Supporting Information for
Impact of Semiconductor Band Tails and Band Filling on Photovoltaic Efficiency Limits

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Section S1. Optoelectronic Reciprocity Relations

The connection between absorption and emission has been known for quite some time. Kirchhoff in 1860\textsuperscript{1} is often cited as being the first to recognize the relation between the two processes, noting that the absorption and emission probability of a photon must be equal, i.e. $a(E) = e(E)$, through arguments of thermal equilibrium. A surface with $e(E) = 1$ for all energies is known as a perfect black body. However, the precise spectral dependence of a perfect black body emitter was not derived until Planck did so in 1906\textsuperscript{2}. He theorized a cavity with perfectly absorbing walls filled with a gas of photons with a small hole that would leak out a spectral flux characteristic of a black body:

$$I_{bb}(E) = E \times S_{bb}(E) = \frac{2\pi}{\hbar^3 c^2} \frac{E^3}{\exp(\frac{E}{kT}) - 1}$$  \hspace{1cm} (1)

which relates the temperature of a black body to its spectral characteristics, often referred to as thermal radiation. Here, $S_{bb}(E)$ is the energy-resolved photon flux per unit area per unit time of black-body radiation and $I_p$ refers to the spectrally resolved intensity of the radiation. The above expression can be generalized to non-black bodies by combining it with Kirchhoff’s law:

$$I(E) = a(E)I_{bb}(E)$$  \hspace{1cm} (2)

To form a general law for thermal radiation with surfaces characterized by an absorptivity. In an analogous manner, van Roosbroeck and Shockley\textsuperscript{3} generalized Planck’s law to semiconductors and related the absorption coefficient ($\alpha$) to the internal photon emission rate per unit volume:

$$R(E) = 4n_e^2(E)\alpha(E)S_{bb}(E) = \frac{8\pi n_e^2}{\hbar^3 c^2} \frac{E^2\alpha(E)}{\exp(\frac{E}{kT}) - 1}$$  \hspace{1cm} (3)

which holds for systems at thermal equilibrium. It was not until Lasher and Stern\textsuperscript{4} considered the situations of spontaneous emission were the above expressions further generalized to include non-equilibrium, steady-state conditions in terms of the quasi-Fermi level splitting $\Delta\mu$, which is exactly equal to the chemical potential of the photon in a spontaneous emission process:

$$R(E) = \frac{8\pi n_e^2}{\hbar^3 c^2} \frac{E^2\alpha(E)}{\exp(\frac{E - \Delta\mu}{kT}) - 1}$$  \hspace{1cm} (4)

Here, we note that that this expression is valid only when quasi-thermal equilibrium holds, where exactly two different quasi-Fermi levels accurately describe the energy dependence of the two separate populations of electrons and holes (e.g. after the electron-electron interactions subsequent to the excitation of carriers, the carriers will be distributed according to the Fermi-Dirac distribution), resulting in a single quasi-Fermi level splitting $\Delta\mu$. We assume this to be true in the case of the carriers in the band tails described here with the carriers above the respective band edges. For example, in the case of band tails caused by some ensemble of defects, an impurity band may be formed. If this impurity band is several $kT$ away from the band edges, the electrons...
in this band would likely thermalize amongst themselves, forming a separate quasi-Fermi level. Therefore, these relations would need to be modified to include this effect.

Wurfel then generalized the Lasher-Stern relation to an external flux of radiative emission from a semiconductor surface:

\[ S_{PL}(E) = a(E)S_{bb}(E, \Delta \mu) \]  

(5)

where

\[ S_{bb}(E, \Delta \mu) = \frac{2\pi}{h^3 c^2} \frac{E^2}{\exp\left(\frac{E - \Delta \mu}{kT}\right) - 1} \]  

(6)

is the spectral flux of a photon gas with chemical potential \( \Delta \mu \) and temperature \( T \). Here, \( E \) is the energy of the emitted photon, \( k \) is the Boltzmann constant, \( h \) is Planck’s constant, and \( c \) is the speed of light. Wurcfel’s expression, with a relation that connects the absorbance to the absorption coefficient (e.g. through the Beer-Lambert law of \( a(E) = 1 - \exp(-\alpha L) \) or more complex light-trapping geometries) suggests a complete set of self-consistent expressions that connect external properties (e.g. absorbance, external luminescence) of the semiconductor to its internal properties (e.g. bandgap, absorption coefficient, quasi-fermi level splitting, internal luminescence). External properties are therefore geometry dependent and can be carefully engineered from the internal properties using photonic design. Moreover, external properties are typically the only properties that are experimentally accessible.

We note that the above expression has an apparent divergence at \( E = \Delta \mu \). The resolution requires including an occupation factor in the absorption coefficient:

\[ a(E) = a_0 K(E) (f_v - f_c) \]  

(7)

Where \( f_v \) and \( f_c \) are the occupation for the holes and electrons, respectively. In the case of a semiconductor with equal effective mass for the holes and electrons and described by a parabolic dispersion, the occupation factor has a simple form:

\[ f_v - f_c = \tanh\left(\frac{E - \Delta \mu}{4kT}\right) \]  

(8)

While real systems may have more complex occupation factors (typically not representable analytically due to a fairly complex band structure), we note that \( f_v - f_c \) is generally a function with limiting values from -1 to 1 with a value of zero at \( E = \Delta \mu \), which is captured by the simple expression above. For simplicity and to capture the physics of the band filling irrespective of other materials properties, we use the simple expression above when calculating band filling effects.

It was suggested more recently by Rau that the principle of optical reciprocity can be further generalized to an optoelectronic reciprocity by including the serial collection/injection with Donolato’s theorem to describe photovoltaic cells and LEDs:

\[ S_{EL}(E) = EQ(E)S_{bb}(E, \Delta \mu) \]  

(9)
Where \( EQE(E) = a(E) \times IQE(E) \) and describes the process of absorbing a photon with probability \( a(E) \) with a subsequent collection probability of \( IQE(E) \). Thus, the LED quantum efficiency \( Q_{LED}(E) = \eta_{inj}(E) \times e(E) \) is a detailed balance pair with the photovoltaic quantum efficiency, taking the injection and collection efficiencies to be detailed balance pairs. We note that while the above generalized Planck’s law (Wurfel’s expression) holds quite generally by any system that can be characterized by two distinct quasi-Fermi levels and a thermodynamic temperature, Rau’s reciprocity relation strictly holds only in systems where carrier transport under illumination is well modelled as a linear perturbation of thermal equilibrium (qualitatively, the law of superposition in the current-voltage curves needs to hold). We also note that previous examples of using optoelectronic reciprocity for photovoltaic analysis (e.g. modified detailed balance models) has often approximated the black-body flux as:

\[
S_{bb}(E, \Delta \mu) \approx S_{bb}(E, 0) \exp\left(\frac{\mu}{kT}\right) = \left(\frac{2\pi}{\hbar^3c^2}\right) \exp\left(\frac{E^2}{2\hbar^2c^2kT} - 1\right) \exp\left(\frac{\mu}{kT}\right)
\]  

(10)

While the above expression has no singularities and generally results in numerically accurate results for most systems of interest (e.g. idealistic systems with \( a(E \leq E_g) = 0 \) will generally have \( \frac{E - \mu}{kT} \gg 1 \)), De Vos and Pauwel\'s noted the subtle differences this approximation has in analyzing entropy generation in the detailed balance limit. We show in this paper that accounting for band filling effects has qualitative and quantitative differences on the luminescence spectra of semiconductors with significant band tailing, which we emphasize in Figure S1 with a-Si:H as an example. Therefore, we use the full expression above without any approximations.

**Section S2. Modified Detailed Balance Limit Calculations**

With the above expressions of optoelectronic reciprocity in hand, we can assemble a modified detailed balance model for solar cells that account for carrier generation, extraction, and recombination:

\[
\int EQE(E, \Delta \mu)S(E)dE = \frac{\int EQE(E, \Delta \mu)S_{bb}(E, \Delta \mu)dE}{\eta_{ext}(\Delta \mu)} + \frac{J(\Delta \mu)}{q}
\]  

(11)

Where the left-hand side describes carrier injection (e.g. from sunlight or other light source) and the right-hand side describes carrier extraction, either through radiative recombination, non-radiative recombination (parametrized by \( \eta_{ext}(\Delta \mu) \)), or usefully as carrier collection \( (J(\Delta \mu)/q) \). In steady state, these populations must be balanced. In our analysis in the main text, we consider the modified detailed balance expression in the radiative limit i.e. \( \eta_{ext} = 1 \), \( \Delta \mu = qV \), \( EQE(E, \Delta \mu) = a(E, \Delta \mu) \) (see Section S8 for a short analysis on non-unity radiative or collection efficiency), with absorptivity described by a Beer-Lambert expression

\[
a(E) = 1 - \exp(-2aL)
\]  

(12)

With a perfect back reflector and perfect anti-reflection coating to describe the optical configuration. To parametrize the band edge density of states, we take inspiration from Katahara
and Hillhouse\textsuperscript{9} and convolve a sub-gap exponential density of states with a parabolic density of states above the bandgap, giving:

\[ \alpha_{0K}(E) = \alpha_0 \sqrt{\frac{\gamma}{kT}} G \left( \frac{E - E_g}{\gamma} \right) \]  
(13)

With

\[ G(x) = \text{real} \left( \frac{1}{2\Gamma(1 + \frac{1}{\theta})} \int_{-\infty}^{\infty} \exp(-|x'|^{\theta}) \sqrt{x - x'} dx' \right) \]  
(14)

And the simplified expression above (Eqn. 8) to account for band filling. Here, \( \gamma \) is the energy width parameter (i.e. the Urbach parameter, for \( \theta = 1 \)). \( E_g \) is the bandgap, \( \Gamma \) is the Gamma function, \( \alpha_0 \) scales the absorption coefficient (i.e. \( \alpha(E = E_g) = \alpha_0 \sqrt{\pi \gamma/16kT} \)), and \( \theta \) describes the power of the sub-gap exponential distribution. Our expression has an extra factor of \( \sqrt{kT} \) compared to the Katahara model, where \( kT \) is the thermal energy, so that \( \alpha_0 \) has the usual units of absorption coefficient. Using a simple piecewise continuous function for the absorption coefficient above and below the gap yields similar results, as long as the absorption coefficient below the gap is still modeled as an Urbach tail. Thus, for a given set of materials parameters (e.g. \( \alpha_0 L, \gamma, E_g \)) and a specific voltage \( V = \Delta \mu \), we can calculate the appropriate absorption coefficient and consequently the absorption and luminescence characteristics. The current-voltage curve of the photovoltaic cell in the detailed balance limit is then calculated using Equation 11. Specific figures of merit can then be extracted from the current-voltage curves.

**Section S3. Band filling Contribution to Photoluminescence**

In general, we are interested in the contribution of including the band filling on the luminescence spectrum of a semiconductor with significant band tails. Let us examine the case where we are weakly absorbing, which is generally true in the spectral region of a band tail. In this limit, we can take \( a \approx 2\alpha L \), where we assume a planar system with a perfect mirror and a perfect antireflection coating as above. In this case, the external luminescence flux by reciprocity becomes

\[ S_{PL}(E, \Delta \mu) = a(E, \Delta \mu)S_{bb}(E, \Delta \mu) \approx 2\alpha(E, \Delta \mu)Ls_{bb}(E, \Delta \mu) \]  
(15)

For systems with intrinsic doping and equal effective masses, we have \( \alpha(E, \Delta \mu) = \alpha(E, 0)(f_v - f_c) = \alpha(E, 0) \tanh \left( \frac{E - \Delta \mu}{4kT} \right) \). To see this, note that generally speaking,

\[ f_v - f_c = \frac{1}{\exp \left( \frac{E_r - E_{fp}}{kT} \right) + 1} - \frac{1}{\exp \left( \frac{E_{fn} - E_e}{kT} \right) + 1} \]  
(16)

And for intrinsic doping and equal effective masses, \( E_{fp} - E_i = -\frac{\Delta \mu}{2} \) and \( E_{fn} - E_i = \frac{\Delta \mu}{2} \) by symmetry arguments. Here, \( E_{fp/n} \) is the quasi-Fermi level for the holes/electrons, \( E_i \) is the Fermi level of the intrinsic system (at mid-gap), and \( \Delta \mu = E_{fn} - E_{fp} \) is the quasi-Fermi level splitting.
By symmetry of the electron and hole in this case, we must have $E_e - E_i = \frac{E}{2}$ and $E_h - E_i = -\frac{E}{2}$, where $E$ is the energy of the photon. Thus,

$$f_v - f_c = \frac{1}{\exp\left(-\frac{E - \Delta \mu}{2kT}\right) + 1} - \frac{1}{\exp\left(\frac{E - \Delta \mu}{2kT}\right) + 1}$$  \hspace{1cm} (17)

For simplicity in analysis, let us set $x = \frac{E - \Delta \mu}{kT}$. Thus,

$$f_v - f_c = \frac{1}{e^{-\frac{x}{4}}(e^{-\frac{x}{4}} + e^{\frac{x}{4}})} - \frac{1}{e^{\frac{x}{4}}(e^{\frac{x}{4}} + e^{-\frac{x}{4}})} = \frac{\exp\left(\frac{x}{4}\right) + \exp\left(-\frac{x}{4}\right)}{\exp\left(\frac{x}{4}\right) + \exp\left(-\frac{x}{4}\right)^2}$$

$$= \frac{\exp\left(\frac{x}{4}\right) - \exp\left(-\frac{x}{4}\right)}{\exp\left(\frac{x}{4}\right) + \exp\left(-\frac{x}{4}\right)} = \frac{\sinh\left(\frac{x}{4}\right)}{\cosh\left(\frac{x}{4}\right)} = \tanh\left(\frac{x}{4}\right) = \tanh\left(\frac{E - \Delta \mu}{4kT}\right)$$

We have argued already above that $\tanh\left(\frac{E - \Delta \mu}{4kT}\right)$ should serve as a good approximation to $f_v - f_c$ for most systems and should capture the main physics of band filling. It may be modified to yield more accurate results in the case of high doping or a large mismatch between the electron and hole effective masses under the parabolic bands approximation. For the purposes of this work, let us proceed with the simple expression so that the luminescence becomes

$$S_{PL}(E, \Delta \mu) = \left(\frac{4\pi L}{h^3 c^2} \alpha(E, 0)E^2\right) \left(\frac{\tanh\left(\frac{E - \Delta \mu}{4kT}\right)}{\exp\left(\frac{E - \Delta \mu}{kT}\right) - 1}\right)$$  \hspace{1cm} (18)

Where the term on the left is a sole function of $E$ and the term on the right includes both $E$ and $\Delta \mu$. Note that by taking $x = \frac{E - \Delta \mu}{kT}$, we have

$$\frac{\tanh\left(\frac{x}{4}\right)}{\exp(x) - 1} = \frac{\sinh\left(\frac{x}{4}\right)}{\cosh\left(\frac{x}{4}\right)(\exp(x) - 1)}$$

$$= \frac{1}{\exp(x) - 1}\left[\exp\left(\frac{x}{4}\right) - \exp\left(-\frac{x}{4}\right)\right] / \left[\exp\left(\frac{x}{4}\right) + \exp\left(-\frac{x}{4}\right)\right]$$

$$= \frac{1}{\exp(x) - 1}\left[\exp\left(\frac{x}{2}\right) - 1\right] / \left[\exp\left(\frac{x}{2}\right) + 1\right]$$

Finally, we use that $\exp(x) - 1 = \left[\exp\left(\frac{x}{2}\right) - 1\right]\left[\exp\left(\frac{x}{2}\right) + 1\right]$, so that

$$\frac{\tanh\left(\frac{x}{4}\right)}{\exp(x) - 1} = \frac{1}{\left(\exp\left(\frac{x}{2}\right) + 1\right)^2}$$
Let us double check that there are no singularities as $x \to 0$, since $\frac{\tanh(0)}{\exp(0) - 1} = \frac{0}{0}$. To do so, we shall use L’Hôpital’s rule, i.e.\[
\lim_{x \to c} \frac{f(x)}{g(x)} = \lim_{x \to c} \frac{f'(x)}{g'(x)}
\]

With $f(x) = \tanh\left(\frac{x}{4}\right)$ and $g(x) = \exp(x) - 1$, giving $f'(x) = \frac{1}{4} \sech^2\left(\frac{x}{4}\right)$ and $g'(x) = \exp(x)$. Thus, $\lim_{x \to 0} \frac{\tanh\left(\frac{x}{4}\right)}{\exp(x) - 1} = \frac{1}{4}$, so that there are no singularities and the luminescence can be rewritten as

\[
S_{PL}(E, \Delta \mu) = \frac{4\pi L}{\hbar^3 c^2} \frac{\alpha(E, 0) E^2}{\left(\exp\left(\frac{E - \Delta \mu}{2kT}\right) + 1\right)^2}
\]

(19)

Which is positive definite and is a good approximation for the luminescence with significant band tailing while explicitly including the band filling effects. Note that when $\frac{E - \Delta \mu}{kT} \gg 1$, we have

\[
S_{PL}(E, \Delta \mu) \approx \frac{4\pi L}{\hbar^3 c^2} \alpha(E, 0) E^2 \exp\left(-\frac{E}{kT}\right) \exp\left(\frac{\Delta \mu}{kT}\right)
\]

(20)

which recovers the expression without band filling contribution, suitable for low injection and sharp band edges and has been the standard expression used in most detailed balance analyses of solar cells. It is clear from Equation 19 that the luminescence spectra and radiative current will scale non-linearly with $\Delta \mu$. Furthermore, for $\alpha(E, 0) \sim \exp\left(\frac{E - E_g}{\gamma}\right)$, as in the case of Urbach tails, we can take a derivative of the luminescence flux and find that the peak position will occur at

\[
E_{PL}^{\text{max}} = \Delta \mu - 2kT \ln \left(1 + \frac{1}{\exp\left(\frac{E_{PL}^{\text{max}}}{kT}\right) - 1}\right)
\]

(21)

For $\gamma > kT$. A simpler but approximate solution can be found by taking $E_{PL}^{\text{max}} \gg kT$, and neglecting that term, so that

\[
E_{PL}^{\text{max}} \approx \Delta \mu + 2kT \ln \left(\frac{kT}{\gamma - kT}\right)
\]

(22)

Which shows that the luminescence peak depends directly on $\Delta \mu$, for $\gamma > kT$

**Section S4. Effects of band tails on $J - V$ characteristics**

While Equation 19 suggests a rather complex dependence of the band filling characteristics on current, we find that the $J - V$ characteristics can be well fitted to a modified diode expression in most cases:
\[
J_{\text{rad}}(V) \sim J_0(\gamma, E_g) \exp\left(\frac{qV}{n_{\text{eff}}(\gamma, E_g)kT}\right)
\]  

(23)

In other words, the effect of band filling and band tails is to modify the recombination current prefactor \(J_0\) and effective ideality factor \(n_{\text{eff}}\), which manifest in the voltage loss as described in the main text and in section S6. Of particular interest is \(n_{\text{eff}}\), which should be measurable in electroluminescence measurements, because non-radiative dark current occurs in parallel to the radiative dark current. Thus, we would expect the \(n_{\text{eff}}\) estimated here in Figure S5 to be accurate even in systems far away from the radiative limit, as long as we measure the radiative current flux through voltage-dependent electroluminescence. We note that the calculated \(n_{\text{eff}}\) for a-Si (assuming \(E_g \sim 1.7\) eV and \(\gamma \sim 50\) meV) is around 1.7, which is quite similar to the value measured by Rau et al\(^{10}\). To get an approximate analytic expression for \(n_{\text{eff}}\), we use Eqn. 19 and assume that \(E^2\) varies slowly compared to the exponentials in the integrand and that we are in the weakly absorbing limit. Thus, \(J_{\text{rad}}(V) \sim \int dE \exp\left(\frac{E-E_g}{\gamma}\right)\left(\exp\left(\frac{E-V}{2kT}\right) + 1\right)^{-2}\) and with some rewriting, we find that \(J_{\text{rad}}(V) \sim \int dx \exp\left(\frac{kTx+E_g}{\gamma}\right)\left(\exp\left(\frac{x}{2}\right) + 1\right)^{-2} \sim \exp\left(\frac{V}{kT}\right)\). That is, we expect

\[
n_{\text{eff}} \approx \frac{\gamma}{kT}
\]  

(24)

Which seems to hold somewhat well for small \(\gamma\) just above \(kT\), as observed in Figure S5. Furthermore, using the diode approximation from above, we can also calculate the modified fill factor expression as

\[
FF(n_{\text{eff}}, V_{\text{oc}}) \approx \frac{qV_{\text{oc}}}{n_{\text{eff}}kT} - \ln\left(1 + \frac{qV_{\text{oc}}}{n_{\text{eff}}kT}\right)
\]  

(25)

which reduces the fill factor slightly compared to the case without band tails and is an additional efficiency loss mechanism.

**Section S5. Two bandgap model for band tails**

To develop a simple picture for the apparent bandgap shift, voltage loss, and effects of band tailing, we use a simplistic model of the absorbance parametrized by two step functions. We will refer to this as the “two bandgap model”, whose absorbance can be seen in Figure S7(a) and is given by:

\[
A(E) = a_1 \theta(E - E_{g,1})\theta(E_{g,2} - E) + a_2 \theta(E - E_{g,2})
\]  

(26)

Where \(a_2 = 1\) and \(E_{g,2} = 1.34\) eV. The above model represents a simplistic picture of a system with band tails as it deviates from the Shockley-Queisser limit. Here, \(E_{g,2}\) defines the absorption bandgap, \(E_{g,1}\) is the lower bandgap that forms as a result of band tailing, and \(a_1\) is the effective sub-gap absorption. We then calculate the typical photovoltaic figures of merit in Figure S7(b) while varying \(\Delta E_g = E_{g,2} - E_{g,1}\) and \(a_1\). The result is qualitatively similar to what is seen with a band tail (e.g. see Figure S2 for comparison), where the efficiency loss is essentially all in the
voltage. Moreover, there is a specific transition point where the voltage loss is linear with the bandgap separation, dependent on the value of $a_1$. To see this, recall that $V_{oc} = \frac{kT}{q} \ln \left( \frac{J_{sc}}{J_0} + 1 \right)$, where $J_0 = \int A(E) S_{BB}(E) dE$ and $A(E)$ is given in Eqn. 26. The loss due to a lower bandgap $E_{g,1}$ is then

$$\Delta V_{oc} = \frac{kT}{q} \ln \left( \frac{J_{sc}}{J_2} + 1 \right) - \frac{kT}{q} \ln \left( \frac{J_{sc}}{J_0} + 1 \right) \approx - \frac{kT}{q} \ln \left( \frac{J_1}{J_2} + 1 \right)$$

(27)

Where $J_{1,2} = \int A_{1,2}(E) S_{BB}(E) dE$. $A_{1,2}(E) = a_{1,2} \theta(E - E_{g,1,2})$, and we have assumed $J_{sc} \gg J_0 J_2$. Thus, from the perspective of the voltage loss in the detailed balance analysis, $E_{g,1}$ does not appear as a photovoltaic bandgap until $J_1 > J_2$. This occurs when

$$\frac{a_1 e^{-\frac{E_{g,1}}{kT}} \left( \left( \frac{E_{g,1}}{kT} \right)^2 + 2 \left( \frac{E_{g,1}}{kT} \right) + 2 \right)}{a_2 e^{-\frac{E_{g,2}}{kT}} \left( \left( \frac{E_{g,2}}{kT} \right)^2 + 2 \left( \frac{E_{g,2}}{kT} \right) + 2 \right)} > 1$$

(28)

Assuming $E_{g,1,2} \gg kT$, we can neglect the terms outside of the exponential to first order because it shows up logarithmically with $\Delta E_g$. Thus, the transition to a new bandgap occurs when

$$\Delta E_g > kT \ln \left( \frac{a_2}{a_1} \right)$$

(29)

In other words, from the perspective of the Shockley-Queisser limit and voltage loss, the Stokes shift is not apparent until Eqn. 29 is satisfied. At this point, the voltage loss scales linearly with increasing $E_{g,1}$. To see this clearly, we plot the voltage loss with bandgap shift with energies and voltages normalized to $kT \ln \left( \frac{a_2}{a_1} \right)$ in Figure S7(c). We see that indeed the transition occurs under the condition of Eqn. 29, where thereafter $\frac{\partial \Delta V_{oc}}{\partial \Delta E_g} \approx 1$. This is true irrespective of the value of $a_1$.

Moreover, while Eqn. 29 is derived for two discrete bandgaps, we can generalize the concept to how sharp a continuous absorption spectrum should be to avoid a Stokes shifted voltage loss. Let us define $a_2 = \Delta a + a_1$ and take the limit as $\Delta a, \Delta E \to 0$. Thus, the generalized continuous form of Eqn. 29 becomes

$$\frac{kT}{a} \frac{\partial a}{\partial E} < 1$$

(30)

In the case of weakly absorbing Urbach band tails, $a \sim aL \sim C \exp \left( \frac{E - E_g}{\gamma} \right)$. Thus, Eqn. 30 predicts a Stokes shift should occur when $\gamma > kT$, which is what we observe in Figure 2 of the manuscript.

**Section S6. General Expression for Voltage Loss due to Nonabrupt Band Edges**

The plots of Figure 3 and Figure S7 in the main text suggests a general relation between bandgap shifts and voltage loss, irrespective of the exact functional form of the band edge. To see this, note that the majority of the luminescence of the step-function absorbance is concentrated within $kT$ of
the band edge and its integral varies exponentially with the bandgap energy. Thus, the effective bandgap of the luminescence, $E_{g,PL}$ must be chosen to integrate to nearly the majority of the luminescence flux. Thus, we pragmatically define it as

$$\max(E_{g,PL}) \geq \frac{\int_{E_{g,PL}}^{\infty} S_{PL}(E, \Delta \mu) dE}{\int_{0}^{\infty} S_{PL}(E, \Delta \mu) dE} \geq 0.90$$  \hspace{1cm} (31)$$

While this definition of $E_{g,PL}$ is not unique, it parametrizes the luminescence typically assumed under step-function absorbance to a greater variety of luminescence spectra and is somewhat less sensitive to noise. We further define the above-gap absorbance as

$$\bar{a}_{AG} = \frac{\int_{E_{g,PL}}^{\infty} a(E, \Delta \mu) bb(E, \Delta \mu) dE}{\exp(\Delta \mu / kT) \int_{E_{g,PL}}^{\infty} bb(E, 0) dE} = \frac{\int_{E_{g,PL}}^{\infty} S_{PL}(E, \Delta \mu) dE}{\exp(\Delta \mu / kT) \int_{E_{g,PL}}^{\infty} bb(E, 0) dE}$$  \hspace{1cm} (32)$$

And below-gap absorbance as

$$\bar{a}_{SG} = \frac{\int_{E_{g,PL}}^{E_{g,PL}} a(E, \Delta \mu) bb(E, \Delta \mu) dE}{\exp(\Delta \mu / kT) \int_{E_{g,PL}}^{E_{g,PL}} bb(E, 0) dE} = \frac{\int_{E_{g,PL}}^{E_{g,PL}} S_{PL}(E, \Delta \mu) dE}{\exp(\Delta \mu / kT) \int_{E_{g,PL}}^{E_{g,PL}} bb(E, 0) dE}$$  \hspace{1cm} (33)$$

Where both values are apparently dependent on $\Delta \mu$. Accurate estimation of the quantity $S_{PL}(E, \Delta \mu) / \exp(\Delta \mu / kT)$ can be achieved by taking $E \gg \Delta \mu$ and fitting the luminescence spectra to the high energy absorption/EQE, or by fitting the full spectrum with the band filling factor. Alternatively, since Eqn. 6 of the main text only requires knowledge of the ratio $\bar{a}_{AG} / \bar{a}_{SG}$, we can simply use the directly measured luminescence spectrum:

$$\frac{\bar{a}_{AG}}{\bar{a}_{SG}} = \frac{\int_{E_{g,PL}}^{\infty} S_{PL}(E, \Delta \mu) dE}{\int_{E_{g,PL}}^{\infty} S_{PL}(E, \Delta \mu) dE} \frac{\int_{E_{g,PL}}^{E_{g,PL}} bb(E, 0) dE}{\int_{E_{g,PL}}^{E_{g,PL}} bb(E, 0) dE}$$  \hspace{1cm} (34)$$

And the definitions of $E_{g,Abs}$ and $E_{g,PL}$ to estimate the weighted absorbance ratio.

These definitions work well because the integrated number of recombination electrons is what matters in the detailed balance analysis, which is achieved by the appropriate definitions of weighted absorption and bandgaps. Therefore, the voltage loss is given by a form that is quite similar to Eqn. 27:

$$\Delta V_{oc} = \frac{kT}{q} \ln \left( \frac{\bar{a}_{SG}}{\bar{a}_{AG}} \exp \left( \frac{E_{g,PL} - E_{g,PL}}{kT} \right) \left[ \left( \frac{E_{g,PL}}{kT} \right)^2 + \frac{2E_{g,PL}}{kT} + 2 \right] + 1 - \frac{\bar{a}_{SG}}{\bar{a}_{AG}} \right)$$  \hspace{1cm} (34)$$
Noting the logarithmic dependence on the argument and assuming $\Delta E_g = E_{g,\text{Abs}} - E_{g,\text{PL}} \ll E_{g,\text{Abs}}$, as well as $E_{g,\text{Abs}}, E_{g,\text{PL}} \gg kT$, we arrive at a simple expression that only depends on the observed bandgap shifts and the ratio of the above-gap and sub-gap absorbances:

$$\Delta V_{oc} \left( \frac{\tilde{a}_{SG}}{\tilde{a}_{AG}}, \Delta E_g \right) \approx \frac{kT}{q} \ln \left( \frac{\tilde{a}_{SG}}{\tilde{a}_{AG}} \exp \left( \frac{\Delta E_g}{kT} \right) + 1 - \frac{\tilde{a}_{SG}}{\tilde{a}_{AG}} \right)$$

(35)

Note that this expression recovers the expected values of voltage loss as $\frac{\tilde{a}_{SG}}{\tilde{a}_{AG}} \to 0, 1$ and as $\Delta E_g \to 0$. Furthermore, the functional form of the sub-gap absorbance is captured by its effect of varying the value of $\tilde{a}_{SG}$. From an experimental standpoint, another method to estimate the voltage loss is by using the modified $I - V$ characteristics found in Section S4. It is clear then that $J_{rad}(V) \approx J_{0,rad} \exp \left( \frac{qV}{n_{eff}kT} \right) = \int S_{EL}(E,V)dE$. Furthermore, it is possible to estimate $n_{eff}$ directly from the slope of voltage-dependent electroluminescence $S_{EL}(E,V)$. Integrating over $S_{EL}(E,V)$ and dividing by $\exp \left( \frac{qV}{n_{eff}kT} \right)$ then yields $J_{0,rad}$. Note that the $V_{oc}$ loss due to an imperfect band edge can be equivalently written in the form of $\Delta V_{oc} = \frac{kT}{q} \ln \left( \frac{J_{sc,\text{SQ}}}{J_{sc}} \right) - \frac{n_{eff}kT}{q} \ln \left( \frac{J_{sc}}{J_{0,rad}} \right)$ using Eqn. 23, which can be expanded to yield

$$\Delta V_{oc} = \frac{kT}{q} \ln \left( \frac{J_{sc,\text{SQ}}}{J_{sc}} \right) + \frac{kT}{q} \ln \left( \frac{J_{0,rad}}{J_{0,rad,\text{SQ}}} \right) - \frac{(n_{eff} - 1)kT}{q} \ln \left( \frac{J_{sc}}{J_{0,rad}} \right)$$

(36)

where the first term is the voltage loss due to incomplete absorption above the bandgap. The second term is the voltage loss due to band tailing, while the third term is a voltage gain due to band filling effects (e.g. see Fig. S4).

**Section S7. Effects of Sub-Unity Radiative and Quantum Efficiencies**

We have thus far only analyzed the situation assuming the reciprocity between absorption and photoluminescence, which holds quite generally but concerns primarily the internal open circuit voltage of a device i.e. the quasi Fermi level splitting. To analyze the effects of a system with sub-unity quantum efficiencies, which may be particularly relevant for localized states below the absorption gap, we assume Donolato’s theorem still holds and apply Eqn. 9. Therefore, by reciprocity, the injection efficiency into these localized states would be relatively low, lowering the electroluminescence recombination rate and increasing the limiting $V_{oc}$ (see Figure S8). This situation would be analogous to considering free carrier absorption in the absorption band tail, where $IQE \to 0$, and therefore the absorption of free-carriers do not lead to photovoltaic current$^{11}$. Thus, photogenerated carriers that do not contribute to photovoltaic current, whether they are localized states or free carriers, would not result in a loss to the open circuit voltage in the radiative limit. In general, the effect of band tails on the radiative limit should be determined via photocurrent spectroscopies, which captures this effect experimentally directly.

To analyze the voltage loss effects away from the radiative limit, i.e., sub-unity radiative efficiency, we note that generally Eqn. 23 holds and the discussion in section S4 suggests that $J(V) = J_{sc} - \frac{J_{0,rad}}{n_{ext}} \left( \exp \left( \frac{qV}{n_{eff}kT} \right) - 1 \right)$ which is quite similar to Eqn. 11. Thus, it is readily
It should be noted that $\eta_{ext}$ is generally a function of voltage as well and should be measured/calculated at the operating voltage. This radiative ideality factor has already been recognized by Rau et al. to be relevant in amorphous Si$^{10}$ when analyzing its non-radiative losses. In many devices, sub-unity radiative efficiencies and sub-unity quantum efficiencies are both present and are likely competing to provide the observed voltage. In contrast, concentration benefits the voltage by a similar factor $\Delta V_{oc,conc} = \frac{n_{eff}kT}{q}|\ln(C)|$, where $C > 1$ is the concentration factor.

**Section S8. Parametrization of the Band edge Functional Form**

In Figure 3 of the main text, we considered various band edge functional forms to argue that there exists a general expression that relates $\Delta V_{oc,rad}$ to the existence of a Stokes shift, i.e. $\Delta E_g$. We considered two main types of band edges: exponential tails and indirect edge power laws. Exponential tails are the main form of band edges we have discussed in this article and we have thus far used the analysis described in Section S2. For the calculations in Figure S9 and Figure 3 of the main text, we vary $\theta$, $E_g$, and $\gamma$ to generate various functional forms for the band tail given by Eqn. 13. Furthermore, we consider only the absorption spectra that yield a luminescence bandgap above $4kT$, since we assume $E_g \gg kT$ in deriving Eqn. 6 of the main text. We further consider a general power law form for a semiconductor band edge that has a weak oscillator strength (e.g. an indirect transition) with a higher energy direct transition with larger oscillator strengths:

$$\alpha(E) = \alpha_{0,ind}(E - E_{g,ind})^n \theta(E - E_{g,ind})\theta(E_{g,dir} - E) + \alpha_{0,dir}\theta(E - E_{g,dir})$$

(39)

Where $E_{g,ind}$ and $E_{g,dir}$ represent the indirect and direct band edge, respectively, while $\alpha_{0,ind}$ and $\alpha_{0,dir}$ represent the absorption coefficients of the indirect and direct gaps, respectively. $n$ parametrizes the different energetic scaling relations of the indirect edge, typically $n < 3$ experimentally.

For both forms of band edges, we calculate $V_{oc,rad}$ from the complete modified detailed balance analysis, including band filling effects and assuming $\eta_{ext} = 1$ (Eqn. 11), $E_g,Abs$ is then derived from the calculated absorption spectrum using Rau’s definition, and therefore $V_{oc,SG}(E_g,Abs)$ is calculated using a step-function at $E_g,Abs$. $E_g,PL$, $\bar{\alpha}_{SG}$, $\bar{\alpha}_{AG}$ is then calculated from the definitions in Section S6 by examining the luminescence spectra, $S_{PL}$. The results of these different band edges map well onto a simple relation described by Eqn. 35, suggesting a two bandgap model is an adequate representation of most experimentally observed band edge forms.
Figure S1. **Accounting for band filling in modified reciprocity relations**: Experimentally measured a-Si:H EQE and EL from ref. 10 (open circles). Solid lines correspond to the Rau reciprocity relation, whereas the dashed line is a fit that includes band filling effects with asymmetric effective masses in the parabolic approximation \(m_h/m_e = 1.818, \Delta \mu = 1.164 \text{ V} \), and \(E_g = 2.439 \text{ eV} \). The dotted line includes band filling with the same fitted parameters except \(m_h/m_e = 1\), i.e. assumes symmetric effective masses. All spectra are normalized by \(\exp(\Delta \mu/kT)\) and the various reciprocity relations overlay for \(E > E_{g,\text{Abs}}\).
Figure S2. Dependence of photovoltaic figures of merit on the Urbach parameter: (a) The detailed balance limited value of conversion efficiency, open circuit voltage, short circuit current, and fill factor for different bandgaps and Urbach parameters assuming a thickness of $\alpha_0 L = 1$. (b) Linecuts of (a) at specific bandgap values.
Figure S3. Effects of thickness on photovoltaic figures of merit: (a) Absorbance of a photovoltaic cell plotted with different normalized thicknesses ($\alpha_0L$) for $\gamma = 0.5kT$ and (b) $\gamma = 2kT$ assuming a bandgap $E_g = 1.5$ eV. (c) Conversion efficiency, open circuit voltage, short circuit current, and fill factor calculated for different normalized thicknesses assuming a bandgap $E_g = 1.5$ eV. The different colored lines correspond to the same legend shown in (a).
Figure S4. **The importance of including band filling effects:** (a) Calculated absorbance (solid line), photoluminescence (dashed line), and distribution of bandgaps (dotted line) for different Urbach parameters ($\gamma$) without including band filling effects. (b) Calculated efficiency, open circuit voltage, short circuit current, and fill factor with (orange solid line) and without (blue solid line) including band filling effects.
Figure S5. **Effects of band tails and band filling on ideality factor and current-voltage relationships:** (a) Fitted $n_{eff}$ for varying Urbach parameter ($\gamma$) and bandgap $E_g$. Fits were performed for the range $3kT < qV < E_g - 3kT$. Linecuts of (a) occur at $E_g = 0.8$ (blue), 1.34 (orange), and 2.0 eV (yellow). (b) Corresponding linecuts of (a) plotted for varying Urbach parameter ($\gamma$). Note the transition that occurs at $\gamma = kT$ to larger effective ideality factors, corresponding to the onset of band tailing and band filling effects. Dashed lines represent the fit, while solid lines represent the 95% confidence interval. $J - V$ characteristics for different bandgaps of 0.8 eV (c), 1.34 eV (d), and 2.0 eV (e). The different lines in a given plot represent different Urbach parameters. The legend in (c) is the same for (d) and (e). All plots have voltages normalized to $kT/q$ and current densities normalized to their radiative dark current $J_0$, which is a function of $\gamma$. Thicknesses were assumed to be $\alpha_0 L = 1$. Note that for Urbach parameters typically observed in experiment (i.e. $\gamma < \sim 3kT$), $n_{eff}$ is generally less than 3. For larger Urbach parameters, a modified ideality factor no longer describes the voltage scaling appropriately, since $E_{g,PL} \rightarrow kT$. 
Figure S6. **Analysis of a Gaussian band tail distribution**: (a) Calculated absorbance (solid line), photoluminescence (dashed line) and distribution of bandgaps (dotted line) for an increasing Gaussian tail ($\gamma$). Here, the Gaussian tail distribution is calculated by taking $\theta = 2$ in Eqn. 14. (b) Fraction of integrated photoluminescence below the band gap (solid blue line) and Stokes shift $\Delta E_g$ (solid orange line) for a Gaussian tail distribution. (c) Calculated detailed balance efficiency for different bandgaps plotted for increasing Gaussian tail widths. The different colored lines correspond to the same values of the Gaussian tail displayed in (a).
Figure S7. **Analysis of a two-bandgap toy model**: (a) Absorbance and emission of the two-bandgap toy model, parametrized by two step-functions. Solid lines correspond to absorbance, whereas dashed lines correspond to emission. (b) Plot of the photovoltaic figures of merit \(\eta, V_{oc}, J_{sc}\) for varying bandgap difference \(\Delta E_g = E_{g,2} - E_{g,1}\) and values of the lower bandgap absorbance \(a_1, a_2\) is assumed to be 1 while \(E_{g,2} = 1.34\) eV. Colors correspond to the same as the legend in (c). (c) Voltage loss versus bandgap difference in normalized units of \(kT \ln \left(\frac{a_2}{a_1}\right)\), showing the transition to the Stokes shift behavior for large enough bandgap separation, dependent on \(a_2/a_1\).
Figure S8. **Effects of a sub-unity collection efficiency below the bandgap**: Calculated power conversion efficiency, open circuit voltage, short circuit current density, and fill factor assuming that the collection efficiency below the bandgap ($IQE_{SG}$) is less than 1 and given by a constant average value. That is, we take the external quantum efficiency to be $EQE(E) = a(E) \left( IQE_{SG} \theta(E_g - E) + \theta(E - E_g) \right)$. The “Urbach” curve is calculated assuming the collection efficiency decays with a similar Urbach parameter to that used in the absorption calculation (i.e. $IQE_{SG}(\gamma, E) = \exp\left(\frac{E - E_g}{\gamma}\right)$), which may approximate the mobility-edge better than a constant.
Figure S9. **Different band edges that map onto a two-bandgap model**: Stokes shift $\Delta E_g$ and radiative voltage loss $\Delta V_{oc}$ calculated from the full detailed balance analysis with the appropriate definitions of $E_{g,\text{Abs}}, E_{g,\text{PL}}, a_{AG}, a_{SG}$, as described in Section S6. We vary the parameters for the exponential band tail model (a) and the indirect edge power law model (b). For the exponential band tail model we take $\alpha_0 L = 10$, whereas for the indirect edge model we take $\alpha_{0,\text{dir}} L = 100$, $\alpha_{0,\text{ind}} L = 0.1$. Both forms map well onto the generalized expression (c). The colorbar for the generalized expression in (c) is $\log_{10} \left( \frac{a_{SG}}{a_{AG}} \right)$, i.e. describes the ratio of the sub-gap to above-gap absorption. The different ratios plots are overlaid, showing the similarity irrespective of $a_{SG}/a_{AG}$, assuming it is sufficiently small.
| Material    | Bandgap (eV) | Urbach Energy (meV) | Calculated Δ𝑉 Ost (mV) | References        |
|-------------|--------------|----------------------|-------------------------|-------------------|
| c-Si        | 1.12         | 9.6                  | 24.0                    | Cody 1992<sup>12</sup> |
| c-Si        | 1.12         | 8.6                  | 20.7                    | Cody 1992<sup>12</sup> |
| c-Si        | 1.12         | 11                   | 28.9                    | Green 2008<sup>13</sup> |
| GaAs        | 1.42         | 6.9                  | 14.6                    | Sturge 1962<sup>14</sup> |
| GaAs        | 1.42         | 7.5                  | 16.3                    | Johnson 1995<sup>15</sup> |
| GaAs        | 1.42         | 5.9                  | 11.7                    | Beaudoin 1997<sup>16</sup> |
| InP         | 1.355        | 9.4                  | 22.5                    | Subashiev 2010<sup>17</sup> |
| InP         | 1.361        | 10.6                 | 26.6                    | Subashiev 2010<sup>17</sup> |
| InP         | 1.34         | 7.1                  | 16.3                    | Beaudoin 1997<sup>16</sup> |
| a-Si:H      | 1.72         | 42                   | 273.1                   | Cody 1984<sup>18</sup> |
| a-Si:H      | 1.64         | 52                   | 382.8                   | Cody 1984<sup>18</sup> |
| a-Si:H      | 1.69         | 67                   | 559.1                   | Cody 1981<sup>19</sup> |
| a-Si:H      | 1.7          | 43                   | 283.5                   | Tiedje 1981<sup>20</sup> |
| a-Si:H      | 1.69         | 47                   | 329.0                   | Tiedje 1983<sup>21</sup> |
| a-Si:H      | 1.7          | 48                   | 341.3                   | van Veen 2003<sup>22</sup> |
| a-Si:H      | 1.8          | 51                   | 385.3                   | van Veen 2003<sup>22</sup> |
| a-Si:H      | 1.85         | 51                   | 389.1                   | van Veen 2003<sup>22</sup> |
| CdTe        | 1.45         | 17                   | 52.8                    | Rakhshani 2001<sup>23</sup> |
| CdTe        | 1.5          | 7.2                  | 15.5                    | Marple 1966<sup>24</sup> |
| CdTe        | 1.5          | 9                    | 21.1                    | Mullins 1997<sup>25</sup> |
| CdTe        | 1.5          | 10.6                 | 26.5                    | Sculfort 1984<sup>26</sup> |
| CIGS        | 1.53         | 24                   | 94.2                    | Heath 2002<sup>27</sup> |
| CIGS        | 1            | 18                   | 56.0                    | Heath 2002<sup>27</sup> |
| CIGS        | 1.18         | 23                   | 84.9                    | Heath 2002<sup>27</sup> |
| CIGS        | 1.2          | 31                   | 143.6                   | Troviano 2011<sup>28</sup> |
| CIGS        | 1.67         | 25                   | 102.5                   | Meeder 2002<sup>29</sup> |
| CIGS        | 1.08         | 9                    | 21.9                    | Shioda 1996<sup>30</sup> |
| Kesterite   | 1.5          | 69                   | 551.4                   | Islam 2015<sup>31</sup> |
| Kesterite   | 1.1          | 54                   | 346.9                   | Islam 2015<sup>31</sup> |
| Kesterite   | 1.38         | 45                   | 286.6                   | Yan 2017<sup>32</sup> |
| Kesterite   | 1.54         | 65                   | 516.7                   | Yan 2017<sup>32</sup> |
| Kesterite   | 1.68         | 56.8                 | 441.8                   | Ng 2017<sup>33</sup> |
| Perovskite  | 1.57         | 15                   | 44.7                    | De Wolf 2014<sup>34</sup> |
| Perovskite  | 2.23         | 23                   | 90.2                    | Sadhanala 2014<sup>35</sup> |
| Perovskite  | 1.57         | 14                   | 40.5                    | Zhang 2015<sup>36</sup> |
| Perovskite  | 1.57         | 14.4                 | 42.2                    | Zhang 2015<sup>36</sup> |
| Perovskite  | 1.57         | 15.8                 | 48.3                    | Zhang 2015<sup>36</sup> |
| Organic     | 1.66         | 37                   | 214.9                   | Gotoh 1997<sup>37</sup> |
| Organic     | 2            | 50                   | 386.8                   | Kronemeijer 2014<sup>38</sup> |
| Organic     | 1.31         | 25.6                 | 104.8                   | Liu 2020<sup>39</sup> |

Table S1
Organic | 1.47 | 27 | 115.9 | Ran 2016
Organic | 1.88 | 36 | 211.0 | Vandewal 2014
Organic | 1.71 | 27 | 118.7 | Liu 2016
Organic | 1.67 | 24 | 95.2 | Qian 2018

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