Maximal Neutrino Mixing and Maximal $CP$ Violation

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Abstract

We propose a phenomenological model of lepton mixing and $CP$ violation based on the flavor democracy of charge leptons and the mass degeneracy of neutrinos. A nearly bi-maximal flavor mixing pattern, which is favored by current data on atmospheric and solar neutrino oscillations, emerges naturally from this model after explicit symmetry breaking. The rephasing-invariant strength of $CP$ or $T$ violation can be as large as one percent, leading to significant probability asymmetries between $\nu_\mu \rightarrow \nu_e$ and $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ (or $\nu_e \rightarrow \nu_\mu$) transitions in the long-baseline neutrino experiments.
The recent observation of atmospheric and solar neutrino anomalies, in particular by the Super-Kamiokande experiment [1], has provided a strong indication that neutrinos are massive and lepton flavors are mixed. As there exist at least three different lepton families, the flavor mixing matrix may in general contain non-trivial complex phase terms. Hence CP or T violation is naturally expected in the lepton sector.

A violation of CP invariance in the quark sector can result in a variety of observable effects in hadronic weak decays. Similarly CP or T violation in the lepton sector can manifest itself in neutrino oscillations [2]. The best (and probably the only) way to observe CP- or T-violating effects in neutrino oscillations is to carry out the long-baseline appearance neutrino experiments [3].

In the scheme of three lepton families, the $3 \times 3$ flavor mixing matrix $V$ links the neutrino mass eigenstates $(\nu_1, \nu_2, \nu_3)$ to the neutrino flavor eigenstates $(\nu_e, \nu_\mu, \nu_\tau)$:

$$
\begin{pmatrix}
\nu_e \\
\nu_\mu \\
\nu_\tau
\end{pmatrix}
= 
\begin{pmatrix}
V_{e1} & V_{e2} & V_{e3} \\
V_{\mu1} & V_{\mu2} & V_{\mu3} \\
V_{\tau1} & V_{\tau2} & V_{\tau3}
\end{pmatrix}
\begin{pmatrix}
\nu_1 \\
\nu_2 \\
\nu_3
\end{pmatrix}.
$$

If neutrinos are massive Dirac fermions, $V$ can be parametrized in terms of three rotation angles and one CP-violating phase. If neutrinos are Majorana fermions, however, two additional CP-violating phases are in general needed to fully parametrize $V$. The strength of CP violation in neutrino oscillations, no matter whether neutrinos are of the Dirac or Majorana type, depends only upon a universal parameter $J$ [4], which is defined through

$$
\text{Im} \left( V_{il} V_{jm}^* V_{im} V_{jl}^* \right) = J \sum_{k,n} \epsilon_{ijk} \epsilon_{lmn}.
$$

The asymmetry between the probabilities of two CP-conjugate neutrino transitions, due to the CPT invariance and the unitarity of $V$, is uniquely given as

$$
\Delta_{CP} = P(\nu_\alpha \rightarrow \nu_\beta) - P(\bar{\nu}_\alpha \rightarrow \bar{\nu}_\beta)
= -16 J \sin F_{12} \sin F_{23} \sin F_{31}
$$

with $(\alpha, \beta) = (e, \mu)$, $(\mu, \tau)$ or $(\tau, e)$, $F_{ij} = 1.27 \Delta m^2_{ij} L/E$ and $\Delta m^2_{ij} = m_i^2 - m_j^2$, in which $L$ is the distance between the neutrino source and the detector (in unit of km) and $E$ denotes the neutrino beam energy (in unit of GeV). The T-violating asymmetry can be obtained in a similar way [4]:

$$
\Delta_T = P(\nu_\alpha \rightarrow \nu_\beta) - P(\nu_\beta \rightarrow \nu_\alpha)
= -16 J \sin F_{12} \sin F_{23} \sin F_{31}.
$$

These formulas show clearly that CP or T violation is a feature of all three lepton families. The relationship $\Delta_T = \Delta_{CP}$ is a straightforward consequence of CPT invariance. The observation of $\Delta_T$ might be free from the matter effects of the earth, which is possible to fake the genuine CP asymmetry $\Delta_{CP}$ in any long-baseline neutrino experiment. The joint measurement of $\nu_\alpha \rightarrow \nu_\beta$ and $\nu_\beta \rightarrow \nu_\alpha$ transitions to determine $\Delta_T$ is, however, a challenging task in practice. Probably it could only be realized in a neutrino factory, whereby high-quality neutrino beams can be produced with high-intensity muon storage rings [4].

3Note that an asymmetry between the probabilities $P(\bar{\nu}_\alpha \rightarrow \nu_\beta)$ and $P(\bar{\nu}_\beta \rightarrow \nu_\alpha)$ signifies T violation too. This asymmetry and that defined in Eq. (4) have the same magnitude but opposite signs.
Analyses of current experimental data [1, 6] yield $\Delta m^2_{\text{sun}} \ll \Delta m^2_{\text{atm}}$ and $|V_{e3}|^2 \ll 1$, implying that the atmospheric and solar neutrino oscillations are approximately decoupled. A reasonable interpretation of those data follows from setting $\Delta m^2_{\text{sun}} = |\Delta m^2_{12}|$ and $\Delta m^2_{\text{atm}} = |\Delta m^2_{23}| \approx |\Delta m^2_{31}|$. In this approximation $F_{31} \approx -F_{23}$ holds. The $CP$- and $T$-violating asymmetries can then be simplified as

$$\Delta_{CP} = \Delta_T \approx 16J \sin F_{12} \sin^2 F_{23}. \quad (5)$$

Note that $\Delta_{CP}$ or $\Delta_T$ depends linearly on the oscillating term $\sin F_{12}$, therefore the length of the baseline suitable for measuring $CP$ and $T$ asymmetries should satisfy the condition $|\Delta m^2_{12}| \sim E/L$. This requirement singles out the large-angle MSW solution, which has $\Delta m^2_{\text{sun}} \sim 10^{-5}$ to $10^{-4}$ eV$^2$ and $\sin^2 2\theta_{\text{sun}} \sim 0.65$ to 1 [1], among three possible solutions to the solar neutrino problem. The small-angle MSW solution is not favored; it does not give rise to a relatively large magnitude of $J$, which determines the significance of practical $CP$- or $T$-violating signals. The long wave-length vacuum oscillation requires $\Delta m^2_{\text{sun}} \sim 10^{-10}$ eV$^2$, too small to meet the realistic long-baseline prerequisite.

In this paper we extend our previous hypothesis of lepton flavor mixing [10], which arises naturally from the breaking of flavor democracy for charged leptons and that of mass degeneracy for neutrinos, to include $CP$ violation. It is found that the rephasing-invariant strength of $CP$ or $T$ violation can be as large as one percent. The flavor mixing pattern remains nearly bi-maximal, thus both atmospheric and solar neutrino oscillations can well be interpreted. The consequences of the model on the future long-baseline neutrino experiments will also be discussed in some detail.

The phenomenological constraints obtained from various neutrino oscillation experiments indicate that the mass differences in the neutrino sector are tiny compared to those in the charged lepton sector. One possible interpretation is that all three neutrinos are nearly degenerate in mass. In this case one might expect that the flavor mixing pattern of leptons differs qualitatively from that of quarks, where both up and down sectors exhibit a strong hierarchical structure in their mass spectra and the observed mixing angles are rather small. A number of authors have argued that the hierarchy of quark masses and the smallness of mixing angles are related to each other, by considering specific symmetry limits [11]. One particular way to proceed is to consider the limit of subnuclear democracy, in which the mass matrices of both the up- and down-type quarks are of rank one and have the structure

$$M_q = \frac{c_q}{3} \begin{pmatrix} 1 & 1 & 1 \\ 1 & 1 & 1 \\ 1 & 1 & 1 \end{pmatrix} \quad (6)$$

with $q = u$ (up) or $d$ (down) as well as $c_u = m_t$ and $c_d = m_b$. Small departures from the democratic limit lead to the flavor mixing and at the same time introduce the masses of the second and first families. Specific symmetry breaking schemes have been proposed in some literature in order to calculate the flavor mixing angles in terms of the quark mass eigenvalues (see, e.g., Ref. [11]).

Since the charged leptons exhibit a similar hierarchical mass spectrum as the quarks, it would be natural to consider the limit of subnuclear democracy for the $(e, \mu, \tau)$ system, i.e.,

\begin{footnote}{Throughout this work we do not take the LSND evidence for neutrino oscillations [7], which was not confirmed by the KARMEN experiment [8], into account.}

\end{footnote}
the mass matrix takes the form as Eq. (6). In the same limit three neutrinos are degenerate in mass. Then we have \[10\]

\[
M_l^{(0)} = \frac{c_l}{3} \begin{pmatrix}
1 & 1 & 1 \\
1 & 1 & 1 \\
1 & 1 & 1
\end{pmatrix},
\]

\[
M_\nu^{(0)} = c_\nu \begin{pmatrix}
1 & 0 & 0 \\
0 & 1 & 0 \\
0 & 0 & 1
\end{pmatrix},
\]

(7)

where \(c_l = m_\tau\) and \(c_\nu = m_0\) measure the corresponding mass scales. If the three neutrinos are of the Majorana type, \(M_\nu^{(0)}\) could take a more general form \(M_\nu^{(0)}P_\nu\) with \(P_\nu = \text{Diag}\{1, e^{i\phi_1}, e^{i\phi_2}\}\). As the Majorana phase matrix \(P_\nu\) has no effect on the flavor mixing and \(CP\)-violating observables in neutrino oscillations, it will be neglected in the subsequent discussions. Clearly \(M_\nu^{(0)}\) exhibits an \(S(3)\) symmetry, while \(M_l^{(0)}\) an \(S(3)_L \times S(3)_R\) symmetry.

One can transform the charged lepton mass matrix from the democratic basis \(M_l^{(0)}\) into the hierarchical basis

\[
M_l^{(H)} = c_l \begin{pmatrix}
0 & 0 & 0 \\
0 & 0 & 0 \\
0 & 0 & 1
\end{pmatrix}
\]

(8)

by means of an orthogonal transformation, i.e., \(M_l^{(H)} = U M_l^{(0)} U^T\), where

\[
U = \begin{pmatrix}
\frac{1}{\sqrt{2}} & -\frac{1}{\sqrt{2}} & 0 \\
\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{6}} & -\frac{\sqrt{2}}{\sqrt{3}} \\
\frac{\sqrt{3}}{\sqrt{3}} & \frac{\sqrt{3}}{\sqrt{3}} & \frac{\sqrt{3}}{\sqrt{3}}
\end{pmatrix}.
\]

(9)

We see \(m_e = m_\mu = 0\) from \(M_l^{(H)}\) and \(m_1 = m_2 = m_3 = m_0\) from \(M_\nu^{(0)}\). Of course there is no flavor mixing in this symmetry limit.

A simple real diagonal breaking of the flavor democracy for \(M_l^{(0)}\) and the mass degeneracy for \(M_\nu^{(0)}\) may lead to instructive results for flavor mixing in neutrino oscillations \[10, 12\]. To accommodate \(CP\) violation, however, complex perturbative terms are required. Let us proceed with two different symmetry-breaking steps in close analogy to the symmetry breaking discussed previously for the quark mass matrices \[13, 14\].

First, small real perturbations to the (3,3) elements of \(M_l^{(0)}\) and \(M_\nu^{(0)}\) are respectively introduced:

\[
\Delta M_l^{(1)} = \frac{c_l}{3} \begin{pmatrix}
0 & 0 & 0 \\
0 & 0 & 0 \\
0 & 0 & \varepsilon_l
\end{pmatrix},
\]

\[
\Delta M_\nu^{(1)} = c_\nu \begin{pmatrix}
0 & 0 & 0 \\
0 & 0 & 0 \\
0 & 0 & \varepsilon_\nu
\end{pmatrix}.
\]

(10)

In this case the charged lepton mass matrix \(M_l^{(1)} = M_l^{(0)} + \Delta M_l^{(1)}\) remains symmetric under an \(S(2)_L \times S(2)_R\) transformation, and the neutrino mass matrix \(M_\nu^{(1)} = M_\nu^{(0)} + \Delta M_\nu^{(0)}\) has an \(S(2)\) symmetry. The muon becomes massive (i.e., \(m_\mu \approx 2|\varepsilon_l|m_\tau/9\)), and the mass eigenvalue \(m_3\) is no more degenerate with \(m_1\) and \(m_2\) (i.e., \(|m_3 - m_0| = m_0 |\varepsilon_\nu|\)). After the diagonalization
of $M_l^{(1)}$ and $M_l^{(2)}$, one finds that the 2nd and 3rd lepton families have a definite flavor mixing angle $\theta$. We obtain $\tan \theta \approx -\sqrt{2}$ if the small correction of $O(m_\mu/m_\tau)$ is neglected. Then neutrino oscillations at the atmospheric scale may arise in $\nu_\mu \to \nu_\tau$ transitions with $\Delta m^2_{31} \approx 2m_0|\varepsilon_l|$. The corresponding mixing factor $\sin^2 2\theta \approx 8/9$ is in good agreement with current data.

The symmetry breaking given in Eq. (10) for the charged lepton mass matrix serves as a good illustrative example. One could consider a more general case, analogous to the one for quarks [13], to break the $S(3)_L \times S(3)_R$ symmetry of $M_l^{(0)}$ to an $S(2)_L \times S(2)_R$ symmetry. This would imply that $\Delta M_l^{(1)}$ takes the form

$$\Delta M_l^{(1)} = \frac{c_l}{3} \begin{pmatrix} 0 & 0 & \varepsilon_l' \\ 0 & 0 & \varepsilon_l' \\ \varepsilon_l' & \varepsilon_l' & \varepsilon_l \end{pmatrix}, \quad (11)$$

where $|\varepsilon_l| \ll 1$ and $|\varepsilon_l'| \ll 1$. In this case the leading-order results obtained above, i.e., $\tan \theta \approx -\sqrt{2}$ and $\sin^2 2\theta \approx 8/9$, remain unchanged.

At the next step we introduce a complex symmetry breaking perturbation, analogous to that for quark mass matrices discussed in Ref. [15], to the charged lepton mass matrix $M_l^{(1)}$:

$$\Delta M_l^{(2)} = \frac{c_l}{3} \begin{pmatrix} 0 & -i\delta_l & i\delta_l \\ i\delta_l & 0 & -i\delta_l \\ -i\delta_l & i\delta_l & 0 \end{pmatrix}. \quad (12)$$

Transforming $M_l^{(2)} = M_l^{(1)} + \Delta M_l^{(2)}$ into the hierarchical basis, we obtain

$$M_l^H = c_l \begin{pmatrix} 0 & -i\frac{1}{\sqrt{3}}\delta_l & 0 \\ i\frac{1}{\sqrt{3}}\delta_l & \frac{2}{\sqrt{3}}\varepsilon_l & -\frac{\sqrt{2}}{9}\varepsilon_l \\ 0 & -\frac{\sqrt{2}}{9}\varepsilon_l & 1 + \frac{1}{9}\varepsilon_l \end{pmatrix}. \quad (13)$$

Note that $M_l^H$, just like a variety of realistic quark mass matrices [11], has texture zeros in the (1,1), (1,3) and (3,1) positions. The phases of its (1,2) and (2,1) elements are $\pm 90^\circ$, which could lead to maximal $CP$ violation if the neutrino mass matrix is essentially real. For the latter we consider a small perturbation, analogous to that in Eq. (10), to break the remaining mass degeneracy of $M_l^{(1)}$:

$$\Delta M_{\nu}^{(2)} = c_\nu \begin{pmatrix} -\delta_\nu & 0 & 0 \\ 0 & \delta_\nu & 0 \\ 0 & 0 & 0 \end{pmatrix}. \quad (14)$$

From $\Delta M_l^{(2)}$ and $\Delta M_{\nu}^{(2)}$ we obtain $m_e \approx |\delta_l|^2 m_\tau^2/(27 m_\mu)$ and $m_{1,2} = m_0 (1 \mp \delta_\nu)$, respectively. The simultaneous diagonalization of $M_l^{(2)} = M_l^{(1)} + \Delta M_l^{(2)}$ and $M_{\nu}^{(2)} = M_{\nu}^{(1)} + \Delta M_{\nu}^{(2)}$ leads to a full $3 \times 3$ flavor mixing matrix, which links neutrino mass eigenstates ($\nu_1, \nu_2, \nu_3$) to neutrino flavor eigenstates ($\nu_e, \nu_\mu, \nu_\tau$) in the following manner:

$$V = U + i \xi_V \sqrt{\frac{m_e}{m_\mu}} + \xi_V \frac{m_\mu}{m_\tau}, \quad (15)$$

where $U$ has been given in Eq. (9), and

$$\xi_V = \begin{pmatrix} \frac{1}{\sqrt{6}} & \frac{1}{\sqrt{6}} & -\frac{2}{\sqrt{6}} \\ \frac{1}{\sqrt{2}} & -\frac{1}{\sqrt{2}} & 0 \\ 0 & 0 & 0 \end{pmatrix}. \quad (16)$$
\[ \zeta_V = \begin{pmatrix} 0 & 0 & 0 \\ \frac{1}{\sqrt{6}} & \frac{1}{\sqrt{6}} & \frac{1}{\sqrt{6}} \\ \frac{1}{\sqrt{12}} & \frac{1}{\sqrt{12}} & \frac{1}{\sqrt{3}} \end{pmatrix}. \] 

(16)

In comparison with the result of Ref. [10], the new feature of this lepton mixing scenario is that the term multiplying \( \xi \) becomes imaginary. Therefore \( CP \) or \( T \) violation has been incorporated.

The complex symmetry breaking perturbation given in Eq. (12) is certainly not the only one which can be considered for \( M_l^{(1)} \). Below we list a number of other interesting possibilities, i.e., the hermitian perturbations

\[
\Delta \tilde{M}_l^{(2)} = \frac{c_l}{3} \begin{pmatrix} 0 & -i\delta_l & 0 \\ i\delta_l & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}, \\
\Delta \hat{M}_l^{(2)} = \frac{c_l}{3} \begin{pmatrix} 0 & 0 & i\delta_l \\ 0 & 0 & -i\delta_l \\ -i\delta_l & i\delta_l & 0 \end{pmatrix};
\]

(17)

and the non-hermitian perturbations

\[
\Delta M_l^{(2)} = \frac{c_l}{3} \begin{pmatrix} -i\delta_l & 0 & i\delta_l \\ 0 & i\delta_l & -i\delta_l \\ i\delta_l & -i\delta_l & 0 \end{pmatrix}, \\
\Delta \tilde{M}_l^{(2)} = \frac{c_l}{3} \begin{pmatrix} -i\delta_l & 0 & 0 \\ 0 & i\delta_l & 0 \\ 0 & 0 & 0 \end{pmatrix}, \\
\Delta \hat{M}_l^{(2)} = \frac{c_l}{3} \begin{pmatrix} 0 & 0 & i\delta_l \\ 0 & 0 & -i\delta_l \\ i\delta_l & -i\delta_l & 0 \end{pmatrix};
\]

(18)

The three hermitian and three non-hermitian perturbative mass matrices obey the following sum rules:

\[
\Delta M_l^{(2)} = \Delta \tilde{M}_l^{(2)} + \Delta \hat{M}_l^{(2)}, \\
\Delta \hat{M}_l^{(2)} = \Delta \tilde{M}_l^{(2)} + \Delta \hat{M}_l^{(2)}. 
\]

(19)

Let us remark that hermitian perturbations of the same forms as given in Eqs. (12) and (17) have been used to break the flavor democracy of quark mass matrices and to generate \( CP \) violation [13, 15]. The key point of this similarity between the charged lepton and quark mass matrices is that both of them have the strong mass hierarchy and might have the same dynamical origin or a symmetry relationship.

To be more specific we transform all the six charged lepton mass matrices

\[
M_l^{(2)} = M_l^{(0)} + \Delta M_l^{(1)} + \Delta M_l^{(2)}, \\
\tilde{M}_l^{(2)} = M_l^{(0)} + \Delta \tilde{M}_l^{(1)} + \Delta \tilde{M}_l^{(2)}, \\
\hat{M}_l^{(2)} = M_l^{(0)} + \Delta \hat{M}_l^{(1)} + \Delta \hat{M}_l^{(2)};
\]

(20)

and

\[
\tilde{M}_l^{(2)} = M_l^{(0)} + \Delta M_l^{(1)} + \Delta \tilde{M}_l^{(2)}, \\
\hat{M}_l^{(2)} = M_l^{(0)} + \Delta \tilde{M}_l^{(1)} + \Delta \hat{M}_l^{(2)}, \\
\hat{M}_l^{(2)} = M_l^{(0)} + \Delta \hat{M}_l^{(1)} + \Delta \hat{M}_l^{(2)}.
\]

(21)
It is straightforward to obtain \( \Delta M_l^{(2)} \), \( \Delta \bar{M}_l^{(2)} \), and its parallelism with \( M_l \). One may also argue that the simplicity of \( \bar{\nu}_l \) and its quark counterpart \([11, 15]\) could provide us a useful hint towards an underlying symmetry of these hierarchical mass matrices is that their (1,1) elements all vanish. For this reason, we can neglect for our present purpose.

The non-hermitian perturbations \( \Delta M_l \) and \( \Delta \bar{M}_l \) respectively, are approximately independent of other details of the flavor symmetry breaking and have the identical strength to a high degree of accuracy. Indeed it is easy to check that all the six charged lepton mass matrices in Eqs. (20) and (21), together with the neutrino mass matrix \( M_\nu^{(2)} = M_\nu^{(0)} + \Delta M_\nu^{(1)} + \Delta M_\nu^{(2)} \), lead to the same flavor mixing pattern \( V \) as given in Eq. (15). Hence it is in practice difficult to distinguish one scenario from another. In our point of view, the similarity between \( M_l^{(2)} \) and its quark counterpart \([11, 13]\) could provide us a useful hint towards an underlying symmetry between quarks and charged leptons. One may also argue that the simplicity of \( \bar{M}_l^{(2)} \) and its parallelism with \( M_\nu^{(2)} \) might make it technically more natural to be derived from a yet unknown fundamental theory of lepton mixing and CP violation.

The flavor mixing matrix \( V \) can in general be parametrized in terms of three Euler angles and one \( CP \)-violating phase \([1]\). A suitable parametrization reads as follows \([13]\):

\[
V = \begin{pmatrix} c_l & s_l & 0 \\ -s_l & c_l & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} e^{-i\phi} & 0 & 0 \\ 0 & c & s \\ 0 & -s & c \end{pmatrix} \begin{pmatrix} c_\nu & -s_\nu & 0 \\ s_\nu & c_\nu & 0 \\ 0 & 0 & 1 \end{pmatrix} = \begin{pmatrix} s_l s_\nu c + c_l c_\nu e^{-i\phi} & s_l c_\nu c - c_l s_\nu e^{-i\phi} & s_l \nu \\ c_l s_\nu c - s_l c_\nu e^{-i\phi} & c_l c_\nu c + s_l s_\nu e^{-i\phi} & c_l \nu s \\ -s_\nu s & -c_\nu s & c \end{pmatrix},
\]

(22)

in which \( s_l \equiv \sin \theta_l, \ s_\nu \equiv \sin \theta_\nu, \ c \equiv \cos \theta, \) etc. The three mixing angles can all be arranged to lie in the first quadrant, while the \( CP \)-violating phase may take values between 0 and \( 2\pi \). It is straightforward to obtain \( J = s_l c_l s_\nu c s^2 c \sin \phi \). Numerically we find

\[
\theta_l \approx 4^\circ, \quad \theta_\nu \approx 45^\circ, \quad \theta \approx 52^\circ, \quad \phi \approx 90^\circ.
\]

(23)

\[5\]For neutrinos of the Majorana type, two additional \( CP \)-violating phases may enter. But they are irrelevant to neutrino oscillations and can be neglected for our present purpose.
The smallness of $\theta_l$ is a natural consequence of the mass hierarchy in the charged lepton sector, just as the smallness of $\theta_u$ in quark mixing \[1\]. On the other hand, both $\theta_u$ and $\theta$ are too large to be comparable with the corresponding quark mixing angles (i.e., $\theta_d$ and $\theta$ as defined in Ref. \[11\]), reflecting the qualitative difference between quark and lepton flavor mixing phenomena. It is worth emphasizing that the leptonic $CP$-violating phase $\phi$ takes a special value ($\approx 90^\circ$) in our model. The same possibility is also favored for the quark mixing phenomenon in a variety of realistic mass matrices \[17\]. Therefore maximal leptonic $CP$ violation, in the sense that the magnitude of $J$ is maximal for the fixed values of three flavor mixing angles, appears naturally as in the quark sector.

Some consequences of this lepton mixing scenario can be drawn as follows:

(1) The mixing pattern in Eq. (15), after neglecting small corrections from the charged lepton masses, is quite similar to that of the pseudoscalar mesons $\pi^0$, $\eta$ and $\eta'$ in QCD in the limit of the chiral $SU(3)_L \times SU(3)_R$ symmetry \[14\]:

\[
\begin{align*}
\pi^0 &= \frac{1}{\sqrt{2}} \left( |\bar{u}u| - |\bar{d}d| \right), \\
\eta &= \frac{1}{\sqrt{6}} \left( |\bar{u}u| + |\bar{d}d| - 2 |\bar{s}s| \right), \\
\eta' &= \frac{1}{\sqrt{3}} \left( |\bar{u}u| + |\bar{d}d| + 2 |\bar{s}s| \right).
\end{align*}
\]

(24)

A theoretical derivation of the flavor mixing matrix $V \approx U$ has been given in Ref. \[18\], in the framework of a left-right symmetric extension of the standard model with $S(3)$ and $Z(4) \times Z(3) \times Z(2)$ symmetries.

(2) The $V_{e3}$ element, of magnitude

\[
|V_{e3}| = \frac{2}{\sqrt{6}} \sqrt{\frac{m_e}{m_\mu}},
\]

is naturally suppressed in the symmetry breaking scheme outlined above. A similar feature appears in the $3 \times 3$ quark flavor mixing matrix, i.e., $|V_{ub}|$ is the smallest among the nine quark mixing elements. Indeed the smallness of $V_{e3}$ provides a necessary condition for the decoupling of solar and atmospheric neutrino oscillations, even though neutrino masses are nearly degenerate. The effect of small but nonvanishing $V_{e3}$ will manifest itself in long-baseline $\nu_\mu \rightarrow \nu_e$ and $\nu_e \rightarrow \nu_\tau$ transitions, as already shown in Ref. \[10\].

(3) The flavor mixing between the 1st and 2nd lepton families and that between the 2nd and 3rd lepton families are nearly maximal. This property, together with the natural smallness of $|V_{e3}|$, allows a satisfactory interpretation of the observed large mixing in atmospheric and solar neutrino oscillations. We obtain \[\footnote{In calculating $\sin^2 2\theta_{\text{sun}}$, we have taken the $O(m_e/m_\mu)$ correction to the expression of $V$ into account.}]

\[
\begin{align*}
\sin^2 2\theta_{\text{sun}} &= 1 - \frac{4}{3} \frac{m_e}{m_\mu}, \\
\sin^2 2\theta_{\text{atm}} &= \frac{8}{9} + \frac{8}{9} \frac{m_\mu}{m_\tau}.
\end{align*}
\]

(26)
solution to the solar neutrino problem \[19\] and might also be able to incorporate the large-angle MSW solution \[20\].

(4) The rephasing-invariant strength of CP violation in this scheme is given as

\[
\mathcal{J} = \frac{1}{3\sqrt{3}} \sqrt{\frac{m_e}{m_\mu}} \left(1 + \frac{1}{2} \frac{m_\mu}{m_\tau}\right).
\]

Explicitly we have \(\mathcal{J} \approx 1.4\%\). The large magnitude of \(\mathcal{J}\) for lepton mixing turns out to be very non-trivial, as the same quantity for quark mixing is only of order \(10^{-5}\) \[11, 17\]. If the mixing pattern under discussion were in no conflict with the large-angle MSW solution to the solar neutrino problem, then the CP- and T-violating signals \(\Delta_{CP} = \Delta_T \propto -16\mathcal{J} \approx -0.2\) could be significant enough to be measured from the asymmetry between \(P(\nu_\mu \to \nu_e)\) and \(P(\bar{\nu}_\mu \to \bar{\nu}_e)\) or that between \(P(\nu_\mu \to \nu_e)\) and \(P(\nu_e \to \nu_\mu)\) in the long-baseline neutrino experiments. In the leading-order approximation we arrive at

\[
\mathcal{A} = \frac{P(\nu_\mu \to \nu_e) - P(\bar{\nu}_\mu \to \bar{\nu}_e)}{P(\nu_\mu \to \nu_e) + P(\bar{\nu}_\mu \to \bar{\nu}_e)} \\
= \frac{P(\nu_\mu \to \nu_e) - P(\nu_e \to \nu_\mu)}{P(\nu_\mu \to \nu_e) + P(\nu_e \to \nu_\mu)} \\
= \frac{-8}{3\sqrt{3}} \sqrt{\frac{m_e}{m_\mu}} \sin F_{12}.
\]

The asymmetry \(\mathcal{A}\) depends linearly on the oscillating term \(\sin F_{12}\), which is associated essentially with the solar neutrino anomaly.

To give one a numerical estimate of the magnitudes of \(\Delta_{CP}\) and \(\mathcal{A}\), we typically take the baseline length to be \(L = 732\) km or \(L = 7332\) km for a neutrino source at Fermilab pointing toward the Soudan mine in Minnesota or the Gran Sasso underground laboratory in Italy \[3\]. The mass-squared differences are chosen as (a) \(|\Delta m^2_{12}| = 5 \times 10^{-5} \text{ eV}^2\) and (b) \(|\Delta m^2_{23}| = 10^{-4} \text{ eV}^2\) versus the fixed \(|\Delta m^2_{23}| = 10^{-3} \text{ eV}^2\). The behaviors of the asymmetries \(|\Delta_{CP}|\) (or \(|\Delta_T|\)) and \(|\mathcal{A}|\) changing with the beam energy \(E\) in the range \(3\) GeV \(\leq E \leq 20\) GeV are then shown in Figs. 1 and 2, respectively. Clearly the asymmetry \(\mathcal{A}\) can be of \(\mathcal{O}(0.1)\) to \(\mathcal{O}(1)\), even though the corresponding magnitude of \(\Delta_{CP}\) (or \(\Delta_T\)) is about two-order smaller. In reality the matter effect on these \(CP\) asymmetries should be taken into account, in order to extract the genuine \(CP\)-odd parameters. For the model under consideration, the smallness of \(|V_{e3}| \approx 0.056\) together with the maximal \(CP\) violating phase \((\phi \approx 90^\circ)\) is expected to make the possible matter effect insignificant, and unable to completely fake the genuine \(CP\)-violating signals \[21\].

If the upcoming data appeared to rule out the consistency between our model and the large-angle MSW solution to the solar neutrino problem, then it would be quite difficult to test the model itself from its prediction for large \(CP\) and \(T\) asymmetries in any realistic long-baseline experiment.

Finally it is worth remarking that our lepton mixing pattern has no conflict with current constraints on the neutrinoless double beta decay \[22\], if neutrinos are of the Majorana type. In the presence of \(CP\) violation, the effective mass term of the \((\bar{\beta}\beta)_{0}\) decay can be written
Figure 1: Illustrative plots for the $CP$-violating asymmetry $|\Delta_{CP}|$ (between $\nu_\mu \rightarrow \nu_e$ and $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ transitions) changing with the neutrino beam energy $E$, where (a) $|\Delta m^2_{12}| = 5 \times 10^{-5} \text{ eV}^2$ and (b) $|\Delta m^2_{12}| = 10^{-4} \text{ eV}^2$ versus the fixed $|\Delta m^2_{23}| = 10^{-3} \text{ eV}^2$ have typically been taken in the case of the baseline length $L = 732 \text{ km}$ or $L = 7332 \text{ km}$. 
Figure 2: Illustrative plot for the $CP$-violating asymmetry $|A|$ (between $\nu_\mu \to \nu_e$ and $\bar{\nu}_\mu \to \bar{\nu}_e$ transitions) changing with the neutrino beam energy $E$, where (a) $|\Delta m^2_{12}| = 5 \times 10^{-5}$ eV$^2$ and (b) $|\Delta m^2_{12}| = 10^{-4}$ eV$^2$ versus the fixed $|\Delta m^2_{23}| = 10^{-3}$ eV$^2$ have typically been taken in the case of the baseline length $L = 732$ km or $L = 7332$ km.

as

$$\langle M \rangle_{(\beta\beta)_{0\nu}} = \sum_{i=1}^{3} \left( m_i \tilde{V}_{ei}^2 \right), \quad (29)$$

where $\tilde{V} = VP_\nu$ and $P_\nu = \text{Diag}\{1, e^{i\phi_1}, e^{i\phi_2}\}$ is the Majorana phase matrix. If the unknown phases are taken to be $\phi_1 = \phi_2 = 90^\circ$ for example, then one arrives at

$$\left|\langle M \rangle_{(\beta\beta)_{0\nu}}\right| = \frac{2}{\sqrt{3}} \sqrt{\frac{m_e}{m_\mu}} \frac{m_i}{m_i}, \quad (30)$$

in which $m_i \sim 1 - 2$ eV (for $i = 1, 2, 3$) as required by the near degeneracy of three neutrinos in our model to accommodate the hot dark matter of the universe. Obviously $|\langle M \rangle_{(\beta\beta)_{0\nu}}| \approx 0.08m_i \leq 0.2$ eV, the latest bound of the $(\beta\beta)_{0\nu}$ decay [22].

In summary, we have extended our previous model of the nearly bi-maximal lepton flavor mixing to incorporate large $CP$ violation. The new model remains favored by current data on atmospheric and solar neutrino oscillations, and it predicts significant $CP$- and $T$-violating effects in the long-baseline neutrino experiments. We expect that more data from the Super-Kamiokande and other neutrino experiments could soon provide stringent tests of the existing lepton mixing models and give useful hints towards the symmetry or dynamics of lepton mass generation.
References

[1] Y. Fukuda et al., Phys. Lett. B 436 (1998) 33; Phys. Rev. Lett. 81 (1998) 1562; Y. Suzuki, talk given at the 17th International Workshop on Weak Interactions and Neutrinos, Cape Town, South Africa, January 1999; and references therein.

[2] N. Cabibbo, Phys. Lett. B 72 (1978) 333; V. Barger, K. Whisnant, and R.J.N. Phillips, Phys. Rev. Lett. 45 (1980) 2084.

[3] See, e.g., M. Tanimoto, Phys. Rev. D 55 (1997) 322; J. Arafune and J. Sato, Phys. Rev. D 55 (1997) 1653; H. Minakata and H. Nunokawa, Phys. Lett. B 413 (1997) 369; S.M. Bilenky, C. Giunti, and W. Grimus, Phys. Rev. D 58 (1998) 033001; V. Barger, Y.B. Dai, K. Whisnant, and B.L. Young, Phys. Rev. D 59 (1999) 113010; K.R. Schubert, hep-ph/9902213; K. Dick, M. Freund, M. Lindner, and A. Romanino, hep-ph/9903308; J. Bernabéu, hep-ph/9904474; Z.Z. Xing, hep-ph/9908381.

[4] C. Jarlskog, Phys. Rev. Lett. 55 (1985) 1039.

[5] V. Barger, S. Geer, and K. Whisnant, hep-ph/9906487; A. De Rújula, M.B. Gavela, and P. Hernández, Nucl. Phys. B 547 (1999) 21.

[6] M. Apollonio et al., Phys. Lett. B 420 (1998) 397.

[7] C. Athanassopoulos et al., Phys. Rev. Lett. 75 (1995) 2650.

[8] B. Zeitnitz, talk given at Neutrino ’98, Takayama, Japan, June 1998.

[9] J.N. Bahcall, P.I. Krastev, and A.Y. Smirnov, Phys. Rev. D 58 (1998) 096016; and references therein.

[10] H. Fritzsch and Z.Z. Xing, Phys. Lett. B 372 (1996) 265; Phys. Lett. B 440 (1998) 313.

[11] H. Fritzsch and Z.Z. Xing, Nucl. Phys. B 556 (1999) 49; hep-ph/9904286; and references therein.

[12] For a brief overview, see: M. Tanimoto, hep-ph/9807517; and references therein.

[13] H. Fritzsch and J. Plankl, Phys. Lett. B 237 (1990) 451; S. Meshkov, in the Proceedings of the Global Foundation International Conference Unified Symmetry in the Small and in the Large, edited by B.N. Kursunoglu and A. Perlmutter (Nova Science, New York, 1994), p. 195.

[14] H. Fritzsch and D. Holtmannspötter, Phys. Lett. B 338 (1994) 290.

[15] H. Lehmann, C. Newton, and T.T. Wu, Phys. Lett. B 384 (1996) 249.

[16] H. Fritzsch and Z.Z. Xing, Phys. Lett. B 413 (1997) 396; Phys. Rev. D 57 (1998) 594.

[17] H. Fritzsch and Z.Z. Xing, Phys. Lett. B 353 (1995) 114.

[18] R.N. Mohapatra and S. Nussinov, Phys. Lett. B 441 (1998) 299.

[19] V. Barger and K. Whisnant, hep-ph/9903262; and references therein.

[20] S.T. Petcov, hep-ph/9907216.

[21] See, e.g., M. Tanimoto, hep-ph/9906516; A. Donini, M.B. Gavela, P. Hernández, and S. Rigolin, hep-ph/9909254.

[22] L. Baudis et al., hep-ph/9902014.