LHC Bounds on UV-Complete Models of Dark Matter

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Abstract

We analyze the sensitivity of searches for dark matter in the jets and missing energy channel in the case where the particle mediating interactions between hadronic matter and DM is collider accessible. We consider all tree level UV completions of interactions between fermion DM and quarks which contribute to direct detection, and derive bounds which apply to elastic or inelastic scattering dark matter explanations of direct detection signals. We find that studies based on effective operators give robust bounds when the mediator is heavy enough to resonantly produce the final state in question.

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I. INTRODUCTION

The need for extra non-baryonic mass to explain galactic and supergalactic dynamics is well known from astrophysical observations [1]. Unfortunately, not much can be inferred about dark matter (DM) beyond its existence. Some of the most popular candidates for dark matter are weakly interacting massive particles or WIMPs, motivated by the WIMP(less) miracle [2]. WIMPs naturally occur in models postulated to solve the gauge hierarchy problem. Many theories have been found to address other issues in particle physics which also provide a viable dark matter candidate. Much effort has been put into understanding the predictions of these theoretically motivated models for experimental probes.

We have many methods to search for dark matter. These include astrophysical experiments which look for dark matter interacting with normal matter or dark matter annihilations, and collider searches for new particles which may make up the missing dark matter. However, understanding implications one experiment has on another has largely been restricted to placing limits, from one, on a favored theory and then deriving its predictions on the remaining parameter space to compare to other experiments [3–7]. While this leaves us with different regions which are “favored” by different models, it does not provide a good guide as to where we should be looking with future experiments in light of current results. Thus, it would be nice to establish a language for comparison between experiments.

Direct detection experiments search for dark matter particles recoiling off of nuclei as they travel through the Earth, while hadron colliders have the potential to produce dark matter from high energy hadronic material. Both types of experiment probe the couplings of dark matter to hadronic particles, and should complement each other, assuming the dark matter is light enough to be produced at colliders. Many recent direct detection experiments have reported interesting, though, conflicting results. These include, CoGeNT [8, 9], DAMA [10], Xenon [11] and CDMS [12] and, in fact, models to reconcile these discrepancies largely focus on the possibility of a light dark matter candidate [13–25].

In previous work [26–28], the overlap between direct detection experiments and colliders was considered in terms of effective four-point interactions between DM and hadronic particles, either quarks or gluons. In our current work, we drop the assumption that the intermediate state is heavy compared to the scales of interest and consider the simplest UV complete models of DM interactions which could give rise to signals in direct detection
experiments and hadron colliders. This was also considered for some special cases in [28].

This paper is organized as follows: In section II we discuss the relevant interactions between dark matter and quarks at the renormalizable level, in section III we derive bounds on those interactions based on null collider searches for monojets and missing energy, in section IV we discuss the impact on direct detection parameters, and in section V we conclude.

II. INTERACTIONS OF DM WITH QUARKS

In this article we work under the assumption that dark matter is composed of one species of Standard Model singlet Dirac fermion, $\chi$. Other candidates are possible, but the modification of previous results due to the introduction of collider-accessible mediators does not depend sensitively on the spin of the dark matter; it is dominated by phase space and kinematical effects.

We consider all tree-level renormalizable interactions of DM particles, $\chi$, with quarks which contribute to direct detection signals in the limit of zero momentum transfer. This generically requires positing a new mediator particle, $\phi$, whose quantum numbers depend on its particular interaction with quarks and $\chi$. We consider s-channel operators with a standard model singlet mediator and t-channel operators with colored mediators. The channel of a process is defined by the momentum combination which appears in the propagator of the mediating particle, either the sum of initial state momenta for s-channel or the difference of one initial and one final state momentum for t-channel. The mediators of t-channel models are color triplets with electric charge of 2/3 or -1/3, such that they couple to up- and down-type quarks, respectively and we will assume, for simplicity, that they have identical masses. We do not assign specific $SU(2)_L$ charges to mediators since QCD dynamics or direct detection are dependent on parity, rather than chirality. As electroweak effects are subdominant to QCD effects for these charged mediators, this is a good approximation for our study.

We assume that the dark matter abundance is protected by a $Z_2$ or larger symmetry which implies that t-channel mediators are also charged under this group. Table I shows the labels we assign to each model of dark matter interaction vertex; the first letter indicates the type of dark matter, in this case Dirac WIMPs, and the second letter indicates the DM annihilation channel, either s or t. These interaction vertices also contribute to indirect detection.
Similar interactions with leptons can also be considered\textsuperscript{31}. While they have interesting motivations in astrophysical observations, their couplings do not significantly contribute to either hadron collider signals or direct detection experiments and so we will not consider them here. It should be noted that one can also consider couplings to other Standard Model particle, e.g. \textsuperscript{32}, as well.

The s-channel models are assumed to be flavor diagonal in couplings and therefore are not significantly constrained by precision measurements. The t-channel models can, in principle, contribute to FCNC decays of hadrons at one loop order, but, we will assume that these decays are suppressed by some mechanism and not concern ourselves further with this. The couplings of DS\textsubscript{1} are scaled by the quark Yukawa coupling appropriate to ensure Minimal Flavor Violation\textsuperscript{33}.\textsuperscript{1} We thus have, in all posited models of DM interactions, three parameters, the DM mass, $M_\chi$, the mediator mass, $M_\phi$, and the coupling, $g$. In principle the couplings in s-channel models to DM and to quarks can be unrelated, but one can only bound the product of the two couplings at colliders and direct detection experiments, so we take them to be identical.

Integrating out the mediator $\phi$, leads to contact interactions which are suppressed by a scale $M_* \equiv M_\phi/g$, which relates to models considered in \textsuperscript{26, 27, 36, 37}. In the case of s-channel mediators these operators are identical to those discussed previously which contribute to direct detection. One can use Fierz transformations to write the t-channel models as linear combinations of models previously considered.

| Name  | $\chi$ Operator | $q$ Operator |
|-------|-----------------|--------------|
| DS\textsubscript{1} | $g\bar{\chi}\chi\phi$ | $g\lambda_q\bar{q}q\phi$ |
| DS\textsubscript{2} | $g\bar{\chi}\gamma^\mu\chi\phi_\mu$ | $g\bar{q}\gamma^\mu q\phi_\mu$ |
| DS\textsubscript{3} | $g\bar{\chi}\gamma^\mu\gamma^5\chi\phi_\mu$ | $g\bar{q}\gamma^\mu\gamma^5 q\phi_\mu$ |

| Name  | Operator      |
|-------|---------------|
| DT\textsubscript{1} | $g\bar{q}\chi\phi$ |
| DT\textsubscript{2} | $g\bar{q}\gamma^5\chi\phi$ |

\textbf{TABLE I:} Models of DM coupling to quarks. The models are named such that the first letter describes the DM particle as being Dirac fermions and the second letter gives what channel DM annihilations proceed through.

\textsuperscript{1} The flavor phenomenology of dark matter models has been more thoroughly considered in \textsuperscript{34, 35}. 
III. COLLIDER LIMITS AND REACHES

We derive limits and future reaches for collider searches looking for missing energy signatures. For all the bounds which we generate from colliders we only consider theories which may possibly be perturbative, that is, models which have \( \alpha \equiv g^2/4\pi < 4\pi \). We consider any model which has a greater coupling to not be the correct picture for understanding the dynamics of dark matter interactions. Note that, while some interactions of the model DS1 remain perturbative for larger values of \( g \) due to the presence of the quark Yukawa coupling, the coupling to top quarks, which is required by our hypothesis of flavor structure, becomes non-perturbative according to this criterion at approximately the same point. We thus assume perturbativity, as defined above, to be a theoretical bound on all models considered and we report only experimental bounds which are more stringent than this.

The simplest signature of an event which contains dark matter visible in direct detection is one of a jet and missing energy. While additional jets are possible, they generally require a more involved analysis, so we focus only on the monojet case here.\(^2\) In the context of the t-channel mediators this amounts to searching for associated production of a mediator and a dark matter candidate, rather than pair production of the mediators. In the special case of Supersymmetry, this process has been considered at NLO in \(^{39}\). A treatment similar to ours, using simplified models to interpret multijet and missing energy searches, has recently been published by the ATLAS collaboration \(^{43}\).

Our signal events were simulated using MadGraph and MadEvent version 4.4.56 \(^{40}\), using CTEQ6l PDFs with the renormalization scale chosen to be \( \mu^2 = M^2_\chi + P^2_{T,j} \). The events were hadronized using Pythia 6.2 \(^{41}\), and detector simulation was done through PGS4 \(^{42}\).

We utilize the recent search for monojets and missing energy, by the ATLAS collaboration \(^{43}\), to derive bounds on the coupling strength for each choice of dark matter mass and the mediator mass. The search includes three different sets of selection criteria which they call LowPt, HighPt, and VeryHighPt. The cuts required of each are shown in table II, along with the 95% CL limit which was generated on the effective cross section from new physics for each selection.

\(^2\) Other final states are possible and this has been explored in \(^{38}\).
We considered specifically the LowPt and VeryHighPt criteria, applying those cuts to our generated signal events and then finding the necessary coupling strength to saturate the bounds quoted above. We present bounds on $M_* \equiv M_{\phi}/g$ from the effective theory previously considered in [26, 27, 36, 37], as we are interested in the change of bounds from previous results due to the introduction of collider-accessible mediators. The resulting 95% CL limits are presented in figures 1-5. We note that, while bounds are generically weakened when mediators are lighter than the characteristic scale of the process which they are involved in ($M_{\phi} \lesssim 2M_\chi$ for s-channel, $M_{\phi} \lesssim M_\chi + E_{T,\text{cut}}$ for t-channel), above those thresholds the bounds are stronger or comparable to those previously derived within the assumption of a collider-inaccessible mediator.

### TABLE II: Cuts used in ATLAS monojet plus missing energy search [43]

| Name           | $E_T$  | monojet $P_t$ | 2nd jet $P_t$ | 3rd jet $P_t$ | 95% limit |
|----------------|--------|---------------|---------------|---------------|-----------|
| LowPt          | $> 120$ GeV | $> 120$ GeV | $< 30$ GeV     | $< 30$ GeV     | 1.7 pb    |
| HighPt         | $> 220$ GeV | $> 250$ GeV | $< 60$ GeV     | $< 30$ GeV     | 0.11 pb   |
| VeryHighPt     | $> 350$ GeV | $> 300$ GeV | $< 60$ GeV     | $< 30$ GeV     | 0.035 pb  |

FIG. 1: Bounds on effective interaction strength $M_* = M_{\phi}/g$ for the operator DS1. Bounds are presented only for massless $\chi$, as all other bounds are weaker than our adopted perturbativity limit of $g < 4\pi$. The red curve shows bounds resulting from the VeryHighPt analysis, and the black curve shows those resulting from the LowPt analysis.
FIG. 2: Bounds on effective interaction strength $M_* = M_{\phi}/g$ for the operator DS2. Note that the perturbativity constraint of $g < 4\pi$ replaces bounds weaker than that constraint. The left figures show bounds resulting from the VeryHighPt analysis, and the right figures show those resulting from the LowPt analysis.

IV. DIRECT DETECTION OF CANDIDATE MODELS

The direct detection calculations for s-channel annihilation models are rather straightforward, as the mediator, even if very light by other standards, is still very massive compared to the scales involved in direct detection, which are typically of order $\sim$ keV-MeV. We then can integrate out the mediator and return to the calculations of [26, 27], using nuclear matrix elements from [44, 45].

For t-channel models, we again must assume we can integrate out the mediator particle, which now requires it to be heavier than the DM candidate. This was already necessary to allow our assumed symmetry to protect the DM abundance. Once we have integrated out
the mediator we can apply the appropriate Fierz transformations to find the contribution to direct detection scatterings in the language of contact interactions. The direct detection cross sections predicted by the various contributing models are

\[
\sigma_{0,SI}^{DS1} = 1.60 \times 10^{-37} \text{cm}^2 \times \left( \frac{\mu_1}{1 \text{GeV}} \right)^2 \left( \frac{20 \text{GeV}}{M_*} \right)^4 \left( \frac{20 \text{GeV}}{v} \right)^2,
\]

(1)

\[
\sigma_{0,SI}^{DS2} = 1.38 \times 10^{-37} \text{cm}^2 \times \left( \frac{\mu_1}{1 \text{GeV}} \right)^2 \left( \frac{300 \text{GeV}}{M_*} \right)^4,
\]

(2)

\[
\sigma_{0,SD}^{DS3} = 9.18 \times 10^{-40} \text{cm}^2 \times \left( \frac{\mu_1}{1 \text{GeV}} \right)^2 \left( \frac{300 \text{GeV}}{M_*} \right)^4,
\]

(3)

\[
\sigma_{0,SI}^{DT1} = 2.21 \times 10^{-37} \text{cm}^2 \times \left( \frac{\mu_1}{1 \text{GeV}} \right)^2 \left( \frac{300 \text{GeV}}{M_*} \right)^4,
\]

(4)

\[
\sigma_{0,SD}^{DT1} = 2.30 \times 10^{-40} \text{cm}^2 \times \left( \frac{\mu_1}{1 \text{GeV}} \right)^2 \left( \frac{300 \text{GeV}}{M_*} \right)^4,
\]

(5)

\[
\sigma_{0,SI}^{DT2} = 8.08 \times 10^{-38} \text{cm}^2 \times \left( \frac{\mu_1}{1 \text{GeV}} \right)^2 \left( \frac{300 \text{GeV}}{M_*} \right)^4.
\]

(6)

Of particular interest is the interference behavior giving rise to the cross sections due to t-channel models. Model DT1 has constructive interference between the Fierz products which contribute to spin-dependent scattering, while in model DT2 they cancel, giving no spin-dependent signal. The same is true for spin-independent scattering, though the cancellation is less effective so the cross section is merely reduced, not eliminated.

We have translated the bounds on coupling strength to bounds on the direct detection cross section for each model considered. The results, for various choices of mediator mass, are compared to the most stringent current bounds from direct detection experiments and presented in figures 6-10. More detailed spectral information from direct detection can also be exploited in an effective theory framework, see [46].

V. CONCLUSIONS

We have determined the collider bounds due to signals of jets and missing energy on models of dark matter which have UV complete interactions with the SM through mediators within the reach of collider experiments. We have also presented these bounds in terms of the upper limit on direct detection cross sections which they impose for various mediator masses. We find that having light mediating particles can, in general, degrade the bounds derived
FIG. 3: Bounds on effective interaction strength $M_*=M_\phi/g$ for the operator DS3. Note that the perturbativity constraint of $g<4\pi$ replaces bounds weaker than that constraint. The left figures show bounds resulting from the VeryHighPt analysis, and the right figures show those resulting from the LowPt analysis.

based on the assumption of contact interactions, in agreement with previous studies [28, 47].

However, the weakening of bounds due to light mediators occurs only within certain kinematic ranges; $M_\phi \lesssim 2M_\chi$ for $s$-channel and $M_\phi \sim M_\chi$ for $t$-channel. Outside of these kinematic ranges we find that bounds are as strong or stronger than those found in previous studies which assumed contact operators. For the purposes of the presented analysis with hard cuts, $t$-channel mediators must be $\sim 300$ GeV heavier than dark matter candidates in order to have bounds comparable to those derived from contact operator analyses.

In figures 2 and 3, a clear line is visible for $M_\phi \sim 2M_\chi$. Above this line, resonant enhancement of the dark matter production process can occur. This strengthens the constraints on $M^*$ and therefore the bounds on direct detection cross sections become more stringent. In
FIG. 4: Bounds on effective interaction strength $M_* = M\phi/g$ for the operator DT1. Note that the perturbativity constraint of $g < 4\pi$ replaces bounds weaker than that constraint. The left figures show bounds resulting from the VeryHighPt analysis, and the right figures show those resulting from the LowPt analysis.

A horizontal line is seen, with its position dependent on the hardness of cuts on $E_T$ and jet $P_t$. This is due to the presence or absence of enough mass splitting between the mediator and the dark matter candidate to generate kinematics beyond the cuts without significant initial boost. When the splitting is large compared to the cut values, the limits are largely insensitive to changes in mediator mass.

Though we have only presented limits here for Dirac dark matter, the key differences between contact interaction models as presented in [26–28, 36, 48], are not strongly dependent on the Lorentz representation of the dark matter, as it is largely an issue of phase space and kinematics. These bounds apply to all models of dark matter without regard to the relic density generation mechanism [49].
FIG. 5: Bounds on effective interaction strength $M_\ast = M_\phi/g$ for the operator DT2. Note that the perturbativity constraint of $g < 4\pi$ replaces bounds weaker than that constraint. The left figures show bounds resulting from the VeryHighPt analysis, and the right figures show those resulting from the LowPt analysis.

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[1] E. Komatsu et al., arXiv:1001.4538 [astro-ph.CO].
[2] J. L. Feng and J. Kumar, Phys. Rev. Lett. 101, 231301 (2008) arXiv:0803.4196 [hep-ph].
FIG. 6: Constraints on direct detection spin-independent cross section as a function of dark matter mass. The black curves correspond to the experimental limits from Xenon 100 and the CDMSII low-threshold analysis, and the red, blue, and magenta curves correspond to the bounds for operator DS2 with mediators of mass $M_\phi$ of 10, 100, and 1000 GeV, respectively.

J. L. Feng, H. Tu and H. B. Yu, JCAP 0810 (2008) 043 [arXiv:0808.2318 [hep-ph]].

[3] A. Birkedal, K. Matchev and M. Perelstein, Phys. Rev. D 70, 077701 (2004) [arXiv:hep-ph/0403004]; P. Konar, K. Kong, K. T. Matchev and M. Perelstein, New J. Phys. 11, 105004 (2009) [arXiv:0902.2000 [hep-ph]]; J. L. Feng, S. Su and F. Takayama, Phys. Rev. Lett. 96, 151802 (2006) [arXiv:hep-ph/0503117].

[4] M. Beltran, D. Hooper, E. W. Kolb and Z. A. C. Krusberg, Phys. Rev. D 80, 043509 (2009) [arXiv:0808.3384 [hep-ph]].

[5] Q. H. Cao, C. R. Chen, C. S. Li and H. Zhang, arXiv:0912.4511 [hep-ph].

[6] M. Beltran, D. Hooper, E. W. Kolb, Z. A. C. Krusberg and T. M. P. Tait, arXiv:1002.4137 [hep-ph].

[7] W. Shepherd, T. M. P. Tait and G. Zaharijas, Phys. Rev. D 79, 055022 (2009) [arXiv:0901.2125].
FIG. 7: Constraints on direct detection spin-dependent cross section as a function of dark matter mass. The black curves correspond to the experimental limits from Xenon 10 on neutron scatterings (solid curve) and PICASSO on proton scattering (dashed curve), and the red, blue, and magenta curves correspond to the bounds for operator DS3 with mediators of mass $M_\phi$ of 10, 100, and 1000 GeV, respectively.

[hep-ph]].

[8] C. E. Aalseth et al. [CoGeNT collaboration], [arXiv:1002.4703] [astro-ph.CO].

[9] C. E. Aalseth, P. S. Barbeau, J. Colaresi, J. I. Collar, J. D. Leon, J. E. Fast, N. Fields, T. W. Hossbach et al., [arXiv:1106.0650] [astro-ph.CO]].

[10] R. Bernabei et al., [arXiv:1002.1028] [astro-ph.GA].

[11] E. Aprile et al. [XENON100 Collaboration], [arXiv:1005.0380] [astro-ph.CO].

[12] Z. Ahmed et al. [CDMS-II Collaboration], Phys. Rev. Lett. 106, 131302 (2011) [arXiv:1011.2482] [astro-ph.CO]].

[13] F. Petriello and K. M. Zurek, JHEP 0809, 047 (2008) [arXiv:0806.3989] [hep-ph]].

[14] J. L. Feng, J. Kumar and L. E. Strigari, Phys. Lett. B 670, 37 (2008) [arXiv:0806.3746]
FIG. 8: Constraints on direct detection spin-independent cross section as a function of dark matter mass. The black curves correspond to the experimental limits from Xenon 100 and the CDMSII low-threshold analysis, and the red, blue, and magenta curves correspond to the bounds for operator DT1 mediators of mass $M_\phi$ of $10 + M_\chi$, $100 + M_\chi$, and $1000$ GeV, respectively.

[hep-ph].

[15] T. Schwetz and J. Zupan, [arXiv:1106.6241] [hep-ph].
[16] P. J. Fox, J. Kopp, M. Lisanti and N. Weiner, [arXiv:1107.0717] [hep-ph].
[17] M. Farina, D. Pappadopulo, A. Strumia and T. Volansky, [arXiv:1107.0715] [hep-ph].
[18] J. L. Feng, J. Kumar, D. Marfatia, D. Sanford, [arXiv:1102.4331] [hep-ph].
[19] S. Chang, J. Liu, A. Pierce, N. Weiner and I. Yavin, [arXiv:1004.0697] [hep-ph].
[20] Z. Kang, J. Li, T. Li, T. Liu, J. Yang, [arXiv:1102.5644] [hep-ph].
[21] E. Del Nobile, C. Kouvaris, F. Sannino, [arXiv:1105.5431] [hep-ph].
[22] X. Gao, Z. Kang, T. Li, [arXiv:1107.3529] [hep-ph].
[23] M. T. Frandsen, F. Kahlhoefer, J. March-Russell, C. McCabe, M. McCullough and K. Schmidt-Hoberg, [arXiv:1105.3734] [hep-ph].
FIG. 9: Constraints on direct detection spin-dependent cross section as a function of dark matter mass. The black curves correspond to the experimental limits from Xenon 10 on neutron scatterings (solid curve) and PICASSO on proton scattering (dashed curve), and the red, blue, and magenta curves correspond to the bounds for operator DT1 with mediators of mass $M_\phi$ of $10 + M_\chi$, $100 + M_\chi$, and $1000$ GeV, respectively.

[24] Y. Gao, J. Kumar, D. Marfatia, [arXiv:1108.0518 [hep-ph]].
[25] S. -L. Chen, Y. Zhang, [arXiv:1106.4044 [hep-ph]].
[26] J. Goodman, M. Ibe, A. Rajaraman, W. Shepherd, T. M. P. Tait and H. B. P. Yu, [arXiv:1005.1286 [hep-ph]].
[27] J. Goodman, M. Ibe, A. Rajaraman, W. Shepherd, T. M. P. Tait, H. -B. Yu, Phys. Rev. D82, 116010 (2010). [arXiv:1008.1783 [hep-ph]].
[28] Y. Bai, P. J. Fox and R. Harnik, [arXiv:1005.3797 [hep-ph]].
[29] C. P. Burgess, M. Pospelov and T. ter Veldhuis, Nucl. Phys. B 619, 709 (2001) [arXiv:hep-ph/0011335];
[30] K. Cheung, P. Y. Tseng and T. C. Yuan, JCAP 1101, 004 (2011) [arXiv:1011.2310 [hep-ph]].
FIG. 10: Constraints on direct detection spin-independent cross section as a function of dark matter mass. The black curves correspond to the experimental limits from Xenon 100 and the CDMSII low-threshold analysis, and the red, blue, and magenta curves correspond to the bounds for operator DT2 with mediators of mass $M_\phi$ of $10 + M_\chi$, $100 + M_\chi$, and 1000 GeV, respectively.

[31] Y. Mambrini and B. Zaldivar, arXiv:1106.4819 [hep-ph].
[32] S. Kanemura, S. Matsumoto, T. Nabeshima and N. Okada, arXiv:1005.5651 [hep-ph].
[33] A. J. Buras, P. Gambino, M. Gorbahn, S. Jager and L. Silvestrini, Phys. Lett. B 500, 161 (2001) [arXiv:hep-ph/0007085].
[34] J. F. Kamenik, J. Zupan, arXiv:1107.0623 [hep-ph].
[35] K. Cheung, K. Mawatari, E. Senaha, P. Y. Tseng and T. C. Yuan, JHEP 1010, 081 (2010) [arXiv:1009.0618 [hep-ph]].
[36] A. Rajaraman, W. Shepherd, T. M. P. Tait, A. M. Wijangco, [arXiv:1108.1196 [hep-ph]].
[37] J. Goodman, M. Ibe, A. Rajaraman, W. Shepherd, T. M. P. Tait, H. -B. Yu, Nucl. Phys. B844, 55-68 (2011). [arXiv:1009.0008 [hep-ph]].
[38] J. Wang, C. S. Li, D. Y. Shao and H. Zhang, arXiv:1107.2048 [hep-ph].
[39] T. Binoth, D. Goncalves Netto, D. Lopez-Val, K. Mawatari, T. Plehn, I. Wigmore, Phys. Rev. D84, 075005 (2011). [arXiv:1108.1250 [hep-ph]].

[40] J. Alwall, P. Demin, S. de Visscher, R. Frederix, M. Herquet, F. Maltoni, T. Plehn, D. L. Rainwater et al., JHEP 0709, 028 (2007). [arXiv:0706.2334 [hep-ph]].

[41] T. Sjostrand, S. Mrenna and P. Z. Skands, JHEP 0605, 026 (2006) [arXiv:hep-ph/0603175].

[42] http://www.physics.ucdavis.edu/~conway/research/software/pgs/pgs4-general.htm

[43] The ATLAS collaboration, ATLAS-CONF-2011-096 (2011).

[44] G. Belanger, F. Boudjema, A. Pukhov, and A. Semenov, Comput.Phys.Commun.180:747-767,2009 [arXiv:0803.2360v2 [hep-ph]].

[45] J. Giedt, A. W. Thomas and R. D. Young, Phys. Rev. Lett. 103, 201802 (2009) [arXiv:0907.4177 [hep-ph]].

[46] J. Fan, M. Reece and L. T. Wang, JCAP 1011, 042 (2010) [arXiv:1008.1591 [hep-ph]].

[47] J. F. Fortin and T. M. P. Tait, [arXiv:1103.3289 [hep-ph]].

[48] P. J. Fox, R. Harnik, J. Kopp, Y. Tsai, [arXiv:1103.0240 [hep-ph]].

[49] M. R. Buckley, [arXiv:1104.1429 [hep-ph]].