On Papkovich-Neuber type representations for solutions of the Navier-Lamé equation in spatial star-shaped domains

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ABSTRACT. We develop a Papkovich-Neuber type representation formula for the solutions of the Navier-Lamé equation of linear elastostatics for spatial star-shaped domains. This representation is compared to the existing ones.

1. Introduction

Papkovich-Neuber type formulae represent the solutions of the Navier-Lamé equation of linear elastostatics via auxiliary harmonic potentials. This technique is also suitable for generating solutions of the Stokes equations for the creeping flow of an uncompressible fluid. It originates in the papers of Papkovich [6] and Neuber [4]. An important variant of this representation was given by Kratz [2] for the Stokes equation, where the uniqueness of the harmonic potentials in the representation was also proved. The Kratz representation [2] is valid for general planar domains, but in the spatial case only for star-shaped ones. Another variant of it can be found in [3] also for spatial star-shaped domains. In [8] the present author proved that the representation formulae in [2] and [3] are equivalent, moreover, generalized them for the Navier-Lamé equation of linear elastostatics. The representation formulae derived in [2] and [8] solve the problem of eliminating the scalar harmonic potential from the general Papkovich-Neuber representation in the special case of spatial star-shaped domains maintaining simultaneously the completeness of the representation, see [5] and also the references given there. Although these formulae are ment to derive analytic solutions of the respective equations, they can also be utilized in numerical methods, see for example [1].

In this paper we derive another representation similar to that in [8], Theorem 5.3 for a modified Stokes type equation also equivalent to the Navier-Lamé system. The connection of this representation to existing ones is also investigated.

2. Main result

In this paper we develop representation formulae for the solutions $u \in C^2(\Omega)$ and $q \in C^1(\Omega)$ of the equation
\[-\Delta u = \operatorname{rot} q \quad \text{and} \quad \operatorname{rot} u = \tilde{v} q, \]  
where $\Omega$ is a spatial star-shaped domain with respect to the origin. (1) is similar to the system
\[-\Delta u = \operatorname{grad} p \quad \text{and} \quad -\operatorname{div} u = \nu p, \]
which was (although not in this form) considered in [8]. Note that both (1) and (2) are equivalent to the Navier-Lamé equation
\[(\lambda + \mu) \operatorname{grad} u + \mu \Delta u = 0 \]  
in linear elastostatics, where $\lambda$ and $\mu$ are the Lamé constants. These constants satisfy usually $\mu > 0$ and $\lambda + \frac{2}{3} \mu > 0$, where the latter quantity is the compression modulus. In order to establish the connection of (3) to (1) and (2) we have to set $q = \frac{\lambda + \mu}{\lambda + 2\mu} \operatorname{rot} u$ and $p = \frac{\lambda + \mu}{\mu} \operatorname{div} u$, respectively. We also have $\nu = -\frac{\mu}{\lambda + \mu}$ in (2) and $\tilde{v} = \frac{\lambda + 2\mu}{\lambda + \mu} = 1 - \nu$ in (1).

**Theorem 1.** Let $\Omega \subset \mathbb{R}^n$ be a star-shaped domain with respect to the origin, and set $\tilde{v} \in \mathbb{R}$, $\tilde{v} > 0$, $\tilde{v} \neq \frac{1}{4}$, 1. The functions $u \in C^2(\Omega)$ and $q \in C^1(\Omega)$ satisfy (1) if and only if there exists a harmonic function $\tilde{h} \in C^2(\Omega)$ such that
\[u(x) = \frac{1}{2} \nabla \left(x \cdot \tilde{h}(x)\right) + \frac{1}{2} \operatorname{rot} \left(x \times \tilde{h}(x)\right) + (1 - 2\tilde{v})\tilde{h}(x), \]
\[q(x) = -2 \operatorname{rot} \tilde{h}(x) - 4\nabla \phi(x) \]
for $x \in \Omega$. The harmonic function $\tilde{h}$ is unique and we have
\[\tilde{h}(x) = \frac{2}{1 - 4\tilde{v}} \left(u(x) + \frac{1}{4} x \times q(x) - \frac{1}{4(1 - \tilde{v})} x \operatorname{div} u(x) + x \times \nabla \phi(x)\right), \]
where the function $\phi$ is harmonic in $\Omega$ and defined by
\[\phi(x) = -\frac{1}{4} \int_0^1 t^{4\tilde{v}} x \cdot q(tx) \, dt. \]

**Proof.** We need the following identities, where we assume the existence and continuity of the involved scalar and vector valued functions. These are the same identities used in [2] and [8].
\[x \operatorname{div} u + x \times \operatorname{rot} u = u + \nabla (x \cdot u) + \operatorname{rot} (x \times u), \]
\[-\Delta u = -\nabla \operatorname{div} u + \operatorname{rot} \operatorname{rot} u, \]
\[\operatorname{div} (x \times u) = -x \cdot \operatorname{rot} u, \]
\[\operatorname{div} (\phi x) = 3\phi + x \cdot \nabla \phi, \]
\[\operatorname{rot} (\phi x) = -x \times \nabla \phi, \]
\[\operatorname{rot} (x \times \nabla \phi) = -\nabla (\phi + x \cdot \nabla \phi) + x \Delta \phi, \]
\[\nabla (r^2 \phi) = 2\phi x + r^2 \nabla \phi, \text{ for } r = |x|, \]
\[\Delta (x \cdot u) = x \cdot \Delta u + 2 \operatorname{div} u, \]
\[
\Delta(x \times u) = x \times \Delta u + 2 \text{rot} u, \quad (16)
\]
\[
\Delta(\phi x) = x\Delta \phi + 2\nabla \phi, \quad (17)
\]
\[
\Delta(x \times \nabla \phi) = x \times \nabla \Delta \phi. \quad (18)
\]

From the expression (7) of \( \phi \) we obtain setting \( k(x) = x \cdot q(x) \) with integration by parts that

\[
x \cdot \nabla \phi(x) = -\frac{1}{4} \int_0^1 t^4 \bar{\nu} x \cdot (\nabla k)(tx) \, dt = -\frac{1}{4} \int_0^1 t^4 \frac{d}{dt} k(tx) \, dt
\]
\[
= -\frac{1}{4} \left[ t^4 \bar{\nu} k(tx) \right]_0^1 + \frac{1}{4} \int_0^1 4 \bar{\nu} t^4 k(tx) \, dt = -\frac{1}{4} x \cdot q(x) - 4 \bar{\nu} \phi(x).
\]

Hence, the function \( \phi \) defined by (7) satisfies the equation

\[
4 \bar{\nu} \phi(x) + x \cdot \nabla \phi(x) + \frac{1}{4} x \cdot q(x) = 0. \quad (19)
\]

Taking the divergence of the second equation of (1) gives \( \text{div} q = 0 \) in a view of \( \bar{\nu} > 0 \). Moreover, by (1) and (9) we have \( \Delta q = - \text{rot rot} q = \text{rot} \Delta u = \Delta \text{rot} u = \bar{\nu} \Delta q \), which means \( \Delta q = 0 \) using \( \bar{\nu} \neq 1 \). This implies by (15) that \( \Delta k = x \cdot \Delta q + 2 \text{div} q = 0 \) and

\[
\Delta \phi(x) = -\frac{1}{4} \int_0^1 t^4 \bar{\nu} + 2(\Delta k)(tx) \, dt = 0.
\]

That is, the function \( \phi \) is harmonic. Note also, that \( \bar{\nu} > 0 \) is sufficient for the integral in (7) to be well defined.

First assume, that \( u \) and \( q \) are given by (4) and (5) with harmonic \( \tilde{h} \) and \( \phi \). Using the identities (9), (15) and (16) along with (19) there follows

\[
-\Delta u = -\frac{1}{2} \nabla(x \cdot \Delta \tilde{h} + 2 \text{div} \tilde{h}) - \frac{1}{2} \text{rot}(x \times \Delta \tilde{h} + 2 \text{rot} \tilde{h}) - (1 - 2\bar{\nu})\Delta \tilde{h}
\]
\[
= -\nabla \text{div} \tilde{h} - \text{rot rot} \tilde{h} = -2 \text{rot rot} \tilde{h} = \text{rot} q,
\]
\[
\text{rot} u = \frac{1}{2} \left( \nabla \text{div}(x \times \tilde{h}) - \Delta(x \times \tilde{h}) \right) + (1 - 2\bar{\nu}) \text{rot} \tilde{h}
\]
\[
= -2\bar{\nu} \text{rot} \tilde{h} + \frac{1}{4} \nabla \left( x \cdot (-2 \text{rot} \tilde{h}) \right) = \bar{\nu}(q + 4\nabla \phi) + \frac{1}{4} \nabla (x \cdot (q + 4\nabla \phi))
\]
\[
= \bar{\nu} q + \nabla \left( 4\bar{\nu} \phi(x) + x \cdot \nabla \phi(x) + \frac{1}{4} x \cdot q(x) \right) = \bar{\nu} q.
\]

Hence \( u \) and \( q \) satisfy (1).

In the opposite direction, we assume that \( u \) and \( q \) satisfy (1). For the calculation we use repeatedly

\[
\nabla \text{div} u = \Delta u + \text{rot rot} u = -\text{rot} q + \text{rot} \bar{\nu} q = -(1 - \bar{\nu}) \text{rot} q
\]
as a consequence of (1). By \( \Delta \phi = 0 \) and (16) we have \( \Delta(x \times \nabla \phi) = 0 \) and

\[
\Delta \tilde{h} = \frac{2}{1 - 4\bar{\nu}} \left( \Delta u + \frac{1}{4} \Delta(x \times q) - \frac{1}{4(1 - \bar{\nu})} \Delta(\text{div} u) + \Delta(x \times \nabla \phi) \right)
\]
\[
= \frac{2}{1 - 4\bar{\nu}} \left( -\text{rot} q + \frac{1}{4}(x \Delta q + 2 \text{rot} q) - \frac{1}{4(1 - \bar{\nu})} (x \text{div} u + 2 \nabla \text{div} u) \right)
\]
\[
= \frac{2}{1 - 4\bar{\nu}} \left( -\text{rot} q + \frac{1}{2} \text{rot} q - \frac{1}{4(1 - \bar{\nu})} (x \text{div} \text{rot} q - 2(1 - \bar{\nu}) \text{rot} q) \right) = 0,
\]
that is, the function $\tilde{h}$ defined by (6) is harmonic. We calculate by (8), (12), (13) and (19) that
\[
\text{rot } \tilde{h} = \frac{2}{1 - 4\bar{v}} \left( \bar{v} q + \frac{1}{4} x \cdot \text{div} q + x \times \text{rot} q - \nabla (x \cdot q) + \frac{1}{4(1 - \bar{v})} x \times \nabla \text{div} u \\
+ (-\nabla (\phi + x \cdot \nabla \phi) + x \Delta \phi) \right)
\]
\[
= \frac{2}{1 - 4\bar{v}} \left( \bar{v} q + \frac{1}{4} x \times \text{rot} q - \frac{1}{4} q - \frac{1}{4} \nabla (x \cdot q) - \frac{1}{4(1 - \bar{v})} x \times (1 - \bar{v}) \text{rot} q \right) \\
- \nabla (\phi + x \cdot \nabla \phi) = \frac{2}{1 - 4\bar{v}} \left( \frac{4\bar{v} - 1}{4} q - \nabla (\phi + x \cdot \nabla \phi + \frac{1}{4} x \cdot q) \right)
\]
\[
= -\frac{1}{2} q - 2\nabla \phi.
\]
From this we obtain (5) by rearrangement. We also calculate by (10) and (11) that
\[
\text{div } \tilde{h} = \frac{2}{1 - 4\bar{v}} \left( \text{div } u - \frac{1}{4} x \cdot \text{rot} q - \frac{1}{4(1 - \bar{v})} (3 \text{div } u + x \cdot \nabla \text{div } u) \right) \\
= \frac{2}{1 - 4\bar{v}} \left( \frac{1 - 4\bar{v}}{4(1 - \bar{v})} \text{div } u - \frac{1}{4} x \cdot \text{rot} q - \frac{1}{4(1 - \bar{v})} x \cdot \nabla (-1 - \bar{v}) \text{rot } q) \right) \\
= \frac{1}{2(1 - \bar{v})} \text{div } u.
\]
We compose the expressions for $\text{rot } \tilde{h}$ and $\text{div } \tilde{h}$ as
\[
\frac{1}{2} (x \text{div } \tilde{h} + x \times \text{rot } \tilde{h}) = \frac{1}{4(1 - \bar{v})} x \text{div } u - \frac{1}{4} x \times q - x \times \nabla \phi.
\]
On the other hand we have by (6) that
\[
\left( \frac{1}{2} - 2\bar{v} \right) \tilde{h} = u + \frac{1}{4} x \times q - \frac{1}{4(1 - \bar{v})} x \text{div } u + x \times \nabla \phi.
\]
Adding the latter two expressions and using (8) there follows
\[
\frac{1}{2} \nabla (x \cdot \tilde{h}) + \frac{1}{2} \text{rot } (x \times \tilde{h}) + (1 - 2\bar{v})\tilde{h} = \frac{1}{2} (x \text{div } \tilde{h} + x \times \text{rot } \tilde{h}) + \left( \frac{1}{2} - 2\bar{v} \right) \tilde{h} = u,
\]
that is, we have obtained (4) as intended.

Finally we prove the uniqueness of the harmonic function $\tilde{h}$ in the representation (4). Assume that the expression (4) for $u$ is valid for two harmonic functions $\tilde{h}_1$ and $\tilde{h}_2$. Subtracting these two expressions and setting $\tilde{h} = \tilde{h}_1 - \tilde{h}_2$ gives
\[
\frac{1}{2} \nabla (x \cdot \tilde{h}) + \frac{1}{2} \text{rot } (x \times \tilde{h}) + (1 - 2\bar{v}) \tilde{h} = 0
\]
for the harmonic function $\tilde{h}$. Taking first the divergence and then the rotation of (20) gives
\[
0 = \frac{1}{2} \Delta (x \cdot \tilde{h}) + (1 - 2\bar{v}) \text{div } \tilde{h} = 2(1 - \bar{v}) \text{div } \tilde{h}
\]
and
\[
0 = \frac{1}{2} \text{rot rot } (x \times \tilde{h}) + (1 - 2\bar{v}) \text{rot } \tilde{h} = \frac{1}{2} \nabla \text{div } (x \times \tilde{h}) - \frac{1}{2} \Delta (x \times \tilde{h}) + (1 - 2\bar{v}) \text{rot } \tilde{h} \\
= -\frac{1}{2} \nabla (x \cdot \text{rot } \tilde{h}) - \text{rot } \tilde{h} + (1 - 2\bar{v}) \text{rot } \tilde{h} = -\frac{1}{2} \left( 4\bar{v} \text{ rot } \tilde{h} + \nabla (x \cdot \text{rot } \tilde{h}) \right).
\]
The first equation implies $\text{div } \tilde{h} = 0$ by $\bar{v} \neq 1$, while from the second equation we obtain
\[
4\bar{v} x \cdot \text{rot } \tilde{h} + x \cdot \nabla (x \cdot \text{rot } \tilde{h}) = 0.
\]
This equation is very similar to (19). Its solution is \( x \cdot \text{rot} \, \tilde{h} = 0 \), which we substitute into the second of the latter equations. There follows

\[
\tilde{\nu} \text{rot} \, \tilde{h} = 0,
\]

which implies \( \text{rot} \, \tilde{h} = 0 \) by \( \tilde{\nu} > 0 \). Comparing (8) and (20) gives

\[
-\frac{1}{2} (-\tilde{h} + x \text{div} \, \tilde{h} + x \times \text{rot} \, \tilde{h}) + (1 - 2\tilde{\nu})\tilde{h} = \left(\frac{1}{2} - 2\tilde{\nu}\right)\tilde{h},
\]

which on the other hand implies \( \tilde{h} = 0 \) by \( \tilde{\nu} \neq \frac{1}{4} \). Hence we have uniqueness for the harmonic function \( \tilde{h} \) in the representation formula (4).

\[\Box\]

Remark 1. The Lamé constants usually satisfy \( \mu > 0 \) and \( \lambda + \frac{2}{3} \mu > 0 \), which mean \( \frac{\lambda}{\mu} > -\frac{2}{3} \).

This implies \( 1 < \tilde{\nu} < 4 \) for the parameter \( \tilde{\nu} = \frac{\lambda + 2\mu}{\lambda + \mu} \) in Theorem 1. Hence, the assumptions for the parameter \( \tilde{\nu} \) in Theorem 1 are satisfied for every pair of Lamé constants.

Remark 2. An analogous representation for the solutions of (2) in [8] is

\[
\begin{align*}
\mathbf{u}(\mathbf{x}) &= -\frac{1}{2} \nabla (\mathbf{x} \cdot \mathbf{h}(\mathbf{x})) - \frac{1}{2} \text{rot}(\mathbf{x} \times \mathbf{h}(\mathbf{x})) + (1 - 2\nu)\mathbf{h}(\mathbf{x}), \\
\mathbf{p}(\mathbf{x}) &= 2 \text{div} \, \mathbf{h}(\mathbf{x}),
\end{align*}
\]

where the unique harmonic function \( \mathbf{h} \) is defined by

\[
\mathbf{h}(\mathbf{x}) = \frac{2}{3-4\nu} \left( \mathbf{u}(\mathbf{x}) + \frac{1}{4} \mathbf{p}(\mathbf{x}) \mathbf{x} - \frac{1}{4(1-\nu)} \mathbf{x} \times \text{rot} \, \mathbf{u}(\mathbf{x}) + \mathbf{x} \times \nabla \psi(\mathbf{x}) \right)
\]

with \( \psi(\mathbf{x}) = -\frac{1}{4(1-\nu)} \int_{0}^{1} \mathbf{t} \cdot \mathbf{u}(\mathbf{t} \mathbf{x}) \, dt \). If we compare this definition of \( \psi \) to (7) and we also take into account \( \tilde{\nu} = 1 - \nu \) and the second equation in (1), then we can conclude that in fact \( \psi = \phi \). Moreover, comparing this with (1), (2), (6) and (23) gives

\[
\mathbf{h}(\mathbf{x}) = -\tilde{\mathbf{h}}(\mathbf{x}) + \frac{1}{1-4\nu} \mathbf{x} \times \mathbf{q}(\mathbf{x}).
\]

Hence the representation for solutions of (1) is in fact the same as the representation in [8], Theorem 5.3 for solutions of (2).

Remark 3. Theorem 1 of this paper (along with the related Theorem 5.3 in [8]) solves the problem of eliminating the scalar harmonic potential from the general Papkovich-Neuber representation for the solution of the Navier-Lamé equations in spatial star-shaped domains. Decisive in this regard is the solvability of equation (19) for star-shaped domains by the harmonic function (7). For the elimination of the scalar Papkovich-Neuber potential a very similar equation is considered in [5], Section 3, see equation (3.1) in [5].
Remark 4. If we impose a homogeneous Dirichlet boundary condition on the function $u$ in (1) and (2), then we can interpret these as eigenvalue problems for the Schur complement operators

$$
S_{\text{div}} = - \text{div}(-\Delta^{-1}) \nabla \quad \text{and} \quad S_{\text{rot}} = \text{rot}(-\Delta^{-1}) \text{rot}
$$

connected to the Stokes problem on a spatial star-shaped domain. They were studied for example in [7], see also the references given there.

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