Reconsidering power functional theory

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ABSTRACT
The original derivation of power functional theory [M. Schmidt and J. M. Brader, J. Chem. Phys. 138, 214101 (2013)] is reworked in some detail with a view to clarifying and simplifying the logic and making explicit the various functional dependencies. We note various issues with the original development and suggest a modification that allows us to avoid them. In the process, we also suggest an alternative interpretation of our results, which bears surprising similarities to classical density functional theory.

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I. INTRODUCTION

Classical Density Functional Theory (cDFT)\textsuperscript{1,2} has proven to be a powerful tool in the study of inhomogeneous classical systems. In his seminal review article that helped to define the field, Evans also discussed a dynamic extension that has come to be known as Dynamical Density Functional Theory (DDFT), which is applicable to particles obeying an over-damped Brownian dynamics. These two are intimately related since the free energy functionals that play a central role in cDFT are used in DDFT to describe the deterministic driving force governing the evolution of the local density.\textsuperscript{3–5} While cDFT is based on mathematically exact theorems—and as such represents a formally equivalent reformulation of some aspects of classical statistical mechanics—DDFT is heuristic, depending, in all of its various derivations, on a kind of uncontrolled, local-equilibrium approximation. It would obviously be desirable if the paradigm of DFT could be extended to non-equilibrium systems, thus giving a formally equivalent theory that, like cDFT, admits of simple but highly useful approximations.

This is precisely the goal of Power Functional Theory (PFT) as first proposed by Schmidt and Brader\textsuperscript{6} (hereafter referred to as SB) for particles subject to an over-damped Brownian dynamics and subsequently extended by Schmidt and co-workers over the last several years to quantum,\textsuperscript{7} Newtonian,\textsuperscript{8} and active-particle\textsuperscript{9} systems. The basic idea of the various developments is to introduce a functional, the minimization of which generates the time-dependent many-particle distribution function for the system. A version of constrained search (described below) is used to project thisonto a functional that depends only on the local density (as in cDFT) and the local current, where the latter is the key to accessing time-dependent quantities. Minimization with respect to these fields yields the actual time-dependent density and current. A significant development inspired by this work, but not directly dependent on it, has been the study of so-called “non-adiabatic” dynamics—basically the difference between some non-equilibrium dynamics followed via simulation and the predictions of DDFT.\textsuperscript{10,11} This has provided insight into the limits of DDFT (which itself has become increasingly popular; see, e.g., the recent review of te Vrugt, Löwen, and Wittkowski\textsuperscript{12}) and some heuristic approaches to modeling these differences. Recently, PFT has also been used to motivate the models of phase coexistence in active particles.\textsuperscript{13–15}

In the following, we review the development of PFT using a notation that is constructed to make particularly evident the functional dependencies that can get lost with a more standard, and less detailed, notation. The necessary calculations are presented in a fairly explicit manner to avoid ambiguities as good as possible. In doing so, certain inconsistencies will be noted that, taken together, throw doubt on the validity of the framework. A modification aimed at overcoming these difficulties is proposed and yields new insights into the possibility of fulfilling the program proposed by SB. Section II of this paper sets the stage with a brief review of Brownian dynamics and also introduces the functional notation. Section III summarizes the key points of the exact PFT formalism, including the generating functional for the time-dependent N-body
distribution function, its projection onto the density–current sub-
space, and some of the structural elements stressed by SB, and in
the processes, various problems in the derivation are highlighted. In
Sec. IV, we present our modification of the original PFT of SB, which
allows us to complete their program with surprising thoroughness.
In the process, we note an alternative interpretation of our results,
which may provide a conceptual reformulation of Brownian dynam-
is that is closer to the goals of SB than was our original development.
This paper ends with a brief summary of our results.

II. BROWNIAN DYNAMICS

The starting point is a system of N identical, classical particles
of mass \( m = 1 \) in D dimensions for which the \( i \)th particle has coor-
dinates \( \mathbf{q}_i^{(0)} \) and velocities \( \mathbf{v}_i^{(0)} \). The particles interact via a
conservative potential \( U(\mathbf{q}_1^{(0)}, \ldots, \mathbf{q}_N^{(0)}) \), which will be written
more briefly as \( U(\mathbf{q}^{(0)}) \), where \( \mathbf{q}^{(0)} \) is the collection of all \( N \) coor-
dinates. In the following, a slightly compressed notation will be used
whereby the position (and time) arguments are written as subscripts so
that the potential will be written as \( U_{\mathbf{r}} \). The particles also experi-
ence a (possibly) time-dependent external one-body potential \( \phi_{\mathbf{r}} \).
Again, what is written here as \( \phi_{\mathbf{r}} \) would more conventionally be writ-
ten as \( \phi(\mathbf{r}, t) \), and in the following, the time argument will always
precede the position argument(s) in the subscript notation. Finally, there are
a stochastic force and corresponding friction proportionality to the velocity
(repre-
senting, e.g., a bath of much smaller particles) so that the equa-
tions of motion are as follows:

\[
\frac{d}{dt} \mathbf{q}_i^{(0)} = \mathbf{v}_i^{(0)},
\]
\[
\frac{d}{dt} \mathbf{v}_i^{(0)} = -\gamma \mathbf{v}_i^{(0)} - \nabla U_{\mathbf{r}} - \sum_{j=1}^{N} \nabla \phi_{\mathbf{r}}^{(0)} + \mathbf{\xi}_i^{(0)},
\]  

(1)

where a caret (hat) indicates a stochastic variable, \( \gamma \) is the
friction, and \( \mathbf{\xi}_i^{(0)} \) are \( D \)-dimensional vectors whose components
are Gaussian-distributed white noise with correlations \( \langle \mathbf{\xi}_i^{(0)}(t) \xi_j^{(0)}(t') \rangle = 2 \gamma k_b T \delta(t - t') \delta_1 \) (where 1 is the unit tensor for the Cartesian
components). Following DDFT, the original PFT was developed for the
over-damped limit in which the time-derivative of the velocity
can be neglected (this can be justified rigorously with certain scal-
ing assumptions; see, e.g., Ref. 16), giving the so-called Brownian
dynamics

\[
\frac{d}{dt} \mathbf{q}_i^{(0)} = \mathbf{v}_i^{(0)},
\]
\[
\frac{d}{dt} \mathbf{v}_i^{(0)} = -\gamma \mathbf{v}_i^{(0)} - \frac{1}{m} \sum_{j=1}^{N} \nabla \phi_{\mathbf{r}}^{(0)} + \frac{1}{m} \mathbf{\xi}_i^{(0)}.
\]  

(2)

The state of the system is entirely specified by the \( N \times D \) phase space
coordinates \( \mathbf{q}_i^{(0)} \equiv \mathbf{q}_i^{(0)} \). In the following, it will be useful
to note that the one-body term can be written as

\[
\sum_{i=1}^{N} \nabla \phi_{\mathbf{r}}^{(0)} = \sum_{i=1}^{N} \nabla \left( \sum_{j=1}^{N} \phi_{\mathbf{r}}^{(0)} \right) = \sum_{i=1}^{N} \nabla \phi_{\mathbf{r}}^{(0)},
\]  

(3)

where the last equivalence defines the total external potential, \( \phi_{\mathbf{r}}^{(0)} \).
The probability to find the system in a given state, \( \mathbf{r}^{(N)} \), at time \( t \) is

\[
\Psi_{\mathbf{r}^{(N)}} = \left\{ \delta(\mathbf{r}^{(1)} - \mathbf{q}_i^{(1)}) \ldots \delta(\mathbf{r}^{(N)} - \mathbf{q}_i^{(N)}) \right\},
\]  

(4)

where the brackets indicate an average over the noise, \( \mathbf{q}_i^{(1)} \), and the
distribution of initial conditions. That \( \Psi_{\mathbf{r}^{(N)}} \) is the distribu-
tion function is evident since the expectation value with respect to the noise
of any function \( f_{\mathbf{q}_i^{(0)}} \) of the stochastic variables \( \mathbf{q}_i^{(0)} \) can be calculated as

\[
\langle f_{\mathbf{q}_i^{(0)}} \rangle = \int f_{\mathbf{q}_i^{(0)}} \Psi_{\mathbf{r}^{(N)}} d\mathbf{r}^{(N)},
\]  

(5)

and as usual,\(^\dagger\), the Brownian dynamics implies that the dis-
tribution satisfies the Fokker–Planck equation

\[
\frac{\partial}{\partial t} \Psi_{\mathbf{r}^{(N)}} = \frac{1}{2} \sum_{i=1}^{N} \nabla \cdot \left( \nabla \left( U_{\mathbf{r}} + \phi_{\mathbf{r}}^{(0)} \right) + k_b T \nabla \right) \Psi_{\mathbf{r}^{(N)}}.
\]  

(6)

Clearly, the distribution at time \( t \) is completely determined by the
interaction potential \( U \), the external field \( \phi \), and an initial condition
that will be denoted by \( \Psi_{\mathbf{r}^{(N)}} \). As is to be expected, the Boltzmann
form \( \Psi_{\mathbf{r}^{(N)}} \sim \exp(-\beta(U_{\mathbf{r}} + \phi_{\mathbf{r}}^{(0)})) \) is a stationary solution if the field
is stationary.

In the following, square brackets will be used to indicate the
parental dependence so that to be precise and to indicate the full dependence
of the distribution on the interaction poten-
tial, initial condition, and the external field, one should write \( \Psi_{\mathbf{r}^{(N)}}[U, \phi, \Psi_{\mathbf{r}^{(N)}}] \). Note that the spatial arguments are not
indicated in the square brackets because, in principle, to know \( \Psi_{\mathbf{r}^{(N)}} \)
at positions \( \mathbf{r} \) requires knowing the various functional inputs (e.g., \( U, \phi \),
and \( \Psi_{\mathbf{r}^{(N)}} \)) at all points in space. That said, in the following, the
dependence on the inter-atomic potential \( U \) will not be explicitly indicated
since it is ubiquitous, never changes, and plays no role in the discus-
sion below. We also note that SB included non-conservative forces in
their analysis although the same could be done here, we have chosen to
omit them for the sake of clarity as they do not change any of the arguments
to follow and only serve to complicate the expres-
sions. We note for later purposes that, in this notation, SB gave the
Fokker–Planck equation in the form

\[
\frac{\partial}{\partial t} \Psi_{\mathbf{r}^{(N)}} = -\frac{1}{2} \sum_{i=1}^{N} \nabla \cdot \left( \psi_{\mathbf{r}^{(0)}}(\phi_i, \Psi_{\mathbf{r}^{(N)}}) \psi_{\mathbf{r}^{(N)}} \right),
\]  

(7)

with the “velocities” \( \psi_{\mathbf{r}^{(0)}} \) and related “forces” \( \mathbf{F}_{\mathbf{r}^{(0)}} \) defined as

\[
\gamma \psi_{\mathbf{r}^{(0)}}(\phi, \Psi) = \mathbf{F}_{\mathbf{r}^{(0)}}^{\text{total}}(\phi, \Psi) = -\nabla U_{\mathbf{r}} - \nabla \phi_{\mathbf{r}}^{(0)} - k_b T \nabla \ln \Psi_{\mathbf{r}^{(N)}}.
\]  

(8)

The local number density (the probability to find a particle in a
given infinitesimal volume) is evaluated as

\[
\rho_t[\Psi] = \sum_{\mathbf{r}^{(N)}} \delta(\mathbf{r}^{(N)} - \mathbf{r}^{(N)}) \Psi_{\mathbf{r}^{(N)}},
\]  

(9)

and when the distribution—and so the density—is time-dependent,
the number current can be defined via the continuity equation as follows:

\[
\frac{\partial}{\partial t} \rho_t[\Psi] = -\nabla \cdot \mathbf{J}_t[\phi, \Psi_{\mathbf{r}^{(N)}}],
\]  

(10)

which, together with the Fokker–Planck equation, gives the explicit
expression

\[
\mathbf{J}_t[\phi, \Psi_{\mathbf{r}^{(N)}}] = \sum_{\mathbf{r}^{(N)}} \delta(\mathbf{r}^{(N)} - \mathbf{r}^{(N)}) \psi_{\mathbf{r}^{(N)}}(\phi, \Psi_{\mathbf{r}^{(N)}}) d\mathbf{r}^{(N)}.
\]  

(11)
An important distinction that is highlighted by this notation is that between the *intrinsic* time dependence and *inherited* time dependence. The inherited time dependence refers to that arising simply because a functional depends on another time-dependent functional. For example, defining the trivial functional \( I_r[\Psi] = \psi \), its evaluation with a time-dependent input, e.g., \( I_r[\psi_t] \), has an inherited time dependence coming solely from its dependence on the time-dependent field. Here, the functional \( I \) only changes in time because its argument changes in time, and otherwise, the time argument is a passive label. An example here is the current \( J_r[\Psi, \phi] \) for which the time dependence is inherited from its arguments and which has no other source of time dependence. On the other hand, the distribution \( \Psi_{\text{tot}}[\psi; \Psi_0] \) has an *intrinsic* time dependence since it is the solution to the Fokker–Planck equation so that even if the external field is constant in time, \( \Psi \) will still change in time if the initial condition is not the equilibrium distribution. On the other hand, when the external field does change with time, the distribution \( \Psi_t \) depends on the values of the external field \( \psi_{\tau t} \) at all times \( \tau < t \), and for this reason, one writes \( \Psi_{\text{tot}}[\psi; \Psi_0] \) (with \( \phi \) rather than \( \psi_t \)) because the functional acts on both the spatial coordinate and the time.

This leads to a final caveat that is important throughout the analysis done by SB and below: namely, the role of causality. The Fokker–Planck equation can formally be solved as

\[
\Psi_{\text{tot}}[\psi; \Psi_0] = \Psi_{\text{tot}}^{0} \cdot \frac{1}{\gamma} \int_0^t \left\{ \sum_{r=1}^N \nabla_i \cdot \left( \frac{2}{\gamma} \nabla \ln \Psi_{\text{tot}}^{0} \cdot \nabla \Psi_{\text{tot}}^{0} \right) \right\} dt',
\]

where \( \Psi_{\text{tot}}^{0} \) represents an initial value that has to be specified. Notice that in the argument of \( \Psi \) on the left hand side, we write \( \phi \) rather than \( \psi_t \) because the right hand side depends on the external field at all times prior to \( t \). SB specified that the integral on the right will be understood to exclude the end point at time \( t \) so that \( t_0 \leq t' < t \), which can be interpreted as saying that the value of the field at time \( t \) is fully determined by its values at previous times corresponding, physically, to the usual understanding of causality. When it is important below to note this dependence on a field at prior times, but not on the present time, we will write, e.g., \( \Psi_{\text{tot}}^{0} [\phi; \Psi_0] \).

### III. POWER FUNCTIONAL THEORY

#### A. Variational formulation

The analysis by SB begins with a quantity modeled on the Rayleigh dissipation function evaluated at a fixed time, \( t \),

\[
\bar{R}_r[\bar{\phi}^N; \psi_t, \phi_t, \Psi_t] = \sum_{\gamma} \left( \frac{\bar{\gamma}^{(i)}_r}{2} \Psi_{\text{tot}}^{0} - \Psi_{\text{tot}}^{0}[\phi_t, \Psi_t] \right) \cdot \bar{\phi}_{\gamma^0}^{(i)} + \phi_{\gamma^0}, \tag{13}
\]

where \( \bar{\phi}_{\gamma^0}^{(i)} \) is a collection of \( N \) test fields (i.e., variational fields), each a function of the \( N \)-positions \( r^N \). This is used to define the functional

\[
R[\bar{\phi}^N; \psi_t, \phi_t, \Psi_t] = \int_{\mathbb{R}^N} \bar{R}_r[\bar{\phi}^N; \psi_t, \phi_t, \Psi_t] d^N r, \tag{14}
\]

which has the obvious property that its absolute, or global, minimum with respect to the test fields \( \bar{\phi}^N \) occurs at

\[
\bar{\gamma}_{\phi}^{(i)}[\phi, \Psi_t] = \Psi_{\text{tot}}^{0}[\phi_t, \Psi_t] - \nabla_i \ln \Psi_{\text{tot}}^{0} - k_B T \nabla_i \ln \Psi_{\text{tot}}^{0} \tag{15}
\]

corresponding to the “physical” fields at this fixed time \( t \), which occur in the Fokker–Planck equation as written in Eq. (7). SB then expressed the variational fields in terms of a new quantity, a variational distribution \( \bar{\Psi}_{\phi} \), via the definition

\[
\bar{\gamma}_{\phi}^{(i)} \rightarrow \bar{\gamma}_{\phi}^{(i)}[\phi, \bar{\Psi}] = \Psi_{\text{tot}}^{0}[\phi, \bar{\Psi}] \tag{16}
\]

in terms of which the functional \( R \) becomes (see Appendix A)

\[
R[\bar{\Psi}, \phi_t, \phi_t, \Psi_t] = (k_B T)^2 2y \int \sum \left( \nabla_i \ln \Psi_{\text{tot}}^{0} \right) \Psi_{\text{tot}}^{0} d^N r
+ \frac{1}{2} \frac{\partial}{\partial t} \Lambda[\phi_t, \psi_t] + \frac{1}{2} \int \phi_{\gamma^0}^{(i)} \Psi_{\text{tot}}^{0} d^N r, \tag{17}
\]

where the dot notation indicates a time-derivative with

\[
\Lambda[\phi_t, \psi_t] = \int (k_B T \Psi_{\text{tot}}^{0} \nabla \ln \Psi_{\text{tot}}^{0} + (U_{\text{tot}} + \psi_t \Psi_{\text{tot}}^{0})) d^N r. \tag{18}
\]

Minimizing the generating function \( R \) with respect to \( \bar{\Psi} \) clearly gives \( \bar{\Psi}_{\text{global–min}}^{\phi} = \Psi_t \), the physical distribution. As observed by SB, the quantity \( \Lambda \) plays a central role in cDFT and is the same functional used by Mermin\(^\text{16}\) to establish its fundamental theorems. Here, it plays no role in the minimization procedure and only serves to establish the value of the generating function at its minimum. Note that at this point, one has not gained much yet: the variational velocity field in Eq. (14) and the variational distribution in Eq. (17) have been introduced as extra fields, and the functional still depends on the physical but unknown distribution \( \Psi_{\text{tot}}^{0} \).

#### B. Power functional

Next, SB split the minimization of \( R \) with respect to \( \bar{\Psi} \) into two steps as

\[
\min_{\bar{\Psi}} R[\bar{\Psi}; \psi_t, \phi_t, \Psi_t] = \min_{\gamma_0} \left\{ \min_{\bar{\Psi}} \left( \bar{\Psi}; \psi_t, \phi_t, \Psi_t \right) \right\}, \tag{19}
\]

where the expression in curly brackets means that \( R \) is first minimized with respect to the subset of possible fields \( \bar{\Psi} \) that satisfy \( \rho_t[\bar{\Psi}] = \bar{\rho}_t \) and \( \bar{J}_r[\phi_t, \bar{\Psi}] = \bar{J}_r \) for any given fields \( \bar{\rho}_t \) and \( \bar{J}_r \). The first (inner) minimization defines a new functional

\[
\bar{R}[\bar{\rho}_t, \bar{J}_r; \phi_t, \psi_t, \Psi_t] = \min_{\bar{\Psi}} R[\bar{\Psi}; \psi_t, \phi_t, \Psi_t] \tag{20}
\]

and also defines one or more fields, giving the minimum, \( \bar{\psi}_{\text{min}}[\bar{\rho}_t, \bar{J}_r; \phi_t, \psi_t, \Psi_t] \), so that

\[
\bar{R}[\bar{\rho}_t, \bar{J}_r; \phi_t, \psi_t, \Psi_t] = R[\bar{\psi}_{\text{min}}[\bar{\rho}_t, \bar{J}_r; \phi_t, \psi_t, \Psi_t]; \phi_t, \psi_t, \Psi_t], \tag{21}
\]
In terms of the variational problem, it is important that \( \vec{r}_t \) and \( \vec{J}_t \) are arbitrary fields and that even though the current corresponding to \( \Psi_r \), \( \vec{J}_t = [j_r, \vec{\Psi}] \), carries an inherited time dependence via the external potential, this is irrelevant to the choice of \( \vec{J}_t \). Instead, the time dependence of the current manifests itself in that the resulting functional \( R \) carries two dependencies on \( \phi_r \): the explicit dependence that comes via \( \Lambda \), and now, an implicit dependence that comes from minimizing under the constraint \( \vec{J}_t = [j_r, \vec{\Psi}] \). For example, a toy model for \( R \), one that is perfectly consistent and illustrates this difference, is

\[
R[\vec{r}_t, \vec{J}_t; \phi_r, \vec{\Psi}] \equiv \int \left\{ (\vec{r}_t - \rho_t[\vec{\Psi}(r)])^2 + (\vec{J}_t - \vec{J}_t[\phi_r, \vec{\Psi}])^2 \right\} dt + \frac{1}{2} \frac{\partial}{\partial t} \Lambda[\phi_r, \vec{\Psi}] + \frac{1}{2} \int \phi_r(\vec{r}) \rho_t[\vec{\Psi}(r)] dt \quad (22)
\]

showing the explicit field dependence in \( \Lambda \) and an implicit field dependence in \( \vec{J}_t \). This toy model minimizes (in the second step) to give the correct, physical results \( \vec{r}_t = \rho_t[\vec{\Psi}] \) and \( \vec{J}_t = \vec{J}_t[\phi_r, \vec{\Psi}] \), and this, in turn (presumably), forces \( \rho_{t, \phi_r} = \rho_t[\vec{\Psi}] \). This happens with no additional information as is the expected result of the two-step—or “constrained search”—minimization procedure since this must ultimately lead to the same result as the one-step minimization, namely, the exact solution \( \rho_{t, \phi_r} = \rho_t \).

Following SB, the second step in the definition of the power functional is the introduction of the quantity,

\[
\Psi_{t, r, \phi_r} \equiv \Psi_{t, r, \phi_r} = \left( \int_0^t f \left( \nabla \phi \min \rho_{t', r, \phi_{t', r}} \right) dt' \right) dt', \quad (23)
\]

which is a many-particle distribution fulfilling the continuity equation if the velocities of the particles follow from the minimal trial distribution. The first line is written as in SB [Eq. (18)], whereas the second is in our extended notation, showing that the minimum for the velocity fields \( \psi_{t, r} \) is determined by \( \Psi_{t, r, \phi_r} \) through Eq. (16). Notice first that a time dependence has now been assigned to the constraints \( \vec{r}_t, \vec{J}_t \). This means that they are being specified not at a single moment but over some range of times \( t \leq t \leq T \). Second, and more problematic, notice that because \( \rho_t[\vec{\Psi}] \) is a result of Eq. (20), it has a dependence on the exact distribution \( \Psi_t \) evaluated at the same time \( t' \), and so \( \Psi_{t, r, \phi_r} \) depends on the exact distribution at all earlier times. While this appears to be the implication of SB equations (11), (14), and (18), it is at odds with subsequent developments in SB since the point of this step is to eventually eliminate the exact distribution from this problem. What seems to eventually be adopted is the elimination of the exact solution \( \Psi_t \) from the functional through replacing \( \Psi_t \) by \( \Psi_{t, r, \phi_r} \), giving the final form of the power functional as follows:

\[
\Psi_{t, r, \phi_r} = \min_{\Psi_{t, r, \phi_r}} R[\vec{r}_t; \vec{J}_t; \phi_r, \vec{\Psi} \psi_{t, r, \phi_r}] = \min_{\Psi_{t, r, \phi_r}} \left( \Psi_{t, \phi_r} \Psi_{t, \phi_r} \right) \Rightarrow \Psi_{t, \phi_r} = \min_{\Psi_{t, \phi_r}} R[\vec{r}_t; \vec{J}_t; \phi_r, \vec{\Psi}], \quad (24)
\]

The second line presents no difficulties in that the evaluation of \( \Psi_{t, \phi_r} \) requires \( \Psi_{t, \phi_r} \) at times \( t' > t \), and this, in turn, only requires \( \Psi_{t, \phi_r} \) at earlier times. At this point, one can simplify the notation since everything at time \( t \) is now defined in terms of \( \vec{r}, \vec{J} \) at earlier times, so we will write a little more compactly

\[
\Psi_{t, r, \phi_r} = \Psi_{t, r, \phi_r} = \left( \int_0^t \left( \nabla \phi \min \rho_{t', r, \phi_{t', r}} \right) dt' \right) dt', \quad (25)
\]

The density implied by this \( \Psi_{t, r, \phi_r} \) is now easily evaluated from the last of these with the result

\[
\rho_{t, \phi_r} = \rho_{t, \phi_r} \rho_{t, \phi_r} = \left( \int_0^t \left( \nabla \phi \min \rho_{t', r, \phi_{t', r}} \right) dt' \right) dt', \quad (26)
\]

which follows because, by definition, \( \rho_{t, \phi_r} \) is precisely the distribution that implies the current \( \vec{J}_t \) while minimizing the generating function. So, the implied density \( \rho_{t, \phi_r} \) satisfies a continuity equation with respect to the current \( \vec{J}_t \), and this is clearly not the case. Notice that, at this point, the temporal evolution of the fields \( \vec{r}_r \) and \( \vec{J}_t \) is completely independent. Indeed, one also sees that in general, \( \vec{J}_t \vec{r}_r = \vec{J}_r \vec{r}_r \), i.e., the one-particle current from the distribution \( \Psi \) is not the constraint current. This turns out to be important.

C. Problems interpreting the power functional

In SB, one of the main results is that the power functional has a Legendre structure with, e.g., the density \( \vec{r}_r \) and \( \vec{J}_t \) playing the role of conjugate variables, thus mirroring the relation between the density and the external field in equilibrium DFT. In order to verify this, we now turn to an examination of the power functional with the aim of making the dependencies on the field at time \( t \) explicit. Starting from the SB functional [Eq. (19)], performing the first minimization with the replacement \( \Psi_{t, r, \phi_r} \rightarrow \Psi_{t, r, \phi_r} \) (as described) and following the steps detailed in Appendix A give
\[ \mathcal{R}_{\text{i}}^{(\text{SB})}[\bar{\rho}, \bar{\phi}, \phi] = \left( \frac{k_B T}{2} \right) \int \Psi_{\text{ne}}[\bar{\rho}, \bar{\phi}, \phi] \times \left( \sum_i \left( \nabla_i \ln - \frac{\Psi_{\text{min}}}{\alpha_{\text{ne}}} \right)[\bar{\rho}, \bar{\phi}, \phi] \right)^2 dr^N \]

\[ - \frac{1}{2\nu} \int \Psi_{\text{ne}}[\bar{\rho}, \bar{\phi}, \phi] \left( \sum_i F_{\text{ne}}^{(i)}[0, \Psi_t] \right)^2 dr^N \]

\[ - \frac{1}{2\nu} \int \rho_t[\Psi_t](\nabla \phi)^2 dr + \int \phi_t \partial_t \rho_t[\Psi_t] dr. \tag{27} \]

Here, all explicit dependence on the external potential at time \( t \) has been exposed. SB argued that this can be written as [see SB equation (25)]

\[ \mathcal{R}_{\text{i}}^{(\text{SB})}[\bar{\rho}, \bar{\phi}, \phi] = W_t^{(\text{SB})}[\bar{\rho}, \bar{\phi}, \phi] - \int \Psi_t \cdot (\nabla \phi_t) dr + \int \phi_t \partial_t \rho_t dr, \tag{28} \]

where \( W_t^{(\text{SB})} \) does not depend on the external field at time \( t \) but only at earlier times. This form is crucial later on in SB for identifying a truly intrinsic functional of the dissipated power, similar to the intrinsic free energy functional in equilibrium DFT. Our form (27) reveals three difficulties with that interpretation. The first is that in Eq. (27), one cannot guarantee that \( \Psi_{\text{min}}^{\text{ne}} \) does not depend on the external field at time \( t \). Indeed, the global minimum of \( \mathcal{R} \) certainly does [see Eq. (15)], so it is entirely possible that the constrained minimum \( \Psi_{\text{min}}^{\text{ne}} \) does as well. The second is that in Eq. (27), one sees an explicit, quadratic dependence on the field at time \( t \), which is ignored in SB. This seems to be because of the SB switch between the representation of the variational field as \( \Psi \) and the original formulation in terms of \( \phi \), and in terms of the latter, there is no quadratic term. However, formulating everything in terms of the densities is not possible at this point as the definition of \( \Psi_{\text{ne}} \) [see Eq. (24) and (25)] involves both \( \Psi_{\text{min}}^{(Q)} \) and \( \rho_t \), which cannot be replaced by \( \Psi_{\text{min}}^{(G)} \) and \( \rho_t \), which cannot. The third problem is that the term involving the gradient of the field in the last line of Eq. (27) is written in terms of \( \Psi_t[0, \bar{\Psi}_t] \) and \( \rho_t[\Psi_t] \) rather than \( \Psi_t[\Psi_t, \phi_t] = \Psi_t \) and \( \rho_t[\Psi_t] = \rho_t \), as tacitly assumed in SB. This means, e.g., that SB equation (26),

\[ \delta \rho \mathcal{R}_{\text{i}}^{(\text{SB})}[\bar{\rho}, \bar{\phi}, \phi] = \bar{\rho}_t, \tag{29} \]

in which \( \mathcal{R}_{\text{i}}^{(\text{SB})} \) is meant to act as a generator for the density, does not hold here. Finally, we observe that several of these problems can be traced to the fact that \( \Psi_{\text{ne}} \) and \( \Psi_t \) are not the same quantities—a fact that we will exploit below to try to repair these problems.

**D. Using the power functional**

Before exploring the modifications of PFT, it is worthwhile to recall the final purpose. SB eventually introduced an ansatz for the power functional, which is [SB equation (30)]

\[ \mathcal{R}_{\text{i}}^{(\text{SB})}[\bar{\rho}, \bar{\phi}, \phi] = \rho_t[\bar{\rho}, \bar{\phi}] + \int \Psi_t \cdot \frac{\partial \mathcal{F}[\rho_t]}{\partial \rho_t} dr - \int \Psi_t \cdot (\nabla \phi_t) dr + \int \phi_t \partial_t \rho_t dr, \tag{30} \]

where \( \rho_t[\bar{\rho}, \bar{\phi}] = W_t^{(\text{SB})}[\bar{\rho}, \bar{\phi}] - \int \Psi_t \cdot (\nabla \phi_t) dr + \int \phi_t \partial_t \rho_t dr \) is the intrinsic functional of dissipated power. This is then used to complete the minimization procedure defined in Eq. (19), giving the equations

\[ \frac{\delta}{\delta \rho_t} \mathcal{R}_{\text{i}}^{(\text{SB})}[\bar{\rho}, \bar{\phi}, \phi] = 0, \tag{31} \]

\[ \frac{\delta}{\delta \phi_t} \mathcal{R}_{\text{i}}^{(\text{SB})}[\bar{\rho}, \bar{\phi}, \phi] = 0; \]

however, such an interpretation is untenable as it leads to unphysical results (see Appendix B). In fact, SB say at this point that a Lagrange multiplier should be introduced prior to minimization so as to enforce the continuity equation relating the density and current. This statement is problematic for three reasons. First, as illustrated above when discussing the toy model [Eq. (22)], the continuity equation is already implicit in the formalism since it reproduces the exact distribution from which the continuity equation for the density is automatically valid. Second, this ad-hoc modification would not be the minimization that was defined in Eq. (19) and that has been used throughout the analysis. Third, if \( \rho_t \) and \( \Psi_t \) are related by the continuity equation, then they could never have been treated as independent constraints—since the density is then fully determined by the temporal history of the current—and the original split of the minimization should have taken the form

\[
\min_{\bar{\Psi}} \left[ \bar{\Psi}; \phi_t, \phi, \psi_t \right] = \min_{\bar{\Psi}} \left[ \min_{\psi_t} \left[ \bar{\Psi}; \phi_t, \psi_t \right] \right], \tag{32}
\]

with no density constraint at all and so leaving only the variational equation with respect to the current. In such a case, with no variation with respect to the density, the relationship with classical DFT (as the equilibrium limit) becomes less clear: it is relegated to the statement—based on the ansatz above—that the current is zero, and so, the density is stationary in time if the equilibrium condition

\[ \nabla \frac{\delta \mathcal{F}[\rho_t]}{\delta \rho_t} = \nabla \mu = 0 \tag{33} \]

for some constant \( \mu \) holds.

**IV. A VARIATION ON PFT**

The above discussion suggests that several of the problems identified can be addressed with a few modifications of the theory, in particular the use of the variational velocities rather than the variational distribution, imposing a constraint on the current but not on the density. Since the distribution is required in the definition of the current, some replacement for the variational distribution \( \bar{\Psi} \) must be found. Recalling as one source of problems that \( \bar{\Psi}_t[\bar{\Psi}; \phi_t] = (\bar{\Psi}_t = \bar{\Psi}_t[\Psi_{\text{min}}]) \), we begin with a redefinition of \( \bar{\Psi}_t \), satisfying the following equation:

\[ \frac{\partial}{\partial t} \Psi_{\text{ne}}[\Psi_t; \phi] = - \sum_i \nabla_i \cdot \Psi_{\text{ne}}^{(i)}[\Psi_{\text{ne}}[\Psi_t; \phi], \Psi_t, \Psi_t] = \Psi_t[\Psi_t; \phi] \tag{34} \]
for any set of variational velocities \( \vec{\psi}_i^{(v)} \) specified for all relevant times. This describes the time evolution of the many-body distribution due to an arbitrary velocity field, the variational field \( \vec{\psi}_v \), with the initial condition \( \vec{\psi}_v^{(t)} \) that is the same initial condition as for the exact distribution. This is of course equivalent to the integral form

\[
\Psi_{\text{tot}}^{(v)}[\vec{\psi}_v; \phi_1, \phi_2, \phi_3] = \psi_{\text{tot}}[\phi] - \int_0^t \left( \sum_{i=1}^N \left( \vec{\psi}_i^{(v)} \cdot \vec{\nabla}_r \vec{\psi}_i^{(v)}[\vec{\psi}_{\text{tot}}; \phi] \right) \right) dr.
\]

(35)

We keep the original definition of \( R \) [Eq. (14)] evaluated using \( \Psi \) in place of the exact distribution, giving

\[
R\left[ \psi^{(v)}_v; \phi_1, \phi_2, \phi_3 \right] = \int_{\psi^{(v)}_v} \Psi_{\phi_1} \sum_i \left( \frac{1}{2} v_i^{(v)} - F_{\phi_1}^{(v)}[\phi_i, \phi_2] \right)
\times \vec{\psi}_i^{(v)} \cdot d\Psi_{\phi_2} + \psi_{\text{tot}}[\phi_1, \phi_2, \phi_3] \left[ \vec{\psi}_v^{(v)} \right] dN
\]

(36)

so that minimization will now give

\[
\vec{\psi}_i^{(v)}_{\text{global--min}}[\phi_1, \Psi] = \frac{1}{y} F_{\phi_1}^{(v)}[\phi_i, \Psi],
\]

\[
= -\vec{\nabla}_v L^{\psi_v} - \vec{\nabla}_v \phi_v \cdot -k_B T \nabla_i \ln \Psi_{\text{tot}}.
\]

(37)

When this is substituted into the evolution equation [Eq. (34)], it becomes the Fokker–Planck equation, and so, with the exact initial condition, one recovers the exact distribution for the system, thus demonstrating that this is an exact reformulation of the problem.

The power functional is now defined using only a constraint on the current as

\[
\mathcal{R}_v \left[ J, \phi, \phi \right] = \min \left[ \frac{R \left[ \psi^{(v)}_v; \phi_1, \phi_2, \phi_3 \right]}{\int \psi^{(v)}_v \psi^{(v)}_v \right] - J \left[ \psi^{(v)}_v; \phi_1, \phi_2, \phi_3 \right] \left[ \vec{\psi}_v^{(v)} \right] \right],
\]

(38)

so here, the current is evaluated from the usual definition [Eq. (11)] using \( \vec{\psi}^{(v)} \) and \( \vec{\psi}_v \) as inputs, and its value is constrained to be the specified \( \vec{J} \) value. At fixed time \( t \), the minimization is only a minimization with respect to \( \vec{\psi}^{(v)} \) at time \( t \) since \( \psi_{\text{tot}}^{(v)} \) only depends on \( \psi_{\text{tot}}^{(v)} \) at earlier times \( t' < t \). In addition, we have written \( \psi_{\text{tot}}^{(v)} \left[ \psi^{(v)}_v \right] \) (equivalent to \( \psi_{\text{tot}}^{(v)} \left[ \psi^{(v)}_v \right] \)) because \( \vec{J} \) depends on \( \psi^{(v)}_v \) at earlier times, and so, we replace these with (indeed independently determined) \( \psi_{\text{tot}}^{(v)} \left[ \vec{J} \right] \) leaving only an overall dependence on the current temporal history. Analyzing this as before, one finds the power functional

\[
\mathcal{R}_v \left[ J, \phi, \phi \right] = \int \psi_{\text{tot}}^{(v)} \sum_i \left( \frac{1}{2} v_i^{(v)} \right) \left[ \vec{\psi}_i^{(v)} \right] \left[ \vec{J}_i \right] ds^2
\]

\[
- \int J_r \left[ \nabla \cdot \vec{J}_r \right] dr + \int \phi_r \rho_r \left[ \psi_{\text{tot}}^{(v)} \right] dr,
\]

(39)

which has the form of SB equation (25) consisting of the sum of a term independent of the field and linear dependences on the field written in terms of the current \( \vec{J} \) and the corresponding density. Thus, some of the structural problems discussed above have been resolved, although once again one cannot say at this point that \( \psi_{\text{tot}}^{(v)} \) is independent of the field at time \( t \).

One could continue by introducing an ansatz for the power functional as in SB, but here we are able to do much more because it turns out that—unlike in SB—the power functional can be evaluated exactly. Let us return to the basic definition in Eq. (38), and note that the constrained minimization can be formulated using a Lagrange parameter (really a vector field \( \lambda \)) by first defining the Lagrangian

\[
L \left[ \psi^{(v)}; \phi_1, \phi_2, \phi_3 \right] = R \left[ \psi^{(v)}; \phi_1, \phi_2, \phi_3 \right] - \int J_r \left( \vec{J}_r - J \left[ \psi^{(v)}; \phi_1, \phi_2, \phi_3 \right] \right) dr
\]

(40)

and then minimizing by solving

\[
0 = \frac{\delta}{\delta \psi^{(v)}_v} L \left[ \psi^{(v)}; \phi_1, \phi_2, \phi_3 \right] \left[ \vec{J}_r \right],
\]

(41)

\[
0 = \frac{\delta}{\delta \phi_1} L \left[ \psi^{(v)}; \phi_1, \phi_2, \phi_3 \right] \left[ \vec{J}_r \right],
\]

(42)

which gives, since \( R \) is quadratic in the velocities,

\[
0 = \psi_{\text{tot}}^{(v)} \left( \psi^{(v)}_v \right) \left( F_{\phi_1}^{(v)} \left[ \psi^{(v)}_v \right] \right) + \int \left( \frac{\delta}{\delta \psi^{(v)}_v} I_r \left[ \psi^{(v)}; \phi_1, \phi_2, \phi_3 \right] \right) \cdot \lambda_r dr,
\]

(44)

Now, it is straightforward to evaluate

\[
\int \left( \frac{\delta}{\delta \psi^{(v)}_v} I_r \left[ \psi^{(v)}; \phi_1, \phi_2, \phi_3 \right] \right) \cdot \lambda_r dr = \int \lambda_r \delta (\vec{r} - \vec{r}) \psi_{\text{tot}}^{(v)} \left[ \psi^{(v)}_v \right] dr = \lambda_r \psi_{\text{tot}}^{(v)} \left[ \psi^{(v)}_v \right]
\]

(43)

so that the system of equations becomes

\[
0 = \psi_{\text{tot}}^{(v)} \left( \psi^{(v)}_v \right) \left( F_{\phi_1}^{(v)} \left[ \psi^{(v)}_v \right] \right) + \lambda_r \psi_{\text{tot}}^{(v)} \left[ \psi^{(v)}_v \right],
\]

\[
= \int \left( \psi^{(v)}_v \right) \left[ \vec{J}_r \right] \cdot \lambda_r \cdot \lambda_r \psi_{\text{tot}}^{(v)} \left[ \psi^{(v)}_v \right],
\]

(44)

The first line gives

\[
\psi_{\text{tot}}^{(v)} \left[ \psi^{(v)}_v \right] = \frac{1}{y} F_{\phi_1}^{(v)} \left[ \psi^{(v)}_v \right] - \frac{1}{y} \lambda_r,
\]

(45)

while the second then evaluates

\[
\psi_{\text{tot}}^{(v)} \left[ \psi^{(v)}_v \right] = \psi_{\text{tot}}^{(v)} \left[ \psi^{(v)}_v \right] - \frac{1}{y} \lambda_r \rho_r \left[ \psi_{\text{tot}}^{(v)} \right] \left[ \vec{J}_r \right],
\]

(46)

so that the solution to the minimization problem is

\[
\lambda_r = \frac{1}{y} \frac{\psi_{\text{tot}}^{(v)} \left[ \psi^{(v)}_v \right] - \psi_{\text{tot}}^{(v)} \left[ \psi^{(v)}_v \right]}{\rho_r \left[ \psi_{\text{tot}}^{(v)} \right] \left[ \vec{J}_r \right]}. \quad \text{(47)}
\]

Using this to evaluate the power functional gives
\[ R_i \left[ \bar{I}_1; \theta_i, \phi_i \right] = -\frac{1}{2\gamma} \int \phi_i \left( \frac{\partial}{\partial r} \bar{\rho}_t \left[ \Psi_r^{\min} \right] \right) d^3r + \frac{\gamma}{2} \int \left( \bar{I}_r - \bar{I}_r \left[ \phi_i, \Psi_r^{\min} \right] \right)^2 d^3r + \rho_t \phi \left( \frac{\partial}{\partial r} \bar{\rho}_t \left[ \Psi_r^{\min} \right] \right) d^3r. \]  

The dependence on the external field \( \phi_i \) appearing in \( \bar{I}_r^{\text{tot}} \left[ \phi_i, \Psi_r^{\min} \right] \) and \( \bar{I}_r \left[ \phi_i, \Psi_r^{\min} \right] \) can be explicitly taken out, and the functional is rewritten as

\[ R_i \left[ \bar{I}_1; \theta_i, \phi_i \right] = -\frac{1}{2\gamma} \int \phi_i \left( \frac{\partial}{\partial r} \bar{\rho}_t \left[ \Psi_r^{\min} \right] \right) d^3r + \frac{\gamma}{2} \int \left( \bar{I}_r - \bar{I}_r \left[ 0, \Psi_r^{\min} \right] \right)^2 d^3r - \int \phi_i \left( \frac{\partial}{\partial r} \right) \phi \left( \frac{\partial}{\partial r} \bar{\rho}_t \left[ \Psi_r^{\min} \right] \right) d^3r, \]  

thus explicitly showing the structure discussed in SB whereby the external field contributions at time \( t \) appear linearly. Finally, we note that the density automatically satisfies the continuity equation since

\[ \frac{\partial}{\partial r} \bar{\rho}_t \left[ \Psi_r^{\min} \right] = \int \frac{\partial}{\partial r} \bar{\rho}_t \left[ \Psi_r^{\min} \right] \sum_i \delta(r - r_i) d^3r_N = -\int \left( \frac{\partial}{\partial r} \sum_i \delta(r - r_i) \right) \frac{\partial}{\partial r} \bar{\rho}_t \left[ \Psi_r^{\min} \right] d^3r_N = -\nabla \cdot \int \delta(r - r_i) \frac{\partial}{\partial r} \bar{\rho}_t \left[ \Psi_r^{\min} \right] d^3r_N = -\nabla \cdot \bar{J}_r, \]  

by definition of \( \bar{\rho}_t^{(\min)} \). This shows that we can replace \( \bar{\rho}_t \left[ \Psi_r^{\min} \right] \) by \( \bar{\rho}_t \left[ \bar{\rho}_t \right] \) since the current (and an initial condition) completely determines the density. We can interpret this construction as follows: \( R_i \left[ \bar{I}_1; \theta_i, \phi_i \right] \) is a functional for the power in a system, which is described by a certain time-dependent current \( \bar{I}_r \). The associated many-body distribution \( \bar{\rho}_t \left[ \bar{I}_r \right] \) is consistent with this current (i.e., they are related by the usual definition), and the corresponding local density fulfills the continuity with \( \bar{I}_r \) as the material-transporting current. Upon minimization with respect to \( \bar{I}_r \), under the usual constraint of causality, the minimal currents \( \bar{\rho}_r^{(\min)} \left[ \bar{I}_r \right] \) become the exact physical solutions corresponding to the applied field \( \phi_i \).

### A. Pessimistic interpretation of these results

Following SB, the power functional is to be minimized with respect to the current so as to get the minimizing current. Here, this is trivial and results in

\[ \bar{\rho}_r^{(\min)} = \bar{I}_r \left[ \theta_i, \bar{\rho}_r^{(\min)} \left[ \bar{I}_r \right] \right], \]  

and from Eq. (47), one then has for the variational velocities

\[ \bar{\rho}_r^{(\min)} \left[ \bar{I}_r^{\text{tot}} \left[ \phi_i, \bar{\rho}_r \left[ \bar{I}_r \right] \right] \right] = \frac{1}{\gamma} \bar{I}_r^{\text{tot}} \left[ \phi_i, \bar{\rho}_r \left[ \bar{I}_r \right] \right]. \]  

As demonstrated at the start of this section [see Eq. (34)], this simply implies that \( \bar{\rho}_r^{(\min)} \) satisfies the original Fokker–Planck equation and, so, is the exact distribution. One arrives at the same conclusion by recognizing that the solution to Eq. (51) for the current at time \( t \) depends on the current at all earlier times, and for finding a self-consistent solution \( \bar{I}_r \left[ \bar{\rho}_r \right] = \bar{I}_r \left[ \bar{\rho}_r^{(\min)} \right] \), one has to solve Eq. (51) at all earlier times successively. This amounts to nothing but the solution of the original Fokker–Planck equation.

Without the freedom to introduce an ansatz by hand, the formalism just reduces to the original exact result. There seems, therefore, to be no advantage to this development since, while one could introduce one of the usual approximations (e.g., local equilibrium) at any point, there seems no particular rationale to do so as opposed to, e.g., simply inserting such an ansatz directly into the continuity equation for the density (as is done in heuristic derivations of DFT–).
\[
\begin{align*}
\frac{\partial}{\partial \tau} \mathcal{F}_{\tau}[\mathbf{r}] &= -\nabla \cdot \mathbf{J}_{\tau}^{\text{min}}, \\
\mathbf{J}_{\tau}^{\text{min}} &= J_{\tau}\left[\phi, \mathbf{\Psi}_{\tau}^{\text{min}}\right],
\end{align*}
\]

which [using the definitions in Eqs. (8) and (11)] can be written as

\[
\begin{align*}
\mathbf{J}_{\tau}^{\text{min}} &= -\frac{k_B T}{\gamma} \nabla \rho_{\tau}[\mathbf{r}] \frac{1}{\gamma} \rho_{\tau}[\mathbf{r}] (\nabla \phi_{\tau}) + K_{\tau}[\mathbf{r}] \left|\right. \nabla \phi_{\tau}^{\text{min}}[\mathbf{r}],
\end{align*}
\]

where the last term is

\[
K_{\tau}[\mathbf{r}] = -\frac{1}{\gamma} \int \delta(\mathbf{r} - \mathbf{r}_i) \left(\left(\nabla_i U(\mathbf{r}_i)\right) \mathbf{\Psi}_{\tau}^{\text{min}}[\mathbf{r}]\right) d\mathbf{r}_i.
\]

Although not indicated, everything also depends on the initial condition for the distribution, \(\mathbf{\Psi}_{\tau_0}^{\text{eq}}\). Thus, we can summarize the evolution of the physical density with the following exact equations:

\[
\begin{align*}
\frac{\partial}{\partial \tau} \mathcal{F}_{\tau}[\mathbf{r}] &= -\nabla \cdot \mathbf{J}_{\tau}^{\text{min}}, \\
\mathbf{J}_{\tau}^{\text{min}} &= -\frac{k_B T}{\gamma} \nabla \rho_{\tau}[\mathbf{r}] \frac{1}{\gamma} \rho_{\tau}[\mathbf{r}] (\nabla \phi_{\tau}) + K_{\tau}[\mathbf{r}] \left|\right. \nabla \phi_{\tau}^{\text{min}}[\mathbf{r}],
\end{align*}
\]

For an ideal gas, \(K_{\tau} = 0\) by definition and this is exact—not as a consequence of an ansatz but directly demonstrated from the microscopical formalism. In equilibrium, the distribution is constant, so \(\mathbf{\Psi}_{\tau}^{\text{eq}}[\mathbf{r}] = \mathbf{\Psi}_{\tau_0}^{\text{eq}}\), and so

\[
K_{\tau}^{\text{eq}}[\mathbf{r}] = -\frac{1}{\gamma} \int \delta(\mathbf{r} - \mathbf{r}_i) \left(\left(\nabla_i U(\mathbf{r}_i)\right) \mathbf{\Psi}_{\tau}^{\text{eq}}[\mathbf{r}]\right) d\mathbf{r}_i
\]

where \(\mathbf{\Psi}_{\tau}^{\text{eq}}[\mathbf{r}]\) is the excess Helmholtz free energy functional of classical DFT. This is an exact result (essentially a consequence of the Yvon-Born-Green (YBG) hierarchy—see, e.g., Ref. 4). Thus, a natural local-equilibrium assumption would be to continue to use this form for non-equilibrium systems, resulting in

\[
\begin{align*}
\mathbf{J}_{\tau}^{\text{min}}[\mathbf{r}] &= -\frac{k_B T}{\gamma} \nabla \rho_{\tau}[\mathbf{r}] \frac{1}{\gamma} \rho_{\tau}[\mathbf{r}] (\nabla \phi_{\tau}) \\
&= -\frac{1}{\gamma} \nabla \mathcal{F}_{\tau}^{\text{ex}}[\rho],
\end{align*}
\]

which, when inserted into the continuity equation, gives a closed dynamics for the density usually referred to as Dynamic Density Functional Theory (DDFT).

The investigation of deviations between the true dynamics and the dynamics predicted by DDFT (adiabatic dynamics in the sense of a dynamics of quasi-static changes) has been pursued by Schmidt and co-workers over the past years (see, e.g., Refs. 19 and 20). Even without knowing the explicit form of the equilibrium free energy functional, the adiabatic dynamics can be determined by simulation and thus also the nonadiabatic differences to the true dynamics. In terms of the ansatz in Eq. (30), these differences are contained in the excess part \(P_{\tau}^{\text{ex}}\) of the functional \(P_{\tau}[\mathbf{r}]\), which is defined as follows:

\[
P_{\tau}[\mathbf{r}] = P_{\tau}^{\text{id}}[\mathbf{r}] + P_{\tau}^{\text{ex}}[\mathbf{r}] \quad \text{with} \quad P_{\tau}^{\text{id}}[\mathbf{r}] = \frac{y}{2} \int \nabla \cdot \mathbf{J}_{\tau} d\mathbf{r}. \tag{62}
\]

The relation to the functional \(K_{\tau}\) defined above is given by

\[
K_{\tau}[\mathbf{r}] = \frac{\delta P_{\tau}^{\text{ex}}[\mathbf{r}]}{\delta \rho_{\tau}[\mathbf{r}]} + \frac{1}{y} \nabla \left(\frac{\delta P_{\tau}^{\text{ex}}[\rho]}{\delta \rho_{\tau}[\rho]}\right) \rho_{\tau}[\rho]. \tag{63}
\]

Note that we have restricted the functional dependence of the \(P_{\tau}\) functional to the current only, in line with our modified formulation of PFT.

\section*{V. CONCLUSIONS}

We have attempted to clarify the construction of power functional theory for Brownian dynamics. The theory begins with a simple variational principle that reproduces the exact statistical description of the system based on a generating function ultimately related to the Rayleigh dissipation function. The minimization is broken into two steps—first, minimization under the constraints of constant density and current, followed by minimization over those fields. The first minimization results in a functional of the current and density fields called the power functional. This is meant to be analogous to the Helmholtz functional of cDFT and to play a similar role, as a starting point for physically motivated approximations to the dynamics of the full \(N\)-body system. Our analysis has identified several problematic points. There is some ambiguity in the development of SB as to whether the power functional is actually the functional of both density and current fields or only of the current. A dependence on both seems to lead to unphysical results so that it should probably be considered as the result of a constrained minimization at fixed current. Independent of this, the resulting expressions do not have the structure described by SB, which, in turn, has been used in practical applications of the theory later on, and thus the status of those is uncertain and the relation to the underlying power functional is unclear.

Having identified the source of the problems in the original development, we have proposed modifications that appear to avoid them. The result is, in some ways, very similar to that discussed by SB but with one critical difference: in our development, it is possible to carry through the first constrained minimization exactly, resulting in the exact power functional. This turns out to be a rather trivial quadratic function that simply forces the current to take on its exact expression. The theory therefore seems to provide no real advantage over, for example, simply introducing approximations directly into the continuity equation.

More optimistically, we noted that our development does result in an interesting reformulation of the dynamics whereby a “universal” functional of the current is defined (i.e., one that is independent of the external field) and can be used to formulate the dynamics governing the local density. Dropping the unnecessary notation, the resulting theory has the form

\[
\text{...}
\]
\[ \frac{\partial}{\partial t} \rho_\text{tr} = -\nabla \cdot J_\text{tr}, \]
\[ J_\text{tr} = -\frac{k_B T}{\gamma} \nabla \rho_\text{tr} - \frac{1}{\gamma} \rho_\text{tr} (\nabla \phi_\text{tr}) + K_\text{tr}[J], \quad (64) \]

with the non-ideal—or excess—contribution to the current given by
\[ K_\text{tr}[J] = -\frac{1}{\gamma} \int \sum_i \delta(r - r_i) \Psi_{tr} \nabla_i U(r) \, dr^N, \]
and finally the “universal” distribution determined by \( \Psi_\text{tr}^{(0)} \) is the distribution specifying initial conditions
\[ \frac{\partial}{\partial t} \Psi_{tr}^{(0)}[J] = -\sum_i \frac{1}{\gamma} \Psi_{tr} \nabla_i [0, \Psi] \Psi_{tr}^{(0)}[J] \]
\[ + \sum_i \nabla_i \left( J_r[0, \Psi] - J_{tr} \Psi_{tr}^{(0)}[J] \right) \]
with
\[ \Psi_{tr} = \Psi_{tr}^{(0)}. \]
\[ (66) \]

This is similar to the structure of classical DFT in which one has a universal functional of the density, so-called Helmholtz functional \( F[\rho] \), and from this, one constructs simple one-body equations for the local density given an external field. Here, \( K_\text{tr}[J] \) plays the role of the Helmholtz functional, and it shares another characteristic as well: its determination depends on solving an \( N \)-body problem that is as complicated as simply solving the original Fokker–Planck equation (or in the case of cDFT calculating the original partition function) that defined the starting point. Whether there is any advantage to such a reformulation of the original problem remains to be seen.

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**APPENDIX A: DERIVATION OF THE FORM OF THE GENERATING FUNCTIONAL**

SB begin by defining [their Eq. (10)]
\[ \tilde{R}_{tr}^{\Psi_{tr}}[\Psi_{tr}] = \sum_i \frac{1}{2} \Psi_{tr}^{(0)} - F_{tr}^{\text{tot}}(\Psi_{tr}) \cdot \Psi_{tr}^{(0)} + \phi_{tr}, \]
\[ (A1) \]
with [SB equation (5)]
\[ F_{tr}^{\text{tot}}(\phi_t, \Psi_t) = -\nabla U_{\text{tr}} - \nabla \phi_{\text{tr}} - k_B T \nabla_l \ln \Psi_{\text{tr}}, \]
\[ (A2) \]
and [SB equation (17)]
\[ \gamma_{tr}^{(1)} = \gamma_{tr}^{(1)}(\phi_t, \Psi_t) = F_{tr}^{\text{tot}}(\phi_t, \Psi_t) \]
\[ (A3) \]
in terms of which the Fokker–Planck equation becomes
\[ \frac{\partial}{\partial t} \Psi_{tr} = -\frac{1}{\gamma} \sum_i \nabla_i \left( \Psi_{tr} F_{tr}^{\text{tot}}(\phi_t, \Psi_t) \right). \]
\[ (65) \]

The functional \( \tilde{R}_{tr}^{\Psi_{tr}} \) can be written as
\[ \tilde{R}_{tr}^{\Psi_{tr}}[\Psi_{tr}] = \frac{k_B T}{2} \int \sum_i \left( \nabla_i \ln \Psi_{tr} \right) \Psi_{tr}^{2} \, dr^N = \frac{k_B T}{2} \int \sum_i \left( \nabla_i \ln \Psi_{tr} \right) \Psi_{tr}^{2} \, dr^N \]
\[ (A4) \]
The SB generating functional is [SB equation (11)]
\[ R[\Psi_t; \Psi_{\text{tr}}] = \int \Psi_{\text{tr}} \tilde{R}_{tr}^{\Psi_{tr}} \, dr^N \]
\[ = \frac{k_B T}{2} \int \sum_i \left( \nabla_i \ln \Psi_{\text{tr}} \right) \Psi_{\text{tr}}^{2} \, dr^N \]
\[ - \frac{1}{2} \int \Psi_{\text{tr}} \Psi_{\text{tr}}^{2} \, dr^N \]
\[ + \int \Psi_{\text{tr}} \phi_{\text{tr}} \, dr^N. \]
\[ (A5) \]
The second term is
\[ - \frac{1}{2} \int \Psi_{\text{tr}} \sum_i \left( F_{tr}^{\text{tot}}(\Psi_{\text{tr}}, \phi_t) \right)^2 \, dr^N \]
\[ = \frac{1}{2} \int \sum_i \left( \nabla_i (U_{\text{tr}} + \phi_{\text{tr}} + k_B T \ln \Psi_{\text{tr}}) \right) \Psi_{\text{tr}} \, dr^N \]
\[ - \frac{1}{2} \int \sum_i \nabla_i \Psi_{\text{tr}}^{2} F_{tr}^{\text{tot}}(\Psi_{\text{tr}}, \phi_t) \, dr^N, \]
\[ (A6) \]
where the second line follows from the divergence theorem and surface terms are ignored. Substituting for the divergence using the Fokker–Planck equation yields
\[ - \frac{1}{2} \int \Psi_{\text{tr}} \sum_i \left( F_{tr}^{\text{tot}}(\Psi_{\text{tr}}, \phi_t) \right)^2 \, dr^N \]
\[ = - \frac{1}{2} \int \left( U_{\text{tr}} + \phi_{\text{tr}} + k_B T \ln \Psi_{\text{tr}} \right) \left( -\frac{\partial}{\partial t} \Psi_{\text{tr}} \right) \, dr^N \]
\[ = \frac{1}{2} \int \left( U_{\text{tr}} + \phi_{\text{tr}} + k_B T \ln \Psi_{\text{tr}} \right) \Psi_{\text{tr}} \, dr^N \]
\[ = \frac{1}{2} \int \left( \frac{\partial}{\partial t} \phi_{\text{tr}} \right) \Psi_{\text{tr}} \, dr^N. \]
\[ (A7) \]
Putting everything together gives the form used in the main text,
\[ R[\Psi_t; \Psi_{\text{tr}}] = \frac{k_B T}{2} \int \sum_i \left( \nabla_i \ln \Psi_{\text{tr}} \right) \Psi_{\text{tr}}^{2} \, dr^N \]
\[ + \frac{1}{2} \int \left( \frac{\partial}{\partial t} \phi_{\text{tr}} \right) \Psi_{\text{tr}} \, dr^N. \]
\[ (A8) \]
In order to arrive at Eq. (27) of the main text, we need to evaluate
\[
\mathcal{R}_e \left[ \tilde{\rho}, \tilde{J}, \phi_t, \phi_t' \right] = R \left[ \tilde{\Psi}_e^{\min} \left[ \tilde{\rho}, \tilde{J} \right]; \phi_t, \phi_t', \tilde{\Psi}_e \left[ \tilde{\rho}, \tilde{J} \right] \right]
\]
or, using Eq. (A5) and dropping all unnecessary functional arguments for the sake of clarity,
\[
\begin{align*}
\mathcal{R}_e \left[ \tilde{\rho}, \tilde{J}, \phi_t, \phi_t' \right] &= \frac{(k_B T)^2 y}{2} \sum_i \left( \ln \frac{\tilde{\Psi}_e^{\min}}{\tilde{\Psi}_e} \right)^2 \tilde{\Psi}_e \, dr^N \\
&- \frac{1}{2y} \int \tilde{\Psi}_e \tilde{\Psi}_e' \sum_i \left( \tilde{F}_e^{\phi_t} \left[ \tilde{\Psi}_e, \phi_t \right] \right)^2 \, dr^N \\
&+ \int \tilde{\Psi}_e \tilde{\Psi}_e' \phi_t \phi_t' \, dr^N.
\end{align*}
\]
\[\text{(A9)}\]

Now, given the linear dependence of \( F_e^{\phi_t} \) on the field,
\[ F_e^{\phi_t} \left[ \phi, \Psi \right] = F_e^{\phi_t} \left[ 0, \Psi \right] - \nabla \phi \psi_{r} \tag{A10} \]
ones that
\[
- \frac{1}{2y} \int \tilde{\Psi}_e \tilde{\Psi}_e' \sum_i \left( \tilde{F}_e^{\phi_t} \left[ \tilde{\Psi}_e, \phi_t \right] \right)^2 \, dr^N \\
= - \frac{1}{2y} \int \tilde{\Psi}_e \tilde{\Psi}_e' \sum_i \left( \tilde{F}_e^{\phi_t} \left[ 0, \Psi \right] \right)^2 \, dr^N \\
+ \frac{1}{y} \int \tilde{\Psi}_e \tilde{\Psi}_e' \sum_i \left( \tilde{F}_e^{\phi_t} \left[ 0, \Psi \right] \cdot \nabla \phi \right) \, dr^N \\
- \frac{1}{2y} \int \tilde{\Psi}_e \tilde{\Psi}_e' \sum_i \left( \nabla \phi \psi_ \right) \, dr^N. \tag{A11}
\]
The second term is
\[
\begin{align*}
\frac{1}{y} \int \tilde{\Psi}_e \tilde{\Psi}_e' \sum_i \left( \tilde{F}_e^{\phi_t} \left[ 0, \Psi \right] \cdot \nabla \phi \right) \, dr^N \\
= \int \tilde{\Psi}_e \tilde{\Psi}_e' \sum_i \left( \frac{1}{y} \tilde{F}_e^{\phi_t} \left[ 0, \Psi \right] \right) \left( \int \left( \nabla \phi \right) \delta(r - r_i) \, dr \right) \, dr^N \\
= \int \left( \nabla \phi \right) \cdot \left( \tilde{\Psi}_e \tilde{\Psi}_e' \sum_i \left( \frac{1}{y} \tilde{F}_e^{\phi_t} \left[ 0, \Psi \right] \right) \, dr \right) \, dr \\
= \int \left( \nabla \phi \right) \cdot \left[ 0, \Psi \right] \, dr.
\end{align*}
\]
\[\text{(A12)}\]

Thus,
\[
\mathcal{R}_e \left[ \tilde{\rho}, \tilde{J}, \phi_t, \phi_t' \right] = \frac{(k_B T)^2 y}{2} \sum_i \left( \ln \frac{\tilde{\Psi}_e^{\min}}{\tilde{\Psi}_e} \right)^2 \tilde{\Psi}_e \, dr^N \\
- \frac{1}{2y} \int \tilde{\Psi}_e \tilde{\Psi}_e' \sum_i \left( \tilde{F}_e^{\phi_t} \left[ 0, \Psi \right] \right)^2 \, dr^N \\
+ \int \left( \nabla \phi \right) \cdot \left[ 0, \Psi \right] \, dr \\
- \frac{1}{2y} \int \left( \nabla \phi \right)^2 \tilde{\Psi}_e \, dr^N \\
+ \int \rho_e \left[ \tilde{\Psi}_e \right] \phi_t \phi_t' \, dr^N. \tag{A13}
\]

which is Eq. (27) in the main text.

**APPENDIX B: MINIMIZING WITH INDEPENDENT DENSITY AND CURRENT**

To be of use, models must be developed for the functionals to be minimized. SB proposed to write the functional \( W_t^{(SB)} \) as
\[
W_t^{(SB)} \left[ \tilde{\rho}, \tilde{J}, \phi_t, \phi_t' \right] = p^{(ex)} \left[ \tilde{\rho}, \tilde{J}, \phi_t, \phi_t' \right] + \int \left( \frac{\tilde{F}_c}{2y} + \frac{J_t}{r} - \nabla \frac{\delta F}{\delta \tilde{r}} \right) \, dr', \tag{B1}
\]
where \( p^{(ex)} \) is the excess “dissipated power functional,” the quadratic term in \( \tilde{J} \) is the “ideal dissipated power functional,” and the third term is the “adiabatic” (i.e., local equilibrium) contribution that involves the DFT free energy functional \( F[\tilde{\rho}] \) [see Eqs. (30) and (62)]. The Euler–Lagrange equations then become
\[
\frac{\delta}{\delta \tilde{\rho}} p^{(ex)} \left[ \tilde{\rho}, \tilde{J}, \phi_t, \phi_t' \right] - \frac{\tilde{J}_t}{2y} = \int \left( \nabla \cdot \tilde{J}_t \right) \frac{\delta^2 F[\tilde{\rho}]}{\delta \tilde{\rho}^2} \, dr', \tag{B2}
\]
\[
\frac{\delta}{\delta \tilde{J}_t} p^{(ex)} \left[ \tilde{\rho}, \tilde{J}, \phi_t, \phi_t' \right] + \frac{\tilde{J}_t}{y} + \nabla \frac{\delta F[\tilde{\rho}]}{\delta \tilde{\rho}} + \nabla \phi_t = 0.
\]

Separating the equilibrium free energy functional into its ideal and excess parts \( F = F^{(ad)} + F^{(ex)} \) with \( F^{(ad)}[\rho] = k_B T / \partial r \left( \ln \left( \rho(r) \lambda^3 \right) - 1 \right) \), where \( \lambda \) is the thermal de Broglie length and rearranging give
\[
\tilde{\rho}_t \frac{\delta}{\delta \tilde{\rho}} p^{(ex)} \left[ \tilde{\rho}, \tilde{J}, \phi_t, \phi_t' \right] - \frac{\tilde{J}_t}{2y} - \tilde{\rho}_t \left( \nabla \cdot \tilde{J}_t \right)
\]
\[
- \tilde{\rho}_t \int \left( \nabla \cdot \tilde{J}_t \right) \frac{\delta^2 f^{(ex)}}{\delta \tilde{\rho}^2} \, dr' = 0,
\]
\[
\tilde{\rho}_t \frac{\delta}{\delta \tilde{J}_t} p^{(ex)} \left[ \tilde{\rho}, \tilde{J}, \phi_t, \phi_t' \right] + \frac{\tilde{J}_t}{y} + \nabla \frac{\delta F[\tilde{\rho}]}{\delta \tilde{\rho}} + \tilde{\rho}_t \nabla \phi_t = 0. \tag{B3}
\]

This completes as much of the general framework as laid out by SB and needed here, and now, we turn to applications to specific systems.

1. **Equilibrium**

   In equilibrium, there is no time dependence and the current vanishes. We assume that the contribution from \( p^{(ex)} \) also vanishes so that the only non-trivial equation is the second (coming from the variation of the current), which can be written as
\[
\tilde{\rho}_t \nabla \frac{\delta}{\delta \tilde{\rho}} \left( F[\tilde{\rho}] + \int \tilde{\rho}_t \phi_t dr' \right) = 0 \tag{B4}
\]
and which has the usual equilibrium DFT solution
\[
F[\tilde{\rho}] + \int \tilde{\rho}_t \phi_t dr' = \text{const.} \tag{B5}
\]
In this sense, the proposed ansatz for \( W \) reproduces equilibrium cDFT.
2. Adiabatic approximation

By the definition of SB, “adiabatic” means that one ignores the excess dissipated power functional \( \delta \mathcal{F}^{\text{ex}} = 0 \), giving

\[
0 = \frac{1}{2u} \mathbf{J} + \mathbf{p}_0 (\nabla \cdot \mathbf{J}) + \mathbf{p}_0 \int \left( \mathbf{v}' \cdot \mathbf{J}' \right) \frac{\delta \mathcal{F}^{\text{ex}}}{\delta \mathbf{p}_0} d\mathbf{r}',
\]

(B6)

\[
\frac{1}{2} \mathbf{J} = -\nabla \mathbf{p}_0 - \mathbf{p}_0 \nabla \mathbf{p}'_0 - \nabla \mathbf{p}_0 \nabla \phi_r.
\]

The second equation gives the current in terms of the density and external field, and if this is inserted into the continuity equation, the resulting equation of motion can be written as

\[
\frac{\partial \mathbf{p}_0}{\partial t} = \nabla \mathbf{p}_0 \nabla \phi_r ( F[\mathbf{r}] + \int \mathbf{p}_0' \phi_r d\mathbf{r} ),
\]

(B7)

which looks like the standard DDFT equation of motion. However, as emphasized above, there is no reason to suppose that the density and current satisfy the continuity equation, and indeed, as the notation makes clear, \( \mathbf{p}_0 \) is a time-independent test field so that this equation makes little sense. Rather, the density should be determined from the first of Eq. (B6), which is purely local in time. Denoting the solution of the Euler–Lagrange equation as

\[
\mathbf{p}_0^* [\phi_1], \mathbf{J}_r^* [\phi_1]
\]

since the only time dependence in Eq. (B6) occurs via the potential and substituting into the continuity equation give

\[
\int \frac{\delta \mathcal{F}^{\text{ex}}}{\delta \phi_r} \frac{\partial \phi_r}{\partial t} d\mathbf{r}' = -\nabla \mathbf{J}_r^* [\phi_1],
\]

(B9)

which only can be satisfied for particular fields. In particular, it is clear that in the case of an external field that does not depend on time, one could start with a density that is not the equilibrium density so that presumably \( \mathbf{p}_0^* \) will relax to its equilibrium value. In such a case, the left-hand side vanishes, but \( \mathbf{J}_r^* \) will be nonzero, and indeed nontrivial, so that the right hand side does not vanish, thus showing that such a solution cannot be consistent with the continuity equation. To understand this result better, we turn to the final special case, the ideal gas.

3. Ideal gas in the adiabatic approximation

For the ideal gas, the excess part of the free energy functional vanishes, leaving

\[
\frac{1}{2u} \mathbf{J} + \mathbf{p}_0 (\nabla \cdot \mathbf{J}) = 0,
\]

(B10)

\[
\frac{1}{2} \mathbf{J} = -\nabla \mathbf{p}_0 - \mathbf{p}_0 \nabla \phi_r ( \beta = 1/(k_B T) ).
\]

The equation for the current is, in fact, the exact result that can be derived directly from the Fokker–Planck equation. It can be used to eliminate the current from the first equation, giving

\[
\mathbf{p}_0 \left( \nabla^2 \mathbf{p}_0 \right) - \frac{1}{2} \left( \nabla \mathbf{p}_0 \right)^2 + \mathbf{p}_0 \left( \nabla^2 \phi_r - \frac{1}{2} \left( \nabla \phi_r \right)^2 \right) = 0,
\]

(B11)

which is satisfied by the adiabatic solution

\[
\mathbf{p}_0^* [\phi_1] = A e^{-\beta \phi_r},
\]

(B12)

for some constant \( A \). However, this gives vanishing current and when substituted into the continuity equation gives

\[
\frac{\partial}{\partial t} \phi_r = 0,
\]

(B13)

indicating that the minimizers \( \mathbf{p}_0^* \) and \( \mathbf{J}_r^* \) are only consistent with the continuity equation in the particular case of a time-independent field.

DATA AVAILABILITY

Data sharing is not applicable to this article as no new data were created or analyzed in this study.

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