Including the Z in an Effective Field Theory for dark matter at the LHC

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Abstract

An Effective Field Theory for dark matter at a TeV-scale hadron collider should include contact interactions of dark matter with the partons, the Higgs and the Z. This note estimates the impact of including dark matter-Z interactions on the complementarity of spin dependent direct detection and LHC monojet searches for dark matter. Their effect is small, because they are suppressed by electroweak couplings and the contact interaction self-consistency condition \( C/\Lambda^2 < \frac{4\pi}{s} \). The contact interactions between the Z and dark matter are parametrised by derivative operators; this is convenient at colliders because such interactions do not match onto the quark-dark matter contact interactions.

1 Introduction

A diversity of cosmological observations imply that a quarter of the current mass of our Universe is unknown “dark matter”\(^1\). Various experiments attempt to detect the particle making it up. For instance, direct detection (DD) experiments \(^2, 3, 4, 5\), search for \(\sim\) MeV energy deposits due to scattering of dark matter particles from the galactic halo on detector nuclei. And the Large Hadron Collider (LHC) searches \(^6, 7\) for dark matter pairs produced in multi-TeV \( pp \) collisions, which would materialise as an excess of events with missing energy and jets. The LHC and DD searches are at very different energy scales, so different Standard Model (SM) particles are present, and also the quantum interferences are different\(^8\). The expected rates can be compared in specific dark matter models \(^9\), or, in recent years, several studies \(^6, 10, 11, 12, 13, 14\) have compared the LHC and DD sensitivities using a contact interaction parametrisation of the dark matter interactions with the standard model particles.

The LHC bounds obtained in this way are restrictive, and probe smaller couplings than direct detection experiments searching for “spin dependent” interactions between partons and dark matter \(^4\). These contact interaction studies are referred to as “Effective Field Theory” (EFT), and considered to be relatively model independent. However, the particle content is an input in EFT, and the restrictive LHC limits assume that the dark matter particle is the only new particle accessible at the LHC. Relaxing this assumption can significantly modify the experimental sensitivities \(^12, 13, 14\). This has motivated various simplified models for dark matter searches at the LHC \(^15, 16, 17\). Retaining this assumption, as will be done in this note, is only marginally consistent, because the contact interactions to which the LHC is sensitive would have to be mediated by strongly coupled particles. As recalled in the next section, this implies that colliders can exclude contact interactions of order their sensitivity, but not much larger.

Effective Field Theory (EFT) is supposed to be a recipe to get the correct answer in a simple way \(^18\). So this note attempts to compare LHC and DD constraints on dark matter, according to the prescriptions of \(^18\). From a “bottom-up” phenomenological perspective, an EFT for dark matter at the LHC should parametrise all possible SM-gauge invariant interactions of the dark matter with other on-shell particles. So first, contact interactions between the dark matter and the Higgs or Z should be included at the LHC. These can interfere with the contact interactions studied in previous analyses, but contribute differently at colliders from in direct detection, so the linear combination of operator coefficients constrained at high and low energy will be different. Secondly, an EFT contains in principle a tower of operators \(^19\) organised in increasing powers of the inverse cutoff scale \( 1/\Lambda \), and higher orders can be neglected if they are “suppressed”. The importance of higher dimensional operators will be left to a subsequent publication \(^1\). This note focuses on the first point, and estimates analytically the consequences of including the lowest dimension operators allowing dark matter interactions with the Z \(^2\) Section 2 outlines a peculiar choice of operators for the Z vertex; they are proportional to the momentum-transfer-squared. This choice appears convenient, because the effects of the Z are not mixed into the dark-matter-quark contact interactions. Section 3 estimates the impact of cancellations

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\(^2\) Higher dimensional operators can contain more fields and be suppressed by phase space, or contain Higgs fields and be suppressed by \( \langle H \rangle^2 / \Lambda^4 \), or contain derivatives and be dangerous.

\(^3\) Contact interactions between dark matter and the Z have been proposed in \(^17\) as a benchmark model, assuming other contact interactions to be absent.
between $Z$ exchange and dark matter contact interactions with quarks at the LHC, and section 4 recalls the direct detection bounds.

2 Assumptions, Operators, and EFT

The low energy consequences of New Physics from above a scale $\Lambda$ can be parametrised by contact interactions of coefficient $C/\Lambda^n$. The New Physics cannot be more than strongly coupled, implying $C < 4\pi$, and the “low” energy scale must be below $\Lambda$. This means that an experiment can exclude

$$\frac{4\pi}{s} > \frac{C}{\Lambda^2} > \text{sensitivity}, \quad (1)$$

where $s$ is the four-momentum-squared of the process. Low energy experiments, where $s \to 0$, therefore can be taken to exclude everything above their sensitivity. However, the upper limit of eqn (1) is relevant for collider dark matter searches, where $s$ is the invariant mass of the invisibles.

This parametrisation in terms of contact interactions is reasonably model-independent, if a complete set is obtained by adding to the Lagrangian (below the scale $\Lambda$) all operators up to some order in $\Lambda$ coefficients $C$. Experimental limits on contact interactions will therefore be presented as limits on the dimensionless $C/\Lambda^2$.

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where to the right of the arrow is the resulting vertex, $B^\mu$ is the hypercharge gauge boson with coupling $g' = e \tan \theta_W \equiv e s_w / c_w$, $B^{\mu
u} = \partial^\mu B^\nu - \partial^\nu B^\mu$, and a term $\propto p_Z \cdot Z$ was dropped after the arrow in $\mathcal{O}_{Z,A}$, assuming the $Z$ was produced by light quarks. There is in addition a “dipole moment” operator $B^{\mu
u} \bar{\chi} \sigma^{\mu\nu} \chi$, which is neglected here because it also induces dark matter interactions with the photon [21] which are more interesting.

The $Z$ operators are chosen $\propto p_Z^2$ so that they are relevant at the LHC, where the $Z$ is an external leg of the EFT, but do not contribute in the low-energy scattering of DD. This choice should be acceptable, because the operator basis can always be reduced by using the equations of motion [23]. These are, neglecting gauge-fixing terms for $B^\mu$ [21]

\[ D_\mu B^{\mu\nu} = g' y_H (H^\dagger D^\nu H - (D^\nu H)^\dagger H) + g' \sum_\psi \bar{\psi} \gamma^\nu \gamma^\mu \psi \]

where $\psi$ is a SM fermion of hypercharge $y_\psi$. Usually [21], the derivative operators are dropped, and the operator proportional to the Higgs v.e.v. squared $\langle H \rangle^2$ is retained. In this usual basis, $Z$ exchange at the LHC is $\propto m_Z^2 / (p_Z^2 - m_Z^2)$, so negligible for $p_Z^2 \gg m_Z^2$. But it should be included in the quark-$\chi$ contact interaction used in direct detection, so the coefficient of the operators of eqn (3) would not be the same in direct detection as at the LHC.

For $m_\chi < m_Z / 2$, the invisible width of the $Z$ (at “2σ”, so [22] $\Gamma(Z \rightarrow \chi \chi) < 3 \text{ MeV}$) imposes that $|C_{Z,\nu}| < 8.9 (\Lambda / \text{TeV})^2$, for $B = V, A$. This is only marginally more restrictive than the LHC limits, despite the 0.1% precision on the invisible width of the $Z$, because the operator is $\propto p_Z^2$, which amplifies its coupling to very energetic off-shell $Z$s at the LHC.

At dimension seven, the equation of motion for the Higgs field

\[ D_\mu D^\mu H = \mu^2 H - \lambda H^2 H - \tau Y_\mu \ell - \tau Y_\mu q_L + \varepsilon_{\mu \nu} Y_{\mu \nu} u \]

($Y_\ell$ are Yukawa matrices), can be used to exchange the more usual $(H^\dagger H)^2 \bar{\chi} \chi$, $(H^\dagger H)\bar{\chi} \chi$ operators for the double-derivative interactions between dark matter and the Higgs:

\[ H^\dagger D^\mu D_\mu H \bar{\chi} \chi \rightarrow - \frac{m_W^2}{2} W^+ \gamma^\mu W^- \bar{\chi} \chi - m_Z^2 Z_\mu \bar{\chi} \chi + \frac{v p_h^2}{\sqrt{2}} h \bar{\chi} \chi \]

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(5)

where $\langle H \rangle = v = 174 \text{ GeV}$, $p_h$ is the four-momentum of the physical Higgs particle $h$, and after the arrow are the interactions induced by the operator. The operators of eqn (3) are interesting, because they give a Higgs coupling to dark matter $\propto p_h^2$, which has the desirable feature of being relevant at the LHC where the Higgs is in the effective theory, but not contributing at low energy. It is possible to use the Equations of motion to replace two operators with one, because I am only interested in the $h$-$\chi$-$\chi$ interaction, due to the minimisation condition of the Higgs potential.

The dark matter interactions to $W$ and $Z$ pairs, given after the arrows in eqns (3), were studied in [24], who used $U(1)_{em} \times SU(3)$ invariant operators such that these contact interactions have dimension five with coupling $1 / \Lambda_{CHLR}$. They find that the 8 TeV LHC with luminosity 25 fb$^{-1}$ could probe $\Lambda_{CHLR} \lesssim \text{TeV}$. This constrains the coefficients of the operators of eqn (3) to be $\lesssim 1 / (\text{TeV} m_W^2)$, which is not restrictive. A more significant limit, of 10 TeV$^{-3}$ arises for $m_\chi < m_h / 2$ from requiring $\Gamma(h \rightarrow \chi \chi) \lesssim \Gamma(h \rightarrow bb)$. This restriction should be reasonable [23] because the Higgs is observed to decay to $bb$.

Dark matter interactions with the Higgs are neglected in the rest of this note, because the operators of eqns (3) and (4) constitute a sufficient “toy model” in which to estimate the impact of including the $Z$. Since the LHC produces more $Z$s than Higgses, one could anticipate that the $Z$ is more likely to have a significant effect on the LHC’s sensitivity to dark matter.

Figure 1: Effective interactions contributing to $q q \rightarrow \chi \bar{\chi}$ at the LHC. The coefficient of the four fermion operator is $C_{q,AH} / \Lambda^2$, and the effective axial vector coupling of the $Z$ to dark matter is $s_{\omega} p_Z^2 C_{Z,A} / \Lambda^2$. 

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Given the operators of eqns (3) and (4) at the LHC, the axial vector dark matter current can interact with quarks $Q$ via the diagrams of figure 1 which will can be written as a four-fermion interaction of coefficient
\[
C_{QX,A} = C_{QX,A} + g \frac{g_{sw}}{p_{\perp}^2 - m_Z^2} C_{Z,A} \rightarrow \rho_Z \gg m_Z^2, C_{QX,A} + g \frac{g_{sw}}{p_{\perp}^2 - m_Z^2} C_{Z,A}, \tag{6}
\]
where $g_X^Q = \{1 - \frac{4}{3}s_w^2, - \frac{4}{3}s_w^2, -1 + \frac{2}{3}s_w^2, \frac{2}{3}s_w^2\}$ for $\{uL, uR, dL, dR\}$. A similar expression can be obtained for the vector $\chi$ current. The $Z$ exchange looks like a contact interactions for large $p_{\perp}^2 = M_{inv}^2 \gg m_Z^2$, where $M_{inv}$ is the invariant mass-squared of the dark matter. This approximation will used below because the best limit arises at larger $M_{inv}$.

3 Estimated limits from the LHC

Dark matter particles are invisible to the LHC detectors, so pair production of $\chi$ can be searched for in missing (transverse) energy ($E_T$) events, which can be identified by jet(s) radiated from the incident partons. The principle Standard Model background for such “monojet” searches is $Z+$ jet production, followed by $Z \rightarrow \nu\bar{\nu}$. The 8 TeV LHC is sensitive to dark matter contact interactions with $C/A^2 \sim$ TeV$^{-2}$.

The aim here is to analytically estimate the invisible four-momentum-squared $M_{inv}^2$, by comparing the partonic cross-sections for $\nu\bar{\nu}$ and $\chi\chi$ production. I assume that the QCD part of the amplitude is identical in both cases, so it does not need to be calculated. This has the advantage of allowing for an arbitrary number of jets which is more difficult to simulate[20] (the data frequently contains more than one jet[6]). In the matrix element for $\nu\bar{\nu}$ will appear
\[
g_X^Q \frac{g^2}{4 Q_X} \frac{1}{2} \frac{1}{p_{\perp}^2 - m_Z^2 + \frac{m_Z^2}{\Gamma_Z}} \left(\bar{Q}\gamma^\alpha P_X Q\right)\left(\nabla\gamma^\alpha P_L\nu\right)
\]
whereas, for DM production via the $\tilde{\chi}^\alpha\gamma_5\chi$ current, this is replaced by:
\[
\frac{C_{QX,A}}{A^2} \left(\bar{Q}\gamma^\alpha P_X Q\right)\left(\nabla\gamma^\alpha P_L\chi\right)
\]
Then the full matrix element must be squared and integrated the phase space of $N$ jets and two invisible particles. The invisibles can be treated as a single particle of variable mass $p_{\perp}^2 = M_{inv}^2$, using the identity
\[
d\Phi_{N+2} = \delta^4(P_{\text{in}} - \sum q_i - p) \prod_{i=1}^{N} \frac{d^3q_i}{2E_{\text{in}}(2\pi)^3} \times
\]
\[
(2\pi)^3 dp^2 \delta^4(p - p_{\text{in}} - p_{\chi}) \frac{d^3p_X}{2E_X(2\pi)^3} \frac{d^3p_\chi}{2E_\chi(2\pi)^3}.
\]
Neglecting spin correlations and the dark matter mass, the invisible phase space integral over the gamma-matrix trace for the invisible fermions gives $M_{inv}^2/(8\pi)$ for $\chi\chi$, and $3M_{inv}^2/(16\pi)$ for neutrinos. For neutrinos in the final state, $M_{inv}^2 = m_Z^2$ due to the delta-function-like behaviour of the $Z$ propagator-squared. However, for dark matter, the $M_{inv}^2$ phase space integral will privilege larger values of $M_{inv}^2$. Treating the $N$ jets of the event as a particle of negligible mass, the upper bound on $M_{inv}^2$ is $\sim 4E_T^2$, where $E_T$ is the invisible transverse energy. The CMS study[3] uses the range $400\text{ GeV} \leq E_T \leq \text{TeV}$. Therefore most of the dark matter signal will come from $M_{inv}^2 \gg m_Z^2$, and the approximation (5) is consistent. Furthermore, the contact interaction approximation requires $\Lambda^2 > M_{inv}^2$ which suggests $\Lambda \gtrsim 1 - 2 \text{ TeV}$.

The CMS collaboration obtains a limit $\Lambda > 950\text{ GeV}$, for the sum of the operators of eqn (3), each with $|C_{QX,A}| = 1$. There is also an upper limit on the Cs which a collider can exclude, eqn (11), from requiring that the contact interaction approximation be self-consistent: $C/A^2 < 4\pi/s$. It will be applied below for $s \sim \text{TeV}^2$. For the axial $\chi$ current with $\Lambda = \text{TeV}$, the CMS limit and eqn (11) give 3 independent bounds on $\{c_{QL,A}, c_{uR_A}, c_{dR_A}\}$:
\[
4\pi \lesssim \sqrt{\frac{2}{3}} |C_{QL,A} + \frac{2}{15} C_{Z,A}|^2 + \frac{1}{3} |C_{QL,A} - \frac{1}{6} C_{Z,A}|^2 \lesssim \sqrt{2}
\]
\[
4\pi \lesssim |C_{uR_A} - \frac{1}{10} C_{Z,A}| \lesssim \sqrt{3}
\]
\[
4\pi \lesssim |C_{dR_A} + \frac{1}{30} C_{Z,A}| \lesssim \sqrt{6}
\]

\[3\]To obtain an estimated bound by comparing partonic cross-sections (as above), an upper limit for the $dM_{inv}^2$ integral is required, and it is unclear how to fix it, because the assumption that the jet emission part of the cross-section is the same as for $\nu\bar{\nu}$ pairs will fail, if $M_{inv}^2$ is a significant fraction of the energy of the event. With the $M_{inv}$ cutoff at 800 GeV, requiring that the dark matter contribute $\lesssim 1/6$ of the SM background, gives an estimated bound twice as good as the CMS limit.
where the first line is the summed contributions of $u_L$ and $d_L$, the fractions are approximations $g g_X s_w/2 c_w$, and the $d$ to $u$ pd ratio is taken $1/2$. Similar limits apply for the operators of eqn (2).

It can be seen already from eqn (7), that including the interactions with the $Z$ will make little differences to the LHC limits on the $C_{Q,\chi, A}$: for the doublet quarks, the $Z$ contribution cannot cancel simultaneously against the $u_L$ and $d_L$ contributions, and the $Z$ contribution is irrelevant for the singlet quarks, because also $C_{Z, A}$ must be $\lesssim 4\pi$. The parameters ruled out by the first and second eqns of (7) are represented as the central regions in figure 2.

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\[ (9) \] is represented in figure 2 as the vertical exclusions.

4 From the TeV to the MeV

In direct detection, the dark matter scatters non-relativistically off nuclei. Therefore, to translate the EFT from the TeV to the MeV, the $Z$ must be removed, the effects of QCD loops in running the operator coefficients should be included, and the quarks must be embedded in the nucleons.

To remove the $Z$, the Greens function for two quarks and two $\chi$s in the Effective theory with a $Z$, should be matching to the same Greens function in the theory without a $Z$. Since the matching is performed at zero momentum for the fermion legs, the contact interactions of eqn (4) do not contribute, and the coefficients of the four-fermion operators of eqns (3,2) remain the same after the $Z$ vertices were taken $\propto p^2 Z$ to obtain this.

The light quark currents $\bar{7} \gamma^{\mu} P_{X} q$ are conserved in QCD, so do not run. Also, since $\chi$ is a SM gauge singlet and the only dark sector particle below the TeV, I suppose that the operators with vector and axial vector $\chi$ currents do not mix below the TeV.

Finally, the quark currents can be embedded in nucleons $N = \{p, n\}$ using identities [29] such as

\[ \langle N|\bar{Q}_i^\mu Q_i|N\rangle = c_{V, i}^N \langle N|\bar{\psi}_N \gamma^\mu \psi_N|N\rangle \]

where $c_{V, i}^p = c_{V, i}^n = 2$, and $c_{V, i}^d = c_{V, i}^u = 1$, because this current counts valence quarks in the nucleon. The axial quark current is proportional to the nucleon spin:

\[ \langle N|\bar{Q}_i^\mu \gamma^5 Q_i|N\rangle = 2 s^i \Delta Q_i^N = \Delta Q_i^N \langle N|\bar{\psi}_N \gamma^\mu \gamma^5 \psi_N|N\rangle \]

where the proportionality constants are measured [25] as $\Delta u^p = \Delta d^p = 0.84$, $\Delta u^n = -0.43$. In the zero-momentum-transfer limit of non-relativistic scattering, the dark matter can have spin-dependent interactions via the axial current, or spin-independent interactions via the first component of the vector current.

The spin-independent scattering amplitude for $\chi$ on a nucleon, is a coherent sum of vector and scalar interactions, for quarks of both chiralities and all flavours. The experimental limit on the cross-section per nucleon is $\sigma_{SI} \lesssim 10^{-44} \text{ cm}^2$ for $m_\chi \sim 100 \text{ GeV}$ [3]. For the proton ($uR \leftrightarrow dR$ for the neutron), with $C_{q, R, V} = \frac{1}{3}(C_{dR, V} + 2C_{uR, V})$, this gives [29]

\[ \sigma_{SI} \sim \frac{1}{\pi} \left[ \frac{3 m_N}{2 A^2} (C_{qL, V} + C_{qR, V} + \ldots) \right]^2 \lesssim 3 \times 10^{-17} \text{ GeV}^{-2} \]

where the $\ldots$ is scalar contact interactions neglected in this note. For $\Lambda = \text{TeV}$, this gives

\[ [C_{qL, V} + \frac{1}{3}(C_{dR, V} + 2C_{uR, V}) + \ldots] \lesssim 10^{-2} \quad (SI). \] (8)

The spin dependent cross-section per proton is [29]

\[ \sigma_{SD} \sim m_p^2 \left[ \frac{42(C_{qL, A} + C_{dR, A} - 2C_{dR, A})}{2 A^2} \right]^2 \lesssim 10^{-10} \frac{4}{\text{ GeV}^{-2}} \]

where the experimental bound is for $m_\chi \sim 100 \text{ GeV}$. For $\Lambda = \text{TeV}$, this gives

\[ |(C_{qL, A} + C_{uR, A} - 2C_{dR, A})| \lesssim 20 \quad (SD). \] (9)

Comparing to eqn (7) shows that the contact interactions explored by SD direct detection experiments are mediated by physics which is not a contact interaction at the LHC, so are not excluded by the limits given in eqn (7). The limit (10) is represented in figure 2 as the vertical exclusions.

5 Discussion

From a bottom-up EFT point of view, it is important to include all operators which can interfere, when computing experimental constraints. This is to allow for cancellations. Including several operators which do not interfere improves
the bound, but is not so well motivated. In this note, operators with vector and axial vector currents for the dark matter fermion $\chi$ were presented as an example, which illustrates two points.

First, the EFT at the LHC contains more particles than the light partons and dark matter that are relevant in direct detection. At the LHC, the Higgs and $Z$ should also be included. Matching the high and low energy EFTs, as done in this note, suggests that the LHC constrains several combinations of operator coefficients that are different from direct detection, as can be seen by comparing eqns (7) and (9). However, the contribution of the $Z$ is relatively unimportant, because its couplings to singlet quarks are small, and it interferes with opposite sign with $u_L$ and $d_L$. The LHC limits on the dark matter couplings to quarks and the $Z$ are represented as the central exclusion areas of figure 2, the coupling to quarks is more constrained than the coupling to the $Z$, and arbitrary axial current dark matter interactions to quarks cannot be allowed by tuning the dark matter coupling to the $Z$. This is because there is a self-consistency upper bound on contact interaction coefficients at colliders $C/\Lambda^2 < 4\pi/\hat{s}$ (see eqn (1)). It is important to notice that this upper bound also implies that the LHC limits do not exclude the parameter space probed by spin dependent direct detection experiments.

Second, an interesting difference between direct detection and collider experiments, is that quarks of different chirality and flavour interfere in direct detection, whereas the LHC can constrain the interactions of dark matter with each flavour and chirality of quark individually. This is related to the relative unimportance of the $Z$: it cannot cancel separately against the contributions of $u_L, d_L, u_R$ and $d_R$.

![Parameter space excluded by spin dependent direct detection experiments](image)

**Figure 2:** Parameter space excluded by spin dependent direct detection experiments (at either side), and the LHC (central region), for a dark matter fermion of mass $\sim 100$ GeV, with contact interactions with the $Z$ parametrised by $C_{Z,A}$ (see eqn (1)), and with $u_R$ quarks in the left plot, and the doublet $q_L$ in the right plot (see eqn (3)). $\Lambda = \text{TeV}$, and all other coefficients are zero. The upper limit of the LHC exclusions is estimated from eqn (1).

In summary, the rules of bottom-up Effective Field Theory say that one should include all operators up to some specified dimension. So to parametrise at dimension six the axial vector interactions of dark matter with quarks, one should include contact interactions of dark matter with the quarks and with the $Z$. Including interactions with the $Z$ that are $\propto p^2_Z$, as done here, suggests that these are not crucial.

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