NAVIER–STOKES–FOURIER FLUIDS INTERACTING WITH ELASTIC SHELLS

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Abstract. We study the motion of a compressible heat-conducting fluid in three dimensions interacting with a nonlinear flexible shell. The fluid is described by the full Navier–Stokes–Fourier system. The shell constitutes an unknown part of the boundary of the physical domain of the fluid and is changing in time. The solid is described as an elastic non-linear shell of Koiter type; in particular it possesses a non-convex elastic energy. We show the existence of a weak solution to the corresponding system of PDEs which exists until the moving boundary approaches a self-intersection or the non-linear elastic energy of the shell degenerates. It is achieved by compactness results (in highest order spaces) for the solid-deformation and fluid-density. Our solutions comply with the first and second law of thermodynamics: the total energy is preserved and the entropy balance is understood as a variational inequality.

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1. Introduction

The interactions of fluids with elastic structures are important for many applications ranging from hydro- and aero-elasticity [16] over bio-mechanics [4] to hydrodynamics [11]. Motivated by these applications and...
the scientific foundations from engineers and physicists also mathematicians became interested in the field. Nowadays there exists a vast body of literature on incompressible fluid structure interaction, where a part of the boundary of the underlying domain is the mid-section of a flexible shell.

The mathematical analysis of continuum mechanical models in fluid mechanics reaches back to the pioneering work of Leray on the existence of weak solutions for the incompressible Navier–Stokes equations [29]. Based on this, various fluid-structure interaction results have been achieved already; we will explain this in more detail below. A similar foundational work in the compressible case is due to Lions [33] with important extensions by Feireisl et al. [22] [14]. Compressible fluids are important for applications in aero-dynamics and mathematical results on their interactions with elastic structures appeared in this context recently in [6] [13]. A next natural step is to study the thermodynamics of fluid structure interactions. In fact, the assumption that a physical process is isentropic can only be valid for a very short period of time. In general it is indispensable to take into account the transfer of heat. Similarly, the linearisation of the shell model, often applied in the mathematical literature, looses its validity as soon as the displacement of the boundary is not on a small scale any more.

The treatment of non-linear shell models in the context of weak solutions is very recent [39] and (up to date) only available for incompressible fluids. In this work we progress on the theory of weak solutions by showing the existence for systems that take into account 1) heat conduction and compression effects for the fluid and 2) a non-linear elastic respond for the solid. More specifically, we use the classical model by Koiter to describe the shell movement which yields a fully nonlinear fourth order hyperbolic equation with a non-convex energy. The main result of this paper is the existence of a global-in-time weak solution to the Navier–Stokes–Fourier system coupled to the motion of a solid shell of Koiter type. This means that a fourth order PDE for the solid is coupled (via the geometry) to a viscous fluid. A special feature of the Navier–Stokes–Fourier system is that even weak solutions can satisfy an energy equality. We produce a respective equality for the energy of the coupled fluid-structure interaction; this includes the full Koiter energy of the solid deformation. In this context it is noteworthy that we consider perfect elastic shells. This means that no heat is produced by the solid, or reversely entropy is only increased via the fluid. Still some viscous effects can be shown to hold for the elastic solid due to the tight coupling between the solid and the fluid. It is this key observation (and the respective estimate in Subsection 5.2) that allows to show that the elastic part of the energy has the necessary compactness in order to prove that the system is indeed closed (energy is preserved). We note that the interval of existence for our weak solutions could be arbitrarily large. In fact, the time of existence is only restricted once either the topology of the fluid domain changes, namely if a self-intersection of the variable boundary (of the elastic shell) is approached, or if the solid energy reaches a point of degeneracy.

1.1. State of art. Incompressible viscous fluids interacting with lower-dimensional linear elastodynamic equations were studied, for instance, in [12] [23] [31] [37] [39] [42]. All but the last result are concerned with the existence of weak solutions which exist as long as the moving part of the structure does not touch the fixed part of the fluid boundary. The analysis in [12] is concerned with a three-dimensional viscous incompressible fluid modelled by the Navier–Stokes equations, which is interacting with a flexible elastic plate located on one part of the fluid boundary. The shell equations is linearised and the shell is assumed to be one-dimensional. The existence of a weak solution to the incompressible Navier–Stokes equation coupled with a plate in flexion was constructed in [23]. The authors in [31] studied the interaction of an incompressible fluid which interacts with a linear elastic shell of Koiter-type. Here, the middle surface of the shell serves as the mathematical boundary of the three-dimensional fluid domain. In [37] the incompressible Navier–Stokes equations are studied in a cylindrical wall which is moving in time. Its elastodynamics is modelled by the one-dimensional cylindrical linearised Koiter shell model. The authors apply a numerical approach and show the existence of weak solution based on semi-discrete operator splitting scheme. The elastodynamics of the cylinder wall in [38] is governed by the one-dimensional linear wave equation modelling the thin structural layer, and by the two-dimension equations of linear elasticity modelling the thick structural layer. In [12] the analysis of uniqueness properties of weak-solution has been initiated in this field. There the authors show a weak-strong uniqueness result for elastic plates interacting with the incompressible Navier–Stokes equations. As far as we know, the only result on the analysis of weak solutions to fluid-structure interaction, where the original Koiter model (to be described below in Section 1.2) with a leading order nonlinear shell energy is considered, is the recent paper [39] by Muha and the second author. Results regarding the short-time existence of strong solutions can be found in [13].

There are much less results concerning the compressible case. In [6] the authors of the present paper showed the existence of a weak solution to the compressible Navier–Stokes equations coupled with a linear elastic shell...
of Koiter type. Eventually, a similar result has been shown by a time-stepping method \[43\], where the interaction of a compressible fluid with a thermoelastic plate is studied (compare also with the numeric results from \[11\]). Results on the short-time existence of strong solutions for compressible fluid models coupled with one-dimensional linear elastic structures can be found in \[34, 36\]. In \[3\] the author studies an elastic structure (with a regularised elasticity law) which is immersed into a compressible fluid and proves the existence of weak solutions to the underlying system. Results concerning the long-time existence of weak solutions about structure interactions with heat conducting fluids are missing so far - even in the incompressible case. The existence of a unique local-in-time strong solution to compressible Navier–Stokes–Fourier system coupled with a damped linear plate equation has been established very recently in \[35\].

1.2. The model. We consider the full Navier–Stokes–Fourier system of a heat-conducting compressible fluid interacting with a nonlinear elastic Koiter shell in \(\mathbb{R}^3\) of thickness \(2\varepsilon_0 > 0\) (see \[25, 29\] and also \[7, 8\]). Here, \(\omega \subset \mathbb{R}^2\) can be associated to the middle surface of the shell and for simplicity, we take \(\omega = \mathbb{R}^2 \setminus \mathbb{Z}^2\) to be the flat torus. Following \[10\] (see also \[39\] and \[5\]) we suppose that \(\partial \Omega\) can be parametrised by an injective mapping \(\varphi \in C^4(\omega; \mathbb{R}^3)\) such that for all points \(y = (y_1, y_2) \in \omega\), the pair of vectors \(\partial_i \varphi(y), i = 1, 2\), are linearly independent. Simply put, \(\varphi\) is an injective map on the mid-section of the shell of the domain \(\Omega\).

This vector pair \([\partial_1 \varphi(y), \partial_2 \varphi(y)]\) is the covariant basis of the tangent plane to the middle surface \(\varphi(\omega)\) of the reference configuration at each point \(\varphi(y)\) and

\[
\nu(y) = \frac{\partial_1 \varphi(y) \times \partial_2 \varphi(y)}{|\partial_1 \varphi(y) \times \partial_2 \varphi(y)|}
\]

is a well-defined unit vector normal to the surface \(\varphi(\omega)\) at \(\varphi(y)\). We now assume that the shell (and in particular, its middle surface) only deforms along the normal direction with a displacement field \(\eta \nu : \omega \to \mathbb{R}^3\) where \(\eta : \omega \to \mathbb{R}\) is considerably smooth. Then, we can parametrized the deformed boundary by the coordinates

\[
\varphi_\eta(y) = \varphi(y) + \eta(y) \nu(y), \quad y \in \omega,
\]

yielding the deformed middle surface \(\varphi_\eta(\omega)\). The covariant components of the “modified” change of metric tensor \(G(\eta)\) are given by

\[
G_{ij}(\eta) = \partial_i \varphi_\eta \cdot \partial_j \varphi_\eta - \partial_i \varphi \cdot \partial_j \varphi,
\]

where \(\partial_i \varphi_\eta \cdot \partial_j \varphi_\eta\) are the covariant components of the first fundamental form of the deformed middle surface \(\varphi_\eta(\omega)\). We denote by \(\nu_\eta\) the normal-direction to the deformed middle surface \(\varphi_\eta(\omega)\) at the point \(\varphi_\eta(y)\) (which is in general not a unit vector). It is given by

\[
\nu_\eta(y) = \partial_1 \varphi_\eta(y) \times \partial_2 \varphi_\eta(y)
\]

and

\[
R^\eta_{ij}(\eta) := \frac{\partial_{ij} \varphi_\eta \cdot \nu_\eta}{|\partial_1 \varphi \times \partial_2 \varphi|} - \partial_{ij} \varphi \cdot \nu, \quad i, j = 1, 2,
\]

are the covariant components of the change of curvature tensor \(R^2(\eta)\). The elastic energy \(K(\eta) := K(\eta, \eta)\) of the deformation is then given by

\[
K(\eta) = \frac{1}{2} \varepsilon_0 \int_\omega \mathcal{C} : G(\eta) \otimes G(\eta) \, dy + \frac{1}{6} \varepsilon_0^3 \int_\omega \mathcal{C} : R^2(\eta) \otimes R^2(\eta) \, dy
\]

\[
:= \sum_{i,j,k,l=1}^2 \frac{1}{\varepsilon_0^2} \int_\omega G_{ijkl}(\eta) G_{ij}(\eta) \, dy + \frac{1}{6} \varepsilon_0^3 \int_\omega C^{ijkl} R^\eta_{kl}(\eta) R^\eta_{ij}(\eta) \, dy
\]

where \(\mathcal{C} = (C^{ijkl})^2_{i,j,k,l=1}\) is a fourth-order tensor whose entries are the contravariant components of the shell elasticity, see \[3\] Page 162]. We remark that for simplicity, we have normalized the measure \(dy\) in \(1.2\) which should have actually been the weighted measure \(|\partial_1 \varphi \times \partial_2 \varphi| \, dy\) with the weight \(|\partial_1 \varphi \times \partial_2 \varphi|\). Next, given the geometric quantity

\[
\eta(\eta) := 1 + \frac{\eta}{|\partial_1 \varphi \times \partial_2 \varphi|} \left[ \nu \cdot (\partial_1 \varphi \times \partial_2 \nu + \partial_1 \nu \times \partial_2 \varphi) \right] + \frac{\eta^2}{|\partial_1 \varphi \times \partial_2 \varphi|} \nu \cdot (\partial_1 \nu \times \partial_2 \nu),
\]

\[1.3\]
one deduces the $W^{2,2}(\omega)$-coercivity of the Koiter energy [122], as long as $\gamma(\eta) \neq 0$, cf. [34] Lemma 4.3 and Remark 4.4. Finally, we remark that the Koiter energy is continuous on $W^{2,p}(\omega)$ for all $p > \beta > 2$ due to the Sobolev embedding $W^{2,2}(\omega) \hookrightarrow W^{1,\infty}(\omega)$.

For a given function $\eta : I \times \Omega \rightarrow \mathbb{R}$ with an interval $I = (0, T)$ we denote by $\Omega_{\eta(t)}$ the variable in time domain. With a slight abuse of notation we denote by $I \times \Omega_{\eta} = \bigcup_{t \in I} \{t\} \times \Omega_{\eta(t)}$ the deformed time-space cylinder, defined via its boundary

$$\partial \Omega_{\eta(t)} = \{\varphi(y) + \eta(t,y)v(y) : y \in \omega\}.$$  

Along such a cylinder we observe the flow of a heat-conducting compressible fluid subject to the volume force $f : I \times \Omega_{\eta} \rightarrow \mathbb{R}^3$ and the heat source $H : I \times \Omega_{\eta} \rightarrow \mathbb{R}$. We seek the velocity field $u : I \times \Omega_{\eta} \rightarrow \mathbb{R}^3$, the density $\varrho : I \times \Omega_{\eta} \rightarrow \mathbb{R}$ and the temperature $\vartheta : I \times \Omega_{\eta} \rightarrow \mathbb{R}$ solving the following system:

\begin{align}
\partial_t \varrho + \text{div}(\varrho u) &= 0, & \text{in } I \times \Omega_{\eta}, \\
\partial_t (\varrho u) + \text{div}(\varrho u \otimes u) &= \text{div} S(\vartheta, \nabla u) - \nabla p(\varrho, \vartheta) + \varrho f, & \text{in } I \times \Omega_{\eta}, \\
\partial_t (\varrho e(\varrho, \vartheta)) + \text{div}(\varrho e(\varrho, \vartheta) u) &= S(\vartheta, \nabla u) : \nabla u - p(\varrho, \vartheta) \text{div } u, & \text{in } I \times \Omega_{\eta}, \\
-\text{div } q(\varrho, \nabla \vartheta) &= H, & \text{in } I \times \Omega_{\eta}, \\
\varrho(t,x + \eta(t,x)v(x)) &= \varrho_0, & \text{in } I \times \omega, \\
\varrho \vartheta(0) &= \varrho_0, & \text{in } I \times \Omega_{\eta}.
\end{align}

In [135] we suppose Newton’s rheological law

$$S(\vartheta, \nabla u) = \mu(\vartheta)\left(\frac{\nabla u + \nabla u^T}{2} - \frac{1}{3} \text{div } u I\right) + \lambda(\vartheta) \text{div } u I$$

with strictly positive viscosity coefficients $\mu$, $\lambda$ (see Remark 1.3 in [6] for the case $\lambda \geq 0$). The internal energy (heat) flux is determined by Fourier’s law

$$q(\varrho, \nabla \vartheta) = -\kappa(\vartheta) \nabla \vartheta = -\nabla K(\vartheta), \quad K(\vartheta) = \int_0^\vartheta \kappa(z) \, dz$$

with strictly positive heat-conductivity $\kappa$. The thermodynamic functions $p$ and $e$ are related to the (specific) entropy $s$ through Gibbs’ equation

$$\partial Ds(\varrho, \vartheta) = De(\varrho, \vartheta) + p(\varrho, \vartheta) D\left(\frac{1}{\varrho}\right) \quad \text{for all } \varrho, \vartheta > 0.$$

The model case is given by

$$p(\varrho, \vartheta) = \varrho^\gamma + \varrho \vartheta + a \varrho^4, \quad e(\varrho, \vartheta) = \frac{1}{\gamma - 1} \varrho^{\gamma - 1} + c_v \vartheta + a \varrho^4, \quad s(\varrho, \vartheta) = \frac{4a \varrho^3}{3} + \log(\varrho^3) - \log \varrho,$$

for $a, c_v > 0$ and $\gamma > 1$. In view of Gibb’s relation (1.11), the internal energy equation (1.10) can be rewritten in the form of the entropy balance

$$\partial_s gs(\varrho, \vartheta) + \text{div}(gs(\varrho, \vartheta) u) = - \text{div } \left(\frac{q(\varrho, \nabla \vartheta)}{\varrho}\right) + \sigma + \vartheta \frac{H}{\varrho},$$

with the entropy production rate

$$\sigma = \frac{1}{\varrho} \left(\text{div } S(\vartheta, \nabla u) : \nabla u - q(\varrho, \nabla \vartheta) \cdot \nabla \vartheta\right).$$

In the weak formulation (1.12) will be replaced by a variational inequality. The shell should respond optimally with respect to the forces, which act on the boundary. Therefore we have

$$\varepsilon_0 \delta_S \partial_t^2 \eta + K'(\eta) = g + \nu \cdot F \quad \text{in } I \times \omega,$$

where $\delta_S > 0$ is the density of the shell. Here, $g : I \times \omega \rightarrow \mathbb{R}$ is a given force and $F$ is given by

$$F := -\tau \nu(\varrho) \circ \varphi_{\eta(t)}|_{\partial \Omega_{\eta(t)}} \det D\varphi_{\eta(t)}|_{\partial \Omega_{\eta(t)}}$$

Here, $\varphi_{\eta(t)} : \omega \rightarrow \partial \Omega_{\eta(t)}$ is the change of coordinates from (1.1) and $\tau$ is the Cauchy stress. To simplify the presentation in (1.14) we will assume

$$\varepsilon_0 \delta_S = 1$$

throughout the paper. We assume the following boundary and initial values for $\eta$

$$\eta(0,\cdot) = \eta_0, \quad \partial_t \eta(0,\cdot) = \eta_1 \quad \text{in } \omega.$$
where \( \eta_0, \eta_1 : \omega \to \mathbb{R} \) are given functions. Here, we assume that
\[
\text{Im}(\eta_0) \subset (a, b).
\]
In view of (1.7) we have to suppose the compatibility condition\(^1\)
\[
\eta_1(y)\nu(y) = \frac{q_0}{\rho_0}(y + \eta(y)\nu(y)) \quad \text{in} \quad \omega.
\]
(1.16)

Our main result is the following existence theorem. The system (1.3)–(1.15) can be written in a natural way as a weak solution. The concept is introduced in the next section, (see 2.1.3–2.1.9), where also the precise formulation of our main result is presented (see Theorem 2.16). It is concerned with the existence of a weak solution up to degeneracy of the geometry and reads in a simplified version as follows.

**Theorem 1.1.** Under natural assumptions on the data there exists a weak solution \((\eta, u, \varrho, \vartheta)\) to (1.3)–(1.15) with satisfies the energy balance
\[
\mathcal{E}(t) = \mathcal{E}(0) + \int_{\Omega_0} \rho z H \, d\mathbf{x} + \int_{\Omega_0} \rho \vartheta \cdot \mathbf{u} \, d\mathbf{x} + \int_{\omega} g \partial_t \eta \, d\mathbf{y},
\]
(1.17)
\[
\mathcal{E}(t) = \int_{\Omega_0(t)} \left( \frac{1}{2} \varrho(t)|\mathbf{u}(t)|^2 + \varrho(t) c(\varrho(t), \vartheta(t)) \right) \, d\mathbf{x} + \int_{\omega} \frac{\partial_t \eta(t)^2}{2} \, d\mathbf{y} + K(\eta(t)).
\]

The interval of existence is of the form \( I = (0, t) \), where \( t < T \) only in case \( \Omega_{\eta(s)} \) approaches a self-intersection when \( s \to t \) or the Koiter energy degenerates (namely, if \( \lim_{s \to t} \gamma(s, y) = 0 \) for some point \( y \in \omega \)).

The function space of existence for a weak solution to (1.3)–(1.15) is determined by the total energy \( \mathcal{E} \) in (1.17) as well as the quantity \( \sigma \) in (1.13) taking into account the variable domain. Theorem 2.16 extends the results from [3] to the case of a heat-conducting fluid but also applies to nonlinear structure equations. As in the case of fixed domains studied in [21] (see also [17] and [20]) the heat-conducting model allows (different to the isentropic equations) the striking feature of an energy equality. Energy, which is lost by dissipation, is transferred into heat, cf. (1.10).

### 1.3. Mathematical strategy

In this paragraph we provide an overview of the developed methodologies. Further we aim to explain all technical novelties and their potential significance.

As is common in the existence theory for weak solutions, the first step is to understand how to prove sequential compactness. Let us assume there is a given sequence of weak solutions \((\eta_n, u_n, \varrho_n, \vartheta_n)\) to (1.3)–(1.15) possessing suitable regularity properties. Deriving a priori estimates using the entropy balance one can control, in addition to the total energy defined in (1.17), first order spatial derivative of \( u_n \) and \( \vartheta_n \) using (1.12) and (1.13). Unlike in the steady domain case, these estimates are not sufficient to show that a subsequence is again converging to a solution. One problem is to derive energy equality (that is expected for closed systems like the Navier–Stokes–Fourier equations considered here). Critical are the kinetic and elastic part of the solid energy. To prove their compactness which does not follow from the energy estimates. In fact, the functional \( K \) is not even well-defined on \( H^{2,2} (\omega) \) recalling the discussion from Section 1.2. Our strategy is to derive fractional estimates for \( \nabla^2 \eta_n \) as a consequence of a testing procedure for (1.14) with difference quotients. Testing the shell equation with suitable test-functions requires in the weak formulation to choose an appropriate test-function for the full momentum equation as well. Technically, this means we have to “extend” functions defined on \( \omega \) to functions defined on the time dependent domain \( \Omega_{\eta_n} \). An obstacle here is that the pressure is only expected to belong to \( L^1 \) in space near the moving boundary (compare with [3] 2). To circumvent the irregularity of the pressure we work with a solenoidal extension \( \mathcal{F}^{\text{div}}_{\eta_n} \) that was recently constructed in [30] (compare also with [31]).

A second related problem is the strong convergence of \( \partial_t \eta_n \) (which is a part of the kinetic energy). Here we use a modified version of the classical argument by Aubin–Lions. Non-standard are uniform continuity estimates in time of the underlying sequence, which rely on the weak coupled momentum equation. Again a carefully chosen test-function is needed. Here, however, we use an extension which has (different to \( \mathcal{F}^{\text{div}}_{\eta_n} \)) a regularizing effect but no solenoidality is needed. What turns out to be the most sensitive point is that the

\(^1\)Note that the above condition is necessary for strong solutions only and hence is not effecting the rest of the paper. In particular, since the assumptions on the initial data are only in Lebesgue spaces, the compatibility condition is void. It does, however, implicitly appear in the construction of the Galerkin bases, where a smooth approximation of the initial values is considered.

\(^2\)As is explained in [3] the usual test with the Bogovskii-operator (that implies higher integrability of the density) fails and we are only able to prove uniform integrability, cf. Lemmas 5.3 and 6.20.
extension is depending on the variable geometry. In particular, the extension of a constant in time function still possesses a non-trivial time-derivative. The essential term is
\[
\int_I \int_{\Omega_{\eta_n}} g_n u_n \cdot \partial_t (\bar{\mathcal{F}}_{\eta_n} b) \, dx \, dt
\]
using the notation from the next section. We observe that \( \partial_t (\bar{\mathcal{F}}_{\eta_n} b) \) (the time-derivative of the extension) is expected to behave like \( \partial_t \eta_n \). Based on the a priori estimates \( g_n \in L^\infty_t (L^2_r) \), \( u_n \in L^2_t (L^6_r) \), we find that \( \partial_t \eta_n \in L^2_t (L^6_r) \) for all \( r < 4 \) uniformly by the trace theorem (see Lemma 2.3). Consequently the bound \( \gamma > \frac{3}{4} \) naturally appears. It is interesting to note that the same bound was needed in [6, Lemma 7.4] in order to avoid concentrations of the approximate pressure at the boundary (an argument that we will use later in Lemma 5.13).

In order to prove Theorem 2.16 we have to work with a multi-layer approximation scheme. As is nowadays standard in the theory of compressible fluids we follow [22] and use an artificial pressure (replace \( \varepsilon \Delta \vartheta \) to the right-hand side of (1.4)). The resulting system is solved by means of a Galerkin approximation. More specifically, we have to solve a finite-dimensional system of ODEs and eventually pass to the limit in the dimension \( N \). It turns out that existence on the basic level, where the parameters \( \varepsilon \) and \( \delta \) are fixed, is quite involving. Troublesome is the derivation of the entropy balance (1.12) (in form of a variational inequality): Though it is suitable to pass to the limit it is not appropriate for the direct construction of solutions due to its highly involving non-linearities. Hence the entropy balance is derived a posteriori by dividing the internal energy equation (1.9) by \( \vartheta \). In order to do this rigorously it has to be shown that the temperature is strictly positive - a property which can only be expected from strong solutions to (1.9). One of the main efforts of this paper is consequently to construct strong solutions to (1.9) for regularized velocity and smooth pressure. New a priori estimates for (1.9) and (1.13) in variable domains are shown that go well beyond the results from [6, Sec. 3] and form one if the main achievements of this paper. Finally, we wish to note that we can shorten the approach from [6] considerably. Different to [6] we decouple the geometry from the fluid system on the Galerkin level and apply the fixed point argument to the resulting semi-discrete problem directly. This allows to remove one regularization level in which the moving boundary and the convective terms are regularised.

1.4. Overview of the paper. In Section 2 we present basics concerning variable domains as well as the functional analytic set-up. In its last subsection the concept of weak solutions for the coupled system and the main theorem are introduced. The preliminary section is rather significant. Indeed, many standard tools of the analysis need an appropriate adaptation to the variable geometry set-up, as well as to the particular non-linear coupling of the PDE system. In particular, in Subsection 2.3 we introduce two different extension operators that are needed for the analysis performed later. In Section 3 we study the (regularized) continuity equation as well as the (regularized) internal energy equation in a time dependent domain. These are non-trivial extensions from the analysis presented in [6, Section 3]. In particular, we provide regularity estimates and minimum and maximum principles. Section 4 is dedicated to the construction of an approximate solution. Different to previous fixed point approaches (see e.g. [6] and [31]) we construct a fixed point on the Galerkin level which we believe to be appropriate also for future applications. A further achievement is the derivation of the entropy inequality which sensitively relies on Section 3 Finally, in Section 5 the two limit passages \( \varepsilon \to 0 \) and \( \delta \to 0 \) are performed which leads to the proof of Theorem 2.16 and the existence of a weak solution is shown. Of particular importance is here Subsection 4.4 where the derivation of an energy equality is performed. Critical is the strong convergence of the elastic energy of the solid deformation. Here we adapt a regularity argument for the shell displacement derived in [39]. As shown in [39] these estimates are crucial to involve non-linear Koiter shell laws in the weak existence theory for incompressible fluids. In the here considered Navier-Stokes-Fourier system the regularity is needed even for linear shell models. Since, even for linear Koiter shell models an energy equality cannot be derived without additional regularity estimates and the related compactness properties.

2. Preliminaries

2.1. Structural and constitutive assumptions. We impose several restrictions on the specific shape of the thermodynamic functions \( p = p(\vartheta, \bar{\vartheta}) \), \( \varepsilon = \varepsilon(\vartheta, \bar{\vartheta}) \) and \( s = s(\vartheta, \bar{\vartheta}) \) which are in line with Gibbs’ relation (1.11). We consider the pressure \( p \) in the form
\[
p(\vartheta, \bar{\vartheta}) = p_M(\vartheta) + p_R(\vartheta, \bar{\vartheta}), \quad p_M(\vartheta) = \bar{\vartheta}^3 + \bar{\vartheta^3}, \quad p_R(\vartheta, \bar{\vartheta}) = \frac{a}{3} \varepsilon^3, \quad a > 0,
\]
the specific internal energy
\[
e(\varrho, \vartheta) = e_M(\varrho) + e_R(\varrho, \vartheta), \quad e_M(\varrho) = \frac{1}{\gamma - 1} \varrho^{\gamma - 1} + c_v \vartheta, \quad e_R(\varrho, \vartheta) = \varrho \frac{\vartheta^4}{\varrho}, \quad c_v > 0,
\]
and the specific entropy
\[
s(\varrho, \vartheta) = \frac{4a}{3} \frac{\vartheta^3}{\varrho} + \log(\vartheta^{c_v}) - \log \varrho.
\]
This is model case for the set-up in [21, Chapter 1], to which we refer for the physical background and the relevant discussion.

The viscosity coefficients $\mu$, $\lambda$ are continuously differentiable functions of the absolute temperature $\vartheta$, more precisely $\mu, \lambda \in C^1([0, \infty))$, satisfying
\[
\mu(1 + \vartheta) \leq \mu(\vartheta) \leq \overline{\mu}(1 + \vartheta),
\]
\[
\sup_{\vartheta \in [0, \infty)} (|\mu'(\vartheta)| + |\lambda'(\vartheta)|) \leq \overline{\mu},
\]
\[
\Lambda(1 + \vartheta) \leq \lambda(\vartheta) \leq \overline{\lambda}(1 + \vartheta),
\]
with positive constants $\underline{\mu}, \overline{\mu}, \underline{\lambda}, \overline{\lambda}$. The heat conductivity coefficient $\kappa \in C^1[0, \infty)$ satisfies
\[
0 < \underline{\kappa}(1 + \vartheta^2) \leq \kappa(\vartheta) \leq \overline{\kappa}(1 + \vartheta^2)
\]
with some positive $\underline{\kappa}, \overline{\kappa}$. We introduce the following regularizations
\[
p_\delta(\varrho, \vartheta) := p_R(\varrho, \vartheta) + p_{M, \delta}(\vartheta), \quad p_{M, \delta}(\vartheta) := p_M(\varrho) + \delta \vartheta^\beta,
\]
\[
e_\delta(\varrho, \vartheta) := e_R(\varrho, \vartheta) + e_{M, \delta}(\vartheta), \quad e_{M, \delta}(\vartheta) := e_M(\varrho) + \frac{\delta}{\beta - 1} \vartheta^{\beta - 1},
\]
\[
\kappa_\delta(\vartheta) = \kappa(\vartheta) + \delta \left( \vartheta^\beta + \frac{1}{\varrho} \right), \quad \Delta_\delta(\vartheta) = \int_0^\varrho \kappa_\delta(z) \, dz,
\]
\[
S^e(\vartheta, \nabla u) = S(\vartheta, \nabla u) + \varepsilon(1 + \vartheta) |\nabla u|^{p-2} \nabla u,
\]
for some $p > \beta > 2$.

2.2. Function spaces on variable domains. The spatial domain $\Omega$ is assumed to be a non-empty bounded subset of $\mathbb{R}^3$ with $C^1$-boundary and an outer unit normal $\nu$. We recall from Section 1.2 that we assume that $\partial \Omega$ can be parametrised by an injective mapping $\varphi \in C^4(\omega; \mathbb{R}^3)$ such that for all points $y = (y_1, y_2) \in \omega$, the pair of vectors $\partial_i \varphi(y)$, $i = 1, 2$, are linearly independent. For a point $x$ in the neighbourhood of $\partial \Omega$ we can define
\[
y(x) = \arg \min_{y \in \omega} |x - \varphi(y)|, \quad s(x)\text{ is defined such that } s(x)\nu(y(x)) + y(x) = x.
\]
Moreover, we define the projection $p(x) = \varphi(p(y(x)))$. We define $L > 0$ to be the largest number such that $s, y$ and $p$ are well-defined on $S_L$, where
\[
S_L = \{x \in \mathbb{R}^3 : \text{dist}(x, \partial \Omega) < L\},
\]
see also Remark 2.18 in connection with this. We remark that due to the $C^2$ regularity of $\Omega$ for $L$ small enough we find that $|s(x)| = \min_{y \in \omega} |x - \varphi(y)|$ for all $x \in S_L$. This implies that $S_L = \{s \nu(y) + y : (s, y) \in [-L, L] \times \omega\}$. For a given function $\eta: I \times \omega \to \mathbb{R}$ we parametrise the deformed boundary by
\[
\varphi_\eta(t, y) = \varphi(y) + \eta(t, y) \nu(y), \quad y \in \omega, \; t \in I,
\]
and the deformed space-time cylinder $I \times \Omega_\eta = \bigcup_{t \in I} \{t\} \times \Omega_{\eta(t)}$ through
\[
\partial \Omega_{\eta(t)} = \{\varphi(y) + \eta(t, y) \nu(y) : y \in \omega\}.
\]
The corresponding function spaces for variable domains are defined as follows.

**Definition 2.1. (Function spaces)** For $I = (0, T)$, $T > 0$, and $\eta \in C(\overline{\mathbb{T}} \times \omega)$ with $\|\eta\|_{L^r_{\text{loc}}(\mathbb{T})} < L$ we set $I \times \Omega_\eta := \bigcup_{t \in I} \{t\} \times \Omega_{\eta(t)} \subset \mathbb{R}^4$. We define for $1 \leq p, r \leq \infty$
\[
L^p(I; L^r(\Omega_\eta)) := \{v \in L^p(I \times \Omega_\eta) : \nu(v(t, \cdot)) \in L^r(\Omega_{\eta(t)}), \text{ for a.e. } t, \; \|v(t, \cdot)\|_{L^r(\Omega_{\eta(t)})} \in L^p(I)\},
\]
\[
L^p(I; W^{1,r}(\Omega_\eta)) := \{v \in L^p(I; L^r(\Omega_\eta)) : \nabla v \in L^p(I; L^r(\Omega_\eta))\}.
\]
For various purposes it is useful to relate the time dependent domains and the fixed domain. This can be done by the means of the Hanzawa transform. Its construction can be found in [31] pages 210, 211]. Note that variable domains in [31] are defined via functions $\zeta : \partial \Omega \to \mathbb{R}$ rather than functions $\eta : \omega \to \mathbb{R}$ (clearly, one can link them by setting $\zeta = \eta \circ \varphi^{-1}$). For any $\eta : \omega \to (-L, L)$ we define the Hanzawa transform $\Psi_\eta : \Omega \to \Omega_\eta$ by

\begin{equation}
\Psi_\eta(x) = \begin{cases} 
p(x) + (s(x) + \eta(y(x))\phi(s(x))) \nu(y(x)), & \text{if } \text{dist}(x, \partial \Omega) < L, \\
x, & \text{elsewhere}
\end{cases}
\end{equation}

Here $\phi \in C^\infty((-\frac{4L}{3}, \infty), [0, 1])$ is such that $\phi \equiv 0$ in $(-\frac{4L}{3}, -\frac{L}{2})$ and $\phi \equiv 1$ in $[-\frac{L}{2}, \infty)$. Due to the size of $L$, we find that $\Psi_\eta$ is a homomorphism such that $\Psi_\eta \vert_{\Omega \setminus S_L}$ is the identity. Moreover, $\eta \in C^k(\omega)$ for $k \in \mathbb{N}$ implies that $\Psi_\eta$ is a $C^k$-diffeomorphism.

We collect a few properties of the above mapping $\Psi_\eta$.

**Lemma 2.2.** Let $1 < p \leq \infty$ and $\sigma \in (0, 1]$.

a) If $\eta \in W^{2,2}(\omega)$ with $\|\eta\|_{L^\infty} < L$, then the linear mapping $v \mapsto v \circ \Psi_\eta$ ($v \mapsto v \circ \Psi_\eta^{-1}$) is continuous from $L^p(\Omega_\eta)$ to $L^p(\Omega)$ (from $L^p(\Omega)$ to $L^p(\Omega_\eta)$) for all $1 \leq r < p$.

b) If $\eta \in W^{2,2}(\omega)$ with $\|\eta\|_{L^\infty} < L$, then the linear mapping $v \mapsto v \circ \Psi_\eta$ ($v \mapsto v \circ \Psi_\eta^{-1}$) is continuous from $W^{1,p}(\Omega_\eta)$ to $W^{1,p}(\Omega)$ (from $W^{1,p}(\Omega)$ to $W^{1,p}(\Omega_\eta)$) for all $1 \leq r < p$.

c) If $\eta \in C^0(\omega)$ with $\|\eta\|_{L^\infty} < L$, then the linear mapping $v \mapsto v \circ \Psi_\eta$ ($v \mapsto v \circ \Psi_\eta^{-1}$) is continuous from $W^{0,\sigma}(\Omega_\eta)$ to $W^{0,\sigma}(\Omega)$ (from $W^{0,\sigma}(\Omega)$ to $W^{0,\sigma}(\Omega_\eta)$).

d) If $\eta \in W^{2,2}(\partial \Omega)$ with $\|\eta\|_{L^\infty} < L$, then the linear mapping $v \mapsto v \circ \Psi_\eta$ ($v \mapsto v \circ \Psi_\eta^{-1}$) is continuous from $W^{0,\sigma}(\partial \Omega_\eta)$ to $W^{0,\sigma}(\partial \Omega)$ (from $W^{0,\sigma}(\partial \Omega)$ to $W^{0,\sigma}(\partial \Omega_\eta)$)

The continuity constants depend only on $\Omega, p, r, \sigma, \theta$, the respective norms of $\eta$.

The following lemma is a modification of [31] Cor. 2.9].

**Lemma 2.3.** Let $1 < p < 3$, $\sigma \in (\frac{1}{2}, 1]$ and $\eta \in W^{2,2}(\omega)$ with $\|\eta\|_{L^\infty} < L$. The linear mapping $\text{tr}_\eta : v \mapsto v \circ \varphi \circ \varphi_\eta \circ \varphi_\eta^{-1}$ is well defined and continuous from $W^{\sigma,p}(\Omega_\eta)$ to $W^{\sigma-\frac{1}{2},r}(\omega)$ for all $r \in (\frac{1}{2}, p)$ and well defined and continuous from $W^{\sigma,p}(\Omega)$ to $L^q(\omega)$ for all $1 < q < \frac{\sigma - \frac{1}{2}}{\sigma - p}$. The continuity constants depend only on $\Omega, p, \sigma$ and $\|\eta\|_{W^{2,2}}$.

**Remark 2.4.** If $\eta \in L^\infty(\Gamma; W^{2,2}(\omega))$ we obtain non-stationary variants of the results stated above.

It will be convenient for our purposes to extend $\Psi_\eta$, originally defined only on $\Omega(L-\eta)$, as follows.

$$
\overline{\Psi}_\eta(x) = \begin{cases} 
p(x) + (s(x) + \eta(y(x))\phi(s(x))) \nu(y(x)), & \text{if } \text{dist}(x, \partial \Omega) < L, s(x) + \eta(y(x)) < L, \\
x, & \text{elsewhere}
\end{cases}
$$

All the above statements are also true for $v \mapsto v \circ \overline{\Psi}_\eta$ and $v \mapsto v \circ \overline{\Psi}_\eta^{-1}$ on their respective domains.

2.3. **Extensions on variable domains.** Since $\Omega$ is assumed to be sufficiently smooth, it is well-known that there is an extension operator $\mathcal{F}_\Omega$ which extends functions from $\partial \Omega$ to $\mathbb{R}^3$ and satisfies

$$
\mathcal{F}_\Omega : W^{\sigma,p}(\partial \Omega) \to W^{\sigma+1/p,p}(\mathbb{R}^3)
$$

for all $p \in (1, \infty)$ and $\sigma \in [0, 1]$, as well as $\mathcal{F}_\Omega v \vert_{\partial \Omega} = v$. Now we define $\mathcal{F}_\eta$ by

\begin{equation}
\mathcal{F}_\eta b = \mathcal{F}_\Omega ((b \circ \varphi^{-1}) \circ \overline{\Psi}_\eta^{-1}), \quad b \in W^{\sigma,p}(\omega),
\end{equation}

where $\varphi$ is the $C^1$-function in the parametrisation of $\Omega$. If $\eta$ is smooth $\mathcal{F}_\eta$ behaves as a classical extension by Lemma 2.2. The following properties can all be easily derived from the formulas

$$
\nabla \mathcal{F}_\eta b = \nabla \mathcal{F}_\Omega ((b \circ \varphi^{-1}) \circ \overline{\Psi}_\eta^{-1}) \circ \nabla \overline{\Psi}_\eta^{-1},
$$

$$
\nabla^2 \mathcal{F}_\eta b = \nabla^2 \mathcal{F}_\Omega ((b \circ \varphi^{-1}) \circ \overline{\Psi}_\eta^{-1}) \circ \nabla \overline{\Psi}_\eta^{-1} + \nabla \mathcal{F}_\Omega ((b \circ \varphi^{-1}) \circ \overline{\Psi}_\eta^{-1}) \circ \nabla^2 \overline{\Psi}_\eta^{-1},
$$

$$
\partial_x \mathcal{F}_\eta b = \nabla \mathcal{F}_\Omega ((b \circ \varphi^{-1}) \circ \overline{\Psi}_\eta^{-1}) \circ \partial_x \overline{\Psi}_\eta^{-1},
$$

where $\nabla \overline{\Psi}_\eta^{-1}, \nabla^2 \overline{\Psi}_\eta^{-1}$ and $\partial_x \overline{\Psi}_\eta^{-1}$ behave as $\nabla \eta, \nabla^2 \eta$ and $\partial_x \eta$ respectively.
Corollary 2.6. Let \( \eta \in C^{0,1}(\omega) \) with \( \| \eta \|_{L^\infty_\omega} < \alpha < L \).

(a) The operator \( \mathcal{F}_\eta \) defined in (2.11) satisfies for all \( p \in [1, \infty) \) and \( \sigma \in [0,1] \)
\[
\mathcal{F}_\eta : W^{\sigma,p}(\omega) \to W^{\sigma+1/p,p}(\Omega \cup S_\alpha)
\]
and \( \text{tr}_\eta \mathcal{F}_\eta b = bv \) for all \( b \in W^{1,p}(\omega) \). In particular, we have
\[
\| \mathcal{F}_\eta b \|_{W^{\sigma+1/p,p}(\Omega \cup S_\alpha)} \leq c \| b \|_{W^{\sigma,p}(\omega)}
\]
for all \( b \in W^{1,p}(\omega) \), where the constant \( c \) depends only on \( \Omega, p, \sigma, \| \nabla \eta \|_{L^\infty_\omega} \) and \( L - \alpha \).

(b) If \( p = \infty \) we have
\[
\| \mathcal{F}_\eta b \|_{W^{1,\infty}(\Omega \cup S_\alpha)} \leq c(1 + \| \nabla \eta \|_{L^\infty_\omega}) \| b \|_{W^{1,\infty}(\omega)}
\]
for all \( b \in W^{1,\infty}(\omega) \), where \( c \) depends only on \( \Omega, p \) and \( L - \alpha \).

Lemma 2.5. Let \( \eta \in C^1(\mathcal{T} \times \omega) \) with \( \| \eta \|_{L^\infty_\omega} < \alpha < L \). Then we have for all \( q < \infty \)
\[
\sup_{t \in I} \| \partial_t \mathcal{F}_\eta b \|_{L^q(\Omega \cup S_\alpha)} \leq c \| b \|_{W^{1,q}(\omega)} \| \partial_t \eta \|_{L^\infty(I \times \omega)}
\]
for all \( b \in W^{1,q}(\omega) \), where the constant \( c \) depends only on \( \Omega, p \) and \( L - \alpha \).

We now turn to the case of a less regular function \( \eta \) and analyse the properties of \( \mathcal{F}_\eta \) given by (2.11) in this case.

Lemma 2.7. Let \( p \in [1, \infty) \) and \( \eta \in W^{2,2}(\omega) \) with \( \| \eta \|_{L^\infty_\omega} < \alpha < L \) and let the operator \( \mathcal{F}_\eta \) by defined by (2.11).

(a) We have for all \( p \in (1, \infty) \) and \( \sigma \in (0,1] \)
\[
\mathcal{F}_\eta : W^{\sigma,p}(\omega) \to W^{\sigma-1/p,p}(\Omega \cup S_\alpha)
\]
for all \( q < \frac{\sigma}{\sigma - 1}p \) and \( \text{tr}_\eta \mathcal{F}_\eta b = bv \) for all \( b \in W^{1,p}(\omega) \). In particular, we have
\[
\| \mathcal{F}_\eta b \|_{W^{\sigma-1/p,p}(\Omega \cup S_\alpha)} \leq c \| b \|_{W^{\sigma,p}(\omega)}
\]
for all \( b \in W^{1,p}(\omega) \).

(b) We have for all \( r < 2 \)
\[
\mathcal{F}_\eta : W^{2,2}(\omega) \to W^{2,r}(\Omega \cup S_\alpha)
\]
and \( \text{tr}_\eta \mathcal{F}_\eta b = bv \) for all \( b \in W^{2,2}(\omega) \). In particular, we have
\[
\| \mathcal{F}_\eta b \|_{W^{2,r}(\Omega \cup S_\alpha)} \leq c \| b \|_{W^{2,2}(\omega)}
\]
for all \( b \in W^{2,2}(\omega) \).

The constants in (a) and (b) depend only on \( \Omega, p, q, \| \eta \|_{W^{2,2}} \) and \( L - \alpha \).

Corollary 2.8. Let \( \eta \in L^2(I; W^{2,2}(\partial \Omega)) \) with \( \| \eta \|_{L^\infty_{I \times \partial \Omega}} \leq \alpha < L \), Suppose that \( \partial_t \eta \in L^q(I \times \omega) \) for some \( q > 1 \).
Then we have uniformly in time
\[
\| \partial_t \mathcal{F}_\eta b \|_{L^p(I \times \Omega \cup S_\alpha)} \leq c \| b \|_{W^{1,q}(\omega)} \| \partial_t \eta \|_{L^\infty(I \times \omega)}
\]
for all \( b \in W^{1,p}(\omega) \), provided \( \frac{1}{p} = \frac{1}{r} + \frac{1}{q} \leq 1 \). The constant \( c \) depends only on \( \Omega, p \) and \( L - \alpha \).

The following is proved in [39, Prop. 3.3]. It provides a solenoidal extension. For that we introduce the solenoidal space \( W^{1,1}_\text{div}(\Omega \cup S_\alpha) := \{ w \in W^{1,1}(\Omega \cup S_\alpha) : \text{div} w = 0 \} \). The corrector \( \mathcal{X}_\eta \) in the below preconditions the boundary data to be compatible with the interior solenoidality.

Proposition 2.9. For a given \( \eta \in L^\infty(I; W^{1,2}(\omega)) \) with \( \| \eta \|_{L^\infty_{I \times \omega}} < \alpha < L \), there are linear operators
\[
\mathcal{X}_\eta : L^1(\omega) \to \mathbb{R}, \quad \mathcal{F}_\eta^\text{div} : \{ \xi \in L^1(I; W^{1,1}(\omega)) : \mathcal{X}_\eta(\xi) = 0 \} \to L^1(I; W^{1,1}_\text{div}(\Omega \cup S_\alpha)),
\]
such that the tuple \( (\mathcal{F}_\eta^\text{div}(\xi - \mathcal{X}_\eta(\xi)), \xi - \mathcal{X}_\eta(\xi)) \) satisfies
\[
\mathcal{F}_\eta^\text{div}(\xi - \mathcal{X}_\eta(\xi)) \in L^\infty(I; L^2(\Omega_\eta)) \cap L^2(I; W^{1,2}_\text{div}(\Omega_\eta)),
\]
\[
\xi - \mathcal{X}_\eta(\xi) \in L^\infty(I; W^{2,2}(\omega)) \cap W^{1,\infty}(I; L^2(\omega)),
\]
\[
\text{tr}_\eta(\mathcal{F}_\eta^\text{div}(\xi - \mathcal{X}_\eta(\xi))) = \xi - \mathcal{X}_\eta(\xi),
\]
\[
\mathcal{F}_\eta^\text{div}(\xi - \mathcal{X}_\eta(\xi))(t, x) = 0 \text{ for } (t, x) \in I \times (\Omega \setminus S_\alpha)
\]
provided we have \( \xi \in L^\infty(I; W^{2,2}(\omega)) \cap W^{1,\infty}(I; L^2(\omega)) \). In particular, we have the estimates

\[
(2.12) \quad \| F^\text{div}_\eta (\xi - \mathcal{N}_\eta(\xi)) \|_{L^p(I, W^{1,p}(\Omega; S_\eta))] \lesssim \| \xi \|_{L^p(I, W^{1,p}(\omega))} + \| \nabla \eta \|_{L^p(I, W^p(\omega))},
\]

\[
(2.13) \quad \| \partial_\xi F^\text{div}_\eta (\xi - \mathcal{N}_\eta(\xi)) \|_{L^p(I, L^p(\Omega; S_\eta))] \lesssim \| \partial_\xi \xi \|_{L^p(I, L^p(\omega))} + \| \xi \partial_\eta \|_{L^p(I, L^p(\omega))},
\]

for any \( p \in (1, \infty), q \in (1, \infty) \).

2.4. Convergence in variable domains. Due to the variable domain the framework of Bochner spaces is not available. Hence, we cannot use the classical Aubin-Lions compactness theorem. In this subsection we are concerned with the question of how to get compactness anyway. We start with the following definition of convergence in variable domains.

**Definition 2.10.** Let \( (\eta_t) \subset C(\overline{T} \times \omega; [-\theta L, \theta L]) \), \( \theta \in (0, 1) \), be a sequence with \( \eta_t \to \eta \) uniformly in \( \overline{T} \times \omega \). Let \( p, q \in [1, \infty) \) and \( k \in \mathbb{N}_0 \).

- (a) We say that a sequence \((g_i) \subset L^p(I, L^q(\Omega_{\eta_t}))\) converges to \( g \) in \( L^p(I, L^q(\Omega_{\eta_t})) \) with respect to \((\eta_t)\), in symbols \( g_i \rightharpoonup^* g \) in \( L^p(I, L^q(\Omega_{\eta_t})) \), if
  \[
  \chi_{\Omega_{\eta_t}} g_i \to \chi_{\Omega_{\eta_t}} g \quad \text{in} \quad L^p(I, L^q(\mathbb{R}^3)).
  \]

- (b) Let \( p, q < \infty \). We say that a sequence \((g_i) \subset L^{p,1}(I, L^q(\Omega_{\eta_t}))\) converges to \( g \) in \( L^{p,1}(I, L^q(\Omega_{\eta_t})) \) weakly with respect to \((\eta_t)\), in symbols \( g_i \rightharpoonup^* g \) in \( L^{p,1}(I, L^q(\Omega_{\eta_t})) \), if
  \[
  \chi_{\Omega_{\eta_t}} g_i \to \chi_{\Omega_{\eta_t}} g \quad \text{in} \quad L^p(I, L^q(\mathbb{R}^3)).
  \]

- (c) Let \( p = \infty \) and \( q < \infty \). We say that a sequence \((g_i) \subset L^\infty(I, L^q(\Omega_{\eta_t}))\) converges to \( g \) in \( L^\infty(I, L^q(\Omega_{\eta_t})) \) weakly* with respect to \((\eta_t)\), in symbols \( g_i \rightharpoonup^* g \) in \( L^\infty(I, L^q(\Omega_{\eta_t})) \), if
  \[
  \chi_{\Omega_{\eta_t}} g_i \rightharpoonup^* \chi_{\Omega_{\eta_t}} g \quad \text{in} \quad L^\infty(I, L^q(\mathbb{R}^3)).
  \]

Note that in the case of one single \( \eta \) (i.e. not a sequence) the space \( L^p(I, L^q(\Omega_t)) \) (with \( 1 \leq p < \infty \) and \( 1 < q < \infty \)) is reflexive and we have the usual duality pairing

\[
(2.14) \quad L^p(I, L^q(\Omega)) \cong L^{p'}(I, L^{q'}(\Omega))
\]

provided \( \eta \) is smooth enough, see \[ \text{[10]} \]. Definition 2.11 can be extended in a canonical way to Sobolev spaces: A sequence \((g_i) \subset L^p(I, W^{1,q}(\Omega_{\eta_t}))\) converges to \( g \) in \( L^p(I, W^{1,q}(\Omega_{\eta_t})) \) strongly with respect to \((\eta_t)\), in symbols \( g_i \rightharpoonup^* g \) in \( L^p(I, W^{1,q}(\Omega_{\eta_t})) \), if both \( g_i \) and \( \nabla g_i \) converges (to \( g \) and \( \nabla g \) respectively) in \( L^p(I, L^q(\Omega_{\eta_t})) \) strongly with respect to \((\eta_t)\) (in the sense of Definition 2.11(a)). We also define weak and weak* convergence in Sobolev spaces with respect to \((\eta_t)\) with an obvious meaning. Note that also an extension to higher order Sobolev spaces is possible but not needed for our purposes.

For the next compactness lemma (see \[ \text{[2]} \) Lemma 2.8) we require the following assumptions on the functions describing the boundary

- (A1) The sequence \((\eta_t) \subset C(\overline{T} \times \omega; [-\theta L, \theta L])\), \( \theta \in (0, 1) \), satisfies
  \[
  \eta_t \rightharpoonup^* \eta \quad \text{in} \quad L^\infty(I, W^{2,2}(\omega)),
  \]
  \[
  \partial_\eta \eta_t \rightharpoonup^* \partial_\eta \eta \quad \text{in} \quad L^\infty(I, L^2(\omega)).
  \]

- (A2) Let \((v_t)\) be a sequence such that for some \( p, s \in [1, \infty) \) and \( \alpha \in (0, 1) \) we have
  \[
  v_t \rightharpoonup^* v \quad \text{in} \quad L^p(I, W^{\alpha,s}(\Omega_{\eta_t})).
  \]

- (A3) Let \((r_t)\) be a sequence such that for some \( m, b \in [1, \infty) \) we have
  \[
  r_t \rightharpoonup^* r \quad \text{in} \quad L^m(I, L^b(\Omega_{\eta_t})).
  \]

Assume further that there are sequences \((H_1^t)\), \((H_2^t)\) and \((h_t)\), bounded in \( L^m(I; L^b(\Omega_{\eta_t})) \), such that

\[
\int_I \int_{\Omega_{\eta_t}} r_t \partial_\phi \, dx \, dt = \int_I \int_{\Omega_{\eta_t}} H_1^t : \nabla^2 \phi \, dx \, dt + \int_I \int_{\Omega_{\eta_t}} H_2^t \cdot \nabla \phi \, dx \, dt + \int_I \int_{\Omega_{\eta_t}} h_t \phi \, dx \, dt
\]

for all \( \phi \in C_c^\infty(I \times \Omega_{\eta_t}) \).

In \[ \text{[6]} \) Lemma 2.8 the corresponding version of [A2] assumes \( \alpha = 1 \). But the very same argument is also valid in case \( \alpha \in (0, 1) \) due to compact embeddings for fractional Sobolev spaces.
Lemma 2.11. Let \((\eta_l), (v_l)\) and \((r_l)\) be sequences satisfying (A1) \((A3)\) where \(\frac{1}{2} + \frac{1}{p} = \frac{1}{q} < 1\) (with \(s^* = \frac{3q}{2q-3}\)) if \(s \in (1, 3/\alpha)\) and \(s^* \in (1, \infty)\) arbitrarily otherwise and \(\frac{1}{m} + \frac{1}{p} = \frac{1}{q} < 1\). Then there is a subsequence with

\[
(2.15) \quad v_l, r_l \rightharpoonup v, r \text{ weakly in } L^q(I, L^p(\Omega_{\eta_l})).
\]

Corollary 2.12. In the case \(r_l = v_l\) we find that

\[
v_l \rightharpoonup v \text{ strongly in } L^2(I, L^2(\Omega_{\eta_l})).
\]

We finish this section by repeating the following Aubin-Lions type lemma which is shown in [39] Theorem 5.1. & Remark 5.2.]

Theorem 2.13. Let \(X, Z\) be two Banach spaces, such that \(X' \subset Z'\). Assume that \(f_n : (0, T) \to X\) and \(g_n : (0, T) \to X'\), such that \(g_n \in L^\infty(0, T; Z')\) uniformly. Moreover assume the following:

(a) The weak convergence: for some \(s \in [1, \infty)\) we have that \(f_n \rightharpoonup f\) in \(L^s(X)\) and \(g_n \rightharpoonup^* g\) in \(L^{s'}(X')\).

(b) The approximability-condition is satisfied: For every \(\kappa \in (0, 1)\) there exists a \(f_{n, \kappa} \in L^s(0, T; X) \cap L^1(0, T; Z)\), such that for every \(\varepsilon \in (0, 1)\) there exists a \(\kappa_\varepsilon \in (0, 1)\) (depending only on \(\varepsilon\)) such that

\[
\|f_n - f_{n, \kappa}\|_{L^s(0, T; X)} \leq \varepsilon \text{ for all } \kappa \in (0, \kappa_\varepsilon]
\]

and for every \(\kappa \in (0, 1)\) there is a \(C(\kappa)\) such that

\[
\|f_{n, \kappa}\|_{L^1(0, T; Z)} dt \leq C(\kappa).
\]

Moreover, we assume that for every \(\kappa\) there is a function \(f_\kappa\), and a subsequence such that \(f_{n, \kappa} \rightharpoonup f_\kappa\) in \(L^s(0, T; X)\).

(c) The equi-continuity of \(g_n\). We require that there exists an \(\alpha \in (0, 1]\) a functions \(A_\alpha\) with \(A_n \in L^1(0, T)\) uniformly, such that for every \(\kappa > 0\) that there exist a \(C(\kappa) > 0\) and an \(n_\kappa \in \mathbb{N}\) such that for \(\tau > 0\) and a.e. \(t \in [0, T - \tau]\)

\[
\sup_{n \geq n_\kappa} \left| \int_0^\tau \langle g_n(t) - g_n(t + s), f_{n, \kappa}(t) \rangle_{X', X} ds \right| \leq C(\kappa)\tau^\alpha(A_n(t) + 1).
\]

(d) The compactness assumption is satisfied: \(X' \rightrightarrows Z'\). More precisely, every uniformly bounded sequence in \(X'\) has a strongly converging subsequence in \(Z'\).

Then there is a subsequence, such that

\[
\int_0^T \langle f_n, g_n \rangle_{X, X'} \, dt \rightharpoonup \int_0^T \langle f, g \rangle_{X, X'} \, dt.
\]

2.5. Weak solutions and main theorem. In accordance with the current state of the art for weak solutions of Navier-Stokes-Fourier law fluids and fluid-structure interactions, we introduce here our concept of weak solutions. For that we introduce the following function spaces, where \(D(u) = \frac{1}{2}(\nabla u + \nabla u^T)\) denotes the symmetric gradient of a given function:

(S1) For the solid deformation \(\eta : I \times \omega \to \mathbb{R}, Y' : = \{\zeta \in W^{1, \infty}(I; L^2(\omega)) \cap L^\infty(I; W^{2,2}(\omega))\}\).

(S2) For the fluid velocity \(u : I \times \Omega_{\eta} \to \mathbb{R}^d, d = 2, 3, X'_H : = \{u \in L^2(I; L^2(\Omega_{\eta})) : D(u) \in L^2(I; L^2(\Omega_{\eta}))\}\).

(S3) For the fluid density \(\rho : I \times \Omega_{\eta} \to [0, \infty), W^1_H : = C_w(I; L^1(\Omega_{\eta})),\) where the subscript \(w\) refers to continuity with respect to the weak topology.

(S4) For the temperature \(\vartheta : I \times \Omega_{\eta} \to [0, \infty)

\[
Z^I_\eta = \{\vartheta \in L^2(I; W^{1,2}(\Omega_{\eta})) \cap L^\infty(I; L^4(\Omega_{\eta})) : \log(\vartheta) \in L^2(W^{1,2}(\Omega_{\eta}))\}.
\]

The definition of the function spaces above depending on \(\eta\) only make sense provided \(\|\eta\|_{L^\infty_{\infty} \omega} < L\).

Definition 2.14. A weak solution to (1.1) \((\eta, u, \rho, \vartheta)\) is a quadruplet \((\eta, u, \rho, \vartheta) \in \times Y' \times X'_H \times W^1_H \times Z^I_\eta\) which satisfies the following.
The momentum equation is satisfied in the sense that
\[
\begin{aligned}
&\int_I \frac{d}{dt} \int_{\Omega_n} \rho \mathbf{u} \cdot \mathbf{\phi} \, dx \, dt - \int_I \int_{\Omega_n} \left( \rho \mathbf{u} \cdot \partial_t \mathbf{\phi} + \rho \mathbf{u} \otimes \mathbf{u} : \nabla \mathbf{\phi} \right) \, dx \, dt \\
&+ \int_I \int_{\Omega_n} \mathbf{S}(\rho, \nabla \mathbf{u}) : \nabla \mathbf{\phi} \, dx \, dt - \int_I \int_{\Omega_n} p(\rho, \theta) \, \text{div} \mathbf{\phi} \, dx \, dt \\
&+ \int_I \left( \frac{d}{dt} \int_{\Omega_n} d \mathbf{\eta} \cdot b \, dy - \int_{\Omega_n} \partial_t \mathbf{\eta} \cdot \mathbf{b} \, dy + \int_{\Omega_n} K'(\mathbf{b}) \, dy \right) \, dt \\
&= \int_I \int_{\Omega_n} \rho \mathbf{f} \cdot \mathbf{\phi} \, dx \, dt + \int_I \int_{\Omega_n} g \mathbf{\phi} \, dx \, dt
\end{aligned}
\]
holds for all \((h, \mathbf{\phi}) \in C^\infty(\omega) \times C^\infty(T \times \mathbb{R}^3)\) with \(\text{tr}_\omega \mathbf{\phi} = \text{div} \mathbf{v}\). Moreover, we have \((\rho \mathbf{u})(0) = \rho_0, \mathbf{v}(0) = \mathbf{v}_0\) and \(\partial_t \mathbf{\eta}(0) = \mathbf{\eta}_1\). The boundary condition \(\text{tr}_\omega \mathbf{u} = \partial_t \mathbf{\nu}\) holds in the sense of Lemma 2.3.

The continuity equation is satisfied in the sense that
\[
\int_I \frac{d}{dt} \int_{\Omega_n} \rho \mathbf{u} \cdot \mathbf{\psi} \, dx \, dt - \int_I \int_{\Omega_n} \left( \rho \partial_t \mathbf{\psi} + \rho \mathbf{u} \cdot \nabla \mathbf{\psi} \right) \, dx \, dt = 0
\]
holds for all \(\psi \in C^\infty(T \times \mathbb{R}^3)\) and we have \(\rho(0) = \rho_0\).

The entropy balance
\[
\begin{aligned}
&\int_I \frac{d}{dt} \int_{\Omega_n} g s(\rho, \theta) \mathbf{\psi} \, dx \, dt - \int_I \int_{\Omega_n} \left( g s(\rho, \theta) \partial_t \mathbf{\psi} + g s(\rho, \theta) \mathbf{u} \cdot \nabla \mathbf{\psi} \right) \, dx \, dt \\
&\geq \int_I \int_{\Omega_n} \frac{1}{\theta} \left( \mathbf{S}(\rho, \nabla \mathbf{u}) : \nabla \mathbf{\psi} + \frac{\gamma(\theta)}{\theta} | \nabla \mathbf{\psi} |^2 \right) \mathbf{\psi} \, dx \, dt \\
&+ \int_I \int_{\Omega_n} \frac{\gamma(\theta) \nabla \theta}{\theta} \cdot \nabla \mathbf{\psi} \, dx \, dt + \int_I \int_{\Omega_n} \frac{\theta}{\gamma} | \mathbf{\nabla} \mathbf{\psi} |^2 \, dx \, dt
\end{aligned}
\]
holds for all \(\psi \in C^\infty(T \times \mathbb{R}^3)\) with \(\psi \geq 0\); in particular, all integrals above are well defined. Moreover, we have \(\lim_{t \to 0} g s(\rho, \theta)(t) = \rho_0 s(\rho_0, \theta_0)\) and \(\partial_t \rho, \theta|_{\partial \Omega_n} \leq 0\).

The total energy balance
\[
\begin{aligned}
&- \int_I \partial_t \mathbf{\psi} \mathbf{E} \, dt = \rho_0 \mathbf{E}(0) + \int_I \int_{\Omega_n} g \mathbf{u} \, dx \, dt + \int_I \int_{\Omega_n} \rho \mathbf{f} \cdot \mathbf{\psi} \, dx \, dt \\
&+ \int_I \psi \int_{\omega} \theta \partial_t \mathbf{\eta} \, dy \, dt
\end{aligned}
\]
holds for any \(\psi \in C^\infty([0, T))\). Here, we abbreviated
\[
\mathbf{E}(t) = \int_{\Omega_n(t)} \left( \frac{1}{2} g(t)| \mathbf{u}(t) |^2 + \rho(t) e(\rho(t), \theta(t)) \right) \, dx + \int_{\omega} \frac{| \partial_t \mathbf{\eta}(t) |^2}{2} \, dy + K(\eta(t)).
\]

As will be apparent by the analysis we shall show that the renormalized continuity equation in the sense of DiPerna and Lions is satisfied, cf. [15, 32].

**Definition 2.15** (Renormalized continuity equation). Let \(\eta \in Y^T\) and \(\mathbf{u} \in X^T\). We say that the function \(\rho \in W^1_T\) solves the continuity equation \([4]\) in the renormalized sense if we have
\[
\int_I \frac{d}{dt} \int_{\Omega_n} \theta(\rho) \mathbf{\psi} \, dx \, dt - \int_I \int_{\Omega_n} \left( \partial_t \mathbf{\eta} \psi + \theta(\rho) \mathbf{u} \cdot \nabla \mathbf{\psi} \right) \, dx \, dt
\]
holds for all \(\psi \in C^\infty(T \times \mathbb{R}^3)\) and all \(\theta \in C^1(\mathbb{R})\) with \(\theta(0) = 0\) and \(\theta'(z) = 0\) for \(z \geq M_\theta\).

We are now ready to formulate our main result.

**Theorem 2.16.** Let \(\gamma > \frac{\alpha}{\theta}\) \((\gamma > 1\) in two dimensions). Assume that we have
\[
\frac{| \rho \mathbf{u} |^2}{\rho_0} \in L^1(\Omega_{\eta_0}), \; \rho_0 \in L^1(\Omega_{\eta_0}), \; \mathbf{v}_0 \in L^1(\Omega_{\eta_0}), \; \eta_0 \in W^{2,2}(\omega), \; \eta_1 \in L^2(\omega),
\]

f ∈ L^2(I; L^∞(R^3)), g ∈ L^2(I × ω), H ∈ L^∞(I × R^3), H ≥ 0 a.e.

Furthermore suppose that ϑ_0 > 0 a.e., ϑ_0 > 0 a.e. and that (1.10) is satisfied. Then there exists a weak solution (η, u, ϑ, ν) to (1.12) in the sense of Definition 2.14. The interval of existence is of the form \( I = (0, t) \), where \( t < T \) only in case \( \Omega_{(s)} \) approaches a self-intersection when \( s \to t \) or the Koiter energy degenerates (namely, if \( \lim_{s \to t} S(s, y) = 0 \) for some point \( y \in ω \)). Moreover, the continuity equation is satisfied in the renormalized sense as specified in Definition 2.17.

Remark 2.17 (Minimal interval of existence). Let us mention that for any admissible initial conditions there is a minimal positive interval of existence. It follows from the fact that \( ϑ \) (and consequently also \( γ \), cf. (1.13)) can be shown to be uniformly continuous in space-time (with bounds depending on the data only). Consequently, for some non-empty open time-interval no self-touching or point of degeneracy can be approached a-priori.

Remark 2.18 (Simplification of notation). We remark that we will assume without further mentioning that the initial conditions for the elastic deformation are within a neighbourhood of the reference configurations. This simplification is, however, without loss of generality. Indeed, by rephrasing the reference geometry accordingly, the existence procedure can be prolonged until a point of self touching or degeneracy (in case of non-linear Koiter energies) is approached.

Remark 2.19 (Properties of the solution). As can be seen from the proof, in particular (5.5), we can control in addition to the energy the quantity

\[
σ = \frac{1}{\varrho} S(\varrho, \nabla u) : \nabla u + \frac{x}{\varrho^2} |\nabla \varrho|^2
\]

in \( L^1 \). This implies that

(a) The symmetric gradient \( D(u) \) belongs to \( L^2 \) due to (4.7)–(4.10). Since the domain is not Lipschitz continuous (at least not uniformly in time) standard results on Korn-type inequalities do not apply. In our context of domains with less regularity, a corresponding inequality is shown in [30] Prop. 2.9 following ideas of [1]. The integrability of the full gradient is, however, less than the one of the symmetric gradient, that is we only have \( \nabla u \in L^q(I, \Omega_η) \) for some non-empty open time-interval no self-touching or point of degeneracy can be approached a-priori.

(b) All involved integrals in (O3) are finite. In particular, the temperature satisfies \( \nabla \log \varrho \in L^2 \) using (2.7). Further one can deduce that \( \log \varrho \in L^2 \) following [21] Section 2.2.4]. This implies \( \varrho > 0 \) a.a. in \( I × Ω_η \).

3. Equations for density and temperature in variable domains

In this section we study the continuity equation (with artificial viscosity) as well as the internal energy equation in variable domains. In Theorems 3.3 and 3.4 we prove the existence of classical solutions to both equations under the assumption the data (the velocity field as well and the variable boundary) are smooth. In particular, we prove that the temperature stays strictly positive on the regularised level. This is a key ingredient for the remainder of the paper.

3.1. The continuity equation.

In this subsection we are concerned with the regularised continuity equation in a (given) variable domain. We assume that the moving boundary is prescribed by a function \( ζ : T × ω → R \). For a given function \( w ∈ L^2(I; W^{1,2}(Ω_ζ)) \) with \( \text{tr}_ω w = \partial_ω ν \) and \( ε > 0 \) we consider the equation

\[
\partial_ω ν + \text{div}(ρw) = ε \Delta ρ \quad \text{in} \quad I × Ω_ζ \quad (3.1)
\]

\[
ρ(0) = ρ_0 \text{ in } Ω_ζ(0), \quad \partial_ω ρ|_{∂Ω_ζ} = 0 \quad \text{on} \quad I × ∂Ω_ζ.
\]

A weak solution to (3.1) satisfies

\[
\int_I \frac{d}{dt} \int_{Ω_ζ} ψ dV dt - \int_I \int_{Ω_ζ} \left( ρ \partial_ω ψ + ρw · ∇ψ \right) dV dt = - \int_I \int_{Ω_ζ} ε \nabla \varrho · \nabla ψ dV dt
\]

for all \( ψ ∈ C^∞(T × R^3) \). The following result has been proved in [3] Thm. 3.1.(for the analogous results for fixed in time domains see [22] section 2.1).

Theorem 3.1. Let \( ζ ∈ C^{2,α}(T × ω, [-\frac{1}{2}, \frac{1}{2}]) \) with \( α \in (0, 1) \) be the function describing the boundary. Assume that \( w ∈ L^2(I; W^{1,2}(Ω_ζ)) \cap L^∞(I × Ω_ζ) \) with \( \text{tr}_ω w = \partial_ω ν \) and \( ρ_0 ∈ L^2(Ω_ζ(0)) \).

a) There is a unique weak solution \( ρ \) to (3.1) such that

\[
ρ ∈ L^∞(I; L^2(Ω_ζ)) ∩ L^2(I; W^{1,2}(Ω_ζ)).
\]
b) Let \( \theta \in C^2(\mathbb{R}_+; \mathbb{R}_+) \) be such that \( \theta'(s) = 0 \) for large values of \( s \) and \( \theta(0) = 0 \). Then the following holds, for the canonical zero extension of \( \varrho \equiv \varrho_{\Omega_0}^c \):

\[
\int_I \frac{d}{dt} \int_{\Omega} \varrho \psi \, dx \, dt - \int_{I \times \Omega} \varrho(\varrho) \partial_t \psi \, dx \, dt = - \int_{I \times \Omega} (\varrho'^{\theta}(\varrho) - \varrho(\varrho)) \, div \psi \, dx \, dt + \int_{I \times \Omega} \varrho(\varrho) \omega \cdot \nabla \psi \, dx \, dt
\]

(3.3)

for all \( \psi \in C^{\infty}(\tilde{T} \times \mathbb{R}^3) \).

c) Assume that \( \varrho_0 \geq 0 \) a.e. in \( \Omega_0(0) \). Then we have \( \varrho \geq 0 \) a.e. in \( I \times \Omega_0 \).

**Remark 3.2.** Observe that:

- The statement in [6] holds without the assumption \( tr_s \varrho \) under the boundary condition \( \partial_\nu \varrho |_{\partial \Omega_0} = \frac{1}{\nu} \varrho(\varrho - (\partial_\nu \varrho) \circ \varrho^{-1}) \cdot \nu \).

- Theorem 3.1 in [6] is formulated with the stronger assumption \( \varrho \in C^2(\tilde{T} \times \Omega; [-\frac{1}{2}, \frac{1}{2}]) \). However, it can be checked that the condition \( \varrho \in C^{2,0}(\tilde{T} \times \Omega; [-\frac{1}{2}, \frac{1}{2}]) \) is sufficient for the proof.

In the following we improve the result from Theorem 3.1 and obtain a classical solution to \( \mathfrak{3} \).

**Theorem 3.3.** Let the assumptions of Theorem 3.1 be satisfied and suppose additionally that \( \partial_\nu \nabla \varrho \) as well as \( \nabla^2 \varrho \) belong to the class \( C^{\infty}(\tilde{T} \times \Omega) \). Furthermore we assume that \( J_{\varrho} := det \nabla \varrho \) is strictly positive, \( \varrho_0 \in C^{2,\alpha}_{\text{loc}}(\Omega_0(0)) \) and \( \varrho \in C^{1,\alpha}_{\text{loc}}(\tilde{T} \times \Omega) \) such that \( \partial_\nu \nabla \varrho \) and \( \nabla^2 \varrho \) belong to the class \( C^{\alpha}(\tilde{T} \times \Omega) \).

(a) The solution \( \varrho \) from Theorem 3.1 satisfies (3.1) in the classical sense and belongs to the regularity class

\[
\mathcal{L}^1 := \{ z \in C^{1}(\tilde{T} \times \Omega) : \nabla^2 z \in C(\tilde{T} \times \Omega) \}.
\]

In particular, we have

\[
\| \varrho \|_{C^1_{t,x}} + \| \nabla^2 \varrho \|_{C^{0}_{t,x}} \leq c(\varrho_0, \varrho, J_{\varrho}^{-1}, \varrho_0),
\]

with dependence via the (semi-)norms in the affirmative function spaces.

(b) Suppose that \( \varrho_0 \geq 0 \). Then we have the estimate

\[
C^{-1} \min_{\Omega_0(0)} \varrho_0 \leq \max_{I \times \Omega_0} \varrho \leq C \max_{\Omega_0(0)} \varrho_0,
\]

where \( C = C(\varrho, J_{\varrho}^{-1}, \varrho_0) \) with dependence via the (semi-)norms in the affirmative function spaces.

**Proof.** We start by transforming (3.2) to the reference domain. For \( \varrho = \varrho \circ \varrho_0, \varrho_0 \) in \( C^{\infty}(\tilde{T} \times \mathbb{R}^3) \) we set \( \varrho = \varrho \circ \varrho_0^{-1} \).

Defining similarly \( \varrho = \varrho \circ \varrho_0^{-1} \) and \( \varrho = \varrho \circ \varrho_0^{-1} \), we obtain from (3.2)

\[
\int_I \frac{d}{dt} \int_{\Omega} \varrho \circ \varrho_0^{-1} \varrho \circ \varrho_0^{-1} \, dx \, dt = \int_I \int_{\Omega} \varrho \circ \varrho_0^{-1} \varrho \circ \varrho_0^{-1} \partial_\nu (\varrho \circ \varrho_0^{-1}) \, dx \, dt
\]

such that

\[
\int_I \frac{d}{dt} \int_{\Omega} J_{\varrho} \varrho \, dx \, dt = \int_I \int_{\Omega} \varrho \partial_\nu \varrho \circ \varrho_0^{-1} \circ \varrho_0^{-1} \partial_\nu \varrho_0^{-1} \, dx \, dt
\]

such that

\[
\int_I \frac{d}{dt} \int_{\Omega} J_{\varrho} \varrho \, dx \, dt = \int_I \int_{\Omega} \varrho \partial_\nu \varrho \circ \varrho_0^{-1} \circ \varrho_0^{-1} \partial_\nu \varrho_0^{-1} \, dx \, dt
\]

where \( J_{\varrho} = det \nabla \varrho \). Finally, we replace \( \varrho \) by \( \varrho \circ J_{\varrho}^{-1} \) to obtain

\[
\int_I \frac{d}{dt} \int_{\Omega} \varrho \, dx \, dt = \int_I \int_{\Omega} \varrho_0^{-1} \partial_\nu \varrho \circ \varrho_0^{-1} \circ \varrho_0^{-1} \partial_\nu \varrho_0^{-1} \, dx \, dt
\]
Now we set
\[ g_\varsigma = J_\varsigma \partial_t J_\varsigma^{-1} + J_\varsigma \nabla J_\varsigma^{-1} \cdot \partial_t \Psi_\varsigma^{-1} \circ \Psi_\varsigma + \nabla J_\varsigma^{-1} \cdot \nabla J_\varsigma^{-1} \]
\[ g_\varsigma = -\varepsilon J_\varsigma (\nabla \Psi_\varsigma)^{-T} (\nabla \Psi_\varsigma)^{-1} \nabla J_\varsigma^{-1}, \]
\[ f_\varsigma = \partial_t \Psi_\varsigma^{-1} \circ \Psi_\varsigma + (\nabla \Psi_\varsigma)^{-T} \nabla \varsigma, \quad A_\varsigma = \varepsilon (\nabla \Psi_\varsigma)^{-T} (\nabla \Psi_\varsigma)^{-1}, \]
such that the equation reads as
\[ -\int_I \int_\Omega \nabla \partial_t \Psi \cdot \nabla \varsigma \, dx \, dt = \int_I \int_\Omega \nabla \varsigma \cdot \nabla \Psi \, dx \, dt + \int_I \int_\Omega \nabla \cdot \nabla \varsigma \cdot \nabla \Psi \, dx \, dt \]
for any \( \varsigma \) with \( \varsigma(0) = \varsigma(T) = 0 \). Choosing \( \varsigma \in C^\infty_c(I \times \Omega) \) arbitrarily we obtain
\[ \partial_t \Psi = \nabla \varsigma \cdot \nabla \varsigma, \quad \partial_t (\Psi_\varsigma \circ \varsigma^{-1}) = \partial_t (\Psi_\varsigma \circ \varsigma^{-1} + \nabla \varsigma^T \circ \varsigma^{-1} \partial_t \varsigma^{-1} \]
such that
\[ \nabla \varsigma^T \partial_t \varsigma^{-1} \circ \varsigma = -\partial_t \varsigma^{-1} \circ \varsigma = -\partial_t (\varsigma \circ \varsigma^{-1}) = -\varphi = -w \circ \varsigma^{-1} = -\nabla \varsigma \]
on \( I \times \partial \Omega \) due to the definition of \( \nabla \varsigma \) from \( \text{[210]} \).
We can rewrite the equation further as
\[ \partial_t \varsigma = \nabla (g_\varsigma - \nabla f_\varsigma) + \nabla \cdot (g_\varsigma - f_\varsigma) + \nabla A_\varsigma \cdot \nabla \varsigma + A_\varsigma : \nabla^2 \varsigma \]
such that we finally obtain
\[ \partial_t \varsigma + b_\varsigma(t, x, \varsigma, \nabla \varsigma) = A_\varsigma : \nabla^2 \varsigma \quad \text{in} \quad I \times \Omega, \]
\[ \nu A_\varsigma \cdot \nabla \varsigma = 0 \quad \text{on} \quad I \times \partial \Omega, \]
where
\[ b_\varsigma(t, x, u, \varsigma) = -u (g_\varsigma - \nabla f_\varsigma) + U \cdot (f_\varsigma - g_\varsigma - \nabla A_\varsigma). \]
By the classical theory from \([28]\) Thm.‘s 7.2, 7.3 & 7.4, Chapter V] the claim of part (a) follows if we can control the following quantities:

- The \( C^{2,\alpha} \)-norm of \( g_\varsigma \);
- The \( \alpha \)-Hölder-semi-norms of \( \nabla_x b_\varsigma \), \( \partial_t b_\varsigma \) and \( \partial_U b_\varsigma \) with respect to \( x \); the constants in
  \[ -u b_\varsigma(t, x, u, \varsigma) \leq c_0 u^2 + c_1 |U|^2 + c_2 \quad \forall (t, x, u, \varsigma) \in I \times \Omega \times \mathbb{R} \times \mathbb{R}^3; \]
  \[ |b_\varsigma(t, x, u, \varsigma) + |\nabla_x b_\varsigma(t, x, u, \varsigma)| \cdot (1 + |U|)|\nabla_U b_\varsigma(t, x, u, \varsigma)| \| \leq c_3 (1 + u^2 + |U|^2); \]
- The coercivity constant of \( A_\varsigma \) and its upper bound; the \( \alpha \)-Hölder-semi-norm of \( \nabla_x A_\varsigma \) and \( \partial_t A_\varsigma \) with respect to \( x \).

Note that this gives the required regularity for \( \varsigma \); however, transforming back by means of \( \varsigma^{-1} \) does not alter it due to the regularity of \( \varsigma \).
Let us assume that there is a point \((v)\). The reason is that we do not know a priori if

\[ \lambda \]

By (3.5) this implies

\[ A \]

Let us now assume that \(A\) is chosen independently of \(\lambda\) (which we will fix below). We have by

\[ \frac{\nabla \varphi \cdot A \nabla \varphi}{\varphi} < 0 \quad \text{on} \quad T \times \partial \Omega. \]

Such a function \(\varphi\) can be defined using the distance function to the boundary with respect to the direction \(A \nu\). By the assumption that \(\|\zeta\|_{L^\infty} \leq \frac{T}{2}\) and the \(C^2\) regularity of \(\zeta\) this is a well defined function. Note that \(\varphi\) is chosen independently of \(\lambda\) (which we will fix below). We have by

\[ \partial_t v = \varphi e^{-\lambda t} \partial_t \varphi - \lambda v = -\varphi e^{-\lambda t} b_\zeta(t, x, \varphi) + \varphi e^{-\lambda t} A \zeta : \nabla^2 \varphi - \lambda v \]

\[ = -b_\zeta(t, x, \varphi) e^{-\lambda t} \nabla \varphi + \varphi e^{-\lambda t} A \zeta : \nabla^2 \varphi - \lambda v \]

\[ = -b_\zeta(t, x, \varphi) \nabla \varphi - \varphi e^{-\lambda t} \varphi + A \zeta : \nabla^2 \varphi - A \zeta : \varphi e^{-\lambda t} \varphi - \lambda v. \]

Let us assume that there is a point \((t_0, x_0) \in T \times \Omega\) with \(v(t_0, x_0) = \max_{t, x} v(t, x)\). We obtain in this point

\[ 0 = -b_\zeta(t, x, \varphi) \nabla \varphi + A \zeta : \nabla^2 \varphi - A \zeta : \varphi e^{-\lambda t_0} \varphi - \lambda v \]

\[ \leq -b_\zeta(t, x, \varphi) \nabla \varphi + \varphi e^{-\lambda t_0} A \zeta : \left( 2 \frac{\nabla \varphi}{\varphi} \otimes \nabla \varphi - \varphi \nabla \varphi - \varphi \right) - \lambda v \]

\[ = \varphi e^{-\lambda t_0} \left( \varphi (g_\zeta - \text{div} f_\zeta) - \lambda v + \nabla \varphi \cdot (g_\zeta - f_\zeta + \text{div} A \zeta) + A \zeta : \left( 2 \frac{\nabla \varphi}{\varphi} \otimes \nabla \varphi - \varphi \right) \right). \]

If we choose \(\lambda\) large (depending on \(\|g_\zeta\|_{L^\infty}, \|g_\zeta\|_{L^\infty}, \|\text{div} f_\zeta\|_{L^\infty}, \|\text{div} A \zeta\|_{L^\infty}\) and \(\varphi\) this leads to a contradiction (note that \(\varphi\) is non-negative by Theorem 3.6 (b)).

Let us now assume that \(x_0 \in \partial \Omega\) and \(t > 0\). Then since \(A \zeta(t_0, x_0) \nu(x_0)\) points outside \(\Omega\) we have

\[ 0 \leq \frac{d}{ds} v(t_0, x_0 + s A \zeta(t_0, x_0) \nu(x_0)) \bigg|_{s=0} = \nabla v(t_0, x_0) \cdot A \zeta(t_0, x_0) \nu(x_0). \]

By this implies

\[ 0 \leq e^{-\lambda t_0} (\nabla \varphi(x_0) \cdot A \zeta(t_0, x_0) \nu(x_0)) + \varphi \nabla \varphi(x_0) \cdot A \zeta(t_0, x_0) \nu(x_0) \]

\[ \|\varphi\|_{L^\infty} \|A \zeta\|_{L^\infty} \|\nabla \varphi\|_{L^\infty} \|\nabla A \zeta\|_{L^\infty} \varphi(x_0) \]

which yields a contradiction by 3.7. We conclude that the maximum of \(v\) is attained at \((0, x_0)\) for some \(x_0 \in \Omega\). By 3.6 the estimate for the maximum follows.

Unfortunately, the approach above used for the maximum principle does not work to achieve a miniminum principle. The reason is that we do not know a priori if \(\varphi\) is strictly positive at a potential maximum at \(v\) by the boundary. We multiply (3.3) by \(-m(\xi + \bar{\eta})^{-m-1}\) where \(0 < \xi \ll 1\) and \(m \gg 1\). This yields

\[ \partial_t (\xi + \bar{\eta})^{-m} = -m(g_\zeta - \text{div} f_\zeta) \xi + \bar{\eta} \bar{\eta}^{-m-1} \xi + \bar{\eta}^{-m-1} - m \bar{\eta} \cdot (g_\zeta - f_\zeta) (\xi + \bar{\eta})^{-m-1} \]

\[ - m \text{div} \left( A \zeta \nabla \bar{\eta} \right) (\xi + \bar{\eta})^{-m-1} \]

and

\[ \frac{d}{dt} \int_{\Omega} (\xi + \bar{\eta})^{-m} dx + m(m - 1) \int_{\Omega} (\xi + \bar{\eta})^{-m} A \zeta \left( \nabla \bar{\eta}, \nabla \bar{\eta} \right) dx \]

\[ = -m \int_{\Omega} (g_\zeta - \text{div} f_\zeta) \xi + \bar{\eta} \bar{\eta}^{-m-1} dx - m \int_{\Omega} (g_\zeta - f_\zeta) \cdot \nabla \bar{\eta} (\xi + \bar{\eta})^{-m-1} dx = (I) + (II) \]

using 3.3.2. Using the boundedness of \(\varphi\) from (b) we obtain

\[ (I) \leq c m \int_{\Omega} (\xi + \bar{\eta})^{-m-1} dx \leq c m \int_{\Omega} (\bar{\eta})^{-m} dx. \]
The constant $c$ depends on $\|\zeta\|C^2_{t,x}, \sup J_{1}^{-1}$, $\|w\|L^\infty_{t,x}$ and $\|\nabla w\|L^\infty_{t,x}$. Similarly, we have for any $\kappa > 0$

$$(II) \leq c m \int_\Omega |\nabla \bar{\theta}|(\xi + \bar{\theta})^{-m+1} \, dx$$

$$\leq \kappa m (m-1) \int_\Omega |\nabla \bar{\theta}|(\xi + \bar{\theta})^{-m} \, dx + c(\kappa) \int_\Omega |\xi + \bar{\theta}|^{-m} \, dx$$

$$\leq \kappa m (m-1) \int_\Omega A_\xi(\nabla \bar{\theta}, \nabla \bar{\theta})(\xi + \bar{\theta})^{-m} \, dx + c \int_\Omega |\xi + \bar{\theta}|^{-m} \, dx$$

with $c = c(\|\zeta\|C^2_{t,x}, \|g_\zeta\|L^\infty_{t,x}, \|f_\zeta\|L^\infty_{t,x})$. If we absorb the terms containing $A_\xi$ and apply Gronwall’s lemma we obtain

$$\int_\Omega (\xi + \bar{\theta}(t))^{-m} \, dx \leq e^{Cm} \int_\Omega (\xi + \bar{\theta}_0)^{-m} \, dx.$$

The constant $C$ depends on $\|\zeta\|C^2_{t,x}, \|\partial_\zeta \nabla^2 \zeta\|C^2_{t,x}, \|\nabla^3 \zeta\|C^0_{t,x}, \sup J_{1}^{-1}, \|w\|C^1_{t,x}$, $\|\partial_t w\|C^0_{t,x}$ and $\|\nabla^2 w\|C^0_{t,x}$, but is independent of $m$. Taking the $m$-th root shows

$$\left(\int_\Omega \left(\frac{1}{\xi + \bar{\theta}(t)}\right)^{-m} \, dx\right)^{\frac{1}{m}} \leq e^C \left(\int_\Omega (\xi + \bar{\theta}_0)^{-m} \, dx\right)^{\frac{1}{m}}.$$

Passing with $m \to \infty$ implies

$$\sup_\Omega \frac{1}{\xi + \bar{\theta}(t)} \leq e^C \sup_\Omega \frac{1}{\xi + \bar{\theta}_0}$$

or, equivalently,

$$e^C \inf_\Omega (\xi + \bar{\theta}_0) \leq \inf_\Omega (\xi + \bar{\theta}(t)).$$

Consequently, passing with $\xi \to 0$ we have $e^C \inf_\Omega \bar{\theta}_0 \leq \bar{\theta}(t, x)$ for all $(t, x) \in \bar{T} \times \bar{\Omega}$. Thus, transforming back to $\rho$, (c) is shown and the proof is complete. \hfill $\Box$

### 3.2. The internal energy equation.

The artificial viscosity of the mollified continuity equation produces some dissipative forces that will turn into heat. This is captured by the internal energy equation which we will solve next. For that we introduce the artificial dissipation as

$$(3.9) \quad \Phi_\delta(\rho) = (\gamma \rho^{-2} + \delta \rho^{\gamma-2})|\nabla \rho|^2$$

Indeed, we observe that by the renormalized continuity equation (3.3), we find that

$$(3.10) \quad \partial_t (\frac{1}{\rho^{\gamma-1}} \rho^\gamma) + \text{div} (\frac{1}{\rho^{\gamma-1}} \rho^\gamma w) = \Delta (\frac{1}{\rho^{\gamma-1}} \rho^\gamma) - \rho^\gamma \text{div} w - e\gamma \rho^{-2} |\nabla \rho|^2$$

Hence the (regularized) internal energy part $\tilde{e}_R(\rho, \vartheta) := e_R(\rho, \vartheta) + c_v \vartheta = \frac{\tilde{e}_R}{\rho} + c_v \vartheta$ with $\tilde{p}_R(\rho, \vartheta) := p_R(\rho, \vartheta) + \tilde{e}_R$ is required to satisfy

$$(3.11) \quad \partial_t (\tilde{e}_R(\rho, \vartheta)) + \text{div} (\tilde{p}_R(\rho, \vartheta) w) - \text{div} (\tilde{\vartheta}_R(\vartheta) \nabla \vartheta)$$

$$= S^\vartheta(\vartheta, \nabla w) : \nabla w - \tilde{p}_R(\rho, \vartheta) \text{div} w$$

$$+ \varepsilon \Phi_\delta(\rho) + \delta \frac{1}{\rho^{\gamma-2}} - e \rho^{\gamma-2} + \rho H \quad \text{in} \quad I \times \Omega_\zeta,$$

$$\partial_\vartheta \vartheta |_{\partial \Omega_\zeta} = 0 \quad \text{on} \quad I \times \partial \Omega_\zeta$$

and we have $\vartheta(0) = 0$ (note that $\vartheta_R$ and $S^\theta$ are defined in (2.25)).

The equation does indeed uniquely define $\vartheta$ provided $\rho$ exists and is satisfyingly smooth. Certainly $\vartheta$ constructed by Theorem 3.3 does inherit the necessary smoothness. Accordingly and similar to Theorem 3.3 we have the following result concerning a classical solution to (3.11).

**Theorem 3.4.** Let $\zeta \in C^{2,\alpha} (\bar{T} \times \Omega_\zeta, [-\frac{1}{2}, \frac{1}{2}])$ with $\alpha \in (0, 1)$ be the function describing the boundary. Suppose additionally that $\partial_\zeta \nabla^2 \zeta$ as well as $\nabla^3 \zeta$ belong to the class $C^\alpha (\bar{T} \times \Omega)$ and suppose that $J_\zeta := \det \nabla \Psi_\zeta$ is strictly positive. Assume that $w \in C^{1,\alpha} (\bar{T} \times \bar{\Omega}_\zeta)$ such that $\partial_t \nabla w$ and $\nabla^2 w$ belong to the class $C^{\alpha} (\bar{T} \times \bar{\Omega}_\zeta)$ and $\partial_\zeta w = \partial_\zeta \vartheta$. Assume further that $\rho, H \in C^{1,\alpha} (\bar{T} \times \bar{\Omega}_\zeta; [0, \infty))$ with $\nabla^2 \rho \in C^{\alpha} (\bar{T} \times \bar{\Omega}_\zeta)$ that $\partial_\vartheta \vartheta |_{\partial \Omega_\zeta} = 0$ and $\vartheta$ strictly positive on $I \times \partial \Omega_\zeta$. 

(a) There is a unique classical solution $\vartheta$ to \ref{6.11} which belongs to the regularity class

$$Z^I_\ell := \{ z \in C^1(\hat{T} \times \hat{\Omega}_\ell) : \nabla^2 z \in C(\hat{T} \times \hat{\Omega}_\ell) \}.$$  

In particular, we have

$$\| \vartheta \|_{C^1_{t,x}} + \| \nabla^2 \vartheta \|_{C^1_{t,x}} \leq c \left( \vartheta_0, \zeta, \sup J^{-1}_\zeta, w, \varrho, H \right),$$

with dependence via the (semi-)norms in the affirmative function spaces.

(b) We have the estimate

$$\min \left\{ C^{-1} \min \vartheta_0, 1 \right\} \leq \min \vartheta \leq \max \vartheta \leq \max \left\{ C \max \vartheta_0, 1 \right\},$$

where $C = C(\zeta, \sup J^{-1}_\zeta, w, \varrho, H)$ with dependence via the (semi-)norms in the affirmative function spaces.

Proof. Equation \ref{6.11} contains several nonlinear terms which blow up for small or large values of $\vartheta$. Hence we replace them with regularized versions. Let $\chi_{\ell} \in C^\infty([0, \infty))$ with $\chi_{\ell}(Z) = Z$ for $Z \in [1/\ell, \ell]$ and $c \ell^{-1} \leq \chi_{\ell} \leq c \ell$ for some positive constants $c, C$ and $\ell \gg 1$. We also define the function

$$b_{(t,x)}(\vartheta) := \varrho(t,x) \tilde{e}_R(\varrho(t,x), \vartheta) = a\vartheta^4 + c_\ell \varrho(t,x) \vartheta, \quad \vartheta \geq 0.$$

Since $b_{(t,x)}'(\vartheta) = 4a\vartheta^3 + c_\ell \varrho(t,x)$ is strictly positive by the assumptions on $\varrho$ the inverse $b_{(t,x)}^{-1}$ satisfies

$$\left( b_{(t,x)}^{-1} \right)' \leq c(\varrho).$$

We define

$$S^{\ell,f}(t,x,\vartheta, \nabla w) = S^\ell(b_{(t,x)}^{-1}(\chi_{\ell}(b_{(t,x)}(\vartheta))), \nabla w), \quad \chi_{\ell}(b_{(t,x)}(\vartheta)) = \frac{\varepsilon b_{(t,x)}^{-1}(\chi_{\ell}(b_{(t,x)}(\vartheta)))}{b_{(t,x)}^{-1}(\chi_{\ell}(b_{(t,x)}(\vartheta)))},$$

and consider the equation

$$\frac{\partial_{t} b_{(t,x)}(\vartheta)}{\partial_{t} b_{(t,x)}(\vartheta)} + \nabla \cdot (b_{(t,x)}(\vartheta)w) - \nabla \cdot (\chi_{\ell}(b_{(t,x)}(\vartheta))) = S^{\ell,f}(\vartheta, \nabla w) : \nabla w - \tilde{\rho}_{R}(\varrho, b_{(t,x)}^{-1}(\chi_{\ell}(b_{(t,x)}(\vartheta)))) \div w + D_{\delta}(\varrho)$$

\begin{equation}
\frac{\partial b_{(t,x)}^{-1}(\chi_{\ell}(b_{(t,x)}(\vartheta)))}{\partial_{t} b_{(t,x)}(\vartheta)} - \varepsilon(b_{(t,x)}^{-1}(\chi_{\ell}(b_{(t,x)}(\vartheta))))^2 = \varepsilon(b_{(t,x)}^{-1}(\chi_{\ell}(b_{(t,x)}(\vartheta))))^4 + \varrho H \quad \text{in} \quad I \times \Omega_{\zeta},
\end{equation}

and we have $\vartheta(0) = \vartheta_0$. We will show that a solution $\varrho$ to \ref{6.13} exists and that

$$\max_{I \times \Omega_{\zeta}} \vartheta \leq \max \left\{ C \max \vartheta_0, 1 \right\}$$

as well as

$$\min_{I \times \Omega_{\zeta}} \varrho \geq \min \left\{ C^{-1} \min \vartheta_0, 1 \right\}$$

with $C = C(\| \zeta \|_{C^2_{t,x}}, \| \partial \nabla^2 \zeta \|_{C^2_{t,x}}^3, \| \partial \nabla^2 \zeta \|_{C^2_{t,x}}^4, \| \nabla^2 w \|_{C^1_{t,x}}, \| H \|_{L^\infty_{t,x}})$ independent of $\ell$.

Consequently, the cut-offs in \ref{6.13} are not seen for $\ell$ are enough and we obtain the result for the original problem \ref{6.11}. Arguing as in the proof of Theorem \ref{5.3} we can transform \ref{6.13} to the reference domain. For this purpose it is useful to work with the weak formulation

$$\int_{\Omega_{\zeta}} \left( \frac{d}{dt} \int_{\Omega_{\zeta}} b_{(t,x)}(\vartheta) \psi \, dx \right) dt = \int_{\Omega_{\zeta}} \left( b_{(t,x)}(\vartheta) \partial_{t} \psi + b_{(t,x)}(\vartheta) \nabla \psi \cdot \nabla \vartheta \right) dx dt$$

$$+ \int_{I \times \Omega_{\zeta}} \chi_{\ell}(\vartheta) \nabla \vartheta \cdot \nabla \psi \, dx dt$$

$$= \int_{I \times \Omega_{\zeta}} \left[ S^{\ell,f}(\vartheta, \nabla w) : \nabla w - \tilde{\rho}_{R}(\varrho, b_{(t,x)}^{-1}(\chi_{\ell}(b_{(t,x)}(\vartheta)))) \div w \right] \psi \, dx dt$$

$$+ \int_{I \times \Omega_{\zeta}} \left[ D_{\delta}(\varrho) + \delta(b_{(t,x)}^{-1}(\chi_{\ell}(b_{(t,x)}(\vartheta))))^2 - \varepsilon(b_{(t,x)}^{-1}(\chi_{\ell}(b_{(t,x)}(\vartheta))))^4 + \varrho H \right] \psi \, dx dt$$

for all $\psi \in C^\infty(\hat{T} \times \mathbb{R}^3)$. Actually we will solve the PDE for $Z := (b_{(t,x)}(\vartheta)) \circ \Psi_\zeta$, which is defined on the reference configuration and hence a PDE for a cylindrical time-space domain can be considered. Accordingly
we are setting $\psi = \overline{\psi} \circ \Psi$ for some $\overline{\psi} \in C^\infty(\overline{T} \times \mathbb{R}^3)$, $\overline{\psi} = \vartheta \circ \Psi$, $\overline{w} = w \circ \Psi$, $\overline{H} = H \circ \Psi$ and $\overline{\theta} = \theta \circ \Psi$; this is equivalent to

$$
\int \frac{d}{dt} \int_{\Omega_t} Z \circ \Psi^{-1} \overline{\psi} \circ \Psi^{-1} \, dx \, dt \\
- \int \int_{\Omega_t} Z \circ \Psi^{-1} \left( \partial_t \overline{\psi} \circ \Psi^{-1} + \nabla \overline{\psi} \circ \Psi^{-1} \cdot \partial_t \Psi^{-1} \right) \, dx \, dt \\
+ \int \int_{\Omega_t} Z \circ \Psi^{-1} \overline{w} \circ \Psi^{-1} \cdot \nabla \Psi^{-1} \nabla \overline{\psi} \circ \Psi^{-1} \, dx \, dt \\
+ \int \int_{\Omega_t} \kappa^2 (\overline{\psi} \circ \Psi^{-1}) (\nabla \Psi^{-1})^T \nabla \overline{\psi} \circ \Psi^{-1} \cdot \nabla \overline{\psi} \circ \Psi^{-1} \, dx \, dt \\
= \int \int_{\Omega_t} (\nabla \Psi^{-1})^T S^{\epsilon, \ell} (\overline{\psi} \circ \Psi^{-1}, \nabla \Psi^{-1} \nabla \overline{\psi} \circ \Psi^{-1}) \, dx \, dt \\
- \int \int_{\Omega_t} \tilde{p} R (\overline{\theta} \circ \Psi^{-1}, b_{(t,x)}^{-1} (\chi _{\ell}(Z \circ \Psi^{-1})) \nabla \overline{w} : (\nabla \Psi^{-1})^T \nabla \overline{\psi} \circ \Psi^{-1} \, dx \, dt \\
+ \int \int_{\Omega_t} \nabla (\overline{\psi} \circ \Psi^{-1}) \circ \Psi^{-1} \, dx \, dt \\
+ \int \int_{\Omega_t} [ -(\nabla (\overline{b}^{-1} (\chi _{\ell}(Z \circ \Psi^{-1}))))^T + \frac{\Gamma T}{\nabla \overline{\psi} \circ \Psi^{-1}} \, dx \, dt \\
$$

and, setting $J = \det \nabla \Psi$ and $\overline{\Sigma} = (\delta \beta \overline{\gamma}^2 - \gamma \overline{\theta}^{-2}) |\nabla \Psi^{-1} \nabla \overline{\theta}|^2$, we find

$$
\int \frac{d}{dt} \int_{\Omega} J \overline{\psi} \overline{\psi} \, dx \, dt - \int \int_{\Omega} J \left( \partial_t \overline{\psi} + \nabla \overline{\psi} \cdot \partial_t \Psi^{-1} \circ \Psi \right) \, dx \, dt \\
+ \int \int_{\Omega} J \overline{w} \cdot (\nabla \Psi)^{-1} \nabla \overline{\psi} \, dx \, dt + \int \int_{\Omega} J \kappa^2 (\overline{\psi} \circ \Psi^{-1}) (\nabla \Psi)^{-T} \nabla \overline{\psi} \cdot \nabla \overline{\psi} \, dx \, dt \\
= \int \int_{\Omega} J \left( (\nabla \Psi)^{-T} S^{\epsilon, \ell} (\overline{\psi} \circ \Psi^{-1}, \nabla \Psi^{-1} \nabla \overline{\psi} \circ \Psi^{-1}) \right) : \overline{\psi} \overline{\psi} \, dx \, dt \\
- \int \int_{\Omega} J \tilde{p} R (\overline{\theta} \circ \Psi^{-1}, b_{(t,x)}^{-1} (\chi _{\ell}(Z))) \overline{w} : (\nabla \Psi)^{-T} \overline{\psi} \, dx \, dt \\
+ \epsilon \int \int_{\Omega} J \overline{w} \, dx \, dt + \int \int_{\Omega} \delta J \left( b_{(t,x)}^{-1} (\chi _{\ell}(Z)) \right)^{-2} \overline{\psi} \, dx \, dt \\
+ \int \int_{\Omega} J \left[ -(\nabla (b_{(t,x)}^{-1} (\chi _{\ell}(Z))))^5 + \frac{\Gamma T}{\overline{\psi}} \right] \, dx \, dt \\
$$

for all $\overline{\psi} \in C^\infty(\overline{T} \times \overline{\Omega})$. Again we replace $\overline{\psi}$ by $\overline{\psi} / J$ to obtain

$$
\int \frac{d}{dt} \int_{\Omega} Z \overline{\psi} \, dx \, dt - \int \int_{\Omega} Z \partial_t \overline{\psi} \, dx \, dt \\
= \int \int_{\Omega} Z \nabla \overline{\psi} \cdot \partial_t \Psi^{-1} \circ \Psi \, dx \, dt - \int \int_{\Omega} Z \overline{w} : (\nabla \Psi)^{-1} \nabla \overline{\psi} \, dx \, dt \\
- \int \int_{\Omega} \kappa^2 (\overline{\psi} \circ \Psi^{-1}) (\nabla \Psi)^{-T} \nabla \overline{\psi} \, dx \, dt \\
- \int \int_{\Omega} J \kappa^2 (\overline{\psi} \circ \Psi^{-1}) (\nabla \Psi)^{-T} \nabla \overline{\psi} \cdot \nabla J^{-1} \overline{\psi} \, dx \, dt \\
+ \int \int_{\Omega} (\nabla \Psi)^{-T} S^{\epsilon, \ell} (\overline{\psi} \circ \Psi^{-1}, \nabla \Psi^{-1} \nabla \overline{\psi}) : \overline{\psi} \overline{\psi} \, dx \, dt \\
- \int \int_{\Omega} \tilde{p} R (\overline{\theta} \circ \Psi^{-1}, b_{(t,x)}^{-1} (\chi _{\ell}(Z))) \overline{w} : (\nabla \Psi)^{-T} \overline{\psi} \, dx \, dt \\
+ \int \int_{\Omega} \left( b_{(t,x)}^{-1} (\chi _{\ell}(Z)) \right)^{-2} \overline{\psi} \, dx \, dt - \int \int_{\Omega} \varepsilon \left( b_{(t,x)}^{-1} (\chi _{\ell}(Z)) \right)^5 \overline{\psi} \, dx \, dt \\
$$
Recalling \ref{21} and that \( Z = b_{t,(x)}(\Psi) \) we set

\[
g'_\ell(Z) = (\nabla \Psi_\zeta)^{-T} \mathbf{S} : \left( b'_{t,(x)}(\chi(t,(x))) (\nabla \Psi_\zeta)^{-1} \nabla \Psi \right) - \mathbf{P}_R (\Psi, b'_{t,(x)}(\chi(t,(x))))(\nabla \Psi_\zeta)^{-1}
\]

\[+ \varepsilon \delta b_{t,(x)}^{-\frac{1}{2}} |(\nabla \Psi_\zeta)^{-1} \nabla \Psi| + \delta b_{t,(x)}^{-\frac{1}{2}} + \varepsilon b_{t,(x)}^{-2} \left( \varepsilon (b_{t,(x)}(\chi(t,(x))))^5 + \sqrt{H} \right)
\]

\[+ Z \left( \partial_t J^{-1}_\zeta + \nabla J^{-1}_\zeta \cdot \partial_t \Psi^{-1} - \nabla \cdot (\nabla \Psi_\zeta)^{-1} \nabla J^{-1}_\zeta \right) J_\zeta,
\]

\[g'_\ell(Z) = -J_\zeta \left( \frac{\varepsilon \delta}{b'_{t,(x)}(\chi(t,(x)))} \nabla J^{-1}_\zeta (\nabla \Psi_\zeta)^{-T} (\nabla \Psi_\zeta)^{-1}
\]

\[+ \frac{\varepsilon \delta}{b'_{t,(x)}(\chi(t,(x)))} (\nabla \Psi_\zeta)^{-T} (\nabla \Psi_\zeta)^{-1} \nabla \theta(\ell(t,(x))) b_{t,(x)}^{-1}(\chi(t,(x))).
\]

\[f'_\ell(Z) = Z \left( \partial_t \Psi^{-1} - (\nabla \Psi_\zeta)^{-1} \mathbf{v} \right),
\]

\[A'_\ell(Z) = \left( \frac{\varepsilon \delta}{b'_{t,(x)}(\chi(t,(x)))} \right) (\nabla \Psi_\zeta)^{-T} (\nabla \Psi_\zeta)^{-1},
\]

such that the equation becomes

\[ \partial_t Z = g'_\ell(Z) + \nabla Z \cdot g'_\ell(Z) - \text{div}(f'_\ell(Z)) + \text{div}(A'_\ell(Z) \nabla Z) \quad \text{in} \quad I \times \Omega,
\]

\[ \nu \cdot A'_\ell(Z) \nabla Z = 0 \quad \text{on} \quad I \times \partial \Omega,
\]

\[ Z(0) = (\vartheta_0^\partial + c_\nu \vartheta_0^\partial(0)) \circ \Psi \quad \text{in} \quad \Omega.
\]

Here the boundary conditions are a direct consequence of the weak formulation and the fact that \( f'_\ell \cdot \nu = 0 \) on the boundary, by the assumed coupling of \( \mathbf{w} \) and \( \partial_t \zeta \) at the boundary, cf. \ref{23}. We can further rewrite \ref{16} as

\[ \partial_t Z + b'_\ell(t,(x),(Z),(\nabla Z)) = A'_\ell(Z) : \nabla^2 Z \quad \text{in} \quad I \times \Omega,
\]

where

\[ b'_\ell(t,(x),(u),(U)) = -g'_\ell(u) + \text{div}(f'_\ell(u)) + ( - g'_\ell(u) + \partial_t f'_\ell(u) - \text{div}(A'_\ell(u))) \cdot U
\]

\[ - \partial_t A'_\ell(u)(U,U).
\]

Analogous to the proof of Theorem 3.3 (a) we can use the theory from \cite{28} Thm. 7.2, 7.3 & 7.4 Chapter V] to infer the existence of a unique classical solution \( Z \) to \ref{19} with the required regularity. We will below show that the same \( Z \) is actually a solution for all \( \ell \) that are sufficiently large. Observe already, that by defining \( \vartheta \) as the respective reverse transformation of \( Z \) implies the existence of a respective regular \( \vartheta \) claimed in (a) (provided we can show that the cut-offs in \ref{13} are not seen). This follows by showing respective independent upper and lower bounds on \( \vartheta \) which then eventually imply (b), as well. First we bound the maximum of \( Z \) and respectively proof \ref{14}. Here we argue as in the proof of Theorem 3.3 and set

\[ v(t,(x)) := \varphi(x) e^{-\lambda_1 t} Z
\]

where \( \varphi \in C^\infty(\mathcal{T} \times \overline{\Omega}) \) is such that

\[ \varphi(x) \geq 1 \quad \text{in} \quad \mathcal{T} \times \overline{\Omega},
\]

\[ \frac{\nabla \varphi \cdot A'_\ell(Z) \nu}{\varphi} < 0 \quad \text{on} \quad \mathcal{T} \times \partial \Omega.
\]

Note that we have \( \frac{\varepsilon \delta}{b'_{t,(x)}(\chi(t,(x)))} \geq \frac{\varepsilon \delta}{4} \) by \ref{24} such that the coercivity constant of \( A'_\ell(Z) \) can be bounded from below independently of \( \ell \). Consequently, the function \( \varphi \) can also be chosen independently of \( \ell \). In an interior maximum point \((t_0,(x_0)) \in I \times \Omega \) of \( Z \) we have again

\[ 0 = -e^{-\lambda_1 t_0} \varphi (t,(x),(Z),(\nabla Z)) + A'_\ell(Z) : \nabla^2 v - \lambda_1 v
\]

\[ - A'_\ell(Z) : (2e^{-\lambda_1 t_0} Z \nabla \varphi \otimes \nabla \varphi + Z e^{-\lambda_1 t_0} \nabla^2 \varphi).
\]
\[
\begin{align*}
\Rightarrow -e^{-\lambda t} \varphi \mu_C(t, x, Z, -\frac{Z}{\varphi} \nabla \varphi) - \lambda_1 v + c(\varphi) e^{-\lambda_1 t_0} (1 + Z) \\
\leq c e^{-\lambda_1 t_0} (1 + Z) + \frac{\delta \varphi e^{-\lambda_1 t_0}}{(b_{t,x}^{-1}(\chi_t(Z)))^2} - \lambda_1 e^{-\lambda_1 t_0} \varphi Z,
\end{align*}
\]

where

\[
(3.23) \quad c = c(\varphi, \|\xi\|_{C^{1,\alpha}_t}, \|\partial_t \nabla Z, \nabla^3 \zeta\|_{C^{1,\alpha}_t}, |\zeta|_{C^{1,\alpha}_t}, \|\omega\|_{C^{1,\alpha}_t}, \|\nabla^2 \omega\|_{C^{1,\alpha}_t}, \|H\|_{L^2})
\]
is independent of $\ell$. Note that we used that the coefficients in the definition of $b_\xi$ have linear growth uniformly in $\ell$ except for $\delta(b_{t,x}^{-1}(\chi_t(u)))^{-2}$, $-c(b_{t,x}^{-1}(\chi_t(u)))^5$ and $\partial_t A\xi(u)(U, U)$.

Fortunately, the first two terms have the correct sign, whereas the second one is evaluated at $U = -\nabla \nabla \varphi$. Now we distinguish two cases. If $Z(t_0, x_0) \leq 1$ there is nothing to show. Otherwise, $Z(t, x) \geq 0$ is bounded (independent of $\ell$) such that we obtain a contradiction in (3.22) by choosing $\lambda_1$ large (depending on the quantities in (3.23)). The case $x_0 \in \partial \Omega$ and $t_0 > 0$ can be ruled out again as in (3.3). Hence (3.12) follows with a constant independent of $\ell$.

In order to prove (3.15) we first establish a lower bound which depends on $\ell$. Choosing first $\ell$ large enough and then $Z \in (0, \inf \Omega_0)$ small enough (depending on $\ell$) we have $g_\xi(Z) - \text{div} (f_\xi(Z)) \geq 0$. This thanks to the term $\delta \chi(Z)^{-1/2}$ in the definition of $g_\xi$. Consequently, we obtain from (3.16)

\[
\partial_t a(Z - Z) \geq g_\xi(Z) - g_\xi(Z) + \nabla Z \cdot g_\xi(Z) - \nabla Z \cdot g_\xi(Z) - \text{div} (f_\xi(Z) - f_\xi(Z))
\]

Multiplying by $(Z - Z)^-\alpha$ and integrating over $\Omega$ implies

\[
\frac{d}{dt} \int_\Omega (Z - Z)^-\alpha dx + \int_\Omega A_\xi(Z)(\nabla(Z - Z)^-\alpha, \nabla(Z - Z)^-\alpha) dx
\]

\[
\leq \int_\Omega (g_\xi(Z) - g_\xi(Z))(Z - Z)^-\alpha dx + \int_\Omega (\nabla Z \cdot g_\xi(Z) - \nabla Z \cdot g_\xi(Z))(Z - Z)^-\alpha dx
\]

\[
+ \int_\Omega (f_\xi(Z) - f_\xi(Z))\nabla(Z - Z)^-\alpha dx
\]

using also (3.17). By the Lipschitz continuity of $g_\xi$, $g_\xi$ and $f_\xi$ (recall (3.12) and the assumptions on $g$) in $Z$ and (3.14) we obtain

\[
\frac{d}{dt} \int_\Omega (Z - Z)^-\alpha dx + \int_\Omega A_\xi(Z)(\nabla Z^-\alpha, \nabla Z^-\alpha) dx
\]

\[
\leq \xi \int_\Omega (\nabla Z - Z)^-\alpha dx + c(\xi, \ell) \int_\Omega (Z - Z)^-\alpha dx
\]

for all $\xi > 0$. Due to (3.14) the first term can be absorbed for $\xi$ small enough, whereas the second one can be handled by Gronwall’s lemma and $\varphi_0 > 0$. We conclude that

\[
(3.24) \quad Z \geq Z > 0 \text{ in } T \times \Omega.
\]

Recall that $Z$ depends on $\ell$. We are now going to prove a uniform lower bound. Similarly to (3.20) and (3.21) we consider a function $\varphi \in C^\infty(T \times \Omega)$ satisfying

\[
\varphi(x) \geq 1 \text{ in } T \times \Omega,
\]

\[
\frac{\nabla \varphi \cdot A_\xi(Z)\nu}{\varphi} \geq 1 \text{ on } T \times \partial \Omega.
\]

Let us first assume that the minimum of $v = \varphi e^{\lambda_1 t} Z$ is attained in an interior point $(t_0, x_0) \in I \times \Omega$. We obtain similarly to (3.22)

\[
(3.27) \quad 0 \geq Z e^{\lambda_1 t_0} \varphi \cdot A_\xi(t_0, x_0)\nu(x_0).
\]

An appropriate choice of $\lambda_1$ contradicts (3.27). In the case of $x_0 \in \partial \Omega$ and $t_0 > 0$ we have similarly to the proof of (b)

\[
0 \geq Z e^{\lambda_1 t_0} \varphi \cdot A_\xi(t_0, x_0)\nu(x_0).
\]
This gives a contradiction by (4.24), (4.25) and (4.20). Consequently, the minimum of $Z$ is attained at a point $(0,x_0)$ for some $x_0 \in \mathcal{I}$. This gives the claim of (b) since $\lambda_1$ is independent of $\ell$. Hence all properties of $Z$ are shown that imply (by transformation) the existence of a function $\vartheta$ with the required properties. \hfill \Box

4. CONSTRUCTION OF AN APPROXIMATE SOLUTION

In this section we construct an approximation of the system, where the continuity equation contains an artificial diffusion ($\varepsilon$-layer) and the pressure is stabilised by a high power of the density ($\delta$-layer). Following [20] we add various regularizing terms depending on $\varepsilon$ and $\delta$ to the equations to preserve the energy balance. One of the regularizing terms can only be shown to belong to $L^1$, which is not enough to conclude uniform continuity in time needed for the application of Theorem 2.13. To overcome this peculiarity we include a further diffusion term of the fluid velocity which is non-linear and of $p$-growth with $p > \beta > 2$. It vanishes in the limit but improves the time integrability mentioned before. Additionally, we regularize the shell equation by replacing the operator $K$ with

$$K_\varepsilon(\eta) = K(\eta) + \varepsilon \mathcal{L}(\eta), \quad \mathcal{L}(\eta) = \frac{1}{2} \int_\omega |\nabla^3 \eta|^2 \, dy,$$

defined for $\eta \in W^{3,2}(\omega)$. Thanks to this we can prove compactness of the shell energy in the Galerkin limit.

**Remark 4.1.** We observe that adding dissipative regularization terms to the shell equation is not possible. This is a special feature for energetically closed systems and in contrast to other fluid systems. Indeed, a dissipation term in the solid creates heat on the surface, which consequently effects the temperature. In the case of shells this yields a non-homogeneous Neumann boundary value for the temperature variable. This non-homogeneity naturally possesses the "wrong sign" in order to attain in the limit the boundary values for the temperature that are in accordance with the concept of weak solutions. In the case of visco-elastic solids, where dissipative terms such as an additional heat source are included (they are physical and not only relaxation terms) our approximation would yield the correct non-homogeneous boundary values. However, we considered here perfectly elastic solids. Hence all energy is supposed to be stored in the elastic potential.

In contrast to [6] and [31] we construct the fixed point on the Galerkin level. This allows to remove one regularization level for the boundary and the convective term that was needed there. The formulation of the Galerkin approximation in our case is more involved since the basis functions are defined on the a priori unknown time dependent domain. The fixed point argument (which is now applied on the Galerkin level) is, however, much easier. After constructing a solution on the basic level, we prove in Subsection 4.2 the energy equality and derive further estimates through the Helmholtz-function. In particular, we derive the approximate system and the a-priori estimates. They are essential for the remainder of the paper and are preserved in all limit procedures.

For the original system we seek a solution of the shell in the class

$$Y^f := W^{1,\infty}(I;L^2(\omega)) \cap L^\infty(I;W^{2,2}(\omega)).$$

However, in this section we are dealing with a regularised system where instead solutions are located in

$$\tilde{Y}^f := W^{1,\infty}(I;L^2(\omega)) \cap L^\infty(I;W^{3,2}(\omega)).$$

For $\zeta \in \tilde{Y}^f$ with $\|\zeta\|_{L^\infty(\omega)} \leq \frac{1}{2}$ we consider

$$\tilde{X}_\zeta^f := L^p(I;W^{1,p}(\Omega_\zeta(\omega))), \quad Z_\zeta^f := L^2(I;W^{1,2}(\Omega_\zeta)) \cap L^\infty(I;L^4(\Omega_\zeta)).$$

The space $X_\eta^f$ is defined in Section 2.5. A solution to the regularized system, in the weak formulation, is a quadruplet $(\eta, u, \varrho, \vartheta) \in \tilde{Y}^f \times \tilde{X}_\zeta^f \times X_\eta^f \times Z_\zeta^f$ that satisfies the following.
Remark 4.2. In order to deal with the term $\int_I \int_{\Omega_n} \varepsilon \nabla \varphi \nabla u \cdot \phi \, dx \, dt$ (appearing in (K1)) in the proof of (K3) we need higher integrability of $\nabla u$ in time. This is achieved by introducing an artificial p-Laplacian term $\varepsilon(1 + \vartheta)(1 + |\nabla u|)^{p-2} \nabla u$ for some $p > \beta > 2$ on the Galerkin approximation in the next section. It gives the additional term $\varepsilon(1 + \vartheta) \max \{|P|, |\nabla u|\}$ in (K3). The term $\varepsilon(1 + \vartheta)P$ in (K3) is the weak limit of the p-Laplacian term and can be seen as the defect in the strong convergence of $\nabla u$. It disappears in the limit $\varepsilon \to 0$.

The rest of this section is dedicated to the proof of the following existence theorem.
Theorem 4.3. Assume that we have for some $\alpha \in (0, 1) \begin{equation}
abla \frac{|q_0|^2}{\varrho_0} \in L^1(\Omega_{\eta_0}), \frac{\eta_0}{\varrho_0} \in C^{2, \alpha}((\mathcal{I}^m)_{\eta_0}), \eta_0 \in W^{3, 2}(\omega; [\frac{1}{2}, 1]), \eta_1 \in L^2(\omega), \begin{aligned}
f \in L^2(I; L^\infty(\mathbb{R}^3)), \quad g \in L^2(I \times \omega), \quad H \in C^{1, \alpha}(T \times \mathbb{R}^3), \quad H \geq 0.
\end{aligned}
Furthermore suppose that $\varrho_0$ and $\vartheta_0$ are strictly positive and that \(1.13\) is satisfied. Then there exists a solution $(\eta, u, \varrho, \vartheta) \in \mathcal{Y}^T \times \mathcal{X}^T \times Z^T \times \mathcal{Y}^T$ to (K1)–(K4). Here, we have $I = (0, T_s)$, where $T_s < T$ only if $\lim_{t \to T_s} \| \eta(t, \cdot) \|_{L^\infty} = \frac{1}{2}$ or the Koiter energy degenerates (namely, if $\lim_{t \to T_s} \tau(s, y) = 0$ for some point $y \in \omega$).

We prove Theorem 4.3 in two steps. First we construct a finite dimensional Galerkin approximation to (K1)–(K3) in the next subsection. Then we derive the energy balance, prove uniform a priori estimates and pass to the limit.

4.1. Galerkin approximation. By solving respective eigenvalue problems we construct a smooth orthogonal basis $(\tilde{X}_k)_{k \in \mathbb{N}}$ of $W^{1, 2}_0(\Omega)$ that is orthogonal in $L^2(\Omega)$ and a smooth orthonormal basis $(\tilde{Y}_k)_{k \in \mathbb{N}}$ of $W^{3, 2}(\omega)$ which is orthogonal in $L^2(\omega)$. We define vector fields $\tilde{Y}_k$ by setting $\tilde{Y}_k = \mathcal{F}_T(\tilde{Y}_k \tilde{\varphi}^{-1})$, where $\mathcal{F}_T$ is the extension operator used in Section 2.3. We recall that the $\tilde{Y}_k$’s are smooth. Now we choose an enumeration $(\tilde{\varphi}_k)_{k \in \mathbb{N}}$ of $(\tilde{Y}_k)_{k \in \mathbb{N}}$ and a basis $(\tilde{Y}_k)_{k \in \mathbb{N}}$ of $W^{3, 2}(\omega)$. We define for $\phi \in W^{3, 2}(\omega)$ the orthogonal projection (in space) $\mathcal{P}_N$ as

$$\mathcal{P}_N(\phi) := \sum_{k=1}^N P_k^N(\phi) w_k := \sum_{k=1}^N \langle \phi, w_k \rangle_{W^{3, 2}(\omega)} w_k,$$

which satisfies the expected stability and convergence properties in all spaces relevant for the analysis. Next we seek for a couple of discrete solutions $(\eta_N, u_N)$ of the form

$$\eta_N = \mathcal{P}_N \eta_0 + \sum_{k=1}^N \int_0^t \alpha_k N w_k \, dt, \quad u_N = \sum_{k=1}^N \alpha_k N \tilde{Y}_k \circ \Psi_{\eta_N}^{-1},$$

with time-dependent coefficients $\alpha_N = (\alpha_k N)_{k=1}^N$, which solve the following discrete version of (1.1):

$$\begin{aligned}
\int_{\Omega_N} g_N(t) u_N(t) \cdot \tilde{Y}_k \circ \Psi_{\eta_N}^{-1}(t) \, dx \\
- \int_0^t \int_{\Omega_N} \left( g_N u_N \cdot \tilde{\varphi}_k \left( \tilde{Y}_k \circ \Psi_{\eta_N}^{-1} \right) + g_N u_N \circ u_N : \nabla \tilde{Y}_k \circ \Psi_{\eta_N}^{-1} \right) \, dx \, dt \\
+ \int_0^t \int_{\Omega_N} \left( \tilde{N} \cdot (\partial_N, \nabla u_N) : \nabla \tilde{Y}_k \circ \Psi_{\eta_N}^{-1} \right) \, dx \, ds \\
- \int_0^t \int_{\Omega_N} (p_0 \circ \Psi_{\eta_N}^{-1} + \varepsilon \nabla g_N \circ \nabla u_N \circ \Psi_{\eta_N}^{-1}) \, dx \, ds \\
+ \int_0^t \int_{\Omega_N} (K_0 \circ \Psi_{\eta_N}^{-1} + \varepsilon \nabla g_N \circ \nabla u_N \circ \Psi_{\eta_N}^{-1}) \, dx \, ds \\
= \int_0^t \int_{\Omega_N} g_N f \cdot \tilde{Y}_k \circ \Psi_{\eta_N}^{-1} \, dx \, ds + \int_0^t \int_{\Omega_N} g N w_k \, dy \, ds \\
+ \int_{\Omega_N} \varrho_0 \circ \tilde{Y}_k \circ \Psi_{\eta_N}^{-1} (0, \cdot) \, dx \, ds + \int_{\Omega_N} E \varrho_0 \circ \tilde{Y}_k \circ \Psi_{\eta_N}^{-1} \, dx \, dy
\end{aligned}$$

(4.6)

Here $g_N = g(\eta_N, u_N)$ and $\vartheta_N = \vartheta(\eta_N, u_N, \eta_0)$ are the unique solutions from Theorems 6.8 and 5.1 subject to the initial data $\varrho_0$ and $\vartheta_0$, where $\xi \equiv \eta_N$ and $w \equiv u_N$. Note that by construction we have $\tau_{\eta_N} u_N = \vartheta_0 \eta_N \nu$ and that we can choose $\alpha_{kN}(0)$ in a way that $u_N(0)$ converges to $\varrho_0$. In order to solve (4.6) we decouple the nonlinearities. Consider a given couple of discrete functions $(\zeta_N, \nu_N)$ of the form

$$\begin{aligned}
\zeta_N &= \mathcal{P}_N \eta_0 + \sum_{k=1}^N \int_0^t \beta_{kN} w_k \, ds, \quad \nu_N = \sum_{k=1}^N \beta_{kN} \tilde{Y}_k \circ \Psi_{\eta_N}^{-1},
\end{aligned}$$

(4.7)
with time-dependent coefficients $\beta_N = (\beta_{kN})_{k=1}^N$. By construction they satisfy $\text{tr} \nabla v_N = \partial_t \zeta_N v$. We aim to solve

$$
\int_{\Omega_N} g_N(t)u_N(t) \cdot \omega_k \circ \Psi_{\zeta_N}^{-1}(t) \, dx
$$

\begin{align}
&- \int_{0}^{t} \int_{\Omega_N} \left( g_N u_N \cdot \partial_t \left( \omega_k \circ \Psi_{\zeta_N}^{-1} \right) + g_N \nabla v_N \otimes u_N : \nabla \omega_k \circ \Psi_{\zeta_N}^{-1} \right) \, dx \, dt \\
&+ \int_{0}^{t} \int_{\Omega_N} \left( S^t(\partial_{N} \nabla u_N) : \nabla \omega_k \circ \Psi_{\zeta_N}^{-1} \right) \, dx \, d\sigma \\
&+ \int_{0}^{t} \int_{\Omega_N} \left( \rho_N g \cdot \omega_k \circ \Psi_{\zeta_N}^{-1} \right) \, dx \, d\sigma + \int_{\omega} \rho_N \partial_t w_k \, dy \\
&+ \int_{\Omega_N} q_0 \cdot \omega_k \circ \Psi_{\zeta_N}^{-1}(0, \cdot) \, dx + \int_{\omega} \eta_1 w_k \, dy.
\end{align}

(4.8)

Here $g_N = \rho(\zeta_N, v_N)$ and $\partial_N = \partial(\zeta_N, v_N, g_N)$ are the unique solutions from Theorems 3.3 and 3.4 subject to the initial data $g_0$ and $\partial_0$, where $\zeta \equiv \zeta_N$ and $w \equiv v_N$. Note that this is possible since $\|P_N \eta_0\|_{L^\infty} \leq \frac{T}{2}$ for $N$ large enough, which implies $\|\zeta_N\|_{L^\infty, \omega_T} \leq \frac{T}{2}$ for $T$ small enough. The system (4.8) is equivalent to a system of integro-differential equations for the vector $\alpha_N = (\alpha_{kN})_{k=1}^N$. It reads as

$$
A(t)\alpha_N(t) = \int_{0}^{t} B(s)\alpha_N(s) \, ds + \int_{0}^{t} \tilde{B} \left( \sigma, \alpha_N(\sigma), \int_{0}^{\sigma} \alpha_N(s) \, ds \right) \, d\sigma + \int_{0}^{t} c(\sigma) \, d\sigma + \tilde{c},
$$

with

$$
A_{ij} = \int_{\Omega_N} g_N(t) \omega_i \circ \Psi_{\zeta_N}^{-1}(t) \cdot \omega_j \circ \Psi_{\zeta_N}^{-1}(t) \, dx + \int_{\omega} w_i w_j \, dy
$$

$$
B_{ij} = \int_{\Omega_N} \left( g_N \omega_i \circ \Psi_{\zeta_N}^{-1} \cdot \partial_t \left( \omega_j \circ \Psi_{\zeta_N}^{-1} \right) + g_N \nabla v_N \otimes \omega_i \circ \Psi_{\zeta_N}^{-1} : \nabla \omega_j \circ \Psi_{\zeta_N}^{-1} \right) \, dx
$$

$$
- \int_{\Omega_N} \varepsilon \nabla g_N \nabla \omega_i \circ \Psi_{\zeta_N}^{-1} \cdot \omega_j \circ \varphi_{\zeta_N}^{-1} \, dx \, d\sigma - \int_{\omega} w_i \partial_t w_j \, dy
$$

$$
\tilde{B}_{ij} = \int_{\Omega_N} K_{ij} \left( \rho_N \eta_0 + \sum_{k=1}^{N} \alpha_{kN}(s) w_k \right) \, w_j \, dy
$$

$$
c_i = \int_{\Omega_N} p_N(\rho_N, \partial_N) \nabla \omega_i \circ \Psi_{\zeta_N}^{-1} \, dx + \int_{\Omega_N} g_N f \cdot \omega_i \circ \Psi_{\zeta_N}^{-1} \, dx \, dt + \int_{\omega} g w_i \, dy
$$

$$
\tilde{c}_i = \int_{\Omega_N} q_0 \cdot \omega_i \circ \Psi_{\zeta_N}^{-1}(0, \cdot) \, dx + \int_{\omega} \eta_1 w_i \, dy.
$$

The matrix $A_{ij}$ is invertible and all non-linear quantities are locally Lipschitz continuous in $\alpha_N$ (compare also with [6 Thm. 4.4]). Also our analysis from Section 3 shows that $\partial_N$ and $g_N$ depend in a smooth way on $v_N$ and $\zeta_N$. By the Picard-Lindelöf theorem there is a unique solution in short time. Consequently, we obtain a solution $(\eta_N, u_N)$ to (4.8) which satisfies the following energy balance (testing (4.8) by $(u_N, \partial_t \eta_N)$

---

Eventually, the solution can be extended for arbitrary times due to the a priori estimates which we derive below in \[\text{(4.13)}\].
and (3.11) by $\frac{1}{2}|u_N|^2$)

$$-\int_I \left( \int_{\Omega_{N}} g_N \frac{|u_N|^2}{2} dx + \int_{\Omega_{N}} \frac{1}{\gamma - 1} \frac{\beta}{\beta + 1} \frac{g_N^2}{\beta N} dx + \int_{\Omega_{N}} \frac{1}{2} \frac{\partial \eta N}{\partial y} dy + K(\eta N) \right) \partial \psi dt$$

$$+ \int_I \psi \int_{\Omega_{c}} \mathbf{S}^c(\partial N, \nabla u_N) : \nabla u_N \psi dx dt$$

$$= \psi(0) \left( \int_{\Omega_{N}} \frac{|q_0|^2}{2g_0} dx + \int_{\Omega_{N}} \frac{|\eta_0|^2}{2} dy + K(\eta_0) \right) + \int_{\Omega_{N}} \left( \frac{1}{\gamma - 1} \frac{\beta}{\beta + 1} \frac{g_N^2}{\beta N} \right) dx$$

$$+ \int_I \psi \int_{\Omega_{N}} g_N f \cdot u_N \psi dx dt + \int_I \psi \int_{\Omega_{c}} \varepsilon \partial_{\theta} \varphi_N \psi dx dt$$

$$+ \int_I \psi \int_{\Omega_{c}} \left( g_N \partial_N + \frac{\delta}{\beta N} \right) \psi dx dt$$

for all $\psi \in C^\infty_c([0, T))$. Testing further the continuity equation by $\frac{\partial \beta g_N^2}{(3-\beta)} + \frac{\gamma^2 - 1}{(3-\gamma)}$ yields

$$= \psi(0) \left( \int_{\Omega_{N}} \frac{|q_0|^2}{2g_0} dx + \int_{\Omega_{N}} \frac{|\eta_0|^2}{2} dy + K(\eta_0) \right) + \int_{\Omega_{N}} \left( \frac{1}{\gamma - 1} \frac{\beta}{\beta + 1} \frac{g_N^2}{\beta N} \right) dx$$

$$+ \int_I \psi \int_{\Omega_{N}} g_N f \cdot u_N \psi dx dt + \int_I \psi \int_{\Omega_{c}} \varepsilon \partial_{\theta} \varphi_N \psi dx dt$$

$$+ \int_I \psi \int_{\Omega_{c}} \left( g_N \partial_N + \frac{\delta}{\beta N} \right) \psi dx dt$$

for all $\psi \in C^\infty_c([0, T))$. We consider the mapping

$$F : D \rightarrow F(D), \quad \beta \mapsto \alpha, \quad D = \left\{ \beta \in C^{1,\alpha}(\tilde{T}_*, \mathbb{R}^N) : \sup_{T_*} \|\beta\|_{\alpha} \leq K^* \right\}$$

where $I_* = (0, T_*)$ and $\alpha \in (0, 1)$. We will choose $K^*$ sufficiently large. In dependence of $K^*$ we find $T_*$ (sufficiently small) but uniform to solve the above ODE uniquely on $I_*$. Note that we may take $T^*$ small enough (in dependence of $K^*$) such that $\zeta_N$ (defined via $\beta$ by (2.17)) satisfies $\|\zeta_N\|_{L^2_{T*}} \leq \frac{1}{2}$ for any $\beta \in D$.

We are going to prove that $F$ has a fixed point. Let us first note that $F$ is upper-semicontinuous. Indeed, if we have a sequence ($\beta^*$) which converges in $C^{1,\alpha}(\tilde{T}_*, \mathbb{R}^N)$ to some $\beta$ such that $\alpha^* = F(\beta^*)$ converges in $C^{1,\alpha}(\tilde{T}_*, \mathbb{R}^N)$ to some $\alpha$, we have $\alpha = F(\beta)$. This is due to the unique solvability of (4.18) and the continuity of the coefficients $A, B, \tilde{B}$ and $c$. In fact, the continuity of $A, B, \tilde{B}$ and $c$ (with respect to $\beta$) can be shown by transforming the integrals to the reference domain and using (2.19) similarly to the proofs of Theorems 3.3 and 3.4. The regularity and continuity of $g_N$ and $\varphi_N$ then implies the continuity of the coefficients.

Next we aim to show that $F(D) \subset D$. The internal energy equation (3.19) for $\varphi_N$ yields

$$-\int_I \int_{\Omega_{N}} (a\varphi_N^4 + c_N \varphi_N \varphi_N) \partial t \psi \right) dx dt - \psi(0) \int_{\Omega_{N}} (a\varphi_0^4 + c_N \varphi_0 \varphi_0) dx$$

$$= \int_I \int_{\Omega_{N}} \left[ \mathbf{S}^c(\varphi_N, \nabla \varphi_N) : \nabla \varphi_N - (\frac{\varphi_N^2}{2} + g_N \varphi_N) \nabla \varphi_N \psi \right) dx dt$$

$$+ \int_I \int_{\Omega_{N}} \left( \epsilon \partial_{\theta} \varphi_N + \frac{\delta}{\varphi_N} + \frac{\delta}{\varphi_N} \right) \psi dx dt$$
for all \( \psi \in C_c^\infty([0,T]) \). Combining this with (4.10) implies

\[
- \int_I \partial_t \psi \mathcal{E}^N_{\varepsilon, \delta} \, dt = \psi(0) \mathcal{E}^N_{\varepsilon, \delta}(0) + \int_I \psi \int_{\Omega_N} \left( \frac{\delta}{\varepsilon\delta_N} - \varepsilon \partial_N^2 \right) \, dx \, dt \\
+ \int_I \psi \int_{\Omega_N} \left( S^\varepsilon(\partial_N, \nabla v_N) : \nabla v_N - S^\varepsilon(\partial_N, \nabla u_N) : \nabla u_N \right) \, dx \, dt \\
+ \int_I \psi \int_{\Omega_N} p_N(\partial_N, \theta_N)(\text{div} u_N - \text{div} v_N) \, dx \, dt \\
+ \int_I \psi \int_{\Omega_N} (\Theta_N + \Theta_N \Pi \cdot u_N) \, dx \, dt + \int_I \psi \int_\omega g \partial_N \eta_N \, dy \, dt
\]

(4.11)

with

\[
\mathcal{E}^N_{\varepsilon, \delta}(t) = \int_{\Omega_N} \left( \frac{1}{2} \partial_N(t) |u_N(t)|^2 + \partial_N(t) \varepsilon \delta(t, \partial_N(t)) \right) \, dx \\
+ \int_\omega \left( \frac{1}{2} |\partial_N(t)|^2 \right) \, dy + K_\varepsilon(\eta_N(t)).
\]

By choosing \( \psi = 1_{(0,T)} \), we find that (4.11) implies uniform a-priori estimates. Note that we can apply Young’s inequality to the forcing terms in (4.11) and absorb terms containing the unknowns in the left-hand side. Moreover, by Theorem 3.3 we obtain bounds for \( \theta_N \) (in dependence of \( \varepsilon, \delta, N, K^* \)) from below such that

\[
\int_I \int_{\Omega_N} \frac{\delta}{\varepsilon \delta_N} \, dx \, dt \leq c(\varepsilon, \delta, N, K^*) \leq 1
\]

for \( K^* \) small enough. So, in order to apply the Gronwall lemma it is enough to control the error term

\[
\int_I \int_{\Omega_N} \left( S^\varepsilon(\partial_N, \nabla v_N) : \nabla v_N - S^\varepsilon(\partial_N, \nabla u_N) : \nabla u_N \right) \, dx \, dt \\
+ \int_I \int_{\Omega_N} p_N(\partial_N, \theta_N)(\text{div} u_N - \text{div} v_N) \, dx \, dt \\
\leq \int_I \int_{\Omega_N} \left( S^\varepsilon(\partial_N, \nabla v_N) : \nabla v_N + p_N(\partial_N, \theta_N)(|\nabla u_N| + |\nabla v_N|) \right) \, dx \, dt.
\]

Using Theorem 3.3 and 3.4 we can bound \( \partial_N \) and \( \partial_N \) in terms of \( K \) such that the above is bounded by

\[
\leq c(K) \int_I \int_{\Omega_N} (1 + |\nabla v_N|^p) \, dx \, dt + c(K) \int_I \int_{\Omega_N} |\nabla u_N|^2 \, dx \, dt \\
\leq c(K, N) T^* \left( 1 + \text{sup}_{t^*} |\beta_N|^p \right) + c(K, N) T^* \sup_{t^*} \int_{\Omega_N} |u_N|^2 \, dx \\
\leq c(K, N) T^* + c(K, N) T^* \sup_{t^*} \int_{\Omega_N} g_N |u_N|^2 \, dx.
\]

We choose \( T^* = T^*(\varepsilon, N, K^*) \) small enough such that \( c(K, N) T^* \leq \frac{1}{2} \) and and obtain

(4.12)

\[
\sup_{t^*} \mathcal{E}^N_{\varepsilon, \delta} \leq c(\varepsilon, H, g, q_0, \eta_0, \eta_1, \eta_0).
\]

In particular, we have

(4.13)

\[
\sup_{t^*} \int_{\Omega_N} |u_N|^2 \, dx + \sup_{t^*} \int_\omega \left( \frac{|\partial_N \eta_N|^2}{2} \right) \, dy + \sup_{t^*} K_\varepsilon(\eta_N) \leq c(\varepsilon, H, g, q_0, \eta_0, \eta_1, \eta_0).
\]

recalling the lower bound for \( \partial_N \) from Theorem 3.3 (b) (which depends on \( N \) here). Consequently, we see that the mapping \( \beta \mapsto \alpha \) satisfies \( F(D) \subset D \), for \( K^* \) large enough.

Now, we need to prove compactness of \( F \) with respect to the \( C^{1,\beta}(\mathcal{T}) \) topology. First we find by Leibnitz rule that

\[
\partial_t \alpha_N = A^{-1} \left( \partial_t (A \alpha_N) - \partial_t A \alpha_N \right).
\]

Due to (4.9) and the regularity of \( \partial_N \) and \( \partial_N \) from Theorems 3.3 and 3.4 we have \( \partial_t (A \alpha_N) \in C^{1}(\mathcal{T}) \). This can be easily seen by transforming the integrals in the definitions of the coefficients \( A, B, \mathcal{B} \) and \( c \) to the reference domain and recalling from (2.10) that \( \Psi_{\varepsilon, N} \) and \( \Psi_{\varepsilon, N}^{-1} \) have the same regularity as \( \varepsilon_N \). Also note that
\( \beta_N \in C^{1,\alpha}(\overline{T}_* \setminus \Omega) \) implies \( \zeta_N \in C^{2,\alpha}(\overline{T}_*) \) by construction. Similarly, we are going to prove that \( \partial_i A_{i,j} \in C^1(\overline{T}_*) \).

By taking the test function \( \tilde{\omega}_i \circ \varphi^{-1}_{\zeta_N} \cdot \omega_j \circ \varphi^{-1}_{\zeta_N} \) in the continuity equation we find that

\[
\partial_t A_{i,j} = \frac{d}{dt} \int_{\Omega_N} \varrho_N \tilde{\omega}_i \circ \Psi^{-1}_{\zeta_N} \cdot \omega_j \circ \Psi^{-1}_{\zeta_N} \, dx
\]

\[
= \int_{\partial \Omega} \tilde{\omega}_i \circ \varphi^{-1}_{\zeta_N} \varrho_N \tilde{\omega}_i \circ \varphi^{-1}_{\zeta_N} \cdot \omega_j \circ \varphi^{-1}_{\zeta_N} \nu_{\Omega_N} \, d\gamma
\]

\[
+ \int_{\Omega_N} \varrho_N \nabla \cdot \left( \tilde{\omega}_i \circ \varphi^{-1}_{\zeta_N} \cdot \omega_j \circ \varphi^{-1}_{\zeta_N} \right) \, dx
\]

\[
+ \varepsilon \int_{\Omega_N} \nabla \varrho_N \cdot \left( \tilde{\omega}_i \circ \varphi^{-1}_{\zeta_N} \cdot \omega_j \circ \varphi^{-1}_{\zeta_N} \right) \, dx
\]

\[
+ \int_{\Omega_N} \varrho_N \partial_i \left( \tilde{\omega}_i \circ \varphi^{-1}_{\zeta_N} \cdot \omega_j \circ \varphi^{-1}_{\zeta_N} \right) \, dx
\]

\[
+ \int_{\Omega_N} \varrho_N \tilde{\omega}_i \circ \varphi^{-1}_{\zeta_N} \cdot \partial_i \left( \omega_j \circ \varphi^{-1}_{\zeta_N} \right) \, dx.
\]

The last two terms containing the time-derivative behave as \( \beta_N \) which is bounded in \( C^{1,\alpha}(\overline{T}_*) \). Consequently, we find that \( \partial_i A_{i,j} \in C^1(\overline{T}_*) \) with bound depending only on \( K \) (and \( N \)). So, the mapping \( F \) is compact by Arcelá-Ascoli’s theorem. Consequently, there is a fixed point \( \alpha^* \) which gives rise to the solution to (4.10) if \( T^* \) is sufficiently small (depending on \( \delta, \varepsilon, K^* \) and \( N \)). The interval of existence can be extended by iterating the procedure and gluing the solutions together.

### 4.2. Uniform estimates–total energy balance

At this stage \( \vartheta_N \) is still strictly positive by Theorem 3.4 (with a bound depending on \( N \)) so we can divide the internal energy defined in (3.11) by \( \vartheta_N \) to obtain the entropy balance

\[
\partial_t \left( \varrho_N s(\varrho_N, \vartheta_N) \right) + \nabla \cdot \left( \varrho_N \vartheta_N \right) \mathbf{u}_N = \nabla \cdot \left( \frac{\varrho_N s(\varrho_N, \vartheta_N) \mathbf{u}_N}{\vartheta_N} \right) - \nabla \left[ \left( \frac{\vartheta_N}{\vartheta_N} + \delta \vartheta_N^{-1} + \frac{1}{\vartheta_N^2} \right) \nabla \vartheta_N \right]
\]

\[
= \frac{1}{\vartheta_N} \left( \left( \frac{\vartheta_N}{\vartheta_N} + \delta \vartheta_N^{-1} + \frac{1}{\vartheta_N^2} \right) \nabla \vartheta_N \right)^2 + \frac{1}{\vartheta_N} \nabla \cdot \mathbf{u}_N + \varepsilon \mathcal{R}_N = \vartheta_N^4
\]

satisfied in \( I \times \Omega_N \), together with the boundary condition \( \nabla \vartheta_N \cdot \nu_{\Omega_N} \big|_{\partial \Omega_{\Omega_N}} = 0 \). In the weak form it reads as

\[
\int_I \frac{d}{dt} \int_{\Omega_N} \varrho_N s(\varrho_N, \vartheta_N) \psi \, dx \, dt - \int_I \int_{\Omega_N} \left( s(\varrho_N, \vartheta_N) \partial_t \psi + \varrho_N s(\varrho_N, \vartheta_N) \mathbf{u}_N \cdot \nabla \psi \right) \, dx \, dt
\]

\[
\geq \int_I \int_{\Omega_N} \frac{1}{\vartheta_N} \mathbf{S}^t(\varrho_N, \vartheta_N) : \nabla \mathbf{u}_N \psi \, dx \, dt
\]

\[
+ \int_I \int_{\Omega_N} \frac{1}{\vartheta_N} \left( \left( \frac{\vartheta_N}{\vartheta_N} + \delta \vartheta_N^{-1} + \frac{1}{\vartheta_N^2} \right) \nabla \vartheta_N \right)^2 + \delta \frac{1}{\vartheta_N^2} \right) \psi \, dx \, dt
\]

\[
+ \int_I \int_{\Omega_N} \left( \frac{\vartheta_N}{\vartheta_N} + \delta \vartheta_N^{-1} + \frac{1}{\vartheta_N^2} \right) \nabla \vartheta_N \cdot \nabla \psi \, dx \, dt + \int_I \int_{\Omega_N} \frac{\varrho_N}{\vartheta_N} \mathbf{H} \psi \, dx \, dt
\]

\[
\geq \int_I \int_{\Omega_N} \varepsilon \left[ \mathcal{R}_N(\varrho_N) - \vartheta_N^4 \right] \psi \, dx \, dt
\]

for all \( \psi \in C^\infty(I \times \mathbb{R}^3) \) with \( \psi \geq 0 \). We combine this with the energy balance proved in (4.11) which reads as (note that in the fixed point we have \( \zeta_N = \eta_N \) and \( \mathbf{V}_N = \mathbf{u}_N \))

\[
- \int_I \partial_t \psi \mathcal{E}_N^t \, dt = \psi(0) \mathcal{E}_N^t(0) + \int_I \int_{\Omega_N} \left( \frac{\delta}{\vartheta_N} - \varepsilon \vartheta_N^2 \right) \, dx \, dt + \int_I \psi \int_{\Omega_N} \varrho_N \mathbf{H} \, dx \, dt
\]

\[
+ \int_I \int_{\Omega_N} \varrho_N \mathbf{f} \cdot \mathbf{u}_N \, dx \, dt + \int_I \int_{\Omega_N} \varrho_N \eta_N \, dx \, dt
\]

\[
(4.15)
\]
with
\[
\mathcal{E}_\delta^N(t) = \int_{\Omega_{\eta_N}(t)} \left( \frac{1}{2} \varrho_N(t) |u_N|^2 + \varrho_N(t) \varepsilon \vartheta_N + \frac{1}{2} \varrho_N(t) \varepsilon \vartheta_N(t) \right) dx + \int_{\omega} \left| \frac{\partial \varphi_N(t)}{\partial y} \right|^2 dy + K(\eta_N(t)).
\]

We follow [24] Chapter 2, Section 2.2.3, and obtain by substracting from this \( \Theta \)-times the integral of (4.14) (or \( \Theta \)-times the weak formulation tested with \( \varphi \equiv 1 \)) to obtain
\[
- \int_I \partial_t \psi \left( \mathcal{E}_{\delta,\varepsilon}^N - \Theta \varrho_N \vartheta_N, \vartheta_N \right) dt + \Theta \int_{\Omega_{\eta_N}} \sigma_{\varepsilon,\delta}^N dx \ dt + \int_I \int_{\Omega_{\eta_N}} \varepsilon \vartheta_N^2 dx \ dt
\]
\[
= \psi(0) \left( \mathcal{E}_{\delta,\varepsilon}^N - \Theta \varrho_N \vartheta_N \right)(0) + \Theta \int_I \int_{\Omega_{\eta_N}} \varepsilon \vartheta_N^2 dx \ dt
\]
\[
+ \int_I \int_{\Omega_{\eta_N}} \varrho_N H dx \ dt + \int_I \int_{\Omega_{\eta_N}} \varrho_N f \cdot u_N dx \ dt + \int_I \int_{\omega} g \partial \varphi_N \ dy \ dt,
\]
where
\[
\sigma_{\varepsilon,\delta}^N = \frac{1}{\varrho_N} \left[ \left( \varrho_N \vartheta_N : \nabla u_N + \varepsilon (1 + \vartheta_N) |\nabla u_N|^2 \right) \right] - \frac{1}{\varrho_N} \left[ \left( \varrho_N \vartheta_N : \nabla u_N + \varepsilon (1 + \vartheta_N) |\nabla u_N|^2 \right) \right] \varrho_N + \frac{\varepsilon}{\varrho_N} \vartheta_N.
\]

Consequently, we obtain the estimates
\[
\sup_I \int_{\Omega_{\eta_N}} \varrho_N |u_N|^2 dx + \sup_I \int_{\Omega_{\eta_N}} \varrho_N^2 dx + \int_I \int_{\Omega_{\eta_N}} |\nabla u_N|^p dx \ dt \leq c,
\]
\[
\varepsilon \sup_I \int_{\omega} |\nabla^3 \eta_N|^2 d\omega + \sup_I \int_{\omega} \left| \frac{\partial \eta_N}{\partial y} \right|^2 d\omega + \sup_I K(\eta_N) \leq c,
\]
\[
\sup_I \int_{\Omega_{\eta_N}} \vartheta_N^4 dx + \int_I \int_{\Omega_{\eta_N}} \frac{1}{\varrho_N} \left( \varepsilon \vartheta_N + \delta \vartheta_N^2 + \frac{1}{\varrho_N} \right) \varrho_N + \varrho_N H dx \ dt \leq c,
\]
where \( c = (f, \varrho, \vartheta, \eta, \eta_0, \eta, \vartheta_0) \) is independent of \( N \). The first estimate together with Poincaré’s inequality, the boundary condition \( tr_{\eta_N} u_N \) and bound for \( \partial \eta_N \) from the second estimate impliethat \( u_N \) is bounded in \( L^p(I; L^p(\Omega_{\eta_N})) \). So, we may choose a subsequence such that
\[
\eta_N \to \eta \text{ in } L^\infty(I, W^{3,2}(\omega)),
\]
\[
\vartheta_N \to \vartheta \text{ in } L^\infty(I, L^2(\omega)),
\]
\[
u_{n_N} \to u \text{ in } L^p(I; L^p(\Omega_{\eta_N})),
\]
\[
\nabla u_N \to \nabla u \text{ in } L^p(I; L^p(\Omega_{\eta_N})),
\]
\[
|\nabla u_N|^p - 2 \nabla u_N \to \nabla p \text{ in } L^p(I; L^p(\Omega_{\eta_N})),
\]
\[
\varrho_N \to \varrho^* \text{ in } L^\infty(I; L^3(\Omega_{\eta_N})),
\]
\[
\vartheta_N \to \vartheta^* \text{ in } L^\infty(I; L^4(\Omega_{\eta_N})),
\]
\[
\vartheta_N \to \vartheta \text{ in } L^{\beta}(I; L^{\beta}(\Omega_{\eta_N})),
\]
\[
\nabla \vartheta_N \to \nabla \vartheta \text{ in } L^2(I; L^2(\Omega_{\eta_N})),
\]
for some \( p \in L^p(I \times \omega) \). This implies
\[
\eta_N \to \eta \text{ in } C(\overline{\Omega} \times \omega).
\]

Compactness of \( \vartheta_N \) can be shown as in [24] Chapter 3, Section 3.5.3., using (4.14). It is based on local arguments, which are not effected by the moving shell. Consequently we have
\[
\vartheta_N \to \vartheta \text{ in } L^4(I; L^4(\Omega_{\eta_N})).
\]
In order to pass to the limit in various terms in the equations we are concerned with the compactness of \( \varrho_N \). Applying Corollary 2.12 yields
\[
\varrho_N \to \varrho \text{ in } L^2(I; L^2(\Omega_{\eta_N})).
\]
We aim at improving the exponent from 2 to \( \beta \) in order to pass to the limit in the pressure. Testing the continuity equation with \( \eta N \) yields

\[
\int_{\Omega_{\eta N}} \eta^2 dx + \int_0^t \int_{\Omega_{\eta N}} \frac{4(\beta - 1)}{\beta} \varepsilon |\nabla \eta N|^2 dx d\sigma
\]

\[
= \int_{\Omega_{\eta N}} \eta^2 dx - \int_0^t \int_{\Omega_{\eta N}} \eta N \eta N \cdot \nabla u_N dx d\sigma.
\]

Since \( p > \beta \), we find that the right hand side is uniformly bounded recalling (4.20) and (4.22). We conclude (for a non-relabelled subsequence)

\[
\eta N \rightarrow \eta \quad \text{in} \quad L^2(I; L^\beta(\Omega_{\eta N})).
\]

We are, however, still concerned with the term

\[
\int_{\Omega_{\eta N}} \eta^2 dx + \int_0^t \int_{\Omega_{\eta}} 2\varepsilon |\nabla \eta|^2 dx d\sigma
\]

and applying Theorem 3.1 (b) to the limit version. Due to (4.20), (4.28) and the strong convergence of \( \eta N \) we can pass to the limit in all terms in (4.29) except for the one containing \( \eta N \) which requires compactness of \( \eta N \). As for (4.29), we have

\[
\lim_{N \to \infty} \int_{\Omega_{\eta N}} |\nabla \eta N|^2 dx d\sigma = \int_{\Omega_{\eta}} |\nabla \eta|^2 dx d\sigma
\]

for all \( t \in I \), which implies strong convergence of \( \eta N \) and hence by (4.20)

\[
\lim_{N \to \infty} \int_{\Omega_{\eta N}} \nabla \eta N \cdot \phi dx dt = \int_{\Omega_{\eta}} \nabla \eta \cdot \phi dx dt.
\]

4.3. Compactness of \( \partial_t \eta N \). The effort of this subsection is to prove that

\[
\partial_t \eta N \to \partial_t \eta \quad \text{in} \quad L^2(I; L^2(\omega)).
\]

We will show this convergence in the generality we will need also in the subsequent limit procedures in the next section. In particular, we will not make use of any higher regularity beyond \( L^\infty(I; L^\gamma(\omega)) \) with \( \gamma > \frac{2p}{p-2} \) for the density.

The following aim is showing

\[
\int_{\Omega_{\eta N}} |\nabla \eta N|^2 dx dt + \int_{\Omega_{\eta}} |\partial_t \eta N|^2 dy dt
\]

\[
\to \int_{\Omega_{\eta N}} |\nabla u|^2 dx dt + \int_{\Omega_{\eta}} |\partial_t \eta|^2 dy dt,
\]

which implies the strong convergence (4.31) by the strict convexity of the \( L^2 \)-norm. Relation (4.32) will be a consequence of

\[
\int_{\Omega_{\eta N}} \eta N \cdot \mathcal{F} \eta N \partial_t \eta N dx dt + \int_{\Omega_{\eta}} |\partial_t \eta N|^2 dy dt
\]

\[
\to \int_{\Omega_{\eta N}} \eta \cdot \mathcal{F} \eta \partial_t \eta dx dt + \int_{\Omega_{\eta}} |\partial_t \eta|^2 dy dt
\]

and

\[
\int_{\Omega_{\eta N}} \eta N \cdot (\nabla u - \mathcal{F} \eta N \partial_t \eta N) dx dt \to \int_{\Omega_{\eta}} \eta \cdot (\nabla u - \mathcal{F} \eta \partial_t \eta) dx dt.
\]
First observe that (due to the trace theorem Lemma 2.3) we find that \( \partial_t \eta_N \) possesses some compactness in space. To be precise, we have
\[
(4.35) \quad \| \partial_t \eta_N \|_{L^2(I, W^{1,2} \circ \eta_N)} + \| \partial_t \eta_N \|_{L^2(I, L^2(\omega))} \leq c
\]
for all \( r < 2 \) and \( \ell < 4 \). The bounds only depend on the \( L^2(W^{1,2}_v) \) bounds of \( u_N \) and hence are uniform by estimates (4.19) and (4.20). We define the projection
\[
P_N w = \sum_{k=1}^N \alpha_k(w) w_k,
\]
where \( \alpha_k(w) = \langle w, w_k \rangle_{W^{1,2}(\omega)} \) if \( w_k = \tilde{Y}_\ell \) for some \( \ell \in \mathbb{N} \) and \( \alpha_k(w) = 0 \) otherwise. Obviously, we have
\[
\text{tr}_\chi P_N w = P_N w \text{ for any } w \in W^{3,2}(\omega).
\]
We have by definition,
\[
(4.36) \quad \| P_N w \|_{W^{3,2}(\omega)} \leq \| w \|_{W^{3,2}(\omega)} \quad \forall w \in W^{3,2}(\omega).
\]
The eigenvalue equation for the basis vectors implies additionally that
\[
(4.37) \quad \| P_N w \|_{L^2(\omega)} \leq c \| w \|_{L^2(\omega)} \quad \forall w \in L^2(\omega).
\]
Moreover, by definition of \( Y_k \) and \( \mathcal{P}_\chi \) (see Section 2.4) we have
\[
(4.38) \quad P_N w = \mathcal{P}_\chi (P_N w)
\]
for all \( w \in W^{3,2}(\omega) \). Finally, we note that \( P_N \eta_N = \eta_N \) such that \( (\eta_N, \mathcal{P}_\eta \eta_N) \) is admissible in (4.4).

Due to the uniform a priori bounds from the last subsection and the respective embeddings, we find that the convergence in (4.34) follows directly from Lemma 2.11 with the choice \( v_N = u_N - \mathcal{P}_\eta \partial_t \eta_N, \tau_N = P_N^{(s)} (g_N u_N) \) (which solves the projected equation (4.35) in the domain \( \Omega_{\eta_N} \)) and the continuity of the projection \( P_N^{(s)} \) defined above (recall also (4.38)). The corresponding uniform estimates are given in the previous subsection and the weak convergence of \( \mathcal{P}_\eta \partial_t \eta_N \) follows from (4.11), (4.12), Lemma 2.7 and Corollary 2.8.

In order to prove (4.33) we need to make use of the coupled momentum equation using Theorem 2.13. We define \( g_N = (\partial_t \eta_N, g_N u_N, \mathcal{P}_\eta \partial_t \eta_N) \) and \( f_N = (\partial_t \eta_N, \mathcal{P}_\eta \partial_t \eta_N) \) noticing that (by construction) \( \Omega_{\eta_N} \subset \Omega \cup S_{L/2} \) as well as for all \( s < \frac{1}{2} \) and \( q < 3 \)
\[
(4.39) \quad f_N \rightharpoonup f \quad \text{in} \quad L^2(I; X),
\]
where \( f = (\partial_t \eta, \mathcal{P}_\eta \partial_t \eta) \) and
\[
(4.40) \quad g_N \rightharpoonup g \quad \text{in} \quad L^2(I; X'),
\]
where
\[
X = L^2(\omega) \times W^{s,q}(\Omega \cup S_{L/2})
\]
with \( s_x < s_y < \frac{1}{2} \) (such that \( X' = L^2(\omega) \times W^{-s,q'}(\Omega \cup S_{L/2}) \)), since
\[
(4.41) \quad \begin{cases} 
\mathcal{P}_\eta u_N \rightharpoonup u & \text{in} \quad L^2(I; L^{\frac{6}{5-s}}(\Omega_{\eta_N})) \\
\end{cases}
\]
and \( L^\frac{6}{5-s} \hookrightarrow W^{-s,q'} \) due to \( \gamma > \frac{12}{7} \) (choosing \( s_x \) sufficiently close to \( 1/2 \) and \( q \) close to \( 3 \)). Further we define
\[
Z = W^{1,2}(\omega) \times W^{1,q}(\Omega \cup S_{L/2})
\]
Boundedness of \( g_N \) in \( L^\infty(I; Z') \) follows now from (4.19), \( g_N u_N \in L^2(L^\frac{6}{5-s}(\Omega_{\eta_N})) \) uniformly and the embedding
\[
L^\frac{6}{5-s} \hookrightarrow W^{-1,2} \hookrightarrow W^{-1,q}(\Omega) \quad \text{for} \quad \beta > \frac{3}{2} \quad \text{and} \quad q \geq 2. \quad \text{The conditions (a) in Theorem 2.13 follow now from (4.19) and (4.40) by weak compactness. For (b) we observe that we may assume that a regularizer \( b \mapsto (b)_\kappa \) exists such that for any \( s, a \in \mathbb{R} \) and \( p \in [1, \infty) \)
\]
\[
(4.42) \quad \| b - (b)_\kappa \|_{W^{s,q}(\omega)} \leq c \kappa^{s-a} \| b \|_{W^{s,p}(\omega)}, \quad b \in W^{s,p}(\omega).
\]

The estimate is well-known for \( a, s \in \mathbb{N}_0 \), while the general case follows by interpolation and duality. Moreover, since we use standard Fourier bases in \( W^{3,2}(\omega) \) for the discretisation of \( \eta_N \), we find by interpolation that the projection error satisfies the following stability estimates for all \( s \in [0,3] \)
\[
(4.43) \quad \| P_N b \|_{W^{s,q}(\omega)} \leq c \| b \|_{W^{s,q}(\omega)}.
\]

\(^*\)Here, this follows easily from (4.28), but it will be critical in the final limit \( \delta \to 0 \).
Next we introduce the mollification operator on $\partial_t \eta_N$ by considering for $\kappa > 0$ and $N \in \mathbb{N}$ $\mathcal{P}^N ((\partial_t \eta_N)_{\kappa})$ and set

$$f_{N,\kappa}(t) := (\mathcal{P}^N ((\partial_t \eta_N(t))_{\kappa}), \mathcal{F}_{\eta_N}(t)(\mathcal{P}^N ((\partial_t \eta_N(t))_{\kappa}))).$$

We find by the continuity of the mollification operator from (4.42), the continuity of the projection operator from (4.32) and the estimate for the extension operator (due to (4.17) and Lemma 2.7) that for a.e. $t \in (0,T)$

$$\|f_{N,\kappa} - f_N\|_{L^2(\omega) \times W^{1,2}((\Omega \cup \mathcal{S}_L)_{\overline{\omega}})} \leq ce^{c_4 - s} \|\partial_t \eta_N\|_{W^{1,2}(\omega)},$$

which can be made arbitrarily small in $L^2$ choosing $\kappa$ appropriately, cf. (4.39). Similarly, we have

$$\|f_{N,\kappa}\|_{W^{1,2}(\omega) \times W^{1,2}((\Omega \cup \mathcal{S}_L)_{\overline{\omega}})} \leq ce^{c_4} \|\partial_t \eta_N\|_{L^2(\omega)}.$$

Moreover, by (4.39) we clearly can deduce a converging subsequence such that $f_{N,\kappa} \to f_\kappa$ (for some $f_\kappa$) in $L^2(I;X)$ for any $\kappa > 0$, which implies (b).

For (c) have to control $(g_N(t) - g_N(s), f_{N,\kappa}(t))$ and hence decompose

$$(g_N(t) - g_N(s), f_{N,\kappa}(t)) = \left((g_N(t), \mathcal{P}^N ((\partial_t \eta_N(t))_{\kappa}), \mathcal{F}_{\eta_N}(t)(\mathcal{P}^N ((\partial_t \eta_N(t))_{\kappa}))), \right.$$

$$- (g_N(s), \mathcal{P}^N ((\partial_t \eta_N(s))_{\kappa}), \mathcal{F}_{\eta_N}(s)(\mathcal{P}^N ((\partial_t \eta_N(s))_{\kappa}))))$$

$$+ (g_N(s), (0, \mathcal{F}_{\eta_N}(t)(\mathcal{P}^N ((\partial_t \eta_N(t))_{\kappa})), \mathcal{F}_{\eta_N}(s)(\mathcal{P}^N ((\partial_t \eta_N(s))_{\kappa})))) \equiv: (I) + (II).$$

We begin estimating (II) using Corollary 2.8 to find that

$$(II) = \int_s^t \int_{\Omega_{\eta_N}(s)} g_N(s)u_N(s) \cdot \partial_\theta \mathcal{F}_{\eta_N}(t)(\mathcal{P}^N ((\partial_t \eta_N)_{\kappa}))(t) \, dx \, d\theta$$

$$\leq c \|g_N u_N(s)\|_{L^\infty(\Omega_{\eta_N}(s))} \|s - t\|^{\frac{1}{2}} \left(\int \|\partial_t \eta_N(t)\|_{L^2(\omega)}^2 \right)^{\frac{1}{2}}$$

$$\|\mathcal{P}^N ((\partial_t \eta_N)_{\kappa})\|_{L^T(\omega)},$$

for some $\ell < 4$ (recall that $\gamma > \frac{12}{7}$). By Sobolev’s embedding’s, (4.42) and (4.43) the last term can be estimated by

$$\|\mathcal{P}^N ((\partial_t \eta_N)_{\kappa})\|_{L^T(\omega)} \leq c \|\mathcal{P}^N ((\partial_t \eta_N)_{\kappa})\|_{W^{1,2}(\omega)} \leq c \|\partial_t \eta_N\|_{W^{1,2}(\omega)}$$

$$\leq ce^{-3} \|\partial_t \eta_N\|_{L^2(\omega)},$$

which is bounded to to (1.18). Using also (4.35) we conclude

$$(II) \leq c(\kappa) \|g_N u_N(s)\|_{L^\infty(\Omega_{\eta_N}(s))} \|s - t\|^{\frac{1}{2}}$$

The term (I) is estimated using the test-function $I_{(s,t)} f_{N,\kappa}$ in (1.16). One obtains the uniform Hölder estimate in a similar sense as for (I) using the various estimates on the extension, projections, embeddings and Hölder’s inequality. We explain here in detail only the two most complicated terms stemming from the time derivative and the pressure. All other terms can be estimated analogously by simpler means. First, we consider the term acting on the time derivative. Observe that this term only appears due to the time-dependent extension. We choose $a$ such that $\frac{1}{\theta} + \frac{1}{\gamma} + \frac{1}{\delta} = 1$. Then by the assumption $\gamma > \frac{12}{7}$, we find that $a < 4$. Hence we can choose $a_0 \in (a,4)$ and $\chi \in (0,1)$ such that $\frac{1}{\theta} = \frac{1}{a} + \frac{1-\chi}{a_0}$ and

$$\|\partial_t \eta_N\|_{L^\infty(\omega)} \leq \|\partial_t \eta_N\|_{L^2(\omega)}^{1-\chi} \|\partial_t \eta_N\|_{L^3(\omega)}^{\chi}.$$
where the constant depends on the a priori estimates only. As far as the pressure is concerned, Hölder’s inequality and Lemma \[2.5\](b) imply
\[
\left| \int_s^t \int_{\Omega_{\eta_N}(t)} p_s(\varrho_N, \vartheta_N) \text{div} \mathcal{F}_{\eta_N}(\theta)(\mathcal{P}^N((\partial_t \eta_N)_\kappa))(t) \, dx \, d\theta \right|
\leq c \|p_s(\varrho_N, \vartheta_N)\|_{L^1(I; L^1(\Omega_{\eta_N}))} \int_s^t \|\nabla \mathcal{F}_{\eta_N}(\theta)(\mathcal{P}^N((\partial_t \eta_N)_\kappa))\|_{L^\infty(\Omega_{\eta_N})} \, d\theta
\leq c \int_s^t (1 + \|\nabla \eta_N\|_{L^\infty(\omega)})\|\mathcal{P}^N((\partial_t \eta_N)_\kappa)\|_{W^{1,\infty}(\omega)} \, d\theta
\leq c \int_s^t (1 + \|\nabla \eta_N\|_{L^\infty(\omega)})\|\partial_t \eta_N\|_{W^{2,\infty}(\omega)} \, d\theta
\leq c \kappa^{-3} \|\partial_t \eta_N\|_{L^2(I; L^2(\omega))} \int_s^t (1 + \|\nabla \eta_N\|_{L^\infty(\omega)}) \, d\theta
\leq c \kappa^{-3} \|\partial_t \eta_N\|_{L^2(I; L^2(\omega))} \int_s^t (1 + \|\nabla \eta_N\|_{L^\infty(\omega)}) \, d\theta
\leq c \kappa^{-3} \|\partial_t \eta_N\|_{L^2(I; L^2(\omega))} \int_s^t (1 + \|\nabla \eta_N\|_{L^\infty(\omega)}) \, d\theta
\leq c \kappa^{-3} |t - s|^{\frac{3}{2}} \left( \int_I (1 + \|\nabla \eta_N\|_{L^\infty(\omega)}) \, d\theta \right)^{\frac{1}{2}} \leq c \kappa^{-3} |t - s|^{\frac{3}{2}}
\]
provided that we have
\[
(4.45) \quad p_s(\varrho_N, \vartheta_N) \in L^\infty(I; L^1(\Omega_{\eta_N})), \quad \partial_t \eta_N \in L^\infty(I; L^2(\omega)),
(4.46) \quad \nabla \eta_N \in L^2(I; L^2(\omega)),
\]
uniformly in \(N\). While \[4.43\] follows here and on the subsequent directly form the energy estimates, we need some further regularity for \[4.46\]. On this level it follows from the regularisation of the shell equation, cf. \[4.11\].

In conclusion, we can now choose \(\alpha \in (0, 1)\) close enough to one and conclude that for \(\tau > 0\) and \(t \in [0, T - \tau]\)
\[
\left| \int_0^\tau \langle g_N(t) - g_N(t + s), f_N,\kappa(t) \rangle \, ds \right| \leq c \kappa^{-3} \tau^{1/2} \left( A_N(t) + 1 \right),
\]
where
\[
A_N(t) = \|g_N(t)\|_{X^*}^2 + \|f_N(t)\|_{X^*}^2 + \|g_N u_N(t)\|_{L^{q(\eta_N)}(\Omega_{\eta_N})} \leq \int_0^\tau \left( \|g_N(s)\|_{X^*}^2 + \|f_N(s)\|_{X^*}^2 + \|g_N u_N(s)\|_{L^{q(\eta_N)}(\Omega_{\eta_N})} \right) \, ds
\]
uniformly bounded in \(L^1(I)\) due to \[4.39\] and \[4.40\] and \[4.41\]. Finally, the condition on (4) follows by the usual compactness in (negative) Sobolev spaces.

### 4.4. Compactness of the shell energy.
In order to complete the proof of \[4.4\] it remains to justify the limit in the shell energy. Since we have a regularized system \[1.17\] yields for any \(q < \infty\)
\[
(4.47) \quad \eta_N \to \eta \quad \text{in} \quad L^q(I; W^{2,q}(\omega)),
\]
which is enough to conclude
\[
(4.48) \quad \lim_{N \to \infty} \int_I \psi K(\eta_N) \, dt = \int_I \psi K(\eta) \, dt
\]
for all \(\psi \in C^\infty_c(I)\) (this step will be much harder on the subsequent levels, see Section \[5.2\]). It remains to show the convergence of the regularizer
\[
(4.49) \quad \lim_{N \to \infty} \int_I \psi \mathcal{L}(\eta_N) \, dt = \int_I \psi \mathcal{L}(\eta) \, dt.
\]
Next, we can assume that
\[
(4.50) \quad \partial_t \eta_N \to \partial_t \eta \quad \text{in} \quad L^2(I; W^{1-1/r,r}(\omega)),
\]
for all $r < 2$ due to (4.35). We infer from (4.6) using $(\psi \eta_N, \psi \mathcal{F}_{\eta_N}(\eta_N))$ as a test-function

$$
\int_I \psi \int_\omega K'_{\varepsilon}(\eta_N) \eta_N \, dy \, dt = \int_I \int_{\Omega_N} g_N u_N : \nabla \mathcal{F}_{\eta_N}(\eta_N) \, dx \, dt
+ \int_I \psi \int_{\Omega_N} g_N u_N \otimes u_N : \nabla \mathcal{F}_{\eta_N}(\eta_N) \, dx \, dt
+ \int_I \psi \int_{\Omega_N} \mathbf{S}^r(\theta_N, \nabla u_N) : \nabla \mathcal{F}_{\eta_N}(\eta_N) \, dx \, dt
$$

(4.51)

$$
\int_I \psi \int_{\Omega_N} \left( \rho_N(\eta_N, \partial_t \eta_N) \text{ div } \mathcal{F}_{\eta_N}(\eta_N) + \varepsilon \nabla \rho_N \nabla u_N : \mathcal{F}_{\eta_N}(\eta_N) \right) \, dx \, dt
+ \int_I \psi \int_\omega \partial_t \eta_N \partial_t (\psi \eta_N) \, dy \, dt
+ \int_I \psi \int_{\Omega_N} g_N \mathbf{F} \cdot \mathcal{F}_{\eta_N}(\eta_N) \, dx \, dt
+ \int_I \psi \int_\omega g \eta_N \, dy \, dt
+ \psi(0) \int_{\Omega_N(0)} q_0 \cdot \mathcal{F}_{\eta_N}(\eta_N)(0, \cdot) \, dx + \psi(0) \int_\omega \eta_1 \eta_N \, dy.
$$

The terms on the right-hand side related to the shell clearly converge to their expected limits because of (4.14) and (4.31). On account of Lemma 2.7 and Corollary 2.8 we have

$$
\| \partial_t (\mathcal{F}_{\eta_N}(\eta_N)) \|_{L^2(I; L^q(\Omega \cup S_{L/2}))} + \| \mathcal{F}_{\eta_N}(\eta_N) \|_{L^\infty W^{1, q_2}} + \| \mathcal{F}_{\eta_N}(\eta_N) \|_{L^\infty W^{1, q_3}} \leq c
$$

uniformly in $N$ for all $q_1 < 4$, $q_2 < \infty$ and $q_3 < 2$, cf. (4.14) and (4.31). In particular, applying standard compact embeddings we can choose a subsequence (not relabelled) such that

$$
\partial_t (\mathcal{F}_{\eta_N}(\eta_N)) \rightharpoonup \partial_t (\mathcal{F}_\eta(\eta)) \text{ in } L^2(I; L^q(\Omega \cup S_{L/2})),
\mathcal{F}_{\eta_N}(\eta_N) \rightharpoonup \mathcal{F}_\eta(\eta) \text{ in } L^2(I; W^{1, q_2}(\Omega \cup S_{L/2})),
\mathcal{F}_{\eta_N}(\eta_N) \rightharpoonup \mathcal{F}_\eta(\eta) \text{ in } L^\infty(I; L^\infty(\Omega \cup S_{L/2})),
$$

for all $q_1 < 4$ and $q_2 < \infty$. Combining these convergences with the convergences form the last subsection we can pass to the limit in the terms on the right-hand side of (4.51) related to the fluid system as well. On the other hand, the resulting expression coincides with $\int_I \psi \mathcal{K}(\eta) \, dt$ as can be seen from testing the limit system with $(\psi \eta, \psi \mathcal{F}_\eta(\eta))$. We conclude that

$$
\varepsilon \int_I \psi \mathcal{L}(\eta_N) \, dt = \frac{\varepsilon}{2} \int_I \psi \mathcal{L}'(\eta_N) \eta_N \, dt = \frac{1}{2} \int_I \psi K'_\varepsilon(\eta_N) \eta_N \, dt - \frac{1}{2} \int_I \psi K'(\eta_N) \eta_N \, dt
\rightarrow \frac{1}{2} \int_I \psi K'_\varepsilon(\eta) \eta \, dt - \frac{1}{2} \int_I \psi K'(\eta) \eta \, dt = \varepsilon \int_I \psi \mathcal{L}(\eta) \, dt
$$

as $N \to \infty$ due to (4.48). Combing this with (4.47) shows that (4.49) must be true. Combining all the convergences proven above allows us to pass to the limit in the energy balance (4.14) and to conclude that

$$
- \int_I \partial_t \psi \mathcal{E}_{\varepsilon, \delta} \, dt = \psi(0) \mathcal{E}_{\varepsilon, \delta}(0) + \int_I \psi \int_{\Omega_{\eta}} \frac{\varepsilon \varepsilon}{2} \psi \int_{\Omega_{\eta}} \frac{\delta}{\delta \eta} - \varepsilon \psi \delta \, dx \, dt
+ \int_I \psi \int_{\Omega_{\eta}} \partial_t H \, dx \, dt + \int_I \psi \int_{\Omega_{\eta}} \partial_t q \, dx \, dt + \int_I \psi \int_{\Omega_{\eta}} \partial_t \eta \, dy \, dt
$$

with

$$
\mathcal{E}_{\varepsilon, \delta}(t) = \int_{\Omega_{\eta(t)}} \left( \frac{1}{2} q(t) |u(t)|^2 + q(t) \varepsilon \delta \left( q(t), \partial(t) \right) \right) \, dx + \int_{\omega} \frac{|\partial_\eta \eta(t)|^2}{2} \, dy + K_\varepsilon(\eta(t)).
$$

4.5. End of the proof of Theorem 4.3 We have shown that a subsequence can be chosen that inhibits the necessary compactness properties to satisfy (K1) and (K4). The entropy inequality (K2) follows by further convergence properties of $\partial$ and weak sequential lower semi-continuity of the various convex terms. Due to its local character, the limit passage can be obtained without further difficulty by applying the methodology developed in [21, Chapter 3]. Finally, we may also pass to the limit with the momentum equation (4.6) and establish (K1). First observe that the necessary convergence of the approximate solutions has been shown in Subsection 4.2. Hence we are left to show the convergence of the test-functions. For that please observe that
with $N \to \infty$ (and $k \in \{1, \ldots, N\}$) in various spaces including $L^q(I, W^{1,q}(\Omega_{\eta_N})) \cap W^{1,2}(I, L^2(\Omega_{\eta_N}))$ for $q < \infty$ and $\alpha < 4$. Indeed, the a-priori estimates on $\eta_N, \partial_t \eta_N, \nabla \eta_N$ transfer to $\Psi^{-1}_N$ by the respective calculations and estimates in Section 2. Finally, we observe that the linear hull of $\{\omega_k, \omega_k \circ \Psi^{-1}_N\}_{k \in \mathbb{N}}$ exhibits the full space of test functions. Hence we conclude that [K1] is satisfied.

5. Construction of a solution.

In this section we pass to the limit in the approximate equations. For technical reasons the limits $\varepsilon \to 0$ and $\delta \to 0$ have to be performed independently from each other. For the greater part of this Section we study the limit $\varepsilon \to 0$ in the approximate system [K1]-[K4] and only highlight the difference in the $\delta$-limit.

5.1. The limit system for $\varepsilon \to 0$. We wish to establish the existence of a weak solution $(\eta, u, \varrho, \vartheta)$ to the system with artificial pressure in the following sense: We define

$$\tilde{W}_0^I = C_w(T^\lambda(\Omega))$$

as the function space for the density, whereas the other function spaces are defined in Section 2. A weak solution is a quadruplet $(\eta, u, \varrho, \vartheta) \in Y^I \times X^I_\eta \times \tilde{W}_0^I \times Z^I_\eta$ that satisfies the following.

(D1) The momentum equation holds in the sense that

$$\begin{align}
\int_I \frac{d}{dt} \int_{\Omega_\eta} \varrho u \cdot \phi \, dx - \int_I \int_{\Omega_\eta} \left( \varrho u \cdot \partial_t \phi + \varrho u \otimes u : \nabla \phi \right) \, dx \, dt + \int_I \int_{\Omega_\eta} \mathbf{S}(\vartheta, \nabla u) : \nabla \phi \, dx \, dt - \int_I \int_{\Omega_\eta} p_b(\varrho, \vartheta) \text{ div } \phi \, dx \, dt \\
+ \int_I \left( \frac{d}{dt} \int_{\omega} \partial_t \eta b \, dy - \int_{\omega} \partial_t \eta \partial_t b \, dy + \int_{\omega} K'(\eta) b \, dy \right) \, dt \\
= \int_I \int_{\Omega_\eta} \varrho \phi \cdot \phi \, dx \, dt + \int_I \int_{\Omega_\eta} \varrho b \phi \, dx \, dt
\end{align}
$$

for all $(b, \phi) \in C^\infty(\omega) \times C^\infty(T \times \mathbb{R}^3)$ with $\text{tr}_\eta b = \varrho \vartheta$. Moreover, we have $(\varrho u)(0) = \varrho_0$, $\eta(0) = \eta_0$ and $\partial_t \eta(0) = \eta_1$. The boundary condition $\text{tr}_\eta u = \partial_t \eta \vartheta$ holds in the sense of Lemma 2.3.

(D2) The continuity equation holds in the sense that

$$\int_I \frac{d}{dt} \int_{\Omega_\eta} \varrho \psi \, dx \, dt - \int_I \int_{\Omega_\eta} \left( \varrho \partial_t \psi + \varrho u \cdot \nabla \psi \right) \, dx \, dt = 0$$

for all $\psi \in C^\infty(T \times \mathbb{R}^3)$ and we have $\varrho(0) = \varrho_0$.

(D3) The entropy balance

$$\begin{align}
\int_I \frac{d}{dt} \int_{\Omega_\eta} \varrho s(\vartheta, \vartheta) \psi \, dx \, dt - \int_I \int_{\Omega_\eta} \left( \varrho s(\vartheta, \vartheta) \partial_t \psi + \varrho s(\vartheta, \vartheta) \varrho \nabla \psi \right) \, dx \, dt \\
\geq \int_I \int_{\Omega_\eta} \frac{1}{\vartheta} \left[ \mathbf{S}(\vartheta, \nabla u) : \nabla \psi \right] + \frac{z(\vartheta)}{\vartheta} \left( \frac{\vartheta^2}{2} \frac{d}{\vartheta^2} \right) \right] \frac{1}{\vartheta^2} \nabla \psi \, dx \, dt \\
- \int_I \int_{\Omega_\eta} \frac{z(\vartheta)}{\vartheta} \left( \frac{\vartheta^2}{2} \frac{d}{\vartheta^2} \right) \frac{1}{\vartheta^2} \nabla \varrho \cdot \nabla \psi \, dx \, dt + \int_I \int_{\Omega_\eta} \frac{\varrho}{\vartheta} H \psi \, dx \, dt
\end{align}
$$

holds for all $\psi \in C^\infty(T \times \mathbb{R}^3)$ with $\vartheta \geq 0$. Moreover, we have $\lim_{\vartheta \to 0} \varrho s(\vartheta, \vartheta)(t) \geq \varrho_0 s(\vartheta, \vartheta_0)$ and $\partial_t \vartheta \partial_t \vartheta_\eta \vartheta \leq 0$.

(D4) The total energy balance

$$\begin{align}
- \int_I \partial_t \psi E_\delta \, dt = \psi(0)E_0(0) + \int_I \psi \int_{\mathbb{R}^3} \frac{\delta}{\vartheta^2} \, dx \, dt + \int_I \psi \int_{\Omega_\eta} \varrho H \, dx \, dt + \int_I \int_{\Omega_\eta} \varrho \phi \cdot u \, dx \, dt \\
+ \int_I \int_{\Omega_\eta} \varrho \partial_t \eta \, dy \, dt
\end{align}
$$

holds for any $\psi \in C^\infty_c([0, T])$. Here, we abbreviated

$$E_0(t) = \int_{\Omega_\eta(t)} \left( \frac{1}{2} \varrho(t) |u(t)|^2 + \varrho(t) c_s(\vartheta(t), \vartheta(t)) \right) \, dx + \int_{\omega} \frac{\left| \partial_t \eta(t) \right|^2}{2} \, dy + K(\eta(t)).$$
Theorem 5.1. Assume that we have for some $\alpha \in (0, 1)$ and $s > 0$

$$\frac{|q_0|^2}{q_0} \in L^1(\Omega_{\eta_0}), \quad q_0, \vartheta_0 \in C^{2,\alpha}(\Omega_{\eta_0}), \quad \eta_0 \in W^{3+s,2}(\omega), \quad \eta_1 \in L^2(\omega),$$

$$f \in L^2(I; L^\infty(\mathbb{R}^3)), \quad y \in L^2(I \times \omega), \quad H \in C^{1,\alpha}(\bar{T} \times \mathbb{R}^3), \quad H \geq 0.$$

Furthermore suppose that $q_0$ and $\vartheta_0$ are strictly positive and that (1.3) is satisfied. There is a solution $(\eta, u, q, \vartheta) \in Y^T \times X_T \times W^I_q \times Z_T^I$ to (D1)–(D4) Here, we have $I = (0, T_*)$, where $T_* < T$ only if $\lim_{s \to T^*} \|\eta(t, \cdot)\|_{L^\infty}(\omega) = \frac{L}{2}$ or the Kato energy degenerates (namely, if $\lim_{s \to T^*} \|\nabla\eta(s, y) = 0$ for some point $y \in \omega$).

Lemma 5.2. Under the assumptions of Theorem 5.1 the continuity equation holds in the renormalized sense as specified in Definition 2.15.

The proof of the above theorem and lemma will be split in several parts. For a given $\varepsilon$ we obtain a solution $(\eta_\varepsilon, u_\varepsilon, q_\varepsilon)$ to (K1)–(K4) by Theorem 1.3. As in the preceding Section we can combine the total energy balance (4.3) with the entropy balance (4.4) to obtain the total dissipation balance

$$\int_{\Omega_{\eta_\varepsilon}} \left( \frac{1}{2} \varepsilon \partial_t |u_\varepsilon|^2 + H_{\delta, \Theta}(\vartheta_\varepsilon, \vartheta_\varepsilon) \right) \, dx + \int_{\omega} \frac{|\partial_t \eta_\varepsilon|^2}{2} \, dy + K(\eta_\varepsilon)$$

$$= \int_{\Omega_{\eta_\varepsilon}} \int_{\eta_\varepsilon} \sigma_{\varepsilon, \Omega} \, dx + \int_{0}^{t} \int_{\Omega_{\eta_\varepsilon}} \varepsilon \vartheta_\varepsilon^5 \, dt$$

(5.5)

for any $0 \leq \tau \leq T$. Here $H_{\delta, \Theta}(\vartheta, \vartheta) = \vartheta (\vartheta_\varepsilon, \vartheta_\varepsilon - \Theta_\varepsilon(\vartheta, \vartheta))$ for some $\Theta > 0$ and

$$\sigma_{\varepsilon, \Omega} = \frac{1}{\varepsilon \vartheta_\varepsilon} \left( \mathbf{S}(\vartheta_\varepsilon, \nabla u_\varepsilon) \cdot \nabla u_\varepsilon \right) + \varepsilon (1 + \vartheta_\varepsilon) \max \{ |\nabla \vartheta_\varepsilon|, |\nabla u_\varepsilon| \} + \frac{\varepsilon \delta}{\vartheta_\varepsilon} \partial_\varepsilon \vartheta_\varepsilon^2 |\nabla \vartheta_\varepsilon|^2$$

$$+ \frac{1}{\varepsilon \vartheta_\varepsilon} \left( \varepsilon \partial_\varepsilon \vartheta_\varepsilon \right) |\nabla \vartheta_\varepsilon|^2 + \left[ \frac{\delta}{\varepsilon \vartheta_\varepsilon} \left( \vartheta_\varepsilon^{\beta-1} + \frac{1}{\vartheta_\varepsilon^2} \right) |\nabla \vartheta_\varepsilon|^2 + \frac{\delta}{\vartheta_\varepsilon^2} \right].$$

Absorbing the final term on the left-hand side of (5.5) into the left-hand side we deduce the bounds

$$\sup_{t \in I} \left[ \int_{\Omega_{\eta_\varepsilon}} \frac{1}{2} \varepsilon \partial_t |u_\varepsilon|^2 + H_{\delta, \Theta}(\vartheta_\varepsilon, \vartheta_\varepsilon) \right] \, dx \leq c$$

(5.6)

$$\sup_{I} \int_{\omega} \frac{|\partial_t \eta_\varepsilon|^2}{2} \, dy + \sup_{I} K(\eta_\varepsilon) + \varepsilon \sup_{I} \mathcal{L}(\eta_\varepsilon) \leq c.$$ (5.7)

In particular, we have

$$\sup_{t \in I} \|\vartheta_\varepsilon\|_{L^p(\Omega_{\eta_\varepsilon})}^p + \sup_{t \in I} \|\vartheta_\varepsilon u_\varepsilon\|_{L^\infty(\Omega_{\eta_\varepsilon})}^{\frac{2\beta}{\beta+1}} + \sup_{t \in I} \|\vartheta_\varepsilon\|_{L^p(\Omega_{\eta_\varepsilon})}^{\frac{2\beta}{\beta+1}} \leq c.$$ (5.8)

Moreover, boundedness of the entropy production rate

$$\|\sigma_{\varepsilon, \Omega}\|_{L^1(I \times \Omega_{\eta_\varepsilon})} \leq c$$

(5.9)

gives rise to

$$\varepsilon \|\nabla u_\varepsilon\|_{L^p(I \times \Omega_{\eta_\varepsilon})}^p + \varepsilon \|\nabla \vartheta_\varepsilon\|_{L^p(\Omega_{\eta_\varepsilon})} \leq c;$$ (5.10)

$$\|\mathcal{D}(u_\varepsilon)\|_{L^2(I \times \Omega_{\eta_\varepsilon})} + \|\nabla \vartheta_\varepsilon^2\|_{L^2(I \times \Omega_{\eta_\varepsilon})} + \|\nabla \vartheta_\varepsilon\|_{L^2(I \times \Omega_{\eta_\varepsilon})} \leq c;$$ (5.11)

whence, by Poincare’s inequality and (5.8),

$$\|u_\varepsilon\|_{L^2(I; W^{1,2}(\Omega_{\eta_\varepsilon}))} + \|\vartheta_\varepsilon\|_{L^2(I; W^{1,2}(\Omega_{\eta_\varepsilon}))} \leq c.$$ (5.12)

Finally, we deduce from the equation of continuity (4.2) (using the renormalized formulation from Theorem 5.1 (b) with $\theta(z) = z^2$ and testing with $\psi \equiv 1$) that

$$\int_{\Omega_{\eta_\varepsilon}(\varepsilon)} \vartheta_\varepsilon(t, \cdot) \, dx = \int_{\Omega_{\eta_\varepsilon}(\varepsilon)} \vartheta_0 \, dx, \quad \|\sqrt{\varepsilon} \nabla \vartheta_\varepsilon\|_{L^2(I \times \Omega_{\eta_\varepsilon})} \leq c.$$ (5.13)
Note that all estimates are independent of $\varepsilon$. Hence, we may take a subsequence such that for some $\alpha \in (0,1)$ we have
\begin{align}
\eta_t &\rightarrow^{*} \eta \quad \text{in} \quad L^\infty(I; W^{2,2}(\Omega)), \\
\varepsilon \eta_t &\rightarrow 0 \quad \text{in} \quad L^\infty(I; W^{3,2}(\Omega)), \\
\eta_t &\rightarrow^{*} \eta \quad \text{in} \quad W^{1,\infty}(I; L^2(\omega)), \\
\varepsilon \eta_t &\rightarrow \eta \quad \text{in} \quad C^\alpha(\overline{T \times \omega}), \\
\mathcal{D}(u_\varepsilon) &\rightarrow^n \mathcal{D}(u) \quad \text{in} \quad L^2(I; L^2(\Omega_\varepsilon)), \\
u_\varepsilon &\rightarrow^n u \quad \text{in} \quad L^2(I; L^2(\Omega_\varepsilon)), \\
\varepsilon \nu v_\varepsilon &\rightarrow^n 0 \quad \text{in} \quad L^p(I; W^{1,p}_g(\Omega_\varepsilon)), \\
\varepsilon \nabla v_\varepsilon &\rightarrow^n 0 \quad \text{in} \quad L^p(I \times \Omega_\varepsilon), \\
p_\varepsilon &\rightarrow^{*} p \quad \text{in} \quad L^\infty(I; L^2(\Omega_\varepsilon)), \\
\partial_\varepsilon &\rightarrow^{*} \partial \quad \text{in} \quad L^\infty(I; L^4(\Omega_\varepsilon)), \\
d_\varepsilon &\rightarrow^{*} d \quad \text{in} \quad L^\infty(I; L^{1,2}(\Omega_\varepsilon)).
\end{align}

We observe that the a-priori estimates (5.8) imply uniform bounds of $\varrho_\varepsilon u_\varepsilon$ in $L^\infty(I, L^{2,2}(\omega))$. Therefore, we may apply Lemma 2.11 with the choice $v_i \equiv u_\varepsilon$, $r_i = \varrho_\varepsilon$, $p = s = 2$, $b = \beta$ and $m$ sufficiently large to obtain
\begin{align}
\varrho_\varepsilon u_\varepsilon &\rightarrow^0 \varrho u \quad \text{in} \quad L^q(I, L^a(\Omega_\varepsilon)),
\end{align}
where $a \in (1, \frac{2\beta}{\beta + 6})$ and $q \in (1, 2)$. We apply Lemma 2.11 once more with the choice $v_i \equiv u_\varepsilon$, $r_i = \varrho_\varepsilon u_\varepsilon$, $p = s = 2$, $b = \frac{2\beta}{\beta + 1}$ and $m$ sufficiently large to find that
\begin{align}
\varrho_\varepsilon u_\varepsilon \otimes u_\varepsilon &\rightarrow^0 \varrho u \otimes u \quad \text{in} \quad L^1(I \times \Omega_\varepsilon).
\end{align}

We also obtain
\begin{align}
\varrho_\varepsilon u_\varepsilon &\rightarrow^0 \varrho u \quad \text{in} \quad L^q(I, L^q(\Omega_\varepsilon)), \\
\varrho_\varepsilon u_\varepsilon &\rightarrow^{0,*} \varrho u \quad \text{in} \quad L^\infty(I, L^{2,2}(\Omega_\varepsilon)),
\end{align}
for all $q < \frac{6\beta}{\beta + 6}$. Moreover, we have as a consequence of (5.18) and (5.24)
\begin{align}
S(\varrho_\varepsilon, \nabla u_\varepsilon) &\rightarrow^0 \mathfrak{S} \quad \text{in} \quad L^{4/3}(I, L^{4/3}(\Omega_\varepsilon))
\end{align}
for some limit function $\mathfrak{S}$. The convergence (5.14) and the assumption on $K$ (see Section 1.2) yields
\begin{align}
K'(\eta_\varepsilon) &\rightarrow^{*} K' \quad \text{in} \quad L^\infty(I, W^{-2,r}(\omega))
\end{align}
for any $r < 2$ with some limit quantity $K$.

At this stage of the proof the pressure is only bounded in $L^1$, so we have to exclude its concentrations. The nowadays common approach from [21] Chapter 3, Section 3.6.3 only works locally where the moving shell is not seen (see Lemma 5.3 below). The problem can be circumvented by excluding concentrations at the boundary (see Lemma 5.4 which is inspired by [27]). The proof is exactly as in [55, Lemma 6.4].

**Lemma 5.3.** Let $Q = J \times B \Subset I \times \Omega_\varepsilon$ be a parabolic cube. The following holds for any $\varepsilon \leq \varepsilon_0(Q)$
\begin{align}
\int_Q p_\varepsilon(\varrho_\varepsilon, \varrho_\varepsilon) \varrho_\varepsilon \, dx \, dt \leq C(Q)
\end{align}
with a constant independent of $\varepsilon$.

**Lemma 5.4.** Let $\kappa > 0$ be arbitrary. There is a measurable set $A_\varepsilon \Subset I \times \Omega_\varepsilon$ such that we have for all $\varepsilon \leq \varepsilon_0(\kappa)$
\begin{align}
\int_{I \times \mathbb{R}^3 \setminus A_\varepsilon} p_\varepsilon(\varrho_\varepsilon, \varrho_\varepsilon) \varrho_\varepsilon \chi_{\Omega_\varepsilon} \, dx \, dt \leq \kappa.
\end{align}

We connect Lemma 5.3 and Lemma 5.4 to obtain the following corollary.
Corollary 5.5. Under the assumptions of Theorem 5.7 there exists a function $\mathbf{\varphi}$ such that
\[ p_s(\varrho_\varepsilon, \vartheta_\varepsilon) \to^p \mathbf{\varphi} \text{ in } L^1(I; L^1(\Omega_\eta)), \]
and we can follow the arguments in [21, Chapter 3, Section 3.7.3] to conclude
\[ \int_{(I \times \Omega_\eta) \setminus A_\varepsilon} \mathbf{\varphi} \, dx \, dt \leq \kappa. \]

Combining Corollary 5.5 with the convergences (5.14)–(5.32) we can pass to the limit in (4.11) and (4.12) and obtain the following. There is
\[ (\eta, \mathbf{u}, \varrho, \vartheta, \mathbf{\varphi}) \in Y^I \times X^I_\eta \times \tilde{W}^I_\eta \times Z^I_\eta \times L^1(I \times \Omega_\eta) \]
that satisfies
\[ \mathbf{u}(\cdot, \cdot + \eta \nu) = \partial_t \eta \varrho_\eta \quad \text{in } I \times \omega, \]
the continuity equation
\[ \int_I \frac{d}{dt} \int_{\Omega_\eta} \rho \mathbf{u} \cdot \mathbf{\varphi} \, dx \, dt - \int_I \int_{\Omega_\eta} \left( \rho \partial_t \psi + \rho \mathbf{u} \cdot \nabla \mathbf{\varphi} \right) \, dx \, dt = 0 \]
for all $\psi \in C^\infty(T \times \mathbb{R}^3)$ and the coupled weak momentum equation
\[ \begin{aligned}
& \int_I \frac{d}{dt} \int_{\Omega_\eta} \rho \mathbf{u} \cdot \mathbf{\varphi} \, dx \, dt - \int_I \int_{\Omega_\eta} \left( \rho \mathbf{u} \cdot \partial_t \mathbf{\varphi} + \rho \mathbf{u} \otimes \varrho \mathbf{u} : \nabla \mathbf{\varphi} \right) \, dx \, dt \\
& + \int_I \int_{\Omega_\eta} \mathbf{S} : \nabla \mathbf{\varphi} \, dx \, dt - \int_I \int_{\Omega_\eta} \mathbf{\varphi} \, \text{div} \mathbf{\varphi} \, dx \, dt \\
& + \int_I \int_{\Omega_\eta} \partial_t \eta \sigma_b \, dy - \int_I \int_{\Omega_\eta} \partial_t \mathbf{\varphi} \cdot \mathbf{b} \, dy + \int_I \int_{\Omega_\eta} \mathbf{K} \cdot \mathbf{b} \, dy \, dt \\
& = \int_I \int_{\Omega_\eta} \rho \mathbf{f} \cdot \mathbf{\varphi} \, dx \, dt + \int_I \int_{\Omega_\eta} g \mathbf{b} \, dx \, dt.
\end{aligned} \]

for all $(\mathbf{f}, \mathbf{\varphi}) \in C^\infty(\omega) \times C^\infty(T \times \mathbb{R}^3)$ with $\text{tr}_\eta \mathbf{\varphi} = b \nu$. It remains to show strong convergence of $\vartheta_{\varepsilon}, \varrho_{\varepsilon}$ and $\nabla^2 \eta_{\varepsilon}$. The convergence proof for $\vartheta_{\varepsilon}$ is entirely based on local arguments. Consequently the shell is not seen and we can follow the arguments in [21 Chapter 3, Section 3.7.3] to conclude
\[ \vartheta_{\varepsilon} \to^p \vartheta \text{ in } L^4(I \times \Omega_{\varepsilon}). \]

This yields $\mathbf{S} = \mathbf{S}(\vartheta, \nabla \mathbf{u})$ in (5.37). Additionally we can pass to the limit in the entropy balance (4.3) using lower semi-continuity. The remainder of this subsection is dedicated to the proof of $\mathbf{\varphi} = p(\varrho, \vartheta)$. Eventually, we will pass to the limit in the shell energy in Section 5.2 which will finish the proof of Theorem 5.1. The proof of strong convergence of the density is based on the effective viscous flux identity introduced in (32) and the concept of renormalized solutions from (13). Arguing locally, there is no difference to the known setting and we can follow the arguments in [21 Chapter 3, Section 3.6.5]. We consider a parabolic cube $\hat{Q} = J \times B$ with $Q \Subset \hat{Q} \Subset I \times \Omega_\eta$. Due to (5.17) we can assume that $\hat{Q} \Subset I \times \Omega_{\varepsilon}$ (by taking $\varepsilon$ small enough). For $\psi \in C^\infty_c(\hat{Q})$ we obtain
\[ \int_{I \times \mathbb{R}^3} \psi^2 \left( p_s(\varrho_{\varepsilon}, \vartheta_{\varepsilon}) - (\lambda(\vartheta_{\varepsilon}) + 2 \mu(\vartheta_{\varepsilon})) \text{div} \, \mathbf{u}_{\varepsilon} \right) \varrho_{\varepsilon} \, dx \, dt \]
\[ \quad \to \int_{I \times \mathbb{R}^3} \psi^2 \left( p - (\lambda(\vartheta) + 2 \mu(\vartheta)) \text{div} \, \mathbf{u} \right) \varrho \, dx \, dt \]
as $\varepsilon \to 0$ (note that the term related to $p_s$ disappears due to (5.20) provided we choose $\beta$ large enough). The proof of Lemma 5.2 follows exactly as in [6 Lemma 6.2]. So, for $\psi \in C^\infty(T \times \mathbb{R}^3)$ we have
\[ \int_I \frac{d}{dt} \int_{\mathbb{R}^3} \theta(\varrho) \psi \, dx \, dt \to \int_I \int_{\mathbb{R}^3} \left( \theta(\varrho) \, \text{div} \mathbf{u} - \partial_t \varrho \theta \psi \right) \, dx \, dt \]
\[ + \int_I \int_{\mathbb{R}^3} \left( \theta' \varrho - \theta(\varrho) \right) \varrho \mathbf{\varphi} \, dy \, dt \]
\[ - \int_I \int_{\mathbb{R}^3} \theta(\varrho) \partial_t \varphi \psi \, dx \, dt \]
\[ = \int_I \int_{\mathbb{R}^3} \theta(\varrho) \mathbf{\varphi} \cdot \nabla \psi. \]
Here $\mathcal{E}_0 : W^{1,2}(\Omega_0) \to W^{1,p}(\mathbb{R}^3)$ is the extension from [4, Lemma 2.5] where $1 < p < 2$ (but may be chosen close to 2). In order to deal with the local nature of (5.39) we use ideas from [18]. First of all, by the monotonicity of the mapping $\varphi \mapsto p(\varphi, \partial \varphi)$, we find for arbitrary non-negative $\psi \in C^\infty(\bar{Q})$

$$
\lim_{\varepsilon \to 0} \inf \int_{I \times \Omega_\varepsilon} \psi(\lambda(\partial \varepsilon) + 2\mu(\partial \varepsilon)) \left( \text{div} \ u_\varepsilon - \text{div} \ u_\varphi \right) \, dx \, dt
$$

$$
= \lim_{\varepsilon \to 0} \inf \int_{I \times \Omega_\varepsilon} \psi(\lambda(\partial \varepsilon) + 2\mu(\partial \varepsilon)) \left( \text{div} \ u_\varepsilon - (\lambda(\partial \varphi) + 2\mu(\partial \varphi)) \text{div} \ u_\varphi \right) \, dx \, dt
$$

$$
= \lim_{\varepsilon \to 0} \inf \int_{I \times \Omega_\varepsilon} \psi \left( (\lambda(\partial \varepsilon) + 2\mu(\partial \varepsilon)) \text{div} \ u_\varepsilon - (\lambda(\partial \varepsilon) + 2\mu(\partial \varepsilon)) \text{div} \ u_\varphi \right) \, dx \, dt
$$

$$
+ \lim_{\varepsilon \to 0} \inf \int_{I \times \Omega_\varepsilon} \psi \left( \left( p(\varphi, \partial \varepsilon) - \bar{p} \right) (\varphi - \varphi) \right) \, dx \, dt
$$

$$
= \lim_{\varepsilon \to 0} \inf \int_{I \times \Omega_\varepsilon} \psi \left( \left( p(\varphi, \partial \varepsilon) - \bar{p} \right) (\varphi - \varphi) \right) \, dx \, dt \geq 0
$$

using (5.39) as well as (5.38) (together with (2.4) and the uniform bounds (5.8) and (5.11)). As $\psi$ is arbitrary and $\mu$ strictly positive by (2.4) we conclude

$$
\overline{\text{div} \ u_\varphi} \geq \text{div} \ u_\varphi \quad \text{a.e. in} \ I \times \Omega_\varphi,
$$

where

$$
\text{div} u_e \varphi \to^\varepsilon \text{div} u_\varphi \quad \text{in} \ L^1(\Omega; L^1(\Omega_\varepsilon)),
$$

recall (5.18) and (5.22). Now, we compute both sides of (5.41) by means of the corresponding continuity equations. Due to Theorem 3.1 (b) with $\theta(z) = z \ln z$ and $\psi = \mathbb{I}_{(0,1)}$ we have

$$
\int_0^t \int_{\mathbb{R}^3} \text{div} \ u_e \varphi \, dx \, ds \leq \int_{\mathbb{R}^3} \varphi_0 \ln(\varphi_0) \, dx - \int_{\mathbb{R}^3} \varphi_e(t) \ln(\varphi_e(t)) \, dx.
$$

Similarly, equation (5.40) yields

$$
\int_0^t \int_{\mathbb{R}^3} \varphi \, dx \, ds \leq \int_{\mathbb{R}^3} \varphi_0 \ln(\varphi_0) \, dx - \int_{\mathbb{R}^3} \varphi(t) \ln(\varphi(t)) \, dx.
$$

Combining (5.41), (5.42) shows

$$
\limsup_{\varepsilon \to 0} \int_{\mathbb{R}^3} \varphi_e(t) \ln(\varphi_e(t)) \, dx \leq \int_{\mathbb{R}^3} \varphi(t) \ln(\varphi(t)) \, dx
$$

for any $t \in I$. This gives the claimed convergence $\varphi_e \to \varphi$ in $L^1(I \times \mathbb{R}^3)$ by convexity of $z \mapsto z \ln z$. Consequently, we have $\overline{\varphi} = p(\varphi, \varphi)$.

### 5.2. Compactness of the shell energy

All the forthcoming effort is to prove

$$
\lim_{\varepsilon \to 0} \int_I \int_{\omega} |\partial_t \eta_e(t)|^2 \, dy \, dt = \int_I \int_{\omega} |\partial_t \eta(t)|^2 \, dy \, dt,
$$

$$
\lim_{\varepsilon \to 0} \int_I K_e(\eta_e(t)) \, dt = \int_I K(\eta(t)) \, dt,
$$

as $\varepsilon \to 0$ at least for a subsequence. This will allow us to pass to the limit in the energy balance as well as in the nonlinear term of the shell equation. In the following we derive a framework to prove (5.44) based on fractional estimates. The same approach will be subsequently used in the limit passage $\delta \to 0$ in Section 5.3. The difference is that the bounds on the density will be more restrictive. We develop the theory here using only these weaker estimates to have it ready for the final limit procedure as well.

A first observation is that $\text{tr}_{\eta_e}(u_e) = \partial_t \eta_e \nu$ implies

$$
\partial_t \eta_e \to \partial_t \eta \quad \text{in} \ L^2(I; W^{1-1/r, r}(\omega)),
$$

for all $r < 2$ by (5.11) in combination with Lemma 2.3. In the following we are going to prove that

$$
\int_I \|\eta_n\|^2_{W^{2+r, 2}(\omega)} \, dt
$$
is uniformly bounded for some \( s > 0 \) using an appropriate test-function in the shell equation. On account of the coupling we need a suitable test-function for the momentum equation for it as well. Hence we set

\[
(\phi_\varepsilon, \varphi_\varepsilon) = \left( \mathcal{F}^{IV}_{\eta_\varepsilon}(\Delta^{s}_h \Delta^{s}_{\eta_\varepsilon} \varphi_\varepsilon - \mathcal{K}_{\eta_\varepsilon}(\Delta^{s}_h \Delta^{s}_{\eta_\varepsilon} \varphi_\varepsilon)), \Delta^{s}_h \Delta^{s}_{\eta_\varepsilon} \varphi_\varepsilon - \mathcal{K}_{\eta_\varepsilon}(\Delta^{s}_h \Delta^{s}_{\eta_\varepsilon} \varphi_\varepsilon) \right),
\]

where \( \mathcal{F}^{IV}_{\eta_\varepsilon} \) and \( \mathcal{K}_{\eta_\varepsilon} \) have been introduced in Proposition 2.9. Here \( \Delta^{s}_h v(y) = h^{-s} (v(y + h\eta_\varepsilon) - v(y)) \) is the fractional difference quotient in direction \( e_\alpha \) for \( \alpha \in \{1, 2, 3\} \). We obtain

\[
\int_I K'_\varepsilon(\eta_\varepsilon) \phi_\varepsilon \, dt
= \int_I \int_{\Omega^{v\varepsilon}(\varepsilon)} \left( \varrho_\varepsilon \mathbf{u}_\varepsilon \otimes \mathbf{u}_\varepsilon - \mathbf{S}(\vartheta_\varepsilon, \mathbf{u}_\varepsilon) - \varepsilon \mathbf{P}_\varepsilon \right) : \nabla \phi_\varepsilon + \mathbf{f} \cdot \phi_\varepsilon \, dx \, dt
+ \int_I \int_{\Omega^{v\varepsilon}(\varepsilon)} \varrho_\varepsilon \mathbf{u}_\varepsilon \cdot \partial_t \phi_\varepsilon \, dx \, dt
- \int_I \int_{\Omega^{v\varepsilon}(\varepsilon)} \varrho_\varepsilon \mathbf{u}_\varepsilon \cdot \phi_n \, dx + \int_\omega \partial_t \eta_\varepsilon \phi_\varepsilon \, dy \, dt
+ \int_\omega \left( \partial_t \eta_\varepsilon \partial_t \phi_\varepsilon + g \phi_\varepsilon \right) \, dy \, dt =: (I)_\varepsilon + (II)_\varepsilon + (III)_\varepsilon + (IV)_\varepsilon
\]

recalling that the function \( \phi_\varepsilon \) is divergence-free such that the pressure term disappears. Since \( \eta_\varepsilon \in L^\infty(I, W^{2,2}(\omega)) \) uniformly, we have

\[
\int_I \| \Delta^{s}_h \nabla^2 \eta_\varepsilon \|^2_{L^2(\omega)} \, dt \lesssim 1 + \int_I K''(\eta_\varepsilon) \phi_\varepsilon \, dt
\]

for every \( h > 0 \) and \( s \in (0, \frac{1}{2}) \) due to [29, Lemma 4.5]. Consequently, it holds

\[
\int_I \| \Delta^{s}_h \nabla^2 \eta_\varepsilon \|^2_{L^2(\omega)} \, dt + \varepsilon \int_I \| \Delta^{s}_h \nabla^3 \eta_{\varepsilon} \|^2_{L^2(\omega)} \, dt \lesssim 1 + (I)_\varepsilon + (II)_\varepsilon + (III)_\varepsilon + (IV)_\varepsilon
\]

and our task consists in establishing uniform estimates for the terms \((I)_\varepsilon, \ldots, (IV)_\varepsilon\). As far as \((I)_\varepsilon\) is concerned the most critical term is the convective term \( \varrho_\varepsilon \mathbf{u}_\varepsilon \otimes \mathbf{u}_\varepsilon \), with integrability \( \frac{\alpha}{\gamma + \alpha} > 1 \). By (5.12) and (5.14)

\[
\| \phi_\varepsilon \|_{L^q(I; W^{1,p}(\omega))} \leq \| \Delta^{s}_h \Delta^{s}_{\eta_\varepsilon} \eta_\varepsilon \|_{L^q(I; W^{1,p}(\omega))} + \| \Delta^{s}_h \Delta^{s}_{\eta_\varepsilon} \eta_\varepsilon \|_{L^q(I; L^p(\omega))}
\leq \| \eta_\varepsilon \|_{L^q(I; W^{1,p}\,\ast(\omega))} + \| \Delta^{s}_h \Delta^{s}_{\eta_\varepsilon} \|_{L^q(I; L^p(\omega))} \| \nabla \eta_\varepsilon \|_{L^q(I; L^p(\omega))}
\leq \| \eta_\varepsilon \|_{L^q(I; W^{1, p}\,\ast(\omega))} + \| \eta_\varepsilon \|_{L^q(I; L^p(\omega))} \| \nabla \eta_\varepsilon \|_{L^q(I; L^p(\omega))}
\leq \| \eta_\varepsilon \|_{L^q(I; W^{1, p}\,\ast(\omega))} + \| \eta_\varepsilon \|_{L^q(I; L^p(\omega))} + c_p
\]

for all \( s < \frac{1}{2}, \; p < \infty \) and \( q \in [1, \infty] \). For \( p = \frac{6}{5 + 2q - 1} \) we can choose \( s > 0 \) small enough such that \( W^{2,2}(\omega) \subset W^{1+2\alpha, p}(\omega) \). Using (5.14) again implies that \( \phi_\varepsilon \) is uniformly bounded in \( L^q(I; W^{1, p}) \). We conclude that \( (I)_\varepsilon \) is uniformly bounded in \( \varepsilon \) and \( h \) if we choose \( s \) small enough. The most critical term is in fact \((II)_\varepsilon\). We note that (5.18) and (5.22) imply

\[
\varrho_\varepsilon \mathbf{u}_\varepsilon \in L^2(I; L^{q_0}(\Omega^{v\varepsilon}(\varepsilon)))
\]

uniformly for all \( q_0 < \frac{6\gamma}{\gamma + \alpha} \). Due to the assumption \( \gamma > \frac{6\gamma}{3} \) we can choose in the above \( q_0 > \frac{3}{3} \). On the other hand we have

\[
\| \partial_t \phi_\varepsilon \|_{L^2(I; L^{q_0}(\Omega^{v\varepsilon}(\varepsilon)))} \lesssim \| \partial_t \Delta^{s}_h \Delta^{s}_{\eta_\varepsilon} \eta_\varepsilon \|_{L^2(I; L^{q_0}(\omega))} + \| \Delta^{s}_h \Delta^{s}_{\eta_\varepsilon} \eta_\varepsilon \|_{L^2(I; L^{q_0}(\omega))}
\lesssim \| \partial_t \eta_\varepsilon \|_{L^2(I; W^{2, q}(\omega))} + \| \Delta^{s}_h \Delta^{s}_{\eta_\varepsilon} \|_{L^2(I; L^{q_0}(\omega))}
\]

Thus, we can choose \( s \) small enough such that \( \partial_t \phi_\varepsilon \) is uniformly bounded in \( L^2(I; L^{q_0}(\omega)) \) thanks to (5.14) and (5.49) (together with Sobolev’s embedding and \( q_0 < 4 \)). We conclude boundedness of \((II)_\varepsilon\). As far as \((III)_\varepsilon\) is concerned, uniform bounds for the first term are easily obtained from (5.19) (choosing \( p > \beta > 2 \) and using Sobolev’s embedding) in combination with (5.27). For the second term we use

\[
\| \phi_\varepsilon \|_{L^2(I; L^{q_0}(\omega))} \lesssim \| \eta_\varepsilon \|_{L^2(I; W^{2, q}(\omega))} \lesssim \| \eta_\varepsilon \|_{L^2(I; W^{2, q}(\omega))}
\]

together with (5.14). The second term in \((IV)_\varepsilon\) is analogous. Finally, we can use again (5.49) to control the first term in \((IV)_\varepsilon\) and the proof of (5.35) is complete. Moreover we have shown

\[
\varepsilon \int_I \| \eta_\varepsilon \|^2_{V^{3,2}(\omega)} \, dt \leq c
\]
uniformly in $\varepsilon$. This, interpolated with $\|1\gamma\|_{L^2(\Omega_{\mbox{u}})} \to 0$ as $\varepsilon \to 0$, which completes the proof of 5.49. Finally we observe that the convergence in 5.44 follows exactly as was done in Subsection 4.3 (in particular, 5.47 implies the required to obtain a counterpart of 4.10). The proof is even slightly simpler since we do not need to project into a discrete space when proving the equi-continuity.

5.3. Proof of Theorem 5.1

We have collected all convergences that are necessary to pass to the limit with all involved terms. In particular, we may pass to the limit to obtain (D2) and (D3) and (by the methods established for the cylindrical domains, see 21, Chapter 3) with (D4). In order to pass to the limit with the weak momentum equation, fix a pair of smooth test-functions for the limit geometry $(b, \phi) \in C^\infty(\omega) \times C^\infty(T \times \mathbb{R}^3)$ with $\text{tr}_{\gamma} \phi = b\nu$. Now since $\nabla \Psi_{\eta_\epsilon}, \nabla \Psi_{\eta_\epsilon}^{-1}$ are strongly convergent in $L^\infty(I; L^2(\Omega_{\eta_\epsilon})) \cap L^2(I; L^\infty(\Omega_{\eta_\epsilon}))$ for all $q < \infty$ and $\partial_t \Psi_{\eta_\epsilon}, \partial_t \Psi_{\eta_\epsilon}^{-1}$ are strongly convergent in $L^2(I; L^a(\Omega_{\eta_\epsilon}))$ for all $a < 4$ for which (5.44) is an admissible test function for the approximate weak momentum equation with respective convergence properties. Hence we may pass to the limit with the approximate momentum equation and obtain (D1).

5.4. Proof of Theorem 2.16

In this section we are ready to prove the main result of this paper by passing to the limit $\delta \to 0$ in the system (D1)–(D4) from Section 5.3. Large parts of the proof are very similar to their counterparts in the limit $\varepsilon \to 0$. In particular, the compactness arguments from 5.2 and 4.3 have been written in such a way that they are directly adaptable for the final layer here (using only the more restrictive bounds on $\gamma$). The main exception is the analysis related to the limit passage in the molecular pressure. This can, however, be adapted from 6, Section 7. As there, we can localise the argument for fixed boundaries from 21. Consequently, parts of the argument are independent from the variable domain and the fluid-structure interaction. Nevertheless we sketch the main steps of the proof for the convenience of the reader.

Given initial data $(q_0, v_0, \theta_0)$ and $H$ belonging to the function spaces stated in Theorem 2.16 it is standard to find regularized versions $(q_{0,\delta}, v_{0,\delta}, \theta_{0,\delta})$ and $H_\delta$ such that for all $\delta > 0$

$\delta_{0,\delta}, \theta_{0,\delta} \in C^{2,\alpha}(\Omega_{\eta_\delta}), \delta_{0,\delta}, \theta_{0,\delta}$ strictly positive, $H_\delta \in C^{1,\alpha}(I \times \mathbb{R}^3)$, $H_\delta \geq 0$,

as well as

$$\int_{\Omega_{\eta_\delta}} \frac{1}{2} |\partial_t \eta_{\delta}|^2_{L^2(\omega)} + \frac{1}{2} |\partial_t \theta_{0,\delta}|^2_{L^2(\omega)} \, dx \to \int_{\Omega_{\eta_\delta}} \left( \frac{1}{2} |\partial_t q_0|^2 + |\partial_t \nu_\delta| + \partial_t \theta_{0,\delta} + \partial_t \theta_{0,\delta} \right) \, dx,$$

$H_\delta \to H$ in $L^\infty(I \times \mathbb{R}^3)$,

as $\delta \to 0$. For a given $\delta$ we gain a weak solution $(\eta_\delta, \vec{u}_\delta, g_\delta, \vartheta_\delta)$ to (5.51)–(5.52) with this data by Theorem 5.3. It is defined in the interval $(0, T_\ast)$, where $T_\ast$ is restricted by the data only. The counterpart of the total dissipation balance from (5.55), that can be derived exactly as in Section 5.3, provides the following uniform bounds:

$$\sup_{t \in I} \|\partial_t \eta_\delta\|^2_{L^2(\omega)} + \sup_{t \in I} \|\theta_0\|^2_{W^{2,2}(\omega)} \leq c,$$

$$\sup_{t \in I} \|\vec{u}_\delta\|^2_{L^2(I \times \Omega_{\eta_\delta})} + \sup_{t \in I} \|\nabla \theta_\delta\|^2_{L^2(I \times \Omega_{\eta_\delta})} \leq c,$$

$$\sup_{t \in I} \|\nabla \vartheta_\delta\|^2_{L^2(I \times \Omega_{\eta_\delta})} + \|\nabla \theta_\delta\|^2_{L^2(I \times \Omega_{\eta_\delta})} \leq c,$$

$$\sup_{t \in I} \|\nabla \vartheta_\delta\|^2_{L^2(I \times \Omega_{\eta_\delta})} + \|\nabla \theta_\delta\|^2_{L^2(I \times \Omega_{\eta_\delta})} \leq c,$$

Finally, we report the conservation of mass principle

$$\|\vartheta_\delta(\tau, \cdot)\|_{L^1(\Omega_{\eta_\delta})} = \int_{\Omega_{\eta_\delta}} \vartheta(\tau, \cdot) \, dx = \int_{\Omega_{\eta_\delta}} \vartheta_0 \, dx \quad \text{for all } \tau \in [0, T].$$

Hence we may take a subsequence, such that for some $\alpha \in (0, 1)$ we have

$$\eta_\delta \rightharpoonup^* \eta \quad \text{in } L^\infty(I; W^{2,2}(\omega))$$
By Lemma 2.11, arguing as in Sections 4.2 and 5.1, we find for all (5.63)
\[ \gamma > 0 \]
we need the assumption with constant independent of (5.62)
\[ \delta > 0 \]
\[ (5.69) \]
\[ (5.72) \]
\[ (5.73) \]
Similarly to Corollary 5.5 we have the following.
\[ (5.72) \]
\[ (5.73) \]
for any \( r < 2 \) with some limit objects \( \overline{S} \) and \( \overline{T} \). As before in Proposition 5.3 we have higher integrability of the density (see [5] Lemma 7.3 for the proof).

**Lemma 5.6.** Let \( \gamma > \frac{2}{7} \) (\( \gamma > 1 \) in two dimensions). Let \( Q = J \times B \subset I \times \Omega \eta \) be a parabolic cube and \( 0 < \Theta \leq \frac{2}{7} \gamma - 1 \). The following holds for any \( \delta \leq \delta_0(Q) \)
\[ (5.69) \]
\[ \int_Q p_\delta(g_\delta, \partial g_\delta) dx \, dt \leq C(Q) \]
with constant independent of \( \delta \).

Similarly to [5] Lemma 7.4] we can exclude concentrations of the pressure at the moving boundary. Here, we need the assumption \( \gamma > \frac{2}{7} \).

**Lemma 5.7.** Let \( \gamma > \frac{12}{7} \) (\( \gamma > 1 \) in two dimensions). Let \( \kappa > 0 \) be arbitrary. There is a measurable set \( A_\kappa \subset I \times \Omega \eta \) such that for all \( \delta \leq \delta_0 \)
\[ (5.70) \]
\[ \int_{I \times \mathbb{R}^3 \setminus A_\kappa} p_\delta(g_\delta, \partial g_\delta) \chi_{\Omega_\delta} dx \, dt \leq \kappa. \]

Lemma 5.6 and Lemma 5.7 imply equi-integrability of the sequence \( p_\delta(g_\delta, \partial g_\delta) \chi_{\Omega_\delta} \). This yields the existence of a function \( \overline{p} \) such that (for a subsequence)
\[ (5.71) \]
\[ (5.72) \]
Similarly to Corollary 5.5 we have the following.

**Corollary 5.8.** Let \( \kappa > 0 \) be arbitrary. There is a measurable set \( A_\kappa \subset I \times \Omega \eta \) such that
\[ (5.73) \]
\[ \int_{I \times \mathbb{R}^3 \setminus A_\kappa} p dx \, dt \leq \kappa. \]
Using (5.71) and the convergences (5.75)–(5.78) we can pass to the limit in (5.81) and (5.82) and obtain
\[
\int_I \frac{d}{dt} \int_{\Omega_\gamma} \varrho \mathbf{u} \cdot \phi \, dx - \int_{\Omega_\gamma} \left( \varrho \mathbf{u} \cdot \partial_t \phi + \varrho u \otimes u : \nabla \phi \right) \, dx \, dt \\
+ \int_I \int_{\Omega_\gamma} \mathbf{S} : \nabla \phi \, dx \, dt - \int_{\Omega_\gamma} \int_{\Omega_\gamma} \mathbf{\Pi} \, div \phi \, dx \, dt \\
+ \int_I \left( \int_{\Omega} \varrho \partial_t \psi y \, dy - \int_{\Omega} \varrho \psi \, d\Omega + \int_{\Omega} x \mathbf{R} \cdot \phi \, dx \, dt \right) \\
= \int_I \int_{\Omega_\gamma} \varrho f \cdot \phi \, dx \, dt + \int_I \int_{\Omega} g \phi \, dx \, dt 
\]
for all test-functions \((b, \phi)\) with \(tr_\gamma \phi = \partial_1 \eta \nu, \phi(T, \cdot) = 0\) and \(b(T, \cdot) = 0\). Moreover, the following holds
\[
\int_I \int_{\Omega_\gamma} \varrho \partial_t \psi \, dx \, dt \rightarrow \int_I \int_{\Omega} \varrho \psi \, d\Omega \quad \text{for all} \quad \psi \in C^\infty(I \times \mathbb{R}^3). 
\]
Consequently we have \(S = S(\psi, \nabla \mathbf{u})\) in (5.75). Moreover, we can pass to the limit in the entropy balance and obtain (O3). Next we aim to prove strong convergence of the density. We define the \(L^\infty\)-truncation
\[
T_k(z) := \frac{k \psi}{T} \quad z \in \mathbb{R}, \quad k \in \mathbb{N}.
\]
Then we have to show that
\[
\left| T_k(\varrho) - T_k(\varrho) \right| \rightarrow 0 \quad \text{in} \quad L^1(I; L^4(\Omega_\gamma)).
\]
Here \(T\) is a smooth concave function on \(\mathbb{R}\) such that \(T(z) = z\) for \(z \leq 1\) and \(T(z) = 2\) for \(z \geq 3\). Now we have to show that
\[
\int_{I \times \Omega_\gamma} \left( a \varrho \bar{\gamma} + \delta \varrho \bar{\gamma} - (\lambda(\varrho) + 2\mu(\varrho)) \div \mathbf{u} \right) T_k(\varrho) \, dx \, dt \\
\rightarrow \int_{I \times \Omega} \left( \mathbf{\Pi} - (\lambda(\varrho) + 2\mu(\varrho)) \div \mathbf{u} \right) T_k^1(\varrho) \, dx \, dt.
\]
For this step we are able to use the theory established in [32] on a local level. Similarly to [3] Subsection 7.1 (see [21] Chapter 3, Section 3.7.4 about how to include the temperature) we first prove a localised version of (5.75) and then use Lemma (5.7) and Corollary (5.8) to deduce the global version. The next aim is to prove that \(\varrho\) is a renormalized solution (in the sense of Definition 2.15). In order to do so it suffices to use the continuity equation and (5.78) again on the whole space. Following line by line the arguments from [3] Subsection 7.2 we have
\[
\partial_t T^{1,k} + \div (T^{1,k} \mathbf{u}) + T^{2,k} = 0 
\]
in the sense of distributions on \(I \times \mathbb{R}^3\). Note that we extended \(\varrho\) by zero to \(\mathbb{R}^3\). The next step is to show
\[
\limsup_{\delta \rightarrow 0} \int_{I \times \mathbb{R}^3} |T_k(\varrho) - T_k(\varrho)|^q \, dx \, dt \leq C,
\]
where \(C\) does not depend on \(k\) and \(q > 2\) will be specified later. The proof of (5.80) follows exactly the arguments from the setting with fixed boundary (see [21] Chapter 3, Section 3.7.5) using (5.78) and the uniform bounds on \(\mathbf{u}_\delta\) (with the only exception that we do not localise). Using (5.80) and arguing as in [3] Sec. 7.2] we obtain the renormalised continuity equation. As in [3] Sec. 7.3] we can use the latter one to show strong convergence of the density. Now we can pass to the limit with the approximate equations and obtain the weak solution, as it was explained in the previous subsection.

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