Generalized Cornu-type Spirals and their Darboux parametric deformations

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We generalize the Fresnel integrals and introduce a class of planar spirals \( \mathcal{F}_n \), which contains the Cornu spiral as the case \( \mathcal{F}_2 \). Their Darboux parametric deformations are also investigated. The \( \mathcal{F}_3 \) spiral and its Darboux deformed counterparts are illustrated.

Keywords: Planar spiral; Cornu spiral; Riccati equation; Argand plot; parametric Darboux distortion

I. INTRODUCTION

In a previous paper [1], two of us considered parametric Darboux deformations based on the general Riccati solution [2–6] of the Cornu spiral [7] and their supersymmetric origin has been revealed. The purpose of this work is to introduce a class of generalized Cornu spirals and to study their properties using the same approach. For this, we consider the following trigonometric functions

\[
\cos \left( \frac{p}{n} z^n \right), \quad (1)
\]

\[
\sin \left( \frac{p}{n} z^n \right), \quad (2)
\]

where \( p > 1 \) is an arbitrary positive constant, which can be naturally chosen as \( p = \pi \). The spirals, \( \mathcal{F}_n \), are defined through the Argand plane representation, \( \mathcal{F}_n = X_n + iY_n \), with \( X_n \) and \( Y_n \), as the following integrals

\[
\mathcal{C}_n(z) \equiv X_n(z) = \int_0^z \cos \left( \frac{p}{n} s^n \right) \, ds, \quad (3)
\]

\[
\mathcal{S}_n(z) \equiv Y_n(z) = \int_0^z \sin \left( \frac{p}{n} s^n \right) \, ds, \quad (4)
\]

which we call the generalized Fresnel integrals and are parametrized by the arclength of the spiral, \( s \). In principle, \( n \) can be any positive integer, but here we will discard the cases \( n = 0 \), which is the logarithmic spiral, and \( n = 1 \) since the spiral \( \mathcal{F}_1 \) is the circle. For the case \( n = 2 \) and \( p = \pi \), \( \mathcal{C}_2(z) \) and \( \mathcal{S}_2(z) \) are the Fresnel integrals and the spiral \( \mathcal{F}_2 \) is the Cornu spiral that were discussed in [1].

We recall here that the planar curves are characterized by the relationship between the curvature and the arc length known as the Cesàro formula, and by the same token the Cornu spiral is defined as the curve whose curvature increases linearly with arc length. This means that the radius of curvature \( \rho \) times the arc length \( s \) is constant at each point of the Cornu spiral. In the case of the generalized Cornu spirals, the Cesàro equation has the form \( \rho(s) = s^{1-n}/p \), which for \( n = 2 \) and \( p = \pi \) reduces to the standard Cesàro equation for the common Cornu spiral. For \( n = 1 \) one obtains the constant radius of curvature of the circle, and for \( n = 0 \), the radius of curvature is proportional to the arc length which yields to the logarithmic spiral.

In this paper, we consider the set of generalized spirals \( \mathcal{F}_n \) by means of the complex parameter which appears in the general solution of the Riccati equation that corresponds to the associated generalized Fresnel integrals. In Section II, we show the reduction of the third order ordinary differential equation (ODE) satisfied by these integrals as particular solutions to the corresponding Riccati equation, whose general solution is obtained explicitly. We then write the solution of the third order ODE based on the general Riccati solution and present Argand plots of this solution. In Section III, the similarity with supersymmetric quantum mechanics is emphasized by means of the factorization approach [2–6] which is applied to the second-order linear ODE that comes into play in the reduction process of Section II.

II. THE \( \mathcal{F}_n \) SPIRALS: THE THIRD AND SECOND ORDER ODES AND THE RICCATI EQUATION

The following linear third order ODE

\[
z w''' - (n - 1) w'' + p^2 z^{2n-1} w' = 0 \quad (5)
\]

is satisfied by the generalized Fresnel integrals \( \mathcal{C}_n(z) \) and \( \mathcal{S}_n(z) \), and so the general solution can be written as the superposition

\[
w(z) = \alpha \mathcal{C}_n(z) + \beta \mathcal{S}_n(z) + w(0). \quad (6)
\]

In addition, \( v = w' \) can be reduced to the second order ODE

\[
v'' - \frac{(n - 1)}{z} v' + \left( \frac{p^2 z^{n-1}}{z} \right)^2 v = 0 \quad (7)
\]

by using \( w'(z) = v(z) \), with \( \prime = \frac{dz}{dz} \). Letting \( z^n = \zeta \) we obtain the simple harmonic oscillator

\[
\frac{d^2v}{d\zeta^2} + \left( \frac{p}{n} \right)^2 v = 0. \quad (8)
\]
Thus, the solution for \( v(z) = c_1 \cos \left( \frac{p}{n} z^n \right) + c_2 \sin \left( \frac{p}{n} z^n \right) \),
(9)
and by one integration, one obtains the solution \( \gamma \).

On the other hand, using the logarithmic derivative
\[
y(z) = \frac{v'(z)}{v(z)},
\]
becomes the Riccati equation
\[
y' + y^2 = \frac{n - 1}{z} y - p^2 z^{2n-2}
\]
(10)
with particular solution
\[
y_p(z) = i p z^{n-1}.
\]
(11)

To construct the general solution of Riccati equation (10), using any particular solution \( y_p \), we write the Bernoulli ansatz
\[
y_B(z) = y_p(z) + \frac{1}{u(z)},
\]
(12)
where \( u \) satisfies the linear equation
\[
u' + \left( \frac{n - 1}{z} - 2y_p \right) u = 1.
\]
(13)
The solution of (13) is
\[
u(z) = \gamma + \int_0^z \frac{\mu(s)ds}{\mu(z)}
\]
(14)
where \( \mu(z) \) is the integrating factor
\[
\mu(z) = z^{n-1} e^{-2i \frac{\pi}{n} z^n},
\]
(15)
which gives the general Riccati solution in the form
\[
y_B(z) = y_p(z) + \frac{\mu(z)}{\gamma + \int_0^z \mu(s)ds},
\]
(16)
and \( \gamma \) arbitrary. Using (15) in (10) and redefining the integration constant as \( \gamma = \frac{-i \pi}{2p} \), provides the explicit form of the general Riccati solution
\[
y_B(z) = ipz^{n-1} \left( \theta e^{2i \frac{\pi}{n} z^n} - 1 \right). \]
(17)

We then find a polar form of the general solution for (7) using \( v_\beta(z) = Re^{i \theta} y_\beta(z)dz \) to obtain
\[
v_\beta(z) = R \left( e^{-i \frac{\pi}{n} z^n} + \theta e^{i \frac{\pi}{n} z^n} \right),
\]
(18)
By one integration, assuming \( w(0) = 0 \) and using Euler’s formula, the deformed solution of (15) is given by
\[
w_B(z; n) = R \left[ (1 + \theta) C_n(z) + i(-1 + \theta) S_n(z) \right]
\]
(19)
and letting \( \theta = a + ib \neq 0 \), we obtain
\[
w_R(z; n) = R \left[ (a + 1) C_n(z) - b S_n(z) \right],
\]
\[w_T(z; n) = R \left[ b C_n(z) + (a - 1) S_n(z) \right].
\]
(20)
Comparing (6) with (20), one can see that we managed to replace the superposition constants \( \alpha \) and \( \beta \) by the real and imaginary components of the parameter entering the general Riccati solution. This replacement helps to disentangle an underlying supersymmetric structure of the solution expressed in this way as it will be shown in the next section.

Argand plots of the generalized Fresnel spirals for \( n = 3 \) are shown in Fig. 1 based on different values of the deformation parameter \( \gamma = \frac{-i \pi}{2p} \). In the center, when \( a = 0 \) and \( b = 0 \) we used the standard undeformed spiral given by \( w_R(z; 3) = C_3(z) \) and \( w_T(z; 3) = S_3(z) \). All figures except \( a = 0 \), \( b = 0 \) are scaled by the factor \( R = 1/\sqrt{a^2 + b^2} \).

To find the analytical expressions of the generalized Fresnel integrals for \( n \geq 2 \), together with their foci corresponding to \( z \rightarrow \pm \infty \) we proceed as follows. First, we let \( \varphi(s) = \frac{p}{n} s^n \), and we notice that \( d\varphi = \frac{ds}{\rho(s)} \), and by inverting we find \( \rho(\varphi) = \left( \frac{n}{p} \right)^{n-1} \varphi^{\frac{n}{n-1}} \), thus (3) and (4) become
\[
C_n(z^*) = \left( \frac{n}{p} \right)^{\frac{n}{2n-1}} \frac{1}{2n-1} \int_0^{z^*} \varphi^{\frac{n}{n-1}} \cos \varphi \ d\varphi,
\]
(21)
\[
S_n(z^*) = \left( \frac{n}{p} \right)^{\frac{n}{2n-1}} \frac{1}{2n-1} \int_0^{z^*} \varphi^{\frac{n}{n-1}} \sin \varphi \ d\varphi,
\]
(22)
with the upper limit given by \( z^* = p \frac{n}{n} z^n \). These integrals can now be found in closed form in terms of the hypergeometric function, as they are
\[
C_n(z^*) = \left( \frac{n}{p} \right)^{\frac{n}{2n-1}} _2 F_1 \left[ \frac{1}{2n}; \frac{1}{2} \right; 1; 1 + \frac{3}{2n}; -\left( \frac{z^*}{2} \right)^2 \right],
\]
(23)
\[
S_n(z^*) = \left( \frac{n}{p} \right)^{\frac{n}{2n-1}} _2 F_1 \left[ \frac{1}{2}; \frac{1}{n} \right; 3; 2 \left( \frac{3}{2} + \frac{1}{n} \right); -\left( \frac{z^*}{2} \right)^2 \right].
\]
(24)
The location of the foci of the generalized undeformed spirals are obtained from (23) and (24) by letting \( z^* \rightarrow \pm \infty \). The focus in the first quadrant has the coordinates
\[
x_{x_n} = \left( \frac{n}{p} \right)^{\frac{n}{2n-1}} \Gamma \left( 1 + \frac{1}{n} \right) \cos \left( \frac{\pi}{2n} \right),
\]
(25)
\[
y_{y_n} = \left( \frac{n}{p} \right)^{\frac{n}{2n-1}} \Gamma \left( 1 + \frac{1}{n} \right) \sin \left( \frac{\pi}{2n} \right),
\]
(26)
whereas the second focus can be found in the second quadrant at \( -(x_{y_n}, y_{x_n}) \) if \( n \) is odd and in the third quadrant at \( -(x_{y_n}, -y_{x_n}) \) if \( n \) is even. Both foci lie on the circle of radius
\[
r_n = \left( \frac{n}{p} \right)^{\frac{n}{2n-1}} \frac{1}{2} \Gamma \left( \frac{1}{n} \right).
\]
Moreover, for $n \gg 1$, we use the expansion
\[
\Gamma\left(\frac{1}{n}\right) \approx n - \gamma + \frac{1}{2} \left(\frac{\pi^2}{6} + \frac{1}{n} + O\left(\frac{1}{n^2}\right)\right),
\]
where $\gamma = 0.577$ is the Euler-Mascheroni constant, and the limit $(n^{1/n})_{n \to \infty} = 1$, to show that $y_{x_n} \to 0$ and $x_{x_n} \to 1$. Consequently, at ever increasing $n$, these type of spirals are stretched more and more along the [-1,1] segment of the abscissa axis with the two foci at the end points. Concerning the deformed spirals, the coordinates of the foci are related to the undeformed foci through
\[
x^d_{x_n} = R\left[(a + 1)x_n - b y_{x_n}\right],
\]
\[
y^d_{y_n} = R\left[b x_n + (a - 1) y_{x_n}\right],
\]
or in matrix form
\[
\begin{pmatrix}
x^d_{x_n} \\
y^d_{y_n}
\end{pmatrix} = \begin{pmatrix}
\cos \delta & -\sin \delta \\
\sin \delta & \cos \delta
\end{pmatrix} \begin{pmatrix}
x_{x_n} \\
y_{x_n}
\end{pmatrix} + \frac{1}{\sqrt{a^2 + b^2}} \sigma_3 \begin{pmatrix}
x_{x_n} \\
y_{x_n}
\end{pmatrix}
\]
with $\cos \delta = a/\sqrt{a^2 + b^2}$, $\sin \delta = b/\sqrt{a^2 + b^2}$, and $\sigma_3$ the Pauli $z$-spin-flip matrix showing the rotation part and the flip part of the deformation.

To illustrate these results, we choose the odd case $n = 3$. Using the value of $z^*$ for $n = 3$ and $p = \pi$, [23] and [24] are
\[
C_3(z) = z_1 F_2\left[\frac{1}{6}, \frac{2}{3} ; \frac{1}{6} ; -\left(\frac{\pi z^3}{6}\right)^2\right],
\]
\[
S_3(z) = \frac{\pi}{12} z^4 F_2\left[\frac{2}{3}, \frac{5}{3} ; \frac{2}{3} ; -\left(\frac{\pi z^3}{6}\right)^2\right]
\]
with the following coordinates of the foci
\[
x_{x_3} = \pm \frac{\sqrt{3}}{6} \left(\frac{3}{\pi}\right) \frac{1}{3} \Gamma\left(\frac{1}{3}\right) \approx \pm 0.762,
\]
\[
y_{x_3} = \frac{1}{6} \left(\frac{3}{\pi}\right) \frac{1}{3} \Gamma\left(\frac{1}{3}\right) \approx 0.440.
\]

In Fig. 2 we show the graphs of the generalized Fresnel integrals for $n = 3$ for the same values of parameters $a$ and $b$ from Fig. 1. The location of the horizontal asymptotes are the foci for $z > 0$ for the center plot when $a = 0, b = 0$ which corresponds to the standard spiral.

III. THE SUPERSYMMETRIC APPROACH

There are hidden supersymmetric features in the deformations of these spirals generated by the complex Riccati parameter. They can be revealed by writing equation (7) in the factorized form $A^- A^+ v = 0$ using the differential operators given by
\[
A^+ = z^{-\frac{n-1}{2}} \frac{d}{dz} + p z^{-\frac{n-1}{2}} \tan \frac{pz^n}{n},
\]
\[
A^- = \frac{d}{dz} z^{-\frac{n-1}{2}} - p z^{-\frac{n-1}{2}} \tan \frac{pz^n}{n}.
\]
To find out the explicit second-order linear ODE that corresponds to the generalized deformed Cornu spirals, we first write the general Riccati solution [17] in the trigonometric form
\[
y_g(z) = -pz^{n-1} \tan \left(\frac{pz^n}{n} + \phi\right), \quad \theta = R e^{i\phi},
\]
and like in supersymmetric quantum mechanics we employ it in the supersymmetric partner equation $A^- A^+ U = 0$, i.e.,
\[
\left[ z^{-\frac{n-1}{2}} \frac{d}{dz} - z^{-\frac{n-1}{2}} y_g(z) \right] \left[ \frac{d}{dz} z^{-\frac{n-1}{2}} + z^{-\frac{n-1}{2}} y_g(z) \right] U = 0.
\]
We then obtain this equation as
\[
\tilde{V}'' - \frac{n-1}{z} \tilde{V}' + \left(\frac{p z^{n-1}}{2} + \Delta_{\text{Darb}}(z; n, \phi)\right) \tilde{V} = 0,
\]
with the Darboux distortion, which depends parametrically on the phase shift $\phi$, given by
\[
\Delta_{\text{Darb}}(z; n, \phi) = -2 \left(\frac{pn}{n} - \frac{n-1}{4z^n}\right) - (n-1)p z^{n-2} \tan \left(\frac{pz^n}{n} + \phi\right). \quad (38)
\]
In Fig. 3, we display various cases of the parametric Darboux distortions $\Delta_{\text{Darb}}(z; \phi)$ of the deformed Cornu spirals presented in Fig. 1. We notice the negative parabolic envelope as given by the first term in (38) together with the singularities due to the terms containing the tangents for nonzero $z$. The singularities at the origin are due to the $\frac{1}{z^2}$ term except for the cases $\phi = \frac{\pi}{2}, \frac{3\pi}{2}$ when the dominant contribution comes from the cotangent terms. For these values of phase the Darboux distortion simplifies to
\[
\Delta_{\text{Darb}}(z; 0, \pi) = \frac{2}{\pi} - 2 \pi^2 z^4 \frac{\frac{1}{2} \sin(2pz^2)}{\cos^2\left(\frac{pz^n}{n}\right)},
\]
\[
\Delta_{\text{Darb}}(z; \frac{\pi}{2}, \frac{3\pi}{2}) = \frac{2}{\pi} - 2 \pi^2 z^4 \frac{\frac{1}{2} \sin(2pz^2)}{\sin^2\left(\frac{pz^n}{n}\right)},
\]
\[
\Delta_{\text{Darb}}(z; \frac{3\pi}{4}, \frac{7\pi}{4}) = \frac{2}{\pi} - 2 \pi^2 z^4 \frac{\frac{1}{2} \sin(2pz^2)}{1+\sin^2\left(\frac{pz^n}{n}\right)}.
\]
The general solution to (37) can be found by letting $A^- \tilde{V} = \tilde{U}_h$, and first solving the homogeneous equation $A^+ \tilde{U} = 0$ with solution
\[
\tilde{U}_h(z; n, \phi) = b_1 \cos \left(\frac{pz^n}{n} + \phi\right). \quad (39)
\]
We then obtain the general solution $\tilde{V}$ in the form

$$\tilde{V}(z; n, \phi) = \frac{1}{4n} z^{\frac{n-1}{2}} \sec \left( \frac{pz^n}{n} + \phi \right)$$

$$\left[ 2n (b_1 z \cos \phi + 2b_2) - b_1 z \left( e^{i\phi} E_{\frac{2ipzn}{n}} \left( -\frac{2ipzn}{n} \right) + e^{-i\phi} E_{\frac{2ipzn}{n}} \left( \frac{2ipzn}{n} \right) \right) \right] .$$

(40)

where $E_m(\tau)$ is the generalized exponential integral function given by $E_m(\tau) = \int_1^\infty \frac{e^{-\tau s}}{s^m} \, ds$.

**IV. CONCLUSION**

The planar spirals $F_n$, of which the particular case $F_2$ is the Cornu spiral, have been defined in this paper and their parametric Darboux deformations generated through the corresponding general Riccati solution have been studied in detail, with focus on the odd case $F_3$. Geometrically, these deformations are generated by the two independent shifts, $a$ and $b$, along the two orthogonal axes of the plane in which the spirals are plotted. These shifts determine both the deformation of the rolls of the spirals and its global rotation as seen in the plots.

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FIG. 1: The generalized $\mathcal{F}_3$ spiral for different values of the parameters $a$ and $b$. 
FIG. 2: Generalized Fresnel integrals for $n = 3$, for the same values of the parameters $a$ and $b$ given in Fig. [1].
FIG. 3: The Darboux distortion of (7) for various phases of $\phi$. 