Re-summation of fluctuations near ferromagnetic quantum critical points

C. J. Pedder, F. Krüger, and A. G. Green

1London Centre for Nanotechnology, University College London, 17-19 Gordon St, London, WC1H 0AH, UK
2SUPA, Department of Physics, University of St. Andrews, North Haugh, St Andrews, Fife. KY16 9SS

(Dated: May 11, 2014)

We present a detailed analysis of the non-analytic structure of the free energy for the itinerant ferromagnet near the quantum critical point in two and three dimensions. We analyze a model of electrons with an isotropic dispersion interacting through a contact repulsion. A fermionic version of the quantum order-by-disorder mechanism allows us to calculate the free energy as a functional of the dispersion in the presence of homogeneous and spiralling magnetic order. We re-sum the leading divergent contributions, to derive an algebraic expression for the non-analytic contribution to free energy from quantum fluctuations. Using a recursion which relates sub-leading divergences to the leading term, we calculate the full $T = 0$ contribution in $d = 3$. We propose an interpolating functional form, which allows us to track phase transition lines at temperatures far below the tricritical point and down to $T = 0$. In $d = 2$, quantum fluctuations are stronger and non-analyticities more severe. Using a similar re-summation approach, we find that despite the different non-analytic structures, the phase diagrams in two and three dimensions are remarkably similar, exhibiting an incommensurate spiral phase near to the avoided quantum critical point.

PACS numbers: 74.40.-n, 74.40.Kb, 75.30.Kz, 75.10.Lp

I. INTRODUCTION

Quantum fluctuations can have a dramatic effect on the phase behavior of the itinerant ferromagnet in the vicinity of the $T = 0$ quantum critical point. In a pioneering paper,[1] Hertz suggested that fluctuations can lead to new scaling in the physical quantities above the quantum critical point, which he termed “quantum critical scaling”. The theory of ferromagnetic quantum criticality was further developed by Moriya,[2] and Millis.[3] More recently, several authors,[4–10] noted that the coupling of soft electronic particle-hole modes to the magnetic order parameter generates non-analytic terms in the free energy that render the magnetic transitions first-order at low temperatures. Such fluctuation-induced first-order behavior is expected for ferromagnets in the Ising, XY, and Heisenberg universality classes in dimensions $1 < d < 3$, regardless of whether the moments are supplied by the conduction electrons or by electrons in another band.[11] Such first-order behaviour is also stable against weak disorder.[12,13] This explains why discontinuous ferromagnetic transitions are seen in various materials, including MnSi (Ref. [14]), Sr$_{1-x}$Ca$_x$RuO$_3$ (Ref. [15]), CoO$_2$ (Ref. [16]), UGe$_2$ (Ref. [16]), and URhGe (Ref. [17]).

It has been argued[6] that fluctuation-driven first-order transitions are indicative of the appearance of precursor, incommensurate magnetic states. Recently, this possibility has been explored in $d = 3$ within a fermionic version of the quantum order-by-disorder mechanism.[13,14,25] The central idea is to self-consistently calculate fluctuations around magnetically ordered states. This approach not only reproduces the non-analytic free energy of the homogeneous ferromagnet within a much simpler calculation, but also predicts the formation of an incommensurate spiral state, pre-empting the first-order transition into the ferromagnet. Spiral formation may be driven in mean-field by e.g. a feature in the density of states[20,21] or a breaking of inversion symmetry by the crystal structure, resulting in a Dzyaloshinskii-Moriya interaction.[22–24] Our mechanism for spiral formation does not require either of these complications, and is driven by quantum fluctuations of itinerant electrons with a simple, isotropic dispersion.

The quantum order-by-disorder phenomenon has much in common with the Coleman-Weinberg mechanism of mass generation in high energy physics[25] and with the Casimir effect.[26] It is familiar in the context of condensed matter physics.[30] However, where the fermionic version differs from all these approaches is that it encodes the quantum fluctuations in fermionic particle-hole excitations, rather than in fluctuations of the bosonic order parameter. The instability towards incommensurate order is associated with particular deformations of the Fermi surface. These alter the spectrum of the electronic soft modes which couple to the magnetic order parameter, leading to a self-consistent lowering of the free energy. The spiral is more stable than the homogeneous ferromagnet because the corresponding elliptical Fermi surface deformations increase the surface-to-volume ratio and therefore enlarge the phase space available for low energy, particle-hole modes.

Fluctuation-driven spiral phases near ferromagnetic quantum critical points are within the realm of experimental detection. The first clear example is PrPtAl, where detection by neutron diffraction is possible because of an amplification of magnetic moments due to the coupling between conduction electrons and local spins.[35] In the presence of disorder, long-range order in the spiral phase is destroyed, leading to a helical glass with a highly anisotropic correlation length.[37] Such a glassy state appears to have been observed recently near the avoided ferromagnet quantum critical point of CeFePO (Ref. [38]).
Since fluctuation-driven first-order behavior is such a robust feature of itinerant ferromagnets\cite{11} we expect that the instability towards incommensurate order should be similarly generic and also occur in quasi two-dimensional systems.

In this paper, we employ the fermionic quantum order-by-disorder approach to demonstrate that the phase diagram in $d = 2$ has the same topology as in $d = 3$ and exhibits a spiral phase at low temperatures. We perform a re-summation of the leading divergences, which enables us to follow the phase boundaries of the spiral phase far from the tricritical point down to $T = 0$. The analytic structure depends crucially on the spatial dimension. In $d = 3$, the re-summation of the leading divergent terms yields a contribution to the free energy $\Delta F_1 \sim \ln(\mathcal{M}^2 + T^2)$, in agreement with previous result\cite{12}. In $d = 2$ we find $\Delta F_1 \sim 2T$ and $\Delta F_1 \sim \ln \mathcal{M}$, in the regimes $T \gg \mathcal{M}$ and $T \ll \mathcal{M}$ respectively. These expressions are generalized for the modulated spiral state. We show that the low-temperature behavior is controlled by a hierarchy of divergences and we derive a recursion relation that relates all sub-leading divergences to the leading term $\Delta F_1$. We use this recursion to re-sum all corrections at $T = 0$ to find the exact location of the quantum critical point in $d = 3$ from the expression $\Delta F \sim \ln(M)$. The key steps in deriving the free energy functional are as follows. i. Starting from a coherent state path integral, we perform a Hubbard-Stratonovich decoupling of the electron interaction term in spin- and charge channels. ii. We decompose the fluctuation fields introduced in this way into zero- and finite-frequency components. The former correspond to static order in the system. Here, we include the spiral magnetic order parameter $M_F(M, \mathbf{r}) = M[n_{\uparrow, \cos(\mathbf{Q} \cdot \mathbf{r})} + n_{\downarrow, \sin(\mathbf{Q} \cdot \mathbf{r})}]$. iii. We re-insert the parameter $\mathbf{Q} = Q \mathbf{n}_z$ in the free-fermion propagator. This facilitates the self-consistent expansion and re-sums particular classes of diagrams to infinite order. iv. We trace over the fermions, keeping all terms up to quadratic order in the finite-frequency fluctuation fields. v. We perform the Gaussian integrals over the fluctuation fields. Finally, we do the summation over Matsubara frequencies.

This procedure yields a general expression for the free energy $\mathcal{F}(M, Q)$ as a functional of the electron mean-field dispersion $\epsilon^\pm_k$ in the presence of spiral order,

$$\epsilon^\pm_k = \frac{k^2}{2} \pm \sqrt{(\mathbf{k} \cdot \mathbf{Q})^2 + (gM)^2}. \quad (3)$$

At mean-field level we obtain

$$\mathcal{F}_{\text{mf}}(M, Q) = gM^2 - T \sum_{\mathbf{k}, \sigma = \pm} \ln[1 + e^{-(\epsilon^\pm_k - \mu)/T}], \quad (4)$$

while the fluctuation contribution is given by the integral

$$\Delta \mathcal{F}(M, Q) = 2g^2 \sum_{\mathbf{k}_1, \ldots, \mathbf{k}_4} \frac{n(\epsilon^+_{\mathbf{k}_1})n(\epsilon^-_{\mathbf{k}_2})[n(\epsilon^+_{\mathbf{k}_3}) + n(\epsilon^-_{\mathbf{k}_4})]}{\epsilon^+_{\mathbf{k}_1} + \epsilon^+_{\mathbf{k}_2} - \epsilon^-_{\mathbf{k}_3} - \epsilon^-_{\mathbf{k}_4}}. \quad (5)$$

The summation runs over the momenta $\mathbf{k}_1, \ldots, \mathbf{k}_4$ subject to the constraint $\mathbf{k}_1 + \mathbf{k}_2 = \mathbf{k}_3 + \mathbf{k}_4$ and $n(\epsilon) = 1/[1 + e^{(\epsilon - \mu)/T}]$ denotes the Fermi function. Note that this result is derived from self-consistent second-order perturbation theory after subtraction of an unphysical UV divergence\cite{13,15}.

C. Angular averages

The free energy is a functional of the electron dispersion $\epsilon^\pm_k$. As a result of this, $M$ and $Q$ enter the free

\begin{align*}
\mathcal{F}(M, Q) &= gM^2 - T \sum_{\mathbf{k}, \sigma} \ln[1 + e^{-(\epsilon^\pm_k - \mu)/T}], \\
\Delta \mathcal{F}(M, Q) &= 2g^2 \sum_{\mathbf{k}_1, \ldots, \mathbf{k}_4} \frac{n(\epsilon^+_{\mathbf{k}_1})n(\epsilon^-_{\mathbf{k}_2})[n(\epsilon^+_{\mathbf{k}_3}) + n(\epsilon^-_{\mathbf{k}_4})]}{\epsilon^+_{\mathbf{k}_1} + \epsilon^+_{\mathbf{k}_2} - \epsilon^-_{\mathbf{k}_3} - \epsilon^-_{\mathbf{k}_4}}.
\end{align*}
energy in a similar manner, and since derivatives of the integrands in Eq. (4) and Eq. (5) are strongly peaked near \( k_F \), there exist simple proportionality rules between the finite-\( Q \) and \( Q = 0 \) coefficients. The proportionality factors are determined by combinatorial factors and angular averages over powers of

\[
\eta_k^2 = \frac{(k \cdot Q)^2}{k_F^2 Q^2}.
\]

(6)

For example, the ratio of the \( Q^2 M^2 \) and \( M^4 \) coefficients is given by \( 2(\eta_k^2) \). The fact that the two coefficients become negative at the same time explains why the fluctuation-driven first-order transition to the homogeneous ferromagnet is pre-empted by a transition into a spiral state.

The angular averages depend on the dimension \( d \) and are defined as \( \langle \ldots \rangle = \frac{1}{\Gamma_d} \int d\Omega_d \ldots \) with \( \Omega_d \) the angular part of the volume element and \( \Gamma_d \) the surface area of a unit-sphere in \( d \)-dimensions. The relevant averages are easily calculated as

\[
\langle \eta_k^{2n} \rangle = \begin{cases} 
(2n - 1)!!/(2n)!! & (d = 2) \\
1/(2n + 1) & (d = 3)
\end{cases},
\]

(7)

where !! denotes the double factorial function. These observations enable us to calculate the free energy \( F(M, Q) \) of the spiral state from the free energy \( F(M) := F(M, Q = 0) \) for the homogeneous ferromagnet by the equation

\[
F(M, Q) = \left( F \left( \sqrt{M^2 + \eta_k^2 Q^2} \right) \right) - \langle F(\eta_k Q) \rangle, \tag{8}
\]

where we have rescaled \( Q \) by \( g \) for convenience. Note that the free energy only contains even powers of \( M \) and \( Q \) since the electronic Hamiltonian is isotropic and does not break inversion symmetry, \( \mathbf{r} \rightarrow -\mathbf{r} \).

D. Mean-Field contribution

The coefficients \( \alpha_{mf} \), \( \beta_{mf} \), and \( \gamma_{mf} \) in the Landau expansion for the homogeneous ferromagnet,

\[
F_{mf}(M) = \alpha_{mf} M^2 + \beta_{mf} M^4 + \gamma_{mf} M^6 + \ldots, \tag{9}
\]

are given by integrals over derivatives of the Fermi function,

\[
\alpha_{mf} = g + \frac{2g^2}{2!} \int d^d \mathbf{k} n^{(1)}(\epsilon_{\mathbf{k}}),
\]

\[
\beta_{mf} = \frac{2g^4}{4!} \int d^d \mathbf{k} n^{(3)}(\epsilon_{\mathbf{k}}),\tag{10}
\]

\[
\gamma_{mf} = \frac{2g^6}{6!} \int d^d \mathbf{k} n^{(5)}(\epsilon_{\mathbf{k}}).
\]

For the isotropic dispersion \( \epsilon_{\mathbf{k}} = k^2/2 \), these reduce to simple one-dimensional integrals. Using Eq. (8), we obtain the mean-field contribution

\[
F_{mf}(M, Q) = \left( \alpha_{mf} + 2\beta_{mf}(\eta_k^2)Q^2 + 3\gamma_{mf}(\eta_k^2)Q^4 \right) M^2 + \left( \beta_{mf} + 3(\eta_k^2)Q^2 \right) M^4 + \gamma_{mf} M^6, \tag{11}
\]

to the free energy of the spiral ferromagnet.

III. RESUMMATION OF FLUCTUATIONS

In this section we analyze the analytic structure of the fluctuation integral \( \Delta F(M, Q) \) Eq. (5) in general dimension to obtain closed-form expressions from a resummation of the leading divergences in \( d = 3 \) and \( d = 2 \). We first focus on the fluctuation integral of the homogeneous ferromagnet, \( \Delta F(M) = \Delta F(M, Q = 0) \), which is given by Eq. (5) with \( \eta_k = k^2/2 - \sigma g M \), and then generalize to finite \( Q \), using Eq. (8).

It is most convenient to express \( \Delta F(M) \) in terms of the components of the particle-hole density of states \( \rho(q, \epsilon) = \rho(q, \epsilon) - \Delta \rho(q, \epsilon) \), where these are defined as

\[
\rho^{\sigma}(q, \epsilon) = \int d^d \mathbf{k} n(\epsilon_{\mathbf{k}}^{\sigma} - \frac{q}{2}) \delta(\epsilon - \epsilon_{\mathbf{k}}^{\sigma} + \epsilon_q^{\sigma} - \frac{q}{2}),
\]

\[
\Delta \rho^{\sigma}(q, \epsilon) = \int d^d \mathbf{k} n(\epsilon_{\mathbf{k}}^{\sigma} - \frac{q}{2}) n(\epsilon_{\mathbf{k}}^{\sigma} + \frac{q}{2}) \delta(\epsilon - \epsilon_{\mathbf{k}}^{\sigma} + \epsilon_q^{\sigma} + \epsilon_q^{\sigma} - \frac{q}{2}), \tag{12}
\]

and calculated in Appendix A for dimension \( d = 3 \) and \( d = 2 \). With these definitions of \( \rho \) and \( \Delta \rho \), the fluctuation contribution to the free energy becomes

\[
\Delta F(M) = 2g^2 \sum_{\sigma = \pm} \int_{\mathbf{q}, \epsilon_1, \epsilon_2} \frac{\Delta \rho^{\sigma}(\mathbf{q}, \epsilon_1) \rho^{-\sigma}(-\mathbf{q}, \epsilon_2)}{\epsilon_1 + \epsilon_2}. \tag{13}
\]

Here the integrals run over \( \mathbf{q} \in \mathbb{R}^d \) and \( \epsilon_1, \epsilon_2 \in \mathbb{R} \).

A. Landau Expansion of \( \Delta F \)

We generate a Landau expansion of the fluctuation contributions for \( Q = 0 \) by Taylor expanding the expression Eq. (13) for small \( M \),

\[
\Delta F(M) = \alpha_{\delta} M^2 + \beta_{\delta} M^4 + \gamma_{\delta} M^6 + \ldots, \tag{14}
\]

where e.g. \( \alpha_{\delta} = \partial_M \Delta F |_{M=0} \). We exchange derivatives with respect to \( M \) for derivatives with respect to \( \mu \) using the relation \( \sigma g \partial_M = -\partial_\mu \) to write these coefficients in terms of integrals of the form

\[
J_{m,n}(T) = \int_{\mathbf{q}, \epsilon_1, \epsilon_2} \frac{\Delta \rho^{(m)}(\mathbf{q}, \epsilon_1) \rho^{-\sigma}(\mathbf{q}, \epsilon_2)}{\epsilon_1 + \epsilon_2}. \tag{15}
\]

These integrals depend only on derivatives of the particle-hole density of states with respect to \( \mu \) for \( M =
0. We use the shorthand notation $\bar{\rho}^{(m)} = \partial^m \bar{\rho}$ and $\Delta \rho^{(m)} = \partial^m \Delta \rho$. Using this notation, we can write the $M^2$ coefficient as

$$\alpha_\beta = \frac{4g^4}{2!} (J_{0,2} - 2J_{1,1} + J_{2,0})$$

$$= \frac{g^2}{2!} (\partial^2 \Delta F|_{M=0} - 16g^2J_{1,1}).$$  \hspace{1cm} (16)

In the second line we have used integration by parts to write the coefficient in terms of the symmetric integrals $J_{1,1}$ and $\Delta F|_{M=0} = 4g^2J_{0,0}$. Repeating this process up to order $M^{10}$, we find the coefficients listed in Table I. In the limit $T \rightarrow 0$, the integrals $J_{n,n}(T)$ diverge for $n \geq 2$ in $d = 2, 3$. The divergence becomes stronger with increasing order $n$ and therefore as we move from left to right in each row of Table I. At each order $M^{2n}$ in the Landau expansion, the leading small-$T$ dependence comes from the term proportional to $J_{n,n}(T)$.

The behavior of phase boundaries very close to the finite-temperature tricritical point is controlled by the smallness of $M$ and therefore by the coefficients $\alpha$, $\beta$, and $\gamma$. The phase boundaries at low temperatures far from the tricritical point are determined by the leading divergences as $T \rightarrow 0$. Using Table I we may begin to spot patterns in the coefficients of $J_{n,n}$. On the leading diagonal of the table, we note that the coefficient of $J_{n,n}$ takes a particularly simple form of $(-1)^n 2^{2n} (2n)!$. Re-summing the terms of all orders in $M$ along this diagonal we obtain

$$\Delta F_1(M) = 4g^2 \sum_{n=1}^\infty \frac{(-1)^n (2gM)^{2n}}{(2n)!} J_{n,n}(T).$$  \hspace{1cm} (17)

We later calculate closed-form expressions for $\Delta F_1$ in $d = 3$ and $d = 2$. Continuing down the chain to the $m$th diagonal, we obtain the free energy contribution

$$\Delta F_m(M) = g^2 \partial^2 \Delta F_{m-1}$$

$$= \sum_{n=1}^\infty \frac{(-1)^n (2g)^{2n}(2n-1)!}{(2n)!}(2n+2m-1)!n!2^{2n} \Delta F_{m-1}(T).$$  \hspace{1cm} (18)

Using this expression, we find a differential equation which relates $\Delta F_m$ to $\Delta F_{m-1}$,

$$\partial_M \left( \frac{\Delta F_m}{M} \right) = \frac{g^2 \partial^2 \Delta F_{m-1}}{2m}.$$  \hspace{1cm} (19)

As we approach $T = 0$, we find that sub-leading corrections become more significant, and so to find the correct quantum critical point, we must be able to calculate them in this limit. Repeated application of Eq. (19) enables us to calculate all sub-leading corrections from the functional form of the leading re-summed correction $\Delta F_1$. Re-summaion of all these sub-leading corrections is a tractable calculation in $d = 3$.

### B. Calculation of $\Delta F_1$ and $\Delta F$ in $d = 3$

We proceed by deriving an explicit expression for the re-summation $\Delta F_1(M)$ of leading divergences in $d = 3$. Details of the calculation of particle-hole density of states $\frac{\alpha}{\beta}$ and the integrals $J_{n,n}(T)$ are given in Appendices A and B, respectively. After summation over $n$, the final result takes the form

$$\Delta F_1(M) = -2\lambda(1 + 2\lambda)M^2 + \lambda M^4 \ln \left( \frac{4g^2 M^2 + T^2}{4\mu^2} \right),$$

where we have defined $\lambda = 4g^6\nu_F^2/(3\mu^2)$ with $\nu_F = k_F/2\pi^2$ and $\mu = k_F^2/2$. Note that we have included the non-divergent fluctuation correction of order $M^2$, corresponding to the $n = 1$ term in the sum. As noted in previous work[10], this term changes the location of the tricritical point and the phase behavior near to it. The logarithmic contribution arises from the summation of all divergent terms $n \geq 2$ and is of the same form as the diagrammatic result[3].

Using this re-summed form and the relation [5], we can also obtain a closed-form expression for the leading fluctuation correction $\Delta F_1(M, Q)$ to the free energy of the spiral ferromagnet,

$$\Delta F_1(M, Q)/\lambda = -2(1 + \ln 2)M^2$$

$$= \left( M^4 + 2 \frac{2}{3} M^2 Q^2 + \frac{1}{5} Q^4 \right) \ln \left[ \frac{4g^2 (M^2 + Q^2) + T^2}{4\mu^2} \right]$$

$$- \frac{1}{5} Q^4 \ln \left( \frac{4g^2 Q^2 + T^2}{4\mu^2} \right) - \frac{14}{45} M^2 Q^2$$

$$+ \left[ \frac{16}{15} M^4 - \frac{8}{15} M^2 T^2 \frac{2}{4g^2} + \frac{2}{5} \left( \frac{T^2}{4g^2} \right)^2 \right]$$

$$\times \left[ \sqrt{M^2 + \frac{T^2}{4g^2}} \arctan \left( \frac{Q}{\sqrt{M^2 + \frac{T^2}{4g^2}}} \right) \right].$$  \hspace{1cm} (21)

Given the form of $\Delta F_1$ in Eqn. (20), we may then use the recursion relation [19] to calculate the full, all-orders quantum correction $\Delta F(M) = \sum_{n=1}^\infty \Delta F_n(M, 0)$ at $T = 0$ (see Appendix C). The result is that

$$\Delta F(M) = 2\sqrt{\pi} \lambda M^4 \ln \left( \frac{g^2 M^2}{\mu^2} \right).$$  \hspace{1cm} (22)

Having calculated the form of the leading correction to the free energy for general $T$ and $M$, $\Delta F_1$, and the re-summation of all the divergent contributions to the free energy at $T = 0$, $\Delta F$, we may now postulate an analytical form for the full, re-summed quantum fluctuation contribution to the free energy that interpolates between these two cases. We suggest the functional form[23]

$$\Delta F(M) = \lambda M^4 \ln \left( \frac{4g^2 M^2}{\mu^2} \right)^{2\sqrt{\pi} + \frac{T^2}{4g^2}}.$$  \hspace{1cm} (23)
The high-temperature form of the free energy, \( \Delta F = 0 \) intercept correct, thereby capturing the important physics in a simple, closed form.

C. Calculation of \( \Delta F_1 \) in \( d = 2 \)

In \( d = 2 \), it has not proved possible to write down a closed, analytic form for the first re-summed correction to the free energy, \( \Delta F_1 \). However, it is still possible to use our formalism to find the asymptotic forms of this correction in the limits \( T \ll M \) and \( T \gg M \). These are given by

\[
\Delta F_1(M) = \begin{cases} 
    c_- M^3 \ln M & \text{for } T \ll M \\
    c_+ M^2 T & \text{for } T \gg M
\end{cases}
\]

with coefficients \( c_- = 16g^4\pi^{3/2}/[(2\pi)^3T(3/2)] \) and \( c_+ = 16\sqrt{2}g^4\pi^{3/2}Li_1(-1)/(2\pi)^3 \). Details of the derivation are given in Appendix D.

Again, we perform the angular averages \( \bar{Q} \) to generalize to finite \( Q \) and investigate possible instabilities towards spiral formation. The high-temperature form modifies the mean field coefficient of \( \gamma \), adding a term linear in \( T \), which cannot contribute to a term \( \sim Q^2 \). In the regime of low temperatures, \( T \ll M \), the fluctuation corrections generate extra terms in the Landau expansion of \( F(M, Q) \),

\[
\Delta F_1(M, Q) = c_- \left[ M^3 \ln M + \frac{1}{2} Q^2 M (1 + 3 \ln M) \right] + \frac{3}{16} Q^4 M^{-1} \left( 1 + \frac{3}{4} \ln M \right) ,
\]

where we have included terms up to order \( Q^4 \).

IV. PHASE DIAGRAMS

We are now in a position to calculate the phase diagrams in \( d = 3 \) and \( d = 2 \) by minimizing of the free energy \( \mathcal{F}(M, Q) = \mathcal{F}_m(M, Q) + \Delta \mathcal{F}(M, Q) \). In \( d = 3 \), we use the proposed interpolating form \( [\hat{c}] \), whereas in \( d = 2 \), we approximate \( \Delta \mathcal{F}(M, Q) \) by the leading re-summed correction \( \Delta \mathcal{F}_1(M, Q) \). We show that despite the different analytic structures of the fluctuation contributions contained in \( \Delta \mathcal{F} \), the phase diagrams in two and three dimensions have the same topology, and both show an instability to a spiral state below the temperature \( T_c \) of the tricritical point. The spiral phase intervenes between the homogeneous ferromagnet at strong electron repulsions \( g \) and the paramagnet at small \( g \). Let us start by writing conditions for the different phase transitions we seek. Note that the free energy is defined such that \( \mathcal{F}(M = 0, Q = 0) = 0 \).

i. The second-order transition from the paramagnet into the ferromagnet for \( T > T_c \) as the coupling \( g \) is increased is obtained by \( \partial_{M^2} \mathcal{F} |_{M=0} = 0 \). (As before we have defined \( F(M) = \mathcal{F}(M, Q = 0) \).)

ii. A first-order transition between the paramagnet and the ferromagnet as \( g \) is increased when \( T < T_c \) occurs when \( F(M^*) = 0 \) and \( \partial_{M^2} F |_{M=M^*} = 0 \) for some \( M^* \neq 0 \).

iii. Allowing for the possibility of spiral states, we find that the first-order transition from paramagnet into ferromagnet is preempted by a first-order transition into a spiral phase. This is slightly more involved; we must first solve \( \partial_{Q^2} \mathcal{F}(M, Q) = 0 \) to obtain the optimal pitch \( \hat{Q}(M) \) for a given magnetization \( M \neq 0 \). We then look for a first-order transition in \( \mathcal{F}(M) = \mathcal{F}(M, \hat{Q}(M)) \) as described in (ii).

iv. A Lifshitz transition from the spiral state into the ferromagnet, where the pitch \( Q \) of the spiral goes smoothly to zero. This is given by the line along which \( \hat{Q} = 0 \).

A. Phase Diagram in \( d = 3 \).

A phase diagram in \( d = 3 \) has been calculated in Ref. [19] taking into account fluctuation corrections to the \( M^2, M^4 \), and \( M^2Q^2 \) coefficients. In this section we will
show that while this approximation is valid in the vicinity of the tricritical point, it fails to describe the behavior at low temperatures. Since the order of the divergences of the coefficients increases with the order in the Landau expansion, it is crucial to use the re-summation \[21\] of leading divergences to obtain the phase boundaries at low temperatures.

For comparison, we recalculate the phase diagram of Ref. \[19\] from the truncated Landau expansion

\[
\mathcal{F}_{\text{trunc}}(M, Q) = \left( \alpha + \frac{2}{3} \beta Q^2 + \frac{3}{5} \gamma Q^4 \right) M^2 + (\beta + \gamma Q^2) M^4 + \gamma M^6, \tag{26}
\]

including fluctuation corrections \( \alpha_0 = -2\lambda(1 + \ln 2) \) and \( \beta_0 = 2\lambda(1 + \ln T) \) up to quartic order. Note that the \( \ln T \)-contribution of \( \beta_0 \) arises from the zeroth order term of an expansion of the logarithm in \( \Delta F_1 \) \[21\] while the non-divergent contribution to \( \beta_0 \) comes from the sub-leading terms in the second row of Table \[\ref{table:parameters}\].

The second-order phase boundary between the ferromagnet and the paramagnet is determined by \( \alpha = 0 \), consistent with condition (i). Fluctuations render the transition first-order below the temperature \( T_c \) of the tricritical point which is determined by the intersection of the \( \alpha = 0 \) and \( \beta = 0 \) lines. For contact repulsion, we obtain \( T_c/\mu \approx 0.3 \), in agreement with previous work.\[11,19\]

The tricritical temperature is considerably reduced by disorder\[14\] and finite range interactions\[14,20\]. With increasing range of electron interactions, the relative strength of fluctuation corrections \( \lambda \) decreases, leading to an exponential suppression of \( T_c \). This is apparent from the condition \( \beta = \beta_{\text{mf}} + 2\lambda(1 + \ln T_c) = 0 \).

In order to determine the first-order transition line between the spiral and the paramagnet, we follow the recipe described under (iii). This leads to the condition \( \alpha \gamma = \frac{1}{g_0} \beta^2 \) for the phase boundary.\[15\] Finally, the condition \( Q = 0 \) defines the Lifshitz line between the spiral and the ferromagnet (iv) coincides with the \( \alpha = 0 \) line below \( T_c \).

The resulting phase diagram is shown in Fig. 1. The low-temperature behavior of the first-order transition line between the paramagnet and the spiral phase is clearly unphysical. The phase boundary does not terminate on the \( T = 0 \) axis but instead approaches zero temperature asymptotically as \( g \to 0 \), suggesting that there exists a transition into a spiral state at low temperatures for arbitrarily small values of \( g \), far from the avoided quantum critical point.

In what follows, we will study the changes to the phase diagram which result from the inclusion of the functional form for \( \Delta F_1(M) \) given in Eqn. \[20\] which gives just the leading re-summed corrections from quantum fluctuations, and the interpolating form \( \Delta F(M) \) given in Eqn. \[23\], which captures the divergences of all the higher-order terms in the Landau expansion. In order to investigate the spiral phase behaviour, we use the finite-\( Q \) generalisation \( \Delta F(M, Q) \) of Eqn. \[23\], which is obtained by carrying out the angular averages.

The main effect of the re-summation is to cut off the \( \ln T \)-divergence by finite \( M \). This fundamentally changes the behavior of the first-order spiral/paramagnet line as illustrated in Fig. 1. The phase boundary has a vertical intercept with the \( T = 0 \) axis at a finite value \( g_c \), consistent with the Clausius-Clapeyron condition. Using just the leading re-summed expression \( \Delta F_1(M) \) \[20\], we are left with a very narrow region of spiral order. However, we know that for small values of \( T \), the sub-leading re-summed corrections will become significant, so we must include all the subleading corrections too. We may explicitly calculate the location of the \( T = 0 \) intercept for the first-order spiral transition line, to find a critical coupling \( 1/g_c \sim 0.1012 \). Previous Monte Carlo analysis\[15\] suggests the transition into the spiral state at \( T = 0 \) will occur at \( 1/g \sim 0.133 \). The numerical disagreement between the two approaches may stem from the fact that to carry out the \( T = 0 \) Monte Carlo calculation, the contact interaction must be replaced by one with a negative finite range.

\[\text{FIG. 1: (color online) Phase diagram in } d = 3 \text{ as a function of inverse electron repulsion } 1/g \text{ and temperature } T/\mu. \]

Above the tricritical point (shown in red) we find a continuous transition between the ferromagnet at large \( g \) and a paramagnet at small \( g \). The spiral forms below the tricritical point between the paramagnet and the ferromagnet. The effect of the re-summation of leading divergencies is illustrated. (i) The dashed line shows the case without re-summation, which shows unphysical behavior as \( T \to 0 \) (ii) Using the re-summation of the leading divergencies \( \Delta F_1 \), the spiral region becomes invisible on the scale of this phase diagram, collapsing into a region very close to the homogeneous magnetic state. (iii) The solid line shows the phase boundary of the spiral found using the form for \( \Delta F \) in Eqn. \[23\], and gives the exact \( T = 0 \) intercept, and the correct behaviour in the vicinity of the tricritical point.
B. Phase Diagram in $d = 2$.

We proceed to calculate the phase diagram in $d = 2$ by minimizing $\mathcal{F}(M, Q) = \mathcal{F}_{\text{mf}}(M, Q) + \Delta \mathcal{F}_1(M, Q)$ with $\mathcal{F}_{\text{mf}}$ given by Eq. (5) with $\langle n_0^2 \rangle = 1/2$ and $\langle n_0^4 \rangle = 3/8$ and $\Delta \mathcal{F}_1$ defined in Eq. (27). Note that since $\Delta \mathcal{F}_1$ is only known in the regimes $T \ll M$ and $T \gg M$ [see Eq. (25)], we will only be able to determine the asymptotic behavior of the phase boundaries and have to interpolate between the two regimes. There are crucial differences between the cases $d = 2$ and $d = 3$:

- In two dimensions, the density of states is constant, leading to an exponentially weak temperature dependence of $\alpha_{\text{mf}}$. As a consequence, the critical interaction strength $g_{\text{c}}(T)$ for the mean-field transition of the ferromagnet is practically constant up to $T \approx 0.2 \mu$ [see Fig. 2a].

- The fluctuation contributions in $d = 2$ are very different. Even in the vicinity of the tricritical point, we do not have a simple Landau expansion of the free energy; the corrections are intrinsically non-analytic across the whole phase diagram.

- The angular averages $\langle n_k^2 \rangle$ are larger in $d = 2$ and decay as $1/\sqrt{n}$ for large $n$, opposed to $1/n$ in $d = 3$. This has a profound effect on how the free energy of the spiral relates to that of the homogeneous ferromagnet.

In spite of these differences, the phase diagram for the two-dimensional case turns out to be remarkably similar to that for three dimensions. We proceed to construct the phase diagram in three steps. i. We first analyze the effects of the fluctuation corrections $\Delta \mathcal{F}_1(M)$ on the continuous transition between the ferromagnet and the paramagnet. This is controlled by the asymptotic form of $\Delta \mathcal{F}_1$, which is valid for $T \ll M$. We extrapolate this first-order line to higher temperatures and estimate the position of the tricritical point. ii. We obtain the first-order spiral-to-paramagnet transition by minimizing $\mathcal{F}(M, Q)$ in the low-temperature regime. iii. Finally, we find that similar to the $d = 3$ case, the Lifshitz line along which $Q \rightarrow 0$ coincides with the line $\alpha = 0$. This line is again extrapolated up to the tricritical point.

i. Second-Order Line. Using the asymptotic expression $\Delta \mathcal{F}_1(M) = c_+ M^2 T$, valid for $\mu \gg T \gg M$, we obtain the fluctuation-corrected transition line between the paramagnet and the ferromagnet by the condition $\alpha = \alpha_{\text{mf}} + c_+ T = 0$. Since $\alpha_{\text{mf}}$ is almost constant at low temperatures and since $c_+ < 0$, the region of stability of the ferromagnet increases with temperature where the phase boundary is almost linear. This is shown in Fig. 2a. At higher temperatures, the fluctuation effects saturate, and the second-order line then tracks the mean-field transition, albeit at lower values of $g$. We expect that the true second-order transition line interpolates between these two asymptotic limits, leading to the re-entrant behavior sketched in Fig. 2b.

ii. First-Order Line. We now focus on the regime $T \ll M$ where the re-summed fluctuations are of the asymptotic form $\Delta \mathcal{F}_1(M) = c_- M^3 \ln M$ with $c_- > 0$. This contribution leads to a first-order transition between the ferromagnet and the paramagnet at low temperatures and values of $g$ that are considerably smaller than those determined by the condition $\alpha = 0$ for the second-order
line. To determine the first-order phase boundary, we numerically search for solutions $M \neq 0$ of the equations $F(M) = 0$ and $\partial_M F = 0$. If we follow this first-order line to higher $T$, we expect a tricritical point at the intersection with the $\alpha = 0$. The most likely scenario is that this point coincides with the tip of the re-entrant $\alpha = 0$ line, sketched in Fig. 2.

iii. Spiral Phase. Finally, we allow for states where $Q \neq 0$. As found previously, the Lifshitz line where $Q \to 0$ is given by the extension of the second-order ferromagnet-to-paramagnet transition line to temperatures below the tricritical point.

We continue to calculate the first-order, paramagnet-to-spiral transition line in the limit $T \to 0$, using the asymptotic form of $\Delta F_1(M, Q)$ in the regime $T \ll M$. We follow the same procedure we used in $d = 3$ and first determine the optimal pitch $Q(M)$ for a given magnetization. This requires an expansion of $F(M, Q)$ up to 4th order in $Q$. We then determine the first-order transition of $F(M) = F(M, Q(M))$. As in the 3d case we find that this first-order transition pre-empts the one into the homogeneous state. From the asymptotic form of the free energy, we can only determine the phase boundary for small values of $T$ and rely on an interpolation up to the tricritical point. The resulting phase diagram is shown in Fig. 3.

V. CONCLUSIONS AND DISCUSSION

Despite its apparent simplicity, the Hamiltonian yields a rich and interesting phase diagram when we include the possibility of fluctuation-driven phases near to the ferromagnetic quantum critical point. Previous application of the quantum order-by-disorder approach to the three-dimensional case showed that quantum fluctuations not only drive the transition first-order at low temperatures, but also stabilize an incommensurate spiral phase below the tricritical point. These results are valid in the vicinity of the tricritical point but fail for much smaller temperatures, where the phase behavior is no longer controlled by the smallness of the magnetization $M$ but instead by the leading divergences of all orders in $M$ as $T \to 0$.

In this work we have presented a detailed analysis of the non-analytic structure of the fluctuation corrections to the free energy in $d = 2$ and $d = 3$. We have demonstrated that there exists an underlying hierarchy of divergences and obtained closed-form expression for the re-summation of the leading terms of all orders in $M$. By adopting the fermionic quantum order-by-disorder approach and self-consistently expanding around an electronic state with a spiral magnetization, we have derived re-summed expressions for the free energy of the spiral ferromagnet. This re-summation of leading divergences allows us to track phase boundaries at low temperatures far away from the tricritical point.

Our results demonstrate that not only the fluctuation-driven first-order transitions but also the instabilities towards spiral order are generic features of itinerant ferromagnets. Despite the different non-analytic structures, we find very similar phase diagrams in $d = 2$ and $d = 3$. There are however differences which can be tested by future experiments. In $d = 3$, the spiral phase is found in a narrow region on the border of ferromagnetism, leading to a sequence of transitions from paramagnet to spiral and finally to ferromagnet as temperature is decreased, consistent with recent experiments on PrPtAl. This order of transitions is not possible in $d = 2$. Because of the re-entrant behavior of the phase boundary of the homogeneous ferromagnet, we predict that the spiral phase is located below the ferromagnetic state and stable over a larger region of the phase diagram.

Expressions for non-analytic contributions to the free energy have been previously obtained by diagrammatic methods. Our work shows that the diagrammatic calculations are equivalent to self-consistent second-order perturbation theory. In $d = 3$, we indeed recover the previously-known result $\Delta F_1 \sim M^2 \ln(M^2 + T^2)$, and find the exact form of the $T = 0$ free energy. This analytic form allows us to find the precise location of the quantum critical point for the ferromagnet in 3d.

In $d = 2$, we break new ground. In contrast to the three-dimensional case, a straightforward calculation of a Landau expansion for the free energy where each coefficient of $M^{2n}$ is a simple function of $T$ is not possible. This signposts the fact that the system is intrinsically non-analytic, even in the vicinity of the tricritical point. Indeed there is a lack of consensus on the form of the quantum corrections we should expect from dia-
We are not so interested in the particle-hole densities \[ \Delta F \sim M^2 \sqrt{M^2 + T^2}, \] which is in agreement with our result \[ \Delta F_1 \sim M^3 T \] in the regime \( T \gg M \). However, the result of Ref. \cite{11} is based on the assumption that the singular behavior of the fluctuation integral for \( T \ll M \) is cut off in the same way. Our work shows that this assumption is incorrect and that the asymptotic behavior for \( T \ll M \) is given by \( \Delta F_1 \sim M^3 \ln M \).

Our results lay the groundwork for studying the multi-critical behavior of itinerant ferromagnets. In \( d = 3 \), the re-summed form for the phase diagram has already been used to study superconducting instabilities mediated by magnetic fluctuations. In this instance, the intertwined magnetic spiral state and the superconducting instability resulted in the formation of a pair-density-wave state at very low temperatures. In two-dimensional systems such exotic states might be stabilized at higher temperatures since quantum fluctuations become more important and the resulting non-analyticities are of a different form. Future applications of the fermionic quantum order-by-disorder approach include the study of multiple-band effects, orbital fluctuations, and the competition between nesting instabilities and fluctuation-driven phase formation.

Acknowledgments

We thank Andrew Berridge, Gareth Conduit, Gregor Hannappel, Chris Hooley, Una Karahasanovic, and Ed Yelland for helpful discussions and suggestions. This work was funded by the EPSRC under grant codes EP/H049584/1, EP/I031014/1 & EP/I004831/2.

Appendix A: Particle-hole density of states

In order to evaluate the fluctuation integral \( \Delta F(M) \), we must first calculate the particle-hole densities of states \( \hat{\rho}(q, \epsilon) \) and \( \Delta \rho(q, \epsilon) \), which are defined in Eq. (12). We will do this separately for the cases \( d = 3 \) and \( d = 2 \).

1. Three dimensions

In \( d = 3 \), and for \( M = 0 \), \( \hat{\rho} \) takes the form

\[
\hat{\rho}(q, \epsilon) = \int_{-1}^{1} d \cos \theta \int_0^\infty \frac{2 \pi k^2 dk}{(2 \pi)^3} n \left( \frac{k^2}{2} + \frac{q^2}{8} - \frac{kq \cos \theta}{2} \right) \times \delta(\epsilon - kq \cos \theta).
\]

(A1)

The integrals over \( \cos \theta \) and \( k \) may be done exactly. We are not so interested in the particle-hole densities themselves, but rather in their derivatives with respect to \( \mu \), which enter the integrals \( J_{n,n} \). For the \( n \)-th derivative we obtain

\[
\partial_\mu^n \hat{\rho}(q, \epsilon) = \hat{\rho}^{(n)}(q, \epsilon) = \frac{1}{(2 \pi)^2 q} \partial_\epsilon^{(n-1)} n \left[ \phi^-(q, \epsilon) \right],
\]

(A2)

where we have defined

\[
\phi^\pm(q, \epsilon) = \frac{1}{2} \left( \frac{\epsilon}{q} \pm \frac{q}{2} \right)^2.
\]

(A3)

Similarly, we find that

\[
\Delta \rho^{(n)}(q, \epsilon) = \frac{1}{(2 \pi)^2 q} \partial_\epsilon^{(n-1)} \left\{ n \left[ \phi^+(q, \epsilon) \right] n \left[ \phi^-(q, \epsilon) \right] \right\}.
\]

(A4)

2. Two Dimensions

In \( d = 2 \), we must calculate the integral

\[
\hat{\rho}(q, \epsilon) = \int_0^{2 \pi} d \theta \int_0^\infty \frac{k dk}{(2 \pi)^2} n \left( \frac{k^2}{2} + \frac{q^2}{8} - \frac{kq \cos \theta}{2} \right) \times \delta(\epsilon - kq \cos \theta).
\]

(A5)

We carry out the \( \theta \) integral to get

\[
\hat{\rho}(q, \epsilon) = \frac{2}{(2 \pi)^2} \int_0^\infty k dk n \left( \frac{k^2}{2} + \frac{q^2}{8} - \frac{q^2}{2} \right) \times \delta(\epsilon - kq \cos \theta) = \frac{2}{(2 \pi)^2} \int_0^\infty k dk n \left( \frac{k^2}{2} + \frac{q^2}{8} - \frac{q^2}{2} \right) \times \delta(\epsilon - kq \cos \theta).
\]

(A6)

By a suitable change of variables, we reduce this to

\[
\hat{\rho}(q, \epsilon) = \frac{\sqrt{2}}{(2 \pi)^2 q} \int_0^\infty \frac{d x}{\sqrt{x}} n[x + \phi^-(q, \epsilon)],
\]

(A7)

where the functions \( \phi^\pm(q, \epsilon) \) are defined as before (A3). An identical calculation for \( \Delta \rho \) yields

\[
\Delta \rho(q, \epsilon) = \frac{\sqrt{2}}{(2 \pi)^2 q} \int_0^\infty \frac{d x}{\sqrt{x}} n[x + \phi^+(q, \epsilon)] n[x + \phi^-(q, \epsilon)].
\]

(A8)

Although it is possible to perform the integrals (A7) and (A8) and write the results in terms of poly-logarithmic functions, the expressions in integral form will prove more useful to us.

Appendix B: Calculation of \( \Delta F_1(M) \) in \( d = 3 \)

It is convenient to define a new function \( \tilde{n}(x) \) to keep our calculations uncluttered; this takes the form

\[
\tilde{n}(x) = \frac{1}{1 + e^{x/T}}
\]

(B1)

and is a Fermi function with chemical potential \( \mu = 0 \). The most divergent part of the integral \( J_{n,n} \) in \( d = 3 \) comes from

\[
J_{n,n} = \int_0^{2 \pi} \frac{d q}{(2 \pi)^2} \frac{d \epsilon_1}{(\epsilon_1 + \epsilon_2)} \frac{d \epsilon_2}{(\epsilon_1 + \epsilon_2)} \delta^{(n-2)}[\phi^-(q, \epsilon_1) - \mu] \times \delta^{(n-2)}[\phi^-(q, \epsilon_2) - \mu] \tilde{n}(\epsilon_2).
\]

(B2)
We integrate by parts \((n-2)\) times each with respect to \(\epsilon_1\) and \(\epsilon_2\) to find
\[
J_{n,n} = \frac{2}{(2\pi)^6} \int \frac{d^q F}{k_F^{2n-2}} dq \frac{1}{\epsilon_1 - \epsilon_2} \left( \frac{\partial_n - \partial_{n-2}}{\partial_{n-2}} \right) \epsilon_1 + \epsilon_2
\]
\times \delta \left[ \left[ \epsilon_1 - \frac{q}{2} (q - 2k_F) \right] \delta \left[ \epsilon_2 - \frac{q}{2} (q - 2k_F) \right] \right] \bar{n}(\epsilon_2),
\]
and use the \(\delta\)-functions to integrate over \(\epsilon_1\) and \(\epsilon_2\),
\[
J_{n,n} = \frac{2(2n-4)!}{(2\pi)^6} \int_{-2k_F}^{2k_F} d\delta q \frac{(2k_F)^{2n-2}}{k_F^{2n-2}(2k_F \delta q)_{2n-3}} \tilde{n}(k_F \delta q),
\]
where \(\delta q = q - 2k_F\) is the deviation of the particle-hole pair's momentum from \(2k_F\), which is expected to be small. We simplify this equation by folding the integral around \(\delta q = 0\) and use the expression \(\tilde{n}(k_F \delta q) = \frac{1}{2} \left[1 - \tanh(k_F \delta q/2T)\right]\). The second term provides a cutoff for the lower limit of the integral, which now yields
\[
J_{n,n} = -\frac{4(2n-4)! \nu_F^2}{(2\mu)^{2n-2}} \int_{-2k_F}^{2k_F} \frac{du}{u^{2n-3}},
\]
where we have made a simple change of variables, \(u = k_F \delta q\), and defined \(\nu_F = k_F / 2\pi^2\). Substituting this expression into Eq. \([17]\), we find
\[
\Delta F_1 = -64g^2 \nu_F^2 \mu^2 \int_{-2k_F}^{2k_F} \frac{du}{u^{2n-3}} \sum_{n=2}^{\infty} \frac{(2n-4)!}{(2n)!} \left( \frac{2gM}{\mu} \right)^{2n}.
\]
Taking sequential derivatives with respect to \(M\), we simplify our expression significantly to obtain
\[
\partial_M^4 \Delta F_1 = \frac{4 \nu_F^3}{\mu^2} \int_{-2k_F}^{2k_F} \frac{du}{u^{2n-3}} \frac{(-1)^n (2gM)^{2n-4}}{(2n)!} \left[ \frac{u^{2n-3}}{u^{2n-3}} - \frac{2gM^2}{2} \right] u^2
\]
\[
= \frac{32g^2 \nu_F^2 \mu^2}{\mu^2} \ln \left( \frac{g^2 M^2}{\mu^2} + \frac{T^2}{4\mu^2} \right),
\]
To get \(\Delta F_1\), we integrate this four times with respect to \(M\), keeping only terms that will give us an overall coefficient of \(M^4\). This gives the final closed form expression
\[
\Delta F_1(M, T) = \frac{4 \nu_F^3 M^4}{3 \mu^2} \ln \left( \frac{g^2 M^2}{\mu^2} + \frac{T^2}{4\mu^2} \right)
\]
for the re-summation of leading divergences in \(d = 3\).

Appendix C: Calculation of \(\Delta F(M)\) in \(d = 3\).

We wish to use the recursion relation given in the text \([19]\) to calculate the complete, all-orders re-summation at \(T = 0\). Firstly, we write \(\Delta F_1\) in terms of the variable \(y = \sqrt{2\mu/gM}\), and define the functions \(f_n(y)\) by
\[
\Delta F_n(M) = (gM)^2 f_n(y).
\]
In terms of this new variable, the recursion relation \([19]\) takes the form
\[
f_n(y) = \frac{1}{n!} \partial_y^{n-2} f_1(y).
\]
By Fourier transforming this expression to \(k\)-space, and after summation over \(n\), we find that \(\Delta \bar{F} = \sum_n \Delta F_n\) is given by
\[
\Delta \bar{F}(k) = g^2 M^2 \frac{1 - e^{-k^2}}{k^2} \bar{f}_1(k).
\]
We approximate the Fourier transform of \((1 - e^{-k^2})/k^2\) by a triangular function with the same total area, to do the convolution and approximate for \(y \gg 1\). Expressing the result in terms of \(M\), we find the full resummation of the \(T = 0\) quantum corrections to all orders, which is given by
\[
\Delta \bar{F}(M) = 2\sqrt{\pi} \lambda M^4 \ln \left( \frac{g^2 M^2}{\mu^2} \right).
\]

Appendix D: Calculation of \(\Delta F_1(M)\) in \(d = 2\).

Using the integral expressions \([A7]\) and \([A8]\) for the particle-hole densities of states in \(d = 2\), \(J_{n,n}\) is given by a five-dimensional integral
\[
J_{n,n} = \frac{2}{(2\pi)^5} \int \frac{dq}{q} \frac{d\epsilon_1 d\epsilon_2 dx d\delta_n}{x + \phi^{-}(q, \epsilon_1)}
\]
\[
\times \delta_n \left\{ n \left[ y + \phi^{-}(q, \epsilon_2) \right] + n \left[ y + \phi^{-}(q, \epsilon_2) \right] \right\}
\]
\[
\times \sqrt{x y}(\epsilon_1 + \epsilon_2),
\]
where \(\phi^{\pm}(q, \epsilon)\) are defined as before and the ranges of integration are \(\epsilon_1, \epsilon_2 \in (-\infty, \infty)\) and \(q, x, y \in [0, \infty)\). In order to do the integrals over \(\epsilon_1\) and \(\epsilon_2\), we first focus on the most divergent term, where all the derivatives hit the second Fermi function, and none hit the third. We make the same approximation as before, namely that \(n'(x) = -\delta(x - \mu)\), and then linearize the arguments of the derivatives of the Fermi functions as for \(d = 3\). We write the derivatives with respect to argument in terms of derivatives with respect to \(\epsilon\), and then integrate by parts with respect to \(\epsilon_1\) and \(\epsilon_2\), \((n - 1)\) times each. This gives
\[
J_{n,n} = \frac{2(2n-2)!}{(2\pi)^5} \int \frac{dq}{q} \frac{d\epsilon_1 d\epsilon_2 dx dy}{x y}(\epsilon_2)
\]
\[
\times q^{2n} \delta(y - k_F \epsilon_2/q + k_F \delta q) \delta(x - k_F \epsilon_1/q + k_F \delta q),
\]
\[
(D2)
\]
Changing variables $u = y + k_F \delta q$ and doing the integral over $x$ we get

$$
\partial_M^2 \Delta F_1 = -\frac{32g^4}{(2\pi)^5 k_F} \int_{-2k_F}^{2k_F} du \int_0^\infty dy \frac{1}{\sqrt{y}} \times \sqrt{u^2 + 4g^2 M^2 + u} \left(\frac{2k_F}{u^2 + 4g^2 M^2} \bar{n}(y + u)\right). 
$$ (D5)

We now do the $y$-integral to get

$$
\partial_M^2 \Delta F_1 = \frac{32g^4 \pi^{3/2}}{2(2\pi)^5 \sqrt{2k_F}} \int_{-2k_F}^{2k_F} du \int_0^\infty dy \frac{1}{\sqrt{y}} \times \sqrt{u^2 + 4g^2 M^2 + u} \left(\frac{2k_F}{u^2 + 4g^2 M^2} \bar{n}(y + u)\right). 
$$ (D6)

Since this integration cannot be done analytically, we approximate it in two limits. Firstly, when $M \gg 1$, we take $M = 0$ inside Eq. (D6), which becomes

$$
\partial_M^2 \Delta F_1 = \frac{32g^4 \pi^{3/2}}{(2\pi)^5 \sqrt{2k_F}} \int_{-2k_F}^{2k_F} du \frac{\bar{n}(u)}{\sqrt{u}} \left(\frac{2k_F}{u^2 + 4g^2 M^2} \bar{n}(u)\right). 
$$ (D7)

Only positive values of $u$ contribute to this integral when taking the limit $T \to 0$; the lower limit becomes $u = 0$. We may then scale out $T$ to get

$$
\partial_M^2 \Delta F_1 = \frac{32g^4 \pi^{3/2}}{(2\pi)^5 k_F} \int_0^\infty du \frac{\bar{n}(u)}{\sqrt{u}} \times \bar{n}(u). 
$$ (D8)

We split the integral into two regions, $u \in [0, 1]$ where we approximate $\bar{n}(u) \approx \bar{n}(1)$ and $u \in [1, \infty)$ where $\bar{n}(u) \approx 0$. Therefore, we have

$$
\partial_M^2 \Delta F_1 = \frac{64g^4 \pi^{3/2}}{(2\pi)^5 k_F} \bar{n}(1). 
$$ (D9)

which corresponds to a term in the free energy

$$
\Delta F_1(M) = \frac{16g^4 \pi^{3/2}}{(2\pi)^5} \bar{n}(1). 
$$ (D10)

Next, we turn our attention to the limit where $T \ll M$. Setting $T = 0$ in Eq. (D6), we now see that only values of $u < 0$ contribute. For this range of $u$, we may approximate the poly-logarithm by $\bar{n}(u) \approx u^n / (n + 1)$ where $\Gamma$ is the Euler gamma function. This gives

$$
\partial_M^2 \Delta F_1 = \frac{-32g^4 \pi^{3/2}}{(2\pi)^5 \sqrt{2k_F} \Gamma(3/2)} \int_0^{4\mu} du \frac{\bar{n}(u)}{u^2 + 4g^2 M^2 - u^2}. 
$$ (D11)

We rescale $u = 2M \bar{u}$, and scale out $\mu$ to get

$$
\partial_M^2 \Delta F_1 = \frac{-32g^4 \pi^{3/2} M}{2(2\pi)^5 \Gamma(3/2)} \int_0^{2\sqrt{M}} d\bar{u} \bar{u} \bar{u}^2 + 4g^2 M^2 - \bar{u}^2 \bar{u}^2 + g^2. 
$$ (D12)

Splitting the integral as before, and approximating the two halves we find

$$
\int_0^{2\sqrt{M}} d\bar{u} \bar{u} \bar{u}^2 + 4g^2 M^2 - \bar{u}^2 \bar{u}^2 + g^2 
$$

$$
\approx \int_0^1 d\bar{u} \bar{u} \bar{u}^2 + 4g^2 M^2 - \bar{u}^2 \bar{u}^2 + g^2 + \int_1^{2\sqrt{M}} d\bar{u} \frac{1}{\bar{u}^2 + g^2}. 
$$ (D13)

The first term gives us a numerical constant, which corresponds to a sub-leading $\partial_M^2 \Delta F_1 \sim M$ piece. The second integral gives us $\text{ArcSinh}(2\mu / M) - \text{ArcSinh}(1)$. Recall that $\text{ArcSinh}(x) = \ln(x + \sqrt{x^2 + 1})$, and that we are working in the limit where $M < 1$, then this first, positive term gives us a leading contribution

$$
\partial_M^2 \Delta F_1 = \frac{32g^4 \pi^{3/2}}{(2\pi)^5 \Gamma(3/2)} M \ln M, 
$$ (D14)

which corresponds to the term

$$
\Delta F_1(M) = \frac{16g^4 \pi^{3/2}}{(2\pi)^5 \Gamma(3/2)} M^3 \ln M, 
$$ (D15)

in the free energy for sufficiently small values of $T$. 


By doing the re-summation for $T = 0$, we remove one of the two energy scales in the problem completely. As a result, we cannot garner any information about the precise nature of the relative scaling between $M$ and $T$ that would be imposed on us by doing the complete re-summation for $T \neq 0$. Therefore, the Lifshitz line for the spiral coincides with $T_c$. 

The double factorial function is defined for odd argument as $(2k-1)!! = (2k)!/2^k k!$ and for even argument as $(2k)!! = 2^k k!$.

Recall that we define the "optimal" $\hat{Q}$ for a given $M$ as $\hat{Q}(M)$ which solves the equation $\partial_M F(M, \hat{Q}(M)) = 0$. Using Eq. (8), we see that this translates into $(\eta k \hat{Q}^2 F(M^2 + \eta k \hat{Q}^2) - \eta k F(\eta k \hat{Q}^2)) = 0$. Taking the limit that $\hat{Q}(M) \to 0$ in this expression, and assuming that the angular average is non-zero, we find that this reduces to the expression $\partial_M F(M)|_{M=0} = 0$, which is the condition that $\alpha = 0$. Therefore, the Lifshitz line for the spiral coincides with the extension of the second-order line below $T_c$. 

By doing the re-summation for $T = 0$, we remove one of the two energy scales in the problem completely. As a result, we cannot garner any information about the precise nature of the relative scaling between $M$ and $T$ that would be imposed on us by doing the complete re-summation for $T \neq 0$. Therefore, the Lifshitz line for the spiral coincides with $T_c$. 

The double factorial function is defined for odd argument as $(2k-1)!! = (2k)!/2^k k!$ and for even argument as $(2k)!! = 2^k k!$.

Recall that we define the "optimal" $\hat{Q}$ for a given $M$ as $\hat{Q}(M)$ which solves the equation $\partial_M F(M, \hat{Q}(M)) = 0$. Using Eq. (8), we see that this translates into $(\eta k \hat{Q}^2 F(M^2 + \eta k \hat{Q}^2) - \eta k F(\eta k \hat{Q}^2)) = 0$. Taking the limit that $\hat{Q}(M) \to 0$ in this expression, and assuming that the angular average is non-zero, we find that this reduces to the expression $\partial_M F(M)|_{M=0} = 0$, which is the condition that $\alpha = 0$. Therefore, the Lifshitz line for the spiral coincides with the extension of the second-order line below $T_c$. 

By doing the re-summation for $T = 0$, we remove one of the two energy scales in the problem completely. As a result, we cannot garner any information about the precise nature of the relative scaling between $M$ and $T$ that would be imposed on us by doing the complete re-summation for $T \neq 0$. Therefore, the Lifshitz line for the spiral coincides with $T_c$. 

The double factorial function is defined for odd argument as $(2k-1)!! = (2k)!/2^k k!$ and for even argument as $(2k)!! = 2^k k!$.

Recall that we define the "optimal" $\hat{Q}$ for a given $M$ as $\hat{Q}(M)$ which solves the equation $\partial_M F(M, \hat{Q}(M)) = 0$. Using Eq. (8), we see that this translates into $(\eta k \hat{Q}^2 F(M^2 + \eta k \hat{Q}^2) - \eta k F(\eta k \hat{Q}^2)) = 0$. Taking the limit that $\hat{Q}(M) \to 0$ in this expression, and assuming that the angular average is non-zero, we find that this reduces to the expression $\partial_M F(M)|_{M=0} = 0$, which is the condition that $\alpha = 0$. Therefore, the Lifshitz line for the spiral coincides with the extension of the second-order line below $T_c$. 

By doing the re-summation for $T = 0$, we remove one of the two energy scales in the problem completely. As a result, we cannot garner any information about the precise nature of the relative scaling between $M$ and $T$ that would be imposed on us by doing the complete re-summation for $T \neq 0$. Therefore, the Lifshitz line for the spiral coincides with $T_c$. 

The double factorial function is defined for odd argument as $(2k-1)!! = (2k)!/2^k k!$ and for even argument as $(2k)!! = 2^k k!$.

Recall that we define the "optimal" $\hat{Q}$ for a given $M$ as $\hat{Q}(M)$ which solves the equation $\partial_M F(M, \hat{Q}(M)) = 0$. Using Eq. (8), we see that this translates into $(\eta k \hat{Q}^2 F(M^2 + \eta k \hat{Q}^2) - \eta k F(\eta k \hat{Q}^2)) = 0$. Taking the limit that $\hat{Q}(M) \to 0$ in this expression, and assuming that the angular average is non-zero, we find that this reduces to the expression $\partial_M F(M)|_{M=0} = 0$, which is the condition that $\alpha = 0$. Therefore, the Lifshitz line for the spiral coincides with the extension of the second-order line below $T_c$. 

By doing the re-summation for $T = 0$, we remove one of the two energy scales in the problem completely. As a result, we cannot garner any information about the precise nature of the relative scaling between $M$ and $T$ that would be imposed on us by doing the complete re-summation for $T \neq 0$. Therefore, the Lifshitz line for the spiral coincides with $T_c$. 

The double factorial function is defined for odd argument as $(2k-1)!! = (2k)!/2^k k!$ and for even argument as $(2k)!! = 2^k k!$.

Recall that we define the "optimal" $\hat{Q}$ for a given $M$ as $\hat{Q}(M)$ which solves the equation $\partial_M F(M, \hat{Q}(M)) = 0$. Using Eq. (8), we see that this translates into $(\eta k \hat{Q}^2 F(M^2 + \eta k \hat{Q}^2) - \eta k F(\eta k \hat{Q}^2)) = 0$. Taking the limit that $\hat{Q}(M) \to 0$ in this expression, and assuming that the angular average is non-zero, we find that this reduces to the expression $\partial_M F(M)|_{M=0} = 0$, which is the condition that $\alpha = 0$. Therefore, the Lifshitz line for the spiral coincides with the extension of the second-order line below $T_c$. 

By doing the re-summation for $T = 0$, we remove one of the two energy scales in the problem completely. As a result, we cannot garner any information about the precise nature of the relative scaling between $M$ and $T$ that would be imposed on us by doing the complete re-summation for $T \neq 0$. Therefore, the Lifshitz line for the spiral coincides with $T_c$. 

The double factorial function is defined for odd argument as $(2k-1)!! = (2k)!/2^k k!$ and for even argument as $(2k)!! = 2^k k!$.