A Gauge and Lorentz covariant Approximation for the Quark Propagator in an arbitrary Gluon Field

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Abstract: We decompose the quark propagator in the presence of an arbitrary gluon field with respect to a set of Dirac matrices. The four-dimensional integrals which arise in first order perturbation theory are rewritten as line-integrals along certain field lines, together with a weighted integration over the various field lines. It is then easy to transform the propagator into a form involving path ordered exponentials. The resulting expression is non-perturbative and has the correct behavior under Lorentz transformations, gauge transformations and charge conjugation. Furthermore it coincides with the exact propagator in first order of the coupling $g$. No expansion with respect to the inverse quark mass is involved, the expression can even be used for vanishing mass. For large mass the field lines concentrate near the straight line connection and simple results can be obtained immediately.
1 The quark propagator

The quark propagator $S(x, y; A)$ for a quark of mass $m$ in the presence of a gluon field $A^\mu$ plays an important role in many investigations of quantum chromodynamics. It appears, e.g., if one considers the quark four-point Green function, the basis of all modern investigations on quark-antiquark interactions, and integrates over the quark fields. It is defined by

$$S_{\alpha\beta}(x, y; A) = -i < 0|T(\psi_\alpha(x)\bar{\psi}_\beta(y))|0 > .$$

(1.1)

The spinor indices $\alpha, \beta$ will be dropped in the following. We recall the relevant properties of the propagator. With the covariant derivative $D_\mu(x)$ which acts on operators to the right, and $\bar{D}_\mu(y)$ which acts on operators to the left, defined by

$$D_\mu(x) = \frac{\partial}{\partial x^\mu} - igA_\mu(x),$$
$$\bar{D}_\mu(y) = \frac{\partial}{\partial y^\mu} + igA_\mu(y),$$

(1.2)

the field equations give

$$[i\gamma^\mu D_\mu(x) - m]S(x, y; A) = \delta^{(4)}(x - y),$$
$$S(x, y; A)[-i\gamma^\mu \bar{D}_\mu(y) - m] = \delta^{(4)}(x - y).$$

(1.3)

These equations may be reformulated as integral equations:

$$S(x, y; A) = S_0(x - y) - g \int S(x, z; A)\gamma^\mu A_\mu(z)S_0(z - y)d^4z = S_0(x - y) - g \int S_0(x - z)\gamma^\mu A_\mu(z)S(z, y; A)d^4z,$$

(1.4)

with $S_0$ the free propagator. From charge conjugation one has

$$S(x, y; A) = \gamma^2\gamma^0 S^T(y, x; - A^T)\gamma^2\gamma^0.$$

(1.5)

Finally, under a gauge transformation $\psi \rightarrow \psi' = e^{i\Theta}\psi, \bar{\psi} \rightarrow \bar{\psi}' = \bar{\psi}e^{-i\Theta}, A_\mu \rightarrow A'_\mu = e^{i\Theta}A_\mu e^{-i\Theta} - (i/g)(\partial_\mu e^{i\Theta})e^{-i\Theta}$, the propagator transforms as

$$S(x, y; A) \rightarrow S'(x, y; A) = e^{i\Theta(x)}S(x, y; A)e^{-i\Theta(y)}.$$

(1.6)

An exact solution for $S(x, y; A)$ for an arbitrary gluon field $A_\mu$ is not available. However, it would be highly desirable to have an approximation which respects the fundamental properties of the propagator, in particular the correct transformation under Lorentz transformations and under gauge transformations.

The two well known approximations, perturbation theory and the static approximation, either violate gauge covariance or Lorentz covariance:

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1Our formulae are given in Minkowski space, we use Bjorken Drell conventions [2], and the field tensor is defined by $F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu - ig[A_\mu, A_\nu]$. 

1
**Perturbation theory:** Iteration of eq. (1.4) gives the perturbation series

\[ S(x, y; A) = S_0(x - y) - g \int S_0(x - z) \gamma^\mu A_\mu(z) S_0(z - y) dz + \cdots. \]  

(1.7)

Any finite order of the perturbation series gives the correct behavior under Lorentz transformations and under charge conjugation. It will, however, never be able to describe central features of QCD like confinement. In particular, it clearly violates gauge invariance; no truncation of the perturbation series (1.7) has the correct transformation property (1.6).

**Static approximation:** Following the pioneering work of Brown, Weisberger [2] and Eichten, Feinberg [3], one neglects the spatial part \( i \gamma^m D_m(x) \) in eq. (1.3). The partial differential equation then becomes an ordinary differential equation which can be solved in closed form. This leads to the static approximation

\[ S_{\text{stat}}(x, y; A) = -i \left\{ \Theta(x_0 - y_0) \frac{1 + \gamma^0}{2} + \Theta(y_0 - x_0) \frac{1 - \gamma^0}{2} \right\} \delta(3)(x - y) e^{-im|x_0 - y_0|} P \exp\{ig \int_y^x A_0(z) dz\}. \]  

(1.8)

The path in the line integral is the straight line from \( y \) to \( x \), and the path ordering orders \( A_0(x) \) to the left, \( \cdots \), \( A_0(y) \) to the right.

The neglected spatial term \( i \gamma^m D_m(x) \) in (1.3) can subsequently be taken into account as perturbation. This approach has been extremely successful (see e.g. the review [4]). It can be easily generalized to quarks moving with four velocity \( v^\mu \), and thus is the direct predecessor of heavy quark effective theory. Successful combination with perturbation theory has also been made more recently (see [5] and references therein). For quarks moving with high momentum a related formula can be derived from the eikonal approximation [6].

The static approximation and its generalizations are non-perturbative and have the correct behavior under gauge transformations and under charge conjugation. However, they drastically violate Lorentz invariance. Therefore the static approximation is useful for heavy quarks only. It needs quite an effort to recover the relations following from the original Lorentz invariance subsequently [1, 7]. Finally the static approximation does not coincide with the exact propagator even in the trivial case of vanishing gluon field.

It is obvious that a gauge covariant propagator should contain path ordered exponentials as in (1.8). Using just the path along the straight line between \( x \) and \( y \) would give a Lorentz covariant result, but such a procedure would be far too simple. It would involve the vector potential only along the straight line connection and nowhere else, which is clearly unphysical. We prefer to proceed systematically by rewriting perturbation theory in a suitable way. It can then easily be transformed into a gauge covariant expression by simple exponentiation, while keeping the correctness of perturbation theory in the
relevant order. The non-perturbative approximation for the quark propagator obtained in this way has the following properties:

- Correct behavior under Lorentz transformations.
- Correct behavior under gauge transformations.
- Correct behavior under charge conjugation.
- Agreement with perturbation theory in first order of the coupling.

The representation is a weighted superposition of path ordered exponentials between \( x \) and \( y \) along well defined field lines. We don’t need to assume that the quark mass is large, we could even put it equal to zero. This opens perspectives to applications which were hard to attack previously.

The paper is organized as follows:

In sect. 2 we decompose the propagator with respect to Dirac matrices, and write the formula of first order perturbation theory in a form which is convenient for the following. In sect. 3 we rewrite the four-dimensional space-time integrals which arise in perturbation theory as weighted superpositions of line integrals over certain field lines which all run from \( x \) to \( y \). From this form one can simply derive a representation in terms of superpositions of path ordered exponentials. This representation coincides with perturbation theory up to order \( g \) and has the correct behavior under gauge transformations. In sect. 4 we evaluate the weight function explicitly and show a plot of the field lines for different masses. We discuss the limit of large mass \( m \), and give some first simple applications. Actual applications will be given in forthcoming papers.

## 2 A useful form of first order perturbation theory

We start with the first order approximation (1.7), and express the free propagators \( S_0 \) by the free scalar propagator \( \Delta \) in the following way:

\[
S_0(x-z) = \Delta(x-z)[-i\gamma_\nu \frac{\partial}{\partial z_\nu} + m], \\
S_0(z-y) = [i\gamma_\lambda \frac{\partial}{\partial z_\lambda} + m]\Delta(z-y).
\]  

(2.1)

The free scalar propagator satisfies the equation

\[
(\partial_\mu \partial^\mu + m^2)\Delta(x) = -\delta^{(4)}(x).
\]  

(2.2)

In the following all differential operators in the integrand are understood as differentiations with respect to the variable \( z \).

Using the well known identities \( \gamma_\nu \gamma_\mu = g_{\nu \mu} - i\sigma_{\nu \mu} \), and \( \gamma_\nu \gamma_\mu \gamma_\lambda = g_{\nu \mu} \gamma_\lambda - g_{\nu \lambda} \gamma_\mu + g_{\mu \lambda} \gamma_\nu + i\epsilon_{\nu \mu \lambda \kappa} \gamma^5 \gamma^\kappa \), one can write the propagator in form of the familiar decomposition

\[
S(x, y; A) = s + p\gamma^5 + v^\mu \gamma_\mu + a^\mu \gamma^5 \gamma_\mu + t^{\mu \nu} \sigma_{\mu \nu}.
\]  

(2.3)
Before giving the expressions for \(s, \ldots, t^\mu{}^\nu\) which result in this way, it is convenient for later use to define a scalar field \(u(z; x, y)\) and a vector field \(u^\mu(z; x, y)\) by

\[
u(z; x, y) = \Delta(x - z)\Delta(z - y), \tag{2.4}\]

and (with \(\vec{\partial} = \vec{\partial}_z / \partial z_\mu - \vec{\partial}_y / \partial z_\mu\))

\[
u^\mu(z; x, y) = -\Delta^{-1}(x - y)[\Delta(x - z) \vec{\partial} \Delta(z - y)]. \tag{2.5}\]

From (1.7), (2.1), and (2.3) we then obtain the following equations:

\[
s = m\Delta(x - y)[1 + ig \int u^\nu(z; x, y)A_\nu(z) d^4z], \tag{2.6}\]
\[
p = 0, \tag{2.7}\]
\[
v^\mu = \frac{i}{2} \vec{\partial} / \partial x_\mu \Delta(x - y) - \frac{i}{2} \Delta(x - y) \vec{\partial} / \partial y_\mu
- g \int \left\{ \Delta(x - z)[m^2 g^\nu{}^\mu + \vec{\partial} \vec{\partial}^\nu - \vec{\partial}_\nu g^\nu{}^\lambda \vec{\partial}^\lambda + \vec{\partial}^\nu \vec{\partial}^\lambda] \Delta(z - y) \right\} A_\nu(z) d^4z, \tag{2.8}\]
\[
a^\mu = i g \epsilon^{\mu\nu\lambda\kappa} \int [\Delta(x - z) \vec{\partial}_\nu \vec{\partial}_\lambda \Delta(z - y)] A_\nu(z) d^4z, \tag{2.9}\]
\[
t^\mu = - \frac{gm}{2} \int \{\partial_\nu u(z; x, y) A_\mu(z) - (\mu \leftrightarrow \nu)\} d^4z. \tag{2.10}\]

We next transform \(v^\mu\) in (2.8), we denote the four terms in the second line by \(v_1^\mu + v_2^\mu + v_3^\mu + v_4^\mu\). In \(v_2^\mu\) one can use \(\vec{\partial}_z = - \vec{\partial}_y\), in \(v_4^\mu\) correspondingly \(\vec{\partial}_z = - \vec{\partial}_x\). The differentiations with respect to \(y\) and \(x\) can then be taken outside of the integral. The term \(v_1^\mu\) is split into two identical contributions, in the first one we use eq. (2.2) for \(\Delta(x - z)\), in the second one for \(\Delta(z - y)\). Next perform a partial integration on one of the derivatives in the d’Alembert operator. The terms where the differentiations act on the other \(\Delta\) cancel against the contribution \(v_3^\mu\), and one has the intermediate result

\[
v_1^\mu + v_3^\mu = - \frac{g}{2} \Delta(x - y)\{A^\mu(x) + A^\mu(y)\}
- \frac{g}{2} \int \partial_\lambda u(z; x, y) [\partial^\lambda A^\mu(z) - \partial^\mu A^\lambda(z)] d^4z
- \frac{g}{2} \int \partial_\lambda u(z; x, y) \partial^\mu A^\lambda(z) d^4z. \tag{2.11}\]

We subtraced and added the term \(\partial^\mu A^\lambda(z)\). In the second line we can then replace \(\partial^\lambda A^\mu(z) - \partial^\mu A^\lambda(z)\) by \(F^\mu{}^\nu(z)\), which only introduces a higher order error \(O(g^2)\). Furthermore the differential operator \(\partial_\lambda = \partial_\lambda^z\) which acts on \(u(z; x, y)\) can be replaced by \(- (\partial_\lambda^x + \partial_\lambda^y)\) and taken outside of the integral. In the third line we perform a partial integration on \(\partial^\mu\), shift the differentiations from the variable \(z\) to \(x\) and \(y\), and take the differentiations out of the integral. After these manipulations \(v^\mu\) can be expressed in a rather compact form if we replace partial derivatives by covariant derivatives which introduces a correction of order \(g^2\) only:
\[ v^\mu = \frac{i}{2} D^\mu(x) \left\{ \Delta(x-y) [1 + ig \int u^\nu(z;x,y) A_\nu(z) d^4z] \right\} - \frac{i}{2} \left\{ \Delta(x-y) [1 + ig \int u^\nu(z;x,y) A_\nu(z) d^4z] \right\} \frac{\tilde{D}^*}{D} \mu(y) \]
\[ - \frac{g}{2} D_\nu(x) \int u(z;x,y) F^\mu\nu(z) d^4z - \frac{g}{2} \int u(z;x,y) F^\mu\nu(z) d^4z \frac{\tilde{D}^*}{D} \nu(y). \] (2.12)

In \( a^\mu \) we perform a partial integration with respect to \( \tilde{\partial}_\kappa \) or \( \tilde{\partial}_\lambda \), take half of the sum of both terms, antisymmetrize the \( \tilde{\partial}_\kappa A_\nu(z) \) or \( \tilde{\partial}_\lambda A_\nu(z) \) in the integrand, and introduce the field strength tensor \( F^\kappa\nu(z) \) or \( F^\lambda\nu(z) \) as before. Shifting again the differentiation from \( z \) to \( x \) and \( y \) we have in order

\[ a^\mu = \frac{ig}{4} \epsilon^{\mu\lambda\kappa} \left\{ D_\lambda(x) \int u(z;x,y) F_{\kappa\nu}(z) d^4z - \int u(z;x,y) F_{\kappa\nu}(z) d^4z \frac{\tilde{D}^*}{D} \lambda(y) \right\}. \] (2.13)

Similar manipulations applied to \( t^\mu\nu \) finally give

\[ t^\mu\nu = -\frac{gm}{2} \int u(z;x,y) F^\mu\nu(z) d^4z. \] (2.14)

The decomposition (2.3), together with the formulae (2.6), (2.7), (2.12), (2.13), (2.14) is now in a form which allows to rewrite the four-dimensional integrations \( d^4z \) as a superposition of line integrals.

### 3 Gauge covariant reformulation

The vector field \( u^\mu(z;x,y) \) defined in (2.5) satisfies the fundamental equation

\[ \frac{\partial u^\mu(z;x,y)}{\partial z^\mu} = \delta^{(4)}(z-y) - \delta^{(4)}(z-x), \] (3.1)

which is a simple consequence of eq. (2.2). Therefore \( u^\mu \) may be interpreted as a four-dimensional velocity field of an incompressible fluid with a point-like source at \( y \) and a sink at \( x \). The stream lines \( z^\mu(s;w) \), which all run from \( y \) to \( x \), are defined by the characteristic equations

\[ \frac{dz^\mu(s,w)}{ds} = u^\mu(z(s,w)). \] (3.2)

Here \( s \) is the parameter which describes the motion along the stream line, while the three dimensional parameter set \( w \) characterizes the various stream lines. To make \( s \) unique, it is convenient to fix \( s = 0 \) at the symmetrical point of the stream line which has equal distance to \( x \) and \( y \). The dependence on \( x, y \) has been suppressed in the notation. There is precisely one field line passing through every space time point, except, of course, for the source points \( x \) and \( y \). After having solved (3.2) there is a unique correspondence between the 4-dimensional space-time coordinates \( z^\mu \) and the parameters \( s,w \), i.e. \( (z^0, z^1, z^2, z^3) \leftrightarrow (s, w^1, w^2, w^3) \).
We next write the four-dimensional integrals over $d^4z$ which appear in $s$ and the first two lines of $v^\mu$ (eqs (2.6) and (2.12)) as integrals over $dsd^3w$, the Jacobian is called $\rho(w)$:

$$
\rho(w) = \frac{\partial(z^0, z^1, z^2, z^3)}{\partial(s, w^1, w^2, w^3)}.
$$

(3.3)

We have anticipated the crucial fact that $\rho$ does not depend on the curve parameter $s$. This is a direct consequence of the incompressibility of the flow, and easily proved from the following geometrical argument. Take an infinitesimal four-dimensional box in $(s, w)$-space with corners $(s_0, w)$ and $(s_0 + ds, w + dw)$. In $z$-space this corresponds to an infinitesimal region with a certain volume. Consider now the motion of this volume along a field line from $s_0$ to $s_1$, keeping $w, dw$ and $ds$ constant. Because of the vanishing divergence (3.1) outside the sources, the volume in $z$-space stays constant, the volume in $(s, w)$-space stays constant anyhow by construction. This demonstrates that the Jacobian (3.3) is indeed independent of $s$.

We can therefore write

$$
\int u^\mu(z; x, y)A_\mu(z)d^4z = \int \rho(w)[\int_y^x A_\mu(z(s, w))u^\mu(z(s, w))ds]d^3w = \int \rho(w)[\int_y^x Adz]d^3w.
$$

(3.4)

Here $\int_y^x Adz$ is a shorthand notation for the line integral of $A_\mu$ from $y$ to $x$ along the stream line characterized by the parameter $w$.

The normalization of $\rho(w)$ is easily obtained from the special case $A_\mu(z) = \partial_\mu \Theta(z)$, where both sides of (3.4) can be immediately integrated. This gives

$$
\int \rho(w)d^3w = 1.
$$

(3.5)

We are now in the position to rewrite the scalar function $s$ in (2.6):

$$
\begin{align*}
\int u^\mu(z; x, y)A_\mu(z)d^4z &= \int \rho(w)[\int_y^x A_\mu(z(s, w))u^\mu(z(s, w))ds]d^3w \\
&= \int \rho(w)[\int_y^x Adz]d^3w.
\end{align*}
$$

(3.6. a)

$$
\begin{align*}
\int \rho(w)d^3w &= 1. \\
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\int \rho(w)d^3w &= 1.
\end{align*}
$$

(3.5)

We are now in the position to rewrite the scalar function $s$ in (2.6):

$$
\begin{align*}
s &= m\Delta(x - y)\int \rho(w)[1 + ig\int_y^x Adz]d^3w \\
&= m\Delta(x - y)\int \rho(w)P \exp[ig\int_y^x Adz]d^3w + O(g^2).
\end{align*}
$$

(3.6. b)

This was the essential step of the approach! Once having written the first order approximation in terms of line integrals, exponentiation allows to promote it into a non-perturbative gauge covariant expression.

One may wonder about the justification of this step which led to a non-perturbative expression simply by exponentiation. It does not make sense to compare the two lines in (3.6), because (3.6.b) is gauge covariant while (3.6.a) is not. In fact one can easily see, e.g. for a suitable pure gauge, that (3.6.a) can be made as large as one likes. The correct question to be asked is whether (3.6.b) is a reasonable approximation to the exact propagator. To answer this question one only needs to check the quality of the approximation in some special convenient gauge. The manifest covariance of the expression then guarantees this quality for any gauge. A sufficient condition is the
smallness of $g \mathbf{W}_\mu A dz$ for all relevant stream lines, which justifies the validity of first order perturbation theory. Clearly one can always find a gauge where $A_\mu dz^\mu = 0$ on one special stream line, but it is not possible to do the same for two or more lines. For heavy quarks the situation is simple. We will show in sect. 4 that in this case all relevant paths lie near the straight line connection. If one chooses a gauge such that $A_\mu(x-y)^\mu$ vanishes along this line, and if the variations of $A_\mu$ near this line are small, the approximation will be justified. In applications we have to perform an integration over the gauge field $A_\mu$ at the end, which involves the whole spectrum in momentum space. Because there are only two scales available, $m$ and $\Lambda_{QCD}$, one expects a good approximation if $m >> \Lambda_{QCD}$.

For light quarks the lines spread out over the whole space, and the above argument cannot be applied. But in any case the step of exponentiation leading from (3.6.a) to (3.6.b) can be interpreted as partially including higher orders of the perturbation series, namely a minimum of those necessary to guarantee gauge covariance. There are good reasons to believe that gauge covariance is such a fundamental principle that it may indeed be used to transform perturbative expressions into non-perturbative ones of physical relevance.

The expressions in the first and second line of $v^\mu$ can be treated in exactly the same way. The integrals in the third line of $v^\mu$, as well as those in $a^\mu$ and $t^\mu$ have a different form, but the structure of all of them is identical. They have a factor $g$ in front, therefore one can introduce path ordered exponentials without changing the result in order $g$. To each $z \neq x, y$ there belongs a unique $s'$ and $w$ which characterize it’s position $s'$ on the field line $w$. One can write

$$ g F^{\mu\nu}(z(s', w)) = P \left\{ g F^{\mu\nu}(z(s', w)) \exp [ig \int_{y}^{x} A dz] \right\} + O(g^2). \quad (3.7) $$

We have chosen the symbol $s'$ in $z(s', w)$ in order not to mix it up with the curve parameter $s$ in the path ordered exponential. The color matrix $F^{\mu\nu}(z(s', w))$ has to be included in the path ordering prescription with respect to $s$ of the field line characterized by $w$. In this way (3.7) behaves correctly under gauge transformations.

Finally one has to multiply by $u(z; x, y)$ and perform the integration over $d^4z = \rho(w) ds'd^3w$, resulting in

$$ \int u(z; x, y) g F^{\mu\nu}(z) d^4z = \int \rho(w) u(z(s', w)) P \left\{ g F^{\mu\nu}(z(s', w)) \exp [ig \int_{y}^{x} A dz] \right\} ds'd^3w + O(g^2). \quad (3.8) $$

This type of integral is a generalization of the operator insertions into a Wilson loop introduced by Eichten and Feinberg and later on widely used in the literature. It is, however, more general, because the insertions are not equally distributed along the path, but weighted by the $s'$-dependence of $u(z(s', w))$.

We summarize our representation for the quark propagator: It has the decomposition (2.3) with
\[ s = m \Delta(x - y) \int \rho(w) P \exp[i g \int_y^x \! Adz] d^3w, \quad (3.9) \]

\[ p = 0, \quad (3.10) \]

\[ v^\mu = \frac{i}{2} D^\mu(x) \left\{ \Delta(x - y) \int \rho(w) P \exp[i g \int_y^x \! Adz] d^3w \right\} \]

\[ -\frac{i}{2} \left\{ \Delta(x - y) \int \rho(w) P \exp[i g \int_y^x \! Adz] d^3w \right\} \hat{D}^\mu(y) \quad (3.11) \]

\[ a^\mu = \frac{i}{4} \epsilon^{\mu\nu\lambda\kappa} \left\{ D_\lambda(x) \int \rho(w) u(z(s', w)) P \left\{ g F^{\mu\nu}(z(s', w)) \exp[i g \int_y^x \! Adz] \right\} ds' d^3w \right\} \]

\[ -\int \rho(w) u(z(s', w)) P \left\{ g F_{\nu\kappa}(z(s', w)) \exp[i g \int_y^x \! Adz] \right\} ds' d^3w \hat{D}^\kappa(y), \quad (3.12) \]

\[ t^{\mu\nu} = -\frac{m}{2} \int \rho(w) u(z(s', w)) P \left\{ g F^{\mu\nu}(z(s', w)) P \exp[i g \int_y^x \! Adz] \right\} ds' d^3w. \quad (3.13) \]

We could have simplified the curly brackets \{\cdots\} in \( v^\mu \) by \( s/m \), but we left it in the present form in order to keep it applicable also in the case of small or vanishing mass. Obviously the last two lines in \( v^\mu \) and the terms \( a^\mu \) and \( t^{\mu\nu} \) contain the same types of insertions into path ordered integrals.

Clearly the representation (2.3), (3.9)-(3.13) fulfills all properties mentioned at the end of sect. 1.

### 4 Weight function, stream lines, and a simple application

It is now appropriate to switch to Euclidean space, i.e. put \( x^0 = -ix_4^E, x^n = x_n^E, \gamma^0 = \gamma_4^E, \gamma^n = i \gamma_n^E \). In the following the index \( E \) will be written explicitly only where it appears appropriate. For the explicit calculations we make use of the rotation symmetry around the vector \((x - y)_\mu\). Choose, just for intermediate simplification of notation, a system where \((x - y)_\mu\) has a four-component only and where \( \mathbf{x} = \mathbf{y} = 0 \). It is then convenient to describe the vector \( z_\mu \) by its four-component \( z_4 \), and ordinary three-dimensional polar coordinates \( r, \Theta, \varphi \), i.e.

\[ z_\mu = (r \sin \Theta \cos \varphi, r \sin \Theta \sin \varphi, r \cos \Theta, z_4). \quad (4.1) \]

Let us now choose an appropriate parametrization of the various stream lines. We classify them by the orthogonal distance \( w \) of the line from the midpoint \((x + y)/2\) between the sources, and by the angles \( \Theta \) and \( \varphi \), thus \( d^3w = dw d\Theta d\varphi \). Fig. 1 shows a stream line together with the parameters introduced above.
The weight function \( \rho(w) \) defined in (3.3) becomes
\[
\rho(w, \Theta) = \frac{\partial(z_1, z_2, z_3, z_4)}{\partial(r, \Theta, \varphi, z_4)} \frac{\partial(r, \Theta, \varphi, z_4)}{\partial(s, w, \Theta, \varphi)} = r^2 \sin \Theta \frac{\partial(r, z_4)}{\partial(s, w)}.
\]  
(4.2)

It is now very convenient that \( \rho \) depends on \( w \) only, but not on the curve parameter \( s \). Therefore we can evaluate it at a suitable point. We choose the symmetry point \( \zeta \) in the middle of the stream line with the coordinates \( r = w, z_4 = (x + y)_4/2 \). At this point we obviously have \( \partial r/\partial s = 0 \) and \( \partial r/\partial w = 1 \), and thus \( \partial(r, z_4)/\partial(s, w) = -\partial z_4/\partial s = -u_4 \), where we used the definition of the stream lines in (3.2). This allows to write down the weight function in closed form.
\[
\rho(w, \Theta) = -w^2 \sin \Theta \ u_4(\zeta; x, y),
\]  
(4.3)

with
\[
\zeta = (w \sin \Theta \cos \varphi, w \sin \Theta \sin \varphi, w \cos \Theta, (x + y)_4/2).
\]  
(4.4)

Obviously \( u_4(\zeta; x, y) \) is independent of the angles \( \Theta, \varphi \).

It is convenient to put
\[
\rho(w, \theta) = \frac{\hat{\rho}(w)}{4\pi} \sin \Theta,
\]  
(4.5)

such that the weight function \( \hat{\rho}(w) \) is normalized to
\[
\int_0^\infty \hat{\rho}(w) dw = 1.
\]  
(4.6)

We first give the resulting formulae for the special case of vanishing quark mass which may be of some general interest:
\[
\Delta^{(0)}(x) = -\frac{1}{4\pi^2 x^2},
\]  
(4.7)
\[
  u^{(0)}(z; x, y) = \frac{1}{16\pi^4(x-z)^2(z-y)^2},
\]
\[
  u^{(0)}_\mu(z; x, y) = -\frac{(x-y)^2}{2\pi^2} \left\{ \frac{(x-z)_\mu}{(x-z)^4(z-y)^2} + \frac{(z-y)_\mu}{(x-z)^2(z-y)^4} \right\},
\]
\[
  \rho^{(0)}(w) = \frac{2w^2|x-y|^3}{\pi[(x-y)^2/4 + w^2]^3/2}.
\]

The maximum of \(\rho^{(0)}\) is at \(w_{\text{max}}^{(0)} = |x-y|/(2\sqrt{2})\).

We now come to the general massive case. The free scalar propagator \(\Delta\) then is
\[
  \Delta(x) = -\frac{m^2}{4\pi^2x} K_1(mx),
\]
which gives
\[
  u(z; x, y) = \frac{m^2}{16\pi^4} \frac{K_1(m|x-z|)K_1(m|z-y|)}{|x-z||z-y|}. 
\]

With the relation \((K_1(z)/z)' = -K_2(z)/z\) for the Kelvin function, one further obtains
\[
  u_\mu(z; x, y) = -\frac{m^2|x-y|}{4\pi^2 K_1(m|x-y|)} \left\{ \frac{K_2(m|x-z|)K_1(m|z-y|)}{(x-z)^2(z-y)} (x-z)_\mu \right. 
  \left. + \frac{K_1(m|x-z|)K_2(m|z-y|)}{|x-z|(z-y)^2} (z-y)_\mu \right\}.
\]
\[
  \hat{\rho}(w) = \frac{m^2w^2(x-y)^2}{\pi} \frac{K_1(m\sqrt{(x-y)^2/4 + w^2})K_2(m\sqrt{(x-y)^2/4 + w^2}^3/2)}{K_1(m|x-y|)((x-y)^2/4 + w^2)^{3/2}}.
\]

The weight function \(\hat{\rho}(w)\) is trivially suppressed for small \(w\) by the volume element in polar coordinates, it rises to a maximum at some \(w_{\text{max}}\), and decreases exponentially for large \(w\).

In Fig. 2 we plot the function \(\hat{\rho}(w)\) for fixed distance \(|x-y|\) for various values of the mass. It is clearly seen how the maximum moves to the left if the quark mass increases.
Fig. 2: The weight function $\hat{\rho}(w)$ defined in (4.6). The distance $|x - y|$ is fixed, we show the curves for the mass values $m|x - y| = 0, 1, 5, 10, 20$. The maximum moves from right to left and increases with increasing mass.

In Fig. 3 we show the stream lines of the vector field $u_\mu$ for four values of $m|x - y|$.

Fig. 3: The stream lines (3.2) for the cases (from left to right, top to bottom) a) $m|x - y| = 0$, b) $m|x - y| = 5$, c) $m|x - y| = 20$, d) $m|x - y| = 100$. The sources at $x$ and $y$ are located at $\pm0.5$. We show the stream lines for the values of $w/w_{max} = 0, \pm0.5, \pm1, \pm1.5$.

For vanishing mass the lines spread out widely in space, up to $w$ of the order of $|x - y|$. For increasing mass they concentrate more and more to the straight line connection. Apparently the product $m|x - y|$ has to become quite large, however, in order to get a sizeable concentration.

The large mass limit will now be investigated analytically. We use the asymptotic behavior $K_n(z) \sim \sqrt{\pi/(2z)}e^{-z}$ for $z \to \infty$ and get
\[ \Delta(x) \sim -\frac{1}{2} \frac{\sqrt{m}}{(2\pi x)^{3/2}} e^{-m x} \quad \text{for } m \to \infty, \quad (4.15) \]

\[ u(z; x, y) \sim \frac{m}{32\pi^3} \frac{e^{-m|x-z|+|z-y|}}{|x-z||z-y|^{3/2}} \quad \text{for } m \to \infty, \quad (4.16) \]

\[ u_\mu(z; x, y) \sim -\frac{1}{2} \left( \frac{m|x-y|}{2\pi|x-z||z-y|} \right)^{3/2} \left\{ \frac{(x-z)_\mu + (z-y)_\mu}{|x-z| |z-y|} \right\} e^{m|x-y|-|x-z|+|z-y|} \quad \text{for } m \to \infty, \quad (4.17) \]

\[ \hat{\rho}(w) \sim \frac{m^{3/2}w^2|x-y|^{5/2}}{\sqrt{2\pi}(x-y)^2/4 + w^2} \exp[m(|x-y| - 2\sqrt{(x-y)^2/4 + w^2})] \quad \text{for } m \to \infty. \quad (4.18) \]

Obviously only \( w \) of order \( \sqrt{|x-y|}/m \) are now of relevance in the weight function. Therefore one may expand the square roots and gets the simple result

\[ \hat{\rho}(w) \sim \frac{16}{\sqrt{2\pi}} \left( \frac{m}{|x-y|} \right)^{3/2} w^2 \exp\left[ -\frac{2mw^2}{|x-y|} \right] \quad \text{for } m \to \infty. \quad (4.19) \]

This is just the picture which one expects for large mass. Only stream lines near the straight line connection essentially contribute, the maximum of \( \hat{\rho}(w) \) is at \( w_{\text{max}} = \sqrt{|x-y|/(2m)}. \)

If the mass is large enough, such that the variation of the gluon field in transversal direction becomes negligible, all line integrals give the same contribution. The weighted superposition over the paths can then simply be replaced by the path along the straight line connection. This means that one has effectively a three-dimensional \( \delta \)-function in transversal direction. The situation looks now similar to the case of the static propagator but with an important difference. While the static propagator (1.8) singles out a special reference frame, our propagator is manifestly Lorentz covariant. It is the vector \( (x-y)_\mu \) which specifies the direction of propagation.

This has a simple consequence. In the limit of large mass, the scalar function \( s \) in (3.9) becomes

\[ s \sim -\frac{1}{2} \left( \frac{m}{2\pi|x-y|} \right)^{3/2} e^{-m|x-y|} P \exp[-i g \int_y^x A dz], \quad (4.20) \]

with the Euclidean path along the straight line connecting \( x \) and \( y. \)

Consider now the term \( s \) plus the first two lines of \( u^\mu \gamma_\mu \) in (3.11). This sum can be written as

\[ s = \frac{1}{2m} (D^E_\mu(x)s - s D^E_\mu(y)) \gamma^E_\mu \]

\[ \sim - \left( \frac{m}{2\pi|x-y|} \right)^{3/2} \frac{1 + \gamma^E_{\mu}(x-y)_\mu/|x-y|}{2} e^{-m|x-y|} P \exp[-i g \int_y^x A dz]. \quad (4.21) \]
In contrast to the $e^{-m|x-y|}$ of the static approximation which falls off with the euclidean time difference, our $e^{-m|x-y|}$ falls off with the euclidean distance. This means that the Hamiltonian is correctly given by the full relativistic energy. Furthermore we also get the correct projection operator for the $\gamma$-matrices.

We finally discuss the terms $a^\mu$, $t^{\mu\nu}$ in (3.12), (3.13) which can be treated rather simply. We specialize to the case $x = y$ and put $x_4 - y_4 = T > 0$. In the large mass limit we can replace $\rho(w)$ by $\delta^{(3)}(w)$. The $s'$-integrations can be written as follows:

$$u(z(s',0))ds' = u(t)\frac{ds'}{dt} = \frac{u(t)}{u_4(t)}dt = -\frac{e^{-mT}}{8\sqrt{2}\pi^{3/2}T^{3/2}\sqrt{m}}dt. \quad (4.22)$$

In the second step we used the 4-component of (3.2), in the third step we introduced the asymptotic formulae (4.16), (4.17) for the special case where $z$ lies on the line connecting $x$ and $y$. In $a^\mu$ we only need to consider the index $\lambda = 4$ which gives a leading factor $(-m)$ from the differentiation of $e^{-mT}$. All other contributions are suppressed by higher powers of $1/m$. In $t^{\mu\nu}\sigma_{\mu\nu}$ only spatial indices $\mu = m, \nu = n$ survive if we concentrate on the diagonal part of the Dirac matrices. The transition to $2 \times 2$-matrices than gives

in the axial vector (3.12): $a^\mu\gamma^5\gamma_\mu \Rightarrow \epsilon_{mnk}F_{nk}\gamma^5\gamma_m \Rightarrow \epsilon_{mnk}F_{nk}\sigma_m \Rightarrow 2B_m\sigma_m,$

in the tensor (3.13): $t^{\mu\nu}\sigma_{\mu\nu} \Rightarrow F_{mn}\sigma_{mn} \Rightarrow \epsilon_{mnk}F_{mn}\sigma_k \Rightarrow 2B_k\sigma_k. \quad (4.23)$

Both contributing expressions give identical magnetic field insertions which add up.

The resulting expressions have to be combined with the leading term (4.21) where we can drop the projector $(1 + \gamma^E_4)/2$. This gives a spin dependent expression of the form

$$-\left(\frac{m}{2\pi T}\right)^{3/2}e^{-mT}P \exp[-ig \int^x_y Adz](1 + \frac{g}{m} \int B(t)sdz). \quad (4.24)$$

If we take the product of the quark- and the antiquark propagator which arises in the four-point Green function, focus on the product term of the two magnetic field insertions, and extract the Hamiltonion in the usual way \[3,\] we immediately obtain the spin spin and the tensor terms with the correct representations for the corresponding potentials $V_4$ and $V_3$. In the static approach these terms only arise as higher order corrections.

Spin orbit terms are momentum dependent and therefore involve moving quarks. These are, of course, contained in our formalism, but the derivation is slightly more complicated. Spin independent corrections are obtained even harder. Therefore we will not discuss those in this first application, but be content with the simple and correct derivation of spin spin and tensor terms given above.

### 5 Conclusions and outlook

Let us first compare our representation (2.3), (3.9)-(3.13) with the static approximation (1.8). The static approximation is, drastically speaking, completely wrong everywhere. It is completely wrong for $x \neq y$ where it vanishes, but it is also completely wrong for $x = y$ because it has a $\delta$-function there which is not present in the exact propagator. These features survive if one treats the neglected spatial part $i\gamma^m D_m$ in the field equation
as perturbation. In any finite order of this perturbation the approximated propagator vanishes for \( x \neq y \), while higher order derivatives of \( \delta^{(3)}(x - y) \) appear. The miracle that one can nevertheless obtain useful results from this propagator is due to the fact that the perturbation series turns out to become an expansion with respect to \( 1/m \). As long as \( <p^2/m^2> \) is small, the results are reliable.

The propagator proposed in the present work is manifestly covariant and appears to have a reasonable structure everywhere. This advantage is payed by a more complicated form, which, however, looks very natural physically. Not only one path ordered integral, but a whole set of them contribute. We saw how the paths near the straight line connection dominate for large mass. We believe that it will also be possible to get useful information for finite mass. For an investigation of the quark-antiquark interaction one should start, as usual, with the gauge invariant \( q\bar{q} \) four-point Green function and integrate out the fermion fields. Instead of the familiar Wegner-Wilson loop one will now obtain a superposition of loops where the straight paths in time direction are replaced by the stream lines making up our propagator. Quite a lot of knowledge how to treat such loops has been accumulated by various authors which can be used for this investigation.

A comparison with the Feynman-Schwinger representation (see e.g. \[3\]) is also instructive. This representation of the propagator is formally exact and has essentially the form of a quantum mechanical Green function. It can therefore be written as a path integral. In the literature \[3\] one also finds approximate path integral representations for the propagator and the quark-antiquark kernel valid up to order \( 1/m^2 \). In both cases one has, as usual in this formalism, a sum over all paths, which is, conceptionally as well as technically, a rather delicate concept. Even for rather simple situations the path integral cannot be evaluated. In our case, on the other hand, we have only line integrals along a well defined set of field lines. This is a much simpler situation. Our representation stands just between approximations which involve a single path only and those requiring a sum over all paths.

Besides the application of the present propagator for large as well as for finite mass, there is another topic which should be worked out. This is the systematic improvement of our propagator. We don’t have a simple differential equation for it, as it is available in the case of the static propagator. Therefore one probably has to improve higher orders of perturbation theory directly and transform them into gauge covariant expressions in an analogous way as done here for the first order. We emphasize, however, that such an improvement does not appear necessary for many purposes. The present form gives already the correct relativistic energy of a free particle together with the correct spin projectors. Furthermore we have seen that it gives the correct spin-spin and tensor forces for heavy quark-antiquark systems. To get these ”relativistic corrections” from the static propagator one has to make quite complicated manipulations. We expect that all other relativistic corrections can also be obtained with some more effort.

Therefore there are good reasons to believe that the suggested expression for the propagator will turn out quite useful already in it’s present form.

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