The quantum open systems approach to the dynamical Casimir effect

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Abstract
We analyze the introduction of dissipative effects in the study of the dynamical Casimir effect. We consider a toy model for an electromagnetic cavity that contains a semiconducting thin shell, which is irradiated with short laser pulses in order to produce periodic oscillations of its conductivity. The coupling between the quantum field in the cavity and the microscopic degrees of freedom of the shell induces dissipation and noise in the dynamics of the field. We argue that the photon creation process should be described in terms of a damped oscillator with non-local dissipation and colored noise.

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1. Introduction

The motion of a neutral body distorts the quantum electromagnetic vacuum and may induce photon creation. This fact was first pointed out many years ago [1]. The radiation of a single mirror is, in general, an extremely small effect. For this reason, during the 1970s, the radiation of moving mirrors was mainly analyzed because of its formal relation with black hole evaporation [2]. However, it was later pointed out that, in a resonant cavity, photon creation could be enhanced by secular effects: if the length of the cavity oscillates with a frequency that is twice the frequency of an eigenmode of the static cavity, resonant effects may produce a number of photons that, under ideal conditions, grow exponentially with time [3].

Although intensively studied from the theoretical point of view, this ‘motion induced radiation’ is extremely difficult to measure. On the one hand, cavities with a very high $Q$ are needed in order to have an exponential growth during a significant amount of time. On the other hand, and more importantly, for microwave cavities the mirror should oscillate at extremely high frequencies, of the order of GHz. This is a challenge from an experimental point of view, although there are feasible proposals based on the use of nanoresonators in order to accomplish it [4].

Photon creation can occur whenever there are time-dependent external conditions: moving boundaries and/or time-dependent electromagnetic properties. This broad class of phenomena is usually termed the ‘dynamical Casimir effect’ (DCE; for a complete list of publications see [5]). In particular, a setup that has attracted both theoretical and experimental attention is the possibility of using short laser pulses in order to produce periodic variations of the conductivity of a semiconductor layer placed inside a microwave cavity. The fast changes in the conductivity induce a periodic variation in the effective length of the cavity and therefore the creation of photon pairs [6]. This setup has been analyzed at the theoretical level [7, 8], and there is an ongoing experiment aimed at the detection of the motion-induced radiation [9]. More recently, the possibility of changing the effective length of a superconducting coplanar waveguide terminated by a SQUID using a time-dependent magnetic field has also been put forward [10].

In this paper, we will be mainly concerned with the irradiated-semiconductor setup. It has been pointed out that dissipative effects may play an important role in this experiment, which could induce a significant deviation from the expected exponential growth in the number of created photons [8]. To our knowledge, up to now there is no theoretical model that includes, from first principles, the effects of dissipation and noise. There are, on the other hand, phenomenological models [11] to estimate these effects. The idea is the following: in a resonant situation, if the spectrum of the cavity is not equidistant, the photons are created in a single mode. The dynamics of this mode can be described, in the absence of dissipation, in terms of a
In this section, we will review the model of \[ 2 \]. A model without dissipation reads

The aim of this paper is to provide a first step towards a description of the DCE from first principles, including dissipation and noise. The natural arena for this description is the theory of quantum open systems. We will consider the electromagnetic field as our ‘system’, while the degrees of freedom on the semiconductor will be the ‘environment’. Under very general assumptions, it is possible to show that the dynamics of the electromagnetic field will be described by a Langevin equation (section 3). In general, this equation will include a non-local dissipative term and a colored noise, both concentrated on the position of the layer. We will use this Langevin equation to describe the dynamics of the resonant mode (section 4). We will see that the evolution of a single mode is in general more complex than that of a damped oscillator with white noise, and that all the effects come from the non-trivial boundary conditions at the position of the slab.

In order to avoid unnecessary complications we will consider several simplifying assumptions. On the one hand, we will consider a quantum scalar field in 1+1 dimensions instead of the electromagnetic field in a 3+1 cavity. On the other hand, we will assume that this quantum field is linearly coupled to the degrees of freedom of the layer. Finally, we will consider a very thin semiconductor layer. Even with these simplifications, we hope the resulting model to contain the main physics of the problem. This model can be considered as a generalization of the theory described in [7], reviewed in section 2, to the dissipative case.

2. A model without dissipation

In this section, we will review the model of [7], which we will use as starting point to include dissipative effects in the rest of the paper.

We consider a massless scalar field within a cavity with perfect conducting walls. For simplicity, we consider the 1+1-dimensional case, where we consider a cavity of size \( L \). At the midpoint of the cavity (\( x = L/2 \)) a thin film of semiconducting material is located. We model the conductivity properties of such material by a potential \( V(t) \): the ideal limit of perfect conductivity corresponds to \( V \to \infty \) and \( V \to 0 \) to a ‘transparent’ material. This potential varies between a minimum value, \( V_0 \), and a maximum, \( V_{\text{max}} \). The Lagrangian of the scalar field within the cavity is given by

\[
\mathcal{L} = \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{V(t)}{2} \delta(x - L/2) \phi^2,
\]

where \( \delta(x) \) is the one-dimensional Dirac delta function. The use of an infinitely thin film is justified as long as the width of the slab is much smaller than the wavelengths of the relevant modes in the cavity. The corresponding Lagrange equation reads

\[
(\partial_x^2 - \partial_t^2) \phi = V(t) \delta(x - L/2) \phi.
\]

As shown in [12], this model describes a plane-polarized electromagnetic field propagating normally to an infinitesimally thin jellium-type plasma sheet. The strength of the potential is given by

\[
V = 4\pi n_s e^2/m,
\]

where \( n_s \) is the surface charge density of free carriers in the sheet, \( e \) is the electron charge and \( m \) is the effective mass of the free carriers. When a laser field suddenly impinges on the plasma sheet, it produces time-dependent changes in the surface charge density \( n_s(t) \) that induce a time variation in the conductivity \( V(t) \).

For the sake of clarity we divide the cavity into two regions: region I \( (0 \leq x \leq L/2) \) and region II \( (L/2 \leq x \leq L) \). Perfect conductivity at the edges of the cavity imposes the following boundary conditions:

\[
\phi_1(x = 0, t) = \phi_1(x = L, t) = 0.
\]

This can be seen by integrating out the field (equation (2)) in the neighborhood of the film.

We consider the set of functions

\[
\psi_m(x, t) = \sqrt{\frac{2}{L}} \sin \left(k_m(t) x \right),
\]

where \( k_m(t) \) is the \( m \)th positive solution to the following transcendental equation:

\[
2k_m = -V(t) \tan \left( \frac{k_m L}{2} \right).
\]

Note that \( \psi_m \) depends on \( t \) through \( k_m(t) \).

Let us define

\[
\Psi_m(x, t) = \begin{cases} 
\psi_m(x, t), & 0 \leq x \leq L/2, \\
-\psi_m(x - L, t), & L/2 \leq x \leq L.
\end{cases}
\]

These functions satisfy the boundary conditions equation (4) and equation (5), and the orthogonality relations \( \langle \Psi_m | \Psi_n \rangle = \int_0^L [1 - \sin(k_m(t)L)/k_m(t)L] k_{m,n} \), where we have used the usual inner product in the interval \( [0, L] \).

For \( t \leq 0 \) the slab is not irradiated; consequently, \( V \) is independent of time and has the value \( V_0 \). The modes of the quantum scalar field that satisfy the Klein–Gordon equation (2) are

\[
u_m(x, t) = \frac{e^{-i\omega_m t}}{\sqrt{2\omega_m}} \Psi_m(x, 0),
\]

where \( \omega_m = k_m^0 \), and \( k_m^0 \) is the \( m \)th solution to equation (7) for \( V = V_0 \). At \( t = 0 \) the potential starts to change in time and the set \( \{ \omega_m \} \) of the eigenfrequencies of the cavity acquires a time dependence through equation (7).
We expand the field operator \( \phi \) as

\[
\phi(x, t) = \sum_m [b_m u_m(x, t) + b_m^+ u_m^*(x, t)],
\]

where \( b_m \) and \( b_m^+ \) are annihilation and creation operators, respectively. There is another set of solutions of the Klein–Gordon equation that satisfy all boundary conditions and have a node at \( x = L/2 \). For this reason their dynamics is not affected by the presence of the slab. Moreover, these modes are decoupled from the modes \( u_m \) thanks to the orthogonality conditions. Therefore, we will only consider the evolution of the modes given in equation \( (9) \).

For \( t \geq 0 \), we write the expansion of the field mode \( u_t \) as

\[
u_t(x, t > 0) = \sum_m \Psi_m(x, t).
\]

Replacing this expression into \( (\partial_t^2 - \omega_c^2)u_t = 0 \), we find

\[
P^{(s)}_n + k_m^2(t) P^{(s)}_n = -\sum_m \left[ \left( 2 P^{(s)}_m k_m + P^{(s)}_m k_m \right) S_m^{(A)} + P^{(s)}_m k_m S_m^{(B)} \right],
\]

where the coefficients \( g_m^{(A)} \) read

\[
S_m^{(A)} = \frac{1}{(\Psi_{\delta_m}, \Psi_n)} \left( \frac{\partial \Psi_m}{\partial k_m}, \Psi_n \right),
\]

\[
S_m^{(B)} = \frac{1}{(\Psi_{\delta_m}, \Psi_n)} \left( \frac{\partial^2 \Psi_m}{\partial k_m^2}, \Psi_n \right).
\]

Note that \( \delta_m = k_m(0) \).

We are interested in the number of photons created inside the cavity. Hence we focus on resonance effects induced by periodic oscillations in the conductivity \( V(t) \), which translates into effective periodic changes in the modes of the scalar field. Therefore, we start by considering a time-dependent conductivity given by

\[
V(t) = V_0 + (V_{\text{max}} - V_0) f(t),
\]

where \( f(t) \) is a periodic and non-negative function, \( f(t) = f(t + T) = 0 \), which vanishes at \( t = 0 \) and attains its maximum at \( f(t_\tau) = 1 \). In each period, \( f(t) \) describes the excitation and relaxation of the semiconductor produced by the laser pulse. Typically, the characteristic time of excitation \( t_\tau \) is the smallest time scale and satisfies \( t_\tau \ll T \). We expand \( f(t) \) in a Fourier series

\[
f(t) = f_0 + \sum_{j=1}^{\infty} f_j \cos(j \Omega t + c_j),
\]

where \( \Omega = 2\pi/T \). Since \( t_\tau \) is the smallest time scale, on general grounds we expect the first \( T/t_\tau \) terms in the above series to be relevant.

Under certain constraints, large changes in \( V \) induce only small variations in \( k \) through the transcendental relation between \( k \) and \( V \) (see equation \( (7) \)). In this case, a perturbative treatment is valid and a linearization of such a relation is appropriate. Accordingly we write

\[
k_n(t) = k_n^0(1 + \epsilon_n f(t)),
\]

where \( \epsilon_n \) is obtained after replacing equations \( (14) \) and \( (16) \) into \( (7) \) and expanding it to first order in \( \epsilon_n \). The result is

\[
\epsilon_n = \frac{V_{\text{max}} - V_0}{L(k_n^0)^2 + V_0 \left( 1 + \frac{V_0L}{4} \right)}.
\]

The restriction for the validity of the perturbative treatment is \( V_0L \gg V_{\text{max}}/V_0 > 1 \). These conditions are satisfied for realistic values of \( L, V_0 \), and \( V_{\text{max}} \). Indeed, from equation \( (3) \), and using known values for the conductivities of good conductors, we can fix \( V_{\text{max}} = 10^{16} \text{ m}^{-1} \). When the laser field is not applied we can set \( V_0 = 10^{10} - 10^{15} \text{ m}^{-1} \), the range of values for different semiconductors. For a cavity of size \( L_r \approx 10^{-2} \text{ m} \), and when the ratio between the maximum and minimum conductivities is in the range \( 10^3 \leq V_{\text{max}}/V_0 \leq 10^9 \), we obtain from equation \( (17) \) small values of \( \epsilon \), \( 10^{-5} \leq \epsilon \leq 10^{-2} \). It is worth noting that we are interested in low eigenfrequencies, for which \( k(t) \sim \mathcal{O}(L^{-1}) \). Nonetheless the perturbative treatment is also valid for \( k \sim \mathcal{O}(V) \).

In what follows we will only consider expressions to first order in \( \epsilon_n \). To analyze the possibility of parametric resonance we write the time-dependent frequency given in equation \( (16) \) as

\[
k_n(t) = k_n^0(1 + \epsilon_n(f - f_0)),
\]

where \( k_n^0 \equiv k_n^0(1 + \epsilon_n f_0) \) is a ‘renormalized’ frequency. The equation for the coefficients \( P^{(s)}_n(t) \) (equation \( (12) \)) can now be written to first order in \( \epsilon_n \) as

\[
P^{(s)}_n + k_n^2(t) P^{(s)}_n = -2\epsilon_n(k_n^0)^2(f - f_0) P^{(s)}_n
\]

\[
- \sum_m \left[ 2 P^{(s)}_m k_m(f - f_0) + P^{(s)}_m k_m S_m^{(A)} \right] S_m^{(B)} + \mathcal{O}^2.
\]

This equation describes a set of coupled harmonic oscillators with periodic frequencies and couplings. It is of the same form as the equations that describe the modes of a scalar field in a three-dimensional cavity with an oscillating boundary. A naive perturbative solution of previous equations in powers of \( \epsilon_n \) breaks down after a short amount of time (this happens for particular values of the external frequency such that there is a resonant coupling with eigenfrequencies of the cavity).

In order to analyze the long-time behavior of the solutions, the equations can be solved using multiple scale analysis (MSA) \[13\].

As shown in detail in \[7\], if the spectrum of the cavity is not equidistant, then the photons are created mainly in the resonant mode. This means that this particular mode decouples from the rest, and the subsequent quantum dynamics can be described starting from the equation

\[
P^{(s)}_n + (k_n^0)^2 P^{(s)}_n + 2\epsilon_n(k_n^0)^2(f - f_0)P^{(s)}_n = 0,
\]

that is, a harmonic oscillator with time-dependent frequency. When \( \Omega = 2k_n^0 \), the number of created photons grows exponentially due to parametric resonance.

3. Quantum open systems: the Langevin equation for the field

In the previous section, the starting point was the action given in equation \( (1) \), which describes a quantum field with
a time-dependent mass term concentrated on the position of the slab. In a more realistic model, the massless scalar field should be coupled to the microscopic degrees of freedom on the slab, and its dynamics will be described in terms of the effective action that results after integrating the microscopic degrees of freedom.

In static situations, it is enough to compute the in–out or the Euclidean effective action, and to read the zero-point energy of the system from the generating functional. However, if we are interested in the temporal evolution of the quantum field, it is necessary to compute the in–in or CTP effective action [14]. In this direction, we will consider the scalar field and for the infinite set of harmonic oscillators with time-dependent conductivity of the slab, which could be scalar field and the microscopic degrees of freedom on the slab. Therefore, we will consider a model composed of a classical action given by

\[
S[\phi, q_n] = S_0[\phi] + S_0[q_n] + S_{\text{in}}[\phi, q_n],
\]

where \(S_0[\phi]\) and \(S_0[q_n]\) are the free actions for the massless scalar field and for the infinite set of harmonic oscillators with which we model the microscopic degrees of freedom on the slab,

\[
S_0[\phi] = \frac{1}{2} \int d^2x \partial_\mu \phi \partial^\mu \phi,
\]

\[
S_0[q_n] = \frac{1}{2} \sum_n \int dt m_n \left( \dot{q}_n^2 - \omega_n^2 q_n^2 \right),
\]

and the interaction term is assumed as

\[
S_{\text{in}}[\phi, q_n] = - \sum_n \int d^2x \lambda(t)\phi(t, x)q_n(t)\delta \left( x - \frac{L}{2} \right),
\]

where \(\lambda(t)\) is the time-dependent coupling between the scalar field and the microscopic degrees of freedom on the slab. With this time-dependent coupling we are modeling the time-dependent conductivity of the slab, which could be produced by a time-dependent charge carrier density within the semiconducting slab.

Our system of interest is the massless scalar field. Therefore, we will proceed to integrating out the harmonic oscillator variables \(q_n\) in order to obtain an effective description of the evolution of the field. This non-unitary evolution is described, mainly, by two environmentally induced effects: dissipation and noise. After tracing over the environment oscillators, we are able to write an effective action for the quantum open system [15] of the form

\[
A[\phi, \phi'] = S_0[\phi] - S_0[\phi] + \delta A[\phi, \phi'],
\]

where \(\phi\) and \(\phi'\) are each of the CTP branches for the field. The influence (Feynmann–Vermor) action is given by

\[
\delta A[\phi, \phi'] = \int d^2x \int d^2x' \Delta(t, x)D(t'; t')
\]

\[
\times \delta \left( x - \frac{L}{2} \right) \delta \left( x' - \frac{L}{2} \right) \Sigma(t', x')
\]

\[
+ i \int d^2x \int d^2x' \Delta(t, x)N(t'; t')
\]

\[
\times \delta \left( x - \frac{L}{2} \right) \delta \left( x' - \frac{L}{2} \right) \Delta(t', x'),
\]

where we have defined \(\Delta(t, x) = \phi(t, x) - \phi'(t, x)\) and \(\Sigma(t, x) = 1/2[\phi(t, x) + \phi'(t, x)]\).

The influence action is in general non-local and complex, and contains all the information about the environment (i.e. temperature, spectral density, dissipation rates and noise properties). The kernel in the real part of \(\delta A\) is the dissipation kernel \(D(t'; t')\) and \(N(t'; t')\) in the imaginary part is related with the quantum fluctuations or noise coming from the external bath. These kernels are usually related by means of a fluctuation–dissipation relation. The time-dependent coupling strength is included in the definition of the noise and dissipation kernels in equation (26). On general grounds, one can expect that \(D(t'; t') = \lambda(t)\lambda(t')D(t - t')\) and \(N(t'; t') = \lambda(t)\lambda(t')N(t - t')\).

Different environments can be considered. Each of them is characterized by a spectral density. The ohmic environment is the most studied case in the literature and produces a dissipative force that is proportional to the velocity. The supraohmic case, on the one hand, is generally used to model the interaction between defects and phonons in metals and also to mimic the interaction between a charge and its own electromagnetic field [16]. On the other hand, the quantum behavior of free electrons in mesoscopic systems is affected by their interaction with the environment, which, for example in such cases, consist of other electrons, phonons, photons or scatterers. Which environment is more relevant for the dissipative phenomena generally depends on the temperature. For instance, the temperature dependence of the weak-localization correction to the conductivity reveals in metals that electron–electron interactions dominate over the phonon contribution to decoherence at the low temperature regime. Therefore, in order to mimic a real material for the slab in the cavity, it will be relevant to consider spectral densities such as the ones mentioned above. Only in the particular case in which the coupling is constant, \(\lambda(t) = \lambda_0\); and we consider an ohmic environment at very high temperature, the noise kernel is \(N(t - t') \sim \lambda_0^2\delta(t - t')\) and the dissipation kernel reads \(D(t - t') \sim \lambda_0^2\delta(t - t')\). Therefore, the influence action is local in time. It is worth to remark that, for these models, the fluctuation–dissipation relation reads

\[
N(t) = 2k_0 T \int_{-\infty}^{\infty} ds D(s),
\]

which is the classical Einstein formula.

In the general case, the equation of motion of the scalar field can be obtained from the effective action \(A[\phi, \phi']\) by \(\delta A[\phi, \phi'] = 0\):

\[
\square \phi + \phi \left( x - \frac{L}{2} \right) \int_0^\infty ds D(s; t) \phi \left( t, \frac{L}{2}, s \right) = 0.
\]

This equation corresponds to an average over all the possible noise realizations. In order to extract the noise information from the effective action, it is possible to write down the imaginary part of the influence action in terms of a stochastic noise force \(F\), which is coupled to the main system of interest, and it is defined by means of a Gaussian probability distribution \(P[F]\) like

\[
P[F] = N_F \exp \left\{ -\frac{1}{2} \int_0^\infty ds \int_0^\infty \text{d} s' F(s)x s' \right\}.
\]
where \( N_F \) is a normalization constant. Therefore, the imaginary part of \( \delta A \) can be written in terms of the stochastic force source as

\[
\int D\mathcal{F}(t)P[\mathcal{F}] \exp \left\{ i \frac{\hbar}{\Delta} \left( t, \frac{L}{2} \right) \mathcal{F}(t) \right\} = \exp \left\{ i \int_0^t ds \int_0^t ds' \Delta \left( s, \frac{L}{2} \right) N(s; s') \Delta \left( s', \frac{L}{2} \right) \right\}.
\]

Finally, from equation (30) we can derive an equation of motion for the massless scalar field that takes into account the noise. This is a Langevin-like equation, where the full dynamics of the field is also determined by the presence of the stochasticity induced by the environment via \( \mathcal{F} \). The Langevin equation now reads as

\[
\Box \phi + \delta \left( x - \frac{L}{2} \right) \int_0^t ds D(t; s) \phi \left( \frac{L}{2}, s \right) = \mathcal{F}(t) \delta \left( x - \frac{L}{2} \right).
\]

The stochastic noise \( \mathcal{F} \) is characterized by the probability distribution \( P[\mathcal{F}] \), and also by the correlation functions

\[
\langle \mathcal{F}(t) \rangle = 0; \quad \langle \mathcal{F}(t) \mathcal{F}(t') \rangle = N(t; t').
\]

4. The dynamics of the resonant mode

The aim of this section is to obtain a dynamical equation for the evolution of a single mode of the quantum field, starting from the effective equation derived in the previous section. We have already seen that, in the absence of dissipation, if the conductivity of the slab has periodic changes with frequency \( \Omega \), only the resonant mode with \( k^0 = \frac{\Omega}{2} \) is relevant at long times. We will now assume that this is true even in the presence of dissipation. Although reasonable, this is an unjustified assumption, and we hope to further analyze it in a forthcoming publication. Even with this simplification, we will see that the derivation of the dynamical equation of the mode is a non-trivial task.

Let us write the dissipation kernel as

\[
D(t; s) = V_0 \delta(t - s) + d(t; s).
\]

In the particular case \( d(t; s) = (V(t) - V_0) \delta(t - s) \), with \( V(t) \) being a periodic function of frequency \( \Omega \) we recover the model of section 2. In that case, we have shown that even when \( V(t) \) differs several orders of magnitude from \( V_0 \), the oscillations in the wavenumber have a relative amplitude \( \epsilon_n \ll 1 \) (see the discussion after equation (17)). Therefore, we will assume that the main contribution to the wavenumber comes from the term proportional to \( V_0 \), and we will treat the contributions of both \( d(t; s) \) and the noise kernel as small corrections. It is worth emphasizing that we are not assuming that \( D(t, s) \approx V_0 \delta(t - s) \), but that the effect of \( d(t; s) \) and the noise on the instantaneous wavenumber \( k(t) \) can be treated perturbatively.

The resonant mode of the field can be written as

\[
u(x, t) = P(t) \Psi(x, t),
\]

where \( \Psi(x, t) \) is given in equation (6). In order to simplify the notation, from now on we omit the subindex \( m \), since we are considering a single mode. Inserting this particular mode into the Langevin equation for the field, we find that the boundary condition at the position of the slab becomes

\[
2P(t)k(t) \cos \left[ \frac{(k(t)L)}{2} \right] = -\int_0^t ds D(t, s)P(s) \sin \left[ \frac{(k(t)L)}{2} \right] + \sqrt{\frac{L}{2}} \mathcal{F}(t).
\]

It is worth stressing that, when the dissipation kernel is local and in the absence of noise, the amplitude of the mode \( P(t) \) factorizes and the boundary condition fixes the value of the time-dependent wavenumber \( k(t) \). In the general case, \( k(t) \) becomes a non-local function of \( P(t) \). To see this explicitly, we solve equation (34) assuming that \( d(t; s) \) and \( \mathcal{F}(t) \) induce small time-dependent corrections to \( k(t) \). We write \( k(t) = k^0 + \Delta k(t) \), where \( k^0 \) is the solution of the unperturbed (static) problem, i.e.

\[
2k^0 = -V_0 \tan \left[ \frac{k^0 L}{2} \right].
\]

To first order in \( d(t; s) \) and \( \mathcal{F} \) we have

\[
\Delta k(t) P(t) = \frac{k^0}{V_0} \left[ 1 + \frac{V_0}{4} + \frac{(k^0)^2 L}{V_0} \right] \times \left[ \int_0^t ds d(t; s)P(s) - \tilde{\mathcal{F}}(t) \right],
\]

where \( \tilde{\mathcal{F}} = \sqrt{L/2} \csc(k^0 L/2) \). As anticipated, the non-local boundary condition becomes a non-local relation between \( \Delta k(t) \) and \( P(t) \).

The next step is to insert the field mode into the Klein–Gordon equation of the field, at both sides of the slab. Of course the single mode will not be an exact solution of the Klein–Gordon equation, which would involve the coupling to the other modes. Therefore, in order to obtain the dynamical equation for \( P(t) \), we impose \( (\Psi(x, t), \Box \nu) = 0 \). After some simple calculations we obtain,

\[
P + (g^{(A)} \Delta k(t) P) + g^{(A)} (2 P \Delta k + P \Delta k) = 0,
\]

where \( g^{(A)} \) corresponding to \( g^{(A)} \) is defined in section 2, and can be evaluated to lowest order, setting \( d(t; s) = 0 \) and \( \tilde{\mathcal{F}} = 0 \). Note that in this approximation this coupling constant is independent of time.

Let us analyze the resulting equation for the field mode. To begin with, it is important to note that the term proportional to \( \Delta k(t) P(t) \) contains non-local dissipation and additive noise. Moreover, if we write \( d(t; s) = (V(t) - V_0) \delta(t - s) + d(t; s) \) the local term corresponds to a time-dependent frequency. Depending on the details of the environment, we expect that, at high temperatures, \( d(t; s) \sim \delta(t - s) \), so the usual dissipative term proportional to \( \dot{P} \) is recovered. Moreover, in this limit the fluctuation–dissipation theorem implies a white noise. In any other case, we expect non-local dissipation and colored noise.

In the absence of dissipation, the last term in the lhs of equation (37), proportional to \( g^{(A)} \), can be neglected when the external frequency satisfies \( \Omega = 2k^0 \). However, in general this
is not the case. Noting that $2\tilde{P}\Delta k + P\tilde{\Delta}k = (P\tilde{\Delta}k) - \tilde{P}\Delta k$, it is rather difficult to find an expression for this additional term by taking two derivatives of equation (36). The result is a rather complicated expression:

$$2\tilde{P}\Delta k + P\tilde{\Delta}k = -\tilde{P}\Delta k + \frac{k^0}{V_0 \left[ 1 + \frac{V_0 L}{4} + \frac{(k_0)^2 L}{V_0} \right]}
\times \left[ P(d(t; t) + \partial_t d(t; t))
+ d(t; t)\tilde{P} + \int_0^t \mathrm{d}s \, \partial_s^2 d(t; s) P(s) - \tilde{F}(t) \right].$$

The main consequence is that, after inserting this result into equation (37), the final equation for $P(t)$ can be written as

$$\tilde{P} + \omega_{\text{eff}}^2(t) P + \int_0^t \mathrm{d}s \, d_{\text{eff}}(t, s) P(s) = \mathcal{F}_{\text{eff}}(t).$$

In summary, we have shown that the amplitude of the mode satisfies a Langevin equation with non-local dissipation and colored noise. The dissipation ($d_{\text{eff}}$) and noise ($\mathcal{F}_{\text{eff}}$) kernels are related to the ones in the Langevin equation for the field derived in section 3, although they are not exactly the same because of the presence of the term proportional to $g^{(A)}$. All the information about dissipation and noise comes from the non-local boundary condition that the field satisfies on the slab.

5. Conclusions

In this paper, we have argued that the natural approach to analyze the DCE is that of quantum open systems. The coupling between the degrees of freedom of the vacuum field and those of the (imperfect) mirrors generates an influence functional for the vacuum field that contains the information about dissipation and noise. The dynamics of the field is described by a Langevin equation with dissipative and noise kernels concentrated on the position of the mirror.

When the system is under the influence of periodic, time-dependent external conditions, it is possible to have parametric amplification in some of the modes of the field. If the spectrum of the unperturbed system is not equidistant, one expects the amplification to occur for a single mode. Assuming that this is the case even in the presence of dissipation, we have shown that the dynamics of the mode can also be described by a Langevin equation with non-local dissipation and colored noise. This equation is similar to the one derived in the context of quantum Brownian motion, but now the frequency and the (non-local) kernels have a periodic time dependence.

We have considered a number of simplifications in order to illustrate the above points: we worked with a scalar field in $1 + 1$ dimensions, and we assumed that the parametric amplification can be described in terms of a single mode. The last assumption deserves further investigation. Moreover, we have not attempted to compute the influence of dissipation and noise on the number of photons created. This seems to be a rather difficult task, because of the complexity of the Langevin equation for the mode. Some additional simplifications could be necessary in order to estimate these effects (along the lines of [11], for instance). However, we think that we have clarified, from a conceptual point of view, the origin of the description of the DCE in terms of a noisy and damped oscillator.

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