Canonical noncommutativity algebra for the tetrad field in general relativity

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Abstract
General relativity under the assumption of noncommuting components of the tetrad field is considered in this paper. Since the algebraic properties of the tetrad field representing the gravitational field are assumed to correspond to the noncommutativity algebra of the coordinates in the canonical case of noncommutative geometry, this idea is closely related to noncommutative geometry as well as to canonical quantization of gravity. According to this presupposition, generalized field equations for general relativity are derived which are obtained by replacing the usual tetrad field by the tetrad field operator within the actions and then building expectation values of the corresponding field equations between coherent states. These coherent states refer to creation and annihilation operators created from the components of the tetrad field operator. In this sense, the obtained theory could be regarded as a kind of semiclassical approximation of a complete quantum description of gravity. The consideration presupposes a special choice of the tensor determining the algebra providing a division of spacetime into two two-dimensional planes.

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1. Introduction
The unification of quantum theory and general relativity is perhaps the most important question in contemporary fundamental physics. One very important approach concerning a quantum theoretical description of general relativity is the canonical quantization of the gravitational field. In particular, the special manifestation given by loop quantum gravity [1, 2] based on the new variables to describe the gravitational field introduced by Ashtekar [3, 4] is considered as a very promising candidate for a quantum description of the gravitational field. Noncommutative geometry represents another very promising concept towards a quantum theory of gravity. Reference [5] shows how to maintain gauge invariance in the context of noncommutative geometry if the usual approach to noncommutative geometry with the star
product is presupposed. Various quantum field theories have been formulated by using the star product [6–31]. Also general relativity and modified gravity theories on noncommutative spacetime [32–74] as well as aspects of cosmology in the context of noncommutative geometry [75–78] have been considered. The relation of noncommutative geometry and quantum gravity has been treated in [79–90].

In this paper, a kind of combination of quantum gravity and noncommutative geometry is suggested. General relativity is explored under the assumption that the components of the gravitational field, or the tetrad field to be more specific, fulfil commutation relations which correspond to the algebra of the canonical case of noncommutative geometry which usually refers to the spacetime coordinates. The relation to canonical quantum gravity consists in the postulated canonical commutation relations for the gravitational field in the sense of a field quantization and the relation to noncommutative geometry consists in the special form of the commutation relations corresponding to the canonical case of noncommutative geometry, since in the approach of this paper, the commutation relations of noncommutative geometry are transferred from the coordinates to the components of the tetrad field. Quantum field theory on noncommutative spacetime in the sense of noncommuting spacetime coordinates is usually treated by using the star product approach which is based on Weyl quantization and maps products of fields depending on noncommuting coordinates to products of fields depending on usual coordinates. Of course, this approach cannot be used with respect to the noncommutativity of the gravitational field considered in this paper, since here the components of a field itself do not commute with each other. This is in contrast to the usual formulation of quantum field theories on noncommutative spacetime, where the components of the coordinates fulfil nontrivial commutation relations. However, there exists another important approach to treat noncommutative geometry which is called the coherent state approach, because there are defined coherent states which refer to creation and annihilation operators obtained from linear combinations of the noncommuting quantities, namely the components of the position vector in the case of usual noncommutative geometry. In this approach, the generalized quantities depending on the noncommuting coordinates are mapped to generalized quantities depending on usual coordinates by building expectation values between such coherent states. The coherent state approach has been developed in [91–93] where an extended expression for plane waves depending on noncommuting coordinates has been calculated. In [94], the coherent state approach has been extended to the case of noncommuting coordinates and momenta and an extended expression for plane waves has also been determined, and using this expression the corresponding generalization of quantum field theory has finally been considered and derived a propagator. Also thermodynamics has been treated within the coherent state approach to noncommutative geometry [95]. In contrast to the star product, the treatment of noncommutative geometry within the coherent state approach can be transferred to the case of noncommuting fields and especially to the case of noncommuting components of the tetrad field as they are treated in this paper.

The structure of this paper is as follows. A short review of the coherent state approach to noncommutative geometry with respect to the usual case of noncommuting coordinates is given at the beginning. Then the noncommutativity algebra of the tetrad field is provided. To enable the application of the coherent state approach to noncommuting components of the tetrad field, it is inevitable to use a special form of the tensor defining the noncommutativity which implies a foliation of spacetime into a couple of two-dimensional submanifolds. Based on this algebra, we define linear combinations of components of the tetrad field which behave like creation and annihilation operators and with respect to them coherent states referring indirectly to the tetrad field can be introduced. According to the considerations mentioned
above, expressions depending on the noncommuting gravitational field will be mapped to expressions depending on the usual commuting gravitational field by building expectation values between coherent states referring to the gravitational field. As a first application of the developed concept, besides the expectation value of the resulting metric operator and the tetrad field operator which is equal to the usual tetrad field itself, the expectation value of the volume element with respect to the coherent states is calculated. Using these concepts, we subsequently calculate the generalized field equation for a matter field coupled to the gravitational field and the generalized Einstein field equation which are obtained by replacing the usual tetrad field by the tetrad field operator obeying the noncommutativity algebra within the corresponding actions, varying the actions with respect to the tetrad field operator and then building the expectation value between coherent states. We perform a calculation to the third order in an expansion of the tetrad field around the Kronecker symbol corresponding to the Minkowski metric in both cases, because this is the lowest order leading to a deformation of the Einstein field equation of the gravitational field. This generalized description of the dynamics of the gravitational field according to general relativity with the noncommutative tetrad field can be considered as a special kind of a semiclassical description of general relativity.

2. Review of the coherent state approach to noncommutative geometry

In the coherent state approach to noncommutative geometry, a canonical algebra between the coordinates is assumed which reads as follows:

\[ [\hat{x}^\mu, \hat{x}^\nu] = i\Theta^{\mu\nu}, \quad \mu, \nu = 1, \ldots, D, \]

where \([\hat{A}, \hat{B}] = \hat{A}\hat{B} - \hat{B}\hat{A}\) and \(\Theta^{\mu\nu}\) is of the following special shape:

\[ \Theta^{\mu\nu} = \text{diag} \left( \Theta_1, \Theta_2, \ldots, \Theta_{D/2} \right), \]

where the \(\Theta_i\) are defined as

\[ \Theta_i = \begin{pmatrix} 0 & \theta_i \\ -\theta_i & 0 \end{pmatrix}, \quad i = 1, \ldots, d = D/2. \]

This means that spacetime is divided into two-dimensional submanifolds by the noncommutativity tensor. If the components of the spacetime coordinate are denoted as \(x^\mu = (x^1, x^2, \ldots, x^{2d-1}, x^{2d})\), then because of (3) it holds for the components of \(\hat{x}^\mu\),

\[ [\hat{x}^{2i-1}, \hat{x}^{2i}] = i\theta_i, \quad i = 1, \ldots, d. \]

Creation and annihilation operators can now be defined according to

\[ \hat{a}_i = \frac{1}{\sqrt{2\theta_i}} (\hat{x}_{2i-1} + i\hat{x}_{2i}) \quad \text{and} \quad \hat{a}_i^\dagger = \frac{1}{\sqrt{2\theta_i}} (\hat{x}_{2i-1} - i\hat{x}_{2i}), \quad i = 1, \ldots, d, \]

which fulfill the algebra

\[ [\hat{a}_i, \hat{a}_j^\dagger] = \delta_{ij}, \quad i, j = 1, \ldots, d, \]

where \(\delta_{ij}\) denotes the Kronecker symbol. Accordingly, we define \(d\) pairs of creation and annihilation operators constructed by the \(d = D/2\) pairs of coordinates referring to the \(d\) two-dimensional submanifolds defined by (5). With respect to (7), coherent states can be defined according to

\[ |a\rangle = \prod_i \exp \left( -\frac{|a_i|^2}{2} \right) \exp (a_i\hat{a}_i^\dagger)|0\rangle = \prod_i \exp \left( -\frac{|a_i|^2}{2} \right) \sum_{n=0}^{\infty} \frac{a_i^n}{\sqrt{n!}} \frac{\hat{a}_i^\dagger n}{\sqrt{n!}} |0\rangle \]

\[ = \prod_i \exp \left( -\frac{|a_i|^2}{2} \right) \sum_{n=0}^{\infty} \frac{a_i^n}{\sqrt{n!}} |n_a\rangle, \]

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where \( a_i \) and \( a_i^* \) describe the corresponding eigenvalues to the operators \( \hat{a}_i \) and \( \hat{a}_i^* \), the \( |n_{\mu}\rangle \) describe the eigenstates of the occupation number operator \( \hat{a}_i^* \hat{a}_i \) and \( |0\rangle \) denotes the vacuum state. For the coherent states (8), eigenvalue equations hold which read

\[
\hat{a}_i |a\rangle = a_i |a\rangle, \quad \langle a|\hat{a}_j^* = \langle a|a_j^*.
\]

By using these coherent states, we can define an expectation value of any function \( f(\hat{x}) \) depending on the noncommuting coordinates according to

\[
\langle f(\hat{x}) \rangle = \langle a|f(\hat{x}_1, \hat{x}_2, \ldots, \hat{x}_{2d-1}, \hat{x}_{2d})|a\rangle = \langle a|f(\hat{a}_1, \hat{a}_1^*, \ldots, \hat{a}_d, \hat{a}_d^*)|a\rangle = F(a_1, a_1^*, \ldots, a_d, a_d^*) = \langle x \rangle,
\]

where \( F(x) \) is a new function depending on the usual coordinates \( x_i \) given by the expectation value of the corresponding operators: \( x_i = \langle a|\hat{x}_i|a\rangle \). This means that it is possible to map arbitrary functions depending on noncommutative coordinates to functions (with additional terms) depending on usual coordinates by using the coherent states and in this sense the coherent states provide an alternative procedure besides the star product to map products of functions on noncommutative spacetime to expressions with additional terms on usual spacetime. The coherent state approach has been developed in [91–93] and extended in [94, 95]. If one refers to the usual case of a (3+1)-dimensional spacetime, one has to choose \( d = 2 \) which means that spacetime is divided into two two-dimensional submanifolds. In the following section, an analogue noncommutativity algebra for the tetrad field will be introduced and after this the coherent state approach presented in this section will be transferred to this case of noncommuting tetrad fields.

3. Noncommutativity algebra for the tetrad field operator

As a basic assumption it is postulated that the components of the tetrad field describing the gravitational field become operators,

\[
e^\mu_m \rightarrow \hat{e}_m^\mu,
\]

which fulfill the following fundamental canonical noncommutativity algebra:

\[
[\hat{e}_m^\mu(x), \hat{e}_n^\nu(y)] = i\Lambda^{\mu\nu} \delta_m^n \delta(x - y),
\]

where \( \Lambda^{\mu\nu} \) denotes the antisymmetric tensor describing the noncommutativity of the tetrad field and \( \delta(x - y) \) denotes the delta function. This kind of quantization of the tetrad field is related to the canonical case of noncommutativity relations between coordinates on the one hand and to the canonical quantization of a local field theory on the other hand, which becomes manifest with respect to the delta function on the left-hand side of the commutation relation or quantization condition (12). In this sense, it represents a combination of noncommutative geometry and a certain kind of quantization of the gravitational field in the sense of a field quantization. The relation between quantum mechanics and noncommutative geometry is analogue to the relation between canonical quantum gravity and the above way of a quantum theoretical treatment of the tetrad field. In noncommutative geometry, the commutation relation between position and momentum, \([\hat{x}_i, \hat{p}_j] = i\delta_{ij}\), is transferred to the several components of position and leads to commutation relations between them: \([\hat{x}^a, \hat{x}^\alpha] = i\theta^{a\alpha}\). In the theory based on algebra (12), the idea of a canonical quantization between a quantity describing the gravitational field (metric, tetrad or connection) and its canonical conjugated momentum, \([\hat{h}_i^a(x), \hat{p}_j^\alpha(y)] = \frac{i}{2} \left( \delta_i^j \delta^a_\alpha + \delta^a_j \delta^\alpha_i \right) \delta(x - y) \) in quantum geometrodynamics for example where \( h_{ab} \) describes the three-metric and \( p_i^a \) its canonical conjugated momentum, is transferred to the several components of the quantity describing the gravitational field, the tetrad field in the
consideration of this paper, leading accordingly to canonical commutation relations between them. In subsequent sections, it will become necessary to perform a series expansion of the gravitational field. To perform a series expansion, the tetrad field operator $\hat{e}_m^\mu$ can be expressed as the sum of the Kronecker symbol $\delta_m^\mu$ and an operator $\hat{h}_m^\mu$ representing a quantum version of a small fluctuation of the gravitational field around flat spacetime,

$$\hat{e}_m^\mu = \delta_m^\mu \mathbf{1} + \hat{h}_m^\mu,$$  \hspace{1cm} (13)

where $\hat{h}_m^\mu$ fulfills the same algebra as $\hat{e}_m^\mu$,

$$\begin{bmatrix} \delta_m^\mu \mathbf{1} + \hat{h}_m^\mu(x), \delta_n^\nu \mathbf{1} + \hat{h}_n^\nu(y) \end{bmatrix} = i\Lambda^{\mu\nu}\delta_{mn}\delta(x-y).$$

This expansion will become important concerning the calculation of the generalized field equations below. In this paper, the usual case of a (3+1)-dimensional Minkowski spacetime is treated. To be able to define coherent states for the tetrad field operator $\hat{e}_m^\mu$, it is of course necessary to divide the spacetime manifold into two two-dimensional submanifolds by defining the following special shape for the noncommutativity tensor $\Lambda^{\mu\nu}$ in (12) according to the consideration of the last section:

$$\Lambda^{\mu\nu} = \begin{pmatrix} 0 & \lambda_a & 0 & 0 \\ -\lambda_a & 0 & 0 & 0 \\ 0 & 0 & 0 & \lambda_b \\ 0 & 0 & -\lambda_b & 0 \end{pmatrix}. $$  \hspace{1cm} (15)

To maintain Lorentz invariance there has to be chosen $\lambda_a = \lambda_b \equiv \lambda$. This condition of Lorentz invariance can only be fulfilled since $\Lambda^{\mu\nu}$ is a tensor rather than a constant matrix. Definition (15) then implies the following commutation relations between the components of the tetrad field being analogue to (5):

$$\begin{bmatrix} \delta_m^\mu \mathbf{1}, \delta_n^\nu \mathbf{1} \end{bmatrix} = i\lambda \delta_{mn}\delta(x-y), \hspace{1cm} \begin{bmatrix} \delta_m^\mu \mathbf{1}, \delta_n^\nu \mathbf{1} \end{bmatrix} = i\lambda \delta_{mn}\delta(x-y).$$

$$\begin{bmatrix} \delta_m^\mu \mathbf{1}, \delta_n^\nu \mathbf{1} \end{bmatrix} = i\lambda \delta_{mn}\delta(x-y), \hspace{1cm} \begin{bmatrix} \delta_m^\mu \mathbf{1}, \delta_n^\nu \mathbf{1} \end{bmatrix} = i\lambda \delta_{mn}\delta(x-y).$$

\hspace{1cm} (16)

4. Definition of coherent states concerning the gravitational field

In this section, we will now define coherent states with respect to operators which are constructed from the components of the tetrad field operator $\hat{e}_m^\mu$ fulfilling the special manifestation of algebra (12) induced by (15) and leading to (16). There can be followed exactly the formalism developed with respect to noncommutative coordinates transferred to the tetrad field for the special case of a (3+1)-dimensional Minkowski spacetime. Accordingly the following operators can be defined:

$$\begin{align*}
\hat{a}_m &= \frac{1}{\sqrt{2\lambda}} (\hat{e}_m^1 + i\hat{e}_m^2), \\
\hat{a}_m^\dagger &= \frac{1}{\sqrt{2\lambda}} (\hat{e}_m^1 - i\hat{e}_m^2), \\
\hat{b}_m &= \frac{1}{\sqrt{2\lambda}} (\hat{e}_m^3 + i\hat{e}_m^4), \\
\hat{b}_m^\dagger &= \frac{1}{\sqrt{2\lambda}} (\hat{e}_m^3 - i\hat{e}_m^4),
\end{align*}$$

fulfilling the commutation relations

$$\begin{bmatrix} \hat{a}_m(x), \hat{a}_n^\dagger(y) \end{bmatrix} = \delta_{mn}\delta(x-y), \hspace{1cm} \begin{bmatrix} \hat{b}_m(x), \hat{b}_n^\dagger(y) \end{bmatrix} = \delta_{mn}\delta(x-y),$$

and thus behave as creation and annihilation operators. Since in the special case of $d = D/2 = 2$ only two pairs of creation and annihilation operators arise, they are denoted
with $a_m, a_m^+$ and $b_n, b_n^+$ instead of distinguishing them by an index. The components of the tetrad field operator $\hat{e}_m$ can be expressed by operators (17) as follows:

\[
\hat{e}_m^1 = \sqrt{\frac{\lambda}{2}} (\hat{a}_m + \hat{a}_m^+), \quad \hat{e}_m^2 = -i \sqrt{\frac{\lambda}{2}} (\hat{a}_m - \hat{a}_m^+), \\
\hat{e}_m^3 = \sqrt{\frac{\lambda}{2}} (\hat{b}_m + \hat{b}_m^+), \quad \hat{e}_m^4 = -i \sqrt{\frac{\lambda}{2}} (\hat{b}_m - \hat{b}_m^+). \tag{19}
\]

To simplify the notation, the quantities $\hat{E}_m^\mu$ and $\hat{E}_m^{\mu\dagger}$ are defined according to

\[
\hat{E}_m^\mu = \sqrt{\frac{\lambda}{2}} (\hat{a}_m - i \hat{a}_m^+, \hat{b}_m, -i \hat{b}_m), \quad \hat{E}_m^{\mu\dagger} = \sqrt{\frac{\lambda}{2}} (\hat{a}_m^+, i \hat{a}_m, \hat{b}_m^+, i \hat{b}_m^+). \tag{20}
\]

which fulfill the commutation relation

\[
[\hat{E}_m^\mu (x), \hat{E}_n^{\mu\dagger} (y)] = i \Gamma^{\mu\nu} \delta_{mn} \delta(x - y), \tag{21}
\]

where the tensor $\Gamma^{\mu\nu}$ has been introduced which is defined as

\[
\Gamma^{\mu\nu} = \frac{\lambda}{2} \begin{pmatrix}
-1 & 0 & 0 \\
0 & -i & 0 \\
0 & 0 & i \end{pmatrix}. \tag{22}
\]

The tetrad field operator $\hat{e}_m^\mu$ can now be expressed by using definition (20),

\[
\hat{e}_m^\mu = E_m^\mu + \hat{E}_m^{\mu\dagger}. \tag{23}
\]

Accordingly, the usual tetrad field is related to the field $E_m^\mu$ and its complex conjugated field $E_m^{\mu\ast}$ according to

\[
e_m^\mu = E_m^\mu + E_m^{\mu\ast}. \tag{24}
\]

If there are defined coherent states within the Hilbert spaces $\mathcal{H}_a$ and $\mathcal{H}_b$ in which the operators $\hat{a}_m, \hat{a}_m^+$ respectively $\hat{b}_m, \hat{b}_m^+$ act and which represent eigenstates of $\hat{a}_m$ and $\hat{b}_m$ as ket vectors and eigenstates of $\hat{a}_m^+$ and $\hat{b}_m^+$ as bra vectors, these states of course also represent eigenstates of the operator $\hat{E}_m^\mu$ and its Hermitian adjoint $\hat{E}_m^{\mu\dagger}$. The coherent states which depend of course on the spacetime point, since they refer to field operators, are defined according to

\[
|a_m(x)\rangle = \exp \left[ -\frac{|a_m(x)|^2}{2} \right] \exp \left[ a_m(x) \hat{a}_m^\dagger (x) \right] |0\rangle = \exp \left[ -\frac{|a_m(x)|^2}{2} \right] \sum_{n=0}^\infty \frac{|a_m(x)|^n}{\sqrt{n!}} |n_{a_m}(x)\rangle, \\
|b_m(x)\rangle = \exp \left[ -\frac{|b_m(x)|^2}{2} \right] \exp \left[ b_m(x) \hat{b}_m^\dagger (x) \right] |0\rangle = \exp \left[ -\frac{|b_m(x)|^2}{2} \right] \sum_{n=0}^\infty \frac{|b_m(x)|^n}{\sqrt{n!}} |n_{b_m}(x)\rangle, \tag{25}
\]

where $|n_{a_m}(x)\rangle = \sqrt{\frac{\lambda}{2^n n!}} |0\rangle$ and $|n_{b_m}(x)\rangle = \sqrt{\frac{\lambda}{2^n n!}} |0\rangle$ are the eigenstates with respect to the occupation number operators $\hat{a}_m^\dagger (x) \hat{a}_m^\dagger (x)$ and $\hat{b}_m^\dagger (x) \hat{b}_m^\dagger (x)$ meaning that

\[
\hat{a}_m^\dagger (x) \hat{a}_m^\dagger (x) |n_{a_m}(x)\rangle = n_{a_m}(x) |n_{a_m}(x)\rangle, \quad \hat{b}_m^\dagger (x) \hat{b}_m^\dagger (x) |n_{b_m}(x)\rangle = n_{b_m}(x) |n_{b_m}(x)\rangle, \tag{26}
\]

where $n_{a_m}(x)$ and $n_{b_m}(x)$ describe the corresponding eigenvalues of the occupation number operators. The coherent states (25) fulfill the following eigenvalue equations:

\[
\langle a_m(x) | \hat{a}_m(x) \rangle = a_m(x) |a_m(x)\rangle, \quad \langle b_m(x) | \hat{b}_m(x) \rangle = b_m(x) |b_m(x)\rangle, \\
\langle a_m(x) | \hat{a}_m^\dagger (x) \rangle = \langle a_m(x) | a_m^\dagger (x) \rangle, \quad \langle b_m(x) | \hat{b}_m^\dagger (x) \rangle = \langle b_m(x) | b_m^\dagger (x) \rangle, \tag{27}
\]

where $\langle \cdot | \cdot \rangle$ denotes the matrix element of the operators $\cdot$. The coherent states (25) fulfill the following eigenvalue equations:

\[
\langle a_m(x) | \hat{a}_m(x) \rangle = a_m(x) |a_m(x)\rangle, \quad \langle b_m(x) | \hat{b}_m(x) \rangle = b_m(x) |b_m(x)\rangle, \\
\langle a_m(x) | \hat{a}_m^\dagger (x) \rangle = \langle a_m(x) | a_m^\dagger (x) \rangle, \quad \langle b_m(x) | \hat{b}_m^\dagger (x) \rangle = \langle b_m(x) | b_m^\dagger (x) \rangle, \tag{27}
\]

where $\langle \cdot | \cdot \rangle$ denotes the matrix element of the operators $\cdot$. The coherent states (25) fulfill the following eigenvalue equations:

\[
\langle a_m(x) | \hat{a}_m(x) \rangle = a_m(x) |a_m(x)\rangle, \quad \langle b_m(x) | \hat{b}_m(x) \rangle = b_m(x) |b_m(x)\rangle, \\
\langle a_m(x) | \hat{a}_m^\dagger (x) \rangle = \langle a_m(x) | a_m^\dagger (x) \rangle, \quad \langle b_m(x) | \hat{b}_m^\dagger (x) \rangle = \langle b_m(x) | b_m^\dagger (x) \rangle, \tag{27}
\]
where \( a, a^*, b, b^* \) are the corresponding eigenvalues to the operators \( \hat{a}, \hat{a}^*, \hat{b}, \hat{b}^* \). By building the tensor product of the several Minkowski components of the eigenstates \( |a_m \rangle \) and \( |b_m \rangle \), respectively,

\[
|A\rangle = \prod_{m=1}^{4} |a_m\rangle,
|B\rangle = \prod_{m=1}^{4} |b_m\rangle,
\]

and then building the tensor product of \(|A\rangle\) and \(|B\rangle\),

\[
|\Psi\rangle = |A\rangle \otimes |B\rangle = \prod_{m=1}^{4} |a_m\rangle \otimes \prod_{n=1}^{4} |b_n\rangle,
\]

one obtains a state \(|\Psi\rangle\) being an eigenstate of all components of \( \hat{E}^\mu_m \) meaning that

\[
\hat{E}^\mu_m(x)|\Psi(x)\rangle = \sqrt{\frac{\gamma}{2}} [\hat{d}_m(x), -i\hat{a}_m(x), \hat{b}_m(x), -i\hat{b}_m(x)] \prod_{n=1}^{4} |a_n(x)\rangle \otimes |b_n(x)\rangle
\]

\[
= \sqrt{\frac{\gamma}{2}} [\hat{d}_m(x), -i\hat{a}_m(x), b_m(x), -i\hat{b}_m(x)] \prod_{n=1}^{4} |a_n(x)\rangle \otimes |b_n(x)\rangle
\]

\[
= E^\mu_m(x)|\Psi(x)\rangle
\]

\[
\langle \Psi(x)|\hat{E}^\mu_m(x)\rangle = \prod_{n=1}^{4} \langle a_n(x)| \otimes \langle b_n(x)| \sqrt{\frac{\gamma}{2}} [\hat{a}_m^*(x), i\hat{a}_m(x), \hat{b}_m(x), i\hat{b}_m(x)]
\]

\[
= \prod_{n=1}^{4} \langle a_n(x)| \otimes \langle b_n(x)| \sqrt{\frac{\gamma}{2}} [\hat{a}_m^*(x), i\hat{a}_m(x), b_m^*(x), i\hat{b}_m(x)]
\]

\[
= \langle \Psi(x)|E^\mu_m(x)\rangle
\]

(30)

This means that an expectation value can be defined for the tetrad field operator \( \hat{e}^\mu_m \) and for all quantities depending on the tetrad field operator \( \hat{e}^\mu_m \) by expressing the tetrad field operator \( \hat{e}^\mu_m \) through \( \hat{E}^\mu_m \) and its Hermitian adjoint \( \hat{E}^\mu_m^\dagger \) and then building expectation values with respect to \(|\Psi\rangle\). Of course, \(|\Psi\rangle\) has to obey the normalization condition \( \langle \Psi|\Psi\rangle = 1 \). According to (13) and (14), we can analogously introduce the operator \( \hat{H}^\mu_m \) being related to \( \hat{h}^\mu_m \) as \( \hat{E}^\mu_m \) is related to \( \hat{e}^\mu_m \) according to (23),

\[
\hat{e}^\mu_m = \hat{E}^\mu_m + \hat{H}^\mu_m,
\]

fulfilling the same commutation relation as \( \hat{E}^\mu_m \) given in (21),

\[
[\hat{H}^\mu_m(x), \hat{H}^\mu_m(y)] = i\Gamma^\mu_\nu \delta_{mn} \delta(x - y).
\]

The corresponding eigenstates \(|\Phi\rangle\) which correspond to the eigenstates \(|\Psi\rangle\) of \( \hat{E}^\mu_m \) can be defined for which the relations hold being analogue to (30),

\[
\hat{H}^\mu_m|\Phi\rangle = H^\mu_m|\Phi\rangle,
\]

\[
\langle \Phi|\hat{H}^\mu_m = \langle \Phi|\hat{H}^\mu_m^\dagger.
\]

(33)

The eigenstates \(|\Phi\rangle\) of course have also to obey the normalization condition \( \langle \Phi|\Phi\rangle = 1 \).

5. Construction of expectation values with respect to coherent states

An expectation value of any function of the tetrad field operator \( f(\hat{e}^\mu_m) \) can be defined with respect to a coherent state analogue to (10) by expressing the tetrad field operator \( \hat{e}^\mu_m \) through the field operator \( \hat{E}^\mu_m \) and its Hermitian adjoint operator \( \hat{E}^\mu_m^\dagger \),

\[
\langle f(\hat{e}^\mu_m) \rangle = \langle \Psi|f(\hat{e}^\mu_m)|\Psi\rangle = \langle \Psi|f(\hat{E}^\mu_m, \hat{E}^\mu_m^\dagger)|\Psi\rangle = F(E^\mu_m, E^\mu_m^\dagger) = F(\hat{e}^\mu_m).
\]

(34)
where $F(e^e_m)$ describes a new function depending on the usual tetrad field and generally containing additional terms. If the function $f(\hat{e}^e_m) = f(\hat{E}^e_m, \hat{E}^{e\dagger}_m)$ depends linearly on the tetrad field operator $\hat{e}^e_m$, then the operators $\hat{E}^e_m$ and $\hat{E}^{e\dagger}_m$ can directly be applied to $|\Psi\rangle$ and the expectation value of the function depending on the tetrad field operator $\hat{e}^e_m$ yields the function depending on the usual tetrad field $e^e_m$ as will be shown below using the example of the expectation value of the single tetrad field operator $\hat{e}^e_m$ yielding the usual tetrad field $e^e_m$ itself. If $f(\hat{e}^e_m) = f(\hat{E}^e_m, \hat{E}^{e\dagger}_m)$ contains products of the tetrad field operator $\hat{e}^e_m$ and thus products of the operators $\hat{E}^e_m$ and $\hat{E}^{e\dagger}_m$, then the products have to be permuted in such a way that all factors $\hat{E}^{e\dagger}_m$ are to the left of the factors $\hat{E}^e_m$. This can of course be performed by using the commutation relation (21) being a direct consequence of (12). Since there are usually just considered pointwise products of fields, the delta function within the commutator does not appear explicitly within the calculations. Then the eigenvalue equations (30) as well as equation (24) can be used to re-express the obtained expression in terms of the eigenvalues $E^e_m$ and $E^{e\dagger}_m$ belonging to the operator $\hat{E}^e_m$ and its Hermitian adjoint $\hat{E}^{e\dagger}_m$ respectively by the usual tetrad field $e^e_m$. In this way, the explicit expression of the expectation value (34) can be calculated which yields to every function depending on the tetrad field operator a generalized function which depends on the usual tetrad field again. The idea is now to obtain a generalization of the dynamics of the gravitational field according to general relativity by using expectation values of the corresponding expressions describing the dynamics of general relativity and depending on the tetrad field operator fulfilling the noncommutativity algebra (12). This means that the extended description of general relativity according to the approach of this paper is obtained by two steps: first within the usual quantities depending on the tetrad field, the usual tetrad field $e^e_m$ is replaced by the tetrad field operator $\hat{e}^e_m$ defined according to (12),

$$f(e^e_m) \rightarrow f(\hat{e}^e_m),$$

(35)

and after this the expectation value is built between coherent states according to (34). Before approaching the generalized description of the dynamics of general relativity as will be done in the next two sections, first some examples as a kind of preparation shall be given in this section.

As already mentioned above, the expectation value for the tetrad field itself corresponds to the usual tetrad field $e^e_m$,

$$\langle e^e_m \rangle = \langle \Psi | \hat{e}^e_m | \Psi \rangle = \langle \Psi | (\hat{E}^e_m + \hat{E}^{e\dagger}_m) | \Psi \rangle = \langle \Psi | (E^e_m + E^{e\dagger}_m) | \Psi \rangle = E^e_m + E^{e\dagger}_m = e^e_m.$$

(36)

In (36), (23), (24) and (30) have been used and of course the normalization condition $\langle \Psi | \Psi \rangle = 1$. If we consider the expectation value of the metric field operator $\hat{\gamma}^{\mu\nu}$ being related to the tetrad field operator $\hat{e}^e_m$ according to $\hat{\gamma}^{\mu\nu} = \hat{e}^e_m \hat{e}^{e\dagger}_n \eta^{mn}$, where $\eta^{mn} = \text{diag} (1, -1, -1, -1)$ denotes the Minkowski metric, the situation becomes already a little bit more complicated. After replacing the tetrad field operator $\hat{e}^e_m$ by using equation (23), one permutation has to be performed and accordingly one obtains an additional constant term,

$$\langle \hat{\gamma}^{\mu\nu} \rangle = \langle \Psi | \hat{\gamma}^{\mu\nu} | \Psi \rangle = \langle \Psi | \eta^{mn} e^m_n e^{e\dagger}_n | \Psi \rangle = \langle \Psi | \eta^{mn} (\hat{E}^e_m + \hat{E}^{e\dagger}_m) (\hat{E}^e_n + \hat{E}^{e\dagger}_n) | \Psi \rangle = \langle \Psi | \eta^{mn} (\hat{E}^e_m \hat{E}^e_n + \hat{E}^{e\dagger}_m \hat{E}^{e\dagger}_n + \hat{E}^e_m \hat{E}^{e\dagger}_n + \hat{E}^{e\dagger}_m \hat{E}^e_n) | \Psi \rangle = \langle \Psi | \eta^{mn} (\hat{E}^e_m \hat{E}^e_n + \hat{E}^{e\dagger}_m \hat{E}^{e\dagger}_n + \hat{E}^e_m \hat{E}^{e\dagger}_n + \hat{E}^{e\dagger}_m \hat{E}^e_n + \hat{E}^e_m \hat{E}^{e\dagger}_n + \hat{E}^{e\dagger}_m \hat{E}^e_n + i\Gamma^{\mu\nu} \delta_{mn}) | \Psi \rangle = \langle \Psi | \eta^{mn} (\hat{E}^e_m \hat{E}^e_n + \hat{E}^{e\dagger}_m \hat{E}^{e\dagger}_n + \hat{E}^e_m \hat{E}^{e\dagger}_n + \hat{E}^{e\dagger}_m \hat{E}^e_n + i\Gamma^{\mu\nu} \delta_{mn}) | \Psi \rangle = \langle \Psi | \eta^{mn} e^e_m e^{e\dagger}_n | \Psi \rangle + \langle \Psi | i\Gamma^{\mu\nu} \eta^{mn} \delta_{mn} | \Psi \rangle = \hat{\gamma}^{\mu\nu} - 2i\Gamma^{\mu\nu}.$$

(37)

As a more complicated example, this procedure of obtaining generalized functions depending on the usual tetrad field by building of an expectation value of the usual function converted
to a function depending on the tetrad field operator shall be applied to the volume element. Replacing the tetrad field $e_m^i$ by the tetrad field operator $\hat{e}_m^i$, building the expectation value and expressing the tetrad field operator $\hat{e}_m^i$ by the operator $\hat{E}_m^i$ lead to the following expression:

$$\langle \mathcal{V} \rangle = \int d^4x \langle \Psi | \det (\hat{e}_m^i) | \Psi \rangle = \int d^4x \langle \Psi | e^{abcd} \epsilon_{\mu\nu\rho\sigma} \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma | \Psi \rangle$$

$$= \int d^4x \langle \Psi | e^{abcd} \epsilon_{\mu\nu\rho\sigma} (\hat{E}_a^\mu + \hat{E}_b^\mu + \hat{E}_c^\mu + \hat{E}_d^\mu) (\hat{E}_b^\nu + \hat{E}_c^\nu + \hat{E}_d^\nu) (\hat{E}_c^\rho + \hat{E}_d^\rho) | \Psi \rangle$$

$$= \int d^4x \langle \Psi | e^{abcd} \epsilon_{\mu\nu\rho\sigma} (\hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma)$$

$$+ \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma)$$

$$\times \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma)$$

$$\times \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma) \rangle | \Psi \rangle,$$

where the determinant of the tetrad field operator can be expressed as follows: $\det (\hat{e}_m^i) = e^{abcd} \epsilon_{\mu\nu\rho\sigma} \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma$, where $\epsilon_{\mu\nu\rho\sigma}$ denotes the total antisymmetric tensor in four dimensions.

Permutation of the operator $\hat{E}_m^i$ and the operator $\hat{E}_m^i$ by using (21) leads to

$$\langle \mathcal{V} \rangle = \int d^4x \langle \Psi | e^{abcd} \epsilon_{\mu\nu\rho\sigma} (\hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma)$$

$$+ i e^{abcd} \epsilon_{\mu\nu\rho\sigma} (\hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma)$$

$$+ i e^{abcd} \epsilon_{\mu\nu\rho\sigma} (\hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma)$$

$$+ i e^{abcd} \epsilon_{\mu\nu\rho\sigma} (\hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma)$$

$$\times \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma) \rangle | \Psi \rangle,$$

and applying the operators to the coherent states by using (30) and then re-expressing the obtained expression depending on the eigenvalues $E_m^\mu$ and $E_m^\mu$ by the tetrad field $e_m^i$ by using (24) finally lead to

$$\langle \mathcal{V} \rangle = \int d^4x \{ \langle \Psi | e^{abcd} \epsilon_{\mu\nu\rho\sigma} (\hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma)$$

$$+ i e^{abcd} \epsilon_{\mu\nu\rho\sigma} (\hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma)$$

$$+ i e^{abcd} \epsilon_{\mu\nu\rho\sigma} (\hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma)$$

$$\times \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma + \hat{E}_a^\mu \hat{E}_b^\nu \hat{E}_c^\rho \hat{E}_d^\sigma) \} \} | \Psi \rangle.$$

(40)
Thus, the expectation value of the volume element operator contains many additional terms which now also depend on the tetrad field $e^\mu_m$. In the next two sections, this procedure to obtain generalized classical expressions for the generalized quantities of general relativity formulated in terms of the tetrad field operator $\hat{e}^\mu_m$ will be used to obtain generalized dynamics of general relativity.

6. Generalized field equation for matter coupled to gravity

The effective field equations of matter fields as well as the gravitational field respectively in the extended description of general relativity with components of the tetrad field which do not commute are according to the considerations of the last section obtained by replacing the usual tetrad field $e^\mu_m$ by the tetrad field operator $\hat{e}^\mu_m$ within the corresponding action, then varying the action with respect to the tetrad field operator and finally building the expectation value of the resulting field equation containing the tetrad field operator. Varying the action with the usual tetrad field replaced by the tetrad field operator yields of course the usual field equation with the tetrad field replaced by the tetrad field operator. The expectation value has then to be built according to (34) and yields the generalized field equation. However, building the expectation value of the action with the usual tetrad field replaced by tetrad field operator and then varying the resulting action which depends again on the tetrad field would lead to other field equations which are not considered in this paper. In this section, we will consider the generalized matter field equation of a fermionic field, whereas in the next section we will consider the generalized Einstein field equation describing the dynamics of the gravitational field itself which are both obtained by performing the procedure mentioned above. The action of a fermionic matter field on curved spacetime which is thus coupled to the gravitational field reads

$$ S_M[e] = \int d^4x \det \left[ e^\mu_m \right] \bar{\psi} \gamma^m e^\mu_m \left( \partial_\mu + \frac{i}{2} \omega^{ab}_m [e] \Sigma_{ab} \right) \psi, \quad (41) $$

where the $\gamma^m$ denote the Dirac matrices, $\bar{\psi} = \psi^\dagger \gamma^0$ and the $\Sigma_{ab} = -\frac{i}{2} \left[ \gamma_a, \gamma_b \right]$ denote the generators of the Lorentz group fulfilling $[\Sigma_{ab}, \Sigma_{cd}] = \eta_{bc} \Sigma_{ad} - \eta_{bd} \Sigma_{ac} - \eta_{ad} \Sigma_{cb}$. The spin connection $\omega^{ab}_m [e]$ depends on the tetrad field $e^\mu_m$ in the following way:

$$ \omega^{ab}_m [e] = 2e^{[a} \partial_{[b] e^{b]}_m} + e^{[a} e^{b]} \partial_{[b] e^{b]}_m. \quad (42) $$

Remark that the brackets denote antisymmetrization meaning that $[ab] = ab - ba$. Replacing the usual tetrad field by the tetrad field operator obeying (12), $e^\mu_m \rightarrow \hat{e}^\mu_m$, within (41) leads to the following action:

$$ S_M[e] \rightarrow \hat{S}_M[\hat{e}] = \int d^4x \det \left[ \hat{e}^\mu_m \right] \bar{\psi} \gamma^m \hat{e}^\mu_m \left( \partial_\mu + \frac{i}{2} \hat{\omega}^{ab}_m [\hat{e}] \Sigma_{ab} \right) \psi. \quad (43) $$

The corresponding field equation containing the tetrad field operator $\hat{e}^\mu_m$ is obtained by varying the resulting action with respect to the matter field. Building of the expectation value by using coherent states (29) leads then to the generalized Dirac equation on curved spacetime containing the usual tetrad field $e^\mu_m$:

$$ \langle \Psi | \frac{\delta \hat{S}_M[\hat{e}]}{\delta \bar{\psi}} | 0 \rangle \quad \leftrightarrow \quad \langle \Psi | \frac{1}{\det \left[ \hat{e}^\mu_m \right]} \frac{\delta \hat{S}_M[\hat{e}]}{\delta \bar{\psi}} | \Psi \rangle = 0. \quad (44) $$

After concrete variation of the matter action on curved spacetime depending on the tetrad field operator $\hat{e}^\mu_m$ obtained in (43) with respect to $\bar{\psi}$, (44) reads

$$ \langle \Psi | i \gamma^m \hat{e}^\mu_m \left( \partial_\mu + \frac{i}{2} \hat{\omega}^{ab}_m [\hat{e}] \Sigma_{ab} \right) \psi | \Psi \rangle = 0, \quad (45) $$
and inserting the explicit term of the spin connection (42) transformed to the corresponding expression depending on the tetrad field operator \( \hat{e}_m^\mu \) leads to

\[
\langle \Psi \mid i\gamma^m \hat{D}_m(h^0) + \hat{D}_m(h^1) + \hat{D}_m(h^2) + \hat{D}_m(h^3) \rangle \psi \mid \Phi \rangle + \mathcal{O}(\hbar^4) = 0. \tag{46}
\]

To be able to treat the calculation, the exact expression of (46) will not be considered, but a series expansion of the tetrad field operator \( \hat{e}_m^\mu \) around \( \delta_m^\mu \) instead will be considered as has been introduced in (13) and (14). As usual, such a series expansion makes sense if the perturbation \( h_m^\mu \) of the classical expansion, \( e_m^\mu = \delta_m^\mu + h_m^\mu \), becomes an operator after postulating (12) is assumed to be very small. After the transition, this relation looks as described by (13), \( \hat{e}_m^\mu = \delta_m^\mu \mathbb{1} + \hat{h}_m^\mu \), where \( \hat{h}_m^\mu \) fulfills according to (14) the same algebra as \( \hat{e}_m^\mu \). Concerning the further calculation, we will refer to the operators and states defined with respect to the expansion \( h_m^\mu \) in (31)–(33) and (29) which are mathematically of course isomorphic to the operators and states defined with respect to \( e_m^\mu \). In particular, we will consider a calculation to the third order in the expansion operator \( h_m^\mu \), since in the case of the generalized Einstein field equation considered in the next section, the terms of the first and second order do not differ from the usual case which means that they yield no additional terms. Accordingly, the expectation value of the equation can be expressed as follows:

\[
\langle \Phi \mid i\gamma^m \hat{D}_m(h^0) + \hat{D}_m(h^1) + \hat{D}_m(h^2) + \hat{D}_m(h^3) \rangle \psi \mid \Phi \rangle = 0, \tag{47}
\]

where \( \hat{D}_m(h^0) \) describes the term of the covariant derivative which does not depend on the perturbation of the tetrad field operator \( h_m^\mu \) and \( \hat{D}_m(h^1), \hat{D}_m(h^2) \) and \( \hat{D}_m(h^3) \) describe the terms of the covariant derivative to the first, second and third order in \( h_m^\mu \). Accordingly \( \hat{D}_m(h^0), \hat{D}_m(h^1), \hat{D}_m(h^2) \) and \( \hat{D}_m(h^3) \) are defined as

\[
\begin{align*}
\hat{D}_m(h^0) &= \hat{e}_m^\mu = \partial_m, \\
\hat{D}_m(h^1) &= \hat{h}_m^\mu \partial_m + \frac{1}{2} \left[ 2 \delta^\mu[a \partial_m \hat{e}_c^a] + \delta^b \hat{h}_m^b \right] \Sigma_{ab}, \\
\hat{D}_m(h^2) &= \frac{1}{2} \left[ 2 h_m^\mu \partial_m \hat{e}_c^a + \hat{h}_m^\mu \delta^{ac} \partial_m \hat{e}_b^c + \delta_m^\mu \hat{h}_m^a \partial_a \hat{e}_c^b + \delta_m^\mu \hat{h}_m^b \partial_b \hat{e}_c^a \right] \Sigma_{abc}, \\
\hat{D}_m(h^3) &= \frac{1}{2} \left[ 2 h_m^\mu \partial_m \hat{e}_c^a + \hat{h}_m^\mu \delta^{ac} \partial_m \hat{e}_b^c + \delta_m^\mu \hat{h}_m^a \partial_a \hat{e}_c^b + \delta_m^\mu \hat{h}_m^b \partial_b \hat{e}_c^a \right] \Sigma_{abc}. 
\end{align*}
\tag{48}
\]

The expectation value of the term \( i\gamma^m \hat{D}_m(h^0) \psi \) can of course be given directly without a long calculation, since it does not contain the operator \( h_m^\mu \).

\[
\begin{align*}
\langle \Phi \mid i\gamma^m \hat{D}_m(h^0) \psi \mid \Phi \rangle &= \langle \Phi \mid i\gamma^m \partial_m \psi \mid \Phi \rangle = i\gamma^m \partial_m \psi. \tag{49}
\end{align*}
\]

To calculate the expectation values of the other terms, it is necessary to treat the commutator of derivatives of \( \hat{E}_m^\mu \) or \( \hat{H}_m^\mu \) respectively, since these terms contain such derivatives. These commutators can be calculated as follows:

\[
\begin{align*}
[\partial_\mu \hat{E}_m^\nu(x), \partial_\alpha \hat{E}_m^{\nu'}(y)] &= \frac{\partial}{\partial x^\mu} \frac{\partial}{\partial y^\alpha} \left[ \hat{E}_m^\nu(x), \hat{E}_m^{\nu'}(y) \right] \\
&= \frac{\partial}{\partial x^\mu} \frac{\partial}{\partial y^\alpha} \left[ i \Gamma^{\nu \mu} \delta_m^{\nu'} \delta(x - y) \right] \\
&= i \Gamma^{\nu \mu} \delta_m^{\nu'} \frac{\partial}{\partial x^\mu} \frac{\partial}{\partial y^\alpha} \delta(x - y)
\end{align*}
\]
operators \hat{E}^\nu (x) \delta(x-y) \frac{\partial}{\partial y^\nu}, \quad (50)

where in the second step of (50), (21) has been used and in the fourth and the fifth step, a special property of the delta function has been used:

\[ \int d^4x \ f(x) \hat{\partial}_\mu \delta(x-a) = - \int d^4x \ \delta(x-a) \partial_\mu f(x). \quad (51) \]

From (50), one can easily see that the following commutation relations are valid as well:

\[ \left[ \partial_\mu \hat{E}^\nu_m (x), \hat{E}^\nu_n (y) \right] = - i \Gamma^{\mu \nu \rho} \delta_{mn} \delta(x-y) \frac{\partial}{\partial x^\rho}, \]
\[ \left[ \hat{E}^\mu_m (x), \partial_\nu \hat{E}^\nu_n (y) \right] = - i \Gamma^{\mu \nu \rho} \delta_{mn} \delta(x-y) \frac{\partial}{\partial y^\rho}, \]
\[ \left[ \partial_\rho \partial_\nu \hat{E}^\mu_m (x), \hat{E}^\nu_n (y) \right] = i \Gamma^{\mu \nu \rho} \delta_{mn} \delta(x-y) \frac{\partial}{\partial x^\rho} \frac{\partial}{\partial x^\nu}, \]
\[ \left[ \hat{E}^\mu_m (x), \partial_\rho \partial_\nu \hat{E}^\nu_n (y) \right] = i \Gamma^{\mu \nu \rho} \delta_{mn} \delta(x-y) \frac{\partial}{\partial y^\rho} \frac{\partial}{\partial y^\nu}. \quad (52) \]

Of course this implies that accordingly also the corresponding commutation relations containing derivatives with respect to \( \hat{H}^\mu_m \) are valid,

\[ \left[ \partial_\mu \hat{H}^\mu_m (x), \partial_\nu \hat{H}^\nu_n (y) \right] = i \Gamma^{\mu \nu \rho} \delta_{mn} \delta(x-y) \frac{\partial}{\partial x^\rho}, \]
\[ \left[ \partial_\mu \hat{H}^\mu_m (x), \hat{H}^\nu_n (y) \right] = - i \Gamma^{\mu \nu \rho} \delta_{mn} \delta(x-y) \frac{\partial}{\partial x^\rho}, \]
\[ \left[ \hat{H}^\mu_m (x), \partial_\rho \partial_\nu \hat{H}^\nu_n (y) \right] = - i \Gamma^{\mu \nu \rho} \delta_{mn} \delta(x-y) \frac{\partial}{\partial y^\rho} \frac{\partial}{\partial y^\nu}, \]
\[ \left[ \partial_\rho \partial_\nu \hat{H}^\mu_m (x), \hat{H}^\nu_n (y) \right] = i \Gamma^{\mu \nu \rho} \delta_{mn} \delta(x-y) \frac{\partial}{\partial x^\rho} \frac{\partial}{\partial x^\nu}. \quad (53) \]

From (50), (52) and (53), it becomes clear that the commutator between the derivatives of \( \hat{E}^\mu_m \) as well as \( \hat{H}^\mu_m \) vanishes if the expression where the commutator appears contains no further field factors. Within the applications of the commutation relations within the calculation of the generalized actions below, the delta functions do not appear explicitly. If there appear several derivatives, they refer to the same variable, since there always appear pointwise products of fields. This has already been the case concerning the calculation of the volume element (38)–(40). Besides the importance of the commutation relations (50), (52) and (53) which contain derivatives, it is further decisive that the application of the derivatives of the operators \( \hat{E}^\mu_m \) or \( \hat{H}^\mu_m \) to a coherent state, \( \ket{\Psi} \) or \( \ket{\Phi} \), yields the derivative of the corresponding eigenvalue which means that

\[ \partial_\mu \hat{E}^\mu_m \ket{\Psi} = \partial_\mu E^\mu_m \ket{\Psi}, \quad \langle \Psi | \partial_\mu \hat{E}^\mu_m = \langle \Psi | \partial_\mu E^\mu_m * \rangle, \]
\[ \partial_\mu \hat{H}^\mu_m \ket{\Phi} = \partial_\mu H^\mu_m \ket{\Phi}, \quad \langle \Phi | \partial_\mu \hat{H}^\mu_m = \langle \Phi | \partial_\mu H^\mu_m * \rangle. \quad (54) \]

The validity of identities (54) can be shown as follows:

\[ \partial_\mu \hat{E}^\mu_m (x) \ket{\Psi(x)} = \lim_{\epsilon \to 0} \frac{\hat{E}^\mu_m (x + \epsilon) - \hat{E}^\mu_m (x)}{\epsilon^\nu} \ket{\Psi(x)} = \lim_{\epsilon \to 0} \frac{E^\mu_m (x + \epsilon) - E^\mu_m (x)}{\epsilon^\nu} \ket{\Psi(x)} = \partial_\mu E^\mu_m (x) \ket{\Psi(x)}, \quad (55) \]
where in the second step \( \lim_{\epsilon \to 0} |\Psi(x + \epsilon)\rangle = |\Psi(x)\rangle \) has been used and therefore
\( \lim_{\epsilon \to 0} \hat{E}^\mu_m(x + \epsilon)|\Psi(x)\rangle = \lim_{\epsilon \to 0} E^\mu_m(x + \epsilon)|\Psi(x + \epsilon)\rangle = \lim_{\epsilon \to 0} E^\mu_m(x + \epsilon)|\Psi(x)\rangle \). To calculate the expectation values of the expressions \( i\gamma^m \hat{D}_m(\hat{h}^1) \psi \), \( i\gamma^m \hat{D}_m(\hat{h}^2) \psi \) and \( i\gamma^m \hat{D}_m(\hat{h}^3) \psi \) defined through (48), the operator \( \hat{h}^\mu_m \) within expressions (48) has to be replaced by \( \hat{H}_m^\mu \) and \( \hat{H}_m^{\mu\dagger} \) due to (31); then the permutation by using the commutation relations (53) has to be performed and after this the operators can be applied to the coherent state \( |\Phi\rangle \). For the expression \( i\gamma^m \hat{D}_m(\hat{h}^3) \psi \) defined through (48), the commutation relations (53) are not necessary, since there only appear linear expressions and therefore \( i\gamma^m \hat{D}_m(\hat{h}^1) \psi \) yields no additional terms compared with the classical case,

\[
|\Phi\rangle |\gamma^m \hat{D}_m(\hat{h}^1) \psi \rangle = \langle \Phi | \gamma^m \left( \hat{h}^\mu_m |\partial_\mu \rangle + \frac{i}{2} \left[ 2 \delta^{[\alpha}_{[\mu} \partial_\mu \hat{h}^\beta_{\nu]} + \partial_\mu \hat{h}^\beta_{\mu]} \right] \right) \rangle \langle \Phi | \psi \rangle \\
= \langle \Phi | \gamma^m \left( \hat{h}^\mu_m |\partial_\mu \rangle + \frac{i}{2} \left[ 2 \delta^{[\alpha}_{[\mu} \partial_\mu \hat{h}^\beta_{\nu]} + \partial_\mu \hat{h}^\beta_{\mu]} \right] \right) \rangle \langle \Phi | \psi \rangle \\
= \langle \Phi | \gamma^m \left( \hat{h}^\mu_m |\partial_\mu \rangle + \frac{i}{2} \left[ 2 \delta^{[\alpha}_{[\mu} \partial_\mu \hat{h}^\beta_{\nu]} + \partial_\mu \hat{h}^\beta_{\mu]} \right] \right) \rangle \langle \Phi | \psi \rangle.
\]

(56)

In the second step of (56), of course (54) has been used. The expectation value of \( i\gamma^m \hat{D}_m(\hat{h}^2) \psi \) defined through (48) is calculated in the following way:

\[
|\Phi\rangle |\gamma^m \hat{D}_m(\hat{h}^2) \psi \rangle = \langle \Phi | \gamma^m \left( \frac{1}{2} \left[ 2 \delta^{[\alpha}_{[\mu} \partial_\mu \hat{h}^\beta_{\nu]} + \partial_\mu \hat{h}^\beta_{\nu]} \right] + \left[ \hat{H}_{\mu\nu} + \hat{H}_{\nu\mu} \right] \right) \rangle \langle \Phi | \psi \rangle \\
= \langle \Phi | \gamma^m \left( \frac{1}{2} \left[ 2 \delta^{[\alpha}_{[\mu} \partial_\mu \hat{h}^\beta_{\nu]} + \partial_\mu \hat{h}^\beta_{\nu]} \right] + \left[ \hat{H}_{\mu\nu} + \hat{H}_{\nu\mu} \right] \right) \rangle \langle \Phi | \psi \rangle \\
= \langle \Phi | \gamma^m \left( \frac{1}{2} \left[ 2 \delta^{[\alpha}_{[\mu} \partial_\mu \hat{h}^\beta_{\nu]} + \partial_\mu \hat{h}^\beta_{\nu]} \right] + \left[ \hat{H}_{\mu\nu} + \hat{H}_{\nu\mu} \right] \right) \rangle \langle \Phi | \psi \rangle.
\]

(57)

where \( \chi_{ab}^m = -\frac{1}{2} \gamma^m \partial_{\nu} \psi \) has been defined. In the last step of (57), the following identity has been used:

\[
|\Phi\rangle |\hat{H}_{\mu\nu} + \hat{H}_{\nu\mu} \rangle \partial_\mu |\partial_\nu \psi \rangle = \langle \Phi | \hat{H}_{\mu\nu} \partial_\mu |\hat{H}_{\nu\mu} \partial_\nu \psi \rangle
\]

(57)
\[ \mathcal{F} = \langle \mathcal{F} \rangle \]

where (53) has been used and \( f(x) \) denotes an arbitrary field. The expectation value of \( i y^\nu D^\mu \hat{h} \psi \) defined through (48) can be calculated in the following way:

\[
\langle \mathcal{F} \rangle = \langle \mathcal{F} \rangle \left[ \left( \hat{h}_{\mu \nu} + \hat{h}_{\nu \mu} \right) \delta^{\mu \nu} \partial_\nu \hat{H}_\mu + \hat{h}_{\mu \nu} \delta^{\mu \nu} \partial_\nu \hat{H}_\mu \right]
\]

In the last step of (59), the following identity has been used:

\[
\langle \mathcal{F} \rangle = \langle \mathcal{F} \rangle \left[ \left( \hat{h}_{\mu \nu} + \hat{h}_{\nu \mu} \right) \delta^{\mu \nu} \partial_\nu \hat{H}_\mu \right]
\]
a fermionic field containing the tetrad field operator

In the previous section, we derived the expectation value of the matter field equation referring to

7. Generalized Einstein field equation

in

equation in the following section. Thus, the generalized matter field equation to the third order

which will become important concerning the derivation of the generalized Einstein field

(49), (56), (57) and (59).

holds:

where (53) has been used again and from which it can be seen directly that the following also

as follows:

which will become important concerning the derivation of the generalized Einstein field equation in the following section. Thus, the generalized matter field equation to the third order in \( h^p_\mu \) is given by (47) with the concrete expressions for the expectation values calculated in (49), (56), (57) and (59).

7. Generalized Einstein field equation

In the previous section, we derived the expectation value of the matter field equation referring to a fermionic field containing the tetrad field operator \( \hat{e}^{\mu \nu}_{ab} \), which is nothing else than a generalized Dirac equation on curved spacetime. In this section, we will consider the generalized dynamics of the gravitational field itself which is usually described by the Einstein field equation. In accordance with the above derivation of the generalized matter field equation, we consider the Einstein–Hilbert action with the tetrad field replaced by the tetrad field operator to obtain the correct generalized Einstein field equation. The usual Einstein–Hilbert action expressed by the tetrad field \( e^{\mu \nu}_{ab} \) reads

\[
S_{EH} = \int d^4x \, \det \left[ e^{\sigma \rho}_{ab} \right] \epsilon^{\mu \nu}_{ab} \epsilon^{\rho \sigma}_{\alpha \beta} [R^{\alpha \beta}_{\mu \nu}] \left[ e \right].
\]

where the Riemann tensor \( R^{\alpha \beta}_{\mu \nu} \) is expressed by the spin connection \( \omega^{ab}_{\nu \mu} \) as follows:

\[
R^{\alpha \beta}_{\mu \nu} = \partial_{\nu} \omega^{ab}_{\alpha \mu} - \partial_{\mu} \omega^{ab}_{\alpha \nu} + \omega^{ac}_{\nu \mu} \omega^{b}_{\alpha} - \omega^{bc}_{\nu \mu} \omega^{a}_{\alpha \nu}.
\]
the spin connection $\omega^{ab}_\mu[e]$ depends on the tetrad field $e^\mu_a$ according to (42) and therefore inserting the corresponding expression for the spin connection (42) into the Riemann tensor (63) yields the following expression for the Riemann tensor in dependence on the tetrad field $e^\mu_a$.

$$R^{ab}_{\mu \nu} = \partial_\rho (2e^{\rho[a} \partial_\nu e^{b]}_\mu) + e^{[a}_\mu e^{\rho]}_\nu h_{[a} e^{b]}_\rho - \partial_\nu (2e^{\mu[a} \partial_\rho e^{b]}_a) + e^{[a}_\mu e^{\rho]}_\nu h_{[a} e^{b]}_\rho$$

$$\text{+} (4e^{\rho[a} \partial_\mu \partial_\nu e^{b]}_\rho + 2e^{[a}_\mu \partial_\nu e^{b]}_\rho + e^{[a}_\mu e^{\rho]}_\nu h_{[a} e^{b]}_\rho + 2e^{[a}_\mu e^{\rho]}_\nu h_{[a} e^{b]}_\rho)$$

$$\text{+} e^{[a}_\mu e^{\rho]}_\nu h_{[a} e^{b]}_\rho$$

$$\text{+} e^{[a}_\mu e^{\rho]}_\nu h_{[a} e^{b]}_\rho$$

(64)

To obtain the dynamics of the generalized theory analogue to the matter action, the transition $e^\mu_a \rightarrow \hat{e}^\mu_a$ has to be performed with respect to the Einstein–Hilbert action (62),

$$S_{EH}[\hat{e}] = \int d^4x \; \det \left[ \hat{e}^\mu_a \right] \hat{e}^\rho_b \hat{R}^{ab}_{\mu \nu} [\hat{e}] .$$

(65)

The obtained action $\hat{S}_{EH}[\hat{e}]$ has to be varied with respect to $\hat{e}^\mu_a$ to derive the corresponding Einstein field equation containing the tetrad field operator $\hat{e}^\mu_a$. The generalized Einstein field equation containing the usual tetrad field $e^\mu_a$ can then be inferred by building the expectation value of this equation in complete analogy to the case of the derivation of the generalized matter field equation,

$$\langle \Psi | \frac{\delta \hat{S}_{EH}[\hat{e}]}{\delta \hat{e}^\mu_a} | \Psi \rangle = 0 \quad \Leftrightarrow \quad \langle \Psi | \frac{1}{\det [\hat{e}^\mu_a]} \frac{\delta \hat{S}_{EH}[\hat{e}]}{\delta \hat{e}^\mu_a} | \Psi \rangle = 0$$

$$\Leftrightarrow \quad \langle \Psi | \hat{G}^a_{\mu}[\hat{e}] | \Psi \rangle = \langle \Psi | \hat{R}^a_{\mu}[\hat{e}] - \frac{1}{2} \hat{R} | \hat{e}^\mu_a | \Psi \rangle = 0 ,$$

(66)

where the definition of the generalized Einstein tensor has been used: $\hat{G}^a_{\mu}[\hat{e}] = \hat{R}^a_{\mu}[\hat{e}] - \frac{1}{2} \hat{R} | \hat{e}^\mu_a |$. Equation (66) describes the expectation value of the Einstein field equation of the free gravitational field expressed in terms of the tetrad field operator $\hat{e}^\mu_a$. To obtain the expectation value of the Einstein field equation containing the tetrad field operator in the presence of matter, the generalized matter action obtained in (43) for example has to be included, leading to

$$\langle \Psi | \frac{1}{\det [\hat{e}^\mu_a]} \left( \frac{\delta \hat{S}_{EH}[\hat{e}]}{\delta \hat{e}^\mu_a} + \frac{\delta \hat{S}_M[\hat{e}]}{\delta \hat{e}^\mu_a} \right) | \Psi \rangle = 0$$

$$\Leftrightarrow \quad \langle \Psi | \hat{T}^a_{\mu}[\hat{e}] | \Psi \rangle = \frac{1}{2} \hat{R} | \hat{e}^\mu_a |$$

$$\Leftrightarrow \quad \langle \Psi | \hat{G}^a_{\mu}[\hat{e}] | \Psi \rangle = \langle \Psi | \hat{R}^a_{\mu}[\hat{e}] - \frac{1}{2} \hat{R} | \hat{e}^\mu_a |$$

(67)

where the definitions of the generalized Ricci tensor, $\hat{R}^a_{\mu} = \hat{e}^\rho_b \hat{R}^{ab}_{\mu \nu}$, and the generalized Ricci scalar, $\hat{R} = \hat{e}^\rho_b \hat{R}^{ab}_{\mu \nu}$, have been used as well as the definition of the energy–momentum tensor transferred to the definition of the generalized energy–momentum tensor depending on the tetrad field operator $\hat{e}^\mu_a$.

$$\hat{T}^a_{\mu}[\hat{e}] = \frac{1}{\det [\hat{e}^\mu_a]} \frac{\delta \hat{S}_M[\hat{e}]}{\delta \hat{e}^\mu_a} ,$$

(68)

where the matter action $\hat{S}_M[\hat{e}]$ depending on the tetrad field operator $\hat{e}^\mu_a$ is defined according to (43), for example, if the matter field is a fermionic field. To calculate the concrete expression of the expectation value of the generalized Einstein field equation (67), again a series expansion
according to (13) has to be performed, \( \hat{\delta}_m^\mu = \delta_m^\mu + \hat{\delta}_m^\mu \), leading to an expectation value of the Einstein field equation containing the tetrad field operator which is of the following shape:

\[
\langle \Phi | \hat{G}_m^\nu [\hat{h}] + \hat{G}_m^\nu [\hat{h}] + \hat{G}_m^\nu [\hat{h}] \rangle \Phi + \mathcal{O}(\hat{h}) = -8\pi G(\Phi | \hat{T}_m^\nu [\hat{h}, \hat{h}, \hat{h}] | \Phi + \mathcal{O}(\hat{h}^2),
\]

(69)

where \( \hat{G}_m^\nu [\hat{h}] \), \( \hat{G}_m^\nu [\hat{h}] \) and \( \hat{G}_m^\nu [\hat{h}] \) denote the expressions of the Einstein tensor depending linearly, quadratically and to the third power on \( \hat{h}_m^\nu \) and have the following shape, where \( \mu \leftrightarrow \nu \) denotes the term in the bracket with the indices \( \mu \) and \( \nu \) exchanged:

\[
\hat{G}_m^\nu [\hat{h}] = \left[ \partial_\mu (2\delta^\rho_{\nu\lambda}a_{\lambda}^\mu \hat{h}_m^\nu + \hat{a}^{\mu\nu^2} + \hat{a}^{\mu\nu} \hat{h}_m^\nu) - \mu \leftrightarrow b \right] - \frac{1}{2} \left[ \partial_\mu (2\delta^\rho_{\nu\lambda}c_{\lambda}^\mu \hat{h}_m^\nu + \hat{a}^{\mu\nu^2} + \hat{a}^{\mu\nu} \hat{h}_m^\nu) - b \leftrightarrow c \right] \delta^\mu_{\nu},
\]

(70)

\[
\hat{G}_m^\nu [\hat{h}] = \left[ \partial_\mu (2\delta^\rho_{\nu\lambda}a_{\lambda}^\mu \hat{h}_m^\nu + \hat{a}^{\mu\nu^2} + \hat{a}^{\mu\nu} \hat{h}_m^\nu) - \mu \leftrightarrow b \right] + \frac{1}{2} \left[ \partial_\mu (2\delta^\rho_{\nu\lambda}c_{\lambda}^\mu \hat{h}_m^\nu + \hat{a}^{\mu\nu^2} + \hat{a}^{\mu\nu} \hat{h}_m^\nu) - b \leftrightarrow c \right] \delta^\mu_{\nu},
\]

(71)
Within expressions (70)–(72), the Kronecker symbols have only been contracted, where this contraction has improved the representation. The expectation values of the terms of the Einstein tensor which are linear in $\hat{h}^a_{\mu}$ and quadratic in $\hat{h}^b_{\mu}$ yield no additional terms arising from the noncommutativity of the tetrad field. This means that the following holds:

$$\langle \Phi | G^a_{\mu} [\hat{h}^1] | \Phi \rangle = G^a_{\mu} [\hat{h}^1], \quad \langle \Phi | G^2_{\mu} [\hat{h}^2] | \Phi \rangle = G^2_{\mu} [\hat{h}^2].$$

(73)

In the case of the tensor $G^a_{\mu} [\hat{h}^1]$, this is obvious since no permutation of $\hat{H}^a_{\mu}$ and $\hat{H}^b_{\mu}$ has to be performed. In the case of $G^2_{\mu} [\hat{h}^2]$, this property arises from the fact that there exist only terms which contain the derivatives of $\hat{h}^a_{\mu}$ or $\hat{H}^a_{\mu}$ and according to (53), the commutators lead to derivatives acting on the other factors of the corresponding term. Since these terms contain no further field factors on the one hand and they are just quadratic in $\hat{H}^a_{\mu}$ on the other hand, the commutators vanish. From (58) one can also see that there arise no additional terms from the expressions quadratic in $\hat{h}^a_{\mu}$ if there appear no further field factors. The expectation value of the term of the Einstein tensor depending on $\hat{h}^a_{\mu}$ to the third order, $G^3_{\mu} [\hat{h}^3]$, yields additional terms depending on the noncommutativity parameter. To calculate the expectation value, identities (60) and (61) have to be used. This leads to the following expression:

$$\langle \Phi | G^3_{\mu} [\hat{h}^3] | \Phi \rangle = \left[ \partial_{\rho} \left( \delta^{ad}_{\rho \sigma} \partial_{\sigma} \rho + \delta^{ba}_{\rho \sigma} \partial_{\sigma} \rho + \partial_{\rho} \partial_{\sigma} \rho + \partial_{\rho} \partial_{\sigma} \rho \right) + 4 \delta^{ad}_{\rho \sigma} \left[ \Omega_{\mu \nu} \rho |_{\mu \nu} + 2 \Omega_{\mu \nu} \rho |_{\mu \nu} \right] \right]_{\rho \sigma}.$$
This means that the generalized Einstein field equation has been calculated which is given by (69) with (73) and thus (70) and (71) transformed to the corresponding expressions depending on the usual expansion field $h^m_i$. Accordingly, we have derived generalized dynamics for the matter field coupled to gravity and for the gravitational field itself by building expectation values of the corresponding matter field equation and the Einstein field equation containing the tetrad field operator $\hat{e}^a_i$ with respect to coherent states referring to creation and annihilation operators being built from the components of the tetrad field operator $\hat{e}^a_i$. There appear imaginary components within the expectation values of the field equations of the matter and the gravitational field. This is because the expressions within the field equations in terms of the tetrad field operator are not Hermitian, although the tetrad field operator is Hermitian, since a function of Hermitian operators is not always Hermitian as well. This means that the field equations contain two independent conditions on the dynamical evolution of the matter field and the gravitational field respectively being related to the real and the imaginary component.

8. Summary and discussion

In this paper, we have suggested a noncommutativity algebra for the components of the tetrad field describing the gravitational field which behaves thus as an operator. This algebra corresponds to the algebra of the canonical case of usual noncommutative geometry referring to the components of the spacetime coordinate. The relation between the usual quantization in quantum mechanics where a nonvanishing commutator between position and momentum is postulated and the concept of noncommutative geometry where nonvanishing commutation relations between the components of the position vector are postulated corresponds to the relation between canonical quantization of general relativity and the noncommutativity of the tetrad field which has been presented in this paper. This is because canonical quantization of general relativity postulates noncommutative relations between the quantities describing the gravitational field and its canonical conjugated momentum and in the approach of this paper nonvanishing commutation relations are postulated between the components of the quantity describing the gravitational field. In this sense, the concept of a description of general relativity with a noncommutative tetrad field as is treated in this paper could be considered as the consequence of an additional aspect of a fundamental quantum theory of gravity with respect to a kind of semiclassical description of general relativity.
To describe the consequences of the noncommutativity of the tetrad field for the description of general relativity, it has been necessary to transfer the coherent state approach to noncommutative geometry to the case of noncommuting components of the tetrad field and thus to the case of noncommuting field components. The necessity to use the coherent state approach has its origin in the fact that in the star product approach, there are treated fields depending on noncommutative coordinates, and therefore in this case it is possible to use a Weyl quantization which is a kind of Fourier transformation between commuting and noncommuting coordinates. After the Weyl quantization, products of fields depending on the noncommutative coordinates are expressed by a sum of products of these fields depending on the usual coordinates and in this way the generalization of a field theory depending on noncommutative coordinates can be expressed by additional products depending on the noncommutativity parameter. But in the scenario considered in this paper, the components of a field themselves, namely of the gravitational field represented as tetrad field, do not commute and therefore a Weyl quantization cannot be performed. In the coherent state approach, we define new expressions for the quantities depending on the noncommutative coordinates as expectation values between coherent states which are defined with respect to operators constructed from the components of the spacetime coordinate. This procedure can be applied to coordinates as well as fields. Therefore, in this paper, the coherent state approach has been transferred to the tetrad field and accordingly the generalized quantities of general relativity depending on the noncommutative tetrad field are defined as expectation values between coherent states which are defined with respect to operators constructed from the components of the tetrad field operator. Based on this concept, the expectation values of the resulting operator of the metric field and the volume element operator have been calculated as examples. After this, the generalized dynamics of a matter field coupled to gravity, a fermionic field especially, and the generalized dynamics of the gravitational field itself have been determined. These dynamics were obtained by replacing the usual tetrad field by the corresponding tetrad field operator within the actions of the fermionic field coupled to gravity and the Einstein–Hilbert action describing the dynamics of the gravitational field, varying these actions by the tetrad field operator and then building the expectation values between the coherent states. Since the expressions in an exact calculation would become very large, an expansion of the gravitational field around the Minkowski metric or the corresponding Kronecker symbol describing the corresponding tetrad field was used and a calculation to the third order in the expansion field performed, since the terms to the first and second order do not lead to a modification of the expressions in the case of the Einstein field equation.

The deviation terms of the field equations, which of course depend on the tensor $\Lambda^{\mu\nu}$ defining the noncommutativity algebra of the tetrad fields, are of lower order in $h_{\mu\nu}$ than the terms from which they arise. Since they are obtained from the commutator of the tetrad field being proportional to $\Lambda^{\mu\nu}$, the power is two orders lower. In the case of the free Einstein equation, this means that the deviation terms are of first order in $h_{\mu\nu}$ and this implies that even in the case of a small $\Lambda^{\mu\nu}$ for a very small perturbation of the gravitational field, the deviation terms yield a bigger contribution than the usual terms to the second and third order in $h_{\mu\nu}$. Thus, in the generalized Einstein equation, it seems to be advantageous to consider only the terms to the first order in $h_{\mu\nu}$ concerning the treatment of further investigations of the theory like the solution of the field equations for the matter field and the gravitational field or the derivation of a propagator. It would be interesting to explore the possible relation between the presented theory and the canonical quantization of general relativity. Noncommutativity relations between coordinates can be derived as a consequence of a generalized uncertainty relation between position and momentum in quantum mechanics. Analogously it could make sense to postulate a generalized quantization rule for the gravitational field and its canonical
conjugated momentum and to derive the noncommutativity of the tetrad field from such a generalized quantum description of the gravitational field.

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