Quantum $1/f$ Noise in Equilibrium:
from Planck to Ramanujan

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Summary. We describe a new model of massless thermal bosons which predicts
an hyperbolic fluctuation spectrum at low frequencies. It is found that the partition
function per mode is the Euler generating function for unrestricted partitions \( p(n) \).
Thermodynamical quantities carry a strong arithmetical structure: they are given
by series with Fourier coefficients equal to summatory functions \( \sigma_k(n) \) of the power
of divisors, with \( k = -1 \) for the free energy, \( k = 0 \) for the number of particles
and \( k = 1 \) for the internal energy. Low frequency contributions are calculated using Mellin transform methods. In particular the internal energy per mode diverges
as \( \frac{\tilde{E}}{kT} = \frac{\pi^2}{6} x \) with \( x = \frac{h\nu}{kT} \) in contrast to the Planck energy \( \tilde{E} = kT \). The theory
is applied to calculate corrections in black body radiation and in the Debye solid.
Fractional energy fluctuations are found to show a \( 1/\nu \) power spectrum in the low
frequency range. A satisfactory model of frequency fluctuations in a quartz crystal resonator follows. A sketch of the whole Ramanujan–Rademacher theory of partitions is reminded as well.

1 Introduction

According to the equipartition law of statistical mechanics, the available noise
power \( \tilde{P} \) in the frequency interval \( d\nu \) is equal to \( kT d\nu \) [1]: this result is es-
sentially Nyquist’s theorem for the voltage noise \( \langle v^2 \rangle \) at a resistor \( R \), i.e.
\[
\tilde{P} = \langle v^2 \rangle = kT d\nu,
\]
where \( \langle \rangle \) means the average value. Since \( \tilde{E} = kT \) is
the mean energy per mode, Nyquist proposed to add quantum corrections as
\[
\frac{\tilde{E}}{kT} = p(x),
\]
with the Planck’s factor \( p(x) = \frac{x}{\exp(x) - 1} \) in which \( x = h\nu/kT \).
This result was generalized as \( \frac{\tilde{E}}{kT} = p(x) + \frac{x}{2} \) to account for the zero point
energy. There are still controversies concerning the physical relevance of these relations: the Planck factor removed the ultraviolet divergence but this was reincorporated in the frame of quantum electrodynamics [2],[3].

At the present stage, quantum statistical mechanics does not include infrared corrections of the \( 1/\nu \) type. The infrared catastrophe was studied pre-
viously in non-stationary processes such as the scattering of electrons in an atomic field [1, 2].

Here we derive an alternative approach in which $1/\nu$ noise is a property of non degenerate bosons in equilibrium. We first observe that the quantum mechanical partition function of a boson gas, with equally spaced energy levels, is the Euler generating function of $p(n)$: the number of indiscernible collections of the integer $n$. It relates to elliptic modular functions which are very exactly known. As a result the main contribution in the mean energy per mode is the infrared term $\tilde{E}_k T = \frac{\pi^2}{6} x$ instead of unity. The new law also leads to $1/\nu$ fractional energy fluctuations of the whole gas.

Using the new approach and the density of states of the conventional approach we calculate the corrections to black-body radiation laws, including the density of photons, the emissivity and infrared fluctuations. We also apply the calculations to the phonon gas in a quartz resonator.

The partition function $Z$ of a non degenerate boson gas is given from

$$\ln Z = -\sum_s \ln[1 - \exp(-\beta \epsilon_s)],$$

where the summation is performed over all the states $s$ of the assembly. In the conventional approach it is thus considered that the partition function $\tilde{Z}$ per mode of frequency $\epsilon_s = h\nu_s$ is such that $\ln \tilde{Z} = -\ln[1 - \exp(-\beta \epsilon_s)]$.

In black-body radiation one accounts for the wave character of the quantum states by counting the number $l$ of wavelengths in a cubic box of size $L$

$$l^2 = \frac{\nu^2 L^2}{c^2} = l_1^2 + l_2^2 + l_3^2,$$

where the summation (1) should be performed over all integers $l_1, l_2, l_3$ obeying (2).

This can be achieved by removing the discreteness of energy levels and replacing the sum (1) by an integral

$$\ln Z = -\int_0^{+\infty} D(\nu) \ln[1 - \exp(-\beta h\nu)] d\nu,$$

with $D(\nu) = 2 \times \frac{4\pi V \nu^2}{c^3}$ the density of states [3]: the factor 2 happens due to the two degrees of freedom of polarization, $c$ is the light velocity and $V = L^3$ the volume of the cavity.

From now we consider that to each mode is associated a set of equally spaced energy levels $n\nu$, $n$ integer, so that the partition per mode becomes

$$\ln \tilde{Z} = -\sum_{n \geq 1} \ln[1 - \exp(-n\beta h\nu)].$$

As shown in Sect. (2.1) this accounts for new multiparticle microstates not considered so far.
The summation above is well known in number theory and can be very accurately described using elliptic modular functions. As will be shown, there are drastic consequences in the low frequency part of the spectrum, while the high frequency part is left unchanged.

In the following the thermodynamical quantities will be defined as usual

\[ N = -\frac{\partial \ln Z}{\partial (\beta \epsilon_a)}, \]  
\[ E = -\frac{\partial \ln Z}{\partial \beta}, \]  
\[ S = \frac{\partial F}{\partial T}, \]  
\[ u = kT^2 \frac{\partial \ln Z}{\partial T}, \]  
\[ F = -kT \ln Z, \]  
\[ \epsilon^2 = kT^2 \frac{\partial E}{\partial T}, \]

(5) \hspace{1cm} (6) \hspace{1cm} (7) \hspace{1cm} (8) \hspace{1cm} (9) \hspace{1cm} (10)

In all the paper the subscript \( \sim \) will indicate that we restrict the calculation to one single mode.

2 Thermodynamics of the Euler Gas

2.1 Euler generating function

The partition function per mode \( \tilde{Z} \) in (4) can be written in the Euler form \[ \tilde{Z}(y) = \prod_{n \geq 1} \frac{1}{1 - y^n} = \sum_{n \geq 1} p(n)y^n, \]

(11)

with \( y = \exp(-x) \) and \( x = \frac{h\nu}{kT} \). This is equivalent to the Boltzmann summation

\[ \tilde{Z}(y) = \sum_{n \geq 1} p(n) \exp(-nx), \]

(12)

where \( p(n) \) is the degeneracy parameter of the energy level \( n h\nu \). It is known in number theory as the number of unrestricted partitions of the integer \( n \), that is the number of different ways of calculating \( n \) as a sum of integers.

For example with \( n = 4 \) we have \( p(4) = 5 \) and the corresponding indiscernible collections are

\[ 4 = 4 + 0 + 0 + 0 \quad (a), \]
\[ 4 = 1 + 1 + 1 + 1 \quad (b), \quad 4 = 2 + 2 + 0 + 0 \quad (c), \]
\[ 4 = 3 + 1 + 0 + 0 \quad (d), \quad 4 = 2 + 1 + 1 + 0 \quad (e). \]

(13)
This can be pictured in terms of the energy levels. The collection (a) means one particle on the level of index 4 and the three remaining particles on the ground state of index 0, i.e. $4\nu = 1 \times 4\nu + 3 \times 0\nu$. This collection is the only one considered in the conventional approach. The others microstates from (b) to (d) corresponds to different possibilities of bunching of the particles. Collection (b) means the four particles on the level of index 1, i.e. $4\nu = 4 \times 1\nu$, collection (c) means two particles on the level 2 and two particles on the ground state, i.e. $4\nu = 2 \times 2\nu + 2 \times 0\nu$, and so on.

Properties of Euler generating function were studied in full details by Ramanujan [7] in 1918 and completed by Rademacher [8] in 1973. An important result is the asymptotic formula

$$p(n) \sim \frac{1}{4n^{3}} \exp(\pi \sqrt{2n/3}) \text{ when } n \to \infty. \quad (14)$$

If instead of eq. (11) one uses the classical partition function

$$Z = \sum_{n \geq 1} \exp(-nh\nu/kT),$$

one recovers the Planck’s factor $E/kT = x/(\exp(x) - 1)$. Such a dependence remains valid at high frequencies $\hbar\nu kT$. If one introduces the zero point energy through the formula $\epsilon_{s,n} = (n + \frac{1}{2})\nu$ with $n \geq 1$ instead of $\epsilon_{s,n} = n\nu$, then the low frequency part of the quantum spectrum is left unchanged but the ultraviolet catastrophe that one gets from the formula $E/kT = x/(\exp(x) - 1) + x/2 = x \coth(x/2)$, is cancelled as shown in Fig. 1.

The new quantum theory thus solves simultaneously the questions concerning the $1/\nu$ power spectrum (which is observed but was not predicted) and the ultraviolet catastrophe (which is not observed but was predicted) [8].
2.2 Riemann zeta function and the free energy of the Euler gas

The partition function $\tilde{Z}(x)$ defined in (4) is related to the Riemann zeta function $\zeta(s)$ through the Mellin transform as follows (Ref. [10], Eq. (6.3) p. 464)

$$\Gamma(s)\zeta(s)\zeta(s+1) = \int_{0}^{\infty} (-\ln \tilde{Z}(x))x^{s-1}dx.$$  \hfill (15)

Introducing $\sigma_k(n)$ as the sum of $k^{th}$ powers of the divisors of $n$ and the Dirichlet series $\zeta(s)\zeta(s-k)$ as follows

$$\zeta(s)\zeta(s-k) = \sum_{n \geq 1} \frac{\sigma_k(n)}{n^s},$$  \hfill (16)

and computing the inverse Mellin transform one obtains (Ref. [10], p. 467) the free energy $\tilde{F}$ as follows

$$\frac{\tilde{F}}{kT} = -\ln \tilde{Z}(x) = \sum_{n \geq 1} \ln(1-\exp(-nx)) = -\sum_{n \geq 1} \sigma_{-1}(n) \exp(-nx).$$  \hfill (17)

with $\sigma_{-1}(n) = \sigma_{1}(n)/n$. There is thus a close relationship between the arithmetic of $p(n)$ and that of divisors. This will be confirmed in the derivation of the others thermodynamical quantities.

Since the main contributions to $\tilde{Z}(x)$ are given by the poles at $s = 0$ of $\Gamma(s)$ and $\zeta(s+1)$ and at $s = 1$ of $\zeta(s)$, the free energy may be approximated in the low frequency part of the spectrum (Ref. [11], p. 58)

$$\frac{\tilde{F}}{kT} \simeq -\frac{\pi^2}{6x} - \frac{1}{2} \ln\left(\frac{x}{2\pi}\right) + \frac{x}{24},$$  \hfill (18)

with the error term $\sum_{l \geq 1} \ln(1-\exp(-4\pi^2 l/x))$.

2.3 Dedekind eta function and the internal energy of the Euler gas

One easily shows that the Mellin transform of the occupation number $\tilde{N}(x)$ is $\Gamma(s)\zeta(s)^2$. A similar derivation to the one performed in Sect. 2.2 leads to

$$\tilde{N}(x) = -\frac{\partial}{\partial(nx)} \ln \tilde{Z}(x) = \sum_{n \geq 1} \frac{1}{\exp(nx)-1} = \sum_{n \geq 1} \sigma_{0}(n) \exp(-nx).$$  \hfill (19)

In the low frequency region one gets (see [9], p. 27)

$$\tilde{N}(x) \simeq -\frac{\ln x + \gamma}{x}.$$  \hfill (20)

where $\gamma(x) \simeq 0.577$ is Euler constant.
The Mellin transform of the internal energy $\tilde{E}(x)$ is
\[
\Gamma(s) \zeta(s) \zeta(s-1)
\]
and from the same method than above
\[
\tilde{E}(x) = -\frac{\partial (\ln \tilde{Z}(x))}{\partial (\beta)} = h\nu \sum_{n \geq 1} \frac{n}{\exp(nx) - 1} = h\nu \sum_{n \geq 1} \sigma_1(n) \exp(-nx). \tag{21}
\]

An alternative derivation involves elliptic modular aspects. According to Ninhm \[10\]: All mathematics is a tautology, and all physics uses mathematics to look in different ways at a fundamental problem of philosophy - how to bridge the discrete and continuous.

The link between the modular group and the Euler generating function is from the equality \[8\]
\[
\tilde{Z}(y) = \prod_{n \geq 1} \frac{1}{1 - y^n} = \sum_{n \geq 1} p(n) y^n = \frac{\exp(i\pi \tau/12)}{\eta(\tau)}, \tag{22}
\]
where the domain of integration of $\tilde{Z}(y)$ is taken to be the upper half complex plane of the new variable $\tau$
\[
y = \exp(2i\pi \tau), \quad \Im(\tau) > 0. \tag{23}
\]

Here we have
\[
\Im(\tau) = \frac{x}{2\pi}, \quad x = \frac{h\nu}{kT}. \tag{24}
\]

As shown in Sect. \[5\] Dedekind eta function acts on the full modular group $SL(2, Z)$. At this stage we do not enter into the full ramifications of the theory and only emphasizes the connexion to the modular Eisenstein function
\[
G_2(\tau) = \sum_{m, n} \frac{1}{(m\tau + n)^2}, \tag{25}
\]
where the summation is performed over all non zero relative integers $m$ and $n$ and $\Im(\tau) > 0$. It can be shown \(12\), p. 29) that $G_2(\tau)$ connects to the logarithmic derivative of $\eta(\tau)$
\[
G_2(\tau) = -4\pi \frac{d (\ln(\eta(\tau)))}{d\tau}, \tag{26}
\]
with a Fourier expansion
\[
G_2(\tau) = 2\zeta(2) + 2(2i\pi)^2 \sum_{n \geq 1} \sigma_1(n) \exp(2i\pi n \tau). \tag{27}
\]

Using \[22\], \[23\] and \[25\]–\[27\] the relation \[21\] is easily recovered.

The low frequency expansion of internal energy is as follows
\[
\frac{\tilde{E}}{kT} \simeq \frac{\pi^2}{6x} - \frac{1}{2} + \frac{x}{24}. \tag{28}
\]
instead of the Planck result $\tilde{E} \approx kT$.

One can also compute the entropy $\tilde{S}$ from

$$\frac{\tilde{S}}{k} = \frac{\tilde{E}}{kT} + \ln(\tilde{Z}) = \frac{\tilde{E} - \tilde{F}}{kT} = \sum_{n \geq 1} \sigma_1(n) (x + 1/n) \exp(-nx)$$

$$\simeq \frac{\pi^2}{3x} + \frac{1}{2} \ln \frac{x}{2\pi} - \frac{1}{2} \quad \text{when} \quad x \to 0. \quad (29)$$

At very low frequency it results that the internal energy equals the opposite of free energy and one half the entropy.

3 Application to Black-body Radiation

3.1 Stefan-Boltzmann constant revisited

The Stefan-Boltzmann constant is an integrated measure of the emissivity of a black body [6]. In the conventional approach the partition function is calculated from the integral (3) using the density of states $D(\nu) = \frac{8\pi V \nu^2}{c^3}$ that is

$$\ln Z = 8\pi V \left( \frac{kT}{ch} \right)^3 \times 2\zeta(4), \quad (30)$$

with $\zeta(4) = \frac{\pi^4}{90}$ and we used the Mellin integral formula

$$\zeta(s + 1) = -\frac{1}{\Gamma(s)} \int_0^\infty x^{s-1} \ln(1 - \exp(-x))dx. \quad (31)$$

The Stefan-Boltzmann constant $\sigma_{SB}$ is defined from the free energy

$$F = -kT \ln Z = -\frac{4\sigma_{SB}}{3c} VT^4 \quad \text{with} \quad \sigma_{SB} = \frac{2\pi^5 k^4}{15 c^2 h^3}. \quad (32)$$

If instead of (30) one uses the general formula

$$\ln Z = -8\pi V \left( \frac{kT}{ch} \right)^3 \int_0^\infty x^2 \sum_{n \geq 1} \ln(1 - \exp(-nx))dx, \quad (33)$$

the interchange of the integral and the sum leads to

$$\ln Z = -8\pi V \left( \frac{kT}{ch} \right)^3 \times 2\zeta(4) \times \left( \sum_{n \geq 1} \frac{1}{n^3} \right). \quad (34)$$

As a result we find a free energy (and a modified Stefan-Boltzmann constant) in excess with a factor $\sum_{n \geq 1} \frac{1}{n^3} = \zeta(3) \simeq 1.20$. If one uses the alternative derivation in terms of the divisors one recovers the mathematical formula [16] with $s = 3$ and $k = -1$. 
3.2 The density of photons

The number of photons in the bandwidth $d\nu$ is

$$dN(\nu, T) = D(\nu)\tilde{N}(\nu, T)d\nu, \quad (35)$$

with the occupation number $\tilde{N}(\nu, T) = (\exp(\beta\nu) - 1)^{-1}$ in the conventional approach. Integrating one gets per unit volume

$$\frac{N(T)}{V} = 8\pi\left(\frac{kT}{ch}\right)^3 \int_0^{+\infty} \frac{x^2dx}{\exp(x) - 1} = 8\pi\left(\frac{kT}{ch}\right)^3 \times 2\zeta(3). \quad (36)$$

In the general approach the occupation number is defined from the summation and we need to evaluate

$$\int_0^{\infty} x^2 \sum_{n\geq 1} \frac{1}{\exp(nx) - 1} = 2\zeta(3) \left( \sum_{n\geq 1} n^{-3} \right) = 2\zeta(3)^2. \quad (37)$$

This is in excess of a factor $\zeta(3) \simeq 1.20$ as for the free energy. If one compares the calculation in terms of divisors one recovers the mathematical formula with $s = 3$ and $k = 0$.

3.3 Planck’s radiation formula revisited

The energy within the bandwidth $d\nu$ is defined as

$$dE(\nu, T) = D(\nu)\tilde{E}(\nu, T)d\nu = u(\nu, T)d\nu, \quad (38)$$

with $\tilde{E}(\nu, T) = h\nu(\exp(\beta h\nu) - 1)^{-1}$ in the conventional approach and with $u(\nu, T)$ the energy spectral density. We get the Planck’s radiation formula

$$u(\nu, T) = \frac{8\pi hV}{c^3} \frac{\nu^3}{\exp(\beta h\nu) - 1}. \quad (39)$$

The black-body emissivity is defined as

$$e_b(\nu, T) = \frac{c}{4V} u(\nu, T). \quad (40)$$

At very low frequency the conventional result leads to the Rayleigh-Jeans formula

$$[e_b(\nu, T)]_{RJ} = 2\pi\frac{k}{c^2\nu^2}T, \quad (41)$$

which is independent of Planck’s constant and is proportional to the inverse of the square of wavelength $\lambda = c/\nu$.

In the new approach the emissivity is
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\[ e_b(\nu, T) = \frac{2\pi\hbar}{c^2} \sum_{n \geq 1} \frac{n\nu^3}{\exp(n\beta\hbar\nu) - 1}. \]  

(42)

At very low frequency one uses \([43]\) with the result

\[ [e_b(\nu, T)]_{LF} = \frac{\pi^3}{3} \frac{k^2}{c^2 \hbar} \nu T^2. \]  

(43)

Thus the \(\nu^2 T\) dependance is replaced by the \(\nu T^2\) dependance, the low frequency emissivity now depends on the Planck constant and on the inverse wavelength; there is a ratio \(\frac{\pi}{6}\) between the new result and the one predicted by the Rayleigh-Jeans formula.

3.4 Radiative atomic transitions

Let us now consider the equilibrium between atoms and a radiation field, allowing the emission or absorption of photons of frequency

\[ \nu = \nu_1 - \nu_2, \]  

(44)

where \(\epsilon_2 = h\nu_2\) is the energy in the upper state and \(\epsilon_1 = h\nu_1\) in the lower state.

The conventional theory, as derived for the first time by Einstein, states that the rate at which atoms make a transition \(1 \rightarrow 2\) in which one photon is absorbed is equal to the rate at which atoms emit photons, so that

\[ B_{12} N_1 u(\nu) = A_{21} N_2 + B_{21} N_2 u(\nu), \]  

(45)

where \(N_1\) and \(N_2\) are the occupation numbers of atoms in levels 1 and 2, \(A_{21}\) is the spontaneous absorption rate, \(B_{12}\) is the induced emission rate and \(u(\nu)\) is the energy density in the radiation field as given in \([39]\).

In thermal equilibrium the occupation numbers in states 1 and 2 obey the Boltzmann law

\[ \frac{N_2}{N_1} = \exp\left(-\frac{\hbar \nu}{kT}\right). \]  

(46)

Using \([46]\) and \([39]\) one gets the well known formulas

\[ A_{21} = A, \quad B_{12} = B_{21} = B \quad \text{and} \quad \frac{A}{B} = \frac{8\pi^4 kT}{\lambda^3}, \]  

(47)

where \(\lambda = c/\nu\) is the wavelength of the radiation field.

If one uses the general formula one gets low frequency corrections in the spontaneous to stimulated emission ratio. Using the low frequency expression \([28]\) for the internal energy this yields

\[ \left[ \frac{A}{B} \right]_{LF} = A \times \frac{\pi^2}{6\pi} = \frac{4\pi^3 kT}{3c\lambda^2}. \]  

(48)
The $A/B$ low frequency ratio now depends on the inverse of the square of wavelength $\lambda$ and is independent on the Planck constant $h$, in contrast to the $h/\lambda^3$ dependence of the standard ratio. The spontaneous to stimulated absorption rate is enhanced a factor $\frac{\pi^2}{6}$ over the conventional one.

### 3.5 Einstein’s fluctuation law revisited

According to the conventional Einstein's approach [13] the energy fluctuations of a system in equilibrium within a larger system of temperature $T$ are

$$
\epsilon^2 = \langle (E - \langle E \rangle)^2 \rangle = \frac{\partial \langle E \rangle}{\partial \beta} = kT^2 \frac{\partial \langle E \rangle}{\partial T}. \quad (49)
$$

One can reformulate this relation for the fluctuations of the energy $dE = u(\nu, T) d\nu$ in the bandwidth $d\nu$

$$
d\epsilon^2 = kT^2 \frac{\partial u}{\partial T} d\nu = S_u(\nu) d\nu, \quad (50)
$$

with $u$ the energy spectral density and $S_u(\nu)$ the power spectral density of the fluctuations of $u$. One gets ([13], p. 429)

$$
d\epsilon^2 = (h\nu u + \frac{c^3}{8\pi^2 V} u^2) V d\nu. \quad (51)
$$

The first term at the right hand side is the one corresponding to the high frequency part of the spectrum (Wien’s law): it is of a pure quantum nature and is corpuscular like; the second one corresponds to the low frequency (Rayleigh-Jeans) region: it is purely classical and wave-like. In the low frequency part of the spectrum they are fractional energy fluctuations of the random walk type

$$
\left[ \frac{S_u(\nu)}{u^2} \right]_{\text{RJ}} = \frac{c^3}{8\pi^2 V} \frac{1}{\nu^2}. \quad (52)
$$

In the new approach one uses the low frequency energy density $u(\nu, T) \approx \frac{4\pi^3 V}{8\pi^2 h \beta^2} \frac{\nu d\nu}{c^3}$ so that instead of (52) one gets

$$
\left[ \frac{S_u(\nu)}{u^2} \right]_{\text{LF}} = \frac{3 h c^3}{2 \pi^3 V k T \nu}. \quad (53)
$$

This is the announced quantum $1/\nu$ fluctuation spectrum. There is a reduced low frequency noise and the ratio between the new result and the Einstein-Rayleigh-Jeans one is $\frac{12\pi}{\pi^2}$. 
4 Application to a Phonon Gas and to the 1/f Frequency Noise of a Quartz Resonator

4.1 The specific heat of a phonon gas revisited

The properties of the phonon gas are quite similar to those of the photon gas except for the new form of the density of modes as

\[ g(\nu) = \frac{12\pi V}{c_{ph}^3 \nu^2} \]

with

\[ \frac{3}{c_{ph}^3} = \frac{2}{c_t^3} + \frac{1}{c_l^3} \]

where \( c_{ph} \) represents the average wave velocity and \( c_t \) and \( c_l \) are the transverse and longitudinal velocities for an isotropic solid \[6\]. The maximal vibrational frequency \( \nu_m \) (Debye frequency) is defined from the total number of allowed quantum states

\[
3N_0 = \frac{12\pi V}{c_{ph}^3} \int_0^{\nu_m} \nu^2 d\nu \quad \text{that is} \quad \nu_m = \left( \frac{3N_0 c_{ph}^3}{4\pi V} \right)^{1/3}, \quad (54)
\]

where \( N_0 \) is the number of atoms in the volume \( V \).

In the conventional theory \[6\] we get

\[
\ln Z_{ph} = -\frac{9N_0}{\nu_m^3} \int_0^{\nu_m} \nu^2 \ln \left[ 1 - \exp(-\beta h\nu) \right] d\nu. \quad (55)
\]

The internal energy follows from the formula

\[
E = \frac{9RT}{x_m^3} D(x_m), \quad (56)
\]

and the constant volume specific heat \( C_v = \partial E/\partial T \) equals

\[
C_v = 3R \left[ D(x_m) - x_m D'(x_m) \right], \quad (57)
\]

with \( D(x_m) = \frac{3}{x_m^3} \int_0^{x_m} \frac{x^3 dx}{\exp(x) - 1} \) the Debye function, and \( x_m = \frac{\theta_D}{T} \), with \( \theta_D = \frac{h\nu_m}{k} \) the Debye characteristic temperature.

The case \( T \theta_D, D(\theta_D) \rightarrow 1 \) corresponds to the Dulong-Petit value \( C_v \sim 3R \).

At very low temperatures one gets the cubic temperature dependence \( C_v \sim \frac{4\pi^4}{3} \times 3R \left( \frac{T}{\theta_D} \right)^3 \).

In the new approach

\[
\ln Z_{ph} = -\frac{9N_0}{\nu_m^3} \int_0^{\nu_m} \nu^2 \ln \left[ 1 - \exp(-n\beta h\nu) \right] d\nu. \quad (58)
\]

Debye results are found unchanged except for an extra multiplicative factor in the specific heat as was the case for the integrated emissivity in Sect. 3.1 that is

\[ \frac{[C_v]_{new}}{C_v} = \zeta(3) \sim 1.20. \quad (59) \]

At very low temperatures the electronic contribution to the specific heat which decreases as \( T \), dominates the lattice contribution, which decreases as \( T^3 \). This is accounted for in the conventional way.
4.2 1/f noise in a quartz resonator

Specific heat is involved in the energy fluctuations of a canonical ensemble from the relation

$$\epsilon^2 = kT^2 C_v. \quad (60)$$

The relative energy fluctuations follows as $$\frac{\epsilon^2}{E^2} = \left(\frac{2}{3N_0}\right)^{1/2}$$ which is on the order $$10^{-11}$$ for $$N_0 = 10^{23}$$, the Avogadro number.

For energy fluctuations in the bandwidth $$d\nu$$, the main difference with the conventional theory lies in the low frequency region, as was the case of the photon gas. We find the quantum $$1/\nu$$ formula

$$\left[ \frac{S_u(\nu)}{u^2} \right]_{LF} = \frac{9hc_{\text{ph}}^3}{4\pi^3 V kT\nu} = \frac{A_{\text{ph}}}{V\nu}. \quad (61)$$

The method can be used to predict fractional frequency fluctuations in a quartz crystal resonator from the formula

$$\frac{S_\omega(\nu)}{\omega^2} = \frac{1}{4Q^2} \frac{A_{\text{ph}}}{V\nu} = \frac{h_{-1}}{\nu}. \quad (62)$$

where $$\omega$$ and $$Q$$ are the frequency and quality factor of the resonator. Using $$c_{\text{ph}} \sim 3.5 \times 10^3 \text{ m/s}$$, we find $$A_{\text{ph}} \sim 5 \times 10^{-4}$$. For a 5 MHz P5 quartz crystal resonator with $$Q \sim 2 \times 10^6$$, the active region under the electrodes has thickness $$t = 5\lambda/2 \sim 3 \text{ mm}$$, and section $$S \sim 3 \text{ cm}^2$$, that is $$V \sim 1 \text{ cm}^3$$. The resulting $$1/\nu$$ factor is $$h_{-1} = \frac{A_{\text{ph}}}{4Q^2V} \sim \frac{10^2}{Q^2} \sim 6 \times 10^{-24}$$. This is the order of magnitude found in experiments [14].

5 Ramanujan–Rademacher Theory of Partitions: a Short Reminder

Besides the low frequency approximations encountered in Sect. (2) there is a an exact method to calculate the number of partitions $$p(n)$$ first discovered by Ramanujan [7] and improved by Rademacher [8] thanks to an integration along Ford circles in the complex half plane. For completeness we remind here the main points of the theory from which the results in Sect. (2) may also be derived.

From well-known mathematical arguments [7] (p. 113) one can get the leading term for the case $$0 < y < 1$$ and $$y \to 1$$ from the expression

1 To establish the formula one writes the equation for a lossy harmonic oscillator and one postulates that the $$1/\nu$$ fluctuations are present in the loss coefficient.

2 We have
\[
\ln \tilde{Z}(y) \sim \frac{\pi^2}{6(1 - y)}. \tag{63}
\]

The use of \( y = \exp(-\hbar \nu/kT) \) corresponds to the low frequency approximation at \( \nu \to 0 \), that is \( 1 - y = 1 - \exp(-\hbar \nu/kT) \sim \hbar \nu/kT \). This leads to the leading low frequency term in the free energy \( (18) \) and internal energy \( (28) \).

They are similar formulas associated with rational points which are located at
\[
y_{pq} = \exp(2i\pi pq), \tag{64}
\]
on the unit circle \(|y| = 1\). The leading term in the expansion of \( \tilde{Z}(y) \) corresponds to the fundamental mode \( \frac{p}{q} = \frac{1}{1} \).

The general method to compute rational contributions is a master piece of twentieth century mathematics \( (7), (8) \). It uses the connexion of \( \tilde{Z}(y) \) to the elliptic modular functions.

### 5.1 The fundamental contribution

To compute the contribution of the fundamental point \( 1/1 \) of the unit circle \(|y| = 1\) one uses the property \( (16), \text{p. 96,} (11), \text{P. 58} \)
\[
\tilde{Z}(y) = \frac{y^{1/24}}{\sqrt{2\pi}} (\ln \frac{1}{y})^{1/2} \exp \left[ \frac{\pi^2}{6 \ln \frac{1}{y}} \right] \tilde{Z}(y') \text{ with } y' = \exp \left[ \frac{4\pi^2}{\ln y} \right]. \tag{65}
\]

In the low frequency region \( y = \exp(-x) \sim 1 \) so that \( y' = \exp(-4\pi^2/x) \sim 0 \) and \( \tilde{Z}(y') \sim 1 \). Low frequency approximations of the free energy \( (18) \) and of the internal energy \( (28) \) follow. There are similar formulas associated with the other rational points of the circle as shown below.

To get the leading term in \( p(n) \) one uses the Cauchy formula
\[
p(n) = \frac{1}{2\pi i} \oint \tilde{Z}(y) \frac{1}{y^{n+1}} dy, \tag{66}
\]
where \( \oint \) means an arbitrary closed loop encircling the origin.

\[
\log Z(y) = \sum_n \log \frac{1}{1 - y^n} = \sum_{m,n} y_m^{\frac{m}{n}} - \sum_m \frac{y_m^m}{m(1 - y^m)},
\]
and
\[
my_m^{m-1}(1 - y) < 1 - y^m < m(1 - y),
\]
so that
\[
\frac{1}{1 - y} \sum_m y_m^m < \log Z(y) < \frac{1}{1 - y} \sum_m \frac{y}{m^2}
\]

Each of the above series has the limit \( \pi^2/6 \) when \( y \to 1 \) and so \( \log Z(y) \sim \frac{\pi^2}{6(1 - y)} \).
Substituting (65) in (66) with \( \tilde{Z}(y') = 1 \) one can obtain

\[
p(n) = \frac{1}{2\pi \sqrt{2}} \frac{d}{dn} \left( \frac{\exp(K\lambda_n)}{\lambda_n} \right) \text{with } \lambda_n = \sqrt{n - \frac{1}{24}} \text{ and } K = \pi \sqrt{\frac{2}{3}}.
\]

(67)

This includes (14) in the limit \( n \to \infty \) but is much more accurate.

5.2 Farey contributions and Ford circles

From now on we extend the domain of definition of \( \tilde{Z}(y) \) to the complex plane and we introduce the new variable \( \tau \) and Dedekind eta function \( \eta(\tau) \) as defined in (22)-(24).

It can be shown that \( \eta(\tau) \) is a modular form of degree \(-1/2\) on the full modular group. It acts on the generators of such a group through the relations

\[
\eta(\tau + 1) = \exp(i\pi/12)\eta(\tau); \quad \eta(-1/\tau) = (\eta/i)^{1/2}\eta(\tau).
\]

(68)

To express the partition function one uses the Cauchy formula

\[
p(n) = \frac{1}{2\pi \tau} \int \frac{\tilde{Z}(y)}{y^{n+1}} dy = \int_\tau^{\tau_0+1} \tilde{Z} \exp(2i\pi \tau) \exp(-2i\pi n) d\tau.
\]

(69)

In the third term above this corresponds to a path of unit length starting at an arbitrary point in \( \mathcal{H} \).

The choice of the integration path comes along in a natural way by using the connexion of \( \tilde{Z}(y) \) to the modular group. Let us observe that the set of images of the line \( \tau = X + i, X \) real, under all modular transformations

\[
\tau' = \frac{p\tau + p'}{q\tau + q'}, \text{ with } p, p', q, q' \text{ integers and } |pq' - q'p| = 1,
\]

(70)

can be written as

\[
\eta^{(p\tau + p')/q\tau + q')} = \epsilon(p, p', q, q') \sqrt{q\tau + q'/i} \eta(\tau),
\]

with \( \epsilon \) a 24th root of unity related to Dedekind sums as defined in (79). See [8], p. 160.

\[^3\text{For more general modular transformations, we have}
\]

\[
\eta^{(p\tau + p')/(q\tau + q')} = \epsilon(p, p', q, q') \sqrt{q\tau + q'/i} \eta(\tau),
\]

with \( \epsilon \) a 24th root of unity related to Dedekind sums as defined in (79). See [8], p. 160.

\[^4\text{Let } f(z) \text{ be an holomorphic function of the complex variable } z. \text{ The Cauchy formula for the derivatives is as follows }
\]

\[
f^n(a) = \frac{n!}{2\pi i} \oint \frac{f(z)}{(z - a)^{n+1}} dz
\]

where \( \oint \) means a closed contour encircling the pole. It is applied in (69) to the partition function \( f \equiv Z \) with \( a = 0 \) so that \( \frac{f^n(a)}{n!} = p(n) \).
\[ \left| \tau - \left( \frac{p}{q} + \frac{i}{2q^2} \right) \right| = \frac{1}{2q^2}. \quad (71) \]

Equation 71 defines circles \( C(p, q) \) centered at points \( \tau = \frac{p}{q} + \frac{i}{2q^2} \) with radius 1/2q^2. They are named after L.R. Ford who first studied their properties in 1938 [9]. Ford circles are easily generated by using the ordered Farey sequence

\[
0 < \frac{p_1}{q_1} < \frac{p_1 + p_2}{q_1 + q_2} < \frac{p_2}{q_2} < \cdots < \frac{1}{1}. \quad (72)
\]

To each \( \frac{p}{q} \) belongs a Ford circle in the upper half plane, which is tangent to the real axis at \( \tau = \frac{p}{q} \). It can be observed that Ford circles never intersect. They are tangent to each other if and only if they belong to fractions which are adjacent in some Farey sequence.

If \( \frac{p_1}{q_1} < \frac{p}{q} < \frac{p_2}{q_2} \) are three adjacent fractions in a Farey sequence then \( C(p, q) \) touches \( C(p_1, q_1) \) and \( C(p_2, q_2) \) respectively at the points

\[
\tau_{pq}^L = \frac{p}{q} + \zeta_{pq}^L \quad \text{and} \quad \tau_{pq}^R = \frac{p}{q} + \zeta_{pq}^R, \quad (73)
\]

where

\[
\zeta_{pq}^L = -\frac{q_1}{q(q^2 + q_1^2)} + \frac{i}{q^2 + q_1^2} \quad \text{and} \quad \zeta_{pq}^R = \frac{q_2}{q(q^2 + q_2^2)} + \frac{i}{q^2 + q_2^2}. \quad (74)
\]

In Rademacher’s approach (which improves Ramanujan’s one) the unit length path on \( \mathcal{H} \) is choosen so as to go along Ford circles

![Rademacher’s path of integration.](image)
\[ p(n) = \sum_{0 \leq p, q \leq N, (p, q) = 1} \int_{\gamma_{pq}} \tilde{Z}[\exp(2i\pi \tau)] \exp(-2i\pi \tau n) d\tau, \quad (75) \]

where \( \gamma_{pq} \) is the upper arc on a Ford circle which connects points of tangency at \( \tau_{pq}^L \) and \( \tau_{pq}^R \).

Each Ford circle \( C(p, q) \) is parametrized by the expression \( \tau = \frac{p}{q} + \zeta \), where the variable \( \zeta \) runs on an arc of the circle \( |\zeta - \frac{i}{2q^2}| = \frac{1}{2q^2} \). If one uses \( z \) such that \( \zeta = \frac{iz}{q^2} \), a Ford circle is mapped onto the circle \( |z - \frac{1}{2}| = \frac{1}{2} \) and (75) transforms as

\[ p(n) = \sum_{0 \leq p, q \leq N, (p, q) = 1} \left\{ \frac{i}{q^2} \exp(-2i\pi n \frac{p}{q}) \right\} \times \int_{z_{pq}^L}^{z_{pq}^R} \tilde{Z} \left[ \exp(2i\pi \frac{p}{q} - \frac{2\pi z}{q^2}) \exp(\frac{2\pi n z}{q^2}) dz \right], \quad (76) \]

where \( z_{pq}^L \) and \( z_{pq}^R \) follows from (73).

### 5.3 Farey contributions and Dedekind sums

To compute (76) one uses the transformation formula (22). After some manipulations and using \( \frac{z}{q} \) instead of \( z \) (see Ref. [3], p. 269), one gets the formula which generalizes (65)

\[ \tilde{Z}(y) = \omega_{pq} \left( \frac{z}{q} \right)^{1/2} \exp(-\frac{\pi}{12z} - \frac{\pi z}{12q^2}) \tilde{Z}(y'), \quad (77) \]

with

\[ y = \exp(\frac{2i\pi p}{q} - \frac{2\pi z}{q^2}), \quad y' = \exp(\frac{2i\pi p'}{q} - \frac{2\pi}{z}) \text{ and } pp' = -1 \text{ (mod } q). \quad (78) \]

The so-called Dedekind sums \( s(p, q) \) are introduced by

\[ \omega_{pq} = \exp(i\pi s(p, q)) \quad \text{with} \quad s(p, q) = \sum_{l=1}^{q} \left( \frac{1}{q} \right) \left( \frac{pl}{q} - \left[ \frac{pl}{q} \right] \right), \quad (79) \]

where \([ ]\) in (79) denotes the integer part.

For the calculation of \( p(n) \) one uses an approximation similar to the one used in (65). If \( z \) is a small positive real number, then \( y \) is near \( \exp(2i\pi \frac{p}{q}) \), the modulus at that point \( |y'| = \exp(-\frac{2\pi}{z}) \sim 0 \) and \( \tilde{Z}(y') \sim 1 \).

As a result (76) can be readily integrated and the final result is

\[ p(n) = \frac{1}{\pi \sqrt{2}} \sum_{q \geq 1} \sqrt{\pi} A_q(n) \frac{d}{dn} \left( \frac{\sinh(K_q \lambda_n)}{\lambda_n} \right), \quad (80) \]
with $K_n = \frac{\pi}{q} \sqrt{\frac{2}{3}}, \quad \lambda_n = \sqrt{n - \frac{1}{24}},$

and $A_q(n) = \sum_{p \mod(q)} \omega_{pq} \exp(-2\pi i n \frac{p}{q}).$ \hspace{1cm} (81)

6 Conclusion

We have found that quantum $1/f$ noise may be a property of photons or phonons in thermal equilibrium. The theory connects elliptic modular functions and quantum statistical mechanics as in superstring theory, but with physical relevance in the macroscopic realm of infrared divergences of solid state physics. Main results are an enhanced energy per mode at low frequency (one gets $\tilde{E}_k T \simeq \frac{\pi^2 k T}{6 h \nu}$ instead of the Planck result $\tilde{E} \simeq k T$) and an enhanced integrated radiation, photon density and phonon specific heat (with a factor of $\zeta(3) \simeq 1.20$). Low frequency fluctuations are reduced and one obtains fractional energy fluctuations with a $1/\nu$ spectrum in contrast to the random walk $1/\nu^2$ of the standard theory.

Main steps of the general mathematical theory of $p(n)$ is reminded in the last section. It is based on the fact that there are rational singularities on the unit circle in the partition function. The fundamental mode in the Farey decomposition leads to a satisfactory account of the infrared part of the spectrum and remains correct in the high frequency region up to $x = x_1 = 4\pi^2$. The physical meaning of higher modes has still to be understood.

Farey series and Ford circles were used recently in the different context of $1/f$ noise in phase locked loops [17]. In this last case they connect to the arithmetical functions found in prime number theory.
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