On Maxwell’s electrodynamics in two spatial dimensions

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We discuss the construction of Maxwellian electrodynamics in 2 + 1 dimensions and some of its applications. Special emphasis is given to the problem of the retarded potentials and radiation, where substantial differences with respect to the usual three-dimensional case arise. These stem from the general form of the solutions of the wave equation in two dimensions, which we discuss using the Green’s function method. We believe the topics presented here could be stimulating additions to an advanced electrodynamics course at the undergraduate level. Keywords: Electromagnetism, 2D, Maxwell’s Equations, Huygen’s principle.

1. Introduction

Maxwell’s electrodynamics is a monumental construction that became one of the most paradigmatic theories of physics, if not the most celebrated one. In undergraduate courses, it is often presented as the definition of a complete theory: a construction whose internal machinery alone allows for spectacular predictions, such as the existence of electromagnetic waves.

In contemporary physics, it has become customary to study successful physical theories in spacetimes with a different number of dimensions. The motivation for studies of this type go beyond mere scientific curiosity. With the advent of Quantum Electrodynamics (QED) one immediately started considering its realisations in two and one spatial dimensions. Schwinger, in 1962, has discussed some of the peculiarities of QED in a single spatial dimension - a theory sometimes referred to as the Schwinger model. It was shown that quantum corrections generate a mass for the photon without breaking gauge invariance, a rather peculiar fact. Also, it is known that QED becomes confining in such a limit, i.e. isolated monopoles cannot be found since they would be confined in neutral bound states and an infinite amount of energy would be required to extract a charge from these “atoms”. Therefore, QED in one spatial dimension is a toy model for Quantum Chromodynamics (QCD), the theory of strong interactions, which does exhibit confinement in our universe. As a matter of fact, as of today, our understanding of the spectrum of quark-antiquark bound states is partially based on a solution of quantum chromodynamics in a certain limit in only one spatial dimension, due to ’t Hooft.

It seems, therefore, that classical electrodynamics would be an excellent framework to introduce undergraduate students to the issues one may face when studying a physical theory in a spacetime with a different number of spatial dimensions. However, a search for references on Maxwell’s equations in two spatial dimensions reveals that the available material is fragmented and scarce. The present work is an attempt to fill part of this gap in the spirit of an advanced undergraduate course on electrodynamics. Our main goal is to point out differences and similarities between Maxwell equations — and part of their phenomenology — in three and two spatial dimensions. We believe this is an interesting topic to be covered in courses at this level, and our experience shows that motivated students find very interesting to understand the construction of this theory and the potential implications for a hypothetical universe with only two spatial dimensions.

Part of the material presented here can be found in the literature, albeit in a fragmented form. About 30 years ago, a number of papers on the quantum solution of the Hydrogen atom in 1 + 1 and 2 + 1 dimensions were published. In these papers, one can find short discussions about the electric field and the Coulomb potential in one and two spatial dimensions, since this is the only necessary electromagnetic input to the wave-mechanics calculation. Classical text books do not discuss the topic. To the best of our knowledge, the topic was discussed in the spirit of our work in the Refs. The unpublished work of Ref. by K. T. McDonald is probably the most self-contained in the literature and it can be considered as the main reference of the present paper. Here we will borrow part of the notation and the strategy employed in . However, we will develop in detail topics that are not discussed in his work such as the retarded potentials and radiation in 2 + 1 dimensions.

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We will refer to the dimensions of space-time as D + 1, where D is the number of spatial dimensions.

After the online publication of our manuscript, a new version of Ref. appeared encompassing some of the topics that we discuss in the present work, as well as correcting an inconsistency found by us in the expressions of the retarded potentials given in the previous version of Ref.
also try to follow a physical reasoning in order to explain some of the results that can appear as counterintuitive when going from three to two spatial dimensions. With this aim, it is useful to carefully consider the parallel between the theory in the 2+1 world and usual electro-dynamics with one dimension “stretched out” to infinity, something that was first exploited by Hadamard [3].

In this work, we follow the idea that electrodynamics — or a theory totally analogous to it — can exist in any spacetime, regardless of the dimensionality [8,9]. In fact, within this assumption, electrodynamics can even be extended to a fractional number of spatial dimensions [10,11,12]. In the construction of electrodynamics in two dimensions one faces a few issues. The first one, as is well known, is related to the fact that the Coulomb force must change. Arguably, the most natural way to construct the theory is to assume that the essence of Gauss’ law remains unaltered, but this requires that the electric field of a point charge now falls off as the inverse of the distance which, in turn, entails a logarithmic electrostatic potential. Additionally, part of the vector calculus must change. The absence of a right-hand rule is obvious and the concept of curl, for example, must be redefined. As a result of this construction, the magnetic field must be qualitatively different; it turns out that it cannot be a vector anymore: it becomes a scalar field. Once these basic issues are figured out, one can ask how would the multipole expansion be in such a universe, what are the new retarded potentials, what happens to electromagnetic waves or, yet, how would dipole radiation arise. In the remainder of this paper, we discuss these questions having in mind a curious undergraduate student, one with particular interest for theoretical physics, or a teacher that would like to use part of this material in a somewhat advanced course on electrodynamics. We will, therefore, stick to usual SI units. Although sometimes this results in somewhat cumbersome expressions, it becomes easier to make contact with the usual undergraduate text books [6].

One of the main differences between electrodynamics in two and three spatial dimensions appears in the retarded potentials. The reason for this difference is directly linked to the fact that the solutions to the wave equation, satisfied by the potentials, change qualitatively when going from three to two dimensions, a fact that is well known in the literature in other contexts [12]. This difference is often discussed in the framework of Huygens’ principle, which states that every point on a wave front is itself the source of (spherical) waves. This principle is valid only in spacetimes with $D \geq 3$ and odd [9]. Huygens’ principle relies on the fact that all waves propagate with a single speed, $c$ in our case. In two dimensions, however, a solution to the wave equation can be understood as a superposition of waves travelling with speeds ranging from zero to the maximum value $c$, with which the first wave front travels [13]. This fact leads to important differences in the retarded potentials as compared with the $D = 3$ case, and it becomes less obvious how these connect with their static counterpart. Here, we discuss this issue and the problem of radiation in the context of a hypothetical universe with $2+1$ dimensions. We show that one benefits from the systematic use of the Green’s function method to find the correct form of the retarded potentials. This is sometimes not emphasized in usual electrodynamics [6] where the idea of time retardation is more intuitive, due to the validity of Huygen’s principle.

We organize this paper as follows. In Sec. 2 we discuss electrostatics in 2+1 dimensions. The starting point is Gauss’ law, from which we can obtain the electric field and the associated scalar potential. In Sec. 3 we will construct the full set of Maxwell’s equations starting from the relativistic formulation. We then discuss Poynting’s theorem, wave equations, the retarded potentials, and electric dipole radiation. The mathematical notation for vector calculus in $2+1$ dimensions will be introduced when necessary. In Sec. 4 we summarize the results and conclude.

2. Electrostatics in 2+1 dimensions

A generalization of Gauss’ law is possibly the best starting point to construct electrostatics in 2+1 dimensions. This route is not unique, as we mentioned in the introduction, but is arguably the most natural way to construct the theory. The idea consists in postulating that the essence of Gauss’ law is unaltered when one considers spacetimes with a different number of spatial dimensions: the flux of electric field lines, $\Phi_E^{(D)}$, through a closed (hyper)surface in $D$ dimensions contains information about the total charge enclosed by the surface. For $D = 3$, with an electric charge $q$ at the origin of the coordinate system, and choosing a spherical surface of radius $r$, this means that

$$\Phi_E^{(3)} = |E| 4\pi r^2 = \frac{q}{\epsilon_0}, \quad (1)$$

from which the usual behavior of the electric field $|E| \sim 1/r^2$ can be read off. Although we are mostly interested in 2+1 dimensions, it is interesting to start with a discussion in $D+1$ dimensions. Considering the Gaussian surface to be a hypersphere of radius $r$, the formulation of Eq. (1) can be straightforwardly generalized to

$$\Phi_E^{(D)} = |E|r^{D-1} \int_S d\Omega_D = |E|r^{D-1} \frac{2\pi^{D/2}}{\Gamma(D/2)} = \frac{q}{\epsilon_0}, \quad (2)$$

where we have used the well known result for the integration over the element of solid angle in $D$ spatial dimensions, $d\Omega_D$, written in terms of the usual $\Gamma$ function [14]. Throughout this work we will use the symbol $E$ for the generalized electric field, $q$ for the generalized charge and so on. Their physical meaning follow...
from usual electrodynamics although there are important differences with respect to their counterpart in 3+1 dimensions, as we discuss below.

Exploiting the hyperspherical symmetry, the electric field of a point charge at the origin can be written as

\[ E = \frac{q}{2\pi\epsilon_0} \frac{1}{r^{D-1}}. \]  

(3)

The dependence of the electric field, and hence of Coulomb’s law, with \( r \) is now \( 1/r^{D-1} \). Thus, the usual 3+1-dimensional fall off as \( 1/r^2 \) can be understood as a consequence of the dimensionality of our space-time. This point, albeit somewhat obvious, has sometimes been missed in the literature, as pointed out in Ref. [3].

After this short digression, we can now make \( D = 2 \) in Eq. (3) to conclude that the electric field \( E \) of a point charge located at \( r' \) in 2 + 1 dimensions is given by

\[ E(r) = \frac{q}{2\pi\epsilon_0} \frac{(r - r')}{|r - r'|^2}, \]  

(4)

which falls off as the inverse of the distance. One should remark that, as Eq. (4) shows, the units of \( \epsilon_0 \) (and \( \mu_0 \)) change in two dimensions.

The generalization to a continuous surface charge distribution, \( \sigma(r) \), is direct

\[ E = \frac{1}{2\pi\epsilon_0} \int d^2r' \sigma(r') \frac{(r - r')}{|r - r'|^2}. \]  

(5)

Taking the divergence of the last equation\(^4\) and using that, for \( D = 2 \),

\[ \nabla \cdot \left( \frac{r - r'}{|r - r'|^2} \right) = 2\pi \delta^2(r - r'), \]  

(6)

we recover the differential form of Gauss’ law

\[ \nabla \cdot E = \frac{\sigma}{\epsilon_0}, \]  

(7)

which is essentially the same as in the usual space-time with the replacement of the volumetric charge density by a surface one.

Next, we look for an electrostatic potential. In two dimensions, a vector field \( \mathbf{F} \) is conservative — and hence can be written in terms of the gradient of a scalar potential — if \( \partial_y(F)_x = \partial_x(F)_y \). To make contact with the idea of the usual curl of a vector, it is useful to start by defining, in two dimensions, a vector \( \mathbf{v}_\perp \) as

\[ \mathbf{v}_\perp = (v_y, -v_x), \]  

(8)

which is, by construction, perpendicular to the vector \( \mathbf{v} = (v_x, v_y) \). We define then the operator \(^5\)

\[ \nabla_\perp = (\partial_y, -\partial_x). \]  

(9)

With this definition, a vector field is conservative if

\[ \nabla_\perp \cdot \mathbf{F} = 0. \]  

(10)

\( \nabla \cdot E = 0 \), and therefore the electrostatic field can be written as \( E = -\nabla V \). We can then conclude that the potential obeys, as expected, Poisson’s equation in two dimensions

\[ \nabla^2 V = -\frac{\sigma}{\epsilon_0}. \]  

(10)

The solution of this equation by the Green’s function method yields

\[ V(r) = -\frac{q}{2\epsilon_0} \ln \left( \frac{r}{a} \right), \]  

(11)

from which we can obtain the potential of a point charge at the origin as

\[ V_\infty(r) = \frac{1}{4\epsilon_0} \int d^2r' \frac{\rho(r')}{|r - r'|}, \]  

(13)

from which the potential of the point charge at the origin is obtained.

In the usual electromagnetism, one meets a logarithmic potential of the type of Eq. (13) when studying the idealised case of an infinite uniformly charged straight wire. Accordingly, the electric field of such a wire falls off as \( 1/r \), as in Eq. (4). In fact, many of the phenomena of electrodynamics in two dimensions can be understood as if spacetime were a conveniently chosen two dimensional slice of the usual spacetime with point charges “stretched” into infinite wires (an idea pioneered by Hadamard [9]). In the case of the field of a point charge, the problem is essentially equivalent to considering the field of the infinite charged wire in three dimensions on a two-dimensional plane perpendicular to the wire. In the static case this may seem obvious, but in Sec. 3 we will show that even radiation can be understood in terms of this system where the point charges are extended to infinity.

One should remark that the potential of a point charge in \( 2 + 1 \) dimensions, Eq. (13), is qualitatively different from the usual \( 3 + 1 \) case. Note that Eq. (13) diverges when \( r \rightarrow \infty \) (in addition to the divergence at the origin, which is also present in \( 3+1 \)). This means that a hydrogen atom in \( 2 + 1 \) dimensions would be a confined system: an infinite amount of energy would be needed to extract the electron from the atom, in direct analogy to what happens in quark-antiquark bound states where, at large distances, a linear potential, usually written as \( V(r) \approx a \), confines the particles inside the bound state \(^{12}\). Therefore, in a universe with \( 2 + 1 \) dimensions chemistry would be

\(^6\)For the sake of clarity, we will often denote three dimensional quantities with an explicit subindex, as in \( V_{3+1} \).
radically different. In addition, the total energy of a point charge is infinite not only due to the divergence for small distances, as in the usual 3+1 case, but also for large distances as well. In fact, any charge distribution with non-zero net charge will carry an infinite amount of energy. This very peculiar fact is in part solved by the hypothesis of confinement: i.e. if charges in such a theory are confined, as color charge is confined in QCD, then all charge distributions would have a zero monopole term and isolated charges would not be observed in the universe.

In the construction of electrodynamics, the discontinuity of the electric field plays an important role. In 3+1 dimensions, the electrostatic field is discontinuous when crossing a surface charge. Accordingly, in 2+1, the discontinuity appears when crossing a line charge. The boundary conditions can be found with the usual method, by considering a Gaussian “box”, which now becomes a rectangle in two dimensions, around the line charge \( \lambda \). Making the lateral size of the box approach zero one finds that the component of the electric field perpendicular to the surface is discontinuous by an amount given by

\[
E_{\text{above}} - E_{\text{below}} = \frac{\lambda}{\epsilon_0} \mathbf{n},
\]

where \( E_{\text{above}} \) and \( E_{\text{below}} \) are defined by the choice of the normal vector \( \mathbf{n} \).

The scalar potential is continuous (as in 3+1 dimensions), but again the gradient of the potential inherits the discontinuity of the electric field

\[
\nabla V_{\text{above}} - \nabla V_{\text{below}} = -\frac{\lambda}{\epsilon_0} \mathbf{n}.
\]

The last two equations are examples of aspects of the construction that are completely straightforward, and follow exactly what one could intuitively expect, with the surface charge replaced by the line charge \( \lambda \).

One can proceed developing the standard electrostatics tool box in 2+1 dimensions. Let us comment, for example, on the multipole expansion of the scalar potential for points away from the source. The result of Eq. (11) can be straightforwardly expanded for \( r > r' \). In polar coordinates one finds

\[
V(r, \phi) = -\frac{Q}{2\pi \epsilon_0} \ln \left( \frac{r'}{r} \right) + \frac{1}{2\pi \epsilon_0} \sum_{n=1}^\infty A_n \cos(n\phi) + B_n \sin(n\phi) \frac{1}{nr^n},
\]

with the coefficients \( A_n \) and \( B_n \) given by

\[
A_n = \int d^2r' (r')^n \sigma(r') \cos(n\phi'),
\]

\[
B_n = \int d^2r' (r')^n \sigma(r') \sin(n\phi').
\]

The structure of the expansions shows that we can still talk about a “monopole term”, a “dipole term”, and so on, although their behaviour with \( r \) is now different compared with usual electrodynamics. The monopole term contains the total charge \( Q \) and is, as expected, logarithmic. Charge confinement would essentially forbid the monopole term in two dimensions and therefore \( Q_T = 0 \). It is followed by a dipole term, which would be the leading effect here that falls off as \( 1/r \), then by a quadrupole term that behaves as \( 1/r^2 \) and so on. By considering only the dipole term one can then write

\[
V_{\text{dip}}(r) = \frac{1}{2\pi \epsilon_0 r} \int d^2r' r'\sigma(r') \cos(\phi - \phi'),
\]

which can be cast as

\[
V_{\text{dip}}(r) = \frac{1}{2\pi \epsilon_0} \frac{\mathbf{r} \cdot \mathbf{p}}{r},
\]

with the dipole vector moment given by

\[
\mathbf{p} = \int d^2r' r' \sigma(r').
\]

One clearly sees that in the multipole expansion, apart from the different power counting in \( 1/r \), the only qualitatively different term is the leading monopole, which becomes logarithmic in \( 2+1 \) dimensions.

The general solution to the Poisson equation in two dimensions using the method of separation of variables can be found in almost any book on mathematical methods for physics. With Eq. (16) it is easy to make contact with these results and establish the standard correspondence between \( V(r) \) obtained from the multipole expansion and Poisson’s equation in polar coordinates. Typical electrostatic problems can be solved with these techniques and their solution do not differ significantly from the familiar ones in \( 3+1 \) dimensions. In the remainder, we will focus on electrodynamics and the problem of radiation, where more subtle issues arise.

3. Electrodynamics and radiation

An elegant strategy to obtain the full set of Maxwell’s equations in \( 2+1 \) dimensions consists in using as a starting point the Faraday tensor, \( F_{\mu\nu} \). For completeness, and to establish the notation, we briefly review in our conventions Maxwell’s equations in \( 3+1 \) dimensions in the manifestly covariant form. We define the four-potential \( A^\mu = (V/c, \mathbf{A}) \) formed from the scalar potential \( V \) and the vector potential \( \mathbf{A} \). We then build the tensor \( F_{\mu\nu} = \partial^\mu A^\nu - \partial^\nu A^\mu \) which explicitly reads, in our conventions,

\[
F_{\mu\nu}^{3+1} = \begin{pmatrix}
0 & -E^x/c & -E^y/c & -E^z/c \\
E^x/c & 0 & -B^y & B^z \\
E^y/c & B^z & 0 & -B^x \\
E^z/c & -B^y & B^x & 0
\end{pmatrix}.
\]

We use \( g_{\mu\nu} = \text{diag}(+1,-1,-1,-1) \). We will keep all factors of \( c \) explicit and, accordingly, we work with the coordinate \( x^0 = ct \) and \( \partial^0 = \frac{\partial}{\partial t} \).
As in three dimensions, electric charge is conserved.

\[ \partial_\mu F^{\mu\nu} = \mu_0 J^\nu, \]

\[ \partial_\mu F^{\mu\nu} = 0, \]

where \( J^\mu \) is the four-current defined as \( J^\mu = (c\rho, J) \) with \( \rho \) and \( J \) being respectively the volume charge and current densities.

In 2+1 dimensions, Faraday’s tensor has now only three independent components. These will contain \( E_x \) and \( E_y \) as well as (the now scalar) magnetic field \( B \). In a world where all electric effects are confined to the \( xy \) plane, the magnetic field would be along the \( z \) direction. This is sufficient to see that the last row and the last column of \( F^{\mu\nu} \) given in Eq. (21) must now absent. Faraday’s tensor in 2+1 dimensions can then be obtained by “removing” the last row and the last column of Eq. (21) \( F^{\mu\nu} \). It then reads \[ F^{\mu\nu} = \begin{pmatrix} 0 & -E^x/c & -E^y/c \\ E^x/c & 0 & -B \\ E^y/c & B & 0 \end{pmatrix}. \] (24)

The immediate consequence is that now the magnetic field has only its would-be \( z \) component; it is no longer a vector field and it becomes a (pseudo) scalar field. The equivalent of Eq. (22) can be obtained in a straightforward manner. To write the 2+1 equivalent of Eq. (23) one must notice that now the dual tensor will become a vector, since the Levi-Civita symbol loses one index. We have then \( F^\mu = \epsilon^{\mu\nu\sigma} F_{\nu\sigma} \). In covariant form, Maxwell’s equations in 2+1 dimensions are then

\[ \partial_\mu F^{\mu\nu} = \mu_0 j^\nu, \]

\[ \partial_\mu F^\nu = 0, \]

where in \( j^\nu \) the surface charge and current densities appear.

Differentiating Eq. (25) and remembering that the electromagnetic tensor is antisymmetric, we obtain

\[ \partial_\mu j^\nu = 0, \]

which is the covariant form of the continuity equation. As in three dimensions, electric charge is conserved.

One can then work out the differential Maxwell’s equations in 2+1 dimensions in terms of the fields as

\[ \nabla \cdot E = \frac{\sigma}{\epsilon_0}, \]

\[ \nabla \times E = \frac{\partial B}{\partial t}, \]

\[ \nabla \times B = \mu_0 j + \frac{1}{c^2} \frac{\partial E}{\partial t}, \]

where \( \sigma \) and \( j \) are the surface charge and surface current, respectively, and we have made use, again, of the operator of Eq. (9). We have also used that in 2+1 it remains true that \[ c_2 = (\mu_0 \epsilon_0)^{-1} \] (as we will show below). One notices that there is no equivalent to the three-dimensional magnetic Gauss’ law.

Electrodynamics becomes complete with the Lorentz force law. From the expression of the Lorentz force in terms of the Faraday tensor and the 4-velocity, \( u_\mu \), which reads \( F^\mu = (q/c) F^{\mu\nu} u_\nu \), one obtains in 2+1 dimensions the following result

\[ f = q(E + \upsilon \times B). \] (31)

An immediate consequence of Maxwell’s equations in 2+1 dimensions is the existence of electromagnetic waves in two spatial dimensions. To follow the usual steps done in the derivation of the wave equations for \( E \) and \( B \) one needs the following two-dimensional vector relation \( \nabla \times (\nabla \times F) = \nabla^2 F - \nabla (\nabla \cdot F) \). Using it in Eq. (29) we get \( \nabla \times (\nabla \times E) = \partial_\mu (\partial_\mu B) \) and consequently

\[ \nabla^2 E - \frac{1}{c^2} \frac{\partial^2 E}{\partial t^2} = \frac{1}{\epsilon_0} \nabla \sigma + \mu_0 \partial_\mu j, \]

which is the general wave equation with sources. A similar procedure gives the wave equation for the magnetic field, which is now a scalar wave

\[ \nabla^2 B - \frac{1}{c^2} \frac{\partial^2 B}{\partial t^2} = \mu_0 \nabla \times j. \] (33)

In the context of electromagnetic waves, a crucial role is played by the Poynting vector. It may not be obvious how to construct this vector in two spatial dimensions. In order to closely follow the familiar derivation, it turns out to be useful to work with the field \( E_\perp = (E_y, -E_x) \). One can then follow the familiar steps of the derivation of energy balance in the electromagnetic field. Starting from the power \( dW/dt = \int \epsilon_0 E^2/2 \mu_0 + \epsilon_0 B^2/2 \mu_0 - \frac{1}{\mu_0} \int d^2r \nabla \cdot (E \times B) \)

\[ \frac{dW}{dt} = - \frac{d}{dt} \int_A d^2r' \left( \frac{B^2}{2\mu_0} + \epsilon_0 E^2/2 \right) - \frac{1}{\mu_0} \int_A d^2r' \nabla \cdot (E \times B). \]

Upon use of the divergence theorem, one can then relate the derivative with respect to time of the mechanical and electromagnetic energies to the flux of the Poynting vector, defined as

\[ S = \frac{1}{\mu_0} E_\perp B \]

through a closed loop. Or, in the usual differential form,

\[ \frac{\partial}{\partial t} (u_{\text{mech}} + u_{\text{em}}) + \nabla \cdot S = 0. \] (35)

The meaning of the Poynting vector given in Eq. (34) can be understood as follows. The vector \( E_\perp B \) is perpendicular to \( E \) and by construction points in a direction that

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\[ \epsilon \]

The same procedure can be applied once more to obtain \( F^{\mu\nu} \) in 1+1 dimensions, as discussed in more detail in [8]. Welegate to the supplementary material an alternative derivation of this result.
would also be perpendicular to the would-be third dimension. In other words, if we consider in 3+1 dimensions the fields \( \mathbf{E} = (E_x, E_y, 0) \) and \( \mathbf{B} = (0, 0, B) \) the Poynting vector would be proportional to \( \mathbf{E} \times \mathbf{B} = (E_y B, -E_x B, 0) \), which has the same \( x \) and \( y \) components as \( \mathbf{E} \perp \mathbf{B} \) in 2+1 dimensions.

The requirement for radiating fields is that they carry energy away from the system to infinity. For this to happen, since now the flux of the Poynting vector is to be computed through a closed loop, one obtains that for radiating fields \( \mathbf{S} \sim 1/r \), suggesting that radiation fields should fall off as \( \mathbf{E}_{\text{rad}} \sim 1/\sqrt{r} \) and \( \mathbf{B}_{\text{rad}} \sim 1/\sqrt{r} \).

As in 3+1 dimensions, it is advantageous to treat the problem of radiation starting from the potentials \( \mathbf{V} \) and \( \mathbf{A} \). To follow a similar route in \( D = 2 \), we must first establish the relation between the vector potential and the magnetic field. This can be done in a way consistent with Eq. (24) through \( 8 \)

\[
\mathbf{B} = -\nabla_{\perp} \cdot \mathbf{A}. \tag{36}
\]

From Faraday’s law one concludes that \( \nabla_{\perp} \cdot (\mathbf{E} + \partial_t \mathbf{A}) = 0 \) and, thus, the quantity \( \mathbf{E} + \partial_t \mathbf{A} \) is conservative and can be written as the gradient of a scalar potential, which implies \( \mathbf{E} = -\nabla \mathbf{V} = \partial_t \mathbf{A} \). As usual, gauge invariance allows one to choose \( \nabla \cdot \mathbf{A} = -\frac{1}{c^2} \partial_t \mathbf{V} \), which corresponds to Lorentz gauge. The equations satisfied by the potentials are then the wave equation with sources. In 2+1 dimensions, and reminding the reader that \( \nabla^2 = \partial_x^2 + \partial_y^2 \), the equations are \( 10 \)

\[
\begin{align*}
\nabla^2 \mathbf{V} - \frac{1}{c^2} \partial_t^2 \mathbf{V} &= -\frac{\sigma}{\epsilon_0}, \tag{37} \\
\nabla^2 \mathbf{A} - \frac{1}{c^2} \partial_t^2 \mathbf{A} &= -\mu_0 \mathbf{j}. \tag{38}
\end{align*}
\]

At this point, we must look for the solutions of Eqs. (37) and (38). In the usual 3+1 case, they would be the standard retarded potentials. Here, an important difference between three and two spatial dimensions occurs. In three dimensions, the retarded potentials can be obtained in a rather intuitive, albeit not very rigorous way, by simply imposing in the static solution that in the dynamical case information travels with the speed of light, as suggested by the wave equations satisfied by \( \mathbf{V} \) and \( \mathbf{A} \). In practice, one is allowed to take the solution to Poisson’s equation, Eq. (13), and simply replace in the integrand \( \rho(r') \rightarrow \rho(r', t_r) \) with the retarded time given by \( t_r = t - |\mathbf{r} - \mathbf{r}'|/c \). This solution is correct in three dimensions and has the advantage of being very intuitive.

A more formal derivation of the retarded potentials can be obtained using the Green’s function method, which will be the basis for our discussion of the two-dimensional case. It is, therefore, useful to briefly discuss this solution in \( D = 3 \). We define the Green’s function of the d’Alembertian operator in three dimensions as \( 17 \)

\[
\left( \nabla^2 - \frac{1}{c^2} \partial_t^2 \right) G_{3+1}(\mathbf{r}, t) \equiv \Box^2 G_{3+1}(\mathbf{r}, t) = -4\pi \delta(t)\delta(\mathbf{r}). \tag{39}
\]

The solution for \( \mathbf{V}(\mathbf{r}, t) \) is obtained, by construction, as

\[
V_{3+1}(\mathbf{r}, t) = \frac{1}{4\pi\epsilon_0} \int d^3 r' \int dt' G_{3+1}(\mathbf{r} - \mathbf{r}', t - t') \rho(\mathbf{r}', t'). \tag{40}
\]

In 3+1 dimensions, the Green’s function of the d’Alembertian is, explicitly,

\[
G_{3+1}(\mathbf{r} - \mathbf{r}', t - t') = \frac{\delta(t' - t_r)}{|\mathbf{r} - \mathbf{r}'|}. \tag{41}
\]

The integration over \( t' \) immediately leads to the usual form of the retarded potentials. For the scalar potential we have

\[
V_{3+1}(\mathbf{r}, t) = \frac{1}{4\pi\epsilon_0} \int d^3 r' \rho(\mathbf{r}', t_r) |\mathbf{r} - \mathbf{r}'| \tag{42}
\]

The Green’s function of Eq. (41) can be interpreted as the potential associated with a point charge that exists at point \( \mathbf{r}' \) in space only at the instant \( t' \), as per Eq. (39). The Dirac \( \delta \) in the numerator shows that at another point in space, \( \mathbf{r} \), at a later time, \( t \), the effect of the existence of this point charge that blinks at \( t = t' \) will only be felt once, precisely at the time \( t = t' + |\mathbf{r} - \mathbf{r}'|/c \). In turn, this is linked to the fact that the general solution to the spherical wave equation in three dimensions is given by functions of the type \( f(r \pm ct)/r \), i.e., waves that propagate with a single speed \( c \). This is sometimes called “sharp” wave propagation, and is a feature that is consistent with Huygens’ principle. \( 17 \)

Let us now turn to our case of interest, namely the retarded potentials in two spatial dimensions. Now, a naïve attempt to obtain the solution by simply replacing \( \sigma(r') \rightarrow \sigma(r', t_r) \) in Eq. (11) is wrong. This is so because the intuitive idea about the way information propagates, embedded in this naïve solution, breaks down in two spatial dimensions — in other words, Huygens’ principle is not valid in two spatial dimensions. \( 13 \). To understand what is happening it is better to use the Green’s function method. Now we have

\[
\Box^2 G_{2+1}(\mathbf{r}, t) = -2\pi \delta^{(2)}(\mathbf{r})\delta(t). \tag{43}
\]

The solution for this Green’s function can be found in the literature \( 12 \), it reads

\[
G_{2+1}(\mathbf{r} - \mathbf{r}', t - t') = \frac{\theta(ct - s)}{\sqrt{2\pi s'/c^2}} \tag{44}
\]

\text{with } \quad s = |\mathbf{r} - \mathbf{r}'|, \quad \tau = t - t'.

This is, of course, inconsistent with charge conservation, but it remains useful as a mathematical device to interpret the Green’s function. The same applies to the non-physical charge distribution of Eq. (45).
The last observation provides us with some additional intuition regarding the Green’s function $G_{2+1}(r, t)$ given in Eq. (44) and suggests the following practical result. From the integral representation we derived above we can write

$$G_{2+1}(r, t) = \int_{0}^{\infty} dz \frac{\delta(t - \sqrt{\overline{z^2 + r^2}}/c)}{\sqrt{\overline{z^2 + r^2}}},$$

(49)

and apply the change of variables $v(z) = \frac{r}{\sqrt{\overline{z^2 + r^2}}}$ suggested by the analogy with the infinite charged wire to obtain

$$G_{2+1}(r, t) = \int_{0}^{\infty} dv \frac{c/v}{\sqrt{c^2 - v^2}} \delta(t - r/v),$$

(50)

in agreement with the interpretation given above, namely that information in two dimensions appears to travel with different speeds, ranging from 0 to $c$, which immediately implies that Huygens’ principle is not valid.

Having discussed the Green’s function in two dimensions in detail, one can then write the solution for the retarded potentials. The analog to Eq. (40) in 2+1 dimensions is

$$V_{2+1}(r, t) = \frac{1}{2\pi\epsilon_0} \int d^2r' \int dt' G_{2+1}(r-r', t-t') \sigma(r', t'),$$

(51)

and the solutions to $V(r)$ and $A(r)$ can be written as

$$V(r) = \frac{1}{2\pi\epsilon_0} \int d^2r' \int dt' \frac{\theta(ct - s)}{\sqrt{c^2 - s^2/c^2}} \sigma(r', t'),$$

(52)

$$A(r) = \frac{1}{2\pi\sigma_0} \int d^2r' \int dt' \frac{\theta(ct - s)}{\sqrt{c^2 - s^2/c^2}} j(r', t'),$$

(53)

with the notation of Eq. (44). In two dimensions it is, therefore, less trivial to make contact with the static case, given for the scalar potential in Eq. (11). In particular, it is not as obvious how the potentials of a point charge appear in the static limit. It is interesting then to consider an unphysical case which violates charge conservation, but that is instrumental in understanding how the new formulae for the retarded potentials are compatible with the static ones. Let us take the potentials for an oscillating monopole with localized charge distribution at an arbitrary point $r_0$ given by

$$\sigma(r, t) = q\delta^2(r - r_0)e^{i\omega t}.$$  

(54)

(One is tacitly assuming that the physical charge is connected to the real part of the last expression.) The static limit can be recovered doing $\omega \rightarrow 0$. Using this charge density in the scalar potential the integrals over $r'$ can

A superseded version of Ref. [8] contained a mistake in the expression of the retarded potentials. The present version of Ref. [8], published after our work was made public, is now correct and in full agreement with our expressions.

This choice is partially motivated by the fact that we will use this oscillating monopole as a building block for the physical dipole when studying radiation. Another possible route is the direct use of the integral representation of the Green’s function in Eq. (49).
be done with the delta function present in $\sigma(r, t)$ and we find

$$V(r, t) = \frac{q}{2\pi\varepsilon_0} \int_{-\infty}^{+\infty} dt' \frac{\theta(t' - d)}{\sqrt{t'^2 - d^2/c^2}} e^{i\omega t'} \frac{\theta(t' - d)}{\sqrt{t'^2 - d^2/c^2}}, \quad (55)$$

where we used the $\theta$ function and $d = |r - r_0|$ is the distance from $r$ to the charge at $r_0$. With a change of variables, this last integral can be cast as

$$V(r, t) = \frac{q}{2\pi\varepsilon_0} \int_{0}^{+\infty} dz \frac{e^{-i\omega zd/c}}{\sqrt{z^2 - 1}}. \quad (56)$$

In this form, it becomes simple to identify that the solution is a combination of Bessel functions of the first and second kind, known as the Hankel function of the second kind, $H_0^{(2)}(z)$, defined in Eq. (1.2) in the supplementary material. Using the relevant integral representations of Eq. (1.1), we explicitly obtain

$$V(r, t) = \frac{q}{4\pi\varepsilon_0} iH_0^{(2)} \left( \frac{\omega|r - r_0|}{c} \right). \quad (57)$$

In the static limit, $\omega \to 0$, and using that $H_0^{(2)}(z \to 0) \sim \frac{\pi}{2} \ln z$, we recover precisely the potential of a point charge located at $r_0$ as

$$V(r, t) = \frac{q}{2\pi\varepsilon_0} \ln \left( \frac{|r - r_0|}{a} \right), \quad (58)$$

which shows that Eq. (52) is perfectly compatible with the static case.

Let us now turn to the problem of dipole radiation. We can model an oscillating electric dipole in two dimensions adapting the usual procedure followed in three dimensions. We consider two tiny disks at points $r_+$ and $r_-$ connected by a thin wire that carries a current back and forth from one to the other, such that if one disk has charge $q(t)$ the other has charge $-q(t)$ and the system is always neutral. More precisely, the charge distribution of such a dipole can be written as

$$\sigma_{\text{dip}}(r, t) = qe^{i\omega t} \delta^{(2)}(r - r_+) - qe^{i\omega t} \delta^{(2)}(r - r_-). \quad (59)$$

Formally, it is the superposition of two oscillating monopoles and the associated exact scalar potential $V(r, t)$ can be obtained from the solution Eq. (57)

$$V_{\text{dip}}(r, t) = -\frac{qe^{i\omega t} i}{4\pi\varepsilon_0} \left[ H_0^{(2)} \left( \frac{\omega}{c} |r - r_+| \right) - H_0^{(2)} \left( \frac{\omega}{c} |r - r_-| \right) \right]. \quad (60)$$

Here we are interested in the radiation zone, where $r \gg r_{\pm}$. We can then use $r_{\pm}/r \ll 1$ and Taylor expand the Hankel functions to find

$$V_{\text{dip}}(r, t) \approx -\frac{e^{i\omega t} i}{4\pi\varepsilon_0} \left[ \frac{-\omega}{c} \hat{r} \cdot \mathbf{p} \frac{dH_0^{(2)}(z)}{dz} \right]_{z = \omega r/c} + \cdots \quad (61)$$

where we used without loss of generality that $r_+ = -r_-$ and the dipole moment is $\mathbf{p} = 2qr_+$. As expected, the monopole term cancels and we are left with a dipole term, which contains the part of $V(r, t)$ that will contribute to radiating fields. We now use the asymptotic form of $H_0^{(2)}(z)$ for $z \to \infty$ of Eq. (1.4) to obtain

$$V_{\text{dip}}(r, t) \approx -\frac{\hat{r} \cdot \mathbf{p}}{2\pi\varepsilon_0} \left( \frac{\omega}{c} e^{i\omega(t-r/c)+\pi/4} \sqrt{r} \right) + \cdots \quad (62)$$

This expression is the part of the scalar potential that contributes to dipole radiation since its gradient will generate terms in $E$ that behave as $1/\sqrt{r}$ as required for radiating fields in two spatial dimensions. To make contact with the familiar expressions in three dimensions, one can consider an electric dipole at the origin with $\mathbf{p}$ aligned along the $\hat{y}$ direction. In this case, taking the real part of Eq. (62), one finds

$$V_{\text{dip}}(r, t) \approx -\frac{q_{\text{rad}} \cos(\phi)}{2\pi\varepsilon_0} \sqrt{2\omega\pi} \sin[\omega(t - r/c) + \delta_0] + \cdots \quad (63)$$

where $\phi$ is the angle between $\hat{r}$ and the $y$-axis and $\delta_0$ is a constant phase. In the final result for the potential in the radiation zone, the usual retarded time with respect to the origin intervene. It is interesting to note that, although it is crucial in the intermediate steps to consider the somewhat counterintuitive solutions of the wave equation in $2+1$ dimensions, the behaviour of $V_{\text{dip}}(r, t)$ in the radiation zone is essentially the same found in three dimensions, and the outgoing electromagnetic wave travels with the speed of light.

4. Conclusions

We have discussed the construction of Electrodynamics in two spatial dimensions. We tried to emphasize mainly the somewhat unexpected differences that arise with respect to the usual three-dimensional case. First, we began with a general discussion of Gauss’ law and how the Coulomb field changes with the dimensions of space-time. It is well known that Coulomb’s force, or Newton’s law of gravitation attraction, is expected to behave as the inverse of the distance (to the first power) in two spatial dimensions. This leads to a logarithmic potential which, in itself, is enough to dramatically alter the phenomenology of electromagnetism in $2+1$ dimensions, since the attractive potential between opposite charges becomes confining, i.e., an infinite amount of energy would be required to extract the electron from the hydrogen atom, for example. It is not difficult to obtain the general solution of Poisson’s equation in this context and to consider its multipole expansion. This expansion retains its main
features, in spite of a logarithmic monopole term and the different power counting in \(1/r\) with the dipole term behaving as \(1/r^2\) and so on.

Next, we considered the construction of the full set of Maxwell’s equations. One must face the fact that the magnetic field becomes a scalar field. With a few definitions for the vector calculus in two dimensions, one is able to construct Maxwell’s equations in a form that is similar to the usual one, although there is no equivalent for the magnetic Gauss’ law. With these equations at hand, it becomes straightforward to find the wave equations satisfied by electric and magnetic fields and by the scalar and vector potentials (in Lorenz gauge). Here, an important difference with respect to the three-dimensional case appears. Huygen’s principle is not valid in spaces with even number of dimensions. In practice, this means that part of our intuition regarding the way information propagates is no longer valid since information travels as a superposition of waves with speeds ranging from 0 to \(c\). The retarded potentials in two dimensions have a rather different form, inherited from the Green’s function of the d’Alembertian in 2+1 dimensions which has a Heaviside theta function instead of the usual Dirac delta. We have then shown that this solution, albeit less intuitive, is fully consistent with the static case.

We then turned to an investigation of radiation, using the paradigmic oscillating dipole. We were able to show in Eq. (63) that, in spite of the qualitative difference in the behaviour of the solutions of the wave equation in two and three dimensions, the scalar potential in the radiation zone retains a form very similar to the one we are used to in three dimensions. Therefore, dipole radiation in two dimensions is much more similar to its counterpart in three dimensions than the form of the retarded potentials may suggest. We should state very clearly that we do not know of any physical system, e.g. a material, where electrodynamics would be described by the two-dimensional Maxwell’s equations we discussed here. Rather, our main purpose with this note was to collect a few results on electrodynamics in two dimensions that could be useful for an advanced course on electromagnetism at the undergraduate level. However, in this spirit, other interesting aspects of the phenomenology of Maxwell’s theory in 2+1 dimensions remain to be investigated such as the radiation of a magnetic dipole, to name one. We believe the issues we discussed here are, nevertheless, sufficient to show that the subject is interesting, and can be the basis, for example, for student’s projects. These would certainly broaden their knowledge of electromagnetism and introduce them to some of the common practices of theoretical physics. Our experience confirms this fact, since the results presented here were, in part, obtained as a project by four of us while undergraduate students.

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Supplementary Material

The following online material is available for this article: Appendix

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