Power–law tail distributions and nonergodicity

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We establish an explicit correspondence between ergodicity breaking in a system described by power–law tail distributions and the divergence of the moments of these distributions.

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Statistical mechanics is a combination of the law of large numbers and the laws of mechanics. Since its foundation in the late 19th century, this theory has been extremely successful in describing equilibrium and nonequilibrium properties of a very large number of macroscopic systems [1]. In the last decade, however, a new class of systems that do not obey the law of large numbers has emerged [2,3,4]. The behavior of these systems is dominated by large and rare fluctuations that are characterized by broad distributions with power–law tails. The hallmark of these statistical distributions, commonly referred to as Lévy statistics [5], is the divergence of their first and/or second moment.

The question we address in this paper is how ergodicity is affected in systems described by power–law tail distributions with diverging moments. Ergodicity is a central concept in statistical physics and is usually stated by saying that ensemble average and time average of observables are equal in the infinite–time limit [6,7]. The ergodic hypothesis has recently been investigated experimentally in two different systems governed by Lévy statistics in time: fluorescence intermittency of nanocrystal quantum dots [8] and subrecoil laser cooling of atoms [9]. Both experiments have found that Lévy statistics induces ergodicity breaking. A precise understanding of the connection between divergent moments and nonergodicity is thus of high interest. A common feature of the systems in the above experiments is their nonstationarity. This in contrast to the system we propose to study, namely atomic transport in an optical lattice, where a steady state does exist. Nonetheless, we will show that this system can exhibit nonergodic behavior.

An optical lattice is a standing wave light field formed at the intersection of two or more laser beams [10]. Due to the spatial periodicity of the potential, an optical lattice is similar in many respects to a solid state crystal. The main advantage of an optical lattice, however, is its high tunability: both the period and the amplitude of the optical potential can be modified in a controlled way. This has the interesting consequence that the exponents of the power–law distributions appearing in this system are not fixed, as in most systems, but can be varied continuously, allowing the exploration of different regimes.

The motion of atoms in a one-dimensional optical lattice formed by two counterpropagating laser beams with linear perpendicular polarization can be described, after spatial averaging, by a Fokker–Planck equation for the semiclassical Wigner function $W(p,t)$ [11,12,13],

$$\frac{\partial W}{\partial t} = -\frac{\partial}{\partial p} [K(p)W] + \frac{\partial}{\partial p} \left[ D(p) \frac{\partial W}{\partial p} \right].$$

(1)

The momentum–dependent drift and diffusion coefficients are respectively given by,

$$K(p) = -\frac{\alpha p}{1 + (p/p_c)^2}, \quad D(p) = D_0 + \frac{D_1}{1 + (p/p_c)^2}.$$

(2)

The drift $K(p)$ corresponds to a cooling force with friction coefficient $\alpha$, while the diffusion coefficient $D(p)$ represents stochastic momentum fluctuations and describes heating processes. It is worth noticing that for large momentum, the range of the drift is limited by the capture momentum $p_c$, while the range of the fluctuations is not. The kinetic equation (1) is valid in a regime where (i) the atomic momentum is large, $p \gg \hbar k$, where $k$ is the wave number of the laser field (this defines the semiclassical limit), (ii) the saturation parameter is small, $s \ll 1$ (this corresponds to low laser intensity) and (iii) the kinetic energy of the atoms is large, $p^2/2m \gg U_0$, where $U_0$ is the depth of the optical potential (this last condition allows spatial averaging). The stationary solution of the atomic transport equation (1) which satisfies natural boundary conditions, $W_s(p \to \pm \infty) \to 0$, is of the form,

$$W_s(p) = \frac{1}{Z} \left[ 1 + \frac{D_0}{D_0 + D_1} \frac{p^2}{\alpha^2} \right]^{-\alpha^2/(2D_0)},$$

(3)

where $Z$ is a normalization constant. The momentum distribution (3) has an asymptotic power–law tail, $W_s(p) \sim |p|^{-(\alpha^2)/D_0}$, with an exponent that can be expressed in terms of the potential depth $U_0$ and the recoil energy $E_R$ as $(\alpha^2)/D_0 = U_0/(22E_R)$ [11]. The statistical properties of the distribution (3) can therefore be easily changed from normal statistics for $U_0 \geq 66E_R$, to Lévy statistics for $U_0 < 66E_R$, by simply modifying the depth of the optical potential. In particular, one should note that the second moment, $\langle p^2 \rangle = \int dp p^2 W_s(p)$, becomes infinite when $U_0 < 66E_R$. In this regime the mean kinetic energy of the system, $E_K = \langle p^2 \rangle/2m$, diverges, clearly signaling unusual thermodynamic behavior.

Transport in an optical lattice has a number of attractive features that make it an ideal case study of the thermodynamical properties of systems described by power–law distributions. On the one hand, the atomic transport equation (1) has been derived from the microscopic
Hamiltonian that describes the interaction with the laser fields and the quantities that appear in this equation can be expressed in terms of the microscopic parameters of the quantum–optical problem \[11\]. Moreover, the regime defined by conditions (i) to (iii) has been implemented experimentally and the divergence of the kinetic energy below a given potential threshold has been observed \[14\]. Finally, Eq. (1) is an ordinary linear Fokker–Planck equation, meaning that standard methods of stochastic calculus can be used to analyse the problem. This is in contrast to most systems with power–law distributions that are often described by nonlinear or fractional kinetic equations \[17\]. In particular, for the case of vanishing \(D_1\) (which we shall consider in the sequel \[16\]), Eq. (1) corresponds to a random process driven by additive Gaussian white noise. The fact that power–law fluctuations with infinite variance occur here in a system subjected solely to additive noise is worth emphasizing. The physical mechanism which gives rise to divergent fluctuations in systems with multiplicative noise, where the noise strength is proportional to the stochastic variable, is well–known \[17\]: it is based on a positive feedback mechanism is at work, based on the steady decrease to zero of the friction force as the value of the momentum increases. The appearance of infinite momentum fluctuations in this system is thus a striking illustration of the complex behavior that can result from the subtle interplay of the noise and the nonlinearity of the drift. In the following, we establish a correspondence between the divergent moments of the power–law distributions of the system and nonergodic behavior.

We begin by transforming the Fokker–Planck equation (1) into a Schrödinger–like equation by writing \(W(p,t) = W_s(p)/(1 + \psi(p,t))\) \[18\]. The function \(\psi(p,t)\) satisfies the imaginary–time Schrödinger equation,

\[
-\frac{\partial \psi}{\partial t} = -D_0 \frac{\partial^2 \psi}{\partial p^2} + V(p)\psi = H\psi ,
\]

with the potential \(V(p) = K'(p)/2 + K(p)^2/(4D_0)\). For the drift coefficient (2), this potential reads

\[
V(p) = \frac{\alpha p_c^2}{4D_0} \frac{p^2 (\alpha p_c^2 + 2D_0) - 2D_0 p_c^2}{(p^2 + p_c^2)^2} .
\]

This transformation reveals the fundamental difference between the case of finite \(p_c\) and the case of infinite \(p_c\), where Eq. (1) reduces to the familiar Ornstein–Uhlenbeck process with linear drift. For infinite \(p_c\), the Schrödinger potential \(V(p)\) asymptotically increases as \(p^2\), whereas for finite \(p_c\), it asymptotically decreases as \(1/p^2\). By contrast, the Fokker–Planck potential, \(\Phi(p) = -\int_0^p dp' K'(p')\), is confining for any value of the capture momentum. As a consequence, the spectrum of the Hamiltonian, \(H\psi_k(p) = E_k\psi_k(p)\), is discrete in the former case, while it is continuous, except for the discrete ground state \(\psi_0(p)\), in the latter. In both cases, the stationary momentum distribution is given by the square of the ground state eigenfunction, \(W_s(p) = \psi_0(p)^2\). Interestingly, we note that a real–time Schrödinger equation with a potential of the form (3) has recently been considered in Ref. \[19\]. To simplify the discussion, we now adopt rescaled variables for which \(\alpha = p_c = 1\) and \(D_0 = D\); the noise intensity \(D\) being then the only remaining parameter in the problem. It will also be convenient to divide momentum space into a low–momentum region \(|p| < 1\), where the drift is approximately linear, \(K_1(p) \approx -p\), and a high–momentum region \(|p| > 1\), with a drift \(K_2(p) \approx -1/p\). The anomalous dynamics of the system is completely determined by the high–momentum region. The eigenvalues and eigenfunctions of the Hamiltonian \(H\) in this region are given by \(E_k = Dk^2\) and \(\psi_k(p) = \sqrt{\gamma}[c_1 J_\nu(kp) + c_2 Y_\nu(kp)]\), where \(J_\nu(p)\) and \(Y_\nu(p)\) are the Bessel functions of the first and second kind of order \(\nu = (D + 1)/2D\). The constants \(c_{1,2}\) are fixed by the boundary conditions.

When discussing the ergodicity of a system, one is typically not interested in the trajectory in the full space–time, but often in the projection of the trajectory onto some subspace of relevant variables \[1\], in the present case momentum. A criterion for the equality of ensemble average and time average of a dynamical quantity \(A\) is then provided by the condition \[2\]

\[
\sigma^2_A(t) = \left( \overline{A} - \langle A \rangle \right)^2 \rightarrow 0 \text{ when } t \rightarrow \infty .
\]

Here \(\overline{A} = t^{-1} \int_0^t \! d\tau \, A(p(\tau))\) is the time average of the observable \(A\) and \(\langle A \rangle = \int dp \, A(p)W(p,t)\) denotes its ensemble average. In the infinite–time limit, the latter tends to the stationary ensemble average \(\langle A \rangle_s = \int dp \, A(p)W_s(p)\). A system that obeys \[4\] is said to be ergodic in the mean square sense. In order to determine the ergodicity of our system, we thus need to compute the long–time behavior of the covariance,

\[
\sigma^2_A(t) = \frac{1}{t^2} \int_0^t \! dt_1 \int_0^t \! dt_2 \left[ \langle A(p(t_1))A(p(t_2)) \rangle - \langle A(p(t_1)) \rangle \langle A(p(t_2)) \rangle \right] .
\]

This can be done by applying the usual theory of stochastic processes \[18\]. The two–time correlation function \(\langle A(p(t_1))A(p(t_2)) \rangle\) is defined by the integral

\[
\langle A(p(t_1))A(p(t_2)) \rangle = \int \! dp_1 dp_2 \, A(p_1)A(p_2) \, W_2(p_1, t_1; p_2, t_2) .
\]

where \(W_2(p_1, t_1; p_2, t_2)\) is the two–point joint probability density function. Since the Fokker–Planck equation (1) describes a stationary Markov process (for any value of \(p_c\)), the probability distribution \(W_2(p_1, t_1; p_2, t_2)\) depends only on the time difference \(|t_1 - t_2|\) and can be expressed in terms of the eigenvalues and eigenfunctions.
of the Schrödinger equation \( \psi_0(p_1)\psi_0(p_2) \) in the form,
\[
W_2(p_1, t_1; p_2, t_2) = \psi_0(p_1)\psi_0(p_2) \times
\left( \psi_0(p_1)\psi_0(p_2) + \int_0^\infty dk \, \psi_k(p_1)\psi_k(p_2) e^{-E_k|t_1-t_2|} \right) .
\]
Combining Eqs. (7), (8) and (9) and introducing the variable \( \tau = t_2 - t_1 \), we arrive at
\[
\sigma_A^2(t) = \frac{2}{\tau^2} \int_0^t d\tau \, (t - \tau) \, C_A(\tau),
\]
where the function \( C_A(\tau) \) is given by
\[
C_A(\tau) = \int_0^\infty dk \left[ \int dp \, A(p) \psi_0(p)\psi_k(p) \right]^2 e^{-E_k \tau} .
\]
The infinite–time limit of the covariance \( \sigma_A^2(t) \) is entirely determined by the asymptotic behavior of \( C_A(\tau) \). This function depends explicitly on the observable \( A(p) \). It is worthwhile to notice that ergodicity will therefore in general depend on the dynamical variable \( A(p) \) under consideration. In the following, we take \( A(p) = p^n \) and evaluate the long-time behavior of \( C_A(\tau) \) following Ref. [13]. We find \( C_A(\tau) \sim \tau^{-\mu} \) with an exponent \( \mu = (1 - (2n + 1)/D)/2D \). As a result, the covariance \( C_A(\tau) \) will converge to zero as \( t \to \infty \), only if \( D < D_n = 1/(2n + 1) \). We thus obtain the important result that there is a noise threshold \( D_n \) above which ergodicity is broken. As already mentioned, this threshold depends explicitly on the parameter \( n \), that is, on the quantity \( A(p) \): the smaller the power \( n \), the larger the value of \( D_n \). On the other hand, the moments of the stationary momentum distribution, \( (p^m) = \int dp \, p^m W_s(p) \), are finite for \( D < D'_n = 1/(m+1) \). We can therefore conclude that there exists a direct relationship between the loss of ergodicity in the system for the variable \( A(p) = p^n \) and the divergence of the 2nth moment of the stationary momentum distribution.

Let us look in more detail at the value \( n = 0 \), which corresponds to the largest noise threshold \( D_{n=0} = 1 \). We first mention that for \( D > D_{n=0} \), the Fokker–Planck equation doesn’t have a normalizable stationary solution anymore and the system is obviously nonergodic. Further, for \( n = 0 \), the function \( C_A(\tau) \) can be rewritten in terms of the conditional probability density \( P_2(p_2, \tau|p_1, 0) \) as
\[
C_A(\tau) = \int dp_1 \, W_s(p_1) \times
\int dp_2 \left( P_2(p_2, \tau|p_1, 0) - W_s(p_2) \right) .
\]
This result is of special interest. Equation (12) shows that, unlike for the Ornstein–Uhlenbeck process, the steady state is here reached in a nonexponential way. Such an asymptotic power–law decay is usually associated with non–Markovian processes and is rather surprising for a stationary Markov process as described by Eq. (1). The origin of this algebraic behavior is of course rooted in the nonlinear drift coefficient (2).

FIG. 1: Probability to be in the low–momentum region (continuous line) and in the high–momentum region (dashed line), in the long–time limit, as a function of the noise strength \( D \).

A closely related quantity is the first–passage time distribution. The first–passage time is defined as the time at which the momentum of the system first exits a certain momentum interval, given that it was originally in that interval. The first–passage time problem for the Fokker–Planck equation (1) can again be treated using standard techniques of stochastic calculus (see for example Ref. [21]). In the high–momentum region, the Laplace transform \( g_2(s) \) of the first–passage time distribution obeys the following backward equation,
\[
D \frac{\partial^2 g_2}{\partial p^2} - \frac{1}{p} \frac{\partial g_2}{\partial p} - sg_2 = 0 .
\]
Solving Eq. (13) with the boundary conditions \( g(1) = 1 \) and \( g(\infty) = 0 \), we obtain
\[
g_2(s) = \frac{K_v(\sqrt{s/D})}{K_v(\sqrt{s/D})} p^\nu ,
\]
where \( K_v(p) \) is the modified Bessel function of the second kind of order \( \nu \). It follows from Eq. (14), that to leading order, \( g_2(s) \sim s^\nu \) as \( s \to 0 \). The first–passage time distribution in region 2 is thus also a power–law tail distribution and it asymptotically behaves as \( g_2(t) \sim t^{-\gamma} \) with an exponent \( \gamma = (3D + 1)/(2D) \). The corresponding moments \( (t^n) = \int dt \, t^n g_2(t) \) converge if \( D < D'_n = 1/(2n + 1) \). The first–passage time distribution \( g_1(t) \) in region 1 can be computed along similar lines and the associated moments can be shown to be all finite. Figure (1) shows the probability to be in region 1 and 2, in the limit of long times, as a function of the noise intensity \( D \). We observe that for small \( D \), atoms are mostly located in the low–momentum region, where they experience the linear part of the drift. On the other hand, for \( D \) close to \( D_{n=0} \), atoms get localized in the high–momentum region, where the drift asymptotically decays as \( -1/p \). Remarkably, atoms in this system can thus be brought in a high–energy state through the sole action of the noise. We note that this problem exhibits an interesting analogy with subrecoil laser cooling, where atoms accumulate in a low–energy state (in a so–called dark state), with infinite mean trapping time (22). To our knowledge, loss of ergodicity in systems with divergent trapping times has been first discussed in the context of spin–glasses (23).
We can now formulate the main result of the paper: Ensemble average and time average of the dynamical variable $A(p) = p^n$ stop being equal in the infinite–time limit — ergodicity of the system is accordingly broken— when the $2n$th moment of the stationary momentum distribution and the $(n + 1)$th moment of the first passage time distribution in the high–momentum region become infinite. An unambiguous correspondence between the nonergodic properties of a system described power–law distributions and the $(n + 1)$th moment of the first passage time distribution in the high–momentum region become infinite. The method of moments in the continuous description, the moments of the power–law tail distributions describe the ordinary Fokker–Planck equation. Further insight can be gained by considering a discretized form of the Fokker–Planck equation. In doing so, we shall obtain a generalization of the Ehrenfest urn model, which has played an important role in clarifying the behavior of the system, both in momentum space and in time.

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