The fate of hints: updated global analysis of three-flavor neutrino oscillations

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Abstract: Our herein described combined analysis of the latest neutrino oscillation data presented at the Neutrino2020 conference shows that previous hints for the neutrino mass ordering have significantly decreased, and normal ordering (NO) is favored only at the $1.6\sigma$ level. Combined with the $\chi^2$ map provided by Super-Kamiokande for their atmospheric neutrino data analysis the hint for NO is at $2.7\sigma$. The CP conserving value $\delta_{CP} = 180^\circ$ is within $0.6\sigma$ of the global best fit point. Only if we restrict to inverted mass ordering, CP violation is favored at the $\sim 3\sigma$ level. We discuss the origin of these results – which are driven by the new data from the T2K and NOvA long-baseline experiments –, and the relevance of the LBL-reactor oscillation frequency complementarity. The previous $2.2\sigma$ tension in $\Delta m^2_{21}$ preferred by KamLAND and solar experiments is also reduced to the $1.1\sigma$ level after the inclusion of the latest Super-Kamiokande solar neutrino results. Finally we present updated allowed ranges for the oscillation parameters and for the leptonic Jarlskog determinant from the global analysis.

Keywords: neutrino oscillations, solar and atmospheric neutrinos
1 Introduction

Global fits to neutrino oscillation data in the last several years have shown persistent hints for the normal neutrino mass ordering and values of the CP phase $\delta_{\text{CP}}$ around maximal CP violation [1–6]. In this article we are going to re-assess the status of those hints in light of the new data released at the Neutrino2020 conference, in particular by the T2K [7, 8] and NOvA [9, 10] long-baseline (LBL) experiments. As we are going to discuss in detail, the hints have mostly disappeared or are significantly decreased: both neutrino mass orderings provide fits of comparable quality to the global data from accelerator and reactor experiments, and the CP conserving value $\delta_{\text{CP}} = 180^\circ$ is within the 1$\sigma$ allowed range.

We discuss in detail the origin of this apparent change of trends and trace back the data samples responsible for the change. We are going to compare the latest status with our pre-Neutrino2020 analysis, NuFIT 4.1, available at the NuFIT website [11]. Most relevant for mass ordering and CP phase are the updates of the neutrino samples for T2K [8], from 1.49 to 1.97 × 10$^{21}$ POT, and NOvA [10], from 0.885 to 1.36 × 10$^{21}$ POT. The T2K and NOvA anti-neutrino exposures are the same as used for NuFIT 4.1, but both collaborations introduced relevant changes in their analysis and hence we have adapted also our anti-neutrino fits correspondingly. In addition we have updated the reactor experiments Double-Chooz [12, 13] from 818/258 to 1276/587 days of far/near detector data and RENO [14, 15] from 2200 to 2908 days of exposure.

Another update concerns the solar neutrino oscillation analysis, to include the latest total energy spectrum and the day-night asymmetry of the SK4 2970-day sample presented
at Neutrino2020 [16]. As we will show, thanks to these new data the tension on the determination of $\Delta m^2_{21}$ from KamLAND versus solar experiments has basically disappeared.

The outline of the paper is as follows. In Sec. 2 we discuss the status of the neutrino mass ordering and the leptonic CP phase $\delta_{\text{CP}}$, focusing on recent updates from T2K, NOvA, as well as the combination of LBL accelerator and reactor experiments. Despite somewhat different tendencies, we will show quantitatively that results from T2K and NOvA as well as reactors are fully statistically compatible. The status of the tension between solar and KamLAND results is presented in Sec. 3. Section 4 contains a selection of the combined results of this global fit, NuFIT 5.0, which updates our previous analyses [1, 2, 17, 18]. In particular we present the ranges of allowed values for the oscillation parameters and of the leptonic Jarlskog determinant.\footnote{Additional figures, $\Delta \chi^2$ maps and future updates of this analysis will be made available at the NuFIT website [11].} Parametrization conventions and technical details on our global analysis can be found in Ref. [2]. In particular, in what follows we use the definition

\begin{equation}
\Delta m^2_{3\ell} \quad \text{with} \quad \begin{cases} 
\ell = 1 & \text{for } \Delta m^2_{31} > 0: \text{normal ordering (NO)}, \\
\ell = 2 & \text{for } \Delta m^2_{31} < 0: \text{inverted ordering (IO)}. 
\end{cases}
\end{equation}

We finish by summarizing our results in Sec. 5. A full list of the data used in this analysis is given in appendix A.

2 Fading hints for CP violation and neutrino mass ordering

2.1 T2K and NOvA updates

We start by discussing the implications of the latest data from the T2K and NOvA long-baseline accelerator experiments, presented at the Neutrino2020 conference.\footnote{During the preparation of this work Ref. [19] appeared presenting related partial results in qualitative agreement with some of our findings for the LBL analysis.} To obtain a qualitative understanding we follow Refs. [2, 20] and expand the oscillation probability relevant for the T2K and NOvA appearance channels in the small parameters $\sin \theta_{13}$, $\Delta m^2_{21} L/E_\nu$, and $A \equiv |2E_\nu V/\Delta m^2_{3\ell}|$, where $L$ is the baseline, $E_\nu$ the neutrino energy and $V$ the effective matter potential [21]:

\begin{align}
P_{\nu_\mu \rightarrow \nu_e} & \approx 4s_{13}^2 s_{23}^2 (1 + 2 o A) - C \sin \delta_{\text{CP}} (1 + o A), \\
P_{\bar{\nu}_\mu \rightarrow \bar{\nu}_e} & \approx 4s_{13}^2 s_{23}^2 (1 - 2 o A) + C \sin \delta_{\text{CP}} (1 - o A),
\end{align}

with $s_{ij} \equiv \sin \theta_{ij}$ and

\begin{equation}
C \equiv \frac{\Delta m^2_{21} L}{4E_\nu} \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23}, \quad o \equiv \text{sgn}(\Delta m^2_{3\ell}),
\end{equation}

and we have used $|\Delta m^2_{3\ell}| L/4E_\nu \approx \pi/2$ for T2K and NOvA. At T2K, the mean neutrino energy gives $A \approx 0.05$, whereas for NOvA we find that with the empirical value of $A = 0.1$
Taking all the well-determined parameters $\theta_{13}$, $\theta_{12}$, $\Delta m_{23}^2$, $|\Delta m_{31}^2|$ at their global best fit points, we obtain numerically $C' \approx 0.28$ with negligible dependence on $\theta_{23}$. The normalization constants $N_{\nu, \bar{\nu}}$ calculated from our re-analysis of T2K and NOvA are given for the various appearance samples in Table 1. Hence the expression in Eqs. (2.4) and (2.5) serve well to understand the main behaviour under varying the parameters $\sin^2 \theta_{23}$, $\delta_{\text{CP}}$, and the mass ordering.

In Table 1 we also show the observed number of events, background subtracted events, as well as the ratio $r = (N_{\text{obs}} - N_{\text{bck}})/N_{\nu(\bar{\nu})}$. In a fit, the values of $r$ have to be accommodated by the expression in the square brackets of Eqs. (2.4) and (2.5). In brackets, we give also the $r$ values for the NuFIT 4.1 data set, to illustrate the impact of the latest data.

Similar information is presented graphically in Figure 1, showing the predicted number of events for the various appearance event samples as a function of $\delta_{\text{CP}}$, changing $\sin^2 \theta_{23}$ as well as the ordering, compared to the observed event number. Here the predictions are calculated using our experiment simulation based on fully numerical oscillation probabilities, while the general behaviour of the curves is well described by Eqs. (2.4) and (2.5).

We can clearly observe a number of tendencies. T2K data has $r > 1$ for neutrinos and $r < 1$ for anti-neutrinos, implying that the square-bracket in (2.4) [(2.5)] has to be enhanced [suppressed]. If $\theta_{13}$ is fixed as determined by reactor experiments this can be achieved by choosing NO and $\delta_{\text{CP}} \simeq 3\pi/2$ (see Sec. 2.2 for a consistent combination of reactor and LBL data). This has been the driving factor for previous hints for NO and maximal CP violation. We observe from the last row in Table 1 that indeed this tendency has become somewhat weaker with the new data, though still clearly present. In this respect an interesting role is played by the CC1$\pi$ event sample. A value $r = 3.6$ for NuFIT 4.1 shows a large excess of events in this sample, which has come down to $r = 2.4$ with the latest data. Figure 1 still shows, that even the most favorable parameter choice cannot
**Figure 1.** Predicted number of events as a function of $\delta_{CP}$ for the T2K (left) and NOνA (right) appearance data sets. $\sin^2 \theta_{23}$ varies between 0.44 and 0.58, where the lower-light (upper-dark) bound of the colored bands corresponds to 0.44 (0.58). Red (blue) bands correspond to NO (IO). For the other oscillation parameters we have adopted $\sin^2 \theta_{13} = 0.0224$, $|\Delta m^2_{32}| = 2.5 \times 10^{-3}$ eV$^2$, $\sin^2 \theta_{12} = 0.310$, $\Delta m^2_{21} = 7.39 \times 10^{-5}$ eV$^2$. The horizontal dashed lines show the observed number of events, with the ±1σ statistical error indicated by the gray shaded band.

accomodate the observed number of events within 1σ. It seems that part of previous hints can be attributed to a statistical fluctuation in this sub-leading event sample. Let us stress, however, that due to the small CC1π event numbers, statistical uncertainties are large. Indeed, CCQE neutrino and anti-neutrino events consistently point in the same direction and they are both fitted best with NO and maximal CP phase.

Moving now to NOνA, we first observe from figure 1 the larger separation between the NO and IO bands compared to T2K. This is a manifestation of the increased matter effect because of the longer baseline in NOνA. Next, neutrino data have $r \approx 1$ which can be accommodated by (NO, $\delta_{CP} \simeq \pi/2$) or (IO, $\delta_{CP} \simeq 3\pi/2$). This behavior is consistent with NOνA anti-neutrinos, however in tension with T2K in the case of NO. We conclude from these considerations that the T2K and NOνA combination can be best fitted by IO and
Figure 2. $\Delta \chi^2$ profiles as a function of $\delta_{CP}$ for different LBL data sets and their combination. We have fixed $\sin^2 \theta_{13} = 0.0224$ as well as the solar parameters and minimized with respect to $\theta_{23}$ and $|\Delta m^2_{3\ell}|$. The black/blue dashed curves correspond to the combination of LBL data with the reactor experiments Day-aBay, RENO, Double-Chooz, and in this case also $\theta_{13}$ is left free in the fit. Left (right) panels are for IO (NO) and $\Delta \chi^2$ is shown with respect to the global best fit point for each curve. Upper panels are for the NuFIT 4.1 data set, whereas lower panels correspond to the current update.

$\delta_{CP} \simeq 3\pi/2$. This is indeed confirmed in figure 2, showing the $\Delta \chi^2$ profiles as a function of $\delta_{CP}$. We observe in the lower-right panel that NOvA disfavors (NO, $\delta_{CP} \simeq 3\pi/2$) by about 4 units in $\chi^2$, whereas in the lower-left panel we see for IO consistent preference of T2K and NOvA for $\delta_{CP} \simeq 3\pi/2$. For the combination this leads to a preferred best fit for IO with $\Delta \chi^2$(NO) $\approx 1.5$ (which of course is not significant). We can also see that this effect was less relevant in NuFIT 4.1 (fig. 2, upper panels) for which we had $r = 1.3$ – compared to current 1.14 – for NOvA neutrino data. This slightly higher ratio allowed some more enhancement of the square-bracket in eq. (2.4) compared to the present situation, leading to less tension between T2K and NOvA for NO. It also lead to a larger significance of NOvA for NO.

The two-dimensional regions for T2K and NOvA in the ($\delta_{CP}$, $\sin^2 \theta_{23}$) plane for fixed $\theta_{13}$ are shown in figure 3. The better consistency for IO is apparent, while we stress that even for NO the 1$\sigma$ regions touch each other, indicating that also in this case the two
experiments are statistically consistent. We are going to quantify this later in section 2.3.

2.2 Accelerator versus reactor

In the previous section we have discussed the status of the hints on CP violation and neutrino mass ordering in the latest LBL data. In the context of $3\nu$ mixing the relevant oscillation probabilities for the LBL accelerator experiments depend also on $\theta_{13}$ which is most precisely determined from reactor experiments (and on the $\theta_{12}$ and $\Delta m^2_{21}$ parameters which are independently well constrained by solar and KamLAND data). So in our discussion, and also to construct the $\chi^2$ curves and regions shown in figs. 2, 3, and 4 for T2K, NOvA, Minos, and the LBL-combination, those parameters are fixed to their current best fit values. Given the present precision in the determination of $\theta_{13}$ this yields very similar results to marginalize with respect to $\theta_{13}$, taking into account the information from reactor data by adding a Gaussian penalty term to the corresponding $\chi^2_{LBL}$.

Let us stress that such procedure is not the same as making a combined analysis of LBL and reactor data, compare for instance the blue solid versus black/blue dashed curves in fig. 2. This is so because relevant additional information on the mass ordering can be obtained from the comparison of $\nu_\mu$ and $\nu_e$ disappearance spectral data [22, 23]. In brief, the relevant disappearance probabilities are approximately symmetric with respect to the sign of two effective mass-squared differences, usually denoted as $\Delta m^2_{\mu\mu}$ and $\Delta m^2_{ee}$, respectively. They are two different linear combinations of $\Delta m^2_{31}$ and $\Delta m^2_{32}$. Consequently, the precise determination of the oscillation frequencies in $\nu_\mu$ and $\nu_e$ disappearance experiments, yields information on the sign of $\Delta m^2_{3\ell}$. This effect has been present already in previous data (see, e.g., Ref. [2] for a discussion). We see from the two lower-left panels of figure 4 that the region for $|\Delta m^2_{3\ell}|$ for IO from the LBL combination (blue curve) is somewhat in tension

Figure 3. 1σ and 2σ allowed regions (2 dof) for T2K (red shading), NOvA (blue shading) and their combination (black curves). Contours are defined with respect to the local minimum for IO (left) or NO (right). We are fixing $\sin^2 \theta_{13} = 0.0224$, $\sin^2 \theta_{12} = 0.310$, $\Delta m^2_{21} = 7.40 \times 10^{-5}$ eV$^2$ and minimize with respect to $|\Delta m^2_{3\ell}|$. 

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\[ \text{Figure 3. 1}\sigma \text{ and 2}\sigma \text{ allowed regions (2 dof) for T2K (red shading), NOvA (blue shading) and their combination (black curves). Contours are defined with respect to the local minimum for IO (left) or NO (right). We are fixing } \sin^2 \theta_{13} = 0.0224, \sin^2 \theta_{12} = 0.310, \Delta m^2_{21} = 7.40 \times 10^{-5} \text{ eV}^2 \text{ and minimize with respect to } |\Delta m^2_{3\ell}|. \]
Figure 4. $\Delta \chi^2$ profiles as a function of $\Delta m_{32}^2$ (left) and $\sin^2 \theta_{23}$ (right) for different LBL data sets and their combination. In the left 4 panels we show also the combined reactor data from Daya-Bay, RENO and Double-Chooz. For all curves we have fixed $\sin^2 \theta_{13} = 0.0224$ as well as the solar parameters and minimized with respect to the other un-displayed parameters. $\Delta \chi^2$ is shown with respect to the best fit mass ordering for each curve. Upper panels are for the NuFIT 4.1 data set, whereas lower panels correspond to the current update.

with the one from the reactor experiments Daya-Bay, RENO and Double-Chooz (black curve), while they are in quite good agreement for NO.

In the accelerator-reactor combination this leads again to a best fit point for NO, with $\Delta \chi^2(\text{IO}) = 2.7$, considerably less than the value 6.2 of NuFIT 4.1. This is explicitly shown, for example, in the LBL-reactor curves in fig. 2. For the NO best fit, a compromise between T2K and NOvA appearance data has to be adopted, avoiding over-shooting the number of neutrino events in NOvA while still being able to accommodate both neutrino and anti-neutrino data from T2K, see figure 1. This leads to a shift of the allowed region towards $\delta_{\text{CP}} = \pi$ and a rather wide allowed range for $\delta_{\text{CP}}$ for NO, see figures 2 and 3. On the other hand, we see from these figures that for IO, both T2K and NOvA prefer $\delta_{\text{CP}} \simeq 270^\circ$. Consequently, if we restrict to this ordering, CP conservation remains disfavored at $\sim 3\sigma$.

The behaviour as a function of $\sin^2 \theta_{23}$ is shown in fig. 3 and the right panels of figure 4. It is mostly driven by the two T2K neutrino samples. As follows from eq. (2.4), their predicted event rate can be enhanced by increasing $\sin^2 \theta_{23}$. Therefore, in order to compensate for the reduction in IO, a slight preference for the second $\theta_{23}$ octant emerges for IO. In case of NO, this is less preferrable, since large $\sin^2 \theta_{23}$ would worsen the T2K anti-neutrino fit as well as NOvA neutrino data.
Let us now address the question of whether some data sets are in tension with each other at a worrisome level. A useful method to quantify the consistency of different data sets is the so-called parameter goodness-of-fit (PG) [24]. It makes use of the following test statistic:

\[ \chi_{\text{PG}}^2 = \chi_{\text{min},i}^2 - \sum_i \chi_{\text{min},i}^2, \]

(2.6)

where \( i \) labels different data sets, \( \chi_{\text{min},i}^2 \) is the \( \chi^2 \) minimum of each data set individually, and \( \chi_{\text{min},\text{glob}}^2 \) is the \( \chi^2 \) minimum of the global data, i.e., \( \chi_{\text{min},\text{glob}}^2 = \min \left( \sum_i \chi_{i}^2 \right) \). Let us denote by \( n_i \) the number of model parameters on which the data set \( i \) depends, and \( n_{\text{glob}} \) the number of parameters on which the global data depends. Then the test statistic \( \chi_{\text{PG}}^2 \) follows a \( \chi^2 \) distribution with \( n \) degrees of freedom, where [24]

\[ n = \sum_i n_i - n_{\text{glob}}. \]

(2.7)

We are going to apply this test now to different combination of the three data sets, “T2K”, “NOvA”, and “React”, where “React” is the joint data set of Daya-Bay, RENO and Double-Chooz. The accelerator samples always include appearance and disappearance channels for both neutrinos and anti-neutrinos. In order to study the consistency of the sets under a given hypothesis for the neutrino mass ordering, all minimizations are preformed restricting to a given mass ordering. Furthermore, the solar parameters are kept fixed and hence, we have \( n_{\text{T2K}} = n_{\text{NOvA}} = n_{\text{glob}} = 4 \) (namely \( \theta_{13}, \theta_{23}, \delta_{\text{CP}}, |\Delta m^2_{3\ell}| \)) and \( n_{\text{React}} = 2 \) (namely \( \theta_{13}, |\Delta m^2_{3\ell}| \)). The results are shown in table 2.

First, we check the pair-wise consistency of two out of the three sets. In all cases we find perfect consistency with \( p \)-values well above 10%. The only exception is NOvA vs React for IO which show tension at the 2\( \sigma \) level. A large contribution to this effect

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### Table 2

| data sets | normal ordering | inverted ordering |
|-----------|----------------|------------------|
|           | \( \chi_{\text{PG}}^2/n \) | p-value | \#\( \sigma \) | \( \chi_{\text{PG}}^2/n \) | p-value | \#\( \sigma \) |
| T2K vs NOvA | 6.7/4 | 0.15 | 1.4\( \sigma \) | 3.6/4 | 0.46 | 0.7\( \sigma \) |
| T2K vs React | 0.3/2 | 0.87 | 0.2\( \sigma \) | 2.5/2 | 0.29 | 1.1\( \sigma \) |
| NOvA vs React | 3.0/2 | 0.23 | 1.2\( \sigma \) | 6.2/2 | 0.045 | 2.0\( \sigma \) |
| T2K vs NOvA vs React | 8.4/6 | 0.21 | 1.3\( \sigma \) | 8.9/6 | 0.18 | 1.3\( \sigma \) |
| T2K vs NOvA | 6.5/3 | 0.088 | 1.7\( \sigma \) | 2.8/3 | 0.42 | 0.8\( \sigma \) |
| T2K vs NOvA vs React | 7.8/4 | 0.098 | 1.7\( \sigma \) | 7.2/4 | 0.13 | 1.5\( \sigma \) |

3We have also checked that the three reactor experiments by themselves are in excellent agreement with each other, see the figure “Synergies: atmospheric mass-squared splitting” available at [11]. This justifies to merge them into a single set.
comes from the determination of $\Delta m^2_{34}$, which agrees better for NO than for IO, see fig. 4 (lower-left panels). The consistency of all three sets (T2K vs NOvA vs React) is excellent for both orderings.

Second, we perform an analysis for fixed $\sin^2 \theta_{13} = 0.0224$ for all data sets. Since the accelerator experiments provide a comparatively weak constraint on $\theta_{13}$ we want to remove this freedom from the T2K and NOvA fits and test the consistency under the hypothesis of fixed $\theta_{13}$. Under this assumption, all $n_i$ as well as $n_{\text{glob}}$ quoted above are reduced by 1. The results of this analysis are shown in the lower part of tab. 2. Testing T2K vs NOvA under this assumption, we find better compatibility for IO, consistent with the discussion above and figs. 2 and 3. Let us stress, however, that even for NO the $p$-value is 9%, indicating consistency at the 1.7$\sigma$ level. Hence, we find no severe tension between T2K and NOvA. Finally, the joint T2K vs NOvA vs React analysis with fixed $\theta_{13}$ reveals roughly equal good consistency among the three sets for both orderings, at around 1.5$\sigma$. For NO the very slight tension is driven by T2K vs NOvA, whereas for IO the reactor/accelerator complementarity in the determination of $\Delta m^2_{34}$ provides a few units to $\chi^2_{PG}$.

To conclude this discussion, we find that all involved data sets are perfectly statistically compatible under the hypothesis of three-flavor oscillations.

### 3 Resolved tension in the solar sector

The analyses of the solar experiments and of KamLAND give the dominant contribution to the determination of $\Delta m^2_{21}$ and $\theta_{12}$. It has been a result of global analyses for the last decade, that the value of $\Delta m^2_{21}$ preferred by KamLAND was somewhat higher than the one from solar experiments. The tension appeared due to a combination of two effects: the well-known fact that the $^8$B measurements performed by SNO, SK and Borexino showed no evidence of the low energy spectrum turn-up expected in the standard LMA-MSW \cite{21, 25} solution for the value of $\Delta m^2_{21}$ favored by KamLAND, and the observation of a non-vanishing day-night asymmetry in SK, whose size is larger than the one predicted for the $\Delta m^2_{21}$ value indicated by KamLAND. In our last published analysis \cite{2} we included the energy-zenith spectra or day/night spectra for SK1–3, together with the 2860-day total energy spectrum of SK4 \cite{26}. This last one made the lack of the turn-up effect slightly stronger. As for the day-night variation in SK4, it was included in terms of their quoted day-night asymmetry for SK4 2055-day \cite{27}

\[ A_{D/N,SK4-2055} = [-3.1 \pm 1.6(\text{stat.}) \pm 1.4(\text{syst.})] \% . \]  

(3.1)

Altogether this resulted in slightly over 2$\sigma$ discrepancy between the best fit $\Delta m^2_{21}$ value indicated of KamLAND and the solar results. For example the best fit $\Delta m^2_{21}$ of KamLAND was at $\Delta_{\text{solar}}^2 = 4.7$ in the analysis with the GS98 fluxes.

Here we update the solar analysis to include the latest SK4 2970-day results\footnote{We do not include here the latest data release from Borexino \cite{28}, which is expected to have a very small impact on the determination of oscillation parameters.} presented in Neutrino2020 \cite{16} in the form of their total energy spectrum and the updated day-night
Figure 5. Left: Allowed parameter regions (at 1σ, 90%, 2σ, 99%, and 3σ CL for 2 dof) from the combined analysis of solar data for GS98 model (full regions with best fit marked by black star) and AGSS09 model (dashed void contours with best fit marked by a white dot), and for the analysis of KamLAND data (solid green contours with best fit marked by a green star) for fixed \( \sin^2 \theta_{13} = 0.0224 \) (\( \theta_{13} = 8.6 \)). We also show as orange contours the previous results of the global analysis for the GS98 model in Ref. [2]. Right: \( \Delta \chi^2 \) dependence on \( \Delta m^2_{21} \) for the same four analyses after marginalizing over \( \theta_{12} \).

We show in fig. 5 the present determination of these parameters from the global solar analysis in comparison with that of KamLAND data. The results of the solar neutrino analysis are shown for the two latest versions of the Standard Solar Model, namely the GS98 and the AGSS09 models [29] obtained with two different determinations of the solar abundances [30]. For sake of comparison we also show the corresponding results of the solar analysis with the pre-Neutrino2020 data [2].

As seen in the figure, with the new data the tension between the best fit \( \Delta m^2_{21} \) of KamLAND and that of the solar results has decreased. Quantitatively we now find that the best fit \( \Delta m^2_{21} \) of KamLAND lies at \( \Delta \chi^2_{\text{solar}} = 1.3 \) (1.14σ) in the analysis with the GS98 fluxes. This decrease in the tension is due to both, the smaller day-night asymmetry (which lowers \( \Delta \chi^2_{\text{solar}} \) of the the best fit \( \Delta m^2_{21} \) of KamLAND by 2.4 units) and the slightly more pronounced turn-up in the low energy part of the spectrum which lowers it one extra unit.

### 4 Global fit results

Finally we present a selection of the results of our global analysis NuFIT 5.0 using data available up to July 2020 (see appendix A for the complete list of the used data including
references). We show two versions of the analysis which differ in the inclusion of the results of the Super-Kamiokande atmospheric neutrino data (SK-atm). As discussed in Ref. [2] there is not enough information available for us to make an independent analysis comparable in detail to that performed by the collaboration, hence we have been making use of their tabulated $\chi^2$ map which we can combine with our global analysis for the rest of experiments. This table was made available for their analysis of SK1–4 corresponding to 328 kton-years data [31]. The collaboration has presented new oscillation results obtained from the analysis of updated SK4 samples, both by itself [32] and in combination with the SK1–3 phases [16]. They seem to indicate that their hint for ordering discrimination has also decreased. Unfortunately the corresponding $\chi^2$ maps of these analyses have not been made public. Hence in what follows we refer as “with SK-atm” to the analysis including the tabulated SK1–4 328 kiloton years data $\chi^2$ map, i.e., the same as in NuFIT 4.0 and 4.1.

Here we graphically present the results of our global analysis in the form of one-dimensional $\Delta \chi^2$ curves (fig. 6) and two-dimensional projections of confidence regions (fig. 7). The corresponding best fit values as well as 1σ and 3σ confidence intervals for the oscillation parameters are listed in table 3. Defining the $3\sigma$ relative precision of the parameter by $2(x_{\text{up}} - x_{\text{low}})/(x_{\text{up}} + x_{\text{low}})$, where $x_{\text{up}}$ ($x_{\text{low}}$) is the upper (lower) bound on a parameter $x$ at the $3\sigma$ level, we obtain the following $3\sigma$ relative precision (marginalizing over ordering):

$$
\begin{align*}
\theta_{12} & : 14\% , & \theta_{13} & : 9.0\% , & \theta_{23} & : 27\% [25\%] , \\
\Delta m_{21}^2 & : 16\% , & |\Delta m_{3\ell}^2| & : 6.7\% [6.5\%] , & \delta_{\text{CP}} & : 100\% [100\%],
\end{align*}
$$

where the numbers between brackets show the impact of including SK-atm in the precision of the determination of such parameter. The $\Delta \chi^2$ profile of $\delta_{\text{CP}}$ is not gaussian and hence its precision estimation above is only indicative.

In table 3 we give the best fit values and confidence intervals for both mass orderings, relative to the local best fit points in each ordering. The global confidence intervals (marginalizing also over the ordering) are identical to the ones for normal ordering, which have also been used in eq. (4.1). The only exception to this statement is $\Delta m_{3\ell}^2$ in the analysis without SK-atm: in this case a disconnected interval would appear above $2\sigma$ corresponding to negative values of $\Delta m_{3\ell}^2$ (i.e., inverted ordering).

Projecting over the combinations appearing on the elements of the leptonic mixing matrix we derive the following $3\sigma$ ranges (see Ref. [33] for details on how we derive the

5For additional figures and tables corresponding to this global analysis we refer the reader to the NuFIT webpage [11].
Table 3. Three-flavor oscillation parameters from our fit to global data. The numbers in the 1st (2nd) column are obtained assuming NO (IO), i.e., relative to the respective local minimum. Note that $\Delta m^2_{32} \equiv \Delta m^2_{31} > 0$ for NO and $\Delta m^2_{32} \equiv \Delta m^2_{31} < 0$ for IO. The results shown in the upper (lower) table are without (with) adding the tabulated SK-atm $\Delta \chi^2$. 

|                  | Normal Ordering (best fit) | Inverted Ordering ($\Delta \chi^2 = 2.7$) | Normal Ordering (best fit) | Inverted Ordering ($\Delta \chi^2 = 7.1$) |
|------------------|----------------------------|-----------------------------------------|----------------------------|-----------------------------------------|
|                  | bfp ±1σ                    | 3σ range                                | bfp ±1σ                    | 3σ range                                |
| sin$^2 \theta_{12}$ | $0.304^{+0.011}_{-0.012}$  | 0.269 → 0.343                           | $0.304^{+0.011}_{-0.012}$  | 0.269 → 0.343                           |
| $\theta_{12}/^\circ$ | 33.44^{+0.78}_{-0.75}    | 31.27 → 35.86                           | 33.45^{+0.78}_{-0.75}    | 31.27 → 35.86                           |
| sin$^2 \theta_{23}$ | $0.570^{+0.018}_{-0.024}$ | 0.407 → 0.618                           | $0.575^{+0.017}_{-0.021}$ | 0.411 → 0.621                           |
| $\theta_{23}/^\circ$ | 49.0^{+1.1}_{-1.4}        | 39.6 → 51.8                             | 49.3^{+1.0}_{-1.2}        | 39.9 → 52.0                             |
| sin$^2 \theta_{13}$ | 0.0222^{+0.00068}_{-0.00062} | 0.02034 → 0.02430                      | 0.02240^{+0.00062}_{-0.00062} | 0.02053 → 0.02436                      |
| $\theta_{13}/^\circ$ | 8.57^{+0.13}_{-0.12}      | 8.20 → 8.97                             | 8.61^{+0.12}_{-0.12}      | 8.24 → 8.98                             |
| $\delta_{\text{CP}}/^\circ$ | 195^{+51}_{-25}          | 107 → 403                               | 286^{+27}_{-32}          | 192 → 360                               |
| $\Delta m_{21}^2/^{10^{-5} \text{eV}^2}$ | 7.42^{+0.21}_{-0.20}      | 6.82 → 8.04                             | 7.42^{+0.21}_{-0.20}      | 6.82 → 8.04                             |
| $\Delta m_{32}^2/^{10^{-5} \text{eV}^2}$ | $+2.51^{+0.028}_{-0.027}$ | $+2.431 → +2.598$                      | $-2.49^{+0.028}_{-0.028}$ | $-2.583 → -2.412$                      |

without SK atmospheric data

|                  | Normal Ordering (best fit) | Inverted Ordering ($\Delta \chi^2 = 2.7$) | Normal Ordering (best fit) | Inverted Ordering ($\Delta \chi^2 = 7.1$) |
|------------------|----------------------------|-----------------------------------------|----------------------------|-----------------------------------------|
|                  | bfp ±1σ                    | 3σ range                                | bfp ±1σ                    | 3σ range                                |
| sin$^2 \theta_{12}$ | $0.304^{+0.012}_{-0.012}$  | 0.269 → 0.343                           | $0.304^{+0.013}_{-0.012}$  | 0.269 → 0.343                           |
| $\theta_{12}/^\circ$ | 33.44^{+0.77}_{-0.74}    | 31.27 → 35.86                           | 33.45^{+0.78}_{-0.75}    | 31.27 → 35.86                           |
| sin$^2 \theta_{23}$ | $0.573^{+0.016}_{-0.020}$ | 0.415 → 0.616                           | $0.575^{+0.016}_{-0.019}$ | 0.419 → 0.617                           |
| $\theta_{23}/^\circ$ | 49.2^{+0.9}_{-1.2}        | 40.1 → 51.7                             | 49.3^{+0.9}_{-1.1}        | 40.3 → 51.8                             |
| sin$^2 \theta_{13}$ | 0.02219^{+0.00062}_{-0.00063} | 0.02032 → 0.02410                      | 0.02238^{+0.00063}_{-0.00062} | 0.02052 → 0.02428                      |
| $\theta_{13}/^\circ$ | 8.57^{+0.12}_{-0.12}      | 8.20 → 8.93                             | 8.60^{+0.12}_{-0.12}      | 8.24 → 8.96                             |
| $\delta_{\text{CP}}/^\circ$ | 197^{+27}_{-24}          | 120 → 369                               | 282^{+26}_{-30}          | 193 → 352                               |
| $\Delta m_{21}^2/^{10^{-5} \text{eV}^2}$ | 7.42^{+0.21}_{-0.20}      | 6.82 → 8.04                             | 7.42^{+0.21}_{-0.20}      | 6.82 → 8.04                             |
| $\Delta m_{32}^2/^{10^{-5} \text{eV}^2}$ | $+2.51^{+0.028}_{-0.027}$ | $+2.435 → +2.598$                      | $-2.49^{+0.028}_{-0.028}$ | $-2.583 → -2.414$                      |
Figure 6. Global 3ν oscillation analysis. We show $\Delta \chi^2$ profiles minimized with respect to all undisplayed parameters. The red (blue) curves correspond to Normal (Inverted) Ordering. Solid (dashed) curves are without (with) adding the tabulated SK-atm $\Delta \chi^2$. Note that as atmospheric mass-squared splitting we use $\Delta m^2_{31}$ for NO and $\Delta m^2_{32}$ for IO.
Figure 7. Global 3ν oscillation analysis. Each panel shows the two-dimensional projection of the allowed six-dimensional region after minimization with respect to the undisplayed parameters. The regions in the four lower panels are obtained from $\Delta \chi^2$ minimized with respect to the mass ordering. The different contours correspond to 1σ, 90%, 2σ, 99%, 3σ CL (2 dof). Colored regions (black contour curves) are without (with) adding the tabulated SK-atm $\Delta \chi^2$. Note that as atmospheric mass-squared splitting we use $\Delta m^2_{31}$ for NO and $\Delta m^2_{32}$ for IO.
Figure 8. Dependence of the global $\Delta \chi^2$ function on the Jarlskog invariant. The red (blue) curves are for NO (IO). Solid (dashed) curves are without (with) adding the tabulated SK-atm $\Delta \chi^2$.

Note that there are strong correlations between these allowed ranges due to the unitary constraint.

The present status of leptonic CP violation is further illustrated in fig. 8 where we show the determination of the the Jarlskog invariant defined as:

$$J_{\text{CP}} \equiv \text{Im} \left[ U_{\alpha i} U_{\beta j}^* U_{\beta i}^* U_{\beta j} \right]$$

$$\equiv J_{\text{CP}}^\text{max} \sin \delta_{\text{CP}} = \cos \theta_{12} \sin \theta_{12} \cos \theta_{23} \sin \theta_{23} \cos^2 \theta_{13} \sin \theta_{13} \sin \delta_{\text{CP}} .$$

It provides a convention-independent measure of leptonic CP violation in neutrino propagation in vacuum \[34\] – analogous to the factor introduced in Ref. \[35\] for the description of CP violating effects in the quark sector, presently determined to be $J_{\text{CP}}^\text{quarks} = (3.18 \pm 0.15) \times 10^{-5}$ \[36\]. From the figure we read that the determination of the mixing angles implies a maximal possible value of the Jarlskog invariant of

$$J_{\text{CP}}^\text{max} = 0.0332 \pm 0.0008 \text{ (} \pm 0.0019\text{)}$$

at 1$\sigma$ (3$\sigma$) for both orderings. Furthermore we see that with the inclusion of the new results, the best fit value $J_{\text{CP}}^\text{best} = -0.0089$ is only favored over CP conservation $J_{\text{CP}} = 0$ with $\Delta \chi^2 = 0.38$, irrespective of SK-atm.
5 Summary

Let us summarize the main findings resulting from the Neutrino2020 updates in neutrino oscillations.

- The best fit in the global analysis remains for the normal mass ordering, however, with reduced significance. In the global analysis without SK-atm, inverted ordering is disfavored only with a $\Delta \chi^2 = 2.7$ (1.6$\sigma$) to be compared with $\Delta \chi^2 = 6.2$ (2.5$\sigma$) in NuFIT 4.1. This change is driven by the new LBL results from T2K and NOvA which indeed by themselves favor IO (with $\theta_{13}$ as determined by the reactor data and $\theta_{12}$ and $\Delta m^2_{21}$ by the solar and KamLAND results). The best fit for NO in the combined global analysis is driven by the better compatibility between the $\Delta m^2_{3\ell}$ determined in $\nu_\mu$ disappearance at accelerators with that from $\nu_e$ disappearance at reactors (see left panel in fig. 4).

- Despite slightly different tendencies in some parameter regions, T2K, NOvA and reactor experiments are statistically in very good agreement with each other. We have performed tests of various experiment and analysis combinations, which all show consistency at a CL below 2$\sigma$ (section 2.3).

- If atmospheric data from Super-Kamiokande is included, inverted ordering is disfavored with a $\Delta \chi^2 = 7.3$ (2.7$\sigma$) compared to $\Delta \chi^2 = 10.4$ (3.2$\sigma$) in NuFIT 4.1. Hence, a modest indication for NO remains. Let us note that in the recent Super-Kamiokande update presented at Neutrino2020 [16] (with increased statistic and improved mass ordering sensitivity) the $\Delta \chi^2$ for IO is reduced by about 1 unit compared to the analysis we are using in our global fit. Therefore we expect that once the $\chi^2$ map for the new SK analysis becomes available, the combined hint in favor of NO may further decrease.

- We obtain a very mild preference for the second octant of $\theta_{23}$, with the best fit point located at $\sin^2 \theta_{23} = 0.57$ (slightly more non-maximal than the best fit of 0.56 in NuFIT 4.1), but with the local minimum in the first octant at $\sin^2 \theta_{23} = 0.455$ at a $\Delta \chi^2 = 0.53$ (2.2) without (with) SK-atm. Maximal mixing ($\sin^2 \theta_{23} = 0.5$) is disfavored with $\Delta \chi^2 = 2.4$ (3.9) without (with) SK-atm.

- The best fit for the complex phase is at $\delta_{\text{CP}} = 195^\circ$. Compared to previous results (e.g., NuFIT 4.1 [11]), the allowed range is pushed towards the CP conserving value of 180$^\circ$, which is now allowed at 0.6$\sigma$ with or without SK-atm. If we restrict to IO, the best fit of $\delta_{\text{CP}}$ remains close to maximal CP violation, with CP conservation being disfavored at around 3$\sigma$.

- New solar neutrino data from Super-Kamiokande lead to an upward shift of the allowed region for $\Delta m^2_{31}$, which significantly decreased the tension between solar and KamLAND data. They are now compatible at 1.1$\sigma$, compared to about 2.2$\sigma$ for the pre-Neutrino2020 situation.
Overall we have witnessed decreasing significance of various “hints” present in previous data. This is consistent with the fate of fluctuations which is that of fading away as time goes by.

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A List of data used in the analysis

Solar experiments

- *External information*: Standard Solar Model [30].
- Chlorine total rate [37], 1 data point.
- Gallex & GNO total rates [38], 2 data points.
- SAGE total rate [39], 1 data point.
- SK1 full energy and zenith spectrum [40], 44 data points.
- SK2 full energy and day/night spectrum [41], 33 data points.
- SK3 full energy and day/night spectrum [42], 42 data points.
- SK4 2970-day day-night asymmetry [16] and energy spectrum [16], 24 data points.
- SNO combined analysis [43], 7 data points.
- Borexino Phase-I 741-day low-energy data [44], 33 data points.
- Borexino Phase-I 246-day high-energy data [45], 6 data points.
- Borexino Phase-II 408-day low-energy data [46], 42 data points.

Atmospheric experiments

- *External information*: Atmospheric neutrino fluxes [47].
- IceCube/DeepCore 3-year data [48, 49], 64 data points.
- SK1–4 328 kiloton years [50], $\chi^2$ map [31] added to our global analysis.
Reactor experiments

- KamLAND separate DS1, DS2, DS3 spectra [51] with Daya-Bay reactor $\nu$ fluxes [52], 69 data points.
- Double-Chooz FD/ND spectral ratio, with 1276-day (FD), 587-day (ND) exposures [13], 26 data points.
- Daya-Bay 1958-day EH2/EH1 and EH3/EH1 spectral ratios [53], 52 data points.
- Reno 2908-day FD/ND spectral ratio [15], 45 data points.

Accelerator experiments

- MINOS $10.71 \times 10^{20}$ pot $\nu_\mu$-disappearance data [54], 39 data points.
- MINOS $3.36 \times 10^{20}$ pot $\bar{\nu}_\mu$-disappearance data [54], 14 data points.
- MINOS $10.6 \times 10^{20}$ pot $\nu_e$-appearance data [55], 5 data points.
- MINOS $3.3 \times 10^{20}$ pot $\bar{\nu}_e$-appearance data [55], 5 data points.
- T2K $19.7 \times 10^{20}$ pot $\nu_\mu$-disappearance data [8], 35 data points.
- T2K $19.7 \times 10^{20}$ pot $\nu_e$-appearance data [8], 23 data points for the CCQE and 16 data points for the CC$\pi$ samples.
- T2K $16.3 \times 10^{20}$ pot $\bar{\nu}_\mu$-disappearance data [8], 35 data points.
- T2K $16.3 \times 10^{20}$ pot $\bar{\nu}_e$-appearance data [8], 23 data points.
- NOvA $13.6 \times 10^{20}$ pot $\nu_\mu$-disappearance data [10], 76 data points.
- NOvA $13.6 \times 10^{20}$ pot $\nu_e$-appearance data [10], 13 data points.
- NOvA $12.5 \times 10^{20}$ pot $\bar{\nu}_\mu$-disappearance data [10], 76 data points.
- NOvA $12.5 \times 10^{20}$ pot $\bar{\nu}_e$-appearance data [10], 13 data points.

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