Bhabha scattering at NNLO with next-to-soft stabilisation

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1. Introduction

Electron-positron or Bhabha scattering is one of the best studied processes in the Standard Model [1]. It is well suited for luminosity measurements at $e^+e^-$ colliders because of its large cross section and clean signature. Furthermore, for energies well below the electroweak scale the radiative corrections are dominated by quantum electrodynamics (QED) which allows for a very precise theory prediction. As a consequence, much work has been put into the calculation of higher-order matrix elements as well as the development of Monte-Carlo event generators.

The next-to-leading order (NLO) matrix elements have been known in the full Standard Model for quite some time [2–5]. At next-to-next-to-leading order (NNLO) the situation is different. In the case of the electroweak corrections only logarithmically enhanced terms have been calculated [6–9]. On the QED side much more is known. The full two-loop matrix element with vanishing electron mass was calculated some time ago [10]. Subsequently, this result was extended to also include leading-order mass effects [11–14]. The subset of the two-loop matrix element containing closed electron loops has been computed without any approximations [15]. Although the exact mass dependence of the full two-loop contribution is still not known, leading power-suppressed mass effects were recently taken into account in [16]. The one-loop corrections to the radiative matrix element were calculated in [17].

In addition to the work that has been put into the calculation of the matrix elements various Monte-Carlo event generators were developed, combining the matrix elements to physical observables such that non-trivial detector geometries and acceptances can be taken into account [4,18–27]. In particular, the BABAYAGA event generator that is based on the matching of the exact NLO results to a parton shower algorithm has achieved a precision of below 0.1% [28]. A detailed analysis of the impact of fixed-order fermion NNLO contributions was presented in [29].

Even though all necessary ingredients are available, a Monte Carlo that includes also NNLO photonic corrections was missing. The main bottleneck in this regard has been the real-virtual contribution that suffers from numerical instabilities when integrated over the phase space of the emitted photon. The source of these instabilities can be traced back to the disparate scales in the process introduced by the small electron mass that acts as a regulator of collinear divergences. This problem is exacerbated in the presence of soft radiation.

In this paper we present a method to reliably integrate the real-virtual matrix element over the full phase space. It is based on the expansion for small photon energies $E_\gamma \equiv \xi \times \sqrt{s}/2$ including the non-universal next-to-soft contribution at $O(\xi^{-1})$. To verify our method we have compared with approximate results from BABAYAGA at the cross section as well as at the differential level and found agreement within the expected 0.1% precision.

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With this method it is therefore possible to make reliable predictions for Bhabha scattering at the differential level including the full set of NNLO QED corrections.

This paper is organised as follows: We begin by briefly introducing our calculational framework in Section 2. The main result is presented in Section 3 where we describe how the stabilisation of the real-virtual matrix element was achieved via the next-to-soft approximation. We verify our method in Section 4 and conclude in Section 5.

2. Overview of the calculation

We consider the scattering process

\[ e^{-}(p_1) e^{+}(p_2) \rightarrow e^{-}(p_3) e^{+}(p_4) (\gamma(p_3) \gamma(p_4)) \]

up to NNLO in QED. As we are mainly interested in establishing the stabilisation method we restrict ourselves to purely photonic corrections, i.e. we do not take into account contributions from closed fermion loops. Ultraviolet (UV) and infrared (IR) divergences are regulated in \( d = 4 - 2\epsilon \) dimensions and the renormalisation is performed in the on-shell scheme.

All tree-level and one-loop matrix elements were calculated with the full electron mass dependence. The corresponding diagrams were generated using QGraf [30] and evaluated with the Mathematica code Package-X [31]. In the case of the numerically delicate real-virtual matrix element this Mathematica calculation serves mostly as a reference calculation. In the bulk of the phase space we instead rely on OpenLoops [32,33]. As we will discuss in Section 3, for small photon energies we switch to a next-to-soft approximation.

As mentioned in the introduction the full mass dependence of the photonic two-loop matrix element is not known. However, for most practical applications we can assume the scale hierarchy \( m^2 \ll Q^2 \ll s, t, u \) with the electron mass \( m \) and the Mandelstam invariants \( s = (p_1 + p_2)^2, t = (p_1 - p_3)^2, \) and \( u = (p_1 - p_4)^2 \). For our purposes it is therefore justified to neglect the power-suppressed terms of \( O(m^2/Q^2) \). Due to the universal structure of collinear divergences the leading mass effects can be straightforwardly included based on the massless result. This massification procedure was developed in the context of Bhabha scattering [11-13] and was recently extended to processes with a heavy mass [34].

The matrix elements are implemented in the integrator McMur, a Monte Carlo for MUnions and other Leptons [35]. This framework is based on the FKS\textsuperscript{S} subtraction scheme [36] which is an extension of the original FKS scheme [37,38] beyond NLO for QED. This subtraction scheme allows to consistently remove the singularities arising from soft photon emission in order to calculate observables in a fully differential way. The simplicity of the FKS\textsuperscript{S} subtraction scheme is due to the absence of collinear and the simple structure of soft singularities. All collinear divergences are regulated by finite fermion masses. Following the notation from [36] the soft singularities exponentiate according to the YFS formula [39]

\[ \sum_{l=0}^{\infty} M^{(l)}_{n} = e^{-\xi} \sum_{l=0}^{\infty} M^{(l)}_{n} \xi^{l}. \]

All soft poles of the \( \xi \)-loop matrix element (squared amplitude) with \( n \) final-state particles \( M^{(n)}_{\xi} \) are absorbed in the universal integrated eikonal factor \( \xi \), rendering \( M^{(l)}_{\xi} \) finite. This formula can be seen as a consequence of the universal behaviour of radiative matrix elements in the soft limit

\[ \lim_{\xi \rightarrow 0} \xi^{2} M^{(l)}_{n+1} = \xi^{2} M^{(l)}_{n}. \]

with the scaled photon energy \( \xi \) = \( 2E_{\gamma}/\sqrt{s} \) and the eikonal factor \( \xi \).

The most challenging part of the calculation presented here is to ensure a reliable integration of the real emission contributions in the phase-space region where the photon becomes collinear to the emitting fermion or where it becomes soft. In the former case the smallness of the electron mass acting as an infrared regulator results in large pseudo-collinear singularities. To address this issue we use a dedicated tuning of the phase-space parametrisation to help the \texttt{vegas} integration [40] find and deal with these problematic regions. In the latter case the integrand develops an unregularised soft singularity that is subtracted with the IR counterterm, resulting in a large cancellation. For this cancellation to work the matrix element has to be evaluated with very high precision. Due to the analytical and algebraic complexity of the real-virtual matrix element this is a highly non-trivial task and has presented the main obstacle to a complete, fully differential NNLO calculation of Bhabha scattering in the past. Our solution to this problem is the main result of this paper and is discussed in detail in the next section.

3. Real-virtual stabilisation via next-to-soft approximation

This section discusses how an implementation of the real-virtual matrix element can be obtained that ensures a stable and efficient integration in the soft phase-space region. As alluded to above, any general-purpose calculation of a one-loop matrix element will run into numerical instabilities at some point. In particular, for processes with an external photon with ever smaller energies, the IR-subtracted matrix element is a typical numerical pitfall whereby two expressions diverging as \( 1/\xi \) are combined to obtain an integrable integrand diverging as \( 1/\sqrt{\xi} \). The crucial question is whether these instabilities appear only for small enough \( \xi \) such that the integration can be done reliably. In this context, QED calculations are particularly delicate since the final states tend to be much less inclusive than jet cross sections computed for hadronic collisions.

We use OpenLoops [33] for the bulk of the phase space since it shows a remarkable numerical stability. In order to test for which values of \( \xi \) the instabilities start to appear, we compare OpenLoops to a dedicated computation of the real-virtual matrix element in Mathematica using arbitrary precision arithmetic. In Fig. 1 we show the deviation of OpenLoops from the ‘exact’ Mathematica result. For illustration we use an arbitrary phase-space point as well as one where the photon is emitted nearly collinear to the initial-state electron. For the former, at \( \xi = 10^{-5} \) the relative error is \( 10^{-8} \). In the collinear case the numerical instabilities are strongly enhanced with a relative difference of \( 10^{-1} \) for \( \xi = 10^{-5} \). All numbers that enter Fig. 1 including the particle momenta \( p_i \) of the two phase-space points are publicly available under [41].

An obvious idea is to expand the real-virtual matrix element for small photon energies and to switch to this approximation for sufficiently small \( \xi \). The leading \( O(\xi^{-2}) \) term in this expansion is given by (3) and can be easily calculated based on the one-loop matrix element \( M^{(1)}_{\xi} \). Using this approximation amounts to using the same algebraic expression for both terms in the subtracted integrand, albeit with different kinematics. As can be seen from Fig. 1 this approach is insufficient in the collinear region. If an accuracy below \( 10^{-3} \) is to be aimed at, in this case one has to switch to the expansion at \( \xi \sim 10^{-3} \). However, the exact matrix element is not sufficiently well approximated by the leading soft contribution in this region. To ensure a decent approximation we have therefore to include the non-universal \( O(\xi^{-1}) \) term in the soft expansion.

The next-to-soft terms of tree-level matrix elements have been considered a long time ago (Low-Burnett-Kroll theorem) [42,43].
Going beyond tree level, it is tempting to try to apply effective-field-theory methods. However, for QED with massive fermions the appropriate effective theory is the QED version of heavy quark effective theory and the genuine one-loop contribution to the next-to-soft effects is expected to be given by the process dependent soft function. From a practical point of view we have thus decided to directly calculate the non-universal $O(\xi^{-1})$ term in the soft expansion.

To be precise, we have computed the real-virtual matrix element in terms of scalar Passarino-Veltman functions using the Mathematica calculation of the real-virtual matrix element described in the previous section. Next, we have employed the power counting

$$p_i \rightarrow p_i \quad \text{for} \quad i \in \{1, 2, 3, 4\},$$

$$p_5 \rightarrow \lambda p_5,$$

$$m \rightarrow m,$$

and expanded in the book-keeping parameter $\lambda$. The expansion of the rational coefficients and simple Passarino-Veltman functions was performed with Mathematica. More complicated triangle- and box-functions were expanded at the (loop-)integrand level using the method of regions [44]. For the purpose of calculational efficiency we have used its formulation in the parametric representation [45]. In this case, the contributing regions can be easily found using the public code asy.m [46]. Most resulting integrals could be straightforwardly computed. The remaining ones were calculated using Mellin-Barnes techniques [47,48]. For the most involved integrals a two-fold Mellin-Barnes representation was necessary. However, the reduction to single contour integrals was possible in this case by resolving the singularity structure with the Mathematica package MBresolve.m [49]. In summary, the main technical difficulties in performing the next-to-soft expansion are the calculation of the integrals and the treatment of large intermediate expressions.

We have checked that the first term of the expansion indeed reproduces the result from (3). The non-universal subleading contribution was verified numerically. This is also shown in Fig. 1 where the inclusion of the next-to-soft $O(\xi^{-1})$ term significantly improves the approximation. This allows us to switch to a reliable expansion as early as $\xi \sim 10^{-3}$. We can therefore conclude that the next-to-soft approach ensures the numerical stability of the real-virtual matrix element for small photon energies which is a prerequisite for the IR subtraction to work. This is further emphasised by comparing integrated results with and without stabilisation. While the next-to-soft stabilisation ensures that results after successive Monte Carlo iterations are in agreement with each other, a drifting mean value is observed otherwise resulting in a significant discrepancy between the two results. Furthermore, the evaluation of the obtained expansion is a few 100 to over a 1000 times faster than OpenLoops, depending on the details of the kinematics. Since vegas tends to sample predominantly in the soft and collinear region this speed-up is noticeable even in the integration over the full phase space. While OpenLoops provides settings to work at higher accuracy this comes at a cost of speed.

4. Results and verification

To test the next-to-soft approach of stabilising the real-virtual contribution we have compared to BABAAYAGA. For all results presented in this section we have switched from OpenLoops to the next-to-soft approximation at $\xi = 10^{-3}$. As mentioned in the introduction, the event generator BABAAYAGA is based on a parton shower algorithm matched to the exact NLO result. Contrary to our complete fixed-order calculation it therefore only gives the logarithmically enhanced contributions at NNLO. For the comparison we use set-up (a) of [28] that is tailored to $\phi$ factories with a centre-of-mass energy of $\sqrt{s} = 1020$ MeV. The detector configuration is approximated with the kinematical cuts

$$E_{\min} = 408 \text{ MeV},$$

$$20^\circ < \theta_\pm < 160^\circ,$$

$$\xi_{\max} = 10^0,$$

where $E_{\min}$ is the minimum energy of the final-state electron/positron, $\theta_\pm(\theta_\mp)$ is the scattering angle in the centre-of-mass frame between the incoming and outgoing electron (positron), and $\xi_{\max}$ is the maximally allowed acollinearity $\xi = |180^\circ - \theta_+ - \theta_-|$. The order-by-order contributions, $\alpha^{(i)}$, to the integrated cross section, $\sigma_\pm = \sigma^{(0)} + \sigma^{(1)} + \delta K^{(i)}$, are presented in Table 1. Additionally, we show the corresponding $K$ factors defined as

$$K^{(i)} = 1 + \delta K^{(i)} = \frac{\sigma_i}{\sigma_{i-1}}.$$
We were therefore able to calculate for the first time the photonic NNLO corrections for Bhabha scattering in a fully differential way. This was implemented in the McMUL framework. We have cross-checked our exact NNLO results at the level of the total cross section as well as for differential distributions with the logarithmic approximation implemented in the parton shower generator BABAYAGA.

Numerical instabilities of the kind described above are a critical point of higher-order QED calculations. We therefore expect that the next-to-soft method will prove useful in other processes as well. This is obvious in the case of Møller scattering which is related to the Bhabha process via crossing. Corresponding results relevant for the experiment PRad II [51] will be presented in a forthcoming paper [52]. Furthermore, in the context of muon-electron scattering our approach could turn out to be invaluable where a fully differential NNLO calculation is highly desirable [53] and therefore aimed at [54].

Because of the wide range of applicability of the next-to-soft expansion, an investigation of a potential universal structure would be desirable to allow for a more efficient calculation of the expansion. This could be done in the framework of heavy quark effective theory. Furthermore, a similar approach could be pursued in the collinear region. Switching to a leading collinear expansion could result in a significant speed-up of the phase-space integration. This would entail the calculation of the currently unknown one-loop splitting functions for massive fermions.

**Declaration of competing interest**

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

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