Chromitite layers require the existence of large, long-lived, and entirely molten magma chambers

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Chromitite layers require the existence of large, long-lived, and entirely molten magma chambers

An emerging and increasingly pervasive school of thought is that large, long-lived and largely molten magma chambers are transient to non-existent in Earth’s history\(^1\)–\(^{13}\). These ideas attempt to supplant the classical paradigm of the ‘big magma tank’ chambers in which the melt differentiates, is replenished, and occasionally feeds the overlying volcanoes\(^{14}\)–\(^{23}\). The stratiform chromitites in the Bushveld Complex – the largest magmatic body in the Earth’s crust\(^{24}\) – however, offers strong contest to this shifting concept. Several chromitites in this complex occur as layers up to 2 metres in thickness and more than 400 kilometres in lateral extent, implying that chromitite-forming events were chamber-wide phenomena\(^{24}\)–\(^{27}\). Field relations and microtextural data, specifically the relationship of 3D coordination number and grain size, indicate that the chromitites grew as a 3D framework of touching chromite grains directly at the chamber floor from a melt saturated in chromite only\(^{28}\)–\(^{30}\). Mass-balance estimates dictate that a 1 to 4 km thick column of this melt\(^{26}\),\(^{31}\),\(^{32}\) is required to form each of these chromitite layers. Therefore, an enormous volume of melt (>1,00,000 km\(^3\))\(^{24}\),\(^{25}\) must have been involved in

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the generation of all the Bushveld chromitite layers, with half of this melt being expelled from the magma chamber. We therefore argue that the very existence of thick and laterally extensive chromitite layers in the Bushveld and other layered intrusions strongly buttress the classical paradigm of ‘big magma tank’ chambers.

For over a century, the classical paradigm of magma chambers has underpinned all models of the Earth’s magmatism. This paradigm envisages a magma chamber as a large body of the molten, long-lived, and slowly fractionating magma (‘a big magma tank’) enclosed in crustal rocks. In recent years, this classic view of a magma chamber garnered increased scrutiny from many geoscientists who contend that such long-lived and largely molten ‘big tank’ magma chambers are either very short-lived or never existed in Earth’s history. For instance, some igneous petrologists maintain that tectonic processes that create open spaces in the crust at any tectonic setting are so slow that the liquid-dominated magma bodies of any composition can hardly form. Instead, they have suggested that plutons grow incrementally from numerous sills/dykes a few meters thick that cool and may even become totally solidified between injections. Similarly, most volcanologists have abandoned the classic paradigm because the geophysical surveys have failed to detect any present-day eruptible magma bodies in the Earth’s crust. As an alternative, they proposed the existence of transcrustal mushy systems (including mushy reservoirs for mafic layered intrusions) that are formed in the crust from numerous collating intrusions. The transcrustal systems contain only small melt lenses that are produced by compaction or tectonic destabilization of the crystal mush and exist for only a very short period of time before accumulating and erupting as lavas on the Earth’s surface. Yet another group of mafic plutonists, influenced by out-of-sequence zircon geochronological data, have proposed that mafic plutons do not require the existence of large magma chambers. These are rather produced as a stack of randomly-emplaced sills,
with successive crystal-rich pulses often invading pre-existing cumulates\textsuperscript{9–13}. In stark contrast to these ‘anti-magma-chamber’ approaches, we argue here that the existence of large magma chambers is indicated by laterally extensive layers of chromite-rich cumulates, which require many times their own volume of magma to supply the key component, chromium (Cr). We present here field and microtextural data from massive chromitites of the Bushveld Complex that indicate that ‘big magma tank’ chambers in the Earth’s crust are a reality that cannot and should not be dismissed.

**The enormous extent of chromitite layers**

The 2.05 Ga Bushveld Complex in South Africa (Fig. 1a) is the largest mafic-ultramafic layered intrusion in Earth’s crust; totalling to about 600,000 km\(^3\) of igneous rocks\textsuperscript{24,25}. The complex consists of several parts, the western, eastern and northern limbs being the largest, and is subdivided stratigraphically into five major units - the Marginal, Lower, Critical, Main, and Upper Zones, comprising a total thickness of 7 to 9 km\textsuperscript{24,25}. The Bushveld Complex contains >80% of the Earth’s known chromium resources\textsuperscript{33}, an element critical to improving the material properties of steel, making this magmatic body an object of perennial study. The chromium is hosted within 14 principal layers of massive chromitites, mostly confined to the Critical Zone\textsuperscript{26,27}. Three major groups of chromitites are recognised: the Lower (LG1 to LG7); Middle (MG0 to MG4); and Upper Groups (UG1 to UG3)\textsuperscript{27}. The thickness of individual chromitite layers ranges from a few decimeters to 2 meters. Mining activities have allowed most of these layers to be traced across the entire Bushveld Complex\textsuperscript{27}. Remarkably, the vertical distribution of platinum-group elements across some of these chromitites are nearly identical in places that are separated laterally by over 300 km\textsuperscript{26}. The vast lateral extent and mineralogical uniformity of chromitite layers indicate that the process responsible for their
Fig. 1. A notable lack of textural evidence for gravity settling of chromite crystals within a UG1 chromitite of the Bushveld Igneous Complex, South Africa. **a**, Location and generalized geological map of the Bushveld Complex emphasizing its enormous size. Note that this is just an erosional remnant of the complex so that its original size was even larger. The immense lateral extent of the chromitite layers (>350-400 km) in this complex is illustrated here by the UG1 chromitite that occurs at the top of the Critical Zone. Also indicated are places from which the studied samples were obtained. Modified from reference 28. **b**, Panoramic view of a few sublayers of the UG1 chromitite in the anorthosite footwall at the Dwars River, Eastern Limb. **c**, A close-up photograph of the top part of the UG1 chromite sublayer. Note large plagioclase oikocrysts enclosing numerous small crystals of cumulus chromite (chadacrysts). **d**, Photograph of a thin-section (under plane polarized light) of the UG1 chromitite showing isolated chromite grains and their loose clusters enclosed by a single large oikocryst of plagioclase. Red arrows emphasize that chromite grains show no tendency to gravitate downwards despite a high porosity of the framework (~65 vol.%). Sample HX-07-153.33, Mototolo mine, Eastern Limb.
formation has been working synchronously in all parts of the superlarge chamber to produce
the same chromitite layer over lateral distances of up to 400 km (e.g., UG1 in Fig. 1a).

Field and textural evidence for in situ growth of chromite
The nature of this chamber-wide process can be constrained from field and textural features of
massive chromitite layers such as the 2 m-thick UG1 chromitite – the thickest and the best
exposed layer in the entire complex (Fig. 1b). This chromitite shows remarkable field
relationships with its respective footwall rocks. In addition to its occurrence on the planar
portions of the chamber floor, this chromitite develops within potholes, roughly circular
depressions in which footwall rocks are missing due to magmatic erosion\textsuperscript{28,34}. In these areas,
the planar UG1 chromitite that occurs along the periphery of the potholes commonly passes,
without any apparent changes in thickness and texture, into the steeply dipping, subvertical
and even overhanging UG1 chromitite in the interior of potholes\textsuperscript{28,34} (Extended Data Fig. 1).
This field observation strongly argues against the formation of the UG1 chromitite, both on the
planar and overhanging portions of the chamber floor, by processes involving gravity-induced
settling of chromite through either the resident melt\textsuperscript{35–38} or a crystal-rich mush\textsuperscript{39,40}. The
simplest alternative mechanism is in situ growth of chromite directly at the chamber floor from
a chromite-only-saturated melt\textsuperscript{28,30}. This is the only process that allows the chromitite layer to
cover all the planar and irregular margins, even the places where gravity-settling of chromite
grains is physically impossible (i.e., “gravity-settling shadows” in which dips are
overturned\textsuperscript{28,34}) (Extended Data Fig. 1).

An intriguing challenge here is to decipher how in situ growth of chromite is recorded in the
texture of massive chromitites themselves. We have re-visited the UG1 chromitite from the
classical Dwars River locality\textsuperscript{29} (Fig. 1b) where it is composed of 25-50 vol\% of cumulus
chromite that occurs as separate idiomorphic grains or clumps of grains that are smaller than 0.1 mm in size (Fig. 1c, d). The chromite grains are enclosed within much larger oikocrysts of plagioclase (up to 5-10 cm in size) that are clearly visible in outcrops (Fig. 1c). The traditional interpretation of such layers in the frame of gravity settling models is that chromite was the first to settle on the chamber floor\textsuperscript{35-38} followed, after some period of post-depositional cooling, by \textit{in situ} growth of plagioclase oikocrysts from the interstitial melt in a mushy chromitite. An important point is that settling chromite grains have enough time to reach the chamber floor and start growing there. The subsequently forming oikocrysts may capture and arrest chromite from experiencing further growth, producing snapshots of an immature solidification front.

A close look at the UG1 texture (Fig. 1d) raises, however, a simple but fundamental quandary. Chromite is almost twice as dense as a basaltic melt (4,800 kg/m\textsuperscript{3} and 2,600 kg/m\textsuperscript{3}, respectively) and is expected to settle to the chamber floor in a random closely-packed lattice in which all adjacent chromite grains are touching each other. However, this is not the case as chromite occurs as individual grains and clumps of grains that are ‘suspended’ within plagioclase oikocrysts (Fig. 1d). This observation leads to a critical question: why have the chromite grains/clumps failed to sink towards the chamber floor despite being much denser than the host melt? A potential clue to this puzzle is that the chromite grains in the UG1 layer appear to be arranged in chain-like aggregates\textsuperscript{41,42}.

**Three-dimensional framework of chromite crystals**

The analysis and quantification of chromitite in three-dimensions (3D) using high-resolution X-ray computed tomography (HRXCT) revealed that nearly all chromite grains (97 vol\%) from the UG1 chromitite are interconnected to form a single continuous 3D framework composed of many thousands grains that extend across multiple plagioclase and pyroxene oikocrysts (Fig.
Fig. 2. Results of high-resolution X-ray computed tomography revealing that nearly all chromite grains in UG1 chromitite are interconnected within a continuous 3D framework. (a) Three orthogonal slices virtually cut through the UG1 sample (HX-07-153.33, Mototolo mine, Eastern Limb) showing chromite in light grey; (b) Segmented chromitite showing how a single interconnected network (coloured in yellow) covers the entire sample volume; (c) Volume rendering of chromite grains displayed using a false 256 colour scale; (d) Expanded view of a volume of interest showing details of the interconnected chromite network (orange) and the isolated chromite grains (in cyan); (e) and (g) Details of selected chromite cluster morphologies within the large interconnected chromite network where only a small number of grains are coloured to improve visibility; (f) Histogram showing the size distribution of chromite grains in the sample (ESD: equivalent sphere diameter).
We can now consider whether these UG1 microstructures (Fig. 2) could be generated by random loose packing of non-interacting particles, i.e. from small independent chromite grains settling from a basaltic melt by gravity settling or kinetic sieving. Theoretical microstructures of random mechanical crystal packs can be predicted by packing theory and characterised by two properties: the packing density (inverse of porosity); and the distribution of coordination numbers, i.e., the number of nearest neighbour grains in contact. Quantification of clustering and chain formation has typically used assumptions of constant grain sizes but these parameters are known to be sensitive to the particle size distribution. No observations or simulations have been made to date using the characteristic negative log-linear particle size distributions of crystals found in cumulates. To address this gap, we measured chromite particle sizes and coordination number distributions from the segmented HRXCT scan of the UG1 chromitite (Methods, Supplementary Data 1), choosing a volume within which chromite grains are primarily enclosed within plagioclase or pyroxene oikocrysts. This is to eliminate possible effects on microstructure caused by later recrystallisation and annealing. We then compared the results with those of a Discrete Element Method computer simulation of a random loose packing, generated by simple settling, of an assemblage of crystals with the same size distribution as the UG1 sample (Methods). Results show two distinct differences between the synthetic pack and the natural sample (Fig. 3): firstly, the packing densities are greatly different, being much higher in the synthetic pack than in the natural sample (60% vs. 27%). Secondly, the distribution of coordination numbers is significantly different (Fig. 3b, c). In the random pack, coordination number increases exponentially with the grain size. This happens because, for the geometrical reasons, larger grains have a larger surface area and hence are likely to be in contact with a larger number of smaller grains filling space between them. In the natural UG1 data set, however, the coordination number flattens out and remains roughly constant at about 10 for grains larger...
than 150 microns in size (noting that only a small proportion of the total number of grains falls in this size range). This results from the chromite grains forming an open cage-like or chain-like structure where gaps in the framework are not occupied by other grains, causing lower coordination numbers at larger grain sizes. Furthermore, the natural sample contains a significant proportion of isolated or nearly isolated grains with coordination numbers of 0, 1 or 2; these are absent in the simulation. We conclude that (a) the low packing density, (b) the presence of isolated individual grains not supported by contact with any other chromite grains, and (c) the relationship between coordination number and crystal size in the UG1 chromitite are not consistent with random mechanical accumulation of non-interacting chromite grains, be it crystal settling in a melt\textsuperscript{35–38} or kinetic sieving in a crystal mush\textsuperscript{39,40}.

There remains a possibility that chromite may settle in the form of chromite chains/clusters produced either by heterogeneous nucleation against chromite grains suspended in the convecting melt\textsuperscript{46} or physical collision of isolated chromite crystals ‘swimming together’ in this melt (i.e. “synneusis”)\textsuperscript{47}. The accumulation of such clustered chains on the chamber floor may, in principle, give rise to the formation of a continuous 3D framework of touching chromite crystals. However, we reject this scenario on the field evidence: neither individual grains nor clustered chains can settle onto overhanging margins of potholes\textsuperscript{28,34} (Extended Data Fig. 1). Sidewall crystallisation thus indicates that the 3D chromite framework has crystallized \textit{in situ}, i.e., directly at the chamber floor. This may only happen by one physical process – heterogeneous/self-nucleation\textsuperscript{22,46,48} of chromite grains on the floor cumulates\textsuperscript{28,34}. To develop this point further, we also compare the microstructure of a Merensky Reef chromitite seam that
occurs on a vertical to overhanging sidewalls of potholes (Extended Data Figs. 3 and 4; Supplementary Video 1), taken as a definitive example of a microstructure that could only have developed in situ\textsuperscript{34,49} (Fig. 3c).

**Fig. 3. Results of numerical simulation compared to the observed natural data for the UG1 chromitite** showing the contrasting relationship between grain size and coordination number between the randomised loose packing simulation (using the observed UG1 grain size distribution) and the two natural samples, the UG1 chromitite and MR overhanging chromitite. (a) 3D perspective view of a simulation pack of spheres having the same size distribution as chromite grains in the UG1 chromitite; (b) Histogram showing the distribution of coordination number in UG1 simulation, UG1 chromitite and MR overhanging chromitite; (c) Plot showing mean coordination number of all grains within each size range bin, as a function of size range of chromite grains in the UG1 simulation, UG1 chromitite and MR overhanging chromitite (ESD: equivalent sphere diameter).
The packing density is likewise lower than the random loose packing simulation (53% vs 60%) (Supplementary Data 1). The coordination number vs grain size curve for this seam shows a steady linear increase with grain size, but, like the UG1 chromitite, has systematically lower coordination number values than the simulation along the entire length of the trend (Fig. 3c). Significantly, the packing and coordination number characteristics of the UG1 chromitite are much closer to those of the indisputably in-situ crystallised sidewall Merensky Reef chromitite than to the random packing simulation. We deduce that the relatively low coordination number values and low packing densities of both natural samples are the result of in situ growth of chromite chains or cages by heterogeneous self-nucleation. We also conclude that chromite grains (Fig. 1d) are not able to settle freely towards the chamber floor simply because they are all bound together in self-supporting frameworks attached to the floor.

A scenario for in situ growth of chromite on the chamber floor

In situ growth of chromite requires crystallisation from a parental melt that was saturated in chromite as the only liquidus phase. Such melts can be produced in response to decompression during their ascent from a deep staging reservoir towards Earth’s surface \(^{30}\). We propose that the Bushveld chamber has been replenished by such melts as basal flows that caused thermochemical erosion of the floor cumulates\(^{28}\), including the excavation of potholes (Fig. 4a). Upon cooling, the melt became saturated in chromite only \(^{30}\) (Fig. 4b), with the first chromite grains being nucleated heterogeneously on pre-existing plagioclase crystals of the floor anorthosites. With further cooling, chromite started preferentially self-nucleating on earlier-formed chromite grains to produce composite 3D clusters which subsequently merged into a continuous 3D framework of touching chromite grains (Fig. 4c). New crystals emerged in the system mostly by self-nucleation because the activation energy for this process is much lower relative to other types of nucleation\(^{22,46}\). A small portion of crystals (3 vol.%) that occur
as entirely discrete grains (Fig. 2d) has likely formed by homogeneous nucleation in the interstitial space. We envisage that chemical differentiation of the resident melt in the chamber at that time occurred by convective removal of a buoyant compositional boundary layer\(^{50}\) from \textit{in situ} growing chromite crystals in a 3D framework (Fig. 4c). The differentiation is aided by high porosity and permeability of a 3D crystal framework that permits the easy chemical exchange of melts between the crystal framework and the main magma body. The remarkable preservation of the nucleation/growth history in the UG1 chromitite is due to the early growth of plagioclase and pyroxene oikocrysts which have ‘frozen in’ a 3D chromite framework at its early immature stage. This continuous 3D framework of chromite grains (Fig. 2) is the first documented example showing how a natural solidification front looks when it develops via self-nucleation/heterogeneous nucleation\(^{48}\). Under other circumstances, the initial framework would have evolved into perfect chromitite adcumulate (up to 100% chromite; e.g. LG chromitite in Extended Data Fig. 5) in which all primary information is lost. We propose that most layers of monomineralic chromitites in the Bushveld Complex started their life as porous, \textit{in situ} produced 3D chromite frameworks that can no longer be seen due to adcumulus overgrowth\(^{51}\). We further propose that the observed shape of the coordination number vs grain size curve (Fig. 3c) may be diagnostic of chromite cumulates formed by \textit{in situ} crystallisation involving heterogeneous self-nucleation\(^{28,34}\).

**Chromium budget requires a large magma volume**

The realization that massive chromitites form by \textit{in situ} growth of chromite directly in the chamber (Fig. 4a) – rather than from chromite phenocrysts brought into the chamber with externally-derived crystal-rich mushes\(^{9,12,38,40}\) – logically brings us to a long-known Cr mass balance issue\(^{32,52,53}\). The stratiform chromitite layers in layered intrusions can be up to 2 m thick and contain 40-50 wt.% Cr\(_2\)O\(_3\), yet have evidently crystallized from a basaltic melt that
was unlikely to have contained more than 1000 ppm Cr\textsuperscript{31}. An implication is that the formation of a thick chromitite layer, such as the UG1 chromitite, requires extraction of Cr from a very large volume of liquid that can be present either as a thick melt layer in the chamber or as the melt flowing through the chamber, or both. To illustrate, given the thickness of a chromite layer, from mass-balance considerations one can calculate the volume of the parental melt in terms of the thickness of an equivalent layer. If a chromite layer crystallized at 1250°C and the \( f_{O_2} \) equivalent to QFM buffer, then based on experimental data\textsuperscript{31,52}, the chromite should contain \(~45\%\) Cr\textsubscript{2}O\textsubscript{3}, and the coexisting melt should carry about 0.10 wt% Cr\textsubscript{2}O\textsubscript{3}. Assuming that chromitites formed from the overlying melt, it can be estimated that a 1 m thick layer of chromitite will require a magma column of about 4 km thick (Fig. 4a). The thickness can be reduced to 2 km\textsuperscript{32} or 1 km\textsuperscript{26} if Cr solubility in a parental melt is to increase by its higher temperature or lower \( f_{O_2} \). These estimations assume 30% of the Cr removal from a parental melt\textsuperscript{32}. One cannot remove any more Cr from the melt than that because otherwise the melt will reach a cotectic with other liquidus phases (e.g., olivine or orthopyroxene) terminating chromitite formation (Fig. 4b).

This one-dimensional modelling illustrates the mass-balance issue. Applying this logic to the Bushveld Complex, it has been estimated that the formation of its most prominent chromitites would have required a column of 13 to 15 km of a parental chromite-saturated liquid\textsuperscript{25,26}. In addressing this mass-balance requirement, Cawthorn and Walraven\textsuperscript{25} modelled the Bushveld chamber as a long-lived flow-through system (~75,000 years life-time) that developed via a large number of injection events, partial crystallization of these magma batches, and removal of their residual liquids from the chamber by the succeeding magma batches. They concluded that the total volume of basaltic magma involved was 740-1200*10\textsuperscript{3} km\textsuperscript{3}, with only ~50% of this being represented by the cumulates now seen within the Bushveld Complex\textsuperscript{25}. The excess
magma has likely escaped from the chamber laterally\textsuperscript{26} to form mafic intrusions, such as the Molopo Farms Complex located about 200 km west of the Bushveld Complex\textsuperscript{54,55}.

Fig. 4. A ‘big-tank’ chamber of the Bushveld Complex filled with a resident melt that crystallizes into the UG1 chromitite at the chamber floor. \textit{a}, Schematic cartoon of the Bushveld chamber that shows crystallization of the UG1 chromitite near the top of the Critical Zone. The formation of a 1 m thick layer of such chromitite requires an equivalent layer of chromite-only-saturated melt of at least 1 to 4 km in thickness\textsuperscript{31,32}. This large volume of liquid may be present in the chamber either as a melt column or as a melt that flows through the chamber for a long period of time. The chromite nucleates and crystallizes directly on the chamber floor, with the resident melt convecting turbulently to deliver Cr for \textit{in situ} chromite growth. \textit{b}, Ol–Chr–Qtz phase diagram illustrating the position of a chromite-only-saturated melt parental to the UG1 chromitite. The diagram is modified from reference\textsuperscript{35}. Ol, olivine; Opx, orthopyroxene, Chr, chromite; Qtz, quartz. \textit{c}, A close-up view of the immature UG1 chromitite that forms a 3D framework of touching chromite crystals which self-nucleate on the floor of a magma chamber. Note the low density packing and the low coordination number of chromite in the framework. A compositional boundary layer of buoyant liquid is produced around crystallizing chromite clusters that migrates towards their apex and is released into the overlying melt in the form of compositional plumes, thus causing chemical differentiation in the resident melt.
A classical magma chamber paradigm affirmed

The enormous lateral extent of *in situ* formed chromitite layers and related mass-balance considerations indicate that during the formation of massive chromitites the Bushveld chamber has been operating as a giant magma body of more than 400 km in diameter, with a column of the resident melt likely attaining a few km in thickness. Thus, starting from this stage the Bushveld Complex has been developed as a large, long-lived and largely molten magma chamber (a true ‘big tank’ reservoir) in Earth's crust (Fig. 4a). The conclusion is further supported by the remarkable homogeneity of Sr isotopes over an interval of more than 2.5 km of the Upper Zone, which indicates a melt column thickness in the chamber being that thick or even thicker. This is in contrast with the recent assertion, mostly based on out-of-sequence geochronology, that depicts this giant complex as a stack of thin crystal-rich sills. Field relationships indicate, however, that zircon isotopic ages in these studies were almost certainly misinterpreted. We thus argue that the field and textural evidence from massive chromitites disprove this radical re-interpretation of the Bushveld Complex. Our inference may be extended to all large mafic-ultramafic layered intrusions that contained thick and laterally extensive layers of monomineralic chromitites (e.g., Stillwater and Great Dyke). Such intrusions are quite rare in the crust through the whole of geological time, so it is not surprising that there are no known examples of equivalent magma chambers that are active and detectable in the present-day Earth’s crust. We conclude that it is too early to discard the classical paradigm of a magma chamber developed by several generations of petrological luminaries. Rather, we suggest re-directing our efforts to find out how new geophysical, geochronological and thermal/diffusion modelling can be logically reconciled with the classical paradigm.
Any methods, additional references, Nature Research reporting summaries, source data, extended data, supplementary information, acknowledgements, peer review information; details of author contributions and competing interests; and statements of data and code availability are available at……………………………

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Methods

Rock sampling and petrography

Documentation of field observations of the UG1 chromitite was undertaken at the Dwars River locality and its sampling at the nearby Mototolo Mine from the HEX 076 drill-core in the Eastern Bushveld Complex. Thin sections and polished blocks were cut from orientated sample blocks to be as close as possible to the original vertical position. Thin sections and polished blocks were studied using a polarised light microscope with a circular stage and photographed using the Olympus 224 BX-63 OM/FM optical microscope housed at the MMU (Microscopy and Microanalysis Unit) of the University of the Witwatersrand, Johannesburg, South Africa.

High resolution X-ray computed tomography and 3D image analysis and quantification

The UG1 sample was scanned using the Zeiss Versa XRM 520 3D x-ray microscope installed at the Australian Resources Research Centre (CSIRO Mineral Resources, Kensington, Western-Australia). The instrument was set to maximize the contrast between chromite and silicates (plagioclase and pyroxene) present in the sample. Two scans at a voxel size of 5 µm were performed along the vertical axis of the samples and were stitched in 3D to maximize the volume of sample used for further analysis. A total of 1601 projections were recorded over 360° degrees rotation for each scan and were used for volume reconstruction. Beam hardening and ring artefacts were minimized during data acquisition and corrected (if necessary) during image reconstruction. After reconstruction, the sample is represented by a regular grid (1998 x 2046 x 3748 voxels) where each voxel has a unique greyscale value. Chromite, plagioclase and pyroxene were segmented from the volumes using a 3D gradient watershed algorithm\textsuperscript{60} to produce binary images. The separation of touching chromite crystals in 3D was done using a modified version of the algorithm used to separate touching chromite in komatitites\textsuperscript{61} and chromite from the normal Merensky Reef\textsuperscript{62} using Avizo2020.1\textsuperscript{TM} and Matlab\textsuperscript{TM} software. The
shape and size characteristics of chromite network and individual grains were computed to
provide quantitative measure of chromite grains above 15 µm equivalent sphere diameter
(ESD). Chromite grains were defined as touching each other using an 18-voxel connectivity
threshold (i.e., voxels are connected if their faces or edges touch). The coordination number of
each chromite grains (i.e., the total number of other chromite grains touching in 3D a given
grain) was also calculated. All results are summarized in Figs. 2 and 3.

Random packing simulation

The simulated Random Packing of UG1 chromite was generated using the Discrete Element
Method as described in ref 63. The individual chromite particles were modelled as spheres with
a size distribution as measured from the UG1 chromite sample ranging from 15 µm to 420 µm,
and an interparticle friction coefficient of 0.9. The simulation box has dimensions 5 mm by 2
mm by 2 mm, with periodic boundaries in the two directions normal to gravity. Particles are
initially distributed randomly in the simulation box and then allowed to slowly settle under
gravity subject to a Stokes’ drag force to form a random loose packing 43. The packing is then
analysed to determine the packing density and the distribution of inter-particle contacts
between particles.

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Data availability

The authors declare that all relevant data are available within the article and its Supplementary Information Files.

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Author Contributions

R.L. and S.C. undertook field work, mapping and geochemical sampling of outcrops as well as conceptualized the original idea and wrote a draft of the paper. V.R. and K.J. performed X-ray microtomography of samples and data processing, took part in conceptualizing the original idea and participated to the text writing. B.G. collected, processed, and analysed the HRXCT data. G. D., P. C. and S.B. have conducted the randomised loose packing simulation of chromite grains in the UG1 chromitite. S.B. and I.C. participated in interpretation of data, editing the paper as well as in improving clarity of figures. All co-authors discussed the results and problems and contributed to producing a final draft for peer reviews.

Supplementary Information

is available in the online version of the paper.

Competing interests

The authors declare no competing interests.
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