CompHEP Package with Light Gravitino and $\Sigma$goldstinos

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Abstract

We present a complete description of a new physical model with goldstino supermultiplet for CompHEP software package. The model contains MSSM content supplemented by three new fields: goldstino – two component spinor, scalar $\Sigma$goldstino and pseudoscalar $\Sigma$goldstino. To the leading order in $1/F$, where $F$ is the parameter of supersymmetry breaking, and to zero order in MSSM gauge and Yukawa coupling constants, the interactions between the component fields of goldstino supermultiplet and MSSM fields are derived. They correspond to the most attractive for collider studies processes where only one of these new particles appears in a final state. In this case light gravitino behaves exactly as goldstino. For $\Sigma$goldstinos, as they are R-even, we include in the package only $\Sigma$goldstino couplings to goldstino and $\Sigma$goldstino couplings to SM fields as the most interesting phenomenologically. All new coupling constants between the components of goldstino supermultiplet and MSSM fields are completely determined by the ratios of soft supersymmetry breaking parameters and $F$. Thus, in addition to the MSSM coupling constants there are three new parameters in the package: the scale of supersymmetry breaking, $\sqrt{F}$, and masses of scalar and pseudoscalar $\Sigma$goldstinos. The CompHEP notations for the new fields and parameters are presented. The package was successfully tested and all relevant results known in literature have been reproduced with good accuracy. It may be used in order to calculate any tree-level process with one gravitino or $\Sigma$goldstino on mass shell. Universality of the Lagrangian allows to apply the package in study of phenomenology of any supersymmetric extension of the Standard Model.

1 Introduction

CompHEP package $^{[1]}$ is adjusted to calculate automatically particle decay rates and cross sections at the tree level $^{[2]}$. It is aimed to improve the accuracy, to cut down the efforts and to shorten the time usually required in study of the collision processes at high energy.

First versions of CompHEP contained physical models based only on the set of particles and interactions forming the Standard Model (SM). In CompHEP any user has the opportunity to change vertices and add new particles into the SM content. Thereby some modified models (based on the SM) have been created and used in calculations (see e.g. Refs. $^{[3]}$ for leptoquark models and Refs. $^{[4]}$ for SM with anomalous couplings). The growth of activity in a study of phenomenology of various extensions of the SM inspired the authors of the package to include the Minimal Supersymmetric extension of the Standard Model (MSSM) into CompHEP $^{[5]}$. Below we will refer to this model as $\text{CompHEP/SUSY}$ one.

In the framework of supersymmetry there are several types of models which give significantly different predictions for the values of low-energy MSSM parameters and for low-energy spectrum, that might be

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$^{[3]}$The total number of particles in initial and final states does not exceed 6.
$^{[4]}$The package is freely available at URL: http://theory.sinp.msu.ru/~semenov/mssm.html
tested at accelerators. The incorporation of different classes of models into CompHEP gives an opportunity to study quantitatively high energy phenomenology of these models getting results with high accuracy and with minimum boring routines. The first step to this direction had been done, when the possibility to work with the set of parameters evaluated in the frameworks of mSUGRA and GMSB became available in CompHEP/SUSY. The incorporation of other extensions, in particular, the MSSM with R-parity violation, are in progress.

This paper is devoted to the description of the supersymmetric extensions of the SM with very light gravitino and light sgoldstinos. These new particles as well as their interactions with MSSM fields have been included in the new version of CompHEP/SUSY model. The paper is organized as follows. In section 2 we describe the supersymmetric model with sgoldstino supermultiplet. Namely, section 2.1 contains the description of goldstino supermultiplet. In sections 2.2 and 2.3 we present the effective Lagrangians for very light gravitino and light sgoldstinos, respectively. In section 2.4 we derive the free Lagrangian for light gravitino and sgoldstino fields, as well as the effective interactions responsible for sgoldstino decays into gravitinos. A sketch of phenomenology of the models with very light gravitino and models with light sgoldstinos is presented in Section 3. It is aimed to outline the experimentally allowed region of parameter space of the model. Section 4 explains the CompHEP notations for the new particles and parameters additional to ones used in CompHEP/SUSY model. Section 5 is devoted to concluding remarks, the tests which have been done to check the incorporated models, and possible ways to improve the package.

2 Description of the model

2.1 Goldstino supermultiplet

In any supersymmetric extension of the Standard Model (SM) of particle physics spontaneous supersymmetry breaking occurs due to non-zero vacuum expectation value of an auxiliary component of some chiral or vector superfield. As the most simple case, let us consider a model where the only chiral superfield

\[ S = s + \sqrt{2} \theta \psi + \theta^2 F_s \]  

obtains non-zero vacuum expectation value \( F \) for the auxiliary component,

\[ \langle F_s \rangle \equiv F. \]  

Then \( \psi \) is two-component Goldstone fermion, goldstino, and its superpartners

\[ S \equiv \frac{1}{\sqrt{2}}(s + s^*) , \quad P \equiv \frac{1}{i\sqrt{2}}(s - s^*) , \]  

are scalar and pseudoscalar sgoldstinos, respectively.

In the framework of supergravity \( \partial_\mu \psi \) becomes a longitudinal component of gravitino, that is the super-Higgs effect. As a result, gravitino gains mass \( m_{3/2} \) which is completely determined, in a realistic model with zero cosmological constant, by the parameter of supersymmetry breaking \( F \):

\[ m_{3/2} = \frac{\sqrt{8\pi} F}{\sqrt{3} M_{\text{Planck}}} . \]  

Sgoldstinos remain massless at tree level and become massive due to corrections from high order terms in the Kähler potential. Provided these terms are sufficiently suppressed sgoldstinos are light and may appear in particle collisions at high energy colliders. Such pattern emerges in a number of non-minimal supergravity models and also in gauge mediation models if supersymmetry is broken via non-trivial superpotential (see, Ref. and references therein). Here we consider sgoldstino masses as free parameters.

Gravitino and sgoldstinos interact with MSSM fields and corresponding coupling constants are inversely proportional to the parameter of supersymmetry breaking \( F \). This implies that gravitino has to be very light (see Eq. 3)), otherwise gravitino as well as sgoldstinos is effectively decoupled from the MSSM fields at the energy scale which is expected to be available for study at the nearest future colliders.

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3The package is freely available at URL http://theory.sinp.msu.ru/~semenov/mssm.htm
2.2 Gravitino effective Lagrangian

The effective Lagrangian for the gravitino $\tilde{G}_\mu$ is obtained from the N=1 supergravity theory \[1\], and may be used to calculate scattering processes involving all helicity components of the massive gravitino. Meanwhile the energy scale attainable at the present and the nearest future generation of accelerators favors the longitudinal component of gravitino as the most promising to be searched for in collision experiments. The matter is that while the effective coupling constants for $\pm \frac{1}{2}$ helicity components of gravitino are suppressed by gravity scale, the longitudinal component of gravitino in accordance with the Equivalence Theorem \[7\], effectively behaves as a two-component fermion, goldstino $\psi$,

$$\tilde{G}_\mu \sim i \frac{\sqrt{2}}{\sqrt{3} m_{3/2}} \partial_\mu \psi,$$

so gets enhancement if the energy $E$ of the process is significantly larger than gravitino mass. Therefore, goldstino couplings are suppressed by the ratio $E/F$ (see Eq. (8) for details), which may be significantly lower than gravity scale. The effective low energy Lagrangian is non-renormalizable, and on general grounds the sensitivity of collider experiments to the scale of supersymmetry breaking $\sqrt{F}$ is expected to be not better than about 10 TeV. Thus, in the models under discussion gravitino is very light, $m_{3/2} \lesssim \text{few} \cdot 10^{-2}$ eV (lighter than a few tens millielectronvolts), and only longitudinal component, goldstino, is relevant. Hereby gravitino is considered as massless ($M_{\text{Planck}} \to \infty$, $F = \text{fixed}$) in this approach.

Further simplification is related to the fact that cross sections of the processes with few goldstinos in a final state are suppressed by additional powers of $F$ in comparison with single goldstino production. Consequently, searches for the latter process are expected to be more sensitive to the scale of supersymmetry breaking, so we retain only one-goldstino couplings in the effective Lagrangian. In this case the interaction between goldstino and other fields is plainly given by Goldberger–Treiman formula

$$\mathcal{L}_{GT} = \frac{1}{F} J^\mu_{\text{SUSY}} \partial_\mu \psi,$$

with $J^\mu_{\text{SUSY}}$ being a supercurrent. For MSSM this interaction reads

$$\mathcal{L}_\psi = -\frac{1}{\sqrt{2} F} \sum_{\text{all gauge fields}} \left( \partial_\mu \psi \sigma^\nu \bar{\sigma}^\rho \sigma^\mu \lambda_\alpha^a - \partial_\mu \bar{\psi} \tilde{\sigma}^\rho \sigma^\mu \lambda_\alpha^a \right) F^a_{\nu \rho}$$

$$+ \frac{i}{F} \sum_{\text{all matter and Higgs fields}} \left( \left( D_\nu \tilde{\phi}_k \right)^\dagger \partial_\mu \psi \sigma^\nu \bar{\sigma}^\rho \sigma^\mu \phi_k - \tilde{\phi}_k \bar{\sigma}^\rho \sigma^\mu \partial_\mu \psi D_\nu \phi_k \right),$$

where $a$ and $k$ run over component fields of vector ($\lambda_\alpha^a$, $\psi_\alpha^a \mu$) and chiral ($\tilde{\phi}_k$, $\phi_k$) MSSM supermultiplets, respectively (see Eq. (8) for details); $\sigma^\mu = (1, \bar{\sigma})$, $\bar{\sigma}^\mu = (1, -\bar{\sigma})$. Throughout this paper we adopt two-component fermions.

Whereas we are interested in processes with light gravitino at an external line of Feynman graphs (i.e., tree-level production of gravitino), we simplify Eq. (7) by making use of the goldstino equation of motion (equation of motion of a free massless two-component spinor field). Finally we obtain

$$\mathcal{L}_\psi = \frac{\sqrt{2}}{F} \sum_{\text{all gauge fields}} F^a_{\nu \rho} \left( \partial_\rho \bar{\psi} \sigma^\nu \lambda_\alpha^a - \partial_\rho \psi \sigma^\nu \lambda_\alpha^a \right) + \frac{2}{F} \sum_{\text{all matter and Higgs fields}} \left( \left( D_\mu \tilde{\phi}_k \right)^\dagger \phi_k \partial^\mu \psi - \partial_\mu \tilde{\phi}_k D^\mu \phi_k \right).$$

Such interaction terms exist in any supersymmetric extension of the SM, so this Lagrangian may be regarded as the universal tool to study the sensitivity of various processes to the scale of supersymmetry breaking.

It is worth noting, that as the total amplitudes with helicity $\pm \frac{1}{2}$ gravitino in a final state have to vanish unless supersymmetry is broken, these amplitudes should be proportional to some supersymmetry breaking parameters. Therefore one can use the equations of motion of MSSM fields to rewrite the effective Lagrangian (8) in the form, where all vertices are proportional to soft breaking terms. Although this form is very eligible for simple estimates, since the leading order dependence of the amplitudes on the soft parameters is explicit, CompHEP operates with the Lagrangian (7) because of some technical reasons.
2.3 Sgoldstino effective Lagrangian

We obtain the low-energy effective Lagrangian for sgoldstino by making use of spurion method \cite{8}. It consists in exploiting the fact that, by definition, sgoldstinos are scalar components of the very supermultiplet $S$ whose auxiliary component gains non-zero vacuum expectation value due to supersymmetry breaking. Then one can consider a simple supersymmetric model with non-renormalizable interactions between $S$ and MSSM superfields, which provide MSSM soft terms if non-zero ($F_{\alpha}$) is taken into account. Consequently, the corresponding coupling constants are fixed by the ratios of the soft terms and $F$.

We begin with the Lagrangian of the MSSM, which consists of two parts: the first part is a minimal supersymmetric generalization of the Standard Model (SM), while the second one represents the supersymmetry breaking, \[ \mathcal{L}_{\text{MSSM}} = \mathcal{L}_{\text{SUSY}} + \mathcal{L}_{\text{breaking}}, \]

where \[ \mathcal{L}_{\text{SUSY}} = \mathcal{L}_{\text{gauge}} + \mathcal{L}_{\text{Kähler}} + \mathcal{L}_{\text{superpotential}}, \]

and \[ \mathcal{L}_{\text{gauge}} = \frac{1}{4} \sum_{\text{all gauge fields}} \left( \int d^2\theta \ Tr W^\alpha W_\alpha + h.c. \right), \]

\[ \mathcal{L}_{\text{Kähler}} = \sum_{\text{all matter and Higgs fields}} \int d^2\theta \ d^2\tilde{\theta} \ \Phi_k e^{g_1 V_1 + g_2 V_2 + g_3 V_3} \Phi_k, \]

\[ \mathcal{L}_{\text{superpotential}} = \int d^2\theta \ \epsilon_{ij} \left( \mu H_D H_U^i + y_{ab} L_a F_b + y_{ab}^{ij} Q_a^i D_b^j + y_{ab}^{ij} U_a^j U_b^i \right) + h.c., \]

where $L$, $E$, and $Q$, $D$, $U$ are lepton and quark superfields, respectively, $H_D$ and $H_U$ are electroweak Higgs superfields; $i, j = 1, 2$ are SU(2) and $a, b, c$ are generation indices, color indices are omitted and two-component antisymmetric tensor is defined as $\epsilon_{12} = -1$; $y_{ab}$ are matrices of Yukawa constants and $\mu$ is a parameter of higgsino mixing. Hereafter we assume that R-parity is conserved. The expansion of superfields in series of grassmanian variables $\theta, \tilde{\theta}$ introduces the component fields:

\[ \Phi_k = \tilde{\phi}_k + \sqrt{2} \theta \phi_k + \theta^2 F_k, \]

\[ V = T^a V_a, \quad Tr T^a T^b = \delta^{ab}, \quad V_a = -\theta \sigma^- \bar{\theta} v_{a\mu} + i \theta^2 \bar{\theta} \bar{\lambda}_{\alpha} + \frac{1}{2} \theta^2 \bar{\theta}^2 D_a, \]

\[ W^a = T^a W_\alpha, \quad W^a_\alpha = -i \lambda_{\alpha} + \theta D_{a\alpha} - \frac{i}{2} \sigma^\mu \bar{\sigma}_\nu \theta F_{a\mu\nu} + \theta^2 \sigma^\mu D_{a\mu} \bar{\lambda}_{\alpha}, \]

\[ F_{a\mu\nu} = \delta_{\mu\nu} F_{a\alpha} - \partial_{\nu} v_{a\alpha} + f_{abc} v_{b\mu} v_{c\nu} \quad D_{\mu} \bar{\lambda}_{\alpha} = \partial_{\mu} \bar{\lambda}_{\alpha} + f_{abc} v_{b\mu} \bar{\lambda}_{\alpha}, \]

where for vector superfields we chose the Wess–Zumino gauge.

Generally, supersymmetry breaking part has the form \[ -\mathcal{L}_{\text{breaking}} = \sum_{k} \frac{m_k^2 |\tilde{\phi}_k|^2}{2} + \frac{1}{2} \sum_{\alpha} M_{\alpha} Tr \lambda^\alpha \lambda^\alpha + h.c. \]

\[ -\epsilon_{ij} \left( B_{ij} h_D^i h_U^j + A_{ab}^{L} \bar{\epsilon}_a^{\dot{b}} h_D^i + A_{ab}^{D} \bar{\epsilon}_a^{\dot{b}} \bar{c}_b^i h_D^j + A_{ab}^{U} \bar{\epsilon}_a^{\dot{b}} \bar{c}_b^i h_U^j \right) + h.c., \]

where $k, (\alpha)$ runs over all scalar $\tilde{\phi}_k$ (gaugino $\lambda_{\alpha}$) fields.

Since supersymmetry is broken spontaneously, Eq. \cite{8} implies the following effective interaction between MSSM superfields and sgoldstino supermultiplet $S$:

\[ \mathcal{L}_{S-MSSM} = \mathcal{L}_{S-Kähler} + \mathcal{L}_{S-gauge} + \mathcal{L}_{S-superpotential}, \]

where \[ \mathcal{L}_{S-Kähler} = - \int d^2\theta \ d^2\tilde{\theta} \ S^i S \cdot \sum_{\text{all matter and Higgs fields}} \frac{m_k^2}{F^2} \Phi_k e^{g_1 V_1 + g_2 V_2 + g_3 V_3} \Phi_k, \]
\[
\mathcal{L}_{\text{S–gauge}} = \frac{1}{2} \int d^2 \theta \left( \sum_{\text{all gauge fields}} \frac{M_\alpha}{F} Tr W^\alpha W^\alpha + \text{h.c.} \right),
\]

\[
\mathcal{L}_{\text{S–superpotential}} = \int d^2 \theta \left( \sum_{\text{all gauge fields}} \epsilon_{ij} \left( \frac{B}{F} H_D^i H_D^j + \frac{A^L_{ab}}{F} \bar{L}_a^i F_b^j + \frac{A^D_{ab}}{F} Q_a^i D_b^j + \frac{A^U_{ab}}{F} U_a^i U_b^j \right) \right) + \text{h.c.}
\]

These terms emerge if the fields from the hidden sector, where supersymmetry breaking occurs, are integrated out. The only remnant of the hidden sector is goldstino supermultiplet, which is supposed to be light.

Integrating over \( \theta, \bar{\theta} \) and taking into account Eqs. (9), (10) we obtain soft supersymmetry breaking terms (11) as well as interaction between the components of goldstino supermultiplet and components of MSSM superfields:

\[
\mathcal{L}_{\text{S–Kähler}} = - \sum_{\text{all matter and Higgs fields}} \left( \frac{m_i^2}{F} s \cdot \tilde{D}_k^i F_k + \text{h.c.} \right),
\]

\[
\mathcal{L}_{\text{S–gauge}} = \sum_{\text{all gauge fields}} \left( -i \frac{M_\alpha}{2F} s \cdot \chi^i_\alpha \sigma^\mu D^\mu \chi^i_\alpha + \frac{M_\alpha}{2F} s \cdot D^\alpha D^\alpha - \frac{M_\alpha}{4F} s \cdot F^\alpha_{\mu \nu} F^\alpha_{\mu \nu} - i \frac{M_\alpha}{8F} s \cdot \epsilon_{\mu \nu \lambda \rho} F^\alpha_{\lambda \rho} + \text{h.c.} \right),
\]

\[
\mathcal{L}_{\text{S–superpotential}} = - \epsilon_{ij} \left( \frac{B}{F} \chi^i_\alpha \chi^j_\alpha + \frac{B}{F} s \cdot \left( \bar{h}^i_D F_{\bar{h}^i_U} + F^i_{\bar{h}^i_D} h^i_U \right) \right) + \frac{A^L_{ab}}{F} \left( \bar{l}_a^i \epsilon_a^i \cdot s h_D^i + \bar{l}_a^i \epsilon_a^i \cdot s \bar{\epsilon}_a^i \epsilon_b^i \cdot \bar{\epsilon}_b^i \epsilon_a^i \cdot \bar{\epsilon}_a^i \epsilon_b^i \cdot \bar{\epsilon}_b^i \right) - \frac{A^D_{ab}}{F} \left( \bar{q}_a^i \epsilon_a^i \cdot s h_D^i + \bar{q}_a^i \epsilon_a^i \cdot s \bar{\epsilon}_a^i \epsilon_b^i \cdot \bar{\epsilon}_b^i \epsilon_a^i \cdot \bar{\epsilon}_a^i \epsilon_b^i \cdot \bar{\epsilon}_b^i \right) + \frac{A^U_{ab}}{F} \left( \bar{u}_a^i \epsilon_a^i \cdot s h_D^i + \bar{u}_a^i \epsilon_a^i \cdot s \bar{\epsilon}_a^i \epsilon_b^i \cdot \bar{\epsilon}_b^i \epsilon_a^i \cdot \bar{\epsilon}_a^i \epsilon_b^i \cdot \bar{\epsilon}_b^i \right) + \text{h.c.}
\]

Here we presented only leading order in \( \frac{1}{F} \) terms; the convention for the Levi–Civita tensor is \( \epsilon^{0123} = -1 \).

Eliminating auxiliary fields, one gains the low-energy effective Lagrangian for the interaction between the components of goldstino supermultiplet and MSSM fields. A few remarks are in order. First, being aimed at getting the most universal Lagrangian applicable to any supersymmetric extension of the SM, we have to consider only interaction terms obtained to the leading order in \( 1/F \) and to zero order in MSSM coupling constants. Second, from the phenomenological point of view the most promising interaction terms are ones which include the minimum new (additional to SM) fields. In model with R-parity goldstino is R-odd, whereas goldstinos are R-even. Consequently, the simplest vertex with goldstino has to contain at least one new field – superpartner, while among goldstino vertices there is a set containing only SM fields, and only these vertices are of interest. Likewise we reject the interactions with \( F \) and \( D \) terms, since, being resolved, these auxiliary fields are proportional to SM gauge or Yukawa couplings. Thus, only minimal goldstino-partner-partner vertices have been included in the presented model:

\[
\mathcal{L}_S = \left( \sum_{\text{all gauge fields}} \frac{M_\alpha}{4 \sqrt{2} F} S \cdot \epsilon_{\mu \nu \lambda} F^\alpha_{\mu \nu} F^\alpha_{\lambda \rho} \right)
\]

\[
\mathcal{L}_P = \left( \sum_{\text{all gauge fields}} \frac{M_\alpha}{8 \sqrt{2} F} P \cdot \epsilon_{\mu \nu \lambda} F^\alpha_{\mu \nu} F^\alpha_{\lambda \rho} \right)
\]
with these rules reads corresponding couplings presented in Ref. [9] except of the terms proportional to the parameter the corresponding Yukawa couplings, \( A_{ab} = A_{ab \mu \nu} \). The sgoldstino Lagrangian transformed in accordance with these rules reads

\[
L_S = - \sum_{all \ gauge \ fields} \frac{M_{S}}{2\sqrt{2}F} S \cdot F_{a \mu \nu} F_{a}^{\mu \nu} - \frac{A_{ab}^L}{\sqrt{2}F} g_{ab}^D \cdot S (\epsilon_{ij} l_{i}^a \epsilon_{j}^b h_D^i + h.c.) - \frac{A_{ab}^D}{\sqrt{2}F} g_{ab}^D \cdot S (\epsilon_{ij} q_{i}^a \epsilon_{j}^b h_D^i + h.c.) ,
\]

\[
L_P = \sum_{all \ gauge \ fields} \frac{M_{P}}{2\sqrt{2}F} P \cdot F_{a \mu \nu} \epsilon_{\mu \nu \lambda \rho} F_{a \lambda \rho} - i \frac{A_{ab}^L}{\sqrt{2}F} g_{ab}^L \cdot P (\epsilon_{ij} l_{i}^a \epsilon_{j}^b h_D^i - h.c.) - i \frac{A_{ab}^D}{\sqrt{2}F} g_{ab}^L \cdot P (\epsilon_{ij} q_{i}^a \epsilon_{j}^b h_D^i - h.c.) .
\]

One can check that resulting interaction terms between sgoldstinos and SM gauge bosons coincide with corresponding couplings presented in Ref. [10] except of the terms proportional to the parameter \( \mu_a \) introduced there. The resulting interaction terms between sgoldstinos and SM fermions coincide with corresponding couplings presented in Ref. [11] up to overall sign.

### 2.4 Gravitino–sgoldstino effective Lagrangian

To obtain the effective Lagrangian for gravitino-sgoldstino sector we again apply the spurion method [3]. It is suitable for our purposes, since the only process we want to describe is the decay of sgoldstinos into gravitinos, i.e., the process with all involved particles being on mass shell. These decays are kinematically allowed in the region of parameter space viable for study at colliders. The relevant terms follow from the effective Lagrangian

\[
L_S = \int d\theta^2 d\bar{\theta}^2 \left(S^\dagger S - \frac{m_S^2}{2} + \frac{m_P^2}{2F^2} (S^\dagger S)^2 \right) \left( - \frac{m_S^2}{12F} \right) \left( \int d\theta^2 S^3 + h.c. \right) ,
\]

where we introduced sgoldstino masses. Recalling the definitions (1), (2), (3) we obtain the Lagrangian with kinetic and mass terms for sgoldstino (longitudinal component of gravitino) and sgoldstinos and the interactions respectable for sgoldstino decays into gravitinos

\[
L_{\psi, S, P} = i \partial_{\mu} \bar{\psi} \partial^{\mu} \psi + \frac{1}{2} \partial_{\mu} S \partial^{\mu} S - \frac{1}{2} m_S^2 S^2 + \frac{1}{2} \partial_{\mu} P \partial^{\mu} P - \frac{1}{2} m_P^2 P^2 + \frac{m_S^2}{2\sqrt{2}F} S (\psi \psi + \bar{\psi} \bar{\psi}) - i \frac{m_P^2}{2\sqrt{2}F} P (\psi \psi - \bar{\psi} \bar{\psi}) .
\]

The results for sgoldstino-gravitino coupling constants coincide with corresponding terms presented in Ref. [3] up to overall sign.

### 3 Phenomenology of the model

Till now, there are no experimental evidence for gravitino or sgoldstinos. The study of their phenomenology results in acquisition of bounds on their coupling constants.

Let us briefly review the current status of experimental limits on \( F \).
If we ignore sgoldstino (or, equivalently, set \( m_{S,P} \to \infty \)), then in models with light gravitino the current direct bound on the supersymmetry breaking scale is obtained from Tevatron, \( \sqrt{F} \geq 217 \text{ GeV} \). The upgraded Tevatron may be able to cover the range of \( \sqrt{F} \) up to \( \sqrt{F} \approx 290 \text{ GeV} \), while LHC will be capable to reach \( \sqrt{F} \approx 1.6 \text{ TeV} \). As we will see, in models with light sgoldstinos, collider experiments become more sensitive to the scale of supersymmetry breaking.

Then, in the model under discussion gravitino is extremely light: \( m_{3/2} \lesssim 10^{-2} \text{ eV} \) and may affect standard Big Bang Nucleosynthesis as a neutrino-like specie. To get a sufficient dilution the gravitino decoupling temperature must exceed \( T \sim 200 \text{ MeV} \), that yields the limit on the scale of supersymmetry breaking, \( \sqrt{F} \gtrsim 600 \text{ GeV} \) at \( m_{\Sigma} \approx 100 \text{ GeV} \). Similar bounds follow from the cooling rate of white dwarfs and \( Z \)-boson width in the models where one neutralino is lighter than \( Z \)-boson. It is worth noting that all these bounds are ensued from the study of the processes with two gravitinos, so they are model dependent and may be invalid in other supersymmetric models.

Let us turn to the models with light sgoldstinos. Note, that all sgoldstino coupling constants introduced in previous sections are completely determined by MSSM soft terms and the parameter of supersymmetry breaking \( F \), but sgoldstino masses \( (m_S,m_P) \) remain free. Depending on the values of these unknown parameters, sgoldstinos may show up in different experiments. Phenomenologically interesting models form four classes:

1) Sgoldstino masses are of order of electroweak scale, \( \sqrt{F} \sim 1 \text{ TeV} \) — sgoldstino may be produced in collisions of high energy particles at colliders \( \text{ [1] [2] [3]} \).

2) Sgoldstino masses \( m_{S,P} \sim 1 \text{ MeV} \), \( \sqrt{F} \sim 1 \text{ TeV} \) — sgoldstinos may emerge in products of rare decays of mesons \( \text{ [18] [11]} \), such as \( \Upsilon \to S(P)\gamma, J/\psi \to S(P)\gamma \).

3) Sgoldstinos in models with flavor violation in the sector of soft trilinear couplings, \( A_{ab} \neq A_{\delta ab} \) — sgoldstino interactions lead to flavor violating processes. Namely, sgoldstinos may contribute to FCNC (mass difference and/or CP-violation in the systems of neutral mesons \( \text{ [14] [20]} \), then, if kinematically allowed, they appear in the product of rare decays, such as \( t \to cS(P) \) \( \text{ [12]} \), \( \mu \to eS(P) \), \( K \to \pi S \) \( \text{ [10]} \), etc.

4) Sgoldstinos are lighter than 1 MeV — these models may be tested in low energy experiments \( \text{ [10]} \), such as reactor experiments, conversion in magnetic field, etc. Also sgoldstinos may play very important role in astrophysics and cosmology \( \text{ [22] [23] [24] [10]} \), they may change the predictions of Big Bang Nucleosynthesis, distort CMB spectrum, affect SN explosions and cooling rate of stars, etc.

There are flavor-conserving and flavor-violating interactions of sgoldstino fields.

As concerns flavor-conserving interactions, the strongest bounds arise from astrophysics and cosmology, that is \( \sqrt{F} \gtrsim 10^6 \text{ GeV} \) \( \text{ [22] [10]} \), or \( m_{3/2} \gtrsim 600 \text{ eV} \), for models with \( m_{S(P)} < 10 \text{ keV} \) and MSSM soft flavor-conserving terms being of the order of electroweak scale. For the intermediate sgoldstino masses (up to a few MeV) constraints from the study of SN explosions and reactor experiments lead to \( \sqrt{F} \gtrsim 300 \text{ TeV} \) \( \text{ [10]} \). For heavier sgoldstinos, low energy processes (such as rare decays of mesons) provide limits at the level of \( \sqrt{F} \gtrsim 500 \text{ GeV} \) \( \text{ [10]} \).

The collider experiments exhibit the same level of sensitivity to light sgoldstinos as rare meson decays. Indeed the studies \( \text{ [26] [27]} \) of the light sgoldstino \( (m_{S,P} \lesssim a few \text{ MeV}) \) phenomenology based on the effective low-energy Lagrangian derived from N=1 linear supergravity yield the bounds: \( \sqrt{F} \gtrsim 500 \text{ GeV} \) (combined bound on \( Z \to S \bar{f}f, P \bar{f}f \) \( \text{ [20]} \), combined bound on \( e^+e^- \to \gamma S, \gamma P \) \( \text{ [13]} \) at \( M_{soft} \sim 100 \text{ GeV} \), \( \sqrt{F} \gtrsim 1 \text{ TeV} \) \( \text{ [27]} \) (combined bound on \( pp \to gS,gP \) at \( M_3 \approx 500 \text{ GeV} \). Searches for heavier sgoldstino at colliders, though exploiting another technique, results in similar bounds on the scale of supersymmetry breaking. Most powerful among the operating machines, LEP and Tevatron, give a constraint of the order of 1 TeV on supersymmetry breaking scale in models with light sgoldstinos. Indeed, the analysis carried out by DELPHI Collaboration \( \text{ [28]} \) yields the limit \( \sqrt{F} > 500 \div 200 \text{ GeV} \) at sgoldstino masses \( m_{S,P} = 10 \div 150 \text{ GeV} \) and \( M_{soft} \sim 100 \text{ GeV} \). The constraint depends on the MSSM soft breaking parameters. In particular, it is stronger by about a hundred GeV in the model with degenerate gauginos. At Tevatron, a few events in \( pp \to S \gamma(Z) \) channel, and about \( 10^4 \) events in \( pp \to S \) channel would be produced at \( \sqrt{F} = 1 \text{ TeV} \) and \( M_{soft} \sim 100 \text{ GeV} \) for integrated luminosity \( L = 100 \text{ pb}^{-1} \) and sgoldstino mass of order of 100 GeV \( \text{ [17]} \). This gives rise to a possibility to detect sgoldstino, if it decays inside the detector into photons and \( \sqrt{F} \) is not larger than \( 1.5 \div 2 \text{ TeV} \).

\(^4\)All these bounds were obtained in the framework of linear N=1 supergravity.
The flavor-asymmetric processes are generally very sensitive to light sgoldstino. Namely, with flavor-changing off-diagonal entries in squark (slepton) squared mass matrix close to the current bounds, direct measurements of decays of mesons (leptons) provide very strong bounds, up to \( \sqrt{F} \gtrsim 900(15000) \text{ TeV} \) (valid at \( m_S \lesssim 5(0.34) \text{ GeV} \)), which is much higher than bounds expected from future colliders. Sgoldstinos may appear in top-quark decay, that could help to probe the scale of supersymmetry breaking at LHC up to about 10 TeV \(^{[22]}\). The possible contributions of sgoldstinos with \( m_{S,P} = 100 \text{ GeV} \) to mass differences in the systems of neutral mesons yield the bounds \( \sqrt{F} \gtrsim 10 \text{ TeV} \) \(^{[20]}\). If off-diagonal entries are small, the limits on \( \sqrt{F} \) become weaker: they scale as square root of the corresponding off-diagonal elements.

Among the laboratory processes, the most sensitive to very light sgoldstinos are propagation of laser beam in magnetic field and reactor experiments. For heavier sgoldstinos the most sensitive processes to flavor-conserving sgoldstino couplings are \( \Upsilon \) decays and associated sgoldstino production in collisions of high energy particles; charged kaon decays are most sensitive to flavor-violating sgoldstino couplings. For sgoldstino masses of the order of electroweak scale the most sensitive process to flavor-conserving sgoldstino couplings is sgoldstino resonant production at hadron colliders, while the most sensitive to flavor-violating couplings is measurement of mass differences in the systems of neutral mesons.

### 4 CompHEP notations

In the CompHEP package gravitino and sgoldstinos (\( \psi, S, P \)) are denoted as \( \tilde{g}, \tilde{sS} \) and \( \tilde{sP} \), respectively. Sgoldstino masses \( m_S \) and \( m_P \) are referred as \( m_S \) and \( m_P \); gravitino is massless. The scale of supersymmetry breaking in the entire theory, \( \sqrt{F} \), corresponds to \( \tilde{g} \). The Feynman rules for the Lagrangians \(^{(7)}, (15), (16), (18)\) have been obtained by making use of LanHEP package \(^{[29]}\). They have been incorporated into CompHEP, as well as aforementioned new parameters with default values \( m_{S,P} = 200 \text{ GeV}, \tilde{g} = \sqrt{3} \text{ TeV} \). Subsequent checks gave results consistent with ones known in literature.

### 5 Concluding remarks

We described the model with very light gravitino and light sgoldstinos, which has been included into CompHEP/SUSY package. This package may be used in order to calculate a rate of any tree-level process with one gravitino or sgoldstino being on mass shell. This includes:

- decay of a superpartner into partner and gravitino
- sgoldstino decays
- gravitino inelastic scattering off MSSM particles into MSSM particles
- sgoldstino inelastic scattering off SM particles into MSSM particles
- gravitino (sgoldstino) production in collisions of MSSM particles.

The effective Lagrangian incorporated into CompHEP/SUSY is universal, so the package may be used to study the phenomenology in any supersymmetric extension of the Standard Model. The only interaction terms to the leading order in \( 1/F \) and to zero order in MSSM gauge and Yukawa coupling constants, relevant for gravitino and sgoldstino phenomenology, are very terms included into this package.

Note, that direct independent measurement of MSSM soft supersymmetry breaking terms and gravitino or/and sgoldstino couplings gives the unique possibility to estimate the scale of supersymmetry breaking \( \sqrt{F} \).

As a final check for gravitino couplings we compared the analytical results (from CompHEP output) for all decay rates of the kind \( \text{superpartner} \to \text{partner} + \text{gravitino} \) with corresponding formulae presented in Ref. \(^{[30]}\) and found no difference. To make sure that sgoldstino vertices are correct we compared the analytical results (from CompHEP output) for all sgoldstino decay rates with corresponding formulae presented in Refs. \(^{[3], [10]}\). The results coincide up to terms proportional to additional parameter \( \mu_a \) introduced in Ref. \(^{[3]}\). To check

\(^{5}\) These values may be changed by user that is a standard CompHEP menu option.
the numerical calculations carrying out by the package, we have reproduced with good accuracy all results for sgoldstino production in various channels presented in Refs. [9, 17].

There are several remarks which may be helpful for the user who are going to work with this package. It is worth recalling that we discussed the model with R-parity being conserved. Gravitino is R-odd, consequently, it might appear in collision of SM particles only accompanied by another R-odd particle. Moreover, gravitino is stable the lightest supersymmetric particle — LSP. Sgoldstinos are R-even, therefore, their production may be not accompanied by production of superpartners. Sgoldstinos may decay into SM particles, and since two-photon channel is always open, they are unstable.

The second remark concerns the treatment of sgoldstino coupling constants to be proportional to the MSSM soft breaking terms. In fact, the Lagrangians (13), (16) describe sgoldstino interactions at the superpartner scale. Sgoldstino coupling constants at other energies may be obtained by making use of renormalization group evolution. As an example, for the gluonic operator one has

$$C_{\mu\nu}^2(M_3) = C_{\mu\nu}^2(\mu) \frac{\beta(\alpha_s(\mu))}{\alpha_s(\mu)} \frac{\alpha_s(M_3)}{\beta(\alpha_s(M_3))},$$

(19)

where $\beta(\alpha_s)$ is QCD $\beta$-function. These corrections (renormalization group enhancement) are neglected in the package, but may be taken into account plainly by making use of appropriate rescaling.

Then, in CompHEP all Yukawa coupling constants are determined as the ratios of fermion pole masses to the corresponding Higgs vacuum expectation value. In order to estimate sgoldstino branching ratios into fermions one should take into account the difference between running Yukawa constants and Yukawas used in CompHEP.

The forth remark is that sgoldstino couplings to superpartners are not included into this package. If open (allowed kinematically), new sgoldstino decay channels distort the pattern of sgoldstino branching ratios calculated in CompHEP. In a given model the level of distortion depends on the set of soft supersymmetry breaking masses and sgoldstino masses and effective couplings.

Finally, we discuss the possible directions to improve this package. The first interesting possibility is to take into account CP- and flavor-violating sgoldstino couplings. This option will be allowed, if complex coupling constants and flavor violating couplings become available in CompHEP. Then, sgoldstino couplings to superpartners may be also included. In the spurion approach corresponding coupling constants are not completely determined by MSSM soft breaking terms, but may depend on new parameters originated from the hidden sector. Similar situation takes place with sgoldstino interaction terms proportional to MSSM gauge or Yukawa couplings. For instance, the account of $F$ and $D$ auxiliary fields (see Eqs. (11), (12)) results, in particular, in sgoldstino-higgs mixing with coefficients depending also on new parameters from hidden sector. The model independence of these coefficients will be clarified elsewhere.

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