Update to single-variable parametric cost models for space telescopes

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Abstract. Parametric cost models are an important tool routinely used to plan missions, compare concepts, and justify technology investments. In 2010, the article, “Single-variable parametric cost models for space telescopes,” was published [H. P. Stahl et al., Opt. Eng. 49(7), 073006 (2010)]. That paper presented new single-variable cost models for space telescope optical telescope assembly. These models were created by applying standard statistical methods to data collected from 30 different space telescope missions. The results were compared with previously published models. A postpublication independent review of that paper’s database identified several inconsistencies. To correct these inconsistencies, a two-year effort was undertaken to reconcile our database with source documents. This paper updates and revises the findings of our 2010 paper. As a result of the review, some telescopes’ data were removed, some were revised, and data for a few new telescopes were added to the database. As a consequence, there have been changes to the 2010 published results. But our two most important findings remain unchanged: aperture diameter is the primary cost driver for large space telescopes, and it costs more per kilogram to build a low-areal-density low-stiffness telescope than a more massive high-stiffness telescope. One significant difference is that we now report telescope cost to vary linearly from 5% to 30% of total mission cost, instead of the previously reported average of 20%. To fully understand the content of this update, the authors recommend that one also read the 2010 paper.

Subject terms: space telescope cost model; parametric cost model; cost model.

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1 Introduction

Parametric cost models are routinely used to plan missions, compare concepts, and justify technology investments. Unfortunately, there is a wide range of parametric models, with no definitive model for the cost of a space telescope optical telescope assembly (OTA). Part of this problem is the limited number of telescopes that have flown in space for which we have data. Another part is the tendency to extrapolate ground telescope rules of thumb to space telescopes. In 2010, these authors published the article “Single-variable parametric cost models for space telescopes.”¹ That paper presented new single-variable cost models for space telescope OTA. These models were created by applying standard statistical methods to data collected from 30 different space telescope missions. The results were compared with previously published models. This paper updates and revises the findings of our 2010 paper.

To fully understand the content of this update, the authors recommend that one also read the 2010 paper.

After the publication of the original paper,¹ we were invited to present our findings to the National Reconnaissance Office (NRO) in September 2010. The NRO Cost Model Office reviewed our database and, while they did not provide us with access to their database or give us any specific data, they did identify specific missions where our two databases disagreed. In response, we conducted a systematic review of all missions in our database and reconciled them with source documents. As a result, some missions have been removed from our cost database. The costs for some missions have been revised, with revisions ranging from slight to dramatic, and several new missions have been added to the database. As a consequence, there have been changes to the 2010 published results and models. But our two most important findings remain unchanged: aperture diameter is the primary cost driver for large space telescopes, and it costs more per kilogram to build a low-areal-density low-stiffness telescope than a more massive high-stiffness telescope. One significant difference is that we now report telescope cost to vary linearly from 5% to 30% of total mission cost, instead of the previously reported average of 20%. The key lesson learned from this process is the need to be precise in all definitions and consistent in their application. Finally, note that the cost model coefficients reported in this paper are inflated to 2011 values.

2 Methodology

The OTA is defined as the subsystem that collects electromagnetic radiation and focuses it (focal) or concentrates it (afocal) into the science instruments (SI). An OTA consists
of the primary mirror, secondary mirror, auxiliary optics, and support structure (such as optical bench or truss structure, primary support structure, secondary support structure or spiders, straylight baffles, mechanisms for adjusting the optical components, and electronics or power systems for operating these mechanisms). An OTA does not include SIs or spacecraft subsystems. An SI is defined as the subsystem that converts electromagnetic radiation into data. An SI includes conditioning optics (e.g., beam splitters, reimaging optics, spectral filters, dispersive elements), mechanisms, detectors, focal planes, and electronics.

OTA cost is defined as the prime contractor’s cost to design, build, and integrate the OTA. OTA cost includes allocated subsystem-level management and systems engineering as well as program-level costs that can be allocated to the subsystem. For all missions in our database except the Orbiting Astronomical Observatory (OAO), OTA cost does not include NASA labor cost. The reason is that, before the James Webb Space Telescope (JWST), NASA’s accounting system did not track the cost of NASA personnel to a given program, such as Hubble or Kepler. NASA did start tracking these costs in the early phases of JWST when it implemented full-cost accounting. But, to compare apples to apples between JWST and prior missions such as Hubble, we remove NASA labor costs from the JWST cost. Regarding the OAO program, NASA personnel performed most of the design, build and integration effort, and these costs were tracked and reported. Therefore, these costs are included in the database. Total mission cost is not life cycle cost. It is defined as only phase A–D cost, excluding launch cost; costs associated with NASA labor (civil servant or support contractors) for program management, technical insight/oversight, or any NASA-provided ground support equipment, e.g., test facilities. Note that excluding NASA labor costs underestimates the true cost of a given mission by at least 15% and maybe by as much as 30%.

Careful review found that we were inconsistent in the application of the above definitions for OTA cost. In the original paper, the database OTA cost for large missions (Hubble, Kepler, JWST, etc.) were correct. They were for only the OTA. But the database cost for the smaller missions (GALEX, IUE, TRACE, and WIRE) were incorrect. The database costs of these missions were not just for the OTA. Rather, they were an “instrument cost,” where an instrument was defined to be an integrated system consisting of an OTA and an SI. Removing the SI cost to yield just an OTA cost dramatically reduced the database values for these small-aperture missions. Additionally, we established a new definition which resulted in significant changes to the database. We decided to exclude thermal/cryogenic control systems from the definition of an OTA. For example, the JWST OTA does not include the cost of the JWST Sunshades. But the old IRAS and Spitzer OTA costs did include the cryogenic system. Removing these costs dramatically changed their previous database values.

Finally, additional resources were used for source documents, including: NAFCOM (NASA/Air Force Cost Model) database, NICM (NASA Instrument Cost Model), NSCKN (NASA Safety Center Knowledge Now), RSIC (Redstone Scientific Information Center), REDSTAR (Resource Data Storage and Retrieval System), SICM (Scientific Instrument Cost Model), project websites, and interviews.

### 2.1 Database Collection

After careful review of source CADRe documents (Cost Analysis Data Requirements), several changes were made to the database. The costs of the Kepler and Wise missions were increased to include program management, systems engineering, and integration and test cost. The costs of GALEX, HiRISE, HUT, OAO-3, UIT, WIRE, and WUPPE were decreased to remove SI costs. The costs of IRAS and Spitzer were decreased by separating cryostat and OTA cost. The cost of SOFIA was decreased by removing the cost of the gimbal structure that holds the SOFIA OTA in the 747 airframe. Several missions were added to the database, specifically: CloudSat, OAO-B/GEP, Herschel, and Planck. JWST costs were updated to the 2011 estimate to complete. Finally, the Hubble OTA and total missions costs were reduced as shown in Table 1. Previously, the cost of the fine guidance sensor (FGS) had been excluded from the OTA cost, but this cost should be allocated to the spacecraft. But management and systems engineering costs allocated to the FGS had not been properly removed. Additionally, the previous total mission cost included phase E operations costs. Finally, the Hubble OTA mass was increased from 2150 to 3180 kg based on better documentation.

### 3 Cost Models

#### 3.1 Pearson Cross-Correlation Analysis of Parameters

Because the database has changed, the cross-correlation matrix has also changed. However, the methodology of using the cross-correlation matrix to guide the statistical analysis has not changed. Therefore, for the sake of brevity, there is no need to show a corrected matrix.

#### 3.2 OTA Cost Versus Total Cost

Changes to the database significantly changed the ratio of OTA cost to total mission cost. Previously, the ratio of OTA cost to total mission cost was spread from a few percent to 65%. The net effect of this spread was to make it appear that on average, the OTA was approximately 20% of total mission cost. But, with the corrections, the OTA cost as a percentage of total mission cost varies linearly as a function of aperture diameter size from a few percent to nearly 30% (Fig. 1). It is hypothesized that the reason for this increase with aperture size is infrastructure and technology reuse. Smaller-aperture missions tend to use existing manufacturing and testing infrastructure, whereas larger-aperture missions often require the design and fabrication of expensive custom infrastructure. Also, smaller missions tend to have higher reuse of existing designs. Finally, the data imply that for small missions, other major subsystems (such as the spacecraft) are a much larger cost for the total mission than the OTA.

An analysis of detailed work breakdown structure documents of the seven missions for which we have such documents shows that, on average, the spacecraft accounts for approximately 25% of the cost, SIs account for 25%, OTAs account for 12%, program management and systems engineering account for 12%, integration and testing account for 10%, and the balance is “other” (Fig. 2). Obviously, there is significant uncertainty in these percentage values.
Using the revised database, single-variable models are created by regressing OTA cost data versus selected parameters in single-variable models (Fig. 3). Note that Fig. 3 shows the regression results for 15 free-flying OTAs in the database. Also, the aperture diameter result changes if CloudSat, Planck, or Herschel is excluded from the regression.

As discussed in the original paper, each regression is evaluated for its goodness of fit and significance via a range of statistical measures, including Pearson $r^2$ coefficient, Student $t$-test $P$ value, and standard percent error (SPE). Pearson $r^2$ (typically denoted as just $r^2$) describes the percentage of agreement between the model and the actual cost in log-log space. For multivariable models, we use adjusted Pearson $r^2$ (or $r^2_{adj}$), which accounts for the number of data points and the number of variables. In general, the closer $r^2$ (or $r^2_{adj}$) is to 1.0% or 100%, the better the model. SPE is a normalized standard deviation of the fit residual (difference between data and fit) to the fit. The closer SPE is to 0, the better the fit. Please note that since SPE is normalized, a small variation divided by a very small parameter coefficient can yield a very large SPE.

The $P$ value is the probability that a fit or correlation would occur if the variables are independent of each other. The closer the $P$ value is to 0, the more significant the fit or correlation. The closer it is to 1, the less significant. For the purpose of our study, we consider any $P$ value of less than 0.10 to be “good” and any $P$ value greater than 0.10 to be “bad.” Also, it is important to consider how many data points are included in a given correlation, fit or regression. Again, for the purpose of our study, based correlation statistical significance, we urge caution before using any regression with fewer than 12 data points.

The variables that yield a significant regression for OTA cost are aperture diameter (meters), primary mirror focal length (meters), volume (cubic meters), pointing stability (arc-seconds), and mass (kilograms). Of these, aperture diameter is the independent variable. The others are dependent variables because they are correlated with diameter. All space telescope OTAs tend to have similar $F/\lambda$ values, so larger apertures have longer focal lengths. Also, pointing stability is directly proportional to resolution, which is defined by aperture diameter. And of course, larger-aperture OTAs have larger volumes and are more massive than smaller-aperture OTAs. Unfortunately, while these authors disagree with the practice, many cost models only use mass to estimate cost. Therefore, we do report mass models in addition to aperture models.

### Table 1: Refinement of Hubble cost knowledge.

| Total Cost Phase A–D (Design and Build) | Old (FY11$) | Revised (FY11$) | Notes |
|----------------------------------------|-------------|-----------------|-------|
| Total optical telescope assembly (OTA) | 4.0 B       | 2.8 B           | Old: NGST cost model database |
| OTA                                    | 0.9 B       | 0.9 B           | Old: allocated fine guidance sensor (FGS) and C&DH PM & SE costs to OTA |
| Optics                                 | 0.7 B       | 0.47 B          | New: REDSTAR 121-4742 |
| Optics control                         | 0.07 B      | 0.08 B          | New: REDSTAR 121-4742 |
| Optical structure                      | 0.08 B      | 0.08 B          | New: REDSTAR 121-4742 |
| Electrical power                       | 0.02 B      | 0.06 B          | New: REDSTAR 121-4742 |
| Structures, mechanisms, support equipment | 0.05 B     | 0.06 B          | New: REDSTAR 121-4742 |
| System level 53%                       | 0.14 B      | 0.15 B          | New: REDSTAR 121-4742 |
| Space telescope level 53%              | 0.01 B      | 0.01 B          | New: REDSTAR 121-4742 |
| FGS                                    | 0.26 B      | 0.26 B          | New: REDSTAR 121-4742 |
| C&DH                                   | 0.08 B      | 0.08 B          | New: REDSTAR 121-4742 |
| Thermal control                        | 0.01 B      | 0.06 B          | New: REDSTAR 121-4742 |
| System level 47%                       | 0.12 B      | 0.12 B          | New: REDSTAR 121-4742 |
| Space Telescope level 47%              | 0.01 B      | 0.01 B          | New: REDSTAR 121-4742 |

### Total cost phase A–E

| Old (FY11$) | Revised (FY11$) | Notes |
|-------------|-----------------|-------|
| Launch      | 0.62 B          | New: REDSTAR 123-1064 (p. 108) |
| Phase E (Operations) | 1.2 B           | New: REDSTAR 123-1064 (pp. 108 and 122) |

### 3.3 Single-Variable Cost Models

Using the revised database, single-variable models are created by regressing OTA cost data versus selected parameters in single-variable models (Fig. 3). Note that Fig. 3 shows the regression results for 15 free-flying OTAs in the database. Also, the aperture diameter result changes if CloudSat, Planck, or Herschel is excluded from the regression.

As discussed in the original paper, each regression is evaluated for its goodness of fit and significance via a range of statistical measures, including Pearson $r^2$ coefficient, Student $t$-test $P$ value, and standard percent error (SPE). Pearson $r^2$ (typically denoted as just $r^2$) describes the percentage of agreement between the model and the actual cost in log-log space. For multivariable models, we use adjusted Pearson $r^2$ (or $r^2_{adj}$), which accounts for the number of data points and the number of variables. In general, the closer $r^2$ (or $r^2_{adj}$) is to 1.0% or 100%, the better the model. SPE is a normalized standard deviation of the fit residual (difference between data and fit) to the fit. The closer SPE is to 0, the better the fit. Please note that since SPE is normalized, a small variation divided by a very small parameter coefficient can yield a very large SPE.

The $P$ value is the probability that a fit or correlation would occur if the variables are independent of each other. The closer the $P$ value is to 0, the more significant the fit or correlation. The closer it is to 1, the less significant. For the purpose of our study, we consider any $P$ value of less than 0.10 to be “good” and any $P$ value greater than 0.10 to be “bad.” Also, it is important to consider how many data points are included in a given correlation, fit or regression. Again, for the purpose of our study, based correlation statistical significance, we urge caution before using any regression with fewer than 12 data points.

The variables that yield a significant regression for OTA cost are aperture diameter (meters), primary mirror focal length (meters), volume (cubic meters), pointing stability (arc-seconds), and mass (kilograms). Of these, aperture diameter is the independent variable. The others are dependent variables because they are correlated with diameter. All space telescope OTAs tend to have similar $F/\lambda$ values, so larger apertures have longer focal lengths. Also, pointing stability is directly proportional to resolution, which is defined by aperture diameter. And of course, larger-aperture OTAs have larger volumes and are more massive than smaller-aperture OTAs. Unfortunately, while these authors disagree with the practice, many cost models only use mass to estimate cost. Therefore, we do report mass models in addition to aperture models.
3.3.1 Cost as a function of aperture diameter

cost-estimating relationship

The difference between this update paper and the original paper is that the diameter exponent has increased from 1.2. The reason for this increase is that the small-aperture mission OTA costs decreased when the SI costs were removed. Also, in the original paper, we asserted that because the diameter exponent of 1.2 was so much less than 2 that the areal cost of large-aperture telescopes was significantly less than the areal cost for small-aperture telescopes. As will be discussed below, the diameter exponent has increased, and it is unclear at this time if this assertion is still valid. We will revisit this issue in a future multivariable paper.

Based on a sample size of 15 free-flying space telescopes, a single-variable cost-estimating relationship (CER) was developed for OTA cost as a function of primary mirror diameter:

\[
\text{OTA cost} \sim \$30\text{M diameter}^{1.4} \quad (N = 15; r^2 = 81\%; \text{SPE} = 122\%).
\]

As indicated by Pearson \(r^2\), diameter is a good predictor of OTA cost. It explains 81% of the OTA cost variation. The most interesting point about this model is the similarity of its coefficient and exponent to our single-variable ground telescope OTA model:

\[
\text{Ground Telescope OTA cost} \sim \$3\text{M diameter}^{1.4}.
\]

One conclusion of this similarity is that, to first order, space telescopes are approximately 10\(\times\) more expensive than ground OTAs as a function of aperture size. One reason for this difference is the mass design rules. Space and ground telescopes are designed to different mass constraints and safety margins. Ground telescopes and their components are more massive, i.e., are stiffer, and are thus easier to handle and manufacture to the required precision.

However, as indicated by the SPE, the new model is a bit noisy. The reason for this noise is that the revised database includes three additional missions that provide significant wavelength diversity because of their longer diffraction limited wavelengths: Cloud Sat \(\Rightarrow 3\) mm, Planck \(= 0.7\) mm, and Herschel \(= 0.08\) mm. Because the original paper did not include CloudSat, Planck, or Herschel, if we exclude these three OTAs, the regression becomes:

\[
y = 0.09\ln(x) + 0.11 \quad R^2 = 0.94
\]

Fig. 1 The corrected database eliminates the data spread of the original paper and indicates that optical telescope assembly (OTA) cost as a percentage of total mission cost varies linearly from a few percent to nearly 30% as a function of aperture diameter.

Fig. 2 Analysis of seven missions indicates that spacecraft and science instruments (SIs) are approximately 50% of total cost and that OTA is 10% to 15%.
OTA cost

\[ \sim 45\text{M diameter}^{2.0} \quad (N = 12; r^2 = 94\%; \text{SPE} = 62\%). \]

In this case, the regression statistics indicate a very good fit to the data. Finally, if we include just Herschel, the regression becomes

OTA cost

\[ \sim 38\text{M diameter}^{1.6} \quad (N = 13; r^2 = 83\%; \text{SPE} = 89\%). \]

Again, the variation of the single-variable model is wavelength diversity. Once a wavelength parameter is added to the model, in a future paper, the diameter exponent should stabilize. Given that long wavelength telescopes are easier to make than short wavelength telescopes. It is likely that the aperture diameter exponent will be between 1.6 and 2.0.

Figure 4(a) shows the data plotted for 17 free-flying telescopes and four attached telescopes (whose data was excluded from the regression). As discussed in the original paper, the attached OTAs are excluded because, while their trend line slope is similar to that of the free-flying telescopes, their leading coefficient is 2 to 3× lower. As a reminder, the attached OTAs are defined as UIT, WUPPE, and HUT, which flies attached to the space shuttle orbiter, and SOFIA, which flies attached to a 747. Note that, in the original paper, the cost for SOFIA was too high because it included the cost of the gimbal. As a result, in the original paper SOFIA’s cost was in family with the three space shuttle missions. However, after removing the cost of the gimbal, SOFIA’s cost dropped to a level between that of the shuttle attached OTAs and the ground OTAs. Regardless, the implication is that the basic engineering issues that drive cost as a function of aperture apply equally to all OTAs but that the lower the mass, i.e., the less stiff the OTA, the more difficult and more expensive it is to fabricate. The relationship between the free-flying and attached OTAs becomes evident if one eliminates the CloudSat, Planck, and Herschel data points [Fig. 4(b)].

Finally, a regression of 21 free-flying missions yields a total mission cost versus aperture diameter CER of

Total cost \[ \sim \text{diameter}^{0.75} \]

\[ (N = 21; r^2 = 80\%; \text{SPE} = 114\%). \]

As indicated by Pearson \( r^2 \), for this regression, diameter explains 80% of the total mission cost variation. As indicated by SPE, this fit has a slightly noisy SNR. The most interesting result of this regression is that the exponent is 0.75. Total mission cost as a function of aperture diameter is flatter than OTA cost. The implication is that for smaller-aperture missions, other costs (maybe spacecraft) dominate the mission cost. This is consistent with the earlier finding that smaller-aperture OTA cost is only 10% of total mission cost.

**3.3.2 Cost as a function of mass**

The difference between this update paper and the original paper is that the mass exponent has increased from 0.7 to 1.0 and the \( r^2 \) value has decreased from 92% to 55%. The reason for the exponent increase is that removing the instruments from the smaller missions lowered the OTA cost more than it lowered the OTA mass, which indicates that instruments have a higher cost per kilogram than OTAs. The reason that the \( r^2 \) decreased is because of the change in the Hubble OTA mass. Previously HST’s OTA mass was approximately the same as JWST’s OTA mass. But the new documentation yields an HST OTA mass that is 150% that of JWST’s OTA. Also, excluding CloudSat, Planck, and Herschel from the OTA mass regression has a negligible effect. Therefore we will not exclude them. Finally, while the mass and cost of the four attached OTAs were also reduced, they continued to lie on a cost
curve parallel to and below that of the free-flying model. Because we removed the gimbal cost and mass from SOFIA, it is no longer on the same line as the three missions that flew attached to the space-shuttle orbiter.

Based on a sample size of 14 free-flying space telescopes, a single-variable CER was developed for OTA cost as a function of OTA mass (Fig. 5):

\[
\text{OTAcost} \sim 0.17 \times \text{OTA mass}^{1.05} (N = 14; r^2 = 57\%; \text{SPE} = 65\%).
\]

Although the new mass model is less noisy than the new aperture model, it accounts for only 57% of the data variation.

A new graphical tool developed since the original paper is cost density (cost per kilogram). Figure 6 plots OTA cost per kilogram versus OTA aperture diameter for free-flying space, attached, and ground OTAs. Several obvious conclusions can be drawn. All free-flying space telescopes have approximately the same cost per kilogram, independent of aperture diameter. All ground telescopes also have approximately the same cost per kilogram, independent of aperture diameter. Space telescopes cost about 1000× to 30× more than ground telescopes. Additionally, UIT, WUPPE, and HUT, which flew attached to the space shuttle, are 5 to 10× less expensive per kilogram. SOFIA, which flies attached to a 747, is 20 to 30× less expensive. One explanation is that each of these mission types are built to different design rules. While all three types need similar wavefront shape and pointing stabilities as a function of aperture diameter, they have different static gravity and dynamic jitter environments. They also have different mass budgets with which to achieve the required wavefront shape and pointing stability. Free-flying telescopes have mass budgets that are severely constrained by the launch vehicle. Therefore, significant engineering cost is required to achieve the required performance for the allowed mass.
While the attached missions did fly on the space shuttle and SOFIA flies on a 747, the carrying capacities of these vehicles allows for different mass margin design rules.

Finally, as shown in Fig. 7, a regression of 26 free-flying missions yields a total mission cost versus total mission mass CER of

\[ \text{Total cost} \sim 0.5M \text{ diameter}^{0.9} \quad (N = 26; r^2 = 56\%; \text{SPE} = 59\%). \]

In this case, while the small SPE indicates a good regression, the model only explains 56% of the data variation. Similar to the aperture model, the total mission mass model exponent is slightly smaller than the OTA exponent. The implication is that as the OTA mass increases, the cost of the rest of the mission increases at a slightly slower rate.

4 Discussion on Mass

As with the original paper, mass appears to be an attractive CER for both OTA cost and total mission cost. however, it is the belief of these authors that in both cases mass is a surrogate for other engineering parameters. In the case of OTA, mass depends on aperture diameter. And, as indicated by Fig. 6, changing the design rule to allow for a larger mass budget might reduce total cost. Additionally, there are multiple cautionary indicators against a pure mass model. For
example, consider JWST and HST. Hubble is more massive
than JWST at both the OTA (150%) and mission (200%)
level (Figs. 5 and 7). But JWST is 2× more expensive
than HST at both the OTA and mission level. A pure
mass model would predict that JWST would be half as
expensive as HST—and it would be wrong. The reason is
that HST is more massive, stiffer, and less complex than
JWST. Next, consider the three shuttle-attached missions.
For the same aperture diameter, attached missions are
~3× more massive and ~10× less expensive. This is an in-
teresting comparison because the attached missions were
spaced out. Of course, there are multiple mitigating factors, such as
their limited design life, but given that most OTAs have
very few limited lifetime components, these authors believes
it to be a valid comparison. Finally, ground-based OTAs are
10 to 100× more massive than free-flying space OTAs and
1000×/kg less expensive.

5 Conclusions
Parametric cost models are an important tool routinely used
to plan missions, compare concepts, and justify technology
investments. In 2010, “Single variable parametric cost models
for space telescopes” presented new single-variable cost
models for space telescope OTAs. Its models were created by
applying standard statistical methods to data collected from
30 different space telescope missions. The results were com-
pared with previously published models. After its publica-
tion, we were invited to present our findings to the NRO.
The NRO Cost Model Office reviewed our database and,
while they did not provide us with access to their database
give us any specific data, they did identify specific mis-
sions where our two databases disagreed. In response, we
conducted a systematic review of all missions in our database
and reconciled them with source documents. This paper
updates and revises the findings of our 2010 paper. To
fully understand the content of this update paper, the authors
recommend that one also read the original 2010 paper.

Our review determined that we had been inconsistent in
the application of our definition as to what is and is not
included in an OTA. For all of our modeling, we define
an OTA to be the subsystem that collects electromagnetic
radiation and focuses it (focal) or concentrates it (afocal)
into the SIs. An OTA consists of the primary mirror, sec-
secondary mirror, auxiliary optics, and support structure (such as
optical bench or truss structure, primary support structure, secondary support structure or spiders, straylight baffles,
mechanisms for adjusting the optical components, and elec-
tronics or power systems for operating these mechanisms).
An OTA does not include SIs or spacecraft subsystems.
In the 2010 paper, the database OTA cost for large missions
(Hubble, Kepler, JWST, etc.) were correct. They were for
only the OTA. But the database cost for the smaller missions
(GALEX, IUE, TRACE, and WIRE) were incorrect. The
database costs of these missions were for both the OTA
and SI. Removing the SI cost to yield just an OTA cost dra-
matically reduced the database values for these small-aper-
ture missions. As a consequence, there have been changes to
the 2010 published results and models. But our two most
important findings remain unchanged.

First, from an engineering and scientific perspective,
aperture continues to be the best parameter with which to
build a space telescope cost model. Aperture defines the
observatory’s science performance (sensitivity and resolu-
tion) and determines the payload’s size and mass. Depending on which telescope OTAs are included or
excluded from the regression, the CER for an OTA as a function
of aperture diameter (inflated to 2011 values) is bounded
between

\[
\text{OTA cost} \sim \$30M \text{ diameter}^{1.4} \quad (N = 15; r^2 = 81\%; \text{SPE} = 122\%)
\]
Second, we continue to find that telescopes designed to a larger mass budget (for a given aperture), i.e., designed to be stiffer, have a lower cost. This finding is supported by comparisons of free-flying space telescopes, Hubble to JWST, and Hubble to SOFIA. It is also supported by analysis of areal cost (cost per kilogram), which indicates that different classes of telescopes are designed to different design rules. The reason for different classes is the launch vehicle. The total mission mass for a space telescope is constrained by the vehicle used to launch the mission. Additionally, there is a direct connection between the available mass budget and the engineering costs that must be expended to design a lightweight telescope of the desired aperture with the required wavefront shape and pointing stability for its operational environment. It is factual to assert that space telescopes are designed to mass.

Finally, one significant difference with the 2010 paper is that we now report telescope cost to vary linearly from 5% to 30% of total mission cost instead of the previously reported average of 20%. The larger the OTA, the greater percentage of the total mission cost it consumes.

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