Collisionless shocks are associated with numerous natural phenomena from the bow-shock ahead of the earth’s magnetosphere\(^1\) to the blast wave in supernova explosions\(^2\). They can accelerate particles to high energies, and are implicated in the generation of cosmic rays\(^3\). It has been suggested that collisionless shocks can be similarly used to create energetic particle beams in the laboratory. In particular, the radiation pressure of powerful lasers has been proposed as a shock-initiating piston\(^4\). A number of measurements of energetic ion beams have been attributed to acceleration by the space charge\(^5\) or critical density,\(^6\) and leads to a shock being driven into the target. Hole-boring has been experimentally identified from the optical\(^7\) probing of the plasma. As the shock speed decreases, the ion mean free path decreases and the shocked shock eventually becomes collisional as it slows further. These conclusions are supported by PIC simulations which show that the initial evolution is dominated by collisionless mechanisms.

Collaboration:\nN. P. Dover\(^1\), C. A. J. Palmer\(^1\), M. Babzien\(^2\), A. R. Bell\(^3\), A. E. Dangor\(^1\), T. Horbury\(^1\), M. Ispiriyan\(^4\), M. N. Polyanskiy\(^2\), J. Schreiber\(^1\), S. Schwartz\(^1\), P. Shkolnikov\(^4\), V. Yakimenko\(^2\), I. Pogorelsky\(^2\), and Z. Najmudin\(^1\)
\(^1\) Blackett Laboratory, Imperial College, London SW7 2AZ, United Kingdom
\(^2\) Accelerator Test Facility, Brookhaven National Laboratory, NY 11973, USA
\(^3\) Clarendon Laboratory, University of Oxford, Oxford OX1 3PU, United Kingdom
\(^4\) Stony Brook University, Stony Brook, NY 11794, USA

(Dated: May 1, 2014)

Observations of the interaction of an intense \(\lambda_0 \approx 10 \mu m\) laser pulse with near-critical overdense plasmas \((n_c = 1.8 - 3n_e)\) are presented. For the first time, transverse optical probing is used to show a recession of the front surface caused by radiation pressure driven hole-boring by the laser pulse with an initial velocity \(> 10^6 \text{ms}^{-1}\), and the resulting collisionless shocks. The collisionless shock propagates through the plasma, dissipates into an ion-acoustic solitary wave, and eventually becomes collisional as it slows further. These conclusions are supported by PIC simulations which show that the initial evolution is dominated by collisionless mechanisms.

Proton probing offer the ability to study to higher density but cannot directly measure density inside solid targets, where scattering is strong.

This letter details a study of the shock driven into a just-overdense plasma by laser hole-boring. The plasma is driven by an infrared (CO\(_2\)) laser, for which \(n_c \approx 10^{19} \text{cm}^{-3}\); > 100 times lower than that for optical wavelengths. This allows direct optical probing of the plasma. By varying the timing of the probe relative to the drive pulse, the shock evolution was studied, and three separate phases were identified. Initially the laser exerts a pressure on the opaque plasma, causing hole-boring and formation of either an electrostatic double layer at the surface, or a collisionless shock moving into the plasma, accelerating background ions. Once the laser stops, particle reflection from the propagating shock structure causes it to decelerate rapidly, transitioning into a solitary ion-acoustic wave which continues to slow in the cooling plasma. As the shock speed decreases, the ion mean free path also decreases and the shock eventually becomes collisional. These findings are supported by particle-in-cell (PIC) simulations, which also help elucidate the mechanism of narrow energy spread ion beam generation.

The experiment was performed with the \(\lambda_0 \approx 10 \mu m\) CO\(_2\) laser at the Accelerator Test Facility at Brookhaven National Laboratory. The laser was circularly polarised. Streak images of the laser time profile showed a train of pulses, caused by spectral modulation in the amplifiers, each of pulse-length \(\tau \approx 6 \text{ps (FWHM)}\) and 18 ps pulse separation. On average 70% of the energy was in the leading two pulses, with a total integrated energy of 2-3 J. The pulse was focused with an \(f/3\) off-axis parabolic mirror to a spot size \(w_0 \approx 70 \mu m\) (FWHM), giving a peak intensity \(I \approx 5 \times 10^{15} \text{Wcm}^{-2}\) and normalised vector potential \(a_0 \approx 0.5\). The laser was focused \(\approx 0.7 \text{mm}\) above a 2 mm diameter nozzle gas jet producing a helium gas neutral density profile, characterised by transverse interferometer\(^8\), approximated by an isosceles trapezoid of plateau width 1200 \(\mu m\) and side length 600 \(\mu m\).
The probe, a frequency-doubled (532 nm) Nd:YAG laser beam of $\tau \approx 6$ ps, was absolutely synchronised with the CO$_2$ pulse. The timing between them was varied using an optical delay line. The probe passed orthogonally through the plasma before being split; with one part imaged directly for shadowgraphy, and the other going to a Mach-Zender interferometer, to diagnose plasma density. Examples are shown in Fig. 1 for a peak initial density of $n_e = 3n_c$, at $t = 30$ ps after the start of the interaction, which is just after the second major pulse in the pulse train has passed. Bright features in the shadowgram (Fig. 1b) highlight regions of rapidly changing refractive index and therefore density. The shadowgram shows a feature at the target front surface of $\sim 100$ $\mu$m transverse size narrowing to $\sim 50$ $\mu$m. Interferometry reveals that this is a high density feature with a region of considerably lower density inside it (fig. 1b), which is the channel formed by the initial 75 $\mu$m laser focal spot boring into the plasma. The feature ends $\sim 45$ $\mu$m longitudinally from the estimated initial critical surface, calculated from the gas density profile and supported by the off-axis plasma density profile. Assuming the front originated at the critical surface implies an average $v = 1.5 \times 10^6$ ms$^{-1}$ ($\pm 30\%$), which agrees reasonably with $v_{hb} = 1.9 \times 10^6$ ms$^{-1}$, obtained using the laser parameters for this shot and taking $\rho = m_{He}n_c$. This calculation ignores self focusing, the pulse train, and thermal effects.

Fig. 2 shows a series of shadowgrams taken with varying probe timings up to $t = 1600$ ps. The front seen in Fig. 2a remains clearly defined, moving into the plasma as time progresses. Horizontal features can be seen further into the plasma, particularly for $t = 200$ ps (fig. 2a). These are likely to have been created as the target was being ionised, since the gas target is initially transparent to the ionising laser. Below the shadowgraphy images are the corresponding density maps derived from interferometry (Fig. 2b-f). The phase shift was assumed to be symmetrical around the laser axis, allowing Abel inversion to retrieve the density profile. The accuracy of the recovered density maps is limited by asymmetry in the interferograms, caused by, for example, slight vertical density gradients and laser non-uniformities.

Once sufficiently ionised, the plasma becomes opaque ($n_e > n_c$) and the laser is reflected at the critical surface. It then exerts radiation pressure on the plasma, launching a shock into the initially relatively cold plasma. The peak density associated with this moving front exceeds the initial background plasma density (fig. 2b,c), and shows a steepening in the feature profile even at late times; clear indication of shock formation. Two probe timing scans were taken for peak densities $n_e \in (1.8, 3)n_c$. The distance of shock propagation from the initial critical surface position thus obtained is plotted in Fig. 3 as a function of time. The shock initially moves rapidly into the plasma before slowing at later times, with the shock velocity initially higher for lower density. The mean velocities for different parts of the shock trajectory are indicated on the plot. The previously estimated hole-boring velocity ($v_{hb} \approx 1.5 \times 10^6$ ms$^{-1}$) is indicated, slowing to $v_{sh} \approx 5 \times 10^4$ ms$^{-1}$ for $t = 500$-$1700$ ps, for both $n_e \approx 1.8$ and $3n_c$. The ion mean-free-path due to (ion-ion) collisions for ions with $v = v_{hb}$ is $\ell_{mf} \approx 30$ mm [17]. This is much larger than the scale of the measured features ($< 100$ $\mu$m) and even the gas jet, indicating that the shock must initially be collisionless. At the latest times measured, as the shock slows, $\ell_{mf} \propto v^4$ decreases significantly, and collisions play a role in dissipation.

To further investigate this behaviour, simulations of the interaction were performed using particle-in-cell (PIC) codes. 2D3V simulations were performed with the code osiris [18]. A circularly polarised $\lambda = 10.6$ $\mu$m pulse, with $\tau = 6$ ps, $a_0 = 0.55$, $w_L = 70$ $\mu$m was focused on a fully ionised, initially cold ($T_e = T_i = 0$) He plasma of density $n_e = 1.8 - 6n_c$ with a step density gradient.

Fig. 4 shows the plasma density profile from the sim-
was irradiated with an a

simulations were performed with the code

lution with a more realistic box size, a series of 1D PIC

of the shock front behaviour. To model the shock evo-

µ

was limited to 60

more rapidly, and the shock moves faster.

why though it is clearly visible from shadowgraphy (2 a-

c), it is not properly resolved in the interferograms (2 d-
f). The potential of this front (0.15 MV) is so large that

and so moves ahead of the critical surface. The density

at this shock can be as high as

= 5 keV), where

v_{hb} is only just higher than the

initial sound speed, c_{s} = \sqrt{kT_{e}/m_{i}}, a series of solitary

waves are pushed into the plasma. Decreasing the initial

temperature of the simulation to 5 keV increases the ra-
tio of piston velocity to the local sound speed, and the

ion density profiles (e.g. fig. 4b) show the formation of

a collisionless electrostatic shock. The shock moves at

v_{sh} \approx 4 \times 10^{6} \text{ms}^{-1}, compared with

v_{hb} \approx 3 \times 10^{6} \text{ms}^{-1},

and so moves ahead of the critical surface. The density

for He

n_{e} \approx 1 \text{keV}.

From 1DPIC at the peak

FIG. 3. (colour online) a) Shock propagation distance z_{sh}

with time from shadowgrams for

n_{e} \in (1.8, 3) n_{c} (red circles,

blue squares). Green solid line indicates initial v_{hb}, and black

(dashed, dotted) lines represent the mean final v_{sh} for

n_{e} \in (1.8, 3) n_{c}; b) z_{sh} from 2DPIC for

n \in (1.8, 3, 6) n_{c} (red circles,

blue squares, green diamonds). The laser stops at t \approx 7.5 \text{ps},

indicated by vertical dotted line

ulation just after the laser pulse. It shows a density in-
crease at the front along with the cavity formed by the

hole-boring, in agreement with what is seen experiment-
ally. The front is formed by radiation pressure displac-
ing plasma electrons to a collisionless skin-depth. The

ions, which due to their greater mass are unaffected by

the laser fields, accumulate at the front due to the space-
charge field of the ensuing electrostatic double layer [14].

The large electrostatic field associated with this front also

accelerates upstream ions. This leads to the formation of

a large population of energetic ions that can be seen to

outrun the shock front in fig. 4a.

Since the plasma ahead of the front is initially cold,

hole-boring causes a compression of the plasma. The

piston-like effect of the radiation pressure launches a

shock into the plasma. The motion of the shock front
during the simulation is shown in fig. 3 for three dif-
frent starting

n_{e}. Although initialised cold, the plasma is

heated by vacuum heating [20], due to the buckling of the

front surface from hole-boring. Hence for lower

n_{e},

where hole-boring is faster, the rate of heating increases

more rapidly, and the shock moves faster.

To obtain high resolution, the simulation box for 2D

simulations was limited to 60 \mu m longitudinally. As a re-

sult a sheath quickly forms at the rear of the simulation

box due to high energy electrons that leave the box. This

sheath reflects electrons back into the plasma much more

efficiently than in the experiment, where the plasma size

was > 500 \mu m. This results in an unphysically high back-
ground temperature

T_{c}, and consequently poor modelling of the shock front behaviour. To model the shock evo-

lution with a more realistic box size, a series of 1D PIC

simulations were performed with the code EPOCH1D. A

fully ionised He plasma with a top hat profile of

n_{e} = 2 n_{c}

was irradiated with an a_{0} = 1 circularly polarised laser,

\tau_{L} = 5 \text{ps} (gaussian) pulse with \lambda = 10.6 \mu m. The simu-
lation box was 200 \mu m, with 5000 cells and

2.5 \times 10^{5}

particles per species. The ions were set to be initially cold

(T_{i} = 0 \text{eV}). 1D simulations cannot model the buckling of

the plasma due to hole-boring. Therefore, to assess the

influence of heating, the initial electron temperature was

varied between

T_{ei} = 0 - 50 \text{keV}. These simulations are comparable to those in [4], as the lower intensity here is

compensated by reduced densities to give similar piston

velocities, and also in [6], which discusses the role of ion

temperature on the shock structure.

Whilst the laser is on, the laser bores a hole into the

plasma irrespective of

T_{ei}, although the hole-boring speed decreases slightly for higher temperatures due to the in-
crease in the electron thermal pressure. For the highest

T_{ei} (= 50 \text{keV}), where

v_{hb} is only just higher than the

initial sound speed, c_{s} = \sqrt{kT_{e}/m_{i}}, a series of solitary

waves are pushed into the plasma. Decreasing the initial

temperature of the simulation to 5 keV increases the ra-
tio of piston velocity to the local sound speed, and the

ion density profiles (e.g. fig. 4b) show the formation of

a collisionless electrostatic shock. The shock moves at

v_{sh} \approx 4 \times 10^{6} \text{ms}^{-1}, compared with

v_{hb} \approx 3 \times 10^{6} \text{ms}^{-1},

and so moves ahead of the critical surface. The density

at this shock can be as high as

n_{e} = 10 n_{c}. Fig. 4 shows that the shock width (\sim c/\omega_{p}) is less than 1 \mu m, which

is why though it is clearly visible from shadowgraphy (2 a-
c), it is not properly resolved in the interferograms (2 d-
f). The potential of this front (0.15 MV) is so large that

the associated electric field, which reaches > 200 \text{GVm}^{-1},

efficiently reflects nearly all upstream ions. Phase space

diagrams (fig. 4b) show that the reflected ions gain a

FIG. 4. (colour online) a) Density distribution from 2DPIC

for

n_{e} = 1.8 n_{cr}, T_{ei} = 0 \text{at } t = 9 \text{ps. b) Evolution of 1D

He}^{2+} \text{charge density for } n_{e} = 2 n_{cr}, T_{ei} = 5 \text{keV, at } t = 5 \text{ps}

(blue, corresponds to end of laser pulse), 15 \text{ps (red)} and 30

ps (green). v_{x} x \text{ phase space for } \text{He}^{2+} \text{from 1DPIC at the peak}

of laser for c) T_{ei} = 5 \text{ and d) } T_{ei} = 1 \text{ keV.}
speed of $2v_{sh}$ by ‘bouncing-off’ the shock front. Whilst the laser is active, it is $v_{sh}$ that determines the accelerated ion velocity (not $v_{bh}$). This results in a remarkably narrow energy spread for the reflected ions $\delta E\ll E$. With $T_{ei} = 5\,\text{keV}$, there is partial reflection and partial transmission through the shock front. The transmitted ions display the characteristic for a shock solution with partial particle reflection, displaying numerous trailing perturbations in the potential, as is evident in Fig. 3: (and 4v) 21). Further reducing $T_{ei} < 1\,\text{keV}$ (so $v_{bh} \gg c_s$, fig. 41)), a region of high ion density is formed right at the hole-boring front, creating a double layer moving at $v_{sh} \approx v_{bh}$, which reflects nearly 100% of upstream ions.

Therefore, the simulations show three separate interaction regimes for $T_e \gg T_i$: 1) for $v_{bh} \ll 3c_s$, the laser launches solitary waves into the plasma from which there is no particle reflection; 2) for $v_{bh} \gtrsim 3c_s$, a shock forms moving ahead of the hole-boring front with partial particle reflection; and 3) for $v_{bh} \gg c_s$ there is 100% reflection and $v_{sh} \approx v_{bh}$ 4 21). When considering the interaction as a laser plasma ion source, there is therefore a tradeoff between higher maximum energies from a hotter plasma, where a small fraction of the particles are accelerated up to higher energies, and higher efficiency from a colder plasma where all the particles are accelerated to a smaller maximum energy. This explains the larger flux of accelerated protons observed with a short, circularly-polarised laser beam 8, as opposed to a longer, linearly-polarised driver where heating is stronger 4.

At the end of the laser pulse, the potential associated with the shock is still sufficient to reflect ions. However, as there is no longer a pressure being exerted by the laser, this particle reflection reduces the energy of the ion front, causing it to slow until reaching a critical Mach number, $M_s$, where particle reflection practically stops. Assuming $T_i = 0$, one expects a lower limit of $M_s = 1.6$ given an isothermal approximation for the plasma, whereas assuming maximum electron trapping behind the shock gives an upper value of $M_s \sim 3$ 21). In the simulations, at the end of the laser pulse the shock speed drops to $v_{sh} \sim 3M_sc_s$. For the ‘cold’ simulations ($T_{ei} = 0\,\text{keV}$), this results in the shock practically stopping, not matching our observations.

For simulations with a $T_{ei} \neq 0, T_e$ decreases with time as energetic electrons leave the simulation box, causing the soliton speed to exceed $M_sc_s$. This stimulates further particle reflection until $v_{sh} \sim M_sc_s$ again. Though artificial, this does partly imitate reality, since the shock will lose energy as it expands transversely whilst moving forward. $M_s$ is seen to decrease from 3 to $< 2$ during the simulated shock propagation, suggesting a transition between the maximal trapping limit and the isothermal limit. By the end of the simulation ($t = 95\,\text{ps}$), the ion soliton is still visible moving in the plasma with a coherent structure. Therefore, in a cooling plasma $v_{sh}$ constantly decreases to maintain $M = M_s$, which is itself decreasing. Eventually $v_{sh}$ reduces such that $\ell_{nf}$ becomes less than the shock thickness, and the shock becomes collisional. Steepening of the shock profile is observed again at very late times due to collisions (fig.4). This phase is not modelled in the PIC simulations.

In conclusion, the dynamics of the shock driven into an overdense target due to hole-boring have been observed, from the double-layer to ion acoustic soliton and finally to collisional phases, for the first time experimentally. These results hence provide insight into the dynamics of shock acceleration in plasmas, which may prove to be an interesting route to high energy ion production.

* Present address: Fakultät für Physik, Ludwig-Maximilians-Universität München and Max-Planck-Institut für Quantenoptik, Garching, Germany

The work was part funded by RUCU Libra Basic Tech and US DOE DE-FG02-07ER41488 grants. We thank the OSIRIS consortium (UCLA/IST) for use of OSIRIS and K. Kusche and the ATF technical staff for experimental assistance.

[1] N. F. Ness et al., J. Geophys. Res., 69, 3531 (1964); J. W. Freeman, J. Geophys. Res., 69, 1691 (1964).
[2] C. F. McKee, Astrophys. J., 188, 335 (1974).
[3] W. I. Axford and G. C. Reid, J. Geophys. Res., 68, 1793 (1963); E. N. Parker, Phys. Rev. Lett., 14, 55 (1965); F. A. Aharonian et al., Nature, 432, 7013 (2004); Y. Uchiyama et al., Nature, 449, 576 (2007).
[4] J. Denavit, Phys. Rev. Lett. 69, 3052 (1992);
[5] L. O. Silva et al., Phys. Rev. Lett., 92, 015002 (2004); M. Chen et al., Phys. Plasmas, 14, 053102 (2007).
[6] A. Macchi et al., Phys. Rev. E., 85, 046402 (2012).
[7] E. L. Clark et al. Phys. Rev. Lett., 84, 670 (2000); A. Maksimchuk et al., Phys. Rev. Lett., 84, 4108 (2000); A. Henig et al., Phys. Rev. Lett., 103, 245003 (2009).
[8] C. A. J. Palmer et al., Phys. Rev. Lett., 106, 014801 (2011).
[9] D. Haberberger et al., Nat. Phys., 8, 1 (2012).
[10] S. C. Wilks et al., Phys. Rev. Lett., 69, 1383 (1992).
[11] R. Kodama et al., Phys. Rev. Lett., 77, 4906 (1996).
[12] M. Zepf et al., Phys. Plasmas, 3, 3242 (1996).
[13] K. Takahashi et al., Phys. Rev. Lett., 84, 2405 (2000).
[14] M. S. Wei et al., Phys. Rev. Lett. 93, 155003 (2004); P. M. Nilson et al., Phys. Rev. Lett. 103, 255001 (2009).
[15] A. R. Bell et al., Phys. Rev. A, 38, 1363 (1988); T. Morita et al., Phys. Plasmas, 17, 122702 (2010); Y. Kuramitsu et al., Phys. Rev. Lett. 106, 175002 (2011); L. Romagnani et al., Phys. Rev. Lett. 101, 025004 (2008).
[16] Z. Najmudin et al., Phys. Plasmas, 18, 056705 (2011).
[17] L. Spitzer, Physics of fully ionized gases, 2nd ed. (Interscience, New York, NY, 1962).
[18] R. A. Fonseca, Lecture Notes in Computer Science (Springer, Heidelberg, 2002), Vol. III-342, p. 2329.
[19] L. P. Block et al., Astrophys. Space Sci., 55, 59(1978).
[20] F. Brunel, Phys. Rev. Lett. 59, 52 (1987).
[21] D. W. Forslund and C. R. Shonk, Phys. Rev. Lett. 25, 1699 (1970); D. W. Forslund and J. P. Freidberg, Phys. Rev. Lett. 27, 1189 (1971).