Improved eV-scale sterile-neutrino constraints from the second KATRIN measurement campaign

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We present the results of the light sterile neutrino search from the second Karlsruhe Tritium Neutrino (KATRIN) measurement campaign in 2019. Approaching nominal activity, $3.76 \times 10^8$ tritium $\beta$-electrons are analyzed in an energy window extending down to 40 eV below the tritium end point at $E_0 = 18.57$ keV. We consider the $3\nu + 1$ framework with three active and one sterile neutrino flavors. The analysis is sensitive to a fourth mass eigenstate $m_4^2 \lesssim 1600$ eV$^2$ and active-to-sterile mixing $|U_{ei}|^2 \gtrsim 6 \times 10^{-3}$. As no sterile-neutrino signal was observed, we provide improved exclusion contours on $m_4^2$ and $|U_{ei}|^2$ at 95% C.L. Our results supersede the limits from the Mainz and Troitsk experiments. Furthermore, we are able to exclude the large $\Delta m^2_{21}$ solutions of the reactor antineutrino and gallium anomalies to a great extent. The latter has recently been reaffirmed by the BEST Collaboration and could be explained by a sterile neutrino with large mixing. While the remaining solutions at small $\Delta m^2_{21}$ are mostly excluded by short-baseline reactor experiments, KATRIN is the only ongoing laboratory experiment to be sensitive to relevant solutions at large $\Delta m^2_{21}$ through a robust spectral shape analysis.

I. INTRODUCTION

The Karlsruhe Tritium Neutrino (KATRIN) experiment \cite{1,2} is designed to determine the absolute neutrino-mass scale via the kinematics of single $\beta$-decay of molecular tritium

$$ T_2 \rightarrow ^3\text{He}^+ + e^- + \bar{\nu}_e $$

with an unprecedented sensitivity 0.2 eV (90% C.L.) after five years of measurement time \cite{1}. This is achieved by measuring the integrated $\beta$-electron spectrum in a narrow energy interval around the tritium end point at $E_0 = 18.57$ keV. The three known neutrino mass eigenstates $m_i$ lead to a reduction of the maximal observed electron energy as well as to a slight spectral shape distortion. As the mass-squared splittings $\Delta m^2_{ij}$ are known to be small compared to the energy resolution of KATRIN \cite{3}, the observable is the squared effective electron antineutrino mass $^{3}m_e^2$

$$ m^2_e = \sum_{i=1}^{3} |U_{ei}|^2 m_i^2, $$

where $U_{ei}$ are the elements of the Pontecorvo-Maki-Nagawa-Sakata (PMNS) matrix \cite{3}. To date, KATRIN provides the most stringent upper limit from a direct measurement of $m_e \leq 0.8$ eV (90% C.L.) \cite{4} following its limit of $m_e \leq 1.1$ eV (90% C.L.) \cite{5} from the first measurement campaign.

Using the same datasets, this work investigates the existence of a fourth neutrino mass eigenstate $m_4$ with active-to-sterile mixing $^{2}|U_{ei}|^2$. Based on the measured width of the $Z^0$ resonance, it is well established that there are only three light active neutrinos \cite{6}. Therefore, $m_4$ would be mostly composed of a sterile neutrino flavor that does not participate in the weak interaction. The signature of a sterile neutrino in KATRIN is a kinklike spectral distortion, that is most prominent at electron energies around $E_0 - m_4$. Since the neutrino-mass campaigns focus on measuring the $\beta$-spectrum in the vicinity of the end point, our sterile-neutrino analysis is restricted to light sterile neutrinos at the eV scale. An extension of the measured energy range to the complete tritium $\beta$-decay spectrum would offer the opportunity to search for keV-scale sterile neutrinos with KATRIN. The TRISTAN project plans to extend the KATRIN setup with a novel detector system after completion of the neutrino-mass campaigns to handle the high rates involved in such a measurement \cite{7}.

Light sterile neutrinos are motivated by accumulating anomalies in short-baseline neutrino oscillation experiments. Studying the appearance of $\nu_e$ from an accelerator

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n_\mu \text{ beam, the results from LSND and MiniBooNE suggest evidence for nonstandard neutrino oscillations involving a new neutrino mass eigenstate [8,9]. Moreover, the gallium anomaly (GA), observed by both GALLEX and SAGE and recently reaffirmed by BEST, describes a } \nu_e \text{ deficit from } ^{37}\text{Ar} \text{ and } ^{51}\text{Cr} \text{ electron capture decays [10–12]. Additionally, a significant discrepancy between the predicted and observed } \bar{\nu}_e \text{ flux from nuclear reactors, denoted as the reactor antineutrino anomaly (RAA), has been found [13]. These neutrino disappearance anomalies could be explained with the existence of a } \gtrsim 1 \text{ eV sterile neutrino. However, due to difficulties in assessing systematic uncertainties, these anomalies are debated. The recent } > 3\sigma \text{ claim on a } \sim 2.7 \text{ eV sterile neutrino with large active-to-sterile mixing by the Neutrino-4 } \bar{\nu}_e \text{ reactor experiment kindled a controversial discussion [14].} 

KATRIN offers an approach complementary to short-baseline neutrino oscillation experiments, to search for light sterile neutrinos. Being sensitive to the same parameters that could explain the GA and the RAA, KATRIN is able to probe both anomalies in an independent way. This was first demonstrated with data from the first science run (KNM1), in which parts of the parameter space covered by the sterile-neutrino anomalies could be constrained [15]. Here, we present the results of the second measurement campaign (KNM2) as well as the combination of the first two campaigns, improving with respect to the previous KATRIN exclusion bounds.

II. EXPERIMENTAL SETUP

The prerequisites to measure the subtle imprints of } m_\ell^2 \text{ and } m_\nu^2 \text{ in the tritium } \beta \text{-decay spectrum are a high source activity } \mathcal{O}(10^{11} \text{ Bq}), \text{ a low background rate } \mathcal{O}(0.1 \text{ cps}), \text{ and an eV-scale energy resolution. To achieve this, the 70 m long KATRIN experiment combines a windowless gaseous tritium source (WGTS) with a high-precision Magnetic Adiabatic Collimation and Electrostatic (MAC-E) filter [1,2]. An overview of the experimental setup is displayed in Fig. 1. High-purity molecular tritium gas (>95%) is continuously injected at the center of the 10 m long WGTS, keeping the source activity stable at the 0.2% level. To minimize thermal gas motion, the WGTS is cooled to 30 K [16], which is above the freezing point of T_2. Moving from the source toward the spectrometer, the partial tritium pressure is reduced by more than 14 orders of magnitude in the transport and pumping section. This reduction is achieved through a differential and a cryogenic pumping section [17]. Electrons from the source are magnetically guided through the setup [18], starting in a magnetic field of } B_{\text{source}} = 2.5 \text{ T.} 

Applying the MAC-E-filter principle [19,20], the energies of the electrons are evaluated in the spectrometer section consisting of the pre- and main spectrometers. By gradually reducing the magnetic field toward the analyzing plane to } B_{\text{ana}} = 6.3 \times 10^{-4} \text{ T in the main spectrometer, the momenta of all electrons are adiabatically collimated. At the exit of the main spectrometer, the flux tube is tapered by the magnetic field } B_{\text{max}} = 4.2 \text{ T of the superconducting pinch magnet. Both spectrometers act as electrostatic high-pass filters, allowing only electrons with kinetic energies larger than the applied retarding energy } qU \text{ to be transmitted. Given the magnetic field configuration, the filter width at the end point is } \Delta E = 18.6 \text{ keV} \cdot (B_{\text{ana}}/B_{\text{max}}) = 2.78 \text{ eV with a maximum accepted angle of } \theta_{\text{max}} = \sqrt{B_{\text{source}}/B_{\text{max}}} = 50.4^\circ. \text{ Electrons transmitted through the main spectrometer are finally counted as a function of } qU \text{ at the focal plane detector (FPD) [21]. The FPD is a monolithic silicon PIN-diode segmented into 148 pixels of equal area. At the upstream end of the experiment, nontransmitted electrons are eventually absorbed by a gold-plated rear wall. To homogenize the source electric potential, the rear wall is biased to a voltage of } \mathcal{O}(100 \text{ mV}), \text{ compensating for the intrinsic potential difference between the rear wall and the beam-tube surface.}

III. KNM2 MEASUREMENT CAMPAIGN

In the following, we describe the second measurement campaign (KNM2). Section III A addresses the operating
conditions of the KATRIN experiment as well as the measurement procedure. The data selection is presented in Sec. III B, and the data combination is presented in Sec. III C.

A. Measurement of the integral tritium spectrum

The second high-purity tritium campaign was conducted in October and November of 2019. We achieved a source activity of \(9.5 \times 10^{10}\) Bq, improving with respect to our first measurement campaign by a factor of 3.8. The operation with \(\rho = 4.2 \times 10^{17}\) molecules/cm\(^2\), close to the designed nominal column density, was possible as we are no longer limited to operation in the “burn-in” configuration [22]. By continuously monitoring the gas composition with a laser Raman system, we determined the isotopic tritium purity as \(\epsilon_T = 98.7\%\) [23]. The background rate was reduced by 25% to 220 mcps following an improvement of the vacuum conditions in the main spectrometer [4]. We recorded the integrated \(\beta\)-spectrum by repeatedly measuring the count rate \(R_{\text{data}}(qU_i)\) at 39 different retarding energies in a range of \([E_0 - 300\ eV, E_0 + 135\ eV]\). One measurement at a given \(qU_i\) is called a scan step and lasts between 17 and 576 s. The ensemble of all 39 scan steps, referred to as a scan, had a duration of 2 h. The measurement time distribution within a scan is displayed in Fig. 2(c) for the analyzed energy range. It was optimized with respect to the neutrino mass sensitivity prior to the measurement campaign. The scan steps above \(E_0\) are used to determine the constant background rate.

B. Data selection

In this work, we limit the analysis to the energy range \([E_0 - 40\ eV, E_0 + 135\ eV]\), shown in Fig. 2, in which the measurement is dominated by statistical uncertainties. Applying data-quality criteria identical to Ref. [4] results in 361 out of 397 recorded scans selected for the analysis. The rejection of the other scans is based on insufficient gas-composition data as well as corrupted settings of the high voltage. The accepted measurement data amount to a total scan time of 694 h. From the available 148 pixels of the FPD, we select 117 pixels for the analysis. The rejected pixels do not satisfy quality requirements, as they exhibit an increased noise level or a broadened energy resolution or suffer from misalignment of the beam line with respect to the magnetic flux tube.

C. Data combination

The sub-ppm precision level of the high-voltage (HV) system, combined with a source-potential stability of \(\sigma < 80\ mV\), allows us to combine the selected scans into one effective spectrum with averaged HV values \((qU_i)\). The temporal source-potential variations are incorporated in our model as an energy broadening in the final-state distribution (see Sec. IV). In addition, the excellent homogeneity of the electric (\(\sigma < 41\ mV\)) and magnetic (\(\sigma < 10^{-6}\ T\)) fields in the analyzing plane as well as consistent pixel characteristics justify the combination of all selected pixels to one effective detector area (uniform fit). Possible radially dependent source-potential variations, estimated with a fit that allows for a radial-dependent effective end point, are found to be negligible (\(<100\ meV\)) [4]. The combined data spectrum that is used for the final spectral fit is displayed in Fig. 2(a). In the 40 eV energy range below the end point, \(3.76 \times 10^6\) tritium \(\beta\)-electrons and \(0.41 \times 10^6\) background electrons\(^3\) were counted. While

\(^3\)If the five scan steps above \(E_0\) are also taken into account, the total number of background electrons increases to \(0.55 \times 10^6\).
the number of background electrons has only a slight retarding energy dependence, the number of signal electrons increases steeply with decreasing \( qU \). Therefore, we observe \( qU \)-dependent signal-to-background ratios of \((242, 88, 21, 1)\) at \((40, 30, 20, 10)\) eV below the fit end point.

\[ \text{IV. EXPERIMENTAL MODELING} \]

The following section will give an overview of the analytical model describing the measured \( \beta \)-spectrum. More details can be found in Ref. [24].

\[ \text{A. Differential spectrum} \]

The differential spectrum \( R_{q}(E) \) of the superallowed tritium \( \beta \)-decay can be derived using Fermi’s golden rule:

\[
R_{\beta}(E, m_{\nu}^{2}) = \frac{G_{F}^{2} \cdot \cos^{2} \Theta_{C} \cdot |M_{\text{nucl}}|^{2} \cdot F(Z', E)}{2\pi}
\cdot (E + m_{e}) \cdot \sqrt{(E + m_{e})^{2} - m_{\nu}^{2}}
\cdot \sum_{f} \xi_{f} \epsilon_{f}(E) \sqrt{(\epsilon_{f}(E))^{2} - m_{\nu}^{2}}
\cdot \Theta(\epsilon_{f}(E) - m_{\nu}).
\]

(3)

Here, \( G_{F} \) is the Fermi constant, \( \Theta_{C} \) is the Cabibbo angle, and \( |M_{\text{nucl}}|^{2} \) is the energy-independent nuclear matrix element. \( F(Z', E) \) is the Fermi function with \( Z' = 2 \) for the helium daughter nucleus. The neutrino energy \( \epsilon_{f} \) is given by \( \epsilon_{f}(E) = E_{0} - E - V_{f} \), with the kinematic end point \( E_{0} \). \( E \) represents the kinetic energy of the electron, and \( m_{e} \) is the electron mass. Rotational, vibrational, and electronic excitations of parent and daughter molecules are taken into account by summing over all molecular final state energies \( V_{f} \) with probabilities \( \xi_{f} \) [22,25,26].

Further corrections are taken into account in the spectrum such as the Doppler broadening due to thermal motion of molecules in the WGTS and theoretical corrections [24]. Spatial and temporal variations of the source potential are incorporated as an effective energy broadening \( \sigma^{2} = (12.4 \pm 6.1) \times 10^{-4} \text{ eV}^{2} \) in the final-state distribution and as a shift in the energy loss function \( \Delta E = (0 \pm 61) \text{ meV} \) [22].

In our sterile-neutrino analysis, we extend the standard \( \beta \)-decay model by a sterile decay branch associated with a fourth neutrino mass eigenstate \( m_{\nu}^{4} \) and active-to-sterile mixing with \( |U_{e4}|^{2} \). The decay spectrum in Eq. (3) is replaced by

\[
R_{\beta}(E, m_{\nu}^{2}, m_{\nu}^{4}; |U_{e4}|^{2})
= (1 - |U_{e4}|^{2}) \cdot R_{\beta}(E, m_{\nu}^{2}) + |U_{e4}|^{2} \cdot R_{\beta}(E, m_{\nu}^{4})
\]

(4)

with the extended \( 4 \times 4 \) unitary PMNS mixing matrix \( U \).

\[ \text{B. Experimental response function} \]

The response function \( f(E, qU_{i}) \) describes the probability of transmission through the beam line \( (T(E, \theta, U)) \) and energy losses \( e \) due to inelastic scattering in the source. In this work, the same description of the response function as in Ref. [4] is used. The response function is given by

\[
f(E, qU_{i}) = \int_{E-qU_{i}}^{E} \int_{0}^{\theta_{\text{max}}} T(E-e, \theta, U_{i})
\cdot \sin \theta \cdot \sum_{x} P_{x}(\theta) f_{s}(e) \mathrm{d} \theta \mathrm{d} e.
\]

(5)

The integrated transmission probability \( T(E, \theta, U) \) is written as

\[
T(E, \theta, U) = \int_{0}^{\theta_{\text{max}}} T(E, \theta, U_{i}) \sin \theta \mathrm{d} \theta
\]

\[
= \begin{cases}
0, & \text{if } E - qU < 0 \\
1 - \sqrt{1 - \frac{(E - qU) B_{\text{source}}}{E} \frac{2}{B_{\text{ana}}} \max + 1}, & \text{if } 0 \leq E - qU \leq \Delta E \\
1 - \sqrt{1 - \frac{B_{\text{source}}}{B_{\text{ana}}} \max }, & \text{if } E - qU > \Delta E
\end{cases}
\]

(6)

with the Lorentz factor \( \gamma \). The transmission function determines the filter width of the main spectrometer, which depends on the magnetic fields in the source \( B_{\text{source}} \), the analyzing plane \( B_{\text{ana}} \), and the maximum magnetic field \( B_{\text{ana}} \). The energy loss due to inelastic scattering with tritium in the source is described by the energy loss function \( f_{s}(e) \) for \( s \)-fold scattering and the associated scattering probability \( P_{s}(\theta) \). The latter depends on the emission angle \( \theta \), which is defined as the initial polar angle of the electron momentum, relative to the magnetic field. More details can be found in Ref. [24].

\[ \text{C. Spectrum prediction} \]

The integrated spectrum rate \( R_{\text{model}}(qU_{i}) \) that is measured at the detector is predicted by

\[
R_{\text{model}}(qU_{i}) = A_{S} N_{T} \int_{qU_{i}}^{E_{0}} R_{\beta}(E) f(E, qU_{i}) \mathrm{d} E + R_{bg}(qU_{i}).
\]

(7)

Here, the differential \( \beta \)-spectrum of the tritium decay \( R_{\beta}(E) \) from Eq. (4) is convolved with the response function \( f(E, qU_{i}) \) from Eq. (5). The integrated spectrum is multiplied by two normalization factors, \( N_{T} \) and \( A_{S} \). First, \( N_{T} \) scales the spectrum to the number of tritium molecules within the flux tube quantified by experimental determination, multiplied with the detection efficiency. Second, the amplitude of the tritium signal, \( A_{S} \), is considered a free fit parameter as the prediction of the absolute decay rate is not
accurate enough. The background rate $R_{bg}(qU_i)$ comprises three components: a dominant flat background base $R_{bg}^{\text{base}}$, a hypothetical retarding-potential-dependent background $R_{bg}^{\text{ret}}(qU_i)$, and a small background contribution from electrons stored in the Penning trap between the pre- and main spectrometers $R_{bg}^{\text{Penning}}(t_i)$ [4].

$$R_{bg}(qU_i) = R_{bg}^{\text{base}} + R_{bg}^{\text{ret}}(qU_i) + R_{bg}^{\text{Penning}}(t_i(qU_i)).$$

The Penning background increases with the mean scan-step duration $t_i(qU_i)$ at a given retarding energy setting $qU_i$. The retarding potential dependence of the background is constrained to $(0.0 \pm 4.7)$ mcps/keV. The scan-step-duration dependence of the Penning background is constrained to $(3.0 \pm 3.0)$ μcps/s. These constraints were determined from independent analyses [4].

V. DATA ANALYSIS

In the following, we describe the analysis methods of our sterile-neutrino search. Section VA is dedicated to the inference of the sterile-neutrino parameters ($m_4^2$ and $|U_{e4}|^2$) from the tritium $\beta$-spectrum using the grid search technique. In this context, the applicability of Wilks’s theorem and the blinding procedure to ensure a robust and unbiased analysis are discussed. Moreover, the propagation of systematic uncertainties using the covariance matrix approach is presented. The procedure of generating the covariance matrices, as well as their application, is described in Sec. VB. Furthermore, the relative contribution of different systematics to the total uncertainty budget is given.

A. Sterile-neutrino analysis

Our sterile-neutrino analysis aims to infer two physics parameters of interest: $m_4^2$ and $|U_{e4}|^2$. Additionally, the four original fit parameters ($m_e^2$, $E_0$, $A_S$, $R_{bg}^{\text{base}}$) of the neutrino-mass analysis [4] are included as nuisance parameters. Therefore, the sterile-neutrino constraints are retrieved solely from the shape information within the experimental spectrum. To infer the sterile-neutrino parameters, we perform fits with fixed ($m_4^2$, $|U_{e4}|^2$) pairs, minimizing the standard function

$$\chi^2(\vec{x}) = (\vec{R}_{\text{data}} - \vec{R}_{\text{model}}(\vec{x}))C^{-1}(\vec{R}_{\text{data}} - \vec{R}_{\text{model}}(\vec{x}))^\top$$

with respect to all other fit parameters for each pair. Here, $\vec{x}$ denotes the parameter set ($m_e^2$, $m_4^2$, $|U_{e4}|^2$, $E_0$, $A_S$, $R_{bg}^{\text{base}}$).

The elements of the vectors $\vec{R}$ give the rates at different retarding potentials for data and model, respectively. Statistical as well as systematic uncertainties are incorporated in the covariance matrix $C$, which is described in detail in Sec. VB.

A $(50 \times 50)$ logarithmically spaced grid over the parameters $m_4^2$ and $|U_{e4}|^2$ ($m_4^2 \in [0.1, 1600]$ eV$^2$, $|U_{e4}|^2 \in [10^{-3}, 0.5]$) was used. $|U_{e4}|^2 = 0.5$ is considered as maximal mixing: If both $m_4^2$ and $m_4^2$ are unconstrained fit parameters, the active and sterile branch in the decay spectrum [Eq. (4)] are interchangeable: $(m_3^2, |U_{e4}|^2) \leftrightarrow (m_4^2, 1 - |U_{e4}|^2)$. Consequently, no additional information can be gained by extending the active-to-sterile mixing beyond 0.5. Different grid layouts were studied to ensure full convergence of the employed grid. We draw the 95% C.L. exclusion contour at

$$\Delta\chi^2 = \chi^2 - \chi^2_{\text{min}} = 5.99$$

following Wilks’s theorem [27] for 2 degrees of freedom.

We verified the applicability of Wilks’s theorem by constructing the $\Delta\chi^2$ probability distribution functions numerically with 1500 randomized Monte Carlo simulations for the null hypothesis and several different sterile-neutrino hypotheses. The former is, by definition, the boundary physics case. As the null hypothesis can be realized with an infinite number of sterile parameter pairs with $|U_{e4}| = 0$ and arbitrary $m_4^2$, a hypothetical deviation from Wilks’s theorem is anticipated to be most prominent here. Short-baseline oscillation experiments observe a deviation from Wilks’s theorem because statistical fluctuations can likely mimic a fake sterile-neutrino oscillation signal [28]. In contrast to that, the sterile-neutrino signal in KATRIN manifests itself as a global spectral distortion. Therefore, Wilks’s theorem is expected to apply for KATRIN. Indeed, we did not observe deviation from the predictions of Wilks’s theorem for any of the Monte Carlo truths.

To mitigate human-induced biases, the full analysis chain is first applied to a simulated dataset without statistical fluctuations. For each experimental scan, we generate a twin spectrum based on the true experimental parameters assuming no sterile neutrino and vanishing neutrino mass. Only after three independent analysis teams, using different analysis codes, obtained consistent sensitivity estimates, the actual data analysis was performed without any subsequent modifications.

B. Systematic uncertainties

Since KATRIN is a high-precision experiment, the accurate description of all systematic effects is crucial. Various systematic effects are taken into account, arising at different points along the electron’s trajectory from the source to the focal plane detector. We consider the same systematic effects as in the neutrino mass analysis [4]. The uncertainty propagation to the integrated spectrum is conducted with the covariance-matrix approach [22]. Each independent systematic effect can be assessed by a separate covariance matrix $C_j$. The latter is estimated by simulating
O$(10^3)$ tritium spectra, varying the relevant set of parameters associated to a particular systematic effect $j$ according to their joint probability distribution function. Since we perform a shape-only analysis, absolute rate uncertainties are eliminated from the covariance matrices by normalizing each sample spectrum to the average statistics. The sum of all covariance matrices $C_j$ and a diagonal matrix describing statistical uncertainties $C_{\text{stat}}$ comprises the total variances and covariances of the data points:

$$C = \sum_j C_j + C_{\text{stat}}.$$  \hfill (11)

Diagonal elements of the covariance matrix give the bin-to-bin uncorrelated uncertainties, whereas off-diagonal elements are $qU$-dependent correlations in the integrated spectrum. The impact on the model spectrum is then taken into account by applying the covariance matrix to the $\chi^2$-function [Eq. (9)] in the final fit. Since the covariance matrices are precalculated, it is an efficient approach to include many systematic effects simultaneously without demanding more computing time in the fit itself.

The analysis of systematic uncertainties is performed on the simulated twin dataset with fixed $m_\nu^2 = 0$ eV$^2$. For each systematic effect, a grid scan considering only the individual systematic uncertainty on top of the statistical uncertainty was carried out. For all systematic effects, the change of the contour is small compared to the sensitivity evaluated with only statistical uncertainties. To assess the relative contribution of each systematic effect to the total uncertainty in a more quantitative way, we perform a raster scan for each effect. For fixed values of $m_4^2$, we calculate the 95% C.L. sensitivity on the mixing, $\sigma(|U_{e4}|^2)$, as a function of $m_4^2$. In this case, we have only 1 degree of freedom; therefore, the critical $\chi^2$ is reduced to 3.84 at 68.3% C.L. This method allows us to assess the systematic-only contribution on $|U_{e4}|^2$ using

$$\sigma_{\text{syst}} = \sqrt{\sigma_{\text{total}}^2 - \sigma_{\text{stat}}^2}.$$  \hfill (12)

This quantity is displayed in the left panel of Fig. 3 for each systematic effect as a function of $m_4^2$. Moreover, the relative contribution of each systematic variance to the total variance $\sigma_{\text{syst}}^2/\sigma_{\text{total}}^2$ is shown in the right-hand panel.

In Table I, the median contribution of each systematic effect is summarized, sorted by magnitude. The median is
TABLE I. Breakdown of the relative uncertainties on $|U_{e4}|^2$, given as the median $(\sigma^2/\sigma^2_{\text{total}})_{\text{median}}$ over all $m^2_4$. The systematic effects are listed in ascending order of the maximal uncertainty $\max(\sigma^2_{\text{stat}}/\sigma^2_{\text{total}})$. The systematic uncertainty inputs are those used in the neutrino-mass analysis published in Ref. [4]. The analysis is dominated by statistical uncertainties $[(\sigma^2_{\text{stat}}/\sigma^2_{\text{total}})_{\text{median}} > 0.5]$.

| Effect                                      | $(\sigma^2/\sigma^2_{\text{total}})_{\text{median}}$ |
|--------------------------------------------|--------------------------------------------------|
| Statistical                                 | 0.86                                             |
| Source-potential variations                 | 0.06                                             |
| Scan-step-duration-dependent background     | 0.04                                             |
| Non-poison background                        | 0.02                                             |
| Magnetic fields                             | 0.03                                             |
| Molecular final-state distribution          | 0.05                                             |
| $qU$-dependent background                   | 0.01                                             |
| Column density $\times$ inelastic scattering cross section | 0.01  |
| Detector efficiency                         | 0.01                                             |
| Activity fluctuations                       | <0.01                                           |
| Energy-loss function                        | <0.01                                           |
| High-voltage stability and reproducibility  | <0.01                                           |
| Theoretical corrections                     | <0.01                                           |
| **Total systematic uncertainty**            | 0.14                                             |

calculated using the same linearly spaced $m^2_4$ values for all systematic effects.

The analysis of the second KATRIN campaign is statistics dominated in the analyzed energy range, with $\sigma^2_{\text{stat}}/\sigma^2_{\text{total}} > 0.5$ for all $m^2_4$. For $m^2_4 \leq 600 \text{ eV}^2$, the systematic effects are dominated by the non-Poisson rate distribution of the background [29], the scan-step-duration-dependent background, and source-potential variations. For larger $m^2_4 > 600 \text{ eV}^2$, all systematic contributions, except for the non-Poisson rate distribution of the background, rapidly increase. The largest systematic contribution for larger $m^2_4$ is given by the molecular final-state distribution. This can be explained by an increased uncertainty on the excited molecular states in this energy region.

VI. RESULTS

Here, we report on the results of our light sterile neutrino search. Before presenting the improved exclusion bounds, we address the correlation between active and sterile neutrino branches in the KNM2 tritium $\beta$-decay spectrum (Sec. VI A). As the former has a strong influence on the sterile neutrino exclusion bounds, we then distinguish between two analysis cases: in the case I analysis (Sec. VI B), we consider $m^2_4 = 0 \text{ eV}^2$, whereas the case II analysis (Sec. VI C) employs $m^2_4$ as a nuisance parameter. In Sec. VI D, the combined analysis of KNM1 and KNM2 is presented. Lastly, the $m^2_4$ sensitivity in the presence of a sterile neutrino is evaluated (Sec. VI E).

A. Correlation between active and sterile neutrino branches

The model spectrum in Eq. (4) consists of the weighted sum of two branches: the active branch with effective electron antineutrino mass $m_\alpha$ and the sterile branch with the fourth mass eigenstate $m_4$. The branches are weighted according to their mixing: $(1 - |U_{e4}|^2)$ for the active branch and $|U_{e4}|^2$ for the sterile branch, respectively. Apart from the different neutrino masses and weights, the two branches are mathematically identical. Since $m^2_4$ is small in the observed data and simulation, the two branches are degenerate in the case of small $m^2_4$ and large mixing $|U_{e4}|^2 \approx 0.5$.

To quantify this relation more generally, we simulate $R_\beta$ for several values of $m^2_4$, $|U_{e4}|^2$, and $m^2_v = 0 \text{ eV}^2$. Then, we perform five fits to each simulated spectrum by varying $m^2_4$ stepwise by $±1 \text{ eV}^2$ around the respective Monte Carlo truths. The fits are optimized with respect to all nuisance parameters, keeping $|U_{e4}|^2$ fixed to its simulated value. For each $(m^2_4, |U_{e4}|^2)$-pair, we determine the approximately linear relationship $m^2_4 = \alpha_{\text{slope}} \cdot m^2_\alpha + \text{const}$ in the vicinity of the MC truth, which serves as a proxy for the correlation between the two masses. Figure 4 shows $\alpha_{\text{slope}}$ in the $(m^2_4, |U_{e4}|^2)$-parameter space. The smaller the active-to-sterile mixing is, the smaller the contribution of the sterile branch to the simulated spectrum is. We find small $|\alpha_{\text{slope}}| < 0.01$ for small $|U_{e4}|^2 < 0.01$. For small $m^2_4 \lesssim 30 \text{ eV}^2$, we observe negative slope values. For large $30 \text{ eV}^2 \lesssim m^2_4 \lesssim 1000 \text{ eV}^2$, the absolute magnitude of the correlation increases linearly with $m^2_4$.

![Graph showing the correlation between active and sterile neutrino mass](072004-8)
slope is reduced and has a positive sign. Due to the correlation between $m_2^2$ and $m_3^2$, the exclusion curves vary significantly for different treatments of $m_3^2$ as discussed in the following sections.

**B. Neutrino mass fixed to 0 eV²**

In our main analysis, denoted case I, we consider the hierarchical scenario $m_{1,2,3} \ll m_4$. This justifies setting $m_4$ to zero, which is consistent with the lower limit derived from neutrino oscillations (0.009 eV [3]) within our sensitivity. We perform a two-dimensional grid search over $(m_3^2, |U_{e4}|^2)$, minimizing the $\chi^2$-function with respect to three free fit parameters ($E_0, A_S, R_{bg}^{\text{base}}$) at each grid point. As $m_3^2 = 0$ eV² is fixed, we extend our grid to $|U_{e4}|^2 = 1.0$. The exclusion curve at 95% C.L. is shown in Fig. 5 (blue line). The global minimum of the $\chi^2$-function $\chi_{\text{min}}^2 = 27.5$ (23 dof, $p = 0.24$) is found at $m_3^2 = 0.28$ eV² and $|U_{e4}|^2 = 1.0$.

The extreme active-to-sterile mixing shows that the observed decay spectrum [Eq. (3)] can be described best by only one branch with one associated free neutrino mass. As the active and sterile branches only differ in their neutrino mass and mixing nomenclature, the two branches are indistinguishable in the scenario at hand. However, since active-to-sterile mixing values of $|U_{e4}|^2 > 0.5$ are excluded by oscillation experiments [3], we interpret our result as signature from the active branch with free $m_3^2$. Indeed, the best-fit value for $m_3^2$ coincides with the best-fit value of $m_3^2 = 0.28$ eV² that was found in our neutrino-mass analysis with the same FPD pixel combination strategy [4]. The significance of the best fit with respect to the null hypothesis is $\Delta \chi^2 = \chi^2_{\text{null}} - \chi^2_{\text{min}} = 0.7$; i.e., the result is not statistically significant ($\Delta \chi^2 < 5.99$). Moreover, we perform a supplementary analysis, extending our modeled sterile-neutrino signal to the nonphysical parameter space: $m_3^2 \in [-40 \text{ eV}^2, +40 \text{ eV}^2]$, $|U_{e4}|^2 \in [-0.5, 1]$. We find a global best fit for negative mixing and positive sterile-neutrino mass squared, which would formally correspond to a negative decay rate. While still not being significant at 95% C.L., this best fit improves the $\chi^2_{\text{min}}$ with respect to the minimum found in the physical region by one unit.

**C. Analysis with free neutrino mass**

Furthermore, we study a more generic scenario, labeled case II, in which we include $m_3^2$ as an additional unconstrained nuisance parameter in our analysis. Due to the correlation between active and sterile neutrino mass (see Fig. 4), we expect a significantly different result than in analysis case I. The associated exclusion contour (solid red) and the isolines of the $m_3^2$ fit values (dotted light red) are displayed in Fig. 5. At each $(m_3^2, |U_{e4}|^2)$ pair in our grid, the fit has gained the freedom to converge to a value of $m_3^2$ that improves the goodness of fit:

\[
\chi^2_{\text{min}}(m_3^2, m_4^2, |U_{e4}|^2) \leq \chi^2_{\text{min}}(m_3^2 = 0 \text{ eV}^2, m_4^2, |U_{e4}|^2).
\]

For small $m_3^2 \leq 0.4 \text{ eV}^2$ and large $|U_{e4}|^2 \geq 0.04$, we find $m_3^2 \leq 0 \text{ eV}^2$ of the same order of magnitude as $m_4^2$, reflecting the expected strong negative correlation between $m_3^2$ and $m_4^2$. For small mixing, $|U_{e4}|^2 < 6 \times 10^{-3}$, the $m_3^2$ fit values lie within the 1σ confidence region of the neutrino-mass analysis [4]. We report a best fit at $m_3^2 = 97.8 \text{ eV}^2$, $|U_{e4}|^2 = 0.027$, and $m_4^2 = 1.1 \text{ eV}^2$ with $\chi^2 = 25.0$ (dof = 22, $p = 0.30$). The best fit improves with respect to the null hypothesis by $\Delta \chi^2 = 2.5$, thus not reaching the $\Delta \chi^2$ threshold at 95% C.L. for a significant result.

**D. Combined analysis**

We combine the 3ν + 1 sterile-neutrino constraints from the first four-week science run (KNM1). Since the first stand-alone analysis presented in Ref. [15], the spectrum calculation has been slightly refined. It now incorporates the nonisotropic transmission of electrons, a possible time-dependency of Penning trap induced background, an improved parametrization of the energy-loss function [30], and reduced systematic uncertainties on the magnetic fields and the column density. The reanalysis of the first science run with these new inputs yields exclusion contours consistent with those of the original publication.

The combined exclusion contours are obtained by minimizing the $\chi^2$-function.
χ²(m2, m2, |Ue|², nKNM1, nKNM2) = χ²(m2, m2, |Ue|², nKNM1) + χ²(m2, m2, |Ue|², nKNM2) (14)

at each (m2, |Ue|²) pair. Due to different experimental conditions, several nuisance parameters are expected to vary between the datasets. As KNM2 was operated at a lower background level and a higher source activity compared to KNM1, we allow for a campaignwise background and signal normalization. To account for an unknown difference in the absolute source potential, both datasets are described with an individual effective end point. The campaignwise fit parameters (E0, A, Rbg) are summarized in Eq. (14) by nKNM1 and nKNM2 for the first and second measurement campaigns. As both datasets are strongly statistics dominated, possible correlations among systematic uncertainties are negligible.

In case I, KNM1 and KNM2 do not share any common nuisance parameter, because m2 is fixed. Therefore, the χ²-functions of the individual analyses in Eq. (14) can be minimized independently from each other. The combined and individual exclusion contours are shown in Fig. 6 in bluish colors. The corresponding best-fit parameters are stated in Table II. We find a best fit of the combined analysis at (m2 = 59.9 eV², |Ue|² = 0.011) with χ² min = 50.4 (dof = 47, p = 0.34), improving with respect to the null hypothesis by Δ χ² = 0.7. To evaluate the compatibility between the two statistically independent datasets, we perform the parameter-goodness-of-fit (PGoF) test [31]. This test quantifies the penalty of combining KNM1 and KNM2 in units of χ² compared to the stand-alone analyses

χ² penalty = χ² min, combi – (χ² min,KNM1 + χ² min,KNM2). (15)

As the datasets share two fit parameters (m2 and |Ue|²), the χ²-penalty can be converted into a p-value using 2 degrees of freedom. We report on a probability of p = 47%, demonstrating a good agreement. Due to statistical fluctuations in both datasets, the combined exclusion improves for m2 ≤ 50 eV² compared to the KNM2 stand-alone result while providing slightly weaker constraints for m2 > 50 eV². The KNM1 exclusion bounds are improved by the combined analysis for the entire mass range. The observed exclusion contour agrees well with our sensitivity estimate, lying within the 1σ band of 95% C.L. sensitivity contours that are obtained from the simulation of 1500 randomized pseudoexperiments.

In case II, KNM1 and KNM2 share m2 as a common nuisance parameter. Therefore, a simultaneous grid search on both datasets minimizing the combined χ²-function in Eq. (14) has to be performed. The exclusion contours of the stand-alone and combined analyses are displayed in Fig. 6 with reddish coloring. The relevant parameters of the best fits are given in Table II. The best fit of the combined exclusion at χ² = 49.9 (46 dof, p = 0.34) improves with respect to the null hypothesis by Δ χ² = 1.7, rendering it not significant at 95% C.L. The PGoF of 20% indicates good compatibility between the KNM1 and KNM2 sterile-neutrino analysis with free m2. We find the best-fit value m2 = 0.57 eV², which agrees within 1σ with the standard neutrino-mass analysis [4].

E. Neutrino-mass sensitivity

As described in Sec. VI.A, we observe a sizable correlation between the effective electron antineutrino mass and the fourth mass eigenstate. This relation results in weaker constraints on the active-to-sterile neutrino mixing when m2 is included as a free fit parameter in the sterile-neutrino search (see Sec. VI.C). Turning the analysis concept upside down, the 3ν + 1 model extension is expected to cause a reduction in m2 sensitivity [32].

To assess the latter, we calculate the χ²-profile as a function of m2, displayed in Fig. 7 for both data and twin analysis. For different fixed m2 ∈ [−1, 2.5] eV², a two-dimensional grid search over the (|Ue|², m2) parameter space is performed, minimizing the χ²-function with respect to all other nuisance parameters. The value χ²(m2) in the χ²-profile corresponds to the global minimum found in the grid search with the respective fixed m2. For m2 ≤ 0 eV², the global χ² minima are located at m2 < 2 eV² and large mixing |Ue|² ≈ 0.5. As m2 and

![Figure 6](image-url)

FIG. 6. Exclusion contours of the KNM1 and KNM2 stand-alone analyses and the combined analysis at 95% C.L. The contour lines shaded in blue consider m2 = 0 eV², whereas the exclusions shaded in red include m2 as an unconstrained fit parameter. Due to statistical fluctuations, some contours exhibit bumpy features in the region 30 eV² ≤ m2 ≤ 100 eV². The location of the (m2, |Ue|²) best-fit values, summarized in Table II, generally cause slightly weaker constraints in their vicinity compared to the sensitivity. For the same reason, the combined exclusion contour of KNM1+2 gives a weaker constraint in certain regions compared to KNM2 stand-alone exclusion contour.

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TABLE II. Results of the KNM1 and KNM2 (Secs. VI B and VI C) stand-alone and combined analyses (Sec. VI D). The first three rows correspond to analysis case I with \( m_2 = 0 \) eV\(^2\), whereas the last three rows show the results of analysis case II with unconstrained \( m_2 \). The first five columns show the best-fit parameter values (\( m_1^2, |U_{e1}|^2, m_2^2 \)) and the associated goodness of fits (\( \chi^2_{\text{min}}/\text{dof} \)) and confidence level. All observed sterile-neutrino signals are compatible with the no-sterile neutrino hypothesis; i.e., no significant spectral distortions at 95% C.L. are found. The last column gives the parameter goodness of fit \( \hat{p} \) for the combined analyses.

| Analysis case | Dataset | \( m_1^2 \) (eV\(^2\)) | \( |U_{e1}|^2 \) | \( m_2^2 \) (eV\(^2\)) | \( \chi^2_{\text{min}}/\text{dof} \) | \( p \) | \( \Delta \chi^2_{\text{null}} \) | Significance | \( \hat{p} \) |
|-------------|---------|------------------|-------------|------------------|------------------|-------------|------------------|-------------|-------------|
| I           | KNM1    | 77.5 eV\(^2\)    | 0.031       | Fixed            | 21.4/22          | 0.50         | 1.43             | 51.0%       | ...         |
|             | KNM2    | 0.28 eV\(^2\)    | 1.0         | Fixed            | 27.5/23          | 0.24         | 0.74             | 31.0%       | ...         |
|             | KNM1 + 2| 59.9 eV\(^2\)    | 0.011       | Fixed            | 50.4/47          | 0.34         | 0.66             | 28.1%       | 0.47        |
| II          | KNM1    | 21.8 eV\(^2\)    | 0.155       | \(-5.3 \text{ eV}^2\) | 19.9/21          | 0.53         | 1.30             | 47.9%       | ...         |
|             | KNM2    | 98.3 eV\(^2\)    | 0.027       | 1.1 eV\(^2\)    | 25.0/22          | 0.30         | 2.49             | 71.2%       | ...         |
|             | KNM1 + 2| 87.4 eV\(^2\)    | 0.019       | 0.57 eV\(^2\)   | 49.5/46          | 0.34         | 1.69             | 57.1%       | 0.20        |

\( m_1^2 \) are strongly correlated in this region (see Fig. 4), the \( \chi^2 \)-profiles are flat. Assuming the existence of a sterile neutrino, this corresponds to a complete loss of sensitivity. The latter can be restored by using external constraints on \( m_1^2 \) or \( |U_{e1}|^2 \). Restricting \( m_1^2 > 20 \text{ eV}^2 \) or \( |U_{e1}|^2 < 0.04 \), we find lower and upper 1σ sensitivities on \( m_1^2 \) of equal size. For \( m_1^2 > 0 \text{ eV}^2 \), the best fits are located at 10 \text{ eV}^2 < \( m_1^2 < 200 \text{ eV}^2 \) and moderate mixings \( O(10^{-2}) \). In this part of the parameter space, the correlation between the two masses is less pronounced. As a result, the 1σ uncertainty on \( m_0^2 \) in the \( 3\nu + 1 \) extension is only increased by a factor of 2 compared to the standard neutrino-mass analysis.

The neutrino-mass sensitivity in the \( 3\nu + 1 \) framework can be fully restored by limiting the active-to-sterile mixing to small values. For \( |U_{e1}|^2 < 10^{-4} \), the sensitivity on \( m_0^2 \) converges to the nominal one in the \( 3\nu \) framework. Using the same constraint, we can also reproduce the central value and uncertainties of our standard neutrino-mass analysis [4].

VII. COMPARISON TO OTHER EXPERIMENTS

To put this work into context, we compare our case I exclusion contours with constraints from a selection of other experiments displayed in Fig. 8, focusing on sterile neutrino searches in the electron disappearance channel. This result improves on the constraints from the completed Mainz and Troitsk experiments for \( m_2^2 \lesssim 300 \text{ eV}^2 \). As short-baseline neutrino oscillation experiments are sensitive to different observables than \( \beta \)-decay experiments, we perform the associated variable transformations to relate the results. While KATRIN is directly sensitive to \( |U_{e4}|^2 \), sterile neutrino oscillations are characterized by \( \sin^2(2\theta_{\text{ee}}) = 4|U_{e4}|^2(1 - |U_{e1}|^2) \). Moreover, the mass splitting can be written as \( \Delta m_{41}^2 \approx m_3^2 - m_2^2 \), which is valid within \( 2 \times 10^{-4} \text{ eV}^2 \) [33]. For our analysis case I, this approximation is equivalent to \( \Delta m_{41}^2 \approx m_0^2 \). We are able to exclude the large \( \Delta m_{41}^2 \) solutions of the combined gallium experiments for \( 20 \text{ eV}^2 \lesssim \Delta m_{41}^2 \lesssim 1000 \text{ eV}^2 \). Moreover, a considerable fraction of the reactor antineutrino anomaly for \( 50 \text{ eV}^2 \lesssim \Delta m_{41}^2 \lesssim 1000 \text{ eV}^2 \) is challenged by our results. Our combined analysis of the first and second science runs disfavors the Neutrino-4 hint of a signal for \( \sin^2(2\theta_{\text{ee}}) \gtrsim 0.4 \) at 95% C.L.

Furthermore, we compare our results to constraints from 0νββ experiments. If neutrinos are Majorana particles and 0νββ is triggered by light Majorana neutrino exchange, \( m_4 \) will contribute to the effective Majorana mass

\[
m_{\beta\beta} = \sum_{i=1}^{4} U_{ei}^2 m_i = (1 - |U_{e4}|^2) \sum_{i=1}^{4} U_{ei}^2 m_i + |U_{e4}|^2 e^{i\gamma} m_4,
\]

with a possible third Majorana phase \( \gamma \in [-\pi, \pi] \). Limiting the active neutrino contribution to \( m_{\beta\beta} \) to the nondegenerate regime, i.e., between 0 and 0.005 eV (0.01 and 0.05 eV) for the normal (inverted) ordering, allows us to translate the
current upper limit of $m_{\beta\beta}$ to constraints on sterile neutrinos. The conversion from the observable half-life to $m_{\beta\beta}$ depends on the nuclear matrix elements. As an illustration, we select the calculation with the nuclear matrix elements that result in the least stringent limit $m_{\beta\beta} < 0.16$ eV [41,42]. The width of the two gray exclusion bands in Fig. 8 reflects the uncertainties on the entries of the PMNS matrix and the unknown Majorana phases [3].

The projected final sensitivity quoted here after Ref. [15] demonstrates that KATRIN constraints will improve the global sensitivity for $\Delta m_{41}^2 \gtrsim 5$ eV$^2$ and will provide complementary results to short-baseline oscillation experiments for smaller masses.

VIII. CONCLUSION

We present the light sterile-neutrino search from the second KATRIN measurement campaign in 2019. Our dataset comprises $3.76 \times 10^6$ signal $\beta$-electrons inside the region of interest, reaching an energy-dependent signal-to-background ratio of up to 235. The analysis is sensitive to the fourth neutrino mass eigenstate $m_4^2 \lesssim 1600$ eV$^2$ and active-to-sterile mixing $|U_{e4}|^2 \gtrsim 6 \times 10^{-3}$ in the $3\nu + 1$ framework. As no significant sterile-neutrino signal is observed, we report on improved exclusion limits with respect to our first measurement campaign. Our results improve on the constraints by previous tritium $\beta$-decay experiments. Moreover, we are able to exclude the large $\Delta m_{41}^2$ solutions of the reactor and gallium anomalies. Combining the datasets from the first and second KATRIN measurement campaigns, our result disfavors the Neutrino-4 signal for $\sin^2(2\theta_{ee}) \gtrsim 0.4$.

The impact of systematic effects on our sterile-neutrino search was studied in detail. We conclude that our analysis is dominated by statistical uncertainties for all $m_4^2$ with a median relative contribution of $\left(\sigma_{\text{stat}}^2/\sigma_{\text{total}}^2\right)_{\text{median}} = 86\%$ with respect to the total uncertainty budget.

Furthermore, we investigated the correlation between active and sterile neutrino mass. We find a negative correlation for $m_4 \lesssim 30$ eV$^2$ with increasing absolute strength for increasing mixing. For larger sterile masses, the correlation is less pronounced and has a positive sign. Assuming the existence of a light sterile neutrino, this correlation translates into a reduction in neutrino-mass sensitivity by a factor of 2 compared to the neutrino-mass analysis in the $3\nu$ framework. By constraining the sterile
neutrino mass or mixing, the nominal sensitivity can be restored.

With hundreds of scheduled measurement days ahead, KATRIN will further improve its statistics by a factor of 50. In combination with a further reduction in background level and systematic uncertainties, this will allow us to cover an even larger fraction of the gallium and reactor antineutrino anomaly regions and the entire Neutrino-4 signal.

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[1] J. Angrik, T. Armbrust, A. Beglarian, U. Besserer, J. Blümer, J. Bonn et al. (KATRIN Collaboration), KATRIN design report, FZKA scientific report, 1 (2005), https://publikationen.bibliothek.kit.edu/27006/0149.

[2] M. Aker, K. Altenmüller, J. Amsbaugh, M. Arenz, M. Babutzka, J. Bast et al. (KATRIN Collaboration), The design, construction, and commissioning of the KATRIN experiment, J. Instrum. 16, T08015 (2021).

[3] P. Zyla, R. Barnett, J. Beringer, R. Bonventre, O. Dahl, D. Dwyer et al. (Particle Data Group), Review of particle physics, Prog. Theor. Phys. 2020, 083C01 (2020).

[4] M. Aker, A. Beglarian, J. Behrens, A. Berlev, U. Besserer, B. Bieringer et al. (KATRIN Collaboration), First direct neutrino-mass measurement with sub-eV sensitivity, Nat. Phys. 18, 160 (2022).

[5] M. Aker, K. Altenmüller, M. Arenz, M. Babutzka, J. Barrett, S. Bauer et al. (KATRIN Collaboration), Improved Upper Limit on the Neutrino Mass from a Direct Kinematic Method by KATRIN, Phys. Rev. Lett. 123, 221802 (2019).

[6] S. Schael, R. Barate, R. Brunelière, D. Buskulic, I. De Bonis, D. Decamp et al. (ALEPH, DELPHI, L3, OPAL, SLD, LEP Electroweak Working Group, SLD Electroweak Group, SLD Heavy Flavour Group Collaborations), Precision electroweak measurements on the Z resonance, Phys. Rep. 427, 257 (2006).

[7] S. Mertens, A. Alborini, K. Altenmüller, T. Bode, L. Bombelli, and T. Brunst, A novel detector system for KATRIN to search for keV-scale sterile neutrinos, J. Phys. G 46, 065203 (2019).

[8] A. Aguilar, L. Auerbach, R. Burman, D. Caldwell, E. Church, A. Cochran et al. (LSND Collaboration), Evidence for neutrino oscillations from the observation of appearance in a beam, Phys. Rev. D 64, 112007 (2001).

[9] A. Aguilar-Arevalo, B. Brown, L. Bugel, G. Cheng, J. Conrad, R. Cooper et al., Significant Excess of Electronlike Events in the MiniBooNE Short-Baseline Neutrino Experiment, Phys. Rev. Lett. 121, 221801 (2018).

[10] W. Hampel, G. Heusser, J. Kiko, T. Kirsten, M. Laubenstein, E. Pernicka et al. (GALLEX Collaboration), Final results of the 51Cr neutrino source experiments in GALLEX, Phys. Lett. B 420, 114 (1998).

[11] J. Abdurashitov, V. Gavrin, V. Gorbachev, P. Gurkina, T. Ibragimova, A. Kalikhov et al. (SAGE Collaboration), Measurement of the solar neutrino capture rate with gallium metal. III. Results for the 2002–2007 data-taking period, Phys. Rev. C 80, 015807 (2009).

[12] V. V. Barinov, B. T. Cleveland, S. N. Danshin, H. Ejiri, S. R. Elliott, D. Frekers et al. (BEST Collaboration), Results from the Baksan Experiment on Sterile Transitions, arXiv:2109.11482 [Phys. Rev. Lett. (to be published)].

[13] G. Mention, M. Fechner, T. Lasserre, T. Mueller, D. Lhuillier, M. Cribier, and A. Letourneau, The reactor antineutrino anomaly, Phys. Rev. D 83, 073006 (2011).

[14] A. P. Serebrov, V. G. Ivochkin, R. M. Samoilov, A. K. Fomin, A. O. Polyushkin, V. G. Zinoviev et al. (NEUTRINO-4 Collaboration), First observation of the oscillation effect in the neutrino-4 experiment on the search for the sterile neutrino, Pis'ma Zh. Eksp. Teor. Fiz. 109, 209 (2019).

[15] M. Aker, K. Altenmüller, A. Beglarian, J. Behrens, A. Berlev, U. Besserer et al. (KATRIN Collaboration), Bound on 3 + 1 Active-Sterile Neutrino Mixing from the First
Four-Week Science Run of KATRIN, Phys. Rev. Lett. 126, 091803 (2021).

[16] S. Grohmann, T. Bode, M. Hötzel, H. Schön, M. Süßer, and T. Wahl, The thermal behaviour of the tritium source in KATRIN, Cryogenics 55–56, 5 (2013).

[17] A. Marsteller, B. Bornschein, L. Bornschein, G. Drexin, F. Friedel, R. Gehring et al., Neutral tritium gas reduction in the KATRIN differential pumping sections, Vacuum 184, 109979 (2021).

[18] M. Arenz, W.-J. Baek, M. Beck, A. Beglarian, J. Behrens, T. Bergmann et al. (KATRIN Collaboration), The KATRIN superconducting magnets: Overview and first performance results, J. Instrum. 13, T08005 (2018).

[19] V.M. Lobashev and P.E. Spivak, A method for measuring the anti-electron-neutrino rest mass, Nucl. Instrum. Methods Phys. Res., Sect. A 240, 305 (1985).

[20] A. Picard, H. Backe, H. Barth, J. Bonn, B. Degen, T. Edling et al., A solenoid retarding spectrometer with high resolution and transmission for keV electrons, Nucl. Instrum. Methods Phys. Res., Sect. B 63, 345 (1992).

[21] J. Amsbaugh, J. Barrett, A. Beglarian, T. Bergmann, H. Bichsel, L. Bodine et al., Focal-plane detector system for the KATRIN experiment, Nucl. Instrum. Methods Phys. Res., Sect. A 778, 10 (2015).

[22] M. Aker, K. Altenmüller, A. Beglarian, J. Behrens, A. Berlev, U. Besserer et al. (KATRIN Collaboration), Analysis methods for the first KATRIN neutrino-mass measurement, Phys. Rev. D 104, 012005 (2021).

[23] M. Aker, K. Altenmüller, A. Beglarian, J. Behrens, A. Berlev et al. (KATRIN Collaboration), Quantitative long-term monitoring of the circulating gases in the KATRIN experiment using ramam spectroscopy, Sensors 20, 4827 (2020).

[24] M. Klesieck, J. Behrens, G. Drexin, K. Eitel, M. Erhard, J. A. Formaggio et al., β-decay spectrum, response function and statistical model for neutrino mass measurements with the KATRIN experiment, Eur. Phys. J. C 79, 204 (2019).

[25] A. Saenz, S. Jonsell, and P. Froelich, Improved Molecular Final-State Distribution of HeT+ for the β-Decay Process of T2, Phys. Rev. Lett. 84, 242 (2000).

[26] N. Doss, J. Tennyson, A. Saenz, and S. Jonsell, Molecular effects in investigations of tritium molecule β decay end-point experiments, Phys. Rev. C 73, 025502 (2006).

[27] S. S. Wilks, The large-sample distribution of the likelihood ratio for testing composite hypotheses, Ann. Math. Stat. 9, 60 (1938).

[28] M. Agostini and B. Neumair, Statistical methods applied to the search of sterile neutrinos, Eur. Phys. J. C 80, 750 (2020).

[29] F. M. Fränkle, A. Schaller, C. Weinheimer, G. Drexin, S. Mertens, K. Blaum et al., KATRIN background due to surface radioimpurities, Astropart. Phys. 138, 102686 (2022).

[30] M. Aker, A. Beglarian, J. Behrens, A. Berlev, U. Besserer, B. Bieringer et al. (KATRIN Collaboration), Precision measurement of the electron energy-loss function in tritium and deuterium gas for the KATRIN experiment, Eur. Phys. J. C 81, 579 (2021).

[31] M. Maltoni and T. Schwetz, Testing the statistical compatibility of independent data sets, Phys. Rev. D 68, 033020 (2003).

[32] A. S. Riis and S. Hannestad, Detecting sterile neutrinos with KATRIN like experiments, J. Cosmol. Astropart. Phys. 02 (2011) 011.

[33] C. Giunti, Y. Li, and Y. Zhang, KATRIN bound on 3 + 1 active-sterile neutrino mixing and the reactor antineutrino anomaly, J. High Energy Phys. 05 (2020) 061.

[34] C. Kraus, A. Singer, K. Valerius, and C. Weinheimer, Limit on sterile neutrino contribution from the Mainz Neutrino Mass Experiment, Eur. Phys. J. C 73, 2323 (2013).

[35] A. I. Belesev, A. I. Berlev, E. V. Geraskin, A. A. Golubev, N. A. Likhovid, A. A. Nozik, V. S. Pantuev, V. I. Parfenov, and A. K. Skasyrskaya, An upper limit on additional neutrino mass eigenstate in 2 to 100 eV region from ‘Troitsk nu-mass’ data, JETP Lett. 97, 67 (2013).

[36] M. Andriamirado, A. Balantekin, H. Band, C. Bass, D. Bergeron, D. Berish et al. (PROSPECT Collaboration), Improved short-baseline neutrino oscillation search and energy spectrum measurement with the PROSPECT experiment at HFIR, Phys. Rev. D 103, 032001 (2021).

[37] M. Danilov (DANSS Collaboration), Recent results of the DANSS experiment, Proc. Sci. EPS-HEP2019 (2020) 401 [arXiv:1911.10140].

[38] P. Adamson, F. An, I. Anghel, A. Aurisano, A. Balantekin, H. Band et al. (MINOS+, Daya Bay Collaboration), Improved Constraints on Sterile Neutrino Mixing from Disappearance Searches in the MINOS, MINOS+, Daya Bay, and Bugey-3 Experiments, Phys. Rev. Lett. 125, 071801 (2020).

[39] T. Abrahão, H. Almazan, J. C. dos Anjos, S. Appel, J. C. Barriere, I. Bekman et al. (Double Chooz Collaboration), Search for signatures of sterile neutrinos with Double Chooz, Eur. Phys. J. C 81, 775 (2021).

[40] H. Almazan, L. Bernard, A. Blanchet, A. Bonhomme, C. Buck, P. delAmoSanchez et al. (STEREO Collaboration), Improved sterile neutrino constraints from the STEREO experiment with 179 days of reactor-on data, Phys. Rev. D 102, 052002 (2020).

[41] A. Gando, Y. Gando, T. Hachiya, A. Hayashi, S. Hayashida, H. Ikeda et al. (KamLAND-Zen Collaboration), Search for Majorana Neutrinos near the Inverted Mass Hierarchy Region with KamLAND-Zen, Phys. Rev. Lett. 117, 082503 (2016).

[42] M. Agostini, G. R. Araujo, A. M. Bakalyarov, M. Balata, I. Barabanov, L. Baudis et al. (GERDA Collaboration), Final Results of GERDA on the Search for Neutrinoless Double-β Decay, Phys. Rev. Lett. 125, 252502 (2020).