Abstract

The Edelweiss collaboration performs light Dark Matter (DM) particles searches with germanium bolometer collecting charge and phonon signals. Thanks to the Neganov-Trofimov-Luke (NTL) effect, a RMS resolution of 4.46 electron-hole pairs was obtained on a massive (200g) germanium detector instrumented with a NbSi Transition Edge Sensor (TES) operated underground at the Laboratoire Souterrain de Modane (LSM). This sensitivity made possible a search for WIMP using the Migdal effect down to 32 MeV/C² and exclude cross-sections down to $10^{-29}$ cm². It is the first measurement in cryogenic germanium with such thermal sensor, proving the high relevance of this technology. Furthermore, such TES have shown sensitivity to out-of-equilibrium phonons, paving the way for EDELWEISS new experience CRYOSEL. This is an important step in the development of Ge detectors with improved performance in the context of the EDELWEISS-SubGeV program.
1 Introduction

There were great progresses in the Dark Matter (DM) direct interaction with nuclei searches \([1–3]\), for particle with masses ranging from 1 GeV\(\cdot\)c\(^{-2}\) to 1 TeV\(\cdot\)c\(^{-2}\) \([4–6]\). Nevertheless, because of the absence of signal in that region, the interest in extending the search to lower masses intensified \([7–13]\). Probing lower masses raise additional experimental constraints, such as the need of lower energy threshold (below 1 keV) as well as ionization or scintillation yield for nuclear recoil. To such constraints, collaborations have to add dealing with new types of low energy backgrounds \([14–19]\).

In order to avoid the issues raised by the very low kinetic energy of the nuclear recoil, it is possible to use the Migdal effect \([20,21,26,27]\). This effect quantifies the probability that an atomic electron is released because of the DM particle scattering on a nucleus. The electrons are emitted with an energy typically much larger than the kinematic energy of the nuclear recoil \([27]\), yielding signals much easier to detect.

In this context, the EDELWEISS collaboration used the Migdal effect to extend the mass range of DM particles search down to 32 MeV\(\cdot\)c\(^{-2}\), using a 200 g cryogenic Ge detector equipped with a NbSi Transition Edge Sensor (TES) and operated underground at the Laboratoire Souterrain de Modane (LSM). This detector achieved a low energy threshold for electron recoil and appeared to be less sensitive to the Heat Only background (HO), an already known background made of events that are not associated with charges.

The development of the NbSi detector is part of the EDELWEISS-SubGeV program and a step toward the new generation of EDELWEISS detectors, Cryosel. Those detectors will be Ge bolometers able to sustain high biases and equipped with a new sensor able to tag down to a single charge and thus veto HO events \([28]\).

2 New results: DM search with a detector equipped with TES

The search was performed at the LSM benefitting from the ultra-low background environment of the EDELWEISS-III cryostat \([29]\) and the 4800 m.w.e. rock overburden. The experiment, its results and their interpretation are described in detail in \([30]\). The detector named NbSi209 is a 200 g Ge cylindrical crystal (48 mm in diameter and 20 mm in height) on top of which was lithographed a Nb\(_x\)Si\(_1-x\) thin film TES \([28]\) used as thermal sensor. The TES is operated around 44 mK, the temperature at which the transition between the normal and the superconducting states start to occur. The data from the ionization and heat channels were digitized at a frequency of 100 kHz, filtered, and continuously stored on disk with a digitization rate of 500 Hz. The bottom side is fully covered by an electrode biased at a voltage varying between \(\pm 66\) V. The bias is applied in order to trigger the so-called Neganov-Trofimov-Luke (NTL) effect \([31,32]\). This effect states that the drift of the \(N\) electron-hole pairs across a voltage difference \(\Delta V\) creates additional phonons of energy \(E_{NTL} = Ne\Delta V\) (\(e\) is the elementary charge). This energy \(E_{NTL}\) adds to the recoil energy. In the case of events associated with charges, the mean gain of the NTL effect is \(<g> = 1 + e\Delta V/\epsilon_{gamma}\), where \(\epsilon_{gamma} = 3.0\) eV is the average ionization energy in Ge for electron recoils \([33]\). The average resolution over the dataset is between 90 and 100 eV. For the NTL gain of \(<g> = 23\) corresponding to a bias of 66 V, this leads to a resolution of approximately 4 eV electron-equivalent (eV\(_{ee}\)).

The detector was activated with a strong AmBe neutron source. This produces short-lived \(^{71}\)Ge isotopes which decay by electron capture in the K, L, and M shells, with de-excitation x-ray lines at 10.37, 1.30, and 0.16 keV, respectively. The isotopes being produce uniformly throughout the detector volume, the absorption of the low energy x-ray lines allows to probe the DM response in the entire detector volume, providing a precise characterization of
Figure 1: Left: energy spectrum recorded at a bias of 66 V following the $^{71}\text{Ge}$ emitted x-ray lines induced by the neutron activation of the Ge detector, resulting in the strong lines characteristic of $^{71}\text{Ge}$ x-ray emission. Right: efficiency as a function of the input phonon heat energy of simulated events inserted in the data stream in units of keV (lower axis), or keV-electron-equivalent (keV$_{ee}$, upper axis). Each color line corresponds to a step in the trigger and data analysis process. Those figures have been taken from [30] under RNP/22/OCT/058564 license.

The data have been interpreted in terms of limits on the spin-independent interaction of DM particles with target atoms through the so-called Migdal effect using calculations from [27, 34]. In addition, because of the underground location of the detector, the action of the stopping power of the rock overburden on the DM particle flux has been taken into account [35–38]. The search was performed with data recorded between March 2019 and June 2020, only considering samples with a baseline heat energy resolution lower than 140 eV RMS. The resulting average resolution is 102 ± 12 eV RMS, corresponding to 4.46 ± 0.54 eV$_{ee}$ RMS once the NTL gain $\langle g \rangle$ is considered. Half of the dataset was blinded to perform the search, the other half was used to set the analysis criteria. The number of events in the chosen energy range (the regions of interest ROI’s), is used to set the 90% C.L. Poisson upper limit on DM particle interacting through Migdal effect with Ge nuclei. The signal energy distributions for a WIMP of 50 MeV, corrected for efficiency and smeared to detector resolution for upper and lower excluded cross-sections are shown in Fig 2 left panel in green and red, respectively, as
Figure 2: Left: energy spectrum after all criteria are applied for the blinded dataset in counts per keV (black histogram). Reference sample data smoothed by analytical function (plain blue). The red and green curves show the excluded Migdal spectra smeared to detector resolution, corrected for the Earth-shielding effect and efficiency corrected for WIMPs of 50 MeV·c² for lower and upper excluded cross-section. Right: 90% C.L. upper limit on the cross-section for Spin-Independent interaction between DM and Ge nuclei through the Migdal effect. The red contour is obtained by taking into account the slowing of the DM particle flux through the material above the detector. These results are compared to other experiments [19, 22, 25, 39–41].

Those figures have been taken from [30] under RNP/22/OCT/058564 license.

well as their corresponding ROI’s. The non-blinded sample was used to derive a data model in order to set the ROI’s prior to unblind the rest of the data. This model corresponds to the blue curve in Fig 2 left panel.

The resulting limits are shown by the red contour (delimited by the thick red line) in Fig 2, right panel. Those DM constraints go down to 32 MeV·c². Below that mass, the large cross-section that would yield an observable signal is prevented by the stopping effect of the overburden and shielding. This contour is compared with results from other experiments [19, 22, 25, 39–41] and with the previous EDELWEISS Migdal search performed above ground. This shows orders of magnitude of improvement with this previous search and that the underground operation of the detector did not jeopardize the potential of a Migdal search.

Nevertheless, despite the efforts made to reject the HO contamination, this analysis was background limited. Although its HO nature is not questionable at high energy because of the ionization signal associated with the heat signal, the ionization resolution of ~200 eV (RMS) prevent such conclusion at lower energy. Fig 3 shows energy distribution of data recorded at 15 V and 66 V in red and black, respectively. The lack of quantitative shape differences of those spectrums is a hint of their HO nature. The fit of a power law (αEβ) yields identical slopes within uncertainties β ~ 3.40 for both spectra. The compatible fitted powers indicate that no NTL amplification occurs, and therefore, that no charges are involved. The flatness of the ratio shown on the bottom pad enhanced this observation. Furthermore, assuming the that both the HO background and the possible electronic background follow the same power law, a limit on the fraction of events associated with charges x can be set, the resulting upper limit is x < 0.0004 at 90% C.L. This confirms that the nature of this background is mainly HO in the energy range of the search between 0.8 to 2.8 keV.
Figure 3: Heat energy spectra of events recorded with RED20 operated above ground at 0 V (green) \cite{24}, events with no ionization ($E_{\text{ion}} < 0$) for the RED30 detector operated at 15 V (blue), NbSi209 operated at 15 V (red) and 66 V (black). The fitted power law on NbSi spectra when operated at 66 V and 15 V in black and red respectively. The lower figure shows the ratio of NbSi209 distributions recorded at 15 and 66 V (blue) and the associated fit of a constant (black line) and its uncertainty band (orange). This figure has been taken from \cite{30} under RNP/22/OCT/058564 license.

3 Prospect: Cryosel

The next step in the EDELWEISS Sub-GeV program is to improve both the phonon resolution of the detectors and the maximum bias at which they can be operated. Developments are ongoing on the study and reduction of the HO background, as it is one of the limiting factor in future searches in EDELWEISS. For example, a new detector with a novel phonon sensor, consisting of a 10 $\mu$m thin NbSi line called Super Conducting Single Electron Device (SSED) evaporated on top of a 40 g Ge crystal is under study. Such sensor on a detector operated at high biases ($\sim 200$ V) has a sensitivity to the athermal phonons triggered by the NTL effect of a single charge drifting in the strong electric field. This leads to the possibility of evaluating the purely thermal nature of HO events and thus vetoing this background. The goal of CRYOSEL is to successfully couple this SSED with a high performance bolometer with 20 eV (RMS) of baseline resolution. The high-resolution thermal measurement is obtained by using a NTD-Ge thermistor with 20 eV phonon resolution and applying a bias of 200 V on the detector. This design is shown in the left panel of Fig 4 along with the structure of the expected electric field. The right panel shows the expected sensitivity of such design to WIMP nuclei interaction, this would be competitive with the performances of much more massive detectors and cover areas of the parameters space that have yet to be excluded. It would also cover all the milestone models \cite{42} for DM interacting with electron. This program is in it R&D stage and prototypes are currently being produced and tested.
Figure 4: Left: Cryosel 30 g Ge cryogenic detector with covering bottom Al electrode and the SSED detector layout including the simulated electric field-lines. The intensity of the electric field rises significantly under the SSED line (red region) Right: thick red line: projected sensitivity for WIMP-nucleon spin-independent interaction (90%CL) with a 1 kg.d exposure of the CRYOSEL detector (~ 1 month with a 30 g detector). These results are compared to other experiments [5,16,19,24,39–41]

4 Conclusion

The EDELWEISS collaboration has searched for DM particle interaction exploiting the Migdal effect with WIMP masses between 32 MeV·c^{−2} and 2 GeV·c^{−2} using a 200 g Ge detector operated underground at the LSM. The search constrains a new region of the parameters space for cross-sections close to 10^{−29} cm^{2}. In the context of its EDELWEISS-SubGeV program, the collaboration is also investigating new methods to significantly reduce HO backgrounds by improving its ionization resolution with the use of new cold preamplifiers [43], and by developing NbSi-instrumented devices able to tag the out-of-equilibrium NTL phonons associated to a single electron.

Acknowledgments

Funding information The EDELWEISS project is supported in part by the French Agence Nationale pour la Recherche (ANR) and the LabEx Lyon Institute of Origins (ANR-10-LABX-0066) of the Université de Lyon within the program “Investissements d’Avenir” (ANR-11-IDEX-00007), by the P2IO LabEx (ANR-10-LABX-0038) in the framework “Investissements d’Avenir” (ANR-11-IDEX-0003-01) managed by the ANR (France), and the Russian Foundation for Basic Research (grant No. 18-02-00159). This project has received funding from the European Union’s Horizon 2020 research and innovation program under the Marie Skłodowska-Curie Grant Agreement No. 838537. B.J. Kavanagh thanks the Spanish Agencia Estatal de Investigación (AEI, MCIU) for the support to the Unidad de Excelencia María de Maeztu Instituto de Física de Cantabria, ref. MDM-2017-0765. We thank J.P. Lopez (IP2I) and the Physics Department of Université Lyon 1 for their contribution to the radioactive sources.
References

[1] M. W. Goodman and E. Witten, Detectability of certain dark-matter candidates, Phys. Rev. D 31, 3059 (1985), doi:10.1103/PhysRevD.31.3059.

[2] A. K. Drukier, K. Freese and D. N. Spergel, Detecting cold dark-matter candidates, Phys. Rev. D 33, 3495 (1986), doi:10.1103/PhysRevD.33.3495.

[3] A. Drukier and L. Stodolsky, Principles and applications of a neutral-current detector for neutrino physics and astronomy, Phys. Rev. D 30, 2295 (1984), doi:10.1103/PhysRevD.30.2295.

[4] E. Aprile et al., Dark matter search results from a one ton-year exposure of XENON1T, Phys. Rev. Lett. 121, 111302 (2018), doi:10.1103/PhysRevLett.121.111302.

[5] D. S. Akerib et al., Results from a search for dark matter in the complete LUX exposure, Phys. Rev. Lett. 118, 021303 (2017), doi:10.1103/PhysRevLett.118.021303.

[6] A. Tan et al., Dark matter results from first 98.7 days of data from the PandaX-II experiment, Phys. Rev. Lett. 117, 121303 (2016), doi:10.1103/PhysRevLett.117.121303.

[7] R. Essig, J. Kaplan, P. Schuster and N. Toro, On the origin of light dark matter species, arXiv:1004.0691.

[8] C. Cheung, J. T. Ruderman, L.-T. Wang and I. Yavin, Kinetic mixing as the origin of a light dark-gauge-group scale, Phys. Rev. D 80, 035008 (2009), doi:10.1103/PhysRevD.80.035008.

[9] D. Hooper and W. Xue, Possibility of testing the light dark matter hypothesis with the alpha magnetic spectrometer, Phys. Rev. Lett. 110, 041302 (2013), doi:10.1103/PhysRevLett.110.041302.

[10] A. Falkowski, J. T. Ruderman and T. Volansky, Asymmetric dark matter from leptogenesis, J. High Energy Phys. 05, 106 (2011), doi:10.1007/JHEP05(2011)106.

[11] K. Petraki and R. R. Volkas, Review of asymmetric dark matter, Int. J. Mod. Phys. A 28, 1330028 (2013), doi:10.1142/S0217751X13300287.

[12] K. M. Zurek, Asymmetric dark matter: Theories, signatures, and constraints, Phys. Rep. 537, 91 (2014), doi:10.1016/j.physrep.2013.12.001.

[13] G. Bertone and T. M. P. Tait, A new era in the search for dark matter, Nature 562, 51 (2018), doi:10.1038/s41586-018-0542-z.

[14] Q. Arnaud et al., First germanium-based constraints on sub-MeV dark matter with the EDELWEISS experiment, Phys. Rev. Lett. 125, 141301 (2020), doi:10.1103/physrevlett.125.141301.

[15] G. Angloher et al., Results on light dark matter particles with a low-threshold CRESST-II detector, Eur. Phys. J. C 76, 25 (2016), doi:10.1140/epjc/s10052-016-3877-3.

[16] R. Agnese et al., Low-mass dark matter search with CDMSlite, Phys. Rev. D 97, 022002 (2018), doi:10.1103/PhysRevD.97.022002.

[17] A. Aguilar-Arevalo et al., First direct-detection constraints on eV-scale hidden-photon dark matter with DAMIC at SNOLAB, Phys. Rev. Lett. 118, 141803 (2017), doi:10.1103/PhysRevLett.118.141803.
[18] O. Abramoff, SENSEI: Direct-detection constraints on sub-GeV dark matter from a shallow underground run using a prototype skipper CCD, Phys. Rev. Lett. 122, 161801 (2019), doi:10.1103/physrevlett.122.161801.

[19] Z. Z. Liu et al., Constraints on spin-independent nucleus scattering with sub-GeV weakly interacting massive particle dark matter from the CDEX-1B experiment at the China Jinping underground laboratory, Phys. Rev. Lett. 123, 161301 (2019), doi:10.1103/physrevlett.123.161301.

[20] J. D. Vergados and H. Ejiri, The role of ionization electrons in direct neutralino detection, Phys. Lett. B 606, 313 (2005), doi:10.1016/j.physletb.2004.11.085.

[21] C. C. Moustakidis, J. D. Vergados and H. Ejiri, Direct dark matter search by observing electrons produced in neutralino-nucleus collisions, Nucl. Phys. B 727, 406 (2005), doi:10.1016/j.nuclphysb.2005.08.033.

[22] G. Angloher et al., Results on MeV-scale dark matter from a gram-scale cryogenic calorimeter operated above ground, Eur. Phys. J. C 77, 637 (2017), doi:10.1140/epjc/s10052-017-5223-9.

[23] D. W. Amaral et al., Constraints on low-mass, relic dark matter candidates from a surface-operated SuperCDMS single-charge sensitive detector, Phys. Rev. D 102, 091101 (2020), doi:10.1103/PhysRevD.102.091101.

[24] E. Armengaud, Searching for low-mass dark matter particles with a massive Ge bolometer operated above ground, Phys. Rev. D 99, 082003 (2019), doi:10.1103/physrevd.99.082003.

[25] Z. Z. Liu et al. (CDEX Collaboration), Studies of the earth shielding effect to direct dark matter searches at the China Jinping Underground Laboratory, Phys. Rev. D 105, 052005 (2022), doi:10.1103/PhysRevD.105.052005.

[26] R. Bernabei et al., On electromagnetic contributions in wimp quests, Int. J. Mod. Phys. A 22, 3155 (2007), doi:10.1142/S0217751X07037093.

[27] M. Ibe, W. Nakano, Y. Shoji and K. Suzuki, Migdal effect in dark matter direct detection experiments, J. High Energy Phys. 03, 194 (2018), doi:10.1007/JHEP03(2018)194.

[28] S. Marnieros et al., High Impedance TES Bolometers for EDELWEISS, J Low Temp Phys 211, 214 (2022), doi:10.1007/s10909-022-02899-2.

[29] E. Armengaud et al., Performance of the EDELWEISS-III experiment for direct dark matter searches, J. Instrum. 12, P08010 (2017), doi:10.1088/1748-0221/12/08/P08010.

[30] E. Armengaud et al., Search for sub-GeV dark matter via the Migdal effect with an EDELWEISS germanium detector with NbSi transition-edge sensors, Phys. Rev. D 106, 062004 (2022), doi:10.1103/PhysRevD.106.062004.

[31] B. Neganov and V. Trofimov, Colorimetric method measuring ionizing radiation, Otkryt. Izobret. 146, 215 (1985).

[32] P.N. Luke, Voltage-assisted calorimetric ionization detector, J. Appl. Phys. 64, 6858 (1988).

[33] G. F. Knoll, Radiation detection and measurement, Wiley & Sons, New York, USA, ISBN 9780470131480 (2010).
[34] R. Essig, J. Pradler, M. Sholapurkar and T.-T. Yu, Relation between the Migdal effect and dark matter-electron scattering in isolated atoms and semiconductors, Phys. Rev. Lett. 124, 021801 (2020), doi:10.1103/physrevlett.124.021801.

[35] B. J. Kavanagh, R. Catena and C. Kouvaris, Signatures of Earth-scattering in the direct detection of dark matter, J. Cosmol. Astropart. Phys. 012 (2017), doi:10.1088/1475-7516/2017/01/012.

[36] T. Emken and C. Kouvaris, DaMaSCUS: The impact of underground scatterings on direct detection of light dark matter, J. Cosmol. Astropart. Phys. 031 (2017), doi:10.1088/1475-7516/2017/10/031.

[37] M. Shafi Mahdawi and G. R. Farrar, Energy loss during dark matter propagation in an overburden, (arXiv preprint) doi:10.48550/arXiv.1712.01170.

[38] D. Hooper and S. D. McDermott, Robust constraints and novel gamma-ray signatures of dark matter that interacts strongly with nucleons, Phys. Rev. D 97, 115006 (2018), doi:10.1103/PhysRevD.97.115006.

[39] D. S. Akerib et al., Results of a search for sub-GeV dark matter using 2013 LUX data, Phys. Rev. Lett. 122, 131301 (2019), doi:10.1103/physrevlett.122.131301.

[40] E. Aprile et al., Search for light dark matter interactions enhanced by the Migdal effect or Bremsstrahlung in XENON1T, Phys. Rev. Lett. 123, 241803 (2019), doi:10.1103/PhysRevLett.123.241803.

[41] A Abdelhameed et al., First results from the CRESST-III low-mass dark matter program, Phys. Rev. D 100, 102002 (2019), doi:10.1103/PhysRevD.100.102002.

[42] J. Alexander et al., Dark sectors 2016 workshop: Community report, (arXiv preprint) doi:10.48550/arXiv.1608.08632.

[43] A. Juillard et al., Low-noise HEMTs for coherent elastic neutrino scattering and low-mass dark matter cryogenic semiconductor detectors, J. Low Temp. Phys. 199, 798 (2019), doi:10.1007/s10909-019-02269-5.