Signatures of New Physics from HBT Correlations in UHECRs

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Abstract

Quantum fields written on noncommutative spacetime (Groenewold - Moyal plane) obey twisted commutation relations. In this paper we show that these twisted commutation relations result in Hanbury-Brown Twiss (HBT) correlations that are distinct from that for ordinary bosonic or fermionic fields, and hence can provide us useful information about underlying noncommutative nature of spacetime. The deviation from usual bosonic/fermionic statistics becomes pronounced at high energies, suggesting that a natural place is to look at Ultra High Energy Cosmic Rays (UHECRs). Since the HBT correlations are sensitive only to the statistics of the particles, observations done with UHECRs are capable of providing unambiguous signatures of noncommutativity, without any detailed knowledge of the mechanism and source of origin of UHECRs.

Cosmic Rays with energies around $10^{18}$ eV and higher are called as Ultra High Energy Cosmic Rays (UHECRs) [1–3]. They are the highest energy particles known to us and often have energies $10^7$ times more than that produced by LHC. Inspite of recent advancements (both theoretical and experimental), UHECRs pose a considerable theoretical challenge: the source and mechanism of origin of such high energy particles [4, 5] and their composition are areas of active research [6–8].

Due to their extremely high energies, UHECRs are not only an excellent arena for testing the validity of known laws of physics [9,10] but are also some of the best places to look for signatures, if any, of new physics e.g. theories with Lorentz violation and/or deformed dispersion relations [11–15]. In this paper we aim to show that UHECRs can be used to look for signatures of a particular model of nonlocalities coming from the underlying noncommutative nature of spacetime at short distances.

Simple intuitive arguments involving standard quantum mechanics uncertainty relations suggest that at length scales close to Plank length, strong gravity effects will limit the spatial as well as temporal resolution beyond some fundamental length scale ($L \approx$ Planck Length), leading to space - space as well as space - time uncertainties [16]. One cannot probe spacetime with a resolution below this scale i.e. spacetime becomes fuzzy below this scale, resulting into noncommutative spacetime. Hence it becomes important and interesting to study in detail the structure of such a noncommutative spacetime and the properties of quantum fields written on it, because it not only helps us improve our understanding of the Planck scale physics but also helps in bridging standard particle physics with physics at Planck scale.

This noncommutative scale need not always be as small as Planck scale. For instance if there are large extra dimensions, then the “effective Planck scale” can be at much lower energies (usually taken between TeV scale and GUT scale, depending on specific models), resulting in appearance of noncommutativity at scales much larger than Planck scale. Such large scale noncommutativity is of particular interest as its effect can be detected in present day or near future experiments. In this paper we are interested in looking for signatures of such large scale noncommutativity in UHECRs.

Of the various approaches to model the noncommutative structure of spacetime, the simplest is one where the coordinates satisfy commutation relations of the form

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\[ [\hat{x}_\mu, \hat{x}_\nu] = i\theta_{\mu\nu} ; \quad \mu, \nu = 0, 1, 2, 3 \quad \text{and} \quad \theta \text{ is a real, constant, antisymmetric matrix.} \quad (1) \]

The elements of the $\theta$ matrix have the dimension of $(\text{length})^2$ and set the scale for the area of the smallest possible localization in the $\mu - \nu$ plane, giving a measure for the strength of noncommutativity \[17\]. The algebra generated by $\hat{x}_\mu$ is usually referred to as Groenewold - Moyal (G.M) plane \[18\], and in this paper we will restrict our attention to this noncommutative spacetime. Equivalently this noncommutative nature of spacetime can be taken into account by defining a new type of multiplication rule ($\ast$ product) to multiply functions evaluated at same point i.e.

\[ f(x) \ast g(x) = f(x)e^{\frac{i}{2} \theta_{\mu\nu} \partial_\mu \partial_\nu} g(x). \quad (2) \]

One particularly important feature of G.M plane, which makes it quite suitable for writing quantum field theories on it, is the restoration of Poincaré-Hopf symmetry as Hopf algebraic symmetry, by defining a new coproduct (twisted coproduct) for action of Poincaré group elements on state vectors \[19\] \[20\] \[21\].

Twisting of the coproduct has immediate implications for the symmetries of multi-particle wave functions describing identical particles \[22\]. For example, on G.M plane the correct physical two-particle wave functions are twisted (anti-)symmetrized ones and are given by

\[
\begin{align*}
\phi \otimes_{S_\theta} \psi & \equiv \left( \frac{1 + \tau_\theta}{2} \right) (\phi \otimes \psi) \\
\phi \otimes_{A_\theta} \psi & \equiv \left( \frac{1 - \tau_\theta}{2} \right) (\phi \otimes \psi)
\end{align*}
\]

where $\phi$ and $\psi$ are single particle wavefunctions (of two identical particles) and $\tau_\theta$ is the twisted statistics (flip) operator associated with exchange given by

\[ \tau_\theta = F^{-1} \tau_0 F, \quad F = e^{\frac{i}{2} \theta_{\mu\nu} \partial_\mu \partial_\nu}. \quad (4) \]

Here $\tau_0$ is the commutative flip operator : $\tau_0 (\phi \otimes \psi) = \psi \otimes \phi$.

The above analysis can also be extended to field theories on G.M plane resulting in changed commutation relations between creation/annihilation operators \[22\], which now become

\[
\begin{align*}
a_p a^\dagger_q & = \eta e^{ip \cdot q} a_p a_p \quad \text{and} \quad a_p a^\dagger_q = \eta e^{-ip \cdot q} a^\dagger_q a^\dagger_p \\
\text{with} \quad p^\mu = (p^0, \mathbf{p}), \quad p \wedge q = p_\mu \theta_{\mu\nu} q_\nu, \quad \text{and} \quad \eta = \pm 1 \quad \text{for bosons/fermions.} \quad (5)\end{align*}
\]

In literature there exist another approach to quantization of noncommutative fields \[23\]. In this approach, the quantization is done according to the usual rules and the quantum fields follow usual bosonic/fermionic statistics. However, such a quantization scheme does not preserve the classical twisted Poincaré invariance and suffers from UV/IR mixing \[24\]. In this paper we only discuss the twisted quantization as it preserves twisted Poincaré invariance in noncommutative quantum field theories.

Because of \[19\] the quantum fields written on G.M plane, unlike ordinary quantum fields, follow an unusual statistics which we call as twisted statistics. Twisted statistics are a unique feature of fields on G.M plane and can be used to search for signals of noncommutativity: because of the twisted commutation relations, interesting new effects like Pauli forbidden transitions \[25\] \[26\] can arise. The effect of twisted statistics also manifests itself in certain thermodynamic quantities \[27\] \[28\]. The two-point distribution functions remain unchanged

\[ \langle a^\dagger_{p_1} a_{p_2} \rangle = 2 \langle p_1 \rangle_0 N^{(T)}_{p_1} \delta^3(p_1 - p_2) \]

where \[ N^{(T)}_{p} = \frac{1}{e^{-\beta p - \eta}} \] is the thermal distribution, but for example the quantity \[ \langle a^\dagger_{p_1} a^\dagger_{p_2} a_{p_3} a_{p_4} \rangle \] gets changed \[27\].
\[ \langle a_{p_1}^\dagger a_{p_2}^\dagger a_{p_3} a_{p_4} \rangle = 2 (p_1)_0 2 (p_2)_0 N_{p_1}^{(T)} N_{p_2}^{(T)} [ \delta^3(p_1 - p_4) \delta^3(p_2 - p_3) + \eta e^{i p_1 \cdot p_2} \delta^3(p_1 - p_3) \delta^3(p_2 - p_4) ] \]  

(7)

The above differs from the commutative expression by the appearance of the factor \( e^{i p_1 \cdot p_2} \) in the second term.

One can easily check, following an analysis similar to that done in [27], that (6) and (7) are true not only for thermal distribution \( N_p^{(T)} \) but for any arbitrary wavepacket \( f(p) \) i.e.

\[ \langle a_{p_1}^\dagger a_{p_2} \rangle = 2 (p_1)_0 f(p_1) \delta^3(p_1 - p_2) \]  

(8)

and

\[ \langle a_{p_1}^\dagger a_{p_2}^\dagger a_{p_3} a_{p_4} \rangle = 2 (p_1)_0 2 (p_2)_0 f(p_1) f(p_2) [ \delta^3(p_1 - p_4) \delta^3(p_2 - p_3) + \eta e^{i p_1 \cdot p_2} \delta^3(p_1 - p_3) \delta^3(p_2 - p_4) ] . \]  

(9)

Here we discuss the consequences of (9) to the HBT correlation functions.

Hanbury-Brown Twiss (HBT) effect [29] is the interference effect between intensities measured by two detectors when a beam of identical particles is projected on them, with the intensities recorded by the two detectors operating simultaneously. This intensity increases for bosons (and decreases for fermions) when compared with the intensities recorded by the same two detectors, if only one is operated at a time. The correlation function for HBT effect is defined as

\[ C = \frac{\langle I_1 \cdot I_2 \rangle}{\langle I_1 \rangle \langle I_2 \rangle} ; \text{ where} \]

\[ I_1, I_2 = \text{intensities recorded by the two detectors respectively, when both are operated simultaneously.} \]

\[ I_1' = \text{intensity recorded by first detector when the second detector is not operating.} \]

\[ I_2' = \text{intensity recorded by second detector when the first detector is not operating.} \]

The HBT correlation function \( C \) obeys

\[ C = 1 \quad \text{for distinguishable particles} \]

\[ C > 1 \quad \text{for bosons (bunching effect)} \]

\[ C < 1 \quad \text{for fermions (anti-bunching effect)} \]  

(11)

In the commutative case, for a beam of identical (massless, scalar) bosons, the HBT correlation function can be written as [30][31]

\[ C^{(B)}_0 = \frac{\langle \phi^{(-)}_0(y_1) \phi^{(-)}_0(y_2) \phi^{(+)}_0(y_2) \phi^{(+)}_0(y_1) \rangle}{\langle \phi^{(-)}_0(y_1) \phi^{(+)}_0(y_1) \rangle \langle \phi^{(-)}_0(y_2) \phi^{(+)}_0(y_2) \rangle} \]  

(12)

where \( y_1 \) and \( y_2 \) are the position of the two detectors, \( \phi^{(+)}_0(y) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{2Ep} c_p e^{-ip \cdot y} \) is the positive frequency part and \( \phi^{(-)}_0(y) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{2Ep} c_p^\dagger e^{-ip \cdot y} \) is the negative frequency part of the bosonic quantum field \( \phi_0 \). Notice that we have deliberately taken massless fields because we are interested in looking for HBT correlation functions of ultra-relativistic particles.

To account for the uncertainties in the energy measurements, which at present are quite significant for UHE-CRs, the incoming beam should be taken as a wavepacket, instead of plane waves. The choice of an appropriate

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1 Scalar bosons are taken for sake of simplicity but the analysis presented here can be easily generalized to higher spin bosons.

2 We denote the usual bosonic (fermionic) fields by \( \phi_0 (\psi_0) \) and their twisted counterparts are denoted by \( \phi_y (\psi_y) \). Also, the usual creation/annihilation operators are denoted by \( a_p \) and \( a_p^\dagger \) whereas the twisted ones are denoted by \( a_y^\dagger \) and \( a_y \) respectively.
wavepacket is potentially the only place where the information about the details of production mechanism or source of origin of UHECRs can reside. Central Limit Theorem tells us that the mean of a sufficiently large number of independent random variables, each with finite mean and variance, will be approximately distributed like a Gaussian, and hence it is a good first approximation to take the wavepacket as a Gaussian wavepacket \( f(p) = Ne^{-\alpha(p-p_0)^2} \) centered around some mean momentum \( p_0 \).

Taking the wavepacket to be this Gaussian, restricting ourself to only coincidence measurements and using the standard integrals \([32]\), the correlation function turns out to be

\[
C_0^{(B)} = 1 + e^{-\frac{\gamma^2}{2}}
\]

(13)

As clear from \([13]\), \( C_0^{(B)} \) depends only on \( y = |y_1 - y_2| \) the separation between detectors and on \( \gamma \). There is no dependence on the mean momentum \( p_0 \) \([33,34]\).

Similarly, for a beam of identical (chiral) fermions, the HBT correlation function can be written as

\[
C_0^{(F)} = \frac{\langle \psi_0^{(-)}(y_1)\gamma \psi_0^{(-)}(y_2)\gamma_0 \frac{1}{2}(1 \pm \gamma_5) \psi_0^{(+)}(y_2)\frac{1}{2}(1 \pm \gamma_5) \psi_0^{(+)}(y_1) \rangle}{\langle \psi_0^{(-)}(y_1)\gamma_0 \frac{1}{2}(1 \pm \gamma_5) \psi_0^{(+)}(y_1) \rangle \langle \psi_0^{(-)}(y_2)\gamma_0 \frac{1}{2}(1 \pm \gamma_5) \psi_0^{(+)}(y_2) \rangle}
\]

(14)

where, as before, \( y_1 \) and \( y_2 \) are the position of the two detectors, \( \psi_0^{(+)}(y) = \int \frac{dp}{(2\pi)^2} \frac{1}{2E_p} \sum_s u_{s,p} e^{ip \cdot y} \) is the positive frequency part of the fermionic quantum field \( \psi_0 \)

Since the whole analysis is done keeping ultra-relativistic particles in back of our mind, here we look for correlation between only chiral fermions (left-left or right-right): at such high energies, the particles are effectively massless and hence chiral fermions are the more appropriate ones to deal with.

Taking the incoming beam as a Gaussian wavepacket \( f(p) = Ne^{-\alpha(p-p_0)^2} \) centered around some mean momentum \( p_0 \) and restricting ourself to only coincidence measurements, the correlation function turns out to be

\[
C_0^{(F)} = 1 - e^{-\frac{\gamma^2}{2}} \frac{2}{9\pi\alpha} e^{-2\alpha p_0^2} (y^2 + 4\alpha^2 p_0^2) _1F_1 \left[ 2; \frac{5}{2}; \frac{1}{4\alpha} (y^2 - 4\alpha^2 p_0^2 - 4i\alpha y \cdot p_0) \right]
\]

(15)

where \( y = y_2 - y_1 \) is the separation between the two detectors and \( _1F_1 (\alpha; \gamma; z) \) is the degenerate hypergeometric function.

We observe that in this case, the correlation function depends on the separation between detectors \( y \), on the mean momentum \( p_0 \) of the wavepacket and the angles between \( y \) and \( p_0 \).

In the noncommutative case, there are two important differences. Firstly, the \((\cdot)\) product between fields evaluated at the same point has to be replaced by \((\ast)\) product and secondly the expectation value is changed in accordance with \([7]\), as the quantum fields are now composed of twisted creation/annihilation operators. Hence in noncommutative case for twisted (massless, scalar) bosonic particles, we have

\[
C_0^{(B)} = \frac{\langle \phi(-) \end{phi}(y_1) \langle \phi(-) \end{phi}(y_2) \ast y_2 \phi(+)(y_2) \rangle \ast y_1 \phi(+)(y_1) \rangle}{\langle \phi(-) \end{phi}(y_1) \ast y_1 \phi(+)(y_1) \rangle \langle \phi(-) \end{phi}(y_2) \ast y_2 \phi(+)(y_2) \rangle}
\]

(16)

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\[C_0^{(F)} = 1 \] between particles with opposite helicities (i.e. between left-right or right-left), as at ultra-relativistic energies, they behave like distinguishable particles.
where $\star_y = e^{\frac{i}{2} \left( \frac{\partial}{\partial \eta} \right)_\mu \theta^{*\mu} \left( \frac{\partial}{\partial \eta} \right)_\nu}$. 

In rest of the paper we restrict to only space-space noncommutativity i.e. we take $\theta_{0i} = \theta_{i0} = 0$. As in commutative case, we take the incoming beam as a Gaussian wavepacket $f(p) = N e^{-\alpha (p - \mathbf{p}_0)^2}$ centered around some mean momentum $\mathbf{p}_0$ and restrict ourselves to only coincidence measurements. $C_\theta^{(B)}$ then turns out to be

$$C_\theta^{(B)} = 1 + \frac{4\alpha}{4\alpha^2 + \lambda^2} \exp \left[ -\frac{4\alpha^2 y^2 - (y \cdot \lambda)^2}{2\alpha(4\alpha^2 + \lambda^2)} \right] \exp \left[ -\frac{2\alpha \mathbf{p}_0^2 \lambda^2 - (\mathbf{p}_0 \cdot \lambda)^2}{4\alpha^2 + \lambda^2} \right] \exp \left[ -\frac{4\alpha y \cdot (\mathbf{p}_0 \times \lambda)}{4\alpha^2 + \lambda^2} \right]$$

(17)

where we have defined $\theta_{ij} = \varepsilon_{ijk} \lambda_k$. In getting (17) we have used the standard result that, for $n$-dim column matrices $X$ and $B$ and a $n \times n$ positive definite, symmetric square matrix $A$, we have

$$\int d^n X e^{-X^T A X + B^T X} = \left( \frac{\pi^n}{\det A} \right)^\frac{1}{2} e^{\frac{1}{4} B^T A^{-1} B}$$

(18)

We see that (17) not only depends on $y$ (the separation between detectors) but also on the mean momentum $\mathbf{p}_0$, the noncommutative length scale $\lambda$, as well as on the angles between $\lambda$ and $y$ and between $\lambda$ and $\mathbf{p}_0$. Moreover, as a check we can see that, in the limit $\lambda \to 0$ we get back (13).

Similarly, for twisted (chiral) fermionic particles the noncommutative HBT correlation function is given by

$$C_\theta^{(F)} = \frac{\langle \Psi^(-)(\mathbf{y}) \gamma_0 \left( \Psi^(-)(\mathbf{y}) \gamma_0 * \mathbf{y} \gamma_0 \frac{1}{2}(1 \pm \gamma_5) \psi^+(\mathbf{y}) \right) \star_y \frac{1}{2}(1 \pm \gamma_5) \psi^+(\mathbf{y}) \rangle_{\theta} \gamma_0 \Psi^(-)(\mathbf{y}) \gamma_0 * \mathbf{y} \gamma_0 \frac{1}{2}(1 \pm \gamma_5) \psi^+(\mathbf{y}) \rangle_{\theta}}{\langle \Psi^(-)(\mathbf{y}) \gamma_0 * \mathbf{y} \gamma_0 \frac{1}{2}(1 \pm \gamma_5) \psi^+(\mathbf{y}) \rangle_{\theta} \langle \Psi^(-)(\mathbf{y}) \gamma_0 * \mathbf{y} \gamma_0 \frac{1}{2}(1 \pm \gamma_5) \psi^+(\mathbf{y}) \rangle_{\theta}}$$

(19)

Again restricting ourselves to only space-space noncommutativity, using Gaussian wavepackets and considering only coincidence measurements, $C_\theta^{(F)}$ can be written as

$$C_\theta^{(F)} = 1 - \frac{1}{2} e^{-2\alpha \mathbf{p}_0^2} \left[ e^{-\frac{\mathbf{y}^2}{2\alpha} - \frac{(\mathbf{y} \cdot \lambda)^2}{8\alpha^3} - \frac{(\mathbf{y} \cdot \lambda)^2}{8\alpha^3} - \frac{(\mathbf{p}_0 \cdot \lambda)^2}{2\alpha} + \frac{(\mathbf{p}_0 \cdot \lambda)^2}{2\alpha} + \frac{(\mathbf{p}_0 \cdot \mathbf{y})^2}{2\alpha^2}} \right] - \left( \frac{2}{9\pi \alpha} \right) e^{-2\alpha \mathbf{p}_0^2} \left[ \left\{ F_1 \left( 2, \frac{5}{2}, -\frac{z_1^2}{4\alpha} \right) \right\} \right]$$

(20)

where $z_1 = y - 2i \alpha \mathbf{p}_0$ and $z_2 = y + 2i \alpha \mathbf{p}_0$.

Expanding this in $\theta$ and taking terms up to second order we get

$$C_\theta^{(F)} = 1 - \frac{1}{2} e^{-\frac{\mathbf{y}^2}{2\alpha} \left[ 1 + \frac{\mathbf{p}_0 \cdot \mathbf{y} \cdot \lambda}{\alpha} - \frac{\mathbf{y}^2 \lambda^2}{4\alpha^2} + \frac{(\mathbf{y} \cdot \lambda)^2}{8\alpha^3} - \frac{(\mathbf{p}_0 \cdot \lambda)^2}{2\alpha} + \frac{(\mathbf{p}_0 \cdot \mathbf{y})^2}{2\alpha} + \frac{(\mathbf{p}_0 \cdot \mathbf{y})^2}{2\alpha} + \frac{(\mathbf{p}_0 \cdot \mathbf{y})^2}{2\alpha^2} \right] - \left( \frac{2}{9\pi \alpha} \right) e^{-2\alpha \mathbf{p}_0^2} \left[ \left\{ F_1 \left( 2, \frac{5}{2}, -\frac{z_1^2}{4\alpha} \right) \right\} F_1 \left( 2, \frac{5}{2}, -\frac{z_2^2}{4\alpha} \right) \right]$$

$$\left( \frac{\mathbf{y}^2 + 4\alpha \mathbf{p}_0^2}{25\alpha^2} \right) + \left( \frac{4}{25\alpha^2} \right) \left\{ \left( \frac{4\mathbf{p}_0 \cdot \mathbf{y} \cdot \lambda}{\alpha} - 2\lambda^2 \right) \left( \frac{\mathbf{y}^2 + 4\alpha \mathbf{p}_0^2}{25\alpha^2} \right) + \left( \frac{\mathbf{p}_0 \cdot \lambda}{\alpha} \right) \right\}$$

$$+ \left[ \frac{6}{11\alpha \mathbf{p}_0^2} \left( \frac{\mathbf{y}^2 \lambda^2 - 4\alpha \mathbf{p}_0^2 \lambda^2 - 4\alpha \mathbf{p}_0 \cdot \mathbf{y} \lambda^2 - (\mathbf{y} \cdot \lambda)^2 + 4\alpha \mathbf{p}_0 \cdot \lambda \right) \right]$$

$$\left( \frac{\mathbf{y}^2 + 4\alpha \mathbf{p}_0^2}{25\alpha^2} \right) F_1 \left( 4, \frac{9}{2}, -\frac{1}{4\alpha} \left( \frac{\mathbf{y}^2 - 4\alpha \mathbf{p}_0^2 - 4\alpha \mathbf{y} \cdot \mathbf{p}_0}{4\alpha^2} \right) \right) F_1 \left( 3, \frac{7}{2}, -\frac{1}{4\alpha} \left( \frac{\mathbf{y}^2 - 4\alpha \mathbf{p}_0^2 - 4\alpha \mathbf{y} \cdot \mathbf{p}_0}{4\alpha^2} \right) \right)$$

$$+ \left( \frac{6}{11\alpha \mathbf{p}_0^2} \right) \left( \frac{\mathbf{y}^2 \lambda^2 - 4\alpha \mathbf{p}_0^2 \lambda^2 + 4\alpha \mathbf{p}_0 \cdot \mathbf{y} \lambda^2 - (\mathbf{y} \cdot \lambda)^2 + 4\alpha \mathbf{p}_0 \cdot \lambda \right)$$

$$\left( \frac{\mathbf{y}^2 + 4\alpha \mathbf{p}_0^2}{25\alpha^2} \right) F_1 \left( 3, \frac{7}{2}, -\frac{1}{4\alpha} \left( \frac{\mathbf{y}^2 - 4\alpha \mathbf{p}_0^2 - 4\alpha \mathbf{y} \cdot \mathbf{p}_0}{4\alpha^2} \right) \right)$$

(10)

We have assumed only space-space noncommutativity for calculational simplicity but one can do similar analysis with both space-time as well as space-space noncommutativity.
\[ \begin{align*}
1F_1 \left( 4; \frac{9}{2}; -\frac{1}{4\alpha} \right) & - \frac{288}{1225\alpha^2} \left( \mathbf{p}_0 \cdot (\mathbf{y} \times \lambda) \right)^2 \left( \mathbf{y}^2 + 4\alpha^2 \mathbf{p}_0^2 \right) \\
1F_1 \left( 4; \frac{9}{2}; -\frac{1}{4\alpha} \right) & \left( \mathbf{y}^2 - 4\alpha^2 \mathbf{p}_0^2 - 4i\alpha \mathbf{y} \cdot \mathbf{p}_0 \right) \right) \right] + O(\lambda^3) \quad (21)
\end{align*} \]

where again in writing (21) we have defined \( \theta_{ij} = \varepsilon_{ijk} \lambda_k \).

As can be clearly seen from (21), the noncommutative HBT correlation function not only depends on \( \mathbf{y} \) and \( \mathbf{p}_0 \) but also on \( \lambda \) and various angles that \( \lambda \) makes with \( \mathbf{y} \) and \( \mathbf{p}_0 \).

A comparison of (13)-(17) and (15)-(21) tells us that the HBT correlation function for twisted bosons/fermions are different from those for ordinary bosons/fermions and become more pronounced with increasing momenta of incident particles. Since \( \lambda \) is expected to be a very small quantity, the deviations will be highly suppressed and one has to look at particles with high enough energy-momentum (w.r.t noncommutative scale), so that the deviations get sufficiently enhanced to be detectable. Hence, as stated in beginning of the paper, the best place to look for the signatures of twisted statistics, is to look at particles in L.H.C or in Ultra High Energy Cosmic Rays (UHECRs). Moreover, since HBT correlations are essentially quantum correlations between free identical particles and are manifestations of the particular statistics followed by the identical particles of a given beam, they are sensitive only to the statistics obeyed by these particles. Therefore despite our present lack of detailed knowledge about production mechanism and source of origin of UHECRs, a study of HBT correlations may still be able to provide unambiguous signatures, of underlying noncommutative nature of spacetime.

The graphs shown below highlight the difference between \( C_0^{(B)} - C_0^{(B)} \) and \( C_0^{(F)} - C_0^{(F)} \).^3

![Graphs](image)

Figure 1: In the above figures, the blue and red line represent the commutative and noncommutative HBT correlation functions respectively. The figures (a), (c) are plotted for ordinary and twisted bosons and (b), (d) are plotted for ordinary and twisted fermions. The figures (a), (b) are plotted with \( |\mathbf{p}_0| = 6 \times 10^{19} \text{eV} \) (i.e. at G.Z.K cutoff), \( \alpha = 2.04 \times 10^{19} \text{eV}^{-2} \) (i.e taken numerically same as the present error in estimation of energy of UHECRs) \( \|\lambda\| = 1.47 \times 10^{-24} \text{m}^2 \) and taken along z-axis and the angles as \( \theta_{\lambda-\mathbf{y}} = \frac{\pi}{4} \), \( \theta_{\lambda-\mathbf{p}_0} = \frac{\pi}{4} \), \( \phi_{\lambda-\mathbf{y}} = \frac{\pi}{3} \) and \( \phi_{\lambda-\mathbf{p}_0} = \frac{\pi}{6} \). The figures (c), (d) are plotted with \( |\mathbf{y}| = 1000 \text{m} \) and same values for rest of the parameters.

^3The persistence of correlations to large distances is attributed to the large uncertainties in our present determination of energy-momentum of UHECRs. With better and more precise knowledge of the energy-momentum, the correlation will decrease significantly resulting in better bounds on noncommutative deviations.
Of particular interest are the various angular dependences of noncommutative correlation functions (17) and (21) which are completely absent in commutative correlation functions (13) and (15). Due to rotation and revolution of earth, the noncommutative correlation functions will show periodic oscillations completely absent in commutative correlations and hence will perhaps provide the best and most unambiguous signal for underlying noncommutative structure of spacetime. Therefore, as claimed earlier, the information about noncommutative structure of spacetime, can be extracted out, from observing the nature of variation of HBT correlation functions (in particular by looking at angular variations) with varying certain experimentally measurable quantities, in a very unambiguous way.

Also it is worth noting that if we take noncommutative length scale same as Planck scale i.e. $10^{-35}$m then due to an upper limit on momenta of UHECRs (GZK cutoff), the deviations will turn out to be $O(10^{-50})$ which are too small to give any unambiguous signatures of it. So the noncommutative deviations are detectable only if the effective noncommutative scale is much larger than Planck scale, which is likely to happen in presence of large extra dimensions. Thus the noncommutative deviations in HBT correlations effectively provide us signatures of large extra dimensions.

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