Perturbations and moduli space dynamics of tachyon kinks

Hindmarsh, Mark and Li, Huiquan (2008) Perturbations and moduli space dynamics of tachyon kinks. Physical Review D, 77 (6). ISSN 1550-7998

This version is available from Sussex Research Online: http://sro.sussex.ac.uk/id/eprint/27328/

This document is made available in accordance with publisher policies and may differ from the published version or from the version of record. If you wish to cite this item you are advised to consult the publisher’s version. Please see the URL above for details on accessing the published version.

Copyright and reuse:
Sussex Research Online is a digital repository of the research output of the University.

Copyright and all moral rights to the version of the paper presented here belong to the individual author(s) and/or other copyright owners. To the extent reasonable and practicable, the material made available in SRO has been checked for eligibility before being made available.

Copies of full text items generally can be reproduced, displayed or performed and given to third parties in any format or medium for personal research or study, educational, or not-for-profit purposes without prior permission or charge, provided that the authors, title and full bibliographic details are credited, a hyperlink and/or URL is given for the original metadata page and the content is not changed in any way.
Perturbations and moduli space dynamics of tachyon kinks

Mark Hindmarsh and Huiquan Li

Department of Physics and Astronomy, University of Sussex, Brighton BN1 9QH, United Kingdom

(Received 2 January 2008; published 6 March 2008)

The dynamic process of unstable D-branes decaying into stable ones with one dimension lower can be described by a tachyon field with a Dirac-Born-Infeld effective action. In this paper we investigate the fluctuation modes of the tachyon field around a two-parameter family of static solutions representing an array of brane-antibrane pairs. Besides a pair of zero modes associated with the parameters of the solution, and instabilities associated with annihilation of the brane-antibrane pairs, we find states corresponding to excitations of the tachyon field around the brane and in the bulk. In the limit that the brane thickness tends to zero, the support of the eigenmodes is limited to the brane, consistent with the idea that propagating tachyon modes drop out of the spectrum as the tachyon field approaches its ground state. The zero modes, and other low-lying excited states, show a fourfold degeneracy in this limit, which can be identified with some of the massless superstring modes in the brane-antibrane system. Finally, we also discuss the slow motion of the solution corresponding to the decay process in the moduli space, finding a trajectory which oscillates periodically between the unstable D-brane and the brane-antibrane pairs of one dimension lower.

DOI: 10.1103/PhysRevD.77.066005 PACS numbers: 11.25.Uv

I. INTRODUCTION

In type II string theories, there are two kinds of branes: stable or BPS branes, which are supersymmetric and charged, and unstable or non-BPS branes, which are non-supersymmetric and uncharged. The $p$-dimensional unstable D ($Dp$) branes eventually decay into stable $D(p-1)$ ones. This decay process is described by the dynamics of a tachyon field $T$. The effective action of this unstable brane system in low energy approximation is conjectured to be of the Dirac-Born-Infeld (DBI) type form [1–3], as derived from string theory:

$$ S = -\int d^{p+1}x \sqrt{-g} V(T) \sqrt{1 + \partial_{\mu} T \partial^{\mu} T}, \quad (1.1) $$

where the effective potential $V(T)$ takes its maximal value at $T = 0$ and vanishes asymptotically as $T$ tends to infinity.

The homogeneous decay of the tachyon field involves the field rolling to its vacuum at $T = \pm \infty$, towards a state which can be characterized as a pressureless fluid without propagating modes. This is consistent with the notion that the open string states disappear from the spectrum as the brane decays.

This decay can lead to a period of inflation [4,5], although realistic models are not easy to realize [6–10]. One reason is that the tachyon field perturbations around the unstable vacuum must contain an unstable mode whose (negative) eigenvalue is the tachyon mass squared, which is of order the string scale. This leads to an $\eta$ problem in tachyonic inflation [6,11].

In order to study the dynamics away from the unstable vacuum, at least two supposed potential forms are often used: $V(T) = V_m/\cosh(\beta T)$ (where $\beta = 1$ for the bosonic string and $\beta = 1/\sqrt{2}$ for superstrings), and $V(T) = \exp(-T^2/2)$. Both potentials lead to the correct value of the mass of the tachyon on the D-branes. Here, we chose the former because it contains as a static classical solution a periodic array of static solitonic kink-antikink solutions, which is known to exist as an exact solution in the deformed conformal field theory [12–15]. These kinks have exactly the right tension to be understood as branes and antibranes with one fewer dimension.

A similar effective action with a complex tachyon field describes an unstable brane-antibrane system decaying to a network of stable branes with two fewer dimensions, and in the case of D3-D3 system, has been used to realize a version of hybrid inflation terminating in the production of cosmic superstrings [16–18].

In this paper we examine the low energy effective field theory following from the tachyon effective action in the background of static brane-antibrane pairs, neglecting the gravitational and Ramond-Ramond fields. In the limit that the thickness of the branes tends to zero, and the tachyon field between them approaches the vacua at $T = \pm \infty$, we should expect interactions between the branes to vanish and the resulting spectrum should represent the effective field theory around isolated branes.

The fluctuation spectrum in an effective action $S = -\int d^{p+1}x \sqrt{-g} V(T) (1 + \partial_{\mu} T \partial^{\mu} T)$, with the potential $V = \exp(-T^2/2)$, has been discussed in [19], noting that the Regge slope $\alpha' = 1$. We find a number of similarities and differences. There are zero modes associated with the translational degree of freedom of the system, and also the arbitrary maximum value of $|T|$ between the branes. We have numerical evidence that the lowest excited states have a fourfold degeneracy in the limit that the brane thickness vanishes. We find that the modes disappear from the bulk as the field tends to the vacuum between the branes, consistent with the absence of propagating modes of the tachyon field in the homogeneous case. In the same limit,
4 degenerate zero modes of the tachyon field appear, two of which can be identified with the transverse fluctuations of the branes.

Finally we study the dynamics of the zero mode corresponding to changes in \( T_0 \), the maximum value of \([T]\) between the branes, deriving the effective action for this modulus. We find a family of slow-motion solutions corresponding to the field oscillating between the unstable \( Dp \) brane and the stable \( D(p-1) \) \( D(p-1) \) pair.

This also has an interesting implication for the dynamical production of \( D(p-1) \)-branes. In the unstable system, small fluctuations drive the tachyon field \( T(x) \) to roll from the top of its potential \( V(T) \). \( D(p-1) \)-branes form as topological defects (kinks or solitons) due to symmetry breaking [20–22]. The time-dependent tachyon field generically reaches a singularity in finite time, either at the kinks [23,24] or on caustics between the kinks [25]. Our slow-motion solution of the time-dependent modulus in the moduli space reaches infinite gradient at the kink in finite time, consistent with Refs. [23,24], but can be evolved through the singularity.

The paper is constructed as follows. We first discuss the static solutions of the tachyon equation of motion derived from the DBI effective action in Sec. I. Based on the solution, fluctuation modes both in approximate analysis and in general form are discussed in Sec. II. In Sec. III, we further investigate the motion in the moduli space. Section IV gives the conclusions of the paper.

II. STATIC SOLUTIONS TO THE FIELD EQUATIONS

The equation of motion derived from the DBI action is

\[
\Box T + (g^{\mu \nu} \partial T \cdot \partial T - \partial \mu T \partial \nu T) \nabla_\mu \nabla_\nu T = -\frac{V'}{V} (1 + \partial T \cdot \partial T) = 0. \tag{2.1}
\]

We are interested in time-independent solutions, neglecting the gravitational and Ramond-Ramond fields of the brane. It is known that there are solutions depending on only one space coordinate, which we denote \( x \) [19,26,27]. In this case, the static but inhomogeneous field equation reduces to

\[
(\partial_x T)^2 = \left( \frac{V'}{V_0} \right)^2 - 1, \tag{2.2}
\]

where \( V_0 \) is the minimum value of \( V \) attained by the potential in the solution. It is not zero providing \( T \) remains finite. The solution is

\[
T(x) = \frac{1}{\beta} \arcsinh \left[ \sin(\beta x - x_m) \sqrt{\left( \frac{V_m}{V_0} \right)^2 - 1} \right]. \tag{2.3}
\]

which is a regular array of kinks and antikinks, with period \( 2\pi/\beta \), located at \( x = 2n \pi/\beta + x_m \) and \( x = (2n - 1) \pi/\beta + x_m \) (\( n \) is integer), respectively. \( x_m \) is set to be the positions where the maximum potential \( V_m \) is located.

The energy density is peaked near the zeros of \( T \), where \( V = V_m \), and in the limit that \( V_0 \to 0 \) the peaks become infinitely thin, and we identify the kinks with daughter branes and antibranes with one dimension lower. The minimum potential \( V = V_0 \) and the maximum amplitude of the tachyon field \( |T_0| = \arccosh(V_m/V_0)/\beta \) is reached at \( x_0 = (n + \frac{1}{2}) \pi/\beta + x_m \), which is the closest approach to the vacuum of the system.

The energy of this system is

\[
E = \int d^p x V(T) \sqrt{1 + (\partial_x T)^2} = \frac{1}{V_0} \int d^{p-1} x \int dV'. \tag{2.4}
\]

with the specific potential adopted in this paper, where we remind the reader that \( \beta = 1 \) for the bosonic string and \( \beta = 1/\sqrt{2} \) for superstring. The tension of the parent \( Dp \)-brane is \( T_p = V_m \), so this is the correct tension for a \( Dp - 1 \)-brane, \( T_{p-1} \).

III. FLUCTUATION MODES

We study the fluctuation modes by linearizing the variation, \( T(t,x) \to T + \delta T(x)e^{-i\omega t} \). Following the procedure in [28], which makes the substitution \( dx = (V_0/V)dz \) and \( \delta T = (V/V_0)\phi \), we have

\[
\left( -\frac{d^2}{dz^2} + U(z) \right) \phi = \omega^2 \phi, \tag{3.1}
\]

where

\[
U(z) = \frac{1}{\phi_0} \frac{d^2 \phi_0}{dz^2}. \tag{3.2}
\]

Here, \( \phi_0 \approx (dT/dx)\sqrt{V_0/V} \) is the zero mode (\( \omega^2 = 0 \)). We can get different eigenstates with different boundary conditions and their corresponding eigenvalues \( \omega^2 \) from this equation.

More precisely, we can get the simple form of \( U(z) \):

\[
U(z) = \frac{\beta^2}{4} \left[ 1 + \left( \frac{V_0}{V_m} \right)^2 - 3 \left( \frac{V(z)}{V_m} \right)^2 + \left( \frac{V_0}{V(z)} \right)^2 \right]. \tag{3.3}
\]

The maximum value of \( U(z) \) is at \( V(z) = \sqrt{V_0} \), where \( U = \beta^2/4 + O((V_0/V_m)^2) \). The minimum potential \( U(z) = -\beta^2/2 + O((V_0/V_m)^2) \) happens at both \( V(z) = V_m \) and \( V(z) = V_0 \). With the fluctuation equation, we can discuss the fluctuation modes and the discrete spectrum \( \omega^2 \).
PERTURBATIONS AND MODULI SPACE DYNAMICS OF...

A. Approximate fluctuation modes near kinks

We start by studying the behavior near the kink or brane, where \( T(z) \approx 0 \) and \( V(z) \approx V_m \) as mentioned above. We will assume that the kinks are well separated in comparison to their width, which implies that \( V_0/V_m \ll 1 \). In the limit \( V(z) \to V_m \), \( dT/dz \approx \pm 1 \) since \( dT(z)/dz = (dT/dx) \times (dx/dz) = \pm \sqrt{1 - (V_0/V)^2} \). Hence, near a kink, \( T(z) \approx \mp (z + c) \) and \( V(z) \approx V_m/cosh(\beta(z + c)) \), where \( c \) is a constant. Therefore, we have

\[
\frac{d^2 \phi}{dz^2} + \left[ \omega^2 + \frac{\beta^2}{4} \left( \frac{3}{4} - \frac{1 - \omega^2}{\beta^2} \right) \right] \phi = 0,
\]

which is a well-explored equation. Following standard procedures [29] we can recast the equation by setting \( \xi = \tanh(\beta(z + c)) \):

\[
\frac{d}{d\xi} \left( 1 - \xi^2 \right) \frac{d}{d\xi} \phi + \left[ \frac{3}{4} - \frac{1 - \omega^2}{\beta^2} \right] \frac{1}{1 - \xi^2} \phi = 0,
\]

which has solutions

\[
\phi = (1 - \xi^2)^{\epsilon/2} F \left[ \epsilon - \frac{1}{2}, \epsilon + \frac{3}{2}; \frac{1}{1 - \xi^2} \right],
\]

where \( \epsilon = \sqrt{1 - \omega^2/\beta^2} \).

If \( \phi \) is finite, we have that \( m = 1/2 - \epsilon \) must be a non-negative integer. The condition \( \epsilon \geq 0 \) or \( \omega^2 < \beta^2/4 \) gives that \( m \) can only be zero, i.e., \( \omega^2 = (1 - m)\beta^2 = 0 \). Thus there is only a zero fluctuation mode with \( \omega^2 < \beta^2/4 \), the maximum of the potential \( U(z) \). This mode is of course just the zero mode associated with the translational parameter of the solutions. Modes with \( \omega^2 < \beta^2/4 \) are bound states, confined to the kink, and so we conclude that there are no bound states other than the zero mode, at least in the thin kink limit.

B. General fluctuation modes

To obtain the generalized solutions, we need to know the form of the \( z \)-dependent potential \( V(z) = V(T(x(z))) \). From \( dz = (V(x)/V_0)dx \), we find

\[
\int dz' = -\int \frac{1}{\beta V_0 \sqrt{(V/V_0)^2 - 1/\sqrt{1 - (V/V_m)^2}}} dV.
\]

With the reparametrization \( t = V/V_0 \in [1, t_m = V_m/V_0] \), it leads to

\[
\int_{z_0}^{z} dz' = -\int_{t}^{t_m} \frac{i}{\beta} \frac{1}{\sqrt{1 - t^2} \sqrt{1 - k^2 t'^2}} dt'.
\]

The integral can be evaluated in terms of an elliptic function of the first kind: \( F(\varphi, K) = \int_{0}^{\sin \varphi} dx (1 - x^2)^{-1/2} (1 - k^2 x^2)^{-1/2} \), where the parameter \( k = V_0/V_m \). With \( t = \sin \varphi \), we have \( \beta(z - z_0) = \beta \Delta z = -i [F(\varphi, K - K(k)) \]
or we can write

\[
F(\varphi, k) = i \beta \Delta z + K(k),
\]

where \( K(k) \) is the complete elliptic function of the first kind: \( K(k) = F(\pi/2, k) \). Its inverse is one of the Jacobi functions,

\[
t = V(z)/V_0 = sn(i \beta \Delta z + K(k), k).
\]

Using the identities of the Jacobi functions [30]

\[
\frac{dn(i \beta \Delta z + K(k), k)}{dn(i \beta \Delta z, k)} = \frac{1}{dn(\beta \Delta z, k')},
\]

where \( k' = \sqrt{1 - k^2} \), we can bring \( U(z) \) into the form

\[
U(z) = \frac{\beta^2}{4} \left[ 2 - k^2 - 3 \left( \frac{k^2}{dn^2(\beta \Delta z, k')} + dn^2(\beta \Delta z, k') \right) \right].
\]

Notice that, in the following numerical calculations, we just adopt \( z \) instead of \( \Delta z \) by setting \( z_0 = 0 \).

The curves of \( T(z), V(z) \) and \( U(z) \) with \( k = 0.01 \) are depicted in Fig. 1. They are all periodic, with periods \( 4K(k)/\beta, 2K(k)/\beta \) and \( K(k)/\beta \), respectively, which comes from \( dn(u) \) being an even function with period \( z = 2K(k)/\beta \).

Since \( K \) increases with \( k \), the separations between two neighboring branes becomes larger as \( k \) decreases when taking \( z \) as the coordinate, while remaining constant in the \( x \) coordinate. The daughter stable \( Dp - 1 \)-branes and \( \text{anti-Dp} - 1 \)-branes are located at \( z = (4n + 1)K/\beta \) and \( z = (4n - 1)K/\beta \), respectively, where the tachyon field \( T(z) \) vanishes and the potential \( V(z) \) peaks. The closest approach to the vacuum occurs at \( z = 2nK/\beta \), where \( V(z) \)

FIG. 1 (color online). Plots of \( T(z), V(z) \) and \( U(z) \) within a period of the tachyon field \([ -2K(k)/\beta, 2K(k)/\beta ] \). Here \( k = 0.01 \).
gets its minimum value $kV_m$. Note that the plots present only one period of the solution $[-2K(k)/\beta, 2K(k)/\beta]$.

From the figure, we can see that $T(z)$ is indeed linearly dependent on $z$ for most part of the numerical curve, especially near the kink positions. The smaller the value of $k$, the larger the proportion of the solution occupied by the linear section, confirming the validity of the approximation adopted in the previous subsection.

C. Numerical solution of eigenvalue equations

The periodicity of the solution and the potential for the fluctuations means that we need consider the range $0 \leq z \leq K/\beta$, with eigenfunction taking either Dirichlet (D) or von Neumann (N) boundary conditions at each boundary. Hence there are four boundary conditions (BCs) which are listed and labeled in Table I.

We solve the eigenvalue equations by numerical integration, “shooting” for the correct boundary condition at $z = K/\beta$, using the 4th/5th order Runge-Kutta integrator built into MATLAB, ode45. We adopted an absolute tolerance of $10^{-9}$ and a relative tolerance of $10^{-6}$ for all calculations in this paper.

1. Fluctuation modes of low-lying states

We first set $k = 0.01$. In Figs. 2–4, normalized numerical solutions are presented in groups of four, whose eigenvalues are close together. We refer to them as zero mode solutions are presented in groups of four, whose eigenfunctions are close together. We refer to them as zero mode solutions.

We denote normalized solutions by a tilde, and plot values are close together. We refer to them as zero mode solutions.

| BC(1): ND | BC(2): NN | BC(3): DD | BC(4): DN |
|-----------|-----------|-----------|-----------|
| $\phi'(0) = 0$ | $\phi'(0) = 0$ | $\phi(0) = 0$ | $\phi(0) = 0$ |
| $\phi(K/\beta) = 0$ | $\phi(K/\beta) = 0$ | $\phi'(K/\beta) = 0$ | $\phi'(K/\beta) = 0$ |
| $\phi(0) = 1$ | $\phi(0) = 1$ | $\phi'(0) = 1$ | $\phi'(0) = 1$ |

Similarities and symmetries are clearer when looking at $\phi(z)$, as $\delta T(z)$ has a factor $V(z)/V_0$ which does not share the periodicity of the eigenfunctions, and suppresses the periodicity of the eigenfunctions, and suppresses
FIG. 3 (color online). Normalized numerical solutions, \( \delta \tilde{T}(z) \) and \( \phi(z) \), of the first excited states with \( k = 0.01 \) under the four boundary conditions: (a) BC(1), (b) BC(2), (c) BC(3) and (d) BC(4). Note that BC(1) and BC(4) have the same eigenvalue.

FIG. 4 (color online). Normalized numerical solutions, \( \delta \tilde{T}(z) \) and \( \phi(z) \), of second excited states with \( k = 0.01 \) under the four boundary conditions: (a) BC(1), (b) BC(2), (c) BC(3) and (d) BC(4). BC(1) and BC(4) have the same eigenvalue.
function away from the kinks. For example, eigenfunctions with BC(1) and BC(4) are related by the symmetry $z \rightarrow K/2\beta - z$, and therefore the eigenvalues are the same.

Because of the factor $V(z)/V_0$, the fluctuations away from the kinks are greatly suppressed when plotted in terms of $\delta T$, and the eigenfunctions look very similar near the (anti)kink at $z = K(k)/\beta$.

Figure 2 contains the four states whose eigenvalues are very close to zero. Indeed, two with BC(1) and BC(4), and eigenvalues $\omega^2 = -1.0672 \times 10^{-8}$ and $\omega^2 = -2.0095 \times 10^{-9}$, are consistent with being numerical approximations to the true zero modes corresponding to the degeneracies in position $x_0$ [Fig. 2(a)] and amplitude $T_0$ [Fig. 2(b)] of the solution. Not only do they have very small eigenvalues, they also have the correct number and position of nodes.

We also have two low-lying modes with BC(2), one with a small negative eigenvalue, which we call $0^-$, and the other with a small positive one $0^+$. The negative eigenvalue reflects an instability in the system towards annihilation, where the kink-antikink pair moves together. There is no low-lying state with BC(3).

We can understand at least the ordering of the eigenvalues by studying the number of nodes of the eigenfunctions, using the standard result that more nodes means a higher eigenvalue.

We note first that the background has periodicity $4K(k)$, in which the eigenfunctions must also be periodic, and therefore have an even number of nodes. The lowest eigenvalue is associated with BC(2), which allows an eigenfunction with no nodes. BC(1, 4) allow only 2 nodes, which are the zero modes. BC(2) also allows the 4-node eigenfunction shown in Fig. 2(d). BC(3) permits a minimum of 4 nodes which, because of its Dirichlet boundary conditions, are forced to be at the minimum of the potential $U(z)$, and it must therefore have a higher eigenvalue than the 4-node BC(2) eigenfunction, which peaks where the potential is lowest.

2. Fluctuations with smaller $k$

Stable BPS daughter $D(p - 1)$-branes form as $V_0$ approaches zero at the end of the tachyon condensation, so we should check the fluctuation behavior as $k \rightarrow 0$. However, we were unable to get numerical solutions with very small $k$. In Fig. 5, we show the zero mode fluctuation with $k = 10^{-6}$. Here it is even more obvious that fluctuations have support mostly near the brane position, both in $\phi$ and $\delta T$. The higher modes have similar properties. Therefore, we can expect that the fluctuations are confined to the branes more and more strongly when $k$ or $V_0$ tend to zero. We can interpret this as being consistent with the observation [27] in the homogeneous case that there are no propagating tachyon modes ($\omega^2 \rightarrow 0$) in the vacuum when $V_0 \rightarrow 0$.

3. Eigenvalues with varying $k$

We also investigate the eigenvalues against $\log k$ for all boundary conditions in Fig. 6, down to $\log k = -7$. We...
could not get eigenvalues for smaller $k$ for reasons of numerical precision.

As shown in Fig. 6(a), eigenvalues $\omega^2$ seem to approach constant values with $k$ decreasing (which corresponds to letting $V_0$ tend to 0 while keeping $V_{\text{max}} = 1$ constant). We also see that the eigenvalues seem to assemble themselves into groups of four, which we referred to earlier as zero modes, first excited and second excited states. The grouping is not so clear for the second excited states, but supporting evidence is shown in Fig. 6(b), where we see an approximate power law dependence for $\omega^2 - \omega_{(n)}^2$ against $k$, where $\omega_{(0)}^2 = 0$ for the zero mode, $\omega_{(1)}^2 = 0.135$ for the first excited state and $\omega_{(2)}^2 = 0.169$ for the second one. Therefore, they should be the degenerate eigenvalues of all the four BCs in the limit of $T_0 \to \infty$ or $k \to 0$.

The only bound states (whose eigenvalues are between 0 and $U_{\text{max}} \approx 1/8$) are those associated with translations and changes in $V_0$, corroborating our analysis of the previous section with the approximate solution $T(z) \approx z$.

4. Eigenvalues at $k = 0$ and correspondence with string ground states

The numerical results displayed in Fig. 5 imply a trend that, towards the end of the tachyon condensation, there are four degenerate zero modes under BC(1, 2, 4) [BC(2) contributes two]. At the end of the decay, a stable solution consisting of an alternating series of $Dp - 1$- and anti-$Dp - 1$-branes form. It is interesting to compare the four zero modes of the tachyon field at $k = 0$ to the bosonic massless states of the open string spectrum.

Instead of a solution consisting of a periodic array of kinks and antikinks separated by distance $\pi/\beta$, we can take the $x$ direction to be compactified on a circle with radius $R = 1/\beta = \sqrt{2}$ for the superstring. The masses of open strings stretched between the parallel branes are

$$m^2 = N + \left(\frac{\Delta \theta R}{2\pi}\right)^2 - a,$$

where $N$ is the level number, $\Delta \theta$ is the angle difference between any two branes and we recall that $a' = 1$. The parameter $a$ is the normal ordering constant which is 0 for the Ramond (R) sector and 1/2 for the Neveu-Schwarz (NS) sector. Between the two branes located on a circle, there are six ways to attach open strings: both ends on (i) D- or (ii) $D$-branes; stretching from $D$ to $\bar{D}$ on both (iii) the right sides and (iv) the left sides of the branes; stretching from $\bar{D}$ to $D$ on both (v) the right sides and (vi) the left sides. The bosonic massless states for the cases (i) and (ii) are the first excited states in the NS sector, which contain the transverse fluctuations of the brane. For (iii), (iv), (v) and (vi), the ground states are massless when $\Delta \theta = \pi$ and $R = \sqrt{2}$. These states become tachyonic if $R < \sqrt{2}$.

Hence we find a total of 6 bosonic massless states. Two are concerned with translations of the branes, two of them should be associated with brane-antibrane annihilation to $Dp - 1$- and Dp - 3-branes and the appearance of a new effective field theory [31], and the other two can be identified with the branes “melting” back into the unstable $Dp$-brane. It is interesting that the original tachyon field, despite being singular, still contains sensible zero modes.

IV. MOTION IN MODULI SPACE

As we have seen, the condensation process is essentially described by any one of the following parameters: $T_0$, $V_0 = V(T_0)$ and $k = V_0/V_m$. In the following, we study the time-dependent decay of the brane into a brane and an antibrane by studying the motion of the $k$ modulus based on the solution of the static equation of motion.

We start from the $1 + 1$ dimensional effective action:

$$S = - \int dt dx V(T)\sqrt{1 - \dot{T}^2 + T'^2}.$$  (4.1)

In the moduli space, we adopt the approximation $\dot{T}^2/(1 + T'^2) \ll 1$, i.e., the tachyon evolves very slowly. Then an
When the second term is small enough, the Langrangian is exactly the energy of the static case in Eq. (2.4). We assume that there is a slow-motion solution with the approximate form of Eq. (2.3), with \( k = V_0/V \) time dependent. Hence

\[
\dot{T} = -\frac{\tanh(\beta T)}{\beta k(1-k^2)} k. \tag{4.3}
\]

Substituting into Eq. (4.2) and performing the integration over \( x \), over a complete period of the tachyon field, we find

\[
S = \int dt L(k, \dot{k})
= -\frac{2\pi V_m}{\beta} \int dt \left[ 1 - \frac{k^2}{2\beta^2 k(1+k)(1-k^2)} \right]. \tag{4.4}
\]

where \( L(k, \dot{k}) \) is the Langrangian with parameters of \( k \) and \( \dot{k} \). The second term is a kinetic term, which can be expressed in terms of a variable \( X(t) \) defined through

\[
\frac{1}{2} X^2 = \frac{1}{2} g(k) k^2, \tag{4.5}
\]

where \( g(k) = 1/[\beta^2 k(1+k)(1-k^2)] \). The solution of \( X \) is trivial: \( X = X_0 + vt \), where \( X_0 \) and \( v \) are constants. Therefore from the relation between \( k \) and \( X \):

\[
X = \int \sqrt{g(k)} dk = 2\sqrt{\frac{\pi V_m}{\beta^3}} \tan^{-1} \sqrt{\frac{2k}{1-k}} \tag{4.6}
\]

and we obtain the solution:

\[
k(t) = \frac{4}{3 + \cos(c_1 t + c_2)} - 1, \tag{4.7}
\]

where \( c_1 = \sqrt{\beta^3/\pi V_m} \) and \( c_0 \) are constants. We discover that the \( Dp \)-branes decay into \( Dp-1 \)-branes in a finite time \( \Delta t = \sqrt{V_m \pi/\beta^3}/v \). This result is consistent with singularities appearing in finite time in time-dependent tachyon simulations [23,24] and in boundary conformal field theory calculations [32]. We note that the solution is in fact periodic, and oscillates between the unstable \( Dp \)-brane system and the daughter \( Dp-1, \bar{D}p-1 \)-brane system.
[1] M. R. Garousi, Nucl. Phys. B584, 284 (2000).
[2] E. A. Bergshoeff, M. de Roo, T. C. de Wit, E. Eyras, and S. Panda, J. High Energy Phys. 05 (2000) 009.
[3] A. Sen, Mod. Phys. Lett. A 17, 1797 (2002).
[4] A. Mazumdar, S. Panda, and A. Perez-Lorenzana, Nucl. Phys. B614, 101 (2001).
[5] G. W. Gibbons, Phys. Lett. B 537, 1 (2002).
[6] M. Fairbairn and M. H. G. Tytgat, Phys. Lett. B 546, 1 (2002).
[7] D. Choudhury, D. Ghoshal, D. P. Jatkar, and S. Panda, Phys. Lett. B 544, 231 (2002).
[8] L. Kofman and A. Linde, J. High Energy Phys. 07 (2002) 004.
[9] M. Sami, P. Chingangbam, and T. Qureshi, Phys. Rev. D 66, 043530 (2002).
[10] G. Shiu and I. Wasserman, Phys. Lett. B 541, 6 (2002).
[11] P. Chingangbam, S. Panda, and A. Deshamukhya, J. High Energy Phys. 02 (2005) 052.
[12] A. Sen, J. High Energy Phys. 09 (1998) 023.
[13] A. Sen, J. High Energy Phys. 12 (1998) 021.
[14] D. Kutasov and V. Niarchos, Nucl. Phys. B666, 56 (2003).
[15] N. Lambert, H. Liu, and J. M. Maldacena, J. High Energy Phys. 03 (2007) 014.
[16] G. R. Dvali and S. H. H. Tye, Phys. Lett. B 450, 72 (1999).
[17] S. Sarangi and S. H. H. Tye, Phys. Lett. B 536, 185 (2002).
[18] D. Choudhury, D. Ghoshal, D. P. Jatkar, and S. Panda, J. Cosmol. Astropart. Phys. 07 (2003) 009.
[19] K. Hashimoto and N. Sakai, J. High Energy Phys. 12 (2002) 064.
[20] G. Shiu, S. H. H. Tye, and I. Wasserman, Phys. Rev. D 67, 083517 (2003).
[21] J. M. Cline, H. Firouzjahi, and P. Martineau, J. High Energy Phys. 11 (2002) 041.
[22] A. Sen, Phys. Rev. D 68, 066008 (2003).
[23] N. Barnaby, A. Berndsen, J. M. Cline, and H. Stoica, J. High Energy Phys. 06 (2005) 075.
[24] J. M. Cline and H. Firouzjahi, Phys. Lett. B 564, 255 (2003).
[25] G. N. Felder, L. Kofman, and A. Starobinsky, J. High Energy Phys. 09 (2002) 026.
[26] C.-j. Kim, Y.-b. Kim, and C. O. Lee, J. High Energy Phys. 05 (2003) 020.
[27] A. Sen, Int. J. Mod. Phys. A 20, 5513 (2005).
[28] P. Brax, J. Mourad, and D. A. Steer, Phys. Lett. B 575, 115 (2003).
[29] L. D. Landau and E. M. Lifshitz, Quantum Mechanics (Pergamon, Oxford, 1977).
[30] I. S. Gradsteyn and I. M. Ryzhik, Table of Integrals, Series, and Products (Academic, New York, 2000).
[31] M. R. Garousi, J. High Energy Phys. 01 (2005) 029.
[32] A. Sen, J. High Energy Phys. 10 (2002) 003.