Efficient magnetohydrodynamic simulations on graphics processing units with CUDA

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Abstract

Magnetohydrodynamic (MHD) simulations based on the ideal MHD equations have become a powerful tool for modeling phenomena in a wide range of applications including laboratory, astrophysical, and space plasmas. In general, high-resolution methods for solving the ideal MHD equations are computationally expensive and Beowulf clusters or even supercomputers are often used to run the codes that implemented these methods. With the advent of the Compute Unified Device Architecture (CUDA), modern graphics processing units (GPUs) provide an alternative approach to parallel computing for scientific simulations. In this paper we present, to the author’s knowledge, the first implementation of MHD simulations entirely on GPUs with CUDA, named \textit{GPU-MHD}, to accelerate the simulation process. \textit{GPU-MHD} supports both single and double precision computation. A series of numerical tests have been performed to validate the correctness of our code. Accuracy evaluation by comparing single and double precision computation results is also given. Performance measurements of both single and double precision are conducted. These measurements show that our GPU-based implementation achieves speedups (in single precision) of about 10 (1D problems with 4096 grid points), 200 (2D problems with 1024\textsuperscript{2} grid points), and 84 (3D problems with 128\textsuperscript{3} grid points), respectively, compared to the corresponding serial CPU MHD implementation. For double precision computation, \textit{GPU-MHD} still can achieve about 60\% speed of the corresponding single precision computation. In addition, we extend \textit{GPU-MHD} to support the visualization of the simulation results and thus the whole MHD simulation and visualization process can be performed entirely on GPUs.

\textit{Key words:} MHD simulations, GPUs, CUDA, Parallel computing
1 Introduction

Magnetohydrodynamic (MHD) equations can be used in modeling phenomena in a wide range of applications including laboratory [5], astrophysical [39], and space plasmas [9]. However, they form a nonlinear system of hyperbolic conservation laws, which is so complex and high-resolution methods are necessary to solve them in order to capture shock waves and other discontinuities. These high-resolution methods are in general computationally expensive and parallel computational resources such as Beowulf clusters or even supercomputers are often utilized to run the codes that implemented these methods [22] [13] [18] [14] [50].

In the last few years, the rapid development of graphics processing units (GPUs) makes them more powerful in performance and more programmable in functionality. By comparing the computational power of GPUs and CPUs, GPUs exceed CPUs by orders of magnitude. The theoretical peak performance of the current consumer graphics card NVIDIA GeForce GTX 295 (with two GPUs) is 1788.48G floating-point operations per second (FLOPS) per GPU in single precision while a CPU (Core 2 Quad Q9650 — 3.0 GHz) gives a peak performance of around 96GFLOPS in single precision. The release of the Compute Unified Device Architecture (CUDA) [24] hardware and software architecture is the culmination of such development. With CUDA, one can directly exploit a GPU as a data-parallel computing device by programming with the standard C language and avoid working with a high-level shading language such as Cg [21], which requires a significant amount of graphics specific knowledge and was previously used for performing computation on GPUs. Detailed performance studies on GPUs with CUDA can be found in [4] and [32].

CUDA is a general purpose parallel computing architecture developed by NVIDIA. It includes the CUDA Instruction Set Architecture (ISA) and the parallel compute engine. An extension to C programming language and its compiler are provided, making the parallelism and high computational power of GPUs can be used not only for rendering and shading, but also for solving many computationally intensive problems in a fraction of the time required on a CPU. CUDA also provides basic linear algebra subroutines (CUBLAS) and fast Fourier transform (CUFFT) libraries to leverage GPUs’ capabilities. These libraries release developers from rebuilding the frequently used basic operations such as matrix multiplication. Graphics cards from G8x series support the CUDA programming mode; and the latest generation of NVIDIA GPUs (GT2x0 series or later) unifies vertex and fragment processors and provides shared memory for interprocessor communication.

A increasing number of new GPU implementations with CUDA in different astrophysical simulations have been proposed. Belleman et al. [2] re-implemented
the direct gravitational $N$-body simulations on GPUs using CUDA. For $N \gtrsim 10^5$, they reported a speedup of about 100 compared to the host CPU and about the same speed as the GRAPE-6Af. A library Sapporo for performing high precision gravitational $N$-body simulations was developed on GPUs by Gaburov et al. [10]. This library achieved twice as fast as commonly used GRAPE6A/GRAPE6-BLX cards. Stantchev et al. [37] [38] implemented a Particle-In Cell (PIC) code on GPUs for plasmas simulations and visualizations and demonstrated a speedup of 11-22 for different grid sizes. Sainio [31] presented an accelerated GPU cosmological lattice program for solving the evolution of interacting scalar fields in an expanding universe, achieving speedups between one and two orders of magnitude in single precision. In the above works, no discussion on using double precision on GPUs was reported. In MHD simulations, the support of double precision is important, especially for nonlinear problems. We will evaluate the performance and accuracy of double precision on GPUs in this work.

In this paper, we present an efficient implementation to accelerate computation of MHD simulations on GPUs, called GPU-MHD. To our knowledge, this is the first work describing MHD simulations on GPUs in detail. The goal of our work is to perform a pilot study on numerically solving the ideal MHD equations on GPUs. In addition, the trend of today’s chip design is moving to streaming and massively parallel processor models, developing new MHD codes to exploit such architecture is essential. GPU-MHD can be easily ported to other many-core platforms such as Intel’s upcoming Larrabee [34], making it more flexible for the user’s choice of hardware. This paper is organized as follows: A brief description of the CUDA programming model is given in Section 2. The numerical scheme in which GPU-MHD adopted is presented in Section 3. In Section 4, we present the GPU implementation in detail. Numerical tests are given in Section 5. Accuracy evaluation by comparing single and double precision computation results is given in Section 6. Performance measurements are reported in Section 7 and visualization of the simulation results is described in Section 8. We conclude our work and point out the future work in Section 9.

2 A brief description of the CUDA

The Compute Unified Device Architecture (CUDA) was introduced by NVIDIA as a general purpose parallel computing architecture, which includes GPU hardware architecture as well as software components (CUDA compiler and the system drivers and libraries). The CUDA programming model [24] consists of functions, called kernels, which can be executed simultaneously by a large number of lightweight threads on the GPU. These threads are grouped into one-, two-, or three-dimensional thread blocks, which are further organized into
one- or two-dimensional grids. Only threads in the same block can share data and synchronize with each other during execution. Thread blocks are independent of each other and can be executed in any other. A graphics card that supports CUDA, for example, the GT200 GPU [20], consisting of 30 streaming multiprocessors (SMs). Each multiprocessor consists of 8 streaming processors (SPs), providing a total of 240 SPs. Threads are grouped into batches of 32 called warps which are executed in single instruction multiple data (SIMD) fashion independently. Threads within a warp execute a common instruction at a time.

For memory access and usage, there are four types of memory, namely, global memory, constant memory, texture memory as well as shared memory. Global memory has a separate address space for obtaining data from the host CPU’s main memory through the PCIE bus, which is about 102 GB/sec in the GT200 GPU. Any valued stored in global memory can be accessed by all SMs via load and store instructions. Constant memory and texture memory are cached, read-only and shared between SPs. Constants that are kept unchange during kernel execution may be stored in constant memory. Built-in linear interpolation is available in texture memory. Shared memory is limited (16 kB for GT200 GPU) and shared between all SPs in a MP. For detailed information concerning memory optimizations, we refer the reader to “CUDA Best Practice Guide” [26].

Double precision is one important concern in many computational physics applications, however, supports of double precision are limited to the NIVDIA cards having Compute Capability 1.3 (See Appendix A in [24]) such as the GTX 260, GTX 280, Quadro FX 5800 (contains one GT200 GPU), and Tesla C1060 (contains one GT200 GPU) and S1070 (contains four GT200 GPUs). In GT200 GPU, there are eight single precision floating point (FP32) arithmetic logic units (ALUs) (one per SP) in SM, but only one double precision floating point (FP64) ALU (shared by eight SPs). The theoretical peak performance of GT200 GPU is 936 GFLOPS in single precision and 78 GFLOPS in double precision. In CUDA, double precision is disabled by default, ensuring that all double numbers are silently converted into float numbers inside kernels and any double precision calculations computed are incorrect. In order to use double precision floating point numbers, we need to call nvcc: “-arch = sm_13”. The flag “-arch = sm_13” in the command tells “nvcc” to use the Compute Capability 1.3 which means enabling the double precision support.

In Sections 6 and 7 we compare the accuracy and actual performance of GPU-MHD in single and double precision.
3 Numerical scheme

The ideal MHD equations with the assumption of the magnetic permeability $\mu = 1$ can be represented as hyperbolic system of conservation laws as follows [11]

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0 \quad (1)$$

$$\frac{\partial \rho \mathbf{v}}{\partial t} + \nabla \cdot (\rho \mathbf{v} \mathbf{v} - \mathbf{B} \mathbf{B}) + \nabla P^* = 0 \quad (2)$$

$$\frac{\partial \mathbf{B}}{\partial t} - \nabla \times (\mathbf{v} \times \mathbf{B}) = 0 \quad (3)$$

$$\frac{\partial E}{\partial t} + \nabla \cdot ((E + P^*) \mathbf{v} - \mathbf{B} (\mathbf{B} \cdot \mathbf{v})) = 0 \quad (4)$$

Here, $\rho$ is the mass density, $\rho \mathbf{v}$ the momentum density, $\mathbf{B}$ the magnetic field, and $E$ the total energy density. The total pressure $P^* \equiv P + \frac{B^2}{2}$ where $P$ is the gas pressure that satisfies the equation of state, $P \equiv (\gamma - 1)(E - \rho \frac{v^2}{2} - \frac{B^2}{2})$. In addition, the MHD equations should obey the divergence-free constraint $\nabla \cdot \mathbf{B} = 0$.

Over the last few decades, there has been a dramatic increase in the number of publications on numerical solution of ideal MHD equations. In particular the development of shock-capturing numerical methods for ideal MHD equations. We do not provide an exhaustive review of the literature here. A comprehensive treatment of numerical solution of MHD equations can be found in [17], for example. Pen et al. [28] proposed a free, fast, simple, and efficient total variation diminishing (TVD) MHD code featuring modern high-resolution shock capturing on a regular Cartesian grid. This code is second-order accuracy in space and time and enforces the $\nabla \cdot \mathbf{B} = 0$ constraint to machine precision. Due to these advantages and convenience for GPU verse CPU comparison, the underlying numerical scheme in GPU-MHD is based on this work. A detailed comparison of shock capturing MHD codes can be found in [43], for example. We plan to explore other recent high-order Godunov schemes such as [19] and [41] for GPU-MHD as our future work.

We briefly review the numerical scheme [28] we adopted in GPU-MHD here. In this numerical scheme, the magnetic field is held fixed first and then the fluid variables are updated. A reverse procedure is then performed to complete a one time step. Three dimensional problem is split into one-dimensional sub-problems by using a Strang-type directional splitting [42].

Firstly, we describe the fluid update step in which the fluid variables are updated while holding the magnetic field fixed. The magnetic field is interpolated
to cell centers for second-order accuracy. By considering the advection along the $x$ direction, the ideal MHD equations can be written in flux-conservative vector form as

$$\frac{\partial \mathbf{u}}{\partial t} + \frac{\partial \mathbf{F}(u)}{\partial x} = 0 \quad (5)$$

where the flux vector is given by

$$\mathbf{F} = \begin{pmatrix} \rho v_x \\ \rho v_x^2 + P^* - B_x^2 \\ \rho v_x v_y - B_x B_y \\ \rho v_x v_z - B_x B_z \\ (E + P^*)v_x - B_x \mathbf{B} \cdot \mathbf{v} \end{pmatrix} \quad (6)$$

Equation (5) is then solved by Jin & Xin’s relaxing TVD method [15]. With this method, a new variable $\mathbf{w} = \mathbf{F}(\mathbf{u})/c$ is defined, where $c(x, t)$ is a free positive function called the flux freezing speed. For ideal MHD equations, we have $\mathbf{u} = (u_1, u_2, u_3, u_4, u_5) = (\rho, \rho v_x, \rho v_y, \rho v_z, E)$ and equations

$$\frac{\partial \mathbf{u}}{\partial t} + \frac{\partial}{\partial x}(c \mathbf{w}) = 0 \quad (7)$$

$$\frac{\partial \mathbf{w}}{\partial t} + \frac{\partial}{\partial x}(c \mathbf{u}) = 0 \quad (8)$$

These equations can be decoupled through a change of left- and right-moving variables $\mathbf{u}^R = (\mathbf{u} + \mathbf{w})/2$ and $\mathbf{u}^L = (\mathbf{u} - \mathbf{w})/2$

$$\frac{\partial \mathbf{u}^R}{\partial t} + \frac{\partial}{\partial x}(c \mathbf{u}^R) = 0 \quad (9)$$

$$\frac{\partial \mathbf{u}^L}{\partial t} - \frac{\partial}{\partial x}(c \mathbf{u}^L) = 0 \quad (10)$$

The above pair of equations is then solved by an upwind scheme, separately for right- and left-moving waves, using cell-centered fluxes. Second-order spatial accuracy is achieved by interpolating of fluxes onto cell boundaries using a monotone upwind schemes for conservation laws (MUSCL) [47] with the help of a flux limiter. Runge-Kutta scheme is used to achieve second-order accuracy of time integration.

We denote $\mathbf{u}_n^t$ as the cell-centered values of the cell $n$ at time $t$, $\mathbf{F}_n^t$ as the cell-centered flux in cell $n$. As an example, we consider the positive advection
velocity, negative direction can be obtained in a similar way. We obtain the first-order upwind flux $F^{(1),t}_{n+1/2}$ from the averaged flux $F^t_n$ in cell $n$. Two second-order flux corrections can be defined using three local cell-centered fluxes as follows

$$
\Delta F^{L,t}_{n+1/2} = \frac{F^t_n - F^t_{n-1}}{2} \quad (11)
$$

$$
\Delta F^{R,t}_{n+1/2} = \frac{F^t_{n+1} - F^t_n}{2} \quad (12)
$$

When the corrections have opposite signs, there is no second-order correction in the case of near extrema. With the aid of a flux limiter $\phi$ we then get the second-order correction

$$
\Delta F^t_{n+1/2} = \phi(\Delta F^{L,t}_{n+1/2}, \Delta F^{R,t}_{n+1/2}) \quad (13)
$$

The van Leer limiter [46]

$$
\text{vanleer}(a, b) = \frac{2ab}{a + b} \quad (14)
$$

is used in GPU-MHD. By adding the second-order correction to the first-order fluxes we obtain second-order fluxes. For example, the second-order accurate right-moving flux $F^{R,t}_{n+1/2}$ can be calculated

$$
F^{R,t}_{n+1/2} = F^t_n + \Delta F^{t}_{n+1/2} \quad (15)
$$

The time integration is performed by calculating the fluxes $F(u^t_n)$ and the freezing speed $c^t_n$ in the first half time step is given as follows

$$
u^t_{n+\Delta t/2} = u^t_n - \left( \frac{F^t_{n+1/2} - F^t_{n-1/2}}{\Delta x} \right) \frac{\Delta t}{2} \quad (16)
$$

where $F^t_{n+1/2} = F^{R,t}_{n+1/2} - F^{L,t}_{n+1/2}$ is computed by the first-order upwind scheme. By using the second-order TVD scheme on $u^t_{n+\Delta t/2}$, we obtain the full time step $u^t_{n+\Delta t}$

$$
u^{t+\Delta t}_n = u^t_n - \left( \frac{F^{t+\Delta t/2}_{n+1/2} - F^{t+\Delta t/2}_{n-1/2}}{\Delta x} \right) \Delta t , \quad (17)
$$

To keep the TVD condition, the flux freezing speed $c$ is the maximum speed information can travel and should be set to $|v_x| + (\gamma p/\rho + B^2/\rho)^{1/2}$ as the maximum speed of the fast MHD wave over all directions is chosen. As the time integration is implemented using a second-order Runge-Kutta scheme,
the time step is determined by satisfying the CFL condition

\[
\Delta t = \frac{cfl}{c_{\text{max}}}
\]

(18)

where \(cfl\) is the Courant-Number and \(cfl \lesssim 1\) is generally set to \(cfl \approx 0.7\) for stability, and \(B\) is the magnitude of the magnetic field. Constrained transport (CT) [8] is used to keep the \(\nabla \cdot B = 0\) to machine precision. Therefore, the magnetic field is defined on cell faces and it is represented in arrays [28]

\[
\begin{align*}
B_x(i, j, k) &= (B_x)_{i-1/2,j,k} \\
B_y(i, j, k) &= (B_y)_{i,j-1/2,k} \\
B_z(i, j, k) &= (B_z)_{i,j,k-1/2}
\end{align*}
\]

(19)

where the cell centers are denoted by \((i, j, k) \equiv (x_i, y_j, z_k)\), and faces by \((i \pm 1/2, j, k)\), \((i, j \pm 1/2, k)\), and \((i, j, k \pm 1/2)\), etc. The cells have unit width for convenience.

Secondly, we describe the update of the magnetic field in separate two-dimensional advection-constraint steps along \(x\)-direction while holding the fluid variables fixed. The magnetic field updates along \(y\) and \(z\)-directions can be handled in a similar matter. We follow the expressions used in [16]. For example, we can calculate the averaging of \(v\) along \(x\) direction as follows

\[
(v_x)_{i,j+1/2,k} = \frac{1}{4} \left[ (v_x)_{i+1,j+1/2,k} + 2 (v_x)_{i,j+1/2,k} + (v_x)_{i-1,j+1/2,k} \right]
\]

(20)

A first-order accurate flux is then obtained by

\[
\begin{align*}
(v_x B_y)_{i+1/2,j+1/2,k} &= \begin{cases} 
(v_x B_y)_{i,j+1/2,k} & \text{, } (v_x)_{i+1/2,j+1/2,k} > 0 \\
(v_x B_y)_{i+1,j+1/2,k} & \text{, } (v_x)_{i+1/2,j+1/2,k} \leq 0
\end{cases}
\end{align*}
\]

(21)

where the velocity average is

\[
\begin{align*}
(v_x)_{i+1/2,j+1/2,k} &= \frac{1}{2} \left[ (v_x)_{i,j+1/2,k} + (v_x)_{i+1,j+1/2,k} \right]
\end{align*}
\]

(22)

\(B_x\) is updated by constructing a second-order-accurate upwind electromotive force (EMF) \(v_y B_x\) using Jin & Xin’s relaxing TVD method [15] in the advection step. Then this same EMF is immediately used to update \(B_y\) in the constraint step.

Extension to three dimensions can be done through a Strang-type directional splitting [42]. Equation (5) is dimensionally split into three separate one-dimensional equations. For a time step \(\Delta t\), let \(\text{fluid}_x\) be the fluid update
along $x$, $B_{x\rightarrow y}$ be the update of $B_x$ along $y$, and $L_i$ be the update operator of $u^t$ to $u^{t+\Delta t}$ by including the flux along $i$ direction. Each $L_i$ includes three update operations in sequence, for example, $L_x$ includes fluid$_x$, $B_{y\rightarrow x}$, and $B_{z\rightarrow x}$. A forward sweep and a reverse sweep are defined as $u^{t+\Delta t} = L_z L_y L_x u^t$ and $u^{t+2\Delta t} = L_x L_y L_z u^{t+\Delta t}$, respectively. A complete update combines a forward sweep and reverse sweep. The dimensional splitting of the relaxing TVD can be expressed as follows [45]

$$u^{t_2} = u^{t_1+2\Delta t_1} = L_x L_y L_z L_y L_x u^{t_1}$$

(23)

$$u^{t_3} = u^{t_2+2\Delta t_2} = L_z L_x L_y L_y L_z u^{t_2}$$

(24)

$$u^{t_4} = u^{t_3+2\Delta t_3} = L_y L_z L_x L_z L_y u^{t_3}$$

(25)

where $\Delta t_1$, $\Delta t_2$, and $\Delta t_3$ are sequential time steps after each double sweep.

4 GPU implementation

In this section, we provide the implementation details of GPU-MHD. With GPU-MHD, all computations are performed entirely on GPUs and all data is stored in the GRAM of the graphics card. Currently, GPU-MHD works on a regular Cartesian grid and supports both single and double precision modes.

4.1 Data storage arrangement

The most intuitive way to write a parallel program to solve a multidimensional problem is to use multidimensional arrays for data storage and multidimensional threads for computation. However, the ability of the current CUDA is limited in supporting multidimensional threads, therefore, we could not implement our code in such a straightforward way. Especially in three dimensions or higher dimensions, there are still some limitations in handling multidimensional arrays and multidimensional threads. As a result, the most primitive way is to store data in one-dimension and perform parallel computation with one-dimension threads. By using an indexing technique, our storage and threading method can be extended to to solve multidimensional problems. Our data storage arrangement is expressed in Fig. 1 and in Equations (26) to (28).

$$\begin{align*}
INDEX_x &= index \mod (SIZE_y \times SIZE_z) \\
INDEX_y &= [index \mod (SIZE_y \times SIZE_z)]/SIZE_z \quad (26) \\
INDEX_z &= index \mod SIZE_z
\end{align*}$$
Fig. 1. Mapping from 3D array to 1D array in column major.

\[ INDEX_x \pm 1 = index \pm (SIZE_y \times SIZE_z) \]  
\[ INDEX_y \pm 1 = index \pm SIZE_z \]  
\[ INDEX_z \pm 1 = index \pm 1 \]  

Here \( INDEX_x, INDEX_y, \) and \( INDEX_z \) are the indexes of a 3D matrix. \( index \) is the 1D index used in GPU-MHD. \( SIZE_y \) and \( SIZE_z \) are the matrix size (number of grid points in our study) of a 3D matrix.

Equation (26) expresses the mapping of three-dimensional indexes to one-dimensional indexes. Equations (28) to (29) express the shift operations. Shift operations are very important in numerical solution of conservation laws because some calculations are based on the neighboring grid points.

4.2 Program flow

A “CUDA kernel” is a function running on GPU [24]. Noted that the CUDA kernel will process all grid points in parallel, therefore, a For instruction is no needed for going through all grid points. GPU-MHD includes the following steps:

1. CUDA initialization
2. Setup the initialize condition for the specified MHD problem:  
\( \mathbf{u} = (u_1, u_2, u_3, u_4, u_5) \) of all grid points, \( \mathbf{B} = (B_x, B_y, B_z) \) of cell faces, and set parameters such as time \( t \), etc.
3. Copy the initialize condition \( \mathbf{u}, \mathbf{B} \) to device memory (CUDA global memory)
4. For all grid points, calculate the \( c_{\text{max}} \) by Equation (18) (implemented with a CUDA kernel)
5. Use \texttt{cublasIsamax} (in single precision mode) function or \texttt{cublasIdamax}
(in double precision mode) function of the CUBLAS library to find out the maximum value of all \( c_{\text{max}} \), and then determine the \( \Delta t \).

6. Since the value of \( \Delta t \) is stored in device memory, read it back to host memory (RAM).

7. Sweeping operations of the relaxing TVD (Calculation of the \( L_i, i = x, y, z \), implemented with several CUDA kernels, will be explained in the next subsection).

8. \( t = t + 2\Delta t \)

9. If \( t \) reaches the target time, go to next step
   
   else repeats the procedure from step (4)

10. Read back data \( u, B \) to host memory

11. Output the result

The program flow of GPU-MHD is shown in Fig. 2. After the calculation of the CFL condition, the sweeping operations will be performed. The sweeping operation \( L_i \) will update both the fluid variables and orthogonal magnetic fields along \( i \) dimension. This is a core computation operation in the relaxing TVD scheme described in Section 3.

The CFL condition for the three-dimensional relaxing TVD scheme is obtained by Equation (18). The procedure is to calculate all the \( c_{\text{max}} \) of each grid point and find out the maximum value. In GPU-MHD, parallel computation power of CUDA is exploited to calculate the \( c_{\text{max}} \) of each grid point in parallel and all the \( c_{\text{max}} \) values are stored in a matrix. Then the cublasIsamax function is used (in double precision mode, the cublasIdamax function is used) to find out the maximum \( c_{\text{max}} \) of the matrix in parallel (called the reduction operation). The cublasIsamax function is provided in the CUBLAS library—a set of basic operations for vector and matrix provided by NVIDIA with the CUDA toolkit [24]. The reason we read the \( \Delta t \) back and store both \( \Delta t \) and \( t \) in host memory is due to the data in device memory cannot be printed out directly in the current CUDA version. This information is useful for checking if there is any problem during the simulation processing. The implementation of sweeping operations will be explained in the next subsection.

4.3 Sweeping operations

Before we start to describe the sweeping operations, consideration of memory arrangement is presented first in the following.

Implementing parallel computation using CUDA kernels is somewhat similar to parallel implementation on a CPU-cluster, but it is not the same. The major concern is the memory constrain in GPUs. CUDA makes parallel computation process on GPUs which can only access their graphics memory (GRAM).
Therefore, data must be stored in GRAM in order to be accessed by GPUs. There are several kinds of memory on graphics hardware including registers, local memory, shared memory, and global memory, etc., and they have different characteristics and usages [24], making memory management of CUDA quite different compared to parallel computation on a CPU-cluster. In addition, even the size of GRAM in a graphics card increases rapidly in newer models (for example, the latest NVIDIA graphics card — GeForce GTX 295 has 1.75G GRAM), but not all the capacity of GRAM can be used to store data arbitrarily. Shared memory and local memory are flexible to use, however, their sizes are very limited in a block and thus they cannot be used for storing data with large size. In general, numerical solution of conservation laws will generate many intermediate results (for example, $u^{t+\Delta t/2}$, $F$, $c$, $w$, etc.) during the computation process, these results should be stored for subsequent steps in the process. Therefore, global memory was mainly used in GPU-MHD.

After the maximum value of $c_{\text{max}}$ in Equation (18) is found, we can get the $\Delta t$ by determining the Courant-Number ($cfl$). The sequential step is the calculation of $L_i$ ($i = x, y, z$). The implementation of $L_i$ includes two parts: update the fluid variables and update the orthogonal magnetic fields. As an example, the process for calculating $L_x$ is shown in Fig. 3 where each block was implemented with one or several CUDA kernels. The process for calculating $L_y$ or $L_z$ is almost the same as $L_x$ except that the dimensional indexes are different.

Fig. 2. The flow chat of GPU-MHD.
The first part of the $L_x$ calculation process is $\text{fluid}_x$. The fluid variables will be updated along $x$. **Algorithm 1** shows the steps and GPU kernels of this process (the data of $u$ and $B$ are already copied to device memory), all the steps are processed on all grid points with CUDA kernels in parallel.

In this process, we have to calculate the magnetic fields of the grid point (Equation 19) first because all the magnetic fields are defined on the faces of the grid cell [28]. To update the fluid variables of $L_x$, the main process, which includes one or even several CUDA kernels, is to calculate the affect of the orthogonal magnetic fields to the fluid variables of Equations (6), (9) and (10). One such main process gives the flux of the $\Delta t/2$ step. After two main processes of flux calculation and the other difference calculations, the value of fluid — $u$ is updated from $u^t$ to $u^{t+\Delta t}$ in one $L_x$ process.

The second part of the $L_x$ calculation process is to update the orthogonal magnetic fields in $y$-dimension ($B_{y\rightarrow x}$), and $z$-dimension ($B_{z\rightarrow x}$) with the fluid along $x$-dimension. The strategy and implementation are similar to those in the first part but with a different algorithm for the orthogonal magnetic fields.

Fig. 3. Calculation process of $L_x$. 
Algorithm 1 Algorithm of fluid\textsubscript{x}, all equations and difference calculations are processed using CUDA kernels

1: load \textbf{u}, \textbf{B} and \Delta \textit{t}
2: memory allocation for the storage of the intermediate results: \textbf{B\textsubscript{temp}}, \textbf{u\textsubscript{temp}}, \textbf{flux\textsubscript{temp}}, \textbf{other\textsubscript{temp}}. (other\textsubscript{temp} includes the storage of \textbf{F}, \textbf{c}, \textbf{w}, etc)
3: \textbf{B\textsubscript{temp}} ← results obtained by Equation (19) with \textbf{B}, (\textbf{B} stored the magnetic field of the cell faces)
4: \textbf{other\textsubscript{temp}} ← results obtained by Equations (6) and (9) with \textbf{u}
5: \textbf{flux\textsubscript{temp}} ← the flux of a half time step: difference calculation (“First-Order Upwind Scheme of Fluid” CUDA kernels in Fig. 3) obtained by Equation (16) using other\textsubscript{temp}
6: \textbf{u\textsubscript{temp}} ← calculate the intermediate result (\textbf{u\textsuperscript{t+\Delta t/2}}) using Equation (16) with \textbf{u} and \textbf{flux\textsubscript{temp}}
7: \textbf{other\textsubscript{temp}} ← results obtained by Equations (6) and (9) with \textbf{u\textsubscript{temp}} (the same algorithm and same CUDA kernels in Step 4)
8: \textbf{flux\textsubscript{temp}} ← the flux of another half time step: difference calculation (“Second-Order TVD Scheme of Fluid” CUDA kernels in Fig. 3) obtained by Equation (17) and the limiter (Equation (14)) using other\textsubscript{temp}
9: calculate the result of \textbf{u\textsuperscript{t+\Delta t}} with \textbf{flux\textsubscript{temp}} using Equation (17) and save it back to \textbf{u}
10: free the storage of the intermediate results
11: (continue to the second part of \textit{L}\textsubscript{x}, update the orthogonal magnetic fields)

In Algorithm 1, the calculations in steps (4) to (9) are the steps for \textit{B}\textsubscript{y→x}, and steps (11) to (16) are the steps for \textit{B}\textsubscript{z→x}. The steps for \textit{B}\textsubscript{y→x} and \textit{B}\textsubscript{z→x} are almost the same, and the only different parts are the dimensional indexes of the difference calculations, and the affected magnetic fields: \textit{B}\textsubscript{y} and \textit{B}\textsubscript{z}. After the first part of \textit{L}\textsubscript{x} the fluid \textbf{u}\textsuperscript{t} is updated to \textbf{u}\textsuperscript{t+\Delta t}. This change of the fluid affects to the orthogonal magnetic fields. Therefore, the corresponding change (flux) of orthogonal magnetic fields can be calculated with the density and velocity of the updated fluid \textbf{u}\textsuperscript{t+\Delta t}. Then the orthogonal magnetic fields are also updated to \textbf{B}\textsubscript{y\textsuperscript{t+\Delta t}} and \textbf{B}\textsubscript{z\textsuperscript{t+\Delta t}}, and also, these changes give effects to \textbf{B}\textsubscript{x}.

After one process of \textit{L}\textsubscript{x}, both fluid and magnetic fields are updated to \textit{t + \Delta t} with the affect of the flow in \textit{x}-dimension. And a sweeping operation sequence includes two \textit{L}\textsubscript{x}, \textit{L}\textsubscript{y}, and \textit{L}\textsubscript{z} (see Equations (23)). So we actually get the updated fluid and magnetic fields of \textit{t + 2\Delta t} after one sweeping operation sequence. Note that the second \textit{L}\textsubscript{x} in the sequence is a reverse sweeping operation, the order of fluid\textsubscript{x}, \textit{B}\textsubscript{y→x} and \textit{B}\textsubscript{z→x} has to be reversed: \textit{B}\textsubscript{y→x} and \textit{B}\textsubscript{z→x} first, and fluid\textsubscript{x} second.

As we mentioned before, numerical solution of conservation laws needs lots of memory because there are many intermediate results generated during the computation process. These intermediate results should be stored for the next
Algorithm 2 Algorithm of \((B_y \rightarrow x)\) and \((B_z \rightarrow x)\), all equations and difference calculations are processed using CUDA kernels

1: (after the processes of fluid, we obtain an updated \(u\))
2: load \(u_1\) (density \(\rho\)), \(u_2\) (\(\rho v_x\)), \(B\) and \(\Delta t\)
3: memory allocation for the intermediate results: \(B_{\text{temp}}, \text{flux}_{\text{temp}}, \text{vx}_{\text{temp}}\)
4: \(\text{vx}_{\text{temp}} \leftarrow\) determine the fluid speed with the updated \(u_1\) and \(u_2\) in \(\text{fluid}_x\), with the difference calculated in \(y\)-dimension
5: \(\text{vx}_{\text{face}} \leftarrow\) Results obtained by Equation (20)
6: \(\text{flux}_{\text{temp}} \leftarrow\) the flux of a half time step: difference calculation of “flux of magnetic field in \(y\)-dimension” (“First-Order Upwind Scheme of Magnetic Field” CUDA kernels in Fig. 3) obtained by Equations (16) and (21)
7: \(B_{\text{temp}} \leftarrow\) calculate the intermediate result \((u^{t+\Delta t/2})\) by applying Equation (16) to \(B_y\) (not by applying Equation (16) to \(u\)) with \(B_y\) and \(\text{flux}_{\text{temp}}\)
8: \(\text{flux}_{\text{temp}} \leftarrow\) the flux of another half time step: difference calculation (“Second-Order TVD Scheme of Magnetic Field” CUDA kernels in Fig. 3) obtained by Equation (16), the limiter of Equation (14) and Equation (21)
9: calculate the result of \(B^{t+\Delta t}_x\) and \(B^{t+\Delta t}_z\) with \(\text{flux}_{\text{temp}}\) by applying Equation (17)) to \(B_y\), and save it back to \(B\)
10: (the following steps is similar to above steps but the affected orthogonal magnetic field is changed from \(y\) to \(z\))
11: \(\text{vx}_{\text{temp}} \leftarrow\) determine the fluid speed with the updated \(u_1\) and \(u_2\) in \(\text{fluid}_x\), with the difference calculated in \(z\)-dimension
12: \(\text{vx}_{\text{face}} \leftarrow\) Results obtained with Equation (20) using index of \(i, j, k+1/2\)
13: \(\text{flux}_{\text{temp}} \leftarrow\) the flux of a half time step: difference calculation of “flux of magnetic field in \(z\)-dimension” (“First-Order Upwind Scheme of Magnetic Field” CUDA kernels in Fig. 3) obtained by Equations (16) and (21)
14: \(b_{\text{temp}} \leftarrow\) calculate the intermediate result \((u^{t+\Delta t/2})\) by applying Equation (16) to \(B_z\) (not by applying Equation (16) to \(u\)) with \(B_z\) and \(\text{flux}_{\text{temp}}\)
15: \(\text{flux}_{\text{temp}} \leftarrow\) the flux of another half time step: difference calculation (“Second-Order TVD Scheme of Magnetic Field” CUDA kernels in Fig. 3) obtained by Equation (16), the limiter of Equation (14) and Equation (21)
16: calculate the results of \(B^{t+\Delta t}_x\) and \(B^{t+\Delta t}_z\) with \(\text{flux}_{\text{temp}}\) by applying Equation (17)) to \(B_z\), and save it back to \(B\)
17: free the storage of the intermediate results

calculation steps which need the information of the neighboring grid points obtained in the previous calculation steps. Otherwise, in order to avoid the asynchronous problem in parallel computation, we have to do many redundant processes. With the purpose to minimizing the memory usage, not only the calculation process of \(L_x\) is divided into several steps (CUDA kernels), but also the intermediate results are stored as less as possible. The processes dealing with the difference calculations are also divided into several steps to minimize the storage of the intermediate results and to guarantee there is no wrong result caused by asynchronous problem.
It should be realized that most of the processes in the three-dimensional relaxing TVD scheme with the dimensional splitting technique is similar. Pen et al. [28] swapped the data of $x$, $y$, and $z$-dimensions while $GPU-MHD$ used one-dimensional arrays. But the similar swapping technique can be applied in our case with some indexing operations. Instead of transposing or swapping the data, we implemented each calculation part of the flux computation with two sets of CUDA kernels: one set is the CUDA kernels for calculating the relaxing TVD scheme (we call it TVD kernel here) and the other set is the CUDA kernels actually called by $L_i$ operations (we call them $L_i$ kernels here). Indexing operations are contained in all $L_i$ kernels. After the index is calculated, TVD kernels are called and the indexes are passed to the TVD kernels, letting the TVD kernels calculate the flux of corresponding dimension. Therefore, the difference among $L_x$, $L_y$, and $L_z$ is the dimensional index. The flux computation of $GPU-MHD$ is shown in Fig. 4.

![Fig. 4. Flux computation in $GPU-MHD$.](image)

The indexing operation swaps the target that will be updated and the neighboring relationship will also be changed accordingly. For example, the calculation that uses $x + 1$ as the neighboring element in $L_x$ will be changed to $y + 1$ in $L_y$. As transposing the data in a matrix needs more processing time, it is efficient and flexible to extend the code to multidimensional by dividing the indexing operation and flux calculation.

After the whole pipeline of Fig. 2 is completed, the MHD simulation results will be stored in GRAM and these results are readily to be further processed by the GPUs for visualization or read back to the CPU for other usage.

5 Numerical tests

In this section, several numerical tests in one-dimensional (1D), two-dimensional (2D), and three-dimensional (3D) for validation of $GPU-MHD$ are given. The
results shown in this section are computed with single precision mode in GPU-MHD. The difference between single precision and double precision computation results will be discussed in Section 6.

5.1 One-dimensional problems

5.1.1 Brio-Wu shock tube

1D Brio-Wu shock tube problem [3] which is a MHD version of the Sod problem [35], consists of a shock tube with two initial equilibrium states as follows

Left side \((x < 0.5)\)

\[
\begin{align*}
\begin{cases}
v_x \\
v_y \\
v_z
\end{cases} &= \begin{cases}
0 \\
0 \\
0
\end{cases} \\
\begin{cases}B_x \\
B_y \\
B_z
\end{cases} &= \begin{cases}0.75 \\
1 \\
0
\end{cases} \\
\rho &= 1, \quad p = 1
\end{align*}
\]

\((30)\)

Right side \((x \geq 0.5)\)

\[
\begin{align*}
\begin{cases}
v_x \\
v_y \\
v_z
\end{cases} &= \begin{cases}
0 \\
0 \\
0
\end{cases} \\
\begin{cases}B_x \\
B_y \\
B_z
\end{cases} &= \begin{cases}0.75 \\
-1 \\
0
\end{cases} \\
\rho &= 0.125, \quad p = 0.1
\end{align*}
\]

\((33)\)

\((34)\)

\((35)\)

Constant value of \(\gamma = 2\) was used and the problem was solved for \(x \in [0, 1]\) with 512 grid points. Numerical results are presented at \(t = 0.08L\) in Fig. 5 and Fig. 6, which include the density, the pressure, the energy, the \(y\)- and \(z\)-magnetic field components, and the \(x\)-, \(y\)- and \(z\)-velocity components. The results are in agreement with those obtained by Brio and Wu [3] and Zachary et al. [49].
Fig. 5. Results (Part I) of Brio-Wu shock tube problem at $t = 0.08L$. The result computed with 512 grid points is shown with circles and solid line shows reference high resolution result of 4096 grid points.

5.1.2 MHD shock tube

The second 1D test is the MHD shock tube problem considered in [6].

Left side ($x < 0.5$)

\[
\begin{align*}
\begin{bmatrix}
  v_x \\
  v_y \\
  v_z 
\end{bmatrix} &= \begin{bmatrix}
  1.2 \\
  0.01 \\
  0.5 
\end{bmatrix} \\
B_x &= 2\sqrt{4\pi} \\
B_y &= 3.6/\sqrt{4\pi} \\
B_z &= 2/\sqrt{4\pi}
\end{align*}
\]

\[\rho = 1.08, \quad p = 0.95\]
Fig. 6. Results (Part II) of Brio-Wu shock tube problem at $t = 0.08L$. The result computed with 512 grid points is shown with circles and solid line shows reference high resolution result of 4096 grid points.

Right side ($x \geq 0.5$)

$$
\begin{align*}
\begin{bmatrix}
v_x \\
v_y \\
v_z
\end{bmatrix}
&= 
\begin{bmatrix}
0 \\
0 \\
0
\end{bmatrix} \\
\begin{bmatrix}
B_x \\
B_y \\
B_z
\end{bmatrix}
&= 
\begin{bmatrix}
2\sqrt{4\pi} \\
4\sqrt{4\pi} \\
2\sqrt{4\pi}
\end{bmatrix} \\
\rho = 1, & p = 1
\end{align*}
$$

Constant value of $\gamma = 5/3$ was used and the problem was solved for $x \in [0, 1]$ with 512 grid points. Numerical results are presented at $t = 0.2L$ in Fig. 7 and Fig. 8, which include the density, the pressure, the energy, the $y$- and
z-magnetic field components, and the $x$, $y$, and $z$-velocity components. The results are in agreement with those obtained by [6] and [30].

Fig. 7. Results (Part I) of MHD shock tube test at $t = 0.2L$. The result computed with 512 grid points is shown with circles and solid line shows reference high resolution result of 4096 grid points.

5.2 Two-dimensional problems

5.2.1 Orszag-Tang problem

The first 2D test is Orszag-Tang problem [27], which is used to study incompressible MHD turbulence. In our test, the boundary conditions are periodic everywhere. The density $\rho$, pressure $p$, initial velocities ($v_x, v_y, v_z$), and mag
Fig. 8. Results (Part II) of MHD shock tube test at $t = 0.2L$. The result computed with 512 grid points is shown with circles and solid line shows reference high resolution result of 4096 grid points.

Magnetic field ($B_x, B_y, B_z$) are given by

$$\begin{align*}
\begin{cases}
    v_x = -\sin(2\pi y) \\
    v_y = \sin(2\pi x) \\
    v_z = 0
\end{cases}
\end{align*}$$

$$\begin{align*}
\begin{cases}
    B_x = -B_0\sin(2\pi y) \\
    B_y = B_0\sin(4\pi x) \\
    B_z = 0
\end{cases}
\end{align*}$$

where $B_0 = 1/\sqrt{4\pi}$

$$\rho = \frac{25}{36\pi}, \quad p = \frac{5}{12\pi}, \quad \gamma = \frac{5}{3}, \quad (0 \leq x \leq 1) \quad (0 \leq y \leq 1)$$

The Orszag-Tang vertex test was performed in a two-dimensional periodic box with $512 \times 512$ grid points. The results of the density and gas pressure
The evolution of the Orszag-Tang problem at \( t = 0.5L \) and \( t = 1.0L \) are shown in Fig. 9, where the complex pattern of interacting waves is perfectly recovered. The results agree well with those in Lee et al. [19].

Fig. 9. Results of the density (top) and gas pressure (bottom) of Orszag-Tang vortex test at \( t = 0.5L \) (left) and \( t = 1.0L \) (right) computed with 512 × 512 grid points.

5.2.2 Two-dimensional blast wave problem

The second 2D test is the MHD blast wave problem. The MHD spherical blast wave problem of Zachary et al. [49] is initiated by an over pressured region in the center of the domain. The result is a strong outward moving spherical shock with rarified fluid inside the sphere. We followed the test suite [36] of Athena [40]. The condition for 2D MHD blast wave problem is listed as
follows [36]

\[
\begin{align*}
\begin{pmatrix}
v_x \\
v_y \\
v_z
\end{pmatrix} &= \begin{pmatrix}
0 \\
0 \\
0
\end{pmatrix}, \\
\begin{pmatrix}
B_x \\
B_y \\
B_z
\end{pmatrix} &= \begin{pmatrix}
1/\sqrt{2} \\
1/\sqrt{2} \\
0
\end{pmatrix}
\end{align*}
\] (45)

\[p = \begin{cases}
10 & \text{inside the spherical region} \\
0.1 & \text{outside the spherical region}
\end{cases} \] (47)

\[
\rho = 1, \quad p = 5/(12\pi), \quad \gamma = 5/3
\]

spherical region center = (0.5, 0.5), \( r = 0.1 \) \( (0 \leq x \leq 1) \) \( (0 \leq y \leq 1) \) (48)

In Fig. 10, we present images of the density and gas pressure at \( t = 0.2L \) computed with 512 × 512 grid points. The results are in excellent agreement with those presented in [36].

![Fig. 10. Results of the density (left) and gas pressure (right) of the 2D blast wave test at \( t = 0.2L \), computed with 512 × 512 grid points.](image)

5.2.3 MHD rotor problem

The third 2D test is the MHD rotor problem. The problem was taken from [1]. It initiates a high density rotating disk with radius \( r_0 = 0.1 \) of fluid measured from the center point \((x, y) = (0.5, 0.5)\). The ambient fluid outside of the spherical region of \( r_1 = 0.115 \) has low density and \( v_x = v_y = 0 \), and the fluid
between the high density disk fluid and ambient fluid \((r_1 > r > r_0, \text{ where } r = \sqrt{(x - 0.5)^2 + (y - 0.5)^2})\) has linear density and angular speed profile with \(\rho = 1 + 9f, v_x = -f v_0(y - 0.5)/r\) and \(v_y = -f v_0(x - 0.5)/r\) where \(f = (r_1 - r)/(r_1 - r_0)\). Two initial value sets of \(v_0, p, B_x\) and \(\gamma\) provided in [1] and [44] were tested. The initial condition for 2D MHD Rotor problem is listed as follows

spherical region center = \((0.5, 0.5), r_0 = 0.1, r_1 = 0.115\)

\[ f = (r_1 - r)/(r_1 - r_0), (0 \leq x \leq 1) (0 \leq y \leq 1) \]

\( r < r_0 \)

\[
\begin{align*}
\begin{cases}
v_x \\
v_y \\
v_z
\end{cases} &= \begin{cases}
-v_0(y - 0.5)/r_0 \\
v_0(x - 0.5)/r_0 \\
0
\end{cases}
\end{align*}
\]

\[ (50) \]

\( r_0 < r < r_1 \)

\[
\begin{align*}
\begin{cases}
v_x \\
v_y \\
v_z
\end{cases} &= \begin{cases}
-f v_0(y - 0.5)/r \\
-f v_0(x - 0.5)/r \\
0
\end{cases}
\end{align*}
\]

\[ (51) \]

\( r > r_1 \)

\[
\begin{align*}
\begin{cases}
v_x \\
v_y \\
v_z
\end{cases} &= \begin{cases}
0 \\
0 \\
0
\end{cases}
\end{align*}
\]

\[ (52) \]

\[
\rho = \begin{cases}
10 & r < r_0 \\
1 + 9f & r_0 < r < r_1 \\
1 & r > r_1
\end{cases}
\]

\[ (53) \]

First rotor problem:

\[ v_0 = 2, p = 1, \gamma = 1.4, t_{max} = 0.15, \begin{cases}
B_x \\
B_y \\
B_z
\end{cases} = \begin{cases}
5/\sqrt{4\pi} \\
0 \\
0
\end{cases} \]

\[ (54) \]

Second rotor problem:

\[ v_0 = 1, p = 0.5, \gamma = 5/3, t_{max} = 0.295 \begin{cases}
B_x \\
B_y \\
B_z
\end{cases} = \begin{cases}
2.5/\sqrt{4\pi} \\
0 \\
0
\end{cases} \]

\[ (55) \]
In Fig. 11, we present images of the density, gas pressure of the two rotor problems computed with $512 \times 512$ grid points. The results are in excellent agreement with those presented in [1] and [44].

Fig. 11. Results of the density (top-left), gas pressure (top-right) of the first MHD rotor test at $t = 0.15L$, results of the density (bottom-left), gas pressure (bottom-right) of the second MHD rotor test at $t = 0.295L$, both computed with $512 \times 512$ grid points.

5.3 Three-dimensional blast wave problem

The 3D version of MHD spherical blast wave problem was also tested. The condition is listed as follows [36]

$$
\begin{align*}
\begin{cases}
    v_x = 0 \\
    v_y = 0 \\
    v_z = 0
\end{cases}
\end{align*}
$$

(56)
\[
\begin{align*}
\begin{bmatrix} B_x \\ B_y \\ B_z \end{bmatrix} &= \begin{bmatrix} 1/\sqrt{3} \\ 1/\sqrt{3} \\ 1/\sqrt{3} \end{bmatrix} \\
p &= \begin{cases} 10 & \text{inside the spherical region} \\ 0.1 & \text{outside the spherical region} \end{cases}
\end{align*}
\]

\[p = \frac{1}{\sqrt{3}}, \frac{1}{\sqrt{3}}, \frac{1}{\sqrt{3}}\]  

\[(57)\]

\[\rho = 1, \quad \gamma = 5/3\]

spherical region center = (0.5, 0.5, 0.5), \( r = 0.1\)

\[(58)\]

\[(59)\]

Fig. 12 and Fig. 13 show the results of 3D blast wave problem, which include the density, gas pressure, and magnetic pressure at \( t = 0.1L \) and \( t = 0.2L \) sliced along \( x-y \) plane at \( z = 0.5 \). The test was computed with 128 \( \times \) 128 \( \times \) 128 grid points. Due to the scarcity of published 3D test results, we do not make direct contact with results presented in the literature here. Considering only the \( u \) and \( B \), the memory requirement of 256\(^3\) MHD problem is about 512MB GRAM for single precision and 1024MB GRAM for double precision, respectively. If the storage of intermediate results such as \( B_{temp}, u_{temp}, \) \( flux_{temp} \) and \( F \) etc. (See Section 4.3) are considered, the amount of memory requirement will be about 2.25GB (single precision). As we mentioned in Section 4.3, not all the capacity of GRAM can be used to store data arbitrarily. There are actually two GPUs on the GTX 295 and the 1.75 GB GRAM is the total amount of the GRAM shared by two GPUs, so that only less than \( 1.75/2 = 0.875 \) GB GRAM can be used. As a result, the test of 3D problem with 256\(^3\) resolution are not able to be provided on a graphics card.

6 Accuracy evaluation

In MHD simulations, accuracy is always to be considered since the error may increase fast and make the simulation crashed if low precision is used for computation. Scientific computations such as MHD simulation mostly use double precision to reduce errors. In this section, the results generated by \textit{GPU-MHD} using single precision and double precision modes are shown and compared.

The difference between the results of double precision and single precision computation of the 512 \( \times \) 1 \( \times \) 1 one-dimensional Brio-Wu shock tube problem is shown in Fig. 14. Two curves are almost the same but there are actually some differences with the amount of \( error \leq \pm 10^{-6} \).
Fig. 12. Results of the density (top-left), gas pressure (top-right) and magnetic pressure (bottom) of 3D blast wave test at $t = 0.1L$ sliced along $x$-$y$ plane at $z = 0.5$ and computed with $128 \times 128 \times 128$ grid points.

In 2D cases, the absolute difference between the results of double precision and single precision computation of MHD Rotor test ($t = 0.15L$) and Orszag-Tang vortex test ($t = 0.5L$ and $t = 1.0L$) are shown in Fig. 15 and Fig 16, respectively. The double precision computation results of both tests are also shown in the left-hand side of these figures.

For the MHD Rotor test, even the resulting image (left in Fig. 15) looks similar to the single precision resulting image (top-left of Fig. 11), the high differences at the dense region can be found. Experimental result shows that the maximum error is larger than $\pm 3.5 \times 10^{-4}$.

Fig. 16 shows the absolute difference between the results of double precision and single precision computation of Orszag-Tang test at $t = 0.5L$ and $t = 1.0L$. As the simulation time increases, the maximum error increases from about $\pm 8 \times 10^{-5}$ to $\pm 0.03$. 
Fig. 13. Results of the density (top-left), gas pressure (top-right) and magnetic pressure (bottom) of 3D blast wave test at $t = 0.2L$ sliced along $x$-$y$ plane at $z = 0.5$ and computed with $128 \times 128 \times 128$ grid points.

Fig. 17 and Fig. 18 show the resulting images of the simulation using double precision and the contours of the absolute differences between the results of double precision and single precision computation of 3D blast wave test with $128^3$ grid points at $t = 0.1L$ and $t = 0.2L$. As it is a high dimension computation in low resolution, the differences between them are clear. The number of grid points having higher difference value increases, and the amount of difference is still kept as $\text{error} \leq \pm 10^{-6}$. Small difference value makes the double precision resulting images (Fig. 17 and Fig. 18) looked similar to the single precision resulting images (Fig. 12 and Fig. 13).

An important point can be realized that not only the grid points at the high density region has high difference value, but also the number of grid points having high difference value and the amount of the difference values are increasing along with the increase of the simulation time. Higher dimension is another factor to introduce noticeable differences between the computation results with different precisions because higher dimension means a grid point...
Fig. 14. Result of $\rho_{\text{double}} - \rho_{\text{single}}$ (top), $E_{\text{double}} - E_{\text{single}}$ (middle) and $p_{\text{double}} - p_{\text{single}}$ (bottom) of 1D Brio-Wu shock tube problem at $t = 0.08L$ with 512 grid points

has more neighbors and more neighbors need more computation steps in one time step. As a result the differences become more obvious. Therefore, for a long-term simulation, double precision computation is a must.
Performance measurements

The performance measurements of the GPU and CPU implementations as well as the computation using double precision and single precision are carried out in this section. Different numbers of grid points and different dimensions were used in the performance tests. We run both GPU-MHD and Pen et al.’s FORTRAN/CPU MHD code [29] to perform the simulations on a PC with Intel Core i7 965 3.20 GHz CPU, 6G main memory, running Microsoft Windows XP 64-bit Professional. The graphics card is NVIDIA Geforce GTX 295 with 1.75G video memory. GPU-MHD was designed for three-dimensional problems, thus the dimensions are expressed in three-dimensional form in all tables. For 1D test, 1D Brio-Wu shock tube problem (see Section 5.1.1) was used. For 2D test, 2D Orszag-Tang problem (see Section 5.2.1) was used. For 3D test, 3D blast wave problem (see Section 5.3) was used.

Table 1 reports the comparison of GPU-MHD and the FORTRAN/CPU code of 1D test with different numbers of grid points. Basically there is only about 10 times speedup (4096 \times 1 \times 1 case) since the number of grid points is small. And it should be realized that the amount of speedup is increased as long as the resolution is increased but dropped when the resolution reaches 512. It is because the “max threads per block” of the NVIDIA Geforce GTX 295 is 512, all the computations are handled within one block and a very high processing speed can be archived.

Table 2 gives the comparison of GPU-MHD using single precision and double precision of 1D test with different numbers of grid points. A similar speed drop happened both in single and double precision modes, but it occurred in different resolutions: 512 in single precision and 256 in double precision. This is not strange and not difficult to understand since the double precision has
Fig. 16. Results of $\rho_{\text{double}}$ (left) and $|\rho_{\text{double}} - \rho_{\text{single}}|$ (right) of Orszag-Tang problem at $t = 0.5L$ (top) and $t = 1.0L$ (bottom) with $512^2$ grid points.

Table 1
The performance results of 1D test of FORTRAN/CPU and GPU-MHD at different resolutions.

| Number of grid points | FORTRAN/CPU (ms/step) | GPU-MHD (ms/step) | Speedup |
|-----------------------|------------------------|-------------------|---------|
| $128 \times 1 \times 1$ | 10.9                   | 3.6               | 3.0277  |
| $256 \times 1 \times 1$ | 21.9                   | 4.0               | 5.4750  |
| $512 \times 1 \times 1$ | 43.8                   | 4.5               | 9.7333  |
| $1024 \times 1 \times 1$ | 87.5                   | 30.0              | 2.9167  |
| $2048 \times 1 \times 1$ | 173.4                  | 32.2              | 5.385   |
| $4096 \times 1 \times 1$ | 350.0                  | 36.4              | 9.6154  |

double size of data to be handled by the processors. Except the special case of 512 resolution, the processing speed in both modes are very closed.
Fig. 17. Results of $p_{\text{double}}$ (top-left), $\rho_{\text{double}}$ (bottom-left) and $|p_{\text{double}} - p_{\text{single}}|$ (top-right), $|\rho_{\text{double}} - \rho_{\text{single}}|$ (bottom-right) of 3D blast wave problem at $t = 0.1L$ with $128^3$ grid points.

The comparison of GPU-MHD and the FORTRAN/CPU code of 2D test with different numbers of grid points is presented in Table 3. In 2D case, a significant performance improvement is observed, especially when the numbers of grid points are $512^2$ and $1024^2$, a speedup of around 150 and around 200 is achieved, respectively.

Table 4 presents the comparison of GPU-MHD using single precision and double precision of 2D test with different numbers of grid points. The significant performance difference is noticeable when the number of grid points is increased. However, it still keeps a ratio increasing slowly from 1.118 to 1.6218 while the resolution increases from $128^2$ to $1024^2$.

Table 5 shows the comparison of GPU-MHD and the FORTRAN/CPU code of 3D test with different numbers of grid points. The performance of GPU-MHD is faster than the FORTRAN/CPU code about 60 times and 84 times when the numbers of grid points are $64^3$ and $128^3$, respectively.

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Table 5 shows the comparison of GPU-MHD using single precision and double precision of 3D test with different numbers of grid points. The ratio is 1.6020 when the number of grid points is $64^3$, and is 1.7389 when the number of grid points is $128^3$.

The performance tests show that when the number of grid points of the test problems is small, such as those in 1D case, GPU-MHD can give a significant performance improvement. When the number of grid points increases, an obvious disparity of performance becomes clear, especially for multidimensional cases (see Table 3 and Table 5). For example, in Table 3, GPU-MHD even obtained a speedup of around 150 to 200 compared to the FORTRAN/CPU implementation when the number of grid points is $512^2$ or $1024^2$. The performance results show that CUDA is an attractive parallel computing environment for MHD simulations.

Computation using double precision on the current GPUs is known to be
Table 2
The performance results of 1D test between single precision and double precision of GPU-MHD at different resolutions

| Number of grid points | Double precision (ms/step) | Single precision (ms/step) | Ratio |
|-----------------------|-----------------------------|-----------------------------|-------|
| 128 × 1 × 1           | 3.9                         | 3.6                         | 1.0833|
| 256 × 1 × 1           | 4.5                         | 4.0                         | 1.1250|
| 512 × 1 × 1           | 29.5                        | 4.5                         | 6.5555|
| 1024 × 1 × 1          | 31.0                        | 30.0                        | 1.0333|
| 2048 × 1 × 1          | 33.0                        | 32.2                        | 1.0248|
| 4096 × 1 × 1          | 39.1                        | 36.4                        | 1.0742|

Table 3
The performance results of 2D test of FORTRAN/CPU and GPU-MHD at different resolutions

| Number of grid points | FORTRAN/CPU (ms/step) | GPU-MHD (ms/step) | Speedup |
|-----------------------|------------------------|-------------------|---------|
| 128 × 128 × 1         | 901.6                  | 32.2              | 28.0    |
| 256 × 256 × 1         | 3639.1                 | 44.8              | 82.9236|
| 512 × 512 × 1         | 14715.6                | 94.0              | 156.5489|
| 1024 × 1024 × 1       | 58276.6                | 295.1             | 197.4809|

Table 4
The performance results of 2D test between single precision and double precision of GPU-MHD at different resolutions

| Number of grid points | Double precision (ms/step) | Single precision (ms/step) | Ratio  |
|-----------------------|-----------------------------|-----------------------------|--------|
| 128 × 128 × 1         | 35.8                        | 32.2                        | 1.1118 |
| 256 × 256 × 1         | 57.3                        | 44.8                        | 1.2790 |
| 512 × 512 × 1         | 142.3                       | 94.0                        | 1.5138 |
| 1024 × 1024 × 1       | 478.6                       | 295.1                       | 1.6218 |

very low performance compared to single precision. However, in the performance comparison between single precision and double precision modes in GPU-MHD, the ratios of the processing speed between two modes show that GPU-MHD is efficient enough in double precision computation. Even when the number of grid points is $1024^2$ or $128^3$, computation using double precision still reaches 62.2% or 57.5% of the corresponding single precision processing speed. It is due to that GPU-MHD uses indexing operations instead of trans-
Table 5
The performance results of 3D test of FORTRAN/CPU and GPU-MHD at different resolutions

| Number of Grid points | FORTRAN/CPU (ms/step) | CUDA/GPU (ms/step) | Speedup |
|-----------------------|------------------------|---------------------|---------|
| 32 × 32 × 32          | 739.1                  | 36.6                | 20.1940 |
| 64 × 64 × 64          | 5521.9                 | 90.7                | 60.8809 |
| 128 × 128 × 128       | 42331.8                | 506.4               | 83.5936 |

Table 6
The performance results of 3D test between single precision and double precision of GPU-MHD at different resolutions

| Number of grid points | Double precision (ms/step) | Single precision (ms/step) | Ratio   |
|-----------------------|----------------------------|---------------------------|---------|
| 32 × 32 × 32          | 44.6                       | 36.6                      | 1.2186  |
| 64 × 64 × 64          | 145.3                      | 90.7                      | 1.6020  |
| 128 × 128 × 128       | 880.6                      | 506.4                     | 1.7389  |

posing the matrix. These operation also helps keeping the efficiency, while read-write operations of transposing the matrix with double precision data are much slower than those with single precision data.

8 Visualization of the simulation results

There is a need to visualize the MHD simulation data, for examples, Daum [7] developed a toolbox called VisAn MHD in MATLAB for MHD simulation data visualization and analysis. With the help of GPUs, Stantchev et al. [38] used GPUs for computation and visualization of plasma turbulence. In GPU-MHD, using the parallel computation power of GPUs and CUDA, the simulation results of one time step can be computed in dozens or hundreds milliseconds. According to the efficiency of GPU-MHD, near real-time visualization is able to be provided for 1D and 2D problems. The motion or attributes of the magnetic fluid can be computed and rendered on the fly. So the changes of the magnetic fluid during the simulation can be observed in real-time.

By adding the real-time visualization, the flow of GPU-MHD, Fig. 2 is extended as Fig. 19:

GPU-MHD provides different visualization methods for one-dimensional, two-dimensional and three-dimensional problems.
Fig. 19. The flow chat of GPU-MHD.

To visualize one-dimensional problems for each time step, the simulation results are copied to the CUDA global memory that mapped to the Vertex Buffer Object (VBO) [48]. For all grid points, one grid point is mapped to one vertex. The position of each grid point is mapped as the $x$-position of the vertex and the selected physical value ($\rho$, $p$, etc.) is mapped as the $y$-position of the vertex. Then a curve of these vertices is drawn. Since the VBO is mapped to CUDA global memory and simulation results are stored in GRAM, the copying and mapping operations are fast. Experimental result shows that GPU-MHD with real-time visualization can achieve 60 frame per second (FPS) in single precision mode and 30 FPS in double precision mode. Fig. 20 shows two example images of 1D visualizations using GPU-MHD.

The operational flow of visualization of 2D problems is similar to that in 1D
Fig. 20. 1D real-time visualizations of the density ($\rho$) of Brio-Wu shock tube problem with 512 grid points using $GPU-MHD$.

visualization. However, instead of Vertex Buffer Object (VBO), Pixel Buffer Object (PBO) [48] is used. For each time step, the simulation results are copied to the CUDA global memory that are then mapped to the PBO. For all grid points, one grid point is mapped to one pixel. The $x$ and $y$ position of each grid point are mapped as the corresponding $x$-position and the $y$-position of the vertex and the selected physical value ($\rho$, $p$, etc.) is mapped as the color of the pixel to form a color image. To render this color image, a transfer function is set to map the physical value to the color of the pixel and then the resulting image is drawn. Similar to VBO, PBO is also mapped to CUDA global memory and the simulation results are stored in GRAM, so the copying and mapping operations are also fast and do not affect too much to the performance. Although the number of grid points in 2D problem is much larger than those in the one-dimension problem, the FPS still reaches 10 in single precision mode and 6 in double precision mode, still giving acceptable performance to the user. Fig. 21 shows two example images of 2D visualizations using $GPU-MHD$.

However, visualization of 3D problem is different to 1D and 2D problems. GPU-based volume visualization method [12] and texture memory (or video memory) are used. Unfortunately, the current version (Version 2.3) of CUDA does not provide the feature to copy the data from the CUDA global memory to texture memory directly, even both of them are in GRAM. On the other hand, texture memory is readable but is not rewritable in CUDA. So the simulation results have to be copied to the main memory first, and then be copied to texture memory. In addition, the number of grid points is usually large compared to 2D problems and volume visualization techniques are somewhat time-consuming. As a result, $GPU-MHD$ only gets 2 FPS in single precision mode and 1 FPS in double precision mode, and it is far from real-time. Never-
Fig. 21. 2D visualizations of the density ($\rho$) of Orszag-Tang vortex problem with $512^2$ grid points using GPU-MHD. Nevertheless, we still get 10 FPS (single precision mode) and 6 FPS (double precision mode) for performing the simulation of problems with resolution of $64^3$ and about 20 FPS (single and double precision modes) for problems with resolution of $32^3$. Fig. 22 shows two example images of 2D visualizations using GPU-MHD.

Fig. 22. 3D visualizations of the density ($E$) of 3D Blast wave problem with $128^3$ grid points using GPU-MHD.

9 Conclusion and future work

In this paper we present, to the author’s knowledge, the first implementation of MHD simulations entirely on GPUs with CUDA, named GPU-MHD, to accelerate the simulation process. A series of numerical tests have been performed to
validate the correctness of our code. Accuracy evaluation by comparing single and double precision computation results is also given, indicating that double precision support on GPUs is a must for long-term MHD simulation. Performance measurements of both single and double precision modes of GPU-MHD are conducted. These measurements show that GPU-MHD achieves speedups (in single precision mode) of about 10 (1D problems with 4096 grid points), 200 (2D problems with $1024^2$ grid points), and 84 (3D problems with $128^3$ grid points), respectively, compared to the corresponding serial CPU MHD implementation. For double precision computation, GPU-MHD still can achieve about 60% speed of the corresponding single precision computation. In order to provide the user better understanding of the problems being investigated during the simulation process, we have extended GPU-MHD to support visualization of the simulation results. With GPU-MHD, the whole MHD simulation and visualization process can be performed entirely on GPUs. As NVIDIA has announced the next generation CUDA GPUs — Fermi [26], which provides 8x speedups in double precision computation compared to the current generation CUDA GPUs. GPU-MHD will benefit this advance and will give better performance on Fermi series GPUs, giving GPU-MHD useful for wide range applications in astrophysics or space physics.

There are two directions in our future work, firstly, we are going to extend GPU-MHD for multiple GPUs and GPU cluster [33] to fully exploit the power of GPUs. Secondly, we will investigate implementing other recent high-order Godunov MHD algorithms such as [19] and [41] on GPUs. These GPU-based algorithms will be served as the base of our GPU framework for simulating large-scale MHD problems in space weather modeling.

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