Search for a dark photon at \textit{BABAR}

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We present the results of a search for a dark photon decaying to $e^+e^-$ or $\mu^+\mu^-$ in the \textit{BABAR} detector [1]. We find no evidence for such a dark photon and set upper limits on the dark mixing as a function mass, tightening the constraints on the parameter space of several dark sector theories.

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1 Introduction

The notion of dark matter was introduced over 80 years ago by Zwicky [2]. In the intervening years, the astrophysical evidence for dark matter has mounted, but the precise nature of this matter is still not understood. In the 1980’s, the idea of a new force corresponding to a new $U(1)'$ symmetry was introduced [3]. This idea was adapted to the dark sector, suggesting a dark gauge boson. Recent astronomical observations, particularly of an excess of electrons from cosmic rays [4], has raised interest in the possibility of a dark gauge boson that may couple via kinetic mixing to Standard Model (SM) particles and could have a mass in the few GeV/$c^2$ range. The gauge boson is generally referred to as a dark photon ($A'$) and can be considered to appear in any process in which a SM photon could appear. The coupling constant for such a dark photon vertex is taken to be $\alpha' = \epsilon^2 \alpha$. Current opinion generally places $\epsilon$ in the range $10^{-2} - 10^{-7}$.

2 The BABAR detector and data sets

The BABAR detector is a general-purpose collider detector consisting of five subdetectors: the Silicon Vertex Tracker (SVT); the Drift Chamber (DCH); the Detector for Internally Reflected Cherenkov radiation (DIRC); the Electromagnetic Calorimeter (EMC); and the Instrumented Flux Return (IFR). The SVT, DCH, DIRC, and EMC reside inside a superconducting solenoid that provides a uniform 1.5-Tesla axial magnetic field. The SVT and DCH combine to provide tracking for charged particles and particle identification through specific ionization measurements. The SVT also provides excellent vertex resolution while the 40-layer DCH allows measurement of the curvature of tracks in the magnetic field, providing transverse momentum determination. The DIRC is used for charged particle identification, especially for separating pions and kaons, and the EMC detects positions and energies of photons and electrons. A detailed description of the BABAR detector can be found in [5].

The data sample used in this analysis corresponds to an integrated luminosity of $514\text{fb}^{-1}$ recorded by the BABAR experiment using the PEP-II $e^+e^-$ storage rings at the SLAC National Accelerator Laboratory. This data set was recorded between 1999 and 2008 and contains data primarily from the $\Upsilon(4S)$ as well as the $\Upsilon(3S)$ and $\Upsilon(2S)$, and data taken off-peak. Off-peak data was taken 40 MeV below the $\Upsilon(4S)$ and 30 MeV below the $\Upsilon(3S)$ and $\Upsilon(2S)$.

3 Analysis Procedure

In this analysis, we search for the process $e^+e^- \rightarrow \gamma A'$ with subsequent decay $A' \rightarrow e^+e^-, \mu^+\mu^-$. The topology of this event is relatively simple and experimentally clean.
We search for events with one photon and two oppositely charged leptons. We search for dark photon candidates in the mass range $0.02 \text{ GeV}/c^2 - 10.2 \text{ GeV}/c^2$.

The signal and background efficiencies are studied using a variety of simulations. Signal events are simulated with MadGraph [6] for 35 different hypothetical $A'$ masses. Primary non-resonant backgrounds to the electron modes are simulated with BHWIDE [7] while those for muon modes are simulated with KK2f [8]. Simulation of the decays from resonances such as $\Upsilon(nS)$, $j/\psi$, etc. formed in initial state radiation is handled using structure function techniques [9]. A GEANT4-based [10] simulation is used to model material interactions and detector response. A subset of approximately 5% of the data sets is used to tune selection criteria and study systematic effects. This data is then discarded and not used in the final results.

The photon used for event selection must have an energy greater than 200 MeV. We allow extra photons as long as their energies do not add to more than 200 MeV. We consider only events in which the invariant mass of the photon and charged particles is consistent with the beam energy spread and in which the interaction point of the particles are consistent with the beam spot location. We accept only those events in which the candidate kinematic vertex fit has a $\chi^2 < 30$. We require at least one charged particle be well-identified as an electron and one loosely consistent with the electron hypothesis, or both tracks to be well-identified as muons.

A significant background comes from radiative Bhabhas – events of the type $e^+e^- \rightarrow e^+e^-\gamma$ – in which one of the final state leptons has radiated the photon. To suppress this background, we additionally require that the final state electron and positron form an angle with their respective incoming beams such that the cosine of this angle is less than 0.5.

A sizable background from converted photons in $e^+e^- \rightarrow \gamma\gamma$, $\gamma \rightarrow e^+e^-$ events still remains for low $e^+e^-$ invariant mass. A neural network is employed to further reduce this background using five kinematic variables. The selected cut on the neural net output keeps over 70% of remaining signal while removing 99.7% of the conversion background in the $m_{A'} \approx 20 - 50 \text{ MeV}/c^2$ range.

The dielectron and reduced dimuon mass distributions are displayed in Figure 1, overlaid with the predictions of various simulated SM processes, including several resonances. The reduced dimuon mass, $m_R = \sqrt{m_{\mu\mu}^2 - 4m_{\mu}^2}$, is used in place of the invariant mass because it is easier to model near threshold. The data are seen to match well to the predictions from simulation, except in the very low $e^+e^-$ region, where BHWIDE is known to have difficulties.

Signal efficiency is determined from Monte Carlo simulation and found to be approximately 15% for the dielectron channel and 35% for the dimuon channel. The difference is mainly attributable to trigger efficiencies, where a trigger filter reduces the rate of radiative Bhabha events.

Signal yield as a function of dark photon mass is extracted by performing a series of independent fits to the mass spectra for each beam energy. The fits are per-
Figure 1: Distribution of the final dielectron (left) and reduced dimuon invariant masses (right) as black points, overlaid with the predictions of various SM processes (red) and ISR production of resonances (green).

formed in the range $0.02 \text{ GeV}/c^2 < m_{A'} < 10.2 \text{ GeV}/c^2$ for the dielectron sample and $0.212 \text{ GeV}/c^2 < m_{A'} < 10.2 \text{ GeV}/c^2$ for the dimuon sample. We search for the dark photon in steps of approximately half the dark photon mass resolution with each fit window typically 30 times broader than the the signal resolution. The signal resolution is estimated from a range of simulated $A'$ samples and varies from 1.5 MeV/$c^2$ to 8 MeV/$c^2$. A total of 5704 fits are performed in the dielectron channel and 5370 are performed in the dimuon channel. The fit bias is estimated from a large collection of toy-Monte Carlo pseudo-experiments and found to be negligible.

Figure 2: Fitted cross-sections for dielectron (top) and dimuon (bottom) modes, together with their statistical significances, as functions of the dark photon mass. Grey bands indicate mass regions that are excluded from the analysis.

A statistical significance is assigned to each fit as $S = \sqrt{2 \log(\mathcal{L}/\mathcal{L}_0)}$, where $\mathcal{L}$ is the likelihood for the fit with background plus signal and $\mathcal{L}_0$ is the likelihood with a background function only. The statistical significance is assigned a sign that is the
same as the sign of the area of the fitted signal function. The largest local significances are $3.4\sigma$ ($2.9\sigma$) near $m_{A'} = 7.02$ GeV/$c^2$ ($6.09$ GeV/$c^2$) for the dielectron (dimuon) final state. Including trial factors, studied via a large number of toy Monte Carlo experiments, we find the corresponding p-value is 0.57 (0.94), consistent with the null hypothesis. We extract the cross-section as a function of dark photon mass by dividing the fitted yields by efficiency and luminosity. The cross-sections are shown in Figure 2 and the distributions of statistical significance are shown in Figure 3.

Using the expected dark photon branching fractions for $A' \rightarrow l^+l^-$ [11], our results can be combined and translated into a 90% confidence level upper limits on the kinetic mixing strength. These results [12] are shown in Figure 4, which illustrates clearly that our measurements significantly increased the constraints on the dark sector parameter space.

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Figure 4: Upper limits at 90% confidence level on the kinetic mixing strength, $\epsilon$, as a function of dark photon mass.

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