Covariance of noncommutative Graßmann star product

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November 20, 2018

Abstract
Using the Coherent states of many fermionic degrees of freedom labeled by Graßmann variables, we introduce the noncommutative (precisely non anticommutative) Graßmann star product. The covariance of star product under unitary transformations, particularly canonical ones, is studied. The super star product, based on supercoherent states of supersymmetric harmonic oscillator, is also considered.
1 Introduction

In 1949, Moyal introduced a new bracket for functions on the classical phase space that replaces the Poisson one in the quantization procedure [1]. This bracket is closely related to Weyl’s correspondence rule between classical and quantum observables [2]. Recently, deformation à la Moyal has been explored in several context as noncommutative approach to string theory [3], matrix model [4], the non-commutative Yang-Mills theories [5] and non-commutative gauge theories [6]. More generally, models on noncommutative space time, obtained by replacing the ordinary product of functions by the noncommutative Moyal product, helped to revive the deformation quantization technique which was elucited further in [7]. Through this article, we will deal with the concept of coherent states which as it is well established have important application building a bridge between the classical and the quantum worlds views [8,9]. In other side, coherent states have fundamental properties like over-completion and non-orthogonality which constitute the main ingredients to formulate the Moyal product in coherent state framework. The first result obtained in this sens is due to Alexanian and al [10](see also [11]). Coherent states of quantum systems with nonlinear spectrum has been also considered to find new kinds of star products [12] and ones related to su(2) algebra has been nicely used to introduce a new star product on the fuzzy sphere [13].

The main task of this work is to derive star product for graßmannian manifold using the fermionic coherent states. Section 2 collects some notes on Graßmann variables [14] and coherent states of a system with finite number of fermions [15]. Section 3 develops a consistant star product using coherent states associated to fermionic degrees of freedom. Covariance with respect the unitary transformation is studied in section 4. It will be shown that the covariance is not guaranteed under arbitrary unitary transformation except in the particular case of canonical ones. In section 5, the correspondence rules between functions and operators are used to generate the "star"-analogue of the supersymmetric oscillator. Concluding remarks close this note.

2 Basic concepts of fermionic coherent states

2.1 Graßmann variables

It is well established that the Graßmann variables serve as classical analogues of fermionic degrees of freedom. Indeed, the phase space, in the classical sens, of a collection of N fermions can be identified with the Graßmann algebra CB_{2N} which is generated by a set of 2N independants Graßmann variables obeying the anti-commutation relation

{\theta_i, \theta_j} = \delta_{ij}, \quad {\bar{\theta}_i, \theta_j} = {\bar{\theta}_i, \bar{\theta}_j} = 0 \quad (1)

where i = 1, 2, ..., N. This algebra is Z_2-graded; i.e the even elements of CB_{2N} are commuting and have even Graßmann parity. The odd elements are anti-commuting
with odd parity. The Graßmann integration and differentiation are given by

\[ \int d\theta = 0 \quad \int d\theta \theta = 0 \quad \frac{d}{d\theta} 1 = 0 \quad \frac{d}{d\theta} \theta = 1 \quad (2) \]

where \( \theta \) stand for \( \theta_i \) and \( \bar{\theta}_i \) \( (i = 1, 2, ..., N) \). Note that the integration and differentiation are treated like odd Graßmann quantities according to \( \mathbb{Z}_2 \)-graduation (for more details see [14]).

### 2.2 Fermionic coherent states

We consider a finite number \( N \) of fermionic degrees of freedom which are characterized by annihilation and creation operators \( f_i^- \) and \( f_i^+ = (f_i^-)^+ \) \( (i = 1, 2, ..., N) \) satisfying the canonical anti-commutation relations

\[ \{ f_i^-, f_j^- \} = \{ f_i^+, f_j^+ \} = 0 \quad \{ f_i^-, f_j^+ \} = \delta_{ij} \quad (3) \]

The corresponding Hilbert space is the \( N \)-fold tensor product \( \mathcal{H}_f = \mathcal{H}_1 \otimes \mathcal{H}_2 \otimes ... \otimes \mathcal{H}_N \) of two dimensional Fock space

\[ \mathcal{H}_i = \left\{ |0\rangle_i, |1\rangle_i \right\} \quad (4) \]

The operators \( f_i^- \) act on \( \mathcal{H}_i \) as

\[ f_i^- |0\rangle_i = 0 \quad f_i^- |1\rangle_i = |0\rangle_i \quad (5) \]

The Hilbert space \( \mathcal{H}_f \) is spanned by the eigenstates

\[ |n_1, n_2, ..., n_N\rangle \equiv |n_1\rangle_1 \otimes |n_2\rangle_2 \otimes ... \otimes |n_N\rangle_N \quad (6) \]

The fermionic coherent states are defined in analogy with the standard canonical harmonic oscillator. They are defined as eigenstates of annihilation operators \( f_i^- \) and are labeled by Graßmann variables. They can be defined also as the action of a displacement operator on the ground state of the fermionic system under consideration. The coherent state of one fermionic degree of freedom is given by

\[ |\theta\rangle \equiv \exp(\theta f^+ + \bar{\theta} f^-)|0\rangle \quad (7) \]

Using the action of annihilation and creation operators on the one fermion Hilbert space, one has

\[ |\theta\rangle = (1 - \frac{1}{2} \bar{\theta} \theta) [|0\rangle + \theta |1\rangle] = \exp(\frac{1}{2} \bar{\theta} \theta) [|0\rangle + \theta |1\rangle] \quad (8) \]

Note that \( |\theta\rangle \) is eigenstate of the operator \( f^- \) with eigenvalue \( \theta \). The normalized states (8) form an overcomplete set and provide a resolution of the identity

\[ \int d\mu(\theta, \bar{\theta}) |\theta\rangle \langle \theta| = 1 \quad (9) \]
where the measure is defined by \( d\mu(\theta, \bar{\theta}) = d\theta d\bar{\theta} \).

For \( N \) fermionic degrees of freedom, we define the coherent states, as tensorial product of the state \( |\theta_i\rangle_i \) of the fermion \( i \)

\[
|\vec{\theta}\rangle = |\theta_1\rangle_1 \otimes |\theta_2\rangle_2 \otimes |\theta_N\rangle_N
\]

where

\[
|\theta_i\rangle_i = (1 - \frac{1}{2} \bar{\theta}_i \theta_i) [0, \theta_i]
\]

The coherent states \( |\vec{\theta}\rangle \) form an overcomplete basis for the Hilbert space \( \mathcal{H}_f \) and the completeness relation is given by

\[
\int d\mu(\vec{\theta}, \vec{\bar{\theta}}) |\vec{\theta}\rangle \langle \vec{\theta}| = 1
\]

where the measure is given now by

\[
d\mu(\vec{\theta}, \vec{\bar{\theta}}) = \prod_{i=1}^{N} d\bar{\theta}_i d\theta_i.
\]

### 3 Grassmann star product

The main ingredient to introduce the Grassmann star product is the set of the fermionic coherent states \( |\vec{\theta}\rangle \). For this end, let us consider an operator \( A \), acting on the Hilbert space \( \mathcal{H}_f \), which can be expanded in terms of ladder fermionic operators as

\[
A = \sum_{\vec{m}, \vec{n}} a_{\vec{m}, \vec{n}} f^{+m_1}_{1} f^{+n_2}_{2} \cdots f^{+n_N}_{N} f^{-m_1}_{1} f^{-m_2}_{2} \cdots f^{-m_N}_{N}
\]

where \( \vec{m} = (m_1, m_2, ..., m_N) \) and \( \vec{n} = (n_1, n_2, ..., n_N) \) are \( N \)-dimensional vectors. According the nilpotency of the fermionic operators, the admissible values of integers \( m_i \) and \( n_i \) are 0 or 1. In the fermionic coherent states representation, we associate to the operator \( A \) the following function

\[
\mathcal{A}(\vec{\theta}, \vec{\bar{\theta}}) = \langle \vec{\theta}|A|\vec{\theta}\rangle = \sum_{\vec{m}, \vec{n}} a_{\vec{m}, \vec{n}} \bar{\theta}_1^{m_1} \bar{\theta}_2^{m_2} \cdots \bar{\theta}_N^{m_N} \theta_1^{n_1} \theta_2^{n_2} \cdots \theta_N^{n_N}
\]

in terms of the Grassmann variables \( \theta_i \) and \( \bar{\theta}_i (i = 1, 2, ..., N) \). As in the bosonic case [10-12], an associative product of two functions is then defined by

\[
\mathcal{A}(\vec{\theta}, \vec{\bar{\theta}}) * \mathcal{B}(\vec{\theta}, \vec{\bar{\theta}}) = \langle \vec{\theta}|AB|\vec{\theta}\rangle
\]

The associativity of this star product originates from the associativity of the algebra generated by fermionic creation and annihilation operators. Using the overcompletion property (12), the star product (16) can be written as

\[
\mathcal{A}(\vec{\theta}, \vec{\bar{\theta}}) * \mathcal{B}(\vec{\theta}, \vec{\bar{\theta}}) = \int d\mu(\vec{\eta}, \vec{\bar{\eta}}) \langle \vec{\theta}|A|\vec{\eta}\rangle \langle \vec{\eta}|B|\vec{\theta}\rangle
\]
Remark that the matrix elements \( \langle \vec{\theta} | A | \vec{\eta} \rangle \) can be expressed as the action of a translation operator on the function \( A(\vec{\theta}, \vec{\bar{\theta}}) \)

\[
e^{-\theta_i \frac{\partial}{\partial \eta_i}} e^{\eta_i \frac{\partial}{\partial \theta_i}} A(\vec{\theta}, \vec{\bar{\theta}}) = \frac{\langle \vec{\theta} | A | \vec{\eta} \rangle}{\langle \vec{\theta} | \vec{\eta} \rangle} \tag{18}
\]

and

\[
e^{-\theta_i \frac{\partial}{\partial \bar{\eta}_i}} e^{\bar{\eta}_i \frac{\partial}{\partial \bar{\theta}_i}} A(\vec{\theta}, \vec{\bar{\theta}}) = \frac{\langle \vec{\eta} | A | \vec{\theta} \rangle}{\langle \vec{\eta} | \vec{\theta} \rangle} \tag{19}
\]

where we have summation over repeated indices. The translation operators appearing in (18) and (19) can be written as ordered exponential

\[
\langle \eta_i - \theta_i \rangle \frac{\partial}{\partial \eta_i} : e^{\eta_i \frac{\partial}{\partial \theta_i}} : \tag{20}
\]

and

\[
\langle \eta_i - \theta_i \rangle \frac{\partial}{\partial \theta_i} : e^{\eta_i \frac{\partial}{\partial \bar{\theta}_i}} : \tag{21}
\]

respectively. So, the Grassmann star product (16) is given now by

\[
\star = \int d\mu(\vec{\eta}, \vec{\bar{\eta}}) : e^{\eta_i \frac{\partial}{\partial \theta_i}} : \langle \vec{\theta} | \vec{\eta} \rangle^2 : e^{\eta_i \frac{\partial}{\partial \bar{\theta}_i}} : \langle \vec{\bar{\theta}} | \vec{\bar{\eta}} \rangle \tag{22}
\]

From the equations (10) and (11), the overlapping of two fermionic coherent states is

\[
|\langle \vec{\theta} | \vec{\eta} \rangle|^2 = e^{-(\eta_i - \theta_i)(\eta_i - \theta_i)} \tag{23}
\]

Substituting (23) in (22) and computing the integral (22) with a simple change of variables, we obtain

\[
\star = e^{(\eta_i - \theta_i)(\eta_i - \theta_i)} \tag{24}
\]

This is the Grassmann star product. The computation of the star product of two arbitrary functions \( A(\vec{\theta}, \vec{\bar{\theta}}) \) and \( B(\vec{\bar{\theta}}, \vec{\bar{\theta}}) \) becomes easy thanks to correspondence between operators acting on the fermionic Hilbert space and functions of Grassmannian variables (Eq.(15)) and the minimal set of the needed relations, to perform the star calculus, are

\[
\bar{\theta}_i \star \theta_j = \bar{\theta}_i \theta_j \tag{25}
\]

\[
\theta_i \star \theta_j = \theta_i \theta_j \tag{26}
\]

\[
\bar{\theta}_i \star \bar{\theta}_j = \bar{\theta}_i \bar{\theta}_j \tag{27}
\]

\[
\theta_i \star \bar{\theta}_j = \theta_i \bar{\theta}_j + \delta_{ij} \tag{28}
\]

Notice that in the star formulation, we have the following star anti-commutation

\[
\{\theta_i, \theta_j\}_\star = \{\bar{\theta}_i, \bar{\theta}_j\}_\star = 0
\]
\[ \{\bar{\theta}_i, \bar{\theta}_j\}_\star = \delta_{ij} \] (29)

which are the star analogue of the anti-commutation relation of the fermionic operators. Note also that the nilpotency of Grassmann variables

\[ \theta_i \star \theta_i = 0 \] (30)

is preserved in the star language.

4 Covariance of Grassmann star product under unitary transformation

In this section, we analyse the covariance of Grassmann star product under unitary transformation. In this order, we consider the unitary transformation of the operator \( A \) under \( O = e^\Lambda \) such that \( \Lambda^+ + \Lambda = 0 \)

\[ A' = O^+ AO = e^{-\Lambda} Ae^\Lambda \]

\[ = \sum_{k=0}^{\infty} \frac{(-1)^k}{k!} \left[ \Lambda, [\Lambda, \ldots, [\Lambda, A]] \right] \] (31)

The function associated to the operator \( A' \) is given by

\[ \mathcal{A}'(\bar{\theta}, \bar{\theta}) = \langle \bar{\theta} | A' | \bar{\theta} \rangle \]

\[ = \sum_{k=0}^{\infty} \frac{(-1)^k}{k!} \left\{ \lambda, \{ \lambda, \ldots, \{ \lambda, A\}_\star \}_\star \right\}_\star \quad (32) \]

where \( \lambda \) is the function associated to the anti-hermitian operator \( \Lambda \)

\[ \lambda \equiv \lambda(\bar{\theta}, \bar{\theta}) = \langle \bar{\theta} | \lambda | \bar{\theta} \rangle \] (33)

In the noncommutative Grassmann space, the function \( \mathcal{A}' \) can be written also as

\[ \mathcal{A}'(\bar{\theta}, \bar{\theta}) = e^{-D_\lambda} \mathcal{A}(\bar{\theta}, \bar{\theta}) \] (34)

where the action of \( D_\lambda \) on the functions defined on the noncommutative Grassmann space is defined by

\[ D_\lambda \mathcal{A}(\bar{\theta}, \bar{\theta}) = \lambda(\bar{\theta}, \bar{\theta}) \star \mathcal{A}(\bar{\theta}, \bar{\theta}) - \mathcal{A}(\bar{\theta}, \bar{\theta}) \star \lambda(\bar{\theta}, \bar{\theta}) \] (35)

Hence, the unitary transformation \( O = e^\Lambda \) in the operator space has unique representation in the noncommutative Grassmann space formulated by equation (34). One can verify also that

\[ e^\Lambda \star e^\Lambda = e^\Lambda \star e^{-\Lambda} = 1 \] (36)
where the star exponent is defined as follows

$$e^x_\star = \sum_{k=0}^{\infty} \frac{1}{k!} (x^*)^k$$

Remark that for two transformations $O_1 = e^{\Lambda_1}$ and $O_2 = e^{\Lambda_2}$ acting on the operator space, one has

$$[D_{\lambda_1}, D_{\lambda_2}] = D_{\{\lambda_1, \lambda_2\}}$$

where the functions $\lambda_1$ and $\lambda_2$ are ones associated to anti-hermitians operators $\Lambda_1$ and $\Lambda_2$, respectively.

Using the previous results, one can write the functions $A'$ in the form

$$A'_{\vec{\theta}, \vec{\bar{\theta}}} = \sum_{\vec{m}, \vec{n}} a_{\vec{m}, \vec{n}} \bar{\psi}^{n_1}_1 \cdots \bar{\psi}^{n_N}_N \psi^m_1 \cdots \psi^m_N$$

where the new variables $\psi$ are related to old ones as follows

$$\psi_i = e^{-D_{\lambda} \theta_i}, \quad \bar{\psi}_i = e^{-D_{\lambda} \bar{\theta}_i}$$

It is clear that, for an arbitrary unitary transformation, we have

$$A'_{\vec{\theta}, \vec{\bar{\theta}}} \neq A_{\vec{\psi}, \vec{\bar{\psi}}}$$

and the covariance is violated. A necessary condition to get a covariance is

$$e \left( \frac{\partial}{\partial \psi_i} \right) \left( \frac{\partial}{\partial \bar{\psi}_i} \right) = e \left( \frac{\partial}{\partial \theta_i} \right) \left( \frac{\partial}{\partial \bar{\theta}_i} \right)$$

with summation over the repeated indices. This situation occurs in the exceptional case in which the transformation $O$ belongs to the group of real orthogonal $2N \times 2N$ matrices, $SO(2N, R)$. To clarify how the classical covariance is preserved under canonical transformations belonging to $SO(2N, R)$, let us consider linear canonical transformations; i.e, the linear homogenous transformations of the fermionic operators $f_i^-$ and $f_i^+$, that do not change the commutations relations (3)

$$F_i^- = u_{ij} f_j^- + v_{ij} f_j^+$$

which can written in compcat matrical form as follows

$$F^- = U f^- + V f^+$$

The matrix elements of $U$ and $V$ are $u_{ij}$ and $v_{ij}$ respectively. The necessary relations for the transformation (44) to be canonical are

$$U V' + V U' = 1 \quad U U'^+ + V V'^+ = 1$$

where $'$ stands for matrix tranposition and $^+$ for hemitian conjugason.

According to Stone-Von Neuman theorem (see [8] for instance), there exist an unitary operator $O$ such that the operators $F_i^-$ and $f_i^-$ are equivalents

$$F_i^- = O^+ f_i^- O$$
It is given by
\[ O = e^{\alpha_{ij} f_i f_j + \alpha_{ij}^* f_i^* f_j^*} = e^\Lambda \] (47)
where \( \alpha_{ij} \) are elements of complex anti-symmetrical matrix \( \alpha \). A direct computation leads to
\[ F^- = \cos \sqrt{\alpha^+ \alpha} f^- + \frac{\alpha}{\sqrt{\alpha^+ \alpha}} \sin \sqrt{\alpha^+ \alpha} f^+ \] (48)
\[ F^+ = \cos \sqrt{\alpha^+ \alpha} f^+ + \frac{\alpha^+}{\sqrt{\alpha^+ \alpha}} \sin \sqrt{\alpha^+ \alpha} f^- \] (49)
Using the operator-function correspondence rule (15), the function associated to the operator \( \Lambda \) is
\[ \lambda \equiv \lambda(\vec{\theta}, \vec{\bar{\theta}}) = \frac{\alpha_{ij}}{2} \bar{\theta}_i \bar{\theta}_j + \frac{\alpha_{ij}}{2} \theta_i \theta_j \] (50)
The action of the operator \( D_\lambda \) on the canonical Grassmann variables gives
\[ D_\lambda \theta_k = -\alpha_{ki} \bar{\theta}_i \quad D_\lambda \bar{\theta}_k = \alpha_{jk} \theta_j \] (51)
which can be written in a compact form as
\[ D_\lambda \theta = -\alpha \bar{\theta} \quad D_\lambda \bar{\theta} = \alpha^+ \theta, \] (52)
and from which one get
\[ e^{-D_\lambda \theta} = \cos \sqrt{\alpha^+ \alpha} \theta + \frac{\alpha}{\sqrt{\alpha^+ \alpha}} \sin \sqrt{\alpha^+ \alpha} \bar{\theta} = \psi \] (53)
\[ e^{-D_\lambda \bar{\theta}} = \cos \sqrt{\alpha^+ \alpha} \bar{\theta} + \frac{\alpha^+}{\sqrt{\alpha^+ \alpha}} \sin \sqrt{\alpha^+ \alpha} \theta = \bar{\psi} \] (54)
At this stage, it is interesting to note that it is allowed to label coherent states associated to the new fermionic algebra \( \{F^-_i, F^+_i\} \) by the variables \( \psi_i \) and \( \bar{\psi}_i \). Then, by virtue of the correspondence rule (15), the variables \( \psi_i \) (resp. \( \bar{\psi}_i \)) are associated to annihilation operators \( F^-_i \) (resp. creation operators \( F^+_i \)) and following the analysis, presented in section 2, a new star product can be defined as
\[ \tilde{\star} = e^{\overleftarrow{\partial} \psi_i \overrightarrow{\partial} \bar{\psi}_i}, \] (55)
and we have
\[ \bar{\psi}_i \tilde{\star} \psi_j = \bar{\psi}_i \psi_j \] (56)
\[ \psi_i \tilde{\star} \psi_j = \psi_i \psi_j \] (57)
\[ \bar{\psi}_i \tilde{\star} \bar{\psi}_j = \bar{\psi}_i \bar{\psi}_j \] (58)
\[ \psi_i \tilde{\star} \bar{\psi}_j = \psi_i \bar{\psi}_j + \delta_{ij} \] (59)
The new canonical variables \( \psi_i \) and \( \bar{\psi}_j \) satisfy the following star anti-commutation
\[ \{\psi_i, \psi_j\}_\star = \{\bar{\psi}_i, \bar{\psi}_j\}_\star = 0 \]
Finally, one conclude that, under canonical transformations, the anti-commutations relations of Graßmann algebra are preserved, the products $\star$ and $\tilde{\star}$ are identical and the covariance is guaranteed

$$ \mathcal{A}'(\vec{\theta}, \vec{\bar{\theta}}) = e^{-D \lambda} \mathcal{A}(\vec{\theta}, \vec{\bar{\theta}}) = \mathcal{A}(\vec{\psi}, \vec{\bar{\psi}}) $$

(61)

counter to the case of arbitrary unitary transformations.

5 \hspace{1em} Super $\star$-analogue of the supersymmetric oscillator

We start by examining the simplest case of fermionic oscillator. The Hamiltonian is given by

$$ H_f = f^+ f^- $$

(62)

where $f^+$ and $f^-$ are creation and annihilation operators (the eigenstates are $|0\rangle$ and $|1\rangle$ with eigenvalues $e_0 = 0$ and $e_1 = 1$, respectively). Using the fermionic coherent states (8), the function associated to the Hamiltonian $H$ is given by

$$ h(\theta, \bar{\theta}) = \langle \theta | H | \theta \rangle = \bar{\theta} \theta $$

(63)

Any operator, acting on the fermionic Hilbert space, can be expanded in term of the elements of the occupation number operators basis \{P_{m,n} = |m\rangle\langle n|\}. The functions associated to elements of this basis are defined by

$$ P_{m,n}(\theta, \bar{\theta}) = \langle \theta | m \rangle \langle n | \theta \rangle $$

(64)

The function $P_{0,0}(\theta, \bar{\theta})$ provide a $\star$-vacuum. Indeed, we have

$$ \theta \star P_{0,0}(\theta, \bar{\theta}) = P_{0,0}(\theta, \bar{\theta}) \star \bar{\theta} = 0 $$

(65)

Furthermore, we have the following completion relation

$$ P_{0,0}(\theta, \bar{\theta}) + P_{1,1}(\theta, \bar{\theta}) = 1 $$

(66)

which is the analogue of the completion relation $|0\rangle\langle 0| + |1\rangle\langle 1| = 1$. One can verify also that

$$ h(\theta, \bar{\theta}) \star P_{m,n}(\theta, \bar{\theta}) = e_m P_{m,n}(\theta, \bar{\theta}) $$

(67)

$$ P_{m,n}(\theta, \bar{\theta}) \star h(\theta, \bar{\theta}) = e_n P_{m,n}(\theta, \bar{\theta}) $$

(68)

$$ \theta \star P_{m,n}(\theta, \bar{\theta}) = P_{m-1,n}(\theta, \bar{\theta})(1 - \delta_{m,0}) $$

(69)

$$ P_{m,n}(\theta, \bar{\theta}) \star \theta = P_{m,n+1}(\theta, \bar{\theta})(1 - \delta_{n,1}) $$

(70)

$$ \bar{\theta} \star P_{m,n}(\theta, \bar{\theta}) = P_{m+1,n}(\theta, \bar{\theta})(1 - \delta_{m,1}) $$

(71)
\[ P_{m,n}(\theta, \bar{\theta}) \star \bar{\theta} = P_{m,n-1}(\theta, \bar{\theta})(1 - \delta_{n,0}) \]  

(72)

It is interesting to note that the functions \( P_{m,n}(\theta, \bar{\theta}) \) are related to the generating function for the fermionic oscillator:

\[ \Phi(\alpha, \beta, \theta, \bar{\theta}) = \sum_{mn} \alpha^n \beta^m \langle n | \exp(\theta f^+ + \bar{\theta} f^-) | m \rangle \]  

(73)

where the integers \( m \) and \( n \) take the values 0, 1 and \( \alpha \) and \( \beta \) are Grassmann variables.

Computing the matrix elements occurring in the last equation, we have

\[ \Phi(\alpha, \beta, \theta, \bar{\theta}) = \exp(-\frac{1}{2}(\bar{\theta} \theta + \alpha \theta + \beta \bar{\theta} + \alpha \beta)) \]  

(74)

One can verify

\[ \frac{\partial^n}{\partial \alpha^n} \frac{\partial^m}{\partial \alpha^m} \Phi(\alpha, \beta, \theta, \bar{\theta}) \big|_{\alpha=\beta=0} = P_{m,n}(\theta, \bar{\theta}) \star \bar{\theta} \]  

(75)

trading the relation between the generating function and the functions associated to elements \( P_{m,n} \).

The Grassmann star product, defined here by means of fermionic coherent states, is equivalent to the following one

\[ \hat{\star} = e^{\frac{1}{2}\left(\frac{\partial}{\partial \theta} \frac{\partial}{\partial \bar{\theta}} + \frac{\partial}{\partial \bar{\theta}} \frac{\partial}{\partial \theta}\right)} \]  

(76)

which seems to be the Grassmann version of Moyal product in the noncommutative Grassmann space. The equivalence relation is

\[ T(A(\bar{\theta}, \bar{\theta})) \star T(B(\bar{\theta}, \bar{\theta})) = T(A(\bar{\theta}, \bar{\theta}) \hat{\star} B(\bar{\theta}, \bar{\theta})) \]  

(77)

where the relevant operator is

\[ T = e^{\frac{i}{2} \frac{\partial}{\partial \theta} \frac{\partial}{\partial \bar{\theta}}} = 1 - \frac{1}{2} \frac{\partial}{\partial \bar{\theta}} \frac{\partial}{\partial \theta} \]  

(78)

The needed relations to compute the \( \hat{\star} \) product of two arbitrary functions are

\[ \theta \hat{\star} \bar{\theta} = \theta \bar{\theta} + \frac{1}{2} \]  

\[ \bar{\theta} \hat{\star} \theta = \bar{\theta} \theta + \frac{1}{2} \]  

(79)

from which one has

\[ \{ \theta, \bar{\theta} \} \hat{\star} = 1 \]  

\[ \theta \hat{\star} \theta = \bar{\theta} \hat{\star} \bar{\theta} = 0 \]  

(80)

reflecting the correspondence with anti-commutation relations of the fermionic operators \( f^+ \) and \( f^- \). Note that one can obtain also a star analogue of the \( su(2) \) algebra using the noncommutative Grassmann variables. Indeed, the functions

\[ j_+ = P_{1,0}(\theta, \bar{\theta}) = \bar{\theta} \]  

\[ j_- = P_{0,1}(\theta, \bar{\theta}) = \theta \]  

\[ j_3 = P_{1,1}(\theta, \bar{\theta}) = \bar{\theta} \theta \]  

(81)
satisfy the relations
\[
[j_-, j_+] = 2j_3, \quad [j_\pm, j_3] = \pm j_\pm
\] (82)

The Casimir operator is defined by
\[
C = \frac{1}{2}(j_- j_+ + j_+ j_-) + j_3^2
\] (83)

By a direct computation, one obtain \(C = \frac{1}{4}\).

Now, we consider the bosonic oscillator described by the Hamiltonian
\[
H_b = b^+ b^-
\] (84)

where \(b^+\) and \(b^\) are the standard creation and annihilation operators. Using the coherent states of the harmonic oscillator
\[
|z\rangle = e^{-\frac{1}{2}z\bar{z}} \sum_{n=0}^{\infty} \frac{z^n}{\sqrt{n!}} |n\rangle
\] (85)

The star product of two functions \(A(z, \bar{z}) = \langle z | A | z \rangle\) and \(B(z, \bar{z}) = \langle z | B | z \rangle\) (\(A\) and \(B\) operators acting on \(H_b = \{ |n\rangle; n \in \mathbb{N} \}\)) is defined by [10]
\[
\bar{z} = e^{\frac{1}{2}z\bar{z}} \prod_{\alpha} \frac{\partial}{\partial z_{\alpha}} \frac{\partial}{\partial \bar{z}_{\alpha}}
\] (86)

This star product is the so-called Voros Product. Every operator \(A\) can be expanded in the occupation number operator basis \(\{ |n\rangle\langle m| \}\) and for which we associate the functions
\[
\Phi_{mn}(z, \bar{z}) = \langle z | n\rangle \langle m | z \rangle = e^{-\frac{1}{2}z\bar{z}} \frac{\bar{z}^m}{\sqrt{m!}} \frac{z^n}{\sqrt{n!}}
\] (87)

satisfying the completion relation \(\sum_{m=0}^{\infty} \Phi_{mn}(z, \bar{z}) = 1\) and the orthogonality property
\[
\Phi_{m,n}(z, \bar{z}) \bar{z} \Phi_{m',n'}(z, \bar{z}) = \delta_{m,n} \Phi_{m,m'}(z, \bar{z}) \delta_{m,n'}
\] (88)

The \(\Phi_{0,0}(z, \bar{z})\) provides a star vacuum
\[
z\bar{z} \Phi_{0,0}(z, \bar{z}) = \Phi_{0,0}(z, \bar{z}) \bar{z} \bar{z} = 0
\] (89)

We have also
\[
z\bar{z} \Phi_{n,m}(z, \bar{z}) = \sqrt{n} \Phi_{n-1,m}(z, \bar{z})
\] (90)
\[
\bar{z} \Phi_{n,m}(z, \bar{z}) = \sqrt{n + 1} \Phi_{n+1,m}(z, \bar{z})
\] (91)
\[
\Phi_{n,m}(z, \bar{z}) \bar{z} \bar{z} = \sqrt{m + 1} \Phi_{n,m+1}(z, \bar{z})
\] (92)
\[
\Phi_{n,m}(z, \bar{z}) \bar{z} \bar{z} = \sqrt{m} \Phi_{n,m-1}(z, \bar{z}).
\] (93)

The function \(\mathcal{H}_{cla}\) associated to the Hamiltonian \(H_b\) is
\[
\mathcal{H}_{cla} = \langle z | H | z \rangle = z\bar{z}
\] (94)
known, in the coherent states language, as the identity action. One has also

\[ \mathcal{H}_{\text{cla}} \Phi_{n,m}(z, \bar{z}) = m \Phi_{n,m}(z, \bar{z}) \]  
(95)

\[ \Phi_{n,m}(z, \bar{z}) \mathcal{H}_{\text{cla}} = n \Phi_{n,m}(z, \bar{z}) \]  
(96)

At this stage, we have the main tools to introduce the star analogue of the superoscillator. The supersymmetric oscillator is characterised by the Hamiltonian

\[ H_{\text{susy}} = b^+ b^- + f^+ f^- \]  
(97)

The supersymmetric character appears from the fact that the Hamiltonian \( H_{\text{susy}} \) can be written as the anti-commutator \( H_{\text{susy}} = \{Q^-, Q^+\} \) of the conserved supercharges \( Q^- \) and \( Q^+ \)

\[ Q^- = b^+ f^- \quad Q^+ = b^- f^+ \]  
(98)

satisfying the nilpotency conditions \((Q^-)^2 = (Q^+)^2 = 0\) and commuting with the supersymmetric Hamiltonian. The operators \( Q^+ \), \( Q^- \) and \( H \) close the \( Z_2 \) graded algebra \( \mathfrak{sl}(1/1) \). Since we are concerned with the superstar product in the supercoherent states framework, we consider the \( Z_2 \)-graded Fock space

\[ \mathcal{H} = \mathcal{H}_b \otimes \mathcal{H}_f = \{ |n, s\rangle; n \in \mathbb{N}, s = 0, 1 \} \]  
(99)

Using the group theoretical approach, based on the Weyl-Heisenberg supergroup \( W \), the action of a unitary representation of \( W \) leads to the supercoherent states \([16-17]\)

\[ |z, \theta\rangle = (1 - \frac{1}{2} \theta \theta)(|z, 0\rangle + \theta |z, 1\rangle) = |z\rangle \otimes |\theta\rangle \]  
(100)

where the bosonic and fermionic variables are commuting. Following the method, presented in the second section leading to the Grassmann star product, the superstar product is defined by means of the supercoherent states (98) as

\[ *_{\text{susy}} = e^\left( \frac{\partial}{\partial z} \frac{\partial}{\partial \bar{z}} + \frac{\partial}{\partial \theta} \frac{\partial}{\partial \bar{\theta}} \right) \]  
(101)

The functions associated to \( H_{\text{susy}} \) and supercharges \( Q^+, Q^- \) are given by

\[ h_{\text{susy}} = \langle z, \theta | H_{\text{susy}} | z, \theta \rangle = \bar{z} z + \bar{\theta} \theta \]  
(102)

\[ q^- = \langle z, \theta | Q^- | z, \theta \rangle = z \bar{\theta} \]  
(103)

\[ q^+ = \langle z, \theta | Q^+ | z, \theta \rangle = \bar{z} \theta \]  
(104)

Moreover, it is easy to verify the following relations

\[ q^- *_{\text{susy}} q^+ + q^+ *_{\text{susy}} q^- = h_{\text{susy}} \]  
(105)

\[ q^\pm *_{\text{susy}} h_{\text{susy}} = h_{\text{susy}} *_{\text{susy}} q^\pm \]  
(106)

and

\[ q^+ *_{\text{susy}} q^+ = q^- *_{\text{susy}} q^- = 0 \]  
(107)

which are similar to ones defined the \( \mathfrak{sl}(1/1) \) superalgebra where the ordinary (anti)-

commutations are replaced by * ones.
6 Concluding remarks

To close this note, recall that the works [10-13] related to star products and Moyal brackets, formulated in the coherent states framework, are exclusively of bosonic type. In the best of our knowledge, the Grassmannian case involving fermionic coherent states has not been considered elsewhere. This is most probably due to the deficiency of a widely accepted convincing description of classical systems by means of Grassmann algebras. An additional reason may arise from the unconventional techniques for integration and differentiation of Grassmann variables. In this work, we tried to fill this gap and proposed a non anti-commutative Grassmann star product. The anti-commutative limit \( \hbar \to 0 \) (classical limit in some sense) is easily recovered by a simple rescaling of the variables \( \theta \to \frac{1}{\sqrt{\hbar}} \theta \). It has been shown that each unitary transformation admits a unique representation in the noncommutative space spanned by Grassmann variables and the covariance holds only and only if the transformation is canonical. The supersymmetric extension of the Voros star product is also considered thanks to the supercoherent states. Finally, we believe that the results of this work can be used to define star or super star products linked to Lie and super Lie algebras, as for instance \( su(p,q) \) and \( Osp(p/q) \), and we hope to report on this subject in a future paper.

Acknowledgements: The author would like to thank the Condensed matter Section of the Abdus Salam-ICTP for hospitality.
References

[1] J. Moyal, Proc. Camb. Phil. Soc. 45 (1949) 99.
[2] H. Weyl, Z. Phys. 46 (1927) 1.
[3] C. S. Chu and P. M. Ho, Nucl. Phys. B 550 (1999) 151.
[4] A. Connes, M. R. Douglas and A. Shwarz, J. High. Energy. Phys. 02 (1998) 003.
[5] B. Jurco and P. Schupp, Eur. J. Phys. C 14 (2000) 367.
[6] T. Askawa and I. Kishimoto, hep-th/0002138.
[7] T. Curtright, D. Fairlie and C. Zachos, hep-th/9711183; T. Curtright and C. Zachos, hep-th/9903254; T. Curtright, T. Uematsu and C. Zachos, J. Math. Phys. 42 (2001) 2396; D. Fairlie, Mod. Phys. Lett A 13 (1998) 263.
[8] A. Perelomov, Generalized Coherent states and their applications, (Springer, Berlin, 1985).
[9] J. R. Klauder and B. S. Skagerstam, Coherent states and their applications, Texts and Monographs in Physics, (World Scientific, Singapore, 1985).
[10] G. Alexanian, A. Pinzul and A. Stern, Nucl. Phys. B 600 (2001) 531.
[11] A. Pinzul and A. Stern, J. High. Energy. Phys. 0203 (2002) 039.
[12] M. Daoud and E. H. El Kinani, J. Phys. A 35 (2002) 2639.
[13] A. B. Hammou, M. Lagraa and M. M. Sheikh-Jabbari, Phys. Rev. D 66 (2002) 0250025.
[14] B. De Witt, Supermanifolds (Cambridge: Cambridge University Press, 1992).
[15] G. Junker and J. Klauder, Eur. Phys. J. C 4 (1998) 173.
[16] B. W. Fatyga, V. A. Kostelecky, M. M. Nieto and D. R. Traux, Phys. Rev. D 43 (1991) 1403.
[17] Y. Bérubé-Lauzière and V. Hussin, J. Phys. A 26 (1993) 6271.