Macroscopic Pure State of Light Free of Polarization Noise

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The preparation of completely non-polarized light is seemingly easy: an everyday example is sunlight. The task is much more difficult if light has to be in a pure quantum state, as required by most quantum-technology applications. The pure quantum states of light obtained so far are either polarized or, in rare cases, manifest hidden polarization: even if their intensities are invariant to polarization transformations, higher-order moments are not. We experimentally demonstrate the preparation of the macroscopic singlet Bell state, which is pure, completely non-polarized, and has no polarization noise. Simultaneous fluctuation suppression in three Stokes observables below the shot-noise limit is demonstrated, opening perspectives for noiseless polarization measurements. The state is shown to be invariant to polarization transformations. This robust highly entangled isotropic state promises to fuel important applications in photonic quantum technologies.

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Polarization properties of light are of interest in quantum optics and quantum information. There are many reasons, such as the relative simplicity of polarization interferometry and detection schemes \(1\), the similarity between the polarization properties of light and the properties of atoms or atomic ensembles \(2\), and several others. Description of polarization in quantum optics is based on the Stokes operators \(\hat{S}_0, \hat{S}_1, \hat{S}_2, \hat{S}_3\) \(3\). Quantum noise in polarization observables is given by the uncertainty relations

\[
\Delta S_i \Delta S_j \geq |\langle S_k \rangle|, \quad i \neq j \neq k = 1, 2, 3,
\]

and it can be suppressed in \(S_j\) at the expense of its increase in \(S_j\) (polarization squeezing \(3, 4\)). In the particular case where \(\langle S_0 \rangle = 0\), there are no fundamental restrictions for suppressing noise in \(S_1\) and \(S_j\) simultaneously. Moreover, one can imagine a situation where the mean values of all three Stokes operators, usually called the Stokes parameters, are equal to zero, \(\langle \hat{S}_1 \rangle = \langle \hat{S}_2 \rangle = \langle \hat{S}_3 \rangle = 0\). This case can be realized for entangled squeezed vacuum states \(6, 7\).

Macroscopic, i.e., containing many photons, squeezed vacuum attracts much attention nowadays as it is among very few ‘available’ macroscopic quantum systems with essentially nonclassical properties \(8\). In particular, it can be considered as a candidate for macroscopic Bell tests \(3, 10\). Besides, macroscopic squeezed vacuum has important applications such as gravitational wave detection \(11\), quantum memory \(12\), super-resolution \(13\), quantum efficiency absolute measurement \(14\), etc. Entangled squeezed vacuum is a particular case of squeezed vacuum involving at least four radiation modes, for instance, two polarization ones and two frequency ones. Its low-intensity analogues are two-photon Bell states, and the preparation relies on the same experimental scheme \(15, 16\).

The preparation scheme involves two collinear nondegenerate optical parametric amplifiers (NOPAs) pumped by coherent beams. The corresponding Hamiltonian takes one of the four possible forms,

\[
\dot{H}_{\psi^\pm} = i\hbar G(a_1^\dagger b_2^\dagger \pm b_1^\dagger a_2^\dagger) + \text{h.c.},
\]

where \(a, a^\dagger, b, b^\dagger\) are photon annihilation and creation operators in the horizontal and vertical polarization modes, respectively, the subscripts 1, 2 denote two frequency modes, with the frequencies \(\omega_1, \omega_2\), and the parameter \(G\) depends on the quadratic susceptibility and the pump power.

The states at the output can be written exactly as

\[
|\psi_{\text{mac}}^{(\pm)}\rangle = e^{\Gamma(a_1^\dagger b_2^\dagger \pm b_1^\dagger a_2^\dagger)}|\text{vac}\rangle,
\]

where \(\Gamma = \int G dt\) is the parametric gain coefficient. Due to their similarity to the two-photon Bell states, they can be called macroscopic Bell states. The Fock-state expansion of \(8\) immediately reveals photon-number correlation between modes \(a_1, b_2\) and between modes \(a_2, b_1\) \(8, 17\). This leads to the suppression of fluctuations in certain Stokes observables. However, for calculating observables such as Stokes parameters and their variances, it is more convenient to use the Heisenberg approach. For instance, the Hamiltonian \(H_{\phi^-}\) leads to the equations of motion

\[
\dot{a}_1 = Gb_2^\dagger, \quad \dot{b}_2 = Ga_1^\dagger, \quad \dot{a}_2 = -Gb_1^\dagger, \quad \dot{b}_1 = -Ga_2^\dagger,
\]

the dot denoting time differentiation. The solutions are

\[
a_1 = a_{10} \cosh \Gamma + b_{20}^\dagger \sinh \Gamma, \quad b_2 = b_{20} \cosh \Gamma + a_{10}^\dagger \sinh \Gamma,
\]

\[
a_2 = a_{20} \cosh \Gamma - b_{10}^\dagger \sinh \Gamma, \quad b_1 = b_{10} \cosh \Gamma - a_{20}^\dagger \sinh \Gamma,
\]
where ‘0’ labels vacuum operators. The mean values and the variances of the Stokes operators are then calculated by averaging the operators \( \hat{S}_i \) and \( \langle \hat{S}_i - \langle \hat{S}_i \rangle \rangle^2 \), respectively, over the vacuum state. Note that the Stokes operators for the output beam should be calculated as the sum of the Stokes operators for the two wavelengths, \( \hat{S}_i = \hat{S}_i^{(1)} + \hat{S}_i^{(2)} \), where \( \hat{S}_i^{(i)} \) is the \( i \)-th Stokes operator for mode \( j \); \( i = 0, 1, 2, 3; j = 1, 2 \). This is because the detectors measuring intensities in various polarization modes (Fig. 1a), even if fast enough to trace intensity fluctuations, are still too slow to measure the beats caused by the \( \omega_1 - \omega_2 \) difference.

For the Stokes parameters, we get

\[
\langle S_0 \rangle = 4 \sinh^2 \Gamma, \quad \langle S_1 \rangle = \langle S_2 \rangle = \langle S_3 \rangle = 0 \quad (4)
\]

for all states \( |\Psi\rangle \), which clearly shows that they are unpol- larized in the first order in the intensity. It is different, however, for higher-order moments. For the variances, \( \text{Var}(S_i) \equiv \langle S_i^2 \rangle - \langle S_i \rangle^2 = \Delta S_i^2 \); we get the results shown in Table 1. We see that for each of the macroscopic Bell states, variance of at least one of the Stokes operators is zero, i.e., noise in the corresponding Stokes observable is completely suppressed \([3, 7]\).

### Table I. Variances of the Stokes observables for the four states \( |\Psi\rangle \), \( n \equiv \sinh^2 \Gamma \). Each state has fluctuations in some Stokes observable suppressed; the \( |\Psi_{mac}\rangle \) state has no fluctuations in all Stokes observables \( S_{1,2,3} \).

| State          | \( \text{Var}(S_0) \) | \( \text{Var}(S_1) \) | \( \text{Var}(S_2) \) | \( \text{Var}(S_3) \) |
|----------------|-----------------------|-----------------------|-----------------------|-----------------------|
| \( |\Psi_{mac}\rangle \) | \( 8n(n + 1) \) | 0 | 0 | 0 |
| \( |\Psi_{mac}\rangle \) | \( 8n(n + 1) \) | 8n(n + 1) | 8n(n + 1) | 8n(n + 1) |
| \( |\Phi_{mac}\rangle \) | \( 8n(n + 1) \) | 8n(n + 1) | 0 | 8n(n + 1) |
| \( |\Phi_{mac}\rangle \) | \( 8n(n + 1) \) | 8n(n + 1) | 8n(n + 1) | 0 |

The noise properties of the macroscopic Bell states can be illustrated by a diagram shown in Fig. 1. For each state, the coloured distribution is centered around the point \( \langle S_1 \rangle; \langle S_2 \rangle; \langle S_3 \rangle \), while the size of each distribution in direction \( S_i \) corresponds to the uncertainty \( \Delta S_i \). For the ‘triplet’ states \( |\Psi_{mac}\rangle, |\Phi_{mac}^{(+)}\rangle, |\Phi_{mac}^{(-)}\rangle \), the distributions have the shapes of discs, as the uncertainty is zero in only one direction. For \( |\Phi_{mac}^{(-)}\rangle \), the distribution is pointlike, as all \( \Delta S_i = 0 \). Formally, these distributions can be described by a quasiprobability function \([21]\) introduced in Ref. \([22, 23]\), similarly to the way distributions of optical quadratures can be described by the Wigner function.

The degree of noise suppression for a Stokes observable \( S_i \) can be characterized by its normalized variance, \( \text{Var}(S_i)/\langle S_0 \rangle \). Normalized this way, the variance turns into the noise reduction factor (NRF) for the beams in two orthogonal polarization modes (horizontal and vertical for \( i = 1 \), linear ±45 deg for \( i = 2 \), and right- and left-circular for \( i = 3 \)). NRF, defined as the variance of the intensity difference for two beams normalized to their mean intensity sum \([24]\), is a parameter commonly used to quantify twin-beam squeezing. The value NRF = 1, often called the shot-noise level, is realized for the case of two coherent beams and sets the boundary between the classical and quantum behaviour. Although theoretically NRF can reach zero values (see Table 1), it is never the case in a real setup, because of the finite quantum efficiencies, inevitable optical losses, and mismatch of signal and idler mode selection \([23]\). If all losses are incorporated into some effective quantum efficiency \( \eta \), equal for both detectors, the zeros in Table 1 turn into \( 4n(1 - \eta) = (1 - \eta)\langle S_0 \rangle \). Of course, losses also influence the anti-squeezed variances of the Stokes observables, which become \( 4n(1 + \eta) + 8n^2\eta \equiv (1 + \eta)(\langle S_0 \rangle + \eta\langle S_0 \rangle^2)/2 \). The corresponding minimum and maximum values of NRF are, respectively, \( 1 - \eta \) and \( 1 + \eta + \eta\langle S_0 \rangle/2 \).

In our experiment (Fig. 1), we generate the macroscopic singlet state \( |\Psi_{mac}^{(-)}\rangle \) via frequency-nondegenerate parametric down-conversion (PDC) in two 2 mm BBO crystals with optic axes oriented in orthogonal (horizontal and vertical) planes placed into a Mach-Zehnder interferometer (MZI) whose input and output beamsplitters are polarizing ones. The crystals are pumped by Nd:YAG laser third harmonic (wavelength \( \lambda_p = 355 \text{nm} \), repetition rate 1 kHz, pulse duration 17 ps, energy per pulse up to 0.2 mJ). The pump at the input of the MZI is 45° polarized and hence it contributes equally to PDC in both crystals. After the crystals, the pump radiation is cut off by dichroic mirrors reflecting 99.5% of the pump and transmitting 95% of the PDC radiation. Further suppression of the residual pump radiation is made by means of a color-glass OG filter. Each crystal is a traveling-wave NOPA producing a two-color bright squeezed vacuum \([25]\) with signal and idler wavelengths \( \lambda_1 = 635 \text{nm} \), \( \lambda_2 = 1370 \text{nm} \).
and therefore provides the necessary transformation from $(A)$, and two detectors. A quarter-wave plate (QWP), a lens, an aperture (A), and two detectors. b: the wavelength-angular spectrum of PDC, the shaded area showing the range selected by the aperture. c: macroscopic Bell states. Three coloured discs show the mean values and the uncertainties of the Stokes observables for $\psi^{+}_{mac}$ (blue), $\psi^{-}_{mac}$ (green), and $\psi^{0}_{mac}$ (red). The yellow point shows the state $\psi^{+}_{mac}$, for which the means and uncertainties of all Stokes observables are zero.

The registration part of the setup (Fig. 1h, left), including a Glan prism, a half-wave plate (HWP) and a quarter-wave plate (QWP), and two detectors, provides a standard Stokes measurement. With a HWP oriented at 22.5°, the difference of detectors’ output signals corresponds to $S_2$, while with a QWP oriented at 45°, the same measurement provides $S_3$. Intermediate directions in the space of the Stokes variables are accessed by continuously rotating the HWP or QWP. Mode selection is performed by means of a lens with the focal length 300 mm, placed at 300 mm from the crystals, and a 10-mm aperture, placed in its focal plane. This way we select a 1.8° angular spectrum width (Fig. 1d), which automatically restricts the PDC frequency spectrum as well. All optical elements are anti-reflection coated. The radiation is focused on the detectors by means of lenses with focal lengths 5 cm. The detectors are pin photodiodes followed by charge-sensitive amplifiers, and their output pulses have a fixed shape (with the duration about 8 μs) and the amplitudes scaling as the number of photons registered during a single pump pulse. The electronic noise is mostly caused by the amplification circuit and is equivalent to 180 input photons. For each pump pulse, the output pulses of the detectors are integrated in time by means of an AD card; then the data are processed to obtain the mean values and variances of the Stokes parameters. The shot-noise level is measured separately using attenuated laser radiation. In more detail, the procedure for measuring the SNR is described in Ref. [28]. However, unlike in Ref. [28], in our present experiment the squeezed-vacuum pulses contain, on the average, more than $10^6$ photons, hence the shot noise exceeds the electronic noise by an order of magnitude.

Using this setup, we have studied the polarization properties of the macroscopic singlet state. We measured NRF for $S_1$, $S_2$, $S_3$, as well as the intermediate Stokes observables corresponding to arbitrary orientations of HWP and QWP (Fig. 2a,b, respectively). Certain points in the obtained dependences correspond to the measurement of noise reduction in the Stokes observables $S_1$ (Fig. 2a,b, blue dashed lines), $S_2$ (Fig. 2a, green dotted lines) and $S_3$ (Fig. 2b, red solid lines). For more clarity, arrows on top of each figure indicate the measurement of $S_1$, $S_2$, $S_3$. We see that all three observables $S_{1,2,3}$ have fluctuations suppressed about 30% below the shot-noise level. In particular, we have measured $NRF(S_1) = 0.72 \pm 0.01$, $NRF(S_2) = 0.72 \pm 0.01$, $NRF(S_3) = 0.73 \pm 0.02$. This moderate noise suppression, despite the overall quantum efficiency/transmittivity of the setup being $\eta \approx 0.65$, is mainly due to the fact that both 635nm and 805nm beams are restricted by the same angular aperture: for proper mode matching, the apertures for different wavelengths should differ in size $\pm 2\mu$m.

In addition to noise reduction, the data clearly show the invariance of the singlet state to polarization transformations with HWP and QWP. The normalized Stokes
The same way as the orientation of the HWP (between the intensities in orthogonal polarization modes relevant for the polarization observables. One can say that this state is unpolarized state of light with noise suppressed in all polarization properties of the macroscopic singlet state $|\Psi^{(-)}_{\text{mac}}\rangle$ (a,b) and another macroscopic Bell state, $|\Psi^{(+)}_{\text{mac}}\rangle$ (c,d). Green large empty circles: NRF of the Stokes observable versus the orientation of the HWP (a,c) and QWP (b,d). Blue dashed lines mark the measurement of Var($S_1$)/$\langle S_0 \rangle$, green dotted lines, of Var($S_2$)/$\langle S_0 \rangle$, and red solid lines of Var($S_3$)/$\langle S_0 \rangle$. Small squares denote normalized signals in both detectors.

variety (large green open circles), as well as the intensities in the two output channels (small red and blue squares), do not depend on the orientations of the plates. This justifies the term ‘polarization-scalar light’. The state is a pure one but, at the same time, completely unpolarized. For comparison, we have also tested the polarization properties of $|\Psi^{(+)}_{\text{mac}}\rangle$, which was generated the same way as $|\Psi^{(-)}_{\text{mac}}\rangle$ but without the dichroic phase plate. The results are shown in Fig. 2d. One can see that although the output intensities are independent of the polarization rotation, the normalized variances show well-pronounced modulation, which proves that the state has nonzero second-order degree of polarization $\langle S_1 \rangle^2 / \langle S_0 \rangle$. The other two macroscopic Bell states also reveal this hidden polarization effect, but these results deserve a separate consideration and will be published elsewhere [29].

In the context of this work, it is important that the singlet state is polarized neither in the first order in the intensity nor in the second one.

Our results demonstrate the existence of pure intense unpolarized state of light with noise suppressed in all polarization observables. One can say that this state manifests the strongest possible nonclassical correlation between the intensities in orthogonal polarization modes regardless of the choice of these modes. This makes macroscopic singlet Bell state a candidate for macroscopic Bell tests and other challenging fundamental experiments. Apart from its fundamental interest, this state will certainly find important applications in quantum technologies based on the interaction between light and matter.

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[1] N. Korolkova et al., Phys. Rev. A. 65, 052306 (2002).
[2] J. L. Sørensen, J. Hald, and E. S. Polzik, Phys. Rev. Lett. 80, 3487 (1998).
[3] A. S. Chirkin, A. A. Orlov, and D. Yu. Parashuk, Quantum Electron. 23, 870 (1993).
[4] W P. Bowen et al., Phys. Rev. Lett. 88, 093601 (2002).
[5] S. L. Braunstein, Phys. Rev. A 71, 055801 (2005).
[6] V. P. Karassiov, J. Phys. A 26, 4945 (1993).
[7] V. P. Karasev, A. V. Masalov, Opt. Spectrosc. 74, 551 (1993).
[8] F. De Martini, F. Sciarrino, and C. Vitelli, Phys. Rev. Lett. 100, 235601 (2008).
[9] Ch. Simon and D. Bouwmeester, Phys. Rev. Lett. 91, 053601 (2003).
[10] P. Sekatski et al., Phys. Rev. Lett. 103, 113601 (2009).
[11] K. McKenzie et al., Phys. Rev. Lett. 88, 231102 (2002).
[12] J. S. Neergard-Nielsen et al., Phys. Rev. Lett. 97, 083604 (2006).
[13] P. M. Anisimov et al., Phys. Rev. Lett. 104, 103602 (2010).
[14] I. N. Agafonov et al., arXiv:1008.0979v1 [quant-ph]; Giorgio Brida et al., arXiv:1005.2937v1 [quant-ph].
[15] Y.-H. Kim, S. P. Kulik, and Y. H. Shih, Phys. Rev. A 63, 060301(R) (2001).
[16] A. V. Burlakov et al., JETP 95, 639 (2002).
[17] Ch. Vitelli et al., Phys. Rev. A 81, 032123 (2010).
[18] D. N. Klyshko, JETP 84, 1020 (1997).
[19] P. Usachev et al., Opt. Commun. 193, 161 (2001).
[20] A. Sehat et al., Phys. Rev. A 71, 033818 (2005).
[21] P. A. Bushev et al., Optics and Spectroscopy 91, 526 (2001).
[22] K. B. Wolf, Opt. Commun. 132, 343 (1996).
[23] V. P. Karassiov and A.V.Masalov, J. Opt. B 4, 366 (2002).
[24] O. Aytur and P. Kumar, Phys. Rev. Lett. 65, 1551 (1990).
[25] I. N. Agafonov, M. V. Chekhova, and G. Leuchs, Phys. Rev. A 82, 011801(R) (2010).
[26] G. Leuchs, R. Dong and D. Sych, New J. of Phys. 11, 113040 (2009).
[27] A. B. Klimov et al., Phys.Rev.Lett. 105, 153602 (2010).
[28] T. Sh. Iskhakov, M. V. Chekhova, and G. Leuchs, Phys. Rev. Lett. 102, 183602 (2009).
[29] T. Sh. Iskhakov, M. V. Chekhova, and G. Leuchs, to be published.