Exploring the multi-humped fission barrier of $^{238}$U via sub-barrier photofission

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The photofission cross-section of $^{238}$U was measured at sub-barrier energies as a function of the $\gamma$-ray energy using, for the first time, a monochromatic, high-brilliance, Compton-backscattered $\gamma$-ray beam. The experiment was performed at the High Intensity $\gamma$-ray Source (HI$\gamma$S) facility at beam energies between $E_\gamma=4.7$ MeV and $6.0$ MeV and with $\sim 3\%$ energy resolution. Indications of transmission resonances have been observed at $\gamma$-ray beam energies of $E_\gamma=5.1$ MeV and $5.6$ MeV with moderate amplitudes. The triple-humped fission barrier parameters of $^{238}$U have been determined by fitting EMPiRE-3.1 nuclear reaction code calculations to the experimental photofission cross section.

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Photofission measurements enable selective investigation of extremely deformed nuclear states in the light actinides and can be utilized to better understand the landscape of the multiple-humped potential energy surface (PES) in these nuclei. The selectivity of these measurements originates from the low and reasonably well-defined amount of angular momentum transferred during the photoabsorption process. The present study is designed to investigate the PES of $^{238}$U through observation of transmission resonances in the prompt photofission cross section. A transmission resonance appears when directly-populated excited states in the first potential minimum overlap energetically with states either in the superdeformed (SD) $2^{nd}$ or hyperdeformed (HD) $3^{rd}$ potential minima [1, 2]. The fission channel can thus be regarded as a tunneling process through the multiple-humped fission barrier as the gateway states in the first minimum decay through states in the other minima of the PES. So far, transmission resonances have been studied primarily in light-particle-induced nuclear reactions through charged-particle, conversion-electron or $\gamma$-ray spectroscopy. These studies do not benefit from the same selectivity found in photonuclear excitation and consequently they are complicated by statistical population of the states in the $2^{nd}$ and $3^{rd}$ minima with a probability of $10^{-4} - 10^{-5}$. This statistical population leads to a typical isomeric fission rate from the ground-state decay of the shape isomer in the $2^{nd}$ minimum of $\sim 1/\text{sec}$. These measurements have also suffered from dominating prompt-fission background.

Until now, sub-barrier photofission experiments have been performed only with bremsstrahlung photons and have determined only the integrated fission yield. In these experiments, the fission cross section is convolved with the spectral intensity of the $\gamma$-ray beam, resulting in a typical effective $\gamma$-ray bandwidth $\Delta E/E$ between $4 \times 10^{-2}$ and $6 \times 10^{-2}$. These experiments observe a plateau, referred to as the “isomeric shelf”, in the fission cross section, resulting from competition between prompt and delayed photofission [3, 4]. Higher-resolution studies can be performed at tagged-photon facilities, though only with marginal statistics, due to the limited beam intensities realizable through tagging [3]. This beam intensity cannot be significantly improved beyond $\sim 10^4 \gamma/(\text{keV} \cdot \text{s})$, since it is determined by the random coincidence contribution in the electron-tagging process. Thus, high statistics photofission experiments in the deep sub-barrier energy region, where cross sections are typically as low as $\sigma=1 \text{ nb-10}\mu\text{b}$, cannot be performed with tagged-photon beams. The relatively recent development of inverse-Compton scattering $\gamma$-ray sources, capable of producing tunable, high-flux, quasi-monoenergetic $\gamma$-ray beams by Compton-backscattering of eV-range photons off a relativistic electron beam, offers an opportunity to overcome previous limitations. The present study was carried out at such a facility: the High Intensity $\gamma$-ray Source (HI$\gamma$S) located at TUNL. It should be emphasized that a measurement of the photofission cross section in the deep sub-barrier energy region will be a crucial step towards a reliable characterization of the PES, including unambiguous determination of the double- or triple-humped nature of the surface and precise evaluation of the barrier parameters. Next-generation Compton-backscattering $\gamma$-ray sources, such as MEGa-ray (Lawrence Livermore National Laboratory, California, US) [5] and ELI-NP (Bucharest, Romania) [6], are anticipated to provide beams with spectral fluxes of $\sim 10^6 \gamma/(\text{eV} \cdot \text{s})$ and energy resolution of $\Delta E \approx 1 \text{keV}$, far superior to those currently available at HI$\gamma$S. The ca-
abilities of these next-generation sources allow one to aim at an identification of sub-barrier transmission resonances in the fission decay channel with integrated cross sections down to $\Gamma \sigma \approx 0.1$ eV·b, whereas the present study is only sensitive to resonances with $\Gamma \sigma \approx 10$ eV·b. The narrow energy bandwidth expected for the new $\gamma$-ray beam facilities will also allow for a significant reduction of the presently dominant background from non-resonant processes. Thus, next-generation $\gamma$-ray sources are expected to allow preferential population and identification of vibrational resonances in the photofission cross section and ultimately to enable observation of the fine structure in the isomeric shelf. This may open the perspective towards a new era of photofission studies.

Sub-barrier photofission of $^{238}U$ so far has only been studied with intense bremsstrahlung, however, without being able to resolve any resonances [8]. A previous $^{236}U$(p,t) measurement showed pronounced resonance structures at excitation energies of $E^* = 5.6 - 5.8$ MeV and at $E^* = 5.15$ MeV, as well as a weaker resonance at $E^* = 4.9$ MeV [9]. A whole sequence of further transmission resonances at lower energies is expected to explain the isomeric shelf [4], but such resonances have not yet been observed. Furthermore, it has been found experimentally in several measurements on $^{232}U$ [10], on $^{236}U$ [11] and most recently on $^{232}U$ [12] in agreement with older theoretical predictions [13], that for the uranium isotopes the HD $3^{rd}$ potential minimum is in fact as deep as the SD $2^{nd}$ minimum. According to this experimental systematics, the existence of a HD $3^{rd}$ minimum is also predicted for $^{238}U$, however, it has not yet been supported experimentally. On the other hand, recent calculations using a macroscopic-microscopic model do not predict the existence of a deep $3^{rd}$ minimum for the even-even uranium isotopes [14, 15]. This puzzle was more recently addressed within a self-consistent theoretical approach, where the conditions for the existence of HD potential minima were studied [16].

The aim of the present study was to measure the $^{238}U(\gamma,f)$ cross-section at deep sub-barrier energies and to search for transmission resonances. The experiment was performed at the Hi2S facility with its Compton-backscattered $\gamma$-ray beam, having a bandwidth of $\Delta E = 150-200$ keV and a spectral flux of about $10^2 \gamma/(\text{eV} \cdot \text{s})$. An array of parallel plate avalanche counters, consisting of 22 electrolytically-deposited $^{238}$UO$_2$ (2 mg/cm$^2$) targets [17], was used to measure the photofission cross section. Both fission fragments were detected in coincidence to suppress the $\alpha$-particle background to an extremely low level, which is required by the particularly low counting rates (typically 0.1-1 Hz at $E_\gamma = 5$ MeV). The total efficiency of the array was estimated to be 70% based on Ref. [17].

The present experimental photofission cross-section of $^{238}U$ as a function of the $\gamma$-ray energy is shown in Fig.1a, along with the experimental data of Ref. [8]. Near the top of the barrier the two data sets are in a good agreement. The present data are extended by about an order of magnitude in cross section to the deep sub-barrier region down to $E_\gamma = 4.7$ MeV. A clear transmission resonance has been observed at $E_\gamma = 5.6$ MeV, which is consistent with the observation of Ref. [8]. A slight deviation from the exponential slope of the cross section indicates the existence of a resonance at $E_\gamma = 5.1$ MeV, however, with only a limited resonance signal contrast due to the moderate bandwidth of the $\gamma$-ray beam.

For the theoretical evaluation of the present $^{238}U$ photofission experimental data, we performed nuclear reaction code calculations using the EMPIRE-3.1 code [20].

![Graphical representation of the photofission cross-section for $^{238}U$.](image-url)
Within the code, the fission transmission coefficients are calculated using the Hill-Wheeler formalism [21], followed by Hauser-Feshbach statistical model calculations [22], allowing the fission channel to compete with emission of particles and photons. The triple-humped fission barrier parameters of $^{238}$U were extracted by tuning the inputs to these calculations and comparing the resulting predictions of the photofission cross section to the experimental data.

The reliability of the code was tested and the relevant model parameters were adjusted using calculations of the total photo-absorption cross section $\sigma_{\gamma,\text{abs}}$ and experimental $(\gamma,n)$ cross section data. First, $\sigma_{\gamma,\text{abs}}$ had to be determined and checked against existing experimental data. In the present evaluation, the modified Lorentzian parameterization (MLO) was chosen for the $\gamma$-ray strength function. Although the experimental data of Ref. [18] are quite well reproduced (solid line in Fig. 1), the experimental results of Ref. [19] are underestimated at lower energies. Yet, we have not attempted to tune the MLO parameters to reproduce the experimental data. The parameterization used is based on a global fit of experimental data over a wide range of isotopes and excitation energies. Attempts to reproduce this dataset would have a drastic impact on the competing reaction channels, leading (especially for fission) to unphysical parameters. The photo-absorption cross sections of Ref. [19] were inferred from the measured energy-averaged, angle-integrated photon elastic-scattering cross sections $\sigma_{\gamma,\gamma}$ employing a complex analysis technique described in details in Ref. [23]. In such an analysis, the measured values are renormalized by an energy-dependent factor to obtain the corrected photo-absorption cross section $\sigma_{\gamma,\text{abs}}$. Our calculated values are located between the measured $\sigma_{\gamma,\gamma}$ and the corrected $\sigma_{\gamma,\text{abs}}$ cross sections, perhaps indicating systematic uncertainties in the aforementioned analysis.

The transmission coefficients for the particle emission were determined using the global optical parameter set of Ref. [24]. The level density parameters were taken from the enhanced generalized super-fluid model [22], adjusted to fit the discrete level scheme of $^{238}$U. Those were taken from the most recent reference input parameter library (RIPL3). In the code, the optical model for fission [26–28] is applied to calculate the fission transmission coefficients. For comparison, both triple- and double-humped fission barriers were used in the calculations.

The parameters of the double-humped fission barrier were taken from the RIPL3 library and were slightly adjusted to achieve a better description of the present data.

In Figure 1, the dashed line shows the calculated $(\gamma,f)$ cross-section using the parameters listed in Table I. The triple-humped barrier parameters were adjusted to best describe the experimental photofission and $(\gamma,n)$ cross sections over the entire energy range. In Figure 1, the solid line represents the best description with the parameters of Table II used in the calculation. The calculated $(\gamma,n)$ cross-sections are shown as the solid line in Fig. 1, together with the available experimental data [29, 30]. The calculated and the experimental values are in a fair agreement. The uncertainties of the barrier parameters were estimated to be 200 keV for the barrier heights and 100 keV for the curvature parameters.

The present model is capable of reproducing the subbarrier fission resonances empirically, while at higher excitation energies it naturally provides the same results for the fission barrier penetration as the classical models. Since photofission occurs only through the giant dipole resonance, it is not important to consider the negative parity states that are involved in the calculations only to achieve consistency with the neutron-induced fission cross sections (e.g. $n + ^{238}$U, where $^{238}$U is involved in the second chance fission).

The experimental data of the present experiment could be reproduced dramatically better with a calculation assuming a triple-humped fission barrier than with a double-humped one. When using a triple-humped barrier, an additional resonance at $E_\gamma = 4.6$ MeV had to be included in the calculations. Experimental evidence for the existence of such a resonance would fully confirm our present theoretical interpretation. It is also evident that the existing $(\gamma,f)$ and $(\gamma,n)$ experimental data suffer from large uncertainties. It would be highly important to im-

### Table I. Double-humped fission barrier parameters of $^{238}$U (in MeV) used in the calculations. The resulting photofission cross section is indicated in Figure 1 by the dashed line.

| $E_1$ | $E_{II}$ | $E_{III}$ | $\hbar\omega_1$ | $\hbar\omega_{II}$ | $\hbar\omega_{III}$ |
|-------|----------|-----------|------------------|-------------------|-------------------|
| 6.3±0.2 | 2.0±0.2 | 5.8±0.20 | 1.1±0.1 | 1.0±0.1 | 0.6±0.1 |

![FIG. 2. The triple-humped fission barrier of $^{238}$U as determined in the present study, using the parameters listed in Table I](image-url)


TABLE II. Triple-humped fission barrier parameters of $^{238}$U (all in MeV) used in the calculation, represented by the solid line in Figure 3.

| $E_A$ | $E_{II}$ | $E_B$ | $E_{III}$ | $E_C$ | $h\omega_A$ | $h\omega_{II}$ | $h\omega_B$ | $h\omega_{III}$ | $h\omega_C$ |
|-------|---------|-------|-----------|-------|-------------|--------------|-------------|---------------|-------------|
| 4.3±0.2 | 2.05±0.20 | 5.6±0.2 | 3.6±0.2 | 5.6±0.2 | 0.4±0.1 | 1.0±0.1 | 0.7±0.1 | 1.0±0.1 | 0.7±0.1 |

FIG. 3. (Color online) The height of the inner barrier $E_A$ and the depth of the third minimum $E_{III}$ for even-even uranium isotopes, shown as red circles and green triangles, respectively. The experimental data for $^{232}$U, $^{234}$U and $^{236}$U were taken from Refs. [10–12].

The present results on the fission barrier parameters of $^{238}$U supplement the previous findings on the systematics of the barrier parameters of the uranium isotopes [12]. Fig. 3 shows the present results on $^{238}$U together with previous experimental results on $^{232}$U [12], $^{234}$U [10] and $^{236}$U [11]. A reversal of the trends followed by the lighter uranium isotopes for the height of the inner barrier $E_A$ and the depth of the third minimum (expressed by $E_{III}$), respectively, as a function of the neutron number, is observed. For $^{238}$U, the data suggests a decreasing barrier height $E_A$ and a decreased depth of the third minimum. Moreover, the particularly low values of the curvature parameters derived from the present data, especially the one for the inner barrier ($h\omega_A=0.4$ MeV), may suggest a need for reconsideration of the well-accepted approximation of the fission barrier with a harmonic oscillator potential curve. An anharmonic, “tower-like” potential, originally suggested by Bowman et al. decades ago [31], would better approximate the potential landscape determined from the current data.

In summary, we measured the photofission cross-section of $^{238}$U in the $\gamma$-ray energy region of $E=4.7-6.0$ MeV with the monochromatic, high-brilliance, Compton-backscattered $\gamma$-ray beam of the HI\(-\gamma\)S facility. With the significantly higher intensity of the beam, when comparing to a tagged-photon facility, the cross-section could be measured at deep sub-barrier energies. EMPIRE-3.1 reaction code calculations were performed to extract the fission barrier parameters of $^{238}$U. Our present results on the fission barrier of $^{238}$U support a deep 3rd minimum with $E_{III}=3.6$ MeV, a low inner barrier height $E_A=4.3$ MeV and outer barrier heights of $E_B=5.7$ MeV and $E_C=5.7$ MeV. Though in line with the extensive body of experimental evidence for deep third potential minima in uranium isotopes acquired over the last 15 years, this result is in disagreement with recent calculations of Ref. [14], a puzzle that still needs to be resolved. Indications of predicted resonance structures have also been observed, however, with moderate amplitudes. The results indicate the need for further investigations at lower $\gamma$-ray energies and using smaller-bandwidth, higher-intensity $\gamma$-ray beams. ELI-NP, MEGA-ray, and other next-generation $\gamma$-ray sources will enable such measurements.

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[16] J.D. MacDonnell, W. Nazarewicz and J.A. Sheikh, arXiv:1302.1165v1 [nucl-th] (2013).
[17] J. Drexler et al., Nucl. Instrum. Methods 220, 409 (1984).
[18] G.M. Gurevich et al., Nucl. Phys. A273, 326 (1976).
[19] Y. Birenbaum et al., Phys. Rev. C 36, 1293 (1987).
[20] M. Herman et al., Nucl. Data Sheets 108, 2655 (2007).
[21] D.L. Hill and J.A. Wheeler, Phys. Rev. 89, 1102 (1953).
[22] W. Hauser and H. Feshbach, Phys. Rev. 87, 366 (1952).
[23] P. Axel, K.K. Min, and D.C. Sutton, Phys. Rev. C 2, 689 (1970).
[24] R. Capote, E.Soukhovitskii, S.Chiba and J.M.Quesada, J. Nucl. Sci. Technol. 45, 333 (2008).
[25] A. D’Arrigo et al., J. Phys. G 20, 305 (1994).
[26] B.S. Bhandary, Phys. Rev. C 22, 606 (1980).
[27] M. Sin, R. Capote, A. Ventura, M. Herman and P. Oblozinsky, Phys. Rev. C 74, 014608 (2006).
[28] M. Sin and R. Capote, Phys. Rev. C 77, 054601 (2008).
[29] V.V. Varlamov, N.N. Peskov, D.S. Rudenko and M.E. Stepanov, Vop. At. Nauki i Tekhn., Ser. Yadernye Konstanty, Vol. 2003, 48 (2003).
[30] V.V. Varlamov and N.N. Peskov, Moscow State Univ. Inst. of Nucl. Phys. Reports, No. 2007, 829 (2007).
[31] C.D. Bowman, I.G. Schröder, C.E. Dick and H.E. Jackson, Phys. Rev. C 12, 863 (1975).