Centrifugal Effects in a Bose-Einstein Condensate in the TOP-Magnetic Trap

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Single particle states in the atomic trap employing the rotating magnetic field are found using the full time-dependent instantaneous trapping potential. These states are compared with those of the effective time-averaged potential. We show that the trapping is possible when the frequency of the rotations exceeds some threshold. Slightly above this threshold the weakly interacting gas of the trapped atoms acquires the properties of a quasi-1D system in the frame rotating together with the field. The role of the atom-atom interaction in changing the ideal gas solution is discussed. We show that in the limit of large numbers of particles the rotating field can be utilized as a driving force principally for the center of mass motion as well as for the angular momentum $L = 2$ normal modes of the Bose condensate. A mechanism of quantum evaporation forced by the rotating field is analyzed.

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I. INTRODUCTION

The novel methods\textsuperscript{[1–3]} for storing atoms at very low densities and temperatures open up new opportunities for studying the role of the atom-atom interaction in macroscopic quantum phenomena\textsuperscript{[4]}. The problems of Bose-Einstein condensate formation\textsuperscript{[5–7]}, the dynamical response of the condensate in the trap\textsuperscript{[8]}, and the interaction of the condensate with light\textsuperscript{[9]} can now be investigated experimentally.

The zero dimensional geometry and the small size of the atomic traps restrict the direct observation of the most spectacular effects known from the history of superfluidity of HeII (see in, e.g., \textsuperscript{[10]}). Therefore devising new practical methods for probing the condensate in the atomic traps becomes of crucial importance. In this regard the recent suggestion\textsuperscript{[11]} to analyze the rotational properties of the trapped atomic cloud appears to be very promising. As was pointed out in Refs.\textsuperscript{[12]}, an analysis of quantum evaporation from the condensate can yield valuable information about the interatomic interaction. Therefore, adopting this analysis to the trapped gases is highly desirable. The very recent theoretical\textsuperscript{[13]} and experimental\textsuperscript{[14]} analyses of a Bose-Einstein condensate undergoing variations of the trapping potential address the long standing question regarding the coherent versus dissipative behavior of a many body system.

The trap\textsuperscript{[15]} where the Bose-Einstein condensation of Rb atoms was first achieved\textsuperscript{[1]}, utilizes a rapidly rotating magnetic field (RMF). This field, if averaged over the rotational period, creates an effective static oscillator potential (which is called the time orbiting potential (TOP)\textsuperscript{[15]} and reduces the escape of atoms from the trap due to spin-flip effects\textsuperscript{[15]}. Other traps\textsuperscript{[2,3]} (see also\textsuperscript{[16]}) do not rely on the RMF. A main assumption made about the RMF is that as long as the frequency of rotations $\omega$ is much larger than the frequency $\omega_o$ of oscillations in the TOP, the trapped atoms are not disturbed by the time variations of the instantaneous potential. Accordingly, the results\textsuperscript{[17,18]} obtained for the Bose-Einstein condensation in a static parabolic potential can be applied to this case as well.

Generally speaking, the RMF should transfer energy and angular momentum to the condensate. Therefore, the RMF can be viewed as a possible tool for studying the dynamical response of the condensate. In this sense, addressing the problem of the exact description of the quantum atomic states in the trap employing the RMF, rather than relying on the time averaging procedure\textsuperscript{[15]}, appears to be quite important.

In this paper we study various aspects of the RMF: we find the exact single particle states in the trap\textsuperscript{[15]} without relying on the time-averaging procedure; it is shown that the RMF, if properly modulated, should excite selectively some modes of the condensate; we derive the Ginzburg-Gross-Pitaevskii (GGP) equation taking into account the effects of the quantum evaporation induced by the RMF. As an application, the decay rate of the condensate due to the RMF is calculated in the case of steady rapid rotations of the RMF.

The outline of the paper is as follows. In Section II we find the exact eigenenergies and eigenstates for a single particle in the trap\textsuperscript{[15]}. This solution is obtained in the frame rotating together with the RMF. The properties of the solution as a function of $\omega$ are analyzed. In Sec.III we consider the limit of large numbers of particles in the condensate and analyze the nondissipative interaction between the RMF and the normal modes of the condensate. In Sec.IV the GGP equation with dissipation due to the RMF is derived under certain approximations. The quasi-static solution for the rate of the centrifugal evaporation of the condensate is derived in the limit of high $\omega$. 

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II. SINGLE PARTICLE STATES IN THE ROTATING FRAME

In our analysis of the behaviour of a single atom trapped by the magnetic field $B$ we follow the approximation that the atomic spin orientation is parallel to $B$. Then the effective potential energy of the atom seeking the low field is essentially the Zeeman energy $U = |\mu_B B|$, where $\mu_B$ stands for the Bohr magneton (we ignore the nuclear magnetic moment). The magnetic field of the trap consists of the static quadrupolar part $B_q$ having axial symmetry with respect to the z-axis, and the RMF $B_b$ rotating in the $x, y$ plane. Representing these explicitly, one finds the components

$$B_{qx} = B'_qx, \quad B_{qy} = B'_qy, \quad B_{qz} = -2B'_qz,$$

$$B_{bx} = -B_b\cos(\omega t), \quad B_{by} = -B_b\sin(\omega t), \quad B_{bz} = 0,$$

where $B'_q, B_b$ stand for the constant gradient of the quadrupolar field, and the constant amplitude of the RMF, respectively. Given Eq.(1), the potential energy

$$U = |\mu_B (B_q + B_b)|$$

depends on time. As suggested in Ref. [15], for high $\omega$ the time dependence can be effectively averaged over, which results in the TOP-potential [13]. In general, the time dependence in $U$ should result also in the nonadiabatic exchange of energy between the external field and the atoms in the trap. However, as will become clear from the following, for a time independent $\omega$ and noninteracting atoms, no such an exchange occurs between the RMF and the atoms in the trap. In fact, in the frame rotating together with the RMF, the single-particle Hamiltonian becomes time independent insuring that the atom once prepared in the pure state (in the rotating frame) will live forever in such a state. Going to the rotating frame implies the coordinate transformation

$$x'' = \cos(\theta(t))x + \sin(\theta(t))y$$

$$y' = -\sin(\theta(t))x + \cos(\theta(t))y$$

where $\theta(t)$ stands for the angle between $B_b$ and the $x$-direction. In the case of steady rotations, $\theta(t) = \omega t$. The transformation (3) results in (2) rewritten in the time independent form as

$$U = |\mu_B B'_q|\sqrt{(x'' - x_o)^2 + y'^2 + 4z^2}, \quad x_o = \frac{B_b}{B'_q}$$

Since the effective size of the atomic cloud is much less then $x_o$, one can expand (4) in terms of $1/x_o$. This gives for the first two terms (linear and quadratic)

$$U = \frac{\omega_{qy}^2 y'^2}{2} + \frac{\omega_{qz}^2 z^2}{2} - x''$$

where we have omitted the unimportant constant $\omega_{qy} = 1/|x_o|, \omega_{qz} = 4/|x_o|$ are introduced, and the units of energy and length are employed as

$$\varepsilon_o = \frac{\hbar^2}{Ml_o^2}, \quad l_o = \frac{\hbar^{2/3}}{(M|\mu_B B'_q|)^{1/3}}$$

respectively. In (6), $M$ stands for the atomic mass. Note that in the rotating frame the stiffness of the potential along the $x''$-coordinate is zero. This would imply that no states localized around the origin exist. However, as we will see below, the finite kinetic energy of the particle changes this conclusion for sufficiently large $\omega$.

In the rotating frame the kinetic energy acquires the Coriolis term $-\omega L$, where $L$ is the $z$ component of the angular momentum operator and $\omega = \dot{\theta}$. Consequently, taking into account (5), one finds the single particle Schrödinger equation ($\hbar = 1$) in the rotating frame

$$i\partial_t \psi = H_\omega \psi,$$

$$H_\omega = -\frac{1}{2}\Delta + i\omega (x''\partial_{y'} - y'\partial_{x''}) + U$$

(7)
We consider first the case $\omega = \text{const}$. Note that Eqs. (5), (7) represent a quadratic form which can be diagonalized explicitly (see Appendix A). Prior to solving it let us eliminate the linear term $-x''$ from (5). This can be accomplished by the transformation

$$\psi \to \exp(-i\frac{y'}{\omega})\psi(x', y', z), \quad x' = x'' - \frac{1}{\omega^2}$$

which results in Eq. (7) being rewritten as

$$i\partial_t \psi = H'_\omega \psi,$$

$$H'_\omega = -\frac{1}{2} \Delta + i\omega(x' \partial_{y'} - y' \partial_{x'}) + \frac{\omega_o^2 y'^2}{2} + \frac{\omega_z^2 z'^2}{2}.$$

For the case $|\omega| < \omega_o$, no discrete states localized near the origin $x' = y' = z' = 0$ exist in the trap. Accordingly, we will not analyze this case any more. For $|\omega| > \omega_o$ such states do exist. Their eigenenergies are (see Appendix A)

$$\varepsilon_{mnl} = \omega_m m - \omega_o n + \omega_o z l, \quad \omega_m = \omega \sqrt{1 + \eta^2/2 \pm \eta \sqrt{2 + \eta^2}/4}, \quad \eta = \omega_o/\omega;$$

where $m, n, l$ are integer nonnegative quantum numbers, and the energy of the state with $m = n = l = 0$ is set equal to zero. The normalized eigenfunctions are (see Appendix A)

$$\psi_{mnl}(x', y', z) = \frac{(\omega_m \omega_1 \omega_2)^{1/4}}{\pi^{3/4} \sqrt{2^{m+n+l} m! n! l!}} e^{-\Xi_o} \left[ \frac{\partial^{m+n+l}}{\partial t_1^m \partial t_2^n \partial t_3^l} e^{\Xi} \right]_{t_1 = t_2 = t_3 = 0}, \quad (11)$$

where we have introduced

$$\Xi_o = \frac{\omega_1 x'^2}{2} + \frac{\omega_2 y'^2}{2} + \frac{\omega_o z'^2}{2} - i\gamma_o x' y',$$

$$\Xi = \frac{\alpha^2 - \alpha^{-2}}{2} (t_1^2 - t_2^2) - t_3^2 - 2\nu t_1 t_2 + 2\sqrt{\omega_o t_3} z +$$

$$+ \sqrt{2}[\alpha^{-1}\sqrt{\omega_1 x'} - i\nu \alpha \sqrt{\omega_2 y'}] t_1 + \sqrt{2}[\nu \alpha \sqrt{\omega_1 x'} + i\alpha^{-1}\sqrt{\omega_2 y'}] t_2. \quad (12)$$

In Eq. (12) the parameters are

$$\gamma_o = \omega \frac{\omega_2 - \omega_1}{\omega + \omega_1}, \quad \omega_2 = \sqrt{1 - \eta^2} \omega_1, \quad \eta = \frac{\omega_o}{\omega}, \quad \nu = \text{sign}(\omega),$$

$$\omega_1 = \sqrt{8 - 8\sqrt{1 - \eta^2} - 4\eta^2 - \eta^4 \eta^{-2} \omega}, \quad \alpha = \frac{(1 - \frac{\eta^2}{2} - \frac{\eta}{2} \sqrt{8 + \eta^2})^{1/2}}{(1 + \frac{\eta^2}{2} - \frac{\eta}{2} \sqrt{8 + \eta^2})^{1/4}(1 - \eta^2)^{1/8}}. \quad (13)$$

Note that the parameter $\alpha$ as a function of $\omega$ (or $\eta$) has the property $\alpha(-\eta)\alpha(\eta) = 1$. In the limit $\eta \to 0 (\omega \to \infty)$ one obtains $\alpha = 1$, and the eigenfunctions (11), (12) as well as the eigenenergies (10) become exactly those characterizing the TOP as seen from the rotating frame. In particular, the spectrum acquires the form

$$\varepsilon'_{mnl} = \omega_0 (m + n) - \omega(n - m) + \omega_o z l \quad (14)$$

where $\omega_o = \omega_o/\sqrt{2}$ is the frequency of the oscillations in the $x, y$-plane of the TOP.

For large $\omega$, Eq. (14) is an approximation of the exact expression (10). The corrections due to the finiteness of $\omega_o/\omega$ turn out to be of the order of $(\omega_o/\omega)^3$, so that one can effectively ignore these even if $\omega$ is only a few times larger than $\omega_o$. Consequently, for such $\omega$ we will employ Eq. (14) instead of the exact form (10).

In the limit $|\eta| \to 1$ from below, the solution (10)-(13) acquires features characteristic of a quasi-1D system. Indeed, taking this limit in Eqs. (10)-(13), one finds

$$\omega_+ = \omega_1 = \sqrt{3}\omega_0, \quad \omega_- = \frac{1}{3}\omega_2 \approx \sqrt{\frac{2}{3}(1 - |\eta|)} \omega_o. \quad (15)$$

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This expression together with Eqs. (11), (12) imply that the typical extension in the \( y' \)-direction diverges as \((1 - |\eta|)^{-1/4} \to \infty\). Accordingly, the excitation spectrum (10) becomes characterized by the soft mode whose energy \( \omega \) goes to zero. This implies that in this region of \( \omega \) the low energy dynamical response of the system of atoms should exhibit the 1D behavior. In this paper we will not focus on the properties of such a 1D atomic gas.

Note that the spectrum (10) and its limiting form (14) have no lower bound. For the model Hamiltonian whose potential is the axially symmetric TOP \([15]\), this is a pure consequence of the coordinate transformation (3) because this Hamiltonian conserves angular momentum. Accordingly, no instability with respect to a spontaneous growth needs to be taken into account. This expression together with Eqs. (11), (12) imply that the typical extension in the \( y' \)-direction diverges as \((1 - |\eta|)^{-1/4} \to \infty\). Accordingly, the excitation spectrum (10) becomes characterized by the soft mode whose energy \( \omega \) goes to zero. This implies that in this region of \( \omega \) the low energy dynamical response of the system of atoms should exhibit the 1D behavior. In this paper we will not focus on the properties of such a 1D atomic gas.

Another interesting feature of the solution (11), (12) is the common phase factor \( \exp(i\gamma_o x' y') \). Its magnitude is controlled by the asymmetry of the eigenfunctions in the \( x, y \) plane (in addition to the squeezing in the \( z \)-direction). Also, this factor implies a very specific pattern for the velocity \( \mathbf{v} = \nabla \text{Im}(\ln(\psi_00l)) \) at the levels with \( m = n = 0 \). Employing (11), (12) for \( m = n = 0 \) one finds

\[
t_{(r)} = \gamma_o r \sin(2\theta), \quad t_{(\theta)} = \gamma_o r \cos(2\theta)
\]

for the radial and polar components of \( \mathbf{v} \), respectively (in the polar coordinates \( x' = r \cos(\theta), \ y' = r \sin(\theta) \)). This expression exhibits quadrupolar symmetry. As long as particles are condensing into the state with opposite vorticity coupled together. This pattern can be thought of as two pairs of vortices of opposite vorticity coupled together.

Above we have shown that no nonadiabatic energy exchange occurs between the RMF and the ideal gas in the trap. In the Sec.IV, we will show that the interaction between particles changes this situation.

### III. CONDENSATE CONTAINING LARGE NUMBERS OF PARTICLES IN THE ROTATING FRAME

In this section we will analyze the case of the condensate containing large numbers of particles in the presence of the RMF. The condensate wave function \( \Phi \) obeys the GGP equation \([19]\). For the single-particle Hamiltonian (7) in the rotating frame, this equation is

\[
i\partial_t \Phi = (H_o - \mu)\Phi + u_o|\Phi|^2\Phi
\]

where \( u_o > 0 \) is the interaction constant and \( \mu \) stands for the chemical potential. Following the approach \([8]\), we will derive approximate hydrodynamical equations for the condensate in the presence of the RMF. We denote

\[
\Phi = \sqrt{\rho}e^{ip}, \quad \int d\mathbf{r}\rho = N_c,
\]

where \( \rho, \phi \) and \( N_c \) are the density, the phase, and the total number of particles in the condensate, respectively. Substituting (18) into (17), one arrives at the expressions

\[
\dot{\rho} - \omega(x''\partial_y^2\rho - y''\partial_x^2\rho) + \nabla(\rho\nabla\phi) = 0,
\]

\[
\dot{\phi} - \omega(x'\partial_y\phi - y'\partial_x\phi) = \frac{\Delta(\sqrt{\rho})}{2\sqrt{\rho}} - \frac{1}{2}(\nabla\phi)^2 + \mu - U - u_o\rho.
\]

In the rotating frame. The main approximation made in the limit of large \( N_c \) is that the term proportional to \( \Delta(\sqrt{\rho}) \) in the second of equations (19) can be neglected \([3,8]\). In the limit \( \omega \to \infty \), one expects to obtain a solution of (19) which is close to that characterizing the TOP \([3,8]\). In order to see it, one should separate a rotationally invariant part from the total potential (5). Specifically,

\[
U = U_{TOP} + \delta U,
\]

\[
U_{TOP} = \frac{\omega_o^2}{2}(x''^2 + y'^2) + \frac{\omega_o^2}{2}z'^2, \quad \delta U = -\frac{\omega_o^2}{2}(x''^2 - y'^2) - x''.
\]

For the sake of convenience we will omit all primes \((x'' \to x, \ y' \to y)\) from the coordinates, implying that we are working in the frame connected with the RMF unless otherwise stated. Note that \( U_{TOP} \) is the time averaged potential.
(TOP) derived in Ref. [15]. The term $\delta U$ describes the deviations of the instantaneous potential (5) from $U_{TOP}$. It is not strictly obvious that $\delta U$ can be treated as a small correction to $U_{TOP}$. However, the exact results obtained above for the single-particle Hamiltonian show that this is true in the limit of large $\omega$ at least. Below we will show that if $\omega >> \omega_o$, corrections to the solution (21) caused by $\delta U$ remain small for large $N_e$ as well.

To the zeroth order with respect to $\delta U$, one obtains from Eq.(19) the solution

$$\phi^{(o)} = 0, \quad \rho^{(o)} = \frac{1}{u_o} (\mu - U_{TOP}),$$

which is valid inside the droplet whose radius is determined by the condition $\rho^{(o)} = 0$ [8]. We represent $\rho = \rho^{(o)} + \delta \rho$, where $\delta \rho$ is a small correction due to $\delta U$. Correspondingly, we ignore the term $\nabla (\delta \rho \nabla \phi)$ in the first equation (19). Linearizing Eqs.(19) in $\delta \rho, \phi$ [8], one obtains

$$\delta \dot{\rho} - \omega (x \partial_y \delta \rho - y \partial_x \delta \rho) + \nabla (\rho^{(o)} \nabla \phi) = 0,$$

$$\delta \phi - \omega (x \partial_y \phi - y \partial_x \phi) = -u_o \delta \rho - \delta U.$$  

(22)

Note that these equations are a close analog to those obtained in [8] for a trapping oscillator potential which is spherically symmetric. The additional feature of Eqs.(22) is the term $\delta U$ which plays the role of an external force. Later we will see that this term under certain conditions can resonantly excite the condensate normal modes with the angular momenta $L = 1, 2$.

A particular solution corresponding to the symmetry of the driving term $\delta U$ (20) can be taken in the form

$$\delta \rho = \rho_2'(x^2 - y^2) + 2 \rho_2'' xy + \rho_1' x + \rho_1'' y,$$

$$\phi = \phi_2'(x^2 - y^2) + 2 \phi_2'' xy + \phi_1' x + \phi_1'' y,$$

(23)

where $\rho_1', \rho_1'', \phi_1', \phi_1'', l = 1, 2$ are the time dependent amplitudes of the dipole $(l = 1)$, and the quadrupole $(l = 2)$ harmonics. Substitution of (21) and (23) into (22) yields

$$\dot{\phi}_l + i \omega \phi_l + u_o \rho_l = \frac{\omega_o^2}{2} \rho_l^{(o)};$$

$$\dot{\rho}_l + i \omega \rho_l - l \frac{\omega_o^2}{u_o} \phi_l = 0.$$  

(24)

for the complex amplitudes

$$\rho_l = (\rho_l' + i \rho_l'') e^{il\omega t}, \quad \phi_l = (\phi_l' + i \phi_l'') e^{il\omega t}.  $$

(25)

Note that the dipole amplitudes $l = 1$ describe essentially the center of mass motion of the whole atomic cloud in the trap [8].

If the RMF frequency $\omega$ does not change in time, one obtains the steady solutions ($\dot{\rho} = \dot{\phi} = 0$)

$$\phi_1' = 0, \quad \phi_1'' = \frac{\omega}{\omega_o^2 - \omega^2}, \quad \rho_1' = \frac{\omega_o^2}{u_o (\omega_o^2 - \omega^2)}, \quad \rho_1'' = 0,$$

(26)

and

$$\phi_2' = 0, \quad \phi_2'' = \frac{\omega}{4 \omega_o (1 - \frac{\omega^2}{\omega_o^2})}, \quad \rho_2' = \frac{\omega_o^2}{2 u_o (2 \omega_o^2 - \omega^2)}, \quad \rho_2'' = 0.$$  

(27)

This implies that in the limit $\omega >> \omega_o$ the corrections due to the RMF to the zeroth order solution (21) [8] are small.

Note that Eqs.(27) and (23) indicate that the phase factor $\exp(i \gamma_o xy)$, discussed in the Sec.II for the ideal gas situation, is not affected much by the interaction as long as $\omega >> \omega_o$. Indeed, comparing Eqs.(11)-(13) with Eqs.(23),(27), one finds that the parameter $\gamma_o$ in Eqs.(12), (13), (16) must be replaced by $2 \delta \phi_2'' = \gamma_o + o((\omega_o/\omega)^3)$. When $\omega \rightarrow \omega_o$, the solutions (26), (27) based on the condition $\rho^{(o)} >> \delta \rho$ become no longer valid.

We now consider the case when $\omega$ depends on time. For concreteness, we assume that the frequency $\omega$ of the RMF is modulated as
\[ \omega = \overline{\omega} + \lambda \sin(\omega't) \]  

(28)

where \(|\overline{\omega}| >> |\lambda|\), and \(\omega'\) are constants. Accordingly, one finds that the angle between the RMF and the \(x\)-axis in the laboratory frame is

\[ \theta(t) = \overline{\omega}t + \frac{\lambda}{\omega'} (1 - \cos(\omega't)) \]  

(29)

Employing the ansatz

\[ \phi_l = e^{-il\theta(t)} \tilde{\phi}_l, \quad \rho_l = e^{-il\theta(t)} \tilde{\rho}_l, \]  

(30)

one obtains from Eqs.(24)

\[ \tilde{\phi}_l = \frac{u_\omega}{l\omega_o} \tilde{\rho}_l, \quad \tilde{\rho}_l + l\omega_o^2 \tilde{\rho}_l = \frac{\omega_o^2}{u_\omega} e^{i\theta(t)}. \]  

(31)

These equations indicate that the resonance condition on \(\omega'\) in (28) is different for the dipole \((l = 1)\) and quadrupole \((l = 2)\) harmonics. Indeed, given (29) and expanding the r.h.s. of the second equation of (31) in the small quantity \(|\lambda/\omega'| << 1\), we get

\[ e^{il\theta(t)} \approx e^{i\overline{\omega}t} (1 + \frac{il\lambda}{\omega} (1 - \cos(\omega't))). \]  

(32)

Then, one obtains that the resonance with the dipole harmonic occurs when the modulating frequency \(\omega'\) obeys the condition

\[ \omega' = \omega'_1 = \overline{\omega} \pm \omega_o. \]  

(33)

Eqs.(31) yield the resonance condition for the quadrupolar harmonic

\[ \omega' = \omega'_2 = 2\overline{\omega} \pm \sqrt{2}\omega_o, \]  

(34)

where \(\sqrt{2}\omega_o\) stands for the frequency of the lowest quadrupolar harmonic of the trapped condensate with large \(N_c\) \[8\]. Note that \(\omega'_2 - \omega'_1 \approx \overline{\omega} >> \omega_o\).

In the following, we will show that the preceding analysis based on the GGP equation does not take into account quantum processes of the creation of pairs out of the condensate by the RMF. These lead to forced evaporation of the condensate even for zero temperature and steady rotations of the magnetic field. As a consequence, the hydrodynamical equations (22) will acquire a dissipative term.

**IV. CENTRIFUGAL INSTABILITY IN THE MANY BODY APPROACH**

In our previous analysis, we neglected quantum fluctuations of the condensate. These fluctuations in the conventional stable condensate can be thought of as the virtual creation and absorption of pairs. In this regard we note that the spectrum (10) has no lower bound, so that the condensate could be unstable with respect to the real creation of pairs even though the RMF is steady. Correspondingly, the GGP equation (17) can acquire a dissipative part.

Consider first the case \(\omega -> \infty\). The term proportional to \(\omega\) in Eq.(14) is a direct consequence of the Galilean transformation into the rotating frame. Indeed, the limit \(\omega -> \infty\) in Eqs.(10)-(13) insures that the term \((n - m)\) is the projection of the angular momentum \(L\) on the \(z\)-axis, so that the \(\omega\)-dependent part in (14) is exactly the Coriolis contribution \(-\omega L\). This implies that no instability should develop because the absence of the lower bound for the spectrum is purely a frame of reference effect. Nevertheless the condensate can be considered as being potentially able to gain high values of \(L\). In this regard we can employ the rotating frame reasoning \[10\] (see also \[10\], Ch.6) for the vortex creation in the rotating vessel containing a superfluid. In the frame connected to the vessel rotating with the frequency \(\omega\) around its axis, the vortex energy is \(E_v - \omega L_v\) where \(E_v\) and \(L_v\) stand for the vortex energy in the laboratory frame and the vortex angular momentum, respectively. The vortex can be created spontaneously if the Coriolis energy exceeds \(E_v\). However, this argument does not indicate what is the probability for developing this centrifugal instability. In fact, in the case of the perfectly symmetric vessel this probability is essentially zero. To make the vortex creation real, the vessel must have some irregularities on the walls breaking the rotational symmetry so that the angular momentum of the vessel could be transferred to the vortex (or vortices).
Returning to our case, we can see that in the case $\omega \to \infty$ the eigenfunctions (11), (12) of the trap are approaching those of the effective time averaged Hamiltonian (the TOP) which is axially symmetric. Therefore, no centrifugal instability of the condensate is expected to occur in this limit. In other words, no energy exchange between the RMF and the condensate happens in the limit $\omega = \infty$.

For finite $\omega$, the functions (11), (12) are not eigenfunctions of the operator $L$. Consequently, the difference $n - m$ can no longer be interpreted as the eigenvalue of $L$. Accordingly the effective vessel can be thought of as having a symmetry breaking deformation, which in turn implies that the energy and the angular momentum can now be given up to the pairs leaving the condensate into the highly excited states whose energies are $\hbar \omega \gg \hbar \omega_o$ (in physical units). In this regard one should distinguish two cases: 1) $\hbar \omega \gg \mu$ and 2) $\hbar \omega \sim \mu$. In the case 1) the pair escapes into states lying far from those effectively involved in the formation of the interacting condensate. Accordingly, the pair escape process can be treated as an incoherent step in the condensate evaporation. In contrast, in case 2) the escape states with the energies $\approx \hbar \omega$ are to be renormalized strongly because of the presence of the condensate. This implies that the multi-pair processes become significant. Correspondingly, the centrifugal instability should be interpreted as a coherent process of vortex formation. In this paper we will not analyze this case.

The process of the escape of pairs represents the nonresonant quantum evaporation of the condensate induced by the RMF. We emphasize the crucial role of the interatomic interaction for realization of this centrifugal evaporation. From the point of view of the rotating observer this process can be described as follows: two atoms in the condensate interact with each other. As a consequence, they jump to a new pair of single-particle states characterized by large quantum numbers, so that their total energy is conserved. Correspondingly, the rotating observer interprets this event as a nearly elastic escape of the pair from the condensate. Note that if the eigenfunctions (11) were eigenfunctions of angular momentum, there would be a selection rule requiring that the angular momentum of the interacting pair not change in the transition. Correspondingly, referring to Eq.(14), one sees that no instability would occur. In fact this is not the case for finite $\omega$ and instability could occur.

To describe the centrifugal instability effect, we proceed to derive a damping term in the GGP equation for the condensate wave function $\Phi$. The many body Hamiltonian in the rotating frame is

$$H = \int d\mathbf{x} [\Psi^{\dagger}(H_\omega - \mu)\Psi + \frac{\hbar^2}{2m}\Psi^{\dagger}\Psi]$$

where primes are omitted from the coordinates and the Bose operators $\Psi^{\dagger}$, $\Psi$ obey the usual Bose commutation rule. The Heisenberg equation is

$$i\partial_t \Psi = (H_\omega - \mu)\Psi + u_o\Psi^{\dagger}\Psi.$$  \hspace{1cm} (36)

Taking into account the explicit form (7) for $H_\omega$, one finds from (36) the current conservation condition

$$\partial_t (\Psi^{\dagger}\Psi) + \nabla J = 0,$$  \hspace{1cm} (37)

where the current operator $J$ in the rotating frame is defined as

$$J = \frac{1}{2i} [\Psi^{\dagger}(\nabla - \mathbf{A})\Psi],$$  \hspace{1cm} (38)

$$A_x = -i\omega y, \quad A_y = i\omega x, \quad A_z = 0.$$

In the presence of the condensate, the condensate wave function $\Phi = <\Psi>$. The noncondensate part $\Psi' = \Psi - \Phi$. From Eq.(36), one finds

$$\partial_t \Phi = (H_\omega - \mu)\Phi + u_o|\Phi|^2\Phi +$$

$$+ u_o[\Phi^{\dagger}\Psi^{\dagger}\Psi' + 2\Phi <\Psi'\Psi'> + <\Psi^{\dagger}\Psi\Psi'>],$$  \hspace{1cm} (39)

and

$$i\partial_t \Psi' = (H_\omega - \mu)\Psi' + u_o[\Phi^{\dagger}\Phi^{\dagger}\Psi' - <\Psi'\Psi'>] +$$

$$+ 2\Phi^{\dagger}\Phi\Psi' + \Phi^{\dagger}2\Psi\Psi' + \Psi^{\dagger}\Psi'\Psi' - <\Psi'\Psi'\Psi'>].$$  \hspace{1cm} (40)

The condensate wave function $\Phi$ is normally viewed as an external classical field in Eq.(40) for the noncondensate part. Employing the Keldysh technique, the system (39), (40) can be expressed in terms of the joint dynamics of $\Phi$ and the population numbers of the excitations. In general, this procedure is very complicated (see also the
the summation is performed over the final states of the escaping pair. Corresponding to \( \omega \) density \( \rho \), where we have used the second relation in Eq.(18). Note that in this equation the integral depends on both the

\[ \sqrt{\text{employ the ideal gas ansatz } \Phi = \text{const.}} \]

We set \( N \) where \( \rho \) is not be possible to obtain the dissipation term in (41) in closed form [5]. Below it will be shown explicitly that the last term in Eq.(41) would have been zero if either the single particle Hamiltonian \( H_\omega \) conserved angular momentum or if the single particle excitation spectrum were positively defined.

As will be seen below, in the limit \( \omega \to \infty \) the escape rate is much smaller than the typical time scale in the trap corresponding to \( \omega \). Consequently, one can employ the quasistatic approximation that the time dependence of \( \hat{N}_c \) is slow. We set \( \hat{N}_c(t') = \hat{N}_c(t) \) in (44) and after the time-integration rewrite it as

\[ \hat{N}_c = -\chi \hat{N}_c^2, \]

where

\[ \chi = 2\pi u_o^2 \sum_{[1],[2]} |M_{12}|^2 \delta(E_{12}). \]
In Eq.(47), the notation $E_{12} = \varepsilon_1 + \varepsilon_2$ for the escaping pair energy is introduced and $\delta(\xi)$ stands for the $\delta$-function. This expression accounts quantitatively for the centrifugal effect discussed above. One can see that the condition $M_{12} \neq 0$ and $E_{12} = 0$, where explicitly

$$E_{12} = \omega_o(m_1 + m_2 + n_1 + n_2) + \omega_o(l_1 + l_2) - \omega(n_1 + n_2 - m_1 - m_2) = 0,$$  \hspace{1cm} (48)

can be satisfied simultaneously because the eigenfunctions (11), (12) utilized in (45) and (47) are not the eigenfunctions of the angular momentum operator. In what follows, we will show that in the limit of large $\omega$, the dominant contribution in (47) comes from the states with the quantum numbers

$$m_1 \approx m_2 \approx l_1 \approx l_2 \approx n_1 \approx n_2 \approx \frac{\omega}{\omega_o} >> 1$$  \hspace{1cm} (49)

corresponding to a pair leaving the condensate into the states characterized by large quantum displacements. Returning to the laboratory frame, this simply means that two atoms absorb the energy $2\hbar\omega$ (in physical units) from the RMF so that this energy is approximately equally distributed between them. As a result, the pair is transferred to highly excited states. The radio-frequency scalpel [22] is assumed to eventually remove this pair insuring the condition a) of zero population of the excited states.

In order to calculate $M_{12}$ explicitly, we employ the representation (compare with [21], Ch.15)

$$\psi_{m_1 n_1 l_1}(x_1)\psi_{m_2 n_2 l_2}(x_2) = \sum_{pq} g_{pq}^{m_1 m_2} g_{pq}^{n_1 n_2} g_{pq}^{l_1 l_2} \psi_{pq}(R) \psi_{p'q'}(r)$$  \hspace{1cm} (50)

where the notations $p' = m - p$, $k' = n - k$, $q' = l - q$ and

$$g_{pq}^{ab} = \sum_{p'} (-1)^{p+p'} \sqrt{\frac{1! b! p!(a+b-p)!}{2^{a+b} p'!(a-p')!(b-p')!(b-p+p')!}}.$$  \hspace{1cm} (51)

$R = \frac{\sqrt{x_1 + x_2}}{\sqrt{2}}$, $r = \frac{\sqrt{x_1 - x_2}}{\sqrt{2}}$, $m = n_1 + m_2$, $n = n_1 + n_2$, $l = l_1 + l_2$

are employed. The summations in Eqs.(50), (51) run over all integer nonnegative numbers obeying the condition that all the numbers under the signs of the factorial are nonnegative as well. The relations (50), (51) were derived from the explicit representation (11),(12) for the eigenfunctions. Employing Eqs.(50), (51) in Eqs.(45),(47) we obtain for Eq.(47)

$$\chi = \frac{\pi u_o^2}{4} \sum_{mnl} |\psi_{mn}(0)|^2 |\psi_{00}(0)|^2 \delta(\omega_o(m+n) + \omega_o l - \omega(n-m)).$$  \hspace{1cm} (52)

Note that $\chi$ is exactly zero for the case $\alpha = 1$ (or $\omega \to \infty$) in Eqs.(11)-(13). As mentioned above, no escape of the pairs occurs in the case when the RMF is so rapidly rotating that the effective trapping (TOP) potential becomes axially symmetric. In the limit of large but still finite $\omega$ one finds from Eq.(13)

$$\alpha^2 - \alpha^{-2} = -\left(\frac{\omega_o}{\omega}\right)^3 + o\left(\frac{\omega_o}{\omega}\right)^5$$  \hspace{1cm} (53)

which implies that the first term only should be kept in (53), and that the exponent $e\bar{m}$ in Eq.(11) can be expanded in terms of the smallness of $\omega_o/\omega$. This expansion represents the eigenfunction (11) in terms of the harmonics of the angular momentum operator. Keeping the first term only, one finds that each eigenfunction can be effectively characterized by three terms: i) the harmonic of the angular momentum operator with the angular momentum $(n-m)$; ii) two functions with the momenta $(n-m) \pm 2$ whose weight is proportional to $(\omega_o/\omega)^3 < < 1$. This implies that in Eq.(47) it is enough to consider the contribution due to the lowest term. Physically this term corresponds to an absorption of the energy $2\hbar\omega$ and the angular momentum 2 by a pair of atoms escaping from the condensate. Finally one finds

$$\chi = \frac{1}{4} \alpha^2 \omega_o^2 \omega_\omega \left(\frac{\omega_o}{\omega}\right)^6 \sum_{m} \frac{(2l)!|m+1|!|m+2|!}{2^{2l} l!^2} \delta(2\omega_0 m + 2\omega_\omega l - 2\omega).$$  \hspace{1cm} (54)

A simple analysis shows that most of the contribution to the sum (54) comes from the region of high $m$, $l$ (see (49)). Accordingly, we replace the summation in (54) by integration. Finally, in the chosen limit and chosen units (6) we find
\[ \chi = e a^2 \omega_0^2 \left( \frac{\omega_0}{\omega} \right)^{7/2}, \quad \epsilon = \frac{27/4}{15\sqrt{2}} \approx 0.13. \]  

(55)

This expression indicates that the centrifugal escape rate is extremely sensitive to the RMF frequency \( \omega \). For the parameters employed experimentally in Ref. [1], the estimate of (55) gives a very small number (the corresponding lifetime is about \( 10^6 \) s for \( N_c = 1000 \)), implying that the centrifugal vaporization can be effectively ignored as a cause for the condensate escape from the trap. However, with decrease of the ratio \( \omega/\omega_o \), the vaporization rate increases strongly. In the case \( \omega \to \sqrt{2}\omega_o \) from above, the approximation (53) we employed is no longer valid. Accordingly, the exact expression (11)-(13) for the eigenfunctions should be utilized in (52). This means that the escaping pairs acquire higher (even) angular momenta. As a result, the lifetime of the condensate can become very short.

Eqs. (39) and (40) can be analyzed for the case of a nonsteady RMF. Especially interesting appears to be the case when the RMF excites resonantly the quadrupolar harmonic of the condensate (see the condition (34)). Generally, it is natural to expect that this resonance would result in the increase of the vaporization rate as a function of the modulating frequency \( \omega' \). In the future, we will consider this case in greater detail.

V. CONCLUSION

The atomic trap [15,1] bears features absent in the static traps [2,3,16]. These features can be accounted for in the frame rotating together with the RMF. For large frequencies of rotation of the RMF, the exact eigenenergies and eigenfunctions of the trap [15,1] approach those characterizing the time averaged potential TOP [15] having axial symmetry with respect to the axis of rotation. For frequencies close to the threshold below which the trapping is impossible, the eigenstates lose their axial symmetry and become elongated in the direction perpendicular to the RMF (in its frame of reference). Very close to the threshold a gas of trapped atoms acquires properties of an essentially 1D system.

Due to the asymmetry introduced by the RMF, the atom-atom interaction results in the induced evaporation of the Bose-Einstein condensate. The time scale for this evaporation is very sensitive to the RMF frequency of rotation. For high frequencies, the lifetime of the condensate increases as a large power of \( \omega \). Close to the trapping threshold the lifetime shortens considerably, implying that the 1D gas formed in the trap [1,15] in this situation is a strongly interacting system.

The RMF can be utilized as a driving force selectively exciting the condensate normal modes. In the limit of large numbers of atoms, when Stringari’s hydrodynamical approximation is valid, two modes can be excited by the RMF whose frequency of rotation is appropriately modulated. The first is a dipole mode accounting for the center of mass motion. The second mode which can be excited by the RMF is the lowest quadrupolar harmonic. The resonance conditions for the RMF modulation period depend on the averaged RMF frequency, in addition to the eigenfrequencies of the harmonics. The effect of the quantum evaporation induced by the RMF opens up a channel for dissipation of the condensate normal modes.

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APPENDIX A: SOLVING THE EIGENPROBLEM IN THE ROTATING FRAME

The Hamiltonian (9) for the function \( \psi'(x',y',z) = \exp(i\omega x'y')\psi \) can be written as

\[
H'_x = H_{xy} + H_z, \quad H_z = -\frac{1}{2} \frac{\partial^2}{\partial z^2} + \frac{\omega_0^2}{2} z^2, \\
H_{xy} = -\frac{1}{2} (\partial_x^2 + \partial_y^2) + \frac{\omega_{xy}^2 - \omega^2}{2} y^2 + \frac{3}{2} \omega^2 x^2 + 2i\omega x \partial_y
\]

where the unimportant constant \( 1/(2\omega^2) \) and the primes from the coordinates are omitted. The eigenfunctions \( \varphi_l(z) \) of \( H_z \) are the well known oscillator states. These are represented in (11), (12) by the generating function of the Hermite polynomials (see the auxiliary variable \( t_3 \) in (11), (12)) so that the \( \psi'(x,y,z) = \varphi_l(z)\psi'_{nm}(x,y) \). Performing the Fourier transform
\[ \tilde{\psi}(x, p) = \int d\tilde{y} e^{-ip\tilde{y}} \psi'(x, \tilde{y}), \quad \tilde{y} = \sqrt{\omega^2 - \omega_{gy}^2} y, \] (A2)

one finds for (A1)

\[ H_{xy} = -\frac{1}{2} \partial_x^2 + \frac{2}{2} \partial_y^2 + \frac{\omega^2 - \omega_{gy}^2}{2} p^2 + \frac{3}{2} \omega x^2 - 2\omega \sqrt{\omega^2 - \omega_{gy}^2} xp \] (A3)

where \( s = \text{sign}(\omega^2 - \omega_{gy}^2) \).

In the case \(|\omega| < \omega_{gy}\) the Hamiltonian (A3) can be diagonalized by implementation of a real rotation in the \((x, p)\)-plane. However, no discrete states exist in this case because the effective potential of (A3) turns out to have a saddle-like shape, with the kinetic part being positively defined. In the opposite limit \((s = 1)\) the discrete states do exist. The diagonalization can be achieved by means of the Lorentz transformation

\[ x = \cosh(\vartheta) \xi + \sinh(\vartheta) p' \]

\[ p = \sinh(\vartheta) \xi + \cosh(\vartheta) p' \] (A4)

leaving the kinetic part \(-\frac{1}{2}(\partial_x^2 - \partial_p^2)\) invariant. In terms of the new variables \((\xi, p')\), (A3) acquires the form

\[ H_{\xi p} = -\frac{1}{2} \partial_\xi^2 + \frac{\omega_+^2}{2} \xi^2 - \left[ -\frac{1}{2} \partial_{p'}^2 + \frac{\omega_-^2}{2} p'^2 \right] \] (A5)

where \(\omega_{\pm}\) are given in Eq.(10), and the angle \(\vartheta\) satisfies the equation

\[ \tanh(2\vartheta) = \frac{4\nu \sqrt{1 - \eta^2}}{4 - \eta^2}, \] (A6)

with \(\nu, \eta\) defined in Eq.(13). The resulting spectrum of the total Hamiltonian is given by Eq.(10). The eigenfunctions of (A5), expressed in terms of the \(\xi, p'\) variables, can be converted into the \(x, p\) coordinates by means of the relations (A4). Finally, performing the inverse of the Fourier as well as the scaling transforms (A2), one finds the normalized eigenfunctions (11) - (13).

**APPENDIX B: DERIVATION OF THE GGP EQUATION WITH THE DISSIPATION DUE TO THE RMF**

Under a)-d), Eqs.(39) and (40) simplify considerably. We need to find the lowest order term which contributes to the imaginary part of Eq.(39). In Eq.(39), the second term in the brackets does not contribute to the imaginary part, so we omit it. The last term in the brackets of Eq.(39) produces the imaginary part. However it can be shown that it is proportional to the population numbers of the excited states. Consequently, we omit this term also and rewrite (39) as

\[ i\partial_t \Phi = (H_\omega - \mu)\Phi + u_\omega |\Phi|^2 \Phi + u_\omega \Phi^* G(x, x') \] (B1)

where the equal-time anomalous Green’s function \[21\] is defined as \( G(x, x') = < \Psi'(x')\Psi'(x') > \) (the overall factor \(i\) is omitted here and below in the Green’s functions definitions \[21\]). Note that because the interaction potential in (35) is chosen in the \(\delta(x)\) form, Eq.(B1) contains \(G\) with \(x = x'\). The equation for \(G\) can be obtained from Eq.(40). In order to accomplish this, we will employ the Bogolubov approximation d). Accordingly, multiplying (40) by \(\Psi^\dagger\) and taking the average, one finds

\[ i\partial_t G(x, x') = (H_\omega - \mu)\xi G(x, x') + (H_\omega - \mu)\xi G(x, x') + 
\]

\[ +u_\omega [2(\Phi^*(x)\Phi(xt) + \Phi^*(x't)\Phi(x't))G(x, x') + 
\]

\[ +\Phi^2(x't)G^+(x, x') + \Phi^2(xt)G^+(x, x')] \] (B2)

where the normal Green’s functions for coinciding times are

\[ G^+(x, x') = < \Psi'(x)\Psi'(x') > \]

\[ G^+(x, x') = < \Psi'(x')\Psi'(x') > \] (B3)
In Eq.(B2) the notation \( (\ldots)_x \) means that the single particle Hamiltonian \( H_\omega \) acts on the coordinate \( x \). These equations should be supplemented by ones for the normal Green’s functions. However, as long as one is only interested in deriving the imaginary contribution to (B1) to lowest order with respect to \( u_\omega \), significant simplification can be achieved. Moreover, for \( \omega >> \omega_\omega \) only the high energy part of the spectrum of the normal excitations contributes to the imaginary part of \( G \) in Eq.(B1). Correspondingly, one can neglect the effect of the condensate on this part of the spectrum (see the case 1) discussed above). The condition when this assumption is valid can be formulated in terms of the smallness of the first term in the square brackets of Eq.(B2) if compared with \( \omega G \). In other words,

\[
\hbar |\omega| >> u_\omega |\Phi|^2 \tag{B4}
\]

in physical units. Employing the variational approach \([15]\), one can estimate \(|\Phi|^2\) and obtain from (B4)

\[
a_0 \sqrt{\omega_\omega} N_c \leq 0.034 \left( \frac{\omega_\omega}{\omega_\omega} \right)^{5/2} \tag{B5}
\]

where we have employed the representation \( u_\omega = 4\pi a \) for the interaction constant \( u_\omega \) in terms of the scattering length \( a \) in the units (6). Actually, for the parameters of the trap \([1]\) the estimate for \( N_c \) gives \( N_c \leq 10^6 \). If this condition holds, the anomalous Green’s function \( G \) can be found by iteration with respect to \( u_\omega \), with the zeroth order approximation being zero. Correspondingly, the normal functions (B3) \( G^{+\pm} \) and \( G^{-\pm} \) should be taken in the zeroth order as \( \Phi_0 \). Assuming that (B4) (or (B5)) is valid, we finally find

\[
G(x,t; x,t') = -iu_\omega \int dt' \int dx' G^{(r)}(x,t; x',t') \Phi^{2}(x',t') \tag{B6}
\]

where the retarded Green’s function \([2]\)

\[
G^{(r)}(x,t; x',t') = \tilde{\theta}(t - t') \sum_1 e^{-i\epsilon_1(t-t')} \psi_1(x)\psi_1^*(x') \tag{B7}
\]

is expressed explicitly in terms of the single-particle eigenfunctions (11), (12) and the eigenvalues (10), with the summation performed over all the single-particle quantum numbers indicated as 1. In (B7), \( \tilde{\theta}(\tau) \) denotes the step function. Finally, substitution of Eqs.(B6) and (B7) into Eq.(B1) yields Eq.(41).

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[1] M.H. Anderson, J.R. Ensher, M.R. Matthews, C.E. Wieman, and E.A. Cornell, Science 269, 198 (1995).
[2] C.C. Bradley, C.A. Sackett, J.J. Tollett, and R.G. Hulet, Phys. Rev. Lett. 75, 1687 (1995).
[3] K.B. Davis, M.-O. Mewes, M.R. Andrews, N.J. van Druten, D.S. Durfee, D.M. Kurn, and W. Ketterle, Phys. Rev. Lett. 75, 3969 (1995).
[4] Bose-Einstein Condensation, ed. by A. Griffin, D.W. Snoke, S. Stringari (Cambrige University Press, Cambridge, 1995).
[5] B.I. Halperin, P.C. Hohenberg, E.D. Siggia, Phys.Rev.B 13, 1299 (1976); K. Damle, S.N. Majumdar and S. Sachdev, cond-mat/9511053.
[6] Yu. Kagan, in \([4]\), p. 202.
[7] H.T.C. Stoof, in \([1]\), p. 226.
[8] S. Stringari, cond-mat/9603126.
[9] B.V. Svistunov and G.V. Shlyapnikov, Sov.Phys.JETP, 71, 71 (1990); H.D. Politzer, Phys.Rev.A 43, 6444 (1991); J. Javanainen, Phys. Rev. Lett., 75, 1927 (1995).
[10] D.R. Tilley and J. Tilley, Superfluidity and Superconductivity ( Adam Hilger, Bristol, 1990).
[11] S. Stringari, Phys. Rev. Lett. 76, 1405 (1996).
[12] P.A. Mulherman and J.C. Inkson, Phys. Rev. B 46, 5454 (1992); F. Dalfovo, A. Fracchetti, A. Lastri, L. Pitaevskii, and S. Stringari, Phys. Rev. Lett. 75, 2510 (1995).
[13] Yu. Kagan, E.L. Surkov, and G.V. Shlyapnikov, atom-ph/9606001.
[14] M. Matthews, D. Jin, J. Ensher, C. Wieman, E. Cornell, Summaries (QPD9-2) of QELS ’96, Anaheim, California, June 2–7, 1996.
[15] W. Petrich, M.H. Anderson, J.R. Ensher, E.A. Cornell, Phys. Rev. Lett. 74, 3352 (1995).
[16] G. Taub, Science 272, 1587 (1996).
[17] V.V. Goldman, I. Silvera, A.J. Leggett, Phys. Rev. B 24, 2870 (1981); V. Bagnato, D.E. Pritchard, and D. Kleppner, Phys. Rev. A 35, 4354 (1987).

[18] G. Baym and C.J. Pethick, Phys. Rev. Lett. 76, 6 (1996).

[19] E.M. Lifshitz and L.P. Pitaevskii, Statistical Physics, Part 2 (Pergamon Press, Oxford, 1980).

[20] S.T. Beliaev, Sov. Phys. JETP 7, 289 (1958).

[21] A.L. Fetter and J.D. Walecka, Quantum Theory of Many-Particle Systems (McGraw-Hill, New York, 1971).

[22] A. Griffin, [cond-mat/9602030].

[23] L.V. Keldysh, Sov. Phys. JETP 20, 1018 (1965); L.V. Keldysh in [4], p. 246.

[24] A. Griffin, Can. J. Phys. 73, 755 (1995).

[25] K.B. Davis, M.-O. Mewes, M.A. Joffe, M.R. Andrews, W. Ketterle, Phys. Rev. Lett. 74, 5202 (1995).