The entanglement spectrum and Rényi entropies of non-relativistic conformal fermions

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We characterize non-perturbatively the Rényi entropies of degree $n = 2, 3, 4,$ and 5 of three-dimensional, strongly coupled many-fermion systems in the scale-invariant regime of short interaction range and large scattering length, i.e. in the unitary limit. We carry out our calculations using lattice methods devised recently by us. Our results show the effect of strong pairing correlations on the entanglement entropy, which modify the sub-leading behavior for large subsystem sizes (as characterized by the dimensionless parameter $x = k_F L_A$, where $k_F$ is the Fermi momentum and $L_A$ the linear subsystem size), but leave the leading order unchanged relative to the non-interacting case. Moreover, we find that the onset of the sub-leading asymptotic regime is at surprisingly small $x \approx 2 - 4$. We provide further insight into the entanglement properties of this system by analyzing the spectrum of the entanglement Hamiltonian of the two-body problem from weak to strong coupling. The low-lying entanglement spectrum displays clear features as the strength of the coupling is varied, such as eigenvalue crossing and merging, a sharp change in the Schmidt gap, and scale invariance at unitarity. Beyond the low-lying component, the spectrum appears as a quasi-continuum distribution, for which we present a statistical characterization; we find, in particular, that the mean shifts to infinity as the coupling is turned off, which indicates that that part of the spectrum represents non-perturbative contributions to the entanglement Hamiltonian. In contrast, the low-lying entanglement spectrum evolves to finite values in the noninteracting limit. The scale invariance of the unitary regime guarantees that our results are universal features intrinsic to three-dimensional quantum mechanics and represent a well-defined prediction for ultracold atom experiments, which were recently shown to have direct access to the entanglement entropy.

I. INTRODUCTION

This is an incredibly exciting time for research in ultracold atomic physics. The degree of control that experimentalists have achieved continues to rise, year after year, along with their ability to measure collective properties in progressively more ingenious ways (see e.g. [1–4]). Indeed, after the realization of Bose-Einstein condensates over two decades ago [5–7] (see also [8]), followed by Fermi condensates in 2004 [9], the field entered an accelerated phase and rapidly developed control of multiple parameters such as temperature, polarization, and interaction strength (in alkali gases via Feshbach resonances, see e.g. [10], and more recently in alkaline-earth gases via orbital resonances, see e.g. [11–13]), as well as exquisite tuning of external trapping potentials. Additionally, multiple properties can be measured, ranging from the equation of state (see e.g. [14–16]) to hydrodynamic response (see e.g. [17, 18]) and, more recently, the entanglement entropy [19, 20].

This sustained progress has strengthened the intersections with other areas of physics, in particular modern condensed matter physics and quantum information [21], as well as with nuclear [22] and particle physics [23–25]. Quantum simulation by fine manipulation of nuclear spins, electronic states, and optical lattices, now appears more realistic than ever [26–28]. At the interface between many of those areas lies a deceptively simple non-relativistic scale invariant system: the unitary Fermi gas, which corresponds to the limit of vanishing interaction range $r_0$ and infinite s-wave scattering length $a$, i.e.

$$0 \leftarrow r_0 \ll n^{-1} \ll a \rightarrow \infty$$

(1)

where $n$ is the density; this regime corresponds to the threshold of two-body bound-state formation.

Both a model for dilute neutron matter and an actually realized resonant atomic gas, this universal spin-1/2 system has brought together the nuclear [30–32], atomic [33], and condensed matter physics areas [34–36], as well as the AdS/CFT area [37–39], due to the underlying non-relativistic conformal invariance [40]. While many properties of this quintessential many-body problem are known (see e.g. [22] for an extensive review), other properties like entanglement and quantum information aspects have thus far remained unexplored, which brings us to our main point.

As this work is being written, quantum information concepts are increasingly becoming part of the modern language of quantum many-body physics (see e.g. Refs. [21, 41–43]), in particular with regards to the characterization of topological phases of matter and quantum computation, but also in connection with black holes (see e.g. [44]) and string theory (see e.g. [45]). In the past decade or so, a large body of work has been produced characterizing the entanglement properties of low dimensional systems (especially those with spin degrees of freedom [46–48]) at quantum phase transitions (in particular those with topological order parameters that defy a local

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description) as well as systems of noninteracting fermions and bosons [49–53], which presented a challenge of their own.

With that new perspective in mind, in this work we set out to characterize the entanglement properties of the unitary Fermi gas using non-perturbative lattice methods. We analyze the reduced density matrix, entanglement spectrum, and associated Rényi entanglement entropies of the two-body problem by implementing an exact projection technique on the lattice. For the many-body problem, we use a Monte Carlo method developed by us in Refs. [54, 55], based on the work of Ref. [56], to calculate the $n$-th Rényi entanglement entropy. We showed in that work that our method overcomes the signal-to-noise problem of naïve Monte Carlo approaches. We did that using the 1D Fermi-Hubbard model as a test case, but to our knowledge no previous calculations have been attempted for the challenging case of 3D Fermi gases.

The remainder of this paper is organized as follows: In Sec. II we present the main definitions and set the stage for Sec. III, where we explain how we carry out our calculations of the entanglement spectrum and entanglement entropies in two- and many-fermion systems. For completeness, we also include in that section a discussion on how to avoid the signal-to-noise issue that plagues entanglement-entropy calculations in the many-body case. We extend that discussion to the case of bosons in the same section. In Sec. IV we show our results for the entanglement spectrum and entropies of the two-body system along the BCS-BEC crossover, and in Sec. V we present the Rényi entanglement entropies of many fermions at unitarity. We present a summary and our main conclusions in Sec. VI. The appendices contain more detailed explanations of our few- and many-body methods.

II. DEFINITIONS: HAMILTONIAN, DENSITY MATRICES, AND THE ENTANGLEMENT ENTROPY

The Hamiltonian governing the dynamics of resonant fermions can be written as

$$\hat{H} = \hat{T} + \hat{V},$$

where the non-relativistic kinetic energy operator is

$$\hat{T} = \sum_{s=\uparrow, \downarrow} \int d^3r \left( \hat{\psi}^\dagger_s(r) \left( -\frac{\nabla^2}{2m} \right) \hat{\psi}_s(r), \right)$$

where $\hat{\psi}^\dagger_s(r)$ and $\hat{\psi}_s(r)$ are the creation and annihilation operators of particles of spin $s = \uparrow, \downarrow$ at location $r$.

The two-body, zero-range interaction operator is

$$\hat{V} = -g \int d^3r \hat{\psi}^\dagger_\uparrow(r) \hat{\psi}^\dagger_\downarrow(r) \hat{\psi}_\downarrow(r) \hat{\psi}_\uparrow(r),$$

where the bare coupling $g$ is tuned to the desired physical situation. By definition, the limit of unitarity is achieved by requiring that the ground state of the two-body problem lies at the threshold of bound-state formation (note that in 1D and 2D bound states form at arbitrarily small attractive coupling, but a finite value is required in 3D). Because our work was carried out in a finite volume with periodic boundary conditions, we used Lüscher’s formalism [57, 58] to relate the bare coupling to the scattering length in the analysis of the BCS-BEC crossover. We describe that procedure below, when showing the results for the two-body problem.

The full, normalized density matrix of the system is

$$\hat{\rho} = \frac{e^{-\beta \hat{H}}}{\mathcal{Q}},$$

where

$$\mathcal{Q} = \text{Tr}_\mathcal{H} \left[ e^{-\beta \hat{H}} \right],$$

is of course the canonical partition function, and $\mathcal{H}$ is the full Hilbert space. In this work we are concerned with systems in a pure state, namely the ground state $|\Xi\rangle$, such that the full density matrix can be written as

$$\hat{\rho} = |\Xi\rangle\langle\Xi|.$$

Both in the few- and many-body systems we analyze here, the ground-state density matrix will be approached by a projection method we describe below.

FIG. 1: The (bipartite) entanglement entropies computed in this work correspond to partitioning the system into a subsystem $A$ (in coordinate space, but it can also be defined in momentum space) and its complement $\bar{A}$. In practice, the calculations are carried out on systems that live in a cubic lattice of side $L$, and the subsystems are defined by cubic subregions of side $L_A \leq L$. The reduced density matrix $\hat{\rho}_A$ of the open system $A$ contains the information about entanglement between $A$ and $\bar{A}$, and is obtained by tracing the full density matrix over the states supported by $\bar{A}$, which form the Hilbert space $\mathcal{H}_\bar{A}$.
The density matrix $\hat{\rho}_A$ of subsystem $A$, usually referred to as the reduced density matrix, is defined by tracing over the degrees of freedom supported by $\bar{A}$, i.e. tracing over the states in $\mathcal{H}_{\bar{A}}$:

$$\hat{\rho}_A = \text{Tr}_{\bar{A}} \hat{\rho}.$$  \hspace{1cm} (9)

Based on this definition, the properties of $A$ as an open subsystem can be formulated and computed using operators with support in $A$. In particular, a quantitative measure of entanglement between $A$ and $\bar{A}$ is given by the von Neumann entanglement entropy,

$$S_{VNA} = -\text{Tr}_{\bar{A}} [\hat{\rho}_A \ln \hat{\rho}_A],$$  \hspace{1cm} (10)

and by the $n$-th order Rényi entanglement entropy,

$$S_{n,A} = \frac{1}{1-n} \ln \text{Tr}_{\bar{A}} [\hat{\rho}_A^n].$$  \hspace{1cm} (11)

Naturally, these entropies vanish when $A$ is the whole system, as then there is full knowledge of the state of the system. In any other case, the entanglement entropy will be non-zero, unless the ground state factorizes into a state living in $A$ and a state living in $\bar{A}$. Because the entanglement between $A$ and $\bar{A}$ happens across the boundary that separates those regions, it is natural to expect $S_{VNA}$ to be extensive with the size of that boundary, i.e. proportional to the area delimiting $A$. This point was the topic of many papers in the last decade or so, especially in connection with quantum phase transitions (see e.g. [59]).

It was rigorously shown in recent years, however, that the Rényi entropy of non-interacting fermions with a well-defined Fermi surface presents a logarithmic violation of the area law [49–53]. This abnormality was confirmed numerically with the aid of overlap-matrix methods [60], which we reproduce in Fig. 2, where we explicitly show said logarithmic dependence (dashed line) as a function of $x = \frac{k_F L_A}{\pi N}$, where $k_F$ is the Fermi momentum and $L_A$ is the linear size of region $A$, such that

$$x = \frac{k_F L_A}{\pi N} = \frac{\pi N L_A}{2 L} \quad \text{in 1D},$$  \hspace{1cm} (12)

$$= (2\pi N)^{1/2} \frac{L_A}{L} \quad \text{in 2D},$$  \hspace{1cm} (13)

$$= (3\pi^2 N)^{1/3} \frac{L_A}{L} \quad \text{in 3D},$$  \hspace{1cm} (14)

where $N$ is the total particle number. Note that, at large enough $x$, finite size effects eventually take over and the entropy quickly tends to zero. The sub-leading oscillations were studied in detail in Ref. [61].

Although resonant fermions are strongly coupled (the regime is non-perturbative and away from any regime with small dimensionless parameters), we can expect $S_{n,A}$ to follow a similar trend as the non-interacting gas, for the following reasons. First, resonant fermions have a distinguishable Fermi surface (note, however, that that is quickly lost as one proceeds towards the BEC side of

A subsystem $A$ and its complement $\bar{A}$ (in coordinate or momentum space, see Fig. 1) support states that belong to Hilbert spaces $\mathcal{H}_A$ and $\mathcal{H}_{\bar{A}}$, respectively, such that the Hilbert space $\mathcal{H}$ of the full system can be written as a direct product space

$$\mathcal{H} = \mathcal{H}_A \otimes \mathcal{H}_{\bar{A}}.$$  \hspace{1cm} (8)

FIG. 2: Second Rényi entropy $S_2$ of $N$ non-interacting fermions in $d = 1, 2, 3$ dimensions (top to bottom) as a function of $x = \frac{k_F L_A}{\pi N}$, where $A$ is a segment, square, and cubic region, respectively, and $L_A$ is the corresponding linear size; $k_F$ is the Fermi momentum. $S_2$ is scaled by the surface area dependence, namely $x$ and $x^2$ in 2D and 3D, respectively. The $x$ axis is plotted logarithmically to show that, up to finite-size effects, the results heal to the expected asymptotic regime of linear dependence with $\log_{10} x$ (dashed line). This regime sets in at $x \approx 2 - 4$ across all $d$. Finite-size effects appear as a sudden drop at large $x$. 

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the resonance), whose role in the entanglement entropy has been emphasized many times (see e.g. [62]). Second, our experience with $S_{n,A}$ for the Hubbard model in other cases [54] indicates that very strong couplings $U/t$ are needed even in 1D (where quantum fluctuations are qualitatively stronger than in 3D) in order for $S_{n,A}$ to noticeably depart from the non-interacting result. Thus, we anticipate a similar behavior for resonant fermions as that of the bottom panel of Fig. 2; the latter provides some qualitative knowledge of where the leading logarithmic and sub-leading dependence sets in for $S_{n,A}$ as a function of $x = k_F L_A$. In fact, as we will see below, the onset of the asymptotic behavior (meaning dominated by leading and sub-leading dependence on $x$) at $x \approx 2 - 4$ is the same for unitarity as for the non-interacting case. This is surprising, as there is no obvious reason for that to be the case: had this onset appeared at $x \approx 10$, the calculations in this work would not have been possible, as they would have required huge lattices. We return to this discussion below, when presenting our results for the many-body case.

III. METHOD

In this section we explain the two approaches used in this work. We address the two-body problem first, which we solved with a direct (i.e. non-stochastic) projection method on the lattice. This problem can be solved exactly by changing to center-of-mass and relative coordinates. However, doing so implies using a method that only works in that case, and we are interested in techniques that can be used in a variety of situations (e.g. in the presence of external fields, more than two particles, time-dependent cases, and so forth). We then address the many-body problem using a method recently put forward by us, which we first presented and tested for one-dimensional systems in Ref. [55].

Although both approaches make use of an auxiliary field transformation, the ultimate utility of this technique is markedly different in each case. We detail below the divergence from common assumptions and notation.

At chemical potential $\mu$ and inverse temperature $β$, the grand canonical partition function $Z$ is defined via

$$Z = \text{Tr} \left[ e^{-β(\hat{H} - \mu \hat{N})} \right]$$

for Hamiltonian $\hat{H}$ and particle-number operator $\hat{N}$. Writing the inverse temperature as an integer number $N_r$ of steps, we implement a symmetric Suzuki-Trotter decomposition with the goal of separating each operator into distinct one- and two-body factors. For the Boltzmann factor, we obtain

$$e^{-β(\hat{H} - \mu \hat{N})} = \prod_{j=1}^{N_r} e^{-τg \hat{n}_s} e^{-τf \hat{s}_s} + O(τ^2)$$

were we define

$$\hat{K} = \hat{T} - \mu \hat{N}.$$  \hspace{1cm} (17)

At each position $\mathbf{r}$ and for each of the $N_r$ factors, we decompose the interaction via the introduction of a Hubbard-Stratonovich auxiliary field $\sigma$ which we choose to be of a continuous and compact form [63, 64]. More specifically for each spacetime position $(\mathbf{r}, \tau_j)$, where $\mathbf{r} \in [0, L]^3$ and $\tau_j = jτ$ for some $1 ≤ j ≤ N_r$, we write

$$e^{τg \hat{n}_s} = \int_{-π}^{π} \frac{dσ}{2π} \left( \mathbb{1} + B \hat{n}_s \sin σ \right) \left( \mathbb{1} + B \hat{n}_s \sin σ \right)$$

having suppressed the spacetime dependence of the field $σ$ and the spatial dependence of the fermion density operators $\hat{n}_s(\mathbf{r}) = \hat{\psi}_s^{\dagger}(\mathbf{r}) \hat{\psi}_s(\mathbf{r})$ where $s = \uparrow, \downarrow$. Knowing that $\hat{n}_s(\mathbf{r})$ is idempotent, it follows that

$$e^{τg \hat{n}_s} = 1 + (e^{τg} - 1)\hat{n}_s \hat{n}_s,$$  \hspace{1cm} (19)

which shows that the constant $B$ satisfies

$$e^{τg} - 1 = \frac{B^2}{2}.$$  \hspace{1cm} (20)

Collecting the integration measures, we obtain a path-integral form of the partition function accurate to quadratic order in the temporal lattice spacing, writing

$$Z = \int \mathcal{D}σ \text{Tr} \hat{U}[σ] + O(τ^2)$$

FIG. 3: Shown here is a representation of the lattice used in our calculations. Each horizontal lattice plane represents the 3D lattice where the system lives, and the vertical stacking of the planes represents the imaginary time direction. Although the original Hamiltonian is time-independent, the auxiliary field $σ$ that represents the interaction is supported by a spacetime lattice and induces a time dependence that disappears upon averaging.
where
\[ \hat{U}[\sigma] = \prod_{j=1}^{N_s} \hat{U}_j[\sigma], \]
and the individual factors are
\[ \hat{U}_j[\sigma] = e^{-\hat{K}_j/2} \prod_r \left( 1 + B \hat{n}_j(r) \sin \sigma(r, \tau_j) \right) \]
\[ \times \left( 1 + B \hat{n}_j(r) \sin \sigma(r, \tau_j) \right) e^{-\hat{K}_j/2}. \]

As the kinetic energy operator \( \hat{T} \) and the number operator \( \hat{N} \) are already written as products of flavor-specific operators, we may partition the operator \( \hat{U} \) into individual factors each of which assumes responsibility for the evolution of a particular fermion species \( s = \uparrow, \downarrow \). We do this by defining operators \( T_s, N_s, \) and \( K_s \) for \( s = \uparrow, \downarrow \) by
\[ \hat{T}_s = \int d^3r \hat{\psi}_s^\dagger(r) \left( -\frac{\nabla^2}{2m} \right) \hat{\psi}_s(r), \]
\[ \hat{N}_s = \int d^3r \hat{\psi}_s^\dagger(r) \hat{\psi}_s(r), \]
and \( \hat{K}_s = \hat{T}_s - \mu \hat{N}_s \). We then write
\[ \hat{U}_{j,s}[\sigma] = e^{-\hat{K}_s/2} \prod_r \left( 1 + B \hat{n}_s(r) \sin \sigma(r, \tau_j) \right) \]
\[ \times e^{-\hat{K}_s/2}, \]
such that
\[ \hat{U}[\sigma] = \hat{U}_t[\sigma] \hat{U}_l[\sigma], \]
where
\[ \hat{U}_s[\sigma] = \prod_{j=1}^{N_s} \hat{U}_{j,s}[\sigma]. \]

Performing the required Fock-space trace, the exponential form of each factor in the above provides (see e.g. [65])
\[ Z = \int \mathcal{D}\sigma \ \det \left( 1 + U^\dagger[\sigma] \right) \det \left( 1 + U[\sigma] \right), \]
where we have suppressed higher-order contributions in \( \tau \) (which are of order \( \tau^2 \)), and written a matrix \( U_s[\sigma] \) for the restriction of each of the operators \( \hat{U}_s[\sigma] \) to the single-particle Hilbert space. Each of those matrices contains an overall factor of the fugacity
\[ z = e^{\beta \mu}. \]

In what follows, we exhibit this factor explicitly and redefine the matrices \( U_s[\sigma] \) to reflect this revision. In this work, we exclusively treat unpolarized systems, and so we may treat the determinants as equivalent in derivations that follow by writing
\[ Z = \int \mathcal{D}\sigma \ \det^2 \left( 1 + z U[\sigma] \right), \]
and neglecting to denote the spin degree of freedom wherever context precludes confusion.

A. Direct lattice approach to the entanglement spectrum of the two-body problem

1. Identifying the transfer matrix

In order to illustrate the details as well as the generality of our technique, we show the main steps here in broad strokes and leave the details for Appendix A.

Using the above path-integral form of \( Z \), we first isolate the two-body sector. From the finite-temperature partition function Eq. (30), we may derive the conventional virial expansion in powers of the fugacity for each spin, which is given by
\[ Z = \sum_{N\uparrow, N\downarrow=0}^{\infty} z_N^\uparrow z_N^\downarrow Q_{N\uparrow, N\downarrow}, \]
where we have identified the coefficient of the \( N_s \)-th power of the fugacity as the \( N_s \)-particle canonical partition function \( Q_{N\uparrow, N\downarrow} \). Expanding the path-integral expression for the grand canonical partition function, we find that in terms of the matrix \( U[\sigma] \), the \((1+1)\)-particle partition function is
\[ Q_{1,1} = \int \mathcal{D}\sigma \ \text{tr}^2 U[\sigma]. \]

The path integral in \( Q_{1,1} \) above can be evaluated directly in a way that elucidates the form of the two-body transfer matrix. To that end, we define a four-index object from which the above squared trace may be obtained by suitable index contraction:
\[ R_{ab,cd} = \int \mathcal{D}\sigma \ U[\sigma]_{ac} U[\sigma]_{bd}. \]
The same four-index object, with indices properly contracted to account for antisymmetry, can be used to analyze the \((2+0)\)-particle case.

We next write out each of the matrices \( U[\sigma] \) in its product form; that is, we reintroduce Eq. (22) in matrix form:
\[ U[\sigma] = \prod_{j=1}^{N_s} U_j[\sigma]. \]
For each contribution to the \( N \)-body transfer matrix, exactly \( N \) factors of the matrix \( U[\sigma] \) appear, and as a result each temporal lattice point appears in the integrand \( N \) times. Turning to the individual factors, we write each of the matrices \( U_j[\sigma] \) in such a way as to exhibit the interaction. That is, we write
\[ U_j[\sigma] = TV_j[\sigma]T, \]
where
\[ T_{kk'} = e^{-\tau k^2/2} \delta_{kk'}. \]
is the single-particle form of the kinetic energy operator defined above (in momentum space), and the (position-space representation of the) auxiliary external potential operator has matrix elements

\[ [V_j[\sigma]]_{rr'} = (1 + B \sin(\mathbf{r}, \tau_j)) \delta_{rr'} . \]  

(39)

At this point, all matrix elements have been written out and can be shifted around as needed to carry out the path integral. The only non-zero results are obtained, of course, when an even number (in this \( N = 2 \) case no more than 2) of fields \( \sigma(\mathbf{r}, \tau_j) \) appear in the integrand for the same values of \( (\mathbf{r}, \tau_j) \).

This undoing of the Hubbard-Stratonovich transformation may seem a cumbersome or convoluted way to proceed, but it is useful in that it mechanically generates the correct expression for the \( N \)-body partition function for any particle content simply by differentiation of the fermion determinants. Moreover, this is accomplished without the need to deal with operator algebra and is easily generalized to bosons. In the 2-body case, in particular, the above procedure results in

\[ R_{ac,bd} = [M_2^{N_e}]_{ac,bd} , \]  

(40)

where we have naturally identified the transfer matrix in the two-particle subspace

\[ [M_2]_{ac,bd} = K_{ak}K_{cd} + (e^{r_g} - 1)I_{abcd} , \]  

(41)

and where

\[ K_{ij} = \sum_p T_{ip}T_{pj} , \]  

(42)

\[ I_{ijkl} = \sum_p T_{ip}T_{pj}T_{kp}T_{pl} . \]  

(43)

The form of the transfer matrix lends itself to a useful diagrammatic representation, which we show for the two- and three-particle cases (the latter derived in Appendix A) in Eqs. (44), (45) and (46).

\[ [M_2]_{ac,bd} = a \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad b + (e^{r_g} - 1) c \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad d \]  

(44)

\[ [M_3]_{abc,def} = a \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad c \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad d + (e^{r_g} - 1) b \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad e \]  

(45)

\[ \begin{array}{c} a \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad b \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad c \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad d \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad e \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad f \end{array} = b \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad c \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad d \]  

(46)

\[ \begin{array}{c} a \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad b \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad c \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad d \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad e \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad f \end{array} = \begin{array}{c} a \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad b \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad c \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad d \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad e \quad \begin{array}{c} \rightarrow \\ \rightarrow \end{array} \quad f \end{array} \]  

(47)

In practice, we compute the position-space wavefunction \( \xi(x_\uparrow, x_\downarrow) = (x_\uparrow, x_\downarrow)|\Xi\rangle \). Wavefunction in hand, we compute the matrix elements of the full density matrix \( \hat{\rho} \) as

\[ \langle x_\uparrow, x_\downarrow| \hat{\rho} | x'_\uparrow, x'_\downarrow \rangle = \langle x_\uparrow, x_\downarrow|\Xi\rangle \langle \Xi|x'_\uparrow, x'_\downarrow \rangle \]  

(48)

\[ = \xi^*(x'_\uparrow, x'_\downarrow) \xi(x_\uparrow, x_\downarrow) . \]  

(49)

From these, the elements of the reduced density matrix \( \hat{\rho}_A \) can be obtained as well. Given two states \( |s\rangle, |s'\rangle \in \mathcal{H}_A \) for the subregion \( A \), each state being specified by
For large imaginary times, we have
\[ \langle s | \hat{\rho}_A | s' \rangle = \sum_{a \in A_{s,s'}} (|s \otimes |a\rangle) \hat{\rho} (|s' \otimes |a\rangle), \]
where, at each fixed pair of two-particle states \( s, s' \), the sum is taken over all states \( |a\rangle \in \mathcal{H}_A \) such that the state \( |s \otimes |a\rangle \in \mathcal{H}_A \otimes \mathcal{H}_A \) (resp. \( |s' \otimes |a\rangle \in \mathcal{H}_A \otimes \mathcal{H}_A \)) is consistent with the first (resp. second) index of the matrix element being evaluated. We have denoted this set as \( A_{s,s'} \). From this matrix, we compute the entanglement spectrum \( \sigma(H_A) \), that is the spectrum of the entanglement Hamiltonian defined
\[ \hat{\rho}_A = e^{-\hat{H}_A}, \]
as well as the von Neumann and Rényi entanglement entropies.

**B. Lattice Monte Carlo approach to the many-body problem**

To address the many-body system, we implement the Monte Carlo version of the algorithm outlined above. The output of this algorithm, however, is not the ground-state wavefunction but rather the expectation value of the desired observable in a projected state. In our case, the observable is of course the entanglement entropy. To obtain it, crucial intermediate steps are required that go beyond conventional Monte Carlo approaches. We therefore outline the basic formalism first, and then proceed to explain the additional steps required to calculate \( S_{n,A} \).

1. **Basic formalism**

Beginning with a largely arbitrary many-body state \( |\Omega_0\rangle \), we evolve the state forward in imaginary time by an extent \( \beta \) via
\[ \langle \Omega(\beta) \rangle = e^{-\beta \hat{H}} |\Omega_0\rangle, \]
For large imaginary times, we have
\[ \langle \Omega(\beta) \rangle \xrightarrow{\beta \rightarrow \infty} |\Omega\rangle, \]
where \( |\Omega\rangle \) is the true ground state provided that \( \langle \Omega_0 | \Omega \rangle \neq 0 \).

For an operator \( \hat{O} \), we may obtain the ground-state expectation value by studying the asymptotic behavior of the function
\[ O(\beta) = \frac{1}{Z(\beta)} \langle \Omega(\beta/2) | \hat{O} | \Omega(\beta/2) \rangle, \]
with the zero-temperature normalization defined as
\[ Z(\beta) = \langle \Omega(\beta/2) | \Omega(\beta/2) \rangle = \langle \Omega_0 | e^{-\beta \hat{H}} | \Omega_0 \rangle. \]

As derived in detail earlier, we implement a symmetric factorization of the Boltzmann weight [c.f. Eq. (16)] in order to separate factors depending only on the one-body kinetic-energy operator from the significantly more complicated two-body potential-energy operator responsible for the effects of the interaction. Following this approximation, we again implement an auxiliary field transformation [c.f. Eq. (18)] to represent the interaction factor. This allows us to write the ground-state estimator of Eq. (54) defined above in path integral form as
\[ O(\beta) = \frac{1}{Z(\beta)} \int D\sigma \, P_\beta[\sigma] \, O_\beta[\sigma], \]
while simultaneously demonstrating that
\[ Z(\beta) = \int D\sigma \, P_\beta[\sigma]. \]

We have identified a naturally emerging probability measure \( P_\beta[\sigma] \) computed as
\[ P_\beta[\sigma] = \langle \Omega_0 | \hat{U}_\beta[\sigma] | \Omega_0 \rangle, \]
with the operator \( \hat{U}_\beta[\sigma] \) defined as in Eq. (22) (setting \( \mu = 0 \) in the kinetic energy factor since particle number is fixed in this formalism). The integrand takes the form
\[ O_\beta[\sigma] = \frac{\langle \Omega_0 | \hat{U}_\beta[\sigma] | \hat{O} \hat{U}_\beta[\sigma] | \Omega_0 \rangle}{\langle \Omega_0 | \hat{U}_\beta[\sigma] | \Omega_0 \rangle}. \]

Taking advantage of the arbitrariness of the initial state, we choose for \( |\Omega_0\rangle \) a Slater determinant for each fermion species constructed from single-particle plane-wave states \( \phi_j \) for \( 1 \leq j \leq N/2 \) with \( N/2 = N_+ = N_- \). With this assumption, we find that the probability takes the form
\[ P_\beta[\sigma] = \det^2 U_\beta[\sigma], \]
with
\[ [U_\beta[\sigma]]_{kk'} = \langle \phi_k | \hat{U}_\beta[\sigma] | \phi_{k'} \rangle, \]
where the indices \( k, k' \) satisfy \( 1 \leq k, k' \leq N/2 \).

2. **Path integral form of the reduced density matrix, replica fields, and the Rényi entropy**

It was shown by Grover in Ref. [56] that the reduced density matrix can be written in terms of the fermionic creation and annihilation operators \( \hat{c}^\dagger, \hat{c} \) as a weighted average with respect to the probability measure \( P_\beta[\sigma] \) derived above. Specifically,
\[ \hat{\rho}_{A,\beta} = \int D\sigma \, P_\beta[\sigma] \, \hat{\rho}_{A,\beta}[\sigma], \]
where
\[ \hat{\rho}_{A,\beta}[\sigma] = \det (\mathbb{1} - G_{A,\beta}[\sigma]) \times \exp \left( -\sum_{i,j \in A} \hat{c}_i^\dagger \left[ \log \left( G_{A,\beta}^{-1}[\sigma] - \mathbb{1} \right) \right]_{ij} \hat{c}_j \right), \]

\[ \sum_{i,j \in A} \]
It is important to note that \( \hat{\rho}_{A,\beta}[\sigma] \) is the reduced density matrix of a system of non-interacting fermions in the external field \( \sigma \). Expressions for non-interacting reduced density matrices were first derived in Refs. [66–68], but it was not until the much more recent work of Ref. [56] that those were combined into the non-perturbative form of Eq. (62) amenable to Monte Carlo calculations.

In the above, \( G_{A,\beta}[\sigma] \) is the spatial restriction of the (equal-time) one-body density matrix for either flavor to the region \( A \) computed as

\[
G_{A,\beta}[\sigma]_{rr'} = \sum_{a,b=1}^{N/2} [U_{\beta}^{-1}[\sigma]]_{ab} \phi_a^*(r', \beta/2) \phi_a(r', \beta/2),
\]

where

\[
\phi_a(r', \beta/2) = \langle r' | \hat{U}_{\beta}[\sigma] | \phi_a \rangle \quad (65)
\]

\[
\phi_a^*(r, \beta/2) = \langle \phi_a | \hat{U}_{\beta}[\sigma] | r \rangle. \quad (66)
\]

We suppress the imaginary-time \( \beta \) dependence in much of what follows with the understanding that calculations are to be performed in the limit of \( \beta \to \infty \).

From this decomposed form of the reduced density matrix, an estimator for the \( n \)-th order R\'enyi entropy can be derived. Because \( n \) powers of \( \hat{\rho}_A \) are needed, an equal number of auxiliary fields will appear (the “replica” fields), which we will denote collectively as \( \sigma \).

The final result (see Refs. [54–56, 69, 70]) takes the form

\[
\exp ((1-n)S_{n,A}) = \text{Tr}_{\hat{H}_A}[\rho^n_A] = \frac{1}{Z_n} \int \mathcal{D}\Sigma P[\sigma] Q[\sigma],
\]

where (note the suppressed \( \beta \) dependence)

\[
P[\sigma] = P[\sigma_1]P[\sigma_2] \ldots P[\sigma_n], \quad (68)
\]

with the observable being

\[
Q[\sigma] = \det^2 W[\sigma], \quad (69)
\]

with

\[
W[\sigma] = \prod_{j=1}^{n} \left( \mathbb{1} - G_A[\sigma_j] \right) \left[ \mathbb{1} + \prod_{k=1}^{n} \frac{G_A[\sigma_k]}{1 - G_A[\sigma_k]} \right]. \quad (70)
\]

We have adopted a notation such that, for functions or integrals of functions of multiple auxiliary fields, we write

\[
F[\sigma] = F[\sigma_1, \sigma_2, \ldots, \sigma_n], \quad (71)
\]

and

\[
\int \mathcal{D}\Sigma F[\sigma] = \int \mathcal{D}\sigma_1 \mathcal{D}\sigma_2 \ldots \mathcal{D}\sigma_n F[\sigma], \quad (72)
\]

respectively.

Equation (70) poses the challenging task of inverting \( \mathbb{1} - G_A \), which can be very nearly singular, as pointed out in Ref. [69]. For \( n = 2 \), no inversion is required, because the equations simplify such that

\[
Q[\sigma] = \det^2 \left[ (\mathbb{1} - G_A[\sigma_1])(\mathbb{1} - G_A[\sigma_2]) + G_A[\sigma_1]G_A[\sigma_2] \right]. \quad (73)
\]

However, for higher \( n \) there is no simplification of that kind and therefore it is less clear how one may avoid the problem. We solved this problem in Ref. [55] (see also [71–74]); the main point is realizing that

\[
det W[\sigma] = \det L[\sigma] \det K[\sigma], \quad (74)
\]

where \( L[\sigma] \) is a block diagonal matrix (one block per replica \( k \)):

\[
L[\sigma] \equiv \text{diag} \left[ \mathbb{1} - G_A[\sigma_k] \right], \quad (75)
\]

and

\[
K[\sigma] \equiv \begin{pmatrix}
R[\sigma_1] & 0 & 0 & \ldots & 0 \\
0 & R[\sigma_2] & 0 & 0 & 0 \\
\vdots & \ddots & \ddots & \ddots & \vdots \\
0 & \ldots & \ldots & 0 & R[\sigma_{n-1}] \\
\end{pmatrix},
\]

where

\[
R[\sigma_k] = \frac{G_A[\sigma_k]}{G_A[\sigma_k] - \mathbb{1}}. \quad (77)
\]

Within the determinant of Eq. (74), we multiply \( K[\sigma] \) and \( L[\sigma] \) and define

\[
T[\sigma] \equiv K[\sigma]L[\sigma] = \mathbb{1} - D \mathcal{G}[\sigma], \quad (78)
\]

where \( \mathcal{G}[\sigma] \) is a block diagonal matrix defined by

\[
\mathcal{G}[\sigma] = \text{diag} \left[ G_A[\sigma_k] \right], \quad (79)
\]

and

\[
D \equiv \begin{pmatrix}
\mathbb{1} & 0 & 0 & \ldots & -\mathbb{1} \\
\mathbb{1} & \mathbb{1} & 0 & \ldots & 0 \\
0 & \mathbb{1} & \mathbb{1} & \ldots & 0 \\
\vdots & \ddots & \ddots & \ddots & \vdots \\
0 & \ldots & \ldots & \ldots & \mathbb{1} \\
\end{pmatrix}. \quad (80)
\]

Equation (83) is the result that allows us to bypass the inversion of \( \mathbb{1} - G_A \). Moreover, the form of \( T[\sigma] \) is clearly simpler than that of \( W[\sigma] \). For those reasons we use \( T[\sigma] \) in all of the many-body calculations presented here. This formulation allowed us to study R\'enyi entropies as high as \( n = 5 \); higher are also possible.

For completeness, we present here the simplification for the bosonic case as well (and add a subindex \( B \) accordingly), for which

\[
Q_B[\sigma] = \det^{-2} W_B[\sigma], \quad (81)
\]
and

$$W_B[\sigma] = \prod_{j=1}^{n} (\mathbb{1} + G_A[\sigma_j]) \left[ \mathbb{1} - \prod_{k=1}^{n} \frac{G_A[\sigma_k]}{\mathbb{1} + G_A[\sigma_k]} \right].$$ (82)

The analogous strategy to avoid inversion leads here to

$$T_B[\sigma] \equiv \mathbb{1} - D_B G[\sigma],$$ (83)

where $G[\sigma]$ is a block diagonal matrix defined by

$$G[\sigma] = \text{diag}[G_A[\sigma_n]],$$ (84)

and

$$D_B \equiv \begin{pmatrix} -\mathbb{1} & 0 & \ldots & 0 \\ \mathbb{1} & -\mathbb{1} & \ldots & 0 \\ 0 & \mathbb{1} & -\mathbb{1} & \ldots \\ \vdots & \vdots & \ddots & \vdots \\ 0 & 0 & \ldots & -\mathbb{1} \end{pmatrix}. $$ (85)

3. Signal-to-noise issues and how to overcome them

The path integral form of the Rényi entropy Eq. (67) has a deceptively simple form: It seems obvious that one should interpret $P[\sigma]$ as the probability density and $Q[\sigma]$ as the observable being averaged. This, in some sense, is a trap: while $Q[\sigma]$ is crucially sensitive to correlations among the replica fields $\sigma_k$, $P[\sigma]$ completely factorizes across replicas (i.e. it is insensitive to said correlations). As a consequence, a Monte Carlo implementation sampling $\sigma$ according to $P[\sigma]$ will give outlandish values of $Q[\sigma]$ that fluctuate wildly and may not converge to the expected value. This feature is what in the lattice QCD area is often called an overlap problem (see e.g. Refs [75, 76]). The present case is especially challenging in 2D and 3D, as the magnitude of $Q[\sigma]$ is expected to grow exponentially with the size of the boundary of the subregion $A$ (see e.g. [71, 72]).

Motivated by the similarity between the numerator of Eq. (67) and the conventional path-integral form of partition functions, we address the overlap problem by first differentiating with respect to a parameter, then using Monte Carlo methods to compute that derivative, and finally integrating at the end. We outline this procedure in detail in Ref. [55], and reproduce part of it here.

We introduce a parameter $0 \leq \lambda \leq 1$ by defining a function $\Gamma(\lambda; g)$ such that

$$\Gamma(\lambda; g) \equiv \int D\Sigma P[\sigma] Q^\lambda[\sigma].$$ (86)

Normalization of $P[\sigma]$ implies that

$$\ln \Gamma(0; g) = 0,$$ (87)

while Eq. (67) implies

$$\ln \Gamma(1; g) = (1 - n) S_{n,A}. $$ (88)

Using Eq. (86),

$$\frac{\partial \ln \Gamma}{\partial \lambda} = \int D\Sigma \tilde{P}[\sigma; \lambda] \ln Q[\sigma],$$ (89)

where

$$\tilde{P}[\sigma; \lambda] \equiv \frac{1}{\Gamma(\lambda; g)} P[\sigma] Q^\lambda[\sigma]$$ (90)

is a well-defined, normalized probability measure which features the usual weight $P[\sigma]$ as well as an entanglement contribution $Q^\lambda[\sigma]$. It is the latter factor that induces entanglement-specific correlations in the sampling of $\sigma$ when probability $\tilde{P}[\sigma; \lambda]$.

Thus, $S_{n,A}$ is calculated by using $\lambda = 0$ as a reference point and computing $S_{n,A}$ via

$$S_{n,A} = \frac{1}{1 - n} \int_0^1 d\lambda \langle \ln Q[\sigma] \rangle_\lambda,$$ (91)

where

$$\langle X \rangle_\lambda = \int D\Sigma \tilde{P}[\sigma; \lambda] X[\sigma].$$ (92)

We thus obtain an integral form of the interacting Rényi entropy that can be computed using any MC method (see e.g. [63–65]), in particular hybrid Monte Carlo [77, 78] to tackle the evaluation of $\langle \ln Q[\sigma] \rangle_\lambda$ as a function of $\lambda$. In practice, we find that $\langle \ln Q[\sigma] \rangle_\lambda$ is a smooth function of $\lambda$, as exemplified in Fig. 4. It is therefore sufficient to perform the numerical integration using a uniform grid.

IV. RESULTS: TWO-BODY SYSTEM

We solve the two-body problem via the projection method outlined previously, which furnishes the full two-body wavefunction on the lattice. We ensure that the
continuum limit is approached by solving the problem for multiple lattice sizes, and by computing the renormalized coupling using the energy spectrum and Lüscher’s formalism [57, 58]. The latter indicates that the relationship between the energy eigenvalues and the scattering phase shift \( \delta(p) \) is given by

\[
p \cot \delta(p) = \frac{1}{\pi L} S(\eta)
\]

where \( \eta = \frac{pL}{2\pi} \) and \( L \) is the box size, such that the energy of the two-body problem is \( E = p^2/m \); and

\[
S(\eta) \equiv \lim_{\lambda \to \infty} \left( \sum_n \Theta(\Lambda^2 - n^2) - 4\pi\Lambda \right),
\]

where the sum is over all 3D integer vectors, and \( \Theta(x) \) is the Heaviside function. In turn, the scattering phase shift determines the scattering parameters via

\[
p \cot \delta(p) = -\frac{1}{a} + \frac{1}{2} r_{\text{eff}} p^2 + O(p^4),
\]

where \( \delta \) is the scattering phase shift, \( a \) is the scattering length, and \( r_{\text{eff}} \) is the effective range.

A. Low-lying entanglement spectrum

Once the matrix elements of \( \hat{\rho}_A \) are calculated from the projected ground state, as shown above, we obtain the eigenvalues using standard diagonalization routines to obtain the entanglement spectrum \( \sigma(\hat{H}_A) \), which is defined as the spectrum of the entanglement Hamiltonian

\[
\hat{H}_A, \quad \rho_A = e^{-\hat{H}_A}.
\]

In Fig. 5, we present our results for \( \sigma(\hat{H}_A) \) for a cubic subregion \( A \) of linear size \( L_A/L = 0.5 \), for two particles in the BCS-BEC crossover, parametrized by the dimensionless coupling \( (k_F a)^{-1} \), where \( k_F \) is the Fermi momentum (merely a measure of the particle density in the periodic box, as for two particles there is of course no Fermi surface) and \( a \) is the s-wave scattering length. The latter was determined using the Lüscher formalism outlined above.

The main features of \( \sigma(\hat{H}_A) \) can be described as follows. We note first that beyond the lowest 4 or 5 eigenvalues, shown as \( \lambda_1 \) to \( \lambda_5 \) in the bottom panel of Fig. 5, the multiplicity of eigenvalues grows dramatically, forming a quasi-continuum. For this reason, we focus here on the lowest 5 eigenvalues and characterize the rest statistically in the next section. As is evident from the figure, the dependence of all \( \lambda_k \) on \( (k_F a)^{-1} \) is rather mild and smooth, although it has at a few crisp features: there is a rather large gap between \( \lambda_1 \) and the next eigenvalue, which implies that the Rényi entanglement entropies are dominated by that eigenvalue; there is a crossing of \( \lambda_2 \), \( \lambda_3 \) and \( \lambda_4 \) on the BEC side of the resonance; after that crossing \( \lambda_2 \) and \( \lambda_3 \) heal to \( \lambda_5 \) and effectively merge into the lower edge of the quasi-continuum part of the spectrum. The evolution of these properties along the crossover is shown in detail in panels a – e of Fig. 5.

In Fig. 6 we show the Schmidt gap \( \Delta \) (see Refs [79]), defined as the separation between the two largest eigenvalues of the reduced density matrix \( \hat{\rho}_A \), for \( L_A/L = 0.1, 0.2, ..., 0.5 \), as a function of \( (k_F a)^{-1} \). Since we do not expect a quantum phase transition as a function of \( (k_F a)^{-1} \), we similarly do not expect the Schmidt gap to vanish. As a result of the eigenvalue crossing explained above, however, there exists a sharp change (in the sense of a discontinuous derivative) in \( \Delta \) in the BCS-BEC crossover, which takes place in the strongly coupled
region $0 < (k_F a)^{-1} < 1$. It is also evident that, because $\lambda_1$ and $\lambda_4$ track each other at a very nearly constant separation, the Schmidt gap becomes constant to the right of the sharp edge in Fig. 6. As with other features of this spectrum, it remains to be determined how $\Delta$ evolves as a function of particle number, in particular as a Fermi surface forms and Cooper pairing correlations emerge.

As mentioned above, our calculations were carried out in a periodic box. We show the corresponding size effects in Fig. 7, where we show the entanglement spectrum of the two-body system as a function of the bare lattice coupling $g$. In that figure, it is clear that finite-size effects are smallest on the BCS side of the resonance, but become considerably more important on the BEC side. This is consistent with the expectation that, once a two-body bound state forms (as the coupling is increased away from the non-interacting point), the sensitivity to lattice-spacing effects is enhanced. It is noteworthy, in particular, that one may identify the unitary regime just by looking at this figure: for any given eigenvalue, the data for different lattice sizes crosses at about the same value of $g$; this is reminiscent of the finite-size scaling behavior of order parameters in critical phenomena, as it is the hallmark of scale invariance at phase transitions.

The process of reducing finite-size effects, at fixed particle number, implies approaching the dilute limit, i.e. using larger lattices. When that limit is approached, the renormalization prescription that replaces $g$ with the physical coupling $(k_F a)^{-1}$ (described above) should force the finite-size calculations to collapse to a single, universal (in the sense of size-independent) curve. This is indeed what we find and what yields the results of Fig. 5.

FIG. 7: Entanglement spectrum of the two-body problem in the BCS-BEC crossover as a function of the bare lattice coupling at different lattice sizes: solid, dashed, dotted, dash-dotted, for $N_f = 4, 6, 8, 10$, respectively. The subsystem size was fixed to $L_A/L = 0.5$. The coupling corresponding to the unitary point is marked with a vertical dashed line. Note how different volumes cross precisely at unitarity, which reflects the property of scale invariance.

FIG. 8: Top: Histogram of the high entanglement spectrum of the two-body problem showing number of counts (cts.) as a function of the coupling $(k_F a)^{-1}$ and the entanglement eigenvalue $\lambda_k$, for region size $L_A/L = 1/2$. The dashed line shows the dependence of the mean (see also middle plot). Middle and bottom: Mean and standard deviation, respectively, of the high entanglement spectrum distribution, as a function of the interaction strength $(k_F a)^{-1}$ (main) and $k_F a$ (inset). In each plot the different curves show results for various $L_A/L$. Note that the weak coupling limit corresponds to $k_F a \to 0^+$.

B. High entanglement spectrum

As mentioned in the previous section, the entanglement spectrum $\sigma(H_A)$ above $\lambda_5$, which we will refer to here as the high entanglement spectrum, displays a rapidly growing multiplicity of eigenvalues which we
deem best to analyze using elementary statistical methods. In Fig. 8, we show the eigenvalue distribution of the high entanglement spectrum for different system sizes, in histogram form. More importantly, we find that the mean and standard deviation of that distribution, shown here in Fig. 8 (middle and bottom), are smooth functions of $(k_Fa)^{-1}$; the mean, in particular, diverges as the coupling is turned off. We interpret this effect as strong evidence that the high sector of $\sigma(H_A)$ is a non-perturbative component of $H_A$ that is entirely due to quantum fluctuations induced by the interaction. Although the two-body system has no Fermi surface, it seems natural to conjecture a link between Cooper pairing and the high entanglement spectrum. Determining whether this is true, however, is a challenging problem that requires studying the high entanglement spectrum in the progression from few to many particles.

Our numerical calculations show a large number of eigenvalues that lie far (at least 9 to 10 orders of magnitude) above the high entanglement spectrum. While we cannot discard that those eigenvalues are consistent with numerical noise (they come from the lowest eigenvalues of the reduced density matrix), there are enough of them to warrant this brief comment. Although there is a large number of such eigenvalues, their contribution to the entanglement entropy is considerably suppressed by their small magnitude. We agree to add this discussion below.

**C. Entanglement entropy**

Using our knowledge of the eigenvalues $\lambda_k \in \sigma(H_A)$, the entanglement entropy of the two-body problem is easily determined. Indeed, the von Neumann entropy is

$$S_{\text{vN},A} = -\text{Tr} H_A [\hat{\rho}_A \ln \hat{\rho}_A] = \sum_k \lambda_k e^{-\lambda_k},$$  

(97)

and the $n$-th order Rényi entanglement entropy is

$$S_{\text{n},A} = \frac{1}{1-n} \ln \text{Tr} H_A [\hat{\rho}_A^n] = \frac{1}{1-n} \ln \sum_k e^{-n\lambda_k}.$$  

(98)

In Fig. 9 (top panel), we show $S_2$ as a function of $x = k_FL_A$ and for several values of the coupling $(k_Fa)^{-1}$. Remarkably, the trend towards the leading asymptotic behavior proportional to $x^2 \ln x$ appears to set in at $x \approx 2$ for all couplings. This is surprising, as there is no obvious reason for this to be the case. As we will see below, we find the same kind of behavior for the many-body Fermi gas at resonance.

To show explicitly the effect of the high entanglement spectrum on $S_2$, which we referred to in the previous section, we show in Fig. 9 (bottom panel) the contribution $\Delta S_2$ of the first entanglement eigenvalue to the full $S_2$. It is clear in that plot that the contribution is at most on the order of 8% for the parameter ranges we studied.

![FIG. 9: Top: Second Rényi entanglement entropy $S_2$ of the two-body problem as a function of $x = k_FL_A$ and for several values of the coupling $(k_Fa)^{-1}$. Inset: $S_2$ scaled by $x^2 \ln x$. Bottom: Relative contribution of the high entanglement spectrum to the second Rényi entanglement entropy $S_2$, as a function of $x = k_FL_A$.](image)

**V. RESULTS: MANY-BODY SYSTEM**

Using the many-body lattice Monte Carlo techniques described above, along with the tuning procedure outlined in the previous section, we computed several entanglement entropies of the unitary Fermi gas, aiming to characterize its leading and sub-leading asymptotic behavior as a function of the subregion size $x = k_FL_A$.

The results shown throughout this section were obtained by gathering 250 decorrelated auxiliary field configurations (where a single “auxiliary field” contains all the replicas required to determine the desired Rényi entropy) for each value of the auxiliary parameter $\lambda$. We used particle numbers in the range $N = 4 - 400$ and cubic lattice sizes in the range $N_x = 6 - 16$ with periodic boundary conditions. The projection to the ground state was carried out by extrapolation to the limit of large imaginary-time direction. The auxiliary parameter $\lambda$ was discretized using $N_\lambda = 10$ points, which we found to be enough to capture the very mild dependence on that parameter, as explained in a previous section (see also Appendix B for further details).

Because the methods we implemented impose a dis-
cretization of spacetime, special attention was given to the ordering of the scales, to ensure that the thermodynamic and continuum limit were approached. Specifically, we required the following ordering:

\[ k_F \ell \ll 1 \ll k_F L_A \ll k_F L, \quad (99) \]

where \( \ell \) is the lattice spacing, \( L_A \) is the subsystem size, and \( L = N_x \ell = N_z \) is the full system size. The first condition on the left of Eq. (99) ensures that the continuum limit is approached; the second condition implies that the region determined by \( L_A \) must contain many particles (since the density is the only scale in the system, this condition defines the large-\( L_A \) regime); and the last condition means that \( L_A \ll L \), to ensure finite-size effects are minimized. This ordering was accomplished by carefully choosing the restrictions on \( L_A \) for a given particle number \( N \), while aiming to maintain a large \( N \). The latter, however, requires \( L \) to be large in order to avoid high densities where \( k_F \approx 1 \), which can be sensitive to lattice-spacing effects. In addition, we set \( L_A \leq 0.45L \) as a compromise to satisfy the last inequality.

In Fig. 10 we show our results of the second Rényi entropy \( S_2 \) of the unitary Fermi gas in volumes of \( N_x^3 \) lattice points, where \( N_x = 6 - 16 \), as a function of \( x = k_F L_A, \) for cubic subsystems of side \( L_A \). Within the statistical uncertainty, shown in colored bands, the results for different volumes coincide, which indicates that our results are in the continuum and thermodynamic regimes.

The inset of Fig. 10 shows \( S_2 \) scaled by \( x^2 \) in a semi-log plot. The fact that the trend is clearly linear supports the assertion that \( x \) is large enough to discern the asymptotic regime, where \( S_2/x^2 \propto \ln x \). As in the case of the non-interacting Fermi gas, mentioned in the Introduction, this onset of the asymptotic regime appears to be at \( x \approx 2 \).

### A. Rényi entanglement entropies

Using the formalism presented above for the determination of Rényi entanglement entropies for \( n \geq 2 \), we computed \( S_{n,A} \) for the resonant Fermi gas for \( n = 2, 3, 4, 5 \), as a function of \( x = k_F L_A \). In Fig. 11 we show our main results. To interpret those results, we briefly discuss the noninteracting case. In Refs. [49–53] it was shown that the leading-order behavior or the entanglement entropy of non-interacting 3D fermions as a function of \( x \) is given by

\[ S_{n,A}(x) = c(n)x^2 \ln x + o(x^2), \quad (100) \]

where

\[ c(n) = \frac{1 + n^{-1}}{24(2\pi)^d-1} \int_{\partial \Omega} \int_{\partial S_x} dS_x dS_k |\hat{n}_x \cdot \hat{n}_k| \quad (101) \]

where \( \Omega \) is the real-space region \( A \) scaled to unit volume with normal \( \hat{n}_z \), \( \Sigma \) is the Fermi volume scaled by the Fermi momentum with unit normal \( \hat{n}_A \). In our case, \( A \) is a cubic subsystem (as in Fig. 1) and a spherical Fermi volume.

The noninteracting case is shown in Fig. 11 in two ways. The asymptotic result at large \( x \) is shown with crosses on the right edge of the plot, extended into the plot (as a visual aid) with dashed black horizontal lines for \( n = 2, 3, 4, 5 \) (top to bottom). With a thick red dashed line we show the case \( n = 2 \) at finite \( x \), as obtained with the overlap-matrix method [60].

Our results for \( S_{n,A} \) for the unitary Fermi gas (data points with error bands) appear to heal to the noninteracting limit when the slow decay (see below) to a constant at large \( x \) is taken into account; this statement holds especially in the \( n = 2 \) case where the sub-leading oscillations allow for a relatively clean fit. Indeed, our fits for \( n = 2 \) give

\[ S_{2,A}(x) = ax^2 \ln x + bx^2, \quad (102) \]

with \( a = 0.114(2) \) and \( b = 0.04(1) \), while Eq. (101) yields \( c(2) = 3/(8\pi) \approx 0.119366 \ldots \). While \( c(2) \) are different to within our uncertainties, they are surprisingly close (between 3 and 6\%). The sub-leading behavior is consistent with an area law \( \alpha \propto x^2 \). As \( n \) is increased, sub-leading oscillations become increasingly apparent; however, they are mild enough that it is still possible to discern the asymptotic behavior at large \( x \). For \( n = 3, 4, 5 \), oscillations notwithstanding, the results in the large-\( x \) limit appear again to be close to the noninteracting case.

Using our results for the entanglement entropies \( S_n \) as a function of \( n \), it is possible to use the power method to extract the lowest eigenvalue \( \lambda_1 \) of the entanglement
spectrum as a function of $x$. We studied the decay of $(1 - n)S_n/n$ to a constant value which, given Eq. (98), we identified as $-\lambda_1$. In Fig. 12 we show the result of using that sole eigenvalue to approximate $S_n$. As expected, higher orders $n$ emphasize the contribution from the lowest entanglement eigenvalue (highest eigenvalue of the reduced density matrix), which progressively dominates $S_n$ as $n$ increases. From the $n$ dependence of $S_n$, it is also possible to study the degeneracy of the lowest entanglement eigenstate; at large $n$,

$$\frac{(1 - n)}{n} S_n \simeq \frac{\ln d_1}{n} - \lambda_1 + \ldots,$$

(103)

where the ellipsis indicates exponentially suppressed terms, and $d_1$ is the degeneracy associated with $\lambda_1$. We find a vanishing first term, which indicates that $d_1$ is consistent with unity.

VI. SUMMARY AND CONCLUSIONS

We implemented two different lattice methods to characterize non-perturbatively the entanglement properties of three-dimensional spin-1/2 fermions in the strongly interacting, resonant regime of short interaction range and large scattering length, i.e. the unitary limit. This regime is scale invariant (in fact, non-relativistic conformal invariant) in the sense that it presents as many scales as non-interacting gases and therefore its properties are universal characteristics of three-dimensional quantum mechanics, i.e. in the same sense as critical exponents that characterize phase transitions.

We analyzed the two-body spectrum of the entanglement Hamiltonian along the BCS-BEC crossover and presented results for the low-lying part, which displays clear features as the strength of the coupling is varied, such as eigenvalue crossing close to the resonance point and merging in the BEC limit. The lowest two eigenvalues in the spectrum correspond to the largest two eigenvalues of the reduced density matrix, which are separated by the Schmidt gap. We found that the latter displays a sharp change at strong coupling, in the vicinity of the conformal point $(k_F a)^{-1} = 0$.

We also carried out a statistical characterization of the high entanglement spectrum, which appears as a quasi-continuum distribution with well defined mean and standard deviation, which we mapped out along the crossover. We found that the mean of the distribution tends to infinity in the noninteracting limit, which indicates that that sector is due to non-perturbative effects in the entanglement Hamiltonian. In contrast, the low-lying part of the spectrum has a finite noninteracting limit. All of the above two-body results were obtained with non-perturbative non-stochastic methods which are easily generalizable to higher particle numbers (as we show analytically and diagrammatically for 3 particles).

In addition, we studied the Rényi entropies of degree $n = 2, 3, 4, and 5$ of many fermions in the unitary limit, which we calculated using a method recently developed by us (based on an enhanced version of the algorithm of Ref. [56]). We found that, remarkably, the large $x = k_F L_A$ (i.e. subsystem size) limit for those entanglement entropies sets in for $x$ as low as 2.0, which allowed us to characterize the leading and sub-leading asymptotic
behavior using $2 \leq x \leq 10$. For entropies of order $n > 2$, on the other hand, we found that sub-leading oscillations are enhanced, but not enough to spoil the visualization of the asymptotic behavior at large $x$.

Our experience with Monte Carlo calculations of $S_{n,A}$ in 1D gave us empirical indication that the entanglement properties of the unitary Fermi gas might not be too different from those of a non-interacting gas. However, since unitarity corresponds to a strongly correlated, three-dimensional point, that intuition could very well have been wrong. Our calculations indicate that the leading-order asymptotic behavior is approximately consistent with that of a non-interacting system, while the sub-leading behavior may be obtained by suitable index contraction.

The recent measurement of the second Rényi entropy of a bosonic gas in an optical lattice [19, 20] shows that intuition could very well have been wrong. Our calculations indicate that the leading-order asymptotic behavior is approximately consistent with that of a non-interacting system, while the sub-leading behavior is clearly different.

The recent measurement of the second Rényi entropy of a bosonic gas in an optical lattice [19, 20] shows that it is possible to experimentally characterize the entanglement properties of the kind of system analyzed here. Our calculations are therefore predictions for such experiments for the case of fermions tuned to the unitary limit.

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Appendix A: Exact evaluation of the path integral for finite systems

In order to illustrate the details as well as the generality of this technique, we evaluate the path integral for a four-component tensor from which each of the above traces may be obtained by suitable index contraction.

To this end, we define

$$ R_{ac,bd} = \int \mathcal{D}\sigma \, U[\sigma]_{ab} U[\sigma]_{cd}. $$

(A1)

We first write out each of the matrices $U[\sigma]$ in its product form. That is, we reintroduce the expression

$$ U[\sigma] = \prod_{j=1}^{N_r} U_j[\sigma]. $$

(A2)

For each contribution to the $N$-body transfer matrix, exactly $N$ factors of the matrix $U[\sigma]$ appear, and as a result each temporal lattice point appears in the integrand $N$ times. Writing out the integrand and grouping by times-lattice, we obtain

$$ R_{ac,bd} = \int \mathcal{D}\sigma \, U[\sigma]_{ab} U[\sigma]_{cd} = \int \mathcal{D}\sigma \left( U_1[\sigma] U_2[\sigma] \ldots U_{N_r}[\sigma] \right)_{ab} \left( U_1[\sigma] U_2[\sigma] \ldots U_{N_r}[\sigma] \right)_{cd} $$

$$ = \sum_{k_1,k_2,\ldots,k_{N_r-1}} \int \mathcal{D}\sigma \left( U_1[\sigma]_{ak_1} U_1[\sigma]_{d_1} \right) \left( U_2[\sigma]_{k_1k_2} U_2[\sigma]_{l_1l_2} \right) \ldots \left( U_{N_r}[\sigma]_{k_{N_r-1}d} U_{N_r}[\sigma]_{l_{N_r-1}d} \right) $$

$$ = \sum_{k_1,k_2,\ldots,k_{N_r-1}} \prod_{l_1,l_2,\ldots,l_{N_r-1}} \left( \int \mathcal{D}\sigma(\tau_j) \, U_j[\sigma]_{k_{j-1}k_j} U_j[\sigma]_{l_{j-1}l_j} \right), $$

(A3)

where we set $k_0 = a$, $l_0 = c$, $k_{N_r} = b$, and $l_{N_r} = d$, and used the notation

$$ \mathcal{D}\sigma(\tau) \equiv \prod_r \frac{d\sigma(x,\tau)}{2\pi}. $$

(A6)

Using the specific form of the individual $U$ factors, we find

$$ \int \mathcal{D}\sigma(\tau_j) U_j[\sigma]_{k_{j-1}k_j} U_j[\sigma]_{l_{j-1}l_j} = \sum_{p,q} \int \mathcal{D}\sigma(\tau_j) \left( T_{kj-1}p V_j[\sigma]_{pq} T_{kj}q \right) \left( T_{j_{j-1}}p' V_j[\sigma]_{p'q'} T_{j_{j-1}}l' \right), $$

(A7)

which using our chosen form of $V$ becomes

$$ = \sum_{p,q} T_{kj-1}p T_{kj}q T_{j_{j-1}}p' T_{j_{j-1}}l' \delta_{pq} \delta_{p'q'} \times $$

$$ \int \mathcal{D}\sigma(\tau_j) \left( 1 + A \sin \sigma(p,\tau_j) \right) \left( 1 + A \sin \sigma(p',\tau_j) \right) $$

$$ = \sum_{p,q} T_{kj-1}p T_{kj}q T_{j_{j-1}}p' T_{j_{j-1}}l' \delta_{pq} \delta_{p'q'} \left( 1 + (e^{r \theta} - 1) \delta_{pp'} \right), $$

where $A = \frac{\pi}{2}$.
where we used

\[
\int D\sigma (\tau_j)(1 + A \sin \sigma (p, \tau_j)) (1 + A \sin \sigma (p', \tau_j)) = (1 + (e^{\tau g} - 1)\delta_{pp'}) . \tag{A8}
\]

Thus, we arrive naturally at the definition

\[
[M_2]_{ac,bd} = K_{ad}K_{cd} + (e^{\tau g} - 1)I_{abcd}, \tag{A9}
\]

as the transfer matrix in the two-particle subspace, where

\[
K_{ij} = \sum_p T_{ip}T_{pj}, \tag{A10}
\]

\[
I_{ijkl} = \sum_p T_{ip}T_{pj}T_{kp}T_{pl}. \tag{A11}
\]

Indeed, this definition of \( M_2 \) as a transfer matrix makes sense, because

\[
R_{ac,bd} = \sum_{k_1,k_2,...,k_{N_r-1}} \prod_{j=1}^{N_r} [M_2]_{kj-1,k_j-1,k_j,l_j}, \tag{A12}
\]

or more succinctly,

\[
R_{ac,bd} = [M_2]_{ac,bd}^{N_r} . \tag{A13}
\]

In a similar fashion, one may show without much difficulty that the transfer matrix of the three-body problem (for distinguishable particles, i.e. no symmetrization or antisymmetrization is enforced) is

\[
[M_3]_{abc,def} = K_{ad}K_{be}K_{cf} + (e^{\tau g} - 1)J_{abc,def}, \tag{A14}
\]

where

\[
J_{ijk,lmn} = K_{il}I_{jkmn} + K_{jm}I_{ikln} + K_{kn}I_{ijlm}. \tag{A15}
\]

The pattern from this point on is clearly visible: there is one term for each ‘spectator’ particle that does not participate in the interaction, while the other two are accounted for by an interacting term governed by the \( I_{abcd} \) object. One may thus infer the form of the transfer matrix for higher particle numbers.

**Appendix B: Auxiliary parameter dependence**

In this Appendix we show a few more examples on the mild dependence of the entanglement-entropy derivative \( \langle \ln \mathcal{Q}(\sigma) \rangle_\lambda \), as other parameters are varied. In all cases, the data shown corresponds to full 3D calculations in the unitary regime.

In Fig. 13 (top) we show the variation of that derivative when the Rényi order is changed from \( n = 2 \) to \( n = 5 \), at fixed particle number and region size. In Fig. 13 (bottom) we show how \( \langle \ln \mathcal{Q}(\sigma) \rangle_\lambda \) changes when the particle number is varied, at fixed Rényi order \( n \).

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FIG. 13: Top: $\lambda$ dependence of $\langle \ln Q[\sigma] \rangle_\lambda$ for several Rényi orders $n = 2, 3, 4, 5$, for subsystem size $L_A = 5/12 L$, for $N = 172$ fermions at unitarity in a box of size $L = N_x \ell$ (where $N_x = 12$ points and $\ell = 1$). Bottom: $\lambda$ dependence of $\langle \ln Q[\sigma] \rangle_\lambda$ for several subsystem sizes $L_A$, for $N = 136$ fermions at unitarity in a box of size $L = N_x \ell$ (where $N_x = 12$ points and $\ell = 1$), and for Rényi order $n = 2$.

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