Isovector EMC effect explains the NuTeV anomaly

I. C. Cloët,1,* W. Bentz,2 and A. W. Thomas3

1Department of Physics, University of Washington, Seattle, WA 98195-1560, USA
2Department of Physics, School of Science, Tokai University, Hiratsuka-shi, Kanagawa 259-1292, Japan
3Jefferson Lab, 12000 Jefferson Avenue, Newport News, VA 23606, USA and College of William and Mary, Williamsburg, VA 23187, USA

A neutron or proton excess in nuclei leads to an isovector-vector mean-field which, through its coupling to the quarks in a bound nucleon, implies a shift in the quark distributions with respect to the Bjorken scaling variable. We show that this result leads to an additional correction to the NuTeV measurement of \( \sin^2 \Theta_W \). The sign of this correction is largely model independent and acts to reduce their result. Explicit calculation within a covariant and confining Nambu–Jona-Lasinio model predicts that this vector field correction accounts for approximately two-thirds of the NuTeV anomaly. We are therefore led to offer a new interpretation of the NuTeV measurement, namely, that it is further evidence for the medium modification of the bound nucleon wavefunction.

PACS numbers: 24.85.+p, 13.60.Hb, 11.80.Jy, 21.65.Cd

Within relativistic, quark-level models of nuclear structure, the mean scalar and vector fields in the medium generate fundamental changes in the internal structure of bound hadrons. These modifications lead to a good description of the EMC effect in finite nuclei and predict unexplored effects connected with the isovector-vector structure function [1–3]. We show that in nuclei with nuclear data from 56Fe or 208Pb where \( N > Z \), the \( \rho^0 \) field will cause the \( u \)-quark to feel a small additional vector attraction and the \( d \)-quark to feel additional repulsion.

In this Letter we explore the way in which this additional vector field modifies the traditional EMC effect. However, there is an even more important issue which is our main focus. Even though the \( \rho^0 \) mean-field is completely consistent with charge symmetry, the familiar assumption that \( u_p(x) = d_n(x) \) and \( d_p(x) = u_n(x) \) will clearly fail for a nucleon bound in a nucleus with \( N \neq Z \). Therefore correcting for the \( \rho^0 \) field is absolutely critical in a situation where symmetry arguments are essential, such as the use of \( N \neq Z \) nuclear data from \( \nu \) and \( \overline{\nu} \) deep inelastic scattering (DIS) to extract \( \sin^2 \Theta_W \) via the Paschos-Wolfenstein relation [4]. Indeed, we show that the deviation from the naive application of charge symmetry to the \( \nu \) and \( \overline{\nu} \) data on 56Fe naturally explains the famous NuTeV anomaly.

The Paschos-Wolfenstein (PW) ratio is defined by [5]

\[
R_{PW} = \frac{\sigma^{NC}_A - \sigma^{\nu A}_C}{\sigma^{CC}_A - \sigma^{\nu\overline{\nu}A}_C},
\]

where \( A \) represents the target, \( NC \) indicates weak neutral current and \( CC \) weak charged current interaction. Expressing the cross-sections in terms of quark distributions and ignoring heavy flavour contributions, the PW ratio becomes

\[
R_{PW} = \frac{\langle x_A u_A \rangle + \langle x_A d_A \rangle}{\langle x_A d_A \rangle - \langle x_A u_A \rangle},
\]

where \( x_A \) is the Bjorken scaling variable of the nucleus multiplied by \( A \), \( \langle \ldots \rangle \) implies integration over \( x_A \), and \( q_A \equiv q_A - \bar{q}_A \) are the non-singlet quark distributions of the target. Therefore, the quantities in the angle brackets are simply the momentum fractions of the target carried by the valence quarks.

Ignoring quark mass differences and possible electroweak corrections the \( u \)- and \( d \)-quark distributions of an isoscalar target will be identical, and in this limit Eq. (2) becomes

\[
R_{PW} \frac{N=Z}{N/Z} \frac{1}{2} - \sin^2 \Theta_W.
\]

If corrections to Eq. (3) are small the PW ratio provides a unique way to measure the Weinberg angle.

Motivated by Eq. (3) the NuTeV collaboration extracted a value of \( \sin^2 \Theta_W \) from neutrino and anti-neutrino DIS on an iron target [6], finding

\[
\sin^2 \Theta_W = 0.2277 \pm 0.0013(\text{stat.}) \pm 0.0009(\text{ syst.}).
\]

The three-sigma discrepancy between this result and the world average [7], namely \( \sin^2 \Theta_W = 0.2227 \pm 0.0004 \), is the NuTeV anomaly. Some authors have speculated that the NuTeV anomaly supports the existence of physics beyond the Standard Model [8]. However, existing high precision data for other electroweak observables places tight constraints on new physics explanations. Standard Model corrections to the NuTeV result have largely been focused on nucleon charge symmetry violating effects [9] and a non-perturbative strange quark sea [8]. Charge symmetry violation, arising from the \( u \)- and \( d \)-quark mass differences, is probably the best understood and constrained correction and can explain approximately one-third of the NuTeV anomaly [10]. Standard nuclear corrections like Fermi motion and binding are found to be
small [11]. However effects from the medium modification of the bound nucleon, which are now widely accepted as an essential ingredient in explaining the EMC effect [12], have hitherto not been explored in relation to the NuTeV anomaly.

In our approach, presented in Refs. [2, 3, 13], the scalar and vector mean-fields inside a nucleus couple to the quarks in the bound nucleons and self-consistently modify their internal structure. The scalar field renormalizes the constituent quark mass, resulting in effective hadron masses in-medium. The influence of the vector fields on the quark distributions arises from the non-local nature of the quark bilinear in their definition [13]. This leads to a largely model independent result for the modification of the in-medium parton distributions of a bound nucleon by the vector mean-fields [13–15], namely

$$q(x) = \frac{p^+}{p^+ - V^+} \ q_0 \left( \frac{p^+}{p^+ - V^+} x - \frac{V^+_q}{p^+ - V^+} \right). \quad (5)$$

The subscript 0 indicates the absence of vector fields and $p^+$ is the nucleon lightcone plus component of momentum. The quantities $V^+$ and $V^+_q$ are the lightcone plus component of the net vector field felt by the nucleon and a quark of flavour $q$, respectively. Importantly Eq. (5) is consistent with baryon number and momentum conservation, and implies that the mean vector field carries no momentum.

Before embarking on explicit calculations, we first explore the model independent consequences of Eq. (5) for the PW ratio and the subsequent NuTeV measurement of $\sin^2 \Theta_W$. The NuTeV experiment was performed on a predominately $^{56}$Fe target, and therefore isoscalarity corrections need to be applied to the PW ratio before extracting $\sin^2 \Theta_W$. Isoscalarity corrections to Eq. (3) for small isospin asymmetry have the general form

$$\Delta R_{PW} \simeq \left( 1 - \frac{7}{3} \sin^2 \Theta_W \right) \frac{(x_A u_A^+ - x_A d_A^+)}{(x_A u_A^+ + x_A d_A^+)}. \quad (6)$$

where the $Q^2$ dependence of this correction resides completely with $\sin^2 \Theta_W$. NuTeV perform what we term naive isoscalarity corrections, where the neutron excess correction is determined by assuming the target is composed of free nucleons [16]. However there are also isoscalarity corrections from medium effects, in particular from the medium modification of the structure functions of every nucleon in the nucleus, arising from the isovector $\rho^0$ field. For nuclei with $N > Z$ the $\rho^0$ field develops a non-zero expectation value that results in $V_u < V_d$, that is, the $u$-quarks feel less vector repulsion than the $d$-quarks. A direct consequence of this and the transformation given in Eq. (5) is that there must be a small shift in quark momentum from the $u$- to the $d$-quarks. Therefore the momentum fraction $(x_A u_A^+ - x_A d_A^+)$ in Eq. (6) will be negative, even after naive isoscalarity corrections are applied. Correcting for the $\rho^0$ field will therefore have the model independent effect of reducing the NuTeV result for $\sin^2 \Theta_W$. As we shall see, this correction largely explains the NuTeV anomaly.

To determine the nuclear quark distributions we use the Nambu–Jona-Lasinio (NJL) model [17], which is viewed as a low energy chiral effective theory of QCD and is characterized by a 4-fermion contact interaction between the quarks. The NJL model has a long history of success in describing mesons as $q\bar{q}$ bound states [18] and more recently as a self-consistent model for free and in-medium baryons [2, 3, 13, 19]. The original 4-fermion interaction term in the NJL Lagrangian can be decomposed into various $q\bar{q}$ and $qq$ interaction channels via Fierz transformations [20]. The relevant terms of the NJL Lagrangian to this discussion are

$$\mathcal{L} = \bar{\psi} \left( i\gamma \gamma - m \right) \psi + G_\omega \left( \bar{\psi} \gamma^\mu \psi \right)^2 - G_\rho \left( \bar{\psi} \gamma^\mu \gamma^5 \tau^\alpha \psi \right)^2 + G_s \left( \bar{\psi} \gamma_5 C \tau_2 \beta^A \psi \right)^2 + G_a \left( \bar{\psi} \gamma_5 C \tau_3 \beta^A \psi \right)^2, \quad (7)$$

where $\beta^A = \sqrt{\frac{2}{3}} \lambda^A \ (A = 2, 5, 7)$ are the the colour 3 matrices [2], $C = i\gamma_5 \gamma_0$ and $m$ is the current quark mass.

The scalar $qq$ interaction term generates the scalar field, which dynamically generates a constituent quark mass via the gap equation. The vector $\bar{q}q$ interaction terms are used to generate the isoscalar-vector, $\omega$, and isovector-vector, $\rho_0$, mean-fields in-medium. The $qq$ interaction terms give the diquark $t$-matrices whose poles correspond to the scalar and axial-vector diquark masses.

The nucleon vertex function and mass are obtained by solving the homogeneous Faddeev equation for a quark and a diquark, where the static approximation is used to truncate the quark exchange kernel [19]. To regularize the NJL model we choose the proper-time scheme, which enables the removal of unphysical thresholds for nucleon decay into quarks, and hence simulates an important aspect of confinement [21, 22].

To self-consistently determine the strength of the mean scalar and vector fields, an equation of state for nuclear matter is derived from the NJL Lagrangian, Eq. (7), using hadronization techniques [22]. In a mean-field approximation the result for the energy density is [22]

$$\varepsilon = \varepsilon_V - \frac{\omega_0^2}{4G_\omega} - \frac{\rho_0^2}{4G_\rho} + \varepsilon_p + \varepsilon_n, \quad (8)$$

where the vacuum energy $\varepsilon_V$ has the familiar Mexican hat shape and the energies of the protons and neutrons moving through the mean scalar and vector fields are labelled by $\varepsilon_p$ and $\varepsilon_n$, respectively. The corresponding proton and neutron Fermi energies are

$$\varepsilon_{F_\alpha} = E_{F_\alpha} + V_\alpha = \sqrt{M_N^2 + p_{F_\alpha}^2} + 3\omega_0 \pm \rho_0, \quad (9)$$
where $\alpha = p$ or $n$, the plus sign refers to the proton, $M^*_N$ is the in-medium nucleon mass and $p_{F\alpha}$ the nucleon Fermi momentum. Minimizing the effective potential with respect to each vector field gives the following useful relations: $\omega_0 = 6 \alpha_\omega (\rho_p + \rho_n)$ and $\rho_0 = 2 \alpha_g (\rho_p - \rho_n)$, where $\rho_p$ is the proton and $\rho_n$ the neutron density. The vector field experienced by each quark flavour is given by $V_u = \omega_0 + \rho_0$ and $V_d = \omega_0 - \rho_0$.

As explained in Ref. [2], the parameters of the model are determined by standard hadronic properties, and the empirical saturation energy and density of symmetric nuclear matter. The new feature of this work is the $\rho^0$ field, where $\alpha_g$ is determined by the empirical symmetry energy of nuclear matter, namely $a_4 = 32$ MeV, giving $G_\rho = 14.2$ GeV$^{-2}$.

Details of our results for the free and $N \simeq Z$ in-medium parton distributions are given in Refs. [2, 3, 19]. For in-medium isospin dependent parton distributions our produce is as follows: Effects from the scalar mean-field are included by replacing the free masses with the effective masses in the expressions for the free parton distributions discussed in Ref. [19]. For asymmetric nuclear matter ($N \neq Z$) the proton and neutron Fermi momentum will differ and therefore so will their Fermi smearing functions. To include the nucleon Fermi motion, the quark distributions modified by the scalar field are convoluted with the appropriate Fermi smearing function, namely

$$f_{\alpha 0}(y_A) = \frac{1}{M_N} \left( \frac{\hat{M}_N}{M_N} \right)^3 \left[ \left( \frac{p_{F\alpha}}{M_N} \right)^2 - \left( \frac{E_{F\alpha}}{M_N} - y_A \right)^2 \right].$$  \hspace{1cm} (10)$$

where $N_p = Z$, $N_n = N$ and $\hat{M}_N = \frac{Z}{A} E_{Fp} + \frac{N}{A} E_{Fn}$. Vector field effects can be included in Eq. (10) by the substitutions $E_{F\alpha} \rightarrow \varepsilon_{F\alpha}$ and $\hat{M}_N \rightarrow \hat{M}_N = \frac{Z}{A} \varepsilon_{Fp} + \frac{N}{A} \varepsilon_{Fn}$.

Our final result for the infinite asymmetric nuclear matter quark distributions, which includes vector field effects on both the quark distributions in the bound nucleon and

$$q_A(x_A) = \frac{M_N}{\bar{M}_N} \left( \frac{M_N}{\bar{M}_N} x_A - \frac{V_q}{\bar{M}_N} \right).$$  \hspace{1cm} (11)$$

The subscript $A0$ indicates a distribution which includes effects from Fermi motion and the scalar mean-field. The distributions calculated in this way are then evolved [23] from the model scale, $Q_0^2 = 0.16$ GeV$^2$, to an appropriate $Q^2$ for comparison with experimental data.

The EMC effect is defined by the ratio

$$R = \frac{F_{2A}}{F_{2A}^{\text{naive}}} = \frac{F_{2A}}{Z F_{2p} + N F_{2n}} \simeq \frac{4 u_A + d_A}{4 u_f + d_f},$$  \hspace{1cm} (12)$$

where $q_A$ are the quark distributions of the target and $q_f$ are the distributions of the target if it was composed of free nucleons. Results for the isospin dependence of the EMC effect are given in Figs. 1 and 2.

Fig. 1 illustrates the EMC effect for proton rich matter, where we find a decreasing effect as $Z/N$ increases. An intuitive understanding of this result may be obtained by realizing that it is a consequence of binding effects at the quark level. For $Z/N > 1$ the $\rho_0$ field is positive, which means $V_u > V_d$ and hence the $u$-quarks are less bound than the $d$-quarks. Therefore the $u$-quark distribution becomes less modified while medium modification of the $d$-quark distribution is enhanced. Since the EMC effect is dominated by the $u$-quarks it decreases. The isospin dependence of the EMC effect for nuclear matter with $Z/N < 1$ is given in Fig. 2. Here the medium modification of the $u$-quark distribution is enhanced, while the $d$-quark distribution is modified less by the medium. Since the EMC ratio is initially dominated by the $u$-quarks the EMC effect first increases as $Z/N$ decreases from one. However, eventually the $d$-quark distribution dominates the ratio and at this stage the EMC effect begins to decrease in the valence quark region. We find a maximal EMC effect for $Z/N \simeq 0.6$, which is slightly less than the proton-neutron ratio in lead. This isospin dependence is

FIG. 1: Isospin dependence of the EMC effect for proton-neutron ratios greater than one. The data is from Ref. [24] and corresponds to $N = Z$ nuclear matter.

FIG. 2: Isospin dependence of the EMC effect for proton-neutron ratios less than one.
clearly an important factor in understanding the A dependence of the EMC effect, even after standard neutron excess corrections are applied.

Now we turn to the consequences of the isospin dependence of the EMC effect for the NuTeV measurement of $\sin^2\Theta_W$. The NuTeV experiment was performed on an iron target, which, because of impurities had a neutron excess of 5.74% [6]. Choosing our iron target, which, because of impurities had a neutron excess corrections are applied.

Using the Standard Model value for the Weinberg angle $\sin^2\Theta_W$, we obtain $\Delta R_{PW} = 0.0139$. If we break this result into three separate isoscalarity corrections, by using Eq. (6) and the various stages of the medium modified quark distributions, we find

$$\Delta R_{PW} = \Delta R_{PW}^{\text{naive}} + \Delta R_{PW}^{\text{Fermi}} + \Delta R_{PW}^{\rho}$$

$$= -0.0017 + 0.0004 + 0.0028.$$  \hfill (13)

Higher order corrections to Eq. (6) do not change this result. The NuTeV analysis includes the naive isoscalarity correction [15] but is missing the medium correction of $-0.0032$ [25]. This new correction accounts for two-thirds of the NuTeV anomaly. If we also include the well constrained charge symmetry violation (CSV) correction, $\Delta R_{PW}^{\text{CSV}} = -0.0017$ [10], which originates from the quark mass differences, we have a total correction of $\Delta R_{PW}^{\text{medium}} + \Delta R_{PW}^{\text{CSV}} = -0.0049$. The combined correction accounts for the NuTeV anomaly [27].

Since CSV and medium modification corrections explain the discrepancy between the NuTeV result and the Standard Model, we propose that this NuTeV measurement provides strong evidence that the nucleon is modified by the nuclear medium, and should not be interpreted as an indication of physics beyond the Standard Model. In our opinion this conclusion is equally profound since it may have fundamental consequences for our understanding of traditional nuclear physics. We stress that the physics presented in this paper, in particular the effects of the $\rho^0$ mean-field, are consistent with existing data [28], but can strongly influence other observables. For example, the $\rho^0$ field gives rise to a strong flavour dependence of the EMC effect, and experimental proposals have been submitted at Jefferson Lab to look for such effects.

I. C. thanks Jerry Miller for helpful discussions. This work was supported by the the U.S. Department of Energy under Grant No. DEFG03-97ER4014 and by Contract No. DE-AC05-06OR23177, under which Jefferson Science Associates, LLC operates Jefferson Laboratory and by the Grant in Aid for Scientific Research of the Japanese Ministry of Education, Culture, Sports, Science and Technology, Project No. C-19540306.

[1] K. Saito, A. Michels and A. W. Thomas, Phys. Rev. C 46, 2149 (1992).
[2] J. R. Smith and G. A. Miller, Phys. Rev. C 72, 022203 (2005).
[3] I. C. Cloët, W. Bentz and A. W. Thomas, Phys. Lett. B 642, 210 (2006).
[4] I. C. Cloët, W. Bentz and A. W. Thomas, Phys. Rev. Lett. 95, 052302 (2005).
[5] E. A. Paschos and L. Wolfenstein, Phys. Rev. D 7, 91 (1973).
[6] G. P. Zeller et al., Phys. Rev. Lett. 88, 091802 (2002) [Erratum-ibid. 90, 239902 (2003)].
[7] D. Abbaneco et al., arXiv:hep-ex/0112021.
[8] S. Davidson, S. Forte, P. Gambino, N. Rius and A. Strumia, JHEP 0202, 037 (2002).
[9] E. Sather, Phys. Lett. B 274, 433 (1992).
[10] J. T. Londergan and A. W. Thomas, Phys. Lett. B 558, 132 (2003).
[11] S. A. Kulagin, Phys. Rev. D 67, 09301 (2003).
[12] J. R. Smith and G. A. Miller, Phys. Rev. C 65, 055206 (2002).
[13] D. F. Geesaman, K. Saito and A. W. Thomas, Ann. Rev. Nucl. Part. Sci. 45, 337 (1995).
[14] H. Mineo, W. Bentz, N. Ishii, A. W. Thomas and K. Yazaki, Nucl. Phys. A 735, 482 (2004).
[15] F. M. Steffen, K. Tsushima, A. W. Thomas and K. Saito, Phys. Lett. B 447, 233 (1999).
[16] W. Detmold, G. A. Miller and J. R. Smith, Phys. Rev. C 73, 015204 (2006).
[17] NuTeV do not directly utilize Eq. (6) for their naive isoscalarity correction, because in their case, details of this correction depend explicitly on the Monte-Carlo routine used to analyze their data.
[18] Y. Nambu and G. Jona-Lasinio, Phys. Rev. 122, 345 (1961).
[19] T. Hutsuda and T. Kunihiro, Phys. Rept. 247, 221 (1994).
[20] I. C. Cloët, W. Bentz and A. W. Thomas, Phys. Lett. B 621, 246 (2005).
[21] N. Ishii, W. Bentz and K. Yazaki, Nucl. Phys. A 587, 617 (1995).
[22] D. Ebert, T. Feldmann and H. Reinhardt, Phys. Lett. B 388, 154 (1996).
[23] W. Bentz and A. W. Thomas, Nucl. Phys. A 696, 138 (2001).
[24] M. Miyama and S. Kumaano, Comput. Phys. Commun. 94, 185 (1996).
[25] I. Sick and D. Day, Phys. Lett. B 274, 16 (1992).
[26] Using the NuTeV CSV functional [26] for a bound proton and neutron we obtain $\Delta R_{PW}^\rho = 0.0021$.
[27] The NuTeV result for $R_{PW}$ after naive isoscalarity correction is 0.2723 which differs from $\frac{1}{2} - \sin^2\Theta_W = 0.2773$ by the amount $-0.005$.
[28] S. Kumano, Phys. Rev. D 66, 111301 (2002).

* Electronic address: icloet@phys.washington.edu