Phase transitions, like the boiling and evaporating of water at a certain temperature and pressure, are common phenomena both in everyday life and in almost any branch of physics. Loosely speaking, a phase transition brings about a sudden change of the macroscopic properties of a system while smoothly varying a parameter (the temperature or the pressure in the above example). The mathematical description of phase transitions is conventionally based either on Gibbs measures on phase space or on (grand)canonical thermodynamic functions, relating their loss of analyticity to the occurrence of a phase transition. Such a nonanalytic behavior can occur only in the thermodynamic limit \[ \lim_{N \to \infty} \frac{1}{N} \sum_{v} \text{energy} \leq \text{critical energy} \], in which the number of degrees of freedom \( N \) of the system goes to infinity. Conceptually, the necessity of the thermodynamic limit is an objectionable feature: firstly, the number of degrees of freedom in real systems, although possibly large, is finite, and, secondly, for systems with long-range interactions, the thermodynamic limit may even not be well defined.

Recently, an alternative approach to phase transitions has been proposed, which connects the occurrence of a phase transition to certain properties of the potential energy \( V \), resorting to topological concepts. The conceptual advantages of this topological approach are twofold: The microscopic Hamiltonian dynamics, which is at the basis of the thermodynamic behavior of the system, can be linked via the Lyapunov exponents to the topological structure considered \[ \text{topological invariants} \], thus rendering the topological approach a very fundamental one. Furthermore, in contrast to the conventional approach, there exists a “natural” generalization of the concept of phase transitions to finite systems.

The topological approach is based on the hypothesis \[ \text{topological invariants} \] that phase transitions are related to topology changes of submanifolds \( M_v \) of the configuration space of the system, where the \( M_v \) consist of all points \( q \) of the configuration space for which \( V(q)/N \leq v \), i.e., their potential energy per degree of freedom is equal to or below a certain level \( v \). This hypothesis has been corroborated by numerical as well as exact analytical results for some model systems showing first-order as well as second-order phase transitions. A major achievement in the field is the recent proof of a theorem, stating, loosely speaking, that, for systems described by smooth, finite-range, and confining potentials, a topology change of the submanifolds \( M_v \) is a necessary criterion for a phase transition to take place \[ \text{topological invariants} \].

Albeit necessary, such a topology change is clearly not sufficient to entail a phase transition. This follows for example from the analytical computation of topological invariants in the XY model \[ \text{topological invariants} \], where the number of topology changes occurring is shown to be of order \( N \), but only a single phase transition takes place. So topology changes appear to be rather common, and only particular ones are related to phase transitions. While some ideas relating the “strength” of a topology change to the occurrence of a phase transition have been put forward \[ \text{topological invariants} \], a sufficient criterion on the quality of a topology change is still lacking, and the quest to seek for one can be considered as the fundamental open problem in the field.

The objective of the present Letter is to shed some light on how such a sufficient criterion might (not) look like. To this purpose, a so-called solid-on-solid model proposed by Burkhardt \[ \text{topological invariants} \] is considered. This model has the—for our purposes desirable—features to be (i) analytically solvable, and (ii) sensitive upon a slight modification of the model, in the sense that in one case it does exhibit a phase transition, whereas in the other case it does not. By investigating analytically the topological properties of the submanifolds \( M_v \) of the configuration space, we find that, in both these cases, a topology change takes place.

Unattainability of a purely topological criterion for the existence of a phase transition for non-confining potentials

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The relation between thermodynamic phase transitions in classical systems and topology changes in their configuration space is discussed for a one-dimensional, analytically tractable solid-on-solid model. The topology of a certain family of submanifolds of configuration space is investigated, corroborating the hypothesis that, in general, a change of the topology within this family is a necessary condition in order to observe a phase transition. Considering two slightly differing versions of this solid-on-solid model, one showing a phase transition in the thermodynamic limit, the other not, we find that the difference in the “quality” or “strength” of this topology change appears to be insignificant. This example indicates the unattainability of a condition of exclusively topological nature which is sufficient as to guarantee the occurrence of a phase transition in systems with non-confining potentials.

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In the case of the model exhibiting a phase transition, this transition is related to the topology change in accordance with the above mentioned topological hypothesis. Comparing, however, this topology change to the one in the model without a phase transition, a significant difference in the “strength” of the topology does not appear to be present. Hence, the results presented in this Letter indicate that a discrimination between topology changes which entail a phase transition, and those which don’t, may in general not be possible on topological grounds exclusively. This result puts the search for a sufficiency criterion on the topology change into a completely new perspective.

Burkhardt model: We consider a one-dimensional model on a lattice with real, continuously varying variables $q_i$. The Hamiltonian function is given by

$$H(q) = \sum_{i=1}^{N} \left[ |q_{i+1} - q_i| + U(q_i) \right], \quad (1)$$

where $q = (q_1, ..., q_N)$ is a state of the system. Periodic boundary conditions $q_{N+1} \equiv q_1$ are assumed. The so-called pinning potential $U$ is a real valued function, bounded below and above, with a unique infimum at zero. This one-dimensional system was introduced in [10] to model the localization-delocalization transition of an interface in a two-dimensional system. The pinning potential tends to localize the “interface” (i.e., the values of the $q_i$) around zero. The above Hamiltonian describes a static model without a kinetic term, but it can be extended straightforwardly to include the dynamics as well.

The thermodynamic behavior of this system can be analyzed analytically by rewriting the partition function in terms of an integral transfer operator. Then, the eigenvalue equation of this operator can be transformed into a one-dimensional Schrödinger type equation. In doing so, the problem of finding a localization-delocalization transition is mapped onto the question whether there exist bound state solutions of the Schrödinger equation for certain potentials. For the example of a square well pinning potential

$$U(x) = \begin{cases} -1 & \text{for } |x| \leq 1, \\ 0 & \text{for } |x| > 1, \end{cases} \quad (2)$$

the latter problem is analyzed explicitly in [11]. It is found that the existence of a phase transition depends on the domain of the Hamiltonian function [11]:

(a) For the $q_i$ taking on values on the semi-infinite line, $q_i \in [0, +\infty) = \mathbb{R}_0^+$, a second-order localization-delocalization transition is observed.

(b) In case of the $q_i$ having values from the real numbers, $q_i \in (-\infty, +\infty) = \mathbb{R}$, no transition takes place.

It is argued in [10] that this result generalizes to a large class of pinning potentials $U$. Note that the existence of a phase transition in such one-dimensional systems is by no means contradictory to van Hove’s theorem: the conditions assumed in van Hove’s work [11] simply don’t apply to the Burkhardt model [10], neither do they to many other models. For a more general theorem and further interesting aspects of phase transitions in one-dimensional systems see [12].

Topological approach: Certain submanifolds $M_v$ of the configuration space are taken as a starting point for the topological approach. Due to the absence of a kinetic term in [11], the configuration space $\Gamma$ is identical to the domain of the Hamiltonian. For a system consisting of $N$ degrees of freedom, we have $\Gamma_a = (\mathbb{R}_0^N)^N$ and $\Gamma_b = \mathbb{R}^N$ for the above cases (a) and (b), respectively. We define the submanifolds

$$M_v^{a,b} = \left\{ q \in \Gamma_{a,b} \left| \frac{H(q)}{N} \leq v \right. \right\} \quad (3)$$

for the cases (a) and (b) as the subsets consisting of all points $q$ from configuration space with potential energies—given, in our case, by the value of the Hamiltonian $H(q)$—equal to or below a certain level $v$. As $\Gamma_a \subset \Gamma_b$, the relation

$$M_v^a = M_v^b \cap \Gamma_a \quad (4)$$

holds which allows us to consider case (b) first and, at the end, infer the results for (a) by simple reasoning. Now our aim is to investigate and characterize the topology of the $M_v^b$ and to determine the critical level(s) $v_c$ at which the topology changes occur. Loosely speaking, two manifolds are said to be topologically equivalent if they can be mapped onto each other by a smooth deformation, i.e., by stretching and bending, but not cutting or tearing. If the manifolds are not topologically equivalent, we say that a topology change takes place. In the following, we will explicitly characterize the topology and its changes for the Burkhardt model with pinning potentials $U(x)$, bounded below and above, with a unique infimum at $x = 0$, which decrease (increase) monotonously for negative (positive) $x$. Without loss of generality we choose $U$ such that its supremum $\sup_q U(x) = 0$.

In previous analytical calculations of topology changes in configuration space [2, 5], Morse theory was employed in order to calculate topological invariants. Within the standard setting of this theory as put forward originally by Morse [13], compact manifolds are considered. The $M_v^{a,b}$ as defined in [3] are not necessarily compact, and we would have to resort to more sophisticated extensions of Morse theory [14]. The simplicity of one-dimensional models, however, may allow for a more direct determination of topological properties [3]. To this purpose, it helps the intuition to plot the submanifolds $M_v^b$ for the simplest non-trivial case of $N = 2$ degrees of freedom.
Figure 1 illustrates for an example of a smooth pinning potential satisfying the above boundedness and monotonicity conditions that, for small values of $v$, the $M_v^b$ are topologically equivalent to the square $\mathbb{I}^2$ (where $\mathbb{I} = [0, 1]$ denotes the unit interval), whereas for $v$ above a certain critical level, topological equivalence to an infinite stripe, $\mathbb{R} \times \mathbb{I}$, is observed. In the following we will show that, in the general case of $N$ degrees of freedom,

$$
M_v^b \sim \begin{cases} 
\emptyset & v < U_{\text{inf}}, \\
\mathbb{I}^N & \text{for } U_{\text{inf}} < v < 0, \\
\mathbb{R} \times \mathbb{I}^{N-1} & 0 < v,
\end{cases}
$$

(5)

where $\sim$ denotes topological equivalence, $\emptyset$ is the empty set, and $U_{\text{inf}} = \inf E_v U(x)$ is the infimum of $U$.

The simple proof goes in two steps: First, $M_v^b$ is shown to be a star convex subset of $\mathbb{R}^N$, i.e., there exists a $\tilde{q} \in M_v^b$ such that the line segment from $\tilde{q}$ to any point in $M_v^b$ is contained in $M_v^b$. This can be proved by showing that, for every state $q = (q_1, ..., q_N) \in \mathbb{R}^N$ and every $\lambda \in [0, 1)$, the energy of $\tilde{q}' = \lambda q$ is smaller than or equal to the energy of $q$, i.e., $H(\tilde{q}') \leq H(q)$. The star convexity of $M_v^b$ implies homotopical equivalence to $\mathbb{I}^N$ (or to an $N$-ball $\mathbb{B}^N$), but not necessarily topological equivalence.

In a second step, the closedness of $M_v^b$ is investigated. This is done, analogously to the treatment in [3], by studying the asymptotic behavior of the Hamiltonian $H(\lambda q)$ in the limit $\lambda \rightarrow \infty$. Depending on the state $q$ considered, we find

$$
\lim_{\lambda \rightarrow \infty} H(\lambda q) = \begin{cases} 
0 & \text{if } q_i = q_j \forall i, j, \\
\infty & \text{else}.
\end{cases}
$$

(6)

Hence, for negative energies, only states $q \in \mathbb{R}^N$ with finite (Euclidean) norm $||q||$ are accessible, whereas states of arbitrarily large norm can be attained in the case of positive energies. From this observation and definition [3] it can be inferred that $M_v^b$ is a bounded and closed subset of $\mathbb{R}^N$ for $v < 0$, and, together with the star convexity shown above, it follows that $M_v^b$ is topologically equivalent to $\mathbb{I}^N$. For $v > 0$, however, $M_v^b$ is unbounded and not closed. Since, for finite positive energies, states of arbitrarily large norm can be attained only “in a single spatial direction”, i.e., in the vicinity of the (hyper)space diagonal $q = \lambda(1, ..., 1)$, $\lambda \in \mathbb{R}$, we conclude that $M_v^b$ is topologically equivalent to the product of an open interval and $N - 1$ closed ones, $M_v^b \sim \mathbb{R} \times \mathbb{I}^{N-1}$. With the immediate observation that $M_v^b = \emptyset$ for $v < U_{\text{inf}}$, we have accomplished a complete characterization of the topology of $M_v^b$ as summarized in (5).

Relation (4) allows to transfer this result for the case (b) of configuration space $\Gamma_b = \mathbb{R}^N$ straightforwardly to case (a) with $\Gamma_a = (\mathbb{R}^+_0)^N$. In the case of $N = 2$ degrees of freedom this transfer simply consists in considering the positive quadrant in figure 1 only, giving topological equivalence of $M_v^a$ to $\mathbb{I}^2$ for small values of $v$ and to $\mathbb{R}^+_0 \times \mathbb{I}$ for values above a critical level. For the general case of $N$ degrees of freedom, this results in a modification of (5), being of the form

$$
M_v^a \sim \begin{cases} 
\emptyset & v < U_{\text{inf}}, \\
\mathbb{I}^N & \text{for } U_{\text{inf}} < v < 0, \\
\mathbb{R}^+_0 \times \mathbb{I}^{N-1} & 0 < v.
\end{cases}
$$

(7)

Releasing the monotonicity condition on the pinning potential $U$, additional topology changes should occur, but the essential one, being related to the localization-delocalization transition, is expected to persist unaltered.

**Discussion of the results:** Having fully characterized the topology of the submanifolds $M_{v,a,b}$ for the Burkhardt model (1), we found that, for our cases (a) and (b), the respective topology changes are very similar, although not identical, in nature. In both cases, these submanifolds are topologically equivalent to a closed $N$-ball below the transition energy (or temperature). It is only above the transition energy that the topology differs slightly: for case (a) with configuration space $\Gamma_a = (\mathbb{R}^+_0)^N$, we find equivalence to the product of $N - 1$ closed intervals and a half-open interval. For case (b) with configuration space $\Gamma_a = \mathbb{R}^N$, equivalence to the product of $N - 1$ closed intervals and an open interval is obtained. These topology changes are clearly not identical, but neither do they
show a striking difference in “strength” like the ones observed for XY models with and without phase transitions \[7\].

We do not have a satisfactory definition at hand for the strength of a topology change. A criterion based on the variation of the Betti numbers as proposed in \[8\] is not general enough, as the topology change reported above does not include a change of homotopy, and the Betti numbers remain unchanged. One might, as done in \[8\], resort to the \emph{surfaces} of the submanifolds \(M_v\) instead of the manifolds itselfs in order to obtain a change in the Betti number, but neither with this trick a significant difference in the strength of the topology change will be detected. Intuitively, the change in case (b) towards a product space including an open interval appears to be even \emph{stronger} than the change where only a half-open interval is involved; it is the latter case, however, which corresponds to a phase transition in the thermodynamic limit.

Our primary intention was to shed some light on the question how a sufficient criterion on the topology change, guaranteeing the existence of a phase transition, might look like. The above considerations suggest that a one-to-one correspondence between the occurrence of a phase transition and the strength of the underlying topology change cannot be established for the Burkhardt model. Hence the possibility to develop a sufficient condition, based exclusively on topological quantities, guaranteeing the existence of a phase transition, appears to be disproved by means of a counterexample for systems with non-confining potentials. Such a condition has, however, been proposed for a certain class of systems with smooth \emph{confining} potentials \[16\], and we suspect that “non-confining” is the crucial property rendering an exclusively topological condition unattainable. Note that the analysis of the thermodynamic as well as of the topological properties of the Burkhardt model presented in this Letter can be straightforwardly extended to a class of systems with smooth potentials. Then, for a suitable choice of the on-site potentials, identical results as in \[7\] and \[8\] can be obtained for systems with and without a phase transition, respectively. In this way it can be shown that smoothness is \emph{not} the crucial property for the attainability of a sufficient condition. Future work should try to shed some light on the question which ingredient, in addition to topological properties, has to be taken into account in order to specify a sufficient condition for the existence of a phase transition in systems with non-confining potentials.

In the introductory part of this Letter, the “natural” generalization of the concept of phase transitions to finite systems within the topological approach was mentioned. Observing that the topology change, which in the thermodynamic limit is related to the localization-delocalization transition of the system, is present for any finite number \(N \geq 2\) of degrees of freedom, one would simply define a transition-like phenomenon in finite systems on the basis of this topology change and identify the critical level of the Hamiltonian with the transition energy. Although this might appear reasonable, it is clearly an unsatisfactory definition due to the lack of a sufficiency condition on the topology change.

Summary: The relation between thermodynamic phase transitions in classical systems and topology changes in their configuration space is discussed for the Burkhardt model, a one-dimensional solid-on-solid model. A complete characterization of the topology of the submanifolds \(M_v\) of the configuration space is accomplished. The hypothesis—proved in \[9\] for a certain class of systems with confining potentials—that in general a change of the topology within the family \(M_v\) is a necessary condition in order to observe a phase transition, is corroborated for a larger class of systems on the basis of this example with non-confining potential. Considering two slightly differing versions of this solid-on-solid model, one showing a phase transition in the thermodynamic limit, the other not, we find that the difference in the “quality” or “strength” of this topology change appears to be insignificant. This example indicates the unattainability of a condition of exclusively topological nature which is sufficient as to guarantee the occurrence of a phase transition in systems with non-confining potentials.

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