Finite temperature Casimir effect for massless Majorana fermions in a magnetic field

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The zeta function regularization technique is used to study the finite temperature Casimir effect for a massless Majorana fermion field confined between parallel plates and satisfying bag boundary conditions. A magnetic field perpendicular to the plates is included. An expression for the zeta function is obtained, which is exact to all orders in the magnetic field strength, temperature and plate distance. The zeta function is used to calculate the Helmholtz free energy of the Majorana field and the pressure on the plates, in the case of weak magnetic field and strong magnetic field. In both cases, simple analytic expressions are obtained for the free energy and pressure which are very accurate and valid for all values of the temperature and plate distance.

PACS numbers: 03.70.+k, 11.10.Wx, 12.20.Ds, 12.39.Ba

I. INTRODUCTION

The Casimir effect is a purely quantum phenomena where an attractive or repulsive force occurs between electrically neutral conducting plates in vacuum. The effect was first discovered by Casimir [1], who calculated the attractive electromagnetic force between two parallel conducting plates, and can be regarded as a quantitative proof of the quantum fluctuations of the electromagnetic field. The repulsive Casimir effect was discovered by Boyer some time later [2], when he showed that if the electromagnetic field is confined inside a perfectly conducting sphere, the wall of the sphere is subject to a repulsive force. The first experimental evidence of the Casimir force was obtained more than fifty years ago by Sparnaay [3] and, since then, many greatly improved experimental observations of this effect have been reported. For a comprehensive review of these experiments see the review article by Bordag et al. [4].

Since Casimir forces have many applications, from nanotubes and nanotechnology [5, 6], to branes and compactified extra dimensions [7–23] and to string theory [24, 25], a great deal of effort has gone into studying the Casimir effect and its generalization to quantum fields other than the electromagnetic: fermions, bosons and other scalar fields have all been investigated extensively [4].

It is well known that the Casimir effect is very sensitive to the boundary conditions for all types of quantum fields, and the most used ones are Dirichlet and Neuman boundary conditions on the plates. These boundary conditions however cannot be used in the case of fermion fields or fields with spin in general [26] and therefore, for fermions, the bag boundary conditions are used, which originally were introduced to find a solution to confinement [27]. In this work I will use the bag boundary conditions for a Majorana fermion field confined between two parallel plates.

Majorana fermions appear in many different areas of physics: in some models of extra dimensions a Majorana fermion, the Kaluza-Klein neutrino, is a leading WIMP candidate [28]. In the field of superconductivity, a Majorana bound state is theoretically predicted in rotating superfluid $^3$He - A between parallel plates and in the presence of a magnetic field [29]. While the Casimir effect for Majorana fermion fields has been studied in vacuum [30, 31] and recently at finite temperature [32], a study that considers magnetic effects in vacuum and thermal and magnetic effects combined has not appeared in the literature. This paper will therefore investigate the Casimir effect for Majorana fermion fields at finite temperature and in the presence of a magnetic field.

Casimir effect calculations must address carefully the issue of the vacuum energy regularization. Many regularization techniques are available nowadays, and two of them have been applied frequently and successfully to the Casimir effect: the cutoff method often used in various piston configurations [31, 33] and the zeta function technique [34–36]. My choice for this paper is the zeta function technique, which is very powerful and is also used in the computation of effective actions [37, 38]. I will apply this regularization method to calculate the free energy and pressure on the plates for a Majorana fermion field with bag boundary conditions confined between two parallel plates, at a distance $a$ from each other. I assume this system to be in thermal equilibrium with a heat reservoir at finite temperature $T$, and will use the imaginary time formalism of finite temperature field theory, which is suitable for a system in thermal equilibrium. A magnetic field $\vec{B}$ is present in the region between the plates and is perpendicular to the plates.

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In Section II I calculate the zeta function for this system, exact to all orders in $eB, T$ and $a$, where $e$ is the Majorana field charge. In Section III I use the zeta function obtained in the previous section, to calculate the Helmholtz free energy of the Majorana fermion field and the pressure on the plates and obtain simple analytic expressions for these quantities in the case of weak magnetic field ($eB \ll T^2, a^{-2}$), and of strong magnetic field ($eB \gg T^2, a^{-2}$). A discussion of my results is presented in Section IV.

II. ZETA FUNCTION EVALUATION

Using the imaginary time formalism of finite temperature field theory I write the partition function $Z$ for a fermionic system in thermal equilibrium at finite temperature $T$

$$Z = N \int_{\text{Antiperiodic}} D\bar{\psi} D\psi \exp\left(\int_0^\beta d\tau \int d^3x L\right)$$ (1)

where $L$ is the Lagrangian density for the fermionic system, $N$ is a constant and ’antiperiodic’ means that this functional integral is evaluated over field configurations satisfying

$$\psi(x, y, z, 0) = -\psi(x, y, z, \beta),$$ (2)

where $\beta = 1/T$ is the antiperiodic length in the Euclidean time axis. In addition to the boundary conditions given by (2), I impose bag boundary conditions [27] for Majorana fermions between two plates. In 3-dimensional space with two very large parallel plates perpendicular to the $z$-axis and located at $z = 0$ and $z = a$, the MIT bag boundary conditions constrain the Majorana fermion fields to the region between the plates. These boundary conditions are expressed in covariant form as

$$n^\mu \bar{\psi} \gamma_{\mu} \psi = 0$$ (3)

where $n^\mu = (0, \vec{n})$ and $\vec{n}$ is the vector normal to the surface of the plates and directed towards the interior of the slab region. The above boundary conditions prevent the flow of fermion current out of the slab region. In the slab region there is also a uniform magnetic field pointing in the $z$ direction, $\vec{B} = (0, 0, B)$. The Majorana fields have charge $e$ and will interact with the magnetic field.

The Helmholtz free energy $F$ and the partition function $Z$ are related by

$$F = -\beta^{-1} \log Z,$$ (4)

therefore I evaluate the functional integral (1) and, after the straightforward gamma algebra, find

$$2 \log Z = \log \det (-D^{(\pm)}_{E}|F_a) + \log \det (-D^{(+)}_{E}|F_a)$$ (5)

where the symbol $F_a$ indicates the set of functions which satisfy boundary conditions (2) and (3), and the operators $D^{(\pm)}_E$ are defined as:

$$D^{(\pm)}_E = \partial_\tau^2 + \partial_z^2 - (\vec{p} - e\vec{A})_\perp^2 \pm eB,$$ (6)

where $\vec{A}$ is the electromagnetic vector potential, the subscript $E$ indicates Euclidean time, and I used the notation $\vec{p}_\perp = (p_x, p_y, 0)$.

The zeta function technique allows me to use the eigenvalues of $-D^{(\pm)}_E$ to evaluate $\log Z$. The bag boundary conditions (3) are satisfied only if the allowed values for the momentum in the $z$-direction are

$$p_z = \frac{\pi}{a} \left( l + \frac{1}{2} \right),$$ (7)

where $l \in \{0, 1, 2, 3, \ldots\}$, and therefore the eigenvalues of $-\partial^2_\tau - \partial^2_z$ whose eigenfunctions satisfy (2) and (3) are:

$$\frac{\pi^2}{a^2} \left( l + \frac{1}{2} \right)^2 + \frac{4\pi^2}{\beta^2} \left( m + \frac{1}{2} \right)^2,$$ (8)
where \( l \in \{0, 1, 2, 3, \ldots \} \) and \( m \in \{0, \pm 1, \pm 2, \pm 3, \ldots \} \). The spectrum of the operator \((\vec{p} - e\vec{A})_z^2\) is well known from

one-particle quantum mechanics, and its eigenvalues are

\[
2eB \left( n + \frac{1}{2} \right),
\]

(9)

with \( n \in \{0, 1, 2, 3, \ldots \} \). Using the eigenvalues (8) and (9), I construct the zeta functions \( \zeta \left( s, -D_{E}^{(+)} \right) \) and \( \zeta \left( s, -D_{E}^{(-)} \right) \), which are given by

\[
\zeta \left( s, -D_{E}^{(\pm)} \right) = L^2 \sum_{l=0}^{\infty} \sum_{m=-\infty}^{\infty} \frac{eB}{2\pi} \left( \frac{\pi^2}{a^2} \left( l + \frac{1}{2} \right)^2 + \frac{4\pi^2}{\beta^2} \left( m + \frac{1}{2} \right)^2 + 2eB \left( n + \frac{1}{2} \right) \pm eB \right)^{-s},
\]

(10)

where \( L^2 \) is the area of the plates. I add them to obtain

\[
\zeta(s) = \zeta \left( s, -D_{E}^{(+)} \right) + \zeta \left( s, -D_{E}^{(-)} \right),
\]

(11)

and find immediately, using the zeta function technique

\[
2 \log Z = -\zeta'(0).
\]

(12)

Using the following identity

\[
z^{-s} = \frac{1}{\Gamma(s)} \int_{0}^{\infty} dt \ t^{s-1} e^{-zt},
\]

(13)

I rewrite \( \zeta(s) \) as

\[
\zeta(s) = \frac{L^2}{2\pi} \frac{1}{\Gamma(s)} \sum_{l=0}^{\infty} \sum_{m=-\infty}^{\infty} \int_{0}^{\infty} dt \ t^{s-2} e^{Bt} \coth eBt \exp \left[ -\frac{\pi^2}{a^2} \left( l + \frac{1}{2} \right)^2 t - \frac{4\pi^2}{\beta^2} \left( m + \frac{1}{2} \right)^2 t \right],
\]

(14)

where I also used

\[
\sum_{n=0}^{\infty} \left[ e^{-2nz} + e^{-2(n+1)z} \right] = \coth z.
\]

(15)

After applying the Poisson resummation formula (39) to the two sums, I obtain

\[
\zeta(s) = \frac{V}{2\pi^2} \frac{\beta}{\Gamma(s)} \int_{0}^{\infty} dt \ t^{s-3} e^{Bt} \coth eBt \left( \frac{1}{2} + \sum_{l=1}^{\infty} (-1)^l e^{-\alpha^2 t^2 / l} \right) \left( \frac{1}{2} + \sum_{m=1}^{\infty} (-1)^m e^{-\beta^2 m^2 / 4t} \right),
\]

(16)

where \( V = L^2 a \) is the volume of the slab. It is evident from (18) that the Poisson resummation formula allows me to naturally divide \( \zeta(s) \) into four parts

\[
\zeta(s) = \zeta_{HE,0}(s) + \zeta_{HE,\beta}(s) + \zeta_{C,0}(s) + \zeta_{C,\beta}(s),
\]

(17)

two of which are vacuum parts that do not depend on \( \beta \) except for the overall multiplicative factor, while the other two are finite temperature corrections that vanish when \( \beta \to \infty \). The four parts of \( \zeta(s) \) are:

\[
\zeta_{HE,0}(s) = \frac{L^2}{8\pi^2} \frac{a\beta}{\Gamma(s)} \int_{0}^{\infty} dt \ t^{s-3} e^{Bt} \coth eBt,
\]

(18)

\[
\zeta_{HE,\beta}(s) = \frac{L^2}{4\pi^2} \frac{a\beta}{\Gamma(s)} \int_{0}^{\infty} dt \ t^{s-3} e^{Bt} \coth eBt \left( \sum_{m=1}^{\infty} (-1)^m e^{-\beta^2 m^2 / 4t} \right),
\]

(19)

\[
\zeta_{C,0}(s) = \frac{L^2}{4\pi^2} \frac{a\beta}{\Gamma(s)} \int_{0}^{\infty} dt \ t^{s-3} e^{Bt} \coth eBt \left( \sum_{l=1}^{\infty} (-1)^l e^{-\alpha^2 l^2 / t} \right),
\]

(20)
\[ \zeta_{C,\beta}(s) = \frac{L^2}{2\pi^2} \frac{a\beta}{\Gamma(s)} \int_0^\infty dt t^{-3} e^{Bt} \coth e^{Bt} \left( \sum_{l=1}^{\infty} (-1)^l e^{-a^2l^2/t} \right) \left( \sum_{m=1}^{\infty} (-1)^m e^{-\beta^2m^2/4t} \right), \] (21)

and each one will contribute differently to the free energy. I will show in the next section that the contribution of \( \zeta_{HE,0} \) to the free energy is the opposite of the unrenormalized Heisenberg-Euler effective lagrangian [41] multiplied by the volume of the slab \( V \), while the contribution of \( \zeta_{HE,\beta} \) is the opposite of the finite temperature correction to the Heisenberg-Euler effective lagrangian, also multiplied by \( V \). The two remaining pieces, \( \zeta_{C,0} \) and \( \zeta_{C,\beta} \), will be shown to be the main contributors to the vacuum Casimir energy and its finite temperature correction.

### III. FREE ENERGY AND PRESSURE

I combine [4] and [12] to obtain the free energy \( F \) in terms of the zeta function

\[ F = \frac{1}{2\beta} \zeta'(0). \] (22)

The derivatives of the four parts of the zeta function are obtained easily by taking advantage of the useful fact [40] that, for a well behaved \( G(s) \), the derivative of \( G(s) / \Gamma(s) \) at \( s = 0 \) is simply \( G(0) \). I find

\[ \zeta_{HE,0}'(0) = \frac{V\beta}{8\pi^2} \int_0^\infty dt t^{-3} e^{Bt} \coth e^{Bt}, \] (23)

and its contribution to the free energy is

\[ F_{HE,0} = \frac{V}{16\pi^2} \int_0^\infty dt t^{-3} e^{Bt} \coth e^{Bt}, \] (24)

where we recognize immediately that \( F_{HE,0} = -V L_{HE,0} \), where \( L_{HE,0} \) is the unrenormalized Heisenberg-Euler effective lagrangian for massless Majorana fermions. Similarly I find

\[ \zeta_{HE,\beta}'(0) = \frac{V\beta}{4\pi^2} \int_0^\infty dt t^{-3} e^{Bt} \coth e^{Bt} \left( \sum_{m=1}^{\infty} (-1)^m e^{-\beta^2m^2/4t} \right), \] (25)

whose contribution to the free energy is

\[ F_{HE,\beta} = \frac{V}{8\pi^2} \int_0^\infty dt t^{-3} e^{Bt} \coth e^{Bt} \left( \sum_{m=1}^{\infty} (-1)^m e^{-\beta^2m^2/4t} \right), \] (26)

and \( F_{HE,\beta} = -V L_{HE,\beta} \), where \( L_{HE,\beta} \) is the finite temperature correction to the Heisenberg-Euler effective lagrangian [42] for massless Majorana fermions. Notice that \( F_{HE,0} \) and \( F_{HE,\beta} \) do not depend on the plates distance \( a \) except for a multiplicative factor of \( a \) contained in \( V \). The remaining two parts of the zeta function, \( \zeta_{C,0} \) and \( \zeta_{C,\beta} \), are the ones mostly contributing to the Casimir force, and their derivatives are given by

\[ \zeta_{C,0}'(0) = \frac{V\beta}{4\pi^2} \int_0^\infty dt t^{-3} e^{Bt} \coth e^{Bt} \left( \sum_{l=1}^{\infty} (-1)^l e^{-a^2l^2/t} \right), \] (27)

and

\[ \zeta_{C,\beta}'(0) = \frac{V\beta}{2\pi^2} \int_0^\infty dt t^{-3} e^{Bt} \coth e^{Bt} \left( \sum_{l=1}^{\infty} (-1)^l e^{-a^2l^2/t} \right) \left( \sum_{m=1}^{\infty} (-1)^m e^{-\beta^2m^2/4t} \right). \] (28)

Their contributions to the free energy are

\[ F_{C,0} = \frac{\zeta_{C,0}(0)}{2\beta} \] (29)

and

\[ F_{C,\beta} = \frac{\zeta_{C,\beta}(0)}{2\beta} \] (30)
respectively. It is not possible to evaluate (28) in closed form for arbitrary values of \( B, a \) and \( \beta \), but it is possible to find simple expressions for (27) and (28) when the magnetic field is weak, \( eB \ll a^{-2}, \beta^{-2} \), and when \( B \) is strong, \( eB \gg a^{-2}, \beta^{-2} \).

In the case of weak magnetic field I can set
\[
eB t \coth eB t \approx 1 + \frac{1}{3} (eB t)^2
\]
and, after substituting (31) into (27), integrate by changing variable from \( t \) to \( \frac{1}{x} \), to find
\[
\zeta'_{C,0}(0) = \frac{V \beta}{4\pi^2} \left[ -\frac{7}{8} a^{-4} \zeta_R(4) \Gamma(2) + \frac{(eB)^2}{3} \frac{\partial f(s = 0, a)}{\partial s} \right],
\]
where \( \zeta_R(4) = \pi^4/90 \) is the Riemann zeta function of number theory, \( \Gamma(2) = 1 \) is the Euler gamma function and
\[
f(s, a) = a^{2s} (2^{2s+1} - 1) \zeta_R(-2s) \frac{\Gamma(-s)}{\Gamma(s)}.
\]
I find
\[
\frac{\partial f(s = 0, a)}{\partial s} = \gamma_E + 2\zeta'_R(0) + \ln(4a),
\]
where \( \gamma_E = 0.5772 \) is the Euler-Mascheroni constant and \( \zeta'_R(0) = -0.9189 \) is the derivative of the Riemann zeta function. The interesting numerical fact
\[
\gamma_E + 2\zeta'_R(0) + \ln 4 \approx \frac{1}{8}
\]
allows me to write
\[
\frac{\partial f(s = 0, a)}{\partial s} = \ln a + \frac{1}{8},
\]
and, after inserting (36) into (32) and then (32) into (29), I find
\[
F_{C,0} = -\left( \frac{7}{8} \right) \frac{\pi^2}{720} a^4 + \frac{(eB)^2}{24\pi^2} V \left( \ln a + \frac{1}{8} \right),
\]
which, for \( B = 0 \), is the vacuum Casimir energy for massless Majorana fermions with bag boundary conditions and is in agreement with [43], and with [31], where the same result is obtained by a different method, using a piston model to derive the vacuum Casimir energy for Majorana fermions. Notice that the second term in (37) is the correction to the Casimir energy in vacuum due to the weak magnetic field.

In the case of weak magnetic field I insert (31) into (28) and integrate as described previously, to obtain
\[
\zeta'_{C,\beta}(0) = \frac{V \beta}{2\pi^2} \left[ h(a, \beta) \Gamma(2) + \frac{(eB)^2}{3} \frac{\partial g(s = 0, a, \beta)}{\partial s} \right],
\]
where the functions \( g(s, a, \beta) \) and \( h(a, \beta) \) are given by
\[
g(s, a, \beta) = \left[ (1 + 4^{s+1}) E_2 \left( -s; a^2, \frac{\beta^2}{4} \right) - 2E_2 \left( -s; a^2, \beta^2 \right) - 2E_2 \left( -s; 4a^2, \frac{\beta^2}{4} \right) \right] \frac{\Gamma(-s)}{\Gamma(s)},
\]
\[
h(a, \beta) = \frac{5}{4} E_2 \left( 2; a^2, \frac{\beta^2}{4} \right) - 2E_2 \left( 2; a^2, \beta^2 \right) - 2E_2 \left( 2; 4a^2, \frac{\beta^2}{4} \right),
\]
and I have expressed the double sum in terms of Epstein functions [36, 40, 44] which for any positive integer \( N \) are defined by
\[
E_N \left( s; a_1, a_2, ..., a_N \right) = \sum_{n_1=1}^{\infty} \sum_{n_2=1}^{\infty} ... \sum_{n_N=1}^{\infty} \frac{1}{(a_1 n_1^2 + a_2 n_2^2 + ... + a_N n_N^2)^s}.
\]
The Epstein functions can be analytically continued to meromorphic functions in the complex plane \([36, 40]\), and for \(N = 2\) their analytic continuation is given by:

\[
E_2(s; a_1, a_2) = -\frac{a_1^{-s}}{2} + 2^{\frac{s}{2}} \pi \Gamma \left( \frac{s}{2} \right) E_1 \left( s - \frac{1}{2}; a_1 \right) + \frac{2\pi}{\Gamma(s) a_2^{s+1}} \sum_{n=1}^{\infty} \frac{m^{s}}{\Gamma(n+1)} K_{s-\frac{1}{2}} \left( 2\pi mn \frac{a_1}{a_2} \right),
\]

where \(K_\nu(z)\) are modified Bessel functions. This analytic continuation of the Epstein functions allows me to find

\[
h(a, \beta) = \frac{7\pi^4}{90} \frac{1}{\beta^4} + \frac{2\pi^2}{a^2 \beta^2} \sum_{n,m=1}^{\infty} \frac{\Gamma(n+m+1)}{\Gamma(n) \Gamma(m+1)} \left[ \frac{5}{\sqrt{2}} K_{\frac{1}{2}} \left( \frac{\pi mn \beta}{a} \right) - K_{\frac{1}{2}} \left( \frac{\pi mn \beta}{2a} \right) - 2K_{\frac{1}{2}} \left( \frac{2\pi mn \beta}{a} \right) \right] - \frac{1}{2} \ln \left( \frac{\beta}{2} \right) + \frac{1}{8},
\]

where I used \([35]\). With the help of the following

\[
K_{n+\frac{1}{2}}(z) = \sqrt{\frac{\pi}{2z}} \sum_{k=1}^{n} \frac{(n+k)!}{k!(n-k)!(2z)^k}
\]

for the modified Bessel functions of half-integral order, I find

\[
h(a, \beta) = \frac{7\pi^4}{90} \frac{1}{\beta^4} + \frac{2\pi^2}{a^2 \beta^2} \sum_{n,m=1}^{\infty} \frac{1}{m^3} \left[ \frac{5}{2} \left( 1 + \frac{\pi mn \beta}{a} \right) e^{-\frac{\pi mn \beta}{a}} - \left( 2 + \frac{\pi mn \beta}{a} \right) e^{-\frac{\pi mn \beta}{2a}} - \left( 1 + \frac{\pi mn \beta}{a} \right) e^{-\frac{\pi mn \beta}{a}} \right],
\]

and

\[
\frac{\partial g(s=0, a, \beta)}{\partial s} = 2 \sum_{n,m=1}^{\infty} \frac{1}{m^3} \left[ \frac{5}{2} e^{-\frac{\pi mn \beta}{a}} - e^{-\frac{\pi mn \beta}{2a}} - e^{-\frac{\pi mn \beta}{a}} \right] + \frac{1}{2} \ln \left( \frac{\beta}{2} \right) + \frac{1}{8}.
\]

From \([46]\) and \([47]\) I can easily obtain the low temperature expression of \(h\) and \(\frac{\partial h}{\partial s}\). For \(aT \ll 1\), I find

\[
h(a, \beta) = \frac{7\pi^4}{90} \frac{1}{\beta^4} + \frac{2\pi^2}{a^2 \beta^2} \left( 1 + \frac{2a}{\pi \beta} \right) e^{-\frac{\pi a}{2\beta}},
\]

and

\[
\frac{\partial g(s=0, a, \beta)}{\partial s} = -2e^{-\frac{\pi a}{2\beta}} - \frac{1}{2} \ln \left( \frac{\beta}{2} \right) + \frac{1}{8}.
\]

I insert \([48]\) and \([49]\) into \([48]\), then \([48]\) into \([30]\) and find immediately the finite temperature correction to the Casimir energy in the case of weak magnetic field and low temperature. For \(eB \ll T^2 \ll a^{-2}\), I obtain

\[
F_{c, \beta} = \frac{7\pi^2}{360} \frac{V}{\beta^4} - \frac{V}{2a^2 \beta^2} \left( 1 + \frac{2a}{\pi \beta} \right) e^{-\frac{\pi a}{2\beta}} - \frac{(eB)^2}{24\beta^2} V \left[ 4e^{-\frac{\pi a}{2\beta}} \ln \left( \frac{\beta}{2} \right) + \frac{1}{8} \right].
\]

In order to calculate the Helmholtz free energy, I need to add the contribution of the Heisenberg-Euler effective lagrangian \([24]\) and of its finite temperature correction \([25]\) to the Casimir energy and its thermal correction. In the case of weak magnetic field, \(eB \ll T^2 \ll a^{-2}\), these contributions are \([39, 43, 46]\)

\[
F_{HE, \beta} + F_{HE, \beta} = \frac{7\pi^2}{360} \frac{V}{\beta^4} + \frac{V(eB)^2}{48\pi^2} \left[ \ln \left( \frac{\beta eB}{4} \right) + \frac{1}{4} \right],
\]

and, once I add \([51]\) to the vacuum Casimir energy for weak magnetic field \([37]\) and to its finite temperature correction \([50]\), I obtain the Helmholtz free energy \(F\) in the limit \(eB \ll T^2 \ll a^{-2}\)

\[
F = -\left( \frac{7}{8} \right) \frac{\pi^2}{720} \frac{V}{a^2} - \frac{V}{2a^2 \beta^2} \left( 1 + \frac{2a}{\pi \beta} \right) e^{-\frac{\pi a}{2\beta}} - \frac{(eB)^2}{24\pi^2} V \left[ 4e^{-\frac{\pi a}{2\beta}} \ln \left( \frac{eBa^2}{4} \right) - \frac{1}{2} \ln \left( eBa^2 \right) \right].
\]
Notice the cancellation of the Stefan-Boltzmann term and of the temperature dependence from the logarithmic term. Due to these cancellations, I find that the leading order correction to the Casimir vacuum energy is the temperature independent and magnetic field dependent logarithmic term which, even in the case of weak magnetic field, is larger than the exponentially suppressed part, for \( Ta \ll 1 \). The pressure \( P \) on the plates is given by

\[
P = -\frac{1}{L^2} \frac{\partial F}{\partial a}.
\]  \hspace{1cm} (53)

and therefore, in the limit \( eB \ll T^2 \ll a^{-2} \), I find

\[
P = -\left(\frac{7}{8}\right) \frac{\pi^2}{240a^4} + \frac{\pi^3}{4\alpha^3} e^{-\frac{\pi a}{2\alpha}} + \frac{(eB)^2}{12\pi^2} \left[ \frac{\pi\beta}{a} e^{-\frac{\pi a}{2\alpha}} - \frac{1}{4} \ln (eBa^2) \right].
\]  \hspace{1cm} (54)

Another interesting limit, for the weak field case, is the high temperature case of \( eB \ll a^{-2} \ll T^2 \). Exploiting the fact that (28) is invariant if I exchange \( a \) and \( \beta/2 \), I write \( h \) and \( g \) in a different form, which is better suited for the high temperature expansion, \( aT \gg 1 \):

\[
h(a, \beta) = \left(\frac{7}{8}\right) \frac{\pi^4}{180a^4} + \frac{\pi^3}{2a^3} \sum_{n,m=1}^{\infty} \frac{1}{m^3} \times \left[ \frac{5}{2} \left( 1 + \frac{4\pi n \alpha a}{\beta} \right) e^{-\frac{4\pi n \alpha a}{\beta}} - \left( 2 + \frac{4\pi n \alpha a}{\beta} \right) e^{-\frac{2\pi n \alpha a}{\beta}} - \left( 1 + \frac{4\pi n \alpha a}{\beta} \right) e^{-\frac{2\pi n \alpha a}{\beta}} \right]
\]  \hspace{1cm} (55)

and

\[
\frac{\partial g(s = 0, a, \beta)}{\partial s} = 2 \sum_{n,m=1}^{\infty} \frac{1}{m} \left( \frac{5}{2} e^{-\frac{4\pi n \alpha a}{\beta}} - e^{-\frac{2\pi n \alpha a}{\beta}} - e^{-\frac{2\pi n \alpha a}{\beta}} \right) - \frac{1}{2} \left( \ln a + \frac{1}{8} \right),
\]  \hspace{1cm} (56)

from which I find immediately the high temperature expansions of \( h \) and \( \frac{\partial g}{\partial s} \)

\[
h(a, \beta) = \frac{7\pi^4}{1,440a^4} - \frac{\pi}{a^3} \left( 1 + \frac{2\pi a}{\beta} \right) e^{-\frac{2\pi a}{\beta}}
\]  \hspace{1cm} (57)

and

\[
\frac{\partial g(s = 0, a, \beta)}{\partial s} = -2e^{-\frac{2\pi a}{\beta}} - \frac{1}{2} \left( \ln a + \frac{1}{8} \right).
\]  \hspace{1cm} (58)

I use the last two equations to obtain the high temperature limit of \( F_{C,\beta} \)

\[
F_{C,\beta} = \left(\frac{7}{8}\right) \frac{\pi^2 V}{720a^4} - \frac{1}{2} \frac{V}{a^2} \left( 1 + \frac{\beta}{2\pi a} \right) e^{-\frac{2\pi \alpha}{\beta}} - \frac{\pi^2}{24\pi^2} \left( 4e^{-\frac{2\pi \alpha}{\beta}} + \ln a + \frac{1}{8} \right).
\]  \hspace{1cm} (59)

Adding (59) to (37) and (51), I obtain the free energy in the weak field and high temperature limit \( eB \ll a^{-2} \ll T^2 \)

\[
F = -\left(\frac{7}{8}\right) \frac{\pi^2 V}{45\beta^4} - \frac{1}{2} \frac{V}{\alpha^2} \left( 1 + \frac{\beta}{2\pi a} \right) e^{-\frac{2\pi \alpha}{\beta}} - \frac{(eB)^2}{24\pi^2} V \left[ 4e^{-\frac{2\pi \alpha}{\beta}} - \frac{1}{2} \ln (\frac{\beta^2 eB}{4}) - \frac{1}{8} \right].
\]  \hspace{1cm} (60)

Notice that the leading term here is the Stefan-Boltzmann term. The pressure in the weak field and high temperature limit is obtained immediately from (60)

\[
P = \left(\frac{7}{8}\right) \frac{\pi^2}{45\beta^4} - \frac{\pi^2}{a^3} e^{-\frac{2\pi \alpha}{\beta}} - \frac{(eB)^2}{24\pi^2} \left[ \frac{8\alpha}{\beta} e^{-\frac{2\pi \alpha}{\beta}} + \frac{1}{2} \ln (\frac{\beta^2 eB}{4}) + \frac{1}{8} \right],
\]  \hspace{1cm} (61)

where the dominant piece is clearly the outward pressure originating from the Stefan-Boltzmann term of the free energy.

In the case of strong magnetic field

\[
eBt \coth eBt \approx eBt
\]  \hspace{1cm} (62)
and, after substituting (62) into (27) and integrating, I find
\[ \zeta_{C,0}(0) = -\frac{V_e B}{8\pi^2} e B \Gamma(1) \zeta(2). \] (63)

From (63) I obtain immediately the vacuum piece of the Casimir energy for strong magnetic field
\[ F_{C,0} = -\frac{V_e B}{96 a^2}. \] (64)

In order to find the thermal correction to \( F_{C,0} \), I substitute (62) into (28), integrate and obtain
\[ \zeta_{C,\beta}(0) = \frac{V_e B}{\pi a^2} e B \left[ E_2 \left(1; a^2, \frac{\beta^2}{4}\right) - E_2 \left(1; a^2, \beta^2\right) - E_2 \left(1; 4a^2, \frac{\beta^2}{4}\right) \right], \] (65)

which, using the analytic continuation (42) of the Epstein functions and the expression of the modified Bessel functions (43), can be written as
\[ \zeta_{C,\beta}(0) = V e B \left[ \frac{1}{12\beta} + \frac{2}{\pi a} \sum_{n,m=1}^\infty \frac{1}{m} \left( e^{-\frac{\pi m a}{\beta}} - e^{-\frac{2\pi m a}{\beta}} - \frac{1}{2} e^{-\frac{3\pi m a}{\beta}} \right) \right]. \] (66)

From (66) I obtain \( F_{C,\beta} \), the thermal correction to the Casimir energy in the case of strong magnetic field
\[ F_{C,\beta} = \frac{V e B}{24\beta^2} + \frac{V e B}{\pi a \beta} \sum_{n,m=1}^\infty \frac{1}{m} \left( e^{-\frac{\pi m a}{\beta}} - e^{-\frac{2\pi m a}{\beta}} - \frac{1}{2} e^{-\frac{3\pi m a}{\beta}} \right). \] (67)

I now use this equation to obtain the low temperature limit of \( F_{C,\beta} \) and, for \( e B \gg a^{-2} \gg T^2 \), find
\[ F_{C,\beta} = \frac{V e B}{24\beta^2} - \frac{V e B}{\pi a \beta} e^{-\frac{\pi a}{\beta}}. \] (68)

In the case of strong magnetic field, \( e B \gg T^2 \), the contribution of the Heisenberg-Euler effective lagrangian and of its finite temperature correction to the free energy is

\[ F_{HE,0} + F_{HE,\beta} = \frac{V e B}{24 \beta^2} - \frac{V (e B)^2}{48 \pi^2} \ln \left( 2 e B \right). \] (69)

I add (69), (68) and (64) and obtain the free energy \( F \) and the pressure \( P \), in the limit \( e B \gg a^{-2} \gg T^2 \)
\[ F = -\frac{V (e B)^2}{48 \pi^2} \ln \left( 2 e B \right) - \frac{V e B}{96 a^2} - \frac{V e B}{\pi a \beta} e^{-\frac{\pi a}{\beta}}, \] (70)
\[ P = \frac{(e B)^2}{48 \pi^2} \ln \left( 2 e B \right) - \frac{1}{96} \frac{e B}{a^2} + \frac{e B}{2 a^2} e^{-\frac{\pi a}{\beta}}. \] (71)

In both of these equations, the dominant term is quadratic in the magnetic field and comes from the Heisenberg-Euler effective lagrangian.

Finally I want to find \( F \) and \( P \) in the case of strong magnetic field and high temperature, \( e B \gg T^2 \gg a^{-2} \). I again exploit the symmetry of (28) for exchange of \( a \) and \( \beta/2 \) to write (67) in an equivalent form, which is better suited for expansion in the case of \( T^2 \gg a^{-2} \)
\[ F_{C,\beta} = \frac{V e B}{96 a^2} + \frac{V e B}{\pi a \beta} \sum_{n,m=1}^\infty \frac{1}{m} \left( e^{-\frac{\pi m a}{\beta}} - e^{-\frac{2\pi m a}{\beta}} - \frac{1}{2} e^{-\frac{3\pi m a}{\beta}} \right), \] (72)

and its high temperature limit is
\[ F_{C,\beta} = \frac{V e B}{96 a^2} - \frac{V e B}{\pi a \beta} e^{-\frac{2\pi a}{\beta}}. \] (73)

By adding (73) to (68) and (64) I obtain the free energy and the pressure in the case of strong magnetic field and high temperature
\[ F = -\frac{V (e B)^2}{48 \pi^2} \ln \left( 2 e B \right) - \frac{V e B}{24 \beta^2} - \frac{V e B}{\pi a \beta} e^{-\frac{2\pi a}{\beta}}, \] (74)
\[ P = \frac{(e B)^2}{48 \pi^2} \ln \left( 2 e B \right) + \frac{1}{24} \frac{e B}{\beta^2} - \frac{2 e B}{\beta^2} e^{-\frac{2\pi a}{\beta}}. \] (75)
IV. DISCUSSION AND CONCLUSIONS

In this paper I used the zeta function regularization technique to study the finite temperature Casimir effect of a massless Majorana fermion field confined between parallel plates and in the presence of a magnetic field perpendicular to the plates. I have obtained an expression for the zeta function (10) which is exact to all orders in the magnetic field strength \(B\), and from it I have derived expressions for the Helmholtz free energy and for the pressure on the plates in the case of weak magnetic field \((eB \ll a^{-2}, T^2)\) and in the case of strong magnetic field \((eB \gg a^{-2}, T^2)\).

In the case of a weak magnetic field, I found the temperature correction to the Casimir energy to be

\[
F_{C,\beta} = \frac{V}{4\pi^2} \left[ h(a, \beta) + \frac{(eB)^2}{3} \frac{\partial g(s = 0, a, \beta)}{\partial s} \right],
\]

and obtained two equivalent expressions for \(h, (40)\) and \((45)\), which are exact to all orders in \(a\) and \(\beta\), and two equivalent expressions for \(\frac{\partial h}{\partial \beta}, (47)\) and \((50)\), which are also exact to all orders in \(a\) and \(\beta\). These expressions involve two infinite double sums, and I have been able to evaluate them numerically with very high precision. I find that, for \(0 \leq aT \leq \frac{1}{2}\), the simple low temperature expression of \(h\) that I obtained in Eq. (48) is within less than one percent of the exact value of \(h\), as given by the infinite double sum of Eq. (10). Similarly I find that, for \(\frac{1}{2} \leq aT \leq \infty\), the high temperature expression of \(h\) that I write in Eq. (57) is within less than one percent of the exact value of \(h\). A similar numerical evaluation of the exact expression of \(\frac{\partial h}{\partial \beta}\) leads me to discover that the low temperature expression \(\frac{\partial h}{\partial \beta}, (49)\), is within less than four percent of its exact value in the range \(0 \leq aT \leq \frac{1}{2}\), while the high temperature expression \(\frac{\partial h}{\partial \beta}, (55)\), is within less than four percent of its exact value in the range \(\frac{1}{2} \leq aT \leq \infty\). I summarize these findings by writing the free energy \(F\) for weak magnetic field as

\[
F = \left\{ \begin{aligned}
&\left( -\frac{(7\pi^2)}{20a^3} - \frac{2V}{27a^2}\right) e^{-\frac{2a}{\pi a}} - \frac{(eB)^2}{24\pi^2} V \left[ 4e^{-\frac{2a}{\pi a}} - \frac{1}{2} \ln\left( eBa^2 \right) \right] & \text{for } 0 \leq aT \leq \frac{1}{2} ; \\
&\left( -\frac{(7\pi^2)}{45a^3} - \frac{2V}{27a^2}\right) e^{-\frac{2a}{\pi a}} - \frac{(eB)^2}{24\pi^2} V \left[ 4e^{-\frac{2a}{\pi a}} - \frac{1}{2} \ln\left( \frac{\pi a}{2} \right) \right] & \text{for } \frac{1}{2} \leq aT \leq \infty .
\end{aligned} \right.
\]

When evaluated at \(aT = \frac{1}{2}\), both expressions of \(F\) yield the same value. Equation (77) is a simple analytic expression of \(F\) for weak magnetic field, valid for all values of the temperature \(T\) and the plate distance \(a\), and with a discrepancy of no more than a few percent from the exact value of \(F\). A similarly accurate expression of the pressure \(P\), valid for weak magnetic field and all values of \(a\) and \(T\), is obtained immediately from (77), since \(P = -\frac{1}{aT} \frac{\partial F}{\partial a}\). Notice that, if we set \(B = 0\) in (77), we obtain the free energy for the finite temperature Casimir effect of Majorana fermion fields and, opposite to what is claimed in [32], the free energy is always negative and does not become positive for \(aT > 0.37\). Even if we do not include the contribution of the thermal part of the effective Lagrangian and consider only \(F_{C,0} + F_{C,\beta}\) at \(B = 0\), we find

\[
F_{C,0} + F_{C,\beta} = \left\{ \begin{aligned}
&\left( -\frac{(7\pi^2)}{20a^3} + \frac{2V}{27a^2}\right) e^{-\frac{2a}{\pi a}} - \frac{V}{27a^2} \left( 1 + \frac{2\pi a}{\pi a} \right) e^{-\frac{2a}{\pi a}} & \text{for } 0 \leq aT \leq \frac{1}{2} ; \\
&\left( -\frac{V}{27a^2} \left( 1 + \frac{\beta}{2\pi a} \right) e^{-\frac{2a}{\pi a}} \right) & \text{for } \frac{1}{2} \leq aT \leq \infty ,
\end{aligned} \right.
\]

which is negative for \(0 \leq aT < \infty\).

In the case of a strong magnetic field, two equivalent expressions of the exact temperature correction to the Casimir energy are presented in Eqs. (67) and (72), both involving two infinite double sums. After a highly precise numerical evaluation of these two expressions, I find that the low temperature limit of \(F_{C,\beta}\) presented in (59), is accurate within less than two percent for \(0 \leq aT \leq \frac{1}{2}\), while the high temperature limit of \(F_{C,\beta}\) shown in (73) is also accurate within less than two percent for \(\frac{1}{2} \leq aT \leq \infty\). These findings allow me to write the free energy for strong magnetic field as

\[
F = \left\{ \begin{aligned}
&-\frac{V(eB)^2}{48\pi^2} \ln\left( \beta eB \right) - \frac{V eB}{2\pi a} - \frac{V eB}{\pi a^2} e^{-\frac{2a}{\pi a}} & \text{for } 0 \leq aT \leq \frac{1}{2} ; \\
&-\frac{V(eB)^2}{48\pi^2} \ln\left( \beta eB \right) - \frac{V eB}{2\pi a} - \frac{V eB}{\pi a^2} e^{-\frac{2a}{\pi a}} & \text{for } \frac{1}{2} \leq aT \leq \infty ,
\end{aligned} \right.
\]

a simple analytic expression of \(F\), valid for all values of \(T\) and \(a\), and with a discrepancy of no more than one or two percent from the exact value of \(F\). The pressure in the case of strong magnetic field is obtained immediately from (74) for all values of \(a\) and \(T\).

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