Quark spin polarization and spontaneous magnetization in high density quark matter

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By using the Nambu-Jona-Lasinio model with a tensor-type four-point interaction between quarks, it is shown that there exists a possibility of a spin polarized phase in quark matter at finite temperature and density. When there exists the spin polarization, the spontaneous magnetization may occur if the effect of the anomalous magnetic moment of quark is taken into account. An implication to the compact star objects with strong magnetic field is discussed when the spin polarization occurs.

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1. Introduction

In recent years, one of the interests for physics governed by the quantum chromodynamics (QCD) may be to clarify the phase structure with respect to temperature, quark and/or isospin chemical potential, external magnetic field and so on [1]. In the region with low temperature and high baryon density, various phases are expected in quark matter, such as inhomogeneous chiral condensed phase [2], quarkyonic phase [3], color superconducting phase [4] and so forth. Recently, it has been shown that there is a possibility of the existence of spin polarized phase by using the Nambu-Jona-Lasinio (NJL) model [5], which is regarded as an effective model of QCD [6, 7, 8], with a tensor-type four-point interaction between quarks at zero temperature and finite baryon density [9, 10, 11]. Also, under the external magnetic field, a possible QCD ground state has been investigated at zero and finite temperature in the case of one flavor [12].

The investigation of a possible phase structure under the magnetic field is one of interesting subjects about QCD phase structure [13]. The existence of a strong magnetic field is found in certain kinds of compact stars such as neutron star and magnetar [14]. Further, in the ultrarelativistic heavy-ion collision experiments, it is indicated that a strong magnetic field may be created in the early stage of nucleus-nucleus collisions [15].

In this paper, we show that the spin polarized phase may be exist even at finite temperature in the region of high baryon density in quark matter due to the tensor-type four-point interaction between quarks [16]. Further, it is shown that a spontaneous magnetization appears by introducing an external magnetic field, only if spin polarization occurs and the effect of an anomalous magnetic moment of quark is taken into account [17].

2. Spin polarized phase originated from the tensor-type four-point interaction

Let us start from the following two-flavor NJL model Lagrangian density with a tensor-type four-point interaction between quarks:

\[
\mathcal{L} = \bar{\psi} i\gamma^\rho \partial_\rho \psi + G_S \left\{ (\bar{\psi} \psi)^2 + (\bar{\psi} i\gamma^5 \vec{\tau} \psi)^2 \right\} - \frac{G_T}{4} \left\{ (\bar{\psi} \gamma^\rho \gamma^\sigma \vec{\tau} \psi) \cdot (\bar{\psi} \gamma_\rho \gamma_\sigma \vec{\tau} \psi) + (\bar{\psi} i\gamma^5 \gamma^\rho \gamma^\sigma \psi)(\bar{\psi} i\gamma^5 \gamma_\rho \gamma_\sigma \psi) \right\},
\]

(2.1)

where \(\psi\) represents the quark field with the two-flavor. Here, the first line represents the original NJL model Lagrangian density and the second line represents the tensor-type interaction introduced in this model. Under a mean field approximation, the above Lagrangian density is recast into

\[
\mathcal{L}_{\text{MFA}} = \bar{\psi} (i\gamma^\rho \partial_\rho - M) \psi - F (\bar{\psi} \Sigma_3 \psi) - \frac{M^2}{4G_S} - \frac{F^2}{2G_T}.
\]

(2.2)

Here, we define the dynamical quark mass \(M\) with the chiral condensate \(\langle \bar{\psi} \psi \rangle\) and the spin polarized condensate \(F\), respectively, as

\[
M = -2G_S \langle \bar{\psi} \psi \rangle, \quad F = -G_T \langle \bar{\psi} \Sigma_3 \psi \rangle, \quad \Sigma_3 = -i\gamma^5 \gamma^2 = \begin{pmatrix} \sigma_3 & 0 \\ 0 & \sigma_3 \end{pmatrix},
\]

(2.3)
where $\sigma_3$ is the third component of the Pauli spin matrix. We can easily derive the Hamiltonian density under the mean field approximation and by diagonalizing the Hamiltonian density, we can also derive the energy eigenvalues or single-particle energies of quark easily as

$$E_p^{(\eta)} = \sqrt{p_3^2 + \left( \sqrt{p_1^2 + p_2^2 + M^2 + \eta F} \right)^2}, \quad (\eta = \pm 1) \quad (2.4)$$

where $p = (p_1, p_2, p_3)$ is momentum.

The thermodynamic potential $\Omega$ at finite temperature $T$ and quark chemical potential $\mu$ can be expressed as

$$\Omega = \mathcal{H}_{\text{MFA}} - \mu (N_p - N_{\bar{p}}) + \mathcal{H}_{\text{vacuum}} - TS, \quad (2.5)$$

where

$$\mathcal{H}_{\text{MFA}} = \sum_{p, \eta, \tau, \alpha} E_p^{(\eta)} \left( n_p^{(\eta)} + \bar{n}_p^{(\eta)} \right) + \frac{M^2}{4G_s} + \frac{F^2}{2G_T},$$

$$\mathcal{H}_{\text{vacuum}} = - \sum_{p, \eta, \tau, \alpha} E_p^{(\eta)}, \quad N_p = \sum_{p, \eta, \tau, \alpha} n_p^{(\eta)}, \quad N_{\bar{p}} = \sum_{p, \eta, \tau, \alpha} \bar{n}_p^{(\eta)},$$

$$n_p^{(\eta)} = \frac{1}{1 + \exp \left( \left( E_p^{(\eta)} - \mu \right) / T \right)}, \quad \bar{n}_p^{(\eta)} = \frac{1}{1 + \exp \left( \left( E_p^{(\eta)} + \mu \right) / T \right)},$$

$$S = - \sum_{p, \eta, \tau, \alpha} \left[ n_p^{(\eta)} \ln n_p^{(\eta)} + (1 - n_p^{(\eta)}) \ln(1 - n_p^{(\eta)}) + \bar{n}_p^{(\eta)} \ln \bar{n}_p^{(\eta)} + (1 - \bar{n}_p^{(\eta)}) \ln(1 - \bar{n}_p^{(\eta)}) \right] \quad (2.6)$$

Here, $\tau$ and $\alpha$ are indices for isospin and color of quark, respectively. Then, the thermodynamic potential can be reexpressed as

$$\Omega = - \sum_{p, \eta, \tau, \alpha} \left[ E_p^{(\eta)} + T \ln \left( 1 + \exp \left( - \frac{E_p^{(\eta)} - \mu}{T} \right) \right) + T \ln \left( 1 + \exp \left( - \frac{E_p^{(\eta)} + \mu}{T} \right) \right) \right]$$

$$+ \frac{M^2}{4G_s} + \frac{F^2}{2G_T}. \quad (2.7)$$

When the sum with respect to three-momentum $p$ is replaced to the integration, we have to introduce the three-momentum cutoff $\Lambda$. Also, the isospin and color degrees of freedom give factor 2 and 3, respectively. As a result, the sum is replaced as following way:

$$\sum_{p, \tau, \alpha} \rightarrow \frac{3}{\pi^2} \int_0^\Lambda dp_T \int_0 \sqrt{\Lambda^2 - p_T^2} \left. dp_3 \right.,$$

where $p_T = \sqrt{p_1^2 + p_2^2}$.

| $\Lambda/\text{GeV}$ | $G_s/\text{GeV}^{-2}$ | $G_T/\text{GeV}^{-2}$ |
|----------------------|----------------------|----------------------|
| 0.631                | 5.5                  | 11.0                 |

**Table 1**: Parameters used here.
Figure 1: The contour map of the thermodynamic potential is depicted. The horizontal and vertical axes represent the dynamical quark mass $M$ and the spin polarization $F$, respectively, with various quark chemical potential $\mu$ and temperature $T$. The darker color represents the lower value of the thermodynamic potential.

We use the parameters in Table 1. The coupling constant $G_S$ and the three-momentum cutoff $\Lambda$ are determined by the quark condensate and the pion decay constant in vacuum. However, the coupling strength of the tensor-type interaction cannot be determined by the physical quantity in vacuum. In this paper, we regard $G_T$ as a free parameter. It is shown in Figure 1 that chiral symmetry is broken for small chemical potential and low temperature, namely the minimum point of the thermodynamic potential has $M \neq 0$ and $F = 0$. However, if the chemical potential or temperature becomes high, chiral symmetry is restored, namely, $M = 0$ and $F = 0$. In the large chemical potential and low temperature region, the spin polarized condensate $F$ appears without $M$. However, for higher temperature, it disappears. In Figure 2, the phase diagram in this model is presented on the $T-\mu$ plane, where the horizontal and vertical axes represent the quark chemical potential $\mu$ and temperature $T$, respectively. It is indicated that, for the boundary of the chiral condensed and the normal phases, there is a critical endpoint for the phase transition near $\mu = 0.31$ GeV and $T = 0.046$ GeV. On the other hand, the phase transition from the normal quark phase to
Figure 2: The phase diagram is shown. The horizontal and vertical axes represent quark chemical potential and temperature, respectively.

the spin polarized phase is always of second order.

3. Spontaneous magnetization

In this section, a possibility of spontaneous magnetization of quark matter is investigated by using the NJL model with the tensor-type four-point interaction between quarks. For this purpose, from beginning, the effect of the anomalous magnetic moment of quark, $\mu_A$, is introduced at the level of the mean field approximation. Here, we use a fact that the anomalous magnetic moment is expressed as

$$\mu_A = \left( \begin{array}{cc} \mu_u & 0 \\ 0 & \mu_d \end{array} \right), \quad \mu_u = 1.85 \mu_N, \quad \mu_d = -0.97 \mu_N, \quad \mu_N = \frac{e}{2m_p} = 3.15 \times 10^{-17} \text{ GeV/T}. \quad (3.1)$$

We introduce the effects of the anomalous magnetic moment in the Lagrangian density within the mean field approximation. As a result, we obtain the following Lagrangian density:

$$\mathcal{L}_A = \mathcal{L}_{\text{MFA}} - \frac{i}{2} \bar{\psi} \gamma^\rho F_\rho \sigma \psi,$$  \quad (3.2)

where $F_\rho (1) = \partial_\rho A_\sigma - \partial_\sigma A_\rho$ and $F_{12} \equiv -B_z = -B$. Here, $\mathcal{L}_{\text{MFA}}$ is nothing but Eq.(2.2) with $M = 0$ because we consider the large chemical potential region. We only take $\rho$ and $\sigma$ as $\rho = 1$, $\sigma = 2$ and $\rho = 2$, $\sigma = 1$ because the magnetic field has only $z$-component. Then, in the mean field approximation, Eq.(3.2) is recast into

$$\mathcal{L}_A = i \bar{\psi} \gamma^\rho D_\rho \psi - \bar{\psi} (F_3 \Sigma_3 + \mu_A B_1) \Sigma_3 \psi - \frac{F^2}{2G_T} = i \bar{\psi} \gamma^\rho D_\rho \psi - \bar{\psi} \tilde{F} \Sigma_3 \psi - \frac{F^2}{2G_T}. \quad (3.3)$$

Here, since $F_3 \Sigma_3 = F \mathbf{1}$ where $\mathbf{1}$ is the $2 \times 2$ identity matrix for the isospin space, which are denoted in Eq.(2.3), then we introduced the flavor-dependent variables $\tilde{F}_f$ as

$$\tilde{F}_f = F + \mu_f B, \quad \text{namely} \quad \tilde{F}_u = F + \mu_u B, \quad \tilde{F}_d = F + \mu_d B. \quad (3.4)$$
Thus, we can derive the thermodynamic potential in the same way developed in the section 2, except for the existence of the external magnetic field $B$. The energy eigenvalues are obtained as

$$E_{p, \nu, \eta}^f = \sqrt{p_r^2 + \left( \tilde{F}_f + \eta \sqrt{2|Q_f|B} \right)^2}, \quad \left\{ \begin{array}{ll} v = 0, 1, 2, \cdots \text{ for } \eta = 1, \\ v = 1, 2, \cdots \text{ for } \eta = -1, \end{array} \right. \quad (3.5)$$

with $Q_u = 2e/3$ and $Q_d = -e/3$. Further, because the Landau quantization is carried out with respect to $p_r$, the integration with respect to $p_r$ is replaced to the sum with respect to integers $v$.

As a result, the thermodynamic potential $\Phi$ at zero temperature with $M = 0$ can be expressed as

$$\Phi = 3 \int_{-p_F}^{p_F} \frac{dp_r}{2\pi} \sum_{f = u,d, \eta = \pm} \frac{|Q_f|B}{2\pi} \sum_{\nu = \nu_{\min}}^{\nu_{\max}} \left[ E_{p, \nu, \eta}^f - \mu \right] + \frac{F^2}{2G_T} + (\text{vacuum term}), \quad (3.6)$$

where $p_F$, $\nu_{\min}$ and $\nu_{\max}$ are determined by the condition $E_{p, \nu, \eta}^f \leq \mu$. Keeping in mind that we take $B \to 0$ finally in order to derive the spontaneous magnetization through the thermodynamic relation, the sum with respect to the Landau level $\nu$ can be replaced into the integration with respect to the continuum variable $a \nu$ with a small quantity $a$. A detailed derivation is found in Ref.[17].

After the calculation, the results are summarized as follows:

$$\Phi = \Phi_\geq = -\frac{1}{2} \sum_{f = u,d} \left[ \tilde{F}_f \mu^3 / 2\pi + \frac{F^2}{2G} + O(B^2) \right], \quad \text{for } F > \mu,$$

$$\Phi = \Phi_\leq = \frac{1}{2} \sum_{f = u,d} \left[ -\sqrt{\mu^2 - \tilde{F}_f^2} \left( \frac{\mu^3}{6} + \frac{\tilde{F}_f \mu}{4} \right) - \frac{\tilde{F}_f \mu^3}{3} \arctan \frac{\tilde{F}_f}{\sqrt{\mu^2 - \tilde{F}_f^2}} \right. \right. \left. \left. + \frac{\tilde{F}_f}{12} \ln \frac{\mu + \sqrt{\mu^2 - \tilde{F}_f^2}}{\tilde{F}_f} \right] + \frac{F^2}{2G} + O(B^2), \quad \text{for } 0 < F \leq \mu, \quad (3.7)$$

Thus, the spontaneous magnetization per unit volume, $\mathcal{M}$, can be derived thorough the thermodynamic relation, namely, $\mathcal{M} = -\partial \Phi / \partial B|_{B=0} = -\sum_{f = u,d} \partial \Phi / \partial \tilde{F}_f|_{B=0} \cdot \partial \tilde{F}_f / \partial B$. From (3.7), we can derive the spontaneous magnetization originated form the anomalous magnetic moment $\mu_u$, $\mu_d$ and the spin polarization $F$:

$$\mathcal{M} = -\left. \partial \Phi_\geq / \partial B \right|_{B=0} = \frac{\mu^3}{4\pi} (\mu_u + \mu_d), \quad (3.8)$$

$$\mathcal{M} = -\sum_{f = u,d} \frac{\partial \Phi_\leq}{\partial \tilde{F}_f} \left. \frac{\partial \tilde{F}_f}{\partial B} = \frac{F}{2G} (\mu_u + \mu_d), \quad \right. \quad (3.9)$$

where the gap equation $\partial \Phi / \partial F = \sum_{f = u,d} \partial \Phi / \partial \tilde{F}_f + \partial (F^2/(2G)) / \partial F = 0$ is used when Eq.(3.9) is derived. It should be noted that if $F = 0$, the spontaneous magnetization disappears because $F = 0$ in Eq.(3.9). In Figure 3, the spin polarization $F$ and the spontaneous magnetization $\mathcal{M}$ are depicted as a function of the quark chemical potential $\mu$. For $\mu < \mu_c \approx 0.407$ GeV, the magnetization does not occur because the spin polarization does not appear, $F = 0$. It is shown that, at $\mu = \mu_c$, the spontaneous magnetization appears connected with the spin polarization $F$. 
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Finally, let us roughly estimate the strength of the magnetic field due to the spontaneous magnetization. If the spontaneous magnetization per unit volume \( \mathcal{M} \) can be regarded as the magnetic dipole moment, we can simply calculate the strength of the magnetic field yielded by the spontaneous magnetization. As is well known in the classical electromagnetism, the magnetic flux density \( B(r) \) created by the magnetic dipole moment \( m \) can be expressed as

\[
B = \frac{\mu_0}{4\pi} \left[ \frac{m}{r^3} + \frac{3r(m \cdot r)}{r^5} \right],
\]

where \( \mu_0 \) represents the vacuum permeability. In our case, \( m = (0, 0, \mathcal{M} \times V) \), where \( V \) represents a volume. Here, let us consider the hybrid star with quark matter in the core of neutron star. Let us assume that the hybrid (neutron) star has radius \( R = 10 \) km. If there exists quark matter in the inner core of star from the center to \( r_q \) km, the strength of the magnetic flux density on the surface at the north or south pole of hybrid star is roughly estimated. Figure 4(a) and (b) shows the surface magnetic flux density as a function of the quark chemical potential \( \mu \) in the case (a) \( r_q = 1 \) km and (b) \( r_q = 2.15 \) km where quark matter occupies (a) 0.1% and (b) 1% of the total volume of star. It is known that the strength of the magnetic field at the surface of magnetar is near \( 10^{15} \) G. Thus, in our calculation, if quark matter exists and the spin polarization occurs, a strong magnetic field about \( 10^{13} \) or \( 10^{14} \) Gauss may be created.

Figure 3: (a) The spin polarization \( F \) and (b) the spontaneous magnetization per unit volume \( \mathcal{M} \) are depicted as functions with respect to the quark chemical potential \( \mu \).

Figure 4: The \( z \)-component of the magnetic flux density is shown in a logarithmic scale as a function with respect to the quark chemical potential \( \mu \). (a) The case that the quark matter occupies 0.1% in the hybrid compact star and (b) the case that the quark matter occupies 1%. 

Figure 3: (a) Spin Polarization : \( F \) and (b) Spontaneous Magnetization : \( M \) are depicted as functions with respect to the quark chemical potential \( \mu \).

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4. Summary

In this paper, it has been shown that a possibility of the spin polarized phase in quark matter at finite temperature and density is indicated by using the NJL model with the tensor-type four-point interaction between quarks as an effective model of QCD. When there exists the spin polarization, the spontaneous magnetization may appear if the effect of the anomalous magnetic moment of quark is taken into account. It is interesting to investigate what phase is realized under the external strong magnetic field created by the ultrarelativistic heavy-ion collision experiments and expected in the inner core of the compact stars such as neutron star and magnetar. The results of this subject will be reported elsewhere [19].

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