Local $\zeta$-functions, stress-energy tensor, field fluctuations, and all that, in curved static spacetime

Dedicated to Prof. Emilio Elizalde on the occasion of his 60th birthday

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Abstract: This is a quick review on some technology concerning the local zeta function applied to Quantum Field Theory in curved static (thermal) spacetime to regularize the stress energy tensor and the field fluctuations.

1 Quasifree QFT in curved static manifolds, Euclidean approach $\zeta$-function technique.

1.1 The $\zeta$-function determinant. Suppose we are given a $n \times n$ positive-definite Hermitian matrix $A$ with eigenvalues $0 < \lambda_1 \leq \lambda_2 \leq \cdots \leq \lambda_n$. One can define the complex-valued function

$$\zeta(s|A) = \sum_{j=1}^{n} \lambda_j^{-s}, \quad (1)$$

where $s \in \mathbb{C}$. (Notice that $\lambda_j^{-s}$ is well-defined since $\lambda_j > 0$.) By direct inspection one proves that:

$$\det A = e^{-\frac{d\zeta}{ds}(s|A)}|_{s=0}. \quad (2)$$

This trivial result can be generalized to provide a useful definition of the determinant of an operator working in an infinite-dimensional Hilbert space. To this end, focus on a non-negative self-adjoint operator $A$ whose spectrum is discrete and each eigenspace has a finite dimension, and consider the series with $s \in \mathbb{C}$ (the prime on the sum henceforth means that any possible null eigenvalues is omitted)

$$\zeta(s|A) := \sum_{j}^{'} \lambda^{-s}. \quad (3)$$

Looking at (2), the idea [Ha77] is to define, once again.

$$\det A = e^{-\frac{d\zeta}{ds}(s|A)}|_{s=0}. \quad (4)$$
where now the function $\zeta$ on the right-hand side is, in the general case, the analytic continuation of the function defined by the series (3) in its convergence domain, since the series may diverge at $s = 0$ – and this is the standard situation in the infinite-dimensional case! – provided that the analytic extension really reaches a neighborhood of the point $s = 0$. The interesting fact is that this procedure truly works in physically relevant cases, related to QFT in curved spacetime, producing meaningful results as we go to discuss in the following section.

1.2 Thermal QFT in static spacetimes. A smooth globally hyperbolic spacetime $(M, g)$ is said to be static if it admits a (local) time-like Killing vector field $\partial_t$ normal to a smooth spacelike Cauchy surface $\Sigma$. Consequently, there are (local) coordinate frames $(x^0, x^1, x^2, x^3) \equiv (t, \vec{x})$ where $g_{0i} = 0$ ($i = 1, 2, 3$) and $\partial_t g_{\mu \nu} = 0$ and $\vec{x}$ are local coordinates on $\Sigma$. Though the results we are going to present may be generalize to higher spin fields, we henceforth stick to the case of a real scalar field $\phi$ propagating in $M$ and satisfying an equation of motion of the form

$$P \phi = 0,$$

where $P := -\nabla_\mu \nabla^\mu + V$, $V$ being a smooth scalar field like

$$V(x) := \xi R + m^2 + V'(x).$$

We also assume that $V'$ satisfies $\partial_t V' = 0$ so that the space of solutions of (4) is invariant under $t$-displacements. Furthermore $\xi \in \mathbb{R}$, $R$ is the scalar curvature and $m^2$ the squared mass of the particles associated to the field. The domain of $P$ is the space of real-valued $C^\infty$ functions compactly supported Cauchy data on $\Sigma$. In the quasifree case, a straightforward way to define a QFT consists of the assignment of a suitable Green function of the operator $P$ [FR87], in particular the Feynman propagator $G_F(x, x')$ or, equivalently, the Wightman functions $W_{\pm}(x, x')$. Then the GNS theorem (e.g. see [KW91]) allows one to construct a corresponding Fock realization of the theory. In a globally hyperbolic static spacetime it is possible to choose $t$-invariant Green functions. In that case, in static coordinates, one performs the Wick rotation obtaining the Euclidean formulation of the same QFT. This means that (locally) one can pass from the Lorentzian manifold $(M, g)$ to a Riemannian manifold $(M_E, g^{(E)})$ by the analytic continuation $t \rightarrow i\tau$ where $t, \tau \in \mathbb{R}$. This defines a (local) Killing vector $\partial_\tau$ in the Riemannian manifold $M_E$ and a corresponding (local) “static” coordinate frame $(\tau, \vec{x})$ therein. As is well-known [FR87], when the orbits of the Euclidean time $\tau$ are closed with period $\beta$, $T = 1/\beta$ has to be interpreted as the temperature of the quantum state because the Wightman two-point function of the associated quasifree state satisfy the KMS condition at the inverse temperature $\beta$. In this approach, the Feynman propagator $G_F(t - t', \vec{x}, \vec{x}')$ determines – and (generally speaking [FR87]) it is completely determined by – a proper Green function (in the spectral theory sense) $S_\beta(\tau - \tau', \vec{x}, \vec{x}')$ of a corresponding self-adjoint extension $A$ of the operator

$$A' := -\nabla^{(E)}_a \nabla^{(E)a} + V(\vec{x}) : C^\infty_0(M_E) \rightarrow L^2(M_E, d\mu_{g^{(E)}}).$$

$S_\beta(\tau - \tau', \vec{x}, \vec{x}')$ is periodic with period $\beta$ in the $\tau - \tau'$ entry and it is said the Schwinger function. As a matter of fact, $S_\beta$ turns out to be the integral kernel of $A^{-1}$ when $A > 0$. 

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The partition function of the quantum state associated to $S_\beta$ is the functional integral evaluated over the field configurations periodic with period $\beta$ in the Euclidean time

$$Z_\beta = \int \mathcal{D}\phi \, e^{-S_E[\phi]},$$

(7)

the Euclidean action $S_E$ being

$$S_E[\phi] = \frac{1}{2} \int_M \mu_g(E) \, \phi(x)(A\phi)(x).$$

(8)

Thus, extending the analogous result for finite dimensional Gaussian integral, one has

$$Z_\beta = \left\{ \det \left( \frac{A}{\mu^2} \right) \right\}^{-1/2},$$

(9)

where $\mu$ is a mass scale which is necessary for dimensional reasons. To give a sensitive interpretation of that determinant, the idea [Ha77] is to try to exploit (2).

If $M_E$ is a $D$-dimensional Riemannian compact manifold and $A'$ is bounded below by some constant $b \geq 0$, $A'$ admits the Friedrichs self-adjoint extension $A$ which is also bounded below by the same bound of $A'$, moreover the spectrum of $A$ is discrete and each eigenspace has a finite dimension. Then, as we said, one can consider the series

$$\zeta(s | A/\mu^2) := \sum_j \left( \frac{\lambda_j}{\mu^2} \right)^{-s}.$$

(10)

Remarkably [Ha77, BCEMZ03], in the given hypotheses, the series above converges for $Re s > D/2$ and it is possible to continue the right-hand side above into a meromorphic function of $s$ which is regular at $s = 0$. Following (2) and taking the presence of $\mu$ into account, we define:

$$Z_\beta := e^{\frac{d}{2} \pi |s=0 \zeta(s | A/\mu^2)},$$

(11)

where the function $\zeta$ on the right-hand side is the analytic continuation of that defined in (10). It is possible the define the $\zeta$ function in terms of the heat kernel of the operator $A$, $K(t, x, y | A)$ [BCEMZ03]. This is the smooth integral kernel of the (Hilbert-Schmidt, trace-class) operators $e^{-tA}$, $t > 0$. One has, for $Re s > D/2$,

$$\zeta(s | A/\mu^2) = \int_M \mu_g(E) \, \int_0^{+\infty} dt \, \frac{\mu^{2ts-1}}{\Gamma(s)} \left[ K(t, x, x | A) - P_0(x, x | A) \right],$$

(12)

$P(x, y | A)$ is the integral kernel of the projector on the null-eigenvalues eigenspace of $A$. When $M_E$ is not compact, the spectrum of $A$ may included a continuous-spectrum part, however, it is still possible to generalize the definitions and the results above considering suitable integrals on the spectrum of $A$, provided $A$ is strictly positive (e.g., see [Wa79]).

Another very useful tool is the local $\zeta$ function that can be defined in two different, however equivalent, ways [Wa79, Mo98, BCEMZ03]:

$$\zeta(s, x | A/\mu^2) = \int_0^{+\infty} dt \, \frac{\mu^{2st-1}}{\Gamma(s)} \left[ K(t, x, x | A) - P_0(x, x | A) \right],$$

(13)
and, φₗ being the smooth eigenvector of the eigenvalue λₗ,

$$\zeta(s, x/A/\mu^2) = \sum_j \left( \frac{\lambda_j}{\mu^2} \right)^s \phi_j(x) \phi_j^*(x).$$  \hspace{1cm} (14)

Both the integral and the series converges for Re s > D/2 and the local zeta function enjoys the same analyticity properties of the integrated ζ function. For future convenience it is also useful to define, in the sense of the analytic continuation,

$$\zeta(s, x, y/A/\mu^2) = \int_0^{+\infty} dt \frac{\mu^{2s} s^{-1}}{\Gamma(s)} \left[ K(t, x, y/A) - P_0(x, y/A) \right]$$  \hspace{1cm} (15)

(see [Mo98, Mo99] for the properties of this off-diagonal ζ-function). In the framework of the ζ-function regularization framework, the effective Lagrangian is defined as

$$L(x/A/\mu^2) := \frac{1}{2} \frac{d}{ds}|_{s=0} \zeta(s, x/A/\mu^2),$$  \hspace{1cm} (16)

and thus, in a thermal theory, $Z_\beta = e^{-S_\beta}$ where $S_\beta = \int d\mu L_{\beta\mu^2}$. A result which generalizes to any dimension an earlier results by Wald [Wa79] is the following [Mo98]. The above-defined effective Lagrangian can be computed by a point-splitting procedure: For D even

$$L(y/A/\mu^2) = \lim_{x \to y} \left\{ - \int_0^{+\infty} dt \frac{\mu^{2s} s^{-1}}{\Gamma(s)} \ln \frac{\mu^2 \sigma(x, y)}{2} \right\}$$

$$+ \sum_{j=0}^{D/2-1} \left( \frac{D}{2} - j - 1 \right)! a_j(x, y/A) \frac{2}{2(4\pi)^{D/2}} \left( \frac{2}{\sigma(x, y)} \right)^{D/2-j}$$

$$- 2\gamma \frac{a_D(y, y/A)}{2(4\pi)^{D/2}},$$  \hspace{1cm} (17)

for D odd (notice that μ disappears)

$$L(y/A/\mu^2) = \lim_{x \to y} \left\{ - \int_0^{+\infty} dt \frac{\mu^{2s} s^{-1}}{\Gamma(s)} \right\}$$

$$+ \sum_{j=0}^{(D-3)/2} \left( \frac{D - 2j - 2}{2} \right)! a_j(x, y/A) \frac{2}{2(4\pi)^{D/2}} \left( \frac{2}{\sigma(x, y)} \right)^{D/2-j}.$$  \hspace{1cm} (18)

Above, σ(x, y) is one half the square of the geodesical distance of x from y and the coefficients $a_j$ are the well-known off-diagonal coefficients of the small-t expansion of the heat-kernel. These coefficients, in spite of their non symmetric definition, turns out to be invariant when interchanging x and y [Mo99, Mo00].

2 Stress-energy tensor and field fluctuations

2.1 Generalizations of the local ζ function technique. Physically relevant quantities are the (quantum) field fluctuation and the averaged (quantum) stress tensor, respectively:

$$<\phi^2(x)> = \frac{\delta}{\delta J(x)}|_{J=0} \ln \int D\phi e^{-S_E + \int d\mu_{\phi(E)} \phi^2 J},$$  \hspace{1cm} (19)
A standard method to compute them is the so-called point-splitting procedure [BD82, Fu91, Wa94, Mo00, Mo03]. It is however possible to extend the \( \zeta \)-function technique [Mo97, IM98, Mo98, Mo99, Mo00] to define suitable \( \zeta \) functions regularizing those quantities directly, similarly to what done for the effective Lagrangian. Consider the stress tensor. The idea relies upon the following chain of formal identities [Mo97]

\[
\sqrt{g(E)(x)} < T_{ab}(x) > = 2 \sqrt{g(E)(x)} \frac{\delta}{\delta g_{ab}(x)} \ln \int D\phi e^{-S_E[g(E)]}.
\]

(20)

Following this route, one define the \( \zeta \)-regularized (or renormalized) stress tensor as

\[
< T_{ab}(x|A) >_{\mu^2} := \frac{1}{2} \frac{d}{ds} |_{s=0} Z_{ab}(s,x|A/\mu^2),
\]

(22)

where, in the sense of the analytic continuation of the left-hand side

\[
Z_{ab}(s,x|A/\mu^2) := 2 \sum_j \mu^{-2s} \frac{\delta \lambda_j^{-s}}{\delta g_{ab}(x)}.
\]

(23)

The difficult problem is now twofold: how to compute the functional derivative in the right-hand side of (23) and whether or not the series in the right-hand side of (23) defines an analytic function of \( s \) in a neighborhood of \( s = 0 \). We have the result [Mo97, Mo99]:

**Theorem 1.** If \( M_E \) is compact, \( A \geq 0 \) and \( \mu^2 > 0 \), then \( Z_{ab}(s,x|A/\mu^2) \) is well-defined and is a \( C^\infty \) function of \( x \) which is also meromorphic in \( s \in \mathbb{C} \). In particular, it is analytic in a neighborhood of \( s = 0 \).

The result above has been checked even in several noncompact manifolds (containing singularities) [Mo97, BCEMZ03]. In that case, the summation in the right-hand side of (23) has to be replaced by a suitable spectral integration. The series in the right-hand side of (23) can be explicitly computed as [Mo97, Mo99, BCEMZ03]:

\[
s \sum_j \left\{ \frac{2}{\mu^2} \left( \frac{\lambda_j}{\mu^2} \right)^{-s-1} T_{ab}[\phi_j, \phi_j^*](x) + g_{ab}(x) \left( \frac{\lambda_j}{\mu^2} \right)^{-s} \right\},
\]

\( T_{ab}[\phi_j, \phi_j^*](x) \) being the classical stress tensor evaluated on the modes of \( A \) (see [Mo97, Mo99, BCEMZ03] for details). The series converges for \( \text{Re } s > 3D/2 + 2 \).

It is similarly possible to define a \( \zeta \)-function regularizing the field fluctuation [IM98, Mo98]:

\[
< \phi^2(x|A) >_{\mu^2} := \frac{d}{ds} |_{s=0} \Phi(s,x|A/\mu^2),
\]

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where
\[ \Phi(s, x|A/\mu^2) := \frac{s}{\mu^2} \zeta(s + 1, x|A/\mu^2). \] (24)

The properties of these functions have been studied in [IM98, Mo98] and several applications on concrete cases are considered (e.g. cosmic-string spacetime and homogeneous spacetimes). In particular, in [Mo98], the problem of the change of the parameter \( m^2 \) in the field fluctuations has been studied.

### 2.2 Physically meaningfulness of the procedures.

We are now interested in the physical meaningfulness of the presented regularization techniques. The following general results strongly suggest that it is the case [Mo97, Mo99, Mo03].

**Theorem 2.** If \( M_E \) is compact, \( A \geq 0 \) and \( \mu^2 > 0 \), and the averaged quantities above are those defined above in terms of local \( \zeta \)-function regularization, then

(a) \( <T_{ab}(x|A)>_{\mu^2} \) defines a \( C^\infty \) symmetric tensorial field.

(b) Similarly to the classical result,
\[
\nabla^b <T_{bc}(x|A)>_{\mu^2} = -\frac{1}{2} <\phi^2(x|A)>_{\mu^2} \nabla_c V'(x). \tag{25}
\]

(c) Concerning the trace of the stress tensor, it is naturally decomposed in the classical and the known quantum anomalous part
\[
g^{ab}(x) <T_{ab}(x|A)>_{\mu^2} = \left( \frac{\xi_D - \xi}{4\xi_D - 1} \Delta - m^2 - V'(x) \right) <\phi^2(x|A)>_{\mu^2} + \delta_D \frac{a_{D/2}(x,x|A)}{(4\pi)^{D/2}} - P_0(x,x|A), \tag{26}
\]

where \( \delta_D = 0 \) if \( D \) is odd and \( \delta_D = 1 \) if \( D \) is even, \( \xi_D = (D - 2)/[4(D - 1)] \).

(d) for any \( \alpha > 0 \)
\[ <T_{ab}(x|A)>_{\alpha \mu^2} = <T_{ab}(x|A)>_{\mu^2} + t_{ab}(x) \ln \alpha, \tag{27} \]
where, the form of \( t_{ab}(x) \) which depends on the geometry only and is in agreement with Wald’s axioms [Wa94], has been given in [Mo99, Mo03].

(e) In the case \( \partial_0 = \partial_\tau \) is a global Killing vector, the manifold admits periodicity \( \beta \) along the lines tangent to \( \partial_0 \) and \( M \) remains smooth (near any fixed points of the Killing orbits) fixing arbitrarily \( \beta \) in a neighborhood and, finally, \( \Sigma \) is a global surface everywhere normal to \( \partial_0 \), then
\[
\frac{\partial}{\partial \beta} \ln Z(\beta)_{\mu^2} = \int_\Sigma \frac{d^D\vec{x}}{\sqrt{\det g(\vec{x})}} <T_0^0(x, \beta|A)>_{\mu^2}. \tag{28}
\]

Another general achievement regards the possibility to recover the Lorentzian theory from the Euclidean one [Mo99]:
Theorem 3. Let $M_E$ be compact, $A \geq 0$, $\mu^2 > 0$. Also assume that $M_E$ is static with global Killing time $\partial \tau$ and (orthogonal) global spatial section $\Sigma$ and finally, $\partial \tau V' \equiv 0$. Then 

\( \partial \tau < \phi^2 (x|A) >_{\mu^2} \equiv 0; \)
\( \partial \tau L(x|A)_{\mu^2} \equiv 0; \)
\( \partial \tau < T_{ab} (x|A) >_{\mu^2} \equiv 0; \)
\( < T_{0i} (x|A) >_{\mu^2} \equiv 0 \) for $i = 1, 2, 3, ..., D - 1$

where the averaged quantities above are those defined above in terms of local $\zeta$-function regularization and coordinates $\tau \equiv x^0, \vec{x} \in \Sigma$ are employed.

These properties allow one to continue the Euclidean considered quantities into imaginary values of the coordinate $\tau \mapsto \text{it}$ obtaining real functions of the Lorentzian time $t$.

Some of the properties above (regarding Thm.1, Thm. 2, Thm.3) have been found to be valid in some noncompact manifolds too (Rindler spacetime, cosmic string spacetime, Einstein’s open spacetime, $H^N$ spaces, Gödel spacetime, BTZ spacetime) [Mo97, IM98, Ca98, Ra98, Ra98b, Ra99, Ra05, BMVZ98, RF02, SS04, AMR05]. In particular, the presented theory has been successfully exploited to compute the quantum back reaction on the three-dimensional BTZ metric [BMVZ98] in the case of the singular ground state containing a naked singularity. A semiclassical implementation of the cosmic censorship conjecture has been found in that case.

2.3. Interplay of zeta-function approach and point-splitting technique. The procedure of the point-splitting to renormalize the field fluctuation as well as the stress tensor [BD82, Wa94, Mo00, Mo03], when the two-point functions are referred to quasifree Hadamard-states, can be summarized as

\[ < \phi^2 (y) >_{\text{ps}} = \lim_{x \to y} \{ G(x,y) - H(x,y) \}, \]
\[ < T_{ab} (y) >_{\text{ps}} = \lim_{x \to y} D_{ab}(x,y) \{ G(x,y) - H(x,y) \} + g_{ab}(y)Q(y), \]

where $G(x,y)$ is the symmetric part the two-point Wightman function of the considered quantum state or, in Euclidean approach, the corresponding Schwinger function. $H(x,y)$ is the Hadamard parametrix which depends on the local geometry only and takes the short-distance singularity into account. $H(x,y)$ is represented in terms of a truncated series of functions of $\sigma(x,y)$. The operator $D_{ab}(x,y)$ is a bi-tensorial operator obtained by “splitting” the argument of the classical expression of the stress tensor (see [Mo99, Mo03]). Finally $Q(y)$ is a scalar obtained by imposing several physical conditions (essentially, the appearance of the conformal anomaly, the conservation of the stress tensor and the triviality of the Minkowskian limit) [Wa94] in the left-hand side of (30) (see [BD82, Fu91, Wa94, Mo99] for details). More recently, in the framework of Lorentzian generally locally covariant algebraic quantum field theory in curved spacetime, it has established [Mo03] that $Q$ can be omitted, redefining the classical stress-energy tensor, and thus $D_{ab}(x,y)$, into a way that it does not affect the classical expression of $T_{\mu\nu}$ when computed on solutions of the equations of motion, improving the point-splitting procedure. See [Hac10] where that point-splitting procedure is discussed and applied especially to cosmology.
convex neighborhoods:

\[
H_\mu(x,y) = \sum_{j=0}^L u_j(x,y)\sigma(x,y)^j + \delta_D \left[ \sum_{j=0}^M v_j(x,y)\sigma(x,y)^j \ln \left( \frac{\mu^2 \sigma(x,y)}{2} \right) \right] + \delta_D \sum_{j=0}^N w_j(x,y)\sigma(x,y)^j.
\]

(31)

\(L, M, N\) are fixed integers (see [Mo99, Mo03] for details), \(\delta_D = 0\) if \(D\) is odd and \(\delta_D = 1\) otherwise. The coefficients \(u_j\) and \(v_j\) are smooth functions of \((x,y)\) which are completely determined by the local geometry. The coefficients \(w_j\) are determined once one has fixed \(w_0\), and they are omitted [Mo03] when dropping \(Q\). Dealing with Euclidean approaches, it is possible to explicitly compute \(u_j\) and \(v_j\) in terms of heat-kernel coefficients [Mo98, Mo99]. One has the following result [Mo98, Mo99].

**Theorem 4.** If \(M_E\) is compact, \(A \geq 0\) and \(\mu^2 > 0\), and the averaged quantities in the left-hand side below are those defined above in terms of local \(\zeta\)-function regularization, then

\[
<\phi^2(y|A)>_\mu^2 = \lim_{x \to y} \left\{ G(x,y) - H_\mu'(x,y) \right\},
\]

(32)

\[
<T_{ab}(y|A)>_\mu^2 = \lim_{x \to y} D_{ab}(x,y) \left\{ G(x,y) - H_\mu'(x,y) \right\} + g_{ab}(y)Q(y),
\]

(33)

where \(G(x,y) = \zeta(1,x,y|A/\mu^2)\) given in (15), \(H_\mu'\) is completely determined by (31) with the requirement

\[
w_0(x,y) := -\frac{a_{D/2-1}(x,y|A)}{(4\pi)^{D/2}}\left[ 2\gamma + \ln \mu^2 \right],
\]

(34)

and the term \(Q\) is found to be

\[
Q(y) = \frac{1}{D} \left( -P_0(y,y|A) + \delta_D \frac{a_{D/2}(y,y|A)}{(4\pi)^{D/2}} \right).
\]

(35)

If \(D_{ab}\) is defined in order to drop \(Q\) in the right-hand side of (33), \(H_\mu^2\) is determined by fixing \(w_0(x,y) = 0\). The scales \(\mu\) and \(\mu'\) satisfies \(\mu = c\mu'\) for some constant \(c > 0\).

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