A Search for Single Photon Events in Neutrino Interactions

The Harvard community has made this article openly available. Please share how this access benefits you. Your story matters

| Citation               | Kullenberg, C.T., S.R. Mishra, D. Dimmery, X.C. Tian, D. Autiero, S. Gninenko, A. Rubbia, et al. 2012. A Search for Single Photon Events in Neutrino Interactions. Physics Letters B 706, no. 4-5: 268–275. |
|------------------------|------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|
| Published Version      | 10.1016/j.physletb.2011.11.049                                                                                                                                                                                                                               |
| Citable link           | http://nrs.harvard.edu/urn-3:HUL.InstRepos:11870365                                                                                                                                                                                                            |
| Terms of Use           | This article was downloaded from Harvard University’s DASH repository, and is made available under the terms and conditions applicable to Other Posted Material, as set forth at http://nrs.harvard.edu/urn-3:HUL.InstRepos:dash.current.terms-of-use#LAA |
A search for single photon events in neutrino interactions

C.T. Kullenberg, S.R. Mishra, D. Dimmery, X.C. Tian, D. Autiero, S. Gninenko, A. Rubbia, S. Alekhin, P. Astier, A. Baldisseri, M. Baldo-Ceolin, M. Banner, G. Bassompierre, B. K. Benslama, N. Besson, I. Bird, B. Blumenfeld, F. Bobisut, J. Bouchez, S. Boyd, A. Bueno, S. Bunyatov, L. Camilleri, A. Cardini, P.W. Cattaneo, V. Cavasinni, A. Cervera-Villanueva, R. Challis, A. Chukanov, G. Collazuol, G. Conforto, C. Conta, M. Contalbrigo, R. Cousins, H. Degaubert, A. De Santo, T. Del Prete, L. Di Lella, E. Do Couto e Silva, J. Dumarchez, M. Ellis, G.J. Feldman, R. Ferrare, V. Flaminio, M. Fraternali, J.-M. Gaillard, A. Gangler, A. Geiser, D. Geppert, D. Gibin, A. Godley, J.-J. Gomez-Cadenas, J. Gosset, C. Gösling, M. Gouanère, A. Grant, G. Graziani, A. Guglielmi, C. Hagner, J. Hernando, P. Hurst, N. Hyett, E. Iacopini, C. Joseph, F. Juget, N. Kent, O. Klimov, J. Kokkonen, A. Kovzelev, A. Krasnoperov, J. Kim, M. Kirsanov, S. Kulagin, I. Lacapra, C. Lachaud, B. Lakić, A. Lanza, L. La Rotonda, M. Laveder, A. Letessier-Selvon, J.-M. Levy, J. Ling, L. Linssen, A. Ljubičić, J. Long, A. Lupi, V. Lyubushkin, A. Marchioni, F. Martelli, X. Méchain, J.-P. Mendiburu, J.-P. Meyer, M. Mezzetto, G.F. Moorhead, D. Naumov, P. Nédélec, Y. Nevedova, C. Nguyen-Mau, D. Orestano, F. Pastore, L.S. Peak, E. Pennacchio, H. Pessard, R. Petti, A. Placci, G. Polesello, D. Pollmann, A. Polyarush, C. Poulsen, B. Popov, L. Rebuffi, J. Rico, P. Riemann, C. Roda, F. Salvatore, O. Samoylov, K. Schahmaneche, B. Schmidt, T. Schmidt, A. Scona, A.M. Scott, M.B. Seaton, M. Sevior, D. Sillo, F.J.P. Soler, G. Sozzi, D. Steele, U. Stiegler, M. Stipčević, Th. Stolarczyk, M. Tareb-Reyes, G.N. Taylor, V. Tereshchenko, A. Toropin, A.-M. Touchard, S.N. Tovey, M.-T. Tran, E. Tsesmelis, J. Ulrichs, L. Vacavanti, M. Valdata-Nappi, V. Valsecchi, F. Vannuccini, K.E. Varvell, M. Veltri, V. Vercesi, G. Vidal-Sitjes, J.-M. Vieira, T. Vinogradova, F.V. Weber, T. Weiss, F.F. Wilson, L.J. Winton, B.D. Yabsley, H. Zaccane, K. Zuber, P. Zuccon
1. The question

The MiniBooNE experiment, in the neutrino mode, has reported an excess of $129 \pm 43$ (Stat $\oplus$ Syst) electron-like events in the 200–475 MeV energy range [1], a range largely unaffected by the $\nu_\mu \rightarrow \nu_e$ oscillations at $\Delta m^2 > 1$ eV$^2$, which is expected to dominate the 475–1250 MeV range [2]. Assuming that the efficiency of these events is 0.25, similar to that of the $\nu_e$-induced electrons, the rate of the excess with respect to $\nu_\mu$-CC is $\approx 3 \times 10^{-3}$. In the antineutrino mode, initially the MiniBooNE data were consistent with background in the 200–475 MeV range [3]. Recently, with additional antineutrino data, they report a 2 standard deviation excess in this region [4]. MiniBooNE lacks the resolution to determine if the excess is due to $e^-$, $e^+$, or a photon. In neutrino interactions single photons do not occur: they occur in pairs from the $\pi^0 \rightarrow \gamma \gamma$ decay. Evidence of events with a single photon and nothing else is in neutrino-induced neutral current (NC) interaction will signal either new or unconventional physics. This anomaly needs to be checked experimentally.

The low-energy MiniBooNE excess has motivated novel hypotheses invoking conventional, to unconventional, to entirely new physics processes. An explanation [5] of the excess attributed to the internal bremsstrahlung associated with muons in $\nu_\mu$-induced charged current interaction (CC) has been refuted by MiniBooNE [6]. Harvey, Hill, and Hill [7,8] postulate a new anomaly-mediated interaction between the $\nu$-induced Z$^0$-boson, the photon, and a vector meson, such as $\omega$, coupling strongly to the target nucleus. A similar, single photon process has been envisioned by Jenkins and Goldman [9]. In the anomaly-mediated neutrino-photon (ANP) interaction, $\nu N \rightarrow \nu N \gamma$, where $N$ is the target nucleus or nucleon, the observable is a single $\gamma$, approximately collinear with the incident neutrino direction in the lab frame with a recoiling $N$ which remains intact and largely undetected; the form factor of $N$ is expected to induce the single $\gamma$ to be pulled forward. The ANP cross-section is expected to increase with neutrino energy ($E_\nu$). The ANP interaction is akin to the exclusive $\pi^0$ production when the neutrino coherently scatters off a target nucleus ($\text{Col}\pi^0$), $\nu + N \rightarrow \nu + N + \pi^0$, where the only observable is $\pi^0 \rightarrow \gamma \gamma$. The phenomenology of ANP developed by Hill in [10], and largely focussed on the low energy neutrino interaction ($E_\nu \sim 50$–100 MeV), predicts a cross-section sufficiently large to explain the MiniBooNE anomaly. If true, the ANP interaction will also impact astrophysical processes such as neutron star cooling. Hypotheses involving new physics include: decay of a heavy neutrino into a light neutrino and a photon, $\nu_h \rightarrow \nu + \gamma$; by Gninenko [11,12]; CP-violation in 3 $\oplus$ 2 model by Maltoni and Schwetz [13], and by Goldman et al. [14]; extra-dimensions in 3 $\oplus$ 1 model by Pas et al. [15]; and models by Giunti and Laveder [16]; Lorentz-violation by Katori et al. [17]; CPT-violation in 3 $\oplus$ 1 model by Barger et al. [18]; new gauge bosons with sterile neutrinos by Nelson and Walsh [19]; and soft-decoherence by Farzan et al. [20]; etc. The NOMAD Collaboration has reported a search of heavy-$\nu$ which mixes with $\nu_\mu$’s produced in the SPS proton-target, and then decays into $e^- e^+$ pair in the detector [21]. The search focussed on highly collinear ($C = [1 - \cos \Theta_{ee}] \leq 2 \times 10^{-5}$) and high energy ($E_{ee} \approx 4$ GeV) $e^- e^+$ events, where $\Theta_{ee}$ is the energy of the converted photon, and resulted in a single event consistent with the estimated background. The present work extends this search with no cut on $C$.

The high resolution NOMAD data allow a sensitive search for neutrino induced single photon events (1$\gamma$). The detector measures the energy and emission angle of the photon. Additionally, the detector affords the redundancy to measure in situ all relevant backgrounds relieving the reliance on Monte Carlo simulation (MC) of the conventional processes. As regards the MiniBooNE low-energy anomaly, we note that the average energy of the SPS neutrino flux at NOMAD, $E_\nu \sim 25$ GeV, is much higher than that of MiniBooNE, $E_\nu \sim 1$ GeV. However, the NOMAD data provide a stringent check in a different energy range. If a mechanism explaining the anomaly were applicable to higher energies, such as ANP or $\nu$-decay hypotheses, then this search, even if negative, will furnish meaningful limits.

2. Beam and detector

The Neutrino Oscillation Magnetic Detector (NOMAD) experiment at CERN used a neutrino beam produced by the 450 GeV protons from the Super Proton Synchrotron (SPS) incident on a beryllium target and producing secondary $\pi^\pm$, $K^\pm$, and $\pi^0$ mesons. The
positively charged mesons were focussed by two magnetic horns into a 290 m long evacuated decay pipe. Decays of $\pi^\pm$, $K^\pm$, and $K^0_L$ produced the SPS neutrino beam. The average neutrino flight path to NOMAD was 628 m, the detector being 836 m downstream of the Be-target. The SPS beamline and the neutrino flux incident at NOMAD are described in [22]. The $\nu$-flux in NOMAD is constrained by the $\pi^\pm$ and $K^\pm$ production measurements in proton-Be collision by the SPY experiment [23–25] and by an earlier measurement conducted by Atherton et al. [26]. The $E_\gamma$-integrated relative composition of $\nu_\mu : \bar{\nu}_\mu : \nu_e : \bar{\nu}_e$ CC events, constrained in situ by the measurement of CC-interactions of each of the neutrino species, is $1.00:0.025:0.015:0.0015$. Thus, 97.5% of the events are induced by neutrinos with a small antineutrino contamination, similar to that reported by MiniBooNE [2].

The NOMAD apparatus, described in [27], was composed of several sub-detectors. The active target comprised 132 planes of $3 \times 3 \text{m}^2$ drift chambers (DC) with an average density similar to that of liquid hydrogen (0.1 gm/cm$^3$). On average, the equivalent material in the DC encountered by particles produced in a $\nu$-interaction was about half of a radiation length and a quarter of a hadronic interaction length ($\lambda$). The fiducial mass of the NOMAD DC-target, 2.7 tons, was composed primarily of carbon (64%), oxygen (22%), nitrogen (6%), and hydrogen (5%) yielding an effective atomic number, $A = 12.8$, similar to carbon. Downstream of the DC, there were nine modules of transition radiation detectors (TRD), followed by a preshower (PRS) and a lead-glass electromagnetic calorimeter (ECAL). The ensemble of DC, TRD, and PRS/ECAL was placed within a dipole magnet providing a 0.4 T magnetic field orthogonal to the neutrino beam line. Two planes of scintillation counters, $T_1$ and $T_2$, positioned upstream and downstream of the TRD, provided the trigger in combination with an anticoincidence signal, $\mathcal{V}$, from the veto counter upstream and outside the magnet. Downstream of the magnet was a hadron calorimeter, followed by two muon-stations each comprising large area drift chambers and separated by an iron filter. The two stations, placed at 8- and 13-$\lambda$’s downstream of the ECAL, provided a clean identification of the muons. The schematic of the DC-tracker with a $\gamma$-conversion event in the Y-Z view is shown in Fig. 1. The charged tracks in the DC were measured with an approximate momentum ($p$) resolution of $\sigma_p/p \sim 0.05/\sqrt{1+0.008p/\sqrt{E}}$ (p in GeV/c and L in meters) with unambiguous charge separation in the energy range of interest. The experiment recorded over 1.7 million neutrino interactions in the active drift chamber (DC) target in the range $0.1 \leq E_\nu \leq 300$ GeV.

3. The exclusive photon signature

In NOMAD the clearest signature of a $\gamma$ is a $V^0$ arising from its conversion in the DC into $e^-$ and $e^+$ tracks as shown in Fig. 1. The $\gamma$-momentum is reconstructed by measuring the 3-momenta of the associated $e^-/e^+$ tracks in the DC. The ECAL measures energy associated with the $e^+/e^-$ tracks plus any unconverted photons. Because the trigger counters are situated at the downstream end of the detector, only a fraction of events containing single photon will trigger the apparatus, namely events where the photon undergoes conversion in DC and the resulting $e^+$ or $e^-$ traverses the trigger counters. Events with unconverted photons reaching the ECAL will not trigger the apparatus. For example, about 29% of the COS $\pi^0$ events, containing a $\pi^0$ and ‘nothing else’, trigger the apparatus; the loss arises from the unconverted photons and, among the converted photons, from the $e^-/e^+$ tracks that do not reach the downstream trigger counters.

Single photon events will manifest as a $V^0$, composed of $e^-$ and $e^+$, with no additional energy in the ECAL. A converted photon is defined as a $V^0$ such that the invariant mass of the $e^-$ and $e^+$ is less than 100 MeV which selects the converted photons with 95% purity and 97% efficiency. Furthermore, if the single photon is approximately collinear with the incident neutrino then the collinearity of the photon-track offers an additional kinematic handle. Because we do not have a definitive simulation for the single $\gamma$ process, we use one of the photons from the COS $\pi^0$, while ignoring the other photon, as a guide for the signal. We define two signal boxes. Box 1 comprises single $\gamma$ events with negligible energy in the ECAL associated with neutral particles ($\gamma$, neutron, etc.) quantified by the quantity, $PAN = (E_\gamma - E_{\text{Neut}})/(E_\gamma + E_{\text{Neut}})$, $\geq 0.9$, where $E_\gamma$ is the energy of the photon derived from the $e^-/e^+$ tracks and $E_{\text{Neut}}$ is the ECAL energy associated with other neutral particles, i.e. not associated with $e^-/e^-$ tracks. (For calculation of $E_{\text{Neut}}$ see, for example, [28].) The $PAN \geq 0.9$ cut is model independent. Box 2, a subset of Box 1, comprises single $\gamma$ events, provided that they are approximately collinear with the incident neutrino, characterized by the variable, $\zeta = E_\nu(1 - \cos(\theta_{\gamma}) \geq 0.05$, where $E_\nu$ and $\theta_{\gamma}$ are the photon’s energy and polar angle with respect to the neutrino beam. (The variable $\zeta$ has the property that its distribution depends weakly on the incident neutrino energy.) Approximately 90% of the photons from COS $\pi^0$ have $\zeta \leq 0.05$. Although the $\zeta \leq 0.05$ cut will be somewhat model-dependent, at the SPS energies ($E_\nu \simeq 25 \text{ GeV}$) most exotic models suggest that the single photon will be approximately collinear with the incident neutrino.

The backgrounds to the single $\gamma$ are all interactions where all daughter particles, except one photon, evade detection. The background includes the COS $\pi^0$ interaction, the deep inelastic $\nu$-NC scattering (NC-DIS), and $\nu$-interactions occurring outside the fiducial volume (OBG). Note, that with a single converted $\gamma$ in DC the coordinates of the $\nu$-interaction cannot be ascertained. The sample with $PAN < 0.9$ is composed of events with a $\nu$ and neutral energy in the ECAL as expected from a $\pi^0$ decay and constitutes an important control sample. All three backgrounds are determined in situ as described below.

The analysis proceeds as follows. First, the sample with a single photon-conversion in DC is selected. Second, the three backgrounds to the single photon sample are calibrated using data outside Box 1. The analysis leads to an experimentally constrained prediction of the background events in the signal region, Box 1 and Box 2. Finally, boxes are opened and the predictions are compared with the observed data.

![Fig. 1. Schematic of the DC tracker with a single $\gamma$-conversion event in NOMAD in Y (vertical) vs Z (horizontal) view. The energy and angle with respect to the neutrino direction of the $\gamma$ are 3.1 GeV and 0.095 rad.](image-url)
4. Selection of single photon events

We select $\nu$-induced events where a single photon converts ($V^0$) within the fiducial volume of the DC target. The analysis uses the entire NOMAD data and the associated Monte Carlo (MC) samples described in [29]. The NC-DIS sample, defined by requiring that the generated invariant hadronic mass squared ($W^2$) be $\geq 1.96$ GeV$^2$, is normalized to 0.53 $\times$ 10$^6$ events which corresponds to 37% of the $\nu_\mu$-CC. The normalization of the NC-Resonance ($W^2 < 1.96$) sample is set at 3.5% of the NC-DIS. The $\gamma$-$\nu_\mu$ interaction is simulated using the Rein–Sehgal (RS) model [30] which agrees well with our measurement [31].

The MC and data samples are subjected to a preselection requiring: (a) the presence of a converted photon whose reconstructed conversion point ($X$, $Y$, $Z$) be within the fiducial volume, ($X$, ($Y - 5$)) $\leq 130$ cm and $Z_{\text{Min}} \leq Z \leq 405$ cm where $Z_{\text{Min}}$ is 5 and 35 cm for the two configurations of the detector composing more than 95% of the NOMAD data; (b) no reconstructed muon ($\mu$ID-veto) and a single photon-induced $V^0$; (c) the $X$ and $Y$ coordinate of the $\gamma$-vector when extrapolated back to the upstreammost DC at $Z_{\text{Min}}$ be within the transverse fiducial volume, ($X$, $Y - 5$)$_{\text{Proj}}$ $\leq 130$ cm, largely eliminating $\gamma$’s that enter from the sides; and (d) the energy of the $e^+e^-$ pair be $\geq 1.5$ GeV which reduces the NC-DIS and OBG background while having a small effect on the $\gamma$-$\nu_\mu$ sample which also serves as a guide for the single $\gamma$ signal. The preselection reduces the NC-DIS and data samples by more than a factor of one hundred. It is noteworthy that only a small fraction of the NC-Resonance pass the selection since most of the photons, coming from $\pi^0$ emitted at large angles, either fail to trigger the apparatus or have energies below 1.5 GeV. Given the paucity of the NC-Resonance events, it has been added to the NC-DIS sample. Finally a negligible fraction of $\nu_\tau$-CC ($< 10^{-5}$) pass the selection; consequently the CC sample has been ignored in this analysis. The preselection is presented in Table 1.

The final event selection follows the preselection with more stringent requirements. The vertex coordinates of $V^0$ are required to be within ($X$, ($Y - 5$)) $\leq 120$ cm and $Z_{\text{Min}} \leq Z \leq 405$ cm. Two additional cuts (Clean-$V^0$) are imposed to reduce outside background by requiring that there be no tracks upstream of the photon conversion and that there be no hits associated with the tracks composing the converted-$\gamma$ in the most upstream DC. Finally, the $M_{ee}$ cut is tightened to $\leq 50$ MeV which increases the photon conversion purity to $\geq 98$% while reducing the efficiency to 93%. Table 2 summarizes the selection of events in the MC samples. The preselected data are subjected to identical cuts. Distributions of the $X$, $Y$, and $Z$ coordinates of the $\gamma$ vertex in data and the corresponding MC prediction, composed of $\gamma$-$\nu_\mu$, OBG, NC-DIS, agree. Only the control sample with $\gamma < 0.9$ is examined to check and constrain the background prediction. The calibration of the three backgrounds is presented next.

5. Constraining the backgrounds and the PAN $< 0.9$ sample

The prediction of the backgrounds to the single $\gamma$ sample is data driven. The estimation of the $\gamma$-$\nu_\mu$-induced $1\gamma$ needs to be based on the observed $2V^0$ sample where both photons convert in the DC. Furthermore, Monte Carlo simulations can neither reliably provide the OBG nor the NC-DIS contribution to the $1\gamma$ sample. These backgrounds need to be determined using data themselves.

First, we present the calibration of the $\gamma$-$\nu_\mu$ contribution to the $1\gamma$ sample because it is the simplest. The analysis of the $2\gamma$ sample, where two photons and nothing else is observed, provides a measurement of the $\gamma$-$\nu_\mu$ process. It yields a rate with respect to the inclusive $\nu_\tau$-CC, $\gamma$-$\nu_\mu$-induced $1\gamma$ sample of 21 $\gamma$-$\nu_\mu$-induced $1\gamma$ sample of 927 $\times$ 10$^{-3}$ events with DCA-Z $\leq Z_{\text{Min}}$ as shown in Table 3. The number of $2\gamma$ events with DCA-Z $> Z_{\text{Min}}$ originating from events occurring at $Z < Z_{\text{Min}}$ but producing no visible tracks, is then calculated as (451,927) $\times$ 169 = 82.2 $\pm$ 6.1. This constitutes the OBG background to the $2\gamma$ sample. The so normalized OBG sample is then subjected to the $1\gamma$ analysis yielding a calibrated OBG prediction, with a 7.7% error (driven by the statistics of 169 events), shown in Table 4.

### Table 1
Preselection of events with a single, converted $\gamma$ in $\gamma$-$\nu_\mu$, NC-Resonance, and NC-DIS MC samples, and data. The MC samples, having much larger statistics than data, are normalized to the expectation as described in the text. The $\gamma$-$\nu_\mu$-induced events where a single photon converts ($V^0$) is 927 events with DCA-Z $\leq Z_{\text{Min}}$.

| Cut | Coh $\gamma^0$ | NC-Resonance | NC-DIS MC | Data |
|-----|---------------|----------------|-----------|------|
| Fiducial cut | 4900 | 20,000 | 530,000 | 4,018,980 |
| $\mu$ID-veto and single $\gamma$-induced $V^0$ ($M_{\ell} \leq 100$ MeV) | 819 | 306 | 4330 | 34,062 |
| ($|X| < 130$ cm, $|Y - 5|_{\text{Proj}} < 130$ cm | 719 | 138 | 3076 | 10,547 |
| $E_{\text{ee}} \geq 1.5$ GeV | 516 | 69 | 1846 | 4543 |

### Table 2
Final selection of single $\gamma$ events in the MC samples: The MC samples have been normalized as presented in Section 4.

| Cut | Coh $\gamma^0$-RS | NC-DIS $\oplus$ Resonance |
|-----|-------------------|--------------------------|
| Start | 516 | 1915 |
| Tighter Fid-Cuts | 483 | 1775 |
| $M_{ee} \leq 50$ MeV and Clean-$V^0$ | 386 | 400 |

### Table 3
Events passing and failing the Z-cut in $2\gamma$ samples in data and OBG-data.

| DCA-Z $\geq Z_{\text{Min}}$ | DCA-Z $< Z_{\text{Min}}$ |
|-----------------------------|-----------------------------|
| $2\gamma$ data | 381 | 169 |
| $2\gamma$ OBG-data | 451 | 927 |
Finally, we present the calibration of the NC-DIS contribution to the 1γ sample using the control region PAN < 0.9 dominated by events containing a photon-conversion accompanied by a neutral energy cluster in the ECAL. Fig. 2 shows the observed PAN distribution for the full 1γ sample although the signal region PAN ≥ 0.9 is not looked at till all backgrounds are finalized. The figure evinces agreement between data and the prediction for PAN < 0.9. Furthermore, the observed kinematic variables associated with the neutral cluster in ECAL are found to be consistent with the prediction. The observed momentum distribution of the photon is consistent with the prediction as evidenced by Fig. 3. Lastly, Fig. 4 shows that the measured ζγ distribution agrees with the MC prediction. Because the π0-induced photons in NC-DIS will have a broader ζγ than those in Coh.π0, we use ζγ > 0.05 to normalize the NC-DIS. The NC-DIS normalization factor is 1.13 ± 0.14. (This normalization factor has been applied to the NC-DIS in the figures.) The logic behind using this control sample to constrain the background in the signal region is that 1γ events come from π0-dominated neutral current events where one of the photons evades detection. One concern, however, is that of a possible systematic bias when using the PAN < 0.9, a region where both photons are reconstructed in DC and ECAL, i.e. the two photons are emitted in the forward direction, to predict the NC-DIS in the PAN ≥ 0.9 where the second photon misses the ECAL. To check this concern, ordinary NC events, where charged tracks define the event, and, thereby, relieving the forward bias in the 1γ reconstruction due to the trigger requirement, are selected in data and MC. A single 1γ is selected downstream of the event. The charged tracks are ignored, the 1γ is subjected to the current analysis. The Data/MC ratio is examined in Eγ and ζγ variables. The check reveals that the shapes of these variables in MC are consistent with those of data: the ratio is unity within ±5% in the kinematic range, well covered by the 12% error ascribed to the NC-DIS. Table 4 presents the NC-DIS contribution to the 1γ sample.

Table 4

| Cut                | Coh.π0, RS | NC-DIS + Resonance | OBG | Total | Data |
|--------------------|-----------|--------------------|-----|-------|------|
| V0 sample          | 385.9     | 400.1              | 341.3 | 1127.3 | 1149 |
| MC error           | 14.5%     | 12.0%              | 7.7% |       |      |
| Background         |           |                    |     |       |      |
| PAN < 0.9          | 353.9     | 347.7              | 296.5 | 998.1 ± 69.9 ± 25.0 | 994 |
| Signal             |           |                    |     |       |      |
| PAN ≥ 0.9          | 32.0      | 52.4               | 44.8 | 129.2 ± 8.5 ± 3.3 | 155 |
| PAN ≥ 0.9 ⊕ ζγ ≤ 0.05 | 22.8      | 22.6               | 31.2 | 76.6 ± 4.9 ± 1.9 | 78  |

6. Results

Fig. 1 shows a representative event in the signal region, PAN ≥ 0.9. The first signal-box, Box1, with PAN ≥ 0.9, yields 155 observed events with a predicted background 129.2 ± 8.5 ± 3.3 events yielding an excess of 25.8 ± 15.5 events. Fig. 5 and Fig. 6 present the Pγ and ζγ comparison between data and MC. Absence of significant excess leads us to interpret this as null evidence for excess of single photon events. Assuming the error on the background-prediction to be Gaussian, we derive an upper limit of < 51 events at 90% CL in Box1.

The second signal-box, Box2, with PAN ≥ 0.9 and ζγ ≤ 0.05, yields 78 observed events with a predicted background 76.6 ± 4.9 ± 1.9 events yielding an excess of 1.4 ± 10.3 events. Events in the Box2 exhibit kinematic characteristics consistent with the background prediction. The vertex-distributions of the 1γ events in the signal-region are in agreement with the MC prediction; as are, within errors, the Pγ and ζγ distributions, shown in Fig. 7 and Fig. 6. In addition, the observed collinearity (C) of the photon matches that of the prediction, see Fig. 8. Assuming the error to be Gaussian, we derive an upper limit of < 18 events at 90% CL. Table 4 presents the final enumeration.
Fig. 3. Comparison of $P_\gamma$ between data and MC in PAN < 0.9 region.

Fig. 4. Comparison of $\zeta_\gamma$ distribution of the 1$\gamma$ sample with PAN = 0.9 in data and MC. The $\zeta_\gamma > 0.05$ region is used to normalize the NC-DIS.

Fig. 5. Comparison of $P_\gamma$ between data and MC in Box1, PAN $\geq$ 0.9 region.

Fig. 6. Comparison of $\zeta_\gamma$ distribution between data and MC in Box1, PAN $\geq$ 0.9 region.
The only remaining task is to set an upper limit on the rate of single $\gamma$ events. To determine the signal efficiency, we assume that the 'signal' photon has kinematics similar to that of one of the photons from the Coh$\pi^0$ interaction (the other photon is removed from the simulation). Using the RS Coh$\pi^0$ model, the signal efficiency for Box1, $PAN \geq 0.9$, is 8.8%. The efficiency for Box2, $PAN \geq 0.9$ and $\zeta \leq 0.05$, is 8.0%. The number of fully corrected $\nu_\mu$-CC in the same fiducial volume is measured to be $1.24 \times 10^6$

We obtain the following upper limits on the rate of single photon events in $\nu$-interactions:

$$\frac{\sigma(\text{Single} - \gamma)}{\sigma(v_\mu A \rightarrow \mu^{-}X)} < 4.0 \times 10^{-4} \quad (90\% \, \text{CL}),$$  

$$\frac{\sigma(\text{Single, Forward} - \gamma)}{\sigma(v_\mu A \rightarrow \mu^{-}X)} < 1.6 \times 10^{-4} \quad (90\% \, \text{CL}).$$

In summary, we have presented a search for single photon events in interactions of neutrinos with average energy $E_\nu \simeq 25$ GeV. All relevant backgrounds are constrained using data control samples. No significant excess is seen. Assuming that the hypothetical signal has kinematics similar to those of a photon from the Coh$\pi^0$ interaction, the analysis imposes an upper limit on the rate of excess of single photon events of $< 4.0 \times 10^{-4}$ per $v_\mu$-CC at 90% CL; with an additional soft collinearity cut (permitting 90% of $\gamma$ from Coh$\pi^0$) the limit is $< 1.6 \times 10^{-4}$ per $v_\mu$-CC at 90% CL.

Acknowledgements

We extend our grateful appreciations to the CERN SPS staff for the magnificent performance of the neutrino beam. We (C.T.K. and S.R.M.) warmly thank Bill Louis, Richard Hill, Chris Hill, James Jenkins and Terry Goldman for many stimulating discussions and insights. The experiment was supported by the following agencies: ARC and DIISR of Australia; IN2P3 and CEA of France, BMBF of Germany, INFN of Italy, JINR and INR of Russia, FNSRS of Switzerland, DOE, NSF, Sloan, and Cottrell Foundations of USA, and VP Research Office of the University of South Carolina.

References

[1] A.A. Aguilar-Arevalo, et al., MiniBooNE Collaboration, Phys. Rev. Lett. 102 (2009) 101802, arXiv:0812.2243 [hep-ex].
[2] A.A. Aguilar-Arevalo, et al., MiniBooNE Collaboration, Phys. Rev. Lett. 98 (2007) 231801, arXiv:0704.1500 [hep-ex].
[3] A.A. Aguilar-Arevalo, et al., MiniBooNE Collaboration, Phys. Rev. Lett. 103 (2009) 111801.
[4] D. Bjørn, MiniBooNE Collaboration, in: NuFact 2011, Geneva, Switzerland.
[5] A. Bodek, arXiv:0709.4004 [hep-ex].
[6] A.A. Aguilar-Arevalo, et al., MiniBooNE Collaboration, arXiv:0710.3897 [hep-ex].
[7] J.A. Harvey, C.T. Hill, R.J. Hill, Phys. Rev. Lett. 99 (2007) 261601, arXiv:0708.1281 [hep-ph].
[8] J.A. Harvey, C.T. Hill, R.J. Hill, Phys. Rev. D 77 (2008) 085017, arXiv:0712.1230 [hep-ph].
[9] J. Jenkins, T. Goldman, Phys. Rev. D 80 (2009) 053005, arXiv:0906.0984 [hep-ph].
[10] R. J. Hill, Phys. Rev. D 81 (2010) 013008, arXiv:0905.0291 [hep-ph].
[11] S. N. Gninenko, Phys. Rev. Lett. 103 (2009) 241802, arXiv:0902.3802 [hep-ph].
[12] S. N. Gninenko, Phys. Rev. D 83 (2011) 015015, arXiv:1009.5536 [hep-ph].
[13] M. Maltoni, T. Schwetz, Phys. Rev. D 76 (2007) 093005, arXiv:0705.0107 [hep-ph].
[14] T. Goldman, G.J. Stephenson Jr., B.H. McKellar, Phys. Rev. D 75 (2007) 091301.
[15] H. Pas, S. Pakvasa, T.J. Weiler, Phys. Rev. D 72 (2005) 095017, arXiv:hep-ph/0504096.
[16] C. Giunti, M. Laveder, Phys. Rev. D 83 (2011) 053006, arXiv:1012.0267 [hep-ph];
C. Giunti, M. Laveder, arXiv:1107.1452 [hep-ph];
C. Giunti, M. Laveder, arXiv:1109.4033 [hep-ph].
[17] T. Katori, V.A. Kostelecky, R. Tayloe, Phys. Rev. D 74 (2006) 105009, arXiv:hep-ph/0506154.
[18] V. Barger, D. Marfatia, K. Whisnant, Phys. Lett. B 576 (2003) 303, arXiv:hep-ph/0308299.
[19] A.E. Nelson, J. Walsh, Phys. Rev. D 77 (2008) 033001, arXiv:0711.1363 [hep-ph].
[20] Y. Farzan, T. Schwetz, A.Y. Smirnov, JHEP 0807 (2008) 067, arXiv:0805.2098 [hep-ph].
[21] P. Astier, et al., NOMAD Collaboration, Phys. Lett. B 506 (2001) 27, hep-ex/0101041.
[22] P. Astier, et al., NOMAD Collaboration, NIM A 515 (2003) 800.
[23] G. Ambrosini, et al., SPY Collaboration, Phys. Lett. B 420 (1998) 225.
[24] G. Ambrosini, et al., SPY Collaboration, Phys. Lett. B 425 (1998) 208.
[25] G. Ambrosini, et al., SPY Collaboration, Eur. Phys. J. C 10 (1999) 605.
[26] H.W. Atherton, et al., CERN, Yellow Report 80-07, 1980.
[27] J. Altegoer, et al., NOMAD Collaboration, NIM A 404 (1998) 96.
[28] P. Astier, et al., NOMAD Collaboration, Nucl. Phys. B 611 (2001) 3.
[29] Q. Wu, et al., NOMAD Collaboration, Phys. Lett. B 660 (2008) 19, arXiv:07111183 [hep-ex].
[30] D. Rein, L.M. Sehgal, Nucl. Phys. B 223 (1983) 29.
[31] C.T. Kullenberg, et al., NOMAD Collaboration, Phys. Lett. B 682 (2009) 177, arXiv:0910.0062 [hep-ex].