Derivation of the cut-off length from the quantum quadratic enhancement of a mass in vacuum energy constant Lambda

Kimichika Fukushima\textsuperscript{1,3,a}, Hikaru Sato\textsuperscript{2,b}

\textsuperscript{1} Advanced Reactor System Engineering Department, Toshiba Nuclear Engineering Services Corporation, 8, Shinsugita-cho, Isogo-ku, Yokohama 235-8523, Japan
\textsuperscript{2} Emeritus, Department of Physics, Hyogo University of Education, Yashiro-cho, Kato-shi, Hyogo 673-1494, Japan
\textsuperscript{3} Theoretical Division, South Konandai Science Research, 9-32-2-701, Konandai, Konan-ku, Yokohama 234-0054, Japan

Received: 21 November 2017 / Accepted: 6 April 2018 / Published online: 19 April 2018
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Abstract Ultraviolet self-interaction energies in field theory sometimes contain meaningful physical quantities. The self-energies in such as classical electrodynamics are usually subtracted from the rest mass. For the consistent treatment of energies as sources of curvature in the Einstein field equations, this study includes these subtracted self-energies into vacuum energy expressed by the constant Lambda (used in such as Lambda-CDM). In this study, the self-energies in electrodynamics and macroscopic classical Einstein field equations are examined, using the formalisms with the ultraviolet cut-off scheme. One of the cut-off formalisms is the field theory in terms of the step-function-type basis functions, developed by the present authors. The other is a continuum theory of a fundamental particle with the same cut-off length. Based on the effectiveness of the continuum theory with the cut-off length shown in the examination, the dominant self-energy is the quadratic term of the Higgs field at a quantum level (classical self-energies are reduced to logarithmic forms by quantum corrections). The cut-off length is then determined to reproduce today’s tiny value of Lambda for vacuum energy. Additionally, a field with nonperiodic vanishing boundary conditions is treated, showing that the field has no zero-point energy.

1 Introduction

Self-interaction energies in field theory, which contain ultraviolet divergences in continuum theory, sometimes reveal meaningful properties in physics \cite{1–8}. In our previous paper \cite{9–12}, we formulated a field theory in terms of the step-function-type basis functions (SFT field theory), which is based on the finite element theory \cite{9,12–14} (the formulation is rather different from that by Bender et al.), and cuts off high-frequency oscillations of wave functions at short distances. Owing to the space-time continuum and differentiable step-function-type basis functions, this formalism is Poincaré covariant and removes ultraviolet divergences at short distances. The advantage of our formalism is the availability to perform self-energy evaluation. (We note that the conventional finite element method is widely used \cite{13}. The validity of theories is of course justified solely by the correctness of the logical deduction. The support based only on the fact, where an article was published, is insufficient for the true justification of the theory. The assessment of the theory is beyond the range of the work by the authors.) The meaningful self-energy appears in the Lamb shift \cite{1}, which is caused by finite parts of the self-energy in higher-order terms, and the divergent parts are subtracted from the rest mass. In contrast, the self-energy also appears in the $\phi^3$ model (the mass is sometimes not renormalized when the mass is a value in vacuum without containing additional interactions). In our previous paper \cite{12}, we derived excited states such as meta-stable states at stationary states, which are not always orthogonal to the ground state.

In the Einstein field equations \cite{15,16}, the rest energy works as a source of the curvature. The mass renormalization in such as electrodynamics subtracts self-energies, which can be finite using the cut-off scheme. It is then expected that the self-energies are involved in the Einstein field equations.

In our formalism, four-dimensional space-time is divided into many hyper-octahedrons, whose shape are arbitrary and have the size $\Delta$ (cut-off length) in four-dimensional space-time. For simplicity, we consider three-dimensional space and divide the region into many cubes. The classical wave function $\phi(x, y, z)$ is expressed in terms of the step-function-
type basis functions $\hat{\Omega}^3_p(x, y, z)$ in three-dimensional space [the step-function-type basis function in one dimension is defined by Eq. (19)]

$$\phi(x, y, z) = \sum_p \phi_p \hat{\Omega}^3_p(x, y, z),$$

where the basis function takes a value of 1 in a cube (each cubic region is identified by index $p$) and vanishes outside the cube. The coefficient $\phi_p$ is a constant within the cubic region identified by the index $p$.

Motivated by the above expectation, this paper is aimed at presenting a formulation to include the subtracted self-energies into vacuum energy with the constant $\Lambda$ (cosmological constant) [16–32] of the macroscopic classical Einstein field equations. The self-energy in classical electrodynamics is calculated by the continuum theory with a finite cut-off length. The self-energy is also derived using the field theory in terms of the step-function-type basis functions, which was developed by the present authors, and the result is compared with that calculated by the continuum theory. We also examine the curvature (gravitational) self-energy of the fundamental particle with the energy of a rest mass. Considering the examinations that the self-energies derived in terms of the step-function-type basis functions and that by the continuum theory with the same cut-off length are not so different, the classical self-energies are reduced to the logarithmic forms. However, the self-energy of the Higgs boson has the larger quadratic form. The derived self-energy is subtracted and involved in the repulsive vacuum energy with the constant $\Lambda$. Under a classical gravitational field, whose strength is small for scales larger than the Planck scale, we consider the contribution from the self-energy of a Higgs boson to vacuum energy. The cut-off length is then determined to reproduce the observed vacuum energy constant $\Lambda$. This theoretical vacuum energy constant $\Lambda$ has today’s tiny value.

This paper is organized as follows: Section 2 presents the formalism and analysis procedure. We exhibit a formalism for the subtraction of the self-energy by involving the energy into vacuum energy constant $\Lambda$. Subsequently, the field theory in terms of the step-function-type basis functions is described to derive finite subtracted self-energies produced by interactions (in such as electrodynamics) into vacuum energy constant $\Lambda$. Throughout this paper, the notation $x^0 = ct$ ($c$ is the velocity of light) is the time coordinate, and the $x^i$ are space coordinates, where $x^1 = x, x^2 = y$ and $x^3 = z$. The infinitesimal squared distance (according to the notation by Bjorken and Drell [33]) is denoted

$$(ds)^2 = g_{\mu\nu}dx^\mu dx^\nu,$$  

where $g_{\mu\nu}$ is the metric tensor and the indices run over 0, 1, 2 and 3. We use the summation conventions such as

$$g_{\mu
u} dx^\nu = g_{\mu 0} dx^0 + g_{\mu 1} dx^1 + g_{\mu 2} dx^2 + g_{\mu 3} dx^3,$$  

(3)

for Greek indices and

$$g_{\mu i} dx^i = g_{\mu 1} dx^1 + g_{\mu 2} dx^2 + g_{\mu 3} dx^3,$$  

(4)

for Latin indices. The metric tensor of $g_{\mu\nu}$ in a flat Minkowski space is given by

$$g_{\mu\nu} = \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{bmatrix}.$$  

(5)

The action functional $S_g$ for the gravity is expressed by

$$S_g = \int L_g \sqrt{-g} dV_4,$$  

(6)

where $g$ is the determinant of $g_{\mu\nu}$, $g = \det(g_{\mu\nu})$, and $dV_4 = dx^0 dx^1 dx^2 dx^3$. For the gravity,

$$L_g = -\frac{c^2}{16\pi G} R,$$  

(7)

where $G$ is the gravitational constant and $R$ is the scalar curvature written by

$$R = g^{\mu\nu} R_{\mu\nu},$$  

(8)

with $R_{\mu\nu}$ being defined by

$$R_{\mu\nu} = R_{\mu\nu}^\rho,$$  

(9)

using the Riemann curvature tensor $R_{\mu\nu\rho}^\sigma$. The tensor $R_{\mu\nu}^\rho$ is expressed in terms of the Christoffel symbol $\Gamma_{\mu\nu}^\lambda$ as

$$R_{\mu\nu\rho}^\sigma = \frac{\partial \Gamma_{\mu\nu}^\sigma}{\partial x^\rho} - \frac{\partial \Gamma_{\mu\rho}^\sigma}{\partial x^\nu} + \Gamma_{\mu\lambda}^\rho \Gamma_{\nu\lambda}^\sigma - \Gamma_{\nu\lambda}^\rho \Gamma_{\mu\lambda}^\sigma,$$  

(10)
where
\[ R_{\mu\nu} = \frac{1}{2} g^{\lambda\rho} \left( \frac{\partial g_{\rho\mu}}{\partial x^\lambda} + \frac{\partial g_{\rho\nu}}{\partial x^\lambda} + \frac{\partial g_{\mu\lambda}}{\partial x^\rho} \right). \] (11)

Meanwhile, the action functional of the matter is denoted
\[ S_m = \frac{1}{c} \int \mathcal{L}_m \sqrt{-g} dV_4, \] (12)
where \( \mathcal{L}_m \) is the Lagrangian density of the matter, and the energy-momentum tensor of the matter is obtained from the relation
\[ \frac{1}{2} \sqrt{-g} T_{\mu\nu} = -\left( \frac{\partial g_{\mu\lambda}}{\partial x^\nu} \frac{\partial g_{\rho\nu}}{\partial x^\rho} \mathcal{L}_m - \frac{\partial g_{\mu\nu}}{\partial x^\rho} \frac{\partial g_{\rho\nu}}{\partial x^\rho} \mathcal{L}_m \right). \] (13)

The variational calculus with respect to \( \delta g^{\mu\nu} \) of the total action functional,
\[ \delta S_g + \delta S_m = -\frac{c^3}{16\pi G} \int \left( R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R - \frac{8\pi G}{c^4} T_{\mu\nu} \right) \delta g^{\mu\nu} \sqrt{-g} dV_4, \] (14)
yields the Einstein field equations,
\[ R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = \frac{8\pi G}{c^4} T_{\mu\nu}. \] (15)

The renormalization of the mass by interactions in such as electrodynamics subtrahs self-energies from the rest mass. Because the energy of the rest mass produces curvature (gravity), the subtracted energies are included in vacuum energy with the constant \( \Lambda \). In the above equation, we then add the following tensor for the removal of self-energies produced by interactions (such as in electrodynamics):
\[ T^{(S)}_{\mu\nu} = \frac{c^4}{8\pi G} g_{\mu\nu} \Lambda^{(S)}, \] (16)
where \( \Lambda^{(S)} \) is regarded as the vacuum energy constant \( \Lambda \) (cosmological constant). The Einstein field equations given by Eq. (15) are rewritten as follows:
\[ R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = \frac{8\pi G}{c^4} \left( T_{\mu\nu} - T^{(S)}_{\mu\nu} \right). \] (17)
We then have
\[ R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R + g_{\mu\nu} \Lambda^{(S)} = \frac{8\pi G}{c^4} T_{\mu\nu}, \] (18)
which corresponds to the Einstein field equations with vacuum energy constant \( \Lambda \). Consequently, subtracted self-energies in the interactions are involved in a vacuum energy with constant \( \Lambda \).

2.2 Field theory in terms of the step-function-type basis functions

In describing physical quantities at short distances, theories are required to remove ultraviolet divergences. We formulated the field theory [9–12], which is expressed in terms of the step-function-type basis functions to realize the removal of the ultraviolet divergences. In this subsection, the formalism is described so as to express the fields in terms of the step-function-type basis functions in the form used by this paper. Our described formalism divides four-dimensional real space-time into hyper-octahedrons with arbitrary shapes of the boundaries. The hyper-octahedron in real space-time is mapped from a hypercube with flat boundary surfaces in a parameter space-time. A basis function defined around a center of a hypercube takes a value of unity and vanishes outside the hypercube.

In this paper, the cubic region in three-dimensional space is approximated by the spherical region for simplicity and convenience. We calculate fields in spherical coordinates and divide the spherical symmetric region into shells. The results can be generalized to the case in which the region is divided into many hyper-octahedrons with arbitrary shapes. Grid (lattice) points along the radial \( r \)-axis \((r = (x^2 + y^2 + z^2)^{1/2})\) are denoted \( r_1, r_2, \ldots, r_k, \ldots, r_{N_r+1} \), with \( k = 1, 2, c, \ldots, N_r+1 \), where \( N_r \) is the number of lattice points and \( k = N_r + 1 \) is the lattice index for a boundary. We here set the radial cut-off length \( \Delta_h \) (corresponding to the cut-off length \( \Delta \) with \( \Delta = 2\Delta_h \)) to the lattice spacing by \( \Delta_h = r_k - r_{k-1} \) and define the notations \( r_k-1/2 = r_k - \Delta_h/2 \) and \( r_k+1/2 = r_k + \Delta_h/2 \). The step-function-type basis function used is defined by
\[ \tilde{\Omega}_k^E(r) = \begin{cases} 1 & \text{for } r_{k-1/2} \leq r < r_k+1/2, \\ 0 & \text{for } r < r_{k-1/2} \text{ or } r \geq r_k+1/2, \end{cases} \] (19)
which has the properties
\[ \frac{d\tilde{\Omega}_k^E(r)}{dr}_{|r=r_{k-1/2}} = \delta(r - r_{k-1/2}), \] (20)
\[ \frac{d\tilde{\Omega}_k^E(r)}{dr}_{|r=r_{k+1/2}} = -\delta(r - r_{k+1/2}), \] (21)
where \( \delta(r) \) is the Dirac delta function.

The field \( \phi(r) \) in spherical coordinates is transformed to
\[ u(r) = r \phi(r), \] (22)
and this wave function \( u(r) \) is expressed in terms of basis functions defined by Eq. (19):
\[ u(r) = \sum_k u_k \tilde{\Omega}_k^E(r). \] (23)

Thus, we have prepared the formalism to the analysis of the self-energies one finds in the next section.
3 Analysis of self-energies from interactions by classical fields

3.1 Self-energy and mass renormalization in classical electrodynamics by the continuum theory

This subsection examines and summarizes the self-energy in classical electrodynamics using the continuum theory [15,34–37]. The mass density \( \mu_m \) of a fundamental particle with a mass \( m_E \) and size \( R_E \) is denoted

\[
\mu_m = \frac{m_E}{(4\pi/3)(R_E)^3}.
\]  (24)

We divide three-dimensional (3D) space into identical cubic elements, which were considered in Sect. 2.2. The cubic region is approximated by a spherical region with radius \( R_0 \). The charge \( Q \) and mass \( M \) of the spherical region occupied by the fundamental particle are expressed by

\[
M = (4\pi/3)(R_0^3)\mu_m, \quad Q = (4\pi/3)(R_0^3)\rho^{(e)},
\]  (25)

respectively, where \( \rho^{(e)} \) is the charge density. A radial cut-off length \( \Delta_h = \Delta/2 = R_0 \) in spherical coordinates, corresponding to the cut-off length \( \Delta \), is introduced for simplicity and convenience. From the conventional energy-momentum tensor of electrodynamics, the self-energy of the static electric field has the form

\[
E_C^{(e)} = \int dV_3 \left( \frac{1}{2} \right) |\mathbf{E}|^2,
\]  (26)

where \( dV_3 = dx dy dz \). The classical electric field \( \mathbf{E} \) is produced as \( \text{div}(\mathbf{E}) = \rho^{(e)} \) from the electric charge density \( \rho^{(e)} \) and is written by \( \mathbf{E} = -\nabla \phi^{(e)} \), where \( \phi^{(e)} \) is the electric potential and satisfies

\[
-\nabla^2 \phi^{(e)} = \rho^{(e)}.
\]  (27)

The above self-energy,

\[
E_C^{(e)} = -\int dV_3 \left( \frac{1}{2} \right) \mathbf{E} \cdot \nabla \phi^{(e)},
\]  (28)

becomes, using Gauss’ theorem and integration by parts,

\[
E_C^{(e)} = \int dV_3 \left( \frac{1}{2} \right) \rho^{(e)} \phi^{(e)}.
\]  (29)

We consider the case, in which the charges exist in the region \( r \leq R_0 \) and \( \rho^{(e)} = 0 \) for \( r > R_0 \) using \( Q \) in Eq. (25). Gauss’ theorem for Eq. (27) then gives the following potential:

\[
\phi^{(e)}(r) = \frac{Q}{4\pi r} \quad \text{for } r > R_0.
\]  (30)

Similarly, for \( r \leq R_0 \), we have

\[
4\pi r^2 \left( -\frac{\phi^{(e)}(r)}{dr} \right) = \frac{4\pi r^3}{3} \rho^{(e)},
\]  (31)

yielding

\[
-\frac{\phi^{(e)}(r)}{dr} = \frac{1}{3} r \rho^{(e)}
\]  (32)

followed by

\[
\phi^{(e)}(r) = \int dr' \left( -\frac{\phi^{(e)}(r')}{dr'} \right) = \frac{1}{6} r^2 \rho^{(e)} \quad \text{for } r \leq R_0.
\]  (33)

To connect \( \phi^{(e)}(r) \) for \( r \leq R_0 \) in Eq. (33) continuously with that in Eq. (30) for \( r > R_0 \) at \( r = R_0 \), we shift \( \phi^{(e)}(r) \) in Eq. (33) to

\[
\phi^{(e)}(r) = \frac{1}{6} r^2 \rho^{(e)} - \frac{1}{6} R_0^2 \rho^{(e)} + \frac{Q}{4\pi R_0} \quad \text{for } r \leq R_0.
\]  (34)

Using Eqs. (29), (34) and \( Q \) in Eq. (25), we obtain the self-energy by the classical electric interaction in the continuum theory:

\[
E_C^{(e)} = \frac{1}{2} \int_0^{R_0} dr (4\pi r^2 \rho^{(e)} \phi^{(e)}(r))
\]  

\[
= \frac{1}{2} \left( \frac{4\pi}{6} \rho^{(e)} R_0^5 - \frac{1}{6} R_0^2 Q \rho^{(e)} + \frac{Q Q}{4\pi R_0} \right)
\]  

\[
= \frac{1}{2} \left( \frac{3}{10} \frac{Q Q}{4\pi R_0} - \frac{1}{2} \frac{Q Q}{4\pi R_0} + \frac{Q Q}{4\pi R_0} \right)
\]  

\[
= \frac{1}{2} \left( \frac{4}{5} \frac{Q Q}{4\pi R_0} \right).
\]  (35)

Under an external force \( f_e \), the classical Newtonian equation of motion for the above charged object, with a small velocity \( v_C \) compared to the speed of light \( c \), is expressed by (small magnetic contributions are dropped) [34–37]

\[
M \frac{d^2 v_C}{dt^2} = f_e + \int dV_3 (\rho^{(e)} \mathbf{E}),
\]  (36)

where \( M \) is the mass of the charged object in Eq. (24). Using the self-energy in Eq. (29), the lower-order terms expanded with respect to \( 1/c \) amounts to

\[
M \frac{d^2 v_C}{dt^2} = f_e - \frac{4}{3e^2} E_C^{(e)} \frac{dv_C}{dt},
\]  (37)

which results in

\[
\left( M + \frac{4}{3e^2} E_C^{(e)} \right) \frac{dv_C}{dt} = f_e.
\]  (38)

(The relativistic version was given by Dirac and Rohrlich, where the factor 1 appears corresponding to the above factor 4/3 [34–37].) Due to the requirement from the continuum relativistic theory, the fundamental particle is considered to be pointlike. Then the above self-energy diverges, which is why mass renormalization is required in electrodynamics. In mass renormalization, the self-energy is subtracted from the term with the mass.
3.2 Self-energy derivation for classical electrodynamics using the field theory in terms of the step-function-type basis functions

In contrast to the analysis of Sect. 3.1, this subsection studies the self-energy of the same object in Sect. 3.1 in classical electrodynamic equations, using the step-function-type basis functions. As mentioned in Sect. 2.2 and by Eqs. (24)–(25), we divide three-dimensional space into cubic elements with the cut-off length $\Delta_h$, and each cubic region is approximated by a sphere. The action functional for the electric field $\phi^{(e)}(r)$ can be written in the form [considering the form $-\nabla^2\phi^{(e)} - \rho^{(e)} = 0$ on the left in Eq. (27)]

$$S^{(e)}_t = -\frac{1}{2} \int dr \left( -\frac{\partial^2}{\partial x^2} \phi^{(e)}(x,y,z) + \frac{\partial^2}{\partial y^2} \phi^{(e)}(x,y,z) + \frac{\partial^2}{\partial z^2} \phi^{(e)}(x,y,z) \right)$$

$$= -\frac{1}{2} \int dr (4\pi r^2) \phi^{(e)}(r) \left( \frac{d^2}{dr^2} + \frac{2}{r} \frac{d}{dr} \right) \phi^{(e)}(r).$$

(39)

Using the transformed potential

$$u^{(e)}(r) = r \phi^{(e)}(r),$$

as Eq. (22) for spherical coordinates and integration by parts, we find

$$S^{(e)}_t = -\frac{1}{2} \int dr (4\pi r^2) \frac{u^{(e)}(r)}{r} \left( \frac{d^2}{dr^2} u^{(e)}(r) \right)$$

$$= -\frac{1}{2} \int dr (4\pi) \frac{du^{(e)}(r)}{dr} \phi^{(e)}(r)$$

$$= \frac{1}{2} \int dr (4\pi) \frac{du^{(e)}(r)}{dr} \phi^{(e)}(r).$$

(41)

As in Sect. 2.2 and by Eqs. (24) and (25), we divide three-dimensional space, containing the above sphere with radius $R_0$ centered at the origin in spherical coordinates, into shells (the number of cells enclosing the central sphere is $N_c - 1$). The radial width (lattice spacing implying the radial cut-off length) of each shell is $\Delta_h$, which is equal to the radius $R_0$ of the enclosed central sphere. As Eq. (23), the above wave function $u^{(e)}(r)$ is expressed by

$$u^{(e)}(r) = \sum_k u^{(e)}_k \tilde{\Omega}^E_k(r),$$

(42)

in terms of the step-function-type basis functions $\tilde{\Omega}^E_k(r)$ in Eq. (19). From Eq. (41), it follows that ($k, K = 1, 2, \ldots, N_c - 1$)

$$S^{(e)}_t = \sum_{k,K} \frac{1}{2} \int dr (4\pi) u^{(e)}_k \tilde{\Omega}^E_k(r) \frac{\partial \tilde{\Omega}^E_k(r)}{dr} \phi^{(e)}(r).$$

(43)

Using Eqs. (20) and (21), $S^{(e)}_t$ above is decomposed into

$$S^{(e)}_t = S^{(e)}_{t-} + S^{(e)}_{t+} + S^{(e)}_{t-+} + S^{(e)}_{t++},$$

(44)

where

$$S^{(e)}_{t-} = 4\pi \frac{1}{2} \sum_{k,K} u^{(e)}_k \tilde{\Omega}^E_k(r) \int d\delta(r-r_{k-1/2}) \delta(r-r_{K-1/2}),$$

(45)

$$S^{(e)}_{t+} = 4\pi \frac{1}{2} \sum_{k,K} u^{(e)}_k \tilde{\Omega}^E_k(r) \int d\delta(r-r_{k-1/2}) \delta(r-r_{K+1/2}),$$

(46)

$$S^{(e)}_{t-+} = 4\pi \frac{1}{2} \sum_{k,K} u^{(e)}_k \tilde{\Omega}^E_k(r) \int d\delta(r-r_{k+1/2}) \delta(r-r_{K-1/2}),$$

(47)

$$S^{(e)}_{t++} = 4\pi \frac{1}{2} \sum_{k,K} u^{(e)}_k \tilde{\Omega}^E_k(r) \int d\delta(r-r_{k+1/2}) \delta(r-r_{K+1/2}).$$

(48)

With the help of the lattice spacing $\Delta_h$ mentioned above Eq. (19), an element such as $S^{(e)}_{t-}$ in Eq. (45) is reduced to

$$S^{(e)}_{t-} = 4\pi \frac{1}{2} \sum_{k,K} u^{(e)}_k \tilde{\Omega}^E_k(r) \delta[(r-r_{k-1/2}) - (r-r_{K-1/2})].$$

(49)

where $\delta_{k,K}$ is the Kronecker delta. By similar calculations for the elements of $S^{(e)}_t$ given by Eqs. (46)–(48), the total $S^{(e)}_t$ in Eq. (44) amounts to

$$S^{(e)}_t = \frac{4\pi}{2} \sum_{k,K} \left( -u^{(e)}_k \tilde{\Omega}^E_k(r) \delta_{k,K-1} + 2u^{(e)}_k \tilde{\Omega}^E_k(r) \delta_{k,K} - u^{(e)}_k \tilde{\Omega}^E_k(r) \delta_{k,K+1} \right).$$

(50)

On the other hand, the action functional of the electric charge of the matter for spherical coordinates is expressed using $u^{(e)}(r)$ in Eq. (40) by [considering also the form $-\nabla^2\phi^{(e)} - \rho^{(e)} = 0$ on the left in Eq. (27)]

$$S^{(e)}_m = -\int dr (4\pi) r^2 \rho^{(e)} \phi^{(e)}(r)$$

$$= -\int dr (4\pi) r \rho^{(e)} u^{(e)}(r).$$

(51)
Subsequently, by the expansion of $u^{(e)}(r)$ given by Eq. (40) in terms of the basis functions in Eq. (19), the above action becomes

$$S_m^{(e)} = -\int dr (4\pi) r \rho^{(e)} \sum_k u_k^{(e)} \tilde{H}_k(r)$$

$$= -(4\pi) \rho^{(e)} \sum_k u_k^{(e)} \frac{r_{k+1/2}^2 - r_{k-1/2}^2}{2}.$$  \hspace{2cm} (52)

Because $R_0 = \Delta_h$ as mentioned below Eq. (24),

$$R_0 = r_{k+1/2}|_{k=1} = r_{1+1/2},$$  \hspace{2cm} (53)

which implies $\rho^{(e)} = 0$ for $k > 1$ (the index 1 is one) in Eq. (52), and $r_{k-1/2} = 0$ (or $r_{k-1/2} = \epsilon$ with $\epsilon \to 0$ after the calculation). Using $Q$ in Eq. (25) and $R_0 = \Delta_h$, we have

$$S_m^{(e)} = -\rho^{(e)} \sum_k u_k^{(e)} \frac{3}{2} \frac{4\pi}{3} \Delta_h^3 \delta_{k,1}$$

$$= -\sum_k u_k^{(e)} \frac{3}{2} \Delta_h Q \delta_{k,1},$$  \hspace{2cm} (54)

where $\delta_{k,1}$ is the Kronecker delta (the index 1 is one).

From Eqs. (50) and (54), the variation with respect to $u_k^{(e)}$,

$$\delta S_l^{(e)} + \delta S_m^{(e)} = 0,$$

leads to

$$-\frac{1}{\Delta_h} \left( u_k^{(e)} - 2u_k^{(e)} + u_k^{(e)} \right) = \frac{3}{2} \frac{Q'}{\Delta_h} \delta_{k,1},$$

with $Q' = \frac{Q}{4\pi}$.  \hspace{2cm} (55)

This equation is equivalent to

$$\frac{u_k^{(e)} - 2u_k^{(e)} + u_k^{(e)}}{2 \Delta_h^2} = \frac{3}{2} \frac{Q'}{\Delta_h} \delta_{k,1},$$  \hspace{2cm} (56)

corresponding to Eq. (27) for $\phi^{(e)}(r) = ru^{(e)}(r)$.

We then have

$$u_{k-1}^{(e)} - 2u_k^{(e)} + u_{k+1}^{(e)} = 0 \text{ for } k > 1,$$  \hspace{2cm} (57)

which is rewritten by

$$u_{k-1}^{(e)} - u_k^{(e)} = u_k^{(e)} - u_{k+1}^{(e)} \text{ for } k > 1.$$  \hspace{2cm} (58)

Additionally, for the boundary $r_{k-1/2}$ with $k = 1$ (the index is 1)

$$r_{1-1/2} = \epsilon > 0,$$  \hspace{2cm} (59)

(we set $\epsilon \to 0$ after the calculation), the basis function is not given in the region for $r < 0$. Considering this boundary for Eq. (57), we obtain

$$-2u_k^{(e)} + u_{k+1}^{(e)} = \frac{3}{2} \frac{Q'}{\Delta_h^2} \text{ for } k = 1.$$  \hspace{2cm} (60)

In contrast, using $Q'$ in Eq. (25) and $R_0 = \Delta_h$ for the charge, the outer boundary condition imposed is

$$u_{N+1}^{(e)} = Q',$$  \hspace{2cm} (61)

which implies $\phi^{(e)}(r_{N+1}) = Q'/r_{N+1}$ in Eq. (40). Then Eq. (59) becomes

$$u_{N-1}^{(e)} - u_N^{(e)} = u_N^{(e)} - Q'.$$  \hspace{2cm} (62)

We consider a solution that takes

$$u_N^{(e)} = (\beta - 1) Q' + Q' \text{ for } k = N,$$  \hspace{2cm} (63)

where $\beta$ is a constant to be determined below. Equations (59) and (62)–(64) lead to

$$u_{N-1}^{(e)} = [((\beta - 1) Q' + Q')$$

$$+ (\beta - 1) Q' - Q']$$

$$= 2(\beta - 1) Q' + Q',$$  \hspace{2cm} (64)

$$u_{N-2}^{(e)} = [2(\beta - 1) Q' + Q']$$

$$+ [2(\beta - 1) Q' - Q'] = 3(\beta - 1) Q' + Q'.$$  \hspace{2cm} (65)

Using Eqs. (58), the sequential manipulations result in

$$u_k^{(e)} = [(N - k + 1)(\beta - 1)]Q' + Q' \text{ for } k > 1.$$  \hspace{2cm} (66)

Because the above solution diverges unless $\beta = 1$ for $k = 2$, we derive the following solution, by setting $\beta = 1$ and using $Q' = Q/(4\pi)$ in Eq. (56):

$$u_k^{(e)} = Q' = \frac{Q}{4\pi} \text{ for } k > 1.$$  \hspace{2cm} (67)

Furthermore, from Eqs. (61) and (68) as well as $Q' = Q/(4\pi)$ in Eq. (56), we have the solution (at the remaining point) for $k = 1$:

$$u_1^{(e)} = \frac{1}{2} \left( \frac{3}{2} Q' + Q' \right) = \frac{5}{4} Q' = \frac{5}{4} \frac{Q}{4\pi} \text{ for } k = 1.$$  \hspace{2cm} (68)

Thus, from Eqs. (29), (40) and (69) with $2r_1 = \Delta_h = R_0 = \Delta/2$, we obtain the following classical electric self-energy in the region with the cut-off length $\Delta$ (whose volume $\Delta^3$ is approximated by $(4\pi/3)\Delta^3$ with the charge density $\rho^{(e)}$ and charge $Q$ in Eq. (25) for this volume):

$$E_{\Omega}^{(e)} = \frac{14\pi}{2} \frac{\Delta_h^3}{3} \rho^{(e)} \frac{5}{4} \frac{Q}{Q_{\Omega}} \left( \frac{1}{4\pi r_1} \right)$$

$$= 1.5 \frac{Q Q_{\Omega}}{24 \pi r_1} = 1 \left[ \frac{5}{2} \frac{Q Q_{\Omega}}{2(4\pi)^2 r_1} \right]$$

$$= \frac{1}{2} \left( \frac{5}{2} \frac{Q Q_{\Omega}}{2(4\pi)^2 \Delta_h} \right) = \frac{1}{2} \left[ \frac{5}{2} \frac{Q Q_{\Omega}}{2(4\pi)(\Delta/2)} \right].$$  \hspace{2cm} (69)
3.3 Self-energy in macroscopic classical Einstein field equations

This subsection presents the analysis of the curvature self-energy in the Einstein field equations. Although the gravitational field is different from the charged particle fields, we treat the Newtonian approximation case, which is similar to the charged particle case. When the renormalization is difficult in this case, it is possible to use the cut-off length. The self-energy is first evaluated by the continuum field theory. Subsequently, the self-energy is evaluated using the formalism in terms of the step-function-type basis functions. As described by Landau and Lifshitz [15](owing to the negligible contributions of higher-order terms with respect to $1/c$ in the Lagrangian with $c$ being the velocity of light), the Newtonian approximation, within the scheme of the Einstein field equations for the matter with the slow velocities compared to $c$, has the infinitesimal squared distance expressed by

$$ (dx)^2 = g_{\mu\nu} dx^\mu dx^\nu = (\eta_{\mu\nu} + h_{\mu\nu}) dx^\mu dx^\nu $$

$$ = \left(1 + \frac{2\phi}{c^2}\right) (dx^0)^2 $$

$$ - \left(1 - \frac{2\phi}{c^2}\right) [(dx^1)^2 + (dx^2)^2 + (dx^3)^2], $$

(71)

where

$$ \eta_{\mu\nu} = \begin{bmatrix}
1 & 0 & 0 & 0 \\
0 & -1 & 0 & 0 \\
0 & 0 & -1 & 0 \\
0 & 0 & 0 & -1 \\
\end{bmatrix}, \quad |h_{\mu\nu}| \ll 1, $$

$$ g_{00} = 1 + \frac{2\phi}{c^2}, $$

(72)

and $\phi$ is the Newtonian potential. We note that the Newtonian potential (field) $\phi$ is distinguished from electric field $\phi^{(e)}$. Letting $\mu_m$ be the density of the mass, we have the energy-momentum tensor

$$ T^\mu_\nu = \begin{cases}
\frac{\mu_m c^2}{4} & \text{for } \mu = 0 \text{ and } \nu = 0 \\
0 & \text{for } \mu \neq 0 \text{ or } \nu \neq 0.
\end{cases} $$

(73)

It is well known that the field equation Eq. (15) can be rewritten

$$ R^\mu_\nu = \frac{8\pi G}{c^4} \left( T^\mu_\nu - \frac{1}{2} T \delta^\mu_\nu \right), $$

(74)

where $\delta^\mu_\nu$ is the unit tensor and

$$ T = g^{\mu\nu} T_{\mu\nu}. $$

(75)

Furthermore, using the well-known relations for Eq. (74)

$$ R^0_0 = \frac{1}{c^2} \frac{\partial^2 \phi}{\partial x^2}, $$

$$ \frac{8\pi G}{c^4} \left( T^0_0 - \frac{1}{2} T^0_0 T \right) = \frac{8\pi G}{c^4} \frac{1}{2} (\mu_m c^2), $$

(76)

and from Eqs. (72)–(75), we obtain the Newtonian equation

$$ \frac{\partial^2 \phi}{\partial x^2} = 4\pi G \mu_m. $$

(78)

From Eq. (72), the term $\delta g^{00} \sqrt{-g}$ in Eq. (14) is approximated by (higher-order terms with respect to $1/c$ in $\sqrt{-g}$ are neglected)

$$ \delta g^{00} \sqrt{-g} \approx +\delta \left( \frac{2\phi}{c^2} \right). $$

(79)

We then approximate the action functional for directly leading to the Newtonian equation as follows. Because the action functional for the matter is linear with respect to $\phi$, this action is approximated by

$$ S^{(N)}_m = -\frac{c^3}{16\pi G} \left( \frac{2}{c^2} \right) \frac{8\pi G}{c^4} \int \frac{1}{2} (\mu_m c^2) \phi dV_4. $$

(80)

Meanwhile, we approximate the following action functional of the gravity, which is consistent with the above equation (the factor 1/2 appears considering the variational of both $\delta^2 \phi/\partial x^2$ and $\phi$ with respect to $\phi$), with integration by parts:

$$ S^{(N)}_g = -\frac{c^3}{16\pi G} \left( \frac{2}{c^2} \right) \frac{1}{2} \int \left( \frac{\partial \phi}{\partial x^1} \left( \frac{\partial \phi}{\partial x^1} \right) \right) dV_4. $$

(81)

By variation with respect to $\phi$, the above action functionals $S^{(N)}_g$ and $S^{(N)}_m$ lead to the Newtonian equation given by Eq. (78).

In the Newtonian approximation within the Einstein scheme, the energy-momentum tensor has a similar form to that in Eq. (29) for the static electric field. Using the notation $\nabla = (\partial x^1, \partial x^2, \partial x^3)$, the static energy is written

$$ E^{(N)} = -\frac{1}{8\pi} \int dr (4\pi r^2 (-\nabla \phi) \cdot (-\nabla \phi) $$

$$ = \frac{1}{2} \int dr (4\pi r^2 \mu_m \phi). $$

(82)

As in Sects. 2.2, 3.1 and 3.2, the 3D sphere, which is the approximation of the cubic element in 3D space with the cut-off length $\Delta$, has the radius $R_0 = \Delta/2$ and mass density $\mu_m$ in Eq. (24) of the fundamental particle. The mass $M$ in Eq. (25) is the product of $\mu_m$ and the volume of the 3D sphere. From Eq. (78), the gravitational potential, which corresponds to Eq. (30) in the electrodynamical case, becomes

$$ \phi(r) = -\frac{GM}{r} \quad \text{for } r > R_0. $$

(83)

A similar correspondence to Eq. (34) for $r \leq R_0$ gives

$$ \phi(r) = -\frac{4\pi}{6} r^2 \mu_m + \frac{4\pi}{6} R_0^3 \mu_m - \frac{GM}{R_0} \quad \text{for } r \leq R_0. $$

(84)

Subsequently, from Eqs. (82), (84) and $M$ in Eq. (25), we have the following gravitational self-energy in the case of the continuum theory:
coordinates, we find a sphere with the radial cut-off length function-type basis functions. We also use the above 3D self-energy in the Einstein field equations, using the step-function-type basis functions in Eq. (19). From Eqs. (87) and (88), we have:

\[ S^{(N)}_g = \sum_{k,K} \left[ -\frac{\gamma^2}{2} \int dr(4\pi r^2) \frac{du_k(r)}{dr} \right] \left( \frac{\partial \hat{\Omega}_E^k(r)}{\partial \mu_K} \frac{\partial \hat{\Omega}_E^k(r)}{\partial \mu_K} \right). \]  

(89)

Similar to the action in Eq. (50) for the electric field, the above action becomes

\[ S^{(N)}_g = -\frac{4\pi}{2} \sum_{k,K} \left[ \frac{1}{\Delta_h} \left( -u_k u_{K-1} \delta_{k,K-1} + 2u_k u_{K} \delta_{k,K} - u_k u_{K+1} \delta_{k,K+1} \right) \right]. \]  

(90)

Meanwhile, using \( u(r) \) in Eq. (22), the action functional of the matter in Eq. (80) for spherical coordinates becomes

\[ S^{(N)}_m = -4\pi G \gamma g \int dr(4\pi r^2) \mu_m \phi(r) \]
\[ = -4\pi G \gamma g \int dr(4\pi r^2) \mu_m u(r). \]  

(91)

By the expression of \( u(r) \) given by Eq. (23) in terms of basis functions denoted in Eq. (19), the above action is written by

\[ S^{(N)}_m = -4\pi G \gamma g \int dr(4\pi r^2) \mu_m \sum_k u_k \delta_k(r) \]
\[ = -4\pi G \gamma g \sum_k u_k \frac{3}{2} \frac{M}{\Delta_h} \delta_{k,1}. \]  

(92)

From Eqs. (90) and (92), the variational calculus with respect to \( u_k \)

\[ \delta S^{(N)}_g + \delta S^{(N)}_m = 0, \]  

(94)

results in

\[ \frac{1}{\Delta_h} (u_{k-1} - 2u_k + u_{k+1}) = \frac{3}{2} \frac{GM}{\Delta_h} \delta_{k,1}. \]  

(95)

Corresponding to the electric field case in Eqs. (67) and (68), the solution obtained for \( k > 1 \) is

\[ u_k = -GM \quad \text{for} \quad k > 1. \]  

(96)

Equations (95) and (96) yield the solution for \( k = 1 \):

\[ u_1 = -\frac{1}{2} \left( \frac{3}{2} GM + GM \right) = -\frac{5}{4} GM \quad \text{for} \quad k = 1. \]  

(97)

Consequently, from Eqs. (22), (82) and (97) with \( r_1 = \Delta_h/2 = (\Delta/2)/2 \), we derive the following classical curvature self-energy in the region with the radial cut-off length \( \Delta_h = R_0 \) (related to the cut-off length \( \Delta \)) and the mass \( M = (4\pi/3)(\Delta_h)^3(\mu_m) \) [in Eqs. (24)–(25)]:

\[ E^{(N)}_\Omega = \frac{1}{2} \frac{14\pi}{3} (\Delta_h)^3 \mu_m \left( \begin{array}{c} 5 \frac{GM}{\Delta_h} \\ \frac{1}{r_1} \end{array} \right) = -\frac{1}{2} \left( \frac{5}{2} \frac{GM}{\Delta_h} \right)^3 = -\frac{1}{2} \left( \frac{5}{2} \frac{GM}{\Delta_h} \right). \]  

(98)

4 Relationship between the subtracted self-energy and vacuum energy constant \( A \)

The continuum relativistic theory requires that a fundamental particle be considered pointlike, and the radius of a pointlike particle leads to ultraviolet divergences. However, our formalism can obtain finite self-energies by expressing fields in terms of the step-function-type basis functions. As in Sect. 2.1, the self-energy subtracted from the energy of the rest mass is included in vacuum energy expressed in terms

\[ S^{(N)}_g + S^{(N)}_m = 0 \]  

and

\[ E^{(N)}_\Omega = E^{(N)} - \sum_k u_k \]  

for the gravit"{a}tion, where \( E^{(N)} \) is the total energy of the system. Therefore, the self-energy in the present static case is rewritten (considering the form \( \nabla^2 \phi - 4\pi G \mu_m = 0 \) of Eq. (78))

\[ S^{(N)}_m = -4\pi G \gamma g \int dr(4\pi r^2) \mu_m \phi(r) \]  

(85)
of the constant \( \Lambda \) (cosmological constant). The self-energy calculated using the step-function-type basis function with the cut-off length \( \Delta \) is not so different from that calculated by the continuum theory with the same cut-off length as was shown in Sect. 3. For a fundamental particle, the self-energy caused by the classical electrodynamics is proportional to \( 1/\Delta \). This self-energy is reduced to the following logarithmic form by quantum corrections \([33,38]\) \((\hbar=\hbar/(2\pi))\) with \( \hbar \) being the Planck constant:

\[
E_Q^{(e)} = \frac{3}{4\pi} \left( \frac{e^2}{4\pi \hbar c} \right) m_E c^2 \ln \left[ \frac{(\hbar c/\Delta)^2}{(m_E c^2)^2} \right] + \frac{1}{2},
\]

(\( \Delta \) is the cut-off length and \( m_E \) is the rest mass of an electro-dynamically interacting fundamental particle.) This reduction of the Coulomb-type self-energy also occurs in chro- 

modynamics with the asymptotic freedom at short distances \([39-41]\). However, a stronger divergence of the self-energy at a quantum level appears in the Higgs boson case. (Fundamental particles (quarks) receive a mass through the coupling to the Higgs field.) The gravitational field is treated at a classical level, because the cut-off length below in this section is longer than the Planck scale and the gravitational strength is small. Under such a small gravitational field, our treatment in this section relates vacuum energy to the Higgs boson self-energy, which is dominant among other interactions at a quantum level. This relation (between vacuum energy and the Higgs self-energy) determines the cut-off length, reproducing today’s tiny value of the cosmological constant \( \Lambda \) as described below.

It is well known that the matter is mainly composed of protons. The averaged energy of the rest mass of the fundamental particles is \( m_E c^2 \approx 3.23 \) [MeV]. Considering that the contribution from the mass of the fundamental particles to that of a proton is very small, we set

\[
\gamma_E = \frac{\text{energy of proton}}{\text{energy of fundamental particles}} \approx 96.7.
\]

(100)

To derive the cut-off length, the ratio \( \gamma_\Lambda \) is defined by

\[
\gamma_\Lambda = \frac{\text{energy density of vacuum energy with } \Lambda}{\text{energy density of matter}} \approx 0.23,
\]

(101)

The fundamental particle (quark) mass \( m_E \) is due to the coupling to the Higgs field with the coupling constant \( \lambda_f \) written by

\[
\lambda_f = \sqrt{\frac{2}{\bar{v}}} m_E.
\]

(102)

where \( v \) is the vacuum expectation value of the symmetry-broken Higgs field. The Higgs self-energy \( E_H \) (included in vacuum energy) for the cut-off length \( \Delta \) and \( \lambda_f \) above is written by

\[
|E_H| = \frac{\sqrt{2}}{4\pi} \frac{\hbar c}{\Delta} \frac{\lambda_f}{v} c^2
\]

\[
= \frac{\sqrt{2}}{4\pi} \frac{\sqrt{2}}{v} \frac{m_E c^2}{\Delta} \frac{\hbar c}{2\pi} \frac{1}{m_E c^2} \frac{\hbar c}{\Delta}.
\]

(103)

Meanwhile, from Eqs. (100) and (101) it follows that

\[
|E_H| = \frac{1}{\gamma_E} \frac{1}{\gamma_\Lambda} \frac{\hbar c}{2\pi} v.
\]

(104)

Combining Eqs. (103) and (104), we have

\[
\Delta = \frac{1}{\gamma_E} \frac{1}{\gamma_\Lambda} \frac{\hbar c}{2\pi} v.
\]

(105)

Because \( v \approx 246 \) [GeV], we derive the cut-off length \( \Delta \approx 7.2 \times 10^{-8} \) [fm], which corresponds to \( \approx 2.7 \times 10^6 \) [GeV], that is, \( \Delta_h \approx 3.6 \times 10^{-8} \) [fm] corresponding to \( \approx 5.5 \times 10^6 \) [GeV].

Even if modifications of values or definitions are required for the above calculations, we obtain a similar value of \( \Delta \). Owing to the above cut-off length \( \Delta \) derived, the field theory may be advanced without ultraviolet divergences. When the renormalization is difficult, the analysis is possible by using the cut-off length. In general, \( g_{\mu\nu}(x) \) is written \( g_{\mu\nu}(x) = \eta_{\mu\nu}(x) + h_{\mu\nu}(x) \) as given in Eq. (71), and the tensor field \( h_{\mu\nu}(x) \) is expressed by

\[
h_{\mu\nu}(x) = \sum_p h_{\mu\nu p} \tilde{\phi}_p^4(x),
\]

(106)

where the coefficients \( h_{\mu\nu p} \) are tensor elements, and \( \tilde{\phi}_p^4(x) \) is the four-dimensional basis function, which takes the value of unity in a hyper-octahedron with the index \( p \) in four-dimensional space-time and vanishes out of the hyper-octahedron. Namely, the present theory divides the space-time continuum of classical general relativity into pieces (hyper-octahedrons). The expression of wave functions in terms of step-function-type basis functions restricts the degrees of freedom of the wave functions in a hyper-octahedron (cuts off high-frequency contributions), meaning the quantization of space-time in classical general relativity. In Sect. 3, it was shown that the quantities calculated using the step-function-type basis functions are similar to the corresponding quantities calculated by using the continuum theory with the cut-off. The formalism and calculated quantities in the continuum theory are mapped to the corresponding formalism and quantities using step-function-type basis functions. Then, from Eqs. (103)–(105), the cut-off length used for the step-function-type basis functions is related to Higgs self-energy, which amounts to the vacuum energy expressed by the constant \( \Lambda \) (of such as Λ-CDM).

In the Einstein field equations, the energy of the rest mass is the source of the curvature, and the renormalization (by...
such as the electrodynamic interaction) subtracts the self-energy from the rest mass. The subtracted self-energies can be involved in vacuum energy constant \( \Lambda \) in the Einstein field equations, as described in Sect. 2. (Concerning the curvature self-energy by the gravitational coupling between the mass and the produced field, it is well known that the general curvature self-energy is not always within the renormalization scheme.) The relatively large cut-off length (compared to the Planck length) of the present theoretical formalism has an advantage with naturalness that the Higgs self-energy is suppressed, and this cut-off is related to today’s tiny vacuum energy expressed by \( \Lambda \), without fine tuning. In contrast, by the relatively small cut-off at the Planck scale in other models arises the huge Higgs self-energy, which needs the following fine tuning. In a highly precise fine tuning, the huge Higgs self-energy for the cut-off at the Planck scale is canceled by another physical quantity to adjust the Higgs mass. Furthermore, the Planck energy composing \( \Lambda \) (cosmological constant) in other models needs fine tuning to obtain the present tiny value of \( \Lambda \) by the cancellation from such as a huge Higgs self-energy for the cut-off at the Planck scale. Therefore, the present formalism provides an answer to the fine-tuning problem. The present formalism also has the possibility to offer a fundamental physical theory, predicting a cut-off length that may play the role of a fundamental physical constant, if experimentally observed. The present model has another merit: that the initial universe has no possibility to form a black hole because of the relatively large cut-off length unlike the cut-off at the Planck scale near the black hole size of the whole universe. Furthermore, the gravity in the present formalism is weak compared to the other fundamental interactions, and space-time coordinates do not largely deviate from classical numbers. Moreover, the present theoretical vacuum energy constant \( \Lambda \) decreases to today’s order of magnitude expressing the vacuum energy density, telling the ratio of vacuum energy density to the energy density of the matter.

If the expansion of the universe is matter dominated (in the present case the vacuum energy caused by the self-energy has the same property of the matter), the vacuum energy constant \( \Lambda \) seems to be proportional to \( t_U^{-2} \) (at least) at present, where \( t_U \) is the age of the universe at each point in time. This is due to the well-known fact that the solution of the Friedmann equation indicates the scale of the universe \( a(t_U) \) as a function of \( t_U \) to be \( a(t_U) \propto t_U^{2/3} \), that is, \( [a(t_U)]^{-3} \propto t_U^{-2} \). The matter density \( \rho_m \) (we can include the dark matter and vacuum energy into the matter) is written by \( \rho_m \propto [a(t_U)]^{-3} \), which leads to \( \rho_m \propto t_U^{-2} \) and \( \Lambda \propto t_U^{-2} \), because the vacuum energy expressed by the vacuum energy constant \( \Lambda \) obeys the relation \( \Lambda \propto \rho_m \) in our scheme. This is the reason why the vacuum energy constant \( \Lambda \) seems to be proportional to \( t_U^{-2} \). In contrast, at the Planck scale the wave packet size for the mass of the Planck energy \( \approx 10^{19} \text{ [GeV]} \), which seems to be the whole energy of the universe, is equal to the gravitational-based radius, and the conditions between the above two cases are quite different. When the present model is generalized to the early universe, the initial size of the universe is the cut-off length of the present model. Because \( \Lambda \propto t_U^{-2} \) mentioned above diverges in the limit as \( t_U \rightarrow 0 \) (\( t_U \) is larger than the cut-off length), the early universe expands rapidly, although the expansion rate is different from the exponential expansion of inflation models.

Finally, we add that, owing to the nonperiodic boundary condition, the zero-point energy for the candidate of vacuum energy constant \( \Lambda \) is not seen in the present formalism. Let us consider a simple action for the wave function \( \Phi(x) \),

\[
S_B = \frac{1}{2} \int dx \left( \frac{d\Phi(x)}{dx} \right)^2.
\]  

(107)

The wave function in terms of the basis functions in Eq. (19)

\[
\Phi(x) = \sum_k \Phi_k \Omega_k(x).
\]  

(108)

Similar to Eq. (50), we write the action functional

\[
S_B = \frac{1}{2} \sum_{k,K} \int dx \left( \Phi_k \Phi_K \frac{d\Omega_k(x)}{dx} \frac{d\Omega_K(x)}{dx} \right)
\]

\[
= \frac{1}{2} \frac{1}{\Delta} \sum_{k,K} (\Phi_k \Phi_K)(-\delta_{k-1,K} + 2\delta_{k,K} - \delta_{k+1,K}).
\]  

(109)

The variation with respect to \( \Phi_k \),

\[
\delta S_B = 0,
\]  

(110)

yields

\[
\frac{1}{\Delta}(\Phi_{k-1} - 2\Phi_k + \Phi_{k+1}) = 0,
\]  

(111)

which leads to

\[
\frac{1}{\Delta^2}(\Phi_{k-1} - 2\Phi_k + \Phi_{k+1}) = 0,
\]  

(112)

where \( k, K = 1, 2, ..., N_x \). For the above equation, the following boundary conditions on the wave function are imposed \((N_x + 2)\) is the number of lattice points, and the lattice indices of the boundary points are denoted by \( k = 0, N_x + 1 \):

\[
\Phi_0 = 0, \quad \Phi_{N_x+1} = 0.
\]  

(113)

Similar to the classical vibrational case [42], the eigenvector for a diagonalization of the action is expressed as

\[
\Phi_k = \frac{1}{C_N} \sin \left( \frac{k K \pi}{N_x + 1} \right).
\]  

(114)
where \( C_N \) is a normalization constant. Then the element \( S_{B,k} \) of the action in Eq. (109) is diagonalized giving

\[
S_{B,k} = \frac{1}{2} \sum_k (-\delta_{k-1,K} + 2\delta_{k,K} - \delta_{k+1,K}) \Phi_K
\]

\[
= \frac{1}{2\Delta C_N} \left[ k(k-1)\pi \right] \sin \left( \frac{k\pi}{N_x+1} \right) + 2 \sin \left( \frac{k\pi}{N_x+1} \right) - \sin \left( \frac{k(k+1)\pi}{N_x+1} \right)
\]

\[
= \frac{1}{\Delta C_N} \left[ 1 - \cos \left( \frac{k\pi}{N_x+1} \right) \right] \sin \left( \frac{k\pi}{N_x+1} \right) \left( \frac{k\pi}{N_x+1} \right)
\]

yielding the eigenenergies

\[
\eta_k \propto 1 - \cos \left( \frac{k\pi}{N_x+1} \right),
\]

with \( k = 1, 2, \ldots, N_x \). Consequently, the zero-point energy for the candidate vacuum energy is not seen in the present system because of the boundary condition in Eq. (113). (Similarly, eigenvalues in higher dimensions are obtained \([10,12]\).)

We note that when the system is considered using a box normalization, in which wave functions are defined in a box with periodic boundary conditions at the box surfaces, eigenvalues may have zero-point energies. However, physical quantities such as the transition amplitude are calculated without using the zero-point energies by expressing plane waves in the form of complex exponential functions. The zero-point energies dropped in this case may not be included in vacuum energy, because the zero-point energies appearing here are due to the non-vanishing periodicity (which seems to lack in the real expanding universe) in approximate calculational manipulations.

As the dark matter, we considered the classical solution with quantum field fluctuations in chromodynamics in Ref. [12]. Although the Big Bang is out of the scope of this paper, an expansion may arise making the Big Bang like the vaporization of water in vacuum by absorbing heat.

5 Conclusions

For the renormalization of the mass, we have considered the subtracted self-energies, which act as sources of the curvature in the Einstein field equations. It was shown that this consistency is satisfied by including these self-energies into the vacuum energy expressed by the constant \( \Lambda \). The self-energies in electrodynamics and Einstein field equations were investigated by using the ultraviolet cut-off length. The field theory, which was developed by the present authors, expresses wave functions in terms of the step-function-type basis functions to cut off oscillations at short distances. In the other continuum theory, we used the same cut-off length as that used for the former theory. From the examination, the continuum theory with the cut-off length is effective. Classical self-energies are reduced to logarithmic forms by quantum corrections, and the quadratic Higgs self-energy is dominant at a quantum level. The cut-off length was determined so as to reproduce the observed vacuum energy constant \( \Lambda \), using the self-energy derived from the above cut-off theories. The derived vacuum energy expressed by the constant \( \Lambda \) is of the order of the matter (composed of the conventional matter such as atoms and dark matter), showing that the vacuum energy constant \( \Lambda \) has today’s tiny value.

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References

1. W.E. Lamb Jr., R.C. Retherford, Phys. Rev. 72, 339 (1947)
2. V.S. Weisskopf, Kongelig Danske Videnskabernes Selskab, Mathematisk-fysiske Meddeleler XIV(6) (1936)
3. V.S. Weisskopf, Phys. Rev. 56, 72 (1939)
4. S. Tomonaga, Prog. Theor. Phys. 1, 27 (1946)
5. S. Tomonaga, Phys. Rev. 76, 224 (1948)
6. J. Schwinger, Phys. Rev. 76, 790 (1949)
7. R.P. Feynman, Phys. Rev. 76, 769 (1949)
8. F.J. Dyson, Phys. Rev. 75, 1736 (1949)
9. K. Fukushima, Phys. Rev. D 30, 1251 (1984)
10. K. Fukushima, H. Sato, Bulg. J. Phys. 41, 142 (2014). (arXiv:1402.0450v5). Freely available at http://www.bjp-bg.com/papers/bjp2014_2_142-171.pdf
11. K. Fukushima, H. Sato, Bulg. J. Phys. 43, 30 (2016). (arXiv:1501.04837v6). Freely available at http://www.bjp-bg.com/papers/bjp2016_1_030-044.pdf
12. K. Fukushima, H. Sato, Int. J. Mod. Phys. A 32, 1730017 (2017). (arXiv:1705.03767v6)
13. A.R. Mitchell, R. Wait, The Finite Element Method in Partial Differential Equations (Wiley, New York, 1977)
14. C.M. Bender, K.A. Milton, D.H. Sharp, Phys. Rev. Lett. 51, 1815 (1983)
15. L.D. Landau, E.M. Lifshitz, The Classical Theory of Fields, 4th Revised English Edition (Elsevier, Amsterdam, 1951)
16. A. Einstein, Sitzungsberichte der Königlich Preussischen Akademie der Wissenschaften Berlin, part 1, 142 (1917)
17. A. Friedmann, Z. Phys. 10, 377 (1922)
18. S.M. Carroll, H.P. William, E.L. Turner, The cosmological constant. Ann. Rev. Astron. Astrophys. 30, 499 (1992)
19. S.M. Carroll, Cosmol. Constant Living Rev. Relat. 4, 1 (2001)
20. S. Carroll, Spacetime and Geometry (Addison Wesley, San Francisco, CA, 2004), p. 171
21. A. Riess et al., Astron. J. 116, 1009 (1998)
22. J.C. Baker et al., Mon. Not. R. Astron. Soc. 308, 1173 (1999)
23. S. Perlmutter et al., Astrophys. J. 517, 565 (1999)
24. P.J.E. Peebles, B. Ratra, Rev. Mod. Phys. 75, 559 (2003)
25. T. Padmanabhan, Phys. Rep. 380, 235 (2003)
26. M. Tegmark et al., Phys. Rev. D 69, 103501 (2004)
27. L. Lombriser, L. Lucas, A. Nelson, Phys. Lett. B 765, 382 (2017)
28. J.D. Barrow, D.J. Shaw, Gen. Relat. Gravity 43, 2555 (2011)
29. S. Rugh, H. Zinkernagel, Stud. Hist. Philos. Modern Phys. 33, 663 (2001)
30. M.P. Hobson, G.P. Efstathiou, A.N. Lasenby, General Relativity: An Introduction for Physicists (Cambridge University Press, Cambridge, 2006)
31. S. Weinberg, Rev. Mod. Phys. 61, 1 (1989)
32. H. Leutwyler, J.R. Klauder, L. Streit, Nuovo Cim. A 66, 536 (1970)
33. J.D. Bjorken, S.D. Drell, Relativistic Quantum Mechanics (McGraw-Hill, New York, 1964)
34. F. Rohrlich, Phys. Today 15, 19 (1962)
35. P.A.M. Dirac, Proc. R. Soc. (Lond.) A167, 148 (1938)
36. F. Rohrlich, Am. J. Phys. 28, 639 (1960)
37. F. Rohrlich, Nuovo Cim. 21, 811 (1961)
38. H. Yukawa, Y. Katayama, Soryushi-ron (Theory of Elementary Particles) (Iwanami, Tokyo, 1974). (in Japanese)
39. D.J. Gross, F. Wilczek, Phys. Rev. Lett. 30, 1343 (1973)
40. H.D. Politzer, Phys. Rev. Lett. 30, 1346 (1973)
41. P. Ramond, Field Theory: A Modern Primer, 3rd prn (Benjamin, MA, 1982)
42. J.C. Slater, N.H. Frank, Mechanics (McGraw-Hill, New York, 1947)