Disappearance of static magnetic order and evolution of spin fluctuations in Fe$_{1+\delta}$Se$_x$Te$_{1-x}$

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We report neutron scattering studies on static magnetic orders and spin excitations in the Fe-based chalcogenide system Fe$_{1+\delta}$Se$_x$Te$_{1-x}$ with different Fe and Se compositions. Short-range static magnetic order with the “bicollinear” spin configuration is found in all non-superconducting samples, with strong low-energy magnetic excitations near the (0.5, 0) in-plane wave-vector (using the two-Fe unit cell) for Se doping up to 45%. When the static order disappears and bulk superconductivity emerges, the spectral weight of the magnetic excitations shifts to the region of reciprocal space near the in-plane wave-vector (0.5, 0.5), corresponding to the “collinear” spin configuration. Our results suggest that spin fluctuations associated with the collinear magnetic structure appear to be universal in all Fe-based superconductors, and there is a strong correlation between superconductivity and the character of the magnetic order/fluctuations in this system.

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I. INTRODUCTION

Since the discovery of the first high temperature superconductor in the 1980’s, there has been a continuing effort to understand the origin of high-$T_c$ superconductivity. Studies on the cuprate systems seem to suggest that there is an intimate relationship between superconductivity and magnetism,$^{1-4}$ and recently this has been shown to be the case also for the newly discovered Fe-based superconductor families.$^{5,10}$ With non-superconducting parent compounds that have static antiferromagnetic (AF) order, charge doping in the Fe pnictides gradually suppress the AF order and induces superconductivity, for both the RFeAsO (“1111” system) and AFe$_2$As$_2$ (“122” system) families.$^{11-16}$ In addition, a “spin resonance” has been observed in the “122” system by inelastic neutron scattering,$^{17,18}$ showing a sharp increase of the magnetic scattering intensity at the “resonance” energy when the system goes into the superconducting phase. For the pnictides, both the static magnetic order in the parent compound and the “resonance” in the superconducting compounds occur around the in-plane wave vector (0.5,0.5) (using the 2-Fe unit cell), suggesting a “collinear” or “C-type” AF structure [see Fig. 1(a)]$^{19,20}$

The situation is slightly different in another Fe-based superconductor family, the iron chalcogenide Fe$_{1+\delta}$Se$_x$Te$_{1-x}$ (the “1:1” compound). Here the parent compound has a “bicollinear” or “E-type” AF order [see Fig. 1(b)]$^{21,22}$ modulated along the (0.5, 0) in-plane direction. Furthermore, the substitution of Se for Te, which induces superconductivity, does not directly modify the density of electrons in the conduction bands. Despite these differences, the spin resonance observed in the superconducting compositions of this family has been found to occur at the same (0.5, 0.5) in-plane wave-vector$^{23,24}$ as in the Fe pnictides, but rotated 45° from the ordering wave-vector of the parent FeTe compound. This suggests that the superconducting mechanism in the Fe pnictides and chalcogenides are likely to be quite similar. Nevertheless, there has also been a report implying that superconductivity can coexist with the “bicollinear” structure on the atomic scale.$^{25}$ Therefore, understanding how the magnetic structure (static or dynamic) evolves from the E-type bicollinear configuration of the parent compound to the C-type collinear configuration in the superconducting region in the 1:1 system is an important problem. There have been a number of theoretical studies$^{26,27}$ that provide some insight on this issue, but clearly a more systematic experimental study is highly desirable.

In this paper, we present our work using neutron scattering to probe the magnetic order/fluctuations in a few samples from the 1:1 family for Se dopings ranging from 30% to 50% and with varying superconducting properties. Our results suggest that static magnetic order exists in all non-superconducting samples. (Here, by non-superconducting we mean an absence of bulk superconductivity.) This order is short-ranged and occur at in-plane wave-vectors of the type (0.5, 0). For the fully superconducting samples, no static magnetic order is found. With the disappearance of static magnetic order, the associated low energy magnetic excitations near (0.5, 0) also go away, as one might expect. Magnetic excitations near (0.5, 0.5) gradually become dominant as the material becomes more superconducting. While Se doping plays an essential role, it is clearly not the only determining factor regarding the superconductivity and the magnetic correlations. Samples with similar Se doping but differing in Fe content can have very different superconducting properties and corresponding magnetic structures/fluctuations. Our results clearly indicate that static bicollinear (E-type) magnetic orders in the 1:1 system compete with and suppress superconductivity. Superconductivity only appears when the system evolves toward a fluctuating collinear C-type magnetic correlations, which appear to be universal across all known Fe-based supercon-
Table I: List of the Fe$_{1+\delta}$Se$_x$Te$_{1-x}$ samples used in our measurements, with their composition (δ, x), superconducting transition temperature ($T_c$), room-temperature lattice parameters (from powder x-ray diffraction), and crystal mass.

| Sample  | δ  | x  | $T_c$ (K) | $a$ (Å) | $c$ (Å) | mass (g) |
|---------|----|----|-----------|--------|--------|----------|
| SC30    | 0.00 | 0.3 | 14        | 3.815  | 6.140  | 12.7     |
| NSC30   | 0.05 | 0.3 | –         | 3.808  | 6.120  | 7.4      |
| SC50    | 0.00 | 0.5 | 14        | 3.811  | 6.129  | 9.0      |
| NSC45   | 0.05 | 0.45| –         | 3.807  | 6.047  | 6.4      |

The single-crystal samples used in the experiment were grown by a unidirectional solidification method. The samples, their nominal compositions, and various characteristic properties are listed in Table I. The bulk susceptibility results in Fig. 2 were obtained using a superconducting quantum interference device (SQUID) magnetometer. From the magnetization measurements we can see that, although both superconducting samples show evidence of diamagnetic response at around 14 K, SC50 is clearly better in quality as far as superconducting volume fraction is concerned. With a considerable portion of its bulk volume being non-superconducting, it is possible that there is phase separation in SC30. In fact, when measuring different small pieces (~ 1 mm size) from the same SC30 sample, the superconducting volume can vary from $\lesssim$ 10% to ~ 80%, suggesting that the superconducting and nonsuperconducting phases could be macroscopically separated in this sample. The other samples, NC30 and NC45, are mostly non-superconducting, with no more than 1% of the volume giving a superconducting response.

![FIG. 1: (Color online) (a) Schematic of the “collinear” C-type AF spin structure, with scattering intensities mainly around (0.5, 0.5). The square shows a unit cell with two Fe atoms. (b) Schematic of the “bicollinear” E-type AF spin structure, which contribute mainly to scattering intensities near (0.5, 0). (c) The schematic diagram of the neutron scattering measurements in the (HK0) zone. Dashed lines denote linear scans performed across (0.5, 0.5) and (0.5, 0) in the text.](image1)

![FIG. 2: (Color online) ZFC magnetization measurements by SQUID for SC30 (red), NSC30 (blue), SC50 (black) and NSC45 (green). The inset shows the same data from the non-superconducting samples with different scale.](image2)

Neutron scattering experiments have been carried out on the triple-axis spectrometers SPINS (inelastic scattering measurements of magnetic excitations for SC30 and NSC30, and all elastic measurements for static magnetic order), and BT-7 (inelastic scattering measurements for SC50 and NSC45) located at the NIST Center for Neutron Research. We used horizontal beam collimations of Guide-80°-S-80°-240° (S represents “sample”) for the inelastic scattering measurements on SPINS with fixed final energy of 5 meV and a cooled Be filter after the sample to reduce higher-order neutrons; collimations of Guide-open-S-80°-240° were used for the elastic measurements on SPINS, with an additional Be filter before the sample. At BT7, we used beam collimations of open-50°-S-50°-240° with fixed final energy of 14.7 meV and two pyrolytic graphite filters after the sample. The inelastic scattering measurements have been performed in the (HK0) scattering plane, as indicated in Fig. II(c). The data are described in reciprocal lattice units (r.l.u.) of $(a^*, b^*, c^*) = (2\pi/a, 2\pi/b, 2\pi/c)$. The elastic scattering measurements
have been taken in the \((H0L)\) scattering plane instead, since the order in the parent compound occurs at half integer \(L\) values (AF order along the \(c\)-axis). All data have been normalized into absolute units \((\mu_B^2\text{eV}^{-1}/\text{Fe})\), using incoherent elastic scattering intensities from the samples.

### III. RESULTS AND DISCUSSION

![Image](https://via.placeholder.com/150)

**FIG. 3:** (Color online) Elastic neutron scattering measurements performed on SC30 (top) and NC45 (bottom) near \((0.5, 0.5, 0)\). (a) and (b) are intensity profiles along \([100]\) direction (\(H\)-scans); (c) and (d) are scans along \([001]\) direction (\(L\)-scans). (e) and (f) show the magnetic peak intensity at \((0.5, 0.5)\) vs. temperature. Corresponding scans measured at \(T = 34\) K are used as background, and have been subtracted from all the data shown. The error bars represent the square root of the number of counts.

The static long-range magnetic order in the parent compound FeTe appears near \((0.5, 0, 0.5)\), suggesting a bicollinear \(E\)-type AF structure. With small Se doping, it was suggested that the order should gradually become short-ranged\(^{27,37}\) and eventually disappear. Our results, however, suggest that the order can still remain with relatively large Se dopings.

We performed elastic magnetic scattering measurements on all samples. For the SC50 sample, there is no elastic magnetic intensity at \((0.5, 0, 0.5)\), while magnetic peaks are observed for all three other samples. In Fig. 3, we plot \(H\) and \(L\) scans through this AF wave-vector for the SC30 and NC45 samples at \(T = 4\) K. The same measurements for NSC30 have been published in a previous report (see Fig. 2 of Ref. 27).

The \(H\) and \(L\) scans performed at higher temperature \((T=34\) K) show no peak structure and are therefore used as backgrounds to be subtracted from the data. All peaks are much broader than the resolution, indicating the short-ranged nature of the magnetic order. The \(H\)-scans are peaked near but not exactly at \(H = 0.5\), similar to the results reported for NSC30 and another sample with 27% Se doping.\(^{27}\) The \(L\)-scans, however, are qualitatively different. In previous reports on lower Se doped 1:1 compounds,\(^{27,32}\) the \(L\)-scan peaks around \(L = 0.5\), and intensity always goes to zero at \(L = 0\). So there the magnetic order is always AF along the \(L\)-direction, whether short- or long-ranged. Here we see that after background subtraction, the scattering intensity at \(L = 0\) is still appreciable. This suggests that although the magnetic order still has a modulation along the \(L\)-direction, which peaks around \(L = 0.5\), favoring an AF configuration between Fe planes, the order has become much more two-dimensional. In the NSC45 sample, the 3D long-range bicollinear AF magnetic order of the parent compound has not been entirely destroyed, but rather greatly reduced to 2D short-range order. The ordered moment per Fe is \(0.015(4)\mu_B\), much less than the value in the parent compound with long-range order.\(^{26}\) It is nevertheless, enough to destroy superconductivity. With this static magnetic order present, even with 45% Se doping, bulk superconductivity is still not achieved. In the SC30 sample, although the sample does show a superconducting phase transition at around 14 K, the superconducting volume is smaller than for the SC50 sample.

The ordered moment is about \(0.006(2)\mu_B/\text{Fe}\), also much less than that in the NSC45 sample, indicating that this order may be coming from only part of the sample. Therefore there is likely a phase separation in this SC30 sample, where two phases, one superconducting and another one with a short-range magnetic order, coexist.

With the tendency of forming static bicollinear magnetic structures in the non-superconducting samples, it is natural to expect to see magnetic excitation spectra around the \((0.5,0)\) in-plane wave-vector as well. Previous work has shown that the energy dispersion and intensity modulation along the \(L\) direction for magnetic excitations in the 1:1 compound is small.\(^{28,32}\) We can therefore choose to perform the inelastic scattering measurements in the \((H00)\) plane for \(L = 0\). In Fig. 4, we plot our results taken near \((0.5, 0, 0)\). The top two panels [Fig. 4(a) and (b)] show energy scans at \((0.5, 0, 0)\) at \(T = 4\) K and 25 K. Measurements for NSC45 and SC50 were taken on BT7 with a relatively coarse energy resolution \((\text{FWHM} \sim 1.7\text{meV})\) compared to those on SPINS (NSC30 and SC30, \(\text{FWHM} \sim 0.3\text{meV})\), and have a large, resolution-limited tail from scattering at \(h\omega = 0\). Constant-energy scans at \(h\omega = 0.5, 2,\) and 5 meV [Fig. 4(c) to (h)] along the \(K\) direction across \((0.5, 0, 0)\) clearly show that for NSC30, SC30 and NC45, there is significant spectral weight at low energies here. For both 30%-Se samples, where we have measurements with higher energy resolution, one can see that the intensity at \(h\omega = 0.5\) meV increases on warming from 4 K to 25 K. The increase is much less pronounced at \(h\omega = 5\) meV. This behavior is likely due to a transfer of spectral weight from the elastic peak into low energy channels when the static order dissolves with heating. For SC50, the low-energy spin excitations near \((0.5, 0, 0)\) are weak, and not strongly temperature dependent, which is consistent with the fact that there is no static order near \((0.5, 0)\) in this sample. The two small peaks near \(K = \pm 0.5\) observed from samples SC30 and SC50 at
$\hbar \omega = 5$ meV, suggest that there is additional spectral weight developing near the $(0.5,0.5,0)$ wave-vector, corresponding to dynamic collinear spin correlations in the superconducting samples.

In Fig. 5 we show measurements near $(0.5,0.5,0)$. For SC50, a clear “resonance” is observed when comparing the energy(98,739),(927,862)

For the non-superconducting samples, there is no temperature effect observed for data taken between 4 K and 25 K. For NSC30, we did a constant-energy scan near $(0.5,0.5)$ only at $\hbar \omega = 6.5$ meV, and the intensity is very low compared to either its own magnetic scattering near $(0.5,0)$ or those from the other samples near $(0.5,0.5)$. It is clear that the low energy spin excitations are mostly focused around $(0.5,0)$ for NSC30. The NSC45 sample has Se doping very close to SC50, and also very similar magnetic excitation spectrum near $(0.5,0.5)$ compared to that from the latter in its normal state ($T = 25$ K). However, with no superconducting transition, its spectrum at low temperature ($T = 4$ K) does not differ much from that at $T = 25$ K.

The implications of our results are very clear for NSC30 and SC50. For NSC30, a short-ranged static magnetic order is present at low temperature near $(0.5,0)$, corresponding to a 3D bccollinear E-type spin structure. Its low energy magnetic excitations are also focused near $(0.5,0)$. With the static order present, no superconductivity is achieved in this sample. For the SC50 sample, there is no static order and the low energy magnetic excitation spectrum is mostly shifted to the $(0.5,0.5)$ region, corresponding to collinear C-type spin correlations. Similar to the situation in the “122” or “1111” systems, this collinear configuration without static order appears to favor superconductivity.

The results for NSC45 are more complicated. Here, with Se doping close to SC50, the magnetic excitations near the $(0.5,0.5)$ point are rather similar to the superconductor, except that the “resonance” feature is missing. Apparently, having magnetic excitations near $(0.5,0.5)$ associated with the collinear spin configuration is not sufficient for superconductivity to emerge. Also having a static 2D-like magnetic order with bccollinear structure is able to completely suppress superconductivity in this sample. Of course, the tetragonal crystal structure gives no energetic distinction between the ordering wave vectors $(0.5,0)$ and $(0.5,0.5)$, so that the magnetic configuration is relatively soft. There is likely a mixed phase where the bccollinear and collinear magnetic configurations coexist, on a microscopic level, similar to that of the mixed C-E phase in manganites as suggested in Ref. 24.

The case for SC30 is, in fact, quite intriguing. A 2D-like...
short-range static order exists at low temperature, while low energy magnetic excitations are found both around (0.5, 0) and (0.5, 0.5) with comparable spectral weight. Therefore, the magnetic excitation spectrum actually looks very similar to that in NSC45, yet there is bulk superconductivity in SC30 when the static magnetic order is also present. Compared to NSC45, the ratio of spectral weight near (0.5,0.5) to that near (0.5,0) is higher in SC30, indicating a larger volume of the sample favoring a collinear spin configuration. The “resonance” occurs below $T_c$, showing an enhancement of spectral weight only near (0.5,0.5). This indicates that superconductivity only exists in the part of the sample with dynamic collinear spin correlations. Although it is conceivable that the static order and superconductivity could coexist in the same domains as suggested by previous $\mu$SR work, it is also possible to have a system with macroscopic phase separation, where the volume of local collinear or bicollinear regions are large enough to form separate domains. In this case, the features near (0.5,0) (elastic magnetic peak and low energy magnetic excitations), and those near (0.5,0.5) come from different regions. This scenario would be consistent with the (varying) susceptibility results for different small pieces taken at different locations from this sample, and agrees with results from all other samples where static magnetic order and superconductivity do not coexist locally.

Why would samples with similar Se content (e.g., NSC30 vs. SC30, NSC45 vs. SC50) show dramatically different behaviors? It is clear from previous work that the Fe content has a significant impact on the magnetism. For example, higher Fe content in the parent compound can drive the order from commensurate to incommensurate, while its effect has been less clear for the superconducting region. It seems unlikely that the excess Fe atoms act simply as isolated magnetic moments that destroy the superconductivity, since our observations of variations in static magnetic order and low energy spin excitations cannot be explained in such a simple manner. Theoretically, it has been predicted that lowering the height of the chalcogen (Te/Se) positions can drive the “1:1” system from the bicollinear to the collinear spin configuration, and Fe interstitials will certainly have an impact on Fe-Te/Se-Fe bond lengths and bond angles. There are several reports concerning the effect of excess Fe on the lattice parameters; our results indicate that the lattice parameters $a$ and $c$ both decrease slightly with increased Fe content (holding Se constant), which does not appear to be entirely consistent with the trends reported by others. To make further progress on this issue, it will likely be necessary to characterize the microstructure associated with specific compositions.

IV. SUMMARY

Despite uncertainties in the underlying cause of differences among our samples, it is evident that the magnetic structure/fluctuations are intimately correlated with the superconductivity in the 1:1 system. If the magnetic correlations in the system favor a bicollinear (E-type) spin configuration, which sometimes can eventually lead to static order, superconductivity is suppressed. The collinear (C-type) spin configuration, which is universal across the superconducting regions of all Fe-based superconductor families, is necessary, but not sufficient, for the emergence of bulk superconductivity. There are cases where magnetic correlations favoring these two configurations coexist and compete. Overall, our results suggest that magnetic correlations are important in the Fe-based superconductor families, and the proper tuning of these correlations may be the key for enhancing superconductivity.

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1. J. M. Tranquada, H. Woo, T. G. Perring, H. Goka, G. D. Gu, G. Xu, M. Fujita, and K. Yamada, Nature 429, 534 (2004).
2. H. F. Fong, P. Bourges, Y. Sidis, L. P. Regnault, A. Ivanov, G. D. Gu, N. Koshizuka, and B. Keimer, Nature 398, 588 (1999).
3. G. Xu, G. D. Gu, M. Hacker, B. Fauque, T. G. Perring, L. P. Regnault, and J. M. Tranquada, Nature Phys. 5, 642 (2009).
4. H. A. Mook, P. Dai, S. M. Hayden, G. Aeppli, T. G. Perring, and F. Dogan, Nature 395, 580 (1998).
5. Y. Kamihara, T. Watanabe, M. Hirano, and H. Hosono, J. Am. Chem. Soc. 130, 3296 (2008).
6. H. Takahashi, K. Igawa, K. Arii, Y. Kamihara, M. Hirano, and H. Hosono, Nature 453, 376 (2008).
7. X. H. Chen, T. Wu, G. Wu, R. H. Liu, H. Chen, and D. F. Fang, Nature 453, 761 (2008).
8. Z.-A. Ren, W. Lu, J. Yang, W. Yi, X.-L. Shen, C. Zheng, G.-C. Che, X.-L. Dong, L.-L. Sun, F. Zhou, et al., Chin. Phys. Lett. 25, 2215 (2008).
9. F.-C. Hsu, J.-Y. Luo, K.-W. Yeh, T.-K. Chen, T.-W. Huang, P. M. Wu, Y.-C. Lee, Y.-L. Huang, Y.-Y. Chu, D.-C. Yan, et al., Phys. Rev. B 84, 14262 (2008).
10. K.-W. Yeh, T.-W. Huang, Y.-I. Huang, T.-K. Chen, F.-C. Hsu, P. M. Wu, Y.-C. Lee, Y.-Y. Chu, C.-L. Chen, J.-Y. Luo, et al., Euro. Phys. Lett. 84, 37002 (2008).
11. C. de la Cruz, Q. Huang, J. W. Lynn, J. Li, W. R. Li, J. L. Zarestky, H. A. Mook, G. F. Chen, J. L. Luo, N. L. Wang, et al., Nature 453, 899 (2008).
12. J. Zhao, Q. Huang, C. de la Cruz, S. Li, J. W. Lynn, Y. Chen, M. A. Green, G. F. Chen, G. Li, Z. Li, et al., Nature Mater. 7, 953
(2008).

13 Q. Huang, J. Zhao, J. W. Lynn, G. F. Chen, J. L. Luo, N. L. Wang, and P. Dai, Phys. Rev. B 78, 054529 (2008).

14 C.-H. Lee, A. Iyo, H. Eisaki, H. Kito, M. T. Fernandez-Diaz, T. Ito, K. Kihou, H. Matsuhata, M. Braden, and K. Yamada, J. Phys. Soc. Jpn. 77, 083704 (2008).

15 J. Zhao, Q. Huang, C. de la Cruz, J. W. Lynn, M. D. Lumsden, Z. A. Ren, J. Yang, X. Shen, X. Dong, Z. Zhao, et al., Phys. Rev. B 78, 032504 (2008).

16 Y. Qiu, M. Kofu, W. Bao, S. H. Lee, Q. Huang, T. Yildirim, J. R. D. Copley, J. W. Lynn, T. Wu, G. Wu, et al., Phys. Rev. B 78, 052508 (2008).

17 J. Zhao, D. T. Adroja, D.-X. Yao, R. Bewley, S. Li, X. F. Wang, G. Wu, X. H. Chen, J. Hu, and P. Dai, Nature Phys. 7, 555 (2009).

18 H. Chen, Y. Ren, Y. Qiu, W. Bao, R. H. Liu, G. Wu, T. Wu, Y. L. Xie, X. F. Wang, Q. Huang, et al., Euro. Phys. Lett. 85, 17006 (2009).

19 A. D. Christianson, E. A. Goremychkin, R. Osborn, S. Rosenkranz, M. D. Lumsden, C. D. Malliakas, I. S. Todorov, H. Claus, D. Y. Chung, M. G. Kanatzidis, et al., Nature 456, 930 (2008).

20 M. D. Lumsden, A. D. Christianson, D. Parshall, M. B. Stone, S. E. Nagler, G. J. MacDougall, H. A. Mook, K. Lokshin, T. Egami, D. L. Abernathy, et al., Phys. Rev. Lett. 102, 107005 (2009).

21 S. Chi, A. Schneidewind, J. Zhao, L. W. Harriger, L. Li, Y. Luo, G. Cao, Z. A. Xu, M. Loewenhaupt, J. Hu, et al., Phys. Rev. Lett. 102, 107006 (2009).

22 S. Li, Y. Chen, S. Chang, J. W. Lynn, L. Li, Y. Luo, G. Cao, Z. Xu, and P. Dai, Phys. Rev. B 79, 174527 (2009).

23 D. S. Inosov, J. T. Park, P. Bourges, D. L. Sun, Y. Sidis, A. Schneidewind, K. Hradil, D. Haug, C. T. Lin, B. Keimer, et al., Nature Phys. 6, 178 (2010).

24 W.-G. Yin, C.-C. Lee, and W. Ku, arXiv:1003.0512v1 (2010).

25 S. Li, C. de la Cruz, Q. Huang, Y. Chen, J. W. Lynn, J. Hu, Y.-L. Huang, F.-C. Hsu, K.-W. Yeh, M.-K. Wu, et al., Phys. Rev. B 79, 054503 (2009).

26 W. Bao, Y. Qiu, Q. Huang, M. A. Green, P. Zajdel, M. R. Fitzsimmons, M. Zhernenkov, S. Chang, M. Fang, B. Qian, et al., Phys. Rev. Lett. 102, 247001 (2009).

27 J. Wen, G. Xu, Z. Xu, Z. W. Lin, Q. Li, W. Ratcliff, G. Gu, and J. M. Tranquada, Phys. Rev. B 80, 104506 (2009).

28 Y. Qiu, W. Bao, Y. Zhao, C. Broholm, V. Stanev, Z. Tesanovic, Y. C. Gasparovic, S. Chang, J. Hu, B. Qian, et al., Phys. Rev. Lett. 103, 067008 (2009).

29 H. A. Mook, M. Lumsden, A. Christianson, B. C. Sales, R. Jin, M. A. McGuire, A. Sefat, D. Mandrus, S. Nagler, T. Egami, et al., arXiv:0904.2178v1 (2009).

30 J. Wen, G. Xu, Z. Xu, Z. W. Lin, Q. Li, Y. Chen, S. Chi, G. Gu, and J. M. Tranquada, Phys. Rev. B 81, 100513(R) (2010).

31 D. N. Argyriou, A. Hiess, A. Akbari, I. Eremin, M. Korshunov, J. Hu, B. Qian, Z. Mao, Y. Qiu, C. Broholm, et al., arXiv:0911.4713v1 (2009).

32 S.-H. Lee, G. Xu, W. Ku, J. S. Wen, C. C. Lee, N. Katayama, Z. J. Xu, S. Ji, Z. W. Lin, G. D. Gu, et al., arXiv:0912.3205v1 (2009).

33 S. Li, C. Zhang, M. Wang, H.-q. Luo, E. Faulhaber, A. Schneidewind, J. Hu, T. Xiang, and P. Dai, arXiv:1001.1505v1 (2010).

34 W. Bao, A. T. Savici, G. E. Granroth, C. Broholm, K. Habicht, Y. Qiu, J. Hu, T. Liu, and Z. Mao, arXiv:1002.1617v1 (2010).

35 R. Khasanov, M. Bendele, A. Amato, P. Babkevich, A. T. Boothroyd, A. Cervellino, K. Conder, S. N. Gvasaliya, H. Keller, H. H. Klauss, et al., Phys. Rev. B 80, 140511(R) (2009).

36 C. Fang, B. A. Bernevig, and J. Hu, Euro. Phys. Lett. 86, 67005 (2009).

37 N. Katayama, S. Ji, D. Louca, S.-H. Lee, M. Fujita, T. J. Sato, J. S. Wen, Z. J. Xu, G. D. Gu, G. Xu, et al., arXiv:1003.4525v1 (2010).

38 T. Hotta, M. Moraghebi, A. Feiguin, A. Moreo, S. Yunoki, and E. Dagotto, Phys. Rev. Lett. 90, 247203 (2003).

39 C.-Y. Moon and H. J. Choi, Phys. Rev. Lett. 104, 057003 (2010).

40 T. M. McQueen, A. J. Williams, P. W. Stephens, J. Tao, Y. Zhu, V. Ksenofontov, C. Felser, and R. J. Cava, Phys. Rev. Lett. 103, 057002 (2009).

41 T. M. McQueen, Q. Huang, V. Ksenofontov, C. Felser, Q. Xu, H. Zandbergen, Y. S. Hor, J. Allred, A. J. Williams, D. Qu, et al., Phys. Rev. B 79, 014522 (2009).

42 R. Viennoisa, E. Gianninia, D. v. d. Marela, and R. Cerny, arXiv:0911.2081v2 (2009).

43 R. Hu, E. S. Bozin, J. B. Warren, and C. Petrovic, Phys. Rev. B 80, 214514 (2009).