Variational Contact Symmetries of Constraint Lagrangians

Petros A. Terzis\textsuperscript{a,}\textsuperscript{*}, N. Dimakis\textsuperscript{b}\textsuperscript{†}, T. Christodoulakis\textsuperscript{a}\textsuperscript{‡}, Andronikos Paliathanasis\textsuperscript{c,d}\textsuperscript{§}, Michael Tsamparlis\textsuperscript{e}\textsuperscript{¶}

\textsuperscript{a}Nuclear and Particle Physics Section, Physics Department, University of Athens, GR 157–71 Athens
\textsuperscript{b}Instituto de Ciencias Fisicas y Matematicas, Universidad Austral de Chile, Valdivia, Chile
\textsuperscript{c}Dipartimento di Fisica, Università di Napoli “Federico II” Complesso universitario Monte S. Angelo, Via Cintia 9, I- 80126 Napoli, Italy
\textsuperscript{d}Istituto Nazionale di Fisica Nucleare (INFN) sezione di Napoli Complesso universitario Monte S. Angelo, Via Cintia 9, I- 80126 Napoli, Italy
\textsuperscript{e}Department of Physics, Section of Astronomy, Astrophysics and Mechanics, University of Athens, Panepistemiopolis, Athens 157 83, Greece

Abstract

The investigation of contact symmetries of re-parametrization invariant Lagrangians of finite degrees of freedom and quadratic in the velocities is presented. The main concern of the paper is those symmetry generators which depend linearly in the velocities. A natural extension of the symmetry generator along the lapse function $N(t)$, with the appropriate extension of the dependence in $\dot{N}(t)$ of the gauge function, is assumed; this action yields new results. The central finding is that the integrals of motion are either linear or quadratic in velocities and are generated, respectively by the conformal Killing vector fields and the conformal Killing tensors of the configuration space metric deduced from the kinetic part of the Lagrangian (with appropriate conformal factors). The freedom of re-parametrization allows one to appropriately scale $N(t)$, so that the potential becomes constant; in this case the integrals of motion can be constructed from the Killing fields and Killing tensors of the scaled metric. A rather interesting result is the non-necessity of the gauge function in Noether’s theorem due to the presence of the Hamiltonian constraint.

\textsuperscript{*}pterzis@phys.uoa.gr
\textsuperscript{†}nsdimakis@gmail.com
\textsuperscript{‡}tchris@phys.uoa.gr
\textsuperscript{§}paliathanasis@na.infn.it
\textsuperscript{¶}mtsampa@phys.uoa.gr
1 Introduction

The Lie symmetry method is a systematic tool for the study of systems of differential equations. The importance of Lie symmetries lies in the fact that they can be used in order to reduce the order of an ordinary differential equation or to assist in the solution of differential systems by reducing the number of independent variables. The Lie symmetries which leave invariant a variational integral are called Noether symmetries. Noether symmetries form a subalgebra of the Lie symmetries of the equations following from the variational integral. The characteristic of Noether symmetries is that to each symmetry there corresponds a divergence free current along with the corresponding charge. The well known conservation laws of energy, angular momentum all follow from Noether symmetries. It is well known that conservation laws are important tools which can be used for the determination of integrable manifolds of a dynamical system in mathematical physics, biology, economics and many others [1, 2, 3, 4, 5, 6, 7, 8, 9].

The determination of the Lie and Noether symmetries of a given differential equation consists of two steps, (a) The derivation of the symmetry conditions and (b) the solution of these conditions. The first step is formal and it is outlined in e.g. [10, 11, 12]. The second step can be a difficult task since the symmetry conditions can be quite involved. Tsamparlis & Paliathanasis in [13, 14] proposed a geometric method for the solution of the symmetry conditions for regular Hamiltonian systems which describe the motion of a particle in a Riemannian space under the action of a potential. It was shown that the Lie and Noether symmetries of that system are related with the special projective algebra of the underlying space. This geometric approach has been extended to the determination of the Lie and Noether point symmetries of some families of partial differential equations [15, 16].

A similar analysis has been established by Christodoulakis, Dimakis & Terzis [17] for constraint Lagrangians quadratic in the velocities, where it was shown that the point symmetries of equations of motion are exactly the variational symmetries (containing the time reparametrization symmetry) plus the scaling symmetry. The variational symmetries was seen to be the simultaneous conformal Killing vector fields of both the metric, defined by the kinetic term, and the potential. These results have been applied in various areas, i.e. in classical mechanics, in general relativity and in cosmology; in each case new dynamical systems which admit symmetries were found for instance see [18, 19, 20, 14, 21, 22, 23, 24, 25] and reference therein.

In this work we extend this geometric approach in order to study the Lie Bäcklund variational symmetries (or dynamical Noether symmetries) of constrained Lagrangian systems. Specifically, we study the symmetries which arise from contact transformations, the contact Noether symmetries. Contrary to the point transformations which are defined in the configuration space manifold, where the dynamical equations are defined, the contact transformations are defined in the tangent bundle thus they depend also in the velocities [26, 27]. Such contact Noether symmetries of regular Lagrangian systems with a potential have been studied in [28] and it was shown that the corresponding conservation laws follow from the Killing tensors of the kinetic metric and the potential. Some important
conservation laws of that kind in physics are: the Runge-Lenz vector field of the Kepler problem, the Ray-Reid invariant, and the Carter constant in Kerr spacetime \cite{29,30,31,32}. As it will be shown, in the case of constrained Lagrangian systems the results are different from that of \cite{28} in that the Noether contact symmetries are related to the Conformal Killing tensors of the underlying space. The method we develop can be used and for regular Lagrangian systems. Furthermore it is interesting to note that, in the presence of constrains, the gauge function of Noether’s theorem does not play a significant role: the presence of the quadratic constraint renders it non–essential in finding the conservation laws. The plan of the paper is as follows.

In section 2 we give the basic theory of the Lie Bäcklund symmetries. In section 3 we study the linear contact symmetries of constraint Lagrangian systems and we prove the main theorem of this work. The conservation laws which follow from the contact symmetries are studied in 4. Furthermore, in section 5 we demonstrate the main results in two applications, (a) the determination of the variational contact symmetries for the geodesic equations of a Biv pp-wave spacetime and (b) the derivation of the Ray-Reid invariant of corresponding constraint Hamiltonian system. Finally, in section 6 we discuss our results.

2 Preliminaries

For the convenience of the reader, in this section we present the basic properties and definitions concerning the generalized symmetries.

Let \( H = H \left( x^i, u^A, u^A_i, u^A_{ij}, \ldots \right) \) be a function which is defined in the space \( A_M = \{ x^i, u^A, u^A_i, u^A_{ij}, \ldots \} \) where \( x^i \) are \( n \) independent variables and \( u^A \) are \( m \) dependent variables and \( u^A_i = \frac{\partial u^A}{\partial x^i} \). The function \( H \) describes a set of differential equations, for \( n = 1 \) of ordinary differential equations and for \( n > 1 \) a set of partial differential equations.

We shall say that the function \( H \) will be invariant under the action of the following infinitesimal transformation

\[
\begin{align*}
\bar{x}^i &= x^i + \epsilon \xi^i (x^i, u^B, u^B_i, u^B_{ij}, \ldots) \\
\bar{u}^A &= u^A + \epsilon \eta^A (x^i, u^B, u^B_i, u^B_{ij}, \ldots)
\end{align*}
\]  

if there exist a function \( \lambda (x^i, u, u_i, u_{ij}, \ldots) \) such as the following condition holds \cite{10}

\[
[X, H] = \lambda H, \quad \text{mod} H = 0.
\] (2.3)

where \( X = \frac{\partial}{\partial x^i} \partial_i + \frac{\partial u^A}{\partial x^i} \partial_A \) is the generator of the infinitesimal transformation \eqref{2.1}, \eqref{2.2}, i.e.

\[
X = \xi^i (x^i, u^B, u^B_i, u^B_{ij}, \ldots) \partial_i + \eta^A (x^i, u^B, u^B_i, u^B_{ij}, \ldots) \partial_A
\] (2.4)

and \([\text{ }, \text{ }]\) is the Lie Bracket. When condition \( \text{[2.3]} \) holds, the vector field \( X \) is called Lie Bäcklund symmetry of the function \( H \); therefore we conclude that a Lie Bäcklund transformation is a transformation in \( A_M \) which preserves the set of
solutions \( u^A \) of \( H \left( x^i, u, u_i, u_{ij}, \ldots \right) \) in \( \mathcal{A}_M \). An alternative way to write condition (2.3) is

\[
X^{[n]} H = \lambda H, \mod H = 0
\]

where \( X^{[n]} \) is the \( n \)th prolongation of \( X \) in the space \( \mathcal{A}_M \).

Every differential equation admits as Lie Bäcklund symmetry the vector field

\[
D_i = \partial_i + u_i \partial_u + u_{ij} \partial_{u_i} + \ldots
\]

which is the total derivative operator \([11]\). Let \( X \) be a Lie Bäcklund symmetry then the vector field

\[
\bar{X} = X - f^i D_i = \left( \xi^k - f^k \right) \partial_k + \left( \eta^A - f^k u_k^A \right) \partial_{u^A} + \ldots
\]

is also a Lie Bäcklund symmetry for arbitrary functions \( f^i \). Without loss of generality we could select \( f^i = \xi^i \) and obtain

\[
\bar{X} = \left( \eta^A - \xi^k u_k^A \right) \partial_{u^A}.
\]

The generator (2.7) is called the canonical form of (2.4). Since we can always absorb the term \( \xi^k u_k \) inside the \( \eta \) we conclude that \( \bar{X} = Z^A \left( x^i, u^B, u_i^B, u_{ij}^B, \ldots \right) \partial_{u^A} \) is the generator of a Lie Bäcklund symmetry. This form of the generator is also suitable for the variational symmetries of the action e.g. see [12] p. 331–333.

A special class of Lie Bäcklund symmetries are the contact symmetries defined by the requirement that the generator depends only on the first derivatives \( u_i \), i.e. it has the general canonical form

\[
X_C = Z^A \left( x^i, u^B, u_i^B \right) \partial_{u^A}.
\]

### 2.1 Noether’s Theorem for contact transformations

Consider now that \( H = H \left( t, q^A, \dot{q}^A, \ddot{q}^A \right) \) where \( \dot{q}^A = \frac{d q^A}{dt} \), and \( H \) follows from a variational principle, that is there exist a Lagrange function \( L \left( t, q^A, \dot{q}^A \right) \) such as \( E_L \left( L \right) = 0 \) where \( E_L \) is the Euler-Lagrange vector. Let \( \bar{X} = Z^A \left( t, q^B, \dot{q}^B \right) \partial_A \) be the generator of a contact transformation. Then, if there exist a function \( F = F \left( t, q^B, \dot{q}^B \right) \) such as [26]

\[
\bar{X} \left[ 1 \right] L = \frac{dF}{dt}
\]

the vector field \( X \) is a contact Noether symmetry of the Lagrangian \( L \left( t, q^A, \dot{q}^A \right) \) [33]. Where \( \bar{X} \left[ 1 \right] \) is the first prolongation of \( \bar{X} \), i.e. \( \bar{X} \left[ 1 \right] = X + \dot{Z}^i \partial_{q^i} \). Furthermore, we have that for every contact Noether symmetry there exist a function \( I = I \left( t, q^B, \dot{q}^B \right) \) [33, 34]

\[
I \left( t, q^B, \dot{q}^B \right) = Z^A \left( t, q^B, \dot{q}^B \right) \frac{\partial L}{\partial \dot{q}^A} - F
\]

such as \( \frac{dI}{dt} = 0 \), i.e. the function \( I \left( t, q^B, \dot{q}^B \right) \) is a first integral of the Lagrange equations and called a contact Noether Integral.

In the following section we study the contact symmetries of singular constrained Lagrangian systems.
3 Contact transformations linear in velocities

We are seeking variational symmetries of the action

\[ A = \int dt L = \int dt \left( \frac{1}{2N} G_{\mu\nu}(q) \dot{q}^\mu \dot{q}^\nu - N V(q) \right), \quad \mu, \nu = 1, \ldots, d \]

\[ \text{det} G_{\mu\nu} \neq 0, \]  

solved for all \( d + 1 \) dependent variables, but only for \( d \) of them; the remaining variable being freely specifiable. If this variable is indeed selected to be a specific function of time, one can integrate (in principle) \( 3.2a \) to obtain the remaining variables, in which case equation \( 3.2a \) reduces to a relation between constants.

As shown in [17], in order to reveal all the symmetries that exist by virtue of the constraint equation \( 3.2a \), \( N \) must be considered on an equal footing with the configuration space coordinates \( \dot{q}^\mu \). Thus, it is imperative to avoid any unnecessary gauge fixing, prior to the derivation of the symmetries. In this regard, the canonical form of generator \( X \) of the symmetry transformation under consideration must be of the form

\[ X = (\xi^\kappa(t,q,N) + K_\alpha^\kappa(t,q,N)\dot{q}^\alpha) \frac{\partial}{\partial \dot{q}^\kappa} + \Omega(t,q,N,\dot{q},\dot{N}) \frac{\partial}{\partial \dot{N}}, \]  

incorporating \( N \) as one of the variables. The infinitesimal criterion that the symmetry generator must satisfy is condition \( 2.9 \).

The first prolongation of the generator is

\[ X^{[1]} = X + \phi^\kappa \frac{\partial}{\partial \dot{q}^\kappa} \]

with

\[ \phi^\kappa = \frac{d\xi^\kappa}{dt} + \frac{d}{dt} (K_\alpha^\kappa \dot{q}^\alpha) \]

\[ = \xi^\kappa_t + \xi^\kappa_\alpha \dot{q}^\alpha + \xi^\kappa_0 \dot{N} + K_\alpha^\kappa_\beta \dot{q}^\alpha \dot{q}^\beta + K_\alpha^\kappa_0 \dot{N} \dot{q}^\alpha - K_\alpha^\kappa \Gamma^\alpha_{\mu\nu} \dot{q}^\mu \dot{q}^\nu + \frac{1}{N} K_\alpha^\kappa \dot{N} \dot{q}^\alpha - N^2 K_\kappa^\nu V_{,\mu} \]
where \( \dot{q}^a \) and equations (3.2b) have been used, so as to substitute the accelerations \( \ddot{q}^a \) with respect to the velocities \( \dot{q}^a \), \( \dot{N} \). Note that (3.1) needs not a prolongation term in \( \frac{\partial}{\partial q^a} \), since the Lagrangian is singular and thus does not involve the time derivative of the lapse \( N \).

It is true that
\[
\phi^a \frac{\partial L}{\partial \dot{q}^a} = \frac{1}{N} \xi_{\mu,t} \dot{q}^\mu + \frac{1}{N} G_{\mu \kappa} \xi_{\alpha} \dot{q}^\alpha \dot{q}^\mu + \frac{1}{N} \xi_{\mu,0} \dot{N} \dot{q}^\mu + \frac{1}{N} K_{\lambda \alpha,t} \dot{q}^\lambda \dot{q}^\alpha + \frac{1}{N} G_{\lambda \alpha} K_{\alpha,\beta} \dot{q}^\alpha \dot{q}^\beta \dot{q}^\lambda \tag{3.5}
\]
and
\[
(\xi^\alpha + K^\alpha_{\alpha} \dot{q}^\alpha) \frac{\partial L}{\partial q^\alpha} = \frac{1}{2N} \xi^\alpha G_{\mu \nu, \kappa} \dot{q}^\mu \dot{q}^\nu - N \xi^\alpha V_{\kappa} + \frac{1}{2N} G_{\mu \nu, \kappa} K^\alpha_{\alpha} \dot{q}^\mu \dot{q}^\nu - NV_{\kappa} K^\alpha_{\alpha} \dot{q}^\alpha. \tag{3.6}
\]

We can easily observe that only up to cubic terms in the velocities appear in the last two equations: \( \dot{N} \dot{q}^2 \) and \( \dot{q}^3 \). Hence, it is reasonable to consider adopting a specific form for \( \Omega \) and \( F \), so that no terms exceeding the required order are generated. However, this is not a restriction: the existence of any term not complying with this selection is trivial, since for any such addition - for example in \( \Omega \) - a proper term can also be used in \( F \) so that they cancel each other out. In this sense we define:
\[
\Omega = \omega(t, q, N) + \omega^{(1)}(t, q, N) \dot{q}^a + \omega^{(2)}(t, q, N) \dot{N} \tag{3.7a}
\]
and
\[
F = f(t, q, N) + f^{(1)}(t, q, N) \dot{q}^a + f^{(2)}(t, q, N) \dot{q}^a \dot{q}^\beta \tag{3.7b}
\]
with \( f^{(2)}_{\alpha \beta} \) being symmetric in its indices. As one can see, the gauge function \( F \) has been assumed to be free of \( \dot{N} \). The reasoning behind this is that no terms involving \( \dot{N} \) should be produced. This acceleration is impossible to be extracted from the derivatives of (3.2), since we have already chosen a replacement rule for the maximal number of accelerations that we are allowed to use by the equations of motion, namely the \( \ddot{q}^a \)’s. Under the above mentioned assumptions, we get
\[
\Omega \frac{\partial L}{\partial N} = - \frac{\omega}{2N^2} G_{\mu \nu} \dot{q}^\mu \dot{q}^\nu - \omega V - \frac{\omega^{(1)}}{2N^2} G_{\mu \nu} \dot{q}^\mu \dot{q}^\nu \dot{q}^a - V \omega^{(1)} \dot{q}^a \tag{3.8}
\]
and
\[
\frac{dF}{dt} = f_{t} + f_{\alpha} \dot{q}^\alpha + f_{0} \dot{N} - f^{(1)}_{\alpha} \Gamma_{\mu \nu}^{\alpha} \dot{q}^\mu \dot{q}^\nu + \frac{1}{N} f^{(1)}_{\alpha} \dot{N} \dot{q}^\alpha - N^2 f^{(1)}_{\alpha} G^{\alpha} V_{\mu} \\
+ f^{(1)}_{\alpha \beta} \dot{q}^\alpha \dot{q}^\beta + \frac{1}{2} f^{(2)}_{\alpha \beta} \Gamma_{\mu \nu}^{\alpha \beta} \dot{q}^\mu \dot{q}^\nu \dot{q}^\alpha - 2 f^{(2)}_{\alpha \beta} \Gamma_{\mu \nu}^{\alpha \beta} \dot{q}^\mu \dot{q}^\nu \dot{q}^\beta + \frac{2}{N} f^{(2)}_{\alpha \beta} \dot{N} \dot{q}^\alpha \dot{q}^\beta \tag{3.9}
\]
\[- 2N^2 f^{(2)}_{\alpha \beta} G^{\alpha \beta} V_{\mu} \dot{q}^\beta + f^{(2)}_{\alpha \beta \gamma} \dot{q}^\alpha \dot{q}^\beta \dot{q}^\gamma + f^{(2)}_{\alpha \beta \gamma} \dot{q}^\alpha \dot{q}^\beta \dot{q}^\gamma + f^{(2)}_{\alpha \beta \gamma} \dot{q}^\alpha \dot{q}^\beta \dot{q}^\gamma.
\]
where again (3.2b) have been applied where necessary.

When relations (3.5), (3.6), (3.8) and (3.9) are substituted in the condition (2.9), one can demand the vanishing of the coefficients appearing in front of terms involving different powers of the velocities. This is so because none of theses quantities depends on either $\dot{q}^\mu$ or $\dot{N}$. By working in this way, from the coefficients of $\dot{N}\dot{q}^2$, we get

$$ K_{(\mu\alpha),0} + \frac{1}{N} K_{(\mu\alpha)} - \frac{\omega^{(2)}}{2N} G_{\mu\alpha} = 2f_{(\mu\alpha)}^{(2)} + Nf_{(\mu\alpha),0}^{(2)}, \quad (3.10) $$

while out of the cubic terms $\dot{q}^3$ we derive

$$ K_{(\alpha\beta\mu)} - \frac{1}{2N} \omega^{(1)}_{(\alpha} G_{\mu\beta)} = Nf_{(\alpha\beta\mu)}^{(2)}, \quad (3.11) $$

From the coefficients of $\dot{N}\dot{q}$ it follows that

$$ f_{(\mu)}^{(1)} + \frac{1}{N} f_{(\mu)}^{(1)} = \frac{1}{N} \xi_{\mu,0}. \quad (3.12) $$

This equation can be easily integrated to give

$$ \xi_{\mu} = Nf_{(\mu)}^{(1)} + \tilde{\xi}_{\mu}(t, q). \quad (3.12) $$

Subsequently, the coefficients of the quadratic terms $\dot{q}^2$ lead to

$$ \frac{1}{2} \mathcal{L}_{\xi^{\alpha}} G_{\alpha\mu} + K_{(\mu\alpha),t} - \frac{\omega}{2N} G_{\alpha\mu} = Nf_{(\alpha\mu)}^{(1)} + Nf_{(\alpha\mu),t}^{(2)} \quad (3.13) $$

Finally, from the linear terms in $\dot{q}$ and $\dot{N}$, as well as the zero order terms involving no velocities we arrive at

$$ \omega^{(2)} = -\frac{1}{V} f_{0} \quad (3.14) $$

$$ \omega_{\mu}^{(1)} = \frac{1}{NV} \xi_{\mu,t} - \frac{2N}{V} K_{(\mu\alpha)} V^{\sigma} - \frac{1}{V} f_{\mu,t} - \frac{1}{V} f_{\mu}^{(1)} + \frac{2N^2}{V} f_{\mu\sigma}^{(2)} V^{\sigma} \quad (3.15) $$

$$ \omega = -\frac{N}{V} \xi^{\alpha} V_{,\alpha} - \frac{1}{V} f_{,t} + \frac{N^2}{V} f_{\alpha}^{(1)} V^{\alpha} \quad (3.16) $$

We proceed in order to solve the system. Substitution of (3.14) in (3.10) turns the latter into

$$ K_{(\mu\alpha)} = -\frac{f}{2NV} G_{\mu\alpha} + Nf_{\mu\alpha}^{(2)} + \frac{1}{N} S_{\mu\alpha}(t, q), \quad (3.17) $$

where $S_{\mu\alpha}$ is a symmetric matrix. By inserting (3.17) in (3.11) and with the help of (3.15) we deduce that

$$ S_{(\alpha\beta\mu)} + \frac{V^{\alpha}}{V} S_{\sigma(\mu} G_{\alpha\beta)} = \frac{1}{2NV} \partial_{t}\xi_{(\mu} G_{\alpha\beta)} - \frac{1}{2V} \partial_{t} f_{(\mu}^{(1)} G_{\alpha\beta)}, $$

which due to (3.12) becomes

$$ S_{(\alpha\beta\mu)}(t, q) + \frac{V^{\alpha}}{V} S_{\sigma(\mu} G_{\alpha\beta)} = \frac{1}{2NV} \partial_{t} \tilde{\xi}_{(\mu} G_{\alpha\beta)}. $$
From this equation we reach the conclusion that - since $S_{\alpha\beta}$, $G_{\alpha\beta}$, $\tilde{\xi}^\mu$, and $V$ are not functions of $N$ - the following relations must hold.

$$\tilde{\xi}_\mu = \tilde{\xi}_\mu(q) \quad (3.18)$$
$$S_{(\alpha\beta;\mu)} = -\frac{V^\sigma}{V} S_{(\mu G_{\alpha\beta})} \quad (3.19)$$

At this point, we substitute (3.16) in (3.13) and by virtue of (3.17) we are led to

$$L_\xi G_{\alpha\mu} + \frac{1}{N} S_{\alpha\mu,t} + \frac{1}{2V} \tilde{\xi}^\sigma V_{,\sigma} G_{\alpha\mu} - \frac{N}{2V} f^{(1)} V^{,\sigma} G_{\alpha\mu} = N f^{(1)}_{(\alpha;\mu)} \quad (3.20)$$

which, by use of (3.12), becomes

$$\frac{1}{2} \left( L_\xi G_{\alpha\mu} + \frac{V^{,\sigma}}{V} \tilde{\xi}_\sigma G_{\alpha\mu} \right) + \frac{1}{N} S_{\alpha\mu,t}(t, q) = 0.$$

Again, due to the fact that none of the remaining functions depends on $N$, we have

$$S_{\alpha\mu} = S_{\alpha\mu}(q) \quad (3.21)$$
$$L_\xi G_{\alpha\mu} = -\frac{V^{,\sigma}}{V} \tilde{\xi}_\sigma G_{\alpha\mu}. \quad (3.22)$$

Finally, we are left with the relations

$$\xi_\mu = N f^{(1)}_{\mu}(t, q, N) + \tilde{\xi}_\mu(q) \quad (3.23a)$$
$$K_{(\alpha\mu)} = N f^{(2)}_{\alpha\mu}(t, q, N) - \frac{f(t,q,N)}{2NV} G_{\alpha\mu} + \frac{1}{N} S_{\alpha\mu}(q) \quad (3.23b)$$
$$\omega = -\frac{N}{V} \tilde{\xi}^\sigma V_{,\sigma} - \left( \frac{f}{V} \right)_t \quad (3.23c)$$
$$\omega^{(1)}_{\mu} = \frac{2}{V} S_{\mu\sigma} V^{,\sigma} - \left( \frac{f}{V} \right)_\mu \quad (3.23d)$$
$$\omega^{(2)} = -\left( \frac{f}{V} \right)_0 \quad (3.23e)$$

together with the conditions

$$L_\xi G_{\alpha\mu} = -\frac{V^{,\sigma}}{V} \tilde{\xi}_\sigma G_{\alpha\mu} \quad (3.24a)$$
$$S_{(\alpha\beta;\mu)} = -\frac{V^{,\sigma}}{V} S_{(\mu G_{\alpha\beta})}. \quad (3.24b)$$

We recognize (3.24a) as the variational Noether (conditional) symmetries found in [17] for singular Lagrangians ensuing from cosmological models. These symmetries are generated by vectors $\tilde{\xi}^\mu$ in the configuration space, that are simultaneous conformal Killing vectors of the mini-supermetric $G_{\alpha\beta}$ and the potential $V$ with opposite conformal factors. The second condition, (3.24b), is the one that has to be satisfied by a symmetric tensor $S_{\alpha\beta}$ in order for a contact symmetry generator
to exist. As can be seen, \( S_{\alpha\beta} \) must be a conformal Killing tensor with a conformal factor that depends on the potential \( V \).

At this point let us make a comparison between regular systems and reparametrization invariant theories (with quadratic Lagrangians): Regarding the former, we know that the relevant conditions require \( \tilde{\xi}^\mu \) to be a Killing field of both the mini-supermetric and the potential \[13\], i.e.

\[
\mathcal{L}_{\tilde{\xi}} G_{\alpha\mu} = 0 \quad \text{and} \quad \tilde{\xi}^\mu V_{,\mu} = 0
\]

and, in the case of linear contact symmetries, \( S_{\alpha\beta} \) to be a Killing tensor \[28\] satisfying

\[
S_{(\alpha\beta;\mu)} = 0 \quad \text{and} \quad S_{\alpha\beta} V_{,\beta} = f_{,\alpha},
\]

where \( f \) is a gauge function. As one can see in each case (either Noether point or contact symmetry), two major equations have to be satisfied for regular systems. On the other hand, singular Lagrangians require one condition for each type of symmetry \((3.24a)\) or \((3.24b)\). This allows for an equal or (in most cases) larger group of symmetries to be admitted in the case of reparametrization invariant systems. A regular and a singular system with the same configuration space metric \( G_{\alpha\beta} \) and the same potential \( V \), do not necessarily exhibit the same number of symmetries; see the relevant discussion in the last section.

We can reformulate the above conditions \((3.24)\) in order to show explicitly the conformal nature of the them as follows

\[
\mathcal{L}_{\tilde{\xi}} G_{\alpha\mu} = \omega G_{\alpha\mu}, \quad \mathcal{L}_{\tilde{\xi}} V + \omega V = 0
\]

\[
S_{(\alpha\beta;\mu)} = \psi_{(\mu} G_{\alpha\beta)}; \quad S_{\sigma\mu} V^{,\sigma} + \psi_{\mu} V = 0.
\]

Additionally, in the case of singular systems one can make use of the reparametrization invariance of the theory to simplify conditions \((3.24)\). This can be done in the constant potential parametrization, where the scaling \( N \mapsto \tilde{N} = NV \) is performed. Under this change, the Lagrangian reads

\[
L = \frac{1}{2N} \tilde{G}_{\mu\nu} \dot{q}^\mu \dot{q}^\nu - \tilde{N},
\]

with \( \tilde{G}_{\mu\nu} = VG_{\mu\nu} \) being the scaled, by the potential, mini-supermetric. In this parametrization, were \( \tilde{V} = 1 \), relations \((3.24)\) reduce to

\[
\mathcal{L}_{\tilde{\xi}} \tilde{G}_{\alpha\mu} = 0, \quad \tilde{S}_{(\alpha\beta;\mu)} = 0, \quad (3.28)
\]

the covariant derivative being now constructed by \( \tilde{G}_{\mu\nu} \) and the two tensors \( S_{\alpha\beta}, \tilde{S}_{\alpha\beta} \) related by \( \tilde{S}_{\alpha\beta} = V^2 S_{\alpha\beta} \). Thus, the symmetry generators become Killing fields and Killing tensors. All the above results can be gathered in order to formulate the following theorem:

**Theorem 1.** For the constrained Lagrangian in action \((3.1)\) and the symmetry generator \((3.3)\) the following statements are equivalent.
1. The vector field $\tilde{\xi}^\alpha$ is a conformal Killing vector field of the metric $G_{\alpha\beta}$ and the tensor $S_{\alpha\beta}$ is a conformal Killing tensor of the metric $G_{\alpha\beta}$ obeying the conditions

$$\mathcal{L}_{\tilde{\xi}}G_{\alpha\mu} = \omega G_{\alpha\mu}, \quad \mathcal{L}_{\tilde{\xi}}V + \omega V = 0 \quad (3.29a)$$

$$S_{(\alpha\beta\mu)} = \psi_{(\mu} G_{\alpha\beta)} , \quad S_{\sigma\mu} V^{\sigma} + \psi_{\mu} V = 0 \quad (3.29b)$$

2. The vector field $\tilde{\xi}^\alpha$ is a Killing vector field of the scaled metric $G_{\alpha\beta} = V G_{\alpha\beta}$ and the tensor $S_{\alpha\beta}$ is a Killing tensor of the conformal metric $G_{\alpha\beta}$ obeying the conditions

$$\mathcal{L}_{\tilde{\xi}}G_{\alpha\mu} = 0, \quad \mathcal{L}_{\tilde{\xi}}S_{(\alpha\beta\mu)} = 0. \quad (3.30)$$

4 Conserved quantities quadratic in velocities

We now investigate the conserved quantities that are produced by the generator we derived in the previous section. From the previous section we have that the general form of the generator $X$ of a contact Noether symmetry for the singular Lagrangian (3.1) is as follows:

$$X = \left( N f_1(t, q, N) + \tilde{\xi}^\alpha(q) \right) \frac{\partial}{\partial q^\alpha} + \left( N f_2(t, q, N) - \frac{f(q, N)}{2NV} \delta^\sigma_\mu \right) \dot{q}^\mu \frac{\partial}{\partial q^\sigma} - \left( \frac{N}{V} \tilde{\xi}^\sigma V_{,\sigma} + \frac{f}{V} \right) \frac{\partial}{\partial N} - \left( \frac{2}{V} S_{\mu\sigma} V^{\sigma} + \left( f \right)_{,\mu} \right) \dot{q}^\mu \frac{\partial}{\partial N} - \left( \frac{f}{V} \right)_{,0} \dot{N} \frac{\partial}{\partial N}. \quad (4.1)$$

However, as we shall immediately see, the terms involving the arbitrary functions $f$, $f_1^{(1)}$ and $f_2^{(2)}$ are trivial. From (2.10) we split the produced quantity $I$ into four parts

$$I = Q + I_0 + I_1 + I_2$$

with

$$Q = \left( \tilde{\xi}^\alpha + \frac{1}{N} S^\alpha_\mu \dot{q}^\mu \right) \frac{\partial L}{\partial q^\alpha} \quad (4.1)$$

$$I_0 = -\frac{f}{2NV} \dot{q}^\alpha \frac{\partial L}{\partial q^\alpha} \quad (4.2)$$

$$I_1 = N f_1^{(1)} G^{\mu\alpha} \frac{\partial L}{\partial q^\alpha} \quad (4.3)$$

$$I_2 = N f_2^{(2)} G^{\mu\alpha} \dot{q}^\sigma \frac{\partial L}{\partial q^\alpha}. \quad (4.4)$$

It is easy to prove that $I_0$, $I_1$ and $I_2$ construct trivial integrals of motion, i.e. the corresponding conserved quantities are zero on the constraint surface.
• For $I_0$:

\[ I_0 - f = - \frac{f}{2NV} \tilde{q}^\alpha \frac{1}{N} G_{\mu \alpha} \dot{q}^\mu - f = - \left( \frac{1}{2N^2 V} G_{\mu \nu} \tilde{q}^\mu \tilde{q}^\nu + 1 \right) = 0 \]

due to the constraint equation (3.2a).

• In the case of $I_1$:

\[ I_1 - f^{(1)}_\mu \dot{q}^\mu = N f^{(1)}_\mu G^{\alpha \sigma} \frac{1}{N} G_{\sigma \alpha} \tilde{q}^{\sigma} - f^{(1)}_\mu \dot{q}^\mu \equiv 0. \]

• Lastly for $I_2$:

\[ I_2 - f^{(2)}_{\mu \nu} \dot{q}^\mu \dot{q}^\nu = N f^{(2)}_{\sigma \mu} G^{\sigma \alpha} \frac{1}{N} G_{\sigma \alpha} \tilde{q}^{\sigma} - f^{(2)}_{\mu \nu} \dot{q}^\mu \dot{q}^\nu \equiv 0. \]

Thus, without any loss of generality we can consider the gauge function $F$ to be constant and deduce that the only existing conserved quantity is

\[ Q = \tilde{\xi}^\alpha \frac{\partial L}{\partial \tilde{q}^\alpha} + S^{\alpha \beta} \frac{\partial L}{\partial \tilde{q}^\alpha} \frac{\partial L}{\partial \tilde{q}^\beta}, \quad (4.5) \]

as derived by the only non trivial part of the generator

\[ X = \left( \tilde{\xi}^\alpha + \frac{1}{N} S^\alpha_{\mu} \tilde{q}^\mu \right) \frac{\partial}{\partial \tilde{q}^\alpha} - \frac{V^\mu}{V} \left( N \tilde{\xi}_\mu + 2 S_{\mu \nu} \tilde{q}^\nu \right) \frac{\partial}{\partial N}, \]

with $\tilde{\xi}^\alpha$ and $S_{\alpha \beta}$ satisfying (3.24a) and (3.24b) respectively. As seen by (4.5), in the phase space, the vector $\xi^\alpha$ gives rise to integrals of motion linear in the momenta $p^\alpha = \frac{\partial L}{\partial \dot{q}^\alpha}$, while the tensor $S^{\alpha \beta}$ quadratic. As a result, we can state the following:

**Theorem 2.** A singular system described by a Lagrangian of the general form

\[ L = \frac{1}{2N} G_{\mu \nu}(q) \tilde{q}^\mu \tilde{q}^\nu - N V(q), \quad (4.6) \]

admits an integral of motion quadratic in the momenta of the form $(p^\mu = \frac{\partial L}{\partial \dot{q}^\mu})$

\[ Q = S^{\alpha \beta}(q) p^\alpha p^\beta, \quad (4.7) \]

when $S_{\alpha \beta}$ obeys condition (3.29b).

Conformal Killing tensors (CKTs) can be divided into two major classes: reducible and irreducible. The first type consists of tensors that are trivially constructed as tensor products of other CKTs of lower rank. In the case of second rank conformal Killing tensors, the reducible CKTs are made up by conformal Killing vectors. Of course, if there exists an integral of motion linear in the momenta, we expect its square to be again a constant of motion, i.e. if there exist $\tilde{\xi}^\mu$’s that satisfy (3.24a), second rank tensor products can be constructed by them satisfying (3.24b). Indeed, let us consider a second rank tensor $S_{\alpha \beta} = \Lambda^{IJ} \tilde{\xi}^\alpha \tilde{\xi}^\beta$, with $\Lambda^{IJ}$ being a symmetric constant matrix and $\tilde{\xi}^\mu, \tilde{\xi}^\mu$ vectors that satisfy (3.24a)
indexes $I, J$ are just used to discriminate among the various vectors). Under these considerations we write

$$S_{(\alpha\beta;\mu)} = \Lambda^{IJ} \left( \tilde{\xi}_{(\alpha;\mu} \tilde{\xi}_{\beta)} + \tilde{\xi}_{(\alpha} \tilde{\xi}_{\beta;\mu) I} \right)$$

$$= \frac{\Lambda^{IJ}}{6} \left[ (\tilde{\xi}_{I;\mu;\beta} + \tilde{\xi}_{I;\beta;\mu}) \tilde{\xi}_{J;\alpha} + (\tilde{\xi}_{I;\alpha;\mu} + \tilde{\xi}_{I;\mu;\alpha}) \tilde{\xi}_{J;\beta} + (\tilde{\xi}_{I;\alpha;\beta} + \tilde{\xi}_{I;\beta;\alpha}) \tilde{\xi}_{J;\mu} 
+ \tilde{\xi}_{I;\alpha} (\tilde{\xi}_{J;\beta;\mu} + \tilde{\xi}_{J;\mu;\beta}) + \tilde{\xi}_{I;\beta} (\tilde{\xi}_{J;\alpha;\mu} + \tilde{\xi}_{J;\mu;\alpha}) + \tilde{\xi}_{I;\mu} (\tilde{\xi}_{J;\alpha;\beta} + \tilde{\xi}_{J;\beta;\alpha}) \right]$$

$$= - \frac{\Lambda^{IJ} V^\sigma}{6 V} \left( \tilde{\xi}_{I;\alpha} \tilde{\xi}_{J;\beta} G_{\alpha\mu} + \tilde{\xi}_{I;\sigma} \tilde{\xi}_{J;\alpha} G_{\beta\mu} + \tilde{\xi}_{I;\alpha} \tilde{\xi}_{J;\mu} G_{\alpha\beta} 
+ \tilde{\xi}_{I;\alpha} \tilde{\xi}_{J;\beta} G_{\beta\mu} + \tilde{\xi}_{I;\beta} \tilde{\xi}_{J;\sigma} G_{\alpha\mu} + \tilde{\xi}_{I;\mu} \tilde{\xi}_{J;\sigma} G_{\alpha\beta} \right)$$

$$= - \frac{\Lambda^{IJ} V^\sigma}{V} \tilde{\xi}_{I;\sigma} \tilde{\xi}_{J;\mu} G_{\alpha\beta}$$

$$= - \frac{V^\sigma}{V} S_{(\sigma;\mu) G_{\alpha\beta}}$$

and thus symmetries defined by $\tilde{\xi}^\alpha$’s can be trivially used to construct contact symmetries generated with the help of reducible conformal Killing tensors. An occurrence quite common in cosmology is that of conformally flat configuration metrics in mini-superspace; in this event it is well known that irreducible CKTs do exist, see [35].

5 Examples

In order to make things more transparent we apply the previously developed general theory to two specific examples; the first is the determination of the geodesics of a pp–wave spacetime while the second concerns the constrained two dimensional Hamiltonian Ermakov–Ray–Reid system.

5.1 Determination of Geodesics

When the potential $V(q)$ in the action (3.1) is constant the corresponding Lagrangian can be used in order to calculate the geodesics of a manifold with metric tensor $G_{\mu\nu}$; equation (3.2b) represents the geodesic equation in the non-affine parameter $t$. Since we have theorem (2) at hand we can fix the gauge, i.e. choose a particular functional form for $N(t)$; the most natural choice for the time variable is $N(t) = 1$ which gives the time variable $t = s$ the role of an affine parameter.

In [36] the authors calculated the Killing tensors for a number of pp–wave spacetimes along with their conformal Killing vector fields. The metric tensor $G$ for the type $Biv$ pp–wave spacetime admits the form

$$G = G_{\alpha\beta} dq^\alpha \otimes dq^\beta = -2 du \otimes dv - \frac{2}{z^2} du \otimes du + dy \otimes dy + dz \otimes dz. \quad (5.1)$$

The conformal algebra of the above spacetime is six dimensional $S_6 \supset H_5 \supset G_4$. 

12
(four Killing vectors, one homothetic, and one conformal) with basis

\[ X_1 = \partial_v, \quad X_2 = \partial_u, \quad X_3 = \partial_y, \quad X_4 = y\partial_v + u\partial_y \]
\[ X_5 = 2u\partial_u + y\partial_y + z\partial_z, \quad X_6 = u^2\partial_u + \frac{1}{2}(y^2 + z^2)\partial_v + u(y\partial_y + z\partial_z). \]

The five irreducible Killing tensors are

\[(S_1)_{\alpha\beta} = -2y^2z^2\delta^\alpha_\beta - 2y\delta^\alpha_\beta, \quad (S_2)_{\alpha\beta} = 2yz\delta^\alpha_\beta + y\delta^\alpha_\beta, \quad (S_3)_{\alpha\beta} = 2uy - \delta^\alpha_\beta, \quad (S_4)_{\alpha\beta} = 2uz^2\delta^\alpha_\beta + uz\delta^\alpha_\beta, \quad (S_5)_{\alpha\beta} = (z^2 + 2u^2z^2)\delta^\alpha_\beta. \]

Since in this case the potential is zero the symmetry generators are the four Killing vectors \( X_i, i = 1, 4 \) and the five Killing tensors \( S_j, j = 1, 5 \), which produce the nine integrals of motion \( \{ I_i = (X_i)\alpha p^\alpha, Q_j = (S_j)_{\alpha\beta}p^\alpha p^\beta \} \) along with the quadratic constraint \( \mathcal{H} \):

\[
I_1 = -u, \quad I_2 = -2z^2\dot{u} - \dot{v}, \quad I_3 = \dot{y}, \quad I_4 = \dot{y}u - \dot{y}u \tag{5.2a}
\]
\[
Q_1 = z^{-2}(-z^4\dot{u}^2 + 2yz\dot{z}^2 - (2\dot{u}^2 + z^4\dot{z}^2)u^2) \tag{5.2b}
\]
\[
Q_2 = z^{-2}(2\dot{u}^2 - z^4\dot{z}^2 + yz^2\dot{z}) \tag{5.2c}
\]
\[
Q_3 = z^{-2}(z^3\dot{u}(\dot{z} + y\dot{z}) + uy(2\dot{u}^2 + z^4\dot{z}^2 - uz\dot{z})) \tag{5.2d}
\]
\[
Q_4 = z^{-2}(2\dot{u}^2 - z^4\dot{z}^2 + uz^2\dot{z}) \tag{5.2e}
\]
\[
Q_5 = z^{-2}(u^2(2\dot{u}^2 + z^4\dot{z}^2) + z^4\dot{u}^2 - 2uz^3\dot{u}\dot{z}) \tag{5.2f}
\]
\[
\mathcal{H} = \frac{1}{2}z^{-2}(z^2 - 2\dot{u}^2 + y^2 + z^2) - 2\dot{u}^2 \tag{5.2g}
\]

where the dot represents differentiation with respect to the affine parameter \( s \).

From the integrals \( I_1, I_3 \) we can solve for \( u(s), y(s) \) i.e.

\[
u(s) = -I_1 + c_u, \quad y(s) = I_3 + c_y, \tag{5.3}
\]

where \( c_u, c_y \) are constants of integration. The next step is to eliminate the \( \dot{z}^2 \) from the integrals \( Q_2, Q_4 \) using a linear combination of them, yielding

\[
d\frac{\dot{z}^2}{ds} = J_1 + 2J_2s \Rightarrow \dot{z}^2 = J_2s^2 + J_1s + c_z,
\]

where \( c_z \) is another constant of integration and the constants \( J_1, J_2 \) are combinations of the existing constants, i.e. \( J_1 = 2(Q_4c_y - Q_2c_u) / (c_yI_1 + c_uI_3) \), \( J_2 = (Q_2I_1 + Q_4I_3) / (c_yI_1 + c_uI_3) \). Finally we are left with the function \( v(s) \) which can be determined from the integral \( I_2 \) which now assumes the form

\[
\dot{v} = \frac{2I_1}{J_2s^2 + J_1s + c_z} - I_2 \Rightarrow v(s) = \frac{4I_1}{\sqrt{4c_zJ_2 - J_1^2}} \arctan \frac{2J_2s + J_1}{\sqrt{4c_zJ_2 - J_1^2}} - I_2s + c_v,
\]

13
with \( c_v \) the last constant of integration. If one counts the constants that appear in the solution space the total number of them is 9 but the actual number that describes the geodesics equation is \( 2 \times 4 = 8 \), thus there must exist a relation between the 9 constants. The desired relation comes from the integral \( Q_2 \) and reads \( c_2 = \left( 8I_1^2 + J_1^2 \right) / (4J_2) \).

Gathering the above results the final form of the solution space is

\[
\begin{align*}
    u(s) &= -I_1 s + c_u, \\
    v(s) &= \sqrt{2} \arctan \frac{2J_2 s + J_1}{2\sqrt{2}I_1} - I_2 s + c_v \\
    y(s) &= I_3 s + c_y, \\
    z^2(s) &= \frac{(2J_2 s + J_1)^2 + 8I_1^2}{4J_2}.
\end{align*}
\] (5.4)

It is interesting to note that, since we have not used the Hamiltonian constraint (5.2g) to solve the system, the solution (5.4) describes an unconstrained system; if someone insists on the validity of the constraint, then as stated before, a relation among the constants emerges, i.e.

\[ J_2 + I_3^2 - 2I_1 I_2 + 2 = 0. \]

### 5.2 Hamiltonian Ermakov–Ray–Reid system.

The two dimensional constraint Hamiltonian Ermakov–Ray–Reid system can be described by the Hamiltonian

\[
H = \frac{1}{2N} \left( p_\alpha^2 + p_\beta^2 \right) + NV, \quad V = \frac{1}{2} \omega (\alpha^2 + \beta^2) + \frac{1}{\alpha^2} J \left( \frac{\beta}{\alpha} \right),
\] (5.5)

with \( \omega, J \) arbitrary functions of their arguments. The origins of this system can be traced back in 1880 in the pioneer work of Ermakov [37] (see [38] for the English translation) along with the generalization of Ray and Reid [39, 31]; for the unconstrained Hamiltonian version see [40, 41]. The theoretical interest in this system resides in its admittance of an integral of motion for every function \( \omega, J \), namely, the Ray-Reid invariant

\[
I = \frac{1}{2} \left( \dot{\alpha} \beta - \alpha \dot{\beta} \right)^2 + \left( \frac{\alpha^2 + \beta^2}{\alpha^2} \right) J \left( \frac{\beta}{\alpha} \right). \] (5.6)

The above integral can be produced from a Lie point symmetry with the help of Noether’s theorem, as first pointed out by Lutzky in [42]. Here we are going to show how the integral (5.6) emerges naturally from the contact symmetries of the Lagrangian.

The metric tensor \( G_{\mu\nu} = \delta_{\mu\nu} \) emerging from (5.5) corresponds to a flat two–dimensional space. It is well known that the Killing and/or conformal Killing tensors span an infinite dimensional space when the dimension of the base manifold space equals to two. The CKTs \( S_{\alpha\beta} \) along with the corresponding conformal factors \( \psi_\alpha \) are

\[
S_{\alpha\beta} = \begin{pmatrix}
    F(\alpha, \beta) & k_1 - \frac{1}{2} (f_1(z) - f_2(\bar{z})) \\
    k_1 - \frac{1}{2} (f_1(z) - f_2(\bar{z})) & F(\alpha, \beta) + f_1(z) + f_2(\bar{z})
\end{pmatrix}, \] (5.7)
and
\[ \psi_\alpha = (\partial_\alpha F(\alpha, \beta), \partial_\beta F(\alpha, \beta) + f'_1(z) + f'_2(\bar{z})), \] 
(5.8)

where \( z = \beta + \dot{\alpha}, \bar{z} = \beta - \dot{\alpha} \) and \( k_1 = \text{const.} \)

In order to apply theorem (11) we must demand the validity of the condition (3.29b) where the potential and the CKTs are given by equations (5.5) and (5.7) respectively. It is interesting that we can satisfy condition (3.29b) for every function \( \omega \) and \( J \) when the functions \( F, f_1, f_2 \) are given by

\[ F(\alpha, \beta) = \frac{4c_2 J(\beta/\alpha) - 2c_2 \beta^2 \omega(\alpha^2 + \beta^2) + c_3}{2J(\beta/\alpha) + \alpha^2 \omega(\alpha^2 + \beta^2)} \alpha^2 \]
\[ f_1(z) = c_1 + c_2 z^2, \quad f_2(z) = -c_1 + c_2 z^2 \]

With the above choice the resulting CKTs are

\[ (S_1)_{\alpha\beta} = \begin{pmatrix} 2\alpha^2 J(w) - \alpha^2 \beta^2 \omega(u) & \alpha \beta \\ 2J(w) + \alpha^2 \omega(u) & 2\beta^2 J(w) - \alpha^4 \omega(u) \end{pmatrix} \alpha \beta \]
\[ (S_2)_{\alpha\beta} = \begin{pmatrix} \alpha^2 & 0 \\ 2J(w) + \alpha^2 \omega(u) & 2J(w) + \alpha^2 \omega(u) \end{pmatrix} \alpha \beta \]

(5.9)

(5.10)

where \( w = \beta/\alpha, u = \alpha^2 + \beta^2 \). From these CKT, the respective integrals of motion along with the Hamiltonian constraint read

\[ Q_1 = \frac{2(\alpha \dot{\alpha} + \beta \dot{\beta})^2 J(w) - \alpha^2 (\dot{\alpha} \beta - \alpha \dot{\beta})^2 \omega(u)}{2J(w) + \alpha^2 \omega(u)} \] 
(5.11a)
\[ Q_2 = \frac{(\dot{\alpha}^2 + \beta^2)}{2J(w) + \alpha^2 \omega(u)} \alpha^2 \] 
(5.11b)
\[ H = \frac{1}{2} \left( \dot{\alpha}^2 + \beta^2 \right) + \frac{1}{2} \omega(u) + \frac{1}{\alpha^2} J(w). \] 
(5.11c)

Solving the Hamiltonian constraint for the function \( \omega(u) \) and substituting in \( Q_1 \) we recover the Ray–Reid invariant (5.6). We would like to remark that with a similar approach we can construct and the corresponding conservation law of the generalized Ermakov system in an \( n \)-dimensional Riemannian space [43].

6 Discussion

In this work we have generalized previous results in [17] concerning the determination of Lie point – Noether symmetries for constraint systems whose action is quadratic in the velocities. The generalization concerns the consideration of contact symmetries, i.e. of generators which depend linearly on the velocities.

The key ingredients of the present approach are:

15
(a) the treatment of the lapse function $N(t)$ and $q^\alpha(t)$’s on an equal footing in
the generator $\mathcal{L}_N = 0$, which reveals a larger number of symmetries and
(b) the dependence of the gauge function $F(t, q^\alpha, \dot{q}^\alpha)$ in the velocities $\dot{q}^\alpha$
which in turn makes it constant, i.e. non–essential in contrast to the regular
case.

If someone chooses not to implement the $N$–dependence in generator (3.3) then
the function $\Omega(t, q^\alpha, \dot{q}^\alpha)$ should be zero, i.e. $\omega = 0, \omega_1 = 0, \omega_2 = 0$. Then
equation (3.23d) $\omega = 0$ yields $f = f(q^\alpha), \mathcal{L}_\xi V = 0$ and equation (3.29a) reads
$\mathcal{L}_\xi G_{\mu\nu} = 0$. Furthermore from equation (3.23d) $\omega_1 = 0$ we have the equality
$S_{\mu\nu} V^\sigma = -V/2 (f/V)_\mu$, which forces the conformal factor $\psi_\mu$ to be $\psi_\mu = \frac{1}{2} (f/V)_\mu$
due to equation (3.29b). At this stage if we redefine the conformal Killing tensor
$S_{\alpha\beta}$ via $S_{\alpha\beta} = S_{\alpha\beta} + \frac{1}{2} (f/V) G_{\mu\nu}$ then we arrive at $\tilde{S}_{\alpha\beta} = 0$. Recapitulating
the above situation we conclude that if we do not use the $N$–dependence the
symmetry generators are described by the Killing vector fields and Killing tensor
fields of the metric $G_{\mu\nu}$ and not the conformal ones. Thus if we had at hand a
metric $G_{\mu\nu}$ that does not admit Killing fields but do admits conformal ones then
one could conclude that there are no symmetries, which is of course a fault result.

In order to give an example let us introduce the 2d metric $g$:
\[ g = g_{\alpha\beta} dx^\alpha \otimes dx^\beta \Rightarrow g = x^2 y^2 - \frac{1}{y^2} (x dx \otimes dx + y dy \otimes dy). \] (6.1)

It is an easy task for one to see that metric (6.1) does not admit Killing vectors but
since it is two dimensional admits an infinite number of conformal Killing fields.
If we choose the potential $V = \left( y(x^2 y^2 - 1) \right)^{-1}$ then there exist the two conformal
Killing fields
\[ \xi_1 = \frac{1}{\sqrt{x}} \partial_x, \quad \xi_2 = x \partial_x + y \partial_y \] (6.2)
that satisfy condition (3.29a), thus generating two symmetries. Equivalently one
can state that fixing the gauge might result in losing symmetries.

As far as the issue concerning the dependence of the gauge function $F$ in the
velocities $\dot{q}^\alpha$, we would like to stress the following facts: Let’s for a moment follow
the common practice and demand no velocity dependence in $F$, i.e. $F = F(t, q^\alpha)$, then
repeating the procedure for the evaluation of the symmetries we arrive instead of
condition (3.29b) to the condition
\[ S_{\sigma\mu} V^\sigma + \psi_\mu V + \frac{1}{2} f_{,\mu} = 0, \]
which might tempt one to think that it is more general, due to the appearance of the
arbitrary function $f(q^\alpha)$; however, this is not the case: If we redefine the conformal
Killing tensor $S_{\alpha\beta}$ in the same spirit as we did before; i.e. $S_{\mu\nu} = \tilde{S}_{\mu\nu} - \frac{1}{2} (f/V) G_{\mu\nu}$
the above equation reduces to condition (3.29b) signaling the non–essentiality of
the gauge function $f$. As we have shown the function $f$ generates the Hamiltonian
constraint $\mathcal{H}$, see equation (4.2), thus one should have expected the non-necessity of the gauge function $f$, due to the presence of the natural "gauge function" of the problem in hand, i.e. the constraint $\mathcal{H}$.

Furthermore, our results can also be used to the unconstrained case, provided that we have enough symmetries to use, so that we do not need the constraint $\mathcal{H}$; as, for example, it happens the case of section 5, where we calculated the geodesics of the pp–wave spacetime.

Finally, we conclude that the determination of the Noether symmetries (point and contact alike) of systems with Lagrangian (3.1), has been reduced to a problem of differential geometry; that is, to the determination of the Killing vectors and Killing Tensors of conformally related spacetimes.

Acknowledgements

ND acknowledges financial support by FONDECYT postdoctoral grant no. 3150016. AP acknowledges financial support of INFN.

References

[1] C. Daskaloyannis and K. Ypsilantis, “Unified treatment and classification of superintegrable systems with integrals quadratic in momenta on a two-dimensional manifold,” *Journal of Mathematical Physics* **47** no. 4, (2006) – [http://scitation.aip.org/content/aip/journal/jmp/47/4/10.1063/1.2192967](http://scitation.aip.org/content/aip/journal/jmp/47/4/10.1063/1.2192967).

[2] U. Camci, “Symmetries of geodesic motion in gdel-type spacetimes,” *Journal of Cosmology and Astroparticle Physics* **2014** no. 07, (2014) 002. [http://stacks.iop.org/1475-7516/2014/i=07/a=002](http://stacks.iop.org/1475-7516/2014/i=07/a=002).

[3] I. Freire and M. Torrisi, “Weak equivalence transformations for a class of models in biomathematics,” *Abstract and Applied Analysis* **2014** no. 2014, (2014) 546083, [http://www.hindawi.com/journals/aaa/2014/546083/cta/](http://www.hindawi.com/journals/aaa/2014/546083/cta/).

[4] R. Z. Zhdanov and V. I. Lahno, “Group classification of heat conductivity equations with a nonlinear source,” *Journal of Physics A: Mathematical and General* **32** no. 42, (1999) 7405. [http://stacks.iop.org/0305-4470/32/i=42/a=312](http://stacks.iop.org/0305-4470/32/i=42/a=312).

[5] D.-j. Huang and S. Zhou, “Group-theoretical analysis of variable coefficient nonlinear telegraph equations,” *Acta Applicandae Mathematicae* **117** no. 1, (2012) 135–183 [http://dx.doi.org/10.1007/s10440-011-9655-1](http://dx.doi.org/10.1007/s10440-011-9655-1).

[6] H. Azad and M. Mustafa, “Symmetry analysis of wave equation on sphere,” *Journal of Mathematical Analysis and Applications* **333** no. 2, (2007) 1180 – 1188 [http://www.sciencedirect.com/science/article/pii/S0022247X06012923](http://www.sciencedirect.com/science/article/pii/S0022247X06012923).
[7] M. Mustafa and A. Y. Al-Dweik, “Noether symmetries and conservation laws of wave equation on static spherically symmetric spacetimes with higher symmetries,” *Communications in Nonlinear Science and Numerical Simulation* **23** no. 13, (2015) 141 – 152. http://www.sciencedirect.com/science/article/pii/S1007570414005310.

[8] R. Gazizov and N. Ibragimov, “Lie symmetry analysis of differential equations in finance,” *Nonlinear Dynamics* **17** no. 4, (1998) 387–407. http://dx.doi.org/10.1023/A:1008304132308.

[9] M. C. Nucci, “Using lie symmetries in epidemiology,” *Electronic Journal of Differential Equations (EJDE) [electronic only]* **2005** (2005) 87–101. http://eudml.org/doc/125885.

[10] G. Bluman and S. Kumei, *Symmetries and Differential Equations*. Springer Verlag, New York, Heidelberg Germany, 1989.

[11] H. Stephani, *Differential Equations: Their Solutions Using Symmetry*. Cambridge University Press, New York, 1989.

[12] P. J. Olver, *Applications of Lie Groups to Differential Equations*. Graduate Texts in Mathematics, Vol. 107. Springer-Verlag, 2000.

[13] M. Tsamparlis and A. Paliathanasis, “Lie and Noether symmetries of geodesic equations and collineations,” *Gen.Rel.Grav.* **42** (2010) 2957–2980. arXiv:1101.5769 [gr-qc].

[14] M. Tsamparlis and A. Paliathanasis, “Two-dimensional dynamical systems which admit lie and noether symmetries,” *Journal of Physics A: Mathematical and Theoretical* **44** no. 17, (2011) 175202. http://stacks.iop.org/1751-8121/44/i=17/a=175202.

[15] A. Paliathanasis and M. Tsamparlis, “Lie point symmetries of a general class of pdes: The heat equation,” *Journal of Geometry and Physics* **62** no. 12, (2012) 2443 – 2456. http://www.sciencedirect.com/science/article/pii/S039304401200160X.

[16] A. Paliathanasis, M. Tsamparlis, and M. T. Mustafa, “Symmetry analysis of the kleingordon equation in bianchi i spacetimes,” *International Journal of Geometric Methods in Modern Physics* **0** no. 0, (0) 1550033. http://www.worldscientific.com/doi/abs/10.1142/S0219887815500334.

[17] T Christodoulakis and N Dimakis and Petros A Terzis, “Lie point and variational symmetries in minisuperspace Einstein gravity,” *Journal of Physics A: Mathematical and Theoretical* **47** no. 9, (2014) 095202. http://stacks.iop.org/1751-8121/47/i=9/a=095202.

[18] P. A. Terzis, N. Dimakis, and T. Christodoulakis, “Noether analysis of Scalar-Tensor Cosmology,” *Phys.Rev.* **D90** no. 12, (2014) 123543. arXiv:1410.0802 [gr-qc].

[19] T. Christodoulakis, N. Dimakis, P. A. Terzis, B. Vakili, E. Melas, et al., “Minisuperspace canonical quantization of the Reissner-Nordstrm black hole via conditional symmetries,” *Phys.Rev.* **D89** no. 4, (2014) 044031. arXiv:1309.6106 [gr-qc].
[20] T. Christodoulakis, N. Dimakis, P. A. Terzis, G. Doulis, T. Grammenos, et al., “Conditional Symmetries and the Canonical Quantization of Constrained Minisuperspace Actions: the Schwarzschild case,” J. Geom. Phys. 71 (2013) 127–138 arXiv:1208.0462 [gr-qc].

[21] M. Tsamparlis, A. Paliathanasis, and L. Karpathopoulos, “Autonomous three-dimensional newtonian systems which admit lie and noether point symmetries,” Journal of Physics A: Mathematical and Theoretical 45 no. 27, (2012) 275201. http://stacks.iop.org/1751-8121/45/i=27/a=275201

[22] T. Christodoulakis, N. Dimakis, P. A. Terzis, and G. Doulis, “Canonical quantization of the BTZ black hole using Noether symmetries,” Phys. Rev. D 90 (Jul, 2014) 024052, arXiv:1405.0363 [gr-qc]. http://link.aps.org/doi/10.1103/PhysRevD.90.024052

[23] N. Dimakis, T. Christodoulakis, and P. A. Terzis, “FLRW metric f(R) cosmology with a perfect fluid by generating integrals of motion,” J. Geom. Phys. 77 (2014) 97–112, arXiv:1311.4358 [gr-qc].

[24] S. Basilakos, M. Tsamparlis, and A. Paliathanasis, “Using the Noether symmetry approach to probe the nature of dark energy.” Phys. Rev. D83 (2011) 103512, arXiv:1104.2980 [astro-ph.CO].

[25] A. Paliathanasis and M. Tsamparlis, “Two scalar field cosmology: Conservation laws and exact solutions,” Phys. Rev. D 90 (Aug, 2014) 043529. http://link.aps.org/doi/10.1103/PhysRevD.90.043529

[26] G. H. Katzin and J. Levine, “Characteristic functional structure of infinitesimal symmetry mappings of classical dynamical systems. i. velocity dependent mappings of second order differential equations,” Journal of Mathematical Physics 26 no. 12, (1985) 3080–3099. http://scitation.aip.org/content/aip/journal/jmp/26/12/10.1063/1.526686

[27] R. Maartens and D. Taylor, “Lifted transformations on the tangent bundle, and symmetries of particle motion,” International Journal of Theoretical Physics 32 no. 1, (1993) 143–158. http://dx.doi.org/10.1007/BF00674402

[28] T. M. Kalotas and B. G. Wybourne, “Dynamical noether symmetries,” Journal of Physics A: Mathematical and General 15 no. 7, (1982) 2077. http://stacks.iop.org/0305-4470/15/i=7/a=018

[29] n. Ballesteros, A. Enciso, F. Herranz, and O. Ragnisco, “Hamiltonian systems admitting a rungelevenz vector and an optimal extension of bertrands theorem to curved manifolds,” Communications in Mathematical Physics 290 no. 3, (2009) 1033–1049. http://dx.doi.org/10.1007/s00220-009-0793-5

[30] R. C. OConnell and K. Jagannathan, “Illustrating dynamical symmetries in classical mechanics: The laplacerungelezen vector revisited,” American Journal of Physics 71 no. 3, (2003) 243–246. http://scitation.aip.org/content/aapt/journal/ajp/71/3/10.1119/1.1524165
[31] J. L. Reid and J. R. Ray, “Ermakov systems, nonlinear superposition, and solutions of nonlinear equations of motion,” *Journal of Mathematical Physics* **21** no. 7, (1980) 1583–1587. [http://dx.doi.org/10.1063/1.524625](http://dx.doi.org/10.1063/1.524625).

[32] B. Carter, “Global structure of the kerr family of gravitational fields,” *Phys. Rev.* **174** (Oct, 1968) 1559–1571. [http://link.aps.org/doi/10.1103/PhysRev.174.1559](http://link.aps.org/doi/10.1103/PhysRev.174.1559).

[33] W. Sarlet and F. Cantrijin, “Generalizations of noether’s theorem in classical mechanics,” *Siam Rev.* **23** no. 4, (1981) 467 – 494. [http://epubs.siam.org/doi/abs/10.1137/1023098](http://epubs.siam.org/doi/abs/10.1137/1023098).

[34] M. Crampin, “Ermakov systems, nonlinear superposition, and solutions of nonlinear equations of motion,” *Reports on Mathematical Physics* **20** no. 1, (1984) 31–40. [http://www.sciencedirect.com/science/article/pii/0034487784900697](http://www.sciencedirect.com/science/article/pii/0034487784900697).

[35] R. Rani, S. B. Edgar, and A. Barnes, “Killing tensors and conformal killing tensors from conformal killing vectors,” *Classical and quantum gravity* **20** no. 11, (2003) 1929.

[36] A. J. Keane and B. O. J. Tupper, “Killing tensors in pp-wave spacetimes,” *Classical and Quantum Gravity* **27** no. 24, (2010) 245011. [http://stacks.iop.org/0264-9381/27/i=24/a=245011](http://stacks.iop.org/0264-9381/27/i=24/a=245011).

[37] V. P. Ermakov, “Second order differential equations: conditions of complete integrability,” *Univ. Izv. Kiev.* **20** (1880) 1.

[38] A. O. Harin, “Second order differential equations: conditions of complete integrability (engl. transl.),” *Appl. Anal. Discrete Math.* **2** (2008) 123. [http://dx.doi.org/10.2298/AADM0802123E](http://dx.doi.org/10.2298/AADM0802123E).

[39] J. R. Ray and J. L. Reid, “More exact invariants for the time-dependent harmonic oscillator,” *Physics Letters A* **71** no. 4, (1979) 317 – 318. [http://www.sciencedirect.com/science/article/pii/0375960179900641](http://www.sciencedirect.com/science/article/pii/0375960179900641).

[40] C. Rogers, B. Malomed, K. Chow, and H. An, “Ermakov – ray – reid systems in nonlinear optics,” *Journal of Physics A: Mathematical and Theoretical* **43** no. 45, (2010) 455214. [http://stacks.iop.org/1751-8121/43/i=45/a=455214](http://stacks.iop.org/1751-8121/43/i=45/a=455214).

[41] C. Rogers and W. Schief, “The pulsrodon in 2 – 1-dimensional magneto–gasdynamics: Hamiltonian structure and integrability,” *Journal of Mathematical Physics* **52** no. 8, (2011) 083701. [http://dx.doi.org/10.1063/1.3622595](http://dx.doi.org/10.1063/1.3622595).

[42] M. Lutzky, “Noether’s theorem and the time-dependent harmonic oscillator,” *Physics Letters A* **68** no. 1, (1978) 3 – 4. [http://www.sciencedirect.com/science/article/pii/0375960178907387](http://www.sciencedirect.com/science/article/pii/0375960178907387).

[43] M. Tsamparlis and A. Paliathanasis, “Generalizing the autonomous keplerermakov system in a riemannian space,” *Journal of Physics A: Mathematical and Theoretical* **45** no. 27, (2012) 275202. [http://stacks.iop.org/1751-8121/45/i=27/a=275202](http://stacks.iop.org/1751-8121/45/i=27/a=275202).