Dynamics of disentanglement, density matrix and coherence in neutrino oscillations.

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Abstract

In charged current weak interaction processes, neutrinos are produced in an entangled state with the charged lepton. This correlated state is disentangled by the measurement of the charged lepton in a detector at the production site. We study the dynamical aspects of disentanglement, propagation and detection, in particular the conditions under which the disentangled state is a coherent superposition of mass eigenstates. The appearance and disappearance far-detection processes are described from the time evolution of this disentangled “collapsed” state. The familiar quantum mechanical interpretation and factorization of the detection rate emerges when the quantum state is disentangled on time scales much shorter than the inverse oscillation frequency, in which case the final detection rate factorizes in terms of the usual quantum mechanical transition probability provided the final density of states is insensitive to the neutrino energy difference. We suggest possible corrections for short-baseline experiments. If the charged lepton is unobserved, neutrino oscillations and coherence are described in terms of a reduced density matrix obtained by tracing out an unobserved charged lepton. The diagonal elements in the mass basis describe the production of mass eigenstates whereas the off diagonal ones provide a measure of coherence. It is shown that coherences are of the same order of the diagonal terms on time scales up to the inverse oscillation frequency, beyond which the coherences oscillate as a result of the interference between mass eigenstates.

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I. INTRODUCTION

Neutrino masses, mixing and oscillations are the clearest evidence yet of physics beyond the standard model [1–3]. They provide an explanation of the solar neutrino problem [4–6] and have other important phenomenological [1–3, 7–10], astrophysical [5, 11, 12] and cosmological [13] consequences. Furthermore, they offer an incredible example of macroscopic quantum coherence, where in long baseline oscillation experiments this coherence is maintained over hundreds of kilometers.

Given the importance of neutrino oscillations in pointing us to the next layer beyond that of the standard model, it is not unreasonable to make sure we understand the domain of validity of the various calculations of oscillation probabilities. In particular, the standard approach of treating neutrino oscillations by analogy with Rabi-oscillations in a two state system (see for example, [1–3, 7, 8] and references therein) while simple and intuitive, is not necessarily the last word on the subject. Deeper investigations of this basic paradigm have already raised a number of important and fundamental questions [14–17] that are still being debated [18–21].

If we truly want to understand the results from oscillation experiments, we need to factor in both the production and detection mechanisms. In term of its detection, the neutrino is never directly detected. We only know that it is there due to the charged leptons associated with its weak interaction at the detector. As for the production mechanism, the neutrino state is produced by the decay of a parent particle via charged current interactions. Interestingly, a discussion of coherence aspects of neutrinos [22–27] has recognized that the neutrino state produced by the decay of a parent particle via charged current interactions is in fact entangled with that of the charged lepton.

How does this entangled state evolve? As usual, there are two choices. If we measure the charged lepton, this disentangles the quantum state at a time determined either by the detection process of the charged lepton at the source or by stopping at a “beam dump” near the source. This disentangled neutrino state then evolves further in time and is detected at the “far detector” via another charged lepton. Therefore the production, disentanglement and detection involve two times: the time at which the charged lepton is observed at the source (if observed at all) or stopped at a beam dump near the source, which is when the neutrino/charged lepton state is disentangled, and the time when the neutrino is observed via a charged lepton in the far detector. We note that for long-baseline experiments these two time scales are widely separated.

If the charged lepton is not observed, the state is described by a reduced density matrix, which is obtained by tracing out the charged lepton. In this case, we can track the behavior of the off-diagonal elements, i.e., the coherences which contain information on the interference between the different mass eigenstates.

What we do in this article is to use a full field theoretic approach to follow this time evolution explicitly. We start from a simplified model that encodes the salient interactions necessary to analyze the development of oscillations. Using this model, we calculate the form of the state produced when the parent particle (be it a W boson or a pion) decays into the flavor neutrino and its associated charged lepton. We then use the time evolution operator, calculated within perturbation theory from the interaction Lagrangian, to evolve this state until it reaches the detector. At this point, we obtain the overlap with the final states observed at the detector.

Our main observation is that oscillation experiments are two-time experiments, the two
times being the time at which the original charged lepton is measured and that when the final one is measured. While this does not seem earth-shattering, what is extremely interesting is the interplay between these times (and/or time differences) and other time scales in the problem such as the oscillation time scale and the time scale upon which it can be said that the two mass eigenstates have been resolved. In fact, one of our results is that the treatment of the flavor eigenstate as the usual coherent mixture of the two mass eigenstates requires that the original entangled state already be disentangled long before an oscillation time. While this is true for long baseline experiments, it is not necessarily so for short baseline ones, and this could give rise to a reinterpretation of their results to take the dependence on the disentangling time into account.

In all of this, we would like to emphasize that we are just following the tenets of quantum mechanics in our calculation: prepare a state, construct and then use the relevant time evolution operator and lastly, make a measurement on that state by overlapping it with various states of interest at the later time. Following the time evolution in detail as we do gives us the access to finer details of the behavior of the state than what would be obtained when one uses Fermi’s Golden Rule, which requires taking an infinite time limit, to directly calculate rates and the like. In our formalism, the finite-time effects become explicit and this allows us to fully understand the domain of validity of the various results in the literature.

II. A MODEL OF “NEUTRINO” OSCILLATIONS

Let us start by introducing our model. In order to exhibit the main results in a clear and simple manner, we introduce a bosonic model that describes mixing, oscillations and charged current weak interactions reliably without the complications associated with fermionic and gauge fields. We can do so because the technical complications associated with spinors and gauge fields are irrelevant to the physics of mixing and oscillations, as is obviously manifest in meson mixing. Our model is defined by the following Lagrangian density

\[ L = L_0[W, l_\alpha] + L_0[\nu_\alpha] + L_{\text{int}}[W, l_\alpha, \nu_\alpha] ; \quad \alpha = e, \mu \]

with

\[ L_0[\nu] = \frac{1}{2} \left[ \partial_\mu \Psi^T \partial^\mu \Psi - \Psi^T M \Psi \right], \]

where \( \Psi \) is a flavor doublet representing the neutrinos

\[ \Psi = \begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix}, \]

and \( M \) is the mass matrix

\[ M = \begin{pmatrix} m_{ee} & m_{e\mu} \\ m_{e\mu} & m_{\mu\mu} \end{pmatrix}. \]

The interaction Lagrangian is similar to the charged current interaction of the standard model, namely

\[ L_{\text{int}}(\vec{x}, t) = g W(\vec{x}, t) \left[ l_e(\vec{x}, t) \nu_e(\vec{x}, t) + l_\mu(\vec{x}, t) \nu_\mu(\vec{x}, t) \right], \]

where \( g \) is the coupling constant. \( W(x) \) represents the vector boson, or alternatively the pion field, and \( l_\alpha, \alpha = e, \mu \) the two charged leptons. The mass matrix is diagonalized by a
unitary transformation
\[
U^{-1}(\theta)M U(\theta) = \begin{pmatrix} m_1 & 0 \\ 0 & m_2 \end{pmatrix} ; \quad U(\theta) = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix}.
\] (2.6)

In terms of the doublet of mass eigenstates, the flavor doublet can be expressed as
\[
\begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = U(\theta) \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}.
\] (2.7)

This bosonic model clearly describes charged current weak interactions reliably as it includes all the relevant aspects of mixing and oscillations.

We consider “neutrino” oscillation experiments following the interaction processes illustrated in Fig. 1,
\[
W \rightarrow l_\alpha + \nu_\alpha \rightarrow \begin{cases} W + l_\beta , \beta \neq \alpha & \text{appearance} \\ W + l_\alpha & \text{disappearance} \end{cases}.
\] (2.8)

FIG. 1: Typical experiment in which the charged leptons are measured at the source and far detector and the neutrino is an intermediate state.

Throughout this work, we will use plane wave states for simplicity of exposition, though we discuss some of the modifications that the use of wave packets would require later in this section.

In order to study aspects of coherence we consider a simplified interaction Lagrangian density
\[
\mathcal{L}_I = g W e \nu_e = g W e (\cos \theta \nu_1 + \sin \theta \nu_2),
\] (2.9)

focusing only on one lepton, which we refer to as the “electron.” The full coupling as in Eq. (2.5) can be treated similarly without modifying the main conclusions. Although \(W\) may be interpreted as a charged vector boson, the analysis is obviously the same if it describes a pion field.

We can study aspects of coherence by focusing on the Fock state obtained upon evolution of the decaying initial state. We consider a plane-wave Fock initial state \(|W(k)\rangle\) at \(t_i = 0\). The time evolution operator is
\[
e^{-iH(t-t_i)} = e^{-iH_0 t_i} U(t, t_i) e^{iH_0 t_i},
\] (2.10)
where
\[
U(t, t_i) = T e^{i \int_{t_i}^t d^3x dt \mathcal{L}_{\text{int}}(\vec{x}, t)} \tag{2.11}
\]
is the time evolution operator in the interaction picture and $H_0$ is the non-interacting Hamiltonian.

By time $t$, the initial state has evolved into
\[
|\Psi(t)\rangle = |W(\vec{k})\rangle e^{-i E^W_k t} + |\Psi_e(t)\rangle. \tag{2.12}
\]
To lowest order in the interaction, we find the second term to be
\[
|\Psi_e(t)\rangle = \frac{ig}{\sqrt{2 \Omega}} e^{-i E^W_k t} \sum_{q} \left[ \frac{\sin \theta}{\sqrt{\Omega_{1,q}}} \left| \nu_{1,q} \right| e_{q} \right]
\]
\[
\times \left[ e^{i(E^W_k - E^e_q - \Omega_{2,q})t} \right]
\]
\[
+ \frac{\cos \theta}{\sqrt{\Omega_{1,q}}} \left| \nu_{1,q} \right| e_{q} \left[ e^{i(E^W_k - E^e_q - \Omega_{1,q})t} \right]
\]
\[
; \quad \vec{p} = \vec{k} - \vec{q} \tag{2.13}
\]
in which the electron and the neutrinos are entangled\(^1\). The neutrino state that is entangled with the muon is obtained from (2.13) by replacing $\cos \theta \to -\sin \theta$; $\sin \theta \to \cos \theta$.

In what follows, we consider ultrarelativistic neutrinos and write
\[
\Omega_1 = \overline{\Omega} - \Delta ; \quad \Omega_2 = \overline{\Omega} + \Delta, \tag{2.14}
\]
where
\[
\overline{\Omega} = \left[ p^2 + \frac{m_1^2 + m_2^2}{2} \right]^{\frac{1}{2}} \; ; \; \Delta = \frac{\delta m^2}{4 \overline{\Omega}} \; ; \; \delta m^2 = m_2^2 - m_1^2, \tag{2.15}
\]
taking $\Delta \ll \overline{\Omega}$ as is the case for ultrarelativistic nearly degenerate neutrinos.

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\(^1\) The result for the wavefunction in Ref. [22] may be understood using a (non-perturbative) Wigner-Weisskopf approximation for the decaying parent particle, replacing $E_W \to E_W - i \Gamma_W$. Taking $t \gg 1/\Gamma_W$ in the integral replaces the brackets in (2.13) by $1/(E_W - E^e - \Omega_j - i \Gamma_W)$ whose absolute value is proportional to $\delta(E_W - E^e - \Omega_j)/\Gamma_W$. 

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A. Unobserved daughter particles: time evolution of the density matrix

If the electrons (or daughter particle in Ref. [22]) are not observed, they can be traced out of the density matrix $\langle \Psi_e(t) | \Psi_e(t) \rangle$. This gives the reduced density matrix

$$\rho_r(t) = \text{Tr}_e \langle \Psi_e(t) | \Psi_e(t) \rangle$$

$$= \frac{g^2}{8VE_\infty^W \sum \frac{\sin^2 \theta}{\Omega_{1p} E_\infty} \left| \nu_{1p} \right\rangle \left\langle \nu_{1p} \right| \left[ \sin \left( \frac{E_\infty^W - E_\infty^e - \Omega_{1p}}{2} \right) \right]^2$$

$$+ \frac{\cos^2 \theta}{\Omega_{1p} E_\infty} \left| \nu_{1p} \right\rangle \left\langle \nu_{1p} \right| \left[ \sin \left( \frac{E_\infty^W - E_\infty^e - \Omega_{1p}}{2} \right) \right]^2$$

$$+ \frac{\sin 2\theta}{2E^W_{\infty} \sqrt{\Omega_{2p}}} \left[ \sin \left( \frac{E_\infty^W - E_\infty^e - \Omega_{1p}}{2} \right) \right] \left[ \sin \left( \frac{E_\infty^W - E_\infty^e - \Omega_{1p}}{2} \right) \right]$$

$$\times \left\{ e^{-\frac{\delta m^2}{4t}} \left| \nu_{2p} \right\rangle \left\langle \nu_{2p} \right| + e^{\frac{\delta m^2}{4t}} \left| \nu_{1p} \right\rangle \left\langle \nu_{2p} \right| \right\} ; \quad \bar{p} = \bar{k} - \bar{q}. \tag{2.16}$$

This expression contains remarkable information. The function $\sin^2(\pi t)/x^2$ is the usual “diffraction” function of Fermi’s Golden rule. In the formal long time limit $\sin^2(\pi t)/x^2 \rightarrow \pi t \delta(x)$, the first two terms of the density matrix, which are the diagonal entries in the mass basis, describe the production process of the mass eigenstates. As will be seen below, in the long time limit, the time derivative of these two terms yields the production rate for the mass eigenstates. In the formal $t \rightarrow \infty$ limit, these are the diagonal terms obtained in Ref. [22], where in that reference, the product of delta functions is again understood as the total time elapsed times an energy conserving delta function.

The off-diagonal terms in the last line of (2.16) describe the “coherences” and display the oscillatory phases from the interference of the mass eigenstates. The time dependent factors of the off-diagonal density matrix elements determine the coherence between the mass eigenstates and will be a ubiquitous contribution in the real time description of oscillations that follows below. The functions

$$f_\pm(x, t; \Delta) = \frac{2 \sin \left( \frac{(x \pm \Delta)}{2} \right)}{(x \pm \Delta)} ; \quad x = E_\infty^W - E_\infty^e - \Omega_{1p} ; \quad \Delta = \frac{\delta m^2}{4\Omega} \tag{2.17}$$

are strongly peaked at $x \pm \Delta = 0$ with height $t$ and width $\sim 2\pi/t$. In the infinite time limit, $f_\pm(x, t, \Delta) \rightarrow 2\pi \delta(x \pm \Delta)$ and therefore their product would vanish in this limit, leading to the vanishing of the coherence. This is the result obtained in Ref. [22]. However, at finite time $t$, they feature a non-vanishing overlap when $2\Delta \lesssim 2\pi/t$. We recognize this as the condition for oscillations. We note that $t \sim \pi/\Delta$ yields a macroscopically large time scale. The functions $f_\pm(x, t, \Delta)$ and their products are depicted in Figs. (23) for the values $\Delta = 0.1, t = 40, 100$, respectively.
It is straightforward to find

\[ f_-(x, t, \Delta) f_+(x, t, \Delta) = \frac{\sin(\Delta t)}{\Delta} \left[ \frac{\sin[(x - \Delta)t]}{x - \Delta} + \frac{\sin[(x + \Delta)t]}{x + \Delta} \right] \]

\[ + 2 \frac{\cos(\Delta t)}{\Delta} \left[ \frac{\sin^2[(x - \Delta)t/2]}{x - \Delta} - \frac{\sin^2[(x + \Delta)t/2]}{x + \Delta} \right]. \tag{2.18} \]

In the long time limit, the terms in the first bracket yield a sum of delta functions at \( x = \pm \Delta \). Upon integrating the product of \( f_- f_+ \) with functions of compact support, the contribution from the second line in (2.18) is negligible in the long time limit. Therefore, the long time limit of \( f_- f_+ \) can be replaced by

\[ f_-(x, t, \Delta) f_+(x, t, \Delta) = \pi \frac{\sin(\Delta t)}{\Delta} \left[ \delta(x - \Delta) + \delta(x + \Delta) \right]. \tag{2.19} \]

During a time interval \( t \ll 2\pi/\Delta \), the product \( f_-(x, t, \Delta) f_+(x, t, \Delta) \sim \pi t [\delta(x-\Delta)+\delta(x+\Delta)] \) grows linearly in time until it begins to oscillate with frequency \( 2\pi/\Delta \) for \( t > 2\pi/\Delta \).
Therefore, we conclude that upon integration with a smooth density of states, the off-diagonal terms in the density matrix grow linearly in time for \( t \ll t_{\text{osc}} = 2\pi/\Delta \), but feature a bound oscillatory behavior of frequency \( \Delta \) for \( t \gtrsim 2\pi/\Delta \).

The diagonal terms, i.e., the first two terms in the reduced density matrix (2.16), are proportional to \( 4\sin^2(x \frac{t}{\Delta})/x^2 \rightarrow 2\pi \, t \delta(x) \). This shows that the coherences or off diagonal terms are linear in time and of the same order as the diagonal elements for \( t \lesssim t_{\text{osc}} = 2\pi/\Delta \), but are of \( O(1/\Delta \, t) \) and oscillate compared to the diagonal terms for \( t \gg t_{\text{osc}} \). This behavior is displayed in Fig. 4, where as an example we consider a smooth density of states and the integral

\[
I(t, \Delta) = \int_{-\infty}^{\infty} e^{-x^2} f_+(x, t, \Delta) f_-(x, t, \Delta) dx. \tag{2.20}
\]

The case \( \Delta = 0 \) describes either of the diagonal terms and is linearly secular in time. This figure clearly shows the slow oscillations for \( t \gtrsim 1/\Delta \).

![Fig. 4: The integral \( I(t, \Delta) = \int_{-\infty}^{\infty} e^{-x^2} f_+(x, t, \Delta) f_-(x, t, \Delta) dx \) vs. \( t \) for \( \Delta = 0, 0.1, 0.3 \)](image)

Therefore, the approximation (2.19) is reliable in the long time limit and upon integration with functions of compact support. We see that for large \( t \), but \( t\Delta \ll 1 \), \( f_+(x, t, \Delta) f_-(x, t, \Delta) \rightarrow \pi t \left[ \delta(x - \Delta) + \delta(x + \Delta) \right] \) and for \( \Delta \rightarrow 0 \), the product yields \( 2\pi t \delta(x) \).

The reduced density matrix (2.16) allows us to obtain the time evolution of the neutrino populations and coherences, namely

\[
n_i(\vec{p}, t) = \text{Tr} \rho_r(t)a_i^\dagger(\vec{p})a_i(\vec{p}) ; \quad C_{ij}(\vec{p}, t) = \text{Tr} \rho_r(t)a_i^\dagger(\vec{p})a_j(\vec{p}) , \quad i \neq j \tag{2.21}
\]

where the annihilation and creation operators are in the Schroedinger picture. In the long time limit and using the results above, we find

\[
n_1(\vec{p}, t) = t \, \Gamma_1(\vec{p}) \cos^2 \theta ; \quad n_2(\vec{p}, t) = t \, \Gamma_2(\vec{p}) \sin^2 \theta, \tag{2.22}
\]

where

\[
\Gamma_{1,2} = \frac{2\pi g^2}{8E_W^Q} \int \frac{d^3 \vec{Q}}{(2\pi)^3} \frac{E_W^Q}{E_Q^Q \Omega_{1,2}} \delta(E_W^Q - E_Q^Q - \Omega_{1,2}) \tag{2.23}
\]
are the partial widths for the decay of the $W$ into the charged lepton and the neutrino mass eigenstates. For the off-diagonal matrix elements or coherences we find,

$$C_{12}(\vec{p},t) = C_{12}^\dagger(\vec{p},t) = \frac{2\pi g^2 \sin 2\theta}{32E_k^W} \frac{\sin|\Delta|}{\Delta} e^{i\Delta t}$$

$$\int \frac{d^3\vec{Q}}{(2\pi)^3 E_Q^e \sqrt{\Omega_1 \Omega_2}} \left[ \delta(E_k^W - E_Q^e - \Omega_1) + \delta(E_k^W - E_Q^e - \Omega_2) \right].$$

(2.24)

d$\frac{dn_i(\vec{p},t)}{dt}$ yields the production rate of the neutrino mass eigenstates from the decay of the $W$ and the coherences $C_{ij}$ are non-vanishing at any finite time. In Ref. [22], these coherences vanish as a consequence of the product of delta functions on the different mass shells of the mass eigenstates. The coherences vanish in the formal infinite time limit because of the oscillatory behavior averages out on time scales $t \gg 1/\Delta$ and they are of $O(\sin(t\Delta)/t\Delta)$ with respect to the diagonal terms. However, we emphasize that experimentally, the time scales involved (or length scales) are of order $1/\Delta$ as these are the scales on which oscillatory phenomena are revealed. Taking $\Omega_1 \sim \Omega_2 \sim \bar{\Omega}$ in the denominators in (2.23,2.24), it follows that

$$C_{12}(\vec{p},t) \approx \frac{\sin 2\theta \sin[t\Delta]}{2\Delta} e^{i\Delta t} \frac{1}{2} \left[ \Gamma_1 + \Gamma_2 \right].$$

(2.25)

Therefore, the coherences are of the same order of the population terms on time scales $t \leq 1/\Delta$, but average out for $t \gg 1/\Delta$, showing that coherence persists over the oscillation time scale.

**B. Disentangling the neutrino: a two-time measurement**

As we have discussed above, a long baseline experiment is actually a two-time measurement process, as the charged lepton produced at the interaction vertex at the source is detected at the source or stopped and absorbed in a nearby “beam dump”.

This “measurement” of the charged lepton disentangles the neutrinos from the charged lepton in the quantum state (2.13) [22]. The detection of the charged lepton at the source (or its stopping at a nearby “beam dump”) projects the quantum state (2.13) at the observation time $t_S$ onto the single particle charged lepton state $e^{-iE_Q^e t_S} |e_{\vec{Q}}\rangle$ disentangling the neutrino states into the “collapsed” state

$$|\mathcal{V}_e(\vec{Q}, t_S)\rangle \equiv \langle e_{\vec{Q}} | \Psi(t_S) \rangle e^{iE_Q^e t_S}$$

$$= i \frac{g}{2 \sqrt{2V E_k^W E_Q^e}} \left\{ \frac{\sin \theta}{\sqrt{\Omega_2, \vec{P}}} \left[ \nu_{2, \vec{P}} e^{-i\Omega_2, \vec{P} t_S} \frac{\sin \left( (E_S - \Omega_2, \vec{P}) t_S \right)}{(E_S - \Omega_2, \vec{P})/2} \right] \right. \right.$$

$$\left. + \frac{\cos \theta}{\sqrt{\Omega_1, \vec{P}}} \left[ \nu_{1, \vec{P}} e^{-i\Omega_1, \vec{P} t_S} \frac{\sin \left( (E_S - \Omega_1, \vec{P}) t_S \right)}{(E_S - \Omega_1, \vec{P})/2} \right] \right\};$$

$$E_S = E_k^W - E_Q^e; \quad \vec{P} = \vec{k} - \vec{Q}.$$
We note that up to a phase the coefficient functions that multiply $|\nu_{1,2}\rangle$ are

$$\sin\left[\frac{(E_S - \Omega_1)t_S}{2}\right], \quad \sin\left[\frac{(E_S - \Omega_2)t_S}{2}\right],$$

(2.27)

respectively. In the limit $t_S \to \infty$, these become $2\pi \delta(E_S - \Omega_1)$ and $2\pi \delta(E_S - \Omega_2)$, respectively. Therefore, in this limit, for a fixed $E_S$, one of the quantum states will be projected out. However, as we insist on keeping a finite time interval, we will keep $t_S$ finite.

The state $|\mathcal{V}_e(\tilde{Q}, t_S)\rangle$ then evolves forward in time with the full Hamiltonian

$$|\mathcal{V}_e(\tilde{Q}, t)\rangle = e^{-iH_0 t} U(t, t_S) e^{iH_0 t_S} |\mathcal{V}_e(\tilde{Q}, t_S)\rangle.$$  

(2.28)

The “free” evolution is obtained by setting to lowest order $U(t, t_S) = 1$, leading to

$$|\mathcal{V}_e(\tilde{Q}, t)\rangle = i \frac{g}{2\sqrt{2VE_W E_k^*}} \left\{ \frac{\sin \theta}{\sqrt{\Omega_{1,\tilde{P}}}} |\nu_{2,\tilde{P}}\rangle \left[ \frac{\sin \left[\frac{(E_S - \Omega_2,t_S)}{2}\right]}{(E_S - \Omega_2)/2} \right] e^{-i\Omega_{2,\tilde{P}} t_S} e^{-i\Omega_{2,\tilde{P}}(t-t_S)} \right\}.$$

(2.29)

The phase factors $e^{-i\Omega_{j}t_S/2}$ multiplying each mass eigenstate are the consequence of the phase build-up during the time evolution from the production vertex until the detection of the charged lepton and the collapse of the wave function. These can be absorbed into the definition of the states $|\nu_{1,2}\rangle$ at the fixed time $t_S$.

The expression (2.29) features the factors

$$\sin\left[\frac{(E_S - \Omega_j)t_S}{2}\right],$$

(2.30)

which as $t_S \to \infty$ becomes $2\pi \delta(E_S - \Omega_j)$. These factors, which are a direct consequence of the neutrino state being produced by the decay of the “parent” particle (here the $W$) into an entangled quantum state, distinguish Eq. (2.29) from the familiar quantum mechanical description. These factors emerge from the (approximate) energy conservation at the decay vertex. Again, in the $t_S \to \infty$ limit, if the energy of the parent particle and the charged lepton are both certain, only one of the mass eigenstates will be produced but not both. However, writing $\Omega_{1,2}$ as in Eq. (2.14), it follows that for $t_S \ll 4\pi/\delta m^2$, the width of the “diffraction” functions is much larger than the frequency difference $\Delta$ and there is a substantial overlap between these “approximate” delta functions (see Fig. 2). Only for $t_S \geq t_{osc} = 2\pi/\Delta$ are the two peaks at $E_S - \Omega = \pm \Delta$ actually resolved, whereas for $t_S \ll t_{osc}$, the two peaks are unresolved, “blurred” into one broad peak at $\Omega$. Thus for $t_S \ll 2\pi/\Delta$, we can use the approximation

$$\frac{\sin\left[\frac{(E_S - \Omega_1)t_S}{2}\right]}{(E_S - \Omega_1)/2} \approx \frac{\sin\left[\frac{(E_S - \Omega_2)t_S}{2}\right]}{(E_S - \Omega_2)/2} \approx \frac{\sin\left[\frac{(E_S - \Omega)t_S}{2}\right]}{(E_S - \Omega)/2}.$$  

(2.31)
To illustrate the validity of the above approximation, let us consider the case in which the typical size of the source or beam dump is a few meters across. In a typical experiment, the charged lepton emerging from the interaction vertex travels this distance within a time scale \( t_S \approx 10^{-8} \text{ s} \), leading to an energy uncertainty above \( \delta E \sim \hbar / t_S \sim 10^{-7} \text{ eV} \). Taking as an example \( \delta m^2 \sim 10^{-4} \text{ eV}^2; \Omega \sim E_S \sim 100 \text{ MeV} \), it follows that \( \delta m^2 / \Omega \sim 10^{-12} \text{ eV} \ll \delta E \).

Therefore, the detection or absorption of the charged lepton near the source, on a time scale much smaller than \( 1 / \Delta \) cannot resolve the energy difference between the mass eigenstates and the approximation (2.31) is justified.

Another approximation we can use in (2.26) is \( \Omega_1 \simeq \Omega_2 \simeq \Omega \), for \( \Omega \gg \Delta \). Absorbing the phase factors \( e^{-i\Omega_j t_S / 2} \) into the definition of the states \( |\nu_j\rangle \), the time evolved disentangled state is then approximately given by

\[
|\nu_e(t_S)| = \frac{ig}{8VE_k W E_{\vec{Q}}^{\Omega / 2}} \left[ \sin \left( (E_S - \Omega) t_S / 2 \right) / (E_S - \Omega) / 2 \right] \left\{ \sin \theta |\nu_2, \vec{\rho}\rangle e^{-i\Omega_2, \vec{\rho}(t-t_S)} + \cos \theta |\nu_1, \vec{\rho}\rangle e^{-i\Omega_1, \vec{\rho}(t-t_S)} \right\},
\]

for \( t_S \ll 2\pi / \Delta \).

The state inside the brackets is identified as the usual quantum mechanical state that is time evolved from the “electron” neutrino state, which is prepared initially at \( t_S \). From this analysis, we see that there are two conditions required for the disentangled neutrino state to be identified with the familiar quantum mechanical state: \( \delta m^2 / \Omega^2 \ll 1 \) and \( t_S \ll t_{osc} \sim 2\pi/\Omega / \delta m^2 \). The former is always satisfied for neutrinos with \( \delta m^2 \sim 10^{-3} - 10^{-4} \text{ eV}^2; \Omega > \text{few MeV} \), while the latter is fulfilled if the charged lepton produced at the source is either observed or stopped at a beam-dump near the production region in long-baseline experiments.

The latter condition implies that the neutrino state is disentangled before oscillations can occur. In a long-baseline experiment this is achieved if the charged lepton, which is entangled with the neutrino at the production vertex, is measured or stopped near the production region. Therefore, we conclude that an important condition for the familiar quantum mechanical description of oscillations to be reliable is that the quantum state must be disentangled on time scales much shorter than the oscillation time.

It is clear from this discussion that the precise value of \( t_S \) is irrelevant as long as \( t_S \ll t_{osc} \). Experimentally, \( t_S \) is of the order of the time of flight of the charged lepton in the production region at the source namely a few \( 10^{-8} \text{ s} \).

C. Transition amplitudes and event rates

The number of charged lepton events with momentum \( \vec{Q} \) produced at the source, at time \( t_S \) is given by

\[
n_e(Q, t_S) = \langle \Psi_e(t_S) | a_e^\dagger(Q) a_e(Q) | \Psi_e(t_S) \rangle = \langle \nu_e(t_S) | \nu_e(Q, t_S) \rangle.
\]

For \( t_S \ll t_{osc} = 2\pi / \Delta \) and \( \Omega \gg \Delta \), using the approximations leading to (2.32), we obtain the charged lepton differential detection rate at the source.
The disappearance and appearance amplitudes are then given by

\[
(2\pi)^3 \frac{dN^S_e}{d^4x\,d^3\vec{Q}} = \frac{dn_e(\vec{Q},t_S)}{dt_S} = \frac{2g^2}{8VE^W_k\,E^e_Q\,\Omega} \sin \left[ \left( E_S - \Omega \right) t_S \right] \sim \frac{2\pi g^2}{8VE^W_k\,E^e_Q\,\Omega} \delta(E_S - \Omega),
\]

where at the last step we have replaced the diffraction function by the delta function. This can be justified as follows. For \( t_S \sim 10^{-8}\) s, the width of this function (the resolution) in energy is \( \sim 10^{-7}\) eV. Since the typical energy in a long-baseline experiment is \( \gtrsim 40 - 100\) MeV, the error incurred in replacing the diffraction function by a delta function is smaller than one part in \( 10^{15}\).

We can also obtain the transition amplitude for the disentangled state to produce a final charged lepton and another W particle at the far detector at time \( t_D \), where \( t_D - t_S \sim L \) and \( L \) is the baseline. This is obtained from the time evolution of the disentangled state with the full Hamiltonian from the time of disentanglement \( t_S \) until the time \( t_D \), at which the detection measurement is carried out by overlapping the time evolved state with the final state.

The transition amplitude is given by

\[
\mathcal{A}_{\alpha \rightarrow \beta} = \langle W(\vec{k}_D), l_\beta(\vec{p}_D) | e^{-iH(t_D-t_S)} | \nu_e(\vec{Q},t_S) \rangle = e^{-iE_D t_D} \langle W(\vec{k}_D), l_\beta(\vec{p}_D) | U(t_D,t_S) e^{iH_0 t_S} | \nu_e(\vec{Q},t_S) \rangle.
\]

The disappearance and appearance amplitudes are then given by

\[
\mathcal{A}_{e \rightarrow e} = -g^2 \Pi (2\pi)^3 \delta(\vec{k}_S - \vec{p}_S - \vec{k}_D - \vec{p}_D) \times
\]

\[
\left\{ \frac{\cos^2 \theta}{2 \Omega_{1,\hat{r}}} e^{-i\Omega_{1,\hat{r}} t_D/2} \left[ \sin \left( \frac{(E_S - \Omega_{1,\hat{r}}) t_D}{2} \right) \right] \left[ \sin \left( \frac{(E_D - \Omega_{1,\hat{r}}) (t_D - t_S)}{2} \right) \right] \right\} \]

\[
+ \frac{\sin^2 \theta}{2 \Omega_{2,\hat{r}}} e^{-i\Omega_{2,\hat{r}} t_D/2} \left[ \sin \left( \frac{(E_S - \Omega_{2,\hat{r}}) t_D}{2} \right) \right] \left[ \sin \left( \frac{(E_D - \Omega_{2,\hat{r}}) (t_D - t_S)}{2} \right) \right],
\]

and

\[
\mathcal{A}_{e \rightarrow \mu} = -g^2 \Pi (2\pi)^3 \delta(\vec{k}_S - \vec{p}_S - \vec{k}_D - \vec{p}_D) \sin \frac{2\theta}{2} \times
\]

\[
\left\{ e^{-i\Omega_{1,\hat{r}} t_D/2} \left[ \sin \left( \frac{(E_S - \Omega_{1,\hat{r}}) t_D}{2} \right) \right] \left[ \sin \left( \frac{(E_D - \Omega_{1,\hat{r}}) (t_D - t_S)}{2} \right) \right] \right\} \]

\[
- \frac{e^{-i\Omega_{2,\hat{r}} t_D/2}}{2 \Omega_{2,\hat{r}}} \left[ \sin \left( \frac{(E_S - \Omega_{2,\hat{r}}) t_D}{2} \right) \right] \left[ \sin \left( \frac{(E_D - \Omega_{2,\hat{r}}) (t_D - t_S)}{2} \right) \right].
\]

(2.36)

(2.37)
with $\vec{P} = \vec{k}_S - \vec{p}_S$, and $\Pi$ is given by

$$\Pi = \left[ \frac{1}{16 V^4} \frac{1}{E_{kS}^W} \frac{1}{E_{kD}^W} \frac{1}{E_{pS}^l} \frac{1}{E_{pD}^l} \right]^{\frac{1}{2}},$$  \hspace{1cm} (2.38)$$

the labels $D, S$ refer to (far) detector and source respectively.

Implementing the same approximations leading to the factorized state (2.32), namely $\Omega \gg \Delta$ and $t_S \Delta \ll 1$, we find the disappearance and appearance probabilities

$$P_{e^{-}\mu}(t_D) = \left( \frac{g^2 \Pi}{2\Omega} \right)^2 (2\pi)^5 V \delta(\vec{k}_S - \vec{p}_S - \vec{k}_D - \vec{p}_D) t_S \delta(E_S - \overline{\Omega}_P)$$

$$\begin{cases} 
\cos^4 \theta f^2(x, t, \Delta) + \sin^4 \theta f^2(x, t, \Delta) + 2 \cos^2 \theta \sin^2 \theta \cos(t\Delta) f_+(x, t, \Delta) f_-(x, t, \Delta)
\end{cases},$$

$$P_{e^{-}\mu}(t_D) = \left( \frac{g^2 \Pi}{2\Omega} \right)^2 (2\pi)^5 V \delta(\vec{k}_S - \vec{p}_S - \vec{k}_D - \vec{p}_D) t_S \delta(E_S - \overline{\Omega}_P) \frac{\sin^2 2\theta}{4}$$

$$\begin{cases} 
f^2_+(x, t, \Delta) + f^2_-(x, t, \Delta) - 2 \cos(t\Delta) f_+(x, t, \Delta) f_-(x, t, \Delta)
\end{cases},$$  \hspace{1cm} (2.39)$$

where $t = t_D - t_S$ and $x = E_D - \overline{\Omega}_P$. Here,

$$E_S = E_{kS}^W - E_{pS}^l; \ E_D = E_{kD}^W + E_{pD}^l,$$  \hspace{1cm} (2.40)$$

and $f_{\pm}$ are given by Eq. 2.17.

In the long time limit, using $f_{\pm}(x, t, \Delta) \rightarrow 2\pi t \delta(x \pm \Delta)$ their product is given by (2.19), and we find

$$P_{e^{-}\mu}(t_D) = \left( \frac{g^2 \Pi}{2\Omega} \right)^2 (2\pi)^5 V \delta(\vec{k}_S - \vec{p}_S - \vec{k}_D - \vec{p}_D) t_S \delta(E_S - \overline{\Omega}_P)$$

$$\begin{cases} 
\cos^4 \theta t \delta(x + \Delta) + \sin^4 \theta t \delta(x - \Delta) + 2 \cos^2 \theta \sin^2 \theta \frac{\sin(2t\Delta)}{2\Delta} \frac{1}{2} \left[ \delta(x + \Delta) + \delta(x - \Delta) \right]
\end{cases},$$

$$P_{e^{-}\mu}(t_D) = \left( \frac{g^2 \Pi}{2\Omega} \right)^2 (2\pi)^5 V \delta(\vec{k}_S - \vec{p}_S - \vec{k}_D - \vec{p}_D) t_S \delta(E_S - \overline{\Omega}_P) \frac{\sin^2 2\theta}{4}$$

$$\begin{cases} 
t \delta(x + \Delta) + t \delta(x - \Delta) - 2 \frac{\sin(2t\Delta)}{2\Delta} \frac{1}{2} \left[ \delta(x + \Delta) + \delta(x - \Delta) \right]
\end{cases}. \hspace{1cm} (2.41)$$

where $\vec{P} = \vec{k}_D + \vec{p}_D$.

These transition probabilities feature the two time scales $t_S$ and $t = t_D - t_S$ associated with the measurements at the near and far detector. They also feature energy conserving delta functions associated with the different mass eigenstates.

There is a further simplification when $\Omega \gg \Delta$. In this regime, when the probabilities (2.41) are integrated over a smooth density of states, the delta functions corresponding to the mass eigenstates yield the density of states at values $E_D = \Omega \mp \Delta$. In typical experiments, where $\Omega \sim 100$ MeV and $\delta m^2 \sim 10^{-3}$ eV$^2$, the density of final states must vary dramatically near $\Omega$ to resolve the small interval $\Delta$, with $\Delta/\Omega \lesssim 10^{-19}$. Therefore, understanding the
probabilities as being integrated with a smooth final density of states insensitive to the mass difference, we can approximate \( \delta(x \pm \Delta) \approx \delta(x) \). In this case, we can approximate the above expressions by

\[
\mathcal{P}_{e \rightarrow e}(t_D) = \left( \frac{g^2 \Pi}{2 \Omega_D} \right)^2 (2\pi)^5 V \delta(k_S - \vec{p}_S - \vec{k}_D - \vec{p}_D) t_S \delta(E_S - \Omega_D) \delta(E_D - \Omega_D)
\]

\[
\left\{ t \left[ \cos^4 \theta + \sin^4 \theta \right] + 2 \cos^2 \theta \sin^2 \theta \frac{\sin(2t \Delta)}{2\Delta} \right\},
\]

(2.42)

\[
\mathcal{P}_{e \rightarrow \mu}(t_D) = \left( \frac{g^2 \Pi}{2 \Omega_D} \right)^2 (2\pi)^5 V \delta(k_S - \vec{p}_S - \vec{k}_D - \vec{p}_D) t_S \delta(E_S - \Omega_D) \delta(E_D - \Omega_D) \frac{\sin^2 2\theta}{2}
\]

\[
\left\{ t - \frac{\sin(2t \Delta)}{2\Delta} \right\}.
\]

(2.43)

The product \( \delta(E_S - \Omega_D) \delta(E_D - \Omega_D) \) is an approximate energy conservation at both production and detection vertices, where we have neglected \( \Delta \), which is twice the energy difference between the mass eigenstates.

Further insight can be gained by obtaining the phase space distribution of the number of charged leptons \( l = e, \mu \) at the far detector

\[
(2\pi)^3 \frac{dN_{FD}^l}{d^3x d^3\vec{p}_D} = n_l(\vec{p}_D, t_D) = \langle V_e(\vec{Q}, t_D, t_S) | a_l^\dagger(\vec{p}_D) a_l(\vec{p}_D) | V_e(\vec{Q}, t_D, t_S) \rangle.
\]

(2.44)

Here,

\[
| V_e(\vec{Q}, t_D, t_S) \rangle = e^{-iH_0 t_D} U(t_D, t_S) e^{iH_0 t_S} | V_e(\vec{Q}, t_S) \rangle
\]

(2.45)

is the neutrino state disentangled at \( t_S \) near the source and has been time-evolved until it is detected at the far detector at time \( t_D \). Not surprisingly, since we only keep terms that are up to order \( g^2 \), the time evolved state contains a single lepton Fock state. We find that

\[
(2\pi)^3 \frac{dN_{FD}^e}{d^3x d^3\vec{p}_D} = \mathcal{P}_{e \rightarrow e}(t_D) \quad ; \quad (2\pi)^3 \frac{dN_{FD}^\mu}{d^3x d^3\vec{p}_D} = \mathcal{P}_{e \rightarrow \mu}(t_D),
\]

(2.46)

with the probabilities \( \mathcal{P}_{e \rightarrow e}(t_D) \) and \( \mathcal{P}_{e \rightarrow \mu}(t_D) \) are given by (2.42, 2.43).

Taking the time derivative with respect to \( t_D \), we obtain the differential charged lepton event rates at the far detector

\[
(2\pi)^3 \frac{dN_{FD}^e}{d^3x dt d^3\vec{p}_D} = \left( \frac{g^2 \Pi}{2 \Omega_D} \right)^2 (2\pi)^5 V \delta(k_S - \vec{p}_S - \vec{k}_D - \vec{p}_D) t_S \delta(E_S - \Omega_D) \delta(E_D - \Omega_D)
\]

\[
\left\{ \cos^4 \theta + \sin^4 \theta + 2 \cos^2 \theta \sin^2 \theta \cos(2t \Delta) \right\},
\]

(2.47)

\[
(2\pi)^3 \frac{dN_{FD}^\mu}{d^3x dt d^3\vec{p}_D} = \left( \frac{g^2 \Pi}{2 \Omega_D} \right)^2 (2\pi)^5 V \delta(k_S - \vec{p}_S - \vec{k}_D - \vec{p}_D) t_S \delta(E_S - \Omega_D) \delta(E_D - \Omega_D) \frac{\sin^2 2\theta}{2}
\]

\[
\left\{ 1 - \cos(2t \Delta) \right\}.
\]

(2.48)
Remarkably, these rates can be simply written as

\[ (2\pi)^3 \frac{dN^D_{\nu}}{d^3x dt d^3p_D} = (2\pi)^3 \frac{dN^S_{\alpha}}{d^3x d^3p_s} P_{\alpha \rightarrow \beta}(t) d\Gamma_{\nu_{\beta} \rightarrow W_{l_\beta}} \]

(2.49)

where we have used the expression (2.34) for the differential charged lepton event rate at the source and integrated in \( t_S \).

\[ d\Gamma_{\nu_{\beta} \rightarrow W_{l_\beta}} = \frac{(2\pi)^4 g^2 V}{8 V^3 E^W E^p \Omega_H} \delta(\vec{k}_S - \vec{p}_S - \vec{k}_D - \vec{p}_D) \delta(E_D - \Omega) \]

(2.50)

is the charged lepton production rate from the reaction \( \nu_{\beta} \rightarrow W_{l_\beta} \) for a flavor neutrino of energy \( \Omega \) and \( P_{\alpha \rightarrow \beta}(t) \) are the disappearance (\( \alpha = \beta \)) or appearance (\( \alpha \neq \beta \)) transition probabilities obtained from the usual quantum mechanical calculations of Rabi-oscillations,

\[ P_{e \rightarrow e} = 1 - \sin^2(2\theta) \sin^2 \left( \frac{\delta m^2}{4\Omega} t \right) \]

\[ P_{e \rightarrow \mu} = \sin^2(2\theta) \sin^2 \left( \frac{\delta m^2}{4\Omega} t \right) \]

(2.51)

The remarkable aspect of the final result (2.50) is the factorization of the different processes contributing to the far detector event rate, namely the number of events at the source multiplies the quantum mechanical transition probability which in turn multiplies the production rate at the vertex in the far detector. This factorization is a distinct consequence of the two time analysis, of the disentanglement of the neutrino near the production region along with the approximations invoked in the resolution of the energy conserving delta functions. We emphasize that the factorization in terms of the quantum mechanical transition probabilities \( P_{\alpha \rightarrow \beta}(t) \) only applies to the detection rate defined by taking the time derivative. The total number of events per phase space volume also factorizes but not in terms of the quantum mechanical transition probabilities but in terms of their integral in time.

D. Possible corrections for short baseline experiments

This analysis also suggests that potentially important corrections may arise in short baseline experiments for much lower energy. In this case, it may occur that the time scale of disentanglement is of the same order as the oscillation scale \( \Omega/\delta m^2 \) so that \( t_S \delta m^2/\Omega \sim 1 \) and the sine functions cannot be factored out of the quantum state as in (2.32). In this case, the resulting disentangled state is

\[ \left| \nu_e(\vec{Q}, t) \right| \propto \left\{ \sin \theta \left| \nu_{\alpha, \vec{p}} \right| e^{-\Omega_{\alpha, \vec{p}}(t-t_S)} f_-(x, t_S, \Delta) + \cos \theta \left| \nu_{\alpha, \vec{p}} \right| e^{-\Omega_{\alpha, \vec{p}}(t-t_S)} f_+(x, t_S, \Delta) \right\}, \]

(2.52)

which is not of the familiar quantum mechanical form of a coherent superposition of the mass eigenstates with the usual \( \cos \theta, \sin \theta \) amplitudes but with the “diffraction functions” multiplying these factors. If the width of these functions is of order \( 2\pi/t_S \lesssim 2\Delta \), interference term in the probability will feature the product (2.19) which would lead to an extra interference terms of the form \( \sin(t_S \Delta)/t_S \Delta \) multiplying the \( \sin(2t \Delta)/2\Delta \) in (2.42,2.43). Such
factor yields an extra modulation with energy which may yield phenomenologically interesting modifications in the interpretation and analysis of data. Although the typical time scale $t_S$ for disentanglement near the source region is approximately the same for short and long baseline experiments, this discussion applies to the possibility in which $t_S$ is not too small compared to the oscillation time scale as could be the case in short-baseline experiments. For example at LSND the source region is a few meters across and the baseline is $\sim 30 \text{ m}$, and at KARMEN where the baseline is $\sim 17.5 \text{ m}$. The baseline and beam energy are designed so that at least an oscillation takes place along the baseline. However in these short-baseline experiments, the distance over which the neutrino state is disentangled from the charged lepton, a few meters, is not too much smaller than the baseline, consequently $t_S$ may not be much smaller than an oscillation time.

The potential phenomenological importance of these effects merit their further study.

E. Wave packet description

As mentioned earlier in the paper, we restricted ourselves to plane-wave states, the main reason being to discuss the main concepts within the simplest setting. Although we postpone a thorough discussion of the more technical aspects including the wave packet description to further study, we can provide physical arguments that allow the extrapolation of the main results obtained above to the wave packet case.

Consider that the initial particle, the charged leptons measured at the near and far detectors and the final particle are all described by wave packets with a macroscopic localization length scale $\sigma$ of the order of the typical scale of the detectors (a few meters) localized at the source and detectors respectively. These are sharply localized in momentum. The transition amplitudes between wave packets are now obtained from (2.35-2.37) by convolution with the wave functions of the initial state, the charged lepton measured near the source and those of the final state at the far detector.

The disentangled neutrino state (2.26) is now described as a propagating wave packet with a typical spatial localization length also of $O(\sigma)$, namely of macroscopic scale. These mildly localized wave packets lead to both momentum and energy uncertainty of the same order $\Delta E \sim \Delta p \sim h/\sigma \sim h/t_S \sim 10^{-7} \text{ eV}$. These uncertainties in energy and momentum are much larger than typical values for the neutrino energy differences $\delta m^2/O \sim 10^{-12} \text{ eV}$. As this wave packet propagates, the mass eigenstates will slowly separate because they feature different group velocities, and coherence will be maintained as long as the separation between the wave packets is much smaller than their localization lengths (neglecting dispersion). For a baseline $L$ and typical energies $O \sim 100 \text{ MeV}$, the separation at the far detector is $\propto (\Delta/O) L \sim 10^{-20}L$. Therefore, coherence is always maintained for macroscopic localization lengths in terrestrial experiments. The detection amplitudes (2.36,2.37) and probabilities (2.42,2.43) vanish until the time when the front of the packet overlaps with the far detector, namely for $t_D - t_S \approx L$, with an uncertainty of $O(\sigma)$, which for long-baseline experiments is always $\sigma \ll L$. The total number of events at the far detector is found by integrating the rate (2.50) from $t = L$, which is when the front of the wave packet overlaps with the detector until $t = L + \sigma$ when the wave packet has completely passed through the detector. For $L \gg \sigma$ and neglecting the separation of wave packets the total number of events at the
The delta functions will also be integrated with the wave packet profiles and volume factors are replaced by $\sigma^3$. Furthermore, at distances from the source $\gg \sigma$, the disentangled neutrino wave packet propagates as a spherical wave. Therefore, for far-detectors in the “radiation zone,” the probability will be suppressed by a geometric flux factor $1/L^2$ [24, 25]. Obviously, there are uncertainties of $O(\sigma)$ in this argument, but these are irrelevant for long-baselines $L \gg \sigma$. The $\sigma$ multiplying the total number of events in (2.53) is expected: it is the “total interaction time” during which the wave packet interacts with the detector. Obviously, these factors modify the overall normalization. However, the distortion of the energy spectrum in disappearance and the appearance probabilities for long-baseline experiments are determined by $P_{\alpha \rightarrow \beta}(L, E)$ (for short baseline there are potential corrections as discussed above). While we deem these arguments describing the wave-packet scenario to be physically sound, we will provide a thorough technical discussion of this case based on the results obtained above in the plane wave approximation in future work. The analysis presented above for the dynamics of disentanglement and propagation provides the fundamental basis to include the wave packet description, which is obtained from simple convolution of these results with wave packet wave functions.

III. CONCLUSIONS AND FURTHER DISCUSSION

In appearance and disappearance experiments, neutrinos produced by a charged current vertex at the source are entangled with the charged lepton. We study the dynamics of this quantum state directly as a function of time.

If the charged lepton (or daughter particle) produced at the source is not measured, tracing out this degree of freedom yields a density matrix for the neutrino. In the mass basis, the time evolution of the diagonal density matrix elements describe the production of mass eigenstates from the decay, whereas the off-diagonal density matrix elements exhibit the oscillations resulting from the interference of mass eigenstates and are a measure of coherence. We find that coherence remains up to a time of the order of the oscillation time scale during which diagonal and off diagonal matrix elements are of the same order.

The measurement of the charged lepton near the source (or its stopping at a nearby “beam dump”) disentangles the neutrino state, and it is the further time evolution of this disentangled state with the total Hamiltonian that leads to the production of charged leptons at the far detector. Thus, the process of production and detection in long-baseline experiments involves two different time scales: the measurement of the charged lepton near the source determines the first time scale at which the neutrino state is disentangled, while the measurement of the charged lepton at the far detector is the second time scale, which is much longer than the first one in long-baseline experiments.

In this time dependent description, we establish that the usual quantum mechanical state emerges if the disentanglement of the charged lepton produced with the neutrino occurs on time scales much shorter than the oscillation scale $\sim E/\delta m^2$. Under these circumstances, the event rate at the far detector factorizes in terms of the usual quantum mechanical probability if the final density of states is insensitive to the difference in the energies of the mass eigenstates in the ultrarelativistic limit $\delta m^2/E$. 

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Although in this article we focused on the study of the dynamics for the case of plane waves to exhibit the main results within the simplest setting, we provide physically motivated arguments to extrapolate the results to the case of wave packets, analyzing the effect of localization of the initial and final state wave functions at the near and far detector on the detection process. Combining the results obtained for plane waves with a wave packet description yields the total number of events detected in terms of the usual transition probability as a function of baseline. A deeper and more detailed analysis of wave packets and localization aspects will be provided in forthcoming work.

The analysis presented here also suggests that there could arise potentially interesting corrections in the case of short baseline experiments such as MiniBoone and/or LSND with baselines of $\sim 500, \sim 30$ meters and typical energies $\sim 1\text{GeV}; \sim 30\text{MeV}$, respectively, wherein the disentanglement of the neutrino could occur on time scales of the same order as the oscillation scale. In this case, the disentangled state will differ from the usual quantum mechanical coherent superposition and this difference introduces extra modulation with energy and could be important in the final detection rate. Such possibility could be relevant in the interpretation and analysis of data, in particular masses and mixing angles and certainly merits further detailed study.

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