Loading and cooling in an optical trap via hyperfine dark states

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We present an optical cooling scheme that relies on hyperfine dark states to enhance loading and cooling atoms inside deep optical dipole traps. We demonstrate a sevenfold increase in the number of atoms loaded in the conservative potential with strongly shifted excited states. In addition, we use the energy selective dark state to efficiently cool the atoms trapped inside the conservative potential rapidly and without losses. Our findings open the door to optically assisted cooling of trapped atoms and molecules which lack the closed cycling transitions normally needed to achieve low temperatures and the high initial densities required for evaporative cooling.

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I. INTRODUCTION

Ultracold quantum gases have attracted much attention in recent decades as versatile platforms for investigating strongly correlated quantum systems [1] and as the basis for quantum technologies exploiting atomic interferometry [2,3]. Cooling of an atomic gas to the required temperatures requires a multistage process: laser cooling in a magneto-optical trap (MOT); sub-Doppler cooling; loading into a conservative magnetic or optical trap; evaporative cooling. Although quite efficient, this process is only possible for a small subset of alkali and alkali-earth-like atoms that can be initially cooled to low temperature by optical means.

The sub-Doppler cooling phase typically uses light red detuned from the $F \rightarrow F = F + 1$ cycling transition of a $D_2$ line $nS_{1/2} \rightarrow nP_{3/2}$ [4], and can be understood in terms of the "Sisyphus effect" [5–8]. Sub-Doppler cooling schemes involving dark states (DSs) [9] have emerged as a powerful alternative; they are known as gray molasses. Very recently they have been pivotal in obtaining an all-optical Bose-Einstein condensate in microgravity [10] and a degenerate Fermi gas of polar molecules [11]. The DSs are coherent superpositions of internal and external states, decoupled from the optical field via electromagnetically induced transparency [12]; their creation does not require cycling $F \rightarrow F' = F + 1$ transitions, but can rely on any transitions of the $F \rightarrow F' \leq F$ form.

To prevent the expansion of the cold atom cloud during cooling, it is tempting to combine DS sub-Doppler cooling with spatial confinement in a far-off-resonance optical dipole trap (FORT). However, the DSs are then strongly modified by the trap potential, which typically shortens their lifetime and eventually couples them to the light field.

In this paper, we show that DS cooling can be used in combination with FORT when strong differential light shifts are present. We use the effect of FORT trapping light close to the $5F_3/2 \rightarrow 4D_{3/2,5/2}$ transitions of rubidium at 1529 nm [13,14] to maximize the cooling action on the surrounding of the trap and at its borders. We observe an order of magnitude improvement in the number of trapped atoms. Additionally, we explore the possibility of cooling the atomic ensemble in the FORT. We achieve a notable temperature reduction of the trapped sample in a few ms and without losing atoms and analyze the limitations of the protocol and the possible workarounds.

Gray molasses commonly rely on Zeeman dark states (ZDSs), i.e., linear combinations of Zeeman sublevels not coupled to lasers blue-detuned with respect to the cooling transition [15]. Generally, ZDSs are not eigenstates of the kinetic-energy operator, and free evolution turns them into bright, absorbing states. Moreover, the presence of a repumper laser spoils these ZDSs, limiting the cooling efficacy [16,17].

However, with a configuration where the repumper and cooler frequencies are equally detuned from the excited state, i.e., a Raman configuration [see Fig. 1(b)], hyperfine dark states (HDSs) can be created. These HDSs have been rigorously studied in the context of coherent population trapping and electromagnetically induced transparency [9,18,19], but not yet for gray molasses cooling.

The existence of these HDSs depends on the number of degenerate ground ($N_g$) and excited ($N_e$) Zeeman states [9,15]: (i) for $N_g > N_e$, a DS always exists provided the laser frequencies match the Raman condition; (ii) for $N_g \leq N_e$, additional conditions on the complex Rabi frequencies must be satisfied. For instance, DSs where the connectivity between ground and excited states forms a loop exist if the summed phase of each complex Rabi frequency is 0 mod$(2\pi)$ [9,20]. These states are particularly relevant for gray molasses cooling because they

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can be eigenstates of the momentum, thus stable both under free evolution and in the presence of a slowly varying external potential.

The DSs for case (ii) can exist in the usual $\sigma_+ - \sigma_-$ cooling configuration, provided that the lasers tuned to the cooling $F = 2 \rightarrow F' = 2$ and repumping $F = 1 \rightarrow F' = 2$ transitions are both phase locked and retroreflected to fulfill the phase requirements on the Rabi frequencies. In addition, the excited Zeeman states $|F' = 2, m'\rangle$ must be degenerate [see Fig. 1(b)], e.g., at zero magnetic field or when the connected states experience the same light shifts in the presence of far off resonance light. We used a numerical approach to ascertain the exact DSs for our configuration where light at 1560 nm is used for trapping (see the Appendix).

II. EXPERIMENT

Our experimental scheme is described in Ref. [21]. We load and cool $^{87}$Rb atoms at the center of an optical cavity used for the 1560-nm FORT. During the entire MOT stage, the FORT depth is maintained at $\sim 27 \mu$K. After 2 s of MOT loading, we realize a compressed MOT phase (CMOT) by detuning for 40 ms the MOT beams to $\Delta = -\hbar \Gamma$ to the red of the $F = 2 \rightarrow F' = 3$ atomic resonance and decreasing the optical power by a factor 10 ($\Gamma = 2\pi 6.066$ MHz is the natural linewidth of the Rb D2 line). Throughout this process, the MOT region overlapping the FORT remains effectively dark due to the large excited state light shifts. The result is a larger density in this region, in a similar fashion to what is reported for the dark-SPOT MOT [22]. At this point we begin the sub-Doppler cooling phase. The FORT is ramped up to $170 \mu$K, simultaneously the cooling beams are detuned to the red in 200 $\mu$s, and we illuminate the atoms for 1 ms. We then turn off the cooling and repumping beams and hold the atoms for 500 ms in the FORT. We finally measure the number of trapped atoms by absorption imaging.

III. RESULTS

A. FORT loading

Figure 2(a) shows the number of captured atoms as a function of $\Delta$. We used two molasses configurations: first [blue points in Fig. 2(a)] with a repumper at fixed frequency, on-resonance to the $F = 1 \rightarrow F' = 2$ transition; second [red points in Fig. 2(a)] with the repumper phased-locked to the cooling laser by using an electro-optic modulator at the hyperfine frequency $\omega_{hf} = 6.83468$ GHz. This latter configuration leads to the Raman configuration of Fig. 1(b). We can identify three regions: (i) $|\Delta|/\Gamma < 30$, where standard $F = 2 \rightarrow F' = 3$ sub-Doppler cooling takes place, leading to optimum loading at $|\Delta|/\Gamma \approx 15$ for both configurations; (ii) $30 < |\Delta|/\Gamma < 60$; (iii) $60 < |\Delta|/\Gamma$, where gray molasses occurs, respectively, on the $F = 2 \rightarrow F' = 2$ and $F = 2 \rightarrow F' = 1$ transitions.

In the first configuration, only ZDSs can form. Gray molasses leads to a 2.5 times more efficient loading at $-40\Gamma$, i.e., blue detuned from the $F = 2 \rightarrow F' = 2$ transition. This coincides with an optimum in density when the cooling is applied without the FORT [7]. We find that almost all the atoms are in the $F = 1$ manifold [blue points in Fig. 2(b)] even though the ZDSs are in the $F = 2$ manifold. This can be understood as the effect of the repumper weakly coupling the DSs [16,17], leading to eventual decay in $F = 1$.

In the second configuration HDSs can form. We observe a more efficient loading, with an outstanding narrow peak at $-43.5\Gamma$, seven times higher than standard red detuned molasses and more than twice higher than ZDS gray molasses.
sweeping the atoms in the FORT from its borders and towards its center, by trapped atoms at specific radial positions, or to progressively cool 700 $\mu$K dependent excited-state light shift, represented for a FORT depth of 43.5\Gamma. The effect can be exploited to selectively cool the optically trapped atoms at specific radial positions, or to progressively cool the atoms in the FORT from its borders and towards its center, by sweeping $\Delta$ as indicated by the gray arrow.

A similar peak is also observed at $-69\Gamma$, i.e., blue detuned from the $F = 2 \rightarrow F' = 2$ transition; its lower loading efficiency is partially due to the ramping process of cooler+repumper frequencies from the MOT frequencies, which involves crossing the $F = 2 \rightarrow F' = 2$ transition with associated heating. The atomic population in this region is balanced for all detunings [red points in Fig. 2(b)]; all atoms are in balanced HDSs, mediated either by the symmetric excited state $m_F = \pm 1$ or via $m_F = 0$.

We now focus on the sharp peak at $-43.5\Gamma$. As in the previous configuration, this coincides with the optimum density reached without FORT-induced light shifts, which is shifted closer to the $F = 2 \rightarrow F' = 2$ transition because of the Raman configuration [7]. This suggests that cooling is mostly efficient outside and on the edges of the FORT. This peak sits atop a broader feature which does not abruptly end at the position of the $F = 2 \rightarrow F' = 2$ resonance [23]. The broader feature shifts and broadens with increasing FORT power while the position of the sharp peak remains constant (Fig. 3). The deleterious effects of the 1560 nm on the formation of the DSSs can be partly offset by tuning the Raman detuning frequency $\delta_R$. Figure 4 shows how the optimum value $\delta_R$ varies with the FORT power. We measure this value by scanning $\delta_R$ to find the maximum peak central density of the atomic cloud. The inset in Fig. 4 shows such a scan for an atomic cloud after 4 ms of gray molasses in free space, without 1560 nm trapping light. The exact Raman condition, $\delta_R = 0$, gives optimal cooling—that is the lowest temperature and highest central density—and the cooling efficiency falls when the Raman detuning is changed, reaching a minimum at $\delta_R = 80$ kHz. Remarkably, the distance between the maximum and the minimum is much smaller than $\Gamma$ as in similar work with Cs atoms [24], but in contrast with Ref. [6].

Increasing the power of the FORT changes how gray molasses works, which is reflected in the loading curves shown in Fig. 3. Two effects could explain our observations. First, the trapping light at 1560 nm induces light shift much larger for the excited-state levels than for the ground-state ones. Off-resonant scattering from excited-state hyperfine levels not involved in the formation of DSSs can hinder ZDSs and HDSs, and the resulting decay rate will be strongly related to the energy shifts. The DSSs’ effective lifetime becomes dependent on the local intensity of the trapped light. Second, HDSs relying on two (or more) magnetic sublevels with different $|m_F|$ are strongly modified and probably inhibited when atoms move in the inhomogeneous FORT intensity. We have performed numerical calculations and found a good agreement between the measured population ratio and the calculated ones [see the violet points in Fig. 2(b) and the Appendix]. A detailed experimental and theoretical study would be needed to further confirm these hypothesis, and is beyond the scope of this work.

FIG. 3. HDS cooling near the $F = 2 \rightarrow F' = 2$ transition at different FORT depths indicated in colors. The sharp peak at $-43.5\Gamma$ is given by atoms cooled just outside the FORT. The broader peak underneath corresponds to atoms residing within the FORT volume. Inset: The radially changing FORT power results in a position dependent excited-state light shift, represented for a FORT depth of 700 $\mu$K. The effect can be exploited to selectively cool the optically trapped atoms at specific radial positions, or to progressively cool the atoms in the FORT from its borders and towards its center, by sweeping $\Delta$ as indicated by the gray arrow.

FIG. 4. Raman detuning optimizing the HDS cooling, as a function of the FORT power. The filled red points correspond to HDSs on the $F = 2 \rightarrow F' = 2$ transition, the empty blue ones to HDSs on the $F = 2 \rightarrow F' = 1$ transition (where no shift is required to optimize cooling). Inset: HDS cooling efficiency as a function of the Raman detuning frequency, without the FORT.

B. HDS and light shifts

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The red points in Fig. 4 refer to the HDS gray molasses on the $F = 2 \rightarrow F' = 2$ transition: an increasing compensation...
shift is required for higher FORT power. Since for the ground level the vector light shifts is zero, the observed shift must be caused by the effect of the excited-state light shifts on the HDSs, however the exact mechanism is still unclear. The blue points refer to the HDS gray molasses on the \(F = 2 \rightarrow F' = 1\) transition \(\Delta/\Gamma = -69\), for which cooling is optimal at \(\delta_R = 0\), independent of the FORT power.

### C. HDS cooling in the FORT

As discussed earlier, Fig. 3 shows that trapped atoms experience position dependent excited-state light shifts, proportional to the local FORT intensity, which allow for addressing spatially these atoms by adjusting \(\Delta\). In addition, the position of atoms within a conservative trap is energy dependent, with higher (lower) energy states spending more time near the edges (center) where the light shift is smaller (larger) [13]. Therefore at any specific detuning \(\Delta\), atoms of a well defined energy should be optimally cooled and fall deeper into the FORT. Consequently, if we progressively increase \(\Delta\), we should be able to continuously cool the atoms from outside the trap to deep inside it, increasing the density while potentially reaching temperatures close to the recoil limit with no atom loss since our cooling relies on DSs.

To investigate this possibility, we adiabatically increase the FORT depth to \(\sim 1\) mK after loading, to increase the excited-state light shifts and, thereby, the energy selectivity. We sweep the detuning in 6 ms from \(-55\Gamma\) to \(-68\Gamma\) and observe a temperature reduction from 198 to \(48\) nK with no atom loss (Fig. 5). A slower sweep does not lower the final temperature, \(\sim 6 \times 10^6\), with no significant difference between the two cases.

![Graph](image)

**FIG. 5.** (a) Ballistic expansion of the atomic cloud released from the FORT: cloud size vs time of flight. The fits (solid lines) yield a temperature of 198 \(\mu\)K before the cooling sweep (black points) and 48 \(\mu\)K after it (red points). Absorption images of the atoms before (b) and after (c) the cooling sweep, taken with a time-of-flight of 700 \(\mu\)s. The atom number in the cross configuration is \(\pm 6 \times 10^6\), with no significant difference between the two cases.

### IV. CONCLUSIONS

To conclude, we have demonstrated the potential of all-optical loading and cooling in a far-off-resonance trap using gray molasses cooling relying on hyperfine dark states, combined with large light shifts of the excited levels. The method could open avenues in the production of ultracold atomic and molecular gases thanks to many key features: high speed, minimal scattering, no atom loss, and no need for cycling transitions. Remarkably, cooling and loading happen simultaneously, which solves the mode matching issue for optical dipole traps preventing the need for an intermediate step often relying on magnetic trapping. The cooling scheme looks optimal for the rapid production of ultracold gases in unusual geometries [25,26] and environments [27,28], and could provide a complementary alternative for all optical degenerate gases production [29,30].

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### APPENDIX: DETERMINATION OF THE HDS

To investigate the formation of DSs, we use a one-dimensional model of the gray molasses [5,6] that includes the presence of two counterpropagating beams with opposite circular polarization \((\sigma_+ / \sigma_-)\) configuration, each carrying both cooler and repumper frequency, and the light shifts due to the trapping potential.

We consider the Hamiltonian as a sum of three terms: (i) the atomic Hamiltonian \(H_a\), taking into account only the \(5S_{1/2} \rightarrow 5P_{3/2}\) states connected by the \(D2\) transition; (ii) the molasses Hamiltonian \(H_{mol}\), modeling the interaction of the atom with the cooler and repumper light; (iii) the second-order Stark Hamiltonian \(H_{\text{dip}}\), describing the interaction of the atom with the dipole trap light at 1560 nm, causing scalar and vector light shifts:

\[
H = H_a + H_{mol} + H_{\text{dip}},
\]

\[
H_a = \sum_{n=SS,SP,PP} \sum_{n,F,m} E(n,F,m)|n,F,m\rangle\langle j,F,m|,
\]

\[
5S_{1/2} \rightarrow 5P_{1/2} D1\text{ transition}, \text{where no higher } F' = 3 \text{ level exists. This solution should permit recoil temperatures in time scales of milliseconds without loss of atoms.}
\]
\[
H_{\text{mol}} = -\frac{\hbar}{2}(\Omega_R e^{-i\omega_R t} + \Omega_c e^{-i\omega_c t}) \\
\times (\hat{e}_+ e^{i\vec{k}z} + \hat{e}_- e^{-i\vec{k}z}) \mathbf{\hat{a}} + \text{h.c.}], \tag{A3}
\]

\[
H_{\text{dip}} = \left[ \frac{|E_0|^2}{2} \right] e^{i\vec{v}\vec{k}z} \sum_{\mu} 2 \frac{P_2(E_\mu - E_k)}{(E_\mu - E_k)^2 - (\hbar\omega_R)^2} d_\mu. \tag{A4}
\]

In the \( H_{\text{dip}} \) term, \( E(n, F, m) \) denotes the energies of the ground and excited hyperfine manifolds, hereafter indicated with \( 5\Sigma \) and \( 5\Pi \), with the definition \( E(5\Sigma, 1, m) = 0 \).

In \( H_{\text{mol}} \), \( \Omega_{R(C)} \equiv \sqrt{\Gamma_{R(C)}}/(2\kappa) \) is the repumper (cooler) Rabi frequency in terms of the saturation intensity \( \kappa = 1.67 \) mW/cm\(^2\) and the excited-state linewidth \( \Gamma/(2\pi) = 6.065 \) MHz. \( \vec{a} \) is the raising operators of atomic levels whose matrix elements are the 6-\( j \) Wigner coefficient, and \( \omega_{R(C)} \) is the repumper (cooler) angular frequency. Each molasses beam carries the repumper and cooler components, with Rabi frequencies \( \Omega_{R(C)} = 4.2 \Gamma \) and \( \Omega_{C} = 1.2 \Gamma \), corresponding to the total intensities, i.e., summed on the six beams in our experiment. Equation (A3) is further simplified by neglecting the coupling of the cooler with the \( F = 1 \rightarrow F' \) transitions, due to very large detuning (~10\(^4\)\( \Gamma \)), and likewise the coupling of the repumper with \( F = 0 \rightarrow F' \) transitions:

\[
\frac{H_{\text{mol}}}{\hbar} \simeq -\frac{\hbar}{2}(\Omega_R e^{-i\omega_R t} P_2 \mathbf{\hat{a}} P_1 + \Omega_c e^{-i\omega_c t} P_2 \mathbf{\hat{a}} P_2) \times (\hat{e}_+ e^{i\vec{k}z} + \hat{e}_- e^{-i\vec{k}z}) + \text{h.c.}], \tag{A5}
\]

\( H_{\text{dip}} \) describes the atomic interaction with the 1560-nm light \( (\vec{E}(t) = \frac{1}{2}\hat{\rho}_0 e^{i\omega_0 t} + \text{c.c.}) \), with \( d_\mu, \epsilon_{\mu}(\mu = 0, \pm 1) \) being the spherical components of the electric dipole operator and of the polarization vector; \( E_\mu \) is the energy of the right ket and \( k \) labels the intermediate state.

One can apply a unitary transformation

\[
U = P_1 + \exp[i(\omega_R - \omega_C)t] P_2 + \exp[i\omega_R t] P_e
\]

where \( P_1, P_2, P_e \) are the projectors on the ground lower \( (|5\Sigma, F = 1, m\rangle) \), ground upper \( (|5\Sigma, F = 2, m\rangle) \), and electronic excited \( (|5\Pi, F', m\rangle) \) hyperfine levels, respectively. Under the unitary transformation \( U \), the Hamiltonian is modified: (i) the time dependence of \( H_{\text{mol}} \) drops, (ii) the atomic energy levels are shifted, \( E'(5\Sigma, 1, m) = 0, E'(5\Sigma, 2, m) = E(5\Sigma, 2, m) - \hbar(\omega_R - \omega_C), E'(5\Pi, F, m) = E(5\Pi, F, m) - \hbar\omega_C \).

We set the frequency difference \( \omega_R - \omega_C \) to match the hyperfine separation of a free atom and then diagonalize the Hamiltonian for different values of the position \( z \). At each position, we identify the DSs as those whose projection in the excited level 5\( \Pi \) is below an arbitrary threshold value, chosen equal to 0.1; in practice a given eigenstate \( |\psi\rangle \) of the total Hamiltonian is considered dark if \( \langle\psi|P_2|\psi\rangle < 0.1 \), which is equivalent to requiring a scattering rate below 0.1\( \Gamma \). We find that, in the \( \sigma_+/-\sigma_- \) configuration of our 1D model, the light shifts are uniform in space, i.e., independent of \( z \). At each point the polarization of the molasses lasers field is linear, with direction varying periodically in space: the polarization vector winds like a helix. As a consequence, it is expected that the number of DSs does not depend on the position \( z \).

In the set of DSs, we operate a further selection by keeping only those composed of hyperfine states \( (F, m = \pm 1) \) and discarding those formed by \( (F, 2) \) and \( (F, 0) \). The former are linear superpositions of states with momenta \( p \pm \hbar k_0 \), that can be stable under kinetic-energy evolution for sufficiently small \( p \). On the other hand, the latter are superpositions of states with momenta \( p, p \pm 2\hbar k_0 \) that cannot be eigenstates of the kinetic energy. For molasses times longer than the inverse kinetic energy associated to the two-photon transitions, \( t_{\text{mol}} > (2\hbar k_0^2/m)^{-1} \approx 10 \) \( \mu \)s, the latter states become bright.

To calculate the expected relative populations of the \( F = 1 \) and \( F = 2 \) hyperfine levels at the end of the molasses, we have to account only the DSs selected as described. On each of these states \( |\text{DS}_j\rangle \), we evaluate \( P_{1,j} = |\text{DS}_j| P_1 |\text{DS}_j\rangle \) and \( P_{2,j} = |\text{DS}_j| P_2 |\text{DS}_j\rangle \) and define the relative population \( N_j/N_i = \sum_j p_{2,j}/(p_{1,j} + p_{2,j}) \), implicitly assuming that all DSs are equally populated. The purple points in Fig. 2 report the relative population in the specific case for zero power of the FORT; this configuration approximates the initial molasses, and the condition experienced by the atoms roaming the edges of the FORT trap. A fully time-dependent model, which is beyond the scope of this work, should consider the FORT intensity increase during the 1-ms-long gray molasses phase.

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