Ab initio study of helium and hydrogen interactions in α-Fe

C. J. Ortiz, R. Vila, and J. M. Pruneda

1Laboratorio Nacional de Fusion por Confinamiento Magnetico, CIEMAT, Avenida Complutense, 40, 28040, Madrid, Spain
2Centre d’Investigació en Nanociència i Nanotecnologia (CSIC-ICN). Campus de la UAB, E-08193 Bellaterra, Spain

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Density Functional Theory (DFT) calculations show a weak interaction between hydrogen and helium in iron, in contrast to previous reports of a strong trapping of hydrogen at helium\footnote{6,7}. The strong preference of He and H to occupy regions with low electronic density (such as vacancies) explains this discrepancy, with vacancy-He and vacancy-H binding forces concealing the repulsive interaction between He and H. Furthermore, Rate Theory simulations based on our DFT-calculated $V_n\text{He}_m\text{H}_p$ cluster energetics predict, as it is observed in some experiments, that synergetic effects could be expected between H and He in iron under irradiation.

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Hydrogen and helium are responsible for more than 98% of nuclear matter in today’s universe. Understanding how they interact is critical for fundamental fields of research such as modeling the evolution of stars\footnote{8–22} or White Dwarfs and the interior of Jovian planets\footnote{23,24}, but also in more technological applications such as sequential H/He implantation in the semiconductor industry,\footnote{25–27} the interior of Jovian planets,\footnote{28} and the direct interaction between helium and hydrogen. By comparing the energetics of these clusters we argue that the direct interaction between helium and hydrogen is marginal (mostly repulsive) and that the strongest possible attraction reported by others\footnote{29} is due to the preferential clustering of helium or hydrogen to iron vacancies and not to the direct interaction between helium and hydrogen. Additional kinetic simulations based on a Rate Theory approach are presented to illustrate the expected effects due to the combined presence of H and He on their evolution under typical irradiation conditions in a fusion environment.

Calculations are performed using the spin-polarized generalized gradient approximation (GGA)\footnote{29} as implemented in the SIESTA code.\footnote{30} Norm-conserving pseudopotentials\footnote{31} with cutoff radii of 1.11, 1.21, 0.71, and 1.06 Å are used for Fe 4s, 4p, 3d and 4f, and a cut-off radius of 0.45 Å for the partial-core correction. H and He pseudopotentials are built with cutoff radii of 0.66 and 0.69 Å, respectively. Valence electrons are described with a linear combination of strictly localized numerical atomic orbitals that vanish outside certain cutoff radii ($r_c$, in Å). Triple-$\zeta$ basis was used for Fe 4s ($r_c = 4.84$) and 3d ($r_c = 3.06$) orbitals and double-$\zeta$ for Fe 4p ($r_c = 3.53$), H and He 1s ($r_c = 2.63$) Additional (single-$\zeta$) polarization orbitals were considered for the H and He electrons. Supercell calculations were first performed on ferromagnetic Fe$_{54}$ boxes to explore the energetics of the different defect configurations, and then repeated for larger Fe$_{128}$ cells. A real space mesh of 0.079 Å was used, and at least $4 \times 4 \times 4$ k-point grid samplings for full relaxation of the atomic positions (forces smaller than 0.04 eV/Å), keeping the volume of the cell fixed.

The zero-point energy (ZPE) is known to be important for H defects\footnote{32} and can be computed from the vibrational frequencies of the normal modes using $ZPE = \frac{1}{2} \sum \nu \hbar \nu$, where the frequencies $\nu$ are obtained assuming the Fe atoms are fixed at their equilibrium positions (Fe mass is much larger than H or He masses), and only the light atoms vibrate (Einstein approximation). The ZPE of $^{1}H_2$ is 0.13 eV, in good agreement with previous results\footnote{33,34} and comparable to the average ZPE for VH$_n$ complexes (0.14 eV/H) reported by Tateyama and Ohno.\footnote{35} Similar values are obtained here for interstitial H in bulk Fe (0.15 eV and 0.19 eV for octahedral and tetrahedral sites), and close to a substitutional He (0.15 eV). This homogeneity of the ZPE for H in different configurations allows to take an average contribution for ZPE of $\sim 0.15$eV per H atom for each complex cluster.\footnote{36}
One approach to quantify the interaction between H and He is through their binding energies. The binding energy of hydrogen to a defect cluster, $V_{n}H_{m}H_{p}$, is defined as the energy gained by trapping an additional H into that cluster to form $V_{n}H_{m}H_{p+1}$. A positive energy indicates an attraction of H to the defect cluster. Based on this representation, Jiang and collaborators suggest a strong hydrogen trapping at helium in W (shown in figure 1 with empty squares), with up to twelve H atoms clustering around one single substitutional He. Trapping at $V_{1}H_{1}$ in iron is notably weaker, with a binding energy of 0.2-0.4 eV with a maximum of five hydrogen atoms around a helium atom. We found that the binding of H to a single vacancy site (no He) is also positive, in good agreement with the results reported by Tateyama and coworkers (light gray circles) that showed that a vacancy can accommodate no more than six hydrogen atoms. Notice however the decrease of ~0.1 eV with respect to the H binding energy to a single vacancy (gray squares) observed when a helium atom is placed at the vacancy (black squares).

We can similarly plot the binding energy of an interstitial He to form a $V_{n}H_{m+1}H_{p}$ cluster (figure 2). In agreement with previous reports, there is a strong self-trapping of He atoms in bulk He (circles in the figure) and an even larger trapping at vacancy complexes. However, the attractive interaction is again decreased (by tenths of eV) when hydrogen atoms are present. In other words, although interstitial He lowers its energy by approaching a vacancy site, the presence of H at the vacancy inhibits He trapping. Further evidence that He and H do not interact strongly is given by the small (even negative) binding energies of He to interstitial H (gray circles). In what follows, we will argue that the positive binding energies for H at $V_{n}H_{m}$ clusters (Fig. 1) or He at $V_{n}H_{p}$ clusters (Fig. 2) are not due to the presence of He or H atoms, but to the vacancy itself.

The interatomic potential for He-H in iron can be determined from the dependence of their binding energy on the He-H distance. We can do this for a relevant subset of the defect configurations studied in this work, which is shown in Fig. 3 together with their corresponding He-H interatomic distances and formation energies. The defect formation energies are defined relative to the chemical potentials taken from bulk $\alpha$-Fe, molecular hydrogen, and the isolated He atom. At the right panels, (a) to (d), different configurations for interstitials of He and H are presented ($V_{0}H_{3}H_{1}$ clusters). Left panels, (e) to (i), show several $V_{1}H_{1}H_{2}$ and $V_{1}H_{1}H_{2}$ clusters, and illustrate how hydrogen affects the stability of $V_{He}$, displacing He from the substitutional site, and moving itself from the preferential tetrahedral occupation of interstitial H, towards the octahedral site.

Central panel in Fig. (3) shows the interatomic potential (binding energies) for He-H pairs computed in bulk $\alpha$-Fe. Filled squares are obtained from the difference between the energy of a HeH complex, and the sum of the energy of interstitial helium and interstitial hydrogen. On the other hand, empty squares are defined from the decrease (per H atom) on the binding energy of one He atom to a vacancy due to the presence of hydrogen atoms at the vacancy (difference between black, dark gray and light gray squares at x=1 in Fig. 2 divided by the number of He atoms). This rough approximation follows with reasonable accuracy the interatomic potential for He and H.

![FIG. 1: Binding energies of interstitial H to a VH (gray squares), VHeH (black squares), and VHe2H (black diamonds) defect clusters. Also shown the binding energies for VH reported by Tateyama et al. ref. 24 (light gray circles) and the binding energies to a single substitutional He in tungsten (empty squares). The horizontal axis refers to the number of H atoms in the initial cluster, x.](image1)

![FIG. 2: Binding energies of interstitial He to $V_{n}H_{m}H_{p}$ complexes. Empty symbols show the binding to interstitials He (circles), VHeH (squares) and $V_{2}He_{x}$ (diamonds) reported in ref. 11. Our results are shown as solid black symbols. Dark (light) gray symbols correspond to the binding energies when one (two) H atoms sit close to the vacancy position. The horizontal axis refers to the number of He atoms in the initial cluster, x.](image2)
computed in vacuum (solid line). It shows that the binding of He and H in bulk iron is only marginal, and could be modeled with the simple two-body potential of isolated He and H atoms. Although there is a small attractive interaction for interstitials in configurations (c) and (d), for which the respective He-H distances are 2.21 Å and 2.86 Å the interaction is clearly repulsive for interatomic distances smaller than \( \sim 2 \) Å, which is the case for all the studied clusters where He and H are trapped in a vacancy.

These \textit{ab initio} calculations show that, although the direct interaction between H and He is weak, they both bind strongly to vacancies, which could be at the origin of the experimentally observed synergic effects.\textsuperscript{25–27} To investigate this point, we simulated the evolution of H and He in \( \alpha \)-Fe under irradiation, in conditions similar to those expected in fusion environment. Nuclear transmutation continuously generates H and He that evolve in the presence of a large amount of defects created by atomic displacements due to collisions between energetic neutrons and lattice atoms. Fusion conditions expected in a DEMO-HCLL-4000MW reactor configuration were considered. A neutron flux of \( 1.52 \times 10^{15} / \text{cm}^2 / \text{s} \) was calculated\textsuperscript{23} and transmutation rates of \( 1.75 \times 10^{12} \) and \( 1.1 \times 10^{13} / \text{cm}^2 / \text{s} \) were obtained for He and H in Fe using the nuclear inventory code ACAB\textsuperscript{24} and the EAF 2007 library\textsuperscript{25}. The generation rate of atomic displacements (Frenkel pairs) corresponding to the neutron spectrum was calculated using the methodology presented in Ref. \textsuperscript{36} and was found to be about \( 1.8 \times 10^{16} / \text{cm}^3 / \text{s} \). To simulate the evolution of H, He and defects we expanded the Rate Theory model developed by Ortiz \textit{et al.}\textsuperscript{26} to predict the diffusion and clustering of He in \( \alpha \)-Fe in the presence of impurities. The various \( V_n \text{He}_m \text{H}_p \) clusters studied from \textit{ab initio} in the present work were added to the model, with their corresponding binding energies and reaction rates, calculated similarly to what is done in Ref. \textsuperscript{25}. It is important to note that the growth of large bubbles containing both H and He is not yet implemented in our model, since their energetics is not known, and only the limited number of small \( V_n \text{He}_m \text{H}_p \) clusters is included.

For our purpose, we considered a temperature of 450°C, which falls in the range of operating temperature (350–500°C) expected for DEMO reactor\textsuperscript{27}, depending on the cooling options. Fig. \textsuperscript{4} reports the evolution of the mean bubble size as a function of irradiation time for the H-He system described above. For comparison, the evolution of the mean bubble size calculated for a system where only He is generated (no H) is also shown. Our simulations clearly predict that a larger mean bubble size is expected when H and He evolve simultaneously. This result follows the trend of the experimental measurements obtained in FeCr alloys by Tanaka \textit{et al.},\textsuperscript{27} who observed larger bubbles when H is present. It is important to note here that only the mean size of He bubbles is shown in Fig. \textsuperscript{4} The contribution of the small \( V_n \text{He}_m \text{H}_p \) clusters to the mean bubble size is minor. Hence, the larger mean size of bubbles that is predicted by our simulations when H is present is not due to large bubbles containing H but to a synergic effect between H and He. Remarkably, a significant effect is predicted though only few \( V_n \text{He}_m \text{H}_p \) configurations are included in the kinetic model.

Our simulations confirm that significant synergic effects between H and He could be expected in some cases, though their direct interaction might be weak, the apparent strong positive interaction reported in the litera-
that conceals the repulsive interaction between He and H. They both tend to be trapped by voids, and this, rather than the He-H interaction, could explain the reduction of hydrogen permeation, and the larger cavity sizes in ferritic alloys simultaneously irradiated with H and He. This has important implications for modeling damage evolution in materials for fusion, that until now have mainly considered the role of isolated He or H aggregates, and not their coexistence. To understand in which conditions synergistic effects could be expected, further work will be needed on the energetics of larger clusters or bubbles containing both He and H in steels, and on the effect of solutes and grain boundaries where bubbles tend to nucleate.

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FIG. 4: Evolution of the mean bubble size in irradiated Fe for a system where only He is present (red dots) and for a system where H and He evolve simultaneously (black squares).

—miguel.pruneda@cin2.es

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