Article

Application of the Generalized Hamiltonian Dynamics to Spherical Harmonic Oscillators

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Abstract: Dirac’s Generalized Hamiltonian Dynamics (GHD) is a purely classical formalism for systems having constraints: it incorporates the constraints into the Hamiltonian. Dirac designed the GHD specifically for applications to quantum field theory. In one of our previous papers, we redesigned Dirac’s GHD for its applications to atomic and molecular physics by choosing integrals of the motion as the constraints. In that paper, after a general description of our formalism, we considered hydrogenic atoms as an example. We showed that this formalism leads to the existence of classical non-radiating (stationary) states and that there is an infinite number of such states—just as in the corresponding quantum solution. In the present paper, we extend the applications of the GHD to a charged Spherical Harmonic Oscillator (SHO). We demonstrate that, by using the higher-than-geometrical symmetry (i.e., the algebraic symmetry) of the SHO and the corresponding additional conserved quantities, it is possible to obtain the classical non-radiating (stationary) states of the SHO and that, generally speaking, there is an infinite number of such states of the SHO. Both the existence of the classical stationary states of the SHO and the infinite number of such states are consistent with the corresponding quantum results. We obtain these new results from first principles. Physically, the existence of the classical stationary states is the manifestation of a non-Einsteinian time dilation. Time dilates more and more as the energy of the system becomes closer and closer to the energy of the classical non-radiating state. We emphasize that the SHO and hydrogenic atoms are not the only microscopic systems that can be successfully treated by the GHD. All classical systems of N degrees of freedom have the algebraic symmetries $O^N + 1$ and $SU^N$, and this does not depend on the functional form of the Hamiltonian. In particular, all classical spherically symmetric potentials have algebraic symmetries, namely $O_4$ and $SU_3$; they possess an additional vector integral of the motion, while the quantal counterpart-operator does not exist. This offers possibilities that are absent in quantum mechanics.

Keywords: generalized Hamiltonian dynamics; spherical harmonic oscillator; classical non-radiating stationary states; algebraic symmetry of classical systems

1. Introduction

The generalized Hamiltonian dynamics (hereafter GHD) was developed by Dirac 70 years ago [1–3]. While the conventional Hamiltonian dynamics employs an assumption that the momenta are independent functions of velocities, Dirac considered a more general situation where momenta are not independent functions of velocities [1–3]. From the physical point of view, the GHD is a purely classical formalism for constrained systems: in the GHD, constraints are incorporated into the Hamiltonian. Dirac developed the GHD with the purpose to apply it to quantum field theory [3].

For the application to the quantum field theory and statistical mechanics, Dirac’s GHD was further developed by a number of authors—see, e.g., papers by Sergi [4–7] and references therein. The focus of Sergi’s works [4–7] was on non-Hamiltonian mathematical structures, including non-Hamiltonian commutators.
In search of a purely classical formalism that can be applied to atomic and molecular physics and can reproduce quantum results classically, Oks and Uzer, in 2002 [8], brought up an idea of choosing integrals of the motion as the constraints. The authors of paper [8] first provided a general description of their formalism. Then they considered hydrogenic atoms as an example. The authors of paper [8] demonstrated that this purely classical formalism allows the existence of classical non-radiating states, so that such states are stable. Remember that, according to the usual classical formalism (including classical electrodynamics), the electron would lose the energy through the radiation and fall into the nucleus: this failure of the usual classical formalism was one of the primary reasons for the birth of quantum mechanics. In distinction, in the purely classical formalism from paper [8] the electron does not fall into the nucleus.

Further, the authors of paper [8] derived the formula for the energy of such classical non-radiating states. They showed that this set of classical energies coincides with the energies of the corresponding quantal stationary states. While obtaining this result, the authors of paper [8] did not “forcefully” quantize any physical quantity describing the atom.

It should be emphasized that the physical interpretation was that the existence of these classical non-radiating states is due to a new kind of a time dilation. This new kind of the time dilation is non-Einsteinian (see paper [9] and book [10]): it has nothing to do with the time-dilation in the theory of relativity.

The purpose of the present paper is to present another application of the Oks–Uzer purely classical formalism [8] within atomic and molecular physics, i.e., to microscopic systems of discrete (rather than continuous) charges. Before proceeding with our presentation, we note in passing that, outside atomic and molecular physics, some authors studied whether there are continuous moving charge distributions that would not radiate—see, e.g., the paper by Goedecke [11] and references therein. It was found that certain continuous charge distributions would not radiate. However, first, the radius of their “orbit” should be less than the size of the charge distribution (so that the the distribution would only “wobble”). Second, this result is valid only for continuous charge distributions, so that atoms and molecules do not qualify.

In our view, the most interesting (and potentially relevant to atomic physics) paper of that series was published by Raju [12]. He considered classical circular orbits of the electron and of the proton in a hydrogen atom. He took into account the relativistic effect of the retardation, due to which the force on the electron is at the “last-seen” position of the proton, while the proton has moved since then. This results in a torque that would initially accelerate the electron and later on decelerate the electron and so on. Then Raju [12] added radiative damping into the consideration, which provides a decelerating torque for the electron. Raju [12] found sets of parameters for which the initially accelerating torque due to the retardation would be totally compensated by the decelerating torque due to the radiative damping. Then Raju [12] stated that “it was prematurely concluded that radiative damping makes the classical hydrogen atom unstable”. However, this statement seems to be incorrect. In reality, within Raju’s concept [12], the radiative damping does make the classical hydrogen atom unstable. Indeed, when the retardation torque compensates the radiative damping torque, this means only that the tangential acceleration of the electron vanishes, but the centripetal acceleration of the electron remains and so does the radiation. Another view of this situation is that the two torques can compensate each other, but one of them is due to the radiation, which carries the energy away from the electron. The electron would therefore continuously lose energy and would fall into the proton. Thus, the concept by Raju [12] did not lead to a non-radiating state of hydrogen atoms.

To avoid any confusion, we also mention that there were attempts to find stable states of hydrogen atoms in frames of a so-called stochastic electrodynamics, where the interaction with the zero-point fluctuations of a vacuum were added into the system to counterbalance the effect of the radiative damping—see, e.g., papers by Puthoff [13], Cole and Zou [14], as well as by Nieuwenhuizen [15], and references from these papers. However, first and foremost, the zero-point fluctuations are purely quantum effects. These kind of works are thus beyond the scope of the present paper devoted to purely
classical description of microscopic systems. Second, the study by Nieuwenhuizen [15] (the latest out of the above three works) showed that this concept leads to the self-ionization of hydrogen atoms from states of relatively low (by absolute value) energy. Thus, this mixed quantum–classical concept actually does not explain the stability of all states of hydrogen atoms.

In the present paper, we apply the GHD to a charged Spherical Harmonic Oscillator (hereafter SHO). The SHO is important both fundamentally and practically. From the theoretical point of view, the SHO is one of the two fundamental microscopic systems (the other one being hydrogenic atoms), characterized by higher-than-geometrical symmetry (i.e., algebraic symmetry) and thus having conserved quantities beyond energy and the angular momentum. The algebraic symmetries of these two fundamental microscopic systems manifest classically by closed orbits and quantally by an “additional” degeneracy of their energy levels. From the practical point of view, the SHO is employed, e.g., in nuclear physics in nuclear shell models.

According to classical physics, the charged SHO is unstable with respect to radiating electromagnetic waves: it would lose its energy and end up in the state of zero energy. Of course, this is contrary to quantum mechanics, according to which the charged SHO has relatively stable stationary states. We show that, similarly to the Oks–Uzer results [8], the GHD—the purely classical formalism—allows the existence of, generally speaking, an infinite number of classical non-radiating states of the SHO.

2. Overview of the General Formalism and of Its Application to Hydrogenic Atoms

This overview is absolutely necessary for readers to facilitate their understanding of the new results (presented in Section 3) and of the conclusions (Section 4). The alternative would be to simply refer to our previous publications [8,10], but, in this case, readers would have to spend lots of time searching through our paper [8] and book [10]. Therefore, out of the respect to readers, we provide here excerpts from our two previous publications [8,10] as quotations (enclosed in the quotation marks).

From [8]:

“Dirac [1–3] considered a dynamical system of N degrees of freedom characterized by generalized coordinates \( q_n \) and velocities \( v_n = dq_n/dt \), where \( n = 1, 2, ..., N \). From the Lagrangian of the system

\[
L = L(q_n, v_n)
\]  

momenta are defined as

\[
p_n = \partial L/\partial v_n \]  

From [10]:

“The quantities \( q_n, v_n, p_n \) can be varied by small amounts \( \delta q_n, \delta v_n, \delta p_n \), respectively. The latter small quantities are of the order of \( \varepsilon \) and the variation should be worked to the accuracy of \( \varepsilon \). As a result of the variation, the set of Equation (2) would not be satisfied any more. This is because their right side would differ from the corresponding left side by a quantity of the order of \( \varepsilon \).

Further, Dirac made a distinction between two types of equations. One type is equations that do not hold after the variation, such as the set of Equation (2). Dirac called them “weak” equations. Below for weak equations, following Dirac, we use an equality sign \( \equiv \) different from the usual equality sign. Another type constitute equations that hold exactly even after the variation, such as Equation (1). Dirac called them “strong” equations. If quantities \( \partial L/\partial v_n \) are not independent functions of velocities, it is possible to exclude velocities \( v_n \) from the set of Equation (2) and obtain one or several weak equations

\[
\varphi(q, p) \equiv 0
\]  

containing only the sets of \( q \) and \( p \) (here and below we skip the suffix of quantities \( q \) and \( p \)).”

From [8]:
"In his formalism, Dirac [1–3] used the following complete system of independent Equations of the type (3):
\[ \varphi_m(q, p) \equiv 0, \quad (m = 1, 2, ..., M) \] (4)

Here the word "independent" means that neither of the \( \varphi \)'s can be expressed as a linear combination of the other \( \varphi \)'s with coefficients depending on \( q \) and \( p \). The word "complete" means that any function of \( q \) and \( p \), which would become zero with the allowance for Equation (4) and which would change by \( \varepsilon \) under the variation, should be a linear combination of the functions \( \varphi_m(q, p) \) from Equation (4) with coefficients depending on \( q \) and \( p \).

Finally, proceeding from the Lagrangian to a Hamiltonian, Dirac [1–3] obtained the following primary result:
\[ H_g = H(q, p) + u_m \varphi_m(q, p) \] (5)

(here and below, the summation over a twice repeated suffix is understood).

From [10]:
"Equation (5) is a strong equation expressing a relation between the generalized Hamiltonian \( H_g \) and the conventional Hamiltonian \( H(q, p) \). Quantities \( u_m \) are coefficients to be determined."

From [8]:
"Generally, they are functions of \( q, v, \) and \( p \); by using the set of Equation (2), they could be made functions of \( q \) and \( p \). It should be emphasized that \( H_g \equiv H(q, p) \) would be only a weak equation - in distinction to Equation (5).

From Equation (5) it is seen that the Hamiltonian is not uniquely determined, because a linear combination of \( \varphi \)'s may be added to it. Equation (4) are called constraints. The distinction between constraints (i.e., weak equations) and strong equations, described above, can be reformulated as follows. Constraints must be employed in accordance to certain rules. Constraints can be added. Constraints can be multiplied by factors (depending on \( q \) and \( p \)), but only on the left side, so that these factors must not be used inside Poisson brackets.

If \( f \) is some function of \( q \) and \( p \), then \( df/dt \) (i.e., a general equation of motion) in the Dirac’s GHD is
\[ df/dt \equiv [f, H] + u_m[f, \varphi_m] \] (6)

where \([f, g]\) is the usual Poisson bracket. Substituting \( \varphi_{m'} \) in Equation (6) instead of \( f \) and taking into account the set of Equation (4), one obtains:
\[ [\varphi_{m'}, H] + u_m[\varphi_{m'}, \varphi_m] \equiv 0. \quad (m' = 1, 2, ..., M) \] (7)

these consistency conditions allow determining the coefficients \( u_m. \)"

From [10]:
"It should be emphasized that the GHD was designed by Dirac specifically for applications to quantum field theory [3], that is, for the purpose totally different from the purpose of Oks-Uzer work [8]."

From [8]:
"The authors of paper [8] reformulated the GHD for atomic and molecular physics where many systems have a higher than geometrical symmetry and therefore possess additional integrals of the motion. Oks and Uzer [8] suggested using integrals of the motion as the constraints in the GHD.

In their general formalism, they considered a classical atomic or molecular system of \( N \) degrees of freedom, possessing \( M \) classical integrals of the motion \( A_m(q, p), \quad m = 1, 2, ..., M. \) They wrote the generalized Hamiltonian in the form (see Equations (4) and (5)):
\[ H_g = H(q, p) + u_m[A_m(q, p) - A_0m], \quad A_0m = \text{const.} \] (8)
Here $A_{0m}$ is the value of $A_m(q, p)$ in a particular state of the motion, so that in this state
\[ A_m(q, p) - A_{0m} = 0. \] (9)

Since the quantities $A_m(q, p)$ are integrals of the motion, their Poisson bracket with $H(q, p)$ vanishes and the consistency condition (7) reduces to the form
\[ u_m[A_{m'}, A_m] = 0. \] (m' = 1, 2, ..., M). (10)

From [10]:
"The set of Equation (10) allows determining the coefficients $u_m$.
Specifically, for a hydrogenic atom of the nuclear charge $Z$, the integrals of the motion (other than the energy) are the angular momentum $L = r \times p$ and the Runge-Lenz vector (see, e.g., [16])
\[ A(r, p) = \frac{(rp^2 - p(rp))}{(\mu Z e^2)} - \frac{r}{r}, \]
where $\mu$ is the reduced mass. Therefore, Oks and Uzer [8] presented the generalized Hamiltonian in the form:
\[ H_g = \frac{p^2}{2\mu} - \frac{Ze^2}{r} + u \times (r \times p - L_0) + w \times (A(r, p) - A_0). \] (11)

Here $L_0$, $A_0$, and the energy $H_0$ are connected by the well-known relation [16]:
\[ L_0^2 = \mu Z^2 e^4 (A_0^2 - 1)/(2H_0). \] (12)

From [8]:
"The consistency conditions $[r \times p, H_g] = 0$, $[A(r, p), H_g] = 0$ resulted into the following equations for the unknown vector-coefficients $u$ and $w$:
\[ u \times L_0 + w \times A_0 = 0, \]
\[ u \times A_0 - 2w \times A_0 H_0/(\mu Z^2 e^4) = 0. \] (13)

By using the consistency Equation (13), Oks and Uzer [8] reduced the number of yet unknown coefficients to just one, which they denoted as $B$. Of course, $B$ was yet unknown function of energy $H_0$ in the particular state of the atom. Oks and Uzer [8] showed that in terms of $B(H_0)$, the generalized Hamiltonian and the equations of the motion take the following form:
\[ H_g = \frac{p^2}{2\mu} - \frac{Ze^2}{r} + 2B(H_0)H_0|\mathbf{L}_0 \times (r \times p)/\mathbf{M}_0^2 - (1 - A_0 \cdot A(r, p))/(1 - A_0^2)| \] (14)
\[ \frac{dr}{dt} = \left[1 + B(H_0)\right]p/\mu \] (15)
\[ \frac{dp}{dt} = -\left[1 + B(H_0)\right]Ze^2 r/3 \]

From [10]:
"The Equation of the motion (15) differ from their conventional form only by the factor $[1 + B(H_0), A_0]$. Therefore, the transformation of the time
\[ t' = [1 + B(H_0)]t \] (16)
the equations of the motion with respect to the new time $t'$ would be formally brought back to their conventional form.

Thus the authors of paper [8] came to the following central point. In the above generalized formalism, the trajectory of the atomic electron remains the same as in the conventional formalism. However, the generalized period $T_g$ and the generalized frequency $\omega_g$ differ from their conventional values $T_0$ and $\omega_0$ as follows:
\[ T_g = T_0/[1 + B(H_0)] \] (17)
\[ \omega_g = \omega_0[1 + B(H_0)] = 1 + B(H_0)|2H_0|^{3/2}/D^{1/2}, \] (18)
(in Equation (18), the explicit expression for the Kepler frequency $\omega_0$ has been used)."

From [8]:

“Equation (18) clearly demonstrates that the generalized formalism allows the existence of such state (or states) of the motion, where $\omega_g = 0$ despite $H_0 \neq 0$ (the conventional formalism allows to be $\omega_0 = |2H_0|^{3/2}/D^{1/2} = 0$ only for $H_0 = 0$). This is a state (or states) where $B(H_0) = -1$. Therefore, such state or states would not emit the electromagnetic radiation, would not lose energy for the radiation, and would thus constitute stable states of the classical atom.”

The authors of paper [8] showed that there is infinite number of the energies of the classical stationary states—just as in the corresponding quantum solution.

From [10]:

“Oks and Uzer [8] pointed out that in a classical non-radiating stable state, one has $dr/dt = dp/dt = 0$, so that $r(t) = r_0$ and $p(t) = p_0$, where $r_0$ and $p_0$ are some vector constants. Thus, the electron is motionless, but its momentum differs from zero. This is not surprising: the momentum $p$ is a more complex physical quantity than the velocity $v = dr/dt$. For example, for a charge in an electromagnetic field characterized by a vector-potential $A$, it is also possible to have $v = [p - eA/(mc)]/m = 0$ while $p = eA/(mc) \neq 0$ [17].

It is also very important to emphasize that the physics behind such classical non-radiating states is a new kind of time-dilation expressed by Equation (16): a non-Einsteinian time-dilation, as pointed out in book [10]. The closer the energy of the system to the energy of the classical non-radiating state, the more dilates the time. At the classical non-radiating state, the time gets dilated infinitely, so that the frequency $\omega_g$ in Equation (18) vanishes and so does the radiation.”

3. New Results

We consider a charged Spherical Harmonic Oscillator (SHO). The “conventional” conserved quantities are the energy $E$ and the angular momentum vector $M$, the conservation of the latter following from the geometrical (spherical) symmetry of this system. It is well-known that the SHO also possesses another set of conserved quantities, whose conservation is the consequence of the higher-than-geometric (algebraic) symmetry:

$$I_{mn} = p_m p_n/\mu + k x_m x_n, m = 1, 2, 3, n = 1, 2, 3$$ (19)

Here, $p_m$ and $x_m$ are the Cartesian components of the momentum $p$ and of the radius-vector $r$, respectively; $\mu$ is the mass of the SHO. Obviously, $I_{mn} = I_{nm}$, so that there are generally only six independent conserved quantities $I_{mn}$. The unperturbed Hamiltonian $H$ can be actually expressed via some of the conserved quantities from Equation (19) as follows:

$$H = (I_{11} + I_{22} + I_{33})/2$$ (20)

It is well known that the motion is limited to a plane. We choose the $x_3$-axis (the $z$-axis) perpendicular to the orbital plane. Then, the dynamical variables are $x_1, p_1, x_2, p_2$.

In order to study whether classical non-radiative states of the SHO are possible, it should be sufficient to consider the generalized Hamiltonian $H_g$, which differs from $H$ only by the addition of the constraints corresponding to the conserved quantities responsible for the algebraic symmetry, i.e., the conserved quantities from Equation (19), but only those of them that are relevant to the motion in the orbital plane:

$$H_g = (I_{11} + I_{22})/2 + B_{11}(E) (I_{11} - I_{11,0}) + B_{22}(E) (I_{22} - I_{22,0}) + B_{12}(E) (I_{12} - I_{12,0})$$ (21)

where $I_{mn,0}$ are the values of these conserved quantities in the particular state of the system; $E$ is the energy of the system in a particular state of the motion.
The conserved quantities $I_{mn}$ “commute” with each other: the Poisson bracket of any two of them vanishes. Therefore, the consistency conditions from Equation (10) in this case reduce to equating to zero the Poisson brackets of the components of the angular momentum $\mathbf{M}$ with the second term in the right side of Equation (21):

$$[M_i, a_{mn} I_{mn}] = [e_{ijq} x_i] p_{jr} B_{mn} (p_m p_n / \mu + k x_m x_n)] = 0 \quad (22)$$

where $e_{ijq}$ is the Levi-Civita symbol.

The calculations of the Poisson brackets from Equation (22), with the subsequent substitution of $I_{mn}$ by $I_{mn,0}$ (as required by the GHD), lead to the following equations:

$$B_{22} I_{12,0} = B_{12} I_{22,0} \quad (23)$$

$$B_{12} I_{11,0} = B_{11} I_{12,0} \quad (24)$$

From Equation (19), it is obvious that the quantities $I_{11}$ and $I_{22}$ are non-negatively defined.

For definiteness, we assume that $I_{11,0}$ differs from zero, i.e., $I_{11,0} > 0$. Then, from Equations (23) and (24), it is easy to obtain

$$B_{12} = B_{11} I_{12,0} / I_{11,0}, \quad B_{22} = B_{11} I_{22,0} / I_{11,0} \quad (25)$$

Thus, the consistency conditions help reduce the unknown coefficients in the generalized Hamiltonian $H_g$ from three to one, so that $H_g$ can be represented in the form:

$$H_g = (I_{11} + I_{22}) / 2 + B_{11} (E) \left[(I_{11} - I_{11,0}) + I_{22,0} / I_{11,0} + I_{22} - I_{22,0} + I_{12,0} / I_{11,0} (I_{12} - I_{12,0})\right] \quad (26)$$

Based on the Hamiltonian $H_g$ from Equation (26) and using $dx_i / dt = \partial H_g / \partial p_i$, $dp_i / dt = -\partial H_g / \partial x_i$, we find the following equations of motion:

$$dx_1 / dt = [(I + 2B_{11})] p_1 + (B_{11} I_{12,0} / I_{11,0}) p_2, \quad dx_2 / dt = (I + 2B_{11} I_{22,0} / I_{11,0}) p_2 + (B_{11} I_{12,0} / I_{11,0}) p_1 / \mu \quad (27)$$

$$dp_1 / dt = -k(I + 2B_{11}) x_1 + (B_{11} I_{12,0} / I_{11,0}) x_2, \quad dp_2 / dt = (I + 2B_{11} I_{22,0} / I_{11,0}) p_2 + (B_{11} I_{12,0} / I_{11,0}) x_1 \quad (28)$$

By differentiation of Equation (27) with respect to time and substituting Equation (28) into the outcome, we obtain the following system of equations:

$$d^2 x_1 / dt^2 = -\omega^2_0 [((I + 2B_{11})^2 + (B_{11} I_{12,0} / I_{11,0})^2] x_1 + 2(B_{11} I_{12,0} / I_{11,0}) [1 + B_{11} (I + 1 I_{22,0} / I_{11,0})] x_2, \quad (29)$$

$$d^2 x_2 / dt^2 = -\omega^2_0 [((I + 2B_{11} I_{22,0} / I_{11,0})^2 + (B_{11} I_{12,0} / I_{11,0})^2] x_2 + 2(B_{11} I_{12,0} / I_{11,0}) [1 + B_{11} (I + 1 I_{22,0} / I_{11,0})] x_2$$

where

$$\omega_0 = (k / \mu)^{1/2} \quad (30)$$

is the “unperturbed” frequency of the oscillator.

We seek a solution of system (29) in the form:

$$x_1 = \exp(i \omega_g t), \quad x_2 = \alpha \exp(i \omega_g t), \quad \alpha = \text{const} \quad (31)$$

here, $\omega_g$ is the (yet unknown) generalized frequency of the oscillator.

Substituting $x_1$ and $x_2$ from Equation (31) into the first equation in Formula (29), we obtain:

$$\omega^2 / \omega_0^2 = (1 + 2B_{11})^2 + (B_{11} I_{12,0} / I_{11,0})^2 + 2\alpha (B_{11} I_{12,0} / I_{11,0}) [1 + 2B_{11} (1 + 1 I_{22,0} / I_{11,0})] \quad (32)$$
Substituting $x_1$ and $x_2$ from Equation (31) into the second equation in Formula (29), we obtain:

$$
\frac{\omega_g^2}{\omega_0^2} = (1 + 2B_{11}I_{22,0}/I_{11,0})^2 + (2/\alpha)(B_{11}I_{12,0}/I_{11,0})(1 + 2B_{11}(1 + I_{22,0}/I_{11,0}))
$$

(33)

For Equations (32) and (33), which have the same left sides, to be compatible with each other, their right sides should also be equal to each other. By equating the right sides of Equations (32) and (34), after some simplifications, we find that the parameter $\alpha$ must satisfy the following quadratic equation:

$$
\alpha^2 - 2\gamma\alpha - 1 = 0, \quad \gamma = (I_{22,0} - I_{11,0})/I_{12,0}
$$

(34)

The two solutions of Equation (34) are

$$
\alpha = \gamma \pm (\gamma^2 + 1)^{1/2}
$$

(35)

Obviously, $\alpha_+ > 0$ while $\alpha_- < 0$. Physically, these two solutions correspond to the two opposite directions of the revolution along the orbit (see Equation (31)).

Equation (33) can be represented in a more explicit form:

$$
\frac{\omega_g^2}{\omega_0^2} = (4 + 2\alpha_+\delta + \delta^2)B_{11}^2 + 2(2 + \alpha_+\delta)B_{11} + 1
$$

(36)

where we temporarily introduce the following notation:

$$
\epsilon = (1 + I_{22,0}/I_{11,0}), \quad \delta = I_{12,0}/I_{11,0}
$$

(37)

Using Equation (35), it is easy to find out that

$$
4 + 2\alpha_+\epsilon\delta + \delta^2 = (2 + \alpha_+\delta)^2
$$

(38)

so that Equation (35) simplifies to

$$
\omega_g^2/\omega_0^2 = [(2 + \alpha_+\delta)B_{11} + 1]^2
$$

(39)

which is equivalent to the following:

$$
\omega_g/\omega_0 = |(2 + \alpha_+\delta)B_{11} + 1|
$$

(40)

Coming back to the original notations, we rewrite Equation (40) in the form

$$
\omega_g/\omega_0 = |[1 + I_{22,0}/I_{11,0} \pm ([I_{22,0}/I_{11,0} - 1]^2 + I_{12,0}^2/I_{11,0}^2)^{1/2}]B_{11}(E) + 1|
$$

(41)

where we have restored the argument $E$ of the coefficient $B_{11}(E)$. It is seen that, for each direction of the revolution of the charged particle in the orbital plane, there is a value of $B_{11}(E)$, for which the generalized frequency is $\omega_g$ vanishes and so is the radiation. These non-radiating (stationary) states correspond explicitly to the following values, $B_{11+}(E_{st})$ and $B_{11-}(E_{st})$ of $B_{11}(E)$, where the subscript “st” stands for “stationary”:

$$
B_{11+}(E_{st}) = -1/[1 + I_{22,0}/I_{11,0} \pm ([I_{22,0}/I_{11,0} - 1]^2 + I_{12,0}^2/I_{11,0}^2)^{1/2}]
$$

(42)

for $\alpha = \alpha_+$ and

$$
B_{11-}(E_{st}) = -1/[1 + I_{22,0}/I_{11,0} \pm ([I_{22,0}/I_{11,0} - 1]^2 + I_{12,0}^2/I_{11,0}^2)^{1/2}]
$$

(43)

for $\alpha = \alpha_-$. Remember that $\alpha_\pm (\gamma)$ is given by Equation (35), where $\gamma = (I_{22,0} - I_{11,0})/I_{12,0}$.
Figure 1 shows a three-dimensional plot of $B_{11^+}$ (denoted in the plot for brevity as $B_+$) versus $I_{22,0}/I_{11,0}$ (denoted in the plot as C) and $I_{12,0}/I_{11,0}$ (denoted in the plot as D).

Figure 2 shows a three-dimensional plot of $B_{11^-}$ (denoted in the plot for brevity as $B_-$) versus $I_{22,0}/I_{11,0}$ (denoted in the plot as C) and $I_{12,0}/I_{11,0}$ (denoted in the plot as D).
Here is an intermediate summary of the above results. By employing the GHD, we have proven the existence of the classical non-radiating states of the charged spherical harmonic oscillator—similarly to the corresponding results of paper [8] for hydrogenic atoms.

Physically, this is the manifestation of a non-Einsteinian time-dilation. Time dilates more and more as the energy of the system becomes closer and closer to the energy of the classical non-radiating state:

\[ t' = \frac{1 + |I_{22,0}|/I_{11,0}}{1 - |I_{22,0}|/I_{11,0}} \pm \sqrt{1 + \left[ |I_{22,0}|/I_{11,0} - 1 \right]^2 + |I_{12,0}|^2/I_{11,0}^2 \sqrt{2}} \pm |B_{11,0}| + 1 |t \] \hspace{1cm} (44)

At the classical non-radiating state, the time gets dilated infinitely. As a result, the frequency of the revolution along an elliptical orbit \( \omega_{st} \) in Equation (41) vanishes—consequently, the radiation also vanishes.

In the important particular case of \( I_{22,0} = I_{11,0} \), corresponding to the circular orbits, the above formulas can be simplified as follows. In this case, from Equation (34) follows \( \gamma = 0 \), so that from Equation (35) we get \( \alpha_+ = 1 \) and \( \alpha_- = -1 \), as it should be for the circular orbits (see Equation (31)). Then Equation (36) simplifies to

\[ \omega_{g}/\omega_{0} = |(2 - |I_{12,0}|/I_{11,0})B_{11}(E) + 1] \] \hspace{1cm} (45)

where the plus sign corresponds to \( \alpha = 1 \) and the minus sign corresponds to \( \alpha = -1 \).

It is seen that, in this particular case, the generalized frequency \( \omega_{g} \) vanishes (and so does the radiation) at the following values of \( B_{11}(E) \):

\[ B_{11,+}(E_{st}) = -1/(2 + |I_{12,0}|/I_{11,0}) \text{ for } \alpha = 1 \] \hspace{1cm} (46)

and

\[ B_{11,-}(E_{st}) = -1/(2 - |I_{12,0}|/I_{11,0}) \text{ for } \alpha = -1 \] \hspace{1cm} (47)

Obviously, Equation (47) is valid, except if \( |I_{12,0}|/I_{11,0} = 2 \). In the exceptional case, Equation (45) yields a trivial result: \( \omega_{g} = \omega_{0} \).

The primary result for the circular orbits is non-trivial. Namely, there are classical non-radiating states, corresponding to \( B_{11}(E) = B_{11,+}(E_{st}) \) for \( \alpha = 1 \) or \( B_{11}(E) = B_{11,-}(E_{st}) \) for \( \alpha = -1 \).

Thus, for each direction of the revolution of the charged particle in the orbital plane, there is one value of \( B_{11}(E) \)—given by Equations (42) and (43) in the general case of the elliptical orbits or by Equations (46) and (47) for the particular case of the circular orbits—for which the radiation vanishes. The fact that, for each direction of the revolution, there is only one value of \( B_{11}(E) \), does not mean that there is only one classical stationary state. Indeed, if the dependence of \( B_{11} \) on the energy \( E \) is oscillatory (with the amplitude greater than or equal to the absolute value of the right side of Equation (42) for \( \alpha = \alpha_+ \), or with the amplitude greater than or equal to the absolute value of the right side of Equation (43) for \( \alpha = \alpha_- \)), then there would be an infinite number of the energies of the classical stationary states \( E_{st} \)—just as in the corresponding quantum solution.

Here is an example, illustrating the statement from the previous sentence for the case of circular orbits—for the subcase of \( \alpha = 1 \) chosen for definiteness. Let us consider the following dependence of \( B_{11,+} \) on the energy \( E \):

\[ B_{11,+}(E) = -|\cos[(\pi(E - C)/(E_{st,0} - C))]|/(2 + |I_{12,0}|/I_{11,0}) \] \hspace{1cm} (48)

where \( E_{st,0} \) is the energy of the lowest non-radiating state (the ground state) and both \( E \) and \( E_{st,0} \) are measured in units of \( \hbar \omega_{0} \). In Equation (48), \( C \) is a constant, which is an analog of the Maslov index [18], which, for spherically symmetric potentials, is equal to 1/2 (see, e.g., the textbook [19]). With \( C = 1/2 \), Equation (48) takes the form

\[ B_{11,+}(E) = -|\cos[(\pi(E - 1/2)/(E_{st,0} - 1/2))]|/(2 + |I_{12,0}|/I_{11,0}) \] \hspace{1cm} (49)
From Equation (49), it is easy to find out that \( E = E_{st,n} \), where

\[
E_{st,n} - 1/2 = (n + 1)(E_{st,0} - 1/2), \quad n = 0, 1, 2, \ldots ,
\]

(50)

where the quantity \( B_{1+} \) satisfies Equation (46), so that the sequence of values \( E_{st,n} \) from Equation (50) is the sequence of the energies of the classical non-radiating stationary states. More explicitly,

\[
E_{st,n} = (n + 1)E_{st,0} - n/2
\]

(51)

If \( E_{st,0} = 3/2 \), then the sequence of values \( E_{st,n} \) from Equation (51) would coincide with the corresponding quantum results.

4. Conclusions

We extended the applications of the GHD to a charged Spherical Harmonic Oscillator (SHO). We demonstrated that, by using the higher-than-geometrical symmetry (i.e., the algebraic symmetry) of the SHO and the corresponding additional conserved quantities, it is possible to obtain classical non-radiating (stationary) states of the SHO. Generally, there is an infinite number of such states of the SHO—just as was the case for hydrogenic atoms, as was shown in paper [8]. Both the existence of the classical stationary states of the SHO and the infinite number of such states are consistent with the corresponding quantum results.

Physically, the existence of the classical stationary states is the manifestation of a non-Einsteian time-dilation. Time dilates more and more as the energy of the system becomes closer and closer to the energy of the classical non-radiating state.

It should be emphasized that we obtained the above new results from first principles. We did not use any quantization postulates or any input from experiments.

It is worth mentioning that the SHO and hydrogenic atoms are not the only microscopic systems that can be successfully treated by the GHD. Indeed, all classical systems of \( N \) degrees of freedom have the algebraic symmetries \( O_{N+1} \) and \( SU_N \), and this does not depend on the functional form of the Hamiltonian. In particular, all classical spherically symmetric potentials have algebraic symmetries, namely \( tO_4 \) and \( SU_3 \); they possess an additional vector integral of the motion, while the quantal counterpart-operator does not exist [20–22]. (This fact was employed in paper [9], where the authors successfully applied the GHD to a modified Coulomb potential.) This offers possibilities that are absent in quantum mechanics, as noted in paper [8].

Since there are lots of classical systems possessing an algebraic symmetry and, therefore, having additional integrals of the motion, as mentioned in the previous paragraph, it should be obvious that the classical systems studied in papers [8,9] and in the present paper do not constitute a comprehensive list. For example, another fundamental physical system—an electron in the field of two stationary nuclei—is a good candidate to be treated by GHD. Indeed, this system has an additional integral of the motion—the projection of the super-generalized Runge–Lenz vector on the internuclear axis, the latter vector being derived in paper [23].

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References

1. Dirac, P.A.M. Generalized Hamiltonian dynamics. *Canad. J. Math.* **1950**, *2*, 129–148. [CrossRef]
2. Dirac, P.A.M. Generalized Hamiltonian dynamics. *Proc. R. Soc. A* **1958**, *246*, 326–332. [CrossRef]
3. Dirac, P.A.M. *Lectures on Quantum Mechanics*; Academic: New York, NY, USA, 1964; Reprinted by Dover Publications, 2001.
4. Sergi, A. Generalized bracket formulation of constrained dynamics in phase space. *Phys. Rev. E* **2004**, *69*, 21109. [CrossRef] [PubMed]
5. Sergi, A. Phase space flows for non-Hamiltonian systems with constraints. *Phys. Rev. E* **2005**, *72*, 31104. [CrossRef] [PubMed]

6. Sergi, A. Non-Hamiltonian commutators in quantum mechanics. *Phys. Rev. E* **2005**, *72*, 66125. [CrossRef] [PubMed]

7. Sergi, A. Statistical mechanics of quantum-classical systems with holonomic constraints. *J. Chem. Phys.* **2006**, *124*, 24110. [CrossRef] [PubMed]

8. Oks, E.; Uzer, T. Application of Dirac’s generalized Hamiltonian dynamics to atomic and molecular systems. *J. Phys. B At. Mol. Opt. Phys.* **2002**, *35*, 165–173. [CrossRef]

9. Camarena, A.; Oks, E. Application of the generalized Hamiltonian dynamics to a modified coulomb potential. *Int. Rev. At. Mol. Phys.* **2010**, *1*, 143–160.

10. Oks, E. *Breaking Paradigms in Atomic and Molecular Physics*; World Scientific: Singapore, 2015.

11. Goedecke, G.H. Classically radiationless motions and possible implications for quantum theory. *Phys. Rev.* **1964**, *135*, B281–B288. [CrossRef]

12. Raju, C.K. The Electrodynamic 2-body problem and the origin of quantum mechanics. *Found. Phys.* **2004**, *34*, 937–962. [CrossRef]

13. Puthoff, H.E. Ground state of hydrogen as a zero-point-fluctuation-determined state. *Phys. Rev. D* **1987**, *35*, 3266–3269. [CrossRef] [PubMed]

14. Cole, C.D.; Zou, Y. Quantum mechanical ground state of hydrogen obtained from classical electrodynamics. *Phys. Lett. A* **2003**, *317*, 14–20. [CrossRef]

15. Nieuwenhuizen, T.M. On the stability of classical orbits of the hydrogen ground state in stochastic electrodynamics. *Entropy* **2006**, *18*, 135. [CrossRef]

16. Landau, L.D.; Lifshitz, E.M. *Mechanics*; Pergamon Press: Oxford, UK, 1960.

17. Landau, L.D.; Lifshitz, E.M. *Classical Theory of Fields*; Pergamon Press: Oxford, UK, 1960.

18. Mukunda, N. Dynamical symmetries and classical mechanics. *Phys. Rev. 1967*, *155*, 1383–1386. [CrossRef]

19. Bacry, H.; Ruegg, H.; Souriau, J.M. Dynamical groups and spherical potentials in classical mechanics. *Commun. Math. Phys.* **1966**, *3*, 323–333. [CrossRef]

20. Fradkin, D.M. Existence of the dynamic symmetries O4 and SU3 for all classical central potential problems. *Progr. Theor. Phys.* **1967**, *37*, 798–812. [CrossRef]

21. Maslov, V.P.; Fedoriuk, M.V. *Semi-Classical Approximation in Quantum Mechanics*; Springer: Berlin, Germany, 1981.

22. Landau, L.D.; Lifshitz, E.M. *Quantum Mechanics*; Pergamon Press: Oxford, UK, 1965.

23. Kryukov, N.; Oks, E. Super-generalized Runge-Lenz vector in the problem of two coulomb or newton centers. *Phys. Rev. A* **2012**, *85*, 54503. [CrossRef]

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