Superconductivity from Repulsion: A Variational View

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Abstract

This work present a new class of variational wave functions for fermi systems in any dimension. These wave functions introduce correlations between Cooper pairs in different momentum states and the relevant correlations can be computed analytically. At half filling we have a ground state with critical superconducting correlations, that causes negligible increase of the kinetic energy. We find large enhancements in a Cooper pair correlation function caused purely by the interplay between the uncertainty principle, repulsion and the proximity of half filling. This is surprising since there is no accompanying signature in usual charge and spin response functions, and typifies a novel kind of many body cooperative behaviour.
Introduction: Variational wave functions have played an important part in our understanding of several phenomena in quantum manybody systems. The role of the BCS wave function in superconductivity, Jastrow-type functions for superfluidity, the Laughlin function in the Fractional Quantum Hall problem, and the Gutzwiller wavefunction in strongly correlated fermi systems have been recognized to be of fundamental importance.

This work presents a new variational wave function and some variants of it, that promise to be of interest in the topical problem of superconductivity arising from repulsion. We have recently argued on the basis of certain inequalities, that projecting out s-wave Cooper pairs in a fermi system on a lattice would lead to enhanced extended s-wave type pairing fluctuations near half filling. These are expected to lead to superconductivity at precisely half filling for a class of Hubbard-type models. In brief the argument involves the recognition that s-wave and extended s-wave pairing are canonically conjugate in the sense of the uncertainty principle, which ultimately drives the instability towards superconductivity near half filling. This conjugacy arises since the s wave and the extended s wave Cooper order parameters $B, A$ (defined below) satisfy the commutation relation $[B, A^\dagger] = 2T$, where the kinetic energy operator $T$ plays a benign role, similar to that of the number operator $\hat{N}$ as in other familiar contexts, such as the conductivity sum rule on the lattice. From the uncertainty principle, it follows that the fluctuation $< A^\dagger A + AA^\dagger >$ is bounded from below by $4 < T >^2 / < B^\dagger B + BB^\dagger >$. The special role played by half filling is due to the fact that the suppression of s-wave fluctuations is possible to a very high degree near half filling, the commutation relation $[B, B^\dagger] = L - \hat{N}$ in a sector $< \hat{N} > \sim L$, permits both $< B^\dagger B >$ and $< BB^\dagger >$ to be simultaneously small. This “squeezing” results in large enhancements in the conjugate variables, remarkably enough, without an appreciable energy cost. This effect has been termed order by projection.

While the arguments in provide a novel direction, the absence of explicit solutions or of good wave functions has proved to be a hindrance in arriving at a thorough understanding of the physics of these models. We provide in this work a wave function which is tractable enough so that calculations of all relevant expectation values can be done analytically, which by itself is rare in many body systems. Further the wave function catches the essence of the enhancements mentioned above, and at half filling is argued to be exact. While we know of no materials to which these models apply, we make bold to present these models and wave functions, since they address and provide what we believe to be amongst the first positive results for the important theoretical question of the possibility of superconductivity without explicit attractive interactions in fact arising out of pure repulsion.

The model and the variational wavefunction: Let us write the Hamiltonian of the model considered in $d$ dimensions as

$$H = T + U_s B^\dagger B .$$  \hspace{1cm} (1)$$

Here $B = \sum b_k$ is the s-wave Cooper pair operator, and $b_k = c_{-k\downarrow} c_{k\uparrow}$ the pair operators, the kinetic energy $T = \sum \epsilon_k n_{k\sigma}$, with the band dispersion $\epsilon_k = -2 \sum_{\alpha=1}^{d} \cos(k_{\alpha})$, although later we will consider a slightly more general form of the band dispersion in 2-dimensions with orthorhombic distortion. The $U_s$ term discourages on site s-wave correlations, and in
fact projects out s-wave order. In order to get a feeling for its effect, consider the estimate of the energy in a BCS state \(|\Psi_{\text{BCS}} >= \prod (u_k + v_k b_k^\dagger) \mid \text{vac} >\). Provided \(\sum u_k v_k = 0\), we find \(E_{\text{BCS}} = \sum [2(\epsilon_k - \mu) v_k^2 + U_s v_k^4]\), subject to \(N = 2 \sum v_k^2\). The role of \(U_s\) is clearly to flatten out \(v_k\) from its step function behaviour in the normal state, and hence a superconducting state with true Long Ranged Order (LRO) and an extensive energy shift arises, at least as a local variational minimum. In Ref. [3], a more general model is introduced including the above term as well as the Hubbard \(U\) term. The general results of [3] imply for these models that away from half filling, the ground state energy density [3] is unshifted by the \(U_s\) term, while the extended s-wave correlations get a large enhancement at the expense of suppressing s-wave correlations. By continuity in density, we expect superconductivity at half filling. In this work we ignore the Hubbard \(U\) term for tractability, and explore the other aspects within the model of Eq. [1]; in particular the origin of the enhancements.

The extended s-wave operator \(A = -2 \sum \epsilon_k b_k\) satisfies the commutation relations \([B, T] = -A\) and \([B, A^\dagger] = 2T\). The main variational wave function proposed here is written in terms of the free fermi wave function \(|\Phi >= \prod_{\theta} b_k^\dagger \mid \text{vac} >\)
as

\[
|\Psi_\theta >= \exp(-\frac{\theta}{\mathcal{L}} S) |\Phi >, 
\]

where the number of lattice sites is \(\mathcal{L}\). The prefactor generates several Cooper particle pairs and Cooper hole pairs in the fermi gas. To motivate this wavefunction, note that the anti Hermitian operator \(S = \sum (\epsilon_{k_1} - \epsilon_{k_2}) b_{k_1}^\dagger b_{k_2}\) can be viewed as the commutator \(\frac{1}{2} [T, B^\dagger B]\), and hence the wave function may be viewed as a “rotation” about a direction orthogonal to the kinetic and potential energies. Such a strategy is familiar e.g. from diagonalizing a sum over two Pauli matrices, as well as quadratic forms in bosons. A similar rotation also gives the exact answer in a further simplified model [3].

**Calculation of expectation values:** We now turn to the calculation of expectation values \(< Q >_\theta = <\Psi_\theta | Q | \Psi_\theta >\) of various operators \(Q\) in the above state. We begin by evaluating the derivative of the kinetic energy \(< T >_\theta\),

\[
\frac{d}{d\theta} < T >_\theta = \frac{1}{\mathcal{L}} [ < A^\dagger A >_\theta - 2 < C^\dagger B + B^\dagger C >_\theta ] \equiv 4 \alpha(\theta) - 2 \nu(\theta), 
\]

where \(C = \sum \epsilon_k^2 b_k\), thus defining the functions \(\alpha, \nu\). We immediately note a feature of considerable importance, namely that unless one of the correlation functions in the right hand side of Eq. [3] possesses explicit Long Ranged Order (LRO), the change in the (extensive) kinetic energy is *not extensive*, and may be neglected in the thermodynamic limit. We will make repeated use of this insight later. We also note that this includes the case where one of the correlators possesses Quasi LRO, i.e. \(\zeta < 1\), indeed we find later that \(\zeta = \frac{1}{2}\) does occur for the case of greatest interest, namely the half filled limit. We also need the expectation of operators \(T_n = \sum \epsilon_k^2 \phi_k\), with \(\phi_k = n_k^\dagger + n_{-k^-} - 1\), so that \(T_1 = T\). The equations are most compactly written in terms of the operators \(I_n = \sum \epsilon_k^2 b_k\) such that \(I_0 = B, I_1 = -\frac{1}{2} A, I_2 = C\), and their correlators \(\Gamma_{n,m}(\theta) \equiv \frac{1}{\mathcal{L}} < I_n^\dagger I_m + I_m^\dagger I_n >_\theta\) as

\[
\frac{d}{d\theta} < T_n >_\theta = 2[\Gamma_{n,1}(\theta) - \Gamma_{n,0}(\theta)]. 
\]
We note that in view of the above definition, LRO corresponds to \( \Gamma \sim O(\mathcal{L}) \), whereas in the absence of true LRO, \( \Gamma \sim o(\mathcal{L}) \), and hence here the correlators \( < T_n >_\theta \) are insensitive to the rotation, at least to leading order in \( \mathcal{L} \).

The correlators \( \Gamma_{n,m} \) satisfy exact equations

\[
d\frac{d}{d \theta} \Gamma_{n,m}(\theta) = \frac{1}{\mathcal{L}^2} < [T_{n+1} I_0^I I_m + I_n^I I_0^I T_{m+1} - I_n^I I_1^I T_m - T_n I_1^I I_m + h.c.] >_\theta, \tag{5}
\]

with initial conditions obtained by taking the expectation value in the free fermi ground state: they read \( \Gamma_{n,m}(0) = 2[\epsilon_k^{n+m} f_k]_k \), with \( f_k \) the usual (non-interacting) fermi function and \( [g_k]_k \equiv \frac{1}{\mathcal{L}} \sum_k g_k \).

The equations are handled assuming a fundamental factorization that arises in the thermodynamic limit. This factorization of hermitean global operators (i.e. sums over all sites of local operators) and may be expressed as the statement that for any two such \( Q_j \) with nonzero averages, we have \( < Q_1 Q_2 >_\theta / [< Q_1 >_\theta < Q_2 >_\theta] \sim o(\mathcal{L}) \). This is also equivalent to the statement that the connected part of the correlator is subleading in powers of \( \mathcal{L} \). The case needed by us corresponds to \( Q_1 \sim \mathcal{L} \Gamma_{n,m} \) and \( Q_2 \sim T_n \). If this is assumed then we immediately see that Equations(3) factor out into linear equations

\[
d\frac{d}{d \theta} \Gamma_{n,m} = \mu_{n+1} \Gamma_{0,m} + \mu_{m+1} \Gamma_{0,n} - \mu_m \Gamma_{1,n} - \mu_n \Gamma_{1,m} \tag{6}
\]

with coefficients \( \mu_n \equiv [\epsilon_k^n (2 f_k - 1)]_k \) obtainable numerically, together with the initial conditions \( \Gamma_{n,m}(0) = 2[\epsilon_k^{n+m} f_k]_k \).

We have checked this assumption to low orders by evaluating the correlation functions \( \Gamma_{n,m} \) for the most relevant cases \((n,m) = (0,0), (1,1) \) out to \( O(\theta^5) \). The calculation proceeds by expanding the expectation value as a nested commutator with \( S \), which is evaluated by Wick's theorem for ground state correlations, leading to \( n! \) contractions that are further multiplied by powers of \( \epsilon \). This results in sums over \( k \) of algebraic functions of the band energies and fermi functions. The series agrees with the one generated from the above factorization procedure exactly to that order, and provides non trivial support for it.

We now consider the correlation functions of interest, and begin by denoting \( \beta(\theta) = \frac{1}{2} \Gamma_{0,0}, \alpha(\theta) = \frac{1}{2} \Gamma_{1,1}, \gamma(\theta) = \Gamma_{1,0} \) and finally \( \nu(\theta) = \Gamma_{2,0} \). The equations of immediate concern are then

\[
\begin{align*}
\frac{d}{d \theta} \alpha & = -2 \mu_1 \alpha + \mu_2 \gamma \\
\frac{d}{d \theta} \beta & = 2 \mu_1 \beta + (1 - \rho) \gamma \\
\frac{d}{d \theta} \gamma & = 2 (1 - \rho) \alpha + 2 \mu_2 \beta. \tag{7}
\end{align*}
\]

We also note for later use that \( \frac{d}{d \theta} \nu = \mu_1 \nu + 2 \mu_3 \beta + (1 - \rho) \Gamma_{1,2} - \mu_2 \gamma \). We denote the particle density as \( \rho \) such that \( \rho \leq 1 \). The initial conditions are \( \beta(0) = \frac{1}{2} \rho, \alpha(0) = \nu(0) = d + \frac{1}{2} \mu_2, \gamma(0) = \mu_1 \). We now examine the behaviour of these coupled equations. Let us note that \( \mu_0 = (\rho - 1) \), further \( \mu_1 \) is negative and non vanishing near half filling, whereas
\( \mu_2 \) also negative, is very small \( \sim O((1 - \rho)^3) \) near half filling. In fact for all even \( n \) we must have \( \mu_n \to 0 \) as we approach half filling, due to particle hole symmetry. If we neglect the cross couplings, then \( \beta \) falls off as \( \theta \) increases from 0 while \( \alpha \) increases.

A neat way to solve these equations is suggested by inspection: we observe that in effect the Eqns(7) are generated by a linear relation \( \frac{d}{dt}I_n = \mu_{n+1}I_0 - \mu_nI_1 \) and the hermitean conjugates of these. The \( \theta \) dependence of all the operators are related thus to the fundamental ones \( I_0, I_1 \), and thus we find linear combinations that have simple evolution: \( Q_{\pm} = -\mu_0I_1 + (\mu_1 \pm \alpha_0)I_0 \) where \( \alpha_0 = \sqrt{\mu_1^2 - \mu_0\mu_2} \). These obey uncoupled equations \( \frac{d}{dt}Q_{\pm} = \pm \alpha_0Q_{\pm} \).

We readily invert and find with \( \xi_\pm = (1 \pm \frac{\mu_0}{\lambda_0}) \), \( I_0 = \frac{1}{2\lambda_0}\{Q_+ - Q_- \} \) and \( I_1 = \frac{1}{2(1-\rho)}\{\xi_-Q_+ + \xi_+Q_- \} \). The four basic correlators \( F_{\sigma_1,\sigma_2}(\theta) = <Q_{\sigma_1}\bar{Q}_{\sigma_2}>_\theta \) can thus be found simply:

\[
F_{\sigma_1,\sigma_2}(\theta) = F_{\sigma_1,\sigma_2}(0)\exp(\lambda_0\theta(\sigma_1 + \sigma_2)).
\]

The initial conditions are given as

\[
F_{\sigma_1,\sigma_2}(0) = (1 - \rho)^2\alpha(0) + (1 - \rho)\{\mu_1 + \frac{\lambda_0}{2}(\sigma_1 + \sigma_2)\}\gamma(0) + (\mu_1 + \sigma_1\lambda_0)(\mu_1 + \sigma_2\lambda_0)\beta(0).
\]

We thus find the expectation values:

\[
\alpha(\theta) = \frac{1}{4(1 - \rho)^2}[\xi_-^2F_{1,1}(0)\exp(2\lambda_0\theta) + \xi_+^2F_{-1,-1}(0)\exp(-2\lambda_0\theta) + 2\xi_+\xi_-F_{1,-1}(0)]
\]

\[
\beta(\theta) = \frac{1}{4\lambda_0^2}[F_{1,1}(0)\exp(2\lambda_0\theta) + F_{-1,-1}(0)\exp(-2\lambda_0\theta) - 2F_{1,-1}(0)]
\]

\[
\gamma(\theta) = \frac{1}{2\lambda_0(1 - \rho)}[\xi_-F_{1,1}(0)\exp(2\lambda_0\theta) - \xi_+F_{-1,-1}(0)\exp(-2\lambda_0\theta) + 2\frac{(\mu_1)}{\lambda_0}F_{1,-1}(0)]
\]

These together with Eq(3) provide the solution to the variational problem since the variational energy density[4] is \( \mu_1 + U_\theta \beta(\theta) \), as well as that of computing the correlations.

We show in Fig(1) the behaviour of the three functions \( \alpha, \beta, \gamma \) for the case of 2-d at a typical value of the density \( \rho = .75 \). Note that \( \beta \) goes through a very shallow minimum with a value \( \beta^*(\rho) \equiv \beta(\theta^*) \sim 0 \), and \( \gamma \) also seems to go through zero at nearly the same value of the minimizing \( \theta = \theta^* \). From the Equations(7), \( \beta^* \) would be actually zero if \( \gamma \) vanishes at exactly \( \theta^* \). To investigate this further, we plot in Fig(2) the minimum value \( \beta^*(\rho) \) as a function of density in 2-d, and the inset shows a similar plot for 1-d. We see that for general densities, although \( \beta^* \) is impressively small, it is non zero. From the general arguments of [3] we know that the exact ground state energy density is unshifted from the non interacting value, corresponding to \( \beta_{exact} = 0 \), and so our wave function gives a very good approximation to the true energy for most densities, provided we restrict to \( U_\theta \sim |\mu_1| \). In Fig(3) we show \( \alpha^*(\rho)/\alpha(0) \), this is enhanced greatly, by an order of magnitude as far as 15% away from half filling. The lower bound \( \alpha_{lb} = \frac{1}{2}(\frac{\mu_2}{\lambda_0} + \mu_2) \) obtained in Ref[4], also has a similar behaviour close to half filling.

**Solution at Half filling:** The discussion of the solution above indicates that for half filling, i.e. \( \rho = 1 \), the wave function discussed here could be exact, since \( \beta^* \) goes to zero here. Our analysis of the orders of magnitude of the correlators was based so far on the implicit assumption that \( \theta \sim O(1) \). From the solution it can be shown that for \( \rho < 1 \), the minimizing \( \theta^* \sim -\log(1 - \rho) \), and hence at \( \rho = 1 \), we must work with greater
precision to see if the assumptions made earlier are consistent. Firstly we assume that Eqns(1) are still valid at half filling, requiring that \( \lim_{L \to \infty} < T_n > / L \) is unchanged from its non-interacting value \( \mu_n \). We verify at the end that this is still true and hence obtain a self consistent argument. With the assumption we set \( \mu_{2n} \to 0 \) and hence the equations decouple and are integrated immediately giving \( \alpha(\theta) = \alpha(0) \exp(-2\mu_1\theta) \) and \( \beta(\theta) = \beta(0) \exp(2\mu_1\theta) \). We also integrate the equation for \( \nu \) and find \( \nu(\theta) = g_1 \exp(\mu_1\theta) + g_2 \exp(2\mu_1\theta) \). The constants \( g_1, g_2 \) involve the initial conditions, but are not important since we see that for large (positive) \( \theta \) we can safely neglect \( \nu(\theta) \) (recall that \( \mu_1 < 0 \)). We substitute for \( \alpha \) into Eqn(3) and integrate over \( \theta \), to find the energy per site

\[
e(\theta) = e_{\text{non}} + \frac{U_s}{2} \exp(2\mu_1\theta) + \frac{2\alpha(0)}{|\mu_1|} \exp(-2\mu_1\theta) + \text{(negligible terms)} \tag{10}
\]

where \( e_{\text{non}} = \mu_1 \) is the ground state energy density of the free fermi gas, and the negligible terms are of the type \( \frac{1}{2} \{1, \exp(\mu_1\theta), \exp(2\mu_1\theta)\} \) etc. We see that minimizing Eq(10) w.r.t. \( \theta \) pushes up the energy as

\[
e(\theta^*) = e_{\text{non}} + 2 \frac{\sqrt{U_s\alpha(0)}}{\sqrt{|\mu_1|L}} + \text{(negligible terms)} \tag{11}
\]

where \( \theta^* = 1/(4|\mu_1|) \log(U_s|\mu_1|L/4\alpha(0)) \). The energy density continues to be independent of \( U_s \). Thus the growth of \( \theta \) stops on the scale of \( \log(1/L) \), and we find \( \alpha^* = \frac{1}{2} \sqrt{\alpha(0)U_sL|\mu_1|} \). Since \( \alpha^* \) does not have true LRO, our original assumption is validated, and hence we have a consistent solution. The solution at half filling thus has Quasi LRO since \( < A^\dagger A > \sim L^{3/2} \), and is superconducting in the sense of having power law order, e.g. in a 2-d superconductor at finite temperatures.

The d-wave superconductor: We now ask the question whether this class of wave functions works for d-wave superconducting correlations as well in 2-d. This case is clearly of topical interest in the context of the 2-d High \( T_c \) cuprates. There is one possibility inherent in the discussion of Ref(4), wherein we take \( \hat{B} = \sum \eta_k b_k \), with \( \eta_k = \text{sign}(w_k) \), and \( w_k = \cos(k_x) - \cos(k_y) \). This choice of the relative phases of the \( k \) state Cooper pairs preserves the important commutation relation \( [\hat{B}, \hat{B}^\dagger] = \hat{L} - \hat{N} \) (where \( \hat{N} \) is the number operator), since \( \eta_k^2 = 1 \), and all the arguments of Ref(4) can be taken over, exactly as for the case of extended s-wave pairing discussed above. The conjugate variable that is enhanced now is \( \hat{A} = -2 \sum \epsilon_k \eta_k b_k \), and the two particle wave function has the signs that are usually associated with d-wave pairing. The “gap” within a weak coupling approach in this case may be taken to be \( \Delta = \Delta_0 \epsilon_k \eta_k \), and thus has jumps in \( k \) space where one expects nodes. Of course this may not be a fatal difficulty since there is no clear connection between the ground state obtained within the present framework, and the low lying excitation spectrum within a BCS framework.

Within the variational framework, we can admit d-wave pairing provided the lattice has a nonzero orthorhombic distortion, i.e. the band dispersion is asymmetric in \( x \) and \( y \) directions. We can model this by setting \( \epsilon_k = -2(1-\epsilon_0) \cos(k_x) - 2(1+\epsilon_0) \cos(k_y) \) with \( \epsilon_0 \) a measure of the orthorhombicity. We choose a variational wave function exactly as in Eq(2), with \( S_{DW} = \sum (w_{k_1} - w_{k_2}) b_{k_1}^\dagger b_{k_2} \) and define the d-wave pairing operator \( D = \sum w_k b_k \).
With this we can essentially borrow all the results of the previous calculation provided we replace \( \mu_n \rightarrow \nu_n \), where \( \nu_n = \frac{1}{L} < \hat{T}_n >_0 \) and \( \hat{T}_n = \sum w_k^n \phi_k \). The fermi groundstate has the symmetry of the \( \epsilon_k \) and hence we see the need for the distortion, for in its absence, \( \nu_n = 0 \) for all \( n \geq 1 \). We must also replace the initial conditions appropriately, thus \( \Delta(\theta) = \frac{1}{L} < D^\dagger D >_\theta \), and \( \Delta(0) = [w^2 f_k]_k \). We need to examine the reality of \( \hat{\lambda}_0 = \sqrt{\nu^2 + (1 - \rho)\nu_2} \). In case \( \hat{\lambda}_0 \) is imaginary, the nature of solution changes drastically, and we have a limited enhancement, since the functions now oscillate as trigonometric functions, as \( \theta \) varies. If \( \hat{\lambda}_0 \) is real, we can repeat everything said before and qualitatively find the same answers as for the extended s-wave case. The numerics require a knowledge of the elements \( \nu_n \) and we will not discuss them further here. At half filling the state is similarly obtained, and the energy is different from that of the s-wave case only in terms of \( O(1/\sqrt{L}) \), it is higher in general since \( \nu_n \) are usually smaller than the \( \mu_n \) for \( n \geq 1 \). It must be remarked that in this case the correlation function \( \alpha \) is also enhanced similarly to \( \Delta \), since these correlations get coupled. The correct symmetry of the order parameter is then of a mixed s and d type, rather than pure s or d\[9\], and indeed one should, in principle, optimize the form of the S operator by taking a suitable mixture of s and d functions to get the best possible energy.

**Summary and Discussion:** In summary, we have presented variational wave functions that seem interesting on several counts. In many regards, we are tempted to say that the wave functions are at least as interesting as the models if not more. Firstly, the expectation values are computable analytically and throw light on an interesting and novel relationship, namely the enhancement in certain correlations at the expense of on site s-wave correlations. At half filling we have a state that is degenerate with the free fermi gas in the energy-density, and yet has critical superconducting correlations. Is this state the exact ground state of the model in Eq(1)? The coincidence of the energy density with the obvious lower bound \( e_{\text{non}} \), is not a clinching argument by itself, particularly in view of the main message of our work. The nature of the correlations, on the other hand, being of the form expected from the uncertainty principle arguments, provide a significant positive indicator. We feel that certain results found here are likely to be true, such as \( \alpha^* \propto L^{1/2} \), and the form of energy corrections as in Eq(11), without however, being able to prove them in a rigorous fashion. Of the values of the coefficients of the leading behaviour, it is less easy to be certain.

The critical behaviour of the Cooper correlation function \( < A^\dagger A > \) is of particular interest. It behaves as \( L^{3/2} \) at \( \rho = 1 \), and is \( \alpha \propto \frac{L}{1 - \rho} \) otherwise, suggesting that the local density \( \alpha(r) \) of the A operator has correlations \( < \alpha^\dagger(r)\alpha(0) > \sim \frac{1}{r^{d+2}} h(r/\xi(\rho)) \), with a density dependent correlation length \( \xi(\rho) \sim \frac{1}{(1 - \rho)^{d+2}} \), that diverges at half filling. Such a crossover is similar to that in the behaviour of the spin correlations in the Hubbard model in 1-d near half filling\[10\]. Note that the behaviour of this function for the free fermi gas is \( \sim \frac{1}{r^{d+2}} \), and hence the correlations in this case are considerably shorter ranged than ours.

We conclude by mentioning another nontrivial implication of the ideas discussed above, but going beyond the context of the models considered here. We have seen that the fermi gas on a lattice has the possibility of novel large enhancements in the ground state correlation functions of Cooper pairs, produced by repulsive terms of certain kinds. These are not reflected in either the total energy (and hence the compressibility) or the momen-
tum distribution function. The latter can be readily seen to be as sharp as in the ideal fermi case. Similarly one can check that the spin correlations are unaffected. The usual fermi liquid enhancements in spin and charge response that herald incipient instabilities are totally missing in these systems. As a result, any bosonic excitations available, including phonons, would find the electronic system much more prone to superconductivity of the usual sort than expected on grounds of the missing fermi liquid enhancements, and presumably be accompanied by enhanced $T_c$’s. Estimates of such effects must await information on the spectral weight in the dynamical Cooper susceptibility.

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[7] The term “energy density” is used to denote \( \lim_{L \to \infty} \frac{E}{L} \). We use the term “energy per site”, on the other hand, to denote \( \frac{E}{L} \) preparatory to taking the thermodynamic limit, and display only \( \theta \) dependent corrections such as \( \frac{1}{L^{\alpha} g(\theta)} \). We omit displaying other finite size corrections that are undoubtedly present but uninteresting, being \( \theta \) independent. We occasionally also omit displaying \( \theta \) dependent but smaller corrections than the ones displayed.

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Figure 1: The three correlations functions at $\rho = 0.75$ in 2-d vs. $\theta$.

Figure 2: The minimum value of $\beta^*$ vs density.
Figure 3: The enhancement factor of the extended s-wave correlator $\alpha^*/\alpha(0)$ vs density in 2-d, and in 1-d (inset).