Editorial: Quasars at All Cosmic Epochs

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Editorial on the Research Topic

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Accretion onto massive black holes—and its associated manifestation as nuclear activity in galaxies—is among the most remarkable phenomena occurring in the Universe, and likely a key factor in galaxy evolution over a broad range of cosmic epochs. From a time when it was believed that accreting black holes were isolated sources affecting their host only within a limited sphere of influence, we have just begun to appreciate the rich and complex phenomenology induced by active nuclei in their host galaxies due to their radiative and mechanical output. Nuclear activity also presents a unique opportunity to probe spatial scales that we cannot yet resolve, as accretion related phenomena provide an immensely rich phenomenology extending from the longest radio wavelengths to the hardest γ-rays within a few parsec from the central black hole. A world of accreting black holes completely hidden at present will emerge in the next few years from the development of gravitational wave observatories. In late 2016, the time was ripe for a critical assessment of our understanding of quasars as accreting black holes and of their evolution across the cosmic time, with an eye on future developments.

This Research Topic is based on contributions presented at a meeting held in Padova in April 2017. Close to 170 participants convened in Padova from research institutes and Universities around the world. The Padova meeting was one of the largest conferences on active galactic nuclei in recent years. With 159 contributions (88 talks), gender balance was achieved (within 2 σ confidence-level stochastic fluctuations expected from binomial statistics), as the percentage of women participant was 43%, the percentage of women first author in a contribution 48%, and the percentage of women speakers was 55%.

The meeting highlighted progress—both observational and theoretical—achieved over the last decade. At the same time, many shortcomings, wishful thinking, and never critically-reanalyzed prejudices emerged in discussions among the participants about our view of quasars and nuclear activity in general.

The organization of the e-book produced from this Research Topic reflects the temporal sequence of the meeting sessions. Each day was organized around a broad theme, focused on one or more overarching questions. On the first day, we aimed at a comprehensive overview of the main observational aspects of active galactic nuclei (AGN), with a focus on several key questions (Chapter 1). Contributions from this first day, organized into Chapter 1, discuss the most relevant photometric, spectroscopic, polarimetric, and variability observations over the full spectral energy distribution (SED), together with their interpretation, with attention to both statistical results and case studies. Overarching issues are some still-enigmatic aspects of the SED, systematic organization, and contextualization of observational properties, radio properties of
jetted and non-jetted quasars, as well as selection effects that still affect major surveys. Chapter 2 contains contributions on accretion disk structure and wind and jet launching processes. The overarching theoretical issues (which remain open to-date) are the connections between disk structure, relativistic ejections, and continuum and line emitting regions. A satisfactory model of the broad line emitting region of quasars—to which almost 2,000 papers were dedicated over 40 years and to which a large part of the Conference was devoted—could pave the road to the use of quasars as distance indicators but remains as yet incomplete. The contributions in Chapter 3 attempt to connect theory and observation. The overarching questions remain, why do the radio properties of quasars divide the quasar population into the two categories radio quiet and radio loud? What is the connection between disk structure, relativistic ejections, and emitting regions structure? We think that one of the more remarkable outcomes of the conference was the realization of the many observational manifestations of accretion in quasars in a broad range of luminosity and cosmic epochs, and of their dependence on accretion parameters such as mass, dimensionless accretion rate, and spin. As the accreting black hole systems only possess axial symmetry and not spherical symmetry, many observational properties are affected by the viewing angle. Our ability to connect observational to physical and aspect parameters is still lacunose but, if filled, the importance could be comparable to the development of the Hertzsprung-Russell diagrams for stars. We miss an equivalent for quasars, although prospective improvements are in sight. Significant attention was devoted to a promising approach, developed by Jack Sulentic, his collaborators, and other research groups, based on the analysis of the “first eigenvector” of a 4D correlation space of quasar properties.

Chapter 4 and in part Chapter 5 are devoted to the role of nuclear activity in shaping the evolution of galaxies. Feedback and environment of active nuclei are seen in the suggested black hole and host galaxy coevolution, but open issues still involve the interplay of black hole fueling (strongly influenced by the environment over cosmic ages), star formation and feedback. The overarching question posed at the meeting was: which are the strongest evidences and the state-of-the-art modeling, and which are the tests and surveys, both observational and theoretical, that can lead to progresses in our understanding of fueling and feedback processes on all scales?

Chapter 5 discussed some interesting developments having taken place in the last two years, as we seek to learn from the observations of the increasing number of high redshift quasars. An open issue remains the role of quasars at the epoch of re-ionization. A more basic question is the role of first seed quasars and their relation to galaxy formation. The overarching question remains whether we understand quasar evolution beyond selection effects.

The last contributions deal with the possibility to exploit quasars as helpful distance indicators, a possibility that was found closely linked to an improvement in our understanding of the quasar emitting region structure and dynamics. Major issues and the recent developments, including several research lines aimed at answering some of the overarching questions, are critically evaluated in the conference closing remarks by Hagai Netzer (Chapter 6).

DATA AVAILABILITY STATEMENT

Slides of most of the presentations are available at zenodo.org on the Quasars at all cosmic epochs page.

AUTHOR CONTRIBUTIONS

PM wrote the paper. The other authors gave significant contributions.

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