Vortex/anti-vortex pair creation in black hole thermodynamics

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An isolated critical point is a peculiar thermodynamic critical point that occurs in the phase diagram of hyperbolic black holes in \(K\)-th order Lovelock gravity in higher dimensions (with \(K\) odd) for special tuned Lovelock coupling constants. It corresponds to a ‘merger’ of two swallowtails and is characterized by non-standard critical exponents. Upon employing a recent proposal for assigning a topological charge to thermodynamic critical points, we argue that the isolated critical point offers an interpretation corresponding to the onset of a topological phase transition of a ‘vortex/anti-vortex pair’.

I. INTRODUCTION

The framework of extended black hole thermodynamics [1] has introduced many new remarkable features into the thermodynamic phenomenology of AdS black holes [2]. Known as black hole chemistry, a panoply of interesting phase transitions have been discovered, ranging from understanding the Hawking–Page phase transition [3] as a solid/liquid transition, to Van der Waals-like phase transitions [4–6], re-entrant phase transitions [7], triple points [8], and superfluid-like features [9]. Via the AdS/CFT correspondence these are expected to be dual to the phase transitions of the corresponding boundary CFT, as has long been known for the Hawking–Page transition [10] and more recently for the Van der Waals transitions of charged AdS black holes [11,12]. Among these, perhaps the most unexpected was the discovery of the isolated critical point in the phase diagram of hyperbolic black holes of the odd-order Lovelock theories [13–15]. This transition corresponds to a merger of two free energy swallowtails and gives the only known example of a critical point in black hole thermodynamics that is characterized by non-standard critical exponents.

In recent years there have been several attempts to uncover the black hole microscopic degrees of freedom responsible for the above phase transitions. For example, one suggestion involved a proposal for calculating the correlation length and its corresponding critical exponent [16–17]. More recently, a new proposal for assigning a topological charge to various critical points was put forward [18], and further studied for Gauss-Bonnet gravity [19]. It was shown that, apart from the standard critical points (with negative topological charge \(Q = -1\)), one can also find a ‘novel’ critical point – characterized by the opposite topological charge [18]. Unfortunately this example suffers from a drawback insofar as the novel critical point is unphysical – it occurs in an unstable branch of the free energy and so does not correspond to a phase transition.

In what follows, we shall show that the isolated critical point (that occurs for a special tuned Lovelock coupling, referred to as \(a\) in the simplest 3rd-order Lovelock case [14]) can be understood as an ‘onset’ of standard (vortex) and novel (anti-vortex) critical point pair creation. As the Lovelock coupling \(a\) is decreased, the two (now physical) critical points separate from each other – the vortex/anti-vortex pair has been created. This suggests that the isolated critical point can be interpreted as a topological phase transition.

II. LOVELOCK THERMODYNAMICS AND THE ISOLATED CRITICAL POINT

A. Lovelock black holes

In what follows we will concentrate on hyperbolic black holes in Lovelock gravity [20]. This is the most general geometric higher-curvature theory that gives rise to the second-order equations of motion for the metric. In \(d\) spacetime dimensions, its Lagrangian reads [20]

\[
\mathcal{L} = \frac{1}{16\pi G_N} \sum_{k=0}^{K} \hat{\alpha}(k) \mathcal{L}^{(k)},
\]

where \(K = \lfloor \frac{d-1}{2} \rfloor\), the \(\hat{\alpha}(k)\) are the Lovelock coupling constants, and \(\mathcal{L}^{(k)}\) are the \(2k\)-dimensional Euler densities, given by

\[
\mathcal{L}^{(k)} = \frac{1}{2k} \mathring{\delta}_{c_1 d_1 \ldots c_k d_k} R_{a_1 b_1} R_{a_2 b_2} \ldots R_{a_k b_k},
\]

with the ‘generalized’ Kronecker delta function \(\mathring{\delta}_{c_1 d_1 \ldots c_k d_k}\) totally antisymmetric in both sets of indices, and \(R_{a_1 b_1} \ldots R_{a_k b_k}\) the Riemann tensor.

To find the corresponding static vacuum spherically symmetric black hole solutions, we employ the following ansatz:

\[
ds^2 = -f(r) dt^2 + \frac{dr^2}{f(r)} + r^2 d\Omega_{d-2}^2,
\]

\[f(r) = 1 - \frac{2M}{r} + \frac{Q^2}{r^2},
\]

\[M = \int_0^\infty ds \frac{\hat{\alpha}(1) \mathcal{L}^{(1)}}{\hat{\alpha}(0) \mathcal{L}^{(0)}}.
\]
where \(d\Omega_{d-2}^2\) denotes the line element of a \((d-2)\)-dimensional space of constant curvature \(\kappa(d-2)(d-3)\)
with \(\kappa = +1, 0, -1\) for spherical, flat, and hyperbolic geometries respectively of finite volume \(\Sigma_{d-2}\), the latter two cases being compact via identification \([21–23]\). The Lovelock equations of motion derived from \([1]\) then reduce (after integration) to the following polynomial equation for \(f(r)\) \([24–26]\):

\[
P(f) = \sum_{k=0}^{K} \alpha_k \left( \frac{\kappa - f}{r^2} \right)^k = \frac{16\pi G N M}{(d-2)\Sigma_{d-2}^2} \equiv m(r),
\]

where \(M\) stands for the ADM mass of the black hole,

\[
\alpha_0 = \frac{\dot{\alpha}(0)}{(d-1)(d-2)} = \frac{16\pi G N P}{(d-1)(d-2)}, \quad \alpha_1 = \dot{\alpha}(1),
\]

\[
\alpha_k = \dot{\alpha}(k) \prod_{n=3}^{2k} (d-n) \quad \text{for} \quad k \geq 2
\]

are the rescaled Lovelock couplings, and \(P = -\Lambda/(8\pi G N)\) is the thermodynamic pressure associated with the (negative) cosmological constant \(\Lambda\) \([1]\).

The black hole given by \([2]\) and \([3]\) is characterized by the following thermodynamic quantities \([26]\):

\[
M = \frac{\Sigma_{d-2}}{16\pi G N} \sum_{k=0}^{K} \alpha_k \kappa^{k-1} r_{+}^{d-2k}, \quad V = \sum_{d-2}^{d-1} r_{+}^{d-1} \frac{d}{d-1},
\]

\[
T = \frac{|f'(r_+)|}{4\pi} = \frac{1}{4\pi r_+ \Delta} \left[ \sum_{k} \kappa \alpha_k (d-2k-1) \left( \frac{\kappa}{r_+} \right)^{k-1} \right],
\]

\[
S = \frac{\Sigma_{d-2}}{4G N} \sum_{k=0}^{K} \frac{k \alpha_k (d-2k)}{d-2k},
\]

where

\[
\Delta = \sum_{k=1}^{K} k \alpha_k \left( \frac{\kappa r_+^{2k}}{2} \right)^{k-1}.
\]

These satisfy the standard (extended) first law

\[
\delta M = T \delta S + V \delta P + \sum_{k=1}^{K} \Psi_k \delta \alpha_k,
\]

where \(V\) is a thermodynamic quantity conjugate to pressure \(P\), and \(\Psi_k\) are conjugates to couplings \(\alpha_k\); explicit expressions for \(\Psi_k\) are known \([13]\) but we will not need them. For our purposes we shall concentrate on a thermodynamic ensemble defined by the following (Gibbs) free energy:

\[
G = M - TS = G(T, P, \alpha_1, \ldots, \alpha_K).
\]

### B. Isolated critical point

Let us now, for a moment, consider a very special case of Lovelock gravity characterized by the following fine tuned Lovelock coupling constants \([13, 14]\):

\[
\alpha_k = \alpha_K A^{K-k} \left( \frac{K}{k} \right), \quad 2 \leq k < K,
\]

with \(\alpha_0\) arbitrary, \(\alpha_1 = 1\), and \(A = (\Lambda \alpha_K)^{\frac{1}{2K-1}}\). In this case the polynomial \(P(f)\) drastically simplifies and yields the following solution for \(f\):

\[
f = \kappa + r_+^2 \left[ 1 - \left( \frac{m(r) - \alpha_0}{\alpha K A^{K-1} + 1} \right)^{1/K} \right],
\]

and the following equation of state

\[
P = \frac{(d-1)(d-2)\alpha}{16\pi G N} \left[ B^{K-1} \left( \frac{2K(2\pi r_+ T + \kappa)}{(d-1)r_+^2} - B \right) + A K \right],
\]

where \(B \equiv \frac{\kappa}{r_+^2} + A\).

Concentrating on the hyperbolic, \(\kappa = -1\), case we then find a very special point given by

\[
r_c = \frac{1}{\sqrt{A}}, \quad T_c = \frac{1}{2\pi r_+}, \quad P_c = \frac{(d-1)(d-2)\alpha K}{16\pi G N},
\]

for which \(\frac{\partial^2 P}{\partial r_+^2} = 0\) for all \(k = 1, \ldots, K-1\) and \(\frac{\partial^2 P}{\partial r_+^2}\) is negative. When \(K\) is odd, this describes an isolated critical point where the two swallowtails in the free energy – temperature diagram merge together. Such a point corresponds to vanishing black hole mass, \(M = 0\), and is characterized by the following non-standard critical exponents \([13, 14]\):

\[
\dot{\alpha} = 0, \quad \dot{\beta} = 1, \quad \dot{\gamma} = K - 1, \quad \dot{\delta} = K
\]

similar to glass phase transitions \([2]\), and in contrast to the standard exponents from mean field theory.

### C. Representative example in \(d = 7\) dimensions

While the above isolated critical point exists for all odd \(K \geq 3\) and all corresponding higher dimensions, we now focus on the ‘simplest’ case with \(K = 3\) and \(d = 7\). Using the following dimensionless quantities:

\[
v = \frac{r_+}{\left( \alpha_3 \right)^{1/4}}, \quad t = 5(\alpha_3)^{1/4} T, \quad s = (\alpha_3)^{-5/4} S,
\]

\[
g = \frac{1}{\Sigma_{d-2}} \frac{3-d}{2} G, \quad p = 4\sqrt{\alpha_3} P, \quad \alpha = \frac{\alpha_2}{\sqrt{\alpha_3}},
\]

and setting \(\alpha_3 = 1\), the isolated critical point occurs at the following critical value for \(\alpha\):

\[
\alpha_c = \sqrt{3}.
\]
isolated critical point (ICP). This is reflected by the corresponding \( p - t \) phase diagram that features two first order phase transitions ‘interrupted’ by the isolated critical point.

Let us next consider different values of \( \alpha \), using the general expressions \cite{5} for the thermodynamic quantities. As shown in \cite{13, 14}, for \( \alpha > \sqrt{3} \) there are no physical critical points present. On the other hand as \( \alpha \) decreases from its critical value, the two swallow tails separate and ‘travel apart’. This gives rise to two new critical points \( (\text{CP1}) \) and \( (\text{CP2}) \), each of which terminates into its own existence line of first order phase transitions. A representative example with \( \alpha = 1.65 \) is shown in Fig. 2. Interestingly, as \( \alpha \) decreases further, the second critical point \( \text{(CP2)} \) ‘travels’ to larger temperature, and eventually disappears (at infinite temperature) for

\[
\alpha = \alpha_T = \sqrt{5}/3, \quad (16)
\]

below which only \( (\text{CP1}) \) remains present.

As we shall now argue, decreasing \( \alpha \) below its critical value to \( \alpha = \sqrt{3} \) yields a topological phase transition similar to vortex/anti-vortex pair annihilation.

### III. ASSIGNING TOPOLOGICAL CHARGES TO CRITICAL POINTS

#### A. Thermodynamic topological charges

Following \cite{18, 19}, let us now assign the topological charges to the above critical points. This is done as follows. The temperature of a black hole \( T = T(S,P,...) \) at a critical point obeys the following relation:\footnote{As shown in \cite{15}, this relation together with a requirement of vanishing of the vector field \( \phi \) (defined below) is equivalent to finding the standard critical points of the system. As we shall see, this is however not the case of the isolated critical point for which the procedure breaks down: \( \phi \) does not vanish and imposing \cite{17} becomes inconsistent.}

\[
\left( \frac{\partial T}{\partial S} \right)_{P,...} = 0. \quad (17)
\]

This allows one to write down a new ‘thermodynamic potential’ (relevant for critical points)

\[
\Phi = \frac{1}{\sin \theta} \tilde{T}(S,...), \quad (18)
\]

where \( \tilde{T} \) is the black hole temperature obtained via \cite{17} upon eliminating \( P \), and \( 1/\sin \theta \) is an ‘auxiliary factor’ allowing to ‘simplification’ of the critical point topology \cite{18}. Defining the corresponding vector field \( \phi^a = (\phi^S, \phi^\theta) \),

\[
\phi^S = \left( \frac{\partial \Phi}{\partial S} \right)_{P,...}, \quad \phi^\theta = \left( \frac{\partial \Phi}{\partial \theta} \right)_{S,...}, \quad (19)
\]

then yields the following topological current:

\[
\mathbf{j}^\mu = \frac{1}{2\pi} \epsilon^{\mu
u\lambda} \epsilon_{abc} n^a \partial_\nu n^b \quad (20)
\]

upon extending the \((S, \theta)\) space to \( x^\mu = (t, S, \theta) \), with \( n^a = \phi^a/||\phi|| \) and \( \partial_\nu = \frac{\partial}{\partial x^\nu} \). The vector field \( \phi \) vanishes at typical critical points of the system; these are referred to as the zero points of \( \phi \).

It is easy to show that \( \partial_\nu \mathbf{j}^\mu = 0 \), from which we can construct a topological charge:

\[
Q = \frac{1}{2\pi} \int_\Sigma \mathbf{j}^\mu d^2\Sigma_{\mu} = \sum_i w_i \quad (21)
\]

contained within a given region \( \Sigma \) in a surface in parameter space with unit normal \( \sigma^\mu \), with \( d^2\Sigma_{\mu} = \sigma_\mu d^2\Sigma \). The quantity \( w_i \) is the winding number for the \( i \)th zero point of \( \phi \).

#### B. Vortex/anti-vortex pair creation

Returning to our subject of interest, critical points will be located along the \( \theta = \frac{\pi}{2} \) axis in the \((s, \theta)\) plane, which we can reparametrize as the \((v, \theta)\) plane using \cite{5}, since \( s \) is a monotonically increasing function of \( v \) if \( \alpha \leq \sqrt{3} \). For a critical point located at \((v_0, \frac{\pi}{2})\), we can write

\[
v = a \cos \theta + v_0 \quad \theta = b \sin \theta + \frac{\pi}{2} \quad (22)
\]

to parametrize a contour that is near a zero point of \( \phi \). From this, the deflection

\[
\Omega(\theta) = \int_0^\theta \epsilon_{abc} n^a \partial_\nu n^b d\theta \quad (23)
\]

of the vector field along the given contour can be computed. For \( \theta = 2\pi \) the contour surrounds the zero point and yields from \cite{21} the topological charge

\[
Q = \frac{1}{2\pi} \int_0^{2\pi} \epsilon_{abc} n^a \partial_\nu n^b d\theta, \quad (24)
\]

of the critical point. The ‘standard’ critical point is endowed with \( Q = -1 \) and the unstable novel one discussed in \cite{18} has \( Q = 1 \).

Equipped with this classification, we now proceed and calculate the topological charges of the isolated critical point and of the two critical points that merge as the parameter \( \alpha \) is increased to its critical value. These critical points together with the corresponding vector field \( n^a \) in the \((v, \theta)\) plane are displayed in Fig. 3 for \( \alpha = \sqrt{3} \) (left) and \( \alpha = 1.65 \) (right). We observe that the critical point CP1 represents a stable fixed point in the \( v \) direction whereas the CP2 is unstable. The corresponding topological charges are given by:

\[
Q(\text{CP1}) = -1, \quad Q(\text{CP2}) = +1. \quad (25)
\]
FIG. 1: **Isolated critical point:** $\alpha = \sqrt{3}$. The free energy (left) displays two swallow tails that emerge from the same isolated critical point. This is reflected by the corresponding $p-t$ phase diagram that features two first order phase transitions ‘interrupted’ by the isolated critical point, which is consequently characterized by non-standard critical exponents. The diagram is displayed for $d = 7$ and $K = 3$.

FIG. 2: **Emergence of two critical points for** $\alpha < \sqrt{3}$. Left: As the coupling constant $\alpha$ is decreased from its critical value $\alpha_c = \sqrt{3}$, the two swallow tails ‘separate’ and terminate in their own critical points. This corresponds (right) to the two first order-phase transitions terminating at their own critical points CP1 and CP2 – the ‘vortex/anti-vortex pair’ has been created. The diagram is displayed for $K = 3, d = 7$ and $\alpha = 1.65$.

We find that the isolated critical point does not occur at a zero of $\phi$ (nor it obeys (17)); instead this point is the limit point approached by both CP1 and CP2 as $\alpha \to \sqrt{3}$ from below. It is endowed (as expected) with zero topological charge:

$$\text{Q(ICP)} = 0.$$  \hspace{1cm} (26)

As the coupling $\alpha$ increases to its critical value, we see that the vortex/anti-vortex pair has been annihilated. Reversing the process (decreasing $\alpha$ from above), we observe the creation of a vortex/anti-vortex pair.

### IV. CONCLUSIONS

Using the novel framework of assigning topological charges to critical points in black hole thermodynamics, we have re-analyzed the physical interpretation of the isolated critical point in Lovelock gravity. We have shown that while this isolated critical point is very special – occurring for fine tuned Lovelock couplings characterized by $\alpha = \sqrt{3}$ (in $d = 7$ and $K = 3$) and having non-standard critical exponents – its associated topological charge vanishes. However, as one decreases the coupling parameter $\alpha$ from this value, two new critical points emerge – one endowed with a negative topological charge (vortex) and the other with an equal but opposite positive topological charge (anti-vortex). This suggests that an isolated critical point can be interpreted as the onset of vortex/anti-
vortex-like pair creation in black hole thermodynamics.

Such an interpretation, however, should be treated with caution. In particular, recall that although the two emergent critical points are displayed in the same Fig. (3), this is only a ‘projection’ and the actual critical points occur at slightly different physical pressures (see Fig. 2), whose spread becomes wider as α decreases, and thence are never present in the phase diagram ‘simultaneously’. Of course, this feature could change upon considering a different thermodynamic ensemble, for example the one introduced in [11].

An examination of the topological charges of the critical points associated with charged black holes in Gauss-Bonnet gravity [19] indicated that some care needs to be taken in the classification of the critical points. For increasing pressure, conventional/novel critical points were associated with the disappearance/appearance of new phases. As Fig. 2 indicates, our results are in accord with this classification.

Let us finally speculate on potential connections of our findings with topological phase transitions in 2-dimensional spin systems. It is well known, that such systems feature vortex-anti-vortex pair creation at any finite temperature. However for low enough temperatures such vortices are not free and quickly recombine. Nevertheless there exists a critical temperature, known as the Kosterlitz–Thouless temperature [27], above which the presence of vortices is thermodynamically favourable and the vortices ‘roam free’ in the system. It would be interesting to see, if one could correspondingly define some critical α for which the two vortices observed in the black hole system could be considered ‘free’, completing the analogy with ‘topological phase transitions’ in the framework of black hole thermodynamics. A very suggestive in this direction seems α = √5/3, below which one of the vortices ‘has moved to infinity’ and no longer appears in the phase diagram.

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[1] D. Kastor, S. Ray and J. Traschen, Enthalpy and the Mechanics of Ads Black Holes, Class. Quant. Grav. 26 (2009) 195011 [0904.2765].
[2] D. Kubiznak, R.B. Mann and M. Teo, Black hole chemistry: thermodynamics with Lambda, Class. Quant. Grav. 34 (2017) 063001 [1608.06147].
[3] S.W. Hawking and D.N. Page, Thermodynamics of Black Holes in anti-De Sitter Space, Commun. Math. Phys. 87 (1983) 577.
[4] A. Chamblin, R. Emparan, C.V. Johnson and
R.C. Myers, Charged AdS black holes and catastrophic holography, *Phys. Rev. D* 60 (1999) 064018 [hep-th/9902170].

[5] M. Cvetic and S.S. Gubser, Phases of R charged black holes, spinning branes and strongly coupled gauge theories, *JHEP* 04 (1999) 024 [hep-th/9902195].

[6] D. Kubiznak and R.B. Mann, P-V criticality of charged AdS black holes, *JHEP* 07 (2012) 033 [1205.0559].

[7] N. Altamirano, D. Kubiznak and R.B. Mann, Reentrant phase transitions in rotating anti-de Sitter black holes, *Phys. Rev. D* 88 (2013) 101502 [1306.5756].

[8] N. Altamirano, D. Kubizňák, R.B. Mann and Z. Sherkatghanad, Kerr-AdS analogue of triple point and solid/liquid/gas phase transition, *Class. Quant. Grav.* 31 (2014) 042001 [1308.2672].

[9] R.A. Hennigar, R.B. Mann and E. Tjoa, Superfluid Black Holes, *Phys. Rev. Lett.* 118 (2017) 021301 [1609.02564].

[10] E. Witten, Anti-de Sitter space, thermal phase transition, and confinement in gauge theories, *Adv. Theor. Math. Phys.* 2 (1998) 505 [hep-th/9803131].

[11] W. Cong, D. Kubiznak and R.B. Mann, Thermodynamics of AdS Black Holes: Critical Behavior of the Central Charge, *Phys. Rev. Lett.* 127 (2021) 091301 [2105.02223].

[12] W. Cong, D. Kubiznak, R. Mann and M. Visser, Holographic CFT Phase Transitions and Criticality for Charged AdS Black Holes, *2112.14848*

[13] A.M. Frassino, D. Kubiznak, R.B. Mann and F. Simovic, Multiple Reentrant Phase Transitions and Triple Points in Lovelock Thermodynamics, *JHEP* 09 (2014) 089 [1406.7015].

[14] B.P. Dolan, A. Kostouki, D. Kubiznak and R.B. Mann, Isolated critical point from Lovelock gravity, *Class. Quant. Grav.* 31 (2014) 242001 [1407.4783]

[15] S.-W. Wei and Y.-X. Liu, Insight into the Microscopic Structure of an AdS Black Hole from a Thermodynamical Phase Transition, *Phys. Rev. Lett.* 115 (2015) 111302 [1502.00386].

[16] S.-W. Wei, Y.-X. Liu and R.B. Mann, Repulsive Interactions and Universal Properties of Charged Anti-de Sitter Black Hole Microstructures, *Phys. Rev. Lett.* 123 (2019) 071103 [1906.10840].

[17] S.-W. Wei, Y.-X. Liu and R.B. Mann, Ruppeiner Geometry, Phase Transitions, and the Microstructure of Charged AdS Black Holes, *Phys. Rev. D* 100 (2019) 124033 [1909.03887].

[18] S.-W. Wei and Y.-X. Liu, Topology of black hole thermodynamics, *Phys. Rev. D* 105 (2022) 104003 [2112.01706].

[19] F.K. Yerra and C. Bhamidipati, Topology of black hole thermodynamics in Gauss-Bonnet gravity, *Phys. Rev. D* 105 (2022) 104053 [2202.10288].

[20] D. Lovelock, The Einstein tensor and its generalizations, *J. Math. Phys.* 12 (1971) 498

[21] S. Aminneborg, I. Bengtsson, S. Holst and P. Peldan, Making anti-de Sitter black holes, *Class. Quant. Grav.* 13 (1996) 2707 [gr-qc/9604006].

[22] W.L. Smith and R.B. Mann, Formation of topological black holes from gravitational collapse, *Phys. Rev. D* 56 (1997) 4942 [gr-qc/9703007].

[23] R.B. Mann, Topological black holes: Outside looking in, *Annals Israel Phys. Soc.* 13 (1997) 311 [gr-qc/9709039].

[24] D.G. Boulware and S. Deser, String Generated Gravity Models, *Phys. Rev. Lett.* 55 (1985) 2656

[25] J.T. Wheeler, Symmetric Solutions to the Maximally Gauss-Bonnet Extended Einstein Equations, *Nucl. Phys. B* 273 (1986) 732

[26] R.-G. Cai, A Note on thermodynamics of black holes in Lovelock gravity, *Phys. Lett. B* 582 (2004) 237 [hep-th/0311240].

[27] J.M. Kosterlitz and D.J. Thouless, Ordering, metastability and phase transitions in two-dimensional systems, *Journal of Physics C: Solid State Physics* 6 (1973) 1181.