Oblique parameters in gauged baryon and lepton numbers with a 125 GeV Higgs

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Abstract: In an extension of the standard model, where baryon number and lepton number are local gauge symmetries, we analyze the effect of corrections from exotic fermions and scalars on the oblique parameters \( S, T, U \). Because a light neutral Higgs \( h_0 \) with mass around 124–126 GeV strongly constrains the corresponding parameter space of this model, we also investigate the gluon fusion process \( gg \rightarrow h_0 \) and two photon decay of the lightest neutral Higgs \( h_0 \rightarrow \gamma \gamma \) at the Large Hadron Collider.

Key words: local gauge symmetry, baryon and lepton numbers, Higgs

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1 Introduction

The main physics goals of the Large Hadron Collider (LHC) are to understand the origin of electroweak symmetry breaking, and to search for the neutral Higgs boson predicted by the standard model (SM) and its various extensions. Recently, ATLAS and CMS have reported significant excess events that are interpreted to be most probably related to a neutral Higgs with mass \( m_{h_0} \sim 124–126 \) GeV. This implies that the Higgs mechanism of breaking electroweak symmetry possibly has a solid experimental cornerstone.

The oblique parameters \( S, T, U \) [1] are extracted from electroweak precision data (EWPD) observations that probe the radiative corrections with sufficient accuracy. A light neutral Higgs with mass 124–126 GeV also affects the theoretical evaluations of the oblique parameters \( S, T, U \) through loop corrections to the gauge boson propagators, which contain the neutral Higgs as a virtual field. In extensions of the SM, the corrections from the exotic fields to the gauge boson propagators can be expressed in terms of shifts of the parameters \( S, T, U \) [2].

A broken baryon number \((B)\) conservation can explain the origin of the matter-antimatter asymmetry in the Universe in a natural way. The heavy majorana neutrinos contained in the seesaw mechanism can induce the tiny observed neutrino masses [3] to explain the results of neutrino oscillation experiments. Hence, the lepton number \((L)\) is also expected to be broken. In Ref. [4], two extensions to the SM are examined, where \( B \) and \( L \) are spontaneously broken gauge symmetries around the TeV scale, while Ref. [5] also investigates the predictions for the Higgs mass and the Higgs decays in a supersymmetric model named BLMSSM, which is a minimal supersymmetric extension of the SM (MSSM) with local gauged \( B \) and \( L \). Within the framework of the first extension of the SM with spontaneously broken \( B \) and \( L \) [4], we analyze the gluon fusion production and then decay into two photons of the Higgs with mass \( m_{h_0} \sim 124–126 \) GeV. Additionally, we also investigate the corrections from exotic fields of the oblique parameters \( S, T, U \).

This paper is organized as follows. In Section 2, we briefly summarize the main ingredients of an extension of the SM where the baryon and lepton numbers are local symmetries, we then present the mass squared matrices for the neutral Higgs sector. Inspired by the new results from the ATLAS and CMS collaborations, in Section 3 we study in great detail the Higgs production through gluon fusion, followed by the decay of the Higgs boson into two photons. We discuss the constraints on the parameter space from the oblique parameters \( S, T, U \) in Section 4. Our conclusions are given in Section 5.
2 An extension of the SM where baryon and lepton numbers are local gauge symmetries

When baryon and lepton numbers are local gauge symmetries, one can write the gauge group as \( SU(3)_C \otimes SU(2)_L \otimes U(1)_Y \otimes U(1)_B \otimes U(1)_L \). In the first extension of the SM proposed in Ref. [4], the exotic particles include new quarks \( Q_L, u_R, d_R \), new leptons \( l_R, \nu_R, e_R \) and three scalar singlets \( S_B, S_L, S \) along with a scalar doublet \( \phi \). The Yukawa couplings are written as

\[
-L_Y = \sum_{i,j=1}^{3} \left[ (Y_{ij})_1 \bar{Q}_L^i \tilde{H} u_R^j + (Y_{ij})_1 \bar{Q}_L^i H d_R^j \right] + \sum_{i=1}^{3} \left[ (Y_{ij})_1 \tilde{Q}_L^i \tilde{H} u_R^j + (Y_{ij})_1 \tilde{Q}_L^i H d_R^j \right] + \text{h.c.} + \sum_{i=1}^{3} \left[ (Y_{ij})_1 \tilde{Q}_L^i \tilde{H} \nu_R^j + (Y_{ij})_1 \tilde{Q}_L^i H \epsilon_R^j \right] + \sum_{i=1}^{3} \left[ (Y_{ij})_1 \tilde{Q}_L^i \tilde{H} \nu_R^j + (Y_{ij})_1 \tilde{Q}_L^i H \epsilon_R^j \right] + \text{h.c.} \right],
\]

The scalar potential is generally given as follows:

\[
-L_S = m_\phi^2 H^\dagger H + m_{S_{B}}^2 S_{B}^* S_{B} + m_{S_{L}}^2 S_{L}^* S_{L} + m_{\phi}^2 \phi^\dagger \phi + \lambda_{HH} (H^\dagger H)(H^\dagger H) + \lambda_{\phi \phi} (\phi^\dagger \phi)(\phi^\dagger \phi) + \lambda_{BB} (S_{B}^* S_{B})(S_{B}^* S_{B}) + \lambda_{LL} (S_{L}^* S_{L})(S_{L}^* S_{L}) + \lambda_{SS} (S^* S)(S^* S) + \lambda_{HH} (H^\dagger H)(H^\dagger H) + \lambda_{\phi \phi} (\phi^\dagger \phi)(\phi^\dagger \phi) + \lambda_{HH} (H^\dagger H)(S_{B}^* S_{B}) + \lambda_{HL} (H L^\dagger H)(S_{L}^* S_{L}) + \lambda_{HS} (H^\dagger H)(S^* S) + \lambda_{\phi \phi} (\phi^\dagger \phi)(S_{B}^* S_{B}) + \lambda_{\phi \phi} (\phi^\dagger \phi)(S_{L}^* S_{L}) + \lambda_{BL} (S_{B}^* S_{B})(S_{L}^* S_{L}) + \lambda_{BS} (S_{B}^* S_{B})(S^* S) + \lambda_{LS} (S_{L}^* S_{L})(S^* S) + \lambda_{HH} (H^\dagger H)(\phi^\dagger \phi) + \mu_1 (H^\dagger \phi)(\phi^\dagger H) + \mu_2 S_{B}^2 S_{B}^2 + \text{h.c.} \right].
\]

When the \( SU(2)_L \) doublet \( H \) and \( SU(2)_L \) singlets \( S_B, S_L \) acquire the nonzero vacuum expectation values (VEVs) \( v, \bar{v}_{B,L} \),

\[
H = \left( \frac{1}{\sqrt{2}} (v + H_0 + iG_0^0) \right),
\]

\[
S_B = \frac{1}{\sqrt{2}} (v_B + S_B^0 + iG_B^0),
\]

\[
S_L = \frac{1}{\sqrt{2}} (v_L + S_L^0 + iG_L^0),
\]

the local gauge symmetry \( SU(2)_L \otimes U(1)_Y \otimes U(1)_B \otimes U(1)_L \) is broken down to the electromagnetic symmetry \( U(1)_e \), where \( G^+, G^0, G_B^0 \) and \( G_L^0 \) denote massless Goldstone bosons. Correspondingly, the mass terms for the neutral Higgs are formulated as

\[
-L^H_{\text{mass}} = \frac{1}{2} \left( H_0, S_B^0, S_L^0 \right) m_{\text{CPM}}^2 \left( \begin{array}{c} H_0 \\ S_B^0 \\ S_L^0 \end{array} \right),
\]

where the symmetric 3x3 mass squared matrix \( m_{\text{CPM}}^2 \) is

\[
m_{\text{CPM}}^2 = \begin{pmatrix} 2\lambda_{HH} v^2 & \lambda_{HB} & \lambda_{HL} v \nu_L \\ \lambda_{HB}^* & 2\lambda_{BB} v^2 & \lambda_{BL} v \nu_L \\ \lambda_{HL} v \nu_L & \lambda_{BL} v \nu_L & 2\lambda_{LL} v^2 \end{pmatrix}.
\]

Through the orthogonal 3x3 transformation matrix \( Z_{\text{CPM}} \), the mass squared matrix \( m_{\text{CPM}}^2 \) can be diagonalized as

\[
Z_{\text{CPM}}^T m_{\text{CPM}}^2 Z_{\text{CPM}} = \begin{pmatrix} m_{H_1}^2 & m_{H_2}^2 & m_{H_3}^2 \end{pmatrix},
\]

where \( m_{H_i} \approx m_{h_0} \approx 125 \text{ GeV} \).

In a similar way, we can write the \( SU(2)_L \) doublet \( \phi \) and the \( SU(2)_L \) singlet \( S \) as

\[
\phi = \begin{pmatrix} \phi^+ \\ \phi^0 + i\phi^0 \end{pmatrix},
\]

\[
S = \frac{1}{\sqrt{2}} \begin{pmatrix} S_R^0 + iS_I^0 \end{pmatrix}.
\]

Since the local gauge symmetry \( SU(2)_L \otimes U(1)_Y \otimes U(1)_B \otimes U(1)_L \) is broken down to the electromagnetic symmetry \( U(1)_e \), the terms in square brackets in Eq. (2) induce mixing among the neutral scalar particles \( \phi_R^0, \phi_I^0, S_R^0, S_I^0 \), and the mass terms are written as

\[
-L^\phi_{\text{mass}} = \frac{1}{2} \left( \phi_R^0, S_R^0, \phi_I^0, S_I^0 \right) m_{\text{CPM}}^2 \left( \begin{array}{c} \phi_R^0 \\ S_R^0 \\ \phi_I^0 \\ S_I^0 \end{array} \right),
\]

with the symmetric 4x4 mass squared matrix \( m_{\text{CPM}}^2 \) being
Meanwhile, the Majorana mass matrix \( m^2_{CPM} \) through the \( 4 \times 4 \) orthogonal rotation \( Z_{CPM} \):

\[
Z^T_{CPM} m^2_{CPM} Z_{CPM} = \begin{pmatrix} m^2_\phi & \sqrt{2} v R(\mu_1) & 0 & -\sqrt{2} v \Im(\mu_2) \\
\sqrt{2} v R(\mu_1) & m^2_3 + 2 \sqrt{2} v b R(\mu_2) & 0 & -2 \sqrt{2} v b \Im(\mu_2) \\
0 & 0 & m^2_\phi & -\sqrt{2} v b R(\mu_1) \\
-\sqrt{2} v \Im(\mu_1) & -2 \sqrt{2} v b \Im(\mu_2) & -\sqrt{2} v R(\mu_1) & m^2_3 + 2 \sqrt{2} v b R(\mu_2) \end{pmatrix}. \tag{9}
\]

We also diagonalize the mass squared matrix \( m^2_{CPM} \) by a 5 \times 5 diagonal matrix \( m^2_{\phi \pm} \) is expressed by

\[
m^2_{\phi \pm} = \frac{1}{2} m^2_\phi - \frac{1}{2} \eta_{\phi \mp} v^2. \tag{11}
\]

Similarly, the mass for the charged scalar \( \phi^\pm \) is expressed by

\[
m^2_{\phi \pm} = \frac{1}{2} m^2_\phi \mp \frac{1}{2} \eta_{\phi \mp} v^2. \tag{11}
\]

Since the field \( \phi \) does not get a nonzero VEV after the electroweak symmetry is broken down, there is no mass mixing between the exotic quarks and the SM quarks.

In the left-handed basis \( (\nu^c_1, \nu^c_2, \nu^c_3, \nu^c_4) \), \( (I=1, 2, 3) \), the mass matrix for neutrinos is given by the \( 8 \times 8 \) matrix

\[
\mathcal{M}_n = \begin{pmatrix}
0_{3 \times 3} & (M_D)_{3 \times 5} \\
(M_D^T)_{5 \times 3} & (M_N)_{5 \times 5}
\end{pmatrix}. \tag{12}
\]

Here, the \( 3 \times 5 \) matrix \( M_D \) is written as

\[
M_D = \begin{pmatrix}
0 & v \\
\frac{v}{\sqrt{2}} Y_{\nu}^{*} & 0_{1 \times 3}
\end{pmatrix}, \tag{13}
\]

and the \( 5 \times 5 \) matrix \( M_N \) is

\[
M_N = \begin{pmatrix}
0 & v \\
\frac{v}{\sqrt{2}} Y_{\nu}^{*} & 0_{1 \times 3} \\
0_{3 \times 1} & \frac{v}{\sqrt{2}} (\lambda_{b})_{3 \times 1} \\
0_{1 \times 3} & \frac{v}{\sqrt{2}} (\lambda_{b})_{3 \times 3}
\end{pmatrix}. \tag{14}
\]

By integrating the heavy freedoms out, we get the following mass matrix for three light neutrinos:

\[
\mathcal{M}_\nu = -M_D M_N^{-1} M_D^T, \tag{15}
\]

which is diagonalized by the Pontecorvo-Maki-Nakagawa-Sakata matrix \( U_{PMNS} \)

\[
U_{PMNS}^T \mathcal{M}_\nu U_{PMNS} = \text{diag}(m_{\nu_1}, m_{\nu_2}, m_{\nu_3}). \tag{16}
\]

Meanwhile, the Majorana mass matrix \( M_N \) is similarly diagonalized by a \( 5 \times 5 \) matrix \( Z_N \)

\[
Z_N^T M_N Z_N = \text{diag}(m_{N_1}, m_{N_2}, m_{N_3}, m_{N_4}, m_{N_5}). \tag{17}
\]

### 3 The \( gg \to h_0 \to \gamma \gamma \) process in gauged baryon and lepton numbers

At the LHC, the Higgs is produced chiefly through gluon fusion. In the SM, the leading order (LO) contributions originate from the one-loop diagram, which involves virtual top quarks. The cross section for this process is known as the next-to-leading order (NNLO) \([6]\), which can enhance the LO result by 80%–100%. Furthermore, any new particle that couples strongly with the Higgs can significantly modify this cross section. In the extension of the SM considered here, the LO decay width for the process \( h_0 \to gg \) is given by (see Ref. \([7]\) and the references therein)

\[
\Gamma_{NP}(h_0 \to gg) = \frac{G_F^2 m^3_{h_0} |(Z_{CPE})_{11}|^2}{64 \sqrt{2} \pi^3} |A_{1/2}(x_t)|^2 + A_{1/2}(x_{t'}) + A_{1/2}(x_{b'})^2, \tag{18}
\]

where \( x_a = m^2_{h_0} / (4 m^2_\phi) \), \( a = t, t', b, b' \), and the loop function \( A_{1/2} \) is defined as given in the Appendix.

The Higgs to diphoton decay is also obtained from loop diagrams. The LO contributions are derived from the one-loop diagrams containing virtual charged gauge bosons \( W^\pm \) or virtual top quarks in the SM. In this model, the additional charged scalar \( \phi^\pm \) and exotic fermions \( t, t', \tau' \) contribute corrections to the decay width of the Higgs to diphoton at LHC. The corresponding expression is written as

\[
\Gamma_{NP}(h_0 \to \gamma \gamma) = \frac{G_F^2 m^3_{h_0}}{128 \sqrt{2} \pi^3} |(Z_{CPE})_{11} |(4 \frac{1}{3} A_{1/2}(x_t)
\]

\[
+ \frac{4}{3} A_{1/2}(x_{t'}) + \frac{1}{3} A_{1/2}(x_{b'})
\]

\[
+ A_{1/2}(x_{t'}) + A_{1/2}(x_{b'})
\]

\[
+ \frac{8 m^2_{W^\pm} m^2_{h_0}}{e^2 m^2_\phi}(\lambda_{h\phi}(Z_{CPE})_{11}
\]

\[
+ \frac{v}{\sqrt{2}} \lambda_{b\phi}(Z_{CPE})_{21}
\]

\[
+ \frac{v}{\sqrt{2}} \lambda_{b\phi}(Z_{CPE})_{31} A_0(x_{\phi^\pm})^2, \tag{19}
\]

the concrete expressions for the loop functions \( A_0, A_1 \).
are given in the Appendix.

The Higgs discoveries from both the ATLAS and CMS experiments have observed an excess in Higgs production and decay into the diphoton channel, which is a factor of 1.4–2 times larger than the SM expectations. The observed signal for the diphoton channels is quantified by the ratio

$$R_{\gamma\gamma} = \frac{\Gamma_{\text{NP}}(h_0 \to gg)}{\Gamma_{\text{SM}}(h_0 \to gg)} \bigg/ \frac{\Gamma_{\text{NP}}(h_0 \to \gamma\gamma)}{\Gamma_{\text{SM}}(h_0 \to \gamma\gamma)},$$  \hspace{1cm} (20)$$

where we assume that all exotic fields are heavier than the lightest Higgs $h_0$. The current value of this ratio is as follows [8, 9]:

ATLAS: $R_{\gamma\gamma} = 1.90 \pm 0.5$,

CMS: $R_{\gamma\gamma} = 1.56 \pm 0.43$,

ATLAS+CMS: $R_{\gamma\gamma} = 1.71 \pm 0.33$.  \hspace{1cm} (21)

Note that the combination of the ATLAS and CMS results is taken from Ref. [10].

4 Correlations to the oblique parameters

A common approach to constrain physics beyond the SM is to use global electroweak fitting through the oblique parameters $S$, $T$, $U$ [1]. In the SM, electroweak precision tests imply a relationship between the oblique parameters and the lightest Higgs $h_0$. The current value of this ratio is as follows [8, 9]:

ATLAS: $R_{\gamma\gamma} = 1.90 \pm 0.5$,

CMS: $R_{\gamma\gamma} = 1.56 \pm 0.43$,

ATLAS+CMS: $R_{\gamma\gamma} = 1.71 \pm 0.33$.  \hspace{1cm} (21)

where we assume that all exotic fields are heavier than the lightest Higgs $h_0$. The current value of this ratio is as follows [8, 9]:

$$R_{\gamma\gamma} = \frac{\Gamma_{\text{NP}}(h_0 \to gg)}{\Gamma_{\text{SM}}(h_0 \to gg)} \bigg/ \frac{\Gamma_{\text{NP}}(h_0 \to \gamma\gamma)}{\Gamma_{\text{SM}}(h_0 \to \gamma\gamma)},$$  \hspace{1cm} (20)$$

where $s_W = \sin \theta_W$ and $c_W = \cos \theta_W$ with the Weinberg angle $\theta_W$ defined at the energy scale $\mu = m_Z$. In the above definitions, $\Pi_{11}$ and $\Pi_{33}$ are the vacuum polarizations of isospin currents, and $\Pi_{3\alpha}$ is the vacuum polarization of one isospin and one hypercharge current.

By comparing the measurable electroweak observables with the theoretical predictions, one finds the fitted values [12]

$$\Delta S = S - S_{\text{SM}} = 0.04 \pm 0.10,$$

$$\Delta T = T - T_{\text{SM}} = 0.05 \pm 0.11,$$

$$\Delta U = U - U_{\text{SM}} = 0.08 \pm 0.11. \hspace{1cm} (23)$$

As mentioned above, there is no mass mixing between the exotic quarks and the SM leptons. The corresponding corrections to the oblique parameters from exotic quarks are

$$\Delta S_{Q'} = \frac{1}{\pi} \left\{ \int_0^1 dx x(1-x) \ln \left( \frac{m_{Q'}^2 - x(1-x)m_{Z}^2}{m_{Q'}^2 - x(1-x)m_{Z}^2} \right) \right\},$$

$$\Delta T_{Q'} = -\frac{3}{4\pi s_W^2 c_W^2} \left\{ \int_0^1 dx x(1-x) \ln \left( \frac{m_{Q'}^2 + (1-x)m_{Z}^2}{m_{Q'}^2 - (1-x)m_{Z}^2} \right) \right\},$$

$$\Delta U_{Q'} = \frac{1}{\pi} \left\{ \int_0^1 dx x(1-x) \ln \left( \frac{m_{Q'}^2 + (1-x)m_{Z}^2}{m_{Q'}^2 - (1-x)m_{Z}^2} \right) \right\}.$$  \hspace{1cm} (24)$$

Here, $m_{Q'}$ and $m_{Q'}$ denote the masses of the charged $-1/3$ exotic quark $b'$ and the charged $2/3$ exotic quark $t'$, respectively.

In a similar way, there is no mass mixing between the exotic charged leptons and the SM leptons. Ignoring the tiny mixing between the left-handed neutrinos and heavy majorana neutrinos, we write the corrections to the oblique parameters from exotic leptons as

\text{the exotic charged leptons and the SM leptons. Ignoring the tiny mixing between the left-handed neutrinos and heavy majorana neutrinos, we write the corrections to the oblique parameters from exotic leptons as}
\[ \Delta S_{L'} = \frac{1}{\pi} \sum_{i,j=1}^{5} (Z_N)_{i1}(Z_N)_{i1}(Z_N)_{i1}(Z_N)_{i1} \left\{ -\frac{1}{2} \int_0^\infty dx \left( \frac{m_{2j}^2}{m_{2j}} + (1-x) \frac{m_{2j}^2}{m_{2j}} \right) \times \ln \frac{x m_{2j}^2 + (1-x) m_{2j}^2}{x m_{2j}^2 + (1-x) m_{2j}^2} \right\} \]

\[ + \frac{1}{2} \int_0^\infty dx (1-x) \ln \frac{x m_{2j}^2 + (1-x) m_{2j}^2}{x m_{2j}^2 + (1-x) m_{2j}^2} \right\} \}

\[ \Delta T_{L'} = -\frac{1}{4\pi s^2} \sum_{i,j=1}^{5} (Z_N)_{i1}(Z_N)_{i1}(Z_N)_{i1}(Z_N)_{i1} \left\{ \frac{m_{2j}^2}{m_{2j}} \int_0^\infty dx \ln \frac{x m_{2j}^2 + (1-x) m_{2j}^2}{x m_{2j}^2 + (1-x) m_{2j}^2} \right\} \]

\[ \Delta U_{L'} = \frac{1}{\pi} \sum_{i,j=1}^{5} (Z_N)_{i1}(Z_N)_{i1}(Z_N)_{i1}(Z_N)_{i1} \left\{ \frac{m_{2j}^2}{m_{2j}} \int_0^\infty dx \ln \frac{x m_{2j}^2 + (1-x) m_{2j}^2}{x m_{2j}^2 + (1-x) m_{2j}^2} \right\} \]

Here, the 5×5 unitary matrix \( Z_N \) is the mixing matrix for heavy majorana neutrinos, \( m_{N_i} \) (i = 1, 2, ..., 5) are the corresponding masses of the heavy neutrinos, and \( m_{\tau'} \) is the mass of the charged lepton \( \tau' \).

Since the radiative corrections to the self energy of gauge bosons originate from three \( CP \)-even Higgs (\( h_0, H^0_2, H^0_3 \)), the corresponding contributions to the oblique parameters are given by

\[ \Delta S_H = \frac{1}{\pi} \sum_{i=1}^{3} (Z_{\text{CPE}})^2_{i1} \left\{ \frac{1}{2} \int_0^\infty dx (1-x) \ln \frac{x^2 m_Z^2 + (1-x) m_{h_0}^2}{m_Z^2} + \frac{1}{2} \int_0^\infty dx \left( 1 - \frac{x}{2} - (1-x) \frac{m_{h_0}^2}{2 m_Z^2} \right) \ln \frac{x^2 m_Z^2 + (1-x) m_{h_0}^2}{x m_Z^2 + (1-x) m_{h_0}^2} \right\} \]

\[ \Delta T_H = -\frac{1}{4\pi s^2} \sum_{i=1}^{3} (Z_{\text{CPE}})^2_{i1} \left\{ -\frac{1}{2} \int_0^\infty dx \left( 1 - \frac{x}{2} - (1-x) \frac{m_{h_0}^2}{2 m_Z^2} \right) \ln \frac{x^2 m_Z^2 + (1-x) m_{h_0}^2}{m_Z^2} \right\} \]

\[ + \frac{1}{2} \int_0^\infty dx \left( (1-x) \frac{2 s^2}{c_W^2} - (1-x) \frac{m_{h_0}^2}{2 m_Z^2} \right) \ln \frac{x m_Z^2 + (1-x) m_{h_0}^2}{m_Z^2} \right\} \}

\[ \Delta U_H = \frac{1}{\pi} \sum_{i=1}^{3} (Z_{\text{CPE}})^2_{i1} \left\{ \frac{1}{2} \int_0^\infty dx \left( x^2 + (1-x) \frac{m_{h_0}^2}{m_Z^2} \right) \ln \frac{x m_Z^2 + (1-x) m_{h_0}^2}{m_Z^2} \right\} \]

\[ - \frac{1}{2} \int_0^\infty dx (1-x) \ln \frac{x m_Z^2 + (1-x) m_{h_0}^2}{x m_Z^2 + (1-x) m_{h_0}^2} \right\} \}

\[ + \frac{1}{2} \int_0^\infty dx \left( (1-x) \frac{2 s^2}{c_W^2} - (1-x) \frac{m_{h_0}^2}{2 m_Z^2} \right) \ln \frac{x m_Z^2 + (1-x) m_{h_0}^2}{x m_Z^2 + (1-x) m_{h_0}^2} \right\} \}

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Here, we adopt the notation $H_0^1$ to represent the lightest neutral Higgs $h_0$. In addition, the contributions from $\Phi^0$ and $\Phi^\pm$ to the oblique parameters are formulated as follows

\[
\Delta S_\Phi = \frac{1}{2\pi} \left\{ \sum_{i,j}^4 (Z_{CPM})^2_{i4}(Z_{CPM})^2_{j4} \left[ -\int_0^1 dx(1-x) \ln \frac{x m_{\phi^0}^4 + (1-x)m_{\phi^0}^2 - x(1-x)m_{\phi^0}^2}{m_{\phi^0}^2} \right] \frac{x m_{\phi^0}^4 + (1-x)m_{\phi^0}^2 - x(1-x)m_{\phi^0}^2}{x m_{\phi^0}^4 + (1-x)m_{\phi^0}^2} \right. \\
+ \int_0^1 dx(1-x) \ln \frac{x m_{\phi^0}^4 + (1-x)m_{\phi^0}^2 - x(1-x)m_{\phi^0}^2}{m_{\phi^0}^2} \right\},
\]

\[
\Delta T_\Phi = \frac{1}{8\pi^2 e^2 W} \left\{ -\sum_{i=1}^4 (Z_{CPM})^2_{i1}(Z_{CPM})^2_{j1} \int_0^1 dx \ln \frac{x m_{\phi^0}^4 + (1-x)m_{\phi^0}^2 - x(1-x)m_{\phi^0}^2}{m_{\phi^0}^2} \right. \\
+ \int_0^1 dx(1-x) \ln \frac{x m_{\phi^0}^4 + (1-x)m_{\phi^0}^2 - x(1-x)m_{\phi^0}^2}{m_{\phi^0}^2} \right\},
\]

\[
\Delta U_\Phi = \frac{1}{2\pi} \left\{ -\sum_{i=1}^4 (Z_{CPM})^2_{i1}(Z_{CPM})^2_{j1} \int_0^1 dx \ln \frac{x m_{\phi^0}^4 + (1-x)m_{\phi^0}^2 - x(1-x)m_{\phi^0}^2}{m_{\phi^0}^2} \right. \\
\right. \\
+ \int_0^1 dx(1-x) \ln \frac{x m_{\phi^0}^4 + (1-x)m_{\phi^0}^2 - x(1-x)m_{\phi^0}^2}{m_{\phi^0}^2} \right\},
\]

5 **Numerical analysis**

As mentioned above, the most stringent constraint on the parameter space is that the 3x3 mass squared matrix in Eq. (5) should produce the lightest eigenvector with a mass $m_{b_0} = 125$ GeV.

In order to make the final results consistent with this condition, we require the self coupling of the Higgs doublet to satisfy

\[
\lambda_{hh} = \frac{A}{B},
\]

where

\[
A = m_{h_0}^6 - 2 \left( \lambda_{BB} v_B^2 + \lambda_{LL} v_L^2 \right) m_{b_0}^4 + \left( 4 \lambda_{BB} \lambda_{LL} - \lambda_{BL}^2 \right) v_B^2 v_L^2 \\
- \lambda_{BB}^2 v_B^2 v_L^2 - \lambda_{LL}^2 v_B^2 v_L^2 \right) m_{b_0}^2 + 2 \lambda_{BB} \lambda_{HL} \lambda_{HB} \\
- \lambda_{HB} \lambda_{HL} \lambda_{BB} \right) v_B^2 v_L^2, \\
B = 2 v^2 \left[ m_{h_0}^4 - 2 \left( \lambda_{BB} v_B^2 + \lambda_{LL} v_L^2 \right) m_{b_0}^2 \\
+ \left( 4 \lambda_{BB} \lambda_{LL} - \lambda_{BL}^2 \right) v_B^2 v_L^2 \right].
\]

The present experimental lower bounds on the fourth generation charged lepton $\tau'$, up-type and down-type quarks $t'$ and $b'$ at 95% C.L. are $m_{\tau'} > 100.8$ GeV, $m_{t'} > 420$ GeV and $m_{b'} > 372$ GeV, respectively. The fourth generation quarks $t'$ and $b'$ acquire nonzero masses $m_{t'} = m_{b'} = \sqrt{E_v} v_{b,\tau}$ when local $U(1)_B$ symmetry is broken. In addition, the charged leptons of the fourth generation $\tau'$ obtain nonzero masses $m_{\tau'} = \sqrt{E_v} v_{\tau}$ when local $U(1)_L$ symmetry is broken.

However, there are too many free parameters in the model considered here. In our numerical analysis, we adopt the assumption on the parameter space

\[
m_{t'} = m_{b'} = m_{\tau'} = m_{\nu}, \\
\lambda_{BB} = \lambda_{LL} = 0.5, \quad \lambda_{HL} = \lambda_{BL} = \lambda_{HB} = \lambda_{NP}, \\
\lambda_{h_0} = \lambda_{\phi B} = \lambda_{\phi L} = \lambda_{\phi N},
\]

(30)

to decrease the number of free parameters in the concerned model. Furthermore, we assume $v \ll v_{b,\tau}$.
(λ_a)_{3x3} = \text{diag}(λ_a, λ_b, λ_c), and choose the hierarchical assumption on Yukawa couplings \(|Y_i| \ll |Y_j| \sim λ_jb_j, (I, J = 1, 2, 3)| to obtain our final results. Applying the assumptions above, we obtain the majorana mass for the lightest exotic neutrino N_1 to be

\[ m_{N_1} \approx \frac{v^2 \lambda_a |Y_e^r|}{2v_L \lambda_b^2}, \]

(31)

with \( \lambda_b^3 = \lambda_b^3 + \lambda_b^2 + λ_b^2 \). Of course, we need this mass to be greater than \( m_2/2 \) in order to be consistent with the measured Z-boson decay width. The masses of other heavy majorana neutrinos are

\[ m_{N_i} \approx \left( \frac{v_L}{\sqrt{2}} λ_a, \frac{v_L}{\sqrt{2}} λ_b (Δ−λ_b), \frac{v_L}{\sqrt{2}} (Δ+λ_b) \right), \]

(32)

for \( i=2, 3, 4, 5 \) and \( Δ = \sqrt{λ_b^2 + 4 λ_b^2} \).

Correspondingly, the \( 5 \times 5 \) mixing matrix \( Z_N \) is approximated as

\[
Z_N \approx \begin{pmatrix}
1, & \frac{λ_a Y_e^r v_L}{λ_b^2 v_L}, & \frac{λ_b Y_e^r v_L}{λ_b^2 v_L}, & \frac{λ_b Y_e^r v_L}{λ_b^2 v_L}, & \frac{λ_b Y_e^r v_L}{λ_b^2 v_L} \\
\frac{λ_a Y_e^r v_L}{λ_b^2 v_L}, & 0, & 0, & -i\sqrt{2+λ_b}, & -i\sqrt{2−λ_b} \\
-\frac{λ_b Y_e^r v_L}{λ_b^2 v_L}, & -\frac{λ_b Y_e^r v_L}{λ_b^2 v_L}, & \frac{λ_b Y_e^r v_L}{λ_b^2 v_L}, & \frac{λ_b Y_e^r v_L}{λ_b^2 v_L}, & \frac{λ_b Y_e^r v_L}{λ_b^2 v_L} \\
-\frac{λ_b Y_e^r v_L}{λ_b^2 v_L}, & 0, & 0, & i\sqrt{2+λ_b}, & i\sqrt{2−λ_b} \\
-\frac{λ_b Y_e^r v_L}{λ_b^2 v_L}, & -\frac{λ_b Y_e^r v_L}{λ_b^2 v_L}, & -\frac{λ_b Y_e^r v_L}{λ_b^2 v_L}, & -\frac{λ_b Y_e^r v_L}{λ_b^2 v_L}, & \frac{λ_b Y_e^r v_L}{λ_b^2 v_L} \\
\end{pmatrix}
\]

(33)

For the relevant parameters in the SM, we take [13]

\[ α(m_Z) = 0.118, \quad α(m_Z) = 1/128, \]

\[ s_W^2(m_Z) = 0.23, \quad m_t = 174.2 \text{ GeV}, \quad m_W = 80.4 \text{ GeV}. \]

5.1 Numerical results of \( R_{γγ} \)

Under our assumptions of the parameter space, the theoretical prediction of \( R_{γγ} \) depends on six parameters in the model: \( m_F \), \( m_{ϕ^±} \), \( λ_{NP} \), \( λ'_{NP} \), \( v_B \) and \( v_L \). Taking \( m_{ϕ^±} = 500 \text{ GeV}, \quad λ_{NP} = 0.5 \) and \( λ'_{NP} = 0.5 \), we plot the variation of \( R_{γγ} \) with the mass scalar of exotic fermions \( M_F \), as shown in Fig. 1. The dotted line corresponds to \( v_B = v_L = 500 \text{ GeV} \), the dashed line corresponds to \( v_B = v_L = 1000 \text{ GeV} \), and the solid line corresponds to \( v_B = v_L = 1500 \text{ GeV} \). In general, the ratio \( R_{γγ} \) depends very weakly on the mass scale \( m_{ϕ^±} \), and the value of \( R_{γγ} \) is about 1.8–1.9 when \( 500 \text{ GeV} \leq v_B \leq v_L \leq 1500 \text{ GeV} \).

In Fig. 2(a), we plot the variation of \( R_{γγ} \) with the VEV \( v_L \) when \( m_{ϕ^±} = 500 \text{ GeV} \), \( λ'_{NP} = −0.5 \) and \( λ_{NP} = 0.5 \). The dotted line corresponds to \( m_F = 500 \text{ GeV} \), the dashed line corresponds to \( m_F = 550 \text{ GeV} \), and the solid line corresponds to \( m_F = 600 \text{ GeV} \). The dependence of \( R_{γγ} \) on \( v_L \) is relatively sensitive for \( v_L \leq 600 \text{ GeV} \) and is weak for \( v_L > 600 \text{ GeV} \). Since the dependence of \( R_{γγ} \) on \( m_F \) and \( v_B \) is very weak, the three lines almost coincide with each other. In Fig. 2(b), we plot the variation of \( R_{γγ} \) with the VEV \( v_B \) when \( m_F = m_{ϕ^±} = 500 \text{ GeV} \), \( λ'_{NP} = 0.5 \). The dotted line corresponds to \( λ_{NP} = 0.5 \), the dashed line corresponds to \( λ_{NP} = 0 \), and the solid line corresponds to \( λ_{NP} = −0.5 \). Generally, there is a weak dependence of the ratio \( R_{γγ} \) on \( v_B \).

In Fig. 3(a), we show the variation of \( R_{γγ} \) with the VEV \( v_B \) when \( m_F = v_L = 500 \text{ GeV} \), \( λ_{NP} = 0.5 \). The dotted line corresponds to \( λ'_{NP} = 0.5 \), the dashed line corresponds to \( λ'_{NP} = 0 \), and the solid line corresponds to \( λ'_{NP} = −0.5 \). The dependence of \( R_{γγ} \) on \( v_B \) is relatively sensitive for \( v_B \leq 600 \text{ GeV} \) and is weak for \( v_B > 600 \text{ GeV} \). In Fig. 3(b), we show the variation of \( R_{γγ} \) with \( v_B \) when \( m_F = v_L = 500 \text{ GeV}, \quad λ'_{NP} = −0.5 \) and \( λ_{NP} = 0.5 \). The dotted line corresponds to \( m_{ϕ^±} = 500 \text{ GeV} \), the dashed line
R presents the variation of the ratio corresponds to $m_t$ the dashed line represents $R$ weak dependence of the ratio represents $\lambda$. Thedotted line represents $R$ dependence of experimental data, as $\nu_\gamma$ and can easily coincide with the present experimental data, as $\nu_\gamma$. 

Choosing $v_B = v_L = 500$ GeV, $\chi'_{NP} = -0.5$, Fig. 4 presents the variation of the ratio $R_{\gamma \gamma}$ with $\lambda_{NP}$. The dotted line represents $m_F = 500$ GeV, $m_{\phi \pm} = 1500$ GeV, the dashed line represents $m_F = 550$ GeV, $m_{\phi \pm} = 1000$ GeV, and the solid line represents $m_F = m_{\phi \pm} = 500$ GeV. As $\Lambda_{NP}$ increases, $R_{\gamma \gamma}$ changes drastically and can easily coincide with the present experimental data, as $-0.5 \leq \Lambda_{NP} \leq 1.0$. Choosing $m_F = v_B = v_L = 500$ GeV, and $\Lambda_{NP} = -0.5$, Fig. 5 shows the ratio $R_{\gamma \gamma}$ versus $m_{\phi \pm}$. The dotted line represents $\chi'_{NP} = 0.5$, the dashed line represents $\chi'_{NP} = 0$, and the solid line represents $\chi'_{NP} = -0.5$. For $\chi'_{NP} = 0$, there is a slight dependence of $R_{\gamma \gamma}$ on the mass $m_{\phi \pm}$. When $\chi'_{NP} = \pm 0.5$, $R_{\gamma \gamma}$ decreases steeply as $m_{\phi \pm}$ increases. 

In Fig. 6, we plot the variation of the ratio $R_{\gamma \gamma}$ with $\chi'_{NP}$ when $m_F = v_B = v_L = 500$ GeV and $\Lambda_{NP} = -0.5$. The dotted line represents $m_{\phi \pm} = 1500$ GeV, the dashed line represents $m_{\phi \pm} = 1000$ GeV, and the solid line represents $m_{\phi \pm} = 500$ GeV. The dependence of $R_{\gamma \gamma}$ on $\chi'_{NP}$ is strong when $m_{\phi \pm} = 500$ GeV but weaker for higher values of $m_{\phi \pm}$.

Generally, the ratio $R_{\gamma \gamma}$ depends strongly on the parameters $\Lambda_{NP}$, $\chi'_{NP}$ and $m_{\phi \pm}$, and depends weakly on $v_B$, $v_L$ and $m_F$. These numerical results can be reasonably explained from Eq. (18) and Eq. (19), where $\Lambda_{NP}$ affects theoretical predictions of $R_{\gamma \gamma}$ through the $3 \times 3$ mixing matrix $Z_{CP}$, while $\chi'_{NP}$ and $m_{\phi \pm}$ affect theoretical predictions of $R_{\gamma \gamma}$ through the last term in Eq. (19).
Fig. 4. Variation of $R_{\gamma\gamma}$ with $\lambda_{NP}$ when $\upsilon_B = \upsilon_L = 500$ GeV, $\lambda_{NP} = -0.5$, where the dotted line represents $m_F = 500$ GeV, $m_{\phi^\pm} = 1500$ GeV, the dashed line represents $m_F = 550$ GeV, $m_{\phi^\pm} = 1000$ GeV, and the solid line represents $m_F = m_{\phi^\pm} = 500$ GeV.

The important point is that the parameters $\lambda_a$, $\lambda_b_i$ ($i=1, 2, 3$) do not affect the theoretical predictions of $R_{\gamma\gamma}$ since there is no correction to the decay widths of $h_0 \to \gamma\gamma$ and $h_0 \to gg$ from the neutrino sector at one-loop level. Similarly, the parameters $M_S, \mu_1, \mu_2$ also do not affect theoretical evaluations of $R_{\gamma\gamma}$ because there is no one-loop correction to the decay widths of $h_0 \to \gamma\gamma$ and $h_0 \to gg$ from virtual $\Phi^0_i$ ($i=1, 2, 3, 4$).

5.2 The constraints on parameter space from oblique corrections

The heavy neutrinos contribute one-loop radiative corrections to the self energies of $ZZ, W^\pm W^\mp$ in this model. This results in the theoretical values of the $S, T, U$ parameters depending on $\lambda_a, \lambda_b_i$ ($i=1, 2, 3$) here. Furthermore, the theoretical values of the $S, T, U$ parameters also depend on $M_S, \mu_1, \mu_2$ through the virtual $\Phi^\pm, \Phi^0_i$ ($i=1, 2, 3, 4$) radiative corrections to the self energies of $ZZ, W^\pm W^\mp$ at one-loop level. So far, fitting
In order to obtain theoretical values of $S, T, U$ within $3\sigma$ deviation indicates

\[-0.26 \leq \Delta S \leq 0.34,\]
\[-0.28 \leq \Delta T \leq 0.38,\]
\[-0.25 \leq \Delta U \leq 0.41.\]  

(35)

In order to obtain theoretical values of $S, T, U$ that satisfy present experimental data, we adopt the following additional assumptions:

\[m_{N_1} \approx \frac{v^2 \lambda_1 Y_1'}{\sqrt{2} v_L} \lambda_0 = m_F,\]
\[\lambda_{b_1} = \lambda_{b_2} = \lambda_{b_3} = \frac{1}{\sqrt{3}} \lambda_0,\]
\[v_B = v_L = m_S = m_F = 500 \text{ GeV},\]
\[|\mu_1| = 20 \text{ GeV}, |\mu_2| = 200 \text{ GeV},\]
\[\lambda = \lambda_b = 0.6, \lambda_{bb} = \lambda_{ll} = 0.5, \lambda_{np} = \lambda'_{np} = 0.01.\]  

(36)

Choosing $\theta_1 = \text{arg}(\mu_1) = \pi, \theta_2 = \text{arg}(\mu_2) = \pi/4$, we depict the theoretical values of $\Delta S, \Delta U, \Delta T$ versus the mass of charged scalar $\phi^\pm$ in Fig. 7, in which the solid line represents $\Delta S$, the dash-dot-dot line represents $\Delta U$, and the dashed line represents $\Delta T$. For our choices of the relevant parameters, the theoretical value of $\Delta T$ is very sensitive to the mass $m_{\phi^\pm}$, while the theoretical values of $\Delta S$ and $\Delta U$ have a weak dependence on the mass $m_{\phi^\pm}$. When the mass of the charged scalar lies in the range $400 \leq m_{\phi^\pm}/\text{GeV} \leq 700$, the theoretical predictions of $\Delta S, \Delta T, \Delta U$ simultaneously satisfy the inequalities in Eq. (35). The CP phases $\theta_1, \theta_2$ also affect the numerical results of $\Delta S, \Delta U, \Delta T$ through the $4 \times 4$ mixing matrix $Z_{CPM}$. Taking $m_{\phi^\pm} = 600 \text{ GeV}$ and $\theta_2 = \pi/4$, we present the theoretical evaluations on $\Delta S, \Delta U, \Delta T$ versus the CP phase $\theta_1$ in Fig. 8. With our assumptions on the parameter space, the theoretical value of $\Delta T$ varies strongly with the CP phase $\theta_1$, while the theoretical values of $\Delta S$ and $\Delta U$ vary weakly with the CP phase $\theta_1$. In the neighbourhoods of $\theta_1 = 0, \pm \pi/2, \pm \pi$, the theoretical predictions on $\Delta S, \Delta T, \Delta U$ simultaneously lie within the ranges presented in Eq. (35).

In Fig. 9, we present the theoretical values of $\Delta S, \Delta T, \Delta U$ varying with the CP phase $\theta_2$ when $m_{\phi^\pm} = 600 \text{ GeV}$ and $\theta_1 = \pi$. As the CP phase $\theta_2$ varies, the theoretical value of $\Delta T$ changes drastically, while the theoretical values of $\Delta S$ and $\Delta U$ change slowly. In the neighbourhoods around $\theta_2 = \pm \pi/4, \pm 3\pi/4$, the theoretical predictions of $\Delta S, \Delta T, \Delta U$ coincide with the present global EWPD fit within $3\sigma$ deviations.

6 Summary

For an extension of the SM with local gauged baryon and lepton numbers, we have discussed the constraints from the oblique parameters $S, T, U$ when the lightest Higgs has a mass around 125 GeV. Considering these constraints, we find that there is parameter space to account for the excess in Higgs production and decay in the diphoton channel observed in the ATLAS and CMS experiments. Of course, our numerical results strongly depend on the assumptions made in the model considered here. In other words, our theoretical prediction cannot be precise because of the theoretical uncertainties. The purpose of our calculation is to show that this extension of the SM may still be right even after the constraints from LHC data on the Higgs and oblique parameters have been taken into account.
Appendix A

Higgs masses and relevant couplings

After diagonalizing the relevant matrix in Eq. (5), we obtain

\[
m_1^2 = \text{Min}(m_i^2, m_j^2, m_k^2)
\]
\[
m_2^2 = \text{Max}(m_i^2, m_j^2, m_k^2)
\]

(A1)

with

\[
m_2 = -\frac{a}{3} + \frac{2}{3} p \cos \phi,
\]
\[
m_3 = \frac{1}{3} p \cos (\phi - \sqrt{3} \sin \phi),
\]
\[
m_3 = \frac{1}{3} p \cos (\phi + \sqrt{3} \sin \phi).
\]

(A2)

To formulate the expressions in a concise form, we define the notations

\[
p = \sqrt{a^2 - 3b}, \phi = \frac{1}{3} \arccos \left( -\frac{1}{p} \left( a^3 - \frac{9}{2} ab + \frac{27}{2} c \right) \right)
\]

where

\[
a = -2 \lambda_{hh} v_1^2 + \lambda_{bb} v_3^2 + \lambda_{ll} v_1^2,
\]
\[
b = 4 \lambda_{hh} \lambda_{bb} v_2^2 + \lambda_{hh} \lambda_{ll} v_2^2 + \lambda_{bb} \lambda_{ll} v_2 v_3,
\]
\[
c = -2 \lambda_{bb} v_2^2 + \lambda_{ll} v_2 v_3 - \lambda_{hh} v_3^2
\]
\[
\lambda_{hh} \lambda_{ll} v_1^2 + 4 \lambda_{hh} \lambda_{bb} \lambda_{ll}
\]
\[
- \lambda_{hh} \lambda_{ll} h \lambda_{bb} v_1^2 v_3^2.
\]

(A3)

The normalized eigenvectors of the mass squared matrix in Eq. (5) are given by

\[
\begin{pmatrix}
(Z_{CPE})_{11} \\
(Z_{CPE})_{21} \\
(Z_{CPE})_{31}
\end{pmatrix} = \frac{1}{\sqrt{|X_1|^2 + |Y_1|^2 + |Z_1|^2}} \begin{pmatrix}
X_1 \\
Y_1 \\
Z_1
\end{pmatrix}
\]

(A4)

with

\[
X_1 = (2 \lambda_{bb} v_2^2 - m_1^2)(2 \lambda_{ll} v_2^2 - m_1^2) - \lambda_{bb} v_2^2 v_3^2,
\]
\[
Y_1 = \lambda_{hh} \lambda_{ll} v_2 v_3 - \lambda_{bb} v_2 v_3 (2 \lambda_{ll} v_2^2 - m_1^2),
\]
\[
Z_1 = \lambda_{hh} \lambda_{ll} v_2 v_3 - \lambda_{bb} v_2 v_3 (2 \lambda_{bb} v_2^2 - m_1^2),
\]
\[
X_2 = (2 \lambda_{bb} v_2^2 - m_2^2)(2 \lambda_{ll} v_2^2 - m_2^2) - \lambda_{bb} v_2^2 v_3^2,
\]
\[
Y_2 = \lambda_{hh} \lambda_{ll} v_2 v_3 - \lambda_{bb} v_2 v_3 (2 \lambda_{ll} v_2^2 - m_2^2),
\]
\[
Z_2 = \lambda_{hh} \lambda_{ll} v_2 v_3 - \lambda_{bb} v_2 v_3 (2 \lambda_{bb} v_2^2 - m_2^2),
\]
\[
X_3 = (2 \lambda_{bb} v_2^2 - m_3^2)(2 \lambda_{ll} v_2^2 - m_3^2) - \lambda_{bb} v_2^2 v_3^2,
\]
\[
Y_3 = \lambda_{hh} \lambda_{ll} v_2 v_3 - \lambda_{bb} v_2 v_3 (2 \lambda_{ll} v_2^2 - m_3^2),
\]
\[
Z_3 = \lambda_{hh} \lambda_{ll} v_2 v_3 - \lambda_{bb} v_2 v_3 (2 \lambda_{bb} v_2^2 - m_3^2).
\]

(A5)

Appendix B

The loop functions

The loop functions in Eq. (18) and Eq. (19) are given as

\[
A_1(x) = -\left[ 2x^2 + 3x + 3 \right] g(x) / x^2,
\]
\[
A_{1/2}(x) = 2 \left( x + (x-1) g(x) \right) / x^2,
\]
\[
A_0(x) = -(x - g(x)) / x^2,
\]

with

\[
g(x) = \begin{cases}
\text{arcsin} \sqrt{x} x \leq 1 \\
\frac{1}{4} \left\{ \frac{1}{1 + \sqrt{1 - 1/x} - x} \right. \\
\left. \frac{1}{1 - \sqrt{1 - 1/x} - x} \right\}^2, x > 1
\end{cases}
\]

(B2)

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