Holographic Lessons for Quark Dynamics

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Abstract

We give a brief overview of recent results obtained through the gauge/gravity correspondence, concerning the propagation of a heavy quark in strongly-coupled conformal field theories (such as $\mathcal{N} = 4$ super-Yang-Mills), both at zero and finite temperature. In the vacuum, we discuss energy loss, radiation damping, signal propagation and radiation-induced fluctuations. In the presence of a thermal plasma, our emphasis is on early-time energy loss, screening and quark-antiquark evolution after pair creation. Throughout, quark dynamics is seen to be efficiently encapsulated in the usual string worldsheet dynamics.
1 Introduction

A substantial amount of evidence from heavy ion collision experiments at RHIC and, recently, LHC, supports the detection of the long-sought quark-gluon plasma (QGP), a hot and dense phase of deconfined strongly-interacting matter [1]. Energetic partons serve as important probes of this thermal medium, and there exists an enormous ‘jet-quenching’ literature dedicated to analyzing the manner in which the plasma damps their motion, and is in turn disturbed by their passage [2]. Given the experimental indications that, in the range of accessible temperatures (of order a few times the deconfinement temperature), the QGP is strongly-coupled, perturbative QCD is believed to be inadequate for at least some aspects of the relevant calculations, creating a demand for new theoretical tools. In the past five years, interesting steps have been taken towards meeting this demand via the AdS/CFT, or, more generally, gauge/gravity or holographic correspondence [3, 4], starting with the seminal works [5, 6, 7, 8] and continuing with a large body of work that has been reviewed in [9].

The gauge/gravity correspondence is by now well-established as a tool that grants access (often analytically) to a large class of strongly-coupled non-Abelian gauge theories, via a drastic and surprising rewriting in terms of string-theoretic (frequently just supergravity) degrees of freedom living on a curved higher-dimensional geometry. The gauge/gravity catalog includes some theories that display characteristic QCD-like physics such as confinement and chiral-symmetry breaking, but not, to date,
QCD itself. Thus, at our present stage of knowledge, contact with the real-world QGP produced at RHIC or LHC will be feasible only if QCD can be reasonably well approximated by at least one of the gauge theories whose dual description is known. In the past few years, encouraging signs in this direction have emerged even for the most rudimentary example \cite{3}, $SU(N_c)$ maximally-supersymmetric ($\mathcal{N} = 4$) Yang-Mills theory (MSYM), a conformal field theory (CFT) that at zero temperature is completely unlike QCD (and in particular, has a coupling that does not run), but at finite temperature is in various respects analogous to deconfined QCD. Besides the gauge field, MSYM contains 6 real scalar fields and 4 Weyl fermions, all in the adjoint representation of the gauge group.

Following the very exciting findings on the shear viscosity in holographic thermal plasmas \cite{10} (which in turn led to calculation of many other transport coefficients \cite{11}), analyses of energy loss have been conducted considering various types of partonic probes of the MSYM plasma, including quarks \cite{5, 6, 7, 12}, mesons \cite{13, 14, 15, 16}, baryons \cite{17, 18, 19}, gluons \cite{17, 20} $k$-quarks \cite{17} and various types of defects \cite{21}. Energy loss studies have also been carried out via the gauge/gravity correspondence in many other theories, including some with properties more akin to those of QCD—see, e.g., \cite{22, 23} and references therein\cite{1}. Other examples of drag force calculations may be found in the reviews \cite{9}. While we are probably still far from achieving firm contact between experimental data and first-principles gauge/gravity calculations, the correspondence has already succeeded in enhancing our general intuition on the behavior of gauge theories at strong coupling, and has additionally provided useful suggestions for phenomenological models of the QGP (see, e.g., \cite{25, 26, 27, 28, 29}). In the present paper, we will review a number of recent results related to the evolution of a quark in a strongly-coupled gluonic plasma.

# Quarks, with Strings Attached

## 2.1 Basic setup

For our analysis we need two ingredients: a strongly-coupled thermal non-Abelian plasma, and a quark that traverses it. Via the gauge/gravity correspondence, a plasma in a $d$-dimensional CFT is known to be described in dual language by a Schwarzschild black hole (more precisely, black brane) in $(d + 1)$-dimensional asymptotically anti-de Sitter (AdS) space. For concreteness, we will frame our discussion in the context of $SU(N_c)$ MSYM in 3 + 1 dimensions, with coupling $g_{YM}$ and temperature $T$, which is equivalent to Type IIB string theory on the (AdS-Schwarzschild)$_5 \times S^5$ geometry

$$
\begin{align*}
\left(1\right) & \\
\text{ds}^2 & = g_{mn} dx^m dx^n = \frac{R^2}{z^2} \left(-h dt^2 + dz^2 + \frac{d z^2}{h}\right) + R^2 d\Omega_5, \\
\text{h} & = 1 - \frac{z^4}{z_h^4}, \\
\frac{R^4}{l_s^4} & = g_{YM}^2 N_c \equiv \lambda, \\
z_h & = \frac{1}{\pi T},
\end{align*}
$$

\footnote{See also \cite{24} for an interesting comparison between lattice results and its holographic counterpart.}
(with a constant dilaton and $N_c$ units of Ramond-Ramond five-form flux through the five-sphere), where $l_s$ denotes the string length, and $z_h$ is the location of the event horizon. The radial direction $z$ is mapped holographically into a variable length scale in the gauge theory, in such a way that $z \to 0$ and $z \to \infty$ are respectively the ultraviolet and infrared limits. The directions $x^\mu \equiv (t, \vec{x})$ are parallel to the AdS boundary $z = 0$ and are directly identified with the gauge theory directions. The five-sphere coordinates are associated with the global $SU(4)$ internal (R-) symmetry of MSYM. They will play no role in our discussion, so all results below will hold equally well in the more general case where the $S^5$ is replaced by a different compact five-dimensional space $X_5$, which corresponds to replacing $N = 4$ SYM with a different $(3+1)$-dimensional CFT.

Notice that the field theory temperature is identified with the Hawking temperature $T_H = 1/\pi z_h$ of the black hole. In the limit of vanishing temperature ($z_h \to \infty$), we are left in (1) with a pure AdS geometry, which is dual to the (symmetry-preserving) vacuum of MSYM. The closed string sector describing (small or large) fluctuations on top of AdS fully captures the nonperturbative gluonic (+ adjoint scalar and fermionic) physics. The generic (AdS-Schwarzschild)$_5 \times S^5$ geometry (1) is a special case of such a fluctuation, and is dual to a thermal ensemble in MSYM. The string theory description is under calculational control only for small string coupling and low curvatures, which translates into $N_c \gg 1$, $\lambda \gg 1$, i.e., a large number of colors and strong (‘t Hooft) coupling. It is also known that one can add $N_f$ flavors of matter in the fundamental representation of the $SU(N_c)$ gauge group by introducing in the string theory setup an open string sector associated with a stack of $N_f$ D7-branes [30]. We will refer to these degrees of freedom as ‘quarks,’ even though, being $N = 2$ supersymmetric, they include both spin 1/2 and spin 0 fields. For $\lambda N_f \ll N_c$, we are allowed to neglect the backreaction of the D7-branes on the geometry [22, 31]; in the gauge theory this corresponds to working in a ‘quenched’ approximation that ignores quark loops. These branes cover the four gauge theory directions $x^\mu$, and are spread along the radial AdS direction from $z = 0$ to $z = z_m$. The D7-brane parameter $z_m$ is related to the Lagrangian mass $m \gg \sqrt{\lambda} T$ of the quark through [31]

$$\frac{1}{z_m} = \frac{2\pi m}{\sqrt{\lambda}} \left[ 1 + \frac{1}{8} \left( \frac{\sqrt{\lambda} T}{2m} \right)^4 - \frac{5}{128} \left( \frac{\sqrt{\lambda} T}{2m} \right)^8 + O \left( \left( \frac{\sqrt{\lambda} T}{2m} \right)^{12} \right) \right],$$

which simplifies to

$$z_m = \frac{\sqrt{\lambda}}{2\pi m}$$

at vanishing temperature. For applications of this formalism to phenomenology, we must choose values of the mass parameter $z_m$ based on the charm and bottom quark masses, $m \simeq 1.4, 4.8$ GeV. Following [32], at the temperatures relevant to RHIC this translates into $z_m/z_h \sim 0.16 - 0.40$ for charm and $z_m/z_h \sim 0.046 - 0.11$ for bottom.
An isolated heavy quark corresponds to a string extending up from the horizon at $z = z_h$ to a location $z = z_m$ where it ends on the stack of $N_f$ D7-branes. The string dynamics is governed by the Nambu-Goto action

$$S_{\text{NG}} = -\frac{1}{2\pi l_s^2} \int d^2\sigma \sqrt{-\det g_{ab}} \equiv \int d^2\sigma L_{\text{NG}}, \tag{4}$$

where $g_{ab} \equiv \partial_a X^M \partial_b X^N G_{MN}(X) \ (a, b = 0, 1)$ denotes the induced metric on the worldsheet. We can exert an external force $\vec{F}$ on the string endpoint by turning on an electric field $F_{0i} = F_i$ on the D7-branes. This amounts to adding to (4) the usual minimal coupling, which in terms of the endpoint/quark worldline $x^\mu(\tau) \equiv X^\mu(\tau, z_m)$ reads

$$S_F = \int d\tau A_\mu(x(\tau)) \frac{dx^\mu(\tau)}{d\tau}. \tag{5}$$

Variation of $S_{\text{NG}} + S_F$ implies the standard Nambu-Goto equation of motion for all interior points of the string, plus the boundary condition

$$\Pi^z_\mu(\tau)|_{z=z_m} = F_\mu(\tau) \ \forall \ \tau, \tag{6}$$

where $\Pi^z_\mu \equiv \partial L_{\text{NG}}/\partial(\partial_\tau X^\mu)$ is the worldsheet current associated with spacetime momentum, and $F_\mu = F_{\mu\nu}\partial_\tau x^\nu = (-\gamma \vec{F} \cdot \vec{v}, \gamma \vec{F})$ the Lorentz four-force.

Notice that the string is being described (as is customary) in first-quantized language, and, as long as it is sufficiently heavy, we are allowed to treat it semiclassically. In gauge theory language, then, we are coupling a first-quantized quark to the gluonic (+ other MSYM) field(s), and then carrying out the full path integral over the strongly-coupled field(s) (the result of which is codified by the AdS spacetime), but, for the time being, treating the path integral over the quark trajectory $x^\mu(\tau)$ in a saddle-point approximation. (The effect of quantum fluctuations about the classical string configuration will be discussed in Sections 2.2 and 3.4.)

A crucial property of the gauge/gravity dictionary is that it identifies the endpoint(s) of the string as being dual to the quark (or antiquark), while the body of the string codifies the profile of the (near and radiation) gluonic (+ other MSYM) field(s) set up by the quark. In other words, this dictionary teaches us that the usual ‘QCD’ string exists even for non-confining theories, is infinitesimally thin, and actually lives in a curved 5 (+ 5)-dimensional geometry. The gluonic profile can be mapped out explicitly by computing one-point functions of local operators ($\langle \text{Tr} F^2(x) \rangle, \langle T_{\mu\nu}(x) \rangle, \ldots$) in the presence of the quark, which, via the GKPW recipe [33], requires a determination of the near-boundary profile of the closed string fields ($\phi(x, z), h_{\mu\nu}(x, z), \ldots$) generated by the macroscopic string.

It is important to keep in mind that the quark described by this string is not ‘bare’ but ‘composite’ or ‘dressed’. In an interacting field theory, a particle of finite mass can only be regarded as strictly pointlike at zeroth-order in a perturbative description. In the nonperturbative framework made available to us by the gauge/gravity duality, a quark with finite mass automatically acquires a finite size. This can be seen most
clearly by working out the vacuum expectation value of the gluonic field surrounding a static quark located at the origin \[34\]. For \( m \to \infty \) (\( z_m \to 0 \)), this is just the Coulombic field expected (by conformal invariance) for a pointlike charge. For finite \( m \), the profile is still Coulombic far away from the origin, but in fact becomes non-singular at the location of the quark. The characteristic thickness of the implied non-Abelian charge distribution is precisely the length scale \( z_m \) defined in \(2\). This is then the size of the gluonic cloud that surrounds the quark, or in other words, the analog of the Compton wavelength for our non-Abelian source.

Below we will also be interested in the description of a quark-antiquark pair. The IIB strings we have introduced above are oriented, and a state with two oppositely oriented purely radial strings would correspond to a quark and antiquark that are merely superposed. With such boundary conditions, however, the configuration with lowest energy is given by a single \( \cup \)-shaped string that has both of its endpoints at \( z = z_m \). This can again be verified by computing the corresponding gluonic profile in vacuum \[35,36\], to verify that the falloff is more rapid than Coulombic, as expected for an overall color-neutral source.

### 2.2 Late-time energy loss, Brownian motion and limiting velocity

Combining the two preceding ingredients, we know that a heavy quark in a strongly-coupled MSYM plasma is described by a string on the Schwarzschild-AdS geometry \(1\). As the string endpoint moves, its body lags behind it, exerting a drag on the tip that is the gravity-side realization of the damping force exerted by the plasma on the quark, as studied in \[5,6\]. For a restricted type of quark trajectories, this mechanism leads to energy loss at a rate

\[
\frac{dE_{\text{rad}}}{dt} = \frac{\pi}{2} \sqrt{\lambda T^2} \frac{v^2}{\sqrt{1 - v^2}}.
\]

\(7\)

Various works have explored the way in which these heavy quark results are modified for heavy sources of the gluonic field in color representations other than the fundamental (including the adjoint) \[17\], as well as for light quarks and gluons \[37,38\].

In the context of weakly-coupled QCD, it is known that energy loss at high parton velocity is dominated by the strong-interaction analog of bremsstrahlung, i.e., medium-induced gluonic radiation, while for low velocities collisional loss becomes important \[2\]. In the strongly-coupled MSYM setup made available to us by AdS/CFT, the perturbative terminology is no longer adequate. The flow of energy/momentum along the body of the trailing string corresponds to energy/momentum transported away from the quark via the gluonic field generated jointly by the quark and the plasma, in a pattern that has been meticulously studied in a large body of work that began with \[39\], has been reviewed in \[40\] and includes the recent additions \[41,42,43,44,45,46\]. For ease of language, throughout this paper we will continue to speak of radiation, even though this concept is not really appropriate within an infinite thermal medium.
We should stress that, within the framework of [5, 6], adopted also in this paper, the entire energy loss calculation is carried out inside the MSYM theory, whose coupling does not run and is taken to be large. For quarks that are extremely energetic, the asymptotic freedom of QCD would lead one to expect deviations between this scenario and the real-world QGP, and seek a hybrid approach describing hard/perturbative emission of a gluon that subsequently propagates through a strongly-coupled medium, as advocated in [8, 47]. It is debatable, however, which of these two scenarios is more appropriate at the not extremely relativistic heavy quark energies achieved at RHIC. At the very least, the calculations discussed in this paper have a direct interpretation in terms of energy loss in a thermal plasma of strongly-coupled gluons and exotic matter (‘XGP’) of MSYM (or other CFTs), which is an interesting theoretical question in its own right.

In [5, 6], the rate (7) was deduced for an isolated quark in the stationary or late-time regimes, i.e., assuming that the quark either moves at constant velocity, under the influence of an external force which precisely cancels the drag force exerted by the plasma, or has decelerated solely under the action of this drag force for a long period of time, and is about to come to rest. In the actual experimental setup, however, the configuration is neither stationary nor asymptotic: the quark is not externally forced, and slows down under the influence of the plasma, whose finite spatial and temporal extent imply that the late-time regime is not generically accessible. Moreover, the real-world quark is not isolated, but is created within the plasma together with an antiquark. In Section 4 we will explore to what extent the rate of energy loss is affected by these issues. Also, as we will seen in Section 3, the above trailing string mechanism is in fact much more general, and accounts even for the radiative damping expected in vacuum [54, 55]. That is, irrespective of whether a spacetime black hole is present or not, the body of the string plays the role of an energy sink, as befits its identification as the embodiment of the gluonic degrees of freedom. On the other hand, energy loss via the string does turn out to be closely associated with the appearance of a worldsheet horizon, as noticed initially in [18, 19] at finite temperature and in [56] (see also [57, 58]) for the zero temperature case.

The appearance of a black hole on the string worldsheet is crucial in particular to reproduce the expected Brownian motion of the dual quark in the hot medium. This connection was first worked out in detail in [59, 60], for a static string on the Schwarzschild-AdS_{d+1} geometry, which as we know from Section 2.1 is dual to a static quark in a thermal bath of the CFT. At finite \( \lambda \), one must take into account small perturbations about the average string embedding. These are described by

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2The approach of [8, 47] involves the extraction from a lightlike Wilson loop of a transport (‘jet-quenching’) parameter \( \hat{q} \), which characterizes the medium and whose value at strong coupling curiously differs from the customary definition of \( \hat{q} \) as the average transverse momentum picked up per unit distance traveled [18, 19]. Some doubts about the procedure proposed in [8, 47] to calculate \( \hat{q} \) on the gravity side of the correspondence had been raised by, e.g., [15, 50, 51], but these concerns appear to have been resolved by a recent correction to the procedure [52], which happens to leave the result unchanged.

3This latter point has been emphasized from the phenomenological perspective in [53].
free scalar fields propagating on the worldsheet black hole geometry, and upon quantization, they are excited by Hawking radiation. This leads to Brownian motion of the endpoint/quark, whose detailed form is captured by a generalized Langevin equation. The authors of [59] reached this conclusion in arbitrary dimension by assuming that the state of the quantized embedding fields is the usual Hartle-Hawking (or Kruskal) vacuum, which describes the black hole in equilibrium with its own thermal radiation. The authors of [60] focused on the case $d = 4$ and followed a different but equivalent route, employing the dual relation between the Kruskal extension of the Schwarzschild-AdS geometry and the CFT Schwinger-Keldysh formalism [61, 62, 67], together with the known connection between the latter and the generalized Langevin equation. These calculations were later generalized and elaborated on in [63, 64, 65, 66].

Another very interesting prediction of AdS/CFT in this thermal context is the existence of a subluminal limiting velocity

$$v_m \equiv \sqrt{1 - \frac{z_m^4}{z_h^4}} \simeq 1 - \left( \sqrt{\lambda} \frac{T}{2m} \right)^4$$

for the quark traversing the plasma. This follows simply from the fact that the quark velocity is dual to the coordinate velocity of the string endpoint, and $v = v_m$ corresponds to a proper velocity at $z = z_m$ equal to that of light [67]. A more general bound involving the external force $\vec{F}$ can be derived by requiring the Nambu-Goto square root to remain real [56]. This same restriction on the velocity can be seen to arise from microscopic calculations of the meson spectrum [68, 69]. Its validity for isolated quarks has been emphasized in [51, 56]. The existence of this limiting velocity might have interesting phenomenological consequences, such as the photon peak predicted in [27] or the Cherenkov emission of mesons analyzed in [28].

3 Isolated Quark at Zero Temperature

3.1 Energy loss in vacuum

To orient ourselves, we will first inquire into the rate of energy loss for an accelerating quark in vacuum, whose dual description involves the string moving on pure AdS spacetime. In this case we will have full analytic control over the system, which will allow us to develop some intuition on the problem at hand. Besides being of theoretical interest in its own right, our vacuum analysis will provide a useful benchmark against which the finite-temperature results can be compared.

A quark that accelerates in vacuum would be expected to emit chromoelectromagnetic radiation. The first definite characterization of the radiation rate off an accelerating quark by means of AdS/CFT was worked out in an important paper by Mikhailov [70]. Remarkably, this author was able to solve the highly nonlinear equation of motion for a string on AdS$_5$ that follows from [4], for an arbitrary timelike
trajectory of the string endpoint dual to an infinitely massive quark. In terms of the coordinates used in (1) (where for now $h = 1$), his solution is

$$X^\mu(\tau, z) = z \frac{dx^\mu(\tau)}{d\tau} + x^\mu(\tau),$$

(9)

with $\mu = 0, 1, 2, 3$, and $x^\mu(\tau)$ the worldline of the string endpoint at the AdS boundary—or, equivalently, the worldline of the dual quark—parametrized by the proper time $\tau$ defined through $\eta_{\mu\nu}(dx^\mu/d\tau)(dx^\nu/d\tau) = -1$. From the structure of (9) we see that the behavior of the string segment located at a given time $t$ and radial depth $z$ (which according to our discussion in the previous section codifies the behavior of the gluonic field at the length scale $z$) is completely determined by the behavior of the quark/string endpoint at the earlier, retarded time $t_{\text{ret}}(t, z)$ defined by

$$t_{\text{ret}} = z \frac{1}{\sqrt{1 - \tilde{v}(t_{\text{ret}})^2}} + t_{\text{ret}},$$

(10)

where the quark worldline has been parametrized by $x^0(\tau)$ instead of $\tau$, and $\tilde{v} \equiv d\vec{x}/dx^0$. The definition (10) can be shown to imply that $t_{\text{ret}}$ is obtained by projecting back to the AdS boundary along a (fixed $\tau$) curve that is null on the string worldsheet, in analogy with the Lienard-Wiechert story in classical electrodynamics. The embedding (9) thus describes a wave on the string that is purely outgoing: it is generated at the string endpoint and then descends along the body of the string, moving into the AdS bulk. This is why, among the infinite number of extremal string embeddings that satisfy the boundary condition associated with the given quark trajectory, the profile (9) is the one that is of physical interest for us: it is dual to a configuration where waves in the gluonic field move out from the quark to infinity.

Working in the static gauge $\sigma^0 = t$, $\sigma^1 = z$, Mikhailov was able to reexpress the total energy of the string embedding (9) (via a change of integration variable $z \rightarrow t_{\text{ret}}$) as a local functional of the quark trajectory,

$$E(t) = \frac{\sqrt{\lambda}}{2\pi} \int_{-\infty}^{\infty} dt_{\text{ret}} \frac{\vec{a}^2 - [\vec{v} \times \vec{a}]^2}{(1 - \tilde{v}^2)^3} + E_q(\vec{v}(t)),
$$

(11)

where of course $\vec{a} \equiv d\vec{v}/dx^0$. The second term in the above equation arises from a total derivative that was not explicitly written down by Mikhailov, but can easily be worked out to be

$$E_q(\vec{v}) = \frac{\sqrt{\lambda}}{2\pi} \left( \frac{1}{\sqrt{1 - \tilde{v}^2 z}} \right)_{z=0}^{z=m} \gamma m,
$$

(12)

which gives the expected Lorentz-invariant dispersion relation for the quark. The energy split achieved in (11) therefore admits a clear and pleasant physical interpretation: $E_q$ (associated only with information of the string endpoint) is the intrinsic energy of the quark at time $t$, and the integral over $t_{\text{ret}}$ (associated with the body of the string) encodes the accumulated energy lost by the quark to its gluonic field over all times prior to $t$. Completely analogous statements can be derived for the spatial

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4 This energy split has been extended to $k$-quarks in the recent work [72].
momentum. No less remarkable is the fact that the rate of energy loss for the quark in this strongly-coupled non-Abelian theory is found to be in precise agreement with the standard Lienard formula from classical electrodynamics! The AdS/CFT correspondence thus teaches us that, in this very unfamiliar nonlinear setting, the energy loss turns out to depend only locally on the quark worldline. This feature has been argued in \[29\] to lead to an upper bound for the energy of a quark at finite temperature. In the next subsection we will see that both the rate of radiation and the dispersion relation are modified when the quark has a finite mass, and is therefore not pointlike.

### 3.2 Radiation damping

The radiation emitted by an accelerated charge inevitably backreacts on the charge. One effect, present already at the classical level, is a reaction force on the charge, that tends to damp its motion. And if the system is quantized, one additionally expects the emission of radiation to induce stochastic fluctuations of the charge’s trajectory. The discovery of the gauge/gravity duality has opened the possibility of exploring these effects in the previously uncharted terrain of strongly-coupled non-Abelian gauge theories. We will discuss here radiation damping, and postpone the analysis of fluctuations until Section 3.4.

In the context of classical electrodynamics, and for an electron modeled as a vanishingly small charge distribution, the effect of radiation damping is incorporated in the classic (Abraham-)Lorentz-Dirac equation \[73\],

\[
m\left( \frac{d^2 x^\mu}{d\tau^2} - t_e \left[ \frac{d^3 x^\mu}{d\tau^3} - \frac{1}{c^2} \frac{d^2 x^\nu}{d\tau^2} \frac{dx^{\mu\nu}}{d\tau} \right] \right) = F^\mu, \tag{13}
\]

with $\tau$ the proper time, $F^\mu \equiv \gamma(\vec{F} \cdot \vec{v}/c, \vec{F})$ the external 4-force, and $t_e \equiv 2e^2/3mc^3$ a characteristic timescale set by the classical electron radius. The second term within the square brackets is the negative of the rate at which 4-momentum is carried away from the charge by radiation (as given by the covariant Lienard formula), so it is only this term that can properly be called radiation reaction. The first term within the square brackets, usually called the Schott term, is known to arise from the effect of the charge’s ‘near’ (as opposed to radiation) field \[74, 75\]. The appearance of a third-order term in \[13\] leads to unphysical behavior, including pre-accelerating and self-accelerating (or ‘runaway’) solutions. These deficiencies are known to originate from the assumption that the charge is pointlike.

Analysis of radiation damping in the quantum non-Abelian case using traditional methods would be a serious challenge, but the AdS/CFT correspondence allows us to address it rather easily in certain strongly-coupled non-Abelian gauge theories \[54, 55\]. By Lorentz invariance, given the description of the static quark, we know that a quark moving at constant velocity corresponds to a purely radial string, moving as a rigid vertical rod. When the quark accelerates, the body of the string will trail behind the endpoint, and will therefore exert a force on the latter. Remembering that the body of the string codifies the SYM fields sourced by the quark, we know
that in the gauge theory this force is interpreted as the backreaction of the gluonic field on the quark. In other words, in the AdS/CFT context the quark always has a ‘tail’, and it is this tail that is responsible for the damping effect we are after. This is of course the same mechanism that yields the drag force exerted on the quark by a thermal plasma \cite{5, 6}, reviewed in the previous section. The analysis of \cite{54, 55} makes it clear that the damping effect is equally present in the gauge theory vacuum.

To obtain a noticeable damping effect, we must extend the analysis of the preceding subsection to the more interesting case with \( z_m > 0 \), where the quark has a finite mass. As we have emphasized in Section 2.1, in this case our non-Abelian source is no longer pointlike but has size \( z_m \). On the string theory side, the string endpoint is now at \( z = z_m \), and we must again require it to follow the given quark trajectory, \( x^\mu(\tau) \). As before, this condition by itself does not pick out a unique string embedding. Just like we discussed for the infinitely massive case, we additionally require the solution to be retarded, in order to focus on the gluonic field causally set up by the quark. As in \cite{56}, we can inherit this structure by truncating a suitably selected retarded Mikhailov solution. The embeddings of interest to us can thus be regarded as the \( z \geq z_m \) portions of the solutions \cite{2}. In \cite{54, 55} it was shown that, with this information, the standard boundary condition \cite{6} for the string endpoint can be rewritten in the form of an equation of motion for the dressed quark,

\[
\frac{d}{d\tau} \left( m \frac{dx^\mu}{d\tau} - \frac{\sqrt{\lambda}}{2\pi m^2} F^\mu \right) = F^\mu - \frac{\sqrt{\lambda}}{2\pi m^2} F^2 \frac{dx^\mu}{d\tau}.
\]  

(14)

Notice that the characteristic length scale appearing in (14) is precisely \( z_m = \sqrt{\lambda}/2\pi m \), which as previously discussed, plays the role of the quark Compton wavelength. Let us now examine the behavior of a quark that is sufficiently heavy, or is forced sufficiently softly, that the condition \( \sqrt{\lambda}|F^2|/2\pi m^2 \ll 1 \) holds. It is then natural to expand the equation of motion in a power series in this small parameter. To zeroth order in this expansion, we correctly reproduce the pointlike result \( m \frac{d^2 x^\mu}{d\tau^2} = F^\mu \). If we instead keep terms up to first order, we find

\[
m \frac{d}{d\tau} \left( \frac{d x^\mu}{d\tau} - \frac{\sqrt{\lambda}}{2\pi m^2} F^\mu \right) \simeq F^\mu - \frac{\sqrt{\lambda}}{2\pi m^2} F^2 \frac{dx^\mu}{d\tau}.
\]

(15)

In the \( \mathcal{O}(\sqrt{\lambda}) \) terms it is consistent, to this order, to replace \( F^\mu \) with its zeroth order value, thereby obtaining

\[
m \left( \frac{d^2 x^\mu}{d\tau^2} - \frac{\sqrt{\lambda}}{2\pi m} d^3 x^\mu \right) \simeq F^\mu - \frac{\sqrt{\lambda}}{2\pi m^2} F^2 \frac{dx^\mu}{d\tau}.
\]

Interestingly, this coincides exactly with the Lorentz-Dirac equation (13), with the Compton wavelength \cite{2} playing the role of characteristic size \( t_c \) for the composite quark. This is indeed the natural quantum scale of the problem. The radiation
reaction force in (15) is correctly given by the covariant Lienard formu- 
la, as expected from the result (11) [70], which we see then arising as the pointlike limit of the full radiation rate encoded in the right-hand side of (14). The Schott term in (15) (associated with the near field of the quark), originated from the terms inside the }\tau}\text{-derivative in the left-hand side of (14), which we understand then to codify a modified dispersion relation for our composite quark.

To second order in \(\sqrt{\lambda}|F|^2/2\pi m^2\), we similarly obtain

\[
\begin{align*}
  m\frac{d^2 x^\mu}{d\tau^2} - &\, \frac{\sqrt{\lambda}}{2\pi} \left( \frac{d^3 x^\mu}{d\tau^3} - \frac{d^2 x^\nu d^2 x_\nu}{d\tau^2 d\tau} \right) \\
  + &\, \frac{\lambda}{4\pi^2 m} \left( \frac{d^4 x^\mu}{d\tau^4} - (1+2) \frac{d^2 x^\nu d^3 x_\nu}{d\tau^2 d\tau^3} \right) \\
  - &\, \frac{\lambda}{4\pi^2 m} \left( \frac{1}{2} + \frac{1}{1} \right) \frac{d^2 x^\nu d^2 x_\nu d^2 x^\mu}{d\tau^2 d\tau^2 d\tau} \simeq F^\mu.
\end{align*}
\]

(16)

For compactness, we have grouped together all terms arising at the same order in the expansion, and used underbraces to mark the radiation reaction terms originating from the right-hand side of (14) (which now include corrections beyond the standard Lienard formula), and overbraces to indicate the near-field terms arising from the left-hand side of (14) (which incorporate corrections to the standard Schott term).

One can continue this expansion procedure to arbitrarily high order in the ratio \(\sqrt{\lambda}|F|^2/2\pi m^2\). The full equation (14) is thus recognized as a compact (reduced-order) rewriting of an infinite-derivative extension of the Lorentz-Dirac equation that automatically incorporates the size \(z_m\) of our non-pointlike non-Abelian source. It is curious to note that (14), which clearly incorporates the effect of radiation damping on the quark, has been obtained from (9), which does not include such damping for the string itself. The latter arises from the backreaction of the closed string fields set up by our macroscopic string, but these are of order \(1/N_c^2\), and therefore subleading at large \(N_c\) (though not necessarily at the real-world value \(N_c = 3\) [76]).

The full physical content of (14) can be made transparent by rewriting it in the form

\[
\frac{dP^\mu}{d\tau} \equiv \frac{dp^\mu_q}{d\tau} + \frac{dP^\mu_{\text{rad}}}{d\tau} = F^\mu,
\]

(17)

recognizing \(P^\mu\) as the total string (= quark + radiation) four-momentum,

\[
p^\mu_q = \frac{m dx^\mu}{d\tau} - \frac{\sqrt{\lambda}}{2\pi m} \frac{F^\mu}{\sqrt{1 - \frac{\lambda}{4\pi^2 m^4} |F|^2}}
\]

(18)

as the intrinsic momentum of the quark including the contribution of the near-field sourced by it (or, in quantum-mechanical language, of the gluonic cloud surrounding
the quark), and
\[
\frac{dP^\mu_{\text{rad}}}{d\tau} = \frac{\sqrt{\lambda} F^2}{2\pi m^2} \left( \frac{dx^\mu}{d\tau} - \frac{\sqrt{\lambda}}{2\pi m^2} F^\mu \right) (1 - \frac{\lambda}{4\pi^2 m^2} F^2) \tag{19}
\]
as the rate at which momentum is carried away from the quark by chromo-electromagnetic radiation.

Unlike its classical electrodynamic counterpart (13), the dressed quark equation of motion (14) has no self-accelerating (runaway) solutions: in the (continuous) absence of an external force, it uniquely predicts that the 4-acceleration of the quark must vanish. Interestingly, the converse to this last statement is not true: constant 4-velocity does not uniquely imply a vanishing force.

It is natural to expect the energy split achieved in (11) or (17) to be somehow reflected in the geometry of the string worldsheet. It is interesting then that, on the string worldsheet dual to a quark undergoing arbitrary accelerated motion, there appears a black hole [56], whose event horizon naturally divides the worldsheet into two causally disconnected regions. The appearance of a worldsheet black hole had also been noted previously at finite temperature [48, 49], so the overall lesson is that such causal structure is intrinsically tied to energy dissipation, be it within a thermal plasma or in vacuum. It seems plausible then to interpret the regions outside and inside the black hole as corresponding respectively to the quark and the gluonic field, and postulate that the energy flow across this horizon should be related to the energy radiated by the quark [57, 56]. This possibility has been studied more closely in [58, 77]. Unfortunately, this interpretation is not always correct: for general quark trajectories, one can show that the energy of the portions of the string outside and inside the dynamical horizon does not actually match the intrinsic and radiated energy of the quark [78].

3.3 Absence of broadening in the radiation pattern

In the previous two subsections we have seen how a study of the string embedding yields important quantitative information on the dual quark dynamics, including its rate of energy loss, based on the natural split achieved in [70] of the total energy of the string. The latter is conserved on the fixed AdS (or AdS-Schwarzschild) background, but would of course decrease steadily if we take into account the gravitational (and dilatonic, etc.) radiation given off by the string in the course of its evolution. As we have noted, from the gravity perspective this radiation is suppressed by a factor of \(1/N_c^2\), so it is consistent to neglect it at large \(N_c\). On the other hand, through the GKPW recipe for correlation functions [33], it is this radiation (or, more precisely, the full metric perturbation produced by the string), evaluated at the AdS boundary, that determines the expectation value of the gauge theory energy-momentum tensor. This tensor contains not just the gross information about the total energy loss rate, but also the fine details about the directionality of the flow and the relative weight of the various dissipation channels. The problem of determining the spacetime profile of the gauge field disturbance produced by an arbitrary charge trajectory was solved
long ago by Lienard and Wiechert for classical electrodynamics, but, under various guises, remains of great interest today in the context of quantum non-Abelian gauge theories. It is therefore significant that the gauge/gravity duality has given us a useful handle on this problem for a varied class of gauge theories in the previously inaccessible regime of strong coupling.

The translation between string and gauge theory disturbances was first explored in [35], which studied the dilaton waves given off by small fluctuations on an otherwise static, radial string in AdS$_5$, and inferred from them the profile of the dual gluonic field observable $\langle \text{Tr} F^2(x) \rangle$ in the presence of an oscillating quark in vacuum. The MSYM waves were found to display significant temporal broadening, just as one would expect given that points on the non-Abelian field arbitrarily far from the quark can themselves reradiate. This feature emerges naturally in the gravity side of the correspondence, because motion of the string endpoint generates waves that move up along the body of the string, and each point on the string then emits a dilaton wave that travels back down to the observation point $x$ on the AdS boundary, where, via the AdS/CFT recipe for correlation functions [33], the value of $\langle \text{Tr} F^2(x) \rangle$ is deduced by assembling together all such contributions, each with a different time delay. While informative, the results of [35] did not allow a definite identification of waves with the characteristic $1/|\vec{x}|^2$ falloff associated with radiation, i.e., contributions that transport energy to infinity. The unambiguous detection of the latter calls for examination of the MSYM energy-momentum tensor, $\langle T_{\mu\nu}(x) \rangle$, which in the gravity side of the correspondence requires a determination of the gravitational waves emitted by the string [79].

In more recent years, motivated by potential contact with the phenomenology of the quark-gluon plasma [1], analyses of both $\langle \text{Tr} F^2(x) \rangle$ and $\langle T_{\mu\nu}(x) \rangle$ have been carried out in the case of a heavy quark moving at constant velocity through a thermal plasma (the only finite-temperature case where an exact solution for the corresponding string embedding is available [5, 6]), in a large body of work that includes [39] and has been reviewed in [40]. The results are again compatible with the expected nonlinear dynamics of the (in this case, finite-temperature) MSYM medium.

Given these antecedents, it came as a surprise when, back at zero temperature, additional calculations going beyond the linearized string approximation found, first for special cases [41, 42] and then for an arbitrary quark trajectory [44] (using the string embedding [9]), that the ensuing energy density $\langle T_{00}(x) \rangle$ displays no temporal broadening, and is in fact as sharply localized in spacetime as the corresponding classical profile.

The apparent tension between the results of [35] and [42, 44] for the gluonic fields in vacuum was resolved recently in [45], which carried out the $\langle \text{Tr} F^2(x) \rangle$ calculation beyond the linearized string approximation, and for an arbitrary quark trajectory, to put it on a par with the computation in [44]. The results show that the gluonic near field at any given observation point depends only on dynamical data evaluated at a single retarded time along the quark trajectory (with luminal and subluminal propagation in the cases of infinite and finite quark mass, respectively). This proves
that, despite appearances, there is in fact no conflict between \cite{35} and \cite{42, 44}. While it is true that the gluonic field emerges as a superposition of contributions with all possible time delays (as shown explicitly by the calculations in \cite{41, 42, 44, 45}, and also by a recent reformulation of these as a superposition of gravitational shock waves emitted by each point along the string \cite{43}), it turns out that the net result evidences only the smallest of these delays. So, contrary to \cite{42, 44}, where it was suggested that the supergravity approximation is somehow leaving out the expected non-Abelian rescattering, in \cite{45} the no-broadening result was interpreted as a surprising prediction of the AdS/CFT correspondence for the net pattern of propagation in the CFT at strong coupling and with a large number of colors, under the assumption of a purely outgoing condition for the gluonic field generated by the quark.

We should perhaps stress that this prediction refers only to propagation in the CFT vacuum. When a beam of radiation enters a plasma, it is expected to remain unbroadened only at distances that are small compared to the thermal wavelength, after which it will diffuse and eventually thermalize, consistent with the results of \cite{39}.

### 3.4 Fluctuations and the Unruh effect

When going beyond the classical description of the string, two new effects are found, both of which are suppressed by a factor of the string length divided by the AdS curvature radius, or, equivalently (via (1)), by an inverse factor of the CFT coupling $\lambda$. On the one hand, we pick up the usual quantum fluctuations arising from the determinant of the path integral over string embeddings. These are present even for a static string (see, e.g., \cite{80}), and lead to spontaneous deviations from the average endpoint/quark trajectory of the type studied, e.g., in \cite{81}. On the other hand, the worldsheet black hole emits Hawking radiation, which populates the various modes of oscillation of the string.

This second effect, which is present only for an accelerated trajectory and is thus associated with the quantum fluctuations induced by the gluonic radiation emitted by the quark, has been studied in \cite{52}, for the special case of uniform proper acceleration $A$. In this case, the general solution \cite{59} takes a form that was found independently in \cite{58, 83}. The geometry induced on the worldsheet contains a black hole that is static, and thus amenable to explicit calculation. The analysis of the effect of Hawking radiation is in complete parallel with the story of Brownian motion \cite{59, 60} recalled in Section 2.2, except that it is complicated by the explicit time dependence present in the worldsheet geometry when presented in inertial coordinates.

The problem simplifies if instead of working in the coordinates appropriate for an inertial observer we transform to a Rindler coordinate system adapted to an observer sitting on the quark. This transformation gives rise to an acceleration horizon both in the boundary and bulk descriptions. As a result, a state that is pure from the inertial perspective will generally be mixed from the point of view of the Rindler observers, because the field degrees of freedom accessible to the latter will be entangled with
degrees of freedom in the region beyond their horizon, which they must trace over. In particular, the pure AdS geometry expressed in Rindler coordinates, which is dual to the Minkowski vacuum of the CFT (as evidenced by the vanishing of the expectation value of the stress-energy tensor), is interpreted as a thermal bath at the temperature $T_U = A/2\pi$. Contact is thus made with the celebrated Unruh effect \cite{84}, i.e., the fact that, from the point of view of a uniformly accelerated observer, the Minkowski vacuum behaves as a thermal medium. In Rindler coordinates, the string embedding is static and bends towards the Rindler horizon. It is well-known that the Rindler horizon of the CFT can be removed via a Weyl transformation, leading to the open Einstein universe. The corresponding bulk transformation drastically alters the radial foliation of the AdS geometry. Even though, by construction, in this new conformal frame the acceleration horizon is no longer visible in the boundary description, one can show that it is still present in the bulk, but lies at the fixed radial position that according to the AdS/CFT dictionary corresponds to the Unruh temperature. In other words, after this second transformation, the thermal character of the CFT state arises not from entanglement with degrees of freedom that lie beyond a spacetime horizon, but from the direct identification of the specific energy scale as the temperature of the CFT, in exact parallel with the dual interpretation of the Schwarzschild-AdS geometry. This exercise thus sheds light on the AdS implementation of the Unruh effect. (Related analyses can be found in \cite{58,83,85}.)

After the bulk diffeomorphism that removes the CFT horizon, the string embedding is not only static but also completely vertical. Interestingly, both the base string embedding and the background geometry at the location of the string are found to coincide exactly with the $d = 2$ thermal setup analyzed in \cite{59}, which allows one to obtain the radiation-induced fluctuations of the quark simply by translating the results of that work. This close relation between the quantum fluctuations of the uniformly accelerated quark on Minkowski spacetime and the thermal fluctuations of a static quark in a thermal medium is evidently a direct consequence of the Unruh effect, but the reader should be aware that, for $d > 2$, the detailed properties of this thermal medium are found to differ from those of the familiar homogeneous and isotropic thermal ensemble dual to the Schwarzschild-AdS$_{d+1}$ geometry. Back in the original inertial frame, the quark fluctuations are governed by Langevin type equations \cite{82}.

## 4 Isolated Quark at Finite Temperature

### 4.1 Constant velocity

Having understood the rate of energy loss for a heavy quark that moves in the MSYM vacuum, in this section we restore $z_h < \infty$— and consequently $h < 1$— in the metric \eqref{1}, to study the same quantity in the case where the quark moves through a thermal plasma. A thorough generalization of Mikhailov’s analytic results \cite{70} to this finite temperature setup would require finding the exact solution to the Nambu-
Goto equation of motion for the string on the AdS-Schwarzschild background, for any given trajectory of the string endpoint at \( z_m \geq 0 \). Sadly, this has not yet been accomplished. Nevertheless, based on the results discussed in the previous section, we expect the total energy of the string at any given time to again decompose into a surface term that encodes the intrinsic energy of the quark and an integrated local term that reflects the energy lost by the quark to the thermal medium.

There are two easy cases where one can show analytically that this expectation is borne out [56]: the static quark, where there is of course no energy loss, and the quark moving at constant velocity. An important difference with respect to the \( T = 0 \) case analyzed in the previous section is that here the surface contribution arises not only from the lower (\( z = z_m \)) but also from the upper (\( z = z_h \)) endpoint of the string. Since \( z_h \) marks the position of an event horizon, for any finite coordinate time \( t \) the value of the surface contribution at \( z = z_h \) is not influenced by the behavior of the \( z < z_h \) portion of the string, but depends only on the string’s configuration at \( t \rightarrow -\infty \).

The same interpretation can be then carried over to the gauge theory: for arbitrary finite-temperature configurations, the terms in the quark dispersion relation arising from the upper string endpoint will encode a contribution to the energy of the state that depends solely on the initial configuration of the quark+plasma system.

### 4.2 Early-time energy loss

As we mentioned in Section 2.2, expression (7) was derived under a number of simplifying assumptions. The one that matters most for our purposes is the restriction made by the authors of [5, 6] to a configuration where the quark is either moving with constant (possibly relativistic) velocity as a result of being pulled by an external force that precisely balances the drag, or is unforced but moving nonrelativistically and about to come to rest. Either of these scenarios requires the interaction between the quark and the medium to occur over a considerable period of time. The actual energy loss might thus be expected to differ from (7) in a situation where the quark moving through the plasma is accelerating, or in the initial period following its production within the thermal medium. This point has been emphasized from the phenomenological perspective in [53]. The estimates there are based on perturbative calculations, so it is interesting to inquire into this issue in the strongly-coupled systems available to us through the AdS/CFT correspondence.

Now, the restriction in [5, 6] to the stationary or asymptotic cases was of course implemented to gain analytic control on the problem of energy loss. The authors of [77] were able to slightly extend the stationary result to the case of a slowly decelerating quark, working in an expansion in powers of \( \sqrt{\lambda T/m} \). Away from these regimes, it is difficult to follow the evolution of the quark in the thermal plasma, or equivalently, of the string on the AdS black hole geometry.

The first gauge/gravity exploration of early-time energy loss was carried out in [56], by numerically integrating the Nambu-Goto equation of motion for the string dual to a quark initially at rest, which is accelerated along one dimension by a time-
dependent external force that is turned off after a short interval, allowing the quark to move thereafter only under the influence of the plasma. The results show a qualitative difference between the initial stage ($0 \leq t < t_{\text{release}}$) where the quark is accelerated by means of the external force $F(t)$, and the second stage ($t_{\text{release}} \leq t < t_{\text{breakdown}}$) where it moves only under the influence of the plasma. In the former stage, the rate of energy loss, for values of $m$ in the neighborhood of the charm or bottom masses, is in fact nearly identical to the corresponding rate in vacuum, given by the modified Lienard formula (19). In the latter stage, which would appear to be more relevant from the phenomenological perspective, the quark was found to dissipate energy at a rate much lower than the late-time result (7). Indeed, for the values $z_m/z_h = 0.2, 0.3, 0.4$ that (as noted below (3)) are appropriate for the charm and bottom quarks, the late-time friction coefficient $\mu \equiv - (1/p_q)(dp_q/dt)$ is respectively $\mu_{\text{late}}/\pi T = 0.25, 0.41, 0.59$, but the numerical results for $v(t)$ at $t_{\text{release}} \leq t < t_{\text{breakdown}}$ are best approximated by $\mu_{\text{early}}/\pi T = 0.08, 0.15, 0.26$. A more detailed discussion of the energy lost by the quark in both stages was given in [56], and a similar comparison of vacuum versus thermal energy loss rates was performed later in [86], for the special case of a quark undergoing uniform circular motion.

Additional control on the problem of early-time quark damping in a strongly-coupled plasma was gained recently in [87], which studied thermal effects analytically, as a small perturbation on the zero-temperature evolution. This is reasonable on general physical grounds during the initial stage of motion through the plasma, and is moreover supported by the numerical results of [56]. Interestingly, this framework naturally makes contact with the previously known limiting velocity $v_m$ (defined in (8)) for a quark immersed in the strongly-coupled plasma. The results of [87] identify

$$E_q = \frac{mv_m}{\sqrt{v_m^2 - v^2}} - \frac{1}{2} \sqrt{\lambda T}$$

(20)

as the thermally corrected intrinsic energy of the quark, and

$$\frac{dE_{\text{rad}}}{dt} = \frac{\pi^4 T^4 z_m^3 m v_m v}{2(v_m^2 - v^2)} = \frac{\pi \lambda^{3/2} T^4 v_m v}{16m^2(v_m^2 - v^2)}$$

(21)

as its rate of energy loss.

For parameter values appropriate to RHIC, this early-time rate can be up to six times smaller than the late-time result (7). This comparison is consistent with the (more limited) numerical results of [56] described above. The functional form of the results is also of interest: from the derivation in [87] it becomes clear that the $T^4$-dependence seen in (21) (and anticipated in an estimate based on saturation physics [57]) follows naturally from the first thermal corrections in the metric, and via dimensional analysis, this fixes the accompanying dependence on the quark mass, which in turn explains the peculiar power of the 't Hooft coupling. Another prominent feature of this formula is its velocity-dependence: whereas the late-time friction coefficient deduced from (7) is momentum-independent, its early-time counterpart depends
strongly on momentum. This feature should constitute an interesting experimental signature when used as input for phenomenological models, as in [25, 29].

The rate (21) is limited to the initial stage where the quark is only starting to feel the effects of the plasma, and should be reliable for times that are not too close to the lifetime of the real-world QGP. This suggests that the experimental results should effectively arise from some sort of average or interpolation between the early and late energy loss rates (21) and (7). For times of order \(1/\pi T \sqrt{v_m^2 - v_0^2}\), one expects the string embedding to be significantly modified by the appearance of \(T^8\) and higher corrections neglected in the linearized analysis of [87]. As the string continues to evolve, a dynamical worldsheet horizon (located at \(z = z_h\) in the remote past) is expected to move up toward smaller values of \(z\), and then (if the quark is undisturbed) move back down towards the spacetime horizon [56]. When the wavefront on the string crosses the worldsheet horizon, one would have to deviate from the purely outgoing condition used in [87], to accommodate the fact that, as the medium is disturbed by the quark, it is expected to radiate gluonic waves back towards the source. The net result should be to yield a damping coefficient \(\mu(t)\) that smoothly interpolates between the early-time result and the late-time form.

5 Pair Creation within the Plasma

5.1 Quark-antiquark evolution

As was mentioned in Section 2.2 in the experimental setup the heavy quark is created within the plasma together with its corresponding antiquark, and the presence of the latter would be expected to substantially modify the gluonic fields in the vicinity of the quark, consequently affecting its evolution. The evolution of a heavy quark and antiquark that are created within the plasma and then separate back to back was first studied in [5], and explored in greater depth in [56]. In dual language, this corresponds to a string with both of its endpoints on the D7-branes at \(z = z_m\), such that the endpoints are initially at the same spatial location but have initial velocities in opposite directions.

In setting up the problem, one realizes that on the gauge theory side there are actually two distinct quark-antiquark configurations: the product of a fundamental \(q\) and an antifundamental \(\bar{q}\) can lead to a \(q\)-\(\bar{q}\) pair either in the singlet or the adjoint representation of the \(SU(N_c)\) gauge group. There is a counterpart to this in the gravity side, because there are precisely two distinct types of consistent initial conditions for the string, which lead it to expand into either a \(\cup\)-shape or a \(\vee\)-shape as its endpoints separate [56]. Curiously, the initial quark velocity is found to be freely adjustable in the adjoint, but not the singlet, configuration— in the latter case it is invariably fixed at the limiting value \(v = v_m\).

As the quark and antiquark separate, we expect them to eventually enter the late-time regime where their rate of energy loss is given by (7). The agreement between the analytic and late-time numeric results is most cleanly seen if instead of
comparing graphs of $x(t)$ or $v(t)$ for the quark (where one would need to look at $t \to \infty$), one examines the plots of $v(x)$. These plots are shown in Fig. 1 for mass parameter $z_m/z_h = 0.2$ (in the neighborhood of the charm quark), corresponding to a limiting velocity $v_m = 0.9992$. It is evident from the figure that the late-time behavior is well-described by (7) in all cases, but there is also an initial period where the behavior is different. This difference is clearly more significant for the singlet than the adjoint case, just as one would expect. For values of the mass in the neighborhood of the charm or bottom masses, the numerical results of [56] indicate that the initial evolution is essentially identical to what it would be in the absence of the plasma. This picture seems rather close to the phenomenological discussion given in [53] (in the context of collisional energy loss): when the singlet quark-antiquark pair is formed within the plasma, there is a delay before the interaction between the newly created sources and the plasma can set up the long range gluonic field profile that is responsible for the late-time dissipation. To examine in more detail the transition to the late-time behavior, for a variety of trajectories one can determine the point $(x_f, v_f)$ beyond which the numeric $v(x)$ curve agrees with the late-time curve to a given accuracy $f$. (The work [56] considered $f = 5$ or $10\%$.) These transition points are schematically indicated by the arrows in Fig. 1.

Figure 1: Quark evolution (velocity as a function of traveled distance, in units of $1/\pi T$) for five different initial conditions. The dotted curves show the results of the numerical integration in [56], contrasted against fits in solid red that use the late-time analytic expression (7). The three dotted curves starting at $v_0 = v_m$ describe singlet configurations with different initial velocity profiles and energies. Notice that the purple curve describes a situation where the quark and antiquark turn around and come to rest while approaching one another. The two remaining curves arise from adjoint configurations with different energies and initial quark velocities. The vertical arrows mark the transition distance where each trajectory enters the late-time regime.
5.2 Screening length

By the time when the quark moves beyond $x_f(v_f)$ and therefore enters the late-time regime, it is certainly insensitive to the presence of the antiquark. It is natural then to ask how the transition distance $x_f(v_f)$ compares against (half of) the length $L_{\text{max}}(v)$ beyond which the quark and antiquark are screened from each other by the plasma. Based on what we just said, we know that for a given velocity we must have $L_{\text{max}}/2 < x_f$, but the actual comparison informs us on whether the transition to the regime where the quark experiences a constant drag coefficient occurs right after the quark and antiquark are screened from each other by the plasma, or if there is an intermediate regime where the quark moves independently from the antiquark but nevertheless feels a drag force that differs from the stationary result of [5, 6, 7], as we found when studying the early-time evolution of an isolated quark in Section 4.2.

The screening length for infinitely massive quarks in MSYM was computed in [14, 15], by considering a quark-antiquark pair moving jointly through the plasma (a related length was plotted in [13], and [15] gave in addition an extensive analysis of the energy of the $q$-$\bar{q}$ pair). This entails a study of a $\cup$-shaped string whose endpoints move in the same direction along the Schwarzschild-AdS geometry, a configuration that, interestingly, feels no drag in the $N_c \to \infty$ limit, on account of its being color-neutral [13, 14, 15].

Over the entire range $0 \leq v \leq 1$ the screening length may be approximated as [15]

$$L_{\text{max}}(v) \approx 0.865 \frac{\pi T}{(1 - v^2)^{1/3}},$$

while in the ultra-relativistic limit, it can be shown analytically that [14]

$$L_{\text{max}}(v) \to \frac{1}{\pi T} \frac{3^{-3/4} 4 \pi^{3/2}}{\Gamma(1/4)^2} (1 - v^2)^{1/4} \approx \frac{0.743}{\pi T} (1 - v^2)^{1/4} \text{ for } v \to 1.$$  \hspace{1cm} (23)

The 1/4 in the exponent here has an intuitive interpretation in terms of the boosted energy density of the thermal medium [47]. The full curve $L_{\text{max}}(v)$ does not deviate far from this asymptotic form, so a decent approximation to it is obtained by replacing 0.743 $\to$ 0.865 in [23], to reproduce the correct value at $v = 0$ (at the expense of introducing a 16% error as $v \to 1$) [14]:

$$L_{\text{max}}(v) \approx 0.865 \frac{\pi T}{(1 - v^2)^{1/4}}.$$ \hspace{1cm} (24)

A comparison between the two approximations (22) and (24) is shown in Fig. 2: overall, the exponent 1/3 is better than 1/4 in the sense that it implies a smaller squared deviation from the numerical results, even though 1/4 leads to a smaller percentage error in the range $v > 0.991 (\gamma > 7.3)$. An attempt to better parametrize the deviation away from the ultra-relativistic behavior was made in [88]. In any case, one should bear in mind that the region of principal interest at RHIC is not really $v \to 1$, but $\gamma v \sim 1$.

\footnote{The $O(1/N_c^2)$ drag force has been worked out by indirect means in [16].}
Figure 2: Screening length $L_{\text{max}}$ (in units of $1/2\pi T$) as a function of velocity (in black) compared against the approximations (22) (in red) and (24) (in blue).

To compare against the transition distance $x_f$ defined in the previous subsection for the back-to-back quark and antiquark, the authors of [56] extended this calculation to the case of finite mass. The main difference with respect to the $m \to \infty$ case is that now the ‘ultra-relativistic’ region would refer to the limit where the pair velocity approaches the limiting velocity $v_m < 1$. The screening length $L_{\text{max}}(v)$ is still relatively well approximated in the full range $0 \leq v \leq v_m$ by the natural modification of the $z_m = 0$ fit (22) (and in all cases, the fit analogous to (24) does a poorer job). The behavior in the $v \to v_m$ region can still be determined analytically, and turns out to be linear in $v_m^2 - v^2$. Over the full velocity interval, the screening length $L_{\text{max}}(v)$ and the transition distance $x_f(v)$ and (half of) are found to be of comparable magnitude and scale with velocity in a similar manner [56]. A similar agreement was reported in [38], and later related work may be found in [86]. Notice that the agreement holds in spite of the fact that the two relevant string configurations are quite different, with motion respectively along and perpendicular to the plane in which the string extends.

The conclusion then is that the transition to the late-time regime (7) takes place immediately after the quark and antiquark lose contact with one another. That is to say, unlike what we found for the forced isolated quark in Section 4.2, here there is no appreciable intermediate stage where the quark and antiquark decelerate independently from one another at a rate that differs substantially from the late-time result of [5, 6]. This then extends and at the same time delimits the region where the analytic results of [5, 6, 7] can justifiably be used to model energy loss in heavy ion collisions.

The way in which the findings reviewed in the present section interface with those of Section 4.2 was discussed in [87]. The upshot is that, in the more realistic situation where the presence of the accompanying antiquark is taken into consideration, the early-time damping results would be most relevant in the stage before the separating quark and antiquark lose contact with one another. This is natural, because screening emerges only when thermal effects are substantial. For a quark ploughing through a significant portion of the plasma, then, the late-time formula (7) actually stands a
chance of controlling the larger portion of the evolution.

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