QFT Over the Finite Line. Heat Kernel Coefficients, Spectral Zeta Functions and Selfadjoint Extensions

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Abstract. Following the seminal works of Asorey–Ibort–Marmo and Muñoz–Castañeda–Asorey about selfadjoint extensions and quantum fields in bounded domains, we compute all the heat kernel coefficients for any strongly consistent selfadjoint extension of the Laplace operator over the finite line [0, L]. The derivative of the corresponding spectral zeta function at s = 0 (partition function of the corresponding quantum field theory) is obtained. To compute the correct expression for the $a_{1/2}$ heat kernel coefficient, it is necessary to know in detail which non-negative selfadjoint extensions have zero modes and how many of them they have. The answer to this question leads us to analyze zeta function properties for the Von Neumann–Krein extension, the only extension with two zero modes.

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1. Introduction: Basic Formulas and Results

The physical system on which we will focus is a free massless scalar quantum field theory defined over the finite interval [0, L]. The quantum Hamiltonian that describes the one particle states of this quantum field theory is given by the Laplace operator over the finite line [0, L]. It is a very well-known fact that the Laplace operator over the finite line [0, L] is not an essentially selfadjoint operator, but instead admits an infinite set of selfadjoint extensions. We will denote by $\mathcal{M}$ the set of all the selfadjoint extensions of $\Delta$ over the finite line [0, L]. Physically speaking this means that there is an infinite set of possible quantum field.
theories that describe the behavior of a free quantum massless scalar field confined to propagate in the interval \([0, L]\). To respect the unitarity principle of quantum field theory we must only take into account those selfadjoint extensions of the Laplace operator that give rise to non-negative selfadjoint operators (see [1]). As described in [1] among the set of non-negative selfadjoint extensions we can distinguish between two different types:

1. Non-negative selfadjoint extensions of \(\Delta\) over \([0, L]\) that are non-negative only for certain values of the finite length \(L\) of the interval. Typically, these selfadjoint extensions are non-negative for \(L \geq L_0\) for a given \(L_0\) that depends on the selfadjoint extension. When \(L < L_0\) these selfadjoint extensions have negative eigenvalues and thus give rise to non-unitary quantum field theories. We will call these selfadjoint extensions weakly consistent selfadjoint extensions.

2. Non-negative selfadjoint extensions of \(\Delta\) over \([0, L]\) that are non-negative for any value of the finite length \(L\) of the finite line. These selfadjoint extensions have only zero and positive eigenvalues for any value of \(L \in (0, \infty)\). We will call these selfadjoint extensions strongly consistent selfadjoint extensions and following [1] we denote by \(M_F\) the set of strongly consistent selfadjoint extension.

In this paper, we will focus only on those free massless scalar quantum field theories defined by selfadjoint extensions of \(\Delta\) over \([0, L]\) that are non-negative for all \(L \in (0, \infty)\). Hence, we will only be interested in the selfadjoint extensions contained in \(M_F\). This restriction is very natural from a quantum field theoretical point of view because the space \(M_F\) is stable under the renormalization group transformations. On the other hand, the whole set of non-negative selfadjoint extensions for fixed length \(L\) is not stable under renormalization group transformations since there are non-negative selfadjoint extensions that will lose the non-negativity condition under a renormalization group transformation giving rise to non-unitary quantum field theories (see Ref. [2]). Typically, one distinguishes separated and coupled boundary conditions [3], but in the formulation of [1] this will not be necessary (an extension of the AIM formalism was first addressed in the first chapter of Ref. [4] and later on reformulated in a more rigorous approach in Ref. [5]).

To be able to characterize the selfadjoint extensions of \(M_F\), we will use the Asorey–Ibort–Marmo (AIM) formalism (see [6]) to characterize the selfadjoint extensions of \(\Delta\) over the finite line \([0, L]\). From the first AIM theorem (see [6,1]), the set of selfadjoint extensions of \(\Delta\) over \([0, L]\) is in one-to-one correspondence with the group \(U(2)\). Given any \(U \in U(2)\), we will denote the corresponding selfadjoint extension by \(\Delta_U\). Each selfadjoint extension \(\Delta_U\) is defined by its domain of functions \(D_U \subset H^2([0, L], \mathbb{C})\) being \(H^2([0, L], \mathbb{C})\) the Sobolev space of functions over the finite interval that are \(L^2\) together with their derivatives up to second order. The domain \(D_U \subset H^2([0, L], \mathbb{C})\) that defines the selfadjoint extension \(\Delta_U\) is given in terms of the matrix \(U \in U(2)\) (see [1,6]) by

\[
D_U = \{\psi \in H^2([0, L], \mathbb{C})/\varphi - i\dot{\varphi} = U(\varphi + i\dot{\varphi})\},
\]  

(1.1)
where $\varphi$ and $\dot{\varphi}$ are the boundary data\(^1\) for $\psi \in H^2([0, L], \mathbb{C})$:

$$
\varphi \equiv \begin{pmatrix} \psi(0) \\ \psi(L) \end{pmatrix}, \quad \dot{\varphi} \equiv \begin{pmatrix} -\psi'(0) \\ \psi'(L) \end{pmatrix}.
$$

(1.2)

Following the notation in [1] for any $\psi \in H^2([0, L], \mathbb{C})$, we introduce the two-dimensional column vectors $\varphi_{\pm}(\psi)$:

$$
\varphi_{\pm}(\psi) \equiv \begin{pmatrix} \psi(0) \mp i \psi'(0) \\ \psi(L) \pm i \psi'(L) \end{pmatrix}.
$$

(1.3)

We can write the boundary condition given in Equation (1.1) in terms of $\varphi_{\pm}(\psi)$ as

$$
\varphi_-(\psi) = U \cdot \varphi_+(\psi).
$$

(1.4)

Following the conventions and notation used in [1] we parameterize the elements $U \in U(2)$ using five parameters:

$$
U(\alpha, \beta, n) = e^{i\alpha}[\cos(\beta)I + i \sin(\beta)(n \cdot \sigma)],
$$

(1.5)

where $I$ is the $2 \times 2$ identity matrix, $\sigma = (\sigma_1, \sigma_2, \sigma_3)$ are the Pauli matrices, $n$ is a three-dimensional unit vector ($n_1^2 + n_2^2 + n_3^2 = 1$) and the angles $\alpha$ and $\beta$ are such that

$$
\alpha \in [-\pi, \pi]; \quad \beta \in [-\pi/2, \pi/2].
$$

(1.6)

Using this parameterization, we can characterize the non-zero part of the spectrum for any $\Delta_U \in \mathcal{M}$ including multiplicities of eigenvalues (see the consistency lemma in [1]). The secular equation obtained in [1] for any $\Delta_U \in \mathcal{M}$ is given by

$$
h_U(k) = 2ie^{i\alpha}[\sin(kL)((k^2 - 1) \cos(\beta) + (k^2 + 1) \cos(\alpha))
-2k \sin(\alpha) \cos(kL) - 2kn_1 \sin(\beta)].
$$

(1.7)

The non-zero part $\tilde{\sigma}(\Delta_U)$ of the spectrum of $\Delta_U \in \mathcal{M}$ is given by

$$
\tilde{\sigma}(\Delta_U) = \{k^2 \in \mathbb{R} - \{0\}/h_U(k) = 0\} = \{k^2 \in \mathbb{R} - \{0\}/k \in Z(h_U) - \{0\}\},
$$

(1.8)

where $Z(h_U)$ denotes the set of zeroes of the function $h_U(k)$. For any non-zero root of $h_U(k)$ the multiplicity $d_U(k^2)$ of the corresponding eigenvalue is

$$
\forall k \in Z(h_U) - \{0\}: \quad d_U(k^2) = \text{Res}\left(\frac{d}{dz} \log(h_U(z))\right)_{z=k}.
$$

(1.9)

\(^1\) It is worth noting that for spacetime dimension higher than $1 + 1$ the maximal domain of the symmetric operator $\Delta$ defined over the compact manifold $M$ with smooth boundary $\partial M$ is the Sobolev space $H^2(M, \mathbb{C})$ and the domain of its adjoint $\Delta^\dagger$ is $H^2(M, \mathbb{C}) \oplus \ker(\Delta^\dagger)$. This is of crucial importance from a physical point of view because boundary values of the fields that have singularities represent important physical situations as for example point charge distributions over the boundary. Nevertheless, in the case of spacetime dimension $1 + 1$ and $M = [0, L]$ things become simpler because the space of boundary data is the linear vector space $\mathbb{C}^2$ (see Refs. [4,5]).
Figure 1. This graphic shows the set $\mathcal{M}_F$ in the $\alpha\beta$-plane. In the top corner of the rhombus are placed the Dirichlet boundary conditions, the bottom corner corresponds to Neumann boundary conditions meanwhile the left and right corners correspond to periodic (left for $n_1 = 1$ and right for $n_1 = -1$) and anti-periodic (left for $n_1 = -1$ and right for $n_1 = 1$).

Let us mention that the bound states of a given selfadjoint extension $\Delta_U \in \mathcal{M}$ are given by zeroes of $h_U(z)$ of the form $k = i\kappa$ with $\kappa > 0$, i.e., $k^2 < 0$. Furthermore, note that from Equation (1.7) it is easy to see that $\lim_{k \to 0} h_U(k) = 0$. This fact does not ensure that the corresponding selfadjoint extension $\Delta_U$ admits zero modes. The question about which selfadjoint extensions of $\mathcal{M}_F$ admit zero modes will be solved in the next section.

Once all the selfadjoint extensions of $\mathcal{M}$ have been explicitly characterized using the AIM formalism (see [6] for details), following [1] we can characterize all the selfadjoint extensions that belong to $\mathcal{M}_F$ and hence that give rise to strongly consistent quantum field theories. One of the main results in [1] is the characterization of the set $\mathcal{M}_F \subset \mathcal{M}$ of non-negative selfadjoint extensions $\forall L \in (0, \infty)$ (“strong consistency lemma”):

$$\mathcal{M}_F = \{ U(\alpha, \beta, n) \in U(2) = \mathcal{M} / 0 \leq \alpha \pm \beta \leq \pi \}. \quad (1.10)$$

In Figure 1, we can see a representation of the set $\mathcal{M}_F$ in the $\alpha\beta$-plane. Whereas extensive results on the spectral zeta functions and the heat kernel are available for the standard boundary conditions like Dirichlet, Neumann, Robin or periodic [7–11], general boundary conditions as described in (1.4) have not been analyzed in comparable detail. This is the topic of the current paper. Generic interest in the

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2 Given that the AIM formalism (first AIM theorem in [1,6]) ensures the one-to-one correspondence between selfadjoint extensions of $\Delta$ over $[0, L]$ and unitary matrices of $U(2)$ from now on we will not make a distinction between selfadjoint extensions $\Delta_U \in \mathcal{M}$ and unitary matrices $U \in U(2)$ (see the appendix).
analysis of spectral functions stems from their relevance in global analysis [8,12] and quantum field theory topics such as the Casimir effect [13–20].

The paper is organized as follows. In Section 2, we will answer the question which selfadjoint extensions within the strongly consistent extensions allow for zero modes. This is necessary as the details of the zeta function analysis depend on this input. Based upon the function $h_U(k)$, Equation (1.7), a contour integral representation of the zeta function for any strongly consistent selfadjoint extension will be derived. As usual, residues and certain values of the zeta function determine the associated heat kernel coefficients. The cases with and without zero modes are treated in different subsections of Section 3. Results for standard boundary conditions are verified as a check. In Section 4, we use the integral representation of the zeta function to compute its derivative at $s = 0$, once again for all possible cases. Checks for known results are provided. In the conclusions, we summarize the most important aspects of our work together with possible future directions of research.

2. Zero Modes of $\Delta_U \in \mathcal{M}_F$

The purpose of this section is to study the zero mode structure of selfadjoint extensions contained in $\mathcal{M}_F$. In particular, we will focus our attention on two main questions:

- Characterize the subset $\mathcal{M}_F^{(0)} \subset \mathcal{M}_F$ of selfadjoint extensions that have zero modes,
  \[ \mathcal{M}_F^{(0)} \equiv \{ \Delta_U \in \mathcal{M}_F / \ 0 \in \sigma(\Delta_U) \}. \]  \hfill (2.1)

- Study the zero mode structure and compute $\dim(\ker \Delta_U)$ of any $\Delta_U \in \mathcal{M}_F^{(0)}$.

The motivation to study these two questions about the zero modes of the selfadjoint extensions contained in $\mathcal{M}_F$ is to obtain a correct result of the $a_{1/2}$ heat kernel coefficient, for which we must know explicitly $\dim(\ker \Delta_U)$. There are no contributions of zero modes to residues of the zeta function.

The differential equation for the zero modes is

\[ \frac{d^2}{dx^2} \psi_0(x) = 0, \]  \hfill (2.2)

and its general solution is given by

\[ \psi_0(x) = a + bx, \]  \hfill (2.3)

where $a$ and $b$ are complex constant numbers. Notice that:

- When $\Delta$ is defined over the whole real line, the only solution to Equation (2.2) given by (2.3) with finite $L^2$ norm is given by $a = b = 0$. Hence, when $\Delta$ is defined over the real line there are no zero modes.
On the other hand, when $\Delta$ is defined over the finite line $[0, L]$, due to the finite length of the interval the general solution (2.3) has always finite $L^2$ norm. Hence when $\Delta$ is defined over the finite interval there exists the possibility of having constant and linear zero modes.

Given a selfadjoint extension $\Delta U \in \mathcal{M}_F$, to decide if it admits zero modes of the general form (2.3) we must impose over (2.3) the corresponding boundary condition given by (1.1). From Equation (1.3) we obtain for $\psi_0(x)$

$$
\varphi_\pm(\psi_0) = \begin{pmatrix} a \mp ib \\ a + b(L \mp i) \end{pmatrix} = \begin{pmatrix} 1 & \mp i \\ 1 & L \pm i \end{pmatrix} \cdot \begin{pmatrix} a \\ b \end{pmatrix}.
$$

(2.4)

Using (2.4) in the boundary condition (1.4), we obtain the linear system

$$
\begin{bmatrix}
\begin{pmatrix} 1 & i \\ 1 & L - i \end{pmatrix} - U \cdot \begin{pmatrix} 1 & -i \\ 1 & L + i \end{pmatrix}
\end{bmatrix} \cdot \begin{pmatrix} a \\ b \end{pmatrix} = 0.
$$

(2.5)

This linear system is nothing else than the boundary condition for the zero modes. Given its importance in this section, let us call the matrix of the linear system $D_U$:

$$
D_U = \begin{bmatrix}
\begin{pmatrix} 1 & i \\ 1 & L - i \end{pmatrix} - U \cdot \begin{pmatrix} 1 & -i \\ 1 & L + i \end{pmatrix}
\end{bmatrix}.
$$

(2.6)

Next, we investigate the solutions of the linear system (2.5).

2.1. THE FIRST QUESTION: CHARACTERIZATION OF $\mathcal{M}^{(0)}_F$

From basic algebra, we know that $\Delta U \in \mathcal{M}_F$ will admit zero modes if and only if the linear system (2.5) has non-trivial solutions, i.e.,

$$
\ker(\Delta U) \neq 0 \iff \ker(D_U) \neq 0 \iff \det(D_U) = 0.
$$

(2.7)

Hence, the characterization of $\mathcal{M}^{(0)}_F$ is given by

$$
\mathcal{M}^{(0)}_F = \{ U \in \mathcal{M}_F / \det(D_U) = 0 \}.
$$

(2.8)

To explicitly compute all the selfadjoint extensions contained in $\mathcal{M}^{(0)}_F$, we need to solve the secular equation of the linear system (2.5)

$$
\det(D_U) = 0.
$$

(2.9)

Introducing the parameterization (1.5) in (2.6) and simplifying we obtain

$$
\det(D_U) = 2e^{ia}[L(\cos(\alpha) - \cos(\beta)) - 2(\sin(\alpha) + n_1 \sin(\beta))].
$$

(2.10)

Therefore, neglecting the global factor $2e^{ia}$ that is never, zero the equation to solve is

$$
L(\cos(\alpha) - \cos(\beta)) - 2(\sin(\alpha) + n_1 \sin(\beta)) = 0
$$

(2.11)
with the restrictions ensuring that the corresponding solution gives a matrix $U$ that is in $\mathcal{M}_F$:

$$n_1 \in [-1, 1]; \quad \alpha \in [0, \pi]; \quad \beta \in [-\pi/2, \pi/2]; \quad 0 \leq \alpha \pm \beta \leq \pi.$$  \hfill (2.12)

The simplest way to solve (2.11) is by imposing

$$\cos \alpha - \cos \beta = 0 \implies \alpha = \pm \beta,$$

which makes

$$\sin \alpha + n_1 \sin \beta = \sin \alpha \pm n_1 \sin \alpha = 0 \implies n_1 = \mp 1$$

necessary. In fact, these are all possible solutions. Because if $\cos \alpha - \cos \beta \neq 0$, we have

$$L = \frac{2 \sin \alpha + n_1 \sin \beta}{\cos \alpha - \cos \beta}$$

with $L > 0$. However, with the parameters confined by the conditions in (2.12) one can show that the right-hand side is always negative. As a consequence, we have shown that all the solutions to (2.11) that satisfy conditions (2.12) are given by

$$\mathcal{M}_F^{(0)} = \{ U \in \mathcal{M}_F / \quad n_1 = \pm 1; \quad \alpha \in [0, \pi/2]; \quad \beta = -n_1 \alpha \}. \hfill (2.13)$$

In terms of the parameterization given in (1.5), the unitary matrices contained in $\mathcal{M}_F^{(0)}$ are given by

$$U \in \mathcal{M}_F^{(0)} \Rightarrow U = e^{i\alpha} [\cos(\alpha) I - i n_1 \sin(\alpha) \sigma_1]; \quad \alpha \in [0, \pi/2]; \quad n_1 \in \{-1, 1\}. \hfill (2.14)$$

Using the expression above for the matrices contained in $\mathcal{M}_F^{(0)}$ and the definition (2.6), we find

$$D_U = \left( \begin{array}{cc} 0 & i e^{i\alpha} (2 \cos(\alpha) + L \sin(\alpha)) \\ 0 & -i e^{i\alpha} (2 \cos(\alpha) + L \sin(\alpha)) \end{array} \right) \quad \forall \ U \in \mathcal{M}_F^{(0)}. \hfill (2.15)$$

As can be seen from this expression above, when $U \in \mathcal{M}_F^{(0)}$ the matrix $D_U$ has indeed zero determinant. In Figure 2 it is shown the space $\mathcal{M}_F^{(0)}$ in the $\alpha \beta$-plane.

### 2.2. THE SECOND QUESTION: $\dim(\ker(\Delta U))$ FOR $\Delta U \in \mathcal{M}_F^{(0)}$

Taking into account Equation (2.5) and the meaning of the constants $a$ and $b$, see Equation (2.3), the second question will be answered by studying the explicit solutions to (2.5) when $D_U$ is given by expression (2.15), i.e., $U \in \mathcal{M}_F^{(0)}$. We will answer this second question in two lemmas with their corresponding demonstrations.

**LEMMA 1.** Any selfadjoint extension $\Delta U \in \mathcal{M}_F^{(0)}$ admits a constant zero mode.
Proof. To proof the lemma, we only need to demonstrate that the column vector
\[ v_c^{(0)} = \begin{pmatrix} a \\ 0 \end{pmatrix}, \quad a \neq 0, \tag{2.16} \]
belongs to \( \text{ker}(D_U) \) for any \( U \in \mathcal{M}_F^{(0)} \) (notice that according to (2.3) when \( b = 0 \) and \( a \neq 0 \) the expression gives rise to the constant function over the interval \([0, L]\)). For any \( U \in \mathcal{M}_F^{(0)} \) the associated matrix \( D_U \) is given by (2.15). Since the first column in (2.15) is identically zero by direct trivial calculation
\[ D_U \cdot v_c^{(0)} = \begin{pmatrix} 0 \\ 0 \end{pmatrix} \quad \forall \ U \in \mathcal{M}_F^{(0)}. \tag{2.17} \]
Therefore, \( v_c^{(0)} \) is a solution to the linear system (2.5). Hence, taking (2.3) into account for any \( \Delta U \in \mathcal{M}_F^{(0)} \) there exists a constant zero mode.

This lemma ensures that any selfadjoint extension \( \Delta U \in \mathcal{M}_F^{(0)} \) has at least a constant zero mode, i.e.,
\[ \forall \Delta U \in \mathcal{M}_F^{(0)}, \quad \dim(\text{ker}(\Delta U)) = \dim(\text{ker}(D_U)) \geq 1. \tag{2.18} \]
Since any \( \Delta U \in \mathcal{M}_F^{(0)} \) has a constant zero mode the only possibility to be explored now is the possibility of having selfadjoint extensions \( \Delta U \in \mathcal{M}_F^{(0)} \) that also admit a linear zero mode. The condition for a selfadjoint extension \( \Delta U \in \mathcal{M}_F^{(0)} \) to admit a linear zero mode is given by
\[ \dim(\text{ker}(\Delta U)) = \dim(\text{ker}(D_U)) = 2. \tag{2.19} \]
Since $D_U$ is a $2 \times 2$ complex matrix
\[
\dim(\ker(D_U)) = 2 \iff D_U = 0. \tag{2.20}
\]
This condition ensures the existence of a linear zero mode for any selfadjoint extension $\Delta_U \in \mathcal{M}_F^{(0)}$ by the following argumentation:

i. For any $\Delta_U \in \mathcal{M}_F^{(0)}$ there is a constant zero mode $\Rightarrow v_c^{(0)}$ given by (2.16) belongs to $\ker(D_U)$ for any $\Delta_U \in \mathcal{M}_F^{(0)}$.

ii. $\Delta_U \in \mathcal{M}_F^{(0)}$ will admit a linear zero mode if and only if the matrix $D_U$ is such that there exists in addition to $v_c^{(0)}$ a solution to the linear system (2.5) with $b \neq 0$ [see Equation (2.3)].

iii. Hence $\Delta_U \in \mathcal{M}_F^{(0)}$ will admit a linear zero mode if and only if
\[
\dim(\ker(D_U)) = 2 \iff D_U = 0, \tag{2.21}
\]
because any solution to (2.5) with $b \neq 0$ will be linearly independent of the vector $v_c^{(0)} \in \ker(D_U) \ \forall \ \Delta_U \in \mathcal{M}_F^{(0)}$.

**Lemma 2.** There are no selfadjoint extensions $\Delta_U \in \mathcal{M}_F^{(0)}$ that admit a linear zero mode.

**Proof.** Given any $\Delta_U \in \mathcal{M}_F^{(0)}$ the necessary and sufficient condition to admit a linear zero mode is (2.20). Since for $\Delta_U \in \mathcal{M}_F^{(0)}$ the associated $D_U$ matrix is given by (2.15) the condition $D_U = 0$ is given by the equation
\[
2 \cos(\alpha) + L \sin(\alpha) = 0 \Rightarrow \tan(\alpha) = -2/L. \tag{2.22}
\]
Because $L$ is the length of the interval $-2/L \leq 0$. Therefore, there is no $\alpha \in [0, \pi/2]$ satisfying (2.22). Hence, no $\Delta_U \in \mathcal{M}_F^{(0)}$ can satisfy the condition $D_U = 0$, i.e., no $\Delta_U \in \mathcal{M}_F^{(0)}$ admits a linear zero mode. \qed

To conclude this section, we compile all the results in the following theorem.

**Theorem 1.** The space $\mathcal{M}_F^{(0)} \subset \mathcal{M}_F$ of non-negative selfadjoint extensions of the Laplace operator $\Delta$ over $[0, L]$ that admit zero modes is given by
\[
\mathcal{M}_F^{(0)} = \{ U \in \mathcal{M}_F / \ n_1 = \pm 1; \ \alpha \in [0, \pi/2]; \ \beta = -n_1 \alpha \}. \tag{2.23}
\]
In addition, $\dim(\ker(\Delta_U)) = 1$ for any selfadjoint extension $\Delta_U \in \mathcal{M}_F^{(0)}$ and the unique zero mode is the constant function over the interval $[0, L]$.

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3 Since $\Delta_U \in \mathcal{M}_F^{(0)}$ the angle $\alpha$ is restricted to lie in the interval $[0, \pi/2]$. 
2.3. A REMARK ABOUT THE VON NEUMANN–KREIN EXTENSION

To complete the study of the zero modes, we will determine the minimal non-negative selfadjoint extension: the so-called Von Neumann–Krein extension. To introduce the general definition of the Von Neumann–Krein extension, we need a quick overview of some general results (see Refs. [21–23]). Let $T_1$ and $T_2$ be two non-negative selfadjoint operators with dense domains in a Hilbert space $\mathcal{H}$. We say that

$$T_1 \leq T_2$$

if and only if

1. $D(T_1^{1/2}) \subseteq D(T_2^{1/2})$,
2. $\langle T_1^{1/2} \psi | T_1^{1/2} \psi \rangle_{L^2} \leq \langle T_2^{1/2} \psi | T_2^{1/2} \psi \rangle_{L^2}$ for all $\psi \in D(T_2^{1/2})$.

Let now $T$ be a non-negative symmetric operator over a Hilbert space $\mathcal{H}$. Then, there exist two unique non-negative selfadjoint extensions $T_{\min}$ and $T_{\max}$ such that

$$T_{\min} \leq T \leq T_{\max}.$$  

(2.25)

The minimal non-negative selfadjoint extension $T_{\min}$ is the so-called Von Neumann–Krein (VNK) extension. From now on, we will denote the Von Neumann–Krein extension with the sub-index VNK.

Following subsection 11.1 in Ref. [23], the VNK extension of the operator $T = -\Delta$ over the finite interval $[0, L]$ is characterized as the unique selfadjoint extension with a maximal number of zero modes. From Equation (2.5), the maximum number of zero modes for the Laplace operator over the finite line is two: a constant zero mode and a linear zero mode. Therefore, the condition that characterizes uniquely the VNK extension is

$$D_U = 0 \Rightarrow U_{\text{VNK}} = \begin{pmatrix} 1 & i \\ 1 & L - i \end{pmatrix} \cdot \begin{pmatrix} 1 & -i \\ 1 & L + i \end{pmatrix}^{-1},$$

(2.26)

$$U_{\text{VNK}} = \frac{1}{L + 2i} \begin{pmatrix} L & 2i \\ 2i & L \end{pmatrix}.$$  

(2.27)

It is straightforward to check that $U_{\text{VNK}}$ is a unitary matrix and therefore defines a selfadjoint extension of the Laplacian over the finite line. To demonstrate whether or not the VNK extension belongs to $\mathcal{M}_F$, we must compute the parameters $\{\alpha_{\text{VNK}}, \beta_{\text{VNK}}, n_{\text{VNK}}\}$ that characterize the VNK extension in the parameterization given by (1.5). Since (2.27) is a symmetric matrix and both diagonal elements are equal we must require $n_2 = n_3 = 0$. Therefore, we can assume without loss of generality that $n_{\text{VNK}} = (1, 0, 0)$. Knowing that

$$U|_{n_1=1} = e^{i \alpha} \begin{pmatrix} \cos(\beta) & i \sin(\beta) \\ \sin(\beta)i & \cos(\beta) \end{pmatrix}$$

(2.28)
and comparing with (2.27), we obtain the following two equations:

\[ e^{i\alpha} \cos(\beta) = \frac{L}{L + 2i}, \]  
\[ e^{i\alpha} \sin(\beta) = \frac{2}{L + 2i}. \]  

(2.29)  

(2.30)

Dividing Equation (2.30) by Equation (2.29), it follows that \( \tan(\beta_{\text{VNK}}) = \frac{2}{L} \).

Since the principal value of \( \arctan(x) \) is in the interval \([-\pi/2, \pi/2]\), we conclude that \( \beta_{\text{VNK}} = \arctan(2/L) \). In addition, it is easy to see that \( \sin(\beta_{\text{VNK}}) = 2/\sqrt{L^2 + 4} \) and \( \cos(\beta_{\text{VNK}}) = L/\sqrt{L^2 + 4} \). To determine \( \alpha_{\text{VNK}} \), we sum (2.29) + i (2.30) to obtain the equation \( e^{i(\alpha_{\text{VNK}} + \beta_{\text{VNK}})} = 1 \Rightarrow \alpha_{\text{VNK}} = -\beta_{\text{VNK}} = -\arctan(2/L) \). Hence, the VNK extension is characterized by:

\[ n_{\text{VNK}} = (1, 0, 0), \quad \alpha_{\text{VNK}}(L) = -\beta_{\text{VNK}}(L), \]  
\[ \beta_{\text{VNK}}(L) = \arctan(2/L). \]  

(2.31)  

(2.32)

Taking into account that \( \arctan(2/L) \) is a non-negative and monotonically decreasing function in the interval \( L \in [0, \infty) \) that goes from the value \( \pi/2 \) for \( L \to 0 \) to the value \( 0 \) when \( L \to \infty \) we conclude that \( \beta_{\text{VNK}} \in [0, \pi/2] \) and \( \alpha_{\text{VNK}} \in [-\pi/2, 0] \) for any \( L \in (0, \infty) \). Therefore, \( U_{\text{VNK}} \notin \mathcal{M}_F \) for any value of \( L \). Nevertheless, since the VNK extension is non-negative, we will be able to compute the heat kernel coefficients and the derivative at zero of the spectral zeta function with the methods we develop in the following sections.

3. The Heat Kernel Expansion of \( \Delta_U \in \mathcal{M}_F \)

Using standard methods described for example in Ref. [10], we will next compute all the coefficients of the asymptotic expansion of the heat kernel corresponding to any selfadjoint extension \( \Delta_U \in \mathcal{M}_F \). Before going over the explicit calculation let us introduce the general results contained in [10] that will be necessary in our calculation.

Let \( \hat{\Delta} \) be an elliptic non-negative selfadjoint second-order differential operator (in one dimension) over a Hilbert space \( \mathcal{H} \). Let \( f_{\hat{\Delta}}(z) \) be a holomorphic function over the complex plane such that for \( k \in \mathbb{R} \)

\[ \lim_{k \to 0} f_{\hat{\Delta}}(k) \neq 0, \infty, \]  

(3.1)

and such that the non-zero part of the spectrum of \( \hat{\Delta} \) is given by\(^4\)

\[ \hat{\sigma}(\hat{\Delta}) = Z(f_{\hat{\Delta}}), \]  

(3.2)

\(^4\)We will denote by \( \sigma(\hat{\Delta}) \) the spectrum of the operator \( \hat{\Delta} \) and \( \hat{\sigma}(\hat{\Delta}) \) the non-zero part of \( \sigma(\hat{\Delta}) \). Given a function \( f(z) \) over the complex plane, we will denote by \( Z(f) \) the set of its zeroes over the complex plane.
where the multiplicities of eigenvalues are reflected in the order of the zeroes. When \( f_\mathcal{O} \) satisfies the conditions stated above, the spectral zeta function of the operator \( \hat{\mathcal{O}} \) can be written as:

\[
\zeta_\mathcal{O}(s) = \frac{\sin(\pi s)}{\pi} \int_0^\infty dk \cdot k^{-2s} \partial_k \log(f_\mathcal{O}(ik)).
\]  

(3.3)

This approach has been used for many examples; see, e.g., [10,24]. The integral in (3.3) in the current context will be convergent in the region \( 1/2 < \Re s < 1 \). However, expression (3.3) admits an analytical continuation to the whole complex plane with, in general, poles at

\[
s = \frac{1}{2} - n; \quad n = 0, 1, 2, 3, \ldots
\]  

(3.4)

The heat kernel coefficients can be computed in terms of the residues at the poles and the values at non-positive integers of \( \zeta_\mathcal{O}(s) \) [25]:

\[
a_{1/2-z}(\mathcal{O}) = \Gamma(z) \text{Res}(\zeta_\mathcal{O}, s = z),
\]  

(3.5)

\[
a_{1/2+q}(\mathcal{O}) = (-1)^q \frac{\zeta_{\mathcal{O}}(-q)}{\Gamma(q+1)} + \delta_q,0 N_\mathcal{Z}(\mathcal{O}).
\]  

(3.6)

In Equation (3.6), \( N_\mathcal{Z}(\mathcal{O}) \) denotes the number of zero modes of the operator \( \mathcal{O} \).

Hence, according to formulas (3.5) and (3.6), to know all the heat kernel coefficients we only need to know the residues at the poles and the values at the non-positive integers of the spectral zeta function \( \zeta_\mathcal{O}(s) \). To use formula (3.3), we will need to use the secular equation given in formula (1.7). Note, however, that \( k \) needs to be replaced by \( ik \) when used in (3.3).

Directly from formula (1.7), it is easy to see that

\[
\lim_{k \to 0} h_U(k) = 0.
\]  

(3.7)

Therefore, using formula (3.3) to compute the residues and the values at the non-positive integers of \( \zeta_U(s) \) for any \( \Delta_U \in \mathcal{M}_F \) is not possible using the function (1.7) because it does not satisfy the condition (3.1). Hence, we need to extract from (1.7) the suitable function by studying the behavior of \( h_U(z) \) as \( z \to 0 \).

3.1. BEHAVIOR OF \( h_U(z) \) AS \( z \to 0 \)

If we perform power series expansion in \( k \) around \( k = 0 \) of the secular equation given by (1.7) up to first order in \( k \) we obtain

\[
h_U(k) = 2ie^{i\alpha}(L(\cos(\alpha) - \cos(\beta)) - 2(n_1 \sin(\beta) + \sin(\alpha))) + O(k^2).
\]  

(3.8)

Taking into account Equation (2.10) for any \( \Delta_U \in \mathcal{M}_F \), we can write the power series expansion above as

\[
h_U(k) = ik \det(D_U) + O(k^2).
\]  

(3.9)
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Hence for any \( \Delta_U \in \mathcal{M}_F - \mathcal{M}_F^{(0)} \) the function that satisfies the required conditions to be used in the representation of the spectral zeta function given by Equation (3.3) is

\[
\Delta_U \in \mathcal{M}_F - \mathcal{M}_F^{(0)} \Rightarrow f_U(k) = \frac{h_U(k)}{2i k e^{i\alpha}}. \tag{3.10}
\]

When the selfadjoint extension has a constant zero mode \( \Delta_U \in \mathcal{M}_F^{(0)} \), the first order in \( k \) of the power expansion (3.9) is zero. Therefore, we must expand \( h_U \) up to order 3 (notice from Equation (1.7) the function \( h_U \) is odd in \( k \)) to study the behavior at the origin:

\[
\det(D_U) = 0 \Rightarrow h_U(k) = \frac{k^3 L}{3} (L(2 \sin(\alpha) - n_1 \sin(\beta)) + 3(\cos(\alpha) + \cos(\beta))) + O(k^5). \tag{3.11}
\]

Hence, to obtain the function that satisfies the conditions under which (3.3) is valid, we must divide by an extra \( k^2 \) when \( \Delta_U \in \mathcal{M}_F^{(0)} \):

\[
\Delta_U \in \mathcal{M}_F^{(0)} \Rightarrow f_U^{(0)}(k) = \frac{h_U(k)}{2k^3 e^{i\alpha}}. \tag{3.12}
\]

3.2. HEAT KERNEL COEFFICIENTS FOR \( \Delta_U \in \mathcal{M}_F - \mathcal{M}_F^{(0)} \)

For this case, when \( \cos(\alpha) + \cos(\beta) \neq 0 \), the appropriate function is given by Equation (3.10). Using (1.7), we can rewrite (3.10) for \( k = ix \) as

\[
f_U(ix) = e^{xL} \frac{\cos(\alpha) + \cos(\beta)}{2} \left[ 1 + \frac{2 \sin(\alpha)}{x \cos(\alpha) + \cos(\beta)} + \frac{1}{x^2} \frac{\cos(\beta) - \cos(\alpha)}{\cos(\alpha) + \cos(\beta)} \right]
\]

\[-e^{-2xL} \left( 1 + \frac{x^{-2}(\cos(\beta) - \cos(\alpha))}{\cos(\alpha) + \cos(\beta)} - \frac{2x^{-1} \sin(\alpha)}{\cos(\alpha) + \cos(\beta)} \right)
\]

\[+ x^{-1} \frac{4n_1 \sin(\beta)}{\cos(\alpha) + \cos(\beta)} \right]. \tag{3.13}
\]

For positive \( L \), the second and third lines in Equation (3.13) represent exponentially damped terms as \( x \to \infty \). These terms do not contribute to the poles and to the values of \( \zeta_{\Delta_U}(s) \) at non-positive integers. Therefore, we can neglect them in the following formulas and just denote them as \( e.s.t. \). Hence \( \log(f_U(ix)) \) will be given by

\[
\log(f_U(ix)) = \log \left( \frac{\cos(\alpha) + \cos(\beta)}{2} \right) + \log(x) + xL + \log(1 + \tau_U(x)), \tag{3.14}
\]

where \( \tau_U(x) \) is given by

\[
\tau_U(x) = \frac{2}{x} \frac{\sin(\alpha)}{\cos(\alpha) + \cos(\beta)} + \frac{1}{x^2} \frac{\cos(\beta) - \cos(\alpha)}{\cos(\alpha) + \cos(\beta)} + \text{e.s.t.} \tag{3.15}
\]
Now, if we take into account the series expansion
\[
\log(1 + \tau) = \sum_{n=1}^\infty (-1)^{n+1} \frac{\tau^n}{n}, \quad (3.16)
\]
we can write
\[
\log(1 + \tau U(x)) = \sum_{n=1}^\infty (-1)^{n+1} \frac{n}{n} \left( \frac{2 \sin(\alpha)}{x \cos(\alpha) + \cos(\beta)} + \frac{1}{x^2 \cos(\alpha) + \cos(\beta)} \right)^n + \text{e.s.t.} \quad (3.17)
\]
Using Newton’s binomial formula, we can write
\[
\tau U(x)^n / n = \sum_{j=0}^n \frac{\Gamma(n) 2^{n-j} \sin^{n-j}(\alpha) (\cos(\beta) - \cos(\alpha))^j}{\Gamma(j+1) \Gamma(n-j+1) (\cos(\alpha) + \cos(\beta))^n x^{-(n+j)}}. \quad (3.18)
\]
After reordering the double summation, we obtain
\[
\log(1 + \tau U(x)) = \sum_{m=1}^\infty b_m x^{-m}, \quad (3.19)
\]
where
\[
b_m = \sum_{j=0}^{[m/2]} (-1)^{m-j+1} \frac{2^{m-2j} \Gamma(m-j) \sin^{m-2j}(\alpha) (\cos(\beta) - \cos(\alpha))^j}{\Gamma(j+1) \Gamma(m-2j+1) (\cos(\alpha) + \cos(\beta))^{m-j}}, \quad (3.20)
\]
and
\[
\frac{\partial_x \log(f_U(ix))}{F} = L + x^{-1} - \sum_{m=1}^\infty m b_m x^{-m-1} + \text{e.s.t.} \quad (3.21)
\]
Taking into account the integral representation (3.3), we can write for any selfadjoint extension \( \Delta_U \in \mathcal{M}_F - \mathcal{M}_F^{(0)} \),
\[
\zeta_{\Delta_U}(s) = \frac{\sin(\pi s)}{\pi} \int_0^1 dk \cdot k^{-2s} \partial_k \log(f_U(ik)) + \frac{\sin(\pi s)}{\pi} \int_1^\infty dk \cdot k^{-2s} \partial_k \log(f_U(ik)). \quad (3.22)
\]
With this splitting all the information about the poles and the values of \( \zeta_{\Delta_U}(s) \) at the non-positive integers is contained in the integration from 1 to \( \infty \). Therefore, to perform the analytic continuation of \( \zeta_{\Delta_U}(s) \) to the complex plane, we have to perform the analytic continuation of
\[
\frac{\sin(\pi s)}{\pi} \int_1^\infty dk \cdot k^{-2s} \partial_k \log(f_U(ik)) \quad (3.23)
\]
to the complex plane. To do so, we must remember the following identities:

\[
\begin{align*}
\int_{1}^{\infty} dz \cdot z^{-2s} &= \frac{1/2}{s - 1/2}, \\
\int_{1}^{\infty} dz \cdot z^{-2s-1} &= \frac{1/2}{s}, \\
\int_{1}^{\infty} dz \cdot z^{-2s-m-1} &= \frac{1/2}{s + m/2}.
\end{align*}
\] (3.24) (3.25) (3.26)

Hence, the relevant information about the analytic continuation of (3.23) is contained in

\[
\frac{\sin(\pi s)}{\pi} \int_{1}^{\infty} dk \cdot k^{-2s} \partial_k \log(f_U(ik))
= \frac{\sin(\pi s)}{\pi} \left( \frac{L/2}{s - 1/2} + \frac{1/2}{s} - \sum_{m=1}^{N-1} b_m \frac{m/2}{s + m/2} + A(s) \right),
\] (3.27)

where \( N \in \mathbb{N} \) and \( A(s) \) in the bracket of the right-hand side represents a meromorphic function of \( s \) analytic for \( \Re s > -N/2 \). The integer \( N \) can be chosen as large as we wish and details of the function \( A(s) \) are irrelevant for our purposes in this section. Using this analytic continuation, the poles of \( \zeta_{\Delta_U}(s) \) can be easily computed:

\[
\begin{align*}
\text{res} \left( \zeta_{\Delta_U}(s), s = 1/2 \right) &= \text{res} \left( L \sin(\pi s)/(2\pi (s - 1/2), s = 1/2 \right) \\
&\Rightarrow \text{res} \left( \zeta_{\Delta_U}(s), s = 1/2 \right) = \frac{L}{2\pi},
\end{align*}
\] (3.28)

\[
\begin{align*}
\text{res} \left( \zeta_{\Delta_U}(s), s = -\frac{(2n+1)}{2} \right) &= \text{res} \left( -\frac{(2n+1)b_{2n+1} \sin(\pi s)}{2\pi (s + (2n+1)/2)}, s = -\frac{(2n+1)}{2} \right) \\
&\Rightarrow \text{res} \left( \zeta_{\Delta_U}(s), s = -\frac{(2n+1)}{2} \right) = (-1)^n b_{2n+1} \frac{(2n+1)}{2\pi}, \quad n = 0, 1, 2, 3, \ldots
\end{align*}
\] (3.29)

Furthermore, it gives the values of \( \zeta_{\Delta_U}(s) \) at the non-positive integers:

\[
\begin{align*}
\zeta_{\Delta_U}(0) &= \frac{1}{2} \lim_{s \to 0} \frac{\sin(\pi s)}{\pi s} = 1/2, \\
\zeta_{\Delta_U}(-n) &= -nb_{2n} \lim_{s \to -n} \frac{\sin(\pi s)}{\pi (s + n)} = (-1)^n b_{2n}, \quad n = 1, 2, 3, \ldots
\end{align*}
\] (3.30) (3.31)

Given Equations (3.28)-(3.31), for any \( \Delta_U \in \mathcal{M}_F - \mathcal{M}_F(0) \) such that \( \cos(\alpha) + \cos(\beta) \neq 0 \) it is easy to compute the heat kernel coefficients using the general formulas (3.5) and (3.6). Namely, we find

\[
\begin{align*}
a_0 &= \frac{L}{2\sqrt{\pi}}, \quad a_{n+1} = -\frac{4^n n! b_{2n+1}}{(2n)! \sqrt{\pi}}, \quad n = 0, 1, 2, 3, \ldots
\end{align*}
\] (3.32)

\[
\begin{align*}
a_{1/2} &= 1/2, \quad a_{n+1/2} = -\frac{b_{2n}}{(n-1)!}, \quad n = 1, 2, 3, \ldots
\end{align*}
\] (3.33)
3.2.1. The case of $\Delta_U \in \mathcal{M}_F - \mathcal{M}_F^{(0)}$ with $\cos(\alpha) + \cos(\beta) = 0$

For this case, the appropriate function reads

$$f_U^{(B)}(ix) = e^{xL} \left[ \sin(\alpha) + \frac{1}{2x} (\cos(\beta) - \cos(\alpha)) + e^{-xL} 2n_1 \sin(\beta) + e^{-2xL} \left( \sin(\alpha) - \frac{1}{2x}(\cos(\beta) - \cos(\alpha)) \right) \right].$$

Following the same procedure as in the general case we expand, for $\alpha \neq \pi$,

$$\log(f_U^{(B)}(ix)) = \log(\sin(\alpha)) + xL + \sum_{m=1}^{\infty} c_m x^{-m} + \text{e.s.t.},$$

$$c_m = -\frac{\cot^m(\alpha)}{m}.$$  

Again, the analytical continuation of

$$\frac{\sin(\pi s)}{\pi} \int_1^{\infty} dk \cdot k^{-2s} \partial_k \log(f_U(ik))$$

provides the residues at the half integers and the values at the non-positive integers of $\zeta^{(B)}_{\Delta_U}(s)$,

$$\text{res}\left(\zeta^{(B)}_{\Delta_U}(s), s = \frac{1}{2}\right) = \frac{L}{2\pi},$$

$$\text{res}\left(\zeta^{(B)}_{\Delta_U}(s), s = \frac{2n+1}{2}\right) = (-1)^n c_{2n+1} \frac{2n+1}{2\pi}, \quad n = 0, 1, 2, 3, \ldots,$$

$$\zeta^{(B)}_{\Delta_U}(0) = 0,$$

$$\zeta^{(B)}_{\Delta_U}(-n) = (-1)^{n+1} n c_{2n}, \quad n = 1, 2, 3, \ldots$$

Once we use formulas (3.5) and (3.6), we obtain the corresponding heat kernel coefficients,

$$a_0^{(B)} = \frac{L}{2\sqrt{\pi}}, \quad a_{n+1}^{(B)} = -\frac{4^n n! c_{2n+1}}{(2n)! \sqrt{\pi}}, \quad n = 0, 1, 2, 3, \ldots,$$

$$a_{1/2}^{(B)} = 0, \quad a_{n+1/2}^{(B)} = -\frac{c_{2n}}{(n-1)!}, \quad n = 1, 2, 3, \ldots$$

Finally, the case $\alpha = \pi$, $\beta = 0$, has to be treated separately and

$$\partial_x \left( \ln f_U^{(B)}(ix) \right)|_{\alpha = \pi} = L - \frac{1}{x} + \text{e.s.t.}$$

From here,

$$\text{res}\left(\zeta^{(B)}_{\Delta_U}(s)|_{\alpha = \pi}, s = \frac{1}{2}\right) = \frac{L}{2\pi}, \quad \zeta^{(B)}_{\Delta_U}(0)|_{\alpha = \pi} = -\frac{1}{2},$$
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and

\[ a_0^{(B)} \big|_{\alpha=\pi} = \frac{L}{2\sqrt{\pi}}, \quad a_{1/2}^{(B)} \big|_{\alpha=\pi} = \frac{1}{2}, \]

(3.44)

with all other residues and relevant values respectively heat kernel coefficients equal to zero.

3.3. HEAT KERNEL COEFFICIENTS FOR \( \Delta_U \in \mathcal{M}_F^{(0)} \)

Taking into account (2.23), we can write (3.12) as

\[ f_U^{(0)}(ix) = \frac{e^{xL}}{x} \cos(\alpha) \left[ 1 + \frac{\tan(\alpha)}{x} - 2e^{-xL}\tan(\alpha) - e^{-2xL} \left( \frac{1}{x} - \tan(\alpha) \right) \right]. \]

(3.45)

Therefore,

\[ \log(f_U^{(0)}(ix)) = \log(\cos(\alpha)) + xL - \log(x) + \sum_{n=1}^{\infty} \frac{(-1)^{n+1}\tan^n(\alpha)}{n} x^{-n} + \text{e.s.t.} \]

(3.46)

\[ \Rightarrow \partial_x \log(f_U^{(0)}(ix)) = \frac{L}{x} - \sum_{n=1}^{\infty} (-1)^n \tan^n(\alpha) x^{-n-1} + \text{e.s.t.} \]

(3.47)

Hence, the analytical continuation gives as before the required residues and values of \( \zeta_{\Delta_U}^{(0)}(s) \),

\[ \text{res}(\zeta_{\Delta_U}^{(0)}(s); s = 1/2) = L/2\pi, \]

(3.48)

\[ \text{res} \left( \zeta_{\Delta_U}^{(0)}(s); s = -\frac{2n+1}{2} \right) = \frac{(-1)^n}{2\pi} \tan^{2n+1}(\alpha), \quad n = 0, 1, 2, 3, \ldots, \]

(3.49)

\[ \zeta_{\Delta_U}^{(0)}(0) = -1/2, \]

(3.50)

\[ \zeta_{\Delta_U}^{(0)}(-n) = \frac{1}{2} (-1)^n \tan^{2n}(\alpha), \quad n = 1, 2, 3, \ldots \]

(3.51)

To obtain the heat kernel coefficients, we must take into account that there is one zero mode in all cases as demonstrated previously. Therefore, we must add 1 to \( a_{1/2}^{(0)} \):

\[ a_0^{(0)} = \frac{L}{2\sqrt{\pi}}, \quad a_{1/2}^{(0)} = \frac{1}{2}, \]

(3.52)

\[ a_{n+1}^{(0)} = -\frac{4^n n! \tan^{2n+1}(\alpha)}{(2n+1)! \sqrt{\pi}}, \quad n = 0, 1, 2, 3, \ldots, \]

(3.53)

\[ a_{n+1/2}^{(0)} = \frac{1}{2} \frac{\tan^{2n}(\alpha)}{n!}, \quad n = 1, 2, 3, \ldots \]

(3.54)

The heat kernel coefficients obtained above for \( \Delta_U \in \mathcal{M}_F^{(0)} \) become singular for \( \alpha = \pi/2 \). In this case, instead

\[ \partial_x \log(f_U^{(0)}(ix)) \big|_{\alpha=\pi/2} = \frac{L}{x} + \text{e.s.t.}, \]

(3.55)
and therefore the spectral zeta function $\zeta^{(0)}_{\Delta U}(s)|_{\alpha=\pi/2}$ will only have a residue at $s=1/2$ and non-zero value at $s=0$, 
\[
\text{res}(\zeta^{(0)}_{\Delta U}(s)|_{\alpha=\pi/2}; s=1/2) = L/2\pi, \\
\zeta^{(0)}_{\Delta U}(0)|_{\alpha=\pi/2} = -1.
\] (3.56)

Hence, the only non-vanishing heat kernel coefficients are given by 
\[
a^{(0)}_{0}|_{\alpha=\pi/2} = \frac{L}{2\sqrt{\pi}}.
\] (3.58)

The Von Neumann–Krein extension The corresponding results for the VNK extension follow from 
\[
\mathcal{f}_{\text{VNK}}(k) = \frac{h_{\text{VNK}}(k)}{2i k^5 e^{i \sigma_{\text{VNK}}}} = \sin(\beta_{\text{VNK}}) (kL \sin(kL) + 2 \cos(kL) - 2).
\] (3.59)

Note, we divided by $k^5$ instead of the $k^3$ as in Equation (3.12), this being necessary because the VNK extension has two zero modes. The large $k$ expansion relevant for the heat kernel coefficients reads 
\[
\mathcal{f}_{\text{VNK}}(ik) = \frac{2 (1 - \frac{kL}{2})}{k^4 \sqrt{L^2 + 4}} e^{kL (1 + \text{e.s.t.})},
\] (3.60)

and the coefficients follow along the lines explained to be 
\[
a^{(\text{VNK})}_{0} = \frac{L}{2\sqrt{\pi}}, \quad a^{(\text{VNK})}_{n+1} = \frac{n! (4/L)^{2n+1}}{(2n+1)! 2\sqrt{\pi}}, \quad n = 0, 1, 2, 3, \ldots,
\] (3.61)

\[
a^{(\text{VNK})}_{1/2} = \frac{1}{2}, \quad a^{(\text{VNK})}_{n+1/2} = \frac{1}{n!} \left(\frac{2}{L}\right)^{2n}, \quad n = 1, 2, 3, \ldots
\] (3.62)

Let us stress, that to obtain $a^{(\text{VNK})}_{1/2}$ we have added $+2$ to $\zeta^{(0)}_{\text{VNK}}(0)$, as is requested by having two zero modes.

### 3.4. HEAT KERNEL COEFFICIENTS FOR COMMON BOUNDARY CONDITIONS.

As a check of our calculations, let us compare the results found for the heat kernel coefficients with the known ones for the most common boundary conditions.

- **Periodic boundary conditions.** The periodic boundary conditions are usually written as\(^5\)

\[
\psi(0) = \psi(L); \quad \psi'(0) = \psi'(L).
\] (3.63)

\(^5\) When it is required that the solutions of the Laplace equation are smooth functions the periodic boundary conditions are given by the condition $\psi(0) = \psi(L)$. However square integrable solutions of the Laplace equation are not necessarily smooth. Therefore the condition $\psi(0) = \psi(L)$ does not necessarily give rise to periodic boundary conditions. As an example it is worth to mention the case of Dirac delta potentials (see references [26–28] for recent developments in the interpretation of Dirac delta potentials as boundary conditions and infinitely thin kinks) where the condition $\psi(0) = \psi(L)$ is satisfied but obviously the system does not satisfy periodic boundary conditions. Therefore in order to distinguish periodic boundary conditions from other types of point interactions it is necessary to include the second condition over the derivatives: $\psi'(0) = \psi'(L)$. 

Equivalently, we can write the following two independent equations for periodic boundary conditions

\[
\begin{align*}
\psi(0) + i \psi'(0) &= \psi(L) + i \psi'(L), \\
\psi(L) - i \psi'(L) &= \psi(0) - i \psi'(0).
\end{align*}
\]

Hence, following the notation of Equation (1.3) we can write the periodic boundary conditions in the form of (1.4) as

\[
\varphi_-(\psi) = \sigma_1 \cdot \varphi_+ (\psi),
\]

being \(\sigma_1\) the corresponding Pauli matrix. Therefore, the unitary matrix that characterizes periodic boundary conditions is given by \(U_p = \sigma_1 \in \mathcal{M}_F^{(0)} \Rightarrow \alpha = \pi/2, \beta = \pm \pi/2, n_1 = \mp 1\). The heat kernel coefficients are given by (3.58).

- **Dirichlet boundary condition.** The usual form of the Dirichlet boundary condition for any manifold \(M\) with boundary \(\partial M\) is

\[
\psi|_{\partial M} = 0.
\]

As can be seen, the normal derivatives \(\partial_n \psi|_{\partial M}\) do not enter in the boundary condition. Form Equation (1.1), the general boundary condition for those unitary operators \(U \in \mathcal{M}\) such that \(1 \notin \sigma(U)\) can be written as

\[
\psi|_{\partial M} = i \frac{1 + U}{1 - U} \cdot \partial_n \psi|_{\partial M}.
\]

From this last expression, it is immediate to notice that the Dirichlet boundary condition is obtained when \(U = -I\). Therefore the Dirichlet boundary condition is given by \(U_D = -I \in \mathcal{M}_F - \mathcal{M}_F^{(0)} \Rightarrow \alpha = \pi, \beta = 0\). The heat kernel coefficients are given by (3.44).

- **Neumann boundary condition.** The usual form of the Neumann boundary condition for any manifold \(M\) with boundary \(\partial M\) is

\[
\partial_n \psi|_{\partial M} = 0,
\]

where \(\partial_n\) denotes the normal derivative to \(\partial M\). As can be seen, the boundary value \(\psi|_{\partial M}\) does not enter in the boundary condition. Form Equation (1.1), the general boundary condition for those unitary operators \(U \in \mathcal{M}\) such that \(-1 \notin \sigma(U)\) can be written as

\[
\partial_n \psi|_{\partial M} = -i \frac{1 - U}{1 + U} \cdot \psi|_{\partial M}.
\]

From this last expression, it is immediate to notice that the Neumann boundary condition is obtained when \(U = I\). Therefore the Neumann boundary condition
is given by $U_N = \mathbb{I} \in \mathcal{M}^{(0)}_F \Rightarrow \alpha = \beta = 0$. It is of note that in this case $\sin(\alpha) = 0$.

Therefore, from (3.52)–(3.54),

$$a_0^{(N)} = -\frac{L}{2\sqrt{\pi}}, \quad a_{1/2}^{(N)} = 1/2, \quad a_n^{(N)} = 0, \quad n = 1, 2, 3, \ldots$$

(3.69)

(3.70)

- **Robin boundary conditions.** The common expression for the family of Robin boundary conditions is given by (see for example reference [29])

$$\psi|_{\partial M} - g \frac{\partial \psi}{\partial n}|_{\partial M} = 0, \quad g \in (-\infty, \infty).$$

(3.71)

For the case in which the boundary manifold $\partial M$ has several disjoint components $\partial M = \bigcup_i \Omega_i$ the family of Robin boundary conditions can be written as

$$\psi|_{\Omega_i} - g_i \frac{\partial \psi}{\partial n}|_{\Omega_i} = 0, \quad g_i \in (-\infty, \infty).$$

(3.72)

The extreme values $g_i = 0, \infty$ correspond to Dirichlet and Neumann boundary conditions respectively in the ith component of $\partial M$. Note that, in the most general case the set of constants $g_i$ do not have to be the same for all the disjoint components $\Omega_i$ of $\partial M$. For $M = [0, L]$, the boundary is formed by two points and therefore it has two disjoint components. The most simple choice of Robin boundary conditions in this case is

$$-\psi'(0) = \tan\left(\frac{\alpha}{2}\right) \psi(0), \quad \psi'(L) = \tan\left(\frac{\alpha}{2}\right) \psi(L), \quad \alpha \in [0, \pi].$$

(3.73)

In a more compact notation, we can write

$$\tan\left(\frac{\alpha}{2}\right) \psi|_{\partial M} - \frac{\partial \psi}{\partial n}|_{\partial M} = 0, \quad \alpha \in [0, \pi].$$

(3.74)

Taking into account Equation (3.68) and comparing it with expression (3.74), the unitary operator $U_R$ for Robin boundary conditions satisfies the equation

$$\tan\left(\frac{\alpha}{2}\right) \mathbb{I} = -i \mathbb{I} - U_R \mathbb{I} + U_R.$$

(3.75)

Therefore, the unitary operator that characterizes the family of Robin boundary conditions given by (3.73) is given by $U_R = e^{i\alpha} \mathbb{I}$ as was firstly pointed out in references [30,31]. Note that $U_R(\alpha = 0) = \mathbb{I} = U_N$ and $U_R(\alpha = \pi) = -\mathbb{I} = U_D$.

In the parameterization (1.5), Robin boundary conditions correspond to $\beta = 0$.

For $\alpha \in (0, \pi)$ $U_R(\alpha) \in \mathcal{M}_F - \mathcal{M}^{(0)}_F$ with $\cos(\alpha) + \cos(\beta) \neq 0$. Therefore, the heat kernel coefficients for Robin boundary conditions are determined by Equations (3.32) and (3.33). From Equation (3.20), it is easy to obtain the coefficients $b_m$ for the Robin boundary conditions:

$$b_m^{(R)} = \tan^m \left(\frac{\alpha}{2}\right) \sum_{j=0}^{[m/2]} (-1)^{m-j+1} \frac{2^{m-2j} \Gamma(m-j)}{\Gamma(j+1) \Gamma(m-2j+1)}, \quad m = 1, 2, 3, \ldots$$

(3.76)
Using Equations (3.32) and (3.33), it is immediate to compute the heat kernel coefficients for Robin boundary conditions to any desired order using any symbolic calculation software. As an example, we show the first ten heat kernel coefficients:

\[ a_{0}^{(R)} = \frac{L}{2\sqrt{\pi}}, \quad a_{1/2}^{(R)} = 1/2, \quad (3.77) \]
\[ a_{1}^{(R)} = -\frac{2\tan\left(\frac{\alpha}{2}\right)}{\sqrt{\pi}}, \quad a_{2}^{(R)} = -\frac{4\tan^{3}\left(\frac{\alpha}{2}\right)}{3\sqrt{\pi}}, \quad (3.78) \]
\[ a_{3}^{(R)} = -\frac{8\tan^{5}\left(\frac{\alpha}{2}\right)}{15\sqrt{\pi}}, \quad a_{4}^{(R)} = -\frac{16\tan^{7}\left(\frac{\alpha}{2}\right)}{105\sqrt{\pi}}, \quad (3.79) \]
\[ a_{5/2}^{(R)} = \tan^{2}\left(\frac{\alpha}{2}\right), \quad a_{5/2}^{(R)} = \frac{1}{2}\tan^{4}\left(\frac{\alpha}{2}\right), \quad (3.80) \]
\[ a_{7/2}^{(R)} = \frac{1}{6}\tan^{6}\left(\frac{\alpha}{2}\right), \quad a_{9/2}^{(R)} = \frac{1}{24}\tan^{8}\left(\frac{\alpha}{2}\right). \quad (3.81) \]

These results coincide with the results obtained by Dowker in Ref. [32] (equations (10) and (14) with \( h_1 = h_2 = \tan(\alpha/2) \)). More recently Fulling has studied the heat kernel coefficients for Robin boundary conditions in the Ref. [33].

**4. The Functional Determinant of \( \Delta_U \): Derivative at \( s = 0 \) of the Spectral Zeta Function**

In this section, we compute the derivative of the zeta function at \( s = 0 \) for each of the different cases considered in Sections 3.2 and 3.3. As is well known, this derivative is a natural constituent when defining functional determinants of elliptic operators [12]. As usual, we subtract and add back a suitable number of the asymptotic \( k \to \infty \) terms of \( \partial_k \log f_\mathcal{O}(ik) \) in (3.3). In the current context, we have to subtract terms up to the order \( 1/k \) to make the integral well defined at \( k = \infty \) once \( s = 0 \) is set. As a technical tool, at the start of this analysis it is convenient to consider a massive scalar field of mass \( m \), where \( m \) will be sent to zero at a suitable point of the computation. In this way, we can avoid splitting the integral representing the zeta function into two pieces and the computation becomes a little easier. The procedure is valid as in the limit \( m \to 0 \) the zeta function for the case with vanishing mass is recovered. A presentation of (3.22) valid about \( s = 0 \) is then given by

\[
\xi_{\Delta_U}(s) = \frac{\sin \pi s}{\pi} \int_{m}^{\infty} dk (k^2 - m^2)^{-s} \partial_k \log \left[ \frac{2 f_\mathcal{O}(ik)}{k e^{kL}(\cos \alpha + \cos \beta)} \right] + \frac{\sin \pi s}{\pi} \int_{m}^{\infty} dk (k^2 - m^2)^{-s} \partial_k \log \left[ \frac{k e^{kL} \cos \alpha + \cos \beta}{2} \right].
\]

The integral in the first line by construction is analytic about \( s = 0 \) and its derivative at \( s = 0 \) is trivially computed. The needed integrals in the second line are known [34],
\[
\int_{m}^{\infty} dk (k^2 - m^2)^{-s} = \frac{m^{1-2s}}{2\sqrt{\pi}} \Gamma(1-s) \Gamma \left( s - \frac{1}{2} \right),
\]
\[
\int_{m}^{\infty} dk (k^2 - m^2)^{-s} \frac{1}{k} = \frac{m^{-2s}}{2 \sin \pi s},
\]
and
\[
\zeta'_{\Delta U}(0) = - \log \left( \frac{2 f_U(i m)}{m e^{m L} (\cos \alpha + \cos \beta)} \right),
\]
is found. As \( m \to 0 \) we use
\[
\lim_{m \to 0} f_U(i m) = L (\cos \alpha - \cos \beta) - 2 (\sin \alpha + n_1 \sin \beta)
\]
to obtain
\[
\zeta'_{\Delta U}(0) = - \log \left( \frac{2 L (\cos \alpha - \cos \beta) - 4 (\sin \alpha + n_1 \sin \beta)}{\cos \alpha + \cos \beta} \right). \tag{4.1}
\]
The case treated in Section 3.2.1, for \( \alpha \neq \pi \), follows along the same lines from
\[
\zeta^{(B)}_{\Delta U}(s) = \frac{\sin \pi s}{\pi} \int_{m}^{\infty} dk (k^2 - m^2)^{-s} \partial_k \log \left[ \frac{f^{(B)}_U(i k)}{e^{k L} \sin \alpha} \right] + \frac{\sin \pi s}{\pi} \int_{m}^{\infty} dk (k^2 - m^2)^{-s} \partial_k \log \left[ e^{k L} \sin \alpha \right].
\]
In the limit as \( m \to 0 \) we obtain
\[
\zeta^{(B)}_{\Delta U}'(0) = - \log \left( \frac{L (\cos \alpha - \cos \beta) - 2 (\sin \alpha + n_1 \sin \beta)}{\sin \alpha} \right).
\]
For \( \alpha = \pi, \beta = 0 \), instead we start with
\[
\zeta^{(B)}_{\Delta U}(s)|_{\alpha=\pi} = \frac{\sin \pi s}{\pi} \int_{m}^{\infty} dk (k^2 - m^2)^{-s} \partial_k \log \left[ \frac{f_U(i k)|_{\alpha=\pi}}{e^{k L}} \right] + \frac{\sin \pi s}{\pi} \int_{m}^{\infty} dk (k^2 - m^2)^{-s} \partial_k \log \left[ \frac{e^{k L}}{k} \right]
\]
to find
\[
\zeta^{(B)}_{\Delta U}'(0)|_{\alpha=\pi} = - \log(2L). \tag{4.2}
\]
We are left to treat the cases with a zero mode dealt with in Section 3.3. Therefore, for \( \alpha \neq \pi/2 \), the starting point is
\[ \zeta^{(0)}_\Delta U(s) = \frac{\sin \pi s}{\pi} \int_0^\infty dk (k^2 - m^2)^{-s} \partial_k \log \left[ \frac{f_U^{(0)}(ik)}{e^{kL} \cos \alpha} \right] \]

\[ + \frac{\sin \pi s}{\pi} \int_0^\infty dk (k^2 - m^2)^{-s} \partial_k \log \left[ \frac{e^{kL} \cos \alpha}{k} \right] \]

leading to

\[ \zeta^{(0)}_\Delta U'(0) = -\log \left| \frac{L(2 \cos \alpha + L \sin \alpha)}{\cos \alpha} \right|. \]

For \( \alpha = \pi/2 \) instead

\[ \zeta^{(0)}_\Delta U(s) \big|_{\alpha=\pi/2} = \frac{\sin \pi s}{\pi} \int_0^\infty dk (k^2 - m^2)^{-s} \partial_k \log \left[ \frac{f_U^{(0)}(ik)\big|_{\alpha=\pi/2} k^2}{e^{kL}} \right] \]

\[ + \frac{\sin \pi s}{\pi} \int_0^\infty dk (k^2 - m^2)^{-s} \partial_k \log \left[ \frac{e^{kL}}{k^2} \right], \]

leading to

\[ \zeta^{(0)}_\Delta U'(0) = -2 \log L. \]

The expression obtained for \( \zeta^{(0)}_\Delta U'(0) \) does not allow one to compute the derivative of the spectral zeta function at \( s = 0 \) for the VNK extension by just replacing \( \{ \alpha, \beta \} \mapsto \{ \alpha_{\text{VNK}}, \beta_{\text{VNK}} \} \) as it does produce an undefined answer. The reason is that there are two zero modes and the formulas have to be adapted; see Equations (3.59) and (3.60). From the large \( k \) expansion (3.60) and from \( f_{\text{VNK}}(im) \sim -L^4/(6L^2 + 4) + O(m) \) the computation explained above leads to the following expression for \( \zeta^{(0)}_{\text{VNK}}'(0) \):

\[ \zeta^{(0)}_{\text{VNK}}'(0) = -\log \left| \frac{f_{\text{VNK}}(im)\big|_{m \to 0}}{-L/\sqrt{L^2 + 4} + 4} \right| = -\log \left( \frac{L^3}{6} \right). \]

These results can be confronted with the easily computed answers for periodic, Dirichlet and Neumann boundary conditions.

For Dirichlet boundary conditions, the spectrum is \( \lambda_n = (\pi n/L)^2, \ n \in \mathbb{N} \), with associated zeta function \( \zeta_{\text{Dir}}(s) = (\pi/L)^{-2s} \zeta_{\text{R}}(2s) \). This gives \( \zeta_{\text{Dir}}'(0) = -\log(2L) \) in agreement with (4.2).

For Neumann boundary conditions, the spectrum is as above but with zero included. For the determinant, the answers therefore again reads \( \zeta_{\text{Neu}}'(0) = -\log(2L) \), which agrees with (4.3), once \( \alpha = \beta = 0 \) has been put.

Finally, for periodic boundary conditions the spectrum is \( \lambda_n = (2\pi n/L)^2, \ n \in \mathbb{Z} \), with associated zeta function (zero mode excluded) \( \zeta_{\text{per}}(s) = 2(2\pi/L)^{-2s} \zeta_{\text{R}}(2s) \). This shows \( \zeta_{\text{per}}'(0) = -2 \log L \), again in agreement with (4.3).
As a specific new result, Robin boundary conditions as described above follow from (4.1) as

\[ \zeta_{\Delta_{UR}}(0) = -\log \left( 2 \tan \left( \frac{\alpha}{2} \right) \left( L \tan \left( \frac{\alpha}{2} \right) + 2 \right) \right). \]

5. Conclusions

In this article, we have analyzed the spectral zeta function resulting from the Laplacian on the interval \([0, L]\) for the case when strongly consistent selfadjoint extensions and the Von Neumann–Krein extension are applied. Contour integral representations for the zeta functions are obtained for this class of selfadjoint extensions. These are used to compute leading heat kernel coefficients and the functional determinant in this context. Our results agree with known results for standard boundary conditions like Dirichlet, Neumann and periodic. The generalization of these results to a scalar quantum field theory in \(D + 1\) spacetime confined between two \(D - 1\) dimensional plane parallel plates is straightforward for the heat kernel coefficients due to the factorization properties of the heat kernel in the same way as it is done in Ref. [1].

The current article represents the start of further investigations into the details of heat kernel coefficients. Heat kernel coefficients are usually represented in terms of geometric invariants with universal multipliers depending on the boundary condition. The question arises how the multipliers depend on the chosen selfadjoint extension. To get some non-trivial boundary geometry involved a similar computation should be done for balls along the lines of [35–37], where choosing general selfadjoint extensions will lead to different combinations of Bessel functions. Furthermore, following [38], surfaces of revolution are possible candidates to analyze how different selfadjoint extensions impact spectral functions.

Finally, the presented analysis could also be done for selfadjoint extensions that allow for finitely many negative eigenvalues by using a variation of the current procedure [39]. (Note, that the results presented in this article actually remain valid beyond strongly consistent selfadjoint extensions as long as the eigenvalues are positive!) We believe that this kind of selfadjoint extensions could provide a natural mechanism for inflation in cosmological models with compact extra dimensions with boundary (Kaluza–Klein cosmology or RS-type scenarios), where the dark energy is interpreted as the quantum vacuum of a fundamental scalar. In this scenario, the inflationary phase is produced by the existence of negative energy modes the existence of which is strongly dependent on the size of the compact extra dimension with boundary.

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Appendix A: Some Remarks on the AIM Formalism

For completeness, we would like to introduce some basic notions about the AIM formalism. For simplicity, we will restrict ourselves to the case of the selfadjoint extensions of the Laplace operator over the interval \([0, L]\). The AIM formalism establishes a one-to-one correspondence between the abstract set \(\mathcal{M}(L)\) of selfadjoint extensions of the Laplace operator over the finite line \([0, L]\) and the unitary group \(U(2)\)

\[ s_L : U(2) \longrightarrow \mathcal{M}(L). \]

The isomorphism is unique for each value of \(L > 0\). The physical realization of an abstract selfadjoint extension in \(\mathcal{M}(L)\) is a boundary condition that determines the dynamics of the free particle on the interval \([0, L]\) together with the equation of motion given by the Laplace operator. Therefore, an abstract selfadjoint extension acquires a physical meaning as a unitary operator \(s_L^{-1}(\Delta_U)\) defining a boundary condition. Hence from a physical point of view what is physically meaningful is not the abstract set \(\mathcal{M}(L)\) but its inverse image through \(s_L\), i.e., \(s_L^{-1}(\mathcal{M}(L)) \simeq U(2)\). With this image in mind, and knowing that the length \(L\) of the interval plays a crucial role in physical problems such as the Casimir effect (see for example [1]), one must assume that the basic object is the group \(U(2)\). In this picture, the space \(\mathcal{M}_F(L)\) is defined as

\[ \mathcal{M}_F(L) \equiv \{ U \in U(2) \mid q_L(U) \in \mathcal{M}_{NN}(L') \ \forall \ L' > 0 \}, \tag{A.1} \]

where \(\mathcal{M}_{NN}(L) \subset \mathcal{M}(L)\) is the subset of non-negative selfadjoint extensions for a given length of the interval \(L\). In this sense, the set \(\mathcal{M}_F\) as a set of boundary conditions is independent of the length of the interval, and it is stable under \(s_L^{-1}\) for any \(L > 0\). Keeping boundary conditions defined by \(U(2)\) matrices as the basic objects, statements concerning all values of \(L\) or particular values of \(L\) made throughout this paper can be reformulated in terms of the isomorphisms \(s_L\). In particular, concerning the VNK extension, one should say that “\(s^{-1}_L(\Delta_{\text{VNK}}) \notin \mathcal{M}_F\) for any \(L > 0\)”.

A more geometric picture can be obtained by constructing a bundle structure over the set of selfadjoint extensions of the Laplace operator over the finite interval. The base space is given by the positive real numbers \(\mathbb{R}^+\), the total space will be given by \(E = \{ \mathcal{M}(L) \}_{L \in \mathbb{R}^+}\)

\[ \pi : E \longrightarrow \mathbb{R}^+, \]

\[ 6 \text{ In this appendix, we will make a small change of notation in that we specify the } L \text{ dependence of the space of selfadjoint extensions of the Laplace operator over the finite interval } [0, L]. \text{ Of course for a pair of lengths } L_1, L_2 > 0, \mathcal{M}(L_1) \simeq \mathcal{M}(L_2) \equiv \mathcal{M} \text{ but it is important for the moment to keep this dependence in mind.} \]
and the fiber is given by \( F = \pi^{-1}(L) = \mathcal{M}(L) \simeq U(2) \). In this picture, the abstract selfadjoint extensions \( \mathcal{M} \) are nothing but sections of the fiber bundle just introduced. Obviously, this bundle has a natural structure of principal bundle with structural group \( U(2) \) acting on \( E \) naturally through the group isomorphisms \( s_L \). The VNK section would be the only section \( w_{\text{VNK}} : \mathbb{R}^+ \to E \) that for every value \( L \in \mathbb{R}^+ \) gives rise to a selfadjoint extension with two zero modes, and the space \( \mathcal{M}_F \) should be defined as those constant sections (the sections such that \( w(L) = (L, U) \) with \( U \in U(2) \) independent of \( L \)) whose image is in \( \{ \mathcal{M}_{NN}(L) \}_{L \in \mathbb{R}^+} \).

References

1. Asorey, M., Munoz-Castaneda, J.M.: Attractive and repulsive Casimir vacuum energy with general boundary conditions. Nucl. Phys. B 874, 852–876 (2013)
2. Asorey, M., Garcia-Alvarez, D., Munoz-Castaneda, J.M.: Vacuum energy and renormalization on the edge. J. Phys. A 40, 6767–6776 (2007)
3. Zettl, A.: Sturm–Liouville theory, mathematical surveys and monographs. American Mathematical Society, Providence (2005)
4. Munoz-Castaneda, J.M.: Boundary effects in quantum field theory. PhD dissertation (in spanish), Zaragoza Univ. (2009)
5. Pérez-Pardo, J.M.: On the theory of self-adjoint extensions of the Laplace–Beltrami operator, quadratic forms and symmetry. ArXiv e-prints (2013)
6. Asorey, M., Ibort, A., Marmo, G.: Global theory of quantum boundary conditions and topology change. Int. J. Mod. Phys. A 20, 1001–1026 (2005)
7. Elizalde, E., Odintsov, S.D., Romeo, A., Bytsenko, A.A., Zerbini, S.: Zeta regularization techniques with applications. World Scientific, Singapore (1994)
8. Gilkey, P.B.: Invariance theory, the heat equation and the Atiyah–Singer index theorem. CRC Press, Boca Raton (1995)
9. Gilkey, P.B.: Asymptotic formulae in spectral geometry. Chapman & Hall/CRC, Boca Raton (2004)
10. Kirsten, K.: Spectral functions in mathematics and physics. Chapman & Hall/CRC, Boca Raton (2002)
11. Vassilevich, D.V.: Heat kernel expansion: user’s manual. Phys. Rep. 388, 279–360 (2003)
12. Ray, D.B., Singer, I.M.: R-torsion and the Laplacian on Riemannian manifolds. Adv. Math. 7, 145–210 (1971)
13. Blau, S., Visser, M., Wipf, A.: Zeta functions and the Casimir energy. Nucl. Phys. B 310, 163 (1988)
14. Bordag, M., Klimchitskaya, G.L., Mohideen, U., Mostepanenko, V.M.: Advances in the Casimir effect. Oxford University Press, New York (2009)
15. Bordag, M., Mohideen, U., Mostepanenko, V.M.: New developments in the Casimir effect. Phys. Rep. 353, 1–205 (2001)
16. Bytsenko, A.A., Cognola, G., Vanzo, L., Zerbini, S.: Quantum fields and extended objects in space-times with constant curvature spatial section. Phys. Rep. 266, 1–126 (1996)
17. Dowker, J.S., Critchley, R.: Effective Lagrangian and energy momentum tensor in de Sitter space. Phys. Rev. D 13, 3224 (1976)
18. Dowker, J.S., Kennedy, G.: Finite temperature and boundary effects in static space-times. J. Phys. A 11, 895 (1978)
19. Hawking, S.W.: Zeta function regularization of path integrals in curved space-time. Commun. Math. Phys. 55, 133 (1977)
20. Milton, K.A.: The Casimir effect: physical manifestations of zero-point. World Scientific, River Edge (2001)
21. Alonso, A., Simon, B.: The Birman–Krein–Vishik theory of selfadjoint extensions of semibounded operators. J. Oper. Theory 4(2), 251–270 (1980)
22. Ashbaugh, M.S., Gesztesy, F., Mitrea, M., Teschl, G.: Spectral theory for perturbed krein laplacians in nonsmooth domains. Adv. Math. 223(4), 1372–1467 (2010)
23. Ashbaugh, M.S., Gesztesy, F., Mitrea, M., Shterenberg, R., Teschl, G.: A survey on the Krein–von Neumann extension, the corresponding abstract buckling problem, and Weyl-type spectral asymptotics for perturbed Krein Laplacians in nonsmooth domains. ArXiv e-prints (2012)
24. Kirsten, K., McKane, A.J.: Functional determinants by contour integration methods. Ann. Phys. 308, 502–527 (2003)
25. Seeley, R.T.: Complex powers of an elliptic operator, singular integrals, Chicago 1966. Proc. Symp. Pure Math. 10, 288–307 (1968)
26. Bordag, M., Munoz-Castaneda, J.M.: Quantum vacuum interaction between two sine-Gordon kinks. J. Phys. A 45, 374012 (2012)
27. Mateos Guilarte, J., Munoz-Castaneda, J.M.: Double-delta potentials: one dimensional scattering. The Casimir effect and kink fluctuations. Int. J. Theor. Phys. 50, 2227–2241 (2011)
28. Munoz-Castaneda, J.M., Mateos Guilarte, J., Moreno Mosquera, A.: Quantum vacuum energies and Casimir forces between partially transparent δ-function plates. Phys. Rev. D 87(10), 105020 (2013)
29. Romeo, A., Saharian, A.A.: Vacuum densities and zero point energy for fields obeying Robin conditions on cylindrical surfaces. Phys. Rev. D 63, 105019 (2001)
30. Asorey, M., Garcia-Alvarez, D., Munoz-Castaneda, J.M.: Casimir effect and global theory of boundary conditions. J. Phys. A 39, 6127–6136 (2006)
31. Asorey, M., Munoz-Castaneda, J.M.: Vacuum boundary effects. J. Phys. A 41, 304004 (2008)
32. Dowker, J.S.: Robin conditions on the Euclidean ball. Class. Quantum Gravity 13, 585–610 (1996)
33. Stephen, A.: Fulling. Local spectral density and vacuum energy near a quantum graph vertex (2005). arXiv:math/0508335
34. Gradshteyn, I.S., Ryzhik, I.M.: Table of integrals, series, and products. Academic Press, New York (1965)
35. Bordag, M., Elizalde, E., Kirsten, K.: Heat kernel coefficients of the Laplace operator on the D-dimensional ball. J. Math. Phys. 37, 895–916 (1996)
36. Bordag, M., Geyer, B., Kirsten, K., Elizalde, E.: Zeta function determinant of the Laplace operator on the D-dimensional ball. Commun. Math. Phys. 179, 215–234 (1996)
37. Bordag, M., Kirsten, K., Dowker, J.S.: Heat kernels and functional determinants on the generalized cone. Commun. Math. Phys. 182, 371–394 (1996)
38. Jeffres, T.D., Kirsten, K., Lu, T.: Zeta function on surfaces of revolution. J. Phys. A 45(34), 345201 (2012)
39. Kirsten, K., McKane, A.J.: Functional determinants for general Sturm–Liouville problems. J. Phys. A 37, 4649–4670 (2004)