The Indian Ocean and the bordering continental margins are characterized by a number of tectonic features consisting of deep ocean basins, aseismic ridges/submarine plateaus and seamounts. Marine geophysical studies carried out over these regions for the past four decades have led to an increased understanding of the formation and evolution of these features. The present paper synthesizes an up-to-date compilation of the results of the salient inferences of those studies to describe the structure and tectonic evolution of these features as well as the Indian Ocean in general. A review of the existing knowledge on the structure and tectonics of the continental margins of India and the adjacent ocean basins reveal the existence of several unresolved problems, primarily due to the inherent tectonic complexities or the paucity of adequate data. Some of the major problems that need to be addressed/examined are: the existence and extent of the Prathap and Kori-Comorin ridges; the locations of the continent-ocean boundaries around the continental margins of India; the timing of formation of the deep offshore regions adjacent to the Indian continental margins; the deeper crustal structure and trend of the 85°E Ridge; the crustal configuration as well as the extent of the postulated micro-continental slivers; and the cause of intense seismicity over the Chagos Bank segment of the Laccadive-Chagos Ridge. In addition to addressing these questions, the present paper suggests ways to go forward by acquiring high-resolution and deep penetration multichannel seismic reflection data, seismic refraction data using Ocean Bottom Seismometers (OBS) and closely spaced sea-surface and deep-tow magnetic profiles along with high-resolution seafloor mapping.

Introduction

The present day configuration of the continents and the deep ocean basins of the Indian Ocean (Figure 1) resulted from the fragmentation and dispersal of the eastern sector of the Gondwanaland since the Late Jurassic. In this framework, the Indian Ocean was formed by rifting and subsequent drifting among Africa, Antarctica, Australia, India, Madagascar and Seychelles. The deeper basins of the Indian Ocean bear the imprints of this plate tectonic history, superposed over time by several other tectonic events such as the Kerguelen, Marion and Réunion hotspots volcanism, the India-Eurasia collision and the deformation of the Central Indian and Wharton basins. A detailed understanding of the evolution of all these resultant features is important to trace the plate kinematic history of the Indian Ocean and the bordering continental margins.

Geophysical investigations carried out in the Indian Ocean and the adjacent continental margins for the past four-odd decades have built a significant knowledge-base on the structure and tectonics of the different sectors of the Indian Ocean. However, partly due to the inherent tectonic complexities of the region and partly on account of the paucity of adequate data, several questions which have a strong bearing on the plate tectonic evolutionary history of the Indian Ocean still persist. This paper is an attempt to highlight some of the more important unresolved questions against the backdrop of our current status of knowledge on the structure and tectonics of the Indian Ocean region with emphasis on the continental margins of India and the adjacent deep ocean basins.

Tectonic framework - current status of knowledge

Continental margins of India

The continental margin of India is subdivided into western and eastern continental margins. Both these are passive margins formed during different episodes of breakup and dispersal of the East Gondwanaland.
The western Indian continental margin consists of (i) a wide continental shelf extending NW-SE, (ii) a remarkably straight shelf edge limited by the 200 m isobath and (iii) a narrow continental slope bounded by 200 m and 2000 m isobaths (Naini, 1980; Biswas, 1989). From north to south, these are the Kutch Basin, Saurashtra Basin, Mumbai Basin, Konkan Basin and the Kerala Basin, delimited by the nearly E-W trending Saurashtra Arch, Son-Narmada Lineament, Vengurla Arch and the Tellicherry Arch, respectively (Sriram et al., 2006; Bastia and Radhakrishna, 2012). The Kutch Basin forms the northwestern part of the western continental margin, between Nagar-Parkar Ridge in the north and the ENE Saurashtra Arch to the south (Biswas, 1982, 1987; Sriram et al., 2006; Bisen et al., 2010). These basement features cut across several coast-parallel horst-graben structures in the continental shelf. The Saurashtra Basin is located between the Kutch Basin in the north and the Mumbai Basin in the south. The basement trends in this basin are nearly E-W, as evidenced by the continuation of Narmada and Kutch rifts and the Saurashtra Arch onto the continental shelf (Biswas, 1982; Bhattacharya and Subrahmanyam, 1986). The Mumbai Basin, located between the Saurashtra Basin and the Konkan Basin, is bounded to the north by the offshore extension of the Narmada-Son lineament and to the south by the Vengurla Arch. The continental shelf in this part contains a system of shelfal horst-graben complex, parallel to the NNW-SSE trending Dharwar trend (Biswas, 1989; Gombos et al., 1995). The main structural elements that constitute the Mumbai Basin are the Surat Depression, Murud Depression, Rajapur Depression, Bombay Platform, Heera-Bassein Block, Ratagirigiri Block and the Shelf Margin Basin, along with the N-S trending Dia Arch and the NE-SW trending Ratagiri Arch (Mathur and Nair, 1993). The Konkan Basin is located between the Mumbai Basin in the north and the Kerala Basin in the south. The Mumbai and Konkan basins are separated by the southwesterly plunging Vengurla arch (Mathur and Nair, 1993), while the Konkan Basin and Kerala Basin are separated by the Tellicherry Arch. Subrahmaniyam et al. (1995) inferred the presence of a coast-parallel horst-graben complex consisting of the inner shelf, mid-shelfal basement ridge sedimentary sub-basins delimited by transverse basement arches or fault-bounded highs characterise the shelf (Figure 2; Ramaswamy and Rao, 1980; Biswas and Singh, 1988; Biswas, 1989; Singh and Lal, 1993).
and a shelf margin basin in the Konkan Basin, between 12°20’N and 15°00’N. Rao et al. (2010) identified a chain of parallel, discrete margin highs existing immediately west of this shelf margin basin in the mid-continental slope region off southwest coast of India.

The Kerala Basin is the southernmost basin in the western continental margin of India, extending from the Tellicherry Arch in the north to Cape Comorin in the south (Figure 2). The entire western continental margin in the northern part shows a regular shelf-slope topography, but the mid-continental slope region south off Cochin shows an anomalous topography due to the presence of two contiguous terraces, the Alleppey Terrace and the Trivandrum Terrace (Figure 3), together known as the Alleppey- Trivandrum Terrace Complex (Yatheesh et al., 2006; Yatheesh et al., 2013b). The seafloor depth of the Alleppey Terrace varies between 300 m and 400 m, while that of the Trivandrum Terrace varies between 1500 m and 1900 m (Yatheesh et al., 2013b). The southerly limit of the Alleppey Terrace is defined by an E-W trending scarp face, referred to as the Quilon Escarpment. The seismic reflection data in and around this region reveals that the Alleppey Terrace is associated with a basement high in the west of its present-day geometry (Unnikrishnan et al., 2018). The Trivandrum Terrace is characterized by the presence of a NNW-SSE trending wide basement high in its central part that is flanked by thick sediment-filled grabens in its easterly and westerly sides (Figure 3b, 3c), referred to as the TT-Eastern Basin and the TT-Western Basin, respectively (Yatheesh et al., 2013b). Both these basins are characterized by the presence of NNW-SSE trending block-faulted basement features, extending across the entire Trivandrum Terrace. The western limit of the Alleppey-Trivandrum Terrace Complex is defined by a ~500 km long steep linear escarpment, known as the Chain-Kairali Escarpment. Some authors (Subrahmanyam and Chand, 2006; Nathaniel et al., 2008; Nathaniel, 2013) considered the linear gravity low associated with this feature to represent the landward extension of the Vishnu Fracture Zone. However, based on updated tectonic elements information from the conjugate continental margins and the adjacent deep offshore regions off southwest coast of India and the southeast coast of Madagascar and with the help of revised plate tectonic reconstruction of the Western Indian Ocean, Shuhail et al. (2018) demonstrated that the Chain-Kairali Escarpment does not represent the landward extension of the Vishnu Fracture Zone.

**Eastern continental margin of India**

The eastern continental shelf of India (Figure 2) trends in a NE-SW direction towards north in the regions off West Bengal, Odisha, and Andhra Pradesh, and in an ~N-S direction towards south, off Tamil Nadu. Compared to the western continental margin of India, the shelf in the eastern continental margin of India is found to be irregular and narrower, with a variable width of ~120 km around Digha, ~60 km off north Andhra Pradesh and ~35 km off Tamil Nadu (Faruque et al., 2014). The water depth of the shelf break varies irregularly from ~220 to ~70 meters. The maximum value of ~220 m depth for the shelf break is observed from Gopalpur to the Chilika Lake while a minimum value of ~70 m is observed between Karaikkal and Nagapattinam as well as between Krishna and Godavari basins.
(Murthy et al., 1993; Murthy et al., 2012). Regional geophysical investigations on the eastern continental margin of India suggest that the margin can be considered to have a northern rifted segment and a southern transform segment (Subrahmanyam et al., 1999; Chaudhary et al., 2001). The interpretation of the transform margin for the southern segment has also been supported by Krishna et al. (2009), based on the presence of normal faults with major slips, narrow stretch of deformed crust and limited lithospheric extension. Recently, Nemocek et al. (2013) reported the presence of a proto-oceanic crust, representing a transitional zone between the extended continental crust and pure oceanic crust, all along the eastern continental margin of India using deep-penetrated multichannel seismic reflection data. Radhakrishna et al. (2012b) have provided the crustal architecture of the central part of the eastern continental margin of India, based on integrated gravity, seismic reflection and seismic refraction data. They also supported the idea of the presence of a proto-oceanic crust in this part of the continental margin.

The eastern continental margin of India consists of five major onshore-offshore sedimentary basins – the Bengal Basin, Mahanadi Basin, Krishna-Godavari Basin, Palar Basin and the Cauvery Basin (Figure 2) - separated by major tectonic elements (Lal et al., 2009; Bastia and Radhakrishna, 2012; Murthy et al., 2012). The Bengal and Mahanadi basins are separated by a major fault, while the Mahanadi, Krishna-Godavari, Palar and the Cauvery basins are separated from the adjacent basins by the Visakhapatnam, Nuyudupeta and Chingleput highs respectively (Lal et al., 2009). The Bengal Basin is bordered by the Indian shield in the west, Shillong Plateau in the north, the Indo-Burma Ranges in the east and the Bay of Bengal to the south. This basin has regional tectonic trends oriented N-S, NE-SW and ENE-WSW (Khan and Chouhan, 1996). Based on the overall geologic settings, the Bengal Basin has been divided into a Stable Continental Shelf and the Bengal Foredeep, respectively to the west and east of the continental slope (also known as Hinge Zone) (Alam et al., 2003; Uddin and Lundberg, 2004; Hossain et al., 2019). The Mahanadi Basin is located between the Bengal Basin in the north and the Krishna-Godavari Basin to the south. Subrahmanyam et al. (2008) inferred the presence of a series of coast-parallel structural highs and depressions and their shearing patterns in the Mahanadi offshore, using bathymetry, magnetic and gravity data. Based on the interpretation of gravity and magnetic data and the basement configuration derived from the multichannel seismic reflection sections, several researchers (Nayak and Rao, 2002; Subrahmanyam et al., 2008; Bastia et al., 2010; Rao and Radhakrishna, 2014 etc.) have inferred the continuity of the 85°E Ridge into the Mahanadi Basin till the coast at Chilka Lake. The Krishna-Godavari (K-G) Basin is located between Nellore and Kakinada in the central part of the eastern continental margin of India. The NE-SW trending Bapatla Horst and the Tanuku Horst divide this basin into three sub-basins, the Krishna, the West Godavari and the East Godavari. Being a prospective petroliferous basin, several studies have been carried out in the KG Basin to understand the basement configuration and to identify the major structural trends (e.g. Sastri et al., 1973; Kumar, 1983; Mohinuddin et al., 1993; Prabhakar and Zutshi, 1993; Rao, 2001; Bastia, 2006; Gupta, 2006). Rajaram et al. (2000) inferred the presence of NW-SE trending structural features in the offshore regions of the KG Basin, suggesting the continuation of the onshore NW-SE trends offshore, up to the continent-ocean boundary. The Palar Basin is located between the Krishna-Godavari Basin and the Cauvery Basin, with its northern part extending to the offshore. This half crescent shaped basin is bordered in the north by the Nuyudupeta High and in the south by the Chingleput High (Mazumder et al., 2013; Prakash et al., 2018; Mazumder et al., 2019). The Palar Basin is dissected into a northern and a southern part by a broadly E-W oriented subsurface ridge, which is considered to have resulted from a post-rift compression. The southern part of the basin shows a N-S trend, while the northern part trends NE-SW (Mazumder et al., 2013). The Cauvery Basin is located in the southeastern part of the Peninsular India. This basin consists of block-faulted basement comprising of horst-graben architecture oblique to the coast, giving rise to several sub-basins separated by NE-SW trending subsurface basement ridges (Sastri et al., 1981). These sub-basins include the Ariyalur-Pondicherry, Tanjavur-Tranquebar, Nagapattinam and Palk Bay Sub-basins, separated respectively by the Kumbhakonam-Madanam, Karaikall and the Vedaranyam ridges (Sastri et al., 1981; Twinkle et al., 2016; DGH, 2018a). Subsurface mapping carried out in the Cauvery Basin indicate that the Vedaranyam-Tiruchirapalli and the Madurai-Rameswaram trends cut across the general NE-SW trending horst-graben structures of the Cauvery Basin (Twinkle et al., 2016). Consequently, the Cauvery Basin is considered to have undergone three major episodes of tectonism that governed the development of basement structures consisting of three major N-S, NE-SW and NW-SE trending faults during three different phases (Mazumder et al., 2019).

**Deep ocean basins adjacent to the Indian continental margins**

The Indian Ocean consists of several deep basins, some of which have developed at the presently active mid ocean ridges, while the others occurred at now-extinct spreading centres or failed rifts. A detailed understanding on the age of the oceanic crust underlying these ocean basins is essential to decipher the plate tectonic evolution of the Indian Ocean. In this section, I discuss the current status of knowledge of the structure and age of the ocean basins associated with the formation and evolution of the Indian continental margin, following the geomagnetic polarity reversal timescale of Cande and Kent (1995).

**Arabian and Eastern Somali basins**

The Arabian and Eastern Somali basins (Figures 1 and 4) are the conjugate ocean basins formed at the Carlsberg Ridge, separating the Laxmi Ridge from the Seychelles Plateau. The oldest magnetic lineation in these basins corresponds to chron C28n~ (~62.5 Ma), located immediately south of the Laxmi Ridge in the Indian side, and north of the Seychelles Plateau in the African side (Chaubey et al., 2002a). Miles and Roest (1993) reported the presence of oblique offsets, which are the diagnostic features of propagating ridge segments, in both these basins. Later studies (Chaubey et al., 1998; Dyment, 1998) established the existence of oblique offsets in these basins with the help of additional magnetic profiles. Based on an up-to-date compilation of the marine magnetic profiles from these conjugate basins, Chaubey et al. (2002a) carried out a detailed mapping of the magnetic lineations, pseudofaults and fracture zones between chron C28n~ (~62.5 Ma) and C20n (42.54 Ma) from these two conjugate basins. The derived tectonic fabric pattern based on their studies establishes three major episodes of ridge propagation in the Arabian and Eastern Somali basins; the first stage of westward
Laxmi and Gop basins

The Laxmi Basin is the deep offshore region between the NW-SE trending segment of the Laxmi Ridge and the northwestern continental margin of India (Figures 4 and 5). The nature of the crust underlying the Laxmi Basin has been a matter of intense academic debate for many years now. Some authors (e.g. Naini and Talwani, 1982; Kolla and Coumes, 1990; Todal and Eldholm, 1998; Krishna et al., 2006) consider that the Laxmi Basin is underlain by continental crust while others (Biswas and Singh, 1988; Bhattacharya et al., 1994a; Talwani and Reif, 1998; Yatheesh, 2007; Corfield et al., 2010; Eagles and Wibisono, 2013; Siawal et al., 2014; Misra et al., 2015) argue for an oceanic crust. The continental nature of the crust underlying the Laxmi Basin has been inferred on the basis of the semi-continental crustal thickness (e.g. Naini and Talwani, 1982; Krishna et al., 2006), lack of identifiable seafloor spreading type magnetic anomalies (Naini and Talwani, 1982) and extension of Precambrian structural grain into the basin (Kolla and Coumes, 1990). On the other hand, support for its oceanic nature stems from the identification of seafloor spreading-type magnetic anomalies in the basin (Bhattacharya et al., 1994a; Talwani and Reif, 1998; Yatheesh, 2007; Eagles and Wibisono, 2013; Bhattacharya and Yatheesh, 2015), hyperbolic reflection pattern in the multi-channel seismic reflection sections (Biswas and Singh, 1988) and presence of seaward dipping reflectors (SDRs) to the east of the Laxmi Ridge and to the west of the Indian continental margin, suggesting continent-ocean boundaries in between (Corfield et al., 2010; Siawal et al., 2014; Misra et al., 2015). The magnetic anomaly sequence in the Laxmi Basin has been interpreted to represent two-limbed seafloor spreading, identified variously as C33n-C28n-C33n (Bhattacharya et al., 1994a), C33n-C27n-C33n (Yatheesh, 2007), C29n-C28n-C29n (Eagles and Wibisono, 2013) and C33r-C28r-C33r (Ramana et al., 2015).

The Gop Basin is the deep offshore region existing between the E-W trending segment of the Laxmi Ridge and the western India-Pakistan margin (Figures 4 and 5), that was created by the crustal divergence of the Seychelles-Laxmi Ridge continental block from the Saurashtra Volcanic Platform. By and large, this basin has been suggested to be underlain by oceanic crust, based on the presence of well-correlatable magnetic anomalies (Malod et al., 1997; Krishna et al., 2006; Yatheesh, 2007; Collier et al., 2008; Yatheesh et al., 2009), the crustal velocity structure derived from the wide angle seismic reflection data (Collier et al., 2004; Minshull et al., 2008; Collier et al., 2009) and the integrated forward modelling of gravity and magnetic anomalies together with the seismic constraints (Yatheesh, 2007; Yatheesh et al., 2009). Recently however, Rao et al. (2018) have proposed a continental origin for the Gop Basin, based on the derived crustal stretching factor and effective elastic thickness. As in the case of the Laxmi Basin, the magnetic anomaly sequence in the Gop Basin also has been interpreted to represent two-limbed seafloor spreading during chron C28n (~62.5 Ma) to C27n (60.92 Ma), the second stage of eastward propagation during C26r (~57.91-60.92 Ma) to C25n (~55.90 Ma) and the third stage of westward propagation during chron C24r (~53.35-55.90 Ma) to chron C21n (~46.26 Ma).
Figure 4. Tectonic fabric map of the western continental margin of India and the adjacent regions (after Bhattacharya and Yatheesh (2015), with permission from Springer Nature). The continuous black lines, thick dashed lines and the dotted lines represent the magnetic lineations, fracture zones and pseudofaults, respectively. The hachured thick black lines represent postulated boundaries of rift graben basins onland and the magenta lines represent onshore and offshore structural trends. ATTC: Alleppey-Trivandrum Terrace Complex; CKE: Chain-Kairali Escarpment; ABHZ: Axial basement high representing the Panikkar Ridge in the Laxmi Basin; PTR: Palitana Ridge; CG: Cambay Rift Graben; KG: Kutch Rift Graben, NG: Narmada Rift Graben; LCP: Laccadive Plateau; MLR: Maldive Ridge; CMR: Comorin Ridge; PB: Padua Bank; LXB: Laxmi Basin; GPB: Gop Basin; LCB: Laccadive Basin; Sau: Saurashtra; TVM: Trivandrum; MR: Murray Ridge; SVP: Saurashtra Volcanic Platform; BH: Bombay High; Explanation of items of the legend - (a) Continental slivers; (b) Extent of axial basement high zones representing the extinct spreading centres in the Laxmi and Gop basins; (c) anomalous gravity high zone (AGHZ); (d) extents of Deccan Flood Basalts; (e) Seamounts in the Laxmi Basin, R Raman Seamount; P Panikkar Seamount; W Wadia Guyot; (f) Cannanore Rift System.

spreading with varied identifications as C29r-C29n-C29r (Malod et al., 1997), C32n.1r-C31r-C32n.1r (Collier et al., 2008), C29r or C31r (Minshull et al., 2008), C31r-C25r-C31r or C29r-C25r-C29r (Yatheesh et al., 2009) and C33r-C31r-C33r (Ramana et al., 2015).

Recently, Bhattacharya and Yatheesh (2015) evaluated all the existing magnetic anomaly identifications in the Laxmi and Gop basins and concluded that the magnetic anomaly sequence in the Laxmi Basin represents C30n-C25r-C30n sequence and those in the Gop Basin represents C29r-C25r-C29r sequence, with the locations of their extinct spreading centres coinciding with the Panikkar and the Palitana Ridges, respectively (Figure 5). They postulated the existence of a fossil triple junction, referred to as the Gop-Narmada-Laxmi (GNL) Triple Junction located off Saurashtra Peninsula, based on the pattern and age of the magnetic lineations in the two basins. Furthermore, they provided detailed plate reconstructions maps to depict the evolution of the Gop Basin, Laxmi Basin and the Narmada Rift, whose axes of divergence represent the three arms of this triple junction (Figure 5).

**Laccadive Basin**

The Laccadive Basin is the deep ocean basin existing between the southwestern continental margin of India and the Laccadive Plateau (Figure 4), which represent the northernmost segment of the Laccadive-Chagos Ridge. The bathymetric map of the Laccadive Basin shows the presence of several isolated highs located over a continuous subsurface ridge, referred to as the Prathap Ridge (Naini, 1980). This inferred ridge is considered to have been formed either by Réunion Hotspot volcanism (Krishna et al., 1992) or by rift-phase volcanism related to India-Madagascar separation (Subrahmanyam et al., 1995). The nature of crust underlying the Laccadive Basin has not yet been confidently established. Chaubey et al. (2002b) considered this basin to represent a failed rift with volcanism on a stretched continental regime, based on their deductions of rotated fault blocks akin to half graben. Considering the presence of seaward
dipping reflectors (SDRs) west of the Laccadive Plateau and the derived crustal configuration from the forward modelling of the gravity data, Ajay et al. (2010) inferred that the continent-ocean boundary in this region lies west of the Laccadive Plateau implying that the Laccadive Basin is continental in nature. Yatheesh et al. (2013b) demonstrated that the Laccadive Basin region could be explained either in terms of a much thinned continental crust or as anomalously thick oceanic crust, based on forward modelling of the gravity profiles.

**Mannar Basin**

Mannar Basin is the deep offshore region between Sri Lanka and the southeastern part of India (Figure 2). This triangular shaped basin is narrower in the north, widening southwards. Mannar 1-1A and Pearl-1 are the first drill wells located respectively in the Indian and Sri Lankan sectors of the Mannar Basin (Premarathne et al., 2013). Stratigraphic section of the Pearl-1 well shows the existence of ~2.9 km Upper Cretaceous to Recent sediments on an igneous basement. This latter has led to a hypothesis that the Mannar Basin is underlain by oceanic crust. In contrast however, based on interpretations of the seismic data, Baillie et al. (2002) inferred the occurrence of sediments beyond this igneous rock. This was later confirmed with the stratigraphic column of another deep drill well, the Barracuda-1G/1, which penetrated about 760 m thick basaltic rock. This igneous rock is overlain by shale, which continues down to the bottom of the Barracuda well. The age of this volcanics was estimated to be ~76.8 Ma and consequently, the volcanics interbedded in the sediments was considered to be the result of several episodes of volcanism related to the onset of rifting and subsequent northward drift of India (Rana et al., 2008; Premarathne et al., 2013). Subsequently, based on the forward and inverse modelling of the satellite-derived free-air gravity anomalies, Herath et al. (2017) inferred that the Moho is elevated beneath the Mannar Basin with a minimum depth of about 16 km and extending to a depth of about 36 km on either sides of the basin, as a failed rift.

**Central Indian, Crozet and Madagascar basins**

The Central Indian, Crozet and Madagascar basins (Figure 1) constitute the Central Indian Ocean, formed by seafloor spreading at the Southeast Indian and Central Indian ridges. The Rodrigues Triple Junction trace divides the Central Indian Basin into the Western Central and the Eastern Central Indian basins. The Eastern Central Indian and the Crozet basins represent conjugate basins formed at
the Southeast Indian Ridge, while the Western Central Indian and the Madagascar basins are the conjugate basins formed at the Central Indian Ridge. Geomagnetic investigations (Fisher et al., 1971; Sclater and Fisher, 1974; Royer and Schlich, 1988; Kamesh Raju and Ramprasad, 1989; Dyment, 1993; Kamesh Raju, 1993; Krishna et al., 1995; Yatheesh et al., 2008; Cande et al., 2010; Yatheesh et al., 2013a; Cande and Patriat, 2015; Yatheesh et al., 2019) carried out in these basins suggest that the conjugate Eastern Central Indian and Crozet basins contain the magnetic lineation sequence C34n-C1n (83.0 to 0.0 Ma), while the conjugate Western Central Indian and the Madagascar basins contain the magnetic lineation sequence C30n-C1n (~65.58 to 0.0 Ma). The magnetic lineations in these basins are divided into several compartments offset by fracture zones, such as the Vishnu, Northern Astrolabe, Northern Boussole and the 86°E fracture zones (FZ) in the Central Indian Basin; the Mauritius and the Southern Boussole fracture zones in the Madagascar Basin; and the Southern Astrolabe FZ in the Crozet Basin, besides several other unnamed fracture zones. The oceanic crust in the Central Indian Basin has been suggested to be undergoing intraplate deformation resulting from high stresses in the oceanic lithosphere due to the plate boundary configuration and the ongoing collision of India with Eurasia (Royer and Gordon, 1997; Deplus et al., 1998; Krishna et al., 2001). This deformation is considered to have initiated at ~11.0 Ma (Royer et al., 1997).

**Bay of Bengal and Enderby Basin**

The Bay of Bengal is bordered by the eastern continental margin of India and Bangladesh in the west/north and the Sunda-Java Trench to the east (Figure 1). Ramana et al. (1994) identified ~NNW-SSE trending Mesozoic magnetic anomaly sequence M11-M0 (132.5 to 118 Ma) in the Bay of Bengal. Subsequently, Ramana et al. (2001) also identified M11-M0 sequence of magnetic anomalies in the Enderby Basin and inferred that both these basins represent conjugate basins formed by seafloor spreading between India and Antarctica during Early Cretaceous. Following this, Desa et al. (2006) identified magnetic anomaly sequence M11-M0 (132.5 to 118 Ma) in the region between Comorin Ridge and the 85°E Ridge, south off Sri Lanka. Subsequently, Gai et al. (2007) identified both the flanks of oceanic crust with magnetic anomaly sequence M9n to M2n (~130.2 to 124.1 Ma) in the Enderby Basin, establishing the existence of an extinct spreading axis within the Enderby Basin. Krishna et al. (2009) revisited the magnetic anomaly identifications in the Bay of Bengal and Enderby Basin and inferred that the oceanic crust in the Bay of Bengal has been generated during the Cretaceous long normal superchron. Subsequently, Radhakrishna et al. (2012a) inferred the presence of oceanic crust younger to M4 in southern Bay of Bengal and younger to M2 in the northern Bay of Bengal along the eastern continental margin of India. Gibbons et al. (2013) revisited the magnetic anomaly identifications in the Eastern Enderby Basin and revised the identification of magnetic anomalies to represent M4-M0 (~126.7–120.4 Ma) sequence. Talwani et al. (2016) identified the conjugate magnetic anomaly sequences M12n-M0n (132.0 to 120.4 Ma) in the Western Basin of the Bay of Bengal and the Western Enderby Basin. According to them, with the arrival of the Kerguelen Hotspot at around M0 time, reorganization of spreading occurred and a new spreading axis opened at or close to the line joining the Rajmahal and Sylhet traps. Based on the existence of the prominent magnetic anomaly doublet connecting the Rajmahal and Sylhet traps, combined with the inferred location of the seaward dipping reflectors, they defined the new line along which seafloor spreading initiated after 118 Ma, forming the oceanic crust underlying Bangladesh and the Eastern Basin of the Bay of Bengal. The seaward dipping reflectors representing the continent-ocean boundary lies onshore Bangladesh due to the later deposition of enormous sediments derived from Himalayan orogeny.

**Mascarene Basin**

The Mascarene Basin (Figure 6) resulted from the breakup of India-Seychelles-Madagascar continental block and subsequent plate tectonic evolution. This basin can be divided into Southern and
Northern Mascarene basin. Whereas the Southern Mascarene Basin formed by the divergence between Madagascar and India, the Northern Mascarene Basin formed from the divergence between Madagascar and Seychelles (Bernard and Munschy, 2000; Bhattacharya and Yatheesh, 2015; Shuhail et al., 2018). The southern and northern domains of the Mascarene Basin contain the magnetic anomaly sequence C34ny-C27ny-C34ny (83.0 to 60.92 Ma) and C34ny-C30no-C34ny (83.0 to ~67.61 Ma), respectively (Dymont, 1991; Bernard and Munschy, 2000; Eagles and Wibisono, 2013; Bhattacharya and Yatheesh, 2015; Shuhail et al., 2018). The nearly NW-SE trending magnetic lineations in both these domains are offset by several nearly NE-SW trending long fracture zones. In the Southern Mascarene Basin, the entire sequence of magnetic lineations C34ny-C27ny have been mapped on both the flanks of the oceanic crust. In contrast however, in the Northern Mascarene Basin, the magnetic anomaly sequence C34ny-C30no have been mapped only on the southern flank (the Madagascar side). The two major fracture zones identified in the Mascarene Basin are the Mahanoro FZ and Mauritius FZ (Figure 6), both of which are important features in constraining the India-Madagascar initial breakup and subsequent evolution of the Western and Central sectors of the Indian Ocean.

**Wharton Basin**

The Wharton Basin (Figure 1) was created by seafloor spreading at the Indian-Australian plate boundary. Several workers (e.g. Sclater and Fisher, 1974; Schlich, 1982; Liu et al., 1983; Krishna et al., 2012; Jacob et al., 2014) have studied and mapped the magnetic lineations and the associated tectonic fabric in the Wharton Basin. These magnetic lineation patterns indicate that the basin consists of both the flanks of oceanic crust containing the magnetic anomaly sequence C34ny-C20ny (83.0 to 42.54 Ma), representing thereby, an extinct spreading centre, however, some part of the oldest oceanic crust in the northern flank has already been subducted through the Sunda-Java Trench. These magnetic lineations are offset by ~N-S trending long fracture zones. The cessation of spreading in the basin is considered to have occurred at around 36.5 Ma (Jacob et al., 2014).

**Aseismic ridges/submarine plateaus**

The Indian Ocean consists of several aseismic ridges/submarine plateaus (Figure 1), most of which are generally considered to have been formed by hotspot volcanism. The features in the Western Indian Ocean comprise the Laxmi Ridge, Laccadive-Chagos Ridge, Comorin Ridge, Seychelles-Mascarene Plateau and the Madagascar Ridge, while those in the Eastern Indian Ocean include the 85°E Ridge and the Ninetyeast Ridge.

**Laxmi Ridge**

The Laxmi Ridge is an aseismic basement high located off the northwestern continental margin of India (Figures 4 and 5). This feature, which is characterized by a gravity low signature, consists of a ~NW-SE and an E-W trending segment bordering the seaward side of the Laxmi and Gop basins, respectively. There is a broad agreement among the researchers that the Laxmi Ridge is underlain by continental crust (Naini and Talwani, 1982; Bhattacharya et al., 1994a; Talwani and Reif, 1998; Collier et al., 2004; Krishna et al., 2006; Yatheesh, 2007; Collier et al., 2008; Minshull et al., 2008; Collier et al., 2009; Bhattacharya and Yatheesh, 2015; Nair et al., 2015; Mishra et al., 2018). This inference is primarily based on the crustal velocity-depth structure (Naini and Talwani, 1982; Collier et al., 2004; Minshull et al., 2008; Collier et al., 2009), crustal configuration based on seismic-constrained gravity modeling (Krishna et al., 2006; Yatheesh, 2007; Yatheesh et al., 2009; Nair et al., 2015), admittance analysis (Mishra et al., 2018) and the presence of seaward dipping reflectors (SDRs) on both the landward and seaward sides of the Ridge (Corfield et al., 2010; Calvés et al., 2011; Siwal et al., 2014). Todal and Eldholm (1998) however, inferred that the Laxmi Ridge is a marginal high complex, comprising both continental and oceanic crusts. Based on the high-resolution and deep penetrated seismic reflection data, Misra et al. (2015) recently concluded that the Laxmi Ridge is composed of oceanic crust formed at an extinct spreading centre.

**Laccadive-Chagos Ridge**

The Laccadive-Chagos Ridge is an ~2500 km long aseismic bathymetric high (Figure 1) located between 12°S and 14°N latitudes in the Western Indian Ocean. This slightly arcuate elongated feature is divisible into three main segments, the Laccadive Plateau, Maldives Ridge and the Chagos Bank, separated by relatively deep saddle-like features (Bhattacharya and Chaubey, 2001). Hypotheses on the origin of the Laccadive-Chagos Ridge range from a leaky transform fault (Fisher et al., 1971; Sclater and Fisher, 1974), to a hotspot trail (Dietz and Holden, 1970; Whitmarsh, 1974), composite structural elements of various origins (Avraham and Bunce, 1977), and to a product of crack propagation (Sheth, 2005). Though the hotspot hypothesis appears to have a broader acceptance for the genesis of the Laccadive-Chagos Ridge as a whole, several evidences point out the possibility of a continental origin for the Laccadive Plateau segment. This inference is mainly based on the estimated crustal thickness from seismic refraction experiments (Naini and Talwani, 1982), presence of clearly identifiable rotated fault blocks that represents extensional tectonic events (Murty et al., 1999), presence of a complex basement structure comprising of single normal faults with half grabens and grabens, referred to as the Cannanore Rift System (Bhattacharya and Yatheesh, 2015; DGH, 2018bb), presence of seaward dipping reflectors (SDRs) west of the Laccadive Plateau (Ajay et al., 2010), crustal configuration derived from the forward modelling of gravity data (Chaubey et al., 2002b; Ajay et al., 2010; Nair et al., 2013) and the admittance analysis of gravity and bathymetry data (Chaubey et al., 2008). Recently, based on the three-dimensional inversion of gravity anomalies, Kunnumal et al. (2018) inferred that the Greater Maldives Ridge, which consists of the Maldives Ridge, deep sea channel and northern limit of the Chagos Bank, is underlain by ~22 km thick crust representing either oceanic crust with magmatic underplating or continental crust.

**Comorin Ridge**

The Comorin Ridge is a ~NNW-SSE trending bathymetric high located southeast of the Alleppey-Trivandrum Terrace Complex (Figures 1 and 4). The topography of this ~500 km long aseismic ridge is defined by 3900 and 2900 m isobaths. The western flank of the Comorin Ridge dips gently, while its eastern side is characterized by a steep scarp. Based on the integrated analysis of bathymetry, gravity and single-channel seismic reflection data, Kahle et al. (1981) considered that the Comorin Ridge was formed on an oceanic crust.
and the eastern edge of the ridge denote the boundary between this oceanic crust and the adjacent rifted continental crust. On the other hand, based on the admittance analysis and forward modelling of gravity data, Sreejith et al. (2008) inferred that the southern part of the ridge was emplaced on the relatively weak oceanic crust, while the northern part was emplaced on a continental crust.

**Seychelles-Mascarene Plateau**

The Seychelles-Mascarene Plateau is a ~2600 km long, arcuate feature consisting of the Amirante Plateau, Seychelles Bank, Seychelles-Saya de Malha Saddle, Mascarene Plateau (comprising of the Saya de Malha Bank, Nazareth Bank, Cargados-Carajos Bank and the Mauritius Island) and the Réunion Island (Figures 1 and 6). Among these, the continental nature of the crust underlying the Seychelles Bank has been established based on the geochronological (Baker and Miller, 1963) and seismic refraction (Plummer and Belle, 1995) investigations. Bhattacharya and Yatheesh (2015) considered that the continental crust underlying the Seychelles Bank extends further southeast to a part of the Seychelles-Saya de Malha saddle. This entire region of the continental crust has been considered by them to represent the Seychelles Plateau continental sliver, which was developed after a series of tectonic events related to India-Laxmi Ridge-Seychelles-Madagascar breakup. On the other hand, the Mascarene Plateau is considered to have been formed by hotspot volcanism, as evidenced by the north to south age progression of the samples obtained from Saya de Malha and the Nazareth Bank (Duncan, 1990). The radiometric results from the ODP 115 leg drill suggest that the Mauritius Island is an eroded volcanic island formed by the Réunion Hotspot activity (Duncan and Hargraves, 1990). Recently, based on the age of zircon xenocrysts recovered from the Mauritian beach sands, Torsvik et al. (2013) suggested that the Mauritius Island and the other segments of the Mascarene Plateau may overlie a Precambrian basement of continental origin, referred to as the Mauritia microcontinent.

**Madagascar Ridge**

The Madagascar Ridge is a N-S trending bathymetry high extending south of the Madagascar Island (Figure 1). This aseismic ridge separates the Mozambique and Madagascar basins and abuts the Southwest Indian Ridge to the South. The Madagascar Ridge consists of a northern and a southern domain, separated at the 31°S latitude (Goslin et al., 1980). The crust underlying the Northern Madagascar Ridge is considered to be anomalous as it has neither purely continental nor oceanic affinity (Bhattacharya and Chaubey, 2001). Many workers (Yatheesh et al., 2006; Yatheesh et al., 2013b; Bhattacharya and Yatheesh, 2015) consider the bathymetric notch defined by 2000 m isobath on the Northern Madagascar Ridge as representing the conjugate of the Alleppey-Trivandrum Terrace Complex located in the southwestern continental margin of India. The velocity-depth structure of the Southern Madagascar Ridge is closely related to that of mean oceanic crust (Goslin et al., 1981).

**85°E Ridge**

The 85°E Ridge is an anomalous basement high feature located in the Bay of Bengal, buried under the sediments in the north and exposed intermittently above the seafloor to the south (Figure 7). The northern part of this aseismic ridge is associated with a gravity low signature with variable widths of 100-180 km and a steeply dipping western flank compared to the eastern flank. Towards north, this ridge continues up to the Mahanadi coast at Chilka Lake (Rao and Radhakrishna, 2014). Towards south, the 85°E Ridge is associated with a positive gravity anomaly, extending till the Afanasy-Nikitin Seamount (Krishna, 2003; Sreejith et al., 2011). Hypotheses about the origin of this ridge range from a hotspot trace (Curray and Manasinghe, 1991; Michael and Krishna, 2011; Sreejith et al., 2011), to the effect of horizontal compressive forces (Ramana et al., 1997; Anand et al., 2009), to a continental origin (Sar et al., 2009), leaky transform fault (Gibbons et al., 2013) and a fracture zone (Talwani et al., 2016). Michael and Krishna (2011) inferred the presence of alternate bands of positive and negative magnetic anomaly signatures over the 85°E Ridge and dated these anomalies to represent 80-55 Ma. Based on the effective elastic thickness estimate and the flexural

![Image](http://example.com/image_url)

**Figure 7.** The satellite-derived free-air gravity anomaly map of the central and eastern sectors of the Indian Ocean, showing the locations, trends and gravity signatures of the 85°E Ridge, Ninetyeast Ridge and the Afanasy-Nikitin Seamount. The thick dashed white line represents the postulated extent of the 85°E Ridge. **WB:** Western Basin; **EB:** Eastern Basin; **CIB:** Central Indian Basin; **ANS:** Afanasy-Nikitin Seamount.
modelling, Rao et al. (2016) inferred that the 85°E Ridge was emplaced in off-ridge environment. Recently, based on the detailed analysis of high resolution and deep-penetration multi-channel seismic reflection sections, Ismaiel et al. (2017) delineated the internal structure of the 85°E Ridge and inferred this feature to represent a volcanic construct built by both subaqueous and multiphase submarine volcanisms.

**Ninetyeast Ridge**

The Ninetyeast Ridge is a ~5600 km long, N-S trending, linear volcanic feature extending from 34°S to 17°N latitudes (Figure 7). South of 10°N, the Ninetyeast Ridge is clearly discernible as a seafloor feature, however, towards north this feature is entirely buried under the thick Bengal Fan sediments. This feature with an average width of ~200 km and a height of ~2000 m from its surroundings along most of its length (Krishna et al., 2012), is associated with a positive residual geoid/gravity anomaly (Sreejith et al., 2013). The Ninetyeast Ridge separates the Central Indian Basin from its west to the Wharton Basin to the east (Sclater and Fisher, 1974). It has been widely accepted that the Kerguelen Hotspot located beneath the Indian Plate resulted in the accretion of the Ninetyeast Ridge during the first major plate reorganization that occurred at around 90 Ma (Duncan, 1978). The basalts recovered from the Ninetyeast Ridge by DSDP and ODP drilling range in age from ~82 to ~37 Ma, with a decrease in age from north to south (Coffin et al., 2002). Based on the comparison of age of the oceanic crust to the east and west of the Ninetyeast Ridge, combined with the radiometric ages at the DSDP and ODP sites, Krishna et al. (2012) inferred that the Ninetyeast Ridge lengthened at a rate twice that of adjacent oceanic crust.

**Seamounts**

Seamounts on the ocean floor are generally assumed to record the passage of the oceanic crust over the loci of magma generation (Bhattacharya and Chaubey, 2001). The various geophysical investigations carried out in the Indian Ocean reveal the presence of several seamounts in the deep offshore regions. However, to date, only a few of them have been systematically mapped to understand their geomorphology and the geophysical characteristics. These seamounts include the Error Seamount, Sagar Kanya Seamount, Raman-Panikkar-Wadia seamount chain and the Afasany-Nikitin Seamount, as well as several other un-named seamounts in the Laccadive and the Central Indian basins. The Error Seamount, located in the northwestern boundary of the Carlsberg Ridge, is about 42 km long and 16 km wide with an elevation of 4000 m (Ramana et al., 1987). The Sagar Kanya Seamount, which has a height of 2464 m and is located in the Arabian Basin about 200 km west of the Laccadive Plateau, is considered to have been formed by Réunion Hotspot volcanism (Bhattacharya and Subrahmanyan, 1991). The Raman Seamount, Panikkar Seamount and the Wadia Guyot are three prominent seamounts (Figure 5) located in the axial part of the Laxmfi Basin, with basal areas and heights ranging from 300 to 1200 km² and 1068 to 2240 m, respectively. These features, which together form a N30°W trending linear seamount chain of about 250 km length, is considered to have been formed by an episode of anomalous volcanism resulting from the interaction of the Réunion Hotspot with the waning spreading centre in the Laxmfi Basin (Bhattacharya et al., 1994b; Bhattacharya and Yatheesh, 2015), coeval with the formation of Ghatkopar-Powai tholeiites on the Indian mainland (Pande et al., 2017). Further south, Bijesh et al. (2018) identified and mapped several bathymetric highs in the Laccadive Basin and the adjacent regions of Laccadive Plateau and southwestern continental margin of India. They classified these features as seamounts, hills, knolls, plateaus and guyots. Similarly, Das et al. (2005) identified several seamounts located in the Central Indian Basin region of latitudes 9°S to 16°S and longitudes 72°E to 80°E.

A major seamount located in the Central Indian Ocean is the Afasany-Nikitin Seamount, which is a ~2000 m high, 400 km long and 150 km wide feature located between 2°S and 6°S latitudes along ~83°E longitude (Krishna, 2003; Krishna et al., 2014). Earlier studies (Curray and Manasinghe, 1991; Krishna, 2003) considered the Afasany-Nikitin Seamount and the 85°E Ridge as representing two continuous parts of the same linear ridge system formed by the same hotspot volcanism; however, recently Krishna et al. (2014) inferred that these features are unrelated and the proximity of the southern end of the 85°E Ridge to the Afasany-Nikitin Seamount is merely coincidental.

**Formation and evolution of the continental margins of India and the adjacent deep ocean basins**

The large-scale plate tectonic evolution models (McKenzie and Sclater, 1971; Norton and Sclater, 1979; Besse and Courtillot, 1988) for the Indian Ocean suggest that the formation and evolution of the continental margins of India can be described in terms of rifting and drifting among India, Antarctica, Sri Lanka, Madagascar and Seychelles. However, recent studies (Yatheesh, 2007; Gibbons et al., 2013; Bhattacharya and Yatheesh, 2015; Shuhail et al., 2018, etc.) have considered various identified microcontinental slivers in the Indian Ocean and have attempted to provide detailed plate tectonic evolution for the different sectors of the Indian Ocean and the adjacent continental margins by including the slivers as well, as intervening microcontinental blocks. Furthermore, considering that the evolution of the western continental margin of India involves the formation and evolution of a triple junction as well, it becomes necessary to consider the Indian continental block as two separate blocks, the Southern Indian Protocontinent (SIP) and the Northern Indian Protocontinent (NIP), with the Saurashtra Volcanic Platform (SVP) forming a part of the NIP. Consequently in this section, an updated concept on the formation and evolution of the continental margins of India and the adjacent ocean basins has been provided in eight stages (Figures 8a-h), comprising the breakup and evolution among the Southern Indian Protocontinent (SIP), Northern Indian Protocontinent (NIP), Antarctica (ANT), Sri Lanka (SLK), Australia (AUS), Madagascar (MAD) and Seychelles (SEY), along with the identified microcontinental blocks of the Elan Bank (ELB), Laxmi Ridge (LAX), Laccadive Plateau (LCP) and the Saurashtra Volcanic Platform (SVP).

(a) Separation of ANT-AUS block from ELB-SIP-NIP-SLK-LAX-LCP-SEY-MAD block:

The separation of ELB-SIP-NIP-SLK-LAX-LCP-SEY-MAD block from ANT-AUS block appears to have initiated sometime around 132 Ma, creating the Western Basin in the Bay of Bengal on the Indian side and its conjugate Western Enderby Basin on the Antarctic
Figure 8. Schematic diagram depicting the major stages of breakup and formation of the continental margins of India and the adjacent ocean basins, based on Bhattacharya and Yatheesh (2015), Shuhail et al. (2018) and Gibbons et al. (2013). (a) ~126.7 Ma; (b) ~115.0 Ma; (c) ~88.0 Ma; (d) ~83.0 Ma; (e) ~68.5 Ma; (f) ~62.5 Ma; (g) ~60.25 Ma; (h) 56.4 Ma. Explanations to the legend are as follows: (1) Major continental blocks, (2) Microcontinents/continental crust that later broke up to form the microcontinents, (3) Ultra-thinned continental crust, (4) Volcanics, (5) Location of the Marion Hotspot, (6) location of the Réunion Hotspot, (7) Rift axis, (8) Ridge axis, (9) Transform fault, (10) Extinct spreading centre/failed rift, (11) Palaeo-transform fault, (12) location of the Gop-Narmada-Laxmi fossil Triple Junction off the Saurashtra Peninsula. SIP: Southern Indian Protocontinent; NIP: Northern Indian Protocontinent; SLK: Sri Lanka. Other details are as in Figures 1 and 4.
side. The constraint to this age comes from the age of the oldest magnetic lineations representing chron M12n (132 Ma) identified from these conjugate basins (Talwani et al., 2016). Subsequently, at around 126.7 Ma (Figure 8a), spreading was initiated between these blocks, creating oceanic crust in the Eastern Enderby Basin since chron M4 (126.7 Ma) and this situation continued till the spreading ceased in the Eastern Enderby Basin, at around 115 Ma (Gibbons et al., 2013). A divergence is believed to have occurred contemporaneously between India and Sri Lanka, creating the Mannar Basin, bringing Sri Lanka into its present position relative to India (Gibbons et al., 2013).

(b) Separation of ANT-ELB and SLK blocks from SIP-NIP-LAX-LCP-SEY-MAD block:

At around 115 Ma, seafloor spreading in the Eastern Enderby Basin and the crustal divergence in the Mannar Basin ceased and the spreading centre in the Eastern Enderby Basin jumped north of the Elan Bank initiating a new spreading centre between Elan Bank and India (Figure 8b). As a result, the Elan Bank and the Eastern Enderby Basin crust got welded with Antarctic Plate and Sri Lanka got welded to the Indian Plate. The seafloor spreading along this new spreading centre continued creating the conjugate flanks of the oceanic crust in the Eastern Basin of Bay of Bengal and the region north of Elan Bank.

(c) Separation of SIP-NIP-SLK-LAX-LCP-SEY block from MAD block:

The breakup of SIP-NIP-SLK-LAX-LCP-SEY block is believed to have been caused by the Marion Hotspot at around 90 Ma, as evidenced by the existence of ~85–92 Ma volcanics on the western side of India (Valsangkar, 1981; Torsvik et al., 2000; Pande et al., 2001; Melluso et al., 2009; Ram Mohan et al., 2016; Sheth et al., 2017) and the eastern side of Madagascar (Storey et al., 1995; Torsvik et al., 2000). At around 88 Ma, a rift was initiated between Madagascar and SIP-NIP-SLK-LAX-LCP-SEY block (Figure 8c) and following this, a strike-slip motion occurred between southeast coast of Madagascar and southwest coast of India, resulting in the formation of the Chain-Kairali Escarpment (Bhatthacharya and Yatheesh, 2015; Shuhal et al., 2018). This motion also resulted in a configuration where the Alleppey-Trivandrum Terrace Complex broke away from Madagascar side forming a bathymetric notch in the Northern Madagascar Ridge. This was followed by seafloor spreading accreting oceanic crust in the Mascarene Basin (Figure 8d).

(d) Formation and evolution of the GNL Triple Junction and the Laccadive Basin:

At ~68.5 Ma, the Gop-Narmada-Laxmi (GNL) Triple Junction was initiated off Saurashtra, with three rift axes. Subsequently this created the Laxmi Basin, Gop Basin and the Narmada Rift Graben (Bhatthacharya and Yatheesh, 2015), dividing India into the Northern Indian Protocontinent and the Southern Indian Protocontinent (Figure 8e). Rift-drift transition occurred in the Laxmi and Gop basins at chron C30n and C29r, respectively. During this period, another rift axis also existed in the region between the southwestern part of India and the Laccadive Plateau, creating the Laccadive Basin. At around 62.5 Ma, a new spreading axis was initiated between Seychelles and the Laxmi Ridge, creating the conjugate Arabian and Eastern Somali basins (Figure 8f). However, the divergence in the Laxmi, Gop, Laccadive and Southern Mascarene basins continued further and sometime around 60.25 Ma, the divergence in the Southern Mascarene and Laccadive basins ceased (Figure 8g). As a result, the Laccadive Plateau and the Laccadive Basin got attached to the Indian Plate, while the Southern Mascarene Basin got welded to Madagascar. Subsequently, at around 56.4 Ma, the axes of divergence in the Gop Basin, Laxmi Basin and the Narmada Rift also became extinct and all these basins got welded to the Indian Plate (Figure 8h). Seafloor spreading continued along the Carlsberg Ridge accreting oceanic crust in the conjugate Arabian and Eastern Somali basins in the Western Indian Ocean.

The above scenario suggests that the development of the northwestern continental margin of India was controlled to a great extent by the evolution of the Laxmi and Gop basins, while the development of the southwestern continental margin was controlled by the evolution of the Laccadive Basin and the easternmost part of the Southern Mascarene Basin. Thus, contrary to the conventional concept, it would appear that the conjugate of northwestern continental margin of India is the eastern flank of the Laxmi Ridge, while the conjugate of southwestern continental margin of India lies partly on the eastern flank of the Laccadive Plateau and partly on the eastern flank of the Northern Madagascar Ridge. This would also imply that the genesis of the structural features identified in the northwestern continental margin of India might be related to the breakup between India and welded Seychelles-Laxmi Ridge block, while those in the southwestern continental margin of India might be related to the breakup among India, Laccadive Plateau and the southeasternmost part of Madagascar.

Some unresolved problems

The breakup of the East Gondwanaland resulted in the formation of the continental margins and the adjacent deep offshore regions of the Indian Ocean. Geoscientific investigations carried out prior to the 1990s have provided a broad understanding on the structure and tectonics of the Indian Ocean as a whole. Subsequently, newer techniques and newly acquired closely-spaced data have built up on the earlier studies, facilitating a better understanding of the history of the Indian Ocean basin and the bordering continental margins. At the same time however, the new information has also thrown up many challenging questions related to the evolutionary history of the Indian Ocean that need to be addressed for a more comprehensive picture to emerge. Some of these questions are detailed in this section.

Existence and extent of the Prathap and Kori-Comorin ridges

The Prathap Ridge and the Kori-Comorin Ridge are the two structural features reported to be present in the western continental margin of India by Naini (1980) and Biswas and Singh (1988), respectively (Figure 9a,b). Naini (1980) identified isolated topographic highs between 2500 m and 3000 m isoliths on the continental slope and rise, between Cochin and Goa, and considered them as continuous sub-bottom highs forming an unbroken structural high between 7°N and 15°N latitudes, which he referred to as the Prathap Ridge. Subsequently, Biswas and Singh (1988) proposed the existence of a very prominent NW-SE trending linear fault-bounded
continuous structural high extending from Kutch shelf to Cape Comorin, termed as the Kori-Comorin Ridge, consisting of the Kori High in the north and the Prathap Ridge in the south. Other workers attempted to refine the extent of the Prathap Ridge with the help of additional regional seismic reflection, gravity and magnetic data and reported that the Prathap ridge is characterized by magnetic and gravity highs discernible on the ship track data. However, the latest available satellite-derived free-air gravity anomalies (Figure 9a) do not suggest the existence of such a prominent linear gravity high of ~1000 km extent coinciding with the postulated location of the Prathap Ridge. Instead, the free-air gravity map shows the existence of several isolated gravity highs representing isolated structural highs. The multibeam bathymetry map of this region (Figure 9b) also suggests the existence of several isolated bathymetry highs, classified to represent seamounts, plateaus, hills, knolls and guyots (Bijesh et al., 2018). It is important to note that the continuity of the sub bottom structural highs representing the existence of Prathap Ridge for a distance of ~1000 km was inferred based on the single channel seismic reflection data along sparsely distributed profiles, especially in the region between 7°N and 13°N latitudes. Therefore, a detailed revisit to the western continental margin of India using the closely spaced and high resolution gravity, magnetic and multichannel seismic reflection data is necessary to evaluate whether the structural highs identified by Naini (1980) represent isolated highs or they form part of a continuous subsurface ridge. This will also help to delineate the extent of the Kori-Comorin Ridge, which has been postulated to extend from the Kutch shelf to Cape Comorin.

**Location of the continent-ocean boundaries around the continental margins of India**

The nature and location of the boundary between the continental and oceanic crust in a passive continental margin setup is important to understand the processes related to rifting and the onset of seafloor spreading at a divergent plate boundary. Delineation of the continent-ocean boundary based only on the geophysical data appears to be
The COB has been inferred to be located west of the Laccadive Plateau, sections. The oceanic crust of the Gop Basin is yet to be imaged from seismic boundary between the continental crust of the Laxmi Ridge and the and Yatheesh, 2015). However, the COB representing the magnetic lineations in the Gop Basin (Malod et al., 1997; Yatheesh, 2007; Collier et al., 2008; Yatheesh et al., 2009; Bhattacharya and Yatheesh, 2015) representing the COB between the Ridge and the oceanic crust of the Gop Basin. Such an inference is further supported by the presence of seafloor spreading lineations in the Laxmi Basin (Bhattacharya et al., 1994a; Yatheesh, 2007; Bhattacharya and Yatheesh, 2015) and the recovery of basalt from the basement retrieved by drilling in the Laxmi Basin during International Ocean Discovery Program Expedition 355 (Pandey et al., 2016).

Recently-published multichannel seismic reflection sections across the E-W trending segment of the Laxmi Ridge, Gop Basin and the Saurashtra Volcanic Platform clearly depict the existence of SDRs south of the Laxmi ridge (Collier et al., 2004; Misra et al., 2015) representing the COB between the Ridge and the oceanic crust in the Arabian Basin; and a separate sequence of SDRs south of the Saurashtra Volcanic Platform (Gaedicke et al., 2002; Calvès et al., 2011) representing the COB between the continental crust of Saurashtra Volcanic Platform and the oceanic crust of the Gop Basin. Such an existence of two COBs in the India/Pakistan continental margin is further supported by the presence of seafloor spreading magnetic lineations in the Gop Basin (Malod et al., 1997; Yatheesh, 2007; Collier et al., 2008; Yatheesh et al., 2009; Bhattacharya and Yatheesh, 2015). However, the COB representing the boundary between the continental crust of the Laxmi Ridge and the oceanic crust of the Gop Basin is yet to be imaged from seismic sections.

In the southern sector of the western continental margin of India, the COB has been inferred to be located west of the Laccadive Plateau, based on the location of SDRs and the nature of crust underlying the Laccadive Basin inferred by forward modelling of gravity anomalies. However, a wide swath of crust representing the Laccadive Basin exist east of this inferred COB. Unlike in the deep offshore regions representing the Laxmi and Gop basins, no seafloor spreading type magnetic lineations have been reported from the Laccadive Basin. However, such a scenario of lack of inferred magnetic lineations cannot be considered as the evidence to the “absence” of such seafloor spreading type magnetic lineations in the Laccadive Basin, since a detailed magnetic investigation using closely spaced magnetic data has not yet been undertaken in the Laccadive Basin, probably due to the unavailability of the adequate data. In addition, no high-resolution and deep penetrated multichannel seismic reflection sections are available in public domain to carry out detailed analysis to derive the crustal configuration and examine the possibility of existence of SDRs within the Laccadive Basin, and to find out their location if this region is underlain by oceanic crust. Moreover, there is no seismic velocity-depth information derived from seismic refraction data available from the Laccadive Basin to infer the detailed crustal structure as well as to use those information as a dependable constraint for the forward modeling of gravity anomalies. Therefore, the current status of the knowledge about the Laccadive Basin warrants a detailed integrated geophysical investigation based on seismic refraction experiment integrated with high resolution seismic reflection, gravity and magnetic data acquired over closely spaced profiles.

The Mannar Basin appears to have been formed by the crustal divergence between southeasternmost part of India and Sri Lanka. Based on the presence of igneous crust under the sediments, this deep ocean basin was originally considered to have underlain by oceanic crust. Followed by this, Baillie et al. (2002) inferred the existence of sediments beyond this igneous crust and this inference was considered against for the oceanic crust hypothesis in the Mannar Basin. Subsequent studies considered that the Mannar Basin is underlain by thinned continental crust with elevated Moho, representing a failed rift dominated with igneous activities causing emplacement of volcanics interbedded in the sediments. So far, no seafloor spreading magnetic lineations were reported in the Mannar Basin. However, one cannot negate the possibility of existence of oceanic crust overlain by thick pile of sediments in the Mannar Basin, based on the above observations. Therefore, as in the case of Laccadive Basin, a revisit to the Mannar Basin using the high-resolution seismic reflection and refraction, gravity and magnetic data will be useful to derive crustal configuration and to obtain the most reasonable explanation for the nature of crust underlying the Mannar Basin. If in case the Mannar Basin is underlain by oceanic crust, such a scenario will necessitate defining continent ocean boundary in the southeastern part of India and the western part of Sri Lanka.

**Timing of formation of the Laxmi, Gop, Laccadive and Mannar basins**

Timing of formation of ocean basins created at various divergent axes in the Indian Ocean are important to understand the spatio-temporal evolution of the Indian Ocean as a whole. These timings are well established for the ocean basins located far away from the continental margins, however, there exists several smaller ocean basins closer to the continental margins of India, whose nature of the underlying crust as well as the timing of formation is not well established. Some of these basins were formed by rifting only.
representing a failed rift, while others were formed by rifting followed by spreading representing extinct spreading centres. There are no direct methods to understand the timing of opening of an ocean basin formed by rifting alone using the geophysical data. However, the linear magnetic anomalies imprinted on the oceanic crust formed at a now-extinct spreading centre can provide important constraints on the age of the underlying oceanic crust and thereby the timing of opening and cessation of seafloor spreading in those ocean basins.

The Laxmi and Gop basins contain linear magnetic anomalies representing seafloor spreading between India/Pakistan margin and the Seychelles-Laxmi Ridge block. Correlation of the linear magnetic anomalies with geomagnetic timescale provides unique age estimate of the oceanic crust in the wide ocean basins containing larger linear magnetic anomaly sequences. But the magnetic lineations in the Gop and Laxmi basins are short and therefore a unique identification appears to be difficult. As a result, researchers proposed different age estimates ranging between chron C33n and C29n for the opening and chron C31r to C25r for cessation of the Gop and Laxmi basins. In case of the Laccadive and Mannar basins, no seafloor spreading type linear magnetic anomalies were reported so far and therefore no age constraints are available from geophysical data to define the timing of opening and cessation of divergence in these basins. As rightly pointed out by Bhattacharya and Yatheesh (2015), until the unambiguous age constraints are obtained from the Gop, Laxmi, Laccadive and Mannar basins, the plate tectonic evolution of the continental margins of India and the adjacent deep offshore regions cannot be established conclusively in proper temporal framework.

**Crustal structure, genesis and trend of the 85°E Ridge**

The 85°E Ridge is an important structural feature, represented as a basement high buried under the sediments, in the Bay of Bengal. This aseismic ridge has got attention of several researchers in the recent past and those studies (Bastia et al., 2010; Michael and Krishna, 2011; Sreejith et al., 2011; Rao and Radhakrishna, 2014; Ismaiel et al., 2017) provided important insights about the geophysical signatures, internal structure, extent and probable genesis of the 85°E Ridge. The extent of this anomalous basement high feature has been considered to start from Mahanadi Basin in the north to the Afanasy-Nikitin Seamount in the south. Several authors considered the genesis of the entire length of this basement high to represent volcanic trace from the same hotspot, while a recent study (Krishna et al., 2014) considered that the 85°E Ridge and the Afanasy-Nikitin Seamount are unrelated features arose from different mantle sources and the proximity of the southern end of the 85°E Ridge to the Afanasy-Nikitin Seamount is coincidental. The extent of the 85°E Ridge north of 6°N and south of 2°N trends in nearly N-S direction, but it shows a major curvature between 2°N and 6°N. If the 85°E Ridge is formed by the hotspot volcanism, such a change in direction implies a major reorientation of Indian plate motion during the time of formation of this segment of the 85°E Ridge. However, such a motion of the Indian plate has not been recorded in any other ocean basins / aseismic ridges in the Indian Ocean and therefore this observation warrants a satisfactory answer for considering a hotspot origin for the entire length of the 85°E Ridge.

Another important aspect of the 85°E Ridge is to understand the deep crustal architecture of the 85°E Ridge and the adjacent ocean basins. Ismaiel et al. (2017) derived the internal structure of the 85°E Ridge by identifying the volcanic successions and inferred that the 85°E Ridge is a volcanic construct built by both subaqueous and multiphase submarine volcanisms. Though the internal structure has been mapped in detail, the deeper structure of the 85°E Ridge awaits to be firmly established. A detailed seismic refraction experiment over this feature using the Ocean Bottom Seismometers will provide important constraints for such an understanding of the deeper crust over which the 85°E Ridge volcanics appears to have emplaced. In addition, the velocity-depth information derived from those seismic experiments can also evaluate the concept of existence of high density metasediments on either side of the 85°E Ridge and a broader Moho depression that are considered to cause the characteristic gravity low associated with this enigmatic feature.

**Establishing the nature and extent of the postulated microcontinental slivers in the Indian Ocean**

The Indian Ocean contains large number of aseismic ridges and submarine plateaus, whose origin are broadly attributed to the hotspot volcanism. However, the detailed investigations (Naini and Talwani, 1982; Plummer and Belle, 1995; Talwani and Reif, 1998; Nicolaysen et al., 2001; Borissova et al., 2003; Collier et al., 2009; Calvès et al., 2011; Bhattacharya and Yatheesh, 2015) carried out over those features reported strong evidences to consider some of those features, in part or full, to represent microcontinents. The broadly agreed major postulated microcontinents existing in the Indian Ocean are the Laxmi Ridge, Laccadive Plateau, Saurashtra Volcanic Platform, Seychelles Plateau, Northern Madagascar Ridge and the Elan Bank. In a recent study, Torsvik et al. (2013) proposed the existence of a Precambrian microcontinent, namely “Mauritia”, which mainly consists of Mauritius Island and the continental fragments from the Southern Mascarene Plateau (e.g. parts of Saya de Malha, Nazareth and Cargados-Carajos banks) and the Laccadive, Maldives and Chagos areas adjacent to Indian margin. Subsequently, Bhattacharya and Yatheesh (2015) critically examined all the available information on the postulated microcontinents and attempted to provide a plate tectonic evolution model for the Western Indian Ocean by considering only those microcontinents whose continental affinity was better-constrained. In their model, they accommodated the postulated continental blocks of the Laxmi Ridge, Seychelles Plateau, Laccadive Plateau and the Saurashtra Volcanic Platform as intervening continental slivers between India and Madagascar in their immediate pre-drift tectonic scenario. This model strongly suggests the existence of the above microcontinents, however, their crustal structure and the extent of the continental crust are yet to be confidently established. To decipher the plate tectonic evolution of the Indian Ocean, it is important to establish their continental nature and delineate the respective continent-ocean boundaries of all these postulated microcontinents since those COBs are the most representative boundary to obtain accurate plate reconstruction models. Therefore, the current status of the knowledge about the nature and extent of various postulated microcontinents necessitate confident establishment of their continental nature and the delineation of their extent by using high-resolution and deep penetrated multichannel seismic reflection data collected along closely spaced profiles, integrated with seismic refraction data, acquired using Ocean Bottom Seismometers.
The intense intraplate seismicity over Chagos Bank, a segment of an “aseismic” ridge

The Chagos Bank represents the southernmost segment of the Laccadive-Chagos Ridge, which is considered as an aseismic ridge formed by Réunion Hotspot volcanism in the past. The drilling at ODP Site 713, located in the northern part of the Chagos Bank, reached basaltic basement that yielded an age of 49 Ma (Duncan and Hargraves, 1990). The plate tectonic evolution models (McKenzie and Selater, 1971; Norton and Selater, 1979; Patriat and Segoufin, 1988) suggest that the Chagos and Saya de Malha banks were created contemporaneously by hotspot volcanism since ~ 49Ma, when the Réunion Hotspot lay beneath or close to the spreading axis of the Central Indian Ridge. Subsequently, at ~35 Ma, the Chagos Bank broke away from the Mascarene Plateau and a reorganization of the spreading centre occurred resulting in transferring of the Réunion Hotspot to the African Plate and forming the younger parts of the Mascarene Plateau. This tectonic framework clearly shows that the Chagos Bank qualifies to be considered as a part of the aseismic ridge formed by past hotspot volcanism that is usually a region devoid of seismic activities. However, Stein (1978) reported an unusual, isolated earthquake swarm consisting of 16 events occurred during 1965-1970, over the steep west-facing scarp of the Chagos Bank. Based on the study of body and surface wave data of the three largest earthquakes, Stein (1978) opined that the swarm might have occurred at depth on a cross fracture remaining from the breakup of the Chagos Bank from the Mascarene Plateau and the formation of the present Central Indian Ridge. Based on the moment variance technique applied to the intraplate earthquakes occurred over the Chagos Bank, Wiens (1986) inferred a regional tectonic deformation extending northeast of Chagos Bank consisting of thrust, normal and strike-slip events, representing repeated reactivation of weak zones (presumably fracture zones) by a regional stress regime characterized by N-S extension with E-W and vertical compression. Based on the moment release computation using the detailed catalog of near-ridge earthquakes occurred between 1912 and 1993, Radhakrishna et al. (1998) suggested that the Chagos Bank seismicity forms part of the plate-wide stress distribution in the Central Indian Ocean. The earthquake data catalogue of International Seismological Centre (http://www.isc.ac.uk/isbulletin/search/catalogue) lists a total number of 209 earthquakes with magnitudes ranging from 3.4 to 6.7, occurred between 1965 and 2018, showing the continuing seismicity over the Chagos Bank. A passive seismic experiment using Ocean Bottom seismometers in this region would provide detailed velocity-depth constraints to derive the crustal structure of the Chagos Bank and the adjoining regions and thereby provide new insights on the cause of intraplate seismicity over the Chagos Bank segment of the Laccadive-Chagos Ridge, which is considered as an aseismic ridge formed by past hotspot volcanism.

Ways to go forward

The marine geophysical investigations carried out in the Indian Ocean for the last four decades resolved several first order problems associated with structure and tectonics of the continental margins of India and the adjacent deep ocean basins. Most of these problems were addressed with the available sea-surface bathymetry, gravity, magnetic and the limited seismic reflection data, integrated with the sonobuoy refraction results of Naini (1980). Though considerable advancement in the understanding of these regions were achieved, there remains several important problems unresolved, which cannot be addressed solely with the existing sparse and conventional geophysical data.

Most of the unresolved problems described in the previous section can be addressed confidently by retrieving basement samples by deep sea drilling, however, it is not possible to ground truth everywhere in the ocean basins as it is too expensive. In absence of such ground truth data, it necessitates acquisition of high-resolution and deep penetrated multichannel seismic reflection data along closely spaced profiles as well as seismic refraction data using Ocean Bottom Seismometers, integrated with closely spaced gravity, magnetic and multi-beam bathymetry data, that can provide important constraints to bridge those knowledge gaps. Acquisition and interpretation of the seismic refraction data over the regions such as the Laccadive Basin, 85°E Ridge, Mannar Basin, Chagos Bank and the several postulated microcontinents, can provide detailed velocity-depth information that can be used to infer the nature of crust underlying those regions and their detailed crustal structure. Similarly, acquisition and interpretation of high-resolution and deep penetrated multichannel seismic reflection data in the western continental margin of India and the adjacent ocean basins can be used to identify the possible seaward dipping reflectors and thereby delineate the continent-ocean boundary around Indian continental margins as well as the postulated microcontinents. These multichannel seismic reflection data along closely spaced profiles can also be used to identify the structures considered to represent the Prathap Ridge, examine the continuity of this feature from profile to profile and to map its linear as well as lateral extents. A passive seismic experiment using Ocean Bottom Seismometers conducted over the Chagos Bank and the adjoining regions can provide important insights on the cause of the recurring intraplate seismicity in this region. Acquisition and interpretation of magnetic data along closely spaced profiles in the Laccadive and Mannar basins will provide an opportunity to examine the possibility of existence of seafloor spreading type linear magnetic anomalies in these ocean basins. If any such linear magnetic anomalies are identified, correlation of them with the geomagnetic polarity reversal timescale can provide age constraints on the formation and cessation of these ocean basin. As explained earlier, the sea-surface magnetic data do not allow to derive a unique age corresponding to the magnetic anomalies observed in the Gop and Laxmi basins. An alternate approach for getting a better age constraint from these basins perhaps would be to use high resolution magnetic data collected by deep tow magnetic system that can provide characteristic signatures of the magnetic anomalies to aid their identification.

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