Majorana approach to the stochastic theory of lineshapes

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Motivated by recent Mössbauer experiments on strongly correlated mixed-valence systems, we revisit the Kubo-Anderson stochastic theory of spectral lineshapes. Using a Majorana representation for the nuclear spin we demonstrate how to recast the classic lineshape theory in a field-theoretic and diagrammatic language. We show that the leading contribution to the self-energy can reproduce most of the observed lineshape features including splitting and lineshape narrowing, while the vertex and the self-consistency corrections can be systematically included in the calculation. This new approach permits us to predict the line-shape produced by an arbitrary bulk charge fluctuation spectrum providing a model-independent way to extract the local charge fluctuation spectrum of the surrounding medium. We also derive an inverse formula to extract the charge fluctuation from the measured lineshape.

I. INTRODUCTION

Resonant spectroscopy of two-level systems is a powerful tool to study the environment in which they are embedded, and include photo and X-ray absorption, electron spin resonance, nuclear magnetic resonance and even current spectroscopy of a quantum dot. An important example is Mössbauer spectroscopy, where nuclear transitions in a solid are studied by the recoil-free resonant absorption/emission of γ-ray photons, providing a sensitive probe of low frequency electric charge fluctuations and the magnetic field texture in the material. Recently, the advent of synchrotron-based-radiation as a new source of hard X-rays has opened up a wide range of materials to Mössbauer study, with the possibility of new insights into strongly correlated systems, such as YbAlB$_4$.

In spite of this general applicability and recent demand, the stochastic theory of lineshapes has been largely unchanged since the seminal works of Anderson, Kubo and later Blume on the topic. The classic theory is model-based, providing information about the lineshape in environments with specialized Gaussian or Markovian dynamics. This motivates us to revisit the problem of spectral lineshapes in two level systems, recasting the problem in a modern framework. Our work leads us to conclude that there is a wide parameter regime in Mössbauer spectroscopy, in which the entanglement between two-level system and the environment is weak enough for the measured absorption lineshape to provide model-independent information about the spectrum of charge fluctuations in the surroundings.

II. THE MODEL

Here, we study the spectroscopic lineshape of a two-level system (probe) immersed in a fluctuating environment [Fig. 1]. Quite generally, the Hamiltonian of a two-level system can be described by a pseudo-spin variable $I_z = \pm 1/2$ in a Zeeman field which is modulated by another degree of freedom

$$H = (\omega_0 + \lambda \sigma_f)(I^2 + 1/2) + H_{\text{rest}}(\sigma_f).$$

The level splitting is modified by the fluctuating bosonic variable $\sigma_f$ whose dynamics is governed by $H_{\text{rest}}$. $I^2$ commutes with both $H_{\text{rest}}$ and $\sigma_f$ but the latter two do not commute, resulting in resonance-frequency fluctuations. A main assumption is that the level-splitting (which corresponds to a Nuclear excitation energy in Mössbauer spectroscopy) is large enough ($\omega_0 \gg T$) so that the two-level system is in its ground state. Without much restriction, we assume that $\langle \sigma_f \rangle = 0$ and its average is absorbed in $\omega_0$. Most of our discussion is applicable to arbitrary $\sigma_f$ and $H_{\text{rest}}$ and the latter can be quite complicated. However, we will make a distinction between the cases when $\sigma_f$ is a continuous variable and when it is has discrete eigenvalues. The latter occurs, for example when $\sigma_f = \pm 1$ is a digital variable caused by the capacitive coupling of the probe to the occupation of a nearby fermionic $f$-level. On the other hand, if the probe is coupled to many independent $f$-levels, the average becomes a continuous variable which according to the central limit theorem, will have Gaussian dynamics.
In spectroscopic studies, the above system is coupled to an additional photonic degree of freedom, \( H' = H + g(I^+a + h.c.) + \omega a^\dagger a \) which is used to probe the line-shape of the excited state at absorption energies \( \omega \sim \omega_0 \). The absorbed power (positive) is written (see app. 1) in terms of the retarded Green’s function of the spins

\[
P(\omega) = -\omega \Omega^2 \chi''_{+}(\omega + i\eta),
\]

where \( \chi''_{+}(\omega + i\eta) \) denotes the imaginary part of the Fourier transform of the retarded function \( \chi''_{+}(t) = -i\theta(t) (\langle I^- (t), I^+ (0) \rangle) \), evaluated in thermal equilibrium, and \( \Omega^2 \) is the field strength of the incoming radiation.

III. MAJORANA REPRESENTATION AND THE SELF-ENERGY DIAGRAMS

To use the machinery of field theory, we need to represent the spin in terms of canonical fields. We use a Majorana representation of the spin,\(^{10}\)\(^{11}\)\( \vec{\eta} = -\frac{1}{\sqrt{2}} \eta \times \vec{\eta} \) which expresses the spin in terms of three\(^{13}\) Majorana fermions \( \eta^i, i = 1, 2, 3 \) obeying the anti-commutation relation \( \{\eta^i, \eta^j\} = \delta^{ij} \). The advantage of this representation, is that it avoids the use of a constraint and furthermore, the spin dynamics can be directly read-off from the one-particle Green’s function of the Majorana fermions.\(^{10}\)\(^{11}\) This is because the Majorana composite \( \phi = -2i\eta_1 \eta_2 \eta_3 \) commutes with the spin \( [\phi, \vec{I}] = 0 \) and the Hamiltonian \( [\phi, H] = 0 \), and is thus a constant of motion. Moreover, since \( 2\phi \vec{I} = \vec{\eta} \) and \( \phi^2 = 1/2 \), it follows that

\[
\langle I^a (t) I^b (t') \rangle = \frac{1}{2} \langle \eta^a (t) \eta^b (t') \rangle. \tag{3}
\]

It is convenient to combine two of the Majoranas into a single Dirac fermion, \( d^\dagger \equiv (\eta^1 + i\eta^2)/\sqrt{2} \), so that \( I^z = d^\dagger d - 1/2 \) and \( I^+ = \sqrt{2} \eta^3 d^\dagger \). The occupied/unoccupied states of the \( d \)-level correspond to the up/down states of the probe isospin, \( I_z \). By Eq. (3), \( I^- (t) I^+ = d(t) d^\dagger \) and \( I^z I^- (t) = d^\dagger (t) d(t) \) are true at all times and therefore, \( \langle I^- (t) I^+ (0) \rangle = \langle d(t) d^\dagger (0) \rangle = \langle d^\dagger (0) d(t) \rangle \). Fourier transform of the left side is the imaginary part of the spin susceptibility, whereas the right-side has the ‘wrong’ sign to be a fermionic retarded function. Instead, it is the Keldysh function of the鬼魂, and by the fluctuation-dissipation theorem can be related to the imaginary part of the retarded function. Therefore, we obtain\(^{10}\) (see app. 2)

\[
\chi''_{+}(\omega + i\eta) = \tanh(\beta \omega/2)G''_R(\omega + i\eta) - \frac{\pi c}{\lambda^2}. \tag{4}
\]

Here, \( \beta = 1/T \) is the inverse temperature and \( G''_R(\omega + i\eta) \) is the imaginary part of the Fourier transform of the retarded Green’s function, \( G''_R(t) = -i\theta(t) \langle d(t), d^\dagger (0) \rangle \). Quite generally, we expect the absorption function, \( \sim G''_R(\omega + i\eta) \) to be narrow function centered at \( \omega_0 \), so combining Eqs. 2 and 4, we have \( P(\omega) \sim -G''_R(\omega + i\eta) \) and the proportionality constant is \( c = \omega_0 \tanh(\beta \omega_0/2)\Omega^2 \). The area under the resonance is constant, leading to the sum rule \( \int d\omega P(\omega) = -c \int d\omega G''_R(\omega + i\eta) = \pi c \{d, d^\dagger\} = \pi c \).

The advantage of the Majorana representation is that one can apply standard field theory techniques, developing a Feynman expansion for the one-particle Green’s function \( G_d(d') = -\langle T, d(d') d^\dagger (0) \rangle \), which we can immediately convert to a spin correlation function of the probe isospin using Eq. 5. Taking advantage of the Dyson’s equation for the Green’s function, \( G_d(z) = [z - \omega_0 - \Sigma_d(z)]^{-1} \), the \( d \)-fermion Green’s function is described in terms of the self-energy \( \Sigma_d(z) \) of the \( d \)-fermion (Fig. 2b). The calculation of the spin dynamics then reverts to a calculation of the self-energy of the \( d \) fermion. The relevant diagrams in the expansion of the self-energy to order \( O(\lambda^4) \) are shown in Fig. 2. Assuming that temperature is much smaller than \( \omega_0 \), we have \( f(\omega_0) \approx 0 \), where \( f(\omega) = [e^{\beta \omega} + 1]^{-1} \) is the Fermi-Dirac distribution. Therefore, to a good approximation, it suffices to consider an exclusively forward-time bare \( d \)-level propagator \( g_{d}(\tau) = [\langle -\theta(\tau) + f(\omega_0)\rangle e^{-\omega_0 \tau}] \approx -\theta(\tau) e^{-\omega_0 \tau} \) in the calculations. This considerably simplifies the diagrams, because it eliminates all diagrams that involve fermions propagating backwards in time, i.e any diagrams with additional fermion loops (e.g. the last two Feynman diagrams of Fig. 2b). Physically, this approximation means that there is no back-action from the \( d \)-level on the charge fluctuations, e.g. by the last diagram in Fig. 2b.

The forward-time restriction on the \( d \)-propagator limits us to a single \( d \)-fermion branch in the Green’s function, and these diagrams re-sum to

\[
G_d(\tau) = -e^{-\omega_0 \tau} \langle T, e^{\lambda \int_0^\tau d\tau' \sigma_f (\tau')} \rangle. \tag{5}
\]

Alternatively, this formula can be obtained using the methods similar to the orthogonality catastrophe problem by writing \( d(\tau) = e^{\tau H_-} de^{-\tau H_+} \) where \( H_{\pm} \) correspond to the \( I_z = \pm 1/2 \) sectors of the Hamiltonian, respectively.

Expanding the exponent inside the bracket of Eq. 4, we have to evaluate \( n \)-point correlation functions of \( \sigma_f \). These have a disconnected part (sum of all possible Wick’s contractions, the Gaussian subset in Fig. 2b) and a connected part caused by the interaction vertices of \( H_{\text{int}} \). For a continuous \( \sigma_f \) variable with Gaussian dynamics, the connected-part contributions to Eq. 5 are zero and the bracket becomes \( \exp[\lambda^2 \int_0^\tau d\tau_1 \int_0^\tau d\tau_2 \langle T, \sigma_f (\tau_1) \sigma_f (\tau_2) \rangle] \) (see app. 3). We can simplify this to write

\[
G_d(\tau) \to -e^{-\omega_0 \tau} \exp \left[ -\lambda^2 \int_0^\tau dx (\tau - x) \chi_C (x) \right]. \tag{6}
\]

Here, \( \chi_C (\tau) = -\langle T, \sigma_f (\tau) \sigma_f (0) \rangle \) is the correlation function of the fluctuations and we used that \( \chi_C (-\tau) = \chi_C (\tau) \). This formula, due to Anderson\(^{12}\) is quite precise within the Gaussian-action assumption, but unfortunately due to the appearance of the \( \chi_C \) within an-
nalizing the exponent in this formula and generalizing it:

\[
\Gamma \text{ contains steady-state probabilities and is the solution to }
\]

Here, \( \sigma \) is ±1. We can write

\[
G_d(\omega + i\eta) = \frac{1}{2} \frac{1}{\omega + i\eta - \omega_0 + \lambda} \Gamma_1 + \frac{1}{2} \frac{1}{\omega + i\eta - \omega_0 - \lambda} \Gamma_2
\]

where we have absorbed the statistical mixture into a perturbative self-energy \( \Sigma_d(\omega + i\eta) = \lambda^2(\omega + i\eta - \omega_0) \) for the ensemble-averaged Green's function. Note that this agrees with Eq. (6) and Eq. (7), while it disagrees with Eq. (6) using \( \chi_C(\tau) = -1 \). The structure of the \( G_d \) function here is reminiscent of the zero-coupling limit of the Anderson impurity problem. Indeed, for the case of \( H_{\text{res}} = \sum_k \epsilon_k f_k^\dagger f_k \) with a bandwidth \( \Gamma \), we have an Ising Kondo (in a Zeeman field) which can be solved exactly\(^1\) and \( \chi_{\rightarrow-}(\omega) \) exhibits orthogonality catastrophe physics. Usually, there is an energy scale \( T_s(\Gamma, \lambda) \sim \Gamma \), below which a coherent entanglement between the probe and \( H_{\text{res}} \) is established. However, at the weak-coupling limit \( T \gg T_s \) there is little entanglement and a perturbative self-energy accurately matches the exact non-perturbative result\(^1\).

One of observations of this paper is that the simplest “single-photon” exchange approximation to the self-energy of the \( d \)-fermion provides an interpolation between motionally narrowed and double-line limits of the absorption line, and is capable of reproducing most of the observed features of the lineshape for general \( H_{\text{res}} \) in the \( T \gg \Gamma \) limit. This is the leading order contribution to the self-energy, \( \Sigma_d(\tau) = -\lambda^2 g_d(\tau) \chi_C(\tau) \), which to order \( \lambda^2 \) is exact and thus it goes beyond both forward-time propagator and Gaussian-action approximations mentioned earlier. Taking the Fourier transform, and using \( g_d(\omega_n) = [\omega_n - \omega_0]^{-1} \) and analytically continuing to real frequencies we obtain (see app. 5)

\[
\Sigma_d''(\omega_0 + \omega + i\eta) = \lambda^2[f(\omega_0) - n_B(-\omega)]\chi_C''(\omega + i\eta) \tag{8}
\]

where \( n_B(\omega) = \frac{e^{\beta\omega} - 1}{e^{\beta\omega} + 1} \) is the Bose-Einstein distribution. Eq. (8) is the central result of this paper. We have related the retarded function of the fluctuations to the imaginary part of the self-energy. The latter is related to the absorbed power by \( \Sigma_d''(\omega) = -cP(\omega)/[P(\omega) + P''(\omega)] \) where \( c \) is the proportionality constant introduced after Eq. (4) and \( P(\omega) \) is the Hilbert transform of the absorbed power \( P(\omega) \approx P''(\omega) \) so that \( P' + iP'' \) obeys the Kramers-Krönig relation

\[
P'(\omega) = \mathcal{P} \int \frac{dw'}{\pi} \frac{P(\omega')}{\omega' - \omega}.
\]

Having obtained \( \Sigma_d'' \) one can then use Eq. (8) to extract the charge susceptibility from the lineshape. A main assumption is that the fluctuations have a small bandwidth \( \Gamma \ll T \ll \omega_0 \). Thus, we can drop \( f(\omega_n) \) and do a high-temperature expansion of \( n_B(-\omega) \approx -1/\beta\omega \) to write \( \Sigma_d''(\omega_0 + \omega) \approx \lambda^2 \chi_C''/\beta\omega \). This together with Kramers-Krönig relation for \( \chi_C(\omega + i\eta) \) provides another sum-rule that relates the area under the dissipative part.
of self-energy (or $\chi''_{\sigma}/\beta\omega$) to the static charge susceptibility $\int d\omega \Sigma_{xy}(\omega) = \pi \lambda^2 T \chi_{0y}^\alpha$. Moreover, combining these equations we find the simple and useful inverse formula

$$\chi_C(\omega) \approx \frac{e}{\lambda^2 T} \frac{\omega}{P'(\omega) + iP(\omega)}. \quad (9)$$

This formula can be used to directly extract the charge susceptibility of the medium from the observed lineshape.

**IV. EXAMPLES**

As a first example, we look at the case when $\chi''_{\sigma}/\beta\omega = -\pi \delta_{f}(\omega)$ is a $\Gamma$-broadened delta function at the origin. A Lorentzian $\delta_{f}(\omega)$ function, defined in Eq. (15), corresponds to an exponential decay in the time-domain $\langle \sigma_f(\tau)\sigma_f(0) \rangle \sim e^{-\Gamma \tau}$ and a single-rate Markovian process in which $\Gamma$ is the rate of switching between the two $\sigma_f = \pm 1$ states (see app. 4). In this case, Kramers-Kronig relation is trivial and we obtain $\Sigma_{xf}(\omega + \omega + i\eta) = \frac{\lambda^2}{(\omega^2 - \lambda^2)^2 + \omega^2 T^2}$. Inserting this into the Green’s function gives us the Archer-Anderson formula:

$$G''_{xf}(\omega_0 + \omega + i\eta) = -\frac{\lambda^2 \Gamma}{(\omega^2 - \lambda^2)^2 + \omega^2 T^2}, \quad (10)$$

previously obtained from the completely different approach of Eq. (7) (see app. 4). This agreement is a remarkable observation whose origin is unclear to us at the moment. Eq. (10) is plotted in Fig. (3a,b). In the slow-switching case $\omega = \omega_0 \pm \lambda$, whereas in the fast-switching case $\omega \ll 1$, we have two well separated $\delta$-peaks at $\omega = \omega_0 \pm \lambda$, and on a rescaled frequency axis $\tilde{\omega} = (\omega - \omega_0)/\Gamma$, we get a single peak at the origin $\Gamma G''_{xf}(\tilde{\omega}) \rightarrow x^2(\tilde{\omega}^4 + \tilde{\omega}^2 + x^4)^{-1}$ which is Lorentzian within $\tilde{\omega} \leq \omega$ and has non-Lorentzian $\tilde{\omega}^{-4}$ tails at $\tilde{\omega} \gg \omega$.

It is instructive to compare Eq. (10) to the Gaussian result, Eq. (6) which can be computed with little effort in this special case. Defining $\alpha = x^2$ and from $\chi(\tau) = -e^{i\tau\tau}$, it is straightforward (app. 4) to find

$$\Gamma G''_{xf}(\tilde{\omega}) = -\pi e^{\alpha} \sum_{n \geq 0} f_n(\alpha) \delta_{(n+\alpha)}(\tilde{\omega}). \quad (11)$$

This is a summation of $(n + \alpha)$-broadened $\delta$-functions all centred at $\tilde{\omega} = 0$ with the coefficients $f_n(\alpha) \equiv \sum_{m} J_m(-\alpha)(-\alpha) J_{n-\alpha}(\tilde{\omega})$, given in terms of ordinary $J_m(x)$ and modified $I_m(x)$ Bessel functions. This function is plotted in Fig. (3c) and (3d) along with Eq. (10). The two functions agree at $x \ll 1$, which can be understood, qualitatively, by the fact that the time-averaged $\langle \sigma_f \rangle$ in both continuous and discrete cases, spend most of the time around zero. We argue below that the rainbow diagrams [e.g. second diagram of Fig. (1b)] and vertex corrections [e.g. third diagram of Fig. (1b)] that are included in the Gaussian subset but not in Eq. (5), are negligible for $x \ll 1$ but become important at finite $x$ in agreement with these plots.

To account for the vertex correction one has to solve the Bethe-Salpeter equation for the vertex function (see appp. 6). To the leading order, and from Fig. (2d), we...
can estimate these vertex corrections to be \( \lambda^2 \rightarrow \kappa \lambda^2 \) where \( \kappa \approx 1 - \lambda^2 \int d\tau \chi_C(\tau) \int_0^\tau d\tau' \approx 1 - x^2 \). This is analogous to the Migdal’s theorem in superconductivity and confirms that vertex corrections are negligible in the fast regime \( x \ll 1 \) but important in the slow regime.

The rainbow diagrams are part of the so-called non-crossing approximation (NCA) class and can be included within our formalism by self-consistently upgrading the bare propagator \( g_d \) in the self-energy to the exact Green’s function \( G_d \). This is schematically shown in Fig. (2c) and can be implemented iteratively. Figs. (3 c,d) include the first two iterations as well as the NCA result, starting from Eq. (10). While nothing changes for \( x \ll 1 \), in the opposite regime of finite \( x \) the lineshape becomes closer to the Gaussian result. To the leading order we can see this, by expanding the self-energy as \( \Sigma_d(\omega) = \Sigma_d(\omega_0) + (\omega - \omega_0) \partial_{\omega} \Sigma_d(\omega_0) + \cdots \) and only keeping the leading terms. Therefore, \( g_d(i\omega_n) \) is replaced by \( G_d(i\omega_n) \approx Z[\omega_n - \omega_0 + i\text{sign}(n)]^{-1} \) where \( Z = [1 - \partial_{\omega} \Sigma_d(\omega_0)]^{-1} \) is the wavefunction renormalization of the \( d \)-level and \( \Lambda = -Z\Sigma_d''(\omega_0) \) is its broadening.

We can estimate \( \Lambda/\Gamma \propto x^2 \) and \( Z = [1 - x^2]^{-1} \) in our example. Calculating the self-energy with this propagator has the effect of renormalizing the coupling constant \( \lambda^2 \rightarrow \lambda^2 Z \), but more importantly the self-energy is blurred by a convolution with a \( \Lambda \)-broadened delta function \( \delta_\Lambda(\omega) \) (see app. 5). Therefore, in the limit of fast/slow charge fluctuations rainbow diagrams have a negligible/important effect, in agreement with the numerical result. The fact that rainbow diagrams and vertex corrections drive the lineshape towards the Gaussian result is not surprising, since these additional diagrams can be understood as if the two-level system is coupled to many independent mutually-non-interacting \( f \)-levels which average into a Gaussian lineshape.

The second example is when a combination of fast and slow switching processes are involved and where the power of the field theoretical approach shows up. These more complex scenarios happen for example in Kondo systems as spin-flip processes involve fast charge switching between unpaired single-spin and empty or doubly occupied states. In Fig. (3), we have considered a case where \( \chi''_C/\beta \omega \) contains a slow peak at the origin with the width \( \Gamma_S \) and a fast mode at a finite frequency with the width \( \Gamma_F \). The charge susceptibility is then

\[
\frac{\chi''_C(\omega)}{\beta \omega} = -\pi \cos^2 \theta \delta_{\Gamma_F}(\omega) - \pi \sin^2 \theta \delta_{\Gamma_F}(|\omega| - \omega_F).
\]  

where \( \omega_F, \Gamma_S \) and \( \Gamma_F \) are kept constant and \( \theta \) is varied in Figs. (3a,b) to change the relative strength of fast and slow fluctuations for different ratios of \( \Gamma_F < \Gamma_S \) and \( \Gamma_F = \Gamma_S \). Figs. (3c,d) shows the corresponding lineshapes we would expect from such a medium. A direct consequence of the sum rules is that the total area under \( \chi''_C/\omega \) is a constant. As is clear from these results, the existence of the fast mode and its breadth can have significant effects on the lineshape, a detailed study of which we leave for the future.

V. CONCLUSION

To conclude, we have revisited the classic stochastic theory of lineshapes and used a Majorana representation of the spin to recast it in a field-theoretical language. We have shown that the leading contribution to the self-energy provides an interpolation between Markovian and Gaussian results as well as featuring a flexibility to study non-trivial charge fluctuations.

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12. The linear conductance through a quantum dot is \( G = \alpha \frac{1}{4} \int d\omega |f(\omega)| G''(\omega + i\eta) \), where \( \alpha \) is a measure of coupling symmetry of the dot to the source and drain. At low temperature, \( G \propto G''(\omega + i\eta) \) and if the dot is in the weak-coupling (to the leads) regime, the system is again described by the Hamiltonian of Eq. (1) but with \( \omega_0 \approx 0 \). Therefore, in this case the forward-time propagator is not a good approximation, and the full Eq. (5) has to be used.
written the perturbation many-body energies of \( \delta H \) where we have employed completeness we can write Hamiltonian in Eq. (1). Then using \( 14 \)
P. Coleman, *Introduction to Many-Body Physics*, Cambridge Univ. Press (2016).

**APPENDIX**

The following appendices include the proof of some of the statements or formulas given in the manuscript.

1. **Absorbed power formula**

From Fermi’s Golden rule, the absorption rate is related to the transition rate, given by

\[
\Gamma(\omega) = \sum_f \frac{2\pi}{\hbar} |\langle f | \delta H | i \rangle|^2 \delta(E_f - E_i) \quad (13)
\]

Here \( \delta H = g(a^\dagger I - h.c.) \) and \( E_i \) and \( E_f \) are the exact many-body energies of \( H_0 = H + \omega a^\dagger a \) where \( H \) is the Hamiltonian in Eq. (1). Then using

\[
\delta(E_f - E_i) = \frac{1}{2\pi\hbar} \int_{-\infty}^{+\infty} dt e^{-it(E_f-E_i)/\hbar} \quad (14)
\]

can write

\[
\Gamma(\omega) = \frac{1}{\hbar^2} \int_{-\infty}^{+\infty} dt e^{-it(E_f-E_i)/\hbar} \sum_f \langle i | \delta H | f \rangle \langle f | \delta H | i \rangle
\]

\[
= \frac{1}{\hbar^2} \int_{-\infty}^{+\infty} dt \sum_f \langle i | e^{itH_0/\hbar} \delta H e^{-itH_0/\hbar} | f \rangle \langle f | \delta H | i \rangle
\]

\[
= \frac{1}{\hbar^2} \int_{-\infty}^{+\infty} dt \langle i | \delta H(t) \delta H(0) | i \rangle, \quad (15)
\]

where we have employed completeness \( \sum_f |f\rangle \langle f| = 1 \) and written the perturbation

\[
\delta H(t) = e^{itH_0/\hbar} \delta H e^{-itH_0/\hbar}
\]

in the interaction representation with respect to \( H_0 = H + \omega a^\dagger a \). As a result of this interaction representation \( I^\pm, a \) and \( a^\dagger \) develop time-dependence. Doing a thermal average over the initial states, the total transition rate is then given by [the frequency dependence on the right is implicit in the time-evolution of the \( \delta H(t) \)].

\[
\Gamma(\omega) = \frac{1}{\hbar^2} \int_{-\infty}^{+\infty} dt \langle \delta H(t) \delta H(0) \rangle \quad (16)
\]

which can be written as

\[
\Gamma(\omega) = \left( \frac{g}{\hbar} \right)^2 \int_{-\infty}^{+\infty} dt \{ n e^{i\omega t} \langle I^- (t) I^+ (0) \rangle + (1 + n) e^{-i\omega t} \langle I^+ (t) I^- (0) \rangle \}
\]

\[
= ig^2 [n \chi^>_{-} (\omega) + (1 + n) \chi^<_{-} (\omega)] \quad (17)
\]

where \( n \equiv \langle a^\dagger a \rangle \) as the average number of photons created at the steady state (a measure of input power). We have used the definitions of the greater/lesser functions

\[
\chi^>_{\pm} (t) = -i \langle I^- (t) I^+ (0) \rangle, \quad \chi^<_{\pm} (t) = -i \langle I^+ (0) I^- (t) \rangle.
\]

We assume in this paper that the two-level system (described by the spin \( \vec{I} \) and weakly probed by the gamma photons) is in thermal equilibrium with its surrounding and considering the \( \omega_0 \gg T \) condition, it is mainly in its ground state. Therefore, the greater/lesser susceptibilities are related to the imaginary part of retarded susceptibility by the fluctuation-dissipation theorem

\[
\chi^>_{\pm} (\omega) = 2i [n_B (\omega)] \chi''_{\pm} (\omega), \quad \chi^<_{\pm} (\omega) = 2i n_B (\omega) \chi''_{\pm} (\omega) \quad (18)
\]

Anticipating that \( \chi''_{\pm} (\omega) \) is a narrow resonance at \( \omega_0 \gg T \), we can approximate the Bose function \( n_B (\omega) \approx [e^{\omega_0} + 1]^{-1} \approx 0 (\omega_0) \). Therefore, \( \chi^>_{\pm} (\omega \gg T) \approx 0 \) and \( \chi^<_{\pm} (\omega \gg T) \approx 2i \chi''_{\pm} (\omega) \). Inserting these into Eq. (17) we conclude

\[
\Gamma(\omega) = -2\Omega^2 \text{Im} \left[ \chi''_{\pm} (\omega) \right] \quad (19)
\]

where we have defined \( \Omega \equiv g \sqrt{n} \). Eq. (19) is the total number of transitions. Dividing this by two (absorptions) and writing \( P(\omega) = \frac{1}{2} \hbar \omega \Gamma(\omega) \) gives Eq. (2).

2. **Majorana representation of the spin**

   a. **Commutation relations**

Using the index notation \( S^l = -(i/2) \epsilon_{lab} \eta^a \eta^b \), and using the anticommutation algebra \( \{ \eta_a, \eta_b \} = \delta_{ab} \), we can confirm that the spin operators faithfully reproduce the SU(2) algebra

\[
[S^l, S^m] = -\frac{1}{4} \epsilon_{labmcd} [\eta_a \eta_b, \eta_c \eta_d] = -\frac{1}{4} \epsilon_{labmcd} \left\{ \delta^{dc} \eta^a \eta^d - \delta^{cd} \eta^a \eta^c - \delta^{ad} \eta^c \eta^d + \delta^{ac} \eta^d \eta^b \right\} = \frac{1}{2} \left[ \eta^l \eta^m - \eta^m \eta^l \right] = i \epsilon_{lmb} S^p. \quad (20)
\]

and that furthermore, \( (S^a)^2 = 1/4, \ S^2 = \frac{3}{4} \), confirming that this is a faithful representation of a spin-1/2 operator.

   b. **Relation between Keldysh and retarded functions**

As a reminder, we can write the Keldysh \( G^K \) and the difference between retarded \( G^R \) and advanced \( G^A \) functions as

\[
G^K_a = G^> + G^<, \quad G^> - G^< = G^R - G^A, \quad (21)
\]
in terms of the greater and lesser functions, defined by
\[ G_0^\omega(t) = -i \langle d(t)d(0) \rangle, \quad G_0^\omega(t) = +i \langle d^\dagger(0)d(t) \rangle. \]

Using the cyclic properties of the trace, \( G_0^\omega(t - i\beta) = -G_0^\omega(t) \) or equivalently, \( G_0^\omega(\omega) = -e^{-i\omega t}G_0^\omega(\omega) \) in the frequency domain. Combining these with Eqs. (21) gives \( G_0^\omega(\omega) = [1 - 2f(\omega)]2iG_0^\omega(\omega) \) which leads to Eq. (4).

3. Diagrammatic proofs

The Hamiltonian in the \( d \)-level representation is
\[ H = (\omega_0 + \lambda \sigma_f)d^d + H_{\text{rest}}[\sigma_f] \]  
and we are interested in \( G_d(\tau) = -\langle T_\tau d(\tau)d(0) \rangle \). A Brute force perturbation theory in \( \lambda \) is
\[ G_d(\tau) = -\sum_{n=0}^{\infty} \frac{(-\lambda)^n}{n!} \int_0^\beta d\tau_1 \cdots d\tau_n \]
\[ \langle T_\tau d(\tau)d(0)d(\tau_1)d(\tau_1) \cdots d^d(\tau_n)d(\tau_n) \rangle \times \]
\[ \langle T_\tau \sigma_f(\tau_1) \cdots \sigma_f(\tau_n) \rangle. \]  
(23)

The \( d \)-operators commute with \( H_{\text{rest}}[\sigma_f] \) and therefore, there is no additional interaction vertex and the above correlation function factorizes into \( d \) and \( \sigma_f \) parts. Therefore, we can write
\[ G_d(\tau) = -\sum_{n=0}^{\infty} \frac{(-\lambda)^n}{n!} \int_0^\beta d\tau_1 \cdots d\tau_n \]
\[ \langle T_\tau d(\tau)d(\tau_1)d(\tau_1) \cdots d^d(\tau_n)d(\tau_n) \rangle \times \]
\[ \langle T_\tau \sigma_f(\tau_1) \cdots \sigma_f(\tau_n) \rangle. \]  
(24)

We can apply Wick’s contraction to these non-interacting \( d \)-levels, but before that we introduce an approximation that simplifies the resulting diagrams.

a. Forward-time propagators and single-branch simplification

The propagator for the free \( d \)-level is
\[ g_d(\tau) \equiv -\langle T_\tau d(\tau)d(0) \rangle \]
\[ = \{-\theta(\tau)[1 - f(\omega_0)] + \theta(\tau)f(\omega_0)\}e^{-\omega_0 \tau} \]
\[ = [f(\omega_0) - \theta(\tau)]e^{-\omega_0 \tau} \approx -\theta(\tau)e^{-\omega_0 \tau}. \]  
(25)

We have used that the (average) resonant frequency is much larger than temperature \( (\omega_0 \gg T) \) and therefore
\[ f(\omega_0) = \frac{1}{e^\beta\omega_0 + 1} \approx e^{-\omega_0/T} \approx 0. \]  
(26)

In the following, we show that this approximation significantly simplifies the Wick’s contractions. We can categorize the Feynman diagrams with the number of inter-connected \( d \)-level propagators. If going from \( d(0) \) to \( d(\tau) \) we pass through all the \( d \)-level propagators, there is only one branch. Inter-disconnected diagrams may appear for example to order \( \lambda^2 \) as the tadpole diagram or \( \lambda^4 \) as the fermion bubbles inside the propagator of the \( \sigma_f \) charge fluctuation \[ H_{\text{term}}(2b) \]. The simplification of the forward-time approximation is that, \( d \)-fermion bubbles are suppressed. For example, the fermion bubbles’ contribution is proportional to \( g(\tau_1 - \tau_2)g(\tau_2 - \tau_1) \approx \theta(\tau_1 - \tau_2)\theta(\tau_2 - \tau_1) = 0 \). Therefore, single-branch assumption is justified.

Moreover, since \( g_d(\tau < 0) = 0 \), in doing Wick’s contraction, we must have the ascending order \( \tau_1 < \tau_2 < \cdots < \tau_n \) and they are all connected in one branch. If the order is violated even once, e.g. \( g(\tau_1 - \tau_3) \) the result will be zero because \( \tau_2 \) will appear somewhere else and make the time-argument of the propagator negative and zero result due to forward-time approximation. With this, we have
\[ G_d(\tau) = \sum_{n=0}^{\infty} \lambda^n \int_0^\tau \int_0^{\tau_1} \int_0^{\tau_2} \cdots \int_0^{\tau_n} \]
\[ \langle \sigma_f(\tau_1) \cdots \sigma_f(\tau_n) \rangle \]
(26)

Using Eq. (24) we find
\[ G_d(\tau) = -e^{-\omega_0 \tau} \sum_{n=0}^{\infty} \frac{(-\lambda)^n}{n!} \int_0^\tau \int_0^{\tau_1} \int_0^{\tau_2} \cdots \int_0^{\tau_n} \]
\[ \langle \sigma_f(\tau_1) \cdots \sigma_f(\tau_n) \rangle \]
(27)

which is Eq. (5).

b. Orthogonality catastrophe proof

An alternative proof of Eq. (27) is by writing the evolution operator of the \( d(\tau) = e^{H_d^{\tau} e^{-\tau H}} \), where \( H \) is given by Eq. (22). To have non-zero result, the \( d \)-state has to be filled prior to the acting with the annihilation operator and empty afterwards. Therefore, we can write
\[ d(\tau) = e^{H_d^{\tau} e^{-\tau H} + \lambda \sigma_f} + H_{\text{rest}}^d e^{-\tau H_{\text{rest}}} \]  
(28)

After this replacement, the \( d \) operator has done its job and can be dropped out \[ H_{\text{rest}} \];
\[ G_d(\tau) = -\langle T_\tau d(\tau)d^\dagger \rangle \]
\[ \approx -\langle e^{H_d^{\tau} e^{-\tau H} + \lambda \sigma_f} \rangle_{H_{\text{rest}}} \]
\[ = -e^{-\omega_0 \tau} \langle e^{H_d^{\tau} e^{-\tau H} + \lambda \sigma_f} \rangle_{H_{\text{rest}}}. \]  
(29)

We have added a subindex \( H_{\text{rest}} \) to the first correlator to indicate the Hamiltonian that appears in the corresponding Boltzman factor. Also, we have used the forward-time approximation to drop a factor proportional to
\( e^{-\beta \omega_0} \) in the expansion of the partition function
\[
Z = \text{Tr} \left[ e^{-\beta H_{\text{rest}}} \right] + e^{-\beta \omega_0} \text{Tr} \left[ e^{-\beta (H_{\text{rest}} + \lambda \sigma_f)} \right] \\
\approx \text{Tr} \left[ e^{-\beta H_{\text{rest}}} \right].
\]  
(30)

The product of the two exponents \( U(\tau) = e^{\tau H_{\text{rest}} - \tau (H_{\text{rest}} + \lambda \sigma_f)} \) in Eq. (29) is the definition of the time-evolution in the interaction picture with respect to \( H_{\text{rest}} \). To see that, just take its derivative with respect to \( \tau \) and observe the corresponding Schrödinger equation.

\[
d\frac{d}{d\tau} U(\tau) = e^{\tau H_{\text{rest}}} [H_{\text{rest}} - (H_{\text{rest}} + \lambda \sigma_f)] e^{-\tau (H_{\text{rest}} + \lambda \sigma_f)} = -\lambda \sigma_f(\tau) U(\tau)
\]  
(31)

where \( \sigma_f(\tau) = e^{\tau H_{\text{rest}}} \sigma_f e^{-\tau H_{\text{rest}}} \). Integrating this equation, we can write

\[
G_d(\tau) \approx -e^{-\omega_0 \tau} \left\langle T_{\tau} e^{\lambda \int_0^\tau d\tau' \sigma_f(\tau')} \right\rangle.
\]  
(32)

c. Gaussian subset

Generally, the \( n \)-point function of \( \sigma_f \)-s can all have non-trivial connected and disconnected ones. If \( H_{\text{rest}} \) is Gaussian, all the connected terms vanish. The set of all disconnected terms [exact for Gaussian \( H_{\text{rest}} \)], e.g. the diagram shown in Fig. (5)a can be written as Anderson wrote

\[
G_D^D(\tau) = -e^{-\omega_0 \tau} \exp \left[ -\frac{1}{2} \int_0^\tau \int_0^\tau d\tau_1 d\tau_2 \langle T_{\tau} \sigma_f(\tau_1) \sigma_f(\tau_2) \rangle \right] = e^{-\omega_0 \tau} \exp \left[ -\lambda^2 \int_0^\tau (\tau - x) \chi_C(x) dx \right].
\]  
(33)

This contains all order of \( \lambda^{2n} \), each being sum of all possible two-point contractions of \( 2n \) \( \sigma_f \)-operators. Some of these are shown in the following figure.

\[ \begin{array}{c}
\text{fermion} \\
\sigma_f \text{-boson}
\end{array} \]

\( a \)

\( \begin{array}{c}
0 \tau_1 \tau_2 \tau_3 \quad \text{f} \\
\text{f} \quad \text{f} \quad \text{f} \quad \text{f}
\end{array} \]

\( b \)

FIG. 5: (a) An example of an \( O(\lambda^n) \) diagram within Gaussian (and forward-time propagator) approximation. (b) The diagrams obtained from the leading self-energy contribution.

4. Markovian case

a. Proof of Eq. (7)

In the Markovian case, the main assumption is that the reduced density matrix of a single \( \sigma_f \) is diagonal and it obeys the classical rate equation. Then Eq. (4) has to be interpreted as a statistical average. Following Anderson we divide the interval \((0, \tau)\) to \( N \) segments and write Eq. (4) as

\[
\left\langle T_{\tau} e^{\lambda \int_0^\tau d\tau' \sigma_f(\tau')} \right\rangle = \exp \left[ \frac{\lambda}{N} \sum_{m=1}^N \sigma_{f,m} \right] = \prod_{m=1}^N \exp \left[ \frac{\lambda}{N} \sigma_{f,m} \right]
\]  
(34)

\( p(\sigma_{f,N}; \sigma_{f,N-1}; \cdots; \sigma_{f,1}) \) is the probability that the variable \( \sigma_f \) is equal to \( \sigma_{f,1} \) at time \( \tau/N \), \( \sigma_{f,2} \) at time \( 2\tau/N \), \( \sigma_{f,n} \) at time \( n\tau/N \) and so on. The Markovian assumption means that the probability of the state \( \sigma_{f,n} \) only depends on the state \( \sigma_{f,n-1} \) and hence

\[
p(\sigma_{f,N}; \sigma_{f,N-1}; \cdots; \sigma_{f,1}) = p(\sigma_{f,N}; \sigma_{f,N-1}; \tau/N) \times p(\sigma_{f,N-1}; \sigma_{f,N-2}; \tau/N) \times \cdots \\
p(\sigma_{f,2}; \sigma_{f,1}; \tau/N) p(\sigma_{f,1}).
\]  
(35)

Using the (imaginary time) rate equation \( d\bar{p}/d\tau = -i\Gamma \bar{p} \), one can write

\[
p(\sigma_{f,m}; \sigma_{f,m-1}; \tau/N) = 1 - i \frac{\tau}{N} \Gamma \approx \exp \left[ -i \frac{\tau}{N} \Gamma \right].
\]  
(36)

This can be combined with Eq. (35) to write Eq. (34) as a product of matrices

\[
\left\langle T_{\tau} e^{\lambda \int_0^\tau d\tau' \sigma_f(\tau')} \right\rangle = \prod_{m=1}^N \exp \left[ -i \frac{\tau}{N} \Gamma \right] \exp \left[ \lambda \frac{\tau}{N} \sigma_z \right] \bar{p}_1,
\]

where \( \sigma_z \) is a Pauli matrix containing the two possible states of each time-slot. We can combine the two matrices and drop the commutator term of order \( 1/N^2 \). Replacing \( \bar{p}_1 \) with the steady-state probabilities and using \( N \to \infty \) limit we get the desired result

\[
G_d(\tau) \to -e^{-\omega_0 \tau} \left\{ \Gamma T \exp[\tau (\lambda \sigma_z - i\Gamma)] \bar{p}_st \right\}.
\]  
(37)

e. Archer-Anderson formula

To calculate the Green’s function from this method, one has to diagonalize the matrix \( \Gamma = \lambda \sigma_z - i\Gamma \) in the exponent and expand the \( \bar{p}_st \) in terms of the eigenstates of that matrix. For a single-rate Markovian model,

\[
\Gamma = \frac{\Gamma}{2} \begin{pmatrix}
-1 & 1 & 1 \\
1 & -1 & 1 \\
1 & 1 & -1
\end{pmatrix},
\]  
(38)
from which using \( \overline{p}_{st} = 0 \) we have
\[
\mathcal{M} = \frac{\Gamma}{2} \begin{pmatrix} 2x + i & -i \\ -i & -2x + i \end{pmatrix}, \quad \overline{p}_{st} = \frac{1}{2} \begin{pmatrix} 1 \\ 1 \end{pmatrix} \tag{39}
\]
where \( x = \lambda/\Gamma \) as before, and the (right-) eigenvalue/vectors of \( \mathcal{M} \), defined by \( \mathcal{M} \overline{u}_{\pm} = v_{\pm} \overline{u}_{\pm} \), are
\[
v_{\pm} = \frac{\Gamma}{2} \begin{pmatrix} i \mp \sqrt{4x^2 - 1} \\ \pm \sqrt{1 \mp y} \end{pmatrix}, \quad \overline{u}_{\pm} = \frac{1}{\sqrt{2}} \begin{pmatrix} i \sqrt{1 + y} \\ \pm \sqrt{1 + y} \end{pmatrix}
\]
with \( y \equiv \sqrt{1 - 1/4x^2} \). In the slow limit \( x \gg 1 \) the first term of the eigenvalue is the width and the second term is the splitting, but in the fast limit \( x \ll 1 \) they merge into one peak. Substituting in Eq. (37) we get
\[
G_d(\tau) = -e^{-\omega_0 + i\tau/2} \left[ q_+ q_+ e^{-(\tau\Gamma/2)\sqrt{4x^2 - 1}} + q_- q_- e^{i(\tau\Gamma/2)\sqrt{4x^2 - 1}} \right]. \tag{40}
\]
Here, \( q_{\pm} \) are the coefficients of the expansion \( \overline{p}_{st} = q_+ \overline{u}_+ + q_- \overline{u}_- \), given by
\[
q_{\pm} = \frac{\pm 1}{2\sqrt{2}y} \left[ \sqrt{1 \pm y} + i \sqrt{1 \pm y} \right], \tag{41}
\]
and \( q'_\pm = i\Gamma \cdot \overline{u}_\pm \) are related to them by \( q'_\pm = \pm 2yq_{\pm} \). Substituting these into Eq. (10), Fourier transforming, and doing the analytical continuation gives
\[
G_d(\omega + i\eta) = \frac{\omega + i\Gamma}{\omega(\omega + i\Gamma) - x^2\Gamma^2} = \frac{1}{\omega - \lambda^2/(\omega + i\Gamma)} \tag{42}
\]
in agreement with our self-energy result. The imaginary part gives the Archer-Anderson formula, Eq. (10). Note that we assumed that \( x > 1 \) in our proof. Interestingly, exactly the same result is obtain in the opposite regime of \( x < 1 \). While this comes as a surprise in the the present derivation, our self-energy derivation in the paper makes it clear that indeed there is no analytical difference between these two limits.

c. Markovian charge susceptibility

We can obtain this from the rate equation for the probabilities \( p_{\pm} \) of being in the states \( \sigma_f = \pm 1 \), i.e. \( \overline{\sigma}_{\pm} = \Gamma_{\pm} p_{\pm} - \Gamma_+ p_- \). Writing \( \Gamma_{\pm} = \frac{1}{2} \Gamma(1 \pm \langle \sigma_f \rangle_s) \), and combining the two equations to describe the dynamics of \( \langle \sigma_f \rangle = p_+ - p_- \) we obtain the Bloch equation
\[
\frac{d}{dt} \langle \sigma_f \rangle = -\Gamma |\langle \sigma_f \rangle - \langle \sigma_f \rangle_s| \tag{43}
\]
where \( \langle \sigma_f \rangle_s = \Gamma_+ - \Gamma_- \) is the steady-state value given by the mismatch in tunnelling rates. A small slowly-varying polarizing field \( h(t) \sigma_f \) changes the probabilities to \( p_{\pm} \approx e^{\pm \delta h(t)}/(e^{\delta h(t)} + e^{-\delta h(t)}) \approx [1 \mp \beta h(t)]/2 \) and leads to the steady-state value \( \langle \sigma_f \rangle_s = \chi_0 h(t) \) with \( \chi_0 = -\beta \). Therefore, taking the Fourier transform of Eq. (43), the susceptibility can be obtained from the ratio
\[
\chi_C(\omega + i\eta) = \frac{\langle \sigma_f \rangle}{i\omega - \frac{i\beta\Gamma}{\omega + i\Gamma}} \tag{44}
\]
so that \( \chi''_C(\omega + i\eta)/\beta\omega = -\pi\delta(\omega) \), where
\[
\delta(\omega) = \frac{\Gamma/\pi}{\omega^2 + \Gamma^2}. \tag{45}
\]
In time domain, the retarded function is \( \chi'^*_C(t) = \theta(t)\chi_0\Gamma e^{-\Gamma t} \).

d. Markovian-Gaussian formula, Eq. (11)

The retarded charge susceptibility can be obtained from the imaginary-time function \( \chi_C(\tau) = -e^{i\tau\Gamma} \). Inserting this into Eq. (6) and doing the integral we find
\[
G_d(\tau) = -e^{\alpha - \omega_0 \tau + i\alpha \Gamma} e^{-\alpha \exp[i\Gamma \tau]} \tag{46}
\]
where \( \alpha = x^2 \). We can expand the last exponent as \( e^{-\alpha \exp[i\Gamma \tau]} = \sum_n e^{i\omega_n \tau} f_n(\alpha) \) where the coefficients \( f_n(\alpha) = \sum M_n(\alpha) / (\alpha) = 0 \) and the (right) eigenfunctions \( g_n(\alpha) \) are real and they are zero for \( n < 0 \). Fourier transforming and analytical continuation gives Eq. (11).

5. Matsubara sum

We start from \( \Sigma_d(\tau) = -\lambda^2 G_d(\tau) \chi_C(\tau) \) and do a Fourier transform to write
\[
\Sigma_d(\omega_n) = -\frac{\lambda^2}{\beta} \sum_n G_d(i\omega_n) \chi_C(\omega_n - i\omega_m). \tag{47}
\]
where \( \omega_n = (2n + 1)\pi/\beta \) are fermionic Matsubara frequencies. Following the standard procedure, the summation on fermionic Matsubara frequencies is written as a contour integral around the poles of \( f(z) \) and the contour is deformed to move the integration parallel to the real frequency axis. The poles of \( G_d(z) \) and \( \chi_C(\omega_n - z) \) are along \( \text{Im} [z] = 0 \) and \( \text{Im} [z] = \omega_n \), respectively. Therefore, after analytical continuation of \( \Sigma_d(\omega_n) \) to real frequency and taking the imaginary part, we obtain
\[
\Sigma''_d(\omega) = \lambda^2 \int_{-\infty}^{+\infty} \frac{dz}{\pi} \left[ f(\omega) G_d''(\omega) \chi_C''(\omega - \omega') - n_B(-\omega') G_d''(\omega - \omega') \chi_C''(\omega') \right]. \tag{48}
\]
The part proportional to \( f(\omega) \) can be safely dropped in the \( T \ll \omega_0 \) limit. Assuming that \( G_d(\omega + i\eta) \) is the bare propagator \( g_d(\omega + i\eta) = [\omega - \omega_0 + i\eta]^{-1} \) we arrive at Eq. (13).
Eq. (48) can be used iteratively for a self-consistent calculation together with the Dyson equation $G_d(z) = [z - \omega_0 - \Sigma_d(z)]^{-1}$. To get a rough idea of the effect of this self-consistent calculation, we give the $d$ level a natural linewidth and a wavefunction renormalization factor by approximating $G_d(\omega + i\eta) \approx Z[\omega - \omega_0 + i\Delta]^{-1}$. Inserting this into Eq. (48) leads to

\[
\Sigma'_d(\omega_0 + \omega) = -\lambda^2 Z \int_{-\infty}^{+\infty} d\omega' \delta(\omega - \omega') n_B(-\omega')C_{\sigma}^{\nu}(\omega').
\]

Therefore, we see that applying the self-consistency leads to broadening of the features in the self-energy. As we showed in the main part of the paper, the self-consistency requires $\Lambda/\Gamma \sim x^2$ and therefore, the slow fluctuation case $x \gg 1$, is affected more by the self-consistent calculations than the fast fluctuation case.

### 6. Vertex correction

Taking into account the vertex correction, Eq. (47) becomes

\[
\Sigma_d(i\nu_n) = -\frac{\lambda^2}{\beta} \sum_{iv_p} \chi_C(iv_n) G_d(i\nu_n - iv_p) \Pi(i\nu_n, iv_p),
\]

where $\nu_p = 2\pi p/\beta$ are bosonic Matsubara frequencies. In time-domain this is

\[
\Sigma_d(\tau) = -\lambda^2 \int_0^\tau d\tau_1 \int_0^{\tau_1} d\tau_2 \chi_C(\tau - \tau_2) \times G_d(\tau - \tau_1) \Pi(\tau_1, \tau_1 - \tau_2),
\]

and assuming that there are no interaction vertices for the charge fluctuation in $H_{\text{rest}}$, the vertex function $\Pi$ obeys the equation

\[
\Pi(\tau_1, \tau_1 - \tau_2) = \delta(\tau_1) \delta(\tau_2)
\]

\[
-\lambda^2 \int_0^{\tau_2} d\tau'_1 \int_0^{\tau_1} d\tau'_2 \int_0^{\tau} d\tau'_3 \int_0^{\tau'_2} d\tau'_4 \left[ G_d(\tau'_1 - \tau'_3) G_d(\tau_1 - \tau'_2) \chi_C(\tau_1 - \tau'_4) \times \Pi(\tau'_3, \tau'_3 - \tau'_4) \Pi(\tau'_2 - \tau'_1, \tau'_2 - \tau_2) \right].
\]

This is shown diagrammatically in Fig. (6).

The leading correction is obtained from the first iteration [shown in Fig. (2d)]

\[
\Sigma_d(\tau) = -\lambda^2 G_d(\tau) \chi_C(\tau)
\]

\[
+ \lambda^4 \int_0^\tau d\tau_1 \int_0^{\tau_1} d\tau_2 \left[ \chi_C(\tau_1) \chi_C(\tau - \tau_2) G_d(\tau - \tau_1) \times G_d(\tau_2) G_d(\tau_1 - \tau_2) \right].
\]

Assuming that $G_d(\tau) = g_d(\tau)$ is the bare propagator and we use the forward-time approximation, the second term becomes

\[
\lambda^4 g_d(\tau) \int_0^\tau d\tau_1 \int_0^{\tau_1} d\tau_2 \chi_C(\tau_1) \chi_C(\tau - \tau_2)
\]

\[
\approx \lambda^4 g_d(\tau) \chi_C(\tau) \int_0^\tau d\tau_1 \int_0^{\tau_1} d\tau_2 \chi_C(\tau_1),
\]

where we have approximated $\chi_C(\tau - \tau_2) \approx \chi_C(\tau) + \cdots$. Assuming that $\chi_C(\tau)$ has a time-scale of $1/\Gamma$, we obtain

\[
\Sigma_d(\tau) = -\lambda^2 (1 - \lambda^2 / \Gamma^2) g_d(\tau) \chi_C(\tau),
\]

resulting in the leading vertex correction $\lambda^2 \to \lambda^2 (1 - x^2)$. Again we see that the fast fluctuation case ($x \gg 1$) is not affected by the vertex corrections. In the opposite limit of fast fluctuations, the result is negative suggesting the break down of the leading order calculation.