Weak–strong uniqueness property for models of compressible viscous fluids near vacuum

Eduard Feireisl\textsuperscript{1,}\textsuperscript{,*} and Antonín Novotný\textsuperscript{2,}\textsuperscript{,**}

\textsuperscript{1} Institute of Mathematics of the Academy of Sciences of the Czech Republic, Žitná 25, CZ-115 67 Praha 1, Czech Republic
\textsuperscript{2} IMATH, EA 2134, Université de Toulon, BP 20132, 83957 La Garde, France

E-mail: feireisl@math.cas.cz and novotny@univ-tln.fr

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Abstract

We extend the weak–strong uniqueness principle to general models of compressible viscous fluids near/on the vacuum. In particular, the physically relevant case of positive density with polynomial decay at infinity is considered.

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1. Introduction

The motion of a compressible viscous fluid in the barotropic regime is described by the time evolution of the mass density $\rho = \rho(t, x)$ and the velocity field $u = u(t, x)$, $t > 0$, $x \in \Omega \subset \mathbb{R}^d$, $d = 1, 2, 3$, satisfying the equation of continuity

$$\partial_t \rho + \text{div}_x(\rho u) = 0, \quad (1.1)$$

and the momentum balance

$$\partial_t (\rho u) + \text{div}_x(\rho u \otimes u) + \nabla_x p(\rho) = \text{div}_x S + \rho f. \quad (1.2)$$

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\textsuperscript{**}Author to whom any correspondence should be addressed.

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Here, \( p = p(\rho) \) is the pressure, \( f \) the external driving force, and \( S \) the viscous stress. To close the system, we suppose that \( S \) is related to the symmetric part of the velocity gradient

\[
\mathbb{D}_x u \equiv \frac{1}{2} (\nabla_x u + \nabla_x ^t u)
\]

through a general rheological law

\[
F(\mathbb{D}_x u) + F^*(S) = S : \nabla_x u \iff S \in \partial F(\mathbb{D}_x u) \iff \mathbb{D}_x u \in \partial F^*(S), \tag{1.3}
\]

where

\[
F : R^{d \times d} \rightarrow [0, \infty] \text{ is a convex l.s.c. function.}
\]

The best known example is the isentropic Navier–Stokes system, where

\[
p(\rho) = a \rho^\gamma, \ a > 0, \ \gamma > 1, \ F = \mu \left| \nabla_x u + \nabla_x ^t u - \frac{2}{d} \text{div}_x u I \right|^2 + \lambda \left| \text{div}_x u \right|^2, \ \mu > 0, \ \lambda \geq 0. \tag{1.4}
\]

The ‘implicit’ rheological law (1.3) covers the class of power law fluids as well as other non-Newtonian fluids, see e.g. Bulíček et al [2].

Our goal is to study stability of strong solutions of the problem in the regime when the fluid density either vanishes or approaches asymptotically the vacuum state. To begin, we point out that the model has been derived for non-dilute fluids out of vacuum. In particular, as strong solutions satisfy the equation of continuity

\[
\partial_t \rho + u \cdot \nabla_x \rho + \rho \text{div}_x u = 0,
\]

the density will always remain positive unless the vacuum state is artificially imposed through the initial or boundary data. However, allowing \( \rho \approx 0 \) is still physically relevant at least in certain asymptotic regimes:

- Solutions of (1.1)–(1.3) with \( f = \nabla_x G, \ G = G(x) \) approach an equilibrium state \( \rho = \tilde{\rho}, \ u = 0 \) as \( t \to \infty \), see [7], where

\[
\nabla_x p(\tilde{\rho}) = \tilde{\rho} \nabla_x G. \tag{1.5}
\]

It is a simple observation that \( \tilde{\rho} \) may vanish on a non-void subset of the physical domain, in particular if the total mass of the fluid is small enough.

- In models of gaseous stars in astrophysics, the physical domain \( \Omega \) is exterior to a rigid object, while

\[
\rho \to 0, \quad u \to 0 \quad \text{as} \quad |x| \to \infty. \tag{1.6}
\]

Thus the density is close to zero at least in the far field. Note that, on the one hand, the isentropic Navier–Stokes system (1.1), (1.2) and (1.4) is globally well posed in this regime if the total energy is small enough, see Huang et al [9], and also Fang et al [5] for a similar result in the non-Newtonian case and \( d = 1 \). On the other hand, Rozanova [15] and Xin [17] showed that fast decay of solutions as \( |x| \to \infty \) is not compatible with global existence. More recently, Merle et al [14] obtained blow-up results in a similar regime for radially symmetric solutions with certain profiles and certain values of the adiabatic exponent \( \gamma \).
• Last but not least, the problem with vanishing initial density is mathematically challenging and has been considered in a large number of recent studies, see Gong et al [8], Li and Xin [11], Liang [12], to name only a few.

Our aim is to clarify in which way the strong solutions obtained in the above references can coexist/coincide with the weak solutions in the sense of Lions [13] or even more general dissipative solutions introduced in [1]. More specifically, we establish the weak–strong uniqueness principle: a weak solution of the problem (1.1)–(1.3) coincides with the strong one with the same initial/boundary data as long as the strong solution exists. As shown in [1, theorem 6.3], the weak–strong uniqueness principle holds for the problem (1.1)–(1.3) in the class of strong solutions away from vacuum, and for $\Omega \subset \mathbb{R}^d$ a bounded domain with general in/out flow boundary conditions:

$$u|_{\partial \Omega} = u_B, \quad \varrho|_{\Gamma_{\text{in}}} = \varrho_B, \quad \Gamma_{\text{in}} = \{x \in \partial \Omega | u_B \cdot n \leq 0\}, \quad (1.7)$$

where $n$ is the outer normal vector to $\partial \Omega$. We show that this result can be extended to the class of strong solutions near/with the vacuum as long as:

(a) The potential $F$ is uniformly strictly convex in $\mathbb{R}^{d \times d}_{\text{sym}}$.
(b) The pressure $p$ is strictly convex in $[0, \infty)_\text{loc}$.
(c) The density component $\tilde{\varrho}$ of the strong solution satisfies

$$\nabla_{x}P'(\tilde{\varrho}) \in L^2(0, T; L^q(\Omega)), \quad q = q(d),$$

where $P$ is the pressure potential,

$$P'(\varrho)\varrho - P(\varrho) = p(\varrho).$$

(d) If $\Omega$ is unbounded, the initial density must decay sufficiently fast as $|x| \to \infty$.

Conditions (a) and (b) are constitutive. Condition (a) holds in particular for the Newtonian stress in the Navier–Stokes system. Condition (b) is satisfied for $p = a\varrho^\gamma$ as long as $\gamma > 1$, which is physically relevant in the isentropic regime. Condition (c) is associated to the flow. If the velocity field $\tilde{u}$ is smooth enough, the Sobolev regularity of $P'(\tilde{\varrho})$ is propagated along streamlines. It is therefore enough to impose 3 on the data only. We point out that

$$\nabla_{x}P'(\tilde{\varrho}) \approx \tilde{\varrho}^{\gamma-2} \nabla_{x}\tilde{\varrho}$$

becomes singular near the vacuum if $\gamma < 2$. Condition (d) is relevant if the physical domain is unbounded. Here again, the decay may be inherited from the initial/boundary data if the velocity field is sufficiently regular.

Our method is based on a general relative energy inequality obtained in [1] that must be adapted to the vanishing density regime. The most delicate issue is the behavior of the fluid velocity on or near the vacuum region. As a matter of fact, the concept of velocity in the absence of matter is meaningless, however, the velocity can be recovered as a solution of an elliptic equation in the whole physical space. This paradox can be seen as a consequence of the infinite speed of propagation due to the viscous stress with constant viscosity coefficients. By the same token, the decay of the velocity for $|x| \to \infty$ is not expected to be faster than that of the Dirichlet kernel at least in the linearly viscous case. This fact makes the analysis on unbounded domains rather delicate and requires certain decay properties of the initial density profile. Seen from this perspective, the hypothesis $u(t, \cdot) \in W^{1,2}$ assumed sometimes in the literature is not very realistic.
The paper is organised as follows. We begin by recalling the basic concepts of strong as well as weak (dissipative) solution in section 2. In section 3, we introduce the relative energy functional. The main results are stated and proved in section 4. We distinguish three cases: bounded domain, general unbounded domain, and the whole space $\mathbb{R}^3$ with compactly supported density or rapidly decaying density.

2. Strong and weak solutions

We recall the definition of strong and weak (dissipative) solutions to the problem (1.1)–(1.3), with the boundary conditions (1.7), and with the far field conditions (1.6) if the domain $\Omega$ is unbounded. For the sake of simplicity, we assume that the boundary data $u_B, \varrho_B$ are independent of $t$. First, extend the function $u_B$ to be smooth in $\Omega$ and

$$u_B(x) = 0 \quad \text{for} \quad |x| > R \quad (2.1)$$

if $\Omega$ is unbounded. To avoid technical difficulties, we suppose that the outer normal is defined whenever $u_B \cdot \mathbf{n} \neq 0$. Moreover, we consider the pressure in the isentropic form

$$p(\varrho) = a\varrho^\gamma, \quad a > 0, \quad \gamma > 1. \quad (2.2)$$

Finally, we impose certain growth restrictions on the dissipation potential $F$,

$$0 \leq F(\mathbb{D}) \lesssim |\mathbb{D}|^2, \quad (2.3)$$

$$F(\mathbb{D} + Q) - F(\mathbb{D}) - S : Q \lesssim |Q - \beta \text{tr}(Q)|^2, \quad \beta = \frac{1}{d} \quad \text{if} \quad d = 2, 3, \quad \beta = 0 \quad \text{if} \quad d = 1, \quad (2.4)$$

for any $\mathbb{D}, Q \in \mathbb{R}^{d \times d}_{\text{sym}}, S \in \partial F(\mathbb{D})$. Note that both (2.3) and (2.4) are compatible with the Newtonian potential (1.4). The condition (2.3) can be relaxed at the expense of several technical difficulties in the proofs. The coercivity assumption (2.4), reflecting non-degeneracy of the viscous stress, is essential.

2.1. Strong solutions

The goal is to identify the largest possible class of strong solutions. Inspired by the existence result by Cho et al [3] we consider the velocity $\tilde{\mathbf{u}}$ in the class

$$(\tilde{\mathbf{u}} - \mathbf{u}_B) \in C([0, T]; D^{0,2}_0(\Omega; \mathbb{R}^d) \cap L^2(0, T; D^{2,\eta}(\Omega; \mathbb{R}^d))),$$

$$\partial_t \tilde{\mathbf{u}} \in L^2(0, T; L^4(\Omega; \mathbb{R}^d)) \quad (2.5)$$

for some $q > d$. Here, $D^{k,q}$ denotes the space of locally integrable functions $v$ with $D^{k,q}_0 v \in L^q(\Omega), D^{k,q}_0$ is the closure of $C_0^\infty(\Omega)$ in $D^{k,q}(\Omega)$. If $\Omega$ is bounded, $D^{k,q}$ coincide with the standard Sobolev spaces $W^{k,q}$. Unbounded domains will be considered only for $d = 3$. Note that the class (2.5) includes the case $q = 6, d = 3$ considered in [3]. In particular, by interpolation and the standard embedding $D^{k,q} \hookrightarrow C, q > d$,

$$\tilde{\mathbf{u}} \in C([0, T]; D^{1,q} \cap D^{0,2}_0(\Omega; \mathbb{R}^d)) \hookrightarrow BC([0, T] \times \overline{\Omega}; \mathbb{R}^d).$$
Similarly, we deduce from (2.5)
\[ \int_0^T \| \tilde{\mathbf{u}} \|^2_{W^{1,\infty}(\Omega; \mathbb{R}^d)} \, dt \lesssim 1 \]  
(2.6)
if either \( \Omega \) is bounded or \( d = 3 \).

In particular, given the velocity field \( \tilde{\mathbf{u}} \), the initial state and the boundary data
\[ \tilde{\varrho}(0, \cdot) = \varrho_0 \in L^1_{\text{loc}}(\Omega), \quad \tilde{\varrho}|_{\Gamma_{\text{in}}} = \varrho_B \in L^1(\partial\Omega), \]
the density \( \tilde{\varrho} \) is uniquely determined by the equation of continuity (1.1) and can be computed explicitly via the method of characteristics. More specifically, supposing \( \Omega \) has a compact Lipschitz boundary, we may extend \( \tilde{\mathbf{u}} \) outside \( \Omega \) to remain in the class
\[ \tilde{\mathbf{u}} \in L^2(0, T; D^{2,q}(\Omega; \mathbb{R}^d)) \cap BC([0, T] \times \mathbb{R}^d; \mathbb{R}^d). \]

Then we define the characteristic curves
\[ \frac{d}{dt} X(t, x) = \tilde{\mathbf{u}}(t, X(t, x)), \quad X(0, x) = x. \]
For any \( t > 0, x \in \Omega \), there is a unique characteristic satisfying
\[ X(t, x_0) = x. \]
Let \( \tau = \inf \{ s \in [0, t] \mid X(z, x_0) \in \Omega \text{ for all } z \in (s, t) \} \). If \( \tau = 0 \), we set
\[ \tilde{\varrho}(t, x) = \varrho_0(x_0) \exp \left( - \int_0^t \text{div}_x \tilde{\mathbf{u}}(s, X(s, x_0)) \, ds \right). \]  
(2.7)
If \( \tau > 0 \), then necessarily
\[ x_\tau = X(\tau, x_0) \in \Gamma_{\text{in}}, \]
and we set
\[ \tilde{\varrho}(t, x) = \varrho_B(x_\tau) \exp \left( - \int_\tau^t \text{div}_x \tilde{\mathbf{u}}(s, X(s, x_0)) \, ds \right). \]  
(2.8)

To avoid jumps in the resulting density, the above procedure requires an obvious compatibility condition
\[ \varrho_0 \in BC(\overline{\Omega}), \quad \varrho_0|_{\Gamma_{\text{in}}} = \varrho_B. \]

Still the construction via characteristic lines may give rise to ambiguous results as soon as part of the boundary is tangent to the characteristic curves. To avoid this problem, several non-degeneracy conditions on \( \Gamma_{\text{in}} \) are usually imposed in the literature to eliminate ingoing characteristics emanating from \( \partial\Gamma_{\text{in}} \), cf e.g. Valli and Zajaczkowski [16]. The density \( \tilde{\varrho} \in BC([0, T] \times \overline{\Omega}) \) associated to \( \tilde{\mathbf{u}} \) is then well defined through formulae (2.7), (2.8). As a direct consequence, we obtain the following standard result concerning propagation of Sobolev regularity. To avoid the aforementioned technical problems and also because the available regularity of \( \partial \tilde{\mathbf{u}} \) is rather limited, we restrict ourselves to the case \( \Gamma_{\text{in}} = \emptyset \).

**Lemma 2.1** Let \( \Omega \subset \mathbb{R}^d \) be a domain with compact (possibly empty) Lipschitz boundary. Let a vector field \( \tilde{\mathbf{u}} \) be given in the class
\[ \tilde{\mathbf{u}} \in L^2(0, T; D^{2,q}(\Omega; \mathbb{R}^d)) \cap BC([0, T] \times \overline{\Omega}; \mathbb{R}^d), \quad d < q \leq \infty \]
and such that $\tilde{u} \cdot n|_{\partial \Omega} \geq 0$. Suppose that

$$\varrho_0 \in BC(\Omega), \quad 0 \geq \varrho_0, \quad \nabla_x (\varrho_0^{-1}) \in L^\gamma(\Omega; R^d) \quad \text{for some } 1 < \gamma \leq 2.$$  

Then the unique solution $\tilde{\varrho}$ of the equation of continuity (1.1) determined by (2.7) satisfies

$$\|\nabla_x (\tilde{\varrho}^{-1})(t, \cdot)\|_{L^q(\Omega; R^d)} \leq c(T, \varrho_0, \tilde{\varrho}) \quad \text{for any } t \in [0, T].$$

**Proof.** Extending $\tilde{u}$ and $\varrho_0$ outside $\Omega$ as the case may be, we may assume $\Omega = R^d$. As $\Gamma_{in} = \emptyset$, the solution $\tilde{\varrho}$ is given by formula (2.7). In particular,

$$\tilde{\varrho}(t, x) = \varrho_0(X^{-1}(t, x)) \exp \left( - \int_0^t \text{div}_x \tilde{u}(s, X(s, X^{-1}(t, x))) ds \right),$$

$$\tilde{\varrho}^{-1}(t, x) = \varrho_0^{-1}(X^{-1}(t, x)) \exp \left( (1 - \gamma) \int_0^t \text{div}_x \tilde{u}(s, X(s, X^{-1}(t, x))) ds \right).$$

Thus the desired result follows by direct differentiation of the above formula. $\Box$

If the exponent $\gamma$ remains in the range $1 < \gamma \leq 2$, differentiability of $\tilde{\varrho}^{-1}$ implies differentiability of $\tilde{\varrho}$. In particular, the equation of continuity (1.1) can be interpreted in the strong sense. Similarly, the left-hand side of the momentum equation can be written in the form

$$\tilde{\varrho} \partial_t \tilde{u} + \tilde{\varrho} u \cdot \nabla_x \tilde{u} + \tilde{\varrho} \nabla_x \tilde{\varrho}^{-1}$$

Thus for $\tilde{u}$ belonging to the class (2.5), the system (1.1)–(1.3) can be written as

$$\begin{align*}
\partial_t \tilde{\varrho} + \tilde{u} \cdot \nabla_x \tilde{\varrho} + \tilde{\varrho} \text{div}_x \tilde{u} &= 0, \\
\tilde{\varrho} \partial_t \tilde{u} + \tilde{\varrho} u \cdot \nabla_x \tilde{u} + \tilde{\varrho} \nabla_x \tilde{\varrho}^{-1} &= \text{div}_x \tilde{S} + \tilde{\varrho} \mathcal{R},
\end{align*}$$

with

$$\tilde{S} \in \partial F(\mathcal{D}_x \tilde{u}) \quad \text{for a.a. } t \in (0, T).$$

**2.2. Dissipative (weak) solutions**

The dissipative solutions to the system (1.1)–(1.3) were introduced in [1]. We denote $\mathcal{M}^+(\overline{\Omega}; R^{d \times d}_{\text{sym}})$ the set of positively semi-definite finite tensorial measures on $\overline{\Omega}$. Specifically, $R \in \mathcal{M}^+(\overline{\Omega}; R^{d \times d}_{\text{sym}})$ if $R$ is a tensor valued finite measure satisfying

$$R : (\xi \otimes \xi) \geq 0 \quad \text{for any } \xi \in R^d.$$

**Definition 2.2. (Dissipative solution).** The functions $(\varrho, u)$ represent dissipative solution of the problem (1.1)–(1.3), with the boundary conditions (1.7), the far field conditions (1.6), and the initial conditions

$$\varrho(0, \cdot) = \varrho_0, \quad (\varrho u)(0, \cdot) = m_0,$$

if the following holds:


- **Integrability.**

  \[ \varrho \geq 0, \varrho \in C_{\text{weak}}([0, T]; L^r(\Omega)) \cap L^r((0, T); L^r(\partial \Omega, |u_B \cdot n| d\sigma)), \]

  \[ (u - u_B) \in L^2(0, T; D_0^1(\Omega; R^d)), \varrho u \in C_{\text{weak}} \left([0, T]; L^{\frac{2r}{r+1}}(\Omega; R^d)\right), \]

  \[ S \in L^2(0, T; L^2(\Omega; R^{d \times d})). \]

- **Equation of continuity.** The integral identity

  \[
  \int_\Omega \varrho \cdot \varphi \, dx \bigg|_{t=0}^{t=T} = \int_0^T \int_\Omega \varrho \left( \frac{\partial u}{\partial t} + u \cdot \nabla u + \varrho \nabla \varphi \right) + \int_\Omega \varrho \left( \psi \varphi + u \cdot \nabla \varphi \right) \, dx \, dt
  \]

  \[
  + \int_\Omega \int_{\partial \Omega} \varrho \left( \frac{\partial u}{\partial n} + u \cdot n \right) \cdot \nabla \varphi + u \cdot \nabla \varphi \cdot n \right) \, dS_t - \int_0^T \int_{\partial \Omega} \varrho u \cdot \nabla \varphi + u \cdot \nabla \varphi \cdot n \right) \, dS_t \, dt,
  \]

  \[ \varphi(0, \cdot) = \varrho_0 \]

  holds for any \( 0 \leq \tau \leq T \), and any test function \( \varphi \in C^1_c([0, T] \times \overline{\Omega}) \).

- **Momentum equation.** There exists

  \[ \mathcal{R} \in L^\infty_{\text{weak}(\cdot)}(0, T; \mathcal{M}^+(\overline{\Omega}, R^{d \times d})) \]

  such that

  \[
  \int_\Omega \varrho \cdot \varphi \, dx \bigg|_{t=0}^{t=T} = \int_0^T \int_\Omega \varrho \left( \frac{\partial u}{\partial t} + u \cdot \nabla u + \varrho \nabla \varphi + p(\varrho) \text{div} u \varphi \right)
  \]

  \[
  - \varrho \partial_t \varphi + u \cdot \nabla \varphi \right) \, dx \, dt
  \]

  \[
  + \int_0^T \int_{\partial \Omega} \varrho \left( \frac{\partial u}{\partial n} + u \cdot n \right) \cdot \nabla \varphi + u \cdot \nabla \varphi \cdot n \right) \, dS_t \, dt, \varrho(0, \cdot) = m_0
  \]

  holds for any \( 0 \leq \tau \leq T \) and any test function \( \varphi \in C^1_c([0, T] \times \overline{\Omega}; R^d), \varphi|_{\partial \Omega} = 0 \).

- **Energy inequality.** There exists

  \[ \mathcal{E} \in L^\infty_{\text{weak}(\cdot)}(0, T; \mathcal{M}^+(\overline{\Omega})) \]

  such that

  \[
  \int_\Omega \left[ \frac{1}{2} \varrho |u - u_B|^2 + P(\varrho) \right] \, dx \bigg|_{t=0}^{t=T} \int_0^T \int_\Omega \left[ F(\varrho u) + \gamma \mathcal{E} \right] \, dx \, dt
  \]

  \[
  + \int_0^T \int_{\partial \Omega} P(\varrho)|u_B - u_B|^2 \, dS_t \, dt
  \]

  \[
  + \int_0^T \int_{\partial \Omega} P(\varrho)\varrho |u_B \cdot n|^2 \, dS_t \, dt + \int_{\partial \Omega} \frac{d\mathcal{E}(\tau)}{d\tau} \]

  \[
  \leq - \frac{1}{2} \int_0^T \int_{\Omega} \left( \varrho u \cdot \nabla u + p(\varrho) I \right) : \nabla u u_B \, dx \, dt
  \]

  \[
  + \int_0^T \left( \frac{1}{2} \int_\Omega \varrho |u - u_B|^2 \right) \, dx \, dt
  \]

  \[
  + \int_0^T \int_{\partial \Omega} \left[ S : \nabla u_B + \varrho f \cdot u \right] \, dx \, dt
  \]

  \[
  - \frac{1}{2} \int_{\Omega} \nabla u_B : d\mathcal{E}(t) \, dt, \quad P(\varrho) \equiv \frac{a}{\gamma - 1} \varrho^{\gamma} \quad (2.13)
  \]
for a.a. $0 \leq \tau \leq T$.

- **Defect compatibility**.
  \[ d\mathcal{E} \leq \text{tr}[\mathcal{R}] \leq d\mathcal{E}, \quad \text{for certain constants } 0 < d < d. \tag{2.14} \]

**Remark 2.3.** The hypothesis
\[ S \in L^2(0, T; L^2(\Omega; \mathbb{R}^{d \times d})) \]
is pertinent to the class of the dissipative potentials $F$ with at most quadratic growth (2.3).

**Remark 2.4.** The dissipative solutions introduced above apparently represent a larger class of generalised weak solutions than the conventional finite energy weak solutions, for which $\mathcal{R} = 0$ and the dissipation in the energy inequality is given by the conventional term $S : D_x u$.

### 3. Relative energy

The relative energy reads:
\[ E(\rho, u | \bar{\rho}, \bar{u}) = \frac{1}{2} \rho |u - \bar{u}|^2 + P(\rho) - P'(\bar{\rho})(\rho - \bar{\rho}) - P(\bar{\rho}). \]

The relative energy inequality associated to the problem (1.1)–(1.3) and (1.7) was derived in [1, section 5, formula (5.5)]. Its generalisation to the case of unbounded domain is straightforward (cf also [6]):

\[
\left[ \int_{\Omega} E(\rho, u | \bar{\rho}, \bar{u}) \, dx \right]_{\tau=0}^{\tau} + \int_0^\tau \int_{\Omega} \left[ F(D_x u) + F'(S) \right] \, dx \, dt - \int_0^\tau \int_{\Omega} \mathcal{S} : \nabla_x \bar{u} \, dx \, dt \\
+ \int_0^\tau \int_{\partial \Omega} \left[ P(\rho) - P'(\bar{\rho})(\rho - \bar{\rho}) - P(\bar{\rho}) \right] [u_B \cdot \mathbf{n}]^+ \, dS_x \, dt \\
+ \int_0^\tau \int_{\Omega} \left[ P(\rho_B) - P'(\bar{\rho})(\rho - \bar{\rho}) - P(\bar{\rho}) \right] [u_B \cdot \mathbf{n}]^- \, dS_x \, dt + \int_1 d\mathcal{E}(\tau) \\
\leq - \int_0^{\tau} \int_{\Omega} \rho(\bar{u} - u) \cdot (\bar{u} - u) \cdot \nabla_x \bar{u} \, dx \, dt \\
- \int_0^{\tau} \int_{\Omega} \left[ p(\rho) - p'(\bar{\rho})(\rho - \bar{\rho}) - p(\bar{\rho}) \right] \text{div}_x \bar{u} \, dx \, dt \\
+ \int_0^{\tau} \int_{\Omega} \rho(\bar{u} - u) \cdot \left[ \partial_t \bar{u} + \bar{u} \cdot \nabla_x \bar{u} + \nabla x P'(\bar{\rho}) \right] \, dx \, dt \\
+ \int_0^{\tau} \int_{\Omega} \rho(\bar{u} - u) \cdot f \, dx \, dt \\
+ \int_0^{\tau} \int_{\Omega} p'(\bar{\rho}) \left( 1 - \frac{\rho}{\bar{\rho}} \right) \left[ \partial_t \bar{\rho} + \text{div}_x(\bar{\rho} \bar{u}) \right] \, dx \, dt \\
- \int_0^{\tau} \int_{\Omega} \nabla_x \bar{u} : d\mathcal{R}(\tau) \, dt \quad \text{for a.a. } \tau \in (0, T), \tag{3.1}
\]
for a.a. $0 < \tau < T$, and any pair of ‘test functions’

$$\tilde{u} \in C^1([0,T] \times \overline{\Omega}; \mathbb{R}^d), \tilde{u}|_{\partial \Omega} = u_B,$$

$$\tilde{\rho} \in C^1([0,T] \times \overline{\Omega}), \quad \tilde{\rho} > 0, \quad (\tilde{\rho} - \varepsilon) \in C_c([0,T] \times \overline{\Omega})$$

for some $\varepsilon > 0$.

(3.2)

Our goal is to extend validity of (3.1) to the class of test functions containing $(\tilde{\rho} + \varepsilon, \tilde{u})$, where $(\tilde{\rho}, \tilde{u})$ is a strong solution. Using the standard regularisation by a convolution kernel in time, we first observe that the class of test functions can be extended to

$$\tilde{u} \in C_c([0,T] \times \overline{\Omega}; \mathbb{R}^d), \quad \tilde{u}|_{\partial \Omega} = u_B,$$

$$\partial_t \tilde{u} \in L^2(0,T; C_c(\overline{\Omega}; \mathbb{R}^d))$$

$$(\tilde{\rho} - \varepsilon) \in C_c([0,T] \times \overline{\Omega}), \quad \tilde{\rho} > 0, \quad \partial_t \tilde{\rho} \in L^\infty_{\text{weak-*(T)}}(0,T; C_c(\overline{\Omega}; \mathbb{R}^d))$$

(3.3)

for some $\varepsilon > 0$.

The class (3.3) covers all interesting cases if $\Omega$ is bounded. If $\Omega$ is an exterior domain or $\Omega = \mathbb{R}^3$, integrability for $|x| \to \infty$ is relevant. In this case, we consider $d = 3$ only, where we consider the class:

$$\tilde{u} \in BC([0,T] \times \overline{\Omega}; \mathbb{R}^3), \quad \tilde{u}|_{\partial \Omega} = u_B,$$

$$\nabla_x \tilde{u} \in L^2(0,T; BC \cap L^2(\overline{\Omega}; \mathbb{R}^d)),$$

$$\partial_t \tilde{u} \in L^2(0,T; BC \cap L^6(\overline{\Omega}; \mathbb{R}))$$

$$\nabla_x \tilde{\rho} \in L^\infty_{\text{weak-*(T)}}(0,T; BC \cap L^6(\overline{\Omega}; \mathbb{R}^3))$$

(3.4)

The extension to this class can be justified by the standard cut-off procedure as long as all integrals in (3.1) are finite. Extensions to other classes of test functions can be obtained in a similar manner as the case may be.

### 4. Weak strong uniqueness

We are ready to establish the main results. We distinguish the cases of $\Omega$ bounded, exterior to a bounded set, or $\Omega = \mathbb{R}^3$. For the sake of simplicity, we suppose here and hereafter that

$$f = 0.$$

#### 4.1. Weak–strong uniqueness–bounded domain

We start with our main result concerning bounded domains, where inflow is absent.

**Theorem 4.1 (Weak–strong uniqueness, bounded domain, no inflow).** Suppose that $\Omega \subset \mathbb{R}^d$, $d = 1, 2, 3$ is a bounded Lipschitz domain. Let the boundary velocity $u_B$ be a twice continuously differentiable function satisfying

$$u_B \cdot n \geq 0 \quad \text{on} \quad \partial \Omega.$$

The class (3.3) covers all interesting cases if $\Omega$ is bounded. If $\Omega$ is an exterior domain or $\Omega = \mathbb{R}^3$, integrability for $|x| \to \infty$ is relevant. In this case, we consider $d = 3$ only, where we consider the class:

$$\tilde{u} \in BC([0,T] \times \overline{\Omega}; \mathbb{R}^3), \quad \tilde{u}|_{\partial \Omega} = u_B,$$

$$\nabla_x \tilde{u} \in L^2(0,T; BC \cap L^2(\overline{\Omega}; \mathbb{R}^d)),$$

$$\partial_t \tilde{u} \in L^2(0,T; BC \cap L^6(\overline{\Omega}; \mathbb{R}))$$

$$\nabla_x \tilde{\rho} \in L^\infty_{\text{weak-*(T)}}(0,T; BC \cap L^6(\overline{\Omega}; \mathbb{R}^3))$$

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### 4. Weak strong uniqueness

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#### 4.1. Weak–strong uniqueness–bounded domain

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$$u_B \cdot n \geq 0 \quad \text{on} \quad \partial \Omega.$$
Let \((\tilde{\rho}, \tilde{u})\) be a strong solution of the problem (1.1), (1.2) and (1.7) in \((0, T) \times \Omega\) in the sense specified in Section 2.1 (see (2.9) and (2.10)) belonging to the class
\[
\tilde{u} \in C([0, T] \times \overline{\Omega}; R^d), \quad \tilde{\rho} \in C([0, T] \times \Omega), \quad \tilde{\rho} \geq 0,
\]
\[
\nabla^2 \tilde{u} \in L^2(0, T; L^p(\Omega; R^{d \times d})), \quad \partial_t \tilde{u} \in L^2(0, T; L^q(\Omega; R^d)),
\]
where
\[
1 < \gamma \leq 2, \quad q \geq \frac{2\gamma}{\gamma - 1}.
\]

Let \((\rho, u)\) be a weak solution of the same problem in the sense of definition 2.2 such that
\[
\rho(0, \cdot) = \tilde{\rho}(0, \cdot) = \rho_0, \quad (\rho u)(0, \cdot) = (\tilde{\rho} \tilde{u})(0, \cdot) = m_0, \quad u|_{\partial \Omega} = \tilde{u}|_{\partial \Omega} = u_0,
\]
where
\[
(\rho_0)^{-1} \in W^{1,q}(\Omega; R^d).
\]
Then
\[
\rho = \tilde{\rho}, \quad u = \tilde{u} \quad \text{in} \quad (0, T) \times \Omega.
\]

Remark 4.2. In the Newtonian case, the local in time existence of strong solutions with vacuum in the case \(u_0 = 0\) (no in/outflow) on smooth bounded domains, say \(\Omega \in C^3\), with the pressure law (2.2) was proved in Cho and Kim [4, theorem 3]. The solutions belong to a regularity class included in (2.4) if \(q = 6\). Local existence for general non-Newtonian viscous stress was recently established by Kalousek et al [10] on periodic spatial domains.

Proof. The strategy is to apply the relative energy inequality (3.1) to \((\tilde{\rho}, \tilde{u})\). This cannot be done in a direct manner as the latter requires \(\inf \tilde{\rho} > 0\). Thus we start with the choice \((\tilde{\rho} + \varepsilon, \tilde{u})\), \(\varepsilon > 0\), obtaining
\[
\int_{\Omega} E(\rho, u | \tilde{\rho} + \varepsilon, \tilde{u})(\tau, \cdot) \, dx + \int_{\Omega} 1 \, d\varepsilon(\tau)
\]
\[
+ \int_{0}^{T} \int_{\Omega} \left[ F(\nabla_s u) + F^*(S) \right] \, dx \, dt - \int_{0}^{T} \int_{\Omega} S : \nabla \tilde{u} \, dx \, dt
\]
\[
\leq \int_{\Omega} P(\rho_0) + \varepsilon P'(\rho_0) - P(\rho_0 + \varepsilon) \, dx
\]
\[
- \int_{0}^{T} \int_{\Omega} \rho(\tilde{u} - u) \cdot (\tilde{u} - u) \cdot \nabla \tilde{u} \, dx \, dt
\]
\[
- \int_{0}^{T} \int_{\Omega} \left[ p(\rho) - p'(\tilde{\rho} + \varepsilon)(\rho - \tilde{\rho} - \varepsilon) - p(\tilde{\rho} + \varepsilon) \right] \, \text{div}_s \tilde{u} \, dx \, dt
\]
\[
+ \int_{0}^{T} \int_{\Omega} \rho(\tilde{u} - u) \cdot \left[ \partial_t \tilde{u} + \tilde{u} \cdot \nabla \tilde{u} + \nabla \xi P'(\tilde{\rho} + \varepsilon) \right] \, dx \, dt
\]
\[
+ \varepsilon \int_{0}^{T} \int_{\Omega} p'(\tilde{\rho} + \varepsilon) \left( 1 - \frac{\rho}{\tilde{\rho} + \varepsilon} \right) \, \text{div}_s \tilde{u} \, dx \, dt
\]
\[
- \int_{0}^{T} \int_{\Omega} \nabla \tilde{u} : d\mathcal{R}(\tau) \, dr \quad \text{for a.a.} \quad \tau \in (0, T),
\]
Moreover, as \( \tilde{u} \) belongs to the class (4.1) and \( q \geq 4 \), we get \( \| \nabla_x \tilde{u} \|_{L^{\infty}(\Omega; R^d)} \in L^2(0, T) \), and the above inequality reduces to

\[
\int_{\Omega} E \left( \varrho, u | \tilde{\varrho} + \varepsilon, \tilde{u} \right) (\tau, \cdot) \, dx + \int_{\Omega} \rho E(\rho) \, dx - \int_{\Omega} \rho E(\rho) \, dx \\
+ \int_{0}^{\tau} \int_{\Omega} \left[ F(\|D_xu\|) + F'(S) \right] \, dx \, dt - \int_{0}^{\tau} \int_{\Omega} S : \nabla_x \tilde{u} \, dx \, dt \\
\leq \int_{\Omega} \rho(\varrho) + \varepsilon \rho'(\varrho) - P(\varrho_0 + \varepsilon) \, dx \\
+ \int_{0}^{\tau} \chi(t) \left[ \int_{\Omega} E \left( \varrho, u | \tilde{\varrho} + \varepsilon, \tilde{u} \right) (t, \cdot) \, dx + \int_{\Omega} 1 \, dx \right] \, dt \\
+ \int_{0}^{\tau} \int_{\Omega} \rho(\tilde{u} - u) \cdot \left[ \tilde{\varrho} \tilde{u} + \tilde{u} \cdot \nabla_x \tilde{u} + \nabla_x \rho' \tilde{\varrho} \right] \, dx \, dt \\
+ \varepsilon \int_{0}^{\tau} \int_{\Omega} \rho'(\tilde{\varrho} + \varepsilon) \left( 1 - \frac{\varrho}{\tilde{\varrho} + \varepsilon} \right) \, dx \, dt, \quad \text{with } \chi \in L^2(0, T). \tag{4.4}
\]

Next, we let \( \varepsilon \to 0 \) in (4.4). Obviously,

\[
\int_{\Omega} E \left( \varrho, u | \tilde{\varrho} + \varepsilon, \tilde{u} \right) (\tau, \cdot) \, dx \to \int_{\Omega} E \left( \varrho, u | \tilde{\varrho}, \tilde{u} \right) (\tau, \cdot) \, dx \quad \text{as } \varepsilon \to 0 \text{ uniformly in } \tau \in (0, T),
\]

\[
\int_{\Omega} \rho(\varrho) + \varepsilon \rho'(\varrho) - P(\varrho_0 + \varepsilon) \, dx \to 0 \quad \text{as } \varepsilon \to 0,
\]

and

\[
\varepsilon \int_{0}^{\tau} \int_{\Omega} \rho'(\tilde{\varrho} + \varepsilon) \, dx \, dt \to 0 \quad \text{as } \varepsilon \to 0.
\]

In addition,

\[
\varepsilon \rho \frac{\rho'(\tilde{\varrho} + \varepsilon)}{\tilde{\varrho} + \varepsilon} \lesssim \varepsilon \rho(\tilde{\varrho} + \varepsilon)^{\gamma - 2} \lesssim \varepsilon^{\gamma - 1} \tilde{\varrho} \to 0 \text{ in } L^{\gamma}(\Omega) \quad \text{as } \varepsilon \to 0.
\]

Finally, by virtue of lemma 2.1 and hypotheses (4.1) and (4.3), we have

\[
\nabla_x (\tilde{\varrho})^{\gamma - 1} \in L^\infty(0, T; L^2(\Omega; R^d)). \tag{4.5}
\]

Seeing that

\[
\varrho u \in L^\infty \left( 0, T; \overline{L^{\frac{2\gamma}{\gamma - 1}}(\Omega; R^d)} \right)
\]

we may use hypothesis (4.2) to conclude

\[
\int_{0}^{\tau} \int_{\Omega} \varrho(\tilde{u} - u) \cdot \nabla_x P'(\tilde{\varrho} + \varepsilon) \, dx \, dt \to \int_{0}^{\tau} \int_{\Omega} \varrho(\tilde{u} - u) \cdot \nabla_x P'(\tilde{\varrho}) \, dx \, dt \quad \text{as } \varepsilon \to 0.
\]
Performing the limit $\varepsilon \to 0$ in (4.4) we may infer that

$$
\int_{\Omega} E \left( \varrho, u \mid \tilde{\varrho}, \tilde{u} \right) (\tau, \cdot) \, dx + \int_{\Omega} \theta \, dx 
+ \int_{0}^{\tau} \int_{\Omega} \left[ F(D_{u} u) + F'(S) \right] \, dx \, dt
- \int_{0}^{\tau} \int_{\Omega} \nabla_{x} \tilde{u} \, dx \, dt
\leq \int_{0}^{\tau} \chi(t) \left[ \int_{\Omega} E \left( \varrho, u \mid \tilde{\varrho}, \tilde{u} \right) (t, \cdot) \, dx + \int_{\Omega} \theta \, dx \right] \, dt
+ \int_{0}^{\tau} \int_{\Omega} \varrho(\tilde{u} - u) \cdot \left[ \tilde{\varrho} \tilde{u} + \tilde{u} \cdot \nabla_{x} \tilde{u} + \nabla_{x} P'(\tilde{\varrho}) \right] \, dx \, dt.
$$

(4.6)

Now, as $(\tilde{\varrho}, \tilde{u})$ is a strong solution, we have

$$
\varrho \tilde{\varrho} \tilde{u} + \varrho \tilde{u} \cdot \nabla_{x} \tilde{u} + \varrho \nabla_{x} P'(\tilde{\varrho}) = \text{div}_{x} \tilde{S}.
$$

(4.7)

Consequently, we may add the integral

$$
\int_{0}^{\tau} \int_{\Omega} (u - \tilde{u}) \cdot \text{div}_{x} \tilde{S} \, dx \, dt = \int_{0}^{\tau} \int_{\Omega} D_{u}(u - \tilde{u}) : \tilde{S} \, dx \, dt
$$

to both sides of the inequality (4.6). Regrouping terms on the left-hand side we get

$$
\int_{0}^{\tau} \int_{\Omega} \left[ F(D_{u} u) + F'(S) - S : D_{u} \tilde{u} + \tilde{S} : (D_{u} \tilde{u} - D_{u} u) \right] \, dx \, dt
= \int_{0}^{\tau} \int_{\Omega} \left[ F(D_{u} u) - F(D_{u} \tilde{u}) - \tilde{S} : (D_{u} u - D_{u} \tilde{u}) \right] \, dx,
$$

where we have used Fenchel–Young inequality

$$
F(D_{u} \tilde{u}) + F'(S) \geq S : D_{u} \tilde{u}.
$$

As $\tilde{S} \in \partial F(D_{u} \tilde{u})$ we can apply hypothesis (2.4) for $D = D_{u} \tilde{u}, S = \tilde{S}, Q = D_{u} u - D_{u} \tilde{u}$ to obtain

$$
\int_{0}^{\tau} \int_{\Omega} \left\| D_{u} \tilde{u} - D_{u} u \right\|^{2} \, dx
\lesssim \int_{0}^{\tau} \int_{\Omega} \left[ F(D_{u} u) + F'(S) - S : D_{u} \tilde{u} + \tilde{S} : (D_{u} \tilde{u} - D_{u} u) \right] \, dx \, dt.
$$

Finally, using the traceless version of Korn’s inequality, we may rewrite (4.6) in the form

$$
\int_{\Omega} E \left( \varrho, u \mid \tilde{\varrho}, \tilde{u} \right) (\tau, \cdot) \, dx + \int_{\Omega} \theta \, dx 
+ \int_{0}^{\tau} \int_{\Omega} |\nabla_{x} u - \nabla_{x} \tilde{u}|^{2} \, dx \, dt
\lesssim \int_{0}^{\tau} \chi(t) \left[ \int_{\Omega} E \left( \varrho, u \mid \tilde{\varrho}, \tilde{u} \right) (t, \cdot) \, dx + \int_{\Omega} \theta \, dx \right] \, dt
+ \int_{0}^{\tau} \int_{\Omega} \varrho(\tilde{u} - u) \cdot \left[ \tilde{\varrho} \tilde{u} + \tilde{u} \cdot \nabla_{x} \tilde{u} + \nabla_{x} P'(\tilde{\varrho}) \right] \, dx \, dt.
$$

(4.8)
Thus it remains to control the last integral in (4.8). Let $\varpi$ be a positive constant chosen so that $\varrho \leq \frac{\varpi}{2}$ in $(0, T) \times \Omega$. We consider two complementary cases. Suppose first that $\varrho \geq \varpi$. By virtue of hypotheses (4.1)–(4.3), we get

$$
\int_{\Omega} 1_{\varrho > \varpi} (\varrho - \varpi) (\tilde{u} - u) \cdot \left[ \partial_t \tilde{u} + \tilde{u} \cdot \nabla_x \tilde{u} + \nabla_x P' (\tilde{\varrho}) \right] \, dx
$$

\[
\leq \left\| 1_{\varrho > \varpi} (\varrho - \varpi) \right\|_{L^2 (\Omega)} \left\| 1_{\varrho > \varpi} (\varrho - \varpi) \right\|_{L^2 (\Omega)} \left\| \partial_t \tilde{u} + \tilde{u} \cdot \nabla_x \tilde{u} + \nabla_x P' (\tilde{\varrho}) \right\|_{L^2 (\Omega)}
\]

\[
\times \left\| \partial_t \tilde{u} + \tilde{u} \cdot \nabla_x \tilde{u} + \nabla_x P' (\tilde{\varrho}) \right\|_{L^2 (\Omega)}
\]

\[
\leq \delta \left\| \nabla_x P' (\tilde{\varrho}) \right\|_{L^2 (\Omega)}
\]

\[
\leq \delta \left\| \partial_t \tilde{u} + \tilde{u} \cdot \nabla_x \tilde{u} + \nabla_x P' (\tilde{\varrho}) \right\|_{L^2 (\Omega)}
\]

\[
\leq \delta \left\| \nabla_x P' (\tilde{\varrho}) \right\|_{L^2 (\Omega)}
\]

for any $\delta > 0$. As $\gamma \leq 2$, the standard Poincaré–Sobolev inequality yields

\[
\delta \left\| u - \tilde{u} \right\|_{L^2 (\Omega)}^2 \leq \frac{1}{2} \int_{\Omega} \left| \nabla_x u - \nabla_x \tilde{u} \right|^2 \, dx
\]

(4.11)

for a suitably small $\delta > 0$. By the same token $P$ is strictly convex on the compact interval $[0, \varpi]$; whence

\[
\left\| 1_{\varrho > \varpi} (\varrho - \varpi) \right\|_{L^2 (\Omega)} \leq \int_{\Omega} E (\varrho, u \mid \varrho, \tilde{u}) \, dx
\]

(4.12)

In view of hypothesis (4.1), and (4.5), we have

$$
\left\| \partial_t \tilde{u} + \tilde{u} \cdot \nabla_x \tilde{u} + \nabla_x P' (\tilde{\varrho}) \right\|_{L^2 (\Omega)} \leq \varpi (0, \Omega).
$$

Thus plugging (4.9)–(4.12) in (4.8) we may use the standard Gronwall argument to conclude

$$
\int_{\Omega} E (\varrho, u \mid \varrho, \tilde{u}) (\tau, \cdot) \, dx = 0, \quad \tau \in (0, T), \quad \int_{0}^{T} \int_{\Omega} \left| \nabla_x u - \nabla_x \tilde{u} \right|^2 \, dx \, dt = 0.
$$

The extension of the above result to more general viscous potentials still satisfying the coercivity condition (2.4) is straightforward, see [1]. The major drawback of theorem 4.1 is the absence of inflow. If $\Gamma_{in} \neq \emptyset$, a slight modification of the above arguments yields a positive result provided the boundary density $\partial_{\nu} u$ is bounded below away from vacuum.

**Theorem 4.3 (Weak–strong uniqueness, bounded domain, int/out flow).** Suppose that $\Omega \subset \mathbb{R}^d$, $d = 1, 2, 3$ is a bounded Lipschitz domain. Let the boundary conditions
be determined by \( \rho_B \in C^1_c(\mathbb{R}^d) \), \( u_B \in C^2_c(\mathbb{R}^d) \), where
\[
\rho_B(x) \geq \underline{\rho} > 0 \quad \text{for any } x \in \Gamma_m. \tag{4.13}
\]

Let \((\tilde{\rho}, \tilde{u})\) be a strong solution of the problem (1.1), (1.2) and (1.7) in \((0, T) \times \Omega\) in the sense specified in section 2.1 belonging to the class
\[
\tilde{u} \in C([0, T] \times \overline{\Omega}; \mathbb{R}^d), \quad \tilde{\rho} \geq 0,
\]
\[
\nabla^2 \tilde{u} \in L^2(0, T; L^q(\Omega; \mathbb{R}^d \times d \times d)), \quad \partial_t \tilde{u} \in L^2(0, T; L^q(\Omega; \mathbb{R}^d)), \tag{4.14}
\]
where
\[
1 < \gamma \leq 2, \quad q \geq \frac{2\gamma}{\gamma - 1}.
\]

Let \((\rho, u)\) be a weak solution of the same problem in the sense of definition 2.2 such that
\[
\rho(0, \cdot) = \tilde{\rho}(0, \cdot) = \rho_0, \quad (\rho u)(0, \cdot) = (\tilde{\rho} \tilde{u})(0, \cdot) = m_0, \quad u|_{\partial \Omega} = \tilde{u}|_{\partial \Omega} = u_B, \quad (\rho u) \cdot n|_{\Gamma_m} = (\tilde{\rho} \tilde{u}) \cdot n|_{\Gamma_m} = \rho_B u_B \cdot n,
\]
where
\[
(\rho_0)^{-1} \in W^{1,q}(\Omega; \mathbb{R}^d). \tag{4.15}
\]

Then
\[
\rho = \tilde{\rho}, \quad u = \tilde{u} \quad \text{in } (0, T) \times \Omega.
\]

Remark 4.4. The class (4.14) is slightly smaller than (4.1) in theorem 4.1 but still large enough to accommodate the strong solutions obtained in the Newtonian non-degenerate case by Valli and Zajaczkowski [16]. Note, however, that quite severe restrictions are imposed on \( \Gamma_m \) in [16].

Proof. The proof can be done following the same arguments as in theorem 4.1 as soon as we observe that
\[
\tilde{\rho}^{-1} \in L^\infty(0, T; W^{1,q}(\Omega)).
\]
To see this, write
\[
\tilde{\rho}^{-1} = r_1 + r_2, \quad r_1 = \delta + [\tilde{\rho}^{-1} - \delta]^+, \quad r_2 = [\tilde{\rho}^{-1} - \delta]^- \quad \text{for a suitable } \delta > 0.
\]
Consequently, it is enough to show
\[
\nabla_i r_i \in L^\infty(0, T; L^q(\Omega; \mathbb{R}^d)), \quad i = 1, 2.
\]
As for \( r_1 \), we have
\[
\nabla_i r_1 = \nabla_i [\tilde{\rho}^{-1} - \delta]^+ = (\gamma - 1)\text{sgn}^+[\tilde{\rho}^{-1} - \delta]^{\gamma-2} \nabla_i \tilde{\rho},
\]
so the desired conclusion follows from the hypothesis (4.14) as long as \( \delta > 0 \).

To obtain a similar estimate for \( r_2 \), we first extend the velocity field \( \tilde{u} \) and the initial density \( \rho_0 \) outside \( \Omega \) so that they satisfy the hypotheses (4.14) and (4.15) on \( \mathbb{R}^d \). Accordingly, the
equation of continuity (1.1) endowed with the extended velocity field \( \tilde{\mathbf{u}} \) admits a solution \( r \) determined uniquely through formula (2.7). By virtue of lemma 2.1, we have

\[ r^{\gamma - 1} \in L^\infty(0, T; W^{1,6}(R^3)). \]  

(4.16)

We say that a point \((t, x) \in [0, T] \times \Omega\) is regular, if the value of \( \tilde{\varrho}(t, x) \) is determined by formula (2.7), otherwise \((t, x)\) is singular. Clearly the backward characteristic curve emanating from a singular point reaches the boundary \( \partial \Omega \) at a positive time, and \( \tilde{\varrho}(t, x) \) is then given by (2.8). Obviously,

\[ \tilde{\varrho}(t, x) = r(t, x) \text{ whenever } (t, x) \text{ is regular}. \]

As stated in (4.13), \( \varrho_B \) is bounded below away from zero. Consequently, there exists \( \varepsilon = \varepsilon(T, \tilde{\mathbf{u}}) > 0 \) such that

\[ \tilde{\varrho}^{\gamma - 1}(t, x) \geq \varepsilon > 0 \text{ whenever } (t, x) \text{ is singular}. \]

Consequently, for \( 0 < \delta < \frac{\varepsilon}{2} \) and any fixed \( t \in [0, T] \), the set \( \{ x \in \Omega \mid \tilde{\varrho}^{\gamma - 1}(t, x) \leq \delta \} \) admits an open neighbourhood in \( \Omega \) that consists of regular points. Consequently,

\[ \nabla_x \tilde{\varrho} = \nabla_x [\tilde{\varrho}^{\gamma - 1} - \delta]^- = (\gamma - 1) \text{sgn}^- [\tilde{\varrho}^{\gamma - 1} - \delta] \nabla_x r^{\gamma - 1}, \]

and the desired conclusion follows from (4.16).

\[ \square \]

5. Weak–strong uniqueness, unbounded domains

The main difficulty when extending the previous results to unbounded domains is the lack of Poincaré–Sobolev inequality. For this reason, we consider only the case \( d = 3 \), where Sobolev’s inequality is available, namely

\[ \|v\|_{L^6(\Omega)} \lesssim \|\nabla_x v\|_{L^2(\Omega; R^3)} \]

(5.1)

for any \( v \in D^{1,2}_0(\Omega) \).

For simplicity, we restrict ourselves to the case \( u_B = 0 \).

5.1. General exterior domain

We look for strong solutions \((\tilde{\varrho}, \tilde{\mathbf{u}})\) belonging to the class introduced by Huang et al [9]. In particular,

\[ \tilde{\mathbf{u}} \in C([0, T]; D^{1,2}_0 \cap D^{3,2}(\Omega; R^3)) \cap L^2(0, T; D^{4,2}(\Omega; R^3)), \]

\[ \partial_t \tilde{\mathbf{u}} \in L^2(0, T; D^{1,2}_0(\Omega; R^3)) \cap L^2(0, T; D^{2,2}(\Omega; R^3)), \]

(5.2)

\[ \tilde{\varrho}, \ p(\tilde{\varrho}) \in C([0, T]; W^{3,2}(\Omega)). \]

We point out that the class is pertinent to the Newtonian case, where the momentum equation can be differentiated with respect to time to obtain higher order estimates on \( \partial_t \tilde{\mathbf{u}} \).

In addition, similarly to theorem 4.1, we impose the condition

\[ (\varrho_0)^{\gamma - 1} \in W^{1,6} \cap W^{1,\infty}(\Omega; R^3). \]  

(5.3)

As shown in lemma 2.1, this regularity will propagate in time; whence we get

\[ \sup_{t \in (0, T)} \|\nabla_x (\tilde{\varrho}^{\gamma - 1})\|_{L^4(\Omega; R^3)} \leq c. \]  

(5.4)
Now, following step by step the proof of theorem 4.1 we arrive at the inequality (4.8), specifically,
\[
\int_{\Omega} E (\rho, u | \tilde{\rho}, \tilde{u})(\tau, \cdot) \, dx + \int_{\Omega} 1 \, E (\tau) \, dx + \int_{0}^{\tau} \int_{\Omega} |\nabla x u - \nabla x \tilde{u}|^2 \, dx \, dt
\]
\[
\lesssim \int_{0}^{\tau} \chi(t) \left[ \int_{\Omega} E (\rho, u | \tilde{\rho}, \tilde{u})(t, \cdot) \, dx + \int_{\Omega} 1 \, E (t) \right] \, dx + \int_{0}^{\tau} \int_{\Omega} (\rho - \tilde{\rho})(\tilde{u} - u) \cdot b \, dx \, dt.
\]
where, by virtue of (5.2) and (5.4),
\[
b = \partial_t \tilde{u} + \tilde{u} \cdot \nabla x \tilde{u} + \nabla x P'(\tilde{\rho}) \in L^2(0, T; L^6 \cap L^\infty(\Omega; \mathbb{R}^3)).
\]
Similarly to the preceding section, we write
\[
\int_{0}^{\tau} \int_{\Omega} (\rho - \tilde{\rho})(\tilde{u} - u) \cdot b \, dx \, dt = \int_{0}^{\tau} \int_{\Omega} 1_{\rho \geq \tilde{\rho}}(\rho - \tilde{\rho})(\tilde{u} - u) \cdot b \, dx \, dt
\]
\[
+ \int_{0}^{\tau} \int_{\Omega} 1_{\rho < \tilde{\rho}}(\rho - \tilde{\rho})(\tilde{u} - u) \cdot b \, dx \, dt,
\]
where, exactly as in (4.9),
\[
\int_{0}^{\tau} \int_{\Omega} 1_{\rho < \tilde{\rho}}(\rho - \tilde{\rho})(\tilde{u} - u) \cdot b \, dx \, dt
\]
\[
\lesssim \|b\|_{L^\infty(\Omega; \mathbb{R}^3)} \int_{\Omega} E (\rho, u | \tilde{\rho}, \tilde{u}) \, dx.
\]

The integral
\[
\int_{0}^{\tau} \int_{\Omega} 1_{\rho < \tilde{\rho}}(\rho - \tilde{\rho})(\tilde{u} - u) \cdot b \, dx \, dt
\]

is more difficult to handle. Indeed, in order to mimic the arguments leading to (4.10), we would need
\[
b \text{ bounded in } L^2(0, T; L^3(\Omega; \mathbb{R}^3)),
\]
which does not follow from (5.2) if \( \Omega \) is an exterior domain and must be imposed as an extra hypothesis.

**Theorem 5.1 (Weak–strong uniqueness, unbounded domains).** Suppose that \( \Omega \subset \mathbb{R}^3 \), \( d = 1, 2, 3 \) is an exterior domain with compact Lipschitz boundary. Let \((\tilde{\rho}, \tilde{u})\) be a strong solution of the problem (1.1), (1.2) and (1.7) with \( u_B = 0 \) in \((0, T) \times \Omega\) in the sense specified in section 2.1 belonging to the class (5.2). In addition, suppose that
\[
\partial_t \tilde{u} \in L^2(0, T; L^3(\Omega; \mathbb{R}^3)).
\]
Let \((\varrho, \u)\) be a weak solution of the same problem in the sense of definition 2.2 such that
\[
\varrho(0, \cdot) = \tilde{\varrho}(0, \cdot) = \varrho_0, \quad (\varrho\u)(0, \cdot) = (\tilde{\varrho}\tilde{\u})(0, \cdot) = \mathbf{m}_0,
\]
where
\[
(\varrho_0)^{-1} \in W^{1,3} \cap W^{1,\infty}(\Omega; \mathbb{R}^3) \tag{5.10}
\]
Then
\[
\varrho = \tilde{\varrho}, \quad \u = \tilde{\u} \text{ in } (0, T) \times \Omega.
\]
The result extends easily to general in/out flow boundary conditions exactly as in theorem 4.3.

5.2. Compactly supported density

In view of the existence class (5.2) identified in [9], the hypothesis (5.9) does not seem very realistic, in particular in the presence of vacuum. To handle the general case, we restrict ourselves to the initial data with compactly supported initial density \(\varrho_0\) considered in [9] or [17]. Moreover, we restrict ourselves to the class of Newtonian fluids, where
\[
\mathcal{G}(\tilde{\nabla}_x \tilde{\u}) = \mu \left( \tilde{\nabla}_x \tilde{\u} + \tilde{\nabla}_t x \tilde{\u} - \frac{2}{3} \text{div}_x \tilde{\u} \right) + \lambda \text{div}_x \tilde{\u}, \quad \mu > 0, \quad \lambda \geq 0.
\]

In view of lemma 2.1, there exists \(R > 0\) sufficiently large such that
\[
\tilde{\varrho}(t, x) = 0 \quad \text{for all } t \in [0, T], \quad |x| \geq R. \tag{5.11}
\]

Going back to the integral (5.8), we get
\[
\int_0^T \int_\Omega 1_{v \in \mathbb{R}}(v - \tilde{\varrho})(\tilde{\u} - \u) \cdot \mathbf{b} \, dx \, dt = \int_0^T \int_{|x| \leq R} 1_{v \in \mathbb{R}}(v - \tilde{\varrho})(\tilde{\u} - \u) \cdot \mathbf{b} \, dx \, dt \\
+ \int_0^T \int_{|x| > R} 1_{v \in \mathbb{R}}(v - \tilde{\varrho})(\tilde{\u} - \u) \cdot \mathbf{b} \, dx \, dt,
\]
where
\[
\int_{|x| \leq R} 1_{v \in \mathbb{R}}(v - \tilde{\varrho})(\tilde{\u} - \u) \cdot \mathbf{b} \, dx \leq \|1_{v \in \mathbb{R}}(v - \tilde{\varrho})\|_{L^2(\Omega)} \|\u - \tilde{\u}\|_{L^p(\mathbb{R}^3)} \|\mathbf{b}\|_{L^q(\Omega \cap \{|x| \leq R\})}.
\]
Seeing that
\[
\|\mathbf{b}\|_{L^q(\Omega \cap \{|x| \leq R\})} \leq c(R)\|\mathbf{b}\|_{L^q(\Omega; \mathbb{R}^3)}
\]
the above integral is controlled by (5.6).

Finally,
\[
\int_{|x| > R} 1_{v \in \mathbb{R}}(v - \tilde{\varrho})(\tilde{\u} - \u) \cdot \mathbf{b} \, dx \leq \delta \|\u - \tilde{\u}\|_{L^p(\mathbb{R}^3)}^2 + c(\delta)\|1_{v \in \mathbb{R}}\|_{L^2(\Omega)}^2 \|\mathbf{b}\|_{L^q(\{|x| \geq R\})}^2 \tag{5.12}
\]
for any $\delta > 0$, where
\[
\frac{1}{q} = \frac{5}{6} - \frac{1}{\gamma}.
\]
Suppose that $1 < \gamma \leq 2$. Consequently,
\[
\|v_{\epsilon \xi_{\tau}}\|^2_{L^2(\Omega)} \leq c(\Omega) \|v_{\epsilon \xi_{\tau}}\|^2_{L^2(\Omega)} \lesssim \int_{\Omega} E(\rho, u | \tilde{\rho}, \tilde{u}) \, dx.
\]
Thus, to close the estimates, we need to control
\[
\int_0^T \|b\|^2_{L^q(|x| \geq R)} \, dt \leq c(q) \text{ for any } q > 3.
\]
To see (5.13), we realise that
\[
b = \partial_t \tilde{u} + \tilde{u} \cdot \nabla x \tilde{u} \quad \text{if } |x| > R.
\]
If the fluid is Newtonian, we get immediately from (4.7) that
\[
\mu \text{div}_x \left( \nabla_x \partial_t \tilde{u} + \nabla_x' \partial_t \tilde{u} - \frac{2}{3} \text{div}_x \partial_t \tilde{u} \right) + \lambda \text{div}_x \partial_t \tilde{u} = 0 \quad \text{for } |x| > R,
\]
where, in view of (5.2)
\[
\partial_t \tilde{u} \in D^{1,2} \cap L^\infty(|x| > R), \quad \partial_t \tilde{u} \in C^\alpha(|x| = R) \quad \text{for some } \alpha > 0 \text{ and a.a. } t \in (0, T).
\]
Using the standard elliptic estimates, we get
\[
\|\partial_t \tilde{u}(t, \cdot)\| \lesssim \frac{1}{|x|} \|\partial_t \tilde{u}(t, \cdot)\|_{C(|x| = R)} \quad \text{for all } |x| \geq R,
\]
which, together with (5.2), yields (5.13). Thus we are able to control the integral (5.12) as soon as $1 < \gamma < 2$.

We have shown the following result.

**Theorem 5.2 (Weak–strong uniqueness, compactly supported density).** Suppose that $\Omega \subset \mathbb{R}^3$ is a Lipschitz exterior domain, $u_B = 0$, and $1 < \gamma < 2$. In addition, let $S$ be Newtonian,
\[
S = \mu \left( \nabla_x u + \nabla_x u - \frac{2}{3} \text{div}_x u I \right) + \lambda \text{div}_x u I, \quad \mu > 0, \quad \lambda \geq 0.
\]
Let $(\tilde{\rho}, \tilde{u})$ be a strong solution of the problem (1.1) and (1.2) in $(0, T) \times \Omega$ in the sense specified in section 2.1 belonging to the class (5.2). Let $(\rho, u)$ be a weak solution of the same problem in the sense of definition 2.2 such that
\[
\rho(0, \cdot) = \tilde{\rho}(0, \cdot) = \rho_0, \quad (\rho u)(0, \cdot) = (\tilde{\rho} \tilde{u})(0, \cdot) = m_0,
\]
where
\[
\rho_0 \in C_s(\overline{\Omega}), \quad (\rho_0)^{-1} \in W^{1,6} \cap W^{1,\infty}(\Omega; \mathbb{R}^3).
\]
Then

\[ \rho = \tilde{\rho}, \ u = \tilde{u} \text{ in } (0, T) \times \Omega. \]

**Remark 5.3.** Local existence of strong solutions as well as global existence for small initial data in the class included in (5.2) was proved by Huang *et al* [9, lemma 2.1], see also Cho and Kim [4, theorem 3] for local existence in the class of more regular solution.

The result can be extended to the case of general in/out flow boundary conditions in the spirit of theorem 4.3. The hypothesis of Newtonian viscous stress is however necessary.

### 5.3. Positive density

Finally, we consider the physically relevant case \( \Omega = \mathbb{R}^3 \),

\[ 0 < \rho_0(x) \leq \overline{\rho}, \]

\[ |\nabla_x (\rho_0)^{-1}(x)| \lesssim \frac{1}{1 + |x|^{\beta}}, \quad \alpha = \alpha(\gamma) > 0, \tag{5.15} \]

meaning vacuum is not present but \( \rho_0 \) decays to zero as \( |x| \to \infty \). We consider the Newtonian viscous stress (5.14) and restrict slightly the class (5.2) of strong solutions

\[ \tilde{\rho} \in C([0, T]; D^{2,2}(\mathbb{R}^3)), \quad \partial_t \tilde{\rho} \in L^1(0, T; L^1(\mathbb{R}^3; \mathbb{R}^3)), \]

\[ \sqrt{\tilde{\rho}} \partial_t \tilde{\rho} \in L^1(0, T; L^2(\mathbb{R}^3; \mathbb{R}^3)). \tag{5.16} \]

introduced by Cho and Kim [4].

The first observation is that any solution belonging to (5.16) satisfies

\[ \tilde{\rho} \in C([0, T]; BC^1(\mathbb{R}^3; \mathbb{R}^3)), \quad \nabla^2 \tilde{\rho} \in L^2(0, T; BC(\mathbb{R}^3; \mathbb{R}^3)). \]

Thus a direct inspection of formula (2.7) reveals that the decay properties of the initial density propagate in time. More specifically, we have

\[ |\nabla_x (\rho_0)^{-1}(x)| \lesssim \frac{1}{1 + |x|^{\beta}} \Rightarrow (\rho_0)^{-1}(x) \lesssim \frac{1}{1 + |x|^{\alpha-1}} \]

\[ \Rightarrow \rho^{-1}(t, x) \lesssim \frac{1}{1 + |x|^{\alpha-1}} \]

\[ \Rightarrow |\nabla_x (\rho(t, x))^{-1}| \lesssim \frac{1}{1 + |x|^{\alpha-1}} \]

In particular, we get

\[ \nabla_x P(\tilde{\rho}) \in L^\infty(0, T; L^1 \cap L^\infty(\mathbb{R}^3; \mathbb{R}^3)) \text{ provided } \alpha > 2, \tag{5.17} \]

and

\[ \tilde{\rho}(t, x) \approx \frac{1}{1 + |x|^{\beta}}, \quad |\nabla_x \tilde{\rho}(t, x)| \lesssim \frac{1}{1 + |x|^{\beta}} \quad \text{for all } t \in [0, T], \ x \in \mathbb{R}^3, \quad \beta = \frac{\alpha - 1}{\gamma - 1}. \tag{5.18} \]
Consequently, all steps in section 5.1 up to formula (5.7) can be performed and we are left with the integral (5.8)

\[
\int_0^T \int_{\mathbb{R}^3} 1_{\nu \leq \tilde{\rho}} (\rho - \tilde{\rho}) (u - \tilde{u}) : \left( \tilde{\rho} \tilde{u} + \tilde{\rho} \nabla_x (\tilde{\rho} U) \right) \, dx \, dt.
\] (5.19)

We proceed in several steps:

Step 1. As \( \tilde{u} \in L^\infty((0,T) \times \mathbb{R}^3; \mathbb{R}^3) \), \( \nabla_x \tilde{u} \in L^\infty \cap L^2((0,T) \times \mathbb{R}^3; \mathbb{R}^3) \), we get, by interpolation,

\[
u \tilde{u} \cdot \nabla_x \tilde{u} \in L^2(0,T; L^3(\mathbb{R}^3, \mathbb{R}^3))
\]

Similarly, it follows from (5.18) that

\[
\nabla_x P'(\tilde{\rho}) \in L^\infty((0,T) \times \mathbb{R}^3, \mathbb{R}^3)
\]

Consequently,

\[
\int_0^T \int_{\mathbb{R}^3} 1_{\nu \leq \tilde{\rho}} (\rho - \tilde{\rho}) (u - \tilde{u}) : \left( \tilde{\rho} \tilde{u} + \tilde{\rho} \nabla_x P'(\tilde{\rho}) \right) \, dx \, dt
\]

\[
\leq \| \rho - \tilde{\rho} \|_{L^2(\mathbb{R}^3)} \| u - \tilde{u} \|_{L^\infty(\mathbb{R}^3, \mathbb{R}^3)} \| \tilde{\rho} \|_{L^\infty(\mathbb{R}^3, \mathbb{R}^3)} \| \nabla_x \tilde{u} \|_{L^2(\mathbb{R}^3, \mathbb{R}^3)} \| \nabla_x P'(\tilde{\rho}) \|_{L^2(\mathbb{R}^3, \mathbb{R}^3)}
\]

and we may use the arguments of section 5.1 to control the expression on the right-side. Thus (5.19) reduces to

\[
\int_0^T \int_{\mathbb{R}^3} 1_{\nu \leq \tilde{\rho}} (\rho - \tilde{\rho}) (u - \tilde{u}) : \left( \tilde{\rho} \tilde{u} + \tilde{\rho} \nabla_x P'(\tilde{\rho}) \right) \, dx \, dt
\]

(5.21)

Step 2. To control \( \tilde{\rho} \tilde{u} \), we differentiate the momentum equation with respect to time obtaining

\[
\mathcal{L}[\tilde{\rho} \tilde{u}] = \tilde{\rho} \tilde{u}^\prime_x \left( \tilde{\rho} \tilde{u} + \tilde{\rho} \nabla_x \tilde{\rho} \right) + \nabla_x \tilde{\rho} P, \quad \text{(5.22)}
\]

where \( \mathcal{L} \) denotes the Lamé elliptic operator

\[
\mathcal{L}[v] = \mu \text{div}_x \left( \nabla_x v + \frac{2}{3} \text{div}_x v \right) + \lambda \nabla_x \text{div}_x v.
\] (5.23)

Now,

\[
\tilde{\rho} (\tilde{\rho} \tilde{u} \otimes \tilde{u}) = \tilde{\rho} \tilde{u}' \left( \tilde{\rho} \tilde{u} \otimes \tilde{u} \right) + \tilde{\rho} \nabla_x \tilde{\rho} \tilde{u} \otimes \tilde{u} + \tilde{\rho} \tilde{u} \otimes \nabla_x \tilde{\rho} \tilde{u}
\]

\[
= -\tilde{u} \cdot \nabla_x \tilde{\rho} \tilde{u} \otimes \tilde{u} - \tilde{\rho} \text{div}_x \tilde{\rho} \tilde{u} \otimes \tilde{u} + \tilde{\rho} \left( \tilde{\rho} \tilde{u} \otimes \nabla_x \tilde{\rho} \tilde{u} \right). \]

By Hölder’s inequality

\[
\| \tilde{u} \cdot \nabla_x \tilde{\rho} \tilde{u} \|_{L^2(\mathbb{R}^3, \mathbb{R}^3)} \leq \| \tilde{u} \|_{L^2(\mathbb{R}^3, \mathbb{R}^3)} \| \nabla_x \tilde{\rho} \|_{L^1(\mathbb{R}^3, \mathbb{R}^3)}
\]

Similarly

\[
\| \text{div}_x \tilde{u} \tilde{u} \|_{L^2(\mathbb{R}^3, \mathbb{R}^3)} \leq \| \tilde{u} \|_{L^2(\mathbb{R}^3, \mathbb{R}^3)} \| \text{div}_x \tilde{u} \|_{L^1(\mathbb{R}^3, \mathbb{R}^3)} \| \tilde{u} \|_{L^2(\mathbb{R}^3, \mathbb{R}^3)}
\]

and

\[
\| \tilde{\rho} \tilde{u}' \left( \tilde{\rho} \tilde{u} \otimes \tilde{u} \right) \|_{L^2(\mathbb{R}^3, \mathbb{R}^3)} \leq 2 \| \tilde{\rho} \|_{L^1(\mathbb{R}^3, \mathbb{R}^3)} \| \tilde{u} \|_{L^2(\mathbb{R}^3, \mathbb{R}^3)} \| \tilde{u} \|_{L^2(\mathbb{R}^3, \mathbb{R}^3)} \| \tilde{\rho} \|_{L^1(\mathbb{R}^3, \mathbb{R}^3)}.
\]
As \((\tilde{\rho}, \tilde{u})\) belongs to the class \((5.16)\) and \(\tilde{\rho}\) satisfies \((5.20)\), it is easy to check that
\[
\partial_t (\tilde{\rho} \tilde{u} \otimes \tilde{u}) \in L^2 \left(0, T; L^{3/2}(R^3; R^9)\right).
\] (5.24)

Next,
\[
\partial_t p(\tilde{\rho}) = -p'(\tilde{\rho}) \left(\nabla_x \tilde{\rho} \cdot \tilde{u} - \tilde{\rho} \text{div}_x \tilde{u}\right),
\]
where
\[
\left\|p'(\tilde{\rho}) \nabla_x \tilde{\rho} \cdot \tilde{u}\right\|_{L^{3/2}(\Omega)} \lesssim \left\|\tilde{\rho}\right\|_{L^6(R^3)} \left\|\nabla x P'(\tilde{\rho})\right\|_{L^3(R^3)} \left\|\tilde{u}\right\|_{L^6(R^3)}.
\]
Similarly to the above, we conclude
\[
\partial_t p(\tilde{\rho}) \in L^\infty \left(0, T; L^2(R^3)\right).
\] (5.25)

Going back to the equation \((5.22)\), we get
\[
\mathcal{L}[\partial_t \tilde{u}] = \partial_t^2 \tilde{u} + \text{div}_x B
\]
where
\[
B \in L^2 \left(0, T; L^{3/2}(R^3; R^9)\right).
\]
In view of the standard elliptic estimates, we may write
\[
\partial_t \tilde{u} = z + v,
\]
where
\[
\mathcal{L}[z] = \partial_t^2 \tilde{u},
\] (5.26)
while
\[
v \in L^2 \left(0, T; D_0^{1,3/2}(R^3; R^3)\right) \hookrightarrow L^2(0, T; L^3(R^3; R^3)).
\]
In view of the arguments employed in step 1, our task reduces to controlling the integral
\[
\int_0^T \int_{R^3} 1_{\text{exc}}(\rho - \tilde{\rho})(\mathbf{u} - \tilde{\mathbf{u}}) \cdot \mathbf{z} \, d\mathbf{x} \, dt.
\] (5.27)

Step 3. First, we write the right-hand side of the elliptic equation \((5.26)\) in the form
\[
\partial_t^2 \tilde{u} = \partial_t^2 \tilde{\rho} \tilde{u} + \tilde{\rho} \partial_t^2 \tilde{u} - 2 \text{div}_x (\tilde{\rho} \tilde{u}) \partial_t \tilde{u}
\]
\[
= -\partial_t \text{div}_x (\tilde{\rho} \tilde{u}) \mathbf{u} + \tilde{\rho} \partial_t \tilde{u} - 2 \text{div}_x (\tilde{\rho} \tilde{u}) \partial_t \tilde{u} - 2 \nabla_x \tilde{\rho} \cdot \tilde{u} \partial_t \tilde{u},
\]
where, furthermore,
\[
\partial_t \text{div}_x (\tilde{\rho} \tilde{u}) \mathbf{u} = \text{div}_x \left(\partial_t (\tilde{\rho} \tilde{u}) \otimes \mathbf{u}\right) - \partial_t (\tilde{\rho} \tilde{u}) \cdot \nabla_x \tilde{u}.
\]
As the term \(\text{div}_x (\partial_t (\tilde{\rho} \tilde{u}) \otimes \mathbf{u})\) can be handled exactly as in step 2, our task reduces to estimating the integral
\[
\int_0^T \int_{R^3} 1_{\text{exc}}(\rho - \tilde{\rho})(\mathbf{u} - \tilde{\mathbf{u}}) \cdot \mathbf{v} \, d\mathbf{x} \, dt.
\] (5.28)
where \( w \) solves the elliptic system

\[
\mathcal{L}[w] = g,
\]

with the right-hand side

\[
g = \tilde{\varrho} \partial u \cdot \nabla u - \tilde{\varrho} \text{div} u (u \cdot \nabla u) + u \nabla \tilde{\varrho} \cdot u \cdot \nabla u + \tilde{\varrho} \Box \tilde{\varrho} u - 2 \tilde{\varrho} \text{div} (u \tilde{\varrho}) u - 2 \nabla \tilde{\varrho} \cdot u \partial_\varrho u.
\]

(5.30)

Our goal is to show that the function \( w \) decays to zero for large \( x \), more specifically, we shall see that

\[
|w(t, x)| \leq \chi(t) \frac{1}{|x|} \quad \text{with} \quad \chi \in L^2(0, T).
\]

(5.31)

Taking (5.31) for granted, we use Hardy’s inequality to estimate the integral (5.28),

\[
\int_{\mathbb{R}} 1_{\varrho \leq \varepsilon} (|\varrho - \tilde{\varrho}|(u - \tilde{\varrho}u) \cdot w) \, dx \leq \delta \int_{\mathbb{R}} \frac{|u - \tilde{\varrho}u|^2}{|x|^2} \, dx + c(\delta) \int_{\mathbb{R}} \chi^2 1_{\varrho \leq \varepsilon} (|\varrho - \tilde{\varrho}|^2)
\]

\[
\lesssim \delta \int_{\mathbb{R}} |\nabla u - \nabla \tilde{\varrho}u|^2 \, dx + c(\delta) \int_{\mathbb{R}} \chi^2 (\varrho, u |\tilde{\varrho}, \tilde{\varrho}|), \quad \delta > 0 \text{ arbitrary}.
\]

Thus the proof of weak strong uniqueness can be completed via Gronwall’s argument exactly as in section 4.1.

It remains to show (5.31). As \( w \) solves the elliptic problem (5.29), we get

\[
w(t, x) = \int_{\mathbb{R}} G(x, y) \cdot g(t, y) \, dy,
\]

where \( G \) is the Green kernel associated to \( \mathcal{L} \). Writing

\[
\int_{\mathbb{R}} G(x, y) \cdot g(t, y) \, dy = \frac{1}{|x|} \int_{\mathbb{R}} (|x| - |y|) G(x, y) \cdot g(t, y) \, dy + \frac{1}{|x|} \int_{\mathbb{R}} |y| G(x, y) \cdot g(t, y) \, dy
\]

and using the fact

\[
|G(x, y)| \lesssim \frac{1}{|x - y|},
\]

we observe that (5.31) follows provided

\[
\|g(t, \cdot)\|_{L^q(\mathbb{R}; L^r)} + \|x |g(t, \cdot)|\|_{L^q(\mathbb{R}; L^r)} \lesssim \chi(t) \quad \text{for some} \quad \chi \in L^2(0, T), \quad q \in \left( \frac{3}{2} - \varepsilon, \frac{3}{2} + \varepsilon \right), \quad \varepsilon > 0.
\]

(5.32)

Thus our ultimate goal is to check (5.32) for all terms appearing on the right-hand side of (5.30). First, we have

\[
\|\tilde{\varrho} \partial_\varrho \tilde{u} \cdot \nabla \tilde{u}\|_{L^1(\mathbb{R})} \leq \|\tilde{\varrho}\|_{L^\infty(\mathbb{R})} \|\partial_\varrho \tilde{u}\|_{L^q(\mathbb{R})} \|\nabla \tilde{u}\|_{L^q(\mathbb{R})},
\]

and

\[
\|x |\tilde{\varrho} \partial_\varrho \tilde{u} \cdot \nabla \tilde{u}\|_{L^1(\mathbb{R})} \leq \|x |\tilde{\varrho}\|_{L^\infty(\mathbb{R})} \|\partial_\varrho \tilde{u}\|_{L^q(\mathbb{R})} \|\nabla \tilde{u}\|_{L^q(\mathbb{R})}.
\]
In view of (5.16), the bound (5.32) holds as soon as $\beta > 3$. The same argument applies to the terms $\tilde{\rho} \, \text{div} \, \mathbf{u} \, \partial \mathbf{u}$, $\tilde{\rho} \, \text{div} \, \mathbf{u} \cdot \nabla \mathbf{u}$, and $\tilde{\rho} \, \mathbf{u} \cdot \nabla \mathbf{u}$. Moreover, seeing that $\nabla \tilde{\rho}$ decays at least as fast as $\tilde{\rho}$, we can handle $\tilde{\rho} \, \nabla \mathbf{u} \cdot \tilde{\mathbf{u}} \cdot \nabla \mathbf{u}$ and $\nabla \tilde{\rho} \cdot \tilde{\mathbf{u}}$ in the same manner. Here, we have used the fact that $\tilde{\mathbf{u}}$ is uniformly bounded.

Finally,

$$\| \tilde{\rho} \partial^2 \mathbf{u} \|_{L^2(\mathbb{R}^3)} \leq \| \sqrt{\tilde{\rho}} \|_{L^2(\mathbb{R}^3)} \| \sqrt{\tilde{\rho}} \partial^2 \mathbf{u} \|_{L^2(\mathbb{R}^3)}.$$ 

Thus the choice $\beta > 3$ yields (5.32).

We have shown the following result:

**Theorem 5.4 (Weak–strong uniqueness, positive density).** Suppose that $\Omega = \mathbb{R}^3$ and $1 < \gamma \leq 2$. In addition, let $\mathbb{S}$ be Newtonian,

$$\mathbb{S} = \mu \left( \nabla \mathbf{u} + \nabla^T \mathbf{u} - \frac{2}{3} \text{div} \mathbf{u} \mathbb{I} \right) + \lambda \text{div} \mathbf{u} \mathbb{I}, \quad \mu > 0, \lambda \geq 0.$$

Let $(\tilde{\rho}, \tilde{\mathbf{u}})$ be a strong solution of the problem (1.1) and (1.2) in $(0, T) \times \mathbb{R}^3$ in the sense specified in section 2.1 belonging to the class (5.16). Let $(\rho, \mathbf{u})$ be a weak solution of the same problem in the sense of definition 2.2 such that

$$\rho(0, \cdot) = \tilde{\rho}(0, \cdot) = \rho_0, \quad (\rho \mathbf{u})(0, \cdot) = (\tilde{\rho} \tilde{\mathbf{u}})(0, \cdot) = \mathbf{m}_0,$$

where

$$0 < \rho_0(x) \leq \overline{\rho},$$

$$|\nabla \rho_0^{-1}(x)| \lesssim \frac{1}{1 + |x|^\alpha}, \quad \alpha > \max \{ 2; 3\gamma - 2 \}.$$

Then

$$\rho = \tilde{\rho}, \quad \mathbf{u} = \tilde{\mathbf{u}} \quad \text{in} \; (0, T) \times \mathbb{R}^3.$$

**Remark 5.5.** Local existence of strong solutions in the class (5.16) was proved by Cho and Kim [4, theorem 3], see also Huang et al [9, lemma 2.1].

Similarly to the preceding results, theorem 5.4 can be extended to exterior domains with general in/out flow boundary conditions. Vacuum can be accommodated at the expense of various technical difficulties. Finally, it is worth noting that the condition $\alpha > 3\gamma - 2$ in theorem 5.4 is critical for the fluid to have finite mass, namely

$$\int_{\mathbb{R}^3} \tilde{\rho} \, dx < \infty.$$

**References**

[1] Abbatiello A, Feireisl E and Novotný A 2021 Generalized solutions to models of compressible viscous fluids *Discrete Continuous Dyn. Syst. A* 41 1–28

[2] Bulíček M, Gwiazda P, Málek J and Świerczewska Gwiazda A 2012 On unsteady flows of implicitly constituted incompressible fluids *SIAM J. Math. Anal.* 44 2756–801
[3] Cho Y, Choe H J and Kim H 2004 Unique solvability of the initial boundary value problems for compressible viscous fluids J. Math. Pures Appl. 83 243–75
[4] Cho Y and Kim H 2006 On classical solutions of the compressible Navier–Stokes equations with nonnegative initial densities Manuscripta Math. 120 91–129
[5] Fang L, Zhu H and Guo Z 2018 Global classical solution to a one-dimensional compressible non-Newtonian fluid with large initial data and vacuum Nonlinear Anal. 174 189–208
[6] Feireisl E, Jin B J and Novotný A 2012 Relative entropies, suitable weak solutions, and weak-strong uniqueness for the compressible Navier-Stokes system J. Math. Fluid Mech. 14 712–30
[7] Feireisl E, Kwon Y-S and Novotný A 2021 On the long-time behavior of dissipative solutions to models of non-Newtonian compressible fluids Arch. Ration. Mech. Anal. 241 475–95
[8] Gong H, Li J, Liu X-G and Zhang X 2020 Local well-posedness of isentropic compressible Navier–Stokes equations with vacuum Commun. Math. Sci. 18 1891–909
[9] Huang X, Li J and Xin Z 2012 Global well-posedness of classical solutions with large oscillations and vacuum to the three-dimensional isentropic compressible Navier–Stokes equations Commun. Pure Appl. Math. 65 549–85
[10] Kalousek M, Mácha V and Nečasová Š 2020 Local-in-time existence of strong solutions to a class of compressible non-Newtonian Navier–Stokes equations (arXiv:2012.01795v1)
[11] Li J and Xin Z 2019 Global well-posedness and large time asymptotic behavior of classical solutions to the compressible Navier–Stokes equations with vacuum Ann. PDE 5 37
[12] Liang Z 2021 Global strong solutions of Navier–Stokes equations for heat-conducting compressible fluids with vacuum at infinity J. Math. Fluid Mech. 23 17
[13] Lions P-L 1998 Mathematical Topics in Fluid Dynamics (Compressible Models vol 2) (Oxford: Oxford University Press)
[14] Merle F, Raphael P, Rodnianski I and Szeftel J 2019 On the implosion of a three dimensional compressible fluid (arXiv:1912.11009)
[15] Rozanova O 2008 Blow-up of smooth highly decreasing at infinity solutions to the compressible Navier–Stokes equations J. Differ. Equ. 245 1762–74
[16] Valli A and Zajaczkowski W M 1986 Navier–Stokes equations for compressible fluids: global existence and qualitative properties of the solutions in the general case Commun. Math. Phys. 103 259–96
[17] Xin Z 1998 Blowup of smooth solutions to the compressible Navier-Stokes equation with compact density Commun. Pure Appl. Math. 51 229–40