Hydrodynamics of granular particles on a line

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We investigate a lattice model representing a granular gas in a thin channel. We deduce the hydrodynamic description for the model from the microscopic dynamics in the large system limit, including the lowest finite-size corrections. The main prediction from hydrodynamics, when finite-size corrections are neglected, is the existence of a steady “uniform longitudinal flow” (ULF), with the granular temperature and the velocity gradient both uniform and directly related. Extensive numerical simulations of the system show that such a state can be observed in the bulk of a finite-size system by attaching two thermostats with the same temperature at its boundaries. The relation between the ULF state and the shocks appearing in the late stage of a cooling gas of inelastic hard rods is discussed.

I. INTRODUCTION

Statistical physics offers a systematic computational scheme for averages and fluctuations, suitable for systems of particles at thermal equilibrium [1]. A similar scheme for systems far from equilibrium is lacking and represents an open challenge [2]. An example of non-equilibrium statistical system pervading many human activities is given by granular fluids, an assembly of inelastic hard particles agitated by some external driving mechanism [3–5].

The fluidised state of granular matter is an excellent testing ground for kinetic theory [6–8]. One of the most widely used model in granular kinetic theory is a system composed of inelastic smooth hard particle particles [9–11]. In this context, the inelastic Boltzmann (or Enskog) equation has been shown to be a powerful tool. However, the structure of the collision term for inelastic hard particles makes the corresponding kinetic equation a tough mathematical problem, for which exact solutions and rigorous results are difficult to obtain.

In view of the above issue, models that simplify the collision term have been proposed to make it easier their analytical investigation, while preserving similar physics. This is the spirit of the Maxwell-like collision model, either elastic [12] or inelastic [13], in which the collision rate is assumed to be independent of the relative velocity. Some experts claim that “What harmonic oscillators are for quantum mechanics, and dumb-bells for polymer physics, is what elastic and inelastic Maxwell models are for kinetic theory” [14]. Indeed, many rigorous results in the inelastic case have been derived, both in the freely cooling [13,15,22] and the uniformly heated [13,23,24] cases. As pointed out in a review article by Villani [25], the most relevant questions in inelastic Maxwell models have been solved in homogeneous situations and, rather than going for refinements, the current priority is to deal with inhomogeneous states.

The nonconservation of energy has several, physically relevant, implications. First, the distribution function is in general non-Gaussian [26,27], even for the stationary states reached when some external mechanism injects energy into the system. Second, and most importantly, kinetic theory establishes a link between the microscopic and the macroscopic, hydrodynamic, descriptions, by making it possible to derive the latter from the former [9,10,11]. Along this route, the nonconservation of energy introduces another time scale that makes more delicate the critical requirement on separation of scales needed to accomplish it [28,30]. However, there are many situations, typically in dimension larger than one and for dilute quasi-elastic systems, in which hydrodynamic equations fairly reproduce real granular experiments qualitatively [31] or even quantitatively [32,33].

Here we aim at building and verifying the hydrodynamic equations for an idealised one-dimensional (1d) lattice model [15,34] with a Maxwell-like collision rule. The model simplifies the dynamics of a granular fluid in a 1d channel with negligible density fluctuations, a condition that reasonably holds in not too dilute systems. A strong connection between this idealised model and an inelastic 1d gas of hard rods [35] has been established in previous studies [15]. Specifically, there appears a non-trivial correspondence between the velocity profiles of the lattice model and the shock-like structures of the granular gas of hard rods in the asymptotic cooling regime. The 1d lattice model has been further studied in [36,37] and also generalised to two dimensions [38].

In spite of the above described strong connection between 1d granular gases and the 1d Maxwell model, the hydrodynamic equations of the latter have been neither written or analysed. Then, in the present study, first we derive the hydrodynamic equations from the microscopic dynamics. Afterwards, we focus on a solution of them:
the steady uniform longitudinal flow (ULF). We argue that such a state resembles the shock profiles observed in the late cooling regime of many 1d granular systems. Also, we incorporate finite-size corrections and compare them with simulations of the microscopic dynamics.

Section II is devoted to a brief overview of previous studies on granular models in 1d and their comparison with hydrodynamic theories. Our model is introduced in Section III in the same Section its hydrodynamic equations are derived and the ULF solution is discussed. Section IV presents the numerical results and their comparison with theory. Finite size corrections and boundary layers are discussed in Section V. Conclusions are drawn in Section VI. Appendices illustrate the technical aspects of the derivation of the hydrodynamics (Appendix A) and of the boundary layer calculations (Appendix B), and give also some numerical results for the local velocity distributions and spatial correlations (Appendix C).

II. 1D GRANULAR HYDRODYNAMICS: A BRIEF OVERVIEW

Despite the progress made during the last decades in the realm of granular kinetic theory and hydrodynamics, the validity of a hydrodynamic description in 1d systems is still under debate. It should be stressed that 1d systems essentially differ from higher dimensional setups in which the gradients have a well-defined direction: particles cannot hop over their nearest neighbours, which hinders the necessary “mixing” to obtain a continuum description.

We briefly review some of the main results on 1d granular hydrodynamics below. The 1d version of the granular Navier-Stokes equations reads

\[
D_t \rho = -\rho \frac{\partial}{\partial x} u, \tag{1a}
\]

\[
\rho D_t u = -\frac{\partial}{\partial x} P, \tag{1b}
\]

\[
\rho D_t T = -2P\frac{\partial}{\partial x} u - 2\frac{\partial}{\partial x} q - \rho \frac{\partial}{\partial x} T, \tag{1c}
\]

where the density \( \rho \), the mean velocity \( u \), the granular temperature \( T \), the pressure \( P \), the heat flux \( q \) and the cooling rate \( \zeta \) depend on \((x,t)\). See, for instance, [39] for their microscopic definitions. As usual, we have used the notation \( D_t \equiv \frac{\partial}{\partial t} + u \frac{\partial}{\partial x} \) for the material derivative.

One of the first tests of 1d granular hydrodynamics was carried out in [10]. Therein, a gas of inelastic hard rods was confined by one or two thermostatted walls. Numerical simulations revealed an incompatibility between Eqs. (1) and the observed steady hydrodynamic profiles. Later, the same system has been demonstrated to lack a pure thermodynamic limit, i.e. infinite size at constant restitution coefficient [41, 42].

Subsequent studies of the 1d granular gas focused on the cooling state, that is, with periodic boundary conditions and no external driving. The 1d gas of hard rods develops strong inhomogeneities, which include clusters and shocks, for long enough times [55]. Extensive numerical simulations [43] have revealed that this 1d gas becomes then indistinguishable from a perfectly “sticky” gas [55–44, 45], regardless of its actual inelasticity, a conjecture analysed also in higher dimensions [46, 47] and for wet particles [48]. Interestingly, the (inviscid) Burgers equation \( D_t u = \rho \frac{\partial}{\partial x} u \) [49, 50], with \( \mu \rightarrow 0 \), describes the sticky gas [51–53]. Notwithstanding, to the best of our knowledge, an analytical justification of the Burgers equation in 1d granular gases has not been obtained yet.

A first step in this direction would be a neat derivation of 1d granular hydrodynamics.

1d continuum equations can appear also in higher dimensions, when only one direction develops a gradient, as in [54] for a steady case under gravity or in cooling systems with high aspect ratio without gravity [55–57]. In the latter case, granular hydrodynamics predicts that the so-called flow by inertia \( D_t u = 0 \) sets in during the highly inhomogeneous stage of cooling. Therefore, there also appear shocks in \( u(x,t) \), which imply a singularity in \( \rho(x,t) \). In addition, Molecular Dynamics simulations of this channel system have made clear that such singularities are approached following the hydrodynamic predictions, that is, the flow by inertia scenario [54]. Eventually, when close-packing is almost reached, the flow by inertia picture breaks down [57, 58].

III. THE MODEL AND ITS HYDRODYNAMICS

Here, we study the 1d lattice model first introduced in Refs. [15, 34] to mimic the evolution of a 1d granular gas, but neglecting density fluctuations. The model is a 1d space-time discretised cellular automaton, where we have one unit-mass particle at each site \( l, l = 1, \ldots, N \). Particles are characterised by their (scalar) “velocity” \( v_l \), which evolves according to a collisional dynamics. Note that particles do not move but remain on the same site; in this way density fluctuations are neglected.

At each discrete time step, a random pair of adjacent sites is chosen and their corresponding velocities, say \( v \) and \( \tilde{v} \), are updated according to

\[
v' = v + \frac{1}{2}(\tilde{v} - v), \quad \tilde{v}' = \tilde{v} - \frac{1}{2}(\tilde{v} - v), \tag{2}
\]

if \( v \) and \( \tilde{v} \) satisfy the so-called kinematic constraint, that

\[
\begin{align*}
\text{Allowed (green, left) and forbidden (red, right) collision, according to the kinematic constraint. Arrows show (precollosional) velocities of a pair of (candidate) colliding particles. Green pairs actually collides, red pairs do not collide.}
\end{align*}
\]
A. Microscopic evolution of velocity and energy

At discrete time $p$, the pair $\{y_p, y_p + 1\}$ that undergo a collision is chosen at random and the velocity on site $l$ at time $p + 1$ is given by

$$v_{l,p+1} - v_{l,p} = -j_{l,p} + j_{l-1,p}, \tag{3}$$

where we have defined the microscopic momentum flux from site $l$ to site $l+1$ at time $p$

$$j_{l,p} = \delta_{y_p} \Theta(v_{l,p} - v_{l+1,p}) \frac{1 + \alpha}{2} (v_{l,p} - v_{l+1,p}). \tag{4}$$

Therein, $\delta_{ij}$ is Kronecker’s delta, which identifies the colliding pair, and $\Theta(x)$ is Heaviside’s step function, which imposes the kinematic constraint. The random integer $y_p$ is uniformly distributed in $[1, L]$, where $L$ is basically equal to $N$ but depends on the boundary conditions. For a thermostat system $L = N + 1$. Obviously, $j_{l,p}$ only differs from zero when the pair $(l, l + 1)$ actually collides. Note that no external volume forces (like gravity, for instance) are being considered, but they could be incorporated by adding a term $f_{l,p}$ to the rhs of Eq. (3).

The evolution of the kinetic energy is obtained by squaring Eq. (3), which after some algebra yields

$$v_{l,p+1}^2 - v_{l,p}^2 = \frac{\alpha^2 - 1}{4} \left[ \delta_{y_p} \Theta(v_{l,p} - v_{l+1,p}) (v_{l,p} - v_{l+1,p})^2 + \delta_{y_p} \Theta(v_{l-1,p} - v_{l,p}) (v_{l-1,p} - v_{l,p})^2 \right] - j_{l,p} + j_{l-1,p}. \tag{5}$$

The microscopic energy current is $j_{l,p} = (v_{l,p} + v_{l+1,p}) j_{l,p}$. The sink terms in squared brackets on the rhs of Eq. (5) stem from the inelasticity of collisions and lead to a monotonic decrease of the total energy. The last two terms are a discrete spatial derivative and correspond to the energy flux between neighbouring sites, which is already present in the conservative case $\alpha = 1$.

B. Equations for the velocity and temperature fields in the continuum limit

The average (over realisations) velocity field $u_{l,p} = \langle v_{l,p} \rangle$ evolves according to

$$u_{l,p+1} - u_{l,p} = -\left(j_{l,p} - j_{l-1,p}\right). \tag{6}$$

To compute the last average, we introduce a local equilibrium assumption

$$P_2(v_l, v_{l+1}) \simeq \frac{1}{2\pi \sqrt{T_{l,p} T_{l+1,p}}} e^{-\frac{(v_{l+1,p} - v_{l,p})^2}{2 T_{l,p}} - \frac{(v_{l+1,p} - v_{l+1,p})^2}{2 T_{l+1,p}}} . \tag{7}$$

We also assume that $u_{l,p}$ and $T_{l,p}$ are smooth functions of $l$ in the large system size limit $L \gg 1$, more specifically that the discrete derivatives $u_{l+1,p} - u_{l,p} = O(L^{-1})$ and $T_{l+1,p} - T_{l,p} = O(L^{-1})$, which gives

$$\langle \Theta(v_{l,p} - v_{l+1,p}) (v_{l,p} - v_{l+1,p}) \rangle \sim \sqrt{T_{l,p}}. \tag{8}$$

This is a “modified” pressure, the usual ideal gas equation of state $\rho T$ is replaced by $\sqrt{T}$ because (i) collisions do not occur at a rate proportional to $\sqrt{T}$ but constant and (ii) density $\rho$ is uniform, $\rho = 1$. Therefore,

$$u_{l,p+1} - u_{l,p} \sim -\frac{1 + \alpha}{2L} \left(\sqrt{T_{l,p}} - \sqrt{T_{l-1,p}}\right). \tag{9}$$

Now we introduce a continuum limit by defining spatial and temporal variables as

$$x = \epsilon l, \quad t = \epsilon^2 p, \quad \epsilon = L^{-1} \ll 1, \tag{10}$$

to obtain

$$\partial_t u(x, t) = -\frac{1 + \alpha}{2} \partial_x \sqrt{T(x, t)} + O(\epsilon). \tag{11}$$

We now discuss the average of Eq. (5) for the kinetic energy, which contains both dissipative and transport
terms. Within the local equilibrium approximation, the
average of the dissipative or cooling term is
\[
\frac{\alpha^2 - 1}{4L} (T_{l,p} + T_{l-1,p}) = \frac{(\alpha^2 - 1)\epsilon}{2} T_{l,p} + O \left( (\alpha^2 - 1)\epsilon^2 \right),
\]
which in the continuum limit reads
\[
\frac{(\alpha^2 - 1)\epsilon}{2} T(x, t) + O \left( (\alpha^2 - 1)\epsilon^2 \right).
\]
This linear cooling, instead of the typical $T^{3/2}$ behaviour,
also stems from the constant collision rate.

The average of the energy current, again under local
equilibrium, reads $\langle J_{l,p} \rangle = (1 + \alpha)\epsilon u_{l,p}\sqrt{T_{l,p}/\pi} + O(\epsilon^2)$, so that the average of the last two terms in Eq. (5) goes
in the continuum limit to
\[
-(1 + \alpha)\epsilon^2 \frac{\partial_x}{\epsilon^2} u(x, t)\sqrt{T(x, t)/\pi} + O(\epsilon^3).
\]

Comparing Eqs. (13) and (14), we see that they are of
the same order only when $\alpha^2 - 1 = O(\epsilon)$: this choice
makes them of order $\epsilon^2$, consistently with the scaling of
the finite-time difference on the lhs of Eq. (3), $\Delta t = \epsilon^2$.
Therefore, we introduce a “macroscopic inelasticity”
\[
\nu = L(1 - \alpha^2)/2 \geq 0.
\]

Taking into account that $\langle v^2 \rangle = u^2 + T$, we get
\[
\partial_t T(x, t) = -\nu T(x, t) - 2\sqrt{T(x, t)/\pi} \partial_x u(x, t) + O(\epsilon).
\]
where we have also made use of $1 + \alpha \sim 2$ for $L \gg 1$ with
constant $\nu$. Moreover, Eq. (11) becomes
\[
\partial_t u(x, t) = -\partial_x \sqrt{T(x, t)/\pi} + O(\epsilon).
\]

Equations (17) and (18) are the hydrodynamic equations
of the model to the lowest order. They coincide with
the hydrodynamic Eqs. (1) after identifying: $D \rightarrow \partial_t$ (Lagrangian coordinate), $\rho \rightarrow 1$, $P \rightarrow \sqrt{T/\pi}$ (Eq. (5)), $\zeta \rightarrow \nu$ and $\partial_x \nu = 0$. The absence of dissipative transport (viscous stress and heat flow) makes these equations close relatives of the so-called Ideal Granular Hydrodynamics [57, 58, 74].

C. Steady uniform longitudinal flow

In the large system size limit, the stationary solution of
Eqs. (17) and (18) can be obtained. It suffices to ask
that the velocity flow in the middle of the system is 0
[75] to get the following steady ULF profiles
\[
T_s(x) = T_0, \quad u_s(x) = \frac{\nu}{2} \sqrt{\pi T_0} \left( \frac{1}{2} - x \right),
\]
in which the constant $T_0$ remains undetermined at this
level of description [76]. The steady ULF in Eq. (18) does
not require the presence of thermostats at the bound-
aries, it is self-sustained. Nevertheless, as demonstrated
numerically below, one can recover the above steady
ULF profile at the system bulk in a finite-size system
by attaching two identical thermostats at the boundaries.
Also, note that this ULF, at difference with the one
discussed in Ref. [39], is incompressible because no mass
flow is allowed in our lattice system by definition.

Quite strikingly, regions with a linear velocity profile
with negative slope are also observed in the Lagrangian
coordinate during the formation of shocks, in numerical
simulations of the inelastic 1d gas of hard rods [15, 35],
the inelastic hard disk gas in a channel [59] and the lattice
model [15], all in the cooling regime at long times. This
suggests an interesting connection between the ULF and
the shocks characterising granular cooling in 1d (or quasi-
1d) systems.

IV. NUMERICAL RESULTS

Numerical simulations of the model have been carried
out and compared with the previous theory. A system
with $N$ particles is initialised with uncorrelated, normally
distributed random velocities. We introduce thermostats
at its boundaries, at sites 0 and $N + 1$, to make the system
reach a steady state. Then, at each time step $p$, the uniformly
distributed random integer $y_p$ choosing the candidate
colliding pair $\{y_p, y_p+1\}$ is drawn between 0 and $N$,
that is, $L = N + 1$. The velocities for the peripheral
sites of indexes 0, $L$ are randomly and inde-
pendently drawn from a normal distribution with zero
average and unit variance [77]. Due to the kinematic
constraint, the collision described by Eq. (2) only takes
place if $v_{y_p} - v_{y_p+1} > 0$. Otherwise, velocities remain
unchanged. A large number of independent long runs
has been performed to get average stationary profiles of
velocity and energy. Typically, our data correspond to
$10^5$ runs of more than $10^5 L^2$ steps each, starting from
independent normal-distributed velocities.

Our numerical results for the average temperature
profiles are shown in Fig. 2 for different values of the macro-
scopic inelasticity $\nu$. From them, we measure the tempera-
ture of the largest system at mid position, denoting it
by $T_{l,1/2} \equiv T(x = 1/2)$. Later, we compare $T_{l,1/2}$ with the
bulk temperature $T_0$ of Eq. (18), which predicts a con-
stant temperature profile, in quite good agreement with
our observed numerical profiles.

Numerical results for the average velocity profiles are
reported in Fig. 3. Again, the comparison with the theory
is satisfactory: in the bulk, the profiles are linear with
a slope almost perfectly matching the one given by the
value of $T_{1/2}$ in Fig. 2. Both for the temperature and the
average velocity, deviations from the theoretical ULF
profiles are apparent near the boundaries, especially for $\nu = 2$. See below for a closer look at this issue.
Figure 2. Average profile for the steady temperature, for different values of the macroscopic inelasticity $\nu$. The solid (red) lines stand for the estimated value of the temperature $T_{1/2}$ in the system bulk.

Figure 3. Average profile for the stationary velocity, for the same values of $\nu$ considered in Fig. 2. The solid (red) lines stand for the theoretical expression for $u_s(x)$ in Eq. (18), with $T_0$ given by the numerical estimate $T_{1/2}$ for the bulk temperature in Fig. 2.

Figure 4. Average temperature profiles $T(x)$ in the stationary state for different values of $\nu$ (blow up of data in Fig. 2). Red lines stand for the numerical estimate $T_{1/2}$ for the bulk temperature, see Fig. 2 whereas black lines give our analytical prediction $T_0$, as given by Eq. (20).

V. FINITE SIZE CORRECTIONS AND BOUNDARY LAYERS

In Appendix A, we show how to incorporate finite size corrections. Still, we do so assuming local equilibrium, in order to estimate in the simplest way the effect of these higher order terms. Specifically, these $O(\varepsilon)$ corrections introduce second-order spatial derivatives in the evolution equations, and these viscous terms make it possible to accommodate all the boundary conditions. The considered thermostats impose that

$$u(0,t) = u(1,t) = 0, \quad T(0,t) = T(1,t) = 1. \quad (19)$$

The smallness of the viscous terms bring to bear two boundary layers close to the thermostats at $x = 0, 1$, see Appendix B for details. Therein, we determine the unknown bulk temperature $T_0$ by asymptotic matching [78], which yields

$$\varphi(T_0) = \sqrt{\frac{\pi}{4}} \nu, \quad \text{with} \quad \varphi(T_0) \equiv \sqrt{\frac{1}{3} + \frac{2}{3} T_0^{-3/2} - T_0^{-1}}. \quad (20)$$

The function $\varphi(T_0)$ decreases monotonically from infinity to zero as $T_0$ is varied from zero to unity. Therefore, Eq. (20) tells us that $T_0$ is a monotonically decreasing function of the macroscopic inelasticity $\nu$, with

$$\lim_{\nu \to 0} T_0 = 1, \quad \lim_{\nu \to \infty} T_0 = 0. \quad (21)$$

These two limit results are expected on a physical basis: in the elastic limit $\nu \to 0$, the system should be in equilibrium at the temperature of the heat baths whereas in the strongly dissipative limit $\nu \to \infty$ all the energy is dissipated in the boundary layers before reaching the bulk, which is then at zero temperature [79].

A more precise inspection of the temperature profiles is shown in Fig. 4 aimed to emphasise the deviation from the uniform bulk profile in Eq. (18). Interestingly, the estimate for the bulk temperature $T_0$ stemming from our finite size analysis, as given by Eq. (20), is quite close to our numerical estimation $T_{1/2}$. Most importantly, the success of our theory is good for very different values of $\nu$, both small and large. The absolute error $|T_0 - T_{1/2}|$ increases very slowly with $\nu$, ranging from 0.01 for $\nu = 0.2$ to 0.02 for $\nu = 20$. The relative error is also reasonable, remaining under ten per cent for $\nu \lesssim 2$.

In Appendix C, we look into possible sources of discrepancy between simulations and theory. We focus on violations of local equilibrium, Eq. (7): both non-Gaussian
VI. CONCLUDING REMARKS

We have analysed a 1d lattice model, which by construction do not contain density fluctuations. Our derivation of hydrodynamics can be thought as if carried out in an off-lattice system but using a Lagrangian coordinate at the microscopic level. The resulting Eqs. (17) and (16), in which finite-size corrections have been neglected, tell us that a self-sustained ULF steady state appears in the bulk. Therein, the granular temperature is uniform and the average velocity has a linear profile, whose slope is directly related to the temperature. In order to observe such a steady state in a finite-size system, thermostats at the boundaries must be introduced, as shown in our numerical simulations which quite compare with our theory. The dependence of the bulk temperature on the temperature of the thermostats is only obtained after incorporating finite-size corrections to the theory.

Our work demonstrates that 1d granular hydrodynamic equations, whose validity is still under debate, can be derived by means of a proper continuum limit that keeps a finite macroscopic inelasticity, but in which the microscopic dynamics is quasielastic. Our main assumption is the local equilibrium approximation, which has been shown to be valid in other simple models for a wide range of the system parameters \[61, 62, 80\]. How to improve upon the current results by going beyond local equilibrium is an open perspective for future work.

An emerging interesting conjecture is whether the system in the cooling regime is still well described by our hydrodynamic equations. In fact, the ULF discussed here seems to be similar to the linear velocity profiles in shock regions of 1d cooling granular gases, specifically for either hard rods on a line \[35\] or hard disks in a channel geometry \[53\]. A verification of this conjecture, together with a route to connect our theory with the Burgers or flow-by-inertia equations, is currently under investigation.

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Appendix A: Hydrodynamic equations with finite size corrections

Here we derive the hydrodynamic equations, incorporating into them finite size, \[O(L^{-1})\], corrections. First, we analyse the evolution equation for the velocity: our starting point is Eq. (6), in which the average momentum current at site \(l\) is exactly given as

\[
\langle j_{l,p} \rangle = \frac{(1 + \alpha)\epsilon}{2} \langle \Theta(v_{l,p} - v_{l+1,p})(v_{l+1,p} - v_{l+2,p}) \rangle, \quad \epsilon = L^{-1},
\]

(A1)

since \(y_p\) is an independent, uniformly distributed, stochastic integer choosing the specific pair that collides. We recall that \(L\) is the total number of possible colliding pairs, whose relation to the number of sites \(N\) depends on the boundary conditions. The average on the rhs of the above equation is done over all the velocities \(\{v_{l,p}, v_{l+1,p}\}\), by assuming the local equilibrium approximation written in Eq. (7) of the main text. The result is

\[
\langle \Theta(v_{l,p} - v_{l+1,p})(v_{l+1,p} - v_{l+2,p}) \rangle = \frac{u_{l,p} - u_{l+1,p}}{2} \left[ \text{erf} \left( \frac{u_{l,p} - u_{l+1,p}}{\sqrt{2(T_{l+1,p} + T_{l,p})}} \right) + 1 \right] + \exp \left( \frac{(u_{l,p} - u_{l+1,p})^2}{2(T_{l+1,p} + T_{l,p})} \right) \frac{T_{l+1,p} + T_{l,p}}{2\pi},
\]

(A2)

where erf\(z\) is the error function defined by

\[
\text{erf}(z) = \frac{2}{\sqrt{\pi}} \int_0^z dt e^{-t^2}.
\]

(A3)

Now we go to a continuum limit in space by assuming that both \(u_{l,p}\) and \(T_{l,p}\) vary smoothly with the site index \(l\). Then, we define a continuous spatial variable \(x = \epsilon l\), with \(\epsilon = L^{-1} \ll 1\) and make the mapping \(u_{l,p} \rightarrow u(x;p)\), \(T_{l,p} \rightarrow T(x;p)\) and, consistently, \(u_{l\pm 1,p} \rightarrow u(x \pm \epsilon;p)\), \(T_{l\pm 1,p} \rightarrow T(x \pm \epsilon;p)\). Of course, this is also done for the current, \(j_{l,p} \rightarrow j(x;p)\) and \(j_{l\pm 1,p} \rightarrow j(x \pm \epsilon;p)\). In addition, we expand all the functions evaluated at \(x \pm \epsilon\) in powers of the small parameter \(\epsilon\), which transforms the exact momentum balance Eq. (6) into

\[
u(x;p + 1) - u(x;p) = - \frac{(1 + \alpha)\epsilon^2}{2} \left[ \partial_x \sqrt{\frac{T(x;p)}{\pi}} - \frac{\epsilon}{2} \partial_x^2 u(x;p) + O(\epsilon^2) \right].
\]

(A4)

It is clearly seen that the evolution of \(u(x;p)\), being proportional to \(\epsilon^2\), is very slow in discrete time. This suggests the introduction of a continuous time scale \(t = p\Delta t\), with \(\Delta t = \epsilon^2\) \[31\]. Over this scale, \(u(x;p + 1) - u(x;p) =\)
\[ \Delta t \partial_t u(x, t) + O(\epsilon^4) \] and we thus finally have
\[ \partial_t u(x, t) = \frac{1 + \alpha}{2} \left[ -\partial_x \sqrt{\frac{T(x, t)}{\pi}} + \frac{\epsilon}{2} \partial_x^2 u(x, t) + O(\epsilon^2) \right]. \tag{A5} \]

Next, we repeat the above procedure for the balance of energy equation. The corresponding expressions are much lengthier than those for the average velocity above and thus we only give the final expressions for the averages of the dissipative and flux terms on the rhs of Eq. (5). For the dissipative term, we have
\[ \frac{(\alpha^2 - 1)\epsilon}{4} (\Theta(v_{l,p} - v_{l+1,p}) (v_{l+1,p} - v_{l,p})^2 + \Theta(v_{l-1,p} - v_{l,p}) (v_{l-1,p} - v_{l,p})^2) \]
\[ = \frac{(\alpha^2 - 1)\epsilon}{2} \left[ T(x; p) - 2\epsilon \sqrt{\frac{T(x; p)}{\pi}} \partial_x u(x; p) \right] + O((\alpha^2 - 1)\epsilon^3), \tag{A6} \]
again within the local equilibrium approximation. For the flux terms,
\[ \langle J_{l,p} - J_{l-1,p} \rangle = (1 + \alpha)\epsilon^2 \left\{ \partial_x \left[ u(x; p) \sqrt{\frac{T(x; p)}{\pi}} \right] - \frac{1}{4} \epsilon \left\{ \partial_x^2 T(x; p) + 2 \partial_x [u(x; p) \partial_x u(x; p)] \right\} + O(\epsilon^2) \right\}. \tag{A7} \]

The evolution equation for the energy \( e \) is then
\[ \langle e(x; p + 1) - e(x; p) \rangle = \frac{(\alpha^2 - 1)\epsilon}{2} \left[ T(x; p) - 2\epsilon \sqrt{\frac{T(x; p)}{\pi}} \partial_x u(x; p) \right] \]
\[ - (1 + \alpha) \left\{ \epsilon^2 \partial_x \left[ u(x; p) \sqrt{\frac{T(x; p)}{\pi}} \right] - \frac{1}{4} \epsilon^3 \left\{ \partial_x^2 T(x; p) + 2 \partial_x [u(x; p) \partial_x u(x; p)] \right\} \right\} \]
\[ + O((\alpha^2 - 1)\epsilon^3) + O(\epsilon^4). \tag{A8} \]

Going again to the continuous time variable,
\[ \partial_t e(x, t) + O(\epsilon^2) = \frac{\alpha^2 - 1}{2\epsilon} \left[ T(x, t) - 2\epsilon \sqrt{\frac{T(x, t)}{\pi}} \partial_x u(x, t) \right] - (1 + \alpha) \partial_x \left[ u(x, t) \sqrt{\frac{T(x, t)}{\pi}} \right] \]
\[ + \frac{1 + \alpha}{4} \epsilon \left\{ \partial_x^2 T(x; p) + 2 \partial_x [u(x; p) \partial_x u(x; p)] \right\} + O(\epsilon^2) + O((\alpha^2 - 1)\epsilon). \tag{A9} \]

In order to have a consistent limit over the continuous time \( t \), \( \alpha^2 - 1 \) must be of the order of \( \epsilon \). Hence, we introduce the macroscopic inelasticity \( \nu \), as defined in Eq. (15), which is assumed to be of order of unity. This implicitly assumes that the underlying microscopic dynamics is quasi-elastic, since \( 1 - \alpha \sim \nu \epsilon = \nu/L \ll 1 \) \cite{[20]}. With this definition,
\[ \partial_t e(x, t) = -\nu T(x, t) - (2 - \nu \epsilon) \partial_x \left[ u(x, t) \sqrt{\frac{T(x, t)}{\pi}} \right] + \epsilon \left\{ 2\nu \sqrt{\frac{T(x, t)}{\pi}} \partial_x u(x, t) + \frac{1}{2} \partial_x^2 T(x, t) + \partial_x [u(x, t) \partial_x u(x, t)] \right\} \]
\[ + O(\epsilon^2). \tag{A10} \]

Finally, the hydrodynamic equations for the average velocity and the temperature up to order \( \epsilon = L^{-1} \) are obtained by taking into account the definition of the macroscopic inelasticity (15) and the identity \( e(x, t) = u^2(x, t) + T(x, t) \), which yields
\[ \partial_t u(x, t) = -\partial_x \sqrt{\frac{T(x, t)}{\pi}} + \frac{\epsilon}{2} \left[ \nu \partial_x \sqrt{\frac{T(x, t)}{\pi}} + \partial_x^2 u(x, t) \right] + O(\epsilon^2), \tag{A11a} \]
\[ \partial_t T(x, t) = \frac{2\nu}{\pi} \partial_x \left[ u(x, t) \sqrt{\frac{T(x, t)}{\pi}} \right] - \frac{1}{\pi} \partial_x^2 T(x, t) + \partial_x [u(x, t) \partial_x u(x, t)] + O(\epsilon^2). \tag{A11b} \]
\[ \partial_t T(x,t) = -\nu T(x,t) - 2\sqrt{\frac{T(x,t)}{\pi}} \partial_x u(x,t) + \epsilon \left\{ 3\nu \sqrt{\frac{T(x,t)}{\pi}} \partial_x u(x,t) + \frac{1}{2} \frac{\partial^2 T(x,t)}{\partial x^2} + [\partial_x u(x,t)]^2 \right\} + O(\epsilon^2). \]  

Note that the parameter \( \epsilon \) can be understood as the ratio of a microscopic length (1 lattice site) to a macroscopic length (the total length \( L \)). In this way, the condition \( \epsilon \ll 1 \) can be conceived as the usual small Knudsen number condition for the validity of a continuum description. A more detailed look at this issue can be found in the following section.

**Appendix B: Boundary layer calculations**

Now we solve the hydrodynamic equations to the lowest order, but incorporating the boundary layers close to the system edges. Thus, we consider Eqs. (A11) with the boundary conditions in the half interval \([0, 1/2]\) and \( T(x,t) \) being an odd and an even function with respect to \( x = 1/2 \), respectively.

We focus on stationary solutions of the hydrodynamic equations. Then, we can simplify the notation by introducing \( t \equiv \partial_x \). Using the terminology in Ref. [78], the “outer” solutions \( \{u_{\text{out}}(x), T_{\text{out}}(x)\} \) satisfy the equations

\[
\left( \sqrt{\frac{T_{\text{out}}}{\pi}} \right)' = 0, \quad -\nu T_{\text{out}} - 2\sqrt{\frac{T_{\text{out}}}{\pi}} \nu u'_{\text{out}} = 0. \quad (B2)
\]

These equations are first-order in space and thus we can only impose the boundary conditions at one of the endpoints of the half interval \([0, 1/2]\), specifically those at the system centre \( x = 1/2 \). Then,

\[
T_{\text{out}} = T_0, \quad u_{\text{out}}(x) = -\nu \frac{\sqrt{\pi}}{2} T_0 \left( x - \frac{1}{2} \right), \quad (B3)
\]

where \( T_0 \) remains undetermined.

The explicit form of the outer (bulk) solutions allows us to have a more precise look at the necessary condition to have a meaningful continuum description. Making use of Eq. (B3), we can readily identify a macroscopic length \( \xi = (\nu \sqrt{\pi}/2)^{-1} \) (the inverse of the factor in front of the expression for \( u_{\text{out}}(x) \), apart from \( \sqrt{T_0} \) that carries with it the velocity dimensions). The macroscopic length scale in the continuum is \( \epsilon = L^{-1} \), corresponding to one lattice site, and then the sought condition would be

\[
\frac{\epsilon}{\xi} = \frac{\nu \sqrt{\pi}}{2L} = \frac{(1 - \alpha^2) \sqrt{\pi}}{4} \ll 1, \quad (B4)
\]

where we have again used the definition of \( \nu \), Eq. (15). This is nothing but a condition of quasielasticity for the microscopic dynamics, a restriction that is already known for other lattice models [62, 80, 83]. Several comments are in order. First, note that this condition has already been imposed when we have derived the hydrodynamic equations, specifically when we have assumed that the macroscopic inelasticity \( \nu \) is of the order of unity. Second, this is physically reasonable: the magnitude of the gradients in granular systems is often typically controlled by the inelasticity, not by the imposed boundary conditions. Therefore, the microscopic inelasticity \( 1 - \alpha \) must be small in lattice models to ensure that the gradients are not so large that the lattice is unavoidable “seen” and a continuum description is not possible. Third, despite the quasielasticity of the microscopic dynamics, the observed macroscopic behaviour does not correspond to a quasielastic granular fluid, because the macroscopic inelasticity \( \nu \) (also called the cooling rate) is not small.

A different distinguished limit is obtained within the boundary layer, with the scaling

\[
x = \epsilon X, \quad \frac{d}{dX} = \epsilon^{-1} \frac{d}{dx}. \quad (B5)
\]

It is \( X \) that the “inner” solution within the boundary layer, which we denote by \( \{u_{\text{inn}}(X), T_{\text{inn}}(X)\} \), depends on. The dominant terms in the balance equations become

\[
- \frac{d}{dX} \left( \sqrt{\frac{T_{\text{inn}}}{\pi}} \right) + 2 \frac{d^2}{dX^2} u_{\text{inn}} = 0, \quad (B6a)
\]

\[
- 2\sqrt{\frac{T_{\text{inn}}}{\pi}} \frac{d}{dX} u_{\text{inn}} + \frac{1}{2} \frac{d^2}{dX^2} T_{\text{inn}} + \left( \frac{d}{dX} u_{\text{inn}} \right)^2 = 0. \quad (B6b)
\]

We have to solve the above equations with the boundary conditions (B1a) at \( X = 0 \), that is,

\[
u_{\text{inn}}(X = 0) = 0, \quad T_{\text{inn}}(X = 0) = 1, \quad (B7)
\]

and the matching conditions [78]

\[
\lim_{X \to \infty} T_{\text{inn}}(X) = \lim_{x \to 0} T_{\text{out}}(x) = T_0, \quad (B8a)
\]

\[
\lim_{X \to \infty} u_{\text{inn}}(X) = \lim_{x \to 0} u_{\text{out}}(x) = \frac{\nu}{4} \sqrt{\pi T_0}. \quad (B8b)
\]

The latter conditions assure that the solution in the boundary layer smoothly matches the outer solution in the bulk. These matching conditions make it possible to determine the unknown bulk temperature \( T_0 \), as shown below.

Equation (B6a) is integrated straightforwardly to give

\[
\frac{d}{dX} u_{\text{inn}} - 2\sqrt{\frac{T_{\text{inn}}}{\pi}} = - \frac{d}{dX} \sqrt{\frac{T_0}{\pi}}, \quad (B9)
\]
where the constant on the rhs has been obtained by matching the outer and inner solutions to the lowest order; note that \( du_{\text{inn}}/dX \to 0 \) for \( X \to \infty \). Substitution of (B9) into (B8b) yields, after some simple algebra

\[
\frac{d^2}{dX^2} \theta + \frac{8}{\pi} \left( 1 - \sqrt{\theta} \right) = 0, \quad \theta = \frac{u_{\text{inn}}}{T_0}.
\]

(B10)

A first integral can be directly derived from this equation,

\[
\frac{1}{2} \left( \frac{d}{dX} \theta \right)^2 + \frac{8}{\pi} \left[ \theta - \frac{2}{3} \theta^{3/2} \right] = D,
\]

(B11)

where \( D \) is a constant. Again, \( D \) is obtained by matching arguments in the limit as \( X \to \infty \), for which we have that \( \theta \to 1 \) and \( d\theta/dX \to 0 \). Therefore, \( D = \frac{8}{3\pi} \) and

\[
\frac{d}{dX} \theta = -\frac{4}{\sqrt{\pi}} \sqrt{\frac{1}{3} + \frac{2}{3} \theta^{3/2} - \theta}.
\]

(B12)

We have chosen the minus sign on the rhs because \( d\theta/dX \) must be negative in the boundary layer, since the bulk temperature \( T_0 < 1 \) as a consequence of the dissipative character of the dynamics.

We do not need to solve Eq. (B12) to obtain \( T_0 \), which is our main goal. Going back to Eq. (B9), we can rewrite it as

\[
\frac{d}{dX} u_{\text{inn}} = 2\sqrt{\frac{T_0}{\pi}} \left( \sqrt{\theta} - 1 \right)
\]

(B13)

and combining it with Eq. (B12),

\[
du_{\text{inn}} = -\frac{\sqrt{T_0}}{2} \frac{\sqrt{\theta} - 1}{\sqrt{\frac{1}{3} + \frac{2}{3} \theta^{3/2} - \theta}} d\theta.
\]

(B14)

This equation allows us to calculate \( u_{\text{inn}} \) as a function of \( \theta \), taking into account that \( u_{\text{inn}}(X = 0) = 0 \) and \( \theta(X = 0) = T_0^{-1} \).

\[
u_{\text{inn}}(\theta) = \sqrt{T_0} \left[ \sqrt{\frac{1}{3} + \frac{2}{3} T_0^{-3/2} - T_0^{-1}} - \sqrt{\frac{1}{3} + \frac{2}{3} \theta^{3/2} - \theta} \right].
\]

(B15)

Now we impose the matching conditions in Eq. (B8), that is, \( \{ u_{\text{inn}} \to \nu \sqrt{\pi T_0}/4, \theta \to 1 \} \) in the limit as \( X \to \infty \). Hence, one can write that

\[
\frac{\sqrt{\pi}}{4} \nu = \varphi(T_0), \quad \varphi(T_0) = \sqrt{\frac{1}{3} + \frac{2}{3} T_0^{-3/2} - T_0^{-1}}.
\]

(B16)

The function \( \varphi(T_0) \) on the rhs is a monotonically decreasing function of \( T_0 \) (recall that \( 0 \leq T_0 \leq 1 \)),

\[
\lim_{T_0 \to 0} \varphi(T_0) = \infty, \quad \lim_{T_0 \to 1} \varphi(T_0) = 0, \quad \frac{d\varphi(T_0)}{dT_0} < 0.
\]

(B17)

Thus, Eq. (B16) is the desired expression for the bulk temperature as a function of the macroscopic inelasticity \( \nu \), since it univocally gives \( T_0 \) for each value of \( \nu \).

In Fig. 4, a good agreement is shown between the prediction for \( T_0 \) obtained here and numerical simulations. The discrepancies remain quite small for \( \nu \leq 2 \), becoming only larger for the highly dissipative case \( \nu = 20 \). It must be taken into account that the bulk temperature \( T_0 \) has been assumed to be of the order of unity in our theory, whereas Eq. (B16) implies that it becomes very small for high \( \nu \), specifically

\[
T_0 \sim \left( \frac{32}{3\pi} \right)^{2/3} \nu^{-4/3}, \quad \nu \gg 1.
\]

(B18)

This means that a more elaborate theory, corresponding to a different dominant balance in the hydrodynamic equations, might be necessary in the highly dissipative limit \( \nu \gg 1 \). See also next section for other possible sources of discrepancy between our theory and the numerical results as \( \nu \) increase.

One can also examine the role of the boundary layers in Fig. 4. The boundary layers are barely noticeable for the smaller values of \( \nu, \nu = 0.2 \) and \( \nu = 0.02 \), for which the bulk temperature is close to unity. In addition, these small values of \( \nu \) necessarily bring about a stronger noise in the averages, since the bulk temperature is larger. As \( \nu \) increases, the bulk temperature decreases and so do the fluctuations while the boundary layers become more visible. In Figs. 2 and 3 it seems that the boundary layer is wider for \( \nu = 2 \) than for \( \nu = 20 \). This is reasonable and consistent with the behaviour observed in other models in the strongly dissipative limit \( \nu \gg 1 \) [80, 84], in which the width of the boundary layer algebraically decreases with \( \nu \) and vanishes in the limit as \( \nu \to \infty \). Besides the above calculations show that, as a function of the system size, the width of the boundary layer in the \( x \) variable is expected to be of order \( L^{-1} \). This scaling is not clearly confirmed in our simulations, which seem to indicate that this width does not go to zero in the limit as \( L \to \infty \), especially for the highest macroscopic inelasticity \( \nu = 20 \). This discrepancy might also be mended by a more elaborate theory in the highly dissipative case, see also next section.

Finally, it should also be noted some kind of “boundary resistance”, that is, a difference between the actual value of \( T \) at the boundary and the value imposed by the thermostat in the simulation (\( T = 1 \)). This phenomenon is known to appear in nonlinear transport problems [85] and has been already observed in other models [80].

**Appendix C: Non-Gaussian local distributions and spatial correlations.**

Here we investigate the validity of the local equilibrium approximation. A first check is obtained by measuring the local velocity pdf. This is done in Fig. 5. In the top row, we consider positions close to the boundaries and deviations from Gaussianity are already apparent for \( \nu = 0.2 \), with an anomalous but localised peak at \( \nu = \)}
Figure 5. Local velocity distributions $P(v, x)$ in the stationary state for different values of $\nu$ and different positions. All the plots correspond to the largest considered system size $N = 1600$. Top row: position close to the left boundary, $x \to 0$ and $x = 0.1$. Bottom row: bulk position $x = 0.5$.

These discrepancies become even more patent as $\nu$ increases, with the emergence of asymmetric tails in the pdf. The pdf in the bulk of the system is presented in the bottom row of the same figure. Again, discrepancies arise and are evident, with an analogous localised peak at $v = 0$ for $\nu = 0.2$ that splits into two symmetric peaks for larger $\nu \gtrsim 1$ [87, 88]. Note that the tails remain symmetric with respect to $v = 0$ in the bulk, however.

Our conclusion is that some of the observed discrepancies as $\nu$ increases stem from the non-Gaussianities described above. Certainly, the non-Gaussianities are more important at the boundary layers than in the bulk, where the deviations seem to be milder, especially for the not-so-large inelasticity $\nu = 2$. Anyhow, non-Gaussianities may be responsible for the deviations of the actual bulk temperature $T_{1/2}$ from the theoretical description $T_0$ for $\nu \gtrsim 1$. It has to be taken into account that we have assumed that the local equilibrium approximation holds both in the bulk and at the boundary layers for deriving Eq. (20). It is remarkable that the relative error between $T_0$ and $T_{1/2}$ remains under ten per cent for $\nu = 2$, despite the large discrepancies at the boundaries that include asymmetry with respect to $v = 0$.

Also, we have looked into the nearest-neighbour correlations $d_l = \langle v_{l-1} v_l \rangle - \langle v_{l-1} \rangle \langle v_l \rangle$, which have been assumed to vanish upon writing the local equilibrium approximation [7]. This assumption is consistent with the Molecular Chaos hypothesis, in which these correlations are assumed to be of the order of $L^{-1}$. Taking into account that we have incorporated $O(L^{-1})$ corrections into our theory, these correlations are another possible source for discrepancies and should be investigated. In Fig. 6, the numerical evaluation of the nearest-neighbour correlations is displayed. It is clearly seen that $d_l$ is always different from zero and for large $L$ (keeping $\nu$ constant) displays a negative plateau in the bulk. On the one hand, the value of the correlations at such a plateau seems to be independent of $L$, which means that Molecular Chaos is violated in the bulk. On the other hand, the measured value is rather small and thus it seems that correlations are not the main source for the observed discrepancies.

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The width of the boundary layers typically scales as $\nu^{-1/2}$, as shown for instance in Ref. [80] for the dissipative version of the Kipnis-Marchioro-Presutti model.

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In other lattice models, the typical scaling has been found to be $t \propto \epsilon^3$ as in Refs. [80, 83, 90].

Again, this scaling for the macroscopic inelasticity is different from the one found in other models, as a consequence of the different scaling of the continuous time variable. Notwithstanding, the underlying microscopic dynamics is quasi-elastic in all cases.

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