The Schrödinger equation, the zero-point electromagnetic radiation and the photoelectric effect.

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Abstract
A Schrödinger type equation for a mathematical probability amplitude $\Psi(x, t)$, is derived from the generalized phase space Liouville equation valid for the motion of a microscopic particle, with mass $M$, moving in a potential $V(x)$. The particle phase space probability density is denoted $W(x, p, t)$ and the entire system is immersed in the “vacuum” zero-point electromagnetic radiation. We show that the generalized Liouville equation is reduced to a non-quantized Liouville equation in the equilibrium limit where the small radiative corrections cancel each other approximately. Our derivation will be based on a simple Fourier transform of the non-quantized phase space probability distribution $W(x, p, t)$. For convenience, we introduce in this Fourier transform an auxiliary constant $\alpha$, with dimension of action, and an auxiliary coordinate denoted by $y$. We shall prove that $\alpha$ is equal to the Planck’s constant present in the momentum operator of the Schrödinger equation. Moreover, we shall show that this momentum operator is deeply related with the ubiquitous zero-point electromagnetic radiation. It is also important to say that we do not assume that the mathematical amplitude $\Psi(x, t)$ is a de Broglie matter-wave, in other words, the wave-particle duality hypothesis is not used within our work. The implications of our study for the standard

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interpretation of the photoelectric effect is discussed by considering the main characteristics of the phenomenon. We also mention, briefly, the effects of the zero-point radiation in the tunneling phenomenon and the Compton’s effect.

Key words: Foundations of quantum mechanics; Schrödinger equation; Zero-point electromagnetic fluctuations; Photoelectric effect.
1. Introduction

In 1932 E. Wigner published an important paper [1, 2] where he introduced what is called today the Wigner’s quasi probability function, or simply Wigner’s phase space function. We shall denote it by $Q(x, p, t)$ and, according to the Wigner proposal, it is connected with a Schrödinger equation solution $\Psi(x, t)$ by the expression:

$$Q(x, p, t) \equiv \frac{1}{\pi \hbar} \int_{-\infty}^{+\infty} \Psi^*(x - y, t) \Psi(x + y, t) e^{-\frac{2ipy}{\hbar}} dy,$$  \hspace{1cm} (1)

where $\hbar$ is the Planck’s constant. Wigner’s intention was to obtain a quantum mechanical description of the phase space microscopic and macroscopic phenomena as demanded by the correspondence principle [3, 4, 5]. We recall that the Schrödinger equation, namely

$$i\hbar \frac{\partial}{\partial t} \Psi(x, t) = \left[ -\frac{\hbar^2}{2M} \frac{\partial^2}{\partial x^2} + V(x) \right] \Psi(x, t),$$  \hspace{1cm} (2)

also depends explicitly on the Planck’s constant $\hbar$ [6].

We want to emphasize that neither Wigner nor Schrödinger clarified the relation of the constant $\hbar$ with the classical stochastic zero-point electromagnetic radiation. Notice, however, that the existence of the classical stochastic zero-point radiation was considered seriously by Albert Einstein and Otto Stern in the period 1911-1913 [7], and by T. W. Marshall in an important paper published in 1963 [8]. Interesting experimental observations of the “vacuum” zero-point fluctuations are described in more recent publications [9].

We also would like to mention that, if applied to an arbitrary solution $\Psi(x, t)$ of the Schrödinger equation (2), the expression (1) may lead to negative phase space probability densities. In other words, the relation (1) should be considered only as the definition of an auxiliary function, which has some properties of a classical distribution, useful to calculate the statistics on the configuration and momentum space. Interesting examples are the excited states of the harmonic oscillator [10].

The definition (1), and the Schrödinger equation (2), do not lead to the classical Liouville equation except when $V(x)$ is quadratic in the variable $x$ (see ref. [3]). However, according to the correspondence principle, one should
necessarily obtain the Liouville equation in the case of a macroscopic particle moving in a generic potential $V(x)$.

We recall that our method was suggested by some authors in the past [4, 11]. Here, in contrast with the original Wigner procedure, we start from a non-quantized Liouville equation in order to obtain a Schrödinger type equation [12, 13, 14].

In other words, starting with the generalized Liouville equation in phase space, we shall obtain the mathematical and the physical conditions necessary to relate it with an equation that is formally identical to the Schrödinger equation for a mathematical probability amplitude $\Psi(x,t)$ appropriately defined in the configuration space.

With the use of the Wigner transform, that is, the inverse Fourier transform of the equation (1), we shall disclose the conditions in which the relation between an approximate Liouville equation and the Schrödinger type equation is valid for a generic potential $V(x)$.

Our presentation is organized as follows. We give, within section 2, the physical background and the mathematical approximations necessary to connect the non-quantized phase space probability distribution to the mathematical configuration space probability amplitude $\Psi(x,t)$. The relation between the corresponding Liouville equation and the Schrödinger equation is explained in details within the section 3. The role of the random zero-point radiation with spectral distribution $\rho_0(\omega)$ given by [11, 12]

$$
\rho_0(\omega) \equiv \lim_{T \to 0} \left[ \frac{\omega^2}{\pi^2 c^3} \left( \frac{\hbar \omega}{2} + \frac{\hbar \omega}{\exp(\frac{\hbar \omega}{kT}) - 1} \right) \right] = \frac{\hbar \omega^3}{2\pi^2 c^3},
$$

is also explained within the section 3. Some implications of our study for the standard interpretation of the photoelectric effect are presented in the section 4. The section 5 is devoted to a brief discussion in which we mention the role of $\rho_0(\omega)$ in the free particle propagator, in the tunneling phenomenon and in the Compton effect.
2. Physical background and mathematical approximations

Our starting point is the non-relativistic Liouville equation for the classical probability distribution in phase space, denoted by \( W(x, p, t) \). This probability density is the average distribution over the realizations of the fluctuating electric field indicated by \( E_x(t) \) in the equation (5) below. The generalized Liouville equation, associated with the non-relativistic motion of a particle with mass \( M \) and charge \( e \), evolves in time according to \([12, 13, 14]\)

\[
\frac{\partial}{\partial t} W(x, p, t) + \frac{p}{M} \frac{\partial W(x, p, t)}{\partial x} + \frac{\partial}{\partial p} \left[ F(x) + \frac{\gamma}{M} F'(x)p \right] W(x, p, t) \approx \frac{e^2}{3 M} \frac{\partial^2}{\partial p^2} W(x, p, t). \quad (4)
\]

Here \( F(x) \) is the deterministic force, \( \gamma \) is a very short time interval (\( \gamma = \frac{2}{3} \frac{e^2}{M c^3} \simeq 6.3 \times 10^{-24} \text{ sec} \) for electrons) and \( e^2 D(t) \) is the time dependent diffusion coefficient. According to de la Peña and Cetto [12] the eq. (4) is a non-Markoffian Fokker-Planck equation with memory (see ref. [12] for the details). We recall that the term proportional to \( \gamma \) in (4) is due to the radiation reaction force.

The equation (4) was derived by assuming that the total force acting on the charged particle is given by

\[
M \frac{d^2 x}{dt^2} = F(x) + eE_x(t) + \frac{2}{3} \frac{e^2}{c^3} \ddot{x}(t) + ... , \quad (5)
\]

where \( E_x(t) \) is the “vacuum” electric field in the dipole approximation. This random field is such that \( \langle E_x(t) \rangle = 0 \) on average. The reader can see the work of T. H. Boyer [15] for an expression of \( E_x(t) \) in terms of plane waves with random phases. The last term in (5) is of order \( e^2 \), and is the first approximation of the radiation reaction force.

According to the analysis presented by L. de la Peña and A. M. Cetto (see the sections II, III and V of their paper [12] ) it is possible to conclude that (4) is equivalent to the non-quantized Liouville equation

\[
\frac{\partial W(x, p, t)}{\partial t} + \frac{p}{M} \frac{\partial W(x, p, t)}{\partial x} + F(x) \frac{\partial W(x, p, t)}{\partial p} \approx 0 , \quad (6)
\]
in the equilibrium limit where the radiation reaction term and the diffusion term in \( \text{(4)} \) compensate each other in an approximate way \[12\].

After the above clarifications we shall use the simpler equation \( \text{(6)} \) in order to relate it with the Schrödinger type equation for the mathematical probability amplitude \( \Psi(x,t) \). The approximate equation \( \text{(6)} \) is much more easy to handle than the generalized Liouville equation \( \text{(4)} \).

Following S. Hayakawa \[4\] we shall consider the following Fourier transform defined by
\[
\tilde{W}(x, y, t) \equiv \int_{-\infty}^{\infty} dp W(x, p, t) e^{\frac{2\pi p y}{\alpha}}, \tag{7}
\]
where \( y \) is the auxiliary coordinate, \( \alpha \) is the auxiliary constant with dimension of action, and \( W(x, p, t) \) is the phase space distribution function (see the equation \( \text{(6)} \)). The constant \( \alpha \) will be easily fixed later with the use of physical and mathematical criteria.

We know that \( W(x, p, t) \) is always real and positive definite and the classical configuration space probability density can be defined mathematically as
\[
P(x, t) \equiv \int dp W(x, p, t) = \lim_{y \to 0} \tilde{W}(x, y, t). \tag{8}
\]

Notice that this probability density is the average distribution over the realizations of the fluctuating electric field indicated by \( E_x(t) \) in the equation \( \text{(5)} \).

Since \( P(x, t) \) is real and positive it is always possible to introduce a mathematical probability amplitude \( \Psi(x, t) \) as \[12, 13, 14\]
\[
P(x, t) \equiv |\Psi(x, t)|^2 = \Psi^*(x, t)\Psi(x, t), \tag{9}
\]
where the factorization \( \Psi^*(x, t)\Psi(x, t) \) is completely general in the limit \( y \to 0 \). In other words, the equations \( \text{(8)} \) and \( \text{(9)} \) can be summarized as
\[
\Psi^*(x, t)\Psi(x, t) \equiv \lim_{y \to 0} \int_{-\infty}^{\infty} dp W(x, p, t) e^{\frac{2\pi p y}{\alpha}} \equiv \lim_{y \to 0} \tilde{W}(x, y, t), \tag{10}
\]
where the limiting process considered above means that the behavior of \( \tilde{W}(x, y, t) \) when \( y \approx 0 \) is very important. However, the above definition \( \text{(10)} \) is not enough to determine a differential equation for the amplitude \( \Psi(x, t) \). In order to achieve this goal we shall use the approximate Liouville equation \( \text{(6)} \). We shall see, in the next section, that the approximate Liouville equation has a close relation with the Schrödinger type equation for the mathematical probability amplitude \( \Psi(x, t) \).
3. The approximate Liouville equation and its connection with the Schrödinger equation

It is a widespread belief [3] that the connection between the Wigner quasi probability function (1), the Liouville equation (6), and the Schrödinger equation (2), is only possible for quadratic potentials in the $x$ variable. We shall extend this connection to a more general potential $V(x)$, that is, beyond the case of the harmonic potential. In order to achieve this goal we shall use a procedure similar to that presented by S. Hayakawa [4] and K. Dechoum, H. M. França and C. P. Malta [5]. Dechoum, França and Malta discussed several classical aspects of the Pauli-Schrödinger equation written in the spinorial notation. Here we are treating the spinless case.

To obtain the differential equation for $\Psi(x, t)$ we shall use the Wigner type transformation defined previously in the equation (7). Regarded as a simple mathematical definition, we stress that this Fourier transform contains the same dynamical information carried by the phase space density $W(x, p, t)$.

We observe that, due to the definitions (7), (8), (9) and (10), the Wigner type transform $\tilde{W}(x, y, t)$ is a complex function that has a physical significance only in the limit $|y| \to 0$, that is,

$$\lim_{y \to 0} \tilde{W}(x, y, t) = |\Psi(x, t)|^2 = \Psi(x, t)\Psi^*(x, t).$$

(11)

For this reason, we shall consider, in what follows, the definition (7) only for small values of $y$ (see also the reference [16] for a similar approach). Our goal is to obtain the differential equation for $\Psi(x, t)$, from the differential equation for $\tilde{W}(x, y, t)$, in the limit of small $y$.

Our first step is to consider the equation

$$\frac{\partial}{\partial t} \tilde{W}(x, y, t) = \int_{-\infty}^{\infty} \frac{\partial W}{\partial t} e^{\frac{2iuy}{\alpha}} dp.$$  

(12)

Using the Liouville equation (6) and after an integration by parts, we obtain

$$\frac{\partial \tilde{W}}{\partial t} = -\int_{-\infty}^{\infty} \left[ \frac{p}{M} \frac{\partial W}{\partial x} + F(x) \frac{\partial W}{\partial p} \right] e^{\frac{2iuy}{\alpha}} dp =$$

$$= \left[ -\frac{i\alpha}{2M} \frac{\partial^2}{\partial x^2} + F(x) \frac{2iy}{\alpha} \right] \int_{-\infty}^{\infty} W(x, p, t) e^{\frac{2iuy}{\alpha}} dp.$$  

(13)

Substituting (7) into (13) we get the Hayakawa [4] equation for $\tilde{W}(x, y, t)$, namely
\[ i \alpha \frac{\partial}{\partial t} \tilde{W}(x, y, t) = \left[ \frac{(-i \alpha)^2}{2M} \frac{\partial^2}{\partial y \partial x} - 2yF(x) \right] \tilde{W}(x, y, t). \quad (14) \]

In order to facilitate the calculations it is convenient to use new variables, namely \( s = x - y \) and \( r = x + y \), so that the equation (14) can be written as

\[ i \alpha \frac{\partial}{\partial t} \tilde{W}(r, s, t) = \left[ \frac{(-i \alpha)^2}{2M} \left( \frac{\partial^2}{\partial r^2} - \frac{\partial^2}{\partial s^2} \right) - (r - s)F \left( \frac{r + s}{2} \right) \right] \tilde{W}(r, s, t). \quad (15) \]

According to (11), we want an equation for \( \tilde{W}(r, s, t) \), when \( y = \frac{r - s}{2} \to 0 \), that is, when \( r \to s \). Consequently we must consider that the points \( r \) and \( s \) are arbitrarily close in the equation (15). Therefore, in accordance to the mean value theorem, one can consider that

\[ - (r - s)F \left( \frac{r + s}{2} \right) \simeq - \int_{s}^{r} F(\xi)d\xi \simeq V(r) - V(s), \quad (16) \]

is a very good approximation [4]. Notice that the mean value theorem requires only that the integrand function (force \( F(x) = -V'(x) \) in this case) is a continuous function in the small integration interval under consideration.

The use of equation (16) will imply in a great simplification of the problem of finding the differential equation for the mathematical probability amplitude \( \Psi(x, t) \), because (15) becomes separable in the variables \( r \) and \( s \) as we shall see below. Substituting (16) into (15) we get the equation

\[ i \alpha \frac{\partial}{\partial t} \tilde{W}(r, s, t) = \left[ \frac{(-i \alpha)^2}{2M} \left( \frac{\partial^2}{\partial r^2} - \frac{\partial^2}{\partial s^2} \right) + V(r) - V(s) \right] \tilde{W}(r, s, t), \quad (17) \]

which was obtained by S. Hayakawa in 1965 (see also ref. [12]). One can verify that the above equation is separable in the variables \( r \) and \( s \) if we introduce a function \( \tilde{W}(r, s, t) \) such that

\[ \tilde{W}(r, s, t) \equiv \Psi^*(s, t)\Psi(r, t). \quad (18) \]

Notice that (18) is in accordance with our previous equation (9). Thus (17) is decomposed into the equation

\[ i \alpha \frac{\partial}{\partial t} \Psi(x, t) = \left[ \frac{1}{2M} \left( -i \alpha \frac{\partial}{\partial x} \right)^2 + V(x) \right] \Psi(x, t), \quad (19) \]
and its complex conjugate. The equation (19) is formally identical to the Schrödinger equation provided that the operator $-i\alpha \partial/\partial x$ is identified with the Schrödinger momentum operator $p = -i\hbar \partial/\partial x$, that is, the auxiliary constant $\alpha$ is identified with $\hbar$. See the paragraphs containing our equations (23), (24) and (25), where we discuss this point using the Heisenberg picture.

Using a more explicit notation we get, from the equation (19), the result

$$i\hbar \frac{\partial \Psi}{\partial t}(x, t) = \left[ \frac{1}{2M} \left( -i\hbar \frac{\partial}{\partial x} \right)^2 + V(x) \right] \Psi(x, t) \equiv \hat{H}\Psi(x, t),$$

(20)

which has exactly the form of the Schrödinger equation for probability amplitude $\Psi(x, t)$. Notice that $\hat{H}$ is the Hamiltonian operator of the system and the equation (20) possesses a complete set of solutions, $\Psi_n(x, t) \equiv \phi_n(x) \exp(-i\epsilon_n t/\hbar)$, where the functions $\phi_n(x)$ are such that

$$\hat{H}\phi_n(x) \equiv \epsilon_n \phi_n(x)$$

(21)

The constants $\epsilon_n$ are the “energy levels” and the functions $\phi_n(x)$ are the “eigenfunctions” of the system [17].

We stress that, due to our approximations, the excited states of the set \{\epsilon_n, \phi_n(x)\} do not decay (see the references [18] and [19] for the physical explanation of the decay processes). It is important to recall that the equation (20) was derived from the non-quantized Liouville equation (6) for the phase space probability distribution $W(x, p, t)$.

We know that the Planck’s constant $\hbar$ is related with the electromagnetic fluctuation phenomena characteristic of Stochastic Electrodynamics (SED) [12, 15] and Quantum Electrodynamics (QED) [19]. The most important of these fluctuations are associated with the zero temperature electromagnetic radiation which has a spectral distribution such that

$$\rho_0(\omega) = \frac{\hbar \omega^3}{2\pi^2 c^3},$$

(22)

in both QED and SED. We recall that these two theories have many features in common [19, 15]. It is also interesting to recall that the classical effects of the zero-point radiation were discovered very early by M. Planck in the period 1911-1912 [11]. The quantum zero-point radiation was discovered latter, in 1927, by Paul Dirac.

Some properties of the spectral distribution $\rho_0(\omega)$ will be explicitly used at this point. For the moment we would like to say that, using the Heisenberg picture, Sokolov and Tumanov [20] and P.W Milonni [21, 22] were able
to show that the commutation relation between the position operator \( x(t) \) and the momentum operator \( p(t) \) are strongly related to \( \rho_0(\omega) \) given in the equation (22). Moreover, according to de la Peña, Valdés-Hernández and Cetto [23], who use a non-quantized system as a starting point, the action of the electric zero-point field on matter is essential and ultimately leads the matrix (or Heisenberg) formulation of quantum mechanics for the motion in an arbitrary potential \( V(x) \). This is an important conclusion. Therefore, according to these authors, a “free” electron (mass \( M \) and charge \( e \)) has a nonrelativistic equation of motion for the position operator \( x(t) \) such that

\[
M\ddot{x}(t) = \frac{2e^2}{3c^3} \ddot{x}(t) + eE_x(t),
\]

within the Heisenberg picture [20] [21] [22] [23]. Recall that, according to the authors quoted in ref. [23], \( E_x(t) \) can be a fluctuating non-quantized electric field. We shall use the Heisenberg approach in what follows.

Within the Heisenberg picture the equation (23) can be easily solved. We get the position operator \( x(t) \) and the canonical momentum operator \( \dot{p}(t) \equiv M\dot{x}(t) + \frac{e}{c}A_x(t) \), where \( A_x(t) \) is the quantized vector potential \( E_x(t) = -\frac{1}{c}\frac{\partial A_x}{\partial t} \).

From the solution for \( x(t) \) we get \( p(t) \). Therefore, it is possible to calculate the commutation relation between the position operator \( x(t) \) and the canonical momentum operator \( p(t) \). The result is

\[
[x(t), p(t)] = [x, M\dot{x}] = 4i\pi^2c^3\gamma \int_0^\infty d\omega \frac{\rho_0(\omega)}{\omega^3(1 + \gamma^2\omega^2)} = i\hbar,
\]

where \( \gamma = \frac{2e^2}{3Mc^3} \). This is an interesting and relevant result because only \( \rho_0(\omega) \) depends on \( \hbar \) in the above integral. Moreover, the result (24) is independent of the charge of the particle. Since the Heisenberg and the Schrödinger pictures are equivalent (see the references [21] and [25] for a good discussion of this point) one concludes that the momentum operator used in the Schrödinger type equation (20), namely

\[
p = -i\hbar\partial/\partial x,
\]

is deeply related with the “vacuum” electromagnetic fluctuations with spectral distribution \( \rho_0 = \hbar\omega^3/2\pi^2c^3 \).
4. The photoelectric effect and the zero-point radiation

A three-dimensional generalization of the one-dimensional Schrödinger type equation (20) is

\[ i\hbar \frac{\partial \Psi(\vec{r}, t)}{\partial t} = \left[ -\frac{\hbar^2}{2M} \nabla^2 - V(\vec{r}) - e\vec{r} \cdot \vec{E}_c(\vec{r}, t) \right] \Psi(\vec{r}, t), \]  

where we have used the dipole approximation. The term \(-e\vec{r} \cdot \vec{E}_c(\vec{r}, t)\) is the interaction energy, \(e\) is the electron charge and \(M\) is the electron mass.

In the above equation \(\vec{E}_c(\vec{r}, t)\) is a classical deterministic electric field, associated with the incident radiation, and \(V(\vec{r})\) is the atomic Coulomb potential for instance. We are not explicitly considering the stochastic “vacuum” fields in the equation (26) because it was concluded, in previous works [24, 25], that this may generate double counting. The operator \(-i\hbar \nabla\) already contains the effects of the zero-point electromagnetic fields necessary to our purposes.

An interesting physical feature of the above three-dimensional generalization of the mathematical Schrödinger type equation (20) is that it leads to a novel interpretation of the photoelectric effect.

The reader can find a good discussion of this point in the subsection entitled “The photoelectric effect” (pgs. 40, 41 and 42) of the article by M. O. Scully and M. Sargent III published in 1972 [26]. Within this paper it is used a classical deterministic electric field, of a monochromatic plane wave, polarized in the \(z\) direction, which is denoted \(\vec{E}_c(\vec{r}, t) = \hat{z}E_0 \cos(\omega t - ky)\), where \(E_0\) is the amplitude of the electric field and \(\omega\) is the corresponding angular frequency (\(\omega = kc\)).

According to an standard analysis, based on the equation (26) and on the “Fermi Golden Rule”, these authors obtained the expression [26, 27]

\[ \frac{dP_f}{dt} = 2\pi |\langle f | e\vec{r} \cdot \hat{z} | g \rangle|^2 \hbar^2 E_0^2 t \delta \left( \omega - \frac{\epsilon_f - \epsilon_g}{\hbar} \right), \]  

for the probabilistic rate of emission of photo-electrons from a bounded initial state of energy \(\epsilon_g\) to the final state of energy \(\epsilon_f = Mv^2/2\), where \(\epsilon_f\) is the continuous energy of the ejected electron, and \(e\vec{r}\) is the electric dipole of the
atom. Notice, however, that the authors mentioned in references [26] and [27] do not mention the role of the zero-point radiation.

From the argument of the delta function in (27), we see that

\[ \omega = \frac{\epsilon_f - \epsilon_g}{\hbar}. \]  

(28)

The Planck’s constant \( \hbar \) appearing above was introduced within the mathematical Schrödinger type equation (26), without using the concept of photon. Moreover, according to the interpretation of the Schrödinger equation presented within our paper (see the section 3) the origin of \( \hbar \) can be traced back to the expressions (24) and (25) which involves \( \rho_0(\omega) \), that is, the Planck’s constant, in the equation (28), has its origin in the zero-point radiation with spectral distribution \( \rho_0 = \hbar \omega^3/(2\pi^2c^3) \). Therefore, the famous Einstein equation for the photoelectric phenomenon follows from (28) and can be put in the form [26, 27].

\[ \hbar \omega = \epsilon_f - \epsilon_g \geq \frac{Mv^2}{2} + \phi, \]  

(29)

where \( \phi \) is the work function characteristic of the material, and \( Mv^2/2 \) is the continuous kinetic energy of the ejected electron. See our figure 1 for an illustration.

![Figure 1: Illustration of the photoelectric effect from a potassium surface (2.0 eV needed to eject electrons). In the figure we indicate the characteristic energy \( \hbar \omega \) of the “photons”. For instance \( \hbar \omega = 1.8 \) eV, \( \hbar \omega = 2.3 \) eV and \( \hbar \omega = 3.1 \) eV. The maximum velocity \( V_{\text{max}} \) of the ejected electrons, is indicated for each frequency \( \omega \) (see the Einstein equation (29)). Notice that the electrons inside the metal move with different energies.](image-url)
In order to better describe the “collision” (or scattering) process we need the equations for the momentum conservation. We recall that Einstein introduced the classical concept of “needle radiation” [11] which has “momentum and energy”. See our figure 2 for an illustration.

![Diagram of momentum conservation in the photoelectric phenomenon.](image)

Figure 2: Illustration of the momentum conservation in the photoelectric phenomenon. In this case the atom will suffer a recoil in the indicated direction. In other words, the parent atom (or crystal) plays an important role in the photoelectric effect (see the section III of the reference [28]).

The role of the momentum conservation equations is very well illustrated in the section III of the paper by R. Kidd, J. Ardini and A. Anton [28]. These authors gave very interesting examples relating the velocity variation of the recoiling electron, and the direction of the recoil of the parent atom. In other words, it is shown that for some values of $\hbar \omega$ and $\phi$ the value of the momentum of the recoiling electron can be much larger than the momentum of the incident “photon” due to the recoil of the parent atom (see the figure 2 in the section III of the reference [28]).

Consequently, the Einstein’s equation for the photoelectric effect can be interpreted in a non-quantized manner where the zero-point electromagnetic field play a very important role. This opens a new way of understanding the photoelectric effect. It is our intention to explore this interpretation in future calculations, by considering the atoms in special physical conditions
(flying between very close to conducting Casimir plates [29], localized close to some element of an electric circuit [30] or moving inside a solid material [31]). In these cases there are important modifications in the spectral distribution \( \rho_0(\omega) \).

5. Brief discussion

We want to recall that within this paper we have clarified the deep relationship between the Schrödinger equation, the zero-point electromagnetic radiation and the photoelectric effect. Another important observation is that the kinetic energy operator present in the unidimensional equation (20), namely

\[
\frac{1}{2M} \left(-i\hbar \frac{\partial}{\partial x}\right)^2,
\]

(30)
generates the spreading of the mathematical probability amplitude \( \Psi(x,t) \). In other words, if \( V(x) = 0 \) we get from the equation (20) the “free” particle propagator [32]

\[
K(x,t|x_0,t_0) = \sqrt{\frac{M}{2\pi i\hbar(t-t_0)}} \exp \left[ \frac{iM(x - x_0)^2}{2\hbar(t-t_0)} \right],
\]

(31)

which connects \( \Psi(x_0,t_0) \) with \( \Psi(x,t) \) in the time interval \( t-t_0 \), that is

\[
\Psi(x,t) = \int_{-\infty}^{\infty} dx_0 \Psi(x_0,t_0) K(x,t|x_0,t_0).
\]

(32)

This solution was obtained from the mathematical Schrödinger equation (20) assuming that \( V(x) = 0 \). The above equations mean that, even in the far “empty” space, for example, the matter particles are always interacting with the vacuum zero-point radiation. The inevitable conclusion is that the fluctuating electromagnetic background with spectral density \( \rho_0(\omega) = \hbar^3/2\pi c^3 \) is the source of the zero temperature diffusion. In other words, the constant \( \hbar \) present in the propagator \( K(x,t|x_0,t_0) \) has its origin in the spectral density \( \rho_0(\omega) \). Notice that the equations (31), (32) and the superposition principle suggest that the interference phenomena of electrons, and other matter particles, may be explained without the use of de Broglie waves.

We also would like to say that, from the connection between the approximate Liouville equation (6) and the Schrödinger type equation (20) presented here, one can conclude that the mathematical probability density \( |\Psi(x,t)|^2 \)
also may be not necessarily associated with real waves. Moreover, according to the *Copenhagen Interpretation*, the amplitude $\Psi(x, t)$ is not a physical wave but a “probability wave”. As a matter of fact, the existence of de Broglie waves was recently questioned by Sulcs, Gilbert and Osborne in their analysis of the experiments, by M. Arndt et al., on the interference of massive particles as the $C_{60}$ molecules (fullerenes). According to M. Arndt et al., “the de Broglie wavelength of the interfering fullerenes is already smaller than the diameter of a single molecule by a factor of almost 400”. This is an important observation which deserves further attention.

It is equally important to mention that, recently, a *tunneling* phenomenon was also interpreted in classical terms, that is, without using the concepts of de Broglie waves and the wave-particle duality hypothesis. The important concept used was, again, the existence of the zero-point radiation with spectral distribution given by $\rho_0(\omega) = \hbar \omega^3 / 2\pi^2 c^3$.

Finally we want to say a few words concerning the relationship between the Compton effect and the zero-point radiation. This relationship was established by H. M. França and A. V. Barranco in 1992, by studying a system of free molecules in equilibrium with thermal and zero-point radiation.

In their work Barranco and França, using a statistical approach, replaced the Einstein concept of *random spontaneous emission* by the concept of *stimulated emission by the random zero-point electromagnetic fields* with spectral distribution $\rho_0(\omega)$. As a result, Compton and Debye’s kinematic relations were obtained within the realm of a completely classical theory, that is, without having to consider the wave-particle duality hypothesis for the molecules or the radiation bath. This is another important dynamical effect of the zero-point radiation. We recall that many other relevant effects of the zero-point electromagnetic fields are presented in the excellent book by L. de la Peña and A. M. Cetto.
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