Fingerprint of different spin–orbit terms for spin transport in HgTe quantum wells

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Abstract. Using \( \mathbf{k} \cdot \mathbf{p} \) theory, we derive an effective four-band model describing the physics of the typical two-dimensional topological insulator (HgTe/CdTe quantum well (QW)) in the presence of an out-of-plane (in the \( z \)-direction) inversion breaking potential and an in-plane potential. We find that up to third order in perturbation theory, only the inversion breaking potential generates new elements to the four-band Hamiltonian that are off-diagonal in spin space. When this new effective Hamiltonian is folded into an effective two-band model for the conduction (electron) or valence (heavy hole) bands, two competing terms appear: (i) a Rashba spin–orbit interaction originating from inversion breaking potential in the \( z \)-direction and (ii) an in-plane Pauli term as a consequence of the in-plane potential. Spin transport in the conduction band is further analysed within the Landauer–Büttiker formalism. We find that for asymmetrically doped HgTe QWs, the behaviour of the spin-Hall conductance is dominated by the Rashba term.
1. Introduction

Two-dimensional (2D) topological insulators form a new state of matter where the gapped bulk state is accompanied by gapless spin edge states that are protected against small perturbations by time reversal symmetry [1]–[7]. In contrast to the chiral nature of the quantum-Hall state, where only one channel propagates at the edge of the system, these new edge states have a helical nature, i.e. there are two counter-propagating spin-edge channels at the edge [1]–[7]. The topological protection of the edge states is connected with Kramer’s theorem and states that a system with an odd number of Kramer’s pairs at a given edge is protected against single-particle excitations [6]. The simplest way to find topologically nontrivial insulators is to look at systems where the conduction and the valence bands have opposite parity and a change in band ordering (band inversion) occurs as a function of a tuning parameter like the strength of spin–orbit (SO) coupling [4, 8]. This criterion leads to the unified Dirac form of the effective Hamiltonian for topological insulators with a spatial inversion centre [4, 9]. In this paper, we will be particularly interested in HgTe/CdTe quantum wells (QWs), which is the first topological insulator discovered in nature [4, 5]. This system can be tuned from the normal to 2D topological insulator phase by changing the thickness $d$ of the HgTe layer [4, 5]. Recent conductance measurements in multi-terminal structures [5, 7] clearly show the existence of 1D helical edge channels in this material for $d$ larger than the critical value $d_c = 6.3$ nm [5, 7].

An effective four-band model introduced by Bernevig, Hughes and Zhang (BHZ) [4] consists of two disconnected blocks, each having the form of the Dirac Hamiltonian in 2D and additional quadratic terms crucial for defining the concept of band inversion. The BHZ model adequately describes the insulating regime in HgTe/CdTe QWs close to the $\Gamma$ point and the topological quantum phase transition near the critical thickness $d = d_c$. It has been extended to include the bulk inversion symmetry breaking effects in [6]. However, this model does not yet include the structural inversion asymmetry (SIA) terms that can be very large in this narrow gap material. Indeed, it was shown experimentally that an external top gate applied to the HgTe/CdTe QWs can change the energy of the Rashba SO splitting in the range from 0 to 30 meV [10] and the samples can be tuned from insulating to metallic regime [11]. Furthermore,
the Aharonov–Casher oscillations [12] as well as the ballistic spin-Hall effect in HgTe/CdTe QWs [11], which occur in the metallic regime, can be well described by an effective two-band (electron or heavy hole) model taking into account the Rashba SO interactions. Also, very recent angle-resolved photoemission experiments on Bi$_2$Se$_3$ thin films [13] show the importance of SIA terms [14]–[17]. Therefore, it is desirable to build a unified Hamiltonian for 2D topological insulators, that includes information about both the band structure and the presence of an inversion breaking potential. This is the purpose of this paper. Using \( \mathbf{k} \cdot \mathbf{p} \) theory we derive an extension of the BHZ model describing a 2D topological insulator, which includes an out-of-plane (in the \( z \)-direction) inversion breaking potential and an in-plane potential. The central result of this paper, the generalized four-band Hamiltonian, is presented in equation (17) and should also be applicable to other 2D topological insulators, such as type II InAs/GaSb QWs [18]. Next, we will use the Foldy–Wouthuysen (FW) transformation to find an effective model describing the electron or heavy hole bands. We show that such an effective model contains two different types of SO interactions; one of them is the well-known Rashba spin–orbit interaction (SOI) induced by the inversion breaking potential in the \( z \)-direction, whereas the other originates from the in-plane potential, and is referred to as the in-plane Pauli term. Although both these terms, for the conduction band, are linear in the wave vector and the spin, they contribute differently to spin transport. The first (Rashba) term does not conserve the \( z \)-component of spin, \( S_z \), causing spin precession, while the in-plane Pauli term conserves \( S_z \). We study the interplay of the Rashba and in-plane Pauli terms. We predict that the spin-Hall conductance will show the precession pattern as a function of the inversion breaking potential in the \( z \)-direction even in the presence of a strong in-plane potential. Further, the strong in-plane potential enhances the spin-Hall conductance generated by the Rashba term, because it partially fixates the direction of the precessing spin. Therefore, the behaviour of the spin transport in asymmetrically doped QWs should be dominated by the Rashba term and it is justified to describe the spin-Hall conductance in the metallic regime through simple effective models for electrons and heavy holes (see equations (30) and (31) of this manuscript), as long as the band gap is nonzero.

The rest of the paper is organized as follows. Section 2.1 gives the derivation of the effective four-band model for a 2D topological insulator with a spatial inversion breaking potential using \( \mathbf{k} \cdot \mathbf{p} \) theory. In section 2.2, we show that the same Hamiltonian can be derived using general symmetry arguments. In section 3, we derive an effective one-band model with competing Rashba-type and in-plane Pauli contributions. Section 4 describes the interplay between both terms within the Landauer–Büttiker formalism. We finish the paper with conclusions.

2. Effective Hamiltonian for HgTe QWs in the presence of the inversion breaking potential in the \( z \)-direction and the in-plane potential

2.1. Derivation of the extended HgTe Hamiltonian within \( \mathbf{k} \cdot \mathbf{p} \) theory

In this section, we will consider the influence of SIA on HgTe/CdTe QWs and derive a corresponding effective \( 4 \times 4 \) model with an out-of-plane (in the \( z \)-direction) inversion breaking potential. Our starting point is the eight-band Kane Hamiltonian \( H_K \), which is described in [10]. The SO split-off bands \( |\Gamma_7, \pm 1/2\rangle \) are far away in energy from the other bands and are not important for the description of the QW. This is in contrast to the bulk case, where SIA terms are nonzero only when the \( \Gamma_7 \) band is taken into account [19].

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Therefore we limit ourselves to the upper $6 \times 6$ block of the Kane model with the basis set of wave functions in the sequence $(1) = |\Gamma_6^-, 1/2\rangle$, $(2) = |\Gamma_6^-, -1/2\rangle$, $(3) = |\Gamma_8^+, 3/2\rangle$, $(4) = |\Gamma_8^+, 1/2\rangle$, $(5) = |\Gamma_8^+, -1/2\rangle$, $(6) = |\Gamma_8^+, -3/2\rangle$ where we use the standard notation with $|\Gamma_6^-, \pm 1/2\rangle$ describing an s-like conduction band, $|\Gamma_8^+, \pm 1/2\rangle$ a p-like light hole band and $|\Gamma_8^+, \pm 3/2\rangle$ a p-like heavy hole band in zinc blende crystal structures [19]. Let us emphasize that the $6 \times 6$ Kane Hamiltonian preserves the bulk inversion asymmetry. We omit here negligible effects of bulk inversion asymmetry introduced by the zinc blende structure of bulk HgTe and CdTe [6].

In the following we always use $|\alpha\rangle$ ($\alpha = 1, 2, \ldots, 6$) to denote the basis set of wave functions. We consider a QW configuration with HgTe layers sandwiched by two CdTe barrier layers along the $z$-direction; hence the parameters of the Kane model $H_K$ have spatial dependence [4]. The matching of wave functions in the $z$-direction for HgTe/CdTe QWs has to be done very carefully because the bulk barrier material CdTe has a normal band structure with $|\Gamma_6^-, \pm 1/2\rangle$ above $|\Gamma_8^+\rangle$ bands while bulk HgTe has inverted band ordering with $|\Gamma_8^+\rangle$ above $|\Gamma_6^-\rangle$ bands [4, 5]. This is exactly the reason for the change from the normal to inverted band structure ordering for the HgTe/CdTe QWs above the critical value of HgTe layer width $d_c = 6.3$ nm (see for example figure 1 in [5]).

The envelope function approximation [20] is applied to solve the eigenproblem of the QW. Since the Kane model preserves inversion symmetry, in order to discuss the SIA, we need to take into account an additional potential $V(\mathbf{r}) = V_0(x, y) + ze\varepsilon z$, where $e > 0$ is the elementary charge and $ze\varepsilon z$ is the inversion breaking potential in the $z$-direction, while $V_0(x, y)$ is the in-plane potential and a possible form will be specified in section 4. Then the full Hamiltonian is

$$\hat{H}_{\text{full}} = H_K(k_\parallel, z) + V(\mathbf{r}).$$

Next we split the Hamiltonian (1) into two parts $\hat{H}_{\text{full}} = H_0 + H'$, where $H_0$ is the Kane Hamiltonian when $k_\parallel = 0$ and is treated as the zero-order Hamiltonian. Explicitly, $H_0$ is given by

$$H_0 = H_K(k_\parallel = 0) = \begin{pmatrix}
T^{(0)} & 0 & 0 & \sqrt{\frac{2}{3}} P \hat{k}_z & 0 & 0 \\
0 & T^{(0)} & 0 & 0 & \sqrt{\frac{2}{3}} P \hat{k}_z & 0 \\
0 & 0 & W^{(0)}_+ & 0 & 0 & 0 \\
\sqrt{\frac{2}{3}} P \hat{k}_z & 0 & 0 & W^{(0)}_- & 0 & 0 \\
0 & \sqrt{\frac{2}{3}} P \hat{k}_z & 0 & 0 & W_-^{(0)} & 0 \\
0 & 0 & 0 & 0 & 0 & W_+^{(0)}
\end{pmatrix},$$

where $\hat{k}_z$ is an operator and the heavy hole bands ($\Gamma_8^+, \pm 3/2$) are completely decoupled from the electron and light hole bands. Here $P = -\frac{\hbar^2}{2m_0} (S[p_x]X)$ is the Kane matrix element between the $\Gamma_6$ and $\Gamma_8$ bands, while the other parameters are given by

$$T^{(0)} = E_c(z) + \frac{\hbar^2}{2m_0} \hat{k}_z (2F(z) + 1) \hat{k}_z,$$

$$W_{\pm}^{(0)} = E_c(z) + \frac{\hbar^2}{2m_0} \hat{k}_z (2\gamma_2(z) \mp \gamma_1(z)) \hat{k}_z,$$

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with \( F(z) = \frac{1}{m_0} \sum_{\Gamma} \frac{|\langle \Gamma | \rho_i | \mu_j \rangle |^2}{E_{\Gamma}(z) - E_{\Gamma}(\gamma)} \) including remote bands \(| \mu_j \rangle \) with \( \Gamma \) symmetry perturbatively. \( E_{\Gamma} \) designates the positions of the conduction/valence band edges and the \( \gamma_i \) are renormalized Luttinger parameters [21]. The axial approximation is adopted [22, 23] in order to keep the in-plane rotation symmetry.

\( H' \) is treated as a perturbation, and is written as

\[
H' = H_0 + V = \begin{pmatrix}
T^{(1)} & 0 & -\frac{p_k}{\sqrt{2}} & 0 & \frac{p_k}{\sqrt{2}} & 0 \\
0 & T^{(1)} & 0 & -\frac{p_k}{\sqrt{6}} & 0 & \frac{p_k}{\sqrt{2}} \\
-\frac{p_k}{\sqrt{2}} & 0 & W^{(1)} & -S^- & R & 0 \\
0 & -\frac{p_k}{\sqrt{6}} & -S^+ & W^{(1)} & C & R \\
\frac{p_k}{\sqrt{6}} & 0 & R^\dagger & C^\dagger & W^{(1)} & S^+_1 \\
0 & \frac{p_k}{\sqrt{2}} & 0 & R^\dagger & S^- & W^{(1)}_1
\end{pmatrix}, \tag{5}
\]

with \( k_\pm = k_x \pm ik_y, C = \frac{\hbar^2}{2m_0} k_\gamma \), \( R = \frac{\gamma k^2}{2m_0}, S_\pm = -\frac{\sqrt{3}\hbar^2}{2m_0} (\gamma_3, k_z) + (\gamma_3, k_z) \), \( T^{(1)} = \frac{\hbar^2}{2m_0} k_\gamma + V \) and \( W^{(1)}_\pm = -\frac{\hbar^2}{m_0} (\gamma_3 \pm \gamma_2) k^2 + V \). Here, \( \gamma_i = (\gamma_3 + \gamma_2)/2 \). \( \kappa \) is the renormalized Luttinger parameter related to the part of the Hamiltonian that is antisymmetric in the components of \( k \). In the original Luttinger model, it was introduced because in the presence of a magnetic field the components of \( k \) do not commute. In our case, it appears because the material parameters are functions of the \( z \)-coordinate.

Now we will generalize the BHZ approach [4] to project the Hamiltonian (1) into the low energy subspace, which can be done in two steps. First, we numerically diagonalize the Hamiltonian \( H_0 \), so that \( H_0 | i \rangle = E_i | i \rangle \), to obtain the eigenenergies \( E_i \) and eigenstates \( | i \rangle \) of the QW. Here the eigenstate \( | i \rangle \) can be expanded in the basis \( | \alpha \rangle \) as \( | i \rangle = \sum_\alpha f_i,\alpha (z) | \alpha \rangle \), where the function \( f_i(z) \) gives the envelope function along the \( z \)-direction for the QW. We use Greek indices to indicate basis functions of the Kane Hamiltonian and Roman indices to denote the subbands. The envelope function components \( f_i,\alpha (z) \) are calculated with the help of the numerical diagonalization of \( H_0 \).

In order to perform the degenerate perturbation calculation, we need to cast the eigenstates of \( H_0 \) into two classes. The first one, denoted as class A, includes the basis wave functions of our final four-band effective model. As shown by BHZ [4], for HgTe/CdTe QWs, it is necessary to take into account the two electron-like subbands \(| E1, \pm \rangle \) and two heavy hole subbands \(| H1, \pm \rangle \), which are expanded explicitly as

\[
| E1, + \rangle = f_{E+,1}(z)|1\rangle + f_{E+,4}(z)|4\rangle, \tag{6}
\]

\[
| H1, + \rangle = f_{H+,3}(z)|3\rangle, \tag{7}
\]

\[
| E1, - \rangle = f_{E-,2}(z)|2\rangle + f_{E-,5}(z)|5\rangle, \tag{8}
\]

\[
| H1, - \rangle = f_{H-,6}(z)|6\rangle. \tag{9}
\]

As pointed out above, for \( H_0 \) the heavy hole bands are decoupled from the electron and light hole bands; therefore the eigenstate \(| H1, +(-) \rangle \) consists only of the basis \(| 3 \rangle \) (|6\rangle) while \(| E1, +(-) \rangle \) is a combination of the basis \(| 1 \rangle \) (|2\rangle) and \(| 4 \rangle \) (|5\rangle). The second class, denoted as class B,
includes the states that need to be taken into account in the following perturbation procedure. Here we consider the first light hole-like subbands \(|LH, \pm\rangle\) and the second and third heavy hole subbands \(|HH2, \pm\rangle\) and \(|HH3, \pm\rangle\), which are written explicitly as

\[
|LH, +\rangle = f_{LH+1}(z)|1\rangle + f_{LH+4}(z)|4\rangle, \tag{10}
\]
\[
|HHn, +\rangle = f_{HHn+3}(z)|3\rangle, \tag{11}
\]
\[
|LH, -\rangle = f_{LH-2}(z)|2\rangle + f_{LH-5}(z)|5\rangle, \tag{12}
\]
\[
|HHn, -\rangle = f_{HHn-6}(z)|6\rangle. \tag{13}
\]

All the other subbands of the QW are neglected here since they are well separated in energy.

Before we go to the next step of the perturbation calculation, it is useful to have a look at the symmetry properties of the relevant states. For Hamiltonian \(H_0\), we have three types of symmetries: the time reversal symmetry \(\mathcal{T}\), the inversion symmetry \(P\) and the in-plane full rotation symmetry \(R_\alpha(\theta)\). For the time reversal operation \(\mathcal{T}\), it is not hard to show that \(|E1, \pm\rangle\) and \(|H1, \pm\rangle\) are Kramer’s partners, i.e. \(\mathcal{T}|E1, +\rangle = |E1, -\rangle\), \(\mathcal{T}|E1, -\rangle = -|E1, +\rangle\), \(\mathcal{T}|H1, +\rangle = |H1, -\rangle\) and \(\mathcal{T}|H1, -\rangle = -|H1, +\rangle\) and we use phase conventions for the envelope functions which yield the relations \(f_{E-2} = f_{E+1}^*\), \(f_{E-5} = -f_{E+4}^*\), \(f_{H-6} = f_{H+3}^*\), \(f_{LH-2} = -f_{LH+1}^*\) and \(f_{HH2-6} = -f_{HH2+3}^*\). Time reversal symmetry relates states with opposite spin to each other; hence when the effective Hamiltonian for one spin is constructed, the Hamiltonian for the opposite spin can be easily obtained through the operation \(\mathcal{T}\). The inversion operation \(P\) defines the parity of each subband, which can greatly simplify the matrix elements in the perturbation procedure below. The parity of the subbands \(|i\rangle\) in the QW is determined by both the envelope function \(f_{i, \alpha}(z)\) and the basis wave function \(|\alpha\rangle\). The parities of the envelope functions can be obtained through numerical calculation [4, 23], and are listed in Table 1. The parities of the basis functions are given by \(P|\Gamma_6^-, \pm 1/2\rangle = -|\Gamma_6^-, \pm 1/2\rangle\), \(P|\Gamma_8^+, \pm\rangle = +|\Gamma_8^+, \pm\rangle\). Thus, the parities of the subbands are \(P|E1\pm\rangle = -|E1\pm\rangle\), \(P|H1\pm\rangle = |H1\pm\rangle\), \(P|LH\pm\rangle = |LH\pm\rangle\), \(P|HH2\pm\rangle = -|HH2\pm\rangle\) and \(P|HH3\pm\rangle = |HH3\pm\rangle\). Due to the in-plane rotation symmetry, the total angular momentum \(J\) along the \(z\)-direction is a good quantum number, which can be used to identify the eigenstates. Since the electron-like subbands have \(J = \frac{1}{2}\), the rotation operator is \(R_z(\theta)|E1\pm\rangle = e^{i\frac{3\pi}{4}}|E1\pm\rangle\), while for the heavy hole subbands with \(J = \frac{3}{2}\), it should be \(R_z(\theta)|H1\pm\rangle = e^{i\frac{3\pi}{4}}|H1\pm\rangle\).

Next, we calculate the effective Hamiltonian of the four states in class A based on quasi-degenerate perturbation theory. All states in classes A and B are eigenstates of Hamiltonian \(H_0\). However when \(H'\) is introduced, they are no longer eigenstates due to mixing between the states of class A and class B. Therefore, treating \(H'\) as a small perturbation, we need to perform a unitary transformation to eliminate the coupling between the states in class A and class B up to the required order. Details of the perturbation procedure can be found in [19], and here we
directly apply the following third-order perturbation formula:

\[ \tilde{H}_{mm'} = E_m \delta_{mm'} + H'_{mm'} + \frac{1}{2} \sum_{l} H'_m H'_{lm'} \left( \frac{1}{E_m - E_l} + \frac{1}{E_{m'} - E_l} \right) \]

\[ - \frac{1}{2} \sum_{l,m''} \left( \frac{H'_m H'_{lm'} H'_{m''m'}}{(E_m - E_l)(E_{m''} - E_l)} + \frac{H'_{mm'} H'_{m''m'} H'_{lm'}}{(E_m - E_l)(E_{m''} - E_l)} \right) \]

\[ + \frac{1}{2} \sum_{l,l'} H'_{ml} H'_{l'l} H'_{lm'} \left( \frac{1}{E_m - E_l} + \frac{1}{E_{m'} - E_{l'}} \right) (14) \]

with

\[ H'_{jk} := \langle f_j | H' | f_k \rangle = \int dz \sum_{\alpha, \beta=1}^{6} f_{j,\alpha}(z)(H')_{\alpha \beta} f_{k,\beta}(z). \] (15)

The summation indices \( m, m', m'' \) are taken from the states in class A \((E1+, E1-, H1+, H1-)\) and indices \( l, l' \) are from the states in class B \((LH+, LH-, HH2+, HH2-, HH3+, HH3-)\). As mentioned above, Greek indices denote entries of the Kane matrix \((5)\). Here we should keep in mind that the order of the matrix elements of \( H' \) in \((14)\) is important, as they may not commute with each other.

The perturbation calculation based on \((14)\) is straightforward but lengthy. The parities of the envelope functions discussed above can be used to reduce the number of matrix elements of \( H' \). For example, the first-order term

\[ \langle E1 + | H' | E1 - \rangle = - \frac{P_{k_+}}{\sqrt{6}} \langle f_{E-,2} | f_{E+,4} \rangle + \frac{P_{k_-}}{\sqrt{6}} \langle f_{E-,5} | f_{E+,1} \rangle + \langle f_{E-,5} | C^\dagger | f_{E+,4} \rangle \]

vanishes completely because all integrands are odd functions of \( z \) (see table 1).

In the four-band basis \((| E1 + \rangle, | H1 + \rangle, | E1 - \rangle, | H1 - \rangle)\), the final effective Hamiltonian is written as

\[ \tilde{H} = \tilde{H}_0 + \tilde{H}_R + V_0(x, y), \] (17)

\[ \tilde{H}_0 = \epsilon(k) I + \begin{pmatrix} M(k) & Ak_+ & 0 & 0 \\ Ak_- & -M(k) & 0 & 0 \\ 0 & 0 & M(k) & -Ak_- \\ 0 & 0 & -Ak_+ & -M(k) \end{pmatrix}, \] (18)

\[ \tilde{H}_R = \begin{pmatrix} 0 & 0 & -iR_0k_- & -S_0k_+^2 \\ 0 & 0 & S_0k_-^2 & iT_0k_+^3 \\ iR_0k_+ & S_0k_+^2 & 0 & 0 \\ -S_0k_-^2 & -iT_0k_-^3 & 0 & 0 \end{pmatrix}, \] (19)

with \( A = A + A_2k^2, M(k) = M - Bk^2, \epsilon(k) = C - Dk^2 \) and \( I \) is the diagonal unit matrix. Here \( k^2 = k_x^2 + k_y^2 \).

We note that \( \tilde{H}_0 \) is equivalent to the BHZ Hamiltonian in \([4]\) if we further omit the \( k \)-dependence of \( A \), by setting \( A_2 = 0 \). We also assume that the reference energy is fixed in the middle of the gap, i.e. \( C = 0 \). Besides the BHZ Hamiltonian, we find a new term \( \tilde{H}_R \), which is
off-diagonal in spin space due to the inversion breaking potential $\mu e\varepsilon_z$. As mentioned above, we have included the subbands $|HH3, \pm\rangle$ in the calculation of the effective parameters $A$, $B$, $D$. However these subbands do not contribute to $\tilde{H}_R$. This is a consequence of the fact that the envelope functions belonging to $|HH3, \pm\rangle$ have parities opposite to the envelope functions of $|HH2, \pm\rangle$; see table 1.

There are three new terms in $\tilde{H}_R$. The first term ($R_0$ term) originates from the second-order perturbation theory and is exactly the electron Rashba term with

$$R_0 = \frac{\text{i}e\varepsilon_z}{3(E_{\text{LH}} - E_{\text{E1}})} \left[ \left( \langle f_{E+,1} | z | f_{\text{LH}+,1} \rangle + \langle f_{E+,4} | z | f_{\text{LH}+,4} \rangle \right) \times \left( \sqrt{6} P \langle f_{\text{LH}+,4} | f_{E+,1}^* \rangle + \sqrt{6} P \langle f_{\text{LH}+,1} | f_{E+,4}^* \rangle + \frac{3\hbar^2}{m_0} \langle f_{\text{LH}+,4} | \langle \kappa, k_z \rangle | f_{E+,4}^* \rangle \right) \right].$$

(20)

Here $(f_{i,\alpha} | O | f_{j,\beta}) = \int dz f_{i,\alpha}^*(z) O f_{j,\beta}(z)$ for an arbitrary operator $O$. The electron Rashba term is linear in $k$ because of the $\frac{1}{2}$ electron spin.

The second term ($T_0$ term) originates from the third-order perturbation and denotes the heavy hole $k^3$ Rashba term with the parameter

$$T_0 = \frac{-i\sqrt{3} e\varepsilon_z \hbar^2}{8m_0^2} (s_1 + s_2 + s_3)$$

(21)

with

$$s_1 = \frac{2}{(E_{\text{E1}} - E_{\text{HH2}})(E_{\text{H1}} - E_{\text{HH2}})} \left[ \left( \sqrt{2} P m_0 \langle f_{\text{H}+,3} | f_{E+,1} \rangle - \sqrt{3} h^2 \langle f_{\text{H}+,3} | (| \gamma_3, k_z \rangle + [\kappa, k_z]) | f_{E+,4} \rangle \right) \times \langle f_{E+,4} | \gamma | f_{\text{HH2}+,3} \rangle \langle f_{\text{HH2}+,3} | z | f_{\text{H}+,3}^* \rangle \right],$$

(22)

$$s_2 = \frac{2}{(E_{\text{E1}} - E_{\text{H1}})(E_{\text{H1}} - E_{\text{LH}})} \left[ \left( \sqrt{2} P m_0 \langle f_{\text{H}+,3} | f_{E+,1} \rangle - \sqrt{3} h^2 \langle f_{\text{H}+,3} | (| \gamma_3, k_z \rangle + [\kappa, k_z]) | f_{E+,4} \rangle \right) \times \left( \langle f_{E+,1} | z | f_{\text{LH}+,1} \rangle + \langle f_{E+,4} | z | f_{\text{LH}+,4} \rangle \right) \langle f_{\text{LH}+,4} | \gamma | f_{\text{H}+,3}^* \rangle \right],$$

(23)

$$s_3 = \frac{4}{(E_{\text{H1}} - E_{\text{HH2}})(E_{\text{H1}} - E_{\text{LH}})} \left[ \langle f_{\text{H}+,3} | z | f_{\text{HH2}+,3} \rangle \langle f_{\text{HH2}+,3} | \gamma | f_{E+,4} \rangle - \sqrt{3} h^2 \langle f_{\text{HH2}+,3} | (| \gamma_3, k_z \rangle + [\kappa, k_z]) | f_{\text{LH}+,4} \rangle \langle f_{\text{LH}+,4} | \gamma | f_{\text{H}+,3}^* \rangle \right].$$

(24)

Since for heavy holes the spin is $3/2$, the change of angular momentum upon a spin-flip is 3, which corresponds to $k^3_{z}$. The third term ($S_0$ term), which is proportional to $k^2$, also comes from the second-order perturbation with the parameter

$$S_0 = -\frac{\sqrt{3} h^2 e\varepsilon_z}{4m_0} \left[ \left( \frac{1}{E_{\text{E1}} - E_{\text{HH2}}} + \frac{1}{E_{\text{H1}} - E_{\text{HH2}}} \right) \langle f_{\text{H}+,3} | z | f_{\text{HH2}+,3} \rangle \langle f_{\text{HH2}+,3} | \gamma | f_{E+,4} \rangle \right.$$

$$+ \left( \frac{1}{E_{\text{E1}} - E_{\text{LH}}} + \frac{1}{E_{\text{H1}} - E_{\text{LH}}} \right) \langle f_{\text{H}+,3} | \gamma | f_{\text{LH}+,4} \rangle \langle f_{\text{LH}+,4} | z | f_{E+,4} \rangle + \langle f_{\text{LH}+,1} | z | f_{E+,1} \rangle + \langle f_{\text{LH}+,4} | z | f_{E+,4} \rangle \right]$$

(25)
Inversion symmetry for a QW with Rotation symmetry or heavy hole subbands (see section by the in-plane potential are also very small. The latter corrections, after folding to the electron \( \tilde{\nabla} \cdot k \) of third or higher order to the element \( (E_A^0 - E_{\tilde{z}}) \) are of higher order in the perturbation than the ones coming from \( \varepsilon_z \). Moreover, they must contain both \( \varepsilon_z = \nabla V_0(x, y) / \epsilon \) and \( \varepsilon_z \) due to the fact that the in-plane field \( \varepsilon_z \) does not break the \( z \rightarrow -z \) inversion symmetry. Furthermore, \( V_0(x, y) \) introduces corrections of third or higher order to the element \( \tilde{H}_0^{(1)}, \) which have the form \[ [k_+, [k_+, V_0(x, y)]]k_- \] and \( \nabla \varepsilon_z \cdot k_z \). These corrections are much smaller than the element \( A k_+ \), which appears already in the first order of perturbation theory in \( \tilde{H}_0 \). Corrections to the diagonal elements of \( \tilde{H}_0 \) induced by the in-plane potential are also very small. The latter corrections, after folding to the electron or heavy hole subbands (see section 3), produce similar contributions to the ones originating from \( \tilde{H}_0 \), but are an order of magnitude smaller due to a large energy separation between main bands and bands that are treated perturbatively. Therefore, we justified that the only significant contribution to \( \tilde{H} \) connected with the in-plane potential comes from the bare diagonal potential \( V_0(x, y) \), as shown in (17).

2.2. Symmetry arguments for the validity of the extended HgTe Hamiltonian

The goal of this subsection is to derive the effective 4 \( \times \) 4 Hamiltonian (17) using the theory of invariants [19]. The theory of invariants states that the Hamiltonian must be invariant under all symmetry operations of the considered system. As discussed in the last section, the system has time reversal symmetry \( T \), space inversion symmetry \( P \) and in-plane rotation symmetry \( R_z(\theta) \). The transformation of the set of basis wave functions under these symmetries for the effective model has been discussed in the last section. The symmetry operations in the matrix form for the basis \( (|E^1+, |H^1+, |E^1-, |H^1-|) \) are given by

- **Time reversal symmetry:** \( T = \Theta K \), where \( \Theta = -i \sigma_2 \otimes 1 \) and \( K \) is the complex conjugate operator.
- **Inversion symmetry:** \( P = -1 \otimes \tau_3 \).
- **Rotation symmetry:** \( R_z(\theta) = e^{i \theta \Sigma_z / 2} \) with \( \Sigma_z = \sigma_3 \otimes (1 + \tau_3) / 2 + \sigma_3 \otimes (3(1 - \tau_3)) / 2 = \sigma_3 \otimes (2 - \tau_3) \)

where the \( \sigma_i \) denote Pauli matrices acting on the spin basis and the \( \tau_i \) represent Pauli matrices acting on the electron or heavy hole subbands.

**Table 2.** Parameters of the effective 4 \( \times \) 4 Hamiltonian, calculated for the QW width 70 Å, at charge density \( n = 2 \times 10^{10} \text{ cm}^{-2} \).

| Parameter | Value |
|-----------|-------|
| \( A \) (nm \( eV \)) | 0.365 |
| \( B \) (nm\(^2\) \( eV \)) | -0.706 |
| \( D \) (nm\(^2\) \( eV \)) | -0.532 |
| \( M \) (eV) | \(-10.09 \times 10^{-3}\) |
| \( R_0/(e\varepsilon_z) \) (nm\(^2\)) | -15.6 |
| \( T_0/(e\varepsilon_z) \) (nm\(^2\)) | -8.91 |
| \( iS_0/(e\varepsilon_z) \) (nm\(^4\)) | -2.10 |

This is a new off-diagonal term between the electron-like \( \frac{1}{2} (-\frac{1}{2}) \) and heavy hole \( -\frac{3}{2} \) \( \frac{3}{2} \) states and the change of the angular momentum is 2, corresponding to \( k_z \). All parameters can be determined by using the numerically calculated \( f_{i,a}(z) \), and are listed in table 2 for a QW with a thickness \( d_0 = 70 \text{ Å} \).

We note that all three terms in the inversion breaking Hamiltonian \( \tilde{H}_R \) are proportional to \( \varepsilon_z \) but are independent of \( V_0(x, y) \). The corrections originating from \( V_0(x, y) \) to the 2 \( \times \) 2 off-diagonal blocks in (19) are of higher order in the perturbation than the ones coming from \( \varepsilon_z \).
Table 3. Summary of the symmetry properties of $\Gamma$ matrices and the tensors formed by $k$.

| $R_z$ | $T$ | $P$ |
|-------|-----|-----|
| $\mathbf{I}$ | 0 | + | + |
| $(\Gamma_1, \Gamma_2)$ | 2 | − | − |
| $(\Gamma_4, \Gamma_3)$ | 1 | − | − |
| $(\Gamma_{15}, \Gamma_{25})$ | 2 | + | − |
| $(\Gamma_{45}, \Gamma_{35})$ | 1 | + | − |
| $\Gamma_3$ | 0 | + | + |
| $\Gamma_{12}$ | 0 | − | + |
| $\Gamma_{34}$ | 0 | − | + |
| $(\Gamma_{14} + \Gamma_{23}, \Gamma_{31} + \Gamma_{24})$ | 1 | − | + |
| $(\Gamma_{23} - \Gamma_{14}, \Gamma_{31} - \Gamma_{24})$ | 3 | − | − |
| $(k_i, k_i)$ | 1 | − | − |
| $(k_i^2 + k_i^7)$ | 0 | + | + |
| $(k_i^3 - 3k_i^7, 2k_i k_i)$ | 2 | + | + |
| $(k_i^3 - 3k_i^7, 3k_i^2 k_i - k_i^7)$ | 3 | − | − |
| $(k_i^3 + k_i^7, k_i^2 k_i + k_i^7)$ | 1 | − | − |

Generally, any $4 \times 4$ Hamiltonian can be expanded using $\Gamma$ matrices as

$$\hat{H}_{\text{eff}} = \epsilon(k)\mathbf{I} + \sum_i d_i(k)\Gamma_i + \sum_{ij} d_{ij}(k)\Gamma_{ij},$$

(26)

where $\mathbf{I}$ is the $4 \times 4$ identity matrix, $\Gamma_i$ ($i = 1, \ldots, 5$) denote five $\Gamma$ matrices, which satisfy $\{\Gamma_i, \Gamma_j\} = 2\delta_{ij}$, and the ten commutators of $\Gamma$ matrices $\Gamma_{ij} = [\Gamma_i, \Gamma_j]/2i\epsilon(k)$, $d_i(k)$ and $d_{ij}(k)$ can be expanded as polynomials of the momentum $k$. The Hamiltonian should be invariant under the symmetry operations $P$, $T$ and $R_z$, which indicates that $d_{ij}(k)$ should behave as the same as $\Gamma_{i(j)}$. Therefore we need to work out the transformation form of the $d_{ij}(k)$ and the $\Gamma$ matrices.

We construct the $\Gamma$ matrices as follows:

$$\begin{align*}
\Gamma_1 &= \sigma_1 \otimes \tau_1, & \Gamma_2 &= \sigma_2 \otimes \tau_1, & \Gamma_3 &= \sigma_3 \otimes \tau_1, \\
\Gamma_4 &= 1 \otimes \tau_2, & \Gamma_5 &= 1 \otimes \tau_3, & \Gamma_{ij} &= \epsilon_{ijk}\sigma_k \otimes 1,
\end{align*}$$

(27)

where $i, j = 1, 2, 3$. It is easy to prove that $(\Gamma_{ab})^2 = 1$, $\{\Gamma_a, \Gamma_{ab}\} = 0$ and $\Gamma_{ab}, \Gamma_{ac}$ for $b \neq c$. For the above $15$ $\Gamma$ matrices, it is easy to calculate the symmetry transformation under the time reversal operation $T$ and inversion operation $P$, which are listed in table 3. For the in-plane rotation operation $R_z(\theta)$, we can calculate the transformation rule $\Gamma'(\theta) = e^{i\frac{\theta}{2}}\Gamma e^{-i\frac{\theta}{2}}$ with the help of the differential equation $\frac{d\Gamma'(\theta)}{d\theta} = \frac{1}{2}[\Sigma, \Gamma'(\theta)]$. The obtained results are also given in table 3, from which we find that $\Gamma_{55}$, $\Gamma_{12}$ and $\Gamma_{34}$ behave as a scalar under the rotation $R_z$, $(\Gamma_4, \Gamma_3)$, $(\Gamma_{45}, \Gamma_{35})$ and $(\Gamma_{14} + \Gamma_{23}, \Gamma_{31} + \Gamma_{24})$ rotate as a vector with angular momenta $1$, $(\Gamma_1, \Gamma_2)$, $(\Gamma_{15}, \Gamma_{25})$ corresponds to angular moment $2$, and $(\Gamma_{23} - \Gamma_{14}, \Gamma_{31} - \Gamma_{24})$ corresponds to angular moment $3$ in Table 3, we also list the corresponding tensors formed by $k$ up to the order $k^3$. From table 3, if we hope to preserve $T$, $P$ and rotation symmetry, then up to $k^3$ the general form of the Hamiltonian is given by

$$H_0 = \epsilon_k + \mathcal{M}(k)\Gamma_5 + \mathcal{A}(k)(\Gamma_4, \Gamma_3)
\begin{pmatrix}
\cos \theta & \sin \theta \\
-\sin \theta & \cos \theta
\end{pmatrix}
\begin{pmatrix} k_x \\ k_y \end{pmatrix},$$

(28)
where \( \mathcal{A}(k) = A + A_2 k^2 \), \( \mathcal{M}(k) = M - B k^2 \) and the phase \( \theta \) represents the relative phase between \(|E1\rangle\) and \(|H1\rangle\), which can be chosen arbitrarily. Taking \( \theta = -\pi/2 \), in (28), we recover the BHZ Hamiltonian [4].

We now consider additional terms that preserve rotation symmetry and time reversal symmetry, but break the inversion symmetry. By inspecting table 3, the following three terms are possible:

\[
\tilde{H}_R = \frac{\mathcal{R}(k)}{2} (\Gamma_{14} + \Gamma_{23}, \Gamma_{31} + \Gamma_{24}) \left( \begin{array}{cc} \cos \phi & \sin \phi \\ -\sin \phi & \cos \phi \end{array} \right) \left( \begin{array}{c} k_x \\ k_y \end{array} \right) \\
+ \frac{T_0}{2} (\Gamma_{23} - \Gamma_{14}, \Gamma_{31} - \Gamma_{24}) \left( \begin{array}{cc} \cos \psi & \sin \psi \\ -\sin \psi & \cos \psi \end{array} \right) \left( \begin{array}{c} k_x^3 - 3k_x k_y^2 \\ 3k_y^2 k_x - k_y^3 \end{array} \right) \\
+ S_0 (\Gamma_{15}, \Gamma_{25}) \left( \begin{array}{cc} \cos \varphi & \sin \varphi \\ -\sin \varphi & \cos \varphi \end{array} \right) \left( \begin{array}{c} k_x^2 - k_y^2 \\ 2k_x k_y \end{array} \right),
\]

(29)

where \( \mathcal{R}(k) = R_0 + R_2 k^2 \). Similar to \( \theta \), the phase factors \( \phi, \varphi \) and \( \psi \) are also arbitrary and will not affect the energy spectra. If we choose \( \theta = -\pi/2, \phi = -\pi/2, \psi = \pi/2 \) and \( \varphi = \pi/2 \), then the Hamiltonian (17) is recovered. This ensures that the derivation in section 2.1 has yielded the Hamiltonian with the correct structure.

3. Foldy–Wouthuysen transformation of the effective HgTe Hamiltonian

The goal of this section is to obtain an effective \( 2 \times 2 \) Hamiltonian (where 2 stands for the spin degree of freedom) for electron \(|E1\rangle\) and heavy hole \(|H1\rangle\) subbands including nonzero in-plane and out-of-plane electric fields. So far, the quantum spin-Hall effect (QSHE) was described by the Hamiltonian used in [4], where only the diagonal blocks of our Hamiltonian (17) were taken to be nonzero, i.e. for \( \varepsilon_z, \varepsilon_{||} = 0 \). Such a block-diagonal Hamiltonian of a HgTe QW is isomorphic to the Dirac Hamiltonian describing the relativistic motion of an electron in two dimensions \( (\hat{p}_z = 0) \), which couples particle and antiparticle components with the same spin direction. Here, we start from the full Hamiltonian (17) and consider the low-energy physics with the energy scale smaller than the gap \( 2M \). In this case, we can apply the perturbation formula (14) to obtain an effective model for electron and hole subbands. This procedure is equivalent to the FW transformation [24], which reduces the relativistic Dirac equation in a potential to the Pauli equation [19]. We keep terms up to linear order in the in-plane \( \varepsilon_z \) and out-of-plane \( \varepsilon_{\parallel} \) electric fields, as well as terms up to the third order in \( k \). Then the effective Hamiltonians for electron (\( \tilde{H}_e \)) and hole (\( \tilde{H}_h \)) subbands are given by

\[
\tilde{H}_e = M + V_0(x, y) + \left( -D - B + \frac{A^2}{2M} \right) k^2 + \frac{A^2}{8M^2} e\nabla\varepsilon_{\parallel} - \mathcal{R}(k)(\sigma \times k)_z + \mathcal{G}(k)(e\varepsilon_{\parallel} \times k)_z \sigma_z,
\]

(30)

\[
\tilde{H}_h = -M + V_0(x, y) + \left( B - D - \frac{A^2}{2M} \right) k^2 + \frac{A^2}{8M^2} e\nabla\varepsilon_{\parallel} + \frac{1}{2} (\mathcal{Q}(k)\sigma_x k_z^2 + \mathcal{Q}(k)^\dagger \sigma_- k_x^2)
\]

\[
- \mathcal{G}(k)(e\varepsilon_{\parallel} \times k)_z \sigma_z + \frac{1}{2} \left( \frac{A S_0}{2M^2} [\sigma_-, [\sigma_-, V_0]] \sigma_+ + \text{h.c.} \right),
\]

(31)
Figure 1. Material parameters characterizing the SO coupling induced by (a) the in-plane field \( (A^2/4M^2) \) and (b) the out-of-plane field \( (R_0/(e\varepsilon_z)) \) as a function of the QW thickness \( d \). (c) Energies of the most relevant subbands as a function of the QW width for the \( \Gamma \) point.

with

\[
G(k) = \frac{A^2}{4M^2},
\]

\[
\mathcal{R}(k) = R_0 + \left( \frac{iAS_0}{M} - \frac{A^2}{4M^2} R_0 \right) k^2 + \frac{iAS_0}{2M^2} e\nabla \varepsilon_{||},
\]

\[
\mathcal{Q}(k) = iT_0 + \frac{AS_0}{M} + \frac{iA^2 R_0}{4M^2}.
\]

The spin-dependent term \( g(e\varepsilon_{||} \times \mathbf{k}) \sigma_z \) and the spin-independent term \( (A^2/8M^2)e\nabla \varepsilon_{||} \) originate directly from the FW transformation from the Dirac-type Hamiltonian in the external potential \( \tilde{H}_0 + V_0(x, y) \) (see equation (17)) to a Pauli-type equation. Therefore, by analogy to the relativistic electron in vacuum, we call \( g(e\varepsilon_{||} \times \mathbf{k}) \sigma_z \) the in-plane Pauli term, whereas we call \( (A^2/8M^2)e\nabla \varepsilon_{||} \) the Darwin term. Note that the Pauli term can also be visualized as resulting from a Rashba field due to the edges of a typical mesa structure used in experiments or as coming from the atomic SO splitting but it is only active at the edges where \( \varepsilon_{||} \) is finite. The Darwin term does not include a contribution from the field in the \( z \)-direction due to the assumption that \( \varepsilon_z \) is constant. The in-plane Pauli and Darwin terms appear in both the electron and hole effective Hamiltonians. The additional terms that are proportional to \( \mathcal{R}(k) \) and \( \mathcal{Q}(k) \) originate from \( \tilde{H}_R \) and are a direct consequence of the broken space inversion symmetry in the \( z \)-direction. These terms are usually called Rashba terms and they give linear and cubic in \( k \) contributions for electron and heavy hole subsystems, correspondingly. In a typical experimental setup, Rashba terms are generated by an asymmetric doping profile surrounding the QW and can be adjusted by a top gate which induces a tunable electric field in the \( z \)-direction. Figures 1(a) and (b) show the magnitude \( A^2/4M^2 \) of the in-plane SOI and electron Rashba coefficient \( R_0/(e\varepsilon_z) \) as a function of the thickness \( d \) of the HgTe/CdTe QW. Note that the coupling strength \( A^2/4M^2 \) for the Pauli term decreases with \( d \), whereas the strength of Rashba coupling \( R_0/(e\varepsilon_z) \) increases with \( d \). The origin of the different behaviour of these two SOIs can be understood from the
plot of energy versus $d$ (see figure 1(c)). The in-plane term $A^2/4M^2$ comes from the coupling between the electron and the heavy hole $|H1\pm\rangle$ subbands, and the energy difference between these bands increases with $d$. The Rashba term $R_0/(e\varepsilon_z)$ originates from the coupling between the electron and the light hole $|L1\rangle$ subbands. Their energy difference decreases with $d$; therefore $R_0/(e\varepsilon_z)$ increases with $d$. Comparing the magnitudes of $R_0/(e\varepsilon_z)$ and $A^2/4M^2$, one can see that close to the critical thickness $d = d_c$ determining the transition from a normal to a topologically nontrivial insulator, the magnitude of the in-plane term is an order of magnitude larger than the SOI term in the $z$-direction, while for $d = 80 \text{ Å}$ the magnitudes of both interactions are comparable.

4. Spin transport within an effective electron model

As described in detail in the previous section, the effective conduction band description of a HgTe QW (30) includes two different SOI terms: the Rashba SO coupling and the in-plane Pauli term. To understand the interplay of both SOIs, we will analyse here the spin-Hall conductance signal numerically using the tight binding version of the Hamiltonian equation (30) within the Landauer–Büttiker formalism [26].

4.1. Description of the model

The Rashba field $\varepsilon_z$ can be applied constant in space and varied in strength easily by an external top gate. By contrast, $\varepsilon_{||}$, generating the in-plane Pauli term, usually originates from impurities or from confinement due to sample boundaries. In our calculations we work in the quasiballistic regime, which is very well justified for HgTe/CdTe QWs with the typical mobilities $1-5 \times 10^5 \text{ cm}^2\text{ V}^{-1}\text{ s}^{-1}$. Consequently, the contribution of the impurities to the in-plane field is negligible and the confining potential is dominant, i.e. $|\varepsilon_{x,y}|$ decreases with the distance from the sample edges. The confining potential requires that an electric field at the boundary always points outside the sample, i.e. its magnitude changes sign at opposite edges.

We use two different setups. In both setups a finite size sample with SOIs (see figures 2(a) and (b)) is attached to four semi-infinite leads of the same width $w$. For the first setup, a square sample, the in-plane field is introduced by tunnelling barriers between the leads and the conductor along the blue lines in figure 2(a). This simple setup has the advantage that the numerical results are easy to interpret, but is not a realistic description of actual experiments. The second setup is a symmetric cross structure, which resembles the experimental Hall bars and is shown in figure 2(b). In this case again the blue lines indicate the sample border, where a confining potential is present (for a form of this potential, see also figure 2(c)).

We construct the confining potential in 1D $V_0(x, y)$ from the one dimensional profile

$$V_{1D}(t) = e^{-t/l} + e^{-(w-t)/l},$$

(35)

where the coordinates in both the quadratic and the cross-shaped samples are chosen such that 0, respectively $w$, mark the $x$- or $y$-coordinates at the edges of the central square. $l$ is the characteristic decay length of the potential. For the square-shaped sample, we define

$$V_0(x, y) = c_{sq} (V_{1D}(x) + V_{1D}(y)).$$

(36)

The maximal field is then $\varepsilon_x^0 = -\partial_x V_0(x, y)|_{x=0}$. We adjust the constant $c_{sq}$ to choose some particular value $\varepsilon_x^0$. For the cross-shaped sample, we use the definition

$$V_0(x, y) = c_{cr} V_{1D}(\min(\max(x, 0), w)) V_{1D}(\min(\max(y, 0), w)).$$

(37)
Figure 2. The two different samples we use for the numerics: (a) a quadratic Hall bar of width $w$ and (b) a cross structure. In both cases the sample (blue) contains SOI, whereas the four semi-infinite leads (black) are SOI free. The numbering of the leads and the used boundary conditions on the currents are indicated in the figure, i.e. we drive a current $I$ from leads 1 to 2 and do not allow for charge currents at the vertical leads 3 and 4. The discretization is shown for a part of each sample. Along the blue lines the confining potential $V_0(x, y)$ is applied, as indicated in the inset of (b), to give rise to the in-plane Pauli interaction. In the case of the quadratic sample, this potential corresponds to a tunneling barrier between leads and sample. (c) Shows the confining potential normalized by its maximal value ($V_0^{\text{max}}$) for the cross structure.

Here the constant $c_{cr}$ is adjusted to choose the desired maximal field $E_x^0 = -\frac{\partial}{\partial x} V_0(x, y)|_{x=y=0}$ and we assume that $E_y^0 = E_x^0$. Further, in the leads the potential is always set as zero. For clarification, the confining potential corresponding to (37) is shown in figure 2(c). We find that the numerical results do not change qualitatively if the boundary field is defined differently, as long as the characteristic decay length $l$ is unchanged. In both setups we assume that the leads do not include SOI, and therefore an analytical form of the eigenmodes [27] and a clear definition of the spin current are available [28].

In the calculations, we set the boundary conditions on the currents $I_1 = -I_2 = I$ and $I_3 = I_4 = 0$, where $I_p = I_p^1 + I_p^\downarrow$ is the total current at lead $p$. The spin-dependent current, $I_p^\sigma$, is calculated by use of the Landauer–Büttiker formula [26, 28, 29]

$$I_p^\sigma = \frac{e}{h} \sum_{q \neq p} \sum_{\sigma' = \uparrow, \downarrow} \left[ T_{qp}^{\sigma\sigma'} \mu_p - T_{pq}^{\sigma\sigma'} \mu_q \right],$$

(38)

which links the spin-resolved current to the chemical potential $\mu_p = eV_p$ via the transmission matrix elements $T_{pq}^{\sigma\sigma'}$. $T_{pq}^{\sigma\sigma'}$ describes the probability that an electron, entering the sample at lead $p$ with spin $\sigma$, will leave the sample through lead $q$ having spin $z$-projection $\sigma'$. For a sample with nonzero SO coupling, applying an electric field between leads 1 and 2 will generate a transverse spin current $I_p^2 = -\frac{\hbar}{2e}(I_p^\uparrow - I_p^\downarrow)$ at leads 3 and 4, which is the so-called spin-Hall effect [28, 30, 31]. We define the spin-Hall conductance as follows:

$$G_p^\sigma = \frac{I_p^\sigma}{V_1 - V_2} = \frac{\hbar}{2e} \frac{I_p^\uparrow - I_p^\downarrow}{V_2 - V_1} = \frac{e}{4\pi} \sum_{\sigma' = \uparrow, \downarrow} \left( T_{p1}^{\uparrow\sigma'} - T_{p1}^{\downarrow\sigma'} \right), \quad p = 3, 4,$$

(39)
where $V_2 - V_1$ is the voltage difference between leads 2 and 1. Due to the symmetries of the Hamiltonian, only a few transmission matrix elements are independent, so that the last equality in (39) follows. The time reversal symmetry of (30) implies $T_{pq}^{\alpha\sigma'} = T_{qp}^{-\alpha'\sigma}$, the fourfold rotational symmetry $c_4$ of the setup about the $z$-axis implies e.g. $T_{23}^{\sigma\alpha'} = T_{14}^{-\sigma\alpha'}$, and the mirror symmetry with respect to the $yz$-plane implies $T_{32}^{\sigma\alpha'} = T_{31}^{-\sigma\alpha'}$.

The transmission matrix elements are computed numerically in a tight binding approach by using the Green’s function method [26]–[28], [31], and the Fisher–Lee relation [32] connecting the Green’s function with the transmission amplitudes. We discretize the sample as indicated in figure 2. By making use of the fermionic field operators $c_{\alpha,\sigma}^\dagger$ ($c_{\alpha,\sigma}$), which create (annihilate) a spin $\sigma$ electron at lattice site $\alpha$, the Rashba and in-plane Pauli interactions take the following form in second quantization:

$$\hat{H}_{\text{Rashba}} = \frac{R_0}{2a} \sum_{\alpha} \left[ i c_{\alpha,\uparrow}^\dagger c_{\alpha+\alpha_x,\downarrow} + i c_{\alpha,\downarrow}^\dagger c_{\alpha+\alpha_x,\uparrow} - c_{\alpha,\uparrow}^\dagger c_{\alpha+\alpha_x,\downarrow} - c_{\alpha,\downarrow}^\dagger c_{\alpha+\alpha_x,\uparrow} + c_{\alpha,\dagger} c_{\alpha+\alpha_x,\dagger} + \text{h.c.} \right],$$

$$\hat{H}_{\text{Pauli}} = \frac{A^2 e^2}{8 M^2} \sum_{\alpha,\sigma} \left[ \epsilon_{\alpha}^{a\uparrow}/\epsilon_{\alpha}^{a\downarrow} c_{\alpha,\sigma}^\dagger c_{\alpha+\alpha_x,\sigma} - \epsilon_{\alpha}^{a\downarrow}/\epsilon_{\alpha}^{a\uparrow} c_{\alpha,\sigma}^\dagger c_{\alpha+\alpha_x,\sigma} - \text{h.c.} \right] \kappa_\sigma,$$

where $R_0/(\epsilon_{\alpha})$, $A$ and $M$ are the material parameters defined in section 2.1. Here $a$ denotes the lattice constant, and $\alpha_x,\alpha_y$ stand for the lattice unit vectors connecting nearest neighbours. To obtain the parameter $R_0$ we must assume a specific value of the perpendicular field $\tilde{E}_z$, $\tilde{E}_x^\sigma$ and $\tilde{E}_y^\sigma$ designate, for a given site $\alpha$, the in-plane electric field components. $\kappa_\sigma = \pm 1$ for spin up and down, respectively, and the symbol h.c. denotes the Hermitian conjugate.

The Rashba SOI does not conserve the $z$-component of the spin and thus leads to spin precession [33]. By contrast, the in-plane Pauli term (41) conserves the $z$-component of the spin, causing a shift in energy for two spin directions. This energy shift, however, must not be mistaken as the Zeeman effect, because (41) does not break time reversal symmetry.

For numerical calculations we consider the Hamiltonian

$$\hat{H} = \hat{T} + \hat{H}_{\text{SO}} + \hat{H}_{\text{Dis}} + \hat{V}_0(x, y) + \hat{H}_{\text{Dar}}.$$

Here $\hat{T} = (-D - B + A^2/(2M)) k^2 = \hbar^2 k_x^2/(2m^*)$ describes the kinetic part of the conduction band Hamiltonian (30). In second quantization, $\hat{T}$ is described by spin-conserving nearest-neighbour hopping [26]–[28]. $\hat{H}_{\text{Dis}} = \text{diag}(\epsilon_i)$ specifies the disorder of the sample, where the diagonal on-site energies $\epsilon_i$ are uniformly distributed between $[-W/2, W/2]$ [34]. The disorder strength $W = \hbar e/(m^* \mu)$ is calculated from the mobility $\mu$. The confining potential is taken into account via $\hat{V}_0(x, y)$. The SO coupling $\hat{H}_{\text{SO}}$ is described by the Rashba term (40), the in-plane Pauli term (41) or a linear superposition of both terms. Hence $\hat{H}$ mirrors the conduction band Hamiltonian (30), where we omit the $k$ dependence of the $\mathcal{R}$ parameter and negligible terms that include the combined effect of the in-plane and out-of-plane electric fields. The spin-independent Darwin term $\hat{H}_{\text{Dar}} \propto \mathbf{V} \cdot \hat{\mathbf{E}}_{||}$ breaks the particle–hole symmetry of the tight binding Hamiltonian, just like any space-dependent in-plane potential would do. Here particle–hole symmetry means the relation $G_{\alpha}(E_l) = -G_{\alpha}^*(\pm E_l)$ if the energy zero point is chosen in the middle of the tight binding band. It originates from the cosine dispersion relation of a free electron on a lattice. $\hat{H}_{\text{SO}}$ does not break this symmetry. The Darwin term does not qualitatively change the spin conductance signal and will be considered after the SO terms are analysed.

We use realistic parameters for the calculations, which are shown in table 2. Here, we assume a thickness of the QW in the $z$-direction of 7 nm, corresponding to the inverted
regime. Although this has no impact on our one-band approach, it guarantees a large coupling strength $R_0/(e\varepsilon_c)$. The carrier density is set to $n = 2 \times 10^{10}$ cm$^{-2}$, whereas the effective mass $m^* = 0.00712m_0$, where $m_0$ is the bare electron mass. We want to point out that we have changed the originally negative sign of the effective mass to achieve a positive curvature of the band dispersion, which is justified in a one-band model. In our calculations, we focus on the regime where the Fermi energy is deep enough in the conduction band so that the spin edge states have already merged into the bulk [6, 25] and the spin-Hall conductance that is the property of the states close to the Fermi level will be dominated by the bulk response.

Assuming a quadratic dispersion around $k = 0$, we determine the Fermi energy to be $E_f = 6.73$ meV. Finally, the assumed mobility $\mu = 25 \times 10^4$ cm$^2$ (V s$^{-1}$) leads to $W = 0.65$ meV. For such a small disorder strength, averaging over ten different disorder configurations is sufficient. We note that the mean spin-Hall conductance deviates from that of a clean sample only by about $\lesssim 1\%$. The parameters are chosen carefully to the restricted range of possible energies where this model is valid, i.e. $E_i < E_f < E_{\text{gap}}$, where $E_{\text{gap}} = 2M$ is the energy gap and $E_i$ is the energy splitting of the band due to Rashba interaction.

We will first focus on the square sample because the scattering barrier in this setup allows us to study the competition between the in-plane and out-of-plane electric fields. The influence of in-plane electric fields is much weaker in the cross structures. Further, the first minimum in the spin-Hall conductance generated by the Rashba contribution is shifted away from the spin precession length to smaller fields due to quantum interference effects in the vertical stubs (see e.g. [27]) and therefore the results are less transparent to interpret.

### 4.2. Numerical results for the quadratic sample

We choose a quadratic sample of width $w = 1000$ nm, which is discretized by 200 x 200 lattice points, so that the Fermi wavelength is about 36 times the lattice constant. The characteristic length scale of the electric field is assumed to be $l = 10$ nm. The computed spin-Hall conductance is presented in figure 3. First, we focus on pure SOI and consequently omit the Darwin term and the potential in figures 3(a) and (b). Figure 3(a) shows $G_3^z(\varepsilon_0)$ for different top gate fields and figure 3(b) presents $G_3^z(\varepsilon_c)$ for different in-plane fields.

#### 4.2.1. Rashba coupling

The spin-Hall conductance signal induced by the Rashba coupling alone, $\varepsilon_\parallel = 0$, is shown in figure 3(b) by blue circles. For a small interaction strength $R_0$, the spin-Hall conductance rises quadratically, saturates and finally starts to precess. The behaviour of the spin-Hall conductance originating from the Rashba model can be understood by the spin force operator

$$\hat{F}_H = -\frac{m^*}{\hbar^2}[[\hat{r}_H, \hat{H}], \hat{H}] = \frac{2m^*R_0^2}{\hbar^3}(\hat{p}_H \times z) \otimes \hat{\sigma}_H^z,$$

where $\hat{r}_H$, $\hat{p}_H$ and $\hat{\sigma}_H^z$ are the position, momentum and spin-operators in the Heisenberg representation (see Nikolić et al [33]) and $z$ is a unit vector. In this simple picture the force acting on electrons due to SO coupling is quadratic in $R_0$, explaining the behaviour of $G_3^z$ as a function of the out-of-plane electric field for low $\varepsilon_c$. The force described in (43) deflects the spin-$\uparrow$ and the spin-$\downarrow$ electrons in opposite transverse directions, leading to the spin-Hall effect. However, the Hamiltonian does not conserve the $z$-component of the spin, leading to a rotation of the spin direction and as a consequence to oscillations in $G_3^z$ as a function of $R_0$. The first maximum of the spin-Hall conductance is reached, when the spin has travelled a distance...
Figure 3. Spin-Hall conductance due to the presence of a superposition of Rashba and in-plane Pauli terms in a quadratic sample with tunneling barriers at the boundaries. Due to the presence of the second interaction, the starting value of spin-Hall conductance can be nonzero. (a) When \( \varepsilon_z \) is nonzero, the magnitude of \( G^3_{\sigma} \) as a function of \( \varepsilon^0_{x,y} \) is reduced. (b) Spin-Hall conductance as a function of \( \varepsilon_z \). For higher values \( \varepsilon^0_{x,y} \) the precession amplitude of \( G^3_{\sigma} \) at \( L_{SO} \) (indicated by a red line) is increased and the minimum is slightly shifted to higher interaction energies. (c) and (d) show the same dependences of \( G^3_{\sigma} \) as above, but here the Darwin term and the confining potential were additionally taken into account.

4.2.2. In-plane Pauli interaction. The spin-Hall conductance shows linear behaviour as a function of the in-plane electric field (see the blue circles in figure 3(a), where only the \( \varepsilon_{x,y} \) components are nonzero). This linear dependence on field strength can be explained within the semiclassical approach, where we have adopted the wave packet dynamics by Sundaram and Niu [35] to obtain equations of motion for the in-plane Pauli term:

\[
\dot{r}_c = \frac{1}{\hbar} \frac{\partial \epsilon}{\partial k_c} - \hat{k}_c \times \Omega_{\sigma}, \tag{44a}
\]

\[
\hbar \dot{k}_c = -e \varepsilon_{||}, \tag{44b}
\]

with the magnetic field set to zero. Here the index \( c \) denotes the center coordinate of the wave packet in position and \( k \) space. The Berry curvature is defined as

\[
\left( \Omega_{\sigma}^\pm \right)_{\alpha} := -\varepsilon_{\alpha\beta\gamma} \text{Im} \left[ \frac{\partial u_\sigma^\pm(k)}{\partial k_\beta} \frac{\partial u_\sigma^\pm(k)}{\partial k_\gamma} \right], \tag{45}
\]
where symbol $\pm$ corresponds to two eigenvalues $E_{\pm} = -Dk^2 \pm \sqrt{A^2k^2 + M^2(k)}$ of the upper spin block of the Hamiltonian $\hat{H}_0$ in (18) with $\sigma_{\pm}(k)$ being corresponding eigenstates for spin $\sigma$. Equation (44b) simply describes the change of the lattice momentum due to the electric field. Equation (44a) describes the time evolution of the position operator due to the band dispersion and the anomalous velocity term $\mathbf{k} \times \Omega_z$ with the spin-dependent nonzero $z$-components $(\Omega^\pm_z)_z = -(\Omega^\pm_z)_z = \pm A^2/2M^2$. The spin-dependent anomalous velocity term shifts the position of the wave packet with different spins in two opposite transverse directions, leading to the spin-Hall effect. Inserting (44b) into (44a) yields the dependence of the anomalous velocity term linear in $\epsilon_1$. Note that the energy range ($\approx 0.0111\tau$) plotted in figures 3(a) and (b) is the same for both interaction terms. Since the coupling parameters can be quite different (in our calculations $21 \times R_0/(e\epsilon_z) \approx A^2/(4M^2)$), the magnitudes of in-plane and out-of-plane electric fields are adjusted so the interaction energies are the same.

4.2.3. Interplay of both interactions. A linear superposition of (40), (41) and $\hat{T}$ leads to the spin signal, which is shown in figures 3(a) and (b), when all three field components are nonzero. The finite value of the spin-Hall conductance in figure 3(a) for $\epsilon_{xy}^0 = 0$ is due to the Rashba coupling. It can be observed that the linear behaviour of the in-plane Pauli term with the electric field is not changed, when Rashba SO coupling is present. However, the slope of the spin-Hall conductance curves decreases with $\epsilon_z \neq 0$, which means that the in-plane Pauli contribution to the spin signal is suppressed by the Rashba interaction. The $z$-component of the spin is not conserved for finite $\epsilon_z$, as can be seen from equation (40). The resulting spin precession implies that generation of a spin current by the anomalous velocity becomes less effective. The smallest slope and therefore smallest in-plane Pauli contribution in figure 3(a) is found for an electric field $\epsilon_z$ corresponding to the precession length, where an expectation value $\langle \sigma_z \rangle$ rotates by $\pi$.

Figure 3(b) shows $G_3^z$ as a function of $\epsilon_z$ for different in-plane electric fields. One can see the typical precession pattern of $G_3^z(\epsilon_z)$ also for $\epsilon_{\parallel} \neq 0$. Moreover, the precession amplitude of the spin-Hall conductance of the Rashba type is enhanced in the presence of the in-plane Pauli term. The origin of this increase can be traced back to the $k$-dependent energy splitting of the spin subbands due to the in-plane Pauli interaction. In order to lower its energy, the electron now prefers to stay in either spin up or down states. The precession of the spin is thus slightly suppressed, as can be seen in a small shift of the minima to higher electric fields. As discussed above, the spin force operator (43) can act more efficiently on electrons with a preferred spin $z$-projection, which leads to the relatively higher magnitude of the spin-Hall conductance caused by Rashba coupling.

Figures 3(c) and (d) also show $G_3^\parallel(\epsilon_\parallel)$ and $G_3^\perp(\epsilon_z)$ respectively, but now with included potential $V_0(x,y)$ and Darwin $A^2/(8M^2)\nabla \cdot \epsilon_\parallel$ terms. In the tight binding approach, both terms renormalize the diagonal on-site energy. They cannot generally be considered to be small, as they scale, like the in-plane Pauli term, with the magnitude and the shape of the confining electric field. The relative magnitudes of the in-plane Pauli, Darwin and potential terms depend on the choice of the functional dependence $V_0(x,y)$. The most important scale is the characteristic length scale $l$, over which the corresponding field drops to $e\epsilon^0/e$. We have performed numerical calculations with different values $\epsilon_{xy}^0$ and $l$, and found that the main features of the results discussed in this paper stay the same. We found that the Darwin term is the main contribution which renormalizes the spin-Hall conductances in figures 3(c) and (d) with respect to $G_3^\parallel(\epsilon_\parallel)$ and $G_3^\perp(\epsilon_z)$ in figures 3(a) and (b).
Figure 4. Spin-Hall conductance for a cross structure, with the Darwin and the potential terms included. In (a) and (b), we have used the same range of the electric field as for the quadratic sample, i.e. $l = 10$ nm. Although the in-plane field was enhanced to very high values, the influence of the in-plane electric field on $G_3^z(E_z)$ is very weak. In (c) and (d), we have increased $l$ to 40 nm to get a higher in-plane Pauli signal. In this case $G_3^z$ behaves similarly as for the quadratic sample.

The divergence of $E_\parallel$ appears as an additional term in the semiclassical equation (44b), and therefore the spin-independent Darwin term can contribute to the anomalous velocity and renormalize the spin-Hall conductance term. This can be seen in figure 3(c) as nonlinear behaviour of $G_3^z(E_\parallel)$. However, in the range of in-plane electric fields shown in figures 3(c) and (d), the qualitative behaviour of the spin-Hall conductance is the same as in the absence of the Darwin term (see figure 3(a) and (b)). Increasing the in-plane electric field to the same magnitude as the electric field perpendicular to the 2DEG has two difficulties. First of all the interaction energy of the in-plane Pauli term exceeds the Fermi energy, which marks the limit of validity of our effective electron model. Secondly, increasing $E_\parallel$ comes along with raising the tunnelling barrier in the quadratic sample. We can omit these difficulties by choosing the sample in the shape of a cross (see the next subsection).

4.3. Numerical results for the cross sample

The cross sample is made up of five square parts: four stubs and the central square (see figure 2(b)). Each part has width $w = 500$ nm and is discretized by $100 \times 100$ lattice sites. The corresponding spin-Hall conductance originating from in-plane Pauli and Rashba terms in the presence of the Darwin term and the confining potential is shown in figures 4(a)–(d).

Figures 4(a) and (b) show $G_3^z(E_\parallel)$ and $G_3^z(E_z)$ for different values of fields in the $z$ and in-plane directions correspondingly and for the characteristic range of the electric field $l = 10$ nm. These figures should be compared with figures 3(c) and (d) correspondingly. The overall
behaviour of spin-Hall conductance in figures 4(a) and (b) is similar to the quadratic sample. However, although we used values of the in-plane electric field around 20 times larger than in figure 3(d), the influence of \( E_\parallel \) on \( G_3^S(E_z) \) is much weaker than for the square structure (compare 4(b) with 3(d)). The probability of scattering from vertical and horizontal walls in the cross structure is much smaller than in the case of square structure where the electron directly hits the wall. Therefore, we find that the influence of an in-plane electric field will be much weaker in the experimentally relevant cross samples.

In order to obtain a contribution to \( G_3^S \) due to the confining potential comparable to that for the quadratic sample, we have increased \( l \) from 10 to 40 nm, in figures 4(c) and (d). This leads to a larger range of lattice sites, which can contribute efficiently to the spin-dependent hopping. In the case of larger \( l \), the results resemble those discussed in the last section for square samples. In figure 4(d) the amplitude of the Rashba contribution to the spin-Hall effect counted to the first minimum is enhanced by the confinement potential. By contrast, the signal due to the in-plane Pauli term decreases until \( E_z \) corresponds to the first minimum, as can be seen in figure 4(c). The magnitude of the slope increases slightly for higher values of the top gate field, but with an inverted sign with respect to the quadratic sample (see figure 3(c)).

At the end of this section let us emphasize that for the experimentally relevant case both interactions are present, but only \( E_z \) can be easily varied, e.g. by a top gate. Therefore, in the experimentally relevant case, the presence of both the in-plane Pauli and the Darwin interactions could lead to an increase of the amplitude of spin signal.

5. Conclusions

We have derived an extension to the BHZ Hamiltonian for the typical 2D topological insulator (HgTe QWs) in the presence of the inversion breaking potential in the \( z \)-direction and in-plane potential. For the derivation, we used two independent methods: \( \mathbf{k} \cdot \mathbf{p} \) perturbation theory and symmetry arguments based on Clifford algebra. We found that to the third order in the perturbation theory, only the inversion breaking potential in the \( z \)-direction generates a new off-diagonal in spin space terms. These terms lead to the Rashba SOI when the FW transformation to the effective electron model is performed. On the other hand, the diagonal-in-spin space part of the Hamiltonian in the presence of the in-plane potential generates an additional term to the one-band model that is also linear in momentum and spin, but conserves the \( z \)-component of the spin. By analogy with the equation for a relativistic electron in vacuum, we call this term the in-plane Pauli term. The presence of both terms in the conduction band Hamiltonian leads to interesting behaviour of the spin-Hall conductance. In particular, the in-plane Pauli contribution to the spin-Hall conductance is suppressed in the presence of the spin precession inducing terms. By contrast, the spin-Hall conductance from the Rashba term preserves the oscillation pattern in the presence of the in-plane Pauli term and its magnitude can be enhanced due to partial pinning of the \( z \)-component of the spin. This latter situation is experimentally relevant since the inversion breaking potential in the \( z \)-direction can be easily tuned by a top gate in experiments. Therefore, we expect that in experiments on asymmetrically doped HgTe/CdTe QWs [11] in the metallic regime (the Fermi level in the conduction or valence band), the behaviour of spin transport and especially the spin-Hall conductance will be dominated by the Rashba SOI. Note that in our derivation we omit the BIA terms since they have already been studied in [6].
Let us also emphasize that our effective four-band Hamiltonian in the presence of inversion breaking potential is not limited only to the HgTe/CdTe QWs and can be easily generalized to other topological insulators such as type II InAs/GaSb/AlSb QWs \cite{18} with correctly adjusted strengths of the Rashba SOIs.

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