AEGIS: an experiment to measure the gravitational interaction between matter and antimatter

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Abstract. Considerable efforts have been made and are still being made to verify the validity of the principle of the equivalence of gravitational and inertial mass, one of the cornerstones of the classical theory of general relativity. Specific attempts at quantum-mechanical formulations of gravity allow for non-Newtonian contributions which might lead to a difference in the gravitational force on matter and antimatter. While it is widely expected that the gravitational interaction is independent of the composition of bodies, this has only been tested for matter systems, but never yet for antimatter systems. By combining techniques from different fields, and relying on recent developments in the production of Positronium and ongoing work to laser-excite Positronium to Rydberg states, such a test with neutral antimatter has become feasible. The primary goal of the AEGIS experiment being built at the Antiproton Decelerator at CERN is to carry out the first direct measurement of the Earth's gravitational acceleration on antihydrogen by means of a classical Moiré deflectometer.

1. Introduction
A hypothetical quantum theory of gravitation necessarily constitutes a departure from the Einsteinian view of gravity as a geometric phenomenon and could potentially constitute a violation of the weak equivalence principle. Attempts at formulating a quantum theory of gravity have mainly been hampered by the fact that such a theory is non-renormalizable, though a renormalization within the framework of an effective field theory may turn out to be feasible [2]. As in other quantum field theories, the interaction is mediated by exchange particles. The spins of these exchange bosons as well as the signs of the charges to which they couple determine whether a force is repulsive or attractive. Modern theories of gravity that attempt to unify gravity with the other forces of nature allow that, at least in principle, antimatter may fall differently from normal matter in the Earth's gravitational field. Specifically, as pointed out by Sherk [3], N=2,...,8 theories of supergravity (where N is the number of supersymmetries) lead to the possibility of (massive, and thus finite range) Kaluza-Klein graviscalar and gravivector components to the gravitational interaction which may lead to different couplings for matter and antimatter.

While it is possible to constrain deviations from standard gravity through e.g. effects of virtual antiparticles in ordinary matter, most of the quantitative limits [4, 5, 6] do not apply to more elaborate models involving vector and scalar gravitons [7]. If the hypothetical vector and scalar charges, as well as their masses (and thus the ranges of the interactions) are carefully chosen, such contributions can be strongly suppressed in ordinary matter. The most stringent limits on such additional components come from an analysis of the Eöt-Wash experiment.
an Eötvös-type torsion-balance experiment using test bodies of different compositions. Unlike ordinary matter, the behavior of antimatter particles in a gravitational field has never been tested experimentally. Two attempts, at Stanford [10] and CERN’s Low-Energy Antiproton Ring [11] were thwarted by the overwhelming effect of stray electric and magnetic fields upon the electrically charged test particles. The recent production of copious amounts of cold antihydrogen $\bar{\text{H}}$ at CERN’s Antiproton Decelerator (AD) [12, 13] has paved the way for high-precision gravity experiments with neutral antimatter. We have proposed the AEGIS experiment (Antimatter Experiment: Gravity, Interferometry, Spectroscopy), to be performed at CERN/AD, in order to address this important question.

The primary scientific goal of AEGIS is the direct measurement of the Earth’s local gravitational acceleration $g$ on $\bar{\text{H}}$. In a first phase of the experiment, a gravity measurement with 1% relative precision will be carried out by observing the vertical displacement (using a high-resolution position sensitive detector) of the shadow image produced by an $\bar{\text{H}}$ beam, formed by its passage though a Moiré deflectometer, the classical counterpart of a matter wave interferometer. This measurement will represent the first direct determination of the effect of gravity on antimatter.

The essential steps leading to the production of a pulsed cold beam of $\bar{\text{H}}$ and the measurement of $g$ with AEGIS are the following (Fig. 1):

- Production of positrons ($e^+$) from a Surko-type source and accumulator;
- Capture and accumulation of $\bar{p}$ from the AD in a cylindrical Penning trap;
- Cooling of the $\bar{p}$ to sub-K temperatures
- Production of positronium (Ps) by bombardment of a cryogenic nanoporous material with an intense $e^+$ pulse;
- Excitation of the Ps to a Rydberg state with principal quantum number $n = 30 \sim 40$;
- Pulsed formation of $\bar{\text{H}}$ by resonant charge exchange between Rydberg Ps and cold $\bar{p}$;
- Pulsed formation of an $\bar{\text{H}}$ beam by Stark acceleration with inhomogeneous electric fields;
- Determination of $g$ in a two-grating Moiré deflectometer coupled with a position-sensitive detector.

The feasibility of the first two points has been conclusively demonstrated by the ATHENA and ATRAP collaborations (see, in particular [14, 15]). In the following, we will discuss the remaining aspects of the proposed technique in more detail.

### 2. Method

#### 2.1. Positronium production and excitation

In recent years, the potential for nanoporous insulator materials to be used as highly efficient Ps converters has been recognized, and the relevant formation mechanisms have been studied extensively [16]. When $e^+$ are implanted into such a material at kinetic energies ranging from several 100 eV to a few keV, they scatter off atoms and electrons in the bulk and are slowed to eV energies within a few ps. With efficiencies ranging from 10% to 50%, the slow $e^+$ capture either bound $e^-$ or those liberated in prior collisions and form Ps. In the pores, Ps repeatedly bounces off the cavity walls and eventually approaches complete thermalization with the target material.

While some ortho-Ps are lost due to so-called pick-off annihilations of $e^+$ with the molecular $e^-$ of the cavity walls, a sizable fraction diffuses out of the film at thermal energies. The overall ortho-Ps yield as well as the final velocity distribution depend upon the characteristics of the target material (in particular, its pore structure), the implantation depth, and the target temperature. For the AEGIS experiment, the precise degree of thermalization is critical, since at too low a Ps temperature, annihilation in the target material will dominate, while at too high
Figure 1. Proposed method for the production of a pulsed beam of cold $\bar{H}$ atoms.

A Ps temperature, the subsequent charge exchange cross section (to form $\bar{H}$) drops rapidly [17]. By carefully tailoring the topology of the target material's pores however, a degree of control appears possible [18].

Measurements using positron annihilation lifetime spectroscopy have shown that the ortho-Ps fraction (outside the sample) can exceed 20% in silicon-based polymer materials cooled to 50 K [19]. This is shown in Fig. 2, where the ortho-Ps fraction is displayed as a function of implantation energy for different target materials and temperatures. In other experimental work, it was shown that the energy profile of Ps emitted from the surface of a silica film at room temperature followed a Maxwell-Boltzmann distribution and was compatible with the Ps being fully thermalized [20]. We are currently conducting experiments in order to determine the optimal converter material and $e^+$ energy in terms of ortho-Ps yield. Furthermore, we are investigating how well the emitted Ps is thermalized at very low target temperatures.

The photo-excitation of Ps to Rydberg states requires photon energies close to the binding energy of 6.8 eV. Laser systems at the corresponding wavelengths (180 nm) are not commercially available. We are therefore planning to perform a two-step excitation, from the ground state to the $n = 3$ state ($\lambda = 205$ nm), and then to the $n = 35$ Rydberg band ($\lambda \sim 1670$ nm) [21]. Two pulsed-laser systems, both of which are pumped by a Q-switched Nd:YAG laser (1064 nm, 4 ns, 140 mJ), are currently under development. Both systems must provide sufficient power to excite the emitted Ps within a few ns and must be geometrically matched to the expanding cloud. Furthermore, the bandwidths of the lasers must be tailored to the transition linewidth broadened due to the Doppler effect as well as level splitting due to the motional Stark effect and the linear and quadratic Zeeman effect.

2.2. Antihydrogen recombination and beam formation

An $\bar{H}$ recombination scheme based on resonant charge exchange with Ps was first proposed almost twenty years ago [22]. The reaction proceeds according to the equation

$$Ps^* + \bar{p} \rightarrow \bar{H}^* + e^-$$

where the star denotes a highly excited Rydberg state. This reaction owes its appeal to the fact that the cross-section scales approximately with the fourth power of the principal quantum
number. In addition, it creates $\bar{H}$ in a narrow and well-defined band of final states. Most importantly, $\bar{H}$ formed with $\bar{p}$ at rest is created with a velocity distribution dominated by the $\bar{p}$ temperature, hence the surrounding (cryogenic) environment [23]. This is in stark contrast to the rather high $\bar{H}$ temperature observed when using the nested-well technique pioneered by ATRAP and ATHENA [24, 25]. Our proposed technique is conceptually similar to a charge exchange technique based on Rydberg cesium [26] which has been successfully demonstrated by ATRAP [27], but offers greater control of the final state distribution of $\bar{H}$ and, more importantly, allows pulsed production of $\bar{H}$.

The principle is illustrated in Fig. 1. The Ps emitted from the porous insulator material are excited to Rydberg states. They then traverse a Penning trap region in which $\sim 10^5 \bar{p}$ have been accumulated, stored and cooled to O(100 mK). The low temperature requirement on the antiprotons comes from the requirement that the antihydrogen atoms that will be formed will have a velocity that is low compared to the velocity of several 100 m/s that they will achieve after acceleration. To reach such a low temperature, the Penning trap is coupled to a 50 mK dilution refrigerator, and the antiprotons are coupled to the low temperature environment by embedding them in an electron plasma. The later will cool through synchrotron radiation, as well as through a tuned circuit; furthermore, evaporative cooling of the pre-cooled antiprotons is being envisaged. The charge exchange cross-section is very large ($\sim 10^7 \text{A}^2$ for $n = 35$) and reaches a maximum when the $e^+$ and $\bar{p}$ relative velocities are matched. Taking into account the corresponding kinetic energy, as well as a smaller contribution due to converted internal energy, $\bar{H}$ is created at velocities of $25 \sim 80 \text{ m/s}^{-1}$.

While neutral atoms are not sensitive (to first order) to constant electric fields, they do experience a force when their electric dipole moment is exposed to an electric-field gradient. Since the dipole moment scale approximately with the square of the principal quantum number, Rydberg atoms are especially amenable to being manipulated in this way. Such so-called Stark acceleration (and deceleration) has been successfully demonstrated, among others, by one of the AEGIS groups with (ordinary) hydrogen after excitation to the $n = 22, 23, 24$ states [28, 29]. In these experiments, accelerations of $2 \times 10^8 \text{ms}^{-2}$ were achieved using two pairs of electrodes.

**Figure 2.** Ortho-Ps fraction as a function of implantation energy into a closed porosity (top figure) and open porosity (bottom figure) target for different target temperatures between 13K and 200K.
Figure 3. Schematic of the principle of the Moiré deflectometer consisting of two gratings coupled to a position-sensitive detector.

at right angles to each other. A hydrogen beam traveling at 700 $ms^{-1}$ was stopped within 5 $\mu$s over a distance of only 1.8 mm. We intend to use a similar field configuration, generated by axially split electrodes within the cylindrical geometry of a Penning trap, to accelerate the formed $\bar{H}$ atoms to about 400 $ms^{-1}$ in the direction of the deflectometer apparatus.

2.3. Gravity measurement

In matter wave interferometers of the Mach-Zehnder type [30, 31], three identical gratings are placed at equal distances $L$ from each other. The first two gratings produce an interference pattern at the location of the third. That pattern has the same period $d$ as the gratings, and its position perpendicular to the diffracted particle beam can be determined precisely by displacing the third grating and recording the overall transmission with a particle detector. Under the influence of gravity, the interference pattern is vertically displaced (it falls) by a distance

$$\delta x = -gT^2$$

where $g$ is the local gravitational acceleration and $T$ is the time of flight $L/v$ between each pair of gratings of a particle beam traveling at velocity $v$.

Contrary to such true interferometers, which place a very stringent limit on the acceptable beam divergence (and thus antihydrogen temperature at production), the so-called Moiré deflectometer, in which diffraction on the gratings is replaced by a (classical) shadow pattern of those particles that converge onto the third grating, works in the classical regime. Interestingly, the gross characteristics of the interferometer are retained [32], in particular, the vertical displacement of the interference pattern according to Eq. (2). A three-grating Moiré deflectometer has been used to measure the local gravitational acceleration to a relative precision of $2 \times 10^{-4}$ with a beam of argon atoms traveling at an average velocity of 750 $ms^{-1}$ [32]. In departing from the three-grating deflectometer, we intend to replace the third grating by a position-sensitive silicon strip detector (see Fig. 3). Antihydrogen atoms impacting on the detector plane annihilate, and the impact point can be reconstructed by means of the energy deposited locally by the annihilation products.
The value of $g$ is extracted from two primary observables ($\text{time of flight } T$ and vertical displacement of the fringe pattern $\delta x$). The periodic nature of the arrangement means that for a given value of $T$, the impact point will have dropped by a well-defined amount $\delta x(T)$, modulo the grating period. By varying the accelerating field gradient (and thus varying the velocity of the antihydrogen atoms), or simply through the spread in velocity of the ensemble of antihydrogen atoms, the quadratic dependence of $\delta x$ on $T$ is probed.

Our simulations have shown that in order to perform a measurement of $g$ to $1\%$ relative precision, about $10^5$ H atoms at a temperature of $100 \text{ mK}$ will be required. equivalent to several months of data taking at the AD. In these simulations, a grating period of $80 \mu m$ was used, and a finite detector resolution of $10 \mu m$ was taken into account.

3. Summary and outlook

Construction has started on the AEGIS experiment, whose design is based upon the broad experience gained with the ATHENA and ATRAP experiments at the AD, a series of ongoing related experiments, tests and developments, as well as extensive simulations of critical processes (charge exchange production of $\bar{H}$, Stark acceleration and propagation through the Moiré deflectometer, resolution of the position-sensitive detector located at the end of the deflectometer). The proposed gravity measurement merges in a single experimental apparatus technologies already demonstrated or based on reasonable additional development.

For the initial phase of the experiment, obtaining samples of anti-atoms at $100 \text{ mK}$ is an essential requirement. Gravity measurements with even higher precision, as well as competitive CPT tests through spectroscopy, are desirable, but will necessitate the development of novel techniques to attain even colder $\bar{H}$ ensembles. The experiment has been designed with flexibility of the apparatus in mind, in order to allow a number of techniques, which may lead to such physics topics, to be implemented. One natural extension of the modular design is to incorporate, in a future stage, a magnetic decelerator and trap for $\bar{H}$, which will be spatially separated from the region where the anti-atoms are produced, similar to the devices currently being used to trap and study H atoms [33, 34]. The experience gained in the first phase of AEGIS with the formation of an $\bar{H}$ beam will be used to optimize the design of such a trapping system.

[1] http://aegis.web.cern.ch/aegis
[2] J.F. Donoghue, Phys. Rev. D 50 (1994) 3874.
[3] J. Sherk, Phys. Lett. B 88 (1979) 265.
[4] L.I. Schiff, Phys. Rev. Lett. 1 (1958) 254.
[5] L.I. Schiff, Proc. Natl. Acad. Sci. 45 (1959) 69. (http://www.pnas.org/cgi/reprint/45/1/69)
[6] R.J. Hughes, M.H. Holzscheiter, Phys. Rev. Lett. 66 (1991) 854.
[7] M.M. Nieto, T. Goldman, Phys. Rep. 205 (1991) 221.
[8] D. Alves, M. Jankowiak, P. Saraswat, arXiv:0907.4110v1
[9] B. Heckel et al., Adv. Space Res. 25 (2000) 1225
[10] W.M. Fairbank et al., in: B. Bertotti (Ed.), Proc. Int. School Phys. Enrico Fermi, Academic Press, New York, 1974, p. 310.
[11] M.H. Holzscheiter et al., Nucl. Phys. A 558 (1993) 709c.
[12] M. Amoretti et al., ATHENA Collaboration, Nature 419 (2002) 456.
[13] G. Gabrielse et al., ATRAP Collaboration, Phys. Rev. Lett. 89 (2002) 213401.
[14] L.V. Jorgensen et al., ATHENA Collaboration, Phys. Rev. Lett. 95 (2005) 025002.
[15] G. Gabrielse et al., ATRAP Collaboration, Phys. Lett. B 548 (2002) 140.
[16] D.W. Gidley, H.-G. Peng, R.S. Vallery, Annu. Rev. Mater. Res. 36 (2006) 49.
[17] C. Canali, publication in preparation.
[18] S. Mariazzi, A. Salemi, R. Brusa, Phys. Rev. B 78 (2008) 085428.
[19] S. Mariazzi, A. Salemi, R. Brusa, Appl. Surf. Sci. 255 (2008) 191.
[20] R.S. Vallery, P.W. Zitzewitz, D.W. Gidley, Phys. Rev. Lett. 90 (2003) 203402.
[21] F. Castelli, L. Boscoli, S. Cialdi, D. Comparat, Phys. Rev. A 78 (2008) 052512.
[22] M. Charlton, Phys. Lett. A 143 (1990) 143.
[23] B.I. Deutch et al., Hyperfine Interact. 76 (1993) 153.
[24] G. Gabrielse et al., ATRAP Collaboration, Phys. Rev. Lett. 93 (2004) 073401
[25] N. Madsen et al., ATHENA Collaboration, Phys. Rev. Lett. 94 (2005) 033403
[26] E.A. Hessels, D.M. Homan, M.J. Cavagnero, Phys. Rev. A 57 (1998) 1668
[27] C.H. Storry et al., ATRAP Collaboration, Phys. Rev. Lett. 93 (2004) 263401
[28] E. Vliegen, F. Merkt, J. Phys. B 39 (2006) L241.
[29] E. Vliegen et al., Phys. Rev. A 76 (2007) 023405.
[30] L. Zehnder, Z. Instrumentenkunde 11 (1891) 275.
[31] L. Mach, Z. Instrumentenkunde 12 (1892) 89.
[32] M.K. Oberthaler et al., Phys. Rev. A 54 (1996) 3165.
[33] S. Hogan and F. Merkt, Phys. Rev.Lett. 100 (2008) 043001
[34] S. Hogan, A. Wiederkehr, H. Schmutz and F. Merkt, Phys. Rev. Lett. 101 (2008) 143001