Measurement of proton-distribution radii of neutron-rich nitrogen isotopes

Soumya Bagchi1,2,3,∗, Rituparna Kanungo1,4, Wataru Horiuchi5, Gaute Hagen6,7, Titus D. Morris6,7, Steven Ragnar Stroberg8, Toshiba Suzuki8, Frederic Ameil1, Joel Atkinson1, Yassid Alyaad9, Dolores Cortina-Gil10, Iris Dillmann1, Alfredo Estrade1,2, Alexey Evdokimov2, Fabio Farinon2, Hans Geissel2,3, Giulia Guastalla2, Rudolf Janik10, Sabir Kaur1,11, Ronja Knobel2, Jan Kurcewicz2, Yuriy Litvinov2, Michele Marta2, Magdalena Mostazo2, Ivan Mukha2, Chiara Nociiforo2, Hooi Jin Ong1,2, Stephane Pietri3, Andrej Prochazka2, Christoph Scheidenberger2,3, Branislav Sitar10, Peter Strmen10, Maya Takechi2, Junki Tanaka12, Yoshiki Tanaka1,2,4, Isao Tanihata12,13, Satoru Terashima12, Jossit Vargas3, Helmut Weick2, and John Stuart Winfield2

1Astronomy and Physics Department, Saint Mary’s University, Halifax, NS B3H 3C3, Canada
2GSI Helmholtzzentrum für Schwerionenforschung GmbH, D-64291 Darmstadt, Germany
3Justus-Liebig University, 35392 Giessen, Germany
4TRIUMF, Vancouver, BC V6T 4A3, Canada
5Department of Physics, Hokkaido University, Sapporo 060-0810, Japan
6Physics Division, Oak Ridge National Laboratory, Oak Ridge, TN 37831, USA
7Department of Physics and Astronomy, University of Tennessee, Knoxville, TN 37996, USA
8Department of Physics, Nihon University, Setagaya-ku, Tokyo 156-8550, Japan
9Universidad de Santiago de Compostela, E-15706 Santiago de Compostela, Spain
10Faculty of Mathematics and Physics, Comenius University, 84215 Bratislava, Slovakia
11Department of Physics and Atmospheric Science, Dalhousie University, Halifax, NS B3H 4R2, Canada
12RCNP, Osaka University, Mihogaoka, Ibaraki, Osaka 567 0047, Japan
13School of Physics and Nuclear Energy Engineering and IRCNPC, Beihang University, Beijing 100191, China

Abstract. Measurement of root-mean-square radii of proton distributions of 17−22N from charge-changing cross section shows the emergence of thick neutron skin towards the neutron-drip line. Signature of N = 14 shell gap has been found in nitrogen isotopes along with the emergence of neutron halo in 22N. The measured radii are in good agreement with the shell model calculations.

1 Introduction

With large neutron-to-proton ratios far from the valley of stability, nuclei develop exotic structures such as neutron skin, nuclear halo, modifications of conventional shells and appearance of new shell gaps. Evidence for breaking down of N = 8 shell gap has been found in 11Li and 11Be [1–3] while a new shell gap at the neutron number N = 16 has been found in carbon, nitrogen and oxygen isotopes at the drip line [4]. Another sub-shell gap at N = 14 has been found in carbon, nitrogen and oxygen isotopes along with the emergence of neutron halo in 22N. The measured radii are in good agreement with the shell model calculations.}

2 Experimental setup

In this work, the determination of root-mean-square radii of proton-distribution (protons treated as point-like particles and therefore referred further as point-proton radii) from the measured charge-changing cross sections (σCC) of neutron-rich isotopes 17–22N and stable isotopes 14,15N is presented. The experiment was performed at the fragment separator FRS at GSI. Beams of 14,15N and 17–22N were produced by fragmentation of 22Ne and 40Ar, respectively, at 1 A GeV, impinging on a 6.3 g/cm2 thick Be target. The isotopes of interest were separated in flight and identified using their magnetic rigidity (Bρ), time-of-flight (ToF) between the focal planes (F2-F4) of the FRS, and the energy-loss (∆E) measured in a multi-sampling ionization chamber (MUSIC) placed at the final focal plane F4. The σCC was measured with a 4.010 g/cm2 thick carbon target placed at F4. The experimental setup is discussed in more details in Ref. [7].

To obtain σCC, the number (Nin) of incident nuclei of interest A Z is determined from event-by-event counting. After the reaction target, the number of nuclei with the same charge Z as the incident nuclei (N~amez) are identified using the MUSIC detector. The σCC is then obtained from the relation σCC = r−1ln(RP~amez/RT~amez). Here RP~amez and RT~amez are the ratios of N~amez / Nin with and without the reaction target, respectively, and r is the target thickness. The term RT~amez attributes for losses due to interactions with non-target materials. Phase-space restriction of the inci-
dent beam at the target is required to eliminate events having large incident angles and positions far from the center. Details of the measurement procedure can be found in Ref. [7].

3 Results and discussion

To obtain the proton-distribution radii from the measured cross section $\sigma_{CC}$, finite-range Glauber model framework is used [8]. For stable isotopes (e.g. $^{10}$B, $^{12-14}$C, $^{14}$N), the obtained proton-distribution radii are in good agreement with the electron-scattering data [7, 9, 10]. The extracted point-proton-distribution radii ($R_p$) from $\sigma_{CC}$ are shown in Fig. 1.

![Figure 1](image.png)

Figure 1. The experimental point-proton radii are shown with red filled circles. Shell model predictions using the YSOX interaction and WS (HO) potentials are shown with black filled squares (black open triangles). RMF theory predictions are given with the black dashed curve from Ref. [11].

The decrease of $R_p$ from $^{17}$N to $^{21}$N within the uncertainties suggests the reduction in deformation and the transition towards sphericity at the $N = 14$ shell closure. An increase of $R_p$ has been observed within uncertainties from $^{21}$N to $^{22}$N. In $^{22}$N, the valence neutron is in the $2s_1/2$ orbital with a closed-shell core of $^{21}$N. As a consequence, the center-of-mass (c.m.) of $^{22}$N is different from that of the $^{21}$N core which leads to the c.m. motion and hence it smears the core density, resulting in a larger core size [7]. Similar phenomenon has been observed in the two neutron halo $^{11}$Li [12] and $^{6}$He [13] isotopes.

Shell model calculations are compared with the experimental data and are shown in Fig. 1. Both harmonic oscillator (HO) and Woods-Saxon (WS) wave functions are used in the YSOX hamiltonian. The combination of both HO and WS wavefunctions depicts that the proton radii of the nitrogen isotopes increase from $^{13}$N to $^{17}$N and then a decrease is seen with a shallow minimum at the neutron number $N = 14$. In Fig. 1, the relativistic mean field (RMF) predictions [11] are shown which are consistent with the experimental data. Ab initio, in-medium similarity re-normalization group, calculations with a state-of-the-art chiral nucleon-nucleon and three-nucleon interaction reproduce well the data towards the neutron drip-line isotopes rather than the whole isotopic trend of the radii [7].

The matter-distribution radii can be obtained from the measured interaction cross sections [14] using the finite-range Glauber model framework. Together with the information on the matter-distribution and proton-distribution radii, thick neutron skins for $^{19-21}$N, have been found, while for $^{22}$N a neutron halo-like structure develops [7].

For nitrogen isotopes, the proton–neutron tensor interaction is attractive for p($1p_1/2$)–n($1d_5/2$) orbitals, thereby reducing the gap between proton $1p_1/2$ and $1p_3/2$ orbitals when more neutrons are added to the $1d_5/2$ orbital, resulting in a small point proton radius in $^{21}$N. This attractive interaction also lowers the neutron $1d_5/2$ orbital leading to the $N = 14$ shell gap [7] in neutron-rich nitrogen isotopes.

References

[1] A. Navin, et al., Phys. Rev. Lett. 85, 266 (2009).
[2] H. Iwasaki, et al., Phys. Lett. B 491, 8-14 (2000).
[3] H. Simon, et al., Phys. Rev. Lett. 83, 496 (1999).
[4] Z. Dlouhý, J. Mrázek and D. Baiborodin, AIP Conf. Proc. 610, 736 (2002).
[5] M.J. Strongman, et al., Phys. Rev. C 80, 021302(R) (2009).
[6] I. Angeli, K.P. Marinova, At. Data Nucl. Data Tables 99, 69 (2013).
[7] S. Bagchi, et al., Phys. Lett. B 790, 251-256 (2019).
[8] Y. Suzuki, et al., Phys. Rev. C 94, 011602(R) (2016).
[9] A. Estradé, et al., Phys. Rev. Lett. 113, 132501 (2014).
[10] R. Kanungo, et al., Phys. Rev. Lett. 117, 102501 (2016).
[11] J. Meng, et al., Phys. Lett. B 532, 209-214 (2002).
[12] R. Sánchez, et al., Phys. Rev. Lett. 96, 033002 (2006).
[13] P. Mueller, et al., Phys. Rev. Lett. 99, 252501 (2007).
[14] A. Ozawa, et al., Nucl. Phys. A 693, 32-62 (2001).