Production cross section of neutron-rich isotopes with radioactive and stable beams

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The production cross section of neutron-rich isotopes of Ca, Zn, Te, Xe, and Pt are predicted in the diffusive multi-nucleon transfer reactions with stable and radioactive beams. With these isotopes one can treat the neutron shell evolution beyond \( N = 28, 50, 82, \) and 126. Because of the small cross sections, the production of nuclei near the neutron drip line requires the optimal choice of reaction partners and bombarding energies.

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I. INTRODUCTION

One of the primary issues of nuclear physics is an expansion of the present limits of the nuclear chart far from the line of stability. The neutron-rich isotopes located close to the neutron-drip line have attracted interests during the last few decades. In the study of exotic neutron-rich nuclei a variety of new aspects have been discovered such as the existence of neutron skins, neutron halos, and shell evolutions leading to new magic numbers. Furthermore, these neutron-rich nuclei play a crucial role in the \( r \)-process in nuclear synthesis.

Nowadays, the nuclei far from stability are accessible in vast region of the nuclear chart by various reactions. Fission should be suitable to generate a certain number of neutron-rich nuclides [1], but this method was found not to be promising for the production of nuclei near the drip line. Instead, the fragmentation reactions can produce very neutron-rich nuclei [2, 3] at energies well above the Fermi energy. However, in the fragmentation reactions, the yields of these nuclei decrease with increasing isospin quantum number and a large excitation energy is available, which can wash out the primary yields of extremely neutron-rich nuclei. On the other hand, the neutron-rich nuclei can be also effectively produced in the transfer-type reactions in which the control of excitation energy is easier [4–17].

The experimental setups, which are presently under construction, would provide quite intensive beams of radioactive nuclei. (See, for example, Ref. [18].) One of the applications of these beams would be the production of new neutron-rich isotopes using either complete fusion or transfer reactions. While one can reach only specific isotopes in the complete fusion reactions, the transfer-type reactions can yield larger variety of isotopes because of the restrictions in the choice of the target. Note that the use of the beam of neutron-rich nuclei in the fragmentation reactions has no advantage compared with the beam of stable isotopes.

The aim of the present paper is to predict the production cross sections of new neutron-rich Ca, Zn, Te, Xe, and Pt isotopes via the diffusive multi-nucleon transfer reactions at incident energies near the corresponding Coulomb barriers (3–5 MeV/nucleon). These very neutron-rich isotopes have the neutron numbers beyond \( N = 28, 50, 82, \) and 126, respectively. So, the study on their productions is expected to provide useful information about the shell evolution near the drip line. Because the cross sections for producing neutron-rich isotopes near the neutron drip line are estimated to be small, the production of neutron-rich isotopes requires an optimal choice of the projectile-target combinations and bombarding energies for future experiments. In the present paper, by extending the work of Ref. [16], we investigate the production of various neutron-rich isotopes and study other possible reactions for the production of the isotopes previously considered.

The present paper is organized as follows. In Sec. II we present the model used for estimating the production cross sections of neutron-rich nuclei. The obtained results are given in Sec. III with detailed discussions. We then summarize in Sec. IV.
II. MODEL

The dinuclear system (DNS) concept \cite{19} was proposed to explain the mechanisms of the complete fusion, quasi-fission, and nucleon transfer reactions. The quasifission and transfer products result from the decays of various DNS configurations. The DNS is formed at the capture stage of the reaction after complete dissipation of the kinetic energy of the relative motion \cite{13,16,21,24}. The dynamics of the process is considered as a diffusion of the DNS in the charge and mass asymmetry coordinates, which are defined by the charge and neutron numbers \(Z\) and \(N\), respectively, of the light or heavy nucleus of the DNS. During the evolution in \(Z\) and \(N\), the excited DNS can decay into two fragments by the diffusion in relative distance \(R\) between the centers of the DNS nuclei.

In our approach, the production of exotic nuclei is considered as a three-step process. First, the initial DNS with the nucleus of \((Z = Z(0), N = N(0))\) is formed in the entrance reaction channel. And then the DNS with exotic nucleus of \((Z = Z^{(ex)}, N = N^{(ex)})\) is created through nucleon transfers. Finally, this DNS is rapidly divided into two fragments. The created exotic neutron-rich nuclei with \((Z^{(ex)}, N^{(ex)})\) are predominantly de-excited \(\gamma\) emissions and neutron evaporation.

In the present work, we firstly consider the reactions leading to the neutron-rich nuclei whose excitation energies are smaller than their neutron separation energies. Then we study the opportunity of feeding the yield of neutron-rich nuclei from the neutron emission of properly excited more neutron-rich isotopes.

A. Potential energy of dinuclear system

In the DNS, the potential energy is defined as the sum of the mass excesses of the both nuclei and their interaction energy. Therefore, the phenomenological potential of the DNS is written as \cite{20, 21, 25}:

\[
U(R, Z, N, J) = B_1 + B_2 + V(R, Z, N, J),
\]

where \(B_1\) and \(B_2\) are the mass excesses of the light and heavy nuclei, respectively, and \(V\) is the nucleus-nucleus interaction potential. In this work, the experimental mass excess is used, if possible. For the nuclei whose empirical mass excesses are not available, we take the values predicted in Ref. \cite{27} within the microscopic-macroscopic model.

The nucleus-nucleus interaction potential between two nuclei can be written as

\[
V(R, Z, N, J) = V_C(R, Z, N) + V_N(R, Z, N) + V_{rot}(R, Z, N, J),
\]

where \(V_C\) includes the Coulomb interaction, the nuclear interaction, and the centrifugal term. The centrifugal potential \(V_{rot}\) in the formed DNS is usually expressed under the assumption that this system has the rigid body moment of inertia \(I\), which leads to

\[
V_{rot} = \frac{h^2 J(J+1)}{2I},
\]

where \(J\) is the orbital angular momentum of the collision between the two nuclei.

In the calculation of the Coulomb and nuclear interactions, the deformations of the DNS nuclei are important. The Coulomb potential for two quadrupole deformed nuclei is analytically calculated following Ref. \cite{25} as

\[
V_C = \frac{e^2 Z_1 Z_2}{R} + \frac{e^2 Z_2}{R^3} \left(\frac{9}{20\pi}\right)^{1/2} \sum_{j=1,2} R_0 j^2 \beta_j P_2(\cos \theta_j).
\]

Here, \(Z_j\) and \(\beta_j\) are the proton number and the quadrupole deformation parameter of the nucleus “\(j\)” \((j = 1, 2)\) in the DNS, respectively, and \(P_2(\cos \theta_j)\) is the Legendre polynomial with the angle \(\theta_j\) between the symmetry axis of the nuclei “\(j\)” and the axis \(z\) connecting the centers of mass of the DNS nuclei. The radius of a nucleus “\(j\)” with mass number \(A_j = Z_j + N_j\) is parameterized as \(R_0_j = r_0 A_j^{1/3}\), where \(r_0 = 1.15\) fm. In the present work, we take the \(\beta_j\) values from Ref. \cite{28}.

The nuclear part of the nucleus-nucleus potential energy reads

\[
V_N = \int \rho_1(r_1) F(r_1 - r_2) \rho_2(R - r_2) \, dr_1 \, dr_2,
\]

which is obtained by using the double folding formalism \cite{25} with the density-dependent nucleon-nucleon interaction. Among those kinds of interactions, we employ the Skyrme-type effective interaction which is well known from the theory of finite Fermi systems \cite{29}:

\[
F(r_1 - r_2) = C_0 \left[ F_{\text{in}} \left( \frac{\rho_1(r_1)}{\rho_{00}} \right) + F_{\text{ex}} \left( 1 - \frac{\rho_0(r_1)}{\rho_{00}} \right) \right] \times \delta(r_1 - r_2),
\]

\[
F_{\text{in,ex}} = f_{\text{in,ex}} + f'_{\text{in,ex}} \frac{(N_1 - Z_1)(N_2 - Z_2)}{A_1 A_2},
\]

\[
\rho_0(r) = \rho_1(r) + \rho_2(r).
\]

For our numerical calculations we use the following parameters: \(C_0 = 300\) MeV fm\(^3\), \(f_{\text{in}} = 0.09\), \(f_{\text{ex}} = 2.59\), \(f'_{\text{in}} = 0.42\), \(f'_{\text{ex}} = 0.54\), and \(\rho_{00} = 0.17\) fm\(^{-3}\). The densities of the nuclei are taken in the two-parameter Saxon-Woods form as

\[
\rho_j(r) = \frac{\rho_{00}}{1 + \exp \left((r - R_j(\theta'_j, \varphi'_j))/a_{0j}\right)},
\]

with the diffuseness parameter \(a_{0j} = 0.55\) fm for the nuclei considered in this work and \(R_j(\theta'_j, \varphi'_j) = R_{0j}(1 + \beta_{2j} Y_{20}(\theta'_j, \varphi'_j))\), where \(\theta'_j\) and \(\varphi'_j\) are the angles defined in the body-fixed system.
B. Production cross sections

The production cross section $\sigma_{Z,N}$ of the nucleus with $Z$ proton and $N$ neutrons can be written as

$$\sigma_{Z,N} = \sigma_{\text{cap}}Y_{Z,N},$$

where $\sigma_{\text{cap}}$ and $Y_{Z,N}$ are the capture cross section and the formation-decay probability of the DNS configuration with the given charge and mass asymmetries, respectively. The capture cross section is expressed as

$$\sigma_{\text{cap}} = \frac{\pi\hbar^2}{2\mu E_{\text{c.m.}}}J_{\text{cap}}(J_{\text{cap}} + 1),$$

where $\mu$ is the reduced mass for the projectile-target system and $E_{\text{c.m.}}$ is the bombarding energy in the center of mass frame. The value of angular momentum $J_{\text{cap}}$ is determined by the bombarding energy. For example, if $E_{\text{c.m.}}$ is 20% larger than the Coulomb barrier energy, the orbital angular momentum is about 90. If the value of $E_{\text{c.m.}}$ is close to the Coulomb barrier energy, $J_{\text{cap}}$ does not exceed 30–40. However, the angular momentum decreases the stability of the exotic nuclei of interest. So, we choose $J_{\text{cap}} = 30$ to produce neutron-rich isotopes of interest with rather small angular momenta.

The formation-decay probability is calculated by the simple statistical method introduced in Ref. [13], which reads

$$Y_{Z,N} \approx 0.5\exp \left[ -\frac{U(R_m(Z,N_0), Z, N_0, J) - U(R_m(Z^{(i)}, N^{(i)}), Z^{(i)}, N^{(i)}, J) - B_{\text{qf}}(Z^{(i)}, N^{(i)})}{\Theta(Z^{(i)}, N^{(i)})} \right].$$

Here, we use the DNS potential energy $U$ at the touching distance $R_m(Z,N) \approx R_0(1 + \sqrt{5/4\pi}\beta_2) + R_0(1 + \sqrt{5/4\pi}\beta_2) + 0.5\text{ fm}$. The decaying DNS with given $Z$ and $N$ has to escape from the local minimum at $R = R_m(Z,N)$ by overcoming the potential barrier at $R = R_b(Z,N)$, where $R_b(Z,N) \approx R_m(Z,N) + 2\text{ fm}$. The height of the barrier which the DNS with $Z$ and $N_0$ should overcome to observe the nucleus with $Z$ and $N$ is given by $B_R(Z,N) = U(R_0(Z,N), Z, N, J) - U(R_m(Z,N_0), Z, N_0, J)$. The neutron number $N_0$ corresponds to the $N/Z$ equilibrium in the DNS at given $Z$, i.e., the conditional minimum of the potential energy surface. In Eq. (11), the value of $B_{\text{qf}} = \min(B_{\text{qf}}^\text{sym}, B_{\text{qf}}^R)$ is the barrier preventing the DNS decay either in $R$ or toward more symmetric configurations. The DNS symmetrization increases the Coulomb repulsion and, thus, leads to its decay from more symmetric configuration. The value of $B_{\text{qf}}^\text{sym}$ measures the barrier for the initial DNS toward more symmetric configurations.

The temperature $\Theta(Z^{(i)}, N^{(i)})$ is estimated using the Fermi-gas expression $\Theta = \sqrt{E^*}/a$, where $a = A_{\text{tot}}/12\text{ MeV}^{-1}$, $A_{\text{tot}} = A_1 + A_2$ is the total mass number of the DNS. The temperature $\Theta(Z,N_0)$ is calculated for the excitation energy $E^*(Z^{(i)}, N^{(i)}) = U(R_m(Z,N_0), Z, N_0, J) - U(R_m(Z^{(i)}, N^{(i)}), Z^{(i)}, N^{(i)}, J)$.

To reduce the number of transferred nucleons and, thus, to increase the yields, we consider only the production of neutron-rich nuclei with $Z = Z^{(i)}$ and $N > N^{(i)}$. In the reactions treated the initial DNS are rather close in energy to the $N/Z$ equilibrium. In this case $\Theta(Z^{(i)}, N^{(i)}) \approx \Theta(Z,N_0)$ and the expression for the formation-decay probability $Y_{Z,N}$ in Eq. (11) is simplified as

$$Y_{Z^{(i)},N} \approx 0.5\exp \left[ -\frac{B_R(Z^{(i)}, N) - B_{\text{qf}}(Z^{(i)}, N^{(i)})}{\Theta(Z^{(i)}, N^{(i)})} \right].$$

III. CALCULATED RESULTS

In the present work, we focus on the multi-nucleon transfers that transform the initial DNS to a larger $N$ when $Z = Z^{(i)}$. In order to supply a larger yield of the neutron-rich nucleus with $(Z^{(i)}, N)$, the potential energy of the DNS containing this nucleus should be close to that of the initial DNS with $(Z^{(i)}, N^{(i)})$. This condition should be one of the main criteria for selecting the projectile and target nuclei. The excitation energy of the initial DNS should not exceed the threshold above which the excitation energy of the neutron-rich product is larger than its neutron separation energy. So, the nuclei produced near the neutron drip line should be quite cold to avoid their losses during the de-excitation process. The primary neutron-rich nucleus excited above the threshold contributes to the yield of the nucleus with smaller number of neutrons. Because the exotic nucleus results from the multi-nucleon transfers between the projectile-like and target-like parts of the DNS, one can assume the thermal equilibrium in the DNS containing the exotic nucleus or in the DNS which is quite far from the initial DNS in the $(Z,N)$ space. Indeed, for the formation of these DNS, quite a long time is needed such as $t_0 \approx 10^{-20}\text{ s}$ at $J \lesssim 30$. Therefore, in the considered reactions the excitation energy of the light or heavy nucleus...
with the mass $A_1$ or $A_2$ is

$$E^{i,2}_{1,2}(Z^{(i)}, N) = \left[ E^*(Z^{(i)}, N_0) - B_R(Z^{(i)}, N) \right] \frac{A_{1,2}}{A_{\text{tot}}}.$$  \hfill (12)

The deviation from the thermal equilibrium is expected only for the DNS near the initial DNS with $(Z^{(i)}, N^{(i)})$ which are formed in a short time. If one of the DNS nuclei is magic, it takes a smaller excitation energy to create Te isotopes close to the neutron drip line. The Ca nuclei are taken from the predictions of the finite-range liquid-drop model presented in Ref. [20].

Through Refs. [15–17, 23] the applicability of the present model to describe the existing experimental data has been tested. The yields of isotopes, which are several nucleons away from the entrance channel, are well described by the DNS potential energy, and our approach therefore seems to be suitable.

Employing the radioactive beam of $^{144}$Xe, one can produce more neutron-rich isotopes of Xe and study their properties. The nuclei $^{148,150,152}$Xe have more than 93 neutrons, i.e., with these nuclei one can trace an evolution of neutron shells well beyond $N = 82$. The predicted cross sections in $0n$ evaporation channel are presented in Fig. 1. Here, we show only the cross sections larger than 0.01 pb. In the reactions considered in this work, the formation of the nuclei of interest occurs in accordance with the following schemes: $^{144}$Xe $+$ $^{48}$Ca $\rightarrow$ $^{142}$Xe $+$ $^{50}$Ca $\rightarrow$ $^{148,150,152}$Xe $+$ $^{44,42,40}$Ca, $^{144}$Xe $+$ $^{70}$Zn $\rightarrow$ $^{136}$Xe $+$ $^{72}$Zn $\rightarrow$ $^{140,150}$Xe $+$ $^{66,64}$Zn, $^{144}$Xe $+$ $^{86}$Kr $\rightarrow$ $^{138}$Xe $+$ $^{92}$Kr $\rightarrow$ $^{142,150,152}$Xe $+$ $^{82,80,78}$Kr, $^{144}$Xe $+$ $^{88}$Sr $\rightarrow$ $^{136}$Xe $+$ $^{96}$Sr $\rightarrow$ $^{148}$Xe $+$ $^{84}$Sr. The $^{48}$Ca target seems to be favorable to produce $^{148,150,152}$Xe, while the $^{86}$Kr target would lead to relatively large cross sections for the yields of $^{148}$Xe. With increasing charge number of the target nucleus the initial DNS becomes unstable with respect to the decay in $R$. So, the use of the combination with a larger $Z_1 \times Z_2$ does not seem to be promising even they supply a gain in $Q$ values.

Figure 2 shows the cross sections predicted for the production of neutron-rich isotopes of Te with the beams of $^{134,136}$Te. Only the yields with the cross sections larger than 0.1 pb are presented. As in the case of Xe beams, the $^{48}$Ca target is expected to be the most appropriate to create Te isotopes close to the neutron drip line.

For these calculations we use the neutron separation energies $S_n$ predicted in Ref. [20]. As a result, the model-dependence of these values is unavoidable. To study the influence of the uncertainties of $S_n = 4 - 4.5$ MeV, in the case of the $^{136}$Te $+$ $^{48}$Ca reaction, we consider the cases of $S_n = 0.5$ MeV. The variations of the obtained cross sections are marked by the error bars in Fig. 2.

FIG. 1. (Color online) The predicted cross sections of the production of neutron-rich isotopes $^{148,150,152}$Xe in the transfer reactions with the radioactive beam of $^{144}$Xe. The cross sections correspond to the maxima of $0n$ evaporation channels. The symbols correspond to the indicated reactions.

FIG. 2. (Color online) The predicted cross sections of the production of neutron-rich isotopes $^{136,138,140,142}$Te in the transfer reactions $^{134,136}$Te $+$ $^{48}$Ca. The cross sections correspond to the maxima of the $0n$ evaporation channels (closed symbols) and $2n$ evaporation channels (open symbols). The symbols correspond to the indicated reactions. Arrows connect the points corresponding to the productions of the same isotope but in different evaporation channels.
We find that 0.5 MeV uncertainty in $S_n$ causes about 20% uncertainty in the cross section and about 1 MeV uncertainty in the corresponding value of $E_{c.m.}$.

If the bombarding energy is higher, then the neutron evaporation process occurs from the primary excited products. Here, we also consider the $2n$ evaporation channels as in the reactions of $^{134}\text{Te} + ^{48}\text{Ca} \rightarrow ^{136,138,140}\text{Te} + 2n$. As shown in Fig. 2, the yields of the primary products with the same $N$ are larger in the $2n$ channel than in the $0n$ channel because of the larger excitation. After the 2-neutron emission, the secondary products ($N-2$) appear with the same cross sections as the primary products ($N$) which are almost one order of magnitude smaller than those for the products produced in the $0n$ evaporation channel. So, the increase of $E_{c.m.}$ feeds the yield of isotopes with smaller number of neutrons. Although this feeding is quite small for the isotopes considered in this present work, one may increase the thickness of the target in future experiments in order to increase the total yields of the isotopes of interest. We note that the role of the feeding becomes more important as decreasing the neutron number of the produced isotope.

With the radioactive beams of $^{142,144}\text{Xe}$ one can produce neutron-rich target-like nuclei with $N > 50$. In Fig. 3 the cross sections are predicted for the reactions of $^{142,144}\text{Xe} + ^{68,70}\text{Zn} \rightarrow ^{82,84,86}\text{Zn}$. The yields of the neutron-rich Zn isotopes increase with neutron numbers of the projectile and target. The yield of neutron-rich isotopes decreases almost by two orders of magnitude when its neutron number increases by 2 units. The reactions of $^{142}\text{Xe} + ^{70}\text{Zn}$ and $^{144}\text{Xe} + ^{68}\text{Zn}$ are found to have the same value of the cross sections but different values of the optimal $E_{c.m.}$. The production of $^{86}\text{Zn}$ occurs with the largest cross section in the $^{144}\text{Xe} + ^{70}\text{Zn}$ reaction. Therefore, the shell evolution beyond $N = 50$ can be traced by studying neutron-rich nuclei in this region of $Z$.

The reactions of $^{134,136}\text{Te} + ^{48}\text{Ca}$ can be used to produce the neutron-rich isotopes $^{52,54,56,58,60}\text{Ca}$ ($N > 28$) with the cross sections larger than 1 pb as shown in Fig. 4. The replacement of $^{134}\text{Te}$ by $^{136}\text{Te}$ leads to about 5 times larger yields of neutron-rich Ca isotopes. However, these cross sections are about 5 times smaller compared with the $^{132}\text{Sn} + ^{48}\text{Ca}$ reaction.

To study the neutron-rich nuclei with $N$ around...
126, we consider the production of $^{202,204,206}$Pt. The unknown platinum isotopes can be produced in the nucleon transfer reactions of $^{198}$Pt + $^{48}$Ca, $^{50}$Ti, $^{54}$Cr, $^{58}$Fe, $^{64}$Ni, $^{70}$Zn, $^{76}$Ge → $^{202,204,206}$Pt with stable beams. The estimated cross sections are presented in Fig. 6. The $^{198}$Pt + $^{48}$Ca reaction is the most favorable to produce the $^{202,204,206}$Pt isotopes. For the projectiles with larger value of Z the cross sections are about two orders of magnitude smaller. In the reactions with $^{64}$Ni, $^{70}$Zn, and $^{76}$Ge the cross sections are similar to or larger than those of the reactions with $^{50}$Ti, $^{54}$Cr, and $^{58}$Fe due to the gains in the Q-values. With the reaction partner heavier than $^{76}$Ge the DSN formed in this case becomes very unstable with respect to the decay in R and, therefore, the nucleon transfers are reduced. This shows that the use of more symmetric reactions does not help to produce isotopes close to the neutron drip line.

In Fig. 6 we compare the cross sections in different neutron evaporation channels. Although the cross section of the production of $^{204}$Pt in the 2n evaporation channel is more than two orders of magnitude smaller than that of the 0n evaporation channel, it is larger than 0.1 nb and could be examined with the current experimental facilities. In order to make these experiments feasible, thicker targets and higher energy will be helpful as they would increase the yield.

The results shown in Fig. 6 show that the use of the radioactive beams of $^{66,68,70,72}$Ni and the $^{198}$Pt target will allow to increase the yields of the neutron-rich nuclei $^{202,204,206}$Pt in comparison with the case of the $^{64}$Ni beam. The increase of the mass number of Ni by 2 units leads to more than 10 times larger cross sections. In the $^{72}$Ni + $^{198}$Pt reaction one can produce even $^{210}$Pt isotope ($N=132$) with the cross section of about 1 pb.

IV. SUMMARY

In this work, we propose the multi-nucleon transfer reactions with radioactive and stable beams to produce the neutron-rich isotopes close to the neutron drip-line. The radioactive beams are expected to be of a good intensity enough to produce such isotopes properly even with a small cross section. Thus we search for the production of neutron-rich isotopes of which cross sections are larger than 1 pb as the experiments for such reactions can be performed at current and future facilities. In the reactions considered in the present work, the neutron-rich nuclei with neutron numbers beyond the magic values $N = 28$ (52,$^{54,56,58,60}$Ca), $N = 50$ ($^{82,84,86}$Zn), $N = 82$ (136,138,140,142,148,150,152Te and Xe), and $N = 126$ ($^{202,204,206,208}$Pt) would be produced with the cross sections larger than 1 pb. The study of these nuclei is, therefore, desirable to trace the shell evolution toward the neutron drip line. The results of the present calculations show that the production cross sections of neutron-rich nuclei increase by about one order of magnitude with increasing mass number by 2 units in the reaction partners. The maximal cross sections are found to be in the 0n evaporation channels. The cross sections in 1n ~ 4n evaporation channels are smaller but still measurable. Then in this case, one can use higher bombarding energies and thicker targets to increase the yields of neutron-rich nuclei.
in this work can be tested by the current experimental facilities. If the expected intensity of the radioactive beam is $10^7 - 10^9$ pps, then the production rate of the nuclides is roughly estimated as 0.01–1 particle per day at production cross section 1 nb. Therefore, the observation of the isotopes produced with the cross section smaller than 1 nb would be difficult in the experiments proposed with the present radioactive beams. The production cross section of 1 pb would lead to the reasonable irradiation time if the intensity of the radioactive beam is at least larger than $10^{10}$ pps.

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