Yang–Mills gravity in biconformal space

Lara B Anderson and James T Wheeler

1 Mathematical Institute, University of Oxford, 24-29 St Giles’, Oxford OX1 3LB, UK
2 Department of Physics, Utah State University, Logan, UT, 84321, USA

E-mail: anderson@maths.ox.ac.uk and jwheeler@cc.usu.edu

Received 10 August 2006, in final form 20 November 2006
Published 15 December 2006
Online at stacks.iop.org/CQG/24/475

Abstract

We write a gravity theory with Yang–Mills-type action using the biconformal gauging of the conformal group. We show that the resulting biconformal Yang–Mills gravity theories describe 4-dim, scale-invariant general relativity in the case of slowly changing fields. In addition, we systematically extend arbitrary 4-dim Yang–Mills theories to biconformal space, providing a new arena for studying flat-space Yang–Mills theories. By applying the biconformal extension to a 4-dim pure Yang–Mills theory with conformal symmetry, we establish a 1-1, onto mapping between a set of gravitational gauge theories and 4-dim, flat-space gauge theories.

PACS numbers: 11.15.−q, 11.30.Pb, 04.50.+h

1. Introduction

Is it possible to write a Yang–Mills theory of gravity? By such a theory we mean a Yang–Mills-type action,

\[ S_{YM} = \int K_{AB} F^A F^B, \]  

(1.1)

where the gauge group of \( F^A \) contains the Poincaré group. While this is a natural action to write for an internal symmetry, the nature of general relativity as a gauge theory makes \( S_{YM} \) unacceptable. The reason is that viewing GR as a gauge theory requires a caveat: the requirement of ‘soldering,’ i.e., the identification of the solder form with the co-tangent basis. This identification has a disturbing effect on the action \( S_{YM} \) above—the volume element of the underlying spacetime is expressed in terms of some of the gauge fields. The result is that the field equations are not of a purely Yang–Mills type, but also include artificial source terms built from the curvature. As a result, standard treatments replace \( S_{YM} \) above by the
Einstein–Hilbert action,

$$S_{GR} = \int R^{ab} e^c e^d F_{abcd}.$$  \hspace{1cm} (1.2)$$

See, e.g., [1–9].

Here, we present a new approach to this question which results from studying a class of conformal geometries called biconformal spaces. We find gravity theories with pure Yang–Mills-type Lagrangians in eight-dimensional biconformal space, thereby constructing Yang–Mills theories of gravity. We refer to these new theories as biconformal Yang–Mills gravity theories and show that in the case of slowly changing fields the theories reduce to general relativity on a four-dimensional submanifold.

An interesting further insight into these gravity theories is gained once we develop a general embedding of four-dimensional flat-space Yang–Mills theories into biconformal space. Applying this embedding to the symmetry of biconformal Yang–Mills gravity leads to a relationship between pure four-dimensional Yang–Mills solutions and biconformal gravity solutions.

We briefly review here the central concepts of biconformal gravity theories. The biconformal approach to conformal theory possesses a number of unique advantages over standard conformal gaugings [10–21]. In the standard approaches, the local symmetry is relativistic similarity group (Poincaré plus dilatations). In all of these approaches, difficulties arise, including the presence of unphysical size changes, the requirement for an invariant action in \( n \) dimensions to be of order \( n^2 \) in the curvature and/or the requirement for auxiliary fields to write a linear action.

An alternative to the standard gaugings was first formulated by Ivanov and Niederle [22, 23], who created an eight-dimensional manifold by gauging the conformal group of a four-dimensional spacetime. In their approach, the local symmetry group was taken to be homothetic. Rather than constraining this space to construct a four-dimensional spacetime, they restricted four of the dimensions as far as possible given the required gauge freedom. Later these results were extended [24], by generalizing to arbitrary dimensions, \( n \), and defining the class of biconformal spaces as a result of the \( 2n \)-dim gauging without imposing constraints. In that work, it was shown that the resulting space possessed symplectic structure and admitted torsion-free spaces consistent with general relativity and electromagnetism. Further work [25] provided the most general class of actions linear in the biconformal curvatures. These models eliminated the problems listed above, and it was demonstrated that the resulting field equations lead to the Einstein field equations. The \( 2n \)-dimensional space constructed by the biconformal gauging is interpreted as a relativistic phase space of an \( n \)-dimensional configuration space. This interpretation is justified by the existence of an integrable symplectic form which guarantees that the space is a symplectic manifold. The first supersymmetric biconformal gravity theory was constructed in [26].

Returning to the results of this paper, we first investigate the issues associated with constructing Yang–Mills gravity theories and the successful model described above. In this discussion we will use unavoidably similar terminology to discuss several distinct gauge theories. For clarity, we stress here that below we will define four distinct gauge theories: (1) pure Yang–Mills theory, (2) Yang–Mills theory coupled to gravity, (3) Yang–Mills gravity theory and (4) biconformal Yang–Mills gravity theory.

Our second result is a systematic program for embedding four-dimensional, flat-space Yang–Mills theories into biconformal space. We provide a consistent definition of eight-dimensional fields in terms of their four-dimensional counterparts. This extension provides a new arena for investigating the structure and properties of standard Yang–Mills theories.
Finally, we find that while most flat-space Yang–Mills theories in four dimensions are mapped to theories in flat biconformal space, a certain class of flat Yang–Mills theory can be extended to a gravity theory (with Yang–Mills Lagrangian) in eight dimensions. We establish an exact correspondence between a non-compact $SU(2,2|N)$ super Yang–Mills theory on flat four-dimensional spacetime and a conformal supergravity theory formulated on a biconformal space constructed from the same supergroup. The result provides a previously unknown relationship between a Yang–Mills gauge theory and a gravity theory and a new class of solutions for biconformal Yang–Mills gravity theories.

In the following section, we discuss our notation and the definitions of the relevant field theories. In particular, we review the definitions of Yang–Mills theories and Yang–Mills theories coupled to gravity. Then we define a new class of models, called Yang–Mills gravity theories which are distinct from both of the theories listed above. In section 3, we explicitly present a new class of supergravity models, biconformal Yang–Mills gravity theories. Developing these models in section 4, we demonstrate that biconformal Yang–Mills gravity theories reproduce general relativity on a 4-dim submanifold of an 8-dim space. In section 5, we develop a formalism for extending four-dimensional, flat-space Yang–Mills theories to biconformal space, beginning with a simple example of a $U(1)$ gauge theory. Section 6 extends the construction to the full non-Abelian case. Further, we show in section 7 that the biconformal field equations can be identified with a non-compact supersymmetric Yang–Mills theory on a flat four-dimensional spacetime. Thus, a correspondence is established between a conformal supergravity theory and a non-compact Yang–Mills gauge theory. Finally, we review the results of this work and state conjectures for the structure of biconformal Yang–Mills supergravity theories.

2. Definitions and notation

We begin with some definitions. Although common usage often restricts Yang–Mills theories to unitary symmetry, we define a Yang–Mills theory \[27–29\] to follow from a functional of the form

$$S_{YM} = \int K_{AB} F^A F^B$$

(neglecting topological terms) where $F^A$ is the curvature 2-form of a principal fibre bundle with arbitrary Lie fibre group, $G$, and $K_{AB}$ is the Killing metric (here $A, B$ are Lie algebra indices). The curvatures are related to a connection (or potential), $A^A$, on the bundle by the Cartan structure equations,

$$F^A = dA^A - \frac{1}{2} c_{BC}^A A^B A^C,$$

where the $c_{BC}^A$ are the structure constants of the group $G$. The integral is over a given fixed background manifold. The symmetry is said to be internal if the Yang–Mills potential is distinct from the solder form connection on the background manifold. Standard examples include the $U(1)$ gauge theory of electromagnetism, the $SU(2) \times U(1)$ electroweak theory or the standard model \[1\]. The theory is a flat-space Yang–Mills theory if there is no curvature of the background space. Then, the action takes the form (in n-dim),

$$S_{YM} = \int K_{\alpha\beta} F^a_{\alpha\beta} F^b_{\mu\nu} \eta^{\alpha\mu} \eta^{\beta\nu} d^n x$$

where $\alpha, \beta = 1, \ldots, n$.

If the background spacetime is curved, then the theory is a Yang–Mills theory coupled to gravity. To accomplish this, we introduce a general metric, $g_{\alpha\beta}$, and a volume form $\sqrt{|g|} d^n x$. 
where \( g = \det(g_{\alpha \beta}) \). In this case, the gravity action is added to \( S_{YM} \),

\[
S_{YM+G} = \int K_{ab} F^a \wedge F^b + R^{ab} e^d e_{abc} d^4x.
\]

Variation of the metric then leads to \( n \)-dim general relativity with the energy–momentum tensor associated with \( F_{\alpha \beta}^a \) as gravitational source. The Yang–Mills field \( F_{\alpha \beta}^a \) evolves in the curved background described by \( g_{\alpha \beta} \). Note that the Hodge dual automatically brings in couplings to the curved metric. In both flat-space Yang–Mills and Yang–Mills coupled to gravity, the symmetry is internal.

Before we define Yang–Mills gravity, we require a digression to characterize gravitational gauge theories. Despite our very general definition of Yang–Mills gauge theory, there are other types of gauge theory. One simple variation is to choose any other group invariant action to replace the Yang–Mills action. But there is a deeper difference that arises when we consider a gravitational gauge theory. Suppose we select the Poincaré group as our symmetry group. In this case, one of the potentials of the Yang–Mills field, the gauge field of translations, \( A^a = e^a = dx^\alpha e^a_{\alpha} \), called the solder form, also describes the space in which the force acts via the relation

\[
g_{\alpha \beta} = e^a_{\alpha} e^b_{\beta} \eta_{ab} \tag{2.5}
\]

where \( \eta_{ab} \) is the Minkowski metric [30]. This means that the symmetry is no longer internal. The identification of the solder form \( e^a \) as basis forms for the co-tangent space (‘soldering’) breaks translational invariance. The degrees of freedom associated with translations are not lost, however, but reappear as general coordinate invariance. The final spacetime does not have full Poincaré invariance, rather the local symmetry is Lorentz only [2, 3, 9].

One drawback of this ‘soldering’ approach to gravitational gauge theory is the redundancy of introducing both a 10-dim symmetry group and a separate 4-dim background space, and then identifying four of the dimensions of the symmetry group with those of the background space. It is tempting to think that the introduction of an independent background space might be avoided. This hope is realized by the group quotient or group manifold method [1, 4, 5], in which the local symmetry and the manifold are simultaneously constructed from an enveloping symmetry. For the Poincaré example, we may construct the base manifold by taking the quotient of the Poincaré group by the Lorentz group. The result is a principal fibre bundle with Lorentz fibres over a 4-dim (or \( n \)-dim) manifold. The Cartan connection on this manifold is generalized by introducing horizontal curvature 2-forms into the Maurer–Cartan equations. This procedure gives not only local Lorentz symmetry as developed by Utiyama [1], but also the solder form as employed by Kibble [2], without the separate introduction of a background space. The model always retains the overall dimension of the original group, with the dimensions partitioned between the local symmetry and the manifold. One can then write an action using the curvatures, solder form and invariant tensors in a way respecting the remaining local Lorentz symmetry. One advantage of these group-theoretic methods is that the fields are determined by the symmetry. This leaves a great deal of freedom in choosing the action, and suggests a way to write a theory that simultaneously looks like Yang–Mills and gravity.

We may now define a Yang–Mills gravity theory to be a Yang–Mills theory with a symmetry group containing the Poincaré group (Poincaré, Weyl, conformal, their covering groups, and their supersymmetric extensions) together with a Yang–Mills action:
\[ S_{\text{YMG}} = \int K_{AB} \Omega^A \Omega^B = \int K_{AB} \Omega^A_{\mu\nu} g^{\mu\nu} \epsilon \, d^6 x \]  

where \( \epsilon = \sqrt{|g|} \).

In light of the preceding paragraphs, we avoid introducing a separate background geometry. Thus, the metric tensor, \( g_{\alpha\beta} \), is constructed from translational gauge fields, \( e^\Sigma, \Sigma = 1, \ldots, n \), according to equation (2.5). As a result, some of the original symmetry is broken, with the corresponding transformations replaced by diffeomorphisms. The field equations differ from a pure Yang–Mills theory, while at the same time differing from general relativity. In particular, the field equations arising from variation of the translational gauge fields will be of the form

\[ D^\mu \Omega^\Sigma_{\alpha\beta} = T^\Sigma_{\beta} \]

where the derivative is covariant with respect to the remaining local symmetry and general coordinate transformations. The appearance of an energy–momentum tensor, \( T^\Sigma_{\beta} \), which makes the theory differ from pure Yang–Mills, is due to the variation of the extra metric terms, \( g^{\mu\nu} \epsilon \). At the same time, the derivative of curvature on the left gives a non-standard gravity theory. Note that the difference between equation (2.6) and the pure Yang–Mills theory of equation (2.3) lies in the dual. Because the solder form that determines the volume form is one of the gauge fields, the dual introduces non-Yang–Mills coupling, \( g^{\mu\nu} \epsilon \).

In addition to the difficulty reproducing either pure Yang–Mills or standard gravity, the Poincaré group encounters an immediate problem with its lack of a Killing metric. Writing the action using the degenerate Killing form leads to vanishing of the energy–momentum tensor constructed from the Riemannian curvature (i.e., \( R_{abc}^d R^{adj} g^{ef} - g_{cd} R_{abc}^d R^{adj} g^{ef} \))—a highly artificial constraint. The other logical symmetry choices of the Weyl group (\( \text{ISO}(3, 1) \times R^* \)) and conformal group with standard gauging (\( \text{SO}(4, 1) \times R^* \)) lead to higher order field equations and non-Yang–Mills couplings. However, as we shall argue below, there exists a Yang–Mills gravity theory which solves the coupling problem.

**Biconformal gauging** of \( \text{SU}(2, 2) \) (or of \( \text{SU}(2, 2|N) \)) is unique in solving some or all of these problems. By taking the quotient of the conformal group by the homogeneous homothetic group (Lorentz plus dilatations) instead of the inhomogeneous homothetic group (Poincaré plus dilatations), biconformal gauging produces a symplectic base manifold with local Lorentz and dilatational symmetry. The symplectic manifold has a natural, dimensionless volume element which may be written without using the gauge fields. This allows the gravitational symmetry to remain internal. Moreover, the reduction of local symmetry reduces the conformally covariant divergence of the Yang–Mills field equation to a Weyl-covariant divergence together with terms algebraic in the curvatures. These algebraic terms can reproduce general relativity in a suitable limit. These improvements are the topic of the next two sections.

### 3. Biconformal Yang–Mills gravity

By **biconformal Yang–Mills gravity**, we mean a Yang–Mills gravity theory of the conformal group, its covering group or its supersymmetric extension in which the gravitational gauging is accomplished by writing the theory on the 2n-bosonic dimensional quotient manifold of the (super)conformal group by its Lorentz plus dilatational sub-(super)group. These gaugings
are described more fully elsewhere [24–26] (and further properties of biconformal space are examined in [31, 32]). For simplicity, we treat only the $n = 4$ case.

Biconformal Yang–Mills gravity makes use of the fact that biconformal spaces have coordinates which are naturally adapted to a symplectic manifold. Half the biconformal coordinates arise from translations, while the other half arise from special conformal transformations (co-translations). Since we may always choose coordinates of these two types, biconformal spaces have almost symplectic structure:

$$\Theta = dx^\alpha dy^\beta.$$ 

The structure is almost symplectic because we are guaranteed these coordinates on each chart of the manifold. The structure may or may not be integrable globally. However, all known classes of biconformal solutions do give rise to a global symplectic form. The symplectic form $\Theta$ enables us to write the volume form as

$$\Phi = (dx^\alpha dy^\beta)^4.$$ 

Taking the dual relative to this volume form does not introduce spurious gauge fields. For a 2-form

$$\Omega^A = \frac{1}{24} \Omega^A_{\mu\nu} dx^\mu dx^\nu + \Omega^A_{\mu\rho} dy^\mu dx^\rho + \frac{1}{2} \Omega^A_{\mu\nu} dy^\mu dy^\nu,$$

(3.1)

the action is of pure Yang–Mills form,

$$S = \frac{1}{4} \int K_{AB} \Omega^A \Omega^B = \frac{1}{4} \int K_{AB} \left( \Omega^A_{\mu\nu} \Omega^{B\mu\nu} - \Omega^A_{\mu\nu} \Omega^{B\nu\mu} \right) \Phi.$$ 

(3.3)

There are no gauge fields in the volume element, $\Phi$, nor is the metric required to form the various contractions. Note that the coordinate-based, metric-free action does not exclude the possibility of an equivalent orthonormal basis.

We define the anti-symmetric symbol in biconformal space to be

$$\varepsilon_{M_1 \cdots M_8} = \varepsilon_{M_1 \cdots 4} \delta_{4+ N_1}^{\delta_{4+ N_1}} \cdots \delta_{4+ N_1}^{\delta_{4+ N_1}}$$

where $M_1 = 1, 2, \ldots, 8$ and $N_1 = 1, 2, \ldots, 4$. It is important to recognize that in this geometry the anti-symmetric symbol is already a tensor, not a tensor density. To see this, recall that in spacetime integrals we require the tensor

$$e_{\mu\nu\gamma\delta} = \sqrt{|g|} \varepsilon_{\mu\nu\gamma\delta}$$

in writing the volume form. But, as indicated above, the full biconformal Levi-Civita tensor reduces to a product of 4-dim Levi-Civita symbols, $e^{\alpha\beta\gamma\delta} e_{\mu\nu\gamma\delta}$. Each of these forms is separately a tensor density, but since they have opposite weights the product is a tensor without the addition of any metric-dependent terms. That is,

$$e^{\alpha\beta\gamma\delta} e_{\mu\nu\gamma\delta} = \frac{1}{\sqrt{|g|}} e^{\alpha\beta\gamma\delta} \sqrt{|g|} \varepsilon_{\mu\nu\gamma\delta} = e^{\alpha\beta\gamma\delta} e_{\mu\nu\gamma\delta}.$$
Moreover, because of the ‘doubled’ volume form, the components of the dual arising from
\[ \Omega^B_{\mu \nu} \]
may be written as
\[ \Omega^B_{\mu \nu} \epsilon_{\mu \nu \alpha \beta} \epsilon_{\rho \gamma \delta} \]
so no metric is required. As a result, \( \Omega^A_{\mu \nu} \) contracts with \( \Omega^B_{\mu \nu} \) and \( \Omega^A_{\mu} \) with itself, again requiring no metric. This means that the usual metric contribution
\[ g^{\alpha \mu} g^{\beta \nu} \sqrt{|g|} \]
to Yang–Mills gravity is absent in biconformal Yang–Mills gravity.

These results carry through when we include supersymmetry. When we gauge the supergroup \( SU(2, 2|N) \) as a biconformal Yang–Mills gravity theory, taking the group quotient, \( SU(2, 2|N)/(R^+ \times SL(2, C) \times SU(N)) \), we also do not require any extra factors involving the gauge fields. In this case, the superspace volume form,
\[ \Phi = (dx^\alpha dy_\alpha)^4 (\theta A \bar{\theta} A)^4 N, \]
is also metric free, since the fermionic variables also occur in dimensionally conjugate pairs.

There is also the question of coordinate invariance. How can the action be written without the usual metric determinants, yet still describe diffeomorphism invariant gravity? The answer lies in the symplectic structure. The volume form \( \Phi = (dx^\alpha dy_\alpha)^4 \) is invariant under 8-dim symplectic transformations, and these include general coordinate transformations of the 4-dim configuration space. When we perform the reduction to the configuration space as described above, setting \( y_\alpha = y_\alpha(x^\beta) \), we must either introduce a connection or continue to transform \( y_\alpha \) inversely to \( x^\alpha \).

Note that conformal supergravity theories exist in 4, 6 and 10 but not in 8 dimensions. The 8-dim biconformal theory uses the superconformal group associated with 4-dim spacetime. It still describes 4-dim gravity.

For simplicity and clarity, we restrict attention to the bosonic conformal sector of the symmetry group for the remainder of the paper. There appears to be no obstacle to writing the theory for the full supergroup. In the following section, we examine the gravity equations that arise from the Yang–Mills action. Remarkably, they still lead to general relativity when the Weyl vector vanishes or in general for vanishing torsion and weak, slowly changing fields.

4. Reduction to general relativity

In this section we produce our next central result: the reduction of biconformal Yang–Mills gravity theory to general relativity in a certain limit. The field equations for the \( SU(2, 2) \) part of the bosonic sector are the usual Yang–Mills field equations,
\[ D_\alpha \Omega^B_{\beta \rho} = D^\sigma \Omega^B_{\beta \sigma} = 0 \]
\[ D^\mu \Omega^B_{a \rho} + D_\alpha \Omega^B_{a \rho} = 0, \]
where the derivatives are conformally covariant with respect to the Latin (group) indices only:
\[ D_\mu \Omega^B_{a \rho} = \partial_\mu \Omega^B_{a \rho} + \Omega^C_{a \rho} \omega^B_{C \mu} \]
\[ = \partial_\mu \Omega^B_{a \rho} - \frac{1}{2} c_{CD}^B \Omega^C_{a \rho} \omega^D_{\mu} \]
where \( c_{CD}^B \) are the structure constants of the conformal group. Normally, this entire expression comprises a single tensor. However, the biconformal quotient breaks the conformal invariance...
to homothetic (Weyl) invariance. As a result, the gauge fields of translations $\omega^a$ (’solder form’) and the gauge fields of special conformal transformations $\omega_a$ (’co-solder form’) become tensors. The conformally covariant derivative then reduces to a Weyl-covariant derivative plus tensor terms linear in the curvatures. Specifically, we have

$$D^\text{Conf}_{S_a bMN} = \partial S /\Omega^a_{bMN} - \omega^a_{bMN} \Omega^a_{cMN} + \Omega^a_{bMN} \omega^a_{cMN} + 4 \Delta^a_{db} \left( \omega^a_{cMN} \omega^a_{dMN} - \omega^a_{cMN} \Omega^a_{dMN} \right)$$

where $a, b = 1, \ldots, 4$ and $\Omega^a_{MN}, \Omega^a_{bMN}$, $\omega^a_{MN}$ are the curvature, torsion, co-torsion and dilatation, respectively, and $\Delta^a_{db} \equiv \left\{ \frac{\partial^a}{\partial y^b} - \eta^a \eta^b \right\}$ (see [25] and [26]). This leads to the field equations

$$0 = D^\text{Conf}_b \Omega^a_{bMN} + D^\text{Conf}_b \Omega^a_{bfaMN} + 4 \Delta^a_{db} \left( \Omega^a_{cMN} \omega^a_{d} - \omega^a_{cMN} \Omega^a_{dMN} \right)$$

and

$$0 = D^\text{Conf}_b \Omega^a_{bMN} + D^\text{Conf}_b \Omega^a_{bfaMN} + 4 \Delta^a_{db} \left( \Omega^a_{cMN} \omega^a_{d} - \omega^a_{cMN} \Omega^a_{dMN} \right)$$

Since the solder and co-solder forms $(\omega_a = \omega_a d y^\beta + \omega_a^a d x^\beta + \omega^a_d d y^\beta)$ are tensors, each field equation has an algebraic piece in addition to the usual divergence terms of a Yang–Mills theory.

In torsion-free biconformal spaces, the second equation becomes purely algebraic:

$$0 = \Omega^a_{bMN} \omega^a_{cMN} + \Omega^a_{bMN} \omega^a_{dMN} - \omega^a_{bMN} \Omega^a_{cMN}$$

(4.1)
This equation is the biconformal equivalent of the Einstein equation. In biconformal spaces with actions linear in the curvature, torsion-free solutions are known to descend to general relativity on a submanifold \([24, 25]\).

One class of exact gravity solutions to the Yang–Mills gravity solutions is known \([39]\). It is found that when the dilatational curvature and torsion vanish the field equations of the linear-curvature theory \([24, 25]\) and the field equations of the full Yang–Mills gravity theory are simultaneously satisfied. The common solution is completely characterized by solutions to the vacuum Einstein equation in 4 dimensions.

In addition, there is a class of solutions to these equations when the Weyl-covariant derivative of the curvatures is small. In this regime, the field equations reduce to the algebraic, curvature-linear equations of familiar gravity theories. For example, consider a solution which reduces under suitable conditions to the Schwarzschild solution. Choosing a local frame in which the connection vanishes at a point, the curvatures are inversely proportional to the square of the radial coordinate, while the derivatives of the curvatures vary as the inverse cube of the radial coordinate. Since the only relevant constant with units of length is the Schwarzschild radius, the condition for the derivative terms to be larger than the curvature terms must be

\[
\frac{1}{r^3} \gtrsim \frac{1}{r_0^2}
\]

or simply

\[
r \lesssim r_0.
\]

It is therefore only within the horizon that the solutions differ significantly from a solution to the algebraic equations.

In any region where we may neglect derivatives of the curvature, the field equations reduce to

\[
0 = 4 \Delta a^d (\Omega_a^c \rho^c_{\beta} \rho^d_{\beta} + \Omega_4^d \rho^d_{\beta} - \omega_{\beta}^c \rho^d_{\beta} - \omega_{\beta}^c \rho^d_{\beta})
\]

\[
0 = \Omega_4^d \rho^d_{\beta} + \Omega_3^d \rho^d_{\beta} - \omega_{\beta}^c \rho^d_{\beta} - \omega_{\beta}^c \rho^d_{\beta}
\]

\[
0 = \Omega^c_{\rho} \rho^c_{\beta} - \omega_{\beta}^c \rho^d_{\beta} - \omega_{\beta}^c \rho^d_{\beta}
\]

\[
0 = 4 \Delta a^d (\Omega_a^c \rho^d_{\beta} + 4 \Delta a^d \Omega_4^d \rho^d_{\beta} - \omega_{\beta}^c \rho^d_{\beta} - \omega_{\beta}^c \rho^d_{\beta} - \omega_{\beta}^c \rho^d_{\beta} - \omega_{\beta}^c \rho^d_{\beta})
\]

\[
0 = \Omega^c_\beta - \omega_{\beta}^c + \Omega_4^d \rho^d_{\beta} - \omega_{\beta}^c \rho^d_{\beta} - \omega_{\beta}^c \rho^d_{\beta} - \omega_{\beta}^c \rho^d_{\beta}
\]

\[
0 = \Omega^c_\beta - \omega_{\beta}^c + \Omega_4^d \rho^d_{\beta} - \omega_{\beta}^c \rho^d_{\beta} - \omega_{\beta}^c \rho^d_{\beta} - \omega_{\beta}^c \rho^d_{\beta}
\]

In addition, we have the biconformal structure equations,

\[
\mathbf{d} \omega_a^c = \omega_a^c \mathbf{d} \omega_a^c + 2 \Delta a^d \omega_b^d \omega_c^f + \Omega_4^d
\]

\[
\mathbf{d} \omega_a^d = \omega_a^d \mathbf{d} \omega_a^d + \omega_b^d \omega_c^f + \Omega_4^d
\]

\[
\mathbf{d} \omega_a^c = \omega_a^c \mathbf{d} \omega_a^c + \omega_b^d \omega_c^f + \Omega_4^d
\]

\[
\mathbf{d} \omega_0^d = \omega_0^d \mathbf{d} \omega_0^d + \Omega_4^d.
\]

To continue, we make two assumptions. First, we assume vanishing torsion

\[
\Omega_\gamma = 0.
\]

This is consistent with our usual expectation for spacetime and also guarantees the existence of a momentum submanifold of biconformal space. Second, we assume the minimum condition
consistent with the existence of a spacetime submanifold. This minimum condition will be discussed below. We show that under these conditions the low-energy field equations reduce to the conformal Einstein equation. The proof is lengthy and will only be summarized here. Details will be found in [33].

We begin by rewriting the curvatures in an orthonormal frame and imposing equation (4.7). Vanishing torsion considerably simplifies the field equations. Obviously, all torsion terms drop out immediately. In addition, the structure equation for the torsion, equation (4.4), becomes

\[ d\omega^a = \omega^b \omega^a_b + \omega^a_0 \omega^a_0 \]  

(4.8)
giving two further conditions. First, we have the corresponding Bianchi identity following from \( d^2 \omega^a = 0 \) which implies

\[
\begin{align*}
\Omega^{abcd}_b & = 0 \\
\Omega^{cd}_0 & = 0 \\
\Omega^{ab}_{cd} & = -2 \Delta^{ac}_{df} \Omega^0_{be} \\
\Omega^{ac}_{bec} - \Omega^{ac}_{cab} & = -(n - 2) \Omega^0_{bc}.
\end{align*}
\]

(4.9)

Second, the structure equation is involute. This means that we can write the solder form as

\[ \omega^a = e^a = e^a_a \text{d}x^a \]  

(4.10)

for some four coordinates \( x^a \). In this coordinate system we have

\[ \omega^{ad} = 0. \]  

(4.11)

Imposing these conditions and writing the resulting field equations in the orthonormal basis \( (\omega^a, \omega_\alpha) \) with

\[
\begin{align*}
\omega^a & = e^a \\
\omega_a & = f_a + b_a = f_a a \text{d}y_a + b_{ab} e^b,
\end{align*}
\]

(4.12)

where \( f_a \) spans the remaining four dimensions, we are left with

\[
\begin{align*}
0 & = (\Omega^{bc}_{ad} + 2 \Omega^{bc}_{ad} h_{ec}) e^c \epsilon^d \beta_f e^\beta + \Omega^{bc}_{ad} f_c e^d \beta e^\beta d b_{\beta} \\
& - b_{ab} \Omega^{bd}_{ad} f_c \epsilon^d \beta e^\beta d - f_a \beta \Omega^{cd}_{ad} + 2 \Omega^{cd}_{ad} b_{bc} e^\beta \epsilon^d e^\beta \\
0 & = \Omega^{bc}_{ad} f_c e^d \beta e^\beta a - \Omega^{ad}_{cd} \epsilon^d \beta e^\beta c \\
0 & = \Omega^{bc}_{ad} f_c \epsilon^d \beta e^\beta d f_a + f_a \beta \Omega^{bc}_{ad} f_c \epsilon^d e^\beta d \\
0 & = \Omega^{bd}_{ad} \epsilon^d \\
0 & = 4 \Delta^{ac}_{db} (\Omega_{cg} f^c + \Omega_{cg} f^c b_{cg}) f_f \beta e^\beta d \\
0 & = \Omega_{cd} e^c + \Omega_{cd} e^c b_{cd} \\
0 & = \Delta^{ad}_{eb} \Omega_e f^d \\
0 & = \Omega^{cd}_{0d}.
\end{align*}
\]

(4.13)

These may be simplified and combined algebraically to show that most components of the curvatures, co-torsion and dilatation vanish:

\[
\begin{align*}
0 & = \Omega^{bc}_{ad} \epsilon^d 0 = \Omega^{bc}_{0d} \\
0 & = \Omega_{ea} 0 = \Omega_{ec} a b \\
0 & = \Omega^{0d}_{0d} 0 = \Omega^{0d}_{00}.
\end{align*}
\]

(4.14)
along with the trace of the spacetime curvature,

\[ 0 = \Omega_{\alpha\beta} \]

Next, the Frobenius theorem implies the existence of submanifolds found by setting \( \omega^a = 0 \). We study the form of these submanifolds. As a consequence of the vanishing ‘momentum’ curvatures,

\[
0 = \Omega_{abcd}^b \\
0 = \Omega_{\alpha\beta}^{ab} \\
0 = \Omega_0^{ab}
\]

(4.16)

the biconformal structure equations (4.3)–(4.6) reduce to the structure equations of a flat, 4-dim Riemannian geometry. This allows us to find a gauge in which the connection on the full biconformal space takes the form

\[
\omega^a_b = \alpha^a_b e^c \\
\omega^a = e^a \\
= e_a^a \delta(x) dx^a \\
\omega^a = e_a^a \delta(x) dy + b_a e_b \\
\omega_0^a = W_c e^c = W_\beta d x^\beta.
\]

(4.17)–(4.20)

With this form of the connection, the full structure equations now reduce to

\[
d\omega^a_b = \omega^a_c \omega^c_b + 2 \Delta_{cb}^d \omega^d e^c + \frac{1}{2} \Omega_{abcd} e^e e^d \\
d e^a = e^a \omega^a + \omega^b e^a \\
d \omega_a = \omega^a_b \omega_b + \omega_a \omega^0 + \frac{1}{2} \Omega_{abc} e^b e^c \\
d \omega_0^a = e^a \omega_0
\]

(4.21)–(4.24)

The rest of the calculation consists in substituting the form of the connection above, equations (4.17)–(4.20), into the structure equations, equations (4.21)–(4.24). After considerable algebra, the connection reduces to

\[
\omega^a_b = \alpha^a_b - 2 \Delta_{db}^a W_c e^d \\
\omega^a = e^a \delta(x) \\
\omega_a = f_a + b_a \\
\omega_0^a = W_c e^c = -y^\beta d x^\beta
\]

(4.25)–(4.28)

where \( \alpha^a_b \) is the spin connection compatible with the solder form \( e^a \delta(x) \). Let \( R^a_{bcd}, \Omega_{ab} \) and \( R \) be the Riemann curvature tensor, Ricci tensor and Ricci scalar, computed from \( e^a \) and \( \alpha^a_b \). Then, the 1-forms \( f_a \) and \( b_a \) are given by

\[
f_a = e^a \delta(x) dy^\beta \\
b_a = R_a - e_a^\nu y_\mu \Gamma^\nu_{\nu \mu} - y_\alpha y_c e^c + \frac{1}{2} \eta_{ac} (\eta^{\beta b} y^h y^h) e^c
\]

(4.29)–(4.30)

where

\[
R_a = -\frac{1}{2} (R_{ab} - \frac{1}{6} R \eta_{ab}) e^b
\]

(4.31)
is the Eisenhart tensor \([34]\) and
\[
\Gamma^\mu_\nu = \Gamma^\mu_{\nu\lambda} dx^\lambda
\]
is the Christoffel connection compatible with \(e^a(x)\). The structure equations take the form
\[
\begin{align*}
\omega^a_b &= \omega^a_c dx^c + 2 \Delta^{ac}_{db} \omega^a_d e^c + C^a_b \\
\omega^a_b &= \delta^a_b + \frac{1}{2} \eta_{ac} (\eta^b_h y_h y_b) e^c
\end{align*}
\]
(4.32)

\[
\begin{align*}
de^a &= e^b \omega^a_b + \omega^a_0 e^a \\
\omega^a_0 &= e^a \omega_a
\end{align*}
\]
(4.33)

\[
\begin{align*}
\omega^a_0 &= \omega^a_0 + \Delta^{ac}_{db} \omega^a_d e^c + 1 \cdot \eta_{ac} (\eta^d_h \phi^c) e^c \\
\omega^a_0 &= \omega^a_0 + \frac{1}{2} \eta_{ac} (\eta^b_h \phi^c) e^c
\end{align*}
\]
(4.34)

\[
\begin{align*}
\omega^a_0 &= \omega^a_0 + \Delta^{ac}_{db} \omega^a_d e^c + 1 \cdot \eta_{ac} (\eta^d_h \phi^c) e^c
\end{align*}
\]
(4.35)

where
\[
C^a_b = R^a_{bc} - 2 \Delta^{ac}_{bd} \omega^a_d e^d
\]
is the Weyl curvature (i.e., the traceless part of the Riemann curvature) and
\[
D(x, \alpha) = \omega^a_0 e^a
\]
for any contravariant vector, \(u^a\).

We now return to our second assumption: the minimal condition necessary to guarantee the existence of a spacetime submanifold. This is provided by a second involution, this time of the co-solder form, \(\omega^a_b\). Setting \(\omega^a_b = 0\) we have
\[
0 = \omega^a_b = e^a_\beta(x) dy^\beta + \Delta^{ac}_b \omega^a_d e^c + C^a_b
\]
(4.36)

To examine the consequences of these equations, we first consider equation (4.36), which has become a differential equation for a hypersurface, \(y_a = y_a(x)\). We first rewrite the derivative term as
\[
e^a_\beta(x) dy^\beta = dy^a - y^a_\beta e_a^\beta
\]
(4.37)

where \(e^a_\beta\) is the spin connection compatible with the solder form \(e^a(x)\). Rearranging, we have
\[
0 = D(x, \alpha) = \omega^a_0 e^a
\]
(4.38)

Because \(y_a\) is the negative of the Weyl vector, equation (4.38) is closely related to the change in the Ricci and Eisenhart tensors under a conformal transformation ([34]),
\[
\tilde{R}_{ab} = R_{ab} - \eta_{ab} \nabla^c \phi - (n - 2) [\phi_{ab} - \phi_{a} \phi_{b} + \eta_{ab} \phi_{c} \phi_{c}]
\]
\[
\tilde{R}_a = R_a + d \phi_a - \phi_{a} \phi_{b} - \phi_{a} \phi_{c} e^c + \frac{1}{2} \eta_{ac} (\eta^d_h \phi^c) e^c
\]
where \(\phi_{ab} = \phi_{a} \phi_{b} + \phi_{b} \phi_{a}\) and \(n\) is the dimension. Specifically, note that if we could replace \(y_a\) with \(\phi_a\) in equation (4.38) we would have exactly the condition \(\tilde{R}_a = 0\), equivalent to the vacuum Einstein equation in the conformally transformed basis. Therefore, equation (4.38), together with
\[
y_a dx^a = y_e e^c = d \phi_c
\]
(4.39)
guarantees the existence of a conformal gauge in which the vacuum Einstein equation holds.
We now show that the reduced structure equations, equations (4.37), provide the integrability conditions for equations (4.38) and (4.39). The integrability conditions are given by the Poincaré lemma, $d^2 = 0$. Applying this first to equation (4.38), we have

$$0 = d^2 y_a = dy_b \alpha^b_a + y_a d \alpha^b_a + dy_a \gamma^e_c e^c + y_a d(y_c e^c) - \eta_{ae} \eta^{bh} d y_b \gamma^e_c e^c - \frac{1}{2} \eta_{ae} \eta^{bh} (\gamma^e_c y_b) d e^c - d R_a.$$ 

Substituting for all occurrences of $dy_b$ we find after some cancellations,

$$0 = y_b R^b_a - D R_a - R_a y_c e^c - (\eta^{bc} y_b R^c) \eta_{ad} e^d + y_a d(y_c e^c).$$ 

Now, we substitute for the curvature 2-form,

$$R^a_b = C^a_b + 2 \Delta^{ab}_{cd} R^e_c e^d.$$ 

This reduces the integrability condition to

$$0 = y_b C^b_a - D R_a + y_a d(y_c e^c).$$ 

Turning now to the integrability condition for equation (4.39), we have

$$0 = d^2 \phi = d(y_c e^c).$$ 

Combining these two conditions as the pair

$$y_b C^b_a - D R_a = 0$$

and recalling that $W_a = -y_a$, we see that the reduced structure equations, (4.37), provide exactly these conditions. Therefore, there exists a choice of the conformal gauge such that the Einstein equation holds on the $\omega_a = 0$ submanifold. This same choice reduces the Weyl vector to zero and the remaining structure equations to

$$d \alpha^a_b = \alpha^a_b \alpha^b_a + \frac{1}{2} C^a_{bcd} e^d$$

and therefore

$$d(y_a e^a) = 0$$

by the symmetry of the Eisenhart tensor.
The meaning of this additional result is clearest if we begin with the expression for the change in the Eisenhart tensor under a conformal transformation,

\[ \tilde{R}^a_{\alpha} = R^a_{\alpha} + \phi_c \epsilon^c_{\alpha} \]

Treatting \( \phi_a = \phi_i \) as a vector, we see that

\[ \phi_b C^b_{\alpha} - DR_{\alpha} = 0 \] (4.40)

is the integrability condition for the existence of a vector field \( \phi_i \) such that \( \tilde{R}_a = 0 \). Then, contracting with the solder form as above, we see that \( d(\phi_i \epsilon^i) = 0 \) so that \( \phi_i \) must be a gradient. Equation (4.40) alone is therefore a sufficient condition for the existence of a conformal transformation to a Ricci flat spacetime. Szekeres [36] uses the spinor representation to show that equation (4.40) may be written as a constraint on the curvatures which is independent of \( y_\alpha \). It follows from the results of [36] that equation (4.40) is also necessary. Further, our result shows the equivalence of certain well-known conditions: the \( C \)-spaces \( (\phi_b C^b_{\alpha} = DR_{\alpha} = 0) \) of Szekeres [36], the \( J \)-spaces \( (DR_{\alpha} = 0) \) of Thompson ([35]) and conformally Ricci flat spaces. This follows because Ricci flatness implies the \( J \)- and \( C \)-conditions, while we have shown that the \( C \)-condition implies conformal Ricci flatness.

5. \( U(1) \) Yang–Mills in biconformal space

We will now pause in our development of a gravity theory to consider the structure of a non-gravitational gauge theory on biconformal space. In particular, in this section we will show that any 4-dim pure Yang–Mills theory can be consistently written in biconformal space. We establish an invertible map between the 4-dim and 8-dim theories and show that the 8-dim field equations are satisfied if and only if the 4-dim equations are satisfied. We provide the construction here in detail for \( U(1) \). The non-Abelian case is discussed in the following section.

Suppose we have a \( U(1) \) gauge theory in 4-dim, with structure equation, field equation and conservation law, given by

\[ A = A_a(x) dx^a \] (5.1)

\[ F = dA \] (5.2)

\[ dF = d^2 A = 0 \] (5.3)

\[ ^* d^* F = J \] (5.4)

\[ ^* d^* J = 0 \] (5.5)

We wish to extend this to biconformal space. First, we show that the co-tangent bundle extends to a biconformal space. Let \( (x^a, y^b) \) be coordinates on the bundle, which we treat as an 8-dim manifold. Extend the bundle by 7-dim fibres, with each fibre isomorphic to the spacetime homothetic group. We may then define a connection on the resulting 15-dim bundle,

\[ \omega = - y_a dx^a \]

\[ \omega^a = dx^a \]

\[ \omega_a = dy_a - \alpha (y_a y_b - \frac{1}{2} y^2 \eta_{ab}) dx^b \]

\[ \omega^a_b = (y_b dx^a - y^a dx_b) \] (5.6)
This connection satisfies the Maurer–Cartan structure equations for the conformal group,
\[ d\omega = \omega^a \omega_a, \]
\[ d\omega^a = \omega^b \omega^a_b + \omega^a \omega^a, \]
\[ d\omega_a = \omega^b \omega_b + \omega_a \omega \]
\[ d\omega^a_b = \omega^c \omega^a_b + 2 \Delta^{ac} \omega^a \omega^c. \]  
(5.7)

Because the bundle is homothetic over an 8-dim base manifold, it is a biconformal space.

Next we extend the field theory to the biconformal space. Let
\[ A = A + \frac{1}{2} F^{ab} y_b d y_b, \]
\[ F = dA, \]
\[ F = F - \frac{1}{2} y_a F^{ab} d y_b d x^c + \frac{1}{2} F^{ab} d y_b d y_b. \]  
(5.8)

It follows that
\[ dF = d^2 A = 0 \]
\[ d^* F = J^b, \]  
(5.9)

(Note that this extension is similar in spirit to the ‘generalized differential calculus’ used by Guo et al to establish correspondences between Chern–Simons and BF theories and in writing supergravity and Yang–Mills theories as generalized topological field theories. See, e.g., [37, 38].)

Now from the original 4-dim Yang–Mills theory we have
\[ F^{ab}, a = J^b \]  
(5.10)

and it follows that
\[ d^* F = \frac{1}{2} J^b d y_b. \]  
(5.11)

Since the right-hand side is constructed purely from the 4-dim current, we may define
\[ J = \frac{1}{2} J^a d y_a - \frac{1}{2} y_a J^a, d x^b \]  
(5.12)

and write the 8-dim biconformal equation
\[ d^* F = J. \]  
(5.13)

This completes the biconformal equations governing \( A, F \) and \( J \). Conservation of \( J \) follows from \( d^* d^* J = 0 \), but is dependent upon the conservation of the original 4-current, since direct calculation shows that
\[ d^* J = \frac{1}{2} J^a d y_a - \frac{1}{2} y_a J^a, d x^b \]  
(5.14)

It is now immediate that for any biconformal current of the form \( J = \frac{1}{2} J^a d y_a - \frac{1}{2} y_a J^a, d x^b \) the biconformal field equation for \( F \) is satisfied if and only if the original Minkowski space Yang–Mills equation is satisfied. By construction, the Minkowski space Yang–Mills solution guarantees that the biconformal equations are satisfied. Conversely, suppose we can write the biconformal current as
\[ J = \frac{1}{2} J^a d y_a - \frac{1}{2} y_a J^a, d x^b. \]  
(5.15)

Then we need only the ansatz \( A = A + \frac{1}{2} F^{ab} y_b d y_b \) to find that the field equation takes the form
\[ -\frac{1}{2} y_a F^{ab}, c d x^c + \frac{1}{2} F^{ab}, d y_b = \frac{1}{2} J^a d y_a - \frac{1}{2} y_a J^a, d x^b. \]  
(5.16)

Equating the coefficients of \( d y_b \) gives
\[ F^{ab}, a = J^b \]  
(5.17)

with the remaining coefficients of \( d x^c \) matching identically. This completes the proof.
6. Non-Abelian Yang–Mills in biconformal space

In this section we generalize the results of section 5, embedding a general four-dimensional Yang–Mills theory into eight-dimensional biconformal space. For non-Abelian Yang–Mills theories, the calculations are similar to those given in the previous section. We will state the result here and provide details in the appendix.

Consider the extension of a generic SUSY Yang–Mills theory to biconformal space. We consider only the bosonic sector for simplicity. Beginning with the 4-dim equations

\[ A^A = A^A_B G_B = A^A_B(x)dx^a \]  
\[ F^B = F^B_B + \frac{1}{2}A^B \wedge A^C [G_B, G_C] \]

\[ DF^B = dF^B - F \wedge A + A \wedge F = 0 \]

where the covariant exterior derivative of a vector \( v^A \) in the Lie algebra is given by

\[ Dv^A = dv^A + c^A_{BC} A^B \wedge v^C. \]

We extend by defining

\[ A^B = A^B + \frac{1}{2}F^B_{ab} y_a d y_b \]

or, more explicitly,

\[ A^B = A^B + \frac{1}{2}F^B_{ab} y_a d y_b \]

\[ F^A = dA^A + A^A \wedge A \]

If in eight-dimensional biconformal space we use these fields to form the pure Yang–Mills action with source,

\[ S_{BC} = \int K_{AB} (F^A \wedge F^B + A^A \wedge J^B), \]

then the eight-dimensional field equations take the form

\[ \ast D^A \ast F^A = J^A \]

where we define \( J^A \) to be the 8-dim current,

\[ J^A = 2 \left( (J^A_{,a} + \frac{1}{2}c_{FG} F^GA^F_{,a}) y_a d x^a \right) + \frac{1}{2} J^A_{,a} y_a (\ast (J^B_{,a} F^C_{,a} - 5 F^B_{,a,cd} F^C_{,d})) d y_a \]

By inspection, we see that if we set \( y = 0 \), we return to the original 4-dim field equations. Furthermore, direct calculation shows that \( J^A \) is conserved if and only if the 4-dim current, \( J^A_{,a} \), is conserved. That is,

\[ \ast D^A \ast J^A = \frac{5}{2} J^A_{,a}. \]

As a result, the 8-dim field equations are satisfied if and only if the 4-dim equations hold. It is of some interest to note that the biconformal current equation (6.11) depends on both the
field and current of the 4-dim theory (see the appendix). This opens up the possibility that vacuum solutions to the biconformal field theory could correspond to Yang–Mills theories with sources—that is, we have a geometric description of certain matter fields. Finally, it is also worth noting that the field equations are a local statement, and thus the integral equation (6.9) may be regarded as convergent even for a field strength of the form equation (6.8) if we take the integration over a compact subset of spacetime, with the variation taken to be zero on the boundary. Since we can do this for any region, and can cover the space with such regions, the field equations will hold at every point.

This construction provides a new arena for investigating the behaviour of such standard four-dimensional theories as $SU(N)$ Yang–Mills or $N = 4$ SUSY Yang–Mills in four dimensions. As essentially the only known finite quantum field theory, the study of $N = 4$ SUSY Yang–Mills might provide insight into the quantum behaviour of gauge theories in biconformal space. Additionally, although the four-dimensional, flat-space, SUSY Yang–Mills theory is embedded into another Ricci flat geometry in eight dimensions, the geometry of biconformal space is non-trivial. For example, the symplectic and phase space structures of biconformal space [32] could possibly yield interesting new effects in the extended Yang–Mills theory.

In summary of the previous two sections, we have shown that the co-tangent bundle of Minkowski space extends to a biconformal bundle and that any Yang–Mills theory on 4-dim Minkowski space correspondingly extends to an equivalent field theory on a biconformal space. This result is important for establishing that a Yang–Mills theory in biconformal space may be an inherently 4-dim theory. A similar result holds for biconformal gravity, where it has been shown that the field equations for a curved biconformal geometry reduce to general relativity on a 4-dim submanifold [24]. Understanding that 8-dim biconformal spaces describe 4-dim physics helps clarify the naturalness of our final result below. We turn now to four-dimensional Yang–Mills theories that lift to curved biconformal spaces.

7. A special class of biconformal gravity solutions

We now investigate how an unusual choice of gauge group in 4 dimensions can lead to a new class of biconformal Yang–Mills gravity solutions. Suppose we begin with a flat-space non-compact $SU(2, 2|N)$ SUSY Yang–Mills theory on 4-dim Minkowski space. That is, 

$$ S_{YM} = \int K_{AB} F^A_{\mu\nu} F^B_{\rho\sigma} \eta^{\mu\rho} \eta^{\nu\sigma} \, d^4x $$

(7.1)

where $K_{AB}$ is the Killing metric of $SU(2, 2|N)$. As was demonstrated in the previous section, we can write an equivalent field theory on biconformal space, equations (6.7), (6.8). However, the gauge group of this Yang–Mills theory is identical to the gauge group of biconformal supergravity [26]. This means that we can identify the Yang–Mills field strengths with the curvatures of the biconformal Yang–Mills (super-)gravity theory.

The fact that makes this possible is that these field strengths, defined on an 8-dim manifold, satisfy the same conformal structure equations as the 4-dim theory,

$$ F^A = dA^A - \frac{1}{2} c_{BC} A^B A^C. $$

(7.2)

This means that we can identify the potentials $A^B$ with the biconformal connection $\omega^A$, and the Yang–Mills curvatures $F^A$ with the biconformal curvatures $\Omega^A$, i.e.,

$$ A^A = \omega^A, \quad F^A = \Omega^A. $$

(7.3)

Since we have chosen a Yang–Mills-type action, equation (3.3), to describe the biconformal gravity theory, and used the metric-independent volume form, the field equations are also identical.
The only difference between the biconformally extended Yang–Mills equations and the biconformal gravity equations is that (before imposing the field equations) the biconformal curvatures are generic 2-forms in 8-dim, while the extended Yang–Mills field strengths take the particular form of equation (6.8). However, identifying the field strengths and curvatures is acceptable as an ansatz for a biconformal solution. Allowing for differences of homothetic gauge, \( g \), conformal gauge, \( h \), and coordinates, this ansatz takes the form

\[
g^A g^{-1} (u(x, y), v(x, y)) = h F^A h^{-1} (x, y). \tag{7.4}
\]

Since the extended Yang–Mills equations are solved if the original 4-dim Yang–Mills theory is satisfied, we have written a class of solutions to the biconformal field equations. It is equation (7.4) that provides the equivalence between a 4-dim gauge theory on flat space and an 8-dim biconformal gravity theory.

Further, the gravity solutions are four-dimensional, because the full curvatures are determined by the connection on the 4-dim submanifold found by holding \( \omega a = 0 \). That is, there are two ways to regard equation (6.8). First, regarded as a Yang–Mills field strength, \( F^A |_{y=0} \) reduces to a Yang–Mills field on 4-dim Minkowski space. Second, regarded as a biconformal curvature, \( \Omega^A |_{y=0} = F^A |_{y=0} \) reduces to a gravity solution on a 4-dim submanifold of a biconformal space. This is a correspondence between a theory containing gravity and a gauge theory without it. It is worth stressing again here that this relationship between flat and curved geometries is only possible because of the symplectic structure of biconformal spaces. It is that structure that allows us to write the action without introducing a metric, so that the resulting field equations are identical to those of the flat-space, non-compact, SUSY Yang–Mills theory.

8. Conclusions

We have accomplished our goal of constructing a Yang–Mills theory of gravity by using the biconformal gauging of the conformal group. In addition, we have provided a systematic extension of flat-space Yang–Mills theories to biconformal space. The resulting geometries are an arena for both biconformal Yang–Mills gravity theory and extended Yang–Mills theories and possess a number of new and interesting properties:

1. Biconformal Yang–Mills gravity theories describe 4-dim, scale-invariant general relativity in the case of slowly changing fields and vanishing torsion.
2. The non-trivial systematic biconformal extension of any 4-dim Yang–Mills theory to a Yang–Mills theory on a 8-dim biconformal space provides a new arena for analysing the properties and behaviour of flat-space Yang–Mills theories.
3. By applying the biconformal extension above to a 4-dimensional pure Yang–Mills theory with conformal symmetry, we establish a 1-1, onto mapping between a set of gravitational gauge theories and 4-dim, flat-space gauge theories.
4. Since the biconformal Yang–Mills current, \( J^A \), depends on both the four-dimensional fields and currents, some vacuum configurations of the 8-dim field strength give rise to a non-trivial source term in the 4-dim Yang–Mills theory. This provides a geometric origin for some matter fields.

The extension in (2) above is obtained by extending the theory from a 4-dim (flat) manifold to its 8-dim co-tangent bundle \( (T^* M) \), then requiring a conformal connection, \( \omega^2 \), on \( T^* M \) to obtain a Yang–Mills theory on a biconformal space. Although the extension is non-trivial, it is readily invertible. After extension, the original theory may be re-obtained by setting the \( y \)-coordinates to zero.
In the special class of Yang–Mills gravity solutions described in (3), we choose simple but non-compact $O(4, 2)$ or $SU(2, 2)$ Yang–Mills theories, and for their supersymmetric extension $SU(2, 2|N)$ non-compact Yang–Mills theory. Then, biconformal Yang–Mills gravity theory has a class of solutions for the curvatures, $\Omega^A$, that satisfy $\Omega^A = F^A$, where $F^A$ is an extended $SU(2, 2|N)$ Yang–Mills field strength. These curvatures, $\Omega^A$, satisfy the gravitational field equations if and only if the original Yang–Mills field strengths, $F^A$, satisfy the super Yang–Mills equations in 4-dim Minkowski space. The fact that the biconformal space described in (3) is a curved geometry, while the equivalent Yang–Mills system has Minkowski base manifold, is made possible by the symplectic structure of biconformal spaces and their resulting dimensionless volume forms.

The field equations for the Yang–Mills theory in 4-dim are completely determined by the potentials $A^A_{\mu}(x)$ and as a result contain a total of 60 degrees of freedom. When it is possible to impose both Dirac- and Coulomb-type gauge conditions, there will remain 30 degrees of freedom, two for each potential. However, the torsion-free biconformal field equations are completely determined by the solder form, $e^a(x)$, which, subtracting 7 for the gauge choice (corresponding to the six local Lorentz symmetries plus one dilatational symmetry), is a total of only nine degrees of freedom. As a result, there may be considerably more freedom in the class of extended Yang–Mills solutions to biconformal gravity than in the torsion-free solutions. Moreover, the condition of vanishing torsion is not gauge invariant in the Yang–Mills theory (though it is in biconformal space). Allowing for homothetic ($g$) and coordinate freedom in the biconformal solutions, and the conformal symmetry of the Yang–Mills solution ($h$) in equating the fields

$$(g\Omega^A g^{-1}(u(x, y), v(x, y)) = h F^A h^{-1}(x, y)),$$

we make the conjecture that all torsion-free solutions to the biconformal Yang–Mills field equations can be identified with extended Yang–Mills theories.

If this conjecture holds, then there is a 1-1 correspondence between solutions for torsion-free biconformal supergravity with a class of 4-dim non-compact SUSY Yang–Mills theories. In any case, we have shown that the weak-field limit of extended Yang–Mills solutions gives rise to a gravity theory on a 4-dim submanifold of biconformal space. Further study is required to know if this set includes torsion-free solutions [39].

While the theories presented in this work have been purely classical, the quantum properties of the biconformal gravity theory are a potentially rich topic for further investigation. Since the field equations obtained from the Yang–Mills gravity actions are linear in the curvatures, we do not expect ghosts for the reasons that are common in standard conformal gravity theories. However, because the theory is formulated in a 8-dim symplectic space, it not clear how to apply the standard techniques of ghost analysis. As a result, the existence of ghosts in generic biconformal supergravity is a topic for further study.

Appendix. Explicit biconformal extension of non-Abelian Yang–Mills theory

Here, we consider the biconformal extension of a non-Abelian Yang–Mills theory. Suppose we have a gauge theory with Lie gauge group $G$ (gauge elements $g$, generators $G_A$, structure constants $c_{ABC}$) in 4-dim Minkowski space, with potential, field strength and Bianchi identity given by equations (6.1)–(6.4). With these fields, the standard Yang–Mills action with source

$$S_M = \int K_{AB} (F^A + F_B)$$

leads to the field equation, the 4-dim equation,

$$^*D^*F = -J$$
\[ D_\mu F^{Ac}_e = -J^A_e \]

where

\[ D_a F^{Ac}_e = F^{Ac}_e,a + c_{BC}^A F^{Cc}_e A^B_a. \]  \hspace{1cm} (A.3)

We wish to extend this to biconformal space. As before, we construct the biconformal space from the co-tangent bundle with coordinates \((x^a, y_b)\) extended by 7-dim homothetic fibres. We may then define a biconformal connection on the full 15-dim bundle, equation (5.6), which satisfies equation (5.7). If the Yang–Mills group is the superconformal group, \(SU(2, 2|N)\), we identify the extended Yang–Mills potentials with the biconformal connection.

In either case, we define the extended Yang–Mills potential,

\[ A = A + \frac{1}{2} F^{Aab} y_a G A d y_b \]  \hspace{1cm} (A.4)

(and hence equation (6.7)). The potential remains a Lie-algebra-valued 1-form, and therefore satisfies the usual Maurer–Cartan structure equations. It follows that the field is given by

\[ F = dA + A \wedge A \]

which leads to equation (6.8). As in the \(U(1)\) case, the biconformal potential and field reduce to the 4-dim potential and field when \(y = 0\).

In general, we may think of the extension from \(A\) to \(A\) as a change of connection, so that the difference is a tensor. The covariant derivative will then change by the addition of a tensorial term, and the curvature by the addition of the covariant derivative of the new tensor, plus the square of the new tensor. The above calculation confirms this.

The Bianchi identity also generalizes to

\[ \mathbf{D} \mathcal{F} = (\mathbf{d} \mathcal{F}^A + \epsilon_{BC}^A A_B^H \wedge \mathcal{F}^C) G_A = 0 \]

where the covariant derivative is given by

\[ \mathbf{D} = D_\mu dx^\mu + D^\rho d y_\rho \]  \hspace{1cm} (A.5)

with

\[ D_\mu X^A = \partial_\mu X^A + \epsilon_{EF}^A A^E X^F \]

\[ D^\rho X^A = \partial^\rho X^A + \frac{1}{2} \epsilon_{EF}^A A^E [X^F, y_\rho]. \]  \hspace{1cm} (A.6)

The field equations from the action written with the 8-dim fields

\[ \mathcal{S}_{BC} = \int K_{AB}(\mathcal{F}^A \mathcal{F}^B + A^A \mathcal{J}^B) \]  \hspace{1cm} (A.7)

then take the form

\[ *\mathbf{D}^* \mathcal{F}^A = \mathcal{J}^A \]  \hspace{1cm} (A.8)

where direct computation yields

\[ *\mathbf{D}^* \mathcal{F}^A = \frac{1}{2} \left( (F^{A}_{\rho a \mu} p_a + \frac{5}{2} \epsilon_{G}^F A^F G^a p_{\mu}) y_a d x^\rho \right. \]

\[ + \left( \frac{3}{2} F^{A \rho a}_{\mu} p_a + \frac{5}{8} \epsilon_{BC}^A y_a y_c (-F^{B}_{\rho a \mu} p F^{C}_{\rho c d} - 5 F^{B}_{\rho a d} F^{C}_{\rho c d}) d y_a \right) \]

and \(\mathcal{J}^A\) is the 8-dim current, equation (6.11). It is evident that if we set \(y = 0\), we return to the original 4-dim field equations.

Furthermore, the 8-dim current, \(\mathcal{J}^A\), is conserved if and only if the 4-dim current, \(J^A_{\mu}\), is conserved. That is,

\[ *\mathbf{D}^* \mathbf{D}^* \mathcal{F}^A = \frac{5}{2} J^A_{\mu \nu}. \]  \hspace{1cm} (A.10)

As a result, the 8-dim field equations are satisfied if and only if the 4-dim equations hold.
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