Minimally implicit Runge-Kutta methods for Resistive Relativistic MHD

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Abstract. The Relativistic Resistive Magnetohydrodynamic (RRMHD) equations are a hyperbolic system of partial differential equations used to describe the dynamics of relativistic magnetized fluids with a finite conductivity. Close to the ideal magnetohydrodynamic regime, the source term proportional to the conductivity becomes potentially stiff and cannot be handled with standard explicit time integration methods. We propose a new class of methods to deal with the stiffness to the system, which we name Minimally Implicit Runge-Kutta methods. These methods avoid the development of numerical instabilities without increasing the computational costs in comparison with explicit methods, need no iterative extra loop in order to recover the primitive (physical) variables, the analytical inversion of the implicit operator is trivial and the several stages can actually be viewed as stages of explicit Runge-Kutta methods with an effective time-step. We test these methods with two different one-dimensional test beds in varied conductivity regimes, and show that our second-order schemes satisfy the theoretical expectations.

1. Introduction

The Relativistic Resistive Magnetohydrodynamic (RRMHD) equations are a hyperbolic system of partial differential equations used to describe the fluid dynamics in the presence of magnetic fields with a finite (but potentially large) conductivity, and when the velocities involved are close to speed of light, $c$, or the fluid specific energies become comparable or larger than $c^2$. In Astrophysical contexts, all these ingredients can be found in scenarios such as compact stars or relativistic jets.

The RRMHD system of equations can be written as a set of evolution equations corresponding to the conservation of rest-mass, momentum, energy, magnetic flux and electric charge. In addition to these conserved variables, Komissarov [1] has extended the system by two extra scalar potentials which control the evolution of the solenoidal constraint for the magnetic field and the divergence of the electric field following the ideas of Dedner et al. [2].

The physical or primitive variables in RRMHD are

$$ W = (B^i, E^i, q, \rho, \epsilon, v^i), $$

where $B^i (i = 1, 2, 3)$ are the components of the magnetic field, $E^i$ are the components of the electric field, $q$ is the electric charge, $\rho$ is the rest-mass density, $\epsilon$ is the specific internal energy.
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In order to derive the MIRK methods, by construction we impose that they must recover the optimal TVD explicit RK methods for the $S_E$ and $S_Y$ operators. We also consider...
updated and non-updated values of the electric field in the evaluation of the stiff source term in Eq. (3), but only non-updated values of the velocity field and Lorentz factor. The later strategy is key to avoid employing extra iterative loops to recover the primitive variables from the conserved ones (see below). In order to simplify the notation, we define \( \bar{\omega} = \sigma W \) and \( \bar{S}_E = \tilde{S}_E - \sigma W [(v \times B)^J - E^j v^j] \).

The proposed method can be written as an usual explicit, \( s \)-stages RK method with an effective time-step in the following way:

\[
\begin{align*}
Y^{(1)} &= Y|_n, \\
(E^k)^{(1)} &= E^k|_n, \\
Y^{(i)} &= Y|_n + \Delta t \sum_{j=1}^{i-1} \tilde{a}_{ij} S_Y^{(j)}, \\
(E^k)^{(i)} &= E^k|_n + \frac{\Delta t}{1 + \Omega_i} \sum_{j=1}^{i-1} \tilde{a}_{ij} (S_Y^{(j)}) \sum_{j,l=1}^{i-1} a_{ijl} (-\bar{\omega})^l (E^k)^{(j)}, \quad i = 2, \ldots, s + 1, \\
\Omega_i &= \Delta t \sum_{j=1}^{i-1} b_{ij} \bar{\omega}^{(j)}, \\
Y|_{n+1} &= Y^{(s+1)}, \\
E^k|_{n+1} &= (E^k)^{(s+1)},
\end{align*}
\]

where \( |_n \) and \( |_{n+1} \) denote the values of variables at time levels \( t_n \) and \( t_{n+1} \), respectively, \( \tilde{a}_{ij} \) denotes the coefficients of the Butcher matrix associated to the explicit RK method, \( a_{ijl}, b_{ij} \) are the coefficients associated to the implicit part of the method, and \( \Delta t \) is the time-step. Note that, differently from standard RK-IMEX methods, there is no final explicit substep. Indeed, Eqs. (9) and (10), simply mean that the updated values of the variables at \( t_{n+1} \) are the ones computed after the previous \( s \)-substeps of the MIRK algorithm. Equivalently, in the standard notation of Butcher tables (e.g., [7]), the coefficients \( \omega \) and \( \bar{\omega} \) are identically zero.

We point out that the time update of the conserved variables is explicit, i.e., all the terms in the right-hand side of the equations are evaluated in time levels of the algorithm that have already been computed. Singularly, the updated values of the electric field depend upon already known values of it and of the primitive variables. This particularity of the MIRK schemes is very useful for the RRMHD system, since it avoids the need of performing a double iterative procedure to recover the primitive variables from the conserved ones. This is a clear advantage over the usage of RK-IMEX schemes to deal with the stiffness of the RRMHD system (e.g., [7]). It is also relevant to point out that, in order to write the implicit scheme in an explicit form, a much simpler matrix than in [7] has been inverted analytically, since the terms \( E^j v^j \) in the expression for \( S_E^l \) are treated explicitly in MIRK schemes.

### 2.1. First-order method

The first-order MIRK method corresponds to \( s = 1 \) in the previous expressions, with all coefficients zero except for

\[
\tilde{a}_{21} = a_{211} = 1, \quad b_{21} = (1 - c_1),
\]

(1 - \( c_1 \) \( \neq 0 \), and \( c_1 \) being a real coefficient to be determined below. With this set of coefficients, the new method provides bounded values to all the conserved variables, even in the limit \( \bar{\omega} \gg 1 \). After linearization of the non-stiff source terms, we study the matrix which updates the values from one time-step to the next one. Linear stability is satisfied when the absolute value of all the eigenvalues of this matrix are bounded by 1. In particular, linear stability implies that
the absolute value of the determinant of the matrix that updates the values of the conserved variables from one timestep to the next one, $|M_{\Delta t}|$, is also bounded by 1. This necessary (not sufficient) condition restricts the range of the coefficients very efficiently and allows to find values for the coefficient(s) very close to the optimal ones. Under the hypothesis that $|M_{\Delta t}|$ is bounded by 1 when the stiff source terms are neglected, it can be checked that the choice $c_1 = 0$ yields a determinant of $M_{\Delta t}$ also bounded by 1 when the stiff source terms are included, for $\Delta t$ sufficient small. The final expression of the method for this optimal value of the coefficient can be written as:

$$Y|_{n+1} = Y|_n + \Delta t \cdot S \cdot Y|_n,$$

$$E^k|_{n+1} = E^k|_n + \frac{\Delta t}{1 + \Delta t \cdot \bar{\sigma}|_n} \left[ S^k_E - \bar{\sigma} \cdot E^k \right]|_n. \quad (12)$$

We note that the factor

$$\frac{\Delta t}{1 + \Delta t \cdot \bar{\sigma}|_n},$$

can be regarded as an effective time step, that decreases in magnitude as $\bar{\sigma}$ increases. Indeed, in the ideal limit, Eq. (13) yields $E^k \simeq -(v \times B)^k$.

2.2. Second-order method

The second-order MIRK method corresponds to $s = 2$ in the general expression of Sect. 2. Following the same prescription as in the first-order method, i.e., imposing that the method recovers the optimal TVD second-order explicit RK method when stiff source terms are neglected, all coefficients are zero except for

$$\bar{a}_{21} = a_{211} = 1, \quad b_{21} = (1 - c_1), \quad \bar{a}_{31} = a_{321} = 1/2,$$

$$a_{311} = c_1/2, \quad a_{321} = (1 - c_1)/2, \quad a_{312} = 1/2,$$

$$b_{32} = (c_1/2 - c_2),$$

$$(1 - c_1) \neq 0 \neq (c_1/2 - c_2). \quad (14)$$

If the values of the conserved variables $(E^i, Y)|_n$ are bounded at the time step $n$, this method guarantees that the values $(E^i, Y)|_{n+1}$ will also be bounded, even when $\bar{\sigma} \gg 1$.

In order to save RAM memory and improve on computational efficiency, the second-order MIRK scheme can be rewritten as:

$$Y|_{(1)} = Y|_n + \Delta t \cdot S \cdot Y|_n,$$

$$E^k|_{(1)} = E^k|_n + \frac{\Delta t}{1 + \Delta t \cdot (1 - c_1) \cdot \bar{\sigma}|_n} \left[ S^k_E - \bar{\sigma} \cdot E^k \right]|_n, \quad (15)$$

$$E^k|_{n+1} = \frac{1}{2}[E^k|_n + Y|_{(1)} + \Delta t \cdot S \cdot Y|_{(1)}],$$

$$Y|_{n+1} = \frac{1}{2}[E^k|_{(1)} + \left( -1 + \Delta t \cdot \bar{\sigma}(1 - c_1) \right) (E^k|_{(1)} - E^k|_n)] + \frac{\Delta t}{2[1 + \Delta t \cdot \bar{\sigma}(c_1/2 - c_2)]} \left[ S^k_E - \bar{\sigma} \cdot E^k \right]|_{(1)}. \quad (16)$$

Written in this form, the second-order MIRK scheme does not need any extra memory storage as compared with an optimal explicit second-order RK TVD scheme.

After linearization of the non-stiff source terms, we focus again on the condition $|M_{\Delta t}| < 1$. We find two possible sets of values for the coefficients. If we choose $c_2 = \frac{(1-c_1)^2}{2c_1}$, we have to restrict to $c_1 < 0$ or $0.67 < c_1 < 0.75$. We find by numerical experimentation (see Sect. 3.3)
that only the region \( c_1 < 0 \) gives stable results, with the optimal value for \( c_1 \) located around \(-0.1\). If we consider the case \( c_1 = 0 \), we have to restrict to \(-1 < c_2 < -0.5\). After numerical experimentation we find that the optimal value for \( c_2 \) is located around \(-0.97\). Values of both coefficients below zero are favored, and all the options have to be verified numerically.

3. Numerical results
The MIRK methods described in the previous sections have been implemented in the MRGENESIS code [9, 10, 11, 12]. In order to compute the numerical fluxes we employ a Local Lax-Friedrichs approximate solver [1], in which the states left and right to each interface are obtained by a second-order monotone central (MC) reconstruction of primitive variables. We have also tested the new methods in combination with a more diffusive minmod reconstruction, but the results are affected by a larger diffusion close to discontinuities or big gradients (as expected).

In order to test the newly proposed MIRK schemes, we have considered two different one-dimensional tests, each of which proves a different regime of conductivities and, hence, different degrees of stiffness of the RRMHD system. In all the tests considered in the next sections, we assume that the equation of state is that of an ideal gas with a fixed adiabatic index \( \gamma \).

3.1. Self-similar current sheet
Komissarov [1] proposed a test to prove the resistive regime in which \( \sigma \simeq 100 \), i.e., relatively far away from the ideal limit. This is the so-called self-similar current sheet, where we have a magnetic field \( \mathbf{B} = (0, B(x, t), 0) \). The magnetic pressure is much smaller than the gas pressure and \( B(x, 0) \) changes sign within a thin current layer. Except for the magnetic field, the rest of the primitive variables are set to be constant in the whole domain: \( p = 50, \rho = 1, \gamma = 4/3, \mathbf{E} = 0, \mathbf{v} = 0 \) and \( \sigma = 100 \). This test has an analytic (self-similar) solution for \( t > 0 \) of the form

\[
B(x, t) = B_0 \text{erf} \left( \frac{x - x_0}{2 \sqrt{\frac{\sigma}{t}}} \right),
\]

Figure 1. Comparison of the numerical and the analytical solution of the self-similar current sheet. The numerical solution (symbols) has been computed employing \( N_x = 200, \text{CFL}=0.7, \) the second-order MIRK scheme with \( c_1 = -0.1, c_2 = -0.97383794 \) (left panel) and \( c_1 = 0, c_2 = -0.85 \) (right panel). The analytic solution (dashed lines; difficult to see since it is below the symbols) is computed according to Eq. (19). The initial time \( t = 1 \) and the final time \( t = 10 \) are displayed with purple and red colors, respectively. There are no obvious differences in the numerical solutions among the two sets of parameters \((c_1, c_2)\) in this resistive (i.e., low \( \sigma \)) test.

\[
B(x, t) = B_0 \text{erf} \left( \frac{x - x_0}{2 \sqrt{\frac{\sigma}{t}}} \right),
\]
where $B_0 = 1$, $x_0 = 1.5$, the initial time is set to $t = 1$, and the error function is defined in the usual way:

$$\text{erf}(x) = \frac{2}{\sqrt{\pi}} \int_0^x e^{-t^2} \, dt.$$  

We employ 200 computational cells to cover the domain $[0, 3]$. We let the system to evolve until $t = 10$ and compare our results with the analytic solution in Fig. 1. In this case we have used the second order MIRK scheme with $c_1 = -0.1$ and $c_2 = -0.97383794$ (Fig. 1; left) and also with $c_1 = 0$ and $c_2 = -0.85$ (Fig. 1; right). As can be observed, there is basically no difference between the numerical and the analytical solutions. The results do not qualitatively change if the MIRK coefficients are slightly changed with respect to their canonical values. We note that the CFL employed in this test is 0.7.

\[ \text{erf}(x) = \int_0^x e^{-t^2} \, dt. \]

**Figure 2.** Comparison of numerical solutions of the circularly polarized Alfvén waves test for different resolutions, $N_x = 50, 100$ and 200 (in orange, red and black, respectively). The numerical solution has been computed employing the second-order MIRK scheme with $c_1 = 0$, $c_2 = -0.85$ and CFL=0.45 (right panel) and with $c_1 = -0.1$, $c_2 = -0.97383794$ and CFL=0.4 (left panel).

### 3.2. Circularly polarized Alfvén waves

A second one-dimensional test that proves the ideal limit assuming a very large value for the conductivity is considered by Palenzuela et al. [7]. This test was originally proposed by Del Zanna et al. [13] and has an exact solution in ideal Relativistic Magnetohydrodynamics. The solution describes the propagation of large amplitude Alfvén waves along a uniform background magnetic field with strength $B_0$, in a domain with periodic boundary conditions. In the ideal limit, the velocity along the $x-$direction is zero ($v_x = 0$), and the exact solution reads:

\[ (B_y, B_z) = \eta_A B_0 \left( \cos \left[ k(x - v_A t) \right], \sin \left[ k(x - v_A t) \right] \right), \tag{20} \]

\[ (v_y, v_z) = -\frac{v_A}{B_0} (B_y, B_z), \tag{21} \]

where $B_x = B_0$, $k$ is the wave vector, $\eta_A$ is the wave amplitude, and $v_A$ is the relativistic Alfvén speed, that in this case is

\[ v_A^2 = \frac{2B_0^2}{h + B_0^2 (1 + \eta_A^2)} \left\{ 1 + \sqrt{1 - \left[ \frac{2\eta_A B_0^2}{h + B_0^2 (1 + \eta_A^2)} \right]^2} \right\}^{-1}. \tag{22} \]
Though there is no exact solution in the RRMHD regime for circularly polarized Alfvén waves, in the limit of very large conductivity, Eqs. (20) and (21) are approximately satisfied. Thus we compute the evolution from the initial state and until the train of waves completes one period. In the ideal limit, the solution should be identical to that in the initial state and we expect that if \( \sigma \gg 1 \), also the resistive solution shall be quite similar to the initial one. For this test we fix \( \gamma = 2, k = 2\pi, B_0^2 = 4/3, p = \rho = \eta_A = 1 \), which leads to a period \( T = 2 \), and an Alfvén speed \( v_A = 1/2 \).

Figure 2 shows the comparison of the analytic solution versus the numerical solutions with different resolutions. It is evident that our numerical solution lies on top of the analytic one, even for very moderate resolutions. As in the previous test, we have computed the solution with different MIRK coefficients. We find that, in this case, much more demanding from the numerical point of view (because of the large value of the conductivity), the fact that we do have only approximately optimal MIRK coefficients (see Sect. 2) has a clear impact on the stability of the method. While the set of coefficients \( (c_1, c_2) = (-0.1, -0.97383794) \) yields unstable results for CFL values larger than approximately 0.4 (Fig. 2; left), the set of parameters \( (c_1, c_2) = (0, -0.85) \) is still stable for CFL=0.45. In any case, the range of allowed values of the CFL yielding stable results depends on the magnitude of \( \sigma \). From an extensive numerical experimentation we conclude that for \( \sigma \lesssim 5 \times 10^6 \), CFL values up to about 0.7 yield stable results for \( (c_1, c_2) = (0, -0.85) \) and small variations of these parameters. For \( \sigma \gtrsim 5 \times 10^6 \) progressively smaller values of the CFL coefficient must be used to maintain the stability of the numerical solution.

### 3.3. Order of convergence

We have also performed a number of convergence tests employing as reference the one dimensional problems at hand, varying the value of the second-order MIRK coefficients, \( c_1 \) and \( c_2 \). In Fig. 3 we display the \( L_1 \)-norm errors for the \( B_y \) component of the magnetic field as a function of the resolution. The precise definition of the error for a given test including \( N_x \) uniform cells reads:

\[
\text{Error}_{N_x} = \frac{L}{N_x} \sum_{i=1}^{N_x} |B_{y,i}^{(\text{ref})} - B_{y,i}^{(\text{ref})}|,
\]

where \( L \) is the domain size, and \( B_{y,(\text{ref})} \) is a numerical solution computed at sufficiently high resolution to be considered as a reference solution. In our case, \( B_{y,(\text{ref})} \) is computed with \( N_x = 3200 \).

Our second-order MIRK scheme converges to a bit more than second-order for the self-similar current sheet test for a relatively large range of coefficients \( c_1 \) and \( c_2 \) (Fig. 3). The order of convergence is slightly smaller than 2 (\( p = 1.83 \)) for \( c_1 = -0.1 \) and \( c_2 = (1 - c_1)^2/(2c_1) = -6.05 \) and a bit larger than 2 (\( p = 2.19 \)) for \( c_1 = 0 \) and \( c_2 = -0.9 \). Thus, in the limit of moderate conductivity the method performs as expected.

In Fig. 4 we display the \( L_1 \)-norm errors for the \( B_y \) component of the magnetic field as a function of the resolution for the test of circularly polarized Alfvén waves. We observe that at low resolution and for very high conductivity values, \( \sigma = 10^6 \), the second-order schemes have an order of convergence closer to first-order than to second-order. However, as the resolution increases, we find that the order of the method evolves towards the nominal second-order of the algorithm. We also find by numerical experimentation that the best values of the second-order MIRK coefficients lie in the range \(-0.1 \leq c_1 \leq 0 \), and \(-1 < c_2 < -0.85 \). We believe that the difference between the theoretical and numerical order of convergence are triggered by the very large conductivity, which not only increases the stiffness of the system of equations, but also introduces severe losses of accuracy as a result of the very different orders of magnitude among different variables in arithmetic operations (e.g., to compute the electric currents).
Figure 3. Comparison of $L_1$-norm errors (defined in Eq. 23) for the $B^y$ component of the magnetic field in the self similar current sheet test. The different line styles correspond to different choices of the parameters $c_1$ and $c_2$ of the second-order MIRK scheme (see legends). All the numerical tests have been computed with $CFL=0.7$ and $\sigma = 10^2$. As a reference, we add two solid lines showing a theoretical evolution of the error as in a first-order scheme ($p = 1$) and in a second-order one ($p = 2$).

Figure 4. Same as Fig. 3 but for the test of large amplitude circularly polarized Alfvén waves. The different line styles correspond to different choices of the parameters $c_1$ and $c_2$ of the second-order MIRK scheme (see legends). All the numerical tests have been computed with $CFL=0.1$ and $\sigma = 10^6$. For clarity, we do not include the case $c_1 = 0$, $c_2 = -0.97383794$, since it basically lies on top of the case $c_1 = -0.1$, $c_2 = -0.97383794$.

4. Conclusions
We propose here a new kind of RK methods, named minimally implicit RK methods, to numerically evolve some systems of hyperbolic equations with stiff source terms. As a first target for the new MIRK methods, we have considered the numerical evolution of the RRMHD equations as a motivation of the development of the methods. However, we foresee applications to other systems of hyperbolic equations with stiff source terms. At the moment, we are working on applying MIRK methods to coupled system of General Relativistic Radiation Hydrodynamics. In this system, the neutrino transport equations become very stiff in the opaque regime (when
neutrinos are trapped in the matter; see, e.g., [14, 15, 16]).

MIRK methods recover the optimal TVD explicit RK algorithm when the stiff source terms are neglected. They have a very simple analytical inversion of the implicit operators included to deal with the stiff source terms, and their structure can be viewed as an explicit RK method with an effective time-step.

Currently, one dimensional numerical test problems have been considered, but we are working on their application to multidimensional scenarios. Our aim is to improve the current treatment of the resistivity employed in ideal MHD simulations (where the source of such resistivity is not physical but numerical). In this line, we aim to improve the treatment of resistivity in some local numerical simulations of the Kevi-Helmholtz instability in the context of neutron star mergers [17], as well as in global models (e.g., [18]). Likewise, we aim to incorporate physical resistive dissipation to the problem of internal shocks in blazars [19, 20, 21, 22] and in the shocks associated to afterglows [23, 24, 25].

MIRK methods present some advantages in comparison with the well-known RK IMEX methods: CPU time saving, RAM memory saving and accuracy; these properties have to be explored in more detail. We also seek to compare our methods with Strang-splitting ones.

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