Electroweak Radiative Corrections to $p\bar{p} \rightarrow W^{\pm} \rightarrow \ell^{\pm} \nu$

Beyond the Pole Approximation

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Abstract

We present a calculation of the complete electroweak $O(\alpha)$ corrections to $p\bar{p} \rightarrow W^{\pm} \rightarrow \ell^{\pm} \nu X$ ($\ell = e, \mu$) in the Standard Model of electroweak interactions, focusing on those corrections which do not contribute in the pole approximation. We study in detail the effect of these corrections on the transverse mass distribution, the $W$-width measurement, and the transverse mass ratio and cross section ratio of $W$ and $Z$ bosons.

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I. INTRODUCTION

The Standard Model of electroweak interactions (SM) so far withstood all experimental challenges and is tested as a quantum field theory at the 0.1% level [1]. However, the mechanism of mass generation in the SM predicts the existence of a Higgs boson which, so far, has eluded direct observation. Direct searches at LEP2 give a 95% confidence-level lower bound on the mass of the SM Higgs boson of $M_H > 114.4$ GeV [2]. Indirect information on the mass of the Higgs boson can be extracted from the $M_H$ dependence of radiative corrections to the $W$ boson mass. With the present knowledge of the $W$ boson and top quark masses [1,3], and the electromagnetic coupling constant, $\alpha(M_Z^2)$ [4], the SM Higgs boson mass can be indirectly constrained to $M_H = 113^{+62}_{-42}$ GeV [1,5] by a global fit to all electroweak precision data. Future more precise measurements of the $W$ boson and top quark masses are expected to considerably improve the present indirect bound on $M_H$: with a precision of 27 MeV for the $W$ boson mass, $M_W$, and 2.7 GeV for the top quark mass, which are target values for the expected integrated luminosity of 2 fb$^{-1}$ of Run II of the Tevatron, $M_H$ can be predicted with an uncertainty of about 35% [6]. In addition, the confrontation of a precisely measured $W$ boson mass with the indirect SM prediction from a global fit to all electroweak precision data, $M_W = 80.386 \pm 0.023$ GeV [1] will provide a stringent test of the SM. A precise measurement of the $W$ width, $\Gamma_W$, and comparison with the SM prediction, will help to further scrutinize the SM. The $W$ width is expected to be measured in Run II with a precision of about $25 - 30$ MeV from the high transverse mass tail when data from both lepton channels and both experiments are combined [7]. It can also be determined indirectly from the cross section ratio [8,9],

$$R_{W/Z} = \frac{\sigma(pp \to W \to \ell\nu X)}{\sigma(pp \to Z \to \ell^+\ell^- X)}, \quad (1)$$

together with the theoretical prediction for the ratio of the total $W$ and $Z$ production cross sections, the LEP measurement of the branching ratio $B(Z \to \ell^+\ell^-)$ [1], and the SM prediction for the $W \to \ell\nu$ decay width.

In hadronic collisions, the $W$ mass is usually determined from the transverse mass distribution of the lepton and neutrino which originate from the $W$ decay, $W \to \ell\nu$. The experimental uncertainty on $M_W$ strongly depends on an accurate measurement of the transverse momentum of the neutrino which requires that the transverse momentum distribution of the $W$ is well understood. In lowest order, the $W$ is produced without any transverse momentum. Only when QCD corrections are taken into account does the $W$ acquire a non-zero transverse momentum, $p_W^T$. For a detailed understanding of the $p_W^T$ distribution, it is necessary to resum the soft gluon emission terms [10], and to model non-perturbative QCD corrections [11]. In addition to QCD corrections, electroweak (EW) radiative corrections play an important role in the $W$ mass and width measurement; final state photon radiation is known to shift both quantities by $\mathcal{O}(100$ MeV) [12–17]. In order to measure the $W$ mass and width with high precision at a hadron collider, it is thus necessary to fully understand and control higher order QCD and electroweak corrections. In the last few years, significant progress in our understanding of the EW corrections to $W$ boson production in hadronic collisions has been made. A calculation of the $\mathcal{O}(\alpha)$ EW corrections to $pp \to W^\pm \to \ell^\pm\nu$ ($\ell = e, \mu$) in the pole approximation was presented in Ref. [18]. The complete $\mathcal{O}(\alpha)$ corrections were calculated in Ref. [19]. Two photon radiation in $W$ production and decay
has been calculated in Ref. [20]. Finally, first steps towards going beyond fixed order in EW radiative corrections were taken in Refs. [21] and [22] by including the effects of final state multiphoton radiation in $W$ production, and in Ref. [23] where final state photon radiation was added to a calculation of $W$ boson production which includes resummed QCD corrections.

In this paper we present an independent calculation of the complete $\mathcal{O}(\alpha)$ EW radiative corrections to $p \bar{p} \rightarrow W^{\pm} \rightarrow \ell^{\pm} \nu$ using the methods developed in Ref. [24], and examine how the non-resonant corrections neglected in Ref. [18] affect several observables of interest, in particular the measurement of the width of the $W$ boson. Preliminary results of our calculation were reported in Ref. [25]. For the numerical evaluation, we use the Monte Carlo phase space slicing method for next-to-leading-order (NLO) calculations described in Ref. [26]. With the Monte Carlo method, it is easy to calculate a variety of observables simultaneously and to simulate detector response. The collinear singularities associated with initial state photon radiation are removed by universal collinear counter terms generated by “renormalizing” the parton distribution functions (PDFs) [18,27–30], in complete analogy to gluon emission in QCD. Final state charged lepton mass effects are included in our calculation in the following approximation. The lepton mass regularizes the collinear singularity associated with final state photon radiation. The associated mass singular logarithms of the form $\ln(\hat{s}/m_\ell^2)$, where $\hat{s}$ is the squared parton center of mass energy and $m_\ell$ is the charged lepton mass, are included in our calculation, but the very small terms of $\mathcal{O}(m_\ell^2/\hat{s})$ are neglected.

The technical details of our calculation are described in Sec. II. The electroweak $\mathcal{O}(\alpha)$ corrections consist of the electroweak one-loop contributions, including virtual photons, and of the emission of a real photon. To regularize the ultraviolet divergences associated with the virtual corrections, we use dimensional regularization in the on-shell renormalization scheme [31]. The non-resonant corrections are small in the $W$ pole region, but become important at high $\ell\nu$ invariant masses, $m(\ell\nu)$, due to the presence of large Sudakov-like electroweak logarithms of the form $(\alpha/\pi) \ln^2(m(\ell\nu)/M_V)$ ($V = W, Z$) [32].

Numerical results for the Tevatron ($p\bar{p}$ collisions at $\sqrt{s} = 2$ TeV) and the CERN Large Hadron collider (LHC, $pp$ collisions at $\sqrt{s} = 14$ TeV) are presented in Sec. III. For $\ell\nu$ transverse masses above the $W$ peak region, the non-resonant corrections reduce the differential cross section by $\mathcal{O}(10\%)$. We study in detail how these corrections affect the measurement of the $W$ width from the tail of the transverse mass distribution. Using the results of Ref. [33], we also consider how the ratio of the $W$ and $Z$ cross sections, $R_{W/Z}$ (see Eq. (1)), and the transverse mass ratio of $W$ and $Z$ bosons are influenced. Finally, our conclusions are presented in Sec. IV.

II. THE $\mathcal{O}(\alpha)$ ELECTROWEAK RADIATIVE CORRECTIONS TO W PRODUCTION: RESONANT AND NON-RESONANT CONTRIBUTIONS

The complete $\mathcal{O}(\alpha^3)$ parton level cross section of $W$ production via the Drell-Yan mechanism $q\bar{q}' \rightarrow f\bar{f}'(\gamma)$ is given by

$$d\hat{\sigma}^{(0+1)}(\hat{s}, \hat{t}) = d\hat{\sigma}^{(0)} + d\hat{\sigma}_{\text{virt}} + \sum_{a=\text{initial, final, interf.}} [d\hat{\sigma}^{(0)} F_{BR}^a + d\hat{\sigma}_{2\rightarrow3}^a],$$  \hspace{1cm} (2)
where the Born cross section, $d\hat{\sigma}^{(0)}$, is of Breit-Wigner form and $\hat{s}$ and $\hat{t}$ are the usual Mandelstam variables in the parton center of mass frame. The corresponding Feynman diagrams are shown in Fig. 1 and Fig. 2. $d\hat{\sigma}_{\text{virt}}$ describes the complete set of virtual $\mathcal{O}(\alpha)$ contributions consisting of the pure weak and the photonic virtual corrections:

$$d\hat{\sigma}_{\text{virt}}(\hat{s}, \hat{t}) = d\hat{\sigma}^{(0)} 2\Re[F_{\text{weak}}(\hat{s}) + F_{\gamma}(\hat{s}, \hat{t})] + d\hat{\sigma}_{WZbox}(\hat{s}, \hat{t}).$$

(3)
Explicit expressions for the pure weak form factor $F_{\text{weak}}$ and for the contribution of the $W, Z$ box diagrams, $d\hat{\sigma}_{WZ\text{box}}$, are given in the Appendix. The form factors $F_{BR}^a$ of Eq. (2) describing the initial, final state and interference contribution of real soft photon radiation, and the photonic form factor $F_{\gamma}$ can be found in Ref. [24] (Appendix D).

The soft photon region is defined by requiring that the photon energy in the parton center of mass frame, $\hat{E}_{\gamma}$, is smaller than a cutoff $\Delta E = \delta_s \sqrt{\hat{s}}/2$. In this phase space region, the soft photon approximation can be used to calculate the cross section as long as $\delta_s$ is sufficiently small. Throughout the calculation the soft singularities have been regularized by giving the photon a fictitious mass. As usual, the unphysical photon mass dependence cancels in the sum of the virtual and soft photon terms, $2\text{Re}F_{\gamma} + \sum_a F_{BR}^a$. The IR finite contribution $d\hat{\sigma}_{2\rightarrow 3}^a$ describes real photon radiation with $\hat{E}_{\gamma} > \Delta E$. The superscript $a$ denotes the initial state, final state or interference contributions. Throughout the calculation of the $O(\alpha)$ corrections we consider the fermions as being massless. We retain finite fermion masses only to regularize the collinear singularities which arise when the photon is emitted collinear with one of the charged fermions. Thus, $d\hat{\sigma}_{2\rightarrow 3}^a$ and the form factors $F_{BR}^a$ and $F_{\gamma}$ contain large mass singular logarithms which have to be treated with special care. For final state photon radiation, the collinear singularity is regularized by the finite lepton mass. In sufficiently inclusive measurements, the mass singular logarithmic terms originating from the collinear singularity cancel [35]. For initial state photonic corrections, however, the mass singular logarithms always survive. These singularities are universal to all orders in perturbation theory and can be absorbed by a redefinition (renormalization) of the PDFs [18,27]. This can be done in complete analogy to the calculation of QCD radiative corrections. As a result, the renormalized parton distribution functions become dependent on the QED factorization scale $\mu_{\text{QED}}$ which is controlled by the well-known Gribov-Lipatov-Altarelli-Parisi (GLAP) equations [36]. These universal photonic corrections can be taken into account by a straightforward modification [28–30] of the standard GLAP evolution equations.

The QED induced terms in the GLAP equations lead to small corrections at the permille level to the distribution functions for most values of $x$ and $\mu_{\text{QED}}^2$ [30,37]. Only at large $x \gtrsim 0.5$ and large $\mu_{\text{QED}}^2 \gtrsim 10^3 \text{ GeV}^2$ do the corrections reach the magnitude of one per cent.

In order to treat the $O(\alpha)$ initial state photonic corrections to $W$ production in hadronic collisions in a consistent way, QED corrections should be incorporated in the global fitting of the PDFs, i.e. all data which are used to fit the parton distribution functions should be corrected for QED effects. Current fits [41,42] to the PDFs do not include QED corrections. The missing QED corrections introduce an uncertainty which, however, is probably much
smaller than the present experimental uncertainties on the parton distribution functions.

Absorbing the collinear singularity into the PDFs introduces a QED factorization scheme dependence. The squared matrix elements for different QED factorization schemes differ by the finite $\mathcal{O}(\alpha)$ terms which are absorbed into the PDFs in addition to the singular terms. Our calculation has been carried out both in the QED $\overline{\text{MS}}$ and DIS schemes, which are defined analogously to the usual $\overline{\text{MS}}$ [43] and DIS [44] schemes used in QCD calculations. All numerical calculations in this paper are performed using the QED DIS scheme. The QED DIS scheme is defined by requiring the same expression for the leading and next-to-leading order structure function $F_2$ in deep inelastic scattering, which is given by the sum of the quark distributions. Since $\tilde{F}_2$ data are an important ingredient in extracting PDFs, the effect of the $\mathcal{O}(\alpha)$ QED corrections on the PDFs should be reduced in the QED DIS scheme.

In the vicinity of the $W$ resonance, the $W, Z$ box diagrams of Fig. 2 can be neglected as non-resonant contributions of higher order in perturbation theory, and, as demonstrated in [24], a gauge invariant decomposition of the complete $\mathcal{O}(\alpha)$ contribution into a QED-like and a modified weak part can be performed. Unlike the $Z$ boson case, the Feynman diagrams of Fig. 1 which involve a virtual photon do not represent a gauge invariant subset. In Ref. [24], it was demonstrated that gauge invariant contributions can be extracted from the infrared (IR) singular virtual photonic corrections, $\tilde{F}_a^{\text{YFS}}$ (the modified Yennie-Frautschi-Suura form factors). These contributions can be combined with the real photon corrections in the soft photon region, $F_{\text{BR}}^a$, to form gauge invariant QED-like contributions corresponding to initial state, final state and interference corrections. The IR finite remainder of the virtual photonic corrections and the pure weak one-loop corrections of Fig. 1 can be combined to separately gauge invariant modified weak contributions to the $W$ boson production and decay processes.

Both the QED-like and the modified weak contributions can be expressed in terms of form factors, $F_{\text{QED}}^a$ and $\tilde{F}_{\text{weak}}^a$, which multiply the Born cross section [24]. Thus, the complete $\mathcal{O}(\alpha^3)$ parton level cross section of resonant $W$ production via the Drell-Yan mechanism $q_i \bar{q}_{i'} \rightarrow ff'(\gamma)$ can be expressed in the form

$$d\hat{\sigma}_{\text{res}}^{(0+1)} = d\hat{\sigma}^{(0)} \left[ 1 + 2 \Re (\tilde{F}_{\text{initial}} + \tilde{F}_{\text{final}}^\text{weak}) (M_W^2) \right]$$

$$+ \sum_{a=\text{initial,final,interf.}} \left[ d\hat{\sigma}^{(0)} F_{\text{QED}}^a (\hat{s}, \hat{t}) + d\hat{\sigma}_{2\rightarrow 3}^a \right].$$  \hspace{1cm} (4)

The modified weak contributions have to be evaluated at $\hat{s} = M_W^2$. Explicit expressions for the form factors $F_{\text{QED}}^a, \tilde{F}_{\text{weak}}^a$ are given in Ref. [24].

The non-resonant part, which was neglected in Refs. [18] and [24], can then be obtained from the resonant contribution, $d\hat{\sigma}_{\text{res}}^{(0+1)}$ (see Eq. (4)), and the complete $\mathcal{O}(\alpha)$ contribution, $d\hat{\sigma}^{(0+1)}$ of Eq. (2):

$$d\hat{\sigma}_{\text{non-res}} (\hat{s}, \hat{t}) = d\hat{\sigma}^{(0+1)} - d\hat{\sigma}_{\text{res}}^{(0+1)}$$

$$= d\hat{\sigma}_{WZbox} (\hat{s}, \hat{t}) + d\hat{\sigma}^{(0)} 2 \Re \left[ \sum_{a=\text{initial,final}} (F_{\text{weak}}^a (\hat{s}) - \tilde{F}_{\text{weak}}^a (M_W^2)) \right]$$

$$+ \left( F_{\gamma} - \sum_{a=\text{initial,final,interf.}} \tilde{F}_{\text{YFS}}^a (\hat{s}, \hat{t}) \right),$$  \hspace{1cm} (5)
where we have used Eq. (A1) and $F_{QED}^a = F_{BR}^a + 2Re F_{YFS}^a (a = initial, final, interf.)$. Equation (5) shows that the non-resonant contribution consists of the $W, Z$ box contributions, the IR finite remnants of the virtual photon one-loop corrections which are not included in the YFS form factor, and the $s$-dependent parts of the pure weak vertex and self-energy one-loop corrections, which have been neglected when evaluating the pure weak form factor at $s = M_W^2$. $d\sigma_{non-res}$ is free of mass singularities (terms proportional to $\ln(s/m_f^2)$, where $m_f$ is the mass of an initial state or final state fermion) and on-shell singularities, i.e. logarithms of the form $\ln(|s - M_W^2|)$. It represents a gauge invariant subset of the complete $O(\alpha)$ contribution to the $W$ production process $q_\ell \bar{q}_{\nu} \rightarrow f f'$, and its numerical impact can thus be studied separately. In Fig. 3 we illustrate the effect of the non-resonant contribution on the total parton level cross section by showing the variation of the relative correction $\delta = \delta_{non-res}/\delta^{(0)}$ (in percent) with $\sqrt{s}$ where $\delta^{(0)}$ is the Born cross section,

$$\delta^{(0)} = \frac{\pi}{36} \alpha^2 \left(1 - \frac{M_W^2}{M_Z^2}\right)^{-2} \frac{\hat{s}}{(\hat{s} - M_W^2)^2 + (\frac{s}{M_W^2} \Gamma_W^2)^2},$$

in the $s$-dependent width scheme [24]. As can be seen, the non-resonant contribution can be neglected in the vicinity of the $W$ resonance but becomes increasingly important at large center of mass energies due to the occurrence of large, Sudakov-like logarithms of the form $(\alpha/\pi) \ln^2(s/M_W^2)$. The kink at $\sqrt{s} \approx 200$ GeV is due to the $WH$ threshold in the $W$ self-energy.

The observable $O(\alpha^3)$ cross section is obtained by convoluting the parton cross section with the quark distribution functions $f_{q/A}(x, Q^2)$ ($\hat{s} = x_1 x_2 s$) and summing over all quark flavors $q$ and $q'$,

$$d\sigma(s) = \sum_{q,q'} \int_0^1 dx_1 dx_2 \left(f_{q/A}(x_1, Q^2) f_{q'/B}(x_2, Q^2) d\sigma^{(0+1)} + (q \leftrightarrow q')\right)$$

with $(A,B) = (p,\bar{p})$ for the Tevatron and $(p,p)$ for the LHC. The parton distribution functions depend on the QCD renormalization and factorization scales, $\mu_r$ and $\mu_f$, which we choose to be equal, $\mu_r = \mu_f = Q$.

### III. PHENOMENOLOGICAL RESULTS

#### A. Preliminaries

We shall now discuss the phenomenological implications of the $O(\alpha)$ non-resonant EW corrections to $\ell\nu$ production at the Tevatron ($p\bar{p}$ collisions at $\sqrt{s} = 2$ TeV) and the LHC ($pp$ collisions at $\sqrt{s} = 14$ TeV). For the numerical evaluation we chose the following set of SM input parameters [45]:

$$G_\mu = 1.16639 \times 10^{-5} \text{ GeV}^{-2}, \quad \alpha = 1/137.0359895,$$

$$M_Z = 91.1867 \text{ GeV}, \quad \alpha_s = \alpha_s(M_W^2) = 0.121,$$

$$m_e = 0.51099907 \text{ MeV}, \quad m_\mu = 0.105658389 \text{ GeV}, \quad m_\tau = 1.777 \text{ GeV},$$

$$m_u = 0.0464 \text{ GeV}, \quad m_c = 1.5 \text{ GeV}, \quad m_t = 174 \text{ GeV},$$

$$m_d = 0.0465 \text{ GeV}, \quad m_s = 0.15 \text{ GeV}, \quad m_b = 4.7 \text{ GeV},$$

$$|V_{ud}| = |V_{cs}| = 0.97 \quad |V_{us}| = |V_{dc}| = 0.22. \quad (8)$$
FIG. 3. The relative size (in percent) of the non-resonant $\mathcal{O}(\alpha)$ corrections to the Born $u\bar{d} \to W^+ \to \ell^+\nu$ parton-level total cross section as a function of the parton center-of-mass energy, $\sqrt{s}$. The parameters used are listed in Eqs. (8) – (10).

The fermion masses only enter through loop contributions to the vector boson self energies and as regulators of the collinear singularities which arise in the calculation of the QED contribution. Non-zero light quark masses are only used in the calculation of the vector boson self energies. The light quark masses are chosen such that the value for the hadronic contribution to the photon vacuum polarization for five active flavors, $\Delta\alpha^{(5)}_{\text{had}}(M_Z^2) = 0.028$ [4], which is derived from low-energy $e^+e^-$ data with the help of dispersion relations, is recovered. $V_{ij}$ are the matrix elements of the quark mixing matrix.

The $W$ mass and the Higgs boson mass, $M_H$, are related via loop corrections. A parametrization of the $W$ mass which, for $10 \text{ GeV} < M_H < 1 \text{ TeV}$, deviates by at most 0.5 MeV from the theoretical value including the full fermionic two-loop contributions is given in Ref. [46]. Here we use the somewhat older parametrization of Ref. [47]

$$M_W = M_W^0 - 0.0581 \ln \left( \frac{M_H}{100 \text{ GeV}} \right) - 0.0078 \ln^2 \left( \frac{M_H}{100 \text{ GeV}} \right) - 0.085 \left( \frac{\alpha_s}{0.118} - 1 \right)$$
$$- 0.518 \left( \frac{\Delta\alpha^{(5)}_{\text{had}}(M_Z^2)}{0.028} - 1 \right) + 0.537 \left( \frac{m_t}{175 \text{ GeV}} \right)^2$$

(9)
with \( M_W^0 = 80.3805 \text{ GeV} \), which was used in the analysis of the LEP data. The parametrization of Eq. (9) reproduces the result of Ref. [47] to 0.2 MeV for \( 75 \text{ GeV} < M_H < 350 \text{ GeV} \). For the numerical discussion we choose

\[
M_H = 120 \text{ GeV},
\]

which is consistent with current direct [2] and indirect bounds [1,5], and work in the \( s \)-dependent width scheme. For the input parameters listed in Eq. (8) we obtain \( M_W = 80.3612 \text{ GeV} \). Higher order (irreducible) corrections connected with the \( \rho \)-parameter are taken into account in our calculation by replacing

\[
\frac{\delta M_Z^2}{M_Z^2} - \frac{\delta M_W^2}{M_W^2} \rightarrow \frac{\delta M_Z^2}{M_Z^2} - \frac{\delta M_W^2}{M_W^2} - \Delta \rho^{HO}
\]

in the renormalized \( Z \) and \( \gamma \), \( Z \) self-energies as described in Appendix B of Ref. [18]. The \( W \) width and the \( q\bar{q} \rightarrow f\bar{f}(\gamma) \) amplitude are calculated in the \( G_\mu \) scheme. In the \( G_\mu \) scheme, the fine structure constant, \( \alpha \), is replaced with

\[
\alpha \rightarrow \sqrt{2} G_\mu M_W^2 \left( 1 - \frac{M_W^2}{M_Z^2} \right),
\]

and \( M_W \) is obtained from Eq. (9). For the \( W \) width, including \( \mathcal{O}(\alpha) \) EW (see also Ref. [38]) and QCD corrections up to \( \mathcal{O}(\alpha_s^3) \) [18,39,40], we obtain \( \Gamma_W = 2.0721 \text{ GeV} \).

In the numerical results presented below, we also take into account the leading \( \mathcal{O}(\alpha^2) \) weak corrections to resonant \( W \) production which can be obtained by performing the following replacement in Eq. (4):

\[
1 + 2\Re(\tilde{F}_{\text{weak}}^{\text{initial}} + \tilde{F}_{\text{weak}}^{\text{final}})(M_W^2) \rightarrow \left( 1 + \frac{1}{2} \tilde{F}_{\text{weak}}^{\text{initial}}(M_W^2) \right) \left( 1 + \frac{1}{2} \tilde{F}_{\text{weak}}^{\text{final}}(M_W^2) \right)^2.
\]

The dominant non-photonic electroweak corrections to resonant \( W \) production can be taken into account in the effective Born approximation (EBA) where the \( W \) cross section, \( \sigma^{EBA} \), is calculated by replacing \( \alpha \) in Eq. (6) with the expression given in Eq. (12) and using the \( W \) mass obtained from Eq. (9). In the following, we shall use the EBA cross section as a reference when discussing how the \( \mathcal{O}(\alpha) \) EW corrections affect physical observables.

Following a brief comparison with the calculation of Ref. [19], we discuss the impact of the non-resonant EW corrections on the transverse mass distribution and how they affect the value of the \( W \)-width extracted from that distribution. We then consider the \( W \) production cross section, the \( W \) to \( Z \) cross section ratio and the \( W \) to \( Z \) transverse mass ratio. To compute the hadronic cross section we use the MRSR2 set of parton distribution functions [48], and take the renormalization scale, \( \mu_r \), and the QED and QCD factorization scales, \( \mu_{\text{QED}} \) and \( \mu_{\text{QCD}} \), to be \( \mu_r^2 = \mu_{\text{QED}}^2 = \mu_{\text{QCD}}^2 = M_W^2 \). The detector acceptance is simulated by imposing the following transverse momentum \( (p_T) \) and pseudo-rapidity \( (\eta) \) cuts:

\[
p_T(\ell) > 20 \text{ GeV}, \quad |\eta(\ell)| < 2.5, \quad \ell = e, \mu,
\]

\[
\not{p}_T > 20 \text{ GeV},
\]

\( p_T(\ell) > 20 \text{ GeV}, \quad |\eta(\ell)| < 2.5, \quad \ell = e, \mu, \quad \not{p}_T > 20 \text{ GeV}, \)
where $p_T$ is the missing transverse momentum originating from the neutrino. These cuts approximately model the acceptance of the CDF II [49] and DØ [50] detectors at the Tevatron, and the ATLAS [51] and CMS [52] detectors at the LHC. Uncertainties in the energy measurements of the charged leptons in the detector are simulated in the calculation by Gaussian smearing of the particle four-momentum vector with standard deviation $\sigma$ which depends on the particle type and the detector. The numerical results presented here were calculated using $\sigma$ values based on the DØ and ATLAS specifications.

The granularity of the detectors and the size of the electromagnetic showers in the calorimeter make it difficult to discriminate between electrons and photons with a small opening angle. In such cases we recombine the four-momentum vectors of the electron and photon to an effective electron four-momentum vector. The exact recombination procedure is detector dependent. For calculations performed at Tevatron energies we use a procedure similar to that used by the DØ Collaboration in Run I, requiring that the electron and photon momentum four-vectors are combined into an effective electron momentum four-vector if their separation in the pseudorapidity – azimuthal angle plane,

$$\Delta R(e, \gamma) = \sqrt{(\Delta \eta(e, \gamma))^2 + (\Delta \phi(e, \gamma))^2},$$

(16)

is $\Delta R(e, \gamma) < 0.2$. For $0.2 < \Delta R(e, \gamma) < 0.4$ events are rejected if $E_\gamma > 0.15 E_e$. Here $E_\gamma$ ($E_e$) is the energy of the photon (electron) in the laboratory frame. For events with $0.2 < \Delta R(e, \gamma) < 0.3$ and $E_\gamma < 0.15 E_e$, the electron and photon momentum four-vectors are again combined. At LHC energies, we recombine the electron and photon four-momentum vectors if $\Delta R(e, \gamma) < 0.07$, similar to the resolution expected for ATLAS [51]. Recombining the electron and photon four-momentum vectors eliminates the mass singular logarithmic terms originating from final state photon radiation and strongly reduces the size of the QED-like final state corrections [18].

Muons are identified by hits in the muon chambers and the requirement that the associated track is consistent with a minimum ionizing particle. This limits the photon energy for small muon – photon opening angles. For muons at the Tevatron, we again adopt the DØ specifications and require that the energy of the photon is $E_\gamma < 2$ GeV for $\Delta R(\mu, \gamma) < 0.2$, and $E_\gamma < 6$ GeV for $0.2 < \Delta R(\mu, \gamma) < 0.6$. At the LHC, following Ref. [51], we require the photon energy to be smaller than $E_\gamma^c = 5$ GeV if $\Delta R(\mu, \gamma) < 0.3$. The cut on the photon energy increases the size of the photonic corrections for $m(\mu\nu) > 100$ GeV [18]. For future reference, we summarize the lepton identification requirements in Table I.

We use the input parameters listed in Eq. (8) and impose the cuts and lepton identification requirements described above in all subsequent numerical simulations, unless explicitly noted otherwise.

**B. Comparison with Ref. [19]**

As mentioned before, the matrix elements of the full $\mathcal{O}(\alpha)$ EW corrections to $q\bar{q} \to \ell\nu + X$ were presented in Ref. [19], together with a discussion of how EW radiative corrections influence the $p_T$ distribution of the charged lepton, and the transverse mass distribution. In this section we compare the integrated cross sections given in Ref. [19] for different ranges in $p_T(\ell)$ at the Tevatron and LHC with the results of our calculation, using the input
TABLE I. Summary of lepton identification requirements.

|              | Tevatron                                                                 | LHC                                                                 |
|--------------|---------------------------------------------------------------------------|----------------------------------------------------------------------|
| electrons    | combine $e$ and $\gamma$ momentum four vectors if $\Delta R(e, \gamma) < 0.2$ and if $E_\gamma < 0.15 E_e$ for $0.2 < \Delta R(e, \gamma) < 0.3$ | combine $e$ and $\gamma$ momentum four vectors if $\Delta R(e, \gamma) < 0.07$ |
|              | reject events with $E_\gamma > 2 \text{ GeV}$ if $\Delta R(\mu, \gamma) < 0.2$ for $\Delta R(e, \gamma) < 0.4$ | reject events with $E_\gamma > 5 \text{ GeV}$ for $\Delta R(\mu, \gamma) < 0.3$ |
| muons        |                                                                          |                                                                      |

parameters, PDFs, cuts, and the lepton – photon recombination procedure of Ref. [19]. The results are shown in Table II. The table lists the lowest order cross section in the $G_\mu$ scheme, $\sigma^{(0)}$, and the relative corrections,

$$\delta = \frac{\sigma^{O(\alpha^3)} - \sigma^{(0)}}{\sigma^{(0)}}$$

(17)

for the full $O(\alpha)$ EW corrections, and

$$\delta_{PA} = \frac{\sigma^{PA} - \sigma^{(0)}}{\sigma^{(0)}}$$

(18)

for the $O(\alpha)$ EW corrections in the pole approximation. Our results are found to agree within the statistical accuracy of the Monte Carlo integration with those obtained in Ref. [19] over the entire lepton $p_T$ range.

C. Non-resonant corrections to the transverse mass distribution

Table II shows that the non-resonant EW corrections quickly become important for transverse momenta above the Jacobian peak region, $p_T(\ell) > M_W/2 \approx 40 \text{ GeV}$. The deviation of the $O(\alpha^3)$ cross section in the pole approximation from the full NLO result at large $p_T(\ell)$ is due to large Sudakov-like electroweak logarithms of the form $(\alpha/\pi) \ln^2(m(\ell\nu)/M_V)$ ($V = W, Z$) [32] which mostly arise from the contribution of the $W, Z$ box diagrams and to a lesser extend from the energy dependence of the form factors $F^a_{\text{weak}}$ and $F_\gamma - \sum_a \tilde{F}^a_{YFS}$. The same qualitative behavior is also expected for the transverse mass distribution, which is used to determine the $W$ mass and width [14]. The transverse mass is defined by

$$M_T = \sqrt{2p_T(\ell)p_T(\nu)(1 - \cos \phi^{\ell\nu})},$$

(19)

where $p_T(\nu)$ is the transverse momentum of the neutrino, and $\phi^{\ell\nu}$ is the angle between the charged lepton and the neutrino in the transverse plane. The neutrino transverse momentum is identified with the missing transverse momentum, $\not{p}_T$, in the event.
TABLE II. Integrated lowest order cross sections in the $G_{\mu}$ scheme, $\sigma^{(0)}$, the relative corrections for the pole approximation, $\delta_{PA}$, and the full $O(\alpha)$ EW corrections, $\delta$, for several ranges of $p_T(\ell)$. Shown are the results of our calculation and of Ref. [19] for a) $p\bar{p} \rightarrow \ell\nu$ at $\sqrt{s} = 2$ TeV and b) $pp \rightarrow \ell\nu$ at $\sqrt{s} = 14$ TeV. The statistical uncertainties of the Monte Carlo integration are also shown. The lepton and photon momenta are recombined for small values of $\Delta R(\ell, \gamma)$ using the simplified procedure described in Ref. [19]. We also use the same input parameters, PDFs and cuts as Ref. [19].

| $p_T(\ell)$ (GeV) | $\sigma^{(0)}$ (pb) Ref. [19] | $\sigma^{(0)}$ (pb) this calc. | $\delta$ (%) Ref. [19] | $\delta$ (%) this calc. | $\delta_{PA}$ (%) Ref. [19] | $\delta_{PA}$ (%) this calc. |
|-------------------|-------------------------------|-------------------------------|------------------------|------------------------|-----------------------------|-----------------------------|
|                   |                               |                               |                        |                        |                             |                             |
| 25 – $\infty$     | 407.03(5)                    | 407.02(7)                     | −1.8(1)                | −1.7(1)                | −1.7(1)                     | −1.8(1)                     |
| 50 – $\infty$     | 2.481(1)                     | 2.4817(6)                     | −2.7(1)                | −2.5(1)                | −1.6(1)                     | −1.5(1)                     |
| 75 – $\infty$     | 0.3991(1)                    | 0.39926(9)                    | −4.8(1)                | −4.7(1)                | −2.3(1)                     | −1.5(1)                     |
| 100 – $\infty$    | 0.1305(1)                    | 0.13058(3)                    | −6.3(1)                | −6.1(1)                | −2.5(1)                     | −2.2(1)                     |
| 200 – $\infty$    | 0.006020(2)                  | 0.006017(2)                   | −10.4(1)               | −10.1(1)               | −3.3(1)                     | −2.4(1)                     |
| 300 – $\infty$    | 0.0004821(2)                 | 0.0004821(3)                  | −13.6(1)               | −13.3(1)               | −3.9(1)                     | −3.7(1)                     |

| $p_T(\ell)$ (GeV) | $\sigma^{(0)}$ (pb) Ref. [19] | $\sigma^{(0)}$ (pb) this calc. | $\delta$ (%) Ref. [19] | $\delta$ (%) this calc. | $\delta_{PA}$ (%) Ref. [19] | $\delta_{PA}$ (%) this calc. |
|-------------------|-------------------------------|-------------------------------|------------------------|------------------------|-----------------------------|-----------------------------|
|                   |                               |                               |                        |                        |                             |                             |
| 25 – $\infty$     | 1933.5(3)                    | 1933.4(3)                     | −1.8(1)                | −1.8(1)                | −1.8(1)                     | −1.8(1)                     |
| 50 – $\infty$     | 11.50(1)                     | 11.499(2)                     | −2.7(1)                | −2.3(1)                | −1.5(1)                     | −1.2(1)                     |
| 75 – $\infty$     | 0.8198(4)                    | 0.8202(1)                     | −6.2(1)                | −6.0(1)                | −1.6(1)                     | −1.5(1)                     |
| 100 – $\infty$    | 0.1015(1)                    | 0.10155(2)                    | −10.2(1)               | −10.1(1)               | −1.6(1)                     | −1.6(1)                     |
| 200 – $\infty$    | 0.005277(1)                  | 0.005277(1)                   | −19.6(1)               | −19.1(1)               | −2.4(1)                     | −2.4(1)                     |
| 500 – $\infty$    | 0.0003019(1)                 | 0.0003019(1)                  | −29.6(1)               | −28.6(1)               | −3.6(1)                     | −3.4(1)                     |
| 1000 – $\infty$   |                               |                               |                        |                        |                             |                             |
FIG. 4. The ratio \( \frac{d\sigma^{\alpha^3}/dM_T}{d\sigma^{EBA}/dM_T} \) as a function of the transverse mass for a) \( p\bar{p} \rightarrow e^+\nu_e(\gamma) \) and b) \( p\bar{p} \rightarrow \mu^+\nu_\mu(\gamma) \) at \( \sqrt{s} = 2 \text{ TeV} \). The solid lines show the ratio of the complete \( \mathcal{O}(\alpha^3) \) electroweak and the EBA differential cross section. The dashed lines display the corresponding ratio for the case where only the resonant \( \mathcal{O}(\alpha) \) EW corrections (see Eq. (4)) are taken into account (pole approximation). The dotted lines show the ratio when the \( p\bar{p} \rightarrow W^+(\rightarrow \ell\nu)Z(\rightarrow \bar{\nu}\nu) \) background is included in addition to the complete \( \mathcal{O}(\alpha) \) EW corrections. The cuts and lepton identification requirements imposed are described in Sec. IIA.

The ratio of the complete \( \mathcal{O}(\alpha^3) \) electroweak and the EBA differential cross section as a function of \( M_T \) is shown in Fig. 4. In order to make the effect of the non-resonant weak corrections more transparent, we also show the corresponding ratio for the case of the \( \mathcal{O}(\alpha) \) EW corrections in the pole approximation (dashed lines) [18]. For \( M_T \leq M_W \), the pole approximation is seen to very well represent the complete electroweak \( \mathcal{O}(\alpha) \) corrections. In this region, the shape change in the \( M_T \) distribution is largely due to the contribution of the final state QED-like corrections. Due to the recombination of electrons and photons, the EW corrections in the pole approximation reduce the \( e\nu(\gamma) \) differential cross section by only \( 2-3\% \) over the transverse mass region considered. In the muon case, the cut on the photon energy for photons which have a small opening angle with the muon reduces the hard photon part of the \( \mathcal{O}(\alpha^3) \mu\nu(\gamma) \) cross section. As a result, the QED-like corrections are much more pronounced and display a much stronger dependence on the transverse mass than in the electron case. Without taking the lepton identification criteria of Table I into account, the QED-like corrections in the electron case are larger, due to the mass singular logarithmic corrections which originate from final state photon radiation. The non-resonant...
FIG. 5. The ratio \( \frac{d\sigma^{O(\alpha^3)}}{dM_T}/[d\sigma^{EBA}/dM_T] \) as a function of the transverse mass for a) \( pp \to e^+\nu_{e}(\gamma) \) and b) \( pp \to \mu^+\nu_{\mu}(\gamma) \) at \( \sqrt{s} = 14 \text{ TeV} \). The solid lines show the ratio of the complete \( O(\alpha^3) \) electroweak and the EBA differential cross section. The dashed lines display the corresponding ratio for the case where only the resonant \( O(\alpha) \) EW corrections (see Eq. (4)) are taken into account (pole approximation). The dotted lines show the ratio when the \( pp \to W^+(\to \ell^+\nu)Z(\to \bar{\nu}\nu) \) background is included in addition to the complete \( O(\alpha^3) \) EW corrections. The cuts and lepton identification requirements imposed are described in Sec. IIIA.

EW corrections are seen to increase rapidly in size with \( M_T \). For \( M_T = 300 \text{ GeV} \), they reduce the cross section by about 4% at the Tevatron. The dotted lines, finally, show cross section ratio when the \( p\bar{p} \to W^+(\to \ell^+\nu)Z(\to \bar{\nu}\nu) \) background is included in addition to the complete \( O(\alpha) \) EW corrections. The \( WZ \) background is seen to be much smaller than the \( O(\alpha) \) EW corrections.

Figure 5 shows the corresponding results for the LHC in the high transverse mass tail. For \( M_T > 1.5 \text{ TeV} \), the non-resonant \( O(\alpha) \) EW corrections reduce the differential cross section by 20% or more and thus are of the same size as the \( O(\alpha_s) \) corrections. This is larger than the expected statistical uncertainty in a 200 GeV bin centered at \( M_T = 1.5 \text{ TeV} \) for 100 fb\(^{-1}\). It will thus be important to take into account the non-resonant weak corrections when searching for new heavy \( W \) bosons, such as Kaluza-Klein excitations of the \( W \) appearing in TeV-scale models with extra dimensions [53], at the LHC. The results shown in Fig. 5, however, should be interpreted with caution. Since the non-resonant weak corrections become large for transverse masses above 1 TeV, they need to be resummed in order to obtain accurate predictions in this phase space region (for a recent review of the resum-
mation of electroweak Sudakov-like logarithms see Ref. [54]). Although the resummation of electroweak Sudakov-like logarithms in general four fermion electroweak processes has been discussed in the literature [55], a calculation of $\ell\nu$ production in hadronic collisions which includes resummation of electroweak logarithms has not been carried out yet.

**D. Non-resonant EW radiative corrections and the $W$ width**

We now discuss how the non-resonant $\mathcal{O}(\alpha)$ EW corrections affect the $W$ width determined from the high transverse mass tail. In Run I of the Tevatron, $\Gamma_W$ has been measured by the CDF and DØ Collaborations using this technique with a combined uncertainty of 105 MeV [14]. In Run II, with an integrated luminosity of 2 fb$^{-1}$, one expects to achieve a precision of 50 MeV per lepton channel and experiment [7]. Assuming that the error correlation of CDF and DØ data, as in Run I, is small [14] for the $W$-width measurement, this results in an expected overall uncertainty of $\delta\Gamma_W = 25 - 30$ MeV. In the experimental analysis, the measured transverse mass distribution is compared with the theoretical prediction for various values of the $W$ width where the total cross section has been normalized to the experimental value [16,17]. This is equivalent to analyzing the normalized $M_T$ distribution,

$$\frac{d\tilde{\sigma}}{dM_T}(\Gamma_W) = \frac{1}{\sigma_{tot}(\Gamma_W)} \frac{d\sigma}{dM_T}(\Gamma_W),$$

(20)

where $\sigma_{tot}$ is the total $\ell\nu(\gamma)$ cross section within cuts. At lowest order, the $W$ width enters the cross section in the form of the squared $W$ propagator (see Eq. (6)),

$$|D_W(s)|^2 = [(s - M_W^2)^2 + s^2 \Gamma_W^2/M_W^4]^{-1}.$$  

(21)

The lowest order total $\ell\nu$ production cross section thus is proportional to $1/\Gamma_W$. As a result, $d\tilde{\sigma}/dM_T$ is proportional to $\Gamma_W$ for $M_T \gg M_W$. This is clearly displayed in the ratio of the normalized $M_T$ distribution for arbitrary $\Gamma_W$ to the normalized $M_T$ distribution in the SM,

$${\mathcal{R}}(\Gamma_W) = \frac{d\tilde{\sigma}}{dM_T}(\Gamma_W),$$

(22)

where $\Gamma_W^{SM} = 2.072$ GeV is the SM $W$ width (see Sec. III A). $\mathcal{R}(\Gamma_W)$ in the pole approximation, $\mathcal{R}_{PA}(\Gamma_W)$, is shown in Fig. 6 for $p\bar{p} \rightarrow e^+\nu(\gamma)$ at the Tevatron with $\Gamma_W = \Gamma_W^{SM} - 10$ MeV = 2.062 GeV (dotted line) and $\Gamma_W = \Gamma_W^{SM} - 30$ MeV = 2.042 GeV (dashed line). Qualitatively similar results are obtained for the muon final state and at LHC energies. $\mathcal{R}_{PA}(\Gamma_W)$ is seen to be almost constant and approximately equal to $\Gamma_W/\Gamma_W^{SM}$ for $M_T > 100$ GeV. For $M_T < M_W$, the ratio is almost independent of $\Gamma_W$, and very close to one. If no detector resolution effects are taken into account, $d\tilde{\sigma}/dM_T$ is proportional to $1/\Gamma_W$ for $M_T = M_W$ at lowest order, resulting in an enhancement proportional to $\Gamma_W^{SM}/\Gamma_W$ in $\mathcal{R}(\Gamma_W)$. However, detector resolution effects largely dilute this effect. Above the $W$ mass, $\mathcal{R}_{PA}(\Gamma_W)$ rapidly drops from $\mathcal{R}_{PA}(\Gamma_W) \approx 1$ to $\mathcal{R}_{PA}(\Gamma_W) \approx \Gamma_W/\Gamma_W^{SM}$. The $M_T$ range over which the transition occurs sensitively depends on the missing transverse momentum resolution of the detector.
FIG. 6. The ratio $\mathcal{R}(\Gamma_W)$ (see text for definition) as a function of $M_T$ for $p\bar{p} \to e^+ \nu (\gamma)$ at the Tevatron. Results are shown for $\Gamma_W = \Gamma_W^{SM} - 10$ MeV = 2.062 GeV (dotted line) and $\Gamma_W = \Gamma_W^{SM} - 30$ MeV = 2.042 GeV (dashed line) with the cross sections calculated in the pole approximation. The solid line displays the ratio $\mathcal{R}_{PA}^{O(\alpha^3)} = [d\bar{\sigma}_{O(\alpha^3)}/dM_T]/[d\bar{\sigma}_{PA}/dM_T]$ in the SM. The data points and error bars indicate the measurements and statistical uncertainties expected for $\mathcal{R}_{PA}^{O(\alpha^3)}$ for 20 GeV bins and an integrated luminosity of 2 fb$^{-1}$, assuming that the data are described by the SM prediction including the full $O(\alpha)$ EW corrections. The cuts and lepton identification requirements imposed are described in Sec. III A.

The solid line in Fig. 6 shows the ratio

$$\mathcal{R}_{PA}^{O(\alpha^3)} = \frac{d\bar{\sigma}_{O(\alpha^3)}}{dM_T} / \frac{d\bar{\sigma}_{PA}}{dM_T}, \quad (23)$$

for $\Gamma_W = \Gamma_W^{SM}$, where $d\bar{\sigma}_{O(\alpha^3)}/dM_T$ and $d\bar{\sigma}_{PA}/dM_T$ are the SM normalized $M_T$ distributions for the complete $O(\alpha^3)$ calculation and in the pole approximation, respectively. The data points and error bars in Fig. 6 indicate the measurements and statistical uncertainties expected for $\mathcal{R}_{PA}^{O(\alpha^3)}$ for 20 GeV bins and an integrated luminosity of 2 fb$^{-1}$, assuming that the data are described by the SM prediction including the full $O(\alpha)$ EW corrections. Due to the
non-resonant EW corrections, $R_{PA}^{O(\alpha^3)}$ gradually decreases for $M_T > 90$ GeV, the region used by the Tevatron experiments to extract the $W$ width from Run I data [16,17]. Although the shape change of the transverse mass distribution due to the non-resonant $O(\alpha)$ EW correction differs significantly from that caused by a non-standard $W$ width, the expected statistical uncertainties make it difficult to distinguish between a small negative shift in $\Gamma_W$ and the effect of non-resonant $O(\alpha)$ EW corrections. A $\chi^2$ analysis shows that, were the non-resonant $O(\alpha)$ EW corrections ignored, the value of the $W$ width extracted from the $M_T > 90$ GeV data region would be shifted by

$$\Delta \Gamma_W = -7.2\text{ MeV}. \quad (24)$$

Since the $M_T$ distribution depends little on the detector resolution for $M_T > 90$ GeV, the shift in $\Gamma_W$ is almost independent of these effects [7]. For the precision of $\Gamma_W$ expected in Run II, a difference of $\approx 7$ MeV in the extracted value of the $W$ width cannot be ignored, and the complete $O(\alpha^3)$ calculation should be used to compare theory and data.

E. Non-resonant corrections to the $W$ boson cross section and the $W$ to $Z$ cross section ratio

As mentioned in the Introduction, the $W$ width can also be determined from the cross section ratio $R_{W/Z}$ (see Eq. (1)), together with the theoretical prediction for the ratio of the total $W$ and $Z$ production cross sections, the LEP measurement of the $Z \rightarrow \ell^+\ell^-$ branching ratio, and the SM prediction for the $W \rightarrow \ell\nu$ decay width. Since the QCD corrections to $W$ and $Z$ production are very similar, they cancel almost perfectly in the $W$ to $Z$ cross section ratio; the $O(\alpha_s)$ corrections to $R_{W/Z}$ are of $O(1\%)$ or less, depending on the set of parton distribution functions used [56]. In addition many experimental uncertainties, such as the luminosity uncertainty, cancel in the cross section ratio. Accurate knowledge of how electroweak corrections affect $R_{W/Z}$ is thus very important.

The $W$ cross section may be used as a luminosity monitor in the future [57,58]. This requires that the $W$ cross section is reliably computed with small uncertainty. In order to achieve this, it is essential to know how EW radiative corrections affect the $W \rightarrow \ell\nu$ cross section.

The size of the $O(\alpha)$ electroweak corrections to the total $p\bar{p} \rightarrow \ell\nu X$ cross section and to $R_{W/Z}$ is sensitive to the acceptance cuts and whether lepton identification requirements are taken into account or not. In Table III, we list the electroweak $K$-factor,

$$K_{EW}^{\ell\nu X} = \frac{\sigma^{O(\alpha^3)}(p\bar{p} \rightarrow W \rightarrow \ell\nu X)}{\sigma^{EBA}(p\bar{p} \rightarrow W \rightarrow \ell\nu)}, \quad (25)$$

and the correction factor for $R_{W/Z}$,

$$K_{R}^{EW} = \frac{R_{W/Z}^{O(\alpha^3)}}{R_{W/Z}^{EBA}}, \quad (26)$$

for the acceptance cuts listed in Eqs. (14) and (15) with and without taking the lepton identification requirements of Table I into account. For $K_{EW}^{\ell\nu X}$, we also list the result obtained
TABLE III. The electroweak $K$-factor $K_{EW}^{\ell} = \sigma^{O(\alpha^3)}(\ell \rightarrow e^+ \ell^- \nu X)/\sigma^{EBA}(\ell \rightarrow e^+ \ell^- \nu)$ $(\ell = e, \mu)$ and the correction factor to $R_{W/Z}$, $K_{R}^{EW} = R_{W/Z}^{O(\alpha^3)}/R_{W/Z}^{EBA}$, with $75 \text{ GeV} < m(\ell^+ \ell^-) < 105 \text{ GeV}$, for $p\bar{p}$ collisions at $\sqrt{s} = 2 \text{ TeV}$. Shown are the predictions without and with the lepton identification requirements of Table I taken into account. The cuts imposed are listed in Eqs. (14) and (15). The energy and momentum resolutions used are described in Sec. III A.

| without lepton id. requirements | with lepton id. requirements |
|---------------------------------|-----------------------------|
| $K_{EW}^{e}$ | 0.963 | 0.984 |
| $K_{EW}^{\bar{p}p \rightarrow e^+ \nu X}$, pole appr. | 0.963 | 0.984 |
| $K_{EW}^{\bar{p}p \rightarrow \mu^+ \nu X}$, pole appr. | 0.979 | 0.944 |
| $K_{R}^{EW}^{e}$ | 1.024 | 0.992 |
| $K_{R}^{EW}^{\mu}$ | 1.002 | 1.045 |

in the pole approximation. To compute the $O(\alpha^3)$ $Z$ boson cross section entering $R_{W/Z}$, we use the full $O(\alpha^3)$ calculation of di-lepton production in hadronic collisions described in Ref. [33]. We include photon exchange and $\gamma Z$ interference effects, and impose a cut on the di-lepton invariant mass of $75 \text{ GeV} < m(\ell^+ \ell^-) < 105 \text{ GeV}$. The definition of the effective Born approximation in the $Z$ case is given in Ref. [33]. The values listed in Table III update the results presented in Ref. [18] which were obtained using the pole approximation for the $W \rightarrow \ell \nu$ cross section and only included QED corrections in the $Z$ boson case.

From the results listed in Table III one observes that the non-resonant $O(\alpha)$ EW corrections have a very small effect on the $W$ boson cross section; they change the electroweak $K$ factor by $O(10^{-3})$ or less. The full $O(\alpha)$ EW corrections decrease the $W$ cross section and increase $R_{W/Z}$ by several per cent for the cuts imposed. When lepton identification requirements are included, the corrections are reduced in the electron case and enhanced in the muon case. Unlike the QCD corrections, the electroweak corrections do not cancel in $R_{W/Z}$. In $Z \rightarrow \ell^+ \ell^-$ both leptons can emit photons, whereas only the charged lepton radiates in $W \rightarrow \ell \nu \ell^- \nu$ decays. Since final state photonic corrections are the dominating contribution to the $O(\alpha)$ EW corrections, the $O(\alpha)$ corrections to the $W$ and $Z$ cross sections are quite different, and therefore do not cancel in $R_{W/Z}$. They are of the same size as the QCD corrections to $R_{W/Z}$. The size of the $O(\alpha)$ EW corrections to the total $p\bar{p} \rightarrow \ell \nu X$ cross section and to $R_{W/Z}$ is similar to the statistical uncertainty from $72 \text{ pb}^{-1}$ of data from Run II [59]. It is thus important to take the EW radiative corrections into account in the Run II data analysis.

F. The full $O(\alpha)$ electroweak corrections to the $W$ to $Z$ transverse mass ratio

Since detectors cannot directly detect the neutrinos produced in the leptonic $W$ boson decays, $W \rightarrow \ell \nu$, and cannot measure the longitudinal component of the recoil momentum, there is insufficient information to reconstruct the invariant mass of the $W$ boson. Instead, the transverse mass distribution of the final state lepton pair, or the transverse momentum
distribution of the charged lepton are used [12,13] to extract $M_W$. The $M_T$ distribution has the advantage of being invariant under transverse boosts to first order in the velocity of the $W$ boson. On the other hand, the transverse mass depends on an accurate reconstruction of the neutrino direction which leads to a set of experimental requirements which are difficult in practice to control [7].

The $W$ mass can also be determined from the ratio of the transverse mass distributions of the $W$ and $Z$ boson [60–62]. The advantage of this method is that one can cancel common scale factors in ratios and directly determine $M_W/M_Z$, which can be compared with the precise value of $M_Z$ from the LEP experiments. The downside of the ratio method is that the statistical precision of the $Z$ sample is directly propagated into the resultant overall uncertainty of $M_W$. At high luminosities, or when a detailed understanding of the detector response is not available, the transverse mass ratio of $W$ to $Z$ bosons offers advantages in determining the $W$ mass over the $M_T$ distribution [7].

The transverse mass ratio of $W$ and $Z$ bosons is defined as

$$R_{M_T}(X_{M_T}) = \frac{A_W(X_{M_W} = X_{M_T})}{A_Z(X_{M_Z} = X_{M_T})},$$

(27)

where $A_V$ ($V = W, Z$) is the differential cross section

$$A_V(X_{M_T}) = \frac{d\sigma_V}{dX_{M_T}},$$

(28)

with respect to the scaled transverse mass,

$$X_{M_T}^V = \frac{M_T^V}{M_V}.$$  

(29)

The transverse mass of the lepton pair in $Z$ boson events is defined in complete analogy to Eq. (19):

$$M_T^Z = \sqrt{2p_T(\ell^+ + p_T(\ell^-)(1 - \cos \phi_{\ell\ell}),$$

(30)

where $\phi_{\ell\ell}$ is the angle between the two charged leptons in the transverse plane. The $O(\alpha)$ EW radiative corrections to $R_{M_T}$ were calculated in Ref. [18] in the approximation where only the QED corrections were taken into account for $Z \rightarrow \ell^+\ell^-$ and the pole approximation was used for $W \rightarrow \ell\nu$. Here we present results which include the complete $O(\alpha)$ EW radiative corrections.

The ratio of the $O(\alpha^3)$ and the $W$ to $Z$ transverse mass ratio in the effective Born approximation is shown in Fig. 7. To calculate the $O(\alpha)$ electroweak corrections to $Z$ boson production, we again use the results of Ref. [33]. As before, photon exchange and $\gamma Z$ interference effects are included and an additional cut on the di-lepton invariant mass of $75 \text{ GeV} < m(\ell^+\ell^-) < 105 \text{ GeV}$ has been imposed. The complete $O(\alpha)$ EW radiative corrections (solid lines) are uniformly about 1% smaller than those obtained when the pole approximation is used for $W$ production and only QED corrections to $Z$ boson production are taken into account (dashed lines). Most of this effect originates from the genuine weak corrections to the $Z$ boson cross section. In the electron case, the $O(\alpha)$ EW corrections
FIG. 7. Ratio of the $O(\alpha^3)$ and the $W^+$ to $Z$ transverse mass ratio in the effective Born approximation as a function of the scaled transverse mass, $X_{M_T}$, for $p\bar{p}$ collisions at $\sqrt{s} = 2$ TeV, a) for electron, and b) for muon final states. The solid lines show the result when the complete $O(\alpha)$ EW radiative corrections are taken into account. The dashed lines give the result for the EW radiative corrections to $W$ production in the pole approximation, and when only the QED corrections to $Z$ boson production are taken into account (as in Ref. [18]). The cuts and lepton identification requirements imposed are described in Sec. III A. For $p\bar{p} \rightarrow \ell^+\ell^-(\gamma)$, we in addition require the di-lepton invariant mass to satisfy the constraint $75 \text{ GeV} < m(\ell^+\ell^-) < 105 \text{ GeV}$.

change $R_{M_T}$ by $1 - 2\%$. They are larger in the muon case ($3 - 7\%$) due to the lepton identification requirements which reduce the hard photon part of the $O(\alpha^3)$ cross section and thus enhance the effect of the virtual corrections. In the resonance region, $X_{M_T} \approx 1$, hard photonic corrections reduce both the $W$ and $Z$ boson cross sections. As mentioned before, in $p\bar{p} \rightarrow \ell^+\ell^-(\gamma)$ only one of the two leptons can emit a photon, whereas both leptons in $p\bar{p} \rightarrow \ell^+\ell^-(\gamma)$ can radiate. As a result, the effect of the hard photonic corrections is more pronounced in the $Z$ case, resulting in a resonance-like enhancement of $R_{M_T}$ for $X_{M_T} \approx 1$.

IV. CONCLUSIONS

We have presented a calculation of the $O(\alpha)$ corrections to $p\bar{p} \rightarrow W^{\pm} \rightarrow \ell^{\pm}\nu$ based on the complete set of one-loop Feynman diagrams contributing to $\ell\nu$ production, using the methods developed in Ref. [24]. The calculation is based on a combination of analytic
and Monte Carlo integration techniques. Lepton mass effects are included in the approximation where only mass singular terms originating from the collinear singularity associated with final state photon radiation are retained. The ultraviolet divergences associated with the virtual corrections are regularized using dimensional regularization and the on-shell renormalization scheme [31]. The cross sections obtained using our matrix elements were found to be in very good agreement with those obtained in Ref. [19].

Since the structure of the full $\mathcal{O}(\alpha^3)$ matrix elements and the $\mathcal{O}(\alpha)$ EW corrections in the pole approximation were discussed in detail in earlier papers [18,19], we concentrated on the phenomenological effects of the non-resonant corrections. The non-resonant corrections were found to have a very small effect on the total $W$ cross section and the transverse mass distribution in the region $M_T \leq M_W$. However, they increase rapidly in magnitude with $M_T$ above the $W$ peak, due to the presence of Sudakov-like electroweak logarithms. Although these corrections are of moderate size for transverse masses accessible at the Tevatron, they induce a shift of $\Delta \Gamma_W \approx -7$ MeV in the $W$ width extracted from the tail of the transverse mass distribution. Comparison with the expected overall precision of about $\delta \Gamma_W = 25 - 30$ MeV in Run II of the Tevatron shows that it will be necessary to take the non-resonant electroweak corrections into account in the data analysis. For transverse masses in the TeV region which play an important role in new physics searches at the LHC, the non-resonant electroweak corrections are of the same size as the $\mathcal{O}(\alpha_s)$ corrections. The strong increase of these corrections with $M_T$ requires that they are resummed. No such calculation exists yet for $\ell \nu$ production in hadronic collisions.

We also updated the results of the electroweak $K$-factor for the $W/Z$ cross section ratio $R_{W/Z}$ and the $W$ to $Z$ transverse mass ratio given in Ref. [18] to include the complete $\mathcal{O}(\alpha)$ EW radiative corrections and the $\mathcal{O}(g^4 m_t^2 / M_W^2)$ corrections to the $Z$ boson cross section [33].

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APPENDIX A: THE PURELY WEAK CONTRIBUTION

The pure weak vertex and self energy one-loop corrections to the process described by the form factor $F_{\text{weak}}$ of Eq. (2) can be decomposed into initial and final state contributions as follows

$$F_{\text{weak}}(\hat{s}) = F_{\text{weak}}^{\text{initial}}(\hat{s}) + F_{\text{weak}}^{\text{final}}(\hat{s})$$

(A1)

with the final state contribution
Performing the substitution \((f, f') \rightarrow (i, i')\) yields the corresponding initial state form factor \(F_{\text{weak}}^{\text{init}}(\hat{s})\). The explicit expressions for the various terms in Eq. (A2) can be found in Ref. [24].

The pure weak box contribution consisting of the \(WZ\) box diagrams of Fig. 2 in the limit of massless external fermions reads

\[
d\hat{\sigma}_{WZ\text{box}}(\hat{s}, \hat{t}) = \frac{\alpha}{\pi} d\hat{\sigma}^{(0)}(\hat{s}, \hat{t})(\hat{s} - M_W^2) \sum_{V_1, V_2 = Z, W} \mathcal{R}\epsilon(\delta B_{V_{1}V_{2}} + \delta B_{\text{crossed}})
\]

with

\[
\delta B_{V_{1}V_{2}}(\hat{s}, \hat{t}) = \lambda_{V_{1}V_{2}} \left[2D_0^0 + \hat{t}(D_1^1 + D_1^2 + D_1^3 + D_2^2 + D_2^{23} + D_2^{12})\right]
\]

\[
\delta B_{\text{crossed}}(\hat{s}, \hat{u}) = \lambda_{V_{1}V_{2}}^c \left[8D_0^0 + \hat{u}(D_1^1 + 2D_1^2 + D_1^3 + 2(D_2^2 + D_2^{23} + D_2^{12})) - 2\hat{s}D_2^{13}\right]
\]

where \(D_{ij}^{(k)} = D_{ij}^{(k)}(\hat{s}, \hat{t}, 0, M_{V_1}, 0, M_{V_2})\) in case of the box, \(\delta B_{V_{1}V_{2}}\), and \(D_{ij}^{c(k)} = D_{ij}^{c(k)}(\hat{s}, \hat{t}, 0, M_{V_1}, 0, M_{V_2})\) in case of the crossed box contribution, \(\delta B_{\text{crossed}}\). \(\lambda_{V_{1}V_{2}}\) and \(\lambda_{V_{1}V_{2}}^c\) contain the dependence on the vector and axial vector parts of the fermion-\(Z\) boson couplings which, in the notation of Fig. 8, are given by

\[
\begin{align*}
\lambda_{WZ} &= 2(v_1 + a_1)(v_3 + a_3), \\
\lambda_{ZW} &= 2(v_2 + a_2)(v_4 + a_4), \\
\lambda_{WZ}^c &= 2(v_1 + a_1)(v_4 + a_4), \\
\lambda_{ZW}^c &= 2(v_2 + a_2)(v_3 + a_3),
\end{align*}
\]

with \(v_i = (I_i^f - 2s_w^2Q_i)/(2s_wc_w)\) and \(a_i = I_i^f/(2s_wc_w)\) parametrizing the \(Zf\bar{f} (f = \ell, q)\) couplings. Here, \(Q_i\) and \(I_i^f\) denote the charge and third component of the weak isospin quantum numbers of the fermion \(i\), and \(s_w = \sin \theta_W, c_w = \cos \theta_W\) with \(\theta_W\) being the weak mixing angle.

The box integrals arising in the calculation of the Feynman diagrams of Fig. 8 are of the form

\[
\frac{i}{16\pi^2}(D_0, D^\mu, D^{\mu\nu}) = \int \frac{d^4k}{(2\pi)^4} \frac{(1, k^\mu, k^\nu)}{[k^2 + i\epsilon][(k - p_2)^2 - M_{V_1}^2][(k + p_k)^2 + i\epsilon][(k + p_1)^2 - M_{V_2}^2]}.
\]

The explicit decomposition of the vectorial and tensorial four point functions

\[
\begin{align*}
D^\mu &= -p_2^\mu D_1^1 + p_k^\mu D_1^2 + p_1^\mu D_1^3 \\
D^{\mu\nu} &= p_2^\mu p_2^\nu D_1^1 + p_k^\mu p_k^\nu D_2^2 + p_1^\mu p_1^\nu D_2^3 + g^{\mu\nu} D_2^0 \\
&\quad - (p_2^\mu p_k^\nu + p_k^\mu p_2^\nu)D_2^{12} - (p_2^\mu p_1^\nu + p_1^\mu p_2^\nu)D_2^{13} + (p_k^\mu p_1^\nu + p_1^\mu p_k^\nu)D_2^{23}
\end{align*}
\]

can be found in Ref. [63].
FIG. 8. Notations used in the calculation of the $V_1, V_2$ box diagrams.

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