Analytic normal mode frequencies for \( N \) identical particles: The microscopic dynamics underlying the emergence and stability of excitation gaps from BCS to unitarity

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The frequencies of the analytic normal modes for \( N \) identical particles are studied as a function of system parameters from the weakly interacting BCS regime to the strongly interacting unitary regime. The normal modes were obtained previously from a first-order \( L = 0 \) group theoretic solution of a three-dimensional Hamiltonian with a general two-body interaction for confined, identical particles. In a precursor to this study, the collective behavior of these normal modes was investigated as a function of \( N \) from few-body systems to many-body systems analyzing the contribution of individual particles to the collective macroscopic motions. A specific case, the Hamiltonian for Fermi gases in the unitary regime was studied in more detail. This regime is known to support collective behavior in the form of superfluidity and has previously been successfully described using normal modes. Two phenomena that could sustain the emergence and stability of superfluid behavior were revealed, including the behavior of the normal mode frequencies as \( N \) increases. In this paper, I focus on a more detailed analysis of these analytic frequencies, extending my investigation to include Hamiltonians with a range of interparticle interaction strengths from the BCS regime to the unitary regime and analyzing the microscopic dynamics that lead to large gaps at unitarity. The results of the current study suggest that in regimes where higher-order effects are small, normal modes can be used to describe the physics of superfluidity from the weakly interacting BCS regime with the emergence of small excitation gaps to unitarity with its large gaps, and can offer insight into a possible microscopic understanding of the behavior at unitarity. This approach could thus offer an alternative to the two-body pairing models commonly used to describe superfluidity along this transition.

I. INTRODUCTION

The evolution of collective behavior in the form of superfluidity for systems of ultracold gaseous fermions from the BCS regime to unitarity has been studied intensely during the last two decades since this transition was first explored in the laboratory\[1–9\]. Typically, theoretical methods assume that fermions are pairing into loosely-bound Cooper pairs in the BCS regime to explain the emergence of superfluid behavior\[11–16\]. As the inter-particle interaction increases toward the unitary regime in ultracold fermion gases, these atomic Cooper pairs decrease in size, ultimately forming diatomic molecules that condense in the BEC regime on the other side of unitarity. In materials supporting superconductivity, the emergence of Cooper pairs of electrons is mediated by interactions with phonons in the underlying material creating a weak attraction that can bind two electrons at long distances\[10–14\]. In an ultracold Fermi gas, neutral atoms are assumed to pair at large distances due to weak interactions when a Feshbach resonance is tuned far from resonance. In this study, I will present an alternative possibility to describe the transition from the weakly interacting BCS regime to the strong interactions of the unitary regime that does not assume that individual fermions are forming pairs. Instead, the proposed model assumes many-body pairing exhibited through normal modes to model the physics i.e. synchronized, collisionless motion of the particles that makes it impos-

sible to know which fermion is paired with any other fermion. Normal mode functions naturally provide simple, coherent macroscopic wave functions with phase coherence that is maintained over the entire ensemble. The excitations between modes define “quasiparticles” of the macroscopic quantum system. I will show how the lowest two normal mode frequencies relevant for ultracold systems, a phonon mode and a particle-hole excitation mode, exhibit, as expected, an extremely small excitation gap in the weakly interacting BCS regime which widens as the interaction increases reaching a maximum in the unitary regime. The physics of this model from BCS to unitarity is a precursor to the physical pairing of atoms in real space that eventually forms diatomic molecules in the BEC region.

Normal mode behavior is ubiquitous in the universe, occurring at all scales from the vibration of crystals\[17\] to the oscillation of rotating stars\[18\]. This universal dynamic reflects the widespread appearance of vibrational motions in nature\[17–29\] which can often be coupled into the simple collective motions of normal modes. These collective motions depend on the interparticle correlations of the system and thus incorporate many-body effects into simple, dynamic motion. If higher order effects are small, these collective motions are eigenfunctions of an approximate Hamiltonian and acquire some degree of stability as a function of time; thus, a system in a single normal mode will tend to stay in that mode unless perturbed. Normal modes manifest the symmetry of this underlying
The approximate Hamiltonian with the possibility of offering analytic solutions to many-body problems and a clear physical picture of the microscopic dynamics underlying diverse phenomena.

In an earlier paper, I studied the character of five types of normal modes previously derived as the $L = 0$ first-order analytic solutions of a general Hamiltonian for confined, identical particles [33] using a perturbation formalism called symmetry-invariant perturbation theory (SPT) [34, 35]. Using the simple analytic expressions for the $N$-body normal mode coordinates, I investigated the evolution of their physical character as a function of $N$, from few-body to many-body, and examined the motion of the individual particles as they contributed to the collective motion. Some general observations were made based on symmetry considerations and then their behavior was analyzed for a specific case, the Hamiltonian of a confined system of fermions in the unitary regime which is known to support superfluid behavior. This study found that the behavior expected for few-body systems, which have the well known motions of molecular equivalents such as ammonia and methane (symmetric stretch, symmetric bend, antisymmetric stretch, antisymmetric bend and the opening and closing of alternative interparticle angles), evolves smoothly as $N$ increases to the collective motions expected for large $N$ ensembles (breathing, center of mass, particle-hole radial and angular excitations and phonon). Furthermore, the transition from few-body behavior to large $N$ behavior was found to occur at quite low values of $N$ ($N \approx 10$). This change in character from small $N$ to large $N$ is dictated by fairly simple analytic forms that nonetheless incorporate the intricate interplay of individual particles as they contribute to the macroscopic motion. The evolution of behavior was found to be determined primarily from the symmetry structure of the Hamiltonian, and thus could be applicable to diverse phenomena at different scales if the same symmetry is present or dominates.

The SPT formalism was developed initially for systems of bosons [31, 37] and more recently applied to fermions [38, 41]. This formalism has also been tested against an exactly solvable model problem of harmonically interacting particles under harmonic confinement [32, 37, 38, 41]. Exact agreement was found (to ten or more digits of accuracy) for the wave function with the exact analytic wave function obtained in an independent solution, confirming this general formalism for a fully interacting, three-dimensional $N$-body system [36] and verifying the analytic expressions for the normal mode frequencies and coordinates.

In the fermion studies, the numerically demanding determination of explicitly antisymmetrized wave functions is avoided by using specific assignments of normal mode occupations to enforce the Pauli principle at first order “on paper” [38, 41]. Ground [38] and excited state [41] beyond-mean-field energies and their degeneracies have been determined allowing the construction of a partition function [40, 41] and the determination of thermodynamic quantities [40, 41]. Constructing the partition function required a large number of excited states from the infinite spectrum of equally spaced states, chosen specifically to comply with the enforcement of the Pauli principle, thus connecting the Pauli principle to many-body interaction dynamics through the normal modes.

The study of the thermodynamic behavior of ultracold fermions in the unitary regime obtained quite good agreement with experimental data for the energy, entropy and heat capacity [41]. Two normal modes, selected by the Pauli principle, were found to play a role in creating and stabilizing the superfluid behavior at low temperatures, a phonon mode at ultralow temperatures and a single particle excitation mode, i.e. a particle-hole excitation, as the temperature increases. This single-particle excitation has a much higher frequency and creates a gap that stabilizes the superfluid behavior. This normal mode description offers an interesting alternative to two-body pairing correlation models commonly used to describe superfluid regimes.

The good agreement with experiment for thermodynamic quantities increased the interest in investigating the physical character of these states, originally obtained simply as a complete basis at first order, since they offer the possibility of acquiring physical intuition into the dynamics of collective motion [32] and insight into the universal behavior of the unitary regime.

My previous study of the five types of normal mode coordinates looked at the evolution of behavior as a function of $N$ for only one specific case, the strongly interacting unitary regime. This region was chosen because of the current experimental and theoretical interest in this regime which is known to exhibit universal behavior and to support superfluidity with large excitation gaps. My analysis of the $N$ dependence for the unitary Hamiltonian revealed two phenomena that have the potential to support the creation and stabilization of collective behavior. First the mixing of radial and angular behavior in the normal modes is found to limit to pure radial or pure angular behavior for very large (or very small) $N$. This results in symmetry coordinates that are eigenfunctions of an approximate Hamiltonian governing the physics of the unitary regime, thus acquiring some amount of stability if the symmetry is unperturbed. Second, for low values of $N$, the five types of normal mode frequencies start out closer in value, but as $N$ increases these frequencies spread out creating large gaps between the values of these five frequencies. These gaps could provide the stability for superfluid behavior if mechanisms to prevent the transfer of energy to other modes exist (such as low temperatures) or could be engineered.

In this paper, I now extend my investigation to regimes other than the unitary regime, studying the evolution of the frequencies as a function of the interparticle interaction strength, $V_0$, from the BCS regime to unitarity. I will focus on larger values of $N$ relevant to experimental investigations of this transition. For this study, I have scaled the value of $V_0$ so $V_0 = 1.0$ corresponds to the uni-
tary regime which has an infinite scattering length. The BCS regime is loosely defined as having extremely weak interparticle interactions, e.g., $V_0 \approx 10^{-6}$. This potential is briefly defined in Section IV with a more detailed description in Appendix A in Ref. [30].

The analytic expressions for the five types of frequencies have a complicated dependence on $V_0$, both explicitly and implicitly through other variables in the formalism that depend on $V_0$. The goal is to determine the interplay of various terms in the Hamiltonian as they respond to the increase in the interparticle interaction and affect the value of the frequencies. This understanding should offer insight into the microscopic dynamics that leads to large gaps as unitarity is approached and offers the possibility of fine tuning the system parameters to control the appearance and stability of excitation gaps.

In the remainder of this Section, I summarize the results of my investigation and state my conclusions.

Section IV gives a brief review of the SPT method including the derivation of the symmetry coordinates, the normal coordinates and the normal mode frequencies, establishing the necessary notation.

Section V looks at the mixing coefficients (defined in Appendix A) that determine the radial/angular mixing of the symmetry coordinates to form a normal coordinate, extending my earlier study in the unitary regime to regimes with weak interactions. Similar behavior is found for all strengths of the interparticle interaction. Specifically, the character of the normal modes $q^i$ evolves to almost purely radial or purely angular as $N$ increases, with very little mixing of the symmetry coordinates, confirming that this phenomena is driven by dynamics other than the universal behavior of a system at unitarity.

This negligible mixing is reflected in the character of the normal mode frequencies, which can be appropriately labelled as radial frequencies or angular frequencies across the entire transition. It also has implications for the ability to tune these frequencies as well as the stability of collective behavior since the symmetry coordinates are eigenfunctions of an approximate underlying Hamiltonian.

In Section VI and Appendices C, D and E I analyze the analytic expressions for the five types of frequencies in terms of their dependence on $V_0$, confirming that the frequencies can be characterized into two types: radial frequencies that have a strong dependence on $V_0$, and angular frequencies with a weaker dependence on $V_0$ that evolve to stable limits insensitive to changes in $V_0$.

Section VI discusses the behavior of the frequencies and the emergence of stable gaps as a function of $V_0$ from the BCS regime to the unitary regime. This analysis shows the emergence of excitations gaps that increase as $V_0$ increases. For extremely weak interactions, the five frequencies converge to identical values at twice the trap frequency which results in infinitesimally small excitation gaps. As $V_0$ increases, the frequencies begin to spread out creating gaps that reach a maximum at unitarity with the angular frequencies approaching stable limits while the radial frequencies continue to gradually change. (These limits are derived in detail in Appendices F and G.) For ultracold systems, the lowest two frequencies are of interest, the phonon frequency which tends to extremely small values and the angular particle-hole frequency which limits to the trap frequency at unitarity. This sets up an excitation gap that stabilizes as the unitary limit is approached. As $N$ increases, this behavior is stabilized at smaller and smaller values of $V_0$. Since $V_0$ appears as a parameter in the analytical expressions for the frequencies, the evolution of these frequencies can be studied as a function of the interparticle interaction without intensive numerical work.

Finally in Section VI the microscopic dynamics underpinning the stable limits of the angular frequencies and the emergence of excitation gaps that could support superfluidity are investigated from two perspectives. First, the relative contributions of various Hamiltonian terms to the evolving, analytic frequencies are tracked as $V_0$ changes. Then, the motion of the individual particles in the corresponding normal mode coordinate is studied to understand how the ensemble is rearranging on a microscopic level as interactions turn on and collective behavior emerges. The excitation gap relevant for ultracold Fermi gases limits to the trap frequency at unitarity, setting up a spectrum of evenly spaced levels identical to the spectrum of the non-interacting regime. This results in dynamics independent of microscopic details of the underlying interactions consistent with the universal behavior of the unitary regime.

In summary, this study of the evolution of the normal mode frequencies from the first-order solution of the SPT equations for confined systems of identical particles as a function of $V_0$ suggests that these normal modes are able to describe the physics of superfluidity from the weakly-interacting BCS regime to the universal behavior of unitarity and to offer a view of the microscopic dynamics without the assumption of two-particle pairing.

II. SYMMETRY-ININVARIANT PERTURBATION THEORY: THE DERIVATION OF THE NORMAL MODES AND THEIR FREQUENCIES

In this Section, I summarize the development of SPT theory and the previous derivation of the normal modes and their frequencies that was presented in Ref. [31-32], indicating the notation required in Sections IV-VI for the analysis of the frequencies.

The normal modes are the *exact* solutions at first order in inverse dimensionality of a first principle, perturbation, many-body formalism called symmetry-invariant perturbation theory (SPT). This formalism uses group theory to solve a fully interacting, many-body, three-dimensional Hamiltonian with a confining potential and an arbitrary interaction potential[32-33]. Using the symmetry of the symmetric group at large dimension[31], this group theoretic approach successfully rearranges...
the many-body work at each perturbation order so that an exact solution can, in principle, be obtained non-numerically, order-by-order, using group theory and graphical techniques. Specifically, the numerical work is rearranged into analytic building blocks resulting in a formulation with a complexity that does not scale with $N$. Group theory is used to partition the $N$ scaling problem away from the interaction dynamics, allowing the $N$ scaling to be treated as a separate mathematical problem (cf. the Wigner-Eckart theorem). The exponential scaling in complexity is shifted from a dependence on the number of particles, $N$, to a dependence on the order of the perturbation series. Exact first-order results that contain beyond-mean-field effects for all values of $N$ can now be obtained from a single calculation, but determining higher order results becomes exponentially difficult. To minimize the work needed for new calculations, the analytic building blocks have been calculated and stored. Strongly interacting systems can be studied since the perturbation does not involve the strength of the interaction.

The Schrödinger equation in $D$ dimensions is defined in Cartesian coordinates for $N$ interacting particles by:

$$H\Psi = \left[ \sum_{i=1}^{N} h_i + \sum_{i=1}^{N} \sum_{j=i+1}^{N} g_{ij} \right] \Psi = E\Psi,$$

(1)

$$h_i = -\frac{\hbar^2}{2m_i} \sum_{\nu=1}^{D} \frac{\partial^2}{\partial x_{i\nu}^2} + V_{\text{conf}} \left( \sqrt{\sum_{\nu=1}^{D} x_{i\nu}^2} \right),$$

$$g_{ij} = V_{\text{int}} \left( \sqrt{\sum_{\nu=1}^{D} (x_{i\nu} - x_{j\nu})^2} \right),$$

(2)

where $h_i$ is the single-particle Hamiltonian, $g_{ij}$ is a two-body interaction potential, $x_{i\nu}$ is the $i^{th}$ Cartesian component of the $i^{th}$ particle, and $V_{\text{conf}}$ is a spherically-symmetric confining potential. The Schrödinger equation is transformed from Cartesian coordinates to internal coordinates using:

$$r_i = \sqrt{\sum_{\nu=1}^{D} x_{i\nu}^2}, \quad (1 \leq i \leq N),$$

$$\gamma_{ij} = \cos(\theta_{ij}) = \left( \sum_{\nu=1}^{D} x_{i\nu}x_{j\nu} \right)/r_ir_j,$$

(3)

$$(1 \leq i < j \leq N),$$

which are the $D$-dimensional scalar radii $r_i$ of the $N$ particles from the center of the confining potential and the cosines $\gamma_{ij}$ of the $N(N-1)/2$ angles between the radial vectors.

The first-order derivatives are removed using a similarity transformation, and dimensionally-scaled oscillator units are defined with a scale factor, $\kappa(D)$, that regularizes the large-dimension limit of the Schrödinger equation. Substituting the scaled variables, $\vec{r}_i = r_i/\kappa(D)$, $\vec{E} = \kappa(D)E$ and $\vec{H} = \kappa(D)H$, into the similarity-transformed Schrödinger equation gives:

$$\vec{H}\Phi = \left[ \delta^2 \vec{T} + \vec{U} + \vec{V}_{\text{conf}} + \vec{V}_{\text{int}} \right] \Phi = \vec{E}\Phi,$$

(4)

where

$$\vec{T} = \sum_{i=1}^{N} \left( -\frac{1}{2} \frac{\partial^2}{\partial \vec{r}_i^2} \right) - \frac{1}{2\kappa^2} \sum_{j \neq i} \sum_{k \neq i} \frac{\partial}{\partial \gamma_{ij}} \left( \gamma_{jk} - \gamma_{ij} \gamma_{ik} \right) \frac{\partial}{\partial \gamma_{ij}},$$

$$\vec{U} = \sum_{i=1}^{N} \left( \frac{\delta^2 N(N-2) + (1 - \delta(N+1))^2 (\Gamma^{(i)})}{8\kappa^2} \right),$$

$$\vec{V}_{\text{conf}} = \sum_{i=1}^{N} \frac{1}{2} \vec{r}_i^2,$$

$$\vec{V}_{\text{int}} = \frac{\vec{V}_0}{1 - 3b\delta} \sum_{i=1}^{N-1} \sum_{j=i+1}^{N} (1 - \tanh \Theta_{ij}),$$

(5)

$$\vec{V}_{\text{eff}}(\vec{r},\gamma;\delta) = \sum_{i=1}^{N} (\vec{U}(\vec{r}_i;\delta) + \vec{V}_{\text{conf}}(\vec{r}_i;\delta))$$

$$+ \sum_{i=1}^{N-1} \sum_{j=i+1}^{N} \vec{V}_{\text{int}}(\vec{r}_i,\gamma_{ij};\delta),$$

(6)

The minimum of this effective potential yields an infinite-dimensional maximally-symmetric structure with all radii, $\vec{r}_i$, and angle cosines, $\gamma_{ij}$, of the particles equal, i.e. when $D \to \infty$, $\vec{r}_i = \vec{r}_\infty$ (1 $\leq i \leq N$) and $\gamma_{ij} = \gamma_\infty$ (1 $\leq i < j \leq N$). The values of these parameters are determined by two minimum conditions:

$$\left( \frac{\partial \vec{V}_{\text{eff}}}{\partial \vec{r}_i} \right)_{\infty} = 0,$$

$$\left( \frac{\partial \vec{V}_{\text{eff}}}{\partial \gamma_{ij}} \right)_{\infty} = 0.$$

(7)
Substituting the above definition of $\tilde{V}_{\text{eff}}$, two equations in $\tilde{r}_\infty$ and $\gamma_\infty$ are obtained which yield:

$$\tilde{r}_\infty = \frac{1}{\sqrt{2}\sqrt{1 + (N-1)\gamma_\infty}}, \quad (13)$$

while $\gamma_\infty$ can be solved from the transcendental equation:

$$\frac{\gamma_\infty (2 + (N-2)\gamma_\infty)}{(1 - \gamma_\infty)^{3/2}\sqrt{1 + (N - 1)\gamma_\infty}} + \tilde{V}_\text{0} \sech^2 (\Theta_\infty) \Theta_\infty' = 0. \quad (14)$$

In the large-$D$ limit ($\delta \to 0$), the argument $\Theta_{ij}$ becomes

$$\Theta_\infty = \Theta_{ij} \bigg|_\infty = \bar{\alpha}_0 \left( \sqrt{1 - \gamma_\infty \tilde{r}_\infty - \bar{\alpha}} \right), \quad (15)$$

The zeroth-order energy at this minimum, $\tilde{E}_\infty = \tilde{V}_{\text{eff}}(\tilde{r}_\infty)$ provides the starting point for the $1/D$ expansion. A position vector of the $N(N+1)/2$ internal coordinates is defined as:

$$\mathbf{y} = \begin{pmatrix} \tilde{r} \\ \gamma \end{pmatrix}, \quad \text{where} \quad \gamma = \begin{pmatrix} \gamma_{12} \\ \gamma_{13} \\ \gamma_{23} \\ \gamma_{14} \\ \gamma_{24} \\ \gamma_{34} \\ \gamma_{15} \\ \gamma_{25} \\ \gamma_{35} \\ \gamma_{45} \\ \gamma_{N-2,N} \\ \gamma_{N-1,N} \end{pmatrix}, \quad (16)$$

and $\tilde{r} = \begin{pmatrix} \tilde{r}_1 \\ \tilde{r}_2 \\ \vdots \\ \tilde{r}_N \end{pmatrix}$. These

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The substitutions: $\tilde{r}_i = \tilde{r}_\infty + \delta^{1/2} \tilde{r}_i'$ and $\gamma_{ij} = \gamma_\infty + \delta^{1/2} \gamma_{ij}'$ set up a power series in $\delta^{1/2}$ about the $D \to \infty$ symmetric minimum.

The first-order, $\delta = 1/D$, equation is a harmonic problem, which is solved exactly and analytically by obtaining the $N$-body normal modes of the system. The first-order Hamiltonian, $\tilde{H}_1$, is defined in terms of constant matrices $G$ and $F$ that are evaluated at the large dimension limit:

$$\tilde{H}_1 = -\frac{1}{2} \partial_y^T G \partial_y + \frac{1}{2} \mathbf{y}^T F \mathbf{y}' + v_0 \quad (17)$$

where $G$ involves kinetic energy terms, $F$ involves derivatives of the effective potential, and $v_0$ is a constant $[31]$. The FG matrix method[20] is used to obtain the normal-mode frequencies and the harmonic-order energy correction[31]. A review of the FG matrix method is presented in Appendix A of Ref. [31]. $N(N+1)/2$ frequencies, $\bar{\omega}$, are obtained from the roots of the FG equation, however only five roots are distinct due to the large degeneracy of the frequencies reflecting the very high degree of symmetry manifested in the $F$, $G$, and FG matrices. The elements of these matrices are evaluated for the large dimension, maximally-symmetric structure with a single value for all radii $\tilde{r}_\infty$ and angle cosines, $\gamma_\infty$. Thus these matrices are invariant under the $N!$ operations of particle interchanges effected by the symmetric group, $S_N$, which results in the highly degenerate eigenvalues. These matrices do not connect subspaces belonging to different irreducible representations of $S_N$[47, 48], thus the normal coordinates must transform under irreducible representations of $S_N$.

There are a total of five irreducible representations: two 1-dimensional irreducible representations, one radial and one angular, labelled by the partition $[N]$, two $(N - 1)$-dimensional irreducible representations, one radial and one angular, labelled by the partition $[N-1, 1]$, and one angular $N(N-3)/2$-dimensional irreducible representation labelled by the partition $[N-2, 2]$. These representations are given shorthand labels: $0^-$, $0^+$, $1^-$, $1^+$ and $2$ respectively, (see Refs. [32, 33]). Thus the energy through first-order in $\delta = 1/D$ can be written in terms of the five distinct normal mode frequencies[31, 49] as:

$$\mathbf{E} = \mathbf{E}_\infty + \delta \left[ \sum_{\mu=(0^-, 1^-, 1^+, 2)} (n_\mu + \frac{1}{2} d_\mu) \bar{\omega}_\mu + v_0 \right]. \quad (18)$$

where $\mu$ is a label which runs over the five types of normal modes, $0^-$, $0^+$, $1^-$, $1^+$ and $2$ (irrespective of the particle number, see Ref. [31] and Ref.[15] in [32]), $n_\mu$ is the total quanta in the normal mode with frequency $\bar{\omega}_\mu$; and $v_0$ is a constant (defined in Ref. [31], Eq.(125)).

The multiplicities of the five roots are: $d_{0^-} = 1$, $d_{0^+} = 1$, $d_{1^-} = N - 1$, $d_{1^+} = N - 1$, $d_2 = N(N-3)/2$.

We define the symmetry coordinate vector, $S$, as:

$$S = \begin{pmatrix} S_{\gamma_{12}}^{[N]} \\ S_{\gamma_{13}}^{[N]} \\ S_{\gamma_{23}}^{[N]} \\ S_{\gamma_{14}}^{[N-1, 1]} \\ S_{\gamma_{24}}^{[N-1, 1]} \\ S_{\gamma_{34}}^{[N-2, 2]} \\ S_{\gamma_{15}}^{[N-1, 1]} \\ S_{\gamma_{25}}^{[N-1, 1]} \\ S_{\gamma_{35}}^{[N-2, 2]} \\ S_{\gamma_{45}}^{[N-2, 2]} \end{pmatrix} = \begin{pmatrix} W_{\gamma_{12}}^{[N]} \tilde{r}' \\ W_{\gamma_{13}}^{[N]} \gamma' \\ W_{\gamma_{14}}^{[N-1, 1]} \tilde{r}' \\ W_{\gamma_{15}}^{[N-1, 1]} \gamma' \\ W_{\gamma_{24}}^{[N-2, 2]} \gamma' \end{pmatrix}, \quad (19)$$

where the $W_{\gamma_{ij}}^{[a]}$ and the $W_{\gamma_{ij}}^{[a]}$ are transformation matrices. This is shown in Ref. [32] using the theory of group characters to decompose $\tilde{r}'$ and $\gamma'$ into basis functions that transform under these five irreducible representations of $S_N$.

The FG method is applied using these symmetry coordinates to determine the eigenvalues, $\lambda_\alpha = \bar{\omega}_\alpha$, frequencies, $\bar{\omega}_\alpha$, and normal modes, $q_{\alpha}^{[a]}$, of the system:

$$q_{\alpha}^{[N]} = c_{\alpha}^{[N]} \left( \cos \theta_{\alpha}^{[N]} [S_{\gamma_{12}}^{[N]}] + \sin \theta_{\alpha}^{[N]} [S_{\gamma_{13}}^{[N]}] \right) \quad (20)$$

$$q_{\alpha}^{[N-1, 1]} = c_{\alpha}^{[N-1, 1]} \left( \cos \theta_{\alpha}^{[N-1, 1]} [S_{\gamma_{14}}^{[N-1, 1]}] \right) + \sin \theta_{\alpha}^{[N-1, 1]} [S_{\gamma_{15}}^{[N-1, 1]}] \quad (21)$$
for the $\alpha = [N]$ and $[N-1, 1]$ sectors, $1 \leq \xi \leq N-1$ and

$$q^{(N-2, 2)} = c^{(N-2, 2)} S_{\eta}^{(N-2, 2)}$$

(22)

for the $[N-2, 2]$ sector.

From Eqs. (20) and (21) above, the symmetry coordinates in the $[N]$ and $[N-1, 1]$ sectors are mixed to form a normal coordinate. Thus, depending on the value of the mixing angles, the normal modes, which are the eigenfunctions at first order of the Schrödinger equation will have mixed radial and angular behavior in the $[N]$ and $[N-1, 1]$ sectors. The $[N-2, 2]$ normal modes have entirely angular behavior since there are no $i\eta$ symmetry coordinates in this sector and so no mixing. The value of the mixing angles and thus the extent of radial/angular mixing in a normal coordinate depends, of course, on the Hamiltonian terms at this first perturbation order.

III. MIXING COEFFICIENTS AS A FUNCTION OF $N$ AND $V_0$.

The mixing coefficients that determine the radial/angular mixing in the normal modes for the $[N]$ and $[N-1, 1]$ sectors are defined in Appendix A and have a complicated dependence on $N$ and $V_0$ that originates in the Hamiltonian terms at first order. In particular, these coefficients have some explicit dependence from the symmetry present in the first-order Hamiltonian as well as dependence from the $F$ and $G$ elements for a particular Hamiltonian.

As shown in Appendix A there are three layers of analytic expressions that can bring in $N$ and/or $V_0$ dependence. When these mixing coefficients were plotted for the unitary ($V_0$) Hamiltonian as a function of $N$ in a recent study, the character of the normal modes was found to evolve to a pure radial or pure angular symmetry coordinate for $N \geq 200$ i.e. no mixing for $N >> 1$. (This was also true for very small $N$ e.g. $N \leq 10$ which are not being studied in this work.)

A bit of inspection revealed that this behavior was being dictated to a large extent by the explicit $N$ dependence in the expressions for $\cos \theta^{[\alpha]}_\pm$ and $\sin \theta^{[\alpha]}_\pm$ (Eqs. (A1) - (A3)). These expressions depend on the symmetry of the first-order Hamiltonian, not the specific details of the potential. The position and shape of the crossover is influenced by the other sources of $N$ and $V_0$ dependence that originate in the specific Hamiltonian.

The emergence of pure symmetry coordinates for large $N$ has implications for the stability of collective behavior since the symmetry coordinates are eigenfunctions of an approximate underlying Hamiltonian and thus have some degree of stability unless the system is perturbed. In addition, this means that the frequencies, $\omega^{[N]}_\pm$ and $\omega^{[N-1, 1]}_\pm$, associated with these normal modes should reflect pure radial or pure angular character for large $N$.

In the current study, I now extend this earlier study to regimes other than the unitary regime, investigating whether this emergence of pure symmetry character in the normal modes and their frequencies is unique to the unitary regime or is driven by dynamics common to regimes throughout the transition from BCS to unitarity.

In Figs. (1a)-(1d), I show the behavior of the mixing coefficients as a function of $N$ for a system of identical fermions in the weakly interacting BCS regime, plotting the square of the mixing coefficients, $|\cos \theta_\pm^{[\alpha]}|^2$ and $|\sin \theta_\pm^{[\alpha]}|^2$ which give the probability associated with each symmetry coordinate, $|S_\mu^{[\alpha]}|^2$ or $|S_\mu^{[\alpha]}|$, in the expression for the normal modes, $q_\pm^{[\alpha]}$.

The plots show that the character of the normal modes $q_\pm^{[\alpha]}$ evolves to almost purely radial or purely angular, as $N$ increases with very little mixing of the symmetry coordinates. This happens in this weakly interacting regime at even lower values of $N$ than in the unitary regime, thus confirming that this phenomena is driven by dynamics other than the universal behavior of a system at unitarity. This also validates the decision to designate the $[N]$ and $[N-1, 1]$ sector normal mode frequencies for the typical many-body ensemble sizes studied in the laboratory, as either a radial frequency or an angular frequency instead of having mixed radial/angular character. Inspecting the plots reveals that $\omega_0^-$ and $\omega_1^-$ are angular frequencies and $\omega_0^+$ and $\omega_1^+$ are radial frequencies. (This designation holds over the entire range of interparticle interaction strengths until the systems are approaching the unitary regime where the large value of $V_0$ results in a crossing of the character at which point the labels are switched.)

IV. THE ANALYTIC EXPRESSIONS FOR THE FIVE NORMAL MODE FREQUENCIES

In this section, I analyze the analytic expressions for the frequencies, investigating the differences between the radial, $\omega_0^-$, $\omega_1^-$, and angular, $\omega_0^+$, $\omega_1^+$, frequencies in the $[N]$ and $[N-1, 1]$ sectors, as well as studying the angular frequency $\omega_2$ in the $[N-2, 2]$ sector. The radial frequencies depend strongly on the interparticle interaction potential, $V_0$, while the angular frequencies which are comprised primarily from centrifugal potential terms have a weaker dependence on $V_0$. Thus all the frequencies will respond to tuning the interaction strength in the laboratory using a Feshbach resonance.

Analytic expressions for the $N$-body normal mode frequencies were derived in Ref. [31] using a method outlined in Appendices B and C of that paper which derives analytic formulas for the roots, $\lambda_\mu$, of the FG secular equation. The normal-mode vibrational frequencies, $\omega_\mu^2$, are related to the roots $\lambda_\mu$ of FG by:

$$\lambda_\mu = \omega_\mu^2,$$

(23)

The two frequencies associated with the $\lambda_0$ roots of mul-
function of $N$ for the weakly interacting BCS regime.

\[ (b) \]

\[ \begin{align*}
\cos \theta^2 \sin^2 \theta^2 & = \cos \theta^2 \sin^2 \theta^2 \\
\cos \theta^2 & = \cos \theta^2 \\
\sin \theta^2 & = \sin \theta^2
\end{align*} \]

where:

\[ \eta_i = \frac{1}{2} \left[ a + (N - 1)b + g + 2(N - 2)h + \frac{(N - 2)(N - 3)}{2} \right] \]

\[ \Delta_0 = (a + (N - 1)b) \left[ g + 2(N - 2)h + \frac{(N - 2)(N - 3)}{2} \right] \]

\[ - \frac{N - 1}{2} (2c + (N - 2)d)(2e + (N - 2)f) \].

For the two $N - 1$ multiplicity roots, the frequencies are:

\[ \bar{\omega}_{1 \pm} = \sqrt{\eta_1 \mp \sqrt{\eta_1^2 - \Delta_1}} \]

where $\eta_1$ and $\Delta_1$ are given by:

\[ \eta_1 = \frac{1}{2} \left[ a - b + g + (N - 4)h - (N - 3)\nu \right] \]

\[ \Delta_1 = -(N - 2)(e - d)(e - f) + (a - b) \times [g + (N - 4)h - (N - 3)\nu] \].

The frequency $\bar{\omega}_2$, associated with the root $\lambda_2$ of multiplicity $N(N - 3)/2$ is given by:

\[ \bar{\omega}_2 = \sqrt{g - 2h + \nu} \]

The quantities $a, b, c, d, e, f, g, h$, and $\nu$, are defined in Appendix B in terms of the $F$ and $G$ elements of the first-order Hamiltonian.

**Explicit $\bar{V}_0$ dependence in the analytic expressions for the frequencies.** Analogous to the mixing coefficients, there are three layers of analytic expressions that define the frequencies: the expressions for $\bar{\omega}_{0 \pm}, \bar{\omega}_{1 \pm}$ and $\bar{\omega}_2$ in Eqs. (24)-(28) above, the expressions for $a, b, c, d, e, f, g, h$, and $\nu$ given in Appendix B and the expressions for the $F$ and $G$ elements of Eq. (17) which are also given in Appendix B. The expressions for $\bar{\omega}_{0 \pm}, \bar{\omega}_{1 \pm}$ and $\bar{\omega}_2$ do not contain $\bar{V}_0$ explicitly, however the next layer, which involves the quantities $a, b, c, d, e, f, g, h$, and $\nu$, does have explicit dependence on $\bar{V}_0$, as well as the third layer involving the $F$ and $G$ elements of the first-order Hamiltonian.

**Implicit dependence of the frequencies on $\bar{V}_0$ through the variables $\bar{r}_\infty$, $\gamma_\infty$, $\tanh \Theta_\infty$ and $\sech^2 \Theta_\infty$.** The frequencies have some implicit dependence on $\bar{V}_0$ from the variables $\bar{r}_\infty$ and $\gamma_\infty$ whose values are obtained as roots of transcendental equations that involve $\bar{V}_0$ (see Eqs. (12)-(14)). The values of $\bar{r}_\infty$ and
\( \gamma_\infty \) are also used to determine \( \tanh \Theta_\infty \) and \( \text{sech}^2 \Theta_\infty \). This implicit dependence on the interparticle interaction strength through the solution of transcendental equations complicates understanding the dependence of the frequencies on \( \tilde{V}_0 \) solely by analytic means, however it is possible with a little numerical work to understand how \( \tilde{V}_0 \) is implicitly affecting the frequencies through these variables.

A. The \([N]\) sector frequencies, \( \bar{\omega}_{0\pm} \).

The normal mode frequencies in the \([N]\) sector, \( \bar{\omega}_{0+} \) and \( \bar{\omega}_{0-} \), are associated with the angular center of mass mode and the radial breathing mode respectively. These two frequencies are the largest frequencies of the five normal modes and so do not come into play in providing an excitation gap in ultracold regimes. It is interesting to analyze the relative contributions of the interaction potential, \( \tilde{V}_0 \), versus terms originating in the kinetic energy for these \([N]\) sector frequencies. The breathing frequency is expected to depend strongly on the strength of \( \tilde{V}_0 \) as the particles spread out and then move back in toward the minimum of the effective potential. For the center of mass frequency, the dependence on \( \tilde{V}_0 \) should drop out of the final simplified analytic expression since the center of mass mode is independent of interparticle interactions. In Appendix C I demonstrate how the quadratic formula for \( \bar{\omega}_{0+} \) and \( \bar{\omega}_{0-} \) results in one frequency, \( \bar{\omega}_0 \), the breathing mode, constructed from the terms at first order that involve \( \tilde{V}_0 \) with the centrifugal terms from the kinetic energy cancelling, while the other frequency, \( \bar{\omega}_{0+} \), the center of mass mode, is constructed from centrifugal terms at first order with the terms that involve \( \tilde{V}_0 \) cancelling.

B. The \([N-1]\) sector frequencies \( \bar{\omega}_{1\pm} \).

The normal mode eigenvalues in the \([N-1,1]\) sector, \( \bar{\omega}_{1+} \) and \( \bar{\omega}_{1-} \), are associated with angular and radial particle-hole excitations. These two frequencies are the closest frequencies to the extremely low phonon frequency occupied in ultracold regimes and thus play a role in setting up an excitation gap for these systems. It is again enlightening to analyze the relative contributions of the interaction potential, \( \tilde{V}_0 \), versus the kinetic energy for these \([N-1,1]\) sector frequencies. The radial particle-hole frequency should depend strongly on the strength of \( \tilde{V}_0 \) as a single particle is excited from the ensemble in a radial direction. For the angular particle-hole excitation frequency, I expect to see strong dependence on \( \tilde{V}_0 \) drop out of the final simplified analytic expression and the centrifugal terms in the effective potential contribute. This is demonstrated in Appendix D.

C. The \([N-2,2]\) sector: the \( \bar{\omega}_2 \) frequency.

The normal mode frequency in the \([N-2,2]\) sector, \( \bar{\omega}_2 \), is associated with the phonon compressional mode which has an extremely small frequency and thus is the only normal mode occupied by a gas of fermions at ultracold temperatures. This is an angular mode, so I expect its frequency to be relatively independent of \( \tilde{V}_0 \) and to have a strong dependence on centrifugal terms. This is shown in Appendix E.

V. THE EVOLUTION OF THE EXCITATION GAPS FROM WEAKLY INTERACTING TO UNITARITY

I discuss the emergence, growth and stability of excitation gaps as the frequencies evolve as a function of the strength of interparticle interactions.

A. \( \tilde{V}_0 \) increases for a fixed ensemble size.

I fix the system size, i.e. the value of \( N \), and let the interaction strength \( \tilde{V}_0 \) increase. This analysis is directly relevant to experiments which use a Feshbach resonance to tune the interaction strength for a particular system. The value of \( \tilde{V}_0 \) is changed from essentially zero, i.e. the case of independent, non-interacting particles trapped in a harmonic potential, to the large interactions \( \tilde{V}_0 = 1.0 \) of the unitary regime.

Figs. (1a)-(1c) show this effect for \( N = 10^3, 10^4 \) and \( 10^5 \) particles respectively as \( \tilde{V}_0 \) is tuned from the BCS regime to the unitary regime. The plots show that at the limit of zero interparticle interactions, the frequencies coalesce to the same value of \( 2\bar{\omega}_{ho} \) as expected and observed in the laboratory. (See an expanded view in Figs. (2a)-(2c) of the region near the independent particle limit.) As interactions are slowly turned on, gaps rapidly emerge reaching values that stabilize for the angular frequencies as unitarity is approached.

Unlike the angular frequencies that quickly converge to limiting values, the radial frequencies continue to slowly increase as unitarity is approached, suggesting that higher order terms are needed to converge the radial frequencies. The angular frequencies approach limits that are integer multiples of the trap frequency: twice the trap frequency for the center of mass angular frequency, \( \bar{\omega}_{0+} = 2\bar{\omega}_{ho} \); equal to the trap frequency for the single particle angular excitation, \( \bar{\omega}_{1+} = \bar{\omega}_{ho} \); and orders of magnitude smaller than the trap frequency for the phonon mode, \( \bar{\omega}_2 = O(10^{-2})\bar{\omega}_{ho} \). For \( N \gg 1 \), stable values for the angular frequencies are reached quite quickly as \( \tilde{V}_0 \) increases from the BCS regime. (See Figs. (2a)-(2c).)

As can be seen in all the above figures, the largest gap forms between the extremely low frequency phonon mode which is the only mode occupied by ultracold gases and...
the next lowest frequency which is a particle-hole excitation, i.e. a single particle excitation. Both of these are angular frequencies which reach stable limits, not changing as \( \tilde{V}_0 \) increases. This particular excitation gap is relevant to the emergence and sustainability of superfluidity and is shown in Figs. (3a) and (3b) for system sizes of \( N = 10^3 \) and \( N = 10^4 \) particles. Note this gap emerges at lower \( \tilde{V}_0 \) i.e. for weaker interactions for larger ensemble sizes.

In the expanded view of the region near the independent particle limit (Figs. (2a)-(2c)) when the interactions have just turned on, one notices two phenomena. First the change in the frequencies is quite rapid as soon as the interactions turn on. Note this gap emerges at

**FIG. 1:** Frequencies as a function of the strength of the interparticle interaction, \( \tilde{V}_0 \), from BCS to unitarity in units of the trap frequency. Note the log scale on the x axis.

**FIG. 2:** Expanded view of Figs. (1a)-(1c) near the independent particle region, i.e. in the deep BCS regime, showing the rapid change in the frequencies when interactions turn on. Note the small linear scale on the x axis.

\[
\begin{align*}
\text{Fig. 1: Frequencies as a function of the strength of the interparticle interaction, } \tilde{V}_0, \text{ from BCS to unitarity in units of the trap frequency. Note the log scale on the x axis.}
\end{align*}
\]

\[
\begin{align*}
\text{Fig. 2: Expanded view of Figs. (1a)-(1c) near the independent particle region, i.e. in the deep BCS regime, showing the rapid change in the frequencies when interactions turn on. Note the small linear scale on the x axis.}
\end{align*}
\]
I will take advantage of the analytic forms for both the Appendices F and G respectively. Then in Section VI, I will discuss both these limits in this section, deriving particle limit to stable limits at unitarity as derivation of this effect.)

approaches and Appendix G for details of the analytic consideration of the microscopic dynamics underpinning these two rating the frequencies quickly.

periencing a fixed interaction has a similar effect in sepa-

ber of particles or increasing the number of particles ex-

increase while the interactions increase. Second, the fre-

quences separate more quickly for larger systems, i.e. as very slowly as the interactions increase. Second, the fre-

creases while the angular frequencies evolve to limits of \( \bar{\omega}_{1+} = 2\bar{\omega}_{ho} \), \( \bar{\omega}_{1-} = \bar{\omega}_{ho} \), and \( \bar{\omega}_2 = O(10^{-2})\bar{\omega}_{ho} \). These limits for the angular frequencies are stabilized at lower values of the interaction strength for larger values of \( N \) as previously discussed and as shown in Figs. \( 1a-1c \). In Figs. \( 1a-1b \) the approach to the unitary regime on a linear scale shows the stability of the angular frequencies and the gradual change in the radial frequencies. As will be demonstrated in the next Section, the stable limits for the angular frequencies as \( \bar{V}_0 \) increases signifies the vanishing of the interparticle interactions for these angular motions. A derivation of these limits from the analytic expressions for the angular frequencies is given in Appendix \( G \).

In the following section, the microscopic behavior that underpins this stability is analyzed using the analytic expressions for the angular frequencies and the motions as analyzed in detail in Ref. \( 30 \).

VI. UNDERSTANDING THE MICROSCOPIC DYNAMICS THAT RESULT IN STABLE LIMITS FOR THE ANGULAR FREQUENCIES.

The behavior of the angular frequencies as shown in Figs. \( 1a-1c \) and discussed in the previous sections has revealed three interesting phenomena.

1) First, the angular frequencies are evolving to stable limits independent of interactions as \( \bar{V}_0 \) increases while the radial frequencies continue to slowly change.

2) Second, the limits for the angular frequencies are integer multiples of the trap frequency.

B. Stable limits for the angular frequencies as a function of \( \bar{V}_0 \).

The angular frequencies evolve from the independent particle limit to stable limits at unitarity as \( \bar{V}_0 \) increases. I will discuss both these limits in this section, deriving them from the analytic expressions for the frequencies in Appendices \( F \) and \( G \) respectively. Then in Section VI I will take advantage of the analytic forms for both the normal mode frequencies and the corresponding normal coordinates to understand the microscopic dynamics underpinning the stability of these limits by tracking the evolution of behavior including the normal mode motions of individual particles as \( \bar{V}_0 \) increases.

The independent particle limit: \( \bar{V}_0 = 0 \). Determining the values of the five frequencies is straightforward in the limit of no interactions between the particles. Setting \( \bar{V}_0 \) equal to zero in the transcendental equations for \( \gamma_\infty \) and \( r_\infty \) (See Eqs. \( 12-14 \)) results in values of \( \gamma_\infty = 0 \) and \( r_\infty = 1/\sqrt{2} \). Using these values in the formulas for the frequencies, (See Appendix \( F \)) yields a value of \( 2\bar{\omega}_{ho} \), an integer multiple of the trap frequency, as expected\( 4, 50, 51 \) for each of the five frequencies since the only potential affecting the particles is the harmonic trap. The individual fermions obey Fermi-Dirac statistics, but have no interactions with the other fermions in the trap. Thus all five frequencies coalesce to the same value. This can be clearly seen in Figs. \( 1a-1c \) and in the expanded view in Figs. \( 2a-2c \).

The unitary limit: \( \bar{V}_0 = 1.0 \). As \( \bar{V}_0 \) is turned on and the particles begin to interact, the frequencies spread apart. The radial frequencies, \( \bar{\omega}_0^- \) and \( \bar{\omega}_1^- \) increase while the angular frequencies evolve to limits of \( \bar{\omega}_{0+} = 2\bar{\omega}_{ho} \), \( \bar{\omega}_{1+} = \bar{\omega}_{ho} \), and \( \bar{\omega}_2 = O(10^{-2})\bar{\omega}_{ho} \). These limits for the angular frequencies are stabilized at lower values of the interaction strength for larger values of \( N \) as previously discussed and as shown in Figs. \( 1a-1c \). In Figs. \( 1a-1b \) the approach to the unitary regime on a linear scale shows the stability of the angular frequencies and the gradual change in the radial frequencies. As will be demonstrated in the next Section, the stable limits for the angular frequencies as \( \bar{V}_0 \) increases signifies the vanishing of the interparticle interactions for these angular motions. A derivation of these limits from the analytic expressions for the angular frequencies is given in Appendix \( G \).

FIG. 3: Excitation gap in units of the trap frequency from the lowest two normal mode frequencies as a function of the strength of the interparticle interaction, \( \bar{V}_0 \), from BCS to unitarity. interactions turn on (note the small scale on the x-axis),

quickly approaching values that will stabilize or change very slowly as the interactions increase. Second, the frequencies separate more quickly for larger systems, i.e. as more and more particles are responding to a particular interaction strength.

Thus, increasing the interaction between a fixed number of particles or increasing the number of particles experiencing a fixed interaction has a similar effect in separating the frequencies quickly. (See Sec. \( VII \) for a discussion of the microscopic dynamics underpinning these two approaches and Appendix \( G \) for details of the analytic derivation of this effect.)
3) Third, the gap between the frequencies emerges at weaker interaction strengths for larger values of $N$.

I will now analyze the origins of these three interrelated phenomena for each of the three angular frequencies by first looking at the analytic expressions for the frequencies in these limits; then by tracking the increase in correlation using the variable $\gamma_\infty$; and finally by analyzing the corresponding motion of the associated normal mode using the analysis in Ref. [50]. The motion of the individual particles offers an understanding of the microscopic dynamics responsible for these phenomena, including the emergence, growth and stability of the excitation gaps as $V_0$ increases.

**Tracking the contributions of terms in the Hamiltonian to the analytic expressions for the angular frequencies.** It is enlightening to analyze the evolution of the various terms in the Hamiltonian as they contribute to the value of the frequencies as $V_0$ increases. I will focus on terms in the effective potential which is composed of three terms: the centrifugal term that originates in the kinetic energy, the trap potential, and the interparticle potential; and focus on their effect on the angular frequencies that are relevant to the emergence of superfluid behavior in ultracold regimes. The trap potential affects all the frequencies of course; the two radial frequencies explicitly, and the angular frequencies implicitly through other variables. If there are no interparticle interactions, all the frequencies would, of course, be integer multiples of the trap frequencies.

The analysis in Appendices C, D and E shows that $V_0$ contributes at first order to the radial frequencies (See Eq. (C5) and (D4)), while cancelling out of the expressions for the angular frequencies at first order which are dominated by the centrifugal potential terms. The remaining explicit dependence on $V_0$ (in $F_\gamma$) for the angular frequencies is small, damped by a factor of $1/N$. However, there are implicit dependences on $V_0$ through the variables $r_\infty$, $\gamma_\infty$, tanh $\Theta_\infty$ and sech$^2\Theta_\infty$. Of these four variables, two of them, tanh $\Theta_\infty$ and sech$^2\Theta_\infty$, play a role only in these damped terms.

Of the two remaining variables, $r_\infty$ and $\gamma_\infty$, the most interesting variable to study is $\gamma_\infty$, the angle cosine of each pair of particles in the large dimension maximally symmetric configuration. This variable was identified in early dimensional scaling work to signify the existence of correlation between the particles [34, 52, 53]. The term, correlation energy, has been defined, for example, by comparing the energies obtained in configuration interaction calculations with Hartree Fock/mean-field energies. The correlation energy reflects the change in energy as the particles in the system move in a correlated way, thus minimizing their interactions. These early studies compared dimensionless scaling results to Hartree Fock mean field results and noted that $\gamma_\infty = 0$ in the Hartree Fock approximation which is an independent particle approximation. The non-zero values of $\gamma_\infty$ at zeroth order in the dimensional expansion thus indicated that some correlation effects were being included even at lowest order underpinning the excellent results obtained by this early work at low order [54]

**Tracking the magnitude of $\gamma_\infty$.** Tracking the magnitude of $\gamma_\infty$, which is a negative quantity, as $V_0$ increases and its effect on the different terms in the expression for the frequencies has the potential to reveal insight into how the ensemble is adjusting to the introduction of interactions between the particles, specifically how the particles are rearranging from "independent" motion to collective, coherent motion as correlation sets in. The Pauli principle is, of course, in effect as this transition occurs. Its role is fundamental and will be addressed in depth in a subsequent study. Each expression for the three angular frequencies involves several terms that involve $\gamma_\infty$ from the $F$ and $G$ elements as seen in Eqs. (C6), (D5) and (E2) and Eqs. (G1)-(G3). These terms evolve as correlation increases and the magnitude of $\gamma_\infty$ increases, changing the relative contributions of

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**FIG. 4:** Expanded view of Figs. (1a) and (1b) as unitarity is approached showing the stability of the angular frequencies $\tilde{\omega}_1^+$, $\tilde{\omega}_1^-$, and $\tilde{\omega}_2$, and the gradual change of the radial frequencies $\tilde{\omega}_0^+$ and $\tilde{\omega}_1^-$. The analysis in Appendices C, D and E shows that $V_0$ contributes at first order to the radial frequencies (See Eq. (C5) and (D4)), while cancelling out of the expressions for the angular frequencies at first order which are dominated by the centrifugal potential terms. The remaining explicit dependence on $V_0$ (in $F_\gamma$) for the angular frequencies is small, damped by a factor of $1/N$. However, there are implicit dependences on $V_0$ through the variables $r_\infty$, $\gamma_\infty$, tanh $\Theta_\infty$ and sech$^2\Theta_\infty$. Of these four variables, two of them, tanh $\Theta_\infty$ and sech$^2\Theta_\infty$, play a role only in these damped terms.

Of the two remaining variables, $r_\infty$ and $\gamma_\infty$, the most interesting variable to study is $\gamma_\infty$, the angle cosine of each pair of particles in the large dimension maximally symmetric configuration. This variable was identified in early dimensional scaling work to signify the existence of correlation between the particles [34, 52, 53]. The term, correlation energy, has been defined, for example, by comparing the energies obtained in configuration interaction calculations with Hartree Fock/mean-field energies. The correlation energy reflects the change in energy as the particles in the system move in a correlated way, thus minimizing their interactions. These early studies compared dimensionless scaling results to Hartree Fock mean field results and noted that $\gamma_\infty = 0$ in the Hartree Fock approximation which is an independent particle approximation. The non-zero values of $\gamma_\infty$ at zeroth order in the dimensional expansion thus indicated that some correlation effects were being included even at lowest order underpinning the excellent results obtained by this early work at low order [54].
the kinetic energy, the trap and the interparticle interaction to the frequency. I will analyze the response of each angular frequency to these changes in \( \gamma_\infty \) as correlation increases below.

**Analyzing the microscopic motions of the angular normal modes as unitarity is approached.**

What are the microscopic dynamics that are controlling this evolution to stable large gaps at unitarity? My analysis of the motions of the normal modes in a recent study\(^{30}\) makes it possible to understand the dynamics at a microscopic level as the ensemble rearranges its motion from the independent particle case to correlated, collective behavior. This motion is determined by an intricate balancing of Kronecker delta functions and Heaviside functions that give rise to stable large gaps at unitarity? My analysis of the normal mode motions in Ref. \(^{30}\), Sections 3.2 and 5.1, to be executing identical very small angular normal modes as unitarity is approached. This motion is determined by an intricate balancing of Kronecker delta functions and Heaviside functions that give rise to stable large gaps at unitarity. The center of mass frequency is in an analytic expression that is insensitive to changes in the trap frequency for all values of \( \gamma \). The only non-zero terms are \( G_g \) and \( F_g \) (See Appendix E) which both originate in the kinetic energy, depend implicitly on the trap potential, and involve a dependence on just two particles, \( i \) and \( j \), through the variable, \( \gamma_{ij} \). As \( V_0 \) turns on, \( \gamma_\infty \) takes on a small nonzero value, signifying that weak correlations exist. This nonzero value now means that all of the terms in \( \bar{\omega}_{0+} \) are nonzero and \( G_g \) and \( F_g \) evolve to new values. Specifically \( G_h \), \( F_h \) and \( F_i \) acquire nonzero values and involve the dependence of the kinetic energy \( (G_h) \) and the centrifugal potential \( (F_h) \) and \( F_i \) on \( \gamma_{ij} \) and \( \gamma_{12} \) involving four particles, \( i, j, k, \) and \( l \). These terms become significant in determining the value of the frequency with factors of \( 2(N-2)/N-3) \) for \( F_h \) and \( 2(N-2)(N-3)/2 \) for \( F_i \) as interparticle correlations increase. Since the value of \( \bar{\omega}_{0+} \) remains fixed at 2, the magnitudes of \( G_g \) and \( F_g \) adjust as correlations spread out throughout the ensemble. Thus, the emergence of interparticle interactions starts an intricate readjustment of the ensemble as the particles governed by the evolving Hamiltonian respond to the other \( N-1 \) particles.

**B. The angular particle-hole excitation frequency \( \bar{\omega}_{1+} \)**

Now consider the angular single-particle excitation frequency \( \bar{\omega}_{1+} \) which can also be described as a particle-hole angular excitation.

**Independent of \( V_0 \).** This frequency reaches a constant value equal to the trap frequency as \( V_0 \) increases reflecting the vanishing of interparticle interactions. The analytic formula for \( \bar{\omega}_{1+} \) analyzed in Appendix D reveals the cancellation of the terms involving \( V_0 \) to first order. Thus this frequency is expected to become constant as the interaction changes if higher order terms are small.

**Microscopic dynamics.** The individual particles in the center of mass normal mode can be seen from my analysis of the normal mode motions in Ref. \(^{30}\), Sections 3.2 and 5.1, to be executing identical very small angular motions as a rigid body with negligible change in their radial interparticle distances. Thus, these particles are not affected by the interparticle interaction resulting in a frequency that is an integer multiple (= 2) of the trap frequency since the trap is the only potential affecting the particles.

**Response of \( \bar{\omega}_{0+} \) to changes in \( \gamma_\infty \).** The center of mass frequency remains at a fixed value of \( \bar{\omega}_{0+} = 2\bar{\omega}_{0} \) as the variables in the analytic expression:

\[
\bar{\omega}_{0+} \approx \sqrt{[G_g + 2(N-2)G_h]}
\times \sqrt{[F_g + 2(N-2)F_h + \frac{(N-2)(N-3)}{2}]} F_i
\]

change in response to system parameters. How does this expression remain fixed as its various terms are changing?

Consider the independent particle limit where \( V_0 = 0 \), \( \gamma_\infty = 0 \), \( r_\infty = 1/\sqrt{2} \), and the particles are affected only by the harmonic trap while obeying the Pauli principle. Most of the terms in the expression for \( \bar{\omega}_{0+} \) are zero. The only non-zero terms are \( G_g \) and \( F_g \) (See Appendix E) which both originate in the kinetic energy, depend implicitly on the trap potential, and involve a dependence on just two particles, \( i \) and \( j \), through the variable, \( \gamma_{ij} \). As \( V_0 \) turns on, \( \gamma_\infty \) takes on a small nonzero value, signifying that weak correlations exist. This nonzero value now means that all of the terms in \( \bar{\omega}_{0+} \) are nonzero and \( G_g \) and \( F_g \) evolve to new values. Specifically \( G_h \), \( F_h \) and \( F_i \) acquire nonzero values and involve the dependence of the kinetic energy \( (G_h) \) and the centrifugal potential \( (F_h) \) and \( F_i \) on \( \gamma_{ij} \) and \( \gamma_{12} \) involving four particles, \( i, j, k, \) and \( l \). These terms become significant in determining the value of the frequency with factors of \( 2(N-2)/N-3) \) for \( F_h \) and \( 2(N-2)(N-3)/2 \) for \( F_i \) as interparticle correlations increase. Since the value of \( \bar{\omega}_{0+} \) remains fixed at 2, the magnitudes of \( G_g \) and \( F_g \) adjust as correlations spread out throughout the ensemble. Thus, the emergence of interparticle interactions starts an intricate readjustment of the ensemble as the particles governed by the evolving Hamiltonian respond to the other \( N-1 \) particles.
Independent of \( N \). It also becomes constant as the system size increases since the frequency is insensitive to all interparticle interactions so additional particles have no effect.

**Microscopic dynamics.** This behavior can be understood from a microscopic view of the motions of the particles. In this case, the motion of the corresponding normal mode is made up of one particle creating a “large” angular displacement with the other particles, while the remaining interparticle angles make a small adjustment.

The first group has \( N - 1 \) interparticle angles, while the second group of \((N - 1)(N - 2)/2\) angles quickly becomes the overwhelming majority of the ensemble with displacements that are smaller by a factor of \((N - 2)/2\). (See Sections 3.4 and 5.1 in Ref. [30].) These two opposing and unequal motions invoke some radial interactions from slight changes in the interparticle distances and thus this mode does have some response to the interaction \( \tilde{V}_0 \).

However for values of \( N \) typical of laboratory ensembles \((10^4 - 10^6)\), the percentage of particles moving in lockstep by a smaller and smaller angular amount becomes so dominant that the radial contribution is insignificant.

Thus the harmonic trap is the dominant effect determining this frequency, analogous to the center of mass frequency. This explains the value of the frequency at an integer multiple \((= 1)\) of the trap frequency and its independence from changes in \( \tilde{V}_0 \) and/or \( N \).

**Response of \( \bar{\omega}_{1+} \) to changes in \( \gamma_\infty \).** When \( \tilde{V}_0 = 0 \) and the particles are moving independently, \( \gamma_\infty = 0 \) and \( \tau_\infty = 1/\sqrt{2} \), only \( g_{ij} \) and \( F_g \) in the expression for \( \bar{\omega}_{1+} \) are nonzero with values determined by the harmonic trap yielding \( 2\bar{\omega}_{ho} \) a integer multiple of the trap frequency for \( \bar{\omega}_{1+} \). (See Appendix [E])

\[
\bar{\omega}_{1+} \approx \sqrt{\left[G_g + (N - 4)G_h\right]} \times \sqrt{\left[F_g + (N - 4)F_h - (N - 3)F_i\right]} \to 2\bar{\omega}_{ho} \quad (G_g = 4, F_g = 1, G_h = F_h = F_i = 0)
\]

As interparticle interactions are introduced, \( \gamma_\infty \) is no longer zero so \( G_h, F_h \) and \( F_i \) contribute as interparticle correlations increase. Smaller factors for \( F_h \) of \((N - 4)\) and a negative factor of \(-(N - 3)\) for \( F_i \) result in a smaller value of \( \bar{\omega}_{1+} \) which evolves to \( \bar{\omega}_{ho} \) from its value of \( 2\bar{\omega}_{ho} \) at the independent particle limit. The magnitudes of \( G_g \) and \( F_g \) adjust as the changing value of \( \gamma_\infty \) signifies longer-range correlations bringing in new contributions involving three and four particles.

**C. The phonon frequency \( \bar{\omega}_2 \).**

Finally consider the angular phonon compressional frequency \( \bar{\omega}_2 \). This frequency reaches a constant value that is two or three orders of magnitude smaller than the trap frequency as \( \tilde{V}_0 \) increases.

**Independent of \( \tilde{V}_0 \).** The analytic formula for \( \bar{\omega}_2 \):

\[
\bar{\omega}_2 = \sqrt{\left[G_g - 2G_h\right] \left[F_g - 2F_h + F_i\right]} \]

analyzed in Appendix [E] reveals the insignificance of the terms involving \( \tilde{V}_0 \) to first order. Thus this frequency is expected to become constant as the interaction increases.

**Independent of \( N \).** It also becomes constant as the system size increases since adding particles has no effect on this frequency which is independent of the interparticle interaction to first order.

**Microscopic dynamics.** In this third case, the motion of the corresponding normal mode is made up of three groups of interparticle angles involving particles that move with different angular motions and amounts: a single dominant interparticle angle which has the largest angular displacement; \( 2(N - 2) \) nearest neighbor angles which move with an opposing angular displacement that is smaller than that of the dominant angle by a factor of \((N - 2)\); and a third group which quickly becomes the dominant group of particles involving \((N - 2)(N - 3)/2\) angles which have a displacement that is a factor of \((N - 2)(N - 3)/2 \) smaller than the dominant angle. (See Sections 3.5 and 5.1.3 in Ref. [30].) These groups move in opposing directions with unequal displacements and thus experience some radial interparticle interactions. However as \( N \) increases, the percentage of particles moving in lockstep in the third group by a smaller and smaller angular amount becomes so dominant that the radial contribution due to the movement of the other two groups is negligible. (See Section 5.1.3 in Ref. [30].) This results in the value of the frequency at an integer multiple \((\approx 0)\) of the trap frequency and independence from changes in \( \tilde{V}_0 \) and/or \( N \).

**Response of \( \bar{\omega}_2 \) to changes in \( \gamma_\infty \).** When \( \tilde{V}_0 = 0 \) and the particles are moving independently affected only by the trap potential, \( \gamma_\infty = 0 \) and \( \tau_\infty = 1/\sqrt{2} \). The non-zero terms, \( G_g \) and \( F_g \), depend on \( \gamma_{ij} \). (See Appendix [E]) As \( \tilde{V}_0 \) turns on, \( \gamma_\infty \) is nonzero so all terms in \( \bar{\omega}_2 \) are now nonzero and involve a dependence on \( \gamma_{ij}\gamma_{jk} \) and \( \gamma_{ij}\gamma_{ki} \). Since the value of \( \bar{\omega}_2 \) evolves to very small values from \( 2\bar{\omega}_{ho} \), the magnitudes of \( G_g \) and \( F_g \) must adjust as longer-range correlations throughout the ensemble reflect the realignment of the particles into collective motion.

**D. A discussion of the microscopic dynamics**

The dynamics that drive the angular frequencies to integer multiples of the trap frequency at unitarity are responsible for both the large excitation gap between the lowest two normal modes and the independence of the ensemble from the microscopic interaction details consistent with the expected universal behavior. There are, in fact, two distinct dynamical effects that can produce this behavior. The discussion of the microscopic dynamics in the above paragraphs assumes large values of \( N \) typical
of experiments with ultracold Fermi gases to understand these stable limits as the interparticle interactions vanish. However, increasing $V_0$ can also result in the angular frequencies approaching integer multiples of the trap frequency for fixed values of $N$. These two effects can be seen in the figures in Section V which show $V_0$ increasing for several fixed ensemble sizes. This complimentary behavior as either $N$ or $V_0$ increases was previously noted at the end of Section V A and an analytic derivation of these two effects is given in Appendix G. I discuss both these behaviors in more detail below.

Two distinct microscopic dynamics. As discussed above, when $N$ increases to large values, the percentage of particles that have very small angular movements, i.e. $\gamma_{ij} \ll 1$ in the $[N-1, 1]$ and $[N-2, 2]$ angular modes become the overwhelming majority of particles in the ensemble. (The center of mass mode, of course, has all the particles moving in lockstep with amounts that are small when $N$ is large.) The angular motion i.e. the magnitude of $\gamma_{ij}$, for this majority of particles becomes smaller and smaller as $N$ increases. (See Section 5.1 in Ref. [30].) Since purely angular motions produce no change in the radial distances from the center of the trap, i.e. $\bar{r}_i$ and $\bar{r}_j$ are constant, this motion yields negligible changes in the interparticle distances, $\bar{r}_{ij} = \sqrt{\bar{r}_i^2 + \bar{r}_j^2 - 2\bar{r}_i\bar{r}_j \gamma_{ij}}$ when $\gamma_{ij}$ is tiny.

Now consider letting $V_0$ increase for fixed system size. The correlation between the particles increases as tracked by the parameter $\gamma_{\infty}$. This happens quite rapidly as $V_0$ increases from zero reflecting the rearrangement of the particles into correlated angular motion as collective behavior sets in. This dynamic is relevant to experiments using Feshbach resonances to tune interactions to the large values of the unitary regime. As correlations spread throughout the ensemble into this rigid angular motion, the interparticle interactions become negligible for this fixed value of $N$ and the system is independent of the details of the microscopic interactions.

Microscopic dynamics of unitarity. The lowest two normal modes in the unitary regime have frequencies that set up an excitation gap that is stable and independent of the microscopic details of the interaction. The spectrum of the $[N-1, 1]$ angular mode has evenly spaced levels at every integer multiple of the trap frequency, identical to the spectrum of the non-interacting regime of independent particles. The strong interactions of the unitary regime result in synchronized, correlated behavior that paradoxically have minimal interparticle interactions.

The unitarity limit is defined as having no interaction length scale that is strong enough to sustain the angular normal mode which has all (center of mass mode) or the overwhelming majority (angular particle-hole excitation and phonon modes) of the particles moving as a rigid body with collisionless motion. The gas is expected to show a universal thermodynamic behavior at zero temperature, independent of any microscopic details of the underlying interactions.

VII. SUMMARY AND CONCLUSIONS

In this study, I have looked in detail at the analytic frequencies for $N$ identical particles as a function of the interparticle interaction strength as it is tuned from weakly interacting regimes to the strong interactions of the unitary regime. The frequencies were obtained previously from the normal mode solutions to the SPT first-order equation in inverse dimensionality for a system of confined, interacting, identical particles. These $N$-body normal modes were determined analytically as a function of various system parameters and used to construct wave functions and density profiles for systems of identical bosons [32, 34, 35] and later energies [38] and thermodynamic quantities for fermions [40, 41].

The current investigation is motivated by a recent study of the evolution of the $N$-body analytic normal mode coordinates as $N$ increases from few-body systems that have good molecular equivalents to the expected behavior of many-body ensembles [30]. A specific Hamiltonian, that of the unitary regime, was investigated and two phenomena were noted that could sustain the emergence and stability of superfluid behavior. In this paper, I have extended the study of these two phenomena to a range of interaction strengths from BCS to unitarity.

In particular, I have investigated closely the behavior of the lowest two angular frequencies that are relevant to the emergence of excitation gaps that could support superfluidity as the interparticle strength is increased from BCS to unitarity. I used both the analytic expressions for the frequencies which allow the different contributions from Hamiltonian terms to be assessed and the simple analytic expressions for the normal mode motions [30] to gain insight into the microscopic dynamics underpinning this evolution.

Summary. In summary, my analysis has resulted in a number of observations that may prove useful in understanding the emergence, growth and stability of excitation gaps as well as offering a possible explanation of the microscopic dynamics responsible for universal behavior at unitarity. I list them below:

1) The analytic expressions for the normal mode frequencies produce behavior that supports the emergence of excitation gaps consistent with the known behavior of ultracold Fermi gases in the laboratory tuned using Feshbach resonances from the weakly interacting BCS regime with small gaps to the large gaps of the strongly interacting unitary regime.
2) The normal modes evolve to almost purely radial or purely angular character as $N$ increases, with very little mixing of the symmetry coordinates, over the entire transition from BCS to unitarity. This confirms that the frequencies can be labelled as radial or angular and affects the stability since the symmetry coordinates are analytic solutions of an underlying, approximate Hamiltonian.

3) As $V_0$ increases from zero at the independent particle limit, these first-order analytic frequencies rapidly separate. As unitarity is approached the angular frequencies stabilize while the radial frequencies continue to slowly change. This suggests that higher order terms may be necessary to converge the radial frequencies.

4) The change in the frequencies emerges at weaker interaction strengths as the ensemble grows. Thus, increasing the number of particles experiencing a fixed interaction or increasing the interaction between a fixed number of particles has a similar effect in separating the frequencies quickly.

5) The largest gap forms between the extremely low frequency angular phonon mode, which is the only mode occupied by ultracold gases, and the next lowest frequency which is an angular particle-hole excitation, i.e. a single-particle excitation. This particular excitation gap is relevant to the emergence and sustainability of superfluidity in ultracold systems.

6) The limits for the three angular frequencies are integer multiples of the trap frequency, reflecting the interaction independence of these frequencies.

7) The lowest two normal mode frequencies relevant to ultracold gases provide a spectrum of evenly spaced levels at integer multiples of the trap frequency at unitarity, identical to the spectrum of the non-interacting regime of independent particles. This spectrum is, of course, independent of any microscopic details of the underlying interactions consistent with the dynamics expected for the unitary regime. Thus, the strong interactions of the unitary regime result in strong, long-range correlated behavior that paradoxically has minimal interparticle interactions.

8) Two distinct dynamical effects were found that can drive the angular frequencies to integer multiples of the trap frequency at unitarity. First, when $N$ increases, the angular phonon and single-particle excitation modes that are involved in creating an excitation gap for ultracold particles now have an overwhelming percentage of particles moving with small, purely angular motions that have a negligible response to interparticle interactions. This results in angular frequencies at integer multiples of the trap frequency since the trap is the only potential affecting the particles. Second, as $V_0$ increases for fixed $N$, correlations increase and become long range as tracked by the parameter, $\gamma_\infty$. The motion evolves into rigid angular motion as interparticle interactions vanish yielding frequencies at integer multiples of the trap frequency.

Conclusions. This analysis of the normal mode frequencies yields consistent, physically intuitive behavior that has been observed in the laboratory. The microscopic dynamic underlying this behavior is based on normal mode motions and thus is different than the accepted view that the relevant particles in a superfluid form loosely bound pairs that decrease in size as a Feshbach resonance is tuned to strong interactions.

Normal modes have an infinite spectrum of evenly-spaced excited states. At unitarity, the spectrum involved in an excitation gap for ultracold fermions consists of integer steps of the trap frequency identical to the spectrum of the non-interacting independent particle limit. This behavior supports dynamics at unitarity that is independent of interparticle interactions. Despite having the same spectrum, the dynamics of independent fermions in a trap are quite different from the dynamics of fermions at unitarity whose behavior reflects the strong interactions that have been encapsulated into normal mode motions. A full understanding of how this spectrum affects the dynamics at unitarity requires an understanding of the role of the Pauli principle which is the subject of a future study.

If higher order effects are small, the normal coordinates whose frequencies and mixing coefficients depend on the interparticle interactions are, in fact, beyond-mean-field analytic solutions to a many-body Hamiltonian. The frequencies and the motions of the normal modes evolve in sync with each other, both responding to the same microscopic dynamics. These analytic forms for the frequencies and coordinates allow the details of the terms in the Hamiltonian that are driving the dynamics to be revealed in a particularly transparent way.

Specifically, I looked at the change in the parameter $\gamma_\infty$ whose magnitude increases as $V_0$ increases signalling an increase in the strength and long-range character of correlation as terms involving three and four particles begin to contribute to the values of the frequencies.

The dynamics revealed by this study are based on an exact solution of the the first-order equation of SPT perturbation theory. If higher order terms are significant in a particular regime along this transition, the dynamics could change. In particular, the radial frequencies (which are not involved in providing excitation gaps for ultracold gases) do not show stable limits as unitarity is approached which suggests that higher order terms are needed for these frequencies. First order SPT results have been tested only in the unitary regime, i.e. for strong interactions, yielding ground state energies comparable to benchmark Monte Carlo results and excellent agreement with experiment for thermodynamic quantities.11 The weakly interacting regime has so far been unexplored using this formalism. This approach also does not offer a mechanism for the pairing in real space that occurs beyond unitarity as the ensemble transitions to the BEC regime.

Normal mode functions provide simple, coherent macroscopic wave functions with phase coherence over
the entire system. The dynamics of a normal mode description of the BCS to unitarity transition with its many-body pairing offers an interesting alternative to the models relying on two-body pairing mechanisms to achieve superfluidity. This approach also offers a possible microscopic understanding of the universal behavior at unitarity which could be applicable to other strongly correlated superfluids in diverse systems.

VIII. ACKNOWLEDGMENTS

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Appendix A: The mixing coefficients for the \([N]\) and \([N-1,1]\) sectors.

The mixing coefficients for the \([N]\) sector are:

\[
\begin{align*}
\cos \theta_+^{[N]} &= \frac{\sqrt{2} \sqrt{N-1} (c + (N/2-1)d)}{\sqrt{2} (N-1) (c + (N/2-1)d)^2 + (-a - (N-1)b + \lambda_+^{[N]})^2} \\
\sin \theta_+^{[N]} &= \frac{-a - (N-1)b + \lambda_+^{[N]}}{\sqrt{2} (N-1) (c + (N/2-1)d)^2 + (-a - (N-1)b + \lambda_+^{[N]})^2} \\
\cos \theta_-^{[N]} &= \frac{\sqrt{2} \sqrt{N-1} (c + (N/2-1)d)}{\sqrt{2} (N-1) (c + (N/2-1)d)^2 + (-a - (N-1)b + \lambda_-^{[N]})^2} \\
\sin \theta_-^{[N]} &= \frac{-a - (N-1)b + \lambda_-^{[N]}}{\sqrt{2} (N-1) (c + (N/2-1)d)^2 + (-a - (N-1)b + \lambda_-^{[N]})^2}
\end{align*}
\]

while the coefficients in the \([N-1,1]\) sector are:

\[
\begin{align*}
\cos \theta_+^{[N-1,1]} &= \frac{\sqrt{N-2} (c - d)}{\sqrt{(N-2) (c - d)^2 + (-a + b + \lambda_+^{[N-1,1]})^2}} \\
\sin \theta_+^{[N-1,1]} &= \frac{-a + b + \lambda_+^{[N-1,1]}}{\sqrt{(N-2) (c - d)^2 + (-a + b + \lambda_+^{[N-1,1]})^2}} \\
\cos \theta_-^{[N-1,1]} &= \frac{\sqrt{N-2} (c - d)}{\sqrt{(N-2) (c - d)^2 + (-a + b + \lambda_-^{[N-1,1]})^2}} \\
\sin \theta_-^{[N-1,1]} &= \frac{-a + b + \lambda_-^{[N-1,1]}}{\sqrt{(N-2) (c - d)^2 + (-a + b + \lambda_-^{[N-1,1]})^2}}
\end{align*}
\]

where \(\lambda_+^{[N]}\) and \(\lambda_+^{[N-1,1]}\) are given by Eqs. (23)- (27) in Section IV.

The above equations have some explicit \(N\) dependence (but no \(V_0\) dependence) that is due to the symmetry present in the first-order Hamiltonian. The quantities \(a, b, c, d, e, f, g, h, i\) in the expressions for the mixing coefficients and the eigenvalues, \(\lambda_+^{[N]}\) and \(\lambda_+^{[N-1,1]}\), are defined in Appendix [3] (See also Eq. (42) in Ref [31].) in terms...
of the $F$ and $G$ elements and have explicit $N$ and $\bar{V}_0$ dependence as well as $N$ and $\bar{V}_0$ dependence from the $F$ and $G$ elements from a particular Hamiltonian.

Thus there are three layers of analytic expressions that can bring in $N$ and/or $\bar{V}_0$ dependence; the expressions for mixing coefficients in Eqs. (A1–A3) above, the expressions for $a, b, c, d, e, f, g, h, i, \lambda^{[N]}_a$ and $\lambda^{[N-1,1]}_a$ and the expressions for the $F$ and $G$ elements for a specific Hamiltonian.

**Appendix B: The FG matrix elements.**

The constants used in the expressions for the mixing coefficients and the frequencies are defined below:

$$
\begin{align*}
 a &= G_a F_a \\
b &= G_b F_b \\
c &= G_g F_e + (N - 2) G_h (F_e + F_f) \\
d &= G_g F_f + 2 G_h (F_e + (N - 3) F_f) \\
e &= G_a F_e \\
f &= G_a F_f \\
g &= G_g F_g + 2 (N - 2) G_h F_h \\
h &= G_g F_h + G_h F_g + (N - 2) G_h F_h + (N - 3) G_h F_e \\
i &= G_g F_i + 4 G_h F_h + 2 (N - 4) G_h F_e.
\end{align*}
$$

The non-zero elements of the $G$ matrix are:

$$
\begin{align*}
G_a &= G_{\bar{r}_i \bar{r}_i} = 1 \\
G_g &= G_{\gamma_{ij} \gamma_{ij}} = 2 \frac{1 - \gamma_{\infty}^2}{\bar{r}_{\infty}^2} \\
G_h &= G_{\gamma_{ij} \gamma_{jk}} = \frac{\gamma_{\infty} (1 - \gamma_{\infty})}{\bar{r}_{\infty}^2} \\
&= 2 [1 + (N - 1) \gamma_{\infty}] (1 - \gamma_{\infty}),
\end{align*}
$$

where the matrix elements have been evaluated at the infinite-$D$ symmetric minimum.

Likewise the non-zero $F$ matrix elements are:

$$
\begin{align*}
F_a &= \left( \frac{\partial^2 \bar{V}_{\text{eff}}}{\partial \bar{r}_i^2} \right)_{\infty} \\
&= 1 + \frac{3}{4 \bar{r}_{\infty}^2} \frac{1 + (N - 2) \gamma_{\infty}}{(1 - \gamma_{\infty})(1 + (N - 1) \gamma_{\infty})} \\
&+ \frac{\bar{V}_0 \bar{c}_o}{2} \frac{(N - 1) \text{sech}^2 \Theta_{\infty}}{(1 - \gamma_{\infty}) \tanh \Theta_{\infty} - 1 + \gamma_{\infty}} \\
&= 1 - \gamma_{\infty} \\
&= 2 \bar{r}_{\infty} \frac{1 + (N - 2) \gamma_{\infty}}{2 \bar{r}_{\infty} \sqrt{1 - \gamma_{\infty}}},
\end{align*}
$$

$$
\begin{align*}
F_b &= \left( \frac{\partial^2 \bar{V}_{\text{eff}}}{\partial \bar{r}_i \partial \bar{r}_j} \right)_{\infty} \\
&= \frac{\bar{V}_0 \bar{c}_o}{2} \frac{\text{sech}^2 \Theta_{\infty}}{(1 - \gamma_{\infty}) \tanh \Theta_{\infty} + \frac{1 + \gamma_{\infty}}{2 \bar{r}_{\infty} \sqrt{1 - \gamma_{\infty}}}},
\end{align*}
$$

$$
\begin{align*}
F_c &= \left( \frac{\partial^2 \bar{V}_{\text{eff}}}{\partial \bar{r}_i \partial \gamma_{\bar{r}i}} \right)_{\infty} \\
&= \frac{\gamma_{\infty}^2}{2 \bar{r}_{\infty}^2 (1 - \gamma_{\infty})^2 (1 + (N - 1) \gamma_{\infty})^2} \\
&+ \frac{\bar{V}_0 \bar{c}_o}{2} \frac{\text{sech}^2 \Theta_{\infty}}{(1 - \gamma_{\infty}) \tanh \Theta_{\infty} + \frac{1}{2 \bar{r}_{\infty} \sqrt{1 - \gamma_{\infty}}}},
\end{align*}
$$

$$
\begin{align*}
F_d &= \left( \frac{\partial^2 \bar{V}_{\text{eff}}}{\partial \gamma_{\bar{r}i} \partial \gamma_{\bar{r}j}} \right)_{\infty} \\
&= \frac{1}{2 \bar{r}_{\infty}^2 (1 - \gamma_{\infty})^3 (1 + (N - 1) \gamma_{\infty})^3} \\
&\times \left[ 1 + 3 (N - 2) \gamma_{\infty} + (13 - 11 N + 3 N^2) \gamma_{\infty}^2 \\
&+ (N - 2) (4 - 3 N + 2 N^2) \gamma_{\infty}^3 \right] + \frac{\bar{V}_0 \bar{c}_o}{2} \frac{\text{sech}^2 \Theta_{\infty}}{(1 - \gamma_{\infty}) \tanh \Theta_{\infty} + \frac{1}{2 \bar{r}_{\infty} \sqrt{1 - \gamma_{\infty}}}},
\end{align*}
$$

$$
\begin{align*}
F_e &= \left( \frac{\partial^2 \bar{V}_{\text{eff}}}{\partial \gamma_{\bar{r}i} \partial \gamma_{\bar{r}j}} \right)_{\infty} \\
&= \frac{1}{2 \bar{r}_{\infty}^2 (1 - \gamma_{\infty})^3 (1 + (N - 1) \gamma_{\infty})^3} \\
&\times \left[ (3 + 5 N - 14) \gamma_{\infty} + (11 - 9 N + 2 N^2) \gamma_{\infty}^2 \right],
\end{align*}
$$

$$
\begin{align*}
F_i &= \left( \frac{\partial^2 \bar{V}_{\text{eff}}}{\partial \gamma_{\bar{r}i} \partial \gamma_{\bar{r}kl}} \right)_{\infty} \\
&= \frac{\gamma_{\infty}^2 (2 + (N - 2) \gamma_{\infty})}{2 \bar{r}_{\infty}^2 (1 - \gamma_{\infty})^3 (1 + (N - 1) \gamma_{\infty})^3}.
\end{align*}
$$

Inspection of the formulas for $F$ elements easily re-
veals the explicit dependence of their terms on the confining (trap) potential, \( V_{\text{conf}} \), the centrifugal potential \( \bar{U} \) and/or the interparticle interaction potential, \( \tilde{V}_0 \). All the terms also have some implicit dependence on all the terms in \( V_{\text{eff}} \) through the variables, \( \bar{\ell}_\infty \) and \( \gamma_\infty \). I list the explicit contributions below:

\[
\begin{align*}
F_a &\leftrightarrow V_{\text{conf}}, \bar{U} \text{ and } \tilde{V}_0(\text{strong}) \\
F_b &\leftrightarrow \tilde{V}_0(\text{weak}) \\
F_c &\leftrightarrow \bar{U} \text{ and } \tilde{V}_0(\text{weak}) \\
F_f &\leftrightarrow \bar{U} \\
F_g &\leftrightarrow \bar{U} \text{ and } \tilde{V}_0(\text{weak}) \\
F_h &\leftrightarrow \bar{U} \\
F_i &\leftrightarrow \bar{U}
\end{align*}
\]

Using the fact that the three nonzero \( G \) elements and the centrifugal “potential” terms originate in kinetic energy terms of the Hamiltonian, it is possible to classify the \( a, b, c, d, e, f, g, h, i \) terms. Since they all contain one of the three \( G \) elements, they all have a contribution from the kinetic energy. Below I list the explicit contributions in terms of the potentials.

\[
\begin{align*}
a &= G_aF_a \leftrightarrow V_{\text{conf}}, \bar{U}, \text{ and } \tilde{V}_0(\text{strong}) \\
b &= G_aF_b \leftrightarrow \tilde{V}_0(\text{weak}) \\
c &= G_cF_c + (N - 2)G_h(F_c + F_f) \leftrightarrow \bar{U} \text{ and } \tilde{V}_0(\text{weak}) \\
d &= G_cF_f + 2G_h(F_c + (N - 3)F_f) \leftrightarrow \bar{U} \text{ and } \tilde{V}_0(\text{weak}) \\
e &= G_cF_c \leftrightarrow \bar{U} \text{ and } \tilde{V}_0(\text{weak}) \\
f &= G_cF_f \leftrightarrow \bar{U} \\
g &= G_cF_g + 2(N - 2)G_hF_h \leftrightarrow \bar{U} \text{ and } \tilde{V}_0(\text{weak}) \\
h &= G_cF_h + G_hF_g + (N - 2)G_hF_h + (N - 3)G_hF_i \\
&\leftrightarrow \bar{U} \text{ and } \tilde{V}_0(\text{weak}) \\
i &= G_cF_i + 4G_hF_h + 2(N - 4)G_hF_i \leftrightarrow \bar{U}
\end{align*}
\]

**Appendix C: The analysis of the \([N]\) sector frequencies.**

The frequency, \( \omega_{0\pm} \), associated with the root \( \lambda_{0\pm} \) of multiplicity 1 is given by:

\[
\omega_{0\pm} = \sqrt{\eta_0 \mp \sqrt{\eta_0^2 - \Delta_0}},
\]

**Defining:**

\[
\begin{align*}
\eta_0 &= \frac{1}{2} \left[ a + (N - 1)b + g + 2(N - 2)h + \frac{(N - 2)(N - 3)}{2} \right] \\
\Delta_0 &= (a + (N - 1)b) \left[ g + 2(N - 2)h + \frac{(N - 2)(N - 3)}{2} \right] \\
&\quad - \frac{N - 1}{2}(2c + (N - 2)d)(2e + (N - 2)f).
\end{align*}
\]

(C2)

\[
\begin{align*}
\eta_0^2 - \Delta_0 &= \frac{1}{4} \left[ a + (N - 1)b - (g + 2(N - 2)h) + \frac{(N - 2)(N - 3)}{2} \right] \\
&\quad + \frac{(N - 1)}{2}(2c + (N - 2)d)(2e + (N - 2)f).
\end{align*}
\]

(C3)

Defining:

\[
\begin{align*}
A &= a + (N - 1)b \\
T &= (g + 2(N - 2)h + \frac{(N - 2)(N - 3)}{2}i)
\end{align*}
\]

and substituting the definitions of \( a, b, c, d, e, f, g, h \) and \( i \) in terms of \( F \) and \( G \) elements:

\[
\begin{align*}
A &= G_aF_a + (N - 1)G_aF_b \\
T &= [G_g + 2(N - 2)G_h] \\
&\quad \times \left[ F_g + 2(N - 2)F_h + \frac{(N - 2)(N - 3)}{2}F_i \right] \\
&= T_1T_2 \\
2c + (N - 2)d &= [G_g + 2(N - 2)G_h][2F_e + (N - 2)F_f] \\
&= T_1T_3 \\
2e + (N - 2)f &= [2F_e + (N - 2)F_f] \\
&= T_3
\end{align*}
\]

So:

\[
\begin{align*}
\omega_{0\pm} &= \sqrt{\eta_0 \mp \sqrt{\eta_0^2 - \Delta_0}} \\
&= \sqrt{\frac{1}{2}[A + T] \mp \sqrt{\frac{1}{4}[A - T]^2 + \frac{(N - 1)}{2}T_1T_3}}
\end{align*}
\]

Thus,

\[
\begin{align*}
\omega_{0^-} &= \sqrt{\frac{1}{2}[A + T] + \frac{1}{2}[A - T](1 + x)^{1/2}} \\
\omega_{0^+} &= \sqrt{\frac{1}{2}[A + T] - \frac{1}{2}[A - T](1 + x)^{1/2}}
\end{align*}
\]

(C4)
where

\[ x = \frac{(N-1)T_1T_3^2}{4(A-T)^2} \]

is small. Thus the two frequencies \( \omega_{0\pm} \) in the \([N]\) sector split into a frequency, \( \omega_{0-} \), which has a leading term (under the square root) \( A \) with a strong dependence on the interparticle interaction potential, \( \bar{V}_0 \), and a frequency \( \omega_{0+} \) in which \( A \) cancels out leaving a leading term \( T \) that depends on the centrifugal terms. The powers of \( x \) bring in higher order terms.

\[
\begin{align*}
\omega_{0-} & \approx \sqrt{A} = \sqrt{a + (N-1)b} \\
& = \sqrt{G_aF_a + (N-1)G_aF_b} \\
\omega_{0+} & \approx \sqrt{T} = \sqrt{g + 2(N-2)h + \frac{(N-2)(N-3)}{2} \iota} \\
& = \sqrt{[G_g + 2(N-2)G_h] + \frac{(N-2)(N-3)}{2} F_i} \\
& \times \sqrt{F_g + 2(N-2)F_h + \frac{(N-2)(N-3)}{2} F_i}
\end{align*}
\]

\[ (C5) \]

\[ (C6) \]

**Appendix D: The analysis of the \([N-1,1]\) sector frequencies.**

The frequency, \( \bar{\omega}_{1\pm} \), associated with the root \( \lambda_{1\pm} \) of

\[
\bar{\omega}_{1\pm} = \sqrt{\eta_1 \pm \sqrt{\eta_1^2 - \Delta_1}},
\]

\[ (D1) \]

\[ \eta_1 = \frac{1}{2} \left[ a - b + g + (N-4)h \right] \\
+ (N-3)\iota
\]

\[ (D2) \]

\[ \Delta_1 = (a - b)[g + (N-4)h - (N-3)\iota] \\
- (N-2)(c - d)(e - f), \]

\[ (D3) \]

and substituting the definitions of \( a, b, c, d, e, f, g, h \) and \( \iota \) in terms of \( F \) and \( G \) elements:

\[
\begin{align*}
B &= G_aF_a - G_aF_b \\
R &= [G_g + (N-4)G_h] \times [F_g + (N-4)F_h - (N-3)F_i] \\
&= R_1R_2 \\
c - d &= [G_g + (N-4)G_h] \times [F_c - F_f] \\
&= R_1R_3 \\
e - f &= [F_c - F_f] \\
&= R_3
\end{align*}
\]

So:

\[
\bar{\omega}_{1\pm} = \sqrt{\eta_1 \pm \sqrt{\eta_1^2 - \Delta_1}}
\]

\[
= \sqrt{\frac{1}{2} [B + R] \pm \sqrt{\frac{1}{4} [B - R]^2 + (N-2)R_1R_3}}
\]

Regrouping the expressions for \( \eta_1 \) and \( \eta_1^2 - \Delta_1 \) using these factors yields:

\[
\begin{align*}
\omega_{1-} &= \sqrt{\frac{1}{2} [B + R] + \frac{1}{2} [B - R](1 + x')^{1/2}} \\
\omega_{1+} &= \sqrt{\frac{1}{2} [B + R] - \frac{1}{2} [B - R](1 + x')^{1/2}}
\end{align*}
\]

where

\[ x' = \frac{\frac{1}{2} [B - R]}{\frac{1}{4} (B - R)^2} \]

is small. Similar to the \([N]\) sector, the two frequencies \( \omega_{1\pm} \) in the \([N-1,1]\) sector split into a frequency, \( \omega_{1-} \), which has a leading term \( B \) with a strong dependence on the interparticle interaction potential, \( \bar{V}_0 \), and a frequency \( \omega_{1+} \) in which \( B \) cancels out leaving a leading term \( R \) that depends on the centrifugal terms. The powers of \( x' \) bring in higher order terms.

\[
\begin{align*}
\omega_{1-} & \approx \sqrt{B} = \sqrt{a - b} \\
& = \sqrt{G_aF_a - G_aF_b} \\
\omega_{1+} & \approx \sqrt{R} = \sqrt{g + (N-4)h - (N-3)\iota} \\
& = \sqrt{[G_g + (N-4)G_h] \times [F_g + (N-4)F_h - (N-3)F_i]}
\end{align*}
\]

\[ (D4) \]

\[ (D5) \]

**Appendix E: The analysis of the \([N-2,2]\) sector frequency.**

The frequency, \( \bar{\omega}_2 \), associated with the root \( \lambda_2 \) of multiplicity \( N(N-3)/2 \) is given by:

\[
\bar{\omega}_2 = \sqrt{g - 2h + \iota}.
\]

\[ (E1) \]
Substituting the definitions of \( g, h \) and \( \iota \) in terms of \( F \) and \( G \) elements, the terms in the expression for \( \tilde{\omega}_2 \) can be factored as:

\[
\tilde{\omega}_2 = \sqrt{[G_g - 2G_h] [F_g - 2F_h + F_i]}
\]

(E2)

Note that the only term, \( F_g \), that has explicit dependence on \( \bar{V}_0 \) has only a weak dependence, while \( F_h \) and \( F_i \) contain contributions from the centrifugal potential. As \( \bar{V}_0 \) approaches unitarity, this frequency decreases becoming a tiny fraction of the trap frequency.

**Appendix F: Limits for the frequencies at the independent particle limit.**

The formulas for the five frequencies from Section IV are summarized below:

\[
\begin{align*}
\tilde{\omega}_0 &= \sqrt{\eta_0 + \sqrt{\eta_0^2 - \Delta_0}} \\
\tilde{\omega}_1 &= \sqrt{\eta_1 + \sqrt{\eta_1^2 - \Delta_1}} \\
\tilde{\omega}_2 &= \sqrt{g - 2h + \iota}.
\end{align*}
\]

\[(F1)\]

where the variables \( \eta_0, \Delta_0, \eta_1 \) and \( \Delta_1 \) are defined in Eqs. (C2) and (D2), in terms of the quantities \( a, b, c, d, e, f, g, h, \) and \( \iota \) given in Appendix E in terms of the \( F \) and \( G \) elements of the first-order Hamiltonian.

At the independent particle limit, also referred to as the non-interacting limit, \( \bar{V}_0 = 0, \gamma_\infty = 0 \) and \( \bar{r}_\infty = 1/\sqrt{2} \). Substituting these values into the \( F \) and \( G \) elements readily gives:

\[
\begin{align*}
F_a &= 4 & G_a &= 1 \\
F_b &= 0 & G_b &= 4 \\
F_c &= 0 & G_c &= 0 \\
F_f &= 0 & G_f &= 0 \\
F_g &= 1 \\
F_h &= 0 \\
F_i &= \iota
\end{align*}
\]

Substituting these values into the expressions for \( a, b, c, d, e, f, g, h, \) \( i \) yields:

\[
\begin{align*}
a &= G_a F_a = 4 \\
b &= G_a F_b = 0 \\
c &= G_g F_c + (N - 2) G_h (F_e + F_f) = 0 \\
d &= G_g F_f + 2G_h (F_c + (N - 3) F_f) = 0 \\
e &= G_a F_e = 0 \\
f &= G_a F_f = 0 \\
g &= G_g F_g + 2(N - 2) G_h F_h = 4 \\
h &= G_g F_h + G_h F_g + (N - 2) G_h F_h + (N - 3) G_h F_i = 0 \\
i &= G_g F_i + 4G_h F_h + 2(N - 4) G_h F_i = 0.
\end{align*}
\]

which gives for \( \eta_0, \Delta_0, \eta_1, \Delta_1: \)

\[
\begin{align*}
\eta_0 &= 4 \\
\Delta_0 &= 16 \\
\eta_1 &= 4 \\
\Delta_1 &= 16
\end{align*}
\]

yielding a value of 2 for each frequency (See Eqs. (E1)) in units of the trap frequency as expected.

**Appendix G: Limits for the analytic, angular frequencies for large values of \( \bar{V}_0 \).**

Now consider the strength of the interparticle interaction, \( \bar{V}_0 \), increasing from the weak interactions of the BCS regime to the strong interactions of unitarity. In the expanded view in the region of weak interactions in Figs. (2a)-(2c), one can see these angular frequencies, \( \omega_0, \omega_1, \) and \( \omega_2 \), separate as the interaction gets stronger and in Figs. (2a) and (1b) stabilize at multiples of the trap frequency. What is happening in the analytic expressions as \( \bar{V}_0 \) is increasing for fixed \( N \) to yield these stable limits?

In this Appendix, I will use the analytic expressions for these three frequencies to derive these limits, working with the roots, \( \lambda_+ = \omega_\infty^2 \) in order to avoid the square root signs in the formulas. The three angular roots, \( \lambda_{0+}, \lambda_{1+}, \) and \( \lambda_2 \), are given in terms of the \( F \) and \( G \) elements by:

\[
\begin{align*}
\lambda_{0+} &= \left[ G_g + 2(N - 2) G_h \right] \times \left[ F_g + 2(N - 2) F_h + \frac{(N - 2)(N - 3) F_i}{2} \right] \\
\lambda_{1+} &= \left[ G_g + (N - 4) G_h \right] \times \left[ F_g + (N - 4) F_h - (N - 3) F_i \right] \\
\lambda_2 &= \left[ G_g - 2G_h \right] [F_g - 2F_h + F_i]
\end{align*}
\]

(G1)

(G2)

(G3)

Using the definitions of \( g, h, F, F, \) and \( F \) found in Appendix E and letting \( \bar{V}_0 \) become large, a little numerical work reveals that the dominant terms will involve powers of \( N \gamma_\infty \) which limits to a value of \( 1 \) as \( \bar{V}_0 \rightarrow 1.0 \). Since \( \gamma_\infty \rightarrow O(-1/N) \) for ensemble sizes relevant to experiment, extra factors of \( \gamma_\infty \) in a term will make it drop out. These limits for \( N \gamma_\infty \) and \( \gamma_\infty \) are easily determined numerically and shown in Figs. (4a)-(4b) for two values of \( N \). Similarly, limits for \( \tanh \Theta_\infty \) and \( \operatorname{sech}^2 \Theta_\infty \) as \( \bar{V}_0 \) increases to unitarity can also be obtained numerically.

\[
\begin{align*}
\gamma_\infty &\rightarrow O(-1/N) \\
N \gamma_\infty &\rightarrow -1. \\
\operatorname{sech}^2 \Theta_\infty &\rightarrow 0 \\
\tanh \Theta_\infty &\rightarrow 1,
\end{align*}
\]

as \( \bar{V}_0 \rightarrow 1.0 \).

Note that the powers of \( N \gamma_\infty \) in the analytic expressions for the angular frequencies can begin to dominate the expressions in two ways: 1) As \( \bar{V}_0 \) increases, the magnitude of \( \gamma_\infty \) increases driving \( N \gamma_\infty \) towards its limit of
Using the above limits and the relation $\bar{r}_\infty = \frac{1}{\sqrt{2\sqrt{1+(N-1)\gamma_\infty}}}$, in the definitions of $G_g, G_h, F_g, F_h,$ and $F_i$ in Appendix B and keeping powers of $\gamma_\infty$ that will contribute when factors of $(N-2), (N-4)$ etc. are included from the expressions for the roots, $\lambda_{0^+}, \lambda_{1^+},$ and $\lambda_2$, the limits for these $G$ and $F$ elements for values of $N$ typical of experiments are:

\[
G_a = 1
\]

\[
G_g = \frac{2}{\bar{r}_\infty^2} \left(1 - \gamma_\infty^2\right) \approx \frac{2}{\bar{r}_\infty^2} \gamma_\infty^2
\]

\[
G_h = \frac{\gamma_\infty (1 - \gamma_\infty)}{\bar{r}_\infty^2} \approx \frac{\gamma_\infty}{\bar{r}_\infty^2}
\]  

Looking at the expressions for the three angular roots, $\lambda_{0^+}, \lambda_{1^+},$ and $\lambda_2$, the following combinations are needed: $G_g + 2(N-2)G_h, G_g + (N-4)G_h, G_g - 2G_h, 2(N-2)F_h, (N-2)(N-3)F_i, (N-4)F_h,$ and $(N-3)F_i$. Taking

\[
F_g = \left[\frac{1}{2}\bar{r}_\infty^2 - (N-1)\gamma_\infty^2 \right] \left[\frac{1}{2}\bar{r}_\infty^2 - (N-1)\gamma_\infty^2 \right]
\]

\[
F_h = \left[\frac{1}{2}\bar{r}_\infty^2 - (N-1)\gamma_\infty^2 \right] \left[3 + (5N-14)\gamma_\infty^2 + (11 - 9N + 2N^2)\gamma_\infty^4 \right]
\]

\[
F_i = \left[\frac{1}{2}\bar{r}_\infty^2 - (N-1)\gamma_\infty^2 \right] \left[\frac{6}{2}\gamma_\infty^2 \right]
\]
these limits gives:

\[
G_g + 2(N - 2)G_h = \frac{2 - \gamma_\infty^2}{\bar{r}_\infty^2} + 2(N - 2)\gamma_\infty(1 - \gamma_\infty) \frac{1}{\bar{r}_\infty^2}
\]

\[
= \frac{2}{\bar{r}_\infty^2} (1 - \gamma_\infty)(1 + (N - 1)\gamma_\infty)
\]

\[
= \frac{(1 - \gamma_\infty)}{\bar{r}_\infty^2} \approx \frac{1}{\bar{r}_\infty^2}
\]

\[
G_g + (N - 4)G_h = \frac{2 - \gamma_\infty^2}{\bar{r}_\infty^2} + (N - 4)\gamma_\infty(1 - \gamma_\infty) \frac{1}{\bar{r}_\infty^2}
\]

\[
= \frac{(1 - \gamma_\infty)}{\bar{r}_\infty^2} (2 + (N - 2)\gamma_\infty)
\]

\[
\approx \frac{1}{\bar{r}_\infty^2}
\]

\[
G_g - 2G_h = \frac{2 - \gamma_\infty^2}{\bar{r}_\infty^2} - 2\gamma_\infty(1 - \gamma_\infty) \frac{1}{\bar{r}_\infty^2}
\]

\[
= \frac{2}{\bar{r}_\infty^2} \left[ (1 - \gamma_\infty^2) - \gamma_\infty(1 - \gamma_\infty) \right]
\]

\[
= \frac{2(1 - \gamma_\infty)}{\bar{r}_\infty^2} \approx \frac{2}{\bar{r}_\infty^2}
\]

\[
2(N - 2)F_h \approx 2(N - 2)(-\gamma_\infty\bar{r}_\infty^2) \approx 2\bar{r}_\infty^2
\]

\[
(N - 2)(N - 3)F_i = \frac{N^2 - 5N + 6}{2} F_i \approx 4\bar{r}_\infty^4
\]

\[
(N - 4)F_h = (N - 4)(-\gamma_\infty\bar{r}_\infty^2) = \bar{r}_\infty^2
\]

\[
(N - 3)F_i \approx (N - 3)8\gamma_\infty^2\bar{r}_\infty^4
\]

\[
\approx -8\gamma_\infty^2\bar{r}_\infty^4
\]  

Substituting these results into the expressions for the roots, Eqs. (G7)-(G9) gives:

\[
\lambda_{0+} = \frac{1}{\bar{r}_\infty^4} \times \left[ \frac{1}{2\bar{r}_\infty} + 2\bar{r}_\infty^2 + 4\bar{r}_\infty^4 \right]
\]

\[
\approx 4
\]  

\[
\lambda_{1+} = \frac{1}{\bar{r}_\infty^4} \times \left[ \frac{1}{2\bar{r}_\infty} + \bar{r}_\infty^2 + 8\gamma_\infty\bar{r}_\infty^4 \right]
\]

\[
= \frac{1}{\bar{r}_\infty^4} + 1 + 8\gamma_\infty\bar{r}_\infty^2
\]

\[
\approx 1
\]  

\[
\lambda_2 = \frac{2}{\bar{r}_\infty^2} \times \left[ \frac{1}{2\bar{r}_\infty} + 2\gamma_\infty\bar{r}_\infty^2 + 8\gamma_\infty^2\bar{r}_\infty^4 \right]
\]

\[
= \frac{1}{\bar{r}_\infty^4} + 4\gamma_\infty + 16\gamma_\infty^2\bar{r}_\infty^2 \approx \frac{1}{\bar{r}_\infty^4}
\]

\[
\approx = O(10^{-4}) \text{ for the unitary regime, (G11)}
\]

yielding values for \(\lambda_{0+}, \lambda_{1+}, \text{ and } \lambda_2\) of:

\[
\omega_{0+} = 2
\]

\[
\omega_{1+} = 1
\]

\[
\omega_2 = O(10^{-2})
\]  

in units of the trap frequency, \(\omega_{ho}\). Thus, as expected from physical arguments (see Section XV), these expressions for the angular frequencies limit to multiples of the trap frequency as \(V_0\) and/or \(\gamma\) increase.

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