THE W MASS AND THE U PARAMETER

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The Z-pole data from $e^+e^-$ colliders and results from the NuTeV experiment at Fermilab can be brought into agreement if (1) the neutrino-Z couplings were suppressed relative to the Standard Model (SM), and (2) the Higgs boson were much heavier than suggested by SM global fits. However, increasing the Higgs boson mass will move the theoretical value of the W mass away from its experimental value. A large and positive $U$ parameter becomes necessary to account for the difference. We discuss what type of new physics may lead to such values of $U$.

1. The NuTeV Anomaly and Neutrino Mixing

The NuTeV experiment at Fermilab has measured the ratios of neutral to charged current events in muon (anti)neutrino–nucleon scattering:

$$R_\nu = \frac{\sigma(\nu_\mu N \rightarrow \nu_\mu X)}{\sigma(\nu_\mu N \rightarrow \mu^- X)} = g_L^2 + r g_R^2,$$

$$R_\bar{\nu} = \frac{\sigma(\bar{\nu}_\mu N \rightarrow \bar{\nu}_\mu X)}{\sigma(\bar{\nu}_\mu N \rightarrow \mu^+ X)} = g_L^2 + \frac{g_R^2}{r},$$

where

$$r = \frac{\sigma(\bar{\nu}_\mu N \rightarrow \mu^+ X)}{\sigma(\nu_\mu N \rightarrow \mu^- X)} \sim \frac{1}{2},$$

and has determined the parameters $g_L^2$ and $g_R^2$ to be

$$g_L^2 = 0.30005 \pm 0.00137,$$

$$g_R^2 = 0.03076 \pm 0.00110.$$
The Standard Model (SM) predictions of these parameters based on a global fit to non-NuTeV data, cited as $\left[ g^2_L\right]_{\text{SM}} = 0.3042$ and $\left[ g^2_R\right]_{\text{SM}} = 0.0301$ in Ref. [2], differ from the NuTeV result by 3$\sigma$ in $g^2_L$. This disagreement between NuTeV and the SM (as determined by non-NuTeV data) is sometimes referred to as the NuTeV “anomaly”.

Note that the NuTeV value of $g^2_L$ is smaller than the SM prediction. In terms of the ratios $R_\nu$ and $R_{\bar{\nu}}$, this means that the neutral current events were not as numerous as predicted by the SM when compared to the charged current events. In addition, the $Z$ invisible width measured at $e^+e^-$ colliders,

$$\Gamma_{\text{inv}}/\Gamma_{\text{lept}} = 5.942 \pm 0.016,$$

is 2$\sigma$ below the SM prediction of $[\Gamma_{\text{inv}}/\Gamma_{\text{lept}}]_{\text{SM}} = 5.9736 \pm 0.0036$. Both these observations seem to suggest that the neutrino couplings to the $Z$ boson are suppressed.

Suppression of the $Z\nu\nu$ couplings occurs most naturally in models which mix the neutrinos with heavy gauge singlet states. For instance, if the $SU(2)_L$ active neutrino $\nu_L$ is a linear combination of two mass eigenstates with mixing angle $\theta$,

$$\nu_L = (\cos \theta)\nu_{\text{light}} + (\sin \theta)\nu_{\text{heavy}},$$

then the $Z\nu\nu$ coupling will be suppressed by a factor of $\cos^2 \theta$ provided the heavy state is too massive to be created in the interaction. The $W\ell\nu$ coupling will also be suppressed by a factor of $\cos \theta$.

In general, if the $Z\nu\nu$ coupling of a particular neutrino flavor is suppressed by a factor of $(1-\varepsilon)$, then the $W\ell\nu$ coupling of the same flavor will be suppressed by a factor of $(1-\varepsilon/2)$. For the sake of simplicity, assume that the suppression parameter $\varepsilon$ is common to all three generations. The theoretical values of $R_\nu$ and $R_{\bar{\nu}}$ are then reduced by a factor of $(1-\varepsilon)$, since their numerators are suppressed over their denominators, while the invisible width of the $Z$ is reduced by a factor of $(1-2\varepsilon)$. Thus, neutrino mixing could, in principle, provide an explanation for both the NuTeV and invisible width discrepancies.

At this point, we recall that one of the inputs used to calculate SM predictions is the Fermi constant $G_F$, which is extracted from the muon decay constant $G_\mu$. The suppression of the $W\ell\nu$ couplings leads to the correction

$$G_F = G_\mu(1 + \varepsilon),$$

(6)
Table 1. The observables used in this analysis. The SM predictions are for inputs of $M_{\text{top}} = 174.3$ GeV, $M_{\text{Higgs}} = 115$ GeV, $\alpha_s(M_Z) = 0.119$, and $\Delta\alpha_{\text{had}}^{(5)} = 0.02761$.

| Observable | SM prediction | Measured Value |
|------------|---------------|----------------|
| $\Gamma_{\text{lept}}$ | $83.998$ MeV | $83.984 \pm 0.086$ MeV |
| $\Gamma_{\text{inv}}/\Gamma_{\text{lept}}$ | $5.973$ | $5.942 \pm 0.016$ |
| $\sin^2\theta_{\text{eff}}^{\text{lept}}$ | $0.23147$ | $0.23148 \pm 0.00017$ |
| $g_L^2$ | $0.3037$ | $0.3002 \pm 0.0012$ |
| $g_R^2$ | $0.0304$ | $0.0310 \pm 0.0010$ |
| $M_W$ | $80.375$ | $80.449 \pm 0.034$ GeV |

which would affect all SM predictions. One might worry that a shift in $G_F$ would destroy the excellent agreement between the SM and the majority of the Z-pole data. However, since the Fermi constant always appears multiplied by the $\rho$-parameter in neutral current amplitudes, the shift can be compensated by the introduction of the $T$ parameter, leaving $\rho G_F$ unaffected. The suppression of $Z\nu\nu$ couplings together with oblique corrections from new physics could thus reconcile the NuTeV result with the Z-pole data.

2. The Fits

To test this idea, we fit the Z-pole, NuTeV, and W mass data with the oblique correction parameters $S$, $T$, $U$ and the $Z\nu\nu$ coupling suppression parameter $\varepsilon$. Table 1 comprises the six observables used in our fit and their SM predictions. The details of the analysis are presented in Ref. [6]; here, we merely outline our results.

That oblique corrections alone cannot account for the NuTeV anomaly is established with a fit using only the oblique correction parameters $S$, $T$, and $U$. Taking $M_{\text{top}} = 174.3$ GeV, $M_{\text{Higgs}} = 115$ GeV as the reference SM, we obtain

$$S = -0.09 \pm 0.10 \,,
T = -0.13 \pm 0.12 \,,
U = 0.32 \pm 0.13 \,.$$  \hfill (7)

The quality of the fit is unimpressive: $\chi^2 = 11.3$ for $6 - 3 = 3$ degrees of freedom. The preferred region on the $S$-$T$ plane is shown in Fig. 1a. As is evident from the figure, there is no region where the $1\sigma$ bands for $\Gamma_{\text{lept}}$, $\sin^2\theta_{\text{eff}}^{\text{lept}}$, and $g_L^2$ overlap.

Next, we fit using $S$, $T$, $U$, and $\varepsilon$. The reference SM is $M_{\text{top}} =$
Figure 1. The fit to the data with only $S$ and $T$ (a), and with $S$, $T$, and $\varepsilon$ (b),(c),(d). The bands associated with each observable show the 1σ limits. (The $M_W$ band is not shown.) The shaded ellipses show the 68% and 90% confidence contours. The unshaded ellipses partially hidden behind the shaded ones in (a) show the contours when only the $Z$-pole data is used. The origin is the reference SM with $M_{\text{top}} = 174.3\text{ GeV}$ and $M_{\text{Higgs}} = 115\text{ GeV}$. The curved arrow attached to the origin indicates the path along which the SM point will move when the Higgs mass is increased from 115 GeV to 1 TeV.

174.3 GeV, $M_{\text{Higgs}} = 115\text{ GeV}$ as before. The result is

$$ S = -0.03 \pm 0.10 , $$
$$ T = -0.44 \pm 0.15 , $$
$$ U = 0.62 \pm 0.16 , $$
$$ \varepsilon = 0.0030 \pm 0.0010 . $$ \hfill (8)

The quality of the fit is improved dramatically to $\chi^2 = 1.17$ for $6 - 4 = 2$ degrees of freedom. The preferred regions in the $S$-$T$, $S$-$\varepsilon$, and $T$-$\varepsilon$ planes are shown in Figs. 1b through 1d. As anticipated, inclusion of both oblique corrections and $\varepsilon$ results in an excellent fit to both the $Z$-pole and NuTeV data.
3. Heavy Higgs and the W Mass

What type of new physics would provide the values of the oblique correction parameters and $\varepsilon$ preferred by the fit? The value of $\varepsilon$ implies a largish mixing angle between the light active and heavy sterile states. In Ref. [8], we discuss how such mixings can be realized within the seesaw framework. For the oblique corrections, the limits on $S$ permit it to have either sign, while $T$ is constrained to be negative by 3$\sigma$. Few models of new physics are available which predict a negative $T$.

A heavy SM Higgs provides a simple starting point. Recall that the effect of a SM Higgs heavier than our reference value (here chosen to be 115 GeV) is manifested as shifts in the oblique correction parameters. The approximate expressions for these shifts are

\begin{align}
S_{\text{Higgs}} & \approx \frac{1}{6\pi} \ln \left( \frac{M_{\text{Higgs}}}{M_{\text{ref Higgs}}} \right), \\
T_{\text{Higgs}} & \approx -\frac{3}{8\pi c^2} \ln \left( \frac{M_{\text{Higgs}}}{M_{\text{ref Higgs}}} \right), \\
U_{\text{Higgs}} & \approx 0.
\end{align}

Thus increasing the Higgs mass generates a negative $T$. Indeed, we have shown in Ref. [6] that the $Z$-pole and NuTeV observables can be fit by $\varepsilon$ alone if the Higgs boson is as heavy as a few hundred GeV. Since the Higgs boson has not been found in the $\sim 80$ GeV range preferred by the SM global fit, the prospect that the data prefer a heavier Higgs is actually welcome.

However, as shown in Fig. 2a, a heavier Higgs will lower the SM prediction of the $W$ mass, shifting it away from the experimental value. We would like to point out that although the experimental value of the $W$ mass differs from the SM global fit (with $M_{\text{Higgs}} \sim 80$ GeV) by only 1.7$\sigma$, if the

![Figure 2](image)
Higgs mass is raised to its lower bound of 115 GeV, the difference is $2.2\sigma$. As the experimental error on the $W$ mass decreases and the lower bound on the Higgs mass increases, the $W$ mass may become the next ‘anomaly’ to be confronted.

Regardless the actual mass of the Higgs, our fits indicate that the presence of $Z\nu\nu$ suppression demands a large and positive $U$ parameter to account for the $W$ mass. (See Fig. 2b.) What new physics predicts a small $T$ and a large $U$? One possibility is that the $U$ parameter is enhanced by the formation of bound states at new particle thresholds. Expressing $T$ and $U$ as dispersion integrals over spectral functions gives

$$T \propto \int_{s_{\text{thres}}}^{\infty} \frac{ds}{s} [\text{Im}\Pi_\pm(s) - \text{Im}\Pi_0(s)],$$

$$U \propto \int_{s_{\text{thres}}}^{\infty} \frac{ds}{s^2} [\text{Im}\Pi_\pm(s) - \text{Im}\Pi_0(s)],$$

(10)

using the notation of Ref. [12]. Because of the extra negative power of $s$ in its integrand, $U$ is more sensitive to the threshold enhancement than $T$. Indeed, it has been shown in Ref. [12] that threshold effects do not enhance the $T$ parameter. This could be an indication that technicolor theories are the most promising candidates. Thus, technicolor theories which were killed by the $S$ parameter could be resurrected by the $U$ parameter.

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References

1. The LEP Collaborations, the LEP Electroweak Working Group, and the SLD Heavy Flavor and Electroweak Groups, hep-ex/0212036.
2. [NuTeV Collaboration] G. P. Zeller et al., Phys. Rev. Lett. 88, 091802 (2002) [hep-ex/0110059]; Phys. Rev. D 65, 111103 (2002) [hep-ex/0203004]; K. S. McFarland et al., hep-ex/0205080; G. P. Zeller et al., hep-ex/0207052.
3. C. H. Llewellyn Smith, Nucl. Phys. B 228, 205 (1983).
4. S. Davidson, S. Forte, P. Gambino, N. Rius and A. Strumia, JHEP 0202, 037 (2002) [hep-ph/0112302].
5. M. Gronau, C. N. Leung and J. L. Rosner, Phys. Rev. D 29, 2539 (1984); J. Bernabeu, A. Santamaria, J. Vidal, A. Mendez and J. W. Valle, Phys. Lett. B 187, 303 (1987); K. S. Babu, J. C. Pati and X. Zhang, Phys. Rev. D 46, 2190 (1992); L. N. Chang, D. Ng and J. N. Ng, Phys. Rev. D 50, 4589 (1994) [hep-ph/9402259]; W. J. Marciano, Phys. Rev. D 60, 093006 (1999) [hep-ph/9903451]; A. De Gouvea, G. F. Giudice, A. Strumia and K. Tobe, Nucl. Phys. B 623, 395 (2002) [hep-ph/0107156]; K. S. Babu and J. C. Pati, hep-ph/0203029.

6. W. Loinaz, N. Okamura, T. Takeuchi and L. C. R. Wijewardhana, hep-ph/0210193 (to appear in Phys. Rev. D); T. Takeuchi, hep-ph/0209109.

7. M. E. Peskin and T. Takeuchi, Phys. Rev. Lett. 65, 964 (1990); Phys. Rev. D 46, 381 (1992); J. L. Hewett, T. Takeuchi and S. Thomas, hep-ph/9603391.

8. W. Loinaz, N. Okamura, S. Rayyan, T. Takeuchi, and L. C. R. Wijewardhana, hep-ph/0304004.

9. S. Bertolini and A. Sirlin, Phys. Lett. B 257, 179 (1991); E. Gates and J. Terning, Phys. Rev. Lett. 67, 1840 (1991); B. Holdom, Phys. Rev. D 54, 721 (1996) [hep-ph/9602248]; B. Holdom and T. Torma, Phys. Rev. D 59, 075005 (1999) [hep-ph/9807561].

10. M. E. Peskin and J. D. Wells, Phys. Rev. D 64, 093003 (2001) [hep-ph/0101342].

11. M. S. Chanowitz, Phys. Rev. Lett. 87, 231802 (2001) [hep-ph/0104024]; Phys. Rev. D 66, 073002 (2002) [hep-ph/0207123].

12. T. Takeuchi, A. K. Grant and M. P. Worah, Phys. Rev. D 51, 6457 (1995) [hep-ph/9403294]; Phys. Rev. D 53, 1548 (1996) [hep-ph/9407349].