Momentum anisotropy in nuclear collisions from quantum mechanics

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We point out that the intrinsic relationship between space and momentum in quantum physics through the uncertainty principle has potential implications for momentum anisotropy in heavy-ion collisions. Using a harmonic oscillator potential we calculate the elliptic anisotropy and find it to be sizeable compared to elliptic flow measurements in nuclear collisions. Our results question the validity of the completely hydrodynamic interpretation of anisotropic flow data, and highlight the importance of including quantum physics in hydrodynamic calculations which has largely been neglected so far.

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Introduction. One of the most spectacular observations in high-energy nuclear collisions is the so-called elliptic flow \[^1\]. This “left-right” versus “up-down” momentum asymmetry of particle production in the plane transverse to the colliding beams, commonly characterized by a Fourier coefficient \(v_2 = \langle \cos 2\phi \rangle\), is understood as the collective dynamical response of hot and dense plasma of quarks and gluon created in a spatially asymmetric overlap region of target and projectile in the collision. In the hydrodynamic picture \[^2\], elliptic flow is generated by spatially asymmetric pressure gradients in the system, resulting in nonzero final momentum anisotropy \((v_2)\) even if the initial momentum distribution is isotropic azimuthally.

Recently, a similar anisotropy has been seen in proton-nucleus \[^4\] and even high-multiplicity proton-proton collisions \[^4\] at the Large Hadron Collider. Hydrodynamic calculations \[^4\] of these collisions appear to qualitatively and semi-quantitatively reproduce the observations. It is astonishing, however, that a macroscopic theory like hydrodynamics applies so well at such short length and time scales of \(O(1\, fm)\). One might expect quantum mechanics to also play a role. The intrinsic quantum connection between momentum and coordinate spaces through the Heisenberg uncertainty relation implies that spatially anisotropic systems are in general anisotropic in momentum as well. It has been stressed in atomic Bose-Einstein condensation studies, for instance, that this is responsible for causing momentum distributions to have the opposite anisotropy compared to the trapping potential anisotropy (see, e.g., Ref. \[^5\]).

The following estimates this intrinsic momentum asymmetry for thermally equilibrated particles trapped in an anisotropic potential. We simplify the analysis by assuming nonrelativistic motion in either an exactly solvable harmonic oscillator or else a square well potential. We simplify the analysis by ignoring \[^4\]. The trap is then suddenly removed \((V \rightarrow 0)\), without affecting the wave function of the system.

Classically, the momentum distribution \[^3\], \(\frac{dN}{dp}\), \(\frac{dN}{dp} = N \int \frac{d\mathbf{p}}{dp} e^{-H_1(p,r)/T} = N \int \frac{d\mathbf{p}}{dp} e^{-K(p)/T}\), is independent of the potential \(V\), and so it is isotropic at all temperatures as long as \(K\) is isotropic. Quantum mechanically, however, the same system exhibits momentum anisotropy in general even if the kinetic energy \(K\) is isotropic. This is apparent in the \(T = 0\) single-particle limit \((N = 1)\) where the uncertainty relation implies that \(\Delta p_x \approx 1/\Delta x\), \(\Delta p_y \approx 1/\Delta y\) for the ground state.

Single-particle system. For one particle \((N = 1)\) at temperature \(T\) the spatial density is given by the canonical

\[ \rho(r) \equiv \frac{dN}{dr} = \frac{1}{Z} \sum_j |\psi_j(r)|^2 e^{-E_j/T} \]

with \(Z \equiv \sum_j e^{-E_j/T}\). The sums are over the complete orthonormal set of eigenstates of \(H_1\), and \(E_j\) and \(\psi_j(r) = \langle r|j\rangle\) with \(r = (r_x,r_y)\) are the energy and coordinate-space wave function for eigenstate \(|j\rangle\). The distribution of momenta \(p = (p_x,p_y)\) is given by the analogous sum

\[ f(p) \equiv \frac{dN}{dp} = \frac{1}{Z} \sum_j |\psi_j(p)|^2 e^{-E_j/T} \]

where \(\psi_j(p) = \frac{1}{Z} \int dr e^{-ipr}\psi_j(r)\) is the Fourier transform of \(\psi_j(r)\).

Next consider the nonrelativistic \(K(p) = (p_x^2 + p_y^2)/2M\) with a harmonic oscillator potential \(V(r) = M(\omega_x^2 r_x^2 + \omega_y^2 r_y^2)/2\), where \(M\) is the particle mass. The momentum anisotropy is characterized by the parameter \(\tilde{v}_2 \equiv \langle \cos(2\phi) \rangle = (p_x^2 - p_y^2)/(p_x^2 + p_y^2)\), and the initial spatial
eccentricity by \( \varepsilon \equiv \langle r_y^2 - r_x^2 \rangle / \langle r_x^2 + r_y^2 \rangle \). At temperature \( T \) the averages that appear in \( \bar{v}_2 \) and \( \varepsilon \) are readily calculated (\( i = x, y \)):

\[
\langle p_i^2 \rangle = \frac{\hbar^2}{2} \coth \frac{\omega_i}{2T} , \quad \langle r_i^2 \rangle = \frac{\hbar^2}{2M \omega_i} \coth \frac{\omega_i}{2T} .
\]

(3)

Figure 1 illustrates the spatial trap size and eccentricity dependence, general features of the average elliptic anisotropy, for particle mass \( M = 0.3 \) GeV and temperature \( T = 0.2 \) GeV. This mass value can be thought of as a constituent quark, or a particle with mass \( M \approx T \). Here \( \bar{v}_2 \) drops monotonically with system size, but increases monotonically with eccentricity. For a system of \( \langle r_x^2 \rangle^{1/2} \approx 1 \) fm size, \( \bar{v}_2 \approx O(10^{-2}) \). Dotted lines correspond to the remarkably accurate approximation in Eq. (4).

It is instructive to see how the classical limit is approached at high temperature. The relevant parameters are \( \omega_i/T \), namely the ratios of the energy level spacing to the typical energy. For \( \omega_i \lesssim T \) Taylor expansion of Eq. (3) directly relates the average anisotropy to the spatial eccentricity \( \varepsilon \):

\[
\bar{v}_2 \approx \frac{\hbar^2}{12k_B TM \langle r_x^2 \rangle} \cdot \frac{\varepsilon}{1 + \varepsilon}
\]

(4)

using traditional units for clarity. The same \( \Delta p_x \Delta x/\hbar \) combination appears here as in the uncertainty relation, since near the classical limit \( \sqrt{\langle p_x^2 \rangle} \approx \sqrt{k_B TM} \approx \hbar p_{ih} \) is just the typical thermal momentum. The average elliptic anisotropy is roughly one-tenth of the eccentricity when \( \bar{r}_0^2 \langle r_x^2 \rangle = \hbar^2 \).

A crosscheck of these features using an infinite square-well potential shows that the onset of classical behavior is controlled by the same scale \( \bar{r}_0^2 \langle r_x^2 \rangle \), but unlike the harmonic oscillator results the square root of the scale now appears. For example, near the classical limit

\[
\bar{v}_2 \approx \sqrt{\frac{\pi}{96TM \langle r_x^2 \rangle}} \left( 1 - \sqrt{\frac{1 - \varepsilon}{1 + \varepsilon}} \right) \approx \frac{0.18}{\sqrt{TM \langle r_x^2 \rangle}} \frac{\varepsilon}{1 + \varepsilon/2} .
\]

(5)

The reason behind the slower approach is that for the infinite square well \( TM \langle r_x^2 \rangle \) corresponds to the ratio of temperature to gap between the lowest two energy states of the system; unlike the harmonic oscillator this is not indicative of the level spacing compared to \( T \) (since levels grow quadratically).

A more differential measure of momentum anisotropy is provided by the Fourier coefficients \( v_n(p_T) \equiv \langle \cos n\phi \rangle_{p_T} = \frac{1}{\Gamma(2\pi)} \int_0^{2\pi} d\phi \cos n\phi f(p_T, \phi) \), where \( p_T = \sqrt{p_x^2 + p_y^2} \) and \( \phi = \arctan(p_y/p_x) \) are the magnitude and azimuth of the momentum vector. In particular, \( v_2(p_T) \) describes the momentum dependence of “elliptic” anisotropy. The calculation of \( v_n(p_T) \) is straightforward, so we only highlight a few key points. Eigenstates are labelled by a pair of non-negative integers, \( |j| \equiv |nm| \), the wave function factorizes \( \psi_{nm}(r) = \psi_n(r_x)\psi_m(r_y) \), and energy is additive \( E_{nm} = E_n + E_m = (n + \frac{1}{2})\omega_x + (m + \frac{1}{2})\omega_y \). Therefore, the canonical partition sum, wave function Fourier transform, momentum distribution, and particle density all factorize, also for the square well, i.e., \( Z = Z_x Z_y \), \( |\psi_{nm}(p)|^2 = |\psi_n(p_x)|^2|\psi_m(p_y)|^2 \), \( f(p) = g_x g_y |\psi_m(p_y)|^2 \), and \( \rho(r) = g_x g_y |\psi_n(r_x)|^2 \). In Eq. (2) the Fourier transform involves the same Hermite polynomials as those in the coordinate-space eigenstates \( |nm| \), while the canonical sums of \( |\psi_n(x)|^2 \) and \( |\psi_n(p)|^2 \) are doable using Mehler’s formula [11].

The density and momentum distributions both turn out to be (asymmetric) Gaussians

\[
\rho(r) \propto \exp \left( -\sum_i \frac{\xi_i^2}{2} \right) , \quad f(p) \propto \exp \left( -\sum_i \frac{p_i^2}{2\xi_i^2} \right) .
\]

(6)

An intuitive reason for why Gaussians appear is that for \( \omega_x = \omega_y \) there is azimuthal symmetry, and a product form such as \( f(p) = g_x g_y |\psi_m(p_y)|^2 \) is azimuthally symmetric only for Gaussian \( g \) [12].

In the \( T \to \infty \) limit momentum anisotropy vanishes because both \( \langle p_x^2 \rangle \) and \( \langle p_y^2 \rangle \) approach \( TM \) and we recover the classical Maxwell-Boltzmann distribution. At any finite \( T \), however, the tail of the momentum distribution is always anisotropic. Deviations from classical Maxwell-Boltzmann are a factor of two or more when either momentum component exceeds approximately \( \sqrt{24 \ln 2} p_{ih} T/\omega \approx 4\sqrt{TM} p_{ih} L^2 / p_{ih} \), where \( L^2 \) and \( \omega \) are the mean-squared system size and trap strength in the corresponding direction. The system is classical over a wider range of \( p/p_{ih} \), the heavier the particle, the larger the temperature, or the larger the system size.

Different differential coefficients readily follow using Eq. (6):

\[
v_{2n}(p_T) = h_n \left( \frac{p_i^2}{2MT} (S_i - S_x) \right) , \quad S_i = T \omega_i / \tanh \omega_i / 2T .
\]

(7)
where $h_n(x) \equiv I_n(x)/I_0(x)$ is a ratio of modified Bessel functions of the first kind. Even harmonics are all monotonic in $p_T$, start out at low $p_T$ as $v_{2n} \propto p_T^{2n}$, and approach unity at infinite momentum. A useful approximation for low $p_T$ and $\omega_i \lesssim T$ is $v_2(p_T) \approx \hat{v}_2 p_T^2/4MT$, i.e., $\hat{v}_2 \approx v_2(2p_{th})$ (note the factor 4MT and not 2MT here). Also, $v_2(p_T) \approx [v_2(p_T)]^n/n!$ in the same limit. Odd harmonics vanish by symmetry for our even anisotropic trap, but those can be produced with a less symmetrical anisotropic trap.

Multi-particle system. The calculation can be easily extended to multi-particle systems in a grand canonical description, i.e., at fixed temperature $T$ and chemical potential (Fermi energy) $\mu$. Results for the number and momentum densities follow by replacing the probability potential $(Fermi$ energy) $\mu$ by $\gamma/[e^{(E_j-\mu)/T}+1]$, where $a = 1$ for fermions, $-1$ for bosons, and 0 for Boltzmann statistics, with inert internal degrees of freedom accounted for via the degeneracy factor $\gamma$.

With the additional scale $\mu$, departure from classical behavior occurs not only when $\omega_i \gtrsim T$ (discrete levels matter) but also when $\mu \gg T$ (Bose/Fermi statistics is important). The momentum anisotropy discussed here is due to the former. Classical phase space integrals, outlined at the end of the Introduction, give zero anisotropy even with Bose/Fermi distribution instead of Boltzmann.

Figure 2 shows the temperature and transverse momentum dependence of the elliptic anisotropy for multi-particle systems at $\mu = 0$, for the same $M = 0.3$ GeV as in Fig. 1. As expected, $\bar{v}_2$ in the left plot decreases monotonically with temperature, and the anisotropy is noticeably higher for Bose statistics than for Boltzmann or Fermi statistics. This is because lower-energy states are more anisotropic (both spatially and in momentum space) on average.

The right plot shows the $p_T$-differential elliptic anisotropy. Boltzmann statistics leads to the analytic result (Eq. 7), while Bose (Fermi) gives considerable (modest) deviations in $\bar{v}_2$ compared to Boltzmann. The curves differ even at high $p_T$ because the relative weights of various anisotropy states are different among the three types of statistics. Therefore, depending on statistics, slightly different trap parameters $\omega_i$ are needed in order to have the same system dimensions $\langle r_T^2 \rangle_{MT}^{1/2}$. For all three statistics, there is considerable “intrinsic” elliptic anisotropy at intermediate and high $p_T$.

The same features are present for the infinite square well potential, with even larger anisotropies at low and intermediate $p_T$ than for the oscillator potential. At high momenta $v_2$ saturates below unity for the square well, unlike for the oscillator potential, because the momentum distribution has a power-law tail.

To illustrate how this intrinsic anisotropy from quantum physics might be relevant for heavy-ion collisions, Fig. 2 shows the $p_T$ dependences of $v_2$ for pions and protons in the left panel and the impact parameter dependence of $\bar{v}_2$ for pions in the right panel, using the Au+Au collision initial geometry. The binary collision transverse profile sets the mean-squared dimensions. The left panel shows considerable deviations from the measured typical $\bar{v}_2 \approx 5\%$ in Au+Au collisions. In addition to $\epsilon$, the quantum effect also depends on the size of the collision zone, i.e., $v_2$ quickly dies off towards central, small impact parameter collisions. This differs from hydrodynamics where the impact parameter dependence is nearly proportional to eccentricity. At large $b$ the collision approaches the limiting case of nucleon-nucleon interactions, where spatial
anisotropy disappears in our implementation of the geometry. This explains the decreasing trend of $\vec{v}_2$ at large impact parameters.

**Anisotropy from classical fields.** Finally, note that a similar anisotropy effect should arise in any wave dynamics. E.g., consider a solution to the source-free Maxwell equations. At $t = 0$ we can write

$$
\begin{pmatrix} E(r) \\ B(r) \end{pmatrix} = \text{Re} \left[ \int d^3p \frac{\sqrt{\omega_p}}{2\pi} \begin{pmatrix} C(p) \\ C_B(p) \end{pmatrix} e^{ipr} \right]
$$

(8)

where $\omega_p \equiv |p|$, $p \cdot C(p) \equiv 0$, and $C_B(p) = p \times C(p)/\omega_p$. The total electromagnetic energy and momentum are

$$
\int d^3r \frac{1}{8\pi} [E^2(r) + B^2(r)] = \int d^3p \omega_p |C(p)|^2,
$$

(9)

$$
\int d^3r \frac{1}{4\pi} E(r) \times B(r) = \int d^3p |C(p)|^2,
$$

(10)

so we can interpret $f(p) = |C(p)|^2$ as the momentum distribution of (on-shell) photons. If the energy density is spatially anisotropic, in general the momentum distribution will be anisotropic as well. It would be interesting to check to what extent differential elliptic anisotropy from the classical Yang-Mills “color glass” approach \[13\] can be attributed to the Fourier transform of the initial shape of the overlap region.

**Discussions and Conclusions.** We conclude that quantum mechanics can be relevant for the momentum anisotropy in hadron and nuclear collisions. Our estimates indicate that the effect is significant and may not be neglected. Our results are necessarily qualitative: i) we used simple harmonic oscillator and nonrelativistic quantum mechanics; ii) we assumed a thermalized system at temperature $T \sim 0.2$ GeV; iii) we used $M \sim 0.3$ GeV with constituent quarks in mind—will the resulting hadron anisotropy be actually larger in a coalescence picture \[14\]? iv) We estimated pion and proton $v_2$ with a bag of hadrons in mind at initial Au+Au encounter—how would subsequent dynamical evolution such as hydrodynamic expansion affect the anisotropy? v) We used a single statistical ensemble—what if we only have local thermalization \[15\], i.e., many small thermal systems distributed spatially? It is not clear, even qualitatively, what fraction of the initial quantum anisotropy remains after hydrodynamic expansion—which is presumably needed, in addition, to produce the measured final-state anisotropy—or how strong quantum effects would be if we imposed quantum uncertainty right at the late breakup stage of hydrodynamic evolution. Many of the open questions concern how to properly include quantum mechanics in relativistic, dissipative hydrodynamics. But the important message of our work is that quantum physics must be present and needs rigorous consideration, especially for small systems such as proton-proton and proton-nucleus collisions.

Strong elliptic anisotropy was also observed in cold atom systems released from anisotropic trap \[16\]. For those systems, the quantum mechanics effect is minuscule and, therefore, the elliptic anisotropy is due to interactions during expansion. E.g., the experiment by O’Hara et al. \[16\] confined fermionic lithium atoms at $T \sim 1$ µK, Fermi temperature $T_F \sim 10^7$, in a trap of dimensions $r_x \sim 20$ µm and $r_y \sim 100$ µm, for which we estimate using Eq. (1) a quantum anisotropy $\vec{v}_2 \sim 10^{-5}$. It is small because the cold atom system’s intrinsic momentum quantum $h/r_x \sim 10^{-8}$ MeV/c is negligible compared to the typical momentum $\sqrt{kbTM} \sim 10^{-6}$ MeV/c. The intrinsic energy quantum $h^2/Mr_y^2 \sim 10^{-20}$ MeV is also much smaller than the temperature $k_BT \sim 10^{-16}$ MeV. In this sense these cold atoms are much “hotter” than the hot quark gluon plasma, where the temperature is comparable to the intrinsic momentum and energy scales. The cold atom system is thus classical in its expansion dynamics (despite being in the $S$-wave scattering limit), while the quark gluon plasma may exhibit intrinsic quantum features. It would be interesting to do an experiment with cold lithium atoms in a $\sim 100$ times smaller trap or at nano-Kelvin temperatures (conditions closer to the few-body limit \[13\]), or a trapped cold electron system.

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\ln g(p). Substituting \( p_y = p_x + \delta \) and Taylor expanding in \( \delta \) gives the differential equation \( p G''(p) = G'(p) \) with solution \( G'(p) \propto p \), so \( g(p) \) is Gaussian.

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