Bulk synthesis of stoichiometric/meteoritic troilite (FeS) by high-temperature pyrite decomposition and pyrrhotite melting

Juulia-Gabrielle MOREAU1*, Argo JÖLEHT1, Jaan ARUVÄLI1, Mikko J. HEIKKILÄ2, Aleksandra N. STOJIC3, Thomas THOMBERG4, Jüri PLADO1, and Satu HIETALA5

1Department of Geology, Institute of Ecology and Earth Science, University of Tartu, Ravila 14A, Tartu 50411, Estonia
2Department of Chemistry, University of Helsinki, P.O. Box 55 (A.I. Virtasen aukio 1), FI-00014, Finland
3Institut für Planetologie, Westfälische Wilhelms Universität Münster, Wilhelm-Klemm-Str. 10, Münster 48149, Germany
4Institute of Chemistry, University of Tartu, Ravila 14A, Tartu 50411, Estonia
5Geological Survey of Finland, Neulamäentie 5, Kuopio FI-70211, Finland

*Corresponding author. E-mail: juulia.moreau@ut.ee

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Abstract–Stoichiometric troilite (FeS) is a common phase in differentiated and undifferentiated meteorites. It is the endmember of the iron sulfide system. Troilite is important for investigating shock metamorphism in meteorites and studying spectral properties and space weathering of planetary bodies. Thus, obtaining coarse-grained meteoritic troilite in quantities is beneficial for these fields. The previous synthesis of troilite was achieved by pyrite or pyrrhotite heating treatments or chemical syntheses. However, most of these works lacked a visual characterization of the step by step process and the final product, the production of large quantities, and they were not readily advertised to planetary scientists or the meteoritical research community. Here, we illustrate a two-step heat treatment of pyrite to synthesize troilite. Pyrite powder was decomposed to pyrrhotite at 1023–1073 K for 4–6 h in Ar; the run product was then retrieved and reheated for 1 h at 1498–1598 K in N2 (gas). The minerals were analyzed with a scanning electron microscope, X-ray diffraction (XRD) at room temperature, and in situ high-temperature XRD. The primary observation of synthesis from pyrrhotite to troilite is the shift of a major diffraction peak from ~43.2° 2θ to ~43.8° 2θ. Troilite spectra matched an XRD analysis of natural meteoritic troilite. Slight contamination of Fe was observed during cooling to troilite, and alumina crucibles locally reacted with troilite. The habitus and size of troilite crystals allowed us to store it as large grains rather than powder; 27 g of pyrite yielded 17 g of stoichiometric troilite.

INTRODUCTION

Iron sulfides play an essential role in understanding the Earth interior and accretional processes of planetesimals at the dawn of the solar system (Brown et al., 1984; Dibb et al., 2021; Fujita et al., 1999; Hirata et al., 1995; Lehner et al., 2010; Louzada et al., 2010; Mare et al., 2014; Morard et al., 2007, 2011; Tomkins, 2009; Xie et al., 2014). They are major components of Earth rocks, chondrites (Luning et al., 2019; McSween et al., 1991), iron meteorites (Oshtrakh et al., 2019), and achondrites (e.g., Martian meteorites, Rochette et al., 2001). Iron sulfides exist within a solid solution in the Fe-S system (Mare et al., 2014; Sharma & Chang, 1979; Tomkins, 2009; Waldner & Pelton, 2005) with the two principal sulfide endmembers: pyrite (FeS2) crystallizing in the cubic and troilite (FeS) in the hexagonal systems (stoichiometric pyrrhotite [Fe1−x,S] where x = 0). Whereas pyrite and pyrrhotite are either found in oxidizing or reducing environments, stoichiometric troilite is only stable in reducing environments and, thus, is rarely found in its pure form on Earth. Troilite with Fe/S ratio between 0.97 and 1.00 is characteristic in H, LL, LL chondrites and a few carbonaceous
chondrites (Schrader et al., 2021). Because troilite is abundant in planetary rocks other than on Earth, its study and direct observation are important.

Recent studies had a particular focus on shock metamorphism in ordinary chondrites (OCs) (Stöfler et al., 2018), where troilite plays a vital role as a marker of impact events (Kohout et al., 2014, 2020; Moreau et al., 2017, 2018, 2019; Moreau & Schwinger, 2021). The average troilite modal abundance in OCs is 5.5–5.9 wt% (McSween et al., 1991) or ~5 vol% (Hutchison, 2007). The shock melting and migration of troilite melt into solid silicates, from high shock pressures (>40 GPa), darkens the lithology of OCs (Britt & Pieters, 1989, 1994; Britt et al., 1989; Heymann, 1967; Keil et al., 1992; Kohout et al., 2014; Stöfler et al., 1991). The interest in reproducing this darkening in laboratory experiments in a controlled environment is high, because, on the one hand, ultraviolet, visible, near infrared (UV/Vis/NIR) and mid-infrared (MIR) spectra are essential to understand impact processes in the solar system and the distribution of asteroids in the Main Belt (DeMeo et al., 2009; DeMeo & Curry 2014; Güldemeister et al., 2021), and on the other hand, spectral properties of iron sulfides, and troilite in particular (Dibb et al., 2021), are an important asset for the study of planetary surface spectral properties and space weathering in general (Kracher & Sears, 2005; Matsumoto et al., 2020). Space weathering research, where physical samples of planetary or asteroidal surfaces are often not available for direct laboratory measurements, also focuses on experiments that are conducted with analog materials and rocks (e.g., Morlok et al., 2020). In order to simulate the asteroidal or planetary surface composition in an authentic way, genuine mineral phases and rocks must be used. Genuine troilite comparable to the meteoritic troilite is hardly available on Earth; therefore, such research would benefit from obtaining troilite material produced at low costs and as coarse grains.

In previous studies, the synthesis of pyrrhotite (Fe$_{1-x}$S) or troilite (FeS) was achieved using chemical (Akhtar et al., 2013; Pedoussaut & Lind, 2008; Roberts et al., 2018) or thermal treatment (Boyabat et al., 2003; Coats & Bright, 1965; de Oliveira et al., 2018; Onufrienok et al., 2020; Selivanov et al., 2008; Xu et al., 2019). For example, the synthesis of pyrrhotite by heating pyrite at high temperature (<1100 K) in CO$_2$ or N$_2$ atmosphere (Boyabat et al., 2003; de Oliveira et al., 2018) or the synthesis of troilite from pyrite at higher temperature under vacuum (1200 K; Onufrienok et al., 2020) was successfully achieved. However, the troilite obtained (quantities not readily known) at higher temperature (1200 K) was not entirely synthesized and X-ray diffraction (XRD) analyses showed the presence of several pyrrhotite compositions ([Fe$_{1-x}$S] with $x = 0.125$ and 0.025) in addition to troilite; the use of a vacuum may not have allowed the complete evacuation of sulfuric gas. Any high-temperature procedure reported in literature included chemical cleaning of the pyrite samples (acetone, HCl leaching, and other chemical treatments), careful preparation of grain sizes, and the use of tiny quantities of material at a time. The systematic study from Boyabat et al. (2003) investigated the influence of gas flow, grain size, temperature, and the exact time of high-temperature exposure. Forty-minute tests at 853–1023 K were identified as ideal experimental parameters for troilite synthesis at a gas flow of 1670 cm min$^{-1}$ and grain sizes below 0.0548 cm. Similar conditions were reproduced by de Oliveira et al. (2018), who also noticed that pyrite oxidation occurs in CO$_2$ atmosphere and that pyrrhotite specimens remain very porous after decomposition. For chemical synthesis (e.g., Pedoussaut & Lind, 2008), near stoichiometric troilite was obtained, but in minimal quantities (e.g., nanoparticles; Roberts et al., 2018).

In this article, we propose to revisit high-temperature treatment procedures with the addition of induced melting above the melting point of troilite (1465 K) after having produced pyrrhotite from pyrite decomposition (Boyabat et al., 2003). By doing so, we wish to:

1. reduce the porosity at the grain scale by melting the mineral product;
2. allow for large quantities of pyrite to be treated, so the troilite product can be batched and used for various experiments (e.g., shock recovery experiments, Langenhorst & Deutsch, 1994, Langenhorst & Hornemann, 2005; spectral analyses for planetary surface characterization and melt migration tests within chondritic-like rocks, Moreau et al., 2021);
3. offer a simplified approach because controlled environments (gas flow, grain size, custom tube furnaces) are often difficult to obtain in general;
4. obtain troilite, as the stoichiometric mineral FeS, and compare our results with natural troilite from meteorites to validate its meteoritical properties; and
5. lead to the complete synthesis of troilite and avoiding pyrrhotite by-products through longer heat treatment and purging of excess sulfur.

In this study, we provide a thorough characterization of the produced pyrrhotite and troilite using scanning electron microscope (SEM) and in situ high-temperature x-ray diffraction. We compiled analyses to characterize the grain shape and surface, identified contamination products, and observed corrosive interactions with the sample holder. Furthermore, we performed XRD analyses to compare the troilite product to the pyrrhotite
product and natural meteoritic troilite. Additionally, we provide structural and spectral in situ analyses of the pyrrhotite-to-troilite transition process using X-ray and spectral analytical tools in tube furnaces. To offer more flexible experimental conditions, we will also enhance the thermal treatment by applying longer test durations, higher temperatures, and sample quantities, in contrast to more controlled experiments (Boyabat et al., 2003; de Oliveira et al., 2018; Onufrienok et al., 2020). We provide these details to ensure relatively easy synthesis procedures to guarantee that troilite is reproduced in standardized quality for any planetary science application.

MATERIALS AND METHODS

Pyrite Decomposition Procedure

For the study, we used fresh pyrite crystals from Navajun, Spain; they were provided by the University of Tartu Natural History Museum (TUG 1608-6992-3, -5, -7) with masses of 40.2, 14.6, and 12.9 g, respectively. After grinding the pyrite crystals in an agate mortar, we rinsed the powder several times with HCl 6 M followed by ethanol (using filters) and dried the mixture overnight at 373 K. Each sample batch was vacuum sealed.

Some amounts of powder were placed in quartz ceramic boats inserted in the middle of a glass tube fitting a horizontal tube furnace. Oxygen was purged for an hour with an Ar gas flow (18 L h\(^{-1}\)) before slowly heating (5 K min\(^{-1}\)) of the samples, which were then kept at 653 K for an hour to purge pyrite from oxygen elements (Boyabat et al., 2003). With continuous Ar flow, the heat was then slowly increased (between 2 and 10 K min\(^{-1}\)) to a maximum temperature of 1023–1073 K and kept at this level for 4–6 h. The samples were cooled at a rate of 5 K min\(^{-1}\), then retrieved and immediately vacuum sealed.

Pyrrohitite Melting Procedure

The decomposed pyrites were placed in alumina (Al\(_2\)O\(_3\)) crucibles, which were inserted in a high-temperature tube furnace (Carbolite Gero 30–3000 °C) that was continuously flushed either with Ar or N\(_2\) gas (20 L h\(^{-1}\)). The assemblage was oxygen purged for 1 h and subsequently heated at a rate of 10 K min\(^{-1}\). The samples were kept at 1498–1598 K for 1 h before cooling them at a rate of 5 K min\(^{-1}\). The samples were retrieved when the oven temperature readings indicated room temperature (298 K). All samples were immediately vacuum sealed.

The decomposition and melting procedures are summarized in Fig. 1.

Analysis of Samples

Pyrite, decomposed pyrite, and melted pyrrhotite were analyzed. The powder samples were studied by XRD using the Bruker D8 Advance diffractometer with Ni-filtered Cu K\(\alpha\) radiation. Scanning steps of 0.013° from 10 to 97° and a total counting time of 522 s per step were used. The diffraction patterns were analyzed using the Diffrac.EVA software with ICDD Powder Diffraction database PDF4+ (2020) and the full profile analysis software Topas 6 (Bruker, Germany).

To produce SEM-BSE (backscattered electron) images of our samples, we used a variable pressure Zeiss EVO MA15 SEM equipped with an Oxford X-MAX energy-dispersive detector system (EDS). Spectra were collected using 20 keV accelerating voltage and focused electron beam, then processed by the Aztec software.

We also carried out an in situ high-temperature XRD (HTXRD) analysis of the pyrrhotite sample using an Anton-Paar HTK1200N furnace attached to a PANalytical X’Pert Pro MPD diffractometer (up to 1473 K). This was done to observe any structural changes in the pyrrhotite and compare the transformation with the general procedure applied in this article. All measurements were conducted using a 10 mm mask, with a programmable divergence slit set to 8 mm length and a programmable antiscatter slit set to ensure the observation of the same sample portion. A PIXcel detector in 1-D mode was used for detecting the scattered radiation, and the pulse height discrimination range was narrowed to reduce the fluorescence caused by iron in the sample. Prior to the measurements, the furnace was heated for 2 h at 575 K in 1e-5 mbar pressure to dispose the moisture absorbed on the
| Sample | Pyrite | Note | Decomposition<sup>a</sup> | Melting<sup>b</sup> | Melting<sup>b</sup> |
|--------|--------|------|---------------------------|------------------|------------------|
|        |        |      | Quantity (g)<sup>c</sup> | Heating (K min<sup>−1</sup>) | Decomposition (K) | Time (min) | Quantity (g)<sup>d</sup> | Gas | Heating (K min<sup>−1</sup>) | Melting (K) | Time (min) |
| Troilite1-1 | TUG 1608-6992-7 | <0.20 mm | 5 | 5 | 653 | 45 | 10 | 1023 | 240 | 0.4 | Ar | 10 | 1498 | 60 |
| Troilite1-2 | TUG 1608-6992-7 | 0.20–0.32 mm | 2 | 5 | 653 | 45 | 10 | 1023 | 240 | 0.4 | Ar | 10 | 1498 | 60 |
| Troilite1-3 | TUG 1608-6992-7 | >0.32–0.50 mm | 4 | 5 | 653 | 45 | 10 | 1023 | 240 | 0.4 | Ar | 10 | 1498 | 60 |
| Troilite2-1 | TUG 1608-6992-7 | <0.20 mm | Using same decomposed pyrrhotite for troilite1 | 0.4 | N<sub>2</sub> | 10 | 1598 | 60 |
| Troilite2-2 | TUG 1608-6992-7 | 0.20–0.32 mm | 0.2 | N<sub>2</sub> | 10 | 1598 | 60 |
| Troilite2-3 | TUG 1608-6992-7 | >0.32–0.50 mm | 0.4 | N<sub>2</sub> | 10 | 1598 | 60 |
| Troilite3-1 | TUG 1608-6992-3 | Unsorted | 8 | 5 | 653 | 60 | 2 | 1073 | 360 | ~5 | N<sub>2</sub> | 10 | 1598 | 60 |
| Troilite4-1 | TUG 1608-6992-3 | Unsorted—grinded for melting | 27 | 5 | 653 | 60 | 2 | 1073 | 360 | ~25<sup>d</sup> | N<sub>2</sub> | 10 | 1598 | 60 |
| Troilite5-1 | TUG 1608-6992-5 | Unsorted | 11 | 5 | 653 | 60 | 2 | 1073 | 360 | In situ high-temperature X-ray diffraction |

<sup>a</sup> Argon, 18 L h<sup>−1</sup> gas flow, 1 h oxygen purge, 5 K min<sup>−1</sup> cooling.
<sup>b</sup> 20 L h<sup>−1</sup> gas flow, 1 h oxygen purge, 5 K min<sup>−1</sup> cooling.
<sup>c</sup> Quantity of pyrite (decomposition) and pyrrhotite (melting) inserted in the oven before heating.
<sup>d</sup> 17.34 g of troilite recovered from alumina crucible.
porous furnace walls. The measurements on the pyrrhotite sample were conducted first at 298 K, then at 373 K, followed by measurements at 50 K intervals up to 1473 K. At the highest temperature, five measurements were conducted for 1 h. The sample was then cooled symmetrically back to room temperature. Longer measurements with a wider angular range were performed at 298 K before and after heating. The height position of the sample was automatically adjusted by the furnace software, but changes at the correct surface position are unavoidable after melting and recrystallization of the sample.

RESULTS

Decomposition and melting tests are compiled in Table 1. During these tests, two heating rates for decomposition, as well as two heat exposure durations for melting, and distinct inert gas species were applied in the ovens. In Figs. 2–7, we illustrate several aspects of the resulting decomposition and melting of pyrite and pyrrhotite. For tests labelled troilite1, troilite2, and troilite4, we provide SEM-BSE analyses with density maps (Figs. 5–7) of selected samples and BSE spectra qualitative analyses (Research Data). For tests labelled troilite1, troilite3, troilite4, and troilite5, we provide XRD analyses (Figs. 8–10). For the test labelled troilite5, we offer HTXRD results (Fig. 11).

Decomposition and Melting

Desulfurization of pyrite (FeS₂) to pyrrhotite (Fe₁₋ₓS) is reactive at the conditions imposed in the oven as seen in Fig. 2. The loss of sulfur leads to a loss of mass. Decomposition of pyrite (troilite4) led to a net loss of 6.6146 g on a 27.6519 g powdered pyrite sample, a loss of ~90% of the total sulfur content of pyrite from a hypothetical loss of 7.3904 g; pyrite (troilite5) had a ~91% loss of sulfur content. Decomposition of pyrite to pyrrhotite is also macroscopically observable by the darkening of the mineral (Fig. 2).

Melting of pyrrhotite (Fig. 3a) is heavily pronounced with the formation of troilite as brownish-golden nuggets (Fig. 3b). Cavities likely formed owing to degassing (Figs. 3c and 3d) of trapped inert gas during melting. A darker and rough crust at the surface of troilite is observed (Figs. 3c, 3d, and 3f). Similar dark and rough droplets cling on the walls of the alumina crucible (Fig. 3e). Thermal treatment of troilite within alumina crucibles may weaken the crucibles because some broke upon cooling and retrieval. Melting of larger quantities of pyrrhotite (~20 g, troilite4-1) leads to visible surficial contamination (Figs. 4b, 4d, and 4e) and precipitation of sulfur minerals at the outlet of the gas flow (Fig. 4c). The contamination shown in Figs. 4b, 4d, and 4e consists of needle-shaped minerals outgrowing from the crucible walls or from a gray and porous residue at the surface of the sample (“cap,” Fig. 4b.1; Fig. 7 where the porous element is displayed from beneath). Dendritic-shaped minerals also grew along with the needle minerals outward of the gray residue. The melted product is exposed after cleaning the surface, revealing coarse-grained crystals, when the crucible is forcefully opened. In this case, 17.34 g of the mineral was recovered, starting from an initial weight of pyrite of 27.65 g.

SEM-BSE Observations

Pyrite, pyrrhotite, and troilite all have very distinct features. From decomposition, pyrrhotite is highly...
porous (Fig. 5b), depicting pitting-like alteration after desulfurization of the sample in the oven, distinguishing it from pyrite, and illustrating mass loss. Cracks allow us to see that this pitting goes deep into the pyrrhotite grain, giving it a coral-like appearance. The pyrite image in Fig. 5a shows that phyllosilicates can also coexist within pyrite, but we rarely observed such occurrences.

The formation of troilite and its textural features is diverse. Troilite displays different shapes such as aggregates of crystals (Figs. 5c and 5d), single angular crystals (Fig. 7f), and rounded crystals (Figs. 5d and 6b). Exsolutions of Fe-metal are observed such as droplets on surface of troilite (Fig. 6a), crystallized grains (Figs. 5d and 6a), skeletal growths (Fig. 5d), and larger deposits on the walls of the crucibles with higher concentration of sulfur (Figs. 6e and 6f). Contamination of Al2O3 in troilite, and/or iron, is observed at crust level or the crucible wall above it (Figs. 6d and 6f), at the contact between the troilite grain and the crucible wall (Figs. 6a and 6c), or as tiny inclusions in the round-shaped troilite (Fig. 6b). Occasionally, we also observed a lining of troilite on needle-shaped alumina crystals (Fig. 5f). Corrosion of the alumina crucible is

Fig. 3. Pictures taken from a Leica M205-A microscope: (a) pyrrhotite powder (0.3–0.5 mm grains) used to synthesize (b) troilite (troilite1-1); (c) troilite (troilite2-2) in its alumina crucible with (d, e) close-ups of the recrystallized melt and its crust, respectively. Picture of (f) troilite (troilite3-1) in its alumina boat crucible. a) Pyrrhotite melted to form (b) a brownish-gold colored troilite. The recrystallized troilite in (c, d) shows cavities from trapped gas and (e) a darker crust that has a rougher surface compared to the recrystallized troilite beneath. (Color figure can be viewed at wileyonlinelibrary.com.)
visible in some places (Fig. 5e). Whereas exsolution of Fe-metal happens within the troilite crystallized melt, Al₂O₃ contamination only happens at crucible walls or the surface of the crystallized melt. Further analyses of troilite4-1 (Fig. 7) show that the needle-shaped mineral contamination atop the melt is titanium oxide minerals with a coating of Fe-metal or iron sulfide. The porous element, which sat atop the melt (Fig. 4) and shown in Figs. 7a–c, is composed mainly of Al₂O₃, with minor troilite inclusions. The dendritic minerals observed in Fig. 4d could not be readily observed in SEM-BSE. The upper layer right beneath the surface of troilite4-1 is mainly composed of troilite grains (Fig. 7f).

**X-Ray Diffraction**

In Fig. 8, the sequence of analyses of pyrite (TUG 1608-6992-7) over pyrrhotite to troilite is shown from the same series of decomposition and melting (troilite1-1,2,3). Pyrite, pyrrhotite (pyrrhotite-11T, approx.
39.2 wt%), and troilite (troilite-2H) are confirmed with the XRD analyses. Only troilite spectra display a peak of iron contamination (~3%).

Compiled analyses of troilite (troilite3) from the troilite sample displayed in Fig. 4f are shown in Fig. 9. The crust layer was delicately handpicked in different places of the sample. Individual grains of troilite were selected and kept as obtained or hand-cleaned from the crust and possible alumina contamination. All analyses show iron contamination (~2–4%), and only in analyses of the uncleaned grains mullite contamination was detected, which is the ceramic material used for the crucible.

We detected slight quartz contamination (~3%) in the intermediate pyrrhotite of troilite4-1 and troilite5-1; we also detected iron silicon oxide contamination in pyrrhotite of troilite5-1 (~0.2%).

More thorough analyses of troilite4-1 melting yielded the following, with spectra compiled in Fig. 10...
1. Troilite is confirmed for the material laying beneath the surface (Fig. 4f), away from the contamination of the needle-shaped minerals and other contamination mentioned earlier (see the SEM-BSE Observations section). For the picked grains, no iron or mullite was detected.

2. Surficial troilite is also confirmed for the material cleaned from major contamination; slight contamination of mullite and iron is detected in this case.

3. The cap element displayed in Fig. 4b.1 and Fig. 7a–c is composed in majority of mullite, and second, of rutile and troilite with slight iron contamination. Glass was also detected but was not analyzed for its composition.

4. The needles observed in Figs. 4b.2/3, 4d, 4e, 7d, and 7e are confirmed to be rutile crystals.

5. Other minerals that were analyzed from surface contamination in addition to the rutile minerals (Fig. 4b.2/3) are mullite, troilite, and, possibly, quartz.

6. The yellow material from the outlet of the oven is sulfur (Fig. 4c).
In Situ High-Temperature X-Ray Diffraction (HTXRD)

HTXRD measurements were conducted on the pyrrhotite sample (troilite5-1) as described in the Analysis of Samples section. First, the major observation (Fig. 11) is the successful synthesis of troilite as previously seen in this work corroborated by diffractograms corresponding to pyrrhotite (PDF 29-726) before heating and troilite (PDF 37-477) after cooling. Second, peaks attributed to troilite appear when the mineral solidifies after being ~100 min in a liquid state from heating stage (~1373 K) to cooling stage (down to ~1223 K); several troilite peaks appear during the last steps of cooling, <423 K, consistent with the change from the MnP-type structure to the space group (P 6 2c) (Oshtrakh et al., 2016). Finally, the troilite iron contamination observed earlier in this work is depicted by a peak appearing right at the crystallization of the troilite. Peaks appearing at the troilite liquid state are believed to be from the alumina crucible and from a hercynite phase FeAl2O4 possibly formed by reaction between the sample and the crucible. Any peak shifts

Fig. 7. Backscattered electron scanning electron microscope (SEM-BSE) images (density) of troilite4-1 test (Fig. 4): (a) microscope picture of a porous element, “cap” that sat atop the melt with troilite inclusions; (b, c) SEM picture of the porous element in (a) composed in majority of alumina minerals; (d, e) needle-shaped titanium oxide minerals that sat atop the melt; (f) troilite grains collected in the upper layer of the molten sample after cleaning of surface. The needle-shaped minerals also show inclusions/coating of iron and iron sulfide. Ti = titanium oxide-based minerals, FeS = iron sulfide-based minerals; Fe = iron-based minerals. The SEM image in (b) is highlighted in (a). (Color figure can be viewed at wileyonlinelibrary.com.)
during heating or cooling are due to the combination of sample displacement and thermal extension in the oven. While the thermal expansion behavior is very interesting, it is beyond the scope of this paper and already partially explained, for example, in Selivanov et al. (2008), although using a narrower temperature range. A few peaks remain unidentified; they could either be oxidation from accidental exposure of the sample to air prior to measurements or the possible presence of greigite, a common mineral inclusion of pyrite (Akhtar et al., 2013).

DISCUSSION

Our results illustrate the successful synthesis of pyrite to troilite with a two-step thermal treatment where pyrrhotite is the intermediate product. Any test allowed us to synthesize troilite, in large quantities, as shown by XRD analyses and with only slight contaminations of iron and/or alumina. The alumina contamination is believed to come from sulfur corrosion of the crucible in which pyrrhotite is melted (sulfur escapes the system). The contamination remained mostly at the contact between the walls of the crucible and troilite or as a “floating cap,” or slag on top of the melt, when large quantities of pyrrhotite were used for melting. The slag possibly formed from a combination of degassing and corrosion (owing to the reducing environment) and floated on top of the melt. In contrast, there was no observation of denser elements at the bottom of the melt (e.g., chromium oxides). The appearance of needle-shaped crystals of alumina within troilite at crucible walls may be a result of such contamination via redox reaction as well. The contamination of iron exsolution is possible as iron and iron sulfide coexist as a solid solution. The degree of iron contamination is constant for most experiments within different troilite samples (except the sublayer troilite for troilite5-1 where iron was not detected), which is not systematic for alumina contamination. The observation of iron exsolution droplets is also observed in cavities within troilite, where N₂ was possibly trapped during melting of pyrrhotite to troilite. Iron is,
therefore, a by-product of troilite synthetization and should appear during the cooling of the troilite melt (Tomkins, 2009). The presence of iron in the final product will affect the eutectic properties of the troilite and iron mixture by lowering the melting point (Mare et al., 2014). Finally, the titanium oxide minerals (rutile) are only observed atop the melting crust of troilite4-1, indicating the pyrite minerals had a certain amount of titanium to start with. They also grew from and outward the crucible or alumina minerals, suggesting a dissolution–reprecipitation mechanism necessary for such needle-shaped minerals to form (Hwang et al., 2007). This contamination was not observed in the 17.34 g of recovered troilite from troilite4-1 test. Indeed, using large quantities of pyrrhotite for melting yields better qualitative results of troilite products. Using Ar or N₂ atmospheres in the highest temperature oven did not affect our results.

The XRD patterns of the experimental products match those of meteoritic troilite exactly (iron meteorite troilite inclusion; Oshtrakh et al., 2016). We observed a shift of the major peak between pyrrhotite and troilite, from ~43.2°₂θ to ~43.8°₂θ. Interestingly, within the troilite synthesis study of Onufrienok et al. (2020), we cannot safely state that the peak shift we observe in their XRD data for troilite after thermal treatment is comparable with the peak shift observed between pyrrhotite and troilite in our own XRD data (Fig. 8); as a note, the thermal treatment for troilite conversion was not complete in their case. Formation of troilite is also illustrated in our HTXRD analyses. Troilite is formed during cooling of the molten precursor solid phase with final peaks appearing <423 K, a change of crystal structure typical to genuine troilite. The mineral is not readily solidified from 1473 K but rather solidifies only at ~1223 K. As stated, the exsolution of iron appears to be an effect of cooling to troilite. Also, the synthesized troilite has a brownish-gold appearance, which is consistent with its appearance in unaltered meteorites (e.g., Oshtrakh et al., 2016; Ray et al., 2016), an important aspect for studying spectral properties of troilite powders (Dibb et al., 2021) if iron contamination is absent (e.g., troilite4-1 deep layer in our experiments, Fig. 4f). As for most recent studies, scientists either used troilite out from iron meteorites (Dibb et al., 2021) to have the most spectrally genuine elements or they purchased synthesized FeS powder (Prince et al., 2020). In the latter case, the FeS powder
Our synthesized troilite is easily reusable for dedicated experiments (e.g., reflectance spectra analyses and shock recovery experiments) if it is properly extracted from the crucible wall to avoid alumina contamination; this contamination can easily be avoided by producing larger quantities of troilite which, then, yield better purity levels. However, the iron contamination may impact properties of troilite such as a reduced Fe-FeS eutectic melting temperature (Mare et al., 2014) or changes in reflectance spectra properties (Dibb et al., 2021). Such contamination may possibly be prevented by assuring more control on the melting environment (gas flow, trapped gas, better preparation of pyrrhotite), on the contamination from corrosion by using crucibles with similar thermal resistance but better corrosion resistance (e.g., tungsten), and by retrieving troilite farther away from surface contamination and crucible walls.

CONCLUSION

We synthesized troilite in large quantities (e.g., 17 g from an initial weight of pyrite of 27 g). The troilite product matched with XRD analyses of meteoritic troilite. The production of meteoritic troilite is a major asset for planetary science research, especially for those interested in shock metamorphism of chondrites or metallic meteorites and spectral analyses of planetary surfaces and space weathering. Our synthesized troilite may present a slight contamination in iron, which will lower the troilite melting temperature; this is not a disadvantage as it may help establish shock recovery experiments to study melting of troilite by lowering its melting temperature. The procedure we propose is simple and only requires fresh pyrite. The two-step heating process allowed us to observe the transition between pyrrhotite product and troilite product, which is marked by the shift of the ~43.2°2θ to ~43.8°2θ peaks in XRD; we also offered direct observations of the mineral grains under electron microscope, depicting their morphology and the chemical reactions that may happen upon heating within the grains or the alumina crucible.

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Data Availability Statement—We provide XRD raw data and selected SEM-BSE images with qualitative spectral analyses as online Mendeley Data: https://doi.org/10.17632/6td5jtmwrn.1.

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