Crossover in the pressure evolution of elementary distortions in RFeO₃ perovskites and its impact on their phase transition

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I. INTRODUCTION

Hydrostatic pressure has been increasingly considered in the study of critical phenomena since it allows one to modify the interatomic distances and, consequently, the interactions to a greater extent than any other external parameter, such as temperature or magnetic field. Thus, the number of published experimental [1–5] and theoretical [6–13] reports concerning temperature or magnetic field. Thus, the number of published theoretical approaches. The similarity of the pressure-induced isostructural insulator-to-metal phase transition, between the two opposite pressure behaviors is discussed in relation to the general rules proposed from different theoretical approaches. The similarity of the pressure-induced isostructural insulator-to-metal phase transition, observed in the whole series, points out that the tilts play a minor role in its driving mechanisms. A clear relationship between octahedral compressibility and critical pressure is ascertained.

This work reports on the pressure dependence of the octahedral tilts and mean Fe-O bond lengths in RFeO₃ (R = Nd, Sm, Eu, Gd, Tb, and Dy), determined through synchrotron x-ray diffraction and Raman scattering, and their role on the pressure-induced phase transition displayed by all of these compounds. For larger rare-earth cations (Nd-Sm), both anti- and in-phase octahedral tilting decrease as pressure increases, whereas the reverse behavior is observed for smaller ones (Gd-Dy). EuFeO₃ stands at the borderline, with nearly pressure-independent tilt angles. For the compounds where the tilts increase with pressure, the FeO₆ octahedra are compressed at lower rates than for those ones exhibiting opposite pressure tilt dependence. The crossover between the two opposite pressure behaviors is discussed in relation to the general rules proposed from different theoretical approaches. The similarity of the pressure-induced isostructural insulator-to-metal phase transition, observed in the whole series, points out that the tilts play a minor role in its driving mechanisms. A clear relationship between octahedral compressibility and critical pressure is ascertained.

The first attempts to formulate general rules regarding the pressure evolution of elementary distortions in perovskites under high pressure, several rules have been proposed, based on experimental results, theoretical models, and density functional theory (DFT) calculations [7–9].

AO₁₂ dodecahedra [14]. The structure of ABO₃ perovskites can be obtained from basic distortions of the ideal Pm̅3m cubic phase [15,16]. For perovskites with a tolerance factor less than unity, the most important distortion is characterized by BO₆ octahedral rotations [17]. In the case of Pnma orthorhombic perovskites, these distortions are in-phase and antiphase tilts about the [010]_c and [101]_c pseudocubic directions [17], which transform according to the M₅⁺ and R₄⁺ irreducible representations of the Pm̅3m space group, respectively [15,16]. These two distortions are the primary order parameters associated with the symmetry lowering, and present the largest amplitudes, which increase as the A-cation size decreases [15]. Other non-symmetry-breaking distortions occur together with octahedral tilting, where the antiparallel motion of the A cations along the z pseudocubic direction with symmetry X₅⁺ is the most relevant, as it bares the largest amplitude among the secondary distortions. Moreover, it couples to both tilts via a specific trilinear coupling term that provides an energy gain crucial to the stabilization of the Pnma phase [15,18].

In order to predict and explain the structural behavior of perovskites under high pressure, several rules have been proposed, based on experimental results, theoretical models, and density functional theory (DFT) calculations [7–9]. The first attempts to formulate general rules regarding the
pressure dependence of the phase sequence in perovskites were based on the octahedral tilting and on the ratios of the compressibilities ($M_A/M_B$) and volume ($V_A/V_B$) of the AO$_{12}$ and BO$_6$ polyhedra [9–13]. If $M_A/M_B > 1$, a transition to a higher-symmetry phase is expected, whereas the opposite should occur if $M_A/M_B < 1$ [10]. Based on the bond-valence concept, Zhao et al. [9] predicted that the AO$_{12}$ dodecahedra are expected to be significantly more compressible than the BO$_6$ octahedra in orthorhombic perovskites, with both A and B cations having the formal charge +3 (3:3 perovskites), as is the case for rare-earth orthoferrites (RFeO$_3$). This model also evidences the correlation between the BO$_6$ and AO$_{12}$ compressibilities and the rate of change of the octahedral tilting [9]. Thus, the octahedral tilts should decrease with increasing pressure [9]. Moreover, the decrease of the octahedral tilt with pressure should yield a structural phase transition from the orthorhombic into a higher-symmetric structure at some critical pressure [9]. As a matter of fact, these rules are not followed by many of 3:3 perovskites. For instance, new results in rare-earth chromites (RCrO$_3$), with small R cations, show that octahedral tilting increases with pressure [19,20]. The other approach, based on the pressure evolution of the $V_A/V_B$ ratio, followed the work of Thomas et al. [12], which stated that in the case of 3:3 perovskites, the $V_A/V_B$ is equal to 5 for the ideal symmetry ($Pm3m$) and a reduction in $V_A/V_B$ to 4.7 is associated with a change to orthorhombic symmetry ($Pmna$), found to be stable for $V_A/V_B$ values down to 3.8 [12,13,14,15].

In order to unravel the mechanisms underlying the distinct pressure dependences of octahedral tilting, Xiang et al. [7] conducted first-principles calculations on representative perovskites and proposed a set of rules governing the tilt evolution with pressure. In this framework, they simulated the pressure dependence of both in-phase and antiphase octahedral tilts of LaFeO$_3$ and LuFeO$_3$, as border cases in the rare-earth orthoferrites series [7]. For LaFeO$_3$, it was found that both octahedral tilts are suppressed as pressure increases, in agreement with Zhao’s prevision [7]. However, pressure suppresses the antiphase but enhances the in-phase tilting in LuFeO$_3$ [7]. This pressure behavior was interpreted by taking into account the contribution of a trilinear coupling between these two rotations and the antipolar mode involving the A cation [7,18]. Thus, a new rule emerges for the orthorhombic $Pmna$ perovskites, which simultaneously exhibit the in-phase and antiphase octahedral tilts, stating that they are not inevitably both suppressed or enhanced by pressure [7]. Where and how the crossover between the two aforementioned distinct pressure behaviors occurs in the rare-earth orthoferrites still remains to be understood.

Despite the intensive research already done [20,22–26], a systematic study of the pressure evolution of the elementary structural distortions in orthorhombic perovskites is still missing, as well as the search for crossover events. In this regard, the study of the RFeO$_3$ series under high pressure is particularly interesting, as it is expected to show different pressure dependences of the elementary distortions across the series [7,8]. Moreover, this system bears a distinct advantage, as it does not show distortions other than those with direct origin in octahedral tilting, such as Jahn-Teller distortion, which can influence the mean B-O bond length [2].

In this work, we present an experimental study of the structure of the RFeO$_3$ (R = Nd, Sm, Eu, Gd, Tb, and Dy) as a function of pressure, by means of synchrotron x-ray diffraction and Raman scattering. We first analyze the pressure evolution of the octahedral tilt angles across the series in order to reach an overall picture of its dependence on the rare-earth cation size. Then, the pressure variation of the mean Fe-O bond length obtained by Raman scattering is explored in order to correlate the pressure tilt behavior and the mechanisms to accommodate pressure. We also discuss the experimental results within the scope of the predictions of theoretical models. Finally, the dependence of the critical pressure on the rare-earth cation size is examined in terms of the pressure behavior of the elementary distortions.

II. EXPERIMENTAL ASPECTS

SmFeO$_3$ single crystals were grown in an optical-floating-zone furnace [27], and NdFeO$_3$ and TbFeO$_3$ powder was prepared using single crystals grown by the floating-zone method in a FZ-T-4000 (Crystal Systems Corp.) mirror furnace. EuFeO$_3$ powder was obtained by conventional solid-state reactions, while GdFeO$_3$ and DyFeO$_3$ powder was prepared using the urea sol-gel combustion method [28]. The quality of the samples was previously characterized by means of x-ray diffraction, Fourier transform infrared spectroscopy, x-ray photoemission spectroscopy, and scanning electron microscopy. The ceramics were manually grinded to acquire a homogeneous powder and loaded in a diamond anvil cell (DAC) with diamond culets of 300 μm diameter, using a stainless-steel gasket and with helium as a pressure-transmitting medium. The pressure was monitored through the standard fluorescence method of a ruby loaded next to the sample [29]. High-pressure synchrotron x-ray diffraction (XRD) experiments on SmFeO$_3$ and TbFeO$_3$ were performed at the European Synchrotron Radiation Facility (ESRF) on the ID27 high-pressure beam line (λ = 0.3738 Å). The diffraction data were analyzed by Le Bail and Amplitudes refinements using both FULLPROF and JANA2006 software [30,31]. The Raman spectra were recorded on a Horiba LabRam at MINEF (Grenoble, France) using a He-Ne laser at 633 nm for TbFeO$_3$ and on a Horiba T64000 spectrometer at Institut Néel (Grenoble, France) using an Ar+ laser at 514.5 nm for RFeO$_3$ (R = Nd, Sm, Eu, Gd, and Dy). In both cases, the laser power was kept below 8 mW on the DAC to avoid sample heating. Raman spectra were fitted by a sum of independent damped oscillators using IGRPRO software.

III. RESULTS

A. High-pressure XRD: SmFeO$_3$ and TbFeO$_3$

Figure 1 shows the pressure dependence of the pseudocubic lattice parameters $a_{pc} = a/\sqrt{2}$, $b_{pc} = b/2$, and $c_{pc} = c/\sqrt{2}$ (values can be found in Table I of the Supplemental Material [32]), and the pseudocubic volume $V_{pc} = V/4$ (Z = 1), with V the volume of the primitive cell of the $Pmna$ structure of SmFeO$_3$ and TbFeO$_3$. See Fig. S1 of the Supplemental Material [32], where representative XRD patterns of both compounds, recorded at different pressures, are shown.
isothermal equation of state \([33,34]\),
\[
P = 3B_0f_E[1 + 2f_E]^{3/2} \left[ 1 + \frac{5}{3}(B'_0 - 4)f_E \right],
\]
where \(B_0\) the bulk modulus and \(B'_0\) is its pressure derivative, all taken at room pressure, and \(f_E\) is given by
\[
f_E = \frac{1}{2} \left( \frac{V_{pc}(0)}{V_{pc}} \right)^{2/3} - 1,
\]
where \(V_{pc}\) is the pseudocubic volume at the pressure \(P\), and \(V_{pc}(0)\) is its room pressure value. Figure S2 of the Supplemental Material \([32]\) shows the best fit of Eq. (1) to the experimental data. Table I presents the values of \(B_0\) and \(B'_0\), obtained from the fit procedure. For both compounds, the values of the bulk modulus are not much different, taking values around 182 GPa, while the \(B'_0\) is about 4, typical of nearly isotropic compressions.

B. Raman scattering

According to group theory, we expect 24 Raman-active vibration modes for the orthorhombic \(Pnma\) space group (see Ref. \([3]\) for the symmetry of the Raman-active modes). The Raman spectra of \(RFeO_3\), with \(R =\) Nd, Sm, Eu, Gd, Tb, and Dy, recorded at different applied pressures are shown in Fig. S3 of the Supplemental Material \([32]\). They simultaneously exhibit all Raman-active modes due to the unpolarized recording condition. A detailed mode assignment of the Raman bands and the corresponding atomic motions for rare-earth orthoferrites is presented elsewhere \([3]\).

The Raman spectra for all the studied materials exhibit similar trends with increasing pressure: Raman bands shift towards higher wave numbers due to an overall pressure-induced bond shortening and volume reduction, and become broader; their intensity reduces and disappears above a certain pressure that depends on the compound. The latter result corroborates the existence of a structural phase transition at high pressures for all the studied compounds. The nature and structure of the high-pressure phase will be discussed in the last section.

In the following, we focus our attention on the Raman-active modes assigned to the \(FeO_6\) octahedral rotations, mirroring the antiphase and in-phase octahedral tilts. Figure 2 presents the pressure dependence of the wave number of the aforementioned Raman modes of the studied compounds. The pressure dependence of the wave number of the Raman modes does not exhibit anomalous behavior up to the critical pressure, corroborating that the \(Pnma\) structure is preserved in the low-pressure range. A linear pressure dependence of the wave number of the Raman bands is observed below to 20 GPa. Concerning \(NdFeO_3\), our results agree with those already reported below 11 GPa \([26]\). From the best fit of a linear function to the experimental data below to 20 GPa, shown in Fig. 2, we have determined the corresponding slopes and wave numbers at room conditions, which are presented in Table II.

The slopes of the linear pressure dependence of the wavenumber of both in-phase and antiphase tilt modes tend to increase as the rare-earth cation size decreases. We can see this result in another way: for larger tilt angles measured at

![FIG. 1. Pseudocubic lattice parameters and volume as a function of pressure for (a) TbFeO₃ and (b) SmFeO₃. The dashed lines are guides for the eyes.](image-url)
the atomic positions to vary in the three-dimensional space without restriction to find the global minimum, AMPLIMODES are used to improve the structure refinement by describing the displacement of the atoms, relative to their positions in the high-symmetry Pmn̅3n structure, as the superposition of symmetry-adapted distortions [35]. The refinements show that the internal octahedral distortions have small amplitudes (less than 0.05 Å), whereas the octahedral tilts distortions and the rare-earth shifts are apparently larger (between 1 and 2 Å value of global amplitude). However, the refined mode amplitudes when plotted versus pressure show dispersion, especially in the modes mainly involving oxygen motions (that is, the tilting modes—see Fig. S4 of the Supplemental Material [32]). This is not surprising since high-pressure powder diffraction presents peak overlapping and the contribution of the oxygen atoms to the diffraction peak intensity is small compared to the rest of the atoms (that is, Sm or Tb, and Fe). The AMPLIMODES refinement results, especially regarding the octahedral tilts, should be contrasted with other observations.

Thus, we estimated the tilt angles from the lattice parameters, using the equations of Megaw et al. [36,37]:

$$\theta = \cos^{-1}(a/b), \quad \varphi = \cos^{-1}(\sqrt{2a/b}),$$

(3)

where $\theta$ is the antiphase tilt angle around the [101]$_{pc}$ axis and $\varphi$ is the in-phase tilt angle around the [010]$_{pc}$ axis. These formulas assume that the change of the unit-cell volume is originated by the octahedral tilting, with no significant additional octahedra distortion [36,37]. The formula is most suitable for larger tilts than for smaller tilts, as for the latter other contributions to the change of the unit-cell volume may be comparable to the contribution from the tilts [36,37]. Therefore, for our compounds, the estimation using Megaw’s formula is better for the antiphase tilt than for the in-phase tilt, and for smaller tolerance factors than for larger ones. The reliability of this approach was assessed by comparing the values of the tilt angles for these compounds, obtained at room conditions using the Megaw’s formula, with their reference ones obtained from the refinement of the atomic positions, as shown in Fig. 3. Furthermore, their pressure behaviors follow similar trends as the ones found by the AMPLIMODES analysis, though with much less dispersion (see Fig. S4 in the Supplemental Material [32]). Moreover, as we shall discuss in the following, this approach is also consistent with the pressure dependence of the spontaneous $e_{44}$ strain, calculated using the refined lattice parameters (see Fig. S5 in the Supplemental Material [32]), which is inherently connected to the two tilts [17]. The good agreement between the literature and our experimental observations ensures the validity of this method for other compounds where similar conditions are expected, such as LuFeO$_3$, EuFeO$_3$, and NdFeO$_3$. For these compounds, we have calculated the tilt angles from the lattice parameters measured at different pressures already published [20,23], as these values are relevant for the discussion of how their pressure behavior depends on the rare-earth ionic radius (see the following section).

The pressure dependence of the antiphase tilt angle for $R =$ Nd, Sm, Eu, Tb, and Lu, and of the in-phase tilt angle for $R =$ Sm, Eu, and Tb, are shown in Figs. 4(a) and 4(b), respectively. The data regarding the in-phase tilt angle for

![FIG. 2. Pressure evolution of the Raman modes wave number associated with (a) in-phase and (b) antiphase octahedral rotations for the different studied compounds.](image)
LuFeO$_3$ and NdFeO$_3$ are not presented because they are too scattered. For the same compound, both antiphase and in-phase tilt angles present similar pressure behaviors, but the pressure trend depends on the $R$ cation. For the compounds with larger rare-earth cations (SmFeO$_3$ and NdFeO$_3$), the applied pressure causes a decrease of both antiphase and in-phase tilt angles, and thus a decrease of the values of atomic displacements associated with the symmetry-adapted $R_4^+$ and $M_3^+$ distortion modes, while the opposite behavior is observed for the compounds with smaller rare-earth cations (LuFeO$_3$ and TbFeO$_3$). A similar trend is ascertained for the pressure dependence of the spontaneous $e_4$ strain in the same pressure range (see Fig. S5 of the Supplemental Material [32]). The EuFeO$_3$ case sits in the transition between these two opposite pressure behaviors, where both tilt angles and, consequently, the spontaneous $e_4$ strain are almost pressure independent. Our experimental results point out that pressure weakens both octahedra tilts for compounds with larger rare-earth cations towards a less distorted structure, while it enhances them for compounds with smaller rare-earth cations. Moreover, a remarkable continuous evolution between the two opposite pressure behaviors is experimentally observed.

As stated by Xiang et al. [7], the stabilization of the $Pnma$ phase in perovskites is predicted to be due to an energy gain coming from a trilinear coupling of both $R_4^+$ and $M_3^+$ distortions with the $X_5^+$ distortion, associated with the displacement of the rare-earth cation [8,18]. In this theoretical framework, the magnitude of the $X_5^+$ distortion mode is predicted to be proportional to the product of both tilt distortions [8]. If this is the case, then this distortion should follow the same trend as the tilts to which it is coupled. In fact, the pressure evolution of the $X_5^+$ distortion amplitude, which we have experimentally obtained by the AMPLIMODES analysis, is consistent with this model since it mimics the behavior of the tilts for the respective rare earth, as seen in Fig. 4(c).

**B. Relation between distortion amplitudes and Raman wave numbers**

Raman scattering was also used to probe the pressure dependence of both tilt angles and octahedra distortions. For the same $RBO_3$ system, with fixed $B$ atom, where quite similar mean $B$-$O$ bond lengths are experimentally evidenced,
The case of EuFeO$_3$ is straightforward because the tilt angles rotation modes, presented in Fig. 5(a), are always positive. For the other compounds, the not constant for these compounds). Ferrites and orthochromites by Weber et al. [3,40]. However, the tilt-mode wave numbers are also dependent on the mean B-O bond length, as it was established by Todorov et al. [41], and more recently by Vilarinho et al. [42]. So, the tilt-mode wave numbers are functions of both the tilt angle and the B-O bond length. In a general approach, the wave number of the tilt mode for FeO$_6$ octahedral volume may have different values for each compound, as it will be shown. According to Eq. (4), we present an estimative of the slope of the pressure dependence of the corresponding wave number has two contributions: $m_{\text{total}} = m_1 + m_2$, where $m_1$ stands for the contribution coming from the actual tilt-angle change with pressure ($\alpha_1$), and $m_2$ is the contribution coming from the isotropic reduction of the FeO$_6$ octahedra volume; i.e., of the average Fe-O distance ($\alpha_2$[Fe – O]) [41,42]. In the following, we present an estimative of the $m_2$ value, and how $m_1$ changes with the rare-earth ionic radius, thus gaining insight on how the mean Fe-O bond length changes with pressure for each rare-earth orthoferrite.

The rare-earth cation-size dependence of the slopes of the linear pressure relation, in the 0–20 GPa range, of the Raman rotation modes, and of the antiphase and in-phase tilt angles, is depicted in Figs. 5(a) and 5(b), respectively. The slopes of the linear pressure dependence of the Raman rotation modes, presented in Fig. 5(a), are always positive. EuFeO$_3$ is straightforward because the tilt angles are pressure independent, and thus $m_1 = 0$ [see Eq. (4)]. So, the linear pressure dependence of both Raman rotation modes in EuFeO$_3$, observed in the 0–20 GPa range, comes from the reduction of the FeO$_6$ octahedra volume; i.e., of the average Fe-O distance ($\alpha_2$[Fe – O]) [41,42]. In this case, the wave number of the rotational modes mirrors the reduction of the unit-cell volume due to the decrease of the mean B-O bond lengths with pressure, and a linear relation between the mode wave number and volume is experimentally evidenced, as shown in Fig. 5(c) (similar plots for TbFeO$_3$ and SmFeO$_3$ are presented in Fig. S6 of the Supplemental Material [32] for which the linear relation does not hold, as the tilts are not constant for these compounds).

In order to unravel the effect of the tilt angles on the pressure dependence of the Raman tilt-mode wave number for the other compounds, the $m_2 = m_{\text{total}}$ value calculated for EuFeO$_3$ might not be the best reference since the reduction of the FeO$_6$ octahedral volume may have different values for each compound, as it will be shown. According to Eq. (4), the pressure derivative of the wave number of the Raman tilt modes is

$$\frac{d\omega}{dP} = [\alpha_1 - \alpha_2(\text{Fe} -\text{O})] \frac{d\varphi}{dP} - \alpha_2 \varphi(P) \left( \frac{d(\text{Fe} -\text{O})}{dP} \right).$$

In the following, we shall consider the values concerning $\frac{d\omega}{dP}$ and $\frac{d\varphi}{dP}$ as the slopes obtained in the 0–20 GPa range, shown in Figs. 5(a) and 5(b), respectively. For EuFeO$_3$, where $\frac{d\alpha_2}{dP} = 0$, we obtained $\frac{d(\text{Fe} -\text{O})}{dP} = -0.0031 \pm 0.0001$ Å/GPa, using either anti- or in-phase tilts. For the remaining compounds, we focus our calculation to the linear range in the vicinity of $P = 0$ GPa, thus taking the values of $\langle\text{Fe} -\text{O}\rangle$ and $\varphi$ at that point. The obtained results for the low-pressure...
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| TABLE III. $\frac{d(\text{Fe-O})}{dP}$ calculated from Eq. (5), for the low-pressure range, using either the anti- or the in-phase tilt angles, for the different rare-earth orthoferrites. |
|-----------------|-----------------|-----------------|-----------------|
| Tilt angle      | Antiphase       | In-phase        |
| TbFeO$_3$       | $-0.0023 \pm 0.0003$ | $-0.0009 \pm 0.0004$ |
| EuFeO$_3$       | $-0.0031 \pm 0.0001$ | $-0.0031 \pm 0.0001$ |
| SmFeO$_3$       | $-0.0071 \pm 0.0003$ | $-0.0058 \pm 0.0006$ |
| NdFeO$_3$       | $-0.0082 \pm 0.0003$ | N/A             |

range are shown in Table III. The estimated $\frac{d(\text{Fe-O})}{dP}$ values varied only around 0.001 Å/GPa when using the two different rotation modes. It is worthwhile to note that Weber et al. [3] have found different proportionality constants ($m_{\text{total}}$) between $\omega$ and $\varphi$ for the anti- and in-phase tilts for the $\text{AFeO}_3$ system. The presented value agrees with that obtained by Todorov et al. [41] for the antiphase tilt, but not for the in-phase tilt. Therefore, the most suitable value for the actual $\frac{2(\text{Fe-O})}{dP}$ should be the one obtained with the antiphase tilt.

These results evidence that at low pressures, as the rare-earth ionic radius increases, the pressure rate at which the FeO$_6$ octahedra reduce also increases. For TbFeO$_3$, the average reduction of the Fe-O bond length is of $-0.0016$ Å/GPa, which is half the value of EuFeO$_3$, and around five times smaller than the largest one of $-0.0082$ Å/GPa for NdFeO$_3$. This is not unexpected since the tilts provide a mechanism of pressure accommodation. Thus, for the smaller rare-earth cations, where the tilts increase with pressure, the FeO$_6$ octahedra are compressed at lower rates. Conversely, for the larger rare earths, where the tilts decrease with pressure, the FeO$_6$ octahedra are compressed at larger rates.

C. Crossover for the behavior of tilts in the rare-earth series: Experiment and theory

We now focus on the evolution of the tilt angles in the vicinity of zero (atmospheric) pressure across the series. From Fig. 5(b), we can observe that the slope of the linear pressure dependence of both antiphase and in-phase octahedral tilt angles is positive for compounds with smaller rare-earth cations ($R = \text{Lu}$ and $\text{Tb}$), while they are negative for those with larger rare-earth cations ($R = \text{Sm}$ and $\text{Nd}$). EuFeO$_3$ stands at the borderline, as for this compound both slopes are negligibly small. The modulus of the slope of the linear pressure dependence of the in-phase tilt angle is smaller than the antiphase tilt angle for SmFeO$_3$, while for TbFeO$_3$ it is the opposite. These results show that the way the pressure is accommodated is different for compounds with different rare-earth cationic sizes.

We now compare this compilation of experimental results with predictions made by theoretical approaches for the pressure evolution of tilt angles. A first approach is based on the compressibility of the polyhedra forming the perovskite structure. In this approach, the ratio of compressibilities $\beta_A/\beta_B$ between the AO$_{12}$ and BO$_6$ polyhedra is used as a predictor of the behavior of tilt angles under pressure, and is calculated by a valence-bond sum model: tilt angles are expected to increase when this ratio is larger than unity, and decrease otherwise. The case of the rare-earth orthoferrites has been treated in Ref. [9], and the corresponding data are reported in Fig. 5(b) for comparison. The ratio is always positive, but decreases as the ionic radius increases and reaches values very close to unity for the smallest cations [9]. According to this criterion, all compounds in the series should see their tilt angles reduced under pressure, with the exception of TmFeO$_3$ [9]. This is not what is found experimentally, but the overall evolution bears a striking resemblance with the experimental one, up to an overall shift.

More recently, in a DFT-based approach, a set of rules for the evolution of tilts under pressure was proposed [7]. In this paper, the compounds with the smallest $A$ cations (Lu, Tm) are predicted to exhibit an unusual behavior whereby the two tilt angles behave in the opposite way with pressure: the antiphase tilts are reduced, whereas the in-phase tilts are enhanced [7]. For larger cations, both tilts behave the same way and are reduced as pressure increases. Here we find no evidence for this behavior, but instead both tilt angles behave the same in all investigate compounds. Also, Ref. [7] does not predict any case where both tilts increase under pressure.

The discrepancies between the theoretical and experimental results seemingly point to a specific difficulty in predicting the tilt behavior for the smallest cations in the series. While we cannot be conclusive at this point about the precise reasons for this discrepancy, we observe that in both approaches, the predicted tilt changes under pressure reach extremely small values, typically below 0.01°/GPa for LuFeO$_3$ in Ref. [7]. For EuFeO$_3$, which in fact has absolute pressure derivatives below 0.01°/GPa, the slopes for in-phase and antiphase tilts are of opposite sign (0.006 and $-0.01°$/GPa, respectively), but with values negligibly small that fall within the experimental uncertainty. In that context, it is conceivable that some additional parameters, legitimately neglected for large cations, become relevant and suffice to change the trend from slightly positive to negative, or vice versa. In particular, $\text{RFeO}_3$ with very small cations (Er-Lu) are known to exhibit some distortions of the AO$_{12}$ polyhedra, so that the classical separation of the 12 $A$-O bonds into four longer bonds and eight shorter bonds becomes less satisfactory [43]. We hypothesize that this has an influence of the bond valence sum and compressibilities calculated in Ref. [7]. In the DFT approach, where calculations are performed at 0 K and the behavior of tilts is the result of a delicate balance between the pressure evolution of the Landau parameters, one might question the role of temperature effects. Differences in tilt angles of the order of 0.1° between 0 and 300 K—which is reasonable following Ref. [44]—would be equivalent to several tens of GPa and cause significant shifts in the predicted behavior. Altogether, further work will be needed to clarify the picture for small cations, which calls for detailed experimental studies and reexamination of theoretical models.

D. High-pressure phase transition

All studied compounds exhibit a phase transition at high pressure. This phase transition is clearly evidenced by the changes on the XRD patterns, mirroring the sudden changes
of the pseudocubic lattice parameters. The structural phase transition reveals itself by the disappearance of the Raman signal above a certain critical pressure, hereafter designated by \(P_{IM}\). The crystallographic structure of the high-pressure phase of SmFeO\(_3\) and TbFeO\(_3\) is \(Pnma\), as has been reported for other rare-earth orthoferrites [20]. This structure allows for a Raman signal. Therefore, the disappearance of the Raman spectra above the critical pressure suggests the metallic character of the high-pressure structural phase. The pressure hysteresis evidenced by the XRD patterns of SmFeO\(_3\) and TbFeO\(_3\), and Raman spectra, for all studied compounds, recorded on increasing and decreasing pressure runs (3 to 6 GPa depending on the rare-earth cation), evidences for the first-order character of the high-pressure structural phase.

Taking into account both the XRD and Raman data, we can estimate the critical pressure \(P_{IM}\) for the studied compounds, which is depicted in Fig. 6. As can be observed, the critical pressure linearly increases as the rare-earth cation size decreases. In spite of this, all transitions are very similar in character, and it is reasonable to assume that all of them result from the same kind of high-spin to low-spin transition revealed by high-pressure Mössbauer spectroscopy [24].

It is remarkable that all the studied compounds exhibit a similar pressure-induced structural phase transition, regardless of the strong differences in their tilt evolution or their mechanisms accommodating the volume reduction. This can be first analyzed with respect to the predictors proposed to analyze the stability of the different tilted phases of perovskites. Thomas et al. [12], and later on Avdeev et al. [13], have determined that the \(Pnma\) phase is stable whenever \(V_A/V_B\) is smaller than 4.8. Since the \(V_A/V_B\) ratio is known to scale with the tilt angles [21,38], we have estimated its pressure dependence following the equations of Wang et al. [21] for TbFeO\(_3\) and SmFeO\(_3\). The obtained results are plotted in Fig. 4(d). For TbFeO\(_3\), and for any other \(RFeO_3\) where the tilt angles increase with pressure, the \(V_A/V_B\) ratio decreases with pressure and drives the crystal away from the stability limit of 4.8. On the other hand, decreasing tilts, such as in SmFeO\(_3\), do drive the perovskite towards the edge of the stability range, but, in SmFeO\(_3\), the high-spin to low-spin transition, which is of totally different nature and beyond the scope of the polyhedra compressibility models, occurs before this limit is actually reached. We could expect the line to be crossed for larger cations, as a result of a quicker tilt reduction. We cannot make any conclusive statement for NdFeO\(_3\), but for LaFeO\(_3\), a transition at 21 GPa has indeed been reported, with a change in tilts, before the high-spin to low-spin transition at 38 GPa. Altogether, we conclude that the predictors by Thomas and Avdeev remain valid under pressure, at least within the range accessible with the orthoferrites. This is also consistent with the idea that the phase transition seen in most orthoferrites has a different origin, where tilts play a minor role. We also note that the high-pressure metallic phases also seem to adopt the \(Pnma\) structure and it would be interesting to check whether or not those predictors, which were derived from analyses of low-pressure, insulating phases, still hold for the metallic \(Pnma\) perovskites with distinct electronic properties. But data are too scarce at this point and this is left for future studies.

As the tilts have little impact on this transition, it is sensible to assume that is depends mostly on the chemistry of the FeO\(_6\) octahedra. We suggest that as pressure increases and the FeO\(_6\) octahedra become smaller, a critical volume of the octahedron is reached where the electronic repulsion between the oxygens \(p\) electrons and the iron \(e_g\) electrons is such that it triggers this electronic reconfiguration, as it is energetically favorable for the \(e_g\) electrons to pair with the \(t_{2g}\) electrons, avoiding coming closer to the oxygen \(p\) electrons. The critical volume can be estimated for EuFeO\(_3\). As the tilt angles do not change with pressure, there is an isotropic volume reduction of the unit cell, already evidenced in Fig. 5(c), and thus the ratio of the FeO\(_6\) octahedron with the pseudocubic unit-cell volume (\(V_{FeO6}/V_{pc}\)) can be assumed to be pressure independent. Considering a regular octahedron at room pressure, the ratio \(V_{FeO6}/V_{pc} = 0.189\) is obtained. Using the value of \(V_{pc}\) before the critical pressure, the critical volume for EuFeO\(_3\) is estimated to be \(V_{FeO6} = 9.5 \pm 0.2\ \text{Å}^3\).

It is expected that for larger \(R\) cations, where from Table III we know the FeO\(_6\) octahedra reduce their volume at a larger rate, the aforementioned phenomenon is promoted at a lower critical pressure. Conversely, for the smaller \(R\) cations, as the FeO\(_6\) octahedra reduce their volume at a lower rate, they present higher values of critical pressure. This prediction explains the behavior of \(P_{IM}\) with the rare-earth size, shown in Fig. 6. Moreover, one can infer that a similar role is played by the octahedral tilting in the similar pressure-driven phase transition observed in \(R\)MnO\(_3\) [2]. The critical pressures for \(R\)MnO\(_3\) have similar dependence on the rare-earth ionic size, being always slightly higher by a constant value of 2 GPa. This difference can be assigned to the Jahn-Teller distortion present in \(R\)MnO\(_3\), as it provides an additional pressure accommodation mechanism than their respective \(RFeO_3\) compounds. It is worthwhile to note that the tilt angles are almost the same, thus allowing for this comparison [42]. This fact also supports the importance of the chemistry inside the octahedra in determining the symmetry of the high-pressure phase. This is because, unlike the isostructural phase transition found for the \(RFeO_3\), the \(R\)MnO\(_3\) present many different symmetries of the high-pressure phase [2].
V. CONCLUSIONS

In this work, we present an experimental study of the pressure dependence of the main structural distortions and lattice dynamics in the $RFeO_3$ by XRD and Raman spectroscopy. First, we have ascertained the crossover of the pressure dependence of both anti- and in-phase octahedral tilting in the $RFeO_3$ series. EuFeO$_3$ stands distinctly at the borderline, where both tilts are pressure independent. For larger rare-earth cations, we have found that the octahedral tilting decreases as pressure increases, whereas the reverse behavior is observed for smaller ones.

Furthermore, the pressure dependence of the mean Fe-O bond length, estimated from Raman data, enabled us to determine the way the pressure is accommodated in the $RFeO_3$ series. We observed that for the compounds where the tilts increase with pressure, the FeO$_6$ octahedra are compressed at lower rates than for those showing opposite pressure tilt dependence.

Finally, we determined that all the compounds of the $RFeO_3$ series undergo a pressure-induced isostructural insulator-to-metal phase transition. Thus, we have to conclude that the different pressure evolutions of the octahedral tilts play a minor role in its driving mechanisms. Moreover, we have interpreted the rare-earth ionic-size dependence of the transition pressure in terms of reaching a critical volume size of the FeO$_6$ octahedra, estimated to be $V_{FeO_6} = 9.5 \pm 0.2 \text{Å}^3$ in EuFeO$_3$. For the smaller $R$ cations, as the FeO$_6$ octahedra reduce their volume at a lower rate, a shift of the transition to higher pressures occurs, contrarily to the case of larger $R$ cations, wherein the volume increases at a higher rate.

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[1] J. Oliveira, J. Agostinho Moreira, A. Almeida, V. H. Rodrigues, M. M. R. Costa, P. B. Tavares, P. Bouvier, M. Guennou, and J. Kreisel, Phys. Rev. B 85, 052101 (2012).
[2] D. A. Mota, A. Almeida, V. H. Rodrigues, M. M. R. Costa, P. Tavares, P. Bouvier, M. Guennou, J. Kreisel, and J. A. Moreira, Phys. Rev. B 90, 054104 (2014).
[3] M. C. Weber, M. Guennou, H. J. Zhao, J. Íñiguez, R. Vilarinho, A. Almeida, J. A. Moreira, and J. Kreisel, Phys. Rev. B 94, 214103 (2016).
[4] T. Aoyama, K. Yamauchi, A. Iyama, S. Picozzi, K. Shimizu, and T. Kimura, Nat. Commun. 5, 4927 (2014).
[5] T. Aoyama, A. Iyama, K. Shimizu, and T. Kimura, Phys. Rev. B 91, 081107(R) (2015).
[6] O. E. González-Vázquez, J. C. Wojdel, O. Diéguez, and J. Íñiguez, Phys. Rev. B 85, 064119 (2012).
[7] H. J. Xiang, M. Guennou, J. Íñiguez, J. Kreisel, and L. Bellaiche, Phys. Rev. B 96, 054102 (2017).
[8] P. Chen, M. N. Grisolia, H. J. Zhao, O. E. González-Vázquez, L. Bellaiche, M. Bibes, B. G. Liu, and J. Íñiguez, Phys. Rev. B 97, 024113 (2018).
[9] J. Zhao, N. L. Ross, and R. J. Angel, Acta Crystallogr. Sect. B Struct. Sci. 60, 263 (2004).
[10] R. J. Angel, J. Zhao, and N. L. Ross, Phys. Rev. Lett. 95, 025503 (2005).
[11] T. Tohei, A. Kuwabara, T. Yamamoto, F. Oba, and I. Tanaka, Phys. Rev. Lett. 94, 035502 (2005).
[12] N. W. Thomas, Acta Crystallogr. Sect. B Struct. Sci. B 52, 954 (1996).
[13] M. Avdeev, E. N. Caspi, and S. Yakovlev, Acta Crystallogr. Sect. B Struct. Sci. 63, 363 (2007).
[14] R. H. Mitchell, Perovskites, Modern and Ancient (Almaz Press, Ontario, Canada, 2002).
[15] J. M. Perez-Mato, D. Orobenosa, and M. I. Aroyo, Acta Crystallogr. Sect. A 66, 558 (2010).
[16] M. A. Carpenter and C. J. Howard, Acta Crystallogr. Sect. B 65, 134 (2009).
[17] A. M. Glazer, Acta Crystallogr. Sect. B 28, 3384 (1972).
[18] N. Miao, N. C. Bristowe, B. Xu, M. J. Verstraete, and P. Ghosez, J. Phys.: Condens. Matter 26, 035401 (2014).
[19] V. S. Bhardam, D. Swain, R. Dhanya, M. Polentarutti, A. Sunandesaran, and C. Narayana, Mater. Res. Express 1, 026111 (2014).
[20] G. K. Rozenberg, M. P. Pasternak, W. M. Xu, L. S. Dubrovinsky, S. Carlson, and R. D. Taylor, Europhys. Lett. 71, 228 (2005).
[21] D. Wang and R. J. Angel, Acta Crystallogr. Sect. B Struct. Sci. 67, 302 (2011).
[22] A. G. Gavriliuk, G. N. Stepanov, I. S. Lyubutin, A. S. Stepin, I. A. Trojan, and V. A. Sidorov, Hyperfine Interact. 126, 305 (2000).
[23] A. G. Gavriliuk, I. A. Trojan, R. Boehler, M. I. Eremets, I. S. Lyubutin, and N. R. Serebryanaya, J. Exp. Theor. Phys. Lett. 77, 619 (2003).
[24] M. P. Pasternak, W. M. Xu, G. K. Rozenberg, and R. D. Taylor, Mat. Res. Soc. Symp. Proc. 718, D2.7.1 (2002).
[25] M. Etter, M. Müller, M. Hanfland, and R. E. Dinnebier, Acta Crystallogr. Sect. B Struct. Sci. Cryst. Eng. Mater. 70, 452 (2014).
[26] M. Mihalik jr., M. Misek, M. Vavra, K. M. Lebecki, D. Legut, M. Mihalik, K. V. Kamenev, and M. Zentková, High Press. Res. 35, 1 (2015).
[27] S. Cao, H. Zhao, B. Kang, J. Zhang, and W. Ren, Sci. Rep. 4, 5960 (2014).
[28] J. Agostinho Moreira, A. Almeida, W. S. Ferreira, M. R. Chaves, J. B. Oliveira, J. M. M. da Silva, M. A. Sá, S. M. F. Vilela, and P. B. Tavares, J. Electroceram. 25, 203 (2010).
[29] H. K. Mao, J. Xu, and P. M. Bell, J. Geophys. Res. 91, 4673 (1986).
[30] J. Rodríguez-Carvajal, An Introduction to the Program FullProf (Laboratoire Leon Brillouin, CEA-CNRS, 2001).
[31] V. Petrícek, M. Dušek, and L. Palatinus, Z. Krist. 229, 345 (2014).
[32] See Supplemental Material at http://link.aps.org/supplemental/10.1103/PhysRevB.99.064109 for XRD patterns and Raman scattering spectra, fit of the Birch-Murnaghan equation of state, distortion modes amplitude, strain and lattice parameter values with pressure, and volume dependence of Raman octahedral rotational modes.
[33] F. Birch, Phys. Rev. 71, 809 (1947).
[34] F. Birch, J. Geophys. Res. 83, 1257 (1978).
[35] D. Orobengoa, C. Capillas, M. I. Aroyo, and J. M. Perez-Mato, J. Appl. Crystallogr. 42, 820 (2009).
[36] H. D. Megaw, Crystal Structures: A Working Approach (W. B. Saunders, Philadelphia, PA, 1973).
[37] Y. Zhao, D. J. Weidner, J. B. Parise, and D. E. Cox, Phys. Earth Planet. Inter. 76, 1 (1993).
[38] N. W. Thomas, Acta Crystallogr. Sect. B Struct. Sci. B 52, 16 (1996).
[39] M. N. Iliev, M. V. Abrashev, J. Laverdière, S. Jandl, M. M. Gospodinov, Y. Q. Wang, and Y. Y. Sun, Phys. Rev. B 73, 064302 (2006).
[40] M. C. Weber, J. Kreisel, P. A. Thomas, M. Newton, K. Sardar, and R. I. Walton, Phys. Rev. B 85, 054303 (2012).
[41] N. D. Todorov, M. V. Abrashev, and V. G. Ivanov, J. Phys.: Condens. Matter 24, 175404 (2012).
[42] R. Vilarinho, D. J. Passos, E. C. Queirós, P. B. Tavares, A. Almeida, M. C. Weber, M. Guennou, J. Kreisel, and J. A. Moreira, Phys. Rev. B 97, 144110 (2018).
[43] M. Marezio, J. P. Remeika, and P. D. Dernier, Acta Crystallogr. Sect. B 26, 2008 (1970).
[44] S. M. Selbach, J. R. Tolchard, A. Fossdal, and T. Grande, J. Solid State Chem. 196, 249 (2012).