Full counting statistics of the subsystem energy for free fermions and quantum spin chains

K. Najafi and M. A. Rajabpour

Phys. Rev. B 96, 235109 — Published 7 December 2017
DOI: 10.1103/PhysRevB.96.235109
Full counting statistics of the subsystem energy in the free fermions and the quantum spin chains

K. Najafi\textsuperscript{1} and M. A. Rajabpour\textsuperscript{2}

\textsuperscript{1} Department of Physics, Georgetown University, 37th and O Sts. NW, Washington, DC 20057, USA
\textsuperscript{2} Instituto de Física, Universidade Federal Fluminense, Av. Gal. Milton Tavares de Souza s/n, Gragoatá, 24210-346, Niterói, RJ, Brazil

Dated: October 29, 2017

We calculate the full counting statistics (FCS) of a subsystem energy in free fermionic systems by means of the Grassmann variables. We demonstrate that the generating function of these systems can be written as a determinant formula with respect to the Hamiltonian couplings and by using the Bell’s polynomials, we derive exact formulas for the subsystem energy moments. In addition, we discuss the same quantities in the quantum XY spin chain, and we demonstrate that at the critical regimes the fluctuations of the energy moments decay like a power-law as we expect from the conformal field theory arguments, while in non-critical regimes, the decay is exponential. Furthermore, we discuss the full counting statistics of subsystem energy in the quantum XX chain.

I. INTRODUCTION

We have removed the part regarding the uncertainty principle, instead we have added the following parts in red and some recent studies about the full counting statistics of energy in other systems.

In quantum many-body physics, the fluctuations of local observables carry fair amount of information regarding the physical properties of the system. The rapid progress in the manipulation of quantum devices in condensed matter physics has made it possible to investigate the nature of fluctuations in quantum many-body systems to even higher degrees of details. In quantum mechanics, one usually is interested in the first few moments of the fluctuations while the full description of the system requires the knowledge of the full distribution function of the observable. This kind of distribution functions are examples of a more general concept known as full counting statistics (FCS) which studies the full distribution of a macroscopic observable in arbitrary systems. The term FCS was first popularized in the study of charge transport in mesoscopic systems\textsuperscript{1–9}. However, since this quantity can be defined naturally in any quantum system, it has been also applied in different areas such as Fermi edge problems\textsuperscript{10}, ultracold atoms\textsuperscript{11–13}, quantum chains\textsuperscript{14–20}, and many-body localization\textsuperscript{21,22}. For a good list of applications and references, see Ref.\textsuperscript{15,23}. In the context of quantum chains, one can simply consider the system in its ground state and then study the full distribution of the number of fermions\textsuperscript{14–16} or magnetization in the subsystem\textsuperscript{17–20,25–27} and find the relevant universal scaling functions\textsuperscript{28–30}. One of the natural quantities that can be defined for any system is the energy of the subsystem. The Hamiltonian truncated to a subsystem does not commute with the Hamiltonian of the full system. Consequently, if the full system is in its ground state, the energy of the subsystem can have many different values with different probabilities. Surprisingly, to the best of our knowledge, the FCS of the subsystem energy has not been studied in the literature. However, in Ref.\textsuperscript{31}, related problems regarding the relation between probability distributions of the measurement outcomes of the local and global Hamiltonians have been discussed. For the distribution of the energy of the full system in the Gibbs state and possible experimental methods to detect it, see Ref.\textsuperscript{32}. There have also been some works regarding the full counting statistics of both energy transport of phonons\textsuperscript{34–36} and electron energy transfer\textsuperscript{37–39}. In this article, we study the FCS of the subsystem energy in generic free fermion Hamiltonian with real couplings, and in particular, in quantum spin chains that can be mapped to the free fermions.

To have the full distribution of the system, one usually needs to calculate the generating function of the distribution functions. In the case of free fermions, the studies of those observables that have quadratic form in the fermionic representation (for example fermion number) always lead to determinant formulas\textsuperscript{1,16,23}. The Hamiltonian truncated to a subsystem in free fermions naturally has a quadratic form and one expects a determinant formula. In this article, we will find a determinant formula in a form which is different but equivalent to the ones that can be derived from earlier approaches, see for example Ref.\textsuperscript{23}. We will implement the fermionic coherent states written in Grassmann representation which is also practical when the state of the system is generic but has a simple form in the local basis. Afterwards, using the Bell’s polynomials, we will provide exact formulas for the moments. In section III, we will apply our determinant formulas to study the FCS of subsystem energy in the quantum XY chain and will remark on different properties of this quantity on different regions of the phase diagram. In particular, we will study the universal properties in the critical regions. Finally, in section IV, we will summarize our results and future directions.
II. FULL COUNTING STATISTICS OF THE
SUBSYSTEM ENERGY IN THE GENERIC FREE
FERMIONS

In this section, we study the full counting statistics of the subsystem energy in the ground state of a generic real free fermion Hamiltonian in an arbitrary dimension. To start, we calculate the first two moments and then provide an exact formula for the generating function of the subsystem energy. Afterward, we show how one can calculate all of the moments exactly by a proper expansion of the generating function. Consider a generic free fermion Hamiltonian with the following truncated Hamiltonian for the subsystem $D$:

$$H_D = c^\dagger A c + \frac{1}{2} c^\dagger B c + \frac{1}{2} c B^T c - \frac{1}{2} t r A,$$  (1)

where $B^T$ is the transpose of $B$ and $c$ is a vector with elements $c_i$ and $i = 1, 2, ..., |D|$, where $|D|$ is the number of sites in the region $D$. Moreover, $A$ and $B$ are symmetric and antisymmetric matrices respectively to keep the Hamiltonian Hermitian. Note that here we just consider real $A$ and $B$. We are interested in calculating the full counting statistics of energy defined as the following expectation value

$$M(\bar{\lambda}_D) = \langle e^{\bar{\lambda}_D H_D} \rangle = t r [\rho_D e^{\bar{\lambda}_D H_D}],$$  (2)

where $\bar{\lambda}_D = \frac{\lambda}{\sqrt{\langle H_D \rangle}}$. For later implications, it is convenient to introduce new operators defined as,

$$a_i = c_i^\dagger + c_i, \quad b_i = c_i^\dagger - c_i,$$  (3)

with the anticommutator relations $\{a_i, a_j\} = 2 \delta_{ij}, \{b_i, b_j\} = -2 \delta_{ij}$, and $\{a_i, b_j\} = 0$, and the well known correlations defined as block Green matrices,

$$G^{(ba)}_{ij} = t r [\rho_D b_i a_j], \quad G^{(ab)}_{ij} = t r [\rho_D a_i b_j], \quad G^{(aa)}_{ij} = t r [\rho_D a_i a_j], \quad G^{(bb)}_{ij} = t r [\rho_D b_i b_j],$$  (4)

where $\rho_D$ is the reduced density matrix of the region $D$ and $G^{(ab)}_{ij} = -G^{(ba)}_{ji}$. Note that the role of the Hamiltonian of the full system is encoded in $\rho_D$. By substituting operators $a$ and $b$ in the Hamiltonian (1), $H_D$ becomes,

$$H_D = \frac{1}{2} [b_i (A_{ij} + B_{ij}) a_j],$$  (5)

where we have used $G^{(aa)} + G^{(aa)T} = -G^{(bb)} + G^{(bb)T} = 2I$. By using the definition of these correlators, we get

$$\langle H_D \rangle = \frac{1}{2} t r [(A - B) G^{(ba)}] = \frac{1}{4} t r [D^T G^{(ba)}],$$  (6)

where we have introduced a new matrix $D = 2(A + B)$. Similarly, by using the Wick’s theorem one can also calculate the $\langle H_D^2 \rangle$. The final result can be written as:

$$\langle H_D^2 \rangle = \langle H_D \rangle^2 + \frac{1}{16} t r [D D^T]$$

$$- \frac{1}{16} t r [D^T G^{(ba)} D^T G^{(ba)}].$$  (7)

It is needless to say that the calculation of the higher moments starts to get cumbersome immediately which makes the forthcoming calculations much more valuable. In any case, the above two direct calculations can be used to check the validity of the general results.

A. Generating function of the subsystem energy in the ground state

In this subsection, we come back to our problem of calculating the generating function of the subsystem energy for a system which is in its ground state. To do that we can use the fermionic coherent state defined as,

$$|\xi > = |\xi_1, \xi_2, ..., \xi_N > = e^{-\sum_{i=1}^{N} \xi_i c_i^\dagger |0 >},$$  (8)

where $\xi_i$’s are Grassmann numbers which satisfy the following properties: $\xi_i \xi_m + \xi_m \xi_n = 0$ and $\xi_n^2 = \xi_n = 0$. Consequently, we can show

$$c_i |\xi > = \xi_i |\xi >.$$  (9)

By using the Grassmann variables the reduced density matrix can be written as$^{40, 41}$

$$\rho_D(\xi, \xi') = < \xi | \rho_D | \xi' > = t r \left[ (\xi - G^{(ba)}) | e^{\frac{1}{2} (\xi - \xi')^T F (\xi + \xi')} \right],$$  (10)

where we have introduced the matrix $F = (G^{(ba)} + \mathbb{1})(\mathbb{1} - G^{(ba)})^{-1}$. The trace in the context of Grassmann variables can be calculated as:

$$t r \circ O = \int d\xi d\xi' e^{-\xi \xi} < -\xi | O | \xi >.$$  (11)

Then the equation (2) can be written as:

$$M(\bar{\lambda}_D) = \langle e^{\bar{\lambda}_D H_D} \rangle = \int d\eta d\eta' e^{-\eta \eta} < -\eta | \rho_D e^{\bar{\lambda}_D H_D} | \eta >.$$  (12)

After using the identity

$$I = \int d\xi d\xi' e^{-\xi \xi} | \xi > \langle \xi |$$  (13)

we get

$$M(\bar{\lambda}_D) = \int d\eta d\eta' \int d\xi d\xi' e^{-\eta \eta' - \xi \xi}$$

$$< -\eta | \rho_D | \xi > | \xi e^{\bar{\lambda}_D H_D} | \eta >.$$  (14)
We can now calculate the two expectations separately. To calculate the second expectation, first we decompose $e^{\lambda_D H_D}$ by means of the Balian-Brezin formula \(^{22}\) as:

$$e^{\lambda_D H_D} = e^{\frac{1}{2} c^\dagger X c^\dagger Y c - \frac{i}{2} X^\dagger Y e^{\frac{1}{2} e^c Z c}} ,$$  

(15)

where $X, Y, Z$ can be calculated from the blocks of matrix $T$ defined as

$$T = e^{\frac{1}{2} c^\dagger X c^\dagger Y c} = \begin{pmatrix} T_{11} & T_{12} \\ T_{21} & T_{22} \end{pmatrix} .$$  

(16)

which leads to

$$X = T_{12} (T^{-1}_{22}), \quad Z = (T^{-1}_{22}) T_{21}, \quad e^{\gamma Y} = T_{22}^T.$$  

(17)

By using the above formulas and the properties of the fermionic coherent states, we have

$$\langle \xi | e^{\lambda_D H_D} | \eta \rangle = e^{\frac{1}{2} \xi \bar{X} \xi + \frac{1}{2} \eta \bar{Z} \eta - \frac{i}{2} \bar{X}^\dagger Y + \xi e^\gamma \eta} .$$  

(18)

After implementing the above formula in the equation (14) we get,

$$M(\tilde{\lambda}_D) = \det[\frac{1}{2}(1 - G^{(ba)})] e^{-\frac{i}{2} \bar{Y} Y} \int d\bar{\eta} \int d\bar{\xi} e^{\frac{1}{2} \xi \bar{X} \xi + \frac{1}{2} \eta \bar{Z} \eta + \xi e^\gamma \eta} \int d\nu e^{-\nu F \nu - \frac{1}{2} \nu \bar{\nu} - \frac{1}{2} \bar{\nu} \nu} .$$  

(19)

To calculate the integrals, we first introduce new variables, $\nu = \frac{\bar{\eta} + \bar{\xi}}{\sqrt{2}}$ and $\bar{\nu} = \frac{\bar{\eta} - \bar{\xi}}{\sqrt{2}}$, then the integral becomes,

$$M(\tilde{\lambda}_D) = (-1)^{|D|} \det[\frac{1}{2}(1 - G^{(ba)})] e^{-\frac{i}{2} \bar{Y} Y} \int d\bar{\eta} \int d\bar{\xi} e^{\frac{1}{2} \xi \bar{X} \xi + \frac{1}{2} \eta \bar{Z} \eta + \xi e^\gamma \eta} \int d\nu e^{-\nu F \nu - \frac{1}{2} \nu \bar{\nu} - \frac{1}{2} \bar{\nu} \nu} .$$  

(20)

Notice that we get an extra factor of $(-1)^{|D|}$, where $|D|$ is the number of cites in the region $D$, which is the Jacobian for the change of variables of $\xi$ and $\bar{\eta}$. Using the formula for Gaussian integrals, we can also calculate the second integral,

$$M(\tilde{\lambda}_D) = \det[F] \det[\frac{1}{2}(1 - G^{(ba)})] e^{-\frac{i}{2} \bar{Y} Y} \int d\bar{\eta} \int d\bar{\xi} e^{\frac{1}{2} \xi \bar{X} \xi + \frac{1}{2} \eta \bar{Z} \eta + \xi e^\gamma \eta} \int d\nu e^{-\nu F \nu - \frac{1}{2} \nu \bar{\nu} - \frac{1}{2} \bar{\nu} \nu} .$$  

(21)

The $(-1)^{|D|}$ cancels out as the det $[F]$ also brings down $(-1)^{|D|}$ factor. Furthermore, we introduce new variables by defining $\bar{\alpha} = \frac{\bar{\xi} + \bar{\eta}}{\sqrt{2}}$ and $\alpha = \frac{\bar{\xi} - \bar{\eta}}{\sqrt{2}}$,

$$M(\tilde{\lambda}_D) = \det[F] \det[\frac{1}{2}(1 - G^{(ba)})] e^{-\frac{i}{2} \bar{Y} Y} \int d\alpha \int d\bar{\alpha} e^{\frac{1}{2} \bar{\alpha} \bar{\gamma} \bar{\alpha} - \frac{1}{2} \alpha \bar{\gamma} \alpha - \bar{\gamma} \bar{\alpha} \alpha - \bar{\alpha} \bar{\gamma} \alpha} .$$  

(22)

Since $e^\gamma$ is symmetric, the two terms $\bar{\alpha} \bar{\gamma} \bar{\alpha}$ and $\alpha \bar{\gamma} \alpha$ have no contribution in the integrals. Finally, after performing the integrals, we have

$$M(\bar{\lambda}_D) = e^{-\frac{i}{2} \bar{Y} Y} \det[\frac{1}{2}(1 - G^{(ba)})] \det[\tilde{\lambda}_D - FX + F e^\gamma] .$$  

(23)

Using the definition of $F = (G^{(ba)} + 1)^{(1 - G^{(ba)})^{-1}}$, we get

$$M(\bar{\lambda}_D) = \det[e^\gamma] \frac{1}{2} \det[\frac{1}{2} - \frac{1}{2} G^{(ba)} e^\gamma + \frac{1}{2} (1 - X^\dagger e^\gamma)] .$$  

(24)

where we have used the identity $[e^{-\frac{i}{2} \bar{Y} Y}]^\frac{1}{2} = (\det[e^\gamma])^\frac{1}{2}$. By introducing the definition of $X$ and $Y$ we can also write:

$$M(\bar{\lambda}_D) = \det[T_{22}]^{-\frac{1}{2}} \det[\frac{1}{2} - G^{(ba)} T_{22} + \frac{1}{2} (1 - T_{12})] .$$  

(25)

Although the above three equations are equivalent to the formula derived in earlier works, see for example Ref.\(^ {23}\), they have different forms. We will use the current forms to calculate the cumulants in the next sections. Note that one can consider the above equations as the generating function of any observable defined in the subsystem $D$ that can be written as the quadratic equation (1). Consequently, all the subsequent formulas are valid for any observable that can be written in a quadratic form in the fermionic representation.

1. $T$ matrix expansion

One can determine the exact form of the matrix $T$ by expanding it with respect to $\lambda_D$ as follows:

$$T = 1 + \lambda_D \begin{pmatrix} A & B \\ -B & -A \end{pmatrix} + \frac{\lambda_D^2}{2!} \begin{pmatrix} A & B \\ -B & -A \end{pmatrix}^2 + \frac{\lambda_D^3}{3!} \begin{pmatrix} A & B \\ -B & -A \end{pmatrix}^3 + ....$$  

(26)

By calculating the right-hand side of the above equation, we can see that there are only two independent block matrices which can be obtained by using the following formulas:

$$T_{12} = \sum_{n=1}^{\infty} \frac{\lambda_D^n}{n!} r^{(1)}_n ,$$  

(27)

$$T_{22} = 1 + \sum_{n=1}^{\infty} \frac{\lambda_D^n}{n!} r^{(2)}_n .$$  

(28)

Then for the other two matrices we simply have

$$T_{11} = T_{22} (\bar{\lambda}_D \rightarrow -\bar{\lambda}_D) ,$$  

(29)
To start, we need to use the expansion of functions of
and calculate all the moments using the equation (25).

\[ \sum_{j=1}^{k/2} n_{2j-1} = \text{even}, \]

\[ \tau_{\text{even}}^{(1)} = \sum_{\{n_i\}} A^{n_1} B^{n_2} \ldots A^{n_k} B^{n_k} \text{sgn}(k), \]

where

\[ \sum_{j=1}^{k/2} n_{2j-1} = \text{odd}, \]

and

\[ \tau_{\text{odd}}^{(2)} = -\sum_{\{n_i\}} A^{n_1} B^{n_2} \ldots A^{n_k} B^{n_k} \text{sgn}(k), \]

where \( \sum_{i=1}^{k} n_i = n \) and there are \( 2^{n-1} \) terms. Moreover, the sign of the terms can be calculated by

\[ \text{sgn}(k) = (-1)^{\sum_{x=0}^{n-2} [\frac{n_{k-j} + 1 - (\frac{n_{k-1}}{2})]}{n_{k-1}}], \]

where \([x]\) is the floor (the largest integer less than or equal to \( x \)) and for \( s < 1 \), we have \( \sum_{j=0}^{s-1} n_{k-j} = 0 \), for more details see Appendix A.

2. Subsystem energy moments from the generating function

Since the exact form of \( T_{22} \) and \( T_{12} \) are known, one can calculate all the moments using the equation (25).

To start, we need to use the expansion of functions of determinant provided in Ref. 43. Then we have

\[ \det[T_{22}] = 1 + \sum_{k=1}^{\infty} t_k \! \tilde{\lambda}_D \frac{k}{k!}, \]

where

\[ t_k = \sum_{j=1}^{k} f_j B_{kj}(g); \]

and \( f_j = \delta^j \) and \( B_{kj}(g) \) is the partial exponential Bell's polynomial (see Appendix B) defined as

\[ \frac{1}{j!} \left( \sum_{k=1}^{\infty} g_k e^k \right)^j = \sum_{j=0}^{\infty} B_{kj}(g) \frac{e^k}{k!}. \]

Here, we list the first few terms,

\[ t_0 = 1, \]

\[ t_1 = f_1 g_1, \]

\[ t_2 = f_1 g_2 + f_2 g_1^2, \]

\[ t_3 = f_1 g_3 + f_2 (3g_1 g_2) + f_3 g_1^3, \]

\[ t_4 = f_1 g_4 + f_2 (4g_1 g_3 + 3g_2^2) + f_3 (6g_1 g_2 g_1) + f_4 g_1^4, \]

and \( g = (g_1, g_2, \ldots) \) with

\[ g_k = (\sum_{j=1}^{k} (\! -1\! )^{j-1} (j - 1)! \text{tr} B_{kj}(\tau^{(2)}), \]

where \( \tau^{(2)} = (\tau_1^{(2)}, \tau_2^{(2)}, \ldots) \). The \( \text{tr} B_{kj}(\tau^{(2)}) \) can be evaluated by calculating \( B_{kj}(g) \) and then symmetrization of all the terms \( G_1 G_2 \ldots G_r \to \frac{1}{r!} \sum_r G_{\pi_1} G_{\pi_2} \ldots G_{\pi_r} \), where the \( \pi_i \)’s are any sequence of the \( \{g_k\} \) and the sum is over all permutations. After having a symmetrized form for \( B_{kj}(g) \), we can now replace \( g_k \) with \( \{\tau^{(2)}_k\} \) and derive the formulas for \( \text{tr} B_{kj}(\tau^{(2)}) \). Here, we list a few of the coefficients,

\[ g_1 = \text{tr} \tau_1^{(2)}, \]

\[ g_2 = \text{tr} \tau_2^{(2)} - (\tau_1^{(2)})^2, \]

\[ g_3 = \text{tr} \tau_3^{(2)} - 3\tau_1^{(2)} \tau_2^{(2)} + 2(\tau_1^{(2)})^3, \]

\[ g_4 = \text{tr}\tau_4^{(2)} - 4\tau_2^{(2)} \tau_2^{(2)} - 3(\tau_2^{(2)})^2 + 12(\tau_1^{(2)})^2 \tau_2^{(2)} - 6(\tau_2^{(2)})^4. \]

Note that in our case, \( \delta = -\frac{1}{2} \). We need to also calculate the second determinant in the equation (25). Inside the determinant can be written as:

\[ \frac{1}{2} - \frac{G^{(ba)}}{2} T_{22} + \frac{1}{2} + \frac{G^{(ba)}}{2} (1 - T_{12}) = \]

\[ \frac{1}{2} - \frac{G^{(ba)}}{2} \left( 1 + \sum_{n=1}^{\infty} \frac{\lambda^n}{n!} \tau_n^{(2)} \right) + \]

\[ \frac{1}{2} + \frac{G^{(ba)}}{2} \left( 1 - \sum_{n=1}^{\infty} \frac{\lambda^n}{n!} \tau_n^{(1)} \right) \]

\[ = 1 + \sum_{n=1}^{\infty} \frac{\lambda^n}{n!} \left( \frac{1}{2} - \frac{G^{(ba)}}{2} \tau_n^{(2)} - \frac{1}{2} + \frac{G^{(ba)}}{2} \tau_n^{(1)} \right). \]

Then, we define the following new matrices:

\[ \tilde{\tau}_n = \frac{1}{2} - \frac{G^{(ba)}}{2} \tau_n^{(2)} - \frac{1}{2} + \frac{G^{(ba)}}{2} \tau_n^{(1)} \]

We can now calculate \( \det[1 + \sum_{n=1}^{\infty} \frac{\lambda^n}{n!} \tilde{\tau}_n] \) as before with the condition \( \delta = 1 \) which implies \( f_j = 1 \). The expansion has the following form:

\[ \det[1 + \sum_{n=1}^{\infty} \frac{\lambda^n}{n!} \tilde{\tau}_n] = 1 + \sum_{k=1}^{\infty} t_k \frac{\tilde{\lambda}_D}{k!}. \]
where
\[ \tilde{t}_k = \sum_{j=1}^{k} B_{kj}(\tilde{g}). \] (43)

The first few terms have the following forms:
\[ \begin{align*}
\tilde{t}_0 &= 1, \\
\tilde{t}_1 &= \tilde{g}_1, \\
\tilde{t}_2 &= \tilde{g}_2 + \tilde{g}_1^2, \\
\tilde{t}_3 &= \tilde{g}_3 + 3\tilde{g}_1\tilde{g}_2 + \tilde{g}_1^3, \\
\tilde{t}_4 &= \tilde{g}_4 + 4\tilde{g}_1\tilde{g}_3 + 3\tilde{g}_2^2 + 6\tilde{g}_1^2\tilde{g}_2 + \tilde{g}_1^4,
\end{align*} \] (44a) (44b) (44c) (44d) (44e)

with
\[ \begin{align*}
\tilde{g}_1 &= \text{tr} \tilde{\tau}_1, \\
\tilde{g}_2 &= \text{tr} [\tilde{\tau}_2 - \tilde{\tau}_1^2], \\
\tilde{g}_3 &= \text{tr} [3\tilde{\tau}_3 - 3\tilde{\tau}_1\tilde{\tau}_2 + 2(\tilde{\tau}_1)^3], \\
\tilde{g}_4 &= \text{tr} [\tilde{\tau}_4 - 4\tilde{\tau}_1\tilde{\tau}_3 - 3(\tilde{\tau}_2)^2 + 12(\tilde{\tau}_1)^2\tilde{\tau}_2 - 6(\tilde{\tau}_1)^4].
\end{align*} \] (45a) (45b) (45c) (45d)

Using the two formulas (A2) and (42) the expansion for \( M(\lambda_D) \) becomes
\[ M(\lambda_D) = \sum_{m=0}^{\infty} \left( \frac{\lambda_D^m}{m!} \right) \left( \sum_{j=0}^{m} \frac{\lambda_D^j}{j!} \right) \sum_{m=0}^{\infty} \frac{\lambda_D^m}{j!(m-j)!} t_j \tilde{t}_{m-j}, \] (46)

which simplifies to
\[ M(\lambda_D) = \sum_{m=0}^{\infty} \frac{\lambda_D^m}{m!} \sum_{j=0}^{m} \binom{m}{j} t_j \tilde{t}_{m-j}. \] (47)

An expansion of the full counting statistics with respect to \( \lambda \) can be written as below:
\[ M(\lambda_D) = 1 + E_1 \lambda_D + E_2 \frac{\lambda_D^2}{2!} + ..., \] (48)

where we have \( E_m = \langle H_D^m \rangle \).

One can directly check that the above equation for \( m = 1 \) and \( m = 2 \) reproduces the equations (6) and (7) respectively. For the future discussion, it is also important to introduce the fluctuation of the \( m \) moment defined as:
\[ \tilde{E}_m = \langle (H_D - E_1)^m \rangle. \] (50)

Consequently, by using the equation (49) and working out some algebra, one can show that
\[ \begin{align*}
\tilde{E}_1 &= 0, \quad (51a) \\
\tilde{E}_2 &= g_2 - \frac{g_2^2}{2}, \quad (51b) \\
\tilde{E}_3 &= g_3 - \frac{g_2^3}{2}, \quad (51c) \\
\tilde{E}_4 - 3\tilde{E}_2^2 &= g_4 - \frac{g_2^4}{2}, \quad (51d) \\
\tilde{E}_5 - 10\tilde{E}_2\tilde{E}_3 &= g_5 - \frac{g_2^5}{2}. \quad (51e)
\end{align*} \] (51)

Finally, the most general case can be written as:
\[ \tilde{E}_m + \sum_{j=2}^{m} (-1)^{j-1}(j-1)!B_{mj}(\tilde{E}_1, \tilde{E}_2, \ldots, \tilde{E}_{m-j+1}) = g_m - \frac{g_m}{2}, \quad m > 1 \] (52)

where the \( B_{mj} \) is the Bell’s polynomial as we mentioned before. The above formulas show that the quantities \( g_i \) and \( \tilde{g}_i \) with \( i > 2 \) represent the fluctuations of the subsystem energy.

### B. Subsystem energy generating function for an arbitrary state

In this section, we study the subsystem energy generating function for an arbitrary state. In other words, we would like to calculate
\[ M(\lambda_D) = \langle \psi \mid e^{\lambda_D H_D} \mid \psi \rangle, \] (53)

for arbitrary state \( |\psi\rangle \). To calculate the above quantity, we assume that we know the form of \( |\psi\rangle \) in the fermion occupation basis, in other words, we have
\[ |\psi\rangle = \sum_{\{C\}} a_C |C\rangle, \] (54)

where \( |C\rangle \) is an arbitrary configuration for fermions in the subsystem. Using the above state in the equation(53), we have
\[ M(\lambda_D) = \sum_{\{C\}} \sum_{\{C'\}} a_C^* a_C \langle C' | e^{\lambda_D H_D} | C \rangle. \] (55)

The quantity \( \langle C' | e^{\lambda_D H_D} | C \rangle \) can be calculated using the equation (18) as follows (see Ref.45): Consider an arbitrary configuration \( C \). Then, the corresponding Grassmann variable for the unoccupied state is zero while for occupied state, one needs to integrate over Grassmann variable. This transformation will effectively create new matrices \( X_{C'}, Z_C \) and \( (e^\lambda C)_{C'C} \) that are dependent on the configurations. Then we can write
\[ M(\lambda_D) = \sum_{\{C'\}} \sum_{\{C\}} a_C^* a_C \int C' d\xi \int C \eta d\bar{\eta} e^{\frac{i}{2} \xi X_{C'} \xi + \frac{i}{2} \eta Z_C \eta - \frac{i}{2} \bar{\eta} Y_{C'C} + (\xi^\dagger C')_{C'C} \eta}. \] (56)
Finally, after performing the Grassmann integration, we have
\[
M(\lambda_D) = e^{-\frac{i}{2}tY} \sum_{\{c^+_C\} \{c^+_C\}} a_{c^+_C} a_C \text{pf} \left[ \begin{array}{cc} X_{c^+_C} & (e^Y)_{c^+_C} \\ -(e^Y)^T_{c^+_C} & -Z_C \end{array} \right],
\]
where \(\text{pf}\) is the Pfaffian of the matrix. The above equation can be very useful for those states that have simple form in the configuration basis. It can be also useful in the study of time dependent systems. For example, for the state without any fermion, we simply have
\[
M(\lambda_D) = e^{-\frac{i}{2}tY},
\]
and for the case with full of fermions, we have
\[
M(\lambda_D) = e^{-\frac{i}{2}tY} \text{pf} \left[ \begin{array}{cc} X & e^Y \\ -(e^Y)^T & -Z \end{array} \right].
\]

III. SUBSYSTEM ENERGY STATISTICS IN THE XY SPIN CHAIN

In this section, we use the equations that we derived in the previous section to study the full counting statistics of the subsystem energy for the XY spin chain. We study different phases of the chain with analytical and numerical techniques. In particular, we study the transverse field Ising chain and XX chain in more details.

A. Definitions and general results

The Hamiltonian of the XY-chain is as follows
\[
H_{XY} = -\frac{J}{2} \sum_{j=1}^{L} \left[ \left( 1 + \frac{a_j^+ a_{j+1}^x}{} \right) a_j^+ a_{j+1}^+ + \left( 1 - \frac{a_j^+ a_{j+1}^y}{} \right) a_j^+ a_{j+1}^+ \right] - h \sum_{j=1}^{L} \sigma_j^z,
\]
where \(\sigma_j^{x,y,z}\) are the Pauli matrices. Here, \(a\) indicates the anisotropy interaction between spins and \(h\) denotes the transverse magnetic field. Using the Jordan-Wigner transformation \(c_j^+ \equiv \prod_{k<j} \sigma_k^+ \sigma_j^+\), one can map the Hilbert space of a quantum chain of a spin 1/2 into the Fock space of spinless fermions. Then, the new Hamiltonian becomes
\[
H = -\frac{J}{2} \sum_{j=1}^{L-1} (c_j^+ c_{j+1} + a_j^+ a_{j+1}^+ + h.c.) - \sum_{j=1}^{L} h(c_j^+ c_j - \frac{1}{2}),
+ \frac{\sqrt{2}}{2} \left[ c_1^+ c_1 + a_1^+ a_1^+ + h.c. \right).
\]

\[
[H_L, H] = \frac{J^2}{4} \left[ -c_{l+2}^+ c_{l+2} c_{l} + c_{l+2}^+ c_{l} c_{l-1} - c_{l+1}^+ c_{l-1} - c_{l+1} c_{l+1} + c_3 \right].
\]

The commutator is dependent just on the sites that are close to the boundary which indicates that by increasing the size of the subsystem, say \(L = 6\), the does not change. Note that since in this work we are considering the full system in its ground state, uncertainty equations corresponding to the above commutator is not that much useful. However, if the system starts with a state which is not the eigenstate of the total Hamiltonian, as we always have in the quantum quench setup, the corresponding uncertainty relations might be a useful quantity to quantify the fluctuations of the energy of the subsystem. In this paper, we will not discuss this
question further.” It is easy to see that the Hamiltonian (62) can be written in the following form,
\[ H_l = c_1 A c + \frac{1}{2} c_1^† B c + \frac{1}{2} c B^† c - \frac{1}{2} \text{Tr} A, \] (64)
where the matrices \( A \) and \( B \) are:
\[
A = \begin{pmatrix}
-h & \frac{1}{2} & 0 & \ldots & 0 \\
\frac{1}{2} & -h & \frac{1}{2} & 0 & 0 \\
0 & \frac{1}{2} & -h & \frac{1}{2} & 0 \\
\vdots & \ddots & \ddots & \ddots & \ddots \\
0 & 0 & \ldots & \frac{1}{2} & -h
\end{pmatrix},
\]
\[
B = \begin{pmatrix}
0 & \frac{j_a}{2} & 0 & \ldots & 0 \\
\frac{j_a}{2} & 0 & \frac{j_a}{2} & 0 & 0 \\
0 & -\frac{j_a}{2} & 0 & \frac{j_a}{2} & 0 \\
\vdots & \ddots & \ddots & \ddots & \ddots \\
0 & 0 & \ldots & -\frac{j_a}{2} & 0
\end{pmatrix}. \quad (65)
\]

Using the results of the previous section and considering \( J = 1 \), first one can calculate \( \langle H_l \rangle \) as follows:
\[
\langle H_l \rangle = \frac{1}{4} \text{tr}[D^T G^{(ba)}] = \frac{1}{4} \left[ -2h \delta_{nm} + (1 + a) \delta_{n,m-1} + (1 - a) \delta_{n,m+1} \right] G^{(ba)}_{nm} + \frac{1}{4} G^{(ba)}_{nn} + \frac{1}{4} G^{(ba)}_{mm}
\]
\[
+ \frac{1}{4} \left[ -2h G^{(ba)}_{00} + (1 + a) G^{(ba)}_{11} - (1 - a) G^{(ba)}_{-1} \right], \quad (66)
\]
Notice that in this case \( G^{(aa)} = -G^{(bb)} = 1 \). Not surprisingly the \( \langle H_l \rangle \) is proportional to the size of the subsystem. We can also calculate the second moment as follows:
\[
\langle H_l^2 \rangle = \langle H_l \rangle^2 + \frac{1}{16} \text{tr}[D^T D G^{(ba)}] - \text{tr}[D^T G^{(ba)} D^T G^{(ba)}]], \quad (67)
\]
where we have \( \text{tr}[D^T D G^{(ba)}] = (2J(1-a^2) + 4h^2)l - 2J(1-a^2) \).
It is easy to see that here, \( \langle H_l^2 \rangle \) is proportional to \( l^2 \). By similar calculation, one can easily show that
\[
\langle H_l^n \rangle \sim l^n. \quad (68)
\]

All of the \( G^{(ba)} \) matrices are known for the XY chain. For example, take a periodic system and then consider the thermodynamic limit then we have
\[
G^{(ba)}_{nm} = \frac{1}{2\pi} \int_0^{2\pi} dx e^{i(n-m)x} \frac{\cos[x] + ia \sin[x] - h}{\sqrt{\cos[x] - h}^2 + a^2 \sin^2[x]}, \quad (69)
\]
At the critical Ising point \( (a = h = 1) \), the above equation takes the following simple form
\[
G^{(ba)}_{nm} = -\frac{1}{\pi} \frac{1}{n - m + 1/2}. \quad (70)
\]
Then, one can easily study the \( M(\tilde{\lambda}) \) for different values of \( a \) and \( h \) using the equation (25).

B. Conformal field theory expectations

At the critical points, it is expected that the system can be described by conformal field theory. In particular, on the XY critical line the system can be described by Ising field theory with central charge \( c = 1 \) and on the XX critical line with the central charge \( c = 1 \). Since at the critical points we have
\[
H^\text{CFT}_l = \int_0^l dx T_{00}, \quad (71)
\]
where \( T \) is the energy-momentum tensor with the scaling dimension 2. Notice that since in CFT the one point function of the energy-momentum tensor is by definition zero, we need to work with \( \tilde{E} \) defined in equation (62) rather than \( E \). Then, by simple dimensional analysis we expect
\[
\tilde{E}_n = \langle (H_l^\text{CFT})^n \rangle \sim c_0 + \frac{c_{-n}}{l^0}, \quad (72)
\]
where \( c_0 \) is a \( n \) dependent constant which is also dependent on the cut-off as \( \frac{1}{l} \) and \( c_{-n} \) is another constant. The above equation is expected to be valid at the critical points and we will show its validity for the critical Ising point through numerical calculation and analytically for the XX chain.

At non-critical points, it is natural to expect
\[
\tilde{E}_n \sim a_n + b_n e^{-\alpha_n l}, \quad (73)
\]
where \( a_n, b_n \) and \( \alpha_n \) are all cut-off dependent constants. We will show the validity of the above result by numerical calculations in the next subsections.

C. Transverse field Ising chain

In this section, we study the statistics of the subsystem energy in the transverse field Ising chain. In the Figures 2, and 3, \( M(\tilde{\lambda}) \) has shown for different values of \( l \) and the transverse magnetic field \( h \). A few comments are in order: first of all, there is no particular difference in the shape of the \( M(\tilde{\lambda}) \) at and outside of the critical point. Secondly, one can see that for the finite values of \( l \), \( M(\tilde{\lambda}) \) for large positive and negative values of \( \lambda \) diverges exponentially. However, the interesting point is that since the coefficient of the exponential for positive values of \( \lambda \) is very small, one can see the effect of the exponential for just relatively large values of \( \lambda \). It seems that the coefficient of the positive exponential decreases like an exponential with respect to the size of the subsystem \( l \) which makes \( M(\tilde{\lambda}) \approx e^{-\lambda} \) to be a very good approximation for a large interval of \( \lambda \). Of course, for a very large \( l \) which is comparable with the system size, one does not actually expect any fluctuation in the energy of the system which is in its ground state. That is why in the limit of large \( l \), we have exactly
\[
M(\tilde{\lambda}) = e^{-\lambda}. \quad (74)
\]
From figures 3, one can observe that even for relatively small subsystem sizes the graphs can be described perfectly by the above equation. In Figure 4, we checked the validity of the equation (72) for $\tilde{E}_n$ with $n = 2$, and 3 at the critical point of the transverse field Ising chain, i.e. $h = 1$. The depicted numerical results confirm the CFT predictions nicely. Note that since equation (69) is the thermodynamic limit of a periodic system while, the thermodynamic limit of the subsystem is an open system, one naturally does not expect $\tilde{E}_n$ to go to zero for $l \to \infty$. In fact, it approaches a $n$-dependent constant which shows that the omitted boundary point that connects the two extremes of the subsystem plays a finite role. Since we are just interested in the decay with respect to the subsystem size, this constant does not play any important role in our current investigation. Outside of the critical regime, as it is depicted in the Figure 5, the $\tilde{E}_n$ decays exponentially with respect to the size of the subsystem $l$, which is consistent with our prediction in equation (73).

D. **XX spin chain**

The critical XX line is particularly interesting because it follows a simple tight binding form. In the free fermion representation, one can calculate many quantities exactly. For this reason, we make a thorough discussion of the full counting statistics in this case. We give an independent derivation of the generating function in this
case which is in full agreement with the results of the previous sections.

1. The generating function

The Hamiltonian of the $XX$ chain or chain of free fermions is:

$$H = - \sum_i (c_i^\dagger c_{i+1} + c_{i+1}^\dagger c_i - 2 \cos n_c c_i^\dagger c_i),$$  \hfill (75)

where $n_c$ plays the role of filling factor. We would like to calculate:

$$M(\bar{\lambda}) = \langle e^{\bar{\lambda} H_T} \rangle = \text{tr} \rho \ e^{\bar{\lambda} H_T},$$  \hfill (76)

where $H_T = - \sum_{i=1}^{l-1} (c_i^\dagger c_{i+1} + c_{i+1}^\dagger c_i - 2 \cos n_c c_i^\dagger c_i) = \sum_{ij} A_{ij} c_i^\dagger c_j$ with $-A_{ij} = \delta_{i,j+1} + \delta_{i+1,j} - 2 \cos n_c \delta_{i,j}$ and $\rho$ is the reduced density matrix given by Ref. \textsuperscript{40},

$$\rho = \det(1 - C) e^{(\ln F)_{ij} c_i^\dagger c_j},$$  \hfill (77)

where $F = C(1 - C)^{-1}$. For infinite system, $C_{ij} = \langle c_i^\dagger c_j \rangle = \sin \pi(i-j)/\pi(i-j)$ and $C_{ii} = \frac{n}{2}$, where we will mostly work with $n_c = \frac{n}{2}$. Then, we need to calculate the following trace

$$M(\bar{\lambda}) = \det(1 - C) \text{tr} \left[ e^{(\ln F)_{ij} c_i^\dagger c_j e^{\bar{\lambda} A_{ij} c_i^\dagger c_j}} \right].$$  \hfill (78)

The trace can be easily calculated

$$M(\bar{\lambda}) = \det(1 - C) \det[1 + e^{(\ln F) e^{\bar{\lambda} A}}] = \det(1 - C + C e^{\bar{\lambda} A}).$$  \hfill (79)

The above equation is consistent with the equation (25), which provides another check for our main formula.

The above determinant does not have a simple Toeplitz form which makes further analytic calculations nontrivial. However, it can be simplified further by diagonalizing the matrix $A$ as $A = V D V^T$ where:

$$V_{ij} = \sqrt{\frac{2}{l+1}} \sin \frac{\pi ij}{l+1},$$  \hfill (80)

$$-D_k = 2 \cos \left[ \frac{\pi k}{l+1} \right] - 2 \cos n_c,$$  \hfill (81)

$$M(\bar{\lambda}) = \det \tilde{C} \left( \det \tilde{F}^{-1} + 1 + \sum_{i=1}^{l-1} \alpha_i \beta_i \right) = \det(1 - C) + \det C + \det \left[ \sum_{i=1}^{l-1} \alpha_i \beta_i \right],$$  \hfill (87)

where $\alpha_i$ and $\beta_i$ with the sizes $l_i = \binom{l}{i}$ are defined by (85) and (86) with $y_s = e^{-2 \lambda \cos \left[ \frac{\pi s}{l+1} \right]} + 2 \lambda \cos n_c$.

where $s = 1, 2, ..., l$ and $X = \tilde{F}^{-1}$. Note that it is easy to see that the possible energies are either $-2 \cos \left[ \frac{\pi s}{l+1} \right] +$
2 cos $n_c$ or different possible summations among them. Although the above equation is useful for numerical reasons, it can not be easily used to calculate the asymptotic values of the moments. However, the formula is useful if one is interested in calculating the probability of finding the system in a single eigenstate of the Hamiltonian truncated to a subsystem. Here, we report the generating function for half filling for small subsystem sizes:

$$M(\bar{\lambda}) = \left(\frac{1}{2} - \frac{2}{\pi^2}\right) + \left(\frac{1}{4} + \frac{1}{\pi} + \frac{1}{\pi^2}\right) e^{-\bar{\lambda}} + \left(\frac{1}{4} - \frac{1}{\pi} + \frac{1}{\pi^2}\right) e^{+\bar{\lambda}}$$

$$M(\bar{\lambda}) = \left(\frac{1}{2} - \frac{4}{\pi^2}\right) + \left(\frac{1}{4} + \frac{2}{\pi^2} + \frac{\sqrt{2}}{\pi}\right) e^{-\sqrt{2}\bar{\lambda}} + \left(\frac{1}{4} + \frac{2}{\pi^2} - \frac{\sqrt{2}}{\pi}\right) e^{\sqrt{2}\bar{\lambda}}$$

$$M(\bar{\lambda}) = \left(1 + \frac{64}{9\pi^4} - \frac{46}{15\pi^2}\right) + \left(\frac{1}{16} + \frac{16}{9\pi^4} - \frac{8}{9\pi^2} - \frac{5}{9\pi^2} + \frac{1}{6\pi}\right) e^{-\frac{\bar{\lambda}}{2}} + \left(\frac{1}{16} + \frac{16}{9\pi^4} + \frac{8}{9\pi^2} - \frac{5}{9\pi^2} - \frac{1}{6\pi}\right) e^{\frac{\bar{\lambda}}{2}}$$

$$+ \left(\frac{1}{8} + \frac{32}{9\pi^4} - \frac{8}{9\pi^3} - \frac{64}{9\pi^5} + \frac{16}{9\pi^3} + \frac{1}{6\pi} + \frac{4}{3\sqrt{5}\pi}\right) e^{\frac{\sqrt{3}\bar{\lambda}}{2} - \frac{\pi}{\sqrt{30}\pi}}$$

$$+ \left(\frac{1}{8} - \frac{32}{9\pi^4} + \frac{8}{9\pi^3} - \frac{64}{9\pi^5} + \frac{1}{6\pi} - \frac{4}{3\sqrt{5}\pi}\right) e^{\frac{\sqrt{3}\bar{\lambda}}{2} + \frac{\pi}{\sqrt{30}\pi}}$$

The numbers in the exponentials are the possible subsystem energy values and the coefficients of the exponentials are the probability of occurrence of the corresponding subsystem energy. The coefficients are highly non-trivial numbers. In Figure 6, we depicted the distribution of the logarithm of the subsystem energy for the XX chain. As it is clear, although the probability decreases exponentially it is not a smooth function. The exponential decrease of the probability just means that the subsystem with high probability is either in its ground state or in its first few excited states. It is natural to expect that this should be true independent of the considered model.

2. Calculation of moments

In this subsection, we calculate the subsystem energy moments in the XX chain. To do that one can expand the equation (79) directly and derive some results for the moments as we did in the section II. First of all, we write

$$M(\bar{\lambda}) = \det(1 + \sum_{i=1}^{\infty} \frac{\bar{\lambda}^m}{m!} C A^m)$$

and

$$M(\bar{\lambda}) = 1 + E_1 \bar{\lambda} + E_2 \frac{\bar{\lambda}^2}{2!} + ...$$

An expansion of the above formula with respect to $\bar{\lambda}$ can be derived as follows: $M(\bar{\lambda}) = 1 + E_1 \bar{\lambda} + E_2 \frac{\bar{\lambda}^2}{2!} + ...$
where we have $E_k = \langle H_k \rangle$. First, we define

$$a_0 = 1$$
$$a_k = CA^k \quad k = 1, 2, 3, \ldots$$

Then, using the equations of section II, the first few $E_k$'s can be written as:

$$E_1 = g_1,$$
$$E_2 = g_2 + g_1^2,$$
$$E_3 = g_3 + 3g_1g_2 + g_1^3,$$
$$E_4 = g_4 + 4g_1g_3 + 3g_2^2 + 6g_1^2g_2 + g_1^4,$$
$$E_5 = g_5 + 5g_1g_4 + 10g_2g_3 + 10g_1^2g_3 + 15g_1g_2^2 + 10g_1^3g_2 + g_1^5,$$

where

$$g_1 = \text{tr}a_1,$$
$$g_2 = \text{tr}[a_2 - a_1^2],$$
$$g_3 = \text{tr}[a_3 - 3a_1a_2 + 2a_1^3],$$
$$g_4 = \text{tr}[a_4 - 4a_1a_3 - 3a_2^2 + 12a_1^2a_2 - 6a_1^4],$$
$$g_5 = \text{tr}[a_5 - 5a_1a_4 - 10a_2a_3 + 30a_1a_2^2 + 20a_1^2a_3 - 60a_1^3a_2 + 24a_1^5].$$

The formula for generic $E_k$ can be written with respect to complete exponential Bell's polynomials as:

$$E_k = B_k(g_1, g_2, \ldots, g_k).$$

As before, one can write similar equations for $\tilde{E}_k$ as follows:

$$\tilde{E}_1 = 0,$$
$$\tilde{E}_2 = \tilde{g}_2,$$
$$\tilde{E}_3 = \tilde{g}_3,$$
$$\tilde{E}_4 = \tilde{g}_4 + 3\tilde{g}_2^2,$$
$$\tilde{E}_5 = \tilde{g}_5 + 10\tilde{g}_2\tilde{g}_3.$$

With the following generalization

$$\tilde{E}_m + \sum_{j=2}^{m} (-1)^{j-1}(j-1)!B_{m-j}(\tilde{E}_1, \tilde{E}_2, \ldots, \tilde{E}_{j-1}) = \tilde{g}_m,$$

$$m > 1.$$ (98)

Consequently, the above formulas show that the quantities $\tilde{g}_i$ with $i > 2$ represent the fluctuations of the sub-system energy.

3. Integral representation of the moments

In this subsection, we introduce an integral representation for the moments that we have calculated in the previous subsection. The new representation helps to find asymptotic value of $E_2$. Note that similar calculations can be done for generic XY chain, see Appendix C. However, for generic case the integral representation is more complicated. A simple calculation shows that

$$E_1 = \frac{2}{\pi}(l - 1).$$ (99)

To calculate $E_2$, we need to evaluate $g_2$. However, to calculate $g_2$ we need $tr\tilde{a}_2$ and $tr\tilde{a}_2^2$. The former one has the following simple form:

$$tr\tilde{a}_2 = l - 1.$$ (100)

The later one has more complicated form:

$$tr\tilde{a}_2^2 = 2 \sum_{j,k=1}^{l-1} C_{j,k+1}C_{k,j+1} + 2 \sum_{j=2}^{l-1} \sum_{k=1}^{l-1} C_{j,k+1}C_{k,j-1}.$$ (101)

where the correlation matrix $C$ can be written as

$$C_{j,k} = \int_{-\pi}^{\pi} \frac{dq}{2\pi} n(q)e^{iq(j-k)},$$ (102)

where $n(q) = 1$ for $q < n_c$ and it is zero otherwise. Here, we focus on the half filling $n_c = \frac{\pi}{2}$. Using the above equation and then performing change of variables, finally, one can write the equation (101) in the following form:

$$tr\tilde{a}_2^2 = \frac{2}{\pi^2} \int_{0}^{\pi} dQ \frac{(\pi - Q + \sin Q)(\sin^2 \frac{1}{2}(l - 1)Q)}{1 - \cos Q}.$$ (103)

Now, we have an integral formula for $g_2$ as follows:

$$g_2 = l - 1 - \frac{l - 1}{\pi^2} \int_{0}^{\pi} dQ (\pi - Q + \sin Q)F_{l-1}[Q],$$ (104)

where $F_n[x] = \frac{1}{n} \sin^2 (\frac{\pi}{n}x)$ is the Fejér kernel. The Fejér kernel has the following representation as a sum:

$$F_n[x] = \sum_{k=-\infty}^{n} (1 - \frac{|k|}{n})e^{ikx}. $$ (105)

Putting the above equation in the formula (104) and doing the integral and then performing the sum, we get

$$g_2 = \frac{1}{2\pi^2} \left( \frac{2(-1)^l}{l - 1} + 4(l - 1)\psi^{(1)}(l - 1) - (-1)^l(l - 1)\psi^{(1)}(l - 1) \left( \frac{l}{2} + 1 \right) \right),$$ (106)
where $\psi^{(1)}$ is the polygamma function. After expansion around large $l$, we have
\[
\lim_{l \to \infty} g_2 = \frac{2}{\pi^2} + \frac{1}{\pi^2} \left( \frac{1}{3} + (-1)^i \right) \frac{1}{l^2} + \ldots \quad \text{(107)}
\]

The above result is consistent with what we expect from conformal field theory calculations. Similar but much more complicated calculation can be also carried out for $g_3$, see Appendix D.

**IV. CONCLUSIONS**

In this paper, we calculated the full counting statistics of the subsystem energy for an arbitrary free fermion system in its ground state. The provided formula can be also used as the full counting statistics of an arbitrary quadratic observable. We have calculated exact formulas for the moments as explicit functions of the Hamiltonian parameters. We have also provided a formula for the moments but the decay is exponential at non-critical points. In the case of the XX chain, we provided more refined formulas and showed the validity of the CFT by exact calculations.

There are many directions that one can expand the current work. For example, It will be very interesting to study the FCS of the energy after quantum quench in one dimension. For some related discussions in this direction see Refs.\textsuperscript{31,46,47}. It is also important to study the distribution of the subsystem energy in more direct way (rather than calculating the generating functions) with numerical and analytical techniques in quantum spin chains and also bosonic systems. This kind of calculations can also be useful in the field theory context when one is interested in the localization of the energy in a domain.

**Acknowledgment:** We are indebted to Pasquale Calabrese for numerous discussions and suggestions since the beginning of the project. We also thank Maurizio Fogotti for many inspiring discussions. The work of KN is supported by National Science Foundation under grant number PHY-1314295. The work of MAR is supported in part by CNPq. MAR thanks ICTP for hospitality during a period which part of this work was completed.

**Appendix A: First few terms of T matrix expansion**

To have a better understanding of the equation 33, we show how to obtain the first few terms.

\[
T = \mathbb{1} + \lambda_D \left( \begin{array}{cc} A & B \\ -B & -A \end{array} \right) + \frac{\lambda_D^2}{2!} \left( \begin{array}{ccc} A^2 - B^2 & AB - BA \\ AB - BA & A^2 - B^2 \end{array} \right) + \frac{\lambda_D^3}{3!} \left( \begin{array}{cccc} A^3 - 2B^2A - AB^2 + BAB & A^2B - B^3 - ABA + BA^2 \\ -A^2B + B^3 + ABA - BA^2 & -A^3 + 2B^2A + AB^2 - BAB \end{array} \right) + \ldots \quad \text{(A1)}
\]

The first two terms are easy to study. From equation (A1) for $n = 0$ we get $\mathbb{1}$. Then for $n = 1$ we have two possible terms: $A^1B^0$ and $A^0B^1$ as the sum over the power of matrices should satisfy $\sum_{n_i} = n$. Moreover, the strings always start with $A$ and ends with $B$. Since $n = 1$ is odd, we have to choose $\tau_1^{(1)}$ with the condition $\sum_{j=1}^k n_{2j-1} = \text{even}$, which implies $n_1$ (the power of matrix $A$) to be equal to 0 and consequently, for $\tau_1^{(2)}$ we have to choose $n_1 = 1$. So we have $\tau_1^{(1)} = A^0B^1$ and $\tau_1^{(2)} = -A^1B^0$. Now, we have to determine the sign which is trivial for this case. For $\tau_1^{(1)}$ we have $n_1 = 0$ and $n_2 = 1$ with $k = 2$(total length of AB string) and $s = 0$ so that $\text{sgn} = (-1)^{[0+0]} = 1$ and for $\tau_1^{(2)}$, we have $n_1 = 1$ and $n_2 = 0$ which implies $\text{sgn} = 1$ which is what we expect. Then the other two elements can be obtained from equations (29) and (30). Similarly, we explain how to extract the terms in element of $T_{12}$ for $n = 3$. For this case, we have $2n-1 = 4$ different terms. Considering all
different \( n_k \) in equation 31, we get,

\[
\begin{align*}
\mathbf{A}^2 \mathbf{B}^1 : & \quad [k = 2, n_1 = 2, n_2 = 1, s = 0], \\
& \quad \operatorname{sgn}(k) = (-1)^{\left\lfloor \frac{n_k}{2} \right\rfloor + \frac{1 - (-1)^n}{4}} = 1 \\
\mathbf{A}^0 \mathbf{B}^3 : & \quad [k = 2, n_1 = 0, n_2 = 3, s = 0], \\
& \quad \operatorname{sgn}(k) = (-1)^{\left\lfloor \frac{n_k}{2} \right\rfloor + \frac{1 - (-1)^n}{4}} = -1 \\
\mathbf{A}^1 \mathbf{B}^1 \mathbf{A}^0 : & \quad [k = 4, n_1 = 1, n_2 = 1, n_3 = 1, \\
& \quad n_4 = 0, s = (0, 2)], \\
& \quad \operatorname{sgn}(k) = (-1)^{\left\lfloor \frac{n_k}{2} \right\rfloor + \frac{1 - (-1)^n}{4}} \\
& \quad \times \left(1 - (-1)^{n_k + n_4} \right) = 1
\end{align*}
\]

where the sum is over all non-negative \( j_1, j_2, \ldots, j_{n-k+1} \) in a way that we have \( j_1 + j_2 + \ldots + j_n = k \) and \( j_1 + j_2 + \ldots + (n - k + 1) j_{n-k+1} = n \). Then the complete exponential Bell polynomial can be defined as:

\[
B_n(x_1, x_2, \ldots, x_{n-k+1}) = \sum_{k=1}^{n} B_{n,k}(x_1, x_2, \ldots, x_{n-k+1}).
\]

Now if we define

\[
y_n = \sum_{k=1}^{n} B_{n,k}(x_1, x_2, \ldots, x_{n-k+1}),
\]

then we have

\[
x_n = \sum_{k=1}^{n} (-1)^{k-1} (k-1)! B_{n,k}(y_1, y_2, \ldots, y_{n-k+1}).
\]

The above formula can be considered inverse relation for Bell polynomials.

### Appendix C: Integral representation of \( \tilde{E}_2 \) for XY chain

Here, we write an explicit integral formula of \( \tilde{E}_2 \) for XY chain. Based on the equation (67) we have

\[
\tilde{E}_2 = \frac{1}{16} (\operatorname{tr} [\mathbf{DD}^T] - \operatorname{tr} [\mathbf{D}^T \mathbf{G}^{(ba)} \mathbf{D}^T \mathbf{G}^{(ba)}]).
\]

To write the integral representation we need to manipulate the second term as follows:

\[
\begin{align*}
\operatorname{tr} [\mathbf{D}^T \mathbf{G}^{(ba)} \mathbf{D}^T \mathbf{G}^{(ba)}] &= 4h^2 \sum_{i,k=1}^{l} G^{(ba)}_{i,k} G^{(ba)}_{k,i} - 4h(1 + a)J \sum_{i=1}^{l-1} \sum_{k=1}^{l} G^{(ba)}_{i,k} G^{(ba)}_{k,i+1} \\
& \quad - 4h(1 - a)J \sum_{i=1}^{l-1} \sum_{k=1}^{l} G^{(ba)}_{i,k} G^{(ba)}_{k+1,i} + (1 + a)^2 J^2 \sum_{i=1}^{l-1} \sum_{k=1}^{l} G^{(ba)}_{i,k+1} G^{(ba)}_{k,i+1} \\
& \quad + (1 - a)^2 J^2 \sum_{i=1}^{l-1} \sum_{k=1}^{l} G^{(ba)}_{i+1,k} G^{(ba)}_{k+1,i} + 2(1 - a^2) J^2 \sum_{i=2}^{l-1} \sum_{k=1}^{l} G^{(ba)}_{i-1,k} G^{(ba)}_{k+1,i}.
\end{align*}
\]

Note that since the \( \mathbf{G}^{(ba)} \) matrix is a Toeplitz matrix, one can understand the above formula as the sum of the elements of the pentadiagonal sub-matrix of the matrix \( (\mathbf{G}^{(ba)})^2 \). Using the above equation and after putting \( J = 1 \), one can simply write the following formula
\[ \text{tr}[D^T G^{(ba)} D^T G^{(ba)}] = \int_0^{2\pi} dx \int_0^{2\pi} dy f(x,y) \frac{(ia \sin(x) - h + \cos(x))(ia \sin(y) - h + \cos(y))}{\sqrt{a^2 \sin^2(x) + (\cos(x) - h)^2} \sqrt{a^2 \sin^2(y) + (\cos(y) - h)^2}} \]  

(C3)

where

\[ f(x,y) = 2 \frac{(1 - a^2) e^{-i(x+y)} (e^{i(lx+y)} - e^{i(lx+y)})^2}{(e^{ix} - e^{iy})^2} - 4(1 - a)he^{-i(l-1)(x+y)} \left( e^{ilx} - e^{ily} \right) \left( e^{i(lx+y)} - e^{i(lx+y)} \right) \]

\[ - 4(a + 1)he^{-i(l)(x+y)} \left( e^{ilx} - e^{ily} \right) \left( e^{i(lx+y)} - e^{i(lx+y)} \right) + (1 - a)^2 e^{-i(l-1)(x+y)} \left( e^{ilx} - e^{ily} \right)^2 \]

\[ + \frac{(a + 1)^2 e^{-i(l+1)(x+y)} \left( e^{ilx} - e^{ily} \right)^2}{(e^{ix} - e^{iy})^2} + \frac{4h^2 e^{-i(l-1)(x+y)} \left( e^{ilx} - e^{ily} \right)^2}{(e^{ix} - e^{iy})^2}. \]  

(C4)

The above equations for \( a = h = 0 \) produces the formulas for the XX chain introduced in the paper. Although for generic XY point the above equation is very complicated, there might be special points that can handle the above equations analytically. For example, for large \( l \), the above equation seems to simplifies further but we were not able to find the exact expansion for large \( l \) in different regimes.

**Appendix D: Integral representation of \( \hat{E}_3 \) for XX chain**

In this Appendix, we provide an integral representation for \( g_3 \) defined as:

\[ g_3 = \text{tr}[a_3 - 3a_1a_2 + 2a_1^3]. \]  

(D1)

The first two terms are easy to calculate. After simple algebra we have

\[ \text{tr} a_3 = \frac{2}{3\pi} l(l + 1), \]  

\[ \text{tr}[a_1a_2] = \frac{8}{3\pi} (2l - 3). \]  

(D2)

(D3)

The last term \( 2\text{tr}[a_3^3] \) is more complicated. It can be written as

\[ \text{tr} a_3^3 = 2 \sum_{m,n,k=1}^{l-1} C_{n,k+1} C_{k,m+1} C_{m,n+1} \]

\[ + 3 \sum_{n=2}^{l-1} \sum_{m=1}^{l-1} C_{n,k+1} C_{k,m+1} C_{m,n-1} \]

\[ + 3 \sum_{k=1}^{l-1} \sum_{n,m=2}^{l-1} C_{n,k+1} C_{k,m-1} C_{m,n-1}. \]  

(D4)

Using the same trick as before and after doing some simplifications, we get

\[ \text{tr} a_3^3 = \frac{1}{(2\pi)^3} \int_0^{2\pi} \frac{d\theta_1}{2} \int_0^{2\pi} \frac{d\theta_2}{2} \int_0^{2\pi} \frac{d\theta_3}{2} \]

\[ \left( e^{i(\theta_2-\theta_1)} - e^{i(\theta_2-\theta_1)} \right) \left( e^{i(\theta_1-\theta_3)} - e^{i(\theta_1-\theta_3)} \right) \left( e^{i(\theta_2-\theta_3)} - e^{i(\theta_2-\theta_3)} \right) \times \]

\[ e^{-i\theta_1} + e^{-i\theta_2} + e^{-i\theta_3} + e^{i\theta_1} + e^{i\theta_2} + e^{i\theta_3} + e^{i(\theta_1+\theta_2+\theta_3)} + e^{-i(\theta_1+\theta_2+\theta_3)} \]

\[ \left( 1 - e^{i(\theta_2-\theta_1)} \right) \left( e^{i(\theta_1-\theta_3)} - 1 \right) \left( 1 - e^{i(\theta_2-\theta_3)} \right). \]  

(D5)

Then, we can substitute all of them in equation D1 to get

\[ g_3 = \frac{2}{3\pi} l(l + 1) - \frac{8}{3\pi} (2l - 3) + \frac{4}{(2\pi)^3} \int_0^{2\pi} \int_0^{2\pi} \int_0^{2\pi} \frac{d\theta_1 d\theta_2 d\theta_3}{2} \]

\[ \left( e^{i(\theta_2-\theta_1)} - e^{i(\theta_2-\theta_1)} \right) \left( e^{i(\theta_1-\theta_3)} - e^{i(\theta_1-\theta_3)} \right) \left( e^{i(\theta_2-\theta_3)} - e^{i(\theta_2-\theta_3)} \right) \times \]

\[ \sin \theta_1 + \sin \theta_2 + \sin \theta_3 - \sin(\theta_1 + \theta_2 + \theta_3) \]

\[ \left( 1 - e^{i(\theta_2-\theta_1)} \right) \left( e^{i(\theta_1-\theta_3)} - 1 \right) \left( 1 - e^{i(\theta_2-\theta_3)} \right). \]  

(D6)
which can be written as
\[
g_3 = \frac{2}{3\pi}(l + 1) - \frac{8}{\pi}(2l - 3) + \frac{4}{(2\pi)^3} \int_0^\pi \int_0^\pi \int_0^\pi dq_1 dq_2 dq_3 \left[ \sin[(l - 1)(q_1 - q_2)] + \sin[(l - 1)(q_3 - q_1)] + \sin[(l - 1)(q_2 - q_3)] \right] \times \\
\sin[q_1 - q_2] + \sin[q_3 - q_1] + \sin[q_2 - q_3] \\
\sin[q_1 + q_2 + q_3 - \sin[q_1 + q_2 + q_3]].
\]  \tag{D7}

After some manipulations, one can also get the following form:
\[
g_3 = \frac{2}{3\pi}(l + 1) - \frac{8}{\pi}(2l - 3) + \frac{16}{(2\pi)^3} \int_0^\pi \int_0^\pi \int_0^\pi dq_1 dq_2 dq_3 \left[ \sin\left[\frac{(l-1)}{2}(q_1 - q_2)\right] \sin\left[\frac{(l-1)}{2}(q_3 - q_1)\right] \sin\left[\frac{(l-1)}{2}(q_2 - q_3)\right] \right] \times \\
\sin\left[\frac{q_1 + q_2}{2}\right] \sin\left[\frac{q_3 + q_1}{2}\right] \sin\left[\frac{q_2 + q_3}{2}\right].
\]  \tag{D8}

Although the above equations have nice symmetric forms, it is not easy to find its asymptotic value. Based on numerical calculations it is easy to see that
\[
g_3 = c_0 + \frac{c_{-3}}{l^3} + \ldots, \quad \text{with } c_{-3} \text{ a bounded but oscillating function.}
\]  \tag{D9}

---

1. L. Levitov and G. Lesovik, JETP LETTERS \textbf{58}, 230 (1993).
2. Y. M. Blanter and M. Büttiker, Physics Reports \textbf{336}, 1 (2000).
3. L. S. Levitov and M. Reznikov, Phys. Rev. B \textbf{70}, 115305 (2004).
4. Sukhorukov, E. V. et al. Nature Phys. \textbf{3}, 243-247 (2007).
5. W. Belzig, Phys. Rev. B \textbf{71}, 161301(R) (2005).
6. M. Vanević, Yu. V. Nazarov, and W. Belzig, Phys. Rev. Lett. \textbf{99}, 076601 (2007).
7. K. Kaasbjerg and W. Belzig, Phys. Rev. B \textbf{91}, 235413 (2015).
8. D. Dassenbrook, C. Flindt, Phys. Rev. Lett. \textbf{117}, 146801 (2016).
9. T. Szczepanski, V. K. Dugaev, J. Barnas, I. Martinez, J. P. Cascales, J. Y. Hong, M. T. Lin, F. G. Aliev, Phys. Rev. B \textbf{94}, 235429 (2016).
10. D. A. Abanin and L. S. Levitov, Phys. Rev. Lett. \textbf{94}, 186803 (2005).
11. V. Gritsev, E. Altman, E. Demler, and A. Polkovnikov, Nature Phys. \textbf{2}, 705 (2006).
12. A. Polkovnikov, E. Altman, and E. Demler, \textit{Proceedings of the National Academy of Sciences} \textbf{103}, 6125 (2006).
13. S. Hofferberth, I. Lesanovsky, T. Schumm, A. Imambekov, V. Gritsev, E. Demler and J. Schmiedmayer, Nature Phys. \textbf{4}, 489 (2008).
14. I. Klich and L. Levitov, Phys. Rev. Lett. \textbf{102}, 100502 (2009).
15. H. F. Song, S. Rachel, C. Flindt, I. Klich, N. Laforencie, and K. Le Hur, Phys. Rev. B \textbf{85}, 035409 (2012).
16. S. Rachel, N. Laforencie, H. F. Song, and K. Le Hur, Phys. Rev. Lett. \textbf{108}, 116401 (2012).
17. D. A. Ivanov and A. G. Abanov, Phys. Rev. E \textbf{87}, 022114 (2013).
18. M. M. Cardoner, J. F. Sherson and G. De Chiara, New J. Phys. \textbf{18} 103015 (2016).
19. J. M. Stéphan, F. Pollmann, Phys. Rev. B \textbf{95}, 035119 (2017).
20. M. Collura, Fabian H.L. Essler, Stefan Groha [arXiv:1706.07939].
21. D. J. Luitz, N. Laforencie, and F. Alet, Phys. Rev. B \textbf{91}, 081103 (2015).
22. R. Singh, H. H. Bardarson, and F. Pollmann, New J. Phys. \textbf{18}, 023046 (2016).
23. I. Klich, J. Stat. Mech \textbf{P11006} (2014).
24. S. Humeniuk, H. P. Bchler, [arXiv:1706.08951].
25. V. Eisler, Z. Rác, and F. van Wijland, Phys. Rev. E \textbf{67}, 056129 (2003).
26. R. W. Cherng and E. Demler, New J. Phys. \textbf{9}, 7 (2007).
27. A. Lamacraft, P. Fendley, Phys. Rev. Lett. \textbf{100}, 165706 (2008).
28. V. Aji and N. Goldenfeld, Phys. Rev. Lett. \textbf{86}, 1007 (2001).
29. A. Bruce, J. Phys. C \textbf{14}, 3667 (1981).
30. K. Binder, Z. Phys. B \textbf{43}, 119 (1981).
31. I. Arad, T. Kuwahara, Z. Landau, J. Stat. Mech.(2016) \textbf{033301}.
32. J. P. Pekola, P. Muratore-Ginanneschi, A. Kupiainen, Y. M. Galperin, Phys. Rev. E \textbf{94}, 022123 (2016).
33. Y. K. Saito and A. Dhar, Phys. Rev. Lett. 99, 180601 (2007).
34. Y K. Saito and A. Dhar, Phys. Rev. E 83, 041121 (2011).
35. A. A. Clerk, F. Marquardt, and J. G. E Harris, Phys. Rev. Lett. 104, 213603 (2010).
36. Ya. A. Clerk, Phys. Rev. A 84, 043824 (2011).
37. Z. Yu, G. Tang, and J. Wang Phys. Rev. B \textbf{93}, 195419 (2016).
38 R. Seoane Souto, R. Avriller, R. C. Monreal, A. Martín-Rodero, and A. L. Yeyati, Phys. Rev. B 92, 125435 (2015).
39 B. K. Agarwalla, J.-H. Jiang, and D. Segal, Phys. Rev. B 92, 245418 (2015).
40 M. C. Chung and I. Peschel, Phys. Rev. B, 64, 064412(2001) and I. Peschel, J. Phys. A: Math. Gen. 36, L205(2003).
41 T. Barthel, and U. Schollwock, Phys. Rev. Lett. 100, 100601(2008).
42 R. Balian and E. Brézin, Nuov. Cim. 64 B, 37 (1969).
43 C. S Withers and S Nadarajeh, Canadian Applied Mathematics Quarterly, 18, 107 (2010).
44 U. Prells, M.I. Friswell and S.D. Garvey, Proc. Roy. Soc. London A 459 273 (2003).
45 K. Najafi, M. A. Rajabpour, Phys. Rev. B 93, 125139 (2016).
46 A. Silva, Phys. Rev. Lett. 101, 120603 (2008)
47 Markus P. Müller, E. Adlam, L. Masanes and N. Wiebe, Communications in Mathematical Physics, 340, (2015)499561.