Future searches for SUSY at the LHC post Fermilab $(g - 2)_\mu$

Amin Aboubrahim$^a$*, Michael Klasen$^a$†, Pran Nath$^b$‡ and Raza M. Syed$^c$§

$^a$Institut für Theoretische Physik, Westfälische Wilhelms-Universität Münster, Wilhelm-Klemm-Straße 9, 48149 Münster, Germany

$^b$Department of Physics, Northeastern University, Boston, MA 02115-5000, USA

$^c$Department of Physics, American University of Sharjah, P.O. Box 26666, Sharjah, UAE

Submitted to the Proceedings of the US Community Study on the Future of Particle Physics (Snowmass 2021)

Abstract

We assess the future directions for the search for supersymmetry at the Large Hadron Collider in view of the new precision results on the muon anomaly by the Fermilab Collaboration. The existence of a deviation of size 4.1σ from the Standard Model prediction points to light sleptons and light weakinos in the mass range of few hundred GeV while the observation of the Higgs boson mass at $\sim 125$ GeV points to squark masses lying in the few TeV range. Thus a split sparticle spectrum is indicated. We discuss the possibility of such a split sparticle spectrum in the supergravity unified model and show that a splitting of the sfermion spectrum into light sleptons and heavy squarks naturally arises within radiative breaking of the electroweak symmetry driven by heavy gluinos ($\tilde{g}$SUGRA). We discuss the possible avenues for the discovery of supersymmetry at the LHC within this framework under the further constraint of the recent muon anomaly result from the Fermilab Collaboration. We show that the most likely candidates for early discovery of a sparticle at the LHC are the chargino, the stau, the smuon and the selectron. We present a set of benchmarks and discuss future directions for further work. Specifically, we point to the most promising channels for SUSY discovery and estimate the integrated luminosity needed for the discovery of these benchmarks at the High Luminosity LHC and also at the High Energy LHC.

*aabouibr@uni-muenster.de
†michael.klasen@uni-muenster.de
‡p.nath@northeastern.edu
§rsyed@aus.edu

arXiv:2107.06021v2 [hep-ph] 24 Aug 2021
1 Introduction

The anomalous magnetic moment is one of the most precisely computed quantities in the Standard Model and is predicted to have the value

\[ a_{\mu}^{\text{SM}} = 116591810(43) \times 10^{-11}. \] (1.1)

The recent combined Fermilab [8] and Brookhaven [9,10] measurement result gives

\[ a_{\mu}^{\text{exp}} = 116592061(41) \times 10^{-11}. \] (1.2)

The difference between the experimental and the Standard Model result is

\[ \Delta a_{\mu}^{\text{FB}} = a_{\mu}^{\text{exp}} - a_{\mu}^{\text{SM}} = 251(59) \times 10^{-11}. \] (1.3)

The above corresponds to a 4.1\(\sigma\) deviation from the Standard Model. This deviation strengthens the previous Brookhaven observation of 3.7\(\sigma\) (see, however, [11]). The implications of these for supersymmetry in the framework of supergravity unified models were analyzed in [12]. These works are motivated by the fact that it has been known since the early eighties [13,14] that a correction to the muon anomaly from supersymmetry can arise which is comparable to the electroweak correction in the Standard Model. Further, it is known that the muon anomaly is sensitive to soft parameters in supergravity models [15–17]. Thus the soft parameters directly affect the sparticle masses that enter in the supersymmetric correction to the muon anomaly, and specifically when the smuon and muon-sneutrino as well as the chargino and the neutralino are light with masses lying in the sub-TeV region, their loop corrections to the muon \(g-2\) anomaly can be sizable. The above implies that within the framework of high scale models, the 4.1\(\sigma\) deviation seen by Fermilab can be explained provided we have light weakinos and light sleptons. On the other hand, the observation of the Higgs boson mass at \(\sim 125\) GeV [18,19] indicates that the size of weak scale supersymmetry lies in the several TeV region in order to generate a large enough loop correction to the Higgs boson mass to lift its tree value that lies below the \(Z\)-boson mass to the experimentally observed value [20,21]. The supergravity based analysis discussed here proposes
a natural solution to these apparently conflicting constraints on the model. Thus a possible solution to this disparity in the scale of the sfermion mass spectrum can come about if the radiative breaking of the electroweak symmetry in supergravity (SUGRA) grand unified models is driven by a large gluino mass lying in the several TeV region while the universal scalar mass is relatively small of order a few hundred GeV. In this case squarks, which carry color, acquire large masses the size of several TeV in renormalization group (RG) running while the color neutral sleptons remain light down to the electroweak scale.

In the numerical analysis we use an artificial neural network to search for regions of the parameter space consistent with the Fermilab result within the SUGRA parameter space and the search leads to the \( \tilde{g}\text{SUGRA} \) model proposed in earlier works [22–24]. We will also discuss a specific \( SO(10) \) model, where a similar conclusion is reached [25].

2 Fermilab \( g-2 \) constraint on SUGRA models

To discuss the Fermilab constraint on SUGRA models, we utilize an artificial neural network to identify the region of the parameter space compatible with the muon anomaly result. Neural networks are efficient in analyses when one deals with a large parameter space (see, e.g., Refs. [26, 27]) and this is the case for models we investigate. Specifically we investigate the SUGRA model [28] with non-universalities [29–32] where the non-universalities are defined by the soft parameters consisting of the set \( m_0, m_1, m_2, m_3, A_0, \tan \beta \) and \( \text{sign}(\mu) \). Here \( m_0 \) defines the universal scalar mass for the sfermions and for the Higgs, \( m_1, m_2, m_3 \) are the soft masses for the \( U(1), SU(2) \) and \( SU(3) \) gauginos, \( A_0 \) is the universal scalar coupling all taken at the grand unification scale. Further, \( \tan \beta = \langle H_2 \rangle / \langle H_1 \rangle \) where \( \langle H_2 \rangle \) gives mass to the up quarks and \( \langle H_1 \rangle \) gives mass to the down quarks and the leptons.

As noted earlier there are apparent conflicting constraints that arise on the sparticle spectrum from the Higgs boson mass measurement at \( \sim 125 \) GeV and the muon \( g-2 \) anomaly measurement. The first one points to a large size of weak scale SUSY lying in the several TeV region while the muon anomaly result points to a relatively light scale for the sparticles in order to produce the desired correction to the anomaly in supersymmetry. Thus on the surface these two constraints appear to be at odds with each other. However, without any a priori assumption on the sizes of the soft parameters one can investigate regions of the parameter space which may lead to consistency using a neural network. An analysis within this framework shows that the preferred region within the prescribed parameter space is one where the universal scalar mass at the GUT scale is relatively small lying in the few hundred GeV region but the gluino mass is large and lying in the few TeV region.

As a consequence of the large gluino mass, in the RG evolution the squark masses are driven to TeV scale due to their color interactions, while the color neutral sleptons remain light. This mechanism then leads to a natural splitting between the slepton and squark masses as shown in Fig. 1 as noted in the early work of [22]. The large mass of the squarks allows one to satisfy the Higgs boson mass constraint, while the relative smallness of the slepton masses allows us to produce a significant correction to the muon \( g-2 \) anomaly indicated by the new measurements. In Fig. 2 we give a scatter plot of SUGRA model points consistent with the
Fermilab constraint with plots of $m_0, m_1, m_2$ vs. $\Delta a_{\mu}$, and in Fig. 3 we give the probability densities of $m_0, A_0/m_0, m_1$ and $m_2$ consistent with the $g-2$ constraint. We note that in these plots a significant population of models exist within the $1\sigma$ range of $\Delta a_{\mu}^{\text{FB}}$ given by Eq. (1.3). Thus the satisfaction of the constraint occurs over a region of the SUGRA parameter space rather than just over few model points. Examining the distributions of Fig. 3, we notice an apparent anticorrelation between $m_0$ and $m_2$ and general broad distributions for $m_1$ and $A_0/m_0$. The reason behind this bimodal distribution in $m_0$ and $m_2$ is the LHC constraints. Limits on sparticle masses from LHC analyses disfavor small $m_0$ while a larger number of points at higher values of $m_0$ remain viable. The opposite is true for $m_2$ where larger $m_2$ values are disfavored. In particular, for $m_2 > m_1$, the right handed slepton becomes lighter than the left handed one which becomes more constrained by LHC analyses.

![Figure 1](image)

**Figure 1:** Exhibition of a natural splitting between the slepton and squark masses in the gluino-driven radiative breaking of the electroweak symmetry. The running of the gaugino masses is also shown. The SUSY breaking scale, $Q_{\text{SUSY}}$, is defined as the geometric average of the two stop masses. Based on the analysis of [22].

### 3 Sparticle mass hierarchy under the $g - 2$ constraint

As noted above the satisfaction of the Higgs boson mass and the muon anomaly constraints leads to a mass hierarchy between the color neutral sfermions and the sfermions with color as seen from Fig. 1. A priori, a large mass gap between the sleptons and the squarks appears puzzling. This is so since sleptons and squarks belong to common multiplets in grand unification, e.g., a single generation of leptons and quarks and their superpartners lie in
one 16-plet representation of $SO(10)$, and soft breaking does not distinguish between colors. Thus the remarkable aspect of $\tilde{g}\text{SUGRA}$ is that renormalization group naturally leads to a split spectrum between sleptons and squarks, because the squarks couple to the gluino at the loop level and the large mass of the gluino thus drives their masses to large values while sleptons are largely unaffected. A display of the sparticle mass hierarchy generated by the RG running is given in Fig. 4. Here the left panel shows the light sparticle spectrum while the right panel shows the heavy sparticle masses. Corrections to the muon $g - 2$ are governed by the light spectrum of the left panel and this mass spectrum lies in the mass range of few hundred GeV while the heavy mass spectrum lies in the mass range of few TeV. Because of this split mass spectrum, the sparticles most accessible for SUSY searches at the LHC are the ones in the left panel. For the low-lying spectrum consistent with $\Delta a_{\mu}^{\text{FB}}$, the mean value (shown by the dashed line) of the chargino mass is $\sim 445$ GeV, the lightest neutralino mass $\sim 235$ GeV, the left handed slepton mass $\sim 480$ GeV, the right handed slepton mass $\sim 590$ GeV, the neutrino mass $\sim 470$ GeV and two stau mass eigenstates, $\tilde{\tau}_1$ and $\tilde{\tau}_2$, are $\sim 300$ GeV and $\sim 670$ GeV, respectively. The heavy spectrum includes the second chargino, the squarks, the gluino, the lightest stop and the CP even Higgs $H_0$. Note that the charged and CP odd Higgses have masses comparable to $H_0$. The dashed lines on either side of the mean correspond to the quartiles of this distribution.

Figure 2: A scatter plot of $m_0, m_1, m_2$ vs. $\Delta a_{\mu}$ using an artificial neural network analysis consistent with the combined Fermilab-Brookhaven result within 1$\sigma$. The model points satisfy the Higgs boson mass, the dark matter relic density and limits from dark matter direct detection experiments and the LHC sparticle mass limits.

It is also of interest to investigate the effect of the $g - 2$ constraint on specific grand unified models. As a model example, we consider the case of an $SO(10)$ model where the doublet-triplet mass splitting is obtained in a natural fashion via the missing partner mechanism.
Figure 3: Probability density vs. the range of the soft parameters, $m_0$, $A_0/m_0$ (top panels) and $m_1$, $m_2$ (bottom panels) that emerge in the artificial neural network analysis of SUGRA parameter space which generates the desired correction to the muon anomaly indicated by the Fermilab experiment.

For $SU(5)$ this is accomplished in [33, 34] while for $SO(10)$ in [24, 25, 35, 36] (for a review of other grand unified models see [37]). The $SO(10)$ model also allows for a $b - t - \tau$ unification [38, 39] along with other desirable properties. A neural network analysis again leads us to conclusions similar to the above, i.e., that the radiative electroweak symmetry breaking is gluino driven. Further, we note that in the analyses of the cases discussed above, the light sparticle spectrum falls in three classes of mass hierarchies which we label as (A), (B) and (C) as given below.

(A): Here $\tilde{\chi}^0_2, \tilde{\chi}^\pm_1$ are essentially degenerate, and $\tilde{\tau}_1$ is the next lightest supersymmetric particle (NLSP). This leads to the mass hierarchy

$$m_{\tilde{\tau}_1} < m_{\tilde{\chi}^0_2}, m_{\tilde{\chi}^\pm_1} < m_{\tilde{\ell}},$$

where $\tilde{\ell}$ stands for selectron or smuon.
Figure 4: A display of the split sparticle spectrum consisting of light weakinos, sleptons, sneutrino and staus (left panel) and the heavy chargino, squarks, gluino, lightest stop and the CP even Higgs (right panel) that emerge in the gluino-driven radiative breaking of the electroweak symmetry in $\tilde{g}$SUGRA grand unified models. Each subplot shows a probability density distribution in the particles’ masses for two cases: the region consistent with $\Delta a_{\mu}^{\text{FB}}$ (orange) and the region outside of the muon $g - 2$ error bars (blue).

(B): Here there is a reversal in the hierarchy for the first two inequalities, i.e., between $\tilde{\tau}_1$ and $\tilde{\chi}_2^0$ or $\tilde{\chi}_1^\pm$ which leads to the following possibilities

\[
m_{\tilde{\chi}_2^0}, m_{\tilde{\chi}_1^\pm} < m_{\tilde{\tau}_1} < m_{\tilde{\ell}},
\]

\[
m_{\tilde{\chi}_1^\pm} < m_{\tilde{\chi}_2^0} < m_{\tilde{\tau}_1} < m_{\tilde{\ell}},
\]

\[
m_{\tilde{\chi}_1^\pm} < m_{\tilde{\tau}_1} < m_{\tilde{\chi}_2^0} < m_{\tilde{\ell}}.
\]

(C): In this case the selectron and the smuon are lighter than the chargino and the second neutralino while the stau is the NLSP. Thus here we have

\[
m_{\tilde{\tau}_1} < m_{\tilde{\ell}} < m_{\tilde{\chi}_2^0}, m_{\tilde{\chi}_1^\pm}.
\]

The Fermilab result also puts constraints on the allowed region of CP phases arising from the soft parameters specifically the gaugino masses. Thus the gaugino masses $m_1$ and $m_2$ can be complex and one may write $m_1 = |m_1|e^{i\xi_1}$ and $m_2 = |m_2|e^{i\xi_2}$. It turns out that the electroweak corrections to the muon $g - 2$ are very sensitive to the CP phases [40–42]. As a consequence the measured $\Delta a_{\mu}^{\text{FB}}$ puts rather stringent constraints on the allowed range of the CP phases. This is shown in Fig. 5 which exhibits the disallowed region (shaded) and the allowed region (white) in the $\xi_1 - \xi_2$ plane. It is to be noted that the allowed region
is further constrained by the experimental limits on the EDM of the electron and of the neutron. Thus the allowed region resulting from imposing both constraints becomes very narrow. We label this region by a star in Fig. 5 which can be obtained via the cancellation mechanism [43]. The CP phases can also affect dark matter analyses (see, e.g., [44]) and proton stability [45] as well as the production cross sections and decays of the sparticles. These effects are also worth investigation in the future in view of the current experimental result on the possible deviation $\Delta a_{\mu}^{FB}$ from the Standard Model prediction.

![Figure 5: A display of the region that is excluded (shaded) and the region that is allowed (white) by the $g_{\mu} - 2$ constraint. The allowed region is further constrained by the limits on the EDM of the electron and the EDM of the neutron. The star in the white region represents the small region of the parameter space where both the electron EDM and the muon $g - 2$ are satisfied. Based on analysis of [12].](image)

### 4 Benchmarks for future SUSY searches at the LHC consistent with $g - 2$

As noted earlier, the Fermilab $g - 2$ measurement has important implications for the discovery of supersymmetry and future searches at colliders and elsewhere. Here we present a set of SUGRA benchmarks consistent with the recent muon $g - 2$ anomaly result to test for a potential SUSY discovery at HL-LHC and HE-LHC. The benchmarks are given in Table 1 and arranged according to the categories (A), (B) and (C) described earlier. Thus benchmark (a) belongs to case (A), benchmarks (b), (c), (d) belong to case (B) and benchmarks (e), (f), (g) belong to case (C). The benchmarks have an $m_0$ of $O(100)$ GeV while $m_3$ lies in the
several TeV range consistent with a gluino-driven radiative electroweak symmetry breaking. It is noted that benchmark (b) is taken from Ref. [12] (labeled there as (d)) while (a), (c), (d), (e), (f), (g) are from Ref. [25] (labeled there as (f), (i), (b), (h), (g), (d), respectively).

**Table 1:** SUGRA benchmarks for future SUSY searches. All masses are in GeV.

| Model | \(m_0\) | \(A_0\) | \(m_1\) | \(m_2\) | \(m_3\) | \(\tan \beta\) |
|-------|--------|--------|--------|--------|--------|------------|
| (a)   | 688    | 1450   | 852    | 634    | 8438   | 16.8       |
| (b)   | 389    | 122    | 649    | 377    | 4553   | 8.2        |
| (c)   | 452    | 648    | 624    | 346    | 4843   | 13.1       |
| (d)   | 673    | 1127   | 939    | 570    | 8833   | 8.2        |
| (e)   | 206    | 603    | 842    | 1298   | 7510   | 8.0        |
| (f)   | 106    | 22.6   | 523    | 1309   | 5240   | 9.3        |
| (g)   | 164    | 197    | 632    | 1539   | 6171   | 12.2       |

Along with the satisfaction of the recent muon \(g-2\) result, the benchmarks satisfy the Higgs boson mass, the dark matter relic density as well as the LHC mass limits constraints. We present in Table 2 the low-lying sparticle spectrum along with the SM-like Higgs boson mass, the relic density and the muon \(g-2\) anomaly calculated at the two-loop level. Here the left handed slepton is lighter than the right handed one except in benchmarks (e), (f) and (g) due to \(m_2\) being significantly greater than \(m_1\) as seen in Table 1. Furthermore, the benchmarks respect current constraints on the proton-neutralino spin-independent cross section and all allow for a multicomponent dark matter scenario.

**Table 2:** The SM-like Higgs mass, the light sparticle spectrum and the dark matter relic density \(\Omega h^2\) for the benchmarks of Table 1. Also shown is the muon \(g-2\).

| Model | \(h^0\) | \(\tilde{\ell}_L\) | \(\tilde{\ell}_R\) | \(\tilde{\nu}_L\) | \(\tilde{\tau}_1\) | \(\chi^0_1\) | \(\chi^\pm_1\) | \(\Omega h^2\) | \(\Delta a_\mu (\times 10^{-9})\) |
|-------|--------|------------------|------------------|----------------|----------------|------------|-------------|-------------|----------------|
| (a)   | 123.0  | 508.1            | 762.0            | 502.3          | 331.9          | 324.2      | 404.3      | 0.004       | 2.11           |
| (b)   | 123.4  | 305.0            | 463.0            | 295            | 251.7          | 237.4      | 237.6      | 0.002       | 2.33           |
| (c)   | 123.7  | 346.8            | 511.9            | 338.0          | 240.3          | 205.6      | 205.8      | 0.001       | 2.67           |
| (d)   | 125.3  | 422.8            | 763.8            | 415.7          | 370.4          | 337.3      | 337.6      | 0.003       | 2.14           |
| (e)   | 124.5  | 628.7            | 402.2            | 623.6          | 338.3          | 326.8      | 998.4      | 0.082       | 1.94           |
| (f)   | 123.4  | 722.8            | 262.9            | 718.2          | 206.5          | 195.5      | 1038.4     | 0.103       | 2.57           |
| (g)   | 123.9  | 856.4            | 327.4            | 852.4          | 243.5          | 240.1      | 1227       | 0.016       | 1.94           |

We investigate three main SUSY production channels at the LHC: slepton and sneutrino pair production and slepton associated production with a sneutrino. The latter has a significant cross section since it proceeds via the charged current. We calculate the production cross sections at 14 TeV and 27 TeV at the aNNLO+NNLL accuracy using Resummino-3.0 [46,47]. The results are displayed in Table 3. Notice that the contribution from the right handed sleptons becomes important in benchmarks (e), (f) and (g) where \(m_{\tilde{\ell}_R} < m_{\tilde{\ell}_L}\).
Table 3: The aNNLO+NNLL production cross-sections of slepton pair ($\tilde{\ell} = \tilde{e}, \tilde{\mu}$) and sneutrino pair as well as slepton associated production with a sneutrino at $\sqrt{s} = 14$ TeV and $\sqrt{s} = 27$ TeV for the benchmarks of Table 1. The cross section is in fb.

| Model | $\sigma(pp \to \ell_L \ell_L)$ | $\sigma(pp \to \ell_R \ell_R)$ | $\sigma(pp \to \tilde{\nu}_L \ell_L)$ | $\sigma(pp \to \tilde{\nu}_L \tilde{\nu}_L)$ |
|-------|-------------------------------|-------------------------------|-------------------------------------|-------------------------------------|
|       | 14 TeV 27 TeV                 | 14 TeV 27 TeV                 | 14 TeV 27 TeV                       | 14 TeV 27 TeV                       |
| (a)   | 1.084 4.515                   | 0.057 0.335                   | 41.78 127.80                        | 10.52 33.10                        |
| (b)   | 9.810 30.80                   | 0.656 2.54                    | 24.40 78.97                         | 4.94 15.10                         |
| (c)   | 5.805 19.31                   | 0.417 1.72                    | 10.44 37.47                         | 2.10 7.21                          |
| (d)   | 2.486 9.188                   | 0.056 0.332                   |                                     |                                     |
| (e)   | 0.388 1.915                   | 1.22 4.32                     |                                     |                                     |
| (f)   | 0.188 1.065                   | 6.91 20.15                    |                                     |                                     |
| (g)   | 0.074 0.506                   | 2.87 9.22                     |                                     |                                     |

In LHC analyses with simplified models, unit branching ratios for specific production channels are often used. However, this simplified assumption does not hold in high scale models where the branching ratios exhibit a rich diversity of final states. Thus the exclusion limits on sparticle masses do not apply directly to high scale models, but rather limits on $\sigma \times \text{BR}$ become the relevant ones. To illustrate this we display the different decay channels of the sleptons and sneutrinos in Table 4.

Table 4: The branching ratios of relevant decay channels of the left and the right handed slepton and the sneutrino for the benchmarks of Table 1.

| Model | $\ell_L \to \ell \tilde{\chi}_0^0$ | $\ell_L \to \ell \tilde{\chi}_2^0$ | $\ell_L \to \nu_\ell \tilde{\chi}_1^\pm$ | $\ell_R \to \ell \tilde{\chi}_0^0$ | $\tilde{\nu}_L \to \tilde{\chi}_1^+ \ell^-$ | $\tilde{\nu}_L \to \tilde{\chi}_0^0 \nu_\ell$ |
|-------|------------------------------------|------------------------------------|--------------------------------------|------------------------------------|--------------------------------------|--------------------------------------|
| (a)   | 22% 26% 52% 100% [-]              | 51% 24%                            |                                      |                                    |                                      |                                      |
| (b)   | 31% 7% 62% - [100%]               | 62% 30%                            |                                      |                                    |                                      |                                      |
| (c)   | 31% 6% 63% - [100%]               | 62% 31%                            |                                      |                                    |                                      |                                      |
| (d)   | 31% 6% 63% - [100%]               | 63% 32%                            |                                      |                                    |                                      |                                      |
| (e)   | 100% - - 100% [-]                | - 100%                             |                                      |                                    |                                      |                                      |
| (f)   | 100% - - 100% [-]                | - 100%                             |                                      |                                    |                                      |                                      |
| (g)   | 100% - - 100% [-]                | - 100%                             |                                      |                                    |                                      |                                      |

5 A deep neural network analysis for SUSY discovery

As can be seen from Table 4, there are numerous decay channels for the sleptons and sneutrinos which result in rich final states. Leptonic channels are the most prevalent and for these we consider final states consisting of two leptons of the same flavour and opposite sign (SFOS) and missing transverse energy (MET) due to the neutralinos (and neutrinos). We design two signal regions, where in one we require exactly one non-b-tagged jet and in the other we require at least two non-b-tagged jets. One can easily see how slepton pair production can give us exactly such final states but it may be less trivial for sneutrino pair...
production and for slepton-sneutrino associated production. In Table 3 the cross section for sneutrino pair production and slepton associated production for benchmarks (b), (c) and (d) are shown. As mentioned earlier, those benchmarks belong to case (B) where the chargino is nearly mass degenerate with the LSP. Thus it is not easy to distinguish between these particles at the LHC and so we assume that the chargino escapes undetected and adds to the MET in an event. With a 62% branching ratio for $\tilde{\nu}_L \rightarrow \tilde{\chi}_1^+ \ell^-$ (see Table 4), sneutrino pair production will result in the same final states with a significant $\sigma \times \text{BR}$. The same applies for slepton associated production with a sneutrino which has the largest cross section among the different production channels considered. The rest of the benchmarks do not have a degenerate chargino-LSP and so the decay of the chargino in those benchmarks will result in very different final states.

The signal, which involves the different production channels presented in Table 3, and the Standard Model background events are simulated at LO with MadGraph5 [48] and showered with PYTHIA8 [49] with the addition of ISR and FSR jets. Cross sections are scaled to their NLO values for the background and to aNNLO+NNLL for the signal. Detector effects are added by the help of DELPHES-3.4.2 [50].

The analysis presented here shows that a discrimination between the signal and background events can be done using deep neural network (DNN) as part of the ‘ Toolkit for Multivariate Analysis’ (TMVA) [51] framework within ROOT6 [52]. Thus the possibility of detection of light sparticles appears strong at HL-LHC and with much greater certainty at HE-LHC.

We use a set of kinematic variables to train the DNN on the signal and background events. The list of variables include

$$\begin{align*}
E_T^{\text{miss}}, \ p_T(j_1), \ p_T(\ell_1), \ p_T(\ell_2), \ p_T^{\text{ISR}}, \ M_{T2}, \ M_T^{\text{min}}, \ m_{\ell\ell}, \ \\
\Delta \phi(p_T^\ell, p_T^{\text{miss}}), \ \Delta \phi_{\text{min}}(p_T(j_i), p_T^{\text{miss}}),
\end{align*}$$

where $M_T^{\text{min}} = \min \{m_{T}(p_T^{\ell_1}, p_T^{\text{miss}}), m_{T}(p_T^{\ell_2}, p_T^{\text{miss}})\}$, $\Delta \phi(p_T^\ell, p_T^{\text{miss}})$ is the opening angle between the MET system and the dilepton system and $\Delta \phi_{\text{min}}(p_T(j_i), p_T^{\text{miss}})_{i=1,2,3}$ is the smallest opening angle between the first three leading jets in an event and the MET system. Note that $p_T^\ell = p_T^{\ell_1} + p_T^{\ell_2}$. After reconstructing the momentum of the dilepton system, we determine the angle between the dilepton system and each non-b-tagged jet in the event. We select up to two jets that are closest to the dilepton system and tag them as possible jets arising from the decay of the SUSY system while the rest are classified as ISR jets. Here $p_T^{\text{ISR}}$ denotes the total transverse momentum of all the ISR jets in an event.
We present in Table 5 the set of preselection criteria as well as the analysis cuts used on the signal and background events. The cuts are optimized depending on the category the benchmarks belong to and the center-of-mass energy.

After training the DNN on the signal and background events and making sure no overtraining has occurred, the testing phase is done on a separate set of signal and background events. The end result is the construction of a new discriminating variable called the ‘DNN response’ which tends to take values closer to one for the signal. Signal (S) and background (B) distributions in the ‘DNN response’ variable are shown in Fig. 6 for benchmarks (a) (top panels) and (c) (bottom panels) at 14 TeV and 27 TeV. In the bottom pad of each panel, we show the distribution in the signal significance $Z$ as a function of the cut on ‘DNN response’ and taking into consideration the systematics, i.e.,

\[ Z = \frac{S}{\sqrt{S + B + (\delta_SS)^2 + (\delta_BB)^2}} \]  

(5.1)

where $\delta_S$ and $\delta_B$ are the systematic uncertainties in the signal and background estimates. The recommendations on systematic uncertainties (known as ‘YR18’ uncertainties) published in the CERN’s yellow reports [53, 54] suggest an overall 20% uncertainty in the background and 10% in the SUSY signal. Notice that a $5\sigma$ value is reached at the integrated luminosities shown which can be attained at HL-LHC and HE-LHC.

| Observable | Preselection criteria (SR-2ℓ2j) | Preselection criteria (SR-2ℓ1j) |
|------------|---------------------------------|---------------------------------|
| $N_{\ell}$ (SFOS) | 2 | 2 |
| $N_{\text{non-}b\text{-tagged}}$ | $\geq 2$ | $\geq 2$ |
| $p_T(j_1)$ [GeV] | $> 20$ | $> 20$ |
| $p_T(\ell_1)$ (electron, muon) [GeV] | $> 15$, $> 10$ | $> 15$, $> 10$ |
| $E_T^{\text{miss}}$ [GeV] | $> 100$ | $> 100$ |
| $m_{\ell\ell}$ [GeV] > | Analysis cuts | Analysis cuts |
| $E_T^{\text{miss}}/p_T >$ | 136 (110) | 150 (110) |
| $\Delta\phi_{\text{min}}(p_T(j_1), p_T^{\text{miss}})$ [rad] | 1.9 (2.8) | 1.0 |
| $p_T(\ell_2)$ [GeV] > | -0.85 (1.5) | -0.85 (1.5) |
| $M_{T2}$ [GeV] > | -190 (370) | -190 (300) |
| DNN response > | 0.95 | 0.95 |

Table 5: The preselection criteria and the analysis cuts on a set of kinematic variables at 14 TeV (27 TeV) grouped by the benchmarks of Table 1 in two signal regions SR-2ℓ1j and SR-2ℓ2j. Entries with a dash (-) imply that no requirement on the variable is considered. Cuts are optimized for each center-of-mass energy.
Figure 6: Distributions in the ‘DNN response’ variable for benchmark (a) (top panels) and benchmark (c) (bottom panels) at 14 TeV (left) and 27 TeV (right). A distribution in the signal significance as a function of the cut on the ‘DNN response’ is shown in the bottom pads of each panel.

Of the two signal regions considered, we find that the most optimal one is the single jet signal region, which has also been shown to be true in previous LHC searches [55]. This is evident from Table 6 where the integrated luminosity for discovery of the benchmarks in SR-2ℓj2j is much larger than in SR-2ℓj1j after combining all production channels. The presented integrated luminosities include the effect of systematics.
Table 6: The estimated integrated luminosities, in fb$^{-1}$, for discovery of benchmarks of Table 1 at 14 TeV and 27 TeV after combining all production channels and including systematics in the signal and background.

| Model | SR-2ℓ1j | SR-2ℓ2j |
|-------|---------|---------|
|       | $\mathcal{L}$ at 14 TeV | $\mathcal{L}$ at 27 TeV | $\mathcal{L}$ at 14 TeV | $\mathcal{L}$ at 27 TeV |
| (a)   | 880     | 310     | 1262     | 694      |
| (b)   | 200     | 50      | 1860     | 715      |
| (c)   | 148     | 75      | 1887     | 1320     |
| (d)   | 425     | 252     | ...      | 2804     |
| (e)   | 1040    | 232     | 1738     | 1194     |
| (f)   | 730     | 152     | 2074     | 689      |
| (g)   | 970     | 202     | ...      | 1031     |

As a comparison between HL-LHC and HE-LHC and to see the effect of systematic uncertainties, we plot the estimated integrated luminosities for discovery of benchmarks (a)–(g) at 14 TeV and 27 TeV in Fig. 7. Benchmarks (b) and (c) require 200 fb$^{-1}$ and 148 fb$^{-1}$ at 14 TeV which should be attained in the coming run of LHC. The rest require more than 400 fb$^{-1}$ but are all within the reach of HL-LHC. The same benchmarks require much smaller integrated luminosities for discovery at HE-LHC.

**Figure 7:** The integrated luminosities, $\mathcal{L}$, needed for discovery of SUSY at HL-LHC and HE-LHC assuming that $\Delta a^\mu_B$ arises from SUSY loops. Values of $\mathcal{L}$ are shown before and after including the ‘YR18’ uncertainties on the signal and background.
We note that recently several works, including those within SUSY, have come out regarding an explanation of the Fermilab muon anomaly [56–91].

6 Conclusion

The recent results from the Fermilab Collaboration confirm the earlier result from the Brookhaven experiment regarding a deviation from the Standard Model prediction of the muon $g - 2$ anomaly. Specifically, the combined Fermilab and Brookhaven results give a $4.1\sigma$ deviation from the Standard Model prediction compared to $3.7\sigma$ prediction for Brookhaven which both supports the observation and further strengthens it.

In this paper we have pointed to the implications of the Fermilab result for SUSY discovery at the LHC. To this end we have analyzed the effect of the Fermilab constraint on the SUGRA parameter space. This is done in a comprehensive way with the help of neural networks where we have used the constraints of the Higgs boson mass and the relic density along with the muon $g - 2$ to delineate the region of the parameter space consistent with those constraints. The neural network analysis identifies the allowed region as the one that is generated by gluino-driven radiative breaking of the electroweak symmetry. This region of the parameter space naturally leads to a split sparticle spectrum constituted of light sleptons and light weakinos but a heavy gluino and heavy squarks which allows one to satisfy muon $g - 2$ constraint and the Higgs boson constraint within a high scale model.

We emphasize the existence of a light spectrum and their possible early discovery at the LHC. We propose a set of benchmarks relevant for Snowmass study related to the energy frontier. Further, we demonstrate that a deep neural network with the ‘DNN response’ variable constitutes a powerful discriminant to reject the Standard Model background which aids in the discovery of new physics signals. We exhibit the importance of this technique by analyzing the integrated luminosities needed for the discovery of the benchmarks, and show that some of the benchmarks can be discovered at HL-LHC with integrated luminosities as low as 150–300 fb$^{-1}$ while much smaller integrated luminosities are required at HE-LHC.

We propose that it would be fruitful if the future SUSY analyses at the LHC are carried out with the inclusion of the Fermilab $g - 2$ constraint and further such analyses need to go beyond generic simplified models to well motivated high scale models. The basic reason for such analyses is that in well motivated models the branching ratio to any production channel is almost never unity as generally assumed in simplified models. Thus dedicated studies for the production and signal analysis of sleptons and weakinos at the LHC within Snowmass study are timely.

The research of AA and MK was supported by the BMBF under contract 05H18PMCC1. The research of PN was supported in part by the NSF Grant PHY-1913328.
References

[1] T. Aoyama, N. Asmussen, M. Benayoun, J. Bijnens, T. Blum, M. Bruno, I. Caprini, C. M. Carloni Calame, M. Ce and G. Colangelo, et al. Phys. Rept. 887, 1-166 (2020) doi:10.1016/j.physrep.2020.07.006 [arXiv:2006.04822 [hep-ph]].

[2] M. Davier, A. Hoecker, B. Malaescu and Z. Zhang, Eur. Phys. J. C 80, no.3, 241 (2020) [erratum: Eur. Phys. J. C 80, no.5, 410 (2020)] doi:10.1140/epjc/s10052-020-7792-2 [arXiv:1908.00921 [hep-ph]].

[3] M. Davier, A. Hoecker, B. Malaescu and Z. Zhang, Eur. Phys. J. C 77, no.12, 827 (2017) doi:10.1140/epjc/s10052-017-5161-6 [arXiv:1706.09436 [hep-ph]].

[4] M. Davier, A. Hoecker, B. Malaescu and Z. Zhang, Eur. Phys. J. C 71, 1515 (2011) [erratum: Eur. Phys. J. C 72, 1874 (2012)] doi:10.1140/epjc/s10052-012-1874-8 [arXiv:1010.4180 [hep-ph]].

[5] A. Crivellin, M. Hoferichter, C. A. Manzari and M. Montull, Phys. Rev. Lett. 125, no.9, 091801 (2020) doi:10.1103/PhysRevLett.125.091801 [arXiv:2003.04886 [hep-ph]].

[6] A. Keshavarzi, W. J. Marciano, M. Passera and A. Sirlin, Phys. Rev. D 102, no.3, 033002 (2020) doi:10.1103/PhysRevD.102.033002 [arXiv:2006.12666 [hep-ph]].

[7] G. Colangelo, M. Hoferichter and P. Stoffer, Phys. Lett. B 814, 136073 (2021) doi:10.1016/j.physletb.2021.136073 [arXiv:2010.07943 [hep-ph]].

[8] B. Abi et al. [Muon g-2], Phys. Rev. Lett. 126, no.14, 141801 (2021) doi:10.1103/PhysRevLett.126.141801 [arXiv:2104.03281 [hep-ex]].

[9] G. W. Bennett et al. [Muon g-2], Phys. Rev. D 73, 072003 (2006) doi:10.1103/PhysRevD.73.072003 [arXiv:hep-ex/0602035 [hep-ex]].

[10] M. Tanabashi, et al., Review of Particle Physics, Phys. Rev. D98 (3) (2018) 030001. doi:10.1103/PhysRevD.98.030001

[11] S. Borsanyi, Z. Fodor, J. N. Guenther, C. Hoelbling, S. D. Katz, L. Lellouch, T. Lippert, K. Miura, L. Parato and K. K. Szabo, et al. Nature 593, no.7857, 51-55 (2021) doi:10.1038/s41586-021-03418-1 [arXiv:2002.12347 [hep-lat]].

[12] A. Aboubrahim, M. Klasen and P. Nath, [arXiv:2104.03839 [hep-ph]]. (To appear in PRD).

[13] D. A. Kosower, L. M. Krauss and N. Sakai, Phys. Lett. B 133, 305-310 (1983) doi:10.1016/0370-2693(83)90152-1

[14] T. C. Yuan, R. L. Arnowitt, A. H. Chamseddine and P. Nath, Z. Phys. C 26, 407 (1984) doi:10.1007/BF01452567

[15] J. L. Lopez, D. V. Nanopoulos and X. Wang, Phys. Rev. D 49, 366-372 (1994) doi:10.1103/PhysRevD.49.366 [arXiv:hep-ph/9308336 [hep-ph]].
U. Chattopadhyay and P. Nath, Phys. Rev. D 53, 1648-1657 (1996) doi:10.1103/PhysRevD.53.1648 [arXiv:hep-ph/9507386 [hep-ph]].

T. Moroi, Phys. Rev. D 53, 6565-6575 (1996) [erratum: Phys. Rev. D 56, 4424 (1997)] doi:10.1103/PhysRevD.53.6565 [arXiv:hep-ph/9512396 [hep-ph]].

G. Aad et al. [ATLAS], Phys. Lett. B 716, 1-29 (2012) doi:10.1016/j.physletb.2012.08.020 [arXiv:1207.7214 [hep-ex]].

S. Chatrchyan et al. [CMS], Phys. Lett. B 716, 30-61 (2012) doi:10.1016/j.physletb.2012.08.021 [arXiv:1207.7235 [hep-ex]].

S. Akula, B. Altunkaynak, D. Feldman, P. Nath and G. Peim, Phys. Rev. D 85, 075001 (2012) doi:10.1103/PhysRevD.85.075001 [arXiv:1112.3645 [hep-ph]].

A. Arbey, M. Battaglia, A. Djouadi, F. Mahmoudi and J. Quevillon, Phys. Lett. B 708, 162 (2012); H. Baer, V. Barger and A. Mustafayev, Phys. Rev. D 85, 075010 (2012); J. Ellis and K. A. Olive, Eur. Phys. J. C 72, 2005 (2012); S. Heinemeyer, O. Stal and G. Weiglein, Phys. Lett. B 710, 201 (2012);

S. Akula and P. Nath, Phys. Rev. D 87, no.11, 115022 (2013) doi:10.1103/PhysRevD.87.115022 [arXiv:1304.5526 [hep-ph]].

A. Aboubrahim and P. Nath, Phys. Rev. D 100, no.1, 015042 (2019) doi:10.1103/PhysRevD.100.015042 [arXiv:1905.04601 [hep-ph]].

A. Aboubrahim, P. Nath and R. M. Syed, JHEP 01, 047 (2021) doi:10.1007/JHEP01(2021)047 [arXiv:2005.00867 [hep-ph]].

A. Aboubrahim, P. Nath and R. M. Syed, JHEP 06, 002 (2021) doi:10.1007/JHEP06(2021)002 [arXiv:2104.10114 [hep-ph]].

J. Hollingsworth, M. Ratz, P. Tanedo and D. Whiteson, [arXiv:2103.06957 [hep-th]].

C. Balázs et al. [DarkMachines High Dimensional Sampling Group], JHEP 05, 108 (2021) doi:10.1007/JHEP05(2021)108 [arXiv:2101.04525 [hep-ph]].

A. H. Chamseddine, R. Arnowitt and P. Nath, Phys. Rev. Lett. 49 (1982) 970; P. Nath, R. L. Arnowitt and A. H. Chamseddine, Nucl. Phys. B 227, 121 (1983); L. J. Hall, J. D. Lykken and S. Weinberg, Phys. Rev. D 27, 2359 (1983). doi:10.1103/PhysRevD.27.2359

J. R. Ellis, K. Enqvist, D. V. Nanopoulos and K. Tamvakis, Phys. Lett. B 155, 381-386 (1985) doi:10.1016/0370-2693(85)91591-6

D. Feldman, Z. Liu and P. Nath, Phys. Rev. D 80, 015007 (2009) doi:10.1103/PhysRevD.80.015007 [arXiv:0905.1148 [hep-ph]].

A. S. Belyaev, S. F. King and P. B. Schaefer, Phys. Rev. D 97, no.11, 115002 (2018) doi:10.1103/PhysRevD.97.115002 [arXiv:1801.00514 [hep-ph]].
[32] A. Corsetti and P. Nath, Phys. Rev. D 64, 125010 (2001); A. Birkedal-Hansen and B. D. Nelson, Phys. Rev. D 67, 095006 (2003); G. Belanger, F. Boudjema, A. Cottrant, A. Pukhov and A. Semenov, Nucl. Phys. B 706, 411 (2005); H. Baer, A. Mustafayev, E. K. Park, S. Profumo and X. Tata, JHEP 04 (2006), 041 doi:10.1088/1126-6708/2006/04/041 [arXiv:hep-ph/0603197 [hep-ph]]; I. Gogoladze, F. Nasir, Q. Shafi and C. S. Un, Phys. Rev. D 90, no. 3, 035008 (2014) doi:10.1103/PhysRevD.90.035008; S. P. Martin, Phys. Rev. D 79, 095019 (2009) doi:10.1103/PhysRevD.79.095019

[33] A. Masiero, D. V. Nanopoulos, K. Tamvakis and T. Yanagida, Phys. Lett. B 115, 380-384 (1982) doi:10.1016/0370-2693(82)90522-6

[34] B. Grinstein, Nucl. Phys. B 206, 387 (1982) doi:10.1016/0550-3213(82)90275-9

[35] K. S. Babu, I. Gogoladze and Z. Tavartkiladze, Phys. Lett. B 650, 49-56 (2007) doi:10.1016/j.physletb.2007.02.050 [arXiv:hep-ph/0612315 [hep-ph]].

[36] K. S. Babu, I. Gogoladze, P. Nath and R. M. Syed, Phys. Rev. D 85, 075002 (2012) doi:10.1103/PhysRevD.85.075002 [arXiv:1112.5387 [hep-ph]].

[37] P. Nath and P. Fileviez Perez, Phys. Rept. 441, 191-317 (2007) doi:10.1016/j.physrep.2007.02.010 [arXiv:hep-ph/0601023 [hep-ph]].

[38] B. Ananthanarayan, G. Lazarides and Q. Shafi, Phys. Lett. B 300, 245-250 (1993) doi:10.1016/0370-2693(93)90361-K

[39] U. Chattopadhyay, A. Corsetti and P. Nath, Phys. Rev. D 66, 035003 (2002) doi:10.1103/PhysRevD.66.035003 [arXiv:hep-ph/0201001 [hep-ph]].

[40] T. Ibrahim and P. Nath, Phys. Rev. D 62, 015004 (2000) doi:10.1103/PhysRevD.62.015004 [arXiv:hep-ph/9908443 [hep-ph]].

[41] T. Ibrahim, U. Chattopadhyay and P. Nath, Phys. Rev. D 64, 016010 (2001) doi:10.1103/PhysRevD.64.016010 [arXiv:hep-ph/0102324 [hep-ph]].

[42] A. Aboubrahim, T. Ibrahim and P. Nath, Phys. Rev. D 94, no.1, 015032 (2016) doi:10.1103/PhysRevD.94.015032 [arXiv:1606.08336 [hep-ph]].

[43] T. Ibrahim and P. Nath, Phys. Rev. D 58, 111301 (1998) [erratum: Phys. Rev. D 60, 099902 (1999)] doi:10.1103/PhysRevD.58.111301 [arXiv:hep-ph/9807501 [hep-ph]].

[44] U. Chattopadhyay, T. Ibrahim and P. Nath, Phys. Rev. D 60, 063505 (1999) doi:10.1103/PhysRevD.60.063505 [arXiv:hep-ph/9811362 [hep-ph]].

[45] T. Ibrahim and P. Nath, Phys. Rev. D 62, 095001 (2000) doi:10.1103/PhysRevD.62.095001 [arXiv:hep-ph/0004098 [hep-ph]].

[46] J. Debove, B. Fuks and M. Klasen, Nucl. Phys. B 849, 64-79 (2011) doi:10.1016/j.nuclphysb.2011.03.015 [arXiv:1102.4422 [hep-ph]].

[47] B. Fuks, M. Klasen, D. R. Lamprea and M. Rothering, Eur. Phys. J. C 73, 2480 (2013) doi:10.1140/epjc/s10052-013-2480-0 [arXiv:1304.0790 [hep-ph]].
[48] J. Alwall, R. Frederix, S. Frixione, V. Hirschi, F. Maltoni, O. Mattelaer, H. S. Shao, T. Stelzer, P. Torrielli and M. Zaro, JHEP 07, 079 (2014) doi:10.1007/JHEP07(2014)079 [arXiv:1405.0301 [hep-ph]].

[49] T. Sjöstrand, S. Ask, J. R. Christiansen, R. Corke, N. Desai, P. Ilten, S. Mrenna, S. Prestel, C. O. Rasmussen and P. Z. Skands, Comput. Phys. Commun. 191, 159-177 (2015) doi:10.1016/j.cpc.2015.01.024 [arXiv:1410.3012 [hep-ph]].

[50] J. de Favereau et al. [DELPHES 3], JHEP 02, 057 (2014) doi:10.1007/JHEP02(2014)057 [arXiv:1307.6346 [hep-ex]].

[51] P. Speckmayer, A. Hocker, J. Stelzer and H. Voss, J. Phys. Conf. Ser. 219, 032057 (2010) doi:10.1088/1742-6596/219/3/032057.

[52] I. Antcheva, M. Ballintijn, B. Bellenot, M. Biskup, R. Brun, N. Buncic, P. Canal, D. Casadei, O. Couet and V. Fine, et al. Comput. Phys. Commun. 182, 1384-1385 (2011) doi:10.1016/j.cpc.2011.02.008.

[53] X. Cid Vidal, M. D’Onofrio, P. J. Fox, R. Torre, K. A. Ulmer, A. Aboubrahim, A. Albert, J. Alimena, B. C. Allanach and C. Alpigiani, et al. CERN Yellow Rep. Monogr. 7, 585-865 (2019) doi:10.23731/CYRM-2019-007.585 [arXiv:1812.07831 [hep-ph]].

[54] M. Cepeda, S. Gori, P. Ilten, M. Kado, F. Riva, R. Abdul Khalek, A. Aboubrahim, J. Alimena, S. Alioli and A. Alves, et al. CERN Yellow Rep. Monogr. 7, 221-584 (2019) doi:10.23731/CYRM-2019-007.221 [arXiv:1902.00134 [hep-ph]].

[55] G. Aad et al. [ATLAS], Eur. Phys. J. C 80, no.2, 123 (2020) doi:10.1140/epjc/s10052-019-7594-6 [arXiv:1908.08215 [hep-ex]].

[56] S. Iwamoto, T. T. Yanagida and N. Yokozaki, [arXiv:2104.03223 [hep-ph]].

[57] Y. Gu, N. Liu, L. Su and D. Wang, [arXiv:2104.03239 [hep-ph]].

[58] M. Van Beekveld, W. Beenakker, M. Schutten and J. De Wit, [arXiv:2104.03245 [hep-ph]].

[59] W. Yin, JHEP 06, 029 (2021) doi:10.1007/JHEP06(2021)029 [arXiv:2104.03259 [hep-ph]].

[60] F. Wang, L. Wu, Y. Xiao, J. M. Yang and Y. Zhang, [arXiv:2104.03262 [hep-ph]].

[61] J. Cao, J. Lian, Y. Pan, D. Zhang and P. Zhu, [arXiv:2104.03284 [hep-ph]].

[62] M. Chakraborti, S. Heinemeyer and I. Saha, [arXiv:2104.03287 [hep-ph]].

[63] P. Cox, C. Han and T. T. Yanagida, [arXiv:2104.03290 [hep-ph]].

[64] C. Han, [arXiv:2104.03292 [hep-ph]].

[65] S. Baum, M. Carena, N. R. Shah and C. E. M. Wagner, [arXiv:2104.03302 [hep-ph]].

[66] W. Ahmed, I. Khan, J. Li, T. Li, S. Raza and W. Zhang, [arXiv:2104.03491 [hep-ph]].
[67] H. Baer, V. Barger and H. Serce, Phys. Lett. B 820, 136480 (2021) doi:10.1016/j.physletb.2021.136480 [arXiv:2104.07597 [hep-ph]].

[68] M. Endo, K. Hamaguchi, S. Iwamoto and T. Kitahara, [arXiv:2104.03217 [hep-ph]].

[69] M. Ibe, S. Kobayashi, Y. Nakayama and S. Shirai, [arXiv:2104.03289 [hep-ph]].

[70] M. Chakraborti, L. Roszkowski and S. Trojanowski, JHEP 05, 252 (2021) doi:10.1007/JHEP05(2021)252 [arXiv:2104.04458 [hep-ph]].

[71] L. A. Anchordoqui, I. Antoniadis, X. Huang, D. Lust and T. R. Taylor, doi:10.1002/prop.202100084 [arXiv:2104.06854 [hep-ph]].

[72] P. Athron, C. Balázs, D. H. Jacob, W. Kotlarski, D. Stöckinger and H. Stöckinger-Kim, [arXiv:2104.03691 [hep-ph]].

[73] C. T. Lu, R. Ramos and Y. L. Sming Tsai, [arXiv:2104.04503 [hep-ph]].

[74] T. Li, M. A. Schmidt, C. Y. Yao and M. Yuan, [arXiv:2104.04494 [hep-ph]].

[75] J. Y. Cen, Y. Cheng, X. G. He and J. Sun, [arXiv:2104.05006 [hep-ph]].

[76] D. Borah, M. Dutta, S. Mahapatra and N. Sahu, [arXiv:2104.05656 [hep-ph]].

[77] S. Zhou, [arXiv:2104.06858 [hep-ph]].

[78] W. Altmannshofer, S. A. Gadam, S. Gori and N. Hamer, [arXiv:2104.08293 [hep-ph]].

[79] A. Dasgupta, S. K. Kang and M. Park, [arXiv:2104.09205 [hep-ph]].

[80] A. Jueid, J. Kim, S. Lee and J. Song, [arXiv:2104.10175 [hep-ph]].

[81] J. A. Carpio, K. Murase, I. M. Shoemaker and Z. Tabrizi, [arXiv:2104.15136 [hep-ph]].

[82] K. S. Babu, S. Jana, M. Lindner and V. P. K, [arXiv:2104.03291 [hep-ph]].

[83] M. Frank, Y. Hiçyılmaz, S. Mondal, Ö. Özdal and C. S. Ün, [arXiv:2107.04116 [hep-ph]].

[84] J. Ellis, J. L. Evans, N. Nagata, D. V. Nanopoulos and K. A. Olive, [arXiv:2107.03025 [hep-ph]].

[85] B. Yu and S. Zhou, [arXiv:2106.11291 [hep-ph]].

[86] K. S. Jeong, J. Kawamura and C. B. Park, [arXiv:2106.04238 [hep-ph]].

[87] B. De, D. Das, M. Mitra and N. Sahoo, [arXiv:2106.00979 [hep-ph]].

[88] D. Zhang, [arXiv:2105.08670 [hep-ph]].

[89] M. D. Zheng and H. H. Zhang, [arXiv:2105.06954 [hep-ph]].

[90] M. Abdughani, Y. Z. Fan, L. Feng, Y. L. S. Tsai, L. Wu and Q. Yuan, Sci. Bull. 66, 2170-2174 (2021) doi:10.1016/j.scib.2021.07.029 [arXiv:2104.03274 [hep-ph]].
[91] M. Abdughani, K. I. Hikasa, L. Wu, J. M. Yang and J. Zhao, JHEP 11, 095 (2019) doi:10.1007/JHEP11(2019)095 [arXiv:1909.07792 [hep-ph]].