Squeezed States
and Helmholtz Spectra

Francisco Delgado C. *, Bogdan Mielnik †

and

Marco A. Reyes. ‡

Depto. de Física, CINVESTAV,
A.P. 14-740, 07000, México D.F., México.

Abstract

The 'classical interpretation' of the wave function $\psi(x)$ reveals an interesting operational aspect of the Helmholtz spectra. It is shown that the traditional Sturm-Liouville problem contains the simplest key to predict the squeezing effect for charged particle states.

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*Depto. de Matemáticas, ITESM, Campus Estado de México.
A.P. 2, 52926, Atizapán, Edo. de Mex., México.
†Institute of Theoretical Physics, Warsaw University.
ul. Hoza 69, 00-681, Warsaw, Poland.
‡Instituto de Física de la Universidad de Guanajuato, 37150, Guanajuato, México.
1.- Introduction.

One of the most transparent pictures of the Schrödinger’s eigenvalue equation in 1-dimension is obtained by reinterpreting the space coordinate $x$ as the "time" ($x = t$), and the values of the wave function $\psi(x)$ and its derivative $\psi'(x)$ as the position and momentum of a classical point particle. The resulting classical trajectory obeys Hamilton equations which can be further simplified by introducing the Prüfer’s angle on the Poincare plane. The integration of the 1-st order "angular equation" then turns out one of the most efficient methods to determine the energy spectra. In this note we show that the 'classical algorithm' works even better for the Helmholtz problem; in addition, it exhibits a new aspect of the Helmholtz spectra.

2.- The Helmholtz spectrum.

Given an interval $[a, b] \subset \mathbb{R}$ and a real function $\phi : [a, b] \to \mathbb{R}$, the Helmholtz equation is a particular case of a Sturm-Liouville problem:

$$\frac{d^2 \psi}{dx^2} + \lambda \phi(x) \psi(x) = 0 \quad (2.1)$$

and the Helmholtz spectrum is defined as the set of all $\lambda \in \mathbb{R}$ for which (2.1) has a non trivial solution $\psi$ vanishing on both ends:

$$\psi(a) = \psi(b) = 0 \quad (2.2)$$

The 'trajectory picture' applies quite naturally to the Helmholtz eigenproblem, and the spectral condition is much simpler than in the Schrödinger’s case. Let $\psi(x)$ be any real solution of (2.1). Define: $x = t$, $\psi = q$, $\psi' = p$; eq. (2.1) then becomes:

$$\frac{dq}{dt} = p \quad \frac{dp}{dt} = -\lambda \phi(t)q \quad (2.3)$$

with the 'spectral condition':

$$q(a) = q(b) = 0 \quad (2.4)$$

Observe, that (2.3) are the canonical equations for a classical point-particle with the Hamiltonian:

$$H(t) = \frac{p^2}{2} + \lambda \phi(t)\frac{q^2}{2} \quad (2.5)$$
Each integral trajectory of (2.3) 'portraits' a wave function $\psi(x)$ and its first derivative. We shall further assume that $0 \neq \phi(t) \geq 0$ is a piece-wise continuous function in $[a, b]$. The attractive oscillator force in (2.3) then drives the classical trajectory clockwise around the origin of the phase plane (Fig. 1). When $\lambda$ grows to $+\infty$, the end point $(q(b), p(b))$ must cross the $p-$axis an infinite number of times (Fig. 2). Each time this happens, $\lambda$ belongs to the point spectrum of (2.1). In this way the Helmholtz eigenproblem (2.1) has a discrete spectrum in the form of an infinite ladder of positive eigenvalues determined by the classical trajectories (compare the theorems [9, 10]).

The calculation of the spectral values can be carried out by introducing the polar variables on the phase plane: $p = \rho \cos \alpha$, $q = \rho \sin \alpha$. The variable $\alpha$ then separates and (2.3) reads

\[
\begin{align*}
\dot{\alpha} &= \cos^2 \alpha + \lambda \phi(t) \sin^2 \alpha \\
\dot{\rho} \rho &= 1 - \frac{\lambda \phi(t)}{2} \sin 2\alpha
\end{align*}
\] (2.6)

with the initial and spectral conditions:

\begin{align*}
\text{initial condition} & \quad \alpha(0) = 0 \\
\text{spectral condition} & \quad \Delta \alpha = \alpha(b) = n\pi \quad (n = 0, \pm 1, \pm 2, \ldots) (2.9)
\end{align*}

Some exact solutions of (2.1-2) were recently studied by Goyal et al to check the validity of the WKB and Langer approximations [7]. For curiosity, we applied the angular method to the same collection of cases. Eq. (2.6) was integrated numerically for a net of $\lambda$-values by applying the Runge-Kutta method and the condition (2.9) selected the eigenvalues $\lambda$ quoted in Table 1.

| $\phi(t)$ | $[a, b]$ | Eigenvalues $\lambda$ | Amplification constant $\sigma$ |
|-----------|----------|----------------------|-----------------------------|
| $(t + \pi)^4$ | $[0, \pi]$ | Exact | Approximate | Angular |
| 0.00174401 | 0.00188559 (WKB) | 0.00174401 | -0.543046 |
| 0.00734865 | 0.00754235 (WKB) | 0.00734843 | 0.051828 |
| 0.0167524 | 0.01697035 (WKB) | 0.0167517 | -0.510087 |
| $1.1e^t - 1$ | $[0, 1]$ | Exact | Approximate | Angular |
| 11.145735 | 11.145736 (LAN) | 11.145735 | -0.591942 |
| 47.301484 | 47.301493 (LAN) | 43.301493 | 0.54752 |
| 108.428802 | 108.428616 (LAN) | 108.428629 | -0.52727 |
| $(1 + \sin 2\pi t)^2$ | $[0, 1]$ | Exact | Approximate | Angular |
| 5.3347146 | 34.1068933 | 3.1455 | -3.42553 |
| 86.8947093 | 86.8947093 | 86.8947093 | -3.42553 |

Table 1: Helmholtz eigenvalues known in the literature compared with the results of the angular method. In addition we have examined $\phi(t) = \lambda(1 +$
The last column gives the 'optical information', disregarded in the literature.

As one can see, for the low eigenvalues the ranking is won by the almost ignored angular algorithm (both, in number of digits and simplicity of operation). Its additional advantage is that it uncovers the links of the Helmholtz spectrum with the techniques of generating the squeezed states.

3. The trajectory optics.

In a typical Sturm-Liouville problem, one looks for a solution of a differential equation which fulfills a definite boundary condition. The second independent solution is dismissed as unphysical. However, we shall show that for (2.1) the second solution too carries an important message. As eqs. (2.3) are linear, the corresponding \( q(t), p(t) \) depend linearly on the initial values \( q(a), p(a) \):

\[
\begin{pmatrix}
q(t) \\
p(t)
\end{pmatrix}
= u(t, a)
\begin{pmatrix}
q(a) \\
p(a)
\end{pmatrix}
\quad (3.1)
\]

where \( u(t, a) \) is the simplectic evolution matrix (also: the transfer matrix). The existence of a non-trivial solution of (2.2) with \( q(a) = q(b) = 0 \) means that the evolution matrix \( u(a, b) \) in the interval \([a, b]\) maps:

\[
\begin{pmatrix}
0 \\
p(a)
\end{pmatrix}
\rightarrow
\begin{pmatrix}
0 \\
p(b)
\end{pmatrix}
\quad (3.2)
\]

and since \( u(a, b) \) is simplectic, it must be of the form:

\[
u(b, a) = \begin{pmatrix}
\sigma & 0 \\
\eta & 1/\sigma
\end{pmatrix} ; \quad \sigma, \eta \in \mathbb{R}, \sigma \neq 0
\quad (3.3)
\]

Once this is known, one can deduce not only the spectral orbit of (2.3), but also the form of other trajectories. Consider therefore the congruence of trajectories which diverge at \( t = a \) from the initial point \((q_0, p)\), where \( q_0 \) is fixed and \( p \) varies. If \( \lambda \) is an eigenvalue of (2.1) and \( u(b, a) \) is of the form (3.3), then at \( t = b \):

\[
\begin{pmatrix}
q(b) \\
p(b)
\end{pmatrix}
= \begin{pmatrix}
\sigma & 0 \\
\eta & 1/\sigma
\end{pmatrix}
\begin{pmatrix}
q_0 \\
p
\end{pmatrix}
= \begin{pmatrix}
\sigma q_0 \\
p'
\end{pmatrix}
\quad (3.4)
\]

i.e., all orbits departing from the same \( q_0 \) (with different momenta \( p \)) meet at \( \sigma q_0 \) (with momenta \( p' = \sigma q_0 + p/\sigma \)). Thus, the eigenvalues \( \lambda \) of (2.1) are the real numbers for which the time evolution (2.3) focuses any congruence of classical orbits diverging from any common initial \( q_0 \). The optical constants \( \sigma, \eta \) can be read from the second (disregarded) solution of (2.1).
Note that the focusing, in general, is not the scale transformation \([12]\), (as the matrix (3.3) is triangular), though it is an example of squeezing \([13, 14, 15]\). The scale transformation \((\eta = 0)\) is an exceptional effect which can happen only for some \(\phi\)'s (see \([18]\)), whereas the focusing is a typical phenomenon which must occur for any positive \(\phi(x)\), on an infinite ladder of \(\lambda\)'s. The converse problem is as well interesting. Instead of fixing an interval and looking for \(\lambda\), fix any \(\lambda > 0\) and \(\beta(t) \geq 0\) and consider an expanding time interval \([0, \tau]\). If \(\phi(t)\) maintains some positive average in \([0, \infty)\), the final point \((q(\tau), p(\tau))\), as \(\tau \to \infty\) must cross infinity of times the \(p\)-axis, each time generating our focusing effect no matter the exact form of \(\beta(t)\). Should \(\beta(t)\) be a random variable performing errating oscillations, the motions (2.3) will be chaotic too; yet, for any congruence of trajectories, a clean 'echo effect' of the initial configuration (an amplified or reduced image) must emerge now and then against the chaotic forces. Note also that the effect must have a quantum equivalent.

4.- The Quantum Images.

Consider the Hamiltonian \(H(t)\) in form (2.5), where \(q\) and \(p\) are position and momentum operators. Since (2.5) is quadratic, the quantum evolution follows faithfully the classical motions. The techniques of the classical "piloting trajectory" gives specially good results for the Gaussian wave packets \([15, 16]\). We shall show, that it provides also an explicit expression for the evolution operator \(U(t, a)\):

\[
\frac{dU(t, a)}{dt} = -iH(t)U(t, a); \quad U(a, a) = 1 \tag{4.1}
\]

With this end define:

\[
W(t) = e^{i\alpha(t)H_0}U(t, a) \tag{4.2}
\]

where \(H_0 = \frac{1}{2}(p^2 + q^2)\) is the traditional oscillator Hamiltonian and \(\alpha(t)\) is a real, differentiable function. Due to the well known transformation rules:

\[
e^{iaH_0}qe^{-iaH_0} = q \cos \alpha + p \sin \alpha, \quad e^{iaH_0}pe^{-iaH_0} = p \cos \alpha - q \sin \alpha,
\]

the operator \(W(t)\) fulfills:

\[
\frac{dW}{dt} = i[(\dot{\alpha} - \cos^2 \alpha - \lambda \phi(t) \sin^2 \alpha)\frac{p^2}{2} + \]

\[
+ (\dot{\alpha} - \sin^2 \alpha - \lambda \phi(t) \cos^2 \alpha)\frac{q^2}{2} + \]

\[
+ (1 - \lambda \phi(t)) \sin \alpha \cos \alpha \frac{qp + pq}{2}]W(t) \tag{4.3}
\]

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Notice, that the time dependent term coupled to \( p^2/2 \) corresponds exactly to the Prüfer equation. Thus, if \( \alpha(t) \) is the Prüfer angle (2.6), then (4.3) reduces to:

\[
\frac{dW}{dt} = i\left(1 - \phi(t)\right) \sin 2\alpha \frac{qp + pq}{2} + (1 - \lambda \phi(t)) \cos 2\alpha \frac{q^2}{2} W(t) \tag{4.4}
\]

and integrates explicitly in terms of the Prüfer’s radius \( \rho \):

\[
W(t) = e^{i \ln(\frac{\rho(t)}{\rho(a)}) \frac{qp + pq}{2} \eta(t) \frac{q^2}{2}} \tag{4.5}
\]

Inserting (4.5) into (4.4) and remembering that the first exponent in (4.4) is the scale operation, one gets:

\[
\eta(t) = \int_a^t (1 - \lambda \phi(t)) \left(\frac{\rho(a)}{\rho(t)}\right)^2 \cos 2\alpha dt \tag{4.6}
\]

and returning to (4.2) one has an explicit formula for the unitary operator \( U(t, a) \) with three exponential terms piloted by Prüfer variables (2.6-7):

\[
U(t, a) = e^{-i\alpha(t)H_0} e^{i \ln(\frac{\rho(t)}{\rho(a)}) \frac{qp + pq}{2} \eta(t) \frac{q^2}{2}} \tag{4.7}
\]

If in addition the number \( \lambda \) is the eigenvalue of the Sturm-Liouville problem (2.1) \( \Rightarrow \alpha(b) = n\pi \ (n = 1, 2, \ldots) \), then the first factor in (4.7) becomes \( P^n \) (where \( P = \text{parity} \)) and

\[
U(b, a) = P^n e^{i \ln(\frac{\rho(b)}{\rho(a)}) \frac{qp + pq}{2} \eta \frac{q^2}{2}} \tag{4.8}
\]

transforms any initial wave packet \( \psi(x) = \psi(x, a) \) into:

\[
\psi_U(x) = [U(b, a)\psi](x) = \frac{1}{\sqrt{\sigma}} e^{i\eta \frac{x^2}{2\sigma^2}} \psi(x/\sigma) \tag{4.9}
\]

where \( \sigma = \pm \rho(a)/\rho(b) \). The probability density \( \rho = |\psi(x)|^2 \) is simultaneously mapped into its own diffractionless image:

\[
\rho_U = |U\psi|^2 = \frac{1}{\sigma}\rho(x/\sigma) \tag{4.10}
\]

It thus turns out that the authors looking for the squeezing parameters were indeed solving the XIX-century spectral problem! The relative simplicity of the three term formula (4.7) (compare \([1, 12, 13, 14]\)) is due to the "bottom structure" (2.6-9). Its additional implication concerns the "potential kicks" \( \delta(t)V(q) \) considered recently as a method of controlling and cooling quantum systems \([17]\). While the "kicks" are mathematically simple, they are overidealized (no...
δ-pulses available in laboratories!). Notice, however, that if \( \lambda \) is an eigenvalue of (2.1) and if \( \phi(t) \) is symmetric \( \phi(-t) = \phi(t) \) in a symmetric operation interval \([-a, a]\), then the eigenvector \( \psi \) (and consistenly the phase trajectory \( q(t) \)) is even/odd in \([-a, a]\), implying the trivial squeezing \( \sigma = \pm 1 \), and the only non-trivial part of (4.7-8) is \( P^m e^{i\eta \frac{\pi}{2}} \). For \( n = 2l \), this is the unitary operator describing the results of a sharp kick of the oscillator potential \(-\eta \delta(t)q^2/2\). Thus, the spectrum of (2.1) tells in addition how to "imitate softly" the effects of δ-kicks converting the results on kicked oscillators \[17\] into a branch of realistic technology.

5.- Magnetic Images.

The results (3.4), (4.3) have a simple laboratory model. Consider a non-relativistic charged particle in 3-dim, in a homogeneous time-dependent magnetic field \( B(t) = B(t)n \), with a vector potential:

\[
\mathbf{A}(x, t) = -\frac{1}{2} \mathbf{x} \times \mathbf{B}(t)
\]  

where \( n \) is a constant unit vector and \( B(t) = B\gamma(t/T) \) is the magnetic field intensity, \( B \) defining 'the amplitude' and \( \gamma \) the 'shape' of the magnetic pulses. If \( n \) point the \( z \)-axis, the Schrödinger’s wave packet obeys the Hamiltonian:

\[
H(\tau) = \frac{1}{2}p_3^2 - \beta\gamma(\tau)M_3 + \frac{1}{2}[p_1^2 + p_2^2 + \beta^2\gamma^2(\tau)(q_1^2 + q_2^2)]
\]  

which we have written in the dimensionless coordinates: \( \tau = t/T, (q_1, q_2, q_3) = \sqrt{m/T}h(x, y, z), (p_1, p_2, p_3) = (\sqrt{m\hbar/T})^{-1}(p_x, p_y, p_z) \); all physical information now contained in the 'shape function' \( \gamma(\tau) \) and in the 'dimensionless amplitude':

\[
\beta = \frac{eBT}{2mc}
\]

In (5.2) the term \( \frac{1}{2}p_3^2 \) means the free propagation along the \( z \)-axis, the \(-\beta\gamma(\tau)M_3 \) generates rotations in the \( q_1, q_2 \)-plane. The last term (a 2-dimensional oscillator), splits into two twin 1-dimensional oscillators of the form:

\[
h(\tau) = p^2/2 + \beta^2\gamma^2(\tau)q^2/2
\]  

where \( (q, p) \) means either \( (q_1, p_1) \) or \( (q_2, p_2) \). Quite evidently, (5.4) is a model of (2.5) with \( \lambda = \beta^2 \) and \( \phi(\tau) = \gamma^2(\tau) \geq 0 \). If \( \gamma(\tau) \neq 0 \), the spectrum of the Helmholtz equation (2.1) forms a ladder of eigenvalues, and if \( \lambda = \beta^2 \)
is one of them, both oscillators yield the simultaneous focus of the trajectory congruencies (in classical case) or probability patterns (in quantum case). To check this, we have asked the computer to draw congruencies of classical trajectories diverging from a common initial point, with a common initial $v_z$, for $\lambda = \beta^2$ coinciding with one of the eigenvalues of (2.1) and with $\phi(\tau) = \beta^2(\tau)$ taken from Table 1. In all cases we have obtained quite precise optical images on planes perpendicular to the solenoid axis (see Fig. 3). The dependence of the phenomenon on the Sturm-Liouville spectrum is illustrated by sequences of amplified focuses corresponding exactly to the nodal points of the wave function (2.1-2). Henceforth, the integration of the first order differential equation of Prüfer (for one variable) offers the simplest algorithm to predict the optical echo (squeezing) for charged particle beams.

The ”reduced images” might imply new schemes to produce narrow particle beams. The ”amplified images” are no less interesting. They can certainly be used to separate the wave packet components in the quantum state engineering [18, 19, 20, 21]). The magnetic pulses $B(t)$ with $\phi(t) = \gamma^2(t)$ symmetric in the operation interval and with $\lambda = \beta^2$ in the spectrum of (2.1), provide a magnetic model for systems under the influence of oscillator kicks.

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**Figure Captions**

**Fig. 1.** The orbit of the classical oscillator (2.5) for $\phi(t) = \sin^2 t$, $\lambda = 10$. The points of intersection with the $p$-axis determine the time moments $t_1 = 1.43610$, $t_2 = 2.53311$, $t_3 = 4.46714$, $t_4 = 5.51549$, ..., and the amplifications $\sigma_1 = -0.52322$, $\sigma_2 = 0.65528$, $\sigma_3 = -0.23823$, $\sigma_4 = 0.27464$, ... of the optical images. The similar phenomenon would be generated by a chaotic $\phi(t)$.

**Fig. 2.** The orbits of (2.3) and the Helmholtz spectrum of (2.1). If the end point of the orbit with $q_{a}(a)=0$ places itself on the $p$-axis, $\lambda$ is an eigenvalue of the Helmholtz equation (1). Above: a sequence of orbits reaching $q_{b}(b) = 0$ after multiple intersections with the $p$-axis illustrates the sequence of Helmholtz eigenstates for $\phi(t) = (1 + \sin 2\pi t)^2$, $[a, b] = [0, 1]$. The corresponding optical constants are reported in Table 1.

**Fig. 3.** The optical effect of the Helmholtz spectrum observed in the solenoid for sinusoidally oscillating $\gamma(\tau) = 1 + \sin 2\pi \tau$. 

a) When $\lambda = \beta^2 = 5.3347146$ is an eigenvalue of (2.1), the congruence of particle trajectories forms an exact optical focus at $\tau = 1$ ($t = T$) (in quantum case, an amplified diffractionless image) with $\sigma = -2.35707$. The integration of the first order differential equation of Prüfer for one angular variable turns out the simplest method to predict the effect. 

b) For $\lambda = \beta^2 = 86.8947093$ the congruence forms 3 increasingly amplified focuses (images of the initial point) in exact correspondence to the 3 nodal points of the wave function $\psi(x)$. 