Classical analog of the quantum metric tensor

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We present a classical analog of the quantum metric tensor, which is defined for classical integrable systems that undergo an adiabatic evolution governed by slowly varying parameters. This classical metric measures the distance, on the parameter space, between two infinitesimally different points in phase space, whereas the quantum metric tensor measures the distance between two infinitesimally different quantum states. We discuss the properties of this metric and calculate its components, exactly in the cases of the generalized harmonic oscillator, the generalized harmonic oscillator with a linear term, and perturbatively for the quartic anharmonic oscillator. Finally, we propose alternative expressions for the quantum metric tensor and Berry’s connection in terms of quantum operators.

I. INTRODUCTION

Two fundamental structures for understanding the geometrical aspects of quantum states are the quantum metric tensor formulated by Provost and Valée [1, 2] and the geometric phases, in particular, the phase discovered by Berry [3]. The quantum metric tensor is defined in the parameter space and measures the distance between two states corresponding to infinitesimally different parameters. Remarkably, the singularities of this metric are associated with quantum phase transitions exhibited by the corresponding system [4, 5]. Further, the geodesics induced by this metric can also indicate the presence of quantum phase transitions [6, 7]. In general, the quantum metric tensor played an essential role in diverse physical phenomena (see Ref. [8] and references therein). Berry’s phase is the extra phase acquired by the wave function when the system undergoes an adiabatic excursion along a closed path in the parameter space and can be understood as an integral of a curvature [9], the so-called Berry curvature. This phase was analyzed in various contexts [10–13], and, interestingly, it is also connected with quantum phase transitions [14]. These approaches to quantum phase transitions based on the metric and the Berry phase can be unified in terms of the critical singular behavior of the quantum geometric tensor [15, 16], whose real part gives the quantum metric tensor whereas the imaginary part gives the Berry curvature.

On the other hand, Berry’s phase possesses a classical counterpart known as Hannay’s angle [17]. For classical integrable systems, it is an extra angle shift picked up by the angle variables of the system when the parameters undergo a closed adiabatic excursion in the parameter space. This classical angle was investigated in a variety of systems [18–21], and the semiclassical relation between it and Berry’s phase was established in Ref. [22] and has been verified in many systems [22–25].

In the light of this and given the close relationship between the quantum metric tensor and Berry’s curvature, a natural question arises: What about the classical analog of the quantum metric tensor? It is well known that, in the context of thermodynamic systems, Weinhold [26] and later Ruppeiner [27] proposed classical metrics in the parameter space which are defined as the Hessians of a thermodynamic potential. For Weinhold’s metric, the potential is the internal energy, whereas for Ruppeiner’s metric the potential is the entropy. In spite of the existence of these classical metrics, there is so far no evidence for that in the context of classical mechanical systems.

In this paper, we present a meaningful metric tensor for classical integrable systems, which is defined in the parameter space and is the classical analog of the quantum metric tensor. These metrics are analogous in the sense that both yield the same parameter structure, modulo the use of the Bohr-Sommerfeld quantization rule for action variables. It means that we can extract out the same (or almost the same) “relevant” information from either of these metrics. This important feature will be exhibited by the three examples that we have considered: the generalized harmonic oscillator, the generalized harmonic oscillator with a linear term, and the quartic anharmonic oscillator. Another important property of this classical metric, which is shared with Hannay’s angle, is that it is gauge invariant in the parameter space in that it does not depend on the choice of the point of origin from which we measure the angle variables. The fundamental building blocks from which the classical metric is constructed are certain functions that generate displacements in the parameter space. By promoting these classical functions to quantum operators, we also find alternative expressions for the quantum metric tensor and Berry’s connection.

The paper is organized as follows. In Sec. II we briefly review some basics about the quantum metric tensor. In Sec. III we define the notion of distance on the parameter space between points in phase space and derive the classical analog of the quantum metric tensor. In Sec. IV we compute and compare this classical metric and the quantum metric tensor for the considered systems. Section V presents alternative expressions for the quantum metric tensor and Berry’s connection. Finally, Sec. VI is
devoted to conclusions and directions for future research.

II. QUANTUM METRIC TENSOR

In this section, we shortly review the definition of the quantum metric tensor. We start by considering a quantum theory defined by a set of phase space operators $\hat{q} = \{q^a\}$ and $\hat{p} = \{p_a\}$ together with a Hamiltonian operator $\hat{H}(\hat{q}, \hat{p}; x)$, that depends of this set and also smoothly depends on a set of $N$ external parameters $x = \{x^i\}$, $i = 1, \ldots, N$). We consider that repeated indices $i, j, \ldots, 1, \ldots, N$ are summed from 1 to $N$, and $\delta_i := \partial/\partial x^i$.

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For the purposes of this paper, it is convenient to cast Eq. (2) in terms of operators. Let $\hat{P}_i$ be Hermitian operators and consider that $\hat{P}_i \delta x^i$ is the generator of the displacement $|n(x)\rangle \rightarrow |n(x')\rangle$. Thus, the translated state is

$$|n(x')\rangle = \exp\left(\frac{-i}{\hbar} \delta x^i \hat{P}_i\right) |n(x)\rangle.$$

From this equation, by considering a Taylor expansion, we have

$$i\hbar \delta |n(x)\rangle = \hat{P}_i |n(x)\rangle,$$

which substituted into Eq. (2) leads to

$$g_{ij}^{(n)}(x) = \frac{1}{\hbar} \text{Re} \left(\langle \hat{P}_j \hat{P}_j \rangle_n - \langle \hat{P}_i \rangle_n \langle \hat{P}_i \rangle_n \right),$$

where $\langle \hat{X} \rangle_n = \langle \hat{n} | \hat{X} | \hat{n} \rangle$ is the expectation value of $\hat{X}$ with respect to the state $|\hat{n}\rangle$. It should be noted that because of the Hermiticity of $\hat{P}_i$, the right-hand side (r.h.s) of Eq. (5) is symmetric. Furthermore, the line element $dl^2 = g_{ij}^{(n)} \delta x^i \delta x^j$ now reads

$$dl^2 = \frac{1}{\hbar^2} \langle \Delta \hat{P}^2 \rangle_n, \quad (\Delta \hat{P} = \Delta \hat{P}_i \delta x^i),$$

where $\Delta \hat{P}_i := \hat{P}_i - \langle \hat{P}_i \rangle_n$. Then, using operators, the distance $dl^2$ can be seen as the variance of the generator $\hat{P}_i \delta x^i$. This last remark will be the key point to obtain the classical counterpart of the quantum metric in the next section.

III. CLASSICAL ANALOG OF THE QUANTUM METRIC TENSOR

We now turn to the classical setting. Let us consider a classical integrable system with $n$ degrees of freedom described by the time-dependent Hamiltonian $H[q, p; x(t)]$, where $q = \{q^a\}$ and $p = \{p_a\}$ are the canonical coordinates and momenta, and $x = \{x^i\} \in \mathcal{M}$ is the set of slow time-dependent parameters.

Since the system is integrable (for all values of $x \in \mathcal{M}$), we can introduce the action-angle variables, $I = \{I_a\}$ and $\varphi = \{\varphi^a\}$, which satisfy Hamilton’s equations of motion with the new Hamiltonian

$$K(\varphi, I; x) = H(I; x) - G_i(\varphi, I; x) \dot{x}^i,$$

where $H(I; x) \equiv H[q(\varphi, I; x), p(\varphi, I; x); x]$ depends only on the action variables and the parameters, and

$$G_i(\varphi, I; x) := \frac{\partial}{\partial I^i} K(\varphi, I; x),$$

where $S^i(\varphi, I; x)$ is the generating function of the canonical transformation $(q, p) \rightarrow (\varphi, I)$. Also, $\dot{x}^i := dx^i/dt$ and $\alpha$ label different branches of the multivalued function $S^i(\varphi, I; x)$.

We recall that the second term in the r.h.s of Eq. (7) comes from $\partial S^i(\varphi, I; x)/\partial I = (\partial S^i(\varphi, I; x))/(\partial \varphi^a)/(\partial x^i)$, which is a consequence of the fact that $H[q, p, x(t)]$ (and hence $S^i(\varphi, I; x(t))$) also depends explicitly on time through the parameters $x$. The explicit form of $G_i$ in terms of the action-angle variables is

$$G_i(\varphi, I; x) = p_a(\partial I^a)_{\varphi^a, I} - (\partial I^a)_{\varphi, I},$$

where $p_a = p_a(\varphi, I; x)$, $q^a = q^a(\varphi, I; x)$ and we defined the single-valued function $S(\varphi, I; x) := S^i(\varphi, I; x), I; x]$ with $0 \leq \varphi < 2\pi$. We use the notation that repeated indices $a, b, \ldots$, are summed from 1 to $n$.

As our first step towards the classical counterpart of Eq. (5), we find that under the action of an infinitesimal displacement of the parameters $x \rightarrow x' = x + \delta x$ in $\mathcal{M}$, the function $G_i \delta x^i$ is the generator of the infinitesimal canonical transformation

$$[q(x), p(x)] \rightarrow [q(x) + \delta q, p(x) + \delta p],$$

where $\delta q = q^a(\varphi, I; x)$ and $\delta p = p_a(\varphi, I; x)$.
where
\[ \tilde{\delta}q^a := q^a(x') - q^a(x) = (\partial_t q^a)_{\varphi_t} \delta x^t, \]  
(11a)
\[ \tilde{\delta}p_a := p_a(x') - p_a(x) = (\partial_t p_a)_{\varphi_t} \delta x^t. \]  
(11b)

Notice that another form of Eq. (11a) is \( \tilde{\delta}q^a = \delta q^a - \tilde{\delta}q^a \), where \( \delta q^a := q^a(x') - q^a(x) \) is the total variation of \( x \)-fixed parameters and \( \tilde{\delta}q^a := q^a(x) - q^a(x) \) is the variation with “frozen” parameters. A similar expression follows for \( \tilde{\delta}p_a \).

To prove the above statement it is sufficient to show that \( G_i \) satisfy
\[ (\partial_t q^a)_{\varphi_t} \delta q^a = \{ q^a, G_i \}_{q,p} = \frac{\partial G_i}{\partial p_a}, \]  
(12a)
\[ (\partial_t p_a)_{\varphi_t} \delta p_a = \{ p_a, G_i \}_{q,p} = -\frac{\partial G_i}{\partial q^a}, \]  
(12b)
which are the equations of the infinitesimal canonical transformation (10) [29]. Here \( \{ , \} \) denotes the Poisson bracket. To do this, we first take the partial derivative with respect to \( x^t \), holding \( \varphi, I \) fixed, of the familiar relation \( p_a \tilde{\delta}q^a - I_a \tilde{\delta}q^a = \tilde{d}F \), where \( F = S^{\alpha \beta}(q, I; x) - I_a \varphi^a \) and \( \tilde{d} \) is the fixed-time differential (or equivalently with fixed parameters). From this we obtain
\[ (\partial_t p_a)_{\varphi_t} \delta p_a + p_a \tilde{d}(\partial_t q^a)_{\varphi_t} = \tilde{d}(\partial_t S)_{\varphi_t}. \]  
(13)
where we used \( (\partial_t \tilde{d})_{\varphi_t} = \tilde{d}(\partial_t I)_{\varphi_t} \). Next, combining Eq. (13) with the differential of Eq. (9) at fixed \( x \), namely
\[ \tilde{d}G_i = \tilde{d}p_a (\partial_t q^a)_{\varphi_t} + p_a \tilde{d}(\partial_t q^a)_{\varphi_t} - \tilde{d}(\partial_t S)_{\varphi_t}. \]  
(14)
we have
\[ dG_i = - (\partial_t p_a)_{\varphi_t} \delta q^a + (\partial_t q^a)_{\varphi_t} \delta p_a + \tilde{d}G_i. \]  
(15)
Then, taking \( G_i \) as a function of \( (q, p) \), it follows that
\[ \tilde{d}G_i = \frac{\partial G_i}{\partial q^a} \delta q^a + \frac{\partial G_i}{\partial p_a} \delta p_a. \]  
(16)
Equating the coefficients of \( \delta q^a \) and \( \delta p_a \) on the r.h.s of Eqs. (15) and (16), we read off Eqs. (12a) and (12b), which completes the proof.

Given the fact that \( G_i \delta x^i \) generates an infinitesimal displacement in \( x \) of points in phase space, and in complete analogy with the quantum case [see Eq. (6)], we can naturally define the distance between the points \( [q(x), p(x)] \) and \( [q(x) + \delta q, p(x) + \delta p] \) as
\[ ds^2 := \langle \Delta G^2 \rangle \quad \Delta G = G_i \delta x^i, \]  
(17)
where \( \Delta G_i := G_i - \langle G_i \rangle \) and
\[ \langle f(\varphi, I; x) \rangle = \frac{1}{(2\pi)^n} \int d\varphi f(\varphi, I; x), \]  
(18)
with \( \int d\varphi = \prod_{a=1}^n \int_{2\pi} d\varphi^a \), is the average of \( f(\varphi, I; x) \) over the (fast) angle variables. Defined in this way, the classical distance \( ds^2 \) is nothing more than the variance of the generator \( G_i \delta x^i \). Clearly, if the parameters \( x \) are frozen, then \( G_i \delta x^i = 0 \), and hence \( ds^2 \) also vanishes, as expected.

Notice that \( ds^2 \) depends only on the action variables \( I \) and the parameters \( x \). In this regard, it is important to emphasize that, according to the classical adiabatic theorem [30], while the parameters vary slowly with time, the action variables are adiabatic invariants \( I_a \approx 0 \). That is, during the adiabatic evolution from \( [q(x), p(x)] \) to \( [q(x) + \delta q, p(x) + \delta p] \) the action variables \( I \) remain constant. This effect is similar to the quantum case where the quantum number \( n \) remains constant as the parameters vary. On the other hand, note also that in this scenario, the average \( \langle \cdot \rangle \) in Eq. (17) is the classical counterpart of the quantum counterpart of the quantum average \( \langle \cdot \rangle_n \) in Eq. (6).

By expanding Eq. (17), we find that the distance \( ds^2 = g_{ij} \delta x^i \delta x^j \) induces the metric
\[ g_{ij}(I; x) := \langle G_i G_j \rangle - \langle G_i \rangle \langle G_j \rangle, \]  
(19)
where \( G_i = G_i(\varphi, I; x) \) is given by Eq. (9). The metric \( g_{ij}(I; x) \) corresponds to the classical analog of the quantum metric tensor (2) [or Eq. (5)], and provides a measure of the distance between the nearby points \( [q(x), p(x)] \) and \( [q(x) + \delta q, p(x) + \delta p] \) on the parameter manifold \( \mathcal{M} \).

It should be pointed out that, in contrast to the quantum metric tensor, the classical metric (19) is restricted to the case where classical motion is integrable. This restriction is to be expected since it is the same as that found in Hannay’s angle [17], which also involves the action variables and is the classical counterpart of Berry’s phase [22].

We now proceed to check some properties of \( g_{ij}(I; x) \).

Let us first show that, under a coordinate transformation, \( g_{ij}(I; x) \) transforms as a tensor. By considering a coordinate change \( y = y(x) \) and using Eq. (9), it follows that the transformation law for \( G_i \)
\[ G_i'(\varphi, I; y) = \frac{\partial x^j}{\partial y^i} G_j[\varphi, I; x(y)]. \]  
(20)
This result, together with Eq. (19), leads to the expected transformation law for the metric
\[ g'_{ij}(I; y) = \frac{\partial x^k}{\partial y^i} \frac{\partial x^l}{\partial y^j} g_{kl}[I; x(y)]. \]  
(21)

We now prove that \( g_{ij}(I; x) \) is positive semidefinite. This is straightforward and follows from the fact that \( ds^2 = \langle \Delta G^2 \rangle \geq 0 \) since the variance is nonnegative. In this light, it is interesting to note that the quantum metric tensor (2) is also positive semidefinite [31, 32].

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1 Nevertheless, for Hamiltonian systems with \( n \geq 2 \) there may exist conditions for which the adiabatic approximation is not optimal [30].
Analogously as the quantum metric \( g_{ij}^{(n)}(x) \) is independent of the gauge transformation\(^2\) \( |n'(x)\rangle = \exp[i\alpha_n(x)]|n(x)\rangle \) where \( \alpha_n(x) \) is an arbitrary real function of \( x \), the classical metric \( g_{ij}(I; x) \) is invariant under the (gauge) canonical transformation

\[
\varphi^a = \varphi^a + \frac{\partial \lambda(I; x)}{\partial I_a}, \quad I'_a = I_a,
\]

which is generated by the function \( F_2 = \varphi^a I'_a + \lambda(I'; x) \) where \( \lambda(I', x) \) is an arbitrary function of \( I' \) and \( x \). The proof of this statement is as follows. The Hamiltonian \( \varphi \) generates a canonical transformation of the same type (in the last line we used the fact that \( G \) from \( \sum_n = \langle \partial \lambda(I; x) \rangle_I - \langle \partial \lambda(I; x) \rangle_I + [\partial \lambda(I'; x)]_I [\partial \lambda(I; x)]_I \).

\[
K'(\varphi', I'; x) = H(I'; x) - G'_i(\varphi', I'; x)\dot{x}^i, \tag{23}
\]

where \( H(I'; x) = H(I; x) \) with \( I' = I \), and

\[
G'_i(\varphi', I'; x) = G_i(\varphi', I'; x) - \{\partial \lambda(I; x)\}_I, \tag{24}
\]

where \( G_i(\varphi', I'; x) := G_i[\varphi(\varphi', I'; x), I'; x] \) are the functions \( G_i \) for \( \varphi, I \) expressed in terms of the variables \( (\varphi', I') \). Since \( G'_i(\varphi', I'; x) \) satisfies Eqs. (12a) and (12b) with \( (\varphi', I') \) instead of \( (\varphi, I) \), it follows that \( G'_i \delta x^i \) generates a canonical transformation of the same type as Eq. (10) with \( \delta q^a = (\partial \varphi^a)_{\varphi', I} \delta x^i \) and \( \delta p_a = (\partial p_a)_{\varphi, I} \delta x^i \) instead of Eqs. (11a) and (11b), respectively. With this in mind, we can apply Eq. (19), and write the classical metric associated with the variables \( (\varphi', I') \) as

\[
g'_{ij}(I'; x) = \langle G'_i G'_j \rangle' - \langle G'_i \rangle' \langle G'_j \rangle', \tag{25}
\]

where \( G'_i = G'_i(\varphi', I'; x) \) and \( \langle \cdot \rangle' \) stands for the average over the angle variables \( \varphi' \).

By using Eq. (24), the average \( \langle G'_i \rangle' \) gives

\[
\langle G'_i(\varphi', I'; x) \rangle' = \frac{1}{(2\pi)^n} \int d\varphi^1 \ldots d\varphi^n G_i(\varphi', I'; x) - \partial \lambda(I; x)]_I, \\
= \frac{1}{(2\pi)^n} \int_{(a)} d\varphi^1 \ldots d\varphi^n G_i(\varphi', I; x) \\
- [\partial \lambda(I; x)]_I, \\
= \langle G_i(\varphi, I; x) \rangle - [\partial \lambda(I; x)]_I, \tag{26}
\]

where in the second line we made the change of variables from \( \varphi' \) to \( \varphi \) and defined \( b_a := \partial \lambda(I; x)/\partial I_a \), whereas in the last line we used the fact that \( p, (\partial \varphi^a)_{\varphi'; I} \), and \( (\partial \lambda)_{\varphi, I} \) are periodic functions of each angle variable \( \varphi^a \) with period \( 2\pi \), which by virtue of Eq. (9) implies that \( G_i(\varphi, I; x) \) are also periodic functions of each \( \varphi^a \). The periodicity of \( (\partial \lambda)_{\varphi, I} \) is easily seen by writing \( S(q, I; x) = \sum_{n=1}^n S_n(q_n, I; x) \) and recalling that each \( S_n(\varphi, I; x) \)

\[ 2 \text{ In Ref. [33] is shown, however, that under a more general gauge transformation, the quantum metric tensor depends on the gauge.} \]

\[ 3 \text{ In the literature, however, it is often found that the term } \langle (\partial \lambda)_{\varphi', I} \rangle \text{ is dropped from Eq. (29) since it does not contribute to Hannay’s angle [91, 34].} \]
Upon using Eqs. (12a) and (12b), we find that these components take the form

\[ F_{ij}(I; x) = -\langle \{G_i, G_j\}_q, p \rangle. \tag{32} \]

In this way the functions \( G_i \) can be regarded as the fundamental building blocks that underlie the classical metric (19) and Hannay’s curvature (32).

IV. ILLUSTRATIVE EXAMPLES

In this section, we set out some examples of classical integrable systems to illustrate the appearance of the metric (19). At the same time, we compare the results of this classical metric with those found by using the quantum metric counterpart of each system. We shall see that these results corroborate that the metric (19) is the classical analog of (2) [or equivalently Eq. (5)].

A. Generalized harmonic oscillator

As our first example, let us take the generalized harmonic oscillator, whose classical Hamiltonian is given by

\[ H = \frac{1}{2} \left( X q^2 + 2Y qp + Z p^2 \right), \tag{33} \]

where \( x = \{x^i\} = (X, Y, Z) \in \mathbb{R}^3 \) \((i, j, \ldots = 1, 2, 3)\) are the adiabatic parameters, which are assumed to satisfy \( XZ - Y^2 > 0 \). The transformation from the variables \((q, p)\) to the action-angle variables \((\varphi, I)\) is well known and turns out to be

\[ q(\varphi, I; x) = \left( \frac{2Z I}{\omega} \right)^{1/2} \sin \varphi, \tag{34a} \]

\[ p(\varphi, I; x) = \left( \frac{2Z I}{\omega} \right)^{1/2} \left( -\frac{Y}{Z} \sin \varphi + \frac{\omega}{Z} \cos \varphi \right), \tag{34b} \]

where \( \omega := (XZ - Y^2)^{1/2} \) is the parameter-dependent angular frequency. Moreover, the generating function of this transformation, in terms of action-angle variables, is

\[ S(\varphi, I; x) = -\frac{Y I}{\omega} \sin^2 \varphi + I(\varphi + \sin \varphi \cos \varphi). \tag{35} \]

Then, putting Eqs. (34a), (34b), and (35) into Eq. (9), we obtain the functions \( G_i(\varphi, I; x) \):

\[ G_1(\varphi, I; x) = \frac{Z I}{2\omega I^2} \sin \varphi \cos \varphi, \tag{36a} \]

\[ G_2(\varphi, I; x) = \frac{I \sin \varphi}{\omega^2} (Y \cos \varphi + \omega \sin \varphi), \tag{36b} \]

\[ G_3(\varphi, I; x) = \frac{I \sin \varphi}{2Z \omega^2} \left( (XZ - 2Y^2) \cos \varphi - 2Y \omega \sin \varphi \right). \tag{36c} \]

Rewriting these functions in terms of the variables \((q, p)\), they satisfy Eqs. (12a) and (12b). With this at hand, Eq. (19) can now be readily applied to Eqs. (36a), (36b), and (36c). This yields the components of the corresponding classical metric \( g_{ij}(I; x) \), which can be expressed as

\[ g_{ij}(I; x) = \frac{I^2}{32\omega^4} \begin{pmatrix} Z^2 & -2YZ & 2Y^2 - XZ \\ -2YZ & 4XZ & -2XY \\ 2Y^2 - XZ & -2XY & X^2 \end{pmatrix}. \tag{37} \]

The idea is now to compare this metric with that coming from the quantum metric tensor (2). In the quantum case, the time-dependent Hamiltonian operator \( \hat{H} \) of the system is

\[ \hat{H} = \frac{1}{2} \left[ X \hat{q}^2 + Y(\hat{q}\hat{p} + \hat{p}\hat{q}) + Z\hat{p}^2 \right], \tag{38} \]

and leads to the Schrödinger equation (with fixed parameters)

\[ -\frac{Z\hbar^2}{2} \frac{d^2\psi_n}{dq^2} - iYq \frac{d\psi_n}{dq} + \left( \frac{Xq^2}{2} - iY \frac{q^2}{2} \right) \psi_n = E_n \psi_n, \tag{39} \]

which has the normalized solution

\[ \psi_n(q; x) = \left( \frac{\omega}{Z\hbar} \right)^{1/4} \chi_n \left( q \sqrt{\frac{\omega}{Z\hbar}} \right) \exp \left( -iYq^2 \frac{\sqrt{\omega}}{2\hbar} \right). \tag{40} \]

where \( \omega := (XZ - Y^2)^{1/2} \) which implies \( XZ - Y^2 > 0 \), and \( \chi_n(\xi) = (2^n n! \sqrt{\pi})^{-1/2} e^{-\xi^2/2} H_n(\xi) \) are the Hermite functions, with \( H_n(\xi) = (-1)^n e^{\xi^2} \frac{d^n}{d\xi^n} e^{-\xi^2} \) being the Hermite polynomials. Furthermore, the energy eigenvalues are given by \( E_n = (n + 1/2)\hbar\omega \) where \( n \) are nonnegative integers.

Substituting the wave function (40) into \( \langle n|\hat{\partial}_n|n \rangle = \int_{-\infty}^{\infty} dq \psi_n^*(q; x) \hat{\partial}_n \psi_n(q; x) \) and \( \langle \hat{\partial}_n|\hat{\partial}_n|n \rangle = \int_{-\infty}^{\infty} dq d\xi \psi_n^*(q; x) \hat{\partial}_n \psi_n(q, \xi) \), and bearing in mind the following properties of the Hermite functions

\[ \int_{-\infty}^{\infty} d\xi \chi_m(\xi) \chi_n(\xi) = \delta_{mn}, \]

\[ \frac{d}{d\xi} \chi_n = \sqrt{\frac{n}{2}} \chi_{n-1} - \sqrt{\frac{n + 1}{2}} \chi_{n+1}, \]

\[ \chi_n = \sqrt{\frac{n}{2}} \chi_{n-1} + \sqrt{\frac{n + 1}{2}} \chi_{n+1}, \]

the components of the quantum metric (2) become

\[ g^{(n)}_{ij}(x) = \frac{n^2 + n + 1}{32\omega^4} \begin{pmatrix} Z^2 & -2YZ & 2Y^2 - XZ \\ -2YZ & 4XZ & -2XY \\ 2Y^2 - XZ & -2XY & X^2 \end{pmatrix}. \tag{42} \]

Comparing the metrics (37) and (42), it is clear that they are related as follows:

\[ g_{ij}^{(n)}(x) = \gamma g_{ij}(I; x), \tag{43} \]

where

\[ \gamma := \frac{n^2 + n + 1}{I^2}. \tag{44} \]
Therefore, for the generalized harmonic oscillator, the quantum metric tensor \( g_{ij}^{(n)}(x) \) can be determined from the classical metric \( g_{ij}(I;x) \), modulo the parameter-independent constant factor \( \gamma \). This result is nontrivial and supports our claim that the metric (19) is the classical counterpart of the metric (2). Note that if we take into account the Bohr-Sommerfeld quantization rule for action variable

\[
I = \left( n + \frac{1}{2} \right) \hbar,
\]

(45)

then \( \gamma \) turns out to be proportional to \( 1/\hbar^2 \). Additionally, by using Eq. (45), the metrics (37) and (42) can also be related by

\[
\frac{\partial}{\partial n} g_{ij}^{(n)}(x) = \frac{1}{\hbar} \frac{\partial}{\partial I} g_{ij}(I;x).
\]

(46)

On another hand, it is worth pointing out that the determinants of the metrics (37) and (42) are zero, which indicates that the corresponding Hamiltonians (33) and (38) involve more parameters than the effective ones. Actually, these metrics have rank two and hence, to have metrics with nonvanishing determinants, we must leave one of the parameters fixed (but different from zero). In this case, the degeneration in the metric only indicates that one of the parameters is redundant and can be set equal to a constant. To show this explicitly, let us consider for a moment that \( \{g_{ij}(y') = (X, Y)' (i', j' = 1, 2) \) are the adiabatic parameters and suppose that \( Z = Z_0 \) is a nonvanishing constant. In this case, the classical metric (19) becomes

\[
g_{ij}(I; y) = \frac{Z_0 I^2}{32 \omega^4} \begin{pmatrix} Z_0 & -2Y \\ -2Y & 4X \end{pmatrix},
\]

(47)

and its determinant, \( \text{det} [g_{ij}(I; y)] = \frac{Z_0^2 I^2}{64 \omega^4} \), is not zero. Besides, the corresponding quantum metric tensor \( g_{ij}^{(n)}(y) \), which can be obtained from Eq. (47) by replacing \( I^2 \) with \( n^2 + n + 1 \), also has a nonvanishing determinant.

Before concluding this example, it may be interesting to obtain the components of the connection and curvature associated with Hanany’s angle through Eqs. (29) and (32), respectively. Using the functions \( G_i \) given by Eqs. (36a), (36b), and (36c), the components of connection (29) lead to

\[
A_1(I;x) = 0, \quad A_2(I;x) = \frac{I}{2\omega}, \quad A_3(I;x) = -\frac{Y I}{2Z \omega},
\]

(48)

whereas the components of the curvature (32) give

\[
F_{12}(I;x) = -\frac{Z I}{4\omega^3}, \quad F_{13}(I;x) = \frac{Y I}{4\omega^3}, \quad F_{23}(I;x) = -\frac{X I}{4\omega^3}.
\]

(49)

It is also instructive to compare Eqs. (48) and (49) with their quantum analogues, namely Berry’s connection and its curvature, respectively. Then, using the wave function (40) and Eq. (41), the components \( A_1^{(n)}(x) := -\text{Im}(\langle n | \partial_n I \rangle) \) of Berry’s connection become

\[
A_1^{(n)}(x) = 0, \quad A_2^{(n)}(x) = \frac{c_n}{2\omega}, \quad A_3^{(n)}(x) = -\frac{c_n Y}{2Z \omega},
\]

(50)

and the components \( F_{ij}^{(n)}(x) := \partial_i A_j^{(n)} - \partial_j A_i^{(n)} \) of its curvature yield

\[
F_{12}^{(n)}(x) = -\frac{c_n Z}{4\omega^3}, \quad F_{13}^{(n)}(x) = \frac{c_n Y}{4\omega^3}, \quad F_{23}^{(n)}(x) = \frac{c_n X}{4\omega^3},
\]

(51)

where \( c_n := (n + 1/2) \). Comparing Eqs. (48) and (50) as well as Eqs. (49) and (51), it is straightforward to see the following relations:

\[
A_1^{(n)}(x) = \beta A_1(I;x), \quad F_{ij}^{(n)}(x) = \beta F_{ij}(I;x),
\]

(52a, b)

where

\[
\beta := \frac{n + \frac{1}{2}}{I}.
\]

(53)

This entails that \( A_1^{(n)}(x) \) and \( F_{ij}^{(n)}(x) \) can be obtained, respectively, from \( A_1(I;x) \) and \( F_{ij}(I;x) \), modulo the parameter-independent constant factor \( \beta \). Moreover, after using Eq. (45), this factor reduces to \( \beta = 1/\hbar \).

Some comments are in order. First, it is noteworthy to emphasize that connection defined by dropping \( \langle (\partial S)_{\chi, \ell} \rangle \) from Eq. (29), namely \( A_i(I;x) = \langle p_\chi (\partial_i q^{(n)})_{\chi, \ell} \rangle \), does not lead to Eqs. (48) and therefore does not satisfy the relation (52a). Of course, the curvature of such a connection, which is also given by Eq. (31), implies Eq. (49). Second, notice that Eq. (52b) is in complete agreement with the semiclassical relation between Berry’s curvature and the curvature associated with Hanany’s angle reported in Ref. [22]. Finally, it is worth mentioning that the multiplicative constants involved in the relation (43) and the relations (52a) and (52b) are different: while \( \gamma \) is proportional to \( 1/I^2 \), \( \beta \) is proportional to \( 1/I \).

B. Generalized harmonic oscillator with a linear term

For our second example we shall consider the generalized harmonic oscillator with a linear term in the position. Thus the Hamiltonian under consideration is

\[
H = \frac{1}{2} (X q^2 + 2Y qp + Z p^2) + W q,
\]

(54)

where \( x = \{x^i\} = (W, X, Y, Z) \) with \( i, j, \ldots = 0, 1, 2, 3 \) are the adiabatic parameters. Assuming \( XZ - Y^2 > 0 \),
we find that the variables \((q, p)\) in terms of action-angle variables \((\varphi, I)\) read

\[
q(\varphi, I; x) = \left(\frac{2ZI}{\omega}\right)^{1/2} \sin \varphi - \frac{WZ}{\omega^2}, \tag{55a}
\]

\[
p(\varphi, I; x) = \left(\frac{2ZI}{\omega}\right)^{1/2} \left(-\frac{Y}{Z} \sin \varphi + \frac{\omega}{Z} \cos \varphi\right) + \frac{WY}{\omega^2}, \tag{55b}
\]

where \(\omega := (XZ - Y^2)^{1/2}\) is the angular frequency of the system, which is independent of \(W\). Furthermore, we get that the generating function \(S(\varphi, I; x)\) of the transformation \((q, p) \to (\varphi, I)\) is

\[
S(\varphi, I; x) = -\frac{Y}{2Z} \left[\left(\frac{2ZI}{\omega}\right)^{1/2} \sin \varphi - \frac{WZ}{\omega^2}\right]^2 
+ I(\varphi + \sin \varphi \cos \varphi). \tag{56}
\]

With these ingredients at hand, it is straightforward to obtain \(G_i(\varphi, I; x)\) from Eq. (9). The resulting functions, in compact form, are

\[
G_i(\varphi, I; x) = f_i(x)q + g_i(x)q^2 + h_i(x) \left(p + \frac{Y}{Z}q\right), \tag{57}
\]

where \(p = p(\varphi, I; x)\) and \(q = q(\varphi, I; x)\) are given by Eqs. (55a) and (55b), respectively, while

\[
f_i(x) := \frac{\omega}{2Z} \frac{\partial_i}{\partial} \left(\frac{Z}{\omega}\right), \tag{58a}
\]

\[
g_i(x) := \frac{Y}{Z} f_i(x) + \frac{1}{2} \frac{\partial_i}{\partial} \left(\frac{Y}{Z}\right), \tag{58b}
\]

\[
h_i(x) := \frac{W}{2\omega} \frac{\partial_i}{\partial} \left(\frac{Z}{\omega^2}\right) - \frac{\omega}{2} \frac{\partial_i}{\partial} \left(\frac{WZ}{\omega^2}\right). \tag{58c}
\]

It can be verified that \(G_i\) given by Eq. (57) satisfy Eqs. (12a) and (12b). In addition, if the parameter \(W\) is fixed to 0, it is not difficult to realize that these functions reduce to those of Eqs. (36a), (36b), and (36c).

By inserting Eq. (57) into Eq. (19), the corresponding components of the classical metric \(g_{ij}(I; x)\) are

\[
g_{ij}(I; x) = \frac{I^2}{32\omega^4} \begin{pmatrix} 0 & 0 & -2YZ & 2Y^2 - XZ \\
0 & Z^2 & 4XZ & -2XY \\
0 & -2YZ & 4XZ & -2XY \\
0 & 2Y^2 - XZ & -2XY & X^2
\end{pmatrix}
+ \frac{I}{\omega^4} \begin{pmatrix} Z\omega^4 & -WZ^2\omega^2 & 2WYZ\omega^2 & -WY^2\omega^2 \\
-WZ^2\omega^2 & W^2Z^2 & -2W^2ZY^2 & W^2Y^2Z^2 \\
2WYZ\omega^2 & -2W^2ZY^2 & W^2Z(3Y^2 + XZ) & -W^2Y(Y^2 + XZ) \\
-WY^2\omega^2 & W^2Y^2Z & -W^2Y(Y^2 + XZ) & W^2XY^2
\end{pmatrix}. \tag{59}
\]

Notice that this metric has an extra term, as compared to the metric (37), which is proportional to \(I/\omega^7\) and is a consequence of the linear modification introduced in the Hamiltonian (54). Certainly, by fixing \(W = 0\) in the above expression and eliminating the corresponding row and column, we can recover the metric (37).

To contrast Eq. (59) with the quantum metric tensor, we consider the following Hamiltonian operator:

\[
\hat{H} = \frac{1}{2} \left[ X\hat{q}^2 + Y(\hat{q}\hat{p} + \hat{p}\hat{q}) + Z\hat{p}^2 \right] + W\hat{q}. \tag{60}
\]

In this case the Schrödinger equation reads

\[
\frac{\hbar^2}{2} \frac{d^2\psi_n}{dq^2} - i\hbar Yq \frac{d\psi_n}{dq} + \left(\frac{Xq^2}{2} + Wq - i\hbar\frac{Y}{2}\right)\psi_n = E_n\psi_n, \tag{61}
\]

and the eigenfunctions \(\psi_n(q; x)\) are of the form

\[
\psi_n(q; x) \equiv \left(\frac{\omega}{Z\hbar}\right)^{1/4} \chi_n \left[\left(\frac{q + WZ}{\omega^2}\right)^{1/4} \frac{\sqrt{\omega}}{Z\hbar} \exp\left(-\frac{iYq^2}{2Z\hbar}\right)\right]. \tag{62}
\]

where once again \(\omega = (XZ - Y^2)^{1/2}\) which entails \(XZ - Y^2 > 0\). By substituting Eq. (62) into Eq. (2) and using Eq. (41), we get that the components of the quantum metric \(g_{ij}^{(n)}(x)\) are given by
We can see that the classical metric (59) and the quantum metric (63) have exactly the same functional dependence on the adiabatic parameters. Hence we corroborate once again that the metric (19) is the classical analog of the quantum metric tensor (2). Remarkably, by using the Bohr-Sommerfeld quantization rule (45), it follows that the relation (46) also holds for the metrics (59) and (63).

Note that in this example, as well as in the previous one, the metrics $g_{ij}(I; x)$ and $g_{ij}^{(n)}(x)$ have vanishing determinant. However, here the rank of the metrics (59) and (63) is three, which shows the existence of a redundant parameter. In particular, if we take \( \{y^i\} = \{W, X, Y\} \) \((i', j', \ldots = 0, 1, 2)\) as the adiabatic parameters and $Z = Z_0$ as a nonvanishing constant, then the classical metric reads

$$g_{ij}(I; y) = \frac{Z_0 I^2}{32 \omega^4} \begin{pmatrix} 0 & 0 & 0 \\ 0 & Z_0 & 2Y \\ 0 & -2Y & 4X \end{pmatrix} + \frac{Z_0 I}{\omega^2} \begin{pmatrix} \omega^4 & -W^2 Z_0 \omega^2 & 2WY \omega^2 \\ -W Z_0 \omega^2 & -W^2 Z^2 Z_0 & -2W^2 Y Z_0 \\ 2WY \omega^2 & -2W^2 Z_0 Y & W^2 (3Y^2 + XZ_0) \end{pmatrix} + \frac{n + \frac{1}{2}}{\hbar \omega^3} \begin{pmatrix} Z_0^4 & -W^2 Z_0 \omega^2 & 2WY Z_0^2 \\ -W Z_0 \omega^2 & W^2 \omega^2 Z_0 & -2W^2 Y Z_0 \\ 2WY Z_0 \omega^2 & -2W^2 Y Z_0 & X \end{pmatrix},$$

and its determinant \( \det[g_{ij}^{(n)}(I; y)] = \frac{Z_0^2 I^4}{32 \omega^4} (I \omega^3 + 8W^2 Z_0) \) is different from zero.

To conclude this example, let us obtain the corresponding classical and quantum connections and curvatures. Classically, by applying Eqs. (29) and (32) to the functions $G_{ij}$ given by Eq. (57), we obtain the components of the connection,

$$A_0(I; x) = A_1(I; x) = 0, \quad A_2(I; x) = \frac{I}{2 \omega} + \frac{W^2 Z}{2 \omega^4},$$

and the components of the curvature, which are displayed in matrix form,

$$F_{ij}(I; x) = \frac{I}{4 \omega^3} \begin{pmatrix} 0 & 0 & 0 \\ 0 & Z & 0 \\ 0 & 0 & -X \end{pmatrix} + \frac{1}{\omega^6} \begin{pmatrix} 0 & 0 & WZ \omega^2 & 0 \\ 0 & 0 & 0 & -W^2 Z^2 \\ -WZ \omega^2 & W^2 Z^2 & 0 & -W^2 Y \omega^2 \\ WY \omega^2 & -W^2 Z^2 Y & W^2 Y^2 & 0 \end{pmatrix},$$

respectively. Here, we used once again Eq. (41). By comparing Eqs. (65) and (67) as well as Eqs. (66) and (68), it turns out that the relations (52a) and (52b) hold provided that the Bohr-Sommerfeld quantization rule (45) is taken into account, i.e., when $\beta = 1/\hbar$ in Eq. (53).

C. Quartic anharmonic oscillator

In this example, we focus on the classical quartic anharmonic oscillator which is defined by the Hamiltonian

$$H = \frac{p^2}{2m} + \frac{k \ell^2}{2} + \frac{\lambda \ell^4}{4!},$$

where $x = \{x'\} = \{m, k, \lambda\}$ with $i = 1, 2, 3$ are the adiabatic parameters. In this case, in contrast to the previous
where we assume $0 \leq \lambda \ll 1$. Here, $H_0$ is the Hamiltonian of the unperturbed problem, for which the action-angle variables $(\varphi_0, I_0)$ are well known and allow us to express the variables $(q, p)$ as

$$q(\varphi_0, I_0; x) = \left( \frac{2I_0}{m\omega_0} \right)^{1/2} \sin \varphi_0,$$

$$p(\varphi_0, I_0; x) = (2m\omega_0I_0)^{1/2} \cos \varphi_0,$$

where $\omega_0 = (k/m)^{1/2}$ is the unperturbed frequency. Furthermore, $H_1$ is the perturbative potential.

Next we assume that the type 2 generating function $W(\varphi, I; x)$ of the canonical transformation from $(\varphi_0, I_0)$ to the action-angle variables $(\varphi, I)$ of the total problem $H(I; x)$ can be expanded in a power series of $\lambda$:

$$W(\varphi_0, I; x) = \varphi \text{I} + \lambda W_1(\varphi_0, I; x) + \lambda^2 W_2(\varphi_0, I; x) + \lambda^3 W_3(\varphi_0, I; x) + O(\lambda^4),$$

where $W_1, W_2, \ldots$ are functions to be determined. Thus, the equations of the canonical transformation, $\varphi(\varphi_0, I; x) = \partial W(\varphi_0, I; x)/\partial I$ and $I_0(\varphi_0, I; x) = \partial W(\varphi_0, I; x)/\partial \varphi$, take the form

$$\varphi(\varphi_0, I; x) = \varphi + \lambda \frac{\partial W_1(\varphi_0, I; x)}{\partial I} + \lambda^2 \frac{\partial W_2(\varphi_0, I; x)}{\partial I} + \lambda^3 \frac{\partial W_3(\varphi_0, I; x)}{\partial I} + O(\lambda^4),$$

and

$$I_0(\varphi_0, I; x) = I + \lambda \frac{\partial W_1(\varphi_0, I; x)}{\partial \varphi_0} + \lambda^2 \frac{\partial W_2(\varphi_0, I; x)}{\partial \varphi_0} + \lambda^3 \frac{\partial W_3(\varphi_0, I; x)}{\partial \varphi_0} + O(\lambda^4),$$

respectively.

Following the canonical perturbation theory and working up to the third order in $\lambda$, the functions $W_1, W_2, W_3$ can be obtained by solving the differential equations [36, 37]

$$\frac{\partial W_1(\varphi_0, I; x)}{\partial \varphi_0} = (\Phi_1(\varphi_0, I; x))_0 - \Phi_1(\varphi_0, I; x),$$

where $(\cdot)_0$ denotes the average with respect to $\varphi_0$ and, in our case, $\Phi_1 = H_1$, $\Phi_2 = \frac{1}{2} \left( \frac{\partial W_1}{\partial \varphi_0} \right)^2 + \frac{\partial W_1}{\partial \varphi_0} \frac{\partial H_1}{\partial I} + \frac{\partial W_2}{\partial \varphi_0} \frac{\partial H_1}{\partial I}$ and $\Phi_3 = \frac{1}{2} \left( \frac{\partial W_1}{\partial \varphi_0} \right)^2 \frac{\partial^2 H_1}{\partial I^2} + \frac{\partial W_1}{\partial \varphi_0} \frac{\partial H_2}{\partial I}$. Explicitly these functions are given by

$$\Phi_1(\varphi_0, I; x) = \frac{I^2 \sin^4 \varphi_0}{6m^2 \omega_0^2},$$

$$\Phi_2(\varphi_0, I; x) = \frac{I^3 \sin^4 \varphi_0}{144m^4 \omega_0^4} (8 \sin^4 \varphi_0 - 3),$$

$$\Phi_3(\varphi_0, I; x) = \frac{I^4 \sin^4 \varphi_0}{13824m^6 \omega_0^6} (320 \sin^8 \varphi_0 - 144 \sin^4 \varphi_0 - 25),$$

and together with Eq. (76) they imply

$$W_1(\varphi, I; x) = \frac{I^2}{192m^2 \omega_0^2} (8 \sin 2 \varphi_0 - 8 \sin 4 \varphi_0),$$

$$W_2(\varphi, I; x) = \frac{I^3}{55296m^4 \omega_0^6} (-384 \sin 2 \varphi_0 + 132 \sin 4 \varphi_0 - 32 \sin 6 \varphi_0 + 3 \sin 8 \varphi_0),$$

$$W_3(\varphi, I; x) = \frac{I^4}{530416m^6 \omega_0^8} (9264 \sin 2 \varphi_0 - 4101 \sin 4 \varphi_0 + 1624 \sin 6 \varphi_0 - 441 \sin 8 \varphi_0 + 72 \sin 10 \varphi_0 - 5 \sin 12 \varphi_0).$$

Having found $W(\varphi, I; x)$, it is straightforward to obtain the generating function $S(\varphi, I; x)$ of the canonical transformation $(q, p) \to (\varphi, I)$. Indeed, since the transformation $(q, p) \to (\varphi, I)$ can be regarded as the composition of the successive canonical transformations $(q, p) \to (\varphi_0, I_0)$ and $(\varphi_0, I_0) \to (\varphi, I)$, generated, respectively, by $S_0(q, I_0; x)$ and $W(\varphi, I; x)$, we have that $S(\varphi, I; x)$ is given by

$$S(q, I; x) = S_0(q, I_0; x) + W(\varphi, I; x) - \varphi_0 I_0,$$

where the function $S_0$ in terms of the variables $(\varphi_0, I_0)$ reads

$$S_0(\varphi_0, I_0; x) = I_0(\varphi_0 + \sin \varphi_0 \cos \varphi_0).$$

Note that substituting Eq. (75) into Eqs. (71), (72), and (79), we can write the variables $q$ and $p$ and the function $S$ in terms of $\varphi_0$ and $I$. Now, to compute the metric (19) we need to first obtain the functions $G_i(\varphi, I; x) = p(\partial_q q), p - (\partial_q S)_{\varphi, I}$, where the derivatives are taken at fixed action-angle variables $(\varphi, I)$. These derivatives can be achieved by employing the following useful formula:

$$(\partial_\varphi \mathcal{F})_{\varphi, I} = (\partial_\varphi \mathcal{F})_{\varphi_0, I_0} - \frac{(\partial_\varphi \mathcal{F})_{\varphi_0, I_0}}{(\partial_\varphi \mathcal{F})_{\varphi_0, I_0}} (\partial_\varphi \mathcal{F})_{\varphi_0, I_0},$$

where $\mathcal{F}(\varphi, I; x) = \mathcal{F}(\varphi_0, I_0; x)$ is $q(\varphi, I; x)$ or $S(\varphi_0, I; x)$. Notice that we can use Eq. (74) to compute $(\partial_\varphi \mathcal{F})_{\varphi_0, I_0}$ and $(\partial_\varphi \mathcal{F})_{\varphi_0, I_0}$. After carrying out these calculations and retaining terms correct to second order in $\lambda$ (since derivatives with respect to this parameter are involved), we arrive at the functions $G_i$ in terms of $\varphi_0$ and $I$, namely

$$G_i(\varphi, I; x) = \alpha_{i0} + \alpha_{i1} \lambda + \alpha_{i2} \lambda^2,$$
where in the case \( i = 1 \):

\[
\begin{align*}
\alpha_{10} &= -\frac{I \sin 2\phi_0}{4m}, \\
\alpha_{11} &= -\frac{I^2 \sin^2 \phi_0}{48\sqrt{k^3}m^3} (\cos 3\phi_0 - 2 \cos \phi_0), \\
\alpha_{12} &= -\frac{I^3}{55296k^3m^3} (318 \sin 2\phi_0 - 204 \sin 4\phi_0 + 95 \sin 6\phi_0 - 27 \sin 8\phi_0 + 3 \sin 10\phi_0),
\end{align*}
\]

in the case \( i = 2 \):

\[
\begin{align*}
\alpha_{20} &= -\frac{I \sin 2\phi_0}{4k}, \\
\alpha_{21} &= \frac{I^2}{384\sqrt{k^5}m} (23 \sin 2\phi_0 - 7 \sin 4\phi_0 + \sin 6\phi_0), \\
\alpha_{22} &= -\frac{I^3}{18432k^4m} (362 \sin 2\phi_0 - 156 \sin 4\phi_0 + 53 \sin 6\phi_0 - 11 \sin 8\phi_0 + \sin 10\phi_0),
\end{align*}
\]

and in the case \( i = 3 \):

\[
\begin{align*}
\alpha_{30} &= \frac{I^2}{192\sqrt{k^3}m} (\sin 4\phi_0 - 8 \sin 2\phi_0), \\
\alpha_{31} &= \frac{I^3}{27648k^2m} (384 \sin 2\phi_0 - 132 \sin 4\phi_0 + 32 \sin 6\phi_0 - 3 \sin 8\phi_0), \\
\alpha_{32} &= \frac{I^4}{1769472\sqrt{k^5}m^2} (-9264 \sin 2\phi_0 + 4101 \sin 4\phi_0 - 1624 \sin 6\phi_0 + 441 \sin 8\phi_0 - 72 \sin 10\phi_0 + 5 \sin 12\phi_0).
\end{align*}
\]

Finally, substituting \( G_i(\phi, I; x) \) into Eq. (19) and writing the average over the angle variable \( \varphi \) as

\[
\langle f(\varphi) \rangle = \frac{1}{2\pi} \int_0^{2\pi} d\varphi f(\varphi) = \frac{1}{2\pi} \int_0^{2\pi} d\varphi_0 \left( \frac{\partial f}{\partial \varphi} \right)_{I,x},
\]

we obtain the components of the classical metric \( g_{ij}(I; x) \) correct to second order in \( \lambda \):

\[
\begin{align*}
g_{11}(I; x) &= \frac{I^2}{32m^2} - \frac{\lambda I^3}{256m^2k^3} + \frac{47\lambda^2 I^4}{32768m^4k^3}, \\
g_{12}(I; x) &= \frac{I^2}{32mk} - \frac{7\lambda I^3}{768m^3k^3} + \frac{347\lambda^2 I^4}{98304m^4k^3}, \\
g_{13}(I; x) &= \frac{I^3}{192\sqrt{m^3k^3}} - \frac{103\lambda I^4}{49152m^2k^3} + \frac{15\lambda^2 I^5}{16384m^3k^3}, \\
g_{22}(I; x) &= \frac{I^2}{32k^2} - \frac{\lambda I^3}{768\sqrt{mk}} + \frac{1919\lambda^2 I^4}{294912mk^5}, \\
g_{23}(I; x) &= \frac{I^3}{192\sqrt{mk^3}} - \frac{439\lambda I^4}{147456mk^4} + \frac{7\lambda^2 I^5}{4608m^{1/2}k^{11}}, \\
g_{33}(I; x) &= \frac{6I^4}{13728mk^3} - \frac{89\lambda I^5}{147456m^{3/2}k^{11}} + \frac{130621\lambda^2 I^6}{382205952mk^{15}}.
\end{align*}
\]

Note that by multiplying the components \( g_{ij}^{(0)} \) in Eq. (87) by \( \hbar^2 \) and comparing the result with the corresponding components \( g_{ij}(I; x) \) in Eq. (86), we have that terms with the same powers of \( \hbar \) and \( I \) have exactly the same functional dependence on the parameters \( m, k, \lambda \). Then, to match Eqs. (86) and (87), it is reasonable to consider \( g_{ij}^{(0)} = \frac{1}{\hbar^2} g_{ij}(I; x) \). By doing this, we find the following identifications for the ground state: \( I^2 = \hbar^2 \), \( I^3 = \frac{2}{3} \hbar^3 \), and \( I^5 = \frac{1200}{139621} \hbar^5 \). In the case of \( I^4 \) and \( I^5 \), the identifications are not unique, but they differ from each other slightly. Actually, for \( I^4 \) we find \( I^4 \approx 2.48 \hbar^4, 2.43 \hbar^4, 2.44 \hbar^4, 2.45 \hbar^4, 2.47 \hbar^4, 2.51 \hbar^4 \), whereas for \( I^5 \) we find \( I^5 \approx 4.18 \hbar^5, 4.29 \hbar^5, 4.35 \hbar^5 \). Therefore, \( g_{ij}^{(0)} \) can only be obtained in an approximate way from \( g_{ij}(I; x) \) through the relation \( g_{ij}^{(0)} \approx \frac{1}{\hbar^2} g_{ij}(I; x) \) with the appropriate identifications. This result is somehow expected because we have been dealing with perturbation theories to arrive at Eqs. (86) and (87). Finally, it is worth mentioning that the determinants of the metrics defined by the components (86) and (87), obtained by keeping terms up to second order in \( \lambda \), are zero. Nonetheless, if one of the parameters is left fixed, then the corresponding (classical or quantum) metric could have a nonvanishing determinant.

V. ALTERNATIVE EXPRESSIONS FOR THE QUANTUM METRIC TENSOR AND BERRY’S CONNECTION

We now extend the use of the classical functions \( G_i \) given by Eq. (8) to the quantum case. To this end we start by promoting the functions \( G_i(q; p; x) \), expressed in terms of the variables \( (q, p) \), to quantum operators \( \hat{G}_i(\hat{q}; \hat{p}; x) \) which we assume to be Hermitian. By analogy with the classical case where \( G_i \delta x^i \) generates a displacement in the parameter space, it is reasonable to
consider $\delta x_i \hat{G}_i(q, p; x)$ as the generator of the infinitesimal displacement of states $|n(x)\rangle \to |n(x')\rangle$, namely

$$|n(x')\rangle = \exp \left(-\frac{i}{\hbar} \delta x_i \hat{G}_i \right) |n(x)\rangle.$$ 

Then, we can replace the operators $\hat{P}_i$ in Eq. (4) by the operators $\hat{G}_i$, obtaining

$$i\hbar |\partial_t n(x)\rangle = \hat{G}_i |n(x)\rangle.$$ \hspace{1cm} (88)

This allows us to write down the quantum metric tensor (2) or Eq. (5) in terms of $\hat{G}_i(q, p; x)$ as

$$g_{ij}^{(n)}(x) = \frac{1}{\hbar^2} \text{Re} \left( \langle \hat{G}_i \hat{G}_j \rangle_n - \langle \hat{G}_i \rangle_n \langle \hat{G}_j \rangle_n \right).$$ \hspace{1cm} (89)

Note that on account of the Hermiticity of $\hat{G}_i$, the r.h.s. of this expression is symmetric.

Similarly, we can recast Berry’s connection in terms of the operators $\hat{G}_i(q, p; x)$. Indeed, using Eq. (88) and recalling that the expectation values $\langle \hat{G}_i \rangle_n$ are real (by virtue of the Hermiticity of $\hat{G}_i$), we can rewrite Berry’s connection, $A_i^{(n)}(x) := -\text{Im} \langle n | \partial n \rangle$, in the following form:

$$A_i^{(n)}(x) = \frac{1}{\hbar} \langle \hat{G}_i \rangle_n.$$ \hspace{1cm} (90)

Finally, we note that that taking into account Eqs. (88) and (90), the action of the operator $\Delta \hat{G}_i := \hat{G}_i - \langle \hat{G}_i \rangle_n$ on the state $|n(x)\rangle$ can be written as

$$\Delta \hat{G}_i |n(x)\rangle = \left(i\hbar \partial_t - \langle \hat{G}_i \rangle_n \right) |n\rangle = i\hbar \left( \partial_i + i A_i^{(n)} \right) |n\rangle,$$ \hspace{1cm} (91)

which resembles the structure of the covariant derivative $D_{ij}^{(n)} = \partial_i + i A_i^{(n)}$ with connection $A_i^{(n)}$.

In the following example we shall see that Eqs. (89) and (90) yield the expected results.

**Example: Generalized harmonic oscillator with a linear term**

For this example we consider the quantum generalized harmonic oscillator with a linear term described by the Hamiltonian operator (60). The notation used here is the same as in Example B of Sec. IV. The starting point is to promote the corresponding classical functions $G_i$ given by (57) to the quantum operators:

$$\hat{G}_i(q, p; x) = \frac{1}{2} f_i(x) (\hat{q} \hat{p} + \hat{p} \hat{q}) + g_i(x) \hat{q}^2 + h_i(x) \left( \hat{p} + \frac{Y}{Z} \hat{q} \right),$$ \hspace{1cm} (92)

where $f_i(x)$, $g_i(x)$, and $h_i(x)$ are given by Eqs. (58a), (58b), and (58c), respectively. Notice that by construction (92) is Hermitian. Then, using the eigenfunctions (62) and the properties of the Hermite functions (41), we compute the quantum metric tensor (89) with the operators (92), obtaining

$$g_{ij}^{(n)}(x) = \frac{1 + n + n^2}{2} \left( f_i(x)f_j(x) + \frac{Z^2}{\omega^2} l_i(x)l_j(x) \right) + \left( n + \frac{1}{\hbar Z} \right) \omega \left( 4W^2Z^4 \omega^6 l_i(x)l_j(x) + m_i(x)m_j(x) \right),$$ \hspace{1cm} (93)

where $l_i(x) := g_i(x) - \frac{Y}{Z} f_i(x)$ and $m_i(x) := h_i(x) - \frac{Y^2}{Z^2} f_i(x)$. It can be verified, by explicit calculation, that Eq. (93) leads directly to Eq. (63), which corroborates the validity of Eq. (89). Finally, we apply Eq. (90) to the operators (92). The result is

$$A_i^{(n)}(x) = \left( g_i(x) - \frac{Y}{Z} f_i(x) \right) \left[ \left( n + \frac{1}{\hbar} \right) \frac{Z}{\omega} + \frac{W^2Z^2}{\hbar \omega^4} \right],$$ \hspace{1cm} (94)

which leads to Eq. (67) and, hence, verifies the validity of Eq. (90).

**VI. CONCLUSION**

In this paper, we have introduced the metric (19) for classical integrable systems and shown through examples that it corresponds to the classical counterpart of the quantum metric tensor (2). The classical metric is defined on the parameter space and provides a measure of the distance between nearby points in phase space, which is induced by the adiabatic evolution of the classical system. We investigate the main features of this classical metric. In particular, we show that this metric is gauge invariant in the parameter space in the sense that it remains unchanged when we perform the canonical transformation (22), meaning that this classical metric is independent of the “zero” point from which we measure the angle variables. Most importantly, we find for the considered examples that this metric agrees with the quantum metric tensor in rank and the functional dependence on the parameters. This allowed us to establish the exact relation between both metrics for the generalized harmonic oscillator and the generalized harmonic oscillator with a linear term, provided the Bohr-Sommerfeld quantization rule for action variable. For the nontrivial example of the quartic anharmonic oscillator, these metrics were calculated by using perturbation theories, and hence we find an approximate relation between them. We use the generating function (9) of translations in the parameter space as the fundamental object to build the aforementioned classical metric and demonstrated that Hannay’s curvature could also be expressed in terms of it; thereby providing a unified treatment for both geometric structures. Finally, we extend the use of this classical generating function to the quantum case and obtain alternative expressions for the quantum metric tensor and Berry’s connection, which are verified for the case of the quantum generalized harmonic oscillator with a linear term.

We would like to close by pointing out some remarks. First, it would be interesting to address the possible existence of the classical analog of the non-Abelian quantum metric tensor proposed in Ref. [38], and the generalization of the metric (19) for the case of classical systems with chaotic dynamics along the lines of Ref. [39]. In particular, those authors first study the quantum case, where they resort to time dependence to find an expression for Berry’s curvature. Then they introduce a semi-
classical approximation and get an expression for the
classical curvature, where instead of employing action-
gle variables, they perform the integration restricted
to a particular energy shell. This classical curvature then
reduces to Hannay’s curvature when the system is inte-
grable. In this spirit, as both the quantum and classical
cases are still tractable, we may as well generalize
our proposed classical metric. Nevertheless, for a nonin-
tegrable Hamiltonian that slightly differs from an inte-
grable Hamiltonian, the classical metric (19) might shed
some light on the chaotic behavior through the applica-
tion of the canonical perturbation theory. A further gen-
eralization of Eq. (19) is one wherein the classical metric
is invariant under a more general gauge transformation
in Eq. (22) also depends on the angle
variables, this motivated by the work of Ref. [33]. An-
other interesting and useful future consideration is how
to generalize the metric (19) for a classical field theory.

Apart from possible generalizations, the metric (19),
being the classical analog of the quantum metric tensor,
may help to provide more insight into the investigation of
quantum phase transitions. Furthermore, the metric (19)
may be relevant in the context of shortcuts to adiabatic
processes in classical integrable systems, which consist of
the use of a control Hamiltonian $K_{c}(\phi, I; x)$ that turns
out to be $K_{c} = G_{i}(\phi, I; x)\dot{x}^{2}$ and achieves a constant action
variable $I$ with arbitrarily fast changing parameters $x$ [40].

In this line of thought, it may be noted that in
Ref. [41] is proposed a metric analogous to Eq. (19) that
emerges in the study of the thermodynamic cost of short-
cuts to adiabaticity and defines a distance between the
initial and final statistical states of the classical system.

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Appendix: Quantum quartic anharmonic oscillator

The Hamiltonian operator of the system is

$$\hat{H} = \frac{\hat{p}^2}{2m} + \frac{k}{2}\hat{q}^2 + \frac{\lambda}{4!}\hat{q}^4.$$  \hspace{1cm} (A.1)

Here we consider a perturbative treatment in the param-
eter $\lambda$ since there is not an exact solution to the resulting
Schrödinger equation. Furthermore, we restrict ourselves
to obtain the ground-state wave function and its energy
up to third order in $\lambda$. To accomplish this task, we will
use the method proposed in Ref. [42], which is suitable
to find corrections to large powers of $\lambda$.

The adimensional form of the eigenvalue problem con-
sidered is

$$\left(-\frac{d^2}{dQ^2} + Q^2 + \Lambda Q^4\right)\psi(Q) = \epsilon \psi(Q),$$  \hspace{1cm} (A.2)

where we defined the following quantities

$$\Lambda := \frac{h\lambda}{12m^2\omega_0^2}, \quad Q := \sqrt{\frac{m\omega_0}{\hbar}}, \quad \epsilon := \frac{2E}{\hbar\omega_0}.$$  \hspace{1cm} (A.3)

with $\omega_0 = (k/m)^{1/2}$. After applying the method, the
resulting nonnormalized ground-state wave function $\psi_0(Q)$,
to third order in $\Lambda^3$, is

$$\psi_0(Q;x) = e^{\frac{\lambda^2}{4}} \left[ 1 - \frac{\Lambda Q^2}{8} (Q^2 + 3) + \frac{\Lambda^2 Q^2}{384} (3Q^6 + 26Q^4 + 93Q^2 + 252) - \frac{\Lambda^3 Q^2}{3072} (Q^{10} + 17Q^8 + 141Q^6 + 813Q^4 + 2916Q^2 + 7992) \right],$$  \hspace{1cm} (A.4)

and the corresponding energy is

$$\epsilon = 1 + \frac{3}{4} \Lambda - \frac{21}{16} \Lambda^2 + \frac{333}{64} \Lambda^3.$$  \hspace{1cm} (A.5)

Returning to our original variables and normalizing, we arrive at

$$\psi_0(q;x) = e^{-\frac{m\omega_0}{\pi\hbar}} \sqrt{\frac{m\omega_0}{\pi\hbar}} \frac{\lambda P_1(q;x)}{884736 \sqrt{\pi m^3 \omega_0^3} h^5} - \frac{\lambda^2 P_2(q;x)}{339738624 \sqrt{\pi m^2 \omega_0^2} h^9} \] \] $$  \hspace{1cm} (A.6)

and

$$E_0 = \frac{\hbar\omega_0}{2} \frac{h^2 \lambda^3}{32 m^2 \omega_0^2} - \frac{7h^3 \lambda^2}{1536m^2 \omega_0^2} + \frac{37h^4 \lambda^3}{24576m^3 \omega_0^2}.$$  \hspace{1cm} (A.7)

where

$$P_1(q;x) = 4m^2 \omega_0^2 q^4 + 12hm\omega_0 q^2 - 9h^2, \quad (A.8)$$

$$P_2(q;x) = 48m^4 \omega_0^4 q^8 + 416hm^3 \omega_0^3 q^6 + 1272h^2 m^2 \omega_0^2 q^4 + 3384h^3 m \omega_0 q^2 - 4677h^3, \quad (A.9)$$

$$P_3(q;x) = 64m^6 \omega_0^6 q^{12} + 1088hm^5 \omega_0^5 q^{10} + 8592h^2 m^4 \omega_0^4 q^8 + 48288h^3 m^3 \omega_0^3 q^6 + 154524h^4 m^2 \omega_0^2 q^4 + 419076h^5 m \omega_0 q^2 - 729153h^6.$$  \hspace{1cm} (A.10)

Plugging Eq. (A.6) into Eq. (2) and keeping terms up to
second order in $\lambda$ (due to derivatives with respect to $\lambda$
are present), we arrive at the components of the quantum
metric tensor given by Eq. (87).
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