Some properties of the holographic fermions in an extremal charged dilatonic black hole

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Abstract

We study the properties of the Greens functions of the fermions in the extremal charged dilatonic black hole proposed by Gubser and Rocha [Phys. Rev. D 81, 046001 (2010)]. We find that many properties seem to be in agreement with that of Fermi liquid. Especially, the dispersion relation is linear. It is very different from that found in the extremal Reissner-Nordström (RN) black hole for massless fermion, which is obviously non-Fermi liquid due to a nonlinear dispersion relation. However, for another scaling behavior of the height of $\text{Im}G_{11}$ at the maximum, the scaling exponent is not one, which is at odds with that found in the RN black hole.

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I. INTRODUCTION

Motivated by the discovery of many exotic but important numerous correlated electron materials, including the cuprate superconductors and other oxides, lots of theoretical exploration on the quantum phases and phase transitions of correlated electron systems are intensively processing and varieties of phenomenological models were proposed in condensed matter theory. However, a general theoretical framework characterizing such exotic states of matter remains a suspense. Recently, many theoretical physicists have resorted to the AdS/CFT correspondence [1] to offer a possible clue to yield the basic principle of such states of matter. Indeed, we have found many of such exotic states of matter [2, 3] using AdS/CFT correspondence. At the same time, a proper condensed matter interpretation of these exotic states has also been explored in recent studies [4, 5]. Especially, in Ref. [4], they have proposed that the holographic non-Fermi liquid model in Ref. [2, 6] had realized an infinite-range limit of the fractionalized Fermi liquid [7]. Here, we must note that the two models share the same characteristics, that is, nonzero ground state entropy density.

However, nonzero ground state entropy density in Reissner-Nordström black hole in Anti-deSitter spacetimes (RNAdS) seems to be inconsistent with the characteristics of the degenerate Fermi liquid at the zero temperature. Therefore, a systematic exploration of the system in which ground state entropy density vanishes will be important and valuable. Such models have been proposed in Refs. [8, 9]. In this paper, we will only focus on the model of Gubser and Rocha [8]. In this model, they construct a charged dilatonic black hole in $AdS_5$ in which the Lagrangian involves a gauge field and a neutral scalar. In addition to the vanishing ground state entropy density mentioned above, there is another important characteristic, that is, linear specific heat at low temperature. It is also the characteristics of a Fermi gas. Therefore, it is important and interesting to explore the properties of fermion response in this background.

Our paper is organized as follows. In Sec. II, a brief review of the model of Gubser and Rocha is given. Following Ref. [2], we obtain the Dirac equation of the probe fermions in this

\[ 1 \text{ We must point out that in Ref. [8], the authors make the stability analysis of their supergravity solution and they find an instability. Therefore, the solution in the model of Gubser and Rocha is not the true ground state. They presumed that it would involve some sort of charged scalar condensate or perhaps something completely different.} \]
model in Sec. III. Some properties are presented in Sec. IV. Conclusions and discussions follow in Sec. VI.

II. A CHARGED DILATONIC BLACK HOLE GEOMETRY

In this section, a brief review of the model of Gubser and Rocha is given as follows. We will discuss the $\text{AdS}_5$ case and the $\text{AdS}_4$ case, respectively.

A. $\text{AdS}_5$ case

We begin with the following action:

$$
S = \frac{1}{2\kappa^2} \int d^5x \sqrt{-g} \left[ R - 12(\nabla_a \alpha)^2 + \frac{1}{L^2} (8e^{2\alpha} + 4e^{-4\alpha}) - \frac{1}{4} e^{4\alpha} F^{ab} F_{ab} \right],
$$

(1)

where $R$ is the Ricci scalar, and $L$ is the AdS radius. As is commonly known, $F_{ab} = \partial_a A_b - \partial_b A_a$. The charged solution is

$$
ds^2 = e^{2A} (-h dt^2 + dx^2) + \frac{e^{2B}}{h} dr^2, \quad A_a dx^a = \Phi dt,
$$

$$
A = \ln \frac{r}{L} + \frac{1}{3} \ln \left(1 + \frac{Q^2}{r^2}\right), \quad B = -\ln \frac{r}{L} - \frac{2}{3} \ln \left(1 + \frac{Q^2}{r^2} \right),
$$

$$
h = 1 - \frac{\mu L^2}{(r^2 + Q^2)^2}, \quad \Phi = \frac{Q\sqrt{2\mu}}{r^2 + Q^2} - \frac{Q\sqrt{2\mu}}{r^2 + Q^2}, \quad \alpha = \frac{1}{6} \ln \left(1 + \frac{Q^2}{r^2} \right).
$$

(2)

The rescaled energy density, entropy density, and charge density can be easily expressed as follows in the microcanonical ensemble:

$$
\hat{\epsilon} = \frac{\kappa^2}{4\pi^2 L^3} \epsilon = \frac{3\mu}{8\pi^2 L^6}, \quad \hat{s} = \frac{\kappa^2}{4\pi^2 L^3} s = \frac{\sqrt{\mu} r_+}{2\pi L^5}, \quad \hat{\rho} = \frac{\kappa^2}{4\pi^2 L^3} \rho = \frac{Q\sqrt{2\mu}}{4\pi^2 L^5}.
$$

(3)

With the help of Eq. (3) and the condition $h(r_+) = 0$, one can obtain the microcanonical equation of state:

$$
\hat{\epsilon} = \frac{3}{2^{5/3}\pi^{2/3}}(\hat{s}^2 + 2\pi^2 \hat{\rho}^2)^{2/3}.
$$

(4)

Therefore, the temperature and chemical potential can be obtained as follows:

$$
T = \left(\frac{\partial \hat{\epsilon}}{\partial \hat{s}}\right)_{\hat{\rho}} = \frac{r_+}{\pi L^2}, \quad \Omega = \left(\frac{\partial \hat{\epsilon}}{\partial \hat{\rho}}\right)_{\hat{s}} = \frac{\sqrt{2Q}}{L^2}.
$$

(5)

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2 This black hole solution can be embedded in string theory. For more discussions, we can refer to Ref.[8].
Then we can easily find
\[ \hat{s} = \pi \sqrt{\frac{2\hat{\epsilon}}{3}} T \approx (\pi^2 \hat{\rho})^{2/3} T \approx \frac{\Omega^2}{4} T. \] (6)
where the approximate equalities hold in the low-temperature limit. In this paper, we will mainly focus on the zero-temperature limit. From the expression of the temperature, we find that the zero-temperature limit is obtained when \( r_+ = 0 \), which implies \( \mu L^2 = Q^4 \). At the same time, we also notice that in this limit, the entropy vanishes, which can have an important effect on the ground state.

In addition, from Eq. (6), we find that the rescaled specific heats at constant charge density and constant chemical potential
\[ \hat{C}_\rho = T \left( \frac{\partial \hat{s}}{\partial T} \right)_\rho, \quad \hat{C}_\Omega = T \left( \frac{\partial \hat{s}}{\partial T} \right)_\Omega, \] (7)
are linear, which similar the case of a Fermi gas.

From the above discussions, we can see that there exists the similarity between the low-temperature thermodynamics of the charged dilatonic black hole and the Fermi liquid. Furthermore, in Ref. [8], they find the Fermi surface for massless, charged bulk fermions at extremal limit by finding the normal modes. It is plausible to claim the dual of the fermions of the charged dilatonic black hole is a Fermi liquid. Therefore, in this paper, we will furthermore explore the characteristics of the fermions response in this background.

B. \textit{AdS}_4 case

The above case in \textit{AdS}_5 can also be extended to the case in \textit{AdS}_4 [8,10]:
\[ S = \frac{1}{2\kappa^2} \int d^4x \sqrt{-g} \left[ R - \frac{3}{2} (\nabla_{\alpha} \alpha)^2 + \frac{6}{L^2} \cosh \alpha - \frac{1}{4} e^\alpha F^{ab} F_{ab} \right], \] (8)
From the above action, one can obtain the charged solution
\[ ds^2 = e^{2A} (-h dt^2 + d\vec{x}^2) + \frac{e^{2B}}{h} dr^2, \quad A_a dx^a = \Phi dt, \]
\[ A = \ln \frac{r}{L} + \frac{3}{4} \ln \left( 1 + \frac{Q}{r} \right), \quad B = -A, \quad h = 1 - \frac{\mu L^2}{(r + Q)^3}, \]
\[ \Phi = \frac{\sqrt{3Q \mu}}{r + Q} - \frac{\sqrt{3Q \mu \pi}}{L^2}, \quad \alpha = \frac{1}{2} \ln \left( 1 + \frac{Q}{r} \right). \] (9)
The extremal limit for \textit{AdS}_4 is also obtained when \( r_+ = 0 \) (corresponding to \( \mu L^2 = Q^3 \)). The case in \textit{AdS}_4 shares the same characteristic of low-temperature thermodynamics and has the Fermi surface.
III. DIRAC EQUATION

In this paper, we explore the fermion response in the charged dilatonic black hole in $AdS_5$ and $AdS_4$, respectively. First, we will briefly derive the Dirac equation in this background. Subsequently, we discuss how to impose the boundary condition at the horizon and read off the Green’s function on the boundary. For more details, we can also refer to Refs. 2, 6.

A. The Dirac equation in a dilatonic black hole

Considering the following bulk fermion action\(^3\)

\[
S_D = i \int d^{d+1}x \sqrt{-g} \left( \Gamma^a D_a - m \right) \zeta, \tag{10}
\]

where $\Gamma^a$ is related to the usual flat space gamma matrix by a factor of the vielbein, $\Gamma^a = (e_\mu)^a \Gamma^\mu$ and $D_a = \partial_a + \frac{1}{4} (\omega_{\mu\nu})_a \Gamma^{\mu\nu} - iq A_a$ is the covariant derivative with $(\omega_{\mu\nu})_a$ the spin connection 1-forms. The Dirac equation derived from the action $S_D$ is expressed as

\[
\Gamma^a D_a \zeta - m \zeta = 0. \tag{11}
\]

Making a transformation $\zeta = (-g g^{rr})^{-\frac{1}{2}} \mathcal{F}$ to remove the spin connection and expanding $\mathcal{F}$ as $\mathcal{F} = F e^{-i \omega t + i k_x x^1}$ in Fourier space, the Dirac equation (11) turns out to be

\[
\sqrt{g^{rr}} \Gamma^r \partial_r F - i(\omega + q A_t) \sqrt{g^{tt}} \Gamma^t F + i k \sqrt{g^{xx}} \Gamma^x F - m F = 0. \tag{12}
\]

where due to rotational symmetry in the spatial directions, we set $k_x = k$ and $k_x \neq 0, i \neq x$ without losing generality. Notice that Eq. (12) only depends on three Gamma matrices $\Gamma^r, \Gamma^t, \Gamma^x$. So it is convenient to split the spinors $F$ into $F = (F_1, F_2)^T$ and choose the following basis for our gamma matrices as in 6:

\[
\Gamma^r = \begin{pmatrix} -\sigma^3 & 0 \\ 0 & -\sigma^3 \end{pmatrix}, \quad \Gamma^t = \begin{pmatrix} i \sigma^1 & 0 \\ 0 & i \sigma^1 \end{pmatrix}, \quad \Gamma^x = \begin{pmatrix} -\sigma^2 & 0 \\ 0 & \sigma^2 \end{pmatrix}, \ldots \tag{13}
\]

where $1$ is an identity matrix of size $2^{\frac{d-3}{2}}$ for $d$ odd (or size $2^{\frac{d-4}{2}}$ for $d$ even). Because the resulting Green’s functions are proportional to such an identity matrix, we will suppress

\(^3\) Although the background spacetime 2 or 9 comes from a UV-complete string theory, the fermion action 10 is ad hoc so that it is not necessary to derive it from type IIB supergravity.
them in the following. So, we have a new version of the Dirac equation as
\[
\sqrt{g^{rr}} \partial_r \begin{pmatrix} F_1 \\ F_2 \end{pmatrix} + m \sigma^3 \otimes \begin{pmatrix} F_1 \\ F_2 \end{pmatrix} = \sqrt{g^{tt}}(\omega + qA_t)\sigma^2 \otimes \begin{pmatrix} F_1 \\ F_2 \end{pmatrix} \mp k\sqrt{g^{xx}}\sigma^1 \otimes \begin{pmatrix} F_1 \\ F_2 \end{pmatrix} .
\] (14)

Furthermore, according to eigenvalues of \( \Gamma^r \), we make such a decomposition \( F_\pm = \frac{1}{2}(1 \pm \Gamma^r)F \). Then
\[
F_+ = \begin{pmatrix} B_1 \\ B_2 \end{pmatrix}, \quad F_- = \begin{pmatrix} A_1 \\ A_2 \end{pmatrix}, \quad \text{with} \quad F_\alpha = \begin{pmatrix} A_\alpha \\ B_\alpha \end{pmatrix}.
\] (15)

Under such decomposition, the Dirac equation (14) can be rewritten as
\[
(\sqrt{g^{rr}} \partial_r \pm m) \begin{pmatrix} A_1 \\ B_1 \end{pmatrix} = \pm(\omega + qA_t)\sqrt{g^{tt}} \begin{pmatrix} B_1 \\ A_1 \end{pmatrix} - k\sqrt{g^{xx}} \begin{pmatrix} B_1 \\ A_1 \end{pmatrix},
\] (16)
\[
(\sqrt{g^{rr}} \partial_r \pm m) \begin{pmatrix} A_2 \\ B_2 \end{pmatrix} = \pm(\omega + qA_t)\sqrt{g^{tt}} \begin{pmatrix} B_2 \\ A_2 \end{pmatrix} + k\sqrt{g^{xx}} \begin{pmatrix} B_2 \\ A_2 \end{pmatrix}.
\] (17)

Introducing the ratios \( \xi_\alpha \equiv \frac{A_\alpha}{B_\alpha}, \alpha = 1, 2 \), and using the method developed in [6, 11, 12], one can package the Dirac equation (16) and (17) into the evolution equation of \( \xi_\alpha \),
\[
(\sqrt{g^{rr}} \partial_r + 2m)\xi_\alpha = \left[ \sqrt{g^{tt}}(\omega + qA_t) + (-1)^\alpha k\sqrt{g^{xx}} \right] \xi_\alpha^2,
\] (18)
which will be more convenient to impose the boundary conditions at the horizon and read off the boundary Green functions.

**B. The boundary condition and the Green’s function**

Near the boundary, a solution of the Dirac equation (14) can be expressed as
\[
F_\alpha \rightrightarrows \infty a_\alpha r^m \begin{pmatrix} 0 \\ 1 \end{pmatrix} + b_\alpha r^{-m} \begin{pmatrix} 1 \\ 0 \end{pmatrix}, \quad \alpha = 1, 2 .
\] (19)

If \( b_\alpha \begin{pmatrix} 1 \\ 0 \end{pmatrix} \) and \( a_\alpha \begin{pmatrix} 0 \\ 1 \end{pmatrix} \) are related by \( b_\alpha \begin{pmatrix} 1 \\ 0 \end{pmatrix} = S a_\alpha \begin{pmatrix} 0 \\ 1 \end{pmatrix} \), then the boundary Green’s functions \( G \) is given by \( G = -iS \gamma^0 \). Therefore
\[
G(\omega, k) = \lim_{r \to \infty} r^{2m} \begin{pmatrix} \xi_1 & 0 \\ 0 & \xi_2 \end{pmatrix},
\] (20)
FIG. 1: The plot of $\text{Im}G_{11}(k)$ for $\omega = -10^{-9}$ (left plot for $AdS_5$ and right plot for $AdS_4$). A sharp peak locates at $k_F \approx 2.12$ for $AdS_5$ and $k_F \approx 2.19$ for $AdS_4$. Improving the accuracy, the Fermi momentum can be furthermore determined as $k_F \approx 2.121305342$ for $AdS_5$ and $k_F \approx 2.19528236825$ for $AdS_4$.

At the same time, the requirement that the solutions of Eqs. (16) and (17) at the horizon be in-falling implies

$$\xi_\alpha r \to 0 = i, \quad \text{for} \quad \omega \neq 0. \quad (21)$$

Finally, we also note that $G_{11}$ and $G_{22}$ are related to each other as $G_{22}(\omega, k) = G_{22}(\omega, -k)$. Therefore, we will focus solely on $G_{11}$ below.

IV. THE PROPERTIES OF THE GREEN’S FUNCTION

In this section, we solve Eq. (18) numerically with the boundary conditions (21) to investigate the properties of the spectral function. In Ref. [8], they find the Fermi surface by exploring the normal modes at $\omega = 0$. Here, we will follow the method developed in Ref. [2] to search for the Fermi surface. We will consider the specific example as [8]: $L = 1$, $m = 0$, $Q = 1$, $q = 2$.

For $AdS_5$, when $\omega \to 0$, a sharp quasiparticlelike peak occurs near $k_F = 2.12130534$ (left plot in Fig. (1)). While for $AdS_4$, such a peak located at $k_F \approx 2.195282368$ (right plot in Fig. (1)). It is consistent with the results found in Ref. [8].

Subsequently, we will explore the characteristics of the Green’s function. We find that:

1. At the Fermi momentum $k = k_F$, the quasiparticle peak is almost a delta function.
FIG. 2: The plot of $\text{Im}G_{11}(k)$ for various $k$ (left plot for $AdS_5$ and right plot for $AdS_4$). They show that the quasiparticle peak approach a delta function at the Fermi momentum $k = k_F$. Left (right) plot: green for $k = 4$ ($k = 4$), black for $k = 3$ ($k = 3$), blue for $k = 2.4$ ($k = 2.5$), red for $k = 2.12$ ($k = 2.195$), orange for $k = 1.8$ ($k = 1.9$) and magenta for $k = 0.8$ ($k = 0.9$).

From Fig. 2, we find that the peak becomes narrower and narrower when $k$ ($k > k_F$) decreases. Finally, their heights approach infinity when $k$ approaches $k_F$. However, when the Fermi surface is crossed, the peak become wide again.

2. The Green’s function vanishes at the Fermi energy (corresponding to $\omega = 0$ in our convention), which has nothing to do with $k$. It is shown in Fig. 3.

3. There exists a linear dispersion relation between small $\tilde{k}$ and $\tilde{\omega}(\tilde{k})$ (Fig. 4), i.e., $\tilde{\omega}(\tilde{k}) \sim \tilde{k}$, where $\tilde{k} \equiv k - k_F$ and $\tilde{\omega}(\tilde{k})$ is the location of the maximum of the peak. It just corresponds to the form of a Fermi liquid.

4. Another important scaling behavior of the height of $\text{Im}G_{11}$ at the maximum is as follows:

$$\text{Im}G_{11}(\tilde{\omega}, \tilde{k}) \sim \tilde{k}^{-\beta},$$  \hspace{1cm} (22)

where $\beta \simeq 3.8$ for $AdS_5$ and $\beta \simeq 5$ for $AdS_4$ (Fig. 5).4

The first three characteristics above, especially the third, the linear dispersion relation, are like that of Fermi liquid. It is very different from that found in the extremal RN

4 Here, we must point out that due to the numerical instability, the momentum can not approach the fermi momentum so close in pinning down the height of $\text{Im}G_{11}$ as pinning down $\tilde{\omega}$. 
FIG. 3: The plot of $\text{Im}G_{11}(k)$ for various $k$ (left plot for $AdS_5$ and right plot for $AdS_4$). They show that independent of $k$, the Green’s function vanishes at the Fermi energy.

black hole for massless fermions, which is obviously non-Fermi liquid due to the nonlinear dispersion relation [2]. However, another important scaling behavior of the height of $\text{Im}G_{11}$ at the maximum is obviously at odds with that found in Ref. [2], in which the scaling exponent $\beta \simeq 1$. It is a peculiar property of Green’s functions of the fermions in this system. Therefore, it is interesting and important to furthermore explore the properties of holographic fermions in the extremal charged dilatonic black hole.
V. CONCLUSIONS AND DISCUSSION

We have studied the main features of the fermions in the charged dilatonic black hole for zero-temperature limit and massless fermions by AdS/CFT correspondence. Many features of the fermions’ Green’s function, especially the linear dispersion relation, and the low-temperature thermodynamics studied in Ref. [8] are plausible that the dual is Fermi liquid. However, for another scaling behavior of the height of $\text{Im}G_{11}(\tilde{\omega}, \tilde{k})$ at the maximum, the exponent $\beta$ is not one, indicating the case may be not so simple. Nevertheless, due to the peculiar properties of this model, it will be valuable to furthermore explore the characteristics of the fermions response in this background. In addition, it is also important to relate this model
to the state of condensed matter.

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