Emergence and decay of turbulence in stirred atomic Bose-Einstein condensates

N.G. Parker and C.S. Adams
Department of Physics, University of Durham, South Road, Durham, DH1 3LE, United Kingdom

We show that a ‘weak’ elliptical deformation of an atomic Bose-Einstein condensate rotating at close to the quadrupole instability frequency leads to turbulence with a Kolmogorov energy spectrum. The turbulent state is produced by energy transfer to condensate fragments that are ejected by the quadrupole instability. This energy transfer is driven by breaking the two-fold rotational symmetry of the condensate. Subsequently, vortex-sound interactions damp the turbulent state leading to the crystallization of a vortex lattice.

Two-dimensional (2D) turbulence has been explored in diverse areas such as soap films, magnetohydrodynamics, and meteorology (see references therein), and can often display additional features not present in 3D. However, the complexities of real fluids mean that the theoretical predictions are often at odds with observed spectra. In this regard there may be some advantages in studying superfluids where the absence of viscosity and the quantization of vorticity can simplify the theoretical picture. For example, superfluid turbulence in liquid Helium is found to exhibit analogous features to classical turbulence, in particular a Kolmogorov energy spectrum. Even more amenable to theoretical description are atomic Bose-Einstein condensates (BECs). In addition, atomic BECs allow the flexibility of studying the transition between 2D and 3D turbulence.

Recent experiments on atomic BECs have generated vortex lattices by thermodynamically condensing a rotating thermal cloud and mechanical rotation of the condensate in an anharmonic trap. In the latter case, a quadrupolar collective mode of the condensate is excited. A dynamical instability leads to the nucleation of vortices, which subsequently crystallise into a lattice configuration. The time-scale for vortex lattice formation has been shown to be insensitive to temperature, suggesting that the process is a purely dynamical effect. The formation of the lattice has been simulated using the time-dependent Gross-Pitaveskii equation (GPE) in 2D with the inclusion of damping effects and in 3D. However, questions remain over the dynamics involved, for example, how the dynamical instability seeds vortex nucleation, what is the damping mechanism leading to the crystallization of the lattice, and the role of dimensionality in the process.

In this paper, we present evidence that 2D turbulence is a key feature of current experiments on vortex lattice formation in atomic BECs. In particular we show that the route to lattice formation can be divided into four distinct stages, as illustrated in Fig. 1. These stages are:

- Fragmentation: The quadrupolar mode breaks down, ejecting energetic atoms to form an outer cloud.
- Symmetry-breaking: The two-fold rotational symmetry of the system is broken in a macroscopic manner, allowing the rotation to couple to additional modes, thereby injecting energy rapidly into the system.
- Turbulence: A turbulent cloud containing vortices and high energy density fluctuations (sound field) is formed, with a Kolmogorov energy spectrum.
- Crystallisation: The loss of energy at short length-scales coupled with the vortex-sound interactions allow the system to relax into an ordered lattice.

Our analysis is based on the classical field method whereby the GPE is used to describe both the condensate mean-field \( \psi \) and fluctuations \( \delta \psi \). For illustrative purposes we use an initial state without excitations. The rotating trap strongly polarizes the dynamics along the \( z \)-axis, such that 2D dynamics in the rotating plane dominates the system. Indeed we have verified that the same results are obtained in the flattened, quasi-2D geometry by solving the 3D GPE. We therefore proceed in the 2D limit by solving the computationally advantageous 2D GPE, as in our previous work. We have performed simulations for a range of condensate sizes, and various grid and box sizes, and find the same qualitative results throughout. We characterise the state of the system in terms of the energy, which is calculated using the time-independent GPE energy functional,

\[
E = \int \left[ -\frac{\hbar^2}{2m} |\nabla \psi|^2 + V_{ext} \psi^2 + \frac{\hbar^2}{2m} |\psi|^4 \right] dx dy, \quad (1)
\]

and subtract the corresponding energy of the initial state.
Here $V_{\text{ext}}$ is the external potential, $m$ is the atomic mass, and $g = 4\pi\hbar^2 Na/m$, where $N$ is the number of atoms and $a$ is the s-wave scattering length. To apply a rotation we use a ‘weakly’ elliptical potential similar to the experiments \cite{3} (the stirring potential used in Ref. \cite{10} is strongly anharmonic). The potential can be expressed as,

$$V_{\text{ext}}(x, y, t) = \frac{1}{2}m\omega_r^2 [(x^2 + y^2) + \varepsilon (X^2 - Y^2)] . \quad (2)$$

The first term represents the static harmonic trap with transverse frequency $\omega_r$. In the second term, the parameter $\varepsilon$ characterises the trap anisotropy in the coordinate system $(X, Y)$ which is aligned with the static $(x, y)$ coordinates at $t = 0$ and subsequently rotates at angular frequency $\Omega$. We employ a trap deformation of $\varepsilon = 0.025$ \cite{3}. Whereas previous theoretical investigations have considered the rotating frame \cite{2, 3, 15}, we perform simulations in the laboratory frame.

Our units of length, time and energy are the healing length $\xi = \hbar/\sqrt{m\mu}$, $(\xi/c)$, and the chemical potential $\mu - n_0 g$. Here $c = \sqrt{\mu/m}$ is the Bogoliubov speed of sound, $\mu = n_0 g$ is the chemical potential and $n_0$ is the peak condensate density. Using typical parameters for a $^{87}\text{Rb}$ condensate, the units of distance and time correspond to $\xi \sim 0.3 \mu\text{m}$ and $(\xi/c) \sim 10^{-4}\text{s}$, respectively.

**Fragmentation:** At $t = 0$, the rotation is turned on. The rotating trap couples energy into the condensate by exciting a quadrupolar shape oscillation, while the axes of the quadrupole rotate at the trap rotation frequency $\Omega$. This excitation mimics rotation yet the system remains irrotational. In the radially-symmetric system, the quadrupole mode is predicted to have a resonant frequency at $\Omega = \omega_r/\sqrt{2}$ \cite{21}, but this is shifted slightly by the ellipticity. Away from the resonant frequency, the quadrupole oscillations have reduced amplitude and are stable. The condensate cycles between the initial circular state and a higher energy elongated state, but over a complete cycle there is no net increase in energy. Such quadrupole oscillations have been observed experimentally \cite{22}.

The condensate is predicted to be dynamically unstable at the quadrupole resonant frequency \cite{13}. The instability arises because the amplitude of the mode becomes so large that the quadrupolar irrotational flow can no longer be supported. As the condensate relaxes from the point of maximum elongation, the fluid cannot adjust sufficiently quickly back towards the centre of trap. This results in the shedding of fluid from the ends of the condensate forming low density tails and giving the condensate a spiral shape (Fig. 2(a)(ii)). The ejected material collapses back onto the condensate edge, forming phase dislocations with the main condensate. This generates surface waves and ‘ghost’ vortices \cite{12}. An outer, low density ($\sim 0.1n_0$) cloud is formed (Fig. 2(a)(iii)). After the condensate has shed material (Fig. 2(b), solid line) its energy no longer decreases back to the initial value - energy has been transferred irreversibly into the condensate. We monitor the relative evolution of the outer, low density cloud and the inner, high density condensate by defining the outer cloud to be where the density is less than a certain value, taken here to be $0.2n_0$. The injected energy goes primarily into the outer cloud (Fig. 2(b), dashed line). Note that, although the energy of the inner and outer clouds are comparable, the outer cloud contains only around 10% of the total atoms, and therefore the average energy per atom is considerably higher. Subsequently, the inner cloud continues to undergo quadrupole oscillations and eject small fragments, as indicated by the energy curves shown in Fig. 2(b). Also, the outer cloud develops more structure at short length-scales (Fig. 2(a)(iv)).

For the parameters employed here, we observe the fragmentation of the cloud (and ultimately the formation of a vortex lattice) for rotation frequencies in the range $0.72 < \Omega/\omega_r < 0.78$, which is in reasonable agreement with the experimental results of Madison et al. \cite{9}. Outside this range the quadrupole mode is stable, and the width of the unstable region increases with the trap ellipticity $\varepsilon$, which is consistent with the experimental observations of Hodby et al. \cite{11}.

**Symmetry breaking:** The condensate and potential have a two-fold rotational symmetry, which is clearly visible in Fig. 2(a). In the experiments \cite{3} \cite{10} this symmetry is absent. Eventually in the simulations, an asymmetry grows out of the numerical noise generated by modelling a rotation using a static square grid. We characterise this asymmetry in terms of an asymmetry parameter defined as,

$$\sigma = \int |\psi(x, y)|^2 - |\psi(-x, -y)|^2 |dx dy| \int |\psi(x, y)|^2 |dx dy|. \quad (3)$$

This asymmetry parameter (Fig. 3(b), solid line) grows...
exponentially over time. When it reaches macroscopic levels ($\sigma \sim 0.1$) we enter a symmetry-breaking phase at $t \sim 8000 \,(\xi/c)$. From then on, additional modes can be excited, and energy is rapidly coupled into the system (Fig. 3(a), solid line). This energy predominantly excites the outer, low density cloud, as shown in Fig. 1 (dashed line). The condensate density and phase during this stage are shown in Fig. 3(c)-(d). The original two-fold rotational symmetry is now clearly broken. Towards the end of this stage the outer cloud strongly couples to, and merges with, the inner condensate. Energy and angular momentum become transferred to the inner cloud, the quadrupole mode finally breaks down, and vortices become nucleated in the higher density regions.

In experiments, symmetry-breaking will occur due to, for example, the thermal cloud, quantum fluctuations, trap imperfections, and fluctuating fields. We model the effect by allowing the trap centre to randomly jump, or jitter, within a region $[-\gamma, +\gamma] \times [-\gamma, +\gamma]$. Fig. 3(a)-(b) show the results for trap jitters of $\gamma = 0.0001$ (dashed line) and $\gamma = 0.1$ (dotted). (c)-(d) Snapshots of the total condensate density and phase, respectively, during the symmetry breaking stage at $t = 9600 \,(\xi/c)$ in the absence of trap jitter.

**Turbulence:** Following the injection of energy into the condensate during the symmetry-breaking stage, a highly excited and energetic condensate containing randomly-positioned vortices is formed, as shown in Fig. 4(a)-(b). We analyse this stage of the evolution by calculating the energy spectrum, shown in Fig. 4(c). During the turbulent phase the system closely follows a Kolmogorov energy spectrum $E(k) \propto k^{-5/3}$ over a range of $k$-values, as shown in Fig 4(c) (bold line) for $t = 11000 \,(\xi/c)$. Such behaviour is a key signature of classical turbulence and also occurs in models of superfluid turbulence [20]. The departure from a $k^{-5/3}$ law occurs at an upper bound of $k \approx 2\,(2\pi/\xi)$, corresponding to the characteristic length-scale of vortex-sound interactions [20], and a lower bound of $k \approx 0.3\,(2\pi/\xi)$, corresponding to the size of the condensate. An additional feature of 2D classical turbulence is that the spectrum is predicted to show $k^{-3}$ behaviour following the $k^{-5/3}$ region [4]. However, we observe a $k^{-6}$ dependence in this range. The power spectrum in the fragmentation stage (Fig. 4(a), intermediate grey) and the succeeding crystallisation stage (Fig. 4(a), light grey) do not show a $k^{-5/3}$ behaviour.

**Crystallisation:** The turbulent state contains a dense sound field and vortices. In previous work we have shown how vortices can both radiate and absorb sound waves in a Bose-Einstein condensate [18, 19]. We propose that this vortex-sound interaction enables the randomly-positioned vortices in our rotating condensate to crystallise into a lattice.

To demonstrate the transfer of energy from the sound field (density fluctuations) to the vortices, we divide the energy into a component due to the vortices $E_V$ and a component due to the sound field $E_s$. We approximate $E_V$ at a particular time by numerically generating a similar condensate (by propagating the GPE in imaginary time) with the same vortex distribution but without sound. The sound energy is then defined as $E_s = E - E_V$. During the fragmentation, symmetry breaking and turbulent stages, both sound energy (Fig. 5(a), dotted line) and vortex energy (Fig. 5(a), dashed line) are fed into the system. At the start of the crystallisation stage the sound energy is considerably higher than the vortex energy. However, as time progresses the sound energy decreases with this energy being transferred to the vortices. Figs. 5(b)-(c) shows the condensate density at the begin-
In our simulations, the finite grid size sets a maximum number of vortices which ultimately form in the lattice.

The turbulent state subsequently evolves into a vortex lattice by vortex-sound interactions. Future work will focus on turbulence in spinor systems and the effect of dimensionality on the turbulence spectrum.

We acknowledge UK EPSRC for funding and S. A. Gardiner for discussions.

References:

[1] Y. Couder, J. M. Chomaz, and M. Rabaud, Physica 37D, 384 (1989); M. Gharib and P. Derango, Physica 37D, 406 (1989).

[2] O. Cardoso, D. Manteau, and P. Tabeling, Phys. Rev. E 49, 454 (1994).

[3] K. S. Smith, J. Atmos. Sci. 61, 937 (2004).

[4] R. H. Kraichnan, Phys. Fluids 10, 1417 (1967); C. E. Leith, Phys. Fluids 11, 671 (1968); G. K. Batchelor, Phys. Fluids (Suppl.) II 12, 233 (1969).

[5] C. F. Barenghi, R. J. Donnelly and W. F. Vinen (Eds.), Quantized Vortex Dynamics and Superfluid Turbulence (Springer, Berlin, 2001).

[6] C. Nore, M. Abid, and M. E. Brachet, Phys. Rev. Lett. 78, 3896 (1997); T. Araki, M. Tsubota, and S. K. Nemirovskii, Phys. Rev. Lett. 89, 145301 (2002); M. Kobayashi and M. Tsubota, Phys. Rev. Lett. 94, 065302 (2005).

[7] F. Dalfovo, S. Giorgini, L. P. Pitaevskii, and S. Stringari, Rev. Mod. Phys. 71, 463 (1999).

[8] P. C. Haljan, I. Coddington, P. Engels, and E. A. Cornell, Phys. Rev. Lett. 87, 210403 (2001).

[9] K. W. Madison, F. Chevy, W. Wohlleben, and J. Dalibard, Phys. Rev. Lett. 84, 806 (2000).

[10] J. R. Abo-Shaeer, C. Raman, J. M. Vogels, and W. Ketterle, Science 292, 476 (2001).

[11] E. Hodby et al., Phys. Rev. Lett. 88, 010405 (2001)

[12] K. W. Madison, F. Chevy, V. Bretin, and J. Dalibard, Phys. Rev. Lett. 86, 4443 (2001).

[13] S. Sinha and Y. Castin, Phys. Rev. Lett. 87, 190402 (2001).

[14] J. R. Abo-Shaeer, C. Raman, and W. Ketterle, Phys. Rev. Lett. 88, 070409 (2002).

[15] M. Tsubota, K. Kasamatsu, and M. Ueda, Phys. Rev. A 65, 023603 (2002); K. Kasamatsu, M. Tsubota, and M. Ueda, Phys. Rev. A 67, 033610 (2003).

[16] A. A. Penckwitt, R. J. Ballagh, and C. W. Gardiner, Phys. Rev. Lett. 89, 260402 (2002).

[17] C. Lobo, A. Sinatra, and Y. Castin, Phys. Rev. Lett. 92, 020403 (2004).

[18] N. G. Parker, N. P. Proukakis, C. F. Barenghi, and C. S. Adams, Phys. Rev. Lett. 92, 160403 (2004).

[19] M. Leadbeater et al., Phys. Rev. Lett. 86, 1410 (2001).

[20] M. J. Davis, S. A. Morgan, and K. Burnett, Phys. Rev. Lett. 87, 160402 (2001); K. Goral, M. Gajda, and K. Razzewski, Opt. Express 8, 92 (2001); A. A. Norrie, R. J. Ballagh, and C. W. Gardiner Phys. Rev. Lett. 94, 040401 (2005).

[21] S. Stringari, Phys. Rev. Lett. 77, 2360 (1996).

[22] F. Chevy, K. W. Madison, and J. Dalibard, Phys. Rev. Lett. 85, 2223 (2000); G. Hechenblakner et al., Phys. Rev. Lett. 88, 070406 (2002).