary differential forms, supplemented with the Dirac quantization conditions and the various ghost towers that construct a Deligne cohomology and allow the description of torsion classes. That is, the field content is the same as in p-form gauge theory. Then one finds the gauge symmetries of the path integral, includes ghosts and calculates the BRST cohomology as usual.
This yields a BRST cohomology which is larger than twisted K-theory. To obtain twisted K-theory, one must then ignore the values of the various Wilson lines, which are exponentials of the integrals of the gauge connections. Alternately, one can choose not to ignore these and one would arrive at a differential version of twisted K-theory. It would be interesting to see if this differential K-theory agrees with that proposed by Freed in [13].

The description of the twisted K-theory of the spacetime or spatial slice $M$ as a BRST cohomology is summarized in the following table. For concreteness we consider type IIA. Stable, consistent D-branes correspond to anomalies, despite the fact that

| BRST | K-theory | IIA Supergravity | IIA String Theory |
|------|----------|-----------------|------------------|
| fields | $\bigoplus_k H^{2k}(M)$ | RR diff. forms $G_{2k}$ | RR int. classes $G_{2k}$ |
| ghosts | $\bigoplus_k H^{2k+1}(M)$ | gauge xforms/branes | gauge xforms/branes |
| BRST operator | $d_{2p+1}$ | $d_3 = H \wedge$ | $d_3 = Sq^3 + H \cup$, $d_5$ |
| Gauge xforms | $d : H^{\text{odd}} \to H^{\text{even}}$ | Wilson loop shift | FW monodromy |
| Constraints | $d : H^{\text{even}} \to H^{\text{odd}}$ | Bianchi identities | source FW anomaly |
| physical fields | $K^0_H(M)$ | orbits of Bianchi solns | orbits of FW solns |
| anomalies | $K^1_H(M)$ | stable p-branes | stable D-branes |

Table 1: K-theory vs BRST in type IIA supergravity and in type IIA string theory

their partition functions are gauge invariant. Indeed they appear in the supergravity equations of motion as violations of the conservation of RR current. They also correspond to anomalies in the sense of Ref. [14] as instantonic D-branes which form, sweep nontrivial cycles and then decay change the cohomology class of the fluxes that they source, and therefore mediate a tunneling between different source-free RR solutions [15]. Notice in particular that on their worldvolumes the unimproved field strengths are not defined and so the gauge transformations are not defined, similarly to the worldvolume of a magnetic monopole in QED where the gauge potential $A$ is not defined. In the quantum theory this may correspond to a nontrivial inner product between states with stable D-branes and states with Ramond-Ramond fields that are pure gauge. The fact that D-branes correspond to elements of the BRST cohomology and in particular are BRST closed implies a sort of Wess-Zumino consistency condition which is enforced classically by the supergravity equations of motion and quantum mechanically by the Freed-Witten (FW) anomaly.
In section 2 we will use the democratic formulation [16, 17, 18] of classical supergravity to derive the large RR gauge transformations of the components of the RR field strengths that are valued in an integral sublattice of the de Rham cohomology. In section 3 we will describe Atiyah and Hirzebruch’s construction of K-theory from cohomology, and show that restricted to a lattice of de Rham cohomology this yields the gauge-invariant field strengths found in the supergravity analysis. Finally, in section 4 we will use Maldacena, Moore and Seiberg’s interpretation [15] of the Freed-Witten anomaly to extend the action of the large RR gauge transformations to the full integral cohomology, that is, we will find the action on torsion-valued fields. We will argue that the corresponding BRST operator is the Atiyah-Hirzebruch differential on both nontorsion and $\mathbb{Z}_2$ torsion cohomology classes. In type IIB the Freed-Witten anomaly needs to be generalized to capture the $\mathbb{Z}_3$ torsion contributions. An example of a brane with a $\mathbb{Z}_3$ generalized Freed-Witten anomaly is provided.

2 Large gauge transformations

In this section we will review the large gauge transformations of the gauge potential in QED, these are gauge transformations that shift the integrals of the gauge potential by integers and so leave the Wilson loops invariant. We will then find the corresponding large RR gauge transformations in the democratic formulation of type II supergravity theories and argue that, unlike the case of QED, these transformations also change the field strengths. We present a classification of gauge equivalence classes of those field strengths that satisfy the RR supergravity equations of motion and Dirac’s quantization condition.

2.1 A Warm up: QED

In the absence of charged matter, Maxwell’s theory of electromagnetism in any number of dimensions is described by the Lagrange density

$$\mathcal{L} = F \wedge \star F$$

(2.1)

where $F$ is a 2-form field strength containing the electric and magnetic field components. As there is no matter, this field strength is the only observable. To define the classical theory we need to supplement this Lagrangian with an additional equation,
for example we may impose that the field strength is closed
\[ dF = 0. \]  \hspace{1cm} (2.2)

In this case on any local patch of spacetime we may define a one-form gauge potential \( A \) by demanding that
\[ F = dA. \] \hspace{1cm} (2.3)

Then Eq. (2.2) becomes a Bianchi identity for \( A \), which is the condition that \( A \) is locally well-defined and so is annihilated by the square of the exterior derivative \( d \).

Eq. (2.3) does not uniquely specify the gauge potential \( A \), but only specifies it up to an additive shift \( \eta \) which is called a gauge transformation. One often says that \( \eta \) may be any exact one-form. However, if the spacetime topology is nontrivial then in the absence of charged matter \( \eta \) is not necessarily exact but only needs to be closed: \( d\eta = 0 \). For example, if the spacetime is a circle parameterized by the coordinate \( \theta \) then one may shift \( A \) by \( \eta = cd\theta \)
\[ A \mapsto A + \eta = A + cd\theta \]  \hspace{1cm} (2.4)

where \( c \) is any nonzero constant. \( \eta \) is then closed but not exact. The transformation (2.4) leaves invariant the field strength \( F \) and therefore also the Lagrange density \( L \), identifying (2.4) as a gauge transformation.

If one includes electrically charged matter in the theory then the set of allowed gauge transformations is restricted. For example, one may couple the theory to a conserved current \( J \) of particles with charge \( q \) by introducing the following Lagrange density
\[ L = F \wedge *F + qJ A. \] \hspace{1cm} (2.5)

Again one may consider gauge transformations of the form \( A \mapsto A + \eta \) and try to find the values of \( \eta \) for which the Lagrange density is invariant. The current \( J \) is an arbitrary closed 3-form and so it is necessary that the gauge transformations separately preserve both terms in the Lagrangian (2.5). We have already found that \( \eta \) needs to be closed to preserve the first term. The second term depends explicitly on \( A \) and so will not be invariant.

We are searching for a condition on \( \eta \) such that both the field strength \( F = dA \) and the action \( S = \int_M L \) are gauge-invariant. In particular, for an orientable spacetime \( M \) one may use the fact that \( J \) is closed for any conserved current to perform the integral
over all spacetime directions except for one, leaving $S = q \int_N A$, where $N \subset M$ is the Poincaré dual of $J$, which is the worldline of an electrically charged particle. The action will therefore be invariant precisely when $\eta$ is exact.

We have imposed too strong of a condition on $\eta$. The action of a quantum theory is not an observable, and does not need to be well-defined. Instead only the path-integral measure
\[ e^S = e^{\int_M F \wedge \star F} e^{q \int_N A} \] (2.6)
must be well-defined, where the first factor on the right hand side is well defined when $\eta$ is closed. This means that those one-forms $\eta$ that shift $q \int_N A$ by an integer are also gauge transformations. In other words $\eta$ does not need to be exact, but may represent any cohomology class that lies on an integer sublattice $\mathbb{Z}^k$ of the first real cohomology of the spacetime. For example, if the spacetime is a 2-torus $T^2$ then the allowed gauge transformations, up to an exact one-form, are
\[ \eta = m d\theta + n d\phi \in \mathbb{Z}^2 \subset H^1(T^2) = \mathbb{R}^2 \] (2.7)
where $m$ and $n$ are integers and $\theta$ and $\phi$ are the coordinates of two cycles that span the torus. $\mathbb{Z}^2$ is then said to be the group of large gauge transformations. These gauge transformations are called “large” because only the identity itself is in the same connected component of the gauge group as the identity.

These large gauge transformations are used, for example, in the construction of the Dirac string. The gauge potential $A$ is not globally defined on a 2-sphere that links a magnetic monopole, however it may be expressed on two patches, one of which intuitively contains the intersection of the Dirac string with the sphere. On the overlap of the two patches the value of $A$ is defined on both patches, but the values on the two patches do not agree. Instead the values on the overlap are related by a gauge transformation which is a large gauge transformation. In this case the large gauge transformation is a transition function which is only defined on the overlap, therefore it is classified by the first cohomology of the overlap, which is related to the second cohomology of the whole spacetime via the Mayer-Vietoris sequence. It is the second cohomology of the spacetime that classifies nontrivial gauge bundles, such as the bundle on the sphere linking a monopole.

An alternate interpretation of the large gauge transformations is as follows. The second factor in the action (2.6) is a Wilson loop, normalized by the electric charge of the fundamental particles. In general, transformations of $A$ by closed forms $\eta$ change
the Wilson loops, which leads to a measurable Berry’s phase when two charged identical particles encircle N in different directions and interfere. Gauge transformations, by definition, leave the observables invariant. Thus the Berry’s phase must be invariant, and so the Wilson loop must be invariant. This is the case precisely for gauge transformations on the above integral lattice.

2.2 Torsion gauge transformations

This subsection is more difficult to read than the previous and one may skip it and still understand the main idea of sections 2 and 3.

The main idea in this subsection, and in section 4 is that the field content of string theory lies not only in the values of the supergravity fields, but also in choices of phases in the path integral, which are invisible in supergravity. While closed differential forms that satisfy Dirac’s quantization condition are classified by an integer lattice of de Rham cohomology, both fields and phases are classified by integral cohomology. An integral cohomology class $H^p(M)$ of the space $M$ may always be written as the direct sum of two terms. The first term is called the free part, and is an integer lattice $\mathbb{Z}^k$ where $k$ is called the $p$th Betti number and is equal to the rank of the $p$th de Rham cohomology group

$$H^p_{\text{de Rham}}(M) = \mathbb{R}^k, \quad H^p(M) = \mathbb{Z}^k \oplus \mathbb{Z}^{q_i}. \quad (2.8)$$

The torsion part is the direct sum of cyclic groups $\mathbb{Z}_{q_i}$, which are the additive groups of integers modulo $q_i$. The supergravity field data is then classified by the free part of the integral cohomology, which is isomorphic to an integer lattice of the de Rham cohomology. However in the full quantum theory we must also consider the torsion part of the cohomology, corresponding to the aforementioned phases. The goal of sections 4 will be to extend the action of our BRST operator to the torsion part of the integral cohomology.

Including torsion terms, there are more gauge transformations than just the integral lattice described above. It is believed that there is a configuration of the gauge field corresponding to every topological circle bundle with connection $A$. The topologies of circle bundles are in one to one correspondence with classes of the second integral cohomology via an isomorphism called the first Chern class. In particular, if $A$ is considered, as usual, to be a one-form on each coordinate patch, then the transition functions of $A$ do not necessarily uniquely identify the topology of the gauge
bundle. Physically this corresponds to the fact that the sector of a gauge theory is not determined entirely by the transition functions of $A$, but also by the transition functions of the charged matter.

For example, consider a spacetime which is topologically the projective space $\mathbb{RP}^2$, which is the quotient of the two-sphere $S^2$ by the $\mathbb{Z}_2$ antipodal map. This spacetime needs to be Euclidean, but the signature is not important for these considerations and one could think of the $\mathbb{RP}^2$ as a timeslice in a Minkowski spacetime. The second integral cohomology group of $\mathbb{RP}^2$ is the cyclic group of order two

$$H^2(\mathbb{RP}^2) = \mathbb{Z}_2$$

and so we expect this theory to contain two topological sectors corresponding to the gauge bundles whose Chern classes are the two elements in (2.9).

$\mathbb{RP}^2$ can be visualized as the northern hemisphere of the two-sphere $S^2$ with a 2-fold antipodal identification of the equator. Then as one descends southward to the equator in the eastern hemisphere, one finds oneself traveling northwards from the equator in the western hemisphere. A QED configuration on this spacetime consists of a circle bundle over the $\mathbb{RP}^2$, and the charged matter fields are sections of an associated bundle. These bundles can be trivialized over the northern hemisphere, thus all of the topological information is contained in the transition function on the equatorial circle, that is, one needs to know what happens to the field as one passes through the equator and jumps to the other side of the sphere. The first de Rham cohomology group is trivial, therefore all closed one-forms $\eta$ are also exact and so correspond to transitions of $A$ that may be continuously deformed away. Thus the information about the topological sector of the theory is not captured by a one-form $\eta$ as in the previous subsection.

Not all transformations of the matter field may be continuously deformed away. In particular, one may consider a gauge transformation in which the matter field changes sign as one crosses the equator, which is similar to the sign change of a fermion when one rotates once all of the way around the $SO(3) \cong \mathbb{RP}^3$ subgroup of the Lorentz group. No more general phase rotation is possible, as the fundamental group of $\mathbb{RP}^2$ is $\mathbb{Z}_2$ and so a loop that crosses the equator twice may be deformed away and so cannot have a topological phase rotation, thus the square of the transition function is the identity. This transformation is nontrivial, and corresponds to the only nontrivial circle bundle on $\mathbb{RP}^2$, that is, to the bundle with Chern class equal to the element $1 \in \mathbb{Z}_2$ in (2.9) and not the identity element $0 \in \mathbb{Z}_2$.  

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The $\mathbb{RP}^2$ example illustrates that in general there are physically inequivalent field configurations which do not correspond to any transition functions of $A$. In fact, the field configurations always correspond to the second integral cohomology. The free part can described by locally defining the gauge potential $A$ and then attaching these patches using transition functions $\eta$ in the integral lattice described above. The torsion terms $\mathbb{Z}_k$, on the other hand, are defined to be choices of phase in the path integral depending on the topology of the nilpotent world line of the electric particle. A nilpotent world line $c$ is one that corresponds to a torsion term $\mathbb{Z}_k$ in the first integral homology group, which means that the trajectory $kc$, which winds around the loop $c$ $k$ times, is a boundary and so may be deformed away. For example, if one encircles the noncontractible circle in $\mathbb{RP}^3$ twice, then one traces out a contractible path.

The different possible phases in the path integral need to be dual to these $\mathbb{Z}_k$’s, and the universal coefficient theorem implies that the torsion part of the second cohomology is equal to the torsion part of the first homology and so classifies these phases. In particular, if the trajectory encircles a $k$-nilpotent cycle $j \in \mathbb{Z}_k$ times, and one is in the $i$th topological sector of the gauge theory, then the phase is given by $ij$.

Each choice of the path integral phases corresponds to a bundle which in turn corresponds to a choice of transition function of the gauge or matter fields on the overlaps of pairs of patches. Thus each bundle corresponds to a gauge transformation and the classification of bundles by first Chern classes is also a classification of gauge transformations. Chern classes, as in the $\mathbb{RP}^2$ case, are not necessarily valued in the integral sublattice of the real cohomology, which is isomorphic to the free part of the integral cohomology. But rather, Chern classes are arbitrary elements of the integral cohomology. This suggests that the group of possible gauge transformations is not only the free part of the integral cohomology described in the previous subsection, but apparently is the full integral cohomology. Of course we do not claim that this observation is new, it goes back at least 20 years [19].

Thus the gauge field configurations are classified by the first Chern class which is an arbitrary element of the second integral cohomology whose the free part is captured by a two-form called the field strength. In what follows, we will often refer to the entire integral class as the field strength despite the fact that the torsion information is encoded in the transition functions of the matter fields and not in the 2-form $F$ itself.
We will be interested in gauge transformations that are defined on the whole spacetime, and not just on overlaps. Such gauge transformations correspond geometrically to changes of trivializations of bundles. It is these transformations that relate physically equivalent configurations. The above argument suggests that these are classified by the classes in the full first integral cohomology. Due to the universal coefficient theorem, this is isomorphic to the integral lattice of de Rham cohomology that we considered in the previous subsection. However this isomorphism fails at the higher dimensions that will be relevant to our discussion of supergravity below. While these gauge transformations can change the gauge connection and some phases in the definition of the partition function, they do not affect the field strength in QED. On the other hand, in IIA supergravity we will see that the field strengths are no longer invariant under the large gauge transformations. We will argue that the gauge-invariant quantities are not differential forms, but rather they are classes in twisted K-theory.

2.3 Type II supergravity

For concreteness we will restrict our attention to type IIA supergravity, but to obtain type IIB one need only shift the dimensions of the fields by one in either direction. Massless type IIA supergravity contains two observable Ramond-Ramond $p$-form field strengths, a 2-form $F_2$ and a 4-form $F_4$. The Lagrangian density contains the usual kinetic terms for these fields

$$\mathcal{L} \supset F_2 \wedge \ast F_2 + F_4 \wedge \ast F_4.$$  \hfill (2.10)

We will use the democratic formulation of type II supergravity \cite{16, 17, 18} in which there are no Chern-Simons terms, and it is conventional to give new names to the Hodge dual field strengths

$$F_6 = \ast F_4, \quad F_8 = \ast F_2.$$  \hfill (2.11)

Furthermore we will define a single form $F$, called the improved field strength, which is the formal sum of all of the field strengths. Notice that $F$ does not have a definite degree.

As in the case of QED, the Lagrangian density does not suffice to determine the theory. One must also impose a constraint on the field strengths. In the case of
supergravity the constraint is a nontrivial generalization of the Bianchi identity in QED

\[(d + H)F = 0\]  

(2.12)

where \( H \) is a closed 3-form known as the NSNS field strength. \((d + H)\) squares to zero as \(dH = 0\) and so the \((d + H)\)-closure of \( F \) in the generalized Bianchi identity (2.12) implies that locally \( F \) is \((d + H)\)-exact. This means that on local patches of spacetime we may define a potential \( C \), which is a form of mixed odd degrees such that

\[F = (d + H)C.\]  

(2.13)

Then Eq. (2.12) becomes a Bianchi identity for \( C \), which is the condition that \( C \) is locally well-defined and so is annihilated by the square of the nilpotent operator \( d + H \). We will often decompose \( C \) into its component \( p \)-forms \( C_1, C_3, C_5 \) and \( C_7 \).

Eq. (2.13) does not uniquely specify the gauge potentials \( C \), but only specifies them up to an additive shift

\[C \mapsto C + \eta\]  

(2.14)

which is called a gauge transformation. As in QED if the spacetime topology is nontrivial then in the absence of charged matter \( \eta \) is not necessarily \((d + H)\)-exact but only needs to be \((d + H)\)-closed: \((d + H)\eta = 0\).

There is a second natural definition of field strength in this theory

\[G_{p+1} = dC_p.\]  

(2.15)

As \( C \) is locally well-defined, this field strength will satisfy the usual Bianchi identity \( dG = 0 \) and so its integral will be used to measure D-brane charge, therefore unlike \( F \) it will be quantized. However \( G \) is not gauge-invariant, instead

\[G_{p+1} \mapsto G_{p+1} + d\eta_p = G_{p+1} - H \wedge \eta_{p-2}\]  

(2.16)

where in the second equality we have used the \((d + H)\) closure of \( \eta \). We may use the Bianchi identity to find a relation between \( G \) and \( F \)

\[0 = dG = dF + H \wedge G.\]  

(2.17)

As \( F \) is gauge-invariant it is not transformed under transition functions and so it is globally defined. Thus \( dF \) is exact and so \( H \wedge G \) is a trivial element of de Rham
cohomology. The triviality of $H \wedge G$ in cohomology will play the role of the constraints in the BRST interpretation.

Analogously to the case of QED, type II supergravity without matter is invariant under gauge transformations with gauge parameter equal to any $(d + H)$-closed $\eta$. However if we add charged matter to the theory than we must check that the resulting phase shift in the path integral is well-defined, and this will impose a further restriction on $\eta$. Charged matter in this case consists of $p$-branes, which in classical supergravity can wrap any cycle $N \subset M$ such that

$$\int_N H = 0. \quad (2.18)$$

The contribution of such a wrapping to the Wess-Zumino terms of the $p$-brane world-volume path integral is

$$S_{WZ} = e^{\int_N e^{F + B} \eta} \quad (2.19)$$

where the closed 2-form $F$ is the field strength of the worldvolume $U(1)$ gauge field and $B$ is the pullback of the NS 2-form, which is well-defined on $N$ because of Eq. (2.18). Under the gauge transformation $C \mapsto C + \eta$ the Wess-Zumino term transforms as

$$S_{WZ} = e^{\int_N e^{F + B} \eta} \mapsto e^{\int_N e^{F + B} (C + \eta)} = e^{\int_N e^{F + B} \eta} S_{WZ}. \quad (2.20)$$

Thus gauge invariance imposes that

$$\int_N e^{F + B} \eta \in \mathbb{Z} \quad (2.21)$$

for every $N$ satisfying (2.18).

At first (2.21) may seem impossible to satisfy, as $\eta$ is not closed under $d$ but only under $d + H$ and so it may appear as though the integral depends on the choice of homological representative $N$. However this is not the case, since the integrand is closed

$$d(e^{F + B} \eta) = (de^F)e^B \eta + e^F de^B \eta + e^{F + B} d\eta = (dF)e^B \eta + e^F H e^B \eta - e^{F + B} H \eta = 0 \quad (2.22)$$

where we have used the closure of $F$ and the fact that $H$ commutes with even forms. Thus on each cycle $N$ the integrand in the phase (2.20) is closed. As each point in $M$ is in some such cycle, the integrand is everywhere closed and so in particular can be lifted, although not canonically, to a cohomology class on $M$. 

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The quantization condition (2.21) is not sufficient to impose that the resulting cohomology class is integral, as it applies only to cycles $N$ that do not support $H$ flux. However, if one projects out the image of the operator $(H \wedge)$ from the de Rham cohomology group then the image of the lift of the gauge transformation under this projection will inhabit an integral lattice of the resulting quotient of the de Rham cohomology. In fact, to compute the gauge transformations of $G$ in (2.16) we are not interested in gauge transformations $\eta$ in the image of $(H \wedge)$, because $G$ shifts by $H \wedge \eta$ which is invariant under shifts $\eta \mapsto \eta + H \wedge \Lambda$. Thus the classes $\eta$ in the gauge transformation of $G$ may be taken to be integral classes. In fact, the lift of the phase in (2.20) from the integral cohomology of $N$ to that of $M$ is not even well-defined on image of $(H \wedge)$, as one needs to paste together the various values of the $B$-field with NS gauge transformations, and so we are fortunate that in considering the gauge transformations of $G$ we do not need to know the component of $\eta$ that is in the image of $H \wedge$. If one wishes to extend this note to find the gauge orbits of $C$ then one will again be confronted with this problem.

In conclusion, it appears that, as in the case of QED, in type II supergravities there are large gauge transformations in which the gauge parameter $\eta_p$ lie, up to a correction in the image of $(H \wedge)$ which does not contribute to the transformation of $G$, on an integral lattice of the $p$th de Rham cohomology group. Similarly, in the quantum theory there will also be torsion terms that can be included by allowing $\eta_p$ to be an arbitrary element of the $p$th integral cohomology group. Ignoring the torsion for now, one may compute the groups of allowed, gauge-inequivalent field strengths $G_{p+1}$. We are not going to be interested in the connections $C_p$, although including them would be interesting and would hopefully lead to a kind of differential K-theory.

We have seen that the equations of motion imply that $H \wedge G_{p+1}$ represents the trivial cohomology class, thus the allowed field strengths form the kernel of the operator

$$H \wedge : H^{p+1}(M) \rightarrow H^{p+4}(M) : x \mapsto H \wedge x.$$  \hfill (2.23)

The gauge transformations $\eta_p$ are classified by the $p$th cohomology. In fact we’ll be interested in $\eta_{p-2}$, which is classified by the $(p - 2)$nd cohomology. These act on the field strengths via Eq. (2.16), which adds $H \wedge \eta_{p-2}$ to $G_{p+1}$. As $\eta_{p-2}$ may be any $(p - 2)$nd cohomology class, $G_{p+1}$ is only defined up to $(p + 1)$-classes which are the wedge product of $H$ with something. That is, the possible shifts of $G_{p+1}$ are the
image of the operator

\[ H \wedge : H^{p-2}(M) \to H^{p+1}(M) : x \mapsto H \wedge x. \]  \hspace{1cm} (2.24)

Quotienting the solutions of the RR supergravity Bianchi identities \( \text{[2.23]} \) by the RR gauge transformations \( \text{[2.24]} \) gives a classification of gauge-invariant RR field strengths

\[ G_{p+1} \in \frac{\text{Ker}(H \wedge : H^{p+1}(M) \to H^{p+4}(M))}{\text{Im}(H \wedge : H^{p-2}(M) \to H^{p+1}(M))}. \]  \hspace{1cm} (2.25)

The classification of field strengths \( \text{[2.25]} \) has an interpretation in terms of the BRST cohomology with respect to the large gauge transformations \( \text{[2.24]} \) where \( \eta \) is valued in integral cohomology. The numerator consists of the field strengths that satisfy the constraints, which are given by the Bianchi identities above, where as the denominator consists of those field strengths that are pure gauge. On the other hand, no physical fields occupy the cohomology classes \( H^{p-2} \) and \( H^{p+4} \), which are even and so have the wrong statistics to be field strengths in type IIB. Yet the introduction of these two classes is crucial, \( H^{p-2} \) because it is the preimage of the pure gauge field strengths under the \( H \wedge \) map \( \text{[2.24]} \) and \( H^{p+4} \) because it is the image of the forbidden field strengths under the \( H \wedge \) map \( \text{[2.23]} \). This leads us to identify \( H^{p-2} \) and \( H^{p+4} \) as topological (anti)ghost fields and \( H \wedge \) as the BRST operator.

Notice that we have not imposed the NSNS equations of motion, nor have we quotiented by the field strengths that may be removed by NSNS gauge transformations. Thus we have only partially classified the field strengths in IIB supergravity.

### 3 The Atiyah-Hirzebruch spectral sequence

#### 3.1 Adding torsion

In the last section we used supergravity to classify gauge orbits of Ramond-Ramond field strengths that satisfy the Bianchi identities. We argued that the field strengths should live in integral cohomology, which is the sum of a free part of the form \( \mathbb{Z}^j \) plus a torsion part that consists of cyclic groups \( \mathbb{Z}_{q_i} \). The free part is equal to the integral lattice of the de Rham cohomology. Our supergravity techniques are rather limited, in that we use a classical action and try to obtain information about a quantum theory. In particular, in the classical limit the torsion part of the cohomology vanishes and
so we cannot simply use the gauge symmetries of the supergravity action to find the action of the BRST operator on the torsion terms. This means that the classification of RR field strengths will not be quite right on a spacetime whose integral cohomology contains a nontrivial torsion piece. In section 4 we will use the Freed-Witten anomaly to recover these missing torsion contributions.

In the present section we will find the torsion contributions from an entirely different perspective. We will review the Atiyah-Hirzebruch spectral sequence (AHSS) construction of twisted K-theory, and show that if the torsion pieces of the cohomology are simply dropped then instead of constructing twisted K-theory, one arrives at the above classification of RR field strengths. In other words, we will prove that our above classification of supergravity configurations is an approximation of twisted K-theory in which one ignores the torsion parts of the integral cohomology, taking only the free part which is a sublattice of de Rham cohomology. In particular, if we consider a theory with a spacetime whose cohomology contains no torsion, such as the SU(2) WZW model, then the classification is already correct as we will see in Subsec. 3.4. This, in turn, will provide us with a conjecture for the torsion corrections to the supergravity results of the previous section. One need only reintroduce the torsion terms that appear in the AHSS. In section 4 we will argue that those torsion terms are in fact required for the cancellation of global anomalies on the worldsheets of open fundamental strings.

### 3.2 What kind of objects are twisted K-groups?

The integral cohomology $H^*(M)$ of a space $M$ is a collection of abelian groups $H^k(M)$ indexed by a nonnegative integer $k$ and endowed with a product

$$
\cup : H^j(M) \otimes \mathbb{Z} H^k(M) \to H^{j+k}(M) : x \otimes y \mapsto x \cup y
$$

called the cup product. The free part $\mathbb{Z}^k$ of the integral cohomology group $\mathbb{Z}^k \oplus \mathbb{Z} q_i$ is isomorphic to an integral lattice of de Rham cohomology $\mathbb{R}^k$. On this integral lattice the cup product $x \cup y$ reduces to the wedge product of differential forms $x \wedge y$. Thus we may guess that the wedge products of the previous section will, when we include torsion, be written as cup products plus torsion corrections.

The twisted K-theory $K^*_H(M)$, with twist $H$, of a space $M$ is another collection of abelian groups $K^k_H(M)$, this time indexed by a general integer $k$ and depending on
an integral three-class \(H\). However, unlike the cohomology groups, the \(K\) groups are not independent. Instead they are related by the Bott periodicity relation

\[
K^j_H(M) \cong K^{j+2}_H(M)
\]

and thus it will suffice to compute \(K^0_H(M)\) and \(K^1_H(M)\). This structure better mimics the charge structure of D-branes. For example there are dielectric and fractional D\(p\)-branes that carry a half unit of D\((p-2)\)-brane charge, meaning that twice the generator of D\(p\) charge should be equal to the generator of D\((p-2)\) charge. Such a relation, between generators of different degrees, would be impossible in cohomology. However in K-theory it is automatic, a single group classifies all even degrees, while another classifies all odd degrees. Thus, in each string theory, one K-group classifies all of the D-branes simultaneously. As D-branes source fluxes, one can arrive at a similar story for fluxes.

While ordinary K-theory admits a multiplication similar to (3.1), the K group twisted by a fixed \(H\) admits no such multiplication. Instead when one multiplies two twisted K-groups the twists add. However the twist corresponds to a fixed NS 3-form flux, and so this multiplication changes the \(H\) flux and so does not correspond to any obvious physical process in string theory. Setting the twist of one of the factors to zero, one finds that the twisted K-theory \(K^*_H(M)\) is a module of the untwisted K group \(K^0_H\), but no physical interpretation of this fact in string theory has appeared to date.

### 3.3 Constructing twisted K-theory

We will now describe the AHSS, which is an algorithm for computing the twisted K groups \(K^i_H(M)\) from the \(H\) flux and the integral cohomology \(H^*(M)\). In type IIB string theory RR field strengths correspond to odd cohomology classes, and so we shall see they will be classified by \(K^1_H(M)\). Similarly in IIA string theory the fluxes are even classes and so will be described by \(K^0_H(M)\).

We first assemble the even and odd cohomologies into two big groups \(H^{\text{even}}\) and \(H^{\text{odd}}\)

\[
H^{\text{even}}(M) = \bigoplus_k H^{2k}(M), \quad H^{\text{odd}}(M) = \bigoplus_k H^{2k+1}(M).
\]

We want to construct the twisted K-group \(K^i_H(M)\) by finding a finite set of improving approximations \(E^j_i\) starting with the odd cohomology and finishing with a set \(E^i_n\) of
the same cardinality as $K^i_H(M)$

$$E^n_i = H^{\text{even}}(M), \quad E^n_i = H^{\text{odd}}(M), \quad |E^n_i| = |K^i_H(M)|. \quad (3.4)$$

In general $E^n_i$ will not be isomorphic to $K^i_H(M)$ as a group, that is, the addition rule will be different. The addition rules are related by an extension problem. However, as there is no addition rule for fluxes in the S-duality covariant case anyway, since the sum of two sets of fields that satisfy the nonlinear supergravity equations of motion generically is not another solution, we will not concern ourselves with reproducing the addition rule in this note.

To get from $E^n_i$ to $E^n_{i+1}$ we will need to introduce a series of differential operators $d_{2j+1}$ which are degree $(2j + 1)$ cohomology operations, in other words

$$d_{2j+1} : H^k(M) \to H^{k+2j+1}(M), \quad d_{2j+1}d_{2j+1} = 0. \quad (3.5)$$

As we will soon see, only the first differential operator, $d_3$, needs to be well-defined on the full integral cohomology. To pass from $E^n_i$ to the next approximation $E^n_{i+1}$ we need to take the cohomology with respect to the differential operator $d_{2j+1}$

$$E^n_{i+1} = \frac{\text{Ker}(d_{2j+1} : E^n_i \to E^n_{i+1})}{\text{Im}(d_{2j+1} : E^{i+1} \to E^n_i)}. \quad (3.6)$$

In particular, we see that while the differentials do not need to be defined on all of the cohomology classes, as some classes will be eliminated earlier in the procedure by the kernel operations, they do need to be well-defined on the equivalence classes obtained by quotienting by the images of their predecessors. This identifies the differentials beyond $d_3$ not as ordinary cohomology operations, but as secondary cohomology operations obtained perhaps from Toda brackets of primary cohomology operations. However the details of the constructions of these $d$’s will not concern us, it will suffice to use the claims of Diaconescu, Moore and Witten [5] and then Maldacena, Moore and Seiberg [15] that they compute Freed-Witten anomalies.

In known examples, the first differential $d_3$ is the only differential that does not vanish in the absence of torsion cohomology. In fact, the torsion free parts of all of the differentials have been computed in [20], where they have been seen to be Massey products of $H$. This implies in particular that they will always be trivial on compact Kahler manifolds. However in principle they could appear in supergravity, and it would be interesting to find the corresponding gauge transformations. $d_3$ may be written

$$d_3x = Sq^3x + H \cup x \quad (3.7)$$
where $H$ is our familiar NS 3-class and $Sq^3$ is a cohomology operation known as a Steenrod square which takes an integral class in the $k$th cohomology to a class in the $(k + 3)$rd cohomology, as does the cup product with $H$. Unlike the cup product with $H$, however, $Sq^3$ is only nontrivial when acting on $\mathbb{Z}_2$ torsion components of $H^k(M)$, and the image is likewise always a $\mathbb{Z}_2$ torsion component of $H^{k+3}(M)$. This means in particular that in the supergravity limit such torsion components disappear because there is no Dirac quantization and so $\mathbb{Z}_2$ torsion fields may always be written as two times another field and so torsion fields are even and therefore equal to zero modulo 2. Therefore in the classical supergravity limit the $Sq^3$ term vanishes and $d_3$ reduces to the operator in Eq. (2.23).

The physical interpretation is that the $Sq^3$ term measures an obstruction to a brane worldvolume being $spin^c$. If the worldvolume is not $spin^c$, then the fermion partition function will be anomalous unless this anomaly is canceled by the $H$ flux. Geometrically this cancellation may be realized by tensoring a $spin^c$ bundle with 3-class $W_3$ by an $LE_8$ bundle with a three-class that cancels $W_3$, thus the tensor product bundle has no 3-class and so no obstruction to choosing a $spin$ structure.

This anomaly, which comes from an ill-definedness in the square root of a determinant of a Dirac operator in the path-integral measure, is not visible in the classical theory. Thus in the classical limit all that remains of the above procedure is $d_3$ which is equal to the wedge product with $H$. One then recovers the classification (2.25).

The appearance of (2.25) as the classical limit of the AHSS construction suggests the conjecture that the AHSS construction provides a quantum completion of the supergravity construction in the previous section. In particular, one may conjecture that the quantum-corrected Bianchi identities imply that physical states are in the kernel of the AHSS differentials $d_{2j+1}$. In addition, one may conjecture that the quantum-corrected gauge transformations are such that the images of the AHSS differentials are pure gauge. Now the $K_{ij}^{i+1}$’s are interpreted as a series of nonpropagating ghosts, and the $d_{2j+1}$’s as a series of BRST operators. Of course, in string theory only $d_3$, $d_5$ and perhaps $d_7$ may ever be relevant for dimensional reasons.
3.4 An example: The twisted K-theory of $S^3$

The supersymmetric SU(2) WZW model at affine level $k-2$ describes string theory on the three-sphere $S^3$ with $k \neq 0$ units of NS 3-form flux

$$\int_{S^3} H = k.$$  \hfill (3.8)

The integer cohomology of the three-sphere is

$$H^0(S^3) = H^3(S^3) = \mathbb{Z}, \quad H^1(S^3) = H^2(S^3) = 0$$  \hfill (3.9)

where 0 is the trivial group, which contains only the identity element. In particular, the cohomology contains no torsion and so $Sq^3$ acts trivially. As the maximum difference in the dimensions of two elements of the cohomology is equal to three, the operators $d_{2j+1}$ are trivial for $j > 1$, leaving only $d_3$, which contains only the cup product with $H$.

If $e_0$ is the generator of $H^0(S^3) = \mathbb{Z}$ and $e_3$ is the generator of $H^3(S^3) = \mathbb{Z}$ then

$$d_3 e_0 = H \cup e_0 = ke_3, \quad d_3 e_3 = H \cup e_3 = 0$$  \hfill (3.10)

where $H$ kills $e_3$ because the cup product of two 3-classes is a 6-class, but the 6-cohomology is trivial. Thus $d_3$ acts nontrivially on all of $H^0(S^3)$, but annihilates all of $H^3(S^3)$. In other words, the kernel of $d_3$ is the third cohomology group, $\mathbb{Z}$. Similarly the image of $d_3$ consists of those elements $k\mathbb{Z} \subset \mathbb{Z}$ of $H^3$ which are multiples of $k$. Note that both the kernel and the image lie in $H^3$, which is in $H^{\text{odd}}$ because 3 is odd. $H^{\text{even}}$, on the other hand, contains no elements of the kernel and so $K^0_H(S^3)$ is trivial. Summarizing, we have found that

$$K^1_H(S^3) = \frac{\ker(d_3 : H^{\text{odd}}(M) = \mathbb{Z} \to H^{\text{even}}(M) = \mathbb{Z})}{\text{Im}(d_3 : H^{\text{even}} = \mathbb{Z} \to H^{\text{odd}} = \mathbb{Z})} = \frac{H^3(M) = \mathbb{Z}}{kH^3(M) = k\mathbb{Z}} = \mathbb{Z}_k.$$  \hfill (3.11)

As was shown in Ref. [2], the twisted K-groups (3.11) reproduce the known symmetric D-branes in the supersymmetric SU(2) WZW model.

Physically the nontrivial element $j \in \mathbb{Z}_k = K^1_H(S^3)$ corresponds to a RR 3-form field strength $G_3$ such that $\int_{S^3} G_3 = j$ in an embedding of the SU(2) WZW model into IIB string theory, for example on a 3-sphere linking an NS5-brane. We recall that $G_3$ is related to the gauge-invariant field strength $F_3$ via

$$G_3 = F_3 - C_0 H$$  \hfill (3.12)

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and so the large gauge transformation corresponds to \( C_0 \mapsto C_0 + 1 \), which is the generator \( T \) of the S-duality group \( SL(2, \mathbb{Z}) \).

The Bianchi identity \( G_0 H = 0 \) implies that \( G_0 = 0 \) and so there is no braneless embedding of the SU(2) WZW model in massive IIA. Instead such an embedding will require \( k \) D6-branes intersecting the 3-sphere. In the above NS5-brane realization, this reflects the familiar fact that an NS5-brane is confined by \( k \) D6’s when the Romans’ mass is equal to \( k \).

4 Torsion corrections from the Freed-Witten anomaly

4.1 The Freed-Witten anomaly

In Ref. [4] Freed and Witten have demonstrated that global worldsheet anomaly cancellation dictates the condition

\[
W_3 + H = dF
\]  

(4.1)

on the worldvolume of a Dp-brane wrapping a compact cycle \( N \subset M \) in the type II string theory spacetime \( M \). Here \( W_3 \) is the third Stiefel-Whitney class of the normal bundle of \( N \) in \( M \), which vanishes if \( N \) is \( spin^c \) as \( M \) and \( N \) are both orientable in type II if the D-brane is to carry a RR charge and \( M \) is \( spin^c \). More generally \( W_3 \) is a \( \mathbb{Z}_2 \)-valued class in the third integral cohomology of \( N \). \( H \) is, as usual, the NSNS 3-form field strength pulled back to \( N \), or rather the associated class in integral cohomology. \( dF \) is, as a differential form, the exterior derivative of the 2-form field strength of the D-brane worldvolume’s \( U(1) \) gauge field, which is the magnetic monopole charge and so is Poincaré dual, in \( N \), to boundary of a D\((p-2)\) brane that ends on our Dp-brane.

All three terms in Eq. (4.1) are integral cohomology classes, thus as usual only the free part of \( dF \) may be locally interpreted as the derivative of a field strength 2-form. In this context the magnetic monopole charge \( dF \) is still Poincaré dual to a codimension 3 submanifold of \( N \) on which a lower-dimensional D-brane ends, however it may be that this submanifold is homologically trivial when included into the spacetime \( M \). An example of such a phenomenon occurs in the SU(3) WZW model and is described in Ref. [15]. In this instance one finds that, instead of a D\((p-2)\)-brane ending on the Dp, a D\((p-4)\)-brane ends on the Dp. In general the differential \( d_{2j+1} \) will give us information about D\((p-2j)\)-brane insertions on the
Dp-brane. The inserted branes extend outward from the Dp-brane until either they hit another brane on which Eq. (4.1) permits them to end, or else until they reach the end of the spacetime. Such configurations of branes ending on branes are referred to as baryons in Ref. [21], as in some examples they represent baryonic vertices in a dual conformal field theory.

4.2 An aside: Physical interpretation of Freed-Witten

While in this note we are only interested in applying the Freed-Witten anomaly cancellation condition (4.1) and not in deriving it, we will now try to provide a physical interpretation of the condition. The uninterested reader may skip to the next subsection. Whenever there is a loop in spacetime, even if the loop is contractible, one has a choice of boundary conditions for the various spinor fields. They may either remain invariant when they encircle the cycle, or else they may change sign. As one deforms a loop, the spinor is transported along a spin bundle, which is a choice of square root of the tangent bundle where the above sign choices are interpreted as sign choices in the square root. However, sometimes a given set of choices is not compatible with this transport. In fact, sometimes no set of boundary conditions is compatible. If there exists a compatible set of fermion boundary conditions, one says that the manifold is spin. Notice that a manifold, such as \( \mathbb{CP}^2 \), can be nonspin even if it is simply-connected. This is because even the choice of boundary condition of the contractible cycle is not invariant as the contractible cycle is transported around the nontrivial 2-cycle of \( \mathbb{CP}^2 \).

We conclude that in general a fermion partition function can only be defined if the manifold is spin. If the fermion is charged under a \( U(1) \) gauge symmetry then the partition function is not a section of the spin bundle, but rather a section of a spin\(^c\) bundle, which is the tensor product of the spin bundle by the \( U(1) \) gauge bundle. In particular, even if the spacetime is not spin and so the spin bundle does not exist, because the product of transition functions on a three-way overlap is \(-1\) instead of \(+1\), there may exist a \( U(1) \) gauge bundle that fails to satisfy the triple overlap condition on the same overlaps. In this case the tensor product bundle satisfies the triple overlap condition and so the fermions exist. Such a choice of \( U(1) \) bundle is called a spin\(^c\) structure, and it exists when \( W_3 = 0 \).

We may understand this cancellation in terms of \( \mathbb{Z}_2 \)-valued cohomology. A bundle
is not spin if it has a second Stiefel-Whitney \( w_2 \neq 0 \in H^2(M;\mathbb{Z}_2) \) in the second cohomology group. This obstruction is canceled by a \( U(1) \) bundle if the image in \( \mathbb{Z}_2 \) cohomology of the Chern class \( F \in H^2(M;\mathbb{Z}_2) \) of the \( U(1) \) bundle is another \( \mathbb{Z}_2 \) torsion class which precisely cancels \( w_2 \)

\[
w_2 + F = 0. 
\]

(4.2)

An honest bundle must have a Chern class valued in integral cohomology, not in cohomology with \( \mathbb{Z}_2 \) coefficients. However using the short exact sequence

\[
0 \longrightarrow \mathbb{Z} \xrightarrow{x^2} \mathbb{Z} \longrightarrow \mathbb{Z}_2 \longrightarrow 0
\]

(4.3)

one can construct the long exact cohomology sequence

\[
\rightarrow H^2(M;\mathbb{Z}) \rightarrow H^2(M;\mathbb{Z}_2) \xrightarrow{\beta} H^3(M;\mathbb{Z}) \rightarrow \]

(4.4)

where \( \beta \) is referred to as a Bockstein homomorphism. The sequence (4.4) implies that an element \( w_2 + F \) of \( H^2(M;\mathbb{Z}_2) \) can be lifted to an honest Chern class, that is an element of \( H^2(M;\mathbb{Z}) \), whenever it is in the kernel of the Bockstein. This means that the image of the Bockstein

\[
\beta(w_2 + F) = W_3 + dF
\]

(4.5)

is the obstruction to the existence of the bundle and therefore to the existence of the fermion. The Freed-Witten anomaly however is not precisely the condition that (4.5) vanishes, there is also an \( H \) term. This \( H \) term does not appear in QED, but is an additional complication that arises in supergravity.

In type II supergravity theories the consistency condition for spinors is weaker than in QED. We have seen that a charged fermion in QED lives in a spin\(^c\) bundle, which is the tensor product of a spin bundle with a \( U(1) \) bundle. Likewise a charged fermion that is described by a fundamental string lives in the tensor product of a spin\(^c\) bundle with a \( PU(\mathcal{H}) \) bundle, where \( PU(\mathcal{H}) \) is the projective unitary group on the Hilbert space \( \mathcal{H} \). While \( U(1) \) bundles are classified by two-classes \( F \) called Chern classes, \( PU(\mathcal{H}) \) bundles are classified by 3-classes \( H \) called Dixmier-Douady classes. The total bundle on which the fermion partition function is defined is then the product of the spin bundle, the \( U(1) \) bundle and the \( PU(\mathcal{H}) \) bundle. Unlike the previous case, the various characteristic classes \( W_3, dF \) and \( H \) are already integral classes, and so we do not need to worry about the obstruction to lifting them to integral classes.
However we do need to impose that $H$ cancels the other characteristic classes, which leads to the Freed-Witten anomaly (4.1).

While the above construction works for spacetimes of any dimension, for those of dimension less than 15, as in string theory, one may replace the $\text{PU}(\mathcal{H})$ bundle with the based loop group of $E_8$. The physical interpretations of the appearance of both of these groups is at best mysterious. However loop groups of $E_8$ appear to be ubiquitous in string theory, for example they appear on the current algebras of the heterotic strings. One day perhaps we will learn that these are all manifestations of the same $E_8$. The $\text{PU}(\mathcal{H})$ bundle, however mysterious, does provide a connection with the definition of twisted K-theory in [22] as equivariant sections of gauge bundles over $\text{PU}(\mathcal{H})$ bundles over spacetime.

4.3 Quantum corrections to the Bianchi identities

We are now ready to use the Freed-Witten anomaly to compute the torsion corrections to the BRST operator. In particular, we need to know the quantum corrections to the constraints, which determine the action of the BRST operator on the physical fields, and we need to know the quantum corrections to the gauge transformations, which determine the action of the BRST operator on the ghosts.

While we are searching for a classification of fluxes, the Freed-Witten anomaly is a condition on branes. To convert a condition on branes into a condition on fluxes we use an argument based on Gauss’ law that has appeared, for example, in Ref. [23]. As a differential form in de Rham cohomology, a field strength is determined by its integral over the cycles which represent various homology classes. This notion is easily extended to the full integral cohomology by replacing integration with the homology/cohomology pairing. Thus when we speak of a $p$-flux on a $p$-cycle, we will really be referring to the pairing of the corresponding cohomology and homology classes. As has been argued in Ref. [24], the RR $p$-flux on a topologically trivial $p$-cycle measures the $D(8 - p)$-brane charge that is linked by the trivial cycle. In particular, D-brane charge in the full quantum theory is classified by integral homology, and so this pairing also includes torsion terms.

Witten has argued in Ref. [23] that even when a cycle is topologically nontrivial, the consistency conditions for the fluxes supported on the cycle, being local, cannot depend on whether at some far away place the cycle degenerates and the flux is...
sourced by a D-brane. Thus the only consistent RR $p$-fluxes, even on noncontractible cycles, are supposedly those fluxes that could be sourced by a D$(8-p)$-brane had the cycle been contractible.

The twisted K-theory classification only applies to fluxes in the absence of D-brane charges. This is because D-brane charges shift the Bianchi identities and so shift the set of consistent fluxes away from twisted K-theory. Thus the $p$-cycle on which we are measuring the $p$-flux must not intersect any D-branes. However, if $W_3 + H$ is nonzero on the worldvolume of the D$(8-p)$-brane that may source this flux, then there will be lower D-brane insertions ending on the D$(8-p)$-brane. The other insertions may either end on other D-branes, whose total $H + W_3$ cancels that of our original brane, or they may continue to the end of spacetime. In the first case, the total $H + W_3$ vanishes, where this addition is defined after the various forms are pushed forward from the brane worldvolumes $N$ on to the spacetime $M$, alternately one may think of both D-branes as a single disconnected D-brane which is therefore subject to Freed-Witten. In the second case, the semi-infinite inserted branes will intersect our $p$-cycle, no matter how distant the D$(8-p)$ is, and so invalidate our initial hypothesis that our spacetime in fact contains no branes. Thus the Freed-Witten anomaly leads to quantum corrections, that is torsion corrections, to the Bianchi identity which impose that the $p$-flux on any cycle can be generated by a brane for which $W_3 + H$ vanishes.

A necessary, but not sufficient, condition for the vanishing of $W_3 + H$ on this D-brane worldvolume is that the $p$-flux be in the kernel of the AHSS differential

$$d_3 = Sq^3 + H \cup .$$

To see this, we simplify matters by considering a spacetime which is homogenous in the radial direction from the D$(8-p)$, which is the direction followed by any brane insertions. In particular if one is classifying fluxes on time slices then one may identify this radial direction with time, as in Ref. [15]. On each 9-dimensional radial slice Gauss’ law then ensures that the $p$-flux is Poincaré dual to the $(9-p)$-dimensional D$(8-p)$-brane worldvolume. The $H$ flux in the Freed-Witten anomaly (4.1) is the pullback of the $H$ flux in the bulk spacetime, thus its pushforward back to the bulk reproduces the $H$ term in $d_3$. $Sq^3$ on the other hand is defined to be the pushforward of the third Stiefel-Whitney class, $W_3$, of the normal bundle of the cycle Poincaré dual to our $p$-flux $G_p$, which is precisely our D-brane worldvolume. Thus the image of $d_3G_p$ is precisely the pushforward of $W_3 + H$ under the inclusion $i : N \hookrightarrow M$, which
must vanish if $W_3 + H$ is indeed zero as the pushforward map is linear

$$d_3G_p = Sq^3G_p + H \cup G_p = i_*(W_3 + H) = 0. \quad (4.7)$$

### 4.4 A Mod 3 Torsion Freed-Witten Anomaly

We have now argued that a necessary condition on any $p$-flux in a brane-free world is that it be in the kernel of $d_3$. We cannot prove that a necessary and sufficient condition is that it be in the kernel of the entire set of $d_{2j+1}$ operators. However the results of Ref. [15] were consistent with the suggestion that the kernel of this collection of operators is precisely the set of fluxes dual to branes for which $W_3 + H = 0$. This conjecture would be very strong, as $d_5$ for example contains the mod 3 Milnor primitive $Q_1$ [20], and at higher order one finds the mod 5 Milnor primitive, which are conditions on the $\mathbb{Z}_3$ and $\mathbb{Z}_5$ torsions. This is in contrast to $W_3$, which is only a condition on the $\mathbb{Z}_2$ torsion, and so it is difficult to see how it may reproduce the other conditions. One physical source of $\mathbb{Z}_3$ and $\mathbb{Z}_5$ torsion conditions in string theory is the set of $E_8$ triples, investigated in Ref. [26]. A similar $\mathbb{Z}_3$ anomaly has appeared in M-theory [27], where it was noted that the Chern-Simons term is at level 1/6 and so only a conspiracy of anomaly-cancellations allows the modulo 2 and modulo 3 parts of the partition function to be well-defined. Perhaps the $\mathbb{Z}_3$ torsion term is a characteristic class of a generalization of a spin$^c$ structure involving the cube root of an $E_8$ bundle that appeared in the M-theory context.

However it may be that in the low dimensions of interest to string theory such higher torsion terms do not appear. In Ref. [15] the authors have demonstrated that in the SU(3) WZW model, which may be embedded in type II string theory, there is only $\mathbb{Z}_2$ torsion and $d_5$ successfully implements the Freed-Witten condition. SU($N$) models at higher $N$ have higher torsions, but cannot be embedded in type II string theory so there is still no counterexample to this conjecture.

To construct a potential counterexample we need a 10-dimensional manifold on which the $\mathbb{Z}_3$ primitive $Q_1$ is nontrivial. One example is the Euclidean space $S^3/\mathbb{Z}_3 \times S^7/\mathbb{Z}_3$ where we have quotiented by the $\mathbb{Z}_3$ subgroups of the free circle actions on the spheres. This is equivalently a $T^2$ bundle over $\mathbb{C}P^1 \times \mathbb{C}P^3$. Such a compactification will never be Ricci flat, and so in practice one may wish to replace $\mathbb{C}P^1$ by a higher genus Riemann surface, but this replacement is inconsequential for a calculation of $Q_1$. Acting on an integer cohomology class, $Q_1$ is just minus a Bockstein of the
Steenrod power $P^1$. If the degree 1 and 2 cohomologies of $S^3/Z_3$ are generated by $a_1$ and $a_2$ and those of $S^7/Z_3$ by $b_1$ and $b_2$ then the Bockstein $\beta$ takes $a_1$ to $-a_2$ and $b_1$ to $-b_2$. Therefore

$$Q_1(a_1b_2 - a_2b_1) = Q_1(a_1b_2) = -\beta(a_1P^1b_2) = -\beta(a_1)b_2^3 = b_1b_2^3 \neq 0. \quad (4.8)$$

In particular this may exclude a D6-brane Poincaré dual to $a_1b_2 - a_2b_1$, although such a brane does not suffer from any known anomaly. In the interpretation of Ref. [15] this anomaly could be canceled by a $Z_3$-charged D2 ending on a $T^2$ with one leg on each sphere, but in this compact spacetime there is no place where the other end of the D2 can terminate.

This example should be investigated in the future. The K-theory classification of D-branes in IIA appears to indicate that this D6-brane should not exist, and the corresponding RR 3-flux should not exist in IIB. If this is indeed the case, it would be interesting to find a topological characterization of this obstruction and its cure. This would provide a 3-torsion generalization of the $\text{spin}^c$ structure required by fermions on the worldvolume of a D-brane, which yields the 2-torsion constraint in the Freed-Witten anomaly.

In summary, we have shown that the Freed-Witten anomaly provides 2-torsion corrections to the Bianchi identity that imply that $p$-fluxes must indeed be in the kernel of the AHSS differential $d_3$, and we have suggested that the higher differentials may yield a necessary and sufficient condition for the vanishing of the Freed-Witten anomalies.

### 4.5 Quantum corrections to the gauge transformations

Our next goal is to try to find the quantum corrections to the gauge transformations, and to compare them with the images of the AHSS differentials.

Again we will consider a $p$-cycle $X^p$ and the $p$-flux $G_p$ that it supports. Now we want to know not if $G_p$ is consistent, but rather we want to know whether it can be gauge transformed out of existence. To construct a quantum gauge transformation, let us imagine that there is a $D(10-p)$ brane far away. One may think that this is problematic as on dimensional grounds the $p$-cycle and $(11-p)$-dimensional world-volume will intersect, but again we will restrict attention to configurations in which there is a radial symmetry about the D-brane, and so the $p$-cycle will intuitively be
kept at a fixed, finite distance. It is not necessary that there be a total D-brane charge, there could be an antibrane nearby whose charges cancel that of the original brane. However, the presence of such a pair should not affect the available gauge symmetries in the system, as they can be spontaneously pair created and destroyed.

If our D(10−p)-brane wraps a cycle on which \( W_3 + H \) is nontrivial, then the Freed-Witten condition implies that it will contain a worldvolume magnetic source \( dF \) equal to \( W_3 + H \). Such a magnetic source is a lower-dimensional D-brane that ends on our D(10−p). In particular, we have argued in the last section that if the pushforward of \( W_3 + H \) onto the bulk is nonzero

\[
i : N \hookrightarrow M, \quad i_* : H^*(N) \longrightarrow H^*(M), \quad i_*(W_3 + H) \neq 0 \tag{4.9}\]

then the inserted brane will be a D(8−p)-brane which extends radially from the D(10−p). As the D(8−p)-brane is extended in the radial direction, we now need to take care that it does not intersect our p-cycle. Fortunately, the intersection of the D(8−p) with each 9-dimensional radial slice is (8−p)-dimensional, and so it can avoid our p-cycle.

We now are ready to calculate the gauge transformations of \( G_p \). These gauge transformations are the admissible transition functions of \( G_p \) when one moves from patch to patch. In particular, one may calculate the holonomy of \( G_p \) as one encircles the D(10−p). Such a journey, which occurs in parameter space and not in time, sweeps out a \((p+1)\)-dimensional patch \( Y_{p+1} \) which is intuitively just our p-cycle times a circle, although the topology of the spacetime may force the topology of the p-cycle to change during the trip. Mathematically, the topology of the p-cycle at each point is a level set in a circle-valued Morse function. As \( X \) interpolates between the p-cycle and itself it has no boundary, however one may use Stoke’s theorem to calculate the monodromy of \( X \) as the p-cycle sweeps out \( X \), encircling the D(10−p).

We will write this monodromy as an integral, although we are really considering a homology/cohomology pairing

\[
\Delta \int_{X^p} G_p = \int_{Y_{p+1}} dG_p. \tag{4.10}\]

We now remember that D(8−p)-brane charge is Poincaré dual to the source \( dG_p \), which is not exact as \( G_p \) is not gauge invariant and so not globally defined. Thus the right hand side of (4.10) is the intersection number of the trajectory \( Y_{p+1} \) and the D(8-p)-brane worldvolume. Eq. (4.10) then implies that the transformation of the
$p$-flux $G_p$ is the Poincaré dual of the D$(8-p)$-brane worldvolume, which we have seen may be obtained from the original D$(10-p)$-brane using the Freed-Witten anomaly

$$\Delta \int_{X^p} G_p = i_*(W_3 + H). \quad (4.11)$$

We may now again use the above definition of the Steenrod squares to write the pushforward of $W_3$ as $Sq^3$ of the Poincaré dual $\eta$ of the D$(10-p)$-brane worldvolume

$$\Delta \int_{X^p} G_p = i_*(W_3 + H) = (Sq^3 + H \cup)x = d_3x. \quad (4.12)$$

Thus we find that quantum gauge transformations allow $G_p$ to vary by any class in the image of $d_3$, identifying the quotient in the first step of the AHSS as the identification of gauge orbits under gauge transformations that are not annihilated by the pushforward. One may again, following Ref. [15], identify the higher differentials with the gauge transformations that are killed by the pushforward operator but are captured by various secondary cohomology operations.

5 Summary

We have argued that in type II supergravity theories the set of RR field strengths that satisfy the quantum-corrected Bianchi identities quotiented by the quantum-corrected large gauge transformations is the twisted K-theory of the spacetime, more precisely, twisted $K^0$ for IIA and $K^1$ for IIB. In particular, we have identified the Atiyah-Hirzebruch spectral sequence (AHSS) construction of twisted K-theory as the usual BRST formulation in quantum field theory, albeit for discrete transformations so the ghosts have no propagating degrees of freedom. The differential operators of the AHSS have been identified with BRST operators for these large gauge transformations.

In the classical limit, the large gauge transformations are just the shifts of the RR gauge connection that preserve the Wilson loops, which form an integral lattice of the de Rham cohomology. The corresponding constraints arise classically from the Bianchi identity. Quantum corrections to the Bianchi identity arise by supposing that the field strength is sourced by a D-brane and considering the Freed-Witten anomaly on the D-brane’s worldvolume. Quantum corrections to the gauge transformations similarly come from considering the holonomy of a flux as one encircles a D-brane which is afflicted with a Freed-Witten anomaly.
We hope that this construction of the twisted K-theory classification entirely within the usual framework of quantum field theory will be useful for attempted constructions of K-theory based quantum field theories, as one sees that no new technology needs to be introduced. In particular, here we have classified only the field strengths, but if one is also interested in the gauge connections then the BRST cohomology will yield a differential twisted K-theory. One may then compare this form of differential twisted K-theory with that which has been conjectured to exist in string theory in Ref. [13], thereby testing the conjecture.

One may also apply this strategy to the longstanding problem of reconciling the twisted K-theory classification with S-duality. To do this, one needs to also consider the Bianchi identities of the NSNS fields and the NSNS gauge transformations. To consider these in parallel with those of the RR fields may require the BV formalism.

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