Cold Nuclear Matter Effects on $J/\psi$ and $\Upsilon$ Production at the LHC

R. Vogt

Lawrence Livermore National Laboratory, Livermore, CA 94551, USA
and

Physics Department, University of California at Davis, Davis, CA 95616, USA

The charmonium yields are expected to be considerably suppressed if a deconfined medium is formed in high-energy heavy-ion collisions. In addition, the bottomonium states, with the possible exception of the $\Upsilon(1S)$ state, are also expected to be suppressed in heavy-ion collisions. However, in proton-nucleus collisions the quarkonium production cross sections, even those of the $\Upsilon(1S)$, are also suppressed. These “cold nuclear matter” effects need to be accounted for before signals of the high density QCD medium can be identified in the measurements made in nucleus-nucleus collisions. We identify two cold nuclear matter effects important for midrapidity quarkonium production: “nuclear absorption”, typically characterized as a final-state effect on the produced quarkonium state and shadowing, the modification of the parton densities in nuclei relative to the nucleon, an initial-state effect. We characterize these effects and study the energy, rapidity, and impact-parameter dependence of initial-state shadowing in this paper.

I. BASELINE TOTAL CROSS SECTIONS

To better understand quarkonium suppression, it is necessary to have a good estimate of the expected yields. However, there are still a number of unknowns about quarkonium production in the primary nucleon-nucleon interactions. In this section, we discuss models of quarkonium production and give predictions for the yields in a number of collision systems.

| A   | $E_A$ (TeV) | $y_A$ | $\sqrt{s_{NN}}$ (TeV) | $y_{diff}^A$ | $\Delta y_{cm}^A$ | $\sqrt{s_{NN}}$ (TeV) | $y_{diff}^{dA}$ | $\Delta y_{cm}^{dA}$ | $\sqrt{s_{NN}}$ (TeV) |
|-----|------------|-------|----------------------|-------------|-----------------|----------------------|-------------|-----------------|------------------|
| O   | 3.5        | 8.92  | 9.9                  | 0.690       | 0.345           | 7                    | 0           | 0               | 7                |
| Ar  | 3.15       | 8.81  | 9.39                 | 0.798       | 0.399           | 6.64                 | 0.052       | 0.026           | 6.3              |
| Kr  | 3.07       | 8.79  | 9.27                 | 0.824       | 0.412           | 6.48                 | 0.077       | 0.038           | 6.14             |
| Sn  | 2.92       | 8.74  | 9.0                  | 0.874       | 0.437           | 6.41                 | 0.087       | 0.043           | 5.84             |
| Pb  | 2.75       | 8.67  | 8.8                  | 0.934       | 0.467           | 6.22                 | 0.119       | 0.059           | 5.5              |

TABLE I: For each ion species at the LHC, we give the maximum beam energy per nucleon and the corresponding beam rapidity. Using the maximum proton or deuteron beam energy: $E_p = 7$ TeV and $y_p = 9.61$; $E_d = 3.5$ TeV and $y_d = 8.92$ respectively, we present the maximum center-of-mass energy per nucleon; rapidity difference, $y_{diff}^A = y_i - y_A$ ($i = p, d$); and center-of-mass rapidity shift, $\Delta y_{cm}^A = y_{diff}^A/2$, for $p + A$, $d + A$ and $A + A$ collisions. Note that there is no rapidity shift in the symmetric $A + A$ case.

Since the LHC can collide either symmetric ($A + A$) or asymmetric ($A + B$) systems, we present results for $p + p$, $p + A$, $d + A$ and $A + A$ collisions. We consider $d + A$ collisions since the $d + A$ center-of-mass energy is closer to the $A + A$ collision energy than top energy $p + A$ collisions. The maximum ion beam energy per nucleon is the proton beam energy, $E_p = 7$ TeV, times the charge-to-mass ratio, $Z/A$, of the ion beam. Thus the maximum deuteron beam energy is half that of the proton beam, $E_d = 3.5$ TeV. The ion beam energies are given on the left-hand side of Table I for five reference nuclei: oxygen, $^{16}$O; argon, $^{40}$Ar; krypton, $^{84}$Kr; tin, $^{116}$Sn; and lead, $^{208}$Pb. Note that we use the average elemental $A$ since a sample may contain an admixture of several isotopes of different $A$.

In addition to the $A + A$ center-of-mass energy, we also show the maximum $p + A$ and $d + A$ per nucleon center-of-mass energies, $\sqrt{s_{NN}} = \sqrt{4E_{p,d}E_A}$. Because $E_{p,d}$ is typically greater than $E_A$, the center-of-mass rapidity can shift away from $y = 0$. The total shift is $y_{diff}^A = y_i - y_A$ ($i = p, d$) while the center of mass shifts by half this amount, $\Delta y_{cm}^A = y_{diff}^A/2$. Table II shows the maximum nucleon-nucleon center-of-mass energy per nucleon, the rapidity difference between the two beams, $y_{diff}^{AA}$, and the center-of-mass shifts for $p + A$ and $d + A$ collisions. (The $Z/A$ ratio is the same for $d$ and O, thus $\Delta y_{diff}^{dO} = 0$.) Only $\sqrt{s_{NN}}$ is given for symmetric $A + A$ collisions since there is no rapidity shift.

If there were no cold nuclear matter effects on the production cross sections at a given energy, the per nucleon cross sections would all be equal. However, the nuclear parton distributions (nPDFs) are known to be modified with...
respect to the free proton PDFs as a function of parton momentum fraction $x$. At low $x$, $x < 0.05$ (shadowing region), and high $x$, $x > 0.2$ (EMC region), the nuclear structure function, $F_2^A(x)$, the weighted sum of the charged parton distributions, is suppressed relative to that of the deuteron, $F_2^D(x)$, while, in the intermediate $x$ region, the ratio $2F_2^A/AF_2^D$ is enhanced (antishadowing) in nuclear deep-inelastic scattering (nDIS). We refer to the modification of the parton densities in the nucleus as a function of $A$, $x$ and $\mu^2$ in general as shadowing. While a combination of nDIS and Drell-Yan data can separate the nuclear valence and sea quark densities, there is no direct probe of the nuclear gluon density, rather it is inferred from the $\mu^2$ scaling violation.

Gluon fusion dominates quarkonium production up to $x_F \sim 0.7$ already at fixed-target energies [1], including over the entire accessible rapidity range at the LHC, see Fig. 1. Thus while the modification of the gluon distributions in nuclei is the most important for quarkonium studies, it is unfortunately the most poorly known. There are, however, a number of indirect constraints on the gluon density. The scale evolution of $F_2^A$ and momentum conservation provide two important constraints. Most of the low-$x$ nDIS data are at relatively low scales, below the minimum scale of a number of PDF sets and therefore less useful for studies of perturbative evolution. RHIC data on hadron production are an exception since intermediate $p_T$ hadron production occurs at relatively low $x$ and at perturbative scales. At relatively high $x$, the shape of the PHENIX midrapidity $\pi^0$ data [2] helps pin down the nuclear gluon density in the EMC region.

Quarkonium production occurs at sufficiently large scales to provide further constraints on the nuclear gluon PDFs. There are some drawbacks however: the quarkonium production mechanism is not fully understood, even in $p+p$ collisions, and the energy dependence of nuclear absorption is not well known. In the remainder of this section, we discuss the quarkonium yields in various collision systems; the implementation of modified PDFs for the nuclear parton densities; and quarkonium absorption by nucleons.

Early studies of high energy quarkonium production, particularly at high $p_T$, were performed in the context of the color singlet model (CSM) which calculates direct production of a quarkonium state with definite total spin, parity and charge conjugation. The CSM predicted that the $\chi_{c1}$ state, produced directly from $gg$ fusion, would have a much larger cross section than direct color-singlet $J/\psi$ production which requires a 3-gluon vertex [3]. Instead, measurements of direct $J/\psi$ and $\chi_{c}$ production showed that the $J/\psi$ cross section was, in fact, larger than the $\chi_{c}$ cross section [4].

Nonrelativistic QCD (NRQCD) is an effective field theory in which short-distance partonic interactions produce $Q\bar{Q}$ pairs in color singlet or color octet states which then evolve into a quarkonium state, as characterized by nonperturbative matrix elements [9]. The first term in the NRQCD expansion is equivalent to the CSM. The octet contributions
are sufficient to explain the $J/\psi$ yield at the Tevatron. However, the NRQCD approach has so far failed to describe quarkonium polarization.

Perhaps the simplest approach to quarkonium production is the color evaporation model (CEM) which treats heavy flavor and quarkonium production on an equal footing. The quarkonium production cross section is some fraction, $F_C$, of all $Q\bar{Q}$ pairs below the $H\bar{H}$ threshold where $H$ is the lowest mass heavy-flavor hadron. Thus the CEM cross section is simply the $Q\bar{Q}$ production cross section with a cut on the pair mass but without any contraints on the color or spin of the final state. The color of the octet $Q\bar{Q}$ state is ‘evaporated’ through an unspecified process which does not change the momentum. The additional energy needed to produce heavy-flavored hadrons when the partonic center of mass energy, $\sqrt{s}$, is less than $2m_H$, the $H\bar{H}$ threshold energy, is nonperturbatively obtained from the color field in the interaction region. Thus the quarkonium yield may be only a small fraction of the total $Q\bar{Q}$ cross section below $2m_H$. At leading order, the production cross section of quarkonium state $C$ in an $A + B$ collision is

$$\frac{d\sigma^{CEM}_{Q\bar{Q}}(s_{NN})}{d^2r d^2b} = F_C \sum_{i,j} \int_{\frac{4m_H^2}{s}} \frac{ds}{4m_H^2} \int dx_1 dx_2 \int dz' dz \times f^A_i(x_1, \mu^2, \bar{r}, z) f^B_j(x_2, \mu^2, \bar{b} - \bar{r}, z') \delta(s - x_1 x_2 s_{NN}) \ ,$$

where $A$ and $B$ can be any hadron or nucleus, $ij = q\bar{q}$ or gg and $\delta(s)$ is the $ij \rightarrow Q\bar{Q}$ subprocess cross section. If one or both of the collision partners, $A$ and $B$, is a proton, then the transverse, $\bar{r}$, and longitudinal, $z$, spatial parameters may be replaced by delta functions, $\int d^2r dz \delta(\bar{r})\delta(z)$, and the parton densities are simply $f^A_i(x_1, \mu^2, \bar{r}, z) \equiv f^A_i(x_1, \mu^2)$. Our calculations use the NLO $Q\bar{Q}$ code of Mangano et al. [11] with the $2m_H$ mass cut in Eq. 1 and use the same parameters as in Refs. [12, 13] with the MRST parton densities [14], optimized to obtain agreement with the $Q\bar{Q}$ cross section, as described in Ref. [13]. The factor $F_C$ can also be fit with other parton densities such as CTEQ6M [16]. When the same mass and scale parameters are used, the energy dependence of the cross section is very similar, see e.g. Ref. [13].

All these formulations: CSM; NRQCD; and CEM assume the validity of collinear factorization which relies on the separation of the initial and final states. Collinear factorization was proven to be effective at all orders for the Drell-Yan process some time ago [17]. A subsequent paper by Collins, Soper and Sterman showed that the factorization process was correct for heavy flavor production up to corrections of order $(1/M)$ [18]. Thus while higher-order corrections to the charm cross section are large, collinear factorization is generally assumed to hold and, indeed, the scale dependence of the approximate NNLO-NNLL charm cross section is seen to stabilize and the next-order corrections are not as large [19]. Higher-twist effects that might signify factorization breaking, such as intrinsic charm, are generally most important at forward rapidities in the light-cone formulation [20, 21]. Factorization has not been strictly proven for quarkonium where the final quarkonium state may be connected to the initial state by soft gluons. In the CSM, the color singlet matrix element is derived from quarkonium decays where the initial state plays no role. Factorization is most difficult to prove for NRQCD. It depends on the universality of the nonperturbative matrix elements. However, recent works have shown that a redefinition of these matrix elements allows factorization to be restored [22, 23]. The CEM is closest in spirit to the calculation of the open heavy flavor cross section so that collinear factorization should work equally well in the two approaches. Since our calculations are in the CEM, we use collinear factorization to calculate quarkonium production at the LHC.

To go beyond $p + p$ collisions, the proton parton densities must be replaced by those of the nucleus. Then the collision geometry and the spatial dependence of the shadowing parameterization also need to be considered. We assume that if $A$ is a nucleus, the nuclear parton densities, $f^A_i(x_1, \mu^2, \bar{r}, z)$, factorize into the nucleon density in the nucleus, $\rho_A(\bar{r}, z)$, independent of the kinematics; the nucleon parton density, $f^N_i(x_1, \mu^2)$, independent of $A$; and a shadowing ratio, $S_{P,S}(A, x_1, \mu^2, \bar{r}, z)$ that parameterizes the modifications of the nucleon parton densities in the nucleus. The first subscript, $P$, refers to the choice of shadowing parameterization, while the second, $S$, refers to the spatial dependence. Thus,

$$f^A_i(x_1, \mu^2, \bar{r}, z) = \rho_A(s) S_{P,S}(A, x_1, \mu^2, \bar{r}, z) f^N_i(x_1, \mu^2) \ ,$$

$$f^B_j(x_2, \mu^2, \bar{b} - \bar{r}, z') = \rho_B(s') S_{P,S}(B, x_2, \mu^2, \bar{b} - \bar{r}, z') f^N_j(x_2, \mu^2) \ ,$$

where $s = \sqrt{\bar{r}^2 + z'^2}$ and $s' = \sqrt{\bar{b} - \bar{r} + z'^2}$.

The nucleon densities of the heavy nucleus are assumed to be Woods-Saxon distributions [24] and are normalized so that $\int d^2b' ds' f^N_i(x_1, \mu^2, \bar{r}, z) f^N_j(x_2, \mu^2, \bar{b} - \bar{r}, z') = AB f^N_i(x_1, \mu^2) f^N_j(x_2, \mu^2)$.
The impact-parameter averaged shadowing parameterization measured in nDIS is recovered by integrating $S_{P,S}$ over the volume, weighted by the nuclear density,

$$\frac{1}{A} \int d^2rdz\rho_A(s)S_{P,S}^A(A, x, \mu^2, \vec{r}, z) = S_{P}^A(A, x, \mu^2). \tag{5}$$

We discuss more details of the spatial dependence of $S_{P,S}$ in Section 2.3. Most available shadowing parameterizations, including the ones used here, ignore the small effects in deuterium. However, we take the proton and neutron numbers of both nuclei into account. The impact-parameter integrated up and down quark distributions, needed for the $q\bar{q}$ contribution to quarkonium production, are calculated as

$$f_q^A(x, \mu^2) = (Z_A S_{P,n}^A(A, x, \mu^2) f_q^p(x, \mu^2) + N_A S_{P,n}^A(A, x, \mu^2) f_q^n(x, \mu^2)) \tag{6}$$

for $q = u$ and $d$, assuming that, as for the proton and neutron parton densities, $S_{P,n}^A = S_{P,p}^A$ and $S_{P,n}^A = S_{P,n}^A$ and similarly for the antiquarks.

To obtain the rapidity distribution from the total cross section, an additional delta function, $\delta(y - 0.5 \ln(x_1/x_2))$, is included in Eq. (1). At leading order, the parton momentum fractions $x_1$ and $x_2$ are simply $x_{1,2} = (\sqrt{s}/s_{NN}) \exp(\pm y)$. In this notation then, in the forward rapidity region of a $p + A$ collision, $x_1$, the proton momentum fraction, is larger and $x_2$, the parton momentum fraction in the nucleus, is smaller than the midrapidity value, $x = \sqrt{s}/s_{NN}$.

Some of the uncertainties in the production model may be overcome by studying ratios, e.g. $(p + A)/(p + p)$, at the same center-of-mass energy since the dominance of $gq$ processes means that the $(p + A)/(p + p)$ ratio is, to a good approximation, the ratio of the gluon distribution in the nucleus relative to the gluon distribution in the proton. We have chosen to use the CEM because it allows predictions of the total cross section and the $p_T$-integrated rapidity distributions where nuclear effects are more prominent. Measuring the $J/\psi$ and $\Upsilon$ ratios simultaneously also provides a means of determining the scale evolution of the nuclear gluon distribution at relatively large, perturbative scales if shadowing is the only cold nuclear matter effect in $p + A$ and $d + A$ collisions.

At fixed-target energies, the $x_F$ dependence clearly shows that shadowing is not the only contribution to the $J/\psi$ nuclear dependence as a function of $x_F$ \cite{23,26}. Indeed, the characteristic decrease of $\alpha(x_F)$ for $x_F \geq 0.25$ cannot be explained by shadowing alone \cite{1}. In fact, the data so far suggest approximate scaling with $x_F$, not the target momentum fraction $x_2$ \cite{27}, indicating the possible importance of higher-twist effects \cite{28}. The preliminary PHENIX data show an increasing suppression at forward rapidity \cite{29}, similar to that seen in fixed-target experiments at large $x_F$.

Effects we have not considered here which may result in $x_F$ rather than $x_2$ scaling and affect the high $x_F$ region are energy loss in cold matter and intrinsic charm, both discussed extensively in Ref. \cite{1}. We do not consider these effects here because, at heavy-ion colliders, the relationship between $x_F$, rapidity, and $\sqrt{s_{NN}}$ suggests that this interesting $x_F$ region is pushed to far forward rapidities. The onset of initial-state energy loss should, in fact, appear at higher $x_F$ at larger $\sqrt{s_{NN}}$ if it depends on the momentum fraction $x_1$. Figure 1 shows the relationship between $x_F$ and $y$ in the center-of-mass frame for $M = 4$ GeV and $20 \leq \sqrt{s_{NN}} \leq 14000$ GeV. Since $x_F = (2m_T/\sqrt{s_{NN}}) \sinh y$, the large center-of-mass energies at the LHC guarantees that the forward $x_F$ region will not be accessible in the central rapidity region of the LHC. Instead, the $x_F$ distribution becomes narrowly peaked with increasing energy while the rapidity distribution becomes broad and flat. At $y = 5$, the largest $x_F$ accessible (at the lowest $\sqrt{s_{NN}}$) is 0.081 for the $J/\psi$ and 0.25 for the $\Upsilon$. The large $x_F$ region is therefore not probed by quarkonium production in $|y| \leq 5$. Thus shadowing and absorption are likely the most important cold nuclear matter effects at the LHC.

To implement nuclear absorption on quarkonium production in $p + A$ and $d + A$ collisions, the production cross section is weighted by the survival probability, $S_{C}^{abs}$, so that

$$S_{C}^{abs}(\vec{b} - \vec{s}, z') = \exp \left\{ - \int_{z'}^{\infty} dz'' \rho_A(\vec{b} - \vec{s}, z'') \sigma_{abs}(z'' - z') \right\} \tag{7}$$

where $z'$ is the longitudinal production point, as in Eq. (3), and $z''$ is the point at which the state is absorbed. The nucleon absorption cross section, $\sigma_{abs}$, typically depends on the spatial location at which the state is produced and how far it travels through the medium.

We could also consider absorption by comover interactions but this cross section is typically smaller than the nuclear absorption cross section. In addition, in $A + A$ collisions, the higher temperatures and larger particle densities would rule out hadronic comovers in the early stages. Therefore, we do not consider hadronic comovers as a cold matter effect in this paper.

If absorption alone is active, i.e. $S_{p,S}(A, x, \mu^2, \vec{r}, z) \equiv 1$, then an effective minimum bias $A$ dependence is obtained after integrating Eqs. (1) and (7) over the spatial coordinates. If $S_{C}^{abs} = 1$ also, $\sigma_{pA} \approx A\sigma_{pp}$ without any cold nuclear
matter effects. (Note that for $gg$-dominated processes, such as quarkonium production, the relationship would be exact. When $q\bar{q}$ or $q\bar{q}$ interactions dominate, as in gauge boson production, the different relative proton and neutron numbers make the above relationship approximate.) If $S_{P,S}^{q}(A,x,\mu^2,\vec{r},z) \equiv 1$ and $S_{C}^{\text{abs}} \neq 1$, $\sigma_{pA} = A^{\alpha} \sigma_{pp}$ where the exponent $\alpha$ can be related to the absorption cross section, as studied in detail for $J/\psi$ and $\psi'$ production by NA50 \cite{30}. For a constant $\sigma_{C}^{\text{abs}}$ with a sharp surface spherical nucleus of density $\rho_{A} = \rho_{0} \theta(R_{A} - b)$, it can be shown that

$$\alpha = 1 - \frac{9\sigma_{C}^{\text{abs}}}{16\pi r_{0}^{2}}$$

(8)

where $r_{0} = 1.2 \text{ fm}$ \cite{31}. The relationship between $\alpha$ and $\sigma_{C}^{\text{abs}}$ is less straightforward in more realistic geometries.

The NA50 \cite{30} and E866 \cite{25} experiments measured a non-negligible difference in the effective $J/\psi$ and $\psi'$ absorption cross sections at $\sqrt{s_{NN}} = 23 - 29 \text{ GeV}$ and $\sqrt{s_{NN}} = 38.8 \text{ GeV}$ respectively. In addition, the difference between $\sigma_{J/\psi}^{\text{abs}}$ and $\sigma_{\psi'}^{\text{abs}}$ seems to decrease with $\sqrt{s_{NN}}$. The NA50 collaboration measured $\Delta \sigma = \sigma_{\psi'}^{\text{abs}} - \sigma_{J/\psi}^{\text{abs}} = 4.2 \pm 1 \text{ mb at } 400 \text{ GeV and } 2.8 \pm 0.5 \text{ mb at } 450 \text{ GeV} \cite{30}$. At $x_{F} \sim 0$, the E866 results imply $\Delta \sigma = \alpha_{J/\psi} - \alpha_{\psi'} < 0.2$ or, using Eq. (8), $\Delta \sigma < 1.6 \text{ mb} \cite{25}$. This suggests that absorption is a final-state effect since an initial-state effect such as shadowing would not discriminate between the asymptotic $J/\psi$ and $\psi'$ final states. Comparing the effective absorption cross sections determined at central rapidities from the CERN SPS to RHIC, absorption seems to decrease with energy \cite{32}.

Fewer $\Upsilon$ $p + A$ data are available. The E772 experiment \cite{33} measured the $A$ dependence of the three $S$ states and found a reduced $A$ dependence relative to $J/\psi$ absorption. The $A$ dependence of the three $S$ states was indistinguishable within the uncertainties. No $\Upsilon$ $A$ dependence was presented by the E866 collaboration. The STAR $d + Au/p + p$ ratio from RHIC suggests that, within large uncertainties, the $\Upsilon$ $A$ dependence is linear \cite{34} and production is not significantly suppressed. Thus absorption seems to be weaker overall for $\Upsilon$ production but there is not clear indication so far of how much weaker it is or whether it has the same energy dependence as the $J/\psi$.

If conventional shadowing parameterizations, such as the ones used in this paper, are included, the effective $J/\psi$ absorption cross section may seem to decrease with energy due to the increased effect of shadowing at low $x$. A decrease in absorption concurrent with increased shadowing as $\sqrt{s_{NN}}$ increases seems to approximately hold, even without shadowing, at fixed-target energies \cite{32}. Such a decrease is consistent with the $J/\psi$ traversing the nucleus as a color singlet. If the nuclear crossing time is shorter than the $J/\psi$ formation time, the effective absorption decreases with $\sqrt{s_{NN}}$ as an ever smaller state passes through the target.

If the effective absorption cross section indeed decreases with energy, then absorption should be a relatively small contribution to the total $A$ dependence at the LHC. This prediction is easy to check: if absorption is negligible, the $J/\psi$ and $\psi'$ $(p + A)/(p + p)$ ratios should depend only on shadowing and should thus be equivalent. The yield is then related to the ratio of the nuclear to proton gluon densities since $gg$ fusion dominates quarkonium production at these energies. In this work, we have assumed that absorption is negligible so that the $(p + A)/(p + p)$ and $(A + A)/(p + p)$ $J/\psi$ and $\Upsilon$ ratios presented here are the same for all charmonium and bottomonium states respectively.

If both the $p + A$ and $p + p$ data are taken at the same $\sqrt{s_{NN}}$, the same $x$ values of the gluon densities will be probed in the nucleus and in the proton. Such energy comparison runs would be an excellent probe of the nuclear gluon distributions because

$$\frac{\sigma_{pA}(\sqrt{s_{NN}})}{\sigma_{pp}(\sqrt{s_{NN}})} \propto \frac{1}{A} \frac{f^{A}_{g}(x,\mu^{2})}{f^{p}_{g}(x,\mu^{2})}.$$  

(9)

However, if the $p + A$ and $p + p$ data are recorded at different energies (and $x$ values), the extraction of the nuclear gluon density is less straightforward since

$$\frac{\sigma_{pA}(\sqrt{s})}{\sigma_{pp}(\sqrt{s})} \propto \frac{1}{A} \frac{f^{A}(x',\mu^{2})}{f^{p}(x,\mu^{2})}.$$  

(10)

In both cases, the $p_{T}$-integrated ratios provide an additional uncertainty because the scale evolution of the gluon density is not well known but is expected to be strong \cite{35,40}. However, the quarkonium $p_{T}$ distribution is steeply falling for $p_{T} \geq m$ so that the $p_{T}$-integrated ratios ratios are a good representation of $\mu^{2} = \langle m_{T}^{2} \rangle$.

The scale evolution of the gluon densities can be probed in part by relative studies of low $p_{T}$ or $p_{T}$-integrated $J/\psi$ ($m_{\psi} = 3.097$ GeV) and $\Upsilon(1S)$ ($m_{\Upsilon(1S)} = 9.46$ GeV) production. To more precisely obtain the scale evolution of shadowing, it would be preferable to bin the $J/\psi$ and $\Upsilon(1S)$ $(p + A)/(p + p)$ ratios in $p_{T}$. One must be careful in the interpretation of such ratios, particularly at $p_{T} < m$, since, at fixed-target energies, the $p_{T}$-dependent $(p + A)/(p + p)$ ratios show that the $J/\psi$ and $\Upsilon$ $p_{T}$ distributions are broader in $p + A$ than in $p + p$ interactions \cite{31,42}. This broadening
has been attributed to intrinsic parton $p_T$ kicks accrued by the interacting parton as it traverses the nucleus before interacting \cite{43,44}. The magnitude of the average $p_T$ kick increases with $A$ so that the $p_T$-dependent $(p + A)/(p + p)$ ratio is less than unity at low $p_T$ and increases above one with increasing $p_T$. This effect is important at low center-of-mass energies where the average $p_T$ of the produced quarkonium state is not large. By LHC energies, while the $p_T$ kick may be rather small relative to $\langle p_T^2 \rangle$, it may still affect the analysis of shadowing effects in $p_T$-binned ratios but not in $p_T$-integrated ratios. We focus on the $p_T$-integrated results here and will present $p_T$-dependent calculations elsewhere.

| System | $\sqrt{s_{NN}}$ (TeV) | $\sigma^{dir}$/nucleon pair (\(\mu b\)) | $B\sigma^{inc}AB$ (\(\mu b\)) |
|--------|-----------------|-------------------------------|-----------------|
| $p + p$ | 14              | 32.9 31.8 52.5 7.43           | 3.15 0.055      |
| $p + p$ | 10              | 26.8 26.0 43.3 6.06           | 2.57 0.044      |
| $p + p$ | 9.9             | 26.6 25.8 42.6 6.02           | 2.55 0.044      |
| $p + O$ | 9.9             | 23.8 23.0 38.0 5.37           | 36.5 0.632      |
| $p + p$ | 9.39            | 25.8 25.0 41.3 5.83           | 2.48 0.043      |
| $p + O$ | 9.39            | 22.0 21.2 35.1 4.96           | 84.1 1.46       |
| $p + p$ | 9.27            | 25.6 24.8 40.9 5.79           | 2.46 0.043      |
| $p + O$ | 9.27            | 20.9 20.2 33.4 4.73           | 168.4 2.92      |
| $p + p$ | 9               | 25.2 24.4 40.2 5.69           | 2.41 0.042      |
| $p + p$ | 8.8             | 20.2 19.6 32.3 4.56           | 230.4 3.99      |
| $p + p$ | 7               | 25.0 24.2 39.9 5.65           | 2.40 0.042      |
| $p + p$ | 7               | 19.5 18.9 31.1 4.40           | 388.8 6.75      |
| $p + p$ | 7               | 21.8 21.1 34.9 4.93           | 2.09 0.036      |
| $p + O$ | 7               | 19.5 19.0 31.3 4.42           | 30.0 0.520      |
| $d + O$ | 7               | 19.5 19.0 31.3 4.42           | 60.0 1.04       |
| $O + O$ | 7               | 17.6 17.0 28.1 3.98           | 432.4 7.51      |
| $p + p$ | 6.64            | 21.2 20.5 33.8 4.78           | 2.02 0.035      |
| $d + Ar$ | 6.64           | 18.1 17.5 28.9 4.09           | 138.5 2.39      |
| $p + p$ | 6.48            | 20.9 20.2 33.3 4.71           | 2.00 0.034      |
| $d + Kr$ | 6.48           | 17.2 16.6 28.0 3.95           | 281.3 4.85      |
| $p + p$ | 6.41            | 20.7 20.1 33.1 4.68           | 1.98 0.034      |
| $d + Sn$ | 6.41           | 16.8 16.2 26.8 3.78           | 378.3 6.52      |
| $p + p$ | 6.3             | 20.5 19.9 32.8 4.63           | 1.97 0.034      |
| $p + Ar$ | 6.3            | 17.6 17.0 28.1 3.97           | 67.3 1.17       |
| $Ar + Ar$ | 6.3           | 15.0 14.5 23.9 3.38           | 2300 40.0       |
| $p + p$ | 6.22            | 20.4 19.7 32.5 4.60           | 1.95 0.34       |
| $d + Pb$ | 6.22           | 16.0 15.5 25.6 3.62           | 637.3 10.98     |
| $p + p$ | 6.14            | 20.2 19.6 32.3 4.56           | 1.94 0.034      |
| $p + Kr$ | 6.14           | 16.6 16.1 26.6 3.76           | 134.0 2.32      |
| $Kr + Kr$ | 6.14          | 13.7 13.2 21.8 3.08           | 9245 160.6      |
| $p + p$ | 5.84            | 19.6 19.0 31.3 4.42           | 1.88 0.033      |
| $p + Sn$ | 5.84           | 15.9 15.4 25.4 3.59           | 181.3 3.14      |
| $Sn + Sn$ | 5.84         | 12.8 12.4 20.4 2.89           | 17391 302.0     |
| $p + p$ | 5.5             | 18.9 18.3 30.2 4.26           | 1.81 0.032      |
| $p + Pb$ | 5.5            | 14.9 14.4 23.8 3.37           | 297.6 5.16      |
| $Pb + Pb$ | 5.5           | 11.7 11.3 18.7 2.64           | 48500 842       |

TABLE II: The direct cross section per nucleon pair (central columns) and the dilepton yield per nucleon multiplied by $AB$. The results are given for the MRST PDFs \cite{14} with $m_c = 1.2$ GeV, $\mu_F = \mu_R = 2m_T$.

As an example of the possible cross sections for quarkonium production at the LHC, we present the total cross sections in $p + p$, $p + A$, $d + A$ and $A + A$ collisions at the relevant energies. To illustrate the effects of shadowing on the total cross section, calculated to next-to-leading order in the CEM \cite{15}, we use the EKS98 parameterization \cite{37,38}. 
For each possible maximum $p + A$, d+$A$ and $A + A$ center-of-mass energy, we also give the $p + p$ cross section at that same energy. In addition, for the $A + A$ center-of-mass energies, we also give the $p + p$ and $p + A$ cross sections at that energy. The results are given in Tables III and IV. The central columns are the direct cross sections per nucleon pair for all charmonium and bottomonium states. The effects are largest for charmonium (lower $x$ and $\mu^2$ than the $\Upsilon$ states) and for the heaviest nuclei (lowest energies – highest $x$ – but largest $A$). On the right-hand side of the tables, the inclusive (direct plus feed down) cross sections are multiplied by the dilepton decay branching ratios. They are also multiplied by $AB$ to obtain the minimum bias total cross sections.

The approximate $A$ dependence of the total cross section relative to the $p + p$ cross section at the same center-of-mass energy, assuming no other cold matter effects, can be obtained from the $(AB)^{\alpha}$ parameterization so that, per nucleon,

$$\alpha(p + A/p + p) \sim 1 + \frac{\ln[f_g^A(x_2, \mu^2)/f_g^B(x_2, \mu^2)]}{\ln A} \tag{11}$$

$$\alpha(A + B/p + p) \sim 1 + \frac{\ln[f_g^A(x_1, \mu^2)f_g^B(x_2, \mu^2)/(f_g^P(x_1, \mu^2)f_g^P(x_2, \mu^2))]}{\ln(AB)} \tag{12}$$

where $x_2' = x_2$ and $x_1' = x_1$ if the center-of-mass energies are the same for the two systems. For $J/\psi$ production in $p$Pb and Pb+Pb collisions relative to $p + p$ collisions at 5.5 TeV, $\alpha \sim 0.76$ and 0.52 respectively. In the case of $\Upsilon$ production, we have $\alpha \sim 0.88$ and 0.76 respectively.

The $p + p$ rapidity distributions for $J/\psi$ and $\Upsilon$ production at $\sqrt{s_{NN}} = 5.5$ and 14 TeV are compared to RHIC distributions at $\sqrt{s_{NN}} = 200$ and 500 GeV in Fig. 2. The LHC distributions are relatively constant over a range of 5 or more units of rapidity, demonstrating that the cross sections are high enough to obtain good statistics for quarkonium states, even for forward production and detection, provided that the decay leptons are of sufficiently high $p_T$ to reach the detectors.$^1$

![Graph](image_url)

**FIG. 2:** (Color online) The $J/\psi$ (left-hand side) and $\Upsilon$ (right-hand side) rapidity distributions at $\sqrt{s_{NN}} = 200$ (dotted), 500 (dot-dashed), 5500 (dashed) and 14000 (solid) GeV calculated as in Refs. [12, 13]. The kinks in the $J/\psi$ distributions at LHC energies are the point where $x < 10^{-5}$. Since the $\Upsilon$ factorization scale is larger, the $\Upsilon$ rapidity distributions are smoother. Note the different scales on the $y$-axes.

$^1$ This will be more difficult for CMS and ATLAS than for ALICE since the minimum muon $p_T$ for detection in the large $p + p$ experiments is 3.5 GeV/c.
| System  | $\sqrt{s_{NN}}$ (TeV) | $\sigma^\text{dir}$/nucleon pair ($\mu$b) | $\text{B} \sigma^\text{inc} AB$ ($\mu$b) |
|---------|----------------------|---------------------------------------|----------------------------------|
| $p + p$ | 14                   | 0.43 0.27 0.16 0.89 0.69              | 0.020 0.0074 0.0036              |
| $p + p$ | 10                   | 0.33 0.21 0.12 0.70 0.54              | 0.016 0.0059 0.0028              |
| $p + p$ | 9.9                  | 0.32 0.20 0.12 0.66 0.51              | 0.015 0.0055 0.0026              |
| $p + O$ | 9.9                  | 0.30 0.19 0.11 0.62 0.48              | 0.23 0.082 0.040                |
| $p + p$ | 9.39                 | 0.30 0.19 0.12 0.63 0.49              | 0.014 0.0052 0.0025              |
| $p + Ar$| 9.39                 | 0.28 0.17 0.11 0.57 0.44              | 0.53 0.19 0.092                 |
| $p + p$ | 9.27                 | 0.30 0.19 0.11 0.62 0.48              | 0.014 0.0052 0.0025              |
| $p + Kr$| 9.27                 | 0.27 0.17 0.10 0.55 0.43              | 1.06 0.39 0.19                  |
| $p + p$ | 9                    | 0.29 0.18 0.11 0.61 0.47              | 0.014 0.0050 0.0024              |
| $p + Sn$| 9                    | 0.26 0.16 0.099 0.53 0.42             | 1.46 0.53 0.26                  |
| $p + p$ | 8.8                  | 0.29 0.18 0.11 0.60 0.47              | 0.014 0.0059 0.0024              |
| $p + Pb$| 8.8                  | 0.25 0.16 0.097 0.52 0.41             | 2.51 0.96 0.45                  |
| $p + p$ | 7                    | 0.23 0.15 0.090 0.48 0.38              | 0.011 0.0043 0.0019              |
| $p + O$ | 7                    | 0.22 0.14 0.085 0.46 0.36              | 0.17 0.061 0.029                 |
| d+O    | 7                    | 0.22 0.14 0.085 0.46 0.36              | 0.34 0.12 0.058                 |
| O+O    | 7                    | 0.21 0.13 0.081 0.44 0.34              | 2.57 0.97 0.46                  |
| $p + p$ | 6.64                 | 0.22 0.14 0.085 0.46 0.36              | 0.011 0.0038 0.0019              |
| d+Ar   | 6.64                 | 0.20 0.13 0.079 0.42 0.33              | 0.78 0.27 0.13                  |
| $p + p$ | 6.48                 | 0.22 0.14 0.083 0.45 0.35              | 0.010 0.0037 0.0018              |
| d+Kr   | 6.48                 | 0.20 0.12 0.076 0.41 0.32              | 1.57 0.56 0.28                  |
| $p + p$ | 6.41                 | 0.21 0.14 0.082 0.44 0.35              | 0.010 0.0036 0.0018              |
| d+Sn   | 6.41                 | 0.19 0.12 0.074 0.40 0.34              | 2.34 0.77 0.41                  |
| $p + p$ | 6.3                  | 0.21 0.14 0.082 0.44 0.34              | 0.010 0.0038 0.0018              |
| p+Ar   | 6.3                  | 0.20 0.12 0.075 0.41 0.32              | 0.37 0.13 0.065                 |
| Ar+Ar  | 6.3                  | 0.18 0.12 0.070 0.38 0.29              | 13.8 5.29 2.43                  |
| $p + p$ | 6.22                 | 0.21 0.13 0.080 0.43 0.34              | 0.010 0.0035 0.0017              |
| d+Pb   | 6.22                 | 0.18 0.12 0.071 0.38 0.30              | 3.68 1.31 0.65                  |
| $p + p$ | 6.14                 | 0.21 0.13 0.080 0.43 0.33              | 0.0099 0.0038 0.0017             |
| p+Kr   | 6.14                 | 0.19 0.12 0.072 0.39 0.30              | 0.75 0.27 0.13                  |
| Kr+Kr  | 6.14                 | 0.17 0.11 0.066 0.35 0.28              | 57.4 21.8 10.1                 |
| $p + p$ | 5.84                 | 0.20 0.12 0.076 0.41 0.32              | 0.0094 0.0035 0.0017             |
| p+Sn   | 5.84                 | 0.18 0.11 0.068 0.37 0.29              | 1.01 0.36 0.18                  |
| Sn+Sn  | 5.84                 | 0.16 0.10 0.062 0.33 0.26              | 108.1 41.3 19.0               |
| $p + p$ | 5.5                  | 0.19 0.12 0.070 0.39 0.30              | 0.0090 0.0029 0.0016             |
| p+Pb   | 5.5                  | 0.17 0.11 0.064 0.34 0.27              | 1.65 0.60 0.29                  |
| Pb+Pb  | 5.5                  | 0.15 0.094 0.057 0.31 0.24              | 304 116.1 53.5               |

TABLE III: The direct cross section per nucleon pair (central columns) and the dilepton yield per nucleon multiplied by AB. The results are given for the MRST PDFs \[14\] with $m_b = 4.75$ GeV, $\mu_F = \mu_R = m_T$.

II. COLD NUCLEAR MATTER EFFECTS

In this section, we describe the cold nuclear matter effect of initial-state shadowing on $J/\psi$ and $\Upsilon$ production at the LHC. We first discuss the shadowing parameterizations used in our calculations. We then show the effect of shadowing on the rapidity distributions in the $p + A$, $d + A$ and $A + A$ collisions available at the LHC. Finally, we discuss the collision centrality dependence on a simple model of inhomogeneous shadowing where the effect depends on the path length of the parton through the nucleus.
A. Shadowing parameterizations

We use several parameterizations of the nuclear modifications in the parton densities to probe the possible range of gluon shadowing effects: EKS98 [35, 36], nDSg [37], HKN [38], EPS08 [39] and EPS09 [40]. All sets involve fits to data, typically nDIS data with additional constraints from other observables such as Drell-Yan dimuon production. Since these provide no direct constraint on the nuclear gluon density, it is obtained through fits to the $\mu^2$ dependence of the nuclear structure function, $F_2^A$, and momentum conservation. The useful perturbative $\mu^2$ range of the nDIS data is rather limited since these data are only available at fixed-target energies. Thus the reach in momentum fraction, $x$, is also limited and there is little available data for $x < 10^{-2}$ at perturbative values of $\mu^2$. This situation is likely not to improve until an $eA$ collider is constructed [42].

The EKS98 parameterization, by Eskola and collaborators, available for $A > 2$, is a leading order fit using the GRV LO [46] proton parton densities as a baseline [35, 36]. The kinematic range is $2.25 \leq \mu^2_{\text{EKS98}} \leq 10^4$ GeV$^2$ and $10^{-6} \leq x < 1$. deFlorian and Sassot produced the nDS and nDSg parameterizations [37] at both leading and next-to-leading order for $4 < A < 208$. The weak gluon shadowing of the nDS parameterization appears to be ruled out by the rapidity dependence of $J/\psi$ production at RHIC [47]. The stronger gluon shadowing of nDSg is used here. Calculations with the nDS parameterization predict negligible shadowing effects. The kinematic reach in $x$ is the same as EKS98 while the $\mu^2$ range is larger, $1 < \mu^2_{\text{nDSg}} < 10^6$ GeV$^2$. Hirai and collaborators produced the leading order HKN parameterization by fitting parton densities for protons, deuterons and 16 heavier nuclei, typically those most commonly used in nDIS experiments. If a particular value of $A$ needed for our calculations is not included, a set with a similar value of $A$ is substituted. The HKN parameterization goes lower in $x$ than the other parameterizations, $10^{-9} < x < 1$, and higher in scale, $1 < \mu^2_{\text{HKN}} < 10^8$ GeV$^2$. The EPS08 parameterization, a fit by Eskola and collaborators that includes the BRAHMS d+Au data on forward rapidity hadron production at RHIC [48], is designed to maximize the possible gluon shadowing\(^2\). The EPS08 $x$ range is the same as EKS98, $10^{-6} \leq x < 1$, while the $\mu^2$ range was extended, $1.96 \leq \mu^2_{\text{EPS08}} \leq 10^6$ GeV$^2$. Very recently, the EPS09 [10] parameterization, which excludes the BRAHMS data from the fits, was introduced. The EPS09 parameterization includes uncertainties on the global analyses, both at LO and NLO, by varying one of the 15 fit parameters within its extremes while holding the others fixed. The upper and lower bounds on EPS09 shadowing are obtained by adding the resulting uncertainties in quadrature [40]. The EPS09 central LO results are in quite good agreement with the older EKS98 parameterization while the maximum possible gluon shadowing effect resulting from their uncertainty analysis is similar to the EPS08 gluon ratio. The minimal amount of gluon shadowing is nearly negligible, similar to nDS [37] and even leaves room for some antishadowing in light ions. We present the central EPS09 ratio as well as the ratios corresponding to the maximum and minimum range of the shadowing effect, obtained by adding the relative differences in quadrature, as prescribed in Ref. [10]. For computational convenience, we use the LO version of the nPDF parameterizations since the NLO CEM calculations give similar shadowing results [50]. This is to be expected since, even though the LO and NLO values of the cross section and the shadowing parameterization are different, when convoluted, they give the same ratios by design, see e.g. Ref. [37].

While the $x$ values probed at midrapidity are $\approx 10^{-4}$ for the $J/\psi$ and $\approx 10^{-3}$ for the $\Upsilon$, well within the $x$ range of the parameterizations, this is not necessarily the case away from midrapidity. At the largest values of $\sqrt{s_{NN}}$, $x$ values lower than the minimum valid $x$ of the parameterization may be reached within the rapidity range of the LHC detectors. In these cases, the shadowing parameterizations are unconstrained by data. However, when $x < 10^{-6}$ the EKS98, nDSg, EPS08 and EPS09 parameterizations return the value of the shadowing ratio at $x = 10^{-6}$. The minimum $x$ value, $10^{-9}$, for the HKN parameterization is small enough that this minimum is not reached, even for the highest energies.

The ratios of the nuclear gluon densities relative to the gluon density in the proton are shown in Fig. 3 for four different ion species available at the LHC: $A = $ O, Ar, Sn and Pb. The calculations for $A = $ Kr, an alternative intermediate mass ion species, are not shown. Results for scales appropriate for $J/\psi$, Fig. 3(a), and $\Upsilon$, Fig. 3(b), production illustrate the scale dependence of the parameterizations. The scales correspond to those used in the calculations of the cross sections in Tables I and II with $\mu = 2m_c$ for charm and $m_b$ for bottom respectively. If a lower scale, $\mu = m_c$, is used for charm, the shadowing effect is stronger since $\mu^2$ is then closer to the minimum scale of the parameterization. Note that in all cases shadowing increases with decreasing $x$ and increasing $A$ while decreasing with scale, $\mu$, as seen by comparing Fig. 3(a) and (b). For example, the EKS98, nDSg and HKN ratios appear to be approximately independent of $x$ for $x < 10^{-3}$ at the $J/\psi$ scale but not at the $\Upsilon$ scale.

The EKS98, EPS08 and EPS09 parameterizations (solid and dotted curves and solid curves with symbols respec-

\(^2\) It has been suggested that the BRAHMS data should not be used to calculate gluon shadowing effects and that the strong shadowing of EPS08 violates a unitarity bound at the minimum scale [49].
FIG. 3: (Color online) The LO shadowing parameterizations for \(J/\psi\) (a) and \(\Upsilon\) (b) scales for O (upper left), Ar (upper right), Sn (lower left) and Pb (lower right) nuclei. The parameterizations are EKS98 (solid), nDSg (dashed), HKN (dot-dashed), EPS08 (dotted) and EPS09 (solid lines with symbols). Note that the lower limit on the y-axis is changed for Sn and Pb on the left-hand side.

tively) exhibit large antishadowing, \(S^g > 1\), in the region \(0.02 < x < 0.2 - 0.3\), becoming more pronounced for larger \(A\). The nDSg parameterization (dashed curves) show very weak antishadowing around \(x \sim 0.1\). At \(x < 10^{-2}\), the nDSg ratios are weakest for \(A = O\) and Ar, similar to HKN for \(A = Sn\) and compatible with EKS98 for \(x < 10^{-3}\). The HKN parameterization (dot-dashed curves), on the other hand, is similar to EKS98 for \(A = O\) but has a weak \(A\) dependence so that HKN shadowing is the weakest at low \(x\) and large \(A\). The EPS08 parameterization is similar to EKS98 for \(x > 0.01\) but exhibits stronger antishadowing at large \(A\). It also has the strongest shadowing at low \(x\) since the low-\(p_T\) forward-rapidity BRAHMS data was included in the fit. The scale dependence of nDSg and HKN appears to be weaker than EKS98. The EPS09 band is obtained by calculating the deviations from the central value for the 15 parameter variations on either side of the central set and adding them in quadrature. The range of the LO EPS09 uncertainty band encompasses all other shadowing ratios, similar to EPS08 for the maximum effect and even leading to antishadowing for lighter ions. (The central ratio is shown with circular symbols on the solid curve while the bounds include diamond symbols.) For smaller nuclei, the upper edge of the EPS09 uncertainty (minimal shadowing effect) gives a bound above unity for \(S^g\).

All the parameterizations increase at large \(x\) with \(S^g > 1\) for \(x > 0.1\) (HKN and nDSg) and \(x > 0.7\) (EKS98 and EPS08). The rise in the HKN parameterization is steepest and occurs at the lowest \(x\), beginning at the \(x\) value of the antishadowing peak in the EKS98 and EPS08 ratios. This high \(x\) region will not be explored by the LHC detectors.
since it is only reached at rapidities outside their acceptance.

Finally, we note that since our $p + A$ calculations assume the ion beam travels in the negative $z$ direction, low $x$ corresponds to large forward rapidity while high $x$ corresponds to large backward rapidity.

![Graphs showing $J/\psi$ and $\Upsilon$ production ratios as a function of rapidity for cold nuclear matter, CNM, effects at the LHC.](image)

**FIG. 4:** (Color online) The $(p + A)/(p + p)$ ratios with both $p + A$ and $p + p$ collisions at the $p + A$ energy in the equal-speed frame. No rapidity shift has been taken into account. The effect of shadowing on $J/\psi$ (a, upper 4 panels) and $\Upsilon$ (b, lower 4 panels) production is shown. Each panel displays the production ratios for $p + O$ at $\sqrt{s_{NN}} = 9.9$ TeV (upper left), $p + Ar$ at $\sqrt{s_{NN}} = 9.39$ TeV (upper right), $p + Sn$ at $\sqrt{s_{NN}} = 9$ TeV (lower left) and $p + Pb$ at $\sqrt{s_{NN}} = 8.8$ TeV (lower right). The calculations are with CTEQ6 and employ the EKS98 (solid), nDSg (dashed), HKN (dot-dashed) and EPS09 (solid curves with symbols) shadowing parameterizations.

### B. Rapidity dependence

We now show predictions of the $J/\psi$ and $\Upsilon$ production ratios as a function of rapidity for cold nuclear matter, CNM, effects at the LHC. If $h + A$ data (where $h = p$ or d) can be taken at the same energy as the $p + p$ and/or $A + A$ data, as at RHIC, it is easier to make comparisons. However, the setup of the LHC makes this ideal situation more difficult. At the nominal injection energy, the proton beam has an energy of 7 TeV while the nuclear beam energy per nucleon is lower by the nuclear charge-to-mass ratio, $Z/A$. To make a $p + p$ comparison, if we are not to rely on calculations extrapolated to lower energy, the $p + p$ collisions have to be run at the $p + A$ or $A + A$ per nucleon energies. For the proton and ion beam energies to be the same, the proton beam must then circulate at lower than optimal energy, decreasing the luminosity. Since sustained low energy $p + p$ runs are unlikely in early LHC running, especially for sufficiently accurate quarkonium data as a function of rapidity, it may be necessary to rely on higher...
energy $p + p$ reference data\(^3\). However, there is a catch. In $p + A$ collisions where a 7 TeV proton beam collides with a 7($Z/A$) TeV per nucleon ion beam, the so-called equal-speed or equal-rigidity frame, the center-of-mass rapidity is not fixed at $y = 0$ but displaced by $\Delta y_{cm}^A$. In pPb collisions, the shift can be nearly 0.5 units, an important difference, see Table II for the magnitude of the possible shifts. To minimize the rapidity shift and to bring the $hA$ comparison energy closer to that of the $A + A$ energy, $d + A$ collisions may be desirable since $E_d = 3.5$ TeV per nucleon relative to $E_{Pb} = 2.75$ TeV, see Table I. Since $d + A$ collisions require a second ion source, this may not be realized in the short term.

We thus study several different possibilities for determining cold nuclear matter effects on nucleus-nucleus collisions at the LHC. We go from ideal to more realistic scenarios. We first show the $(p + A)/(p + p)$ per nucleon ratios at the same per nucleon center-of-mass energy for both systems, assuming the appropriate $p + p$ energies are available, Figs. [4] and [5]. In the case where $p + A$ and $p + p$ interactions are compared at the $A + A$ energy, we assume zero rapidity gap, $\Delta y_{cm}^A = 0$, between the colliding beams since the proton beam energy is reduced to match that of the nucleus. In the more likely scenario, Figs. [6] and [7] the $p + A$ data will be taken in the equal-speed frame at a higher energy than the $A + A$ collisions. Therefore, we next show the $(p + A)/(p + p)$ per nucleon ratios with respect to $p + p$ collisions at $\sqrt{s} = 14$ TeV with $\Delta y_{cm}^A = 0$ for $p + A$ collisions both in the equal-speed frame and at the $A + A$ center-of-mass energy. The final $p + A$ calculations shown are the most realistic: the $p + A$ cross section in the equal-speed frame with finite $\Delta y_{cm}^A$ is shown relative to the $p + p$ cross section at 14 TeV. In this case, Fig. 8 the numerator and denominator are calculated with different energies and different center-of-mass rapidities. Next, the $(d+A)/(p+p)$ per nucleon ratios are presented for two cases: with the $d + A$ and $p + p$ collisions at the $d + A$ center-of-mass energy and with $d + A$ collisions in the equal-speed frame with $\Delta y_{cm}^A \neq 0$ and $p + p$ collisions at 14 TeV. Finally, we present the baseline $(A + A)/(p + p)$ per nucleon ratios with the $p + p$ center-of-mass energy tuned to the $A + A$ energy and at the nominal 14 TeV $p + p$ energy. In the case of symmetric $p + p$ and $A + A$ collisions, there is no rapidity gap.

We present the rapidity dependence of $p + A$, $d + A$, and $A + A$ collisions for $A = O$, Ar, Sn and Pb relative to $p + p$ collisions, both at the same energy as the nuclear system and at 14 TeV. The $A + B/p + p$ ratios are shown for the EKS98, nDSg and HKN parameterizations. The EPS09 central ratios and the associated uncertainty bands are also shown. Since the EPS08 ratios are similar to the lower limit (strongest shadowing) of the EPS09 uncertainty band at small $x$, we do not show any further calculations with EPS08. We use the CTEQ6 parton densities to calculate the ratios shown in the next two sections. We have checked that the ratios with the MRST densities are essentially identical since the same mass and scale parameters are used in the two calculations\(^4\). Both $J/\psi$ and $\Upsilon$ results are shown in each figure. To guide the reader and clarify the discussion, we first present a table of the figures with the center-of-mass energy of the $A + B$ and $p + p$ collisions and the rapidity shift.

\[
\begin{array}{|c|c|c|c|c|c|c|c|c|}
\hline
\text{Figure number} & \text{Cross Section Ratio} & \sigma_{pp} & \sigma_{BA} & \sqrt{s_{HO}} (\text{TeV}) & \sqrt{s_{BA}} (\text{TeV}) & \sqrt{s_{BSg}} (\text{TeV}) & \sqrt{s_{BHV}} (\text{TeV}) & \Delta y_{cm}^A \\
\hline
\text{ } & (B + A)/(p + p) & \sqrt{s_{pp}} & \sqrt{s_{BA}} & \sigma_{pp} & \sqrt{s_{pp}} & \sqrt{s_{BA}} & \sigma_{pp} & \sqrt{s_{pp}} \\
\hline
4 & \sigma_{pA}(\sqrt{s_{PA}}, y)/[\Lambda \sigma_{pp}(\sqrt{s_{PP}}, y)] & \sqrt{s_{PA}} & 9.9 & 9.39 & 9 & 8.8 & 0 \\
5 & \sigma_{pA}(\sqrt{s_{PA}}, y)/[\Lambda \sigma_{pp}(\sqrt{s_{PP}}, y)] & \sqrt{s_{PA}} & 7 & 6.3 & 5.84 & 5.5 & 0 \\
6 & \sigma_{pA}(\sqrt{s_{PA}}, y)/[\Lambda \sigma_{pp}(\sqrt{s_{PP}}, y)] & \sqrt{s_{PA}} & 9.9 & 9.39 & 9 & 8.8 & 0 \\
7 & \sigma_{pA}(\sqrt{s_{PA}}, y)/[\Lambda \sigma_{pp}(\sqrt{s_{PP}}, y)] & \sqrt{s_{PA}} & 14 \text{ TeV} & 7 & 6.3 & 5.84 & 5.5 & 0 \\
8 & \sigma_{pA}(\sqrt{s_{PA}}, (y - \Delta y_{cm}^A))/[\Lambda \sigma_{pp}(\sqrt{s_{PP}}, y)] & \sqrt{s_{PA}} & 14 \text{ TeV} & 9.9 & 9.39 & 9 & 8.8 & \Delta y_{cm}^A \\
\hline
B = d & & & & & & & & & \\
9 & \sigma_{dA}(\sqrt{s_{dA}}, y)/[\Lambda \sigma_{pp}(\sqrt{s_{PP}}, y)] & \sqrt{s_{dA}} & 7 & 6.64 & 6.41 & 6.62 & 0 & \Delta y_{cm}^A \\
10 & \sigma_{dA}(\sqrt{s_{dA}}, (y - \Delta y_{cm}^A))/[\Lambda \sigma_{pp}(\sqrt{s_{PP}}, y)] & \sqrt{s_{dA}} & 14 \text{ TeV} & 7 & 6.64 & 6.41 & 6.62 & \Delta y_{cm}^A \\
\hline
B = A & & & & & & & & & \\
11 & \sigma_{AA}(\sqrt{s_{AA}}, y)/[\Lambda^2 \sigma_{pp}(\sqrt{s_{PP}}, y)] & \sqrt{s_{AA}} & 7 & 6.3 & 5.84 & 5.5 & 0 & & \Delta y_{cm}^A \\
12 & \sigma_{AA}(\sqrt{s_{AA}}, y)/[\Lambda^2 \sigma_{pp}(\sqrt{s_{PP}}, y)] & \sqrt{s_{AA}} & 14 \text{ TeV} & 7 & 6.3 & 5.84 & 5.5 & 0 & \\
\hline
\end{array}
\]

\(^3\) Although the startup LHC $p + p$ runs are at lower energies, it is not clear how much quarkonium data will be extracted during these runs. Thus we show our results relative to the maximum $p + p$ energy of 14 TeV.

\(^4\) The ratios would only differ if these parameters were changed. However, we leave them fixed since they were optimized to other data for a given set of parton densities.
TABLE IV: Summary of the contents of Figs. 4 - 12 in Section II B. Here B is the identity of the collision partner, B = p for p + A, d for d + A and A for A + A collisions. The value of the center-of-mass energy for p + p collisions used in the calculation of the baseline p + p cross section is given in the third column: $\sqrt{s_{NN}} = \sqrt{s_{pp}}$ for p + A; $\sqrt{s_{dA}}$ for d + A; $\sqrt{s_{AA}}$ for A + A; and 14 TeV for maximum energy p + p collisions. The center-of-mass energy for the B + A cross sections are given in the next four columns. Finally, whether or not the rapidity shift is included is indicated in the last column. The value of $\Delta y_{\text{cm}}^{BA}$ is given in Table I. Note that all ratios are given for the per nucleon B + A cross section.

The (p + A)/(p + p) ratios with equal p + A and p + p center-of-mass energies, shown in Figs. 4 and 5, illustrate the direct shadowing effect. The ratios are given both at the energy in the equal-speed frame, the likely $\sqrt{s_{NN}}$ for p + A collisions (Fig. 4), and at the same $\sqrt{s_{NN}}$ as the corresponding A + A collisions (Fig. 5). The results are shown for all shadowing parameterizations. The nuclear beam is assumed to be moving from positive to negative rapidity so that the smallest values of $x$ probed in the nucleus are at large, positive $y$.

The LHC could be run as either a p + A or an A + p collider. Since the ATLAS and CMS detectors are symmetric around $y = 0$ with central muon detectors in the range $|y| \leq 2.4$, ALICE is the only experiment that could benefit from running in both modes because their dimuon spectrometer covers $-4 < y < -2.4$ in these coordinates [51]. However, since ALICE has muon coverage in the largest $y$ region, running in both modes could be an advantage for reconstructing the nuclear effects in quarkonium measurements, especially since the $y$ distributions are rather flat over a broad rapidity range. The large rapidity rates are thus non-negligible.

FIG. 5: (Color online) The (p + A)/(p + p) ratios with both p + A and p + p collisions at the A + A center-of-mass energy. The effect of shadowing on $J/\psi$ (a, upper 4 panels) and $\Upsilon$ (b, lower 4 panels) production is shown. Each set of panels displays the production ratios for p + O at $\sqrt{s_{NN}} = 7$ TeV (upper left), p + Ar at $\sqrt{s_{NN}} = 6.3$ TeV (upper right), p + Sn at $\sqrt{s_{NN}} = 6.14$ TeV (lower left) and p + Pb at $\sqrt{s_{NN}} = 5.5$ TeV (lower right). The calculations are with CTEQ6 and employ the EKS98 (solid), nDSg (dashed), HKN (dot-dashed) and EPS09 (solid curves with symbols) shadowing parameterizations.
The $J/\psi$ ratios are shown in the upper half of the figures while the Υ results are on the lower half. Since shadowing is an initial-state effect, the same ratios would also be expected for the $\chi_c$ and $\psi'$ on the left and the higher Υ states ($\Upsilon'$, $\Upsilon''$, $\chi_b(1P)$ and $\chi_b(2P)$) on the right. The ratios in Figs. 4 and 5 are stretched mirror images of the gluon shadowing ratios in Fig. 3. The lowest $x$ values are probed by the lightest nuclei since the center of mass energy is higher for nuclei with $Z/A \sim 0.5$ than heavier, neutron-rich nuclei with lower $Z/A$. The differences in the shadowing ratios for a given parameterization are greatest at large negative $y$ where $x$ is largest. As $A$ increases and $\sqrt{s_{NN}}$ decreases, the antishadowing peak moves closer to midrapidity (less negative $y$). Increasing the scale from that appropriate for the $J/\psi$ to that for the Υ also moves the antishadowing peak closer to $y = 0$. For example, the EKS98 antishadowing peak is fully visible for Υ production, occurring at $y \sim -3.5$, while they only appear at $y \leq -5$ for the $J/\psi$. As $\mu^2$ increases, the differences in the EPS09 sets becomes more pronounced at large $x$, leading to the more irregular shapes of the upper and lower limits of the EPS09 uncertainty range at negative rapidity. Note that the central ratio is smooth. Thus, the results in Figs. 4 and 5 suggest that by running the LHC in both ALICE muon coverage.

Finally, we note that at $y = 6$, corresponding to $x < 10^{-6}$, the EKS98 and nDSg shadowing ratios are outside their range of validity. This is also near the region where DGLAP evolution of the parton densities is likely to break down. Nonlinear evolution of the proton parton densities is expected at sufficiently small $x$. The onset of these nonlinearities is predicted to be at larger $x$ for nuclei. However, it is not obvious that nonlinear parton evolution automatically leads to a reduction of the small $x$ gluon density even though the nonlinear term in the gluon evolution has a negative sign [52]. The behavior of the low $x$ gluon density cannot be determined without a complete re-evaluation of all the parton densities since the sea quark evolution is coupled to that of the Υ and overall momentum conservation must be maintained along with the integrity of the global fit. See Ref. [53] for details of modified parton densities based on nonlinear DGLAP evolution and Refs. [54, 55] for a discussion of the possible effect on charm production at the LHC.

Since it is more likely that the best $p + p$ reference data will be at $\sqrt{s} = 14$ TeV or 10 TeV for the initial LHC run, Figs. 6 and 7 show the $(p + A)/(p + p)$ ratios with the $p + p$ reference at 14 TeV. The magnitude of the two ratios (for $p + A$ collisions in the equal-speed frame, Fig. 6) and at the same energy as the corresponding $A + A$ collisions, Fig. 7, is due to the difference in $\sqrt{s_{NN}}$ relative to 14 TeV. The $p + A$ ratios with the $p + A$ center-of-mass energies equal to those of $A + A$ collisions are lower. The rapidity distributions narrow when their magnitudes are reduced with decreasing $\sqrt{s_{NN}}$. Thus the $(p + A)/(p + p)$ ratios without shadowing decrease steadily from 9.9 to 5.5 TeV while the narrowing of the ratios becomes more pronounced.

The symmetric solid curves in Figs. 6 and 7 are the $(p + A)/(p + p)$ ratios without shadowing. Shadowing results in asymmetric ratios but since the $p + A$ phase space is narrower than that of 14 TeV $p + p$ collisions, the ratios in these figures turn over and drop to zero at large $|y|$. The narrower phase space has a bigger effect on the Υ production ratios since the full Υ rapidity range is within $|y| < 6$ while the $J/\psi$ $y$ distribution is broader. The antishadowing peak is lowered and broadened when dividing by the 14 TeV $p + p$ rapidity distribution and is only really apparent for the EKS98 and EPS09 parameterizations. The maximum shadowing allowed by EPS09 shows the most asymmetric curvature, especially for $J/\psi$. The EPS09 ratios suggest that the effect could either be large, as suggested by the EPS08 analysis, or small enough to be effectively indistinguishable from no shadowing. It will thus be harder to differentiate between shadowing parameterizations when employing the higher energy $p + p$ reference.

As discussed previously, there is an additional complication due to the rapidity shift of the $p + A$ center of rapidity in the equal-speed frame. The shift increases with $A$ as $Z/A$ decreases, reducing the energy of the ion beam relative to the proton beam. This results in nearly half a unit rapidity shift in $p+$Pb collisions, as shown in the center part of Table 1 labeled $p + A$. The $(p + A)/(p + p)$ ratios including the rapidity shift and the maximum energy $p + p$ reference are shown in Fig. 8. Note that only the $p + p$ results in the equal-speed frame are shown. Since the proton beam momentum in $p + A$ collisions at the $A + A$ center-of-mass energy must be the same as that of the ion beam, $\Delta \eta_{cm} = 0$. The $p + A$ rapidity distribution is given a positive shift, to the right, since the proton beam, at higher $y$, is assumed to come from the left and move to the right. Thus, at large negative $y$, the ratios are lower than in Figs. 6 and 7 and are flatter as a function of rapidity. While the nuclear effects on the parton densities are most difficult to disentangle here, this scenario is the most realistic. (It may be possible to eliminate or reduce the effect of the rapidity shift by employing different rapidity cuts to compare distributions.) If the LHC is run with the proton and ion beam directions reversed, the antishadowing peak may be enhanced and the large positive rapidity ratios decreased.

The effect of the rapidity shift is reduced if d+$A$ collisions are run instead of $p + A$ collisions. The d+$A$ center-of-mass energy is closer to that of $A + A$ collisions since $Z/A < 1$ for the deuteron rather than equal to 1 as for protons. The ratios with the d+$A$ and $p + p$ collisions at the same center-of-mass energy per nucleon are shown in Fig. 9 (similar to Fig. 5 for $p + A$). They are like those in Fig. 5 with equal $p + A$ and $A + A$ center-of-mass energies since $\sqrt{s_{NN}}$ is similar for d+$A$ and $A + A$ collisions. Shadowing effects on the deuteron are assumed to be negligible.

The results with a 14 TeV $p + p$ reference and the small rapidity shift taken in account, see Table 1 for $\Delta \eta_{cm}$,
FIG. 6: The \((p + A)/(p + p)\) ratios with the \(p + p\) rapidity distributions calculated at \(\sqrt{s} = 14\) TeV. While the \(p + A\) distributions are calculated in the equal-speed frame, no rapidity shift has been taken into account. The effect of shadowing on \(J/\psi\) (a, upper 4 panels) and \(\Upsilon\) (b, lower 4 panels) production is shown. Each set of panels displays the production ratios for \(p + O\) at \(\sqrt{s_{NN}} = 9.9\) TeV (upper left), \(p + Ar\) at \(\sqrt{s_{NN}} = 9.39\) TeV (upper right), \(p + Sn\) at \(\sqrt{s_{NN}} = 9\) TeV (lower left) and \(p + Pb\) at \(\sqrt{s_{NN}} = 8.8\) TeV (lower right), all calculated in the equal-speed frame. The calculations are with CTEQ6 and employ the EKS98 (solid), nDSg (dashed), HKN (dot-dashed) and EPS09 (solid curves with symbols) shadowing parameterizations. The solid curve symmetric around \(y = 0\) is the \((p + A)/(p + p)\) ratio without shadowing.

are shown in Fig. 10. Recall that there is no rapidity shift for \(d + O\) collisions since \(Z_d/A_d = Z_O/A_O = 0.5\). Thus the equal-speed and center-of-rapidity frames coincide. In \(d + Pb\) collisions, since \(\Delta y_{cm} < 0.06\), the shift is negligible. Thus the \(d + A\) rapidity distributions relative to the 14 TeV \(p + p\) reference with the rapidity shift, shown in Fig. 10, are similar to those in Fig. 6 with \(\Delta y_{cm} = 0\) and the same \(\sqrt{s_{NN}}\) in \(p + A\) and \(A + A\) collisions. Note, however, that the ratios in Fig. 10 are somewhat closer to unity since the \(d + A\) center-of-mass energy is larger. Thus the more realistic \(d + A\) scenario shown in Fig. 10 would be preferable for determining nuclear effects on the parton densities both because of the relatively similar center-of-mass energies and the smaller rapidity shift.

We now extrapolate to \(A + A\) interactions to show the projected CNM effects from shadowing alone. The results for \(A + A\) collisions are shown in Figs. 11 and 12. The \((A + A)/(p + p)\) ratio with both systems calculated at the \(A + A\) center-of-mass energy are shown in Fig. 11 while the 14 TeV \(p + p\) reference is employed to obtain the ratios in Fig. 12. The results in Fig. 11 are essentially the convolutions of the \((p + A)/(p + p)\) ratios (with the same \(\sqrt{s_{NN}}\) for both systems and no rapidity shift) shown in Fig. 5 with their mirror image \(Ap/pp\) ratios. While the \(A + A\) ratios exhibit antishadowing peaks at \(y \approx \pm (4 - 5)\), the \((A + A)/(p + p)\) ratios are less than unity everywhere because the product of the \((p + A)/(p + p)\) ratios at positive \(y\) and the \((A + p)/(p + p)\) ratios at negative \(y\) is always smaller than one, e.g. \([(p + A)/(p + p)]_{y \sim 5} \sim 0.6 - 0.75\) while \([(A + p)/(p + p)]_{y \sim -5} \sim 1.2\). Thus, when all ratios are calculated
FIG. 7: The \((p + A)/(p + p)\) ratios with the \(p + p\) rapidity distributions calculated at \(\sqrt{s} = 14\) TeV. The \(p + p\) distributions are calculated at the \(A + A\) center-of-mass energy. The effect of shadowing on \(J/\psi\) (a, upper 4 panels) and \(\Upsilon\) (b, lower 4 panels) production is shown. Each set of panels displays the production ratios for \(p + O\) at \(\sqrt{s_{NN}} = 7\) TeV (upper left), \(p + Ar\) at \(\sqrt{s_{NN}} = 6.3\) TeV (upper right), \(p + Sn\) at \(\sqrt{s_{NN}} = 6.14\) TeV (lower left) and \(p + Pb\) at \(\sqrt{s_{NN}} = 5.5\) TeV (lower right). The ratios are calculated at the \(A + A\) energy. The calculations are with CTEQ6 and employ the EKS98 (solid), nDSg (dashed), HKN (dot-dashed) and EPS09 (solid curves with symbols) shadowing parameterizations. The solid curve symmetric around \(y = 0\) is the \((p + A)/(p + p)\) ratio without shadowing.

Calculations of color singlet \(J/\psi\) interactions in matter using the dipole approximation of the \(J/\psi\)-hadron cross section suggest that factorization is inapplicable due to the coherence of the interaction [57]. These gluon saturation models assume the dominance of higher-twist effects enhanced by powers of \(A^{1/3}\) in \(p + A\) interactions. If these models were valid, enhanced suppression of the \(J/\psi\) should set in at large rapidity. This does indeed seem to be the case at RHIC where \(1.2 < y < 2.2\) corresponds to \(0.0045 > x_F > 0.0017\) [13]. However, the forward \(x_F\) data at \(\sqrt{s} = 38\) GeV [0.2 < \(x_F\) 0.8 and 0.027 > \(x_F\) > 0.008] and 17 GeV [0.1 < \(x_F\) < 0.35 and 0.13 > \(x_F\) > 0.07], in particular, exhibit the same trends as at RHIC [58]. These fixed-target \(x_F\) ranges lie in the transition region from antishadowing to shadowing (38 GeV) and in the antishadowing region (17 GeV), see Fig. 4, seemingly too large to be in the saturation region, especially at \(\sqrt{s} = 17\) GeV.

As is the case for the RHIC \(A + A\) calculations at \(\sqrt{s_{NN}} = 200\) GeV [59], there is typically more suppression at the \(A + A\) center-of-mass energy, assuming factorization of \(A + A\) collisions into a convolution of \(p + A\) and \(A + p\) collisions,

\[
\frac{A + A}{p + p}\big|_{y \sim \pm 5} = \frac{p + A}{p + p}\big|_{y \sim -5} \times \frac{A + p}{p + p}\big|_{y \sim -5} < 1 .
\]

(13)
FIG. 8: (Color online) The \( (p + A)/(p + p) \) ratios with the \( p + p \) rapidity distributions calculated at \( \sqrt{s} = 14 \) TeV. The \( p + A \) rapidity distributions are calculated in the equal-speed frame with the rapidity shift taken into account. The effect of shadowing on \( J/\psi \) (a, upper 4 panels) and \( \Upsilon \) (b, lower 4 panels) production is shown. Each set of panels displays the production ratios for \( p + O \) at \( \sqrt{s_{NN}} = 9.9 \) TeV (upper left), \( p + Ar \) at \( \sqrt{s_{NN}} = 9.39 \) TeV (upper right), \( p + Sn \) at \( \sqrt{s_{NN}} = 9 \) TeV (lower left) and \( p + Pb \) at \( \sqrt{s_{NN}} = 8.8 \) TeV (lower right). The calculations are with CTEQ6 and employ the EKS98 (solid), nDSg (dashed), HKN (dot-dashed) and EPS09 (solid curves with symbols) shadowing parameterizations. The upper solid curve at \( y > 0 \) is the shifted \( (p + A)/(p + p) \) ratio without shadowing.

predicted at \( y = 0 \) than at more forward and backward rapidities for all the shadowing parameterizations as well as for both \( J/\psi \) and \( \Upsilon \) production. At RHIC, the \( A + A \) data are more suppressed at forward rapidity than at central rapidity, both in the minimum bias data as a function of rapidity and as a function of collision centrality, as quantified by the number of participant nucleons. Standard models of shadowing alone or shadowing with absorption by nucleons in cold nuclear matter or shadowing combined with dissociation in a quark-gluon plasma leads to strong suppression at central rapidities. However, \( J/\psi \) regeneration by coalescence of \( c \) and \( \bar{c} \) quarks in the medium [12, 60] is biased toward central rapidities and could lead to more suppression at forward rapidity relative to central rapidity since the rapidity distribution of \( J/\psi \) production by coalescence is expected to be narrower than the initial \( J/\psi \) rapidity distribution [60]. Thus, with coalescence, there should be more suppression at forward \( y \) than at midrapidity. The same trend should hold at the LHC. Coalescence production of the \( J/\psi \) should be even more important than at RHIC since more \( c\bar{c} \) pairs are created in a central Pb+Pb collision at \( \sqrt{s_{NN}} = 5.5 \) TeV. We can also expect that \( \Upsilon \) production by coalescence may be similar to that expected for the \( J/\psi \) at RHIC since the \( b\bar{b} \) production cross section at the LHC will be similar to the \( c\bar{c} \) production cross section at RHIC [13].

The \( (A + A)/(p + p) \) ratios with the 14 TeV \( p + p \) reference, shown in Fig. 12 are relatively flat. The dip around midrapidity has been washed out, except for the \( J/\psi \) ratios calculated with the EKS98 and EPS09 (central and maximum shadowing) parameterizations where some indication remains. For comparison, the \( (A + A)/(p + p) \) ratios
FIG. 9: (Color online) The (d+A)/(p+p) ratios with both d+A and p+p collisions at the d+A energy in the equal-speed frame. No rapidity shift has been taken into account. The effect of shadowing on \( J/\psi \) (a, upper 4 panels) and \( \Upsilon \) (b, lower 4 panels) production is shown. Each set of panels displays the production ratios for d+O at \( \sqrt{s_{NN}} = 7 \) TeV (upper left), d+Ar at \( \sqrt{s_{NN}} = 6.64 \) TeV (upper right), d+Sn at \( \sqrt{s_{NN}} = 6.41 \) TeV (lower left) and d+Pb at \( \sqrt{s_{NN}} = 6.2 \) TeV (lower right). The calculations are with CTEQ6 and employ the EKS98 (solid), nDSg (dashed), HKN (dot-dashed) and EPS09 (solid curves with symbols) shadowing parameterizations.

without shadowing are shown in the upper solid curves.

C. Impact parameter dependence

We now discuss the impact parameter dependence of quarkonium production at the LHC. Unfortunately, there is little relevant data on the spatial dependence of shadowing. Fermilab experiment E745 studied the spatial distribution of nuclear structure functions with \( \nu N \) interactions in emulsion. The presence of one or more dark tracks from slow protons is used to infer a more central interaction \([61]\). For events with no dark tracks, no shadowing is observed while, for events with dark tracks, shadowing is enhanced over spatially-independent measurements from other experiments. Unfortunately, this data is too limited to be used in a fit of the spatial dependence.

The minimum bias shadowing we have discussed up to now is homogeneous, impact parameter-integrated shadowing. The impact parameter-dependent results shown in this section portray inhomogeneous shadowing. In central collisions, with small impact parameter, \( b \), we can expect inhomogeneous shadowing to be stronger than the homogeneous result. In peripheral (large impact parameter) collisions, inhomogeneous effects are weaker than the homogeneous results but some shadowing is still present due to the overlapping tails of the density distributions. The stronger the homogeneous
FIG. 10: (Color online) The $(d+A)/(p+p)$ ratios with the $p+p$ distributions calculated at $\sqrt{s} = 14$ TeV and the $d+A$ rapidity shift taken into account. The effect of shadowing on $J/\psi$ (a, upper 4 panels) and $\Upsilon$ (b, lower 4 panels) production is shown. Each set of panels displays the production ratios for $d+O$ at $\sqrt{s_{NN}} = 7$ TeV (upper left), $d+Ar$ at $\sqrt{s_{NN}} = 6.64$ TeV (upper right), $d+Sn$ at $\sqrt{s_{NN}} = 6.41$ TeV (lower left) and $d+Pb$ at $\sqrt{s_{NN}} = 6.2$ TeV (lower right). The calculations are with CTEQ6 and employ the EKS98 (solid), nDSg (dashed), HKN (dot-dashed) and EPS09 (solid curves with symbols) shadowing parameterizations. The symmetric solid curve is the result without shadowing.

shadowing, the larger the difference between the central and peripheral results.

We assume that the shadowing is proportional to the parton path through the nucleus [62],

$$S_{t,\rho}^i(A, x, Q^2, \vec{r}, z) = 1 + N_{\rho}(S_{t,\rho}^i(A, x, Q^2) - 1) \int dz \rho_A(\vec{r}, z) \int dz \rho_A(0, z),$$

(14)

where $N_{\rho}$ is chosen to satisfy the normalization condition in Eq. [5]. The integral over $z$ in Eq. [14] includes the material traversed by the incident nucleon. At large distances, $s \gg R_A$, the nucleons behave as free particles, while in the center of the nucleus, the modifications are larger than the average value $S_{t,\rho}^i$.

We calculate the nuclear suppression factor, $R_{AB}$, for $p + A$, $d + A$ and $A + A$ collisions. The suppression factor is defined as the ratio [63]

$$R_{AB}(N_{\text{part}}; b) = \frac{d\sigma_{AB}/dy}{T_{AB}(b)d\sigma_{pp}/dy}$$

(15)

where $d\sigma_{AB}/dy$ and $d\sigma_{pp}/dy$ are the quarkonium rapidity distributions in $A + B$ and $p + p$ collisions and $T_{AB}$ is the
FIG. 11: (Color online) The \((A + A)/(p + p)\) ratios with both \(A + A\) and \(p + p\) collisions calculated at the \(A + A\) center-of-mass energy. The effect of shadowing on \(J/\psi\) (a, upper 4 panels) and \(\Upsilon\) (b, lower 4 panels) production is shown. Each set of panels displays the production ratios for \(O + O\) at \(\sqrt{s_{NN}} = 7\) TeV (upper left), \(Ar + Ar\) at \(\sqrt{s_{NN}} = 6.3\) TeV (upper right), \(Sn + Sn\) at \(\sqrt{s_{NN}} = 6.14\) TeV (lower left) and \(Pb + Pb\) at \(\sqrt{s_{NN}} = 5.5\) TeV (lower right). The calculations are with CTEQ6 and employ the EKS98 (solid), nDSg (dashed), HKN (dot-dashed) and EPS09 (solid curves with symbols) shadowing parameterizations.

nuclear overlap function,

\[ T_{AB}(b) = \int d^2s dz' \rho_A(s, z) \rho_B(|\vec{b} - \vec{s}|, z') . \]  

(16)

In \(p + A\) collisions, we assume that the proton has a negligible size, \(\rho_A(s, z) = \delta(s)\delta(z)\) so that \(T_{AB}(b)\) collapses to the nuclear profile function \(T_B(b) = \int dz' \rho_B(b, z')\). The deuteron cannot be treated like a point particle since it is large and diffuse. We use the Hülthen wave function \[64\] to calculate the deuteron density distribution. However, we do not include shadowing effects on the deuteron.

We show the \(p + Pb\) and \(d + Pb\) suppression factors as a function of impact parameter with \(\sqrt{s_{NN}} = 8.8\) TeV and 6.2 TeV in the numerator and denominator in Figs. 13 and 14 respectively. We concentrate on the largest \(A\) ion, \(Pb\), to maximize the relevant impact parameter range. The results in Fig. 13 are given for three values of rapidity: \(y = -4\) (backward rapidity, in the antishadowing region for \(\Upsilon\)), \(y = 0\) (midrapidity) and \(y = 4\) (forward rapidity, where fairly strong shadowing is expected). We present \(J/\psi\) ratios on top and \(\Upsilon\) ratios on the bottom. For comparison, the horizontal lines, centered around the average path length through the lead nucleus, \(b \sim (3/4)R_{Pb}\), show the impact parameter-integrated ratios in Fig. 4. The \(b\) dependence is strong, resulting in \(R_{pPb} \sim 1\) for \(b > R_{Pb}\). Shadowing is stronger in central collisions than the average integrated value, as expected. Because the average decreases at forward rapidities while the spatial dependence is relatively unchanged, the strongest \(b\) dependence is seen for the most forward
FIG. 12: (Color online) The \((A + A)/(p + p)\) ratios with the \(p + p\) rapidity distributions calculated at \(\sqrt{s} = 14\) TeV. The effect of shadowing on \(J/\psi\) (a, upper 4 panels) and \(\Upsilon\) (b, lower 4 panels) production is shown. Each set of panels displays the production ratios for O+O at \(\sqrt{s_{NN}} = 7\) TeV (upper left), Ar+Ar at \(\sqrt{s_{NN}} = 6.3\) TeV (upper right), Sn+Sn at \(\sqrt{s_{NN}} = 6.14\) TeV (lower left) and Pb+Pb at \(\sqrt{s_{NN}} = 5.5\) TeV (lower right). The calculations are with CTEQ6 and employ the EKS98 (solid), nDSg (dashed), HKN (dot-dashed) and EPS09 (solid curves with symbols) shadowing parameterizations. The upper solid curve is the \((A + A)/(p + p)\) ratio without shadowing.

The results for nucleus-nucleus collisions are presented as a function of the number of participant nucleons, \(N_{\text{part}}\),
FIG. 13: (Color online) The suppression factor $R_{pPb}$ at $y = -4$ (left), 0 (center) and 4 (right) as a function of $b$. The result is shown for $J/\psi$ (top) and $\Upsilon$ (bottom) in $p+Pb$ relative to $p+p$ collisions at the same energy, $\sqrt{s_{NN}} = 8.8$ TeV, and employ the EKS98 (solid), nDSg (dashed) and EPS09 (solid curves with symbols) shadowing parameterizations. The horizontal lines show the impact-parameter integrated results.

which depends on $b$ as

$$N_{\text{part}}(b) = \int d^2s \left[ T_A(s)(1 - \exp[-\sigma_{\text{inel}}(s_{NN})T_B(|\vec{b} - \vec{s}|)]) 
+ T_B(|\vec{b} - \vec{s}|)(1 - \exp[-\sigma_{\text{inel}}(s_{NN})T_A(s)]) \right].$$

(17)

Large values of $N_{\text{part}}$ are obtained for small impact parameters with $N_{\text{part}}(b = 0) = 2A$ for spherical nuclei. Small values of $N_{\text{part}}$ occur in very peripheral collisions. Figure 15 shows $R_{AA}(N_{\text{part}})$, at $y = 0$ for the four $A + A$ systems where the $p + p$ and $A + A$ rapidity distributions are calculated at the same center-of-mass energy. A similar pattern is observed for other values of $y$ since the $(A + A)/(p + p)$ ratios are approximately independent of rapidity over a rather broad range. The $(A + A)/(p + p)$ ratio at $y = 0$ from Fig. 11 is indicated by a horizontal line. Note that $R_{AA}(N_{\text{part}})$ in Fig. 15 is equal to $(A + A)/(p + p)$ in Fig. 11 for $N_{\text{part}}(b \approx R_A)$. In small systems, $R_{AA}(N_{\text{part}})$ is almost linear with more curvature appearing for larger collision systems.

Since the $p + p$ reference is not likely to be immediately available at the $A + A$ center-of-mass energy for $R_{AA}$ studies, Eq. (15), it may be preferable to study ratios of two quantities measured at the same energy in $A + B$ collisions where $B = p, d,$ or $A$. In this case, we utilize $R_{CP}$, the ratio of $A + B$ cross sections in central relative to peripheral collisions,

$$R_{CP}(y) = \frac{T_{AB}(b_P) d\sigma_{AB}(b_C)/dy}{T_{AB}(b_C) d\sigma_{AB}(b_P)/dy},$$

(18)

where $b_C$ and $b_P$ correspond to the central and peripheral values of the impact parameter. Indeed, shadowing may best be probed by $R_{CP}$ measurements in asymmetric systems since the most peripheral collisions are a good approximation to nucleon-nucleon collisions. The same rapidity shift is common to both central and peripheral collisions. We note, however, that an experimental measurement will not be able to define a precise impact parameter but will instead define impact parameter bins of finite width. Thus any comparison of calculations to data must be integrated over
the width of the impact parameter bin which will average the impact parameter dependence of the shadowing over the bin width. Our calculations include a width of 0.2R_A for the impact parameter bins.

In fact, studying R_{CP} in p + A and d + A collisions could provide a direct measure of shadowing if absorption is negligible since higher-order corrections unrelated to shadowing cancel in the ratio [62]. As an example of an asymmetric system, Fig. 16 presents R_{CP}(y) for d + Pb collisions with b_C = 0 and b_P \approx R_A^5. As expected, the resulting R_{CP}(y) are very similar to the impact-parameter averaged (d + A)/(p + p) ratios shown in Fig. 9. Since R_{CP}(y) with b_P \approx 2R_A are similar to those in Fig. 16, they are not shown.

Figures 17 and 18 show the values of R_{CP} for b_P \approx R_A and 2R_A relative to b_C = 0 in the four A + A systems studied for the J/ψ (Fig. 17) and the Υ (Fig. 18). Since the change in R_{AA}(N_{part}) between b_C = 0 and b_P \approx R_A is small (see Fig. 15), these ratios are almost independent of rapidity and give R_{CP} close to unity. On the other hand, the weaker shadowing effect at b_P \approx 2R_A produces a stronger rapidity dependence and a lower R_{CP}. Note that, as in Fig. 16, R_{CP}(y) for b_P \approx 2R_A is similar to [(A + A)/(p + p)]_{y>0} in Fig. 11. Thus, if no other medium effects are present, it is possible to trace the shadowing effect rather accurately by determining R_{CP} for sufficiently narrow centrality bins.

III. SUMMARY

We have provided a survey of the quarkonium total cross sections to next-to-leading order in the color evaporation model for all A + B combinations and energies at the LHC. We have included initial-state shadowing, employing several parameterizations of the nuclear modifications of the parton densities, but assumed final-state absorption is

---

5 We do not show R_{CP} for p + A collisions.
FIG. 15: (Color online) The suppression factor $R_{AA}$ at $y = 0$ as a function of $N_{\text{part}}$. The effect of shadowing on $J/\psi$ (a, upper 4 panels) and $\Upsilon$ (b, lower 4 panels) production is shown. Each set of panels displays the suppression factor for O+O at $\sqrt{s_{NN}} = 7$ TeV (upper left), Ar+Ar at $\sqrt{s_{NN}} = 6.3$ TeV (upper right), Sn+Sn at $\sqrt{s_{NN}} = 6.14$ TeV (lower left) and Pb+Pb at $\sqrt{s_{NN}} = 5.5$ TeV (lower right). The calculations are with CTEQ6 and employ the EKS98 (solid), nDSg (dashed) and EPS09 (solid curves with symbols) shadowing parameterizations.

negligible. If the nuclear absorption of quarkonium production can indeed be ignored at LHC energies, it may be possible to use the different mass scales for $J/\psi$ and $\Upsilon$ production to study the scale dependence of the gluon density in the nucleus as well as in the proton. There are considerable uncertainties in the predictions due to the incomplete knowledge of the nuclear gluon distribution. Indeed, at midrapidity, the range of the EPS09 ($p + A$)/($p + p$) ratios differs by a factor of two.

To illustrate the range of predictions for the different systems, we have calculated ($p + A$)/($p + p$) and (d+$A$)/($p + p$) ratios from the most naive (both systems at the same energy) to the most realistic (the $p + p$ reference at 14 TeV and the rapidity shift of $p + A$ interactions in the equal-speed frame). The most naive ratios are most straightforward for extracting the nuclear gluon distributions. It is still possible to use the most realistic ratios using a combination of experimental cuts on the rapidity distributions when $\Delta y_{\text{cm}} \neq 0$ and modeling the appropriate $x$ values for comparing
FIG. 16: (Color online) The central-to-peripheral ratios, \( R_{CP} \), as a function of rapidity for \( b_P \approx R_A \) relative to \( b = 0 \) for \( d+Pb \) collisions at \( \sqrt{s_{NN}} = 6.2 \) TeV. The calculations are with CTEQ6 and employ the EKS98 (solid), nDSg (dashed) and EPS09 (solid curves with symbols) shadowing parameterizations.

FIG. 17: (Color online) The central-to-peripheral ratios, \( R_{CP}(y) \), for \( b = R_A \) (a) and \( b = 2R_A \) (b) relative to \( b = 0 \). The effect of shadowing on \( J/\psi \) production is shown for O+O at \( \sqrt{s_{NN}} = 7 \) TeV (upper left), Ar+Ar at \( \sqrt{s_{NN}} = 6.3 \) TeV (upper right), Sn+Sn at \( \sqrt{s_{NN}} = 6.14 \) TeV (lower left) and Pb+Pb at \( \sqrt{s_{NN}} = 5.5 \) TeV (lower right). The calculations are with CTEQ6 and employ the EKS98 (solid), nDSg (dashed) and EPS09 (solid curves with symbols) shadowing parameterizations.
FIG. 18: (Color online) The central-to-peripheral ratios, $R_{CP}(y)$, for $b = R_A$ (a) and $b = 2R_A$ (b) relative to $b = 0$. The effect of shadowing on $\Upsilon$ production is shown for $O+O$ at $\sqrt{s_{NN}} = 7$ TeV (upper left), $Ar+Ar$ at $\sqrt{s_{NN}} = 6.3$ TeV (upper right), $Sn+Sn$ at $\sqrt{s_{NN}} = 6.14$ TeV (lower left) and $Pb+Pb$ at $\sqrt{s_{NN}} = 5.5$ TeV (lower right). The calculations are with CTEQ6 and employ the EKS98 (solid), nDSg (dashed) and EPS09 (solid curves with symbols) shadowing parameterizations.

$p+p$ collisions at $\sqrt{s} = 14$ TeV with lower energy $p+A$ collisions. As is clear from the RHIC analyses [13], the $A+A$ studies require a good understanding of the nuclear gluon distribution to extract hot and dense matter effects. Thus the more realistic $d+A$ scenario shown in Fig. 10 would be preferable for determining nuclear effects on the parton densities both because of the relatively similar $d+A$ and $A+A$ center-of-mass energies as well as the smaller rapidity shift relative to $p+A$ collisions in the equal-speed frame.

To more cleanly extract the parton densities at LHC energies, it would be preferable to have $e+p$ and $e+A$ data at the appropriate $x$ and $\mu^2$ range of the LHC data. (The HERA $x$ range reaches to approximately the value appropriate for $J/\psi$ production in 5.5 TeV/nucleon collisions at midrapidity. Unfortunately, the $\mu^2$ probed at these $x$ values is smaller than the $J/\psi$ mass scale.) So far, the nDIS data is not available at small enough $x$ values and, simultaneously, large enough $\mu^2$ to be relevant for quarkonium production at high energies. While electron-proton collisions, as studied at HERA, would be useful for obtaining the baseline in $p+p$, they are not sufficient for defining the modification of the nuclear gluon distributions for $p+A$ collisions, $e+A$ studies are needed.

The shadowing parameterizations used in our study exhibit a wide range of behavior for the nuclear gluon density at low $x$, outside the current range of the fits from fixed-target nDIS data at higher $x$ and low $\mu^2$. If nuclear data were available from high energy $e+A$ collisions, the nuclear gluon densities could be more precisely pinned down by global analyses of the scale dependence of the nuclear structure functions. In hadroproduction, direct photon or open charm production, dominated by gluon-induced processes but without the additional complexities of nuclear absorption,
could be utilized to study the nuclear gluon density. Any new $e + A$ data before an electron ring is available at the LHC will be at lower energies than previously available at HERA, reducing the potential overlap of the low $x$ range between an electron-ion collider and the LHC.

We note that, since we have assumed absorption is negligible at the LHC and include no other cold nuclear matter effect, the uncertainties on the ratios can be obtained from the EPS09 bands shown in the figures. However, if other effects are incorporated, a more extensive error analysis, including the uncertainties on these other effects, is necessary.

Finally, we note that the central-to-peripheral ratio, $R_{CP}$, may be useful for extracting the shadowing effect at a given collision energy if the experimental resolution of the impact parameter bins is narrow enough. Indeed, $R_{CP}$ measurements may be a superior method of studying asymmetric systems since very peripheral collisions are a good approximation to nucleon-nucleon collisions. This ratio is advantageous because it can be made at the same collision energy with a common rapidity shift.

Acknowledgements

The numerical values of the ratios shown in this paper are available from the author.

We thank K. J. Eskola, H. Paukkunen and C. Salgado for providing the EPS09 files and for discussions. This work was performed under the auspices of the U.S. Department of Energy by Lawrence Livermore National Laboratory under Contract DE-AC52-07NA27344 and was also supported in part by the National Science Foundation Grant PHY-0555660.
[31] R. Vogt, Phys. Rept. 310, 197 (1999).
[32] C. Lourenço, R. Vogt and H. Wüthri, JHEP 0902, 014 (2009) [arXiv:0901.3054 [hep-ph]].
[33] D. M. Alde et al. [E772 Collaboration], Phys. Rev. Lett. 66, 2285 (1991).
[34] H. Liu [STAR Collaboration], Nucl. Phys. A 830, 235c (2009) [arXiv:0907.4538 [nucl-ex]].
[35] K. J. Eskola, V. J. Kolhinen and P. V. Ruuskanen, Nucl. Phys. B 535, 351 (1998) [arXiv:hep-ph/9802350].
[36] K. J. Eskola, V. J. Kolhinen and C. A. Salgado, Eur. Phys. J. C 9, 61 (1999) [arXiv:hep-ph/9807297].
[37] D. de Florian and R. Sassot, Phys. Rev. D 69, 074028 (2004) [arXiv:hep-ph/0311227].
[38] M. Hirai, S. Kumano and T. H. Nagai, Phys. Rev. C 70, 044905 (2004) [arXiv:hep-ph/0404093].
[39] K. J. Eskola, H. Paulukonen and C. A. Salgado, JHEP 0807, 102 (2008) [arXiv:0802.0139 [hep-ph]].
[40] K. J. Eskola, H. Paulukonen and C. A. Salgado, JHEP 0904, 065 (2009) [arXiv:0902.4154 [hep-ph]].
[41] P. Charpentier et al. [NA3 Collaboration], Z. Phys. C 20, 101 (1983).
[42] M. Hirai, S. Kumano and T. H. Nagai, Phys. Rev. C 70, 044905 (2004) [arXiv:hep-ph/0404093].
[43] S. Gavin and M. Gyulassy, Phys. Lett. B 214, 241 (1988).
[44] J. Hufner, Y. Kurihara, and H. J. Pirner, Phys. Lett. B 214, 241 (1988); J.-P. Blaizot and J.-Y. Ollitrault, Phys. Lett. B 217, 92 (1989).
[45] T. Ullrich, J. Phys. G 35, 140410 (2008).
[46] M. Gluck, E. Reya and A. Vogt, Z. Phys. C 53, 127 (1992).
[47] A. Adare et al. [PHENIX Collaboration], Phys. Rev. C 77, 024902 (2008) [arXiv:0711.3917 [nucl-ex]]; erratum ibid. C 79, 059901 (2009) [arXiv:0903.4845 [nucl-ex]].
[48] I. Arsene et al. [BRAHMS Collaboration], Phys. Rev. Lett. 93, 242303 (2004) [arXiv:nucl-ex/0403005].
[49] B.Z. Kopeliovich, E. Levin, I.K. Potashnikova and I. Schmidt, Phys. Rev. C 79, 064906 (2009) [arXiv:0811.2210 [hep-ph]].
[50] R. Vogt, J. Phys. G 31, S773 (2005) [arXiv:hep-ph/0412303].
[51] A. J. Baltz et al., Phys. Rev. 458, 1 (2008) [arXiv:0706.3356 [nucl-ex]].
[52] L. V. Gribov, E. M. Levin and M. G. Ryskin, Nucl. Phys. B 188, 555 (1981); Zh. Eksp. Teor. Fiz. 80, 2132 (1981); A. H. Mueller and J. W. Qiu, Nucl. Phys. B 258, 427 (1986).
[53] K. J. Eskola, H. Honkanen, V. J. Kolhinen, J. W. Qiu and C. A. Salgado, Nucl. Phys. B 660, 211 (2003) [arXiv:hep-ph/0211239].
[54] K.J. Eskola, V.J. Kolhinen and R. Vogt, Phys. Lett. B 582 (2004) 157 [arXiv:hep-ph/0310111].
[55] A. Dainese, R. Vogt, M. Bondila, K.J. Eskola and V.J. Kolhinen, J. Phys. G 30, 1787 (2004) [arXiv:hep-ph/0403009].
[56] J. Baines et al., Heavy Quarks Summary Report for the HERA-LHC Workshop Proceedings, arXiv:hep-ph/0601164.
[57] D. Kharzeev, E. Levin, M. Nardi and K. Tuchin, Nucl. Phys. A 826, 230 (2009) [arXiv:0809.2933 [hep-ph]].
[58] C. Lourenco, talk at ECT* workshop on Quarkonium Production in Heavy-Ion Collisions, Trento (Italy), May 25-29, 2009 and at Joint CATHIE-INT mini-program “Quarkonia in Hot QCD”, June 16-26, 2009 http://www.int.washington.edu/talks/Workshops/int09f2w.
[59] R. Vogt, Phys. Rev. C 71, 054902 (2005) [arXiv:hep-ph/0411378]; Heavy Ion Phys. 25 (2006), 97 [arXiv:nucl-th/0507027].
[60] R. L. Thews and M. L. Mangano, Phys. Rev. C 73, 044904 (2006) [arXiv:nucl-th/0505055].
[61] T. Kitigaki et al. [E745 Collaboration], Phys. Lett. B 214, 281 (1988).
[62] S. R. Klein and R. Vogt, Phys. Rev. Lett. 91, 142301 (2003) [arXiv:nucl-th/0305046].
[63] K. Adcox et al. [PHENIX Collaboration], Phys. Rev. Lett. 88, 122301 (2002) [arXiv:nucl-ex/0109003].
[64] D. Kharzeev, E. M. Levin and M. Nardi, arXiv:hep-ph/0212316; L. Hulthen and M. Sagawara, Handbuch der Physik, 39 (1957).