Dynamical signature: complex manifolds, gauge fields and non-flat tangent space

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(Dated: September 13, 2022)

Theoretical possibilities of models of gravity with dynamical signature are discussed. The different scenarios of the signature change are proposed in the framework of Einstein-Cartan gravity. We consider, subsequently, the dynamical signature in the model of the complex manifold with complex coordinates and complex metric introduced, a complexification of the manifold and coordinates through new gauge fields, an additional gauge symmetry for the Einstein-Cartan vierbein fields and non-flat tangent space for the metric in the Einstein-Cartan gravity. A new small parameter, which characterizes a degree of the deviation of the signature from the background one, is introduced in all models. The zero value of this parameter corresponds to the signature of an initial background metric. In turn, in the models with gauge fields present, this parameter represents a coupling constant of the gauge symmetry group. The mechanism of metric’s determination through induced gauge fields with defined signature in the corresponding models is considered. The ways of the signature change through the gauge fields dynamics are reviewed, the consequences and applications of the proposed ideas are discussed as well.

I. INTRODUCTION

The idea of a metric with changing signature, looking unusual, attracts a lot of attention in the quantum cosmology and quantum gravity, see different aspects of the problem in [1–15]. Whereas all experiments and observations do not question the fact that the classical metric of the Universe has Lorentzian signature, these are two windows which we can not look through to check the signature. We do not know a lot about the quantum gravity world, there are theoretical models that allows change of the signature at the quantum level, see for example [1]. The very beginning of the Universe is another corner which can hide the possible change of the signature, see for example [2].

Among other, there are two parameters of the classical gravity which is very interesting to explore in particular. They are the number of space-time dimensions and signature of the metric, the last one is related as well to the arrow of time. The main purpose of the proposed article is an investigation of the signature issue. We consider mathematical possibilities of the formulation of the gravity formalisms with a metric which can change dynamically i.e. we discuss possible scenarios of gravity with signature which takes on values in the field of complex numbers. Particularly, a determination of the signature as Euclidean or Lorentzian is happening in the models due to some additional mechanisms. There are a few possibilities which will be discussed. The first one is a complexification of the metric and manifold’s coordinates. The gravity theories with complex metric introduced are not new of course, there is a complex metric calculation and use in [16] for example. The complex manifold also arises naturally in the twistor space formulation, see [17, 18] for the different examples of projection of complex Minkowski space in order to describe gravitons outside of a linearized gravity framework. The examples of the introduction of the complex manifold in the strings frameworks sector can be found in [19, 20] as well. Nevertheless, in the present approach we consider the problem differently. There is no classical eight dimensional complex world around, therefore we introduce a small parameter which zero value corresponds to the real manifold and real coordinates. In turn, the non-zero value of the parameter adds additional dimensions to the manifold as well as an additional complex part to the metric tensor. These additional contributions to the four dimensional gravity are proportional to the powers of the parameter, therefore the smallness of the parameter allows to establish a perturbative scheme related to the expansions of objects of interest with respect to it. We consider a factorization of the real four dimensional gravity from the additional four dimensional space of complex phases. If we assume that the small parameter of the problem is not small only at some special conditions then we get that the additional contributions are important only in these extremal cases, this issue we discuss in the next section. We do not consider a reduction of the complex manifold to the four dimensional real one, but instead we assume an coexisting of an additional contribution to the usual metric. The signature of this contribution is dynamical and it’s value is limited by the value of the new parameter introduced.

The natural next step in this direction is a complexification of the manifold with the help of new gauge fields. In this case the phases of the coordinates are defined by the gauge fields. The introduced parameter in this picture is a coupling constant of the corresponding gauge group. Considering the Einstein-Cartan gravity as the base of the approach, we, consequently, obtain an additional correction to the vierbein field which depends on the introduced phases and is proportional to the new parameter. Therefore, the metric components acquire a phase factor which makes it’s signature complex and non-determined in general. Whereas the phases of the metric’s components are defined by the gauge fields values, we discuss a possibility of a determination of any requested value of the signature.
with the help of induced values of the gauge fields which satisfy some boundary conditions. In this scenario, the signature’s value is fixed by these induced fields. The mechanism can be defined in the case of a consideration of the overall metric signature as well as for the case of description of a signature’s fluctuations above some background metric.

The change of the signature in the approach through the gauge fields is considered in two different formulations in turn. As mentioned above, the first possibility is a complexification of the manifold and metric through a complexification of the coordinates achieved by the gauge fields. The other possibilities are related to the redefinition of the structure of the Einstein-Cartan gravity with the use of the new gauge fields, i.e. we consider possible generalizations of vierbein fields. The first possibility we consider is the simplest one, we complexify the vierbein by the gauge field as for the manifold’s coordinates. In this case the metric obtains an additional part which signature depends on the value of the gauge fields. This gauge field is a new degree of freedom in this set-up. Another possibility is the interesting one; we introduce a non-flat tangent space of vierbein fields through the additional scalar fields with indexes related to the Lorentz group and new gauge group. The usual vierbein in this case arises as a projection of the another vierbein which we can call as gauge one. This set-up is equivalent to the introducution of kind of a metric in the tangent space.

The new scalar field is the metric there. This non-flatness allows to define the usual metric and it’s signature in terms of the scalar fields which values, in turn, will depend on the values of induced gauge fields. For both cases, the usual metric can be formulated in a non-perturbative and in a perturbative manner. In the non-perturbative framework the manifold’s metric will be defined fully in terms of the gauge fields involved non-linearly through some 4D non-linear sigma model, i.e. the action for these fields will depend on the metric which in turn is defined in terms of the gauge fields. For the perturbative case, the gauge fields provide a fluctuation of the metric with undefined signature above some fixed background. In this case the action for the gauge fields can be considered in the flat space-time in the first approximation.

There are interesting additional problems which we do not consider in the article but which may have relation with the proposed idea. The complexification of the coordinates leads to the eight dimensional manifold, in this context a generalization of the approach can be achieved, for example, by the introduction of the coordinates considered as p-adic number field on the manifold, see [21]. This construction can lead to the manifold with dimension larger than 4D dimensions, in this case more that one small parameter can be introduced. The zero values of all parameters will lead to usual real metric in this case as well, otherwise some complicated variant of the proposed framework will be obtained. Therefore, the metric and its signature will be valued in the field of p-adic numbers instead of complex ones. Another face of the complexification is a similarity of the introduced phases to the "fast" variables of t’Hooft, [22], introduced in his generalization of quantum mechanics. Formally speaking, the "fast" variables are fields in the t’Hooft approach, their counterpart in the given framework are gauge fields. Therefore, in general, it is interesting to understand the consequences of the manifold’s complexification on the formulation of quantum mechanic approaches.

The paper is organized as follows. In the next section we discuss basic ideas of a definition of the complex coordinates and metric for a complex manifold. In the Section III we consider a simplest variant of the Einstein-Cartan gravity for the complex manifold with complex coordinates, whereas in the Section IV we investigate the gravity for the coordinates complexified by the gauge fields. In the Section V and Section VI vierbein based approaches to the problem are considered, firstly a investigation of the additional gauge symmetry for a new vierbein field and further a construction of a non-flat tangential space for the Einstein-Cartan gravity. The last section is a Conclusion of the paper.

II. COMPLEX METRICS FOR A COMPLEX MANIFOLD

In order to clarify the ideas of the framework we, first of all, consider the following simple construction. Let’s define a complex manifold on the base of the usual real four-dimensional manifold by simple complexification of it’s coordinates:

\[ p = (p^1, \ldots, p^n) \rightarrow z = (z^1, \ldots, z^n) = (p^1 e^{i\phi_1}, \ldots, p^n e^{i\phi_n}). \]  \hspace{1cm} (1)

Defining the tangential vector fields in each \( z \) of the complex and each \( p \) of the real manifolds

\[ X_z \in T_z^C, \quad X_p \in T_p^R \]  \hspace{1cm} (2)

we observe that the fields are connected as

\[ X_z = M X_p \]  \hspace{1cm} (3)

with \( M \) as \( U(4) \) diagonal matrix, see Appendix A example. Using the usual definition metric for the real manifold

\[ g(X_{p_1}, Y_{p_2}) = g((x^1, \ldots, x^n), (y^1, \ldots, y^n)) = g_{ij} x^i y^j, \]  \hspace{1cm} (4)
we correspondingly define the quadratic complex form on the complex manifold as
\[ g(X, Y) = g((x^1 e^{i\phi_1}, \ldots, x^n e^{i\phi_n}), (y^1 e^{i\phi_1}, \ldots, y^n e^{i\phi_n})) = e^{i(\phi_i + \phi_j)} g_{ij} x^i y^j. \] (5)

Now we introduce the following complex metric in a local coordinate basis defining it as
\[ g = e^{ia_\phi (\phi_i + \phi_j)} g_{ij} dx^i \otimes dx^j + \cdots \] (6)

with the \( g_{ij} \) as a metric field of the real manifold and \( a_\phi \) as some parameter. The additional parts of the metric are proportional to the new parameter \( a_\phi \), redefining the angles in this expressions for the usual complex coordinates we obtain:
\[ \phi \rightarrow a_\phi \phi. \] (7)

In order to stay in the situation with four dimensional classical world we need to push all the effects of the additional dimensions to the areas of some special regimes. In the formalism it means that the parameter must be extremely small for the case of the classical world. We define correspondingly the following dimensionless parameter:
\[ a_\phi = \frac{l_0}{R_0} \] (8)

with \( l_0 \) as Planck length and \( R_0 \) as curvature of the manifold, i.e. the parameter proposed is extremely small indeed in the present physical reality. It’s smallness has two purposes, first of all, the real metric appears in the model as the first term of the expansion of the complex one with respect to \( a_\phi \). The limit
\[ a_\phi \rightarrow 0 \] (9)

provides the expansion with the usual metric as a leading contribution term. The second important role of this parameter, as mentioned above, is that it’s smallness allows to decouple the additional metric’s components in the corresponding expression Eq. (6). Namely, for a general complex metric in eight dimensional space we have \( g_{\phi \phi} \propto a_\phi \) that provides \( a_\phi^2 \) order contribution in the corresponding gravity action. Therefore, preserving everywhere \( a_\phi \) order, i.e. with precision linear with respect to parameter, we can limit calculations by the four dimensional metric of the real Riemann manifold modified in correspondence to Eq. (6) prescription. Due the smallness of the value of the \( a_\phi \) parameter we note also that the corrections related to the complexifications can contribute only at some extremal conditions. We assume that it can be important at the level of Planck length, in quantum gravity consideration and in the situation of the extremely strong gravity appearance. This smallness, as well, provides a simple rule for the use of diffeomorphisms to change the coordinates. For the global gauge symmetry we require that
\[ \frac{\partial z^\mu}{\partial y^\nu} \rightarrow e^{ia_\phi (\phi_z - \phi_p)} \frac{\partial x^\mu}{\partial y^\nu} \] (10)
equivalent to the firstly done diffeomorphism transform and further complexification:
\[ \frac{\partial x^\mu}{\partial y^\nu} \rightarrow e^{-ia_\phi (\phi_z - \phi_p)} \frac{\partial z^\mu}{\partial p^\nu}. \] (11)

Of course in the case when \( a_\phi \rightarrow 1 \) we will need to account all eight coordinates and in this case the commutation between the complexification and real diffeomorphism transformations may not work. We do not discuss this case in the article. The inverse metric field is defined correspondingly, in the dual basis it reads as
\[ g^{-1} = e^{-ia_\phi (\phi_i + \phi_j)} g^{ij} e_i \otimes e_j \] (12)

where
\[ g_{ij} g^{jk} = \delta^k_i \] (13)

for the real manifold.

We note, that we can obtain an additional example of the complex metric if we consider the angles as some internal parameters related to the additional symmetry of the covariant and contravariant bases of the real manifold:
\[ e^i \rightarrow e^{ia_\phi} e^i, \quad e_i \rightarrow e^{-ia_\phi} e_i, \] (14)
i.e. the complexification of the basic vectors leads to the same results as complex coordinates in Eq. (1) and Eq. (5). In the following we will use the Einstein-Cartan formulation of the gravity. Therefore, considering a four dimensional real Riemann manifold and introducing the real Lorentz vierbein (tetrad) $e^a_\mu$ as usual

$$g_{\mu\nu} = \eta_{ab} e^a_\mu e^b_\nu. \tag{15}$$

we can consider the corresponding complex metric defined as following:

$$g_{\mu\nu} = \eta_{ab} e^{i(a+\phi_a)}_\mu e^a_\nu. \tag{16}$$

This metric can be obtained from the real one by the complexification of the tetrad:

$$e^a_\mu \rightarrow e^{i\phi_a} e^a_\mu \tag{17}$$

and

$$E^\mu_a \rightarrow e^{-i\phi_a} E^\mu_a \tag{18}$$

for the inverse vierbein. This construction can be considered as a particular example of the non-flat tangent space which we will discuss further.

It is important to notice, that the Eq. (13) expression can be considered as an approximate one and metric’s complexification in this formulation arises as consequence of the decoupling of the corresponding coordinates. The Eq. (14)-Eq. (18) construction, in contrast to that, is precise and the angles arise there due the introduced additional symmetries related to vierbein’s complexification. Therefore, in the first case we consider a complex manifold with real functions which depend on the complex coordinates, there is a integration over the angles in the action. In the second case, correspondingly, we have a real manifold with angles as parameters of corresponding independent $U(1)$ symmetry groups for each real coordinate with action invariant with respect to the symmetries, there is still only four real coordinates to integrate. Also, whereas in the first case we need to consider small $a_{\phi}$ parameter in the problem because so far we have no observable complex manifolds, in the second case case we can take $a_{\phi} = 1$ of course. Consequently, this formulation of the approach will lead to the variant of the framework with complex metric which was defined and discussed in [16] for example (see also [19, 23]). We do not consider this case further.

III. EINSTEIN-CARTAN ACTION FOR THE COMPLEX METRIC IN THE COMPLEX MANIFOLD

In order to derive the analog of the Einstein-Cartan action for the complex metric introduced above, we, first of all, consider the transformation of the vector with Lorentz index projected with the help of the new vierbein:

$$\delta \tilde{e}^a = \delta \left( e^{i(a+\phi_a)} e^a \right) = e^{i(a+\phi_a)} \delta e^a = e^{i(a+\phi_a)} \omega^a_b e^b = \left( e^{i(a+\phi_a)} \omega^a_b e^{i(a+\phi_b)} \right) \left( e^{i(a+\phi_b)} e^b \right) = \tilde{\omega}^a_b \tilde{e}^b, \tag{19}$$

we see that the expression is invariant in respect to the internal symmetry transformation of the covariant and contravariant Lorentz indices performed in correspondence to the Eq. (17)-Eq. (18) definitions. In the following, denoting the complex vierbein as the usual one, we will remember that the Lorentz indices allows to rotate the corresponding objects in correspondence to this symmetry without mixing with the Lorentz transformations.

Now, as mentioned above, we have to distinguish between the cases when we consider a complex manifold or we introduce the additional symmetry in the problem related to the $U(1)$ global gauge symmetry for each Lorentz index. There is the following form of Palatini action we have for the first case:

$$S = C \frac{m^2 p^2}{2} \int d^4 z \epsilon^{\mu\nu\rho\sigma} \tilde{e}_{abcd} e^c_\rho e^d_\sigma \left( D_\mu \omega_{ab}^{\mu} \right) \tag{20}$$

with $C$ as some normalizations constant, $\partial_z$ in the covariant derivative and $z$ as complex coordinates, see the use of the complex coordinates in the formulation of the Quantum Mechanics in [19] for example. Our next step is an assumption that the integration functions are analytical in the whole region of interest, except, perhaps, some extreme boundary points. Consequently, to first approximation, we can choose the integration paths for each variable.

1 Discussions about complex path integral trajectories for Lorentz path integrals can be found in [23].
$z_\mu$ taking $x_\mu \in [-\infty, \infty]$ at some fixed constant $\phi_{\mu 0}$ angles. Therefore, assuming a smallness of $a_\phi$ parameter\(^2\) we write the action till the requested precision order as following:

\[
S \approx C e^{x^{a_3}_0 \sum_{\mu=0}^3 \phi_{\mu 0}} \frac{m_p^2}{2} \int d^4 x \varepsilon^{\alpha\beta\rho\sigma} \varepsilon_{abcd} e^c_\rho e^d_\sigma (D_\alpha \omega^{ab}_\beta) +
\]

\[
+ i a_\phi C \frac{m_p^2}{2} \sum_{\mu, \nu \neq \mu,} e^{x^{a_3}_\nu \sum_{\phi_{\mu 0}}} \int d^4 x e^{x^{a_3}_\nu \phi_\mu} \varepsilon^{\alpha\beta\rho\sigma} \varepsilon_{abcd} e^c_\rho e^d_\sigma (D_\alpha \omega^{ab}_\beta) .
\]

(21)

The condition when the usual Einstein-Cartan formalism is reproduced in the first order approximation is the following one:

\[
C e^{x^{a_3}_0 \sum_{\mu=0}^3 \phi_{\mu 0}} = 1 .
\]

(22)

There are some arbitrary constant angles $\phi_{\mu 0}$ introduced here and this condition can be considered as definition of the $C$ constant as well. We also note, that the expression under the integration is function of $z$ and in general it must be expanded as well in order to provide all $a_\phi$ order corrections to the real action.

The interesting consequence of the form of Eq. (21) action is that it does not define any preferable direction of time or preferable value of the signature. Indeed, let’s choose the special coordinate system, x-system, with

\[
\phi_{\mu 0} = 0, \quad \mu = 0 \ldots 3; \quad C = 1 .
\]

(23)

In the same way we can choose any other angles such that

\[
\sum_{\mu=0}^3 \phi_{\mu 0} \neq 0, \quad \phi_{\mu 0} \neq 0, \quad C = 1 ,
\]

(24)

the angles define a new coordinate system different from the special one. Namely, for the non-zero $\phi_{i 0}$ there are new coordinates

\[
y^\mu = x^\mu e^{x^{a_3}_0 \phi_{\mu 0}} .
\]

(25)

In terms of the new coordinates the action acquires the following form:

\[
S = \frac{m_p^2}{2} \int d^4 y \varepsilon^{\alpha\beta\rho\sigma} \varepsilon_{abcd} e^c_\rho e^d_\sigma (D_\alpha \omega^{ab}_\beta) +
\]

\[
+ i a_\phi m_p^2 \sum_{\mu=0}^3 \int y^\mu_0 d\chi_\mu d^3 y e^{x^{a_3}_\nu \chi_\mu} \varepsilon^{\alpha\beta\rho\sigma} \varepsilon_{abcd} e^c_\rho e^d_\sigma (D_\alpha \omega^{ab}_\beta) .
\]

(26)

where

\[
\phi_\mu = \phi_{\mu 0} + \chi_\mu .
\]

(27)

With redefinition of the arguments of the integral functions performed in Eq. (26) and subsequent deformation of the integration contours, the form of the redefined action is the same as Eq. (21) with Eq. (23) values of the angles. The only different contribution into the action, therefore, can come from the end points of the integration over real $y^\mu$ which acquire complex phases in the case of Eq. (25) variables change. We assume that these contributions are zero. As a result of the complexification of the manifold and its symmetry we have an infinite number of the equivalent directions of time and foliations of the spatial coordinate\(^3\) for the given value of a signature. An another interesting consequence of the model is that the Eq. (26) action quite naturally acquires a small term additional to the leading one. This term can be considered as a type of the cosmological constant in the action in the framework of the perturbative scheme based on the $a_\phi$ smallness. In general it means that the term must be finite after the integration over $\chi_\mu$ angle at some $y_\mu \rightarrow y_\mu 0$ limits taken in the corresponding contour integral.

\(^2\) We note here, that due the rotation of the manifold’s coordinates, the direction of the rotation is important. Namely, taking $e^{-x^{a_3}_0 \phi}$ (conjugated) definition of the complex coordinate, we will have to change the integration limits as $0 \rightarrow 0$ and $2\pi \rightarrow -2\pi$ that will lead to the invariance of the Eq. (21) expression with respect to the definition of the $z$ coordinate.

\(^3\) This statement can be understood in terms of any evolution equations, the equations will have the same form for any redefined $x$ coordinate.
Now, expanding the vierbein field in the new perturbative scheme as
\[ e^c_{\nu} = e^c_{\nu0} + e^c_{\nu1} \] (28)
and taking a variation of the Eq. (21) action with respect to \( \omega \) connections we will obtain:
\[ \partial_{[\mu} e^{c}_{\nu]1} = -i a_{\phi} \sum_{\rho=0}^{3} x^{\rho} \delta(x^{\rho} - x^{\rho}_{0}) \int d\phi_{\rho} e^{i a_{\phi} \phi_{\rho}} \partial_{[\mu} e^{c}_{\nu]0} \] (29)
or, equivalently
\[ \int d^{4}x \left( \partial_{\mu} e^{c}_{\nu1} + i a_{\phi} \sum_{\rho=0}^{3} x^{\rho} \delta(x^{\rho} - x^{\rho}_{0}) \int d\phi_{\rho} e^{i a_{\phi} \phi_{\rho}} \partial_{\mu} e^{c}_{\nu0} \right) = 0 . \] (30)
Providing some initial value of \( e^{c}_{\nu1}(x) \) at \( x^{\mu}_{0} \) we can write the solution of Eq. (29) as
\[ e^{c}_{\nu1} = e^{c}_{\nu1}(x^{\mu}_{0}) - i a_{\phi} \sum_{\rho=0}^{3} \int_{x^{\mu}_{0}}^{x^{\mu}} dx^{\rho} (x^{\rho} \delta(x^{\rho} - x^{\rho}_{0})) \int d\phi_{\rho} e^{i a_{\phi} \phi_{\rho}} \partial_{\mu} e^{c}_{\nu0} \] (31)
This additional vierbein’s part provides a correction to the metric through Eq. (15) definition. It can be of any signature depending on the value of the integral in Eq. (31) r.h.s..

IV. COMPLEXIFICATION OF THE MANIFOLD THROUGH GAUGE FIELDS

If we consider the simplest generalization of the Eq. (11) complexification through the replacement of the \( \phi_{\mu} \) angles by \( \phi_{\mu}(p) \) functions then we immediately realize that this construction does not work. Namely, in this setup there is no self-consistent definition of the coordinates and corresponding functions in the integrals. Therefore, in order to discuss the case of the local complexification of the real manifold, we will consider the following model. We introduce a set of real coordinates \( x^{\mu} \) and determine the new coordinates of the manifold as transform of \( x \):
\[ z^{\alpha} = M^{\alpha}_{\mu}(x) x^{\mu} \]
\[ z_{\alpha} = M_{\alpha}^{\mu}(x) x^{\mu} \] (32)
with new vierbein like gauge fields \( M \), where the new indices are transforming in correspondence to some group \( G \). Accordingly, we will consider the integrals as taken over the Riemann \( x^{\mu} \) with functions defined as depending on \( z^{\alpha} \) similarly to done before. Introducing an another form of the gauge fields, suitable for the perturbative calculations\(^4\), we can define the complex coordinates in this case as
\[ z^{\alpha} = \left( \delta^{\alpha}_{\mu} + i a_{\phi} A^{\alpha}_{\mu}(x) \right) x^{\mu} \] (33)
with \( A \) as some gauge field related to the \( G \) group. The coupling constant \( a_{\phi} \) is small again and determines a measure of the complexification. In this setup it plays a role of the coupling constant of the new gauge group \( G \).

The new gravity action of the model preserves the form of the Eq. (20) action and we have:
\[ S = C \frac{m^{2}_{P}}{2} \int d^{4}x \varepsilon^{\mu\nu\rho\sigma} \varepsilon_{abcd} e^{d}_{\mu} e^{\alpha}_{\rho} \left( D_{\mu} \omega_{\nu}^{ab} \right) \] (34)
with functions in the integral depending on \( z \) coordinates. There is, correspondingly, an additional part in the action which corresponds to the new introduced field which we consider as a gauge one. Using an usual determination of the field’s strength of the new gauge field \( A \) through the covariant derivative
\[ [D G_{\mu}, D G_{\nu}] = -i a_{\phi} G^{\beta}_{\mu\nu} t^{\alpha} \] (35)
for some representation of the \( G(N) \)
\[ [t^{\alpha}, t^{\beta}] = i f^{\alpha\beta\gamma} t^{\gamma} \] , (36)

\(^4\) We note that there is a difference between gauge actions written in terms of \( M \) and \( A \) fields.
we define this action as

\[ S_A = \kappa \int d^4x \, e \text{tr} \left[ G_{\mu\nu} G_{\mu'\nu'} \right] e^\mu_{\alpha} e^\nu_{\beta} e^{\mu'}_{\alpha'} e^{\nu'}_{\beta'} F^{ab;C}_{1} \]  

(37)

with mostly general

\[ F^{ab;C}_{1} = \kappa_1 \eta^{a_{1}\alpha} \eta^{b_{1}\beta} + \kappa_2 \eta^{a_{1}\alpha} \eta^{b_{1}\beta} + \kappa_3 \eta^{a_{1}\alpha} \eta^{b_{1}\beta} \]  

(38)

with \( \eta \) as Lorentz metric of the flat space. The action is similar, for example, to the QCD action in the curved space time. The interaction between these two parts of the action can be written in terms of the expansion of Eq. (34) functional with respect to the complex part of the \( z \) coordinates.

Our next step is an introduction of the non-trivial \( A \) gauge fields in the action. The idea is the following. We can introduce in the action a following additional term:

\[ S_{\text{ind}} = \frac{m^2}{2} \sum_i \int d^4x T_{\mu i}(A) \omega_i^{\mu} \]  

(39)

which we call an induced part of the action in correspondence to the definition of \cite{27, 28}. The purpose of this part of the action is to introduce in the equations of motion the classical values of the \( A \) gauge field, denoted as \( \omega \), which satisfy some boundary conditions at the edges of time interval. Namely, we define the boundary conditions at some different limits of \( t \) coordinate:

\[ \delta_\omega T_{\mu i}(A) = J_{\mu i}(A) \sigma A \]  

(40)

\[ J_{\mu i}(t_j) \rightarrow 0, \quad t \rightarrow t_0 j \neq i \]  

(41)

with the last equation fixed by the structure of Eq. (39) term and obtained from the usual equations of motion:

\[ \delta_\omega (S_A + S_{\text{ind}}) = 0. \]  

(42)

In general, the effective currents \( T_{\mu i} \) can be consequently reconstructed by requests of the gauge invariance of the induced part of the action, with details that can be found in \cite{27, 28}. Fixing the boundary conditions, i.e. fixing the values of \( A_{\text{cl}} \) at the edges of the overall time interval for example, we will obtain an action with some space-time foam above the background space-time of a fixed signature. The example of this construction is given in the Appendix B.

Taking only one boundary field from Eq. (B.7) expression for example, we will have

\[ A_{\text{cl} \mu} = \omega^{\mu}. \]  

(43)

Now, writing the equations of motion for \( \omega \) connections

\[ D_{[\mu} e^c_{\nu]} = 0, \quad \frac{\partial}{\partial x^\mu} = \frac{\partial z^\nu}{\partial x^\mu} - \frac{\partial z^\mu}{\partial x^\nu}, \]  

(44)

and expanding the vierbein in a perturbative scheme related with the parameter \( a_\phi \) as

\[ e_c^\nu = e_c^\nu_{0} + e_c^\nu_{1}, \]  

(45)

we will obtain then

\[ \partial_{[\mu} e_{\nu]}^{c} = - i a_\phi \partial_{[\mu} \left( \omega^{\alpha}_{\rho} x^\rho \right) \partial_{\alpha} e_{\nu]}^{c}_{0} \]  

(46)

with solution

\[ e_{\nu}^{c} = e_{\nu}^{c}_{0} (x^\rho_{0}) - i a_\phi \int_{x^\mu_{0}}^{x^\mu_{1}} dx^\mu \partial_{\mu} \left( \omega^{\alpha}_{\rho} x^\rho \right) \partial_{\alpha} e_{\nu}^{c}_{0} \]  

(47)

for the additional vierbein’s part. We see, that the vierbein acquires a correction which structure is defined by the value of the boundary \( \omega_{\mu} \) fields. The corresponding Eq. (15) metric obtains an additional part as well and the signature of this metric’s correction is depending on the \( \omega_{\mu} \) fields. This situation, as we will see further, will realized in other models with gauge field involved.
We note, that the Eq. (43) $A_{c\mu}$ field can appear in the equation as a solution of classical equations of motion which contributes mostly in the generating functional of the theory, i.e. as semi-classical solution of the theory. This possibility is a purely dynamical one and requires an analysis of the classical dynamics of whole system under consideration. For example, it is not clear, how dynamically solutions of following type

$$A_{c\mu} = A_{1\mu} + A_{2\mu}$$

with two or more different classical $A_{c\mu}$ fields will arise for a connected manifold. Further we discuss this mechanism only in the Conclusion of the paper.

It is interesting to note also, that assuming the existence of a transform opposite to Eq. (32)

$$x^\mu = N_\mu \alpha (z) z^\alpha$$

we can rewrite the Eq. (34) action fully in terms of $z$ variable as follows

$$S = C \frac{m_p^2}{2} \int d^4z \tilde{N} \varepsilon^{\mu\nu\rho\sigma} \varepsilon_{abcd} e^e_\rho e^d_\sigma (D_\mu \omega_{\nu}^{ab})$$

with $\tilde{N}$ as Jacobian of the Eq. (B.7) coordinates transform given by the

$$\tilde{N}^\mu_\alpha = N^\mu_\alpha + \frac{\partial N^\mu_\beta}{\partial z^\alpha} z^\beta$$

matrix. In this case the value of the factor in front of the action in the path integral will be determined by the value of the $\tilde{N}$. Therefore, again, the redefinition of the $C$ factor in Eq. (50) will lead to the equivalent actions for the different metrics with different phases for their components.

V. GAUGE SYMMETRY FOR THE VIERBEIN FIELD

The complexification mechanism discussed in the previous Section can be applied for the vierbein fields as well. Considering the usual vierbein use in definition of the metric

$$g_{\mu\nu} = \eta_{ab} e^a_\mu e^b_\nu,$$

we can generalize the Eq. (17)-Eq. (18) definition of the complex vierbein:

$$e^a_\mu = M_\mu^\alpha (x) e^a_\alpha$$

or similarly to done before as

$$e^a_\mu = (\delta^a_\mu + i a_\phi A_\mu^\alpha (x)) e^a_\alpha.$$

In these cases the Eq. (52) metric acquires an additional part with signature which depends on the value of the fields. We have for the first metric:

$$g_{\mu\nu} = \eta_{ab} M_\mu^\alpha (x) M_\nu^\beta (x) e^a_\alpha e^b_\beta$$

and correspondingly for the second at linear approximation with respect to $a_\phi$ parameter:

$$g_{\mu\nu} = \eta_{ab} e^a_\mu e^b_\nu + i a_\phi \eta_{ab} \left( A_\mu^\alpha (x) e^a_\alpha e^b_\beta + A_\nu^\alpha (x) e^a_\alpha e^b_\mu \right).$$

The Eq. (55)-Eq. (56) expressions are different in general. Whereas the Eq. (55) metric describes a manifold with arbitrary signature which depends on the value of $M$ matrix, the Eq. (56) metric determines a manifold with an additional part above the background metric with given signature. We note that this additional metric’s part can be of any signature as well, it depends on the value of $A$ gauge field. In both cases, the values of the gauge fields are determined dynamically through the corresponding Lagrangians.

The Einstein-Cartan action can be easily rewritten in terms of new vierbein in this case. We require that the additional metricity condition must be satisfied:

$$\nabla_\mu (A_\nu^\alpha e^a_\alpha) = (\nabla_\Gamma_\mu A_\nu^\alpha) e^a_\alpha + A_\nu^\alpha (D_\mu e^a_\alpha) = 0$$

(57)
with
\[
\nabla_{\mu} A_{\nu}^\alpha = \partial_\mu A_{\nu}^\alpha - \Gamma_{\mu\nu}^\rho A_\rho^\alpha
\]
\[
D_\mu e_\alpha^a = \partial_\mu e_\alpha^a + \omega_\mu^{\alpha b} e_\alpha^b,
\]
here \(\Gamma\) and \(\omega\) are Christoffel and Lorentz connections correspondingly. We will obtain then:
\[
S = \frac{m_P^2}{2} \int d^4x \varepsilon_{\mu\nu\rho\sigma} e_{abcd} M_{\rho\sigma}^{\alpha\beta} e_{\alpha\beta}^{ef} (D_\mu \omega_{\nu}^{ab})
\]
with obvious corresponding redefinition in terms of \(A\) field. This part of the full action is correct for any form of the metric, the non-triviality of the construction, therefore, is manifested through the additional gauge field. Introducing the gauge field strength
\[
[D_{G\mu} D_{G\nu}] = -i a_\phi G_{\mu\nu}
\]
for some symmetry group \(G\), we define the additional part of the action as
\[
S_A = \kappa \int d^4x e M \text{ tr } [G_{\mu\nu} G_{\mu_1\nu_1}] \mathcal{F}^{\mu\nu;\mu_1\nu_1}
\]
with
\[
\mathcal{F}^{\mu\nu;\mu_1\nu_1} = \kappa_1 g^{\mu\nu} g_{\mu_1\nu_1} + \kappa_2 g^{\mu\nu_1} g_{\mu_1\nu} + \kappa_3 g^{\mu\mu_1} g_{\nu\nu_1},
\]
and
\[
M = \det(M_\mu^\alpha), \quad e = \det(e_\alpha^a),
\]
with matrix \(M\) determined or through Eq. (55) or either the Eq. (56) expressions. The metric \(g^{\mu\nu}\), in turn, as well depends on the corresponding gauge field. The Eq. (52) Lagrangian describes a variant of 4D non-linear gravitational sigma-model. Correspondingly, further, we will consider only the Eq. (55) formulation of the metric, since there is no a simply perturbative expansion in respect to \(a_\phi\) for the \(M\) field in the Eq. (55) metric. Therefore, a self-consistent solution of the equations of motion for \(M\) field through Eq. (52) is a non-trivial task. We will consider it in a separate publication. Concerning the Eq. (56) metric and the \(A\) gauge field, there is the \(a_\phi\) parameter in the Eq. (56) definition of the metric so we need to know a solution for the gauge field till \(a_\phi^0\) precision only and, therefore, for our purposes it will be enough to consider the Eq. (52) action in the flat space-time. In this case, with the help of Appendix B results, we obtain:
\[
A_\mu^{cl} = A_1^{\mu} + A_2^{\mu}.
\]
Correspondingly, the Eq. (56) metric will acquire an additional part determined by the \(A_1^{\mu}\) fields:
\[
g_{\mu\nu} = \eta_{ab} e_\mu^a e_\nu^b + i a_\phi \eta_{ab} \left( A_{\mu cl}^\alpha (x) e_\alpha^a e_\nu^b + A_{\nu cl}^\alpha (x) e_\alpha^b e_\mu^a \right),
\]
we see that the metric’s fluctuations as well as their signature are determined by the induced boundary fields of the problem.

Considering the same approach for the Eq. (53) \(M\) fields, we will have a difference between covariant and contravariant gauge fields. The \(M\) gauge fields provide a non-flat metric initially, the field will appear in the relations between covariant and contravariant components in Eq. (B.2) action therefore. It makes the problem even more non-linear. There are different vectors in the power of the Eq. (B.8) ordered exponential and in the induced action.

### VI. NON-FLAT TANGENTIAL SPACE CONSTRUCTION

Geometrizing the proposed ideas, we can define the Lorentz vierbein and it’s inverse as a projection of other vierbein fields:
\[
e_\mu^a = M_\mu^\alpha e_\mu^\alpha, \quad e_\mu^\alpha = M_\mu^\alpha E^\mu; \quad e_\alpha^a E_\alpha^b = \delta_\alpha^b,
\]
\[
\delta_\alpha^\beta = M_\alpha^\beta, \quad E^\mu E_\mu^\beta = M_\alpha^\beta,
\]

\(e_\mu^a\) is a projector onto the tangent space at a point, \(E^\mu\) is a projector onto the normal space. This construction provides a way to define a non-flat spacetime geometry in terms of a flat metric. The projector fields \(e_\mu^a\) and \(E^\mu\) are related to the tangent and normal vectors at a point, respectively. The metric \(g_{\mu\nu}\) is then constructed from these projectors, ensuring that the geometry is not flat. The non-flatness is induced by the gauge fields \(A_\mu^{cl}\) and \(M_{\mu\nu}\).
here Greek indices $\alpha, \beta$ belong to some group $G$, Latin indices denote the Lorentz transforms, the $\mu, \nu$ are used as usual Riemann type indices $^5$. We note, that there is no general prescription to consider the dimension of the $G$ equal to the one of the Lorentz group, see for example $^{12}$. Nevertheless, in the following we will take $\alpha, a = 0 \ldots 3$. Also, unlike the previous Chapter, the $M$ field here is a scalar one, there is an additional dynamics present therefore. Now, using again the usual definition of flat metric in terms of Lorentz vierbein

$$g_{\mu\nu} = \eta_{S_{ab}} e_\mu^a e_\nu^b, \quad (68)$$

with $\eta_{S_{ab}}$ as a flat metric of the tangent space with some signature $S$, we rewrite it as follows:

$$g_{\mu\nu} = \eta_{S_{ab}} M^{a\alpha} M^{b\beta} e_{\mu\alpha} e_{\nu\beta} = M^{\alpha\beta}_{S'} e_{\mu\alpha} e_{\nu\beta}, \quad M^{\alpha\beta} = M^{\beta\alpha}_{S'}, \quad (69)$$

and

$$g^{\mu\nu} = \eta^{S_{ab}} M_{a\alpha} M_{b\beta} E^{\mu\alpha} E^{\nu\beta} = M^{\alpha\beta}_{S'} E^{\mu\alpha} E^{\nu\beta} \quad (70)$$

with signature $S'$ which can be different from $S$. Here we define

$$M_{a\alpha} M^{b\beta} = \delta^b_a, \quad M_{a\alpha} M^{a\beta} = \delta^\beta_\alpha. \quad (71)$$

The invariance of the new scalar product with respect to the new group of the symmetry is provided by the ordinary transformation rules for the new upper and lower indices

$$e_{\mu\alpha} = G_{\alpha \beta} e_{\mu\beta}, \quad E^{\alpha}_\mu = \tilde{G}^{\alpha \beta} E^{\beta}_\mu, \quad (72)$$

with $\tilde{G}$ matrix as inverse to $G$

$$G_{\alpha \beta} \tilde{G}^{\alpha \gamma} = \left(G \tilde{G}^T\right)^\gamma_\beta = \left(\tilde{G}^T G\right)^\beta_\gamma = \delta^\beta_\gamma. \quad (73)$$

Both $G$ and $\tilde{G}$ matrices belong to the group of interest of course. Correspondingly, we introduce the new covariant derivatives of the vierbein and $M$ fields with respect to the $G$ symmetry group:

$$D_{G \mu} e_{\nu\alpha} = \partial_\mu e_{\nu\alpha} - \Omega^{\beta}_\mu e_{\nu\beta} \quad (74)$$

and

$$D_{G \mu} M^{a\alpha} = \partial_\mu M^{a\alpha} + \Omega^{\alpha}_\mu M^{a\beta}. \quad (75)$$

Here

$$\Omega^{\alpha}_\mu = i a_\phi \Omega^{a}_\mu (t_a)^\alpha_\beta \quad (76)$$

is a new gauge field additional to the usual connection field in the corresponding covariant derivative of the Einstein-Cartan gravity Lagrangian.

The form of the Einstein-Cartan gravity action is changing trivially in this version of the formalism. We require the metricity property of the new vierbein in respect to the full covariant derivative

$$\nabla E^{\mu}_a = \nabla (M_{a\alpha} E^{\mu\alpha}) = 0 \quad (77)$$

and obtain

$$S_\omega = - m^2_p \int d^4x \sqrt{-g} E^{\mu}_a E^{\nu}_b R_{\mu\nu}^{ab} \to S = - m^2_p \int d^4x e M (M_{a\alpha} E^{\mu\alpha}) (M_{b\beta} E^{\nu\beta}) R_{\mu\nu}^{ab} \quad (78)$$

in correspondence to the Eq. $(69)$ definition, here

$$M = \det (M^{a\alpha}) \quad (79)$$

$^5$ An another variant of the non-flat tangent space is simply define $g_{\mu\nu} = M_{a\beta} e^a_\mu e^\beta_\nu$ with $M$ belonging to some extended symmetry group with changing signature, see $^{24}$. 

The invariant action for the $M$ field we can write as the usual action for a scalar field:

$$S_M = \int d^4x \, e \, M e^\mu \epsilon^\nu c^\gamma (D_G e M)_{\epsilon^\rho} F^\alpha^\beta F_{\gamma^\rho}$$

with

$$F^\alpha^\beta = \alpha_1 M^\alpha^\beta + \alpha_2 M^\alpha \gamma M^\beta^\rho + \alpha_3 M^\alpha^\rho M^\beta^\gamma.$$

A new, in comparison to the previous section, term of the action is a free action term of the $\Omega$ gauge field. Determining the field’s strength of the new gauge field

$$[D_G e M, D_G e M] = -G_{\mu\nu}$$

we define this action as

$$S_{\Omega} = \kappa \int d^4x \, e \, M e^\mu \epsilon^\nu \epsilon^\rho c^\gamma \text{tr} [G_{\mu\nu} G_{\mu_1\nu_1}] F^\alpha^\beta; \alpha^1 \beta^1$$

with

$$F^\alpha^\beta; \alpha^1 \beta^1 = \kappa_1 M^\alpha^\alpha M^\beta^\beta + \kappa_2 M^\alpha^\beta M^\beta^\alpha + \kappa_3 M^\alpha^\beta M^\beta^\gamma.$$

The action is similar, for example, to the QCD action in the curved space time, we do not consider a torsion and a cosmological constant terms in the action.

The dynamics of the theory given by Eq. (80) and Eq. (83) actions is non-linear and pretty complicated. Therefore, postponing the precise derivation for an additional publication, we can understand a dynamical signature in this variant of the theory by the following simple observations. First of all, we assume that for the Eq. (83) action there exists a classical solution for the gauge fields provided by the mechanism described in the Appendix B. We will have then:

$$\Omega^\alpha^\mu^\beta \text{cl} = \alpha_{\mu^\beta},$$

where the $\alpha_{\mu}$ fields, again, are known and satisfy some boundary conditions. This result, of course, is a consequence of the constant form of the vierbeins fields $e$ and $M$ in Eq. (83), we take these fields as normalized to the delta functions with respect to the corresponding indices in the first approximation. In this case the Eq. (83) will acquire the form of Eq. (B.2) action. Secondly, the classical solution of the Eq. (80) action is provided by the following equation:

$$D_G e M^\alpha^\alpha = 0$$

the solution can be written:

$$M^\alpha^\alpha (x) = M^\alpha^\beta_0 \left( P e^{-a_0} \int^z_{-\infty} dz^\alpha \alpha^\alpha_0 (z) \right)^\alpha_0^\beta.$$  

Again, the signature of the Eq. (69) metric is determined by the values of the boundary gauge fields $\alpha_{\mu}$

$$g_{\mu\nu} = \eta_{ab} M^a_{\mu} M^b_{\nu} \left( P e^{-a_0} \int^z_{-\infty} dz^\mu \alpha^\mu_0 (z) \right)^\alpha_0^\gamma \left( P e^{-a_0} \int^z_{-\infty} dz^\nu \alpha^\nu_0 (z) \right)^\beta_0^\gamma e_{\mu^\alpha} e_{\nu^\beta}$$

and in principle can be arbitrary, see Appendix A for the similar simple example.

VII. CONCLUSION

In this note we consider approaches where the signature of the metric is undefined and takes values in the field of complex numbers. We discuss a few possibilities for the definition of this type of metric, with signs of it’s components are not fixed in general. The change of the signature was widely discussed in the literature, see for example [1–7], for a description of the transition form Lorentzian to Euclidean manifold types in the quantum gravity and quantum cosmology. Nevertheless, mostly, this transition was introduced by the time’s coordinate Wick rotation. We, instead, propose the formalism where the domain of the metric’s signature is expanded. The metric in the proposed approaches is a dynamical object with signature determined by the complexification of the space-time manifold or by new gauge fields. Therefore, the signature can be changed smoothly between any predefined signatures, Lorentzian and Euclidean.
for example, with the help of the gauge fields. For example, introducing the Euclidean $\mathcal{A}$ and Lorentzian $\mathcal{B}$ gauge fields we can calculate an effective action defined in respect to these fields:

$$\Gamma(\mathcal{A}, \mathcal{B}) = \sum \mathcal{A}_1 \cdots \mathcal{A}_l \Gamma^{1 \cdots l; 1 \cdots k} \mathcal{B}_1 \cdots \mathcal{B}_k.$$  

(89)

The corresponding generating functional

$$Z(\mathcal{A}, \mathcal{B}) = Z_0^{-1} \int D\Phi e^{i\Gamma(\mathcal{A}, \mathcal{B})}$$  

(90)

with $\Phi$ as all other fields in the framework, will determine a transitional amplitude (S-matrix) between the manifolds with different signature. A calculation of those S-matrix elements we reserve for the future research. Another possibility for such S-matrix construction is an appearance of the $\mathcal{A}_l$ field as a semi-classical solution, i.e. saddle point, of the equations of motion for the total action in the generating functional, see [13, 23] for the corresponding discussion. In this case, instead the predefined induced values of the fields on corresponding boundaries, some dynamical transition from $\mathcal{A}$ field to the $\mathcal{B}$ must exist as result of a classical equations of motion for the gauge field. Namely, it is a problem of the existence of a classical two-valued boundary solution. Such solutions are known in high-energy scattering, see [25] for examples. This question we plan to investigate in an additional publication.

The simplest from the possibilities we consider is a direct complexification of the manifold by complexification of the manifold’s coordinates. The additional phase, i.e. additional coordinate, is factorized in the equations in this case with the help of the parameter assumed to be extremely small at the present:

$$a_\phi \propto \frac{l_0}{R_0}.$$  

(91)

As mentioned in the Introduction, the obvious choice of the lengths in the definition is $l_0$ as Planck length and $R_0$ a manifold’s curvature. A consequence of that is a factorization of the real and complex parts of the metric, i.e. factorization of real and complex parts of the corresponding complex manifold. Namely, the smallness of the parameter guarantees that the complexification is important and not small only when both parameters are of the same order, i.e. when $a_\phi \propto 1$. In this case we have to consider an eight dimensional manifold instead the four dimensional one. That situation is possible only at some extremal points of the manifold’s evolution. Otherwise the complexification is pure small distance effect, i.e. effect of quantum gravity. Namely, the proposed mechanism allows to determine the small contributions of the complex phases of an eight dimensional manifold to the quantities of the present classical four dimensional world. This is similar to the compatifications of the additional dimensions in the string theory but not quite the same of course. The main difference between these mechanisms of the account of the classically non-observable dimensions is the following. In the present framework the contribution of the additional dimensions is factorized and can be treated perturbatively if the almost flat manifold without strong gravity fields is considered, whereas in the string’s approach the contribution of the additional dimensions is always present, we can not take it equal to zero. Therefore, the framework with small $a$ does not require the compatification of the additional phase dimensions, instead it provides a smallness of their contributions in any expressions which we can treat perturbatively with respect to the parameter when the parameter is small.

The proposed complexification of the manifold through the complex coordinates is interesting also from the point of view of the symmetry group of the manifold. Namely, having the Poincare group representations as determination of the rule of the classification of the existing particles, the natural question is about the allowed transformation group of the new 4D complex manifold. For the $a_\phi \propto 1$ limit there are a plenty of possibilities for the group’s generalization, see an example and discussions in [17, 18]. Anyway the final step in the complex twistor construction is a projection of the extended group on the real slice endowed with Poincare group symmetry. The proposed case with small $a_\phi$ we treat differently. First of all, the $a_\phi \propto 0$ limit is well defined and determines a restoration of the Poincare symmetry. Secondly, considering the complex coordinates and expanding them with respect to $a_\phi$ we will obtain small corrections to the proposed classical transformations. Formally it means that to this precision it is enough to replace the real coordinates on the complex ones in the expressions for the group’s representations and algebra of the real manifold and expand them in respect to $a_\phi$. In this case some quantum corrections will arise in the expressions of interests. Still, the symmetry group for the initial manifold can be any corresponding to the complex Minkowski space symmetry, see [17, 18]. This symmetry in the framework will restore at the $a_\phi \propto 1$ limit, of course in this case the perturbative expansion can not be used. Therefore, in the proposed framework we discuss not the projection of the complex manifold on the real slice, but a small complex corrections to the real metric, i.e. we consider a general framework with Lorentzian and different signatures metrics coexisting.\footnote{The twistor space, definitely, as well describes manifolds endowed with metrics with different signatures, but it is not clear if it can be formulated as a dynamical model with simultaneous inclusion of metrics of different signatures.}
metric with Lorentz signature is defined as a classical one and the metrics with other signatures contribute only at some special condition or at the quantum level.

In any case, the complexification of the gravity action by the complexification of the coordinates results in the additional part to the "bare", real Einstein-Cartan action. This additional part provides a complex part to the classical "bare" vierbein and consequently a complex additional part to the usual metric. For this type of the complexification we need to separate two cases. When we introduce a global phase factor for the coordinates then the additional metric's part is a complex fluctuation above the usual metric, see Eq. (20). In general, for the non-expanded with respect to $a_\phi$ metric, the action is a functional of the complex Eq. (20) Lagrangian, that in fact is not unusual, see [19] and [23]. The interesting question, therefore, is a proper definition and properties of such complex action in the path integral, see discussions in [23]. Introducing a local complex phase in the definition of the complex coordinates, see Eq. (22)–Eq. (53), we introduce new gauge fields and their symmetry group G, in this case the $a_\phi$ parameter can be considered as a coupling constant of the group. This complexification of the manifold is more complicated than the first one of course, there is a possibility to obtain again complex fluctuations above the real metric, see Eq. (143), but, additionally, there is a possibility to introduce a metric with non-determined signature from the very beginning with the help of the Eq. (62) $M$ field. This last case is non-perturbative and complicated. We do not discuss it much in the paper postponing it for an additional publication.

There are following interesting properties of the actions Eq. (20) and Eq. (53) we obtained. First of all, there is no preferable axis of time direction, the metric's component can be of any sign in the situation with an undefined signature and any coordinate can serve as the time coordinate therefore. Moreover, fixing the metric, as usual Lorentzian for example, we still have a freedom to change the phases of the metric's components, i.e. to rotate the coordinate axes determining infinitely many ways of a foliation of the space-time. Each of these possibilities is described by the same action and, therefore, provides the same physics. In this case the preferable signature can be given by a random selection from the infinitely many possibilities or by some fixation procedure similar to some extent to the spontaneous symmetry breaking. The later is possible when we talk about the complexification by the gauge fields. Namely, in this case there is a possibility to define the classical values of the fields as a projection of some predefined boundary fields. The approach is described in Appendix H and the procedure provides a signature of the bulk by the value of the gauge fields on the boundaries of the manifold. The mutual property of all the actions with gauge fields involved is an appearance of the new factor in the front of the actions which is a determinant of the gauge fields. The value of the factor is determined by the boundary values of the fields and defines the relative weight of the action in the corresponding generating functional, otherwise it is arbitrary. The dynamics of such complex systems with many different parts of the general action in the generating functional is not clear and requires an additional investigation.

A different way to introduce the dynamical signature of the metric is a generalization of the tangent space and an introduction of an additional, auxiliary, metric in the tangent space which makes the tangent space curved. This can be achieved by the complexification of the vierbein with the help of the gauge fields, see Eq. (63) and Eq. (53), or by the direct definition of the usual vierbein fields as projection of some "gauge" vierbein performed by the gauge field of some symmetry group G, see Eq. (67). In both cases the final signature is dynamical and determined directly by the gauge fields, see Eq. (69), or by scalar fields and gauge fields together, see Eq. (88). Again, for these mechanisms the projection procedure of Appendix H is important. Without it the dynamics and correspondingly the signature can be arbitrary. Namely, as in the previous cases, there is neither preferable geometrical time nor preferable spatial coordinates and the given and only foliation in the approaches must be fixed separately, if required. We did not consider a matter issue in the frameworks, see [23] for the discussion about the possibilities of a proper definition of matter fields for the manifolds with complex metric. It is interesting to understand in general how the quantum matter fields behave in respect to the change of signature of the manifold and if there exists some dynamical mechanisms which relate a foliation of the space-time and it's signature with properties of the matter. This problem requires an additional research and clarification of the properties and definition of the matter fields in respect to the manifold's symmetries and signature.

Some interesting questions we can ask are about an existence of the different time's arrows directions in the manifolds with different signatures and corresponding issues related to it. First of all, we note that if we stay in the framework of a perturbative approach with respect to the $a_\phi$ parameter, the possible additional contributions to any quantity of interests are extremely small. There is only one time's arrow on the classical level. Namely, the additional contributions are effectively pushed in the region of the quantum gravity regime, therefore any statements about the behavior of the time's arrow at this scale must operate with a quantum gravity theory which we have no. Nevertheless, if we assume that the proposed approach correctly describes the quantum gravity regime or at least some of it's details, we can conclude that on this quantum level it possible that there is no any preferable time or spatial directions, unless some mechanisms fix the corrections to the metric as real with preferable signature. Considering the causality as a definition of the form of the corresponding propagators, we have no any problems with that till we do not consider some special regimes when $a_\phi$ is not small. In turn, an arbitrary value of the $a_\phi$ means the non-perturbative calculations for the eight dimensional manifold with an arbitrary metric's tensor which has complex
components. The definition of the propagators in this case and their reduction to the usual ones is very interesting question which we hope to explore in the future.

Another important problem is about a co-existing of the different regions with different signatures and possible different time’s arrows directions outside the perturbative regime. In this case we have two possibilities. The first one was considered in [29], there a case without the complex coordinates and/or metric was discussed with some separation arises between the Euclidean and Lorentzian regions in a form of a hypersurface. In this set-up the hypersurface plays a role of a domain wall which separates the regions with different signatures and, in some extend, it defines an initial or final singularities for the time’s arrows, see details in [29]. As it seems, such hypersurfaces are unavoidable in the situation with co-existing of real metrics with different signatures, see also [24, 30]. From the point of view of QFT, it can be considered also as Lorentzian space-time ↔ Euclidean space-time geometrical transition vertices between separated parts of some mutual manifold. Due the discontinuities of the Einstein’s tensor components on the hypersurface of separation, we again have no any problems with causality. We have two classically disconnected regions of space-time which possible connection can be perhaps established only on the level of quantum gravity effects. More complicated picture arises when we allow the complexification of the metric through some mechanisms. In this case we have no separating hypersurface between regions with different signatures due the complex phases of the metric. In this situation we again obtain some eight dimensional manifold with two, or more, different time coordinates exist simultaneously, if the time’s arrows can be defined for the metric’s tensor with arbitrary complex components of course. The notion of the causality in this case and possible mechanisms of the reduction of the manifold to the usual four dimensional one with one time’s arrow are complicated problems which requires an additional investigation.

Discussing the applications of the proposed approaches we note that they can be useful in an investigation of different aspects of the topology transition in both quantum gravity and cosmology through Eq. (90) expression for example. In the paper we discussed a few possible mechanism of the Eq. (59) effective action construction. It is interesting to understand which one can be realized in the nature. For that we need to understand the dynamics of the models with matter fields included. Namely, there is an interesting problem to determine the form of the action for the spinor of scalar fields in the new curved spaces of different signatures and investigate the dynamics of these fields in corresponding models, see different aspects of this problem in [8–13, 26] references. Another interesting application of the dynamical signature is a clarification of it’s possible correspondence to the new approaches to the classical gravity introduced and discussed at the last decade, see for example [31–41]. We find an investigations of these ideas and possibilities very interesting.
Appendix A: Complex metric through complex vierbein

In order to illustrate how the Eq. (67) and Eq. (69) construction reproduce the [14] set-up, we firstly can consider as example the action of the following two-dimensional fixed unitary matrix

\[ M = \begin{pmatrix} 1 & 0 \\ 0 & i \end{pmatrix}, \quad \tilde{M} = M^T = \begin{pmatrix} 1 & 0 \\ 0 & -i \end{pmatrix}, \quad \tilde{M} M = 1 \]  
(A.1)

on flat metric:

\[ M \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} M = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}. \]

Therefore, considering as \( G \) the \( U(4) \) group for example, we will obtain complex phases for the metric’s components. Namely, consider the spectral decomposition for the unitary matrix

\[ M^{a\alpha} = \sum_{i=1}^{4} \lambda_i u^a_i \tilde{u}^\alpha_i \]  
(A.2)

with \( \lambda \) and \( u \) as eigenvalues and eigenvectors, we obtain for the metric:

\[ g_{\mu\nu} = \frac{1}{2} \sum_{i,j=1}^{4} \lambda_i \lambda_j \left( u^a_i \eta S_{ab} u^b_j \right) \tilde{u}^\alpha_i \tilde{u}^\beta_j \left( e_{\mu\alpha} e_{\nu\beta} + e_{\nu\alpha} e_{\mu\beta} \right). \]  
(A.3)

Defining the local set of the vierbein through the identities

\[ e_{\mu\alpha} \tilde{u}^\alpha_i = \delta_{\mu i} \]  
(A.4)

we will have finally

\[ g_{\mu\nu} = \frac{1}{2} \lambda_\mu \lambda_\nu \left( u^a_\mu \eta S_{ab} u^b_\nu + u^a_\nu \eta S_{ab} u^b_\mu \right). \]  
(A.5)

Now, if we restrict ourselves by the diagonal unitary matrices, the corresponding eigenvectors are real and orthonormal. Therefore, for the arbitrary four dimensional diagonal unitary matrix

\[ M^{a\alpha} = \begin{pmatrix} e^{i \alpha_1} & 0 & 0 & 0 \\ 0 & e^{i \alpha_2} & 0 & 0 \\ 0 & 0 & e^{i \alpha_3} & 0 \\ 0 & 0 & 0 & e^{i \alpha_4} \end{pmatrix} \]  
(A.6)

we obtain a simple expression for the generalized flat metric:

\[ g_{\mu\nu} = e^{i (\alpha_\mu + \alpha_\nu)} \eta S_{\mu\nu}. \]  
(A.7)

We see, that in terms of Eq. (67) transform we simply can write the vierbein transform as

\[ e^a_\mu = e^{i \phi_a} \delta^{a\alpha} e_{\mu\alpha} \]  
(A.8)

obtaining for the metric

\[ g_{\mu\nu} = e^{i (\phi_a + \phi_b)} \eta S_{ab} e^a_\mu e^b_\nu \]  
(A.9)

which describes, at a first sight, a metric with indefinite complex signature. Nevertheless, we remind that the \( \phi \) angles are dynamical fields in the approach, therefore the final leading order expression for the metric will be determined by the classical values of these fields.
Appendix B: Induced part of the action

Following [27, 28] we consider the action given in Eq. (B2). For this action in the flat-space time we have:

$$S_A = -\frac{1}{4} \int d^4x \ tr [G_{\mu\nu} G^{\mu\nu}]$$  \hspace{1cm} (B.1)

and for the induced part of the action

$$S_{ind} = -\sum_i \int d^4x \ tr \left[ (\partial_\mu O(A_\mu)) \left( \partial_\nu^2 \mathcal{A}_i^{\mu} \right) \right],$$  \hspace{1cm} (B.2)

where the Riemann indexes are summed up through the Minkowski metric as usual. The $\mathcal{A}^{\mu}$ fields are defined at the boundaries, and they intend to provide the signature of the additional part in the Eq. (56) metric. For example, we can define the two complete sets of the boundary fields which satisfy

$$\partial_\mu \mathcal{A}_i^{\mu} = 0$$  \hspace{1cm} (B.3)

and the following boundary conditions:

$$\begin{aligned}
\mathcal{A}_1^{\mu}(x) &\to 0 \quad x^0 \to \infty, \\
\mathcal{A}_2^{\mu}(x) &\to 0 \quad x^0 \to -\infty.
\end{aligned}$$  \hspace{1cm} (B.4)

Also the following term must be added to the action

$$S_{\mathcal{A}} = \sum_i \int \mathcal{A}_i^{\mu} \partial_\nu^2 \mathcal{A}_\mu i$$  \hspace{1cm} (B.5)

which preserves the correct form of the propagators in the full action, see discussions in [27, 28]. Therefore, for the gauge fields

$$\partial_\mu A^{\mu} = 0$$  \hspace{1cm} (B.6)

we obtain as a solution of the equations of motion

$$A_{\mu}^{cl} = \mathcal{A}_1^{\mu} + \mathcal{A}_2^{\mu}.$$  \hspace{1cm} (B.7)

The operator $O$ in the Eq. (B2) action is defined similarly to definitions of [27, 28]. In the simplest variant it is

$$O(A_\mu) = \frac{1}{a_\phi} \ P e^{a_\phi \int_{-\infty}^{\infty} dx^\mu A_\mu(x')} C(R).$$  \hspace{1cm} (B.8)

There is no summation on $\mu$ index in the ordered exponential and the index is fixed in correspondence to the Eq. (B2) expression, $C(R)$ is the eigenvalue of Casimir operator in the representation $R$ for the chosen gauge symmetry group. The different form of this operator and discussion about can be found in [27, 28].

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7 Following the analogy with the high energy scattering approach, we can consider the Eq. (B1) action with an additional induced term as describing a “scattering” between two boundary fields with boundaries defined at the edges of time.
9, 630, (2017).

[28] L. N. Lipatov, Nucl. Phys. B 365, 614 (1991); L. N. Lipatov, Theor. Math. Phys. 169, 1370 (2011); L. N. Lipatov, Phys. Part. Nucl. 44, 391 (2013); L. N. Lipatov, Subnucl. Ser. 50, 213 (2014); L. N. Lipatov, EPJ Web Conf. 125, 01010 (2016); L. N. Lipatov, EPJ Web Conf. 164, 02002 (2017); S. Bondarenko, S. Pozdnyakov and M. A. Zubkov, Eur. Phys. J. C 81, no.7, 613 (2021).

[29] S. Bondarenko and V. De La Hoz-Coronell. [arXiv:2204.07528] [gr-qc].

[30] D. Kothawala, Class. Quant. Grav. 35 (2018) no.3, 03LT01; D. Kothawala, Phys. Rev. D 97 (2018) no.12, 124062; R. Singh and D. Kothawala, [arXiv:2010.01822] [gr-qc].

[31] S. Bondarenko, Mod. Phys. Lett. A 34, no. 11, 1950084 (2019); Universe 6, no.8, 121 (2020); S. Bondarenko, Eur. Phys. J. C 81, no.3, 253 (2021).

[32] M. Villata, EPL 94, no. 2, 20001 (2011); Annalen Phys. 527, 507 (2015); N. Debergh, J. P. Petit and G. D’Agostini, J. Phys. Comm. 2, no.11, 115012 (2018); H. Socas-Navarro, Astron. Astrophys. 626, A5 (2019); G. J. Ni, Rel. Grav. Cosmol. 1 (2004), 123-136.

[33] G. Chardin, Hyperfine Interact. 109, no. 1-4, 83 (1997). J.P.Petit, G.D’Agostini Astrophysics And Space Scicence, A 29, 145-182 (2014); R. J. Nemiroff, R. Joshi and B. R. Patla, JCAP 1506, 006 (2015); G. Kofinas and V. Zarikas, Phys. Rev. D 97, no. 12, 123542 (2018); G. Manfredi, J. L. Rouet, B. Miller and G. Chardin, Phys. Rev. D 98, 023514 (2018); G. Chardin and G. Manfredi, Hyperfine Interact. 239, no.1, 45 (2018).

[34] J.-P. Petit, Astrophys. Space Sci. 226, 273 (1995); J. P. Petit and G. d’Agostini, [arXiv:0903.1362] [math-ph]; J. Petit and G. D’Agostini, Mod. Phys. Lett. A 29, no.34, 1450182 (2014); J. P. Petit and G. d’Agostini, Astrophys. Space Sci. 354, no. 2, 2106 (2014); J. P. Petit and G. D’Agostini, Astrophys. Space Sci. 357, no.1, 67 (2015); J.P.Petit and G.D’Agostini, Mod. Phys. Lett. A 30, no.9, (2015); G. D’Agostini and J.P.Petit, Astrophysics and Space Science, (2018); N. Debergh, J. P. Petit and G. D’Agostini, J. Phys. Comm. 2, no.11, 115012 (2018).

[35] S. Hossenfelder, Phys. Lett. B 636, 119-125 (2006); S. Hossenfelder, [arXiv:gr-qc/0605083] [gr-qc].

[36] A. A. Baranov, Izv. Vuz. Fiz. 11, 118 (1971); A. D. Dolgov, [arXiv:1206.3725] [astro-ph.CO].

[37] L. Boyle, K. Finn and N. Turok, Phys. Rev. Lett. 121, no. 25, 251301 (2018); J. L. Alonso and J. M. Carmona, Class. Quant. Grav. 36, no. 18, 185001 (2019).

[38] D. E. Kaplan and R. Sundrum, JHEP 0607, 042 (2006); A. Hebecker, T. Mikhail and P. Soler, Front. Astron. Space Sci. 5, 35 (2018).

[39] A. D. Linde, Phys. Lett. B 200, 272 (1988).

[40] V. Gorini, A. Kamenshchik, U. Moschella, Phys. Rev. D 67, 063509 (2003); Z. Keresztes, L.A. Gergely, A.Y. Kamenshchik, Phys. Rev. D 86, 063522 (2012); Z. Keresztes, L.A. Gergely, A.Y. Kamenshchik, V. Gorini, D. Polarski, Phys. Rev. D 88, 023535 (2013); A.Y. Kamenshchik, Class. Quantum Gravity 30, 173001 (2013); A.Y. Kamenshchik, E.O. Pozdeeva, S.Y. Vernov, A. Tronconi, G. Venturi, Phys. Rev. D 95(8), 083503 (2017); A.Y. Kamenshchik, Found. Phys. 48(10), 1159–1176 (2018); O. Galkina, A.Y. Kamenshchik, Phys. Rev. D 102(2), 024078 (2020).

[41] I. Ben-Dayan, M. Hadad and A. Michaelis, [arXiv:2110.06240] [hep-th].