ELLiptic flow and the symmeT ry energy
at supra-saturation density

W. TRAUTMANN
GSI Helmholtzzentrum für Schwerionenforschung GmbH
Planckstr. 1, D-64291 Darmstadt, Germany
w.trautmann@gsi.de

H.H. WOLTER
Fakultät für Physik, Universität München
Am Coulombwall 1, D-85748 Garching, Germany
hermann.wolter@lmu.de

Received (received date)
Revised (revised date)

The elliptic flow in collisions of neutron-rich heavy-ion systems at intermediate energies emerges as an observable sensitive to the strength of the symmetry energy at supra-saturation densities. First results obtained by comparing ratios or differences of neutron and hydrogen flows with predictions of transport models favor a moderately soft to linear density dependence, consistent with ab-initio nuclear-matter theories. Comprehensive data sets of high accuracy can be expected to improve our knowledge of the equation of state of asymmetric nuclear matter.

1. Introduction

Heavy-ion reactions at sufficiently high energy represent the only means for compressing nuclear matter in laboratory experiments and for studying the nuclear equation of state (EoS) at supra-saturation densities. From the extensive search for observables suitable for probing the brief high-density phase of the collision, collective flows and sub-threshold production of strange mesons have appeared as most useful. The present consensus that a soft EoS, corresponding to a compressibility $K \approx 200$ MeV and including momentum dependent interactions, best describes the high-density behavior of symmetric nuclear matter is based on studies of flow and kaon production within the framework of transport theory.

In recent years, motivated by the impressive progress made in observing properties of neutron stars and in understanding details of supernova explosion scenarios, the EoS of neutron-rich asymmetric matter has received increasing attention. The symmetry energy, which is the difference between the energy per nucleon of neutron matter and of symmetric matter, is considered to be one of the biggest unknowns in this context. We have precise information on the symmetry energy near...
saturation density from the knowledge of nuclear masses. For densities below saturation, investigations of heavy ion reactions in the Fermi energy regime have constrained the symmetry energy considerably\cite{7} and the importance of clustering for the symmetry energy at very low density has recently been demonstrated\cite{8,9,10}. At super-saturation density, however, the symmetry energy is still largely unknown for several reasons. Phenomenological forces, even though well constrained near saturation, yield largely diverging results if they are extrapolated to higher densities\cite{11,12}. Many-body calculations with realistic potentials face the problem that three-body forces and short-range correlations are not sufficiently constrained at higher densities at which their importance increases\cite{13,14,15}. Even the magnitude of the kinetic contribution, related to the nuclear Fermi motion and considered as principally understood, is possibly modified by a redistribution of nucleon momenta due to short-range correlations in high-density nuclear matter\cite{16}.

\begin{figure}[h]
\centering
\includegraphics[width=0.5\textwidth]{isospin asymmetry}
\caption{Isospin asymmetry $\delta = (\rho_n - \rho_p)/\rho$ as a function of the normalized density $\rho/\rho_0$ at time $t = 20$ fm/c in the $^{132}$Sn + $^{124}$Sn reaction with a stiff ($E_a^{\text{sym}}$) and with a soft ($E_b^{\text{sym}}$) density dependence of the nuclear symmetry energy. The corresponding correlation for neutron stars in $\beta$-equilibrium is shown in the inset (reprinted with permission from Ref.\cite{20}; Copyright (2002) by the American Physical Society).}
\end{figure}

The appearance of the symmetry energy in many aspects of nuclear structure and reactions would seem to imply that constraints of various kinds should be available. However, quantities as, e.g., the thickness of the neutron skin in heavy nuclei or the isospin transport in reactions of isospin-asymmetric nuclei in the Fermi-energy domain are predominantly sensitive to the strength of the symmetry term at densities below the saturation value\cite{5,7,17,18,19}. Extrapolations of parameterizations to higher density raise the question of up to where they remain realistic. This emphasizes the need for more direct high-density probes. As in the case of
symmetric nuclear matter, collective flows and sub-threshold particle production are obvious candidates.

A strong motivation for exploring the information contained in isotopic flows was provided by Bao-An Li when he pointed to the parallels in the density-dependent isotopic compositions of neutron stars and of the transient systems formed in collisions of neutron-rich nuclei as a function of the EoS input used in the calculation.\(^{20}\) Figure 1 illustrates the remarkable fact that the same physical laws are confidently applied to objects differing by 18 orders of magnitude in linear scale or 55 orders of magnitude in mass. It will allow us to infer properties of these exotic astrophysical objects from data obtained in laboratory experiments. The main difficulty is the comparatively small asymmetry of available nuclei. Symmetry effects are, therefore, always small relative to the dominating isoscalar forces which one hopes will cancel in differences or ratios of observables between isotopic partners. The observable proposed by Li is the so-called differential directed flow which is the difference of the multiplicity-weighted directed flows of neutrons and protons. Directed flow describes the rapidity dependence of the mean in-plane transverse momenta of observed reaction products.

A further encouragement was provided by transport model calculations according to which the elliptic flow of free neutrons and protons responds significantly to variations of the parameterization of the symmetry energy.\(^{21}\) Elliptic flow relates to the azimuthal anisotropy of particle emissions. This has motivated a reanalysis of the FOPI/LAND data for \(^{197}\text{Au} + {^{197}\text{Au}} \) collisions at 400 MeV/nucleon, collected many years ago and used to demonstrate the existence of neutron squeeze-out in this energy regime.\(^{22,23}\) Squeeze-out refers to a dominant out-of-plane emission of particles, relative to in-plane emission, and is considered as evidence for the pressure buildup in the collision zone. The analysis favors a moderately soft to linear density dependence of the symmetry energy.\(^{21}\)

This finding had a particular significance, in spite of a large statistical uncertainty. Rather different conclusions, suggesting either a super-stiff or super-soft behavior of the symmetry energy, had previously been reached in analyses of the \(\pi^-/\pi^+\) production ratios, measured by the FOPI Collaboration for the same \(^{197}\text{Au} + {^{197}\text{Au}} \) reaction with different transport models.\(^{24,25,26,27}\) In particular, the super-soft result has initiated a broad discussion of how it might be reconciled with other observations as, e.g., observed properties of neutron stars.\(^{26,28,29}\) The FOPI/LAND elliptic-flow data were found to be inconsistent with this extreme assumption also when compared to recent QMD transport-model calculations, fairly independent of particular choices made for other model parameters.\(^{30}\)

The obvious need to improve the statistical accuracy beyond that of the existing data set has initiated a dedicated measurement of collective flows in collisions of \(^{197}\text{Au} + {^{197}\text{Au}} \) as well as of \(^{96}\text{Zr} + {^{96}\text{Zr}} \) and \(^{96}\text{Ru} + {^{96}\text{Ru}} \) which has been carried out in 2011 with the LAND\(^{31}\) detector coupled to a subset of the CHIMERA\(^{32}\) detector array.\(^{33}\) In the following sections, the present situation will be described in more detail.
2. Present Knowledge of the Symmetry Energy

Our knowledge of the symmetry energy is originally based on nuclear masses whose dependence on the isotopic composition is reflected by the symmetry term in the Bethe-Weizsäcker mass formula. A density dependence is already indicated by the use of separate bulk and surface terms in refined mass-formulae. In the Fermi-gas model, it is given by a proportionality to $(\rho/\rho_0)^\gamma$ with an exponent $\gamma = 2/3$, where $\rho_0 \approx 0.16$ nucleons/fm$^3$ is the saturation density. This so-called kinetic contribution to the symmetry energy, however, amounts to only about 1/3 of the symmetry term of approximately 30 MeV for nuclear matter at saturation. The major contribution is given by the potential term reflecting properties of the nuclear forces.

Nuclear many-body theory has presented us with a variety of predictions for the nuclear equation of state. The examples shown in Fig. 2 for the two cases of symmetric nuclear matter and of pure neutron matter demonstrate that, overall, the results are quite compatible among each other, except for densities exceeding saturation at which the predictions diverge. The symmetry energy $E_{\text{sym}}$ can be defined as the coefficient of the quadratic term in an expansion of the energy per particle in the asymmetry $\delta = (\rho_n - \rho_p)/\rho$, where $\rho_n, \rho_p,$ and $\rho$ represent the neutron, proton, and total densities, respectively,

$$E/A(\rho, \delta) = E/A(\rho, \delta = 0) + E_{\text{sym}}(\rho) \cdot \delta^2 + O(\delta^4).$$  \hspace{1cm} (1)

In the quadratic approximation, the symmetry energy is the difference between the energies of symmetric matter ($\delta = 0$) and neutron matter ($\delta = 1$). Also the symmetry energy diverges at high density, as expected from Fig. 2, while most empirical models coincide near or slightly below saturation, the density range at which constraints from finite nuclei are valid (Fig. 3).
Elliptic Flow and the Symmetry Energy

Fig. 3. Symmetry energy as a function of density as predicted by different models. The left panel zooms the low density range up to saturation. The full lines represent the DBHF and variational approaches using realistic forces (from Ref. 12, reprinted with kind permission from Springer Science+Business Media).

The behavior of the symmetry energy at very low densities is not correctly described in homogeneous mean-field approaches which do not include clustering effects. This is a topic of considerable current interest with good agreement being reached by experimental and theoretical investigations. The uncertainty at supra-saturation density in phenomenological models may be considered as the expected consequence of extrapolations leading beyond the range at which the chosen parameterizations are effectively tested.

In calculations using realistic forces fitted to two- and three-nucleon data, the uncertainty is mainly related to the short-range behavior of the nucleon-nucleon force and, in particular, to the three-body and tensor forces. The three-body force has been shown to make an essential but quantitatively small contribution to the masses of light nuclei. The extrapolation of the partly phenomenological terms used there to higher densities is, however, highly uncertain. The general effect of including three-body forces in the calculations is a stiffening of the symmetry energy with increasing density. Short-range correlations become also increasingly important at higher densities; results from very recent new experiments will, therefore, have a strong impact on predictions for high-density nuclear matter.

At higher energies, the momentum dependence of the nuclear forces becomes important. It is well known that nuclear mean fields are momentum dependent, as seen, e.g., in the energy dependence of the nuclear optical potential. The dominating effect is in the isoscalar sector but there is also an important isovector momentum dependence. It manifests itself as an energy dependence of the isospin-
dependent part of the optical potential but can also be expressed in terms of a
difference of the effective masses of protons and neutrons.\textsuperscript{12} Even the ordering of
these effective masses is still an open problem. It has, moreover, been shown that
the effective mass differences and the asymmetry dependence of the EoS are both
influencing particle yields and flow observables, and that additional observables will
be needed to resolve the resulting ambiguity.\textsuperscript{6,13,39,40}

Fig. 4. Parameterizations of the nuclear symmetry energy as used in transport codes: three pa-
rameterizations of the potential term used in the UrQMD (Ref.\textsuperscript{41}) with power law coefficients
$\gamma = 0.5, 1.0,$ and $1.5$ (lines with symbols as indicated), the result with $\gamma = 0.69$ obtained from
analyzing isospin diffusion data with the IBUU04 (full line, Ref.\textsuperscript{43}), and the super-soft and stiff
parameterizations obtained from analyzing the $\pi^-/\pi^+$ production ratios with the IBUU04 (dotted
line, Ref.\textsuperscript{26}) and the LQMD (dashed line, Ref.\textsuperscript{27}) transport models (from Ref.\textsuperscript{44}, reprinted
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Transport theories needed for calculating the temporal evolution of nuclear reac-
tions often use simplified parameterized descriptions of the composition-dependent
part of the nuclear mean field. In the UrQMD of the group of Li and Bleicher,\textsuperscript{41} the
potential part of the symmetry energy is defined with two paramet ers, the value at
saturation density, usually taken as $22$ MeV in their calculations, and the power-law
coefficient $\gamma$ describing the dependence on density,

$$E_{\text{sym}} = E_{\text{sym}}^{\text{pot}} + E_{\text{sym}}^{\text{kin}} = 22 \text{ MeV} \cdot (\rho/\rho_0)^\gamma + 12 \text{ MeV} \cdot (\rho/\rho_0)^{2/3}. \quad (2)$$

In other codes the nuclear potential of Das \textit{et al.} with explicit momentum dependence in the isovector sector is used.\textsuperscript{12} There, as in the IBUU04 developed by the
groups of Li and Chen,\textsuperscript{43,44} the density dependence of the symmetry energy is char-
acterized by a parameter $x$ appearing in the potential expressions. Examples of
these parameterizations and of results obtained from the analysis of experimental reaction data are given in Fig. 1. The stiff ($E_{\text{sym}}^a$) and soft ($E_{\text{sym}}^b$) density dependences of Fig. 1 correspond approximately to the cases $\gamma = 1$ and $x = 1$ shown here.

Parameterizations of this kind have the consequence that, once the symmetry energy at the saturation point is fixed, a single value at a different density or, alternatively, the slope or curvature at any density will completely determine the parameterization. Measurements of a variety of observables in nuclear structure and reactions have, therefore, been used to obtain in this way results for the density dependence of the symmetry energy. They are often expressed in the form of the parameter $L$ which is proportional to the slope at saturation,

$$L = 3\rho_0 \cdot \frac{dE_{\text{sym}}}{d\rho}|_{\rho_0}. \quad (3)$$

Most results with their errors fall into the interval $20 \text{ MeV} \leq L \leq 100 \text{ MeV}$ and are compatible with a most probable value $L \approx 60 \text{ MeV}$, roughly corresponding to a power-law coefficient $\gamma = 0.65$. The full line in Fig. 1 represents, e.g., the result deduced by Li and Chen from the MSU isospin-diffusion data and the neutron-skin thickness in $^{208}\text{Pb}$. The corresponding slope parameter is $L = 65 \text{ MeV}$. Rather similar constraints have been deduced from very recent investigations and observations of neutron-star properties.

Considerable progress regarding the correlation of the symmetry energy with particular observables in different models has been made by the Florida and Barcelona groups. In continuation of the work of Typel and Brown and of Furnstahl, a universal correlation between the thickness of the neutron-skin of $^{208}\text{Pb}$ and the $L$ parameter has been found for empirical mean-field interactions. The determination of the neutron-skin thickness by measuring the parity-violating contribution to electron scattering at high energy will thus offer a practically model-free access to the slope at saturation, even though it is obtained by probing nuclear matter at an average density of typically $2/3$ of this value.

3. Dynamics of Heavy-Ion Reactions

Densities of two to three times the saturation density may be reached on time scales of $\approx 10 \sim 20 \text{ fm/c}$ in the central zone of heavy-ion collisions at relativistic energies of up to $\approx 1 \text{ GeV/nucleon}$. The resulting pressure produces a collective outward motion of the compressed material whose strength will be influenced by the symmetry energy in asymmetric systems. At the same time, the excitation of $\Delta$ resonances in hard nucleon-nucleon scatterings leads to the production and subsequent emission of charged and neutral $\pi$ and $K$ mesons. The relative intensities of isovectors pairs of mesons depend directly on the proton-neutron content of the site where they are produced, suggesting them as sensitive probes for the high-density symmetry energy.

Measurements of $K^+/K^0$ production ratios have been performed by the FOPI Collaboration for the neutron-rich and neutron-poor $A = 96$ systems $^{96}\text{Zr} + ^{96}\text{Zr}$.
and $^{96}$Ru + $^{96}$Ru. The experimental acceptances for these particles were considerably different, because of their different decay channels, and the precision required for an isotopic study was more reliably achieved with double ratios obtained from the data for the two systems with different isotopic compositions. A significant sensitivity to the symmetry energy was expected from calculations under equilibrium conditions but was found to be considerably reduced when actual collisions were modeled.

The FOPI Collaboration has also presented an extensive account of the $\pi^-/\pi^+$ production ratios measured for the reaction $^{197}$Au + $^{197}$Au at several energies up to 1.5 GeV/nucleon. Theoretically, these data, however, come to rather conflicting conclusions, suggesting everything from a rather stiff to a super-soft behavior of the symmetry energy. The reason may lie in the treatment of the $\Delta$ dynamics and in competing effects of the mean fields and the $\Delta$ thresholds which will require further studies. In the analysis of Xiao et al. with the IBUU04 transport model, the measured system and impact-parameter dependences have been reproduced, however, only by assuming a super-soft dependence of the symmetry energy on density, close to the $\chi = +1$ case shown in Fig. 4. The presently inconclusive situation for pion ratios is very unfortunate because variations of up to 20% for soft versus stiff parameterizations are expected.

Dynamical flow observables have been proposed by several groups as probes for the equation of state at high density, among them the so-called differential neutron-proton flow which is the difference of the parameters describing the collective motion of free neutrons and protons weighted by their numbers. As pointed out by Yong et al., this observable minimizes the influence of the isoscalar part in the EoS while maximizing that of the symmetry term. Its proportionality to the particle multiplicities, however, makes its determination very dependent on the experimental efficiencies of particle detection and identification and on the precise procedure for distinguishing free and bound nucleons in calculations. Therefore, also flow differences or ratios have been considered.

4. Elliptic Flow

The so-called squeeze-out of nuclear matter from the compressed interaction zone has first been observed in experiments at the Bevalac. In a sphericity analysis, the event shape in three dimensions was characterized by a kinetic-energy flow tensor whose main orientation with respect to the beam direction represents the collective sideways flow and whose cross section, if it is non-isotropic, indicates the existence of elliptic flow. At the bombarding energies of up to 1.05 GeV/nucleon investigated in these studies, a preferential emission of charged particles perpendicular to the reaction plane has been observed. The shadowing by the spectator remnants as they pass each other during the collision reduces the in-plane flow, so that the strength of the off-plane emission, as quantified by the azimuthal anisotropy, reflects the internal pressure.
Elliptic Flow and the Symmetry Energy

Fig. 5. Elliptic flow parameter \( v_2 \) at mid-rapidity for \(^{197}\text{Au}+^{197}\text{Au} \) collisions at intermediate impact parameters (about 5.5-7.5 fm) as a function of incident energy, in the beam frame. The filled and open circles represent the INDRA and FOPI data, respectively, for \( Z = 1 \) particles, the triangles represent the EOS and E895 data for protons and the square represents the E877 data for all charged particles (from Ref. 69, reprinted with kind permission from Springer Science+Business Media).

It has become customary to express both, directed and elliptic flows, and possibly also higher flow components by means of a Fourier decomposition of the azimuthal distributions measured with respect to the orientation of the reaction plane \( \phi_R \):

\[
\frac{dN}{d(\phi - \phi_R)} = \frac{N_0}{2\pi} \left( 1 + 2 \sum_{n \geq 1} v_n \cos n(\phi - \phi_R) \right),
\]

where \( N_0 \) is the azimuthally integrated yield. The coefficients \( v_n \equiv \langle \cos n(\phi - \phi_R) \rangle \) are functions of particle type, impact parameter, rapidity \( y \), and the transverse momentum \( p_t \); \( v_1 \) and \( v_2 \) represent the directed and elliptic flows, respectively.

Elliptic flow has become an important observable at other energy regimes as well. At ultrarelativistic energies, the observation of the constituent-quark scaling of elliptic flow is one of the prime arguments for deconfinement during the early collision phase, and properties of the formed quark-gluon liquid are deduced from the observed magnitude of collective motions. It implies that elliptic flow develops very early in the collision which is valid also in the present range of relativistic energies, as confirmed by calculations. Isotopic flow differences appear thus very suitable for studying mean-field effects at high density.

An excitation function of the elliptic flow of \( Z = 1 \) particles in \(^{197}\text{Au}+^{197}\text{Au} \) collisions from various experiments is shown in Fig. 5. Squeeze-out perpendicular to the reaction plane, i.e. \( v_2 < 0 \), as a result of shadowing by the spectator remnants is
observed at incident energies between about 150 MeV/nucleon and 4 GeV/nucleon with a maximum near 400 MeV/nucleon. At lower energies, the collective angular momentum in the mean-field dominated dynamics cause the observed in-plane enhancement of emitted reaction products. The figure also shows that elliptic flow can be measured quite precisely. The reliability of the applied methods is being demonstrated by the good agreement of data sets from different experiments in the overlap regions of the studied intervals in collision energy.

Fig. 6. Measured flow parameters $v_1$ (top) and $v_2$ (bottom) for mid-peripheral ($5.5 \leq b \leq 7.5$ fm) $^{197}$Au + $^{197}$Au collisions at 400 MeV/nucleon for neutrons (dots), protons (circles), and hydrogen isotopes ($Z = 1$, open triangles) integrated within $0.3 \leq p_t/A \leq 1.3$ GeV/c/nucleon as a function of the normalized rapidity $y_{lab}/y_p$. The UrQMD predictions for neutrons are shown for the FP1 parametrization of the in-medium cross sections and for a stiff ($\gamma = 1.5$, full lines) and a soft ($\gamma = 0.5$, dashed) density dependence of the symmetry term. The experimental data have been corrected for the dispersion of the reaction plane (reprinted from Ref. 21, Copyright (2011), with permission from Elsevier).

5. Results from FOPI/LAND

The squeeze-out of neutrons has first been observed by the FOPI/LAND Collaboration who studied the reaction $^{197}$Au + $^{197}$Au at 400 MeV/nucleon. The squeeze-out of charged particles reaches its maximum at this energy (Fig. 5), and similarly large anisotropies were observed for neutrons. The neutrons had been detected with the Large-Area Neutron Detector, LAND, while the FOPI Forward Wall, covering the forward hemisphere of laboratory angles $\theta_{lab} \leq 30^\circ$ with more than 700 plastic scintillator elements, was used to determine the modulus and azimuthal orientation of the impact parameter.
The motivation for returning to the existing data set has been provided by studies performed with the UrQMD transport code for this fairly neutron-rich system \((N/Z = 1.49)\) which indicated a significant sensitivity of the elliptic-flow parameters to the assumptions made for the density dependence of the symmetry energy\(^{[21]}\). In calculations with power-law coefficients \(\gamma = 0.5\) and \(1.5\) (cf. Eq. 2 and Fig. 4), the relative strengths of neutron and proton elliptic flows were found to vary on the level of 15%.

The performed reanalysis of the data consisted mainly in choosing equal acceptances for neutrons and hydrogen isotopes with regard to particle energy, rapidity and transverse momentum (energy and momentum per nucleon for deuterons and tritons). The results obtained for a mid-peripheral event class are shown in Fig. 6 as a function of the rapidity \(y\), normalized with respect to the projectile rapidity \(y_p = 0.896\). Their asymmetry with respect to mid-rapidity, \(y/y_p = 0.5\), is caused by the LAND detector whose kinematic acceptance in \(p_t\) increases with \(y\). With the yields dropping at large \(p_t\), the statistical errors become large at forward rapidity.

The theoretical predictions have been obtained simulating the LAND acceptance and the experimental analysis conditions. The results, shown for neutrons in Fig. 6, follow qualitatively the experimental data.

According to the UrQMD model, the sensitivity of the directed flow of neutrons to the stiffness of the symmetry energy is nearly negligible (Fig. 6 top panel) while it is significant for the elliptic flow (bottom panel). This is evident also from the dependence of the elliptic flow parameter \(v_2\) on the transverse momentum per nucleon, \(p_t/A\), shown in Fig. 7 upper panel, for the full statistics of central and mid-peripheral collisions \((b \leq 7.5\ \text{fm})\) collected in this experiment. The measured values are approximately reproduced by the UrQMD predictions which are significantly different for the stiff \((\gamma = 1.5)\) and soft \((\gamma = 0.5)\) density dependences.

6. Elliptic-Flow Ratios

For the quantitative evaluation, the ratio of the flow parameters of neutrons versus protons or versus \(Z = 1\) particles has been proposed as a useful observable\(^{[21]}\). Systematic effects influencing the collective flows of neutrons and charged particles in similar ways should thereby be minimized, on the experimental as well as on the theoretical side. These include, e.g., the existing uncertainties of the isoscalar EoS which will affect the absolute magnitude of calculated flows and the matching of the impact-parameter intervals used in the calculations with the corresponding experimental event groups. To demonstrate this kind of insensitivity, the calculations were performed with two parameterizations of the momentum dependence of the elastic nucleon-nucleon cross section, labelled FP1 and FP2, which differ in their absolute predictions of \(v_2\) at mid-rapidity by \(\approx 40\%\) for this reaction\(^{[71]}\).

In Fig. 7 (lower panels), the results for the ratio with respect to the total hydrogen yield are shown. The calculated ratios exhibit clearly the sensitivity of the elliptic flow to the stiffness of the symmetry energy predicted by the UrQMD but
Fig. 7. Elliptic flow parameters $v_2$ for neutrons (dots) and hydrogen isotopes (open triangles, top panel) and their ratio (lower panels) for moderately central ($b < 7.5$ fm) collisions of $^{197}$Au + $^{197}$Au at 400 MeV/nucleon, integrated within the rapidity interval $0.25 \leq y/y_p \leq 0.75$, as a function of the transverse momentum per nucleon $p_t/A$. The symbols represent the experimental data. The UrQMD predictions for $\gamma = 1.5$ (a-stiff) and $\gamma = 0.5$ (a-soft) obtained with the FP1 parameterization for neutrons (top panel) and for the ratio (middle panel), and with the FP2 parameterization for the ratio (bottom panel) are given by the dashed lines (reprinted from Ref. [21], Copyright (2011), with permission from Elsevier).

The error is larger for FP2 because the sensitivity is somewhat smaller (Fig. 7, bottom panel).

This analysis was repeated in various forms. With the squeeze-out ratios $v_2^n/v_2^p$ of neutrons with respect to free protons, similar results were obtained, however with larger errors. The study of the impact parameter dependence indicated a slightly
Fig. 8. Sensitivity of the neutron-proton elliptic flow difference \( v_2^{n-p} = v_2^n - v_2^p \) (\( v_2 \) as defined in Eq. 4) to the isospin independent EoS and to the symmetry term in the parameterization of Ref. 43 as indicated. The widths of the bands represent the uncertainty arising from using either a soft (\( K = 210 \) MeV with momentum dependence) or a hard (\( K = 380 \) MeV) version for the isospin-independent part of the EoS (reprinted from Ref. 30, Copyright (2011), with permission from Elsevier).

smaller value \( \gamma \approx 0.5 \) for the mid-peripheral event group, again with larger errors. It was also tested to which density region around \( \rho_0 \) the elliptic-flow ratios are sensitive, with the result that both, sub- and supra-saturation densities are probed with this observable.\(^{21}\)

In consideration of the apparent systematic and experimental errors, a value \( \gamma = 0.9 \pm 0.4 \) has been adopted by the authors as best representing the power-law exponent of the potential term resulting from the elliptic-flow analysis. It falls slightly below the \( \gamma = 1.0 \) line shown in Fig. 4 but, with the quoted uncertainty, stretches over the interval from \( \gamma = 0.5 \) halfway up to \( \gamma = 1.5 \). The corresponding slope parameter is \( L = 83 \pm 26 \) MeV. The squeeze-out data thus indicate a moderately soft to linear behavior of the symmetry energy that is consistent with the density dependence deduced from experiments probing nuclear matter near or below saturation. Comparing with the many-body theories shown in Fig. 3, the elliptic-flow result is in good qualitative agreement with the range spanned by the DBHF and variational calculations based on realistic nuclear potentials.

7. Elliptic-Flow Differences

In an independent analysis, Cozma has used data from the same experiment and investigated the influence of several parameters on the difference between the elliptic flows of protons and neutrons using the Tübingen version of the QMD trans-
port model. They included the parameterization of the isoscalar EoS, the choice of various forms of free or in-medium nucleon-nucleon cross sections, and model parameters as, e.g., the widths of the wave packets representing nucleons. The interaction developed by Das et al. was used which contains an explicit momentum dependence of the symmetry energy part. In Fig. the difference of neutron and proton elliptic flows for five choices of the $x$ parameter describing the density dependence of the symmetry energy is reproduced from this work. It shows the large sensitivity to the stiffness of the symmetry energy (different $x$-parameters) and the reduced influence of the isoscalar EoS (widths of the bands). A super-soft behavior of the symmetry energy was confirmed to be excluded by the comparison with the experimental flow data.

Because the flow difference is still proportional to the overall magnitude of the elliptic flow predicted by the model, also the ratios of the neutron-versus-proton elliptic-flow parameters have been calculated and compared to the FOPI/LAND data. The best description is obtained with a density dependence slightly stiffer than the $x = 0$ solution which approximately corresponds to the $\gamma = 0.5$ case shown in Fig. It is thus rather close to the UrQMD result $\gamma = 0.9$ which may represent an important step towards the model invariance that one would ultimately like to achieve, not only because of the invariance with respect to the treatment of the nucleon-nucleon cross sections observed in the two studies but also because of the explicit momentum dependence of the isovector potentials which is implemented in the Tübingen QMD but not in the UrQMD version used.

8. Conclusion and Outlook

According to the predictions of transport models, the relative strengths of neutron and proton elliptic flows represent an observable sensitive to the symmetry energy at densities near and above saturation. By forming ratios or differences of neutron versus proton or neutron versus hydrogen flows, the influence of isoscalar-type parameters of the model descriptions can be minimized.

The comparison of the results obtained from the FOPI/LAND data for $^{197}$Au + $^{197}$Au reactions at 400 MeV/nucleon with UrQMD transport calculations favors a moderately soft to linear symmetry term with a density dependence of the potential term proportional to $(\rho/\rho_0)^\gamma$ with $\gamma = 0.9 \pm 0.4$, compatible with predictions of ab-initio calculations. With the Tübingen QMD model, a similar result is obtained.

The explicit proof that the elliptic-flow ratios are probing the nuclear mean field at super-saturation densities is, so far, limited to test calculations made with the UrQMD. They indicate that the strength of the symmetry energy at densities both, below and above saturation, are essential. It will be useful to study this in more detail.

The statistical uncertainty of the existing FOPI/LAND data is larger than the investigated systematic effects of model parameters and analysis techniques. More significant data can thus be expected from a new experiment delivering a compre-
hensive data set with sufficiently high statistical accuracy. Explicitly testing the predicted dependences on rapidity and transverse momentum will provide useful constraints for the models. Additional observables as, e.g., neutron-to-proton or $^3$H-to-$^3$He yield ratios may serve in resolving ambiguities caused by the effective-mass splitting between neutrons and protons. It will be interesting to see first results appearing from the recently completed measurements of the ASYEOS Collaboration at the GSI Laboratory.

Because of the quadratically rising importance of the symmetry energy, the continuation of this program with systems of larger asymmetry is very promising and important. The lower luminosities to be expected from the use of secondary beams and isotopically enriched targets will have to be compensated with efficient detector setups. The neutron detector NeuLAND proposed for experiments at FAIR will offer a highly improved detection efficiency for neutrons in the energy range 100 to 400 MeV. This will be essential for extending the program also to reactions at lower energies for which significant mean-field effects are predicted for directed and elliptic flows.

Stimulating and fruitful discussions with M.D. Cozma, M. Di Toro, W. Reisdorf, and with the authors of Ref. 21, P. Russotto, P.Z. Wu, M. Zoric, M. Chartier, Y. Leifels, R.C. Lemmon, Q. Li, J. Lukasik, A. Pagano, and P. Pawłowski are gratefully acknowledged.

References

1. P. Danielewicz, R. Lacey, W.G. Lynch, Science 298 (2002) 1592.
2. C. Sturm et al., Phys. Rev. Lett. 86 (2001) 39.
3. C. Fuchs, A. Faessler, E. Zabrodin, Y.M. Zheng, Phys. Rev. Lett. 86 (2001) 1974.
4. J.M. Lattimer and M. Prakash, Phys. Rep. 442 (2007) 109.
5. Bao-An Li, Lie-Wen Chen, Che Ming Ko, Phys. Rep. 464 (2008) 113.
6. M. Di Toro, V. Baran, M. Colonna, V. Greco, J. Phys. G 37 (2010) 083101.
7. M.B. Tsang et al., Phys. Rev. Lett. 102 (2009) 122701.
8. J.B. Natowitz et al., Phys. Rev. Lett. 104 (2010) 202501.
9. C.J. Horowitz and A. Schwenk, Nucl. Phys. A 776 (2006) 55.
10. S. Typel et al., Phys. Rev. C 81 (2010) 015803.
11. B.A. Brown, Phys. Rev. Lett. 85 (2000) 5296.
12. C. Fuchs and H.H. Wolter, Eur. Phys. J. A 30 (2006) 5.
13. R. Subedi et al., Science 320 (2008) 1476.
14. Chang Xu and Bao-An Li, Phys. Rev. C 81 (2010) 064612.
15. A.W. Steiner and S. Gandolfi, Phys. Rev. Lett. 108 (2012) 081102.
16. A. Carbone, A. Polls, A. Rios, Europhys. Lett. 97 (2012) 22001.
17. V. Baran, M. Colonna, V. Greco, M. Di Toro, Phys. Rep. 410 (2005) 335.
18. A. Klimkiewicz et al., Phys. Rev. C 76 (2007) 051603(R).
19. A. Tamii et al., Phys. Rev. Lett. 107 (2011) 062502.
20. Bao-An Li, Phys. Rev. Lett. 88 (2002) 192701.
21. P. Russotto et al., Phys. Lett. B 697 (2011) 471.
22. Y. Leifels et al., Phys. Rev. Lett. 71 (1993) 963.
23. D. Lambrecht et al., Z. Phys. A 350 (1994) 115.
24. W. Reisdorf et al., Nucl. Phys. A 781 (2007) 459.
25. G. Ferini et al., *Nucl. Phys. A* **762** (2005) 147; *Phys. Rev. Lett.* **97** (2006) 202301.
26. Zhigang Xiao et al., *Phys. Rev. Lett.* **102** (2009) 062502.
27. Zhao-Qing Feng and Gen-Ming Jin, *Phys. Lett. B* **683** (2010) 140.
28. Bao-An Li et al., *J. Phys. Conf. Ser.* **312** (2011) 042006.
29. De-Hua Wen, Bao-An Li, Lie-Wen Chen, *Phys. Rev. Lett.* **103** (2009) 211102.
30. M.D. Cozma, *Phys. Lett. B* **700** (2011) 139.
31. Th. Blaich et al., *Nucl. Instrum. Methods Phys. Res. A* **314** (1992) 136.
32. A. Pagano et al., *Nucl. Phys. A* **734** (2004) 504.
33. R.C. Lemmon et al., proposal for SIS experiment S394 (2009).
34. M. Baldo, C. Maieron, P. Schuck, X. Vinas, *Nucl. Phys. A* **736** (2004) 241.
35. R.B. Wiringa and S.C. Pieper, *Phys. Rev. Lett.* **89** (2002) 182501.
36. F. Burgio, *J. Phys. G* **35** (2008) 014048.
37. K. Hebeler et al., *Phys. Rev. Lett.* **105** (2010) 161102.
38. A.W. Steiner, J.M. Lattimer, E.F. Brown, *ApJ.* **722** (2010) 33.
39. V. Giordano et al., *Phys. Rev. C* **81** (2010) 044611.
40. Zhao-Qing Feng, *Phys. Lett. B* **707** (2012) 83.
41. Q. Li et al., *J. Phys. G* **31** (2005) 1359; ibid. **32** (2006) 151; ibid. **32** (2006) 407.
42. C.B. Das, S. Das Gupta, C. Gale, Bao-An Li, *Phys. Rev. C* **67** (2003) 034611.
43. Bao-An Li and Lie-Wen Chen, *Phys. Rev. C* **72** (2005) 064611.
44. Chen Chen Guo et al., *Science China Physics, Mechanics & Astronomy* **55** (2012) 252.
45. J.M. Lattimer and Y. Lim, preprint arXiv:1203.4286[nucl-th] (2012).
46. B.G. Todd-Rutel and J. Piekarewicz, *Phys. Rev. Lett.* **95** (2005) 122501.
47. X. Roca-Maza, M. Centelles, X. Vinas, M. Warda, *Phys. Rev. Lett.* **106** (2011) 252501.
48. S. Typel and B.A. Brown, *Phys. Rev. C* **64** (2001) 027302.
49. R.J. Furnstahl, *Nucl. Phys. A* **706** (2002) 85.
50. C.J. Horowitz et al., *Phys. Rev. C* **85** (2012) 032501(R).
51. Bao-An Li, *Nucl. Phys. A* **708** (2002) 365.
52. R. Stock, *Phys. Rep.* **135** (1986) 259.
53. X. Lopez et al., *Phys. Rev. C* **75** (2007) 011901(R).
54. V. Greco et al., *Phys. Lett. B* **562** (2003) 215.
55. Gao-Chan Yong, Bao-An Li, Lie-Wen Chen, *Phys. Rev. C* **74** (2006) 064617.
56. H.H. Gutbrod et al., *Phys. Rev. C* **42** (1990) 640.
57. S. Voloshin and Y. Zhang, *Z. Phys. C* **70** (1996) 665.
58. J.-Y. Ollitrault, preprint arXiv:nucl-ex/9711003 (1997).
59. A.M. Poskanzer and S.A. Voloshin, *Phys. Rev. C* **58** (1998) 1671.
60. B.I. Abelev et al., *Phys. Rev. Lett.* **99** (2007) 112301.
61. R. Fries, V. Greco, P. Sorensen, *Annu. Rev. Nucl. Part. Sci.* **58** (2008) 177.
62. R. Snellings, *New J. Phys.* **13** (2011) 055008.
63. A. Adare et al., *Phys. Rev. C* **85** (2012) 064914.
64. P. Danielewicz, *Nucl. Phys. A* **673** (2000) 375.
65. J. Lukasik et al., *Phys. Lett. B* **608** (2005) 223.
66. A. Andronic et al., *Phys. Lett. B* **612** (2005) 173.
67. C. Pinkenburg et al., *Phys. Rev. Lett.* **83** (1999) 1295.
68. P. Braun-Munzinger and J. Stachel, *Nucl. Phys. A* **638** (1998) 3c.
69. A. Andronic, J. Lukasik, W. Reisdorf, W. Trautmann, *Eur. Phys. J. A* **30** (2006) 31.
70. W. Reisdorf et al., *Nucl. Phys. A* **876** (2012) 1.
71. Qingfeng Li, Caiwan Shen, M. Di Toro, *Mod. Phys. Lett. A* **25** (2010) 669.
72. Qingfeng Li et al., *Phys. Rev. C* **83** (2011) 044617.
73. M.D. Cozma, private communication (2011).
74. NeuLAND Technical Design Report, submitted to FAIR (2011).