Nucleation versus Spinodal decomposition in a first order quark hadron phase transition

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We investigate the scenario of homogeneous nucleation for a first order quark-hadron phase transition in a rapidly expanding background of quark gluon plasma. Using an improved preexponential factor for homogeneous nucleation rate, we solve a set of coupled equations to study the hadronization and the hydrodynamical evolution of the matter. It is found that significant supercooling is possible before hadronization begins. This study also suggests that spinodal decomposition competes with nucleation and may provide an alternative mechanism for phase conversion particularly if the transition is strong enough and the medium is nonviscous. For weak enough transition, the phase conversion may still proceed via homogeneous nucleation.

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I. INTRODUCTION

The hadronization of Quark Gluon Plasma (QGP) possibly produced in the early universe or expected to be formed in relativistic heavy-ion collisions has been the focus of much attention during the past few years. The quark gluon plasma (QGP), if formed, would expand hydrodynamically and would cool down until it reaches a critical temperature \( T_C \) where a phase transition from the quark matter to the hadron matter begins. Although the plasma has to hadronize, the mechanism of hadronization still remains an open question. The percolation model calculations are used in the case of a second order phase transition. In a first order scenario, the dynamics of the phase transition has been modeled in several ways. In the most idealized picture, the temperature of the plasma is held fixed at \( T = T_C \) until the phase conversion is completely over. Assuming an isentropic expansion in \( (1+1) \) dimension, the hadronization in the above picture gets completed at \( \tau_h = r \tau_C \) where \( r \) is the ratio of the degrees of freedom of QGP and hadronic phases and \( \tau_C \) is the proper time at which the QGP cools down to the temperature \( T_C \). In reality, a first order phase transition is characterized by a large nucleation barrier that separates the two phases at \( T = T_C \) and the hadronization will not begin unless the matter supercools below \( T_C \). Alternatively, the theory of homogeneous nucleation has been invoked to study the first order phase transition which is more realistic than the above idealized adiabatic scenario and has been in use for quite some time in the cosmological context. In this picture, the transition is initiated by the nucleation of critical-size hadron bubbles from a supercooled metastable QGP phase. These hadron bubbles can grow against surface tension, converting the QGP phase into the hadron phase as the temperature drops below the critical temperature, \( T_C \). For strong enough transition, the large amplitude fluctuations are suppressed so that the nucleation begins from a (nearly) homogeneous background of supercooled metastable phase. This has been the basis of homogeneous nucleation theory based on which the QCD phase transition has been studied extensively. However, for a weak enough transition, the matter may not remain in a pure homogeneous state even at \( T = T_C \) due to the transitions that may occur above \( T_C \). Evidence of such pretransitional phenomena are quite common in several condensed matter studies particularly in the case of isotropic to nematic transition in liquid crystals. The possibility of such a transition had also been investigated during the cosmological electroweak phase transition in the early universe. Due to high rate of thermal fluctuations and slow cooling of the universe, a strong phase mixing is expected even at temperature above \( T_C \) where the new phase is highly metastable. In such situation, the phase transition either proceeds through percolation, if not, the dynamics of the first order transition is going to be quite different than the standard theory of homogeneous nucleation. Similar phenomena is also expected in the case of a quark hadron phase transition both in the early universe as well as in the plasma expected to be formed during the relativistic heavy ion collisions. An ideal quark gluon plasma in \( (1+1) \) dimension expands as per the Bjorken scaling where \( T^3 \tau \) is constant. This scaling would mean that the rate of change of temperature is higher for the plasma at earlier time as compared to the plasma at later time. In the case of early universe, due to high initial temperature, the rate of cooling becomes quite slow by the time it approaches \( T_C \). However, the QGP produced during the heavy ion collisions will have smaller initial temperature and will cool much more rapidly as compared to the early universe. An-
other difference as compared to the early universe is that
the QGP produced at RHIC and LHC energies may at-
tain kinetic equilibrium in a very short time \( t_{\text{CM}} \) of the order of a few fm/c but will remain far off from chemical equilibration \( [19,20] \). Such a plasma will be more gluon rich and many more quark and anti-quark pairs will be needed before the plasma achieves chemical equilibrium. The chemi-
cally unsaturated plasma will cool still at a faster rate since additional energy will be consumed in approaching the chemical equilibration \( [19,20] \). Recently, we have in-
vestigated the effect of thermal fluctuations leading to
phase mixing by modeling them as subcritical hadron
bubbles \( [19] \). Although the equilibrium density distri-
bution of these subcritical bubbles can be quite large, their equilibration time scale is larger than the cooling time scale for the QGP. As a consequence, for RHIC and LHC energies, they will not build up to a level capable of
modifying the predictions from homogeneous nucleation theory: the QGP has to supercool and the hadronization may still proceed through the nucleation of critical bub-
bles within a nearly homogeneous background of quark-
gluon plasma.

For the QGP formed at RHIC and LHC energies, the
question to be asked is how long the plasma would con-
tinue to supercool. It is expected that the plasma will
supercool up to some temperature \( T_{\text{CM}} \) until the density of nucleated hadron bubbles become sufficient to heat up the medium due to the release of latent heat \( [19] \). Since the medium is heated up, the nucleation is again switched off at some point and the hadronization continues due to
the growth of previously nucleated hadron bubbles. Re-
lated to this another crucial aspect which needs investiga-
tion is that how fast the barrier between the metastable
QGP phase and the stable hadron phase decreases as
the system supercools. For strong enough supercooling, the barrier between the stable and metastable minima
completely vanishes leading to a point of inflection at
\( T = T_{I} \) known as spinodal instability. Thus, the rapidly
quenched system leaves the region of metastability and
enters the highly unstable spinodal region before a
substantial amount of nucleation begins. The spinodal decomposition has also been suggested as a possible mecha-
nism of phase conversion for a rapidly expanding system of quarks and gluons \( [21] \). The second scenario of phase
transition may lead to coherent pion emission due to the
formation of disoriented chiral condensates \( [21] \).

For the case of QGP produced in heavy ion collision,
the route to hadronization either through nucleation or
through spinodal decomposition depends sensitively on
several factors like nucleation rate, effective potential
used to model the dynamics of the phase transition and
also on the hydrodynamical evolution that decides the ex-
pansion rate of the matter. The homogeneous nucleation
rate is estimated from the factor
\[
I = I_{0} \exp(-\Delta F_{C}/T)
\]
where \( I_{0} \) is the product of dynamical and statistical prefactors and \( \Delta F_{C} \) is the minimum energy needed to cre-
ate a critical hadron bubble. The exponential factor
dominates the rate at temperature close to \( T_{C} \) whereas

the prefactor \( I_{0} \) plays the significant role away from \( T_{C} \).

The most commonly used expression for the prefactor is
\( I_{0} \sim T^{4} \) \( [22] \) which may be alright for the case of early
universe but not for the heavy ion collisions where we
expect large amount of supercooling. The theory of ho-
ogeneous nucleation assumed significance for the study
of phase transition in QGP produced in the Relativistic
heavy ion collision after Csernai and Kapusta \( [1] \) derived
an improved prefactor using coarse grained field theory.
An important aspect of their formalism is the presence of
viscosity which is essential for the removal of latent heat
away from the hadron bubble. The prefactor vanishes
for an ideal plasma. This would suggest that a dissipative
dynamics should be followed for the phase transition to be completed through nucleation \( [23] \). In a subsequent
work \( [8] \), Ruggeri and Friedman, on the other hand, de-
\( \text{fined a prefactor which does not vanish in the limit of zero}
viscosity. Recently, we have also derived a prefactor
which is more general and also reproduces the Csernai-
Kapusta and Ruggeri-Friedman results in the limiting
cases \( [23] \). The prefactor \( I_{0} \) in all these estimates is found
to be less as compared to \( T^{4} \). In this work, we estimate
the nucleation rate using an improved prefactor as de-
\( \text{rived in Ref.}\ [23] \). The argument \( \Delta F_{C}/T \) that appears
in the exponential depends on the phenomenological po-
tential model used to describe a first order phase tran-
sition. Since the lattice predictions \( [24,25] \) are not yet
conclusive about the nature and the strength of the tran-
sition, we use an effective potential that follows the bag
equation of state and also covers a wide range from very
weak to very strong first order phase transitions. Finally,
we study amount of supercooling and rate of hadroniza-
tion by solving self consistently the nucleation rate along
with energy momentum conserving hydrodynamic equa-
tion both for dissipative and nondissipative plasma. The
dissipative plasma makes the evolution slow and also gen-
erates extra entropy. From this study we show that spin-
odal decomposition may compete with homogeneous nu-
cleation if the plasma is nondissipative and the transition
is relatively stronger. Since the amount of supercooling
is less in case of a dissipative plasma, the phase conver-
sion may still proceed through the nucleation of critical
size hadron bubbles from a homogeneous background of
supercooled quark gluon plasma.

The paper is organized as follows: In Sec. II, we briefly
review the bag equation of state and an effective potent-
tial used to model the phase transition. The nucleation
rate with various prefactors are discussed in Sec. III. Fi-
\( \text{nally, we present our results in Sec. IV followed by the}
conclusion in Sec. V.

II. THE EFFECTIVE POTENTIAL AND THE
EQUATION OF STATE

The spinodal instability corresponds to the inflection
point in the effective potential (below \( T_{C} \)) used to study
the dynamics of a phase transition. Obviously, the temperature $T_S$ at which the instability occurs will depend on the parameters of the effective potential. Since the order as well as strength of the quark hadron phase transition is still an unsettled issue, we consider a more generic form of the potential which covers a wide range from very strong to very weak first order phase transition. The parameters of the effective potential are determined from the physical observables like the surface tension $\sigma$, the correlation length $\xi$ and also from the requirement that the potential difference between the two minima should correspond to the pressure difference between the quark and hadron phases. We follow the bag equation of state to estimate various thermodynamical observables. For example, the pressure can be estimated from the relation $P = T \Sigma \ln Z_i / V - B$ where the partition function $\ln Z_i$ for a single fermion or boson is given by \[27\]

$$ \ln Z_i = \pm \frac{g V}{2\pi^2} \int_0^\infty dk k^2 \left[ \ln \left( 1 \pm e^{-\sqrt{k^2 + m_i^2}/T} \right) \right], \quad (1) $$

where the + and − stand for fermions and bosons respectively. In the above, the chemical potential $\mu$ has been set to zero. Assuming, a baryon free quark gluon plasma which consists of $u$, $d$, and $s$ quarks and gluons, the total pressure can be written as

$$ p_q = \frac{16 \pi^2}{90} T^4 + \sum_{i=1}^{n_f} \frac{16 T}{2\pi^2} \int_0^\infty dk k^2 \ln \left( 1 + e^{-\sqrt{k^2 + m_i^2}/T} \right) - B, $$

$$ = \frac{16 \pi^2}{90} T^4 + \frac{7 \pi^2}{60} T^4 \sum_{i=1}^{n_f} f(m_i/T) - B, \quad (2) $$

where $n_f$ is the number of flavours and $m_i$, $(i=u,d,s)$ are their masses. The function $f(m_i/T) = (360/7\pi^4) \int_0^\infty du u^2 \ln \left( 1 + e^{-\sqrt{u^2 + (m_i/T)^2}} \right)$ and $B$ is the bag constant. In Eq. $2$, the first term is due to massless gluons while the second needs to be evaluated for different physical masses of the quarks. If the quarks are taken as massless, we get

$$ p_q = a_q T^4 - B \quad \text{where} \quad a_q = \frac{19 \pi^2}{36}. \quad (3) $$

If the physical masses of quarks are used as $m_u \simeq m_d \simeq 8$ MeV and $m_s \simeq 160$ MeV, then at the critical temperature ($T_C = 160$ MeV) we get $f(m_u/T) \simeq 1$ and $f(m_s/T) \simeq 0.8$. Thus,

$$ a_q = \frac{37 \pi^2}{90} + 0.8 \frac{7 \pi^2}{60}. \quad (4) $$

which is not much different from the one for the massless gas. Similarly, the total pressure in hadron phase consisting of $n$ different types of particles and resonances is given by

$$ p_h = -T \sum_{i=1}^n \frac{g_i}{2\pi^2} \int_0^\infty dk k^2 \ln \left( 1 - e^{-\sqrt{k^2 + m_i^2}/T} \right), $$

$$ = \frac{\pi^2}{90} T^4 \sum_{i=1}^n g_i h(m_i/T). \quad (5) $$

Here,

$$ h(m_i/T) = -\frac{4\pi}{\sqrt{\pi}} \int_0^\infty du u^2 \ln \left( 1 - e^{-\sqrt{u^2 + (m_i/T)^2}} \right). $$

If we consider only a massless pion gas, the above summation contributes only a factor of 3. Assuming the hadron gas consisting of $\pi$, $K$, $\eta$, $\rho$ and $\omega$ with appropriate mass and degeneracy $[m_{\pi} = 137.0 \text{ MeV} \ (g=3), \ m_K = 496.0 \text{ MeV} \ (g=4), \ m_\eta = 547.45 \text{ MeV} \ (g=1), \ m_\rho = 768.50 \text{ MeV} \ (g=9), \ m_\omega = 781.94 \text{ MeV} \ (g=3)]$ we get

$$ p_h \simeq a_h T^4, \quad \text{where} \quad a_h = 4.5 \frac{\pi^2}{90} \quad \text{at} \quad T = 160\text{MeV}. \quad (6) $$

In addition to the above equation of state, we use the following phenomenological potential to describe the phase transition $28$

$$ V(\phi, T) = a(T) \phi^2 - bT \phi^3 + c \phi^4, \quad (7) $$

where $b$ and $c$ are positive constants. Although, the exact knowledge of the order parameter $\phi$ is not necessary for the present purpose, it can be related to entropy or energy density $4$ or to the sigma and pion field $28$. The above potential has two minima, one at $\phi = 0$ and the other one at

$$ \phi_m = (3bT + \sqrt{9b^2 T^2 - 32ac})/8c. \quad (8) $$

In the present case will represent quark and hadron phases respectively. These phases are separated by a maximum that occurs at $\phi_m$ given by

$$ \phi_m = (3bT - \sqrt{9b^2 T^2 - 32ac})/8c. \quad (9) $$

In the thin wall approximation $22$, $b$ and $c$ can be expressed in terms of surface tension $\sigma$ and correlation length $\xi$ as $14$

$$ b = \frac{1}{\sqrt{64\xi^2 T_C^2}}; \quad c = \frac{1}{12\xi^2 \sigma}. \quad (10) $$

The height of the degenerate barrier at $T = T_C$ or at $a(T_C) = b^2 T_C^2 / 4c$ is given by

$$ V_b(T_C) = \frac{3\sigma}{16\xi^2 T_C}. \quad (11) $$

Thus, the potential can describe first-order phase transitions, with varying strength by either changing $\sigma$ or $\xi$ or both. The latent heat is also a measure of the strength of the first order phase transition which in the bag model is given by $4B$. The requirement that at all the temperatures, the difference between the two minima should be equal to the pressure difference between the two phases $14$ fixes the third parameter $a(T)$ as
\[ \Delta p = p_h - p_q = -[V(\phi_h) - V(0)] = -[a(T) - bT\phi_h + c\phi_h^2] \phi_h^2. \]  

Scavenius and Dumitru [24] have used a linear sigma model (LSM) potential and fitted the potential of the form given by Eq. (4) to fix the parameters \( a, b \) and \( c \). With this form of potential they obtain surface tension \( \sigma \) and correlation length \( \xi \). We directly use surface tension \( (\sigma) \) and correlation length \( (\xi) \) to fix the parameters \( b \) and \( c \). For any value of \( b \) and \( c \) the nonperturbative vacuum effect is taken care of by bag constant \( B \). At \( T = T_C \), \( \phi_h = bT_C/2c = \sqrt{6} \sigma \xi \). For \( \sigma = 10 \) MeV/fm\(^2 \) and \( \xi = 0.7 \) fm (typical hadron size) we get \( \phi_h = 91 \) MeV which is nothing but the pion decay constant.

Figure 1 shows the plot of \( V(\phi) \) as a function of \( \phi \) at four different temperatures for typical values of \( \sigma = 30 \) MeV/fm\(^2 \) and \( \xi(T_C) = 0.7 \) fm. At \( T = T_C \), the potential is degenerate with \( V(\phi_q) = V(\phi_h) = 0 \) being separated by a barrier at \( \phi = \phi_m \). As the matter supercools, the hadron phase has lower free energy as compared to the QGP phase which is held fixed at the hadron phase. If the QGP supercools upto the transition and need to be evaluated properly. While \( T_S \) can be estimated directly from Eq. (13), \( T_m \) requires a self consistent solution of a set of equations involving the nucleation rate and energy momentum conserving hydrodynamical equations.

The homogeneous nucleation theory assumes the formation of nucleating clusters within the initially homogeneous metastable state. One can use the Langer formalism [3], to calculate the probability of nucleation of hadron bubbles per unit time per unit volume from a homogeneous background of quark gluon plasma, given by

\[ I = I_0 \exp \left( -\frac{\Delta F_C}{T} \right). \]

Here, \( \Delta F_C \) is the minimum energy needed to form a critical hadron bubble and \( I_0 \) is the prefactor which can be written as \( I_0 = \kappa \Omega/2\pi \). The statistical factor \( \Omega \) is a measure of both the available phase space as the system goes over the saddle and of the statistical fluctuations at the saddle relative to the equilibrium states. The dynamical prefactor \( \kappa \) gives the exponential growth rate of the bubble or droplet sitting on the saddle.

### A. Statistical prefactor

To understand the meaning of the statistical prefactor, consider a classical system with \( N \) degrees of freedom described by a set of \( N \) collective coordinates \( \eta_i, \ i = 1, \ldots, N \). The coarse-grained free energy functional \( F(\eta) \) of the system has local minima \( F(\eta_i) \) in the \( \{\eta\} \)-space, corresponding to the metastable \( \{\eta^0\} \) and stable states, separated by the energy barrier. The point of minimal energy along the barrier is the so-called saddle point \( \{\eta^S\} \). The rate of the decay of the metastable
state is determined by the steady-state current across the saddle point from the metastable to the stable minimum of \( F[\eta] \). According to \([28, 29]\), the statistical prefactor can be written as

\[
\Omega = \frac{8\pi\sigma}{3|\lambda_1|^{3/2}} \left( \frac{2\pi T}{|\lambda_1|} \right)^{1/2} \left[ \text{det}(M_0/2\pi T) \right].
\]  

(15)

where \( \lambda_1 \) is the negative eigen value of the matrix \( M_{ij} = \partial^2 F/\partial \eta_i \partial \eta_j \), evaluated at the saddle point. The subscript 0 denotes the metastable state and the prime indicates that the negative eigen value \( \lambda_1 \) as well as the zero eigen values of the matrix \( M_{ij} \) is omitted. The calculation of the fluctuation determinant in Eq. (15) is extremely difficult and adds to the uncertainty of the calculation of \( \Omega \). In Ref. \([30]\), the above prefactor has been estimated under harmonic approximation using the product of \( \eta \) energies of the metastable \((\Delta F_{1}(R))\) and the saddle point \((\Delta F_{3}(R))\) are similar if they have same similarity number. In the harmonic approximation, Eq. (18) can be expanded around the critical radius as

\[
\frac{\Delta F}{T} = \left( \frac{\Delta F}{T} \right)_{R=R_C} + \frac{1}{2T} \left( \frac{\partial^2 \Delta F}{\partial R^2} \right)_{R=R_C} (R-R_C)^2
\]

(20)

and we get finally for the negative eigen value \( \lambda_1 \)

\[
\lambda_1 = -2T\lambda_2^2.
\]

(21)

Following \([7, 8]\), we estimate \( \Omega_2 \) using Eq. (17) and Eq. (21).

B. Dynamical prefactor

Csernai and Kapusta \([4]\) have derived a dynamical prefactor \( \kappa \) by generalizing the coarse-grained effective field theory of Langer to a relativistic case. For a baryon-free plasma, where the thermal conductivity vanishes, the dynamical prefactor is found to depend on the viscosity coefficients and is given by

\[
\kappa_{CK} = \frac{4\sigma (4\eta_q/3 + \zeta_q)}{(\Delta \omega)^2 R_C^2}.
\]

(22)

Here \( \eta_q \) and \( \zeta_q \) are the shear and bulk viscosity coefficients of the quark phase respectively and the \( \Delta \omega \) is the difference between the enthalpy densities of the QGP and the hadron phases, \( \omega = \epsilon + p \). According to Csernai and Kapusta (CK), there will be no bubble growth in the case of an ideal plasma with zero viscosity. Since viscosity is an essential ingredient in the above expression, for consistency, dissipative hydrodynamics should also be used to describe the space time evolution of the matter \([\bar{9}]\). The CK approach has also been extended to include thermal dissipation in addition to viscous damping for the case of baryon rich quark gluon plasma \([22]\). On the contrary, Ruggeri and Friedman \([3]\) argued that for relativistic hydrodynamics the energy flow does not vanish in the absence of any heat conduction or viscous damping. Since the change of the energy density \( \epsilon \) in time is given in the low velocity limit by the conservation equation \([4]\), \( \partial \epsilon / \partial t = -\nabla \cdot (\omega \nu) \) which implies that the energy flow \( \propto \omega \nu \) is always present. Therefore, following a different approach, Ruggeri and Friedman derived an expression for \( \kappa \), given by (for zero viscosity)

\[
\kappa_{RF} = \left( \frac{2\pi \omega_q}{R_C^2 (\Delta \omega)^2} \right)^{1/2}.
\]

(23)

The above result is in contradiction with the expression given by Eq. (22). Recently, following the Langer’s procedure, we have solved the relativistic hydrodynamic in
the hadron, quark and the interfacial regions to obtain the dynamical prefactor. The $\kappa$ as obtained in Ref. [23], is given by

$$\kappa = \kappa_0 + \kappa_v; \kappa_0 = c_s \sqrt{\frac{\xi}{3 R_C}}, \kappa_v = \frac{\xi}{6 R_C^2} \sqrt{\omega_q} \left( \frac{4 \eta_q}{3} + \zeta_q \right).$$

(24)

Here $c_s$ is the velocity of sound in the massless gas. Under certain assumption for $c_s$, the above prefactor can also be written as [23]

$$\kappa \approx \left( \frac{2 \sigma \omega_q}{R_C^3 (\Delta \omega)^2} \right)^{1/2} + \frac{1}{c_s^2} \left( \frac{4 \eta_q/3 + \zeta_q}{(\Delta \omega)^2 R_C^3} \right).$$

(25)

As can be seen, the first term in the above equation is the same as obtained by Ruggeri and Friedman corresponding to the case of nonviscous plasma [see Eq. (23)]. The second term is similar to the result obtained by Csernai and Kapusta. Two important aspects of our result are (a) the prefactor $\kappa$ can be written as a linear sum of a nonviscous ($\kappa_0$) and a viscous ($\kappa_v$) components and (b) the nonviscous component ($\kappa_0$) which depends on two parameters $R$ and $\xi$ is finite in the limit of zero viscosity. Further, it has been argued in Ref. [23] that the nonviscous part of the prefactor basically arises due to nonuniform pressure across the interface. In the present work, we will use Eq. (24) to calculate the dynamical prefactor.

IV. NUCLEATION AND THE EXPANSION DYNAMICS

To calculate the fraction of space converted to hadron phase using the nucleation rate $I(T)$, one requires a kinetic equation. The simplest of such equation is given in Ref. [6]. According to it, if the QGP cools to $T_C$ at a proper time $\tau_C$, then at some later time $\tau$ the fraction of space which has been converted to hadronic gas is given by

$$h(\tau) = \int_{\tau_C} d\tau' I(T(\tau'))[1 - h(\tau')]V(\tau', \tau).$$

(26)

Here $V(\tau', \tau)$ is the volume of a bubble at time $\tau$ which had been nucleated at an earlier time $\tau'$; this takes into account the bubble growth. The factor $[1 - h(\tau')]$ is the available space for new bubbles to nucleate. The model for bubble growth can be taken as [8, 23]

$$V(\tau', \tau) = \frac{4 \pi}{3} \left( R(T(\tau')) + \int_{\tau'}^{\tau} d\tau'' v(T(\tau'')) \right)^3,$$

(27)

where $v(T)$ is the velocity of the bubble growth at temperature $T$ and is taken to be $v(T) = 3[1 - T/T_C]^{3/2}$ [6, 34]. This expression is intended to apply only when $T > \frac{2}{3}T_C$ so that the growth velocity stays below the speed of sound of a massless gas, $c/\sqrt{3}$. At the critical temperature, both the nucleation rate and growth rate vanish. One can also write $v$ in terms of pressure difference between the two phases as used in Ref. [23].

The evolution of the energy density is given by [5, 36, 37]

$$\frac{d\omega}{d\tau} + \frac{D \omega}{\tau} = \frac{4 \eta q/3 + \zeta q}{\tau^2} \equiv \frac{\mu}{\tau^2},$$

(28)

where $D = 1$ for the expansion in $(1 + 1)$ dimension. The factors $\eta$ and $\zeta$ are the shear and the bulk viscosity of the medium. For nonviscous plasma (for zero viscosity), the above equation follows the Bjorken’s scaling solution where $T^3 \tau$ is a constant. The energy momentum equation needs to be solved numerically for expansion in $(3 + 1)$ dimensions. However, retaining the simplicity, we can still use Eq. (28) for spherical expansion [8, 34] with the choice of $D = 3$, although the scaling $u^\mu = x^\mu/\tau$ (where $u^\mu$ is the four velocity of the fluid) may not be valid if viscosity is included.

In the transition region, the energy density at a time $\tau$ can be written in terms of hadronic fraction $h(\tau)$ as

$$\epsilon(\tau) = \epsilon_q(T) + [\epsilon_h(T) - \epsilon_q(T)] h(\tau).$$

(29)

Here $\epsilon_q = 3p_q + 4B$ and $\epsilon_h = 3p_h$ are energy densities in the QGP and hadron phase respectively. The viscosity at any time $\tau$ can also be written [10] as

$$\mu(\tau) = \mu_q(T) + [\mu_h(T) - \mu_q(T)] h(\tau),$$

(30)

where $\mu_q = 4\eta q/3 + \zeta q$ and $\mu_h = 4\eta h/3 + \zeta h$. The temperature dependence of the viscosities are taken as $\sim T^3$ [30].

V. RESULTS AND DISCUSSIONS

We compare the nucleation rates using different dynamical prefactors $\kappa$ as given by

$$I_1 = \frac{\Omega_2 \kappa C F}{2\pi} \exp \left( -\frac{\Delta F_C}{T} \right),$$

$$I_2 = \frac{\Omega_2 \kappa_0}{2\pi} \exp \left( -\frac{\Delta F_C}{T} \right),$$

$$I_3 = \frac{\Omega_2 \kappa_v}{2\pi} \exp \left( -\frac{\Delta F_C}{T} \right),$$

$$I_4 = \frac{\Omega_2 (\kappa_0 + \kappa_v)}{2\pi} \exp \left( -\frac{\Delta F_C}{T} \right),$$

$$I_5 = T^4 \exp \left( -\frac{\Delta F_C}{T} \right).$$

(31)

In all the cases, we use the same statistical prefactor $\Omega_2$ which is more physical except for $I_5$ where the total prefactor is of the order of $\sim T^4$ as commonly used.
in the literature \[24\]. Further, we choose \( \xi = 0.7 \) fm and use Eq. (14) to estimate \( \Delta F_C/T \). The nucleation rates \( I_2 \) and \( I_5 \) correspond to the case of a nondissipative plasma whereas viscosity enters as an essential ingredient for the evaluation of \( I_1, I_3 \) and \( I_4 \). Figures 2 and 3 show the plot of above nucleation rates as a function of \( T/T_C \) for \( \sigma = 10 \) MeV/fm\(^2\) and \( \sigma = 30 \) MeV/fm\(^2\) respectively each at two typical values of viscosity coefficient \( \eta_q \). We consider only the shear viscosity \( \eta_q \), since the bulk viscosity \( \zeta_q \) that provides resistance to expansion (or contraction) does not exist in the case of an incompressible fluid. Further, the viscosity coefficients \( \eta_q \) and \( \zeta_q \) in hadron phase are neglected since the amount of supercooling is affected by viscosity of quark phase only. Notice that the most commonly used prefactor \( \sim T^4 \) overestimates the nucleation rate \( (I_5) \) over a wide range of temperature particularly for small amount of supercooling and also when the transition is relatively stronger (see Fig. 3). Further notice, that both for strong and weak transitions but for moderately viscous medium, the nucleation rates are dominated by the nonviscous components (since \( I_4 \) is nearly the same as \( I_2 \), see Figs 2(a) and 3(a) for \( \eta_q = 5T^3 \)). The rates \( I_1 \) and \( I_3 \) which contain viscosity have only a second order effect unless the medium is highly dissipative. At \( \eta_q = 14.4T^3 \), the viscous component \( I_3 \) competes with the nonviscous component \( I_2 \) although \( I_1 \) is still lower. [see Figs. 2(b) and 3(b)]. From the above studies we may conclude that the choice of \( T^4 \) overestimates whereas the rate \( (I_1) \) with the Csernai-Kapusta prefactor underestimates the nucleation rates in the region of interest as compared to \( I_4 \). Since the nucleation rate \( I_4 \) has both dissipative and nondissipative components, we will use it to estimate the nucleation rate and supercooling in the subsequent study.

Next, we solve the coupled equations Eq. (28) and Eq. (23) to study the space time evolution of the matter. The critical temperature is fixed at \( T_C = 160 \) MeV which gives \( B^{1/4} = 236 \) MeV. The strength of the transition depends on \( \sigma \) and \( \xi \) or more explicitly on the ratio \( \sigma/\xi \). Recent lattice calculations [25] predict \( \sigma \) between 2 to 10 MeV/fm\(^2\). These calculations are without any dynamical quarks and thus are only indicative. Therefore \( \sigma \) is treated as a free parameter in the present study and \( \xi = 0.7 \) fm seems the most reasonable value for the correlation length.

Another important parameter that affects the evolution is the time \( \tau_C \) that the plasma takes to cool down to \( T = T_C \). Obviously, \( \tau_C \) will depend on the initial temperature \( (T_i) \), the formation time \( (\tau_1) \) as well as on the expansion dynamics of the plasma. In the Bjorken scenario [13], \( \tau_C = (T_i/T_C)^3\tau_1 \). It is quite reasonable to assume that the plasma will expand in \((1+1)\) dimension between \( \tau_1 \) and \( \tau_C \) until the longitudinal dimension becomes comparable to the size of the colliding nuclei. Assuming \( T_i = 320 \) MeV and \( \tau_1 = 1 \) fm/c, the Bjorken scaling predicts \( \tau_C = 8 \) fm/c for expansion in \((1+1)\) dimension. As mentioned before, \( \tau_C \) is a crucial parameter that affect the solution of the coupled equations which in turn depends on the initial conditions as well as on the expansion scenario of the plasma. Since \( \tau_C \) is not known exactly, we assume it of the order of 6 to 8 fm/c. From \( \tau_C \) onwards, we consider the spherical expansion. This expansion scenario corresponds to the fastest cooling rate resulting in maximum supercooling. Therefore the choice of \( D = 3 \) provides a lower bound on \( T_m \) which we compare with the spinodal temperature \( T_S \).

Figure 4 shows the plot of \( T/T_C \) as a function of \( \tau \) for a nonviscous plasma at a typical value of \( \tau_C = 8 \) fm/c. Unlike the ideal Maxwell construction (where the temperature is held fixed at \( T = T_C \) until phase conversion
is over), the system supercools up to $T_m$. At $T_m$, the number density of the nucleated hadron bubbles is sufficient to raise the temperature again due to the release of the latent heat. The amount of supercooling will obviously depend on the strength of the transition being more for a stronger transition (large value of $\sigma$) as compared to that for the weaker one (small value of $\sigma$). A faster expansion also reduces the number density of the nucleated hadron bubbles resulting in a larger supercooling. Although shown for $\tau_C = 8 \text{ fm/c}$, the amount of supercooling will increase further if $\tau_C$ is reduced due to faster expansion. We have also done the calculations for $\tau_C = 6 \text{ fm/c}$ which are summarized in Fig. 6. For $\tau_C$ still smaller the expansion in (1+1) should be used.

Next, we consider the dissipative hydrodynamics. As mentioned before, the viscous contribution to the nucleation rate is not very significant if the transition is strong enough and also the medium is moderately viscous. However, the presence of a small amount of viscosity will affect the hydrodynamic evolution of the plasma. The temperature of the plasma will fall at a slower rate due to viscous heating of the medium. Figure 5 shows a plot of $T/T_C$ as a function of $\tau$ for $\eta_q$=0, 5$T^3$ and 14.4$T^3$. Here we have considered the case of a relatively stronger transition ($\sigma = 30 \text{ MeV/fm}^2$) where the viscosity contribution to the nucleation rate is not very significant. However, significant effect can be seen on $T_m$ which increases (less supercooling) with increasing viscosity as expected [3].

Figures 6 and 7 show the plot of $T_m/T_C$ as a function of $\sigma/\xi$ at $\tau_C= 6 \text{ fm/c}$ and $8 \text{ fm/c}$ respectively both for ideal as well as for dissipative plasma. The amount of supercooling becomes less as the medium becomes more viscous. The spinodal temperature $T_S$ which depends on the ratio $\sigma/\xi$ has also been plotted in the same figures for comparison. First consider the case of nondissipative plasma with zero viscosity. The curves $T_m$ and $T_S$ show a cross over point depending on the choice of $\tau_C$. The nucleation rate is suppressed for a strong enough transition due to large nucleation barrier resulting in a higher amount of supercooling. On the other hand for weak enough transition, the amount of supercooling is smaller and also well above the point of spinodal decomposition. Since $T_m$ depends on expansion rate of the medium, the cross over point will sensitively depend on $\tau_C$; moving towards left for the faster expansion. Qualitatively it can be concluded here that the homogeneous nucleation is still the dominant mechanism of phase conversion if the transition is weak enough. For stronger transition and fast enough cooling, the system may reach $T_S$ and the phase conversion may proceed through the spinodal decomposition. If the plasma is viscous, the amount of supercooling is reduced further due to slow evolution of the medium. Even the cross over point also shifts towards right showing that the nucleation is the dominant mechanism over a wide range of $\sigma/\xi$ ratios. Due to uncertainties in $\tau_C$, $\sigma/\xi$ and $\eta_q$, it is difficult to say what is the exact scenario at RHIC and LHC energies. Assuming a weakly first order transition for $\sigma$ in the range of 2 to 5 MeV/fm$^2$, $\xi \sim 0.7 \text{ fm}^2$, $\eta_q \sim 2T^3$ (lower limit [1]) and $\tau_C \sim 4$ to 8 $\text{ fm/c}$ [18], the nucleation still seems to be the dominant mechanism of phase conversion as opposed to the spinodal decomposition.
It may be mentioned here that we have considered a dynamical growth rate which is exponential in nature. This may not be true at the later stage of the expansion. The fusion of bubbles may also play a role at the later stage when the density of bubbles is sufficiently high. However, in the study of supercooling, we are very much in the domain of exponential region where the effect of bubble fusion can be ignored.

VI. CONCLUSIONS

We have investigated the mechanism of phase conversion from quark gluon plasma phase to hadron phase via two routes: the standard homogeneous nucleation and spinodal decomposition. The point of spinodal decomposition depends on the strength of the transition or more precisely on the ratio $\sigma/\xi$ where $\sigma$ and $\xi$ are the surface tension and correlation length respectively at $T = T_C$. The nucleation and supercooling, on the other hand, depend on the strength of the transition as well as on the expansion dynamics of the medium. We have solved a set of coupled equations to estimate the amount of maximum supercooling under spherical expansion scenario. Which way the hadronization will proceed depends sensitively on the nucleation and expansion dynamics since the nucleation and expansion time scales are comparable. Qualitatively we can describe the results as follows: For strong enough transition with zero or very small amount of viscosity, the system reaches the spinodal instability before the amount of nucleated hadron bubbles become significant to begin phase conversion. The phase conversion in such a case will proceed via spinodal decomposition. If the medium is viscous or the transition is weak enough or both, the supercooling is much less; The phase conversion may still proceed through homogeneous nucleation. However, depending on the range of the parameters, there could be a competition between the homogeneous nucleation and the spinodal decomposition. A definite answer can be provided only when the parameters such as surface tension, $\tau_C$, viscosity and also the expansion scenario are known precisely.

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