Lepton flavor violating decays of vector quarkonia and of the Z boson

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Abstract

We address the impact of sterile fermions on the lepton flavor violating decays of quarkonia as well as of the Z boson. We compute the relevant Wilson coefficients and show that the $B(V \to \ell_\alpha \ell_\beta)$, where $V = \phi, \psi^{(n)}, \Upsilon^{(n)}, Z$ can be significantly enhanced in the case of large sterile fermion masses and a non-negligible active-sterile mixing. We illustrate that feature in a specific minimal realization of the inverse seesaw mechanism, known as (2, 3)-ISS, and in an effective model in which the presence of non-standard sterile fermions is parameterized by means of one heavy sterile (Majorana) neutrino.

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1 Introduction

So far no signal of new physics has been observed but its search is important in order to understand how to enlarge the Standard Model (SM) to solve both the hierarchy and flavor problems. One of the most significant observations requiring to go beyond the Standard Model (BSM) is the assessment that neutrinos are massive and that they mix \cite{1}. Possible SM extensions aiming at incorporating massive neutrinos give rise to interesting collider signatures and open the door to new phenomena such as lepton flavor violating (LFV) decays.

Currently, the search for manifestations of LFV constitutes a goal of several experimental facilities dedicated to rare lepton decays, such as $\ell_a \rightarrow \ell_B \gamma$ and $\ell_a \ell_B \ell_C \gamma$, and to the neutrinoless $\mu - e$ conversion in muonic atoms. One of the most stringent bounds from these searches is the one derived by the MEG collaboration, $B(\mu \rightarrow e\gamma) < 5.7 \times 10^{-13}$ \cite{2}, which is expected to be improved to a planned sensitivity of $6 \times 10^{-14}$ \cite{3}. Moreover, the bound $B(\mu \rightarrow eee) < 1.0 \times 10^{-12}$, set by the SINDRUM experiment \cite{4}, is expected to be improved by the Mu3e experiment where a sensitivity $\sim 10^{-16}$ is planed \cite{5}. Limits on the $\tau$ radiative decays \cite{6} and the 3-body decays of $\tau \rightarrow e\gamma \nu$ \cite{7,8} appear to be less stringent right now, but are likely to be improved at Belle II \cite{8}, where the search of LFV decays of $B$-meson will be made too \cite{9}. The most promising developments regarding LFV are those related to the $\mu - e$ conversion in nuclei. The present bound for $\mu^- \text{Ti} \rightarrow e^- \text{Ti}$ conversion rate is $4.3 \times 10^{-12}$ \cite{10}, and the planned sensitivity is $\sim 10^{-18}$ \cite{11}. Similar is the case for gold and aluminum \cite{12,13}.

Searches for LFV are also conducted in high-energy experiments and a first bound on the Higgs boson LFV decay $h \rightarrow \mu \tau$ has been reported by the CMS collaboration \cite{14}. The LHCb collaboration, instead, reported the bound $B(\tau \rightarrow 3\mu) < 8.0 \times 10^{-8}$ \cite{15}, which is likely to be improved in the near future \cite{16}. Notice also that they already improved the bounds on $B(B_{(s)} \rightarrow e\mu)$ by an order of magnitude \cite{17}.

In this work we will focus on the indirect probes of New Physics (NP) through the LFV processes of neutral vector bosons, namely $V \rightarrow \ell_a \ell_B$, with $\ell_{a,B} \in \{e,\mu,\tau\}$, and $V \in \{\phi, \psi^{(n)}, \Upsilon^{(n)}, Z\}$, where $\psi^{(n)}$ stands for $J/\psi$ and its radial excitations, and similarly for $\Upsilon^{(n)}$. Most of the research in this direction reported so far is related to the $Z \rightarrow \ell_a \ell_B$ decay modes. More specifically, the experimental bounds, obtained at LEP are found to be $B(Z \rightarrow e^+e^-) < 1.7 \times 10^{-6}$ \cite{18}, $B(Z \rightarrow \mu^+\mu^-) < 1.2 \times 10^{-5}$ \cite{18,19}, and $B(Z \rightarrow e^+\tau^\pm) < 9.8 \times 10^{-6}$ \cite{19,20}. One of these bounds has been improved at LHC, namely $B(Z \rightarrow e^+\mu^\pm) < 7.5 \times 10^{-7}$ \cite{21}. On the theory side, the $Z$ decays have been analyzed in the extensions of the SM involving additional massive and sterile neutrinos that could mix with the standard (active) ones and thus give rise to the LFV decay rates \cite{22,24}. A similar approach has been also adopted in ref. \cite{25}, in the perspective of a Tera-$Z$ factory FCC-ee \cite{26} for which a targeted sensitivity is expected to be $B(Z \rightarrow e^+\mu^\pm) \sim 10^{-13}$ \cite{27}.

Lepton flavor conserving decays of quarkonia have been measured to a high accuracy which can be actually used to fix the hadronic parameters (decay constants). Otherwise, one can use the results of numerical simulations of QCD on the lattice which are nowadays accurate as well \cite{28,31}. The experimentally established bounds for the simplest LFV
decays of quarkonia are \[16\]:

\[
\begin{align*}
B(\phi \to e\mu) &< 2.0 \times 10^{-6}, \\
B(J/\psi \to e\mu) &< 1.6 \times 10^{-7}, \\
B(J/\psi \to \mu\tau) &< 2.0 \times 10^{-6}, \\
B(\Upsilon \to \mu\tau) &< 6.0 \times 10^{-7}, \\
B(\Upsilon(2S) \to e\tau) &< 8.3 \times 10^{-6}, \\
B(\Upsilon(2S) \to \mu\tau) &< 2.0 \times 10^{-6}, \\
B(\Upsilon(3S) \to e\tau) &< 4.2 \times 10^{-6}, \\
B(\Upsilon(3S) \to \mu\tau) &< 3.1 \times 10^{-6},
\end{align*}
\]

where each mode is to be understood as \( B(V \to \ell_\alpha \ell_\beta) = B(V \to \ell_\alpha^+ \ell_\beta^-) + B(V \to \ell_\alpha^- \ell_\beta^+) \).

Despite the appreciable experimental work on the latter observables, only a few theoretical studies have been carried out so far. The authors of ref. \[37\] applied a vector meson dominance approximation to \( \mu \to 3e \) and expressed the width of the latter process, \( \Gamma(\mu \to 3e) = \Gamma(V \to eee) \Gamma(V \to ee) \). Since the values of \( \Gamma(V \to ee) \) are very well known experimentally \[16\], the experimental bound on \( \Gamma(\mu \to 3e) \) is then used to obtain an upper bound on the phenomenological coupling \( g_{V\mu\ell} \), which is then converted to an upper bound on \( \Gamma(V \to \mu e) \). Similar approach has been used in ref. \[38\] where instead of \( \mu \to 3e e \), the authors considered the \( \mu - e \) conversion in nuclei \((N)\), which they described in terms of a product of couplings \( g_{V\mu e} \) and \( g_{VNN} \). The latter could be extracted from the experimentally measured \( \Gamma(V \to p\bar{p}) \), and with that knowledge the experimental upper bound on \( R(\mu Ti \to e Ti) \) results in an upper bound on \( \Gamma(V \to \mu e) \). A more dynamical approach in modeling the \( V \to \ell_\alpha \ell_\beta \) processes has been made in a supersymmetric extension of the SM with type-I seesaw \[39\].

Sterile fermions were proposed in various neutrino mass generation mechanisms, but the interest in their properties was further motivated by the reactor/accelerator anomalies \[40–43\], a possibility to offer a warm dark matter candidate \[44–46\], and by indications from the large scale structure formation \[47–49\].

Incorporating neutrino oscillations (masses and mixing \[1\]) into the SM implies that the charged current is modified to

\[
-\mathcal{L}_{cc} = \frac{g}{\sqrt{2}} U^{\alpha i} \bar{\ell}_\alpha \gamma^\mu P_L \nu_i W^-_\mu + c.c.,
\]

(1)

\( U \) being the leptonic mixing matrix, \( \alpha \) the flavor of a charged lepton, and \( i = 1, \ldots, n_\nu \) denotes a physical neutrino state. If one assumes that only three massive neutrinos are present, the matrix \( U \) corresponds to the unitary PMNS matrix. In that situation the GIM mechanism makes the decay rates \( B(V \to \ell_\alpha^+ \ell_\beta^-) \) completely negligible, \( \lesssim 10^{-50} \). That feature, however, can be drastically changed in the presence of a non-negligible mixing with heavy sterile fermions. In what follows we will consider such situations, derive analytical expressions for \( B(V \to \ell_\alpha \ell_\beta) \), and discuss a specific realization of the inverse seesaw mechanism, known as (2,3)-ISS \[50\]. We will also discuss a simplified model in which the
effect of the heavy sterile neutrinos is described by one effective sterile neutrino state with non-negligible mixing with active neutrinos.\footnote{In this work, due to the tension between the most recent Planck results on extra light neutrinos (relics) and the reactor/accelerator anomalies, we will consider the effect of (heavier) sterile neutrinos not contributing as light relativistic degrees of freedom \cite{51}. We will require our models to be compatible with current experimental data and constraints and to fulfill the so-called perturbative unitary condition which}

Despite several differences, our approach is similar to the one discussed in ref. \cite{53}, where the SM has been extended by new, heavy, Dirac neutrinos, singlets under $SU(2) \times U(1)$, and applied to a number of low energy decay processes. Our sterile neutrinos are Majorana and we apply the approach to the leptonic decays of quarkonia for the first time.

The remainder of this paper is organized as follows: In Sec. 2 we formulate the problem in terms of a low energy effective theory of a larger theory which contains heavy sterile neutrinos, we derive expression for $B(V \to \ell_\alpha \ell_\beta)$ and compute the Wilson coefficients. In Sec. 3 we briefly describe the specific models with sterile neutrinos which are used in this paper to produce our results presented in Sec. 4. We finally conclude in Sec. 5.

2 LFV decay of Quarkonia - Effective Theory

In this section we formulate a low energy effective theory of the LFV decays of quarkonia of type $V \to \ell_\pm \ell_\mp$, and express the decay amplitude in terms of the quarkonium decay constants and the corresponding Wilson coefficients. The latter are then computed in the extensions of the SM which include the heavy sterile neutrinos. We also derive the expression relevant to $\Gamma(Z \to \ell_\pm \ell_\mp)$.

2.1 Effective Hamiltonian

Keeping in mind the fact that we are extending the SM by adding sterile fermions, without touching the gauge sector of the theory, the decays of vector quarkonia, $V(q) \to \ell_\alpha(p) \ell_\beta(q-p)$, can only occur through the photon and the Z-boson exchange at tree level. In the lepton flavor conserving processes the Z-exchange terms are very small with respect to those arising from the electromagnetic interaction and are usually neglected. The generic effective Hamiltonian can be written as

$$H_{\text{eff}} = \frac{e^2 g^2}{2 m_V} \bar{q} \gamma_\mu Q \cdot \bar{\ell}_\alpha \left[ C_{VL} \gamma^\mu P_L + C_{VR} \gamma^\mu P_R + \frac{p^\mu}{m_W} (C_R P_R + C_L P_L) \right] \ell_\beta,$$  \hspace{1cm} (2)$$

where $Q_Q$ is the electric charge of the quark $Q$, $m_V$ is the mass of quarkonium $V$ which is dominated by the valence quark configuration $Q\bar{Q}$, $C_{VL,VR,L,R}$ are the Wilson coefficients, $p$ is the momentum of one of the outgoing leptons, and $P_{LR} = \frac{1}{2}(1 \mp \gamma_5)$. Contributions to the scalar (left and right) terms are suppressed by $m_{\alpha,\beta}/m_W$, where $m_{\alpha,\beta}$ are the charged lepton masses. In this section we will keep such terms so that our expressions can be useful to approaches in which the scalar bosons are taken in consideration. For our phenomenological discussion, however, it is worth emphasizing that $C_{L,R,VR} \ll C_{VL}$.\footnote{In this work, due to the tension between the most recent Planck results on extra light neutrinos (relics) and the reactor/accelerator anomalies, we will consider the effect of (heavier) sterile neutrinos not contributing as light relativistic degrees of freedom \cite{51}. We will require our models to be compatible with current experimental data and constraints and to fulfill the so-called perturbative unitary condition which}
Figure 1: Diagrams contributing the LFV decay of quarkonia $V \to \ell_\alpha \ell_\beta$. The blob in the first diagram is related to the penguin loop that generates the LFV, and the box diagram is particularly important to be included in the case of $\Upsilon(n) \to \ell_\alpha \ell_\beta$ because of $V_{tb} \simeq 1$ and of the top quark mass, making the box diagram contribution to the Wilson coefficient significant.

Without entering the details of calculation it is easy to verify that the only relevant diagrams are those shown in Fig. 1, and therefore the structure of the Wilson coefficients $C_i$ reads,

$$C_i = C_i^{\gamma} + C_i^{Z} \frac{1}{\sin^2 \theta_W \cos^2 \theta_W} \frac{m_V^2}{m_W^2 - m_Z^2} \frac{g_V^Q}{Q_Q} + C_i^{\Box} |V_{Qq}|^2 \frac{1}{\sin^2 \theta_W} \frac{m_W^2}{Q_Q},$$

where $C_i^{\gamma,Z}$ are the contributions arising from the photon and the $Z$-boson exchange, while $C_i^{\Box}$ comes from the box diagram that involves the Cabibbo–Kobayashi-Maskawa coupling $V_{Qq}$. In the above expressions $g_V^Q = \frac{1}{2} I_3^Q - Q_Q \sin^2 \theta_W$. The blob in the diagram shown in Fig. 1 stands for the lepton loop diagrams that may contain one or two neutrino states and which, in the extensions of the SM involving a heavy neutrino sector, will give rise to the LFV decay due to the effect of mixing which is parametrized by the matrix $U$ [see eq. (1)]. Separate contributions coming from different diagrams can be further reduced by factoring out the neutrino mixing matrix elements, namely

$$C_i^{\gamma,\Box} = \sum_{k=1}^{n_\nu} U_{\beta k} U_{\alpha k}^* C_i^{\gamma,\Box;k}, \quad \text{and} \quad C_i^{Z} = \sum_{k=1}^{n_\nu} U_{\beta k} U_{\alpha k}^* C_i^{Z,k} + \sum_{k=1}^{n_\nu} \sum_{j=1}^{n_\nu} U_{\beta k} U_{\alpha j}^* C_i^{Z,kj},$$

where we see that the term involving two neutrino eigenstates appears only in the $Z$ coefficient because it is related to the vertex $Z\nu_\mu \nu_j$.

Using the effective Hamiltonian $[2]$ and parameterizing the hadronic matrix as

$$\langle 0|\bar{Q}\gamma_\mu Q|V(q,\sigma)\rangle = f_V m_V \varepsilon_\mu^\sigma,$$

puts a strong constraint on the models for the very heavy sterile fermion(s) $[52]$. We remind the reader that the ground vector meson $s\bar{s}$, $c\bar{c}$, $b\bar{b}$ states are $\phi$, $J/\psi$, $\Upsilon$, respectively, and the corresponding charges are $Q_{s,b} = -1/3$ and $Q_c = 2/3$.

$^3$The box diagram contribution to $V \to \ell_\alpha \ell_\beta$ in the case of $V = \Upsilon$ is dominated by the top quark ($|V_{tb}| \simeq 1$); for $V = \psi$ it is negligible because the contribution of the $b$ quark is Cabibbo suppressed ($|V_{cb}| \simeq 0.004$) while the Cabibbo allowed one ($|V_{cs}| \simeq 0.99$) is suppressed by the strange quark mass; for $V = \phi$, the contributions of the charm and top quarks are comparable but overall smaller than in the $\Upsilon \to \ell_\alpha \ell_\beta$ case.
where \( f_V \) is the decay constant of a quarkonium \( V \) with momentum \( q \) and in a polarization state \( \sigma \), we can write the decay rate as,

\[
\Gamma(V \to \ell^-_a \ell^+_\beta) = \frac{8\pi Q^2 \alpha^2}{3m^3_V} G^2_F m^4_W \left( \frac{f_V}{m_V} \right)^2 \lambda^{1/2}(m^2_V, m^2_\alpha, m^2_\beta) \phi_C, \tag{6}
\]

with

\[
\lambda(a^2, b^2, c^2) = \frac{a^2 - (b - c)^2}{a^2 - (b + c)^2}, \tag{7}
\]

and

\[
\phi_C = \left(-g^{\mu\nu} + \frac{q^{\mu} q^{\nu}}{m^2_V}\right) \text{tr} \left[ (\not{\phi} - \not{m}_\beta) \cdot (C_{\nu L}^* \gamma^\mu P_L + C_{\nu R}^* \gamma^\mu P_R + C^*_{\nu L} \frac{p^\mu_{\nu L}}{m_W} P_L + C^*_{\nu R} \frac{p^\mu_{\nu R}}{m_W} P_R) \right. \\
\left. \cdot (\not{\phi} - \not{m}_\alpha) \cdot (C_{\nu L}^* \gamma^\nu P_L + C_{\nu R}^* \gamma^\nu P_R + C^*_{\nu L} \frac{p^\nu_{\nu L}}{m_W} P_L + C^*_{\nu R} \frac{p^\nu_{\nu R}}{m_W} P_L) \right], \tag{8}
\]

which gives,

\[
\phi_C = \frac{1}{4m^2_V m^2_W} \left\{ \lambda(m^2_V, m^2_\alpha, m^2_\beta) \left[ (m^2_V - m^2_\alpha - m^2_\beta)(|C_L|^2 + |C_R|^2) - 4\text{Re}(C^*_{\nu L} C_{\nu R}) m_\alpha m_\beta \right. \right.
\]
\[+ 4m_W \text{Re}(C^*_{\nu L} (C_{\nu L} m_\beta + C_{\nu R} m_\alpha) + C^*_{\nu R} (C_{\nu L} m_\alpha + C_{\nu R} m_\beta)) \left. \right. \]
\[+ 4m^2_W (|C_{\nu L}|^2 + |C_{\nu R}|^2) \left[ 2m^4_V - m^2_V (m^2_\alpha + m^2_\beta) - (m^2_\alpha - m^2_\beta)^2 \right] \right. \]
\[\left. + 48m^2_W m^2_\alpha m_\beta \text{Re}(C^*_{\nu L} C_{\nu R}) \right\}. \tag{9}
\]

As we mentioned above, we consider in our framework \( C_{\nu L} \gg C_{\nu R,R,L} \), and therefore we can write,

\[
\Gamma(V \to \ell^+_\alpha \ell^-_\beta) = \frac{32\pi Q^2 \alpha^2}{3m^3_V} f_V G^2_F m^4_W |C_{\nu L}|^2 \lambda^{1/2}(m^2_V, m^2_\alpha, m^2_\beta) \left[ 1 - \frac{(m^2_\alpha + m^2_\beta)}{2m^2_V} - \frac{(m^2_\alpha - m^2_\beta)^2}{2m^4_V} \right], \tag{10}
\]

where \( \lambda(a^2, b^2, c^2) \) is given in eq. \((7)\). In this last expression we also used \( \Gamma(V \to \ell^+_\alpha \ell^-_\beta) = \Gamma(V \to \ell^-_\alpha \ell^+_\beta) = \Gamma(V \to \ell^+_\alpha \ell^+_\beta) \).

Besides quarkonia we will also revisit the issue of adding extra species of sterile neutrinos to the decay of \( Z \to \ell^+_\alpha \ell^-_\beta \). In that case the effective Hamiltonian can be written as,

\[
\mathcal{H}_{\text{eff}}^Z = \frac{g^2}{2 \cos^2 \theta_W} \bar{\ell}_\alpha \left[ D_{\nu L} \gamma^\mu P_L + D_{\nu R} \gamma^\mu P_R + D_L P_L + D_R P_R \right] \ell_\beta Z^\mu, \tag{11}
\]

where the Wilson coefficients are now denoted by \( D_i \) and take the form,

\[
D_i = \sum_{k=1}^{n_L} U^*_{\beta k} U_{\alpha k} C^Z_{i,k} + \sum_{k=1}^{n_L} \sum_{j=1}^{n_L} U^*_{\beta k} U^*_{\alpha j} C^Z_{i,kj}. \tag{12}
\]

The decay rate in the similar limit, \( D_{\nu L} \gg D_{\nu R,R,L} \), reads,

\[
\Gamma(Z \to \ell^-_\alpha \ell^+_\beta) = \frac{8\sqrt{2} C^3_F m^6_W}{3\pi m_Z \cos^2 \theta_W} |D_{\nu L}|^2 \lambda^{1/2}(m^2_Z, m^2_\alpha, m^2_\beta) \left[ 1 - \frac{(m^2_\alpha + m^2_\beta)}{2m^2_Z} - \frac{(m^2_\alpha - m^2_\beta)^2}{2m^4_Z} \right]. \tag{13}
\]
2.2 Wilson coefficients

Concerning the computation of the Wilson coefficients we stress again that our results are obtained in a theory in which the Standard Model is extended to include extra species of sterile fermions, without changing the gauge sector. The origin of the leptonic mixing matrix $U$ is model dependent and in order to be able to do a phenomenological analysis, we will have to adopt a specific model which will be discussed in the next Section.

The blob in the diagram shown in Fig. [1] stands for a series of diagrams such as those displayed in Fig. [2]. All of them, including the box diagram in Fig. [1], have been computed in the Feynman gauge and the results are collected in appendix A. Here we focus on the most important contributions in the case of large masses of sterile (Majorana) neutrinos. Contributions to the Wilson coefficients coming from vertex diagrams can be divided into two pieces: those involving only one neutrino in the loop, $C_{Z\gamma}^{VL}(x_i)$, where $x_i = m_i^2/m_W^2$, and those with two neutrinos in the loop, $C_{Z\gamma}^{VL}(x_i, x_j)$. In the limit of large values of $x_{i,j} \gg 1$, we find the following behavior

$$C_{Z\gamma}^{VL}(x_i) \xrightarrow{x_i \gg 1} \frac{5}{32\pi^2} \log x_i + \text{finite term} + \mathcal{O}(1/x_i) \sim \log x_i,$$

$$C_{Z\gamma}^{VL}(x_i, x_i) \xrightarrow{x_i \gg 1} \frac{C_{ii}}{64\pi^2} \left\{ (2x_i + 3 - 4 \log x_i) + x_i \left( \log x_i - \frac{7}{2} \right) \right\} + \ldots$$

To illustrate the relative contribution of the different diagrams we fix the values of the coefficients $C_{ij} = 10^{-5}$, and plot $|C_{VL}(x_i) - C_{VL}(0)|$ and $|C_{Z\gamma}^{VL}(x_i, 2x_i) - C_{Z\gamma}^{VL}(0, 0)|$ for the case of $\Upsilon \rightarrow \mu\tau$, cf. Fig. [3] [4]. We see that only for very large masses the diagrams with two neutrinos in the loop become more important than those with one neutrino state. We should stress that each contribution to $C_{VL}(x_i)$, i.e. $C_{VL}^{box}(x_i)$ and $C_{VL}^{Z\gamma}(x_i)$, scales as $\log x_i$ for large values of $x_i$, except for $C_{VL}^{\gamma}(x_i)$ which goes to a constant in the same limit. That can also be seen in Fig. [3] where in the left plot we show the dependence of the total $C_{VL}(x_i)$ on $x_i$ and in the right plot we show $C_{VL}^{\gamma}(x_i)$ and its

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4Due to unitarity of the mixing matrix $U$, the terms in the Wilson coefficients that do not depend on neutrino masses give a vanishing contribution after summing over all neutrino states. We thus subtract the constant terms in the plots in order to better appreciate the dependence on neutrino masses. Notice also that $C_{ij} = 10^{-5}$ is in agreement with all constraints discussed in the text when the neutrino masses are below $\mathcal{O}(100)$ TeV.
dependence on the mass of the initial decaying meson, $\phi$, $J/\psi$ and $\Upsilon$. The contribution of sterile neutrinos to the LFV decay of $\Upsilon$ is larger than the one to lighter mesons, since the Wilson coefficients are also proportional the mass of the initial particle.

Figure 3: On the left panel are shown $C_{VL}(x_i)$ and $C_{VL}^Z(x_i, x_j)$, for $x_j = 2x_i$, as functions of $m_i = m_W \sqrt{x_i}$, mass of the heavy sterile neutrino propagating in the loops. For illustration purpose, the couplings $C_{ij}$ were fixed to a common value, $10^{-5}$, and the example corresponds to the $\Upsilon \rightarrow \mu \tau$ decay. Right panel: $C_{VL}^Z(x_i)$ is plotted as a function of $m_i$ for the case of $V \rightarrow e\mu$ in three specific cases $V \in \{\phi, J/\psi, \Upsilon\}$. In both cases the value of functions at $x_{i,j} = 0$ have been subtracted away.

3 SM in presence of sterile fermions

With the expressions derived above, we now have to specify a model for lepton mixing (couplings) $U_{ai}$ in the presence of heavy sterile neutrinos propagating in the loops. We opt for a minimal realization of the inverse seesaw mechanism for generation of neutrino masses, which is nowadays rather well constrained by the available experimental data. Furthermore, we will use a parametric model containing one effective sterile neutrino, which essentially mimics the behavior at low energy scales of mechanisms involving heavy sterile fermions.

3.1 The (2,3)-inverse seesaw realization

Among many possible realizations to account for massive neutrinos, the inverse seesaw mechanism (ISS) [54] offers a possibility to accommodate the smallness of the active neutrino masses for a comparatively low seesaw scale, but still with natural $\mathcal{O}(1)$ Yukawa couplings, which renders this scenario phenomenologically appealing. Indeed, depending on their masses and mixing with active neutrinos, the new states can be produced in collider and/or low-energy experiments, and their contribution to physical processes can be
sizable. ISS, embedded in the SM, results in a mass term for neutrinos of the form,

\[ -\mathcal{L}_{\text{mass}} = \frac{1}{2} n_L^T C M n_L + \text{h.c.}, \]  

(15)

where \( C \equiv i\gamma^2\gamma^0 \) is the charge conjugation matrix and \( n_L \equiv (\nu_{L,\alpha}, \nu_{R,i}^\alpha, s_j)^T \). Here \( \nu_{L,\alpha}, \alpha = e, \mu, \tau \), denotes the active (left-handed) neutrino states of the SM, while \( \nu_{R,i}^\alpha \) \( (i = 1, \#\nu_R) \) and \( s_j \) \( (j = 1, \#s) \) are right-handed neutrino fields and additional fermionic gauge singlets, respectively. The neutrino mass matrix \( M \) then has the form,

\[
M \equiv \begin{pmatrix}
0 & d & 0 \\
d^T & 0 & n \\
0 & n^T & \mu
\end{pmatrix},
\]

(16)

where \( d, n, \mu \) are complex matrices. \footnote{5}{It is in general possible to consider also a nonzero value for the central entry of the matrix (16), with elements at a mass scale similar to the one of \( \mu \). These parameters, however, only affect neutrino masses and mixing at loop level \cite{55}, which is why we do not consider them here.
}

The Dirac mass matrix \( d \) arises from the Yukawa couplings to the SM Higgs boson, \( \tilde{H} = i\sigma^2 H \),

\[
Y_{\alpha i} \bar{\nu}_L^\alpha \tilde{H} \nu_{R,i} + \text{h.c.}, \quad \ell_L^\alpha = \begin{pmatrix} \nu_L^\alpha \\ e_L^\alpha \end{pmatrix}, \]

(17)

while the matrix \( \mu \), instead, contains the Majorana mass terms for the sterile fermions \( s_j \). By assigning a leptonic charge \( L = +1 \) to both \( \nu_R \) and \( s \), one makes sure that the off diagonal terms are lepton number conserving, while \( s^T C s \) violates the lepton number by two units. Furthermore, the interesting feature of this seesaw realization is that the entries of \( \mu \) can be made small in order to accommodate for \( \mathcal{O}(\text{eV}) \) masses of active neutrinos, with large Yukawa couplings. This is not in conflict with naturalness since the lepton number is restored in the limit of \( \mu \to 0 \). \footnote{6}{In this work we consider configurations in which the entries in the above matrices fulfill a naturalness criterion, \( |\mu| \ll |d| < |n| \) \cite{50}.}

Concerning the additional sterile states \( \nu_R \) and \( s \), since up to now there is no direct evidence for their existence and because they do not contribute to anomalies, their number is unknown. In ref. \cite{50}, it was shown that it is possible to construct several minimal distinct realizations of ISS, each reproducing the correct neutrino mass spectrum and satisfying all phenomenological constraints. More specifically, it was shown that, depending on the number of additional fields, the neutrino mass spectrum obtained for each ISS realization is characterized by either 2 or 3 mass scales, one corresponding to \( m_\nu \approx \mu d^2/n^2 \) (light neutrino masses), one corresponding to the heavy mass eigenstates [mass scale of the matrix \( n \) of eq. (16)], and finally an intermediate scale \( \sim \mu \), only present if \( \#s > \#\nu_R \). This allows to identify two truly minimal ISS realizations that comply with all experimental bounds, namely the (2,2)-ISS model, which corresponds to the SM extended by two RH neutrinos and two additional sterile states, leading to a three-flavor mixing scheme, and the (2,3)-ISS realization, where the SM is extended by two RH neutrinos and three sterile states leading to a 3+1-mixing scheme. Interestingly, the lightest sterile neutrino with a mass
around eV in the (2,3)-ISS can be used to explain the short baseline (reactor/accelerator) anomaly [40–43] if its mass lies around eV, or to provide a dark matter candidate if the lightest sterile state were in the keV range [56].

### 3.2 A model with one effective sterile fermion

Since the generic idea of obtaining a significant contribution to our observables applies to any model in which the active neutrinos have sizable mixing with some additional singlet states (sterile fermions), we can use an effective model with 3 light active neutrinos plus 1 extra sterile neutrino.

The introduction of this extra state implies three new active-sterile mixing angles ($\theta_{14}, \theta_{24}, \theta_{34}$), two extra Dirac CP violating phases ($\delta_{14}, \delta_{34}$) and one additional Majorana phase ($\phi_{41}$). The lepton mixing matrix is then a product of 6 rotations times the Majorana phases, namely,

$$U = R_{34}(\theta_{34}, \delta_{34}) \cdot R_{24}(\theta_{24}) \cdot R_{14}(\theta_{14}, \delta_{14}) \cdot R_{23} \cdot R_{13} \cdot R_{12} \cdot \text{diag}(\phi_{21}, \phi_{31}, \phi_{41})$$

where the rotation matrices $R_{34}, R_{24}, R_{14}$ can be defined as:

$$R_{34} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & \cos \theta_{34} & \sin \theta_{34} \cdot e^{-i \delta_{34}} \\ 0 & 0 & -\sin \theta_{34} \cdot e^{i \delta_{34}} & \cos \theta_{34} \end{pmatrix}$$

$$R_{24} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & \cos \theta_{24} & 0 & \sin \theta_{24} \\ 0 & 0 & 1 & 0 \\ 0 & -\sin \theta_{24} & 0 & \cos \theta_{24} \end{pmatrix}$$

$$R_{14} = \begin{pmatrix} \cos \theta_{14} & 0 & 0 & \sin \theta_{14} \cdot e^{-i \delta_{14}} \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\sin \theta_{14} \cdot e^{i \delta_{14}} & 0 & 0 & \cos \theta_{14} \end{pmatrix}$$

In the framework of the SM extended by sterile fermion states, which have a non-vanishing mixing with active neutrinos, the Lagrangian describing the leptonic charged currents becomes

$$-\mathcal{L}_{cc} = \frac{g}{\sqrt{2}} U^{\alpha i} \bar{\ell}_\alpha \gamma^\mu P_L l_i W^-_\mu + \text{c.c.},$$

where $i = 1, \ldots, n_\nu$ denotes the physical neutrino states, and $\alpha = e, \mu, \tau$ are the flavors of charged leptons. In the case of the SM with three neutrino generations, $U$ is the PMNS matrix, while in the case of $n_\nu \geq 4$, the $3 \times 3$ submatrix ($\tilde{U}_{\text{PMNS}}$), is not unitary anymore and one can parameterize it as

$$U_{\text{PMNS}} \rightarrow \tilde{U}_{\text{PMNS}} = (1 - \eta) U_{\text{PMNS}}.$$
where $\tilde{\eta}$ is a matrix that accounts for the deviation of $\tilde{U}_{\text{PMNS}}$ from unitarity \cite{57,58}, due to the presence of extra fermion states. Many observables are sensitive to the active-sterile mixing and their current experimental values can be used to constrain the $\tilde{\eta}$-matrix \cite{59}.

In order to express the deviation from unitarity in terms of a single parameter, we define

$$
\eta = 1 - |\det \tilde{U}_{\text{PMNS}}|,
$$

which, in the case of the extension of the SM by only one sterile fermion and in terms of the mixing angles defined above, reads

$$
\eta = 1 - |\cos \theta_{14} \cos \theta_{24} \cos \theta_{34}|.
$$

4 Results and discussion

In this section we present and discuss our results.

Since the Wilson coefficients of the processes discussed here are proportional to the mass of the decaying particle, it is quite obvious that the most significant enhancement of $B(V \to \ell_\alpha \ell_\beta)$ will occur for $V = \Upsilon$ and its radial excitations. For this reason we will present plots of our results for this decay channel. Plots for other channels are completely similar which is why we do not display them. Before we further discuss the impact of the active-sterile neutrino mixing on the LFV decay rates, we first specify the constraints on parameters of both of our models.

In Fig. 4 (left panel), we plot the dependence of $\eta$ with respect to the mass of the effective sterile neutrino $m_4$. Gray points in that plot are obtained by varying the mass of the lightest neutrino, $m_{\nu_e} \in (10^{-21}, 1)$ eV, and by imposing the following constraints: (i) Neutrino data (masses and mixing angles) respect the normal hierarchy, with $\Delta m_{12}^2 = 7.5(2) \times 10^{-5}$ eV, and $\Delta m_{21}^2 = 2.46(5) \times 10^{-3}$ eV \cite{1}. We checked that our final results do not change in any significant manner if the inverse hierarchy is adopted. Furthermore, we vary the three mixing angles with the fourth neutrino by assuming $\theta_{14} \in (0, 2\pi]$, while keeping the other three mixing angles to their best-fit values, namely $\sin^2 \theta_{12} = 0.30(1)$, $\sin^2 \theta_{23} = 0.47(4)$, $\sin^2 \theta_{13} = 0.022(1)$ \cite{1}. (ii) The selected points satisfy the upper bound $B(\mu \to e\gamma) < 5.7 \times 10^{-13}$ \cite{2}. (iii) The results for $R_\pi = \Gamma(\pi \to e\nu_e)/\Gamma(\pi \to \mu\nu_\mu)$, $R_K$, $\Gamma(W \to \ell \nu_\ell)$ and for $\Gamma(Z \to \text{invisible})$, remain consistent with experimental findings. We see that for all (heavy) sterile neutrino masses the unitarity breaking parameter is $\eta \lesssim 0.005$. That parameter space is not compatible with the perturbative unitarity requirement, which for $m_4 \gg m_W$ translates into

$$
\frac{G_F m_4^2}{\sqrt{2}\pi} \sum_\alpha |U_{\alpha 4}|^2 < 1.
$$

The resulting region, i.e. the one that satisfies constraints (i,ii,iii) and eq. (24), is depicted by blue points (dark region) in Fig. 4 where we see that the parameter $\eta$ is indeed diminishing with the increase of the heavy sterile mass $m_4$. In other words, the decoupling of a

\footnote{To write it in the form given in eq. (24), we replaced $\alpha_W = g^2/(4\pi) = \sqrt{2}G_F m_W^2/\pi$.}
Figure 4: $\eta$-parameter, that parametrizes the size of mixing between the active and heavy sterile states, is plotted vs. the mass of the heavy sterile state. The gray points (left panel) correspond to solutions complying with all experimental data and constraints discussed in the text except for perturbative unitary condition (24), which we then applied to obtain the region of dark-blue points. In the case of the (2,3)-ISS model (right-panel), we further imposed constraints of ref. [59] on the matrix $\tilde{\eta}$, as well as the bound $B(\mu \to eee) < 10^{-12}$, resulting in the bright-blue region of points.

very heavy sterile neutrino entails the unitarity of the $3 \times 3$ submatrix $\tilde{U}_{\text{PMNS}}$. Decoupling from active neutrinos for very large masses was also explicitly emphasized in ref. [60]. We should mention that, besides the above constraints, we also implemented the constraint coming from $B(\mu \to eee) < 10^{-12}$ [4], but it turns out that the present experimental bound does not bring any additional improvement.

By imposing the constraints (i) and eq. (24) on the (2,3)-ISS model, we get a similar region of allowed (blue) points in the right plot of Fig. 4. A notable difference with respect to the situation with one effective sterile neutrino is that the region of very small mixing angles is excluded due to relations between the active neutrino masses and the active-sterile neutrino mixing, cf. ref. [50]. For very heavy $m_5$, on the other hand, the range of allowed $\eta$'s shrinks and eventually vanish with $m_5 \to \infty$. Furthermore, we use the results of ref. [59] which are derived in the Minimal Unitarity Violation scheme in which the heavy sterile neutrino fields are integrated out, and therefore the observables computed in that scheme are functions of the deviation of PMNS matrix from unitarity only [61]. We adapt and apply them to our (2,3)-ISS model and get a region of the bright blue points, as shown in the right panel of Fig. 4. To further constrain the parameter space we find it useful to account for the experimental bound on $B(\mu \to eee) < 1 \times 10^{-12}$, as discussed in refs. [23,60,62]. This latter constraint appears to be superfluous in most of the parameter space, once the constraints of eq. (24) and ref. [59] are taken into account, except in the range $10 \text{ TeV} \lesssim m_5 \lesssim 100 \text{ TeV}$, where the bound $B(\mu \to eee) < 1 \times 10^{-12}$ restricts the parameter space relevant to $B(V \to e\mu)$.

We recall that in the (2,3)-ISS model, $m_4$ stands for the mass of the light sterile state whose impact on the decays discussed here is negligible (as seen from eq. 14), while $m_5 > m_4$ can be large and is important for $B(V \to \ell_\alpha \ell_\beta)$. 

\footnote{We recall that in the (2,3)-ISS model, $m_4$ stands for the mass of the light sterile state whose impact on the decays discussed here is negligible (as seen from eq. 14), while $m_5 > m_4$ can be large and is important for $B(V \to \ell_\alpha \ell_\beta)$.}
Figure 5: $B(\Upsilon \to \mu \tau)$ and $B(Z \to \mu \tau)$ are shown as functions of the heavy sterile neutrino(s) mass, and in both models considered in this paper. The points are selected in such a way that the models are consistent with constraints discussed in the text and shown in Fig. 4. Dashed horizontal lines correspond to the current experimental upper bounds for these decay rates. Notice again that the mass of the heavy sterile neutrino is denoted as $m_4$ in the effective model, and $m_5$ in the (2,3)-ISS model because the latter contains a lighter sterile neutrino state the impact of which is negligible on the decay modes discussed here.

We also mention that we attempted implementing the constraints coming from various laboratory experiments, summarized in ref. [63], but since those results only impact the region of relatively small sterile neutrino masses ($m_5 \lesssim 100$ GeV), they are of not relevance for the present study.

After having completed the discussion on several constraints, we present our results for branching fractions $B(V \to \mu \tau)$ depending on the mass of heavy sterile neutrino(s). In Fig. 5 we plot our results for $V = \Upsilon$ and $V = Z$, for which the enhancement is more pronounced. Other cases of $V$ result in similar shapes but the upper bound becomes lower. In Tab. 1 we collect our results for three values of the heavy sterile neutrino(s) mass.

To better appreciate the enhancement of the LFV decay rates shown in Fig. 5 we
emphasize that both of them are $\text{B}(V \rightarrow \mu \tau) < 10^{-50}$ in the absence of heavy sterile neutrinos. Current experimental bounds in both cases are shown by dashed lines. Since those bounds are expected to improve in near future, a possibility of seeing the LFV modes discussed in this paper might become realistic. Conversely, an observation of the LFV modes $V \rightarrow \ell_\alpha \ell_\beta$, with branching fractions significantly larger than the bounds presented in Tab. 1 would be a way to disfavor many of the models containing heavy sterile neutrinos as being the unique source of lepton flavor violation. In obtaining the bounds presented in Tab. 1 we used masses and decay constants listed in Appendix B.

Finally, we compare in Tab. 2 our upper bounds for the modes for which we could find predictions in the literature.

| $V$ | $\ell_\alpha \ell_\beta$ | $m_4 = 1$ TeV | 10 TeV | 100 TeV | $m_5 = 1$ TeV | 10 TeV | 100 TeV |
|-----|--------------------------|---------------|--------|---------|---------------|--------|---------|
| $\phi$ | $e \mu$ | $1 \times 10^{-24}$ | $5 \times 10^{-24}$ | $3 \times 10^{-24}$ | $1 \times 10^{-23}$ | $6 \times 10^{-23}$ | $5 \times 10^{-23}$ |
| $\psi(2S)$ | $e \mu$ | $9 \times 10^{-22}$ | $1.5 \times 10^{-20}$ | $3 \times 10^{-21}$ | $4 \times 10^{-21}$ | $3 \times 10^{-20}$ | $2 \times 10^{-20}$ |
| $\Upsilon$ | $e \mu$ | $7 \times 10^{-18}$ | $2 \times 10^{-17}$ | $6 \times 10^{-18}$ | $2 \times 10^{-19}$ | $2 \times 10^{-17}$ | $2 \times 10^{-17}$ |
| $\Upsilon(2S)$ | $e \mu$ | $5 \times 10^{-18}$ | $5 \times 10^{-18}$ | $1.5 \times 10^{-18}$ | $2 \times 10^{-19}$ | $2 \times 10^{-17}$ | $2 \times 10^{-17}$ |
| $\Upsilon(3S)$ | $e \mu$ | $1.5 \times 10^{-17}$ | $3 \times 10^{-17}$ | $1.5 \times 10^{-17}$ | $5 \times 10^{-19}$ | $5 \times 10^{-17}$ | $4 \times 10^{-17}$ |
| $Z$ | $e \mu$ | $1.2 \times 10^{-14}$ | $7 \times 10^{-13}$ | $4 \times 10^{-13}$ | $9 \times 10^{-14}$ | $8 \times 10^{-13}$ | $6 \times 10^{-13}$ |

Table 1: Upper bound on $\text{B}(V \rightarrow \ell_\alpha \ell_\beta)$ for three values of the mass $m_{4,5}$. The numbers in three column referring to $m_4$ are obtained by using the effective model discussed in the text, while the other three, referring to $m_5$, are results of the (2,3)-ISS model (also discussed in the text).
Table 2: Upper bounds $B(V \to \ell_\alpha \ell_\beta)$: Comparison of the results reported in the literature with the bounds obtained in this work by using two different models (last two columns). The bounds for other similar decay modes that have not been discussed in the literature can be found in Tab. 1.

| Mode        | Ref. 37 | Ref. 38 | Ref. 39 | Eff.-model | (2,3)-ISS |
|-------------|---------|---------|---------|------------|-----------|
| $B(\phi \to e\mu)$ | $< 4 \times 10^{-17}$ | $< 1.3 \times 10^{-21}$ | $< 5 \times 10^{-20}$ | $< 5 \times 10^{-24}$ | $< 6 \times 10^{-23}$ |
| $B(J/\psi \to e\mu)$ | $< 4 \times 10^{-13}$ | $< 3.5 \times 10^{-13}$ | $< 1.9 \times 10^{-18}$ | $< 3 \times 10^{-20}$ | $< 9 \times 10^{-20}$ |
| $B(J/\psi \to \mu\tau)$ | - | - | $< 1.6 \times 10^{-7}$ | $< 6 \times 10^{-16}$ | $< 4 \times 10^{-18}$ |
| $B(\Upsilon \to e\mu)$ | $< 2 \times 10^{-9}$ | $< 3.8 \times 10^{-6}$ | $< 3.6 \times 10^{-18}$ | $< 2 \times 10^{-17}$ | $< 2 \times 10^{-17}$ |
| $B(\Upsilon \to \mu\tau)$ | - | - | $< 5.3 \times 10^{-7}$ | $< 2.5 \times 10^{-13}$ | $< 8 \times 10^{-16}$ |
| $B(Z \to e\mu)$ | $< 5 \times 10^{-13}$ | $< 8 \times 10^{-15}$ | - | $< 7 \times 10^{-13}$ | $< 8 \times 10^{-13}$ |

5 Conclusions

In this paper we discussed the enhancement of the LFV decays of flavorless vector bosons, $V \to \ell_\alpha \ell_\beta$, with $V \in \{\phi, \psi^{(n)}, \Upsilon^{(n)}, Z\}$, induced by mixing between the active and sterile neutrinos. The enhancement grows with mass of the heavy sterile neutrino(s), as can be seen from the mass dependence of Wilson coefficients that we explicitly calculated. We find that the most significant diagram that gives rise to the LFV decay amplitudes is the one coming from the $Z\nu\nu$-vertex, which suggests a steady growth of the decay rate with the mass of sterile neutrino(s). In the physical amplitude, however, the region of very large mass of the sterile neutrino(s) is suppressed as the decoupling takes place, i.e. mixing between the active and sterile neutrinos rapidly falls.

We illustrated the enhancement of $B(V \to \ell_\alpha \ell_\beta)$ in two scenarios: a model with one effective sterile neutrino that mimics the effect of a generic extensions of the SM including heavy sterile fermions, and in a minimal realization of the inverse seesaw scenario compatible with current observations. Our results for upper bounds on $B(V \to \ell_\alpha \ell_\beta)$ [$V \in \{\phi, J/\psi, \psi(2S), \Upsilon(1S), \Upsilon(2S), \Upsilon(3S), Z\}$] are still considerably smaller than the current experimental bounds (when available), but that situation might change in the future as more experimental research will be conducted at Belle II, BESIII, LHC, and hopefully at FCC-ee (TLEP). If one of the decays studied here is observed and turns out to have a branching fraction larger than the upper bounds reported here, then sources of LFV other than those coming from mixing with heavy sterile neutrinos must be accounted for.

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Appendix A: Wilson Coefficients

In this appendix we present the detailed expressions for Wilson coefficients. All computations have been made in the Feynman gauge. Contributions coming from the penguin and self energy diagrams are shown in Fig. 6 whereas the box diagrams are shown in Fig. 7.

We use the standard notation, \( x_i = m_i^2/m_W^2 \), \( x_t = m_t^2/m_W^2 \), \( x_q = q^2/m_W^2 = m_V^2/m_W^2 \), and write

\[
C_{rVL}^i = \sum_{i,j=1}^{n_\nu} U_{3i}^* U_{\alpha j} C_{r,ij}^{VL}(x_i, x_j),
\]

(25)

where \( r \in \{ \gamma, Z, \text{box} \} \). The coefficients \( C_{r,ij}^{VL} \) related to \( \gamma \) and box contributions are diagonal, \( C_{r,ij}^{VL} = \delta_{ij} C_{r,i}^{VL} \), while those related to the \( Z \) penguins can also involve a coupling to two different neutrinos, since the \( 3 \times 3 \) mixing matrix is no longer unitary. We therefore separate the diagonal and non-diagonal parts of the corresponding coefficient

\[
C_{Z,i}^{VL}(x_i) = \delta_{ij} C_{Z,i}^{VL} + \tilde{C}_{Z,ij}^{VL},
\]

(26)

where the second term depends on the parameter \( C_{ij} \) defined by

\[
C_{ij} = \sum_{\alpha=e,\mu,\tau} U_{\alpha i}^* U_{\alpha j},
\]

which, in the presence of sterile neutrinos, is generally different from \( \delta_{ij} \). Furthermore, from the plots presented in the body of the present paper we see that the region of \( m_4,5 \gg m_V \) is particularly interesting because there occurs the enhancement of the LFV decay rate. For the sake of clarity we thus expand our expressions in \( x_q \) and present here only the dominant terms. We also neglected, in the denominators of the loop integrals, the external momenta since they are negligible with respect to heavy neutrino masses. Therefore, up to terms \( O(x_q^2) \), our results read:

\[
C_{\gamma,i}^{VL}(x_i) = -\frac{1}{16\pi^2} + x_q -\frac{43x_i^3 + 108x_i^2 + 6(5x_i - 6)x_i^2 \log x_i - 81x_i + 16}{288\pi^2(x_i - 1)^4},
\]

(27)

\[
C_{Z,i}^{VL}(x_i) = -\frac{1}{64\pi^2(x_i - 1)^2} + \cos^2 \theta_W C_{\gamma,i}^{VL}(x_i),
\]

(28)

\[
\tilde{C}_{Z,ii}^{VL}(x_i, x_i) = \frac{x_i - 2}{128\pi^2(x_i - 1)^2} \left[ 3(x_i^2 - 1) + 2(x_i - 4)x_i \log x_i \right]
\]

\[
- x_q \left[ x_i - 1 \right] \left[ x_i(2x_i - 47) + 25 \right] + 14 + 6(x_i(12x_i - 13) + 2) \log x_i,
\]

(29)

\[
\tilde{C}_{Z,ij}^{VL}(x_i, x_j) = \frac{\delta_{ij}}{1152\pi^2(x_i - 1)^4},
\]

(30)
Figure 6: Penguin and self-energy diagrams contributing the LFV decay in Feynman gauge.
Figure 7: Box diagrams contributing the LFV decay $\Upsilon^{(n)} \to \ell_\alpha \ell_\beta$ in Feynman gauge.

\[
\hat{C}^{Z,ij}_{VL}(x_i, x_j) = \frac{\sqrt{x_i x_j} C_{ij}^*}{64 \pi^2} \left[ \frac{x_i (x_i - 4)}{(x_i - 1)(x_i - x_j)} \log x_i + \frac{x_j (x_j - 4)}{(x_j - 1)(x_j - x_i)} \log x_j - \frac{3}{2} \right] \\
+ \frac{C_{ij}}{64 \pi^2} \left[ \frac{2 x_i^2 (x_j - 1)}{(x_i - 1)(x_i - x_j)} \log x_i + \frac{2 x_j^2 (x_i - 1)}{(x_j - 1)(x_j - x_i)} \log x_j + 3 \right] \\
+ \frac{1}{192 \pi^2} \left\{ \sqrt{x_i x_j} C_{ij}^* \left[ \frac{x_i^2 (x_i - 3 x_j + 2 x_i x_j)}{(x_i - 1)^2(x_i - x_j)^3} \log x_i + \frac{x_j^2 (x_j - 3 x_i + 2 x_i x_j)}{(x_j - 1)^2(x_j - x_i)^3} \log x_j \right] \\
- \frac{x_i^3 (x_j - 1) - x_j^3 (x_i - 1) + x_i x_j (x_i - x_j)}{(x_i - 1)(x_j - 1)(x_i - x_j)^3} + 2 C_{ij} \left[ \frac{x_i^7 (3 x_i^2 + 3 x_j^2 + 3 x_j - x_i - 8 x_i x_j)}{(x_i - 1)^2(x_i - x_j)^3} \log x_i \right] \right\},
\]

\[
\hat{C}^{\text{Box},i}_{VL} = \frac{1}{256 \pi^2} \left\{ \frac{x_i (x_i - 8) + 4 x_i^2 \log x_i}{(x_i - 1)^2(x_i - x_i)} + \frac{x_t (x_t - 8) + 4 x_t^2 \log x_t}{(x_t - 1)^2(x_t - x_t)} + \frac{7 x_i x_t - 4}{(x_i - 1)(x_t - 1)} \right\}.
\]

Appendix B: Formulas and hadronic quantities

In this appendix we collect the expressions used to constrain the parameters of the models discussed in the present paper, as well as the values of the masses and decay constants used in our numerical analysis. In the expressions below we used the value of $G_F = \ldots$
\( G_\mu = 1.166 \times 10^{-5} \text{ GeV}^{-2} \) as extracted from \( \mu \to e\nu_\mu\bar{\nu}_e \). In our scenarios, in which we extended the neutrino sector by adding heavy sterile neutrinos, the Fermi constant becomes 

\[
G_F = G_\mu / \sqrt{\sum_{i,j} |U_{ei}|^2 |U_{ej}|^2}.
\]

For the models used in this paper, we checked that \( G_F = G_\mu \) remains an excellent approximation.

- \( \mu \to e\gamma \): We use the experimentally established upper bound \( B(\mu \to e\gamma) < 5.7 \times 10^{-13} \), and the expression \([23]\):

\[
B(\mu \to e\gamma) = \frac{\sqrt{2} G_F s_W^2 m_W^2}{128 \pi^5 \Gamma_\mu} m_\mu^5 |U_{\mu 4} U_{e 4} G_\gamma(x_4)|^2 ,
\]

\[
G_\gamma(x) = -\frac{2x^3 + 5x^2 - x}{4(1 - x)^3} - \frac{3x^3}{2(1 - x)^3} \log x ,
\]

(32)

to get one of the most significant constraints in this study. Notice that we use \( s_W^2 = 1 - m_W^2/m_Z^2 \), and we kept the dominant contribution with \( x_4 \).

- \( W \to \ell_\alpha \nu \): Combining the measured \( B(W \to e\nu) = 0.1071(16) \) and \( B(W \to \mu\nu) = 0.1063(15) \), with the expression

\[
B(W \to \ell_\alpha \nu) = \frac{\sqrt{2} G_F m_W}{24 \pi \Gamma_W} \sum_{j=1}^{4} \lambda(m_{\alpha j}^2, m_j^2, m_W^2) \begin{pmatrix} 2 - \frac{m_\alpha^2 + m_j^2}{m_W^2} - \frac{(m_\alpha^2 - m_j^2)^2}{m_W^4} \end{pmatrix} |U_{\alpha j}|^2 ,
\]

(33)

we further restrain the possible values of \( m_4 \) while varying the mixing angles in the largest possible range.

- \( \Delta r_{K,\pi} = R_{K,\pi}^{\exp.}/R_{K,\pi}^{\text{SM}} - 1 \): Ratio of the leptonic decay widths of a given meson \( P \),

\[
R_P = \Gamma(P \to e\nu_\ell) / \Gamma(P \to \mu\nu_\mu) ,
\]

was recently shown to be quite restrictive onto the possible values of \( m_{4,5} \) and \( \eta \) \([64]\). The most significant constraints actually come from \( \Delta r_\pi = 0.004(4) \) and \( \Delta r_K = 0.004(3) \), and the corresponding formula reads,

\[
\Delta r_P = -1 + \frac{m_{\pi}^2 (m_{\mu}^2 - m_{\pi}^2)^2}{m_\mu^2 (m_\pi^2 - m_\mu^2)^2} \sum_i |U_{\mu i}|^2 \begin{pmatrix} m_{\mu}^2 (m_{\pi}^2 + m_{\ell}^2) - (m_{\nu_i} - m_\mu^2)^2 \end{pmatrix} \lambda^1/2 (m_{\pi}^2, m_{\nu_i}^2, m_\mu^2) .
\]

(34)

- \( Z \to \nu\nu \): To saturate the experimental \( \Gamma(Z \to \text{invisible}) = 0.499(15) \text{ GeV} \), we sum over the kinematically available channels involving active and sterile neutrinos,

\[
\Gamma(Z \to \nu\nu) = \sum_{i,j} \left(1 - \frac{\delta_{ij}}{2}\right) \frac{G_F}{12 \sqrt{2} \pi m_Z} \lambda^{1/2} (m_Z^2, m_i^2, m_j^2) |C_{ij}|^2 \times \left[ 2m_Z^2 - m_i^2 - m_j^2 - 6m_i m_j - \frac{(m_i^2 - m_j^2)^2}{m_Z^2} \right] .
\]

(35)
• \( \mu \to eee \): We use the experimental upper bound \( B(\mu \to eee) < 1 \times 10^{-12} \), and the expression [23]:

\[
B(\mu \to eee) = \frac{G_F^4 m_W^4 m_\mu^5}{6144 \pi^7 \Gamma_\mu} \left\{ \frac{1}{2} F_{Box}^{\mu e e} + F_Z^{\mu e} - 2 \sin^2 \theta_W (F_Z^{\mu e} - F_\gamma^{\mu e}) \right\}^2 + 4 \sin^4 \theta_W |F_Z^{\mu e} - F_\gamma^{\mu e}|^2 \\
+ 16 \sin^2 \theta_W \text{Re} \left[ (F_Z^{\mu e} + \frac{1}{2} F_{Box}^{\mu e e}) G_\gamma^{\mu e s} \right] - 48 \sin^4 \theta_W \text{Re} \left[ (F_Z^{\mu e} - F_\gamma^{\mu e}) G_\gamma^{\mu e s} \right] \\
+ 32 \sin^4 \theta_W |G_\gamma^{\mu e}|^2 \left[ \ln \left( \frac{m_\mu^2}{m_e^2} - \frac{11}{4} \right) \right],
\]

with the loop functions \( F_{Box}^{\mu e e}, F_Z^{\mu e}, F_\gamma^{\mu e}, G_\gamma^{\mu e s} \) defined in [60].

Finally, the values of hadronic quantities, not discussed in the body of the paper but used in our numerical analysis, are listed in Tab. 3.

| quantity \( m_{\phi} \) | value (GeV) | Ref. | quantity \( f_{\phi} \) | value (MeV) | Ref. |
|---------------------|-------------|------|---------------------|-------------|------|
| \( m_{J/\psi} \)    | 3.0969      | [16] | \( f_{J/\psi} \)  | 418(9) MeV | [28] |
| \( m_{\psi(2S)} \)  | 3.6861      | [16] | \( f_{\psi(2S)}/f_{J/\psi} \) | 0.713(16) | [16] |
| \( m_{\Upsilon} \)  | 9.460       | [16] | \( f_{\Upsilon} \) | 649(31) MeV | [30] |
| \( m_{\Upsilon(2S)} \) | 10.023     | [16] | \( f_{\Upsilon(2S)} \) | 481(39) MeV | [30] |
| \( m_{\Upsilon(3S)} \) | 10.355     | [16] | \( f_{\Upsilon(3S)} \) | 539(84) MeV | [31] |

Table 3: Masses and decay constants used in numerical analysis.

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9 Notice that the ratio of decay constants \( f_{\psi(2S)}/f_{J/\psi} \) has been obtained from the corresponding (measured) electronic widths and the expression \( \Gamma(\psi_n \to e^+e^-) = 16\pi\alpha^2_{em} f_{\psi_n}^2 / (27 m_{\psi_n}^2) \).
References

[1] M. C. Gonzalez-García, M. Maltoni and T. Schwetz, JHEP 1411 (2014) 052 [arXiv:1409.5439 [hep-ph]], regularly updated at http://www.nu-fit.org/.

[2] J. Adam et al. [MEG Collaboration], Phys. Rev. Lett. 110 (2013) 201801 [arXiv:1303.0754 [hep-ex]].

[3] A. M. Baldini, F. Cei, C. Cerri, S. Dussoni, L. Galli, M. Grassi, D. Nicolo and F. Raffaelli et al., arXiv:1301.7225 [physics.ins-det].

[4] U. Bellgardt et al. [SINDRUM Collaboration], Nucl. Phys. B 299 (1988) 1.

[5] A. Blondel, A. Bravar, M. Pohl, S. Bachmann, N. Berger, M. Kiehn, A. Schoning and D. Wiedner et al., arXiv:1301.6113 [physics.ins-det].

[6] B. Aubert et al. [BaBar Collaboration], Phys. Rev. Lett. 104 (2010) 021802 [arXiv:0908.2381 [hep-ex]].

[7] K. Hayasaka et al., Phys. Lett. B 687 (2010) 139 [arXiv:1001.3221 [hep-ex]].

[8] T. Aushev et al., arXiv:1002.5012 [hep-ex].

[9] A. J. Bevan et al. [BaBar and Belle Collaborations], Eur. Phys. J. C 74 (2014) 11, 3026 [arXiv:1406.6311 [hep-ex]].

[10] C. Dohmen et al. [SINDRUM II. Collaboration], Phys. Lett. B 317 (1993) 631.

[11] A. Alekou et al., arXiv:1310.0804 [physics.acc-ph].

[12] W. H. Bertl et al. [SINDRUM II Collaboration], Eur. Phys. J. C 47 (2006) 337.

[13] Y. Kuno [COMET Collaboration], PTEP 2013 (2013) 022C01.

[14] V. Khachatryan et al. [CMS Collaboration], arXiv:1502.07400 [hep-ex].

[15] R. Aaij et al. [LHCb Collaboration], Phys. Lett. B 724 (2013) 36 [arXiv:1304.4518 [hep-ex]].

[16] K. A. Olive et al. [Particle Data Group Collaboration], Chin. Phys. C 38 (2014) 090001.

[17] R. Aaij et al. [LHCb Collaboration], Phys. Rev. Lett. 111 (2013) 14, 141801 [arXiv:1307.4889 [hep-ex]].

[18] P. Abreu et al. [DELPHI Collaboration], Z. Phys. C 73 (1997) 243.

[19] R. Akers et al. [OPAL Collaboration], Z. Phys. C 67 (1995) 555.

[20] O. Adriani et al. [L3 Collaboration], Phys. Lett. B 316 (1993) 427.
[21] G. Aad et al. [ATLAS Collaboration], Phys. Rev. D 90 (2014) 7, 072010 [arXiv:1408.5774 [hep-ex]].

[22] G. Mann and T. Riemann, Annalen Phys. 40 (1984) 334.

[23] A. Ilakovac and A. Pilaftsis, Nucl. Phys. B 437 (1995) 491 [hep-ph/9403398].

[24] J. I. Illana and T. Riemann, Phys. Rev. D 63 (2001) 053004 [hep-ph/0010193]; J. I. Illana, M. Jack and T. Riemann, In “2nd ECFA/DESY Study 1998-2001” 490-524 [hep-ph/0001273].

[25] A. Abada, V. De Romeri, S. Monteil, J. Orloff and A. M. Teixeira, arXiv:1412.6322 [hep-ph].

[26] A. Blondel, A. Chao, W. Chou, J. Gao, D. Schulte and K. Yokoya, arXiv:1302.3318 [physics.acc-ph].

[27] A. Blondel et al. [team for the FCC-ee study Collaboration], arXiv:1411.5230 [hep-ex].

[28] D. Becirevic, G. Duplancic B. Klajn, B. Melic and F. Sanfilippo, Nucl. Phys. B 883 (2014) 306 [arXiv:1312.2858 [hep-ph]]; D. Becirevic and F. Sanfilippo, JHEP 1301 (2013) 028 [arXiv:1206.1445 [hep-lat]].

[29] G. C. Donald, C. T. H. Davies, R. J. Dowdall, E. Follana, K. Hornbostel, J. Koponen, G. P. Lepage and C. McNeile, Phys. Rev. D 86 (2012) 094501 [arXiv:1208.2855 [hep-lat]].

[30] B. Colquhoun, R. J. Dowdall, C. T. H. Davies, K. Hornbostel and G. P. Lepage, arXiv:1408.5768 [hep-lat].

[31] R. Lewis and R. M. Woloshyn, Phys. Rev. D 85 (2012) 114509 [arXiv:1204.4675 [hep-lat]].

[32] M. N. Achasov et al., Phys. Rev. D 81 (2010) 057102 [arXiv:0911.1232 [hep-ex]].

[33] M. Ablikim et al. [BESIII Collaboration], Phys. Rev. D 87 (2013) 11, 112007 [arXiv:1304.3205 [hep-ex]].

[34] M. Ablikim et al. [BES Collaboration], Phys. Lett. B 598 (2004) 172 [hep-ex/0406018].

[35] W. Love et al. [CLEO Collaboration], Phys. Rev. Lett. 101 (2008) 201601 [arXiv:0807.2695 [hep-ex]].

[36] J. P. Lees et al. [BaBar Collaboration], Phys. Rev. Lett. 104 (2010) 151802 [arXiv:1001.1883 [hep-ex]].

[37] S. Nussinov, R. D. Peccei and X. M. Zhang, Phys. Rev. D 63 (2001) 016003 [hep-ph/0004153].

[38] T. Gutsche, J. C. Helo, S. Kovalenko and V. E. Lyubovitskij, Phys. Rev. D 81 (2010) 037702 [arXiv:0912.4562 [hep-ph]], ibid 83 (2011) 115015 [arXiv:1103.1317 [hep-ph]].
[39] K. S. Sun, T. F. Feng, T. J. Gao and S. M. Zhao, Nucl. Phys. B 865 (2012) 486 [arXiv:1208.2404 [hep-ph]].

[40] T. A. Mueller et al., Phys. Rev. C 83 (2011) 054615 [arXiv:1101.2663 [hep-ex]]; P. Huber, Phys. Rev. C 84 (2011) 024617 [Erratum-ibid. C 85 (2012) 029901] [arXiv:1106.0687 [hep-ph]]; G. Mention, M. Fechner, T. Lasserre, T. A. Mueller, D. Lhuillier, M. Cribier and A. Letourneau, Phys. Rev. D 83 (2011) 073006 [arXiv:1101.2755 [hep-ex]].

[41] A. A. Aguilar-Arevalo et al. [LSND Collaboration], Phys. Rev. D 64 (2001) 112007 [hep-ex/0104049].

[42] A. A. Aguilar-Arevalo et al. [MiniBooNE Collaboration], Phys. Rev. Lett. 98 (2007) 231801 [arXiv:0704.1500 [hep-ex]]; A. A. Aguilar-Arevalo et al. [MiniBooNE Collaboration], Phys. Rev. Lett. 105 (2010) 181801 [arXiv:1007.1150 [hep-ex]]; A. A. Aguilar-Arevalo et al. [MiniBooNE Collaboration], Phys. Rev. Lett. 110 (2013) 161801 [arXiv:1207.4809 [hep-ex], arXiv:1303.2588 [hep-ex]].

[43] M. A. Acero, C. Giunti and M. Laveder, Phys. Rev. D 78 (2008) 073009 [arXiv:0711.4222 [hep-ph]]. C. Giunti and M. Laveder, Phys. Rev. C 83 (2011) 065504 [arXiv:1006.3244 [hep-ph]].

[44] S. Dodelson and L. M. Widrow, Phys. Rev. Lett. 72 (1994) 17 [hep-ph/9303287].

[45] K. Abazajian, G. M. Fuller and M. Patel, Phys. Rev. D 64 (2001) 023501 [astro-ph/0101524].

[46] A. D. Dolgov and S. H. Hansen, Astropart. Phys. 16 (2002) 339 [hep-ph/0009083]; A. Boyarsky, J. Lesgourgues, O. Ruchayskii and M. Viel, Phys. Rev. Lett. 102 (2009) 201304 [arXiv:0812.3256 [hep-ph]]; A. Boyarsky, O. Ruchayskii and M. Shaposhnikov, Ann. Rev. Nucl. Part. Sci. 59 (2009) 191 [arXiv:0901.0011 [hep-ph]].

[47] A. A. Klypin, A. V. Kravtsov, O. Valenzuela and F. Prada, Astrophys. J. 522 (1999) 82 [astro-ph/9901240]; B. Moore, S. Ghigna, F. Governato, G. Lake, T. R. Quinn, J. Stadel and P. Tozzi, Astrophys. J. 524 (1999) L19 [astro-ph/9907411]; L. E. Strigari, C. S. Frenk and S. D. M. White, Mon. Not. Roy. Astron. Soc. 408 (2010) 2364 [arXiv:1003.4268 [astro-ph.CO]]; M. Boylan-Kolchin, J. S. Bullock and M. Kaplinghat, Mon. Not. Roy. Astron. Soc. 415 (2011) L40 [arXiv:1103.0007 [astro-ph.CO]].

[48] A. Kusenko, Phys. Rept. 481 (2009) 1 [arXiv:0906.2968 [hep-ph]].

[49] K. N. Abazajian, M. A. Acero, S. K. Agarwalla, A. A. Aguilar-Arevalo, C. H. Albright, S. Antusch, C. A. Arguelles and A. B. Balantekin et al., “Light Sterile Neutrinos: A White Paper,” arXiv:1204.5379 [hep-ph].

[50] A. Abada and M. Lucente, Nucl. Phys. B 885 (2014) 651 [arXiv:1401.1507 [hep-ph]].

[51] P. A. R. Ade et al. [Planck Collaboration], Astron. Astrophys. 571 (2014) A16 [arXiv:1303.5076 [astro-ph.CO]].
[52] M. S. Chanowitz, M. A. Furman and I. Hinchcliffe, Nucl. Phys. B 153 (1979) 402.

[53] A. Ilakovac, Phys. Rev. D 62 (2000) 036010 [hep-ph/9910213].

[54] R. N. Mohapatra and J. W. F. Valle, Phys. Rev. D 34 (1986) 1642; M. C. Gonzalez-Garcia and J. W. F. Valle, Phys. Lett. B 216 (1989) 360; F. Deppisch and J. W. F. Valle, Phys. Rev. D 72 (2005) 036001 [hep-ph/0406040].

[55] P. S. B. Dev and A. Pilaftsis, Phys. Rev. D 86 (2012) 113001 [arXiv:1209.4051 [hep-ph]].

[56] A. Abada, G. Arcadi and M. Lucente, JCAP10(2014)001 [arXiv:1406.6556 [hep-ph]].

[57] J. Schechter and J. W. F. Valle, Phys. Rev. D 22 (1980) 2227.

[58] M. Gronau, C. N. Leung and J. L. Rosner, Phys. Rev. D 29 (1984) 2539.

[59] S. Antusch and O. Fischer, JHEP 1410 (2014) 94 [arXiv:1407.6607 [hep-ph]].

[60] R. Alonso, M. Dhen, M. B. Gavela and T. Hambye, JHEP 1301 (2013) 118 [arXiv:1209.2679 [hep-ph]].

[61] S. Antusch, C. Biggio, E. Fernandez-Martinez, M. B. Gavela and J. Lopez-Pavon, JHEP 0610 (2006) 084 [hep-ph/0607020]; E. Fernandez-Martinez, M. B. Gavela, J. Lopez-Pavon and O. Yasuda, Phys. Lett. B 649 (2007) 427 [hep-ph/0703098]; S. Antusch, J. P. Baumann and E. Fernandez-Martinez, Nucl. Phys. B 810 (2009) 369 [arXiv:0807.1003 [hep-ph]].

[62] F. Deppisch, T. S. Kosmas and J. W. F. Valle, Nucl. Phys. B 752 (2006) 80 [hep-ph/0512360]; D. N. Dinh, A. Ibarra, E. Molinaro and S. T. Petcov, JHEP 1208 (2012) 125 [Erratum-ibid. 1309 (2013) 023] [arXiv:1205.4671 [hep-ph]].

[63] A. Atre, T. Han, S. Pascoli and B. Zhang, JHEP 0905 (2009) 030 [arXiv:0901.3589 [hep-ph]].

[64] A. Abada, A. M. Teixeira, A. Vicente and C. Weiland, JHEP 1402 (2014) 091 [arXiv:1311.2830 [hep-ph]].

[65] G. C. Donald et al. [HPQCD Collaboration], Phys. Rev. D 90 (2014) 7, 074506 [arXiv:1311.6669 [hep-lat]].