An ultrashort pulse ultra-violet radiation undulator source driven by a laser plasma wakefield accelerator

Anania, M. P.; Brunetti, E.; Wiggins, S. M.; Grant, D. W.; Welsh, G. H.; Issac, R. C.; Cipiccia, S.; Shanks, R. P.; Manahan, G. G.; Aniculaesei, C.; van der Geer, S. B.; de Loos, M. J.; Poole, M. W.; Shepherd, B. J. A.; Clarke, J. A.; Gillespie, W. A.; MacLeod, A. M.; Jaroszynski, D. A.

Published in:
Applied Physics Letters

DOI:
10.1063/1.4886997

Publication date:
2014

Document Version
Publisher's PDF, also known as Version of record

Link to publication in Discovery Research Portal

Citation for published version (APA):
Anania, M. P., Brunetti, E., Wiggins, S. M., Grant, D. W., Welsh, G. H., Issac, R. C., ... Jaroszynski, D. A. (2014). An ultrashort pulse ultra-violet radiation undulator source driven by a laser plasma wakefield accelerator. Applied Physics Letters, 104(26), [264102]. https://doi.org/10.1063/1.4886997

General rights
Copyright and moral rights for the publications made accessible in Discovery Research Portal are retained by the authors and/or other copyright owners and it is a condition of accessing publications that users recognise and abide by the legal requirements associated with these rights.

- Users may download and print one copy of any publication from Discovery Research Portal for the purpose of private study or research.
- You may not further distribute the material or use it for any profit-making activity or commercial gain.
- You may freely distribute the URL identifying the publication in the public portal.
An ultrashort pulse ultra-violet radiation undulator source driven by a laser plasma wakefield accelerator

M. P. Anania, E. Brunetti, S. M. Wiggins, D. W. Grant, G. H. Welsh, R. C. Issac, S. Cipiccia, R. P. Shanks, G. G. Manahan, C. Aniculaesei, S. B. van der Geer, M. J. de Loos, M. W. Poole, B. J. A. Shepherd, J. A. Clarke, W. A. Gillespie, A. M. MacLeod, and D. A. Jaroszynski

Citation: Applied Physics Letters 104, 264102 (2014); doi: 10.1063/1.4886997
View online: http://dx.doi.org/10.1063/1.4886997
View Table of Contents: http://scitation.aip.org/content/aip/journal/apl/104/26?ver=pdfcov
Published by the AIP Publishing

Articles you may be interested in
Giga-electronvolt electrons due to a transition from laser wakefield acceleration to plasma wakefield acceleration
Phys. Plasmas 21, 123113 (2014); 10.1063/1.4903851

Injection and acceleration of electron bunch in a plasma wakefield produced by a chirped laser pulse
Phys. Plasmas 21, 063108 (2014); 10.1063/1.4884792

Induction of electron injection and betatron oscillation in a plasma-waveguide-based laser wakefield accelerator by modification of waveguide structure
Phys. Plasmas 20, 083104 (2013); 10.1063/1.4817294

Self-mode-transition from laser wakefield accelerator to plasma wakefield accelerator of laser-driven plasma-based electron acceleration
Phys. Plasmas 17, 123104 (2010); 10.1063/1.3522757

Wakefield driven by Gaussian (1,0) mode laser pulse and laser-plasma electron acceleration
Appl. Phys. Lett. 95, 091501 (2009); 10.1063/1.3187221
An ultrashort pulse ultra-violet radiation undulator source driven by a laser plasma wakefield accelerator

M. P. Anania,1,2 E. Brunetti,1 S. M. Wiggins,1 D. W. Grant,1 G. H. Welsh,1 R. C. Issac,1 S. Cipiccia,1 R. P. Shanks,1 G. G. Manahan,1 C. Aniculaesei,1 S. B. van der Geer,3 M. J. de Loos,3 M. W. Poole,4 B. J. A. Shepherd,4 J. A. Clarke,4 W. A. Gillespie,5 A. M. MacLeod,6 and D. A. Jaroszynski1,a)

1SUPA, Department of Physics, University of Strathclyde, Glasgow G4 0NG, United Kingdom
2INFN, Laboratori Nazionali di Frascati, I-00044 Frascati, Italy
3Pulsar Physics, Burgstraat 47, 5614 BC Eindhoven, The Netherlands
4ASTeC, STFC, Daresbury Laboratory, Warrington WA4 4AD, United Kingdom
5SUPA, School of Engineering, Physics and Mathematics, University of Dundee, Dundee DD1 4HN, United Kingdom
6School of Computing and Creative Technologies, University of Abertay Dundee, Dundee DD1 1HG, United Kingdom

(Received 14 May 2014; accepted 23 June 2014; published online 1 July 2014)

Narrow band undulator radiation tuneable over the wavelength range of 150–260 nm has been produced by short electron bunches from a 2 mm long laser plasma wakefield accelerator based on a 20 TW femtosecond laser system. The number of photons measured is up to $9 \times 10^{18}$ per shot for a 100 period undulator, with a mean peak brilliance of $1 \times 10^{18}$ photons/s/mrad$^2$/mm$^2$/0.1% bandwidth. Simulations estimate that the driving electron bunch r.m.s. duration is as short as 3 fs when the electron beam has energy of 120–130 MeV with the radiation pulse duration in the range of 50–100 fs.

The laser wakefield accelerator (LWFA) utilizes the strong electrostatic forces of femtosecond laser-driven plasma density waves as an accelerating medium to support strong electrostatic forces of femtosecond laser-driven sources at shorter wavelengths.8,9 However, of immediate interest in the VUV range are applications in ultrafast studies. Furthermore, sufficient bunch charge ($\geq 3$ pC) allows the peak current to reach kA levels, where single-pass high-gain FEL may become feasible.15

The Advanced Laser-Plasma High-energy Accelerators towards X-rays (ALPHA-X) accelerator beam line,3 as shown in Fig. 2(a), has been used for the investigations presented here. A Ti:sapphire laser pulse (central wavelength $\lambda_0 = 800$ nm, full-width at half-maximum duration = 36 fs, and peak intensity = $2 \times 10^{18}$ W/cm$^2$) is focused to a 20 μm waist (radius at $1/e^2$) at the leading edge of a 2 mm diameter helium gas jet to form a relativistic self-guided plasma channel with a relativistic plasma wavelength $\lambda_p = 2\pi c/\omega_p \approx 10.6 \mu$m, where $c$ is the speed of light, $\omega_p = (n_e e^2/\pi \kappa_m e)^{1/2}$, $n_e$ is the electron density, $\kappa_m$ is the permittivity of free space, and $e$ and $m_e$...
are the electron charge and mass, respectively. The normalized laser vector potential, initially, $a_0 = eA/m_ec^2 \approx 1$, where $A$ is the vector potential, grows to $a_0 > 3$ due to non-linear self-focusing and photon acceleration,\textsuperscript{16} which results in a trailing plasma bubble into which electrons are injected from the background plasma.

Electron beams exiting the accelerator are initially collimated using a triplet of miniature permanent magnet quadrupoles (PMQs).\textsuperscript{17,18} The field gradient of each quadrupole is $\sim 500$ T/m, and the triplet entrance is 30 mm from the accelerator exit. A triplet of electromagnetic quadrupoles (EMQs) then focuses the beam through the undulator. The respective EMQ field gradients are 2.47 T/m, 2.20 T/m, and 2.47 T/m. The quadrupoles are set for optimal transport of 130 MeV energy electrons (Fig. 2), and, within $\pm 10$ MeV of this design energy, the simulated electron bunch duration at the undulator entrance is predicted to be $\sim 3$ fs (Fig. 1). Experimental measurements of the duration that detect transition radiation generated by the beam passing through a metal foil perturb the beam too strongly for simultaneous use with the undulator, however, other studies on ALPHA-X\textsuperscript{19} and elsewhere\textsuperscript{20} show that the duration of the electron beam within 1 m of the accelerator is $\sim 1-2$ fs, and this is the basis for the beam transport simulations. Beam profile monitors at positions L1, L2, L3, and L4 comprise phosphor Lanex screens and 12-bit charge-coupled device (CCD) cameras.

An imaging dipole magnetic electron spectrometer (ES1) provides strong focusing in the horizontal and vertical planes thus enabling excellent energy resolution ($\sim 0.5\%–1.0\%$), which can be maintained over a wide energy range (83–196 MeV at the field strength of 0.52 T). Ce:YAG crystals positioned in the focal plane image electrons dispersed by the spectrometer magnetic field with the image captured on a 14-bit CCD camera. The electron beam dump after the undulator is a simple permanent dipole bending magnet that acts as a rudimentary compact electron spectrometer (ES2). This allows UV radiation and electron spectra to be captured simultaneously. The on-axis magnetic field strength of ES2 is 0.75 T suitable for electrons in the range of 20–250 MeV to be imaged on a Lanex screen by a 12-bit CCD camera. All Lanex and Ce:YAG screens (except L1) have been cross-calibrated against imaging plate measurements to determine the absolute electron beam charge.\textsuperscript{21}

The undulator (length 1.5 m, $N_u = 100$ periods, and $\lambda_u = 15$ mm) has a slotted pole planar design and the adjustable pole gap is set at 8.0 mm for these experiments (vacuum tube inner diameter is 6 mm). This gives a peak on-axis magnetic field strength $B_u = 0.27$ T and undulator deflection parameter $K = 0.38$. The slotted pole design of the undulator features a 5 mm by 1 mm slot cut out of the central section of the magnets. This provides a radial focusing force for electrons of energy up to $\sim 100$ MeV. Full details are given elsewhere.\textsuperscript{22} The distance from accelerator exit to undulator entrance is 3.52 m.

Undulator output radiation is detected using a vacuum scanning monochromator (with platinum-coated toroidal mirror and 300 lines/mm grating) and 16-bit CCD camera. The grating is positioned for a 344 nm detection bandwidth centred on 220 nm with a resolution of about 5 nm. Three elements attenuate the radiation signal: the toroidal mirror (peak reflectivity of 65%), the grating (peak efficiency of 25% at 150 nm), and finally the quantum efficiency of the camera (25% across the relevant spectral range). Laser light and plasma emission has been blocked by an aluminium foil (thickness 800 nm) positioned before the undulator at Lanex screen L3.

Removal of the PMQs enables the intrinsic divergence and profile of the electron beam to be observed on Lanex screen L1. The mean r.m.s. divergence is 3.5 mrad (Fig. 2(b)), which is reduced to 1 mrad (Fig. 2(c)) upon insertion of the PMQs, i.e., near collimation of the central part of the beam. The PMQs act as an energy bandpass filter, imparting large angle trajectories on electrons outside of their acceptance range. Hence, outlying swirls that are evident in the Lanex image are related to the low energy “tail” or pedestal of the electron beam. The main central part of the beam, comprising the higher energy quasi-monoeenergetic “main peak” electron bunch, is the sole part of the beam that is preferentially transported through the undulator. Electron energy spectra obtained with ES1 (Fig. 2(d)) illustrate the

![Image](image-url)
broad range of beam energies from the accelerator that are due to laser and plasma density fluctuations. Typically, \(\sim 30\%\) of the charge is contained in the main high energy peak and \(\sim 70\%\) in the (mainly) lower energy pedestal. At ES1, the main peak has a mean central energy of 104 ± 9 MeV, with a 5\% relative energy spread, and contains a mean charge of (1.1 ± 0.8) pC.

The L3 and L4 images (Figs. 2(e) and 2(f), respectively, before and after the undulator) indicate reasonable focusing and transport of the main peak electrons through the undulator. At L3, the mean r.m.s. width is 580 \(\times\) 510 \(\mu\)m (smallest 240 \(\times\) 290 \(\mu\)m). At L4, the mean width is 800 \(\times\) 710 \(\mu\)m (smallest 360 \(\times\) 400 \(\mu\)m). The GPT transport simulations also predict sub-mm beam widths at the screen positions but the energy-dependence inherent in transport and the fact that the electron beam energy is not captured simultaneously with either Lanex image makes direct comparison with the experimental mean values difficult. Further simulations estimate an acceptable focal waist of less than 150 \(\mu\)m close to the centre of the undulator\(^2\) \(^\text{2}\) when L3 and L4 widths are in the range of 200–400 \(\mu\)m, and we estimate that this condition has been satisfied experimentally for at least 25\% of all shots. An estimated beta function of 0.2 m indicates that matching to the undulator length (1.5 m) is typically not optimal. The mean measured charge, averaged over the 10 shots of highest charge, is 4.5 pC at L3 and 2.3 pC at L4, respectively. However, ES2 spectra (Fig. 3) contain a lower mean charge of (0.8 ± 0.4) pC but without a pedestal. Therefore, we conclude that beam loss along the undulator is dominated by the loss of low energy pedestal electrons and that beam loss for the “main peak” electrons is low (<30\%). The lower charge limits the peak current at the undulator entrance to \(\sim 0.35\) kA, according to Fig. 1.

Examples of measured radiation spectra and their corresponding ES2 electron spectra are shown in Fig. 3. No radiation is detected in the absence of an electron beam propagating through the undulator, while the classic radiation wavelength dependence on electron energy (wavelength scaling inversely with the square of the energy)\(^8\) is obtained. The ES2 beam dump demonstrated its value as a crude electron spectrometer with a mean measured central energy of (102 ± 8) MeV that agrees well with the expected mean energy of (99 ± 4) MeV, obtained from the mean radiation wavelength \(\lambda\) of (216 ± 16) nm. This electron energy is less than the 130 MeV design energy such that, from Fig. 1, the mean electron bunch duration at the undulator entrance is estimated to be 28 fs. The mean spectral bandwidth of the radiation is (69 ± 11) nm or (32 ± 7) %, decreasing to as low as 16\%, which is related to the electron beam properties\(^23\) such that (\(\Delta\lambda/\lambda\)\(^2\) \(\approx\) (2\(\sigma_r/\gamma\))\(^2\) + (\(\sigma_{\theta r}/\gamma\))\(^2\)), where \(\sigma_r/\gamma\) is the relative energy spread, \(\theta\) is the divergence, and the natural bandwidth \(1/N_u = 1\%\) has been neglected. Applying an electron beam divergence of 0.8 mrad from the L4 data, the dominant contribution to the spectral bandwidth is seen to be the electron energy spread (\(\sim 15\%\), in agreement with deconvoluted ES2 spectra). This is a larger spread than that measured for the main peak at ES1, which indicates that a significant proportion of the radiation may originate from the pedestal electrons that are lost in transit through the undulator.

The scaling of the number of detected photons with electron charge (taking into account the attenuation by the grating, etc.), as shown in Fig. 4(a), confirms that the undulator radiation emission is incoherent spontaneous synchrotron-like radiation. A non-linear scaling would have

![FIG. 3. False color images of four unprocessed undulator radiation spectra with corresponding ES2 electron spectra indicated. Respective values for number of detected photons (after processing for toroidal mirror, grating, and camera response), electron beam charge, and central energy are (a) 1.2 × 10\(^6\), 0.9 pC, and 92 MeV, (b) 7.7 × 10\(^6\), 1.6 pC, and 95 MeV, (c) 6.1 × 10\(^6\), 2.0 pC, and 108 MeV and (d) 4.0 × 10\(^6\), 1.3 pC, and 122 MeV.](image)

![FIG. 4. (a) Dependence of the number of photons \(N_{\text{phot}}\) on the electron beam charge where the solid line is a linear best-fit and (b) average \(N_{\text{phot}}/\text{charge}\) as a function of the electron energy binned at 5 MeV intervals except in the high energy range (115–139 MeV), where eight shots have been binned together. The total dataset comprises 145 shots.](image)
been evidence of coherent FEL emission. A significant increase in the efficiency of photon production has been obtained at lower electron energy, as shown in Fig. 4(b). This is relatively far from the nominal optimal transport energy (130 MeV) and can be attributed to the ever greater focusing effect imparted on electrons of less than 100 MeV energy by the slotted undulator field. Note that the total number of detected photons is up to $9 \times 10^6$, which is $\sim 1-2$ orders of magnitude greater than that obtained in the two previous experiments.\textsuperscript{5,7}

In summary, a bright tunable source of ultrashort pulse UV radiation has been demonstrated. The estimated mean number of photons per shot per mrad\textsuperscript{2} per 0.1% bandwidth is 2200 with a mean energy of 2.6 pJ and mean peak brilliance of $1 \times 10^{18}$ photons/s/mrad\textsuperscript{2}/mm\textsuperscript{2}/0.1% bandwidth (based on the measured radiation divergence of 2 mrad, estimated source diameter of 300 μm), and mean r.m.s. duration of 100 fs. This is higher than the estimated values of $6.5 \times 10^{16}$ and $1.3 \times 10^{17}$ photons/s/mrad\textsuperscript{2}/mm\textsuperscript{2}/0.1% bandwidth, respectively, obtained in the visible\textsuperscript{3} and extreme UV\textsuperscript{5} spectral ranges where larger photon beam size estimates were applied. In terms of laser-driven light sources, higher peak brilliance of $\sim 1 \times 10^{19}$ photons/s/mrad\textsuperscript{2}/mm\textsuperscript{2}/0.1% bandwidth is reported for an X-ray LWFA-Compton source (100 TW laser, 30 fs radiation pulse duration, and 50% radiation spectral bandwidth).\textsuperscript{24} Our relatively long radiation pulse duration is a sum of the mean electron duration (28 fs simulated) and mean radiation slippage duration ($N_e \lambda_e/c = 72$ fs). At the shortest observed wavelength of 150 nm, the radiation slippage duration is 50 fs. For shots predicted to have the shortest electron bunch lengths (down to 3 fs from Fig. 1), the radiation pulse duration could reduce to $\sim 10$ fs for user applications at the expense of fewer undulator periods and even shorter wavelength radiation, for example, $N_e = 40$ and $\lambda_e = 80$ nm. Furthermore, the average brilliance is limited by our pulse repetition frequency (PRF) of 1 Hz due to vacuum pump gas loading, but the current technological upper limit is the laser system PRF of 10 Hz. Future advances in repetition rates of gas delivery\textsuperscript{25} and high-power femtosecond laser systems\textsuperscript{26} will enable the average brilliance to be greatly improved.

Electron beam transport that minimizes bunch lengthening would reduce the mean electron bunch duration in the undulator and, hence, the radiation pulse duration from shot-to-shot. This would entail tuning the quadrupole settings (PMQ axial spacings, EMQ field strengths) for the given electron beam energy (dependent on laser and accelerator parameters). The desired wavelength for any particular application governs the precise experimental setup. A strong motivation for minimizing bunch duration, besides improving the temporal resolution for ultrafast spectroscopy, is the opportunity to improve the coherence properties of the radiation. Very short bunches or even those with a rapid longitudinal structural variation, such as a sharp rise or fall time, will have finite Fourier components at the radiation wavelength $\lambda$ that could drive coherent spontaneous emission (SASE) in FELs.\textsuperscript{28,29} At $\lambda = 150$ nm, for example, this corresponds to a threshold rise time $\sim 5$ fs. For our beam line setup as modelled in Fig. 1, the SASE seeding condition is close to being fulfilled around the resonant electron energy of $\approx 120$ MeV, however, the energy spread reported here would be around one order of magnitude too large for SASE to occur (the slice energy spread may be lower). Clearly, as $\lambda$ decreases, the requirement for a resonant Fourier beam component will be more difficult to satisfy, but very short (femtosecond) sub-structure has already been observed in LWFA electron beams.\textsuperscript{30}

We acknowledge support of the U.K. EPSRC (Grant No. EP/I018171/1), the EC’s LASERLAB-EUROPE (Grant Agreement No. 284464, Seventh Framework Programme), EuCARD-2 (Grant No. 312453, FP7) and the Extreme Light Infrastructure (ELI) European Project. We thank D. Clark and T. McCanny for technical support.

\textsuperscript{1}T. Tajima and J. M. Dawson, \textit{Phys. Rev. Lett.} 43, 267 (1979).
\textsuperscript{2}E. Esarey, C. B. Schroeder, and W. P. Leemans, \textit{Rev. Mod. Phys.} 81, 1229 (2009).
\textsuperscript{3}D. A. Jaroszynski, R. Bingham, E. Brunetti, B. Ersfeld, J. Gallacher, B. van der Geer, R. Issac, S. P. Jamison, D. Jones, M. de Loos, A. Lyachev, V. Pavlov, A. Reitsma, Y. Saveliev, G. Vieux, and S. M. Wiggins, \textit{Philos. Trans. R. Soc. A} 364, 689 (2006).
\textsuperscript{4}F. Gruener, S. Becker, U. Schramm, T. Eichner, M. Fuchs, R. Weingartner, D. Habs, J. Meyer-ter-Vehn, M. Geissler, M. Ferrario, L. Serafini, B. van der Geer, H. Backe, W. Lauth, and S. Reiche, \textit{Appl. Phys. B} 86, 431 (2007).
\textsuperscript{5}C. B. Schroeder, W. M. Fawley, F. Gruener, M. Bakeman, K. Nakamura, K. E. Robinson, C. Toth, E. Esarey, and W. P. Leemans, \textit{AIP Conf. Proc.} 1086, 637 (2009).
\textsuperscript{6}H.-P. Schlenvoigt, K. Haupt, A. Debus, F. Budde, O. Jäckel, S. Pfotenhauer, H. Schwoerer, E. G. R. Hoijer, J. G. Gallacher, E. Brunetti, R. P. Shanks, S. M. Wiggins, and D. A. Jaroszynski, \textit{Nat. Phys.} 4, 150 (2008).
\textsuperscript{7}M. Fuchs, R. Weingartner, A. Popp, Zs. Major, S. Becker, J. Osterhoff, I. Cortrie, B. Zeiter, R. Hoerlein, G. D. Tsakiris, U. Schramm, T. P. Rowlands-Rees, S. M. Hooker, D. Habs, F. Krausz, S. Karsch, and F. Gruener, \textit{Nat. Phys.} 5, 826 (2009).
\textsuperscript{8}P. Rebernik Ribic and G. Margaritondo, \textit{J. Phys. D: Appl. Phys.} 45, 213001 (2012).
\textsuperscript{9}B. W. J. McNeil and N. R. Thompson, \textit{Nat. Photonics} 4, 814 (2010).
\textsuperscript{10}G. C. Walker, W. Jarzeba, T. J. Kang, A. E. Johnson, and P. E. Barbara, \textit{J. Opt. Soc. Am. B} 7, 1521 (1990).
\textsuperscript{11}T. Kobayashi and Y. Kida, \textit{Phys. Chem. Chem. Phys.} 14, 6200 (2012).
\textsuperscript{12}A. H. Zewail, \textit{J. Phys. Chem. A} 104, 5660 (2000).
\textsuperscript{13}V. Sundstrom, \textit{Phys. Chem. Phys.} 59, 53 (2008).
\textsuperscript{14}S. B. van der Geer, O. J. Luiten, M. J. de Loos, G. Poeplau, and U. van Rienen, “3D space-charge model for GPT simulations of high brightness electron bunches,” in \textit{Institute of Physics Conference Series No. 175} (Institute of Physics, Bristol, UK, 2005), p. 101.
\textsuperscript{15}M. Gallans, G. Penn, J. S. Wurtele, and M. Zolotorev, \textit{Phys. Rev. Spec. Top. - Accel. Beams} 11, 060701 (2008).
\textsuperscript{16}A. J. W. Reitsma, R. A. Cairns, R. Bingham, and D. A. Jaroszynski, \textit{Phys. Rev. Lett.} 94, 085004 (2005).
\textsuperscript{17}E. Esarey, F. Gruener, S. Becker, M. Fuchs, D. Habs, R. Weingartner, U. Schramm, H. Backe, P. Kunz, and W. Lauth, \textit{Phys. Rev. Spec. Top. Accel. Beams} 10, 082401 (2007).
\textsuperscript{18}M. P. Anania, D. Clark, S. B. van der Geer, M. J. de Loos, R. Issac, A. J. W. Reitsma, G. H. Welsh, S. M. Wiggins, and D. A. Jaroszynski, \textit{Proc. SPIE} 7359, 735916 (2009).
\textsuperscript{19}M. R. Islam, E. Brunetti, R. P. Shanks, B. Ersfeld, R. C. Issac, S. Cipiccia, M. P. Anania, G. H. Welsh, S. M. Wiggins, A. Noble, R. A. Cairns, G. Raj, and D. A. Jaroszynski, “Near-threshold electron injection in the laser-plasma wakefield accelerator leading to femtosecond bunches,” Nature Physics (unpublished).
\textsuperscript{20}O. Lundh, J. Lim, C. Rechatin, L. Ammoura, A. Ben-Ismael, X. Davoine, G. Gallot, J.-P. Goddet, E. Lefebvre, V. Malka, and J. Faure, \textit{Nat. Phys.} 7, 219 (2011).
\textsuperscript{21}S. M. Wiggins, R. C. Issac, G. H. Welsh, E. Brunetti, R. P. Shanks, M. P. Anania, S. Cipiccia, G. G. Manahan, C. Aniculaesei, B. Ersfeld, M. R.
Islam, R. T. L. Burgess, G. Vieux, W. A. Gillespie, A. M. MacLeod, S. B. van der Geer, M. J. de Loos, and D. A. Jaroszynski, Plasma Phys. Controlled Fusion 52, 124032 (2010).

22B. J. A. Shepherd and J. A. Clarke, Nucl. Instrum. Methods Phys. Res., Sect. A 654, 8 (2011).

23J. G. Gallacher, M. P. Anania, E. Brunetti, F. Budde, A. Debus, B. Ersfeld, K. Haupt, M. R. Islam, O. Jaeckel, S. Pfotenhauer, A. J. W. Reitsma, E. G. Rohwer, H.-P. Schl envoigt, H. Schwoerer, R. P. Shanks, S. M. Wiggins, and D. A. Jaroszynski, Phys. Plasmas 16, 093102 (2009).

24N. D. Powers, I. Ghebregziabher, G. Golovin, C. Liu, S. Chen, S. Banerjee, J. Zhang, and D. P. Umstadter, Nat. Photonics 8, 28 (2014).

25Z. He, J. A. Nees, B. Hou, B. Bearepaire, V. Malka, K. M. Krushelnick, J. Faure, and A. G. R. Thomas, Proc. SPIE 8779, 877905 (2013).

26G. Mourou, B. Brocklesby, T. Tajima, and J. Limpert, Nat. Photonics 7, 258 (2013).

27B. W. J. McNeil, G. R. M. Robb, and D. A. Jaroszynski, Opt. Commun. 165, 65 (1999).

28D. A. Jaroszynski, R. J. Bakker, A. F. G. van der Meer, D. Oepts, and P. W. Amersfoort, Phys. Rev. Lett. 71, 3798 (1993).

29S. M. Wiggins, D. A. Jaroszynski, B. W. J. McNeil, G. R. M. Robb, P. Aitken, A. D. R. Phelps, A. W. Cross, K. Ronald, N. S. Ginzburg, V. G. Shpak, M. I. Yalandin, S. A. Shunailov, and M. R. Ulmaskulov, Phys. Rev. Lett. 84, 2393 (2000).

30C. Lin, J. van Tilborg, K. Nakamura, A. J. Gonsalves, N. H. Matlis, T. Sokollik, S. Shiraiishi, J. Osterhoff, C. Benedetti, C. B. Schroeder, C. Toth, E. Esarey, and W. P. Leemans, Phys. Rev. Lett. 108, 094801 (2012).