I. INTRODUCTION

Standard quantum theory is readily applied to measurement situations. This requires additional (ad hoc— for this case only) information be supplied for each situation. One may hope the theory could be extended to describe reality independent of experiment and without the need for ad hoc information. Given a state vector for any physical system, no matter how large or complex, one may hope for a theory that specifies the state vectors corresponding to the possible realizable states of nature, and their probabilities of realization.

Why can’t this be done with standard quantum theory? Given a state vector,

$$\psi, t = \sum |a_n\rangle \langle a_n| \psi, t \rangle,$$  

one would like to say that the $|a_n\rangle$ correspond to the possible realizable states of nature, and that $|\langle a_n| \psi, t \rangle|^2$ are their probabilities of realization.

There are two problems with this. The first is the preferred basis problem. Why $|a_n\rangle$ and not another orthonormal basis $|b_n\rangle$? No one has been able to specify the needed preferred basis. Of course, in experimental situations and special model situations this is possible, but it has not been possible in general.

The second I like to call the hopping problem. If $|\langle a_n| \psi, t \rangle|^2$ is the probability of $|a_n\rangle$ being realized in nature at time $t$, and $|\langle a_m| \psi, t \rangle|^2$ is the probability of $|a_m\rangle$ ($m \neq n$) being realized in nature at time $t + dt$, then

$$|\langle a_n| \psi, t \rangle|^2 |\langle a_m| \psi, t + dt \rangle|^2$$

is the probability of occurrence of both these events which, of course, is not seen. The difficulty is that quantum theory just gives the probabilities of these events: it does not give the probability of transitions between these events.

The resolution chosen by the Founding Fathers was to restrict quantum theory to experimental situations, with empirically defined preferred bases, and to adopt the ‘collapse postulate,’ to ensure that a preferred basis state, once chosen by nature, would remain chosen. They thereby relinquished the hope to describe reality.

An alternative resolution, described here, does not give up on this hope. It alters quantum theory so that it may be realized. The Schrödinger equation is modified by adding a randomly fluctuating term to the Hamiltonian, to account for the probabilistic behavior of nature. There is a preferred basis built into this term. The state vector dynamically collapses toward one of these basis states, very slowly for micro-objects, very rapidly for macro-objects.

The preferred basis is essentially the mass density basis. This choice could not be made for standard quantum theory, since collapse to position eigenstates gives the particles infinite energy. It works here because the collapse never goes all the way. First, it would take infinite time for that to occur. Second, it doesn’t happen because of the interaction of the usual Hamiltonian dynamics with the collapse dynamics (e.g., the example in Section VII).

II. DERIVING CSL

For collapse dynamics, one wants the following behavior.

Given the state vector Eq. (11) at time $t = 0$, the squared amplitudes $x_n(t) \equiv |\langle a_n| \psi, t \rangle|^2$ should fluctuate until eventually one amplitude becomes equal to 1, and the rest equal to 0: that is the collapse.

Moreover, with repeated evolutions, the nth amplitude should eventually reach 1 for a fraction $x_n(0)$ of the evolutions: that is the Born rule.

A. Gambler’s Ruin Game

There is a rather precise analogy to this behavior, the ‘gambler’s ruin game.’ Consider two gamblers, with $\$100 between them. Gambler 1 starts with $\$X_1(0)$, gambler 2 with $\$X_2(0) = 100 - X_1(0)$. They toss a fair coin: ‘fair’ is crucial, making the game what mathematicians call a ‘Martingale.’ Heads, gambler 1 gives a dollar to gambler 2, tails, the reverse. The amount each possesses, $X_n(t_k)$, fluctuates. Eventually, the game ends, with one gambler in possession of all the money.
Define $P(X_1)$ as the conditional probability that gambler 1 eventually wins, given that he has $\$X_1$. Then,

$$P(X_1) = \frac{1}{2}P(X_1 - 1) + \frac{1}{2}P(X_1 + 1).$$

That is, when the coin is tossed, gambler 1 can either lose the toss but win from there, or win the toss and win from there. The solution of this difference equation is $P(X_1) = AX_1 + B$ (A and $B$ are constants). With boundary conditions $P(0) = 0$ and $P(100) = 1$, the solution is $P(X_1) = X_1/100$.

Define $x_n(t_k) = X_n(t_k)/100$. Thus, we have seen that if gambler 1 starts out with a fraction $x_1(0)$ of the total amount of money, that is the fraction of repeated games he wins.

This is just the behavior we want for a state vector which is the sum of two basis vectors $|a_1\rangle$, $|a_2\rangle$. Thus, the two basis vectors may be thought of as competing in a continuous version of the gambler’s ruin game.

For, the gambler’s ruin game is a zero-sum, discrete, fair, random walk with absorbing barriers (at $x = 0, 1$) in discrete time, for $x_n(t_k) \equiv X_n(t_k)/100$.

Collapse is a zero-sum, continuous, fair, random walk with absorbing barriers in continuous time, for $x_n(t) \equiv \left(|a_n\rangle\langle\phi|, t\right)^2$.

The continuous limit of discrete random walk is Brownian motion. Thus, it is natural to regard the $x_n(t)$ as undergoing some kind of Brownian motion. We shall replace the probabilities associated with a single coin toss, and with a sequence of coin tosses respectively by

$$P(dB) = \frac{1}{\sqrt{2\pi\lambda}}e^{-dB^2/2\lambda dt}, \quad (2a)$$

$$P(B(t)) = \frac{1}{\sqrt{2\pi\lambda}}e^{-B(t)^2/2\lambda t}, \quad (2b)$$

where $\lambda$ is a constant diffusion rate. Just as the gambler’s ruin dollar count $X_n(t_k)$ depends upon a sequence of coin tosses, so we shall take the collapse $x_n(t)$ to depend upon $B(t)$.

B. Postulates

We shall derive the CSL collapse dynamics from two postulates:

1) gambler’s ruin behavior.
2) a linear, real, Schrödinger equation.

2) could use some explanatory remarks.

In order to achieve the linearity part of 2), it shall be necessary to relinquish the condition that the state vector norm is 1. In standard quantum theory, the unit norm condition is mandatory, since it ensures that probabilities add up to 1. Here there is no need for the unit norm condition: probabilities are provided by $B(t)$’s probabilities. Physical information is carried by a state vector’s direction in Hilbert space, not its norm. To calculate expectation values, one can always normalize a state vector. The un-normalized state vector shall be denoted $|\phi, t\rangle$.

Regarding the real part of 2), we wish to consider a different dynamics from the usual Hamiltonian dynamics $d|\psi, t\rangle = -iHdt|\psi, t\rangle$. That utilizes a hermitian Hamiltonian $H$, so we shall consider dynamics of the form $d|\psi, t\rangle = H_{\text{c}}dt|\psi, t\rangle$ where the ‘collapse Hamiltonian’ $H_{\text{c}}$ is hermitian (we note that the most general linear equation is of the form $d|\psi, t\rangle = -(H + iH_{\text{c}})dt|\psi, t\rangle$).

Now, for the purposes of this derivation, for simplicity and appropriate to the generalization of a classical game, we shall further restrict the state vector components to be real numbers. Therefore, we must restrict $H_{\text{c}}$ to be a real symmetric operator in the chosen basis $|a_n\rangle$. Once we find the unique form of $H_{\text{c}}$ that allows postulate 1), it may readily be seen that that complex state vector components and the most general hermitian $H_{\text{c}}$ also allow postulate 1). Because the state vector components are real, we may write

$$x_n(t) = \frac{\langle a_n|\phi, t\rangle^2}{\langle\phi, t|\phi, t\rangle} = \frac{\langle a_n|\phi, t\rangle^2}{\sum_m\langle a_m|\phi, t\rangle^2}. \quad (3)$$

Also in order to achieve the linearity part of 2), the Schrödinger equation shall be linear in a Brownian motion $B'(t)$. However, the relation assumed between $B'(t)$ and $B(t)$ is allowed to be non-linear in the state vector:

$$dB'(t) = dB(t) + f(x)dt. \quad (4)$$

where $f$ is an arbitrary real function of the $x_n(t)$. For, collapse violates the superposition principle, and this somehow requires a non-linearity. What 2) really means, then, is that we look to have a linear Schrödinger equation and isolate all the non-linearity in the probability. [2]

The derivation, which is presented in the next few sections, is rather lengthy, and it involves stochastic differential equations–but the result does not. Some readers may wish to move immediately to section [11] where the results obtained here are utilized.

Postulate 1) is implemented by an Itô equation for $x_n(t)$:

$$dx_n(t) = b_n(x)dB(t), \quad (5)$$

where $b_n$ is an arbitrary real function of the $x_m(t)$’s.

We shall denote by an overline the ensemble average of a quantity, e.g., from Eq. [2a], $\overline{dB} = \int d(dB)P(dB) = 0$.

It immediately follows from Eq. [5] that $\overline{dx_n(t)} = 0$. This says that it is a ‘fair game’ for each ‘player’ which, as we have said, is crucial for gambler’s ruin behavior. [3]

This is why Eq. [5] is chosen to be an Itô equation.

The other necessary condition for gambler’s ruin behavior is the end-game condition, that $b_n(x) = 0$ when one $x_n(t)$ is equal to 1 and the rest vanishes. However, that does not need to be separately imposed since it automatically occurs, as shall be seen.
Postulate 2) is implemented by the Stratonovich Schrödinger equation for the un-normalized state vector amplitudes:

$$d|\phi, t\rangle = [RdB' + Sdt]|\phi, t\rangle$$

where $R$ and $S$ are arbitrary symmetric real operators, as discussed above. This is chosen to be a Stratonovich equation because the calculus manipulations (e.g., derivative of the product of functions) are the usual ones, which would not be the case for an Itô equation.

However, a rather tedious calculation (Appendix A) shows that $R$ and $S$ have to be diagonal in the $|\alpha_n\rangle$ basis if postulate 2) is to imply postulate 1). Therefore, we shall write this equation as

$$d(a_n|\phi, t\rangle) = [a_n dB' + \beta_n dt]|a_n|\phi, t\rangle$$

(6)

where $\alpha_n, \beta_n$ are real constants.

C. Derivation

We proceed to find $dx_n(t)$ from Eq. (4). With use of Eqs. (4), (6), we obtain the Stratonovich equation

$$dx_n(t) = 2\{(\alpha_n - \alpha \cdot x)dB + \lambda dB^2\} + 2\{(\alpha_n - \alpha \cdot x)dB + \lambda dB^2\} |dx_n(t)\rangle$$

(7)

where $\alpha \cdot x \equiv \sum m \alpha_m x_m(t)$. Now we may use the rule for converting a Stratonovich equation to an Itô equation, which in this case means adding

$$\frac{\lambda dt}{2} \sum m 2\{(\alpha_n - \alpha \cdot x)\frac{\partial}{\partial x_m} 2\{(\alpha_n - \alpha \cdot x)$$

to the right side of Eq. (7). The result is

$$dx_n(t) = 2\{(\alpha_n - \alpha \cdot x)dB + \lambda dB^2\} + 2\{(\alpha_n - \alpha \cdot x)dB + \lambda dB^2\} |dx_n(t)\rangle$$

(8)

where $\lambda dB^2 \equiv \sum m \alpha_m x_m^2(t)$. In order that the Itô Eq. (8) (consequence of postulate 2) agree with the Itô Eq. (5) (consequence of postulate 1), the coefficient of $dB$ and the coefficient of $dt$ in both equations must be equal, so

$$b_n(x) = 2(\alpha_n - \alpha \cdot x)x_n(t), \quad (9a)$$

$$\lambda = \sum m \alpha_m^2 x_m^2(t)$$

(9b)

To find $f$ and $\beta_n$, operate on Eq. (9b) with $\sum x_n \partial^2/\partial x_m^2$. Remembering that $\sum_n x_n = 1$, we obtain

$$2\alpha_n \frac{\partial}{\partial x_m} f = 4\lambda \alpha_m^2, \quad \text{or} \quad f = 2\lambda \alpha \cdot x + c.$$
has these two moments.

Define $N = \sum t/dt$. We multiply Eq. (15) by itself $N + 1$ times with successively smaller values of $t$, obtaining the joint probability of the independent increments $d\mathbf{B}$ at successive values of $t$:

$$
\prod_{n=0}^{N} P[d\mathbf{B}'(t - ndt)] = \prod_{n=0}^{N} \frac{1}{\sqrt{2\pi \lambda dt}} e^{-\frac{(dB'(t-ndt))^2}{2\lambda dt}} \\
\cdot \frac{\langle \phi, t + dt | \phi, t + dt \rangle}{\langle \phi, 0 | \phi, 0 \rangle}.
$$

(16)

Since $\langle \phi, 0 | \phi, 0 \rangle = 1$, and using Eq. (13) to write

$$
\langle \phi, t + dt | \phi, t + dt \rangle = \langle \phi, 0 | e^{2\lambda t (t + dt)} - 2A^2 (t + dt) | \phi, 0 \rangle
$$

(17)

(taking $B(0) = 0$), it follows from Eqs. (10), (17) that the joint probability of the values of $\mathbf{B}$ at successive values of $t$ is

$$
P[B'(t + dt), B'(t), ... B'(dt)] = \prod_{n=0}^{N} \frac{1}{\sqrt{2\pi \lambda dt}} \cdot \langle \phi, 0 | e^{-\frac{(B'(t + dt - ndt) - B'(t - ndt))^2}{2\lambda (t - ndt)}} | \phi, 0 \rangle.
$$

(18)

We have written $d\mathbf{B}'(t - ndt) = B'(t + dt - ndt) - B'(t - ndt)$. We have also written $P$ as the joint probability of $\{B'(t + dt), B'(t), ... B'(dt)\}$, instead of the joint probability of $\{B'(t), B'(t-ndt), ... dB'(0)'\}$, which we can do since the Jacobian determinant for the change of variables is 1.

Now, we are interested in finding the probability $P[B'(t + dt)]$, regardless of what Brownian path leads to the value of $B'(t + dt)$. To obtain this, we integrate Eq. (13) over all $B$’s except $B'(t + dt)$. These integrals are easily done, since each $B'(ndt)$ appears in just two (adjacent) gaussians in the product, and

$$
\int_{-\infty}^{\infty} dB \frac{1}{\sqrt{2\pi c_1}} e^{-\frac{(B-a_1)^2}{2c_1}} \cdot \frac{1}{\sqrt{2\pi c_2}} e^{-\frac{(B-a_2)^2}{2c_2}} = \frac{1}{\sqrt{2\pi (c_1 + c_2)}} e^{-\frac{(a_1-a_2)^2}{2(c_1+c_2)}}.
$$

The result for $P[B'(t + dt)]$ is

$$
P = \frac{1}{\sqrt{2\pi \lambda (t + dt)}} \langle \phi, 0 | e^{-\frac{(B'(t+dt)-2A^2(t+dt)\mathbf{A})^2}{2\lambda (t+dt)}} | \phi, 0 \rangle
$$

(19)

$$
= \frac{1}{\sqrt{2\pi \lambda (t + dt)}} e^{-\frac{(B'(t+dt))^2}{2\lambda (t+dt)}} \langle \phi, t + dt | \phi, t + dt \rangle.
$$

III. CSL

We have arrived at CSL’s two equations in the form of the Schrödinger equation (19) and the Probability Rule (19). It is useful to incorporate the exponential factor in (19) in the Schrödinger equation (and un-prime $B$, and replace $t + dt$ by $t$), so that its solution then becomes

$$
|\phi, t\rangle = e^{-\frac{1}{\sqrt{2\lambda t}} [B(t)-2\lambda t a_n]^2} |\phi, 0\rangle,
$$

(20)

leaving the Probability Rule in the simple form

$$
P[B(t)]dB(t) = \frac{dB(t)}{\sqrt{2\pi \lambda t}} \langle \phi, t | \phi, t \rangle.
$$

(21)

Eqs. (20), (21) define CSL. Everything that follows is based upon these two equations.

A. Collapse mechanism

Let’s see how they contrive to collapse the state vector. Suppose the initial state vector is

$$
|\phi, 0\rangle = \sum_n c_n |a_n\rangle.
$$

(22)

$|\phi, 0\rangle$ is assumed normalized to 1, so $\sum_n |c_n|^2 = 1$. Eqs. (20), (21) then become

$$
|\phi, t\rangle = \sum_n c_n |a_n\rangle e^{-\frac{1}{\sqrt{2\lambda t}} [B(t)-2\lambda t a_n]^2}
$$

(23)

and

$$
P[B(t)]dB(t) = \frac{dB(t)}{\sqrt{2\pi \lambda t}} \sum_n |c_n|^2 e^{-\frac{1}{\sqrt{2\lambda t}} [B(t)-2\lambda t a_n]^2}.
$$

(24)

Now, assume all the $a_n$ are unequal. Then, Eq. (24) describes a bunch of gaussians whose centers at $2\lambda t a_n$ drift further and further apart, while their widths $\sqrt{\lambda t}$ spread much more slowly. As $t$ increases, the gaussians have less and less overlap.

Then, for the set of $B(t)$’s which lie within the $m$th gaussian, the state vector and probability are, to an excellent approximation (which becomes exact for $t \to \infty$),

$$
|\phi, t\rangle \approx c_m |a_m\rangle e^{-\frac{1}{\sqrt{2\lambda \infty}} [B(t)-2\lambda t a_m]^2}
$$

(25)

and

$$
P[B(t)]dB(t) \approx \frac{dB(t)}{\sqrt{2\pi \lambda \infty}} |c_m|^2 e^{-\frac{1}{\sqrt{2\lambda \infty}} [B(t)-2\lambda t a_m]^2}.
$$

(26)

For this set of $B(t)$’s, the integrated probability in (26) is $|c_m|^2$, giving the Born Rule.

For any $B(t)$ in this set, the normalized state vector is

$$
|\psi, t\rangle \equiv \frac{|\phi, t\rangle}{\sqrt{\langle \phi, t | \phi, t \rangle}} \approx |a_m\rangle
$$

(27)

giving the collapsed state vector.
B. Refinements: Density Matrix

The density matrix, as usual, is constructed from the state vectors and their associated probabilities, here given by Eqs. (20), (21):

\[
\rho(t) = \int_{-\infty}^{\infty} P[B(t)] dB(t) \langle \phi, t | | \phi, t \rangle \\
= \int_{-\infty}^{\infty} dB(t) \frac{dB(t)}{\sqrt{2\pi} \lambda} \langle \phi, t | | \phi, t \rangle \\
= \int_{-\infty}^{\infty} dB(t) e^{-\frac{1}{\lambda}[B(t) - 2\lambda t A]^2} | \phi, 0 \langle | e^{-\frac{1}{\lambda}[B(t) - 2\lambda t A]^2} | \phi, 0 \rangle.
\]

In Eq. (28), \( A_L, A_R \) mean that these operators act on the left or the right respectively of the initial density matrix \( | \phi, 0 \rangle \).

With the initial density matrix (22), the matrix elements of the density matrix at time \( t \) is found from (28) to be

\[
(a_n | \rho(t) | a_m) = c_n c_m^* e^{-\frac{\lambda}{2}[a_n - a_m]^2},
\]

showing how the off-diagonal elements decay while the diagonal elements remain constant, the collapse rate increasing as the eigenvalue differences increase.

According to Eq. (28), the differential equation satisfied by the density matrix is

\[
\frac{d\rho(t)}{dt} = -\frac{\lambda}{2} [A, [A, \rho(t)]]
\]

which is the simplest possible Lindblad equation.

C. Refinements: Hamiltonian

To add the Hamiltonian to the state vector dynamics, consider the evolution over an infinitesimal time interval: it and the probability rule become

\[
| \phi, t \rangle = e^{-iHdt - \frac{1}{2\lambda}[dB(t) - 2\lambda dt A]^2} | \phi, t - dt \rangle \\
= e^{-dt[iH + \frac{1}{\lambda}[w(t) - 2\lambda A]^2]} | \phi, t - dt \rangle
\]

\[
P(w) dw = \langle \phi, t | | \phi, t \rangle \frac{d\rho(t)}{dt} = -\frac{\lambda}{2} [A, [A, \rho(t)]]
\]

where \( w(t) \equiv dB(t)/dt \) is called white noise.

Over a finite time interval, Eqs. (31a), (31b) imply

\[
| \phi, t \rangle = T e^{-\int_{0}^{t} dt'[iH(t') + \frac{1}{\lambda}[w(t') - 2\lambda A]^2]} | \phi, 0 \rangle
\]

\[
P(w) Dw = \langle \phi, t | | \phi, t \rangle \prod_{t'=0}^{t} \frac{d\rho(t')}{dt'}
\]

where \( T \) is the time-ordering operator.

To summarize, for each white noise function \( w(t) \) there is a corresponding state vector \( | \phi, t \rangle \) given by Eq. (32a), one of which is supposed to be realized in nature with probability (32b). Generally, the hamiltonian evolution and the collapse-hamiltonian evolution compete against each other. This can give rise to effects which suggest experimental tests of the collapse theory vis-à-vis standard quantum theory/collapse postulate.

White noise was named after the sound which has all frequencies in equal amounts, in analogy to white light. It was named in a paper on the acoustics in airplanes, where the authors wrote:

That white noise is annoying needs little argument. No one has been found who really enjoys it.

However, here it is enjoyed, in its role as the “chooser” of the collapsed state.

D. Refinements: More Collapse-Generating Operators

It is a straightforward generalization to describe collapse to a joint basis of operators \( A^\alpha \) which commute, \( [A^\alpha, A^\beta] = 0 \). This requires one white noise function \( w^\alpha \) for each \( A^\alpha \). The state vector evolution is

\[
| \phi, t \rangle = T e^{-\int_{0}^{t} dt'[iH(t') + \frac{1}{\lambda} \sum_{x}[w^\alpha(t') - 2\lambda A^\alpha]^2]} | \phi, 0 \rangle
\]

and the corresponding density matrix evolution is

\[
\frac{d\rho(t)}{dt} = -i[H, \rho(t)] - \frac{\lambda}{2} \sum_{\alpha} [A^\alpha, [A^\alpha, \rho(t)]]
\]

The ensemble average of an operator \( \mathcal{O} \) shall be denoted \( \mathcal{O}(t) \equiv \text{Tr} \rho(t) \), where Tr is the trace operation. Then, Eq. (34) gives

\[
\frac{d\mathcal{O}(t)}{dt} = -i[H, \mathcal{O}](t) - \frac{\lambda}{2} \sum_{\alpha} [A^\alpha, [A^\alpha, \mathcal{O}]](t).
\]

IV. NON-RELATIVISTIC CSL

Finally, here is the CSL proposal to describe the non-relativistic world. The index \( \alpha \) in Eq. (33) is changed to a continuum index \( x \), so the ‘chooser’ \( w(x, t) \), rather than being a set of random functions, is a random field:

\[
| \phi, t \rangle = T e^{-i \int_{0}^{t} dt'[iH(t') + \frac{1}{\lambda} \sum_{x}[w(x', t') - 2\lambda A(x')]^2]} | \phi, 0 \rangle
\]

The set of collapse-generating operators are mass-density operators, ‘smeared’ over a sphere of radius \( a \):

\[
A(x) = \sum_{n} \frac{m_n}{M} \frac{1}{(\pi a^2)^{3/4}} \int d^{3}z e^{-\frac{m_n}{\lambda}[x - z]^2} \xi_{n}(z).
\]

Here, \( \xi_{n}(z) \) is the creation operator for a particle of type \( n \) at \( z \). \( m_n \) is the mass of this particle and \( M \) is the
mass of a nucleon (say, the neutron). Thus, in ordinary matter, it is the nucleons which are mostly responsible for collapse. Experimental results\[7\] have dictated that the effective collapse rate in Eq. (36) be mass-proportional, \(\sim \lambda m_n\).

Assuming the theory is correct, the parameter values of \(\lambda, a\) should be determined by experiment.\[8\] Until then, we shall provisionally adopt the parameter values chosen by Ghirardi, Rimini and Weber\[9\] in their instantaneous collapse theory, \(\lambda \approx 10^{-16}\text{sec}^{-1}, a \approx 10^{-5}\text{cm}\). However, it should be mentioned that Adler\[10\] has given an argument for \(\lambda\) to be as large as \(\approx 10^{-11}\text{sec}^{-1}\).

The density matrix evolution equation (34) becomes, using (37),

\[
\frac{d\rho(t)}{dt} = -i[H, \rho(t)] - \lambda \sum_{k,n} \frac{m_k m_n}{M^2} \xi_k^* \xi_n .
\]

\[
\int d\mathbf{z}' e^{-\frac{i}{\lambda a^2}[x-z']^2} \sum_{k,n} \phi_k \phi_n \rho(t)
\]

\[
= -i[H, \rho(t)] - \lambda \sum_{k,n} \frac{m_k m_n}{M^2} \xi_k^* \xi_n .
\]

\[
\int d\mathbf{z}' e^{-\frac{i}{\lambda a^2}[x-z']^2} \sum_{k,n} \phi_k \phi_n \rho(t).
\]

V. FREE SMALL CLUMP

In the rest of this paper, we shall illustrate CSL by discussing the force-free behavior of the center of mass (cm) of a small (dimensions \(< a\)) clump of ordinary matter.\[1\]

For simplicity we shall neglect the electrons, and take there to be \(N\) nucleons, regarded as a single type of particle, of mass \(M\), which are very good approximations for our calculations. Then, in the particle position basis, \(|x\rangle \equiv |x_1, \ldots, x_N\rangle\), using \(\xi_k^* \xi_k |x\rangle = \sum_{i=1}^N \delta(z-x_i)|x\rangle\), Eq. (33) becomes

\[
\frac{d\langle x|\rho(t)|x'\rangle}{dt} = -i[x|H, \rho(t)|x'] - \langle x|\rho(t)|x'\rangle .
\]

\[
\lambda \sum_{i=1}^N \left[ e^{-\frac{i}{\lambda a^2}[x-x_i]^2} + e^{-\frac{i}{\lambda a^2}[x'_i-x_i]^2} - 2e^{-\frac{i}{4\lambda a^2}[x-x_i]^2} \right].
\]

Define the center of mass coordinate \(X \equiv N^{-1} \sum_i x_i\) and the relative coordinates \(y_i \equiv x_i - X\). Because it is a ‘small’ clump, exp \(-|y_1 - y_2|^2/4a^2 \approx 1\) and exp \(-|X+y_1 - X' - y'_1|^2/4a^2 \approx \exp(-|X - X'|^2/4a^2).\] With the density matrix assumed to have the form of the direct product of cm and internal coordinate density matrices, we can take the trace over the internal coordinates in Eq. (39) to obtain the equation for the evolution of \(\langle X|\rho(t)|X'\rangle\).

It is useful to express this equation in operator form, writing the cm operator as \(\hat{P}\),

\[
\frac{d\rho(t)}{dt} = -\lambda \frac{\hat{P}^2}{2MN} \rho(t) - \lambda N^2 \left[1 - e^{-\frac{2\pi a^2}{\lambda N A(x)}|x-x'|^2}\right] \rho(t).
\]

The associated state vector evolution equation (36) is

\[
|\phi, t\rangle = T e^{-i\int_0^t dt' \frac{\hat{P}^2}{\lambda N A(x')/4} - \frac{1}{\lambda N A(x')} \int d\mathbf{x}' ||w(x',t') - 2\lambda N A(x')^2|| |\phi, 0\rangle,
\]

where

\[
A(x') \equiv \frac{1}{(\pi a^2)^{3/4}} e^{-\frac{1}{2\lambda a^2|x-x'|^2}}.
\]

To illustrate the use of the collapse part of Eq. (10), consider the initial wave function

\[
|\phi, 0\rangle = \frac{1}{\sqrt{2}} (|L\rangle + |R\rangle),
\]

where the states describe the clump to the left or right, with the two wave packet cm’s separated by the distance \(D\). Then (ignoring the kinetic energy),

\[
\frac{d\langle L|\rho(t)|R\rangle}{dt} = -\lambda N^2 \left[1 - e^{-\frac{2\pi a^2}{\lambda N A(x)}}\right] \langle L|\rho(t)|R\rangle.
\]

Thus, for \(D >> a\), the collapse is described by exponential decay of the off-diagonal density matrix element with characteristic time \(\lambda^{-1} = 10^{16}\text{sec}\) for a single nucleon. For a \(10^{-5}\text{cm}\) cube of gold, where \(N \approx 10^8\), the characteristic collapse time is \(1/\lambda N^2 = 1\text{sec}\).

One might very well extend this theory to include massless particles by replacing mass-density of \(A(x)\) in Eq. (37) by energy-density/e\(^2\). One might then regard it as holding in the co-moving frame of the universe\[12\], or as the limit of a relativistic CSL\[13\].

VI. COLLAPSE OF A PACKET

We shall consider how a single wave packet undergoes collapse.

A. Big Packet

Consider a spread-out, real, positive, initial wave function such as

\[
\langle X|\phi, 0\rangle = \frac{1}{(2\pi D^2)^{3/4}} e^{-\frac{X^2}{4D^2}},
\]
where $D >> a$. We shall neglect the effect of the Hamiltonian. We shall see that the wave function collapses fairly rapidly to an approximately spherical wave function of size $a$, center location consistent with the Born Rule, and thereafter collapses more and more slowly to a smaller and smaller radius.

First we calculate the ensemble average of the operator $O \equiv |X\rangle\langle X|$, so $\overline{O}$ is the ensemble probability density at $X$. For any density matrix, it follows from the collapse part of Eq. (40), using Eq. (35), that

$$\frac{d\langle X \rangle}{dt} = -\lambda N^2 \langle 1 - e^{-\frac{2}{\lambda^2}|X_L - X_n|^2} \rangle \rho(t) = 0. \quad (46)$$

This, of course, doesn’t say that collapse occurs but, if there is collapse, it says that the ensemble position probability distribution does not change from the initial distribution (Born Rule).

In order to see that there is indeed collapse, consider the ensemble average of the modular momentum [14] operator $O \equiv \cos \hat{P}\cdot \hat{n} e$, where $n$ is a unit vector pointing in some direction. This is $1/2$ the sum of two operators, one which translates the wave function by distance $L$ in the $\hat{n}$ direction and the other in the $-\hat{n}$ direction. Thus, its expectation value gives the overlap of the wave function with itself (all the wave functions are real and positive) when translated. For any density matrix, it follows from the collapse part of Eq. (40) that

$$\frac{d\cos \hat{P}\cdot \hat{n} e}{dt} = -\lambda N^2 \left[ 1 - e^{-\frac{2}{\lambda^2}|a|^2} \right] \cos \hat{P}\cdot \hat{n} e \quad (47)$$

Thus, for $L >> a$, the ensemble average of the overlap rate of the collapsing wave functions decreases as $\approx \lambda N^2$ but then it slows, e.g., for $L = a$, the collapse rate is $\approx 2\lambda N^2$.

### B. Small Packet

If the size of the wave function is less than $a$, one can utilize an approximate density matrix evolution equation obtained by expanding the exponential in Eq. (40), retaining only the leading term:

$$\frac{d\rho(t)}{dt} = -i \left[ \hat{P}^2 \overline{2MN}^{-1}, \rho(t) \right] - \frac{\lambda N^2}{4a^2} \sum_{1=1}^{3} [\hat{X}_i, [\hat{X}_i, \rho(t)]] \quad (48)$$

The state vector evolution which yields this density matrix evolution is

$$|\phi, t\rangle = T e^{-i \int_0^t dt' \sum_{i=1}^{3} \left[ \cos \hat{X}_i, \hat{X}_i, \rho(t) \right]} T^{-1} |\phi, 0\rangle \quad (49)$$

When the initial wave function is a gaussian, such as Eq. (13) with $D < a$, since the Schrödinger equation is quadratic in $\hat{P}$ and $\hat{X}$, the solution is a gaussian. The exact solution to this problem can be found [15][16]. We shall arrive at it here using the formalism we have presented. It suffices to solve the one-dimensional problem, with initial wave function (45), Eq. (49) is the product of three terms, one for each dimension.

We assume that the wave function at any time has the form

$$\psi(X, t) = e^{-A(t)X^2 + B(t)X + C(t)}, \quad (50)$$

and proceed to solve the Schrödinger equation which follows from the time derivative of (49):

$$\frac{\partial \psi(X, t)}{\partial t} = i \frac{\partial^2}{2m \partial X^2} \psi(X, t) - \left[ \frac{1}{2} w^2(t) - \frac{\hat{X}}{\lambda} w(t), X + \frac{\hat{X}^2}{\lambda} \right] \psi(X, t) \quad (51)$$

where $m \equiv NM$ and $\hat{\lambda} \equiv \lambda N\overline{\sqrt{2}a}$. Inserting (50) into (51) we obtain

$$\hat{A} = \frac{-2i}{m} \hat{A}^2 + \frac{\hat{X}^2}{\lambda}, \quad (52a)$$

$$\hat{B} = \frac{-2i}{m} \hat{A} \hat{B} + \frac{\hat{X}}{\lambda} w(t). \quad (52b)$$

Eq. (52a) is a Ricatti equation, and can be solved by the ansatz $A = (m/2i) \hat{F} / \hat{F}$. It follows from (52a) that $\hat{F} = F(2i\hat{\lambda}^2/m\lambda)$. Thus, $F = \exp \pm \alpha(1 + it)$, where

$$\alpha \equiv \hat{\lambda}/\sqrt{m\lambda}$$

and

$$A = \frac{m\alpha(1+i)t}{2} e^{\alpha(1+i)t} - Ke^{-\alpha(1+i)t} - \frac{K}{2} e^{\alpha(1+i)t} + Ke^{-\alpha(1+i)t} \quad (53)$$

where $K$ is a constant depending upon $D$. We see that $A$, which characterizes the squared standard deviation of $X$, is the same for all $w(t)$. Thus, the wave function approaches an equilibrium size, independently of its initial spread $D$. The equilibrium occurs because the Schrödinger evolution tends to spread the

---

2 How can Eq. (49), where $w$ is just a function of $t$, arise from Eq. (40), where $w$ is a field, depending upon $x$ as well? As far as I am aware, this has not been discussed before, so we treat it in Appendix [18]. More generally, it involves changing the collapse-generating operators $A^n$ to a new, equivalent set, with concomitant change of white noise functions $w^\alpha(t)$ to a new, equivalent set.
wave function while the collapse evolution tends to narrow it. This takes place in characteristic time \(\alpha^{-1} \approx 5 \times 10^4 / \text{Nsec} \). The equilibrium spread in \(X\) (its standard deviation) is \(1 / \sqrt{2(\lambda + A^*)} = 1 / \sqrt{2m\alpha} \approx 4 / \text{N}^{1/2} \text{cm} \).

We shall henceforth assume either that the collapse process starts at negative times so that equilibrium is reached at time 0, or that \(A\) initially has its equilibrium value. Putting that value into (52d) gives

\[
\dot{B} = -\alpha(1 + i)B + \frac{\lambda}{\Lambda}w(t),
\]

with solution

\[
B(t) = \frac{\lambda}{\Lambda} \int_0^t dt' w(t') e^{-\alpha(1+i)(t-t')}.
\]

Knowing \(A\) and \(B\), the expectation values of position and squared position can be found from (50):

\[
\langle X \rangle = \frac{\langle \psi, t| X| \psi, t \rangle}{\langle \psi, t| \psi, t \rangle} = \frac{B + B^*}{2(\lambda + A^*)} + \frac{\alpha}{\lambda} \int_0^t dt' w(t') e^{-\alpha(1+i)(t-t')} \cos \alpha(t-t'),
\]

(56a)

\[
\langle X^2 \rangle = \frac{\langle \psi, t| X^2| \psi, t \rangle}{\langle \psi, t| \psi, t \rangle} = \langle X \rangle^2 + \frac{1}{2(1 + \lambda^2)(\lambda + A^*)},
\]

(56b)

To complete the solution, we need to find \(C(t)\) but, since it is used to find the probability density \(\langle \psi, t| \psi, t \rangle\), it is best that we calculate that directly from the Schrödinger equation:

\[
\frac{d}{dt} \langle \psi, t| \psi, t \rangle = -\frac{\lambda^2}{2\alpha} \langle \psi, t| \psi, t \rangle + 2\frac{\lambda}{\Lambda} w(t) \langle \psi, t| X| \psi, t \rangle
\]

\[
-2\frac{\lambda^2}{\Lambda^2} \langle \psi, t| X^2| \psi, t \rangle
\]

\[
= -\frac{1}{2\alpha} \left[ \langle \psi, t| \psi, t \rangle - 2\frac{\lambda^2}{\Lambda} \langle \psi, t| X^2| \psi, t \rangle \right]
\]

\[
- \frac{\lambda^2}{\Lambda} \frac{1}{1 + \lambda^2} \langle \psi, t| \psi, t \rangle.
\]

(57)

Therefore, omitting the time-dependent factor arising from the last term of (57) (which is absorbed in the normalization of the probability), and defining a new set of white noise functions

\[
v(t) = w(t) - 2\lambda \langle X \rangle
\]

\[
= w(t) - 2\alpha \int_0^t dt' w(t') e^{-\alpha(1+i)(t-t')} \cos \alpha(t-t')
\]

(58)

the probability density is simply

\[
\langle \psi, t| \psi, t \rangle = e^{-\frac{\lambda^2}{\alpha} \int_0^t dt' v^2(t')}.
\]

(59)

We note that \(Dw = Dv\), since it follows from (58) that the Jacobian of the transformation from \(w\)’s to \(v\)’s has 1’s on the diagonal and 0’s above the diagonal.

In order to use (59), it is necessary to obtain the inverse of the transformation (58). This can be done by taking the second derivative of (58), with the result

\[
\frac{d^2w(t)}{dt^2} = \frac{d^2v(t)}{dt^2} + 2\alpha \frac{dv(t)}{dt} + 2\alpha v(t).
\]

(60)

Defining \(v(t)\)’s Brownian motion \(\tilde{B}(t)\) by \(v(t) = dB(t)/dt\). It then follows from (60) that \(w(t)\) can variously be written as

\[
w(t) = v(t) + 2\alpha \int_0^t dt_1 v(t_1)
\]

\[+ 2\alpha^2 \int_0^t dt_1 \int_0^{t_1} dt_2 v(t_2), \quad (61a)
\]

\[= v(t) + 2\alpha \int_0^t dt_1 v(t_1)[1 + \alpha(t - t_1)], \quad (61b)
\]

\[= v(t) + 2\alpha \tilde{B}(t) + 2\alpha^2 \int_0^t dt' \tilde{B}(t'). \quad (61c)
\]

It then follows from the first equation in (58) that \(\langle X \rangle\) can be written as

\[
\langle X \rangle = \frac{1}{\sqrt{m^2\lambda}} \left[ \tilde{B}(t) + \alpha \int_0^t dt' \tilde{B}(t') \right].
\]

(62)

One can then show, using (50), (55), (61a) and (62), that

\[
\langle P \rangle = 2iA\langle X \rangle - iB(t) = \frac{\lambda}{\Lambda} \tilde{B}(t).
\]

(63)

This problem is completely solved. We see from Eqs. (62), (63) that, after the equilibrium packet size is achieved, the momentum expectation value undergoes Brownian motion and the position expectation value undergoes a motion that can be described as Brownian++.

Any expectation value can be calculated, and any ensemble average expectation value can be calculated. For example, although it can readily be found using the density matrix, the ensemble average of the squared position expectation value can be found from Eq. (65b), using \(\langle X^2 \rangle = (1/2\lambda)|w(t) - v(t)|^2\) (Eq. (58)), Eq. (61a) and \(v(t)v(t') = \lambda\delta(t - t')\) (which follows from (59)):

\[
\langle X^2 \rangle = \frac{1}{2m^2\lambda} + \frac{1}{m^2\lambda} \left[ \int_0^t dt' v(t')[1 + \alpha(t - t')] \right]^2
\]

\[= \frac{1}{2m^2\lambda} + \frac{1}{m^2\lambda} \int_0^t dt' [1 + \alpha(t - t')]^2
\]

\[= \frac{1}{2m^2\lambda} + \frac{1}{m^2\lambda} [t + \alpha t^2 + \alpha^2 t^3/3].
\]

(64)

\(\nabla^2 \sim t\) behavior occurs for classical Brownian motion, modeled as a particle undergoing Newtonian dynamics with a random force and a viscous damping force. In this case, the average Brownian ‘step’ size is constant in time.
\( X^2 \sim t^4 \) behavior occurs for classical Brownian motion when the viscous damping is removed. This is essentially because the average Brownian ‘step’ size increases with time.

So, we have the picture of the final result of collapse, a wave packet of equilibrium size which undergoes classical random walk without viscous damping, with momentum generally increasing as it undergoes classical random walk.

**VII. COLLAPSE OF INTERFERING PACKETS**

It follows from the density matrix evolution Eq. (60) that the interaction picture density matrix \( \hat{\rho}(t) \equiv U_t(t)\rho(t)U(t) \) (\( U(t) \equiv \exp(-i\hat{P}^2/2m) \)) satisfies

\[
\frac{d\hat{\rho}(t)}{dt} = -\lambda N^2 U_t(t) \left[ 1 - e^{-\frac{\lambda N^2}{m^2} |\hat{X}_L - \hat{X}_R|^2} \right] U(t)\hat{\rho}(t),
\]

with solution

\[
\hat{\rho}(t) = T e^{-\lambda N^2 \int_0^t dt' U_t(t-t') \left[ 1 - e^{-\frac{\lambda N^2}{m^2} |\hat{X}_L - \hat{X}_R|^2} \right] U(t-t') \rho(0)}.
\]

or, going back to the density matrix \( \rho(t) \),

\[
\rho(t) = T e^{-\lambda N^2 \int_0^t dt' U(t-t') \left[ 1 - e^{-\frac{\lambda N^2}{m^2} |\hat{X}_L - \hat{X}_R|^2} \right] U_t(t-t') \rho_0(t)},
\]

where \( \rho_0(t) \equiv \exp(-iHt)\rho(0)\exp(iHt) \) is the density matrix without collapse. In the position representation, Eq. (67) is

\[
\langle X|\rho(t)|X' \rangle = T e^{-\lambda N^2 \int_0^t dt' \left[ 1 - e^{-\frac{\lambda N^2}{m^2} (X(t-t') - X(t-t')^2)} \right] \rho_0(t)}\langle X|\rho_0(t)|X' \rangle.
\]

(68a)

\[
Z(t-t') \equiv \langle X - \frac{t-t'}{m^2} \nabla \rangle - \langle X' + \frac{t-t'}{m^2} \nabla \rangle.
\]

(68b)

We now note that, because \( [X_i, X_j, \nabla_j + \nabla_j'] = 0 \), it follows that \( Z(t-t'), Z(t-t') = 0 \) and so the time-ordering operation \( T \) may be removed from Eq. (68a).

Also because this commutator vanishes, any product of powers of Z’s can be written in ‘normal-ordered form,’ by which we mean that the X’s are to the left of the \( \nabla \)’s. Denoting the normal ordered form by \( \langle X|\rho_0(t)|X' \rangle \) becomes

\[
\langle X|\rho(t)|X' \rangle = e^{-\lambda N^2 \int_0^t dt' \left[ 1 - e^{-\frac{\lambda N^2}{m^2} (X(t-t') - X(t-t')^2)} \right] \rho_0(t)}\langle X|\rho_0(t)|X' \rangle.
\]

(69)

We shall apply Eq. (69) to the case where the uncollapsed density matrix \( \rho_0(t) \) is constructed from a number of wave packets,

\[
\langle X|\rho_0(t)|X' \rangle = \sum_{n,n'} c_n c_{n'} \phi_n(X,t) \phi_{n'}^*(X',t).
\]

(70)

The wave packets \( \phi_n(X,t) \) are to have well-defined momenta \( k_n(X) \) at (almost) each point of the wave packet, which itself has dimensions large compared to the wavelength. Thus, a wave packet could be a laboratory ‘plane wave,’ a good approximation to an eigenstate of momentum \( k \). It could be a cylindrical wave packet or a spherical wave packet of momentum magnitude \( k \) such as might be obtained by putting the ‘plane’ wave packet through a slit or a circular hole.

An important feature of such a packet \( \phi_j(X,t) \) is that

\[
\langle X|\hat{P}|\phi_n(t) \rangle = \frac{i}{\hbar} \nabla \phi_n(X,t) \approx k_n(X) \phi_j(X,t)
\]

(71)

is a very good approximation. Another important feature of such a packet is that (almost) each point in each wave packet can be considered as moving on a straight-line trajectory with constant velocity \( k_n(X)/m \).

Putting together Eqs. (68a), (69), (70), we obtain for the ensemble’s probability density at \( X \):

\[
\langle X|\rho(t)|X' \rangle = \sum_{n,n'} c_n c_{n'} \phi_n(X,t) \phi_{n'}^*(X,t)
\]

\[e^{-\lambda N^2 \int_0^t dt' \left[ 1 - e^{-\frac{\lambda N^2}{m^2} (X(t-t') - X(t-t')^2)} \right]^2}\]

(72)

where

\[
X_n(t-t') \equiv X - \frac{k_n(X)}{m} (t-t')
\]

(73)

That is, consider a point on the \( n \)th packet which is located at \( X \) at time \( t \). Then, \( X_n(t-t') \) is the location that point had on the \( n \)th packet at the earlier time \( t' \).

To summarize, we have seen in Eq. (40) or (44) that, when a clump is put into a superposition of two places with constant separation \( D \), the two states play the gambler’s ruin game, so that the off-diagonal elements of the density matrix decay at the rate \( \lambda N^2[1 - \exp(-D^2/4a^2)] \).

Eq. (73) says that, for a superposition of packets, the points on the packets, which end up at the same place \( X \) at time \( t \), may be thought of as playing the gambler’s ruin game with each other on the way to \( X \), with the above-mentioned distance-determining rate now varying with time, governing the collapse all along the way.

Although it is not of concern here, we mention that, of course, the spatially separated points of a single packet, or of multiple packets, likewise mutually play the gambler’s ruin game, and that description is obtained by considering the off-diagonal elements of the density matrix.

**A. Mach-Zender Interference**

As is well known, the Mach-Zender interferometer has a rectangular shape, say, with half-silvered beam-splitters at the lower left and upper right corners, and fully-silvered mirrors at the other two corners. An incoming wave packet splits into two equal packets at the first beam splitter. The packet going →, ↑ has its sign reversed when it reflects at 90◦ from the front-surfaced mirror.
The packet going $\uparrow$, $\to$, $\uparrow$ gets no net sign change: one sign change at the first, front-surfaced, beam splitter, one at the front-surfaced mirror, none at the second, back-surfaced, beam splitter. Thus, without collapse, there is no output in the $\uparrow$ direction.

Although there are certainly velocity changes of the packets, they take place over a relatively brief time interval, so Eq. (72) may be applied seriatim. Let $t$ be the time interval separating emergence from the two beam splitters. At time 0, the two packets start off with $c_1 = c_2 = 1/\sqrt{2}$. Thereafter, $|X_1(t-t') - X_2(t-t')| >> a$. Moreover, the collapse rate is unaffected if a packet changes sign. Finally, at time $t$, the second beam splitter has just made the amplitudes change sign. Therefore, in Eq. (72), since $\exp\left(-\lambda tN^2(1 + \frac{b}{a}\chi(b/a))\right)$

$$\langle X | \rho(t) | X \rangle = \frac{A^2}{2} \left[ 2 + (e^{2ikb\theta} + e^{-2ikb\theta}) \right]$$

$$-e^{-\lambda tN^2\left(1 - \frac{b}{a}\chi(b/a)\right)}$$

$$= A^2 \left[ 1 + \cos(2kb\theta)e^{-\lambda tN^2(1 - \frac{b}{a}\chi(b/a))} \right]$$

$$= 2A^2 \cos^2(kb\theta)e^{-\lambda tN^2(1 - \frac{b}{a}\chi(b/a))}$$

$$+ A^2 \left[ 1 - e^{-\lambda tN^2(1 - \frac{b}{a}\chi(b/a))} \right]$$

where $\chi(b/a) \equiv \text{erf}(b/a)$.

Thus, we see that the two-slit two-packet interference pattern decays while the single packet non-interference pattern builds up as time increases. For $b >> a$, the packet separation is $>> a$ for almost all the time and the collapse rate is $\lambda N^2$, as in the previous section. For $b << a$, the collapse rate is $\lambda N^2b^2/3a^2$.

This concludes our discussion of free particle collapse dynamics.

### Appendix A: Proof That $R$ and $S$ must be diagonal

We prove here that the real symmetric operators $R$ and $S$ in the Stratonovich Schrödinger equation for the un-normalized state vector,

$$d\langle \phi, t \rangle = \{ RdB' + Sdt \} \langle \phi, t \rangle \quad (A1)$$

must be diagonal in the $|\alpha_n\rangle$ basis. This is in order that Eq. (A1) give rise to the Itô gambler’s ruin condition Eq. (35).

$$d\alpha_n(t) = b_n(x)dB(t) \quad (A2)$$

After putting Eq. (4), $dB' = dB + fdt$, into Eq. (A1), we convert that Stratonovich equation to an Itô equation, with the result

$$d\langle \phi, t \rangle = \{ RdB + Vdt \} \langle \phi, t \rangle$$

where $V \equiv S + Rf + \frac{\lambda}{2} R^2$. (A3)

We note that $V$ is also a real symmetric operator and, if we show $R$ and $V$ must be diagonal, then $S$ must also be diagonal.

Using the rules for manipulating Itô equations, it is straightforward to find

$$d\langle \phi, t \rangle \langle \phi, t \rangle = \{ RdB + Vdt \} \langle \phi, t \rangle \langle \phi, t \rangle$$

$$+ \lambda dt R \langle \phi, t \rangle \langle \phi, t \rangle R, \quad (A4a)$$

$$d\langle \phi, t | \phi, t \rangle = 2\langle \phi, t | R \langle \phi, t \rangle dB + \langle \phi, t | V \langle \phi, t \rangle dt \rangle$$

$$+ \lambda dt \langle \phi, t | R^2 \langle \phi, t \rangle R, \quad (A4b)$$

where $\{ M, N \} \equiv MN + NM$. Defining the density matrix $\rho(t) \equiv |\phi, t\rangle \langle \phi, t|$, we obtain from Eqs. (4) and the Itô rules:

$$d\rho = \left[ (R, \rho) - 2\rho R dB + dt \{ V, \rho \} - 2dV \right]$$

$$+ \lambda dt \left[ R \rho R - \rho R^2 - 2R \{ R, \rho \} - 2\rho R \right]. \quad (A5)$$
Now, $x_n(t) = \langle a_n^\dagger \rho(t) a_n \rangle$. Thus, in order that the diagonal elements of Eq. (A6) agree with Eq. (A2), we see that the diagonal elements of Eq. (A5) which do not multiply $dB$ must vanish for arbitrary $\rho$:

$$0 = \left[ [V, \rho]_{nn} - 2 \rho_{nn} \mathcal{V} + \lambda \left[ (\rho \rho \rho \rho)_{nn} - \rho_{nn} \mathcal{V}^2 \right] - 2 \mathcal{V} \left[ \left( \rho \rho \rho \rho \right)_{nn} - 2 \rho_{nn} \mathcal{V} \right] \right].$$

(A6)

where $M_{nm} \equiv \langle a_n | M | a_m \rangle$

First, suppose that $\rho_{nm} = 1$, where $m \neq n$, and all other matrix elements of $\rho$ vanish. It follows from Eq. (A6) that

$$0 = (R \rho R)_{nn} = (R_{nm})^2.$$  

(A7)

That is, all the off-diagonal elements of $R$ vanish, so $R$ is diagonal.

Second, choose a density matrix for which $\rho_{nn}, \rho_{mm} = 1 - \rho_{nm}, \rho_{nm}$ do not vanish, but all other matrix elements of $\rho$ do vanish. Then, using the diagonal nature of $R$, Eq. (A6) may be written as

$$0 = 2V_{nm} \rho_{nm} [1 - 2 \rho_{nn}] + \rho_{nn} [1 - \rho_{nm}] \left[ 2 (V_{nn} - V_{nm}) + \lambda (R_{nn} - R_{nm}) (R_{nn} + R_{nm}) - 4 R_{nn} \rho_{nn} + R_{nn} (1 - \rho_{nn}) \right].$$

(A8)

For fixed $\rho_{nn}$, a viable density matrix (non-negative eigenvalues which add up to 1) exists for $|\rho_{nm}| \leq \sqrt{\rho_{nn} (1 - \rho_{nn})}$. But, as $\rho_{nm}$ is varied, the first term in Eq. (A8) varies while the rest of the terms remain fixed. Thus, the first term must vanish, and this means that $V_{nn} = 0$ for $n \neq m$, i.e., $V$ is diagonal as well as $R$.

### Appendix B: Transformation of Operators and White Noise

Consider the general CSL form for the evolution of the state vector, Eq. (B1)

$$|\phi, t\rangle = T e^{\int_0^t dt' (i H (t') + i \lambda \sum \omega_n (t') - 2 \lambda A_n) |\phi, 0\rangle}$$

We introduce a real orthonormal set of vectors $u_{\beta}^\alpha$, i.e.,

$$\sum \omega_{\beta}^\alpha u_{\beta}^\alpha = \delta_{\beta \alpha}, \sum \omega_{\beta}^\alpha u_{\beta}^\alpha = \delta_{\alpha \alpha}. \quad \text{Defining white noise functions} \quad v^n(t) \quad \text{and complete commuting set of operators} \quad Z^n \quad \text{by} \quad w^n(t) \equiv \sum \omega_{\beta}^\alpha v^n(t) \quad \text{and} \quad A^n(t) \equiv \sum \omega_{\beta}^\alpha Z^n(t), \quad \text{one readily sees that, in the exponent of Eq. (B1),}$$

$$\sum \omega^{\alpha} (t) - 2 \lambda A^{\alpha} = \sum \left[ v^{\beta} (t) - 2 \lambda Z^{\beta} \right]^2.$$  

(B2)

The Jacobian of the transformation from $w$‘s to $v$‘s is 1 so, in using the Probability Rule (B1), $Dw = Dv$.

We wish to apply such a transformation to Eqs. (11), (12) which, for simplicity, we limit to one-dimensional space:

$$\langle \phi, t \rangle = T e^{-i \int_0^t dt' \frac{\hat{P}^2}{2M} \rho^2} \rangle - \frac{i}{4 \lambda} \int_0^t dt' T \int dx' |w(x', t') - 2 \lambda N A(x')|^2 |\phi, 0\rangle, (B3a)$$

$$A(x') \equiv \frac{1}{(\pi a^2)^{1/4} e^{-\frac{1}{4 \pi^2} (x' - \hat{X})^2}}.$$  

(B3b)

We shall use as orthonormal functions the harmonic oscillator wave functions

$$u_n(x) \equiv C_n H_n(x/a) e^{-\frac{x^2}{2a^2}} \quad \text{where} \quad C_n \equiv \frac{1}{\sqrt{\pi^{1/2} 2^n n! a}},$$

(B4)

With the definitions $v_n(t) \equiv \int dx w(x, t) u_n(x)$ and $\hat{Z}_n \equiv \int dx A(x) u_n(x)$, the exponent in Eq. (B3a) may be written as

$$- \frac{1}{4 \lambda} \int_0^t dt' \int dx' [w(x', t') - 2 \lambda N A(x')]^2 =$$

$$- \frac{1}{4 \lambda} \sum_{n=0}^\infty \int_0^t dt' [v_n(t') - 2 \lambda N \hat{Z}_n]^2.$$  

(B5)

Thus, we see how a white-noise field gets converted to an equivalent sum of white noise functions.

Using the identity $\exp(-t^2 + 2t \hat{X}) = \sum_{n=0}^\infty t^n H_n(z)/n!$, with $t \equiv \hat{X}/2a, z \equiv x'/a$, we find

$$\hat{Z}_n = \int dx \frac{1}{(\pi a^2)^{1/4} e^{-\frac{1}{4 \pi^2} (x' - \hat{X})^2} u_n(x)$$

$$= \frac{1}{C_n(\pi a^2)^{1/4} e^{-\frac{1}{4 \pi^2} x^2}} \sum_{m=0}^\infty \frac{(\hat{X}/2a)^m}{m!} \int dx u_m(x) u_n(x)$$

$$= e^{-\frac{1}{4 a^2} \hat{X}^2} \langle \hat{X}/2a \rangle^n \frac{1}{\sqrt{n!}}.$$  

(B6)

This leads to the density matrix evolution equation

$$\frac{d \rho}{dt} = -i \left[ \hat{P}^2 \frac{1}{2 M N}, \rho(t) \right]$$

$$- \frac{\lambda N^2}{2} \sum_{n=0}^\infty \left[ e^{-\frac{1}{4 a^2} (\hat{X}/2a)^2} - \frac{e^{-\frac{1}{4 a^2} \hat{X}^2} (\hat{X}/2a)^n}{\sqrt{n!}} \}, \rho(t) \right].$$

(B7)

If we expand $-\hat{X}^2/4a^2$, we see that the $n = 0$ term goes as $(\hat{X}/a)^4$ and the rest of the terms go as $(\hat{X}/a)^n$ to lowest order. Therefore, the lowest order term comes from $n = 1$. Upon neglect of the higher order terms, this gives the density matrix evolution equation

$$\frac{d \rho}{dt} = -i \left[ \hat{P}^2 \frac{1}{2 M N}, \rho(t) \right] - \frac{\lambda N^2}{4a^2} [\hat{X}, [\hat{X}, \rho(t)]].$$  

(B8)

which is identical to the one-dimensional version of Eq. (B8).
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