Calibrating the photon detection efficiency in IceCube

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The IceCube neutrino observatory is composed of more than five thousand light sensors, Digital Optical Modules (DOMs), installed on the surface and at depths between 1450 and 2450 m in clear ice at the South Pole. Each DOM incorporates a 10-inch diameter photomultiplier tube (PMT) intended to detect light emitted when high energy neutrinos interact with atoms in the ice. Depending on the energy of the neutrino and the distance from secondary particle tracks, PMTs can be hit by up to several thousand photons within a few hundred nanoseconds. The number of photons per PMT and their time distribution is used to reject background events and to determine the energy and direction of each neutrino. The detector energy scale was established from previous lab measurements of DOM optical sensitivity, then refined based on observed light yield from stopping muons and calibration of ice properties. A laboratory setup has now been developed to more precisely measure the DOM optical sensitivity as a function of angle and wavelength. DOMs are calibrated in water using a broad beam of light whose intensity is measured with a NIST calibrated photodiode. This study will refine the current knowledge of the IceCube response and lay a foundation for future precision upgrades to the detector.
1. Introduction

IceCube is a neutrino telescope installed in the South Pole ice. As pictured in Figure 1, the detector is made of 86 strings, mostly spaced 125 m apart. Each string has 60 Digital Optical Modules (DOMs) deployed at depths between 1450 and 2450 m, for a total instrumented volume of about 1 km$^3$. A DOM consists of a pressure resistant glass sphere containing a 10-inch diameter R7081-02 photomultiplier tube (PMT) made by Hamamatsu Photonics, and a custom-design single board data acquisition computer. Neutrino interactions in ice produce secondary charged particles which are sufficiently energetic to generate Cherenkov light that reaches the DOMs. Some of the photons striking the PMTs’ photocathode sensitive area are detected and amplified, resulting in waveforms that record the amount and timing of arriving light. The DOMs digitize the PMT waveforms and send them to a counting house at the surface above the center of the array, where waveform information from multiple strings is combined for each interaction into an "event".

Every recorded event is reconstructed to establish the direction of the neutrino and deposited energy. The understanding of the energy scale in IceCube is a very important issue which involves the sensitivity of the PMTs and the properties of the ice, known to be inhomogeneous in the detector volume [1]. Currently the energy scale is based on different sets of measurements performed both in laboratory and in-situ (explained in detail respectively in [2] and in [3]).

In this paper we describe a new laboratory setup designed and built at the University of Wisconsin, Madison$^1$, in order to perform the absolute calibration of light sensors.

2. Laboratory setup

The lab setup is designed to measure the single photon detection efficiency of a DOM when illuminated in a similar way as from neutrino events in IceCube, where light typically travels 5-200 m from its source before possible detection. Because the DOM diameter is only 13”, at such distances it can be accurately described by its efficiency to detect photons arriving from a particular angle, averaging over all possible points of arrival at the photocathode. Such an average is implicit in our measurements which use a uniform light beam from a source 6.4 m away.

In IceCube, each DOM has its PMT facing downward, and the sensitivity does not vary significantly with rotation around this vertical axis. The sensitivity does depend on the polar angle of illumination relative to the DOM axis, so the lab setup includes a motorized mounting shaft for

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inclining the DOM axis relative to a fixed vertical beam. The DOM is immersed in water in order to closely model reflection and refraction effects occurring at the optical interface where the DOM would be embedded in ice. Effects of PMT gain, discriminator threshold and electronics calibration are accounted for by using the same software and procedures as in standard IceCube operations.

The illumination system and water tank are sketched in Figure 2. There are several choices of light source, any of which can be directed to shine into a small diffuser box. The diffuser box acts as an integrating sphere, coated inside with white barium sulfate paint and with an optical baffle on the entrance port. The diffuser’s main output port is aimed horizontally towards a mirror 5.3 m away, which reflects the beam downward 1.1 m towards the DOM. Photodiode PD1 is mounted on a secondary port of the source diffuser and is used to precisely measure changes in beam intensity, including programmed brightness changes of order $10^5$. The beam geometry is defined by a series of apertures in the tunnel, creating a uniform circular disc of light with a 16” diameter at the DOM position. Surfaces in the tunnel and around the DOM are blackened to control stray reflected light.

Currently installed light sources include a pulsed diode laser (405 nm), pulsed LEDs (400 nm), continuous LEDs (370 nm, 400 nm, 450 nm), and lamp with monochromator (320-700 nm). Beam intensity entering the diffuser box can be varied over a wide range by means of LED current, repetition rate of pulsed sources, neutral density optical filters, and a simple shutter.

Photodiode PD3 is in a waterproof housing and mounted on a robotic arm, so it can be moved into the water above the DOM. When sources are operated in a bright mode, PD3 is used to precisely measure the light reaching the DOM, as well as to test uniformity and boundaries of the beam cross section. When sources are switched to their dim mode, PD3 is moved out of the way and PD1 measures the intensity ratio relative to the bright mode (section 3).

Closer to the diffuser output, a calibration transfer station allows temporary mounting of PD3 side by side with a NIST calibrated photodiode, which has a response known to better than 0.5% when illuminated near its center at wavelengths 300 – 700nm. For this cross calibration of PD3, an aperture is inserted in front of the transfer station, reducing the beam cross section to a small spot that can be positioned on center or scanned across the two photodiodes. PD1 is used to correct for small variations of source intensity during the scan. The calibrated response of PD3 averaged over
its active area is then converted to a value appropriate to a uniform beam.

An optical fiber pickup can also be moved into the beam at the calibration transfer station, with the fiber routed to a separate monochromator and PMT. This allows measurement of the wavelength spectrum of the beam for any source configuration.

Photodiode currents are measured with custom preamplifier and ADC boards. Figure 3 shows the PD3 preamplifier, which is built on a 1"-diameter circuit board and mounted directly on the PD3 connector before encapsulation in a waterproof housing. The photodiode current is amplified with gain $1 \mu V/fA$ and sent to an ADC board in the light sources box via a shielded twisted pair cable, along with another copy of the signal amplified $\times 101$. Low noise and offset of the AD8627 operational amplifier, together with the high shunt impedance of the photodiode, allows measurement of currents around $100fA$ with good precision when an on-off technique is employed for subtracting baseline offsets. The 16-bit ADCs (ADS1110) are configured with least significant bit $8–63\mu V$. A similar circuit is used for PD1 and the NIST photodiode (Hamamatsu S2281), except that relays are provided to switch the first stage between gains in the range $10\mu V/nA$ to $1\mu V/fA$.

The tank water is continuously pumped through an external chiller and water purification system. Maintaining the temperature around $5^\circ C$ avoids introducing high humidity to the tunnel and source region, where it might affect electronic or optical elements. The low temperature also reduces the DOM dark noise rate and discourages bacterial growth in the water. Wetted mechanical elements are made of stainless steel, glass, anodized aluminum, or plastics compatible with high purity deionized water.

Surrounding the water tank, Helmholtz coils are arranged for separate control of the magnetic field in each direction. These coils cancel the ambient field in the lab and create a field relative to the DOM axis (in its rotated position) which matches that experienced at the South Pole.

A blackout curtain hangs from the open end of the tunnel and is secured around the water tank by a belt. For replacing the DOM, the curtain is raised and the tunnel assembly is wheeled aside.

All system elements are controlled and monitored by Python scripts running on a standard IceCube data acquisition computer (DOMHub), which also embodies the interface for control and readout of the DOM[4]. A separate DOM-style data acquisition board is connected to the DOMHub for recording electronic synchronization pulses when using the laser or pulsed LEDs, allowing us to reject out-of-time DOM hits in data analysis.
Table 1: Approximate response of PD1, PD3 and DOM for bright and dim beams.

| Source Setting | Photon flux | PD1    | PD3    | DOM               |
|----------------|-------------|--------|--------|-------------------|
| Bright         | 10⁶/cm²/sec | 50 nA | 100 fA | saturation        |
| Dim            | 10/cm²/sec  | 500 fA | too low| 100-1000 Hz       |

3. Measurements

Our fundamental reference for measuring light levels is a NIST calibrated photodiode [5]. As explained in section 2, we use this precision photodiode to calibrate PD3 which has been encapsulated for underwater use. We then install PD3 in the water tank and turn on one of the light sources, resulting in a PD3 current around 100 fA, measured to within 0.5%. This gives a direct measurement of beam flux in the tank, typically 10⁶ photons/cm²/sec at high brightness settings.

However, the PD3 current signal cannot be precisely measured at the much lower brightness settings needed for DOM sensitivity calibration. For beam fluxes below 100 photons/cm²/sec, we instead rely on PD1 which is much closer to the source. The larger PD1 signal is also proportional to the beam flux, but with an unknown scale factor that must be calibrated against the direct PD3 measurement. For this purpose we perform a bright PD3 measurement simultaneously with a PD1 measurement, yielding the beam flux scale factor:

\[
\text{Flux} = 0.0226 \text{ photons/cm}^2/\text{sec} \cdot (\text{PD1 current}/\text{fA})
\]

Since PD1 and PD3 are both observing outputs of the source diffuser box, this scale factor applies equally when the source is operated in dim mode. The switchable gain preamp used with PD1 facilitates measurement of both bright and dim beams with good precision (Table 1). For example, a dim beam with 10 photons/cm²/sec at 400 nm gives DOM count rates up to 1000 Hz (depending on orientation), with PD1 currents around 500 fA. Such dim measurements are then made for each polar angle and wavelength and used to calculate the final DOM photon sensitivity.

In more detail, Figure 4 shows typical signals observed in PD1 and the DOM when dimly illuminated with steady LED light at 400 nm. Using the relation between PD1 current and photon flux in the tank, the ratio of these signals leads directly to the DOM’s photon detection efficiency. Each measurement block provides an average over 20 seconds with source on, and generally requires subtraction of a baseline value determined during source-off periods before and after. With source off, the PD1 current readout is nonzero due to slowly drifting electronic offsets (possibly negative as in Figure 4), whereas the DOM shows its PMT dark noise rate. Statistical fluctuations are observed to be quite small, and can be reduced further by averaging results from multiple such data blocks. The single photoelectron (SPE) count rate in the DOM is defined by discriminator triggers above a threshold of 0.25 times the peak charge observed in the SPE charge spectrum, which is the same condition used to count photons in the IceCube detector. In order to avoid counting afterpulses as part of the detection efficiency, each discriminator trigger starts a programmed dead time of 12.8 μs, and the remaining live time is used to calculate the corrected SPE rate.

Figure 5 shows that similar results can be obtained from a pulsed light source. In this case the DOM was programmed to report individual PMT hit times and only hits coincident with light pulses were counted in the analysis. While the time coincidence mostly avoids the effects of PMT
dark noise and afterpulses, a statistical correction factor has to be applied when individual pulses are bright enough that often two photons are detected simultaneously without being resolved (blue diamond data points). The figure shows that proper scaling between DOM counts and PD1 current is maintained while changing the per-pulse brightness or the overall rate. With pulse rate 5 kHz and DOM occupancy 0.203 photoelectrons per pulse, the ratio is measured with statistical precision better than 0.5% by averaging 3 time blocks similar to Figure 4.

4. Photodiode linearity

As explained in the previous section, the DOM measurement is done at much lower light level than can be directly measured with PD3 mounted near the DOM in the water tank. The calibration and subsequent use of PD1 for measuring photon fluxes involves a range of photodiode current spanning over five orders of magnitude, as indicated by the ratio between the photodiode currents shown in Table 1. Therefore, we need to check the linearity of the photodiode and its readout over this range. The Hamamatsu S2281 is known for its linearity [5, 6], but we still need to test for non-ideal behavior in our specific application, which includes the amplifier and readout chain.

In order to detect small nonlinear effects, we study the measured response to two steady LEDs (at 400 nm) with similar brightness. We calculate the ratio of current measured with both LEDs turned on to the sum of currents measured when each LED is turned on by itself. Any departure of this ratio from unity indicates accumulation of error from nonlinearity over a factor of two in brightness delivered to the photodiode. The accumulated nonlinearity between two very different brightnesses is then approximately the sum of contributions from each factor two between those brightnesses (assuming only small deviations), so we need to measure each contribution very precisely. This method does not require any particular linearity in the source brightness control, only that sources have stable and reproducible output when enabled.
The light from both LEDs is arranged to shine into the source diffuser entrance port, so a given measurement cycle requires only changes to the LED drive currents and no changes in optical paths. Each LED is driven independently by a current source which can be programmed by a 16-bit DAC and a bright/dim range setting, yielding very stable light outputs spanning more than four orders of magnitude. In order to avoid any effect on one LED’s brightness when the other LED is switched on, each LED is mounted with its DAC control on a separate circuit board with individual voltage regulation and digital communications to a host microcontroller. Generally we use the same brightness DAC settings for both LEDs in a given measurement. To reach the lowest light levels of interest, we insert an optical attenuator rather than reducing the LED drive current so much that noise becomes an issue.

Figure 6 illustrates the method in detail. With both LEDs off, there is an offset that must be measured and subtracted. We first measure the increment in current when LED-A only is turned on (segment 1); then the same with both LED-A and LED-B on (segment 2), and finally with only LED-B on (segment 3). For every segment we have an off-time of 10 s, an on-time of 20 s, and every time the source switches on or off, we record but do not use data for a transient time of 10 s. One full segment therefore lasts 70 s, one cycle (three segments) lasts 210 s; we record at least 30 cycles for every light setting and compute average ratios as described earlier.

The result of the linearity measurements as a function of the photodiode current is shown in Figure 7(a). The deviation of the ratio from unity is plotted in Figure 7(b). Generally the measurements obey linearity within uncertainties well below 0.1%; statistical uncertainties and subtle effects at very dim light levels are currently the limiting factor, but still contributing only 0.1-0.2%. Considering the sum of nonlinearities from each doubling of intensity from 500 fA to 50 nA, we expect the total nonlinearity error to be ≤0.2% when comparing dim and bright light levels with PD1. There is also some uncertainty in the relative gains of the PD1 preamplifier which need to be chosen for the different light levels; theoretically those are set by high precision resistors but their ratios can also be measured with precision comparable to the nonlinearity results using the same LED sources.

5. Summary

A lab setup has been developed to measure the DOM absolute optical sensitivity as a function of incident polar angle and wavelength. An optical system illuminates the DOM uniformly from a certain direction, with several pulsed and steady light sources available for wavelength selection and systematic studies. The calibration is transferred from a NIST calibrated photodiode to the
DOM by using additional photodiodes which monitor the beam, as its intensity is varied over five orders of magnitude. We verified that the error introduced by nonlinearities of the system, operated at the brightest and dimmest limit of the dynamic range, is less than 0.2%. In the near future, we plan to calibrate 10 to 20 of about 100 leftover DOMs, and we foresee each measurement to take about one week. Since relative sensitivities of all IceCube DOMs were measured in pre-deployment testing, a DOM dependent correction factor can be used to improve the understanding of the IceCube detector and the energy scale. The calibration system will be also available for calibration of optical sensors to be used in HAWC¹ and future in-ice detectors such as PINGU [7].

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