Magic Supergravities, $N = 8$ and

Black Hole Composites

Sergio Ferrara$^1$, Eric G. Gimon$^2$ and Renata Kallosh$^3$

$^1$ Physics Department, Theory Unit, CERN, 1211 Geneva 23, Switzerland
INFN, Laboratori Nazionali di Frascati, Via Enrico Fermi, 40,00044 Frascati, Italy
Department of Physics & Astronomy, University of California, Los Angeles, CA 90095, USA
Sergio.Ferrara@cern.ch

$^2$ Department of Physics, University of California, Berkeley, CA 94720, USA
Theoretical Physics Group, LBNL, Berkeley, CA 94720, USA
eggimon@lbl.gov

$^3$ Physics Department, Stanford University, Stanford CA 94305-4060, USA
kallosh@stanford.edu

ABSTRACT

We present explicit U-duality invariants for the $\mathbb{R}, \mathbb{C}, \mathbb{Q}, \mathbb{O}$ (real, complex, quaternionic and octonionic) magic supergravities in four and five dimensions using complex forms with a reality condition. From these invariants we derive an explicit entropy function and corresponding stabilization equations which we use to exhibit stationary multi-center 1/2 BPS solutions of these $N = 2$ $d = 4$ theories, starting with the octonionic one with $E_{7(-25)}$ duality symmetry. We generalize to stationary 1/8 BPS multicenter solutions of $N = 8$, $d = 4$ supergravity, using the consistent truncation to the quaternionic magic $N = 2$ supergravity. We present a general solution of non-BPS attractor equations of the STU truncation of magic models. We finish with a discussion of the BPS-non-BPS relations and attractors in $N = 2$ versus $N = 5, 6, 8$. 
1 Introduction

The purpose of this paper is to present explicit stationary multicenter solutions of $N=8$, $d=4$ supergravity [1] and all magic $N=2$ supergravities [2]. All these models are associated with the Jordan algebras of $3 \times 3$ Hermitian matrices: $J_3^\mathbb{R}$, $J_3^\mathbb{C}$, $J_3^\mathbb{Q}$, $J_3^\mathbb{O}$, and $J_3^\mathbb{O}_1$. Here $A = \mathbb{R}, \mathbb{C}, \mathbb{Q}, \mathbb{O}$.
are the four division algebras, with $\dim A = 1, 2, 4, 8$, while $\mathbb{O}_s$ is a the split form of $\mathbb{O}_3$ with a quadratic norm invariant under $O(4, 4)$ (the indefinite signature means this is no longer a division algebra). The octonionic magic $N = 2$ model goes by the name exceptional since it is the only one of the four magic $N = 2$ supergravities which is not known to be a consistent reduction of $N = 8$: its defining Jordan algebra $J^5_3$ involves the real octonions with a quadratic norm invariant under $O(8)$, and is the only algebra mentioned above which one cannot get from truncating $J^5_3$. Note that the connection to Jordan algebras based on the division algebras defines a relation between the four $N = 2$ magic supergravities and the magic square. From a physics perspective, the main feature of the $N = 8$ and the magic $N = 2$ supergravities which interests us is that for each case all the vector fields, including the graviphoton, transform in a single irreducible representation of U-duality group; these symmetries place strong constraints on the entropy formula and the stabilization equations derived from it.

The black hole entropy formula for $N = 2$ supergravities based on symmetric spaces is either known or can be established. We extend this list to include magic supergravities, for which we construct these entropies. We can then give a complete list of all symmetric spaces with their entropy formulas.

The duality symmetry of $N = 8$, $d = 4$ and $d = 5$ supergravities is $E_{7(7)}$ and $E_{6(6)}$, respectively. The duality symmetry of the exceptional magic $N = 2$, $d = 4$ and $d = 5$ supergravities is $E_{7(-25)}$ and $E_{6(-26)}$, respectively. The number in the brackets stands for the difference between the number of non-compact minus compact generators. For example, in $E_{7(7)}$ and in $E_{7(-25)}$ the total number of non-compact and compact generators of $E_7$ is the same, namely, $70 + 63 = 133$ and $54 + 79 = 133$. However, the difference between them is either $70 - 63 = 7$ for $E_{7(7)}$ or $54 - 79 = -25$ for $E_{7(-25)}$.

For the octonionic $N = 2$ supergravity we will find a new explicit entropy function derived from the quartic invariant of the fundamental 56-dimensional representation of $E_{7(-25)}$ and the relevant “electric” and “magnetic” cubic invariants of $E_{6(-26)}$. This entropy function then provide us with the most general explicit multicenter BPS solution of the octonionic (exceptional magic) $N = 2$ supergravity in terms of a 56-dimensional harmonic function with an arbitrary number of centers. We will also present explicit entropy formulas for all the remaining magic supergravities.

Recently, the generalization of the $N = 2$ attractor equations in [5]-[8] was established for $N > 2$ by a generalization of the special geometry symplectic structure to all extended supergravities [9]. In particular, for $N = 8$ supergravity a simple set of algebraic equations was established which describe regular BPS and non-BPS extremal black hole solutions. In this later paper, a natural truncation of the $N = 8$ theory to it’s largest consistent $N = 2$ truncation, the quaternion magic $N = 2$ supergravity, made it’s appearance. To proceed with an explicit demonstration of stationary multicenter 1/8 BPS solutions of $N = 8$ $d = 4$ supergravity we use this reduction to $N = 2$. Hence, in addition to the $N = 8$ model and the octonionic $N = 2$ model, the N=2 magic supergravity based on the Jordan algebra of quaternions, $J^5_3$, is of particular interest in this paper.
We will argue that our new solution of quaternionic supergravity, depending on a 32-dimensional harmonic function with any number of centers, are also the most general 1/8 BPS stationary multicenter solution of $N = 8$ $d = 4$ supergravity, up to an overall $E_{7(7)}$ rotation.

We also will analyse multiple relations between BPS and non-BPS solutions in $N = 2$ and $N = 5, 6, 8$ supergravities. It is known that the same bosonic solution may be BPS or non-BPS depending on how it is embedded in a given supergravity theory or how one identifies the bosonic vector fields with the specific supersymmetric multiplets, see for example [10] where many such examples were given. In some $N = 8$ examples in [10] the relation between BPS and non-BPS solutions uses the same embedding but requires a flip of a sign from some of the charges. More recently such examples were discovered in the context of non-supersymmetric attractors, see e. g. [11].

We will describe the general solutions of $N = 2$ attractors and, in particular, the general solution for non-BPS attractors of magic supergravities.

In summary, our paper is organized as follows. In section 2 we summarize and expand on known results for multi-center solutions in $N = 2$ SUGRA with special emphasis on the role of the a set of harmonic functions transforming in an irreducible representation of the U-duality group. Writing down the metric, gauge and scalar fields depends crucially on a writing down a quartic invariant, or an entropy function, constructed from this representation. In section 3 we describe two approaches for writing down the quartic invariants for $N = 8$ and the magic $N = 2$ supergravities. The first approach uses connections to the five-dimensional U-duality group and makes novel use of a set of complex matrices with a reality condition in expressing both the entropy function and the attractor equations. The second approach stresses the connection of our quartic invariant to a set of invariants of the six-dimensional U-duality group which allows us to use real matrices (the quaternionic magic case is a pseudo-real exception). In section 4 we develop, as an example for the material in sections 2 and 3, the solutions for BPS composites of octonionic magic $N = 2$. Section 5 follows with a description of N=8 1/8 BPS composites in terms of a quaternionic $N = 2$ sub-algebra. In section 6 we expand our circle of consideration to non-BPS extremal multi-center solutions before closing with some final thought in section 7.

2 A description of stationary multicenter solutions: N=2 black hole composites

The exact BPS multicenter stationary black hole solutions in $N = 2$ supergravity have been worked out as a general case whenever there is a known expression for the explicit single center black hole entropy, $S(p, q)$, as a function of quantized charges [12]. The entropy formula is given by the minimal value of the BPS black hole mass via

$$S(p, q) = \pi M^2(p, q; t, \bar{t})_{\text{attr}} = \pi I_1(p, q),$$

(2.1)

where the moduli, $t(p, q), \bar{t}(p, q)$, are fixed near the black hole horizon by the attractor mechanism [5]-[8]. There are three invariants that the reader should be aware of in the context of
black hole attractors, it is important not to confuse them:

- \( I_1((p, q; t, \bar{t})) \): This is the general symplectic invariant defined in \([13, 6]\). It can be written explicitly as:

\[
I_1(p, q; t, \bar{t}) = -\frac{1}{2}(p \bar{q})M(N(t, \bar{t}))(p \bar{q})^T = |Z(\Gamma)|^2 + |\mathcal{D}Z(\Gamma)|^2.
\] (2.2)

where \( N(t, \bar{t}) \) is the metric on the vector fields. Here \( Z(\Gamma) = \langle \Gamma, \Omega \rangle \) and \( \Gamma = (p^A, q_A) \), \( \Omega = (L^A, M_A) \). This scalar quantity is an invariant of the group \( OSP(2(n_v + 1), \mathbb{R}) \) which acts on both the charges and the symplectic sections. It is manifestly quadratic in the charges and non-negative. Near the black hole horizon \( N = 2 \) supersymmetry is restored and \( \mathcal{D}Z = 0 \) which is equivalent to the requirement that

\[
\partial_t I_1(\Gamma; t, \bar{t}) = 0 \quad \partial_{\bar{t}} I_1(\Gamma; t, \bar{t}) = 0 \tag{2.3}
\]

This equation defines the moduli near the horizon as functions of charges, \( t(p, q), \bar{t}(p, q) \).

- \( I_1(p, q) = I_1(p, q; t(p, q), \bar{t}(p, q)) \): This is the entropy function, \( I_1(\Gamma) \), derived from the function above restricted to \( t \)'s and \( \bar{t} \)'s implicitly defined in terms of \( p \)'s and \( q \)'s through their attractor values. This scalar is no longer invariant under \( OSP(2(n_v + 1), \mathbb{R}) \) since we have chosen particular values of \( t \) and \( \bar{t} \) for given \( p \)'s and \( q \)'s. It is, however, invariant under the U-duality group if such exists, but it has no particular polynomial properties as a function of the charges. The moduli near the horizon are defined by equation \([6]\)

\[
p^\Sigma + i \frac{\partial I_1(p, q)}{\partial q^\Sigma} = 2i\bar{Z}(\Gamma)L^\Sigma, \quad q^\Sigma - i \frac{\partial I_1(p, q)}{\partial p^\Sigma} = 2i\bar{Z}(\Gamma)M^\Sigma
\] (2.4)

which follows from \( \mathcal{D}Z(\Gamma) = 0 \) at the horizon. The holomorphic special coordinates \( t^A = \frac{x^A}{\varphi_0} \) at the horizon follow by dividing eq. (2.4) on the zero component of the same equation.

\[
t^A(p, q) = \frac{p^A + i \frac{\partial I_1(\Gamma)}{\partial q^A}}{p^0 + i \frac{\partial I_1(\Gamma)}{\partial \varphi_0}} \tag{2.5}
\]

- \( J_4(p, q) \): This is a quartic polynomial in the charges, invariant under a U-duality group. It appears in all the theories discussed in this paper, it is related to the entropy function via \( I_1(p, q) = \sqrt{|J_4(p, q)|} \).

Using the entropy function, the equations of motion for general black hole solutions in \( N = 2 \) supergravities were solved in \([12, 14]\). Examples of non-static multicenter solutions were given earlier in \([13]\) where it was also explained on the basis of these examples that the most general expression for the metric can only depend on duality invariants, the Kähler potential \( K(X, \bar{X}) \) and the Kähler connection \( A_\mu(X, \bar{X}) \). The general form of multi-center BPS black hole solutions was presented in \([14]\).
The explicit BPS multicenter stationary black hole solutions in $N = 2$ supergravity were found in [12, 14]. One can solve the 1/2 BPS equations by introducing a harmonic symplectic doublet:

$$H(\vec{x}) = (H^\Lambda, H_\Lambda) = h + \sum_{s=1}^n \frac{\Gamma_s}{|\vec{x} - \vec{x}_s|}. \quad (2.6)$$

From BPS equations one can prove that this doublet is proportional to the imaginary part of the covariantly holomorphic symplectic section $\Omega = (L^\Lambda, M_\Lambda)$:

$$H(\vec{x}) = 2\text{Im} \langle H(\vec{x}), \Omega \rangle \Omega \equiv i(L^\Lambda M_\Lambda - H_\Lambda L^\Lambda)\Omega, \quad \langle \Omega, \overline{\Omega} \rangle = -i \quad (2.7)$$

We will refer to $h$ and $\Gamma$ as fundamentals since for the theories which appear in our paper they transform as fundamentals under the $E_7$ U-duality group. A symplectic invariant $I_1(\Gamma, t, \overline{t})$ is now replaced by the $\vec{x}$-dependent symplectic invariant $I_1(H(\vec{x}), t, \overline{t})$. The other basic ingredient is the symplectic pairing of two charges which induces one on the harmonic functions:

$$\langle \Gamma_1, \Gamma_2 \rangle = p_1^A q_2^A - p_2^A q_1^A \leftrightarrow \langle H_1, H_2 \rangle.$$

Once equipped with these structures, one can prove that the BPS equations require that

$$\langle H, D\Omega \rangle = 0 \quad (2.8)$$

which results in

$$\partial_t I_1(H(\vec{x})); t, \overline{t}) = 0 \quad \partial_{\overline{t}} I_1(H(\vec{x})); t, \overline{t}) = 0 \quad (2.9)$$

This is the stabilization equation for the symplectic invariant which can also be translated into a relation

$$I_1((H(\vec{x})); t, \overline{t})|_{H(\vec{x}), D\Omega)=0 = I_1((H(\vec{x})) \quad (2.10)$$

This leads to

$$H^\Sigma + i \frac{\partial I_1(H)}{\partial H^\Sigma} = 2i \overline{Z}(H)L^\Sigma, \quad H_\Sigma - i \frac{\partial I_1(H)}{\partial H_\Sigma} = 2i \overline{Z}(H)M_\Sigma \quad (2.11)$$

The special coordinates $t^\Lambda = \frac{X^\Lambda}{X^0}$ solve the equations of motion, stabilization equations of the same form as the attractor equations for moduli near the horizon, and are given by

$$t^\Lambda(\vec{x}) = \frac{H^\Lambda + i \frac{\partial I_1(H)}{\partial H^\Lambda}}{H^0 + i \frac{\partial I_1(H)}{\partial H^0}} \quad (2.12)$$

The stationary metric is

$$ds_4^2 = -I_1^{-1}(\vec{x})(dt + \omega d\vec{x})^2 + I_1(\vec{x})d\vec{x}^2, \quad \nabla \times \omega = \langle H, \nabla H \rangle. \quad (2.13)$$

Note that $e^{-2K(X, \overline{X})}$ may or may not be equal to $I_1(\vec{x})$, depending on the choice of the Kähler gauge. The symplectic invariant $I_1$ is invariant under Kähler transformations, $K(t, \overline{t}) \rightarrow K(t, \overline{t}) + f(t) + \overline{f}(\overline{t})$, whereas the Kähler potential can be changed by a sum of the holomorphic and anti-holomorphic function. In [14] an ansatz was used with $K = -2U$ and
\[ H = i\Omega_0 - i\overline{\Omega}_0, \] where \( \Omega_0 = (X^\Lambda, F_\Lambda) \) is the holomorphic section. This section also transforms under Kähler transformations, \( \Omega_0 \rightarrow \Omega_0 e^{-f(t)} \). In [12] the ansatz for the harmonic doublet is \[ H = ie^{-U+K/2}(e^{-i\alpha\Omega_0} - e^{i\alpha\overline{\Omega}_0}) \] where \( \alpha \) is the argument of the central charge. We prefer to codify the solutions by the choice of the harmonic function at infinity, \( h \). Using this boundary value, we can give the values of the special coordinates, the metric and vector field strength and do not have to define combinations which are not invariant under Kähler transformations. For example, the relation between our \( h \) and the quantities in [12] is

\[ h = e^{-i\alpha/2\Omega_0} \]

In fact, only the product of all 3 terms is invariant under Kähler transformations. Therefore we can simply codify our solution with arbitrary \( h \). Note also that in our equation (2.11) the analogous term \( \overline{Z}(H)L^2 \) is Kähler invariant since the transformations on \( \overline{Z}(H) \) and on \( L \) cancel and we do not have to identify these terms separately.

The integrability condition for the multi-center solution is

\[ \langle H, \Delta H \rangle = 0 \quad \iff \quad \langle H(x_s), \Gamma_s \rangle = 0 \quad \iff \quad \sum_{t=1}^{n} \frac{\langle \Gamma_s, \Gamma_t \rangle}{|x_s - x_t|} + \langle \Gamma_s, h \rangle = 0 \quad (2.14) \]

Hence when our generic set of \( n+1 \) fundamentals, \( (\Gamma_s, h) \), satisfies the integrability conditions, we constrain the distances between the \( \Gamma_s \) fundamentals, \( |x_s - x_t| \). Only when two of these fundamental charge vectors are mutually local is their relative distance unconstrained. For the case when all the fundamentals are mutually local there are no constraints on the distances; they are just moduli.

The fundamental which sets the asymptotic behavior, \( h \), also must satisfy some conditions if the metric is to have the correct normalization and the integrability conditions are to be solved:

\[ I_1(h) = 1, \quad <h, \Gamma> = 0, \quad (\Gamma = \sum_s \Gamma_s) \quad (2.15) \]

The vector fields in electric basis, using a mixture of spherical coordinates \( (r_s, \theta_s, \phi_s) \) around each center \( x_s \), are:

\[ A_{\Lambda} = \partial_{H^\Lambda} \left( \ln I_1(H) \right)(dt + \omega) - \sum_s \cos \theta_s d\phi_s \otimes \Gamma_{\Lambda}. \quad (2.16) \]

The features of the geometry at very large \( |x| \gg |x_s - x_t| \) can be read off from the asymptotic form of the solution. In this approximation

\[ H(x) \approx h + \frac{\Gamma}{|x|}, \quad \Gamma = \sum_{s=1}^{n} \Gamma_s \quad (2.17) \]

and

\[ \bar{\omega} \approx \sum_{s<t} \langle \Gamma_s, \Gamma_t \rangle \bar{e}_{st} \times \frac{x_s}{|x|}, \quad \bar{e}_{st} = \frac{x_s - x_t}{|x_s - x_t|} \quad (2.18) \]

Far from the core, this looks like a spherically symmetric black hole (see [16] for more details):

\[ ds^2_4 = -e^{2U} dt^2 + e^{-2U} d\bar{x}^2, \quad e^{-2U(\bar{x})} = I_1(h + \frac{\Gamma}{|x|}) \quad (2.19) \]
whose entropy is equal to

\[ S(\Gamma) = \pi I_1(\Gamma). \] (2.20)

Note that any measured angular momentum derived from our expansion of \( \omega \) in (2.18) comes from terms which are order \( 1/|\vec{x}|^2 \) in the metric and should be matched with a measurement at a similar order for the dipole moments of the gauge fields in (2.16). Thus there is no contradiction with our knowledge that simple charged supersymmetric black holes have no angular momentum; if you measure angular momentum you also measure dipoles.

To recap, we would like to stress the special features of these \( N = 2 \) solutions. For a solution with \( n \) centers we needed \( n + 1 \) constant fundamentals (symplectic doublets):

\[ h = (h^\Lambda, h_\lambda), \quad \Gamma_s = (p^\Lambda, q_\Lambda)_s \quad s = 1, \ldots, n. \] (2.21)

giving these important qualities to our solutions:

- **An attractor at infinity:** This means that the values of all our special coordinates at infinity are specified by the values of the 1st fundamental, \( h \) and are completely independent of all the other fundamentals, \( \Gamma_s \). This follows from the limit of eq. (2.12) at infinity.

\[ t^\Lambda(\vec{x})|_{|\vec{x}|\to\infty} \to h^\Lambda + i \frac{\partial I_1(h)}{\partial h_\Lambda} \]

(2.22)

- **An attractor at each center:** this is a familiar feature from BPS 1-center black holes where the values of the moduli near the black hole horizon are independent on the values of moduli at infinity. In our case, with many centers, the values of special coordinates at each center are independent of all fundamentals except the one at the given center. The limit of eq. (2.12) near each center is given by

\[ t^\Lambda(\vec{x})|_{|\vec{x}|\to|x_s|} \to \Gamma^\Lambda_s + i \frac{\partial I_1(\Gamma_s)}{\partial \Gamma^\Lambda_s} \]

(2.23)

Apart from the nice clarity of the attractor behavior above, at each center and at infinity, writing our solution in terms of \( n + 1 \) fundamentals of our duality group yields two more simple features not yet discussed in the literature:

- The constant term in our integrability condition, (2.14), at each center is simply written in terms of the symplectic pairing between \( h \) and \( \Gamma_s \) without any reference to the phase of the central charge such as in [14, 12] (this simplification is also observed in [16]). This makes the construction, in principle, easier to extend to a non-BPS form.

- The mass of the black hole in terms of the charges and asymptotic moduli can be written very compactly in terms of \( h \) and \( \Gamma \) by expanding \( I_1(\vec{x}) \) to leading order in \( 1/|\vec{x}| \).
In what follows we will find that for most of the magic supergravities cases it is most convenient to generalize the form of the black hole composite solutions above: magic symmetries are best manifested using a double index, i.e matrix, notation instead of a single index Λ in coordinates and charges. We write real quantities, such as charges, as constrained complex matrices with a reality condition of type, $t^* = Ω t Ω^T$, and holomorphic quantities using the associated notion of complex conjugation.

### 3 Black hole entropies of magic $N = 2$ and of $N = 8$ supergravities

In order to find black hole composite solutions in four dimensions we require an explicit entropy formula in terms of generic black hole charges. Such formulas are known for symmetric spaces; the appropriate $N = 2$ supergravity coupled to vector multiplets are classified in [17]-[19]. These models include the reducible spaces $\left[ SU(1,1)/U(1) \right]^3$, $SU(1,1)/U(1) \times SO(P+2)/SO(2)$, and the complex projective space $SU(n)/SO(n)\times U(1)$. In Table 2 of [19] the first two cases are $L(0,0)$ and $L(0,P)$ respectively. For these three classes of $N = 2 d = 4$ supergravities the entropy formula is known for generic set of charges since the attractor equations [5]-[8] have been solved for these three spaces in [20]-[22], respectively. The remaining symmetric spaces for which the entropy formulas have not been established so far are connected to Jordan algebras, according to [17]. These are the magic supergravities.

We will present here the new $d = 4$ entropy formulas for magic $N = 2$ supergravities using the relation between the $d = 4$, $d = 5$ and $d = 6$ dualities [1]. We will also give some new forms of the entropy formula for $N = 8$ supergravity which can be easily compared with the entropy of the octonionic $N = 2$.

#### 3.1 Universal $d = 4 \iff d = 5$ relation

Consider the splitting

$$G_4 \rightarrow G_5 \times SO(1,1)$$

(3.1)

where $G_5$ is the duality group in $d = 5$ and $G_4$ is the duality group in $d = 4$. The quartic invariant of the relevant duality groups $G_4$ can be constructed using the cubic invariant of $G_5$. In all cases we have to split the total set of electric and magnetic charges into the zero component and the rest,

$$(p, q) = (p^0, p^I; q_0, q_I) , \quad I = 1, \ldots, n_v$$

(3.2)

This splitting is of the type

$$R_{p,q} = R_p + 1_p + R_q' + 1_q'$$

(3.3)

1Recently the $d = 4 \iff d = 5$ relation between particular black hole solutions has been used in counting of the BPS black hole degeneracy [23].
where $\mathbf{R}_p$ is the five-dimensional representation and $\mathbf{R}'_q$ is its contravariant representation:

\begin{align}
\mathbf{R}_p & \rightarrow (p^I) & \mathbf{R}'_q & \rightarrow (q_I) \\
\mathbf{1}_p & \rightarrow (p^0) & \mathbf{1}'_q & \rightarrow (q_0)
\end{align}

(3.4)

All of these are real representations of $G_4$ or $G_5$ as appropriate. In the magic models $\mathbf{R}_{p,q}$ refers to the following representations:

\begin{align}
J^O_3 & \rightarrow 56 \quad \text{fundamental of } E_{7(-25)} \\
J^Q_3 & \rightarrow 32 \quad \text{chiral spinor of } SO^*(12) \\
J^C_3 & \rightarrow 20 \quad \text{threefold antisymmetric (selfdual) of } SU(3,3) \\
J^R_3 & \rightarrow 14' \quad \text{threefold antisymmetric traceless of } Sp(6,\mathbb{R})
\end{align}

(3.6)

The general formula for the quartic invariant of the four-dimensional duality groups $G_4$ related to its $d=5$ cubic invariants is:

\[
J_4(p^0, q_0, p^I, q_I) = -(p \cdot q)^2 + 4 \left(q_0 I_3(p) - p^0 I_3(q) + \{I_3(q), I_3(p)\}\right)
\]

(3.7)

where the cubic invariants of the five-dimensional duality groups $G_5$ are given by

\[
I_3(p) = \frac{1}{3!} d_{IJK} p^Ip^jp^K, \quad I_3(q) = \frac{1}{3!} d^{IJK} q_Iq_Jq_K
\]

(3.8)

and the scalar product of charges and the Poisson bracket of cubic invariants are defined as follows

\[
p \cdot q \equiv p^0 q_0 + p^I q_I, \quad \{I_3(q), I_3(p)\} \equiv \frac{\partial I_3(q)}{\partial q_I} \frac{\partial I_3(p)}{\partial p^I}
\]

(3.9)

This structure of quartic invariants in terms of the cubic ones is valid for all cubic systems, where both $d_{IJK}$ as well as $d^{IJK}$ are known, in particular, it can be used for $N = 8$ and for all the magic $N = 2$ theories. Each model corresponds to specific values for $d_{IJK}$ and $d^{IJK}$ in equation (3.8) which form the cubic invariants of $G_5$. The quartic formula above can be derived following [3, 24] using the Freudenthal triple system that applies for any decomposition of the type given in eq. (3.1). The properties of the coset spaces which one encounters in magical supergravities can be inferred from [25, 26].

For $N = 8$ the classic example is $E_{7(7)} \rightarrow E_{6(6)} \times SO(1,1)$. The quartic invariant of the $E_{7(7)}$ and the cubic invariant of the $E_{6(6)}$ have been in the literature a long time: they were used for the single black hole entropy in $N = 8, d = 4$, [27, 28] and $N = 8, d = 5$, [7], respectively. In this section, we will exploit the relationship between $N = 8$ and the octonionic magic $N = 2$ to efficiently present invariants for all the magic supergravities. We can use these, in turn, to write explicit black hole composite solutions.

Let us quickly review the magic $N = 2$ supergravities. These are labeled using the four division algebras $\mathbb{A}$: $\text{dim}\mathbb{A}$ is 1 for real numbers $\mathbb{R}$, 2 for complex numbers $\mathbb{C}$, 4 for quaternions $\mathbb{H}$, and 8 for octonions $\mathbb{O}$. $N = 8$ supergravity corresponds to the split octonions, which do not form a division algebra.
The scalars in $N = 8$ supergravity are in the cosets $\frac{G_5}{H_5} = \frac{E_6(6)}{USp(8)}$ and $\frac{G_4}{H_4} = \frac{E_7(7)}{SU(8)}$ in $d = 5$ and $d = 4$, respectively. In each case, the vector fields are in the real representations 27 of $E_6(6)$ and 56 of $E_7(7)$.

For magic supergravities the cosets in $d = 5$ are, in the order of increasing dimension of the division algebra $\dim A$:

$$d = 5 : \quad \frac{G_5}{H_5} \Rightarrow \frac{SL(3; \mathbb{R})}{SO(3)} , \quad \frac{SL(3; \mathbb{C})}{SU(3)} , \quad \frac{SU^*(6)}{USp(6)} , \quad \frac{E_6(-26)}{F_4} \quad (3.10)$$

and in four-dimensional theories

$$d = 4 : \quad \frac{G_4}{H_4} \Rightarrow \frac{Sp(6; \mathbb{R})}{SU(3) \times U(1)} , \quad \frac{SU(3, 3)}{SU(3) \times SU(3) \times U(1)} , \quad \frac{SO^*(12)}{SU(6) \times U(1)} , \quad \frac{E_7(-25)}{E_6 \times U(1)} \quad (3.11)$$

The number of vector fields in the five-dimensional versions of magic supergravities is

$$d = 5 , \quad n_v = 3(\dim A + 1) \Rightarrow \quad n_v = 6, 9, 15, 27 \quad \text{for } \mathbb{R}, \mathbb{C}, \mathbb{H}, \mathbb{O} , \quad (3.12)$$

with one of the vectors coming from the gravity multiplet and transforming as a singlet of $H_5$.

The “magic” comes from representing the charges for all these vector fields as real parameters of a self-adjoint 3x3 matrix over the appropriate algebra (elements of the Jordan algebra $J^3_A$). In $d = 4$, Kaluza-Klein reduction adds one more vector field, and so the numbers are:

$$d = 4 , \quad \tilde{n}_v = 3(\dim A + 1) + 1 \Rightarrow \quad \tilde{n}_v = 7, 10, 16, 28 \quad \text{for } \mathbb{R}, \mathbb{C}, \mathbb{H}, \mathbb{O} , \quad (3.13)$$

with one of the vectors coming from the gravity multiplet and transforming as a singlet of $H_4$.

We will demonstrate two complementary methods for establishing the explicit form of the quartic/cubic invariants and related $d=4,5$ black hole entropies.

1. Starting in five-dimensions, we use the manifest five-dimensional duality group $G_5$ for all of our models, with $E_{6(6)}, E_{6(-26)}, SU^*(6), SL(3; \mathbb{C}), SL(3; \mathbb{R})$-symmetry, and the simple features of the corresponding Jordan algebras to write down the appropriate cubic invariants. Operationally, this entails representing the real coordinates of the N=8 theory and the very special real geometry of the $N = 2$ theories in terms of complex anti-symmetric matrices obeying a reality condition. We then obtain the holomorphic coordinates of the corresponding four dimensional scalar manifold geometry by relaxing the reality condition. To the reader, this may seem different from the well known case where one starts with real matrices in $d = 5$ complexify them term by term to get the holomorphic coordinates in $d = 4$. In fact, the difference is cosmetic rather than structural, due to the convenience of using a double index notation; we will make this clear in an explicit mathematical way.

2. We rearrange our charges and coordinates using just $G_6$ manifestly-symmetric representations. In this case we have to further split the five-dimensional duality symmetry down to the six-dimensional one:

$$G_5 \quad \Rightarrow \quad G_6 \times SO(1, 1) \quad (3.14)$$
The duality group, \( G_6 \), for magic supergravities is \( SO(1, \dim \mathcal{A} + 1) \), similar to the duality group \( SO(5, 5) \) for the \( N = 8 \) model. For a IIB oxidation, tensor charges (self-dual strings) will appear in vector representations of \( G_6 \) and vector charges (point charges and their magnetic duals) will appear as real spinors. The advantage to this method is that, in all but one case, we can reconstruct the cubic invariant in \( d=5 \) using just the metric and gamma matrices of \( G_6 \). For the quaternionic magic the eight dimensional real representation of \( SO(1, 5) \) is really the combination of 2 pseudo-real spinors of the same chirality; the answer in this case is only slightly more complicated. In the first cases \( (\dim \mathcal{A} \neq 4) \), the \( d=4 \) holomorphic coordinates are a simple complexification of the \( d=6 \) real coordinates. The details for the more subtle quaternionic case are presented in the Appendix B.

While overall not as manifestly symmetric, the six-dimensional approach provides us with more than a cross-check for our first method. The choice of splitting for \( G_5 \to G_6 \times SO(1, 1) \) also yields some valuable insight into the oxidation process, e.g. the split into tensor and vector multiplets above.

### 3.2 \( G_5 \) manifestly-symmetric entropy of \( d = 4 \) BPS black holes

Our goal at this point is to develop a simple expression for the cubic invariants of the five-dimensional magic supergravity; using these we can then extract a \( G_5 \) manifestly-symmetric formula for the quartic invariant in four dimensions using the formula in (3.7). A very efficient way to accomplish our task is to use the relationship of these theories to Jordan algebras and the N=8 maximally supersymmetric theory.

Let us demonstrate with the familiar case of \( N = 8 \) supergravity

\[
E_{7(7)} \to E_{6(6)} \times SO(1, 1)
\]

and the octonionic magic \( N = 2 \) supergravity with

\[
E_{7(-25)} \to E_{6(-26)} \times SO(1, 1).
\]

The respective five-dimensional charge spaces for these theories correspond to the Jordan algebras \( J_3^{\mathbb{O}} \) and \( J_3^{\mathbb{Q}} \). We can represent the 27 elements for both algebras conveniently using the (faithful) traceless anti-symmetric representations of \( Usp(8) \) and \( Usp(6, 2) \). This is similar in spirit to the use of \( SU(8) \) and \( SU(6, 2) \) to construct quartic invariants in [32], but our results for the cubic invariants of \( E_{6(6)} \) and \( E_{6(-26)} \) do not descend straightforwardly from these quartic invariants. It would be interesting to have a better understanding of the connection between the two.

We will now quickly review the details necessary to build these representations. We start by defining \([25, 26]\) the \((8 \times 8)\) symplectic matrix \( \Omega \) and a metric preserved by \( Usp(8) \) and \( Usp(6, 2) \)

\[
\Omega = I_4 \otimes \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix}, \quad K^{(a, b)} = \alpha \otimes \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}
\]
which will allow us to treat both cases, \(E_{6(6)}\) in \(Usp(8)\) basis and \(E_{6(-26)}\) in \(Usp(6, 2)\) basis, simultaneously. The \((4 \times 4)\) matrix \(\alpha\) is

\[
\alpha = \begin{pmatrix}
0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 \\
0 & 0 & 0 & 1 \\
0 & 0 & 1 & 0 \\
\end{pmatrix}
\]

for \(Usp(8)\), \(\alpha = \begin{pmatrix}
-1 & 0 & 0 & 0 \\
0 & -1 & 0 & 0 \\
0 & 0 & -1 & 0 \\
0 & 0 & 0 & 1 \\
\end{pmatrix}
\)

for \(Usp(6, 2)\), \(\alpha = \alpha^t\), \(\alpha^2 = 1\). We will label the two different metrics either \(K^{(0,4)}\) or \(K^{(3,1)}\).

We are now ready to exhibit the real representation 27 of \(E_{6(6)}, E_{6(-26)}\) taken in the \(Usp(8), Usp(6, 2)\) basis by using the map \(t^I \rightarrow t^{ab}\) where \(t^{ab}\) is the antisymmetric traceless matrix, \(t^{ab} = -t^{ba}\), \(Tr t = t^{ab} \Omega_{ab} = 0\) satisfying a reality condition

\[
t^*_a = \tilde{\Omega}_{aa'} \tilde{\Omega}_{b'b} t^{a'b'} \quad \Leftrightarrow \quad t^* = \tilde{\Omega} t \tilde{\Omega}^T,
\]

where

\[
\tilde{\Omega} = \begin{pmatrix} 0 & -\alpha \\ \alpha & 0 \end{pmatrix} = K \Omega = \begin{pmatrix} \alpha & 0 \\ 0 & \alpha \end{pmatrix} \begin{pmatrix} 0 & -I \\ I & 0 \end{pmatrix}, \quad \Omega = \begin{pmatrix} 0 & -I \\ I & 0 \end{pmatrix}
\]

With our framework, we quickly write down our cubic invariant as:

\[
D_{IJK} t^I t^J t^K = Tr(\Omega t \tilde{\Omega}^T)
\]

which is a real polynomial. For the contragradient representation we can take

\[
D_{IJK} t'_I t'_J t'_K = Tr(\Omega' t' \Omega' t' \Omega' t'), \quad t' = t^{ab}_a, \quad \Omega' = \tilde{\Omega}^T.
\]

Using these representations, the cubic invariants of \(E_{6(6)}\) and \(E_{6(-26)}\) in \(Usp(8)\) and \(Usp(6, 2)\) basis, respectively, are given by

\[
I_3(p) = \frac{1}{3!} Tr(\Omega p \Omega p \Omega p), \quad I_3(q) = \frac{1}{3!} Tr(\Omega' q \Omega' q \Omega' q)
\]

Here

\[
p^* = \tilde{\Omega} p \tilde{\Omega}^T, \quad Tr(p \Omega) \equiv p^{ab} \Omega_{ab} = 0, \quad q^* = \tilde{\Omega} q \tilde{\Omega}^T, \quad Tr(q \Omega) = 0.
\]

The reality and trace condition give exactly 27 real entries for the complex matrices \(p^{ab}\) and \(q_{ab}\), the complex two index notation is simply a convenient way to represent the real numbers \(p^I\) and \(q_I\). The only difference between the the cubic invariants of \(E_{6(6)}\) and \(E_{6(-26)}\) comes from the reality condition, driven by a different choice of \(\Omega\) coming from different versions of the metric \(K\). The the fact the cubic invariants are not explicitly different in terms of complex matrices, but only different through their respective reality constraint is a direct consequence of the fact that \(E_{6(6)}\) and \(E_{6(-26)}\) are different real forms of the same complex \(E_6\) algebra.

The reality constraint \(\text{(3.19)}\) and the tracelessness condition \(t^{ab} \Omega_{ab} = 0\) can be solved and one finds that

\[
\text{Re} t = \begin{pmatrix} A_1 \\ -\alpha S \alpha \end{pmatrix}, \quad \text{Im} t = \begin{pmatrix} A_2 \\ \alpha A_1 \alpha \end{pmatrix}
\]

\[A_1 \begin{pmatrix} \alpha A_1 \alpha \\ -\alpha A_2 \alpha \end{pmatrix}
\]

12
where \( S' = \alpha S \alpha, \) \( Tr S = 0 \) and \( A' = -\alpha A \alpha. \) This means that \( \text{Re} \ t \) depends on \( n^2 - 1 = 15 \) real entries since \( A_1 \) and \( A_2 \) and \( A_0 \) are antisymmetric, and \( S \alpha \) is symmetric. \( \text{Im} \ t \) depends on \( n^2 - n = 12 \) real entries. Together \( t \) has 27 real entries. The same forms apply for the electric and magnetic charges, in agreement with \( n_v = 3(\text{dim} A + 1) = 3(8 + 1) \) for octonions.

If we want to use only real entries for magnetic and electric charges, where \( p = \text{Re} \ p + i \text{Im} \ p = p_1 + ip_2, \) we can use

\[
p^* = \tilde{\Omega} p \tilde{\Omega}^T \Rightarrow p_1 = \tilde{\Omega} p_1 \tilde{\Omega}^T, \quad p_2 = -\tilde{\Omega} p_2 \tilde{\Omega}^T
\]

(3.26)

The cubic invariants in terms of the real entries are

\[
I_3(p) = \frac{1}{3!} Tr(\Omega p \Omega p \Omega p) = \frac{1}{3!} Tr(\Omega p_1 \Omega p_1 \Omega p_1) - \frac{1}{2} Tr(\Omega p_1 \Omega p_2 \Omega p_2)
\]

(3.27)

The electric charges are in the contragradient representation, and so we expand \( q \) as \( q = \text{Re} q - i(\text{Im} q) = q_1 - iq_2, \) with the analogous equations to eqs. (3.26) and (3.27).

**Three other \( G_5 \)-invariant magic entropies**

Consider the quaternionic magic case with \( G_5 = SU^*(6), H_5 = USp(6). \) Our charges are in the \( 15, \) the two-fold anti-symmetric representation \( SU^*(6), \) which decomposes as the \( 14 \) traceless anti-symmetric of \( USp(6) \) plus a singlet. This anti-symmetric representation of \( USp(6) \) is not the truncation to the common \( 6 \times 6 \) matrix of the two types of \( 8 \times 8 \) matrices above (as opposed to what happens in [33]).

We write the quaternionic magic charges using the matrix \( U^{\Lambda \Sigma}, U = -U^t \) with the reality condition \( U^* = \Omega U \Omega^T, \) giving us \( n_v = 3(\text{dim} A + 1) = 3(4 + 1) = \frac{65}{2}, \) so we that we end up with 15 electric and 15 magnetic charges. The real cubic forms are:

\[
\frac{1}{3!} t_I t_J t_K = \frac{1}{3!} 2^3 \epsilon_{\Lambda \Sigma \Delta \Gamma \Pi \Omega} U^{\Lambda \Sigma} U^{\Delta \Gamma} U^{\Pi \Omega} \equiv \text{Pf} U
\]

(3.28)

\[
I_3(p) = \frac{1}{3!} 2^3 \epsilon_{\Lambda \Sigma \Delta \Gamma \Pi \Omega} p^{\Lambda \Sigma} p^{\Delta \Gamma} p^{\Pi \Omega} \equiv +\text{Pf} p
\]

(3.29)

\[
I_3(q) = \frac{1}{3!} 2^3 \epsilon_{\Lambda \Sigma \Delta \Gamma \Pi \Omega} q^{\Lambda \Sigma} q^{\Delta \Gamma} q^{\Pi \Omega} \equiv -\text{Pf} q
\]

(3.30)

In terms of real charges \( p_1 = \Omega p_1 \Omega^T \) and \( p_2 = -\Omega p_2 \Omega^T \) where \( p^* = \Omega p \Omega^T \) and \( p = p_1 + ip_2 \) we have

\[
I_3(p_1, p_2) = \frac{1}{3!} 2^3 \epsilon_{\Lambda \Sigma \Delta \Gamma \Pi \Omega} p_1^{\Lambda \Sigma} (p_1^{\Delta \Gamma} p_1^{\Pi \Omega} - 3p_2^{\Delta \Gamma} p_2^{\Pi \Omega})
\]

(3.31)

and analogous for \( q = q_1 - iq_2. \) When we use these expressions (cf. section 5) to obtain a quartic invariant, we get an expression which matches exactly with the quartic invariant for quaternionic magic in [33]. This quartic invariant also appears explicitly in [35] within the context of \( N = 6 \) supergravity.
The next case is complex magic with \( G_5/H_5 = SL(3,C)/SU(3) \) and \( n_v = 3(2+1) = 9 \). We truncate \( U^\Lambda^\Sigma \) by writing it in terms of a symmetric 3x3 matrix \( U^{II} \):

\[
U^\Lambda^\Sigma \rightarrow U^{II}, \quad U_{\Lambda\Sigma} \rightarrow U_{II}.
\] (3.32)

The reality condition makes \( U_{II} \) hermitian (the trace is an invariant of \( H_5 \)). The cubic invariants in terms of real charges, where \( p^{I} = \text{Re} p^{II} + i \text{Im} p^{I} = p_1^{II} + ip_2^{II} \) and \( q_{II} = \text{Re} q_{II} - i(\text{Im} q_{II}) = q_1^{II} - iq_2^{II} \) are:

\[
I_3(p) = \frac{1}{3!} \epsilon_{IJK} \epsilon_{IJK} p^{I} p^{J} p^{K} = \det p \tag{3.33}
\]

\[
= \frac{1}{3!} \epsilon_{IJK} \epsilon_{IJK} p_1^{II} (p_1^{J} p_1^{K} - 3p_2^{JJ} p_2^{KK})
\]

\[
I_3(q) = \frac{1}{3!} \epsilon^{IJK} \epsilon^{IJK} q_{II} q_{JJ} q_{KK} = \det q \tag{3.34}
\]

\[
= \frac{1}{3!} \epsilon^{IJK} \epsilon^{IJK} q_1^{II} (q_1^{JJ} q_1^{KK} - 3q_2^{JJ} q_2^{KK}).
\]

The last case, real magic has \( t^I \) in 6 of \( SL(3,\mathbb{R}) \), and \( n_v = 3(1+1) = 6 \).

\[
\frac{1}{3!} \epsilon^{IJK} t^{I} t^{J} t^{K} = \frac{1}{3!} \epsilon_{LMN} \epsilon_{PQR} t^{LP} t^{MQ} t^{NR} = \det t.
\] (3.35)

Here the symmetric matrix \( t^{LP} = t^{PL} \) is real. The cubic invariants in terms of real charges are

\[
I_3(p) = \frac{1}{3!} \epsilon_{LMN} \epsilon_{PQR} p^{LP} p^{MQ} p^{NR}, \quad I_3(q) = \frac{1}{3!} \epsilon_{LMN} \epsilon_{PQR} q_{LP} q_{MQ} q_{NR}.
\] (3.36)

### 3.3 Magic prepotentials from \( d = 5 \) and the attractor equations

To connect the \( d = 5 \) construction with real coordinates \( Y^I \) to four-dimensional holomorphic coordinates \( X^I \) we have to give a prescription for complexifying the real coordinates. We can then present the holomorphic prepotentials \( F = \frac{I_3(X_1, X_2)}{X^5} \) of \( d = 4 \) \( N = 2 \) supergravities for all of our models. For the magic supergravities, the prescription is sweet and simple: we just relax the reality condition on our defining matrices, keeping it only when we need to define a complex conjugate quantity. In the quaternionic case, for example, we take our reality-constrained matrices \( Y \) and split them once again into real matrices \( Y = Y_1 + iY_2 \) satisfying \( Y_1 = \Omega Y_1 \Omega^T, Y_2 = -\Omega Y_2 \Omega^T \). Regular complexification now requires us to turn the real matrices \( Y_1, Y_2 \) into complex matrices \( X_1, X_2 \) satisfying the linear constraint \( X_1 = \Omega X_1 \Omega^T, X_2 = -\Omega X_2 \Omega^T \). If we write \( X = X_1 + iX_2 \) we now have a new complex matrix independent of any linear constraints, the new conjugate variable is \( \bar{X} = X_1^* + iX_2^* = \Omega(X^*)\Omega^T \).

To summarize, in general we apply the standard complexification to our real coordinates \( Y_1 \) and \( Y_2 \) to get a holomorphic coordinate \( X = X_1 + iX_2 \). In the bi-index notation, what remains of the reality constraint on \( Y \) is a different definition for complex conjugation on our vector space:

\[
\implies Y^\Lambda^\Sigma \rightarrow X^\Lambda^\Sigma, \quad \bar{X}^\Lambda^\Sigma = \Omega(X^\Lambda^\Sigma)^*\Omega^T.
\] (3.37)
Thus the “bar” operation is defined to be not just complex conjugation but also a matrix multiplication as shown above.

The prepotentials for all magic $N = 2$ supergravities the their cubic invariant in five dimensions. These prepotentials may be used in future for the analysis of the Legendre transform and relation to the generalized Hitchin functional, as in [29],[24].

\[ J^0 \Rightarrow F(X) = \frac{1}{3! X^0} \text{Tr}(\Omega(X_1 + iX_2)\Omega(X_1 + iX_2)\Omega(X_1 + iX_2)) = \]
\[ \frac{1}{3! X^0} [\text{Tr}(\Omega X_1 \Omega X_1 \Omega X_1) - 3\text{Tr}(\Omega X_1 \Omega X_2 \Omega X_2)] \] (3.38)

Here $X_1$ and $X_2$ are the complexified fields corresponding to the real and imaginary parts of the antisymmetric traceless matrix $Y^{ab}$ with $Y^* = -\bar{\Omega} Y \bar{\Omega}$, $a, b = 1, \ldots, 8$.

\[ J^Q \Rightarrow F(X) = \frac{1}{X^0} \text{Pf}(X_1 + iX_2) = \]
\[ \frac{1}{3!} \cdot 2^3 X^{0} \varepsilon_{\Lambda\Sigma\Delta\Gamma\Pi\Omega} X^{\Lambda\Sigma}_1 (X^{\Delta\Gamma}_1 X^{\Pi\Omega}_1 - 3X^{\Delta\Gamma}_2 X^{\Pi\Omega}_2) \] (3.40)

Here $X_1$ and $X_2$ are antisymmetric and $X_1 = \Omega X_1 \Omega^T$, $X_2 = -\Omega X_2 \Omega^T$, with $\Lambda, \Sigma = 1, 6$.

\[ J^C \Rightarrow F(X) = \frac{1}{X^0} \det(X_1 + iX_2) = \]
\[ \frac{1}{3! X^0} \varepsilon_{IJK} \varepsilon_{IJK} X^{IJK}_1 (X^{IJK}_1 - 3X^{IJK}_2) \] (3.41)

Here $X_1$ is symmetric and $X_2$ is antisymmetric. We can put them together in a single 3x3 matrix, $X$ in the $(3, \bar{3})$ representation of $H_4 = SU(3) \times U(3))$.

\[ J^R \Rightarrow F(X) = \frac{1}{X^0} \det X = \frac{1}{3! X^0} \varepsilon_{LMN} \varepsilon_{PQR} X^{LP} X^{MQ} X^{NR} \] (3.42)

Here $X^{LM}$ is in the symmetric representation of $H_4 = U(3)$, $L, M = 1, 2, 3$.

In each case, the $X^0$ field is the usual extra projective coordinate which is a singlet of $H_5$. We note the parallel with the five dimensional case. In that case, the real cubic invariant was determined by a cubic invariant of the complex form of $G_5$ plus a reality condition. Here the holomorphic pre-potential is determined by a cubic invariant of $H_4$ supplemented by a different notion of complex conjugation.

For completeness we add here the prepotentials for other $N = 2$ cosets for which the entropy formula is known. For the $L(0, P)$ models the holomorphic prepotential is

\[ F(X) = \frac{X}{2X^0} [Y^r Y^s \eta_{rs}] \quad \eta_{rs} = (1, -1, \ldots, -1) , \quad r, s = 1, \ldots, P + 2. \] (3.43)

The corresponding coset space in $D = 4$ is $SU(1,1)/U(1) \times SO(2, P+2)/SO(1, P+2)$ and in $d = 5$ we have $SO(1, 1) \times SO(1, P+1)/SO(P+1)$ and the entropy was found in [21].
The case of the complex projective space \( SU(1,n) \) is described by a quadratic prepotential

\[
F(X) = \frac{-i}{4} [(X^0)^2 - \sum_{i=1}^{i=n} (X^i)^2],
\]

(3.44)

the entropy was found in [22].

The BPS attractor equations for all cases can be nicely written by once more embedding our charges into constrained complex matrices (here we use quaternionic magic to illustrate):

\[
p^{\Lambda\Sigma} + 2i \frac{\partial \sqrt{J_4(p,q)}}{\partial q_{\Lambda\Sigma}} = 2i \bar{Z} L^{\Lambda\Sigma}
\]

(3.45)

In terms of strictly real charges \( p_1, p_2 \) and \( q_1, q_2 \) these equations can be rewritten as:

\[
p_1 + i \frac{\partial \sqrt{J_4(p,q)}}{\partial q_1} = i \bar{Z} [L + \Omega L \Omega^T] = 2i \bar{Z} L_1
\]

(3.46)

\[
p_2 + i \frac{\partial \sqrt{J_4(p,q)}}{\partial q_2} = i \bar{Z} [L - \Omega L \Omega^T] = 2i \bar{Z} i L_2
\]

(3.47)

Solving these attractor equations yields (in terms of either the constrained complex form or split real form of the charges):

\[
t^{\Lambda\Sigma} = \frac{X^{\Lambda\Sigma}}{X^0} \rightarrow \frac{p^{\Lambda\Sigma} + 2i \frac{\partial \sqrt{J_4(p,q)}}{\partial q_{\Lambda\Sigma}}}{p^0 + i \frac{\partial \sqrt{J_4(p,q)}}{\partial q_0}}, \quad \bar{t}^{\Lambda\Sigma} = \frac{X^{\Lambda\Sigma}}{X^0} \rightarrow \frac{p^{\Lambda\Sigma} - 2i \frac{\partial \sqrt{J_4(p,q)}}{\partial q_{\Lambda\Sigma}}}{p^0 - i \frac{\partial \sqrt{J_4(p,q)}}{\partial q_0}}
\]

(3.48)

\[
t_1^{\Lambda\Sigma} = \frac{X_1^{\Lambda\Sigma}}{X^0} \rightarrow \frac{p_1^{\Lambda\Sigma} + i \frac{\partial \sqrt{J_4(p,q)}}{\partial q_{1\Lambda\Sigma}}}{p^0 + i \frac{\partial \sqrt{J_4(p,q)}}{\partial q_0}}, \quad t_2^{\Lambda\Sigma} = \frac{X_2^{\Lambda\Sigma}}{X^0} \rightarrow \frac{p_2^{\Lambda\Sigma} - i \frac{\partial \sqrt{J_4(p,q)}}{\partial q_{2\Lambda\Sigma}}}{p^0 + i \frac{\partial \sqrt{J_4(p,q)}}{\partial q_0}}
\]

(3.49)

where we also exhibit the matching anti-holomorphic equation to re-emphasize the different complex conjugation. These equations are a slight generalization of the one in eq. (2.12) which allow us accommodate a more algebraic notion of “reality” acting on matrices and will allow us to present the black hole composite solutions for \( N = 8 \) and magic \( N = 2 \) supergravities.

Finally, we can give an explicit expression for the Kähler potential of all the magic supergravities in terms of a modified notion of “reality” as follows. In standard holomorphic coordinates, the Kähler potential depends on the imaginary section. We then get the usual looking Kähler potential:

\[
e^{-K} = \frac{i}{3!} |X^0|^2 \text{Tr} [\Omega (t - \bar{t}) \Omega (t - \bar{t}) \Omega (t - \bar{t})].
\]

(3.50)

One has to keep in mind here that we use a definition of complex conjugation where our anti-holomorphic special coordinates are related to the holomorphic as follows

\[
t = \Omega^T t^* \Omega
\]

(3.51)
This expression for the K"ahler potential again shows that it is not invariant under K"ahler transformations. One can use the gauge $X^0 = 1$, but in this gauge one doesn’t necessarily have $2U = K$. Moreover, it is known from the example of the STU attractor that in this gauge $e^{-2U}$ is not equal to $e^K$, see [20]. On the other hand the freedom in the choice of the gauge for $X^0$ may be used to identify $K$ with $2U$ as in [14].

3.4 $G_6$ manifestly-symmetric entropy of $d = 4$ BPS black holes

The explicit expressions for each case can be established using the very special real geometry constructions of [18, 19]. The corresponding homogeneous special real spaces in five dimensions (very special geometry) are

$$\frac{G_5}{H_5} = L(P, 1), \quad P = \text{dim}A = 1, 2, 4, 8.$$  \hspace{1cm} (3.52)

in the Table 2 of [19]. Their corresponding K"ahler spaces with special geometry are our magic cosets $G_{P+1}$ in $d = 4$. We uplift our models to $d = 6$ (2, 0) supergravities, following [30], [31] (see also [34]). We introduce in $d = 6$ a number of tensor multiplets, $N_t = \text{dim}A + 1$, and a number of vector multiplets $N_v = 2\text{dim}A$ and we split the symmetric tensor $d_{IJK}$ as follows:

$$d_{IJK} \Rightarrow (d_{zrs}, d_{r\alpha\beta}) , \quad d^{IJK} \Rightarrow (d^{zrs}, d^{r\alpha\beta})$$  \hspace{1cm} (3.53)

where $r = 0, \ldots, \text{dim}A + 1$, $z = 1$, $\alpha, \beta = 1, \ldots, 2\text{dim}A$. This corresponds to the split of the real five-dimensional coordinates into $(z, b^r, a^\alpha)$ so that $z$ is the KK singlet scalar, $b^r$ is the coordinate of the coset space $SO(1, \text{dim}A + 1)/SO(\text{dim}A + 1)$ and $a^\alpha$ are spinors of $SO(1, \text{dim}A + 1)$. We have an analogous split for the electric and magnetic charges.

$$p^I = (p^z, p^r, p^\alpha) , \quad q_I = (q_z, q_r, q_\alpha) , \quad r = 0, \ldots, \text{dim}A + 1, \quad \alpha = 1, \ldots, 2\text{dim}A$$  \hspace{1cm} (3.54)

For the $L(0, P)$ models one should take $r, s = 0, \ldots, P + 1$ and $p^\alpha = q_\alpha = 0$, following Table 2 of [19]. The cubic invariants are

$$I_3(p) = \frac{1}{2} \left( p^z \eta_{rs} p^r p^s + (\gamma_r)_{\alpha\beta} p^r p^\alpha p^\beta \right)$$  \hspace{1cm} (3.55)

$$I_3(q) = \frac{1}{2} \left( q_z q^r q_s + (\gamma^r)_{\alpha\beta} q_r q_\alpha q_\beta \right)$$  \hspace{1cm} (3.56)

Here $\eta_{rs}$ is the Lorentzian metric of $SO(1, \text{dim}A + 1)$ and $(\gamma_r)_{\alpha\beta}$ are the $\gamma$-matrices of the groups $SO(1, \text{dim}A + 1)$. The details magic quaternionic supergravity are given in Appendix A. For the groups $SO(1, 2), SO(1, 3)$ and $SO(1, 9)$ we can have real spinors and therefore the cubic invariants (3.55), (3.56) above applies to them.

The relation between the magic exceptional octonionic $N = 2$ and $N = 8$ supergravity has to do with the change from the real octonions with the norm invariant under $SO(8)$ to the split
octonions with the norm invariant under $SO(4.4)$. In $d = 6$ this means that

$$N = 2 \quad \text{octonionic magic, real octonions} \quad \Rightarrow \quad \frac{SO(1, 9)}{SO(9)}$$

$$N = 8 \quad \text{octonionic, split octonions} \quad \Rightarrow \quad \frac{SO(5, 5)}{SO(5) \times SO(5)}$$

To find the cubic/quartic invariants of $N = 8$ theory in the $G_6$ manifest basis all we have to do it to take the cubic invariants for octonionic magic $N = 2$ in eq. (3.55) and (3.56) with the associated metric $\eta_{rs}$ for $SO(5, 5)$ instead of that for $SO(1, 9)$. The $\gamma$-matrices in the second term of cubic invariants change accordingly so that the corresponding Clifford algebra has the $\eta_{rs}$ signature for $SO(5, 5)$ instead of $SO(1, 9)$.

The case of quaternionic magic supergravity with $SO(1, 5) \sim SU^*(4)$ duality group we have to treat separately. We present this case in Appendix A.

## 4 BPS composites of octonionic magic $N = 2$

**Manifest $E_{6(-26)}$**

The quartic invariant of $E_{7(7)}$ is a well known invariant related to the entropy of $N = 8$ black holes. We would like to give a similar explicit formula for the entropy of the octonionic magic $N = 2$ which is a quartic invariant of $E_{7(-25)}$. In this section we will extract this invariant using the procedure given in sec. 3, but in more detail. We will then use this expression to write down the general multi-black hole solution following the framework in sec. 2, with all the modifications necessary to use our the bi-index formalism which proved so useful in deriving the cubic and quartic invariants of octonionic magic.

Recapping the procedure in sec. 3, we introduce the charges with a $d = 4 \rightarrow d = 5$ split corresponding to the $E_{7(-25)} \rightarrow E_{6(-26)} \times SO(1, 1)$ split; in the $N = 8$ case this is an $E_{7(7)} \rightarrow E_{6(6)} \times SO(1, 1)$ split. The real 56 of $E_{7(-25)}$ which combines electric and magnetic charges is split into 2 real charges $p^0, q_0$ and 54 electric and magnetic charges in the $27'$ and 27 representations of $E_{6(-26)}$, written as representations of $USp(6, 2)$. This means we have a complex antisymmetric skew-traceless $p^{ab} = (p_1^{ab} + ip_2^{ab})$ and $q_{ab} = (q_{1ab} - iq_{2ab})$ satisfying a reality constraint $p^* = \tilde{\Omega} p \tilde{\Omega}^T$ and analogous for $q$. Here

$$\tilde{\Omega} = K^{(3,1)} \Omega$$

where the matrices $K^{(3,1)}$ and $\Omega$ are defined in eqs. (3.17), (3.18). Alternatively, one could use the real charges, $p_1^{ab}, p_2^{ab}$ and $q_1, q_2, q_{1ab}, q_{2ab}$, as is and lose explicit duality symmetry.

The quartic invariants of $E_{7(7)}$ and $E_{7(-25)}$ are both

$$J_4(p^0, q_0, p^{ab}, q_{ab}) = -(p \cdot q)^2 + 4 \left(q_0 I_3(p) - p^0 I_3(q) + \{I_3(q), I_3(p)\} \right).$$
with cubic invariants given by

\[ I_3(p) = \frac{1}{3!} Tr(\Omega p \Omega p) = \frac{1}{3!} Tr(\Omega p_1 \Omega p_1 \Omega p_1) - \frac{1}{2} Tr(\Omega p_1 \Omega p_2 \Omega p_2) \] (4.3)

\[ I_3(q) = \frac{1}{3!} Tr(\Omega^p q \Omega^p q \Omega^p q) = \frac{1}{3!} Tr(\Omega^p q_1 \Omega^p q_1 \Omega^p q_1) - \frac{1}{2} Tr(\Omega^p q_1 \Omega^p q_2 \Omega^p q_2) \] (4.4)

and the scalar product of charges and the Poisson bracket of cubic invariants are defined as follows

\[ p \cdot q \equiv p^0 q_0 + \frac{1}{2} p^{ab} q_{ab} = p^0 q_0 + \frac{1}{2} p_1^{ab} q_{1ab} + \frac{1}{2} p_2^{ab} q_{2ab} , \] (4.5)

\[ \{ I_3(q), I_3(p) \} \equiv 2 \frac{\partial I_3(q)}{\partial q_{ab}} \frac{\partial I_3(p)}{\partial p^{ab}} = \frac{1}{2} \frac{\partial I_3(q)}{\partial q_{1ab}} \frac{\partial I_3(p)}{\partial p_1^{ab}} + \frac{1}{2} \frac{\partial I_3(q)}{\partial q_{2ab}} \frac{\partial I_3(p)}{\partial p_2^{ab}} . \] (4.6)

All the formulas above with complex matrices are real due to the reality condition on the charge matrices. We note that any derivatives we write in the octonionic case need to take into account the fact that we differentiate with respect to skew-traceless matrices. This corresponds to restricting the tangent space using the tracelessness condition and operationally gives the following definition for the derivative:

\[ \frac{\partial p^{ab}}{\partial p^{cd}} = \delta^a_c \delta^b_d - \delta^a_d \delta^b_c - \frac{1}{4} \Omega_{ab} \Omega_{cd} . \] (4.7)

It would be interesting to understand exactly which change of variable would be necessary to match with our expression for the quartic invariant, \( J_4 \), with the one in [33] written explicitly in terms of four dimensional charges.

Overall, the only difference between the endpoint expressions for \( E_7(7) \) and \( E_7(-25) \) is exactly in the reality constraint on the charges, i.e. which parts of \( p \) get slotted in \( p_1 \) or \( p_2 \) by our choice of \( \tilde{\Omega} \):

\[ N = 8 : \ E_7(7) \to E_6(6) \times SO(1,1) \Rightarrow p = \pm \tilde{\Omega}^T p^* \tilde{\Omega} = K^{0,4} \Omega^T p^* \Omega K^{0,4} \] (4.8)

\[ N = 2 \text{ magic} : \ E_7(-25) \to E_6(-26) \times SO(1,1) \Rightarrow p = \Omega^T p^* \tilde{\Omega} = \pm K^{3,1} \Omega^T p^* \Omega K^{3,1} . \] (4.9)

Following sec. 2, to describe the BPS composites of octonionic magic \( N = 2 \) we need a set of \( 1 + n \) fundamental 56-dimensional representations of \( E_7(-25) \). Here \( n \) is the number of centers of our multicenter solution

\[ \mathbf{h} \equiv (h^0, h_0, h^{ab}, h_{ab}) , \quad \mathbf{\Gamma}_s \equiv (p^0, q_0, p^{ab}, q_{ab})_s \quad a, b = 1, \ldots, 8, \quad s = 1, \ldots, n . \] (4.10)

We introduce a 56-real-dimensional harmonic function in terms of two real harmonics and two constrained complex matrix harmonics:

\[ \mathbf{H}(\vec{x}) = (H^0, H_0, H^{ab}, H_{ab}) = \mathbf{h} + \sum_{s=1}^n \frac{\mathbf{\Gamma}_s}{|\vec{x} - \vec{x}_s|} . \] (4.11)
Here we have introduced \( n \) constant \( E_{7(-25)} \) fundamentals \( \Gamma_s \equiv (p, q)_s \) with \( s = 1, \ldots, n \) and
one constant \( E_{7(-25)} \) fundamental, \( h \equiv (p_\infty, q_\infty) \) which is the asymptotic value of the harmonic function. Note that once again “reality” for \( H \) in this notation means that \( (H^{ab}, H_{ab}) \) are
complex matrices, they satisfy the same reality condition as the corresponding charge matrices. With the information above, we are almost ready to write down the solution. The final ingredient we need is a definition of the symplectic invariant for our solution, which is:

\[
< \Gamma_s, \Gamma_t > = p_s \cdot q_t - p_t \cdot q_s.
\]  

(4.12)

We can now write the stationary metric for the BPS multicenter solution with \( J_4(\vec{x}) > 0 \):

\[
ds^2_4 = -J_4^{-1/2}(\vec{x}) (dt + \vec{\omega} \vec{d} \vec{x})^2 + J_4^{1/2}(\vec{x}) d\vec{x}^2
\]  

(4.13)

where

\[
J_4(\vec{x}) \equiv J_4 \circ H(\vec{x}) \quad \nabla \times \vec{\omega} = [H, \nabla H].
\]  

(4.14)

As before, we define \( J_4 \circ H(\vec{x}) \) by replacing the set of charges in eqs. (4.10) by the harmonic functions in eq. (4.11). The integrability condition for the solution is

\[
\langle H, \Delta H \rangle = 0, \quad \sum_{t=1}^{n} \langle \Gamma_s, \Gamma_t \rangle + \langle \Gamma_s, H \rangle = 0
\]  

(4.15)

For the asymptotically flat geometry one has to require that \( J_4(h) = 1 \). The values of the vector fields (written as constrained complex matrices again!), in spherical coordinates \( (r_s, \theta_s, \phi_s) \) around each center \( \vec{x}_s \) are:

\[
A^{ab} = \frac{\partial}{\partial H_{ab}} \left( \ln J_4(H) \right) (dt + \psi) - \sum_s \cos \theta_s d\phi_s \otimes \Gamma^s_{ab}
\]  

(4.16)

The 54 scalars of magic octonionic model represent the coset space \( E_{7(-25)} / \text{E}_6 \times SO(2) \). We give their solution in terms of the holomorphic coordinates with the special complex conjugation explained in subsection (3.3):

\[
X^{ab}_0(\vec{x}) = \frac{H^{ab} + 2i \frac{\partial J_4(H)}{\partial H_{ab}}}{H^0 + i \frac{\partial J_4(H)}{\partial H_0}}.
\]  

(4.17)

If we want to write things in terms a purely real decomposition of \( H \), \( H^{ab} = H_1^{ab} + i H_2^{ab}, H_{ab} = H_1^{ab} - i H_2^{ab} \) with \( H_1 = H_1^* = \tilde{\Omega} H_1 \tilde{\Omega}^T \) and \( H_2 = H_2^* = -\tilde{\Omega} H_2 \tilde{\Omega}^T \) we can split the equation above into two parts involving only real harmonic functions and two separate holomorphic functions \( t_1 = \tilde{\Omega} t_1 \tilde{\Omega}^T \) and \( t_2 = -\tilde{\Omega} t_2 \tilde{\Omega}^T \):

\[
t^{\alpha \Sigma}_1 \equiv \frac{X^{\alpha \Sigma}_1}{X^0} \to \frac{H_1^{\alpha \Sigma} + i \frac{\partial J_4(H)}{\partial H_{1 \alpha \Sigma}}}{H^0 + i \frac{\partial J_4(H)}{\partial H_0}}, \quad t^{\alpha \Sigma}_2 \equiv \frac{X^{\alpha \Sigma}_2}{X^0} \to \frac{H_2^{\alpha \Sigma} + i \frac{\partial J_4(H)}{\partial H_{2 \alpha \Sigma}}}{H^0 + i \frac{\partial J_4(H)}{\partial H_0}}
\]  

(4.18)

This completes the multicenter composite solution of the magic octonionic supergravity.

\[
\textit{Manifest } SO(1,9)
\]

20
By uplifting the magic octonionic supergravity to \( \mathcal{d} = 6 \) and breaking \( E_{6(-26)} \rightarrow SO(1,9) \times SO(1,1) \) we can present the entropy formula using only real unconstrained charges. The set of 27 real \( \mathcal{d} = 5 \) coordinates includes a singlet of \( SO(1,9) \) \( z \), a 10-component vector \( b^r \), and a chiral Majorana-Weyl 16-component spinor \( \psi^\alpha \). The quartic invariant of \( E_{7(-25)} \) in this basis can be presented as a function of real 56 charges \((p^0, p^I; q_0, q_I)\) where

\[
p^I = (p^z, p^r, p^\alpha) \quad q_I = (q_z, q_r, q_\alpha)
\]

Here \( p^\alpha \) are 16\(_L\) Majorana-Weyl spinors of \( SO(1,9) \) and \( q_\alpha \) are 16\(_R\) Majorana-Weyl spinors of \( SO(1,9) \).

The relation between the entropy of the magic exceptional octonionic \( N = 2 \) and \( N = 8 \) supergravity in this basis is simple. The metric \( \eta_{rs} \) for \( N = 8 \) is that of \( SO(5,5) \) instead of \( SO(1,9) \) and the \( \gamma \)-matrices in the second term of cubic invariants are the one for \( SO(5,5) \) instead of \( SO(1,9) \).

The choice of the most appropriate basis for the solutions, the one with manifest \( E_{6(-26)} \) or with manifest \( SO(1,9) \) may depend on the problem.

\begin{equation}
J_4(p^0, p^I; q_0, q_I) = -(p \cdot q)^2 + 4 \left( q_0 I_3(p) - p^0 I_3(q) + \{ I_3(q), I_3(p) \} \right) \tag{4.20}
\end{equation}

where the cubic invariants of the \( E_{6(-26)} \) in this basis are given by

\[
I_3(p^I) = \frac{1}{2} \left( p^z \eta_{rs} p^r p^s + (\gamma_r)_{\alpha\beta} p^r p^\alpha p^\beta \right) \quad I_3(q_I) = \frac{1}{2} \left( q_z \eta_{rs} q_r q_s + (\gamma^r)^{\alpha\beta} q_r q_\alpha q_\beta \right)
\]

and the scalar product of charges and the Poisson bracket of cubic invariants are defined as follows

\[
p \cdot q \equiv p^0 q^0 + p^r q^r + p^\alpha q^\alpha \quad \{ I_3(q), I_3(p) \} \equiv \frac{\partial I_3(q)}{\partial q_I} \frac{\partial I_3(p)}{\partial p^I}
\]

Here \( \eta_{rs} \) is the Lorentzian metric for the space \( SO(1,9) \) and \( \gamma \) are the chiral \( \gamma \)-matrices.

The 5\(_N=8\) BPS composites

5.1 Solutions via truncation to quaternionic magic \( N = 2 \)

\( N = 8 \) supergravity with its \( \frac{E_{7(7)}}{SU(8)} \) coset space can be consistently truncated to \( N = 2 \) quaternionic magic supergravity with a \( \frac{SO^*(12)}{U(6)} \) coset space [2]. Hence 1/2 BPS multi-center solutions of quaternionic magic supergravity are simultaneously 1/8 BPS multi-center solutions of \( N = 8 \) \( d = 4 \) supergravity. This fact has the useful consequence that to generate multi-center solutions
of $N = 8$ we can use magic $N = 2$ and the procedure summarized in the section 2 to extend the black hole attractor equations to stabilization equations in terms of harmonic functions.

Our first step is to identify the proper set of 32 charges which transforms as a real spinor representation of the $SO^*(12)$ duality group and are truncated from the 56-dimensional fundamental representation of the $E_7(7)$ duality group of the $N = 8$ model. Note that $E_7(7)$ decomposes into $SU(2) \times SO^*(12)$ and therefore we can use the following split

$$E_7(7) \to SU(2) \times SO^*(12), \quad 56 \to (2, 12) + (1, 32)$$

(5.1)

Thus we keep all 32 charges of $SO^*(12)$ and take as vanishing the remaining 24 charges of $E_7(7)$. For this quaternionic magic model we can use the formulas in sec. 3 to write our expressions with the $G_5 = SU^*(6)$ duality manifest or with the $G_6 = SO(1, 5)$ manifest. In both cases we naturally work with complex coordinates in $d = 5$ satisfying a reality constraint. We will quickly recap here the solution with manifest $G_5 = SO^*(12)$ duality, and then discuss some interesting features of the solution obtained from the truncated $N = 8$ structure.

One 32-dimensional set of charges of $SO^*(12)$ is given by

$$\Gamma = (p^0, p^{\Lambda \Sigma}; q_0, q_{\Lambda \Sigma})$$

(5.2)

where the 15 $p^{\Lambda \Sigma}$s satisfy the constraint $p = \Omega p^T \Omega$ and similarly for the electric charges $q_{\Lambda \Sigma}$. These charges transform as an anti-symmetric plus singlet of the compact subgroup $H_5 = USp(6)$. As in the octonionic case, we can split our charges as $p^{\Lambda \Sigma} = p^{\Lambda \Sigma}_1 + ip^{\Lambda \Sigma}_2$ and $q_{\Lambda \Sigma} = q_{1 \Lambda \Sigma} - iq_{2 \Lambda \Sigma}$. The cubic invariants are

$$I_3(p) = +Pf p, \quad I_3(q) = -Pf q,$$

(5.3)

and the scalar product of charges and the Poisson bracket of cubic invariants are defined as follows

$$p \cdot q \equiv p^0 q_0 + \frac{1}{2} p^{\Lambda \Sigma} q_{\Lambda \Sigma} = p^0 q_0 + \frac{1}{2} p_1^{\Lambda \Sigma} q_{1 \Lambda \Sigma} + \frac{1}{2} p_2^{\Lambda \Sigma} q_{2 \Lambda \Sigma},$$

(5.4)

$$\{I_3(q), I_3(p)\} \equiv 2 \frac{\partial I_3(q)}{\partial p^{\Lambda \Sigma}} \frac{\partial I_3(p)}{\partial q_{\Lambda \Sigma}} = \frac{1}{2} \frac{\partial I_3(q)}{\partial q_{1 \Lambda \Sigma}} \frac{\partial I_3(p)}{\partial p_1^{\Lambda \Sigma}} + \frac{1}{2} \frac{\partial I_3(q)}{\partial q_{2 \Lambda \Sigma}} \frac{\partial I_3(p)}{\partial p_2^{\Lambda \Sigma}}.$$

(5.5)

allowing us to define $J_4(p^0, q_0, p^{\Lambda \Sigma}, q_{\Lambda \Sigma})$ (if we map $p^0$ and $q_0$ to $Z^{78}$ and $Z_{78}$ in \cite{33} we get a matching expression for $J_4$).

The quaternionic multi-center solution starts with $(1 + n)$ of these spinor representations of $SO^*(12)$. Here $n$ is the number of centers of our multicenter solution

$$h \equiv (h^0, h^{\Lambda \Sigma}; h_0, h_{\Lambda \Sigma}) , \quad \Gamma_s \equiv (p^0, p^{\Lambda \Sigma}; q_0, q_{\Lambda \Sigma})_s \quad s = 1, \ldots, n.$$  

(5.6)

The first set $h$ defines the value of the harmonic spinor at infinity. The other $n$ spinors are the quantized electric and magnetic charges at each center. We can now introduce a constrained complex harmonic function:

$$H(\vec{x}) = (H^0, H^{\Lambda \Sigma}; H_0, H_{\Lambda \Sigma}) = h + \sum_{s=1}^n \frac{\Gamma_s}{|\vec{x} - \vec{x}_s|},$$

(5.7)
which gives the stationary metric for the BPS multicenter solution as
\[ ds_4^2 = -J_4^{-1/2}(\vec{x})(dt + \vec{\omega}d\vec{x})^2 + J_4^{1/2}(\vec{x})d\vec{x}^2. \]  
(5.8)
The standard definitions still apply:
\[ J_4(\vec{x}) \equiv J_4 \circ \mathbf{H}(\vec{x}), \quad J_4(\mathbf{h}) = 1, \quad \nabla \times \vec{\omega} = \langle \mathbf{H}, \nabla \mathbf{H} \rangle, \]  
(5.9)
as well as the usual integrability condition:
\[ \sum_{t=1}^{n} \frac{\langle \Gamma_s, \Gamma_t \rangle}{|\vec{x}_s - \vec{x}_t|} + \langle \Gamma_s, \mathbf{h} \rangle = 0. \]  
(5.10)
The vector fields are combined into a constrained complex matrix defined just as in eq.(4.16). The 30 scalars of magic quaternionic model represent the coset space \( SO^*(12)U(6) \). We will give the solution for them using our modified attractor equations (3.49)
\[ \frac{X^{\Lambda \Sigma}}{X^0}(\vec{x}) = \frac{H^{\Lambda \Sigma} + 2i\sqrt{J_4(H)}}{p^0 + i\sqrt{J_4(H)}} \]  
also written as
\[ \frac{X_1^{\Lambda \Sigma} + iX_2^{\Lambda \Sigma}}{X^0}(\vec{x}) = \frac{H_1^{\Lambda \Sigma} + i\sqrt{J_4(H)}}{p^0 + i\sqrt{J_4(H)}} + i\frac{H_2^{\Lambda \Sigma} + i\sqrt{J_4(H)}}{p^0 + i\sqrt{J_4(H)}}. \]  
(5.11)
At infinity the scalars take values defined by the harmonic function at infinity, when \( \mathbf{H} = \mathbf{h} \). Near each center at \( \vec{x} = \vec{x}_s \) the attractor values of the scalars are defined as follows
\[ \frac{X^{\Lambda \Sigma}}{X^0}(\vec{x}_s) = \frac{H^{\Lambda \Sigma} + 2i\sqrt{J_4(H)}}{p^0 + i\sqrt{J_4(H)}} \bigg|_{H \to r_s^+}. \]  
(5.12)
Note that if we would keep only canonical 3+3 of our 15 \( p \) and \( q \) we would reproduce the Caley’s Hyperdeterminant of the 2x2x2 matrix, which will also give us the STU multicenter black hole solutions.

5.2 On the uniqueness of the \( N = 8 \) multi-center BPS solution

We want to argue that the most general 1/8 BPS multicenter solution is given by the 1/2 BPS solution of the quaternionic magic \( N=2 \) supergravity, modulo and overall \( SU(8) \) rotation. Indeed, using an \( SU(8) \) rotation one can diagonalize the \( N = 8 \) central charge matrix so that it has only the non-vanishing complex eigenvalues \( z_1 = Z_{12}, z_2 = Z_{34}, z_3 = Z_{56}, z_4 = Z_{78} \). As argued in [9] this leads to a condition at each attractor point that
\[ z_1z_2 + z^*z^4 = 0 \]
\[ z_1z_3 + z^*z^4 = 0 \]
\[ z_2z_3 + z^*z^4 = 0 \]  
(5.14)
The 1/8 BPS condition for the one center solution requires that
\[ z_1 = \rho_{BPSE}e^{i\varphi_1} \neq 0 \quad z_2 = z_3 = z_4 = 0 \quad J_4^{BPSE} = \rho_{BPSE}^A > 0 \quad M_{BPSE} = |z_1| \quad (5.15) \]

The largest central charge \( z_1 \) belongs naturally to a unique \( N = 2 \) subalgebra. This defines a decomposition of the \( N = 8 \) supergravity. In the compact basis \( SU(2) \times SU(6) \) under \( SU(8) \) of \( E_7(7) \)
\[ 28 \text{ of } SU(8) \rightarrow 28 = (1, 15) + (2, 6) + (1, 1) \text{ of } SU(2) \times SU(6) \quad (5.16) \]

The 1 + 15 are antisymmetric and make a total of 32 (16 electric and 16 magnetic). This is precisely like 28 + 28 combine in 56 of \( E_7(7) \). However, now the 32 combine into a single chiral spinor of \( SO^*(12) \), which is the U duality group of the truncated \( N = 2 \) theory!

In general finding 1/8 BPS solutions of \( N = 8 \) supergravity, 1/6 BPS solutions of \( N = 6 \) supergravity, 1/4 BPS solutions to \( N = 4 \) supergravity always involves a consistent truncation to the appropriate \( N = 2 \) supergravity. For the case of \( N = 8 \), this gives the quaternionic magic \( N = 2 \) supergravity.

**Multi-center case**

It is plausible that for the solution to be BPS each center must be 1/2 BPS with respect to the same \( N = 2 \) algebra. In such a case, we may truncate our theory to a single copy of the magic supergravity with vector multiplets only and ignore all the hypermultiplets, as we did in the previous section, and we may claim that there are no other solutions. The question which one still may like to ask here is the following. Is it necessary that near each center the BPS solution can only belong to the same \( N = 2 \) supergravity? Could it be that more general solution has different \( N = 2 \) parts of \( N = 8 \) as unbroken supersymmetry? For example, at some center instead of eq. (5.15) we may have
\[ z_1 = z_3 = z_4 = 0 \quad z_2 = \rho_{BPSE}e^{i\varphi_1} \neq 0 \quad J_4^{BPSE} = \rho_{BPSE}^A > 0 \quad M_{BPSE} = |z_2| \quad (5.17) \]

If this is possible, one would not be able to provide the most general 1/8 BPS solution of \( N = 8 \) by truncating it to quaternionic magic \( N = 2 \). However, we find this scenario extremely unlikely.

One way to resolve this issue and get a definite statement on the uniqueness/ non-uniqueness of our solution is by using the study of the Killing spinors of 1/8 BPS multicenter solutions in \( N = 8 \) supergravity. These Killing spinors have not been constructed so far. We find it plausible to conjecture that such Killing spinors have the structure seen in many other examples, namely: the Killing spinor depends on \( \vec{x} \) via a common factor \( e^{U(\vec{x})/2} \), where the \( g_{tt} = e^{2U(\vec{x})} \), and a phase. The Killing spinor also satisfies some constraint imposed globally. Thus the projector usually acts on a part of the spinor which is \( \vec{x} \)-independent. The reason is that the supersymmetry transformations [1] are given by,
\[
\delta \Psi_{\mu A} = D_\mu \epsilon_A + Z_{AB \mu \nu} \gamma^\nu \epsilon_B , \quad (5.18)
\]
\[
\delta \chi_{ABC} = \gamma^\mu P_{\mu A B C D} \epsilon_D + Z_{[AB \mu \nu} \sigma^{\mu \nu} \epsilon_C ] . \quad (5.19)
\]

They consist of two parts, the first is the gravitino in eq. (5.18) while the second is for the spin 1/2 fields in eq. (5.19). We deal only with the global part of the spinor for the spin 1/2
equation while for the gravitino we have a spin connection which tells us how the Killing spinor acquires an $\vec{x}$-dependent factor, which depends on some harmonic functions. The projector, however, is the same for all centers if the Killing spinor in $N = 8$ has the same features as the one in truncated models. If our expectation about the structure of Killing spinors are justified, we will be able to confirm definitely the uniqueness of our solution for 1/8 BPS multicenter black holes in $N = 8$ supergravity.

5.3 Alternate $N = 8$ structure and the black hole mass formula

In the material above, we have written the gauge fields, prepotential and complex scalars using a symplectic section nicely adapted to an invariance under the five-dimensional U-duality group $G_5$. We will not give any description of the gauge fields or the coset which is explicitly invariant under $G_4$, but is useful to point out such a description is easily available for understanding the metric. This description uses some of the inherent structures of the $N = 8$ U-duality group, and by truncation applies to the quaternionic magic $N = 2$. It is likely that the modification of a reality structure yields a similar structure for the octonionic magic $N = 2$, but we will not explore this here.

The $N = 8$ U-duality group in four dimensions, $G_4 = E_{7(7)}$ has a covariant cubic map [32] from the the fundamental representation back into itself and an invariant symplectic form. Each $56 (X^{ij}, X_{ij})$, is spanned by the two antisymmetric real tensors $X^{ij}$ and $X_{ij}$ and the action of $E_{7(7)}$ on them is realized in a standard way [1], [32].

The quadratic invariant of $E_{7(7)}$, an invariant symplectic form, can be constructed from any two distinct fundamentals as follows.

$$\langle X, Y \rangle = -\langle Y, X \rangle := X^{ij}Y_{ij} - X_{ij}Y^{ij} \quad (5.20)$$

Note that such quadratic invariant vanishes for a single fundamental. A triple product of any three $56 \left( (X, Y, Z)^{ij}, (X, Y, Z)_{ij} \right)$, gives a trilinear map $56 \times 56 \times 56 \to 56$.

A unique quartic invariant can be constructed from a single fundamental in the following way. One can build out of $X$ a triple which is different from the original $56$. Afterwards one can use the quadratic invariant for 2 distinct fundamentals: a fundamental and its triple: this gives the standard form of a quartic invariant via the following symplectic invariant

$$J_4(X) = \frac{1}{48} \langle (X, X, X), X \rangle \quad (5.21)$$

The triple product obeys some relations

$$(X, Y, Z) = (Y, X, Z) + 2\langle X, Y \rangle Z$$
$$(X, Y, Z) = (Z, Y, X) - 2\langle X, Z \rangle Y \quad (5.22)$$

A symplectic product of a fundamental and a triple satisfies the following relation

$$\langle (X, Y, Z), W \rangle = \langle (X, W, Z), Y \rangle - 2\langle X, Z \rangle \langle Y, W \rangle \quad (5.23)$$
The operations above have their origin in the Jordan algebra structure, $J_3^{Q_8}$ behind the magic square construction of $E_7$, and so carry over nicely to the duality group based on the restriction of this algebra to $J_3^{Q}$, the quaternionic U-duality group $G_4 = SO^*(12)$. This allows us to write for all magic supergravities:

$$J_4(\vec{x}) = \frac{1}{48} < (H(\vec{x}), H(\vec{x}), H(\vec{x})), H(\vec{x})>.$$  \hspace{1cm} (5.24)

Here the corresponding harmonic function $H(\vec{x})$ is in the fundamental representation of $E_7(-25)$ for octonionic magic and in the appropriate representation for all its truncations. In particular, in the quaternionic case this expression provides us with the mass of the N=8 supergravity solution. We can efficiently extract this mass for our solution using the form for $J_4$ above, $J_4$'s fall-off at infinity, the properties of the cubic map and the fact that $<h, \Gamma> = 0$. The ADM mass is defined via the asymptotic behavior of the time-time component of the metric, as

$$g_{tt} = J_4^{-1/2} = 1 - \frac{2M_{ADM}}{|\vec{x}|} + ...$$ \hspace{1cm} (5.25)

and we take into account that in our case

$$<h, h, h, \Gamma> = <(\Gamma, h, h), h> = <(h, \Gamma, h), h> = <(h, h, \Gamma), h>.$$ \hspace{1cm} (5.26)

This gives us a magic black hole mass formula:

$$M_{ADM} = \frac{1}{48} <(h, h, h), \Gamma>.$$ \hspace{1cm} (5.27)

6 BPS and non-BPS, $N = 2$ versus $N = 5, 6, 8$

6.1 $N = 2$ attractors, general case

In [9], simple algebraic attractor equations were derived for regular black holes of $N = 8$ supergravity. These equations have 2 solutions: one BPS with 1/8 of unbroken supersymmetry and one non-BPS solution with all supersymmetries broken. In the case of the STU truncation the $N = 8$ attractor equations correspond to the attractor equations of the $N = 2$ model.

Here we would like to analyse the generic attractor equations for N=2 theory and apply this analysis to BPS and non-BPS attractors of magic supergravities. For this purpose we start with a special geometry identity:

$$p^\Sigma + i\frac{\partial I_1}{\partial q_\Sigma} = 2i\bar{Z}L^\Sigma + 2iG^{ij}\mathcal{D}_j Z \mathcal{D}_j \bar{L}^\Sigma$$

$$q_\Sigma - i\frac{\partial I_1}{\partial p^\Sigma} = 2i\bar{Z}M_\Sigma + 2iG^{ij}\mathcal{D}_j Z \mathcal{D}_j \bar{M}_\Sigma$$ \hspace{1cm} (6.1)

The identity follows from the definition

$$\mathcal{D}_j Z = (\partial_j + \frac{1}{2}K_j)Z$$ \hspace{1cm} (6.2)
where the central charge $Z = L^\Lambda q_\Lambda - M_\Lambda p^\Lambda$. One contracts eq. (6.2) with $D_j \bar{L} G^{ij}$ and proceed as in eqs. (40)-(45) in [8] where the BPS attractor equations were derived assuming $DZ = 0$. A simplified form of these equations, which uses the derivatives of the $I_1(p,q)$-invariant, was explained in Sec. 3.2 in [36]. Equations (6.1) supplemented by a vanishing of the first derivative

$$2 \bar{Z} D_i Z + i c_{ijk} G^{ijl} G^{kl} D_j \bar{Z} D_k \bar{Z} = 0 \quad [8],$$

have 3 types of solutions.

1. 1/2 BPS, well known

$$p^\Sigma + i \frac{\partial I_1}{\partial q_\Sigma} = 2i \bar{Z} L^\Sigma, \quad q_\Sigma - i \frac{\partial I_1}{\partial p_\Sigma} = 2i \bar{Z} M_\Sigma,$$

(6.3)

2. Non-BPS, $Z = 0$, less known

$$p^\Sigma + i \frac{\partial I_1}{\partial q_\Sigma} = 2i G^{ij} D_j Z D_j \bar{L}^\Sigma, \quad q_\Sigma - i \frac{\partial I_1}{\partial p_\Sigma} = 2i G^{ij} D_j Z D_j \bar{M}_\Sigma,$$

(6.4)

If $2i G^{ij} D_j Z D_j \bar{L}^\Sigma = 2i \sum_{i=2}^4 \bar{Z}_i D_I \bar{L}_I$ we find a solution of the following type

$$Z_2 \neq 0 \quad Z_3 = 0 \quad Z_4 = 0$$

(6.5)

and we can solve the attractor equations as follows

$$\frac{p^\Lambda + i \frac{\partial I_1}{\partial q_\Lambda}}{p^0 + i \frac{\partial I_1}{\partial q_0}} = \frac{\bar{D}_2 \bar{L}^\Lambda}{\bar{D}_2 \bar{L}^0}, \quad \frac{q_\Lambda - i \frac{\partial I_1}{\partial p_\Lambda}}{q_0 - i \frac{\partial I_1}{\partial p_0}} = \frac{\bar{D}_2 \bar{M}_\Lambda}{\bar{D}_2 \bar{M}_0}$$

(6.6)

3. Non-BPS, $Z \neq 0$. This case is particularly tractable for the models with the coset spaces $G_4/H_4$ of rank 3. All magic models have rank 3. The matrix of central charge derivatives $DZ$ has an $H_4$ symmetry:

- $J^O \Rightarrow H_4 = E_6 \times U(1)$
- $J^Q \Rightarrow H_4 = SU(6) \times U(1)$
- $J^C \Rightarrow H_4 = SU(3) \times SU(3) \times U(1)$
- $J^R \Rightarrow H_4 = SU(3) \times U(1)$

With an $H_4$ rotation we can skew-diagonalize the matrix $DZ$ so that it has 3 eigenvalues, $\bar{D}_I Z \equiv Z_I, I = 2, 3, 4$. Together with $N = 2$ central charge $Z = -iZ_1$ we have 4 charges,

$$Z_1 = iZ, Z_I = \bar{D}_I \bar{Z}$$

(6.7)

The extremality condition is given by

$$2 \bar{Z} D_I Z + i d_{IJK} \bar{Z}_J \bar{Z}_K \delta^{IJ} \delta^{K\bar{K}} = 0$$

(6.8)

and it becomes for magic models in normal frame

$$Z_1 Z_2 + Z^* Z^4 = 0$$
$$Z_1 Z_3 + Z^* Z^4 = 0$$
$$Z_2 Z_3 + Z^* Z^4 = 0$$

(6.9)
exactly as in $N = 8$ case [9]. The solution is as in [9]

$$Z_1 = \rho e^{i(\pi - 3\phi)}, \quad Z_4 = \rho e^{i\phi}, \quad Z_1 Z_2 Z_3 Z_4 = \rho^4 e^{i\pi} = -\rho^4$$  \hspace{1cm} (6.10)

Here we have satisfied the attractor equations: $Z_I Z_J + Z^K_M Z^*_M = 0, \quad I \neq J \neq K \neq M$.

We may use this in an identity (6.1) and we find in case that $\phi = 0$

$$p^\Sigma + i \frac{\partial I_1}{\partial q^\Sigma} = 2\rho (L^\Sigma + i \sum I \bar{D}_I \bar{L}^\Sigma)$$  \hspace{1cm} (6.11)

$$q_{\Sigma} - i \frac{\partial I_1}{\partial p^\Sigma} = 2\rho (M_{\Sigma} + i \sum I \bar{D}_I \bar{M}_{\Sigma})$$  \hspace{1cm} (6.12)

and

$$\frac{p^\Sigma + i \frac{\partial I_1}{\partial q^\Sigma}}{p^0 + i \frac{\partial I_1}{\partial q^0}} = \frac{L^\Sigma + i \sum I \bar{D}_I \bar{L}^\Sigma}{L^0 + i \sum I \bar{D}_I \bar{L}^0}, \quad \frac{q_{\Sigma} - i \frac{\partial I_1}{\partial p^\Sigma}}{q_0 - i \frac{\partial I_1}{\partial p^0}} = \frac{M_{\Sigma} + i \sum I \bar{D}_I \bar{M}_{\Sigma}}{M_0 + i \sum I \bar{D}_I \bar{M}_0}$$  \hspace{1cm} (6.13)

Equations (6.13) define the values of moduli fields in terms of charges in the non-BPS case. We have verified [37] that these equations provide a non-BPS attractor solution of the STU model in example studied in [11].

### 6.2 $N = 2$ versus $N = 5, 6, 8$ and BPS versus non-BPS

The extended supergravities $N = 5, 6$ as well as $N = 8$ have no matter multiplets, only the gravitational one. Therefore they are very restricted. BPS black holes in these theories have been studied before in [38]. Since now we understand certain features on $N = 8$ BPS and non-BPS black holes and their relation to $N = 2$ BPS and non-BPS ones, as shown in [9], we may extend this relations also to include $N = 6$ and $N = 5$ theories.

The $N = 8$ supergravity has a consistent truncation both to $N = 6$ and $N = 2$ supergravity theories, depending on whether one keeps six or two of the eight gravitinos. In both cases one ends up with $J^Q$ based on the $SO^{*}(12)/U(6)$ manifold. This manifold, because it is consistent with $N=2$, is indeed a special Kähler symmetric space. The charge vector is in the $32$ of $SO^{*}(12)$ (chiral spinor). From an $N = 2$ point of view

$$32 = 15 + \bar{15} + 1 + \bar{1}$$  \hspace{1cm} (6.14)

where $15$ are in the matter multiplet and $1$ is the graviphoton.

In $N = 6$ supergravity the central charge is $Z^{AB} = -Z^{BA}$ with $A, B = 1, \ldots, 6$ and there is also a singlet charge $Z$ [38]. The black hole potential is

$$V_{BH} = \frac{1}{2} Z_{AB} Z^{AB} + Z \bar{Z},$$  \hspace{1cm} (6.15)
The 1/6 BPS solution of N=6 theory was identified in [38]: it has $Z = 0$ and a skew-diagonal $Z_{AB}$ with $Z_{12} \neq 0$ and $Z_{34} = Z_{56} = 0$. From the $N = 2$ point of view such solution is a non-BPS solution (with vanishing central charge). On the other hand, the 1/2 BPS solution of $N = 2$ theory with $Z \neq 0$ and $Z_{AB} = 0$ corresponds to a non-BPS solution of the N=6 theory (with vanishing $N = 6$ central charge).

Let us now comment on $N = 5$ theory and relation of 1/5 BPS solution to 1/2 BPS of the relevant $N = 2$ theory. The $N = 5$ special Kähler geometry theory is based on the $\text{SU}(5,1)/\text{U}(5)$ symmetric space with 5 complex scalars. The central charge $Z_{AB}$ has a non-vanishing component $Z_{12} \neq 0$ whereas $Z_{34} = 0$ is vanishing: this is a 1/5 BPS solution. Despite the fact that the $N = 5$ model has the same sigma model as the corresponding $N = 2$ theory, the vector part of these models is different: $N = 2$ has six vectors and $N = 5$ has ten vectors. Therefore, as different from $N = 6$ case, the 1/5 BPS solutions of $N = 5$ are different from the 1/2 of $N = 2$ supergravity based on $\text{SU}(5,1)/\text{U}(5)$ symmetric space.

7 Discussion

We have given a complete description of the cubic and quartic invariants for magic supergravities. Using these we were able to write down multi-center 1/2 BPS solutions for the exceptional octonionic magic supergravity with a 56-component charge vector at each center. This model has $E_{7(-25)}$ duality symmetry, which is slightly different from the duality symmetry of $N = 8$ theory, $E_{7(7)}$. We have also demonstrated how multi-center quaternionic magic solutions with 32-component charge at each center can be used to describe a broad (possibly complete) array of $N = 8$ 1/8 BPS multi-center solution.

In [28], it was shown that any single center 1/8 BPS solution of $N = 8$ could be rotated using an $E_{7(7)}$ duality transformation to a form involving just a single center solution the $N=2$ STU truncation of $N = 8$. For the non-BPS attractors of magic supergravities we have made used of this truncation to the STU model. Using the solutions of the non-BPS attractors for $N = 8$ supergravity presented in [9] we have found the general solutions of the non-BPS attractors for the $N = 2$ STU model.

The duality rotation which takes a single center solution to an STU form is not powerful enough for a generic 1/8 BPS solution with multiple centers: such a rotation is clearly not enough to rotate every center to STU form. For solutions generated via quaternionic magic supergravity, it would be interesting to work out which is the minimal $N = 2$ truncation that admits generic cases with two, three, four etc... centers, just as STU supergravity is minimal truncation of $N = 8$ which describes all single center solutions.

Clearly, $N = 2$ truncations play a major role in understanding 1/8 BPS states of $N = 8$ supergravity. Apart from clarifying if $N = 2$ truncation is the only way to achieve 1/8 BPS configurations, a clearer look at these truncations in context of black holes might be productive. Different truncations allow a varying number of hypermultiplet scalars (see e.g. [39]) from the $N = 8$ coset to survive. The hyper scalars vevs typically play a spectator role, since they will
not vary in BPS black hole solutions, but for an observer near infinity they may reveal some information about just which $N = 2$ truncation is in operation for a given multi-center solution.

For future reference, we summarize the hyper-multiplet scalars which the survive from the $N = 8$ coset in the truncation to the various magic supergravities (for details see [39]). In quaternionic magic, we have $E_7(7) \to SO^*(12) \times SU(2)$ and no hypers survive. In the complex case, $E_7(7) \to SU(3,3) \times SU(2,1)$ and we get four hyper scalars parameterizing the coset $SU(2,1)/SU(2) \times U(1)$; the real case corresponds to $E_7(7) \to Sp(6,\mathbb{R}) \times G_{2(2)}$ and has eight hyper scalars, $G_{2(2)}/SO(4)$. The STU truncations has hyper scalars in $SO(4,4)/SO(4) \times SO(4,4)$ (see [28]). As we restrict the vector moduli space, the hyper moduli space increases in multiples of four (it is quaternionic).

Finally, given the relationship between extremal non-BPS and BPS solutions of $N = 2$ magic quaternionic supergravity and those of $N = 6$, it would be interesting to understand just how many such links exists between various extremal solutions of different supergravities, perhaps along lines related to [36].

Acknowledgments

It is a pleasure to thank R. D’Auria, V. Balasubramanian, O. Ganor, M. Günaydin, T. Levi, T. Ortín, B. Pioline, A. Sen, A. Tomasiello and M. Trigiante for useful conversations. We are grateful to participants of 2006 Frascati Winter school on Attractor mechanism for the interest to this work. The work of S.F. has been supported in part by the European Community Human Potential Program under contract MRTN-CT-2004-005104 “Constituents, fundamental forces and symmetries of the universe”, in association with INFN Frascati National Laboratories and by D.O.E. grant DE-FG03-91ER40662, Task C. The work of EG was supported by the US Department of Energy under contracts DE-AC03-76SF00098 and DE-FG03-91ER-40676 and by the National Science Foundation under grant PHY-00-98840. The work of R.K. was supported by NSF grant PHY-0244728.

Appendix A: $d = 6$ duality and quaternionic magic model

This model even if we break the manifest $G_5$ duality down to $G_6$ duality still does not give us a very special real geometry in $d = 5$ and the relevant complexified special geometry in $d = 4$. The reason is that the spinors of $SO(1, 5) \sim SU^*(4)$ as opposite to other magic models with $SO(1, 2)$, $SO(1, 3)$, $SO(1, 9)$ do not have real representations. Thus we have to use complex coordinates constrained by reality conditions.

Both the vectors and the spinors are complex and satisfy the reality condition. For vectors of $SU^*(4)$ we use an antisymmetric matrix $V_{AB} = -V_{BA}$ and $V^* = -\Omega V \Omega$ where $\Omega = -\Omega^t$, $\Omega^2 = -1$. Thus $V_1 = \text{Re} V = -\Omega \text{Re} V \Omega$ and $V_2 = \text{Im} V = \Omega \text{Im} V \Omega$. We also introduce $\tilde{V}^{AB} \equiv \frac{1}{2} \epsilon^{ABCD} V_{CD}$. Thus the first term in the cubic invariant, $z[(V^0)^2 - \sum_{i=1}^{5} (V_i)^2]$ can be rewritten as $z V_{AB} \tilde{V}^{AB} = z (V_{1AB} \tilde{V}_{1AB} - V_{2AB} \tilde{V}_{2AB})$ in terms of real entries only. For spinors $\Psi_{Ai}$ of $SU^*(4) \times SU(2)$ with $A = 1, \ldots, 4, i = 1, 2$ there is a reality condition $\Psi^*_{Ai} = \Omega_{AB} \epsilon_{ij} \Psi_{Bj}$ and
\[ \epsilon^2 = -1. \] We also split these spinors into real and imaginary parts which we call \( \Psi_1 \) and \( \Psi_2 \) respectively, \( \Psi_1 = \Omega \epsilon \Psi_1 \) and \( \Psi_2 = -\Omega \epsilon \Psi_2. \) The second term of the cubic invariant can now be presented as a function of only real coordinates, \[ V_{1AB}(\Psi_1^{A_i} \Psi_1^{B_j} - \Psi_2^{A_i} \Psi_2^{B_j}) - V_{2AB}(\Psi_1^{A_i} \Psi_2^{B_j} + \Psi_2^{B_j} \Psi_1^{A_i}) \epsilon_{ij}. \]

Now we need the cubic invariants in terms of charges. We take \( p^z, p^r, p^\alpha \Rightarrow p^z, p_{AB}, p^{Ai} \), \( q_z, q_r, q_\alpha \Rightarrow q_z, q^{AB}, q_{Ai} \), \( p_{AB} = -p_{BA} \) \( q^{AB} = -q^{BA} \) (7.1) and \[ p \cdot q = p^z q_z + p_{AB} q^{AB} + p^{Ai} q_{Ai} \]
\[ \bar{p}^{AB} = \frac{1}{2} \epsilon^{ABCD} p_{CD} , \]
\[ \bar{q}_{AB} = \frac{1}{2} \epsilon^{ABCD} q_{CD} \]

The cubic invariants are \[ I_3(p) = \frac{1}{2} \left( p^z p_{AB} \bar{p}^{AB} + p_{AB} \bar{p}^{Ai} p^{Bj} \epsilon_{ij} \right) \] (7.2) \[ I_3(p) = \frac{1}{2} \left( q_z q^{AB} \bar{q}_{AB} + q^{AB} q_{Ai} q_{Bj} \epsilon_{ij} \right) \] (7.3) where the reality conditions are \[ p^* = -\Omega p \Omega , \]
\[ p^{*Ai} = (\Omega \epsilon p)^{Ai} \] (7.4) As the consequence of the reality conditions the cubic invariants are real \[ I_3^*(p) = I_3(p^*) = I_3(p) \] (7.5) and the same for electric charges. We may split the charges into real and imaginary parts \[ p_{AB} = p_{1AB} + ip_{2AB} \]
\[ p^{Ai} = p_1^{Ai} + ip_2^{Ai} \] (7.6) and find for \( I_3(p) \)
\[ \frac{1}{2} \left( p^z(p_{1AB} \bar{p}_1^{AB} - p_{2AB} \bar{p}_2^{AB}) + [p_{1AB}(p_1^{Ai} p_1^{Bj} - p_2^{Ai} p_2^{Bj}) - p_{2AB}(p_1^{Ai} p_2^{Bj} + p_2^{Ai} p_1^{Bj})] \epsilon_{ij} \right) \] (7.7) and analogous for \( I_3(q) \).

**References**

[1] E. Cremmer and B. Julia, “The SO(8) Supergravity,” Nucl. Phys. B 159, 141 (1979); B. de Wit and H. Nicolai, “N=8 Supergravity,” Nucl. Phys. B 208, 323 (1982).

[2] M. Gunaydin, G. Sierra and P. K. Townsend, “Exceptional Supergravity Theories And The Magic Square,” Phys. Lett. B 133, 72 (1983). M. Gunaydin, G. Sierra and P. K. Townsend, “The Geometry Of N=2 Maxwell-Einstein Supergravity And Jordan Algebras,” Nucl. Phys. B 242, 244 (1984); M. Gunaydin, G. Sierra and P. K. Townsend, “Gauging The D = 5 Maxwell-Einstein Supergravity Theories: More On Jordan Algebras,” Nucl. Phys. B 253, 573 (1985).
[3] S. Ferrara and M. Gunaydin, “Orbits of exceptional groups, duality and BPS states in string theory,” Int. J. Mod. Phys. A 13, 2075 (1998) [arXiv:hep-th/9708025].

[4] H. Freudenthal, Proc. Konink.Ned. Akad. Wetenschap A62 (1959) 447; B.A. Rozenfeld, Dokl. Akad. Nauk. SSSR 106 (1956) 600; J. Tits, Mem. Acad. Roy. Belg. Sci. 29 (1955) fasc. 3.

[5] S. Ferrara, R. Kallosh and A. Strominger, “N = 2 extremal black holes,” Phys. Rev. D 52, 5412 (1995) [arXiv:hep-th/9508072]; A. Strominger, “Macroscopic Entropy of N = 2 Extremal Black Holes,” Phys. Lett. B 383, 39 (1996) [arXiv:hep-th/9602111].

[6] S. Ferrara and R. Kallosh, “Supersymmetry and Attractors,” Phys. Rev. D 54, 1514 (1996) [arXiv:hep-th/9602136].

[7] S. Ferrara and R. Kallosh, “Universality of Supersymmetric Attractors,” Phys. Rev. D 54, 1525 (1996) [arXiv:hep-th/9603090].

[8] S. Ferrara, G. W. Gibbons and R. Kallosh, “Black holes and critical points in moduli space,” Nucl. Phys. B 500, 75 (1997) [arXiv:hep-th/9702103].

[9] S. Ferrara and R. Kallosh, “On N = 8 attractors,” Phys. Rev. D 73, 125005 (2006) [arXiv:hep-th/0603247].

[10] R. R. Khuri and T. Ortin, “A Non-Supersymmetric Dyonic Extreme Reissner-Nordstrom Black Hole,” Phys. Lett. B 373, 56 (1996) [arXiv:hep-th/9512178]; T. Ortin, “Extremality versus supersymmetry in stringy black holes,” Phys. Lett. B 422, 93 (1998) [arXiv:hep-th/9612142]; T. Ortin, “Non-supersymmetric (but) extreme black holes, scalar hair and other open [arXiv:hep-th/9705095].

[11] P. K. Tripathy and S. P. Trivedi, “Non-Supersymmetric Attractors in String Theory,” arXiv:hep-th/0511117; R. Kallosh, N. Sivanandam and M. Sorosh, “The non-BPS black hole attractor equation,” JHEP 0603, 060 (2006) [arXiv:hep-th/0602005]; S. Bellucci, S. Ferrara and A. Marrani, “On some properties of the attractor equations,” Phys. Lett. B 635, 172 (2006) [arXiv:hep-th/0602161].

[12] B. Bates and F. Denef, “Exact solutions for supersymmetric stationary black hole composites,” arXiv:hep-th/0304094; F. Denef, “On the correspondence between D-branes and stationary supergravity solutions of type II Calabi-Yau compactifications,” arXiv:hep-th/0010222; F. Denef, “Supergravity flows and D-brane stability,” JHEP 0008, 050 (2000) [arXiv:hep-th/0005049].

[13] A. Ceresole, R. D’Auria and S. Ferrara, “The Symplectic Structure of N=2 Supergravity and its Central Extension,” Nucl. Phys. Proc. Suppl. 46, 67 (1996) [arXiv:hep-th/9509160].

[14] K. Behrndt, D. Lust and W. A. Sabra, Nucl. Phys. B 510, 264 (1998) [arXiv:hep-th/9705169]; G. Lopes Cardoso, B. de Wit, J. Kappeli and T. Mohaupt, JHEP 0012, 019 (2000) [arXiv:hep-th/0009234].
[15] E. Bergshoeff, R. Kallosh and T. Ortin, “Stationary Axion/Dilaton Solutions and Super-symmetry,” Nucl. Phys. B 478, 156 (1996) [arXiv:hep-th/9605059].

[16] V. Balasubramanian, E. G. Gimon and T. S. Levi, “Four Dimensional Black Hole Microstates: From D-branes to Spacetime Foam,” arXiv:hep-th/0606118.

[17] E. Cremmer and A. Van Proeyen, “Classification Of Kahler Manifolds In N=2 Vector Multiplet Supergravity Couplings,” Class. Quant. Grav. 2, 445 (1985).

[18] B. de Wit and A. Van Proeyen, “Special geometry, cubic polynomials and homogeneous quaternionic spaces,” Commun. Math. Phys. 149, 307 (1992) [arXiv:hep-th/9112027].

[19] B. de Wit, F. Vanderseypen and A. Van Proeyen, “Symmetry structure of special geometries,” Nucl. Phys. B 400, 463 (1993) [arXiv:hep-th/9210068].

[20] K. Behrndt, R. Kallosh, J. Rahmfeld, M. Shmakova and W. K. Wong, “STU black holes and string triality,” Phys. Rev. D 54, 6293 (1996) [arXiv:hep-th/9608059].

[21] R. Kallosh, M. Shmakova and W. K. Wong, “Freezing of moduli by N = 2 dyons,” Phys. Rev. D 54, 6284 (1996) [arXiv:hep-th/9607077].

[22] K. Behrndt and W. A. Sabra, “Static N = 2 black holes for quadratic prepotentials,” Phys. Lett. B 401, 258 (1997) [arXiv:hep-th/9702010].

[23] D. Shih, A. Strominger and X. Yin, [arXiv:hep-th/0506151]. D. Shih, A. Strominger and X. Yin, [arXiv:hep-th/0505094]. D. Gaiotto, A. Strominger and X. Yin, “New connections between 4D and 5D black holes,” JHEP 0602, 024 (2006) [arXiv:hep-th/0503217].

[24] B. Pioline, “BPS black hole degeneracies and minimal automorphic representations,” JHEP 0508, 071 (2005) [arXiv:hep-th/0506228].

[25] S. Helgason, “Differential Geometry, Lie Groups, and Symmetric Spaces”, Academic Press, 1978.

[26] R. Gilmore, “Lie Groups, Lie Algebras, and Some of Their Applications”, Dover Publications, 2006.

[27] R. Kallosh and B. Kol, “E(7) Symmetric Area of the Black Hole Horizon,” Phys. Rev. D 53, 5344 (1996) [arXiv:hep-th/9602014]; M. Cvetic and C. M. Hull, “Black holes and U-duality,” Nucl. Phys. B 480, 296 (1996) [arXiv:hep-th/9606193]; V. Balasubramanian, F. Larsen and R. G. Leigh, “Branes at angles and black holes,” Phys. Rev. D 57, 3509 (1998) [arXiv:hep-th/9704143]; M. Bertolini and M. Trigiante, “Regular BPS black holes: Macroscopic and microscopic description of the generating solution,” Nucl. Phys. B 582, 393 (2000) [arXiv:hep-th/0002191].

[28] L. Andrianopoli, R. D’Auria, S. Ferrara, P. Fre and M. Trigiante, “E(7)(7) duality, BPS black-hole evolution and fixed scalars,” Nucl. Phys. B 509, 463 (1998) [arXiv:hep-th/9707087].
[29] H. Ooguri, A. Strominger and C. Vafa, “Black hole attractors and the topological string,” Phys. Rev. D 70, 106007 (2004) [arXiv:hep-th/0405146]; V. Pestun, “Black hole entropy and topological strings on generalized CY manifolds,” arXiv:hep-th/0512189. G. Lopes Cardoso, B. de Wit, J. Kappeli and T. Mohaupt, “Black hole partition functions and duality,” JHEP 0603, 074 (2006) [arXiv:hep-th/0601108]; J. P. Hsu, A. Maloney and A. Tomasiello, “Black hole attractors and pure spinors,” arXiv:hep-th/0602142.

[30] A. Keurentjes, “The group theory of oxidation,” Nucl. Phys. B 658, 303 (2003) [arXiv:hep-th/0210178]; A. Keurentjes, “The group theory of oxidation. II: Cosets of non-split groups,” Nucl. Phys. B 658, 348 (2003) [arXiv:hep-th/0212024]; A. Keurentjes, “Oxidation = group theory,” Class. Quant. Grav. 20, S525 (2003) [arXiv:hep-th/0212050].

[31] L. Andrianopoli, S. Ferrara and M. A. Lledo, “No-scale D = 5 supergravity from Scherk-Schwarz reduction of D = 6 theories,” JHEP 0406, 018 (2004) [arXiv:hep-th/0406018].

[32] M. Gunaydin, K. Koepsell and H. Nicolai, “Conformal and quasiconformal realizations of exceptional Lie groups,” Commun. Math. Phys. 221, 57 (2001) [arXiv:hep-th/0008063].

[33] M. Gunaydin and O. Pavlyk, “Minimal unitary realizations of exceptional U-duality groups and their subgroups as quasiconformal groups,” JHEP 0501, 019 (2005) [arXiv:hep-th/0409272].

[34] M. Gunaydin and O. Pavlyk, “Generalized spacetimes defined by cubic forms and the minimal unitary realizations of their quasiconformal groups,” JHEP 0508, 101 (2005) [arXiv:hep-th/0506010].

[35] L. Andrianopoli, R. D’Auria and S. Ferrara, “U-invariants, black-hole entropy and fixed scalars,” Phys. Lett. B 403, 12 (1997) [arXiv:hep-th/9703156].

[36] R. Kallosh, “From BPS to non-BPS black holes canonically,” arXiv:hep-th/0603003.

[37] R. Kallosh, N. Sivanandam and M. Soroush, “Exact attractive non-BPS STU black holes,” arXiv:hep-th/0606263.

[38] L. Andrianopoli, R. D’Auria and S. Ferrara, “U-duality and central charges in various dimensions revisited,” Int. J. Mod. Phys. A 13, 431 (1998) [arXiv:hep-th/9612105]; L. Andrianopoli, R. D’Auria and S. Ferrara, “U-invariants, black-hole entropy and fixed scalars,” Phys. Lett. B 403, 12 (1997) [arXiv:hep-th/9703156].

[39] L. Andrianopoli, R. D’Auria and S. Ferrara, “Supersymmetry reduction of N-extended supergravities in four dimensions,” JHEP 0203, 025 (2002) [arXiv:hep-th/0110277].