Digital Preservation and Astronomy: Lessons for funders and the funded

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Abstract. Astronomy looks after its data better than most disciplines, and it is no coincidence that the consensus standard for the archival preservation of all types of digital assets – the OAIS Reference Model – emerged originally from the space science community.

It is useful to highlight both what is different about astronomy (and indeed about Big Science in general), what could be improved, and what is exemplary, and in the process I will give a brief introduction to the framework of the OAIS model, and its useful conceptual vocabulary. I will illustrate this with a discussion of the spectrum of big-science data management practices from astronomy, through gravitational wave (GW) data, to particle physics.

1. Introduction

In this paper we will briefly discuss some of the issues surrounding the ‘long-term’ management of data for ‘big science’ projects. It is therefore intended to illuminate the ways in which scientists in these areas have distinctive data management requirements, and a distinctive data culture, which contrasts informatively with other disciplines.

Below, we give a brief account of what we mean by ‘big science’, illustrate some of the features of astronomy, particle-physics and gravitational wave data, present an overview of the OAIS model, and summarise the work we have done in this area.

2. Big data in big science

What is ‘big science’?

Big science projects tend to share many features which distinguish them from the way that experimental science has worked in the past. Such projects share (non-independent) features such as:

- **big money** These are decades-long projects, supported by country-scale funders and billion-currency-unit budgets;
- **big author lists** The (gravitational wave) LSC author list runs at around 0.8 kAuthor, and the LHC’s ATLAS detector is 3 kAuthor;
- **big data rates** Projects produce petabytes of data per year (ATLAS will flatten out at about 10 PB/yr, and Advanced LIGO at around 1 PB/yr);
- **big administration** These include MOUs, councils, and more.
There is an interesting historical discussion of the features of ‘big science’, and LIGO’s progress towards that style of working, in Collins (2003), with an extended history of the sub-discipline in Collins (2004). The study we are reporting on is restricted to the case of such large-scale projects.

What is the ‘long term’?

Astronomy has a long tradition of preserving data, of exploiting old data, and also, in the teeth of centuries of technological change, of finding old data to be readable. The ‘Venus of Ammisaduqa’ cuneiform tablet (a 7th century BCE media refresh of 17th century observations) is intelligible only to antiquaries, but the 12th century Toledan tables are intelligible, with help, to a contemporary astronomer, and Kepler’s 17th century Rudolphine tables are readable without difficulty, suggesting that the intelligibility timescale for astronomy is of order one millennium (note that each of these examples are disseminated publications, not raw data). Particle physics is more mayfly-like, with HEP data becoming unreadable and uninteresting on a timescale of a few decades. Gravitational Wave astronomy is somewhere in the middle, with an experimental practice reminiscent of particle physics, but (eventual) data products which match astronomy.

Data preservation requires a consideration of both media preservation and data intelligibility on timescales potentially ranging from months (in the case of researchers wanting to re-find their own data) to multiple millennia (in the case of very long term nuclear waste disposal (US Department of Energy 2004)). The Open Archival Information System (OAIS) standard, however, pragmatically defines the ‘long-term’ to mean, essentially, ‘long enough for storage technology to change’.

The OAIS standard (CCSDS 650.0) is a high-level model for archives (see Fig. 1). Since the OAIS model ultimately emerged from the space-data community, it is a natural fit to astronomical data. Although the OAIS model has been criticised for being excessively general – it is hard to think of any half-way respectable archive that is not at some level ‘conformant’ with OAIS (cf., Rosenthal et al. 2005) – it is nonetheless very useful as a conceptual toolbox, and as the basis for costings and other deliberations.

Once data sets have been deposited in an archive, and thus reprocessed from Submission Information Packages (SIP) to Archive ones (AIP), they become the exclusive responsibility of the archive, which must plan, and be funded, accordingly. One of the
key questions to be decided by the archive is the nature of the Designated Communities who are expected to retrieve information from the archive, and the description of these communities – who they are, what they know, how hard they can be expected to work in order to retrieve archived data – is a key part of the archive’s documentation. That documentation is written with enough detail that anyone in the designated communities can use it to make the data intelligible, without recourse to any of the data producers, who are presumed to have died, retired, or forgotten all they know, and the creation of this ‘Representation Information’ is a key part of the archive’s initial negotiation with the data producers.

3. Sharing and preserving astronomy data

In science, we preserve data so that we can make it available later. This is on the grounds that scientific data should generally be universally available, partly because it is usually publicly paid for, but also because the public display of corroborating evidence has been part of science ever since the modern notion of science began to emerge in the 17th century (CE) – witness the Royal Society’s motto, which loosely translates as ‘take nobody’s word for it’. Of course, the practice is not quite as simple as the principle, and a host of issues, ranging across the technical, political, social and personal, make this more complicated than it might be, but it is worth noting that the physical sciences generally perform better here than other disciplines, both in the technical maturity of the existing archives and in the community’s willingness to allocate the time and money to see this done effectively.

In 2009–2010 the authors were commissioned by JISC, which is responsible for the exploitation of digital technology in the UK HE system, to examine the way in which the gravitational wave community, as a proxy for big science in general, managed its data, and to make recommendations as appropriate to JISC, the relevant funding councils, and to the GW community. The results of this work will appear at the URL http://purl.org/nxg/projects/mrd-gw in due course. The brief summary of the conclusions is that this community, accurately representing the broader astronomical community, is functioning very well in this respect.

In particular, the commendable features of the community’s approach to data management are:

• There is an expectation of explicit and costed data management planning (though this could arguably benefit from being more systematic).

• There is implicit identification of ‘designated communities’.

• There has always been a recognition that clearly available and described science-data products (AIPs in OAIS-speak) are a vital part of the communication between observers and their colleagues.

• The community is honest about both its need for proprietary periods, and the acknowledgement that these restrictions must only be temporary and short.

All of these features are naturally re-expressible in terms of the OAIS model, which conveniently couples them to the work being done elsewhere on the costing of OAIS models, and on developing tests of OAIS conformance.
Our project has not yet come to specific conclusions about funding. However, it becomes clear that funders are, at some level, as concerned with predictability and auditability as with economy, so that if conventional practice can be demonstrated to be additionally good practice, or within easy reach of it, then this facet of funders’ goals can be marked as achieved.

Finally, we believe that our recommendations will resemble the following.

1. Big-science funders should require projects to develop plans based on the OAIS model, or profiles of it;
2. they should additionally develop or support expertise in criticising the result;
3. and use that modelled plan as a framework for validation of the project’s efforts, both during the lifetime of the project and at its end.

Glossary

**ATLAS**: one of the detectors at the LHC; **LIGO**: the US-based gravitational wave experiment; **LSC**: the multinational LIGO Scientific Collaboration;

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References

CCSDS 650.0 2002, Reference model for an open archival information system (OAIS) – CCSDS 650.0-B-1, CCSDS Recommendation. Identical to ISO 14721:2003, URL [http://public.ccsds.org/publications/archive/650x0b1.pdf](http://public.ccsds.org/publications/archive/650x0b1.pdf)

Collins, H. 2004, Gravity’s shadow: the search for gravitational waves (University of Chicago Press)

Collins, H. M. 2003, Historical Studies in the Physical and Biological Sciences, 33, 261

Rosenthal, D. S. H., Robertson, T., Lipkis, T., Reich, V., & Morabito, S. 2005, D-Lib Magazine, 11. URL [http://www.dlib.org/dlib/november05/rosenthal/11rosenthal.html](http://www.dlib.org/dlib/november05/rosenthal/11rosenthal.html)

US Department of Energy 2004, Permanent Markers Implementation Plan, Tech. Rep. DOE/WIPP 04-3302, United States Department of Energy. URL [http://www.wipp.energy.gov/picsprog/test1/Permanent_Markers_Implementation_Plan_rev1.pdf](http://www.wipp.energy.gov/picsprog/test1/Permanent_Markers_Implementation_Plan_rev1.pdf)