Soliton Formation in Neutral Ion Gases: Exact Analysis

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Abstract

It is shown here that in neutral ion gases the thermal energy transport can occur in the form of new types of thermal soliton waves. The solitons can form under a vanishing net heating function, and for a quadratic net heating. It is predicted that these solitons play an important role in a diversity of terrestrial and astrophysical phenomena. We claim that the reported soliton waves can be observed under ordinary laboratory conditions.

Keywords: Nonlinear Waves, Ion Gases, Thermal Effects, Solitons, Nonlinear PDEs, Exact solutions

1 Introduction

Compressive wave-front propagation is one of the major causes of energy transport in thermally conducting gases, and gives rise to a diverse range of physical phenomena (see [1-3] for further references). It is known that except for the case of subsonic wave motion, all types of associated temperature perturbation cause thermal instability in such systems. This behavior is known as the Field criterion [4], which allows only highly subsonic wave motion in a thermally stable medium. A direct consequence of the Field criterion is that the pressure variations remain small, hence can be neglected in the analysis of compressive thermal waves. The resulting waves formation is highly nonlinear, and depends on the heating function. Linearization and phase space analysis have been used to determine the nature of such waves, which have indicated that apart from travelling wave formations, there are steady wave fronts in such cases.

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In this work we show that in thermally conducting ion gases, the compressive wave propagation can occur in the form of stable, shape preserving wave forms, known as soliton waves. Solitons can be identified as travelling wave solutions that retain their amplitude throughout their propagation, and have constant amplitude in the asymptotic limits. We show that these soliton waves are very general feature of thermally conducting gases, both in astrophysical and under normal laboratory conditions. These play an important role in the energy transport in the media. We state the conditions under which such soliton waves can form in a neutral ion gas, under a linear or a quadratic or a zero net heating function.

2 Basic Equations

The propagation of a wave in an initially uniformly distributed gas, is governed by the basic equations of mass, momentum, and thermal energy conservation. With gas density $\rho$, at pressure $p$, and temperature $T$ these equation are:

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0,$$

$$\rho \frac{D\mathbf{v}}{Dt} + \nabla p = 0,$$

$$\left(\frac{1}{1 - \gamma}\right) \frac{Dp}{Dt} - \left(\frac{\gamma}{1 - \gamma}\right) \frac{p}{\rho} \frac{D\rho}{Dt} = \nabla \cdot (\lambda \nabla T) + Q(\rho, T),$$

where $\mathbf{v}$ is the velocity of the wave in the gas, and $Q(\rho, T)$ is the net heat gain per unit volume per unit time. It expresses the difference between the heating and cooling rates of the medium. Here $\lambda$ is the thermal conductivity coefficient, given by the ratio of the of specific heats $\gamma = c_p/c_v$.

For a gas at fixed volume and mass, the net heat gain is a function of temperature only. It therefore follows from equation (1) to (3) that instability occurs for the gaseous medium when $\left(\frac{\partial Q}{\partial \ln T}\right)_p > 0$ (Field 1965). This condition is the Field stability criterion. In one spatial dimension the Field criterion implies that wave motion of heating or cooling interfaces must be highly subsonic. This means that we can neglect the velocity term in the momentum equation, hence obtain from equation (2):

$$\frac{\partial p}{\partial x} = 0.$$  \hspace{1cm} (4)

implying that pressure is constant within the medium. A change to Lagrange variable $\eta(x, t)$ defined by $\eta(x, t) = \int_{-\infty}^{x} \rho(x', t) dx'$, gives for the mass conservation equation (1):

$$\frac{\partial \eta}{\partial t} \bigg|_{x} = -\rho \mathbf{v},$$  \hspace{1cm} (5)
where \( v \) is the velocity in \( x \)-direction. This gives \( \partial/\partial x |_t = \rho \partial/\partial \eta |_t \) and \( \partial/\partial t |_x = \partial/\partial \eta | -\rho u \partial/\partial \eta |_t \). The Lagrange derivative \( D/Dt \) can now be expressed as:

\[
\frac{\partial f}{\partial t} |_x + v \frac{\partial f}{\partial x} |_t = \frac{\partial f}{\partial t} |_\eta .
\] (6)

Using these results in the energy equation (3) we obtain,

\[
\frac{1}{\gamma - 1} \frac{Dp}{Dt} - \frac{\gamma}{\gamma - 1} \rho \frac{D\rho}{Dt} = \frac{\partial}{\partial x} \left( \lambda \frac{\partial T}{\partial x} \right) + Q,
\] (7)

which becomes

\[
- \frac{\gamma}{\gamma - 1} \rho \frac{D\rho}{Dt} |_\eta = \rho \frac{\partial}{\partial \eta} \left( \lambda \frac{\partial T}{\partial \eta} \right) + Q.
\] (8)

In the subsonic case the equation of state is \( p = (\mathcal{R}/\mu) \rho T \), it therefore follows that \( \rho \) is proportional to \( 1/T \). Thus both \( \lambda \) and \( Q \) are functions of temperature only. Let now the new time variable \( \tau = \left( (\gamma - 1)p/\gamma (\mathcal{R}/\mu)^2 \right) t \), and denote \( \mathcal{L}(T) = Q(T) \left( \frac{\mathcal{R}/\mu}{p} \right)^2 / T \). Then equation (10) gives the governing equation for the evolution of the temperature front in a thermally conducting gas:

\[
\frac{\partial T}{\partial \tau} = \frac{\partial}{\partial \eta} \left( \frac{\lambda(T)}{T} \frac{\partial T}{\partial \eta} \right) + \mathcal{L}(T).
\] (9)

### 3 Soliton Wave Analysis

The conductivity function \( \lambda(T) \) for a neutral ion gas is proportional to \( \sqrt{T} \). In the case of steadily moving wave front in \( \eta \) space, with velocity \( V \), we set \( \xi = c(\eta - K \tau) \). This gives in equation (12):

\[
c^2 u^{-1} \frac{d^2 u}{d\xi^2} + cK \frac{du}{d\xi} + F(u) = 0,
\] (10)

where \( u = \sqrt{T} \), and \( F(u) = 4T^{-1/2} \mathcal{L}(T) \) is the net heating function. We now consider equation (13) for the case of linear, and quadratic net heating functions.

If the net heating function is linear, we have from equation (10) the nonlinear wave equation:

\[
c^2 \frac{1}{u} \frac{d^2 u}{d\xi^2} + cK \frac{du}{d\xi} + Au + B = 0.
\] (11)

We seek the solution to this equation such that \( u'(\xi) \) is a function of the separable form,

\[
u'(\xi) = a^2 - u(\xi)^2.
\] (12)

This implies that

\[
u''(\xi) = -2a^2 u(\xi) + 2u(\xi)^3.
\] (13)
Therefore equation (11) becomes after substitution from equations (12) and (13), and collecting terms of the like powers in \( u \):

\[
(2c^2 - cK)u^3 + Au^2 + (cKa^2 + B - 2a^2c^2)u = 0.
\]  

(14)

Equating the coefficients of \( u^2 \), \( u^1 \), and \( u^0 \) we get

\[
K = 2c, \quad A = 0, \quad B = 0.
\]  

(15)

This case correspond to the vanishing net heating function, and from equations (12) and (15) the solution can be written as

\[
u(\xi) = a \tanh a(\xi + \xi_0),
\]  

(16)

where \( \xi_0 \) corresponds to the initial value of the travelling wave variable. Equation (16) gives for the temperature

\[
T(\eta, \tau) = a^2 \tanh \left( a(\eta - 2c\tau + \xi_0/c)\right).
\]  

(17)

Equation (17) implies that as \( \xi \rightarrow \pm \infty \), \( T \rightarrow T_{\pm \infty} = a^2 \). The asymptotic limit of temperature must correspond to the equilibrium state. We re-scale the temperature as \( \hat{T}(\eta, \tau) = T_{\pm \infty} - T(\eta, \tau) \), so that equation (17) becomes

\[
\hat{T}(\eta, \tau) = T_{\pm \infty} \left( \text{sech}(a(\eta - 2c\tau + \xi_0/c)) \right)^2.
\]  

(17a)

This represents a free soliton wave solution when \( F(u) = 0 \). Physically this corresponds to the case when heating of the gas occurs at the same rate as cooling. The soliton amplitude in this case is \( a^2 \) and its speed is \( 2ac^2 \). As \( \xi \rightarrow \pm \infty \), the soliton attains a fixed amplitude \( a^2 \), whereas it retains its profile during propagation. Figure (1) shows plot of the soliton (17a) as function of variables \( \eta \) and \( \tau \).

For a quadratic heating function \( F(u) = Au^2 + Bu + C \), we have by a similar procedure:

\[
(A + 2c^2 - cK)u^3 + Bu^2 + (cKa^2 + C - 2a^2c^2)u = 0.
\]  

(18)

Equating coefficients of the various powers of \( u \), gives the determining equations,

\[
A + 2c^2 - cK = 0, \quad cKa^2 + C - 2a^2c^2, \quad B = 0,
\]  

(19)

we get

\[
c = \frac{K}{4} \pm \frac{1}{4} \sqrt{K^2 - 8A}, \quad a = \pm \sqrt{C/A}, \quad B = 0,
\]  

(20)

Thus corresponding to the net heating function \( Au^2 + C \) we have the soliton solution

\[
u(\xi) = \pm \sqrt{C/A} \tanh \pm \sqrt{C/A}[c(\eta - K\tau) + \xi_0],
\]  

(21)

or in terms of the temperature variable

\[
T(\eta, \tau) = \frac{C}{A} \tanh^2 \pm \sqrt{\frac{C}{16A}}[(K + \sqrt{K^2 - 8A})(\eta - K\tau) + 4\xi_0].
\]  

(22)
Re-scaling the temperature as \( \tilde{T}(\eta, \tau) = T_{\pm \infty} - T(\eta, \tau) \), where \( T_{\pm \infty} = C/A \), we obtain

\[
T(\eta, \tau) = T_{\pm \infty} \tanh \left( \sqrt{\frac{C}{16A}} \left[ (K + \sqrt{K^2 - 8A})(\eta - K\tau) + 4\xi_0 \right] \right). \tag{22a}
\]

Equation (22a) shows that the soliton solution for a purely quadratic function has a strong amplitude dependence on the heating function parameters. In this way it represents a totally distinct soliton solution than that for the previous case. The thermal soliton (22a) has amplitude \( C/A \), which increases as the parameter \( A \) decreases. Also the wave speed is proportional to \( \sqrt{C/16A} \), which shows that the soliton wave moves faster if the quadratic term in the net heating function is adjusted to smaller values as compare to the constant terms in the heating function. Another interesting feature of the soliton solution (22a) is that the argument of the \( \tanh \) function involves the square root \( \sqrt{K^2 - 8A} \), which can become imaginary when \( K^2 < 8A \), thus exhibit drastically different behavior depending on the threshold value \( K = \pm \sqrt{8A} \).

4 Conclusions

In this communication it is reported that thermally conducting gases in neutral ionized states contain thermal solitons that participate in the general process of energy transfer in the medium. For a given linear or quadratic net heating function these soliton waves can form as single solitary waves or, depending on the parameters of the heating function, can appear as multiple soliton waves. In the quadratic case, the net heating function parameters not only determine their amplitude but their speed in the gas as well. The existence of free solitons shows that there can be energy transfer via solitons even when the cooling of the gas occurs at the same rate as its heating. In this case the amplitude as well as the speed of the soliton wave depends on the equilibrium temperature of the gas. Since temperature is inversely proportional to the density of the gas, the soliton effects can be observed in density profile of a neutral ion gas, under appropriate conditions.

Neutral ion gases are most prevalent in astrophysical and atmospheric conditions, such as the solar corona, interstellar medium, and galactic clusters, ionosphere, upper atmosphere of the Earth and planets.

However the thermal solitons can also be observed under ordinary laboratory conditions. It is predicted here that for the vanishing and quadratic net heating functions, cases which are easier to be produced in a lab with adjustable heating parameters for neutral ion gases, the existence of these thermal solitons can be demonstrated. As far as the author knows this experimental proof yet remains to be given.
References

[1] Meerson, B., 1996. Rev. Mod. Phys. 68, 215.

[2] Chin, R., Verwichte, E., Rowlands, G., and Nakariakov, V.M., 2010. Phys. Plasmas. 17, 032107.

[3] Iwasaki, K., and Inutsuka, S., 2012. MNRAS. 4, 3638.

[4] Field, G.B., 1965. ApJ. 142, 531.

Figure Caption:
Figure 1: A free soliton wave based on equation (17a), formed at equilibrium temperature $T_{\pm\infty} = 0.1$, $\xi_0 = 0$, and $c = 0.2$. 
This figure "fig1.JPG" is available in "JPG" format from:

http://arxiv.org/ps/1209.3077v3