Large atom number Bose-Einstein Condensate machines

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(Dated: March 23, 2022)

We describe experimental setups for producing large Bose-Einstein condensates of $^{23}$Na and $^{87}$Rb. In both, a high flux thermal atomic beam is decelerated by a Zeeman slower and is captured and cooled in a magneto-optical trap. The atoms are then transferred into a cloverleaf style Ioffe-Pritchard magnetic trap and cooled to quantum degeneracy with radio frequency induced forced evaporation. Typical condensates contain 20 million atoms. The design includes a second vacuum chamber into which ultracold atoms can be transported with an optical tweezers. This allows the flexibility to rapidly prepare and perform a wide variety of experiments.

PACS numbers: 03.75.Hh, 39.10.+j, 39.25.+k, 07.30.-t

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ness in day-to-day operation. Since 1996, the MIT sodium BEC setups have featured the largest alkali condensates. Our three setups routinely produce condensates with atom numbers between 20 and 100 million. Since diode lasers for rubidium are less expensive than the dye lasers for sodium, most new groups have chosen to work in rubidium. The most popular laser cooling setups for rubidium involve vapor cell MOTs, which do not obtain the condensate size of the MIT sodium experiments. There has been a widespread perception in the field that Zeeman slowing is the technique of choice for sodium and vapor cell traps for rubidium. The construction of vapor cell MOT rubidium condensate machines is extensively detailed in the complementary work of Ref. [15].

When the Center of Ultracold Atoms was created at MIT and Harvard, one major funded project was the translation of the techniques we had developed for sodium to rubidium and to create a rubidium BEC experiment with enormous condensates. The successful accomplishment of these goal is described in this paper. Since we are in the unique situation to have two similar experimental setups for Rb and Na, we are able to discuss similarities and differences between the optimization for the two species. Our conclusion is that the technique of using an atomic beam and Zeeman slower works as well for rubidium as for sodium, and we present the technical details of how to build an intense slow beam for both atomic species. Other figures-of-merit besides atom number include simplicity and reliability. In our experience, the Zeeman slowing technique is by far the simplest technique to generate an intense slow beam since it requires only a single laser beam of modest power. The length of the slower and therefore the overall size of the vacuum setup may look intimidating, but once it is built, it provides simple and reliable day-to-day operation without need for realignment.

The third-generation sodium experiment and the rubidium experiment were both designed with an additional vacuum chamber (“science chamber”) into which the BEC or evaporatively cooled atoms can be moved using optical tweezers. The multi-chamber design allows us to rapidly reconfigure the experimental setup in the science chambers while keeping the BEC production chambers under vacuum. This has allowed us to perform very different experiments in rapid succession [16, 17, 18, 19, 20, 21, 22, 23, 24, 25, 26, 27].

II. SYSTEM OVERVIEW

![Vacuum system diagram and major subsystems. $^{87}$Rb apparatus shown.]

Fig. 1 illustrates the layout of our system. A thermal atomic beam emanates from the oven and is decelerated with the Zeeman slower. In the main chamber, the slowed atoms are captured and cooled with a six-beam magneto-optical trap (MOT) [28]. To load the magnetic trap, the atoms are optically pumped into the $F=1$ hyperfine ground state. Atoms in the $F=1, m_F=-1$ state are held by their attraction to the field minimum of the Ioffe-Pritchard magnetic trap.

The trapped sample is evaporatively cooled by removing hotter atoms through radio frequency (RF) induced transitions into untrapped states. Reducing the RF frequency lowers the effective depth of the magnetic trap, allowing us to progressively cool to higher densities and lower temperatures until the atoms reach BEC. Magnetically trapped atoms in the $F=2, m_F=+2$ state have also been evaporated down to BEC.

Ultracold atoms can be transported from the main chamber into an adjoining auxiliary “science chamber” by loading the atoms into the focus of an optical tweezer and then translating the focus. In this manner we have transported $^{23}$Na BECs [16]. Vibrational issues during transport cited in [16] were

| Stage | $n$ ($/\text{cm}^3$) | Temperature | $\rho$ |
|-------|-----------------|-------------|-------|
| Oven  | $10^{13}$       | 383 K       | $10^{-14}$ |
| Thermal Beam | $10^7$ | $v_{mb} = 334 \text{ m/s}$ | $10^{-20}$ |
| Slowed Beam | $10^7$ | $v_{mb} = 43 \text{ m/s}$ | $10^{-18}$ |
| Loading MOT$^a$ | $10^9$ | 150µK | $10^{-8}$ |
| Compressed MOT$^b$ | $10^{10}$ | 300µK | $4 \times 10^{-8}$ |
| Molasses$^c$ | $10^{10}$ | 10µK | $6 \times 10^{-6}$ |
| Magnetic trap | $10^{11}$ | 500µK | $2 \times 10^{-7}$ |
| BEC Transition | $3 \times 10^{13}$ | 500 nK | 2.61 |
| Pure BEC | $10^{14}$ | (250 nK$^b$) | (100) |

TABLE I: Typical phase space densities ($\rho$) during BEC production in the $^{87}$Rb apparatus.

$^a$Typical values, not measured separately
$^b$Chemical potential

![Experimental setup in the science chambers while keeping the...](image)
reduced by the use of Aerotech ABL2000 series air bearing translation stages.

Technical issues related to the greater mass and higher three body recombination rate in \(^{87}\)Rb were overcome by transporting ultracold atoms just above the transition temperature \(T_r\), and then evaporating to BEC at the destination. The oven and Zeeman slower are tilted by 57° from horizontal to allow a horizontal orientation for the weak trapping axes of both the optical tweezers and magnetic trap.

Trapping ultracold atoms requires that they be isolated from the surrounding environment. The laser and magnetic trapping techniques confine the atoms in the center of the chamber, out of contact with the room temperature chamber walls. The atoms are still exposed to thermal black body radiation from the chamber walls, but are transparent to most of the spectrum. The transitions which the black body radiation can couple to are the optical transitions used for laser cooling and the microwave hyperfine transitions. For optical transitions, which have energies much greater then \(k_B T\) the excitation rate is \(\frac{4}{\tau_{hfs}} \frac{k_B T}{\hbar \omega_{optical}}\), where \(\omega_{optical}\) is the frequency of the transition and \(\tau_{hfs}\) the lifetime of the excited state. For rubidium in a 25°C chamber this gives a characteristic excitation lifetime of ~56 billion years. Raising the chamber temperature to 680°C increases the optical excitation rate into the experimentally relevant domain of once per minute. The hyperfine transitions are significantly lower in energy compared to \(k_B T\) and have an excitation rate of \(\frac{4}{\tau_{hfs}} \frac{k_B T}{\hbar \omega_{hfs}}\), which is once per year at 25°C in \(^{87}\)Rb since the ground state hyperfine spontaneous decay lifetime \(\tau_{hfs}\) in an alkali atom is thousands of years. Neither of these excitation rates are limitations on current experiments.

Collisions with background gas molecules result in loss from the trap, necessitating low vacuum pressure for long atom cloud lifetime. We can magnetically trap ultracold atoms with lifetimes of several minutes in the < 10\(^{-11}\) torr ultrahigh vacuum (UHV) environment of the main production chamber. To achieve this vacuum performance we have followed the general guidelines set out in Ref. [29] for constructing vacuum systems. The main chamber body was constructed of nonmagnetic 304 stainless steel and then electropolished to reduce the surface roughness. The only component placed inside the chamber was the RF evaporation antenna coil (Fig. 2).

The cloverleaf-style Ioffe-Pritchard magnetic trap coils fit inside two re-entrant bucket windows, allowing them to be outside the chamber with an inter coil spacing of 25 mm (Fig. 2). The Zeeman slower tube is mounted between the main chamber and the oven chamber. The Zeeman slower coils that are around the Zeeman slower tube are also outside of the vacuum system but cannot be removed without breaking vacuum.

After assembling the chamber, we pumped out the system and reached UHV conditions by heating the system to accelerate outgassing. We heated the main chamber to 230°C and the Zeeman slower to 170°C (limited by the coil epoxy). Using a residual gas analyzer to monitor the main chamber, we "baked" until the partial pressure of hydrogen was reduced to less than 10\(^{-7}\) torr and was at least ten times greater than the partial pressure of other gases. A typical bakeout lasted between 3 and 9 days, with temperature changes limited to less than 50°C/hour. While we acknowledge the merit of using dry pumps as recommended in Ref. [15], we have not had any detrimental experiences using oil sealed rotary vane roughing pumps to back our turbo pumps. The vacuum in the main chamber is preserved after bakeout with a 75 L/s ion pump and a titanium sublimation pump. Refer to Sec. 3.4 of Ref. [30] for more details of our bakeout procedures.

### III. OVEN

We generate large fluxes of thermal atoms for Zeeman slowing from effusive atomic beam ovens. An effusive beam is created by atoms escaping through a small hole in a heated chamber [31]. The higher vapor pressure of rubidium requires a more complicated design, but lower operating temperature (110-150°C Rb, 260°C Na.) At room temperature, the vapor pressure of sodium (\(\approx 2 \times 10^{-11}\) torr [32]) is roughly compatible with our UHV main chamber environment, while that of rubidium (\(\approx 4 \times 10^{-7}\) torr [33]) is not. This dictated that the design of the rubidium oven prevent contaminating the main chamber with rubidium. Because of its greater complexity, further discussion will focus on the rubidium oven (Fig. 3). We expect that the rubidium oven would work as well for sodium, but instead we used a simpler design described in Ref. [30].

A combination of active pumping and passive geometrical techniques were used to reduce extraneous rubidium transfer to the main chamber. A cold cup (I) is used to reduce rubidium vapor in the oven chamber by almost completely surrounding the oven aperture (J) with a cold surface. After bakeout, the combination of cold cup and oven chamber ion pump has achieved pressures as low as \(\approx 10^{-9}\) torr, although we have successfully made BECs with pressures of up to \(\approx 10^{-6}\) torr in this region. A combination of a differential pumping tube, ion pump, and the Zeeman slower tube provides a pressure differential of over 3 orders of magnitude between the oven and main chamber. This is sufficient to isolate the UHV envi-
The oven is loaded with a sealed glass ampoule containing 5 g of rubidium in an argon atmosphere. To add rubidium, the ampoule is cleaned, placed in the oven, and baked out under vacuum while still sealed. We then break the ampoule under vacuum and heat the oven to 110°C to produce the atomic beam. During operation, the machine is run as a sealed system, without the turbo-mechanical pump, to prevent accidental loss of the main chamber vacuum. Oven temperatures from 150°C down to 110°C produced similar sized ⁸⁷Rb BECs. Reducing the oven temperature increased the time between rubidium changes to greater than 1000 hrs of operating time. This long operating cycle precluded the need for more complex recycling oven designs [33].

IV. ZEEMAN SLOWER

The atomic beams are slowed from thermal velocities by nearly an order of magnitude by scattering photons from a resonant, counter-propagating laser beam. When a photon with momentum \( \hbar k = 2\pi/\lambda \) is absorbed or emitted by an atom with mass \( m \), the atom will recoil with a velocity change of

\[
\Delta v = \frac{\bar{h}k}{m} \text{ to conserve momentum. Atoms can resonantly scatter photons up to a maximum rate of } \frac{1}{\Gamma} = 2, \text{ where } 1/\Gamma = \tau \text{ is the excited-state lifetime. This results in a maximum acceleration } \dot{a}_{\text{max}} = \frac{\hbar k^2}{2\bar{h}} (1.1 \times 10^5 \text{m/s}^2 \text{ Rb, } 9.3 \times 10^5 \text{m/s}^2 \text{ Na} ). \text{ As the atoms decelerate, the reduced Doppler shift is compensated by tuning of the Zeeman shift with a magnetic field } [33] \text{ to keep the optical transition on resonance. We designed our slower to decelerate the atoms at a reduced rate } \dot{a}_{\text{max}} \text{ where } f \sim 50\% \text{ is a safety factor to allow for magnetic field imperfections and finite slower laser intensity.}

Our slowers are designed along the lines of Ref. [33], with an increasing magnetic field and \( \sigma^- \) polarized light scattering off the \( F=2, m_F=-2 \rightarrow F'=3, m_{F'}=-3 \) cycling transition. Before the slowing begins, the atoms are optically pumped into the \( F=2, m_F=-2 \) state. The large magnetic field at the end of the slower corresponds to a large detuning from the low velocity, low magnetic field resonance frequency. This large detuning allows the slowing light to pass through the MOT with-
out distorting it due to radiation pressure. Within the slower, the quantization axis is well-defined by the longitudinal magnetic field and optical pumping out of the cycling transition is strongly suppressed by the combination of light polarization and Zeeman splitting.

We slow $^{87}$Rb atoms from an initial velocity of $\sim 350$ m/s with a tailored 271 G change in magnetic field (Fig. 4). An additional uniform $\sim 200$ G bias field was applied along the length of the slower to ensure that neighboring hyperfine levels were not near resonance in either the slower or the MOT. The slower cycling transition light is detuned -687 MHz from the $F=2 \rightarrow F'=3$ transition. The slowing laser intensity is $I/I_{sat} \approx 8$, giving a maximum theoretical deceleration of 89% of $a_{max}$. “Slower repumping” light copropagates with the cycling transition light and is detuned -420 MHz from the $F=1 \rightarrow F'=1$ transition to match the Doppler shift of the unslowed thermal atoms from the oven. A flux of $\sim 10^{11}$ $^{87}$Rb atoms/s with a peak velocity of 43 m/s was measured from our slower with an oven temperature of 150°C. This is significantly greater flux then the $8 \times 10^8$ Rb/sec vapor cell loading rate quoted by [15].

The higher temperature of the sodium oven, along with the atoms’ lower mass, results in a greater initial velocity of 950 m/s. This requires a slower with a much larger magnetic field change of 1150 Gauss. To reduce the maximum magnitude of the magnetic fields we use the spin flip variant of the increasing field design by shifting the zero crossing of the magnetic field from the beginning of the slower to the middle. The first segment then becomes a decreasing field slower, with current flowing in the opposite direction of the second, increasing field segment. In the low magnetic field region near the zero crossing, the atoms are theoretically vulnerable to optical pumping out of the cycling transition, but we did not find this to be a problem experimentally. Ref. [30] has demonstrated a high flux spin flip style slower for $^{87}$Rb. The sodium slowing beam is detuned -1.0 GHz from the $F=2 \rightarrow F'=3$ transition and has an intensity of $I/I_{sat} \approx 4$, giving a laser power limited maximum deceleration of 80% of $a_{max}$. Unlike the rubidium slower, light for optical pumping was generated by adding 1.75 GHz sidebands to the slowing light using an electro-optical modulator.

The sodium slower coils were broken up such that the first segment had an initial field of 440 G and a length of 52 cm and the second segment had a final field of 710 G and a length of 43 cm. The sodium slower was tested as depicted in Fig. 5 with a measured flux of $3 \times 10^{11}$ $^{23}$Na atoms/s with a peak velocity of 100 m/s.
Relevant transitions, hyperfine splittings, and laser frequencies.

**FIG. 6:** Simplified level structure of $^{87}$Rb (A) and $^{23}$Na (B) with relevant transitions, hyperfine splittings, and laser frequencies.

V. LASERS

Resonant laser light is used to slow, cool, trap, and detect the atoms. All laser light is prepared on a separate optics table and delivered to the apparatus (Fig. 1) through single-mode optical fibers. Because stray resonant light can heat the atoms during evaporation, black cloth separates the two tables. All frequency shifting and attenuation of the light is done with acousto-optic modulators. Mechanical shutters are also placed in front of each fiber coupler to block any light which might leak through the modulators and disturb the atoms. Atomic energy levels and laser frequencies used are indicated in Fig. 6.

We use different techniques for generating laser light at the resonant wavelengths of $^{87}$Rb (780 nm) and $^{23}$Na (589 nm). Commercially available (Toptica DL100, TA100) external cavity diode lasers and semiconductor tapered amplifiers are used to create 350 mW and 35 mW of light resonant with the $^{87}$Rb $F=2 \rightarrow F'=3$ and $F=1 \rightarrow F'=1$ transitions at 780 nm respectively. The lasers are stabilized with a polarization sensitivity saturated absorption spectroscopy lock [37, 38]. This modulation-free technique optically creates a derivative signal of the absorption spectra that is locked with a proportional+integral gain servo loop. The locking signal fluctuation indicates a frequency jitter of $< 1$MHz over several seconds, which is much less than the 6.1 MHz natural linewidth of $^{87}$Rb.

The $^{87}$Rb MOT uses a total of 60 mW of light near the $F=2 \rightarrow F'=3$ cycling transition for trapping/cooling. The $F=2 \rightarrow F'=3$ transition in the MOT is only approximately a closed cycle and atoms are often optically pumped into the $F=1$ ground state. To “repump” these atoms back into the $F=2$ state we use 10 mW of light on the $F=1 \rightarrow F'=1$ transition. Likewise, to deliberately transfer atoms from the $F=2$ to $F=1$ manifold we can introduce a few mW of “depumping” light resonant with the $F=2 \rightarrow F'=2$ transition. Powers are quoted after fiber coupling, measured as delivered to the apparatus table. After frequency shifting the slower cycling and repumping light to their desired detunings only a few mW of power are available. Each of these beams is then amplified to 35-40 mW by injection locking [39] a free running Sanyo DL7140-201 laser diode. The two amplified beams are then overlapped and coupled into a fiber, which delivers 18 mW of slower cycling light and 6 mW of slower repumping light.

For $^{23}$Na we use a Coherent 899 dye laser pumped by a Spectra Physics Millenia laser (532 nm, 8.5 W). Typically 1.2 W of 589 nm light is generated by the dye laser. The laser frequency was referenced to an external saturation-absorption lock-in scheme and locked to a Fabry-Perot cavity. Stable operation was improved by using a precision dye nozzle (Radiant Dyes, Germany), high pressure dye circulator at 12 bars, and stabilized temperatures for the room and dye.

For more detailed information on the generation of the laser light for sodium MOT’s, see Sec. 3.4 of Ref. [40]. Typical delivered laser powers are 80 mW for the MOT light, 20 mW for the repumping light, 40 mW for the slowing light and less than one mW for the imaging beam. Electro-optic modulators allow the addition of high frequency sidebands ($\sim 1.8$ GHz) on the slowing and MOT light for repumping without the use of an additional laser beam. Recent advances in single frequency high power fiber and diode pumped solid state lasers [60] have made nonlinear techniques such as sum frequency generation [41, 42] and frequency doubling [43] interesting alternatives as resonant light sources.

VI. MAGNETO OPTICAL TRAP

The MOT [28] is the workhorse of atomic physics for creating large samples of ultracold atoms. We use a six-beam MOT, which doubles as an optical molasses when the magnetic gradient field is off. Similar to Ref. [15] the $^{87}$Rb apparatus uses a bright MOT. The $^{87}$Rb MOT equilibrates to around $4 \times 10^{11}$ atoms after ~ 2 s of loading, operating in a magnetic field gradient of 16.5 G/cm with cycling beams detuned -18 MHz from the $F=2 \rightarrow F'=3$ transition and a peak intensity 5.3 mW/cm$^2$. To increase the efficiency of the transfer into the magnetic trap, we briefly compress the $^{87}$Rb MOT and then switch off the magnetic field gradient to cool the atoms with optical molasses. The $^{87}$Rb MOT is compressed by linearly ramping the gradient to 71 G/cm in 200 ms and simultaneously sweeping the detuning to -45 MHz in 400 ms. We use 5 ms of “gray” molasses, where the repumper power is dropped by 95%, the optical trapping power is ramped down to 50%, and the detuning is swept from -18 MHz to -26 MHz. The molasses phase requires reduction of imbalances in intensity between beams and also to residual magnetic fields [44].
After the molasses phase, 0.5-1 ms of “depumping” light is applied to put all the $^{87}\text{Rb}$ atoms into the F=1 level before loading into the magnetic trap. Exact MOT and molasses parameters were found through empirical optimization, and all listed numbers should be considered as guides.

The $^{23}\text{Na}$ apparatus uses a dark-spot MOT [45], with a detuning of -15MHz, peak beam intensity of 8.8 mW/cm$^2$ and a magnetic field gradient of 11G/cm. A 4 mm diameter opaque circle blocks light in the middle of a single repumper beam, creating a region at the center of the MOT where trapped atoms are optically pumped into the F=1 state. The $^{23}\text{Na}$ MOT equilibrates after a few seconds of loading. The effectiveness of the dark-spot in $^{23}\text{Na}$ has precluded the need for the compression and molasses phases as in $^{87}\text{Rb}$. 99% of trapped atoms reside inside the dark spot, and the $^{23}\text{Na}$ atoms are sufficiently cold and dense to be directly loaded into the magnetic trap.

VII. MAGNETIC TRAP

Atoms in weak magnetic field seeking states can be trapped in a magnetic field minimum. Our magnetic trap is a high current Ioffe-Pritchard (IP) trap with a cloverleaf style winding that can hold F=1, $m_F= -1$ or F=2, $m_F= +2$ ground state atoms of $^{87}\text{Rb}$ and $^{23}\text{Na}$ with long lifetimes. An IP trap has an anisotropic, “cigar”-shaped, 3D harmonic trap for energies which are small compared to the trap minimum $g_F m_F \mu_B B_0$ and a 2D linear/1D harmonic trap at higher energies (See Fig. 7 and Appendix C). This linear regime at higher energies (higher cloud temperatures) is more efficient for evaporatively cooling hot atoms [46], while the finite bias field at the minimum prevents Majorana spin flip loss of colder atoms.

![Fig. 7: Profile of the Ioffe-Pritchard trap magnetic field magnitude. The trap parameters are $B'=223 \text{ G/cm}$, $B''=100 \text{ G/cm}^2$, and $B_0=1 \text{ G}$.](image)

When assembled the anti bias coils enclose the cloverleaf coils and the MOT coils surround the curvature coils.

When the magnetic trap is initially turned on, the strength and shape of the confinement are adjusted to match that of the laser cooled atoms to preserve phase space density. Additional current is applied to the curvature coils, increasing the residual bias field and decreasing the radial confinement to make a roughly spherical magnetic trap that match the spherical MOT. After loading atoms in the trap, the additional curvature coil current is reduced over one second to adiabatically change the trap geometry to the tightly confining cigar shape, favorable for evaporative cooling. Sec. 2.3.2 of Ref. [47] has an extensive discussion of mode matching magnetic traps to MOTs. The adiabatic compression technique is reviewed in Ref. [46].

VIII. CONTROL AND IMAGING

Two computers run the apparatus; one controls the various parts of experiment and the other processes images from a camera which images the atoms. The control computer has custom built National Instruments (NI) LabWindows based software to drive analog (2 NI Model PCI 6713, 8 channels of 12 bit analog, 1MS/s update) and digital output (2 NI Model PCI-6533, 32 channels of binary TTL, 13.3 MS/s update) boards. The control computer also controls an Agilent 33250A 80MHz function generator through a GPIB interface, and triggers a Princeton Instruments NTE/CCD-1024-ED camera through a ST-133 controller to capture the absorption images.

BECs are typically imaged 10-40 ms after release from the trap. Ref. [47, 48] provides the details of analyzing condensates after free expansion. Atoms are first optically pumped into the F=2 state in 200[μs] and then an absorption image is taken using on resonance F=2 $\rightarrow$ F′=3 light. Detuning off res-
onance causes dispersion (lensing) as the light passes through the cloud of atoms and can distort the image. The intensity of the imaging probe is kept lower than the saturation intensity to prevent bleaching of the transition, which would lead to errors in atom number counting. Typical exposure times are 50-200 µs. Sec. 3 of Ref. [47] discusses other imaging techniques that can also be used to probe BECs.

IX. EVAPORATION

Evaporative cooling works by selectively removing hot atoms from the trapped cloud, while the remaining atoms rethermalize to a lower temperature. The efficiency of cooling depends on $\eta$, the ratio of trap depth or energy of the escaping atoms to the temperature $k_B T$, and is reduced by the rate of heating. The speed of this process depends on how quickly the atoms rethermalize. In a magnetic trap evaporation is implemented through RF induced transitions between trapped and untrapped states. A given RF frequency corresponds to a shell of constant $\mu_m |B|$ where the transitions occur. Atoms that pass through this shell enter untrapped states and are lost; thus RF provides a flexible mechanism to control the magnetic trap depth. Our RF antenna consists of two rectangular loops of wire, 10 cm x 2 cm, positioned 3 cm above and below the condensate as depicted in Fig. 2.

![Fig. 9: Typical temperature and trap depth during evaporation to BEC in $^{87}$Rb.](image)

To evaporate thermal atoms to a BEC, we sweep the RF frequency over several seconds using an Agilent 33250A synthesizer amplified with a 5 W RF amplifier (Mini-Circuits ZHL-5W-1). Typical evaporation curves for $^{87}$Rb would ramp from 60 MHz down to ~0.8 MHz in 15 to 40 seconds. Forced RF evaporative cooling is very efficient, increasing phase space density by $> 10^6$ (Table I). Fig. 2 shows the drop in temperature as the trap depth (calculated from the RF frequency) is lowered during evaporation of $^{87}$Rb. Evaporation curves are frequently adjusted in the interest of tuning evaporation speed, atom number, density, and/or reproducibility. For instance, the atom number can be increased by decompressing the magnetic trap near the end of the evaporation. This reduces the effects of three body recombination heating by lowering the final condensate density. Such decompression techniques have allowed us to create nearly pure condensates with $N_e \approx 20 \times 10^6$ in both $^{87}$Rb and $^{23}$Na with lifetimes in excess of 5 seconds.

Decompressing the trap shifts its center due to gravitational sag and imperfections in the balance of magnetic fields between the coils. Such movements can excite oscillations in the cloud, which results in the condensation of BECs which are not rest. Even in the absence of excitations, the magnetic field gradients must exert a force on the atoms which is greater then gravity for them to remain trapped. This limits the extent to which magnetic traps can be decompressed. Specially designed gravito-magnetic traps have been decompressed down to 1 Hz [19] to investigate very cold, dilute BECs.

X. DEEP TRAP LIMITATIONS

A major difference we have observed between $^{87}$Rb and $^{23}$Na condensates is the unexpectedly high decay rate of $^{87}$Rb in tightly confining deep traps, such as those used for transport in an optical trap [16]. At the typical densities of condensates, the lifetime and heating are usually determined by three-body recombination decay. However, the factor of four difference in the three-body rate coefficients (Table II) was insufficient to explain this major discrepancy in behavior.

We investigated this issue in a magnetic trap instead of an optical trap. While it is easier experimentally to create tight trapping and hence high densities in an optical trap, both the trap frequencies and trap depth are functions of the optical power. This makes it difficult to separate density dependent effects, which are strongly affected by the trap frequency, from trap depth effects in an optical trap. In contrast, in a magnetic trap the trap depth can be controlled independently of the trap frequencies by adjusting the RF frequency which flips atoms to untrapped states.

There are two possible processes, both involving secondary collisions, which can greatly enhance the heating and losses due to the primary three-body collisions.

The first process is collisional avalanches, similar to a chain reaction, where the energetic products of three-body recombination collide with further atoms while they leave the condensate. This process would depend on the collisional opacity $\sim n a t$ (where $a$ is the atom atom scattering cross section) and would increase dramatically when it exceeds the critical opacity of 0.693 [50]. This process is independent of trap depth.

The second possible process can already happen at lower collisional opacities and relies on the retention of primary or secondary collision products by the trap in the so-called Oort cloud [47, 51]. Subsequently, when those atoms slosh back into the trapped condensate, they can cause heating and trap loss. The retention of collision products in the Oort cloud should depend on whether the trap depth is larger or smaller than their energies.

Fig. 10 shows the initial loss rates measured for a large and
FIG. 10: Initial loss rates for $^{87}\text{Rb}$ BEC in deep traps. Trap depth dependence of the loss for large and small $^{87}\text{Rb}$ BECs in a 220 Hz x 220 Hz x 9 Hz magnetic trap. The trap depth was controlled by RF truncation. Condensates were nearly pure ($N_c/N > 90\%$) and consisted of $F=1$, $m_F=-1$ atoms. Solid triangles are data for a large condensate, $N_c = 2.7 \times 10^{14}$ atoms, peak density $n_p = 6.1 \times 10^{14}$ cm$^{-3}$, expected three body decay time $\tau_3 = 0.85 \pm 0.22$ s (dashed horizontal lines) [49], and calculated collisional opacity of 0.88 [50]. Open squares are data for a small condensate, $N_c = 5 \times 10^9$ atoms, peak density $n_p = 3.1 \times 10^{14}$ cm$^{-3}$, expected three body decay time $\tau_3 = 3.3 \pm 0.8$ s (dotted horizontal lines), and a calculated collisional opacity of 0.32. The error bars represent the statistical uncertainty in the decay curves. Additional scatter in the data is due to fluctuations in the atom number.

a small BEC as a function of the magnetic trap depth. At low trap depths (5 $\mu$K) both the large and small condensate decay rates are in agreement with established three body recombination rates [49]. Therefore, the avalanche effect does not significantly contribute to the observed decay rate, although the calculated opacity for the larger condensate was 0.88 and may not be far away from the onset of avalanches. Evidence for avalanches was obtained at an opacity of 1.4 in [50].

In the larger condensate at higher trap depth, the decay rate strongly increases, supporting the second process involving the Oort cloud as the likely mechanism. For times longer than 500 ms, the large $^{87}\text{Rb}$ condensate in the deeper traps was heated away and only a few thermal atoms remained. In contrast, at low trap depths the large condensate was still fully condensed after 20 seconds, and had an atom number in agreement with the expected losses from three body decay.

We speculate that this enhancement of the three-body losses was not observed in $^{23}\text{Na}$ for the following reasons. Three body recombination results in a diatomic molecule and an atom which fly apart with a total kinetic energy (2/3 to the atom, 1/3 to the molecule to conserve momentum) equal to the binding energy of the diatomic molecule in the highest vibrational state. This binding energy can be estimated from the scattering length as $E_0 \sim \hbar^2/ma^2$ [52] ($\sim 200\mu$K in $^{87}\text{Rb}$, $\sim 2.7$ mK in $^{23}\text{Na}$). The direct decay products will only be retained if the trap depth is greater than their kinetic energies (min $\sim 70\mu$K for $^{87}\text{Rb}$, $\sim 900\mu$K for $^{23}\text{Na}$).

In addition, $^{23}\text{Na}$ decay products are less likely to undergo secondary collision processes from either primary three body products or hot Oort cloud atoms due to an elastic scattering cross section $\sigma$ which is 3.6 times smaller than for $^{87}\text{Rb}$. The individual products of secondary collisions can have a spectrum of energies, further lowering the energy threshold for their retention by the Oort cloud. The combination of these three factors, three body rate, scattering cross section, and binding energy result in an estimated increase in the loss rate for $^{87}\text{Rb}$ condensates over $^{23}\text{Na}$ by a factor of $\sim 200$.

The optical trap depths needed for transporting condensates in our system are significant fraction of the primary $^{87}\text{Rb}$ decay product energy, but a small fraction of that for $^{23}\text{Na}$. Therefore $^{23}\text{Na}$ condensates can be easily transported using optical tweezers. For $^{87}\text{Rb}$ the preferred method is to transport a cloud at temperatures just above condensation, where the density is lower, and evaporate to BEC after transport.

### XI. DISCUSSION

|                      | $^{87}\text{Rb}$ | $^{23}\text{Na}$ |
|----------------------|------------------|------------------|
| $D_2$ line $\lambda$ (nm) | 780              | 589              |
| $D_2$ linewidth $\Gamma/2\pi$ (MHz) | 6.1              | 9.8              |
| Gravity $mg/k_B$ (nK/$\mu$m) | 102             | 27               |
| Gravity $mg/g_F m_F u_B$ (G/cm) | 30              | 8.1              |
| Three body constant $K_3$ (cm$^3$/s) | $8 \times 10^{-30}$ [49] | $2 \times 10^{-30}$ [53] |
| Scattering length $a$ (nm) | 5.3 [54]         | 2.8 [55]         |
| Recoil velocity $v_r$ (mm/s) | 5.9             | 29               |

$^{87}\text{Rb}$ and $^{23}\text{Na}$ are the two most popular species for BEC research. We have constructed two machines with nearly identical designs and can discuss differences in performance and operation. Key properties of the two atoms are highlighted in Table II. The four principal differences are in vapor pressure, resonant wavelength, recoil velocity, and collisional properties.

The high vapor pressure of rubidium allows the operation of the oven at lower temperatures, but requires a more elaborate design of cold plates to avoid deposition of rubidium on surfaces of the UHV chamber which are at room temperature. An optimized Zeeman slower for rubidium will be about twice as long than that for sodium at similar oven temperatures, the stopping length $L$ for the most probable velocity in a beam of temperature $T$ being $L = \frac{3}{2} \frac{a}{K_B} T$, assuming the maximum spontaneous light force. In our systems the gain from the greater light force is balanced out by the higher operating temperatures required of the sodium oven to produce comparable flux, resulting in both the rubidium and sodium slower being about 1 m in length.

Due to the higher recoil velocity, the slow sodium beam has a larger divergence than the rubidium beam. By keeping the distance between the end of the slower and the MOT to a minimum, we estimate quantitative transfer of atoms from the
slower to the MOT. Our setup for rubidium was almost identical, but we expect that the requirement of keeping the slower and the MOT so close could be more relaxed for rubidium. Although we have not tried it, we expect that our rubidium experiment would work for sodium without changes to the oven, vacuum or magnet designs.

On the laser side, a major difference is the availability of low cost high power laser diodes in the near infrared region around 780 nm. In our experience a well run dye laser system provides similar or even superior performance to a diode laser system with several master and slave lasers. Our dye lasers at 589 nm tend to be more stable than semiconductor lasers during a long run of the experiment and need less day-to-day tweaking. However, occasionally they require major maintenance in terms of dye changes or full optical realignment. Another advantage is the visibility of the laser light and the atomic fluorescence. The near infrared 780 nm light is only modestly visible, whereas the sodium line at 589 nm is near the peak of human eye sensitivity and allows fine alignments of the laser beams and the magneto-optical trap without cameras, IR cards, or IR viewers.

$^{87}$Rb has favorable properties for laser cooling and atom interferometry because of its greater mass, lower recoil velocity, and larger excited state hyperfine structure. While greater mass and longer resonant wavelength give $^{133}$Cs an even lower recoil velocity, its complicated collisional behavior at low magnetic fields makes it difficult to cool to BEC. The lowest molasses temperature in rubidium is a factor of ten lower than for sodium. However, in BEC experiments the laser cooling is optimized to for large atom numbers and high initial elastic collision rates in the magnetic trap, and not for the lowest temperature. For laser cooling sodium at high atom numbers, the Dark SPOT technique is crucial to avoid rescattering of light, whereas it is only used in some rubidium experiments. One possible reason is that the larger excited state hyperfine structure allows for larger detunings from the cycling transition without exciting other hyperfine states. At the end of the day, although with somewhat different techniques, the laser cooling part works equally well for both atoms.

Both atomic species have favorable collisional properties for evaporative cooling. The elastic scattering cross section of $^{87}$Rb atoms at low temperature is four times higher than in $^{23}$Na. However, elastic collision rates after laser cooling are comparable since $^{23}$Na atoms are faster. A peculiarity of $^{87}$Rb is that the two ground electronic state hyperfine levels have similar scattering lengths, which can be advantageous for studies on spinor condensates and atomic clock transitions. Related to that, spin relaxation between the two hyperfine levels is almost completely suppressed. Mixtures of F=1 and F=2 atoms can be kept for seconds, whereas in $^{23}$Na they decay on ms time scales. Both atoms have several Feshbach resonances below 1100 G. Here, $^{87}$Rb has the disadvantage, that the widest known resonance is only 200 mG wide compared to 1 G for $^{23}$Na and requires more stable magnetic fields. Another difference is the higher rate of three-body collisions for $^{87}$Rb atoms. As we discussed in Sec. X this imposes limitations on trapping and manipulating dense $^{87}$Rb condensates.

XII. CONCLUSIONS

In this paper, we have presented details for designing BEC machines with high performance and flexibility, and we hope that this description is useful for designing new experiments. Given the recent developments in the field, there is more than enough room for new experiments to join in the exploration of atom optics and many-body physics with quantum-degenerate atomic gases.

Funding for the $^{87}$Rb machine was provided by the NSF MIT-Harvard Center for Ultracold Atoms. Funding for the $^{23}$Na machine came from NSF, the ARO MURI program, NASA, and the ONR.

We thank S. Gupta, A. Görlitz, and A. E. Leanhardt for their contributions in the construction of the $^{23}$Na machine; J. C. Mun and P. Medley for ongoing contributions to the $^{87}$Rb machine; and MIT UROP students P. Gorenlik and X. Sun for various contributions to the $^{87}$Rb machine. The authors would also like to thank M. Saba and D. Kielpinksi for critical reading of this manuscript.

APPENDIX A: OVEN

To sustain a high flux atomic beam, the background vacuum pressure must be low enough that the mean free path between collisions is much greater than the length of the beam. To generate an effusive beam with a thermal distribution of velocities, the size of the hole through which the atoms escape must be smaller than the mean free path inside the oven. We observed in sodium that at higher pressures (e.g. temperatures) the flux of slowable atoms does not increase and that the velocity distribution narrows. This phenomena is well understood, and limits the diffusive flux from a single aperture oven.

During servicing, a clean ampoule is essential for rapid recovery of good vacuum pressure. The ampoule is cleaned by submerging it in a 50/50 mixture by volume of acetone and isopropanol for 20 minutes, air drying it, and then by placing it in the oven while still sealed. This removes most of the water from the glass surface, which would otherwise require more time to pump away. After installation the ampoule is baked for 24 hours under vacuum at 150–180°C to remove the remaining contaminates before it is broken.

To prevent accumulation of metal at the aperture (Fig. 3 J), the oven nozzle temperature (Fig. 3 K) is kept hotter (∼ 10°C in rubidium and ∼ 90°C in sodium) than the rest of the oven. The velocity distribution of the beam is determined by the nozzle temperature (Fig. 3 K). On the other hand, the vapor pressure in the oven, which controls the beam flux, is dominated by the coldest spot in the elbow and bellows. The factor of two discrepancy between the observed and calculated (Table II) rubidium oven lifetimes at 110°C can be accounted for by a spot ∼ 10°C colder then the lowest measured oven.
| Temp (°C) | Velocity (m/s) | Pressure (torr) | Total Flux (#/sec) | Lifetime (hours) |
|----------|---------------|----------------|-------------------|-----------------|
| -30      | 266           | 2.5E-10        | 1.4E+10           | –               |
| 25       | 295           | 4.0E-07        | 2.0E+13           | –               |
| 40       | 302           | 2.0E-06        | 9.7E+13           | –               |
| 90       | 325           | 1.2E-04        | 5.4E+15           | 1815            |
| 100      | 330           | 2.3E-04        | 1.1E+16           | 926             |
| 110      | 334           | 4.5E-04        | 2.0E+16           | 489             |
| 120      | 339           | 8.3E-04        | 3.7E+16           | 267             |
| 130      | 343           | 1.5E-03        | 6.5E+16           | 151             |
| 140      | 347           | 2.6E-03        | 1.1E+17           | 87              |
| 150      | 351           | 4.4E-03        | 1.7E+17           | 52              |

* Most probable, 3D Beam, Sec 5.2 Ref. [50]
* 5mm aperture
* 5g Rb
* Melting point

**TABLE III: Design Parameter for Rb oven**

The specifics of this cold spot depend on how the oven is insulated.

**APPENDIX B: ZEEMAN SLOWER**

Every photon which scatters off an atom to slow the atom is radiated in a random direction, increasing the atoms spread in transverse velocity. The beam emerging from the tube needs to have sufficient forward mean velocity to load the MOT efficiently. Because of the random direction of the emission recoil, N photon scatterings increase the transverse velocity by \( v_r \sqrt{N/3} \), or \( \sqrt{v_r \Delta v/3} \). The \(^{23}\)Na slower operates with a recoil induced transverse exit velocity of \( \approx 3 \text{m/s} \), with a final forward velocity of \( 30 \text{ m/s} \) so that the spatial transverse spread in the slowed beam matches the MOT capture area. The smaller initial and recoil velocities in the \(^{87}\)Rb slower reduce the transverse velocity to \( \approx 0.8 \text{m/s} \), making MOT capture matching less critical.

An additional concern in both slowers is the fate of atoms not captured by the MOT. In \(^{87}\)Rb we were concerned with the potential adverse impact a deposited film may have on the vapor pressure, and installed a cold plate near the slower window on the main chamber to capture desorbed Rb. Vacuum pressure has not been an issue and we have never needed to chill this cold plate. The opposite problem arises in \(^{24}\)Na, where metal deposition on the slower window reduces the transmission of slower light. We have found heating the slowing beam vacuum port window to 90°C prevents long term buildup.

### 1. Slower Construction

The vacuum portion of the \(^{87}\)Rb slower is a 99 cm long nonmagnetic 304 stainless steel tube with a 19 mm OD and 0.9 mm wall. The rear end of the tube is connected to the main chamber by a DN 16 CF rotatable flange, while the oven end of the tube has a narrow, 50mm long flexible welded bellows ending in another DN 16 CF rotatable flange. The retaining ring on this flange was cut in half for removal, so that the premounted coil assembly could be slid over the vacuum tube.

As shown in Fig. 1, the slower tube enters the main chamber at an angle of 33° from the vertical to accommodate access for optical tweezers. The oven and the Zeeman slower are supported two meters above the the experimental table in order to preserve the best optical and mechanical access to the main chamber. Aluminum extrusion from 80/20 Inc. was used to create the support framework.

**Section 1, 5.0 Amps**

**Section 2, 10.0 Amps**

**Section 3, 30.0 Amps**

**FIG. 11: Winding pattern cross section for \(^{87}\)Rb Zeeman slower consisting of three solenoids. Each drawing represents half of the cross section of each a solenoid. The “O”s represent wires, while the spaces between the wires were meant to be smoothed out to an average value during construction. Each character in the drawing represents a physical size of 3.5 mm. The wire is hollow core water cooled copper, identical to that used in construction of the magnetic trap as described in [6]. The high current coil is closest to the main chamber. The single layer uniform bias coil is not depicted.**

Our \(^{87}\)Rb slower was fabricated with a single layer bias solenoid and three increasing field segments (Fig. 1 and II). The optimum configuration of currents and solenoid winding shapes was found by computer simulated winding of the solenoids one loop at a time, starting at the high field end and tapering the last few loops to best match the desired field profile. An alternative fabrication technique would be to apply a large uniform bias field and subtract away unwanted field with counter current coils. Straw fields from the Zeeman slower can have a detrimental effect on the MOT, particularly during sudden turnoff. An additional bias coil around the main chamber along the axis of the slower can compensate for this effect.

**APPENDIX C: MAGNETIC TRAP APPENDIX**

**1. Ioffe Pritchard Trapping Potential**

The field near the minimum of a Ioffe-Pritchard trap is approximately

\[
\mathbf{B} = B_0 \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix} + B' \begin{pmatrix} x \\ -y \\ 0 \end{pmatrix} + \frac{B''}{2} \begin{pmatrix} -xz \\ -yz \\ z^2 - \frac{1}{2} (x^2 + y^2) \end{pmatrix} \quad \text{(C1)}
\]
which realizes trap frequencies

\[ m\omega_B^2 = \mu m \frac{B'^2}{B_0} \quad (C2) \]

\[ m\omega_{z}^2 = \mu m B'' \quad (C3) \]

Typical trap parameters of \( B' = 223 \, \text{G/cm}, B'' = 100 \, \text{G/cm}^2, B_0 = 1 \, \text{G} \) (Fig. 7, Table VIII) have frequencies of \( \langle \omega_B, \omega_z \rangle / 2\pi \) of \( (200, 9) \) Hz for \( ^{87}\text{Rb} \) and \( (390, 18) \) Hz for \( ^{23}\text{Na} \). Further details of Ioffe-Pritchard magnetic traps are discussed in Sec 2.3.2 of [47] and Ch. 5. of [62].

### 2. Circuitry

![Axial coils circuit diagram](image)

FIG. 12: Axial coils circuit diagram. High current wires are heavy black lines.

Fig. 12 is representative of the magnetic trap circuit. We drive the magnetic trap coils with Lambda EMI DC power supplies in fixed current mode. Current to the cloverleaf coils is supplied from a Model EMS 20-500 15kW power supply, while the axial currents are driven with two Model EMS 20-250 5kW power supplies. Each power supply is protected against damage from reverse current with an International fuse insulation by Essex Group Inc., Magnet Wire & Insulation of Charlotte, NC. The coils are held together with Hysol Epoxi-Patch IC White high temperature epoxy that is bakable to 170°C. Chilled water is forced through the hollow core of the copper wires to dissipate the \( \approx 10 \, \text{kW} \) of power generated from resistive heating in the magnetic trap and Zeeman slower coils. 200 psi of differential pressure is required for sufficient coolant flow. We designed all our coils to increase the cooling water temperature by less than 50°C. Ch. 3 of Ref. [64] has an extensive discussion of water cooling in continuously powered resistive magnets. For our wire the following empirical values were measured.

\[
\rho [\Omega/m] = 2.65 \times 10^{-3} \quad (C4)
\]

\[
Q [ml/sec] = 2.07 \sqrt{\frac{\Delta P [psi]}{L [m]}} \quad (C5)
\]

\[
\Delta T [^\circ C] = 259I^2 [Amps] \rho V \sqrt{\frac{L^3 [m]}{\Delta P [psi]}} \quad (C6)
\]

where \( Q \) is the water flow rate in ml/sec, \( I^2 \rho L \) is the power dissipated by the coil, \( \Delta P \) the pressure drop in psi (1 psi=6.89 kPa), and \( L \) the length of the coil in meters.

### 3. Wire

Both the slower and the magnetic trap coils were fabricated using square hollow core (0.125 in./side, 0.032 in. wall) Alloy 101 soft temper copper tubing from Small Tube Products, Inc. of Altoona, PA, wrapped with double Dacron glass fuse insulation by Essex Group Inc., Magnet Wire & Insulation of Charlotte, NC. The coils are held together with Hysol Epoxi-Patch IC White high temperature epoxy that is bakable to 170°C. Chilled water is forced through the hollow core of the copper wires to dissipate the \( \approx 10 \, \text{kW} \) of power generated from resistive heating in the magnetic trap and Zeeman slower coils. 200 psi of differential pressure is required for sufficient coolant flow. We designed all our coils to increase the cooling water temperature by less than 50°C. Ch. 3 of Ref. [64] has an extensive discussion of water cooling in continuously powered resistive magnets. For our wire the following empirical values were measured.

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\]

where \( Q \) is the water flow rate in ml/sec, \( I^2 \rho L \) is the power dissipated by the coil, \( \Delta P \) the pressure drop in psi (1 psi=6.89 kPa), and \( L \) the length of the coil in meters.

### 4. Fabrication

All of the components for each half of the magnetic trap were epoxied together for stability. Each assembly was then mounted in the bucket windows with an aluminum mounting plate backed by four threaded Alloy 316 stainless steel rods. No ferromagnetic materials were used in the mounting because of concern for irreproducibility from hysteresis effects. Table VIII lists the windings and typical parameters for each coil.
### TABLE IV: Magnetic trap coil winding and performance specifications.

| Coil | Winding | Current | Inner θ | Field |
|------|----------|---------|---------|-------|
| | Turns | Layers | (A) | (cm) | \(B'' = +9 \text{ G/cm}^2\) |
| | | | | | \(B_0 = -243 \text{ G}\) |
| | | | | |
| | | | | \(B'' = +90 \text{ G/cm}^2\) |
| | | | | \(B_0 = +251 \text{ G}\) |
| | | | | |
| | | | | \(B'' = +9 \text{ G/cm}^2\) |
| | | | | \(B_0 = +9 \text{ G/cm}^2\) |
| | | | | |
| | | | | \(B'' = +9 \text{ G/cm}^2\) |
| | | | | \(B_0 = +9 \text{ G/cm}^2\) |
| | | | | |
| | | | | \(B'' = +223 \text{ G/cm}^2\) |
| | | | | \(B_0 = +223 \text{ G/cm}^2\) |
| | | | | |
| | | | | \(B'' = +223 \text{ G/cm}^2\) |
| | | | | \(B_0 = +223 \text{ G/cm}^2\) |

aSegmented for improved cooling.

Fig. 8 illustrates their assembly and direction of current flow.

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