How bright is the proton?
A precise determination of the photon parton distribution function

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It has become apparent in recent years that it is important, notably for a range of physics studies at the Large Hadron Collider, to have accurate knowledge on the distribution of photons in the proton. We show how the photon parton distribution function (PDF) can be determined in a model-independent manner, using electron–proton (ep) scattering data, in effect viewing the ep → e+X process as an electron scattering off the photon field of the proton. To this end, we consider an imaginary, beyond the Standard Model process with a flavour changing photon–lepton vertex. We write its cross section in two ways, one in terms of proton structure functions, the other in terms of a photon distribution. Requiring their equivalence yields the photon distribution as an integral over proton structure functions. As a result of the good precision of ep data, we constrain the photon PDF at the level of 1–2% over a wide range of momentum fractions.

A fast-moving particle generates an associated electromagnetic field which can be interpreted as a distribution of photons, as originally calculated by Fermi, Weizsäcker and Williams [11, 12] for point-like charges. The corresponding determination of the photon distribution for hadrons, specifically fγ/p, for the proton, has however been the subject of debate over recent years.

The photon distribution is small compared to that of the quarks and gluons, since it is suppressed by a power of the electromagnetic coupling α. Nevertheless, it has been realised in the past few years that its poor knowledge is becoming a limiting factor in our ability to predict key scattering reactions at CERN’s Large Hadron Collider (LHC). Notable examples are the production of the Higgs boson through W/Z fusion [4], or in association with an outgoing weak boson [5]. For W±H production it is the largest source of uncertainty [6]. The photon distribution is also potentially relevant for the production of lepton-pairs [7, 11, 19], top-quarks [12], pairs of weak bosons [15, 18] and generally enters into electroweak corrections for almost any LHC process. The diphoton excess around 750 GeV seen by ATLAS and CMS [19, 20] has also generated interest in understanding fγ/p.

The two most widely used estimates of fγ/p are those included in the MRST2004QED [21] and NNPDF23QED [22] parametrisations of the photon structure. In the NNPDF approach, the photon distribution is constrained mainly by LHC data on the production of pairs of leptons, pp → ℓ+ℓ−. This is dominated by q̄q → ℓ+ℓ−, with a small component from γγ → ℓ+ℓ−. The drawback of this approach is that even with very small uncertainties in ℓ+ℓ− production data [8, 30], in the QCD corrections to q̄q → ℓ+ℓ− and in the quark and anti-quark distributions, it is difficult to obtain high-precision constraints on fγ/p.

In the MRST2004QED approach, the photon is instead modeled. It is assumed to be generated as emissions from free, point-like quarks, using quark distributions fitted from deep-inelastic scattering (DIS) and other data. The free parameter in the model is an effective mass-scale below which quarks stop radiating, which was taken in the range between current-quark masses (a few MeV) and constituent-quark masses (a few hundred MeV). A more sophisticated approach [23] supplements a model of the photon component generated from quarks (“inelastic” part) with a calculation of the “elastic” component (whose importance has been understood at least since the early 1970’s [24]) generated by coherent radiation from the proton as a whole. This was recently revived in Refs. [25, 26]. Such an approach was also adopted for the CT14qed_inc [28] set, which further constrains the effective mass scale in the inelastic component using ep → eγ+X data [29], sensitive to the photon in a limited momentum range through the reaction eγ → eγ [30].

In this article we point out that electron-proton (ep) scattering data already contains all the information that is needed to accurately determine fγ/p. It is common to think of ep scattering as a process in which a photon emitted from the electron probes the structure of the proton. However one can equivalently think of it as an electron probing the photon field generated by the proton itself. Thus the ep scattering cross section is necessarily connected with fγ/p. (This point of view is implicit also in Refs. [31, 32].) A simple way to make the connection manifest is to consider, instead of ep scattering, the fictitious process ℓ+p → L+X, where ℓ and L are neutral leptons, with ℓ massless and L massive with mass M. We assume a transition magnetic moment coupling of the form L_{int} = (e/\Lambda L) \sigma^{\mu\nu} F_{\mu\nu} l. Here \sigma^2(\mu^2)/(4\pi) \equiv \alpha(\mu^2) is the MS QED coupling evaluated at the scale \mu, and the arbitrary scale \Lambda \gg \sqrt{s} (where \sqrt{s} is the centre-of-mass energy) is introduced to ensure the correct dimensions.

The crucial observation that we rely on is inspired in part by Drees and Zeppenfeld’s study of supersymmetric particle production at ep colliders [33]: there are two
ways of writing the heavy-lepton production cross section $\sigma$, one in terms of standard proton structure functions, $F_2$ and $F_L$ (or $F_1$), the other in terms of the proton PDFs $f_{a/p}$, where the dominant flavour that contributes will be $a = \gamma$. Equating the latter result with the former will allow us to determine $f_{\gamma/p}$.

We start with the inclusive cross section for $l(k) + p(p) \to L(k') + X$. Defining $q = k - k'$, $Q^2 = -q^2$ and $x_{ni} = Q^2/(2pq)$, we have

$$\sigma = \frac{1}{4p \cdot k} \int \frac{d^4q}{(2\pi)^4} q^2 c_{\gamma\phi}(q^2) \left[ 4\pi W_{\mu\nu}(p, q) L^\mu L^\nu(k, q) \right] \times 2\pi \delta((k - q)^2 - M^2), \quad (1)$$

where the proton hadronic tensor (as defined in [52]) is given by $W_{\mu\nu}(p, q) = -g_{\mu\nu} F_1(x_{ni}, Q^2) + p_{\mu} p_{\nu}/(pq) F_2(x_{ni}, Q^2)$ up to terms proportional to $q_{\mu}, q_{\nu}$, and the leptonic tensor is $L^\mu L^\nu(k, q) = \frac{1}{2} (c_{\gamma\phi}(q^2)/4!) \text{Tr} \left[ \{\gamma \mu, \gamma \nu\} \{k + M\} \{\gamma \nu, \gamma \mu\} \right]$. In Eq. (1) we introduced the physical QED coupling

$$c_{\gamma\phi}(q^2) = e^2(\mu^2)/(1 - \Pi(q^2; \mu^2, e^2(\mu^2))), \quad (2)$$

where $\Pi$ is the photon self energy and $\mu$ is the renormalisation scale. We stress that Eq. (1) is accurate up to corrections of order $\sqrt{\alpha}/\Lambda$, since neither the electromagnetic current nor the $L\gamma\ell$ vertex are renormalised.

We find

$$\sigma = c_0 \frac{1}{2\pi} \int_x^{1 - \frac{z_{\text{min}}}{z}} \frac{dz}{z} \int_{Q_{\text{max}}^2}^{Q_{\text{min}}^2} \frac{dQ^2}{Q^2} F_2(x, Q^2) \left( 2 - 2z + z^2 - \frac{2x^2m_p^2}{Q^2} + \frac{2Q^2}{M^2} - \frac{2x^2Q^2m_p^2}{M^4} \right) F_2(x/z, Q^2) \left[ 2 - 2z + z^2 \right] \left( -\frac{x^2Q^2}{2M^2} + \frac{z^2Q^4}{2M^4} \right) F_L(x/z, Q^2) \right], \quad (3)$$

where $x = M^2/(s - m_p^2)$, $m_p$ is the proton mass, $F_L(x, Q^2) = (1 + 4m_p^2x^2/Q^2) F_2(x, Q^2) - 2x F_1(x, Q^2)$ and $c_0 = 16\pi^2/\alpha^2$. Assuming that $M^2 \gg m_p^2$, we have $Q_{\text{min}}^2 = x^2m_p^2/(1 - z)$ and $Q_{\text{max}}^2 = M^2(1 - z)/z$.

The same result in terms of parton distributions can be written as

$$\sigma = c_0 \sum_a \int_x^1 \frac{dz}{z} \hat{s}_a(z, \mu^2) M^2 \frac{dz}{z} f_{a/p} \left( \frac{M^2}{z^2}, \mu^2 \right), \quad (4)$$

where in the $\overline{\text{MS}}$ factorisation scheme

$$\hat{s}_a(z, \mu^2) = \alpha(\mu^2) \delta(1 - z) \delta_{a\gamma} + \frac{\alpha^2(\mu^2)}{2\pi} \left[ -2 + 3z + z p_{\gamma q}(z) \ln \frac{M^2(1 - z)^2}{z^2} \right] \sum_{i \in \{q, \bar{q}\}} e_i^2 \delta_{ai} + \ldots, \quad (5)$$

where $e_i$ is the charge of quark flavour $i$ and $z p_{\gamma q}(z) = 1 + (1 - z)^2$. To understand which terms we choose to keep, observe that the photon will be suppressed by $\alpha L$ relative to the quark and gluon distributions, which are of order $(\alpha s L)^n$, where $L = \ln \mu^2/m_p^2 \sim 1/\alpha_s$. The contribution proportional to $F_2$ in Eq. (3) is of order $\alpha^2 L(\alpha_s L)^n$, while that proportional to $F_L$ is of order $\alpha^2(\alpha_s L)^n$. We neglect terms that would be of order $\alpha^2 L(\alpha_s L)^n$ or $\alpha^2(\alpha_s L)^n$.

By requiring the equivalence of Eqs. (3) and (4) up to the orders considered, one obtains (in the $\overline{\text{MS}}$ scheme):

$$xf_{\gamma/p}(x, \mu^2) = \frac{1}{2\pi\alpha(\mu^2)} \int_x^1 \frac{dz}{z} \left\{ \int_x^{z^2} \frac{dQ^2}{Q^2} \alpha^2(Q^2) \left[ \left( z p_{\gamma q}(z) + 2x^2m_p^2 \right) F_2(x/z, Q^2) - z^2 F_L \left( \frac{x}{z}, Q^2 \right) \right] - \alpha^2(\mu^2) z^2 F_2 \left( \frac{x}{z}, \mu^2 \right) \right\}, \quad (6)$$

where the result includes all terms of order $\alpha L(\alpha_s L)^n$, $\alpha^2(\alpha_s L)^n$ and $\alpha^2 L^2(\alpha_s L)^n$ [30]. Within our accuracy $\alpha_{\text{em}}(Q^2) \approx \alpha(Q^2)$. The conversion to the $\overline{\text{MS}}$ factorisation scheme, the last term in Eq. (4), is small (see Fig. 2).

From Eq. (6) we have derived expressions up to order $\alpha_s$ for the $F_{\gamma q}, F_{\gamma g}$ and $F_{\gamma s}$ splitting functions using known results for the $F_2$ and $F_L$ coefficient functions and for the QED $\beta$-function. Those expressions agree with the results of a direct evaluation in Ref. [37].

The evaluation of Eq. (6) requires information on $F_2$ and $F_L$. Firstly (and somewhat unusually in the context of modern PDF fits), we will need the elastic contributions to $F_2$ and $F_L$:

$$F_{\text{el}}^2(x, Q^2) = \frac{[G_E(Q^2)]^2 + 2[G_M(Q^2)]^2}{1 + \tau} \delta(1 - x), \quad (7a)$$

$$F_{\text{el}}^L(x, Q^2) = \frac{[G_E(Q^2)]^2}{\tau} \delta(1 - x), \quad (7b)$$

where $\tau = Q^2/(4m_p^2)$ and $G_E$ and $G_M$ are the electric and magnetic Sachs form factors of the proton (see e.g. Eqs. (19) and (20) of Ref. [33]). A widely used approach for $G_{E,M}$ is the dipole form $G_{E,M}^d(Q^2) = 1/(1 + Q^2/m_\text{dip}^2)^2$, $G_M(Q^2) = \mu_p G_E(Q^2)$ with $m_\text{dip} = 0.71$ GeV$^2$ and $\mu_p \approx 2.793$. This form is of interest for understanding qualitative asymptotic behaviours, predicting $f_{\gamma/p}(x) \sim \alpha(1 - x)^4$ at large $x$ dominated by the magnetic component, and $xf_{\gamma/p}(x) \sim \alpha \ln 1/x$ at small $x$ dominated by the electric component. However for accurate results, we will rather make use of a recent fit to precise world data by the A1 collaboration [39], which shows clear deviations from the dipole form, with an impact of up to 10% on the elastic part of $f_{\gamma/p}(x)$ for $x < 0.5$. The data constrains the form factors for $Q^2 \leq 10$ GeV$^2$. At large $x$, Eq. (6) receives contributions only from $Q^2 > x^2m_p^2/(1 - x)$, which implies that the elastic contribution to $f_{\gamma/p}$ is known for $x < 0.9$.

Note that the last term in Eq. (6) does not have an elas-
The inelastic components of $F_2$ and $F_L$ contribute for $W^2 = m_p^2 + Q^2(1 - x)/x > (m_p + m_\pi)^2$. One needs data over a large range of $x$ and $Q^2$. This is available thanks to a long history of $ep$ scattering studies. We break the inelastic part of the $(x, Q^2)$ plane into three regions, as illustrated in Fig. 1. In the resonance region, $W^2 \lesssim 3.5$ GeV$^2$ we use a fit to data by CLAS [40], and also consider an alternative fit to the world data by Christy and Bosted (CB) [41]. In the low-$Q^2$ continuum region we use the GD11-P fit by Hermes [12] based on the ALLM parametric form [13]. Both the GD11-P and CB resonance fits are constrained by photoproduction data, i.e. they extend down to $Q^2 = 0$. The CLAS fit also behaves sensibly there. (Very low $Q^2$ values play little role because the analytic properties of the $W^{\mu\nu}$ tensor imply that $F_2$ vanishes as $Q^2$ at fixed $W^2$.) These fits are for $F_2(x, Q^2)$. We also require $F_L$, or equivalently $R = \sigma_L/\sigma_T$, which are related by

$$F_L(x, Q^2) = F_2(x, Q^2) \left( 1 + \frac{4m_p^2x^2}{Q^2} \right) \frac{R(x, Q^2)}{1 + R(x, Q^2)},$$

and we use the parametrisation for $R$ from HERMES [12], extended to vanish smoothly as $Q^2 \to 0$. The leading twist contribution to $F_L$ is suppressed by $\alpha_s(Q^2)/(4\pi)$. At high $Q^2$ we determine $F_2$ and $F_L$ from the PDF4LHC15_nnlo100 [44] merger of next-to-next-to-leading order (NNLO) [45, 46] global PDF fits [47–49], using massless NNLO coefficient functions [50, 53] implemented in HOPPET [54, 56].

In Fig. 2 we show the various contributions to our photon PDF, which we dub “LUXqed”, as a function of $x$, for a representative scale choice of $\mu = 100$ GeV. There is a sizeable elastic contribution, with an important magnetic component at large values of $x$. The white line represents contributions arising from the $Q^2 < 1$ region of all the structure functions, including the full elastic contribution. For the accuracy we are aiming at, all contributions that we have considered, shown in Fig. 2, have to be included, and inelastic contributions with $Q^2 < 1$ cannot be neglected. The photon momentum fraction is 0.43% at $\mu = 100$ GeV.

In Fig. 3 we show the sources contributing to the uncertainty on our calculation of $f_{\gamma/p}$ at our reference scale $\mu = 100$ GeV. They are stacked linearly and consist of: a conservative estimate of $\pm50\%$ for the uncertainty on $R = \sigma_L/\sigma_T$ at scales $Q^2 < 9$ GeV$^2$ (R); standard 68% CL uncertainties on the PDFs, applied to scales $Q^2 \geq 9$ GeV$^2$ (PDF); a conservative estimate of the uncertainty on the elastic form factors, equal to the sum in quadrature of the fit error and of the estimated size of the two-photon exchange contribution in [39] (E); an
estimate of the uncertainty in the resonance region taken as the difference between the CLAS and CB fits (RES); a systematic uncertainty due to the choice of the transition scale between the HERMES $F_2$ fit and the perturbative determination from the PDFs, obtained by reducing the transition scale from 9 to 5 GeV$^2$ (M); missing higher order effects, estimated using a modification of Eq. (6), with the upper bound of the $Q^2$ integration set to $\mu^2$ and the last term adjusted to maintain $\alpha^2(\alpha_s L)^n$ accuracy (HO); a potential twist-4 contribution to $F_L$ parametrised as a factor $(1 + 5.5 \text{ GeV}^2/Q^2)^{57}$ for $Q^2 \geq 9 \text{ GeV}^2$ (T). One-sided errors are all symmetrised. Our final uncertainty, shown as a solid line in Fig. 3 is obtained by combining all sources in quadrature and is about 1-2\% over a large range of $x$ values.

In Fig. 4 we compare our LUXqed result for the $\overline{MS}$ $f_{\gamma/p}$ to determinations available publicly within LHAPDF 58. Of the model-based estimates, CT14qed_inc 28 and MRST2004 21, CT14qed_inc is in good agreement with LUXqed within its uncertainties. Its model for the inelastic component is constrained by $ep \rightarrow e\gamma + X$ data from ZEUS 29 and includes an elastic component. Note however that, for the neutron, CT14qed_inc neglects the important neutron magnetic form factor. As for the model-independent determinations, NNPDF30 59, which

![Graph showing the ratio of common PDF sets to our LUXqed result.](image)

**FIG. 4.** The ratio of common PDF sets to our LUXqed result, along with the LUXqed uncertainty band (light red). The CT14 and MRST bands correspond to the range from the PDF members shown in brackets (68% cl. in CT14’s case). The NNPDF bands span from $\max(\mu_r - \sigma_r, \mu_1)$ to $\mu_r + \sigma_r$, where $\mu_r$ is the average (represented by the blue line), $\sigma_r$ is the standard deviation over replicas, and $\mu_1$ denotes the 16$^\text{th}$ percentile among replicas. Note the different $y$-axes for the panels.

notably extends NNPDF23 22 with full treatment of $\alpha(\alpha_s L)^n$ terms in the evolution [60], almost agrees with our result at small $x$. At large $x$ its band overlaps with our result, but the central value and error are both much larger.

Similar features are visible in the corresponding $\gamma\gamma$ partonic luminosities, defined as

$$\frac{dL_{\gamma\gamma}}{d\ln M^2} = \frac{M^2}{s} \int \frac{dz}{z} f_{\gamma/p}(z, M^2) f_{\gamma/p}(\frac{M^2}{2s}, M^2),$$

and shown in Fig. 5 as a function of the $\gamma\gamma$ invariant mass $M$, for several centre-of-mass energies.

As an application, we consider $pp \rightarrow HW^+(-\rightarrow \ell^+\nu) + X$ at $\sqrt{s} = 13 \text{ TeV}$, for which the total cross section without photon-induced contributions is $91.2 \pm 1.8 \text{ fb}$ [61], with the error dominated by (non-photonic) PDF uncertainties. Using HAWK 2.01 [62], we find a photon-induced contribution of $5.5^{+4.3}_{-2.9} \text{ fb}$ with NNPDF30, to be compared to $4.4 \pm 0.1 \text{ fb}$ with LUXqed.

In conclusion, we have obtained a formula (i.e. Eq. 6) for the $\overline{MS}$ photon PDF in terms of the proton structure functions, which includes all terms of order $\alpha L (\alpha_s L)^n$, $\alpha (\alpha_s L)^n$ and $\alpha^2 L^2 (\alpha_s L)^n$. Our method can be easily generalised to higher orders in $\alpha_s$ and holds for any hadronic bound state. Using current experimental information on $F_2$ and $F_L$ for protons we obtain a photon PDF with much smaller uncertainties than existing determinations, as can be seen from Fig. 4. The photon PDF has a substantial contribution from the elastic form factor ($\sim 20\%$) and from the resonance region ($\sim 5\%$) even for high values of $\mu \sim 100-1000 \text{ GeV}$.

Our photon distribution, incorporating quarks and gluons from PDF4LHC15.nmlo_100 [44] and evolved with a QED-extended version of HOPPET is available as part of

![Graph showing $\gamma\gamma$ luminosity in $pp$ collisions as a function of the $\gamma\gamma$ invariant mass $M$, at four collider centre-of-mass energies.](image)

**FIG. 5.** $\gamma\gamma$ luminosity in $pp$ collisions as a function of the $\gamma\gamma$ invariant mass $M$, at four collider centre-of-mass energies. The NNPDF30 results are shown only for 8 and 100 TeV. The uncertainty of our LUXqed results is smaller than the width of the lines.
the LHAPDF library as the LUXqed\_PDF4LHC15\_nnlo\_100 set and from http://cern.ch/luxqed. Note that it is only valid for scales $\mu > 10$ GeV.

More details of our analysis, including a derivation using PDF operators, computation of splitting functions, higher order corrections to Eq. (6), as well as an extension to the polarized case will be given in a longer publication [53].

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