We study the relationship between space-time-matter (STM) and brane theories. These two theories look very different at first sight, and have different motivation for the introduction of a large extra dimension. However, we show that they are equivalent to each other. First we demonstrate that STM predicts local and non-local high-energy corrections to general relativity in 4D, which are identical to those predicted by brane-world models. Secondly, we point out that in brane models the usual matter in 4D is a consequence of the dependence of five-dimensional metrics on the extra coordinate. If the 5D bulk metric is independent of the extra dimension, then the brane is void of matter. Thus, in brane theory matter and geometry are unified, which is exactly the paradigm proposed in STM. Consequently, these two 5D theories share the same concepts and predict the same physics. This is important not only from a theoretical point of view, but also in practice. We propose to use a combination of both methods to alleviate the difficult task of finding solutions on the brane. We show an explicit example that illustrate the feasibility of our proposal.

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1 INTRODUCTION

The basic ideas of what today is called space-time-matter theory (STM) have been developed by a number of people [1]-[10]. In this theory, our four-dimensional world is embedded in a five-dimensional spacetime, which is a solution of the five-dimensional Einstein’s equations in vacuum. The extra dimension is not assumed to be compactified, which is a major departure from earlier multidimensional theories where the cilindricity condition was imposed. In this theory, the original motivation for assuming the existence of a large extra dimension was to achieve the unification of matter and geometry, i.e., to obtain the properties of matter as a consequence of the extra dimensions.

Recent results in string theories suggest that gravity is indeed a multidimensional interaction, and that usual general relativity in 4D is the low energy limit of some more general theory. In these theories the matter fields are confined to our 4D spacetime, embedded in a 4 + d dimensional spacetime, while gravity fields propagate in the extra d dimensions as well. In 5D a number of works model our 4D universe as a domain wall in a five-dimensional anti-de Sitter spacetime. In this context, the motivation for large extra dimensions is to solve the hierarchy problem [11]-[12].

This has promoted an increasing interest in gravity theories formulated in spacetimes with large extra dimensions. Brane-world models are inspired by these theories. In these models our universe is a four dimensional singular hypersurface, or “brane”, in a five-dimensional spacetime, or bulk [13]-[16].

Although STM and brane theory have different physical motivations for the introduction of a large extra dimension, they share the same working scenario. Namely, (i) they allow the bulk metrics to have non-trivial dependence of the extra dimension; (ii) the 4D metric is obtained by evaluating the background metric at some specific 4D hypersurface that we identify with our physical spacetime; (iii) the matter fields are confined to a 3-brane in a 5D spacetime, which is a solution to Einstein’s equations; (iv) observers are bound to the brane, unable to access the bulk.

From a practical point of view, they share the same goals. Among them, to predict the effects of the bulk geometry on the brane geometry and dynamics.

Despite of these common grounds, both theories remain as separated, unrelated, entities. The aim of this work is to remedy this situation. Our first goal is, therefore, to show that both theories are equivalent. We will show that STM includes the so-called local high-energy corrections, and non-local Weyl corrections typical of brane-world scenarios. Also that the matter in the brane is purely geometric in nature.

The difference in the motivation for large extra dimensions is reflected in the techniques adopted by authors in each theory. In STM the authors start from the geometry of the bulk and construct the physics in 4D [7]. In brane models, the opposite point of view is taken. Namely, the effective equations in 4D are solved in
the brane for some matter distribution on it. Then the solution is matched to some appropriate bulk metric that satisfies Israel’s boundary conditions [17]-[19].

The problem of finding a complete solution in brane theory is a really involved one. Therefore, our second goal in this work is to propose we use the formal equivalence between the two theories to alleviate this problem. Technically, what we propose is that we incorporate the physics of brane models to interpret and determine the final form of solutions in STM. We show the feasibility of this approach by means of an explicit example.

2 Equivalence Between STM and Brane Theory

In this Section we first show that STM incorporates and predicts the same physics as brane-world models. We then show that brane models include and share with STM the same philosophy about the geometrical nature of matter in $4D$. Finally, we discuss the different points of view adopted in STM and brane-world models, regarding physics in $4D$.

2.1 Equations In Space-Time-Matter Theory

The metric is taken as

$$dS^2 = g_{\mu\nu}(x^\rho, y) dx^\mu dx^\nu + \epsilon \Phi^2(x^\rho, y) dy^2,$$

where $\epsilon$ is $-1$ or $+1$ depending on whether the extra dimension is spacelike or timelike, respectively. In what follows $f^* = \partial f/\partial y$, and the covariant derivatives are taking with respect to $g_{\mu\nu}$. For the signature of the spacetime and definition of tensor quantities we follow Landau and Lifshitz [20].

The field equations in STM theory are the Einstein equations in vacuum, viz.,

$$(5) R_{AB} = 0.$$ The $4 + 1$ splitting of these equations provides a definition for an effective energy-momentum tensor [3]

$$(4) G_{\alpha\beta} = \frac{\Phi_{\alpha\beta}}{\Phi} - \frac{\epsilon}{2\Phi^2} \left[ \Phi g_{\alpha\beta}^{**} - \Phi g_{\alpha\beta}^{**} g_{\lambda\mu} g_{\alpha\lambda} g_{\beta\mu} - \frac{1}{2} g_{\mu\nu} g_{\alpha\beta}^{**} + \frac{1}{4} g_{\alpha\beta} \left( g_{\mu\nu}^{**} + (g_{\mu\nu} g_{\mu\nu})^2 \right) \right],$$

an equation governing the scalar field $\Phi$

$$(3) \epsilon \Phi \Phi_{\mu\nu}^{**} = -\frac{1}{4} g_{\lambda\mu}^{**} g_{\lambda\mu} - \frac{1}{2} g_{\lambda\mu} g_{\lambda\mu} + \frac{\Phi}{2\Phi^2} g_{\lambda\mu}^{**} g_{\lambda\mu},$$

and a conservation equation, which is usually written as

$$(4) P_{\alpha\beta} = 0$$
where
\[ P_{\alpha \beta} = \frac{1}{2\Phi} (g^\ast_{\alpha \beta} - g_{\alpha \beta} g^{\mu \nu} g^\ast_{\mu \nu}), \] (5)

All the above quantities are evaluated at the 4D hypersurface \( y = y_0 = \text{Constant} \), which is identified with the physical spacetime \( \Sigma \). Thus 4D quantities depend on \( x^0, x^1, x^2, x^3 \) only, but not on \( y \).

The above equations form the basis of STM. From a four-dimensional point of view, the empty 5D equations look as the Einstein equations with (effective) matter. In the case where \( g_{\mu \nu} \) is independent of \( y \), equations (2) and (3) show that the effective energy-momentum tensor is traceless, \( T_{(\text{eff})\mu} = 0 \). In other words, independence of the 4D metric from the extra coordinate implies a radiation-like equation of state. Thus the existence of other forms of matter crucially depend on the derivatives of \( g_{\mu \nu} \) with respect to the extra dimension. In this case the symmetric tensor (5) is a non-trivial conserved matter quantity (4). A detailed investigation shows that we can recover all of the equations of state commonly used in astrophysics and cosmology [2], [21]-[22].

The right-hand-side of (2) can be expressed in terms of geometrical quantities. For this we introduce the normal vector orthogonal to spacetime \( \Sigma \). It is
\[ n_A = (0, 0, 0, 0, \epsilon \Phi). \] (6)

Then, the first partial derivatives of the metric with respect to \( y \) can be interpreted in terms of the extrinsic curvature of \( \Sigma \). Namely,
\[ K_{\alpha \beta} = \frac{1}{2} \mathcal{L}_a g_{\alpha \beta} = \frac{1}{2\Phi} g^\ast_{\alpha \beta}, \] (7)
and \( K_{A4} = 0 \). We thus have \( K^\lambda_\alpha = (g^\alpha_\beta g^\ast_{\alpha \beta} / 2\Phi), K_{\alpha \beta} K^\alpha_\beta = -(g^\ast_{\mu \nu} g_{\mu \nu} / 4\Phi^2) \) and \( K_{\mu \alpha} K^\alpha_\nu = (g^\ast_{\mu \alpha} g^\ast_{\nu \rho} g^{\rho \alpha} / 4\Phi^2) \). For the second derivatives \( g_{\mu \nu} \), we evaluate (5) \( R_{\mu \nu A4} \). We obtain,
\[ (5) R_{\mu \nu A4} = -\epsilon \Phi \mathcal{F}_{\mu \nu} = -\frac{1}{2} g_{\mu \nu} g^\ast_{\mu \nu} + \frac{1}{2} g_{\mu \nu} \frac{\Phi}{\Phi} + \frac{1}{4} g^\alpha_\rho g_{\mu \rho} g^\ast_{\sigma \nu} . \] (8)

Now substituting this expression into (2) and using (6) we get
\[ (4) G_{\mu \nu} = \epsilon \left[ K^\alpha_\alpha K_{\mu \nu} - K_{\mu \alpha} K^\alpha_\nu + \frac{1}{2} g_{\mu \nu} \left( K_{\alpha \beta} K^{\alpha \beta} - (K^\alpha_\alpha)^2 \right) - E_{\mu \nu} \right], \] (9)
where
\[ E_{\mu \nu} = \frac{(5) R_{\mu \nu A4}}{\Phi^2}. \] (10)
Equation (9) suggests to us that matter may be purely geometrical in origin. This interpretation is the backbone of STM.

Notice that, as a consequence of the field equations, \((5) R_{AB} = 0\), in STM theory the Riemann tensor and the Weyl tensor become identical to each other, viz, \((5) R_{ABCD} = (5) C_{ABCD}\). Consequently, \(E_{\mu\nu} = (5) C_{A\mu B\nu n^A n^B}\). Now, by virtue of (3), \(E_{\mu\nu}\) is traceless, as one expected

\[
E_\alpha^\alpha = 0. 
\]

Let us now consider the tensor quantity (5). In terms of the extrinsic curvature it becomes

\[
P_{\alpha\beta} = K_{\alpha\beta} - g_{\alpha\beta} K^\mu_{\mu}. 
\]

We now decompose it as

\[
P_{\mu\nu} = -\frac{k^2}{2} (\lambda g_{\mu\nu} + T_{\mu\nu}),
\]

where \(k^2\) is a constant introduced for dimensional reasons. Although this decomposition can be ambiguous, we will see that in brane-world models \(\lambda\) is the tension of the brane in five-dimensions, and \(T_{\mu\nu}\) the energy-momentum tensor of the matter in the brane. From the above equations we get

\[
K_{\mu\nu} = -\frac{1}{2} k^2 (\lambda g_{\mu\nu} + T_{\mu\nu}) - \frac{1}{3} g_{\mu\nu} (T - \lambda).
\]

Substituting (14) into (9), we obtain the STM effective energy-momentum tensor as

\[
(4) G_{\mu\nu} = -\Lambda_{(4)} g_{\mu\nu} + 8\pi G_N T_{\mu\nu} + \epsilon k^4 (\Pi_{\mu\nu} - \epsilon E_{\mu\nu}),
\]

where

\[
\Lambda_{(4)} = \epsilon \frac{\lambda^2 k^4}{12},
\]

\[
8\pi G_N = \epsilon \frac{\lambda k^4}{6},
\]

and

\[
\Pi_{\mu\nu} = -\frac{1}{4} T_{\mu\alpha} T_{\nu}^\alpha + \frac{1}{12} T T_{\mu\nu} + \frac{1}{8} g_{\mu\nu} T_{\alpha\beta} T^{\alpha\beta} - \frac{1}{24} g_{\mu\nu} T^2.
\]

From (4) and (13) it follows that

\[
T^\mu_{\nu;\mu} = 0.
\]

Also, \((4) G_{\nu;\mu}^\mu = 0\) implies

\[
E_{\nu;\mu}^\mu = k^4 (\Pi_{\nu;\mu}).
\]

Equations (14), (19) and (20) are equivalent to the original set (2), (3) and (4). We stress the fact that the above equations contain neither reference, nor any particular assumption, specific to the brane-world scenario.
2.2 STM as Generating 5D Space for Brane-World Models

It should be noted that (15), (19) and (20) look exactly as the equations for gravity in brane-world models, with \( \lambda \) and \( T_{\mu\nu} \) as the vacuum energy and energy-momentum tensor, respectively in the brane [16]. Equation (15) includes the five-dimensional corrections to the field equations in 4D general relativity. These are the local quadratic corrections given by \( \Pi_{\mu\nu} \) and non-local Weyl corrections from the free gravitational field in the bulk given by \( E_{\mu\nu} \).

The usual scenario in brane-world models is that matter fields are confined to a singular 3-brane. Therefore, to proceed with our discussion we need to construct such a brane from STM. For convenience the coordinate \( y \) is chosen such that the hypersurface \( \Sigma : y = 0 \) coincides with the brane, which is assumed to be \( \mathbb{Z}_2 \) symmetric in the bulk background [11]-[16]. The brane is obtained by a simple “copy and paste” procedure. Namely, we cut the generating 5D spacetime, with metric \( g_{\alpha\beta} \), in two pieces along \( \Sigma \), then copy the region \( y \geq 0 \) and paste it in the region \( y \leq 0 \). The result is a singular hypersurface in a \( \mathbb{Z}_2 \) symmetric universe with metric

\[
dS^2 = g_{\mu\nu}^{\text{bulk}}(x^\rho, y) dx^\mu dx^\nu + \epsilon dy^2,
\]

where

\[
g_{\alpha\beta}^{\text{bulk}} = g_{\alpha\beta}(x^\mu, +y) \quad \text{for} \quad y \geq 0, \\
g_{\alpha\beta}^{\text{bulk}} = g_{\alpha\beta}(x^\mu, -y) \quad \text{for} \quad y \leq 0.
\]

(22)

Therefore, the field equations in the bulk exhibit a delta-like singularity, viz., \( ^{(5)}G_{\mu\nu}^{\text{bulk}} = k_5^2 [T_{\mu\nu}^{\text{bulk}} + \delta(y) S_{\mu\nu}] \), where \( S_{\mu\nu}n^A = 0 \) represents the total energy-momentum in the brane.

Israel’s boundary conditions imply

\[
K_{\mu\nu}|_{\Sigma^+} - K_{\mu\nu}|_{\Sigma^-} = -k_5^2 \left[ S_{\mu\nu} - \frac{1}{3} g_{\mu\nu} S \right].
\]

(23)

This equation, together with the imposition of the \( \mathbb{Z}_2 \)-symmetry, keeping the brane fixed, leads to

\[
K_{\mu\nu}|_{\Sigma^+} = -K_{\mu\nu}|_{\Sigma^-} = -\frac{1}{2} k_5^2 \left( S_{\mu\nu} - \frac{1}{3} g_{\mu\nu} S \right).
\]

(24)

Then from (12) we get

\[
P_{\mu\nu} = -\frac{1}{2} k_5^2 S_{\mu\nu}.
\]

(25)

Consequently, from (13) and (23) we find

\[
S_{\mu\nu} = -\lambda g_{\mu\nu} + T_{\mu\nu},
\]

(26)
which is the usual separation for the energy-momentum tensor in brane models, where \( \lambda \) and \( T_{\mu\nu} \) are the vacuum energy and the energy-momentum tensor, respectively (from a 5D point of view, \( \lambda \) is the tension of the brane).

Thus the symmetric tensor \( P_{\mu\nu} \) from STM can be interpreted as (proportional to) the total energy-momentum in a \( \mathbb{Z}_2 \) symmetric brane universe. This identification is also suggested by the Hamiltonian treatment of five-dimensional Kaluza-Klein Gravity where \( P_{\mu\nu} \) is the momentum conjugate to the induced metric \( g_{\alpha\beta} \).

With this identification, (15), (19) and (20) become the dynamical equations for gravity in the brane. These are identical to the ones developed in usual brane theory.

Consequently, we conclude that the STM equations can be interpreted as the equations for gravity in a \( \mathbb{Z}_2 \) symmetric universe whose matter content is described by (25), with the usual decomposition (26), plus local and non-local corrections given by \( \Pi_{\mu\nu} \) and \( E_{\mu\nu} \), respectively. In other words, the STM picture forms the generating space for brane-world solutions.

The effective matter content of the spacetime will be the same whether we interpret it as induced matter, as in STM, or as the “total” matter in a \( \mathbb{Z}_2 \) symmetric brane universe. Again, this is a consequence of the identification (25) and the field equation (15).

Since the dynamics of the spacetime is determined by its total matter content, we conclude that these two theories generate the same physics, although STM and brane theories look very different at first sight.

### 2.3 Geometrization of Matter in Brane Models

From (12) and (25) we see that the matter terms arise from the extrinsic curvature of the brane \( \Sigma \). Equation (1) then shows that the dependence of five-dimensional metrics on the extra coordinate leads to usual matter in 4D. In the case where the 5D bulk metric is independent of the extra coordinate, \( S_{\mu\nu} = 0 \) and there is no matter on the brane.

In this case, there is only an effective energy momentum induced on the brane by the Weyl curvature in the bulk projected onto the brane, viz., \( G_{\mu\nu} = -\epsilon E_{\mu\nu} \). The induced matter is called Weyl (or dark) radiation, because it satisfies the “radiation-like” equation of state \( \rho = \sum_{i=1}^{3} p_i \), where \( \rho \) and \( p_i \) are the energy density and principal pressures, respectively. The Davidson and Owen solution [24] is a perfect example of this and has been much discussed in STM theory.

Thus, although it is not mentioned in an explicit way, brane-world models incorporate the concept that matter in 4D can be regarded as the effect of curvature in the extra dimension in a five-dimensional bulk. This is the typical point of view adopted in STM theory [7].
2.4 Different Points of View in STM and Brane Theories

We already mentioned that these two theories have different motivation for the introduction of a large extra dimension. Besides this, the main difference is that authors in STM and brane theory adopt different points of view regarding the matter content and dynamics of our 4D spacetime.

In STM theory the attention is on the geometry of the bulk. The bulk is a solution of the 5D Einstein’s equations in vacuum. The matter content of the spacetime (brane) Σ is regarded as the effect of the curvature in the extra dimension. The geometry of the brane is determined by the 5D line element evaluated at Σ defined as $y = y_0$. As an illustration of this procedure we mention the cosmological solution

$$dS^2 = y^2 dt^2 - t^{2/\alpha} y^{2/(1-\alpha)} [dr^2 + r^2 (d\theta^2 + \sin^2 \theta d\phi^2)] - \alpha^2 (1 - \alpha)^{-2} t^2 dy^2,$$

(27)

where $\alpha$ is a constant. In four-dimensions (on the hypersurface $\Sigma : y = y_0$) this metric corresponds to Friedmann-Robertson-Walker models with flat 3D sections. The equation of state of the effective perfect fluid in 4D is: $p = n \rho$ with $n = (2\alpha/3 - 1)$ ($\alpha = 2$ for radiation, $\alpha = 3/2$ for dust, etc.).

In brane theory the main thrust is the matter content of the brane. When we assume some equation of state we impose conditions on the geometry of the bulk (specifically on $g_{\mu\nu}^s$). The “backreaction” from the bulk shows up in the gravity on the brane through the nonlocal Weyl anisotropic radiation field described by the trace-free tensor $E_{\mu\nu}$ (which contains $g_{\mu\nu}^s$). Therefore, the equations in the brane, for some specific $T_{\mu\nu}$, should be solved together with the equations in the bulk, which are assumed to be the 5D Einstein’s equation with negative cosmological constant.

In short we can say that authors in STM construct 4D from 5D, while authors in brane-world scenarios seek to reconstruct 5D from 4D.

3 Combining The Two Theories

In this Section we approach the question of how can we benefit from the equivalence between these two theories.

In STM theory, the solutions of the field equations contain certain number of arbitrary functions. The interpretation of these solutions in terms of induced matter, as well as the physical properties, crucially depends on the choice of these functions. However, the number of physical restrictions, in the theory, are not in general sufficient to determine all of them.

In brane theory, the system of equations (15), (19) and (20) is not in general closed, since (20) does not determine $E_{\mu\nu}$, in general. If $E_{\mu\nu} = 0$, then the field
equations on the brane form a closed system. However, there is no guarantee that the resulting brane metric can be embedded in a regular bulk. Simply there is no enough information available in the brane as to reconstruct the bulk [19].

Here we propose an alternative method to alleviate the task of finding solutions in the brane. Our proposal combines both theories. The idea is to capitalize from the rich physics in brane models and the freedom in STM. Specifically, we can impose the physics on the brane to restrict the freedom characteristic of many solutions in STM [25] - [26].

3.1 An Exact Solution in STM

As an illustration of our proposal, let us consider the STM solution [25]

\[ dS^2 = B^2(t, y)dt^2 - A^2(t, y) \left[ \frac{dr^2}{1 - kr^2} + r^2 d\Omega^2 \right] - dy^2, \]  

with

\[ B = \frac{1}{\mu(t)} \frac{\partial A}{\partial t}, \]  

and

\[ A^2 = [\mu^2(t) + k]y^2 + 2\nu y + \frac{\nu^2(t) + K}{\mu^2(t) + k}. \]  

Here the extra dimension is spacelike, \( \epsilon = -1 \). This is an exact cosmological solution, with curvature index \( k = -1, 0, +1 \), which contains two arbitrary functions \( \mu(t) \) and \( \nu(t) \). The constant \( K \) is related to the Kretschmann scalar, namely,

\[ I = R_{ABCD}R^{ABCD} = \frac{72K^2}{A^8}. \]  

The effective total energy density and isotropic pressure are

\[ 8\pi G_N \rho_{\text{eff}} = {}^{(4)}G_0^0 = \frac{3(\mu^2 + k)}{A^2}, \]  

\[ 8\pi G_N p_{\text{eff}} = -{}^{(4)}G_1^1 = -{}^{(4)}G_2^2 = -{}^{(4)}G_3^3 = -\frac{2\mu \dot{\mu}}{AA} - \frac{\mu^2 + k}{A^2}. \]  

This solution has been studied in detail [25] for various choices of \( \mu(t) \) and \( \nu(t) \) and imposing the equation of state \( p_{\text{eff}} = n\rho_{\text{eff}} \), with \( n = \text{Constant} \). Certainly, this physical assumption decreases the number of unknown functions from two to one. But still there is a lot of freedom.
3.2 An Exact Solution on The Brane

Here we will use the brane-world paradigm to determine the unknown functions. The usual assumption is that normal matter, except for gravity, “lives” only on the $\Sigma : y = 0$ brane $[13]-[19]$. The appropriate $\mathbb{Z}_2$ symmetric brane universe, from the generating space $(28)-(30)$, is obtained following $(22)$.

The junction conditions across the brane imply that the metric is continuous. Specifically, on $\Sigma$

$$ds^2 = dt^2 - a^2(t) \left[ \frac{dr^2}{1 - kr^2} + r^2 d\Omega^2 \right], \quad (34)$$

where

$$A^2(t, 0) = a^2(t) = \frac{\nu^2 + K}{\mu^2 + k}, \quad (35)$$

$$B(t, 0) = \left( \frac{\dot{A}}{\mu} \right)_{\Sigma} = \frac{\dot{a}}{\mu} = 1. \quad (36)$$

Thus the continuity of the metric defines $\mu$ and $\nu$ through $a$,

$$\mu = \dot{a}$$

$$\nu^2 = a^2(\dot{a}^2 + k) - K. \quad (37)$$

We now use the identification of $P_{\mu\nu}$ with the energy-momentum tensor on the brane. From equations $(5)$, $(25)$ and $(26)$ we get

$$-\lambda + \rho = 3\alpha \left( \frac{\dot{A}}{A} \right)_{\Sigma} = 3\frac{\nu\alpha}{a^2}, \quad (38)$$

$$-\lambda - p = \alpha \left( \frac{\dot{B}}{B} + 2 \frac{\dot{A}}{A} \right)_{\Sigma} = \frac{\nu\alpha}{\dot{a}a} + \frac{\nu\alpha}{a^2}, \quad (39)$$

where $\alpha = 2/k^2_{(5)}$. Here we have assumed that the energy-momentum tensor $T_{\mu\nu}$ in the brane is a perfect fluid with density $\rho$ and isotropic pressure $p$. This fluid is “at rest” in the frame of $(34)$ because $(4)G_{0j} = 0$. In other words, the system of reference in $(34)$ is comoving with the matter.

As in brane models we impose some physics on the matter quantities $\rho$ and $p$ (not on the effective ones). We will adopt the usual equation of state in cosmology

$$p = n\rho, \quad (40)$$
which for \( n = 1, 3, 0 \) gives the stiff, radiation and dust equations of state. Substituting into (38) and (39) we get

\[
\frac{d\nu}{da} + (3n + 1)\frac{\nu}{a} = -\frac{\lambda(n + 1)}{a}\alpha a. \tag{41}
\]

From which we find

\[
\nu = -\frac{\lambda}{3\alpha} a^2 + \frac{C}{a^{(1+3n)}}, \tag{42}
\]

where \( C \) is a constant of integration. The matter energy density becomes

\[
\rho = \frac{3\alpha C}{a^{3(1+n)}}. \tag{43}
\]

The evolution equation for \( a \) is obtained from (37) and (42) as

\[
\dot{a}^2 = -k + \frac{K}{a^2} + \frac{\lambda^2}{9\alpha^2} a^2 + \frac{C^2}{a^{2(2+3n)}} - \frac{2C\lambda}{3\alpha a^{(1+3n)}}, \tag{44}
\]

For a perfect fluid, \( T_{\mu\nu} = (\rho + p)u_\mu u_\nu - pg_{\mu\nu}, \) the quadratic corrections in the brane (18) become

\[
\Pi_{\mu\nu} = -\frac{\rho^2}{12} u_\mu u_\nu - \frac{1}{12} \rho(\rho + 2p) h_{\mu\nu}, \tag{45}
\]

where we have introduced the projector \( h_{\mu\nu} = u_\mu u_\nu - g_{\mu\nu}. \) Thus, for perfect fluid \( \Pi_0 = -\rho^2/12. \) Therefore, the quadratic contribution to the total energy density is always positive, which is physically reasonable (we remind that \( \epsilon = -1 \)) here. The quadratic correction to the pressure is also positive and proportional to \( (2n+1)\rho^2/12. \)

The black or Weyl radiation coming from the bulk is described by \( E_{\mu\nu}. \) It does not depend on the equation of state in the brane. The corresponding energy density is

\[
8\pi G_N \rho_{Weyl} = -\epsilon E_0^0 = -\frac{**}{B} = \frac{3K}{a^4}, \tag{46}
\]

where the factor \( 8\pi G_N \) is introduced for dimensional reasons. The radiation pressure is isotropic, viz., \( 8\pi G_N p_{Weyl} = -E_1^1 = -E_2^2 = -E_3^3 = K/a^4, \) and \( E_0^0 = 0, \) as expected.

Collecting results, we find the effective energy density as follows

\[
8\pi G_N \rho_{eff} = \Lambda + 8\pi G_N \rho + \frac{k_{(5)}^4}{12} \rho^2 + 8\pi G_N \rho_{Weyl}, \tag{47}
\]

where \( \Lambda = \lambda^2 k_{(5)}^2/12; \) and the densities \( \rho, \rho_{Weyl} \) are given by (13) and (16), respectively. We should remark again that the effective matter content is the same whether calculated from STM equations or from the \( \mathbb{Z}_2 \)-symmetric brane perspective. This is
verified by our example, viz., the STM equation (32) and the brane-world equation (47) yield the same expression for $\rho_{\text{eff}}$, in terms of $a$. We also note that (47) is a general expression valid for any perfect fluid and not only for this model. Another important point to notice here is that the Weyl contribution to the effective energy density does not have to be positive. It can be positive, negative or zero [27].

For very high energies in the brane, when the quadratic term dominates over the other terms, the effective equation of state is stiffened, viz.,

$$p_{\text{eff}} \approx (2n + 1)\rho_{\text{eff}}.$$  \hspace{1cm} (48)

This is a general feature of brane models [16]. It indicates that quadratic corrections might have a dramatic influence in the early universe and during late-stages of the gravitational collapse. Even in the case of “non-gravitating” matter, for which $(\rho + 3p) = 0$, the effective matter is radiation-like, viz., $p_{\text{eff}} \approx \rho_{\text{eff}}/3$. For dust $p_{\text{eff}} \approx \rho_{\text{eff}}$. This feature presents a number of intriguing possibilities that is worth to study in detail.

The above equations represent a cosmological model where the brane is appropriately matched to the bulk metric and the matter in the brane satisfies physical conditions. This model has a number of interesting properties but we leave their discussion to another opportunity.

This model is an example that clearly illustrates the feasibility of our proposal here. Namely, that a combination of the two approaches, STM and brane theory, can be very helpful to find solutions on the brane. This is a working alternative to the direct approach in brane theory of solving the equations from “scratch”.

4 Summary and Conclusions

We have presented here a new interpretation of space-time-matter theory. In this interpretation the equations of STM are identical to those in $\mathbb{Z}_2$-symmetric brane-world models. This is a consequence of the connection between the symmetric tensor $P_{\alpha\beta}$ from STM and the energy-momentum tensor $S_{\mu\nu}$ in the brane. This, in turn is a result of the $\mathbb{Z}_2$ symmetry used in brane-world theory. Therefore, STM incorporates all the local and non-local energy corrections to GR in $4D$, typical of brane-world scenarios.

In both theories the presence of matter crucially relies on the assumption that the metric is a function of the extra variable. For metrics with no dependence on $y$ there is no matter, just Weyl radiation with anisotropic pressures. In this sense, matter in $4D$ is a consequence of the extra dimension. Thus, brane models incorporate the concept that matter can be regarded as effect of the geometry in $5D$, which is the trademark of STM theory.

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Going a little aside, we mention that the motion of test particles presents similar characteristics in both theories. This has been studied in a number of papers [28]-[30]. However, this similarity has nothing to do with the dynamics of the field, but it is a result of the assumption that test particles move along geodesics in both theories.

From a theoretical point of view it is important that we understand the connection between theories that seek to explain the same subject. STM and brane models represent two opposite approaches to the same problem. In STM the physics in 4D is constructed from the bulk. In brane models we should find the solution to the equations in 4D, which has to be matched with the bulk geometry satisfying the appropriate boundary conditions. In other words, the goal in brane theory is to use physical information in 4D to reconstruct the generating bulk.

From a practical point of view, a combined approach of both theories seems to be promising. The incorporation of the rich physics of branes to STM may allow us to obtain interesting physical models. The solution discussed in Section 3 is a clear example of this. It neatly shows how the physics on the brane unambiguously determines the arbitrary functions. It also serves as an example of our main point here. Namely, that effective matter quantities do not depend on whether we calculate them using the STM or brane-world paradigm; only the interpretation is different in both theories. This example-solution presents good physical properties, but their discussion is beyond the scope of the present work.

The main difference between the two models resides in the motivation they have for the introduction of a large extra dimension. Besides the bulk geometry in STM satisfies \(^{(5)}G_{AB} = 0\), while in brane theory \(^{(5)}G_{AB} = -\Lambda_{(5)}\gamma_{AB}\). For this reason there is a “missing” term in the definition of \(\Lambda_{(4)}\) in (16). The introduction of this term presents no problem and one would get

\[
\Lambda_{(4)} = \frac{1}{2} k_{(5)}^2 \left( \Lambda_{(5)} + \epsilon \frac{k_{(5)}^2 \lambda^2}{6} \right).
\]  

(49)

Despite of these differences, we conclude that STM and \(\mathbb{Z}_2\)-symmetric brane-world theories are equivalent to each other. They predict the same corrections to general relativity in 4D, and incorporate the concept that matter can be regarded as the effect from an extra dimension.

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