Towards a standard procedure for the measurement of the multi-photon component in a CW telecom heralded single-photon source

E Rebufello, F Piacentini, M López, R A Kirkwood, I Ruo Berchera, M Gramenegna, G Brida, S Kück, C J Chunnilall, M Genovese and I P Degiovanni

1 Istituto Nazionale di Ricerca Metrologica (INRiM), Strada delle cacce 91, 10135 Torino, Italy
2 Politecnico di Torino, Corso Duca degli Abruzzi 24, 10129 Torino, Italy
3 Physikalisch-Technische Bundesanstalt (PTB), Bundesallee 100, 38116 Braunschweig, Germany
4 National Physical Laboratory (NPL), Hampton Road, Teddington, TW11 0LW, United Kingdom

E-mail: f.piacentini@inrim.it

Received 19 September 2018, revised 20 December 2018
Accepted for publication 28 January 2019
Published 22 February 2019

Abstract

Single-photon sources are set to be a fundamental tool for metrological applications as well as for quantum information related technologies. Because of their upcoming widespread dissemination, the need for their characterization and standardization is becoming of the utmost importance. Here, we illustrate a strategy to provide a quantitative estimate of the multi-photon component of a single-photon source, showing the results achieved in a pilot study for the measurement of the second-order autocorrelation function $g^{(2)}(0)$ of a low-noise CW heralded single photon source prototype (operating at telecom wavelength $\lambda = 1550$ nm) realized in INRiM. The results of this pilot study, involving INRiM, NPL and PTB, will help to build a robust and unambiguous procedure for the characterization of the emission of a single-photon source.

Keywords: single-photon sources, quantum technologies, quantum characterization

(Some figures may appear in colour only in the online journal)
practice, blurs their deterministic behavior with respect to intrinsically probabilistic heralded single-photon sources (HSPSSs) such as those based on four-wave mixing [40, 41] and spontaneous parametric down-conversion (SPDC) [21, 24, 25, 27, 42–44].

Single-photon technologies are nowadays of the utmost relevance in several research branches, and even for upcoming commercial implementations (e.g. secure communication); this makes their characterization a crucial metrological task. The multi-photon component in SPSs emission is one of the key parameters of such devices, and so a standard methodology to properly quantify this parameter would be of major interest for a large scientific community.

In a joint effort to provide a robust procedure suited for this task, three metrological institutes (INRiM, NPL and PTB) have participated in a measurement campaign addressed to the measurement of the multi-photon emission of a particular prototype of SPDC-based continuous wave (CW) HSPS [45, 46], characterized by a very low number of residual non-heralded ‘noise’ photons granting a small multi-photon output. This research aims to prepare a robust procedure for an international metrological comparison on the $g^{(2)}(0)$ measurements of extremely faint light CW sources (i.e. at the single photon level) in the telecom C-band.

An analogous effort has been addressed to establish a proper procedure for the measurement of the $g^{(2)}$ function of a single-emitter pulsed SPS (based on nitrogen vacancies in diamonds emitting in the visible range); the related results can be found in [47].

These joint efforts will be crucial for the development of an international metrological infrastructure for the characterization of both CW and pulsed SPSs, paving the way to the commercial success of the forthcoming quantum photonics related technologies.

In the following, the results achieved for the CW telecom HSPS will be presented.

2. The single-photon source

In our experimental setup (figure 1) a CW laser ($\lambda = 532$ nm) pumps a 10 mm $\times$ 1 mm $\times$ 10 mm periodically-poled lithium niobate (PPLN) crystal, producing non-degenerate SPDC. We choose signal and idler photons with wavelengths of $\lambda_s = 1550$ nm and $\lambda_i = 810$ nm, respectively. The idler photon is sent to an interference filter (IF) with a full width at half maximum (FWHM) of 10 nm, then fiber-coupled and sent to a silicon single-photon avalanche detector (Si-SPAD), heralding the arrival of a 1550 nm signal photon. The signal photon is addressed to a 30 nm FWHM IF and (Si-SPAD), heralding the arrival of a 1550 nm signal photon.

For the $i$th SPAD of each HBT ($i = 1, 2$), we define the total photon detection probability $P_i^{(\text{Ph})}$ as the ratio between the photon count events $N_i^{(\text{Ph})}$ and the trigger events $N_i^{(\text{Trig})}$, i.e.:

$$P_i^{(\text{Ph})} = \frac{N_i^{(\text{Ph})}}{N_i^{(\text{Trig})}}.$$

Analogously, the dark count probability will be $P_i^{(\text{Dark})} = N_i^{(\text{Dark})} / N_i^{(\text{Trig})}$, and the overall detection probability will be $P_i^{(\text{To} \text{g})} = P_i^{(\text{Ph})} + P_i^{(\text{Dark})}$. By carefully tuning the ‘sleep’ time $t_{\text{dead}}$ of the FPGA controlling the OS, in case of gated SPADs or dark-count-free detectors we can always have both detectors ready (i.e. out of their dead time) for each accepted heralding count, thus $N_i^{(\text{Trig})} = N_i^{(\text{Trig})} = N_i^{(\text{Ph})}$. With free-running
detectors, instead, \( N(\text{Trig}) \) is just the number of heralding signals accepted by the FPGA (indicating the number of times the OS has been open), and a dead-time correction must be introduced (unless the dark counts are negligible).

Note that we have made the assumption that all events represented by \( P \)'s are mutually exclusive and independent of each other. This is a reasonable assumption in the situation where these probabilities are \( \ll 1 \), and the HBT SPADs settings are tuned in order to make afterpulses indistinguishable from dark counts. To evaluate the photon count and dark count probabilities, we look at the temporal histograms of the SPADs detection events (as in figure 2).

Actually, one might point out that, even with the three orders of magnitude extinction ratio \( r \) granted by the optical switch, the possibility of unwanted photons leaking through the OS while the HSPS shutter is closed is still present,

---

**Figure 1.** Experimental setup. PPLN: periodically-poled lithium niobate. FC: fiber coupler. SMF: single-mode fiber. \( \lambda/4 \): quarter-wave fiber paddle. \( \lambda/2 \): half-wave fiber paddle. OS: optical shutter. FPGA: field programmable gate array. FBS: fiber beam splitter. HBT: Hanbury-Brown & Twiss interferometer. Coinc: time-tagging coincidence electronics.

**Figure 2.** Histogram of the detections acquired in one run by one of the INRiM HBT gated SPADs, with a 30 ns detection window. The true heralded photons (Her), background photons (Bkg) and dark count (Dark) contributions can be clearly seen. The green shaded area indicates the dark counts forbidding the SPAD to detect a photon from the HSPS.
affecting the dark counts level estimation. Concerning this, we carefully evaluated the leaking photons contribution within the detection window and found it to be way below the statistical uncertainty associated to the dark counts. This means that the photon leakage effect in our HSPS is substantially negligible.

The following analysis allows us to use the time-tagged counts from individual detectors in combination with a time-correlated measurement from both to calculate Glauber’s normalized second-order autocorrelation function:

\[
g^{(2)}(\tau) = \lim_{t \to \infty} \frac{\langle I(t)I(t+\tau) \rangle}{\langle I(t) \rangle \langle I(t+\tau) \rangle},
\]

with \( I(t) \) being the photon emission intensity at time \( t \). For our purpose, the most interesting case appears at zero time difference (\( \tau = 0 \)), since an ideal SPS should present \( g^{(2)}(0) = 0 \), meaning that the source is producing anti-bunched single photons. Hence, the smaller the \( g^{(2)}(0) \), the better the performances in terms of single-photon emission. Actually, what we evaluate for our HSPS is the parameter \( \alpha \) (approximating the second order correlation function \( g^{(2)}(0) \) for \( P_i^{\text{Tot}} \) \( \ll 1 \) [48]:

\[
\alpha = \frac{P_i^{\text{Ph;Ph}}}{P_i^{\text{1}} \cdot P_i^{\text{2}}} \approx g^{(2)}(0),
\]

where \( P_i^{\text{Ph;Ph}} \) is the probability of a coincidence photon count between the two HBT SPADs (dark counts subtracted). Since \( P_i^{\text{Her}} \) is negligible (since we have a probability below \( 10^{-6} \) of getting more than one heralded photon within a 1 ns time interval, which is a generous estimate of the detector jitter time), and \( P_i^{\text{Bkg}} \) is independent from \( P_i^{\text{Dark}} \), one has:

\[
P_i^{\text{Ph;Ph}} = P_i^{\text{Tot;Tot}} - P_i^{\text{2}}P_i^{\text{2}}\left(N_i^{\text{Peak}}\right) + P_i^{\text{1}}P_i^{\text{2}}\left(N_i^{\text{Peak}}\right).
\]

For each HBT SPAD, the quantity \( N_i^{\text{Ph}} \) is extracted as the difference between the counts of the photon peak region (\( N_i^{\text{Peak}} \), integrated on a 10 ns time window) and the counts belonging to a corresponding number of time bins in the dark count region (\( N_i^{\text{Dark}=\text{Peak}} \)), i.e. \( N_i^{\text{Ph}} = N_i^{\text{Peak}} - N_i^{\text{Dark}=\text{Peak}} \). From this, we can compute \( N_i^{\text{Null}} = N_i^{\text{Tot}} - N_i^{\text{Ph}} \), and then we have all we need to evaluate the single-count probabilities in equation (4). The total coincidence probability \( P_i^{\text{Tot;Tot}} = \frac{N_i^{\text{Conc}}}{N_i^{\text{Dark}}} \) is being \( N_i^{\text{Conc}} \), the number of two-photon counts events in the HBT, is independently evaluated by the coincidence electronics associated to it.

A further refinement for the \( \alpha \) parameter evaluation arises when looking at the multichannel picture in figure 2: it is in fact evident that the portion of dark counts on the left of the photon peak plays a different role with respect to the rest of them, since they occur before the arrival of the heralded photons. Indeed, this portion of dark counts forbids the SPAD to detect the heralded photons. This implies that such events (that we will indicate with \( N_i^{\text{Null}} \)) should be removed from the valid trigger events as well as from the SPAD counts, and thus the true valid trigger counts, dark counts and total counts will be, respectively: \( N_i^{\text{Trig}} = N_i^{\text{Trig}} - N_i^{\text{Null}} \), \( N_i^{\text{Dark}} = N_i^{\text{Dark}} - N_i^{\text{Null}} \), \( N_i^{\text{Null}} = N_i^{\text{Null}} \) = \( N_i^{\text{Tot}} - N_i^{\text{Null}} \).

This will affect all the different count probabilities, and the new formula for the \( \alpha \) parameter will be:

\[
\alpha = \frac{P_i^{\text{Trig;Trig}}}{P_i^{\text{1}} \cdot P_i^{\text{2}}}.
\]

By introducing the coefficient \( q_i = N_i^{\text{Trig}} / (N_i^{\text{Trig}} - N_i^{\text{Null}}) \), one can write \( P_i^{\text{Ph}} = q_i P_i^{\text{Trig}} \). Analogously, one can derive for the coincidence count probability the relation \( P_i^{\text{Ph;Ph}} = q_i q_i P_i^{\text{Trig;Trig}} \). This means that for the \( \alpha \) parameter one has:

\[
\alpha = \frac{P_i^{\text{Trig;Trig}}}{P_i^{\text{1}} \cdot P_i^{\text{2}}} = \frac{q_i q_i P_i^{\text{Trig;Trig}}}{p_i^{\text{1}} \cdot p_i^{\text{2}}} = \frac{P_i^{\text{Trig;Trig}}}{P_i^{\text{1}} \cdot P_i^{\text{2}}} \]

that shows how such correction, even though sensible for both single and coincidence count probabilities, does not affect \( \alpha \).

4. The HBT interferometers

INRIM’s HBT comprises a 50:50 telecom FBS and two calibrated Micro Photon Devices single photon counters based on InGaAs/InP SPADs, operating in gated mode with a 30 ns detection window. The two SPADs outputs, together with the FPGA gating signal triggering them, are addressed to a time-tagging coincidence electronics (PicoQuant HydraHarp400) with a 2.5 ps time-bin resolution as well as to a time-amplitude converter (TAC) module, in order to have a separate evaluation of the single counts and coincidence counts of the HBT SPADs. Both SPAD detection windows being 30 ns long, we set a 100 ns coincidence window both in the time-tagging electronics and the TAC, to not miss any coincidence event.

Rearranging equation (3) gives:

\[
\alpha = \frac{P_i^{\text{Tot;Tot}} - P_i^{\text{2}} \cdot P_i^{\text{2}} \left( Q_i^{\text{2}} - Q_i^{\text{1}} \right)}{P_i^{\text{1}} \cdot \left( 1 - Q_i^{\text{1}} \right) \cdot P_i^{\text{2}} \cdot \left( 1 - Q_i^{\text{2}} \right)},
\]

being \( P_i^{\text{1}} = \langle N_i^{\text{Peak}} - N_i^{\text{Dark}=\text{Peak}} \rangle \) and \( Q_i^{\text{1}} = N_i^{\text{Dark}=\text{Peak}} / N_i^{\text{Dark}} \), respectively, the total count probability and the fraction of dark counts registered by the \( i \)th SPAD of the HBT. The \( Q_i^{\text{2}} \) quantities are extracted by the multi-channel analyzer (MCA) histograms of the two detectors forming the HBT, integrating all the counts registered by the time-tagging coincidence electronics within the whole acquisition time. On the other hand, the \( P_i^{\text{1}} \) and \( P_i^{\text{Tot;Tot}} \) probabilities are obtained from the TAC counts, acquired in sets lasting 100 s each (the \( \langle N \rangle \) quantities are the mean values of these repeated acquisitions).

The NPL HBT hosts two different ID Quantique detectors, a gated SPAD (ID210, with a 25 ns detection window) and a low-noise free-running SPAD (ID230). The SPADs outputs, together with the FPGA valid gate output (i.e. the heralding counts validated by the FPGA and used for the HBT SPADs gating), are routed to the HydraHarp400 time-tagger. Since one of the SPADs is free-running, to have a reliable estimate
of the single- and two-photon count events we choose a 200 ns time window in the time-tagging electronics. A first 1000 s acquisition is run feeding the HydraHarp400 with the HBT SPADs output going to the HydraHarp 400 channels 1 and 2, respectively, while the FPGA valid gate output is used as reference external clock. Then, a second 2000 s run is instead performed with the ID210 as the external clock and the ID230 as channel 1 input. Finally a third 1000 s acquisition is collected with the same settings as the first one. This method gives the possibility of making a time-correlated evaluation of the HBT coincidence counts (during the 2000 s acquisition), whilst minimising the (eventual) temporal fluctuations and drifts of the photon source under test. Since one of the HBT detectors is free-running, we need to set a sensible collection time window on the HydraHarp400; to be sure not to loose any significant events, we choose a 200 ns collection time window both for single- and two-photon events acquisition.

Here, the total single-photon count probabilities will be

\[ P_{1}^{(\text{Tot})} = \frac{N_{1}^{(\text{Tot})}}{N_{1}^{(\text{Dark})}}, \]

with \( N_{1}^{(\text{Tot})} \) the number of FPGA valid gates belonging to the single-photon event acquisitions, while the two-photon count probability will be

\[ P_{12}^{(\text{Tot:Tot})} = \frac{N_{12}^{(\text{Coinc})}}{N_{12}^{(\text{Dark})}}, \]

being \( N_{12}^{(\text{Coinc})} \) the FPGA valid gates related to the two-photon events acquisition. The \( Q_{i}^{(\text{Dark})} \) estimation remains instead unchanged.

The PTB HBT hosts two ID Quantique gated SPADs (one ID210 and one ID201, with, respectively, 25 ns and 50 ns detection windows), both triggered by the same FPGA valid gate signal. Since the time-tagging system chosen is a PicoQuant PicoHarp300, hosting only one input channel plus an external clock, the simultaneous evaluation of the single- and two-photon counts is not feasible. For this reason, we use the same method as the second INRIM-NPL measurement session, with a first 1000 s acquisition in which the single

Table 1. Experimental results obtained within the pilot comparison on the parameter \( \alpha \), with a coverage factor \( k = 1 \).

| Session   | \( \alpha \) (INRIM) | \( \alpha \) (NPL) | \( \alpha \) (PTB) |
|-----------|---------------------|-------------------|-----------------|
| INRIM-NPL | 0.013 ± 0.008       | 0.02 ± 0.02       | —               |
| INRIM-PTB | 0.016 ± 0.006       | —                 | 0.04 ± 0.05     |

Table 2. INRIM uncertainty budget related to the INRIM-NPL measurement session. Coverage factor \( k = 1 \). ‘Unc.’: uncertainty. ‘Sens. Coeff.’: sensitivity coefficient. ‘Unc. Contr.’: contribution to the global uncertainty on \( \alpha \).

| Quantity       | Unc.     | Sens. Coeff. | Unc. Contr. |
|----------------|----------|--------------|-------------|
| \( Q_{1}^{(\text{Dark})} \) | 0.05604 | 0.00008      | 1.0461      | 0.00010 |
| \( Q_{2}^{(\text{Dark})} \) | 0.05607 | 0.00008      | 1.0462      | 0.00010 |
| \( N_{1}^{(\text{Trig})} \) | 6.0133 × 10⁶ | 2.6 × 10¹ | 2.242 × 10⁻⁸ | 0.00006 |
| \( N_{1}^{(\text{Tot})} \) | 18.261 | 31           | 7.382 × 10⁻⁶ | 0.0002 |
| \( N_{2}^{(\text{Tot})} \) | 19.396 | 32           | 6.950 × 10⁻⁶ | 0.0002 |
| \( N_{12}^{(\text{Coinc})} \) | 7.1    | 0.4          | 0.01905     | 0.008  |
| \( \alpha \) | 0.013   | 0.008        | —           | —      |

Table 3. INRIM uncertainty budget related to the INRIM-PTB measurement session. Coverage factor \( k = 1 \). ‘Unc.’: uncertainty. ‘Sens. Coeff.’: sensitivity coefficient. ‘Unc. Contr.’: contribution to the global uncertainty on \( \alpha \).

| Quantity       | Unc.     | Sens. Coeff. | Unc. Contr. |
|----------------|----------|--------------|-------------|
| \( Q_{1}^{(\text{Dark})} \) | 0.04525 | 0.00008      | 1.0302      | 0.00008 |
| \( Q_{2}^{(\text{Dark})} \) | 0.04875 | 0.00009      | 1.0340      | 0.00009 |
| \( N_{12}^{(\text{Trig})} \) | 6.1885 × 10⁶ | 2.4 × 10¹ | 1.898 × 10⁻⁸ | 0.00005 |
| \( N_{1}^{(\text{Tot})} \) | 22.490 | 41           | 5.223 × 10⁻⁶ | 0.0002 |
| \( N_{2}^{(\text{Tot})} \) | 23.407 | 43           | 5.018 × 10⁻⁶ | 0.0002 |
| \( N_{12}^{(\text{Coinc})} \) | 9.1    | 0.5          | 0.01294     | 0.006  |
| \( \alpha \) | 0.016   | 0.006        | —           | —      |

Table 4. Uncertainty budget related to the NPL measurement of the INRIM-NPL measurement session. Coverage factor \( k = 1 \). For the evaluation of the global uncertainty on \( \alpha \), we set \( \xi = 2 \). ‘Unc.’: uncertainty. ‘Sens. Coeff.’: sensitivity coefficient. ‘Unc. Contr.’: contribution to the global uncertainty on \( \alpha \).

| Quantity       | Unc.     | Sens. Coeff. | Unc. Contr. |
|----------------|----------|--------------|-------------|
| \( N_{12}^{(\text{Trig})} \) | 1.20348 × 10⁶ | 1.1 × 10¹ | 1.042 × 10⁻⁹ | 0.00002 |
| \( N_{1}^{(\text{Tot})} \) | 304.900 | 60          | 1.726 × 10⁻⁷ | 0.00010 |
| \( N_{2}^{(\text{Tot})} \) | 283.300 | 60          | 1.003 × 10⁻⁷ | 0.00005 |
| \( N_{1}^{(\text{Dark})} \) | 3100   | 60           | 3.256 × 10⁻⁶ | 0.00018 |
| \( N_{2}^{(\text{Dark})} \) | 9600   | 100          | 3.590 × 10⁻⁶ | 0.0004 |
| \( N_{12}^{(\text{Coinc})} \) | 1.20184 × 10⁶ | 1.1 × 10¹ | 5.218 × 10⁻¹⁰ | 0.00001 |
| \( \alpha \) | 0.02    | 0.02         | —           | —      |

5. Results

Table 1 hosts the results obtained by INRIM, PTB and NPL during the different measurement sessions, namely, the concurrent measurements of the \( g^{(2)}(0) \) of the HSPS carried out firstly by INRIM and NPL and then by INRIM and PTB.

The uncertainty budgets related to the INRIM measurements are listed in tables 2 and 3, showing the single uncertainties, sensitivity coefficients and contributions to the total uncertainty on \( \alpha \) for each measurement session.
The uncertainty budget of the INRiM-NPL and INRiM-PTB measurement sessions are instead given in tables 4 and 5, respectively. Since here we do not have repeated measurements providing information on the photon counts fluctuations (in our regime, dark counts obey a Poisson distribution), we give an upper bound for them considering a superpoissonian behavior by putting $u(N) = \xi \sqrt{N}$, with $\xi > 1$.

6. Conclusions

We have illustrated a strategy to properly evaluate the multi-photon component of a CW light source, applied to a low-noise prototype of a fiber HSPS @1550 nm particularly suited for metrological and quantum-communication-related purposes and adaptable to a large variety of detectors and other devices. The results of the whole measurement campaign, carried out with different measurement setups and data collection methodologies, are all in agreement within the experimental uncertainties, even with coverage factor $k = 1$.

The proposed strategy may pave the way to a standardization of the characterisation of single-photon sources, a task of the utmost relevance for present and future metrology for quantum technologies.

Acknowledgments

This work has been supported by EMPIR-14IND05 ‘MIQC2’, EMPIR-17FUN06 ‘SIQUST’ and EMPIR-17FUN01 ‘BECOME’ (the EMPIR initiative is co-funded by the EU H2020 and the EMPIR Participating States).

Appendix. Uncertainties evaluation

A.1. INRiM uncertainty budgets

For what concerns the uncertainty budget related to INRiM measurements, we can identify three different contributions:

- $u(P_i^\text{(Tot)}(\alpha))$: uncertainty derived from the double ($P_i^\text{(Tot)}(\alpha)$) and single ($P_i^\text{(Tot)}$) photon count probabilities, given by the TAC counts;
\[ u(\alpha) = \sqrt{\frac{\partial^2 N_i}{\partial N_i^2}(\alpha) + \frac{\partial^2 N_{i+1}}{\partial N_i^2}(\alpha) + \frac{\partial^2 N_p}{\partial N_i^2}(\alpha)}. \] (A.5)

A.2. NPL and PTB uncertainty budgets

Concerning NPL and PTB measurements, for evaluating the uncertainty \( u(\alpha) \) we use:

\[ u(\alpha) = \sqrt{\sum_{i=0}^{N} \left( \frac{\partial \alpha}{\partial N_i} \right)^2 u^2(N_i)}, \] (A.6)

with \( N_0 = N_{i0}^{\text{Trig}} \), \( N_i = N_i^{\text{Tot}} \) and \( N_{2+i} = N_i^{\text{Dark}} (i = 1, 2) \), \( N_4 = N_C^{\text{Trig}} \) and, finally, \( N_0 = N^{\text{Conc}} \).

Since here we do not have repeated measurements, we provide a sensible upper bound to the measurement uncertainties by statistical considerations on the physics of the source and measurement device. The dark counts of the HBT SPADs are known to follow a Poisson distribution. The statistics of a multi-mode SPDC, like the one exploited in our HSPS, is almost indistinguishable from a Poissonian distribution. For these reasons, we can assume the fluctuations of the \( N_i \) quantities in (A.6) to be Poissonian as well, and choose as related uncertainties \( u(N_i) = \xi \sqrt{N_i} \), being \( \xi \) a coefficient to be tuned taking into account the (eventual) superpoissonian behavior of the system. In our case, having observed some unexpected fluctuations in the CW pump power, to be conservative we choose \( \xi = 2 \).

**ORCID iDs**

E Rebufello https://orcid.org/0000-0002-8374-1976
F Piaccintini https://orcid.org/0000-0002-8098-5692
I Ruo Berchera https://orcid.org/0000-0003-3131-8866
M Gramené https://orcid.org/0000-0002-5725-0444
G Brida https://orcid.org/0000-0003-2935-1544
M Genovese https://orcid.org/0000-0001-9186-8849
I P Degiovanni https://orcid.org/0000-0003-0332-3115

**References**

[1] Zwinkels J C, Ikoven E, Fox N P, Ulm G and Rastello M L 2010 Metrologia 47 R15
[2] Brida G, Genovese M and Gramené M 2006 Laser Phys. Lett. 3 115–23
[3] Polyakov S V and Migdall A L 2009 J. Mod. Opt. 56 1045–52
[4] Avella A et al 2011 Opt. Express 19 23249
[5] Schmunk W, Rodenberger M, Peters S, Hofer H and Kueck S 2011 J. Mod. Opt. 58 1252–9
[6] Brask J B, Chaves R and Brunner N 2013 Phys. Rev. A 88 012111
[7] Brida G, Degiovanni I P, Genovese M, Schettini V, Polyakov S and Migdall A 2008 Opt. Express 16 11750–8
[8] Brida G, Degiovanni I P, Genovese M, Piaccintini F, Schettini V, Gisin N, Polyakov S V and Migdall A 2009 Phys. Rev. A 79 0444102
[9] Genovese M 2005 Phys. Rep. 413 319–96
[10] Brida G, Degiovanni I P, Genovese M, Migdall A, Piaccintini F, Polyakov S V and Traina P 2013 J. Phys. Soc. Japan 82 034004
[11] Piaccinti F et al 2016 Phys. Rev. Lett. 116 180401
[12] Piaccinti F et al 2016 Phys. Rev. Lett. 117 170402
[13] Avella A, Piaccinti F, Borsarelli M, Barbieri M, Gramené M, Lussana R, Villa F, Tosi A, Degiovanni I P and Migdall A 2017 Phys. Rev. A 96 052123
[14] Piaccinti F et al 2017 Nat. Phys. 13 1191–4
[15] Piaccinti F, Avella A, Gramené M, Lussana R, Villa F, Tosi A, Brida G, Degiovanni I P and Genovese M 2018 Sci. Rep. 8 6959
[16] Thew R and Gisin N 2007 Nat. Photon. 1 165–71
[17] O’Brien J L, Furusawa A and Vickovic J 2009 Nat. Photon. 3 687–95
[18] Lasota M, Demkowicz-Dobrzański R and Banaszek K 2013 Int. J. Quantum Inf. 11 1350034
[19] Goldschmidt E A, Piaccinti F, Ruo Berchera I, Polyakov S V, Peters S, Kück S, Brida G, Degiovanni I P, Migdall A and Genovese M 2013 Phys. Rev. A 88 013822
[20] Scheel S 2009 J. Mod. Opt. 56 141–60
[21] Eisaman M D, Fan J, Migdall A and Polyakov S V 2011 Rev. Sci. Instrum. 82 071101
[22] Pomarico E, Sanguinetti B, Guerreiro T, Thew R and Zbinden H 2012 Opt. Express 20 23846
[23] De Martini F, Di Giuseppe G and Marrocco M 1996 Phys. Rev. Lett. 76 900–3
[24] Ramelow S et al 2013 Opt. Express 21 6707–17
[25] Krapick S, Herrmann H, Quiring V, Brecht B, Suche H and Silberhorn C 2013 New J. Phys. 15 033010
[26] Mazzarella L, Ticozzi F, Sergienko A V, Vallone G and Villioresi P 2013 Phys. Rev. A 88 023848
[27] Förtsch M, Fürst J, Wittmann C, Strekalov D, Aiello A, Chekhova M V, Silberhorn C, Leuchs G and Marquardt C 2013 Nat. Commun. 4 1818
[28] Kurtsiefer C, Mayer S, Zarda P and Weinfurter H 2000 Phys. Rev. Lett. 85 290
[29] Aharonovich I, Englund D and Toth M 2016 Nat. Photon. 10 631
[30] Gatto Monticone D et al 2014 New J. Phys. 16 053005
[31] Forneris J et al 2015 Sci. Rep. 5 15901
[32] Shields A J 2007 Nat. Photon. 1 215–23
[33] Arita M, Le Roux F, Holmes M J, Kako S and Arakawa Y 2017 Nano Lett. 17 2902–7
[34] Senellart P, Solomon G and White A 2017 Nat. Nanotechnol. 12 10261039
[35] Farrara P, Heinze G, Albrecht B, Ho M, Chavez M, Teo C, Sangoudar N and de Riedmatten H 2016 Nat. Commun. 7 13556
[36] Dietrich F and Walther H 1987 Phys. Rev. Lett. 58 203–6
[37] Higginbottom D B, Sodicka L, Arnedda G, Lachman L, Filip R, Henmrich M and Blatt R 2016 New J. Phys. 18 093038
[38] Louvis B and Moerner W E 2000 Nature 407 491
[39] Chu X-L, Goetzinger S and Sandeghvar V 2017 Nat. Photon. 11 5862
[40] Goldschmidt E A, Eisaman M D, Fan J, Polyakov S V and Migdall A 2008 Phys. Rev. A 78 013844
[41] Spring J B et al 2013 Opt. Express 21 13522–32
[42] Fasel S, Alibart O, Tanzilli S, Baldi P, Beveratos A, Gisin N and Zbinden H 2004 New J. Phys. 6 163
[43] Zhong T, Hu X, Wong F N C, Berggren K K, Roberts T D and Battle P 2010 Opt. Lett. 35 1392
[44] Montaut N, Sansoni L, Meyer-Scott E, Ricken R, Quiring V, Herrmann H and Silberhorn C 2017 Phys. Rev. Appl. 8 024021
[45] Brida G, Degiovanni I P, Genovese M, Migdall A, Piaccinti F, Polyakov S V and Ruo Berchera I 2011 Opt. Express 19 1484–92
[46] Brida G et al 2012 Appl. Phys. Lett. 101 221112
[47] Moreva E et al 2019 Metrologia 56 015016
[48] Grangier P, Roger G and Aspect A 1986 Europhys. Lett. 11 173–9