Detailed measurement of the $e^+e^-$ pair continuum in $p+p$ and $\text{Au}+\text{Au}$ collisions at $\sqrt{s_{NN}}=200$ GeV and implications for direct photon production

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PHENIX has measured the $e^+e^-$ pair continuum in $\sqrt{s_{\text{NN}}}=200$ GeV Au + Au and $p+p$ collisions over a wide range of mass and transverse momenta. The $e^+e^-$ yield is compared to the expectations from hadronic sources, based on PHENIX measurements. In the intermediate-mass region, below the masses of the $\phi$ and the $J/\psi$ meson, the yield is consistent with expectations from correlated $c\bar{c}$ production, although other mechanisms are not ruled out. In the low mass region, below the $\phi$, the $p+p$ inclusive mass spectrum is well described by known contributions from light meson decays. In contrast, the Au + Au minimum bias inclusive mass spectrum in this region shows an enhancement by a factor of $4.7 \pm 0.4_{\text{stat}} \pm 1.5_{\text{syst}} \pm 0.9_{\text{model}}$. At low mass ($m_{\text{c.m.}} < 0.3$ GeV/$c^2$) and high $p_T$ ($1 < p_T < 5$ GeV/$c$) an enhanced $e^+e^-$ pair yield is observed that is consistent with production of virtual direct photons. This excess is used to infer the yield of real direct photons. In central Au + Au collisions, the excess of the direct photon yield over the $p+p$ is exponential in $p_T$, with inverse slope $T = 221 \pm 19_{\text{stat}} \pm 19_{\text{syst}}$ MeV. Hydrodynamical models with initial temperatures ranging from $T_{\text{init}} \simeq 300-600$ MeV at times of 0.6-0.15 fm/c after the collision are in qualitative agreement with the direct photon data in Au + Au. For low $p_T < 1$ GeV/$c$ the low-mass region shows a further significant enhancement that increases with centrality and has an inverse slope of $T \simeq 100$ MeV. Theoretical models under predict the low-mass, low-$p_T$ enhancement.

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I. INTRODUCTION

Experimental results from the Relativistic Heavy Ion Collider (RHIC) have established that in Au + Au collisions at $\sqrt{s_{\text{NN}}} = 200$ GeV matter is created with very high energy density [1], as indicated by the large energy produced transverse to the beam direction [2], as well as by the large energy loss of light [3, 4] and heavy quarks [5, 6], and is thermalized rapidly, as indicated by the large elliptic flow of these partons [6, 7, 8, 9]. Such a high density thermalized medium is expected to emit thermal radiation [10] in the form of direct photons and dileptons.

Electron-positron pairs, or dileptons in general, are excellent tools for studying collisions of heavy ions at ultra-relativistic energies. Since they are not affected by the strong interaction, and therefore can escape from the dense medium without final state interaction, dilepton spectra can probe the whole time evolution and dynamics of the collision. Dileptons can also be used to study the properties of low-mass vector mesons $\rho$, $\omega$, and $\phi$ in the medium, since their lifetime is shorter or similar ($\phi$) to that of the medium. Their mass and width inside the dense medium can be directly measured through their dilepton decay channels, and thereby one can study the effect of chiral symmetry restoration on these mesons. Furthermore, production of photons can be measured through their conversion to dileptons.

As schematically shown in Fig. 1, the dilepton spectra can be classified into the high-mass region (HMR; $m > 3.2$ GeV/$c^2$) from $J/\psi$ mass and above, the low-mass region (LMR; $m < 1.2$ GeV/$c^2$) from the $\phi$ mass and below, which is further subdivided in LMR I and LMR II as described below, and the intermediate mass region (IMR; $1.2 < m < 2.9$ GeV/$c^2$) between them.

In the HMR hard scattering on partons in colliding nuclei produces dileptons through the Drell-Yan process ($\bar{q}q \rightarrow l^+l^-$) and correlated semi-leptonic decays of heavy quark pairs ($b\bar{b} \rightarrow l^+l^-, c\bar{c} \rightarrow l^+l^-$). Dileptons from these hard processes are expected to dominate in the HMR since their mass spectra are harder than that from other possible sources. Thus dileptons in the HMR probe the initial stage of the collision [1]. Charmonia ($J/\psi, \psi'$) and Upsilon are in this mass region and deconfinement effects can be studied from their yields. Little contribution from thermal radiation is expected in the HMR at RHIC energies [11].

In the IMR theoretical models predict that dileptons from the thermalized deconfined phase, the quark gluon plasma (QGP), are the dominant source of dileptons [12, 13, 14]. The measurement of thermal dileptons from QGP can be used to determine the initial temperature of the matter. Here a competing source of dileptons is the semi-leptonic decay of $c$ and $\bar{c}$, correlated through flavor conservation. The continuum yield in this mass region is sensitive to the energy loss of charm quarks in the medium.

In the LMR dilepton production is expected to be dominated by in-medium decay of $\rho$ mesons in

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The discovery of a large enhancement of the dilepton yield in the LMR in ion-ion collisions by HELIOS/3 [24] and CERES [25] at the CERN SPS has triggered a broad theoretical investigation of modifications of properties of hadrons in a dense medium and of how these modifications relate to chiral symmetry restoration [21, 26, 27]. These studies advanced with the availability of more precise data from NA60 and CERES [28, 29] and HADES [30]. Most theoretical studies suggest that in-medium modifications of the $\rho$ meson, with its short lifetime and its strong coupling to the $\pi\pi$ channel, are primarily responsible for the enhancement.

In the LMR II (marked with II in Fig. 1) the CERES data show that the enhancement increases significantly faster than linearly with charged particle density and is concentrated at very low pair-$p_T$ [29]. This behavior is consistent with the interpretation that the excess is due to annihilation processes. NA60 recently confirmed the excess of dileptons in the LMR in In+In collisions at 158 A GeV with a high statistics dimuon measurement [28]. NA60 also observed that the inverse slope parameter $T_{\text{eff}}$ of the pair-$p_T$ spectra rises with dimuon mass in the LMR [31].

An enhanced yield was also observed in the IMR by HELIOS/3 [24, 25], NA38/50 [32, 33], and NA60 [28, 31]. NA60 data suggest that the enhancement cannot be attributed to decays of $D$ mesons but may result from prompt production, as expected for thermal radiation [33]. Furthermore NA60 measures the inverse slope parameter of pair-$p_T$ spectra in the LMR around 190 MeV, independent of mass and significantly lower than those found at masses below 1 GeV/$c^2$ [28, 33].

The PHENIX experiment at the Relativistic Heavy Ion Collider (RHIC) has measured the dilepton continuum in a new energy regime, $\sqrt{s_{NN}} = 200$ GeV, for $p+p$ and $Au + Au$ collisions. In this article we present results from $Au + Au$ collisions taken in 2004 and $p+p$ collisions taken in 2005. We show the results in the LMR and in the IMR as well as the result of direct photon measurement from the analysis of quasi-real virtual photons. The main physics results in the LMR and IMR in Au + Au and $p+p$ have been reported in [34] and [35], respectively, and the results of the virtual photon analysis have been reported in [36]. New results on the centrality and $p_T$ dependence of the $e^+e^-$ pairs in the LMR are presented in this paper.

This article is organized as follows. Sec. II describes the PHENIX detector system related to the analysis. Sec. III presents the analysis details including the systematic uncertainties. Sec. IV describes the methods used to calculate the pair yield expected from hadronic decays. Sec. V shows the $e^+e^-$ results as a function of $m_{ee}$ and $p_T$, which are then discussed in Sec. VI and compared to available theoretical predictions. Finally Sec. VII provides a conclusion.
II. PHENIX DETECTOR

A detailed description of the complete PHENIX detector system can be found elsewhere [39, 41, 42, 43, 44, 45]. Here we describe the parts of the detector system that are used in this analysis, namely, two global detectors and two central arm spectrometers. The global detectors are the beam-beam counters (BBC) and the zero-degree calorimeters (ZDC). Each central arm covers the pseudorapidity range $|\eta| < 0.35$ and an azimuthal angle of $\pi/2$, and includes a drift chamber (DC) and multi-wire proportional pad chambers (PC) for charged particle tracking, the ring-imaging Čerenkov counter (RICH) for electron identification, and an electromagnetic calorimeter (EMCal) for energy measurement and further electron identification. Figure 2 shows a beam view of the PHENIX detector.

![Diagram of PHENIX Detector](image)

(a) Beam View of PHENIX Detector

FIG. 2: (Color online) Beam view (at $z = 0$) of the PHENIX central arm detector in Run-4 Au + Au and Run-5 $p + p$. The detectors used in the present analysis are the drift chamber (DC) and the multi-wire proportional pad chamber (PC) for charged particle tracking, the ring-imaging Čerenkov counter (RICH) for electron identification, and the electromagnetic calorimeter (lead-scintillator (PbSc) and lead-glass (PbGl)) for energy measurement.

A. Global Detectors

The BBC and the ZDC measure the start time and the collision vertex position $z$-vertex along the beam axis and are used to determine the centrality of the collision [45]. They also provide first level trigger information.

The BBC consists of two sets of 64 Čerenkov counter modules, located $\pm 1.44$ m from the nominal interaction point along the beam line and measure the number of charged particles in the pseudorapidity region $3.1 < |\eta| < 3.9$. They provide the start-time of the collision with a resolution of $\sim 2$ cm in $p + p$ collisions. For Au + Au central collisions we achieve a resolution of $\sim 0.6$ cm.

The ZDC consists of two hadronic calorimeters, located $\sim 18$ m from the interaction point, that measure neutral energy emitted within $\sim 2$ milliradians of the beam direction. This energy is carried by neutrons produced either by Coulomb dissociation of the beam particle or by evaporation from beam spectators. The energy resolution of the ZDC is $\delta E/E \sim 218% / \sqrt{E(\text{GeV})}$ [46].

The centrality of Au + Au collisions is determined by the correlation between the BBC charge sum and the ZDC total energy [47].

B. Central Magnet

The Central Magnet (CM) is an axial field magnet energized by two pairs of concentric coils, roughly in a Helmholtz configuration, which can be run separately, together, or in opposition so that the momentum and the charge of a particle can be determined by its bending curvature [11]. In the mode in which both coils are running in the same direction, the single particle momentum resolution is better than 1% between 0.2 and 1 GeV/c. During the Au + Au measurement in Run-4 and the $p + p$ measurement in Run-5, the field component parallel to the beam axis had an approximately Gaussian dependence on the radial distance from the beam axis, dropping from 0.9 T at the center to 0.096 T (0.048 T) at the inner (outer) radius of the DCs. The total field integral is $\int B \cdot dl = 1.15$ Tm.

C. Tracking Detectors

The drift chambers (DCs) and the pad chambers (PCs) [12] in the central arms measure charged particle trajectories in the azimuthal direction to determine the transverse momentum ($p_T$) of each particle. The DC provides the most precise measurement of particle trajectories in the plane perpendicular to the collision axis. The first layer of PC provides the most precise measurement of the track space point along the collision axis. Additional layers principally supply pattern recognition support.

The DCs are located in the radial region $2.02 < r < 2.46$ m. Each DC volume consists of 20 sectors, each of which covers 4.5 degrees in azimuth and $|\eta| < 0.35$. Each sector has six types of wire modules stacked radially: X1, U1, V1, X2, U2 and V2. Each module is further divided into 4 drift cells in the $\phi$ direction. A plane of sense wires is at the center of each drift cell, with 2 to 2.5 cm drift space on either side. The X wires run parallel to the beam axis and measure the particle trajectory in the $r-\phi$ plane. The U and V wires have a stereo angle of about 6.0 degrees relative to the X wires in order to measure the $z$-coordinate of the track. Each X- and U, V-stereo cell contains 12 and 4 sense wires, respectively.
The single X wire position resolution is $\sim 150 \mu m$. The intrinsic tracking efficiency of the X modules is greater than 99%.

The pad chambers (PC) are multi-wire proportional chambers that form three separate layers. They determine space points along the straight line particle trajectories outside the magnetic field. The first PC layer (PC1) is located between the DC and the RICH, the second layer (PC2) is placed behind RICH (west arm only) and the third layer (PC3) is located in front of the EMCal. PC1 and the DC, along with the first layer (PC1) is located between the DC and the RICH, the second layer (PC2) is placed behind RICH (west arm only) and the third layer (PC3) is located in front of the EMCal. The construction to determine the polar angle of each charged track. The position resolution is $\pm 1.7$ mm for PC1 along the wire (z-direction).

Helium bags were installed between the beam pipe and the DCs to reduce the conversion material prior to the first tracking layer to $\sim 0.4\%$ of a radiation length. The material budget is known with an uncertainty of $\sim 5\%$.

D. Ring-imaging Čerenkov Counter (RICH)

The RICH is a threshold gas Čerenkov detector and is the primary detector to identify electrons in PHENIX \cite{42}. It is located in the radial region 2.5 $< r < 4.1$ m, just outside PC1. Each arm contains spherical mirror panels (0.53% of a radiation length) which focus Čerenkov light onto two arrays of $80(\phi) \times 16(z) = 1280$ PMTs. The PMTs are located outside the acceptance, on either side of the RICH entrance window, and are shielded to allow operation in a magnetic field up to 100 Gauss. The Čerenkov radiator gas, CO$_2$ at atmospheric pressure, has $n = 1.000410$ ($\gamma = 35$) which corresponds to a momentum threshold of 20 MeV/c for an electron and 4.65 GeV/c for a pion.

The average number of hit PMTs per electron track is about 5, and the average number of photo-electrons detected is about 10. Simulation studies show that pion rejection by the RICH alone, for isolated tracks, is limited by the production rate of “co-linear” delta electrons to one part in $10^4$. However in high multiplicity collisions pion tracks may be mistaken for electrons via overlap of their trajectory with a true electron’s ring. This effect worsens the RICH-alone pion rejection to roughly one part in $10^3$ for central Au + Au collisions and requires additional cuts in the offline analysis as described below.

E. Electromagnetic Calorimeter (EMCal)

The EMCal \cite{43} provides a measurement of the energies and the spatial positions of photons and electrons. Each arm consists of four rectangular sectors in $\phi$: the two bottom sectors of the east arm are lead-glass (PbGl) calorimeters; whereas, the remaining sectors are lead-scintillator (PbSc) calorimeters. The radial distance from the $z$ axis is 510 cm for the PbSc and 550 cm for the PbGl.

The PbSc is a Shashlik-type sampling calorimeter made of alternating tiles of lead and scintillator. It consists of $10.5 \times 10.5 \times 37$ cm$^3$ (18.2 $X_0$) modules, constructed of alternating layers of 1.5 mm thick lead, reflecting paper, and 4 mm thick polystyrene-based scintillator. Each module is divided into four equal towers, from which the light is collected separately by scintillating fibers. Each PbSc sector consists of $36(\phi) \times 72(z) = 2592$ towers. The nominal energy resolution is $\delta E/E \sim 4.5\% \pm 8.3\%/\sqrt{E}$ (GeV).

The PbGl is a Čerenkov counter that measures the light emitted by the particles in an electromagnetic shower and collected by one PMT at the back end. Each PbGl sector consists of 4608 $4 \times 4 \times 40$ cm$^3$ (16 $X_0$) modules made of lead-glass crystals. The PbGl has a nominal energy resolution of $\delta E/E \sim 4.3\% \pm 7.7\%/\sqrt{E}$ (GeV).

III. ANALYSIS

In this Section we present all steps of the data analysis. We start by introducing the data set, the event selection for $p + p$ and Au + Au and the centrality definition for Au + Au collisions \cite{1A}. We present the single electron analysis, including track reconstruction \cite{1B} and electron identification \cite{1C}. We present the details of the pair analysis \cite{1D} including pair cuts and photon rejection \cite{1E}, combinatorial and correlated background subtraction. An alternative method to subtract combinatorial and correlated background together is described \cite{1F}. Section III.F presents the final raw mass spectrum. In Section III.F we present the mass spectrum obtained from the analysis of runs with increased conversion material, a technique employed to estimate the systematic uncertainty on the background subtraction. Next we describe the efficiency \cite{1G} and acceptance \cite{1H} corrections, trigger efficiency \cite{1I} (for $p + p$ collisions) and occupancy correction (for Au + Au collisions) \cite{1J}. Finally we describe the calculation of the associated systematic uncertainties involved in the analysis \cite{1K}.

A. Data Sets and Event Selection

The data for $p + p$ collisions at $\sqrt{s} = 200$ GeV were collected during the polarized $p + p$ run in 2005. For this analysis two data sets were used: a reference sample of events selected with a minimum bias interaction trigger (Min. Bias) and a data set selected with a single electron trigger (ERT: EMCal and RICH Trigger). The Min. Bias trigger for $p + p$ requires at least one hit in both the North and South BBC detectors in coincidence with the beam bunch crossing and a $z$-vertex position (determined online by the BBCs) within 38 cm:

$$\text{Min. Bias} = (\text{BBCN} \geq 1) \times (\text{BCS} \geq 1) \times (|z| < 38 \text{ cm}).$$ (1)
The Min. Bias trigger cross section is measured to be $\sigma_{\text{BBC}} = 23.0 \pm 2.2 \text{ mb}$ or $54.5 \pm 5\%$ of the total inelastic $p + p$ cross section at this center of mass energy $\sigma_{\text{inel}}^{p+p} = 42 \pm 3 \text{ mb}$ [10]. Data collected without requiring the BBC trigger show that the BBC fires on 79\% of events with particles in the central arm acceptance. We assume that regardless of the electron $p_T$ or electron pair $p_T$ and invariant mass, the BBC always fired with the same probability of 79\%. Therefore, in the $p + p$ data the yield is divided by $0.79 / 0.545$ to account for the fraction of tracks (0.79) and inelastic $p + p$ collisions (0.545) missed by the Min. Bias trigger.

The ERT trigger requires a minimum energy deposit of 0.4 GeV in a tile of $2 \times 2$ EMCal towers matched to a hit in the RICH, in coincidence with the Min. Bias trigger. In the active area the ERT trigger has a very high efficiency for electrons; it reaches approximately 50\% at a hit in the RICH, in coincidence with the Min. Bias trigger, and saturates at $\sim 1$ GeV/$c$ close to 100\% (for the EMCal) and close to 90\% (for the RICH).

The offline Min. Bias trigger also requires one hit in one of the BBCs. The same $z$-vertex cut $|z| < 25 \text{ cm}$ as in the $p + p$ data is applied offline. This corresponds to 92.3\% of the Au + Au inelastic cross section.

The centrality is determined for each Au + Au collision by the correlation in the measurement of the BBC charge and ZDC energy [47]. Specifically we consider the following variables:

- $N_{\text{coll}}$ is the number of participants
- $N_{\text{part}}$ is the number of binary collisions
- $\phi_{\text{min}}$ is the minimum emission angle
- $\phi_{\text{max}}$ is the maximum emission angle

Table II summarizes the average $N_{\text{part}}$ and $N_{\text{coll}}$ and the corresponding systematic uncertainties in each centrality class used in the analysis.

We analyzed a sample of $8 \times 10^8$ minimum bias events, divided into five centrality classes (0-10\%, 10-20\%, 20-40\%, 40-60\%, and 60-92\%) for which the number of events is summarized in Table II.

### B. Track Reconstruction

The PHENIX tracking system reconstructs charged particles with momentum above 0.2 GeV/$c$ with a momentum resolution of $\sigma_{p_T}/p_T = 0.7\% \pm 1\% \cdot p_T$ for $p_T$ in GeV/$c$. A track is reconstructed by 2 sets of at least 4 hits in the X1 and X2 plane separated by 20 cm in radial direction, i.e. in the main bend plane of the central magnet, using a Hough transform performed over all possible hit combinations. The UV1 and UV2 wires provide up to 6 measurements in the $z$ direction, which are associated with the three-dimensional space point provided by PC1. After the pattern recognition and track reconstruction by the Hough transform technique, the initial momentum vector of the track at the $z$-vertex is calculated. Each reconstructed track is then associated with hit information from the outer detectors (PC2, PC3, RICH, and EMCal).

The transverse momentum ($p_T$) is determined by measuring the angle $\alpha$ between the reconstructed particle trajectory and a line that connects the $z$-vertex point to the particle trajectory at a reference radius $R=220 \text{ cm}$. The angle $\alpha$ is approximately proportional to $|p_T|$. Note that this procedure assumes tracks originate from the vertex. As a result, tracks which originate off vertex are reconstructed with an incorrect momentum. Conversion pairs are reconstructed with invariant mass $m_{ee} > 0$ and contaminate the spectrum up to $m_{ee} \sim 0.3 \text{ GeV}/c^2$ (see Section II D 2).

Because charged particles are deflected in the azimuthal direction by the magnetic field, the single-track acceptance depends on the momentum and charge of the particle, and also on the radial location of the detector component (DC and RICH). The acceptance for a track with charge $q$, transverse momentum $p_T$ and azimuthal emission angle $\phi$ can be described by the logical AND of these conditions:

$$\phi_{\text{min}} \leq \phi \leq \phi_{\text{max}}$$

where $k_{\text{DC}}$ and $k_{\text{RICH}}$ represent the effective azimuthal bend to DC and RICH ($k_{\text{DC}} = 0.206 \text{ rad} \cdot \text{GeV}/c$ and $k_{\text{RICH}} = 0.309 \text{ rad} \cdot \text{GeV}/c$). One arm covers the region from $\phi_{\text{min}} = 220 \text{ cm}$ to $\phi_{\text{max}} = 309 \text{ rad}$, the other arm from $\phi_{\text{min}} = 11 \frac{\pi}{16} \text{ rad}$ to $\phi_{\text{max}} = 19 \frac{\pi}{16} \text{ rad}$.

Results in Sec. V will show the dilepton invariant mass spectrum “in the PHENIX acceptance,” where the data will be compared to the expectations filtered according to this simple parameterization of the acceptance.

### C. Electron Identification

Electrons in the range $0.2 < p_T < 20 \text{ GeV}/c$ are identified by hits in the Ring Imaging Čerenkov detector (RICH) and by matching the momentum with the energy measured in the electromagnetic calorimeter (EMCal) [4]. Specifically we consider the following variables for electron identification (eID), as summarized in Table II.

- **track quality**: A bit pattern representing the reconstruction quality of the track. If the track is reconstructed by both of the X1 and X2 sections of the DC and is uniquely associated with hits in U or...
TABLE I: \( N_{\text{part}} \), \( N_{\text{coll}} \) for Au + Au collisions at \( \sqrt{s_{NN}} = 200 \text{ GeV} \) with the corresponding uncertainties derived from a Glauber calculation \[2\] and the number of events and pairs for each centrality class. Note that the uncertainties are correlated \[48\].

| Centrality class | \( \langle N_{\text{part}} \rangle \) (syst) | \( \langle N_{\text{coll}} \rangle \) (syst) | \( N_{\text{events}} \) | Signal Pairs |
|------------------|---------------------------------|---------------------------------|-----------------|-------------|
| 0-10 %           | 325.2 (3.3)                     | 955.4 (93.6)                    | \( 8.6 \times 10^7 \) | 9.2 \times 10^4 |
| 10-20 %          | 234.6 (4.7)                     | 602.6 (59.3)                    | \( 8.6 \times 10^7 \) | 6.6 \times 10^4 |
| 20-40 %          | 140.4 (4.9)                     | 296.8 (31.1)                    | \( 1.7 \times 10^8 \) | 8.1 \times 10^4 |
| 40-60 %          | 59.95 (3.6)                     | 90.70 (11.8)                    | \( 1.7 \times 10^8 \) | 3.3 \times 10^4 |
| 60-92 %          | 14.50 (2.5)                     | 14.50 (4.00)                    | \( 2.9 \times 10^8 \) | 1.1 \times 10^4 |
| 0-92 %           | 109.1 (4.1)                     | 257.8 (25.4)                    | \( 8.1 \times 10^8 \) | 28.3 \times 10^4 |
| \( p + p \) (Min. Bias) | 2                              | 1                               | \( 1.5 \times 10^9 \) | 1.4 \times 10^4 |
| \( p + p \) (ERT) | 2                              | 1                               | \( 2.7 \times 10^8 \) | 22.8 \times 10^4 |

V stereo wires, the value of \textit{quality} is 63 (in case a unique PC1 hit is found too) or 31 (in case the PC1 hit is found but ambiguous). If there are no UV hits found, but a PC1 hit is, \textit{quality} is 51.

\textbf{EMCal match (\( \sigma_{\Delta \phi} \))} Displacement in \( \phi \) between the position of the associated EMCal cluster and the projection of the track onto the EMCal. The quantity is measured in units of momentum-dependent resolution. For example, \( \Delta \phi < 2 \) means that the position of the associated EMCal cluster in \( \phi \) is within \( 2\sigma \) of the projected track position. The particle hit position of an EMCal cluster is particle-species dependent due to different shower shapes. Here the parameterization has been optimized for electrons.

\textbf{EMCal match (\( \sigma_{\Delta z} \))} Analogous to the previous variable, for the \( z \) coordinate.

\textbf{n0} Number of hit RICH PMTs in an annular region with an inner radius of 3.4 cm and outer radius of 8.4 cm around the track projection on the RICH. The expected radius of a Čerenkov ring emitted by an electron is 5.9 cm.

\textbf{chi2/npe0} A \( \chi^2 \)-like shape variable of the RICH ring associated with the track divided by the number of photo-electrons measured in a given ring (npe0).

\textbf{RICH match} The displacement of the RICH ring center from the projected track position. Units are cm.

\textbf{E/p or dep} A variable quantifying energy-momentum matching. This variable is calculated as \( \text{dep} = (E/p - 1)/\sigma_{E/p} \), where \( E \) is the energy measured by EMCal; \( p \) is the momentum of the track; and \( \sigma_{E/p} \) is the standard deviation of the Gaussian-like \( E/p \) distribution calculated for electrons.

![Image](image-url)

**FIG. 3:** (Color online) \( E/p \) distribution in minimum bias Au + Au for all charged tracks and for tracks after applying all the RICH cuts in Table II except the \( E/p \) or dep. The contribution from randomly associated hadrons is shown by the filled histogram.

**TABLE II:** Electron ID cuts used in the Au + Au and \( p + p \) analyses.

| eID cuts | for Au + Au | for \( p + p \) |
|----------|-------------|-----------------|
| track quality = 63 || 31 || 51 | 63 || 31 || 51 |
| \( \sqrt{\sigma_{\Delta \phi}^2 + \sigma_{\Delta z}^2} < 3.0 \) | 5.0 |
| n0 \( \geq 2 \) | 1 |
| chi2/npe0 \( < 10.0 \) | 15.0 |
| RICH match \( < 5.0 \) | 10 |
| dep \( > -2.0 \) | - |
| \( E/p > \) | 0.5 |

The contribution from randomly associated hadrons is shown by the filled histogram.
D. Backgrounds

The source of any particular electron or positron in an event is unknown; therefore, all electrons and positrons are combined into a foreground of pairs, like-sign $N_{++}, N_{--}$ and unlike-sign $N_{+-}$. This results in a large combinatorial background which must be removed. In the following we will use the notation for the foreground $N_{\pm\pm} = N_{++} + N_{--}$ and for the background $B_{\pm\pm} = B_{++} + B_{--}$. The analysis steps to achieve this are outlined here and presented in detail in the subsections below. We can distinguish our background sources in two types:

Type I background (III D 1) consists of two classes of fake combinations that can be identified on a pair-by-pair basis:

- Overlapping pairs are fake electron pairs that arise from overlapping hits in the detectors, mostly in the RICH.
- Photon conversions are fully reconstructed pairs originating from photon conversions in the detector material are removed by a cut on the orientation of the pairs in the magnetic field.

Type II background (III D 2) consists of all those pairs that cannot be identified on a pair-by-pair basis and are therefore removed statistically:

- Combinatorial background $B_{\text{comb}}$ arises from all the combinations where the origin of the two electrons is totally uncorrelated.
- Correlated background $B_{\text{corr}}$ occurs if there are two $e^+e^-$ pairs in the final state of a meson or when two hadrons, either within the same jet or in back-to-back jets, decay into electron pairs.

Since accurate background subtraction is essential for this analysis, we have developed two independent methods to subtract the type II background. In the first method we calculate the shapes of combinatorial and correlated background with event mixing or simulations and use the yield of the like-sign spectra for the normalization. In the second method we use measured the real like-sign distributions, corrected for the acceptance difference, without making any assumption about the mass dependence nor about the decomposition of the background into correlated and uncorrelated component.

1. Type I Background

Overlapping Pairs: Fake electron pairs can be created if two particles are in close proximity in any of the detectors. These correlations within an event are particularly noticeable when two tracks share the same RICH ring. This issue can be illustrated with a simplified model of the RICH based on spherical mirror optics. In this case, tracks that are parallel to each other while passing through the RICH radiator, i.e., after they have been bent in the magnetic field, share the same search region for Čerenkov light in the RICH PMT plane. Therefore an overlapping pair is created whenever a track after the field is parallel to a true electron. These overlapping pairs typically have a small opening angle and are therefore reconstructed with small invariant mass. Because like-sign pairs are bent in the same direction, in contrast to unlike-sign pairs which are bent in opposite directions, like-sign overlapping pairs have smaller mass than unlike-sign overlapping pairs. Figure 3 shows the invariant mass distribution for all like- and unlike-sign pairs, the distributions for overlapping pairs, and the distributions after removing overlapping pairs.

Overlapping pairs are eliminated by applying a cut on the physical proximity of every pair projected to every detector. The cut value on every detector is determined by the corresponding double hit resolution. In the RICH the cut is applied at 36 cm, which is roughly twice the predicted maximum diameter ($\sim 16.8$ cm) of a RICH ring. In the Pad Chambers the cut is applied at $\Delta z \leq 0.5$ cm and $\Delta \phi \leq 20$ mrad. In the EMCal the cut is applied to a $3 \times 3$ tower region around the hit. These cuts remove a fraction of real pairs which varies from 4% in the most central to 2% in the most peripheral collisions, estimated using mixed events. These cuts remove 10% more real like-sign than unlike-sign pairs. The ratio of like- to unlike-sign pairs lost was determined with an uncertainty of 50% by varying the cut values chosen. While the efficiency of the pair cut depends on the centrality, the ratio of like- to unlike-sign pairs lost was found to be independent of centrality.

Because these overlapping pairs are rare, whenever we encounter one, we remove the entire event. This results

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1 Data tables for this and other data plots are available at https://www.phenix.bnl.gov/WWW/p/info/ppg/088/datatables.
in a loss of $\sim 0.08\%$ of all events. In mixed events the same cuts are applied and whenever an event is discarded due a pair cut, another event is generated with the same electron multiplicity and in the same centrality, $z$-vertex and reaction plane class.

Photon Conversions Since the tracking algorithm assumes that all tracks originate from the collision vertex, pairs from photons that convert off-vertex are reconstructed with an artificial opening angle, which leads to an invariant mass that increases with the radius at which the conversion occurs.

Conversion pairs have no intrinsic opening angle (i.e. their opening angle is exactly zero at the point of creation), they are bent only in the azimuthal direction by the magnetic field, which is parallel to the beam axis $\hat{z}$.

We can define unit vectors $\hat{u}$ in the direction of the pair momentum and $\hat{v}$ perpendicular to the plane defined by the pair

$$\hat{u} = \frac{\vec{p}_+ + \vec{p}_-}{|\vec{p}_+ + \vec{p}_-|} \quad (4)$$

$$\hat{v} = \frac{\vec{p}_+ \times \vec{p}_-}{|\vec{p}_+ \times \vec{p}_-|} \quad (5)$$

where $\vec{p}_\pm = \vec{p}_+ / |\vec{p}_+|$.

We can define the orientation of the actual opening angle as

$$\hat{w} = \hat{u} \times \hat{v} \quad (6)$$

We can also define the expected orientation of the opening angle for conversion pairs

$$\hat{w}_c = \hat{u} \times \hat{z} \quad (7)$$

Finally we can define $\phi_V$ as the angle between these two vectors

$$\cos \phi_V = \hat{w} \cdot \hat{w}_c \quad (8)$$

For pairs originating from photon conversions $\phi_V$ is zero. (By consistently ordering positive and negative tracks within the pair we avoid $\phi_V = \pi$ as a solution for photon conversions.) In contrast, $e^+e^-$ pairs from hadron decays, as well as combinatorial pairs, have no preferred orientation. Figure 5 presents a comparison of the $\phi_V$ angle distributions of real data and Monte Carlo simulated data in a mass bin, $10 < m_{ee} < 30$ MeV/c$^2$, dominated by photon conversion in the beam-pipe. In the simulations we can distinguish which $e^+e^-$ pairs originate from a $\pi^0$ Dalitz decay (dotted-dashed line) and which originates from photon conversion (dotted line). The simulations show that the distribution for all unlike-sign pairs originating from photon conversion is strongly peaked at $\phi_V = 0$. In contrast, simulated pairs originating from $\pi^0$ Dalitz decays have no strongly preferred orientation.
The sum of the Monte Carlo data can be compared to the signal, which in this mass region contains conversion photons, as well as $\pi^0$'s. Figure 6 shows that the agreement between the signal and the simulation is good. The width of the $\phi_V$ peak for photon conversions increases with the path length of the $e^+e^-$ pair in the magnetic field, where residual field in the polar direction as well as multiple scattering cause the pair to lose its perfect alignment perpendicular to the $\vec{z}$ axis.

The contribution from conversion pairs as a function of the (mis-reconstructed) invariant mass is shown in Figure 6 (filled histogram). As the (mis-reconstructed) mass is essentially proportional to the radius where the conversion happens, the mass spectrum of those pairs allows a “tomography” of the material in the spectrometer. The peaks correspond to the conversions in the beam-pipe material ($r = 4$ cm, or $m_{ee} = 20$ MeV/c$^2$) and detector support structures ($r = 25$ cm, or $m_{ee} = 125$ MeV/c$^2$). Conversions in the He bag generates pairs with $m_{ee} \lesssim 0.3$ GeV/c$^2$. For this value of mass the corresponding radius would be the entrance window of the DC. At this point though the electrons do not bend anymore because the region is field-free, and are therefore removed with a $p_T$ cut ($p_T < 20$ GeV/c) on the single electrons. The $\phi_V$ resolution improves for increasing conversion radius because electrons are less modified from their original direction by multiple scattering or the residual polar field. The cut on $\phi_X$ (indicated in Fig. 5 by the line at $\phi_X > 0.25$ for $m_{ee} < 30$ MeV/c$^2$) is reduced for larger masses due to the improved resolution of $\phi_V$ at larger radii. By varying the cut values on real pairs we estimate that our $\phi_V$ cut removes more than 98% of the conversion pairs with a mass-dependent efficiency of more than 90%. The uncertainty of on the final $e^+e^-$ pair signal is 6%.

### 2. Type II Background

After removing type I background, the unlike-sign foreground spectrum $N_{++}$ measures the physical signal plus background, while the like-sign spectra $N_{++}, N_{--}$ measure only background. We have developed two methods to measure and subtract the unlike-sign background

- One solution is to use a mixed-event technique, which combines tracks from different events. With this method the background has much larger statistics than the foreground. The accuracy in the determination of the shape of the background is tested by comparing the like-sign distribution in real and mixed events. We find good agreement between real and mixed-events like-sign spectra in some regions of the $(m_{ee}, p_T)$ plane, while in others they clearly deviate. This indicates that not all the type II background is of uncorrelated origin, but there are also some correlated pairs in the background (II D2) that one needs to separately account for.

- Another solution is to still use the measured like-sign spectra and correct them for the different acceptance (II D3). This solution does not require any assumption on the decomposition of combinatorial and correlated since the like-sign spectra measures all the backgrounds simultaneously. In experiments with equal acceptance for electrons and positrons, the background can be measured directly through the geometric mean of the like-sign pair distributions $2\sqrt{N_{++}N_{--}}$. With this method the background has similar statistics as the foreground. In PHENIX however the mass and $p_T$ dependence of the background is different for the two charge combinations (see Equation 3 and Fig. 4).

We have used both these methods, which are compared in Section II D 4 to assign a systematic uncertainty to the background subtraction. The first method is used for the final analysis.

The structure of the type II background was studied using minimum bias events generated with PYTHIA [41] with the branching ratio of the $\pi^0$ Dalitz decay set to 100% to enhance the sample of $e^+e^-$ pairs per event. All the electrons are filtered through the PHENIX acceptance. From PYTHIA events we make real and mixed-event like- and unlike-sign distributions as in the real data. Now we analyze like-sign pairs, that contain only background. The $\Delta\phi$ distribution shown in Fig. 7 compares the difference of azimuthal emission angle of the two electrons (or positrons) in real and mixed PYTHIA events. If the two shapes were identical, we would conclude that all the background come from uncorrelated sources. However the shapes clearly deviate at $\Delta\phi \sim 0$ and $\Delta\phi \sim \pi$. This indicates the presence of combinations that arise from the same jet ($\Delta\phi \sim 0$) or back-to-back ($\Delta\phi \sim \pi$) jets (see Section II D3). In addition,
correlations can occur if there are two $e^+e^-$ pairs in the final state of a meson, e.g. double Dalitz decays, Dalitz decays followed by a conversion of the decay photon or two-photon decays followed by conversion of both photons.

Therefore, we decompose the type II background into two components: a combinatorial background made of uncorrelated pairs and a background of correlated pairs. The distributions for combinatorial and correlated background are determined with methods that are explained below (III D 3) for like and unlike-sign spectra. The like-sign background distributions are fit to the measured like-sign spectra (that contain only background). The same fit parameters are then applied to the unlike-sign background. The normalized unlike-sign background is finally subtracted from the foreground of all pairs $N_{++}$ to obtain the signal.

**Combinatorial Background** The combinatorial background $B_{comb}$ is determined with a mixed-event technique, which combines tracks from different events with similar centrality, z-vertex and reaction plane. In the $p+p$ data, where we use a triggered data set, the mixed-event pairs are constructed from the Min. Bias data set requiring that at least one of the two partners has fulfilled the ERT trigger condition. Since the tracks are from different events, this technique reproduces the uncorrelated background by definition. This technique also allows computation of background spectra with negligible statistical errors.

We compare like-sign spectra in real and mixed-event data to locate a region in the $(m_{ee}, p_T)$ plane where their shapes agree. In this region we normalize the combinatorial background spectra $B_{comb}$ to the measured like-sign pairs $N_{++}$. We define:

$$A_+ = \frac{\int_{N.R.} N_{++}(m_{ee}, p_T) dm_{ee} dp_T}{\int_{N.R.} B_{++}(m_{ee}, p_T) dm_{ee} dp_T}$$

$$A_- = \frac{\int_{N.R.} N_{--}(m_{ee}, p_T) dm_{ee} dp_T}{\int_{N.R.} B_{--}(m_{ee}, p_T) dm_{ee} dp_T} \quad (9)$$

where $N.R.$ is the chosen normalization region. Then we calculate the integral of the normalized like-sign background over the full phase-space:

$$B_{++} = \int_0^\infty A_+ \cdot B_{++}(m_{ee}, p_T) dm_{ee} dp_T$$

$$B_{--} = \int_0^\infty A_- \cdot B_{--}(m_{ee}, p_T) dm_{ee} dp_T$$

The unlike-sign background is then normalized such that its yield equals the geometric mean of the like-sign pairs $2\sqrt{B_{++}B_{--}}$:

$$B_{--}(m_{ee}, p_T) = \frac{2\sqrt{B_{++}B_{--}}}{\int_0^\infty B_{++}(m_{ee}, p_T) dm_{ee} dp_T} B_{++}(m_{ee}, p_T) \quad (10)$$

Appendix A shows that as long as electrons and positrons are produced in pairs the absolute normalization of the unlike-sign background is given by the geometric mean of the observed positive and negative like-sign pairs $2\sqrt{N_{++}N_{--}}$, without any further assumption about efficiencies, acceptances or probability distribution functions for the pair. Using $B_{++}$ and $B_{--}$ instead of directly taking $N_{++}$ and $N_{--}$ simply avoids counting correlated pairs, which will be measured and normalized separately (III D 3).

The systematic uncertainty of the normalization is therefore determined by the statistical accuracy of the measured like-sign yield in the region chosen for the normalization.

**Combinatorial Background in $p+p$ data**

In our PYTHIA [3] studies we found (Fig. 5) that at $\Delta\phi \sim 0$ and $\Delta\phi \sim \pi$ the real events background deviates from the shape of uncorrelated sources, while at $\Delta\phi \sim \pi/2$ it looks consistent with the shape from mixed events. Considering that $m_{ee}^2 = 2p_T p_z (1 - \cos \Delta\omega)$, with $\Delta\omega$ being the opening angle, this condition corresponds to a region in the $(m_{ee}, p_T)$ plane where $m_{ee} \sim p_T$.

We define this region empirically by a set of equations:

$$m_{ee} > 0.3 \text{ GeV}/c^2$$

$$m_T < 1.2 \text{ GeV}/c^2$$

$$p_T/c - 1.5m_{ee} \leq 0.2 \text{ GeV}/c^2$$

$$p_T/c - 0.75m_{ee} \geq 0 \text{ GeV}/c^2 \quad (12)$$

where $m_T = \sqrt{p_T^2 + m_{ee}^2}$ is the transverse mass of the pair. This region is shown in Fig. 8 by the dashed area.

Figure 8 shows the difference between the like-sign distributions in real and mixed-events, as a function of $m_{ee}$ and $p_T$. The background $B_{comb}$ is normalized to the foreground $N_{++}$ in the normalization region (from Equation 43). The absolute normalization of the unlike-sign combinatorial background is determined with an uncertainty given by the statistical error of the measured like-sign spectra in this region of 3%.
deviation. Figure 8 shows that in this region the background does not deviate from the foreground by more than \( 2 \times \sigma \). The stability of the results has been checked by varying the normalization region, and the difference is included in the systematic uncertainty.

![FIG. 8: Difference between real and mixed-events like-sign distributions divided by its standard deviation \((N_{\pm\pm} - B_{\pm\pm}^{\text{comb}}) / \sigma(N_{\pm\pm} - B_{\pm\pm}^{\text{comb}})\). The background \( B_{\pm\pm}^{\text{comb}} \) is normalized to the foreground \( \bar{N}_{\pm\pm} \) in the normalization region shown by the dashed area.](image)

**Combinatorial Background in Au + Au data**

Figure 8 shows a comparison between the like-sign distribution from real and mixed events. The comparison shows that the mixing technique reproduces the mass dependence within the statistical accuracy of the data not only in the normalization region for the \( p + p \) data, but also for all masses above 0.55 GeV/c\(^2\). In general the larger combinatorial background produced in the Au + Au environment reduces the capability to distinguish between different background shapes. Also the agreement at high mass, which was not observed in \( p + p \) (see Fig. 8), can be qualitatively explained by the suppression of away-side jets observed in Au + Au.

A small signal from correlated background remains at low-masses (see Section III D 2). To quantitatively compare the mass dependence of the data to the mixed events we calculate the ratio \((N_{\pm\pm} - B_{\pm\pm}^{\text{comb}}) / B_{\pm\pm}^{\text{comb}}\) shown in the bottom panels of Fig. 8 and fit it with a constant above the \( \eta \) mass (0.55 GeV/c\(^2\)). The result is \((-2.59 \pm 6.33) \times 10^{-4} \) with \( \chi^2/\text{NDF} = 27.6/19 \). Panel (b) shows the entire mass range and allows to distinguish the signal from correlated background at very low mass. Panel (c) shows a zoom in the region where we fit.

Two statistical tests are performed to test the hypothesis that the two distributions \((N_{\pm\pm} \pm B_{\pm\pm}^{\text{comb}})\) represent identical distributions: the Pearson \( \chi^2 \) test and the Kolmogorov-Smirnov test.

The Pearson test statistic \( \chi^2 \) resulting from the comparison of real and mixed events for \( m_{ee} > 0.55 \) GeV/c\(^2\) returns a \( p \)-value greater than 0.83.

The Kolmogorov-Smirnov test is used to decide if a sample comes from a population with a specific distribution, either comparing one fluctuating distribution (the test) to a truth hypothesis (reference), or querying two fluctuating distributions as to whether they have a common truth origin. This latter one was done in our case.
TABLE III: Fit parameters for the mass dependence comparison of real and mixed-events like-sign pairs for different centrality and $p_T$ bins. The second column reports the results of $(N_{\pm \pm} - B_{\pm \pm}^{\text{comb}})/B_{\pm \pm}^{\text{comb}}$ distributions to a constant and the third is $\chi^2$ value divided by the number of degrees of freedom. The fourth and fifth columns report the result of a $\chi^2$ statistical test and the corresponding $p$-value for $m_{ee} > 0.55 \text{ GeV}/c^2$. The last one gives the maximum deviation of the $N_{\pm \pm}$ and $B_{\pm \pm}^{\text{comb}}$ distribution in a Kolmogorov-Smirnov test.

| Centrality   | $p_T$ < 1 GeV/$c$ | 1 < $p_T$ < 2 GeV/$c$ | $p_T$ > 2 GeV/$c$ |
|--------------|-------------------|------------------------|-------------------|
| $0-10\%$     | $6.3 \pm 8.8 \times 10^{-4}$ | $9.2 \pm 5.1 \times 10^{-4}$ | $-3.4 \pm 1.6 \times 10^{-3}$ |
| $10-20\%$    | $-9.4 \pm 1.4 \times 10^{-4}$ | $18.6/19$               | $27.9/18$          |
| $20-40\%$    | $-2.4 \pm 1.8 \times 10^{-3}$ | $18.7/19$               | $21.9/19$          |
| $40-60\%$    | $-8.5 \pm 4.9 \times 10^{-3}$ | $1.02$                  | $1.65$             |
| $60-92\%$    | $-1.8 \pm 1.6 \times 10^{-2}$ | $1.59$                  | $1.51$             |
| $00-92\%$    | $2.6 \pm 6.3 \times 10^{-4}$ | $27.6/19$               | $0.92$             |
| $p_T < 1 \text{ GeV}/c$ | 18.9/18 | 0.91 | 0.84 |
| $1 < p_T < 2 \text{ GeV}/c$ | 15.2/18 | 0.97 | 0.63 |
| $p_T > 2 \text{ GeV}/c$ | 21.5/14 | 0.92 | 0.83 |

TABLE IV: Systematic uncertainty on the combinatorial background for $p+p$ and various centrality classes of $\text{Au + Au}$ collisions.

| Centrality | $\delta B_{\pm \pm}^{\text{comb}}/B_{\pm \pm}^{\text{comb}}$ (%) |
|------------|---------------------------------------------------------------|
| $0-92\%$   | 0.25                                                          |
| $0-10\%$   | 0.25                                                          |
| $10-20\%$  | 0.30                                                          |
| $20-40\%$  | 0.35                                                          |
| $40-60\%$  | 0.85                                                          |
| $60-92\%$  | 2.50                                                          |
| $p + p$    | 3.00                                                          |

since we do not have a truth reference distribution without fluctuations. The test is based on the maximum cumulative difference between the distribution under test and a specified reference distribution. The left panel of Fig. 10 shows the maximum cumulative difference $d$ of the two like-sign distributions $N_{\pm \pm}$ and $B_{\pm \pm}^{\text{comb}}$ for $m_{ee} > m_{\text{low}}$, where $m_{\text{low}}$ is the lower endpoint chosen between 0.25 and 0.7 GeV/$c^2$. The Kolmogorov-Smirnov test gives a maximum deviation of 0.1% which is small compared to the uncertainty of the absolute normalization of the mixed-event background. The corresponding Kolmogorov-Smirnov $p$-value is also shown in the right panel of Fig. 10 as a function of mass. It is small below 0.55 GeV/$c^2$, where there is some contribution of the correlated background. However, for $m_{\text{low}} > 0.55 \text{ GeV}/c^2$, the $p$-value is ~90%, therefore confirming that the hypothesis of compatibility of the two distributions for $m_{\text{low}} > 0.55 \text{ GeV}/c^2$ is valid for any commonly used significance level.

Figure 11 shows the like-sign mass distribution for real and mixed events in the different centrality bins used in the analysis. The bottom panels show the ratio $(N_{\pm \pm} - B_{\pm \pm}^{\text{comb}})/B_{\pm \pm}^{\text{comb}}$ which fit to a constant for $m_{ee} > 0.55 \text{ GeV}/c^2$. The fit results for all centralities are reported in Table II together with the results of the statistical tests described above. The results reported in Table II demonstrate that the agreement between real and mixed-event like-sign mass spectra demonstrated for minimum bias collisions also holds for all centrality classes.

Figure 12 shows the like-sign mass distribution for real and mixed events in different $p_T$ bins. The ratio $(N_{\pm \pm} - B_{\pm \pm}^{\text{comb}})/B_{\pm \pm}^{\text{comb}}$ shows a good agreement between real and mixed-event distribution for all bins. This demonstrates that, within the statistical error of the foreground, there is no deviation from uncorrelated combinatorial behavior for masses above 0.55 GeV/$c^2$ in any $p_T$ range.

The absolute normalization of the unlike-sign combinatorial background is determined with an uncertainty given by the statistical error of the measured like-sign spectra for $m_{ee} > 0.7 \text{ GeV}/c^2$ of 0.12%. The mass interval ($m_{ee} > 0.7 \text{ GeV}/c^2$) is sufficiently large to achieve the desired statistical accuracy and is conservatively chosen to exclude any possible region which may be contaminated by the correlated background. The results are stable when varying the lower end of the normalization region between 0.55 and 0.7 GeV/$c^2$, and the difference is included in the systematic uncertainty.

Because the pair cuts remove more like-sign than unlike-sign pairs, (see Section III D 3) the normalization factor is corrected by this asymmetry, which is estimated with mixed events to be 1.004 ± 0.002, independent of centrality.

For the various centrality bins and minimum bias collisions, the 0.2% uncertainty on the event rejection is added in quadrature to the uncertainty on the normalization, which is determined by the statistics of the like-sign pairs. Since the ratios shown in Figs. 11, 12, and 13 show no systematic deviation in shape between the like-sign real- and mixed-event distribution for $m_{ee} > 0.55 \text{ GeV}/c^2$, the uncertainty due to the shape is negligible compared to the uncertainty on the normalization. The total uncertainty on the combinatorial background is given in Table IV.
This translates into a systematic uncertainty $\delta S$ on the signal $S$ of $\delta S/S = (\delta B_{\text{comb}}^+/B_{\text{comb}}^+)/S/B_{\text{comb}}^+$. $\delta B_{\text{comb}}^+ / B_{\text{comb}}^+$ for minimum bias collisions and all centrality bins are listed in Table I. Figure 13 shows $S/B_{\text{comb}}^+$ for different ranges of $p_T$ and centrality.

Correlated Background After subtracting the combinatorial background, the remaining pair distributions, like- and unlike-sign, are considered correlated pairs, where the like-sign distribution only contains correlated
Correlated Background in $p+p$ data

Panel a) of Fig. 14 shows like-sign pair distributions: $N_{\pm\pm}$, $B_{\pm\pm}^{\text{comb}}$ and $B_{\pm\pm}^{\text{corr}} = N_{\pm\pm} - B_{\pm\pm}$. Panel b) shows analogous distributions for unlike-sign pairs: $N_{+-}$, $B_{+-}$ and $B_{+-}^{\text{corr}} = N_{+-} - B_{+-}$.

The like-sign correlated background distribution is fit to

$$B_{\pm\pm}^{\text{corr}} = A \cdot \text{cross} + B \cdot \text{jet}$$

The resulting normalization factors, $A$ and $B$, are then applied to the unlike-sign correlated background. The unlike-sign signal results from subtracting the correlated background from all correlated unlike-sign pairs.

Correlated Background in Au + Au data

In Au + Au data the like-sign mixed-event distribution reproduces the mass dependence of the real-event distribution not only in region of Equation 12, but for all masses above 0.55 GeV/$c^2$ for every centrality, as shown in Fig. 11, and at every $p_T$, as shown in Fig. 12. This means that there is no room for correlated background for $m_{ee} > 0.55$ GeV/$c^2$. Here the contribution that typically arises from back-to-back jets in the Au + Au data is indeed expected to be different than $p+p$ because of the observed jet modifications.

We therefore separate the jet distribution into “near-side” ($\text{jet}_{\text{near}}$: $\Delta \phi < \pi/2$) and “away-side” contributions ($\text{jet}_{\text{away}}$: $\Delta \phi > \pi/2$) and we fit the like-sign correlated background distribution to the sum of

$$B_{\pm\pm}^{\text{corr}} = A \cdot \text{cross} + B \cdot \text{jet}_{\text{near}} + C \cdot \text{jet}_{\text{away}}$$

Panels c) and d) of Fig. 14 show the like- and unlike-sign pair distributions, the normalized mixed-event background and the distributions after subtraction. We note that the like-sign yield is well described by the sum of combinatorial and correlated background, and that the contribution from “away-side-jet pairs” is consistent with
FIG. 14: (Color online) Raw dielectron spectra in p+p (left) and Au + Au (right) collisions. The top panels show like-sign pairs \( N_{+\pm} \) as measured in the experiment, the combinatorial background from mixed-events \( B_{+\pm}^{\text{comb}} \), the correlated pair background \( B_{+\pm}^{\text{corr}} \) obtained by subtracting the combinatorial background, and the individual contributions from cross and jet pairs to the correlated background (see text). The bottom panels show the same distributions for unlike-sign pairs. The correlated like-sign background \( B_{\pm\pm}^{\text{corr}} \), is normalized to the measured like-sign pairs remaining after subtracting the combinatorial background \( N_{\pm\pm} - B_{\pm\pm}^{\text{comb}} \), and the same factors are applied to the unlike-sign distribution \( B_{\pm\pm}^{\text{corr}} \).

The measured distributions however are different because of the acceptance.

The acceptance difference between like- and unlike-sign pairs can be measured with the ratio of mixed-event distributions, i.e.

$$\frac{B^{\text{comb}}_{++}}{2 \sqrt{B^{\text{comb}}_{++} B^{\text{comb}}_{-+}}}$$

where \( B^{\text{comb}}_{++, B^{\text{comb}}_{+-}} \) are the number of \( e^+e^-, e^+e^+, \) and \( e^-e^- \) pairs in a given mass and \( p_T \) bin. The like-sign distribution can therefore be acceptance-corrected and then subtracted from the unlike-sign pairs.

3. Like-sign Subtraction Method

The event-mixing technique, along with the determination of correlated backgrounds, used to estimate the background contribution to the measured dilepton mass spectra, was developed in order to get around the problems introduced in a traditional like-sign background calculation by the asymmetric PHENIX acceptance. With a like-sign calculation we need to make only the assumption that the mass dependence of the correlated background is symmetric for like- and unlike-sign pairs. We do not need to make any assumption about the decomposition of the background into “cross”, “jet-near-side” and “jet-away-side” pairs. The like-sign distribution measures only the background. Correlated and uncorrelated backgrounds are charge symmetric, i.e. they yield the same number of like-and unlike-sign pairs, with the same distribution.

The measured distributions however are different because of the acceptance.

The acceptance difference between like- and unlike-sign pairs can be measured with the ratio of mixed-event distributions, i.e.

$$\frac{B^{\text{comb}}_{++}}{2 \sqrt{B^{\text{comb}}_{++} B^{\text{comb}}_{-+}}}$$

where \( B^{\text{comb}}_{++, B^{\text{comb}}_{+-}} \) are the number of \( e^+e^-, e^+e^+, \) and \( e^-e^- \) pairs in a given mass and \( p_T \) bin. The like-sign distribution can therefore be acceptance-corrected and then subtracted from the unlike-sign pairs.

$$S_{+-} = N_{+-} - 2 \sqrt{N_{++} N_{-+}} \cdot \frac{B^{\text{comb}}_{++}}{2 \sqrt{B^{\text{comb}}_{++} B^{\text{comb}}_{-+}}}$$  \hspace{1cm} (17)$$

This technique however measures the background distribution with similar statistical error as the foreground, therefore the resulting error on the signal is much larger than when using the mixed events technique.

One can also use a “hybrid” method where first the mixed events are normalized in the region where real and mixed-event distributions agree (region in Equation 12 for \( p + p \) and \( m_{ee} > 0.7 \) GeV/\( c^2 \) for \( \text{Au} + \text{Au} \)). Then the correlated background is obtained by correcting the subtracted like-sign distribution by the acceptance difference

$$\frac{B^{\text{comb}}_{++}}{2 \sqrt{B^{\text{comb}}_{++} B^{\text{comb}}_{-+}}}.$$
4. Systematic Uncertainty of the Background Subtraction

The background subtraction is the most critical aspect of this analysis, therefore it is crucial to assign a proper systematic uncertainty to it.

In Section II D 2 we reported the systematic uncertainty on the normalization of the combinatorial background determined by the statistical accuracy of the measured like-sign yield (see Table IV). The uncertainty on the shape of combinatorial background is everywhere negligible compared to the uncertainty on the normalization.

To evaluate the systematic uncertainty due to the correlated pairs we compare the results obtained with the two subtraction methods. It is important to note that the contribution of the correlated background is everywhere small compared to the signal. Therefore any uncertainty in its estimate would result in a small uncertainty on the signal.

In the first subtraction method the correlated background $B^+$ was calculated using Monte Carlo simulations. We have estimated potential spectral shape modifications of $B^+$ in several ways, shown in Fig. 15. In Au + Au

- we have not subtracted any jet background (this would clearly result in an upper limit);
- we have subtracted the near side normalizing it for $m_{ee} > 0.55 \text{ GeV}/c^2$;

  - we have subtracted both near and away-side using for the away-side the same normalization as for the near-side (this would clearly result in a lower limit);
  - we have subtracted the near side normalizing it for $70^\circ < \Delta \phi < 110^\circ$;
  - we have subtracted the near side normalizing it for $70^\circ < \Delta \phi < 110^\circ$ and we have corrected the shape for the eID efficiency.

In the second subtraction method, both in $p + p$ and Au + Au we use the like-sign corrected for the different acceptance, or a hybrid method which uses the mixed-event distribution for the combinatorial background and the like-sign distribution corrected for the different acceptance for the correlated background only.

Since the acceptance for pairs is a function of mass and $p_T$, we have checked that for different $e^+e^-$ pair sources, which span reasonable variations in mass and $p_T$ shapes of the $e^+e^-$ pairs, the relative acceptance is unchanged. For this purpose in $p + p$ we have calculated the relative acceptance:

- using data from $p + p$ run only;
- using data from Au + Au run only;
- using data from PYTHIA real events;

In Au + Au

- we have subtracted near and away jet;
- we have subtracted like-sign (HYBRID);
- we have subtracted like-sign (HYBRID) subtract near (norm: $m_{ee} > 0.55 \text{ GeV}/c^2$);
- we have subtracted near and away jet subtract like-sign;
- we have subtracted near jet subtraction;
- we have subtracted near and away jet subtract like-sign.

FIG. 15: (Color online) Unlike-sign dielectron spectra obtained by subtracting the background (combinatorial and correlated) with the methods explained in the text for $p + p$ (left) and Au + Au (right). The RMS/MEAN of all the spectra shown in the bottom panels allows to assign a systematic uncertainty on the signal due to correlated background subtraction.
• using data from PYTHIA mixed events.

We added the results from these two methods in Fig. 13 as well. The agreement within the different methods (see RMS/MEAN in the bottom panel) allows to assign a systematic uncertainty on the signal.

E. Raw Mass Spectrum

Figure 14 shows the mass distribution of $e^+e^-$ pairs, the normalized mixed-event background ($B$), and the signal yield ($S$) obtained by subtracting the mixed-event background and the correlated background (cross and jet pairs) for Min. Bias $Au + Au$ collisions. The right panel shows the signal-to-background ratio ($S/B_{\text{comb}}$). The systematic uncertainties (boxes) reflect the uncertainty from the background subtraction, which is given by $\delta S/S = 0.25\% \cdot B_{\text{comb}}/S$, added in quadrature to the uncertainty due to the correlated background subtraction, conservatively assumed to be around $10\%S$ below 0.55 GeV/$c^2$. Despite the small $S/B_{\text{comb}}$ ratio, an $e^+e^-$ pair continuum is visible up to 4.5 GeV/$c^2$. Due to the limited statistical precision in Au + Au binning was chosen such that the bins near the $\phi$ and $\omega$ meson are centered at the nominal meson mass. The bin width was chosen to correspond approximately to twice the mass resolution observed in $p + p$ collisions. In this way the $\phi$ and $\omega$ mesons can be resolved from the continuum.

F. Runs with Increased Conversion Material

In order to check the background subtraction, a subset of data ($5 \times 10^7$ events) was collected with a brass sheet of 1.68% radiation length ($X_0$) wrapped around the beam pipe to increase the number of photon conversions. We make an estimate on the uncertainty in the radiation length ($5\%$) by comparing photon conversions in data and simulation. The additional conversion leads to an increased electron multiplicity, in this data set the combinatorial and correlated background ($B$) contribution is larger by a factor of $\sim 2.5$.

If there is a systematic bias in background normalization, and the yield of background is off by a small fraction $f$, it can lead to significant difference between the “apparent” signal and the “true” signal $S_{\text{true}}$. The difference should be larger in the converter run since signal-to-background is smaller. Thus the “apparent” signal $S_C$ in the converter run should become larger than the “apparent” signal $S_{NC}$ in normal run without the converter. The relation between $S_{\text{true}}$, $S_C$, $S_{NC}$ can be written as:

$$S_{NC} = N_{NC} - (1 - f) \cdot B = S_{\text{true}} + fB$$

$$S_C = N_C - 2.5 \cdot (1 - f) \cdot B = S_{\text{true}} + 2.5fB$$

where $N_C$ ($N_{NC}$) is the foreground of all $e^+e^-$ pairs in converter (non-converter) runs. If we divide Equation 18 by $B$, we obtain:

$$\frac{S}{B} = \frac{S_{NC}}{B} = \frac{S_{\text{true}} + fB}{B}$$

Then the ratio between the apparent signal in converter and non-converter runs is given by:

$$\frac{S_C}{S_{NC}} = \frac{S_{\text{true}} + 2.5fB}{S_{\text{true}} + fB} = \frac{S_{\text{true}} + fB + 1.5fB}{S_{\text{true}} + fB} = \frac{(S/B)B + 1.5fB}{S/B}$$

$$= 1 + 1.5f \frac{B}{S}$$

Therefore

$$f = \left(\frac{S_C}{S_{NC}} - 1\right) \frac{S}{B} \frac{1}{1.5}$$

This means we can use the agreement between converter and non-converter runs $S_C/S_{NC}$, shown in Fig. 14, and the signal-to-background ratio $S/B$ also shown in the insert of Fig. 14 to constrain a potential bias $f$ in the background normalization.

In $0.3 < m_{ee} < 0.75$ MeV/$c^2$

$$\frac{S_C}{S_{NC}} = 1.29 \pm 0.92$$

and

$$\frac{S}{B} = (7.47 \pm 0.55) \times 10^{-3}$$

constrain $f = (0.14 \pm 0.46) \times 10^{-2}$. If we extend the mass range to lower values $0.04 < m_{ee} < 0.75$ MeV/$c^2$,

$$\frac{S_C}{S_{NC}} = 1.08 \pm 0.28$$

and

$$\frac{S}{B} = (2.75 \pm 0.05) \times 10^{-2},$$

the constraint becomes $f = (0.15 \pm 0.51) \times 10^{-2}$.

The agreement between the converter run and the normal run confirms that the systematic bias in the background normalization is small ($0.15 \pm 0.51\%$). The result is consistent with our estimate of the systematic uncertainty in the background normalization ($0.25\%$). It should be noted that the converter run provides an independent test on the background normalization.
The occupancy efficiency ($\epsilon$) (Section III A).

The Au + Au centrality class and is described in Section III J.

The total pair reconstruction efficiency $\epsilon_{\text{total}}$ to give the $e^+e^-$ pair yield inside the PHENIX aperture as defined by Equation 23:

$$\frac{dN}{dm_{ee}} = \frac{1}{N_{\text{evt}}} \frac{N_{ee}}{\Delta m_{ee}} \frac{1}{\epsilon_{\text{pair}}^{\text{total}}}$$

The $p_T$ spectra are further corrected for the pair geometric acceptance ($\epsilon_{\text{geo}}^{\text{pair}}$) to give the $e^+e^-$ yield in one unit of rapidity:

$$\frac{1}{2\pi p_T^{ee}} \frac{d^2N}{dp_T^{ee} dy_{ee}} = \frac{1}{2\pi p_T^{ee}} \frac{N_{ee}}{N_{\text{evt}}} \frac{\Delta p_T^{ee}}{\Delta y} \frac{1}{\epsilon_{\text{pair}}^{\text{total}}} \frac{1}{\epsilon_{\text{geo}}^{\text{pair}}}$$

The $p + p$ data are further corrected by the factor $\epsilon_{\text{BBC}}$, where $\epsilon_{\text{BBC}} = 54.5 \pm 5\%$ is the BBC efficiency and $\epsilon_{\text{bias}} = 79 \pm 2\%$ is the BBC trigger bias, described above (Section III A).

The total pair reconstruction efficiency $\epsilon_{\text{total}}$ depends on the single electron efficiency for reconstruction and eID ($\epsilon_{\text{single}}^{\text{ID}}$), the efficiency from the detector live area ($\epsilon_{\text{live}}^{\text{pair}}$), the combinatorial efficiency ($\epsilon_{\text{comb}}^{\text{pair}}$) for Au + Au. The occupancy efficiency ($\epsilon_{\text{occ}}^{\text{single}}$) is different for each Au + Au centrality class and is described in Section III J. The $p + p$ data are also corrected for the ERT trigger efficiency, described in Section III B. In addition, the effect of the pair cuts ($\epsilon_{\text{pair}}^{\phi}$ and $\epsilon_{\text{pair}}^{\text{ghost}}$) is taken into account:

$$\epsilon_{\text{total}}^{\text{pair}} = \epsilon_{\text{pair}}^{\text{live}} \epsilon_{\text{pair}}^{\phi} \epsilon_{\text{pair}}^{\text{ghost}}$$

These efficiencies depend on the eID cuts used to determine the electron sample, therefore they do factorize only on the condition that the electron sample used to calculate them is the same for all them.

The pair-eID and reconstruction efficiency $\epsilon_{\text{pair}}^{\text{ID}}$ and pair-live efficiency $\epsilon_{\text{pair}}^{\text{live}}$, as well as the ERT efficiency $\epsilon_{\text{pair}}^{\text{ERT}}$ for $p + p$ data are derived, as a function of mass and pair-$p_T$, from the corresponding single electron efficiency ($\epsilon_{\text{single}}^{\text{ID}}$, $\epsilon_{\text{single}}^{\text{live}}$ and $\epsilon_{\text{single}}^{\text{ERT}}$ respectively) using the pair kinematic properties implemented in our hadron decay generator, as explained below.

The $\epsilon_{\text{single}}^{\text{ID}}$ is the fraction of signal loss due to track reconstruction and eID cuts within the detector active area. It depends only on the momentum of the track. The shape is very similar in $p + p$ and Au + Au, but the scale is different since different cut values are used. It is calculated as follows:

$$\epsilon_{\text{single}}^{\text{ID}}(p_T) = \frac{dN_{\text{eout}}^{\text{in}}/dp_T}{dN_{\text{eout}}^{\text{in}}/dp_T}.$$ 

The $dN_{\text{eout}}^{\text{in}}/dp_T$ is the $p_T$ distribution of the input electron yield that falls into the real PHENIX acceptance, which includes the boundary described by Equation 3 and the active areas of the detector. $dN_{\text{eout}}^{\text{in}}/dp_T$ is the
$p_T$ distribution of the output electron yield in the same acceptance after passing all the eID cuts.

The input distribution comes from a simulation of 450M $\pi^0$s flat in phase space ($0 < p_T < 25$ GeV/c, $|y| < 0.5$, and $0 < \phi < 2\pi$) with the branching ratio of the $\pi^0$ Dalitz decay set to 100% to enhance the sample of $e^+e^-$ pairs per event. These events were processed by the full GEANT simulation program of the PHENIX detector [2] that includes the details of the detector response. The output simulation data files were processed by the event reconstruction chain of PHENIX. Standard eID cuts are applied to the output. The reconstructed $p_T$ of each output electron is weighted according to $p_T^{2}$ with a realistic exponential $p_T$ weight.

Figure 17 shows the single electron (positron) efficiency as a function of $p_T$ for $p + p$ and Au + Au. The different scales corresponding to the two data sets are due to more stringent cuts applied in the Au + Au data set. The band around the curve shows only the $p_T$ dependence of the systematic uncertainty corresponding to a shift of $0.1$ GeV/c of the efficiency curve. This in turns leads to a distortion of the $e^+e^-$ mass shape shown by the band in Fig. 18. The total uncertainties on the pair reconstruction, including the range of applicability, are reported in Table VI.

The $\epsilon_{\text{live}}$ single is the fraction of signal loss due to inactive areas of the detector. The active areas of the detector are parameterized as a function of the particle momenta and azimuth using real data. There are small differences between $p + p$ and Au + Au.

In the EXODUS cocktail of hadron decays we have implemented the parameterization of the single electron (positron) efficiency (shown as a curve in Fig. 17). The single electron (positron) efficiency was applied as a weight to each track. The pair will therefore get a weight given by the product of the electron and the positron weight. This weight, a function of mass and pair-$p_T$, represents $\epsilon_{\text{ID}}^{\text{pair}}$. To calculate $\epsilon_{\text{live}}^{\text{pair}}$ a fiducial cut corresponding to the detector active areas is also implemented in EXODUS and a pair is rejected if at least one track falls out of the active areas. We therefore determine the product $\epsilon_{\text{ID}}^{\text{pair}} \times \epsilon_{\text{live}}^{\text{pair}}$ double differentially as a function of mass and pair-$p_T$.

Figure 18 shows $\epsilon_{\text{ID}}^{\text{pair}} \times \epsilon_{\text{live}}^{\text{pair}}$ as a function of mass. At high masses the efficiency is constant with a value nearly equal to the square of the single electron efficiency. At low mass the pair efficiency results from the convolution of the single electron reconstruction efficiency, which drops toward low momentum, and the geometric acceptance, which effectively truncates the single electron $p_T$ distribution ($p_T > 0.2$ GeV/c). For pairs with $0.4 < m_{ee} \leq 0.8$ GeV/$c^2$ the efficiency drops as a consequence of the drop at low-$p_T$ of the single electron efficiency. However, for $m_{ee} \leq 0.4$ GeV/$c^2$, the lower limit on single electron $p_T$ results in a larger average momentum; and the pair-efficiency consequently increases. The band around the $\epsilon_{\text{ID}}^{\text{pair}} \times \epsilon_{\text{live}}^{\text{pair}}$ curve in Fig. 18 again shows only the possible distortion of the mass distribution, due to the uncertainty shown in Fig. 17. It reaches a maximum of 8% around $m_{ee} = 0.4$ GeV/$c^2$.

The $\epsilon_{\text{ghost}}^{\text{pair}}$ represents the loss of real pairs which are accidently fulfill the detector overlap criteria. It is determined by the corresponding loss in mixed events as:

$$\epsilon_{\text{ghost}}^{\text{pair}} = \frac{d\text{B1}_{\text{pair}}}{\text{dB}_{\text{pair}}}(m_{ee}dp_T)$$

where $\text{dB1}$ (dB2) are the mixed-event unlike-sign pair (mass, $p_T$) distribution with (without) applying the overlap pair cuts.

The $\epsilon_{\text{IV}}^{\text{pair}}$ represents the loss of real pairs which are accidentally oriented like conversion pairs in the magnetic field. At low masses it is calculated using the $\pi^0$ GEANT simulations described above. Above the pion mass, we use our hadron decay generator EXODUS where we have implemented an empirical smearing for the detector resolution in the determination of the magnitude and direction of the momentum vector.

Figure 19 shows the total pair-efficiency as a function of invariant mass for different ranges of pair-$p_T$ for Au + Au collisions.

The systematic uncertainty on $\epsilon_{\text{ID}}^{\text{pair}} \times \epsilon_{\text{live}}^{\text{pair}}$ is given by twice the uncertainty on the single-particle efficiency, determined by varying the cut values. In addition, uncertainties in the active areas have been determined by varying the active areas in the simulations.

The systematic uncertainties depend somewhat on $p_T$ and mass, they are largest at low $p_T$ and low mass where they reach 14.4% for $p + p$ and 13.4% for Au + Au. We use these errors as an estimate for all momenta and masses. Since the pair-efficiency was derived using realistic kinematics, the uncertainty arises mostly from low-$p_T$ tracks and neglecting the mass and pair-$p_T$ dependence is a conservative approach.

The $\epsilon_{\text{ghost}}^{\text{pair}}$ and $\epsilon_{\text{IV}}^{\text{pair}}$ affect only the LMR. Therefore, the corresponding systematic uncertainties have been independently evaluated by varying the cut values. We estimate an uncertainty of 5% due to the ghost cut and 6% due to the photon conversion cut on the final $e^+e^-$ yield and we apply it to the LMR by varying the cut value in a suitable range.

The (mass, $p_T$) distributions from every individual centrality bin has been corrected using the same 2D efficiency correction function calculated for Min. Bias. This procedure avoids an assumption for the input kinematic distribution. However, since the individual centrality bins have limited statistics compared to the Min. Bias sample, this procedure may suffer statistical fluctuations. Alternatively we use an effective efficiency correction, as a function of mass only, obtained by weighting the 2-dimensional (mass, $p_T$) corrections with a realistic $p_T$ distribution provided by the minimum bias data set. This curve is given by the product of all the curves shown in Fig. 18. We take the difference of 10% between the 1D and the 2D corrected result as an additional systematic uncertainty for the efficiency correction of the different centrality data sets (see Fig. 25).
FIG. 17: (Color online) Single electron (positron) efficiency as a function of $p_T$ for electrons (left) and positrons (right) in $p+p$ (top) and Au + Au (bottom) collisions.

We verified this efficiency using a GEANT simulation of 1 million $e^+e^-$ pairs. Roughly half were generated flat in mass (0–4 GeV/$c^2$), $p_T$ (0–4 GeV/$c$), azimuthal angle (0–2π), and rapidity ($|y| < 0.5$). The other half were generated with a probability inversely proportional to $p_T$ in order to enhance the statistics in the low-mass and low-$p_T$ region, where the efficiency varies most. Only pairs with both an electron and a positron in the ideal acceptance (given by Equation 3) are processed by GEANT and reconstructed with the same analysis chain. The
efficiency is determined double differentially in $p_T$ and mass of the $e^+e^-$ pair. This second method gives consistent results, but is limited by MC statistics. It was used as cross-check in the final analysis.

H. Acceptance Correction

In addition to the efficiency corrections, the $p_T$ spectra are also corrected for the detector geometric acceptance $\epsilon_{\text{geo}}$ to give the $e^+e^-$-pair yield over the full azimuth in one unit of rapidity. $\epsilon_{\text{pair}}$ accounts for the fraction of pairs produced in one unit of rapidity over the full azimuthal range that are lost because either one or both particles from the pair miss the PHENIX detector

$$\epsilon_{\text{pair}} = \frac{d^2N_{ee}}{dp_T^2 dm_{ee}} \frac{\text{in PHENIX acceptance}}{d^2N_{ee_{\text{geo}}}} \left| y_{ee} \right| < 0.5$$

$\epsilon_{\text{geo}}$ has been calculated using a Monte Carlo simulation of Dalitz decays of pseudoscalar mesons ($n^0, \eta, \eta'$) and direct decays of vector mesons ($\rho, \omega, \phi, J/\psi, \psi'$). For all mesons the rapidity distribution is assumed to be flat around mid-rapidity. This assumption is well justified as PHENIX measures in $|y_{ee}| < 0.35$, where the natural distribution is flat. The acceptance is therefore uniform in $|y_{ee}| < 0.35$ and we do not assign a systematic uncertainty to it. The $p_T$ distributions are taken from PHENIX measurements (see Section V), and meson polarizations are taken from [53]. The acceptance correction is performed double differentially in pair-mass and $p_T$ with 0.005 GeV/$c^2$ bins in mass and 0.1 GeV/$c$ bins in $p_T$.

The systematic uncertainty due to the $p_T$ parameterization and the polarization of the Dalitz pairs is studied with a simulation of unpolarized pairs with a flat distribution in mass and pair-$p_T$. In Fig. 19 we compare the acceptance for the full hadronic cocktail and the simulation of unpolarized pairs as a function of pair-$p_T$ in 0.1 GeV/$c^2$ wide mass bins. The cocktail consists of the sum of polarized Dalitz decays and unpolarized vector mesons, while the flat simulations are always unpolarized. The shape of the acceptance is very similar, and the relative normalization agrees within 5% in the lowest mass bins, and better for higher bins. Based upon this comparison, illustrated by the ratio in the bottom figure, we assign an upper limit of 10% (marked by the lines in the figure) for the systematic uncertainty of $\epsilon_{\text{geo}}$.

Since it arises from the independent fragmentation of two charm quarks, the contribution of charmed meson decays has a different acceptance. This component has been simulated with PYTHIA [77] and normalized according to the cross section measured in [16] scaled by $N_{\text{coll}}$. However, due to the observed modifications of charm quarks in the medium [4], the acceptance could potentially be different than what is simulated by PYTHIA. For $m_{ee} < 0.5$ GeV/$c^2$ the charm contribution is negligible. For $0.5 < m_{ee} < 1$ GeV/$c^2$ a systematic uncertainty of 5% due to the uncertainty of the charm cross section ($\sigma_{cc} = N_{\text{coll}} \times 567 \pm 57^{\text{stat}} \pm 22^{\text{sys}}$ mb [77,49]) has been added in quadrature to the other systematic uncertainties on $\epsilon_{\text{geo}}$.

I. Trigger Efficiency ($p+p$)

The efficiency of the ERT trigger ($\epsilon_{\text{ERT}}$) in $p+p$ collisions, as a function of pair mass and $p_T$, is determined with a fast Monte Carlo simulation of pairs which uses a parameterization of the single electron ERT efficiency $\epsilon_{\text{ERT single}}$. The single electron ERT efficiency is determined using the Min. Bias data set. The online level-1 trigger decision is recorded in the Min. Bias data set even though it is not used to select the event. We require that the trigger tile that fires the ERT trigger is used by the electron candidate selected by the offline analysis. The ratio of triggered electrons relative to all electrons candidates gives the trigger efficiency as function of the single electron $p_T$ and is shown in Fig. 21. The trigger efficiency at the plateau is $\approx 60\%$, consistent with the fraction of the active trigger tiles. The trigger part and the offline part of the RICH and EMCAL read-out are handled in separate electronics chains, and the trigger part has more noisy tiles which are masked out. This results in less active areas in the trigger. The efficiency is fit to the sum of two Fermi functions

$$f(p_T) = \frac{\epsilon_0 \cdot \theta(p_T - 0.5)}{e^{-\beta (p_T - p_0)}} + \frac{\epsilon'_0 \cdot \theta(p_T + 0.5)}{e^{-\beta (p_T + p_0)}} + 1$$

where $\epsilon_0$ ($\epsilon'_0$), $p_0$ ($p'_0$) and $k$ ($k'$) are free fit parameters and $\theta$ is the usual Heaviside theta function.

This parameterization is used in a fast Monte Carlo simulation of hadron decays into $e^+e^-$. Each electron fires the trigger with a probability given by the trigger efficiency. We require that both electrons are within the PHENIX acceptance and at least one of them fires the trigger. The mass and $p_T$ distribution of those pairs is
FIG. 20: (Color online) $e^+e^-$-pair acceptance as a function of pair-$p_T$ for different mass ranges for the hadronic cocktail (line) and a flat distribution of $e^+e^-$ pairs (empty circles) and ratio of the two (bottom figure). The lines mark the limits of the systematic uncertainty.

compared to the one obtained with no trigger requirement. The ratio is the pair trigger efficiency, and is shown in Fig. 22 as a function of $e^+e^-$ invariant mass. The structures in Fig. 22 result from the turn-on curve of the trigger threshold convoluted with the acceptance of the detector.

The systematic uncertainty has been studied by varying the parameterization of the single electron ERT efficiency, as well as varying the eID cuts which define the reference sample. The effect of changing the eID cuts leads to a larger uncertainty at low-$p_T$ and low-masses. We therefore assign a systematic uncertainty of 20% for
$m_T < 1 \text{ GeV}/c^2$ and 5% elsewhere. Finally, the systematic uncertainty includes potential shape distortion to the efficiency correction due to the variation of active trigger tiles during the data collection.

Fig. 23 compares the invariant mass spectra for the $p + p$ data obtained with the Min. Bias and the ERT data sets. The ERT data set has been corrected by ERT trigger efficiency $\epsilon_{\text{ERT}}$ and the total pair reconstruction efficiency $\epsilon_{\text{pair}}$, while the Min. Bias is corrected only for the $\epsilon_{\text{pair}}$. This comparison of the two data sets confirms that the Min. Bias and the ERT agree within their respective statistical errors.

### J. Occupancy Correction ($\text{Au} + \text{Au}$)

In $\text{Au} + \text{Au}$ collisions there is an additional efficiency loss of particle detection due to the presence of other particles nearby. To study this effect single electrons and positrons are simulated through the GEANT simulator of PHENIX and then embedded into data files containing detector hits from real $\text{Au} + \text{Au}$ events. Next, these new files containing the embedded $e^\pm$ are run through the entire reconstruction software. As the particle density reduces the efficiency but does not introduce additional $p_T$ dependence, the occupancy correction can be factored out. Since all the detectors used in the analysis are located after the pair has been opened by the magnetic field, the pair embedding efficiency in each centrality bin is defined as the square of the single electron embedding efficiency

$$\epsilon_{\text{occ}} = \left(\epsilon_{\text{occ}}\right)^2 = \left(\frac{\# \text{ reconstructed } e^\pm \text{ from embedded data}}{\# \text{ reconstructed } e^\pm \text{ from single track data}}\right)^2$$

where a reconstructed particle from embedded data is required to have most of its detector hits associated with hits from the simulated particle.

### Table V: Embedding efficiency for different centrality classes of $\text{Au} + \text{Au}$ collisions.

| Centrality   | $\epsilon_{\text{occ}}^{\text{single}}$ | $\epsilon_{\text{occ}}^{\text{pair}}$ |
|--------------|----------------------------------------|----------------------------------|
| 0-10%        | 0.86                                   | 0.74                             |
| 10-20%       | 0.91                                   | 0.83                             |
| 20-40%       | 0.93                                   | 0.87                             |
| 40-60%       | 0.97                                   | 0.95                             |
| 60-92%       | 0.99                                   | 0.98                             |

Table V displays the embedding efficiencies for the centrality classes used in the analysis. For the minimum bias results we have weighted the occupancy correction by the fraction of pairs in each centrality class. Since most of the yield is concentrated in the most central classes, the resulting pair efficiency is 0.81 instead of the square of the minimum bias value $0.96^2 = 0.92$. 

FIG. 21: (Color online) Trigger efficiency $\epsilon_{\text{ERT}}$ for single electrons as a function of $p_T$ of the ERT trigger in $p + p$ collisions determined from the Min. Bias data set.

FIG. 22: (Color online) Trigger efficiency $\epsilon_{\text{ERT}}$ for $e^+e^-$-pairs as a function of $e^+e^-$-pair invariant mass.

FIG. 23: (Color online) Invariant mass spectra for $p + p$ data with the Min. Bias (hollow) and ERT (solid) data sets. The agreement between the two data sets is excellent.
A second, data-driven method was developed to determine \( e_{\text{pair}} \). This method uses the conversion pairs created at the beam-pipe to select a track sample of pure \( e^+ \) or \( e^- \) from real data. The conversion pairs originating at the beam-pipe result in a clear invariant mass peak around \( m_{ee} \sim 20 \text{ MeV}/c^2 \) (see Fig. [3]). We assign tight eID cuts to only one track of a pair and measure the efficiency of the other track of the same pair without any eID cut applied. Differences between the two methods are accounted for in a 3% systematic uncertainty [6].

### K. Systematic Uncertainty

The systematic uncertainties are summarized in Table [VII]. The uncertainties in Table [VII] are categorized by type:

- **A**: point-to-point uncertainty uncorrelated between mass or \( p_T \) bins,
- **B**: correlated uncertainty, all points move in the same direction but not by the same factor,
- **C**: an overall normalization uncertainty in which all points move by the same factor independent of mass or \( p_T \).

The uncertainty on pair reconstruction efficiency includes eID cuts, geometric acceptance, and run-by-run fluctuations. The uncertainty due to the conversion rejection and the geometric acceptance, we added \( 10\% \) uncertainty arising from the trigger bias is included. For individual centrality bins, we use the \( \text{ERT} \) trigger efficiency, as well as on the BBC and the occupancy in \( \text{Au + Au} \) are normalization uncertainties in which all points move by the same factor independent of mass and \( p_T \). Since those uncertainties are small compared to the total uncertainty, they are included in the total uncertainty, without plotting it separately.

### IV. COCKTAIL OF HADRONIC SOURCES

In this Section we describe the methods used to calculate the pair yield expected from hadronic decays which will be compared to the experimental data.

We model the \( e^+e^- \) pair contributions from hadron decays using the decay generator EXODUS. EXODUS is a phenomenological event generator that allows to simulate the phase-space distribution of all relevant sources of electrons and electron pairs and the decay of these sources. Also it allows to include the filtering for the geometrical acceptance and the detector resolution. The relevant primary mesons that involve electrons in the final state are \( \eta^0, \eta, \eta', p, \omega, \phi, J/\psi \) and \( \psi' \).

We assume that all hadrons have a constant rapidity density in the range \( |\Delta \eta| \leq 0.35 \) and a uniform distribution in azimuthal angle. Transverse momentum distributions are largely based on measurements in PHENIX. The key input is the rapidity density \( dN/dy \) of neutral pions, which we determine from a fit to PHENIX data on charged and neutral pions [54–56, 60, 61] with a modified Hagedorn function, given by:

\[
E \frac{d^3 \sigma}{dp^3} = A(e^{-(ap_T + bp_T^2)} + p_T/p_0)^{-n} \tag{30}
\]

Fit parameters and \( dN/dy \) for \( p + p \) and \( \text{Au + Au} \) are given in Table [VII].

For all other mesons we assume \( m_T \) scaling, replacing \( p_T \) by \( \sqrt{m_T^2 - m^2} + (p_T/c) \), where \( m \) is the mass of the meson and we fit a normalization factor to PHENIX data [54, 55, 60, 61] and Au + Au data. For the \( \rho \) meson we assume \( \sigma_{\rho}/\sigma_\omega = 1.15 \pm 0.15 \), consistent with values found in jet fragmentation [53]. The \( \eta' \) yield is scaled to be consistent with jet fragmentation \( \sigma_{\eta'}/\sigma_\eta = 0.15 \pm 0.15 \) [53]. The \( \psi' \) is adjusted to be \( \sigma_{\psi'}/\sigma_{J/\psi} = 0.14 \pm 0.03 \) in agreement with PHENIX measurements [57].

For the \( \eta, \omega, \phi \), and \( J/\psi \), and also \( \eta' \), and \( \psi' \) (in \( p + p \)) the quoted uncertainties include those on the data as well.
TABLE VI: Systematic uncertainties on the dilepton yield and mass range of applicability.

| Syst. Err. component                  | $p + p$ | Au + Au | Mass Range | Type |
|---------------------------------------|---------|---------|------------|------|
| pair reconstruction                   | 14.4%   | 13.4%   | 0–4 GeV/c² | B    |
| conversion rejection                  | 6%      | 6%      | 0–0.6 GeV/c² | B    |
| pair cuts                             | 5%      | 5%      | 0.4–0.6 GeV/c² | B    |
| occupancy efficiency                  | -       | -       | 0–4 GeV/c² | C    |
| BBC and trigger bias                  | 11.3%   | -       | 0–4 GeV/c² | C    |
| ERT efficiency                        | 5% (20% for $m_T < 1$ GeV/c²) | -       | 0–4 GeV/c² | B    |
| combinatorial background              | 3%·B/S  | 0.25%·B/S | 0–4 GeV/c² | B    |
| correlated background                 | mass-dependent (Fig. 15) | mass-dependent (Fig. 15) | 0–4 GeV/c² | B    |
| centrality                            | -       | -       | 0–4 GeV/c² | B    |
| acceptance correction                 | 10%     | 10%     | 0–4 GeV/c² | B    |
| charm acceptance                      | 5%      | 5%      | >0.5 GeV/c² | B    |

FIG. 24: (Color online) Compilation of meson production cross sections in $p + p$ (left) and Au + Au (right) collisions at $\sqrt{s_{NN}} = 200$ GeV. Shown for $p + p$ are data for neutral $[54]$, charged pions $[55]$, kaons $[56]$, $\eta$ $[57]$, $\omega$ $[58]$, $\phi$ $[59]$, and $J/\psi$ $[60]$. Shown for Au + Au are data for neutral $[61]$, charged pions $[62]$, $\eta$ $[63]$, $\omega$ $[64]$, $\phi$ $[65]$, and $J/\psi$ $[66]$. The data are compared to the parameterization based on $m_T$ scaling used in our hadron decay generator.

as those using different shapes of the $p_T$ distributions to extrapolate to zero $p_T$. Specifically we have fit the functional form given in Equation [50] with all parameters free and also with an exponential distribution in $m_T$. For the $\rho$, which is not measured in $p + p$ nor in Au + Au, and $\eta'$ and $\psi'$, which are not measured in Au + Au, the uncertainty in the table represent the quadrature sum of the uncertainty of the cross section and the uncertainty relative to other mesons.

All the mesons shown in Fig. 24 can be described by the $m_T$-scaling parameterization of the pion spectrum. The fact that the $\eta$'s and the kaons follow the same $m_T$-scaling prediction over all $p_T$ appears to be due to the fact that the masses of the particles are almost the same. In Au + Au however $\eta$ and kaons do not follow the same $m_T$-scaling prediction. At high $p_T$, where we measure $\eta$'s, we see that they are suppressed as much as pions and the trend of $R_{AA}$ for these two mesons looks identical. However we have observed a different trend (i.e. a smaller suppression) for strange particles and $\eta$ has a strangeness content too $[67, 68]$. Therefore, since the $\eta$ cannot be measured at low-$p_T$, we take as systematic uncertainty in the low-$p_T$ region (and consequently on the extrapolated $dN/dy$), the difference of the two spec-
TABLE VII: Fit parameters from the modified Hagedorn function (Equation 28) for \( p + p \) and \( \text{Au} + \text{Au} \) pion spectra (\( \pi^+ \) and \( \pi^- \)) and the corresponding rapidity density \( dN/dy \).

| Parameter | \( p + p \) | \( \text{Au} + \text{Au} \) |
|-----------|-------------|-----------------|
| \( dN/dy \) | 1.06±0.11 | 95.7±6.9 |
| \( A[\text{mb} \text{GeV}^{-2} e^2] \) | 377±60 | 504.5±10 |
| \( a[\text{GeV}/e]^{-1} \) | 0.356±0.014 | 0.52±0.007 |
| \( b[\text{GeV}/e]^{-2} \) | 0.068±0.019 | 0.16±0.010 |
| \( p_0[\text{GeV}/e] \) | 0.7±0.02 | 0.7±0.005 |
| \( n \) | 8.25±0.04 | 8.27±0.02 |

tra (\( \sim 30\% \)) in the low-\( p_T \) range. We note that this is a conservative estimate of the systematic uncertainty. Statistical models which reproduce well the particle spectra and ratios measured at RHIC [30, 31] calculate a \( dN/dy \) for the \( \eta \) which is well within the assigned systematic uncertainty. At high-\( p_T \) we assign a smaller systematic uncertainty of 7\% (17\%) for \( p + p \) (\( \text{Au} + \text{Au} \)) collisions arising from the asymptotic value of the \( \eta/\pi^0 \) ratio of 0.48 ± 0.03 (0.08) based on PHENIX measurement [5].

Once the meson yields and \( p_T \) spectra are known the dilepton spectrum is given by decay kinematics and branching ratios, which are implemented in our decay generator EXODUS following earlier work published in [10]. The branching ratios are taken from the compilation of particle properties in [32]. For the Dalitz decays \( \pi^0, \eta, \eta' \rightarrow e^+e^-\gamma \) and the decay \( \omega \rightarrow e^+e^-\pi^0 \) we use the Kroll-Wada expression [4] with electromagnetic transition form factors measured by the Lepton-G collaboration [73, 74]. For the decays of the vector mesons \( \rho, \omega, \phi \rightarrow e^+e^- \) we use the expression derived by Gounaris and Sakurai [72], extending it to 2 GeV/\( e^2 \), slightly beyond its validity range. For the \( J/\psi \) and \( \psi' \rightarrow e^+e^- \) we use the same expression discussed in [32] modified to include radiative corrections. All vector mesons are assumed to be unpolarized. For the Dalitz decays, where the third body is a photon, the angular distribution is sampled according to 1 + cos²\( \theta_{CS} \), where \( \theta_{CS} \) is the polar angle of the electrons in the Collins-Soper frame.

The resulting systematic uncertainties on the mass spectrum depend on mass, and range from 10 to 30\%. They result primarily from the uncertainty on the measured pion yield and on the meson-to-pion ratios. The uncertainty from the measured electromagnetic transition form factors, in particular for the \( \omega \rightarrow e^+e^-\pi^0 \) decay, is also included but contributes significantly only in the range around 0.5 to 0.6 GeV/\( e^2 \). The uncertainty from polarization is negligible but also included.

V. RESULTS

This Section presents the results for the \( p + p \) and the \( \text{Au} + \text{Au} \) analyses. The \( p + p \) data provide a good baseline for understanding the results of the \( \text{Au} + \text{Au} \) analy-
TABLE VIII: Hadron rapidity densities used in our hadron decay generator. For the \( \omega \) and \( \phi \), data from this analysis were used together with data from the quoted references.

| \( \frac{dN}{dy} \) at \( y=0 \) | relative uncertainty | \( \text{meson}/\pi^0 \) | data used |
|------------------------|---------------------|----------------|------------|
| \( \pi^0 \)           | 1.065 \( \pm 0.11 \) | 10\%           | 1.0        | PHENIX [54], [55] |
| \( \eta \)            | \( (1.1 \pm 0.3) \times 10^{-1} \) | 30\%           | 1.032 \( \times 10^{-1} \) | jet fragmentation [53] |
| \( \rho \)            | \( (8.9 \pm 2.5) \times 10^{-2} \) | 28\%           | 8.34 \( \times 10^{-2} \) | jet fragmentation [53] |
| \( \omega \)          | \( (7.8 \pm 1.8) \times 10^{-2} \) | 23\%           | 7.32 \( \times 10^{-2} \) | PHENIX [57] |
| \( \phi \)            | \( (9.0 \pm 2.0) \times 10^{-3} \) | 24\%           | 8.4 \( \times 10^{-3} \) | PHENIX [58] |
| \( \eta' \)           | \( (1.3 \pm 0.5) \times 10^{-2} \) | 40\%           | 1.27 \( \times 10^{-2} \) | PHENIX [54] |
| \( J/\psi \)          | \( (1.77 \pm 0.27) \times 10^{-5} \) | 15\%           | 1.66 \( \times 10^{-5} \) | PHENIX [59] |
| \( \psi' \)           | \( (2.5 \pm 0.7) \times 10^{-6} \) | 27\%           | 2.3 \( \times 10^{-6} \) | PHENIX [58], [59] |

\( \text{Au + Au} \)

| \( \pi^0 \)           | \( (9.572 \pm 0.95) \times 10 \) | 10\%           | 1.0        | PHENIX [60], [61] |
| \( \eta \)            | \( (1.077 \pm 0.32) \times 10 \) | 30\%           | 1.12 \( \times 10^{-1} \) | PHENIX [60] |
| \( \rho \)            | 8.60 \( \pm 2.8 \) | 33\%           | 8.98 \( \times 10^{-2} \) | jet fragmentation [53] |
| \( \omega \)          | 9.88 \( \pm 3.0 \) | 30\%           | 1.03 \( \times 10^{-1} \) | PHENIX [58] |
| \( \phi \)            | 2.05 \( \pm 0.6 \) | 30\%           | 2.14 \( \times 10^{-2} \) | PHENIX [58] |
| \( \eta' \)           | 2.05 \( \pm 0.2 \) | 100\%          | 2.15 \( \times 10^{-2} \) | PHENIX [58], [59] |
| \( J/\psi \)          | \( (1.79 \pm 0.26) \times 10^{-3} \) | 15\%           | 1.82 \( \times 10^{-5} \) | PHENIX [59] |
| \( \psi' \)           | \( (2.6 \pm 0.7) \times 10^{-4} \) | 27\%           | 2.70 \( \times 10^{-6} \) | PHENIX [58], [59] |

\( \sigma_{c\bar{c}} = 567 \pm 57^{\text{stat}} \pm 224^{\text{sys}} \mu \text{b} \) [46] has been scaled by \( N_{\text{coll}} \) (given in Table I). For each centrality class, the data and the cocktail are absolutely normalized. Each data set is compared with two corresponding cocktail lines, shown in solid and dotted curves. The difference between the cocktails is due to uncertainty in the \( c\bar{c} \) contribution (see discussion below).

Unlike the \( p + p \) mass spectrum, the \( \text{Au + Au} \) mass spectra show enhancement above the cocktail, in particular for the LMR (0.15 – 0.75 GeV/\( c^2 \)). There is little enhancement for peripheral (60-92\%) data, but very strong enhancement for two most central classes (10-10\% and 10-20\%). The enhancement increases rapidly with increasing centrality.

In order to quantitatively describe this enhancement, more information is needed about other components that can potentially contribute to the LMR, namely the open heavy flavor and internal conversion of real direct photons. We discuss them in the next sections.

### B. Open Heavy Flavor Contribution

The dilepton yield in the IMR is dominated by semi-leptonic decays of charm hadrons correlated through flavor conservation. Small contributions also arise from bottom hadrons and Drell-Yan. For \( p + p \) data we determine the heavy flavor contribution by subtracting the hadronic cocktail from the dilepton data. We integrate the subtracted yield in the IMR, extrapolate to zero \( e^+ e^- \) pair mass to get the entire cross section, correct for geometric acceptance, and convert to a production cross section using known branching ratios of semi-leptonic decays [53]. Details of the analysis of the charm cross section are reported in [37].

We find a rapidity density of \( c\bar{c} \) pairs at mid rapidity:

\[
\frac{d\sigma_{c\bar{c}}}{dy} \bigg|_{y=0} = 118.1 \pm 8.4^{\text{stat}} \pm 30.7 \text{ syst} \pm 39.5^{\text{model}} \mu \text{b}.
\]

This corresponds to a total charm cross section of \( \sigma_{c\bar{c}} = 544 \pm 39^{\text{stat}} \pm 142^{\text{sys}} \pm 200^{\text{model}} \mu \text{b} \), consistent with previous measurement of single electrons by PHENIX \( \sigma_{c\bar{c}} = 567 \pm 57^{\text{stat}} \pm 224^{\text{sys}} \mu \text{b} \) [46] and with a fixed-order-plus-next-to-leading-log (FONLL) pQCD calculation \( \sigma_{c\bar{c}} = 256^{+490}_{-146} \mu \text{b} \) [71].

In \( \text{Au + Au} \) the dynamic correlation of \( c \) and \( \bar{c} \), which is essential to determine the mass spectral shape, could be modified compared to \( p + p \) collisions. The observed suppression and the elliptic flow of non-photonic electrons indicates that charm quarks interact with the medium [3], which should change the correlations between the produced \( c\bar{c} \) pairs. We also note that the \( pp \) distribution for electrons generated by PYTHIA [7] is softer than the spectrum measured in \( p + p \) data but coincides with that observed in minimum bias \( \text{Au + Au} \) collisions. Thus we compare our \( \text{Au + Au} \) data to two extreme scenarios that bracket the charm contribution:

(i) The correlation is unchanged by the medium and
the back-to-back correlation predicted by uncorrelated pairs is smaller than the one resulting from mental data [6]. Because the average opening angle of and keep the overall cross section fixed to the experi-

We sample from the heavy flavor pairs in the PHENIX acceptance in minimum-bias Au + Au collisions (as given in Table I).

(i) The $c\bar{c}$ dynamical correlation is washed out by medium interactions, i.e., the direction of $c$ and $\bar{c}$ quarks are uncorrelated. We sample from the heavy flavor single-electron $p_T$ spectra, choose the angle randomly and keep the overall cross section fixed to the experimental data [1]. Because the average opening angle of uncorrelated pairs is smaller than the one resulting from the back-to-back correlation predicted by PYTHIA, the mass spectral shape of uncorrelated pairs is much softer than the one calculated by PYTHIA.

The charm contribution determined by case (i) is shown as the upper dashed curve in Fig. 25 or the upper solid curves in Fig. 27. The charm contribution determined by case (ii) is shown as the dotted curve in Fig. 26 or the dashed curves in Fig. 23. In both cases the total yield of charm is normalized to the value measured by PHENIX.

In the Min. Bias Au + Au data set, the IMR seems to be well described by the continuum calculation based on case (i). This is somewhat surprising, since single electron distributions from charm show substantial medium modifications [3]. Thus, it is hard to understand how the dynamical correlation at production of the $c\bar{c}$ remains unaffected by the medium. Case (ii) leads to a much softer mass spectrum, as shown by the dotted curve in Fig. 26. This would leave significant room for other contributions, e.g., thermal radiation.

We have integrated the yield in the mass region 1.2 to 2.8 GeV/$c^2$ and normalized to the number of binary collisions \(N_{\text{coll}}\) (Fig. 28). The systematic uncertainty due to \(N_{\text{coll}}\) as indicated in Table 1 [1] has been included in the overall systematic uncertainty. Within uncertainties \(N_{\text{coll}}\) scaling is observed for the production of non-photic electrons, i.e., for those electrons arising from decays of heavy-flavor hadrons [9]. The normalized yield shows no significant centrality dependence and is consistent within systematic uncertainties with the expectation based on \(N_{\text{coll}}\)-scaled PYTHIA, with the cross section measurement of [10] (case (i)). However the scaling with \(N_{\text{coll}}\) may be a mere coincidence resulting from two balancing effects: the energy loss of charm, which increases with \(N_{\text{part}}\), would lead to a softer mass distribution and therefore less yield in the IMR (case (ii)), while a ther-
C. Low-Mass Excess in Au + Au Data

Figure 27 shows that the low-mass enhancement is concentrated in the first two centrality classes, i.e. 0-10% and 10-20%. For more peripheral collisions the enhancement diminishes. Only some small excess is visible for 20-40% and 60-92% while no deviation is observed in 40-60% with respect to the cocktail beyond systematic uncertainties.

To quantify the centrality dependence of the enhancement, we have integrated the yield in two mass windows: below 0.1 GeV/$c^2$, and 0.15 to 0.75 GeV/$c^2$. Since the cocktail yield in these regions arises mostly from hadrons (more than 90% from $\pi^0$ below 0.1 GeV/$c^2$ and more than 99% from the sum of light hadrons in $0.15 < m_{ee} < 0.75$ GeV/$c^2$) we compare the measured yield to the rate of pion production. Pions were found to scale approximately with $N_{\text{part}}$ [78], therefore we com-
pare the measured yield in data to $N_{\text{part}}$.

The top panel of Fig. 24 shows the centrality dependence of the yield in the mass region 0.15–0.75 GeV/c² divided by the number of participating nucleon pairs ($N_{\text{part}}/2$). The systematic uncertainty due to $N_{\text{part}}$ (as indicated in Table III) has been included in the overall systematic uncertainty. For comparison the yield below 0.1 GeV/c², which is dominated by low-$p_T$ pion decays, is shown in the lower panel.

For both mass intervals the yield is compared to the yield calculated from the hadron cocktail. Two solid curves on each panel show the upper and lower limit of the expected yield from the cocktail. The deviation includes the uncertainty in the charm contribution discussed in the previous section. In the lower mass range the yield agrees with expectations, and is proportional to the pion yield (bottom panel of Fig. 24). In contrast, in the range from 0.15 to 0.75 GeV/c² the observed yield rises significantly above expectations.

The enhancement factor, defined as the ratio between the measured yield and the expected yield for $m_{\text{ee}} < 0.15$ GeV/c², is $7.6 \pm 0.5 \pm 1.5 \pm 1.5$ for the upper limit of the measured yield. The first error is the statistical error, the second the systematic uncertainty of the data, and the last error is an estimate of the uncertainty in the cocktail, i.e. the expected yield from hadronic sources. For the various centrality bins the enhancement factor is reported in Table IX.

| Centrality | Enhancement (±stat ±syst ±model) |
|------------|----------------------------------|
| 00-10 %    | $7.6 \pm 0.5 \pm 1.5 \pm 1.5$    |
| 10-20 %    | $3.2 \pm 0.4 \pm 0.1 \pm 0.6$    |
| 20-40 %    | $1.4 \pm 1.3 \pm 0.02 \pm 0.3$   |
| 40-60 %    | $0.8 \pm 0.3 \pm 0.03 \pm 0.2$   |
| 60-92 %    | $1.5 \pm 0.3 \pm 0.001 \pm 0.3$  |
| Min. Bias  | $4.7 \pm 0.4 \pm 1.5 \pm 0.9$    |

The increase is qualitatively consistent with the conjecture that an in-medium enhancement of the dielectron continuum yield arises from scattering processes like $\pi\pi$ or $q\bar{q}$ annihilation. In this case the enhancement would scale proportional to $N_{\text{part}}^{-2}$, different from the hadronic cocktail which scales proportional to $N_{\text{part}}$.\[1.000\]

\[\text{D. } p_T \text{ Dependence of the Mass Spectra}\]

Figure 24 compares $e^+e^-$ invariant mass spectra measured in $p+p$ and in Min. Bias Au + Au collisions to the corresponding expectations from the cocktail of hadron decays and open charm, in different ranges of $p_T$. Data and cocktail are absolutely normalized.

The “knee” in the mass spectra beginning at $m_{ee} \approx 0.1$ GeV/c² corresponds to the cut-off of the $\pi^0$ Dalitz decays. For $m_{ee} > m_{\eta}$, the $\eta$ Dalitz decay ($\eta \rightarrow e^+e^-\gamma$) is the dominant hadronic source of $e^+e^-$ pairs, followed by the $\omega$ Dalitz decay ($\omega \rightarrow e^+e^-\pi^0$), with small contributions from other sources such as $\eta'$ and $\phi$. The detector acceptance and resolution effects broaden the low-mass Dalitz peak and smear the “knee” of the $\pi^0$ Dalitz contribution. These detector effects are included in the cocktail calculation.

The $p + p$ data are consistent with expectations from the cocktail over the full mass range in the low $p_T$ bin. In the highest $p_T$ bins however the data are enhanced with respect to the cocktail. The deviations are however small in contrast to the Au + Au data which show a large enhancement in the LMR above $m_{ee}^0$, which is concentrated at low-$p_T$. For $p_T > 1.0$ GeV/c the enhancement becomes smaller than at lower $p_T$ but it is still larger than the one observed in the $p + p$ data.

In the introduction, we classified the low mass region into LMR I (low mass, high $p_T$) and LMR II (low mass, low $p_T$) (see Fig. 3). The behavior of the low mass excess in Au + Au shown in Fig. 32 is different for LMR I and LMR II. In LMR I, the excess have a similar shape to the cocktail and the level of the excess with respect to the cocktail is approximately constant. In this region, a contribution from inelastic conversion of virtual direct photons is expected. In LMR II, the excess increases with increasing mass and decreasing $p_T$.

In the following we first analyze the data in the low mass high $p_T$ region (LMR I), which allows the measurement of direct photons; then we study how much they contribute to the inclusive enhancement, dominated by the yield in the LMR II.

\[\text{E. Measurement of Direct Photons}\]

In LMR I (Fig. 3), where the $p_T$ of the $e^+e^-$ pair is much greater than its mass ($m_{ee} \ll p_T$), the yield of the virtual photons is approximately the same as that of real photons. Therefore, in this quasi-real virtual photon region, the production of real direct photons can be deduced from measurements of $e^+e^-$ pairs. Theoretical details are given in Appendix B. In this Section, we determine the real direct photon cross section for $1 < p_T < 5$ GeV/c from the data shown in Fig. 32.

We use the mass range $0.1 < m_{ee} < 0.3$ GeV/c².

The relation between real photon production and the associated $e^+e^-$ production can be written as (Equation B3)

\[
\frac{d^2N_{ee}}{dm_{ee}dp_T} = \frac{2\alpha}{3\pi} \frac{1}{m_{ee}} L(m_{ee}) S(m_{ee}, p_T) \frac{dN_\gamma}{dp_T}, \tag{31}
\]

\[
L(m_{ee}) = \sqrt{1 - \frac{4m_{ee}^2}{m_{\gamma}^2} \left(1 + \frac{2m_{ee}^2}{m_{\gamma}^2}\right)}. \tag{32}
\]

Here $\alpha$ is the fine structure constant, $m_{ee}$ is the mass
FIG. 30: (Color online) $e^+e^-$ pair invariant mass distributions in $p+p$ (left) and minimum bias Au + Au collisions (right). The $p_T$ ranges are shown in the legend. The solid curves represent the cocktail of hadronic sources (see Section IV) and include contribution from charm calculated by PYTHIA using the cross section from [4] scaled by $N_{\text{coll}}$.

FIG. 31: (Color online) Electron pair mass distribution for Au + Au (Min. Bias) events for $1.0 < p_T < 1.5$ GeV/c. The two-component fit is explained in the text. The fit range is $0.12 < m_{ee} < 0.3$ GeV/c$^2$. The dashed (black) curve at greater $m_{ee}$ shows $f(m_{ee})$ outside of the fit range.

of $e^+e^-$ pair, $m_e$ is the electron mass, and $S(m_{ee},p_T)$ is a process-dependent factor that accounts for differences between real and virtual photon production, such as form factors, phase space, and the spectral function. Equation (33) holds for any process emitting real photons, in particular direct or thermal emission. For high $p_T$ ($p_T \gg m_{ee}$) the process dependence becomes negligible and the factor $S(m_{ee},p_T)$ becomes 1 as $m_{ee} \to 0$ or $m_{ee}/p_T \to 0$. For $m_{ee} \gg m_e$, the factor $L(m_{ee})$ also becomes very close to unity. Thus the relation simplifies to

$$\frac{d^2N_{ee}}{dm_{ee}dp_T} \approx \frac{2\alpha}{3\pi} \frac{1}{m_{ee}} \frac{dN_{\gamma}}{dp_T} \quad (33)$$

Here the mass distribution of electron pairs for a given $p_T$ bin takes on a very characteristic $1/m_{ee}$ shape. If there is real direct photon production in a given $p_T$ bin, there should be a corresponding electron pair contribution that behaves like $1/m_{ee}$ in the same $p_T$ bin. Therefore, the real direct photon production can be determined from the yield of the excess electron pairs.

For Dalitz decays, the $1/m_{ee}$ behavior is truncated by the kinematic limit and $S(m_{ee})$ becomes zero for $m_{ee} > m_h$, where $m_h$ is the mass of the hadron. The functional form of $S(m_{ee})$ for Dalitz decays is given in Appendix B. In contrast, the factor $S(m_{ee},q)$ for the direct photon process is unity for $p_T \gg m_{ee}$. We exploit this difference to separate the direct photon signal from the hadronic background. Since 80% of the hadronic photons are from $\pi^0$ Dalitz decays, the signal to background (S/B) ratio for the direct photon signal improves by a factor of five for $m_{ee} > m_{\pi^0} \approx 0.135$ GeV/c$^2$, thereby allowing a real direct photon signal that is 10% of the yield of hadronic decay photons to be observed as a 50% excess of $e^+e^-$ pairs for this mass range.

Figure 32 shows a visible excess above the $\pi^0$ cutoff for all $p_T$ bins of the Au + Au data. For $p_T > 1$ GeV/c, the excess is almost a constant factor above the cocktail. As
we examine later the mass distribution for \( p_T > 1 \text{ GeV/c} \) is consistent with the \( 1/m_{ee} \) shape expected for the electron pairs from internal conversion of virtual direct photons.

In the following we assume the excess for \( p_T > 1 \text{ GeV/c} \) and \( m_{ee} < 0.3 \text{ GeV/c}^2 \) is entirely due to internal conversion of virtual direct photons and deduce the real direct photon yield from the \( e^+e^- \) pair yield using Equation \([31] \). We demonstrate the validity of this assumption later. Although the data are consistent with \( 1/m_{ee} \) over a wider mass range \( (m_{ee} \sim 0.7 \text{ GeV/c}^2) \), as can be seen in Fig. \([2] \), we limit our analysis for \( 0.1 < m_{ee} < 0.3 \text{ GeV/c}^2 \). We do so in order (i) to ensure the condition \( m_{ee} \ll p_T \) for the lowest \( p_T \) bin \( (1.0 < p_T < 1.5 \text{ GeV/c}) \), (ii) to keep the correction factor \( S(M,q) \) close to unity, and (iii) to minimize uncertainty due to \( cc \). In this kinematic range, the contribution of \( cc \rightarrow e^+e^- \), estimated by PYTHIA, is less than 5% of the excess.

In order to quantify the excess, we fit a two-component function,

\[
f(m_{ee};r) = (1-r)f_c(m_{ee}) + r f_{\text{dir}}(m_{ee}) \tag{34}
\]

to the mass distribution. Here \( f_c(m_{ee}) \) is the shape of the cocktail mass distribution (shown in Fig. \([30] \)), \( f_{\text{dir}}(m_{ee}) \) is the expected shape of the virtual direct photon internal conversion mass distribution, and \( r \) is the only fit parameter. In the low mass region used for the fit, the functional form of \( f_c(m_{ee}) \) is the sum of Dalitz decay mass distributions of hadrons (Equation \([33] \) – Equation \([34] \)) filtered through the PHENIX acceptance and smeared by the detector effects. It is calculated by a Monte Carlo simulation which takes into account detector effects such as finite mass resolution. The functional form of \( f_{\text{dir}}(m_{ee}) \) corresponds to Equation \([31] \) with \( S(m_{ee}) = 1 \). It is also filtered through the PHENIX acceptance and smeared by detector effects.

Both \( f_c(m_{ee}) \) and \( f_{\text{dir}}(m_{ee}) \) are separately normalized to the data for \( m_{ee} < 30 \text{ MeV/c}^2 \). In this mass region \( S(m_{ee}) \) of \( \pi^0 \) Dalitz decays is very close to unity. Thus the functional shapes of \( f_c \) and \( f_{\text{dir}} \) are essentially identical and equal to \( L(m_{ee})/m_{ee} \) smeared by the detector effects. This means that the fit function \( f(m_{ee};r) \) in this mass range is independent of the fit parameter \( r \) as \( (1-r)L(m_{ee})/m_{ee} + r L(m_{ee})/m_{ee} = L(m_{ee})/m_{ee} \). Thus this normalization ensures that the yield of fit function \( f(m_{ee};r) \) is always normalized to that of the data for \( m_{ee} < 30 \text{ MeV/c}^2 \). The parameter \( r \) can be interpreted as the direct photon fraction of the inclusive photon yield.

This fitting method has the advantage of canceling most of the systematic uncertainties of the cocktail normalization relative to the data. The PHENIX acceptance for electron pairs with \( m_{ee} < 0.3 \text{ GeV/c}^2 \) and \( p_T > 1 \text{ GeV/c} \) is almost constant, and its shape can be calculated accurately as a function of mass. Many systematic effects, such as electron identification efficiency, detector dead area, etc., can influence the absolute value of the acceptance but not its shape.

For each \( p_T \) bin, \( f(m_{ee}) \) is fit to the data for several mass ranges with \( r \) the only fit parameter. Figure \([31] \) shows \( f_{\text{dir}}(m_{ee}) \) and \( f_c(m_{ee}) \) together with the fit result for \( \text{Au + Au Min. Bias data for } 1.0 < p_T < 1.5 \text{ GeV/c} \) and the cocktail components. The dashed curve shows \( f(m_{ee}) \) extended outside of the fit range. Although the mass region is not used in the fit, the fit function describes the data for \( m_{ee} > 0.3 \text{ GeV/c}^2 \).

The fit shown in Fig. \([31] \) has \( \chi^2/\text{NDF} = 12.2/6 \). The somewhat large \( \chi^2 \) values is due to the large contribution from the lowest mass bins, where statistical errors are small and systematic errors due to the detector resolution are significant. The \( \chi^2 \) value is calculated from the statistical errors only. The results for the fit range \( 0.12 < m_{ee} < 0.3 \text{ GeV/c}^2 \) are summarized in Table \(\text{X} \). For \( p_T > 1.5 \text{ GeV/c} \) the fit gives good \( \chi^2/\text{NDF} \), demonstrating that the shape of the excess is consistent with \( 1/m_{ee} \) as expected for internal conversion.

To evaluate the systematic uncertainty due to the mass range used for the fit, the fit was repeated for three mass ranges: \( 0.08 < m_{ee} < 0.3 \text{ GeV/c}^2 \), \( 0.1 < m_{ee} < 0.3 \text{ GeV/c}^2 \), and \( 0.12 < m_{ee} < 0.3 \text{ GeV/c}^2 \). The value of \( r \) is taken as the average of the results for these three fit ranges.

The sources of systematic uncertainty on the fit include (1) the fit range, (2) the mass spectrum of the data, and (3) the cocktail. The sources of the systematic uncertainty on the mass spectrum relative to the cocktail include (1) uncertainties on the correlated background due to jet pairs \( (\sim 2\% \text{ for } p+p) \), (2) uncertainties in the acceptance and efficiency in the mass range of the signal \( (0.1 < m_{ee} < 0.3 \text{ GeV/c}^2) \) relative to the mass range \( m_{ee} < 30 \text{ MeV/c}^2 \) used for the normalization \( (\sim 1\% \text{ for acceptance and } \sim 1\% \text{ for efficiency}) \), and (3) uncertainty in the mixed-event normalization \( (0.25\%/(S/B)) \), where \( S/B \) is the signal-to-background ratio in \( 0.1 < m_{ee} < 0.3 \text{ GeV/c}^2 \). The uncertainty (2) is small since uncertainties in the absolute normalization cancel when the cocktail is normalized to the data in the low-mass peak \( (m_{ee} < 30 \text{ MeV/c}^2) \). The largest source of uncertainty is the particle composition in the hadronic cocktail, namely the \( \eta/\pi^0 \) ratio. This corresponds to a \( \sim 7\% \) \( (\sim 17\%) \) uncertainty in the particle composition in the hadronic cocktail, namely the \( \eta/\pi^0 \) ratio. This corresponds to a \( \sim 7\% \) \( (\sim 17\%) \) uncertainty in the particle composition in the hadronic cocktail, namely the \( \eta/\pi^0 \) ratio.

| \( p_T \) (GeV/c) | \( r \) | \( \chi^2/\text{NDF} \) |
|-----------------|-----|----------------|
| 1.0-1.5         | 0.189 ± 0.021 | 12.2/6 |
| 1.5-2.0         | 0.165 ± 0.022 | 4.6/6  |
| 2.0-2.5         | 0.146 ± 0.029 | 6.6/6  |
| 2.5-3.0         | 0.165 ± 0.040 | 3.3/6  |
| 3.0-4.0         | 0.224 ± 0.048 | 3.7/6  |
| 4.0-5.0         | 0.206 ± 0.093 | 4.2/3  |

TABLE X: Summary of the fits to Equation \([34] \) in the range \( 0.12 < m_{ee} < 0.3 \text{ GeV/c}^2 \).
Since the $\eta/\pi^0$ ratio is the largest source of uncertainty, we also studied fits with a three component function, $f_3(m_{ee}) = (1 - r - r_\eta)f_\eta(m_{ee}) + rf_{dir}(m_{ee}) + r_\eta f_{dir}(m_{ee})$, with a constraint on $r_\eta$ such that $\eta/\pi^0 = 0.48\pm0.03(0.08)$ for $p + p$ (Au + Au) \cite{53}. These alternative fits give consistent results for $r$ within statistical uncertainties.

So far we have assumed that the excess in $p_T > 1$ GeV/c and $0.1 < m_{ee} < 0.3$ GeV/c$^2$ is entirely due to internal conversion of virtual direct photons. This means that we assume $S(m_{ee}, q)$ for excess virtual photons to be unity. In the following, we examine the validity of the assumption from the data.

As shown in Appendix \ref{app:3}, the shape of the virtual photon spectrum as a function of mass can be obtained from the electron pair yield as

$$
\frac{dN^*_\gamma}{dq} = \frac{3\pi}{2\alpha}m_{ee} \times q_0 \frac{dn_{\mu\mu}}{d^3qdm_{ee}}. \quad (35)
$$

Since the shape of $f_{dir}(m_{ee})$ is $1/m_{ee}$ smeared by the detector effects, a fit of $R = (\text{data} - \text{cocktail})/f_{dir}(m_{ee})$ to a constant can be used to test that the excess has the same constant behavior as $S(m_{ee}, q)$ for internal conversion of direct photons. Note that the detector effects in the numerator and denominator cancel in the ratio.

Furthermore, since $f_{dir}(m_{ee})$ is normalized to the data for $m_{ee} < 30$ MeV/c$^2$, $R$ can be interpreted as the ratio of the virtual photon yield to the inclusive real photon yield:

$$
R(m, p_T) \simeq \frac{dN^*_{\gamma}}{dp_T}(m, p_T)/dN_{\gamma}^{incl}(p_T) \quad (36)
$$

$$
= S(m, p_T) dN_{\gamma}^{direct}(p_T)/dN_{\gamma}^{incl}(p_T). \quad (37)
$$

Figure \ref{fig:32} shows $R$ as a function of $m_{ee}$ for $0.1 < m_{ee} < 0.7$ GeV/c$^2$. The ratio cannot be measured for $m_{ee} < 0.1$ GeV/c$^2$ because in this mass region the signal is masked by large background from $\pi^0$ Dalitz decays. The distributions are consistent with a constant for the five $p_T$ bins. For the highest $p_T$ bin ($p_T > 4$ GeV/c) the shape of the virtual photon mass spectrum is not well constrained due to limited statistics. However, it is reasonable to expect that the same constant behavior continues for higher $p_T$. The $\pm1\sigma$ band of a constant value fit are shown in each panel. Table \ref{tab:11} summarizes the results of the fits. For all $p_T$ bins, the constant value fit gives a good $\chi^2$/NDF value. This demonstrate that the data are consistent with a constant $S(m_{ee})$ for these $p_T$ bins.

As discussed in Appendix \ref{app:3}, the ratio is expected to be a smooth function of $m_{ee}$. Hadronic and partonic direct photon contributions to $S(m_{ee}, q)$ are expected to be nearly constant in this range. $q\bar{q}$ annihilation can make a contribution proportional to $m_{ee}^2$, but should be much smaller than these two components for $0.1 < m_{ee} < 0.3$ GeV/c$^2$ \cite{18,19,20}. There is no sign of a component that scales with $m_{ee}^2$ for $m_{ee} < 0.3$ GeV/c$^2$, suggesting that the $q\bar{q}$ contribution is indeed small. Figure \ref{fig:32} illustrates that $S(M, q)$ is indeed constant, and supports the use of $r$ obtained from our fit as the direct photon fraction in the inclusive photon spectrum.

The absence of any increase in $R(m_{ee})$ for $m_{ee} > 0.5$ GeV/c$^2$ is somewhat surprising. A constant $R$ as function of $m_{ee}$ implies that the $S(m_{ee})$ is also a constant. If the excess electron pairs are thermal pairs from the medium, we expect an increasing contribution from $q\bar{q}$ annihilation or from the tail of the (possibly modified)
The fit range is $0 < T < 7$ GeV/c.

$\rho$ resonance, which leads to increase of $S(m_{ee})$ in higher mass. The data show no indication of such increase. For thermal radiation, $S(m_{ee})$ is the space time average of a product of electromagnetic spectral function and the Boltzmann factor, see Equation B24. We note that the Boltzmann factor $f^B = 1/(e^{E/T} - 1)$ can cause significant suppression of $S(m_{ee}, q)$ for $m_{ee} > 0.5$ GeV/c$^2$. Thus the nearly constant behavior of $R(m_{ee})$ for $m_{ee} > 0.3$ GeV/c should not be interpreted as absence of any contribution other than internal conversion of virtual direct photons.

Figures 33(a) and (b) show the fraction $r = \frac{\text{direct } \gamma}{\text{inclusive } \gamma}$ of the direct photon component determined by the two-component fit (see Equation 34 in $p+p$ and Au + Au collisions, respectively. The curves represent the expectations from a next-to-leading-order perturbative QCD (NLO pQCD) calculation [39]. For $p+p$, the curves show the ratio $d\sigma^{\text{NLO}}(p_T)/dN^\text{incl}(p_T)$, and $d\sigma^\text{incl}(p_T)$ is the $p+p$ inclusive photon cross section (obtained from the data as described later). For Au + Au, the curves represent $T_A d\sigma^{\text{NLO}}(p_T)/dN^\text{incl}(p_T)$, where $T_A$ is the Glauber nuclear overlap function and $dN^\text{incl}(p_T)$ is the Au + Au inclusive photon yield. The three curves corresponding (from top to bottom) to the theoretical scales set to $\mu = 0.5 p_T$, $p_T$, and $2 p_T$, respectively, show the scale dependence of the calculations. While the fraction $r$ is consistent with the NLO pQCD calculation in $p+p$, it is larger than the calculation in Au + Au for $p_T < 4.5$ GeV/c.

The direct photon fraction $r$ in Fig. 33 is converted to the direct photon yield using $dN^\text{dir}(p_T) = r \times dN^\text{incl}(p_T)$. Here we determine the inclusive photon yield for each $p_T$ bin from the yield of low mass $e^+e^-$ pairs in the range $m_{ee} < 30$ MeV/c$^2$ using the following method. The differential yield of electron pairs is related to that of photons by Equation 32. The process dependent factor $S(m_{ee}, q)$ is unity within a few percent for any source of photon for $m_{ee} < 30$ MeV/c$^2$. Thus the measured yield of electron pairs ($N^\text{data}_{ee}$) in $m_{ee} < 30$ MeV/c$^2$ for a given $p_T$ bin is proportional to that of inclusive photons in the same $p_T$ bin.

$$N^\text{data}_{ee}(p_T) = e^{acc}K \frac{dN^\text{incl}}{dp_T}.$$  

$$K = \int_0^{30\text{MeV/c}^2} \frac{2\alpha L(m_{ee})}{3\pi m_{ee}} dm_{ee}.$$  

$TABLE XI$: Summary of a constant fit to the ratio data shown in Fig. 32. The fit range is $0.11 < m_{ee} < 0.7$ GeV/c$^2$.

| $p_T$ (GeV/c) | $\langle R \rangle$ | $\chi^2$/NDF |
|---------------|-----------------|-------------|
| 1.0 - 1.5     | 0.173 ± 0.015   | 16.9/12     |
| 1.5 - 2.0     | 0.149 ± 0.016   | 14.8/12     |
| 2.0 - 2.5     | 0.125 ± 0.020   | 14.7/12     |
| 2.5 - 3.0     | 0.115 ± 0.028   | 9.1/12      |
| 3.0 - 4.0     | 0.210 ± 0.033   | 4.4/12      |
Here $e^{ee}$ represents the acceptance of PHENIX. The same relation holds for the cocktail calculation of photon and electron pairs.

$$N_{ee}^{cocktail}(p_T) = e^{acc} K \frac{dN_{ee}^{\gamma^{cocktail}}}{dp_T}. \quad (40)$$

Here $N_{ee}^{\gamma^{cocktail}}$ is the yield of electron pairs for $m_{ee} < 30 \text{ MeV}/c^2$ in the hadronic cocktail calculation, and $dN_{ee}^{\gamma^{cocktail}}/dp_T$ is the yield of photons in the same calculation. Thus we have

$$\frac{dN_{ee}^{incl}}{dp_T} = \frac{N_{ee}^{data}}{N_{ee}^{cocktail}} \times \frac{dN_{ee}^{\gamma^{cocktail}}}{dp_T}. \quad (41)$$

The systematic uncertainty in the inclusive photon spectra equals the systematic uncertainty in $N_{ee}^{data}$, which is summarized in Table XI. The total systematic uncertainty in $N_{ee}$ is approximately 20%.

In Fig. 34, the direct photon spectra thus obtained are compared with the direct photon data from [73, 80] and NLO pQCD calculations. The systematic uncertainty of the inclusive photon yield is added in quadrature with the systematic uncertainties of the data. The $p+p$ data are shown as an invariant cross section using $d\sigma = \sigma_{pp}^{incl}dN$ using $\sigma_{pp} = 42 \text{ mb}$.

The direct photon data of this analysis are obtained from the yield of $e^+e^-$ pairs using Equation 31 under the assumption $S(m_{ee}, q) = 1$ for $0.1 < m_{ee} < 0.3 \text{ GeV}/c^2$. Although we have shown that our data are consistent with this assumption, what we actually measure is the yield of $e^+e^-$ pairs in this mass range. For completeness, we give the relation between the direct photon yield deduced by the analysis and the electron pair yield that is actually measured. The relation between the real photon yield and the $e^+e^-$ pair yield in this mass range is given by Equation 32. Thus the yield of the excess $e^+e^-$ pairs for $0.1 < m_{ee} < 0.3 \text{ GeV}/c^2$ can be obtained by multiplying the photon yield by a factor of $\frac{dN(\gamma)}{dN(ee)}(c^{\gamma}_{ee}(c^{ee}_{ee})^{-1})$.

The pQCD calculation is consistent with the $p+p$ data within the theoretical uncertainties for $p_T > 2 \text{ GeV}/c$. A similarly good agreement is observed for $p_T < 2 \text{ GeV}/c$. The $p+p$ data can be well described by a modified power-law function $(A_{pp}(1 + p_T^2/b)^{-n})$ as shown by the dashed curve in Fig. 34. The Au + Au data are above the $p+p$ fit curve scaled by $T_{AA}$ for $p_T < 2.5 \text{ GeV}/c$, indicating that the direct photon yield in the low-$p_T$ range increases faster than the binary-scaled $p+p$ cross section.

We fit an exponential plus the $T_{AA}$-scaled $p+p$ fit function $(A e^{-p_T/T} + T_{AA} \times A_{pp}(1 + p_T^2/b)^{-n})$ to the Au + Au data. The only free parameters in the fit are $A$ and the inverse slope $T$ of the exponential term. The systematic uncertainties in $T$ are estimated by changing the $p+p$ fit component and the Au + Au data points within the systematic uncertainties. The results of the fits are summarized in Table XII, where $A$ is converted to $dN/dy$ for $p_T > 1 \text{ GeV}/c$. For central collisions, $T = 221 \pm 19(\text{stat.}) \pm 19(\text{syst.}) \text{MeV}$. If an unmodified power-law function $(\propto p_T^{-n})$ is used to fit the $p+p$ spectrum, we find $n = 5.40 \pm 0.15$, and $T = 240 \pm 21 \text{ MeV}$. 

### Table XII: Summary of the fits to the Au + Au data with the exponential plus the modified power-law function $(A e^{-p_T/T} + B(T_{AA} (1 + p_T^2/b)^{-n})$ as explained in the text. The first and second errors are statistical and systematic, respectively.

| Centrality | $dN/dy(p_T > 1 \text{ GeV}/c)$ | $T(\text{MeV})$ | $\chi^2$/NDF |
|------------|--------------------------------|----------------|--------------|
| 0-20%      | $1.50 \pm 0.23 \pm 0.35$     | $221 \pm 19 \pm 19$ | 4.7/4        |
| 20-40%     | $0.65 \pm 0.08 \pm 0.15$     | $217 \pm 18 \pm 16$ | 5.0/4        |
| Min. Bias  | $0.49 \pm 0.05 \pm 0.11$     | $233 \pm 14 \pm 19$ | 3.2/4        |

### F. $p_T$ Dependence of Low Mass Excess

The shape of the enhancement in Au + Au data in LMR II (low mass low $p_T$) is quite different from that in LMR I (low mass high $p_T$), where it behaves like $1/m_{ee}$ and is consistent with internal conversion of direct photons. In LMR II, the enhancement is larger, as seen in the two lowest $p_T$ bins of Fig. 34. For these bins, no excess is observed in the $p+p$ data. In the lowest $p_T$ bin the enhancement in the Au + Au data is approximately a factor of five above the expectations from the cocktail. The data are significantly above the cocktail up to $m_{ee} = 1 \text{ GeV}/c^2$, reaching their maximum around...
Figure 36(b) suggests that there is an enhancement for still consistent with a constant as a function of mass, large to be conclusive. Figure 36(c) suggests a large and consistent with a constant. For each \( p_T \) function is modified at low constant, suggesting that the electromagnetic spectral large and the data are statistically inconsistent with a \( m_{ee} \) ratio summarized in Table XIII. For \( 0 < p_T < 0.8 \) \( \text{GeV}/c \) cannot be well determined. The results of the fits are shown corresponds to the direct photon fraction \( r \). The fit value for \( 0.8 < p_T < 1.0 \) \( \text{GeV}/c \) the fit gives good \( \chi^2/\text{NDF} \). Thus the data are consistent with the expected \( 1/m_{ee} \) behavior. The next \( p_T \) bin, \( 0.6 < p_T < 0.8 \) \( \text{GeV}/c \), gives marginally satisfactory \( \chi^2/\text{NDF} \). For \( 0.4 < p_T < 0.6 \) \( \text{GeV}/c \) the \( \chi^2/\text{NDF} \) is large and the data are statistically inconsistent with a constant, suggesting that the electromagnetic spectral function is modified at low \( p_T \). However, due to the large uncertainty of the point at \( m_{ee} \approx 0.4 \) \( \text{GeV}/c^2 \) the shape cannot be well determined.

The value \( \langle R \rangle \) obtained from the constant fit corresponds to the direct photon fraction \( r \). The fit value for \( 0.8 < p_T < 1.0 \) \( \text{GeV}/c \), \( \langle R \rangle = 0.177 \pm 0.032 \), is consistent with the values of \( r \) for higher \( p_T \) shown in Fig. 33. If we extrapolate the \( p_T \) spectrum of direct photons deduced from the previous section to lower \( p_T \), the expected direct photon fraction for \( p_T < 1 \) \( \text{GeV}/c \) is \( \approx 0.17 \) or less, since the spectrum of decay photons is steeper. The \( R(m) \) values for \( 0.4 < p_T < 0.6 \) \( \text{GeV}/c \) are larger than this expectation for \( m_{ee} < 0.4 \) \( \text{GeV}/c^2 \), suggesting that the enhancement in the low \( p_T \) region is larger than that expected from internal conversion of direct photons.

In principle, the distribution of \( R \) shown in Fig. 36 can be extrapolated to \( m_{ee} = 0 \) to obtain the fraction of real direct photons, even if the distribution of \( R \) is not flat. However, due to large uncertainties in \( R \) for \( p_T < 0.8 \) \( \text{GeV}/c \) arising from the combination of multiple dilepton sources, we cannot reliably extrapolate the virtual photon yield to \( m_{ee} = 0 \) to determine the real direct photon yield for these two \( p_T \) bins.

We test whether the \( R \) distributions in Fig. 36 are consistent with a constant. For each \( p_T \) bin, we fit a constant to the data. The results of the fits are shown as the horizontal band in each panel and are summarized in Table XIII. For \( 0.8 < p_T < 1.0 \) \( \text{GeV}/c \) the fit gives good \( \chi^2/\text{NDF} \). Thus the data are consistent with the expected \( 1/m_{ee} \) behavior. The next \( p_T \) bin, \( 0.6 < p_T < 0.8 \) \( \text{GeV}/c \), gives marginally satisfactory \( \chi^2/\text{NDF} \). For \( 0.4 < p_T < 0.6 \) \( \text{GeV}/c \) the \( \chi^2/\text{NDF} \) is large and the data are statistically inconsistent with a constant, suggesting that the electromagnetic spectral function is modified at low \( p_T \). However, due to the large uncertainty of the point at \( m_{ee} \approx 0.4 \) \( \text{GeV}/c^2 \) the shape cannot be well determined.

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**G. \( p_T \) Spectra for Different Mass bins**

Figure 35 shows the transverse momentum spectra of dileptons in different mass windows for Au + Au and
pythia contributions. The charm contribution is calculated with hadronic cocktail, which includes also the charm decay $\phi$ of $m$. The systematic uncertainty due to $N_{\text{part}}$ has not been included. In order to avoid the influence of $e^+e^-$ decays of narrow vector mesons, the mass regions around the $\omega$ meson ($0.78 \pm 0.030$ GeV/$c^2$) and the $\phi$ meson ($0.1020 \pm 0.030$ GeV/$c^2$) are excluded. The solid curves in Fig. 37 represent the expectations from the hadronic cocktail, which includes also the charm decay contributions. The charm contribution is calculated with PYTHIA. Using the random $c\bar{c}$ correlation makes no difference for $m_{ee} < 0.3$ GeV/$c^2$ since the charm contribution is negligible at low masses. This difference increases for the higher mass bins, leading to spectra lower by $\sim 10\%$, 20% and 30% for the mass bins $0.3 < m_{ee} < 0.5$ GeV/$c^2$, $0.5 < m_{ee} < 0.75$ GeV/$c^2$, $0.81 < m_{ee} < 0.99$ GeV/$c^2$ respectively, when the random $c\bar{c}$ correlation is used.

First we concentrate on the comparison between data and the sum of cocktail and charm. In the low-$p_T$ region ($p_T < 1$ GeV/$c$) all the $p+p$ spectra are consistent with the expectations from the cocktail alone for every mass window. In the high-$p_T$ region however the $p+p$ data show a small excess above the cocktail. The $Au + Au$ data are in agreement with the cocktail in the mass region $m_{ee} < 0.1$ GeV/$c^2$. In higher mass bins the $Au + Au$ data show a large excess both at low- and at high-$p_T$.

As discussed in section V, we have extracted the direct photon yield from the dilepton spectra in the mass range of 0.1 < $m_{ee}$ < 0.3 GeV/$c^2$. The excess in this mass range is consistent with internal conversion of direct photons. As shown in Figure 24, the direct photon component, which appears as a constant $R$, extends to the $m_{ee}$ > 0.3 GeV/$c^2$. Therefore, there should be sizable contribution from direct photons in the dilepton spectra for $m_{ee} > 0.3$ GeV/$c^2$. The relation between real direct photons and virtual photons is presented in Appendix 3. Here we use a constant factor $S(m_{ee}, q) = 1$ to extend the direct photon component to higher mass ($m_{ee} > 0.3$ GeV/$c^2$). The dashed curves in Fig. 37 show the sum of the cocktail, charm and direct photon contributions to the dilepton spectra for $p_T > 1$ GeV/$c$.

The dashed curves describe the data well for all mass bins both in the $Au + Au$ and the $p+p$ data. This indicates that the excess above the cocktail and charm at high $p_T$ ($p_T > 1$ GeV/$c$) is consistent with the contribution from direct photons. It is surprising that the agreement holds even for $m_{ee} > 0.5$ GeV/$c^2$, where significant modifications of the spectral function may be expected due to the presence of the vector mesons. However the data have large statistical errors for $m_{ee} > 0.5$ GeV/$c^2$ and additional enhancement over the direct photon contribution is not excluded. The data at high $p_T$ are also consistent with the cocktail alone for $m_{ee} > 0.5$ GeV/$c^2$.

In the $Au + Au$ data, the enhancement over the cocktail is approximately a constant factor for $p_T > 1$ GeV/$c$. It grows towards low-$p_T$. All the $Au + Au$ spectra for every mass bin above 0.3 GeV/$c^2$ seem to indicate that the enhancement with respect to the cocktail below 1 GeV/$c$ is significantly larger than above 1 GeV/$c$. For $p_T > 1$ GeV/$c$, the data has a slope similar to the cocktail, as shown by solid curves. For $p_T < 1$ GeV/$c$, the slope of the data is much steeper than the cocktail.

In order to study this change of the slope in the $Au + Au$ data more quantitatively, we subtract the cocktail plus charm from the data and examine the shape of the excess. The $p_T$ spectra are combined in the mass range 0.3 < $m_{ee} < 0.75$ GeV/$c^2$. In this mass range, the low-$p_T$ cutoff which artificially truncates the $p_T$ spectra at lower mass is avoided. The combined data also have increased statistical significance.

Figure 35 shows the pair $m_T - m_0$ spectrum in $Au + Au$ for the pair mass range 0.3 < $m_{ee} < 0.75$ GeV/$c^2$. Here $m_T = \sqrt{p_T^2 + m_0^2}$ is the transverse mass of the pair and $m_0$ is the mean value of $dN/dm_{ee}$ in the given mass range ($0.3 < m_{ee} < 0.75$ GeV/$c^2$ in this case). We plot the data as function of $m_T - m_0$ since invariant differential cross sections of hadrons in $p+p, p+A$, and $A+A$ collisions are generally well described by exponential functions in $m_T$. Thus the change in the slope can be seen more clearly in the $m_T$ spectrum. The $m_T$ spectrum shows a clear change in the slope around 1 GeV/$c^2$. The slope below $m_T - m_0 < 1$ GeV/$c^2$ is much steeper than that above 1 GeV/$c^2$. In order to characterize the change of the slope in the two $m_T$ regions, we fit the $m_T$ spectrum with the sum of two exponentials:

$$\frac{d^2N}{2\pi m_T dm_T dy} = A_1 \cdot \exp\left(-\frac{m_T}{T_1}\right) + A_2 \cdot \exp\left(-\frac{m_T}{T_2}\right)$$  \hspace{1cm} (43)
the value of $T_2$, is somewhat higher than, but consistent with, the inverse slope of the exponential component of the direct photon spectrum, $T = 221 \pm 19^{\text{stat}} \pm 19^{\text{syst}}$ MeV, obtained in the previous section.

In Fig. 38 the same Au + Au $p_T$ spectrum is fit with the sum of direct photons and an exponential function in $m_T$. The exponential function is to characterize the low $p_T$ component. The direct photon component is obtained from the direct photon spectrum in Fig. 38, and extended to the larger mass region assuming $S(m, q) = 1$. We then convert the photon yield to the $e^+e^-$ pair yield using Equation 34. Thus the direct photon component is fixed and the only free parameters of the fits are the normalization and the inverse slope $T$ of the exponential component.

The $p_T$ spectrum and the individual components of the fit function, i.e. the direct photon component, and the exponential component are shown in Fig. 39. The spectrum is well reproduced by the fit and $\chi^2$/NDF = 16.6/11. The systematic uncertainty accounts for the uncertainty on the data and the uncertainty ($\sim 20\%$) of the cocktail normalization. From the fit, we extract a value of $T = 86.5 \pm 12.7^{\text{stat}} +11.0-28.4^{\text{syst}}$ MeV. The yield of the low-$p_T$ exponential extracted from the fit contributes more than 50% of the yield of the spectrum.

Both the two exponentials fit shown in Fig. 38 and the exponential + direct photon fit shown in Fig. 39 show that there is a low inverse slope component with $T \approx 100$ MeV for $m_T - m_0 < 0.6$ GeV/$c^2$. In order to further study the mass dependence of the inverse slope, we calculated the local slopes of the invariant pair $m_T$ spectra obtained from the $p_T$ spectra shown in Fig. 37. For all the mass bins the cocktail and charm are subtracted from the data and the average inverse slope ($\langle T(m_0) \rangle$) of the excess at mass $m_0$ has been numerically calculated as

$$\langle T(m_0) \rangle = \frac{\Sigma(m_T^i - m_0)f(m_T^i)\Delta m_T^i}{\Sigma f(m_T^i)\Delta m_T^i},$$

(44)
where

\[ f(m_T) = \frac{1}{2\pi m_T} \frac{d^2 N}{dm_T dy} \]  
\[ = \frac{1}{2\pi p_T} \frac{d^2 N}{dp_T dy} \]

is the invariant spectrum of the electron pairs after cocktail subtraction shown in Fig. [2].

Figure [10] shows the inverse slopes calculated in two ranges, namely \(0 < m_T < 0.6 \text{ GeV}/c^2\) and \(0.6 < m_T < 2.5 \text{ GeV}/c^2\). For \(m_0 < 0.4 \text{ GeV}/c^2\) the spectra are truncated due to the acceptance; therefore, we do not quote any slope here. Also for \(m_0 < 0.1 \text{ GeV}/c^2\) the slope in the range \(0.6 < m_T < 2.5 \text{ GeV}/c^2\) is not shown because the cocktail subtraction has too large systematic uncertainty in this region. The solid and dashed lines show the inverse slope of the cocktail, calculated in the same way as the data, for the same mass ranges. The inverse slope parameters obtained from the two-exponential fit as well as the exponential + direct photon fit in the same mass ranges are also shown.

Figure [11] suggests qualitatively different behavior for low and high \(m_T\). In the large \(m_T - m_0\) range \((0.6 < m_T - m_0 < 2.5 \text{ GeV}/c^2)\) the inverse slope \(T\) is approximately 300 MeV and similar to that of the cocktail (shown in the dashed-dotted line). In the low \(m_T - m_0\) range the inverse slope is approximately 100 MeV, similar to that obtained with the two fit methods. This latter one (shown at \(m_T - m_0 \sim 0.45 \text{ GeV}/c^2\)) is lower than that of the cocktail in similar kinematic range (shown in the solid line).

The effective temperature of the lower inverse slope component \(T \simeq 100 \text{ MeV}\), obtained from the two fit methods as well as the numerical calculation, is much lower than the inverse slope of hadrons with similar masses (kaons) measured by PHENIX [61]. The slope of the kaon spectrum is larger than 200 MeV. The hadron slopes rise linearly with mass, consistent with the expectations from radial expansion of the hadronic source. If arising from thermal radiation of the fireball, dominated by pion-pion annihilation \(\pi^+\pi^- \rightarrow \rho \rightarrow e^+e^-\), the excess yield in the LMR would show similar temperatures and a similar linear rise, reminiscent of radial flow of a hadronic source [62]. The value of the low-\(p_T\) inverse slope is lower than or similar to the freeze-out temperature. Also the inverse slope of dileptons with an average
mass smaller than 0.5 GeV/c² is more than a factor 2 smaller than that of kaons. The simplistic expectation, that 0.5 GeV/c² dilepton emission is created similarly to kaons (from an equilibrated flowing source), is not supported by the data.

VI. THEORY COMPARISON

The Au + Au e⁺e⁻ spectra are now compared to different models of e⁺e⁻ production in the LMR and in the IMR. These models, employed at SPS energies, identified the pion annihilation process as the main source of thermal dileptons in the hadronic phase of the fireball. However, this process, mediated by the intermediate ρ meson, failed to describe the observed enhancement in the LMR [28] at the SPS energy when vacuum properties of the ρ are used. This suggested that in-medium modifications of the ρ spectral function could be responsible for the enhancement of dilepton yield below the ρ mass. The proposed modifications mostly followed two different approaches:

- The **dropping mass** scenario followed the scaling conjecture of G.E. Brown and M. Rho [21], which postulates that the mass of vector mesons decreases in dense matter as the quark condensate ⟨q̅q⟩ decreases due to partial restoration of chiral symmetry. This leads to a decrease of the mass of ρ meson from its vacuum value (0.77 GeV/c²) and causes enhancement of the dilepton yield below the ρ mass.

- The **broadening mass** scenario explains the LMR dilepton enhancement by hadronic many-body interactions [29]. The spectral function in a hot and strongly interacting hadron resonance gas is calculated. The many-body interactions cause the broadening of the ρ resonance, leading to enhancement of dilepton yield below ρ mass. Other hadronic many-body interactions contribute to low mass enhancement.

There seems now to be consensus that the dropping mass scenario alone cannot adequately reproduce the SPS data. These two scenarios, sometimes in combination, are used by different groups to calculate the dilepton yield at the top RHIC energy (√sNN = 200 GeV). Here we compare calculations by the following three groups to the PHENIX data.

- Rapp and van Hees [15, 18, 83] calculate the rate of dilepton emission from a hadronic gas in thermal equilibrium. In the calculation, the electromagnetic spectral function in the vacuum is constrained by the data of e⁺e⁻ annihilation into hadrons. The spectral function in the medium is modified by hadronic many-body interactions. The spectral function is characterized by the light vector resonances ρ(770), ω(782) and φ(1020) at low-mass according to the vector dominance model (VDM), and a perturbative quark-antiquark continuum at higher masses. Dilepton production from a partonic phase outshines the hadronic gas radiation for m_{ee} > 1.5 GeV/c² at RHIC energy due to high initial temperatures.

- Dusling and Zahed [13, 84, 85] use a chiral reduction formalism to calculate the electromagnetic current-current correlation function in the medium. The experimental data of e⁺e⁻ annihilation, τ-decay, two-photon fusion reactions, and pion radiative decays are used to constrain the correlation function. The dilepton emission rates from hadronic gas at finite temperature and baryon density are then computed from a hydrodynamic evolution model of Au + Au collisions. The partonic contribution, which does not become dominant below the φ mass, is computed using the Born q̅q annihilation term.

- Cassing and Bratkovskaya [20, 27, 84, 87] use a microscopic relativistic transport model (HSD) that incorporates the relevant off-shell dynamics of the vector mesons. This model is well established to describe the yields, the rapidity distributions and the transverse momentum spectra of hadrons in p + A and A + A collisions from SIS to RHIC energies. The model reproduces well the dilepton mass spectrum in p + p collisions. In Au + Au the model includes a modified ρ spectral function according to a collisional broadening scenario as well as a tunable
FIG. 41: (Color online) Invariant mass spectra of $e^+e^-$ pairs in Min. Bias Au + Au collisions in the IMR. (top left) The data are compared to the sum of cocktail+charm. The data are also compared to the sum of cocktail-charm and partonic contributions from different models. The calculations are from (center) Rapp and van Hees [15, 18, 83] and (right) Dusling and Zahed [19, 84, 85]. The partonic yields (PY) have been added to the two scenarios for charmed mesons decays, i.e. (i) PYTHIA and (ii) random $c\bar{c}$ correlation.

dropping $\rho$ mass scenario. No yield from the partonic QGP phase is available at the moment from the HSD model.

We received numerical values of these model calculations from the authors. The $e^+e^-$ rates from these calculations are filtered into the PHENIX acceptance and compared to the data.

The theory calculation by Rapp and van Hees [15, 18, 83] is done for a fixed impact parameter $b = 8$ fm. In this calculation the number of charged tracks $N_{ch} = 230$, which is consistent with the measured $N_{ch}$ in Au + Au MB data [2]. The theory calculation by Dusling and Zahed [19, 84, 85] is done for a fixed impact parameter for $b = 0$ fm or $N_{part} = 378$.

Since the calculations were provided for central collisions, or for collisions with a fixed impact parameter, in the comparison to our Min. Bias data we normalize the theory calculations by $N_{model}/N_{part, Min, Bias}$, i.e. assuming that the dielectron yield scales with $N_{ch}$. However, this scaling procedure may have introduced some bias in the comparison, as the data show (Fig. 29) that the dielectron yield increases faster than $N_{part}$ (and $N_{ch}$ is proportional to $N_{part}$). A more detailed comparison requires knowledge of centrality dependence of the dielectron yield both in the data and in the theoretical model.

For comparison to the data, we add these calculations to the hadronic cocktail and the charm decays. The contribution from the freeze-out $\rho$ meson is subtracted from the cocktail to avoid double-counting. Because the calculation with the HSD transport model by Cassing and Bratkovskaya [20, 27, 86, 87] can sample any impact parameter $b$ with a proper geometrical weight. In this case the model also calculates the hadronic contributions, which are in a good agreement with the cocktail. Only the charm contribution is taken from the PYTHIA calculation [77] or from our calculation with random correlation.

In Section VII the Au + Au photon spectrum is compared to several theoretical predictions. These employ hydrodynamical models to calculate thermal photon emission from the thermalized partonic and hadronic
FIG. 42: (Color online) Invariant mass spectra of $e^+e^-$ pairs in Au + Au collisions in the LMR. The data are compared to the sum of cocktail+charm (top left). The data are also compared to the sum of cocktail+charm and hadronic+partonic contributions from different models. The calculations are from The calculations are from (top right) Rapp and van Hees [15, 18, 83], (bottom right) Dusling and Zahed [19, 84, 85], and Cassing and Bratkovskaya [20, 27, 86, 87].

phases of the reaction, added to NLO pQCD calculations that describe prompt photon emission from perturbative parton-parton scatterings in the first tenths of fm/c of the collision process.

The experimental conditions reached at mid-rapidity in central heavy ion collisions at RHIC of nearly zero net baryon density and longitudinally boost-invariance in the initial conditions facilitate the applicability of hydrodynamics to describe the reaction evolution. In addition, the thermalization times usually assumed in the hydrodynamical models ($\tau_{\text{therm}} \lesssim 1$ fm/c) are, for the first time at RHIC, above the lower limit imposed by the transit time of the two colliding nuclei ($\tau_0 = 2R/\gamma \approx 0.15$ fm/c for Au + Au at 200 GeV).

Hydrodynamical approaches describe, under the assumption of local conservation of energy and momentum, the evolution of the system using the equations of motion of perfect relativistic hydrodynamics complemented with a set of initial conditions (e.g. initial temperature $T_{\text{init}}$, thermalization time $\tau_0$, the equation-of-state of the system, and the freeze-out conditions. These models have been very successful in describing quantitatively most of the differential observables of bulk hadronic production (in particular those sensitive to early-times pressure gradients).

The same hydrodynamical models, with initial conditions chosen so as to reproduce the bulk hadron data, are now employed to carry out the description of thermal photon production over the whole space-time evolution of the system.

A. Comparison in the IMR and constraint on possible QGP radiation

In this Section we compare the model calculations with the data in the IMR and investigate whether the experimental data can constrain the QGP radiation in the IMR. We also study whether we can constrain contributions from charm and QGP radiation in the LMR.

All theoretical models predict that there is large contribution from QGP radiation in the IMR. The QGP radiation competes with the dileptons from correlated charm, which contribute much more than any other cocktail con-
tribution in the IMR. There is a large uncertainty in the charm contribution in Au + Au, since charm quarks are known to suffer energy loss in the medium. Therefore the $e^+e^-$ mass shape from semi-leptonic decays of charmed quarks may be modified. The shape calculated from PYTHIA [77] provides an upper limit for the expected $e^+e^-$ yield. The shape with random $c\bar{c}$ correlation is softer for $m_{ee} > 0.5$ GeV/$c^2$, and this provides the lower limit. Below 0.5 GeV/$c^2$ the two shapes are almost identical.

These two scenarios for the open charm contribution are added to the predictions for the QGP radiation provided by the models described above and are compared to the experimental data in Fig. [41]. In all the models the yield in the QGP phase arises entirely from the $q\bar{q} \rightarrow e^+e^-$ annihilation process. The magnitude of the yield is closely linked to the thermalization time $\tau_0$. A larger $\tau_0$ translates into a reduction of the initial temperature and thus of the total QGP radiation. In Fig. [41] the differences between the models are attributed to different initial conditions used for the hydrodynamic evolution of the QGP phase ($\tau_0 = 0.2$ fm/$c$ for Dusling and Zahed [43, 54, 55], $=0.6$ fm/$c$ or Rapp and van Hees [13, 18, 83]).

In the IMR the data have large statistical errors and systematic uncertainties. Therefore, there do not allow discrimination either between the two proposed scenarios for charm production (PYTHIA [77] or random $c\bar{c}$ correlation) or among the three theoretical models.

B. Inclusive Low Mass Excess

The data in the IMR do allow setting an upper limit on the contribution arising from charm or from $q\bar{q} \rightarrow e^+e^-$ going to the LMR. Indeed we can saturate the IMR yield either with charm or with the partonic yield calculated by the theorists and see what their contribution would be in the LMR. Thus, neither the charm nor the contribution from $q\bar{q} \rightarrow e^+e^-$ can be solely responsible of the LMR enhancement.

In the LMR the shape and the yields of the $e^+e^-$ mass spectra from charm calculated by PYTHIA [77] are not very different from those given by a calculation which assumes random $c\bar{c}$ correlation, because the shape is mostly determined by the geometrical acceptance. Thus the dilepton yield measured in the IMR gives strong constraint on the charm contribution in the LMR. Since the calculated charm contribution is consistent with the data and is less than the hadronic cocktail below $m_{ee} < 0.5$ GeV/$c^2$, we conclude that charm contribution alone cannot explain the large enhancement observed in the LMR. A similar consideration can be given for the QGP radiation. The contribution from $q\bar{q} \rightarrow e^+e^-$ process is negligible in LMR I.

Figure [42] compares the inclusive mass spectrum in the LMR with the cocktail+charm only and with cocktail+charm and the calculations by three groups: Rapp and van Hees [15, 18, 83], Dusling and Zahed [19, 84, 55], and Cassing and Bratkovskaya [20, 27, 88, 87]. In all three models the $e^+e^-$ yield in the LMR arises mostly from the hadronic phase. Rapp and van Hees [15, 18, 83] propose three different scenarios of vector mesons spectral functions: (i) no medium effects, (ii) dropping $\rho$ mass, and (iii) broadening $\rho$ mass. Dusling and Zahed [19, 84, 55] use a broadening $\rho$ mass scenario in the hadronic phase and a pion chemical potential of $\mu_\pi = 50$ MeV. Cassing and Bratkovskaya [20, 27, 88, 87] propose two scenarios: (i) broadening $\rho$ mass, and (ii) dropping $\rho$ mass and broadening mass.

The common characteristic of the in-medium effects in these models is a slight suppression of the yield in the $\rho-\omega$ region compared with the unmodified $\rho$ scenario and an enhancement in the region $0.4 < m_{ee} < 0.7$ GeV/$c^2$. The $\phi$ survives as a pronounced resonance, although its width is broadened. These features become less distinct once the cocktail contribution and the smooth yield from the QGP (which constitutes 15-20%) are added. The differences in the yields of $e^+e^-$ pairs in the various models are attributed to differences in the medium effects on the spectral function, different durations of the lifetime of the

![FIG. 43: (Color online) Invariant mass spectra of $e^+e^-$ pairs in Min. Bias Au + Au collisions for different $p_T$ windows compared to the expectations from the calculations of Rapp and van Hees [15, 18, 83], separately showing the partonic and the hadronic yields and the different scenarios for the $\rho$ spectral function, namely “Hadron Many Body Theory” (HMBT) and “Dropping Mass” (DM). The calculations have been added to the cocktail of hadronic decays (where the contribution of the freeze-out $\rho$ meson is subtracted) and charmed meson decays products.](image-url)
fireball in the hadronic phase, and different evolutions of the temperature as a function of time.

While the calculations proposed by Rapp and van Hees [14, 15, 18, 27, 86] agree with the data for \( m_{ee} > 0.5 \text{ GeV}/c^2 \), the ones of Cassing and Bratkovskaya [20, 27, 86, 87] touch the lower end of the systematic uncertainty in the same mass region. The yields calculated by Dusling and Zahed [19, 84, 85] appear everywhere too low to add significant contribution in the LMR, where the shape of the distribution is essentially determined by the detector acceptance.

All of the models under predict the data for \( 0.2 < m_{ee} < 0.5 \text{ GeV}/c^2 \) by at least a factor of two. It should be noted that the contributions in the region \( m_{ee} < 0.4 \text{ GeV}/c^2 \) are very different in the three models compared. In Rapp and van Hees [14, 15, 18, 27, 86] this contribution arises from processes like \( \alpha \Lambda \rightarrow \pi \gamma^* \rightarrow \pi e^+e^- \) or \( N \rightarrow N\gamma^* \rightarrow Ne^+e^- \). Those processes in the HSD model are suppressed by a few orders of magnitude and are not seen at all compared to the major Dalitz decays. In Dusling and Zahed [19, 84, 85] the main contribution below the two-pion threshold comes from \( \pi_A \), the axial-vector contribution in medium. However the absolute yield of this process is too low because it is concentrated at very low-\( p_T \) which is suppressed by our acceptance cut.

\[
(\vec{p}_T)_{\text{single}} > 0.2 \text{ GeV}/c.
\]

C. \( p_T \) Dependence of Low Mass Excess

Figs. 43, 44, and 45 show the \( e^+e^- \) invariant mass spectra in different \( p_T \) windows from data compared to the sum of all these contributions for the predictions of Rapp and van Hees [14, 15, 18, 27, 86], Dusling and Zahed [19, 84, 85], and Cassing and Bratkovskaya [20, 27, 86, 87], respectively. The contribution from the hadronic and the partonic/medium/charm/expectations from PYTHIA [77] are shown separately. The charm distribution from PYTHIA is somewhat harder than the calculations which assume random correlation of the \( c\bar{c} \) pair but not very different in the LMR, where the shape of the distribution is essentially determined by the detector acceptance.

From the comparison we learn that in general the yield from these theoretical predictions is insufficient to explain the observed enhanced dilepton production, both at low- and high-momenta. At low-\( p_T \), where the enhance-
moment reaches approximately a factor of five, the shape of the enhancement shown by the data is quite different from any of the theoretical models.

At high-$p_T$ ($p_T > 1.0$ GeV/c) the enhancement is about a factor two over the cocktail and its shape is quite similar to that of the cocktail. In the previous Section we showed that this enhancement can be attributed to internal conversion of virtual direct photons. In the theoretical calculations of direct photon emission at RHIC energies, the contribution from the QGP phase, e.g. quark-gluon Compton scattering, is the dominant source of real thermal photons for $p_T > 1$ GeV/c. The process that produces real photons in the QGP should also contribute low mass $e^+e^-$ pairs at high $p_T$. However, none of the three models includes such processes (e.g. $q + g \rightarrow q + \gamma^* \rightarrow q + e^+e^-$). The QGP radiation in these models only include $q+\bar{q}$ annihilation. This could explain the discrepancy between the models and the data for $p_T > 1.0$ GeV/c.

D. $p_T$ Spectra in the Low Mass Region

The $p_T$ spectra of the excess (i.e. after subtracting the hadron cocktail and the charm from the dilepton spectra) can be also compared to the theoretical models. We already noted that the yields from the partonic medium in the theoretical models are produced only via the $q\bar{q} \rightarrow e^+e^-$ annihilation process. Processes like $q + g \rightarrow ge^+e^-$ are not included.

Figure 34 shows the $p_T$ spectrum in the mass window $0.3 < m_\text{ee} < 0.75$ GeV/$c^2$ after subtracting the contribution from the cocktail and the charm. The spectrum is compared to the theoretical calculation from Rapp and van Hees [13, 83], Dusling and Zahed [13, 84, 85], and Cassing and Bratkovskaya [24, 27, 50, 57], respectively. The figure shows separately the $e^+e^-$ yields from the partonic phase and the hadronic phase (with two possible implementations of the $p$ spectral function) and their sum is compared to the data. In all the models the sum of the cocktail contribution and the $e^+e^-$ yield from medium-effects is insufficient to explain the experimental data, and divergences are observed both at low- and high-$p_T$. While for Rapp and van Hees [13, 83] and for Dusling and Zahed [13, 84, 85] the disagreement with the data is strong at low-$p_T$, for Cassing and Bratkovskaya [24, 27, 50, 57] a better agreement is achieved over the full $p_T$ range. However, the data seem still higher than the theoretical calculations.

E. Theoretical Comparison to Direct Photon Measurement

In section V E we have extracted the direct photon yield from the analysis of LMR I. The obtained direct photon spectrum in central Au + Au, shown in Figure 34, shows excess over $T_{\text{AA}}$ scaled $p+p$ data, and the shape of the excess is well described by a pure exponential with inverse slope $T \approx 220$ MeV. If the direct photons in Au + Au collisions are of thermal origin, the inverse slope $T$ is related to the initial temperature $T_{\text{init}}$ of the dense matter. In hydrodynamical models, $T_{\text{init}}$ is 1.5 to 3 times $T$ due to the space-time evolution.

Figure 47 compares the direct photon data in central 0-20% Au + Au collisions with several theoretical calculations of thermal photon emission added to the pQCD calculations. Note that the curve by [94] includes pQCD contributions, while the others do not. For $p_T < 3$ GeV/c, the thermal contribution dominates over pQCD. These hydrodynamical models can reproduce the high $p_T$ central Au + Au data within a factor of two. These models assume formation of a hot QGP with initial temperature ranging from $T_{\text{init}} = 300$ MeV at thermalization time $\tau_0 = 0.6$ fm/c to $T_{\text{init}} = 600$ MeV at $\tau_0 = 0.15$ fm/c [38, 39, 40, 41, 42, 43, 44]. Figure 48 summarizes the $T_{\text{init}}$ and $\tau_0$ for theoretical calculations shown in Fig. 47. There is a clear anti-correlation between $T_{\text{init}}$ and $\tau_0$. Lattice QCD predicts a phase transition from hadronic phase to quark gluon plasma at $\approx 170$ MeV.

VII. SUMMARY AND CONCLUSIONS

PHENIX has measured dilepton production in Au + Au and $p+p$ collisions at $\sqrt{s_{NN}} = 200$ GeV. The measured $e^+e^-$ yield is compared with the cocktail of known sources of light hadron decays. Cocktail sources are mostly measured by PHENIX in the same experimental run via hadronic decay channels. Extrapolations to low-$p_T$, where experimental data are not always available, are obtained using an $m_T$-scaling procedure.

In the $p+p$ data the $e^+e^-$ invariant mass spectrum has been measured in the mass range from 0 to 8 GeV/$c^2$ for all pair-$p_T$. The intermediate-mass-region (IMR) is dominated by semi-leptonic decays of heavy flavor mesons for which the extracted production cross sections is consistent with fixed-order next-to-leading-log (FONLL) predictions and with the PHENIX measurement of single electrons [45].

The low-mass-region (LMR) can be described by known contributions from light meson decays, and virtual direct photons for which the extracted cross section is consistent with NLO pQCD calculations and with PHENIX measurements of real photons.

In Au + Au collisions, the data are consistent with the expectations from correlated $c\bar{c}$ production for $m_{ee} > 0.5$ GeV/$c^2$. However, this interpretation is ambiguous, due to the interplay between possible two different medium effects: energy loss of charm quarks in the medium which would deplete the yield in the IMR, and QGP radiation, which would increase the yield in the IMR.

In the low mass region the Au + Au Min. Bias inclusive mass spectrum shows an enhancement by a factor of $4.7 \pm 0.4_{\text{stat}} \pm 1.5_{\text{syst}} \pm 0.9_{\text{model}}$ compared to the expect-
tation from the hadronic cocktail. The enhancement is concentrated at low $p_T$ ($p_T < 1$ GeV/c). The integrated yield increases faster with the centrality of the collisions than the number of participating nucleons.

At low mass ($m_{ee} < 0.3$ GeV/$c^2$) and high $p_T$ ($1 < p_T < 5$ GeV/c) an enhanced $e^+e^-$ pair yield is observed both in $p+p$ and Au + Au collisions. The mass dependence of the excess is consistent with that expected for virtual direct photon production. This excess is used to infer the yield of real direct photons by extrapolating to $m_{ee} = 0$. A perturbative QCD calculation is consistent with the real direct photon cross section in $p+p$ extracted by this method, while in central Au + Au collisions much larger yields compared with the $p+p$ cross section scaled with $T_{AA}$ are observed. In central Au + Au collisions, the excess over the $p+p$ cross section scaled by $T_{AA}$ is exponential in $p_T$, with inverse slope $T = 221 \pm 19_{\text{stat}}^{+19_{\text{sys}}} \text{ MeV}$.

In Au + Au collisions at very low $p_T$ there is a further, very significant, enhancement that increases strongly with centrality. The $p_T$ spectrum of dileptons in this region has been analyzed with two fit methods and a numerical calculation. An inverse slope of $T \approx 100$ MeV has been extracted for $m_T < 0.6$ GeV/c, lower than for hadrons with similar mass and similar to the freeze-out temperature.

The Au + Au data are compared to different models which provide additional $e^+e^-$ yield in both the LMR and the IMR. In the IMR the data have too large uncertainty to discriminate different possible scenarios of charm production and QGP radiation. In the LMR no quantitative agreement has been found yet with the models.

The yield of direct photons in Au + Au collisions is compared with several hydrodynamical models of thermal photon emission at RHIC energies. The models assuming the formation of a hot system with initial temperature ranging $T_{\text{init}} \approx 300 - 600$ MeV at times $\tau_0 \approx 0.6 - 0.15$ fm/c are in qualitative agreement with the data. Lattice QCD predicts a phase transition from hadronic phase to quark gluon plasma at $T \approx 170$ MeV.

In conclusion, we presented measurements of the $e^+e^-$ continuum in $p+p$ and Au + Au at $\sqrt{s_{NN}} = 200$ GeV in a wide range of mass and transverse momenta. In Au + Au collisions, a large enhancement of the yield of $e^+e^-$ pairs is observed at low mass and low $p_T$ in Au + Au. The yield of direct photons is deduced from low mass, high $p_T$ $e^+e^-$ pairs. Future measurements with an upgraded PHENIX detector with higher statistics, together with further advance in theory, will allow more detailed study of the properties of the hot dense matter formed in heavy ion collisions at RHIC.

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The black solid curve shows the pQCD calculation, scaled by the others, the curve by \[94\] includes pQCD contributions. The direct photon data in central 0-20% Au + Au collisions shown in Fig. 47 are compared with the QCD scale uncertainty, with the upper curve and the lower curve corresponding to \(\mu = 1/2\cdot pt\) and \(\mu = 2\cdot pt\), respectively.

**APPENDIX A: BACKGROUND NORMALIZATION**

1. **Pairing of Electrons and Positrons**

In the following we assume that, as dictated by the charge conservation law, \(e^-\) and \(e^+\) are always produced in pairs and that most of these pairs are produced statistically independent of each other. Let us say \(N\) pairs are produced in a particular event and \(N\) is given by a probability distribution \(P(N)\). Of the \(N\) pairs only a fraction \(\varepsilon_p\) is fully reconstructed, and then the number of reconstructed pairs \(n_p\) is given by a binomial distribution \(B\) sampling out of \(N\) “events” with a probability \(\varepsilon_p\).

- Probability to get \(n_p\) pairs from \(N\) true pairs: \(\omega(n_p) = B(n_p, N, \varepsilon_p)\)
- with an average: \(\langle n_p \rangle = \varepsilon_p N\)
- and variance: \(\sigma^2_p = \varepsilon_p N (1 - \varepsilon_p)\)

Of the remaining pairs one track is reconstructed with a probability \(\varepsilon_+\) or \(\varepsilon_-\). For a given \(N\) and \(n_p\) the number of additional single positive tracks \(n_+\) and negative tracks \(n_-\) follow a multinomial distribution \(M\) with three possible outcomes for each of the \(N - n_p\) unreconstructed pairs: no track, one + track or one − track.

The probability to get \(n_+\) and \(n_-\) single tracks from \(N\) true pairs with \(n_p\) reconstructed pairs, i.e., from \((N - n_p)\) not fully reconstructed pairs is:

\[
\omega(n_+, n_-) = \sum_{n_-=1}^{N-n_p} M(n_+, n_-; N - n_p, \varepsilon_+, \varepsilon_-)
\]

\[
\omega(n_+) = \sum_{n_-=1}^{N-n_p} M(n_+, n_-; N - n_p, \varepsilon_+, \varepsilon_-)
\]

\[
\omega(n_-) = \sum_{n_+=1}^{N-n_p} M(n_+, n_-; N - n_p, \varepsilon_+, \varepsilon_-)
\]

- with average: \(\langle n_{\pm} \rangle = \varepsilon_{\pm}(N - n_p)\)
- variance: \(\sigma^2_{\pm} = \varepsilon_{\pm}(N - n_p)(1 - \varepsilon_{\pm})\)

---

**FIG. 47:** (Color online) Theoretical calculations of thermal photon emission \([32, 33, 34, 35, 36, 37]\) are compared with the direct photon data in central 0-20% Au + Au collisions shown separately and added to pQCD calculations. In contrast to the others, the curve by \([34]\) includes pQCD contributions. The black solid curve shows the pQCD calculation, scaled by \(T_{\text{AA}}\). The QCD scale \(\mu\) is set to \(pt\) for this calculation. The two black dashed curves around the black solid curve show the scale uncertainty, with the upper curve and the lower curve corresponds to \(\mu = 1/2 \cdot pt\) and \(\mu = 2 \cdot pt\), respectively.

**FIG. 48:** (Color online) \(T_{\text{AA}}\) vs. \(\tau_0\) for various theoretical calculations shown in Fig. 47.

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• and covariance: \( \text{cov}(n_+, n_-) = -(N - n_p)\varepsilon_+\varepsilon_- \)

In this case the number of unlike-sign pairs for a given \( N \) and \( n_p \) is:

\[
\langle n_{+-} \rangle = n_p^2 + n_p \sum_{n_i = 1}^{N-n_p} n_i \omega(n_i) + n_p \sum_{n_i = 1}^{N-n_p} n_i \omega(n_i) \\
+ \sum_{n_i = 1}^{N-n_p} \sum_{n_i = 1}^{N-n_p} n_i n_j \omega(n_i, n_j)
\]

\[
\langle n_{++} \rangle = \sum_{n_i = 1}^{N-n_p} (n_p + n_i)(n_p + n_i - 1)\omega(n_i)
\]

\[
\langle n_{--} \rangle = n_p^2 - n_p + \langle n_i^2 \rangle - \langle n_{+-} \rangle + 2n_p\langle n_{+-} \rangle
\]

Similarly we can calculate the number of like-sign pairs:

\[
2\langle n_{++} \rangle = \sum_{n_i = 1}^{N-n_p} (n_p + n_i)(n_p + n_i - 1)\omega(n_i)
\]

\[
2\langle n_{--} \rangle = n_p^2 - n_p + \varepsilon_+^2(N - n_p) - 2\varepsilon_+\varepsilon_-(N - n_p) + \varepsilon_+^2(N - n_p) - 2\varepsilon_+\varepsilon_-(N - n_p) + \varepsilon_+^2(N - n_p)
\]

Now we calculate the like-sign background:

\[
2\langle N_{++} \rangle = \sum_{n_i = 1}^{N-n_p} (n_p + n_i)B(n_p, N, \varepsilon_p)
\]

\[
= \varepsilon_p^2N^2 + \varepsilon_p(1 - \varepsilon_p)N - \varepsilon_pN + \varepsilon_p\varepsilon_+^2N^2 \\
+ \varepsilon_p^2(1 - \varepsilon_p)N - 2\varepsilon_+\varepsilon_pN^2 + \varepsilon_+^2N^2 \\
- \varepsilon_+^2N + \varepsilon_+\varepsilon_pN + 2\varepsilon_+\varepsilon_pN^2 \\
- 2\varepsilon_+\varepsilon_p^2N^2 - 2\varepsilon_+\varepsilon_+(1 - \varepsilon_p)N
\]

\[
\langle N_{++} \rangle = \sum_{n_i = 1}^{N-n_p} (n_p + n_i)B(n_p, N, \varepsilon_p)
\]

\[
= \varepsilon_+^2(N^2 - N) + \varepsilon_+\varepsilon_-(N^2 - N) + \varepsilon_+\varepsilon_-(N^2 - N) + \varepsilon_+\varepsilon_-(N^2 - N)
\]

\[
+ 2\varepsilon_+\varepsilon_pN^2 - 2\varepsilon_+\varepsilon_pN^2 - 2\varepsilon_+\varepsilon_pN + 2\varepsilon_+\varepsilon_pN^2
\]

\[
= \varepsilon_+^2(N^2 - N) + \varepsilon_+\varepsilon_-(N^2 - N) + \varepsilon_+\varepsilon_-(N^2 - N)
\]

\[
2\langle n_{--} \rangle = n_p^2 - n_p + \varepsilon_+^2(N - n_p)^2 \\
- \varepsilon_+^2(N - n_p) + 2\varepsilon_+\varepsilon_p(N - n_p)
\]

To obtain the expected number of like- and unlike-sign pairs for a fixed number of real pairs \( N \) we need to average over all possible reconstructed pairs \( n_p \):

\[
\langle N_{++} \rangle = \sum_{n_i = 1}^{N-n_p} (n_p + n_i)B(n_p, N, \varepsilon_p)
\]

\[
= (1 - \varepsilon_+ - \varepsilon_- + \varepsilon_+\varepsilon_-)(n_p^2) \\
+ \varepsilon_+\varepsilon_+N^2 - 2\varepsilon_+\varepsilon_-N + \varepsilon_+\varepsilon_-\langle n_p \rangle \\
+ \varepsilon_+\varepsilon_-N^2 - \varepsilon_+\varepsilon_-N
\]

\[
= (1 - \varepsilon_+ - \varepsilon_- + \varepsilon_+\varepsilon_-)(\varepsilon_+^2N^2 + \varepsilon_+(1 - \varepsilon_p)N) \\
+ (\varepsilon_+N + \varepsilon_-N - 2\varepsilon_+\varepsilon_-N + \varepsilon_+\varepsilon_-)\langle n_p \rangle \\
+ \varepsilon_+\varepsilon_-N^2 - \varepsilon_+\varepsilon_-N
\]

\[
= \varepsilon_+^2\varepsilon_+\varepsilon_- - \varepsilon_+\varepsilon_- + \varepsilon_+\varepsilon_-N + \varepsilon_+\varepsilon_-(N^2 - N) + \varepsilon_+\varepsilon_-N
\]

\[
= (\varepsilon_+ + \varepsilon_-)(1 - \varepsilon_p)\langle (\varepsilon_+ + \varepsilon_-)(1 - \varepsilon_p) \rangle \\
(N^2 - N) + \varepsilon_pN
\]
Let us compare the background to the product of the average track multiplicities. For a fixed $n_p$:

$$
\langle n_+ \rangle = \sum_{n_p=1}^{N-n_p} (n_p + n_+) \omega(n_p)
= n_p + \langle n_+ \rangle
= n_p + \varepsilon_+(N - n_p)
$$

(A12)

averaged over all possible $n_p$:

$$
\langle N_+ \rangle = \sum_{n_p=0}^{N} \langle n_+ \rangle \omega(n_p)
= \varepsilon_p N + \varepsilon_+ N - \varepsilon_+ \varepsilon_p N
= (\varepsilon_p + \varepsilon_+(1 - \varepsilon_p)) N
$$

(A13)

or averaged over all possible $N$:

$$
\langle FG_+ \rangle = \sum_{N} \langle N_+ \rangle P(N)
= (\varepsilon_p + \varepsilon_+(1 - \varepsilon_p)) \langle N \rangle
$$

(A14)

and thus:

$$
\langle FG_+ \rangle \langle FG_- \rangle = (\varepsilon_p + \varepsilon_+(1 - \varepsilon_p))(\varepsilon_p + \varepsilon_-(1 - \varepsilon_p)) \langle N \rangle^2
$$

(A15)

or

$$
\frac{\langle BG_+ \rangle}{\langle FG_+ \rangle \langle FG_- \rangle} = 1 + \frac{\sigma^2 - \langle N \rangle}{\langle N \rangle^2}.
$$

(A16)

So in general $\langle BG_+ \rangle \neq \langle FG_+ \rangle \langle FG_- \rangle$, except for the special case that $P(N)$ is a Poisson distribution. Note this is the opposite conclusion one derives in the case that the sources of + and − tracks are independent, i.e., + and − tracks are produced as singles and not as pairs as in the case of muons. In that case $\langle FG_+ \rangle \langle FG_- \rangle$ is the correct background normalization.

APPENDIX B: RELATION BETWEEN REAL PHOTONS, VIRTUAL PHOTONS, AND ELECTRON PAIRS

1. Introduction

Figure 49 illustrates that, in general, any source of high energy real photons can also emit virtual photons which materialize into electron pairs. On the left side a real photon is emitted by a source labeled as $M(Q^2 = 0)$. On the right side is an analogous diagram, where a virtual photon with mass $m_{\gamma'}$ is emitted. The virtual photon can then convert to an $e^+e^-$ pair if $m_{\gamma'} > 2m_e$. This $e^+e^-$ pair production process is a QED correction to the real photon production process and is often called internal conversion.

In the energy region, where electroweak effects are negligible, an electron pair can only be produced through a virtual photon. (Here we don’t include $e^+e^-$ pairs from correlated weak decays such as $\bar{c}c \rightarrow e^+e^- \gamma$.) Thus any electron pair production process can be described as production of a virtual photon and its subsequent decay into an $e^+e^-$ pair.

The yields of virtual photons $dN_{\gamma^*}$ and electron pairs $dN_{ee}$ are related:

$$
\frac{d^2 N_{ee}}{dM^2} = \frac{\alpha L(M)}{3\pi M^2} dN_{\gamma^*},
$$

(B1)

$$
L(M) = \sqrt{1 - \frac{4m_e^2}{M^2}} \left(1 + \frac{2m_e^2}{M^2}\right),
$$

(B2)

where $M$ is the mass of the virtual photon or the electron pair ($M = m_{\gamma^*} = m_{ee}$) and $\alpha$ is the fine structure constant ($\alpha \approx 1/137$). The factor, $\frac{\alpha L(M)}{3\pi M^2}$, is a universal factor describing the decay of the virtual photon into an $e^+e^-$ pair. This relation is exact to first order in the electromagnetic coupling $\alpha$.

Equation (B1) can be written as

$$
\frac{d^2 N_{ee}}{dM^2} = \frac{\alpha L(M)}{3\pi M^2} S(M, q) dN_{\gamma^*}.
$$

(B3)

Here we have introduced $S(M, q) = dN_{\gamma^*}(M)/dN_{\gamma}$ to factor out the difference between real photon emission and virtual photon emission. The factor $S(M, q)$ is process dependent and accounts for effects such as form factors, phase space, and spectral functions. $S(M, q)$ approaches 1 for small $M$, $S(M, q) \rightarrow 1$ for $M \rightarrow 0$. Additionally, since $L(M) \simeq 1 - 6m_e^2/4M^2$ for $m_e \ll M$, $L(M) = 1$ is a very good approximation. Thus the relationship between the electron pair yield and the direct photon yield simplifies to

$$
\frac{d^2 N_{ee}}{dM^2} \simeq \frac{\alpha}{3\pi M^2} \frac{1}{4M^2} dN_{\gamma^*},
$$

(B4)

$$
\frac{d^2 N_{ee}}{dM^2} \simeq \frac{2\alpha}{3\pi M} dN_{\gamma^*}.
$$

(B5)

The relation between real photon production and electron pair production shown by Equation (B3) is valid if $M \ll E_\gamma$, i.e., if the virtual photon is quasi-real. In this region, the yield of electron pairs in the mass range
\( m_1 < M < m_2 \) is related to the photon yield as

\[
dN_{ee}(m_1 < M < m_2) \simeq \int_{m_1}^{m_2} \frac{2\alpha}{3\pi M} dN_\gamma \quad \text{(B6)}
\]

\[ \simeq \frac{2\alpha}{3\pi} \log \frac{m_2}{m_1} dN_\gamma. \quad \text{(B7)} \]

In the following, we discuss two examples of internal conversions: Dalitz decays and high \( p_T \) Drell-Yan production. We then discuss the relationship between direct photons and electron pairs from thermal sources. At the end, we illustrate the relationship between electron pairs and virtual photons using a theoretical model calculation.

2. Dalitz decays

Dalitz decays of pseudo-scalar and vector mesons are prime examples of internal conversion. In these processes, a virtual photon, instead of a real photon, is emitted in the decay of a hadron and subsequently decays into an \( e^+e^- \) pair. The relation between \( A \rightarrow Be^+e^- \) and \( A \rightarrow B\gamma \) is given by 74

\[ \frac{d\Gamma(A \rightarrow Be^+e^-)}{dM^2} = \frac{\alpha}{3\pi M} S_{AB}(M) \Gamma(A \rightarrow B\gamma), \quad \text{(B8)} \]

where \( m_A \) and \( m_B \) are the mass of hadrons \( A \) and \( B \), and \( F_{AB} \) is the electromagnetic transition form factor. \( S_{AB}(M) \) here is an example of \( S(M, q) \) in Equation B3.

For \( S(M, q) \) one can write the relation between the photonic decay \( (P \rightarrow \gamma\gamma) \) and the corresponding Dalitz decay \( (P \rightarrow e^+e^-\gamma) \) is given by the well known Kroll-Wada formula 72, 74.

\[ \frac{d\Gamma(P \rightarrow e^+e^-\gamma)}{dM^2} = \frac{2\alpha}{3\pi M} S_{KW}(M) \Gamma(P \rightarrow \gamma\gamma), \quad \text{(B9)} \]

\[ S_{KW}(M) = |F_p(M^2)|^2 \left(1 - \frac{M^2}{m_p^2}\right)^3, \quad \text{(B10)} \]

where \( m_p \) is the meson mass and \( F_p(M^2) \) is the electromagnetic form factor. Note that the factor 2 in \( \frac{2\alpha}{3\pi} \) accounts for the fact that each of the two decay photons can convert to an electron pair. The form factor is usually parameterized as \( F_p(Q^2) = 1/(1 - Q^2/A_p^2) \). Experimental measurements of the transition form factor by Lepton-G 73 and Cello 95 show \( \Delta_p \simeq M_p \), consistent with the vector meson dominance model (VDM).

3. High \( p_T \) Drell-Yan process

In \( p + p \) collisions, the cross section for Drell-Yan electron pair production can be expressed in terms of the cross section for virtual photon production as follows 60:

\[ \frac{d^2\sigma_{pp-i\to X}}{dM^2 dp_T^2 dy} = \frac{\alpha}{3\pi} \frac{L(M)}{M^2} \frac{d^2\sigma_{pp-i\to X}}{dp_T^2 dy}. \quad \text{(B11)} \]

Where \( M, p_T, \) and \( y \) are the mass, the transverse momentum, and the rapidity of the virtual photon. For \( p_T \gg M \), the virtual photon cross section becomes equal to the real photon cross section \( (d\sigma_{\gamma\gamma} \rightarrow d\sigma_\gamma \rightarrow \gamma\gamma \rightarrow e^+e^-) \). Thus the electron pair cross section can be described as

\[ \frac{d\sigma_{pp-i\to X}}{dM^2 dp_T^2 dy} \simeq \frac{\alpha}{3\pi} \frac{L(M)}{M^2} \frac{d\sigma_{pp-i\to X}}{dp_T^2 dy}. \quad \text{(B12)} \]

Direct photon production via gluon-Compton scattering \((q + g \rightarrow q + \gamma)\) has an associated electron pair production process \((q + g \rightarrow q + \gamma^* \rightarrow q + e^+e^-)\). For the lowest order pQCD calculation, the following relation between the two processes is obtained by an explicit calculation.

\[ \frac{d^2\sigma_{ee}}{dM^2 dt} = \frac{\alpha}{3\pi} \frac{L(M)}{M^2} \frac{d\sigma_\gamma \times S_{qq}(u, t, s)}{dt}, \quad \text{(B13)} \]

where

\[ S_{qq}(u, t, s) = (1 + \frac{2u}{t^2 + s^2} M^2) \]

\[ = 1 - \frac{2x}{(x + \sqrt{1 + x^2})(3x^2 + 1 + 2x\sqrt{1 + x^2})}. \quad \text{(B14)} \]

Here \( x = p_T/M \) and \( s, t, u \) are the Mandelstam variables defined as \( s = (p + k)^2 \), \( t = (p - k')^2 \), \( u = (p - p')^2 \). \( k, k', p' \) are 4-momenta of the incoming quark, the incoming gluon, the outgoing (virtual or real) photon, and the outgoing quark, respectively. The factor \( S_{qq}(u, t, s) \) accounts for the difference between the virtual photon cross section and the real photon cross section, and \( S_{qq}(u, t, s) \) becomes unity for small \( M \). It is an example of \( S(M, q) \) in Equation B3.

4. Thermal radiation

In heavy ion collisions, thermal radiation from the hot and dense matter formed in the collision may contribute to both real direct photon production and electron pair production. The emission rate of electron pairs per space-time volume from a thermal source can be described...
in terms of the electromagnetic (EM) spectral function as \[ dR_{ee} \frac{d^4 q}{d^3 q} = -\frac{\alpha^2}{3\pi^3} \frac{L(M)}{M^2} \text{Im} \Pi_{em,\mu}(M, q; T) f^B(q_0, T), \] (B15)

Here \( \Pi_{em}^T \) and \( \Pi_{em}^L \) are the transverse and the longitudinal components of the in-medium photon self-energy tensor \( \Pi_{em,\mu} \), and \( f^B(q_0, T) = 1/(e^{q_0/T} - 1) \) is the Boltzmann factor.

Using the same notation, the emission rate of virtual photons is described as \[ q_0 \frac{dR_{\gamma^*}}{d^3 q} = -\frac{\alpha}{2\pi^2} \text{Im} \Pi_{em,\mu}(M, q; T) f^B(q_0, T). \] (B17)

The real photon emission rate is obtained in the limit of \( M \to 0 \). The longitudinal polarization contribution \( \Pi_{em}^L \) vanishes for real photons: \( \Pi_{em}^T(M, q; T) \to \Pi_{em}^L(0, q; T), \Pi_{em}^L(M, q; T) \to 0 \) for \( M \to 0 \). The real photon emission rate is given by

\[ q_0 \frac{dR_{\gamma}}{d^3 q} = -\frac{\alpha}{2\pi^2} \text{Im} \Pi_{em}^T(M = 0, q; T) f^B(q_0, T). \] (B18)

The virtual photon and electron pair emission rates are thus related by

\[ q_0 \frac{dR_{ee}}{dM^2 d^3 q} = \frac{1}{2} \frac{dR_{\gamma}}{d^3 q}, \]

\[ = \frac{\alpha}{3\pi} \frac{L(M)}{M^2} q_0 \frac{dR_{\gamma^*}}{d^3 q}, \] (B19)

\[ q_0 \frac{dR_{ee}}{dM^2 d^3 q} \approx \frac{\alpha}{3\pi} \frac{L(M)}{M^2} q_0 \frac{dR_{\gamma}}{d^3 q}. \] (B20)

Note that Equation (B14) describing the relationship between the virtual photon and electron pair emission, is exact to order \( \alpha \) in QED and is exact to all orders of the strong coupling constant. Note also that this relationship between electron pairs and photons is the same as that shown in Equation (B11). This reflects the fact that an \( e^+e^- \) pair can only be produced through virtual photon and that the conversion rate of a virtual photon into an \( e^+e^- \) pair is described by a universal factor, \( \frac{\alpha L(M)}{3\pi M^2} \). Equation (B24) is equivalent to Equation (B12) and is an approximation for small \( M \), where \( dR_e \approx dR_{\gamma} \).

The relations above are for the emission rates per space-time volume. The yields \( dN_{ee} \) and \( dN_{\gamma} \) are obtained from space-time integral over the rates.

\[ \frac{d^2 N_{ee}}{dM^2} = \frac{\alpha}{3\pi} \frac{L(M)}{M^2} dN_{\gamma^*}, \] (B21)

\[ \frac{d^2 N_{ee}}{dM} = \frac{2\alpha}{3\pi} \frac{L(M)}{M} dN_{\gamma^*}, \] (B22)

\[ = \frac{2\alpha}{3\pi} \frac{L(M)}{M} S(M, q) dN_{\gamma}. \] (B23)

These are the same equations as Equation (B1) – Equation (B3). Here, following Equation (B3) the difference between real and virtual photons is factored out in \( S(M, q) \). The \( S(M, q) \) factor for radiation from thermal sources can be written as

\[ S(M, q) = \frac{\langle \text{Im}(2\Pi_{em}^T(M, q) + \Pi_{em}^L(M, q)) f^B(q_0) \rangle}{\langle \text{Im}(2\Pi_{em}^T(0, q) f^B(q_0; M = 0) \rangle} \] (B24)

Here \( \langle \rangle \) indicates the space-time average.

Deviation of \( S(M, q) \) from unity can arise from non-zero \( \Pi_{em}^L(M, q) \) in the medium and a change of \( \Pi_{em}^T(M, q) \) from \( \Pi_{em}^L(0, q) \). The behavior of \( \Pi_{em}^T \) and \( \Pi_{em}^L \) is model-dependent. However, on very general grounds we can conclude that the in-medium spectral function \( \Pi_{em} \) is a smooth function of \( M \) for the low-mass region \( (M < \text{a few} 100 \text{ MeV}/c^2) \).

Hadronic interactions yield and almost flat behavior in \( S(M) \) as we see in a model calculation by Rapp later. \( q+g \) scattering gives an almost flat contribution in \( S(M) \); see \( S_{qg} \) of gluon Compton scattering. As we see later, the contribution from \( qg \) annihilation is not constant and is \( \propto M^2 \) (see Equation (B27)). This means that it is strongly suppressed in the low-mass region relative to hadronic and \( qg \) scattering contributions. Thus \( S(M, q) \) is a smooth and almost constant function of \( M \) at low masses. Furthermore, the short time \( (\leq 1 \text{ fm}/c) \) between rescattering among hadrons and partons in the medium should smear any feature in \( S(M) \) smaller than a mass scale of a few 100 MeV/c².

5. Conversion of \( e^+e^- \) pairs to virtual photons

Equation (B23) can be used to determine the virtual photon yield from the electron pair yield. The equation can be rewritten as

\[ q_0 \frac{dN_{\gamma^*}}{d^3 q} = S(M, q)q_0 \frac{dN_{ee}}{d^3 q}, \] (B25)

\[ \approx \frac{3\pi}{2\alpha} M q_0 \frac{dN_{ee}}{d^3 q} q dM. \] (B26)

Using this relationship, the virtual photon spectrum as a function of mass \( M \) and \( p_T \) can be determined from the double differential electron pair spectrum. In particular, the mass dependence for a given \( p_T \) bin can be measured. This is a direct measurement of the shape of the \( S(M, q) \) function and the shape of \( \langle \text{Im}(2\Pi_{em}^T(M, T) f^B(T) \rangle \) as a function of \( M \). The real photon yield is then obtained by extrapolation of \( dN_{\gamma^*}(M, T) \) to the \( M \to 0 \) limit.

6. Electron pairs and virtual photons in a theoretical model calculation

Figure (a) illustrates the relationship between the electron pair mass spectrum and the virtual photon cross section. Figure (a) shows the double differential
the low-mass region, the mass spectrum steeply increases out the low-mass regions below two-pion threshold. In

\[ q \rightarrow \rho \]

which is augmented with a

\[ q \rightarrow \rho \rightarrow \pi^+ \pi^- e^+ e^- \]

electron pair spectrum, \( (1/pt)dN_{ee}/dMdp_Tdy \) at \( pt = 1.525 \text{ GeV/c} \) from a model calculation of electron pair production by Rapp \[83\]. The dashed and solid curve show electron pairs from the hadronic gas. The calculation shown in the dashed curve uses the spectral function \( \Pi_{em} \) that is unchanged from its vacuum value so the line shapes of vector mesons (\( \rho, \omega, \phi \)) are unmodified. The calculation shown in the solid curve, uses a spectral function calculated by a model Lagrangian of hadronic many-body interactions, and the line shapes of vector mesons are broadened due to the interactions. It also includes the contributions like \( a_1(1260) \rightarrow \pi^+ e^- e^- \), \( \rho \rightarrow \pi^+ e^- e^- \), and \( N + \pi \rightarrow N^* \rightarrow Ne^+ e^- \) that fill out the low-mass regions below two-pion threshold. In the low-mass region, the mass spectrum steeply increases with decreasing \( M \). This steep behavior is due to the \( 1/M \) behavior in \( \gamma^* \rightarrow e^+ e^- \). The dotted curve shows the contribution from the leading order (LO) \( q\bar{q} \) annihilation which is augmented with a \( q = 0 \) Hard Thermal Loop (HTL) correction factor \[124\].

The solid curve shown in Fig. 50(b) shows the same calculations presented as the differential yield of virtual photons as a function of mass. The electron pair yield shown in Fig. 50(a) is converted to the virtual photon yield using Equation B26. In this plot, the solid curve becomes almost constant for \( M < 0.3 \text{ GeV} \). The steep \( 1/M \) behavior of the electron pair spectrum is removed, and much more smooth behavior of the virtual photon spectrum is revealed. The plot shows that the virtual photon yield is almost constant as function of \( M \). The value of the solid curve at \( M = 0 \) corresponds to the real photon yield. Thus in this model calculation, the yield of virtual photons is a smooth function of \( M \) and it becomes the real photon yield in the limit of \( M = 0 \). \( dN_{ee}^*(M, pt) \rightarrow dN_e(pt) \) as \( M \rightarrow 0 \).

The flat behavior of the solid curve in this low-mass region comes from the fact that hadronic scattering processes such as \( a_1(1260) \rightarrow \pi^+ e^- e^- \), \( \rho \rightarrow \pi^+ e^- e^- \), and \( N + \pi \rightarrow N^* \rightarrow Ne^+ e^- \) dominate this low-mass region. These processes are internal conversion of the corresponding real photon production processes, i.e., \( a_1 \rightarrow \pi^+ \gamma, \rho \rightarrow \pi^+ \gamma, N + \pi \rightarrow N + \gamma \). Thus production of virtual photons from these processes should be very close to that of real photons at low mass. In the language of the spectral function \( \Pi_{em} \), these processes contribute to \( \Pi_{em} \) at \( M = 0 \) as well as \( M > 0 \). Their contribution to \( \Pi_{em} \) is a smooth, almost a constant function of virtual photon mass \( M \).

In Fig. 50(b), the contribution from \( q\bar{q} \) annihilation is shown by the dotted curve. In perturbation theory, the \( q\bar{q} \) contribution is given by

\[
\text{Im}\Pi_{em}(M) = \frac{M^2}{12\pi}(1 + \frac{\alpha_s}{\pi} + ...)N_e\Sigma(e_q)^2. \tag{B27}
\]

The quark annihilation contribution behaves as \( \propto M^2 \) in the low-mass region. Thus it is strongly suppressed and has little contribution in the low-mass region. In the high-mass region, the \( M^2 \) behavior of the quark annihilation is suppressed by the Boltzmann factor.

There is a large uncertainty in the approximation used in the dotted curve, and it is shown here just to illustrate that the \( q\bar{q} \) annihilation contribution is sub-leading contribution in the low-mass region (\( M < \text{ a few } 100 \text{ MeV/c}^2 \)). Many effects can alter the shape and magnitude of the \( q\bar{q} \) contribution. The calculation shown in Fig. 50 uses zero quark mass. The effective quark mass in the medium is uncertain, but in most theoretical calculations it is the order of the temperature of the medium. Quarks in the medium should also have a large...
width (more than a few 100 MeV/$c^2$) due to the short time between interactions. A non-zero effective quark mass would further suppress the quark annihilation contribution in the very low-mass region while a finite quark width would smear the $M^2$ behavior and can cause non-zero $q\bar{q}$ contribution at $M = 0$.

It should be noted that the dotted curve does not include processes like $q + g \rightarrow q + \gamma^*$ that are associated with real direct photon production. This is because HTL calculation of thermal radiation from QGP is only available in the real photon case. Contributions from processes associated with real photon production can be much larger than those from the LO $q\bar{q}$ annihilation shown in Fig. 50 in the low-mass region. Turbide, Gale, and Rapp [40] calculated real photon production in a hadronic gas using the same model and compared it with real photon production in the QGP phase using the complete leading order HTL analysis. They found that real photons from the QGP outshine those from a hadronic gas for $p_T > 1.5$ GeV/$c$ in Au + Au collisions at RHIC. Since the virtual photon yield should be continuous from $M = 0$ (i.e. real photons) to $M > 0$, this implies that the contribution from the QGP, including processes like $q + g \rightarrow q + \gamma^*$, is as large as or even greater than that shown in the solid curve in Fig. 50(b) in the low-mass and high-$p_T$ region (i.e. $M < 1$ a few 100 MeV/$c^2$ and $p_T > 1.5$ GeV/$c^2$) at RHIC energies.

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