Synoptic-Scale Variability and Its Relationship with Total Ozone and Antarctic Vortex Displacements

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ABSTRACT

The main synoptic-scale circulation anomaly pattern over extratropical South America during the austral spring (September–November) is identified by means of rotated extended empirical orthogonal function techniques, applied to the meridional wind perturbation time series at 300 hPa. The dataset is based on 15 spring seasons (1979–93) of meteorological data from the National Centers for Environmental Prediction–Department of Energy Atmospheric Model Intercomparison Project version-2 daily averaged reanalyses, given in 17 vertical levels from 1000 to 10 hPa. The total-ozone daily measurements for the same period are from the Total Ozone Mapping Spectrometer instrument (version 7). The principal synoptic-scale anomaly pattern is associated with an anticyclone–cyclone pair evolving eastward along subpolar latitudes (and hence it is termed the subpolar mode), with a typical length scale of 5000 km and a phase velocity of 8 m s⁻¹. The subpolar-mode waves, which display the main characteristics of midlatitude baroclinic waves, typically maximize near or above the tropopause and propagate upward into the lower stratosphere, showing large amplitudes even at 50 hPa and above.

Subpolar-mode-related circulation anomalies are found to be responsible for large total-ozone daily fluctuations near southern South America and nearby regions. In the positive phase of the subpolar mode, total-ozone fluctuations, which are negative, adopt a sigmoid structure, with a zonal scale as large as the anticyclone–cyclone pair. Moreover, it is herein shown that the associated anticyclone produces a local ozone-column decrease to the north and east of its center, due to adiabatic uplift of air parcels in the upper troposphere and lower stratosphere. At the same time, the downstream cyclonic disturbance is responsible for large negative total-ozone anomalies to the west and south of its center. As the cyclone develops in the lower stratosphere, it promotes the northward incursion of the Antarctic vortex up to about 55°S, along with air masses of highly depleted ozone levels.

1. Introduction

Maximum areas of short time-scale variability are widely known as storm tracks, which are closely related to the preferred track of transient disturbances (of both negative and positive geopotential height perturbation centers; Trenberth 1991). The storm-track maximum in the Southern Hemisphere (SH) is located over the southern Indian Ocean, and a relative minimum is typically found over the eastern South Pacific. Berbery and Vera (1996) showed that during the winter there is also a region of conspicuous high-frequency variability along the subtropical latitudes of the Pacific Ocean. They also showed that baroclinic waves evolve eastward in packets along the storm-track regions, in which downstream baroclinic development is part of their life cycle. During the winter season, storm-track distribu-
tion in southern South America (SSA) and adjacent oceans presents relative maximum values over both subtropical and midlatitudes. Vera et al. (2002) and Vigliarolo et al. (2001, hereafter VI01), showed that the leading modes of high-frequency variability over this region are characterized by baroclinic wave trains evolving to the east along upper-tropospheric subtropical and subpolar jet latitudes, although subpolar waves propagate faster and exhibit lower wavenumbers than the subtropical waves.

The activity of the transient disturbances in the SH may vary according to the season, in association with seasonal changes of the mean flow in which they are embedded (Trenberth 1991). While in August the maximum wind speeds are found in relation to polar-night jet maturity state, in October, the jet experiences a conspicuous deceleration due to the progressive warmth of the high-latitude temperatures as the summer approaches. The jet deceleration sets favorable conditions to allow the vertical propagation of waves of intermediate zonal wavenumbers. According to the Charney and Drazin (1961) criterion, total wavenumbers 5 or less at 45° of latitude are able to propagate vertically with westerlies lower than 15 m s\(^{-1}\) (James 1994). In this sense, during winter, subpolar baroclinic waves in the vicinity of SSA maximize close to the tropopause and do not extend much farther than the lowermost stratospheric levels (VI01). On the other hand, at least to our knowledge, there is a lack of literature about the characteristics of the transient variability over SSA in spring. In particular, questions such as the following still need to be answered: (a) Which are the most active synoptic-scale patterns over extratropical South America; (b) What characterizes their temporal variation and structure; and (c) Assuming that they are able to extend significantly upward into the lower stratosphere, what are the possible consequences for the stratospheric variability and ozone?

It is well known that baroclinic waves are characterized by potential vorticity anomalies near the tropopause, which typically decrease the tropopause height in regions of cyclonic vorticity and raise it where the vorticity is of anticyclonic sign (Hoskins et al. 1985, among others). Moreover, potential vorticity high-frequency variability and ozone are also well correlated near the tropopause (e.g., Reed and Danielsen 1959; Vaughan and Price 1991; Salby and Callaghan 1993). Since a great contribution to total-ozone content comes from the lower stratosphere (Salby and Callaghan 1993), the processes that tend to diminish the tropopause height also will tend to replace ozone-poor tropospheric air by ozone-rich stratospheric air, and thus total-ozone content will increase. Therefore, baroclinic wave ridges (troughs) will be associated with enhanced (diminished) total-ozone levels. This mechanism is suitable to explain ozone-column variability due to synoptic activity in regions of smooth mean ozone meridional gradients, as observed at midlatitudes in winter (e.g., VI01, and references therein). Nevertheless, during the SH spring, the ozone meridional gradient experiences a substantial increase south of 50°S, in association with high-latitude ozone-depleted levels (Figs. 1a,b). Therefore, during this season the relationship between midlatitude baroclinic perturbations and ozone may not develop as simply as described before.

Ozone and weather relationships have been largely investigated through individual case studies (e.g., Orlanski et al. 1989; McKenna et al. 1989; Rood et al. 1992; Orsolini et al. 1995). Mote et al. (1991) for the Northern Hemisphere and VI01 for the SH are some of the few papers that relate baroclinic wave evolution to ozone-column fluctuations during the winter season from a statistical point of view. Nevertheless, up to now there are no such studies for the spring.

Hence, the main objectives of this paper are the following: (i) to characterize high-frequency variability over SSA and adjacent regions during the spring months, and (ii) to address the role that baroclinic waves play in the total-ozone daily distribution in a season of strong mean total-ozone depletion and favored conditions to allow a greater vertical extent of waves. The organization of this paper is as follows: the next section describes datasets and methodology used; in section three, we introduce subpolar-mode main characteristics; while in the fourth and fifth sections we concentrate on ozone-column variations related to subpolar patterns. Finally, in the last section the main results are summarized and discussed.

2. Data and methodology

The dataset is based on 15 austral spring seasons (1 September–30 November), spanning from 1979 to 1993 of National Centers for Environmental Prediction–Department of Energy Atmospheric Model Intercomparison Project version-2 (NCEP–DOE AMIP-II) daily averaged reanalyses (hereafter R-2), given on a regular 2.5° × 2.5° grid resolution and in 17 vertical levels from 1000 to 10 hPa. This period was chosen because during 1979–93 total-ozone global measurements are continuous. The R-2 database is a version that improves upon the widely known NCEP–National Center for Atmospheric Research (NCAR) version-1 reanalyses (hereafter R-1; Kalnay et al. 1996) by fixing some known errors and by updating the parameterizations of the physical processes. The differences between both
datasets are referred to in Kanamitsu et al. (2002). In R-2, the misallocation of the Australian Surface Pressure Bogus Data for the Southern Hemisphere (PAOBS), which significantly affected the middle and high latitudes at synoptic time scales, is fixed, suggesting an improvement of daily analyses (Kistler et al. 2001). Nevertheless, R-2 is also subject to errors; perhaps one of the most relevant includes the use of an overturned north–south ozone zonal mean climatology. Some experiments suggested that this would have a minor impact on the quality of the dataset, despite the introduction of some errors to the radiative fluxes in the stratosphere (Kanamitsu et al. 2002; W. Ebizusaki 2002, personal communication). In this sense, we examined the synoptic-scale anomalous circulation pattern in the lower stratosphere during the austral spring season based on three different datasets: NCEP analyses, R-1, and R-2. We found no significant differences among them, and it is suggested that the ozone climatology misallocation does not affect any of the results from this study.

Daily ozone-column data are from the Total Ozone Mapping Spectrometer (TOMS) instrument, version 7, onboard Nimbus-7 and Meteor-3 satellites for 1979–93 (McPeters et al. 1996); whenever possible, Nimbus-7 data were preferred, especially regarding the overlap period (August 1991–April 1993). The horizontal resolution of total-ozone data is 1° in latitude and 1.25° in longitude. October mean ozone-mixing ratio data (Fig. 1a) correspond to the Halogen Occultation Experiment (HALOE), version 19, spanning the 1991–97 period, combined with polar measurements of the Microwave Limb Sounder (MLS) instrument, version 4, with a time period spanning November 1991–April 1993, both onboard the Upper Atmosphere Research Satellite (UARS). Details of the UARS HALOE and MLS ozone-mixing ratio dataset can be found in Fishbein et al. (1996), Park et al. (1996), and Randel et al. (1998).

The 300-hPa meridional wind perturbation time series was utilized to compute the spatial patterns of synoptic-scale waves over extratropical South America. For any grid point and any variable, the perturbation time series (denoted by a "/\)") was obtained by removing from the daily average values the corresponding climatological mean seasonal cycle (depicted by the climatological daily averages) and the interannual variability (calculated as the difference between the seasonal mean for each particular year and the climatological seasonal mean). Berbery and Vera (1996) have shown that no further data-filtering procedure is necessary as the meridional wind perturbation defined in that way concentrates most of its variability at synoptic scales. Like so, the calculations presented in this paper are not affected by the decreasing ozone trends observed in the 1980s.

An extended empirical orthogonal function (EEOF; Weare and Nasstrom 1982) technique for four 1-day lag units was applied to the 300-hPa \( v' \) temporal time series, over a domain enclosed by 80°–10°S and 130°–20°W. EEOF modes were computed from the correlation matrix. To maximize the local variance within the
domain, the EEOF modes were further rotated using the Varimax method (Horel 1981) and hereafter will be referred to as REEOFs. The choice of the effective number of eigenmodes to be rotated follows O’Lenic and Livezey (1988) criteria.

The principal REEOF modes explain 4.94%, 4.26%, 3.64%, and 3.63% of the total variance, respectively, with each individual variance being about one-fourth of what would be obtained from a traditional EOF analysis (Lau and Chan 1985). First and fourth leading REEOFs constitute a pair of modes that are in quadrature phase shift (and thus represent the propagation of the same local wave). The first REEOF pattern (Fig. 2) was chosen for this study as it better represents the baroclinic waves evolving along mid- and high latitudes over the subpolar jet location, and hence is termed the

Fig. 2. Positive composites for the subpolar mode, showing the time evolution from day −1 to day +2 of perturbations of geopotential height (contours) and temperature (shaded) at (left) 850 hPa, (middle) 300 hPa, and (right) 50 hPa. Heavy (light) shading represents positive (negative) temperature perturbation regions. Contour intervals are (a)–(d) 20 m and 0.5 K; (e)–(h) 40 m and 0.3 K; (i)–(l) 40 m and 0.5 K, respectively. All 0 contours have been omitted.
3. Subpolar mode

3.1. Horizontal and vertical structure

In the following section, only results for subpolar-mode positive composites are shown because of the high degree of symmetry with the corresponding negative ones. Figure 2 displays the spatial patterns and temporal evolution of temperature and geopotential height perturbations associated with the subpolar mode. In the lower troposphere (850 hPa), the left column shows the time evolution of an anticyclonic center over the southeast Pacific Ocean, and a cyclonic one just over SSA. The wave pattern has a characteristic length scale of about 5000 km (zonal wavenumber 4–5) and a phase speed of 8 m s\(^{-1}\). The center trajectories are mainly zonal and along 55°S over the Pacific Ocean, but once on the lee side of the Andes they detour to the northeast. Although SSA is a region in which meridional wave propagation is possible (Ambrizzi et al. 1995), such behavior could be also promoted by the interaction between transient perturbations and the orography (Gan and Rao 1994). Vera et al. (2002) showed that near the Andes, the lower-tropospheric portion of the synoptic wave tends to advance following the shape of the terrain, while exhibiting an equatorward migration on the lee side. This behavior is suggested in the evolution of the anticyclonic center from day +1 to day +2 (Figs. 2c,d).

Both the horizontal and vertical structure of the waves associated with the subpolar mode (Figs. 2, 3) display the main characteristics of midlatitude baroclinic waves (Lim and Wallace 1991). These are summarized as follows. (i) The anomalous temperature pattern peaks both at low-tropospheric and low-stratospheric levels (Figs. 2a–d, 2i–l, 3b), while it exhibits a sign change at the tropopause level (Figs. 2e–l, 3b). (ii) The geopotential height perturbation centers are tilted southwestward with altitude, while the temperature perturbation centers are tilted eastward (Figs. 2, 3b); this configuration is favorable for developing baroclinic waves through the intensification of poleward heat fluxes (Figs. 3b,c). (iii) The progressive upward increase of the geopotential height perturbation amplitudes to the east strongly suggests a vertical dispersion of energy (Fig. 3a; Lim and Wallace 1991).

A comparison between the waves associated with subpolar mode and those typical of the extratropical South American region during the austral winter season (VI01; Vera et al. 2002) show that they share essentially the same characteristics, although the latter are more affected by the presence of the Andes, which produces a strong modification of the wave structure at low levels. In addition, the spring subpolar waves attain their extremes at higher altitudes compared to their winter counterparts, extending far above the tropopause with prominent amplitudes in the lower stratosphere (Figs. 2i–l, 3a). Moreover, across the tropopause, they undergo a noticeable change in their horizontal structure, mainly through a lowering of their zonal wavenumber from 4–5 in the troposphere to 3–4 in the lower stratosphere (Fig. 2, left and middle columns, respectively). At 50 hPa, while the stratospheric portion of the cyclone is located just over or only a few degrees to the west of its upper-tropospheric counterpart, the anticyclone over the South Pacific displays, in contrast, a conspicuous shift to the southwest of about 10° (Fig. 2, middle and right columns).

3.2. Vertical propagation of the synoptic signal

The vertical structure of the vertical component of the energy flux (\(F_v\)) associated with subpolar mode has been considered as well (Fig. 3d). Following Orlanski and Sheldon (1993), the expression for \(F_v\) flux is

\[
F_v = \rho \omega \frac{\partial \phi}{\partial z}
\]
where $\rho$ stands for the density, $\omega'$ and $\phi'$ are the vertical velocity and geopotential perturbations, respectively, and the overbar represents the composite of cases for the subpolar mode. From Fig. 3d, it can be seen that maximum Fv fluxes are found at midtropospheric levels, with absolute values rounding 0.2 Pa m$^{-2}$ s$^{-3}$. In addition, Fv fluxes extend well above the tropopause, indicating that the eddy energy is flowing upward at the leading fringe of both the cyclonic and anticyclonic perturbations (where the sign is negative), from the surface to the lower stratosphere. Chang and Orlanski (1994) have shown that Fv divided by the energy represents a good approximation to the group velocity of wave packets, even if they are linear or not. Hence, Fig. 3d also suggests that the synoptic signal is propagating from the troposphere to the lower stratosphere, contributing to the growth of the waves above the tropopause. In this sense, we examined Hovmöller diagrams of composites of geopotential height perturbation averaged in 55°–70°S for the subpolar-mode extreme events (Fig. 4). An extreme event is defined when the absolute value of the subpolar-mode time series coefficient is larger than 1.8 times the standard deviation. Using this criterion, 29 statistically significant events were found for positive composites. However, it should be noted that considering all the cases that make up the positive composite set (101, with the extreme cases also included) provided almost identical results, although with lower amplitudes. By day +1, the main cyclonic disturbance at 300 hPa minimizes to ~210 m slightly east of 60°W (Fig. 4b, left and right panels), while it attains a minimum in the lower stratosphere one day after (~240 m near 40°W), supporting the notion of upward propagation of the synoptic wave signal.
same behavior is observed for the anticyclone, which the lower-stratospheric portion maximizes about 1 or 2 days after its upper-tropospheric counterpart.

According to Haynes and Shuckburgh (2000), the lowest boundary of the Antarctic polar vortex lays close to 400 K (100–120 hPa). Subpolar waves exhibiting large amplitudes at altitudes that concur with the lower vortex region has important implications for the Antarctic vortex dynamics and will be discussed in subsequent sections.

4. Total-ozone composites

In this section, the focus is on the analysis of composites for subpolar pattern of total ozone and its relationship with atmospheric fields. As mentioned in section 2, the 5-day time evolution depicted in Fig. 5 was obtained by averaging daily total-ozone perturbation fields over those days in which the REEOF temporal coefficients are larger than 0.8 times the standard deviation of the series. Figures 5a–c display the evolution of ozone-column perturbations from day −1 to day +1, when since the beginning of the sequence, a negative center is found near 57°S, 100°W and a positive one is developing over the Patagonian region. By day 0 (Fig. 5b), the negative total-ozone perturbation minimizes to
–26 Dobson units (DU) slightly west of 90°W, while the positive center also attaches a maximum of 16 DU off the coast of Argentina. These perturbation values are comparable to those observed for winter total-ozone disturbances over the same region (VI01). Also during day 0, a southeastward extension of the negative ozone center becomes apparent, extending well beyond the Antarctic Peninsula to the east and getting deeper the day after (Fig. 5c).

When considering individual spring months, we found that the positive composites for the subpolar mode of total-ozone perturbations display the largest amounts in October. Total-ozone perturbation composites for October, which include 34 cases (which represents about one-third of the total number of cases for the whole spring), show larger disturbances though they have almost the same pattern (cf. Fig. 6a with Fig. 5c). Hence, in what follows, discussion will be centered on the October positive composites for the subpolar mode.

By day +1 (Fig. 6a), the October ozone-column negative perturbations near SSA display a conspicuous southeastward lengthening, and the whole pattern resembles a sigmoid feature, with the upper branch lying over the southeastern Pacific and the bottom branch over the Weddell Sea. This feature in the total-ozone perturbation field is in contrast with what is observed during the winter, when total-ozone perturbations typically display shorter zonal scales (cf. Fig. 6a with Figs. 7 and 9 from VI01).

October composites of geopotential height perturbations at day +1 are shown as well (Figs. 6b,c) and should be compared with the composites for the whole spring months (Figs. 2g,k). At 300 hPa, ozone-column negative perturbations lying over the southeastern Pacific Ocean (light-gray-shaded regions in Fig. 6b) are associated, north of 65°S, with the anticyclonic center, while negative total-ozone values to the south of SSA and east of the Antarctic Peninsula are in relation to the western and southern portion of the cyclone downstream. While typically during spring geopotential height zonal wavenumbers 4–7 correlate negatively with ozone at the troposphere level over midlatitudes, south of 60°S, such correlation turns out to be positive and maximum over ∼50–20 hPa (Stanford et al. 1995), near the region where ozone partial pressure maximizes during that season. The positive correlation between ozone and geopotential height is not only due to synoptic disturbances but is also attributed to the presence of planetary waves very typical of the spring stratosphere (Wirth 1993).

Composites of 150-hPa temperature perturbations

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**Fig. 6.** October (positive) composites for the subpolar mode of perturbations of (a) ozone, (b) geopotential height at 300 hPa, (c) geopotential height at 50 hPa, and (d) temperature at 150 hPa. All perturbation fields are shown for day +1. Contour intervals correspond to (a) 4 DU; (b), (c) 30 m; and (d) 0.75 K with 0 contours omitted. In (b)–(d), light (heavy) gray shading represents negative (positive) ozone perturbation values.
during day +1 (Fig. 6d), show a pattern of remarkable correspondence with the negative total-ozone center, particularly to the north of 65°S, suggesting a high and positive correlation. This is in agreement with the results of Ziemke et al. (1997), who noted that ozone and temperature correlation remains positive above the tropopause almost all year round, with extremes near 100 hPa and over the 60°–40°S latitude band. Nevertheless, southward of 65°S, the ozone–temperature correlation diminishes, particularly regarding the polar side of the ozone negative center. The ratio between total ozone and 150-hPa temperature perturbations was of the order of 11 DU K⁻¹; while slightly larger, this value compares with the range of ratios 6–10 DU K⁻¹, resulting if total-ozone fluctuations are only forced by dynamics (Ziemke et al. 1997; Randel and Cobb 1994).

Ozone-column and temperature correlation is associated with both distributions of ozone-mixing ratio and potential temperature in the lower and midstratosphere. According to Wirth (1991), the highly positive correlation follows from the fact that surfaces of potential temperature and ozone-mixing ratio are approximately parallel, a situation that is observed for October at midlatitudes between 100 and 20 hPa (not shown). This would imply that lower-stratospheric temperature perturbations due to the dynamics of synoptic waves could be used as proxies to infer the behavior of the total-ozone anomalies in the region. On the other hand, south of 65°–60°S ozone-mixing ratio surfaces undergo an abrupt upward migration in relation to high-latitude low-ozone content (Fig. 1a), with the result that they are no longer congruent with potential temperature surfaces; this explains why ozone and temperature correlation diminishes over that region.

5. Potential vorticity outlook

Potential vorticity on isentropic coordinates (hereafter PV), which is negative in the SH, is associated with upper-tropospheric synoptic disturbances (Orlanski and Katzfey 1991). Areas with minimum PV are typically found along with air masses with cyclonic (negative) relative vorticity in pressure coordinates, while negative but with locally higher PV amounts are associated with the presence of an anticyclone. Furthermore, the PV field also experiences a conspicuous variation with latitude and height, with minimum values at polar latitudes and higher stratospheric levels, due to the static stability parameter that dominates over the absolute vorticity distribution. Therefore, within one isentropic level above the tropopause, regions with significant negative PV may be suggesting the presence of air masses from higher in the stratosphere.

Figure 7 shows October composites for the subpolar mode of PV, both at the 360-K isentropic level (which corresponds to approximately 170 hPa at midlatitudes) and at 490 K (~60 hPa). By day −2 in the lowest stratosphere (Fig. 7a), the local maximum of PV over 35°–55°S, 120°–100°W denotes the presence of the anticyclonic perturbation. Immediately downstream, a tongue of PV minimum values along 80°W corresponds to the location of the cyclonic disturbance. Two days later (Fig. 7b), the trough intensifies and experiences a meridional elongation off the east coast of Patagonia, reaching a peak value of −13 PVU over the Weddell Sea, in agreement with the cyclonic behavior depicted in Fig. 6b.

It is well known that ozone and PV are highly correlated fields in the lower stratosphere (Vaughan and Price 1991; Salby and Callaghan 1993; among others); in this sense, composites of total-ozone have been considered as well and are shown as shaded regions in Fig. 7. During lag −2 (top panels of Fig. 7), the total-ozone field displays minimum values from the south of the Antarctic Peninsula through midlatitudes of the southwestern Pacific Ocean, with a clear NW–SE direction. At 360 K (Fig. 7a), the total-ozone configuration is associated, in its northwestern extreme with the PV ridge, while its southeast section concurs with the location of the PV trough, with the axis of minimum total ozone being across the PV gradient. By day +1 (Fig. 7b), an overall decrease in total-ozone levels is observed over 45°–60°S, 100°–70°W, in the southeast Pacific, and from central Patagonia to the Weddell Sea, which reflects the sigmoid feature of ozone-column negative anomalies (Fig. 6a).

The occurrence of the total-ozone minimum in relation to the presence of the Pacific anticyclone could be understood as follows. In the midlatitude lower stratosphere, the ozone-mixing ratio is nearly conserved following the motion for time scales of 1 or 2 weeks (Solomon and García 1985; Salby and Callaghan 1993; among others). Moreover, also in this atmospheric section for synoptic time scales, the flow could be considered adiabatic and frictionless, leading to PV conservation (Andrews et al. 1987). When air parcels approach the PV ridge, they typically experience uplift, adiabatic expansion, and cooling, which in turn leads to a decrease of its ozone number density in the lower stratosphere (Perez et al. 2000).

To examine the climatological characteristics of both polar stratospheric vortices, Waugh and Randel (1999) utilized the elliptical diagnosis applied to PV data, which enables us to define some useful vortex parameters as well as to quantify their structure and evolution. In this sense, the Antarctic vortex boundary

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shown throughout was calculated first considering the “modified” PV contours used by those authors to define the vortex edge, then converted to “real” PV values, and, last, interpolated onto the isentropic levels used in this study. At 490 K (right panels, Fig. 7), the −53 PVU contour denotes the vortex edge (note that the vortex boundary defined in that way also lies close to the maximum gradient of PV), with strongly negative PV inside the vortex. By day +1, the intensification of the trough at 490 K seems to promote the northward displacement of the lower stratospheric vortex to the east of SSA (Fig. 7d). As the mean position of the vortex boundary roughly coincides with the maximum mean ozone-mixing ratio meridional gradient up to about 30–40 hPa (Fig. 1a), an equatorward displacement of the lower-stratospheric vortex could bring about highly ozone-depleted air masses. In the subpolar composites, the lower portion of the vortex is able to reach 55°S over the southwestern Atlantic, and the associated low total-ozone levels are evident (Fig. 7d). In agreement, Perez et al. (2000) reported two episodes of reduced ozone levels over SSA caused by the advance of the low-stratospheric Antarctic vortex over that region.

Figure 8 displays longitude– and latitude–height sections in isentropic coordinates for both PV and PV perturbation composites. The longitude–height section, at 60°S (Fig. 8a), shows that by day +1, total-ozone negative perturbation extends from 120°W to the date line (as seen by comparison with Fig. 6a) with the minimum over 60°W, slightly on the cyclonic side of the baroclinic wave couplet. As mentioned before, to the west (and north) the ozone-column negative anomalies are related to the anticyclone over 100°–80°W, with positive anomalies of PV from 380 to 600 K. In addition, in the downstream side, negative PV anomalies coincide with the downward displaced PV surfaces, with minimum values of both PV and PV′ inside the vortex area (Fig. 8a). The large negative values of PV anomalies relate to the cyclonic disturbance, which displays large amplitudes even up to 550 K (~40 hPa) as observed in Fig. 3a. Hence, the cyclonic development is associated with the equatorward penetration of the low-stratospheric vortex, which in turn brings about high-latitude depleted ozone, as seen from the presence of total-ozone negative anomalies from about 70°W eastward to 0°. A latitude–height perspective along 50°W (Fig. 8b) shows that the minimum PV values at high latitudes remain inside the vortex, while the vortex boundary is located along the axis of minimum PV anomalies. Furthermore, in the lower stratosphere, the meridional advance of the vortex north of 55°S becomes evident, in agreement with Fig. 7d. In this cross section, the ozone-column
minimum is found somewhat southward of 60°S, inside the vortex area.

6. Discussion and conclusions

A statistical–dynamical diagnosis of the high-frequency variability over SSA and adjacent regions during the austral spring is presented in this paper, as well as a discussion about the role such baroclinic waves play in the total-ozone daily distribution.

Main patterns of synoptic-scale variability were identified by means of REEOFs techniques applied to the 300-hPa meridional wind perturbation time series, over a domain enclosed by 80°–10°S and 130°–20°W. The first REEOF describes a wave train that propagates to the east along upper-tropospheric subpolar jet latitudes (55°–60°S) and hence is termed the subpolar mode. Composites for the subpolar mode (based on subpolar-mode temporal coefficients larger than a threshold) show the typical structure of midlatitude baroclinic waves. The waves, which maximize in the upper troposphere and lower stratosphere, display significant amplitudes up to about 40 hPa.

The temporal evolution of the composite waves showed that they reach maximum amplitude first in the upper troposphere and 1–2 days later in the lower stratosphere. This fact, together with the corresponding energy flux pattern, suggests an upward propagation of the synoptic signal from the troposphere (where the waves originate) deep into the lower stratosphere.

Although winter and spring subpolar modes (as analyzed in VI01 and in this paper, respectively) display most of the same characteristics of midlatitude baroclinic waves, only during spring are baroclinic waves able to extend upward into the lower stratosphere. In that region, they could further influence the dynamics of the lower-stratospheric polar vortex, promoting its meridional lengthening over SSA. Therefore, in this transition season the relationship between baroclinic waves and ozone seems not to be as direct as that observed during winter. The analysis of October composites for the subpolar mode (which is the month that displays the largest total-ozone anomalies among the other spring ones) reveals that ozone-column negative fluctuations are related to the location of an anticyclone–cyclone system near SSA. The basic features of

Fig. 8. (a) Longitude–height cross section at 60°S, in isentropic coordinates, of October (positive) composites for the subpolar mode of potential vorticity perturbation (heavy black contours) and potential vorticity (light gray contours and gray shading), during day +1. (b) As in (a), but for the latitude–height cross section at 50°W. The PV contour interval corresponds to 25 PVU (drawn from −25 to −375 PVU). For clarity, only selected PV contours are shown, ranging from −3 to −75 PVU. The thickest black solid contours designate the location of the vortex boundary. Continuous black contours near the bottom of the panels denote the approximate position of the tropopause, as given by the −2.5, −2, and −1.5 PVU contours (James 1994). Capital letters indicate the location of the anticyclonic (A) and cyclonic (C) disturbances. Horizontal thick dashed lines near 350 K stand for the approximate extension of the ozone negative perturbation (see Fig. 6a), and the arrows point out the location of the ozone minimum during that day.
the aforementioned relationship are summarized in Fig. 9 as follows.

North of 65°S, the development of the anticyclonic perturbation center in the lower stratosphere is accompanied by both upward-sloped surfaces of potential temperature and potential vorticity. As air parcels moving along isentropes approach the anticyclone, they experience uplift, adiabatic expansion, and a decrease of their ozone number density—resulting in the lifting of the ozone-mixing ratio isopleths above the anticyclone. As a result, total-ozone anomalies become negative over the northwestern portion of the anticyclone. It is noteworthy that the same mechanism has been proposed to account for total-ozone fluctuations in relation to baroclinic waves during winter, as discussed by VI01 and references therein.

Meanwhile downstream, the outstanding cyclonic development throughout the upper troposphere and lower stratosphere is accompanied by strongly negative PV amounts, indicating the meridional advance of the lower-stratospheric vortex. This also brings about highly depleted ozone-column levels, as across the vortex boundary and inside, ozone-mixing ratio surfaces slope abruptly upward. Therefore, negative total-ozone perturbations are generated over the southeastern section of the cyclone. In conclusion, the entire configuration for the total-ozone anomalies appears as a sigmoid pattern, with a minimum between the anticyclone and cyclonic centers.

The results presented in this paper show the prominent role of spring midlatitude baroclinic waves in driving total-ozone changes. In this sense, taking into account the record low-ozone mean levels in recent years, one could speculate with the possibility of the referred waves to promote the incursion of the ozone hole into permanently inhabited areas of SSA, a topic that is currently being researched.

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