Dynamical heavy-quark recombination and the non-photonic single electron puzzle at RHIC

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We show that the single, non-photonic electron nuclear modification factor $R^{p}_{AA}$ is affected by the thermal enhancement of the heavy-baryon to heavy-meson ratio in relativistic heavy-ion collisions. We make use of the dynamical quark recombination model to compute such ratio and show that this produces a sizable suppression factor for $R^{p}_{AA}$ at intermediate transverse momenta. We argue that such suppression factor needs to be considered, in addition to the energy loss contribution, in calculations of $R^{p}_{AA}$.

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I. INTRODUCTION

The suppression of single, non-photonic electrons at RHIC [1, 2] is usually attributed to heavy-quark energy losses. However, calculations that successfully describe the nuclear modification factor of hadrons fail to describe the single, non-photonic electron nuclear modification factor $R^{p}_{AA}$ [3, 4, 5, 6]. This has prompted a great deal of effort aimed to better describe the heavy-quark energy loss mechanisms to include not only the radiative part [6, 7, 8] but also the collisional [9] and the medium dynamical properties to compute the radiative piece [10]. As a result, although some improvement in the description of the nuclear modification factor has been gained, it is not yet clear whether the anomalous suppression can be completely attributed to energy losses.

Working along a complementary approach to describe the non-photonic electron yield at RHIC, it has been argued [11, 12] that under the assumption of an enhancement in the heavy-quark baryon to meson ratio, analogous to the case of the proton to pion and the A to kaon ratios in Au+Au collisions [13, 14, 15, 16], it is possible to achieve a larger suppression of the nuclear modification factor. The rationale behind the idea is that heavy-quark mesons have a larger branching ratio to decay inclusively into electrons as compared to heavy-quark baryons, and therefore, when the former are less copiously produced in a heavy-ion environment, the nuclear modification factor decreases, even in the absence of heavy quark energy losses in the plasma.

In order to give a qualitative argument that shows how an enhancement in the heavy-quark baryon to meson ratio can suppress the single, non-photonic electron nuclear modification factor, let us look at the $p_t$ integrated $R^{p}_{AA}$ and to consider that the heavy hadrons are only those containing a single charm,

$$R^{p}_{AA, int} = \frac{1}{\langle n_p \rangle} \frac{N^{D}_{AA}B^{\Lambda-e} + N^{D}_{AA}\Lambda^{B-e}}{N^{D}_{pp}B^{\Lambda-e}}. \tag{1}$$

where $\langle n_p \rangle$ is the average number of participants in the collision for a given centrality class, $N^{D}_{AA}$ ($pp$) refers to the number of $x$-particles produced in $A + A$ ($p + p$) collisions and $B^{\Lambda-e}$ is the branching ratio for the inclusive decay of $x$-particles into electrons. Let us bring Eq. (1) into a form that contains the corresponding $p_t$ integrated nuclear modification factor for particles containing charm. We write

$$R^{p}_{AA, int} = \frac{1}{\langle n_p \rangle} \left( \frac{N^{D}_{AA}}{N^{D}_{pp}} \right) \left( \frac{B^{\Lambda-e} + N^{\Lambda}_{AA}B^{\Lambda-e}}{B^{\Lambda-e} + N^{\Lambda}_{pp}B^{\Lambda-e}} \right). \tag{2}$$

Let us introduce the shorthand notation

$$C = \frac{N^{\Lambda}_{AA}/N^{D}_{AA}}{N^{\Lambda}_{pp}/N^{D}_{pp}}, \quad x = \frac{B^{\Lambda-e}}{B^{\Lambda-e}}, \tag{3}$$

where $C$ represents the enhancement factor for the ratio of charm baryons to mesons in $A + A$ as compared to $p + p$ collisions and $x$ is the charm baryon to meson relative branching ratios for their corresponding inclusive decays into electrons. With these definitions, and after rewriting the factor $N^{D}_{AA}/N^{D}_{pp}$ in the form

$$\frac{N^{D}_{AA}}{N^{D}_{pp}} = \left( \frac{N^{D}_{AA} + N^{\Lambda}_{AA} - N^{\Lambda}_{AA}}{N^{D}_{pp} + N^{\Lambda}_{pp} - N^{\Lambda}_{pp}} \right) = \left( \frac{N^{D}_{AA} + N^{\Lambda}_{AA}}{N^{D}_{pp} + N^{\Lambda}_{pp}} \right) \times \left( 1 - \frac{N^{\Lambda}_{AA}/(N^{D}_{AA} + N^{\Lambda}_{AA})}{1 - N^{\Lambda}_{pp}/(N^{D}_{pp} + N^{\Lambda}_{pp})} \right), \tag{4}$$
we can express Eq. (2) as

$$R_{AA}^{pT \text{int}} = \frac{1}{\langle n_p \rangle} \left( \frac{N_{AA}^D + N_{AA}^A}{N_{pp}^D + N_{pp}^A} \right) \times \left( \frac{1 - N_{AA}^A/(N_{AA}^D + N_{AA}^A)}{1 - N_{pp}^A/(N_{pp}^D + N_{pp}^A)} \right) \times \left( \frac{1 + C_a N_{pp}^D/N_{pp}^A}{1 + x N_{pp}^D/N_{pp}^A} \right) \equiv \frac{1}{\langle n_p \rangle} \left( \frac{N_{AA}^D + N_{AA}^A}{N_{pp}^D + N_{pp}^A} \right) T_{AA}^{pT \text{int}}. \quad (5)$$

When not integrated over transverse momentum, the factor $1/\langle n_p \rangle \left[ (N_{AA}^D + N_{AA}^A)/(N_{pp}^D + N_{pp}^A) \right]$ represents the nuclear modification factor for particles with charm. Let us not assume any particular value for this factor and instead concentrate in the other one in Eq. (5), which can be written as

$$T_{AA}^{pT \text{int}} = \frac{(1 + a)(1 + x C_a)}{(1 + C_a)(1 + x a)} \quad (6)$$

where $a = N_{pp}^A/N_{pp}^D$. The above quantity is plotted in Fig. 1 as a function of $x$ for different combinations of $C a$ and $a$. Notice that the function $T_{AA}^{pT \text{int}}$ is less than 1 when $x < 1$ provided that $C a > a$.

In this work, we want to quantitatively address the question of whether the enhancement factor $C$ times $a$—namely, the heavy-baryon to heavy-meson ratio in Au + Au collisions—can indeed be larger than $a$—namely, the heavy-baryon to heavy-meson ratio in p + p collisions—and if so, how this affects the behavior of the factor $T_{AA}^{pT \text{int}}$ as a function of $p_T$. For these purposes, we use a dynamical recombination scenario that accounts for the fact that the probability to form baryons and mesons can depend on a different way on the evolving density during the collision.

The work is organized as follows: After presenting a brief introduction to the dynamical quark recombination model in Sec. II, we proceed in Sec. III to compute the probabilities to form mesons and baryons containing a heavy quark in a relativistic heavy-ion collision environment. In Sec. IV we use these probabilities to write expressions for the meson and baryon transverse momentum distributions. In Sec. V we compute such distributions as well as the baryon to meson ratio. We convolute such ratio with the branching ratios of charmed baryons and mesons to decay into electrons to obtain the $p_T$ unintegrated function $T_{AA}^{pT \text{int}}$ and show that this can be indeed less than 1. Finally we summarize and conclude in Sec. VI.

II. DYNAMICAL QUARK RECOMBINATION

Recall that hadronization is not an instantaneous process. In fact, lattice calculations [17] show that the phase transition from a deconfined state of quarks and gluons to a hadron gas is, as a function of temperature, not sharp. Working along this line of thought, it has recently been shown [18] that the features of the proton to pion ratio can be well described by means of the so called dynamical quark recombination model (DQRM) that incorporates how the probability to recombine quarks into mesons and baryons depends on density and temperature. Other approaches toward a dynamical description of recombination have been recently formulated [19].

The upshot of the DQRM is that the density evolving probability differs for hadrons made up by two and three constituents with the same mass, that is to say, the relative population of baryons and mesons can be attributed not only to flow but rather to the dynamical properties of quark clustering in a varying density scenario. A natural question is whether those features remain true for baryons and mesons with one constituent heavy-quark and whether a computed, as opposed to assumed, baryon to meson ratio, can at least partially explain the anomalous single, non-photonic electron suppression at RHIC.

The invariant transverse momentum distribution of a given hadron can be written as an integral over the freeze-out space-time hypersurface $\Sigma$, of the relativistically invariant phase space particle density $F(x, P)$,

$$E \frac{dN}{d^3P} = g \int_{\Sigma_f} d\Sigma \frac{P \cdot u(x)}{(2\pi)^3} F(x, P), \quad (7)$$

where $P$ is the hadron’s momentum, $u(x)$ is a future oriented unit four-vector normal to $\Sigma$ and $g$ is the de-
generality factor for the hadron which takes care of the spin degree of freedom. The function $F(x, P)$ contains the information on the probability that the given hadron is formed.

To allow for a dynamical recombination scenario in a thermal environment, let us assume that the phase space particle density $F(x, P)$ can be factorized into the product of a term containing the thermal occupation number, including the effects of a possible flow velocity, and another term containing the system energy density $\epsilon$ driven probability $\mathcal{P}(\epsilon)$, for the coalescence of partons into a given hadron. We thus write

$$F(x, P) = e^{-P \cdot v(x)/T} \mathcal{P}(\epsilon),$$

where $v(x)$ is the flow velocity. As we will show, the probability $\mathcal{P}(\epsilon)$ incorporates in a simple manner the information that the coalescing partons need to be close in configuration space as well as to have a not so different velocity.

To compute the probability $\mathcal{P}(\epsilon)$, it has been shown in Ref. [18] (where we refer the reader to for details) that use can be made of the string flip model [20, 21, 22] in order to get information about the likelihood of clustering of constituent quarks to form hadrons from an effective quark-quark interaction. In short, the model is a variational quantum Monte Carlo simulation that, taking a set of equal number of all color quarks and antiquarks at a given density, computes the optimal configuration of colorless clusters (baryons or mesons) by minimizing the potential energy of the system. At low densities, the model or less clusters (baryons or mesons) by minimizing the potential of the system made up of mesons is given by:

$$V_{mes} = V_{BB} + V_{GG} + V_{RR}$$

where the individual terms are given by Eq. [9] for the corresponding colors. $R(\bar{R})$, $B(\bar{B})$ and $G(\bar{G})$ are the labels for red, blue and green color (anticolor) respectively. Note that this potential can only build pairs.

ii) Baryon-like. In this case the pairing is imposed to be between the different colors in all the possible combinations. In this manner, the many-body potential is:

$$V_{bar} = V_{RB} + V_{BG} + V_{RG}$$

which can build colorless clusters by linking $3\{RBG\}$, $6\{RBGRBG\}$, etc., quarks. Since the interaction is pair-wise, the 3-quark clusters are of the delta (triangular) shape.

According to QCD phenomenology, the formed hadrons should interact weakly due to the short-range nature of the hadron-hadron interaction. This is partially accomplished by the possibility of a quark flipping from one cluster to another. At high energy density, asymptotic freedom demands that quarks must interact weakly. This behavior is observed once the average inter-quark separation is smaller than the typical confining scale.

To describe the evolution of a system of $N$ quarks as a function of the particle density we consider the quarks moving in a three-dimensional box, whose sides have length $L$, and the system described by a variational wave function of the form:

$$\Psi_\lambda(x_1, ..., x_N) = e^{-\lambda V(x_1, ..., x_N)} \Phi_{FG}(x_1, ..., x_N),$$

where $\lambda$ is the single variational parameter, $V(x_1, ..., x_N)$ is the many-body potential defined in Eqs. [11] and [12] for mesons and baryons respectively, and $\Phi_{FG}(x_1, ..., x_N)$ is the Fermi-gas wave function given by a product of Slater determinants, one for each color-flavor combination of quarks, which are built up of single-particle wave functions describing a free particle in a box [22]. The square of the variational wave function is the weighting probability in the sampling, which we carry out using metropolis algorithm.

The variational wave function is taken to have the form given in Eq. [13] since we are interested in the evolution of the system from low to high energy densities. The exponential term is responsible of the clustering correlations. At low energy density, the system is formed by isolated color-singlet hadrons and quarks strongly interacting inside each cluster; in this case, the exponential term of the wave function has a big contribution since the average interquark distance is of the order of the confining scale. In contrast, at high energy density, where asymptotic freedom takes place, the interaction between quarks is weak and the system looks like a Fermi gas of quarks. In this case, the inter-quark separation is much smaller than the confining scale and the effect of the exponential term vanishes. Notice that these features allow us to identify the value of the variational parameter $\lambda$ as being directly proportional to the probability to form a cluster. This fact will be latter exploited to define the density dependent probability $\mathcal{P}(\epsilon)$ since, as we show be-
low, \(\lambda\) changes from a fixed value at low density (isolated clusters) to zero at high density (Fermi gas).

III. PROBABILITIES

All the results we present here come from simulations done with 384 particles, 192 quarks and 192 antiquarks, corresponding to having 32 light quarks and 32 heavy quarks, plus their antiquarks in the three color charges (anti-charges). Hereafter we refer to light quarks as quarks, plus their antiquarks in the three color charges.

We set \(m_c = 10m_u\). We have checked that variations of this particular choice do not affect our relative probabilities.

To determine the variational parameter as a function of density we first select the value of the particle density \(\rho\) in the box, which, for a fixed number of particles, means adjusting the box size. Then, we compute the energy of the system as a function of the variational parameter using the Monte Carlo method described in the previous section. The minimum of the energy determines the optimal variational parameter. We repeat the procedure for a set of values of the particle densities in the region of interest. To get a measure of the probability to form a cluster, we take the variational parameter and divide it by its corresponding value at the lowest density. Notice that since the heavy quarks are not as abundant as the light ones, they do not contribute to the energy density and thus, within the model, this last can be computed by assuming that only light flavors contribute.

The information contained in the variational parameter is global in the sense that it only gives an approximate idea about the average size of the inter-particle distance at a given density, which is not necessarily the same for quarks in a single cluster. To correct for this, and in order to find an appropriate measure of the probability to form baryons and mesons, we need to multiply these variational parameters by the likelihood to find clusters of baryons made up of two-light, one-heavy quark and mesons made up of one-light, one-heavy quark. This likelihood has to consider the fact that the thermal plasma is mainly made up of light quarks and thus that the number of produced heavy quarks is relatively small. To accomplish this, notice that in a model where the interaction between quarks to form clusters is flavor (as well as color) blind, this likelihood should account only for the combinatorial probabilities.

Consider the case where one starts out with a set of \(n\) \(u\)-quarks and \(m\) \(c\)-quarks each coming in three colors.

The number of possible colorless baryons containing three quarks of all possible flavors that can be formed are

\[
\begin{align*}
\text{kind} & \quad \text{number} \\
\text{uuu} & \quad n^3 \\
\text{uuc} & \quad 3n^2m \\
\text{ucc} & \quad 3nm^2 \\
\text{ccc} & \quad m^3,
\end{align*}
\]

and the total number of possible baryons is \((n+m)^3\). The same counting applies for antibaryons when one starts from the same numbers of antiquarks instead of quarks.

Now, consider the case where one starts with a set of \(n\) \(u\)-quarks, \(n\) \(\bar{u}\)-antiquarks, \(m\) \(c\)-quarks and \(m\) \(\bar{c}\)-antiquarks, each coming in three colors. The number of possible colorless mesons containing quark-antiquark pairs of all possible flavors that can be formed are

\[
\begin{align*}
\text{kind} & \quad \text{number} \\
\text{u\bar{u}} & \quad 3n^2 \\
\text{u\bar{c}} & \quad 3nm \\
\text{\bar{u}c} & \quad 3nm \\
\text{\bar{c}c} & \quad 3m^2,
\end{align*}
\]

and the total number of possible mesons is \(3(n+m)^2\).

We now ask for the relative abundance of baryons with respect to mesons computed under the above assumptions on the number of light and heavy quarks that we start from. Since in the case of mesons we are allowing to consider the case \(uc\) as well as \(\bar{u}c\), we need to include in the counting of the groups of three quarks also the antibaryons. Thus the relative abundance is

\[
\frac{c - \text{baryons} + c - \text{antibaryons}}{c - \text{mesons} + c - \text{antimesons}} = \frac{2 \times 3n^2m/(n + m)^3}{2 \times nm/(n + m)^2} = \frac{3n}{2(n + m)}.
\]

Let us now impose that the number of \(u\)-quarks be a multiple \(l\) of the number of \(c\)-quarks, namely, \(n = lm\). Therefore the above relative abundance can be written as

\[
\frac{c - \text{baryons} + c - \text{antibaryons}}{c - \text{mesons} + c - \text{antimesons}} = \frac{3l}{2(l + 1)}.
\]

Notice that in the plasma, the number of \(u\)-quarks greatly exceeds the number of \(c\)-quarks. Therefore a good analytical estimate of the above relative abundance can be obtained by taking \(l \to \infty\) which gives

\[
\frac{c - \text{baryons} + c - \text{antibaryons}}{c - \text{mesons} + c - \text{antimesons}} \overset{l \to \infty}{\to} \frac{3}{2}.
\]

It can be checked that the asymptotic value 3/2 is rapidly reached, for instance, by taking \(l = 30\), the above fraction already becomes 1.475.

Figure 2 shows the probability parameter \(P^{B,M}(\epsilon)\) for baryons and mesons, obtained by multiplying the variational parameter with the corresponding fraction of baryons/mesons formed at the given energy density. In the case of mesons it corresponds to 1/4 irrespective of the density, while for baryons it has a functional form, since the kind of clusters can be different as density increases. For low densities the ratio of the probabilities becomes 3/2, as expected from the combinatorial described above. Shown in the figure is also a fit to the variational parameters with the functional form

\[
f(x) = a_1 + \frac{a_2}{1 + \exp((x - x_0)/dx)}.
\]
FIG. 2: (Color online) Probabilities \( P_{B,M} \) to produce charmed baryons and mesons as a function to the energy density \( \epsilon \). Shown are the results of the Monte Carlo simulation for baryons (full circles) and mesons (open circles) together with a fit to these.

\[
\begin{align*}
  a_B^1 &= 0.0294 \\
  a_B^2 &= 0.3374 \\
  x_0^B &= 0.8604 \\
  dx^B &= 0.0078,
\end{align*}
\]

For baryons

\[
\begin{align*}
  a_M^1 &= 0.0496 \\
  a_M^2 &= 0.1953 \\
  x_0^M &= 0.4812 \\
  dx^M &= 0.0813.
\end{align*}
\]

whereas for mesons

We will use this analytical expression to carry out the calculation of the spectra that we proceed to describe.

IV. BARYON TO MESON RATIO

In order to quantify how the different probabilities to produce sets of three quarks as compared to sets of two quarks affect the particle’s yields as the energy density changes during hadronization, we need to resort to a model for the space-time evolution of the collision. We take Bjorken’s scenario which incorporates the fact that initially, expansion is longitudinal, that is, along the beam direction which we take as the \( \hat{z} \) axis and include transverse flow as a small effect on top of the longitudinal expansion. In this scenario, the relation between the temperature \( T \) and the 1+1 proper-time \( \tau \) is given by

\[
T = T_0 \left( \frac{\tau_0}{\tau} \right)^{v_s^2},
\]

(22)

where \( \tau = \sqrt{t^2 - z^2} \). Equation (22) assumes that the speed of sound \( v_s \) changes slowly with temperature. For simplicity we take \( v_s \) as a constant equal to the ideal gas limit \( v_s^2 = 1/3 \).

We also consider that hadronization takes place on hypersurfaces \( \Sigma \) characterized by a constant value of \( \tau \) and therefore

\[
d\Sigma = \tau \rho \, d\rho \, d\phi \, d\eta,
\]

(23)

where

\[
\eta = \frac{1}{2} \ln \left( \frac{t + z}{t - z} \right),
\]

(24)

is the spatial rapidity and \( \rho, \phi \) are the polar transverse coordinates. Thus, the transverse spectrum for a hadron species \( H \) is given as the average over the hadronization interval of the right hand-side of Eq. (17), namely

\[
E \frac{dN_H}{d^2P} = \frac{g}{\Delta \tau} \int_{\tau_0}^{\tau_f} d\tau \int_{\Sigma} \frac{P \cdot u(x)}{(2\pi)^3} F^H(x,P),
\]

(25)

where \( \Delta \tau = \tau_f - \tau_0 \).
where the magnitude of the transverse flow velocity $v_T$ where

$$\eta$$

freeze-out hypersurfaces of constant $W$ we write the momentum four-vector in components as

$$\tau$$

the probability momenta) to 0.4 (lower curve at low momenta).

in the calculation are

$$f$$

a final time $\tau_f = 8\text{ fm}$. Shown is a range when varying the transverse expansion velocity $v_T$ from 0 (upper curve at low momenta) to 0.4 (lower curve at low momenta).

To find the relation between the energy density $\epsilon$—that the probability $P$ depends upon and $T$, we resort to lattice simulations. For the case of two flavors (since the heavy quark does not thermalize), a fair representation of the data \cite{17} is given by the analytic expression

$$\epsilon/T^4 = a \left[ 1 + \tanh \left( \frac{T - T_c}{bT_c} \right) \right],$$

with $a = 4.82$ and $b = 0.132$. We take $T_c = 175\text{ MeV}$.

The flow four-velocity vector $u^\mu$ is given by

$$u^\mu = (\cosh \eta \cosh \eta_T, \sinh \eta_T \cos \phi, \sinh \eta_T \sin \phi, \sinh \eta \cosh \eta_T),$$

where the magnitude of the transverse flow velocity $v_T$ and $\eta_T$ are related by $v_T = \tanh \eta_T$. The normal to the freeze-out hypersurfaces of constant $\tau$, $u^\mu$, is given by

$$u^\mu = (\cosh \eta, 0, 0, \sinh \eta).$$

We write the momentum four-vector in components as

$$P^\mu = (m_T \cosh \eta, p_T \cos \Phi, p_T \sin \Phi, m_T \sinh \eta),$$

where $y$ is the 1 + 1 momentum rapidity given by

$$y = \frac{1}{2} \ln \left( \frac{E + p_z}{E - p_z} \right),$$

and $\Phi$ the azimuthal angle of the momentum components in the transverse plane.

Therefore, the products $P \cdot u$ and $P \cdot v$ appearing in Eq. \cite{29} can be written as

$$P \cdot v = m_T \cosh(\eta - y) \cos \eta_T - p_T \cos(\phi - \Phi) \sinh \eta_T,$$

$$P \cdot u = m_T \cosh(\eta - y),$$

(31)

Considering the situation of central collisions, we can assume that there is no dependence of the particle yield on the transverse polar coordinates. Integration over these variables gives

$$\frac{dN}{p_T d\eta} = \frac{g m_T^2}{4\pi} \rho_{\text{nucl}}^2 \int_{\tau_0}^{\tau_f} \tau d\tau \mathcal{P}(\tau) I_0(p_T \sinh \eta_T/T)$$

$$\times \int d\eta \cosh(y - \eta) e^{-[m_T \cosh(y - \eta) \cosh \eta_T]/T}.$$  

(32)

where $\rho_{\text{nucl}}$ is the radius of the colliding nuclei and $I_0$ is the Bessel function $I$ of order zero.

We now consider as a further simplification that the space-time and momentum rapidities are completely correlated, that is $\eta \sim y$. Under this assumption, the integral over $\eta$ in Eq. \cite{32} can be performed and we finally get

$$\frac{dN}{p_T d\eta} = g m_T^2 \Delta \eta \rho_{\text{nucl}}^2 \int_{\tau_0}^{\tau_f} \tau d\tau \mathcal{P}(\tau) I_0(p_T \sinh \eta_T/T)e^{-[m_T \cosh(y - \eta) \cosh \eta_T]/T}.$$  

(33)

Armed with the expression to compute the hadron transverse momentum distribution, we now proceed to apply the analysis to the computation of the charmed meson and baryon distributions.

V. RESULTS

Figure \ref{fig:3} shows examples of the transverse momentum distributions for mesons and baryons obtained from Eq. \cite{33}. We set the masses of the charmed baryons and mesons as $m_B = 2.29\text{ GeV}$ (corresponding to $\Lambda_c$) and $m^M = 1.87\text{ GeV}$ (corresponding to $D$). We take the initial hadronization time as $\tau_0 = 1\text{ fm}$, at an initial temperature $T_0 = 200\text{ MeV}$ and the final hadronization temperature as $T_f = 100\text{ MeV}$, corresponding, according to Eq. \cite{22}, to a final time $\tau_f = 8\text{ fm}$. Shown are the cases with $v_T = 0$ and $v_T = 0.4$. Notice that a finite transverse expansion velocity produces a broadening of the distributions, as expected.

Figure \ref{fig:4} shows the charmed baryon to meson ratio obtained from the ratio of the above transverse momentum distributions. Shown is a range for this ratio when varying the transverse expansion velocity $v_T$ from 0 to 0.4. Notice that for a finite $v_T$, this ratio goes above 1 for $p_T \gtrsim 3.5\text{ GeV}$. 

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**FIG. 4:** (Color online) Charmed baryon to meson ratio, $Ca$, as a function of transverse momentum. The parameters used in the calculation are $m_B = 2.29\text{ GeV}$, $m^M = 1.87\text{ GeV}$, $\tau_0 = 1\text{ fm}$, $T_0 = 200\text{ MeV}$, $T_f = 100\text{ MeV}$, corresponding to a final time $\tau_f = 8\text{ fm}$. Shown is a range when varying the transverse expansion velocity $v_T$ from 0 (upper curve at low momenta) to 0.4 (lower curve at low momenta).
FIG. 5: (Color online) Suppression factor $T_{AA}^e$ as a function of transverse momentum. The parameters used in the calculation are $m_D^p = 2.29$ GeV, $m_D^p = 1.87$ GeV, $\tau_0 = 1$ fm, $T_0 = 200$ MeV, $T_f = 100$ MeV, corresponding to a final time $\tau_f = 8$ fm, $x = 0.14$, $a = 0.073$. Shown is a range for the transverse expansion velocity form $v_T = 0$ (upper curve at low $p_T$) and $v_T = 0.4$ (lower curve at low $p_T$).

We now proceed to compute the $p_T$ unintegrated function $T_{AA}^e$. For this purpose, we take that the possible charmed mesons decaying inclusively into electrons or positrons are $D^\pm (B^{D^\pm} \rightarrow e^\pm = 16.0\%)$, $D^0$, $\bar{D}^0 (B^{D^0} \rightarrow e^\pm = 6.53\%)$, $D_s^\pm (B^{D_s^\pm} \rightarrow e^\pm = 8\%)$ and that the possible charmed baryons decaying inclusively into electrons or positrons are $\Lambda_c, \bar{\Lambda}_c (B^{\Lambda_c, \bar{\Lambda}_c} \rightarrow e^\pm = 4.5\%)$. Thus we get

$$x = 0.14.$$ (34)

We also approximate the masses of all the charmed mesons considered to be equal to the mass of the $D^{\pm}$. From Eq. 34 we see that, without integrating over $p_T$, the dependence on the transverse momentum comes from $a = (dN_{pp}^\Lambda / d\tau_T) / (dN_{pp}^D / d\tau_T)$ and the product $C a = (dN_{AA}^\Lambda / d\tau_T) / (dN_{AA}^D / d\tau_T)$. The integrated ratio $a^{int}$ has been computed in Ref. 12 using a Pythia simulation, with the result $a^{int} = 0.073$. We have also performed a simulation using Pythia at NLO with 100,000 events and have found that with such distributions, the ratio of charmed baryons to charmed mesons in $p + p$ collisions at $\sqrt{s_{NN}} = 200$ GeV is flat up to $p_T \simeq 5$ GeV and consistent with the value reported in Ref. 12. Therefore, for simplicity we take $a$ as a constant equal to the above quoted number. Thus

$$T_{AA}^e \simeq \frac{1 + x a^{int}}{1 + x a^{int}} 1 + (dN_{AA}^\Lambda / d\tau_T) / (dN_{AA}^D / d\tau_T).$$ (35)

Figure 5 shows $T_{AA}^e$ as a function of $p_T$. We have used a range of values for the transverse expansion velocity between $v_T = 0$ and $v_T = 0.4$. We see that for the chosen evolution parameters, $T_{AA}^e$ is indeed smaller than 1 and thus it contributes to the suppression of the single non-photonic electron nuclear modification factor $R_{AA}^e$.

VI. CONCLUSIONS

In this work we have shown that the anomalous suppression of the single non-photonic electron nuclear modification factor $R_{AA}^e$ can be partially understood by realizing that this quantity is affected by an enhancement in the charmed baryon to meson ratio at intermediate $p_T$ in $Au + Au$ collisions. This enhancement is due to the fact that in this region, thermal recombination becomes the dominant mechanism for hadron production. We have made use of the DQRM to calculate this ratio and have shown that for moderate and even for vanishing transverse expansion velocities, it indeed can be larger than the charmed baryon to meson ratio in $p + p$ collisions. This enhancement in turn produces that the function $T_{AA}^e$ is below 1 and thus contributes to the suppression factor introduced by considering energy losses due to the propagation of heavy flavors in the plasma.

It is worth to keep in mind some important features concerning the results of the present calculation: First, notice that we have not included the momentum shift introduced by energy losses when computing the transverse distributions of charmed mesons and baryons. This is so because for $R_{AA}^e$, energy losses should be included in the prefactor of the function $T_{AA}^e$. In this sense, in order to avoid a double counting of the effect, the ratio that goes into the calculation of this last function is the $raw$ ratio. Second, it is expected that at some value of $p_T$, fragmentation becomes the dominant mechanism for hadron production and therefore that the charmed baryon to meson ratio decreases above that $p_T$ value, given that fragmentation produces more mesons than baryons. Third, we have considered finite values of transverse flow for charmed mesons and baryons even though it might be questionable that heavy flavors also flow as light flavors do. Nevertheless, there seems to be some experimental support for heavy quark flow [24]. In this sense, the flow strength range we have considered is only for moderate values. Notice however that even in the absence of flow the suppression factor keeps being less than 1. Some of these issues will be the subject of a future work to appear elsewhere.
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