On Epsilon Expansions of Four-loop Non-planar Massless Propagator Diagrams

R.N. Lee\textsuperscript{1,a}, A.V. Smirnov\textsuperscript{2,4,b}, and V.A. Smirnov\textsuperscript{3,4,c}

\textsuperscript{1} Budker Institute of Nuclear Physics and Novosibirsk State University, 630090, Novosibirsk, Russia
\textsuperscript{2} Scientific Research Computing Center, Moscow State University, 119992 Moscow, Russia
\textsuperscript{3} Skobeltsyn Institute of Nuclear Physics of Moscow State University, 119992 Moscow, Russia
\textsuperscript{4} Institut für Theoretische Teilchenphysik, KIT, 76128 Karlsruhe, Germany

Abstract. We evaluate three typical four-loop non-planar massless propagator diagrams in a Taylor expansion in dimensional regularization parameter $\epsilon = (4 - d)/2$ up to transcendentality weight twelve, using a recently developed method of one of the present coauthors (R.L.). We observe only multiple zeta values in our results.

PACS. XX.XX.XX No PACS code given

Analytic results for one-scale multiloop Feynman integrals in a Laurent expansion in $\epsilon = (4 - d)/2$ are expressed as linear combinations of transcendental constants with rational coefficients. The set of these constants essentially depends on the type of Feynman integrals. Probably, the simplest type of one-scale Feynman integrals are massless propagator integrals depending on one external momentum. Here the world record is set at four loops — see Ref. \textsuperscript{[1]} where all the corresponding master integrals were analytically evaluated in an epsilon expansion up to transcendentality weight seven.

Practical calculations show that only multiple zeta values (MZV) (see, e.g., \textsuperscript{[2]}) appear in results. Brown proved \textsuperscript{[3]} that convergent scalar massless planar propagator diagrams with the degree of divergence $\omega \equiv 4h - 2L = -2$ (where $h$ and $L$ are numbers of loops and edges, correspondingly) up to five loops contain only MZV in their epsilon expansions. (In two loops, a proof was earlier presented in Ref. \textsuperscript{[4].}) He also proved that for the three non-planar diagrams depicted in Fig. 1 every coefficient in a Taylor expansion in $\epsilon$ is a rational linear combination of MZV and Goncharov’s polylogarithms \textsuperscript{[5]} with sixth roots of unity as arguments.

The goal of this brief communication is to study these diagrams experimentally. We present results in an epsilon expansion up to transcendentality weight twelve. To do this we apply the DRA method recently suggested by one of the authors (R.L.), Ref. \textsuperscript{[6]}. The method is based on the use of dimensional recurrence relations (DRR) \textsuperscript{[7]} and analytic properties of Feynman integrals as functions of the parameter of dimensional regularization, $d$, and was already successfully applied in previous calculations \textsuperscript{[8,9,10,11,12]}. To apply this method it is essential to perform an integration by parts (IBP) \textsuperscript{[13]} reduction of integrals that participate in dimensional recurrence relations to master integrals. To do this, we use the C++ version of the code FIRE \textsuperscript{[4]}.

To study analytic properties of solutions of dimensional recurrence relations, i.e. to reveal the position and the order of poles in $d$ in a basic stripe, we used a sector decomposition \textsuperscript{[15,16,17]} implemented in the code FIESTA \textsuperscript{[17,18]}. To fix remaining constants in the homogeneous solutions of dimensional recurrence relations it was quite sufficient for us to use analytic results for the four-loop massless propagators master integrals \textsuperscript{[1]} (confirmed numerically by FIESTA \textsuperscript{[19]}). Finally, after obtaining results for master integrals in terms of multiple series we calculated resulting coefficients at powers of $\epsilon$ numerically with a high precision and then applied the PSLQ algorithm \textsuperscript{[20]}. We also applied the code HPL \textsuperscript{[21]}, a Mathematica code dealing with harmonic polylogarithms \textsuperscript{[22]} and MZV.

For all the three diagrams of Fig. 1, we performed evaluation up to transcendentality weight twelve where
the basis of transcendental numbers we used includes 48 constants. Coefficients in our results turn out to be cumbersome, in particular, such are the coefficients at $\pi^{12}$ in subsequent formulæ. Therefore, to achieve a successful calculation by PSLQ, we were forced to perform numerical calculations with a high accuracy. In fact, the accuracy of 800 digits was enough to obtain results by PSLQ. However, to be on the safe side, we checked our results by numerical evaluation with the accuracy of 1500 digits.

The first diagram of Fig. 1 is a master diagram. This is nothing but $M_{45}$ in Fig. 2 of Ref. 11 where all the master integrals for four-loop massless propagators are shown. Following the method of 12 we needed, first to calculate lower master integrals $M_{01}, M_{11}, M_{13}, M_{14}, M_{21}, M_{27}, M_{45}$. Eventually, we arrived at the following result which is made homogeneously transcendental by pulling out an appropriate rational function of $\epsilon$ and normalizing it at the fourth power of the one-loop integral of (i.e. $M_{31}$ according to the notation of 12):

$$\frac{M_{45}(4 - 2\epsilon)}{\epsilon^4 M_{31}(4 - 2\epsilon)} = \left(1 - 2\epsilon\right)^3 \left(\frac{36\pi^2}{5} + 378\pi \right) \epsilon + \left(-\frac{427\pi^8}{1500} + \frac{3024\zeta_5}{5} - \frac{22}{3} \pi^6 \zeta_3 \right) \epsilon^2 + \left(-\frac{60329\pi^{10}}{24948} - \frac{183}{5} \pi^4 \zeta_3 \right) \epsilon^3 + \left(-\frac{29578\pi^6 \zeta_5}{315} - \frac{101952\zeta_5^2}{5} - \frac{19771\pi^4 \zeta_7}{50} - \frac{325152\pi^2 \zeta_9}{500} \right) \epsilon^4 + \left(-\frac{9898891597\pi^{12}}{170270100000} - \frac{19242\pi^6 \zeta_3}{3} + \frac{37216\pi^2 \zeta_7}{3} \right) \epsilon^5 + \left(-\frac{10280\zeta_5^3}{9} + \frac{210112\zeta_5^2 \zeta_3}{3} + \frac{660320\zeta_5 \pi^2 \zeta_7}{3} + \frac{210079\zeta_5 \pi^4 \zeta_9}{9} + \frac{945570\zeta_5}{9} + \frac{160424}{75} \pi^4 \zeta_5 \right) \epsilon^6 + O(\epsilon^7). \quad (1)

Here $\zeta_{m_1, \ldots, m_k}$ are MZV given by

$$\zeta(m_1, \ldots, m_k) = \sum_{m_1=1}^{\infty} \sum_{m_2=1}^{m_1-1} \cdots \sum_{m_k=1}^{m_{k-1}-1} \prod_{j=1}^{k} \frac{\text{sgn}(m_j)}{i^{m_j}}. \quad (2)

The other two diagrams of Fig. 1 are not master integrals and, consequently, they do not appear in Fig. 2 of Ref. 11. We found their reduction to lower master integrals by FIRE and then evaluated resulting master integrals by our technique. Let us stress that the master integrals involved in the IBP reduction of the second and the third non-planar diagrams of Fig. 1 are all planar diagrams. However, the result of Brown 13 quoted in the beginning of our paper is not applicable to such planar diagrams. (Two of them are $M_{35}$ and $M_{46}$ in Fig. 2 of Ref. 11, with $\omega = 0$, rather than $-2$) So, a priori, one could admit the presence of something additional to MZV in their epsilon expansions.

These are our results for them in the same normalization where we again managed to reveal homogenous transcendentality:

$$\frac{M_{46,2}(4 - 2\epsilon)}{\epsilon^4 M_{31}(4 - 2\epsilon)} = \left(1 - 2\epsilon\right)^3 \left(\frac{36\pi^2}{5} + \frac{36\pi^4 \zeta_3}{5} + \frac{189\pi \zeta_7}{2} \right) \epsilon + \left(-\frac{1559\pi^8}{700} + \frac{144\zeta_5}{5} - \frac{2916\zeta_5^3}{5} \right) \epsilon^2 + \left(-\frac{4\pi^6 \zeta_3}{21} - \frac{1392\pi^3}{3} \right) \epsilon^3 + \left(-\frac{51\pi^4 \zeta_5}{6} + \frac{2559\pi \zeta_9}{84} \right) \epsilon^4 + \left(-\frac{7584\pi^2 \zeta_9}{4} + \frac{20665\zeta_5}{84} \right) \epsilon^5 + \left(-\frac{313039\pi^8 \zeta_5}{31500} \right) \epsilon^6 + \frac{29412\pi^2 \zeta_9}{315} + \frac{299512\zeta_5}{160} + \frac{19068}{5} \zeta_5 \zeta_3 + \frac{234339\pi^4 \zeta_7}{200} \epsilon^7 + \left(-\frac{10793028885712}{68108040000} + \frac{42968}{315} \right) \epsilon^8 + \frac{85696\pi^2 \zeta_9}{3} \epsilon^9 + \left(-\frac{23510}{9} \right) \epsilon^{10} + \frac{31822 \pi^2 \zeta_5}{3} + \frac{90920}{3} \pi^2 \zeta_5 \zeta_7 + \frac{165244\pi \zeta_7}{6} \epsilon^{11} + \left(-\frac{6321395\zeta_9}{9} + \frac{45614}{75} \right) \epsilon^{12} + \frac{803569\zeta_9}{6} \epsilon^{13} + \left(-\frac{18184\zeta_6 \zeta_{4,1,1}}{1} \right) \epsilon^{14} + O(\epsilon^{15}). \quad (3)

$$\frac{M_{46,3}(4 - 2\epsilon)}{\epsilon^4 M_{31}(4 - 2\epsilon)} = \left(1 - 2\epsilon\right)^3 \left(\frac{36\pi^2}{5} + \frac{6\pi^4 \zeta_5}{5} + \frac{567\pi \zeta_7}{5} \right) \epsilon + \left(-\frac{211\pi^8}{1750} + \frac{3744\zeta_5 \pi^2 \zeta_7}{5} + \frac{1944\zeta_5^3}{5} \right) \epsilon^2 + \left(-\frac{68\pi^6 \zeta_3}{7} + \frac{288\pi^3}{5} \right) \epsilon^3 + \left(-\frac{30\pi^4 \zeta_5}{7} + \frac{2208\pi \zeta_9 \zeta_7}{7} \right) \epsilon^4 + \left(-\frac{12484\pi^6 \zeta_5}{125} - \frac{65952\zeta_5^2 \zeta_7}{25} + \frac{17538\pi \zeta_9^2 \zeta_7}{25} + \frac{65232\pi^2 \zeta_9}{25} \right) \epsilon^5 + \left(-\frac{36009\zeta_5}{10} + \frac{54432}{5} \zeta_5 \zeta_3 + \frac{43488}{5} \zeta_3 \zeta_5 \right) \epsilon^6 + \left(-\frac{2972813873\pi^{12}}{1064118125} + \frac{15056}{105} \pi^6 \zeta_3^2 - \frac{29608\zeta_4}{3} \epsilon^7 + \frac{3640}{3} \pi^4 \zeta_5 \zeta_3 \epsilon^8 - \frac{8176\pi^2 \zeta_5^2}{2} + \frac{23360\pi^2 \zeta_3 \zeta_7}{5} + \frac{1091724\zeta_5 \zeta_7}{1} \epsilon^9 + \frac{4198640\zeta_9 \zeta_5}{3} \epsilon^{10} + O(\epsilon^{11}) \right). \quad (4)
\[-\frac{6752}{25} \pi^4 \zeta_{3,3} - 7008 \pi^2 \zeta_{7,3} + 95936 \zeta_{9,3} + 14016 \zeta_{6,4,1,1}\right) \epsilon^6 + O(\epsilon^7) \right]. \tag{4}

We see that only MZV are present in our results. Although Goncharov’s polylogarithms at sixth roots of unity were allowed to appear according to the analysis of Ref. [3] they have not appeared. In fact, such transcendental numbers do appear in epsilon expansions of some classes of massive Feynman integrals [23, 24, 25, 26].

Taking our results into account one could try to prove that there are only MZV in massless propagator diagrams. Unfortunately, the way Brown proceeded [4] in the case of a few concrete diagrams can hardly be generalized to other situations because his analysis was based on the possibility to perform recursive integrations in Feynman parameters, as long as the denominator of the integrand is linear with respect to them, so that some other approach is needed. On the other hand, attempts to discover new types of transcendental numbers, in addition to MZVs, meet more and more difficulties because we have to turn to more loops and more terms of ε-expansions. Still this is certainly possible at the current level of our possibilities so that we are planning to do this in the nearest future.

Acknowledgments.

This work was supported by the Russian Foundation for Basic Research through grant 11-02-01196 and by DFG through SFB/TR 9 “Computational Particle Physics”. The work of R.L. was also supported through Federal special-purpose program “Scientific and scientific-pedagogical personnel of innovative Russia”. R.L. gratefully acknowledges Karlsruhe Institut für Theoretische Teilchenphysik for warm hospitality and financial support during his visit. We are grateful to David Broadhurst and Mikhail Kalmykov for instructive discussions.

References

1. P. A. Baikov and K. G. Chetyrkin, Nucl. Phys. B 837 (2010) 186 [arXiv:1004.1153 [hep-ph]].
2. J. Blumlein, D. J. Broadhurst and J. A. M. Vermaseren, Comput. Phys. Commun. 181 (2010) 582 [arXiv:0907.2557 [math-ph]].
3. F. Brown, Commun. Math. Phys. 287, 925 (2009) [arXiv:0804.1660 [math.AG]].
4. I. Bierenbaum and S. Weinzierl, Eur. Phys. J. C 32 (2003) 67.
5. A. B. Goncharov, “Multiple polylogarithms, cyclotomy and modular complexes,” Math. Research Letters 5 (1998) 497. “Multiple polylogarithms and mixed Tate motives,” arXiv:math/0110350.
6. R. N. Lee, Nucl. Phys. B 830 (2010) 474 [arXiv:0911.0252 [hep-ph]].
7. O. V. Tarasov, Phys. Rev. D 54 (1996) 6479.
8. R. N. Lee, A. V. Smirnov and V. A. Smirnov, JHEP 1004 (2010) 020 [arXiv:1001.2887 [hep-ph]].
9. R. N. Lee, A. V. Smirnov and V. A. Smirnov, Nucl. Phys. Proc. Suppl. 205-206 (2010) 308 [arXiv:1005.0362 [hep-ph]].
10. R. N. Lee, Nucl. Phys. Proc. Suppl. 205-206 (2010) 135 [arXiv:1007.2256 [hep-ph]].
11. R. N. Lee and I. S. Terekhov, JHEP 1101 (2011) 068 [arXiv:1010.6117 [hep-ph]].
12. R. N. Lee and V. A. Smirnov, JHEP 1102 (2011) 102 [arXiv:1010.3554 [hep-ph]].
13. K. G. Chetyrkin and F. V. Tkachov, Nucl. Phys. B 192 (1981) 159.
14. A. V. Smirnov, JHEP 0810, 107 (2008) [arXiv:0807.3243 [hep-ph]].
15. T. Binoth and G. Heinrich, Nucl. Phys. B, 585 (2000) 741; Nucl. Phys. B, 680 (2004) 375; Nucl. Phys. B, 693 (2004) 134; G. Heinrich, Int. J. of Modern Phys. A, 23 (2008) 10. [arXiv:0803.4177], J. Carter and G. Heinrich, arXiv:1011.5493 [hep-ph].
16. C. Bogner and S. Weinzierl, Comput. Phys. Commun. 178 (2008) 596 [arXiv:0709.4092 [hep-ph]]; Nucl. Phys. Proc. Suppl. 183 (2008) 256 [arXiv:0806.4307 [hep-ph]].
17. A. V. Smirnov and M. N. Tentyukov, Comput. Phys. Commun. 180 (2009) 735 [arXiv:0807.4129 [hep-ph]].
18. A. V. Smirnov, V. A. Smirnov and M. Tentyukov, Comput. Phys. Commun. 182 (2011) 790 [arXiv:0912.0158 [hep-ph]].
19. A. V. Smirnov and M. Tentyukov, Nucl. Phys. B 837 (2010) 40 [arXiv:1004.1149 [hep-ph]].
20. H. R. P. Ferguson, D. H. Bailey and S. Arno, Math. Comput. 68, (1999) 351, NASA–Ames Technical Report, NASA–96–005.
21. D. Maitre, Comput. Phys. Commun. 174 (2006) 222 [arXiv:hep-ph/0507152].
22. E. Remiddi, J. A. M. Vermaseren, Int. J. Mod. Phys. A15 (2000) 725 [hep-ph/9905237].
23. D. J. Broadhurst, Eur. Phys. J. C 8 (1999) 311.
24. J. Fleischer and M. Yu. Kalmykov, Phys. Lett. B 470 (1999) 168.
25. M. Yu. Kalmykov, Nucl. Phys. B 718 (2005) 276.
26. M. Yu. Kalmykov and B. A. Kniehl, Nucl. Phys. Proc. Suppl. 205-206 (2010) 129 [arXiv:1007.2373 [math-ph]].