Strain-enhanced giant Rashba spin splitting in ultrathin KTaO$_3$ films for spin-polarized photocurrents

Ning Wu$^{ab}$, Xue-Jing Zhang$^{ab\dagger}$ and Bang-Gui Liu$^{a*bc}$

Strong Rashba effects at semiconductor surfaces and interfaces have attracted great attention for basic scientific exploration and practical applications. Here, we show through first-principles investigation that applying biaxial stress can cause tunable and giant Rashba effects in ultrathin KTaO$_3$ (KTO) (001) films with the most stable surfaces. When increasing the in-plane compressive strain to $-5\%$, the Rashba spin splitting energy reaches $E_R = 140$ meV, corresponding to the Rashba coupling constant $a_R = 1.3$ eV Å. We investigate its strain-dependent crystal structures, energy bands, and related properties, and thereby elucidate the mechanism for the giant Rashba effects. Further calculations show that the giant Rashba spin splitting can remain or be enhanced when capping layer and/or Si substrate are added, and a SrTiO$_3$ capping can make the Rashba spin splitting energy reach the record 190 meV. Furthermore, it is elucidated that strong circular photogalvanic effect can be achieved for spin-polarized photocurrents in the KTO thin films or related heterostructures, which is promising for future spintronic and optoelectronic applications.

I. Introduction

The Rashba spin–orbit interaction$^{1-3}$ due to the broken inversion symmetry and the atomic spin–orbit coupling (SOC) can bring about the momentum-dependent spin splitting of the electron states. Rashba effect can play key roles in quantum wells,$^4$ two dimensional (2D) electron gases (2DEG),$^2$ and thin films based on traditional III–V semiconductors.$^6$ After intensive investigations, one can tailor the Rashba coupling by electric field and strain, and design artificial microstructures for various applications. External electric field can be used to modulate the magnitude of Rashba spin splitting in LaAlO$_3$/SrTiO$_3$ (LAO/STO) interface$^7$ and InSe multilayer.$^8$ The Rashba spin splitting can be effectively tuned by varying the interlayer distance in graphene/As–I van der Waals heterostructure$^9$ and adjusting the halogen doping concentration in doped PtSe$_2$ monolayer.$^{10}$ In addition, it is very interesting to manipulate the Rashba spin–orbit coupling by applying stress, as were done in 2D LaO$^{11}$ binary alloyed hexagonal nanosheets,$^{12}$ 2D heterostructures,$^{13}$ Bi$_2$S$_3$ monolayer,$^{14}$ and BiTeI monolayer.$^{15}$ Large Rashba interaction can be used to realize spin polarization and spin injection by applying electric field$^{16,17}$ and produce spin-polarized photocurrents through polarized light.$^{18-21}$

Recently, a 2DEG was observed at KTaO$_3$ (KTO) (100) surface$^{24,25}$ in terms of the angle-resolved photoemission (ARPES) spectrum.$^{24}$ For another 2DEG at an amorphous-LAO/KTO interface, an experimental analysis of the weak anti-localization effect resulted in a Rashba coupling constant 0.1 eV Å and a 50-fold enhanced Hall mobility of charge carriers.$^{26}$ Surprisingly, hysteretic magnetoresistance up to 25 K and anomalous Hall effect up to 70 K were observed at an EuO/KTO interface.$^{27}$ It was shown that LaVO$_3$/KTO heterostructure with strong spin–orbits interactions can be used to produce planar Hall effect, anisotropic magnetoresistance, and 2DEG.$^{28,29}$ Theoretically, the magnitude of Rashba spin splitting in a KTO surface was studied by applying external electric fields in a symmetric slab model.$^{30}$ It is highly desirable to investigate the effects of in-plane strain fields on the strength of Rashba spin splitting for KTO surfaces. Actually, strain (stress) is a wonderful approach to manipulate the crystal structures of KTO and thus control their electronic structures and functional properties. Recent studies have demonstrated that the strain can affect the formation and migration of oxygen vacancies in KTO$^{31}$ and induce electron–hole interchanging of the two opposite surface 2D carrier gases in KTO ultrathin film.$^{32}$ It has been also shown, experimentally and theoretically, that the KTO (001) $1 \times 1$ surfaces terminated with KO and TaO$_2$ are most stable compared to others.$^{33,34}$ Therefore, structurally stable ultrathin KTO (001) films with the two stable surfaces can be manipulated to realize giant Rashba spin effects.
Here, through first-principles calculations and theoretical analyses, we investigate the in-plane strain dependencies of the structural features, intrinsic electrostatic potentials, band edges, carrier concentrations, carrier effective masses, and Rashba parameters of the ultrathin KTO (001) films. We show that their Rashba spin splitting can be controlled by applying biaxial stress, and giant Rashba spin splitting can be obtained by applying compressive biaxial stress, being able to reach 140 meV. In addition, we explore effects of capping layers and/or Si substrate on the giant Rashba spin splitting, and show that a STO capping can make the Rashba spin splitting reach the record 190 meV, as shown in Table 2. Such giant Rashba effects in the ultrathin KTO films can be used to generate spin-polarized photocurrents through circular photogalvanic effect. More detailed results will be presented in the following.

II. Computational method and parameters

Our first-principles calculation is performed using the projector-augmented wave method within the density-functional theory, as implemented in the Vienna Ab initio Simulation Package (VASP). To describe the exchange-correlation energy, we used the general gradient approximation (GGA) with the Perdew–Burke–Ernzerhof for solids (PBEsol) parametrization. The on-site Coulomb interaction in 5d states of transition-metal ions is corrected by the DFT+U (where U is the Hubbard energy) method. The effective value \( U_{\text{eff}} = 3 \) eV is employed for Ta 5d states in this work, as it is well established that such a value is appropriate to describe these strongly-correlated states. An Monkhorst–Pack k-point grid of \( 4 \times 4 \times 1 \) is used for reciprocal space integration, and the plane wave energy cutoff is set to 500 eV. Our convergence standard requires that the Hellmann–Feynman force on each atom is less than 0.01 eV Å\(^{-1}\) and the absolute total energy difference between two successive consistent loops is smaller than \( 1 \times 10^{-5} \) eV. A fully converged electronic structure is used for further calculation including SOC. Band calculations with SOC are confirmed with larger k-point grid \( 6 \times 6 \times 1 \) in cases of zero and \(-5\%\) strain values, and the resulting changes are very small. A 20 Å thick vacuum layer is used in the KTO-slab geometry. Additional calculations with vacuum layer of 30 Å and dipole corrections are made for confirmation. When a biaxial stress is applied, the in-plane strain is defined as \( \epsilon_a = (a - a_0) / a_0 \times 100\% \), where \( a_0 \) is the experimental lattice constant of bulk KTO without strain (\( a_0 = 3.989 \) Å (ref. 43)) and \( a \) is the in-plane lattice constant of strained KTO slab. We also optimize cubic KTO bulk by using the above parameters and better convergence standards: k-point grid \( 15 \times 15 \times 15 \), the Hellmann–Feynman force \( 0.001 \) eV Å\(^{-1}\), and total energy \( 1 \times 10^{-8} \) eV. This optimization produces the theoretical lattice constant of KTO: \( a_T = 3.996 \) Å. Because \( a_T \) is only 0.2% larger than \( a_0 \), we use the experimental \( a_0 \) as our reference when defining strain values. Given an in-plane strain value, the out-of-plane lattice and all the internal atomic positions are allowed to relax sufficiently during optimization.

III. Results and discussion

A. KTO slab under biaxial stress

We construct a KTO slab model to describe the KTO ultrathin film under different biaxial stresses. The slab consists of \( m = 12 \) KTO unit cells along the vertical [001] axis. Fig. 1(a) shows the optimized structure of the KTO slab at the in-plane strain \( \epsilon_a = 0\% \).
0% (zero stress). We study the strained KTO slabs with the in-plane strain $\varepsilon_\parallel$ ranging from $-5\%$ (compressive) to $+8\%$ (tensile). With a given in-plane strain, the system is fully optimized, with the out-of-plane strain being determined by requiring that the out-of-plane stress is zero, and thus we can determine the in-plane stress. Actually, this is a system with biaxial stress. With the condition that the out-of-plane stress is zero, however, the in-plane stress is determined by the in-plane strain. Therefore, for convenience, we shall use the in-plane strain $\varepsilon_\parallel$ to characterize the strained slabs in the following. It is confirmed that the dipole correction has little effect in these results. In Fig. 1(b–m), we plot the representative electronic band structures along $M \,(\pi, \pi) \rightarrow \Gamma \,(0,0) \rightarrow X \,(0,\pi)$ of the optimized KTO slabs for $\varepsilon_\parallel = 0\%, -1\%, -2\%, -3\%, -4\%, -5\%$, 1%, 2%, 3%, 4%, 5%, and 6%, respectively. Here, SOC is taken into account, and the $k$ vector is in units of $1/a$, where $a$ is the calculated lattice constant of the strained KTO slab. When $\varepsilon_\parallel$ is larger than $-2\%$, the KTO slab remains metallic. For the metallic state, there are electron carriers near the $\Gamma$ point and hole carriers near the $M$ point, and they form a 2DEG at the TaO$_2$-terminated surface and a 2D hole gas (2DHG) at the KO-terminated surface. It should be pointed out that the electron concentration in the 2DEG is equivalent to the hole concentration in the 2DHG.\textsuperscript{32} There is a critical strain $\varepsilon_\parallel = -2\%$, as shown in Fig. 1(d). From Fig. 1(g, f and e), it is clear that the KTO slab is insulating for $\varepsilon_\parallel = -5\%, -4\%$, and $-3\%$. Therefore, there is a strain-driven metal–insulator transition (MIT) at $\varepsilon_\parallel = -2\%$. For the insulating phase, there is an indirect semiconductor gap increasing with $|\varepsilon_\parallel + 0.02|$, with the conduction band minimum (CBM) being at $\Gamma$ point and the valence band maximum (VBM) at the $M$ point.

To show the stress-driven structural features, we present in Fig. 2 the monolayer-resolved intra-monolayer ionic bucklings ($B_i$, defined as the maximal cation–anion out-of-plane difference within the monolayer), out-of-plane cation displacements with respect to the centers of the nearest O anions ($D^R_i$ and $D^T_i$), out-of-plane Ta displacements with respect to the centers of the nearest K atoms ($D^R_i$, defined only for even $i$), and inter-monolayer spacings ($S_i$) of the slab under the in-plane strain values between $-5\%$ and $6\%$. It is clear that the atomic positions of the surfaces change substantially with respect to the internal region in the four aspects, and in the internal region the four aspects are made nearly independent of monolayers at strong tensile strains. It is interesting that all the four values monotonically decrease with tensile strain, but increase with compressive strain. When the compressive in-plane strain becomes strong, the bucklings of the surface monolayers are substantially enhanced, and the displacements in the KO monolayers (the TaO$_2$ monolayers) converge to a nearly constant value $d_1 (d_2)$, with $d_2 > d_1$. By combining the buckling and spacing values, it is visible that the two surface single-unit-cells are a little separated from the main body, which is useful to reduce the effect of polarity and stabilize the slab structure. $D^R_i$ is frequently used in experimental measurement reflecting polarization. It is enhanced by compressive stress and reduced by tensile stress.

Furthermore, we plot in Fig. 3(a) the plane-averaged electrostatic potentials for $\varepsilon_\parallel = -5\%, -3\%, 0$, and $4\%$ as representative strain values. It is clear that the maximal (or minimal) value increases from the left to right hand side in the cases of $\varepsilon_\parallel = -3\%$, 0, and 4%, but remains the same for $\varepsilon_\parallel = -5\%$. The internal electric field $E_{\text{int}}$ can be estimated from the slope of the plane-averaged electrostatic potential shown in Fig. 3(a).\textsuperscript{43} The calculated results as a function of $\varepsilon_\parallel$ are presented in Fig. 3(b). The internal electric field at the unstrained KTO slab is $7.1 \times 10^{-2} \text{ V \ Å}^{-1}$, comparable with a previous study.\textsuperscript{43} It is clear that $E_{\text{int}}$ slowly decreases with tensile strain, and accelerates with compressive strain, nearly reaching zero at $\varepsilon_\parallel = -5\%$. It is expected that the out-of-plane cation displacements with respect
atomic SOC further splits the inversion symmetry breaking is taken into account, the Ta reduces the initial cubic symmetry of the Ta t\(_{2g}\) orbitals in the bulk perovskite. The triple degeneracy (excluding spin) of the t\(_{2g}\) bands are labeled by E\(_1\) and E\(_2\).

minimum (at the zone center) lies 0.089 eV below E\(_c\). The corresponding electron concentration of the 2DEG is 2.67 \(\times\) 10\(^{13}\) cm\(^{-2}\), which is an order of magnitude smaller than 2 \(\times\) 10\(^{14}\) cm\(^{-2}\) of the 2DEG formed at an experimental KTO surface from the ARPES measurements.\(^{44}\) This difference can be interpreted by the low formation energy for the oxygen vacancies at the KTO surface,\(^{30,48}\) which allows much more electrons in the 2DEG. For the \(\epsilon_s = 5\%\) case shown in Fig. 4(c), the minima of the occupied lowest and second lowest d\(_{xy}\) states (at the \(\Gamma\) point) are 0.182 and 0.027 eV below E\(_{f}\), in which the summed electron concentration is 6.98 \(\times\) 10\(^{13}\) cm\(^{-2}\), larger than that in the unstrained system.

For the \(\epsilon_s = 4\%\) case shown in Fig. 4(a), the lowest d\(_{xy}\) band lies 0.31 eV above the E\(_{f}\), which means that there are no carriers (without doping), in contrast with those of the \(\epsilon_s = 0\%\) and 5\% cases.

It is obvious that there are some Rashba-like spin splitting in the conduction bands of the KTO slab for the \(\epsilon_s\) between −5\% and 8\%. To further investigate the Rashba effects, we also present in Fig. 4 the magnified electron band structures with SOC of the KTO slab near the \(\Gamma\) point for \(\epsilon_s = 0\%\).

To better describe the properties of the lowest and second lowest d\(_{xy}\) conduction bands near the \(\Gamma\) point respectively, defined by E\(_1\) and E\(_2\) in Fig. 4(c), we investigate and show the band edge positions, electron effective masses, and 2DEG concentrations of the E\(_1\) and E\(_2\) bands as functions of \(\epsilon_s\) in Fig. 5. In Fig. 5(a), as \(\epsilon_s\) changes from −5\% to 4\%, the band edge positions of E\(_1\) and E\(_2\) decrease rapidly with large compressive strain, but they change slowly with tensile strain. As \(\epsilon_s\) varying from 5\% to 8\%, the band edge positions of E\(_1\) and E\(_2\) are almost unchanged. It should be noted that the bottom of the conduction band is less affected by the tensile strain, while it is significantly changed by the compressive strain. In detail, the band edge of the E\(_1\) band is below E\(_f\) for \(\epsilon_s \geq −1\%\), and that of E\(_2\) becomes below E\(_f\) for \(\epsilon_s \geq 1\%\).

In Fig. 5(b), the effective mass (\(m^*\)) is evaluated from a second-order fit of the band energies using

\[
m^* = \frac{\hbar^2}{\frac{d^2E(k)}{dk^2}}.
\]

Remarkably, the values of \(m^*_1\) for E\(_1\) and \(m^*_2\) for E\(_2\)
for $E_s$ in the unstrained KTO system are 0.35 and 0.32$m_e$ ($m_e$ is the mass of the free electron), respectively, which are both in excellent agreement with $0.30m_e$ for the KTO surface 2DEG measured by ARPES.25 When the strain changing from $\varepsilon_s = -5\%$ to 8%, $m_1^*$ and $m_1^*$ increase, with $m_1^*$ being always smaller than $m_1^*$. Under $\varepsilon_s = 8\%$, $m_1^*$ and $m_1^*$ reach the maximum values 0.38 and 0.37$m_e$, respectively, which are still smaller than $\sim 0.5$ and 0.6$m_e$ recently determined for a surface 2DEG on STO.25 This suggests that developing high-mobility oxide electronics by KTO is better than by STO.

Fig. 5(c) shows the relationship between $\varepsilon_s$ and the carrier concentrations. For $\varepsilon_s \leq -2\%$, the $E_1$ and $E_2$ bands are empty and the KTO film is insulating, which is consistent with the critical strain of insulator–metal transition shown in Fig. 1. The carrier concentrations $n_1$ and $n_2$ for the $E_1$ and $E_2$ bands have the maximum values at the $\varepsilon_s = 5\%$ and 4%, respectively, and the total 2DEG concentration $n$ of the $E_1 + E_2$ bands reaches the maximum values of $7.06 \times 10^{13}$ cm$^{-2}$ at $\varepsilon_s = 4\%$. This indicates that the conductivity of the 2DEG formed at the surface can be effectively modulated by the in-plane strain.

In addition, we summarize in Table 1 the energy differences ($E_{R1}$ and $E_s$) and the spin–orbit splitting energy $\Delta E_R$ between the lowest and the second lowest Rashba doublets, as defined in Fig. 4.

C. Rashba spin splitting

Since the KTO slab obeys the $C_{nv}$ point group symmetry, the symmetry-allowed linear spin-momentum coupling can be expressed as $H_R = \alpha_R (k_x\sigma_y - k_y\sigma_x)$. According to the linear Rashba model, the dispersion due to the Rashba spin splitting can be described by

$$ E = \frac{\hbar^2}{2m^*}(k \pm k_R)^2 - E_R, $$

where $k$ is the magnitude of the electron wave vector, $m^*$ is the electron effective mass, $E_R = \hbar^2 k_R^2/2m^*$ is the Rashba spin splitting energy, and $k_R$ is the momentum offset $k_R = 2E_R/m^*$. The in-plane spin polarizations of the “+” and “−” eigenstates are oppositely aligned and normal to the electron wave vector. In the isotropic case, the Rashba coupling constant can be estimated by $\alpha_R = 2E_R/k_R$, and $\alpha_R$ depends on the strength of SOC and inversion asymmetry.49

For the KTO slab, the lowest Rashba spin split bands near the $\Gamma$ point are similar to those defined by eqn (1), and we present $E_R$ and $k_R$ in Fig. 5(d) for different in-plane strains. The calculated values of $E_R$ and $k_R$ are summarized in Table 1 for $\varepsilon_s = 5\%$, 0%, −1%, −2%, −3%, −4%, and −5%. In Fig. 5(d), $E_R$ and $k_R$ increase drastically with the compressive strain increasing, but they both remain almost unchanged for increasing tensile strain. In Table 1, noticeably, $E_R$ and $k_R$ are 140 (90) meV and 0.21 (0.24) Å$^{-1}$ for the KTO slab at $\varepsilon_s = −5\%$ (−4%). It is clear that compressive in-plane strain can enhance the Rashba spin splitting energy $E_R$. Because $E_{int}$ is near zero at $\varepsilon_s = −5.0\%$, $E_{int}$ becomes negative when $\varepsilon_s < −5.0\%$, and consequently the conduction bands are reconstructed, which leads to smaller $E_R$ or substantial deformation of the Rashba bands. Actually, this means that the maximal $E_R$ is reached at $\varepsilon_s = −5\%$.

To elucidate the mechanism responsible for the giant Rashba spin splitting, we investigate the strain-dependent structural parameters and electrostatic potentials. For the unstrained KTO slab, there is a strong intrinsic electric field $e_0$ due to the out-of-plane alternate alignment of negative KO and positive TaO$_2$ monolayers. When compressive biaxial stress is applied, there are out-of-plane displacements of cations with respect to the neighboring anions driven by the compressive in-plane strain $\varepsilon_s$ and tensile out-of-plane strain, and in addition the ionic displacements cause an out-of-plane electric field $e_d$.

| Strain  | $E_{R1}$ (meV) | $E_s$ (meV) | $\Delta E_R$ (meV) | $E_R$ (eV) | $k_R$ (Å$^{-1}$) |
|---------|----------------|------------|-------------------|-----------|----------------|
| $\varepsilon_s = 5\%$ | 0.188 | 0.822 | 0.339 | 7 | 0.046 |
| $\varepsilon_s = 0\%$ | 0.466 | 1.144 | 0.298 | 9 | 0.054 |
| $\varepsilon_s = −1\%$ | 0.603 | 1.282 | 0.309 | 20 | 0.096 |
| $\varepsilon_s = −2\%$ | 0.810 | 1.516 | 0.309 | 33 | 0.151 |
| $\varepsilon_s = −3\%$ | 1.161 | 1.947 | 0.314 | 51 | 0.154 |
| $\varepsilon_s = −4\%$ | 1.618 | 2.652 | 0.325 | 90 | 0.24 |
| $\varepsilon_s = −5\%$ | 2.228 | 4.136 | 0.093 | 140 | 0.21 |

This journal is © The Royal Society of Chemistry 2020
antiparallel to $e_a$. Our calculated results show that $e_d$ increases with $e_s$, reaching the maximum nearly at $e_s = -5\%$. Consequently, we can attribute the enhanced giant Rashba spin splitting energy to the strong intrinsic out-of-plane electric field $e_d$ due to the large compressive biaxial stress.

For comparison, we summarize the $E_R$, $k_R$, and $\alpha_R$ values of some typical Rashba systems in Table 2. For brevity, we can take $E_R$ as the key parameter to characterize such Rashba systems. $E_R$ can reach 100 meV for BiTeI van der Waals bulk,29 or 190 meV for $\alpha$-GeTe(111) film.31 In contrast, for perovskites, the previous maximal $E_R$ is 15 meV for KTaO$_3$/BaHfO$_3$ interface.32 It is clear that our strategy is very efficient to promote the Rashba spin splitting energy in perovskite oxides because our maximal $E_R$ value reaches 140 meV at $e_s = -5\%$. Such large Rashba effect is useful to realize strong circular photogalvanic effect.18–23

D. Heterostructures and photogalvanic effect

Recent STM experiment and DFT calculation have already shown that the KTO (001) $1 \times 1$ surfaces terminated with KO and TaO$_2$ are the most stable compared to others.33,34 For more applications, the KTO films could be capped with STO overlayers, and/or should be grown on good substrates. The metallic interfaces can be obtained by generating carriers through experimentally applying gate voltage.26 Furthermore, the STO capping can change the Rashba spin splitting in the bare TaO$_2$-terminated surfaces of KTO films. For the in-plane strains of $-3\%$, $-4\%$, and $-5\%$, our calculations show that the maximal Rashba spin splitting energies are 25, 95, and 190 meV, respectively, and the corresponding $k$ vector offsets are 0.12, 0.19, and 0.24 Å$^{-1}$. It is a surprise that the STO capping can make the Rashba spin splitting enhanced to 190 meV at strain of $-5\%$! We also study STO/KTO/Si trilayer to simulate capped KTO films on Si substrate. For STO/KTO/Si trilayer, the lattice mismatch is 3.6% for STO/KTO on Si substrate, and then $e_s$ is $-3.6\%$ and the Rashba spin splitting energy is 51 meV, with the $k$ vector offset being 0.16 Å$^{-1}$. It is clear that the Rashba effects are still very strong after capping layers and/or substrates are added.

Because of the giant Rashba spin splitting, the KTO ultrathin films can be used for achieving circular photogalvanic effect (CPGE) to generate spin-polarized photocurrents.18–23 For the right-handed (left-handed) circularly polarized light, its photon has the angular momentum of $+1$ ($-1$), labeled by $\sigma_+$ ( $\sigma_-$), and the selection rule for necessary transitions is that the allowed $z$-component change of the total angular momentum is $\Delta m_z = 1$ ($-1$). The valence band edge, originating from the $J = 3/2$ states, has $J_m = \pm 3/2$, and the conduction band edge, from the $d_{3/2}$ states of $l_m = 2$ and $-2$, consists of $J_m = \pm 3/2$ and $\pm 5/2$ states. Because the $d_{3/2}/d_{5/2}$ bands have $l_m = 1$ and $-1$, the Rashba split bands consist of $J_m = \pm 3/2$ and $\pm 1/2$ states. Therefore, for achieving the CPGE, the electrons can transit from the valence band top (with photon energy $E_1$) or the filled conduction band edge ($E'_R$), to the $d_{3/2}/d_{5/2}$-based bands with giant Rashba spin splitting as the final states, as shown in Fig. 6(a). When the electron concentration is small, $E'_R$ is a little lower than $E_{R1}$.

For generating a net spin-polarized photocurrent, both circularly-polarized light and the Rashba split bands are necessary.18 Upon illumination with a circularly polarized light with photon energy $h\omega$ and given helicity, the energy and angular momentum conservations require that the transition happens only at the two asymmetric $k$ values: $k_{s+}$ and $k_{s-}$.18 This makes the average electron velocity in the excited state become nonzero and the contributions of $k_{s-}$ photoelectrons to the current do not cancel each other.18 Changing the photon helicity from $+1$ to $-1$ inverts the current because the “center-of-mass” for this transition is shifted in the opposite direction. This results in the generation of the spin polarized CPGE current of the Rashba split $d_{3/2}/d_{5/2}$ bands, as shown in Fig. 6(b)

![Fig. 6 Schematic for CPGE: photon energy for transition from the valence band edge ($E_1$) or from the conduction band edge ($E'_R$) to the Rashba bands (a); the two permitted transitions (only two $k$ values, $k_{s+}$ and $k_{s-}$, for a given $e_s$) from the conduction band edge through circularly polarized light with $\sigma_+$ (red) or $\sigma_-$ (purple) (b), and the similar transitions from the valence band edge (c); and the strain dependence of the needed photon energies $E'_R$ and $E_1$ (d).](image-url)
and c). The photon energies needed for the CPGE from the valence band edge ($E_v$) and the conduction band edge ($E_{c1}$) are shown in Fig. 6(d). In principle, it can be described by $j = \gamma E \mathbf{E}_c$, where $\gamma$ is the second-rank pseudotensor, $E$ is the amplitude of the electric field of the light, $\mathbf{e}$ is the unit vector pointing in the direction of the light propagation and $P_c$ is the helicity of the light beam, and the pseudotensor $\gamma$ has non-zero element for the $C_{3v}$ point group, which results in a non-zero spin-polarized photoelectron current.\textsuperscript{19,79,69}

IV. Conclusion

In summary, through the first-principles calculations, we have systematically investigated the effect of the biaxial stress on the Rashba spin splitting of the KTO slabs for modelling strained KTO [001] ultrathin films with the most stable surfaces.\textsuperscript{39,44} The calculated results reveal that tensile in-plane strain changes the Rashba spin splitting little, but the Rashba spin splitting energy $E_R$ increases with compressive stress increasing, which is in reasonable agreement with recent experimental result. When the compressive in-plane strain approaches $\varepsilon = -5\%$, $E_R$ is 140 meV and the $k$ vector offset is $k_R = 0.21 \text{Å}^{-1}$, which corresponds to a large Rashba coupling constant $1.3 \text{ eV Å}$. Our analysis indicates that the enhanced giant Rashba spin splitting energy can be attributed to the strong out-of-plane electric field generated by the large compressive biaxial stress. Our further calculations show that the Rashba spin splitting energy can be enhanced to the record 190 meV by adding a STO capping layer. We also show that such giant Rashba effect can be used to generate spin-polarized photocurrents in terms of the circular photogalvanic effect.\textsuperscript{18–23} Therefore, these giant Rashba phenomena may open a new door to promising spintronic and optoelectronic applications.

Conflicts of interest

There are no conflicts to declare.

Acknowledgements

This work is supported by Department of Science and Technology of China (Grant No. 2016YFA0300701), Nature Science Foundation of China (Grant No. 11974393 and 11574366), and the Strategic Priority Research Program of the Chinese Academy of Sciences (Grant No. XDB33000000). The calculations were performed in the Milky Way #2 supercomputer system at National Supercomputer Center of Guangzhou, Guangzhou, China.

References

1 E. I. Rashba, Sov. Phys. Solid State, 1960, 2, 1109.
2 Y. A. Bychkov and E. I. Rashba, J. Phys. C: Solid State Phys., 1984, 17, 6039.
3 Y. A. Bychkov and E. I. Rashba, JETP Lett., 1984, 39, 78.
4 M. O. Nestoklon, E. L. Ivchenko, J.-M. Jancu and P. Voisin, Phys. Rev. B: Condens. Matter Mater. Phys., 2008, 77, 155328.
5 E. Lesne, Y. Fu, S. Oyarzun, J. C. Rojas-Sánchez, D. C. Vaz, H. Naganuma, G. Sicoli, J.-P. Attané, M. Jamet, E. Jacquet, J.-M. George, A. Barthélémy, H. Jaffrès, A. Fert, M. Bibles and L. Vila, Nat. Mater., 2016, 15, 1261.
6 J. Nitta, T. Akazaki, H. Takayanagi and T. Enoki, Phys. Rev. Lett., 1997, 78, 1335.
7 A. D. Caviglia, M. Gabay, S. Gariglio, N. Reyren, C. Cancellieri and J.-M. Triscone, Phys. Rev. Lett., 2010, 104, 126803.
8 K. Premasiri, S. K. Radha, S. Sucharitakul, U. R. Kumar, R. Sankar, F.-C. Chou, Y.-T. Chen and X. P. A. Gao, Nano Lett., 2018, 18, 4403.
9 N. Yu, J. Yuan, K. Li and J. Wang, Appl. Surf. Sci., 2018, 427, 10.
10 M. A. U. Absor, I. Santoso, Harsojo, K. Abraha, H. Kotaka, F. Ishii and M. Saito, Phys. Rev. B, 2018, 97, 205138.
11 Q. Liu, Y. Guo and A. J. Freeman, Nano Lett., 2013, 13, 5264.
12 L. Zhu, T. Zhang, G. Chen and H. Chen, Phys. Chem. Chem. Phys., 2018, 20, 30133.
13 Q. Zhang and U. Schwingenschlögl, Phys. Rev. B, 2018, 97, 155415.
14 S. Singh and A. H. Romero, Phys. Rev. B, 2017, 95, 165444.
15 S.-H. Zhang and B.-G. Liu, Phys. Rev. B, 2019, 100, 165429.
16 Z. Kovacs-Krausz, A. M. Hoque, P. Makk, B. Szentpeteri, M. Kocsis, B. Fulop, M. V. Yakushev, T. V. Kuznetsova, O. E. Tereshchenko, K. A. Kokh, I. E. Lukacs, T. Taniguchi, K. Watanabe, S. P. Dash and S. Csonka, Nano Lett., 2020, 20, 4782.
17 D. Culcer, A. C. Keser, Y. Q. Li and G. Tkachov, 2D Mater., 2020, 7, 022007.
18 S. D. Ganichev, E. L. Ivchenko, S. N. Danilov, J. Eroms, W. Wegscheider, D. Weiss and W. Prett, Phys. Rev. Lett., 2001, 86, 4358.
19 M. M. R. Moayed, T. Bielewicz, M. S. Zöllner, C. Herrmann and C. Klinke, Nat. Commun., 2017, 8, 15721.
20 Y. Zhang, T. Holder, H. Ishizuka, F. de Juan, N. Nagaosa, C. Felser and B. H. Yan, Nat. Commun., 2019, 21, 3783.
21 H. T. Duc, C. Ngo and T. Meier, Phys. Rev. B, 2019, 100, 045308.
22 M. Chen, K. Lee, J. Li, L. Cheng, Q. Wang, K. Cai, E. E. M. Chia, H. Chang and H. Yang, ACS Nano, 2020, 14, 3539.
23 D. Rees, K. Manna, B. Lu, T. Morimoto, H. Borrmann, C. Felser, J. E. Moore, D. H. Torchinsky and J. Orenstein, Sci. Adv., 2020, 6, eaba0509.
24 P. D. C. King, R. H. He, T. Eknapakul, P. Buaphet, S.-K. Mo, Y. Kaneko, S. Harashima, Y. Hikita, M. S. Bahrarny, C. Bell, Z. Hussain, Y. Tokura, Z.-X. Shen, H. Y. Hwang, F. Baumberger and W. Meevasana, Phys. Rev. Lett., 2012, 108, 117602.
25 A. F. Santander-Syro, C. Bareille, F. Fortuna, O. Copie, M. Gabay, F. Bertran, A. Taleb-Ibrahimi, P. Le Fèvre, G. Herranz, N. Reyren, M. Bibles, A. Barthélémy, P. Lecoer, J. Guevara and M. J. Rozenberg, Phys. Rev. B: Condens. Matter Mater. Phys., 2012, 86, 121107.
26 H. Zhang, X. Yan, X. Zhang, S. Wang, C. Xiong, H. Zhang, S. Qi, J. Zhang, F. Han, N. Wu, B. Liu, Y. Chen, B. Shen and J. Sun, ACS Nano, 2019, 13, 609.
27 H. Zhang, Y. Yun, X. Zhang, H. Zhang, Y. Ma, X. Yan, F. Wang, G. Li, R. Li, T. Khan, Y. Chen, W. Liu, F. Hu, B. Liu, B. Shen, W. Han and J. Sun, Phys. Rev. Lett., 2018, 121, 116803.
28 N. Wadehra, R. Tomar, R. Varma, R. Gopal, Y. Singh, S. Dattagupta and S. Chakraverty, Nat. Commun., 2020, 11, 874.
29 S. Goyal, N. Wadehra and S. Chakraverty, Adv. Mater. Interfaces, 2020, 7, 2000646.
30 K. V. Shanavas and S. Satpathy, Phys. Rev. Lett., 2014, 112, 086802.
31 J. Xi, H. Xu, Y. Zhang and W. J. Weber, Phys. Chem. Chem. Phys., 2017, 19, 6264.
32 X.-J. Zhang and B.-G. Liu, Phys. Chem. Chem. Phys., 2018, 20, 24257.
33 M. Setvin, M. Reticcioli, F. Poelzleitner, J. Hulva, M. Schmid, L. A. Boatner, C. Franchini and U. Diebold, Science, 2018, 359, 572.
34 Y. Wang, J. Cheng, M. Behtash, W. Tang, J. Luo and K. Yang, Phys. Chem. Chem. Phys., 2018, 20, 18515.
35 P. Hohenberg and W. Kohn, Phys. Rev., 1964, 136, B864.
36 W. Kohn and L. J. Sham, Phys. Rev., 1965, 140, A1133.
37 G. Kresse and D. Joubert, Phys. Rev. B: Condens. Matter Mater. Phys., 1999, 59, 1758.
38 P. E. Blöchl, Phys. Rev. B: Condens. Matter Mater. Phys., 1994, 50, 17953.
39 J. P. Perdew, K. Burke and M. Ernzerhof, Phys. Rev. Lett., 1996, 77, 3865.
40 J. P. Perdew, A. Ruzsinszky, G. I. Csonka, O. A. Vydrov, G. E. Scuseria, L. A. Constantin, X. Zhou and K. Burke, Phys. Rev. Lett., 2008, 100, 136406.
41 V. I. Anisimov, F. Aryasetiawan and A. Lichtenstein, J. Phys.: Condens. Matter, 1997, 9, 767.
42 B. Lennart, Phys. Rev. B: Condens. Matter Mater. Phys., 1999, 59, 12301.
43 R. Peng, X. P. Shen, X. Xie, H. C. Xu, S. Y. Tan, M. Xia, T. Zhang, H. Y. Cao, X. G. Gong, J. P. Hu, B. P. Xie and D. L. Feng, Phys. Rev. Lett., 2014, 112, 107001.
44 D.-H. Choe, D. West and S. Zhang, Phys. Rev. Lett., 2018, 121, 196802.
45 C. Cancellieri, D. Fontaine, S. Gariglio, N. Reyren, A. D. Caviglia, A. Fête, S. J. Leake, S. A. Pauli, P. R. Willmott, M. Stengel, P. Ghosez and J.-M. Triscone, Phys. Rev. Lett., 2011, 107, 056102.
46 N. Wadehra, R. Tomar, S. Halder, M. Sharma, I. Singh, N. Jena, B. Prakash, A. De Sarkar, C. Bera, A. Venkatesan and S. Chakraverty, Phys. Rev. B, 2017, 96, 115423.
47 C. Xia, W. Xiong, J. Du, T. Wang, Y. Peng and J. Li, Phys. Rev. B, 2018, 98, 165424.
48 W. Meevasana, P. D. C. King, R. H. He, S.-K. Mo, M. Hashimoto, A. Tamai, P. Songsiriritthigul, F. Baumberger and Z.-X. Shen, Nat. Mater., 2011, 10, 114.
49 Z. Zhong, L. Si, Q. Zhang, W.-G. Yin, S. Yunoki and K. Held, Adv. Mater. Interfaces, 2015, 2, 1400445.
50 K. Ishizaka, M. S. Bahramy, H. Murakawa, M. Sakano, T. Shimojima, T. Sonobe, K. Koizumi, S. Shin, H. Miyahara, A. Kimura, K. Miyamoto, T. Okuda, H. Namatame, M. Taniguchi, R. Arita, N. Nagaosa, K. Kobayashi, Y. Murakami, R. Kumai, Y. Kaneko, Y. Onose and Y. Tokura, Nat. Mater., 2011, 10, 521.
51 J. Krempaský, H. Volfová, S. Muff, N. Pilet, G. Landolt, M. Radović, M. Shi, D. Kriegner, V. Holý, J. Braun, H. Ebert, F. Bisti, V. A. Rogalev, V. N. Strocov, G. Springholz, J. Minár and J. H. Dil, Phys. Rev. B, 2016, 94, 205111.
52 M. Kim, J. Ihm and S. B. Chung, Phys. Rev. B, 2016, 94, 115431.
53 Y. Ma, Y. Dai, W. Wei, X. Li and B. Huang, Phys. Chem. Chem. Phys., 2014, 16, 17603.
54 Y.-H. Meng, W. Bai, H. Gao, S.-J. Gong, J.-Q. Wang, C.-G. Duan and J.-H. Chu, Nanoscale, 2017, 9, 17957.
55 L. G. D. da Silveira, P. Barone and S. Picozzi, Phys. Rev. B, 2016, 93, 245159.
56 S. Thiel, G. Hammerl, A. Schmehl, C. W. Schneider and J. Mannhart, Science, 2006, 313, 1942.
57 S. Zhang, N. Tang, W. Jin, J. Duan, X. He, X. Rong, C. He, L. Zhang, X. Qin, L. Dai, Y. Chen, W. Ge and B. Shen, Nano Lett., 2015, 15, 1152.
58 H. Yuan, X. Wang, B. Lian, H. Zhang, X. Fang, B. Shen, G. Xu, Y. Xu, S.-C. Zhang, H. Y. Kwong and Y. Cui, Nat. Nanotechnol., 2014, 9, 851.
59 S. Vajna, E. Simon, A. Szilva, K. Palotas, B. Ujfalussy and L. Szunyogh, Phys. Rev. B: Condens. Matter Mater. Phys., 2012, 85, 075404.
60 L. E. Golub, L. E. Ivchenko and B. Z. Spivak, JETP Lett., 2017, 105, 782.