Flavon VEV Scales in U(3) × U(3)' Model

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Abstract

We have recently proposed a quark and lepton mass matrix model based on U(3) × U(3)' family symmetry as the so-called Yukawaon model, in which the U(3) symmetry is broken by VEVs of flavons (Φ\textsubscript{f\textsuperscript{α}i})\textsuperscript{α} which are (3,3') of U(3) × U(3)'. The model has successfully provided the unified description of quark and lepton masses and mixings by using the observed charged lepton masses as only family-number dependent input parameters. However, our final goal is not only to give well-satisfied fitting of quark and lepton masses and mixings, but to investigate physics behind such a successful parameter fitting. Therefore, our next concern is scales of VEVs of the flavons because we have not paid attention to scales of flavons in the previous study. In order to give consistency among the scales, the previous flavon model is drastically changed together with economizing number of the flavons, but with keeping previous phenomenological success. We estimate that VEVs of flavons with (8 + 1, 1), (3,3'), and (1, 8 + 1) are of the orders of 10 TeV, 10\textsuperscript{4} TeV, and 10\textsuperscript{7} TeV, respectively.

PCAC numbers: 11.30.Hv, 12.60.-i, 12.15.Ff, 14.60.Pq,

1 Introduction

The greatest concern in the flavor physics is how to understand the origin of the observed hierarchical structures of masses and mixings of quarks and leptons. Recently, we have proposed a quark and lepton mass matrix model based on U(3) × U(3)' flavor symmetry \cite{1,2}. The model is an extended version of the so-called “Yukawaon” model \cite{3}, which is a kind of flavon model \cite{4} to understand origin of masses (mass matrices) by vacuum expectation values (VEVs) of flavons.

In the past Yukawaon models, VEV scales of flavons have not been discussed. In this paper, we take the energy scales of VEV of flavons into consideration. In order to discuss the VEV scales consistently, we have rebuilt the previous model. Especially, in the present model, “Yukawaons” are absent in the limit of unbroken U(3) × U(3)' symmetry. We assume fermions \(F_\alpha\) \((\alpha = 1, 2, 3)\) which are triplet of U(3)' in addition to quark and leptons \(f_i\) \((i = 1, 2, 3)\) which are triplet of U(3), and we consider that the effective Yukawa coupling terms are caused by the \(f-F\) mixing. As a result, the basic formulation of the model is drastically changed. However, the successful phenomenological study of the masses and mixings in the previous works \cite{1,2}...
is substantially kept as far as quarks and leptons are concerned. Therefore, the purpose of the present paper is not to discuss masses and mixings of quarks and leptons. Our concern is in the scales of $U(3)'$ and $U(3)$ and a scale of the right-handed Majorana neutrino mass.

Let us present the basic structure of the model:

(i) In the previous model with $U(3)\times U(3)'$ flavor symmetry, the quarks and leptons are assigned to triplets of $U(3)$ family symmetry.

(ii) The Majorana neutrino mass matrix is considered based on the conventional seesaw mechanism [5],

\[ (M^\nu_M)_{ij} = (\hat{M}_\nu)_{ik}(M_{R}^{-1})_{kl}(\hat{M}_\nu)_{lj}, \]

where $i,j = 1,2,3$ are indexes of $U(3)$ symmetry. Here the right-handed Majorana neutrino mass matrix $M_R$ will take a scale of $U(3)$ symmetry breaking $\mu \sim \Lambda_3$.

(iii) The difference among sectors $f = (u,d,\nu,e)$ seen in the observed masses and mixings is brought by scalars $S_f = (1,8+1)$ of $U(3)\times U(3)'$ by assuming that the $U(3)'$ symmetry is broken into a discrete symmetry $S_3$ at a scale $\mu = \Lambda_1$. The VEV forms of $S_f$ are given by

\[ \langle \hat{S}_f \rangle = v_S(1 + b_f X_3), \]

where $1$ and $X_3$ are defined by

\[ 1 = \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}, \quad X_3 = \frac{1}{3} \begin{pmatrix} 1 & 1 & 1 \\ 1 & 1 & 1 \\ 1 & 1 & 1 \end{pmatrix}, \]

and $b_f$ are complex parameters.

(iv) The Dirac mass matrices $(\hat{M}_f)_{ij}$ ($f = u,d,\nu,e$) for quarks and leptons have been given by a common form

\[ (\hat{M}_f)_{ij} = \langle \Phi_f \rangle_i^\alpha \langle S^{-1}_f \rangle_{\alpha}^\beta \langle \bar{\Phi}_f \rangle^j_\beta, \]

where $i,j$ and $\alpha,\beta$ are indexes of $U(3)$ and $U(3)'$, respectively. In this model, all of scalars (except for Higgs scalars $H$) are singlets of $SU(3)_c \times SU(2) \times U(1)_Y$. Such a unified description of quark and lepton mass matrices based on the form (1.4) with matrix forms (1.2) and (1.3) has first been proposed in Ref. [6] in order to get top quark mass enhancement. (However, their work was not based on the symmetry $U(3)\times U(3)'$.) We consider that the form (1.4) comes from a seesaw-like mechanism for quarks and leptons $f_i$ and hypothetical heavy fermions $F_\alpha$. (The detail formulation is discussed in Sec.2.)

1 In this paper, we denote flavons with $(8+1,1)$ and $(1,8+1)$ of $U(3)\times U(3)'$ as $\hat{A}_i^\alpha$ and $\hat{A}^\alpha_i$ with “hat”, respectively. We also denote a flavon $(3^*,3^*)$ as $\hat{A}^{\alpha i}$ in contrast to $A_\alpha$ of $(3,3)$. However, for flavons with $(3,3^*)$ and $(3^*,3)$, we denote those as $A^{\alpha i}$ and $A^{i\alpha}$, respectively, giving priority to $U(3)$ index.
(v) The U(3) is broken by VEVs of $\Phi_f$ with $(3, 3^*)$ of $U(3) \times U(3)'$, the VEV forms of which are assumed to be diagonal as given by

$$\langle \Phi_f \rangle = v_\Phi \text{diag}(z_1 e^{i\phi_1}, z_2 e^{i\phi_2}, z_3 e^{i\phi_3}).$$

(1.5)

We shall see that $\phi_f^i = 0$ for $f = d, e, \nu$ in later discussions. Therefore, we hereafter denote $(\phi_1^u, \phi_2^u, \phi_3^u)$ as $(\phi_1, \phi_2, \phi_3)$ simply. Namely, only for up-quark sector $f = u$, $\langle \Phi_u \rangle$ takes phase factors as $\langle \Phi_u \rangle = v_\Phi \text{diag}(z_1 e^{i\phi_1}, z_2 e^{i\phi_2}, z_3 e^{i\phi_3})$.

(vi) We will give $b_e = 0$ for the parameter $b_e$ in Eq.(1.2), i.e. $\hat{S}_e = v_S 1$ as seen in Eq.(2.9) in the next section. Thereby the charged lepton mass matrix is given by

$$\hat{M}_e \propto \text{diag}(z_1^2, z_2^2, z_3^2),$$

(1.6)

from which we get a relation

$$z_i = \frac{\sqrt{m_{ei}}}{\sqrt{m_e} + m_\mu + m_\tau}$$

(1.7)

Here $m_{ei} = (m_e, m_\mu, m_\tau)$ are charged lepton masses, and the real parameters $z_i$ are normalized as $z_1^2 + z_2^2 + z_3^2 = 1$. Note that the parameter values $(z_1, z_2, z_3)$ are universal for every sector $f = u, d, e, \nu$. On the other hand, the parameters $b_f$ in Eq.(1.2) are sector-dependent ($f$-dependent) but family-independent. Family-dependent parameter $\phi_i$ are described in terms of $z_i$ as seen in Sec.3. Therefore, family dependent parameters are only $z_i$. In this model, parameters which we can adjust are only the family-independent parameters $b_f$. Note that the parameters $b_f$ determine not only mass eigenvalues $(m_{f1}, m_{f2}, m_{f3})$ but also family mixing matrix $U_f$. The origin of the parameter values (1.6) has not been given in the Yukawaon model. The relation (1.7) is nothing but an ad hoc assumption.

(vii) Finally let us emphasize the strategy of the Yukawaon model based on the $U(3) \times U(3)'$ symmetry. In most mass matrix models, a symmetry for quarks and leptons is given, and thereby the masses and mixings are derived. (It is investigated how the family symmetry breaks into a suitable sub-symmetry or discrete symmetry.) In the Yukawaon model, we do not propose an explicit symmetry breaking mechanism for $U(3)$ symmetry. We only use the observed values of the charged lepton masses for the relation (1.7). The mechanism which gives charged lepton masses is left for a future task, and, for the moment, we do not ask the origin. Instead, we try to give unified description of all mass ratios and mixings in terms of the charged lepton mass values, and without using any family-number dependent parameters at all. The purpose of the Yukawaon model is to show that all the quark and lepton mass ratios and mixings can be described when we accept the observed charged lepton mass values as only family-dependent parameters.

Anyhow, we could successfully obtain a unified description of quark and lepton masses and mixing by using the observed charged lepton masses as only family-number dependent input.
parameters. The aim has been almost accomplished in the phenomenological level. Therefore, the purpose of the present paper is not to give parameter fitting of quark and lepton masses and mixing, but to take the energy scales of VEV of flavons into consideration consistently. In order to discuss the VEV scales consistently, the basic formulation of the model is drastically changed.

The new model in this paper is characterized as follows: (i) There are no Yukawaons, although the flavons $\Phi_f$ and $\hat{S}_f$ still play an essential role in the new model, too. (ii) There are three type of VEV scales. We denote these scales as $\langle \hat{S}_f \rangle \sim \Lambda_1$, $\langle \Phi_f \rangle \sim \Lambda_2$, $\langle E \rangle \sim \Lambda_3$. (1.8)

More exactly speaking, we define $\Lambda_1$ as $\Lambda_1 = v_S$ in Eq.(1.2), and $\Lambda_2$ as $\Lambda_2 = v_\Phi$ (except for $f = \nu$) in Eq.(1.5), respectively. We have also denote a scale $\Lambda_3$ as $v_E = \Lambda_3$ which appears in Eq.(2.15) later. Here and hereafter, in order to be easy to see a scale of flavons, we will sometimes show the indexes of $U(3)$ and $U(3)'$, by “◦” and “•”, instead of $i, j, k, \cdots$ and $\alpha, \beta, \gamma, \cdots$, respectively. We consider $\Lambda_1 \gg \Lambda_2 \gg \Lambda_3$. The purpose of the present paper is to fix those scales by the observed data. (iii) The most troublesome problem is VEV scales in neutrino sector. In order to make the model anomaly free, the right-handed neutrinos $\nu_R$ belong to triplet of $U(3)$ as well as $\nu_L$, so that the right-handed neutrino Majorana mass matrix ($M_R^{ij}$) are given by VEV of a flavon $\bar{Y}_R$, $\langle \bar{Y}_R \rangle$. We demand that Majorana mass matrix of the left-handed neutrinos is given by a seesaw mechanism

$$ (M^{\text{Maj}}_\nu)_{\alpha\beta} = (\hat{M}_\nu)_{\alpha\cdot} \langle \bar{Y}_R^{-1} \rangle_{\alpha\cdot} (\hat{M}_T^{\text{Maj}}\nu)_{\cdot\beta}. \quad (1.9) $$

However, a scale of the right-handed neutrino Majorana mass matrix is considerably small differently from the conventional neutrino seesaw [5], because our flavon seems to have $\langle \bar{Y}_R \rangle \sim \Lambda_3$. We will squarely grapple with this problem.

In Sec.2, we discuss our Dirac mass matrix forms which are given in Eq.(1.4). In Sec.3, we discuss relations between phase parameters $\phi_i$ given in Eq.(1.5) and the charged lepton masses $m_{ei}$. That is, we show that the apparently “family-number dependent parameters $\phi_i$ can be described in terms of charged lepton mass parameters ($m_e, m_\mu, m_\tau$). This was first pointed out in Ref.[7]. In our present model, since the VEV relations are completely revised, the relations between $\phi_i$ and $m_{ei}$ are also completely different from previous ones. The new relations lead us not only to that the relations give an understanding of the values $\phi_i$ on the basis of the charged lepton masses $m_{ei}$, but also to that we can predict the scales $\Lambda_1$, $\Lambda_2$ and $\Lambda_3$ by using the result of data-fitting in the $\phi_i$-$m_{ei}$ relations. In Sec.4, we discuss slight VEV deviations among flavons with the same transformation of $U(3)xU(3)'$. This is not essential to predict $\Lambda_2/\Lambda_1$ and $\Lambda_3/\Lambda_2$. However, in order to obtain more consistency in our scenario, we will discuss somewhat detailed VEV relations. In Sec.5, we discuss a Majorana mass matrix form of the right-handed neutrinos $\nu_R$ as a preparation for predicting the scales $\Lambda_1$, $\Lambda_2$ and $\Lambda_3$. In Sec.6, we discuss the VEV scales
Finally, Sec. 7 is devoted to concluding remarks. In Appendix A, the detail parameter fitting for the masses and mixings of quarks and leptons presented in Ref. [1, 2] is reviewed.

## 2 Dirac mass matrices

### 2.1 Seesaw-type mechanism in the Dirac mass matrix

In our model based on $U(3) \times U(3)'$ symmetry, we consider hypothetical fermions $F_\alpha$ ($\alpha = 1, 2, 3$), which belong to $(1, 3)$ of $U(3) \times U(3)'$, in addition to quarks and leptons $f_i$ ($i = 1, 2, 3$) which belong to $(3, 1)$. In this model, differently from the Yukawaon formulation, we assume that the VEV form (1.4) originates from the following $6 \times 6$ mass matrix model:

$$
\begin{pmatrix}
\bar{f}_L F_\alpha^\dagger & (0)_i^j \\
(\hat{\Phi} f)_i^j & - (S f)^\alpha_\beta
\end{pmatrix}
\begin{pmatrix}
\bar{f}_{Rj} \\
F_{R\beta}
\end{pmatrix}.
$$

Here, $F_{L(R)}$ are heavy fermions with $(1, 1, 3)$ of $SU(2)_L \times U(3) \times U(3)'$. On the other hand, $f_R$ are right-handed quarks and leptons, $f_R = (u, d, \nu, e^-)_R$, while $f_L$ are not physical fields. They are given by the following combinations:

$$f_L \equiv (f_u, f_d, f_\nu, f_e)_L \equiv \left( \frac{1}{L_U} \begin{pmatrix} H^u_d q_L & H^d_u q_L & H^\nu_\ell L & H^\ell_\nu L \end{pmatrix} \right)
\left( \begin{array}{c} f_R \\ F_{R\beta} \end{array} \right)
\tag{2.2}
$$

where

$$
q_L = \begin{pmatrix} u_L \\ d_L \end{pmatrix}, \quad \ell_L = \begin{pmatrix} \nu_L \\ e^-_L \end{pmatrix}, \quad H_u = \begin{pmatrix} H^u_u \\ H^0_u \end{pmatrix}, \quad H_d = \begin{pmatrix} H^+_d \\ H^0_d \end{pmatrix}
\tag{2.3}
$$

In other words, the matrix given in Eq.(2.1) denotes would-be Yukawa coupling constants.

After the $U(3)$ and $U(3)'$ have been completely broken, the quarks and leptons are described by the effective Hamiltonian

$$
\mathcal{H}_Y = (\bar{\nu}_L)^i_j (\hat{M}_\nu)_i^j (\nu_R)_j + (\bar{e}^0_i (\hat{M}_e)_i^j (e_R)_j + y_R (\bar{\nu}_R)^i_j (Y_R) (\nu_R)_j + (\bar{u}_L)^i_j (\hat{M}_u)_i^j (u_R)_j + (\bar{d}_L)^i_j (\hat{M}_d)_i^j (d_R)_j.
\tag{2.4}
$$

Of course, $(\hat{M}_f)_i^j$ do not mean flavons although we have called those “Yukawaons” in the past Yukawaon model [3]. Note that the quarks and leptons $f_i$ are not $U(3)$ family triplet any more in the exact meaning, but they are mixing states between $f$ and $F$. However, we will still use the index of $U(3)$ family for these fermion states. Also, note that there are no Yukawaons in the present model. By performing a seesaw-like approximation with $\Lambda_2 \ll \Lambda_1$, the mass matrix
(2.1) leads to the following Dirac mass matrices of quarks and leptons:

\[
(\tilde{M}_f)^{ij}_f \simeq \frac{(H_u/d)_i}{\Lambda_H} (\Phi_f)_j^{\alpha} \langle (\hat{S}_f)^{-1} \rangle_\alpha^{\beta} (\bar{\Phi}_f)_j^\beta.
\]  

(2.5)

In this paper, we search VEV scales of flavons by assuming that our flavon VEV scale is given by one of \(\Lambda_1, \Lambda_2\) and \(\Lambda_3\), except for \(H_u/d\). As we stated in subsection 2.3, we put \(1/(\Lambda_1)^n\) for superpotential term with dimension \((3 + n)\). However, we assume that these rules are exceptional for the factor \(H_u/d/\Lambda_H\). The VEV of \(H_u\) and \(H_d\) are fixed as

\[
v_{H_u} = v_{H_d} \equiv v_H \equiv \frac{1}{\sqrt{2}} \times 246 \text{ GeV}.
\]

(2.6)

(The reason of \(v_{H_u} = v_{H_d}\) will be given in Eq.(4.6) later.) We will take \(\Lambda_H\) as a different value from \(\Lambda_1, \Lambda_2\) and \(\Lambda_3\). That is, we consider that the additional factor \(H_u/d/\Lambda_H\) in Eq.(2.5) is factor at the electroweak scale after our family symmetries \(U(3) \times U(3)\) are completely broken.

2.2 \(R\)-charge assignment of \(\Phi_f\)

In this section, we will often mention an effective superpotential. However, differently from the conventional SUSY scenario, our SUSY is already broken at a high energy scale, for example, \(\mu \sim 10^7\) TeV. Nevertheless, we assume that interaction forms at the unbroken SUSY scale are still kept in low energy (smaller than \(\mu < 10^4\) TeV) phenomenology.

We assume that \(R\) charges are still conserved under the block diagonalization of (2.1), so that we consider an \(R\) charge relation

\[
R(\hat{S}_f) = R(\Phi_f) \equiv r_f.
\]

(2.7)

Here we denote \(R\) charge of the flavon \(\hat{S}_f\) as \(R(\hat{S}_f)\) and so on. We have taken \(R\) charge of the Higgs fields as \(R(H_{u/d}) = 0\). Then, the Dirac mass matrix \(\tilde{M}_f\) has effectively a \(R\) charge \(r_f\) (although \(\tilde{M}_f\) is not a flavon). Note that if there is a flavon combination \((\Phi_f \Phi_{f'})^j_1/\Lambda_1\) with \(r_f + r_{f''} = r_f\), the combination can couple to \(\hat{f}_L^j f_{Ri}\) as well as \((\tilde{M}_f)^{ji}_f\) with the same energy scale \((\Lambda_2)^2/\Lambda_1\). Besides, if \((\Phi_f \Phi_{f'})\) and another \((\Phi_{f'} \Phi_{f''})\) have the same \(R\) charge, those can have the same VEV structure. This is an unwelcome situation, because we demand that \(\Phi_u\) and \(\Phi_{\nu}\) have somewhat different VEV structures from \(\Phi_e\) and \(\Phi_d\) as we state later. A simple way to avoid appearance of such unwelcome combinations is to take \(R\) charge difference among \(\Phi_f\)'s completely different from each other: \(r_e - r_u = 2/3, r_d - r_e = 3/3, r_{\nu} - r_d = 4/3\), i.e.

\[
(r_u, r_e, r_d, r_{\nu}) = \left(\frac{1}{3}, 1, 2, \frac{10}{3}\right).
\]

(2.8)
The reason $r_u < r_e$ will also be discussed later.

The reason why we choose $r_e = 1$ is as follows: Only when $r_e = 1$, we can write down a superpotential for $\hat{S}_e$,

$$W_{Se} = \lambda_1^S [(\hat{S}_e)_\alpha^\beta (\hat{S}_e)^\alpha_\beta] + \lambda_2^S [(\hat{S}_e)_\alpha^\alpha][(\hat{S}_e)^\beta_\beta].$$

Supersymmetric vacuum condition for this $W_{Se}$ leads to $\hat{S}_e = v_S \mathbf{1}$, so that we obtain $b_e = 0$. This means that the charged lepton mass matrix $\hat{M}_e$ is proportional to $\text{diag}(z_1^2, z_2^2, z_3^2)$. Therefore, our parameters $z_i$ are given by (1.7). The phase parameters $\phi_i^e$ do not have physical meaning because $\langle \Phi_e \rangle$ is commutable with $\langle \hat{S}_e \rangle = v_S \mathbf{1}$, so that we can hereafter put $\phi_i^e = 0$, i.e. $\langle \Phi_e \rangle = v_e \text{diag}(z_1, z_2, z_3)$.

In Table.1 we list $R$ charges of leading flavons.

### 2.3 VEV relation among $\Phi_f$

In this paper, we will often give effective superpotential terms. We write down those effective superpotentials according to the following rules: (i) For effective superpotential terms with a higher mass-dimension $(n + 3)$, we put a factor $(1/\Lambda_1)^n$ in order to adjust the term to the dimension 3. (ii) In order to avoid unwelcome terms with higher dimension, we assume $R$ charge conservation and positive $R$ charge assignment (except for special flavons $\Theta$ as we show later), so that superpotential terms with $R > 2$ are forbidden.

A VEV of $\Phi_d$ is obtained by the following superpotential

$$W_d = \left\{ (\Phi_d)_i^\beta (\hat{E})_\beta^\alpha + \frac{1}{\Lambda_1} (\Phi_e)_i^\beta (\hat{S}_e)_\beta^\gamma (\hat{E})_\gamma^\alpha \right\} (\hat{\Theta})_\alpha^i.$$  

(2.10)

Here and hereafter, we assume that the $\Theta$ fields always take $\langle \Theta \rangle = 0$. Therefore, from $\partial W_d / \partial \Theta_d = 0$, we obtain

$$(\Phi_d)_i^\beta (\hat{E})_\beta^\alpha = -\frac{1}{\Lambda_1} (\Phi_e)_i^\beta (\hat{S}_e)_\beta^\gamma (\hat{E})_\gamma^\alpha.$$  

(2.11)

Therefore, when the scale of $(\hat{E})_\gamma^\alpha$ is $\Lambda_1$, we obtain the VEV $\langle \Phi_d \rangle$ with the same order as $\langle \Phi_e \rangle$, i.e. $\langle \Phi_d \rangle = v_d \text{diag}(z_1, z_2, z_3)$ with $\phi_i^d = 0$.

However, for $\Phi_u$ with $R = 1/3$, we have to consider another type of superpotential,

$$W_u = \left\{ \frac{\lambda_{u1}}{\Lambda_1} (\Phi_u)_\alpha^k (\Phi_u)_k^\gamma (\hat{E})_\gamma^\beta + \frac{\lambda_{u2}}{\Lambda_1} (\Phi_d)_\alpha^k (\Phi_d)_k^\beta \right\} (\hat{\Theta}_u)_\beta^\alpha.$$  

(2.12)

Here, we have introduced a new flavon $\hat{E}_\bullet$ with $R = 5/3$ and a VEV form $\langle \hat{E} \rangle = v_E \mathbf{1}$. Since this potential gives a relation between $\Phi_u \Phi_u$ and $\Phi_d \Phi_d$, $\langle \Phi_u \rangle$ can have phase factors differently from real VEV matrix $\langle \Phi_d \rangle$. Hereafter, we simply denote phase parameters $\phi_i^u$ as $\phi_i$:

$$\langle \Phi_u \rangle = v_u \text{diag}(z_1 e^{i\phi_1}, z_2 e^{i\phi_2}, z_3 e^{i\phi_3}).$$  

(2.13)
Table 1: R charges and VEV scales of leading flavons. Transformation property under U(3)×U(3) is indicated by “•” and “◦”, respectively.

| flavon | (Φ_u)◦ | (Φ_e)◦ | (Φ_d)◦ | (Φ_ν)◦ | (S_u)◦ | (S_e)◦ | (S_d)◦ | (S_ν)◦ |
|--------|---------|---------|---------|---------|---------|---------|---------|---------|
| R      | 1/3     | 3/3     | 3/3     | 10/3    | 1/3     | 3/3     | 9/3     | 10/3    |
| Scale  | Λ_2     | Λ_2     | Λ_2     | Λ_3     | Λ_1     | Λ_1     | Λ_1     | Λ_2     |

| E_oo  | E_•    | E_◦    | E_′    | (Y_eu)◦ | (S_u)◦ | (Y_R)◦ |
|-------|--------|--------|--------|---------|--------|--------|
| 2/3   | 2/3    | 1/3    | 5/3    | 5/3     | 13/3   |        |
| Λ_3   | Λ_3    | Λ_2    | Λ_1    | Λ_1     | Λ_2    | Λ_3    |

| (Θ_d)◦ | (Θ_u)◦ | (Θ_ν)◦ | (Θ_e)◦ | (Θ_R)◦ | (Θ_eu)◦ | (Θ_S_u)◦ | (Θ_E)◦ | (Θ_E)◦ |
|--------|--------|--------|--------|--------|--------|----------|--------|--------|
| -1/3   | -2/3   | -4/3   | -4/3   | -3     | 2/3    | -1       | -3/3   | 2/3    |
| 0      | 0      | 0      | 0      | 0      | 0      | 0        | 0      | 0      |

Meanwhile, in this model, flavons E with VEV matrix form \( v_E \text{diag}(1, 1, 1) \) frequently appear. For \( E \) with the VEV scale \( \Lambda_3 \) and \( R = 2/3 \), we consider a superpotential given by

\[
W_E = \lambda_1 \text{Tr}[\hat{E}_o^\circ E_oo \bar{E}^\circ] + \lambda_2 \text{Tr}[\bar{E}_o^\circ \text{Tr}[E_oo \bar{E}^\circ]].
\]

SUSY vacuum condition leads to

\[
\hat{E}_o^\circ = v_E 1, \quad E_oo \bar{E}^\circ = (v_E)^2 1.
\]

For \( E_oo \), we take a special solution \( E_oo = v_E 1 \) in the relation \( E_oo \bar{E}^\circ = (v_E)^2 1 \). For \( \hat{E}_o^\circ \) and \( E_• \), there is not such a simple superpotential. We have to assume the following superpotential

\[
W_E = \lambda_1 \text{Tr}[\hat{E}_o^\circ \hat{E}_o^\circ (\hat{\Theta}_{-4/3})^\circ] + \lambda_2 \text{Tr}[E_oo E_•^\circ (\hat{\Theta}_{-4/3})^\circ] + \ldots,
\]

where \( (\hat{\Theta}_{-4/3})^\circ \) has \( R = -4/3 \). We also assume

\[
W_E = \text{Tr}[\hat{E}_o^\circ E_oo \bar{E}^\circ (\hat{\Theta}_{2/3})^\circ] + \ldots,
\]

where \( (\hat{\Theta}_{2/3})^\circ \) has \( R = 2/3 \).

Finally, let us discuss a VEV relation of \( \langle \Phi_\nu \rangle \). We cannot consider the similar mechanism as \( \langle \Phi_f \rangle \) using \( \hat{E} \), since we need to give a small scale compared with \( \langle \Phi_e \rangle \sim \Lambda_2 \) in order to satisfy the seesaw approximation (1.9). (Since the scale \( \Lambda_2 \) means the maximal scale of a flavon with \( A_\circ \), the requirement \( \langle \Phi_\nu \rangle < \Lambda_2 \) has no problem.) Therefore, we assume the following superpotential term:

\[
W_\nu = \left\{ \mu_\nu (\Phi_\nu)\circ + \lambda_\nu (\frac{1}{\Lambda_1})^2 (\Phi_\nu)\circ \hat{E}^\circ E_oo \bar{E}^\circ \right\} (\hat{\Theta}_2)^\circ,
\]
where new flavons $E_{oo}$ and $E_{••}$ with VEV form $v_E 1$ has $R$ charges

$$R(E_{oo}) = \frac{2}{3}, \quad R(E_{••}) = \frac{5}{3}, \quad R(E_{o•}) = \frac{1}{3}. \quad (2.19)$$

Therefore, a VEV scale of $(\Phi_{ν})_{••}$ with $R = 10/3$ is given by

$$\|\Phi_{ν}\| = \frac{(Λ_2)^3}{μ_ν Λ_1}, \quad (2.20)$$

differently from other $(\Phi_f)_{••}$. Hereafter, for convenience, we denote a VEV scale of a flavon $A$ as a notation $\|A\|$. Also, in the estimation of VEV scales, for simplicity, we put $λ = 1$ for all dimensionless coefficients $λ$ in superpotentials (2.9), (2.10), and so on. We denote Eq.(2.20) as

$$\langle \Phi_{ν} \rangle = ξ_{ν} \langle \Phi_e \rangle, \quad ξ_{ν} = \frac{(Λ_2)^2}{μ_ν Λ_1}. \quad (2.21)$$

We have obtained $b_e = 0$, i.e. $\langle \hat{S}_e \rangle = v_S 1$ by assigning $r_e = 1$ as shown in Eq.(2.9). However, since $\hat{S}_ν$ has $R = 10/3$, we cannot assert $b_ν = 0$ by means of a similar way to Eq.(2.9). We want a similar VEV structure of $\hat{M}_ν$ to $\hat{M}_e$ except for its VEV scale. Therefore, we assume a superpotential similar to (2.18):

$$W_{Sν} = \left\{ μ_{Sν}(\hat{S}_ν)_{••} + λ_ν \frac{1}{(Λ_1)^2}(\hat{S}_e)_{••} E_{••} E_{o•} \right\}(\hat{Θ}_ν)_{••}. \quad (2.22)$$

Then, we obtain a result similar to (2.21):

$$\langle \hat{S}_ν \rangle = ξ_{Sν} \langle \hat{S}_e \rangle, \quad ξ_{Sν} = \frac{(Λ_2)^2}{μ_{Sν} Λ_1}, \quad (2.23)$$

so that we can obtain $\hat{M}_ν$ with the same form as $\hat{M}_e$:

$$\hat{M}_ν = ξ_M \hat{M}_e, \quad ξ_M = (ξ_ν)^2 (ξ_{Sν})^{-1} = \frac{μ_{Sν}(Λ_2)^2}{(μ_ν)^2 Λ_1}. \quad (2.24)$$

2.4 Parameter values in the quark sector

As far as quark sector is concerned, the VEV matrices are exactly the same as those in the previous paper, although the model is completely different from the previous one.
We choose parameter values\footnote{In the parameter fitting, we have used running masses at $\mu = m_Z$ as input quark and lepton masses.} as
\[
    b_u = -1.011, \quad b_d = -3.522 e^{i 17.7^\circ}, \quad (2.25)
\]
and
\[
    (\tilde{\phi}_1, \tilde{\phi}_2) = (-176.05^\circ, -167.91^\circ), \quad (2.26)
\]
where we have fitted $(\phi_1, \phi_2, \phi_3)$ as $(\tilde{\phi}_1, \tilde{\phi}_2, 0)$ without losing generality. Then, we can obtain reasonable quark masses and Cabibbo-Kobayashi-Maskawa (CKM) mixing \footnote{In the parameter fitting, we have used running masses at $\mu = m_Z$ as input quark and lepton masses.} as shown in the previous papers \cite{1, 2} as shown in Appendix A. Note that the parameter $\phi_0$ cannot be fixed in the CKM fitting, i.e, it is an unobservable parameter as we discuss in the next section.

3 Relation between phase parameters $\phi_i$ and charged lepton masses $m_{ei}$

The parameters $(\phi_1, \phi_2, \phi_3)$ were typical family-number dependent parameters. Now we try to describe $(\phi_1, \phi_2, \phi_3)$ in terms of charged lepton masses $m_{ei}$ and two family-number independent parameters. This has first been pointed out by the authors \cite{7}. However, since the $R$-charge assignment in the present model is completely different from that in the previous model, the relations between the phase parameters $\phi_i$ and charged lepton masses $m_{ei}$ is also changed. New relations will be simply denoted with fractional coefficients and thereby it will make possible to estimate the scales of $U(3) \times U(3)'$.

We assume the following superpotential
\[
    W_\phi = \left\{ \frac{\lambda_1}{A_1} \left[ (\Phi_u)_{\circ}^\circ (\bar{E})_{\bullet}^\circ + h.c. \right] + \frac{\lambda_2}{(A_1)^2} (\Phi_u)_{\circ}^\circ (E)_{\circ} (\bar{E})_{\bullet}^\circ (\Phi_u)_{\circ}^\circ \right. \]
\[
    \left. + \frac{\lambda_2}{(A_1)^3} (\Phi_u)_{\circ}^\circ (\Phi_u)_{\circ}^\circ (\Phi_u)_{\circ}^\circ \right\} (\hat{\Theta}_\phi)_{\bullet}^\circ; \quad (3.1)
\]
where $R(\Theta_\phi) = 2/3$. Note that in addition to the second term $(\Phi_u)_{\circ}^\circ (E)_{\circ} (\bar{E})_{\bullet}^\circ (\Phi_u)_{\circ}^\circ$ in Eq.(3.1), the following additional terms are allowed: $(\Phi_u)_{\circ}^\circ (\Phi_u)_{\circ}^\circ (\bar{E})_{\bullet}^\circ (\Phi_u)_{\circ}^\circ$, $(\Phi_u)_{\circ}^\circ (E)_{\circ} (\bar{E})_{\bullet}^\circ (\Phi_u)_{\circ}^\circ$, $(\Phi_u)_{\circ}^\circ (\Phi_u)_{\circ}^\circ (E)_{\circ} (\bar{E})_{\bullet}^\circ$, and $(E)_{\circ} (\bar{E})_{\bullet}^\circ (\Phi_u)_{\circ}^\circ (\Phi_u)_{\circ}^\circ$. Here, some remarks are in order: (i) we regard those additional terms as “substantially same terms”, (ii) but, we count a flavon $A^\dagger$ as a different field from $A$, and (iii) the coefficient $\lambda$ is defined as follows: the $\lambda$ is a coefficient with a factor $1/n$ for sum of $n$ substantially same terms. For example, in the second term in (3.1), the factor $\lambda_2$ is defined as one for sum of the five terms with $1/5$. However, for simplicity, we denote only representative one even if there are many equivalent terms, and give the coefficient $\lambda$ instead of $\lambda/n$. Also note that the $h.c.$ term in the first term is different from the original one according to our counting rule.
The flavon $\hat{\Theta}_{\phi}$ in (3.1) has VEV value of zero. The SUSY vacuum condition $\partial W/\partial \Theta = 0$ leads to a condition

$$
2c_1 z_i \cos \phi_i = c_2 z_i^2 + c_3 z_i^4,
$$

(3.2)

where parameters $c_1$, $c_2$ and $c_3$ are family-number independent parameters and they have scales

$$
c_1 = \frac{(\Lambda_2)^2}{(\Lambda_1)^2}, \quad c_2 = \frac{(\Lambda_2)^4}{(\Lambda_1)^2}, \quad c_3 = \frac{(\Lambda_2)^4}{(\Lambda_1)^2}.
$$

(3.3)

When we denote these parameters $\phi_i$ as

$$
\phi_1 = \phi_0 + \tilde{\phi}_1,
\phi_2 = \phi_0 + \tilde{\phi}_2,
\phi_3 = \phi_0,
$$

(3.4)

the parameter $\phi_0$ is unobservable in the CKM parameter fitting, while it is not unobservable in the $U(3) \times U(3)'$ model. From the input values $(\tilde{\phi}_1, \tilde{\phi}_2) = (-176.05^\circ, -167.91^\circ)$ given in (2.26), we obtain

$$
\phi_0 = 86.69^\circ, \quad \frac{c_2}{c_1} = 1.368, \quad \frac{c_3}{c_2} = -0.967,
$$

(3.5)

which leads to

$$
(\phi_1, \phi_2, \phi_3) = (-89.36^\circ, -87.25^\circ, 86.69^\circ).
$$

(3.6)

Thus, the family-number dependent parameters $(\phi_1, \phi_2, \phi_3)$ can be reduced into family-number independent parameters $(c_2/c_1, c_3/c_2)$. (Note that the parameter $\phi_0$ is not unobservable any longer in this model.)

Note that the numerical results (3.5) suggests $|c_3/c_2| = 1$ if we take $\lambda_2 = \lambda_3$. Then, it is natural that we consider $\lambda_1 = \lambda_2 = \lambda_3$. Only the ratio $c_2/c_1$ is not one. When we compare (3.3) with the numerical result $c_2/c_1 \equiv \rho \simeq 1.37$ given in (3.5), we obtain

$$
R_\Lambda \equiv \frac{\Lambda_2}{\Lambda_1} = \rho \frac{\Lambda_2}{\Lambda_2}.
$$

(3.7)

We will see $R_\Lambda \sim 10^{-3}$ in Sec.5.

4 Slight deviation of the scales in the lepton sector

4.1 $\hat{M}_{\text{lepton}}$ versus $\hat{M}_{\text{quark}}$
Let us comment on the ratio $v_{Hu}/v_{Hd}$. Usually, it is understood that the value of $\tan \beta \equiv v_{Hu}/v_{Hd}$ is $\tan \beta \sim 10 - 40$. However, from (2.5), our quark masses are predicted as

\[
(m_u, m_c, m_t) = (0.0003964, 0.106411, 29.743) m_{0u},
\]

\[
(m_d, m_s, m_b) = (0.0007249, 0.01467, 0.5365) m_{0d},
\]

(4.1)

where $m_{0f}$ are defined by

\[
m_{0f} = \frac{v_{Hf}(v_f)^2}{\Lambda H v_{Sf}} \frac{v_{Hf}}{\Lambda H} \rho \Lambda_3.
\]

(4.2)

In the last term in Eq.(4.2), we have used the relation (3.7). The numerical values in (4.1) are eigenvalues of the dimensionless matrix

\[
Z(1 + b_f X_3)^{-1} Z,
\]

(4.3)

where

\[
Z = \text{diag}(z_1, z_2, z_3),
\]

(4.4)

and we have used input values $b_u = -1.011$ and $b_d = -3.3522 e^{i17.7^\circ}$ [1].

The prediction (4.1) can give reasonable values not only for mass ratios $m_u/m_c$, $m_c/m_t$, $m_d/m_s$, and $m_s/m_b$, but also for the mass ratio of up-quark mass to down-quark mass

\[
\frac{m_t}{m_b} = 55.44,
\]

(4.5)

if we take $m_{0u} = m_{0d}$. The value (4.5) roughly agrees with the observed value [1] $(m_t/m_b)^{obs} \simeq 59.41$. Therefore, in the present model, we can regard

\[
\frac{m_{0u}}{m_{0d}} = 1, \quad i.e \quad v_{Hu} = v_{Hd}.
\]

(4.6)

Therefore, we have already taken $v_{Hu} = v_{Hd}$ in Eq.(2.6).

Also, since $m_\tau = (z_3)^2 m_{0e} (z_3 = 0.97170)$, we obtain

\[
m_{0e} = 1.8499 \text{ GeV},
\]

(4.7)

so that we get

\[
\frac{m_{0u}}{m_{0e}} = 3.121.
\]

(4.8)

This suggests

\[
\frac{m_{0u}}{m_{0e}} = 3,
\]

(4.9)
in contrast to Eq.(4.6). We consider that such the factor 3 originates in a slight difference
between flavon VEV scales in the lepton sector and the quark sector. In the next subsection,
we will discuss such an additional VEV scale difference.

4.2 Slight deviation of the scale $\|\Phi_{\text{lepton}}\|$ from $\|\Phi_{\text{quark}}\|$

So far, we have not mentioned the origin of the parameter value of $\rho$ defined by Eq.(3.7)
and the factor 3 in Eq.(4.9). In order to understand those numerical factors, in this subsection,
we define additional scale deviation factors of $\Phi_\ell$ and $\hat{S}_\ell$ ($\ell = e, \nu$ and $q = u, d$) from $\Phi_q$ and $\hat{S}_q$, respectively. However, since the deviations are very small (smaller than $O(1)$), arguments
in this subsection will not give an essential influence on our main purpose which is to estimate
the scales of $U(3) \times U(3)'$.

We consider the following VEV deviation factors, $\eta_\Phi$ and $\eta_S$, of the lepton sector from the
quark sector:

$$
\|(\Phi_\ell)_\circ\| = \eta_\Phi \Lambda_2, \quad \|(\hat{S}_\ell)_\circ\| = \eta_S \Lambda_1, \tag{4.10}
$$

which is slightly different from $\|(\Phi_q)_\circ\| = \Lambda_2$ and $\|(\hat{S}_q)_\circ\| = \Lambda_1$. Here, we have consider that factors $\eta_\Phi$ and $\eta_S$ are orders of one in contrast to factors $\xi_\nu$ and $\xi_{S\nu}$ with orders of $10^{-3}$ as shown later.

The modification (4.10) gives

$$
\|(\bar{M}_\ell)_\circ\circ\| = (\eta_\Phi)^2 (\eta_S)^{-1} \|(\hat{M}_q)_\circ\|, \tag{4.11}
$$

so that (4.9) demands

$$
\frac{(\eta_\Phi)^2}{\eta_S} = \frac{1}{3}, \quad \text{i.e.} \quad 3(\eta_\Phi)^2 = \eta_S. \tag{4.12}
$$

On the other hand, the VEV relation (2.11) does not hold unless $\|(\Phi_\ell)_\circ\| \cdot \|(\hat{S}_\ell)_\circ\| = \Lambda_2 \Lambda_1$. Therefore, we put additional relation

$$
\eta_\Phi \eta_S = 1, \tag{4.13}
$$

so that we obtain

$$
3(\eta_\Phi)^3 = 1 \quad \Rightarrow \quad \eta_\Phi = \frac{1}{\eta_S} = \frac{1}{3^{1/3}}. \tag{4.14}
$$

As a result, we can give the parameter value $\rho$ defined in Eq.(3.7) as $3^{1/3} = 1.44$ as seen in
Eq.(5.30).

5 Flavons in the neutrino sector

We discuss Majorana mass matrix, $Y_R$, of the right-handed neutrinos $\nu_R$. Note that the
mass matrix $\tilde{Y}_R$ is $(6^*, 1)$ of $U(3) \times U(3)'$, i.e. $(\tilde{Y}_R)^\circ$. 

13
In the present neutrino mass matrix model, the Dirac neutrino mass matrix $\hat{M}_\nu$ is given by Eq.(2.5), i.e.

$$\langle \hat{M}_\nu \rangle_\circ = \langle H_u \rangle \langle \Phi_\nu \rangle_\circ \langle \hat{S}_\nu^{-1} \rangle_\circ \langle \Phi_\nu \rangle_\circ.$$

(5.1)

In this section, since we pay attention to VEV scales of flavons, it is important whether those flavons belong to $U(3)$ or $U(3)'$. From Eq.(5.1), we find that the scale of $\langle \hat{M}_\nu \rangle_\circ$, $\|\langle \hat{M}_\nu \rangle_\circ\|$, is given by

$$\|\hat{M}_\nu\| \equiv \xi_M \|\hat{M}_\nu\| = \xi_M \langle H_u \rangle \langle \Lambda_2 \rangle^2 \Lambda_1. \quad (5.2)$$

Since $\|\hat{M}_\nu\|$ is given by the order of $\langle (H_u)/\Lambda_H \rangle (\Lambda_2)^2/\Lambda_1$ and we consider an additional seesaw (1.9), a ratio of the scales $\|\hat{M}_\nu^{\text{Major}}\|/\|\hat{M}_\nu\|$ is given by

$$R_{\nu/e} \equiv \frac{\|\hat{M}_\nu^{\text{Major}}\|}{\|\hat{M}_\nu\|} = (\xi_M)^2 \frac{v_H (\Lambda_2)^2}{\Lambda_1} \frac{1}{\|Y_R\|}, \quad (5.3)$$

where $v_H = v_{Hd} \equiv v_H$ is defined in (2.6). In order to estimate the ratio (5.3), we have to build a model for $Y_R$.

5.1 VEV structure of Majorana mass matrix $\hat{Y}_R$

For convenience, we denote the form $\langle Y_R \rangle$ by the following terms

$$\mu_R(\hat{Y}_R)^{\circ\circ} = \mu_R \left[ (\hat{Y}_R^{1st})^{\circ\circ} + (\hat{Y}_R^{2nd})^{\circ\circ} \right], \quad (5.4)$$

with $\|\langle Y_R^{1st} \rangle\| \gg \|\langle Y_R^{2nd} \rangle\|$. In Eq.(5.4), we denote the first and second terms in $Y_R$ as $\langle Y_R^{1st} \rangle$ and $\langle Y_R^{2nd} \rangle$, but it does not mean that they are two new flavons. Considering that the mass hierarchy in the neutrino sector is mild compared with other sectors $f = u, d, e$, we assume that the first term of $M_\nu^{\text{Major}}$, $\hat{Y}_\nu (\hat{Y}_R^{1st})^{-1} \hat{Y}_\nu$ takes a form of diagonal matrix, and the observed mass differences and the PMNS mixing come form in the second term $\hat{Y}_R^{2nd}$. Since $\langle \hat{Y}_\nu \rangle \propto Z^2$ ($Z$ is defined in (4.4)), we take a form $\hat{Y}_R^{1st} \propto Z^2$.

Since we want that the $R$ charge of $\hat{Y}_R$ is as smaller as possible in order to avoid appearance of many combinations with the same $R$ charge, we take a form of $\langle Y_R^{1st} \rangle^{\circ\circ}$ with $R = 13/3$, $(\hat{\Phi}_e^T)^\circ \hat{E}_\circ (\hat{\Phi}_e)_\circ (\hat{\Phi}_u)_\circ$, i.e.

$$\mu_R(\hat{Y}_R^{1st})^{\circ\circ} = \frac{1}{2} \langle \Lambda_1 \rangle^2 \left\{ (\hat{\Phi}_e^T)^\circ \hat{E}_\circ (\hat{\Phi}_u)_\circ + (\text{transposed}) \right\}. \quad (5.5)$$

Here, in order to adjust the scale of $\hat{Y}_R^{1st}$ compared with $\hat{Y}_R^{2nd}$, we introduce a new flavon $\hat{Y}_{eu}$ defined by

$$\mu_{eu}(\hat{Y}_{eu})_\circ = (\hat{\Phi}_e)_\circ (\hat{\Phi}_u)_\circ. \quad (5.6)$$
Then, we obtain
\[
\mu_R \|(\bar{Y}_{R}^{1st})^\circ\circ\| = (\eta\Phi)^2 \frac{A_1}{\mu_{eu}} \frac{(A_2)^4}{(A_1)^2}. \tag{5.7}
\]

Next, we discuss a term \(\bar{Y}_{R}^{2nd}\). We suppose that the deviation term \(\bar{Y}_{R}^{2nd}\) gives Pontcorvo-Maki-Nakagawa-Sakata (PMNS) \[10\] mixing and neutrino mass ratios. Since we want to inherit the form of \(\langle \bar{Y}_R \rangle\) from the previous model, in which the form of \(Y_R\) has included a term proportional to up-quark mass matrix \(\hat{M}_u\) \[11\]. However, in the present model, there is no Yukawaon \(\hat{Y}_u\). Therefore, we introduce a new flavon whose VEV is proportional to \(\langle \hat{S}_{u}^{-1} \rangle\) with use of a superpotential given by
\[
W_{S_u'} = \left\{ \bar{E}^\circ\circ (\hat{E})^\circ + (\hat{S}_{u}')^\circ\circ (\hat{S}_u)^\circ\circ \right\} (\Theta_{S_u}'). \tag{5.8}
\]
where, for simplicity, we have drop the coefficients \(\lambda_1\) and \(\lambda_2\) although we suppose \(\lambda_1 \simeq \lambda_2 \simeq 1\). Here, since \(R(\hat{E}^\circ\circ) = 10/5\) and \(R(E_{o}^\circ) = 1/3\), the flavon \(\hat{S}_{S_u}'\) has zero VEV value and \(R\) charge \(R = 0\). The flavon \((\hat{S}_{u}')^\circ\circ\) with \(R = 5/3\) can obtain the VEV form proportional to \(\langle \hat{S}_{u}^{-1} \rangle\). Thus, we take a small deviation term \(\bar{Y}_{R}^{2nd}\) as follows:
\[
\mu_R (\bar{Y}_{R}^{2nd})^\circ\circ = \frac{1}{2} \frac{1}{(A_1)^2} \left\{ (\hat{\Phi}_d)^\circ\circ (\hat{\Phi}_u^T)^\circ\circ (\hat{S}_{u}')^\circ\circ (\hat{\Phi}_u)^\circ\circ + (\text{transposed}) \right\}. \tag{5.9}
\]
Then, we estimate of a scale of (5.9) as
\[
\mu_R \|(\bar{Y}_{R}^{2nd})\| = \frac{(A_2)^4}{(A_1)^2}. \tag{5.10}
\]
Therefore, we obtain the ratio
\[
R_{2/1} \equiv \frac{\|\bar{Y}_{R}^{2nd}\|}{\|\bar{Y}_{R}^{1st}\|} = \frac{\mu_{eu}}{(\eta\Phi)^2 A_1}. \tag{5.11}
\]

### 5.3 Parameter \(\xi_R\)

Parameter fitting for neutrino data such as neutrino masses and PMNS lepton mixing matrix is done under the following dimensionless re-expression:
\[
\tilde{M}_{\nu}^{Maj} = (\xi_M)^4 Z^2 \bar{Y}_R^{-1} Z^2, \tag{5.12}
\]
where
\[
\bar{Y}_R = Z^4 + \xi_R Z^2 P(1 + b_u X_3)^{-1} P^d Z, \tag{5.13}
\]
\[ Z = \text{diag}(z_1, z_2, z_3), \]
\[ P = \text{diag}(e^{i\tilde{\phi}_1}, e^{i\tilde{\phi}_2}, 1). \]  
(5.14)

\( Z \) and \( P \) do not mean new flavons. Those are nothing but dimensionless \( 3 \times 3 \) matrices. The parameter \( \xi_R \) corresponds to the ratio \( R_{2/1} \) defined in (5.11). Since the parameter values \( b_u \) and \( (\tilde{\phi}_1, \tilde{\phi}_2) \) have already determined from the CKM fitting as shown in (2.25) and (2.26), only a free parameter in the neutrino sector is \( \xi_R \) in (5.13). The best fitting value of \( \xi_R \) is [1, 2]

\[ \xi_R = 0.9806 \times 10^{-3}. \]  
(5.15)

The detail fitting is reviewed in Appendix A.

5.4 Scales of \( \mu \) parameters

We have four flavon mass parameters \( \mu_\nu, \mu_{S\nu}, \mu_{eu} \) and \( \mu_R \) defined by (2.18), (2.22), (5.6) and (5.5), respectively. So far, we have considered three VEV scales \( \Lambda_1, \Lambda_2 \) and \( \Lambda_3 \) as shown in (1.8) and also in Table 1. Therefore, let us put the following selection rules for \( \mu \) parameters defined by

\[ \mu_A \| A \| = \| G \|, \]  
(5.16)

where \( G \) is a combination of some flavons including factor \( 1/(\Lambda_1)^n \) : (i) We assume that our parameters \( \mu_A \) such as \( \mu_\nu, \mu_{S\nu}, \mu_{eu} \) and \( \mu_R \) are given by some of those three scales \( \Lambda_1, \Lambda_2 \) and \( \Lambda_3 \). (ii) If our choice \( \mu_A = \Lambda_a \) \( (a = 1, 2, 3) \) as a result, gives a VEV value of \( \| A \| \) which is larger than the maximally allowed scale of \( \| A \| \), e.g. \( \Lambda_1 \) for \( \| (\hat{Y}_{eu})^\circ \| \), \( \Lambda_2 \) for \( \| (\Phi_\nu)^\bullet \| \), and \( \Lambda_3 \) for \( \| (\hat{Y}_R)^\circ \| \), then, the choice is ruled out. (iii) If the choice \( \mu_A = \Lambda_a \) gives a VEV value \( \| A \| \) which is two rank lower compared with \( \| A \|_{\text{max}} \), the choice is also ruled out.

For example, let us see the case (5.6):

\[ \mu_{eu} \| (\hat{Y}_{eu})^\bullet \| = \eta_\Phi (\Lambda_2)^2 = \eta_\Phi \rho \Lambda_1 \Lambda_3 = \eta_\Phi (\Lambda_2)^2. \]  
(5.17)

If we take \( \mu_{eu} = \Lambda_3 \), we obtain \( \| \hat{Y}_{eu} \| = \eta_\Phi \rho \Lambda_1 \) from Eq.(5.17). Since \( \| (\hat{Y}_{eu})^\bullet \| \leq \Lambda_1 \), the case \( \mu_{eu} = \Lambda_3 \) is not ruled out by the rule (ii). However, the case gives \( R_{2/1} = (\eta_\Phi)^{-2} \sim O(1) \) from (5.11). This contradicts with the parameter fitting result (5.15). Therefore, we rule out this case \( \mu_{eu} = \Lambda_3 \). On the other hand, if we take \( \mu_{eu} = \Lambda_1 \), we obtain \( \| (\hat{Y}_{eu})^\bullet \| = \rho \Lambda_3 \). However, the value is two rank small compared with the maximal value \( \Lambda_1 \), so that we rule out the case \( \mu_{eu} = \Lambda_1 \) by the rule (iii). As a result, we choose \( \mu_{eu} = \Lambda_2 \), and then we obtain

\[ \| (\hat{Y}_{eu})^\bullet \| = \eta_\Phi \Lambda_2. \]  
(5.18)

Then, from the relation (5.11), we obtain

\[ \xi_R \equiv R_{2/1} = \frac{1}{(\eta_\Phi)^2} \frac{\Lambda_2}{\Lambda_1}. \]  
(5.19)
Similarly, from the relation (2.18), we have

$$\mu_\nu \| (\Phi_\nu)_o \| = \eta \phi \frac{(\Lambda_2)^2}{\Lambda_1} \Lambda_2 = \eta \phi \rho \Lambda_3 \Lambda_2 = \rho \Lambda_3 \| (\Phi_e)_o \|.$$  \hfill (5.20)

If we take \( \mu_\nu = \Lambda_3 \), we obtain \( \| (\Phi_\nu)_o \| = \rho \| (\Phi_e)_o \| \), i.e. \( \xi_\nu = \rho > 1 \). The result is not our desired one, because our aim is to understand tiny neutrino masses by \( \xi_\nu \ll 1 \). Therefore, we choose \( \mu_\nu = \Lambda_2 \), and we get

$$\| (\Phi_\nu)_o \| = \frac{\Lambda_3}{\Lambda_2} \eta \phi \Lambda_2 = \frac{\Lambda_2}{\Lambda_1} \| (\hat{S}_e)_o \|,$$  \hfill (5.21)

so that we have

$$\xi_\nu = \frac{\Lambda_2}{\Lambda_1} = (\eta \phi)^2 \xi_R.$$  \hfill (5.22)

from (5.11).

Similarly, when we choose \( \mu_{S\nu} = \Lambda_2 \) in Eq.(2.23), from

$$\mu_{S\nu}(\hat{S}_\nu)_o = \eta_S \Lambda_1 \frac{(\Lambda_2)^2}{\Lambda_1} = \eta_S \rho \Lambda_1 \Lambda_3,$$  \hfill (5.23)

we obtain

$$\| (\hat{S}_\nu)_o \| = \eta_S \frac{\Lambda_2}{\Lambda_1} \Lambda_1 = \frac{\Lambda_2}{\Lambda_1} \| \hat{S}_e \|,$$  \hfill (5.24)

so that we have

$$\xi_{S\nu} = \frac{\Lambda_2}{\Lambda_1} = (\eta \phi)^2 \xi_R = \xi_\nu,$$  \hfill (5.25)

from Eq.(5.22). Therefore, from the relation (2.24), we obtain

$$\xi_M = (\xi_\nu)^2 (\xi_{S\nu})^{-1} = \xi_\nu = (\eta \phi)^2 \xi_R.$$  \hfill (5.26)

Finally, we discuss a scale of \( \mu_R \). From (5.7), by regarding \( Y_R \) as \( Y_R \simeq Y_R^{1st} \), we can write

$$\mu_R \| (\bar{Y}_R)^{oc} \| = \frac{1}{(\Lambda_1)^2} (\eta \phi)^2 \Lambda_2 \Lambda_1 \Lambda_2 \Lambda_2 = \rho (\eta \phi)^2 \Lambda_2 \Lambda_3.$$  \hfill (5.27)

If we take \( \mu_R = \Lambda_2 \), we obtain

$$\| (\bar{Y}_R)^{oc} \| = \rho (\eta \phi)^2 \Lambda_3.$$  \hfill (5.28)
Since $\|\bar{Y}^\circ \| \leq \Lambda_3$ cannot have a larger scale than $\Lambda_3$, we have a constraint
\[ \rho (\eta) \leq (\eta)^{-2} = \frac{3^{2/3}}{3}. \]  
\hspace{1cm} (5.29)
Comparing the fitting value $\rho = 1.37$ with the value $(\eta)^{-1} = 3^{1/3} = 1.44$ in (4.14), it is likely that the value of $\rho$ is given in unit of $3^{1/3}$, so that we regard the value of $\rho$ as
\[ \rho = \frac{1}{\eta} = 3^{1/3}. \]  
\hspace{1cm} (5.30)
Then, this choice satisfies the condition $\|Y_R\| \leq \Lambda_3$ in (5.28).

6 Estimate of flavon VEV scales

Similarly to Eq.(4.1), from the diagonalization of (5.12) with the parameter values (2.25), (2.26) and (5.15), we obtain
\[ (m_{\nu 1}, m_{\nu 2}, m_{\nu 3}) = (0.30892, 0.31689, 0.51026) m_{0\nu}. \]  
\hspace{1cm} (6.1)
Here, we have defined $m_{0\nu}$ by
\[ m_{0\nu} = (\xi_M)^2 \frac{m_{0e}^2}{m_{0R}}. \]  
\hspace{1cm} (6.2)
From the observed values $\Delta m_{22}^2 = 0.000244$ eV$^2$ and $\Delta m_{21}^2/\Delta m_{32}^2 = 0.0309$, we estimate $m_{\nu 3} = 0.063$ eV. Thus we obtain
\[ m_{0\nu} = 1.235 \times 10^{-13} \text{ TeV}. \]  
\hspace{1cm} (6.3)
Then, from Eq.(6.2) with (6.3) and (4.7), we obtain
\[ m_{0R} = (\xi_M)^2 \frac{m_{0e}^2}{m_{0\nu}} = (\xi_M)^2 \times 2.7710 \times 10^7 \text{ TeV}. \]  
\hspace{1cm} (6.4)
From Eq.(5.26) and (4.14), we estimate
\[ m_{0R} = (\eta \Phi)^4 (\xi_R)^2 \times 2.7710 \times 10^7 \text{ TeV} = 6.158 \text{ TeV}. \]  
\hspace{1cm} (6.5)
Therefore, from (5.28) with (5.30), i.e. $\|Y_R\| = \eta \Phi \Lambda_3$, we obtain
\[ \Lambda_3 = (\eta \Phi)^{-1} m_{0R} = 8.883 \text{ TeV}. \]  
\hspace{1cm} (6.6)
In conclusion, we obtain
\[ \Lambda_3 = 8.883 \text{ TeV}, \]
\[ \Lambda_2 = \rho \Lambda_3/\xi_R = 1.306 \times 10^4 \text{ TeV}, \]
\[ \Lambda_1 = \Lambda_2/\xi_R = 1.332 \times 10^7 \text{ TeV}. \]  
\hspace{1cm} (6.7)
Finally, we estimate the value of $\Lambda_H$ defined in (2.5). From (2.5), we use a relation

$$m_{0e} = \|\Psi_e(\tilde{S}_e)^{-1}\tilde{\Phi}_e\| = \frac{v_H^2(\eta_\Phi)^2(\eta_S)^{-1}(b_2^2)}{\Lambda_H} = \frac{v_H^2(\eta_\Phi)^2(\eta_S)^{-1} \rho \Lambda_3}{\Lambda_H}, \quad (6.8)$$

where $m_{0e} = 1.8499 \times 10^{-3}$ TeV, $v_H = 173.9 \times 10^{-3}$ TeV, and $(\eta_\Phi)^2(\eta_S)^{-1} \rho = 3^{-2/3}$, so that we obtain

$$\Lambda_H = 401.4 \text{ TeV}. \quad (6.9)$$

The result (6.9) gives $(v_H)/\Lambda_H = 0.4332 \times 10^{-3}$.

7 Concluding remarks

In conclusion, we have investigated a unified quark and lepton mass matrix model on the basis of $U(3) \times U(3)'$ family symmetry. We have inherited a basic aim of a series of the so-called Yukawaon models [3]. However, in the present model, we have not assumed existence of Yukawaons. Instead, we have introduced triplet fermions $F_\alpha$ of $U(3)'$ in addition to quarks and leptons $f_i$ which are triplets of $U(3)$, and we assumed a seesaw-like mechanism in (2.1).

The $U(3)'$ symmetry is broken into $S_3$ at a scale $\mu = \Lambda_1$, and the $U(3)$ symmetry is broken at $\mu = \Lambda_2$ by VEVs of flavons $\langle (\Phi_f)_i^a \rangle$, which are $(\mathbf{3}, \mathbf{3}^*)$ of $U(3) \times U(3)'$. However, in this paper, we do not ask the origin of the VEV form $\langle \Phi_f \rangle \propto \text{diag}(\sqrt{m_e}, \sqrt{m_\mu}, \sqrt{m_\tau})$, and do not discuss what mechanism leads to such a VEV form. We consider that it is too early to investigate the origin of the values $m_{ei}$. It is our future task.

The purpose of the present paper is not to give parameter fitting for quark and lepton masses and mixing because the VEV structures are the same as previous model, although the present model is completely different from the previous one. In the present model, we have only six free family-number-independent parameters for quark and lepton mass ratios and mixing as well as in the previous model. We have used the values $b_u = -1.011$, $b_d = -3.522 e^{17.7^\circ}$, $(\tilde{\phi}_1, \tilde{\phi}_2)$ and $\xi_R = 0.4903 \times 10^{-3}$, whose values have been quoted from the previous paper [1].

In this paper, we have investigated the symmetry-breaking scales $\Lambda_1$ and $\Lambda_2$ from phenomenological study in the neutrino data. Our essential hypothesis is that all flavon VEVs (and also $\mu$ parameters) take one of three scales $\Lambda_1$, $\Lambda_2$ and $\Lambda_3$ except for $H_{u/d}$ and $\Lambda_H$. First, we have fixed the ratio $\Lambda_2/\Lambda_1 \sim \Lambda_3/\Lambda_2 \sim 10^{-3}$ from the $\phi_i-m_{ei}$ relation with observed data as discussed in Sec.3. Here, we would like to emphasize that the relations (3.3) with the numerical fitting values (3.5) are essential for our result $\Lambda_2/\Lambda_1 \sim \Lambda_3/\Lambda_2$, and the relations (3.3) are completely different from that in the previous model [1][2]. Then, we have obtain the value of $\Lambda_1 \sim 4 \times 10^7$ TeV, $\Lambda_2 \sim 2 \times 10^4$ TeV, and $\Lambda_3 \sim 10^4$ TeV. Here, the value $\xi_R \sim 10^{-3}$ has been obtained for the parameter $\xi_R$ which is only the free parameter in the neutrino sector after remaining free parameters have been fixed by fitting the observed quark mass ratios and CKM mixing. The scale $\Lambda_3$ has been fixed from the scale of the Majorana mass matrix $\langle (\bar{Y}_R)^{\circ\circ} \rangle$. 

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Table 2: Predicted values vs. observed values in quark sector. Input values are given by Eq.(2.25) and (2.26). $r_{12}^q$ and $r_{23}^q$ are defined by $r_{12}^q = \sqrt{m_u/m_c}$, $r_{23}^q = \sqrt{m_c/m_t}$, $r_{12}^d = m_d/m_s$ and $r_{23}^d = m_s/m_b$. The predicted values are same as those in the previous paper [1].

|       | $|V_{us}|$ | $|V_{cb}|$ | $|V_{ub}|$ | $|V_{td}|$ | $\delta_{CP}^{\nu}$ | $r_{12}^q$ | $r_{23}^q$ | $r_{12}^d$ | $r_{23}^d$ |
|-------|-----------|-----------|-----------|-----------|----------------|-----------|-----------|-----------|-----------|
| Predicted | 0.2257    | 0.03996   | 0.00370   | 0.00917   | 81.0           | 0.061     | 0.060     | 0.049     | 0.027     |
| Observed | 0.22536   | 0.0414    | 0.00355   | 0.00886   | 69.4           | 0.045     | 0.060     | 0.053     | 0.019     |
| ±0.00061  | ±0.0012   | ±0.00015  | ±0.00033  | ±0.00032  | ±3.4           | ±0.013    | ±0.005    | ±0.005    | ±0.006    |

Note that our estimate of $\Lambda_1$, $\Lambda_2$ and $\Lambda_3$ is highly dependent on the VEV structure of the flavon $\tilde{Y}_R$, which corresponds to the Majorana mass matrix of right-handed neutrinos $\nu_R$. We have inherited the VEV structure from the previous model, which has given excellent description of neutrino masses and mixing with only one parameter $\xi_R$ phenomenologically. The structure is dependent on the $R$ charge assignment. The present $R$ charge assignment $(r_u, r_e, r_d, r_{\nu}) = (1/3, 1, 6/3, 10/3)$ is one of a few possible cases which do not cause theoretical trouble, although it has no theoretical ground. Investigation of the logical necessity of this assignment is a future task to us.

In Sec.4, we have discussed somewhat trifling parameterization with an additional ansatz. The orders of three VEV scales can roughly be obtained without these parameters in Sec.4. However, owing to this additional parametrization, we can understand $m_{0u}/m_{0e} = 3$ and $\rho \equiv c_2/c_1 \simeq 3^{1/3}$ consistently.

Although the present model still leaves some of tasks in future, we consider that the outline of the model is worthy to notice.

Acknowledgments

The authors thank T. Yamashita and Y. Sumino for helpful discussions on the effective superpotential and helpful comments. They also thank T. Asaka and K. Kaneta for valuable comments on a light right-handed Majorana neutrino mass and $\nu_L$-$\nu_R$ mixing. The work was supported by JSPS KAKENHI Grant number JP16K05325 (Y.K).

Appendix A: Parameter fitting for masses and mixings

Since the purpose of the present paper is not to give parameter fitting of masses and mixings of quarks and leptons, so far, we have not give the details of the parameter fittings. The present is drastically changed from the previous model [1] [2]. Nevertheless, the main part of the VEV relations are substantially the same as the previous those. Therefore, we have quoted the fitting values $(b_u, b_d)$ in Eq.(2.25), $(\tilde{\phi}_1, \tilde{\phi}_2)$ in Eq.(2.26) and $\xi_R$ in Eq.(5.15) from the previous paper [1] [2]. However, here, in order to help for reader’s understanding, we would like to give a brief review of the parameter fitting.
First, note that the parameter fitting in Ref.[1][2] has been done at $\mu = m_Z$. (For the charged lepton mass values, too, we have used the values at $\mu = m_Z$, not pole mass values.)

In our model, the parameters $(b_u, b_d)$ and $(\tilde{\phi}_1, \tilde{\phi}_2)$ are fixed by quark masses and CKM quark mixing fitting. We have chosen the parameter values as

$$b_u = -1.011, \quad b_d = -3.522 e^{i 17.7^\circ}, \quad (A.1)$$

and

$$(\tilde{\phi}_1, \tilde{\phi}_2) = (-176.05^\circ, -167.91^\circ). \quad (A.2)$$

Then, we can obtain reasonable quark mass ratios and CKM mixing as shown in Table.2.

Parameter fitting for neutrino data is done under the following dimensionless re-expression:

$$\tilde{M}_\nu^{Maj} = (\xi_M)^4 Z^2 \tilde{Y}_R^{-1} Z^2, \quad (A.3)$$

where

$$\tilde{Y}_R = Z^4 + \xi_R Z^2 P(1 + b_u X_3)^{-1} P^\dagger Z, \quad (A.4)$$

$$Z = \text{diag}(z_1, z_2, z_3), \quad P = \text{diag}(e^{i\phi_1}, e^{i\phi_2}, 1). \quad (A.5)$$

($Z$ and $P$ do not mean new flavons. Those are noting but dimensionless $3 \times 3$ matrices.)

Since the parameter values $b_u$ and $(\tilde{\phi}_1, \tilde{\phi}_2)$ have already determined from the CKM fitting as shown in (A.1) and (A.2), only a free parameter in the neutrino sector is $\xi_R$ in (A.4) as discussed in subsection 5.3. (The parameter fitting is practically the same as one in the previous model[7]. However, note that the present parameter $\xi_R$ corresponds to $2/\xi_R$ in the previous model.)

Therefore, as shown in the previous paper, the PMNS mixing parameters $\sin^2 2\theta_{12}$, $\sin^2 2\theta_{23}$, $\sin^2 2\theta_{13}$, $CP$ violating Dirac phase parameter $\delta^\ell_{CP}$, and the neutrino mass squared difference ratio $R_\nu \equiv \Delta m^2_{21}/\Delta m^2_{32}$ are turned out to be functions of the remaining only one parameter $\xi_R$. In Fig. 1, we draw the curves as functions of our new $\xi_R$ with taking $b_u = -1.011$, and $b_d = -3.3522$, $\beta_d = 17.7^\circ$, and $(\phi_1, \phi_2) = (-176.05^\circ, -167.91^\circ)$. From Fig.1, it turns out that the best fitting value of $\xi_R$ in the present paper is

$$\xi_R = 0.9806 \times 10^{-3}, \quad (\frac{2}{\xi_R} = 2039.6), \quad (A.6)$$

which gives the predictions

$$\sin^2 2\theta_{12} = 0.8254, \quad \sin^2 2\theta_{23} = 0.9967, \quad \sin^2 2\theta_{13} = 0.1007, \quad \delta^\ell_{CP} = -68.1^\circ, \quad R_\nu = 0.03118, \quad (A.7)$$
Figure 1: $\xi_R$ dependence of the lepton mixing parameters $\sin^2 2\theta_{12}$, $\sin^2 2\theta_{23}$, $\sin^2 2\theta_{13}$, and the neutrino mass squared difference ratio $R_\nu \equiv \Delta m_{31}^2 / \Delta m_{32}^2$. For the input values of $b_u$ and $(\phi_1, \phi_2)$, see the text.

Table 3: Predicted values vs. observed values in neutrino sector.

| Predicted | Observed | $\sin^2 2\theta_{12}$ | $\sin^2 2\theta_{23}$ | $\sin^2 2\theta_{13}$ | $R_\nu \times 10^{-2}$ | $\delta^\ell_{CP}(^\circ)$ | $m_{\nu_1}$ (eV) | $m_{\nu_2}$ (eV) | $m_{\nu_3}$ (eV) | $\langle m \rangle$ (eV) |
|-----------|----------|-----------------------|-----------------------|-----------------------|--------------------------|-----------------|-----------------|-----------------|-----------------|-----------------|
| 0.8254    | 0.846    | 0.9967                | 0.1007                | 3.118                 | +0.001                   | -68.1           | 0.038           | 0.039           | 0.063           | 0.021           |
| ±0.021    | ±0.008   | +0.001                | -0.018                | ±0.015                | ±0.15                    | < O(10^{-1})    | no data         | no data         | no data         | no data         |

as shown in Table 3. These predictions are in good agreement with the observed values [12].

The neutrino masses are predicted as $m_{\nu_1} \approx 0.038$ eV, $m_{\nu_2} \approx 0.039$ eV, and $m_{\nu_3} \approx 0.063$ eV, by using the input value $\Delta m_{32}^2 \approx 0.00244 \text{ eV}^2$. We have also predicted the effective Majorana neutrino mass $\langle m \rangle$ in the neutrinoless double beta decay as

$$\langle m \rangle = |m_{\nu_1}(U_{e1})^2 + m_{\nu_2}(U_{e2})^2 + m_{\nu_3}(U_{e3})^2| \approx 21 \text{ meV}. \quad (A.8)$$

It is interesting that the model predicts $\delta^\ell_{CP} = -68^\circ$, which shows $\delta^\ell_{CP} \simeq -\delta^\phi_{CP}$. It is also worthwhile noticing that we obtain a large value 21 meV for the effective Majorana neutrino mass $\langle m \rangle$ in spite of the normal hierarchy for the neutrino mass in our model.

References

[1] Y. Koide and H. Nishiura, Phys. Rev. D 92, 111301(R) (2015).
[2] Y. Koide and H. Nishiura, Mod. Phys. Lett. A 31, 1650125 (2016).

[3] Y. Koide, Phys. Rev. D 79, 033009 (2009); Phys. Lett. B 680, 76 (2009); H. Nishiura and Y. Koide, Phys. Rev. D 83, 035010 (2011); Y. Koide and H. Nishiura, Euro. Phys. J. C 72, 1933 (2012); Y. Koide, J. Phys. G 38, 085004 (2011); Y. Koide and H. Nishiura, Euro. Phys. J. C 73, 2277 (2013); JHEP 04, 166 (2013); Phys. Rev. D 88, 116004 (2013); Phys. Rev. D 90, 016009 (2014).

[4] C. D. Froggatt and H. B. Nelsen, Nucl. Phys. B 147, 277 (1979). For recent works, for instance, see R. N. Mohapatra, AIP Conf. Proc. 1467, 7 (2012); A. J. Buras et al., JHEP 1203, 088 (2012).

[5] P. Minkowski, Phys. Lett. B 67, 421 (1977); M. Gell-Mann, P. Ramond, and R. Slansky, Proceedings of the Supergravity Stony Brook Workshop, New York, 1979, edited by P. Van Nieuwenhuizen and D. Freedman (North-Holland, Amsterdam, 1979); T. Yanagida, Proceedings of the Workshop on Unified Theories and Baryon Number in the Universe, Tsukuba, Japan 1979, edited by A. Sawada and A. Sugamoto [KEK Report No. 79-18, Tsukuba]; R. Mohapatra and G. Senjanovic, Phys. Rev. Lett. 44, 912 (1980); J. Schechter, J.W.F. Valle, Phys. Rev. D 22, 2227 (1980).

[6] Y. Koide and H. Fusaoka, Z. Phys. C 71, 459 (1996).

[7] Y. Koide and H. Nishiura, Phys. Rev. D 91, 116002 (2015).

[8] N. Cabibbo, Phys. Rev. Lett. 10, 531 (1963); M. Kobayashi and T. Maskawa, Prog. Theor. Phys. 49, 652 (1973).

[9] Z.-z. Xing, H. Zhang, and S. Zhou, Phys. Rev. D 77, 113016 (2008). Also see H. Fusaoka and Y. Koide, Phys. Rev. D 57, 3986 (1998).

[10] B. Pontecorvo, Zh. Eksp. Teor. Fiz. 33, 549 (1957); 34, 247 (1957); Z. Maki, M. Nakagawa, and S. Sakata, Prog. Theor. Phys. 28, 870 (1962).

[11] Y. Koide, J. Phys. G 35, 125004 (2008); Phys. Lett. B 665, 227 (2008).

[12] K. A. Olive et al. (Particle Data Group), Chinese Phys. C, 38, 090001 (2014).

[13] M. Doi, T. Kotani, H. Nishiura, K. Okuda, and E. Takasugi, Phys. Lett. B 103, 219 (1981) and B 113, 513 (1982).