Dynamics of levitated nanospheres: towards the strong coupling regime

T S Monteiro\textsuperscript{1,5}, J Millen\textsuperscript{1}, G A T Pender\textsuperscript{1}, Florian Marquardt\textsuperscript{2,3}, D Chang\textsuperscript{4} and P F Barker\textsuperscript{1}

\textsuperscript{1}Department of Physics and Astronomy, University College London, Gower Street, London WC1E 6BT, UK
\textsuperscript{2}Institute for Theoretical Physics, Universität Erlangen-Nürnberg, Staudtstraße 7, D-91058 Erlangen, Germany
\textsuperscript{3}Max Planck Institute for the Science of Light, Günther-Scharowsky-Straße 1/Bau 24, D-91058 Erlangen, Germany
\textsuperscript{4}ICFO—Institut de Ciencies Fotoniques, Av. Carl Friedrich Gauss, E-308860 Castelldefels, Barcelona, Spain
E-mail: t.monteiro@ucl.ac.uk

\textit{New Journal of Physics} \textbf{15} (2013) 015001 (21pp)
Received 3 July 2012
Published 3 January 2013
Online at http://www.njp.org/
doi:10.1088/1367-2630/15/1/015001

\textbf{Abstract.} The use of levitated nanospheres represents a new paradigm for the optomechanical cooling of a small mechanical oscillator, with the prospect of realizing quantum oscillators with unprecedentedly high quality factors. We investigate the dynamics of this system, especially in the so-called self-trapping regime, where one or more optical fields simultaneously trap and cool the mechanical oscillator. The determining characteristic of this regime is that both the mechanical frequency $\omega_M$ and single-photon optomechanical coupling strength parameters $g$ are a function of the optical field intensities, in contrast to usual set-ups where $\omega_M$ and $g$ are constant for the given system. We also measure the characteristic transverse and axial trapping frequencies of different sized silica nanospheres in a simple optical standing wave potential, for spheres of radii $r = 20–500$ nm, illustrating a protocol for loading single nanospheres into a standing wave optical trap that would be formed by an optical cavity.
We use these data to confirm the dependence of the effective optomechanical coupling strength on sphere radius for levitated nanospheres in an optical cavity and discuss the prospects for reaching regimes of strong light–matter coupling. Theoretical semiclassical and quantum displacement noise spectra show that for larger nanospheres with $r \gtrsim 100 \text{ nm}$ a range of interesting and novel dynamical regimes can be accessed. These include simultaneous hybridization of the two optical modes with the mechanical modes and parameter regimes where the system is bistable. We show that here, in contrast to typical single-optical mode optomechanical systems, bistabilities are independent of intracavity intensity and can occur for very weak laser driving amplitudes.

Contents

1. Introduction ........................................ 2

2. Theory: quantum Hamiltonian for a nanosphere in a cavity .......... 4

3. Towards strong light–matter coupling with two optical cavity modes .... 5

4. Dynamics .......................................................... 7

  4.1. Optomechanical damping ........................................ 7

  4.2. Quantum and semiclassical noise spectra ......................... 10

  4.3. Comparison between perturbation theory, semiclassical and quantum results .................................................. 12

5. Strong coupling regimes: triple mode splitting and bistability ...... 13

6. Experiment ......................................................... 16

  6.1. Current experimental status: loading protocols and variation of trap frequency with radius ........................................ 16

7. Conclusions ...................................................... 20

Acknowledgments ................................................. 20

References .......................................................... 20

1. Introduction

Extraordinary progress has been made in the last half-dozen years [1–4] towards the goal of cooling a small mechanical resonator down to its quantum ground state and hence to realize quantum behaviour in a macroscopic system. Implementations include cavity cooling of micromirrors on cantilevers [5–8]; dielectric membranes in Fabry–Perot cavities [9]; radial and whispering gallery modes of optical microcavities [10] and nano-electromechanical systems [11–13]. Indeed the realizations span 12 orders of magnitude in size [2], up to and including the Laser Interferometer Gravitational-wave Observatory (LIGO) gravity wave experiments. In 2011 two separate experiments [14, 15] achieved sideband cooling of micromechanical and nanomechanical oscillators to the quantum ground state. In [16], spectral signatures (in the form of asymmetric displacement noise spectra) of quantum ground state cooling were further investigated. Corresponding advances in the theory of optomechanical cooling have also been made [17–21].

Over the last year or so, a promising new paradigm has been attracting much interest: several groups [22–26, 28] have now investigated schemes for optomechanical cooling of levitated dielectric particles, including nanospheres, microspheres and even viruses. The important
advantage is the elimination of the mechanical support, a dominant source of environmental
noise which can heat and decohere the system.

In general, these proposals involve two fields, one for trapping and one for cooling. This
may involve an optical cavity mode plus a separate trap, or two optical cavity modes: the
so-called ‘self-trapping’ scenario.

Mechanical oscillators in the self-trapping regime differ from other optomechanically
cooled devices in a second fundamental respect (in addition to the absence of mechanical
support): the mechanical frequency, $\omega_M$, associated with centre-of-mass oscillations is not an
intrinsic feature of the resonator but is determined by the optical field. In particular, it is a
function of one or both of the detuning frequencies, $\Delta_1$ and $\Delta_2$, of the optical modes. In
previous work [28], we analysed cooling in the self-trapped regime and found that the optimal
condition for cooling occurs where both fields competitively cool and trap the nanosphere.

This happens when $\omega_M$ is resonantly red detuned from both the detuning frequencies i.e.
$\omega_M(\Delta_1, \Delta_2) \sim -\Delta_{1,2}$ so the relevant resonant frequencies are mutually interdependent. Most
significantly, the effective light–matter coupling strength $g$ also depends on the detunings.
However, the quantum dynamics and in particular the behaviour of the quantum noise spectra
in the self-trapped regime, were not previously considered.

The effective coupling strength, $\tilde{g} = g\sqrt{n}$ (the optomechanical coupling rescaled by
the square root of photon number) determines whether one can attain strong coupling regimes
in levitated systems such as recently observed in a non-levitated set-up [29]. There is
currently intense interest in experimental investigation of strong and ultra-strong light–matter
coupling [30–34]. It determines too whether one may access other interesting dynamics, both
in the semiclassical and quantum regimes. In the present work, we investigate the possibility
of simultaneous hybridization of the two optical modes with the mechanical mode; we show
also that novel strong-coupling regimes can be present in the presence of a static bistability. We
show also that bistability occurs for quite weak driving (low photon numbers) in the levitated
self-trapped system.

We investigate theoretically and experimentally the strength of the optomechanical
coupling. In particular, we present experimental measurements of the mechanical frequency
of a nanosphere trapped in an optical standing wave in order to investigate the optical coupling
as a function of the size of the nanosphere.

In section 2 we review the theory of the cavity cooling and dynamics of a self-trapped
system, and in section 3 we employ the experimentally measured size dependence of the
coupling to determine the range of optomechanical coupling strengths accessible in a cavity.
The data suggest that the most effective means to attain stronger coupling will be to employ
larger nanospheres of radii up to $r \approx 200$ nm. Our work suggests that increasing photon number
by stronger driving (and by implication increasing rescaled coupling strengths) will not prove
an effective alternative, since in the present system we show $\tilde{g} \propto n^{1/4}$ rather than $\sqrt{n}$, so the
rescaled coupling increases very slowly with laser input power.

In section 4, we investigate the cooling and dynamics. In section 4.1, we review the
corresponding cooling rate expressions obtained from quantum perturbation theory (or linear
response theory). In section 4.2 we report a study of the corresponding semiclassical Langevin
equations and compare them with fully quantum noise spectra; we compare also quantum,
semiclassical and perturbation theory results for levitated nanospheres. In section 5 we
investigate novel regimes of triple mode hybridization, coincident with static bistabilities, which
the present study shows are experimentally accessible given the large optomechanical coupling
strengths associated with \( r = 100–200 \) nm nanospheres. The requirement to avoid excessive recoil heating and scattering places an upper limit on the sphere radius. In section 6 we describe the experimental study which provides data from which the size-dependence of the coupling may be inferred. In section 7, we conclude.

2. Theory: quantum Hamiltonian for a nanosphere in a cavity

We approximate the equivalent cavity model by a one-dimensional system, with centre-of-mass motion confined to the axial dimension. In this simplified study, we consider only the axial dynamics: for the cavity system, we will have a much smaller transverse frequency relative to the axial frequency, i.e. \( \omega_t \ll \omega_a \), and there is little mixing between these transverse and axial degrees of freedom.

We consider the dynamics of the following Hamiltonian:

\[
\frac{\hat{H}}{\hbar} = -\Delta_1 \hat{a}_1^\dagger \hat{a}_1 - \Delta_2 \hat{a}_2^\dagger \hat{a}_2 + \frac{\hat{P}^2}{2m\hbar} - A \hat{a}_1^\dagger \hat{a}_1 \cos^2(k_1 \hat{x} - \phi_1) \\
- A \hat{a}_2^\dagger \hat{a}_2 \cos^2(k_2 \hat{x} - \phi_2) + E_1(\hat{a}_1^\dagger + \hat{a}_1) + E_2(\hat{a}_2^\dagger + \hat{a}_2).
\]  

Two optical field modes of a high finesse cavity \( \hat{a}_{1,2} \) are coupled to a nanosphere with centre-of-mass position \( x \). The parameter \( A \) (dependent on the nanosphere polarizability), determines the depth of the optical standing-wave potentials. We investigate the case where both modes competitively cool and trap the nanosphere, in contrast to previous schemes [22, 23] where one optical field is exclusively responsible for trapping, while the other is exclusively responsible for cooling. \( \hat{H} \) is given in the rotating frame of the laser which drives the modes with amplitudes \( E_1 \) and \( E_2 \) respectively, where \( R \) represents the ratio of driving amplitudes for the two modes. We restrict ourselves to the regime \( 0 \leq R \leq 1 \), since we consider the most general case where both optical modes contribute to the trapping as well as the cooling. Thus we can define mode 1 simply as the mode which is more strongly driven and mode 2 as the mode which is (except where \( R = 1 \)) more weakly driven. The detunings \( \Delta_j = \omega_j - \omega_c \) for \( j = 1, 2 \) are between the input lasers and the corresponding cavity mode of interest, and \( \phi_{1,2} \) represents the phases of the optical potentials.

The two fields could represent two modes generated by the same laser field, or they could be generated by two independent lasers. Nonetheless, since the particle motion is confined to within one wavelength, one can make the approximation \( k_1 \approx k_2 \approx k \). Previous studies generally consider \( \phi_1 = 0, \phi_2 = \pi / 4 \) to be convenient, since then the anti-node of one field coincides with a purely linear potential of the other optical field, but we may also consider general values of \( \phi_1 - \phi_2 \). One can write corresponding equations of motion:

\[
\begin{align*}
\ddot{x} &= -\frac{\hbar k A}{m} \sum_j \hat{a}_j^\dagger \hat{a}_j \sin 2(k \hat{x} - \phi_j) - \Gamma_M \dot{x}, \\
\dot{\hat{a}}_j &= i \Delta_j \hat{a}_j - i E_j + i A \hat{a}_j \cos^2(k \hat{x} - \phi_j) - \frac{\kappa}{2} \hat{a}_j,
\end{align*}
\]  

where \( j = 1, 2 \) for the two optical-mode realization. Additional damping terms have also been added: \( \xi \hat{a}_1 \) accounts for photon losses due to mirror imperfections and the \( \Gamma_M \dot{x} \) term for mechanical damping. The above should also include quantum noise terms arising from (say) shot noise or gas collisions: for brevity, the quantum noise terms are left out until section 4.
We consider here the linearized dynamics; we replace operators by their expectation values and perform the shifts about equilibrium values such as \( \hat{a}_j(t) \rightarrow \alpha_j + \hat{a}_j(t) \), and \( \hat{x} \rightarrow x_0 + \hat{x}(t) \). The values for the equilibrium photon fields (e.g. for the two-mode case) are:

\[
\alpha_1 = -iE_1 \left[ \frac{\pi}{2} - i \Delta_j^1 \right]^{-1} \quad \text{and} \quad \alpha_2 = -iRE_1 \left[ \frac{\pi}{2} - i \Delta_j^2 \right]^{-1}.
\]

The equilibrium position is then given by the relation

\[
- \frac{\sin 2(kx - \phi_1)}{\sin 2(kx - \phi_2)} = |\alpha_2|^2 / |\alpha_1|^2,
\]

by numerical solution of the equation

\[
- \frac{\sin 2(kx - \phi_1)}{\sin 2(kx - \phi_2)} = R^j \left[ \frac{\pi}{2} - i \Delta_j^1 \right]^{-2} \left[ \frac{\pi}{2} - i \Delta_j^2 \right]^{-2},
\]

where \( \Delta_j^s = \Delta_j + A \cos^2(kx_0 - \phi_j) \). As usual we consider the dynamics of the fluctuations via the linearized equations. To first order, the linearized equations of motion, in the shifted frame, are:

\[
\ddot{x} = -\omega_M^2 \dot{x} - \frac{\hbar k A}{m} \sum_j (\alpha_j^* \dot{a}_j + \alpha_j \dot{a}_j^*) \sin 2(kx_0 - \phi_j) - \Gamma_M \dot{x},
\]

\[
\dot{a}_j = i \Delta_j^s \dot{a}_j - i k A \alpha_j \dot{x} \sin 2(kx_0 - \phi_j) - \frac{\kappa}{2} \dot{a}_j.
\]

The resulting effective mechanical harmonic oscillator frequency is

\[
\omega_M^2 = \frac{2\hbar k A^2}{m} \sum_j |\alpha_j|^2 \cos 2(kx_0 - \phi_j).
\]

We can restrict ourselves to real equilibrium fields. We take \( \alpha_j = \tilde{\alpha}_je^{-i\theta_j} \) then transform \( \hat{a}_j \rightarrow \hat{a}_je^{i\theta_j} \). Thus \( (\alpha_j^* \dot{a}_j + \alpha_j \dot{a}_j^*) = \tilde{\alpha}_j (\dot{\tilde{a}}_j + \dot{\tilde{a}}_j^*) \). We also rescale the mechanical oscillator coordinates \( \tilde{x} \rightarrow \sqrt{2}X_{zpf} \tilde{x} \) and \( \tilde{p} \rightarrow \sqrt{\hbar m \omega_M} \tilde{p} \), where \( X_{zpf} = \sqrt{\frac{\hbar}{2m\omega_M}} \) is the zero-point fluctuation length scale. Hence, \( \frac{\tilde{p}^2}{2m} + \frac{1}{2}m \omega_M^2 \tilde{x}^2 \rightarrow \frac{\hbar \omega_M}{2} (\tilde{x}^2 + \tilde{p}^2) \).

Below we drop the tilde so the equilibrium field values \( \tilde{\alpha}_j \equiv \alpha_j \) are real. Using field operators \( \tilde{x} = (\tilde{b} + \tilde{b}^\dagger) / \sqrt{2} \), the linearized dynamics for a two-optical mode system would correspond to an effective Hamiltonian:

\[
\frac{\hat{H}_{1,\text{lin}}}{\hbar} = -\Delta_j^s \tilde{a}_j^\dagger \tilde{a}_j - \Delta_j^2 \tilde{a}_2^\dagger \tilde{a}_2 + \omega_M (\Delta_j^1, \Delta_j^2) \tilde{b}^\dagger \tilde{b}
\]

\[
+ g_1 (\Delta_j^1, \Delta_j^2) \alpha_1 (\tilde{a}_1 + \tilde{a}_1^\dagger) (\tilde{b} + \tilde{b}^\dagger) + g_2 (\Delta_j^1, \Delta_j^2) \alpha_2 (\tilde{a}_2 + \tilde{a}_2^\dagger) (\tilde{b} + \tilde{b}^\dagger).
\]

3. Towards strong light–matter coupling with two optical cavity modes

The Hamiltonian in equation (6) appears analogous in form to standard, well-studied optomechanical Hamiltonians, albeit with two optical modes rather than one. However, it differs in one important respect: in this case, both the mechanical frequency \( \omega_M (\Delta_j^1, \Delta_j^2) \) and the optomechanical coupling strengths \( g_{1,2} \equiv g_{1,2} (\Delta_j^1, \Delta_j^2) \) are not fixed and depend on the detunings. The fact that the frequencies of the three modes (two optical, one mechanical) are interdependent makes the dynamics different from other optomechanical set-ups, where the equilibrium mechanical frequency (i.e. excluding shifts arising from the fluctuations) is intrinsic to the mechanical oscillator.

There is considerable interest in achieving strong coupling, which leads to regimes of light–matter hybridization. The corresponding mode splitting has been observed experimentally...
[29]. In typical set-ups, these regimes are reached if the rescaled effective optomechanical coupling exceeds the damping rates, i.e. \( \tilde{g} = g \sqrt{n} \gtrsim \kappa, \Gamma_M \), where \( n \sim |\alpha|^2 \) is the cavity photon number. Even if the unnormalized coupling is weak, strong-coupling regimes may be achieved by increasing the driving power and thus increasing intracavity photon numbers.

In the present levitated system, a particularly interesting regime would involve triple-mode hybridization enabling, for example, the coupling of the two modes of light via the mechanical mode. However, here, mode hybridization (for which \( \tilde{g}_{1,2} = g_{1,2} \alpha_{1,2} \gtrsim \kappa, \Gamma_M \)) depends non-trivially on the detunings.

We argue that large coupling cannot be easily achieved by increasing the driving power, since \( \tilde{g}_{1,2} \propto n^{1/4} \) and thus increases slowly with the driving strength. We can show, by a simple argument, that increasing the nanosphere size provides the most effective means to attain strong coupling.

For the self-trapped system, optomechanical coupling strengths are \( g_j = \sqrt{2k_A X_{zpf}} \sin(2k_{x0} - \phi_j) \) and depend on the detunings via \( x_0 \). Note also that \( X_{zpf} = \sqrt{\hbar / (2m \omega_M)} \) here too depends on the detunings via \( \omega_M \).

For triple mode hybridization, \( \omega_M \sim \Delta_1 \sim \Delta_2 \). For convenience, we also take \( \phi_1 = 0, \phi_2 = \pi / 4 \). Then, since \( 2k_{x0} = |\alpha_2|^2 / |\alpha_1|^2 \), we can re-write equation (5):

\[
\omega_M^2 = \frac{2\hbar A k^2}{m \cos 2k_{x0}} |\alpha|^2.
\]  

We consider near symmetric driving of the two optical modes for which \( R \sim 1 \) and thus \( k_{x0} \approx \phi / 2 = \pi / 8 \) so \( \omega_M \sim \left( \frac{2\hbar A k^2}{m} \right)^{1/2} n^{1/2} \). Hence,

\[
\tilde{g}_{1,2} \sim \left( \frac{\hbar k^2}{4} \right)^{1/4} \left( \frac{A^3 n_{1,2}}{m} \right)^{1/4}.
\]  

Since the optomechanical coupling increases only very slowly with cavity photon number the most effective means to reach strong coupling regimes is to increase the nanosphere size to the maximum practical values, of order \( r \sim 200 \text{ nm} \), given the need to avoid losses and heating due to photon recoil.

For the ideal case where the nanosphere radius \( r \) is small [23], i.e. \( \lambda \gg r \), then \( A(r \ll \lambda) \equiv A_0(r) \) where the small nanosphere coupling takes the form

\[
A_0(r) = \frac{3 \epsilon_r - 1}{2 \epsilon_r - 2} V_c \omega_L,
\]  

where \( V_c = 4/3 \pi r^3 \) is the sphere volume (and hence \( m = V_c \rho \) where the density \( \rho = 2000 \text{ kg m}^{-3} \) for silica). In turn, \( V_c = \pi (w/2)^2 L \) is the cavity volume, where \( w \approx 40 \mu \text{m} \) is the cavity waist and \( L \simeq 0.5–1 \text{ cm} \) is the cavity length.

For larger nanospheres, the measured size-dependent corrections must be applied. In the experiments described below, we find that the mechanical oscillation frequency is modulated by a finite size correction \( \omega_M(r) = \omega_M(r \approx 0) f(r) \) (see figure 12 and description of the measurement of \( f(r) \) in section 6 below). Thus, since

\[
\omega_M^2 \propto \frac{A(r)}{m},
\]  

then \( A(r) = A_0(r) f^2(r) \) and the coupling is in turn modulated by the finite size correction. The experimental results suggest that for \( r \lesssim 200 \text{ nm} \), then \( f(r) \sim 1 \) and \( A(r) \sim A_0(r) \).
Figure 1. Size dependent effects in the magnitude of the optomechanical coupling parameter $\tilde{g}_1$. It is assumed that cavity parameters would correspond to $\tilde{g}_1 = 10^6$ Hz at $r = 150$ nm, for photon numbers $n_1 = 10^9$. For comparison, the value of $A$ is also shown, as are the experimental and simulated frequencies $\omega_a \equiv \omega_M = 2\pi f_M$. The $\omega_M$ are scaled by a factor of 10 for clarity and, in fact, values of $\omega_M \sim 1$ MHz are quite realistic in optical cavities.

For example, for $r = 150$ nm, $L = 1$ cm and $w = 40 \mu m$, then $A_0 \simeq 8 \times 10^5$ Hz. For reasonable values of cavity decay constants $\kappa \simeq 2 - 8 \times 10^5$ Hz, then for $n_1 \sim 10^9$,

$$\tilde{g}_{1,2} \sim 5.4 \times 10^{-6} \left(\frac{A^3 n_{1,2}}{m}\right)^{1/4} \simeq 10^6 \text{ Hz} \gg \kappa. \tag{11}$$

For $r \lesssim 200$ nm, $\tilde{g}_{1,2} \propto r^{3/2}$. Thus a 200 nm sphere provides an optomechanical coupling about an order of magnitude larger than a 40–50 nm sphere. To achieve a comparable increase in coupling by photon number enhancement would require increasing the driving power by a factor of order $10^4$.

A more careful analysis, including the effects of the finite-size correction function $f^2(r)$, is shown in figure 1. We see that $\tilde{g}_1$ (and for $n_1 \sim n_2$, also $\tilde{g}_2$) reaches a maximum value for $r \simeq 300$ nm before falling to zero. Other maxima for larger $r$ do not provide a larger value of $\tilde{g}_{1,2}$. Furthermore, they have the disadvantage that they may enhance photon recoil heating effects. For comparison, the value of $A$ is also shown, overlaid on the experimental and simulated frequencies.

4. Dynamics

4.1. Optomechanical damping

A previous study [28], using rescaled coordinates, investigated the full parameter space of two optical mode cooling. Here we investigate more carefully the effect of non-zero mechanical damping. Using linear response theory, we can extract cooling rates from equations (4):

$$\Gamma_{\text{opt}} = [S_1(\omega_M) + S_2(\omega_M) - S_1(-\omega_M) - S_2(-\omega_M)], \tag{12}$$
Maps of cooling rates calculated from equation (12) for parameters $R = 1.0$ and $0.5$. Blue corresponds to cooling, yellow/white to heating. The white lines indicate the locus of the single field resonances (where $-\Delta_1 = \omega_M$ or where $-\Delta_2 = \omega_M$). The detunings are given in units of $A$ and are dimensionless.

For $R = 1$ it is clear that there is a deep, maximum cooling region at a double resonance where the two white lines intersect and both optical fields cool simultaneously. It is also evident that there is a strong cooling resonance for $+\Delta_{1,2} = \omega_M$. For $R = 0.5$, three cooling resonances $-\Delta_{1,2} = \omega_M$, $-\Delta_1 = \omega_M$ and $+\Delta_1 = \omega_M$ merge to give a very broad strong-cooling region, quite insensitive to detuning $\Delta_2$ over a range of over 1 MHz. Here $A = \kappa/2 = 0.3$ MHz and the input power into mode 1 corresponds to 2 mW.

Figure 2. Maps of cooling rates calculated from equation (12) for parameters $R = 1.0$ and $0.5$. Blue corresponds to cooling, yellow/white to heating. The white lines indicate the locus of the single field resonances (where $-\Delta_1 = \omega_M$ or where $-\Delta_2 = \omega_M$). The detunings are given in units of $A$ and are dimensionless.

For $R = 1$ it is clear that there is a deep, maximum cooling region at a double resonance where the two white lines intersect and both optical fields cool simultaneously. It is also evident that there is a strong cooling resonance for $+\Delta_{1,2} = \omega_M$. For $R = 0.5$, three cooling resonances $-\Delta_{1,2} = \omega_M$, $-\Delta_1 = \omega_M$ and $+\Delta_1 = \omega_M$ merge to give a very broad strong-cooling region, quite insensitive to detuning $\Delta_2$ over a range of over 1 MHz. Here $A = \kappa/2 = 0.3$ MHz and the input power into mode 1 corresponds to 2 mW.

Figure 2. Maps of cooling rates calculated from equation (12) for parameters $R = 1.0$ and $0.5$. Blue corresponds to cooling, yellow/white to heating. The white lines indicate the locus of the single field resonances (where $-\Delta_1 = \omega_M$ or where $-\Delta_2 = \omega_M$). The detunings are given in units of $A$ and are dimensionless.

For $R = 1$ it is clear that there is a deep, maximum cooling region at a double resonance where the two white lines intersect and both optical fields cool simultaneously. It is also evident that there is a strong cooling resonance for $+\Delta_{1,2} = \omega_M$. For $R = 0.5$, three cooling resonances $-\Delta_{1,2} = \omega_M$, $-\Delta_1 = \omega_M$ and $+\Delta_1 = \omega_M$ merge to give a very broad strong-cooling region, quite insensitive to detuning $\Delta_2$ over a range of over 1 MHz. Here $A = \kappa/2 = 0.3$ MHz and the input power into mode 1 corresponds to 2 mW.

For $R = 1$ it is clear that there is a deep, maximum cooling region at a double resonance where the two white lines intersect and both optical fields cool simultaneously. It is also evident that there is a strong cooling resonance for $+\Delta_{1,2} = \omega_M$. For $R = 0.5$, three cooling resonances $-\Delta_{1,2} = \omega_M$, $-\Delta_1 = \omega_M$ and $+\Delta_1 = \omega_M$ merge to give a very broad strong-cooling region, quite insensitive to detuning $\Delta_2$ over a range of over 1 MHz. Here $A = \kappa/2 = 0.3$ MHz and the input power into mode 1 corresponds to 2 mW.

For $R = 1$ it is clear that there is a deep, maximum cooling region at a double resonance where the two white lines intersect and both optical fields cool simultaneously. It is also evident that there is a strong cooling resonance for $+\Delta_{1,2} = \omega_M$. For $R = 0.5$, three cooling resonances $-\Delta_{1,2} = \omega_M$, $-\Delta_1 = \omega_M$ and $+\Delta_1 = \omega_M$ merge to give a very broad strong-cooling region, quite insensitive to detuning $\Delta_2$ over a range of over 1 MHz. Here $A = \kappa/2 = 0.3$ MHz and the input power into mode 1 corresponds to 2 mW.

where

$$S_j(\omega) = \frac{|\alpha_j|^2 g_j^2 \kappa}{[\Delta_j - \omega]^2 + \frac{\kappa^2}{4}},$$

for $j = 1, 2$. Net cooling occurs for $\Gamma_{\text{opt}} < 0$. Although the above is quite similar in form to standard optomechanical expressions, as explained previously, rather different behaviour is observed since here $\omega_M$ and $g_j$ are both dependent on the $\Delta_j$.

From quantum perturbation theory we can show that $R_{n \rightarrow m}$, the rate of transition from state $n$ to $n+1$, is $R_{n \rightarrow n+1} = (n+1) (S_1(\omega_M) + S_2(\omega_M))$ while $R_{n \rightarrow n-1} = n (S_1(-\omega_M) + S_2(-\omega_M))$.

For $n \gg 1$, then $R_{n \rightarrow n+1} - R_{n \rightarrow n-1}$ gives the cooling rate of equation (12). However, with the exact expressions we can show that the equilibrium mean phonon number is

$$\langle n \rangle_{\text{PT}} = \frac{S_1(\omega_M) + S_2(\omega_M)}{S_1(-\omega_M) + S_2(-\omega_M) - S_1(\omega_M) - S_2(\omega_M)}.$$ 

In figure 2 we show colour maps comparing the cooling and minimum phonon numbers for both $R = 0.5$ and $1$ respectively. The cooling behaviour was investigated previously in [28]. In this case, for each fixed detuning $\Delta_1$ there are up to three cooling resonances (at three different
values of $\Delta_2$), where strong damping is observed (and similarly for each fixed $\Delta_2$). This is in contrast to single optical mode schemes where there is a single cooling resonance for which $\Delta_1 = -\omega_M$ or $\Delta_2 = -\omega_M$. For the $R = 0.5$ map the three cooling resonances merge, giving a single extended cooling region of about 1 MHz width.

For $R = 1$ the map has a high degree of symmetry, since the role of the two optical modes is interchangeable. The figures show that the largest cooling rates are found in the double resonance region, making it the most favourable region to work in.

The equilibrium phonon number in equation (14) concerns only the idealized situation where there is a very good vacuum, negligible photon recoil heating and thus no mechanical damping or heating effects.

Recoil heating depends strongly on the sphere size [23, 27]. A recent study [27] obtained expressions for recoil heating rates $\Gamma_{\text{rec}} = X_{\text{zpf}} cn \pi V_c^{-1} F_{\text{rec}}(r)$, where $F_{\text{rec}}(r)$ is a size-dependent envelope, given in terms of Bessel functions (noting also that $X_{\text{zpf}} \propto r^{-3}$). A similar envelope was obtained in [27] for photon losses $\kappa_{\text{rec}} = \frac{c r^2}{2 \lambda V_c} F_{\text{rec}}(r)$. For the cavities we investigate below, with waist 40 $\mu$m, length $L = 1$ cm, for spheres in the range 100–200 nm and mechanical frequency $\omega_M \simeq 1$ MHz, we obtain $\Gamma_{\text{rec}} \sim 10^{-4} n F_{\text{rec}}$. Typically $n = 10^8$–$10^9$.

For optimal cooling rates $\Gamma_{\text{opt}} \sim \kappa \sim 3 \times 10^5$ Hz, the recoil heating is weak if $\Gamma_{\text{rec}} \ll \kappa$ and hence if $F_{\text{rec}} \ll 1$. Figure 3 plots the form of $F_{\text{rec}}$ and shows that $\Gamma_{\text{rec}} \ll 1$ corresponds to $r \lesssim 200$ nm. For example, for our key results (shown in figure 8), $n \simeq 10^8$ and $r < 150$ nm thus $\Gamma_{\text{rec}} \sim 5 \times 10^3 \ll \kappa = 2 \times 10^5$ Hz.

For the case of the correction to the cavity decay, $\kappa_{\text{rec}} \simeq 10^9 F_{\text{rec}}(r)$, which is easily satisfied for $r = 150$ nm.

For small $r \lesssim 200$ nm spheres, we assume recoil heating is negligible [23], for reasonable choices of experimental parameters. Below, we investigate dynamics for $r = 100$–150 nm and the dominant source of mechanical damping is background gas collisions, which provide an

*Figure 3.* The radial dependence of the recoil effects. The regime $F_{\text{rec}} \lesssim 1$ corresponds approximately to $r \ll 200$ nm. The inset amplifies this region, highlighting the $100 \ll r \ll 200$ nm region for which we investigate the quantum dynamics. For double resonance, $k x_0 \simeq \pi/4$ so is intermediate between the two curves.
effective mechanical damping $\Gamma_M = \frac{8}{\pi} \pi m_r r^2 n_g \bar{v}_g$ [22, 23] where $m_g/m_s$ is the ratio of the gas particle’s mass to that of the sphere, $n_g$ is the gas number density and $\bar{v}_g$ is the mean gas velocity for a room temperature thermal distribution.

It can be shown that the perturbation theory argument above can be adapted to obtain equilibrium phonon numbers for a given cooling rate $\Gamma_{\text{opt}}$:

$$\langle n \rangle_{\text{PT}} = \frac{k_B T_B}{\hbar \omega_M} \Gamma_M + \left[ S_1(\omega_M) + S_2(\omega_M) \right] \frac{\Gamma_M}{\Gamma_{\text{opt}}},$$

(15)

where $T_B \simeq 300$ K. Alternatively, the final equilibrium temperatures $T_{eq} = \frac{\Gamma_M T_B + |\Gamma_{\text{opt}}| T_{\text{vac}}}{\Gamma_{\text{opt}}}$, where $T_{\text{vac}}$ is the equilibrium oscillator temperature which would have been obtained in a perfect vacuum.

4.2. Quantum and semiclassical noise spectra

Although we investigate only a two optical mode system, generalization to more optical modes is straightforward. We consider a set of equations of motion, for $j = 1, \ldots, N$:

$$\dot{\hat{b}} = - \left( i \omega_M (\Delta_1^*, \ldots, \Delta_j^*) + \frac{\Gamma_M}{2} \right) \hat{b} + i \sum_j g_j (\Delta_1^*, \ldots, \Delta_j^*) (\hat{a}_j + \hat{a}_j^*) + \sqrt{\Gamma_M} \hat{b}_\text{in},$$

$$\dot{\hat{a}}_j = \left( i \Delta_j^* - \frac{\kappa}{2} \right) \hat{a}_j + i g_j (\Delta_1^*, \ldots, \Delta_j^*) \hat{a}_j (\hat{b} + \hat{b}^*) + \sqrt{\kappa} \hat{a}_{\text{in}}^{(j)},$$

(16)

where the optomechanical strengths $g_j (\Delta_1^*, \ldots, \Delta_j^*) = -k A X_{2P} \sin(2kx_0 - \phi_j)$ depend on the detunings (as does the mechanical frequency $\omega_M$). In the two mode case we consider, we take $\phi_1 = 0$ and $\phi_2 = \pi/4$.

The optical modes are subject to photon shot noise, while the mechanical modes are subject to Brownian noise from collisions with gas molecules in the cavity. For the photon shot noise, we assume independent lasers and uncorrelated zero temperature noise for which $\langle \hat{a}_{\text{in}}^{(i)} (t') \hat{a}_{\text{in}}^{(j)} (t) \rangle = 0$.

The above equations can be integrated in frequency space to obtain analytical expressions for the displacement noise spectra for the arbitrary mode case. We can evaluate the displacement spectrum $S_{xx} (\omega) \equiv \langle |x(\omega)|^2 \rangle_{\text{QM}} = \frac{1}{2\pi} \int e^{-i \omega t} \langle x(t) x(t) \rangle dt$. We obtain

$$\langle |x(\omega)|^2 \rangle_{\text{QM}} |M(\omega)|^2 = \Gamma_M \left[ |\chi_M(\omega)|^2 n_B + |\chi_M(-\omega)|^2 (n_B + 1) \right] + \frac{\kappa}{2} |\mu(\omega)|^2 \sum_{j=1,2} g_j^2 |\chi_{jo}(-\omega)|^2,$$

(17)

where the $\chi (\omega)$ represent optical and mechanical susceptibilities:

$$\chi_{jo} (\omega) = \left[ -i(\omega + \Delta_j^*) + \frac{\kappa}{2} \right]^{-1}, \quad \chi_M (\omega) = \left[ -i(\omega - \omega_M) + \frac{\Gamma_M}{2} \right]^{-1},$$

(18)

with $\mu(\omega) = \chi_M (\omega) - \chi_M^*(\omega)$ and $\eta_j (\omega) = \chi_{jo} (\omega) - \chi_{jo}^*(\omega)$; then also $M(\omega) = 1 + \mu(\omega) \sum_j g_j^2 |\eta_j (\omega)|^2$.

We compare the quantum displacement with corresponding semiclassical solutions in the steady state. The linearized two mode system equation (6), in matrix form corresponds to a standard problem [39]. Inclusion of the noise arising from gas collisions or laser shot noise.
Figure 4. Comparison between quantum and semiclassical displacement spectra for gas pressures of 1 mbar and at near vacuum pressure in a strong cooling region. At high vacuum, the ground state is approached and thus, for the quantum spectrum, the blue sideband vanishes. At higher pressure there is good agreement between the quantum and classical results. Spectra near double resonance for input power $P_1 = 7$ mW, $A = \kappa = 3 \times 10^5$ Hz, $\Delta_1 = -1.5$ MHz, $\Delta_2 = -0.68$ MHz, $R = 0.5$. Some hybridization between the mechanical mode and optical mode 1 is seen in the characteristic double-peak sideband structure.

Yields a set of corresponding Langevin equations: $\frac{dX(t)}{dt} = AX + BE(t)$, where $A$ is termed the drift matrix. Its eigenvalues give the stabilities and eigenfrequencies of the system’s normal modes, while the noise is determined by $B$, a constant diagonal matrix. The elements of the random noise matrix are assumed to be $\delta$-correlated $\langle E_i(t)E_j(t') \rangle = \delta(t-t')\delta_{ij}$. Methods for obtaining the solution for the steady state correlation functions of this system, under conditions of stability, i.e. where all the eigenvalues of $A$ have negative real parts, are well known [39]. The required noise spectra, or autocorrelation functions, in frequency space are

$$S(\omega) = \frac{1}{2\pi} \int_{-\infty}^{\infty} e^{-\text{i}\omega \tau} \langle X(t+\tau)X^T(t) \rangle d\tau,$$

(19)

where

$$S(\omega) = \frac{1}{2\pi} (A + i\omega I)^{-1} BB^T (A^T - i\omega I)^{-1},$$

(20)

where the diagonal matrix $BB^T$ has elements $(n_B + \frac{1}{2})\Gamma_M, (n_B + \frac{1}{2})\Gamma_M, \kappa, \kappa, \kappa, \kappa)$. From the above, the noise spectra of all modes may be calculated. Equation (20) yields semiclassical sideband spectra, symmetrical in $\omega$.

In Figure 4, we compare semiclassical displacement spectra calculated from equation (20) with corresponding quantum results obtained from equation (17). At high pressures (and hence high phonon occupancy) there is excellent agreement between semiclassical and quantum...
For the broad cooling region formed from three overlapping resonances seen in figure 2(b), we show equilibrium phonon numbers obtained from equations (15), (17) and (20), i.e. perturbation theory, the analytical quantum noise formula and semiclassical Langevin equations respectively. Agreement between quantum and semiclassical results is excellent, less so for perturbation theory at low pressures.

At low pressures (near ground state cooling) however, the quantum spectrum shows a characteristic asymmetry, such as was observed recently in experiments on photonic cavities [16].

4.3. Comparison between perturbation theory, semiclassical and quantum results

The equilibrium variance (and hence the final phonon number) of the mechanical oscillator is

$$\langle x^2 \rangle = \frac{1}{2\pi} \int_{-\infty}^{\infty} \langle |x(\omega)|^2 \rangle d\omega,$$

thus the final equilibrium temperature of the mechanical oscillator after optomechanical cooling is $k_B T_{eq} = 1/2 m\omega_M^2 \langle X^2 \rangle$. Noting the rescaling $\langle X^2 \rangle = 2X_{ZPF} \langle x^2 \rangle$ and setting $k_B T_{eq} = (\langle n \rangle + 1/2)\hbar\omega_M$, we can write $\langle x^2 \rangle = \langle n \rangle + 1/2$.

Using equations (16), (20) and (15), we can investigate final equilibrium phonon numbers (and the minimum achievable for levitated self-trapped spheres) comparing quantum, semiclassical and perturbation theory respectively. In figure 5, we compare the corresponding equilibrium phonon numbers, $\langle n \rangle_{QM}$, $\langle n \rangle_{SC}$ and $\langle n \rangle_{PT}$ respectively for the unusual triple cooling resonance region shown in figure 2. Cooling to near the ground state $\langle n \rangle \sim 0$ is possible for a pressure of order $10^{-6}$ mbar, even for modest driving powers of 2 mW and values of $A \simeq 3 \times 10^5$ Hz corresponding to spheres of order $r \simeq 100$ nm.
Figure 6. Triple mode splitting. For $A = \kappa = 0.3$ MHz, even at quite high pressures (here 1 mbar), mode splitting is seen in the noise spectra of the optical modes. In all the plots, $\Delta_1 = -1.15$ MHz is held fixed while $\Delta_2$ is swept from 0 to $-1.6$ MHz (for an input power of 2 mW into mode 1, while $R = 0.5$). (a) Noise spectra for both optical mode 1 and mode 2. Three way hybridization between the mechanical and both optical modes appears clearly (highlighted by the bold blue line). For clarity, some of the strongest peaks have been truncated in height. In (b) three avoided crossings are apparent. The dominant character of each normal mode is indicated by the colour (black is mechanical, blue is optical mode 1, red is optical mode 2). When $-\Delta_2$ is large, there is no mixing. However, as $\Delta_2 \to 0$, there is strong mixing and the dominant character of each normal mode changes from light to matter (or vice versa) as an avoided crossing is encountered. Panel (c) shows the cooling and indicates strong cooling at each of the avoided crossings.

5. Strong coupling regimes: triple mode splitting and bistability

The multimode, or at least two mode, self-trapping regime may permit new possibilities for position sensing and for controlling entanglement between two optical modes and the mechanical resonator. Here we investigate regimes where such effects are clearly apparent. The implications of the measurement for the accessible range of optomechanical coupling strengths suggests that multiple hybridization and bistability are quite accessible with reasonable cavity parameters.

In figure 6 we investigate the complex behaviour of the eigenmode frequencies of the self-trapped, levitated system. On the left panels (figure 6(a)) we plot the noise spectra of the two optical modes, which exhibit sidebands near $\omega \simeq \omega_M$ since the corresponding optical
Figure 7. Maps of the displacement noise spectra $S_{xx}(\omega)$, showing mode splitting for similar parameters to figure 4, near the quantum limit, except that here $\Delta_1 = -1.5$ MHz is held fixed while $\Delta_2$ (vertical axes) is swept. (a) The semiclassical spectrum which is symmetric in frequency $\omega$. (b) The quantum spectrum which is asymmetric. In both cases, triple hybridization appears clearly and is seen near $\Delta_2 \approx -1.0$ MHz. Note that $\log S_{xx}(\omega)$ is plotted.

fields are modulated by the motion of the mechanical oscillator. Here we fix one detuning ($\Delta_1 = -1.15$ MHz) and look at the behaviour as the other detuning is varied. The sidebands are displaced in frequency and split: one effect is simply due to the dependence of $\omega_M$ on $\Delta_j$ (unique to the levitated system); it arises from the calculation of the equilibrium fields and frequencies. The other effect is due to normal mode mixing (hybridization of light and matter modes) arising from the linearized equations. If $\omega_M \approx \Delta_1 \approx \Delta_2$ simultaneous hybridization is observed, provided $g_{1,2} \gtrsim \kappa$. Figure 6(b) shows that there are several distinct avoided anti-crossings, where the dominant character of each eigenmode changes; if two crossings coincide, the spectra show a characteristic triple-peak structure (symmetric about $\omega = 0$ in the semiclassical regime shown here). Panel (c) shows that the corresponding cooling rate is enhanced at each avoided level crossing.

In figure 7 we plot displacement spectra corresponding to figure 4, but over a range of values of $\Delta_2$. A log-scale is used for $S_{xx}(\omega)$. The triple mixing which can appear when two avoided crossings nearly coincide is clearly apparent at $\Delta_2 \approx 1$ MHz.

Static bistability in a cavity of varying length has been seen experimentally [40]. The potential for generating entanglement has recently been investigated in an optomechanical system [41]; however, a relatively high laser power $P \sim 50$ mW is required.

For the self-trapped systems, the incoherent sum of the optical standing-wave potentials $\cos^2(kx - \phi_1)$ and $\cos^2(kx - \phi_2)$ does not by itself produce a double-well structure; nevertheless, as we see below, in combination with optomechanical shifts, bistabilities are observed, even for weak driving. Whether a double-well structure emerges, or not, is completely independent of the driving power (where $P \propto E_f^2$) and can emerge at very low input powers, as we demonstrate below. It is easy to see that the levitated particle moves in an effective static
Figure 8. Mode mixing and bistability for $A = 3\kappa = 6 \times 10^5$ Hz. We consider a relatively low input power of 0.37 mW. (a) Plot of the optomechanical cooling rate (blue indicates cooling, brown indicates heating). The red dashed line indicates the locus of bistability as a function of the detunings $\Delta_1$ and $\Delta_2$. The discontinuity in the cooling can be discerned near the strong doubly resonant cooling region. (b) Displacement noise spectra $S_{xx}(\omega)$ as a function of $\omega$ plotted along $\Delta_1 = -1$ MHz (along black horizontal line in (a)). We sweep in increasing $\Delta_2$. At the point where $\Delta_1 = -1$ MHz intersects the bistability, a discontinuity in the noise spectra is apparent. On one side of the discontinuity there is strong hybridization between the mechanical mode and optical mode 2; this changes abruptly across the discontinuity to hybridization between the mechanical mode and optical mode 1 for larger $\Delta_2$. This may allow for control of entanglement between the modes. The map corresponds to near-vacuum conditions so the system is near the quantum ground state in this regime (as evidenced by the asymmetric sidebands).

The potential $V(x)$ where

$$\frac{dV(x)}{dx} = \hbar k A E_1^2 \left[ \frac{\sin 2(kx - \phi_1)}{|(\kappa/2) - i \Delta_1(x)|^2} + R^2 \frac{\sin 2(kx - \phi_2)}{|(\kappa/2) - i \Delta_2(x)|^2} \right]$$

(22)

and

$$V(x) = \hbar A E_1^2 \left[ \tan^{-1} \left( \frac{\Delta_1(x)}{\kappa/2} \right) + R^2 \tan^{-1} \left( \frac{\Delta_2(x)}{\kappa/2} \right) \right].$$

(23)

Here, note that the shifted detuning $\Delta_j(x) = \Delta_j + A \cos^2(kx - \phi_j)$ is dependent on $x$, not the equilibrium displacement $x_0$. This potential admits two stable equilibrium points over parameter regimes where $A \gg \kappa$ (in practice, such bistability is observed already for $A/\kappa \simeq 3$). It is evident that the driving power factors out, so does not affect the shape of the potential, providing only a scaling factor. We show in figure 8 that for a high $A/\kappa$ ratio simultaneous hybridization and bistability coexist. Here we take input power $P_1 = 0.37$ mW, $A = 3\kappa$, $R = 0.15$. Since $P_1 = \frac{2k\epsilon_i^2 \hbar}{(\kappa/2)}$, this implies $E_1 = 10^{10}$. The photon number $n_1 \simeq \frac{E_1}{\omega_{34}} \simeq 10^8$. Thus for quite modest photon numbers $n_1 \sim 10^8 \sim n_2$, we can switch discontinuously from hybridization
between the mechanical mode and optical mode 1, to hybridization between the mechanical mode and mode 2. We take $\phi_1 = 0, \phi_2 = \pi/4$. In the noise spectra, the switch is heralded by a large zero-frequency peak in the displacement spectra, which is clearly apparent in figure 8.

6. Experiment

6.1. Current experimental status: loading protocols and variation of trap frequency with radius

We have built a simple standing wave dipole trap to develop protocols for loading a single nanosphere into the trap to confirm that nanospheres with a range of radii around 100 nm can be trapped. Importantly, we have measured the variation in trap frequency with sphere radius so that realistic values of optomechanical coupling strengths, for a given nanosphere radius, can be included in our models. In this section we explain how the size-dependent modulation function $f(r)$, used in the theory section above to obtain $A(r)$ and the optomechanical coupling strengths, was measured.

A schematic of the standing wave trap used in our experiments is shown in figure 9. The standing wave trap consists of two focused beams that counter-propagate and overlap near their foci. The two laser beams, derived from the same laser at a wavelength of 1064 nm, enter the trapping region via optical fibres. The light exiting the fibres is focused using aspheric lenses (Thorlabs C140TME) with a focal length of 1.45 mm and numerical aperture 0.55. The power in each trapping beam after it has passed through each lens is measured to be $150 \pm 10$ mW, and

Figure 9. Schematic diagram of the standing wave trap. It is formed from two counter-propagating 1064 nm beams focused inside a vacuum chamber. Light at 532 nm enters via one fibre to image the sphere. Images and measurement of the axial and transverse position of the trapped nanosphere as a function of time are measured by a CCD camera and quadrant cell photodiode (QPD) respectively, through a long working distance microscope outside the vacuum chamber.
Figure 10. An image of a string of 100 nm diameter beads trapped in the standing wave trap. The single exposure of 60 ms is much longer than the trap oscillation period of 29 \( \mu s \). A single bead is trapped by continually blocking and unblocking one of the trapping beams until only one sphere is trapped.

the best focused beam waist (radius) is theoretically 1.7 \( \mu m \). To optimize the alignment of the trap, we maximize the light through-coupled from one fibre into the other. This is accomplished by mounting one optical fibre and its aspheric lens on an XYZ flexure stage. The alignment is done inside the vacuum chamber at atmospheric pressure.

A long working distance microscope (Navitar Zoom 6000 system, with up to 45\( \times \) zoom) is used to image the trapped sphere. The image is split into two using a beamsplitting plate, with one image directed to a CCD camera for diagnostics and the other aligned onto a quadrant cell photodiode (QPD) which measures position fluctuations as a function of time in two orthogonal axes. We define the axial direction as that along which the trapping light propagates, and the transverse direction as the orthogonal axis in the focal plane of our imaging system. Light at 532 nm is used to illuminate the trapped sphere, as the QPD is more sensitive at this wavelength. The green beam enters the system via one of the optical fibres, as shown in figure 9, and a filter is used to stop 1064 nm light reaching the detectors. The power of the 532 nm beam is 10 mW.

Silica (\( \text{SiO}_2 \)) nanospheres, manufactured by Microspheres–nanospheres and Bangs Laboratories, are introduced into the trapping region at atmospheric pressure via an ultrasonic nebulizer (Omron NE-U22). These spheres range in radius from 26 to 510 nm and are suspended in methanol. The nanosphere solution is sonicated using an ultrasonic bath for at least an hour before trapping to prevent clumping. The spheres are held in the methanol liquid in the nebulizer and are forced through the 3 \( \mu m \) holes of the nebulizer. The droplets released from the nebulizer, which is several centimetres from the trap, fall into the trapping region and are trapped. This process is aided by the dissipation of the surrounding gas typically at atmospheric pressure. Once trapped, the methanol around the spheres rapidly evaporates leaving only the trapped sphere. As many particles are released a number of spheres are trapped over many fringes of the standing wave, as shown in figure 10. As our imaging system does not have single fringe resolution we cannot determine if more than one sphere is trapped in a single fringe by this method. However, this information can be inferred from the relative intensity of the light scattered from the trapped spheres and also by the reduced stability of the particles in the trap when more than one particle is trapped. To reduce the number of trapped particles the trapping light is briefly blocked and unblocked. This is repeated until a single sphere is visible in the trap. At this pressure, where there is a strong damping force from the air, the sphere can be held in the trap indefinitely.

New Journal of Physics 15 (2013) 015001 (http://www.njp.org/)
Figure 11. Power spectra at 5 mbar calculated from a measurement of the position of a trapped 200.1 nm diameter nanosphere as a function of time, using a QPD. (a) The transverse frequency and (b) an axial frequency. Outlier points are due to electronic noise. Red line fitted with the formula described in the text, from which the trap frequencies are extracted.

To measure the trap frequency the air is pumped from the system, and at this point no more spheres enter the trap, as without air-damping their velocity is too high. The air pressure in the trap is reduced to 5 mbar, so that clear trap frequencies can be obtained from the power spectrum of the position fluctuations of the trapped sphere, as recorded on the QPD. Example power spectra are shown in figure 11. Above 5 mbar the damping of the motion in the trap due to air broadens the peak in the power spectrum so that finding an accurate trap frequency is difficult. Below pressures of 5 mbar the spheres become unstable in the trap and escape. This is most likely due to radiometric forces which have been compensated for in other experiments using feedback techniques [35, 36]. At 5 mbar the damping rate due to gas collisions is significantly less than our lowest measured trap frequencies, and thus the measured frequency at this pressure is a good approximation to the bare trap frequency which would be measured in vacuum without damping.

The angular axial trap frequency for a small polarizable particle in a standing wave is \( \omega_a = 2\pi f_a = \sqrt{\frac{4\kappa k_I_0}{m\epsilon_0c} \omega_a^2 - \omega_t^2} \), where the polarizability of a sphere of refractive index \( n \) is \( \alpha = 4\pi\epsilon_0r^3 \left( \frac{n^2-1}{n^2+2} \right) \). The maximum intensity in the radial centre of each equal intensity beam is given by \( I_0 \), and \( k \) is the magnitude of the wavevector of each beam. The sphere has mass \( m = \frac{4}{3}\pi\rho r^3 \), radius \( r \) and density \( \rho \simeq 2000 \text{ kg m}^{-3} \). The transverse trap frequency is given by \( \omega_t = \sqrt{\frac{8\alpha I_0}{m\epsilon_0c}} \), where \( w \) is the focused spot size (radius) of the two counter-propagating beams. From these expressions the ratio of the trap frequencies is given by \( \omega_a/\omega_t = kw/\sqrt{2} \).

The trap frequencies on each axis are determined by fitting measured position fluctuation power spectra using \( \frac{2k_B T}{m} \frac{\Gamma_0}{(\omega - \omega_m)^2 + \omega^2\Gamma_0^2} \), where \( k_B \) is Boltzmann’s constant and \( \Gamma_0 \) is the damping rate. The fit to the data is shown in figure 11 with \( \Gamma_0/2\pi \simeq 2.4 \text{ kHz} \) for the transverse trap frequency. Several sets of data were taken for different spheres of the same nominal radius and the measured axial and transverse trap frequencies for each size sphere are shown in figure 12. The derived trap frequencies for each sphere radius are the average over different experiments at each radius, and the errors are the standard errors in the mean. The uncertainty in the sphere radius is taken from the information supplied by the manufacturer. Two axial (red and green data in figure 12) frequencies and one transverse frequency (blue data points) were measured.
Figure 12. Measured trap frequencies as a function of sphere radius. Points plotted in green are the axial trap frequency, blue are the transverse trap frequency and the red data points are the higher axial trap frequencies which are believed to be due to optical binding. The solid black line is a theoretical curve derived from a numerical calculation [38].

When a particle is tightly trapped by the optical field only one axial frequency is expected from a single sphere in a standing wave. The lower axial frequency (in green in figure 12) is always observed in the data and this is taken as the true axial frequency. The higher frequency, which is often present in the data, may be due to the trapping of two spheres in a single anti-node, with the higher frequency occurring due to optical binding, which requires further study [37]. The higher axial frequency also changes rapidly with sphere size, indicating that it is not the true axial trap frequency, which should be almost constant for the small spheres. The presence of a single axial frequency is, we believe indicative of having trapped a single sphere.

Although we do not know the radial dimensions of the beam within the trap we can estimate this value from the ratio of the axial to transverse trap frequencies for small spheres. The spot size from this ratio is $w = \sqrt{\frac{2}{k}} \frac{\omega_a}{\omega_t}$ and for $\frac{\omega_a}{\omega_t} = 9.8$ this gives a spot size of $w = 2.3 \, \mu m$. Since the trap frequency with two overlapping beams of size $2.3 \, \mu m$ would be equal to $207 \, kHz$ with a power in each beam of $150 \, mW$, and we only measure a maximum axial trap frequency of approximately $40 \, kHz$, we conclude that the particles are trapped in a standing wave formed where the waist of one beam is much larger. If one spot size is $2.3 \, \mu m$ the other would have to be $15 \, \mu m$. A plot of the calculated axial trapping frequency, found by calculating Maxwell’s stress tensor [38], is also shown in figure 12. Like the experimental data the trapping frequency is constant for small spheres and decreases to approximately zero when the particle size is comparable to the size of the interference pattern produced by the standing wave. At larger radii the force on the particle changes sign and a stable trap is formed in a node of the standing wave, as shown for the particle of radius $510 \, nm$. Our measurements confirm that for particle radii less that $200 \, nm$ the simple dipole model for the nanospheres is adequate for modelling the cooling and dynamics of the nanospheres in an optical cavity utilizing $1064 \, nm$ radiation.

New Journal of Physics 15 (2013) 015001 (http://www.njp.org/)

$IOP$ Institute of Physics $\bullet$ DEUTSCHE PHYSikalische GESELLSCHAFT
Our experiments have shown that optical traps without feedback are currently limited to operation at pressures down to a few millibar for all particles that we have measured. In addition, this limiting pressure did not change by reducing the intensity by 50%. This radiometric force is due to localized heating of nanosphere and the subsequent heating of the surrounding air. At low pressures, when the mean free path is comparable to the size of the nanosphere, the radiometric force competes with, and eventually dominates, the dipole force which traps the particle. The radiometric effect does give rise to a supplementary force and additional fluctuations. It is, however, generally not dissipation but heating. The nanospheres are heated through laser absorption and, in turn, the nanospheres heat the colliding gas particles. We are currently studying this effect near the lowest pressures by varying the intensity of the trapping light. However, the fits to the axial frequency data in figure 11 agree well with predictions from gas theory [36] giving a $\Gamma$ of $2\pi \times 15.6 \pm 1.0$ kHz (theory value is $2\pi \times 15.4$ kHz) at 300 K. We note that in much higher vacuum the effect of collisions is small even if the particle temperature is high since there are few collisions and cavity cooling rates can be high, in the MHz range. The radiometric effect prevents us from pumping the system below approximately 5 mbar. To circumvent this problem we have already used an ion trap to bring these particles in the $10^{-5}$ mbar range, limited only by current vacuum equipment. This step allows us to begin optical trapping from within the ion trap in a pressure regime where radiometric forces are no longer important. We note that while feedback techniques have been successful [36] in taking particles to higher vacuum, decreasing the absorption of the nanospheres is another important route to minimizing radiometric effects. This is feasible since all the spheres we have used in this study are not made of optical quality glass but from less pure colloidally grown silica nanospheres.

7. Conclusions

We have described a study of the dynamics and noise spectra of self-trapped levitated optomechanical systems. We have been able to show, by combining experimental measurements and theoretical calculations, that strong light–matter coupling is attainable over a wide range of particle sizes, and that these can be trapped. The interdependence of the mechanical and optical mode frequencies, unique to self-trapped levitated systems, provides a complex and interesting side-band structure, including multi-mode mixing and bistabilities which we aim to explore experimentally. These conclusions are supported by measurements of trap frequency made in an optical standing trap where we have demonstrated a protocol for loading a single nanosphere in a single antinode.

Acknowledgments

We acknowledge support for the UK’s Engineering and Physical Sciences Research Council. FM acknowledges support by an ERC starting grant, the DARPA ORCHID program and a DFG Emmy-Noether grant.

References

[1] Marquardt F and Girvin S 2009 Physics 2 40
[2] Kippenberg T and Vahala K 2008 Science 321 1172

New Journal of Physics 15 (2013) 015001 (http://www.njp.org/)
[3] Favero I and Karrai K 2009 *Nature Photon.* 3 201–5
[4] Aspelmeyer M, Gröblacher S, Hammerer K and Kiesel N 2010 *J. Opt. Soc. Am. B* 27 A189–97
[5] Metzger C H and Karrai K 2004 *Nature* 432 1002
[6] Arcizet O et al 2006 *Nature* 444 71
[7] Gigan S et al 2006 *Nature* 444 67
[8] Corbitt T et al 2007 *Phys. Rev. Lett.* 98 150802
[9] Thompson J D et al 2008 *Nature* 452 72
  Jayich A M et al 2008 *New J. Phys.* 10 095008
[10] Schliesser A et al 2009 *Nature Phys.* 5 509
[11] Armour A D et al 2002 *Phys. Rev. Lett.* 88 148301
[12] Naik A et al 2006 *Nature* 444 193
[13] Teufel J D, Donner T, Li D, Harlow J H, Allman M S, Cicak K, Sirois A J, Whittaker J D, Lehnert K W and Simmonds R W 2011 *Nature* 475 359
[15] Chan J, Mayer Alegre T P, Safavi-Naeini A H, Hill J T, Krause A, Groblacher S, Aspelmeyer M and Painter O 2011 *Nature Phys.* 4 87
[16] Safavi-Naeini A H, Chan J, Hill J T, Mayer Alegre T P, Krause A and Painter O 2012 *Phys. Rev. Lett.* 108 033602
[17] Braginski V B and Khalili F Y 1992 *Quantum Measurement* (Cambridge: Cambridge University Press)
[18] Bose S, Jacobs K and Knight P L 1999 *Phys. Rev. A* 59 3204
[19] Paternostro M et al 2006 *New J. Phys.* 8 107
[20] Marquardt F et al 2007 *Phys. Rev. Lett.* 99 093902
[21] Wilson-Rae I et al 2007 *Phys. Rev. Lett.* 99 093901
[22] Romero-Isart O, Juan M L, Quindat R and Cirac J I 2010 *New J. Phys.* 12 033015
[23] Chang D E, Regal C A, Papp S B, Wilson D J, Painter O, Kimble H J and Zoller P 2010 *Proc. Natl Acad. Sci. USA* 107 1005
[24] Barker P F and Shneider M N 2010 *Phys. Rev. A* 81 023826
[25] Schulze R J, Genes C and Ritsch H 2011 *Phys. Rev. A* 81 063820
[26] Romero-Isart O et al 2011 *Phys. Rev. A* 83 013803
[27] Pflanzer A, Romero-Isart O and Cirac I 2012 *Phys. Rev. A* 86 013802
[28] Pender G A T, Barker P F, Millen J, Marquardt F and Monteiro T S 2012 *Phys. Rev. A* 85 021802
[29] Groblacher S et al 2009 *Nature* 460 724
[30] Eichenfield M, Camacho R, Chan J, Vahala K J and Painter O 2009 *Nature* 459 550–5
[31] Ding L, Baker C, Senellart P, Lemaitre A, Ducci S, Leo G and Favero I 2010 *Phys. Rev. Lett.* 105 263903
[32] Roh Y-G, Tanabe T, Shinya A, Taniyama H, Kuramochi E, Matsuo S, Sato T and Notomi M 2010 *Phys. Rev. B* 81 121101
[33] Gavartin E, Braive R, Sagnes I, Arcizet O, Beveratos A, Kippenberg T J and Robert-Philip I 2011 *Phys. Rev. Lett.* 106 203902
[34] Teufel J D, Li D, Allman M S, Cicak K, Sirois A J, Whittaker J D and Simmonds R W 2011 *Nature* 471 204–8
[35] Ashkin A and Dziedzic J M 1976 *Appl. Phys. Lett.* 28 333
[36] Li T, Kheifets S and Raizen M G 2011 *Nature Phys.* 7 527
[37] Zemanek P, Jonas A, Sramek L and Liska M 1998 *Opt. Commun.* 151 273
[38] Barton J P, Alexander D R and Schaub S A 1989 *J. Appl. Phys.* 66 4594
[39] Walls D F and Milburn G J 2008 *Quantum Optics* (Heidelberg: Springer)
[40] Dorsel A, McCullen J D, Meystre P, Vignes E and Walther H 1983 *Phys. Rev. Lett.* 51 1550
[41] Ghobadi R, Bahrampour A R and Simon C 2011 arXiv:1104.4145

New Journal of Physics 15 (2013) 015001 (http://www.njp.org/)