LAGRANGIAN RELAXATION SCHEMES FOR CALCULATING FORCE-FREE MAGNETIC FIELDS, AND THEIR LIMITATIONS

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Abstract

Force-free magnetic fields are important in many astrophysical settings. Determining the properties of such force-free fields—especially smoothness and stability properties—is crucial to understanding many key phenomena in astrophysical plasmas, for example, energy release processes that heat the plasma and lead to dynamic or explosive events. In the present work we discuss a serious limitation on the computation of force-free fields, within the context of a Lagrangian relaxation scheme that conserves magnetic flux and \( \nabla \cdot \mathbf{B} \) identically. This issue has the potential to invalidate the results produced by numerical force-free field solvers even for cases in which they appear to converge (at fixed grid resolution) to an equilibrium magnetic field. Error estimates are introduced to assess the quality of the calculated equilibrium. We go on to present an algorithm, based on rewriting the curl operation via Stokes’ theorem, for calculating the current which holds great promise for improving dramatically the accuracy of the Lagrangian relaxation procedure.

Key words: magnetic fields – methods: numerical – stars: coronae

1. INTRODUCTION

Force-free magnetic fields, \( \mathbf{B} \), satisfy \( \mathbf{J} \times \mathbf{B} = 0 \), or equivalently

\[
(\nabla \times \mathbf{B}) \times \mathbf{B} = 0. \tag{1}
\]

Calculation of such force-free fields is of importance in many astrophysical settings, for example, accretion disks around various objects (e.g., Frank et al. 2002; Uzdensky et al. 2002), neutron stars (McKinney 2006), pulsars (Mestel 1973), magnetic clouds (Burlaga 1988), and solar and stellar coronae (e.g., Anzer 1968).

A particular application in solar physics is the controversial “topological dissipation” model proposed by Parker (1972). The assertion of this model is that if an equilibrium magnetic field is perturbed by arbitrary motions at a line-tied boundary, then the subsequent field cannot relax to a smooth force-free equilibrium. Rather, the equilibrium must contain tangential discontinuities—corresponding to current sheets. Doubt has been cast upon the model however, as a number of authors have demonstrated the existence of smooth solutions in the scenario posed (van Ballegooijen 1985; Zweibel & Li 1987; Longcope & Strauss 1994; Craig & Sneyd 2005). The question as to whether current sheets form spontaneously in the coronal magnetic field is key to understanding the so-called coronal heating problem. This is just one example which demonstrates that determining both the structure and stability of force-free magnetic fields is of fundamental importance.

There are different approaches that one may take when searching for force-free magnetic fields. One method, often used when modeling the solar corona, is to solve a boundary value problem (Amari et al. 1997), and see Schrijver et al. (2006) for a comparison of numerical schemes). In such nonlinear force-free extrapolation methods, the force-free field is reconstructed from boundary data, provided, for example, by a vector magnetogram. An alternative approach is to impose an initial magnetic field over the entire volume that is not in force-free equilibrium, and then to perform a relaxation procedure (see below). This is the natural approach if one wants to investigate the properties of particular magnetic topologies. As long as the relaxation procedure can be guaranteed to be ideal, then the topology will be conserved during the relaxation.

One powerful computational approach for investigating the properties of force-free fields is to employ an ideal Lagrangian relaxation scheme. Such schemes exploit the property that under ideal magnetohydrodynamics (MHD) the vector \( \mathbf{B}/\rho \) evolves according to the equation

\[
\frac{D}{Dt} \frac{\mathbf{B}}{\rho} = \left( \frac{\mathbf{B}}{\rho} \cdot \nabla \right) \mathbf{v}, \tag{2}
\]

where \( D/Dt \) is the material derivative, \( \rho \) is the plasma density, and \( \mathbf{v} \) is the plasma velocity. This is of exactly the same form as the evolution equation of a line element \( \delta x \) in a flow (see, e.g., Moffatt 1978), and thus a Lagrangian description facilitates a relaxation that is, by construction, ideal. These schemes can be used to investigate the structure and (ideal MHD) stability of force-free fields. The latter is guaranteed by the iterative convergence of the scheme provided that the resolution is sufficient. The primary variables that the numerical scheme dynamically updates are the locations of the mesh points, with the quantities \( \mathbf{B} \) and \( \mathbf{J} \) being calculated via matrix products involving the initial magnetic field and derivatives of the mapping that describes the mesh deformation. Beginning with a nonequilibrium magnetic field, the computational mesh is evolved in such a way as to deform the associated magnetic field toward a force-free state. An artificial frictional term is included in the equation of motion (see also Chodura & Schlueter 1981) which guarantees a monotonic decrease of the energy. Two implementations of this method are described in Craig & Sneyd (1986) and Longbottom et al. (1998). The method has been used extensively to investigate the stability and equilibrium properties of various different magnetic configurations, such as the kink instability of magnetic flux tubes (Craig & Sneyd 1990), line-tied collapse of two-dimensional and three-dimensional magnetic null points (Craig & Litvinenko 2005; Pontin & Craig 2005) and the Parker problem (Longbottom et al. 1998; Craig & Sneyd 2005).
In the following section we describe a test problem that illustrates one major difficulty in the computation of force-free fields, in the context of the Lagrangian relaxation scheme outlined above. In Section 3, we present two possible extensions of the numerical scheme. In Section 4 we describe our results, and in Section 5 we present our conclusions.

2. THE PROBLEM

2.1. Outline of the Problem

In a numerical relaxation experiment using braided initial fields (Wilmot-Smith et al. 2009) we came across an inconsistency of the resulting numerical force-free state, which is best explained with the help of the following example. Consider a magnetic field obtained from the homogenous field by a simple twisting deformation as shown in Figure 1(a). Obviously, an ideal relaxation toward a force-free state must end again in a homogenous state (\(J = 0\)). During this process the Lagrangian relaxation leads to a deformation of the initial computational mesh which exactly cancels the initial deformation applied to the homogenous field. This is a well defined setup in which we know exactly the initial and final states. We now employ the implicit (alternating direction implicit—ADI) relaxation scheme in what follows, it is worth emphasizing here the distinction between iterative convergence (at fixed resolution \(N\)) and real convergence, i.e., convergence toward a “correct” solution as the resolution \(N\) becomes sufficiently large.

2.2. Analysis

In order to investigate the source of the inconsistency described in the previous section, we consider the test problem outlined there. Specifically, we begin with an initially uniform magnetic field (\(B = b_0 \hat{z}\)), and superimpose two regions of toroidal field, centered on the \(z\)-axis at \(\pm z_0\), with exactly the same functional form, but of opposite signs:

\[
B = b_0 \hat{z} + \sum_{i=1}^{2} \frac{2 b_0 \phi_i}{\pi a_r} \exp\left(-\frac{x^2 + y^2}{a_r^2} - \frac{(z - L_i)^2}{a_z^2}\right)(-y \hat{x} + x \hat{y}),
\]

with \(b_0 = 1\), \(a_r = \sqrt{2}\), \(a_z = 2\) and \(\phi_1 = \pi\), \(\phi_2 = -\pi\), \(L_1 = -4\), \(L_2 = 4\). We refer to this field in the following as \(T2\). The field \(T2\) (see Figure 1(a)) is constructed such that the two regions of twisted field, which are of opposite sign, should exactly cancel one another under an ideal relaxation, approaching the uniform field (with \(J = 0\)) as the equilibrium. Note that \(\phi\) is the maximum turning angle of field lines around the \(z\)-axis.

One of the great advantages of an ideal Lagrangian relaxation is that it is possible to extract the paths of the magnetic field lines of the final state if one knows them in the initial state, simply by interpolating over the mesh displacement. Calculating the field lines in this way, no error is accumulated as occurs when integrating numerically along \(B\). Given knowledge of the field line paths, one can test the quality of the force-free approximation by plotting \(\alpha\) along field lines. For a force-free field

\[
\nabla \times B = \alpha B, \tag{4}
\]

and \(\alpha\) should be constant along field lines since taking the divergence of the above yields

\[
B \cdot \nabla \alpha = 0. \tag{5}
\]

We begin by defining the variable \(\alpha^*\), motivated by Equation (4), as

\[
\alpha^* = \frac{J_1}{|B|}. \tag{6}
\]
We find that for the magnetic field calculated by the relaxation scheme, the value of $\alpha^*$ changes dramatically along field lines. Of course the relaxation gives a magnetic field for which $\nabla \cdot J = 0$, where $J$ is the exact value. The lower number is $100 \times |B_i - \hat{B}_i|_{\text{max}}/|\hat{B}_i|_{\text{max}}$, where $\hat{B}_i$ is the exact value. The upper number shows the maximum relative percentage error in the domain over all components of $B$. It is clear that there are large errors in the current calculated by the numerical scheme. While errors in $B$ and in each individual term in the mesh distortion Jacobian are much smaller, it turns out that the combination in which they are multiplied, summed, and divided to calculate $J$ incurs large errors. The calculation has been meticulously checked such that we are certain that the error appears not due to mathematical or coding error, but rather due to an accumulation of numerical truncation errors in the process of calculating $J$ (via Equation (2.10) in Craig & Sneyd 1986). Note that the errors increase as the mesh distortion ($\psi$) increases, and decrease with resolution ($N$).

### 3. MORE SOPHISTICATED NUMERICAL SCHEMES

#### 3.1. Higher-Order Derivatives

Clearly, the accuracy of the force-free approximation is impaired by the accuracy of the curl operation performed in the numerical scheme. One way to increase the accuracy of spatial derivatives could be to use higher-order finite-difference expressions. Existing versions of the scheme use conventional second-order centered differences involving two nearest-neighbor values. If we expect smooth solutions (without grid-scale features) then increasing to fourth-order finite-difference expressions (using four nearest neighbors) is expected to increase the accuracy. Recall that in the frictional Lagrangian method fluid displacements are determined from an equation of the form

$$\frac{\partial x_i}{\partial t} = A_{ij\beta\gamma} x_{\alpha,\beta\gamma} + C_i,$$

where $A$ and $C$ are prescribed tensor and vector functions and summation over repeated (Greek) indices is assumed. Note that the partial differentiation with respect to the background Cartesian coordinates $(X_1, X_2, X_3)$ is indicated using the comma notation (i.e., $\partial^2 f/\partial \beta \partial \gamma = f_{,\beta\gamma}$).

The important point for us is that the method involves two spatial derivatives of the Lagrangian variables $x_\alpha$. This reflects the fact that the Lorentz force is computed using first- and second-order derivatives of the Lagrangian mesh. Now “diagonal derivatives” such as $x_{i,ij}$ are relatively easy to compute using finite differences: they involve the point

$$\nabla \cdot (B / |B|) \equiv \sum_{i=1}^{3} \frac{\partial B_i}{\partial x_i} = \frac{\nabla \cdot \mathbf{B}}{|B|} = 0,$$

and it is here that these discretizationerrors arise. This is demonstrated below.

#### 2.3. Accuracy Test: Artificially Imposed Deformation

To ascertain the source of the errors, we take our initial state $T2$ and instead of performing the relaxation procedure, we artificially apply a deformation to the mesh which we can write down as a closed form expression, and moreover for which we can obtain the derivatives of the mesh displacement, and thus the resultant $B$ and $J$ fields, as closed form expressions. Motivated by the results of the relaxation, we impose a similar rotational distortion of the mesh which acts to “untwist” the field, via the transformation

$$(x, y, z) \rightarrow (x \cos \theta - y \sin \theta, y \cos \theta + x \sin \theta, z),$$

where

$$\theta = \psi \exp \left(-\frac{x^2 + y^2 - z^2}{4} - \frac{z^2}{16}\right).$$

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### Table 1

| $N$ | $\psi = \pi$ | $\psi = 3\pi/4$ | $\psi = \pi/2$ |
|-----|--------------|-----------------|----------------|
| 21  | 29.5         | 15.3            | 6.46           |
|     | 670          | 159             | 37.8           |
| 41  | 7.22         | 3.93            | 1.72           |
|     | 144          | 42.4            | 10.6           |
| 61  | 3.30         | 1.80            | 0.793          |
|     | 64.7         | 19.9            | 4.99           |
| 81  | 1.86         | 1.02            | 0.451          |
|     | 37.0         | 11.3            | 2.83           |

Notes. $N^3$ is the mesh resolution. In each case, the upper number shows the maximum relative percentage error in the domain over all components of $B$, i.e., $100 \times |B_i - \hat{B}_i|_{\text{max}}/|\hat{B}_i|_{\text{max}}$, where $\hat{B}_i$ is the exact value. The lower number is $100 \times |J_i - \hat{J}_i|_{\text{max}}/|\hat{J}_i|_{\text{max}}$. The important point for us is that the method involves two spatial derivatives of the Lagrangian variables $x_\alpha$. This reflects the fact that the Lorentz force is computed using first- and second-order derivatives of the Lagrangian mesh. Now “diagonal derivatives” such as $x_{i,ij}$ are relatively easy to compute using finite differences: they involve the point

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$$\nabla \cdot (B / |B|) \equiv \sum_{i=1}^{3} \frac{\partial B_i}{\partial x_i} = \frac{\nabla \cdot \mathbf{B}}{|B|} = 0,$$
itself and two/four nearest neighbors depending on whether the scheme is second or fourth order. It is these derivatives that are handled implicitly (via tridiagonal and pentadiagonal implementations of the ADI method) to provide, formally at least, the unconditional stability of the numerical scheme. Note that the computation of the mixed derivatives can be more complicated: terms such as \( x_{i,j,k} \) involve 16 terms in the fourth-order scheme, as opposed to just four when the method is second order. However, irrespective of the formal accuracy, perhaps the main drawback of the scheme is that, unlike second-order scheme, as opposed to just four when the method is second order.

Furthermore, the area of the enclosed quadrilateral is

\[
A = \frac{1}{2} \left| \mathbf{dx}_1 \times \mathbf{dx}_2 \right| + \frac{1}{2} \left| \mathbf{dx}_2 \times \mathbf{dx}_3 \right| + \frac{1}{2} \left| \mathbf{dx}_3 \times \mathbf{dx}_4 \right| + \frac{1}{2} \left| \mathbf{dx}_4 \times \mathbf{dx}_1 \right|. 
\]

(14)

We can now define the direction perpendicular to this mesh surface as

\[
\mathbf{n}^{(1)} = \frac{1}{4} \left( \frac{\mathbf{dx}_1 \times \mathbf{dx}_2}{|\mathbf{dx}_1 \times \mathbf{dx}_2|} + \frac{\mathbf{dx}_2 \times \mathbf{dx}_3}{|\mathbf{dx}_2 \times \mathbf{dx}_3|} + \frac{\mathbf{dx}_3 \times \mathbf{dx}_4}{|\mathbf{dx}_3 \times \mathbf{dx}_4|} + \frac{\mathbf{dx}_4 \times \mathbf{dx}_1}{|\mathbf{dx}_4 \times \mathbf{dx}_1|} \right) 
\]

(15)

and (comparing Equations (12)–(14)) the current in this direction as

\[
J_n^{(1)} = \mathbf{I}/A.
\]

(16)

In the same way, we can define \( J_n^{(2)} \) and \( J_n^{(3)} \) perpendicular to the other two mesh surfaces, with normal vectors \( \mathbf{n}^{(2)} \) and \( \mathbf{n}^{(3)} \), that pass through \( X_{i,j,k} \).

Now, the \( J_n^{(p)} \) are projections of the current we require (denoted \( J_p \)) in such a way that

\[
\mathbf{n}^{(p)} \cdot \mathbf{J}_i = J_n^{(p)}, \quad p = 1, 2, 3.
\]

(17)

Denoting by \( \mathcal{N} \) the matrix whose rows are the row vectors \( \mathbf{n}^{(1)}, \mathbf{n}^{(2)}, \) and \( \mathbf{n}^{(3)} \), and by \( \mathbf{J}_n \) the vector with components \( J_n^{(1)}, J_n^{(2)}, J_n^{(3)} \) we can re-write Equation (17) as \( \mathcal{N} \mathbf{J}_n = \mathbf{J}_i \).

Finally, we obtain the current at point \( X_{i,j,k} \) via

\[
\mathbf{J}_i = \mathcal{N}^{-1} \mathbf{J}_n
\]

(18)

Note that \( \mathcal{N} \) is always invertible assuming that the \( \mathbf{n}^{(p)} \) are linearly independent, i.e., as long as the grid cells have nonzero volume. It is straightforward to verify that this procedure reduces to the standard second-order centered difference expression in each direction for a rectangular (undeformed) mesh.

Since this method makes use of Stokes’ theorem, which is “topological” in the sense that it does not depend on a deformation of the loop we are integrating over, we expect the method to be more robust and accurate for deformed grids (see also Hyman & Shashkov 1997). The algorithm has not yet been implemented in the full numerical scheme, as it requires the method to be more robust and accurate for deformed grids.

4. COMPARISON OF METHODS

We now return to the test problem with two equal and opposite twists centered on \( z = \pm z_0 \) described above.

4.1. Relaxing Toward a Rectangular Mesh

First, to demonstrate that in principle the original (second-order) relaxation scheme is sound, we perform the relaxation “experiment” in the following way. We begin with a uniform field \( \mathbf{B} = b_0 \hat{z} \) on a uniform mesh, and then deform this mesh in such a way that the resulting magnetic field is \( T^2 \) (with \( \phi = \pm \pi \)). This configuration is then relaxed, to a level where the Lorentz force calculated by the numerical scheme is reduced to \( J \times \mathbf{B} = 10^{-6} \). The equilibrium field corresponding to \( T^2 \) is the uniform field \( \mathbf{B} = b_0 \hat{z} \), and in this case (due to the frozen-in condition) the straight field should correspond to the rectangular field.
mesh. (The choice $|\phi| = \pi$ is motivated by the braiding example discussed in Wilmot-Smith et al. 2009 since this is the minimum level of twist which yields a magnetic field whose field lines are truly braided in that case.)

To diagnose the success of the relaxation, referring back to Equation (9), we calculate

$$
\epsilon^* = d \frac{\Delta a^*}{\Delta l},
$$

(19)

taking $d = \sqrt{2}$ (radius of twist regions) and $\Delta a^*$ and $\Delta l$ as the maximum change in $a^*$ over a given length, which occurs along the central field line—the $z$-axis—by symmetry. This expression puts a true value on the “quality” $\epsilon^*$ of the force-free approximation: for a given variation in $a^*$, $\epsilon^*$ provides a lower bound for the maximum value of $|J \times B|/|B|^2$ within our domain, for a current free of errors (i.e., setting $\delta = 0$ in Equation (9)). We obtain a value of $\epsilon^* = 9.5 \times 10^{-7}$ for resolution $N = 61$, demonstrating that discretization errors in the scheme are very small when the relaxed state has an approximately rectangular mesh.

### 4.2. Accuracy Test: Artificially Imposed Deformation

We now investigate the promise of the two extensions to the scheme described in the previous section. In our test case ($T^2$) the topology of the field is simple, and the equilibrium field known, permitting the above approach (i.e., relaxation toward a uniform mesh). However, to investigate magnetic fields with non-trivial topology we must approach the problem in a different way. We therefore return to the case where we begin with $T^2$ on a rectangular mesh, setting $\phi_i = \pm \pi$.

Performing the artificially imposed analytical (“untwisting”) deformation instead of relaxation, we see that derivatives calculated with the two new methods (fourth-order finite differences and the Stokes-based method) both give smaller maximum errors than with the original second-order scheme (see Table 2). For a less deformed mesh ($\psi = \pi/2$), the fourth-order finite differences perform better than the Stokes-based routine. However, for a more distorted mesh ($\psi = \pi$), the Stokes routine gives significantly lower errors than either of the finite-difference methods. It is particularly interesting to note that the errors for the Stokes-based method seem to scale relatively weakly with the mesh deformation, suggesting it to be a good choice for highly deformed meshes. Both new schemes, for reasonable resolution and levels of deformation ($N \geq 41$, $\psi = \pi$) give errors that are an order of magnitude lower than the original (second-order) method. This suggests these methods are worth pursuing, so we go on to perform relaxation simulations using them.

### Table 2

Errors in $J$ for Deformations with $\psi = \pi$ and $\pi/2$ (in Equation (11)) Using Second- and Fourth-Order Finite Differences (Two n.n./Four n.n., respectively) and the Stokes-Based Routine

| $N$ | $\psi = \pi$ | $\psi = \pi/2$ |
|-----|-------------|---------------|
|     | Two n.n. | Four n.n. | Stokes | Two n.n. | Four n.n. | Stokes |
| 21  | 670      | 222       | 22.1   | 73.8    | 12.2     | 18.7   |
| 41  | 144      | 19.4      | 5.85   | 10.6    | 1.01     | 4.71   |
| 61  | 64.7     | 4.14      | 2.69   | 4.99    | 0.551    | 2.13   |
| 81  | 37.0     | 1.61      | 1.57   | 2.83    | 0.898    | 1.20   |

Notes. $N^3$ is the mesh resolution. In each case, the value shown is the maximum relative percentage error in the domain over all components of $J$, i.e., $100 \times |J_i - J^*_i|/|J^*_i|_{\text{max}}$, n.n. = nearest neighbor.

### Table 3

Values of the Force-Free Quality Parameter $\epsilon^*$ and the Lorentz Force Calculated via the Stokes-Based Routine, $J \times B/|B|^2$, for Simulation Runs with Second- and Fourth-Order Finite Differences (Two n.n./Four n.n.) and Resolution $N^3$, to Two Significant Figures

| $N$ | $\epsilon^*$ | $J \times B/|B|^2$ | $\epsilon^*$ | $J \times B/|B|^2$ |
|-----|--------------|-------------------|--------------|-------------------|
| 21  | 0.11         | 0.053             | 0.048        | 0.063             |
| 41  | 0.054        | 0.046             | 0.0067       | 0.018             |
| 61  | 0.032        | 0.035             | 0.0042       | 0.0050            |
| 81  | 0.021        | 0.027             | 0.0015       | 0.0019            |

Note. Twist parameter $\phi_i = \pm \pi/2$.

### Table 4

As Table 3, with Twist Parameter $\phi_i = \pm \pi$

| $N$ | $\epsilon^*$ | $J \times B/|B|^2$ | $\epsilon^*$ | $J \times B/|B|^2$ |
|-----|--------------|-------------------|--------------|-------------------|
| 21  | 0.28         | 0.18              | 0.17         | 0.21              |
| 41  | 0.16         | 0.17              | 0.071        | 0.13              |
| 61  | 0.11         | 0.13              | 0.026        | 0.062             |
| 81  | 0.074        | 0.11              | 0.021        | 0.023             |

### 4.3. Relaxation with Initially Rectangular Mesh

We now leave the artificially imposed deformation, and relax $T^2$ using both the second- and fourth-order schemes. The relaxation is allowed to run until $|J \times B|$ as calculated by the relevant numerical scheme is reduced to $10^{-5}$. We compare this value with $\epsilon^*$ (defined by Equation (19)), and also the maximum value of the Lorentz force obtained by calculating $J$ via the Stokes-based routine, denoted $J_f$.

The results for two levels of initial twist ($\phi_i = \pm \pi$, $\phi_i = \pm \pi$) are displayed in Tables 3 and 4. A number of points are immediately clear. First, in no simulation do we approach the apparent value of $\epsilon^* = 10^{-5}$. However, the use of fourth-order rather than second-order finite differences improves the quality of the relaxed field by an order of magnitude in $\epsilon^*$ when $|\phi| = \pi/2$. Increasing the deformation in the final state (by increasing $|\phi|$ in the initial state $T^2$) has a strong adverse effect on the relaxation process. This is found to be because spurious (unphysical) current concentrations arise where none should reasonably be expected. Examining the corresponding mesh, we find that these “false” current regions appear where the grid is most distorted—see Figure 3, and compare with Figure 1(b). The “current shards” shown in Figure 3(b) actually intensify as the relaxation proceeds. We find that this is possible since $\nabla \cdot J$ (approximated by interpolating $J$ onto a uniform mesh) is not sufficiently close to zero. As a result there is no “return current” associated with these localized current regions, which might be expected to generate a Lorentz force that would act against the further intensification of the current shards (if they have no physical basis). In previous studies using such codes (e.g., Craig & Litvinenko 2005; Pontin & Craig 2005), intensification of $|J|$ as $|J \times B|$ decreased in time was associated with current singularities, so at first sight it appears that these current shards could naively be interpreted as “current sheets”, which would of course be unphysical. Note, however, that the most important signature of current singularity in previous studies was a power-law proportionality of the peak current with mesh resolution (for given $\epsilon$). We have found that in fact the current shards become
less intense as \( N \) is increased (as required by the convergence of the scheme), so there is a clear distinction between the two phenomena.

Finally, consider the values of \( \mathbf{J}_s \times \mathbf{B}/|\mathbf{B}|^2 \) we find for the relaxed fields (Tables 3 and 4). They are clearly of the same order as \( \epsilon^* \) (note that \( \epsilon^* \) based on \( \mathbf{J} \) from the numerical scheme and \( \epsilon^* \) based on \( \mathbf{J}_s \) are of the same order), and thus it seems that an implementation involving the Stokes-based routine has the capacity to yield a magnetic field that is much closer to being force-free (with lower \( \epsilon^* \)). This is illustrated in Figure 4. We consider the observed value of \( \alpha^* \) based on Equation (6) (dashed line), compared with a maximum allowable value for \( \alpha^* \). Since \( \alpha^* \) is antisymmetric about \( z = 0 \), the maximum value allowed, based on Equation (9), is

\[
\alpha_{\text{max}}^* = \frac{L\epsilon}{d},
\]

and we take \( |\mathbf{B}| = 1 \) and \( d = \sqrt{2} \) as before, and \( L = 20 \), the length of the domain. Then we obtain the maximum possible \( \alpha^* \) by taking \( \epsilon \) to be the maximum value during the relaxation of \( \mathbf{J} \times \mathbf{B} \) (solid line) or \( \mathbf{J}_s \times \mathbf{B} \) (dot-dashed). We see that very early in the relaxation the actual value of \( \alpha^* \) becomes greater than the maximum allowed by \( \mathbf{J} \times \mathbf{B} \) from the numerical scheme (second-order). Moreover, the discrepancy grows steadily. However, \( \alpha^* \) always remains less than the maximum allowed by the Stokes-based method, implying that this may be a more sound method to calculate the current and resulting Lorentz force.

5. CONCLUSIONS

Force-free magnetic fields are important in many astrophysical applications. Determining the properties of such force-free fields—especially smoothness and stability properties—is key to understanding energy release processes that heat the plasma and lead to dynamic events such as flares in the solar corona. We have investigated the properties of different relaxation procedures for determining force-free fields based on a Lagrangian mesh approach. These techniques have previously been shown to have many powerful and advantageous properties. Previous understanding was that such schemes would iteratively converge (i.e., \( \mathbf{J} \times \mathbf{B} \) decreasing monotonically to a given level) up to a certain degree of mesh deformation. Beyond this level of mesh deformation the scheme no longer converges (\( \mathbf{J} \times \mathbf{B} \) oscillates or grows), and it is this phenomenon that was thought to limit the method. However, we have shown above that even when the numerical scheme iteratively converges, the accuracy of the force-free approximation can become seriously compromised for even “moderate” mesh deformations. This error is an accumulation of numerical discretization errors resulting from the calculation of \( \mathbf{J} \) via combinations of first and second derivatives of the mesh deformation Jacobian—which are calculated using finite differences. The result is that neither \( \mathbf{J} = \nabla \times \mathbf{B} \) nor subsequently \( \nabla \cdot \mathbf{J} = 0 \) are well satisfied at practical levels of the numerical resolution.

It was demonstrated that a result of the breaking of the solenoidal condition for \( \mathbf{J} \) can be the development of spurious (unphysical) current structures. However, we note that these rogue currents do diminish with resolution (\( N \)), so when using these schemes this property should always be checked where
possible. We expect that, as a result, if it were possible to systematically increase $N$ indefinitely the rogue currents would eventually vanish. In other words, the real problem is that the iterative convergence (i.e., monotonic decrease of $|J \times B|$ with $t$) is not compromised by the rogue currents but the real convergence to a correct solution is severely impaired.

A force-free field is defined by $\nabla \times B = \alpha B$. One key result of this equation is that $\alpha$ must be constant along magnetic field lines. We therefore argued that a correct diagnostic to measure the quality of a force-free approximation is the constancy of the parameter $\alpha^* = J_1/|B|$ along field lines. An appropriate normalization is given in Equation (19). The results of our investigations suggest that for Lagrangian schemes the $|J \times B|$ measure does not provide a good indicator of true convergence—a better measure is the constancy of $\alpha^*$ along $B$. We note that other authors have proposed measures other than the maximum of $J \times B$ for testing a force-free approximation—for example, Wheatland et al. (2000) introduced the “mean current-weighted angle between $J$ and $B$.” However, calculation of this measure still relies upon the value of $J \times B$ in the numerical scheme, and in the present scenario we have shown that the errors arise not because $J$ and $B$ are not parallel, but because $\nabla \times J \neq 0$.

Since errors in the (Lagrangian) numerical scheme investigated here arise as the mesh becomes increasingly distorted, a natural choice is to begin with a nonequilibrium field on a non-rectangular mesh, and relax toward a (perhaps approximately) rectangular one. However, this approach is not feasible if the field has complex topology. In the case of a braided field—which is of particular interest to the theory of the solar corona—we find that for our realization of such a field (Wilmot-Smith et al. 2009) there is no escaping having at least a moderately distorted mesh in the final state.

We proposed two possible extensions to the numerical method. The first was to increase the order of the finite differences used. It was found that for certain levels of deformation this can give an order of magnitude improvement in the quality of the force-free approximation obtained. It is therefore certainly a good approach to use in some circumstances. As the mesh became more and more highly deformed, the advantage of the scheme with fourth-order finite differences was lost for our test case T2. Furthermore, we found that for relaxation of the braided field described in Wilmot-Smith et al. (2009) no appreciable improvement arose from using the fourth-order scheme.

The other extension that we proposed to the scheme seems very promising. In Section 3.2 we presented an algorithm for calculating the curl of a vector field on an arbitrary mesh, based on Stokes’ theorem. For increasing levels of mesh deformation, this performed progressively better than the finite-difference methods. What’s more, in all of our tests the resultant Lorentz force $J \times B$ had lower errors than that calculated by the traditional finite difference. In order for a relaxation experiment to remain accurate as it proceeds, the maximum allowed value of $\alpha^*$ based on $J \times B$ (see Equation (20)) must always remain greater than the maximum observed value of $\alpha^*$. We found that this is the case for the Stokes-based $\alpha^*$ down to at least an order of magnitude lower in $J \times B$ than for the finite-difference methods (see Figure 4).

All of the above leads us to believe that the Stokes-based algorithm is a highly promising one for improving the accuracy of Lagrangian relaxation schemes. At present it has not been implemented (i.e., the code does not act to minimize $J_1/|B|$) because this requires a complete rewriting of the implicit (ADI) time stepping, and a simple explicit implementation turns out to be prohibitively computationally expensive. However, our intended next step in this investigation is to implement this scheme, either by introducing the Stokes-based current calculation as a correction term in the existing scheme or by employing a more sophisticated explicit time stepping to reduce the computational expense to acceptable levels. We note that while the algorithm at present only uses two nearest-neighbor points in each direction, it could be extended to include further line integrals as corrections to the present formula for $J_1$, in much the same way as is done by increasing the order of finite-difference derivatives.

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REFERENCES

Amari, T., Aly, J. J., Luciani, J. F., Boulemezoud, T. Z., & Mikic, Z. 1997, Sol. Phys., 174, 129
Anzer, U. 1968, Sol. Phys., 3, 298
Burlaga, L. F. 1988, J. Geophys. Res., 93, 7217
Chodura, R., & Schlueter, A. 1981, J. Comput. Phys., 41, 68
Craig, I. J. D., & Litvinenko, Y. E. 2005, Phys. Plasmas, 12, 032301
Craig, I. J. D., & Sneyd, A. D. 1986, ApJ, 311, 451
Craig, I. J. D., & Sneyd, A. D. 1990, ApJ, 357, 653
Craig, I. J. D., & Sneyd, A. D. 2005, Sol. Phys., 232, 41
Frank, J., King, A., & Raine, D. J. 2002, Accretion Power in Astrophysics (Cambridge: Cambridge Univ. Press)
Hyman, J. M., & Shashkov, M. 1997, Comput. Math. Appl., 33, 81
Hyman, J. M., & Shashkov, M. 1999, J. Comput. Phys., 151, 881
Longbottom, A. W., Rickard, G. J., Craig, I. J. D., & Sneyd, A. D. 1998, ApJ, 500, 471
Longcope, D. W., & Strauss, H. R. 1994, ApJ, 437, 851
McKinney, J. C. 2006, Mon. Not. R. Astron. Soc., 368, L30
Mestel, L. 1973, Astrophys. Space Sci., 24, 289
Moffatt, H. K. 1978, Magnetic Field Generation in Electrically Conducting Fluids (Cambridge: Cambridge Univ. Press)
Parker, E. N. 1972, ApJ, 174, 499
Pontin, D. I., & Craig, I. J. D. 2005, Phys. Plasmas, 12, 072112
Schrijver, C. J., et al. 2006, Sol. Phys., 235, 161
Udzenisky, D. A., Königl, A., & Litwin, C. 2002, ApJ, 565, 1191
van Ballegooijen, A. A. 1985, ApJ, 298, 421
Wheatland, M. S., Sturrock, P. A., & Roumeliotis, G. 2000, ApJ, 540, 1150
Wilmot-Smith, A., Hornig, G., & Pontin, D. I. 2009, ApJ, 696, 1339
Zweibel, E. G., & Li, H.-S. 1987, ApJ, 312, 423