The $\beta$-decay of $^{22}$Al

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Abstract. In an experiment performed at the LISE3 facility of GANIL, we studied the decay of $^{22}$Al produced by the fragmentation of a $^{36}$Ar primary beam. A $\beta$-decay half-life of $T_{1/2} = 91.1 \pm 0.5$ ms was measured. The $\beta$-delayed one- and two-proton emission as well as $\beta$-α and $\beta$-delayed γ decays were measured and allowed us to establish a partial decay scheme for this nucleus. New levels were determined in the daughter nucleus $^{24}$Mg. The comparison with model calculations strongly favours a spin-parity of $I^\pi = 4^+$ for the ground state of $^{22}$Al.

PACS. 21.10.-k Properties of nuclei; nuclear energy levels – 23.90.+w Other topics in radioactive decay and in-beam spectroscopy – 23.50.+z Decay by proton emission

1 Introduction

Our understanding of the structure of the atomic nucleus is mainly based on studies of stable nuclei or nuclei close to stability. However, it is now well established that the nuclear structure changes with the addition of the isospin degree of freedom, i.e. when studying nuclei far away from the valley of stability. The study of these exotic nuclei is today a well established tool for a deeper understanding of nuclear structure.

The exotic odd-odd nucleus $^{22}$Al ($Z=13$, $N=9$) is situated at the proton drip line of the nuclear chart and is evaluated to be bound by only 20 keV [1]. It was observed for the first time in 1982 by Cable et al. [2]. The spin of the ground state of $^{22}$Al is not known. In the mirror nucleus $^{24}$F, the spin of the ground state is $I^\pi = 4^+$, but this level is very close to a $3^+$ excited state ($E^* = 71.6$ keV) and these states could be inverted in $^{22}$Al. As the $\beta$-decay energy of $^{22}$Al is high ($Q_{\beta^+} = 17.56$ MeV) [1], different decay channels are open: $\beta$-γ, $\beta$-p (the one-proton separation energy $S_p$ in $^{22}$Mg is equal to 5501.6 keV), $\beta$-2p ($S_{2p} = 7931$ keV) and $\beta$-α ($S_{\alpha} = 8138.7$ keV) decays are energetically possible.

Cable et al. [2] were the first to study the $\beta$ decay of this nucleus. In their experiment, $^{22}$Al was produced by a $^{24}$Mg($^3$He, $p\alpha n$)$^{22}$Al reaction at 110 MeV. The aluminum atoms were transported using a helium-jet technique. In the decay-energy spectrum, only two peaks at high energy were clearly seen, the low-energy part being contaminated by the decay of other nuclei. The measured energies correspond to 8212 ± 16 keV and 8537 ± 22 keV in the center of mass frame. Using energy and intensity considerations, these peaks were assigned to the isospin-forbidden proton decay from the lowest $T=2$ state in $^{22}$Mg - the isobaric analogue state (IAS) of the $^{22}$Al ground state - to the ground and the first excited states in $^{21}$Na. This IAS is measured to be at an excitation energy of 13650 ± 15 keV and is fed by a superallowed $\beta^+$ decay. Based on the $T=2$ isospin multiplet, a $I^\pi = 4^+$ was assigned to the $^{22}$Al ground state. A rough half-life value of $70^{+50}_{-35}$ ms was also determined. Due to the technique used, no absolute branching ratios could be determined. In another study of this nucleus, the first observation of the exotic $\beta$-2p decay was published by Cable et al. [3]. Two $\beta$-delayed 2p decay branches were observed. From energy considerations, these two branches were attributed to transitions from...
the IAS in $^{22}Mg$ to the first excited state and the ground state of $^{20}Ne$. In a recent experiment performed by Blank et al. [9], the $\beta$-$\alpha$ decay of this nucleus was observed for the first time using a technique based on the implantation of $^{22}Al$ into a silicon detector and into a micro-strip gas counter. A half-life of $T_{1/2} = 59 \pm 3$ ms was measured. A more detailed experimental decay scheme was obtained with absolute branching ratios for $\beta$-$\alpha$, $\beta - p$ and $\beta - 2p$ emission. However, this experiment suffered from contamination. The $^{22}Al$ nuclei were produced by fragmentation of a $^{36}Ar$ primary beam and only about 30% of the nuclei detected were $^{22}Al$ nuclei. Shell model calculations were also reported in this publication. From several arguments, the hypothesis that the ground state of $^{22}Al$ has a spin-parity of $3^+$ was favoured. Finally, a detailed 1s-0d shell-model calculation was carried out by Brown [8] for the isospin-forbidden $\beta$-delayed proton emission from the IAS in $^{22}Mg$. A first complete set of calculations for this particle decay was presented.

The $\beta$-decay of this nucleus is quite well known but still suffers from uncertainties and limitations. Due to strong contaminations in the previous measurements, we can expect that several transitions are still missing and probably hidden by much stronger transitions from contaminants. In addition, the $\gamma$ rays were not measured in any of these experiments. The aim of our experiment was to improve the $\beta$-decay measurement by means of different advantages: a better purity of the $^{22}Al$ secondary beam, higher statistics, greatly improved energy resolution in the measurement of the charged particles emitted, a $\gamma$-ray measurement and $\gamma$-particle coincidences to give a more certain assignment for many transitions.

A comparison of the experimental results with detailed shell model calculations is also proposed in the present paper. For the first time, a complete set of calculations for all particle decay channels from relevant levels in $^{22}Mg$ is done. This will allow us to investigate the Gamow-Teller quenching in $\beta$-decay, to test the isobaric multiplet mass equation (IMME), and to probe the isospin symmetry in $\beta$-transitions of mirror decays. Another challenge is to determine the spin of the $^{22}Al$ ground state.

2 The experimental setup

The study of the $^{22}Al$ $\beta$ decay was performed at GANIL with the LISE3 zero-degree achromatic recoil spectrometer [7,8]. It was part of a systematic $\beta$-decay study of neutron deficient nuclei: $^{22}S$, $^{25}Si$ and $^{26}P$ [10] where the same experimental setup was used. The nucleus $^{22}Al$ was produced by the fragmentation of a 95 A MeV $^{30}Ar$ primary beam on a 528 ng/cm$^2$ carbon target located in the SISSI device. The $^{22}Al$ secondary beam had an energy of 48 A MeV. Using a beryllium degrader at the intermediate LISE focal plane and the LISE3 Wien filter, this isotope was produced with an intensity of 40 pps and a purity of 93%. The main contaminants were $^{21}Mg$ (1.5%), $^{20}Na$ (3.5%) and $^{19}Ne$ (2%). A total number of $2.5 \times 10^6$ $^{22}Al$ was recorded in this experiment. Other settings of the spectrometer were meant to select and measure the $^{21}Mg$ and $^{24}Al$ $\beta$ decays with the same detection system for energy and efficiency calibration.

The experimental setup shown in figure [11] consisted of two sets of detectors. The identification of the incident ions was performed with two 300 $\mu$m silicon detectors (E1 and E2xy), generating an energy-loss signal and a multi-channel plate (MCP) detector situated at the first LISE focal plane which, together with the E1 detector, produced a time-of-flight signal. The $\beta$-decay detection system was composed of two 500 $\mu$m silicon detectors (E3, E4), a 6 mm Si(Li) detector (E5) and an EXOGAM Germanium clover detector in close geometry. The fragments were implanted at the downstream edge of the fourth silicon detector E4, which served also to measure the charged-particle decay-energy spectrum. The energy calibration of the E4 detector was performed using the known $\beta$-proton transitions of $^{21}Mg$ [11]. The E3 and E5 detectors were used as $\beta$-particle detectors.

3 Monte Carlo simulation

A Monte Carlo simulation of the experiment was performed using the GEANT code [12]. This simulation has produced four main conclusions about:

- The proton detection efficiency:
  High-energy protons can escape from the E4 implantation detector. The simulation allowed us to extract the proton detection efficiency of the E4 detector as a function of proton energy assuming an isotropic proton emission distribution. For energies from 2 MeV to 8 MeV, the efficiency decreases from 100% to 52% for an implantation depth of 420 ± 10 $\mu$m. The relative uncertainty of the proton detection efficiency depends on the uncertainty of the implantation depth in the silicon detector and on the uncertainty of the detector thickness. After a simulation using the extreme values of the implantation profile, the relative uncertainty of the proton detection efficiency was found to be less than 6%.

- The proton energy resolution:
  This simulation and previous experiments have shown that a condition on the last silicon detector (E5 >0)
leads to a better energy resolution in the decay-energy spectrum. In figure 2, we have simulated the $\beta$ decay of $^{21}\text{Mg}$ to one state at 4.468 MeV in $^{21}\text{Na}$ followed by the emission of a 2.036 MeV proton. In the E4 spectrum, we can see that the $\beta$ particle associated with the transition has a strong effect. The proton peak is shifted by 36 keV when the positrons escape on the downstream side of E4 and by 158 keV when they escape in the other direction. The dotted line represents the proton spectrum conditioned by a $\beta$ detection in E5 ($E_5 > 0$). The dashed line is the Gaussian fit of this conditioned spectrum peak.

So, to extract the peak area we have used the method described here. First of all, we performed a Gaussian fit of the low energy part of the simulated proton spectrum, conditioned by the E5 detector ($E_5 > 0$) as shown by the dashed line in fig. 2. This fit allowed us to calculate the corresponding area. By normalising this area with the total peak area without any condition (the total number of simulated particles for this proton energy), we obtain the ratio of events selected by such a procedure. We have introduced two ratios:

1. $R_1 = \frac{N_{p(E_5 > 0)}}{N_{p{\text{total}}}}$

This ratio is determined by dividing the total number of counts in the conditioned proton spectrum ($E_5 > 0$) by the total number of counts in the proton spectrum without any condition. However, the charged-particle spectrum contains a $\beta$ contribution from $^{22}\text{Al} \beta\gamma$ transitions. To get the total number of protons, one has to subtract this contribution from the spectrum. In fact, simulations have shown that the $\beta$ contribution can be fitted by the sum of two exponential functions and subtracted from the unconditioned proton spectrum.

2. $R_2 = \frac{N_{p(E_5 > 0)Gauss}}{N_{p(E_5 > 0)}}$

$R_2$ is the ratio between the Gaussian area of the proton peak and the total proton peak area in the conditioned spectrum. This takes into account that the Gaussian does not perfectly describe the peak in the conditioned spectrum.

In practice, the low-energy part of the proton peak in the conditioned spectrum ($E_5 > 0$) is fitted by a Gaussian. The area obtained is divided by $R_2$. This yields the total area of the proton peak in the conditioned spectrum. This value is divided by $R_1$ to obtain the area of the total proton peak (without condition). Finally, the branching ratio is determined by dividing this area by the total number of $^{22}\text{Al}$ implanted.

The simulations showed that the ratios $R_1$ and $R_2$ depend weakly on the $\beta$-particle energy (a maximum of 10% relative error). However, since they depend on the geometry, we need to determine these ratios from experimental data at least for one proton peak (for one $\beta$-particle energy) and use these ratios for the determination of the other proton peak areas.

- The $\beta$-trigger efficiency:
In the present experiment, the acquisition system was triggered by charged particles detected in E3, E4 or E5. If the $\beta$ particle is followed by proton emission, the energy loss is always sufficient to trigger the acquisition. Otherwise, in the case of $\beta\gamma$ transitions, the energy lost by $\beta$ particles in E3 or E4 is not sufficient to trigger the acquisition. In this case, the acquisition may be triggered by $\beta$ particles detected in the E5 detector. This efficiency must be determined to calculate the branching ratio for the $\beta\gamma$ transitions.

To study the behaviour of the $\beta$-trigger efficiency, a simulation was performed. It has shown that this effi-
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4 Experimental results

In the following, we present our experimental results: half-life, β-delayed γ decays, β-delayed charged-particle emissions as well as coincidences between charged particles and γ rays and γ-γ coincidences. These results allow us to determine a detailed experimental decay scheme and to compare it with previous results.

1. $^{22}\text{Al}$ half-life

The $^{22}\text{Al}$ half-life was measured in beam-on/beam-off mode (120 ms of implantation, 300 ms beam-off). The inset in figure 3 shows the decay-time spectrum in the beam-off period fitted by an exponential function added to a constant for the background. In the main figure, the half-lifes obtained from the fit of different decay-time spectra in coincidence with protons are drawn as a function of proton energies. The half-life given in the main figure is the recommended value (see text).

Fig. 3. The inset shows the decay-time spectrum in the beam-off period fitted by an exponential function added to a constant for the background. In the main figure, the half-lifes obtained from the fit of different decay-time spectra in coincidence with protons are drawn as a function of proton energies. The half-life given in the main figure is the recommended value (see text).

The branching ratio for a transition is defined as the total number of counts in the conditioned proton spectrum divided by the total number of counts in the unconditioned proton spectrum after subtraction of the background contribution.

After implantation of $^{22}\text{Al}$ in the E4 detector, an energy spectrum of β-delayed charged particles was obtained as shown in figure 4. The figure shows the spectrum without any condition and, shaded, the spectrum in coincidence with β particles detected in the last detector. As discussed in the simulation section (section 3), this condition provides a better energy resolution of the proton peaks. Up to 25 peaks are visible. The energy measured is the total decay energy which was corrected for the pulse-height defect. The branching ratio for a transition is defined as the total number of counts in the condition proton spectrum divided by the total number of counts in the unconditioned proton spectrum after subtraction of the background contribution.

The ratio $R_2 = 53.5\%$ is determined from a fit of a relatively isolated proton peak in the conditioned proton spectrum (E5 > 0). We have chosen the peak labelled 1 at E=2.583 MeV. The low energy part of the peak was fitted by a Gaussian while the total peak was fitted by a Gaussian added to an exponential function for the high-energy tail.

Each proton peak in the conditioned spectrum was fitted by a Gaussian, the area obtained was divided by the total number of counts in the unconditioned spectrum (E5 > 0) by the total number of counts in the unconditioned proton spectrum after subtraction of the background contribution.

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Each proton peak in the conditioned spectrum was fitted by a Gaussian, the area obtained was divided by the number of $^{22}\text{Al}$ implanted. The branching ratios obtained were corrected for the proton detection efficiency.

The half-life was measured in beam-on/beam-off mode (120 ms of implantation, 300 ms beam-off). The inset in figure 3 shows the decay-time spectrum in the beam-off period fitted by an exponential function added to a constant for the background. The decay of the daughter $^{22}\text{Mg}$ is the main contributor to this background ($T_{1/2} = 3.857 \text{ s}$). The half-life obtained with this method is $T_{1/2} = 91.9 \pm 1.4 \text{ ms}$.

A second analysis was performed by generating, from the same data set, the decay-time spectra in coincidence with the different proton groups. Each point in the figure represents in y axis the half-life obtained from the fit of a decay-time spectrum in coincidence with a range of proton energies in x axis. The error-weighted mean value for the half-life from this analysis is $T_{1/2} = 91.1 \pm 0.5 \text{ ms}$. The two points at low energy were not used to avoid the contribution from the contaminants. In fact, the main contaminant $^{22}\text{Mg}$ decays only by β−γ and the β particles associated contributes to the proton spectrum at low energies. However, this value is more accurate than the previous one which contains a higher contamination.

A third analysis with a continuous implantation mode has produced a value of $T_{1/2} = 87.3 \pm 1.1 \text{ ms}$. Since the second method is less subject to errors than the other methods, we recommend the value of $T_{1/2} = 91.1 \pm 0.5 \text{ ms}$ as the half-life of $^{22}\text{Al}$. 

2. Beta-delayed charged particle decay

The areas of the proton peaks were determined using the method described in the simulation section (see text). The ratio $R_1$ was determined by dividing the total number of counts in the unconditioned proton spectrum (E5 > 0) by the total number of counts in the unconditioned proton spectrum after subtraction of the background contribution. The ratio $R_2 = 53.5\%$ is determined from a fit of a relatively isolated proton peak in the conditioned proton spectrum (E5 > 0). We have chosen the peak labelled 1 at E=2.583 MeV. The low energy part of the peak was fitted by a Gaussian while the total peak was fitted by a Gaussian added to an exponential function for the high-energy tail.

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Each proton peak in the conditioned spectrum was fitted by a Gaussian, the area obtained was divided by the number of $^{22}\text{Al}$ implanted. The branching ratios obtained were corrected for the proton detection efficiency.

The branching ratios of the peaks 15, 17 and 23 are given in table 1 and compared with previous measurements of Blank et al. [3].

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have not been corrected for proton detection efficiency as the other peaks because, as it will be shown later, peak 17** is assigned to a β-2p transition to the 20Ne first excited state and peak 15* corresponds to a β-α decay to the 18Ne first excited state. In these two cases, the β-2p and β-α peaks are superimposed on proton transitions. The only way to distinguish between these transitions and the proton emission, to get the branching ratios, is to use the coincidences with γ-rays.

The peak 23** with $E = 6.085\pm0.008$ MeV is compatible with the β-2p decay measured by Cable et al. [3] to the 20Ne ground state. For the β-2p transitions, the proton detection efficiency was taken at half of the transition energy.

3. γ-ray spectrum

Figure 5 shows eleven γ rays detected in the germanium clover detector during the 22Al measurement. Only six of them (labelled 2, 4, 6, 9, 10 and 11) are in coincidence with a low-energy event in the E4 detector (E4 < 400 keV). This energy range corresponds mainly to β particles emitted without a coincident proton. The other γ rays 1, 5, 7 and 8 will be discussed below.

No γ-γ coincidences could be established using the diagonally opposite detectors of the clover. This lack of coincidences does not mean that there are no γ cascades, but it is rather due to the relatively low efficiency of each individual segment of the germanium detector and to low statistics.

– β-γ decay

In figure 5 the γ ray 2 at 351 ± 2 keV comes from the β-γ decay of 21Na [4]. The γ ray 4 at 583.7 ± 1.6 keV corresponds to the β-γ decay of
This work & Blank et al. [4]

| Peak | Energy (MeV) | Br (%) | Energy (MeV) | Br (%) |
|------|-------------|--------|-------------|--------|
| 1    | 0.475±0.008 | 4.73±0.63 | 0.45±0.04 | 6.4±1.2 |
| 2    | 0.721±0.008 | 7.39±1.01 | 0.72±0.04 | 6.8±1.2 |
| 3    | 0.975±0.008 | 0.25±0.05 |             |        |
| 4    | 1.033±0.008 | 3.00±0.34 | 1.04±0.04 | 3.9±1.2 |
| 5    | 1.223±0.008 | 0.75±0.10 |             |        |
| 6    | 1.299±0.008 | 18.5±1.74 | 1.32±0.04 | 18.0±1.0 |
| 7    | 1.551±0.010 | 0.81±0.16 |             |        |
| 8    | 1.753±0.008 | 0.45±0.08 |             |        |
| 9    | 2.072±0.008 | 0.48±0.07 |             |        |
| 10   | 2.503±0.010 | 0.64±0.13 |             |        |
| 11   | 2.583±0.008 | 4.89±0.24 |             |        |
| 12   | 2.838±0.008 | 2.11±0.09 |             |        |
| 13   | 3.088±0.008 | 1.89±0.07 |             |        |
| 14   | 3.484±0.008 | 2.18±0.15 |             |        |
| 15   | 4.017±0.008 | 1.04±0.33 | 3.997±0.049 | 0.31±0.09 |
|      |             |        | 4.10±0.06 | 2.1±1.0 |
| 16   | 4.224±0.009 | 0.84±0.11 |             |        |
| 17   | 4.464±0.008 | 2.52±0.14 |             |        |
| 17** | 4.464±0.008 | 0.69±0.08 | 4.48±0.25 | 0.9±0.5 |
| 18   | 4.912±0.010 | 0.27±0.32 | 4.92±0.07 | 2.4±1.0 |
| 19   | 5.177±0.013 | 0.29±0.11 |             |        |
| 20   | 5.667±0.008 | 0.35±0.11 |             |        |
| 21   | 5.808±0.049 | 0.18±0.55 |             |        |
| 22   | 5.909±0.056 | 0.21±0.62 |             |        |
| 23** | 6.085±0.008 | 0.41±0.07 |             |        |
| 24   | 6.774±0.008 | 0.41±0.12 | 6.62±0.10 | 0.7±0.3 |
| 25   | 7.517±0.011 | 0.33±0.07 | 6.93±0.10 | 0.1±0.05 |
|      |             |        | 8.29±0.10 | 0.2±0.05 |

Table 1. Total decay energies of β-delayed protons and their branching ratios corrected for the proton detection efficiency except for the transitions labelled by * and by **. The transition labelled by * is a β-α transition superimposed on a proton transition. The branching ratio of the α transition was derived from the intensity of a γ-ray observed in coincidence (see text). The two transitions labelled by ** are β-2p transitions (see text). The present results are compared to a previous measurement by Blank et al. [4].

$^{22}Mg$ [15]. The other γ-rays 6, 9, 10, 11 at 1248.5±2.0 keV, 1985.6±1.3 keV, 2062.3±1.5 keV and 2145±5 keV are attributed to the β-γ decay of $^{22}Al$. In fact, four γ-ray transitions at 1246.98±0.03 keV, 1984.8±0.1 keV, 2061.09±0.05 keV and 2143.5±0.6 keV were previously established in $^{22}Mg$ (see [16], [14]). The γ-rays 6 and 10 should correspond to these known transitions. However, according to the reference [16], a known γ-ray at 1984 keV should be accompanied by another γ-ray at 891 keV with a relative branching ratio of 67%. From 657 counts observed in the γ peak at 1985.6 keV we would expect 837 counts at 891 keV. This is not the case. We deduce therefore that the γ-ray 9 measured at 1985.6 keV is compatible with the new adopted γ transition in reference [15] with 100% relative branching ratio. From our measurement, this γ-ray is most likely emitted from a level situated at 5.294±0.003 MeV in $^{22}Mg$. This assumption is in agreement with the adopted value of 5.29311±0.00015 MeV and supported by theoretical calculations (see section [15] below). The weak γ-ray labelled 11 in fig. [3] is situated at 2145±5 keV. It may correspond to a known γ-ray in $^{22}Mg$ at 2143±5 keV. Since the energy range is limited to 3800 keV, this fact cannot be checked. We assume that the γ-ray 11 comes from this known level. The absolute intensities of the γ-transitions are given in table 2 corrected for γ efficiency and β-trigger efficiency. The β-trigger efficiency has been determined from the well known β-γ transitions of $^{24}Al$, also produced in this experiment for detector calibration. Two $^{24}Al$ γ-rays at 1077 keV and 1368 keV [16] have been used and lead to a value of 39±3% for the β-trigger efficiency.
### Table 2. Measured energies and absolute intensities of γ-rays from the β-γ decay of $^{22}\text{Al}$.

| Energy (MeV) | Intensity (%) | Parent nucleus |
|--------------|---------------|-----------------|
| 583.7 ± 1.6  | 47.4 ± 9.3    | $^{22}\text{Mg}$ |
| 1248.5 ± 2.0 | 38.2 ± 6.9    | $^{22}\text{Al}$ |
| 1985.6 ± 1.3 | 31.1 ± 5.4    | $^{22}\text{Al}$ |
| 2062.3 ± 1.5 | 34.1 ± 5.8    | $^{22}\text{Al}$ |
| 2145 ± 5     | 1.7 ± 0.7     | $^{22}\text{Al}$ |

The intensities are corrected for the detection and the trigger efficiencies. The first γ-ray comes from the decay of $^{22}\text{Mg}$ and corresponds to the total β-γ decay probability of $^{22}\text{Al}$.

#### γ-particles coincidences

Four γ-rays noted 1, 5, 7 and 8 in figure 4 were measured at 332±1.2 MeV, 1112.9±2.4 MeV, 1385.5±1.3 MeV and 1633.8±2.2 MeV. The three first rays are known as γ transitions in $^{21}\text{Na}$, which suggest a β-proton-γ decay from $^{22}\text{Al}$ to $^{21}\text{Na}$. The E4 spectrum conditioned by these 3 γ-rays shows coincidences. We have observed that the γ-ray at 332 keV is in coincidence with proton lines 1, 4, 6 and 11 (figure 4). The coincidence spectrum with the γ-ray at 1112.9 keV is not conclusive. Moreover, the γ-ray at 1385.5 keV is in coincidence with the proton peak 20. The charged-particle peak 17* with $E = 4.464 ± 0.008$ MeV is compatible with the β-2p decay energy measured by Cable et al. [3] and Blank et al. [4] at $E = 4.48 ± 0.06$ MeV. The γ-ray spectrum observed in coincidence with this peak is shown in figure 6. It clearly shows a coincidence with a γ-ray at 1633.8 keV, corresponding to the transition from the first excited state to the ground state in $^{20}\text{Ne}$. This is a firm confirmation of the β-2p decay to the $^{20}\text{Ne}$ first excited state. The branching ratio of this β-2p decay is calculated from the number of counts in the 2p peak in the charged-particle spectrum conditioned by the γ-ray at 1633.8 keV. After correction for the γ and proton detection efficiencies and normalisation to the total number of $^{22}\text{Al}$ implanted, we find a branching ratio of 0.69 ± 0.08%. The energy of the charged-particle peak 15* ($E = 4.017 ± 0.008$ MeV) is compatible with the β-α transition measured by Blank et al. [4] at $E = 4.01 ± 0.05$ MeV. The γ spectrum conditioned by this peak shows a clear γ-ray at 1887 keV (see figure 7). This γ-ray energy corresponds to the γ transition from the first excited state to the $^{18}\text{Ne}$ ground state. Again, this observation is the first firm confirmation of this β-α transition. The branching ratio of the β-α decay to the first excited state of $^{18}\text{Ne}$ is 0.038 ± 0.017%.

### Fig. 6. Gamma energy spectrum in coincidence with the charged-particle peak at 4.464 MeV in the E4 detector. The coincidence of this charged-particle peak with the 1633.8 keV γ-ray from $^{20}\text{Ne}$ demonstrates that the peak 17* (at 4.464 MeV) is a β-2p decay to the first excited state in $^{20}\text{Ne}$.

### Fig. 7. Gamma energy spectrum in coincidence with the charged-particle peak 15* at 4.017 MeV in the E4 detector which confirms the β-α transition to the first excited state of $^{18}\text{Ne}$.

### 5 Shell model calculations

Shell model calculations were performed to investigate the β-decay of $^{22}\text{Al}$. One aim of these calculations was to determine the branching ratios to the different states in $^{22}\text{Mg}$. Different calculations already exist for this isotope [4, 5]. We have performed new calculations with an isospin non-conserving interaction. No significant difference has been observed compared to an isospin-conserving calculation.

These calculations were also performed for the following reason. To construct an experimental decay scheme, we have to interpret the different experimental decay peaks in terms of their absolute energy position as states in $^{22}\text{Mg}$. This procedure is subject to uncertainties, because we deal with a large number of peaks and a huge number of states and transitions. It is obvious that errors in the interpretation can result in an apparent change of the Gamow-Teller strength. To avoid this problem, we have determined the different branching ratios for the emission of γ rays, protons and sequential two protons from $^{22}\text{Mg}$ states. Brown [5] has already done this type of calculations in the case of $^{22}\text{Al}$ for the particle emission from
one unique state, the IAS in $^{22}Mg$. For the first time, we have calculated these branching ratios for all interesting states in $^{22}Mg$, from the ground state to levels up to a $Q_{\beta^+} = 17.56$ MeV. Thus, we are able to compare directly the experimental spectrum with the predicted one using a comparison between the branching ratios measured and those calculated (see figure 4).

We performed $sd$ shell-model calculations using the OXBASH code [17]. Two different interactions were used, the USD interaction [18] and the isospin non-conserving Ormand-Brown OB interaction [19]. Only the OB interaction allowed us to calculate the particle decay from the IAS in $^{22}Mg$, because it is isospin forbidden. Since the ground state spin of $^{22}Al$ is not known, we performed two calculations, one for a spin-parity of $3^+$ and one for a spin-parity of $4^+$. For the $\beta$ transitions, we used effective single-particle matrix elements -an overall quenching factor of 0.76 ± 0.03 is included- for the Gamow-Teller operator [20].

The phase-space factor $f$ for $\beta^+$-decay was calculated by means of a program available at [21]. For the particle transitions, it is necessary to calculate spectroscopic factors $\theta^2$ for all possible transitions from states in $^{22}Mg$ to states in the daughter nuclei. The partial widths were estimated in the standard way from the expression $\Gamma = 2\hbar^2\gamma^2 P(l, Q)$ where $Q$ is the particle-decay energy, $\gamma^2$ is the Wigner single-particle width, $P(l, Q)$ is the penetrability, and $l$ is the angular momentum of the transition. The penetrabilities are calculated in a Woods-Saxon well using the correct number of nodes of the wave functions.

The calculated decay schemes are similar to those presented in [4]. The calculated $^{22}Al$ half-life is 80 ms in the $3^+$ case and 85 ms in the $4^+$ case. It is similar to a previous study [6], where half-lives of $T_{1/2} = 78$ ms for a $3^+$ state and of $T_{1/2} = 85$ ms for a $4^+$ state were found. The difference between $3^+$ and $4^+$ calculations is not enough significant to prefer one or the other hypothesis.

However, the differences in the calculated branching ratios are more significant. For a $4^+$ ground state of $^{22}Al$, the branching ratio of $\beta$-$\gamma$ decay is 39.4%, whereas the same branching ratio is 48.9% for a $3^+$ ground state. Similarly, a $\beta$-$p$ branching ratio of 53% is obtained in the $4^+$ case, while this number drops to 39.5% in the $3^+$ case.

$\beta$-$2p$ emission has also been investigated as a sequential emission, a first proton is emitted from $^{22}Mg$, followed by a second proton emitted from an excited state in $^{21}Na$. If we consider each transition individually, weak branching ratios are obtained. Only one transition could be measured by our setup and corresponds to a branching ratio of 0.84% for a total proton energy of 4.215 MeV for a $4^+$ ground state of $^{22}Al$. However, our setup can not distinguish between transitions to different intermediate states in $^{21}Na$.

The measured energy is simply the difference between the initial state in $^{22}Mg$ and the final state in $^{20}Ne$. Thus the branching ratios of the transitions from the same initial state to the same final state were summed and tabulated in table 3. For both cases, the two transitions which can be detected are emitted from the calculated IAS in $^{22}Mg$ to the $^{20}Ne$ ground state and to the $^{20}Ne$ first excited state. In the $4^+$ case, the transition to the ground state is weaker than the transition to the first excited state while the opposite is calculated for the $3^+$ case.

| Initial | Final | Transition | Br(%) |
|--------|-------|------------|-------|
| 4^+   | 13910 | 0          | 5980  | 0.26 |
|       | 13910 | 1765      | 4215  | 1.19 |
|       | 13910 | 4206      | 1774  | 1.30 10^-2 |
| 3^+   | 14101 | 0          | 6171  | 0.61 |
|       | 14101 | 1765      | 4406  | 0.36 |

Table 3. Branching ratios for sequential $\beta$-$2p$ emissions calculated for a ground state spin-parity of $4^+$ and $3^+$. For each initial state in $^{22}Mg$ and final state in $^{20}Ne$ the branching ratios of different transitions were summed. The energies tabulated are the calculated excitation energies. The total proton energy of the transition in the fourth column can be obtained by the subtraction of the final state energy, added to the $2p$ separation energy, from the initial state energy. All energies are in keV.

6 Results and discussion

In this section, we compare our experimental results to previous measurements as well as to model predictions. This comparison will allow us to favour one of the two possible spin-parity assignments for the ground state of $^{22}Al$ and to construct a new $\beta$-decay scheme.

1. Half-life

The $^{22}Al$ half-life has been measured twice before this experiment. The first measurement was done by Cable et al. [2] and they found $T_{1/2} = 70^{+50}_{-35}$ ms. The second measurement was performed by Blank et al. [4]. The value measured was smaller $T_{1/2} = 59 \pm 3$ ms but consistent with the first value. This last value is in conflict with our value of $T_{1/2} = 91.1 \pm 0.5$ ms. In reference [4], the half-life was measured in a beam-on/beam-off mode where the beam was switched off in a time interval of only 100 ms. If we assume a half-life close to 90 ms, this time interval is too short to correctly fit the decay curve when contamination is present. A fit was performed with the present data on a range of 100 ms. The half-life obtained was $73.4 \pm 2.6$ ms. This indicates that performing a fit in such a short time interval reduced artificially the half-life value obtained, a fact also confirmed by a simulation. The values calculated with the $4^+$ (85 ms) and $3^+$ (80 ms) hypotheses are close to the present experimental value, slightly favouring a $4^+$ ground-state spin. However, no conclusion can be drawn from the half-life alone.

2. Experimental decay scheme
The informations extracted from the charged-particle spectrum (position of the peaks and their absolute intensity), from the measurement of γ rays and the coincidences allow us to establish a new experimental β-decay scheme for 22Al as shown in figure 8. For transitions other than from the IAS, only those with a branching ratio higher than 1% are shown. Otherwise, all the 22Mg levels deduced from this experiment are tabulated in table 4, with their branching ratios and log(ft) values.

Starting from the energy measured in the E4 detector, the position of the levels in 22Mg is calculated assuming that the levels are placed above the particle threshold, by adding the particle energy and the energy of the following γ ray when observed. Thus, the transitions 1, 4, 6, 11 in table 4 in coincidence with the

**Fig. 8.** 22Al decay scheme deduced from the present experiment. Arrows are drawn for transitions with clearly identified coincidences with γ-rays except for the transition to the 20Ne ground state. Levels supposed to emit protons to the 21Na ground state are just drawn as dashed lines. Since no γ coincidences have been determined for β-p transitions with a branching ratio less than 1%, these transitions have not been drawn for the clarity of the scheme. All the energies, the branching ratios and the log(ft) of the excited states in 22Mg determined from table 4 are tabulated in table 4. The branching ratios between brackets correspond to β-γ decays deduced from the measured intensities of the γ-rays assuming the relative γ intensities adopted (see reference [15]).
\( \gamma \)-ray at 332 keV are coming from the levels situated at 6.307, 6.865, 7.132 and 8.416 MeV, respectively. The transition 20 is in coincidence with the \( \gamma \)-ray at 1384 keV but no coincidences were clearly established with the \( \gamma \)-rays at 1113 keV and 332 keV. However, the energy level determined by adding the transition energy above the proton threshold to the energy of these three \( \gamma \)-rays leads to an energy compatible with the energy of the IAS. From this energy consideration, we assign the line 20 to a transition from the IAS to the \( ^{21}Na \) third excited state.

For the transition 24 at 6.774 MeV, no \( \gamma \)-rays coincidences were determined but if we add the energy of the \( \gamma \)-rays 332 keV and 1384 keV, we obtain an energy of a \( ^{22}Mg \) level in agreement with the energy of the IAS. In addition, the shell model calculations predict a proton transition of 6.637 MeV from the IAS to the \( ^{21}Na \) second excited state. So the line 24 can be assigned to a transition from the IAS to the \( ^{21}Na \) second excited state.

Finally, as shown in figure 8, five transitions can be attributed to an emission from the IAS in \( ^{22}Mg \). The IAS energy is determined as the mean value of level energies deduced from these five measured transitions. The \( \beta \) feeding to the IAS is determined as the sum of the branching ratios of these transitions.

3. Particle emission

In figure 9 the summed branching ratios for charged-particle emission is shown as a function of the proton energy transition. This sum allowed us to compare the experimental results with the calculations without any interpretation of the data. The comparison concerns only the branching ratio related to a transition without any assignment to a level. This is not possible in the case of the summed Gamow-Teller strength where a level scheme have to be assumed before comparison. The full points represent the experimental sum, while the lines represent the calculations performed with a OB interaction for both \( 3^+ \) and \( 4^+ \) spin-parities. The theoretical sum is limited to branching ratios larger than 0.25\% which is our experimental limit. There is a significant difference between the \( 3^+ \) and \( 4^+ \) case. The calculation for the \( 4^+ \) case is in much better agreement with the experiment.

For the \( \beta-2p \) emission, two transitions were established in this experiment: to the \( ^{20}Ne \) ground state, a transition of 6085 keV with a branching ratio of 0.41±0.07 \% and to the \( ^{20}Ne \) first excited state, a transition of 4464 keV with a branching ratio of 0.69±0.08 \%. According to table 3 the calculated transition energies for the \( 3^+ \) case are closer to experiment, but in the \( 4^+ \) case the branching ratio to the \( ^{20}Ne \) ground state is weaker than the branching to the \( ^{20}Ne \) first excited state as measured in the experiment. The calculated branching ratios in both cases are of the same order of magnitude. The comparison of the \( \beta-2p \) emission with the calculation do not give additional information on the spin-parity of \( ^{22}Al \) ground state. However, since the transition energies and the branching ratios calculated are close to the measured ones, we can conclude that the \( \beta-2p \) emission observed is probably mainly a sequential emission.

4. Gamow-Teller strength

Figure 10 shows that the calculated summed Gamow-Teller strength in the case of a \( 4^+ \) ground state is in better agreement with the experimental values than in the case of a \( 3^+ \) ground state. The experimental strength was deduced from the experimental decay scheme established in the present work. The theoretical sum is obtained from branching ratios above 0.25\%. For higher energies, the experimental values are lower than the calculated sum. This is probably due to an incomplete measurement of \( \beta \) decays to higher-lying states where the density of states increases and the branching ratios are small. Since figure 8 does not exhibit such kind of difference, it is also possible that an error in the interpretation of the origin of the exper-

| \( ^{22}Mg \) levels (MeV) | Br (%) | Log(\( f \)) |
|-------------------------|--------|------------|
| 5.294 ± 0.003           | 31.1 ± 5.4 | 4.87 ± 0.08 |
| 5.455 ± 0.004           | 5.86 ± 2.41 | 5.56 ± 0.20 |
| 6.221 ± 0.008           | 7.39 ± 1.01 | 5.33 ± 0.06 |
| 6.307 ± 0.008           | 4.73 ± 0.63 | 5.51 ± 0.06 |
| 6.476 ± 0.008           | 0.25 ± 0.05 | 6.75 ± 0.09 |
| 6.724 ± 0.008           | 0.75 ± 0.10 | 6.23 ± 0.06 |
| 6.865 ± 0.008           | 3.00 ± 0.34 | 5.60 ± 0.05 |
| 7.052 ± 0.008           | 0.81 ± 0.16 | 6.13 ± 0.09 |
| 7.132 ± 0.008           | 18.51 ± 1.74 | 4.75 ± 0.04 |
| 7.254 ± 0.008           | 0.45 ± 0.08 | 6.34 ± 0.08 |
| 7.573 ± 0.008           | 0.48 ± 0.07 | 6.25 ± 0.07 |
| 8.004 ± 0.009           | 0.64 ± 0.13 | 6.03 ± 0.09 |
| 8.339 ± 0.009           | 2.11 ± 0.09 | 5.44 ± 0.02 |
| 8.416 ± 0.008           | 4.89 ± 0.24 | 5.06 ± 0.02 |
| 8.589 ± 0.008           | 1.89±0.07   | 5.43 ± 0.02 |
| 8.985 ± 0.008           | 2.18±0.15   | 5.28 ± 0.03 |
| 9.518 ± 0.008           | 1.04 ± 0.33 | 5.47 ± 0.14 |
| 9.725 ± 0.010           | 0.84 ± 0.11 | 5.51 ± 0.06 |
| 9.965 ± 0.008           | 2.52±0.14   | 4.97 ± 0.03 |
| 10.413 ± 0.010          | 0.28±0.32   | 5.80 ± 0.50 |
| 10.678 ± 0.012          | 0.29±0.11   | 5.70 ± 0.17 |
| 11.309 ± 0.049          | 0.18±0.55   | 5.70 ± 1.40 |
| 11.410 ± 0.008          | 0.21±0.62   | 5.60 ± 1.30 |
| 13.018 ± 0.056          | 0.33±0.07   | 4.81 ± 0.10 |
| IAS: 14.012 ± 0.003     | 1.90 ± 0.20 | 3.56 ± 0.05 |

Table 4. Energy of the \( ^{22}Mg \) levels deduced from the \( ^{22}Al \) \( \beta \) decay. The two first levels were determined from \( \beta-\gamma \) transitions. The other ones were determined from the delayed charged particles detected and the \( \gamma \)-rays in coincidence when observed. For each level, the branching ratio from \( ^{22}Al \) is given as well as the \( \log(\!f) \) values calculated using the program available at [27]. The IAS energy was calculated as the mean value of all the transitions coming from this level.
The parity of $3^-$ dashed line is the result calculated with a ground-state spin-parity for the calculated sums. Only branching ratios above 0.25% were considered for the calculated sums.

**Fig. 9.** Summed delayed proton branching ratios. The full circles represent the experimental values, the full line shows the calculated values with a ground-state spin-parity of $4^+$ and the dashed line is the result calculated with a ground-state spin-parity of $3^+$. Both calculations were performed with the OB interaction. Only branching ratios above 0.25% were considered for the calculated sums.

![Branching ratios vs Proton energy](image)

**Fig. 10.** Summed Gamow-Teller strength for the $\beta$ decay of $^{22}$Al. The black circles represent the experimental values. The full line shows the shell-model results with a ground-state spin-parity of $4^+$, whereas the dashed line is due to a calculation with a ground-state spin-parity of $3^+$. Both calculations were performed with the OB interaction.

![Gamow-Teller strength vs Excitation energy](image)

5. The ground-state spin-parity of $^{22}$Al

Prior to the present work, no convincing determination of the $^{22}$Al ground-state spin was presented. The mirror nucleus $^{22}F$ has a $4^+$ ground state, but also a low-lying $3^+$ state. The shell model calculations suggest two main decays from $^{22}$Al to $^{22}Mg$ levels below the proton threshold. In the case of a $3^+$ ground state spin-parity, $^{22}$Al decays to a $^{22}Mg$ $2^+$ level at 4.401 MeV and to a $4^+$ level at 5.311 MeV. In the case of a $4^+$ ground state spin-parity, it decays to a $^{22}Mg$ $4^+$ level at 5.311 MeV, to a $3^+$ level at 5.525 MeV and weakly to a $4^+$ level at 3.313 MeV as shown by the decay scheme in the middle of figure [11] (the calculation using a $3^+$ value for the $^{22}$Al ground state is not shown on this figure). The mirror nucleus $^{22}F$ decays to two $^{22}Ne$ $4^+$ levels and to a $^{22}Ne$ $3^+$ level in the same range of excitation energy as shown by the partial $\beta - \gamma$ decay scheme on the left of figure [11]. The mirror nucleus $^{22}F$ decays to two $^{22}Ne$ $4^+$ levels and to a $^{22}Ne$ $3^+$ level in the same range of excitation energy as shown by the partial $\beta - \gamma$ decay scheme on the left of figure [11]. Experimental results show two $\gamma$-rays at 1985.6 keV and 2145 keV. Two levels can be deduced from these measurements at 5.294 MeV and 5.453 MeV as shown on the right scheme of figure [11].

A simple comparison between these decay schemes favours again the $4^+$ case because both $^{22}$Al and $^{22}F$ do not decay to a $2^+$ level around 4 MeV as calculated for the $3^+$ case.

Another way to analyse these results is to compare quantitatively the absolute $\gamma$-ray intensities measured and the intensities deduced from shell model calculation. The $\gamma$-ray intensities are calculated as the product of the calculated $\beta$-decay branching ratios from the $^{22}$Al ground state to the $^{22}Mg$ levels and the relative intensities of the $\gamma$ ray from these levels. The relative intensities are assumed to be the same as the intensities of the analogue transitions in the $^{22}Ne$ mirror nucleus.

In figure [12], the calculated intensities of the expected $\gamma$ rays - taking into account our limited range of $\gamma$ energies - for both cases are compared to the experimental results. We notice that only three $\gamma$-rays are predicted in the $3^+$ case whereas four $\gamma$-rays are predicted in the $4^+$ case as for the experimental results. A much better agreement is observed between the calculated intensities for a $4^+$ spin-parity and our experimental results.

Therefore, from the different pieces of evidence previously discussed, we deduce that the $^{22}$Al ground state spin-parity is most likely $4^+$.

6. $^{22}Mg$ levels assignment

Here we assume a $4^+$ spin assignment for the ground state of $^{22}$Al. With this condition, the following levels in $^{22}$Mg are discussed and are tentatively assigned:
Comparison of the partial $\beta$-$\gamma$ decay scheme of $^{22}F$ - the mirror nucleus of $^{22}Al$ - (on the left) with the calculated decay scheme obtained with a spin-parity of 4$^+$ for $^{22}Al$ ground state (in the middle) and the decay scheme deduced from this experiment (on the right). The partial $\beta$-$\gamma$ decay scheme of $^{22}F$ was taken from the reference [12].

- $E_x = 5.294$ MeV. Experimentally we concluded that the observed $\gamma$ ray at 1895.4 keV is compatible with the adopted $\gamma$ ray at 1894.8 keV which is emitted from a level situated at 5.293 MeV with a spin-parity from 3$^+$ to 5$^+$. In addition, the comparison between the $^{22}Mg$ and $^{22}Ne$ $\gamma$ decay schemes and the shell model calculations (see figure 11) shows that this ray is probably emitted from a 4$^+$ level. So we can establish that the level measured at 5.294 ± 0.003 MeV is a 4$^+$ one.

- $E_x = 5.453$ MeV. There is a known state positioned at $E_x = 5.452$ MeV with a spin-parity 2$^+$ or 3$^+$. A comparison with the mirror decay scheme and the shell model calculations (figure 11) allowed us to assign a spin 3$^+$ for this level.

- $E_x = 6.221$ MeV. This level is fed with a branching ratio of 7.4% and a log(ft) of 5.32. Shell model calculations predict the existence of a proton transition from a 4$^+$ level at 6.315 MeV with a branching ratio of 5.74%. In the mirror nucleus $^{22}Ne$, a 4$^+$ level is situated at 6.346 MeV and is fed by $\beta^-$ decay with a branching ratio of 7.0% and a log(ft) of 5.34. These facts are strong evidences that this level has spin 4$^+$.  

- $E_x = 6.307$ MeV and $E_x = 6.865$ MeV. These levels have branching ratios of 4.7% and 3.0%. Shell model calculations predict the existence of only two levels in this range of energy, positioned at 6.394 MeV and 6.844 MeV, with spin 3$^+$ and 4$^+$ and with branching ratios 2.67% and 8.24%. There are evidences (see [22]) that the first state corresponds to a 4$^+$ state measured at $E_x = 6.267$ MeV. In this case, the second state is assigned to a spin-parity 3$^+$.  

- $E_x = 7.132$ MeV. The strongest proton transition is predicted to originate from a 5$^+$ level at 7.270 MeV with a branching ratio of 13.2% while experimentally it decays from a level at 7.132 MeV with a branching ratio of 18.51% and a log(ft) of 4.75. In the mirror nucleus, a 5$^+$ level is situated at 7.422 MeV and is fed by $\beta^-$ decay with a branching ratio of 8.7% but with a log(ft) of 4.7. It is the strongest $\beta^-$ feeding after the two levels corresponding to those decaying by $\gamma$ emission. So the level measured at 7.132 MeV can be assigned to a spin-parity 5$^+$.  

7. Mass excess of $^{22}Al$

The mass excess of the $^{22}Al$ ground state can be deduced from the measurement of the $^{22}Mg$ IAS energy. Indeed, four states are experimentally observed for this $T=2$ isobaric multiplet:

(a) the $^{22}Al$ ground state with $T_Z = -2$.

(b) the $^{22}Mg$ excited state measured at an energy of $14011.7 \pm 3.5$ keV with $T_Z = -1$ and a mass excess of $-396.8 \pm 1.4$ keV [23].

(c) the $^{22}Ne$ excited state measured at an energy of $14070 \pm 40$ keV with $T_Z = +1$ and a mass excess of $-8024.3 \pm 0.22$ keV [23].

(d) the $^{22}F$ ground state with $T_Z = +2$ and a mass excess of $2794 \pm 12$ keV [23].

The last three states are experimentally measured and permit the determination of the IMME parameters. This equation gives the mass of states which belong to the same isospin multiplet: 

$$M(A, T, T_Z) = a(T, Z) + b(A, T) T_Z + c(A, T) T_Z^2.$$ 

By using the mass excess of the three known states, we calculated the IMME coefficients as follows: $a = 9652.7 \pm 40.2$, $b = -3784.6 \pm 20.1$
and $c = 177.6 \pm 20.4$. With these values, we determined the mass of the $^{22}$Al to be $17932 \pm 99 \text{ keV}$. This mass excess is close to the extrapolation from Audi et al.\cite{Audi1999} which predicts a mass excess of $18180 \pm 90 \text{ keV}$.

8. Mirror asymmetry

The spin-parity assignment of four levels in $^{22}$Mg allowed us to determine the asymmetry factor $\delta$ for analogue $\beta$ decays to $^{22}$Mg and to $^{22}$Ne states. This parameter is calculated as:

$$\delta = \left( \frac{ft}{ft^-} - 1 \right)$$

$f$ is the Fermi function and $t$ the partial half-life, $+(-)$ stands for $\beta^+(-)$ decays. Table 5 gives the asymmetry parameters determined for the decay to $^{22}$Mg and $^{22}$Ne mirror levels. The error bars of these numbers are relatively large especially for the $3^+$ state due to rather large uncertainty for the $\gamma$-ray intensity. No conclusive information can be drawn from these values. Nevertheless, for the other states, the asymmetry factors are in agreement with the values determined in the $sd$ shell.

Table 5. Asymmetry parameters $\delta$ of several analogue levels in $^{22}$Mg and $^{22}$Ne fed in the $\beta$ decays of $^{22}$Al and $^{22}$F mirror nuclei.

| spin | $^{22}$Mg | $^{22}$Ne | $\delta$ |
|------|-----------|-----------|---------|
| $4^+$ | 5.294 MeV | 5.324 MeV | 0.202±0.037 |
| $3^+$ | 5.453 MeV | 5.641 MeV | 0.995±0.460 |
| $4^+$ | 6.221 MeV | 6.346 MeV | -0.023±0.003 |
| $5^+$ | 7.132 MeV | 7.422 MeV | 0.122±0.014 |

7 Conclusion

In the present work, the decay of $^{22}$Al was studied after its implantation in a silicon detector. The nuclei were produced by fragmentation of a $^{36}$Ar primary beam in a carbon target and selected by the LISE3 facility at GANIL. The high purity of the secondary beam (93%) permits a precise measurement of the $^{22}$Al half-life (91.1±0.5 ms) and a good energy resolution for the measurement of $\beta$-delayed particle transitions.

The measurement of $\gamma$ rays has confirmed a $\beta - 2p$ decay branch to the $^{20}$Ne first excited state and a $\beta$-$\alpha$ decay to the $^{18}$Ne first excited state. In addition, it allowed for the spin-parity assignment of six levels in $^{22}$Mg.

A complete calculation of the $\beta$-delayed particle decay was performed using shell model calculations in the $sd$ shell. The comparison of the theoretical summed Gamow-Teller strength and the experimental values shows good agreement. As a conclusion, the $^{22}$Al $\beta$ decay is well described by the matrix elements developed by Brown and Wildenthal\cite{Brown1997} where the quenching factor is included.

The $^{22}$Al ground-state spin-parity was assigned to be most likely $4^+$. This assignment is based on a comparison of the present experimental results with theoretical calculations and the mirror nucleus. The measurement of the $\gamma$ rays and the Gamow-Teller strength distribution was essential to reach this conclusion.

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