Containing intense laser light in circular cavity with magnetic trap door

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It is shown by particle-in-cell simulation that intense circularly polarized (CP) laser light can be contained in the cavity of a solid-density circular Al-plasma shell for hundreds of light-wave periods before it is dissipated by laser-plasma interaction. A right-hand CP laser pulse can propagate almost without reflection into the cavity through a highly magnetized overdense H-plasma slab filling the entrance hole. The entrapped laser light is then multiply reflected at the inner surfaces of the slab and shell plasmas, gradually losing energy to the latter. Compared to that of the incident laser, the frequency is only slightly broadened and the wave vector slightly modified by appearance of weak nearly isotropic and homogeneous higher harmonics.

Recently, intense laser interaction with plasma has been widely studied in connection with applications such as inertial confinement fusion and laser-driven particle acceleration. Particle-in-cell (PIC) simulations and theoretical studies have shown that isolated intense light pulses on the scale of the laser wavelength can be self-consistently trapped in plasmas as solitons. Trapping of electromagnetic (EM) waves in plasma can also be from excitation of surface plasma waves, beam-plasma and/or parametric instabilities. Recently, it is found that an intense laser pulse that has entered into a hollow shell through a hole can also be trapped when a near-critical density plasma layer is applied or self-generated to seal the entrance hole. In all these cases, there is much loss of the laser energy to the plasma. On the other hand, there has also been considerable research on trapping and localizing light in atomic media or photonic structures. However, in these works no intense laser-plasma interaction is involved because of the relatively weak laser intensities used.

Intense magnetic fields of $10^3 - 10^5$ T and above are ubiquitous in the cosmic environment. Strong magnetic fields can also be self-generated during ultraintense-laser interaction with matter by compression of seeded magnetic fields, baroclinic effect, high-current electron beams, etc. In the presence of kilotesla magnetic fields, the electron gyrofrequency can be comparable to or even larger than the laser frequency. In this case, a right-hand circularly polarized (RHCP) laser can propagate deep into an overdense plasma along an intense embedded magnetic field without encountering cutoff or resonance. Moreover, transmission of RHCP laser into dense plasma can be controlled by an intense magnetic pulse in the plasma.

In this Letter, we propose a scheme for trapping intense light in a hollow shell of high density plasma by sending a RHCP laser light pulse through a strongly magnetized slab that serves as a transparent trap door. The laser pulse can propagate through the slab with almost no reflection. Once the entire laser pulse (which can be suitably long) is inside the shell cavity, the magnetic field in the slab is turned off, so that the light becomes contained. It is then multiply reflected by the unmagnetized shell plasma or propagates along its surface. As a result, the light can survive for hundreds of light-wave periods until all its energy is absorbed by the slab and shell plasma.

The optical properties of a linear medium can be described by its refractive index $N = c k_L / \omega_L = \epsilon^{1/2}$, where $k_L = 2 \pi / \lambda_L$, $\omega_L$, and $c$ are the wave number, frequency, and vacuum speed of the light and $\epsilon$ is the dielectric constant of the medium. For simplicity, in the following we shall normalize the plasma density $n_e$ by the critical density $n_c = m_e \omega_L^2 / 4 \pi e^2$, where $e$ and $m_e$ are the electron charge and mass, respectively. Accordingly, for unmagnetized plasma we have $\epsilon = 1 - \omega_p^2 / \omega_L^2 = 1 - n_e$. For magnetized plasma with the external magnetic field $B_0$ (normalized by $m_e c \omega_L / e$) along the laser propagation direction, the dielectric constant can be written as $\epsilon_B = 1 - n_e / (1 \pm B_0)$, where the plus and minus signs correspond to the so-called $L$ and $R$ electromagnetic waves, respectively. Thus, for the $R$ wave $\epsilon_R$ is always larger than unity if $B_0 > 1$, and the plasma is transparent at any density. In contrast, the $L$ wave has a cutoff density at $n_L = 1 + B_0$, where it will be reflected.

The proposed trapping scheme is illustrated in Fig. 1(a). The circular plasma shell has on its left wall a small section replaced by a highly-magnetized overdense flat slab that serves as a trap door for the incident laser pulse. In order to see the function of the slab, it is instruc-
tive to first consider the one-dimensional linear theory of light transmission through a thin highly-magnetized high-density plasma layer, as shown in Fig. 1(b). The total wave electric field $E_y(x)$ in each of the three regions can be written as $28, 29^{28, 29}$ $E_1 \exp(-ik_L x) + E_1 \exp(i k_L x)$ for $x < 0$, $E_2 \exp(-ik_s x) + E_3 \exp(i k_s x)$ for $0 < x < d$, and $E_4 \exp(-ik_s (x - d))$ for $x > d$, where $E_1$, $E_2$, $E_3$, and $E_4$ are the amplitudes of the laser light incident on and reflected from the front surface of the slab, transmitted into the slab, reflected from the back surface of the slab, and transmitted into the vacuum, respectively, $k_s = \frac{1}{2} k_L$ is the wave number in the slab. Applying the boundary conditions that at each of the boundaries (the total) $E_y$ and $\partial_x E_y$ should be continuous, one obtains after some algebra

$$
E_1 = \frac{-i2RE_0 \sin(k_s d)}{R^2 e^{ik_s d} - e^{-ik_s d}},
$$

$$
E_2 = \frac{2RE_0}{R(N_s + 1) - (N_s - 1) e^{-2ik_s d}},
$$

$$
E_3 = \frac{2E_0}{R(N_s + 1) e^{2ik_s d} - (N_s - 1)},
$$

$$
E_4 = \frac{4N_s E_0}{(N_s + 1) e^{ik_s d} - (N_s - 1) e^{-ik_s d}},
$$

where $N_s = \frac{1}{2}$ and $R = (N_s + 1)/(N_s - 1)$. Additional internal reflections are negligible. Fig. 1(c) shows the profile of the electric field obtained from Eq. (1) for $n_e = 20 n_c$, $d = 0.2 \mu m$, $B_0 = 5$, and $E_1 = 0.27$ (the same as the corresponding parameters used in the PIC simulation below). It can be seen that the reflected light ($E_1 = 0.017$) is only 6% of that ($E_i = 0.27$) of the incident, and the transmitted light ($E_4 = 0.269$) is very close to the latter.

The relativistic 2D3V PIC simulation code EPOCH is used to investigate our scheme. The shell consists of Al plasma (ion mass $m_{Al} = 26.98m_p$, where $m_p = 1836m_e$ is the proton mass) with charge $Ze = 10e$ and density $50n_c$. Its inner radius and thickness are $14.8 \mu m$ and $0.2 \mu m$, respectively. The center of the shell cavity is located at $(x, y) = (20, 20) \mu m$, so that the left front of the slab is at $x = 5.86 \mu m$. An external magnetic field of strength $B_0 = 5$ (corresponding to $5.36 \times 10^4 T$ in dimensional units) along the laser axis is embedded in the slab plasma, which is shut down after the laser pulse injection is complete. To account for laser-plasma interaction induced plasma expansion, the extent of $B_0$ is slightly wider than the slab. The initial temperature of the slab and shell is $100 \text{ eV}$. The simulation box is $40 \mu m \times 40 \mu m$ with $4000 \times 4000$ cells. Each cell contains 50 macro particles per species. At $t = 0$, a RHCP laser pulse of wavelength $\lambda = 1 \mu m$ and intensity $I_0 = 2 \times 10^{17} W/cm^2$ (or laser parameter $a_L = eE_L/m_e c \omega L = 0.27$, where $E_L$ is the peak laser electric field) enters from the left side ($x = 0$) of the simulation box and is focused on of the slab at $x = 5.86 \mu m$. Its front rises with a Gaussian profile for $17 \text{ fs}$ to the peak intensity and remains constant for $150 \text{ fs}$. The transverse profile of the laser pulse is Gaussian, with spot radius $2 \mu m$.

Figure 2 shows the electric field $E_y$, magnetic field $B_z$, and axial Poynting vector $S_z$ at $t = 100 \text{ fs}$. We see that the laser pulse can easily enter the overdense slab into the cavity. It propagates through the slab almost without reflection, and the amplitude of the transmitted laser in the vacuum is very close to that of the incident laser, in good agreement with the theoretical analysis (see Fig. 1(c)). One can see from Figs. 2 for $t = 100 \text{ fs}$ that there is little energy loss by the laser as it passed the slab. The axial Poynting flux is everywhere positive except at the shell-slab boundary, where some light is scattered.

Reflections from the cavity wall and wave interference lead to restructuring of the EM fields in the cavity as well as loss of wave energy to the slab and shell plasmas. Figure 3 for $t = 350 \text{ fs}$ shows the transverse electric field $E_{y'}$, magnetic field $B_z'$, and EM field energy density $W = \frac{1}{2}(\varepsilon_0 |E|^2 + \mu_0^{-1} |B|^2)$, where $\varepsilon_0$ and $\mu_0$ are the free-space permittivity and permeability, respectively. We see that light is trapped in the cavity and its structure is rather complex. Because of constructive interference of the reflecting light in the cavity, locally the magnitudes of the wave electric and magnetic fields can be up to 1.61 and 2.35, respectively, times that of the incident...
laser. Figure 3(d) for the evolution of the energies of the trapped light and the plasmas shows that the former increases with time until the laser pulse has completely entered the cavity (at t ~ 190 fs, recall that the slab is 5.86 µm away from the left side of the simulation box). Then it gradually decreases. One can also see that the energy of the plasmas first increases rapidly, mainly due to light absorption by the slab electrons, which is heated and expands, as shall be discussed later. Figure 3(d) also shows that after the laser pulse has completely entered the cavity at t ~ 190 fs, energy absorption by the plasma continues at a slower rate, and the total energy of the light in the cavity and the plasma in the slab and shell is well conserved.

On the other hand, Fig. 3(d) also shows that even though transfer of light energy to the slab and shell plasmas starts as soon as the laser enters the slab, even at t = 500 fs the trapped light still retains about 78.6% of its maximum energy, namely at that t ~ 190 fs. We have not attempted to optimize the light trapping, which has to be done by trial and error adjustment of the initial parameters. For example, simulations indicate that the energy of the trapped light can increase with the cavity size. It should also be possible to improve the design of the slab so that absorption of reflected light by its cavity facing side is minimized.

Figure 4 for the plasma density and light Poynting flux distributions shows that the slab-shell boundaries are modified and strong local electric and magnetic fields are generated around them. One can also see in Fig. 4(a) for t = 250 fs that, as expected, at this instant the flow of light energy is still mainly around the ±x directions. Figure 4(b) for t = 350 fs shows that at this later time interaction of the trapped light with plasma takes places mainly in three regions: at the slab and two regions on the shell wall. That is, the Poynting flux is at first localized along the axis region, but then spreads to other regions at latter times. Focusing and defocusing of the reflected light on the cavity axis can be observed at x ~ 28 µm and x ~ 17 µm. The 20n_c highly magnetized H slab plasma at the laser entrance is strongly heated and expands, but the very high charge density Al shell plasma remains almost unaffected. The reason is that for the level of light intensity here, energy absorption by the shell plasma is mainly due to vacuum heating [31, 32], whose efficiency depends on the skin depth, which for the 50n_c Z = 10 Al shell plasma is 5 times smaller than that of the 20n_c H slab plasma, where efficient resonant absorption dominates, especially after the expansion [32, 33]. At still later times the interaction regions on the right inner shell wall are wider spread, but the structure and behavior of the trapped light in the cavity remain similar. Due to continuous energy transfer to the slab and shell plasma, after the laser pulse has fully entered the cavity at t ~ 190 fs, the total light/plasma energy decreases/increases almost linearly with time, but as mentioned, the total energy remains constant. However, no stationary or quasistationary state was found even at t ≥ 500 fs, when the light field has become weak.

Figure 5 shows the spatial EM field spectra E_y(k) and B_z(k) at two instants after the laser pulse has fully entered the cavity. One can see that even though the spatial wave structure in the real space is rather complex, its Fourier space is relatively simple. As mentioned, initially the ±k_x components from the just-transmitted laser light dominates (not shown), but k_y components also appear as reflections take place since the laser is of finite spot-size and the cavity wall is curved. Figures 5(a)–(c) show that except around θ ~ π/2 and 3π/2 (apparently due to less light reflections occurring around these angles in the physical space), the |k| = 1 mode (small dark partial circles in the figures) associated with that (k = i_e x) of the input laser remains dominant at all times, which can also be roughly seen from the angles of the Poynting vectors in Fig. 4. In fact, close examination shows that the |k| < 1 region is more highly populated around the ±x directions, as to be expected. Weak but distinguishable spatial harmonics |k| = 2, 3, 4, ... can also be observed. The spectra continue to evolve with time as energy loss to the slab and shell plasma continues, but their overall profiles remain almost unchanged. In particular, they remain peaked at |k| = 1 (except around θ ~ π/2 and 3π/2) and k = 0. Figures 5(e) and (f) show the frequency spectra E_y(ω) and B_z(ω) at the center of the
cavity. At other locations in the cavity they (not shown) are very similar. We see that the frequency spectrum is highly peaked at the incident laser frequency $\omega_L$, which together with the the dominance of the wavelength near that of the incident laser, confirms that the light of the latter is trapped without altering its basic characteristics. The slight frequency broadening can be attributed to the presence in the cavity of low-density electrons driven out by the laser-wall plasma interaction, so that the vacuum light-wave dispersion relation is slightly modified by inclusion of the plasma frequency $\omega_p$ ($< k_L c$). However, as already pointed out, we were unable to identify the EM field or Poynting vector structures with that of an eigenmode solution of the vacuum wave equation in circular geometry \[34\], even though the trapped light has quite well defined frequency and wavelength, as well as $E_y(\omega, k)$ and $B_z(\omega, k)$ relationship. The reason could be that there is insufficient time to form an eigenmode, the cavity is not really circularly symmetric because of the slab, and/or the wave-plasma interaction at the slab and cavity wall cannot be ignored. Furthermore, one can see in Fig. 3(a)-(d) that the region near $k = 0$ is also enhanced, which can be attributed to the multiple light reflections at the slab and shell plasma boundaries, where $k$ changes sign. Finally, we note that weak oscillations (not shown) are also detected in the electron density, which can be attributed to excitation of standing electron plasma oscillations by the laser-plasma interaction.

In conclusion, we have considered a scheme for containing intense light in the cavity of a solid-density Al plasma shell. PIC simulations demonstrate that the laser light can enter the cavity with the help of an axial magnetic field embedded in an overdense H slab covering a hole in the shell. Inside the cavity the light wave is reflected as well as partially absorbed by the slab and shell plasma. Because of the multiple reflections, the wave frequency is slightly broadened but remains close to that of the incident laser. The wave vector, originally only along the laser axis, also acquires new directions and weak spatial harmonics are generated, which then become azimuthally isotropic except around $\theta \sim \pi/2$ and $3\pi/2$. However, the wavelength of the incident laser still remains the dominating spatial scale. Because of energy transfer to the slab and shell, the trapped light continues to decay until it vanishes. The results here may be useful for understanding storage of intense light, as well as interpretation of phenomenon associated with highly localized light \[35\] \[36\].

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FIG. 5. Fast Fourier transform of $E_y$ (top row) and $B_z$ (bottom row) at $t = 300$ fs (left column) and 400 fs (right column). Frequency spectra $|E_y(\omega)|$ (c) and $|B_z(\omega)|$ (d) (arbitrary units) at the center ($x = 20\, \mu m$, $y = 20\, \mu m$) of the cavity.

[1] J. Lindl, O. Landen, J. Edwards, E. Moses, and NIC Team, Phys. Plasmas 21, 020501 (2014).
[2] A. Macchi, M. Borghesi, and M. Passoni, Rev. Mod. Phys. 85, 751 (2013).
[3] X. H. Yang, W. Yu, H. Xu, H. B. Zhuo, Y. Y. Ma, D. B. Zou, T. P. Yu, Z. Y. Ge, Y. Yin, F. Q. Shao, and M. Borghesi, Phys. Plasmas 21, 063105 (2014).
[4] X. H. Yang, H. B. Zhuo, Y. Y. Ma, H. Xu, T. P. Yu, D. B. Zou, Z. Y. Ge, B. B. Xu, Q. J. Zhu, F. Q. Shao, and M. Borghesi, Plasma Phys. Control. Fusion 57, 025011 (2015).
[5] N. M. Naumova, S. V. Bulanov, T. Zh. Esirkepov, D. Farina, K. Nishihara, F. Pegoraro, H. Ruhl, and A. S. Sakharov, Phys. Rev. Lett. 87, 185004 (2001).
[6] M. Y. Yu, P. K. Shukla, and K. H. Spatschek, Phys. Rev. A 18, 1591 (1978).
[7] L. Stenflo, Phys. Scr. T63, 59 (1996).
[8] M. I. Bakunov and A. V. Maslov, Phys. Rev. Letts. 79, 4558 (2001).
[9] N. Kumar and V. K. Tripathi, Phys. Plasmas 14 103108 (2007).
[10] P. Kaw, G. Schmidt, and T. Wilcox, Phys. Fluids 16, 1522 (1978).
[11] V. P. Silin and A. N. Starodub, Zh. Eksp. Teor. Fiz. 73, 1379 (1977).
[12] S. X. Luan, W. Yu, M. Y. Yu, S. M. Weng, J. W. Wang, H. Xu, H. B. Zhuo, and A. Y. Wong, Phys. Plasmas 22, 093110 (2015).
[13] M. Störzer, P. Gross, C. M. Aegerter, and G. Maret, Phys. Rev. Lett. 96, 063904 (2006).
[14] L. V. Hau, S. E. Harris, and C. H. Behroozi, Nature 397, 594 (1999).
[15] M. Fleischhauer and M. D. Lukin, Phys. Rev. Lett. 84, 5094 (2000).
[16] T. Tanabe, M. Notomi, E. Kuramochi, A. Shinya, and H. Taniyama, Nat. Photon. 1, 49 (2007).
[17] P. N. Melentiev, A. E. Afanasiev, A. A. Kuzin, A. S. Baturin, and V. I. Balykin, Opt. Lett. 38, 2274 (2013).
[18] V. Canuto and H. Y. Chiu, Space Sci. Rev. 12, 3 (1971).
[19] J. P. Knauer, O. V. Gotchev, P. Y. Chang, D. D. Meyerhofer, O. Polomarov, R. Betti, J. A. Frejje, C. K. Li, M. J. E. Manuel, R. D. Petrasso, J. R. Ryyg, and F. H. Seguin, Phys. Plasmas 17, 056318 (2010).
[20] P. Y. Chang, G. Filkse, M. Hohenberger, J. P. Knauer, R. Betti, F. J. Marshall, D. D. Meyerhofer, F. H. Seguin, and R. D. Petrasso, Phys. Rev. Lett. 107, 035006 (2011).
[21] H. Yoneida, T. Namiki, A. Nishida, R. Kodama, Y. Sakawa, Y. Kuramitsu, T. Morita, K. Nishio, and Takao Ide, Phys. Rev. Lett. 109, 125004 (2012).
[22] U. Wagner, M. Tatarakis, A. Gopal, F. N. Beg, E. L. Clark, A. E. Dangor, R. G. Evans, M. G. Haines, S. P. D. Mangles, P. A. Norreys, M. S. Wei, M. Zepf, and K. Krushelnick, Phys. Rev. E 70, 026401 (2004).
[23] S. Fujioka, Z. Zhang, H. Ishihara, K. Shigemori, Y. Hironaka, T. Johzaki, A. Sunahara, N. Yamamoto, H. Nakashima, T. Watanabe, H. Shiraga, H. Nishinuma, and H. Azechi, Sci. Rep. 3, 1170 (2013).
[24] Ph. Korneev, E. d’Humières, and V. Tikhonchuk, Phys. Rev. E 91, 043107 (2015).
[25] F. F. Chen, Introduction to Plasma Physics and Controlled Fusion, 2nd Ed., Springer, New York, 2006.
[26] X. H. Yang, W. Yu, H. Xu, M. Y. Yu, Z. Y. Ge, B. B. Xu, H. B. Zhuo, Y. Y. Ma, F. Q. Shao, and M. Borghesi, Appl. Phys. Lett. 106, 224103 (2015).
[27] G. J. Ma, W. Yu, M. Y. Yu, S. X. Luan, and D. Wu, Phys. Rev. E 93, 053209 (2016).
[28] J. A. Kong, Electromagnetic Wave Theory, Wiley, New York, Wiley, 1990.
[29] R. W. Scharstein, IEEE Trans. Education 35, 170 (1992).
[30] T. D. Arber, K. Bennett, C. S. Brady, A. Lawrence-Douglas, M. G. Ramsay, N. J. Sircombe, P. Gillies, R. G. Evans, H. Schmitz, A. R. Bell and C. P. Ridgers, Plasma Phys. Control. Fusion 57, 113001 (2015).
[31] F. Brunel, Phys. Rev. Lett. 59, 52 (1987).
[32] P. Gibbon and E. Förster, Plasma Phys. Control. Fusion 38, 769 (1996).
[33] J. Friedberg, R. W. Mitchell, R. L. Morse, and L. I. Rudsinski, Phys. Lett. 28, 795 (1972).
[34] M. J. Padgett and L. Allen, Optics Comm. 121, 36 (1995).
[35] M. Stenhoff, Ball Lightning: Unsolved Problem in Atmospheric Physics, Kluwer, New York, 1999.
[36] H. C. Wu, Sci. Rep. 6, 28263 (2016).