A Nonequilibrium Field Theory Description of the Bose-Einstein Condensate

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We study the detailed out of equilibrium time evolution of a homogeneous Bose-Einstein condensate. We consider a nonrelativistic quantum theory for a self-interacting complex scalar field, immersed in a thermal bath, as an effective microscopic model for the description of the Bose-Einstein condensate. This approach yields the following main results: (i) the interaction between fluctuations proves to be crucial in the mechanism of instability generation; (ii) there are essentially two regimes in the $k$-space, with a crossover for $k^2/2m \sim 2\lambda|\phi_0|^2$, where, in our notation, $\lambda$ is the coupling constant and $|\phi_0|^2$ is the condensate density; (iii) a set of coupled equations that determines completely the nonequilibrium dynamics of the condensate density as a function of the temperature and of the total density of the gas. PACS number(s): 03.75.Fi, 05.30.Jp, 11.10.Wx

The experimental verification of the phenomenon of Bose-Einstein condensation in weakly interacting gases has boosted a large number of theoretical investigations on the dynamics of weakly-interacting dilute gas systems [for a recent review, see e.g. Ref. \cite{1} and references therein]. Current experiments and planned ones make it possible to probe different aspects of the Bose-Einstein condensate formation, with great control over interactions, trapping potentials, etc. Nevertheless, a basic problem not yet fully understood is the following: given an initial state, how will the condensate evolve with time? In special, the time scales for the condensate formation and its final size are important quantities involved in recent experiments with dilute atomic gases \cite{2}.

On the theoretical side, however, only restrict progress has been achieved concerning the problems above. Previous studies by Stoof \cite{3} were able to give a qualitative idea of the various time scales involved during the condensate formation. In fact, they were the first attempts to an-
the fluctuation field. In this way, as we show below, one can make clear the appearance of instability modes towards the condensation formation once the interactions between fluctuations are taken into account. With the decomposition above, the interaction term for fluctuations in the Lagrangian becomes \( \lambda (\phi^* \phi)^2 \). The mean-field approximation amounts to the following:

\[
\lambda (\phi^* \phi)^2 = 4\lambda (\phi^* \varphi)\varphi^* \varphi + \left[ \lambda (\phi^* \phi)^2 - 4\lambda (\phi^* \varphi)\varphi^* \varphi \right],
\]

where the first term in the rhs is taken as part of the quadratic Lagrangian for fluctuations, and the term inside the square brackets is taken as part of the interaction Lagrangian.

From the decomposition of the fields \( \phi^* \phi \), the quadratic part of the Lagrangian density for the fluctuations, \( \mathcal{L}_0(\varphi, \varphi^*) \), may be written as

\[
\mathcal{L}_0(\varphi, \varphi^*) = \varphi^* \left[ \frac{d}{dt} + \frac{1}{2m} \nabla^2 \right] \varphi + \varphi^* (-\lambda \varphi_0^2) \varphi
+ \varphi (-\lambda \varphi_0^2) \varphi^*.
\]

Here, we have used the fact that, under the field decomposition in the condensate and out of the condensate modes, the density constraint then becomes \( \langle \phi^* \phi \rangle = |\varphi_0|^2 + \langle \phi^* \varphi \rangle = n \). Additionally, assuming that at the initial time the system is mostly composed of particles outside the condensate, \( \langle \phi^* \phi \rangle \approx (\phi^* \varphi) \) (at \( t = 0 \)), simple relations involving the generating functional for the correlation functions (see, for instance, the last section of chapter 2 in [1]) allow us to write the total number density \( n \) of particles in terms of the chemical potential \( \mu \), valid in the mean-field approximation for the potential, as: \( \mu = 4\lambda n \). Note that this is just the expression obtained also in the Hartree-Popov approximation [1], which turns out to satisfy the Hugenholtz-Pines relation [1] that would be obtained in the equilibrium problem. These considerations lead to the quadratic Lagrangian for the fluctuations shown above.

The scenario we have in mind is that for time \( t < 0 \) the initial state is in equilibrium at a temperature \( T_i \gg T_c \). At \( t = 0 \) the system is then abruptly quenched to a much lower temperature \( T_f \ll T_c \). \( T_f \) is the temperature of the thermal bath in which the system is immersed and, of course, it will be the equilibrium temperature which the system will reach asymptotically. This kind of quench is easily attained in the experiments of Bose-Einstein condensation of atomic gases, where the typical relaxation time scales are long enough (around \( 0.1s \), depending on the temperature [2]) to allow for a fast drop in the temperature of the system that evolves afterwards out of equilibrium. With this choice of initial state, it is reasonable to approximate the dynamics of the build up of the condensate state, which at the initial time is \( n_{\text{cond.}}(t = 0) = |\varphi_0(t = 0)|^2 \approx 0 \), and the depletion of the excited states (which at \( t = 0 \) it is given by \( n \approx n_{\text{exc.}}(t = 0) = \langle \varphi^* \varphi \rangle \)) as essentially a two-level problem. It is clear that this approximation breaks down for temperatures close to the critical temperature, where the detailed treatment would require a thorough study of the dynamics among the many levels of excited states. In the above approximation, the condensate builds up subject to the density constraint relation, which may be expressed in terms of the averages of the real and the imaginary parts of \( \varphi \) and \( \varphi^* \) \( (\varphi = \varphi_1 + i\varphi_2 \) and \( \varphi^* = \varphi_1 - i\varphi_2 \), respectively). Spatial translational invariance yields:

\[
|\varphi_0(t)|^2 + n_{\text{exc.}}(t) = n,
\]

\[
n_{\text{exc.}}(t) = \langle \varphi_1(t)\varphi_1(t) \rangle + \langle \varphi_2(t)\varphi_2(t) \rangle.
\]

The field averages above can be expressed in terms of the Green’s functions for \( \varphi_1 \) and \( \varphi_2 \) as \( (j = 1, 2) \)

\[
\langle \varphi_j(t)\varphi_j(t) \rangle = \int \frac{d^3k}{(2\pi)^3} \left[ -iG_j^\rightarrow(k, t, t) \right],
\]

where \( G_j^\rightarrow(k, t, t) \) is defined from the Green’s functions for the fields in the closed-time path \( \mathbb{I} \) in (momentum space)

\[
G_{jj}^\rightarrow(k, t, t') = G_{jj}^\rightarrow(k, t, t')\theta(t - t') + G_{jj}^\rightarrow(k, t, t')\theta(t' - t''),
\]

\[
G_{jj}^\rightarrow(k, t, t') = -G_{jj}^\rightarrow(k, t, t'),
\]

\[
G_{jj}^\rightarrow(k, t, t') = -G_{jj}^\rightarrow(k, t, t').
\]

The functions \( G^\rightarrow \) and \( G^\leftarrow \) satisfy the property \( G_j^\rightarrow(k, t, t') = G_j^\leftarrow(k, t-i\beta, t') \), which is recognized as the periodicity condition in imaginary time (Kubo-Martin-Schwinger (KMS) condition). Here \( \beta \) is the inverse of the temperature of the thermal bath and appears here as a consequence of the boundary conditions arising from the construction of the complex time path. \( G^\rightarrow \) and \( G^\leftarrow \) are constructed from the homogeneous solutions to the operator of quadratic fluctuation modes, which, using Eq. [3] expressed in terms of \( \varphi_1 \) and \( \varphi_2 \), are given by (in momentum space)

\[
\frac{d\chi_2(k, t)}{dt} + \left( \frac{k^2}{2m} + 2\lambda |\varphi_0|^2 \right) \chi_1(k, t) = 0,
\]

\[
\frac{d\chi_1(k, t)}{dt} + \left( \frac{k^2}{2m} - 2\lambda |\varphi_0|^2 \right) \chi_2(k, t) = 0.
\]

The boundary conditions for the solutions of the equations above are such that, for \( t < 0 \), \( |\varphi_0(t)|^2 = 0 \), \( \chi_1(k, t) = \cos(\varepsilon_k t) \) and \( \chi_2(k, t) = -\sin(\varepsilon_k t) \), where \( \varepsilon_k = k^2/(2m) \). In terms of these fluctuations modes, the Green’s functions are expressed as

\[
G_{jj}^\rightarrow(k, t, t') = \frac{i}{2(1 - e^{-\beta\varepsilon_k})} \times \left[ \chi_j(k, t)\chi_j^\rightarrow(k, t') + e^{-\beta\varepsilon_k}\chi_j^\rightarrow(k, t)\chi_j(k, t') \right].
\]
and \( G_{ij}^0(k, t, t') = G^0_{ji}(k, t', t) \).

By decoupling the set of equations in (6), we can readily identify that those modes with \((k^2 / 2m) < 2|\varphi_0|^2\) are unstable and drive the excited particles towards condensation. Note also that not all the excited particles condense, since there will always be a fraction (which depends on various parameters for a particular system and on the temperature of the thermal bath) of excited modes, with high enough frequency, that remains stable. This will be clear from our numerical results shown later.

Using Eqs. (6), (8) and the boundary conditions on the Green’s functions, Eq. (3), together with the initial condition on the density (at \( t = 0 \), as defined before), one can then show that \( n_{\text{exc}}(t) \) can be expressed as (by subtracting the zero point divergent contribution)

\[
n_{\text{exc}}(t) = \left( \frac{\beta}{\beta_c} \right)^{3/2} \int \frac{d^3k}{(2\pi)^3} \left[ |\chi_1(k, t)|^2 + |\chi_2(k, t)|^2 \right] n_k(\beta),
\]

where \( n_k(\beta) = (e^{\beta\epsilon_k} - 1)^{-1} \), and \( \beta_c \) is the inverse of the equilibrium critical temperature, defined in terms of the total gas density, \( n \). It should also be noted that in our out of equilibrium approach there are no infrared divergences since the finite time is a natural regulator. However for the equilibrium \( t \to \infty \), the critical temperature will be modified by the interactions as pointed out in Ref. [4].

The expression above can also be obtained directly in terms of the Green’s functions for the complex fields \( \varphi, \varphi^* \):

\[
\langle \varphi(t) \varphi^*(t) \rangle = \int \frac{d^3k}{(2\pi)^3} \left[ -iG_{\varphi \varphi^*}(k, t, t) \right],
\]

\[
\langle \varphi^*(t) \varphi(t) \rangle = \int \frac{d^3k}{(2\pi)^3} \left[ -iG_{\varphi^* \varphi^*}(k, t, t) \right].
\]

In terms of (10) and (11), we have \( \langle \varphi_1(t) \varphi_1(t) \rangle + \langle \varphi_2(t) \varphi_2(t) \rangle = \langle (\varphi(t) \varphi^*(t)) + (\varphi^*(t) \varphi(t)) \rangle / 2 \), and the KMS condition can be expressed, in this case, as \( G_{\varphi \varphi^*}(k, t - i\beta, t') = \frac{1}{2} [G_{\varphi \varphi^*}(k, t, t')] + [G_{\varphi^* \varphi^*}(k, t, t')]^* \), or \( G_{\varphi^* \varphi^*}(k, t - i\beta, t') = [G_{\varphi \varphi^*}(k, t, t')]^* \).

Equation (12) is the first order term in the finite temperature quantum many-body perturbation expansion for \( \langle \varphi^2 \rangle \). Higher-order corrections for the equal-time two-point field averages can be expressed in terms of the coincidence limit of the (causal) two-point Green’s functions \( G_{\varphi \varphi^*} \) and \( G_{\varphi^* \varphi} \), which satisfy the Dyson equations (the indices stand for \( \varphi \) and \( \varphi^* \)):

\[
G_{ij} = G^0_{ij} + G^0_{ik} G\Sigma \delta_{kj} G_{lj},
\]

where \( \Sigma \) is the (matrix) self-energy, and \( G^0_{ij} \) is the zeroth-order non-interacting Green’s function, satisfying the equation (in momentum space)

\[
\left[ \pm \frac{i}{\beta} \frac{d}{dt} - \epsilon_k \right] G^0_{\varphi \varphi^*}(\varphi, \varphi^*)(k, t, t') = \delta(t - t').
\]

One of the advantages of expressing the Green’s functions in terms of the solutions of (7) is the possibility of obtaining, in an unambiguous way, all higher-order corrections to the two-point and many-point functions (3).

By using Eq. (4), we can rewrite the constraint on the density as

\[
|\varphi_0(t)|^2 = \frac{(\beta/\beta_c)^{3/2}}{2\pi^2} \int_0^{16\pi a |\varphi_0(t)|^2} \frac{dk}{k^2} \times \left[ 1 - \left( |\chi_1(k, t)|^2 + |\chi_2(k, t)|^2 \right) n_k(\beta) \right],
\]

that reproduces the result obtained by Stoof (3) for the limit \( t \to \infty \). Equations (6) and (14) form an integro-differential system that may be solved for \( \varphi_0(t) \) numerically, given the initial conditions for \( \varphi_0(t) \), \( \chi_1(k, t) \) and \( \chi_2(k, t) \) mentioned before. Indeed, this system of equations determines completely the time evolution of the condensate density as a function of the temperature and of the total density of the gas. Explicit results for different temperatures are shown in Fig. 1. It is important to point out at this stage that the evolution of the condensate is completely driven by the interactions between the microscopic fluctuations of the field around the condensate.

Throughout this letter, we have developed an out-of-equilibrium non-perturbative quantum field theory description of the condensation process of an interacting homogeneous Bose-Einstein gas quenched below the critical temperature. In summary, this approach yielded the following main results: (i) The interaction between fluctuations proved to be crucial in the mechanism of instability generation; without it, there is simply no macroscopic condensate at all. (ii) There are essentially two regimes in the \( k \)-space: for \( (k^2 / 2m) \ll 2\lambda|\varphi_0|^2 \), we have unstable modes that decay exponentially, while for \( (k^2 / 2m) \gg 2\lambda|\varphi_0|^2 \), we have stable modes that oscillate, with a crossover for \( (k^2 / 2m) \sim 2\lambda|\varphi_0|^2 \). (iii) Equations (6) and (14) come from a microscopic model for the weak-interaction gas, and determine completely the dynamics of the condensate. In fact, they are non-perturbative and certainly implement a resummation of the ladder Feynman diagrams mentioned by Stoof in (3). Indeed, the highly nonequilibrium character of this description should complement the usual approach via Boltzmann equation.

Although we focused this letter on the instability process that generates the condensate (i.e., the short time behavior), for \( t \to \infty \) our results confirm the behavior predicted in Ref. (3) for this limit. However, the equilibrium \( (t \to \infty) \) values of the condensate fraction are lower than the experimental results (2) and the calculations of Dalfovo et. al. (1). This may be due to our approximation of neglecting incoherent collisional processes, which is a
valid approximation in an infinite homogeneous gas at very low temperatures and densities but otherwise may give an important contribution. We expect that the self-consistent inclusion of pair terms should account for most of these contributions. We will report on these improvements of our mean-field approximation in a future publication [15].

In spite of the absence of non-homogeneity effects, we hope that the approach developed here may be useful in the analysis of transients in realistic Bose-Einstein condensation experiments with atomic gases. Moreover, with a suitable generalization of the formalism presented above, we could be able to develop a theoretical description of the dynamical aspects of a recently proposed experiment [16] (currently in progress), regarding the Bose-Einstein condensation in a weakly-interacting photon gas in a nonlinear Fabry-Perot cavity [17].

![Graph](image)

**FIG. 1.** Condensate density as a function of time for na^3 = 0.01 and T_1/T_c = 0.06, T_2/T_c = 0.08 and T_3/T_c = 0.1. Here, \( \tau \equiv (\hbar/ma^2) t \) is a dimensionless time and \( \rho_0 \equiv a^3 |\varphi_0|^2 \) is a dimensionless density.

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