Saturation properties and incompressibility of nuclear matter:
A consistent determination from nuclear masses

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Abstract

Starting with a two-body effective nucleon-nucleon interaction, it is shown that the infinite nuclear matter model of atomic nuclei is more appropriate than the conventional Bethe-Weizsacker like mass formulae to extract saturation properties of nuclear matter from nuclear masses. In particular, the saturation density thus obtained agrees with that of electron scattering data and the Hartree-Fock calculations. For the first time using nuclear mass formula, the radius constant $r_0=1.138$ fm and binding energy per nucleon $a_v = -16.11$ MeV, corresponding to the infinite nuclear matter, are consistently obtained from the same source. An important offshoot of this study is the determination of nuclear matter incompressibility $K_{\infty}$ to be $288\pm 28$ MeV using the same source of nuclear masses as input.

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I. INTRODUCTION

The binding energy, saturation density and compression modulus of infinite nuclear matter are fundamental constants of nature. Traditionally, the first two quantities termed as saturation properties are determined from two different sources, namely the volume coefficient $a_v$ of the Bethe-Weizsacker (BW) like mass formulae and the electron scattering data on heavy nuclei respectively. Although the Coulomb coefficient $a_C(= 0.6 e^2/r_0)$ in BW-like mass formulae specifies the density $\rho = 3/(4\pi r_0^3)$, it is not accepted as the density of nuclear matter. This is because, the corresponding radius constant $r_0 \approx 1.22 \text{ fm}$ obtained in a totally free fit is much higher than the value $1.12 - 1.13 \text{ fm}$ obtained from the electron scattering data on heavy nuclei and Hartree-Fock calculations. As yet no mass formula fit to nuclear masses has yielded a value of $r_0$ in this range. This is the so-called “$r_0$– paradox”, which has been a subject of investigation over the years by many. Since, the two properties are highly inter-related, the above constrained practice of their determination from two different sources, has been a serious discomfeature in our understanding of nuclear dynamics. Coupled to this, the incompressibility of nuclear matter has posed a much serious problem with regard to its determination, both theoretically and experimentally.

In this work, we report our attempt to determine all the three properties of nuclear matter using a single model, and one kind of experimental data, namely the nuclear masses, which are abundant in nature and are the best known properties of nuclei. We use the infinite nuclear matter (INM) model based on the generalised Hugenholtz-Van Hove (HVH) theorem of many-body theory, whose success has been well tested through its unique ability to predict masses of nuclei far from stability, masses of $Na$ isotopes and other light nuclei, and finally through the 1986-87 mass predictions of the entire periodic table. In the formulation of INM model, it was claimed that this model is more suitable than the traditional (BW) ones to extract the properties of nuclear matter, as it is exclusively built in terms of infinite nuclear matter at ground-state. In the present work, we have improved the model and show conclusively, starting from two-body
effective interaction within the energy density formalism, that the saturation properties
derived through this improved model are closer to the true properties of the nuclear matter
than those derived using BW model based mass formulae. Then this model is fitted to
the experimental masses, which yields a value of $r_0$ to be 1.138 fm, in close agreement
with that obtained from the electron scattering data, and with the empirical value found
through many-body mean-field approaches[9]. Further, using these saturation properties
determined from the same set of data on nuclear masses, we arrive at a value of about
288 MeV for the incompressibility $K_\infty$, which is of equal fundamental importance in the
realm of nuclear physics and astrophysics.

In Sec. II, the improvements we have made in the INM model are presented. In Sec.
III, we show at a microscopic level that the improved INM model is more appropriate than
the BW-like model for the extraction of saturation properties of infinite nuclear matter
from the nuclear masses. Determination of such properties are presented in Sec. IV. In
Sec.V, the value of nuclear compression modulus is determined from the nuclear masses.
Finally, we conclude in Sec.VI.

II. THE IMPROVED INM MODEL

We recall here the essential features of the INM model[8] which we have now improved.
In this model, the ground-state energy $E^F(A, Z)$ of a nucleus(A,N,Z) with asymmetry $\beta$
is considered equivalent to the energy $E^S$ of a perfect sphere made up of infinite nuclear
matter at ground-state density with same asymmetry $\beta$ plus the residual energy $\eta$, called
the local energy, which contains all the characteristic contributions like shell, deformation
etc. So,

$$E^F(A, Z) = E_{INM}^S(A, Z) + \eta(A, Z)$$

with $E_{INM}^S(A, Z) = E(A, Z) + f(A, Z)$, where

$$f(A, Z) = a_s^I A^{2/3} + a_C^I (Z^2 - 5(3/(16\pi))^{2/3} Z^{4/3}) A^{-1/3} + a_{ss}^I A^{2/3} \beta^2 + a_{cv}^I A^{1/3} - \delta(A, Z)$$

denotes the finite-size effects and $E(A, Z)$ is the energy of the infinite part. The super-
script $I$ refers to the INM character of the coefficients. Here $a_s^I$, $a_C^I$, $a_{ss}^I$ and $a_{cv}^I$ are
the surface, Coulomb, surface-symmetry and curvature coefficients. $\delta(A, Z)$ is the usual
pairing term, given as

\[ \delta(A, Z) = \begin{cases} +\Delta A^{-1/2} & \text{for even } - \text{even nuclei} \\ 0 & \text{for odd } - A \text{ nuclei} \\ -\Delta A^{-1/2} & \text{for odd } - \text{odd nuclei} \end{cases} \]

Eq.(1) now becomes,

\[ E^F(A, Z) = E(A, Z) + f(A, Z) + \eta(A, Z) \]

Thus, the energy of a finite nucleus is given as a sum of three distinct parts; an infinite part \( E(A, Z) \), a finite-size component \( f(A, Z) \) and a local energy part \( \eta(A, Z) \). The term \( E(A, Z) \) being the property of infinite nuclear matter at ground-state, will satisfy the generalised HVH theorem[7].

\[ E/A = [(1 + \beta)\epsilon_n + (1 - \beta)\epsilon_p]/2 \]  

where \( \epsilon_n = (\partial E/\partial N)_Z \) and \( \epsilon_p = (\partial E/\partial Z)_N \) are the neutron and proton Fermi energies respectively. Using Eq.(3), the INM Fermi energies \( \epsilon_n \) and \( \epsilon_p \) can be expressed in terms of their counterparts for finite nuclei as

\[ \epsilon_n = \epsilon_n^F - (\partial f/\partial N)|_Z - (\partial \eta/\partial N)|_Z; \epsilon_p = \epsilon_p^F - (\partial f/\partial Z)|_N - (\partial \eta/\partial Z)|_N \]  

where \( \epsilon_n^F = (\partial E^F/\partial N)_Z \) and \( \epsilon_p^F = (\partial E^F/\partial Z)_N \). Using (3) and (5), Eq.(4) is rewritten as

\[ E^F/A = [(1 + \beta)\epsilon_n^F + (1 - \beta)\epsilon_p^F]/2 + S(A, Z) \]

where, \( S(A, Z) = f/A - (N/A)(\partial f/\partial N)_Z - (Z/A)(\partial f/\partial Z)_N \) is a function of all the finite-size terms \( a_s^I, a_c^I, a_{ss}^I \) and \( a_{cv}^I \), which are global in nature. As discussed earlier[7], the \( \eta \) terms in Eq.(6) drops out, which plays a crucial role in the success of INM model, and whose validity has been amply demonstrated[3, 8]. It must be noted that Eq.(6) does not contain the infinite part \( E \) as well as the \( \eta \) terms. Thus, through Eq.(6), the decoupling of the finite component \( f \) from the infinite one \( E \) has been acheived. The coefficients \( a_s^I, a_c^I, a_{ss}^I \) and \( a_{cv}^I \) can be determined by fitting \( S(A, Z) \) function with the combination of
data $E_F/A - [(1 + \beta)\epsilon_n^F + (1 - \beta)\epsilon_p^F]/2$ obtained from the nuclear masses. We would like to mention here that in the earlier work[6], due to the use of the expressions for Fermi energies, $\epsilon_n^F = E_F(N, Z) - E_F(N - 1, Z)$ and $\epsilon_p^F = E_F(N, Z) - E_F(N, Z - 1)$, a small contribution $a_a(\beta^2 - 1)/(A - 1)$ survives (of the order of $a_a/A$) in Eq.(6), whereas in the present work by using the better formulae

\begin{align*}
\epsilon_n^F &= \frac{\partial E_F}{\partial N} \bigg|_{\frac{Z}{2}} \frac{1}{2} \left[ E_F(A + 1, Z) - E_F(A - 1, Z) \right] \\
\epsilon_p^F &= \frac{\partial E_F}{\partial Z} \bigg|_{\frac{N}{2}} \frac{1}{2} \left[ E_F(A + 1, Z + 1) - E_F(A - 1, Z - 1) \right]
\end{align*}

(7)

the following important improvements are achieved.

(i). The decoupling of the infinite part (asymmetry term) from the finite part in Eq.(6) occurs up to an order of $a_a/A^2$, which can be considered perfect at the numerical level.

(ii). The pairing term $\delta$ contained in $f$ effectively drops out in Eq.(6), thereby rendering the determination of other coefficients with greater accuracy due to less correlation.

(iii). The exchange Coulomb term of the standard form $O(Z^{4/3}A^{-1/3})$ (2) exactly cancels in Eq.(6). This cancellation gives rise to a more reliable determination of INM saturation density through $a_C^I$.

(iv) The other factors which might affect the determination of density such as proton-form-factor $[O(Z^2/A)]$ and Nolen-Schiffer anomaly $[O(\beta A)$, referred to as charge-asymmetry energy] also cancel exactly.

Thus all the finite-size coefficients contained in $S(A, Z)$, which are global in nature, are determined from nuclear masses by a fit to Eq.(6). Now, of the three distinct parts of the energy $E_F$ of a finite nucleus(3), the infinite part $E$ and the local energy part $\eta$ remain to be determined. The infinite part $E$ must satisfy the generalised HVH theorem (4), whose solution is of the form,

$$E = -a_v I A + a_a I \beta^2 A$$

(8)
where $a^I_v$ and $a^I_a$ are the global parameters which can be identified as volume and symmetry coefficients corresponding to INM. Using (5) and (8) in the right and left hand sides of Eq.(4) respectively, one obtains

$$-a^I_v + a^I_a \beta^2 = \frac{1}{2} \left[ (1 + \beta)\epsilon^F_n + (1 - \beta)\epsilon^F_p \right] - \left[ \frac{N}{A} \frac{\partial f}{\partial N} \right]_{\text{Z}} + \left[ \frac{Z}{A} \frac{\partial f}{\partial Z} \right]_{N}$$

(9)

where the contribution from the local energy part (of the order of $\eta/A$) is neglected, which in the limit of large $A$ goes to zero. Since $f$ is known from Eq.(6), the above equation can be used to determine the two parameters $a^I_v$ and $a^I_a$ with the combination of data $\left[ (1 + \beta)\epsilon^F_n + (1 - \beta)\epsilon^F_p \right] / 2$ obtained from nuclear masses. Thus all the global parameters are determined essentially in two fits: Eq.(6) determines the finite-size coefficients like $a^I_s$, $a^I_C$, etc, and Eq.(9) determines the INM coefficients $a^I_v$ and $a^I_a$. Since the present study is intended for the determination of the properties of nuclear matter, we do not discuss the determination of $\eta$ and consequently the masses, the details of which can be seen in Refs.[6, 8].

III. IMPROVED INM MODEL VERSUS BW MODEL

In this section, we would like to make a comparative study of the improved INM model and BW model, in regard to their suitability for the determination of saturation properties of infinite nuclear matter from nuclear masses.

As noted in the introduction, it has not been possible to determine both the energy and density of infinite nuclear matter in the BW model based mass formulae. Further, it has been hoped only that the volume coefficient determined in the BW model corresponds to nuclear matter at ground state. On the other hand, in the INM model, this fact has been ensured by the explicit use of HVH theorem, which is valid only at the ground-state of infinite nuclear matter. Since in this model, the binding energy of a nucleus is written in terms of the properties of INM, it is expected that the INM model is well equipped to extract infinite nuclear matter properties from nuclear masses. We demonstrate this by predicting the a priori known INM properties for a given effective interaction.

In this regard, we make use of the extended Thomas-Fermi(ETF) calculation[10] of nuclear binding energies with Skyrme-like forces, which over the years has been firmly established. In such calculations, one obtains the smooth part of the energy corresponding
to the liquid-drop nature of nuclei. This smooth part, hereafter referred to as macroscopic
part describes the energy $E_{INM}^S$ of the INM sphere as defined in Eq.(1). For the purpose
of making a comparative study of the INM model and the BW-like model, the appropriate
BW mass formula is

$$E_{BW} = -a_L A + a_L A^{2/3} + a_C (Z^2 - 5(\frac{3}{16\pi})^{2/3}Z^{4/3})A^{-1/3} + a_L A^{2/3} + a_L A^{1/3} - \delta(A, Z)$$

(10)

In the case of the ETF calculations, nuclear curvature coefficient comes out to be about
10 MeV as against the BW-like mass formula fit to real nuclei, which gives a value close
to zero. For this reason, we have included higher-order terms like curvature and surface-
symmetry terms in both the INM and BW models.

The macroscopic or the ETF nuclear ground state energies used here for the compar-
ative study of the INM and BW models are taken from the calculations of Aboussir et al.[11]. In their calculation, they used a generalised Skyrme force SkSC4 of the form

$$v_{ij} = t_0[1 + x_0 P_a \delta(\vec{r}_{ij})]$$

+ $t_1(1 + x_1 P_a)[p_{ij}^2 \delta(\vec{r}_{ij})^2 + h.a.] / 2h^2$

+ $t_2(1 + x_2 P_a) p_{ij} \cdot \delta(\vec{r}_{ij}) p_{ij} / h^2$

+ $(t_3 / 6)(1 + x_3 P_a)[\rho_{qi}(\vec{r}_i) + \rho_{qj}(\vec{r}_j)] \gamma \delta(\vec{r}_{ij})$

+ $(i / h^2) W_0(\vec{\sigma}_i + \vec{\sigma}_j) \cdot \vec{p}_{ij} \times \delta(\vec{r}_{ij}) p_{ij}$. 

Then, the macroscopic part of the total energy for a given nucleus is calculated using
the energy density formalism, i.e. $E = \int \mathcal{E}(\vec{r}) d^3r$, where $\mathcal{E} = \tau(\vec{r}) + v(\vec{r})$. The potential
energy density $v$ is derived using the two-body force given above. For the kinetic part $\tau$, they use the full fourth order ETF kinetic functional[10]. It must be noted that realistic
nuclear ground state energies contain shell effects. To incorporate this characteristic
feature in a self-consistent way, a Hartree-Fock calculation is performed for the same
generalised Skyrme force. Now, using the single particle states obtained within the HF
approximation, the shell corrections can be calculated by directly making use of the
Strutinsky procedure. Then, the total energy is given as the sum of the macroscopic part
and these shell corrections.
We have made an exhaustive study using those macroscopic energies provided by them for 1492 nuclei. We fitted both the INM formulae given by (6) and (9) and the BW one by (10) to the above macroscopic part of the nuclear masses to determine the corresponding global parameters. The results so obtained, together with their respective errors, are given in Table I. The values obtained directly by Aboussir et al with the SkSC4, force performing nuclear matter and semi-infinite nuclear matter calculations for the various coefficients (hereafter referred to as exact values) are also presented in Table I. It is gratifying to find that the values obtained in the INM fit for the principal coefficients like $a_I$, $a_s$ and $a_C$ agree better with the exact values, compared to that of the BW fit. The symmetry coefficient $a_B$ agrees reasonably well with the exact value, although somewhat inferior to the BW value. Even though the agreement of the higher-order terms like surface-symmetry and curvature in the BW-fit agree better, it must be noted that, because of correlations amongst the coefficients, they significantly affect the principal term like surface, and to a lesser extent the other ones also. In case of the INM fit, since the infinite and finite parts are determined in two separate fits, the principle coefficients are not influenced by the higher-order terms. In any case, these two coefficients contribute insignificantly in real nuclei, and are normally ignored. Thus, the saturation properties of nuclear matter, which are a priori known for a given effective interaction like SkSC4 is relatively well reproduced by the INM model than the BW-like model. This gives us more confidence in the INM model in extracting real saturation properties from experimental nuclear masses which is done in the next section.

The success of the INM model over that of the BW-like model is essentially due to the following. As also discussed in Ref.[3], the BW-like mass formulae use only the average property of nuclear matter, namely the average energy per nucleon. However, as demonstrated by Hugenholtz-Van Hove[12], an interacting Fermi system has an additional property namely the single-particle property. Such a system has one true single-particle state i.e. Fermi state, which has infinite life time, while other low lying ones are metastable. In other words, the lifetime of the single particle state approaches infinity in the limit $k \rightarrow k_F$. This important property is additionally taken into account in the INM model,
which is not present in BW-like mass formulae.

IV. DETERMINATION OF NUCLEAR MATTER SATURATION PROPERTIES

Before coming to the actual determination of the various parameters of the INM model, and thereby the saturation properties of nuclear matter, it is essential to assess the relative importance of the possible higher-order terms, which is somewhat different in this model.

The two saturation properties, namely the density $\rho_\infty$ given by $a^I_C$ and the volume energy $a^I_v$, are determined in two different fits; Eqs.(6) & (9) respectively. The first fit determines the crucial quantity $a^I_C$; and hence, it is imperative that we analyse the role of other finite-size effects which may influence the determination of the saturation properties. The finite-size terms which are directly related to the Coulomb effect are: exchange Coulomb, proton-form-factor correction and charge-asymmetry energy. It may be recalled here that in the INM model, the binding energies and Fermi energies are used in the particular combination, $E^F/A = [(1 + \beta)\epsilon^F_n + (1 - \beta)\epsilon^F_p]/2$, in Eq.(6), as dictated by the HVH theorem. As a result, the above stated three effects exactly cancel in Eq.(6) rendering a clean determination of $a^I_C$, and hence the density $\rho_\infty$. This is indeed a very fortunate situation.

The other two higher-order terms which may indirectly affect the value of $a^I_C$ are the curvature $a^I_{cv}$ and surface-symmetry $a^I_{ss}$ coefficients. In real nuclei, the curvature coefficient comes out to be nearly zero, and is normally not included. So, we have dropped it. In regard to the surface-symmetry coefficient $a^I_{ss}$, it has been recognised that its value is somewhat difficult to determine from nuclear masses. Even at the theoretical level, the value of $a^I_{ss}$ determined from various effective interactions differ widely. Further, in the modern BW-like mass formulae, this coefficient is fixed from considerations other than the ground-state nuclear masses, such as fission barrier heights. Since, in the present study, we address ourselves to the determination of the properties of INM at ground-state, it is essential that we only use the ground-state masses, and not any other property which may drift the system from the ground-state and jeopardize the determination of $a^I_v$ and $a^I_C$. Therefore, in the present context, it is proper that the important coefficients are treated as free parameters to be fixed by nuclear masses through Eq.(6).
But, fortunately, the surface-symmetry coefficient $a_{ss}^I$ cancels to a major extent ($\sim 66\%$) due to the particular combination of data used in Eq.(6). Although this does not fully cancel like the exchange Coulomb, proton-form-factor correction terms etc, $a_{ss}^I$ being a second-order term, such cancellation renders it relatively insignificant as compared to $a_{s}^I$. At numerical level, it may be considered to be virtually cancelled. Nevertheless, since our main goal is to determine the saturation properties of nuclear matter, which are of fundamental importance, we are anxious to check if any semblance of survival of $a_{ss}^I$ term can affect the results.

Hence, we carried out calculations retaining this term as a free parameter in our fit to Eq.(6). It is found that while the values of the other coefficients remain almost unaltered, the value of $a_{ss}^I$ widely varies from -30 MeV to -11 MeV with accompanied error of about 50-100% as number of data varies from 1085 to 1371. This fact is also true when one uses presupposed values for $a_{ss}^I$, while other coefficients are being fitted to the 1371 masses. As the value of $a_{ss}^I$ is varied from -10 MeV to -30 MeV, it was found that $\chi_{rms}$ shows a minimum at $a_{ss}^I \sim -12$ MeV. However, this optimum value of $a_{ss}^I$ fluctuates with the variation of the number of data, resulting in no definite value. These two features are reminiscent of its insignificant presence in Eq.(6). Hence we have omitted this term. The same is true also for the curvature term. Therefore, the optimum representation for the finite-size function $f(A)$ defined in Eq.(1) is

$$f(A,Z) = a_{s}^I A^{2/3} + a_{C}^I (Z^2 - 5(3/(16\pi)))^{2/3}Z^{4/3}A^{-1/3}.$$ 

Now coming to the actual determination of the saturation properties of INM, we use all the nuclear masses with experimental error $\leq 60$ keV from the recent mass table of Wapstra and Audi[15]. There are 1371 cases, which have been used in our study. As mentioned earlier the universal parameters in this model are determined in a two-step process. In the first step, we determine the finite-size coefficients $a_{C}^I$ and $a_{s}^I$ by making a least-square fit to Eq.(6) using all the 1371 masses. Then these parameters so determined are further used in the second step to obtain the coefficients corresponding to the infinite part $a_{v}^I$ and $a_{b}^I$, by a fit to Eq.(9) using the same set of data. The $\chi_{rms}$ obtained for these two fits are 371 keV and 372 keV respectively, which are substantially lower than the corresponding ones 460 keV and 506 keV obtained in the earlier study[6, 8]. The
lowering of $\chi_{rms}$ is almost entirely due to the improvements made in the model, and not as the result of the use of recent masses. To check the goodness and the stability of the parameters obtained in our fits, we have carried out five sets of calculations by varying the number of data randomly considered throughout the mass table, choosing them on the basis of experimental error ranging from 20 keV to 60 keV, and these are presented in Table II. One can clearly see that almost all the the four parameters are quite stable in spite of widely varying data. Especially remarkable is the stability of the two crucial nuclear parameters namely, the Coulomb coefficient $a_C^I$, and and the volume coefficient $a_V^I$. It is satisfying to note that the degree of stability of these two important parameters, in which we are specifically interested, is relatively better than in $a_s^I$ and $a_\beta^I$. The final values obtained for these two coefficients with maximum number of data (1371 nuclei), and the corresponding values for $r_0$ and $\rho_\infty$ are:

$$a_V^I = 16.108\text{MeV} \quad \text{and} \quad a_C^I = 0.7593\text{MeV}$$

$$r_0 = 1.138\text{fm} \quad \text{and} \quad \rho_\infty = 0.1620\text{fm}^{-3}.$$

We quote no errors for our parameters as they are firmly determined, say with errors less than 1%. The saturation properties $a_V^I = 18.335 \text{ MeV}$ and $a_C^I = 0.841 \text{ MeV}$, determined earlier[6] are inaccurate due to the use of the expressions $\epsilon_N^F = E_N^F(N, Z) - E_N^F(N-1, Z)$ and $\epsilon_P^F = E_P^F(N, Z) - E_P^F(N, Z-1)$ for finite nuclei Fermi energies. It is indeed remarkable that the saturation density $\rho_\infty = 0.162 fm^{-3}$ and the corresponding $r_0 = 1.138 fm$ found here agree quite well with that obtained from the fit of electron scattering data. This value of $r_0$ is also close to 1.13 fm obtained in the HF studies, which has been widely accepted in literature[4]. It may be noted that our value of $r_0$ is quite similar to the value $1.140 \pm 0.005$ obtained from the fit of nuclear charge radii[6] extracted from the recent electron scattering data[16]. Thus, the two important ground-state properties, i.e $a_V$ and $\rho_\infty$ which are inter-related, are consistently determined from one kind of data using a single model.

**V. INCOMPRESSIBILITY OF NUCLEAR MATTER**

To determine $K_\infty$, we note that INM model determines binding energy per nucleon $a_v$ and saturation density $\rho_\infty$ at ground-state. The BW model based mass formulae give
the value of $a_L^v$ at a different density $\rho_0$ corresponding to their $a_L^c$, since they do not have any ingredient to ensure that these parameters pertain to nuclear matter at ground-state. Hence, using the values of $a_v$ and densities from INM model as well as from BW formula, one can determine $K_\infty$ using the relation

$$a_L^v(\rho_\infty) + (K_\infty/18)(\rho_0/\rho_\infty - 1)^2 = a_L^v(\rho_0)$$

as shown in Ref.[17].

In order to determine the optimum number of parameters in the BW mass formula given by Eq.(10), we have carried out least-squares fit with varying number of parameters, the results of which are presented in Table III. We have used the same 1371 nuclear masses mentioned in our earlier section. It can be seen that the values of the principal five coefficients are not affected when the surface-symmetry $a_{ss}^L$ and curvature $a_{cv}^L$ terms are successively included. Hence, the $a_{ss}^L$ term is well supported and should be retained. The curvature term, in spite of its smallness and relatively large error, can be included as it does not affect the leading terms much. However, the inclusion of the Gauss curvature $a_{gc}^L$ term, the next higher-order term in the model, completely destablises the fit by violently disturbing the leading order coefficients. The surface coefficient has even become negative. This may be due to the very small value[13, 14] of this term $a_{gc}^L$, which is of the order of 6 MeV.

Quite importantly, the above result is contrary to the common belief that the inclusion of more and more higher-order terms in a liquid-drop model like expansion would result in progressively refined values of the leading order terms. Therefore, one should be judicious in retaining higher-order terms in such models. In the present study, we consider Eq.(10) having six parameters to be the optimum representation, where we have dropped the curvature term as it comes out to be nearly zero.

With this view, we carried out a least-square fit to Eq.(10) (without the curvature term) using the same 1371 masses. As in the case of INM model calculation, we have varied the number of data to arrive at stable values of $a_L^v$ and $a_L^c$, with similar accuracies of second and third decimal places respectively, since the value of $K_\infty$ is sensitive to these values. The results are given in Table IV. Now, $K_\infty$ is computed using these values in
Eq. (11) together with the values of $a_1^I$ and $a_1^C$ from Table II for the corresponding set of data, and are presented in the last column of Table IV. It is remarkable that inspite of variation of the number of input data ranging from 1085 to 1371, the value of $K_\infty$ comes out to be in between 288 and 305. The average value thus obtained in Table IV is about 294 MeV which is very close to 288 MeV obtained with the maximum number of data used in the fitting procedure, which further substantiates the stability of our result with respect to the variation of data.

We then attempt to get an estimate of the error in this value of $K_\infty$ arising out of the limitations of the model, which may be due to the inclusion/non-inclusion of higher-order terms like curvature, exchange Coulomb and proton-form-factor. The results of our calculation of $K_\infty$ with inclusion/non-inclusion of these three effects are presented in Table V. We have calculated the error $\chi$ using the expression 

$$\chi^2 = \frac{1}{N} \sum_{i=1}^{N} (K_\infty^i - K_\infty^{opt})^2,$$

where $K_\infty^{opt} = 288$ MeV, $N = 5$ and $i$ stands for the five values, other than $K_\infty^{opt}$, tabulated in Table V. The error thus calculated comes out to be 28 MeV.

The recent BW model based mass formulae usually use presupposed value of $r_0$ determined from other considerations. The one which treats $r_0$ as an adjustable parameter and more or less looks similar to Eq.(10) is by Myers and Swiatecki[11], where $r_0$ is determined by using the data on both the nuclear masses and fission barriers. Using their values of $a_0$ and the density, and the present values of INM, we obtain $K_\infty$ to be about 299 MeV. Hence, we would like to firmly state that, if one allows $r_0$ as a free parameter in the fit to nuclear masses, one would invariably arrive at a value of about 288± 28 MeV for $K_\infty$.

VI. CONCLUSIONS

In conclusion, we have improved the INM model by using better Fermi energies for neutron and proton, which has resulted in a cleaner decoupling of the finite-size effects and the INM part of the ground-state energies of nuclei. Unlike in the BW-like mass formulae, the Coulomb related higher-order terms such as exchange Coulomb, proton-form-factor correction and charge-asymmetry energy cancel exactly, rendering accurate determination of the most important quantity, namely the saturation density. More importantly, we have demonstrated at a fundamental level starting with effective two-body interaction,
the appropriateness of the INM model over that of the BW-like models to determine the ground-state properties of INM.

The saturation density $\rho_\infty$ and binding energy per nucleon $a_v$ of nuclear matter, the two highly inter-related quantities, are extracted consistently for the first time from a single source, i.e. nuclear masses, through a mass formula. It is particularly satisfying to find that the radius constant corresponding to $\rho_\infty$ determined here agrees quite well with that obtained from electron scattering data and Hartree-Fock calculations. These have been possible because of taking into account additionally the single particle property of the interacting Fermi system through the use of the generalised HVH theorem in the INM model. Thus, the $r_0-$ anomaly is resolved here satisfactorily.

An important offshoot of this study is the determination of the value of nuclear matter incompressibility starting from nuclear masses, which are the best measured and most abundant data in nuclear physics. The value so obtained for $K_\infty$ is $288 \pm 28$ MeV. We finally comment that inclusion of surface-symmetry term $a_{ss}I$ leads to fluctuation of the value of $K_\infty$ to a larger side of the above value.

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Table Captions

Table I. Values obtained for the global parameters (Eqs. 6, 9-10) in the infinite nuclear matter (INM) model and Bethe-Weizsacker (BW) mass formula fit using the macroscopic part of nuclear energies (see text). Exact values determined directly using INM and semi-INM calculations are also given. All quantities are in MeV.

Table II Values obtained for the global parameters (Eq. 6,9) in the present study using the experimental data [15] on nuclear masses are given for the various sets of data. All quantities are in MeV.

Table III Values obtained for the parameters (Eq. 10) of BW model using the experimental data [15] with varying number of higher order terms in the model (see text). All quantities are in MeV.

Table IV Same as Table II, but using a Bethe-Weizsacker like mass formula. Values obtained for incompressibility $K_\infty$ using these in Eq.(11) together with the values from Table III for the corresponding set of data are given. All quantities are in MeV.

Table V Values of $K_\infty$ obtained with inclusion/non-inclusion of higher-order effects like curvature $a_{cv}^L$, proton-form-factor (p.f.f.) and exchange Coulomb. INM4 stands for INM model mass fit with 4 parameters, namely $a_v^I$, $a_s^I$, $a_a^I$ and $a_C^I$. And, BW6 stands for BW model fit with 6 parameters, the two additional parameters in this case are the pairing $\Delta$ and surface-symmetry $a_{ss}^I$ terms.
### Table I

| Parameters | Exact Values | INM  | BW  |
|------------|--------------|------|-----|
| $a_v$      | 15.87        | 15.925 | 14.769 |
| $a_C$      | 0.757        | 0.7360 | 0.6945 |
| $a_s$      | 17.3         | 18.10  | 11.15  |
| $a_a$      | 27.0         | 29.80  | 25.41  |
| $a_{ss}$   | -36.0        | -31.37 | -17.77 |
| $a_{cv}$   | 11.1         | 5.06   | 16.43  |

### Table II

| No. of nuclei | $a_v^I$ | $a_C^I$ | $a_s^I$ | $a_a^I$ |
|---------------|---------|---------|---------|---------|
| 1085          | 16.101  | 0.7592  | 19.18   | 24.65   |
| 1191          | 16.115  | 0.7596  | 19.25   | 24.56   |
| 1252          | 16.112  | 0.7589  | 19.23   | 24.66   |
| 1294          | 16.096  | 0.7572  | 19.23   | 24.32   |
| 1371          | 16.108  | 0.7593  | 19.27   | 24.06   |
### Table III

| No. of Para. | $a_v^L$ | $a_s^L$ | $a_a^L$ | $a_C^L$ | $\Delta$ | $a_{ss}^L$ | $a_{cv}^L$ | $a_{gc}^L$ |
|--------------|--------|--------|--------|--------|--------|----------|----------|----------|
| 5            | -15.80 | 18.4   | 23.0   | 0.733  | 11.9   |          |          |          |
| 6            | -15.64 | 18.2   | 26.6   | 0.713  | 11.2   | -22.3    |          |          |
| 7            | -15.48 | 17.2   | 26.3   | 0.707  | 11.2   | -21.9    | 1.5      |          |
| 8            | -12.66 | -10.6  | 21.7   | 0.651  | 10.7   | -7.9     | 102.7    | -127.3   |

### Table IV

| No. of nuclei | $a_v^L$ | $a_C^L$ | $K_\infty$ |
|---------------|--------|--------|------------|
| 1085          | 15.648 | 0.7142 | 291        |
| 1191          | 15.640 | 0.7134 | 290        |
| 1252          | 15.651 | 0.7141 | 297        |
| 1294          | 15.634 | 0.7128 | 303        |
| 1371          | 15.635 | 0.7131 | 288        |
|                     | Model Set                                      | $K_\infty$ |
|---------------------|------------------------------------------------|------------|
| **With Exc. Coul.** | INM4 & BW6                                     | 288        |
|                     | INM4 & BW6+$a_{cv}^L$                           | 302        |
|                     | INM4+p.f.f. & BW6 + p.f.f.                      | 326        |
| **Without Exc. Coul.** | INM4 & BW6                                     | 303        |
|                     | INM4 & BW6+$a_{cv}^L$                           | 309        |
|                     | INM4+p.f.f. & BW6 + p.f.f.                      | 330        |