Some aspects of field equations in generalized theories of gravity

T. Padmanabhan
IUCAA, Pune University Campus,
Pune, India.

Abstract
A class of theories of gravity based on a Lagrangian \( L = L(R^{a}_{bcd}, g^{ab}) \) which depends on the curvature and metric — but not on the derivatives of the curvature tensor — is of interest in several contexts including in the development of the paradigm that treats gravity as an emergent phenomenon. This class of models contains, as an important subset, all Lanczos-Lovelock models of gravity. I derive several identities and properties which are useful in the study of these models and clarify some of the issues that seem to have received insufficient attention in the past literature.

1 Introduction

The simplest geometrical theory of gravity is Einstein’s theory for which the gravitational Lagrangian is proportional to the Ricci scalar. A more general class of theories can be obtained from Lagrangians which have an arbitrary dependence on the curvature tensor and the metric but do not depend on the covariant derivatives of the curvature tensor. In recent years there is a growing recognition that the field equations of gravity may be emergent and may have the same conceptual status as the equations of fluid mechanics of elasticity (for a recent review, see [1]). In particular, the Lanczos-Lovelock models of gravity [which is a subset of theories built from Lagrangians of the form \( L(R^{a}_{bcd}, g^{ab}) \)] exhibit thermodynamic properties that are strikingly similar to those in Einstein’s theory of gravity. The main purpose of this paper is to describe several curious and useful identities and mathematical results which arise in this general class of theories. To the extent I know, they do not seem to have been emphasized in a coherent manner in the existing literature.

1.1 Motivation
Consider a generally covariant metric theory of gravity in \( D \) dimensional spacetime based on the action functional \( A_{\text{tot}} = A_{g} + A_{\text{matter}} \) where the gravitational action is given by

\[
A_{g} = \int_{\mathcal{V}} d^{D}x \sqrt{-g} L[g^{ab}, R^{a}_{bcd}] \tag{1}
\]

The variation of this action under the variation \( g^{ab} \rightarrow g^{ab} + \delta g^{ab} \) can be computed in a straightforward manner [2] to give the result:

\[
\delta A_{g} = \delta \int_{\mathcal{V}} d^{D}x \sqrt{-g} L = \int_{\mathcal{V}} d^{D}x \sqrt{-g} E_{ab} \delta g^{ab} + \int_{\mathcal{V}} d^{D}x \sqrt{-g} \nabla_{j} \delta v^{j} \tag{2}
\]
where
\[ \sqrt{-g}E_{ab} \equiv \left( \frac{\partial \sqrt{-g}L}{\partial g^{ab}} - 2 \sqrt{-g} \nabla^{m} \nabla^{n} P_{amnb} \right) ; \quad P_{abcd} \equiv \left( \frac{\partial L}{\partial R_{abcd}} \right) \] (3)
and
\[ \delta v^{j} \equiv [2P^{i}bjd(\nabla^{b} \delta g^{di}) - 2 \delta g^{di}(\nabla^{c}P^{ij}cd)] \] (4)
The variation of the matter term will give
\[ \delta A_{m} \equiv -\frac{1}{2} \int T_{ik} \delta g^{ik} \sqrt{-g} d^{4}x. \] (5)

where \( T_{ik} \) is the symmetric stress tensor of matter defined by this relation. To obtain the field equations from this variation, it is usual to assume that the boundary term involving \( \delta v \) which arises in Eq. (2) can be ignored. In that case, the vanishing of the integrand of \( \delta A_{g} \) requires the condition \[ E_{ab} - (1/2)T_{ab} \delta g_{ab} = 0. \] Since \( \delta g_{ab} \) is symmetric, this leads to the result that the symmetric part of \( E_{ab} \) should be equal to \( (1/2)T_{ab} \).

In the Einstein’s theory — which is, of course, a special case of the above — we have \( E_{ab} = R_{ab} - (1/2)g_{ab}R \), which is known to be symmetric because Ricci tensor is symmetric. In the more general class of theories we are considering, it is not obvious from the expression in Eq. (3) that \( E_{ab} \) is symmetric. A survey of the literature shows that this issue has not been clearly discussed and the field equation is written both in terms of \( E_{ab} \) as well as in terms of the explicitly symmetrized form \( E_{(ab)} \) by different authors.

While carrying out the differentiation with respect to \( g^{ab} \) in the first term of \( E_{ab} \) in Eq. (3), one is led to the (symmetric) combination \( P_{a}^{ijk}R_{bijk} + P_{b}^{ijk}R_{aijk} \). If the tensor \( P_{a}^{ijk}R_{bijk} \) is symmetric in \( a \) and \( b \), then these two terms can be combined into a single term, as is sometimes done in the literature (e.g., in my papers and textbook but without elaboration). However, it is not obvious why \( P_{a}^{ijk}R_{bijk} \) should be symmetric in \( a \) and \( b \) in general and I could not find an explicit proof in published literature.

The purpose of this note is to clarify many of these issues and prove several useful identities which help in the study of the generalized class of theories. These identities are non-trivial in the sense that they do not emerge from the algebraic symmetry properties of \( P_{abcd} \); instead they arise from the fact that when scalar quantities are constructed from several tensorial objects (like in the case of \( L \) built from \( R_{abcd} \) and \( g^{ij} \)), the derivatives of the scalar quantity with respect to these tensors must obey certain identities. While some of these results can be proved by other methods in special cases, the procedure I follow to derive these identities is very powerful and will be of use in more general contexts. (For example, the same procedure can be used when the Lagrangian depends on the covariant derivatives of the curvature and I hope to discuss this in a future work.)

With this motivation, let us take a closer look at the variational principle and the resulting expressions mentioned above.

1.2 Summary of results
While varying \( L\sqrt{-g} \) to obtain the equations of motion, we encounter two tensors closely related to the following partial derivatives
\[ P^{abcd} = \left( \frac{\partial L}{\partial R_{abcd}} \right)_{g_{ij}} ; \quad P_{a}^{b} = \left( \frac{\partial L}{\partial g_{ab}} \right)_{R_{ijkl}} \] (6)
The \( P^{abcd} \) may be called the entropy tensor of the theory because the integration of this tensor over a horizon (contracted with a pair of binormals of the horizon) gives the entropy.
of the horizon in the generalized theories of gravity \[4, 1\]. The derivatives appearing in Eq. (6) are defined within the subspace of infinitesimal deformations which preserve the symmetries of the independent variable. Therefore, \(P_{ab}\) is symmetric while \(P_{abcd}\) is assumed to inherit the algebraic symmetries of the curvature tensor:

\[
P_{abcd} = -P_{bacd} = -P_{abdc}; \quad P_{abcd} = P_{cdab}; \quad P_{a[bc]} = 0 \quad (7)
\]

One can construct several such tensors — which have the properties listed in Eq. (7) — from the curvature tensor and the metric. But these properties alone do not guarantee that such a tensor can be expressed as the derivative of a scalar with respect to the curvature tensor. The fact that \(P_{abcd}\) has the form in Eq. (6) leads to several further properties and inter-relationships which form the main thrust of this paper. In particular, we will prove the following results:

1. Let us construct the tensor

\[
R_{ab} \equiv P_{aijk} R_{bijk} \quad (8)
\]

which plays the role of generalized Ricci tensor in these theories. We will prove that \(P_{ab} = -2R_{ab}\). That is, the derivatives of an arbitrary scalar function with respect to the metric and the curvature are not independent but are connected by the curious relation:

\[
\left( \frac{\partial L}{\partial g_{im}} \right)_{R_{ijkl}} = -2R_{ijkm} \left( \frac{\partial L}{\partial R_{aijk}} \right)_{g_{ab}} \quad (9)
\]

It immediately follows from Eq. (9) that the tensor \(R_{ab}\) is symmetric. (This was noted earlier in a different context in \[5\]). This result does not follow merely from the algebraic symmetries of \(P_{abcd}\) given in Eq. (7) above. The nature of the proof given in the next section indicates that it holds only because \(P_{abcd}\) is expressible as a derivative of a scalar with respect to the curvature tensor.

2. From Eq. (9) we can also obtain several other relations connecting the partial derivatives when different independent variables like \((g^{ab}, R_{abcd})\) or \((g^{ab}, R^{a}_{bcd})\) or \((g^{ab}, R^{ab}_{cd})\) are used to describe the system. First, if we use the pair \((g^{ab}, R_{abcd})\) as independent variables, we have

\[
\left( \frac{\partial L}{\partial g^{im}} \right)_{R_{abcd}} = - \left( \frac{\partial L}{\partial g_{im}} \right)_{R_{abcd}} = 2R_{im} \quad (10)
\]

Further, if we use the pair \((g^{ab}, R^{a}_{bcd})\) as independent variables, one can show

\[
\left( \frac{\partial L}{\partial g^{im}} \right)_{R^{a}_{bcd}} = R_{im} \quad (11)
\]

and, more interestingly, if we use the pair \((g^{ab}, R^{ab}_{cd})\) as independent variables, we get

\[
\left( \frac{\partial L}{\partial g^{im}} \right)_{R^{ab}_{cd}} = 0 \quad (12)
\]

An intuitive way of understanding these results is as follows: To construct scalar polynomials of arbitrary degree in the curvature tensor, using just the curvature tensor \(R_{abcd}\) and metric \(g^{ij}\) (with index placements as indicated), one requires two metric tensors to contract out the four indices of each curvature tensor in each term. So one can understand Eq. (10) if \(L\) was a polynomial in \(R_{abcd}\). An arbitrary scalar function of curvature and metric can be thought of having (possibly infinite) series expansion in powers of the
curvature tensor. Since Eq. (10) is linear in \( L \), if it holds for an arbitrary polynomial in curvature tensor, then it should hold for arbitrary scalar functions which possess a series expansion. This intuitive argument finds strength in the result in Eq. (12) which says, for example, that in scalar polynomials constructed from \( R^{abcd} \) the metric cannot appear and all the contractions must be with Kronecker delta. This is obvious from the fact that if we use \( g^{ab} \) to contract any two lower indices, that will leave two upper indices hanging loose which cannot be contracted out because we do not have \( g^{ab} \) available as an independent variable. In fact, if we assume Eq. (12) the other results can be obtained from it without further ado.

(3) Using the fact that \( P^{bcid} \) is anti-symmetric in \( b \) and \( c \), one can show
\[
\nabla_b \nabla_c P^{bcid} = \frac{1}{2} (R^{id} - R^{di}) = 0
\]
where the last equality follows from the symmetry of \( R^{ab} \). Thus we get another counterintuitive result that, for any scalar \( L \) built from curvature tensor and the metric, we have the identity
\[
\nabla_a \nabla_b \left( \frac{\partial L}{\partial R^{abcd}} \right) g_{ij} = 0
\]

We shall now provide the proofs for these results. The discussion in this paper will be based on the metric formalism in which the metric is treated as the fundamental variable in the Lagrangian. It is possible to approach these field theories in terms of affine-metric formalism or affine formalism wherein one deals with variational principles in which metric and the connection, say, are treated as independent (see e.g., [3]). I will not discuss these approaches though it should be possible to obtain similar results in these formalisms as well.

2 Rigorous derivation of the results

2.1 The basic identities

Consider an infinitesimal diffeomorphism \( x^a \rightarrow x^a + \xi^a(x) \) which changes \( L, g_{ab} \) and \( R_{ijkl} \) by infinitesimal amounts. The idea behind the proof of the identities is to express the Lie derivative of \( L \) in two different ways and equate the results. First, because \( L \) is a scalar which depends on \( x^i \) only through \( g_{ab}(x) \) and \( R_{ijkl}(x) \), if follows that
\[
\mathcal{L}_\xi L = \xi^m \nabla_m L = \xi^m P^{ab} \nabla_m g_{ab} + \xi^m P^{abcd} \nabla_m R_{abcd} = P^{abcd} \xi^m \nabla_m R_{abcd}
\]
where we have used \( \nabla_m g_{ab} = 0 \). On the other hand, if we think of the change \( \delta L \) in \( L \), due to small changes in \( \delta g_{ab} \equiv \mathcal{L}_\xi g_{ab} \) and \( \delta R_{ijkl} \equiv \mathcal{L}_\xi R_{ijkl} \), we also know that
\[
\mathcal{L}_\xi L = P^{ab} \mathcal{L}_\xi g_{ab} + P^{ijkl} \mathcal{L}_\xi R_{ijkl}
\]
We will now work out the right hand side of Eq. (16) bringing it into a form similar to the right hand side of Eq. (15) and equate the two expressions. That will lead to the result in Eq. (9).

The first term in the right hand side of Eq. (16) is easy; using \( \mathcal{L}_\xi g_{ab} = \nabla_a \xi_b + \nabla_b \xi_a \) and the symmetry of \( P^{ab} \) we get
\[
P^{ab} \mathcal{L}_\xi g_{ab} = 2P^{ab} \nabla_a \xi_b
\]
The term involving Lie derivative of curvature tensor requires more work. The Lie derivative of the curvature tensor will have one term with the structure
\[ \xi^m \nabla_m R_{ijkl} \]
and four other terms having the structure \( R_{\nabla \xi}^{\cdot} \). On contracting with \( P_{ijkl} \) and using Eq. (7) we can combine these four terms pairwise and obtain
\[ P_{ijkl} \xi^m \nabla_m R_{ijkl} + 2 P_{ijkl} \left[ (\nabla_i \xi^m) R_{mjkl} + (\nabla_k \xi^m) R_{ijml} \right] \] (18)
Relabeling the dummy indices on \( \nabla \xi \) and using the symmetries \( R_{mlij} = R_{ijml} \) and \( P_{kjil} = P_{ilkj} \) we get
\[ P_{ijkl} \xi^m \nabla_m R_{ijkl} + 4 \left( \nabla_i \xi^m \right) R_{im} \] (19)
where we have used the definition \( R_{im} = P_{ijkl} R_{mjkl} \). Using Eq. (19) and Eq. (17) in Eq. (16) we get
\[ \xi\mathcal{L} = P_{ijkl} \xi^m \nabla_m R_{ijkl} + 2 \left( \nabla_i \xi^m \right) \left[ P_{im} + 2 R_{im} \right] \] (20)
Since \( \xi_m \) is arbitrary, it follows that the term within square brackets in the second line has to vanish, which leads to the condition in Eq. (9):
\[ P_{ijmn} = -2 P_{ijkl} R_{mjkl} = -2 R_{im} \] (21)
Equation (13) can be obtained immediately from the above result. To do this, we start with the relation:
\[ \nabla_b \nabla_c P_{bcd} = \frac{1}{2} \left[ \nabla_b, \nabla_c \right] P_{bcd} \]
\[ = \frac{1}{2} \left\{ R_{kbc}^{\cdot} P_{dckid} + R_{kbc}^{\cdot} P_{dbckid} + R_{kbc}^{\cdot} P_{dckid} + R_{kbc}^{\cdot} P_{dbckid} \right\} \] (22)
The first equality uses \( P_{bcd} = -P_{bdc} \) and second is a standard result for commutator of covariant derivatives. The first two terms vanish since \( R_{kbc}^{\cdot} = -R_{kbc}^{\cdot} = R_{kbc} \) is symmetric while \( P_{ijmn} = -P_{ijmn} \). Hence
\[ \nabla_b \nabla_c P_{bcd} = \frac{1}{2} \left\{ R_{kbc}^{\cdot} P_{dckid} + R_{kbc}^{\cdot} P_{dbckid} \right\} = \frac{1}{2} \left\{ -R_{kbc}^{\cdot} P_{dckid} + R_{kbc}^{\cdot} P_{dbckid} \right\} \]
\[ = \frac{1}{2} \left( -R_{d}^{\cdot} + R_{d}^{\cdot} \right) = 0 \] (23)
To obtain the first equality we have used the pair exchange symmetry of \( P_{ijkl} \) in both terms and to obtain the second equality we have used \( P_{dkbc} = -P_{dkbc} \). From Eq. (21), one can derive a few corollaries which are useful in expressing the field equations that arise from our Lagrangian. We first note that since \( \delta g_{ab} = -g_{ai} g_{bj} \delta g^{ij} \) we have:
\[ \left( \frac{\partial L}{\partial g_{im}} \right)_{R_{abcd}} = \left( \frac{\partial L}{\partial g_{im}} \right)_{R_{abcd}} = 2 P_{ijkl} R_{mjkl} = 2 R_{im} \] (24)
We next consider how our results change if we use the pair \( (g^m, P^a_{\cdot bcd}) \) or the pair \( (g^a_b, R^m_{\cdot ij}) \) as the independent variables. It is obvious that \( (\partial L/\partial g^m) \) with \( R_{abcd} \) held fixed is not the
same as \( \frac{\partial L}{\partial g^{im}} \) with \( R^a_{bcd} \) held fixed. To find their inter-relationship, we note that

\[
dL = \left( \frac{\partial L}{\partial g^{im}} \right)_{R_{abcd}} dg^{im} + P^{abcd}_{\alpha} dR_{abcd} = \left( \frac{\partial L}{\partial g^{im}} \right)_{R^a_{bcd}} dg^{im} + P^a_{\alpha c} d(g^{\alpha k} R_{kbcd})
\]

or

\[
p^{mbcd}_{\alpha} dR_{mbcd} + dg^{im} \left\{ \left( \frac{\partial L}{\partial g^{im}} \right)_{R^a_{bcd}} + R_{im} \right\}
\]

It follows that

\[
\left( \frac{\partial L}{\partial g^{im}} \right)_{R^a_{bcd}} = \left( \frac{\partial L}{\partial g^{im}} \right)_{R_{abcd}} - R_{im}
\]

or

\[
= 2R_{im} - R_{im} = R_{im} = P^a_{\alpha c} R_{mbcd}
\]

More interestingly, if we work with \((g^{ab}, R^{ij}_{kl})\) as the independent variables we get in place of Eq. (25) the relation:

\[
dL = \left( \frac{\partial L}{\partial g^{im}} \right)_{R_{abcd}} dg^{im} + P^{abcd}_{\alpha} dR_{abcd} = P^{abcd}_{\alpha} dR_{abcd} + dg^{im} \left\{ \left( \frac{\partial L}{\partial g^{im}} \right)_{R^a_{bcd}} + 2R_{im} \right\}
\]

leading to

\[
\left( \frac{\partial L}{\partial g^{im}} \right)_{R^a_{bcd}} = \left( \frac{\partial L}{\partial g^{im}} \right)_{R_{abcd}} - 2R_{im} = 0
\] (28)

As mentioned earlier, this result shows that scalars built from \(g^{ij}, R^{ab}_{cd}\) are actually independent of the metric tensor when \(R^{ab}_{cd}\) is held fixed.

For the sake of completeness, I should mention the following subtlety as regards the very first equation Eq. (15) which, I believe, the reader thought was obvious; but it requires some pedagogical discussion. Consider a scalar function \(f(u^i, A_{ij}, B_{jk}^i, \ldots)\) built out of several tensorial quantities \(u^i(x), A_{ij}(x), B_{jk}^i(x), \ldots\) We assume that the \(f\) depends on \(x^i\) only implicitly through these tensorial objects. It then follows from standard laws of calculus that the partial derivative \(\partial_i f\) can be expressed in terms of the partial derivatives of \(f\) with respect to \(u^i, A_{ij}, B_{jk}^i, \ldots\) and the partial derivatives of the tensorial independent variables. Can one write a similar relation using covariant derivatives rather than partial derivatives? It turns out that one indeed can though the result is not trivial because the simultaneous validity of the two results — one involving ordinary derivatives and one involving covariant derivatives — again leads to certain identities amongst the derivatives of \(f\) with respect to the tensorial independent variables. The validity of the ‘function of the function’ rule for covariant differentiation — which again I have not seen discussed explicitly in text books — can be proved in two ways: (a) In the local inertial frame, the relation with covariant derivatives reduces to the one with partial derivatives, the validity of which is obvious from ordinary laws of calculus. The general covariance of the expression assures the validity of the result in any coordinate system. (b) A more explicit proof which explains the “mechanism” behind the result begins by noticing that to construct a scalar out of tensorial quantities one needs to first construct a set of scalars involving products of the tensorial quantities by contracting on all indices like, for example, \(\phi_1 = u^i u^j A_{ij}, \phi_2 = B_{ij}^k A_{jk} u^i, \ldots\), and then construct our scalar as a function \(f(\phi_1, \phi_2, \ldots)\) of primitive scalars which only involve products of tensorial quantities. We know, of course, that any derivative (partial, covariant, Lie) obeys the product rule. Given the fact that any scalar \(f\) has to be a function of primitive scalars \((\phi_1, \phi_2, \ldots)\) which must be products of the
tensorial quantities, we obtain the function of the function differentiation rule for covariant derivatives of scalars from the product rule of covariant differentiation. Once again we see the importance of combinatorics ensuring that all the indices must be contracted in the construction of scalars.

2.2 Field equations in the generalized theory of gravity

To obtain the field equation by varying the Lagrangian, the most convenient set of variables to use happens to be the pair \((g^{im}, R^a_{bcd})\). This is because \(R^a_{bcd}\) can be expressed entirely in terms of \(\Gamma^i_{jk}\) without the use of a metric which makes the variations simple to perform. As mentioned earlier, the equations of motion obtained by varying \(\sqrt{-g} L\) can be expressed in the form

\[
E_{ab} = (1/2) T_{ab}
\]

where \(T_{ab}\) is the energy-momentum tensor of matter and

\[
\sqrt{-g} E_{ab} = \left( \frac{\partial L}{\partial g^{ab}} \right) R_{ijkl} - 2 \sqrt{-g} \nabla^m \nabla^nP_{amnb}
\]

which does not involve the derivatives of the Lagrangian. This is the form which I believe is most useful for computational purposes, though in the literature one often finds more complicated expressions possibly because the issues addressed above were not clearly understood.

The equation \(E_{ab} = (1/2) T_{ab}\) with a symmetric energy momentum tensor requires \(E_{ab}\) to be symmetric and have already proved that \(R_{ab}\) is symmetric. This implies that the last term \(\nabla_m \nabla_n P_{amnb}\) in Eq. (29) is also symmetric in \(a\) and \(b\). It is easy to see that the anti-symmetric part of this tensor is given by

\[
\nabla_m \nabla_n \left( P_{amnb} - P_{bmna} \right) = \nabla_m \nabla_n \left( P_{amnb} + P_{panb} \right) = -\nabla_m \nabla_n P_{abmn}
\]

In arriving at the first equality we have used the pair exchange symmetry and the anti-symmetry of the first two indices \(P_{ijkl}\); in arriving at the second equality we have used the cyclic relation \(P_{ijkl} = 0\). It follows that the last term in Eq. (30) is symmetric only if the right hand side of Eq. (31) vanishes which is assured by Eq. (23) and the symmetry of \(P_{abcd}\) under pair exchange. Thus each of the terms in Eq. (30) is individually symmetric.

From the expression Eq. (2) for the variation of the action, one can obtain a generalized Bianchi identity satisfied by \(E^a_b\). If we consider the variation \(\delta g^{ab} = \nabla^a \xi^b + \nabla^b \xi^a\) arising from the coordinate change \(x^a \to x^b + \xi^a\) and assume that \(\xi^a\) and its derivatives vanish sufficiently fast close to the boundary of the region, we get the result

\[
\nabla_a E^a_b = \nabla_a \left[ R^a_b - \frac{1}{2} \delta^a_b L - 2 \nabla^m \nabla^n P_{mnab} \right] = 0
\]

Once again it is not obvious that the expression in the square bracket is divergence-free, though it is. Explicit covariant differentiation of the expression will again lead to derivatives of \(L\) and the generalized Bianchi identity will be satisfied for reasons similar to the one described above.

7
The first two terms in Eq. (30) for $E_{ab}$ involve at most second derivatives of the metric tensor with respect to the coordinates while the third term $\nabla \nabla P$ could involve up to fourth derivative of the metric tensor with respect to the coordinates. A very important subclass of generalized theories of gravity is obtained by ensuring that the field equations do not contain more than second order derivatives. Obviously, such a class of theories can be obtained from the Lagrangian which satisfy the additional constraint $\nabla_a P^{abcd} = 0$. Given the symmetries of $P^{abcd}$, it follows that $\langle \partial L/\partial R_{abcd} \rangle$ must be a divergence-free tensor in all indices. In such a case, the field equations simplify significantly and we get

$$E_{ab} = R_{ab} - \frac{1}{2} g_{ab} L = P_\alpha^{ijk} R_{\beta jk} - \frac{1}{2} g_{ab} L = \frac{1}{2} T_{ab}$$

(33)

The similarity with Einstein’s theory is obvious. In this context we have constructed from $L$ a second rank tensor $E_{ab}$ which (a) is symmetric, (b) divergence-free and (c) does not contain more than second derivatives of the metric. It can be proved [8] that the most general tensor which satisfies these criteria is the Lanczos-Lovelock tensor obtained from the Lanczos-Lovelock Lagrangian. Therefore, the theory in which $P^{abcd}$ is divergence-free is the same as Lanczos-Lovelock models of gravity and the corresponding $L$ must be the Lanczos-Lovelock Lagrangian.

All these theories possess a Noether current which is conserved off-shell due to the diffeomorphism invariance of the Lagrangian. The identities obtained above have important implications for the structure of Noether current and for derivation of field equations in the emergent paradigm [1, 6, 7] starting from the Noether current. This will be examined in a separate publication.

3 Some further curiosities and an alternative derivation

While the analysis given above seems both algebraically and logically straightforward, there are some subtleties about these results which we will now discuss. First, one would have thought that some physical meaning or intuitive content could be attributed to the result in Eq. (23) which looks strikingly similar to some kind of Bianchi identity. Unfortunately, I could not find any simple interpretation of this very general result. Obviously, the result relies on the symmetries of $R_{ab}$ which, in turn, arises from the relationship in Eq. (9) and does not have a simple intuitive interpretation. This aspect deserves further examination. (This result is, of course, related to the generalized Bianchi identity, $\nabla_a E^{ab} = 0$ but I could not make the connection self-evident.)

The second subtlety about these results is the occurrence of the factor $\nabla \xi_m$ in Eq. (19) which allows us to obtain Eq. (9). This is surprising for the following reason: We know that, under the diffeomorphism $x^i \to x^i + \xi^i$, the metric changes by

$$\delta g_{ab} = -[\nabla_a \xi_b + \nabla_b \xi_a] = -\xi^i g_{ab}$$

(34)

showing that $\delta g_{ab}$ and $\xi^i g_{ab}$ depend only on the symmetric part of $\nabla \xi_m$. It immediately follows that $\delta R_{ijk}$ as well as $\delta R_{abcd}$ can depend only on the symmetric part of $\nabla \xi_m$. Therefore, the left hand side of Eq. (19) can depend only on the symmetric part of $\nabla \xi_m$.

Note that if we had only the symmetric part of $\nabla \xi_m$ in the second term of Eq. (19) — rather than $\nabla \xi_m$ itself — we could not have obtained the result in Eq. (21); instead we would have found that the symmetric part of $R^m$ is related to $P^m$, and one cannot prove the symmetry of $R^m$. It is only the separation of the left hand side of Eq. (19) into
two terms on the right hand side with a very specific structure for the first term, which
introduces $\nabla_i \xi_m$ in Eq. (19). In other words, it is the first term in the right hand side of
Eq. (19) which breaks the inherent symmetry and allows us to obtain the result which, of
course, is closely related to the fact that $L$ is a scalar. So, this subtlety, fortunately finds
a resolution in the above algebraic fact.

It is nevertheless worth checking this fact explicitly by evaluating $L_\xi R_{abcd}$ in terms
of $L_\xi g_{ab}$ and rewriting the expression to obtain Eq. (19). Since the variations in the
curvature tensor and the metric are related in the form $L_\xi R = \nabla\nabla (L_\xi g)$ we would expect
$P \nabla_\xi R = P \nabla_\xi \nabla_\xi \xi$, which can depend only on the symmetric part of $\nabla_\xi \xi$ (when we re-
introduce indices correctly!). But at this stage it has no $P \xi \nabla R$ type of terms which is
the first term in the right hand side of Eq. (19). When we commute the $\nabla$s properly and
use the identities obeyed by the curvature tensor, one can rewrite this expression with a
$P \xi \nabla R$ term separated out. But then, the remaining terms depend on full $\nabla_\xi$ rather than
on the symmetric part alone!

This is to be expected because the two ways of computing the Lie derivative of curvature
tensor must give the same result. The first way is to treat it as a fourth rank tensor which
will indeed introduce a $\xi \nabla R$ term in the Lie derivative almost by definition. The second
way is to express $R$ in terms of the metric and compute the result which has to come
from an expression with the structure $\nabla \nabla \nabla \xi$. Either way, the result for $L_\xi R_{abcd}$ can only
depend on the symmetric part of $\nabla_\xi \xi$. But if we subtract out the $\xi \nabla R$ term, the remaining
terms, of course, can depend on the full $\nabla_\xi$ rather than on its symmetric part. This is
what happens.

We will now outline the steps in this derivation for the sake of completeness. It is
easier to start with the expression for $P_{a}^{bcd} \delta R_{bcd}$ rather than from $P^{abcd} \delta R_{abcd}$. (These
two are, of course, not the same when the metric is varied; see Eq. (26) below). We begin
with the standard result (see, e.g.,[2])

$$- P_{a}^{bcd} \delta R_{bcd} = -2 P^{ibcd} \nabla_i \nabla_b \delta g_{di} = (2 P^{ibcd} \nabla_i \nabla_b \nabla_d \xi_l + 2 P^{ibcd} \nabla_i \nabla_b \nabla_c \xi_d)$$  (35)

In the triple $\nabla$s we want to commute the indices so that the anti-symmetry of $P^{ijkl}$ can be utilized. Doing this and using $[\nabla_i, \nabla_j] \xi^l = R^l_{kij} \xi^k$ we get

$$- P_{a}^{bcd} \delta R_{bcd} = 2 P^{ibcd} \nabla_c \left( \nabla_d \nabla_b + [\nabla_b, \nabla_d] \right) \xi_l + P^{ibcd} \nabla_c \left( [\nabla_b, \nabla_i] \xi_d \right)$$

$$= P^{ibcd} \nabla_c \left( \nabla_b \xi_l \right) - P^{ibcd} \nabla_c \left( 2 R^m_{ibd} + R^m_{dbi} \right) \xi_m$$  (36)

The second term in the right hand side can be replaced by $- P^{ibcd} \nabla_c \left( 2 R^m_{dbi} \xi_m \right)$. To see this, we use the anti-symmetry of $P^{abcd}$ to write $2 R^m_{ibd}$ as the sum of $R^m_{ibd}$ and $-R^m_{bid} = R^m_{dbi}$, and use the cyclic property of the curvature tensor. This gives

$$P^{ibcd} \nabla_c \left( 2 R^m_{ibd} + R^m_{dbi} \right) \xi_m \right) = P^{ibcd} \nabla_c \left( R^m_{ibd} + R^m_{dbi} + R^m_{dbi} \right) \xi_m$$

$$= P^{ibcd} \nabla_c \left( -R^m_{dbi} + R^m_{dbi} \right) \xi_m$$

$$= P^{ibcd} \nabla_c \left( 2 R^m_{dbi} \xi_m \right)$$  (37)

This will expand to a structure like $P \xi \nabla R + P R \nabla \xi$. In the term involving $\xi \nabla R$ we want
to bring in the directional derivative $\xi_m \nabla_m$. This can be done using the result

$$2 P^{ibcd} \xi_m \nabla_c R^m_{dbi} = - P^{ibcd} \xi_m \nabla_m R_{ibcd}$$  (38)
The proof of Eq. (38) is similar to that for Eq. (37). We have:

\[
2P^{abcd} \nabla_c R_{mdbi} = P^{abcd} (\nabla_c R_{mdbi} - \nabla_d R_{mbci}) = P^{abcd} (\nabla_c R_{bimd} - \nabla_d R_{binc})
\]

In arriving at the last step we have used

\[
P^{abcd} \nabla_m R_{abcd}
\]

and cancelled out one term using the anti-symmetry in \(d\) and \(m\). Plugging all these back, we get

\[
- \nabla_d R_{binary} = \nabla_c R_{badm} + \nabla_m R_{bcd}
\]

The first term is in the required form and the second term does depend on \(\nabla_m \xi\) rather than on its symmetric part alone. We will now simplify this term a little bit.

In spite of appearances, we know that the right hand side can depend only on the symmetric part of \(\delta R_{ijkl}\) which matches with Eq. (19). Once again, from the very nature of the derivation, we know that the right hand side can depend only on the symmetric part of \(\nabla \xi\). But the algebra was simpler. The derivation of Eq. (45) is more involved but makes clear that the right hand side can only depend on the symmetric part of \(\nabla \xi\).

In deriving either of these, Eq. (21) or Eq. (45), we have only used the symmetries of \(P^{abcd}\) but not the fact that it is the derivative of a scalar with respect to the curvature tensor. It is this vital input (plus the facts that \(L\) has no explicit dependence on \(x^i\) and \(\nabla g = 0\)) which allows us to separate out the first term and obtain the condition in Eq. (21).
Acknowledgments

I thank D. Kothawala, D. Lynden-Bell, A. Paranjape for useful and extensive discussions on these results and Hamsa Padmanabhan and Krishna Parattu for comments on an earlier draft. This work was completed while I was visiting Institute of Astronomy, Cambridge during May-July, 2011 and the hospitality of IOA, Cambridge is gratefully acknowledged. My research is partially supported by J.C. Bose Fellowship of DST, India.

References

[1] T. Padmanabhan, Rep. Prog. Phys., 73, 046901 (2010) [arXiv:0911.5004]; T. Padmanabhan, Lessons from Classical Gravity about the Quantum Structure of Spacetime, J.Phys. Conf.Ser. 306 012001 (2011) [arXiv:1012.4176].

[2] For a textbook derivation, see page 657 of T. Padmanabhan, Gravitation: Foundations and Frontiers, (Cambridge University Press, 2010).

[3] G. J. Olmo, 2011, Int.J.Mod.Phys, D20, 413 [arXiv:1101.3864]; S. Capozziello, M. De Laurentis, 2011, Extended Theories of Gravity, [arXiv:1108.6266].

[4] Wald R M 1993 Phys. Rev. D 48 3427 [gr-qc/9307028]; Iyer V and Wald R M 1995 Phys. Rev. D 52 4430 [gr-qc/9503052].

[5] J. Koga and K. Maeda, Phys. Rev. D 58, 064020 (1998)

[6] T. Padmanabhan, A Physical Interpretation of Gravitational Field Equations, AIP Conference Proceedings, 1241, 93-108 (2010) [arXiv:0911.1403]

[7] T. Padmanabhan, Phys.Rev., D 81., 124040 (2010) [arXiv:1003.5665]

[8] Lanczos C Z. Phys., 73, 147 (1932); Lovelock D J. Math. Phys. 12 498 (1971); D. Lovelock and H.Rund Tensors, Differential forms and Variational Principles (Dover, 1989).