Azimuthally integrated HBT parameters for charged pions in nucleus-nucleus interactions versus collision energy

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In the paper energy dependence of space-time extent of emission region obtained from Bose–Einstein correlations is studied for charged pions in various ion collisions for all experimentally available energies. There is no dramatic change of HBT parameters with increasing of collision energy per nucleon-nucleon pair, \(\sqrt{s_{NN}}\), in domain of energies \(\sqrt{s_{NN}} \geq 5\) GeV. Energy dependence of estimations for emission duration is almost flat for all energy domain under study within large error bars. Analytic function is suggested for smooth approximation of energy dependence of main HBT parameters. Fit curves demonstrate reasonable agreement with experimental data for most HBT parameters in energy domain \(\sqrt{s_{NN}} \geq 5\) GeV. Estimations of some observables are obtained for energies of the LHC and FCC project.

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I. INTRODUCTION

At present femtoscopic correlations in particular that based on Bose–Einstein correlations are unique experimental method for the determination of sizes and lifetimes of sources in high energy and nuclear physics. The discussion below is focused on specific case of femtoscopy, namely, on correlations in pairs of identical charged pions with small relative momenta – HBT-interferometry – in nucleus-nucleus collisions. Space-time characteristics for emission region of secondary particles created in (heavy) ion collisions are important for study of deconfinement state of strongly interacting matter – strong-coupling quark-gluon plasma (sQGP). Furthermore the study of energy dependence of femtoscopic observables can be useful for understanding in detail the transition from sQGP produced at higher energies to confined hadronic resonance matter created in final state at lower energies. HBT analysis allows studying dynamic features of interaction process at late, i.e. soft, stage of space-time evolution of multiparticle final state. Therefore the study of nucleus-nucleus collisions in wide energy domain by correlation femtoscopy seems important for better understanding both the equation of state (EOS) of strongly interacting matter and general dynamic features of soft processes.

The paper is organized as follows. In Sec. II, definitions of main observables for correlation femtoscopy are briefly described. The Sec. III devotes discussion of experimental energy dependence for space-time extent of source of charged pions and corresponding fits. Also estimations for femtoscopic observables are shown for the LHC and the Future Circular Collider (FCC) project energies. Some final remarks and conclusions are presented in Sec. IV.

II. METHOD AND VARIABLES

In general phenomenological parameterization of correlation function (CF) for two identical particles with 4-momenta \(p_1, p_2\) and with taking into account different forms of corrections on Coulomb final state interaction (FSI) can be written as follows [1]:

\[
C_{2,(m)}^{ph}(q, K) = \epsilon \mathcal{P}^{(m)}(q) [\epsilon^{-1} + K_2^{ph}(A)] , \quad \epsilon = \begin{cases} 
\lambda, & \text{at } m = 1, 2; \\
1, & \text{at } m = 3.
\end{cases}
\] (1)

where \(m = 1\) corresponds to the standard Coulomb correction, \(m = 2\) – the dilution procedure and \(m = 3\) – the Bowler–Sinyukov correction, \(q \equiv (q^0, \vec{q}) = p_1 - p_2\) is the relative 4-momentum, \(K \equiv (K^0, \vec{K}) = (p_1 + p_2)/2\) – the average 4-momentum of particles in pair (pair 4-momentum), for standard simplest (Gaussian) case

\[
K_2^{ph}(A) = \prod_{i,j=1}^{3} K_2^{ph}(A_{ij}) = \exp \left( - \sum_{i,j=1}^{3} q_i^2 R_{ij}^2 q_j \right).
\] (2)

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Here $A \equiv \mathbf{q} \mathbf{R}^2 \mathbf{q}^T$ and $\mathbf{R}^2$ are the matrices $3 \times 3$, $\mathbf{q}^T$ – transposed vector $\mathbf{q}$, $\forall i, j : \mathbf{R}_{ij}^2 = \mathbf{R}_{ji}^2$, $\mathbf{R}_{ii}^2 = \mathbf{R}_{ii}^2$, where $R_i = R_i(K)$ are parameters characterized the linear scales of homogeneity region [2]; the products are taken on space components of vectors, $\lambda(K) = \mathbf{K}_\lambda(0, K), 0 \leq \lambda \leq 1$ is the parameter which characterizes the degree of source chaoticity. Taking into account the hypothesis of cylindrical symmetry of source the volume of homogeneity region is derived as follows [2]

$$V = (2\pi)^{3/2} \prod_{i=1}^{3} R_i. \tag{3}$$

The space component of pair 4-momentum ($\mathbf{K}$) is decomposed on longitudinal $k_{\parallel} = (p_{\parallel,1} + p_{\parallel,2})/2$ and transverse $k_{\perp} = (p_{\perp,1} + p_{\perp,2})/2$ parts of pair momentum. In the paper the decomposition of Pratt–Bertsch [4, 5] is used for $\mathbf{q}$ as well as the longitudinal co-moving system (LCMS) frame. The radius $R_o$ contains additional contribution from the temporal extent of the source. Therefore this parameter is usually excluded from calculation of $V$ and the volume of source can be written as follows

$$V = (2\pi)^{3/2} R_o^2 R_l. \tag{4}$$

As seen $V_{\infty} = R_o/R_s$, where $V_{\infty}$ denotes the source volume calculated from eq. (3) and the volume of emission region defined in accordance with eq. (4) is designated by $V_{\infty}$. But it should be emphasized that in the limit for absolute value of transverse pair momentum vector $k_{\perp} \to 0$, no transverse vector allows to distinguish between out- and side-components [4, 5]. This implies that $\lim_{k_{\perp} \to 0} R_o(K) = \lim_{k_{\perp} \to 0} R_e(K)$. Consequently, it is expected $V_{\infty} \approx V_{\infty}$ for particles with low $k_{\perp}$ and both the relations (3) and (4) for freeze-out volume are valid for such particles.

The one of the important additional observables is the following difference $[6, 7]$,

$$\delta \equiv R_o^2 - R_s^2. \tag{5}$$

If the emission function features no position-momentum correlation, then $\delta$ is finite at non-zero $\mathbf{K}$ only due to explicit $\mathbf{K}$-dependence (resulting from the mass-shell constraint $q^0 = \mathbf{q} \mathbf{K}/K^0$) [6]. In this case

$$\delta \approx \beta_1^2 (\Delta \tau)^2, \tag{6}$$

where $\beta_1 = k_{\perp}/m_{\perp}$ is the transverse velocity of pair of particles with mass $m$, $m_{\perp}^2 = k_{\perp}^2 + m^2$, $\Delta \tau$ – the emission duration for the particle type under discussion. It should be stressed the last relation is valid in some specific cases 1D hydrodynamics while is violated in both the cascade approaches and multidimensional hydrodynamic models. Thus in the framework of some assumptions the $\delta$ gives direct access to the emission duration of the source and allows to partially disentangle the spatial and temporal information contained in radii parameters $R_i$. [6]. The sensitivity to the $\Delta \tau$ is the main advantage of the observable (5).

In the paper the following set of main femtoscopic observables $G_i \equiv \{G_i^1\}_{i=1}^4 = \{\lambda, R_e, R_o, R_l\}$ is under consideration as well as the set of some important additional observables which can be calculated with help of HBT radii $G_2 \equiv \{G_i^2\}_{i=1}^3 = \{R_o/R_s, \Delta \tau, V\}$. The set of parameters $G_1$ characterizes the chaoticity of source and its 4-dimensional geometry at freeze-out stage completely. Scaled femtoscopic parameters $G_1^i, i = 2, 4$, and $G_2$ are calculated as follows [1]:

$$R_i^0 = R_i/R_A, \quad i = s.o.l; \quad \delta^n = \delta/R_A^3 ; \quad V^n = V/V_A. \tag{7}$$

Here $R_A = r_0 A^{1/3}$, $V_A = 4\pi r_0^3/3$ is radius and volume of spherically-symmetric nucleus, $r_0 = (1.25 \pm 0.05)$ fm [8, 9]. The change $R_A \to (R_A) = 0.5(R_A + R_A)$ is made in the relation (7) in the case of non-symmetric nucleus-nucleus collisions [1]. One needs to emphasize the most central collisions are usually used for study the space-time characteristics of final-state matter, in particular, for discussion of global energy dependence of femtoscopic observables (see below Sec. 3). Thus the using of radius of all the nucleus in (7) seems reasonable. In general case the scale factor in (7) for calculation of normalized femtoscopic radii, $\delta$ and volume should takes into account the centrality of nucleus-nucleus collisions. The normalization procedure suggested in [1] allows to consider two data samples, namely, (i) only (quasi)symmetric heavy ion collisions and (ii) all available data for nucleus-nucleus collisions. Ensemble of experimental data reviewed in [1] with replacement of Au+Au points at $\sqrt{s_{NN}} \geq 11.5$ GeV by the recent STAR results of high-statistics analysis for Au+Au collisions at $\sqrt{s_{NN}} = 7.7, 11.5, 62.4$ and 200 GeV [10] is used in the present study. It should be noted that estimation based on [11] is used for $\lambda = \sqrt{s_{NN}} = 2.76$ TeV here as well as in the previous analysis [1]. But recent study at the LHC energy [12] obtains slightly larger value of chaoticity ($\lambda \sim 0.6$) for two-pion correlations in kinematic domain under study. The change of the $\lambda$ value in TeV-region of collision energies can influence on general trend and quantitative results discussed below for this femtoscopic parameter. Therefore this discrepancy should be investigated additionally. Such study is in the progress.
III. ENERGY DEPENDENCE OF SPACE-TIME EXTENT OF PION SOURCE

Dependencies of femtoscopic parameters $G_i^l(\sqrt{s_{NN}})$, $i = 1 - 4$ and $R_o/R_s(\sqrt{s_{NN}})$ are shown in Figs. 1a–d and Fig. 1I respectively. It should be stressed that the STAR results considered here were obtained for fit function (2) with taking into account two additional cross-terms for $R_o$ and $R_s$ as well as with improved Coulomb correction $P_{\text{coul}}^q(q)$. The two sets of STAR results for $\{G_i^l\}_{i=1}^4$ are in a good agreement for previously study (1) and for present analysis for most femtoscopic parameters under consideration in Fig. 1I. The some decreasing is seen for $\lambda$ and transverse radii $R_o$, $R_s$ for Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV with respect to the earlier STAR results from 13. The new results from 10 agree better with both the general trends and the results of other experiments (PHENIX and PHOBOS) at top RHIC energy $\sqrt{s_{NN}} = 200$ GeV. The chaoticity parameter $\lambda$ decreases with increasing $\sqrt{s_{NN}}$ relatively fast at lower (AGS) energies and shows the weak changing at $\sqrt{s_{NN}} > 4$ GeV (Fig. 1I). Femtoscopic radii of source in transverse plane with respect to the beam direction, $R_o$ (Fig. 1I) and $R_s$ (Fig. 1I), show little change over a wide range of energy $5 \lesssim \sqrt{s_{NN}} \lesssim 200$ GeV which corresponds to highest AGS–SPS–RHIC beam collision energies. On the other hand, the value of source size in longitudinal direction, $R_l$ (Fig. 1I), appears to reach a minimum around $\sqrt{s_{NN}} = 5$ GeV, rising in energy domain available at RHIC. As seen there is increasing of HBT radii (Figs. 1I–d) at growth of collision energy from $\sqrt{s_{NN}} \sim 20$ GeV up to maximum available LHC energy $\sqrt{s_{NN}} = 2.76$ TeV. The significant increasing of HBT radii is seen for much broader energy range (on about two order of magnitude $\sqrt{s_{NN}} \sim 0.02 - 3$ TeV) only than it was expected early at the beginning of RHIC operation. Therefore the space-time extent of emission region at freeze-out changes slowly at increasing of collision energy. The transverse radius $R_s$ reflects the spatial extent of particle source, whereas $R_o$ is also affected by dynamics 14, 15 and is believed to be related to the duration of particle emission 16. As indicated, for example, in 10, the ratio $R_o/R_s$ was predicted to increase with beam energy by hydrodynamical calculations and might shows an significant enhancement if the life-time of the collision evolution (and, within these models, the duration of particle emission as a result) was to be extended by entrance into a different phase 10. There is no significant increasing of ratio $R_s/R_o$ in all experimentally available energy domain (Fig. 1I). Recent developments, in particular in viscous hydrodynamics, allow to get reasonable agreement between experimental and model values of $R_o/R_s$ at top RHIC energy and demonstrate that the behavior of experimental dependencies of $R_o/R_s$ on kinematic variables can be explained in particular by realistic EoS with crossover phase transition and sQGP at high temperatures 17–22. Therefore the soft femtoscopic observables confirm the phase transition and creation of deconfinement state of strongly interacting matter in collider experiments.

Taking into account the view of experimental dependencies in Figs. 1I–d the following function is suggested

$$f(\sqrt{s_{NN}}) = a_1 [1 + a_2(\ln \epsilon)^{a_3}]$$

(8)

for smooth approximation of $G_i^l(\sqrt{s_{NN}})$, $i = 1 - 4$, where $\epsilon \equiv s_{NN}/s_0$, $s_0 = 1$ GeV$^2$. Also the specific case of (8) at $a_3 = 1.0$ is under consideration. As seen from Figs. 1I–d there is indication on change of behavior of energy dependence (inflection point) for $\{G_i^l\}_{i=2}^4$ at $\sqrt{s_{NN}} \approx 5$ GeV. This inflection point is seen most clear for $R_l$ (Fig. 1I). Therefore the fit function (8) is used for approximation of the energy dependence of HBT radii in the energy domain $\sqrt{s_{NN}} \geq 5$ GeV only. Experimental energy dependence of $\lambda$ is fitted by general function (8) at all available energies. As seen the point from the WA97 experiment 22 differs significantly from other results at close energies for $\lambda$ (Fig. 1I) and longitudinal radius (Fig. 1I). Thus for these parameters fits are made for data sample (i) with exception of the point from 22. For each main HBT parameters $\{G_i^l\}_{i=1}^4$ fits are made for both the statistical and total errors, where total errors of experimental points include available clear indicated systematic errors added in quadrature to statistical ones. The numerical values of fit parameters are presented in Table I where the second line for each HBT parameter $\{G_i^l\}_{i=1}^4$ corresponds to the approximation by specific case of (8). Fit curves are shown in Fig. 1I by solid lines for (8) and by dashed lines for specific case of fit function at $a_3 = 1.0$ with taking into account the statistical errors. In general fit function described above agrees reasonably with experimental dependence $G_i^l(\sqrt{s_{NN}})$, $i = 1 - 4$ (Fig. 1I–d). But the fit qualities are poor for all the main HBT parameters, especially for $\lambda$, with statistical errors taken into account (Table I). Spread of experimental points leads to the statistically unacceptable values of $\chi^2/\text{n.d.f.}$ In the case of $\lambda$ inclusion of estimation for systematic uncertainty for Pb+Pb collisions at $\sqrt{s_{NN}} = 2.76$ TeV leads to both the dramatic growth of $a_2$ and improvement of the fit quality at transition from statistical errors of experimental points to total errors in data sample (i) (Table I). Inclusion of total errors allows to get a statistically acceptable fit qualities for HBT radii for both the function (8) and its specific case. It seems more complex fit function should be used in order to describe energy dependence of HBT radii at all available collision energies. This study is in the progress. Taking into account the similar behavior of the energy dependence of HBT radii (Figs. 2I–d) and elliptic flow $v_2$ 24 at qualitative level the following functional form can be suggested $g(\sqrt{s_{NN}}) = a_1 + a_2(\sqrt{s} - a_3)^{a_4} + \sum_{i=5,6} a_i e^{a_i i+1} + a_6 (\ln \epsilon)^{a_7}$ as first approach for description of $G_i^l(\sqrt{s_{NN}})$, $i = 1 - 4$ in all experimentally available energy domain. Smooth solid and dashed curves shown in Fig. 1I: are calculated for the ratio $R_o/R_s$ from the fit results for $R_s$ and $R_o$ (Table I). As seen
these curves agree with experimental points reasonably at $\sqrt{s_{NN}} \geq 5$ GeV. In general fits by the function (8) at free $a_3$ and fixed $a_3$ show close behavior for all the main HBT parameters from $G_1$ with some differences at intermediate ($\sqrt{s_{NN}} \lesssim 10$ GeV) and high ($\sqrt{s_{NN}} > 200$ GeV) energies. These differences result in more significant discrepancy between fit curves for $R_o/R_s$ (Fig. 2) and for other parameters from the set $G_2$ (see below).

Table I: Values of fit parameters for approximation of the data sample (i)

| HBT parameter | Fit with statistical errors | Fit with total errors |
|---------------|-----------------------------|-----------------------|
| $a_1$        | $a_2$ | $a_3$ | $\chi^2$/n.d.f. | $a_1$ | $a_2$ | $a_3$ | $\chi^2$/n.d.f. |
| $\lambda$    | 0.36 ± 0.02 | 1.90 ± 0.12 | 0.91 ± 0.14 | 534/23 | 0.008 ± 0.002 | 104 ± 21 | 0.291 ± 0.011 | 373/23 |
| $R_s$        | 4.77 ± 0.02 | 1.3 ± 0.8 | 0.0001 | 580/19 | 0.621 ± 0.005 | 0.032 ± 0.001 | 1.0 (fixed) | 266/19 |
| $R_o$        | 0.59 ± 0.06 | 7.4 ± 1.0 | 0.0097 ± 0.009 | 302/2 | 0.38 ± 0.18 | 0.038 ± 0.008 | 0.03 ± 0.003 | 31.9/22 |
| $R_I$        | 5.38 ± 0.04 | 0.0120 ± 0.0012 | 1.0 (fixed) | 411/23 | 5.3 ± 0.2 | 0.003 ± 0.002 | 1.0 (fixed) | 31.9/23 |

Fig. 2 demonstrates the energy dependence of $\Delta \tau$ for (quasi)symmetric heavy ion collisions. The emission duration in these collisions is calculated based on known HBT-radii (Figs. 11 – d), kinematic regime for pion pairs and on $c_3$. The $\langle k_1 \rangle \approx 0.82$ for pion pairs with $\langle k_1 \rangle \approx 0.2$ GeV/c. Value $\Delta \tau = (0.53 \pm 9.15)$ fm/c at $\sqrt{s_{NN}} = 130$ GeV derived from the PHENIX results at this energy is not shown due to extremely large errors. As seen the emission duration for pions extracted from $\delta$ $c_3$ is about 2 – 4 fm/c for any energies under consideration. The visible energy dependence of emission duration is absent, $\Delta \tau(\sqrt{s_{NN}})$ is close to flat within large error bars. One can see more interesting behavior for this dependence for the STAR high-statistics data only. But additional precise measurements are necessary in order to confirm the change of $\Delta \tau(\sqrt{s_{NN}})$ at $\sqrt{s_{NN}} \sim 10 – 20$ GeV and locate the possible knee in the experimental dependence. Smooth solid and dashed curves shown in Fig. 2 are calculated for $\Delta \tau$ from the fit results for $R_s$ and $R_o$ (Table I). It seems the function (8) at free $a_3$ agrees better with experimental points at $\sqrt{s_{NN}} \leq 200$ GeV than that at fixed $a_3$. But large error bars do not allow to choice preferable curve unambiguously. Moreover the general function (8) underestimates $\Delta \tau$ in TeV-region significantly.

Figure 12 demonstrates the energy dependence of $\Delta \tau$ for (quasi)symmetric heavy ion collisions. The emission duration in these collisions is calculated based on known HBT-radii which are shown in Figs. 11 – d. Both the equations (8) and (11) are used for verification and increasing of reliability of results. The first approach is the same as in the previous study [1]. As expected the values of $V$ are similar for both the equations (8) and (11) that confirms the validity of the results from [1]. The energy dependence of estimations (11) for volume of emission region is shown in Fig. 12. It would be emphasized that values of $V$ calculated at $\sqrt{s_{NN}} = 62.4$ and 200 GeV from the recent STAR high-statistics results [10] agree better, especially at top RHIC energy, with results of other experiments (PHENIX and PHOBOS) than that in earlier study [1]. The values of $V$ (8) which are derived here from the STAR results obtained in the framework of the phase-I beam energy program (BES) at RHIC in energy domain $\sqrt{s_{NN}} = 7.7 – 39$ GeV form the trend lines some higher than most of the results from AGS and SPS. But on the other hand the spread of results at $\sqrt{s_{NN}} = 7 – 20$ GeV is within total errors if systematic uncertainties will be taken into account also. The values of $V$ (8) from the STAR results at energies $\sqrt{s_{NN}} = 7.7 – 39$ GeV (11) agree noticeably better with AGS and SPS results (Fig. 13) than that for $V$ (8). Results for volume at $\sqrt{s_{NN}} = 7.7 – 39$ GeV which are derived from the STAR high-statistics data (10) based on (8) as well as on (11) agree better with (quasi)linear growth $V$ with ln $\varepsilon$ from $\sqrt{s_{NN}} \approx 5$ GeV up to highest RHIC energy than that the earlier data at close energies. Smooth solid and dashed curves shown in Fig. 13 are calculated for $V$ from equation (11) and the fit results for $R_o$, $R_s$ (Table I). Both the curves are very close at $\sqrt{s_{NN}} \leq 200$ GeV but function (8) at $a_3 = 1.0$ underestimates $V$ (11) in TeV-region significantly. Therefore the general function (8) is the preferable approximation of the experimental $V$ (11) at $\sqrt{s_{NN}} \geq 5$ GeV.

Predictions for values of the femtoscopic observables from sets $G_m$, $m = 1, 2$ are obtained for heavy-ion mode energies of the LHC and the FCC project based on the fit results for the main HBT parameters. Estimations are shown in Table I for fits with inclusion of statistical errors, the second line for each collision energy corresponds to the using of the specific case of (8), the volume of homogeneity region is estimated with help of (11). Large uncertainties obtained for estimations based on the function (8) do not allow to distinguish predictions from (8) with free $a_3$ and with fixed $a_3 = 1.0$. One can expect the volume of homogeneity region $V \sim 6000$ fm$^3$ at $\sqrt{s_{NN}} = 5.52$ TeV (LHC).
and $V^{3+} \sim 9000 \text{ fm}^3$ at $\sqrt{s_{NN}} = 39.0 \text{ TeV}$ (FCC) based on the reasonable agreement between experimental data and solid curve at Fig. 3.

### Table II: Estimations for observables based on fit results for the data sample (i)

| $\sqrt{s_{NN}}$ (TeV) | $\lambda$ | $R_a$, fm | $R_o$, fm | $R_1$, fm | $R_o/R_a$ | $\Delta \tau$, fm/GeV | $V \times 10^{-3}$, fm$^3$ |
|------------------------|-----------|-----------|-----------|-----------|------------|----------------------|------------------------|
| 5.52                   | 0.41 ± 0.03 | 6.8 ± 1.9 | 6.3 ± 1.0 | 7.6 ± 1.5 | 0.9 ± 0.3 | –                    | 6 ± 3                  |
| 39.0                   | 0.362 ± 0.009 | 5.79 ± 0.10 | 6.49 ± 0.12 | 8.20 ± 0.16 | 1.12 ± 0.03 | 3.6 ± 0.4 | 4.33 ± 0.17          |
| 0.315 ± 0.011         | 6.11 ± 0.12 | 6.74 ± 0.15 | 9.04 ± 0.19 | 1.10 ± 0.03 | 3.5 ± 0.5 | –                    | 5.3 ± 0.2              |

Fig. 4 shows the energy dependence of $\lambda$ (a), scaled HBT-radii (b – d) and $R_o/R_a$ ratio (e) for both the symmetric and non-symmetric collisions of various nuclei. Fits of experimental dependencies for the data sample (ii) are made by [8] in the same energy domains and with the same error types as well as for the data sample (i) above. It seems the $\lambda$ value from the WA97 experiment [22] can not be excluded from the data sample (ii) because there are the STAR results $\lambda \sim 0.3$ for Cu+Cu collisions also (Fig. 4h). There are no physics reasons in order to exclude the points of these experiments from the fitted data sample (ii) in the case of all available nucleus-nucleus collisions. Furthermore the scaled value of longitudinal radius $R_o^0$ from [22] agrees better with results of other experiments at close energies (Fig. 4i) than that for the data sample (i). Therefore there is no exception of any experimental point from fitted ensemble for any HBT observable in Fig. 4 in contrast with the fitting procedure for the data sample (i).

The numerical values of fit parameters are presented in Table III where the second line for $\lambda$ and each normalized HBT radius corresponds to the approximation by specific case of [8]. Fit curves are shown in Fig. 4 by solid lines for [8] and by dashed lines for specific case of fit function at $a_2 = 1.0$ with taking into account statistical errors. Fit qualities are improved for $R_o^0$ in the case of total errors of experimental point and for $R_o^0$ at any error types of experimental point with respect to the corresponding fit results for the data sample (i) shown in Table II. There is dramatic growth of $\chi^2$/n.d.f. values for fits of $\lambda$ data (Fig. 4i) despite of qualitative agreement between smooth approximations and experimental $\lambda$ values for range $10 \lesssim \sqrt{s_{NN}} \lesssim 200 \text{ GeV}$. The fit by [8] at $a_3 = 1.0$ underestimates the $\lambda$ value at the LHC energy $\sqrt{s_{NN}} = 2.76 \text{ TeV}$ significantly. The $\lambda$ values for asymmetric nucleus-nucleus collisions at intermediate energies $\sqrt{s_{NN}} \lesssim 20 \text{ GeV}$ agree well with values of $\lambda$ in symmetric heavy ion collisions at close energies. On the other hand the $\lambda$ for Cu+Cu collisions is smaller systematically than $\lambda$ in Au+Au collisions in energy range $\sqrt{s_{NN}} = 62 – 200 \text{ GeV}$ (Fig. 4a). New experimental data are important for verification of the suggestion of separate dependencies $\lambda(\sqrt{s_{NN}})$ for moderate and heavy ion collisions. Also the development of some approach is required in order to account for type of colliding beams in the case of $\lambda$ parameter and improve quality of smooth approximation. In this case significant growth of $a_2$ as well as improvement of the fit quality at transition from statistical errors of experimental points to total errors in the data sample (ii) (Table III) is dominated by inclusion of estimations for systematic uncertainties for Cu+Cu collisions and/or Pb+Pb ones at $\sqrt{s_{NN}} = 2.76 \text{ TeV}$. Smooth curves for normalized HBT radii and ratio $R_o/R_a$ are in reasonable agreement with experimental dependencies in fitted domain of collision energies $\sqrt{s_{NN}} \geq 5 \text{ GeV}$ (Figs. 4b – e). Parameter values obtained for fit of $R_o^0$ with total uncertainties by [8] at $a_3 = 1.0$ are equal within errors with results from [22] accounting for that experimental results studied here are obtained at $\langle m_\perp \rangle \approx 1.75m_\pi$. Dramatic improvement of the fit qualities for scaled HBT radii at transition from the data sample (ii) with statistical errors to the data sample with total errors (Table III) is dominated mostly by the uncertainty in $r_0$ which leads to additional errors due to scaling [7]. At the same time inclusion of total uncertainties for Au+Au collisions at $\sqrt{s_{NN}} = 19.6 \text{ GeV}$ results in significant decreasing of $a_2$ parameter in the case of $R_o^0$ scaled radius.

The corresponding dependencies for $\delta^n$ and $V^n$ are demonstrated in Fig. 5 and Fig. 6 respectively. The definition [4] for source volume is used in Fig. 6 as well as in Fig. 5 above. As well as in [1] results for $\pi^+\pi^+$ pairs are shown in Figs. 5 and 6 also because femtoscopic parameters from the set $G_1$ depend on sign of electrical charge of secondary pions weakly. The relation $R_o < R_a$ is observed for $\approx 11\%$ of points in Fig. 5. In general the $\delta < 0$ can be possible in the model of opaque source with surface dominated emission [26, 27]. But possibly results should be similar for both the same ion beams and close kinematic regimes in various experiments. Therefore additional study is required in order to distinguish the physical and technique sources of negative values of the $\delta^n$ in Fig. 5 and to get a more definite explanation. The dependence $\delta^n(\sqrt{s_{NN}})$ is almost flat within large error bars in all energy domain under consideration. Taking into account the STAR high-statistics results [10] only one can see the indication on change of behavior of $\delta^n(\sqrt{s_{NN}})$ inside the range of collision energy $\sqrt{s_{NN}} = 11.5 – 19.6 \text{ GeV}$. This observation is in agreement with features of behavior of emission duration dependence on $\sqrt{s_{NN}}$ (Fig. 2) discussed above. The estimation of
energy range agrees well with results of several studies \(^{10, 28, 33}\) in the framework of the phase-I of BES program at RHIC which indicate on the transition from dominance of quark-gluon degrees of freedom to hadronic matter at \(\sqrt{s_{NN}} \lesssim 19.6\) GeV. But future precise measurements are crucially important for extraction of more definite physics conclusions. Smooth solid and dashed curves shown in Fig. 5 are calculated for \(\delta^n\) from the fit results for \(R_a^n\) and \(R_0^n\) (Table III). The situation is similar to that for \(\Delta t\): calculation based on the fit function \(\tilde{S}\) at free \(a_3\) agrees reasonably with experimental points at \(\sqrt{s_{NN}} \lesssim 200\) GeV but underestimates \(\delta^n\) in TeV-region significantly. The large errors in Fig. 5 for strongly asymmetric nuclear collisions is dominated by large difference of radii of colliding moderate and heavy nuclei and corresponding large uncertainty for \(R_A\). Smooth solid and dashed curves shown in Fig. 5 are calculated for \(V^n\) from equation \(4\) and the fit results for \(R_A^n\), \(R_0^n\) (Table III). The fit results for normalized HBT radii obtained with general function \(\tilde{S}\) lead to very good agreement between smooth curve and experimental data in TeV-region in contrast with the curve obtained from corresponding fit results for \(\tilde{S}\) at \(a_3 = 1.0\).

Estimations for \(\lambda, R_a/R_0\), and normalized femtoscopic parameters at the LHC and the FCC energies are shown in Table IV for fits of various nucleus-nucleus collisions with inclusion of statistical errors, the second line for each collision energy corresponds to the using of the specific case of \(\tilde{S}\) at \(a_3 = 1.0\), the volume of homogeneity region is estimated with help of 4. All the smooth approximations discussed above predict amplification of coherent pion emission with significant decreasing of \(\lambda\). Uncertainties are large for estimations obtained on the basis of results of fits by function \(\tilde{S}\) at free \(a_3\). Thus values of femtoscopic observables in Table IV are equal within errors for general and specific case of \(\tilde{S}\) at \(\sqrt{s_{NN}} = 5.52\) TeV (LHC) and \(\sqrt{s_{NN}} = 39.0\) TeV (FCC) as well as for estimations obtained on basis of the data sample (i) above.

| HBT parameter | Fit with statistical errors | Fit with total errors |
|---------------|-----------------------------|-----------------------|
| \(\lambda\)   | \(a_1\) \(a_2\) \(a_3\) \(\chi^2/\text{n.d.f.}\) | \(a_1\) \(a_2\) \(a_3\) \(\chi^2/\text{n.d.f.}\) |
| \(R_a^n\)     | 1.21 \pm 0.09 -0.30 \pm 0.04 0.38 \pm 0.04 3656/29 0.60 \pm 0.02 \(-0.014 \pm 0.008\) 1.3 \pm 0.2 780/29 |
|               | 0.717 \pm 0.003 -0.051 \pm 0.001 1.0 (fixed) 3786/23 0.631 \pm 0.005 \(-0.034 \pm 0.001\) 1.0 (fixed) 706/23 |
| \(R_0^n\)     | 0.656 \pm 0.002 \((6 \pm 3) \times 10^{-5}\) 3.11 \pm 0.19 195/25 0.63 \pm 0.02 \((6 \pm 5) \times 10^{-4}\) 2.4 \pm 0.9 26.8/25 |
|               | 0.599 \pm 0.003 0.019 \pm 0.001 1.0 (fixed) 280/26 0.56 \pm 0.03 0.029 \pm 0.008 1.0 (fixed) 28.9/26 |
| \(R_1^n\)     | 0.10 \pm 0.02 6.3 \pm 1.7 0.068 \pm 0.006 402/25 0.019 \pm 0.003 30 \pm 9 0.12 \pm 0.05 23.9/26 |
|               | 0.758 \pm 0.004 0.008 \pm 0.001 1.0 (fixed) 415/26 0.67 \pm 0.04 0.017 \pm 0.008 1.0 (fixed) 24.4/26 |
|               | 0.022 \pm 0.002 23 \pm 3 0.258 \pm 0.005 502/25 0.23 \pm 0.04 0.8 \pm 0.2 0.57 \pm 0.05 66.0/25 |
|               | 0.634 \pm 0.004 0.043 \pm 0.001 1.0 (fixed) 615/26 0.47 \pm 0.03 0.089 \pm 0.014 1.0 (fixed) 66.7/26 |

TABLE IV: Estimations for observables based on fit results for the data sample (ii)

| \(\sqrt{s_{NN}}\) | HBT parameter |
|-----------------|----------------|
| TeV             | \(\lambda\) \(R_a^n\) \(R_0^n\) \(R_1^n\) \(R_0/R_a\) \(\delta^n\) \(V^n\) |
| 5.52            | 0.16 \pm 0.19 0.9 \pm 0.2 0.8 \pm 0.3 1.06 \pm 0.16 0.9 \pm 0.4 \(-0.2 \pm 0.6\) 3.5 \pm 1.6 |
|                 | 0.091 \pm 0.004 0.792 \pm 0.009 0.860 \pm 0.010 1.099 \pm 0.013 1.086 \pm 0.018 0.11 \pm 0.02 2.59 \pm 0.07 |
| 39.0            | 0.07 \pm 0.21 1.2 \pm 0.4 0.9 \pm 0.3 1.11 \pm 0.16 0.7 \pm 0.3 \(-0.7 \pm 1.1\) 6 \pm 4 |
|                 | 0.836 \pm 0.011 0.883 \pm 0.012 1.205 \pm 0.015 1.06 \pm 0.02 0.08 \pm 0.03 3.17 \pm 0.09 |

The energy dependencies for sets \(G_m, m = 1, 2\) of femtoscopic parameters with taking into account the scaling relation \(7\) and the high-statistics STAR data \(10\) demonstrate the reasonable agreement between values of parameters obtained for interactions of various ions (shown in Figs. 4 - 6). The observation confirms the suggestion \(1\) that normalized femtoscopic parameters allow to unite the study both the symmetric and asymmetric nucleus-nucleus collisions in the framework of united approach. This qualitative suggestion is confirmed indirectly by recent study of two-pion correlations in the collisions of the lightest nucleus (d) with heavy ion (Au) at RHIC. Estimations of space-time extent of the pion emission source in d+Au collisions at top RHIC energy \(14\) in dependence on kinematic observables (collision centrality, the mean transverse momentum for pion pairs) indicate similar patterns with corresponding dependencies in Au+Au collisions and indicate on similarity in expansion dynamics in collisions of various systems (d+Au and Au+Au at RHIC, p+Pb and Pb+Pb at LHC). The scaling results for some radii indicate...
that hydrodynamic-like collective expansion is driven by final-state rescattering effects \[34\]. On the other hand the normalized femtoscopic parameters allow to get the common kinematic dependencies only without any additional information about possible general dynamic features in different collisions. Thus the hypothesis discussed above is qualitative only. The future quantitative theoretical and phenomenological studies are essential for verification of general features of soft stage dynamics for different collisions at high energies.

IV. SUMMARY

The following conclusions can be obtained by summarizing of the basic results of the present study.

Energy dependence is investigated for range of all experimentally available initial energies and for estimations of the main femtoscopic parameters from set the \( G_1 \) (\( \lambda \) and radii) derived in the framework of Gauss approach as well as for the set of important additional observables \( G_2 \) contains ratio of transverse radii, emission duration (or \( \delta \)) and HBT volume. There is no dramatic change of femtoscopic parameter values with increasing of \( s_{NN} \) in domain of collision energies \( s_{NN} \geq 5 \) GeV. The estimation of emission duration of pions is about 2 – 4 fm/c for any energies under consideration. The energy dependence is almost flat for both the emission duration and the \( \delta^n \) parameter within large error bars. The indication on possible curve knee at \( s_{NN} \sim 10 – 20 \) GeV obtained in the STAR high-statistics data agree with other results in the framework of the phase-I of the beam energy scan program at RHIC. But additional precise measurements are crucially important at various \( s_{NN} \) in order to confirm this feature in energy dependence of additional femtoscopic parameters (\( R_0/R_s, \Delta \tau, \delta^n \)).

Analytic function is suggested for approximation of energy dependence of main HBT parameters. The fit curves demonstrate qualitative agreement with experimental data for \( \lambda \) at all available collision energies and for both the absolute and normalized HBT radii in energy domain \( s_{NN} \geq 5 \) GeV. Reasonable fit qualities are obtained for HBT radii at approximation of experimental points with total errors. Smooth curves calculated for energy dependence of the set \( G_2 \) of additional femtoscopic parameters agree reasonably with corresponding experimental data in the most cases. Estimations of femtoscopic observables are obtained on the basis of the fit results for energies of the LHC and the FCC project. For multi-TeV energy domain the emission region of pions will be characterized by decreased chaoticity parameter, linear sizes about 8.5 – 9.5 fm in longitudinal direction and 7 – 8 fm in transverse plane, volume of about \( 10^4 \) fm\(^3\).

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FIG. 1: Dependence of chaoticity parameter (a), HBT-radii (b – d) and ratio $R_o/R_a$ (e) on collision energy for central heavy ion Au+Au, Au+Pb, Pb+Pb interactions at midrapidity and $(k_\perp) \simeq 0.2$ GeV/$c$ [1, 10]. Experimental results are demonstrated for pairs of $\pi^-$ mesons (in the cases of ALICE, STAR at $\sqrt{s_{NN}} = 7.7 - 62.4$ and 200 GeV – for $\pi^+\pi^-$ pairs) and for standard Coulomb correction $P_C^{(1)}(q)$ (in the cases of ALICE, NA44, NA45, PHOBOS and STAR at $\sqrt{s_{NN}} = 7.7, 11.5 - 62.4$ and 200 GeV – for correction $P_C^{(3)}$). Statistical errors are shown (for NA44 – total uncertainties). The solid lines (a – d) correspond to the fits by function $G_i \propto \ln \varepsilon$, $i = 1 - 4$. Fitted data samples for $\lambda$ (a) and for $R_t$ (d) do not include the point of the WA97 experiment [23] while the fits for transverse HBT radii (b, c) are shown for samples with point from [23]. Smooth solid and dashed curves at (e) correspond to the ratio $R_o/R_a$ calculated from the fit results for $R_a$ and $R_o$, dotted line is the level $R_o/R_a = 1$. 
FIG. 2: Energy dependence of emission duration for secondary charged pions in central heavy ion collisions Au+Au, Au+Pb, Pb+Pb in midrapidity region and at \( \langle k_\perp \rangle \simeq 0.2 \text{ GeV/c} \). Experimental results are shown for the same particle types and Coulomb corrections as well as in Fig. 1. Error bars are only statistical (for NA44 – total uncertainties). Smooth curves are derived from (6) and the fit results for \( R_s, R_o \) without the point of the WA97 experiment [23]. The solid line corresponds to the fits of HBT radii by function (5) and dashed line – to the fits by specific case \( R_i \propto \ln \varepsilon, i = s, o \).
FIG. 3: Energy dependence of volume of emission region at freeze-out for secondary charged pions in central heavy ion collisions \( \text{Au+Au, Au+Pb, Pb+Pb} \) in midrapidity region and at \( \langle k_\perp \rangle \simeq 0.2 \text{ GeV/c} \). The equation (4) is used for calculation of volume values. Experimental results are shown for the same particle types and Coulomb corrections as well as in Fig. 1. Error bars are only statistical (for NA44 – total uncertainties). Smooth curves are derived from (4) and the fit results for \( R_s, R_l \) with taking into account the point of the WA97 experiment [23]. The solid line corresponds to the fits of HBT radii by function (8) and dashed line – to the fits by specific case \( R_i \propto \ln \varepsilon, i = s, l \).
FIG. 4: Energy dependence of λ parameter (a), normalized HBT-radii (b – d) and ratio $R_o/R_s$ (e) in various nucleus-nucleus collisions at $\langle k_{\perp} \rangle \simeq 0.2 \text{ GeV/c}$ [1]. Experimental results are shown for central collisions (for minimum bias event in the case of E802 for Al+Si), for pairs of $\pi^-$ mesons (in the cases of ALICE and STAR for both the Cu+Cu and Au+Au at $\sqrt{s_{NN}} = 7.7 - 62.4$ and 200 GeV – for $\pi^+\pi^-$ pairs, E802 for Al+Si, NA44 for S+Pb – for pairs of $\pi^+$ mesons) and for standard Coulomb correction $P_C^{(1)}(q)$ (in the cases of ALICE, NA44, PHOBOS, STAR for both the Cu+Cu and Au+Au at $\sqrt{s_{NN}} = 7.7, 11.5 - 62.4$ and 200 GeV – for correction $P_C^{(3)}$). Statistical errors are shown (for NA44 – total uncertainties). The solid lines (a – d) correspond to the fits by function (8) and dashed lines – to the fits by specific case of (8) at fixed $a_3 = 1.0$. Smooth solid and dashed curves at (e) correspond to the ratio $R_o/R_s$ calculated from the fit results for $R^n_o$ and $R^n_s$, dotted line is the level $R_o/R_s = 1$. 
FIG. 5: Dependence of scaled difference of squares of transverse radii on beam energy for emission region of secondary charged pion in various nucleus-nucleus collisions at \( \langle k_{\perp} \rangle \approx 0.2 \text{ GeV/c} \). Experimental results are shown for the same particle types and Coulomb corrections as well as in Fig. 4. Error bars are only statistical (for NA44 – total uncertainties). Dotted line is the level \( \delta^n = 0 \). Smooth curves are derived from (7) and the fit results for \( R_n^s, R_n^o \). The solid line corresponds to the fits of normalized HBT radii by function (8) and dashed line – to the fits by specific case \( R_i^\alpha \propto \ln \varepsilon, i = s, o \).
FIG. 6: Energy dependence of normalized volume of emission region at freeze-out for secondary charged pions in various nucleus-nucleus collisions at \( \langle k_\perp \rangle \simeq 0.2 \text{ GeV}/c \). The equation (4) is used for calculation of volume values. Experimental results are shown for the same collision, particle and Coulomb correction types as well as in Fig. 4. Error bars are only statistical (for NA44 – total uncertainties). Smooth curves are derived from (4) and the fit results for \( R_n^s, R_l^n \). The solid line corresponds to the fits of normalized HBT radii by function (8) and dashed line – to the fits by specific case \( R_i^n \propto \ln \varepsilon, i = s, l \).