Advancing numerics for the Casimir effect to experimentally relevant aspect ratios

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Abstract. Within the scattering theoretical approach, the Casimir force is obtained numerically by an evaluation of the round trip of an electromagnetic wave between the objects involved. Recently [Hartmann M et al 2017, Phys. Rev. Lett. 119 043901] it was shown that a symmetrization of the scattering operator provides significant advantages for the numerical evaluation of the Casimir force in the experimentally relevant sphere-plane geometry. Here, we discuss in more detail how the symmetrization modifies the scattering matrix in the multipole basis and how computational time is reduced. As an application, we discuss how the Casimir force in the sphere-plane geometry deviates from the proximity force approximation as a function of the geometric parameters.

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1. Introduction

When Casimir first derived a force between two objects induced by quantum fluctuations of the electromagnetic vacuum [1], he considered a setup consisting of two infinitely extended, perfectly reflecting parallel plates and obtained a universal expression for the pressure pushing the plates towards each other. In a finite-size system, the Casimir force can thus be increased by increasing the size of the plates.

Instead of the plane-plane geometry, most modern experiments rather employ the sphere-plane geometry, thus avoiding the problem of misalignment. In experiments with a particularly large sphere radius, only a section of a sphere is used. The geometrical parameter characterizing the sphere-plane setup displayed in figure 1 is the aspect ratio $R/L$, i.e. the ratio between the sphere radius $R$ and the minimal distance $L$ between sphere and plane. Note that in figure 1 the sphere radius is chosen to be rather small. The corresponding aspect ratio of $4/3$ is more than two orders of magnitude smaller than common aspect ratios used in experiments. An intuitive idea of a typical aspect ratio is obtained by imagining the ratio of the earth radius and the typical cruising altitude of a commercial jet.

From an experimental point of view, the aspect ratio should be chosen large in order to obtain a sizeable force and to allow for precise measurements which are relevant in various respects. Precise Casimir force measurements are crucial to detect possible deviations from the gravitational interaction at submicrometer distances [2, 3] and thus to exclude or possibly support proposed mechanisms for a fifth fundamental interaction [4, 5]. Precise measurements render the results also sensitive to aspects of the experimental setup beyond the geometrical features, notably the material properties of sphere and plate [6]. This has led to the so-called Drude-plasma controversy. While the materials used in experiments clearly have a finite dc conductivity which can be accounted for within the Drude model, the plasma model with its infinite dc conductivity describes better most Casimir force measurements (for a recent experiment addressing this issue see Ref. [7]).

While on the experimental side, one aims for relatively large aspect ratios, on the theoretical side, the situation is a bit more complicated. The cases of extremely large aspect ratios and of very small aspect ratios can be treated rather easily while intermediate aspect ratios can be very demanding to cope with.

For the theoretical description of experiments, frequently the proximity force approximation (PFA) has been employed. It assumes that the Casimir force between a sphere and a plate can be decomposed into contributions from many small plane-plane segments which are integrated over. As the Casimir force is not additive, this approach can only be approximate. For perfect reflectors at zero temperature, the PFA yields the Casimir force in the sphere-plane geometry

$$F_{\text{PFA}} = -\frac{\pi^3 \hbar c}{360 L^3} R.$$  \hspace{1cm} (1)

The proximity force approximation can be generalized to account for arbitrary electromagnetic response of sphere and plate as well as arbitrary temperatures by using the corresponding expression for the force in the plane-plane geometry. Recently, it has been proven for the sphere-sphere geometry that the proximity force approximation yields the leading small-distance behaviour for arbitrary temperatures and materials [8]. In particular, the proximity force approximation for the sphere-plane geometry is obtained by taking one sphere radius to infinity. Furthermore, Ref. [8] has established a connection between the specular-reflection limit of Mie scattering and the proximity force approximation.

In the opposite limit of small aspect ratios, it is convenient to express the Casimir force in terms of
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We start in Section 2 by explaining how the symmetrization of the round-trip operator became possible by using the appearance of a negative Casimir entropy. Sections 3 and 4 provide technical details allowing to set up the required matrix elements. Numerical aspects related to hierarchical matrices and to the performance achieved by making use of them are given in Section 5. As a physical application, we discuss in Section 6 the size of the corrections to PFA as a function of the geometrical parameters in the sphere-plane geometry before presenting our conclusions in Section 7. Some more technical details are given in the appendices.

2. Symmetrization of the round-trip operator

The Casimir force for the setup originally considered by Casimir, i.e., two parallel perfectly reflecting plates, can be evaluated by summing the vacuum energy over all modes of the electromagnetic field. For more general situations like the experimentally relevant sphere-plane geometry discussed here, the scattering approach has turned out to be very well suited.

Within the scattering approach to the Casimir effect in imaginary frequencies $\omega = i\xi$, the free energy is expressed as a sum 

$$
\mathcal{F} = \frac{k_B T}{2} \sum_{n=-\infty}^{\infty} \log \det \left[ 1 - \mathcal{M}(|\xi_n|) \right] 
$$

over the Matsubara frequencies $\xi_n = 2\pi nk_B T/\hbar$. The round-trip operator

$$
\mathcal{M} = \mathcal{R}_S \mathcal{T}_{SP} \mathcal{R}_P \mathcal{T}_{PS}
$$

represents a complete round-trip of an electromagnetic wave between the sphere and the plane as indicated in Figure 3. The operator $\mathcal{T}_{SP}$ describes a translation from the reference frame of the plane to that of the sphere, and vice versa for $\mathcal{T}_{PS}$. $\mathcal{R}_S$ denotes the reflection at the sphere, while $\mathcal{R}_P$ denotes the reflection at the plane. The Matsubara sum (3) together with (4) holds even if sphere, plane or the medium in between are dissipative.

We express the round-trip operator $\mathcal{M}$ in the multipole basis $|\ell, m, P\rangle$ described by the angular momentum quantum numbers $\ell$ and $m$ ($\ell = 1, 2, \ldots$ and $|m| \leq \ell$). The polarization represents electric multipoles for $P = E$ and magnetic multipoles for $P = M$, respectively. Due to the rotational symmetry of the plane-sphere setup about the $z$-axis, the round-trip operator is diagonal in $m$ and every block $\mathcal{M}^{(m)}$ yields an independent contribution to the free energy

$$
\mathcal{F} = \frac{k_B T}{2} \sum_{n=-\infty}^{\infty} \sum_{m=-\infty}^{\infty} \log \det \left( 1 - \mathcal{M}^{(m)}(|\xi_n|) \right). 
$$

The numerical problems associated with the definition (3) of the round-trip operator become clear from Figure 4, where we depict the values of the matrix elements $\langle \ell_1, m, P_1 | \mathcal{M} | \ell_2, m, P_2 \rangle$ on a
logarithmic colour scale. Even though the matrix elements depend on the parameters chosen here, $R/L = 50$, $\xi(L + R)/c = 1$, $m = 1$, and perfect conductors, the data shown are typical. Already for the relatively small aspect ratio used in figure 4a, the round-trip operator (3) clearly results in an ill-conditioned matrix with elements differing by hundreds of orders of magnitude. As a consequence, a fast and stable numerical evaluation of the determinant becomes extremely difficult. When the determinants in (4) are evaluated, the combination of very small and very large matrix elements can yield contributions of the order one. Small perturbations in the matrix elements may then cause large errors. Furthermore, common computer number formats cover a range of numbers from about $10^{-324}$ to $10^{308}$ [38] which is not sufficient to represent all matrix elements in figure 4a. Instead, one has to use number formats that cover a wider range of numbers, but are also significantly slower.

To overcome these problems, we make use of the fact that the round-trip operator is not uniquely defined. Instead of (3), we choose the symmetrized form for the round-trip operator \( \tilde{M} \) which is not defined. Instead of (3), we choose the symmetrized form for the round-trip operator \( \tilde{M} \) as illustrated in figure 3b. The scattering operator at the sphere \( R_S \) is diagonal in the multipole basis and therefore the matrix square root exists. With this choice for the round-trip operator, the matrix elements take their maximum on the diagonal of each polarization block and decrease away from it, as can be seen in figure 4b. Here, the same parameters have been used as in figure 4a.

In figure 3 we depict a polarization-conserving block (left) and a polarization-mixing block (right) with the matrix elements taken from figure 4, but now on a linear colour scale. This representation emphasizes the fact that a sizeable fraction of the matrix elements off the diagonal is numerically irrelevant. Furthermore, the contribution of the polarization-mixing blocks is significantly smaller than that of the polarization-conserving blocks.

3. Reflection and translation operators

The reflection operator at the sphere is diagonal in the multipole basis
\[
R_S|\ell, m, P\rangle = r^{(S)}_{\ell, P}|\ell, m, P\rangle.
\]
(6)
The reflection coefficients
\[
r^{(S)}_{\ell, E} = -a_\ell, \quad r^{(S)}_{\ell, M} = -b_\ell
\]
(7) are given by the Mie coefficients \( a_\ell \) and \( b_\ell \), where the minus sign is a consequence of the definition of the Mie coefficients employed here [39]. For explicit expressions, we refer the reader to Appendix B.

While the multipole basis is well adapted for the scattering at the sphere, plane waves are better suited to describe the reflection at the plate and the translation between the reference frames of plate and sphere. A plane wave is characterized by the wave vector \( \mathbf{K} \) and the polarization \( p \). Given the special
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4. Matrix elements of the round-trip operator

Having discussed the individual building blocks in the previous section, we now combine them to obtain the matrix elements of the symmetrized round-trip matrix \(^{(3)}M\) in the multipole basis

\[
\hat{M}_{j_1j_2}(P_1, P_2) = \langle \ell_1, m, P_1 | \hat{M}| \ell_2, m, P_2 \rangle \nabla \ktm \int_0^\infty \frac{d^2k}{(2\pi)^2} r_p^{(p)} e^{-2\kappa(L+R)} \times \langle \ell_1, m, P_1 | k, p, + \rangle \langle k, p, -| \ell_2, m, P_2 \rangle .
\]

The last two factors arise due to a change from the multipole basis to the plane-wave basis and vice versa between the translation operators and the reflection operators at the sphere. Explicit expressions for these matrix elements are given in Ref. [11].

We organize the round-trip matrix in the form of a block matrix

\[
\hat{M}^{(m)} = \begin{pmatrix} \hat{M}^{(m)}(E,E) & \hat{M}^{(m)}(E,M) \\ \hat{M}^{(m)}(M,E) & \hat{M}^{(m)}(M,M) \end{pmatrix}
\]

where the diagonal blocks correspond to matrix elements preserving polarization, and the off-diagonal blocks correspond to matrix elements with a change of polarization. Expressing the double integral in [13] in polar coordinates, the integral over the angle can be carried out. After a similarity transform of the round-trip matrix in order to remove phase factors, we finally obtain the matrix elements

\[
\hat{M}^{(m)}_{j_1j_2}(M,M) = \sqrt{b_{j_1}b_{j_2}} \left( A^{(m)}_{j_1j_2,TM} - B^{(m)}_{j_1j_2,TE} \right)
\]

where the corresponding basis \(\{A_{j_1j_2,TE}, B_{j_1j_2,TE}\}\) is then labeled by the projection of the wave-vector \(K\) onto the \(x-y\)-plane, \(k = (k_x, k_y, 0)\), the polarization \(p = TE, TM\), and the direction of propagation \(\phi = \pm 1\) in \(\pm z\)-direction. The latter fixes the sign ambiguity when solving [9] for \(\kappa\).

We are therefore free to always choose \(\kappa\) as positive.

Plane waves propagating in \(-z\)-direction are reflected by the plate

\[
R_p[k, p, -] = r_p^{(P)}(i\xi, k) |k, p, +\rangle
\]

where the reflection coefficients \(r_p^{(P)}\) are the Fresnel coefficients. Explicit expressions are given in Appendix C.

Finally, the translation operators are diagonal in the plane-wave basis

\[
T_{SP}[k, p, +] = e^{-\kappa(L+R)} |k, p, +\rangle, \quad T_{PS}[k, p, -] = e^{-\kappa(L+R)} |k, p, -\rangle
\]

with matrix elements given by exponential factors. Here, we have assumed vacuum between sphere and plane as we will do in the rest of the paper.

\[
\text{role of the z-axis as symmetry axis, we choose the basis vectors for transverse electric (TE) and transverse magnetic (TM) modes as}
\]

\[
\hat{\varepsilon}_{TE} = \frac{\hat{z} \times \hat{K}}{|\hat{z} \times \hat{K}|}, \quad \hat{\varepsilon}_{TM} = \hat{\varepsilon}_{TE} \times \hat{K}.
\]

Here, unit vectors are denoted by a hat. Within the Matsubara formalism, all quantities are expressed in terms of imaginary frequencies \(\xi\). For the wave vector \(K\), we choose an imaginary \(z\)-component \(\kappa\) while the projection \(k\) onto the \(x-y\)-plane is kept real. The dispersion relation then reads

\[
\xi^2 = c^2(\kappa^2 - |k|^2)
\]

with the speed of light \(c\). As the frequency remains constant during a round trip, it is convenient to make use of the angular spectral representation [10].

The corresponding basis \(\{A, B, C\}\) is then labeled by the projection of the wave-vector \(K\) onto the \(x-y\)-plane, \(k = (k_x, k_y, 0)\), the polarization \(p = TE, TM\), and the direction of propagation \(\phi = \pm 1\) in \(\pm z\)-direction. The latter fixes the sign ambiguity when solving [9] for \(\kappa\).

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\]
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A (3) = A 3 \times A 2 \times A 1 \times A 0

Note that for the associated Legendre polynomials \( P^m_n \), we use an uncommon phase convention defined in Appendix A. The dimension of the matrices \([15]–[18]\) is infinite. For a numerical evaluation, the vector space has to be truncated in the angular momentum.

The matrix elements for the round-trip operator \( \mathcal{M} \) differ from those of the symmetrized round-trip operator \( \hat{\mathcal{M}} \) only with respect to the Mie coefficients. While for the symmetrized round-trip operator according to \([15]–[18]\) the matrix elements are proportional to the square root of a product of Mie coefficients with different angular momenta, the matrix elements of \( \mathcal{M} \) are proportional to one Mie coefficient with angular momentum \( \ell_1 \), thus resulting in the numerical problems discussed above.

As the Fresnel coefficients are positive for \( p = TM \) and negative for \( p = TE \), the matrix elements of the round-trip operator \( \hat{\mathcal{M}}(m) \) are positive. Also, the integrals \( A^{(m)}_{\ell_1 \ell_2 \ell_3} \) and \( B^{(m)}_{\ell_1 \ell_2 \ell_3} \) are symmetric with respect to \( \ell_1 \) and \( \ell_2 \), the round-trip matrix is symmetric. Numerical tests suggest that the scattering matrix \( 1 - \hat{\mathcal{M}}(m) \) is diagonally dominant. Together with the positivity of the diagonal entries, it follows that the scattering matrix \( 1 - \hat{\mathcal{M}}(m) \) is positive definite. These properties ensure the stability of the numerical evaluation of the determinant. Small perturbations in the matrix elements thus only cause small changes in the value of the determinant \([42]\).

5. Hierarchical matrices

The symmetrization allows to exploit further properties of the round-trip matrices. As the matrix elements decrease away from the diagonal, one might think that the dominant contribution to the determinant comes from matrix elements close to the diagonal. In fact, it turns out that the matrices \( \mathcal{M}(m) \) are hierarchical off-diagonal low-rank (HODLR) matrices \([43]\). This means that the round-trip matrices can be sub-divided into a hierarchy of rectangular blocks which can be approximated by low-rank matrices.

A low-rank matrix \( M \) of dimension \( N \times N \) can be efficiently approximated by

\[
M \approx UV^T
\]

where \( U \) and \( V \) are matrices of dimension \( N \times p \) with \( p \ll N \). The best rank \( p \) approximation of \( M \) can be obtained using the singular value decomposition \([44]\). Instead of a computationally expensive full singular-value decomposition, low-rank approximations can be computed using fast algorithms like adaptive cross approximation, pseudo-skeletal approximations, interpolatory decompositions or rank revealing QR and LU (see \([43]\) and references therein).

A HODLR matrix \( A \) can be factored into \( n + 1 \) block-diagonal matrices

\[
A \approx A^{(n)} = A_n A_{n-1} \ldots A_0
\]

as sketched in figure 6 for \( n = 3 \). The matrix \( A_n \) consists of \( 2^n \) full-rank blocks around the diagonal while the other matrices \( A_{n-1} \) to \( A_0 \) represent low-rank updates to the identity. The error of this approximation can be made negligibly small by choosing appropriate ranks \( p \).

The factorization \([25]\) allows for a fast computation of the determinant \( A^{(n)} \) for two reasons. Firstly, one can exploit the multiplicativity of determinants. Secondly, the block matrices appearing in \( A_0 \) to \( A_{n-1} \) are of the form

\[
B = \begin{pmatrix} 1 & U_1 V_1^T \\ U_2 V_2^T & 1 \end{pmatrix}
\]

requiring at first sight the evaluation of the determinant of an \( N \times N \)-matrix according to

\[
det(B) = det \left( 1 - U_1 V_1^T U_2 V_2^T \right).
\]

However, exploiting Sylvester’s determinant identity

\[
det(1 + AB) = det(1 + BA),
\]

we obtain

\[
det(B) = det \left( 1 - V_2^T U_1 V_1^T U_2 \right).
\]

It is thus sufficient to evaluate the determinant of a \( p \times p \)-matrix, resulting in a significant speed-up.

In order to assess the numerical advantages of the approach discussed above, we compute the determinants of the scattering matrices either by means of a Cholesky decomposition or the implementation \([45]\) of the algorithm for HODLR matrices described in \([43]\). The Cholesky decomposition factorizes a symmetric positive-definite matrix into the product of a triangular matrix and its transpose allowing for a simple computation of the determinant. The factorization requires \( O(N^3) \) of time for an \( N \times N \) matrix and is about twice as fast as an LU decomposition. The computation of determinants using the HODLR approach takes
6. Corrections to the PFA for the force

As an application of our numerical approach, we consider the corrections to the Casimir force

\[ F = -\frac{\partial F}{\partial L} \]  \hspace{1cm} (30)

Figure 7. Average runtime to compute \( \log \det (1 - M^{(m)}(\xi)) \) as a function of the aspect ratio \( R/L \) for \( m = 1, \xi = c/(L + R) \), and perfect reflectors. The angular momentum is truncated at \( \ell_{\text{dim}} = 5R/L \), yielding a scattering matrix of dimension \( 2\ell_{\text{dim}} \times 2\ell_{\text{dim}} \). Squares and triangles correspond to a computation using Cholesky decomposition and the HODLR algorithm [45], respectively. The lines represent fits and correspond to the asymptotic scaling of the two algorithms. The computations were carried out on an Intel Core i7 with 3.4 GHz. For the Cholesky decomposition LAPACK [46] in combination with ATLAS [47] was used.

\( \mathcal{O}(\rho^2 N \log^2 N) \) steps where, depending on the nature of the problem, \( \rho \) may be a function of \( N \).

In figure 7 we compare the average time to compute \( \log \det (1 - M^{(m)}(\xi)) \) depending on the aspect ratio using the HODLR approach and a Cholesky decomposition. We specifically choose \( m = 1, \xi = c/(L + R) \), and perfect reflectors, but other parameters yield similar results. For aspect ratios \( R/L \lesssim 100 \), both algorithms take about the same time. For larger aspect ratios \( R/L \gtrsim 200 \), the computational time using the Cholesky decomposition scales as \( \propto (R/L)^{2.56} \). This is faster than the theoretical complexity \( \mathcal{O}((R/L)^3) \) of the Cholesky decomposition, because our numerical implementation saves and reuses intermediate values so that the computational time per matrix element is not constant. In contrast, the HODLR algorithm becomes significantly faster for \( R/L \gtrsim 200 \) and the computational time scales only as \( \propto (R/L)^{1.31} \). For the largest aspect ratio displayed in figure 7 \( R/L = 2000 \), we find a speed-up by a factor 33. At even larger aspect ratios, the time required by the Cholesky decomposition becomes prohibitively long.

Figure 8. The correction \( 1 - F/F_{\text{PFA}} \) of the force with respect to the corresponding PFA result is displayed as a function of the radius \( R \) of the sphere and its distance \( L \) from the plane. The numerical data depicted by symbols corresponding to different values of the correction have been obtained for gold at \( T = 300 \text{K} \) as explained in the text. The solid lines indicate the corresponding high-temperature result according to Ref. [50].

\begin{itemize}
  \item \( 1 \% \)
  \item \( 0.5 \% \)
  \item \( 0.25 \% \)
\end{itemize}

For a given value of the correction to the PFA force, the data in figure 8 display three different regimes. In the upper right corner, i.e. for large values of \( L \) and \( R \), the thermal wavelength \( \lambda_T = \hbar c/k_B T \) can be taken to be small, thus indicating the high-temperature regime. In contrast, the low-temperature behaviour is found in the opposite corner, i.e. the lower left part. The two regimes are joined by a transition region visible in the middle of figure 8.

In the high-temperature regime, only the term with \( n = 0 \) in the Matsubara sum (2) contributes and thus only the zero-frequency limit of the dielectric function is relevant. This special case within the Drude model allows for an analytical solution for the
free energy in the sphere-plane setup [50]. As $R$ and $L$ are the only length scales remaining in the high-temperature limit, a fixed value of the correction to the PFA force yields a straight line of slope 1 in the presentation of figure [3]. The high-temperature limits for the chosen values of the correction are indicated by solid lines. As expected, our numerical data approach the high-temperature limit for large sphere radii and large distances between sphere and plane. We note that with decreasing sphere radius, the distance $L$ required for a given value of the correction to the PFA force exceeds the distance obtained within the high-temperature limit. However, decreasing the radius further, the distance falls below the prediction of the high-temperature limit.

7. Conclusions

The standard approach to calculating Casimir forces and free energies within the scattering theory in the multipole basis has been plagued with ill-conditioned round-trip matrices resulting in various numerical problems. We have shown that these problems can be eliminated by a symmetrization of the round-trip operator which physically amounts to splitting the round-trip of an electromagnetic wave between the two scattering objects in the middle of the reflection at one of these objects. The symmetrization of the round-trip operator allows to significantly reduce computational time and thus to perform calculations in the experimentally relevant regime of aspect ratios in the sphere-plane geometry. Furthermore, numerical errors become controllable.

As large aspect ratios are now numerically accessible, it has become possible to assess the quality of the proximity force approximation by determining the deviations from the exact result (cf. figure [3]). This is particularly relevant as the analysis of experimental results so far has relied on the proximity force approximation. Specifically, precise numerical results are of importance in the ongoing Drude-plasma debate. As an example, with the numerical approach detailed here it has become possible recently [32] to demonstrate that the experimental bounds for the Casimir force gradient found in Ref. [24] are violated both for the Drude and the plasma prescription, even though the violation for the latter is found to be significantly larger.

Acknowledgments

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Appendix A. Associated Legendre polynomials

We define the associated Legendre polynomials as [51]

$$P^m_\ell(x) = (x^2 - 1)^{m/2} \frac{d^m}{dx^m} P_\ell(x)$$

(A.1)

where $P_\ell(x)$ denotes the ordinary Legendre polynomial of degree $\ell$. The phase convention employed here differs from the common choice usually made in physics: We omit the Condon-Shortley phase, and change the sign in the first term. With this choice the associated Legendre polynomials are real and non-negative functions for $x \geq 1$.

Appendix B. Mie coefficients

The Mie coefficients are given by [39]

$$a_\ell(x) = (-1)^\ell \frac{\pi}{2} \frac{n^2 s_\ell^{(a)} - s_\ell^{(b)}}{s_\ell^{(c)} + s_\ell^{(d)}},$$

(B.1)

$$b_\ell(x) = (-1)^{\ell+1} \frac{\pi}{2} \frac{s_\ell^{(b)} - s_\ell^{(a)}}{s_\ell^{(c)} + s_\ell^{(d)}},$$

(B.2)

where

$$s_\ell^{(a)}(x) = I_{\ell+\frac{1}{2}}(nx) \left[ x I_{\ell-\frac{1}{2}}(x) - \ell I_{\ell+\frac{1}{2}}(x) \right],$$

(B.3)

$$s_\ell^{(b)}(x) = I_{\ell+\frac{1}{2}}(nx) \left[ nx I_{\ell-\frac{1}{2}}(nx) - \ell I_{\ell+\frac{1}{2}}(nx) \right],$$

(B.4)

$$s_\ell^{(c)}(x) = I_{\ell+\frac{1}{2}}(nx) \left[ x K_{\ell-\frac{1}{2}}(x) + \ell K_{\ell+\frac{1}{2}}(x) \right],$$

(B.5)

$$s_\ell^{(d)}(x) = K_{\ell+\frac{1}{2}}(nx) \left[ nx K_{\ell-\frac{1}{2}}(nx) - \ell I_{\ell+\frac{1}{2}}(nx) \right].$$

(B.6)

The Mie coefficients are evaluated at $x = \xi R/c$ and $n = \sqrt{\epsilon(i\xi)}$ denotes the refractive index. The dielectric function $\epsilon(i\xi)$ is evaluated at imaginary frequencies. The coefficients $s_\ell^{(a)}$, $s_\ell^{(b)}$, $s_\ell^{(c)}$, $s_\ell^{(d)}$, and the numerators in (B.1) and (B.2) are positive independently of $\ell$ and $\xi$.

Appendix C. Fresnel coefficients

The Fresnel coefficients are given by

$$r_{\text{TE}}^{(p)}(i\xi, k) = \frac{e^{i\xi} + \sqrt{c^2 k^2 + \xi^2} [\epsilon(i\xi) - 1]}{e^{i\xi} + \sqrt{c^2 k^2 + \xi^2} [\epsilon(i\xi) + 1]},$$

(C.1)

$$r_{\text{TM}}^{(p)}(i\xi, k) = \frac{\epsilon(i\xi) e^{i\xi} - \sqrt{c^2 k^2 + \xi^2} [\epsilon(i\xi) - 1]}{\epsilon(i\xi) e^{i\xi} + \sqrt{c^2 k^2 + \xi^2} [\epsilon(i\xi) + 1]}.$$

(C.2)
For perfect reflectors they simplify to $r^{(P)}_{TM} = 1$ and $r^{(P)}_{TE} = -1$.

References

[1] Casimir H B G 1948 Proc. K. Ned. Akad. Wet. 51 793

[2] Decca R S, Fischbach E, Klimchitskaya G L, Krause D E, Lópe D and Mostepanenko V M 2003 Phys. Rev. D 68 116003

[3] Decca R S, López D, Fischbach E, Klimchitskaya G L, Krause D E and Mostepanenko V M 2007 Phys. Rev. D 75 071701

[4] Fischbach E and Talmadge C 1992 Nature 356 207

[5] Antoniadis I et al 2011 C. R. Physique 12 755

[6] Klimchitskaya G L, Mohideen U and Mostepanenko V M 2009 Mod. Phys. 81 1827

[7] Bimonte G, López D and Decca R S 2016 Phys. Rev. B 93 184434

[8] Spreng B, Hartmann M, Henning V, Maia Neto P A and Ingold G-L 2018 Proximity force approximation and specular reflection: Application of the WKB limit of Mie scattering to the Casimir effect [arXiv:1803.02254]

[9] Ingold G-L, Umrath S, Hartmann M, Guéroult R, Lambrecht A, Reynaud S and Milton K A 2015 Phys. Rev. E 91 033203

[10] Umrath S, Hartmann M, Ingold G-L and Maia Neto P A 2015 Phys. Rev. E 92 042125

[11] Canaguier-Durand A 2011 Ph.D. thesis (Paris: Université Pierre et Marie Curie)

[12] Masuda M and Sasaki M 2009 Phys. Rev. A 92 062504

[13] Sushkov A O, Kim W J, Dalvit D A R and Lamoreaux S K 2011 Nat. Phys. 7 230

[14] Garcia-Sanchez D, Hong S, Bhaskaran H, Lamoreaux S and Tang H X 2012 Phys. Rev. Lett. 109 027202

[15] van Zwal P J, Palasantzas G, van de Schootbrugge M and De Hosson J Th M 2008 Appl. Phys. Lett. 92 054101

[16] van Zwal P J, Palasantzas G and De Hosson J Th M 2008 Phys. Rev. B 77 075412

[17] de Man S, Hecck K, Wijnhagden R J and Iannuzzi D 2009 Phys. Rev. Lett. 103 040402

[18] Decca R S, Lópe D, Fischbach E and Krause D E 2003 Phys. Rev. Lett. 91 050402

[19] Chan H B, Aksyuk V A, Kleinman R N, Bishop D J and Capasso F 2001 Science 291 1941

[20] Munday J N, Capasso F, Parsegian V A and Bezsukov S M 2008 Phys. Rev. A 78 032109

[21] Munday J N, Capasso F and Parsegian V A 2009 Nature 457 170

[22] Mohideen U and Roy A 1998 Phys. Rev. Lett. 81 4549

[23] Krause D E, Decca R S, López D and Fischbach E 2007 Phys. Rev. Lett. 98 050403

[24] Elzbieciak-Wodka M, Popescu M N, Ruiz-Cabello F J M, Trefalt G, Maroni P and Varmovc M 2014 J. Chem. Phys. 140 104906

[25] Banishev A A, Klimchitskaya G L, Mostepanenko V M and Mohideen U 2013 Phys. Rev. Lett. 110 137401

[26] Jourdan G, Lambrecht A, Comin F and Chevrier J 2009 EPL 85 31001

[27] Chang C-C, Banishev A A, Castillo-Garza R, Klimchitskaya G L, Mostepanenko V M and Mohideen U 2012 Phys. Rev. B 85 165443

[28] Torricelli G, Pirozhkova I, Thornton S, Lambrecht A and Binns C 2011 EPL 93 51001

[29] Ether Jr. D S et al 2015 EPL 112 44001

[30] Garrett J L, Somers D A T and Munday J N 2018 Phys. Rev. Lett. 120 040401

[31] Hartmann M, Ingold G-L and Maia Neto P A 2017 Phys. Rev. Lett. 119 043901

[32] Schleich W 2001 Quantum Optics in Phase Space (Berlin: Wiley-VCH)

[33] For an elementary introduction, see Ingold G-L and Lambrecht A 2015 Am. J. Phys. 83 156

[34] Lambrecht A, Maia Neto P A and Reynaud S 2006 New J. Phys. 8 243

[35] Emig T, Graham N, Jaffe R L and Kardar M 2007 Phys. Rev. Lett. 99 170403

[36] Guéroult R, Ingold G-L, Lambrecht A and Reynaud S 2018 Symmetry 10 37

[37] IEEE Computer Society 2008 IEEE Standard for Floating-Point Arithmetic: IEEE Std 754-2008

[38] Bohren F C and Huffman D R 1983 Absorption and Scattering of Light by Small Particles (Weinheim: Wiley-VCH) ch. 4.

[39] Nieto-Vesperinas M 2006 Scattering and Diffraction in Physical Optics (Singapore: World Scientific)

[40] Messina R, Maia Neto P A, Guizal B and Antezza M 2015 Phys. Rev. A 92 062504

[41] Dailey M, Domico F M and Ye Q 2014 SIAM J. Matrix Anal. Appl. 35 1303

[42] Ambikasaran S and Darve E 2013 J. Sci. Comput. 57 477

[43] Eckart C and Young G 1936 Psychometrika 1 211

[44] Ambikasaran S 2013 A fast direct solver for dense linear systems, https://github.com/sivaramambikasaran/HUDLR

[45] Anderson E et al 1999 LAPACK Users’ Guide 3rd edition (Philadelphia: Society for Industrial and Applied Mathematics)

[46] Whaley R C and Petitet A 2005 Softw. Pract. Exper. 35 101

[47] Palik E (ed) 1998 Handbook of Optical Constants of Solids vol 1 (New York: Academic Press)

[48] Lambrecht A and Reynaud S 2000 Eur. Phys. J. D 8 309

[49] Bimonte G and Emig T 2012 Phys. Rev. Lett. 109 160403

[50] Zhang S and Jin J 1996 Computation of Special Functions (New York: Wiley)