Nonequilibrium Thermodynamical Inequivalence of Quantum Stress-energy and Spin Tensors

F. Becattini, L. Tinti

Università di Firenze and INFN Sezione di Firenze, Florence, Italy

It is shown that different pairs of stress-energy and spin tensors of quantum relativistic fields related by a pseudo-gauge transformation, i.e., differing by a divergence, imply different mean values of physical quantities in thermodynamical nonequilibrium situations. Most notably, transport coefficients and the total entropy production rate are affected by the choice of the spin tensor of the relativistic quantum field theory under consideration. Therefore, at least in principle, it should be possible to disprove a fundamental stress-energy tensor and/or to show that a fundamental spin tensor exists by means of a dissipative thermodynamical experiment.

I. INTRODUCTION

In recent years, there has been a considerable interest in theoretical relativistic hydrodynamics and its most general form including dissipative terms [1]. This renewed interest has been mainly triggered by its successful application to the description of the Quark Gluon Plasma dynamical evolution in ultrarelativistic heavy ion collisions [2]. Relativistic hydrodynamics can be seen as the theory describing the dynamical behaviour of the mean value of the quantum stress-energy tensor $\hat{T}^{\mu\nu}$, that is $\text{tr}(\hat{\rho}\hat{T}^{\mu\nu})$. This tensor is generally assumed to be symmetric, although in special relativity it does not need to be such if it is accompanied by a non-vanishing rank 3 tensor, the so-called spin tensor $\hat{S}^{\lambda,\mu\nu}$.

In fact, in special relativistic quantum field theory, starting from particular stress-energy and spin tensors, different pairs can be generated (and are generally related) by means of a pseudo-gauge transformation [3, 4] preserving the total energy, momentum and angular momentum:

$$\begin{align*}
\hat{T}'^{\mu\nu} &= \hat{T}^{\mu\nu} + \frac{1}{2} \partial_\alpha \left( \hat{\Phi}^{\alpha,\mu\nu} - \hat{\Phi}^{\mu,\alpha\nu} - \hat{\Phi}^{\nu,\alpha\mu} \right), \\
\hat{S}'^{\lambda,\mu\nu} &= \hat{S}^{\lambda,\mu\nu} + \partial_\alpha \hat{Z}^{\alpha\lambda,\mu\nu}
\end{align*}$$

where $\hat{\Phi}$ is a rank three tensor field antisymmetric in the last two indices (often called and henceforth referred to as superpotential) and $\hat{Z}$ a rank four tensor antisymmetric in the pairs $\alpha\lambda$ and $\mu\nu$.

In a previous paper [3] we have shown that indeed different pairs $(\hat{T}, \hat{S})$ and $(\hat{T}', \hat{S}')$ are in general thermodynamically inequivalent as they imply different mean values of physical quantities for a rotating system at equilibrium. Particularly, for the free Dirac field, we showed that the canonical and Belinfante (obtained from the canonical one by setting $\hat{\Phi} = \hat{S}$ and $\hat{Z} = 0$ in $\hat{1}$), hence with a vanishing new spin tensor $\hat{S}'$) quantum stress-energy tensors result in different mean values for the momentum density and the total angular momentum density.

The thermodynamical inequivalence is (at least in our view) surprising because it was commonly believed that the only physical phenomenon which can discriminate between stress-energy tensors of a fundamental quantum field theory related by a transformation like $\hat{1}$ is gravity, or, in other words, the coupling to a metric tensor. In this paper we reinforce our previous finding by showing that the inequivalence extends to nonequilibrium thermodynamical quantities, specifically entropy production and transport coefficients. In summary, we will show that the use of different stress-energy tensors, related by $\hat{1}$, to calculate transport coefficients with the relativistic Kubo formula leads, in general, to different results. Therefore, at least in principle, an extremely accurate measurement of transport coefficients or total entropy in an experiment where dissipation is involved, would allow to disprove a candidate stress-energy or spin tensor, with obvious important consequences in relativistic gravitational theories. This finding means, in other words, that the existence of a fundamental spin tensor affects the microscopic number of degrees of freedom, or at least on how quickly macroscopic information gets converted into microscopic, namely on entropy generation.

The paper is organized as follows: in Sect. $\hat{1}$ we will extend the framework of the nonequilibrium density operator introduced by Zubarev [6] to the case of a non-vanishing spin tensor. In Sect. $\hat{1}$ it will be shown that the nonequilibrium density operator is not invariant under a pseudo-gauge transformation $\hat{1}$, that is it does depend on the chosen couple of stress-energy and spin tensor. In Sect. $\hat{1}$ we will provide a general formula for the change of mean values of observables and we will determine how entropy is affected by a pseudo-gauge transformation. In Sect. $\hat{1}$ we will show that transport coefficients are also modified and, particularly, we will focus on the modification of the Kubo formula for shear viscosity. Finally, in Sect. $\hat{1}$ we will discuss the implications of this finding and draw our conclusions.
II. NONEQUILIBRIUM DENSITY OPERATOR

A suitable formalism to calculate transport coefficients for relativistic quantum fields without going through kinetic theory was developed by Zubarev \([4, 5]\), extending to the relativistic domain a formalism already introduced by Kubo \([10]\). In this approach, a non-equilibrium density operator is introduced which reads \([11]\):

\[
\hat{\rho} = \frac{1}{Z} \exp[-\hat{\Upsilon}] = \frac{1}{Z} \exp \left[ - \lim_{\epsilon \to 0} \frac{\epsilon}{\epsilon} \int_{-\infty}^{t'} dt \; e^{\epsilon(t-t')} \int d^3x \left( \hat{T}^{0\nu} \beta_\nu(x) - \hat{j}^0 \xi(x) \right) \right]
\]  

(2)

where \(\hat{\Upsilon}\) is a conserved current, the four-vector field \(\beta\) is a point-dependent inverse temperature four-vector \((\beta = u/T_0, \; u\) being a four-velocity field and \(T_0\) the comoving or invariant temperature\) and \(\xi = \mu_0/T_0\) a scalar function whose physical meaning is that of a point-dependent ratio between comoving chemical potential \(\mu_0\) and comoving temperature \(T_0\); the \(Z\) factor is analogous to a partition function, i.e. a normalization factor to have \(\text{tr}\hat{\rho} = 1\). The operators in the exponential of Eq. (2) are in the Heisenberg representation. It should be stressed that in the formula (2) covariance is broken from the very beginning by the choice of a specific inertial frame and its time. However, it can be shown that the operator \(\hat{\rho}\) is in fact time-independent \([11]\), namely independent of \(t'\), so that \(\hat{\rho}\) is a good density operator in the Heisenberg representation.

In the formula (2) the possible contribution of a spin tensor is simply disregarded; therefore, the formula is correct only if the stress-energy tensor is the symmetrized Belinfante one (or improved ones, see last section), whose associated spin tensor is vanishing. It is the aim of this Section to find the appropriate extension of the formula (2) with a spin tensor.

Using the identity:

\[
e^{\epsilon(t-t')} \left( \hat{T}^{0\nu} \beta_\nu(x) - \hat{j}^0 \xi(x) \right) = \left( \frac{\partial}{\partial x^\mu} \frac{e^{\epsilon(t-t')}}{\epsilon} \right) \left( \hat{T}^{\mu\nu} \beta_\nu(x) - \hat{j}^\mu \xi(x) \right)
\]

integrating by parts and taking into account the continuity equations \(\partial_\mu \hat{T}^{\mu\nu} = \partial_\mu \hat{j}^\mu = 0\), the operator \(\hat{\Upsilon}\) in Eq. (2) can be rewritten as:

\[
\hat{\Upsilon} = \int d^3x \left( \hat{T}^{0\nu} \beta_\nu(t', x) - \hat{j}^0 \xi(t', x) \right) + \lim_{\epsilon \to 0} \int_{-\infty}^{t'} dt \; e^{\epsilon(t-t')} \int dS \; n_i \left( \hat{T}^{\mu\nu} \beta_\nu(x) - \hat{j}^\mu \xi(x) \right)
\]

(3)

The first term the so-called local thermodynamical equilibrium one, which is defined by the same formula of the global equilibrium \([4, 5]\) with \(x\)-dependent four-temperature and chemical potentials, whereas the term dependent on their derivatives is interpreted as a perturbation.

At equilibrium, the right hand side should reduce to the known form, which, at least for the most familiar form of thermodynamical equilibrium with \(\beta^{eq} = (1/T, 0) = \text{const}\) and \(\xi^{eq} = \mu/T = \text{const}\) is readily recognized in the first

---

1 Throughout the paper, the four-vector \(x\) implies the time \(t\) and position vector \(x\), i.e. \(x = (t, x)\). The dependence of the stress-energy and spin tensor on \(x\) will always be understood.
term setting $β = β^{eq}$ and $ξ = ξ^{eq}$:

$$\hat{\mathbf{\omega}}^{eq} = \int d^3x \left( \hat{T}_{\mu\nu}^{\beta eq} - \hat{j}_0^{eq} \right) + \lim_{\epsilon \to 0} \int_{-\infty}^{t'} dt \, e^{\epsilon(t-t')} \int dS \, n_i \left( \hat{T}_{\mu
u}^{\beta eq} - \hat{j}_i^{eq} \right)$$

$$- \lim_{\epsilon \to 0} \int_{-\infty}^{t'} dt \, e^{\epsilon(t-t')} \int d^3x \left( \hat{\mathbf{\omega}}^{eq} \partial_\mu \beta^{eq} - \hat{j}_\mu \partial_\mu \zeta^{eq} \right) = \hat{H}/T - \mu \hat{Q}/T$$

$$+ \lim_{\epsilon \to 0} \int_{-\infty}^{t'} dt \, e^{\epsilon(t-t')} \int dS \, n_i \left( \hat{T}_{\mu\nu}^{\beta eq} - \hat{j}_\nu^{eq} \right) - \lim_{\epsilon \to 0} \int_{-\infty}^{t'} dt \, e^{\epsilon(t-t')} \int d^3x \left( \hat{T}_{\mu\nu} \partial_\mu \beta^{eq} - \hat{j}_\mu \partial_\mu \zeta^{eq} \right)$$

(4)

Hence, the two rightmost terms of (4) must vanish at equilibrium. Indeed, the surface term is supposed to vanish through a suitable choice of the field boundary conditions while the third term vanishes in view of the constancy of $\beta^{eq}$ and $\zeta^{eq}$. However, this is not the case for the most general form of equilibrium; in the most general form (see discussion in ref. [8]), whilst the scalar $\zeta^{eq}$ stays constant the four-vector $\beta$ fulfills a Killing equation, whose solution is [12]:

$$\beta^{eq}_\nu(x) = b^{eq}_\nu + \omega^{eq}_{\mu\nu} x^\mu$$

with both the four-vector $b^{eq}$ and the antisymmetric tensor $\omega^{eq}_{\mu\nu}$ constant. Therefore:

$$\partial_\mu \beta^{eq}_\nu = -\omega^{eq}_{\mu\nu}$$

which in general is non-vanishing, so that the third term on the right hand side of Eq. (4) survives. For instance, for the thermodynamical equilibrium with rotation [8], the tensor $\omega$ turns out to be:

$$\omega^{eq}_\nu = \omega/T \left( \delta^{3}_{\lambda} \delta^{2}_{\beta} - \delta^{3}_{\lambda} \delta^{2}_{\beta} \right)$$

(6)

$\omega$ being the angular velocity and $T$ the temperature measured by the inertial frame.

In order to find the appropriate generalization of the operator $\hat{\mathbf{\omega}}$, let us plug the formula (5) of general thermodynamical equilibrium into the (4):

$$\hat{\mathbf{\omega}}^{eq} = \int d^3x \left( \hat{T}_{\mu\nu}^{\beta eq} - \hat{j}_0^{eq} \right) + \lim_{\epsilon \to 0} \int_{-\infty}^{t'} dt \, e^{\epsilon(t-t')} \int dS \, n_i \left( \hat{T}_{\mu\nu}^{\beta eq} - \hat{j}_i^{eq} \right)$$

$$+ \lim_{\epsilon \to 0} \int_{-\infty}^{t'} dt \, e^{\epsilon(t-t')} \int d^3x \hat{T}_{\mu\nu} \omega^{eq}_{\mu\nu}$$

(7)

where $\partial_\nu \epsilon^{eq} = 0$ has been taken into account. For a symmetric stress-energy tensor $\hat{T}$, the last term vanishes, but if a spin tensor is present $\hat{T}$ may have an antisymmetric part. Particularly, from the angular momentum continuity equation:

$$\hat{T}_{\mu\nu} \omega^{eq}_{\mu\nu} = \frac{1}{2} \left( \hat{T}_{\mu\nu} - \hat{T}_{\nu\mu} \right) \omega^{eq}_{\mu\nu} = -\frac{1}{2} \partial_\lambda \hat{S}^{\lambda\mu\nu} \omega^{eq}_{\mu\nu}$$

(8)

so that the last term on the right hand side of Eq. (7) can be rewritten as:

$$\lim_{\epsilon \to 0} \int_{-\infty}^{t'} dt \, e^{\epsilon(t-t')} \int d^3x \hat{T}_{\mu\nu} \omega^{eq}_{\mu\nu} = \frac{1}{12} \partial_\lambda \hat{S}^{\lambda\mu\nu} \omega^{eq}_{\mu\nu}$$

$$= \frac{1}{2} \omega^{eq}_{\mu\nu} \lim_{\epsilon \to 0} \int d^3x \int_{-\infty}^{t'} dt \, e^{\epsilon(t-t')} \int d^3x \hat{T}_{\mu\nu} \omega^{eq}_{\mu\nu}$$

(9)

The first term on the right hand side of (9) can be integrated by parts, yielding:

$$-\frac{1}{2} \omega^{eq}_{\mu\nu} \lim_{\epsilon \to 0} \int d^3x \int_{-\infty}^{t'} dt \, e^{\epsilon(t-t')} \frac{\partial}{\partial t} \hat{S}^{\mu\nu}(t', x) + \frac{1}{2} \omega^{eq}_{\mu\nu} \lim_{\epsilon \to 0} \int_{-\infty}^{t'} dt \, e^{\epsilon(t-t')} \int d^3x \hat{S}^{0\mu\nu}(x)$$

(10)

Plugging the Eq. (10) into (9) and this into (7) we obtain:

$$\hat{\mathbf{\omega}}^{eq} = \int d^3x \left( \hat{T}_{\mu\nu}^{\beta eq} - \hat{j}_0^{eq} \right) + \lim_{\epsilon \to 0} \int_{-\infty}^{t'} dt \, e^{\epsilon(t-t')} \left[ b^{eq}_\nu \int dS \, n_i \hat{T}_{\mu\nu} - \zeta^{eq} \int dS \, n_i \hat{j}_i \right]$$

$$- \frac{1}{2} \omega^{eq}_{\mu\nu} \int dS \, n_i \left( \hat{S}^{0\mu\nu}(x) + \hat{S}^{1\mu\nu}(x) \right) + \frac{1}{2} \omega^{eq}_{\mu\nu} \lim_{\epsilon \to 0} \int_{-\infty}^{t'} dt \, e^{\epsilon(t-t')} \int d^3x \hat{S}^{0\mu\nu}(x)$$

(11)
where the surface term involving \( \hat{T} \) in Eq. (7) has been rearranged taking advantage of the antisymmetry of the \( \omega \) tensor. The surface terms in the above equations now are manifestly the total momentum flux, the charge flux and the \textit{total} angular momentum flux through the boundary. All of these terms are supposed to vanish at thermodynamical equilibrium through suitable conditions enforced on the field operators at the boundary, so that the (11) reduces to:

\[
\hat{\Upsilon}^{\text{eq}} = \int d^3x \left( \hat{T}^{\mu \nu} \beta^{\text{eq}}_{\nu} - \hat{J}^{0} \xi^{\text{eq}} - \frac{1}{2} \omega^{\text{eq}}_{\mu \nu} \hat{S}^{0,\mu \nu} \right) + \frac{1}{2} \omega^{\text{eq}}_{\mu \nu} \lim_{\varepsilon \to 0} \varepsilon \int_{-\infty}^{t'} dt \ e^{\varepsilon(t-t')} \int d^3x \ \hat{S}^{0,\mu \nu}(x)
\]  

(12)

The first term on the right hand side just gives rise to the desired form of the equilibrium operator. For instance, for a rotating system with \( \omega \) as in Eq. (6) one has [8]:

\[
\int d^3x \ \left( \hat{T}^{0 \nu} \beta^{\text{eq}}_{\nu} - \hat{J}^{0} \xi^{\text{eq}} - \frac{1}{2} \omega^{\text{eq}}_{\mu \nu} \hat{S}^{0,\mu \nu} \right) = \hat{H}/T - \mu \hat{Q}/T - \omega \hat{J}/T
\]

\( \hat{J} \) being the total angular momentum, which is the known form [13]. Nevertheless, the second term in Eq. (12) does not vanish and, thus, must be subtracted away with a suitable modification of the definition of the \( \hat{\Upsilon} \) operator. The form of the unwanted term demands the following modification of (4):

\[
\hat{\rho} = \frac{1}{Z} \exp[-\hat{\Upsilon}] = \frac{1}{Z} \exp \left[ - \lim_{\varepsilon \to 0} \varepsilon \int_{-\infty}^{t'} dt \ e^{\varepsilon(t-t')} \int d^3x \ \left( \hat{T}^{0 \nu} \beta_{\nu}(x) - \hat{J}^{0} \xi(x) - \frac{1}{2} \hat{S}^{0,\mu \nu} \omega_{\mu \nu}(x) \right) \right]
\]

(13)

where \( \omega_{\mu \nu}(x) \) is an antisymmetric tensor field which must reduce to the constant \( \omega_{\mu \nu}^{\text{eq}} \) tensor at equilibrium. It is easy to check, by tracing the previous calculations, that the equilibrium form of \( \hat{\Upsilon} \) reduces to the desired form:

\[
\hat{\Upsilon}^{\text{eq}} = \int d^3x \ \left( \hat{T}^{0 \nu} \beta^{\text{eq}}_{\nu} - \hat{J}^{0} \xi^{\text{eq}} - \frac{1}{2} \omega^{\text{eq}}_{\mu \nu} \hat{S}^{0,\mu \nu} \right)
\]

as the spin tensor term in Eq. (12) cancels out. Therefore, the operator [13] is the only possible extension of the nonequilibrium density operator with a spin tensor.

The new operator \( \hat{\Upsilon} \) can be worked out the same way as we have done when obtaining Eq. (4) from Eq. (2):

\[
\hat{\Upsilon} = \int d^3x \ \left( \hat{T}^{\mu \nu} \beta_{\nu}(t', x) - \hat{J}^{0}(t', x) - \frac{1}{2} \hat{S}^{0,\mu \nu} \omega_{\mu \nu}(t', x) \right)
\]

\[
+ \lim_{\varepsilon \to 0} \varepsilon \int_{-\infty}^{t'} dt \ e^{\varepsilon(t-t')} \int dS \ n_i \left( \hat{T}^{\mu \nu} \beta_{\nu}(x) - \hat{J}^{0} \xi(x) - \frac{1}{2} \hat{S}^{0,\mu \nu} \omega_{\mu \nu}(x) \right)
\]

\[
- \frac{1}{2} \lim_{\varepsilon \to 0} \varepsilon \int_{-\infty}^{t'} dt \ e^{\varepsilon(t-t')} \int d^3x \ \left( \hat{T}^{\mu \nu}_{S} (\partial_{\mu} \beta_{\nu}(x) + \partial_{\nu} \beta_{\mu}(x)) + \hat{T}^{\mu \nu}_{A} (\partial_{\mu} \beta_{\nu}(x) - \partial_{\nu} \beta_{\mu}(x) + 2 \omega_{\mu \nu}(x)) - \hat{S}^{\lambda,\mu \nu} \partial_{\lambda} \omega_{\mu \nu}(x) - 2 \hat{J}^\mu \partial_{\mu} \xi(x) \right)
\]

(14)

where:

\[
\hat{T}^{\mu \nu}_{S} = \frac{1}{2} (\hat{T}^{\mu \nu} + \hat{T}^{\nu \mu}) \quad \hat{T}^{\mu \nu}_{A} = \frac{1}{2} (\hat{T}^{\mu \nu} - \hat{T}^{\nu \mu})
\]

and the continuity equation for angular momentum has been used. The first term on the right hand side is the new local thermodynamical term whilst the third term can be further expanded to derive the relativistic Kubo formula of transport coefficients (see Appendix A).

**III. NONEQUILIBRIUM DENSITY OPERATOR AND PSEUDO-GAUGE TRANSFORMATIONS**

A natural requirement for the density operator [13] would be its independence of the particular couple of stress-energy and spin tensor, because one would like the mean value of any observable \( \hat{O} \):

\[
O \equiv \text{tr}(\hat{\rho} \hat{O})
\]
to be an objective one. In ref. [3] we showed that even at thermodynamical equilibrium with rotation this is not the case for the components of the stress-energy and spin tensor themselves because they change through the pseudo-gauge transformation [11]. However, at equilibrium, \( \hat{\rho} \) itself is a function of just integral quantities (total energy, angular momentum, charge) which are invariant under a transformation [11] provided that boundary fluxes vanish, so a specific operator \( \hat{O} \), including the components of a specific stress-energy tensor, does not change under [11]. However, it is not obvious that this feature persists in a nonequilibrium case, in fact we are going to show that, in general, this is not the case.

Let us consider the operator \( \hat{\Upsilon} \) in [13] and how it gets changed under a pseudo-gauge transformation [11] with \( \hat{Z} = 0 \). The new operator \( \hat{\Upsilon}' \) reads:

\[
\hat{\Upsilon}' = \hat{\Upsilon} + \frac{1}{2} \lim_{\varepsilon \to 0} \varepsilon \int_{-\infty}^{t'} dt \, e^{\varepsilon(t-t')} \int d^3x \left( \partial_\lambda \varphi^{0,\mu}_\nu(x) + \tilde{\Phi}^{0,\mu}_\nu \omega_{\mu\nu}(x) \right)
\]

where:

\[
\varphi^{0,\mu}_\nu = \hat{\Phi}^{\lambda,\mu\nu} - \hat{\Phi}^{\mu,\lambda\nu} - \hat{\Phi}^{\nu,\mu\lambda}
\]

is antisymmetric in the first two indices. We can rewrite Eq. (15) as:

\[
\hat{\Upsilon}' - \hat{\Upsilon} = \frac{1}{2} \lim_{\varepsilon \to 0} \varepsilon \int_{-\infty}^{t'} dt \, \int d^3x \, e^{\varepsilon(t-t')} \left[ \partial_\lambda \left( \varphi^{0,\mu}_\nu(x) \right) - \varphi^{0,\mu}_\nu \partial_\lambda \beta_\nu + \tilde{\Phi}^{0,\mu}_\nu \omega_{\mu\nu}(x) \right]
\]

after integration by parts. Let us now write the general fields \( \beta \) and \( \omega \) as the sum of the equilibrium values and a perturbation, that is:

\[
\beta(x) = \beta^\text{eq}(x) + \delta \beta(x) \quad \omega(x) = \omega^\text{eq} + \delta \omega(x)
\]

and work out first the equilibrium part of the right hand side of Eq. (17). As \( \partial_\lambda \beta^\text{eq} = -\omega^\text{eq}_\lambda \) one has:

\[
(\hat{\Upsilon}' - \hat{\Upsilon})|_{\text{eq}} = \frac{1}{2} \lim_{\varepsilon \to 0} \varepsilon \int_{-\infty}^{t'} dt \, \int d^3x \, e^{\varepsilon(t-t')} \left[ \int dS_n \varphi^{0,\mu}_\nu(x) + \int d^3x \left( \tilde{\Phi}^{0,\mu}_\nu \omega_{\mu\nu}(x) \right) \right]
\]

where we have used the Eq. (16) and the antisymmetry of indices of the superpotential \( \tilde{\Phi} \). By using the Eq. (5), the last expression can be rewritten as:

\[
\lim_{\varepsilon \to 0} \varepsilon \int_{-\infty}^{t'} dt \, e^{\varepsilon(t-t')} \left[ b_\nu \int dS_n \varphi^{0,\mu}_\nu + \frac{1}{2} \omega^\text{eq}_\mu \int dS_n (x^\nu \varphi^{0,\mu} - x^\nu \varphi^{0,\mu}) \right]
\]

The two surface integrals above are the additional four-momentum and the additional total angular momentum, in the operator sense, after having made a pseudo-gauge transformation [11] of the stress-energy and spin tensor. If the boundary conditions ensure that the momentum and total angular momentum fluxes vanish (in order to have conserved energy and momentum operators) for any couple \( (\hat{T}, \hat{\mathcal{S}}) \) of tensors, then the two fluxes in the above equations must vanish as well. Therefore, we can conclude that:

\[
\hat{\Upsilon}'|_{\text{eq}} = \hat{\Upsilon}|_{\text{eq}}
\]

---

2 It should be pointed out that the mean value of operators involving quantum relativistic fields are generally divergent (e.g. \( T^{00} \) for a free field has an infinite zero point value). To remove the infinities, the mean values must be renormalized, what can be simply done for free fields by using normal ordering in all expressions, including the density operator itself. Henceforth, it will be understood that all the mean values of operators are the renormalized ones.
Now, let us focus on the nonequilibrium perturbation of the \( \hat{T} \) operator.

\[
(\hat{T}' - \hat{T})|_{\text{non-eq}} = \frac{1}{2} \lim_{\varepsilon \to 0} \int_{-\infty}^{t'} dt \ e^{\varepsilon (t - t')} \left[ \int dS n_i \hat{\omega}^{0,\nu} \partial_\nu \delta \beta - \int d^3x \hat{\omega}^{0,\nu} \partial_\nu \delta \beta - \hat{\Phi}^{0,\mu \nu} \delta \omega_{\mu \nu} \right]
\]

\[
= \frac{1}{2} \lim_{\varepsilon \to 0} \int_{-\infty}^{t'} dt \ e^{\varepsilon (t - t')} \left[ \int dS n_i \hat{\omega}^{0,\nu} \partial_\nu \delta \beta - \int d^3x \left( \hat{\Phi}_{\lambda,0\nu} - \hat{\Phi}^{0,\lambda \nu} \right) \partial_\beta \delta \beta - \hat{\Phi}^{0,\mu \nu} \delta \omega_{\mu \nu} \right]
\]

\[
= \frac{1}{2} \lim_{\varepsilon \to 0} \int_{-\infty}^{t'} dt \ e^{\varepsilon (t - t')} \left[ \int dS n_i \hat{\omega}^{0,\nu} \partial_\nu \delta \beta - \int d^3x \hat{\Phi}_{\lambda,0\nu} (\partial_\lambda \delta \beta + \partial_\nu \delta \beta) - \hat{\Phi}^{0,\lambda \nu} \left( \frac{1}{2} (\partial_\lambda \delta \beta + \partial_\nu \delta \beta) + \delta \omega_{\lambda \nu} \right) \right]
\]

(20)

where the dependence of \( \delta \beta \) and \( \delta \omega \) on \( x \) is now understood. It can be seen that it is impossible to make this difference vanishing in general. One can get rid of the surface term by choosing a perturbation which vanishes at the boundary and the last term by locking the perturbation of the tensor \( \omega \) to that of the inverse temperature four-vector:

\[
\delta \omega_{\lambda \nu} (x) = -\frac{1}{2} (\partial_\lambda \delta \beta_{\nu} (x) - \partial_\nu \delta \beta_{\lambda} (x))
\]

(21)

but it is impossible to cancel out the term:

\[
\delta \hat{T} = -\frac{1}{2} \lim_{\varepsilon \to 0} \varepsilon \int_{-\infty}^{t'} dt \ e^{\varepsilon (t - t')} \int d^3x \hat{\Phi}_{\lambda,0\nu} (\partial_\lambda \delta \beta_{\nu} (x) + \partial_\nu \delta \beta_{\lambda} (x))
\]

(22)

unless in special cases, e.g. when the \( \hat{\Phi} \) is also antisymmetric in the first two indices.

We have thus come to the conclusion that the nonequilibrium density operator does depend, in general, on the particular choice of stress-energy and spin tensor of the quantum field theory under consideration. Therefore, the mean value of any observable in a non-equilibrium situation shall depend on that choice. It is worth stressing that this is a much deeper dependence on the stress-energy and spin tensor than what we showed in ref. [3] for thermodynamical equilibrium with rotation. Therein, mean values of the angular momentum densities and momentum densities were found to be dependent on the pseudo-gauge transformation [11] because the relevant quantum operators could be varied, but not because the density operator \( \hat{\rho} \) was dependent thereupon. In fact, at non-equilibrium, even \( \hat{\rho} \) varies under a transformation [11]. Note that, in principle, even the mean values of the total energy and momentum could be dependent on the quantum stress-energy tensor choice although boundary conditions ensure, as we have assumed, that the total energy and momentum operators are invariant under a transformation [11]. Again, this comes about because the density operator is not invariant under [11], in formula:

\[
\text{tr}(\hat{\rho} \hat{\Phi}^\mu) = \text{tr}(\hat{\rho} \hat{\Phi}^\mu) \neq \text{tr}(\hat{\rho} \hat{\Phi}^\mu)
\]

It must be pointed out that the variation of the Zubarev non-equilibrium density operator [22] depends on the gradients of the four-temperature field and it is thus a small one close to thermodynamical equilibrium. In the next Section we will show in more details how the mean values of observables change under a small change of the nonequilibrium density operator, or, in other words, when the system is close to thermodynamical equilibrium.

IV. VARIATION OF MEAN VALUES AND LINEAR RESPONSE

We will first study the general dependence of the mean value of an observable \( \hat{O} \) on the spin tensor by denoting by \( \delta \hat{T} \) the supposedly small variation, under a transformation [11], of the operator \( \hat{T} \). This can be either the one in Eq. (22) or the more general (only bulk terms) in Eq. (20). We have:

\[
\text{tr}(\hat{\rho} \hat{\Phi}^\mu) = \frac{1}{Z'} \text{tr}(\exp[-\hat{T}+\delta \hat{T}])\hat{O} = \frac{1}{Z'} \text{tr}(\exp[-\hat{T}+\delta \hat{T}])\hat{O}
\]

(23)

being \( Z' = \text{tr}(\exp[-\hat{T}+\delta \hat{T}]) \). We can expand in \( \delta \hat{T} \) at the first order (Zassenhaus formula):

\[
Z' \simeq Z - \text{tr} (\exp[-\delta \hat{T}])
\]

\[
\text{tr}(\exp[-\delta \hat{T}]) \simeq \text{tr} \left( \exp[-\hat{T}][(I - \frac{1}{2} [\hat{T}, \delta \hat{T}] - \frac{1}{6} [\hat{T}, [\hat{T}, \delta \hat{T}]] + \ldots)\hat{O} \right)
\]

(24)
hence, with \( \langle \rangle = \text{tr}(\rho) \), at the first order in \( \delta \tilde{Y} \):

\[
\text{tr}(\rho' \tilde{O}) \equiv \langle \tilde{O} \rangle' \simeq \langle \tilde{O} \rangle(1 + \langle \delta \tilde{Y} \rangle) - \langle \delta \tilde{Y} \rangle \tilde{O} + \frac{1}{2} \langle [\tilde{Y}, \delta \tilde{Y}] \tilde{O} \rangle - \frac{1}{6} \langle [\tilde{Y}, [\tilde{Y}, \delta \tilde{Y}]] \tilde{O} \rangle + \ldots
\]

which makes manifest the dependence of the mean value on the choice of the superpotential \( \tilde{\Phi} \).

As has been mentioned, close to thermodynamical equilibrium, the operator \( \delta \tilde{Y} \) is “small” and one can write an expansion of the mean value of the observable \( \tilde{O} \) in the gradients of the four-temperature field, according to relativistic linear response theory \[\text{[11]}\]. This method, just based on Zubarev’s nonequilibrium density operator method, allows to calculate the variation between the actual mean value of an operator and its value at local thermodynamical equilibrium for small deviations from it. In fact, it can be seen from Eq. (22) that the operator \( \delta \tilde{Y} \), from the linear response theory viewpoint, is an additional perturbation in the derivative of the four-temperature field and therefore the difference between actual mean values at first order turns out be (see Appendix A for reference):

\[
\Delta(\tilde{O}) \simeq - \lim_{\varepsilon \to 0} \frac{T}{2}\int_{-\infty}^{t'} dt \, e^{\varepsilon(t-t')} \int d^3x \, \langle [\tilde{\Phi}^\lambda, \rho (x) \tilde{O}] \rangle_0 (\partial_j \delta \beta_\lambda(x) + \partial_\nu \delta \beta_\lambda(x))
\]

(25)

where \( \langle \ldots \rangle_0 \) stands for the expectation value calculated with the equilibrium density operator, that is:

\[
\tilde{\rho}_0 = \frac{1}{Z_0} \exp[-\hat{H}/T + \mu \hat{Q}/T]
\]

(26)

Since \( \text{tr}(\tilde{\rho}_0 \tilde{\Phi}^{\lambda, 0\nu}, \tilde{O}) = \text{tr}(\tilde{\Phi}^{\lambda, 0\nu}[\tilde{O}, \tilde{\rho}_0]) \) the right hand side of (25) vanishes for all quantities commuting with the equilibrium density operator, notably total energy, momentum and angular momentum. Nevertheless, in principle, even the mean values of the conserved quantities are affected by the choice of a specific quantum stress-energy tensor, though at the second order in the perturbation \( \delta \tilde{\Phi} \).

We now set out to study the effect of the transformation \( \text{[11]} \) on the total entropy. In nonequilibrium situation, entropy is usually defined as \( \text{[13]} \) the quantity maximizing \( -\text{tr}(\tilde{\rho} \log \tilde{\rho}) \) with the constraints of fixed mean conserved densities. The solution \( \tilde{\rho}_{\text{LE}} \) of this problem is the local thermodynamical equilibrium operator, namely:

\[
\tilde{\rho}_{\text{LE}}(t) = \frac{\exp[-\int d^3x \, (\tilde{T}^{0\nu} \beta_\nu(x) - \tilde{j}^0 \xi(x) - \frac{1}{2} \tilde{S}^{0, \mu\nu} \omega_{\mu\nu}(x))]}{\text{tr}(\exp[-\int d^3x \, (\tilde{T}^{0\nu} \beta_\nu(x) - \tilde{j}^0 \xi(x) - \frac{1}{2} \tilde{S}^{0, \mu\nu} \omega_{\mu\nu}(x))])}
\]

(27)

which - as emphasized in the above equation - is explicitly dependent on time, unlike the Zubarev stationary nonequilibrium density operator \( \text{[13]} \); of course the time dependence is crucial to make entropy

\[
S = -\text{tr}(\tilde{\rho}_{\text{LE}} \log \tilde{\rho}_{\text{LE}})
\]

(28)

increasing in nonequilibrium situation. In order to study the effect of the transformation \( \text{[11]} \) on the entropy it is convenient to define:

\[
\tilde{T}_{\text{LE}} = \int d^3x \, (\tilde{T}^{0\nu} \beta_\nu(x) - \tilde{j}^0 \xi(x) - \frac{1}{2} \tilde{S}^{0, \mu\nu} \omega_{\mu\nu}(x))
\]

(29)

for which it can be shown that, with calculations similar to those in the previous section, the variation induced by the transformation \( \text{[11]} \) is:

\[
\delta \tilde{T}_{\text{LE}} = \frac{1}{2} \left\{ \int dS \, n_i \tilde{\chi}^{0, \nu} \delta \beta_\nu - \int d^3x \, [\tilde{\Phi}^{\lambda, 0\nu}(\partial_\lambda \delta \beta_\nu + \partial_\nu \delta \beta_\lambda) - \tilde{\Phi}^{0, \lambda\nu} \left( \frac{1}{2} \partial_\lambda \delta \beta_\nu - \partial_\nu \delta \beta_\lambda + \delta \omega_{\lambda\nu} \right)] \right\}
\]

(30)

As has been mentioned, it is possible to get rid of the surface and the last term in the right hand side of above equation through a suitable choice of the perturbations, but not of the second term.

Since \( \delta \tilde{T}_{\text{LE}} \) is a small term compared to \( \tilde{T}_{\text{LE}} \) we can determine the variation of the entropy \( \text{[28]} \) with an expansion in \( \delta \tilde{T}_{\text{LE}} \) at first order. First, we observe that (see also Eq. (24)):

\[
Z'_{\text{LE}} \equiv \text{tr} \left( \exp[-\tilde{T}_{\text{LE}} - \delta \tilde{T}_{\text{LE}}] \right) \simeq \text{tr} \left( \exp[-\tilde{T}_{\text{LE}}] (I - \delta \tilde{T}_{\text{LE}}) \right) = Z_{\text{LE}}(1 - \langle \delta \tilde{T}_{\text{LE}} \rangle / \tilde{T})
\]
where $\langle \cdot \rangle_T$ stands for the averaging with the original $\hat{\mathcal{T}}_{LE}$ local equilibrium operator. Hence, the new entropy reads:

$$S' = \frac{1}{Z_{LE}} \text{tr} \left( \exp[-\hat{\mathcal{T}}_{LE} - \delta \hat{\mathcal{T}}_{LE}](\hat{\mathcal{T}}_{LE} + \delta \hat{\mathcal{T}}_{LE}) \right) + \log Z_{LE}'$$

$$\simeq \frac{1}{Z_{LE}} \left( 1 + \langle \delta \hat{\mathcal{T}}_{LE} \rangle_T \right) \text{tr} \left( \exp[-\hat{\mathcal{T}}_{LE} - \delta \hat{\mathcal{T}}_{LE}](\hat{\mathcal{T}}_{LE} + \delta \hat{\mathcal{T}}_{LE}) \right) + \log Z_{LE} + \log(1 - \langle \delta \hat{\mathcal{T}}_{LE} \rangle_T)$$

(31)

We can now further expand the exponentials as we have done in Eq. (31). First:

$$\text{tr} \left( \exp[-\hat{\mathcal{T}}_{LE} - \delta \hat{\mathcal{T}}_{LE}](\hat{\mathcal{T}}_{LE} + \delta \hat{\mathcal{T}}_{LE}) \right) \simeq \text{tr} \left( \exp[-\hat{\mathcal{T}}_{LE} - \delta \hat{\mathcal{T}}_{LE}](I - \delta \hat{\mathcal{T}}_{LE} + \frac{1}{2}[\hat{\mathcal{T}}_{LE}, \delta \hat{\mathcal{T}}_{LE}] - \frac{1}{6}[\hat{\mathcal{T}}_{LE}, [\hat{\mathcal{T}}_{LE}, \delta \hat{\mathcal{T}}_{LE}]] + \ldots) \hat{\mathcal{T}}_{LE} \right)$$

$$= \text{tr}(\exp[-\hat{\mathcal{T}}_{LE}]) - \text{tr}(\exp[-\hat{\mathcal{T}}_{LE}]) \delta \hat{\mathcal{T}}_{LE} = Z_{LE}\langle \hat{\mathcal{T}}_{LE} \rangle_T - Z_{LE}\langle \delta \hat{\mathcal{T}}_{LE} \rangle_T$$

(32)

where, in the second equality, we have taken advantage of commutativity and cyclicity of the trace. Then:

$$\text{tr} \left( \exp[-\hat{\mathcal{T}}_{LE} - \delta \hat{\mathcal{T}}_{LE}] \delta \hat{\mathcal{T}}_{LE} \right) \simeq \text{tr} \left( \exp[-\hat{\mathcal{T}}_{LE} - \delta \hat{\mathcal{T}}_{LE}] (I - \delta \hat{\mathcal{T}}_{LE} + \frac{1}{2}[\hat{\mathcal{T}}_{LE}, \delta \hat{\mathcal{T}}_{LE}] - \frac{1}{6}[\hat{\mathcal{T}}_{LE}, [\hat{\mathcal{T}}_{LE}, \delta \hat{\mathcal{T}}_{LE}]] + \ldots) \delta \hat{\mathcal{T}}_{LE} \right)$$

$$\simeq \text{tr}(\exp[-\hat{\mathcal{T}}_{LE}]) \delta \hat{\mathcal{T}}_{LE} = Z_{LE}\langle \delta \hat{\mathcal{T}}_{LE} \rangle_T$$

(33)

keeping only first order terms. Thus, Eq. (31) can be rewritten as:

$$S' \simeq \frac{1}{Z_{LE}} \left( 1 + \langle \delta \hat{\mathcal{T}}_{LE} \rangle_T \right) \text{tr} \left( \exp[-\hat{\mathcal{T}}_{LE} - \delta \hat{\mathcal{T}}_{LE}](\hat{\mathcal{T}}_{LE} + \delta \hat{\mathcal{T}}_{LE}) \right) + \log Z_{LE} + \log(1 - \langle \delta \hat{\mathcal{T}}_{LE} \rangle_T)$$

$$\simeq \frac{1}{Z_{LE}} \left( 1 + \langle \delta \hat{\mathcal{T}}_{LE} \rangle_T \right) \left( Z_{LE}\langle \hat{\mathcal{T}}_{LE} \rangle_T - Z_{LE}\langle \delta \hat{\mathcal{T}}_{LE} \hat{\mathcal{T}}_{LE} \rangle_T + Z_{LE}\langle \delta \hat{\mathcal{T}}_{LE} \rangle_T \right) + \log Z_{LE} + \log(1 - \langle \delta \hat{\mathcal{T}}_{LE} \rangle_T)$$

$$= (1 + \langle \delta \hat{\mathcal{T}}_{LE} \rangle_T) \left( \langle \hat{\mathcal{T}}_{LE} \rangle_T - \langle \delta \hat{\mathcal{T}}_{LE} \hat{\mathcal{T}}_{LE} \rangle_T + \langle \delta \hat{\mathcal{T}}_{LE} \rangle_T \right) + \log Z_{LE} + \log(1 - \langle \delta \hat{\mathcal{T}}_{LE} \rangle_T)$$

(34)

Retaining only the first order terms in $\delta \hat{\mathcal{T}}_{LE}$, expanding the logarithm for $\langle \delta \hat{\mathcal{T}}_{LE} \rangle_T \ll 1$ and inserting the original expression of entropy:

$$S' \simeq S - \langle \delta \hat{\mathcal{T}}_{LE} \hat{\mathcal{T}}_{LE} \rangle_T$$

(35)

Therefore, the variation of the total entropy is, to the lowest order, proportional to the correlation between $\hat{\mathcal{T}}$ and $\delta \hat{\mathcal{T}}$, which is generally non-vanishing.

We can expand the above correlation to gain further insight. For the $\delta \hat{\mathcal{T}}_{LE}$, let us keep only the second term of the right hand side of Eq. (30):

$$\delta \hat{\mathcal{T}}_{LE} = -\frac{1}{2} \int d^3 x \tilde{\Phi}^a \partial \beta^a + \partial \beta^a \lambda$$

(36)

By using the (29) and the (30), the Eq. (35) can be rewritten as:

$$S'(t) \approx S(t) - \frac{1}{2} \int d^3 x \int d^3 x' \left( \langle \tilde{\Phi}^a \partial \beta^a \rangle_T \rangle_T - \langle \tilde{\Phi}^a \partial \beta^a \rangle_T \langle \partial \beta^a \rangle_T \rangle_T \right) \beta^a \partial \beta^a + \partial \beta^a \lambda$$

$$= \frac{1}{2} \int d^3 x \int d^3 x' \left( \langle \tilde{\Phi}^a \partial \beta^a \rangle_T \langle \partial \beta^a \rangle_T \rangle_T - \langle \tilde{\Phi}^a \partial \beta^a \rangle_T \langle \partial \beta^a \rangle_T \rangle_T \right) \xi \partial \beta^a + \partial \beta^a \lambda$$

$$- \frac{1}{4} \int d^3 x \int d^3 x' \left( \langle \tilde{\Phi}^a \partial \beta^a \rangle_T \langle \partial \beta^a \rangle_T \rangle_T - \langle \tilde{\Phi}^a \partial \beta^a \rangle_T \langle \partial \beta^a \rangle_T \rangle_T \right) \omega \partial \beta^a + \partial \beta^a \lambda$$

(37)

where $x$ and $x'$ have equal times. The above expression could be further simplified by e.g. approximating the local equilibrium mean $\langle \cdot \rangle_T$ with the global equilibrium one $\langle \cdot \rangle_0$, but this does not lead to further conceptual insight. The physical meaning of Eq. (37) is that the entropy difference depends on the correlation between local operators in two different space points multiplied by a factor which is at most of the second order in the perturbation $\delta \beta$. This kind of expression resembles the product of transport coefficients expressed by a Kubo formula times the squared gradient of the perturbation field. Therefore, the difference between entropies suggest that the introduction of a superpotential may lead to a modification of the transport coefficients. We will show this in detail in the next Section.
V. TRANSPORT COEFFICIENTS: SHEAR VISCOITY AS AN EXAMPLE

As has been mentioned, a remarkable consequence of the transformation \( \delta \beta \) is a difference in the predicted values of transport coefficients calculated with the relativistic Kubo formula, which is obtained by working out the mean value of the stress-energy tensor itself with the linear response theory and the nonequilibrium density operator in Eq. \( \text{(2)} \). For this purpose, the derivation in ref. \( \text{(1)} \) must be extended to the most general expression of the nonequilibrium density operator including a spin tensor, that is, Eq. \( \text{(13)} \); it can be found in Appendix A.

The equation \( \text{(25)} \), yielding the difference of mean values of a general observable under a transformation \( \delta \beta \), cannot be straightforwardly used to calculate the mean value of the stress-energy tensor setting \( \delta \beta = \frac{\pi}{L} \) because \( \hat{T}_{\mu \nu}(y) \) gets transformed itself. It is therefore more convenient to work out the general expression of the Kubo formula and study how it is modified by \( \delta \beta \) thereafter.

We will take shear viscosity as an example, the transformation of other transport coefficients can be obtained with the same reasoning. Shear viscosity, in the Kubo formula, is related to the spatial components of the symmetric part of the stress-energy tensor. It is worth pointing out that, since a non-vanishing spin tensor can make the stress-energy tensor non-symmetric, there might be a new transport coefficient related to the antisymmetric part of the stress-energy tensor.

For the symmetric part of the stress-energy tensor \( T_{\mu \nu}^{\text{sym}} \equiv (1/2)(T_{\mu \nu}^{\text{sym}} + T_{\nu \mu}^{\text{sym}}) \), using the general formula of relativistic linear response theory (Eq. \( \text{(70)} \) of Appendix A), the difference \( \delta T_{S}^{\mu \nu}(y) \) between actual mean value and local equilibrium value reads, at the lowest order in gradients:

\[
\delta T_{S}^{\mu \nu}(y) = \lim_{\epsilon \to 0} \frac{T}{i} \int_{-\infty}^{t'} dt \frac{1 - e^{\epsilon(t-t')}}{\epsilon} \int d^3x \left\langle \left[ \hat{T}_{\mu \nu}^{\text{sym}}(x), \hat{T}_{S}^{\mu \nu}(y) \right] \right\rangle_{0} \partial_{\rho} \delta \beta_{\rho}(x)
\]

\[
-\frac{1}{2} \lim_{\epsilon \to 0} \frac{T}{i} \int_{t}^{t'} dt e^{\epsilon(t-t')} \int d^3x \left\langle \left[ \hat{S}^{\text{sym}}_{\rho \sigma}(x), \hat{T}_{S}^{\mu \nu}(y) \right] \right\rangle_{0} \partial_{\rho} \omega_{\sigma \rho}(x)
\]

In order to obtain transport coefficients, a suitable perturbation must be chosen which can be eventually taken out from the integral. Physically, this corresponds to enforcing a particular hydrodynamical motion and observing the response of the stress-energy tensor to infer the dissipative coefficient. The perturbation \( \delta \beta = \frac{\pi}{L} \delta u \) is taken to be a stationary one and non-vanishing only within a finite region \( V, \) at whose boundary it goes to zero in a continuous and derivable fashion. The perturbation \( \delta \omega \) is also taken to be stationary and it can be chosen either to vanish or like in Eq. \( \text{(21)} \); in both cases, one gets to the same final result.

Let us then set \( \delta \omega = 0 \) and expand the perturbation \( \delta \beta = (0,0,\delta \beta^{2}(x^{1}),0) \) dependent on \( x^{1} \) in a Fourier series (it vanishes at some large, yet finite boundary). Since we want the higher order gradients of the perturbation to be negligibly small (the so-called hydrodynamic limit), the Fourier components with short wavelengths must be correspondingly suppressed. The component with the longest wavelength will then be much larger than any other and, therefore, \( \delta \beta^{2} \) can be approximately written, at least far from the boundary, as \( A \sin(\pi x^{1}/L) \) where \( L \) is the size of the region \( V \) in the \( x^{1} \) direction and \( A \) is a constant. The derivative of this perturbation reads:

\[
\partial_{1} \delta \beta_{2}(x) = \frac{\pi}{L} A \cos(\pi x^{1}/L) = \partial_{1} \delta \beta_{2}(0) \cos(\pi x^{1}/L) = \partial_{1} \delta \beta_{2}(0) \cos(k x^{1})
\]

where \( k \equiv \pi/L \). Therefore, by defining \( k = (k,0,0) \) and plugging the last equation in Eq. \( \text{(38)} \):

\[
\delta T_{S}^{\mu \nu}(y) = \lim_{\epsilon \to 0} \frac{T}{i} \partial_{1} \delta \beta_{2}(0) \int_{-\infty}^{t} dt \frac{1 - e^{\epsilon(t-t')}}{\epsilon} \int_{V} d^{3}x \cos k \cdot x \left\langle \left[ \hat{T}_{12}^{\text{sym}}(x), \hat{T}_{S}^{\mu \nu}(y) \right] \right\rangle_{0}
\]

\[
= \lim_{\epsilon \to 0} \frac{T}{i} \partial_{1} \delta \beta_{2}(0) \text{ Im} \int_{-\infty}^{t} \frac{1 - e^{\epsilon(t-t')}}{\epsilon} \int_{V} d^{3}x e^{ik \cdot x} \left\langle \left[ \hat{T}_{12}^{\text{sym}}(x), \hat{T}_{S}^{\mu \nu}(y) \right] \right\rangle_{0}
\]

(39)

taking into account that the commutator is purely imaginary. To extract shear viscosity we have to evaluate the stress-energy tensor in \( y = 0 \) to make it proportional to the derivative of the four-temperature field in the same point and we have to take the limit \( L \to \infty \) which implies \( V \to \infty \) and \( k \to 0 \) at the same time:

\[
\delta T_{S}^{\mu \nu}(t_{y},0) = \lim_{\epsilon \to 0} \frac{T}{i} \partial_{1} \delta \beta_{2}(0) \text{ Im} \int_{-\infty}^{t} \frac{1 - e^{\epsilon(t-t')}}{\epsilon} \int d^{3}x e^{ik \cdot x} \left\langle \left[ \hat{T}_{12}^{\text{sym}}(x), \hat{T}_{S}^{\mu \nu}(t_{y},0) \right] \right\rangle_{0}
\]

(40)
where it has been assumed that the integration domain goes to its thermodynamic limit independently of the integrand. Because of the time-translation symmetry of the equilibrium density operator $\hat{\rho}_0$, the mean value in the integral only depends on the time difference $t-t_y$. Thus, choosing the arbitrary time $t'=t_y$ and redefining the integration variables, the Eq. (40) can be rewritten as:

$$
\delta T^{\mu\nu}_{S}(t_y,0) = \lim_{\epsilon \to 0} \lim_{k \to 0} T \partial_1 \delta \beta_2(0) \int_{-\infty}^{0} dt \frac{1-e^{\epsilon t}}{\epsilon} \int d^3 x e^{i k x} \left[ \hat{T}^{12}(x), \hat{T}^{\mu\nu}_{S}(0) \right]_0 \tag{41}
$$

which shows that the mean value $\delta T^{\mu\nu}_{S}(t_y,0)$ is indeed independent of $t_y$, which is expected as $\delta \beta$ is stationary.

We can now take advantage of the well known Curie symmetry “principle” which states that tensors belonging to some irreducible representation of the rotation group will only respond to perturbations belonging to the same representation and with the same components \(^3\). In our case the Curie principle implies that only the same component of the symmetric part of the stress-energy tensor, i.e. $\hat{T}^{12}_{S}$, will give a non-vanishing value:

$$
\delta T^{12}_{S}(t_y,0) = \lim_{\epsilon \to 0} \lim_{k \to 0} T \partial_1 \delta \beta_2(0) \int_{-\infty}^{0} dt \frac{1-e^{\epsilon t}}{\epsilon} \int d^3 x e^{i k x} \left[ \hat{T}^{12}(x), \hat{T}^{12}_{S}(0) \right]_0 \tag{42}
$$

From the above expression, a Kubo formula for shear viscosity can be extracted setting $\delta \beta = (1/T)\delta u$:

$$
\eta = \lim_{\epsilon \to 0} \lim_{k \to 0} \Im \int_{-\infty}^{0} dt \frac{1-e^{\epsilon t}}{\epsilon} \int d^3 x e^{i k x} \left[ \hat{T}^{12}(x), \hat{T}^{12}_{S}(0) \right]_0 \tag{43}
$$

which, after a little algebra, can be shown to be the same expression obtained in ref. [11]. Because of the rotational invariance of the equilibrium density operator, shear viscosity is independent of the particular couple (1, 2) of chosen indices. It is worth pointing out that, had we started from Eq. (71) instead of Eq. (70), choosing $\delta \omega = 0$ or like in Eq. (21), we would have come to the same formula for shear viscosity; in the latter case, the third contributing term in Eq. (71) would have been of higher order in derivatives of $\delta \beta$, hence negligible.

Now, the question we want to answer is whether equation (43) is invariant by a pseudo-gauge transformation \([11]\), which turns the symmetric part of the stress-energy tensor into:

$$
\hat{T}^{\mu\nu}_{S} = \hat{T}^{\mu\nu}_{S} - \frac{1}{2} \partial_\lambda (\hat{\Phi}^{\mu,\lambda \nu} + \hat{\Phi}^{\nu,\lambda \mu}) = \hat{T}^{\mu\nu}_{S} - \partial_\lambda \hat{\Xi}^{\lambda \mu \nu} \tag{44}
$$

where:

$$
\frac{1}{2} (\hat{\Phi}^{\mu,\lambda \nu} + \hat{\Phi}^{\nu,\lambda \mu}) \equiv \hat{\Xi}^{\lambda \mu \nu} \tag{45}
$$

$\hat{\Xi}$ being symmetric in the last two indices. We will study the effect of the transformation on the mean value of the stress-energy tensor in the point $y=0$ starting from the formula Eq. (71) instead of Eq. (70) with $\delta \omega = 0$ or like in Eq. (21), which allows us to retain only the first contributing term to $\delta T^{12}_{S}(0)$. The perturbation $\delta \beta$ is taken to be stationary and $t'$ is set to be equal to $t_y = 0$. Eventually, the appropriate limits will be calculated to get the new shear viscosity. Thus:

$$
\delta T^{12}_{S}(0) = \delta T^{12}_{S}(0) + \lim_{\epsilon \to 0} \int_{-\infty}^{0} dt \frac{1-e^{\epsilon t}}{\epsilon} \int d^3 x \left( \partial_\alpha \hat{\Xi}^{\alpha 12}(x) \right)_0 \partial \beta \hat{\delta \beta}(x) + \partial \beta \hat{\delta \beta}(x)) \tag{46}
$$

We can simplify the above formula by noting that the mean value of two operators at equilibrium can only depend on the difference of the coordinates, so:

$$
\left( \partial \hat{O}_1(y), \partial \hat{O}_2(x) \right)_0 = \frac{\partial}{\partial y^{\mu}} \left( \partial \hat{O}_1, \partial \hat{O}_2 \right)_0 (y-x) = -\frac{\partial}{\partial y^{\mu}} \left( \partial \hat{O}_1, \partial \hat{O}_2 \right)_0 (y-x).
$$

\(^3\) This is true provided that the right hand side of Eq. (41) is a continuous function of \(k\) for \(k = 0\) or that its limit for \(k \to 0\) exists, i.e. it is independent of the direction of \(k\)
hence, the Eq. (46) can be rewritten as:

\[
\delta T_S^{12}(0) = \delta T_S^{12}(0) - \lim_{\varepsilon \to 0} \int_{-\infty}^{0} dt \frac{1 - e^{\varepsilon t}}{\varepsilon} \int d^3x \frac{\partial^2}{\partial x^\alpha \partial x^\beta} \langle \hat{\Xi}^{\alpha 12} (x), \hat{\Xi}^{\beta 12} (0) \rangle_0 (\partial_1 \delta T_2 (x) + \partial_2 \delta T_1 (x))
\]

\[
- \lim_{\varepsilon \to 0} \int_{-\infty}^{0} dt \frac{1 - e^{\varepsilon t}}{\varepsilon} \int d^3x \frac{\partial^2}{\partial x^\alpha \partial x^\beta} \langle \hat{\Xi}^{\alpha 12} (x), \hat{T}_S^{12} (0) \rangle_0 - \langle \hat{T}_S^{12} (x), \hat{\Xi}^{\alpha 12} (0) \rangle_0 (\partial_1 \delta T_2 (x) + \partial_2 \delta T_1 (x))
\]

We are now going to inspect the two terms on the right-hand side of the above equation. If the hamiltonian is time-reversal invariant, it can be shown (see Appendix B):

\[
\langle \hat{T}_S^{ij} (t, x), \hat{\Xi}^{\alpha ij} (0, 0) \rangle_0 = (-1)^{n_0} \langle \hat{T}_S^{ij} (-t, x), \hat{\Xi}^{\alpha ij} (0, 0) \rangle_0 = (-1)^{n_0} \langle \hat{\Xi}^{\alpha ij} (t, -x), \hat{T}_S^{ij} (0, 0) \rangle_0
\]

where \(n_0\) is the total number of time indices among those in the above expression. Similarly, if the hamiltonian is parity invariant, then:

\[
\langle \hat{\Xi}^{\alpha ij} (t, -x), \hat{T}_S^{ij} (0, 0) \rangle_0 = (-1)^{n_s} \langle \hat{\Xi}^{\alpha ij} (t, x), \hat{T}_S^{ij} (0, 0) \rangle_0
\]

where \(n_s\) is the total number of space indices. Using the last two equations to work out the last term of Eq. (47) one gets:

\[
\delta T_S^{12}(0) = \delta T_S^{12}(0) - \lim_{\varepsilon \to 0} \int_{-\infty}^{0} dt \frac{1 - e^{\varepsilon t}}{\varepsilon} \int d^3x \frac{\partial^2}{\partial x^\alpha \partial x^\beta} \langle \hat{\Xi}^{\alpha 12} (t, x), \hat{\Xi}^{\beta 12} (0, 0) \rangle_0 (\partial_1 \delta T_2 (x) + \partial_2 \delta T_1 (x))
\]

\[
-2 \lim_{\varepsilon \to 0} \int_{-\infty}^{0} dt \frac{1 - e^{\varepsilon t}}{\varepsilon} \int d^3x \frac{\partial^2}{\partial x^\alpha \partial x^\beta} \langle \hat{\Xi}^{\alpha 12} (t, x), \hat{T}_S^{12} (0, 0) \rangle_0 (\partial_1 \delta T_2 (x) + \partial_2 \delta T_1 (x))
\]

Now, the two terms on the right-hand side of (48) can be worked out separately. Using invariance by time-reversal and parity, one has:

\[
\langle \hat{\Xi}^{\alpha ij} (t, x), \hat{\Xi}^{\beta ij} (0, 0) \rangle_0 = (-1)^{n_0} \langle \hat{\Xi}^{\beta ij} (t, -x), \hat{\Xi}^{\alpha ij} (0, 0) \rangle_0 = (-1)^{n_0 + n_s} \langle \hat{\Xi}^{\beta ij} (t, x), \hat{\Xi}^{\alpha ij} (0, 0) \rangle_0 = \langle \hat{\Xi}^{\beta ij} (t, x), \hat{\Xi}^{\alpha ij} (0, 0) \rangle_0
\]

being \(n_0 + n_s = 6\). Hence, the first term on the right-hand side of (48) can be decomposed as:

\[
- \lim_{\varepsilon \to 0} \int_{-\infty}^{0} dt \frac{1 - e^{\varepsilon t}}{\varepsilon} \int d^3x \left( \frac{\partial^2}{\partial t^2} \langle \hat{\Xi}^{ij} (x), \hat{\Xi}^{0 ij} (0) \rangle_0 + 2 \frac{\partial}{\partial t} \frac{\partial}{\partial x^k} \langle \hat{\Xi}^{k ij} (x), \hat{\Xi}^{0 ij} (0) \rangle_0 \right)
\]

\[
+ \frac{\partial}{\partial x^k} \frac{\partial}{\partial t} \langle \hat{\Xi}^{k ij} (x), \hat{\Xi}^{ij} (0) \rangle_0 \left( \partial_1 \delta T_2 (x) + \partial_2 \delta T_1 (x) \right)
\]

and, similarly, the second term as:

\[
-2 \lim_{\varepsilon \to 0} \int_{-\infty}^{0} dt \frac{1 - e^{\varepsilon t}}{\varepsilon} \int d^3x \left( \frac{\partial}{\partial t} \langle \hat{\Xi}^{ij} (x), \hat{T}_S^{12} (0) \rangle_0 + \frac{\partial}{\partial x^k} \langle \hat{\Xi}^{k ij} (x), \hat{T}_S^{12} (0) \rangle_0 \right) \left( \partial_1 \delta T_2 (x) + \partial_2 \delta T_1 (x) \right)
\]

All terms in Eqs. (50) and (51) with a space derivative do not yield any contribution to first-order transport coefficients. This can be shown by, firstly, integrating by parts and generating two terms, one of which is a total derivative and the second involves the second derivative of the perturbation \(\delta \beta\). The total derivative term can be transformed into a surface integral on the boundary of \(V\) which vanishes because therein the perturbation \(\delta \beta\) is supposed to vanish along with its first-order derivatives. The second term, involving higher order derivatives, does not give contribution to transport coefficients at first order in the derivative expansion. Altogether, the Eq. (48) turns into:
which can be further integrated by parts in the time \( t \), yielding:

\[
\delta T_{S}^{12}(0) = \delta T_{S}^{12}(0) - \lim_{\varepsilon \to 0} \int_{-\infty}^{0} dt \left( \delta(t) - \varepsilon e^{t} \right) \int d^{3}x \left\{ \left[ \hat{\Xi}^{012}(x), \hat{\Xi}^{012}(0) \right] \right\}_{0} (\partial_{1} \delta \tilde{\beta}_{2}(x) + \partial_{2} \delta \tilde{\beta}_{1}(x))
\]

\[-2 \lim_{\varepsilon \to 0} \int_{-\infty}^{0} dt \varepsilon e^{t} \int d^{3}x \left\{ \left[ \hat{\Xi}^{012}(x), \hat{T}_{S}^{12}(0) \right] \right\}_{0} (\partial_{1} \delta \tilde{\beta}_{2}(x) + \partial_{2} \delta \tilde{\beta}_{1}(x)) + \mathcal{O}(\varepsilon^{2} \delta \beta) \tag{53}\]

provided that, for general space-time dependent operators \( \hat{O}_{1} \) and \( \hat{O}_{2} \)

\[
\lim_{t \to -\infty} \int_{V} d^{3}x \varepsilon^{3t} \frac{\partial}{\partial t} \langle \hat{O}_{1}(t, x), \hat{O}_{2}(0, 0) \rangle = 0
\]

\[
\lim_{t \to -\infty} \int_{V} d^{3}x \varepsilon^{3t} \langle \hat{O}_{1}(t, x), \hat{O}_{2}(0, 0) \rangle = 0
\]

with \( n = 0, 1 \), which is reasonable because thermodynamical correlations are expected to vanish exponentially as a function of time for fixed points in space \(^4\).

From Eq. (53), the variation of the shear viscosity can be inferred with the very same reasoning that led us to formula (13), that is:

\[
\Delta \eta = \eta' - \eta = - \lim_{\varepsilon \to 0} \lim_{k \to 0} \text{Im} \int_{-\infty}^{0} dt \langle \delta(t) - \varepsilon e^{t} \rangle \int d^{3}x \varepsilon^{ikx} \langle \left[ \hat{\Xi}^{012}(t, x), \hat{\Xi}^{012}(0, 0) \right] \rangle_{0}
\]

\[-2 \lim_{\varepsilon \to 0} \lim_{k \to 0} \text{Im} \int_{-\infty}^{0} dt \varepsilon e^{t} \int d^{3}x \varepsilon^{ikx} \langle \left[ \hat{\Xi}^{012}(t, x), \hat{T}_{S}^{12}(0, 0) \right] \rangle_{0} \tag{54}\]

If the first integral is regular, then the \( \varepsilon \to 0 \) limit kills one term and the (54) reduces to:

\[
\Delta \eta = \eta' - \eta = - \lim_{k \to 0} \int_{V} d^{3}x \cos kx \langle \left[ \hat{\Xi}^{012}(0, x), \hat{\Xi}^{012}(0, 0) \right] \rangle_{0}
\]

\[-2 \lim_{\varepsilon \to 0} \lim_{k \to 0} \text{Im} \int_{-\infty}^{0} dt \varepsilon e^{t} \int d^{3}x \varepsilon^{ikx} \langle \left[ \hat{\Xi}^{012}(x), \hat{T}_{S}^{12}(0, 0) \right] \rangle_{0} \tag{55}\]

In general, this difference is non-vanishing, leading to the conclusion that the specific form of the stress-energy tensor and, possibly, the existence of a spin tensor in the underlying quantum field theory affects the value of transport coefficients. The relative difference of those values depends on the particular transformation (11), hence on the particular stress-energy tensor. In the next Section a specific instance will be presented and discussed.

An important point to make is that the found dependence of the transport coefficients on the particular set of stress-energy and spin tensor of the theory is indeed physically meaningful. This means that the variation of some coefficient is not compensated by a corresponding variation of another coefficient so as to eventually leave measurable quantities unchanged. This has been implicitly proved in Sect. (17) where it was shown that total entropy itself undergoes a variation under a transformation of the stress-energy and spin tensor (see Eq. (55)).

## VI. DISCUSSION AND CONCLUSIONS

As a first point, we would like to emphasize that in our arguments space-time curvature and gravitational coupling have been disregarded. On one hand, this shows that the nature of stress-energy tensor and, possibly, the existence of a fundamental spin tensor could, at least in principle, be demonstrated independently of gravity. On the other hand, for each stress-energy tensor created with the transformation (11), it should be shown that an extension of general relativity exists having it as a source, which could not be always possible.

An important question is whether a concrete physical system indeed exists for which the transformation (11) leads to actually different values for e.g. transport coefficients, entropy production rate or other quantities in nonequilibrium

---

\(^4\) There might be singularities on the light cone, however for fixed \( x \) and \( 0 \) and integration over a finite region \( V, \) in the limit \( t \to -\infty \) light cone is not involved
situations. For this purpose, we discuss a specific instance regarding spinor electrodynamics. Starting from the symmetrized gauge-invariant Belinfante tensor of the coupled Dirac and electromagnetic fields, with associated \( \tilde{S} = 0 \):

\[
\tilde{T}^{\mu\nu} = \frac{i}{4} \left( \overline{\Psi} \gamma^{\mu+\nu} \nabla \Psi + \overline{\Psi} \gamma^{\nu+\mu} \nabla \Psi \right) + \tilde{F}^{\mu}_\lambda \tilde{F}^{\lambda\nu} + \frac{1}{4} \tilde{F}^{\mu\nu} \tilde{F}^{\mu\nu} \tag{56}
\]

where \( \nabla_{\mu} = \partial_{\mu} - ieA_{\mu} \) is the gauge covariant derivative, one can generate other stress-energy tensors with suitable rank three tensors and then setting \( \tilde{\Phi} = -\tilde{S} \) where \( \tilde{S} \) is the new spin tensor, according to \( \{1\} \). One of the best known is the canonical Dirac spin tensor:

\[
\tilde{\Phi}^{\lambda,\mu\nu} = -\frac{i}{8} \nabla \{\gamma^{\lambda}, [\gamma^{\mu}, \gamma^{\nu}]\} \Psi
\]

\( \{\ } \) stands for anticommutator which is gauge-invariant and transforms the Belinfante tensor \( \tilde{T} \) back to the canonical one obtained from the spinor electrodynamics lagrangian (see also \( \{2\} \) for a detailed discussion). However, this is totally antisymmetric in the three indices \( \lambda, \mu, \nu \) and thus the variation of \( \tilde{T} \) operator (see Eq. \( \{22\} \)) as well as transport coefficients, which depend on the symmetrized \( \tilde{S} \) tensor \( \{13\} \) vanish. Nevertheless, other gauge-invariant \( \tilde{\Phi} \)-like tensors can be found. For instance, one could employ a superpotential:

\[
\tilde{\Phi}^{\lambda,\mu\nu} = \frac{1}{8m} \nabla \left( \gamma^{\mu\nu} - \gamma^{\nu\mu} \right) \gamma^{\lambda} \Psi + h.c. = \frac{1}{8m} \nabla \left( [\gamma^{\mu}, \gamma^{\lambda}]^{\nu\mu} - [\gamma^{\nu}, \gamma^{\lambda}]^{\mu\nu} \right) \Psi
\]

which is the gauge-invariant version of the one used in ref. \( \{12\} \) to obtain a conserved spin current. This superpotential gives rise to a non-vanishing spin tensor as well as a \( \tilde{\Sigma} \) tensor (see Eq. \( \{14\} \)):

\[
\tilde{\Sigma}^{\lambda\mu\nu} = \frac{1}{16m} \nabla \left( [\gamma^{\lambda}, \gamma^{\mu}]^{\nu\mu} + [\gamma^{\lambda}, \gamma^{\nu}]^{\mu\nu} \right) \Psi
\]

hence a variation of thermodynamics. By noting that the structure of the above tensor is very similar to the Belinfante stress-energy tensor \( \{15\} \), it is not difficult to find a rough estimate of the variation of e.g. shear viscosity induced by the transformation. Looking at Eq. \( \{55\} \) we note that \( \tilde{\Sigma}^{012} \) mainly differs from \( \tilde{T}^{012} \) in Eq. \( \{56\} \) by the factor \( 1/m \). The last term on the right hand side of Eq. \( \{56\} \) tells us that the dimension of \( \tilde{\Sigma} \) is that of a stress-energy tensor multiplied by a time, and therefore this term must be of the order of \( \eta \hbar/mc^2\tau \) where \( \tau \) is the microscopic correlation time scale of the original stress-energy tensor or the collisional time scale in the kinetic language and \( \eta \) the shear viscosity obtained from the original stress-energy tensor. Thus, the expected relative variation of shear viscosity from Eq. \( \{55\} \) in this case is of the order:

\[
\frac{\Delta \eta}{\eta} \approx O \left( \frac{\hbar}{mc^2\tau} \right)
\]

which is (as it could have been expected) a quantum relativistic correction governed by the ratio \( \lambda_e/c/\tau \), \( \lambda_e \) being the Compton wavelength. For the electron, the ratio \( \lambda_e/c \approx 10^{-21} \) sec, which is a very small time scale compared to the usual kinetic time scales, yet it could be detectable for particular systems with very low shear viscosity.

It is also interesting to note that the “improved” stress-energy tensor by Callan, Coleman and Jackiw \( \{14\} \) with renormalizable matrix elements at all orders of perturbation theory, is obtained from the Belinfante’s symmetrized one in Eq. \( \{50\} \) with a transformation of the kind \( \{1\} \) setting (for the Dirac field and vanishing constants \( \{14\} \)):

\[
\tilde{Z}^{\alpha\lambda\mu\nu} = -\frac{1}{6} (g^{\alpha\mu} g^{\lambda\nu} - g^{\alpha\nu} g^{\lambda\mu}) \nabla \Psi
\]

and requiring \( \tilde{S} = \tilde{\Sigma} = 0 \) so that \( \tilde{\Phi}^{\lambda,\mu\nu} = \partial_{\alpha} \tilde{Z}^{\alpha\lambda,\mu\nu} \), hence:

\[
\tilde{\Phi}^{\lambda,\mu\nu} = -\frac{1}{6} (g^{\lambda\nu} \partial^{\mu} - g^{\lambda\mu} \partial^{\nu}) \nabla \Psi
\]

\[
\tilde{\Sigma}^{\lambda\mu\nu} = \frac{1}{2} (\tilde{\Phi}^{\lambda,\mu\nu} + \tilde{\Phi}^{\lambda,\nu\mu}) = -\frac{1}{6} \left[ g^{\mu\nu} \partial^{\lambda} - \frac{1}{2} (g^{\lambda\nu} \partial^{\mu} + g^{\lambda\mu} \partial^{\nu}) \right] \nabla \Psi
\]

\[
\tilde{T}^{\mu\nu} = \tilde{T}^{\mu\nu} - \partial_{\lambda} \tilde{\Sigma}^{\lambda\mu\nu} = \tilde{T}^{\mu\nu} + \frac{1}{6} (g^{\mu\nu} \Box - \partial^{\mu} \partial^{\nu}) \nabla \Psi
\]

which is just the improved stress-energy tensor \( \{14\} \). It is likely (to be verified though) that the aforementioned modified stress-energy tensors imply a different thermodynamics with respect to the original Belinfante symmetrized tensor. This problem has been recently pointed out in ref. \( \{15\} \).
To summarize, we have concluded that different quantum stress-energy tensors imply different values of nonequilibrium thermodynamical quantities like transport coefficients and entropy production rate. This reinforces our previous similar conclusion concerning differences of momentum and angular momentum densities in rotational equilibrium [3]. The existence of a fundamental spin tensor has, thus, an impact on the microscopic number of degrees of freedom and on how quickly macroscopic information is converted into microscopic. The difference of transport coefficients depends on the particular form of the tensors and in the examined case it scales like a quantum relativistic effect with $\hbar/c$. Therefore, at least in principle, it is possible to disprove a supposed stress-energy tensor with a suitably designed thermodynamical experiment.

Acknowledgments

We are grateful to F. Bigazzi, F. W. Hehl and D. Seminara for useful discussions and suggestions.

References

[1] G. S. Denicol, T. Kodama, T. Koide and P. .Mota, J. Phys. G 35, 115102 (2008); P. Romatschke, Int. J. Mod. Phys. E 19, 1 (2010); P. Romatschke, Class. Quant. Grav. 27, 025006 (2010); K. Jensen, M. Kaminski, P. Kovtun, R. Meyer, A. Ritz and A. Yarom, Phys. Rev. Lett. 109, 101601 (2012); P. Van and T. S. Biro, Phys. Lett. B 709, 106 (2012); J. Bhattacharya, S. Bhattacharyya and M. Rangamani, arXiv:1211.1020.

[2] D. H. Rischke, Nucl. Phys. A 610, 85C (1996); L. P. Csernai et al., Heavy Ion Phys. 17, 271 (2003); H. Song and U. W. Heinz, Phys. Rev. C 77, 064901 (2008); P. Huovinen and D. Molnar, Phys. Rev. C 79, 014906 (2009); B. Schenke, S. Jeon and C. Gale, Phys. Rev. C 82, 014903 (2010); P. Bozek, W. Broniowski and I. Wyskiel-Piekarska, arXiv:1207.3176.

[3] F. Becattini and L. Tinti, Phys. Rev. D 84, 025013 (2011).

[4] F. Halbwachs, Theorie relativiste des fluids `a spin, Gauthier-Villars, Paris, (1960).

[5] F. W. Hehl, Rept. Math. Phys. 9, 55 (1976).

[6] D. N. Zubarev, Sov. Phys. Doklady 10, 850 (1966); D. N. Zubarev and M. V. Tokarchuk, Theor. Math. Phys. 88, 876 (1992) [Teor. Mat. Fiz. 88N2, 286 (1991)].

[7] H. A. Weldon, Phys. Rev. D 26, 1394 (1982).

[8] F. Becattini, Phys. Rev. Lett. 108, 244502 (2012).

[9] D. N. Zubarev Non Equilibrium Statistical Thermodynamics, Nauka, Moscow (1971). (english translation: New York, Consultant Bureau, 1974); D. N. Zubarev, V. G. Morozov, G. Röpke, Statistical Mechanics of Nonequilibrium Processes, Volume 2: Relaxation and Hydrodynamic Processes, Akademie Verlag, Berlin (1997).

[10] R. Kubo, M. Toda, N. Hashitsume, Statistical Physics II. Nonequilibrium Statistical Mechanics, Springer-Verlag, Berlin (1985).

[11] A. Hosoya, M. Sakagami and M. Takao, Annals Phys. 154, 229 (1984).

[12] S. R. De Groot, W. A. van Leeuwen, Ch. G. van Weert, Relativistic kinetic theory, North Holland (1980).

[13] R. Balian, From microphysics to macrophysics, Springer, Heidelberg (2007).

[14] C. G. Callan, Jr., S. R. Coleman and R. Jackiw, Annals Phys. 59, 42 (1970).

[15] Y. Nakayama, Int. J. Mod. Phys. A 27, 1250125 (2012).

APPENDIX A - Relativistic linear response theory with spin tensor

We extend the relativistic linear response theory in the Zubarev’s approach to the case of a non-vanishing spin tensor. The (stationary) nonequilibrium density operator is written in Eq. [13], with $\hat{\Upsilon}$ expanded as in Eq. [14]. As has been shown in Sect. [11] at equilibrium, only the first term of the $\hat{\Upsilon}$ operator survives in Eq. [14]; therefore, one can rewrite that equation using the perturbations $\delta\beta$, $\delta\xi$ and $\delta\omega$ which are defined as the difference between the
actual value and their value at thermodynamical equilibrium:

\[
\hat{\Upsilon} = \int d^3x \left( \hat{T}^{0\nu} \beta_\nu(t', x) - \bar{J}^0(\xi(t', x) - \frac{1}{2} \hat{S}^{0,\mu\nu} \delta \omega_{\mu\nu}(t', x) \right) \\
+ \lim_{\varepsilon \to 0} \int_{-\infty}^{t'} dt \ v^{(t-t')} \int dS \ n_i \left( \hat{T}^{\mu\nu} \delta \beta_\nu(x) - \bar{J}^0 \delta \xi(x) - \frac{1}{2} \hat{S}^{0,\mu\nu} \delta \omega_{\mu\nu}(x) \right) \\
- \frac{1}{2} \lim_{\varepsilon \to 0} \int_{-\infty}^{t'} dt \ d^3x \ e^{v^{(t-t')} \left( \hat{T}^{\mu\nu} (\partial_\mu \delta \beta_\nu(x) + \partial_\nu \delta \beta_\nu(x)) + \hat{T}^{\mu\nu} (\partial_\mu \delta \beta_\nu(x) - \partial_\nu \delta \beta_\nu(x) + 2 \delta \omega_{\mu\nu}(x) \right) \\
- \hat{S}^{0,\mu\nu} \delta \omega_{\mu\nu}(x) - 2 \hat{S}^{3,\mu\nu} \delta \omega_{\mu\nu}(x) \right)
\]

(57)

where it is understood that \( x = (t, x) \).

In fact, we will use a rearrangement of the right-hand-side expression which is more convenient if one wants to work with an unspecified, yet small, \( \delta \omega \). Therefore, the above equation is rewritten as:

\[
\hat{\Upsilon} = \int d^3x \left( \hat{T}^{0\nu} \beta_\nu(t', x) - \bar{J}^0(\xi(t', x) - \frac{1}{2} \hat{S}^{0,\mu\nu} \omega_{\mu\nu}(t', x) \right)
\]

(58)

what it can be easily obtained from Eq. (13) integrating by parts in time.

For the sake of simplicity we calculate the linear response with \( \xi_{eq} = \delta \xi = 0 \), but it can be shown that our final expressions hold for \( \xi_{eq} \neq 0 \) (in other words with a non-vanishing chemical potential \( \mu \neq 0 \)). Let us now define:

\[
\hat{A} = - \int d^3x \left( \hat{T}^{0\nu} \beta_\nu(t', x) - \frac{1}{2} \hat{S}^{0,\mu\nu} \omega_{\mu\nu}(t', x) \right)
\]

and:

\[
\hat{B} = \lim_{\varepsilon \to 0} \int_{-\infty}^{t'} dt \ e^{v^{(t-t')} \partial / \partial t} \int d^3x \left( \hat{T}^{0\nu} \delta \beta_\nu(x) - \frac{1}{2} \hat{S}^{0,\mu\nu} \delta \omega_{\mu\nu}(x) \right)
\]

so that:

\[
\bar{\rho} = \frac{1}{Z} \exp[\hat{\Upsilon}] = \frac{1}{Z} \exp[\hat{A} + \hat{B}]
\]

(59)

with \( Z = \text{tr}(\exp[\hat{A} + \hat{B}]) \).

The operator \( \hat{B} \) is the small term in which \( \bar{\rho} \) is to be expanded, according to the linear response theory. It can can be rewritten in a way which will be useful later on. Since:

\[
\int d^3x \ \partial / \partial t \left( \hat{T}^{0\nu}(x) \delta \beta_\nu(x) \right) = \int d^3x \ \partial_\mu \left( \hat{T}^{\mu\nu}(x) \delta \beta_\nu(x) \right) - \int d^3x \ \partial_\nu \hat{T}^{\mu\nu}(x) \delta \beta_\nu(x)
\]

\[
= \int d^3x \ \hat{T}^{\mu\nu}(x) \partial_\mu \delta \beta_\nu(x) - \int dS \ n_i \hat{T}^{\mu\nu}(x) \delta \beta_\nu(x)
\]

then:

\[
\hat{B} = \lim_{\varepsilon \to 0} \int_{-\infty}^{t'} dt \ e^{v^{(t-t')} \int d^3x \left( \hat{T}^{\mu\nu} \partial_\mu \delta \beta_\nu(x) - \frac{1}{2} \partial / \partial t \left( \hat{S}^{0,\mu\nu} \delta \omega_{\mu\nu}(x) \right) \right) - \int dS \ n_i \hat{T}^{\mu\nu}(x) \delta \beta_\nu(x)
\]

The perturbation \( \delta \beta \) must be chosen such that \( \delta \beta |_{\partial V} = 0 \) so that only the bulk term survives in the above equation:

\[
\hat{B} = \lim_{\varepsilon \to 0} \int_{-\infty}^{t'} dt \ e^{v^{(t-t')} \int d^3x \left( \hat{T}^{\mu\nu} \partial_\mu \delta \beta_\nu(x) - \frac{1}{2} \partial / \partial t \left( \hat{S}^{0,\mu\nu} \delta \omega_{\mu\nu}(x) \right) \right)}
\]

(60)

At the lowest order in \( \hat{B} \):

\[
Z = \text{tr}(e^{\hat{A} + \hat{B}}) \simeq \text{tr}(e^{\hat{A}} \left[ 1 + \hat{B} \right]) = Z_{LE}(1 + \langle \hat{B} \rangle_{LE}) \Rightarrow \frac{1}{Z} \simeq \frac{1}{Z_{LE}}(1 - \langle \hat{B} \rangle_{LE})
\]

(61)
where the subscript LE stands for Local Equilibrium and implies the calculation of mean values with the local equilibrium density operator (see Sect. IV). Thereby, putting together (61) and (62) and retaining only first-order terms in $\hat{B}$:

$$\hat{\rho} \simeq \left(1 - \langle \hat{B} \rangle_{\text{LE}} \right) \hat{\rho}_{\text{LE}} + \int_0^1 dz \, e^{z\hat{A}} \hat{B} e^{-z\hat{A}} \hat{\rho}_{\text{LE}},$$

hence the mean value of an operator $\hat{O}$ becomes:

$$\langle \hat{O}(y) \rangle \simeq \left(1 - \langle \hat{B} \rangle_{\text{LE}} \right) \langle \hat{O}(y) \rangle_{\text{LE}} + \langle \hat{O}(y) \rangle \int_0^1 dz \, e^{z\hat{A}} \hat{B} e^{-z\hat{A}} \langle \hat{O}(y) \rangle_{\text{LE}}.$$

(63)

Let us focus on the last term, which, by virtue of (60), contains expressions of this sort:

$$\langle \hat{O}(y) \hat{X}'(z, t, x) \rangle_{\text{LE}} = \langle \hat{O}(y) e^{z\hat{A}} \hat{X}(t, x) e^{-z\hat{A}} \rangle_{\text{LE}}$$

where $\hat{X}$ stands for components of either $\hat{T}$ or $\hat{S}$ or $\partial_b \hat{S}$. From the identity:

$$\langle \hat{O}(y) \hat{X}'(z, t, x) \rangle_{\text{LE}} = \int_{-\infty}^{t} d\tau \, \langle \hat{O}(y) \partial_\tau \hat{X}'(z, \tau, x) \rangle_{\text{LE}} + \lim_{\tau \to -\infty} \langle \hat{O}(y) \hat{X}'(z, \tau, x) \rangle_{\text{LE}},$$

and the observation that correlations vanish for very distant times (check footnote 4), one obtains:

$$\langle \hat{O}(y) \hat{X}'(z, t, x) \rangle_{\text{LE}} = \int_{-\infty}^{t} d\tau \, \langle \hat{O}(y) \partial_\tau \hat{X}'(z, \tau, x) \rangle_{\text{LE}} + \lim_{\tau \to -\infty} \langle \hat{O}(y) \hat{X}'(z, \tau, x) \rangle_{\text{LE}},$$

(64)

where we have also taken advantage of the commutation between $\exp[\hat{A}]$ and $\exp[\pm z\hat{A}]$.

We now approximate the local equilibrium density operator with the nearest equilibrium operator $\hat{\rho}_0$ in Eq. (26), which also implies that:

$$\hat{A} \simeq -\hat{H}/T$$

where $\hat{H}$ is the hamiltonian operator (which ought to exist given the chosen boundary conditions). The straightforward consequence of this approximation is that the second term on the right hand side in Eq. (54) can be written as:

$$\langle \hat{X}(\infty, x) \rangle_{\text{LE}} \simeq \langle \hat{X}(\infty, x) \rangle_0$$

because the mean value is stationary under the equilibrium distribution. Therefore, the Eq. (64) can be approximated as:

$$\langle \hat{O}(y) \hat{X}'(z, t, x) \rangle_{\text{LE}} \simeq \int_{-\infty}^{t} d\tau \, \langle \hat{O}(y) \partial_\tau \hat{X}'(z, \tau, x) \rangle_0 + \langle \hat{O}(y) \rangle_0 \langle \hat{X}(t, x) \rangle_0,$$

(65)

and the (63) as:

$$\langle \hat{O}(y) \rangle \simeq (1 - \langle \hat{B} \rangle_0) \langle \hat{O}(y) \rangle_0 + \int_0^1 dz \, \langle \hat{O}(y) \rangle_0 e^{-z\hat{H}/T} \hat{B} e^{z\hat{H}/T} \langle \hat{O}(y) \rangle_0,$$

(66)

Once integrated, the second term in (65) gives rise to a term which cancels out exactly the $\langle \hat{B} \rangle_0 \langle \hat{O}(y) \rangle_0$ in the equation above, which then becomes:

$$\langle \hat{O}(y) \rangle \simeq \langle \hat{O}(y) \rangle_0 + \int_0^1 dz \int_{-\infty}^{t} d\tau \, \langle \hat{O}(y) \rangle_0 \langle \hat{X}'(z, \tau, x) \rangle_0,$$

(67)
Let us now integrate the last term on the right hand side in $z$:

$$
\int_0^1 dz \int_{-\infty}^t d\tau \langle \bar{O}(y)\partial_{\tau} \bar{X}'(z, \tau, x)\rangle_0 = \frac{1}{\beta} \int_0^t du \int_{-\infty}^t d\tau \langle \bar{O}(y)\partial_{\tau} e^{-u\bar{H}} \bar{X}(\tau, x)e^{u\bar{H}}\rangle_0
$$

where $\beta = 1/T$ and $\bar{\beta} z = u$. As $\bar{H}$ is the generator of time translations:

$$
\frac{1}{\beta} \int_0^t du \int_{-\infty}^t d\tau \langle \bar{O}(y)\partial_{\tau} e^{-u\bar{H}} \bar{X}(\tau, x)e^{u\bar{H}}\rangle_0 = \frac{1}{\beta} \int_0^t du \int_{-\infty}^t d\tau \langle \bar{O}(y)\partial_{\tau} \bar{X}(\tau + iu, x)\rangle_0
$$

$$
= \frac{1}{i\beta} \int_0^t du \int_{-\infty}^t d\tau \langle \bar{O}(y)\partial_{\bar{u}} \bar{X}(\tau + iu, x)\rangle_0 = \frac{1}{i\beta} \int_0^t du \int_{-\infty}^t d\tau \frac{\partial}{\partial u} \langle \bar{O}(y)\bar{X}(\tau + iu, x)\rangle_0
$$

$$
= \frac{1}{i\beta} \int_{-\infty}^t d\tau \int_0^t du \frac{\partial}{\partial u} \langle \bar{O}(y)\bar{X}(\tau + iu, x)\rangle_0 = \frac{1}{i\beta} \int_{-\infty}^t d\tau \left( \langle \bar{O}(y)\bar{X}(\tau + i\bar{\beta}, x)\rangle_0 - \langle \bar{O}(y)\bar{X}(\tau, x)\rangle_0 \right)
$$

On the other hand:

$$
\langle \bar{O}(y)\bar{X}(\tau + i\bar{\beta}, x)\rangle_0 = \text{tr}(\bar{\rho}_0 \bar{O}(y)e^{-\bar{\beta}\bar{H}} \bar{X}(\tau, x)e^{\bar{\beta}\bar{H}}) = \frac{1}{Z_0} \text{tr}(e^{-\bar{\beta}\bar{H}} \bar{O}(y)e^{-\bar{\beta}\bar{H}} \bar{X}(\tau, x)e^{\bar{\beta}\bar{H}})
$$

$$
= \frac{1}{Z_0} \text{tr}(\bar{O}(y)e^{-\bar{\beta}\bar{H}} \bar{X}(\tau, x)) = \text{tr}(\bar{X}(\tau, x))\bar{\rho}_0 \bar{O}(y)) = \langle \bar{X}(\tau, x)\bar{O}(y)\rangle
$$

Hence, putting the last three equations together, we have:

$$
\int_0^1 dz \int_{-\infty}^t d\tau \langle \bar{O}(y)\partial_{\tau} \bar{X}'(z, \tau, x)\rangle_0 = \frac{1}{i\beta} \int_{-\infty}^t d\tau \langle [\bar{X}(\tau, x), \bar{O}(y)]\rangle_0 \tag{68}
$$

Substituting now $\bar{X}$ with its specific operators, Eq. (67) can be expanded as:

$$
\delta(\bar{O}(y)) = \langle \bar{O}(y)\rangle - \langle \bar{O}(y)\rangle_0 = \lim_{\varepsilon \to 0} \frac{1}{i\beta} \int_{-\infty}^t dt \int_{-\infty}^{t-\varepsilon} d\tau e^{(t-\tau)} \int_{\infty}^t d\tau \int d^3x \left( \bar{T}^{\mu\nu}(\tau, x), \bar{O}(y) \right)_{0\mu\nu}(x)
$$

$$
- \frac{1}{\varepsilon} \int_{-\infty}^t dt \langle e^{(t-\tau)} \frac{\delta}{\delta \tau} \left( \bar{T}^{\mu\nu}(\tau, x), \bar{O}(y) \right)_{0\mu\nu}(x) \rangle_0 \partial_{\mu\nu}(x)
$$

$$
= \lim_{\varepsilon \to 0} \frac{1}{i\beta} \int_{-\infty}^t dt \int_{-\infty}^{t-\varepsilon} d\tau e^{(t-\tau)} \int_{\infty}^t d\tau \int d^3x \left( \bar{S}^{\mu\nu}(\tau, x), \bar{O}(y) \right)_{0\mu\nu}(x)
$$

$$
- \frac{1}{2} \lim_{\varepsilon \to 0} \frac{1}{i\beta} \int_{-\infty}^t dt \int_{-\infty}^{t-\varepsilon} d\tau e^{(t-\tau)} \int_{\infty}^t d\tau \int d^3x \left( \bar{S}^{\mu\nu}(\tau, x), \bar{O}(y) \right)_{0\mu\nu}(x)
$$

$$
\partial_{\mu\nu}(x) \delta_{\mu\nu}(x) \langle \bar{O}(y)\rangle_0 \partial_{\mu\nu}(x)
$$

The first term on the right hand side of the above equation can be integrated by parts using:

$$
\int_{-\infty}^t dt \int_{-\infty}^{t-\varepsilon} d\tau f(\tau) = \int_{-\infty}^{t-\varepsilon} d\tau f(\tau) - \int_{-\infty}^{t-\varepsilon} d\tau \frac{\partial}{\partial \tau} \left( \frac{e^{(t-\tau)}}{\varepsilon} \right) f(\tau)
$$

$$
= \frac{1}{\varepsilon} \int_{-\infty}^{t-\varepsilon} d\tau f(\tau) - \int_{-\infty}^{t-\varepsilon} dt \frac{e^{(t-\tau)}}{\varepsilon} f(t) = \int_{-\infty}^{t} dt \frac{1}{\varepsilon} e^{(t-\tau)} f(t)
$$

so that the Eq. (69) can be finally written:

$$
\delta(\bar{O}(y)) = \lim_{\varepsilon \to 0} \frac{1}{i\beta} \int_{-\infty}^t dt \int_{-\infty}^{t-\varepsilon} d\tau \frac{1}{\varepsilon} \int d^3x \left( \bar{T}^{\mu\nu}(x), \bar{O}(y) \right)_{0\mu\nu}(x)
$$

$$
- \frac{1}{2} \lim_{\varepsilon \to 0} \frac{1}{i\beta} \int_{-\infty}^t dt \int_{-\infty}^{t-\varepsilon} d\tau e^{(t-\tau)} \int d^3x \left( \bar{S}^{\mu\nu}(x), \bar{O}(y) \right)_{0\mu\nu}(x)
$$

$$
\partial_{\mu\nu}(x) \delta_{\mu\nu}(x) \langle \bar{O}(y)\rangle_0 \partial_{\mu\nu}(x) \tag{70}
$$
Another useful (equivalent) expression of $\delta\langle \hat{O}(y) \rangle$ can be obtained starting from the expression (14) of $\hat{T}$, where the continuity equation for angular momentum is used from the beginning. Repeating the same reasoning as above, it can be shown that one gets:

$$\delta\langle \hat{O}(y) \rangle = \lim_{\varepsilon \to 0} \frac{1}{2\mu\beta} \int_{-\infty}^{t'} dt \frac{1 - e^{\varepsilon(t-t')}}{\varepsilon} \int d^3x \left( \left[ \hat{T}^{\mu}_{\nu}(x), \hat{O}(y) \right] \right) \langle 0 | (\partial_{\mu} \delta_{\nu}(x) + \partial_{\nu} \delta_{\mu}(x)) \rangle$$

Then, for Hermitian operators, what gets changed is the order of the operators besides their time argument. For $m$ follows:

$$\hat{O}_m^{(1)}(t, \mathbf{x}) \hat{O}_n^{(2)}(0, \mathbf{0}) = \hat{O}_n^{(2)}(0, \mathbf{0}) \hat{O}_m^{(1)}(t, \mathbf{x})$$

where $\hat{O}_m^{(1)}$ and $\hat{O}_n^{(2)}$ are physical tensor densities of rank $m$ and $n$, respectively.

The equilibrium density operator $\hat{\rho} = \exp(-\hat{H}/T)/Z$ is symmetric for space-time translations and rotations, as well as time reversal and parity if the hamiltonian is itself parity and time reversal invariant. The symmetry under this class of transformations allows to simplify the above expression. For any linear unitary transformation $\hat{U}$ which commutes with $\hat{\rho}$ one has:

$$\langle \hat{O} \rangle_0 = \text{tr} \left( \hat{\rho}_0 \hat{O} \right) = \text{tr} \left( \hat{\rho}_0 \hat{U} \hat{O} \hat{U}^{-1} \right) = \langle \hat{U} \hat{O} \hat{U}^{-1} \rangle_0$$

Taking $\hat{U} = \hat{T}(a)$ with $\hat{T}(a)$ a general translation operator:

$$\langle \left[ \hat{O}_1^{\mu_1 \cdots \mu_m}(t, \mathbf{x}), \hat{O}_2^{\nu_1 \cdots \nu_n}(0, \mathbf{0}) \right] \rangle_0 = \langle \hat{O}_2^{\nu_1 \cdots \nu_n}(0, \mathbf{0}), \hat{O}_1^{\mu_1 \cdots \mu_m}(t + a^0, \mathbf{x} + \mathbf{a}) \rangle_0$$

and so, setting $(a^0, \mathbf{a}) = (-t, -\mathbf{x})$:

$$\langle \left[ \hat{O}_1^{\mu_1 \cdots \mu_m}(t, \mathbf{x}), \hat{O}_2^{\nu_1 \cdots \nu_n}(0, \mathbf{0}) \right] \rangle_0 = \langle \hat{O}_2^{\nu_1 \cdots \nu_n}(0, \mathbf{0}), \hat{O}_1^{\mu_1 \cdots \mu_m}(-t, -\mathbf{x}) \rangle_0$$

Similarly, for a space inversion:

$$\langle \left[ \hat{O}_1^{\mu_1 \cdots \mu_m}(t, \mathbf{x}), \hat{O}_2^{\nu_1 \cdots \nu_n}(0, \mathbf{0}) \right] \rangle_0 = (-1)^{n_s + m_s} \langle \left[ \hat{O}_1^{\mu_1 \cdots \mu_m}(t, -\mathbf{x}), \hat{O}_2^{\nu_1 \cdots \nu_n}(0, \mathbf{0}) \right] \rangle_0$$

where $m_s$ and $n_s$ are the number of space indices among $\mu_1, \cdots \mu_m$ and $\nu_1, \cdots \nu_n$ respectively.

The time reversal operator $\hat{\Theta}$ is antiunitary, thus a point-dependent physical scalar operator $\hat{A}(t, \mathbf{x})$ transforms as follows:

$$\hat{\Theta} \hat{A}(t, \mathbf{x}) \hat{\Theta}^{-1} = \hat{A}^\dagger(-t, -\mathbf{x})$$

whence, for commutators:

$$\hat{\Theta} \left[ \hat{A}(t, \mathbf{x}), \hat{B}(t, \mathbf{x}) \right] \hat{\Theta}^{-1} = \left[ \hat{A}^\dagger(-t, x), \hat{B}^\dagger(-t, x) \right]$$

Then, for Hermitian operators, what gets changed is the order of the operators besides their time argument. For tensor hermitian observables and time-reversal symmetric hamiltonian, one obtains:

$$\langle \left[ \hat{O}_1^{\mu_1 \cdots \mu_m}(t, \mathbf{x}), \hat{O}_2^{\nu_1 \cdots \nu_n}(0, \mathbf{0}) \right] \rangle_0 = (-1)^{m_0 + n_0} \langle \left[ \hat{O}_2^{\nu_1 \cdots \nu_n}(0, \mathbf{0}), \hat{O}_1^{\mu_1 \cdots \mu_m}(-t, \mathbf{x}) \right] \rangle_0$$

where $m_0$ and $n_0$ are the number of time indices among $\mu_1, \cdots \mu_m$ and $\nu_1, \cdots \nu_n$ respectively.