THE CALABI METRIC AND DESINGULARIZATION OF EINSTEIN ORBIFOLDS

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Abstract. Consider an Einstein orbifold \((M_0, g_0)\) of real dimension \(2n\) having a singularity with orbifold group the cyclic group of order \(n\) in \(SU(n)\) which is generated by an \(n\)th root of unity times the identity. Existence of a Ricci-flat Kähler ALE metric with this group at infinity was shown by Calabi. There is a natural “approximate” Einstein metric on the desingularization of \(M_0\) obtained by replacing a small neighborhood of the singular point of the orbifold with a scaled and truncated Calabi metric. In this paper, we identify the first obstruction to perturbing this approximate Einstein metric to an actual Einstein metric. If \(M_0\) is compact, we can use this to produce examples of Einstein orbifolds which do not admit any Einstein metric in a neighborhood of the natural approximate Einstein metric on the desingularization. In the case that \((M_0, g_0)\) is asymptotically hyperbolic Einstein and non-degenerate, we show that if the first obstruction vanishes, then there does in fact exist an asymptotically hyperbolic Einstein metric on the desingularization. We also obtain a non-existence result in the asymptotically hyperbolic Einstein case, provided that the obstruction does not vanish. This work extends a construction of Biquard in the case \(n = 2\), in which case the Calabi metric is also known as the Eguchi-Hanson metric, but there are some key points for which the higher-dimensional case differs.

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1. Introduction

Let \( \Gamma_n = \{ e^{\frac{2\pi ki}{n}} : k = 0, \ldots, n - 1 \} \) and consider the action of \( \Gamma_n \) on \( \mathbb{C}^n \) by
\[
e^{\frac{2\pi ki}{n}} \cdot (z^1, \ldots, z^n) = (e^{\frac{2\pi ki}{n}} z^1, \ldots, e^{\frac{2\pi ki}{n}} z^n).
\]
The Calabi metric \( g_{\text{cal}} \) is a Ricci-flat Kähler asymptotically locally Euclidean (ALE) metric on the total space \( X \) of the line bundle \( \mathcal{O}(-n) \to \mathbb{P}^{n-1} \), which is \( U(n) \)-invariant. The group at infinity is \( \Gamma_n \subset SU(n) \).

Let \( (M_0, g_0) \) denote an Einstein orbifold of real dimension \( m = 2n \) satisfying
\[
Ric(g_0) = \Lambda g_0,
\]
and which has a singular point with group \( \Gamma_n \). One can desingularize the orbifold metric by gluing on a Calabi metric in the following manner.

Since \( g_{\text{cal}} \) is Ricci-flat and ALE, there exists a coordinate system \( \{ x^i \}, i = 1 \ldots 2n, \) and a diffeomorphism
\[
\Psi : (\mathbb{R}^{2n} \setminus B(0, R))/\Gamma_n \to X \setminus K,
\]
where \( B(0, R) \) is a ball of radius \( R \), and \( K \) is a compact subset of \( X \) so that
\[
\Psi^* g_{\text{cal}} = g_{\text{euc}} + O(r^{-2n}),
\]
\[
\partial^\alpha \Psi^* g_{\text{cal}} = O(r^{-2n-|\alpha|}),
\]
where \( \alpha \) is a multi-index of order \( |\alpha| \), \( r = |x| \), and \( g_{\text{euc}} \) denotes the Euclidean metric. In the following we extend \( r \) to a globally defined function on \( X \) which satisfies \( 0 < r < R \) on the interior of the compact subset \( K \).

Similarly, since \( p_0 \) is a smooth orbifold point of \( (M_0, g_0) \), there exists a coordinate system \( \{ z^i \}, i = 1 \ldots 2n, \) and a diffeomorphism
\[
\tilde{\Psi} : B(0, \epsilon) \setminus \{ 0 \} \to B(p_0, \epsilon) \setminus \{ p_0 \}
\]
such that
\[
\tilde{\Psi}^* \pi^* g_0 = g_{\text{euc}} + O(|z|^2),
\]
where \( \pi : B(p_0, \epsilon) \to B(p_0, \epsilon) \) is the universal covering mapping, and such that \( \tilde{\Psi}^* \pi^* (g) \) extends to a smooth metric on \( B(0, \epsilon) \).

Let \( t > 0 \) be small, and consider the following regions
\[
X^t = \{ x \in X : r < 2t^{-1/4} \}
\]
\[
M^t_0 = M \setminus B(p_0, (1/2)t^{1/4})
\]
Define a mapping
\[
\phi_t : \{ x \in X \mid (1/2)t^{-1/4} < r < 2t^{-1/4} \} \to \{ m \in M_0 \mid (1/2)t^{1/4} < |z| < 2t^{1/4} \}
\]
by \( \phi(x) = \sqrt{t}x \). Define \( M^t \) to be the union of \( X^t \) and \( M^t_0 \) along the attaching map \( \phi_t \).

**Lemma 1.1.** If \( A \in SO(2n) \) satisfies \( A\gamma = \gamma A \) for \( \gamma \) a generator of \( \Gamma_n \), then \( A \in U(n) \), if \( n \geq 3 \).
This lemma is elementary, and the proof is omitted. It illustrates a key difference between the case \( n = 2 \) and the higher-dimensional cases: if \( n = 2 \), then one can modify the attaching map \( \phi \) by 2 nontrivial rotational parameters, but there is no rotational freedom if \( n > 2 \) (since the Calabi metric is \( U(n) \)-invariant). Another consequence of this lemma is that there is a well-defined almost complex structure \( J \) at \( p_0 \). The corresponding Kähler form at \( p_0 \), will be denoted by \( \omega_{p_0} \).

An approximate Einstein metric on \( M^t \) is defined as follows. Let \( \chi \) be a smooth function satisfying \( 0 \leq \chi \leq 1 \),

\[
\chi(s) = \begin{cases} 
1 & s \leq 1/2 \\
0 & s \geq 2
\end{cases},
\]

and define \( \chi_t(s) \equiv \chi(t^{1/4}s) \). Then

\[
g_t = \begin{cases} 
g_0 & M_0 \setminus B(p_0, 2t^{1/4}) \\
(1 - \chi_t(r))\phi^*_t g_0 + \chi_t(r) t \cdot g_{cal} & \left(1/2\right)t^{-1/4} < r < 2t^{-1/4} \\
t \cdot g_{cal} & r < \left(1/2\right)t^{-1/4}
\end{cases}
\]

is a natural “approximate” Einstein metric on \( M^t \).

\[\text{Figure 1.1. The desingularization procedure.}\]

We define a weight function on \( M^t \) as follows.

\[
w = \begin{cases} 
1 & z \in M_0 \setminus B(p_0, i_{g_0}) \\
|z| & \left(1/2\right)t^{1/4} < |z| < \frac{1}{2}i_{g_0} \\
rt^{1/4} & 2 < r < 2t^{-1/4} \\
t^{1/4} & r \leq 1
\end{cases},
\]

where \( i_{g_0} \) is the injectivity radius of \( g_0 \), and such that \( w \) is increasing in \( r \) for \( 1 < r < 2 \), and increasing in \( |z| \) for \( \frac{1}{2}i_{g_0} < |z| < i_{g_0} \).
Our first result is in the case of a compact Einstein orbifold, which is an extension of [Biq13, Theorem 1.2]. We define

\[ \langle R(\omega), \omega \rangle(p_0) = \sum_{i,j,k,l} R_{ijkl}(p_0)(\omega_{p_0})_{ij}(\omega_{p_0})_{kl}, \]

and let \( R(p_0) \) denote the scalar curvature at \( p_0 \).

**Theorem 1.1.** Let \((M_0, g_0)\) be a compact Einstein orbifold having a singular point \( p_0 \) with orbifold group \( \Gamma_n \subset SU(n) \) with \( n \geq 3 \), and assume that

\[ n \langle R(\omega), \omega \rangle(p_0) + 2(n - 2)R(p_0) \neq 0. \]

Then there exist constants \( \epsilon > 0 \) and \( \delta_0 > 0 \) small, such that there does not exist any Einstein metric \( g_E \) near \( g_t \) satisfying

\[ \| \nabla^\ell(g_t - g_E) \|_{C^0(g_t)} < \epsilon w^{-\delta_0 - \ell}, \]

for \( 0 \leq \ell \leq 3 \), and sufficiently small gluing parameter \( t \).

This theorem does give some new understanding of possible degenerations of Einstein metrics, but we do not obtain an existence result in the compact case because this gluing problem is obstructed, and the criterion in (1.15) is only the first obstruction. To find an Einstein metric in this manner, one would have to show the vanishing of infinitely many obstructions (both of a local and global nature). This would be difficult, if not impossible, to check in any given case. However, it is possible that this procedure could work in a specific example, an example of which is the following (and is a generalization of the well-known Kummer construction in complex dimension 2).

**Example 1.2.** This example is due to Peng Lu [Lu99] (and is proved using a gluing argument for Kähler-Einstein metrics). Let \( \Sigma \) denote the elliptic curve with a non-trivial automorphism of order 3. Consider the 6-torus \( \Sigma \times \Sigma \times \Sigma \). Let \( \mathbb{Z}_3 \) act diagonally on the product. Then \( T^6/\mathbb{Z}_3 \) has 27 singular points of type \( \Gamma_3 \). The flat metric on the quotient can be resolved to a Calabi-Yau metric on the resolution by attaching a Calabi metric at each of the 27 singular points. Clearly, the obstruction (1.15) is satisfied at the singular points, since the orbifold metric is flat.

Theorem 1.1 implies several non-existence results. The first “non-example” is the following.

**Non-example 1.3.** Let \((M, g)\) be a compact Kähler-Einstein orbifold with Einstein constant \( \Lambda \neq 0 \), and singular point \( p_0 \) of type \( \Gamma_n \) and \( n \geq 3 \). In the Kähler-Einstein case, \( \langle R(\omega), \omega \rangle = 2R \) (see [Bes08, page 77]), so the obstruction in (3.32) is equal to \( 4(n - 1)R(p_0) \). Consequently, if \( \Lambda \neq 0 \), then there does not exist any Einstein metric near the approximate metric \( g_t \), for \( t \) sufficiently small.

This “non-example” should be contrasted with the case of \( n = 2 \), in which case there are known examples of sequences of positive Kähler-Einstein metrics limiting to orbifolds with Eguchi-Hanson metrics bubbling off [Tia91, OSS16]. The “explanation” is that for \( n = 2 \), the rotational parameters come into play, and the obstruction can
be made to vanish upon rotating the complex structure, but this phenomenon does not happen in higher dimensions.

**Non-example 1.4.** The football metric $g_S$ on $S^{2n}/\Gamma$ is defined as the quotient of the round metric on $S^{2n}$ by $\Gamma$, where $\Gamma$ acts in the above manner on the first $2n$-coordinates of $\mathbb{R}^{2n+1}$. Since $g_S$ has constant curvature equal to 1, the obstruction (1.15) does not vanish at either of the singular points. Consequently, there does not exist any Einstein metric near the approximate metric $g_t$, for $t$ sufficiently small.

We emphasize that Theorem 1.1 only can be used to rule out an Einstein metric near the specific approximate metric $g_t$ defined above. Of course, if the first obstruction in (1.15) is non-zero, it is possible that there could be some other procedure to desingularize the orbifold points which could produce an Einstein metric, but the above approximate metric seems to be the most “natural”.

1.1. **Asymptotically hyperbolic Einstein metrics.** We next consider a class of complete, non-compact Riemannian metrics.

**Definition 1.5.** Let $\overline{M} = M \cup \partial M$ be a smooth, compact manifold with boundary, and let $g$ be a Riemannian metric on $M$. We say that $g$ is *conformally compact* if there exists a function $x \in C^\infty(\overline{M}, \mathbb{R})$ with $x > 0$ on $M$ such that

$$\partial M = \{ p \in \overline{M} : x(p) = 0 \},$$

and $\overline{g} = x^2 g$ extends to a Riemannian metric on $\overline{M}$ of class $C^2$ with $dx \neq 0$ on $\partial M$.

A function $x$ with the above properties is called a *boundary defining function*. Note that $\overline{g}|_{\partial M}$ is not well-defined since there might be another candidate for a boundary defining function $x$ but its conformal class $[\overline{g}|_{\partial M}]$ is well-defined and we call it the conformal infinity of $g$.

**Definition 1.6.** Let $(M, \partial M, g, x)$ be a smooth, conformally compact manifold. We say that $(M, \partial M, g, x)$ is *asymptotically hyperbolic (AH)* if $|dx|_{\overline{g}} = 1$, where $x$ is a boundary defining function as above. If moreover $g$ is Einstein with $\text{Ric}(g) = -(m-1)g$, then $g$ is called *asymptotically hyperbolic Einstein (AHE)*.

Note that the AH condition is independent of the defining function $x$. Furthermore, if $m = \text{dim}(M)$ is even, the metric $\overline{g}$ extends smoothly to $\overline{M}$ [CDLS05]. For more background on the theory of AHE metrics, we refer the reader to [Biq06, GL91, Lee06, Maz91b]. We say that $g$ is non-degenerate if the space of $L^2$ infinitesimal Einstein deformations is trivial (see Definition 2.2).

Our main existence result is the following, which is an extension of [Biq13, Theorem 1.1] to the higher-dimensional case.

**Theorem 1.2.** Let $(M_0, g_0)$ be an asymptotically hyperbolic Einstein orbifold having a singular point $p_0$ with orbifold group $\Gamma_n \subset \text{SU}(n)$, with conformal infinity $[g_\infty]$ and $n \geq 3$. If $g_0$ is non-degenerate and

$$n\langle \mathcal{R}(\omega), \omega \rangle(p_0) + 2(n-2)R(p_0) = 0,$$

then...
then there exists an AHE metric $\hat{g}_t$ on $M^4$ for all $t$ sufficiently small. The conformal infinity $[\hat{g}_\infty, t]$ of the Einstein metric is a small perturbation of $[g_\infty]$, in the sense that
\[(1.19) \quad \|\hat{g}_\infty, t - g_\infty\|_{C^2(\partial M)} < Ct.\]
for some constant $C$.

The Einstein metric $\hat{g}_t$ is close, in a weighted norm, to a certain perturbation of $g_t$ which changes the conformal infinity. For the precise statement, see Section 9. The only “explicit” example of this phenomenon that we know of is the following.

**Example 1.7.** In [PP87], Page and Pope wrote down an explicit 1-parameter family of $U(n)$-invariant Einstein metrics $\mathcal{O}(-n) \to \mathbb{P}^{n-1}$, see also [BB82]. In [MS07], it was pointed out by Mazzeo and Singer that at one end of the parameter space, the Page-Pope family degenerates to an Einstein orbifold, with a Ricci-flat ALE metric bubbling off (which is in fact the Calabi metric). It can be shown that the limiting orbifold is the quotient of a Page-Pope metric on $\mathbb{D}^{2n}$, which also comes in a 1-dimensional family, and the parameter is exactly that for which the obstruction (1.18) is satisfied.

Of course, Theorem 1.2 could be used to produce the Page-Pope metrics on $\mathcal{O}(-n)$, but since these metrics are already known explicitly, it is not necessary to analyze this example in more detail here.

We note that the boundary conformal class in the previous example is that of a Berger sphere. The boundary Berger sphere conformal classes in the Page-Pope family on $\mathcal{O}(-n)$ are those with smaller length of the Hopf fiber than the one for which the obstruction vanishes. The ones with longer fiber length probably do not have any AHE filling, but the proof of this would require a more detailed analysis than given here (in particular, it would require a computation of the second obstruction).

There is also a corresponding non-existence result in the AHE case, provided that the obstruction in (1.18) does not vanish. But to state this properly, we need the following. Use the boundary defining function $x$ to identify a neighborhood of the boundary, call it $U_2\epsilon$ with $(0,2\epsilon) \times \partial M_0$, and if necessary, we modify the defining function $x$ to a new defining function $\tilde{x}$ satisfying
\[(1.20) \quad \tilde{x}(z) = \begin{cases} x & z \in U_2\epsilon \\ 2\epsilon \leq \tilde{x} \leq 1 & z \in M_0 \setminus \{U_2\epsilon\}. \end{cases} \]

Given a boundary metric $\tilde{g}_\infty$ close to $g_\infty$ using the identification above we extend $\tilde{g}_\infty$ to an asymptotically hyperbolic metric on $M_0$ by
\[(1.21) \quad \tilde{g} = g + \chi(x\epsilon^{-1})(\tilde{g}_\infty - g_\infty).\]

Our non-existence result in the AHE case is the following.

**Theorem 1.3.** Let $(M_0, g_0)$ be an asymptotically hyperbolic Einstein orbifold having a singular point $p_0$ with orbifold group $\Gamma_n \subset SU(n)$, with conformal infinity $[g_\infty]$ and $n \geq 3$. Assume that
\[(1.22) \quad n\langle R(\omega), \omega \rangle(p_0) + 2(n - 2)R(p_0) \neq 0.\]
Fix $\delta_0 > 0$ and $\delta_\infty > 0$ small. Then given $\epsilon > 0$, there exists a constant $\delta > 0$ such that if
\begin{equation}
\| \tilde{g}_\infty - g_\infty \|_{C^2(\partial M)} < \delta,
\end{equation}
then there is no AHE metric $g_E$ near $\tilde{g}_t$ satisfying
\begin{equation}
\| \nabla^\ell (\tilde{x}^{-\delta_\infty} (\tilde{g}_t - g_E)) \|_{C^0(g_t)} < \epsilon^\ell_{\delta_0 - \ell}
\end{equation}
for $0 \leq \ell \leq 3$, and sufficiently small gluing parameter $t$.

Using this result, we have the following “non-example”.

**Non-example 1.8.** Let $(\mathbb{D}^{2n}, g_H)$ denote the hyperbolic space of real dimension $2n$. This is an AHE metric, with conformal infinity the round sphere $S^{2n-1}$. The group $\Gamma_n$ acts on $\mathbb{D}^{2n}$, and the hyperbolic metric descends to the quotient. Since $g_H$ has constant curvature equal to $-1$, the obstruction (1.22) does not vanish at the singular point. Therefore, there exists no AHE metric near the approximate metric $\tilde{g}_t$, with boundary conformal class near the round metric.

The obstruction can be interpreted as defining a “wall” in the space of boundary conformal classes. That is, in the space of conformal infinities which have orbifold Einstein fillings, the ones which have the obstruction (1.18) vanishing at the orbifold point should generically define a hypersurface. It is expected that only the conformal classes on one side of this “wall” have smooth fillings. This has been completely analyzed by Biquard in the 4-dimensional case [Biq16]. However, due to space limitations, we will not go into more detail about the wall-crossing in the higher-dimensional case in this paper.

We end the introduction with an outline of the paper. In Section 2 we will recall some basic details about the Calabi metric, then we will show that the space of decaying infinitesimal Einstein deformations is one-dimensional. In Section 3 we will identify the obstruction for finding a solution of the gauged linearized Einstein equation with prescribed quadratic leading term, and also give a useful reformulation of the obstruction. In Section 4 we will define the “refined” approximate metric which is an improvement of the “naïve” approximate metric which was defined above. Then in Section 5 we will present the main Lyapunov-Schmidt reduction which yields a solution of the Einstein equation modulo the space of obstructions, which we will use to prove Theorem 1.1.

The remainder of the paper will then focus on the asymptotically hyperbolic Einstein case. In Section 6 we will give some background on the theory of AHE Einstein metrics, and in particular compute the indicial roots at infinity for several important operators. In Section 7 we will define a mapping from germs to boundary tensors, which considers solutions which blow-up at $p_0$ at a certain rate with prescribed leading term. This will be used in Section 8 together with a duality argument, to prove that there is a solution of the linearized equation with a certain prescribed quadratic leading term at $p_0$, and which is bounded at infinity (in the interior metric). This solution will then be used in Section 9 to prove the main existence theorem, Theorem 1.2. Finally, in Section 9 the non-existence result Theorem 1.3 will be proved.
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2. **Infinitesimal Einstein deformations of the Calabi metric**

We begin with a brief description of the Calabi metric.

### 2.1. The Calabi metric.

The cone $\mathbb{C}^n/\Gamma_n$ admits a crepant resolution $X$ which is the total space of the bundle $\mathcal{O}(-n) \to \mathbb{P}^{n-1}$. In [Cal79], Calabi proved the following.

**Theorem 2.1.** The space $X$ admits a $U(n)$-invariant Ricci-flat Kähler ALE metric denoted by $g_{\text{cal}}$.  

**Proof.** This metric is well-known, so we only present some details in the proof that we need, following the presentation of [KR97]. Let  

$$u = z_1^1 + \ldots + z_n^1,$$

and  

$$F : \mathbb{C}^n \to \mathbb{R},$$

be a radial Kähler potential to be determined later. The corresponding Kähler metric is  

$$g_{\overline{j}j} = \partial_i \partial_{\overline{j}} F = \partial_i (z^j F'(u)) = \delta_{ij} F'(u) + z^j \overline{z}^i F''(u).$$

For a Kähler metric, the components of the Ricci tensor are given by  

$$R_{\overline{j}j} = -\partial_i \partial_{\overline{j}} \log \det(g_{\alpha \overline{\beta}}).$$

A computation shows that a solution of $R_{\overline{j}j} = 0$ is given by  

$$\det(g_{\alpha \overline{\beta}}) = (F')^{n-1}(F'' + u F''') = 1.$$  

By solving (2.3) we obtain  

$$F(u) = \int_1^u (a^n + x^n)^{\frac{1}{n}} dx,$$

$$F'(u) = \frac{(a^n + u^n)^{\frac{1}{n}}}{u},$$

$$F''(u) = \frac{u^{n-1}(a^n + u^n)^{\frac{1-n}{n}}}{u} - \frac{(a^n + u^n)^{\frac{1}{n}}}{u^2}.$$  

We will fix the constant of integration $a = 1$. The metric then becomes  

$$g_{\overline{j}j} = \frac{(1 + u^n)^{\frac{1}{n}}}{u} \left\{ \delta_{ij} - \frac{1}{u(1 + u^n)} z^j \overline{z}^i \right\}.$$  

We next write the metric in radial form. Consider the following 1-forms  

$$dr, J(dr),$$
where $dr$ is the radial form and $J$ is the standard complex structure. We denote the dual vectors (with respect to the Euclidean metric) by
\begin{equation}
\partial_r, J(\partial_r).
\end{equation}
Let $\theta = -J(\frac{dr}{r})$ and $\partial_{\theta}$ denote the dual of $\theta$ (with respect to the Euclidean metric). It follows that
\begin{equation}
J(r \partial_r) = \partial_{\theta}.
\end{equation}
Since the Calabi metric is $U(n)$-invariant, we only need to compute the metric at the point $(r, 0, 0, \ldots, 0)$. On this axis, we have
\begin{equation}
\partial_r = \partial_1 + \overline{\partial}_1,
\end{equation}
and
\begin{equation}
g(\partial_r, \partial_r) = g(\partial_1 + \overline{\partial}_1, \partial_1 + \overline{\partial}_1) = 2g_{11} = 2u^{n-1}(1 + u^n)^{\frac{1-n}{n}},
\end{equation}
and also
\begin{equation}
g(\partial_{\theta}, \partial_{\theta}) = g(J(r \partial_r), J(r \partial_r)) = g(r \partial_r, r \partial_r) = 2u^n(1 + u^n)^{\frac{1-n}{n}}.
\end{equation}
Also at $(r, 0, 0, \ldots, 0)$ and for $2 \leq i \leq 2n - 2$ it follows that
\begin{equation}
g(\partial_i + \overline{\partial}_i, \partial_i + \overline{\partial}_i) = 2(1 + u^n)^{\frac{1}{n}}.
\end{equation}
Finally we obtain the metric in radial coordinates
\begin{equation}
2g_{\text{cal}} \equiv g = 2u^{n-1}(1 + u^n)^{\frac{1-n}{n}} dr^2 + 2(1 + u^n)^{\frac{1}{n}} g_{FS} + 2u^n(1 + u^n)^{\frac{1-n}{n}} \theta^2,
\end{equation}
where $u = r^2$, and $g_{FS}$ denotes the Fubini-Study metric on $\mathbb{P}^{n-1}$. There is an apparent singularity at $u = 0$, which can be resolved by replacing the origin with the $\mathbb{P}^{n-1}$ as the zero section of $O(-n) \rightarrow \mathbb{P}^{n-1}$, and the metric then extends smoothly to the resolution. $\square$

We will next set some notation which will be used below. Let $\omega$ denote the Fubini-Study 2-form on $\mathbb{P}^{n-1}$. Recall that on $\mathbb{P}^{n-1}$ we have
\begin{equation}
d\theta = 2\omega, \quad d\omega = 0,
\end{equation}
where $\theta$ is viewed as a connection form for the Hopf fibration $S^{2n-1} \rightarrow \mathbb{P}^{n-1}$. Next, we define
\begin{equation}
A(r) = \frac{r^{n-1}}{(1 + r^{2n})^{\frac{n}{2n}}} - 1, \quad B(r) = (1 + r^{2n})^{\frac{1}{n}}, \quad C(r) = (1 + r^{2n})^{\frac{1-n}{2n}} r^n,
\end{equation}
so that
\begin{equation}
\|Adr\|_{g_{\text{cal}}} = \|C\theta\|_{g_{\text{cal}}} = 1, \quad \|B\omega\|^2_{g_{\text{cal}}} = n - 1.
\end{equation}
We also define
\begin{equation}
d\bar{r} \equiv Adr, \quad \bar{\theta} \equiv C\theta, \quad \overline{\omega} \equiv B\omega, \quad g_{FS} \equiv Bg_{FS},
\end{equation}
\footnote{There is a typo in KR97 for $g_{\text{cal}}$ in radial form.}
so that the Calabi metric may be written as
\begin{equation}
    g_{\text{cal}} = d\tau \otimes d\tau + \rho F S + \vartheta \otimes \bar{\vartheta}.
\end{equation}

2.2. Infinitesimal Einstein deformations. In this section we prove that the space of decaying infinitesimal Einstein deformations of \((X, g_{\text{cal}})\) is one-dimensional and we explicitly identify a generator. We first define the operators that we need. Let \((M^m, g)\) be a Riemannian manifold of real dimension \(m\). For a 1-form \(\omega\),
\begin{align*}
    \delta_g : \Omega^1(M) &\to \Omega^0(M) \quad \delta_g(\omega) = -g^{ij} \nabla_i \omega_j \\
    \delta^*_g : \Omega^1(M) &\to S^2(T^*M) \quad \delta^*_g(\omega) = \frac{1}{2} (\nabla_i \omega_j + \nabla_j \omega_i) \\
    K_g : \Omega^1(M) &\to S^2_0(T^*M) \quad K_g(\omega) = \delta^*_g(\omega) + \frac{1}{m} \delta_g(\omega) g.
\end{align*}
For a symmetric 2-tensor \(h\),
\begin{align*}
    \delta_g : S^2(T^*M) &\to \Omega^1(M) \quad (\delta_g(h))_{ij} = -g^{ik} \nabla_j h_{ki} \\
    B_g : S^2(T^*M) &\to \Omega^1(M) \quad B_g(h) = \delta_g(h) + \frac{1}{2} d(\text{tr}_g h) \\
    P_g : S^2(T^*M) &\to S^2(T^*M) \quad P_g(h) = \frac{1}{2} \nabla^* \nabla h - \hat{R}(h),
\end{align*}
where \(\nabla^*\) is the formal \(L^2\)-adjoint of \(\nabla\), and \(\hat{R}\) is the action of curvature tensor on symmetric two tensors defined by
\begin{equation}
    \hat{R}(h)_{ij} = g^{kp} g^{lq} R_{iklj} h_{pq}.
\end{equation}
We note that the operator \(\nabla^* \nabla = -\Delta\), is the negative rough Laplacian of \(h\). For simplicity of notation we will often omit the metric \(g\) when it is clear from context. We also recall the following definition from [Bes08].

**Definition 2.2.** An infinitesimal Einstein deformation of an Einstein metric \(g\) is a symmetric 2-tensor field \(h\) such that
\begin{equation}
    \delta_g h = 0, \quad P_g(h) = 0, \quad \text{tr}_g(h) = 0.
\end{equation}
We say that a tensor \(h\) is decaying if \(h = O(r^{-\epsilon})\) as \(r \to \infty\) for some \(\epsilon > 0\).

**Theorem 2.1.** The space of decaying infinitesimal Einstein deformations of the Calabi metric is one-dimensional, and is spanned by the element
\begin{equation}
    o = \frac{1}{1 + r^{2n}} \left( -\rho F S + (n - 1) d\tau \otimes d\tau + (n - 1) \vartheta \otimes \bar{\vartheta} \right).
\end{equation}
To prove this, we will first characterize the decaying harmonic \((1, 1)\)-forms on the Calabi metric, which will then be used to characterize the decaying infinitesimal Einstein deformations. Recall that \(n\) denotes the complex dimension, while \(m = 2n\) denotes the real dimension.

**Proposition 2.3.** On \((X, g_{\text{cal}})\), the space of decaying harmonic \((1, 1)\)-forms is one-dimensional and is spanned by the element
\begin{equation}
    \Omega = \frac{1 - n}{1 + r^{2n}} d\tau \wedge \vartheta + \frac{1}{1 + r^{2n}} d\bar{\tau} \wedge \bar{\vartheta}.
\end{equation}
Proof. Define
\begin{equation}
\mathcal{H}^{p,q} = \{ \eta \in C^\infty(\Lambda^{p,q}X) \mid \eta = O(r^{1-2n}), d\eta = 0, \delta\eta = 0 \},
\end{equation}
and let $H^{p,q}(X)$ denote the image of $\mathcal{H}^{p,q}(X)$ in $H^{p+q}(X, \mathbb{C})$ under the natural mapping $\eta \mapsto [\eta]$. By [Joy00, Theorem 8.4.2], the decomposition
\begin{equation}
H^2(X, \mathbb{C}) = H^{2,0} \oplus H^{1,1} \oplus H^{0,2}
\end{equation}
holds, which implies that
\begin{equation}
\dim(H^{1,1}) = 1.
\end{equation}
By a standard argument (see for example [HV16, Section 3]), if $\eta \in C^\infty(\Lambda^{p,q}X)$ satisfies $\eta = O(r^{-\epsilon})$ for some $\epsilon > 0$, then $\eta = O(r^{-2n+1})$ as $r \to \infty$.

The remainder of the proof is therefore to show that the form in (2.24) is both closed and co-closed, since it is obviously a $(1,1)$-form. Using (2.16) it is easy to see that $\Omega$ is closed. To prove co-closedness, we recall that for a Kähler manifold $(M^n, g_0, J_0, \omega_0)$ ($n$ is complex dimension) we have the following (see [MK06])
\begin{equation}
dV_{g_0} = \frac{\omega_0^n}{n!}.
\end{equation}
Using (2.28) it follows that
\begin{equation}
* \omega = \frac{1}{(n-2)!} B^{n-3} \omega^{n-2} \wedge d\bar{r} \wedge \bar{\theta}
\end{equation}
\begin{equation}
dV_g = \frac{B^{n-1} \omega^{n-1}}{(n-1)!} \wedge d\bar{r} \wedge \bar{\theta}.
\end{equation}
Now straightforward calculation using (2.16) and (2.29) shows that
\begin{equation}
\delta\Omega = - * d * \Omega = 0.
\end{equation}
Finally, it is easy to check that this form extends across the zero section to define a smooth form on $X$. \hfill \Box

Next, we complete the proof of Theorem 2.1. If $h$ is an infinitesimal Einstein deformation, then write
\begin{equation}
h = h_+ + h_-,
\end{equation}
where $h_+$ is the $J$-invariant part, and $h_-$ is the $J$-anti-invariant part. It is proved in [Bes08] that
\begin{equation}
\Delta \Omega = 0,
\end{equation}
\begin{equation}
\Delta I = 0,
\end{equation}
where $\Omega$ is the $(1,1)$-form associated to $h_+$, and $I$ is the section of $\Lambda^{0,1} \otimes T^{1,0}$ associated to $h_-$. Since the canonical bundle is trivial, the conjugate Hodge star operator applied to $I$ is a harmonic $(1,n-1)$-form. By [Joy00, Theorem 8.4.2], this implies that $h_- \equiv 0.$
By Proposition 2.3, the space of decaying infinitesimal Einstein deformations is spanned by
\( (2.34) \quad o(., .) = \Omega(I(., .)) \).

Observe that \( o \) is a symmetric tensor
\( (2.35) \quad o(Y_1, Y_2) = \Omega(I(Y_1), Y_2) = \Omega(I(Y_2), Y_1) = o(Y_2, Y_1) \).

Noting that \( I(\partial_r) = \partial_\theta \), where \( \partial_\theta, \partial_r \) are the duals of \( \theta, d\theta \) with respect to the Calabi metric, we obtain
\( (2.36) \quad o(\partial_\theta, \partial_\theta) = \Omega(I(\partial_\theta), \partial_\theta) = \frac{n-1}{1+r^{2n}} \)
\( (2.37) \quad o(\partial_r, \partial_r) = \Omega(I(\partial_r), \partial_r) = \frac{n-1}{1+r^{2n}}. \)

Also recall that \( \omega \) is the Fubini-Study form, and that \( I \) agrees with standard complex structure on \( \mathbb{P}^{n-1} \), so
\( (2.38) \quad \omega(., .) = g_{FS}(I(., .)) \quad \Rightarrow \quad \omega(I(., .)) = -g_{FS}(., .) \quad \Rightarrow \quad \varphi(I(., .)) = -g_{FS}(., .) \)
(for \( \varphi \) we have \( \langle \varphi, \varphi \rangle_{cal} = n - 1 \)) and \( o \) becomes
\( (2.39) \quad o = \frac{1}{1+r^{2n}} \left( -g_{FS} + (n-1)d\bar{\theta} \otimes d\theta + (n-1)\bar{\theta} \otimes \bar{\theta} \right), \)
and clearly
\( (2.40) \quad tr(o) = \frac{1}{1+r^{2n}}(-2(n-1) + 2(n-1)) = 0. \)

Now observe that we can rewrite \( o \) as:
\( (2.41) \quad o = \frac{1}{1+r^{2n}} \left( -g_{FS} - d\bar{\theta} \otimes d\theta + n(d\bar{\theta} \otimes d\theta) - \bar{\theta} \otimes \bar{\theta} + n(\bar{\theta} \otimes \bar{\theta}) \right), \)

Also since \( \Omega \) is divergence-free (\( \Omega \) is co-closed) and \( I \) is parallel, it follows that \( o \) is divergence-free, so \( o \) is both traceless and divergence-free, as required. This completes the proof of Theorem 2.1.

We end this section with a slightly modified statement of Theorem 2.1.

**Theorem 2.2.** On \( (X, g_{cal}) \), the space of decaying solutions of
\( (2.42) \quad P_g h = 0 \)
\( (2.43) \quad tr_g h = 0, \)
is one-dimensional, and is spanned by the element \( o \).

**Proof.** The only thing we need to prove is that a solution of \( P_{cal} h = 0 \) which is decaying must also be divergence-free. To see this, since Calabi metric is Ricci-flat we get the following for linearized operator
\( (2.44) \quad d_{cal} Ric(h) = \frac{1}{2} \nabla_{cal}^* \nabla_{cal}(h) - \delta_{cal} B_{cal}(h). \)
Applying the divergence operator to the equation $P_{\text{cal}} h = 0$ yields the equation
\begin{equation}
P_1 \delta g h = 0,
\end{equation}
where $P_1 \equiv B \delta^*$. Since $g_{\text{cal}}$ is Ricci-flat, this implies that
\begin{equation}
\Delta_H (\delta g h) = 0,
\end{equation}
where $\Delta_H$ is the Hodge Laplacian. By [Joy00, Theorem 8.4.1], there are no decaying harmonic 1-forms on $X$, which implies that $\delta g h = 0$. \qed

3. Linearized equation on the Calabi metric

In this section we apply a technique in [Biq13] to find the obstruction for solving linearized equation with prescribed leading quadratic term. Let
\begin{equation}
H = H_{ijkl} x^i x^j dx^k \otimes dx^l
\end{equation}
be a given quadratic tensor defined on $\mathbb{R}^{2n}$ which is symmetric in the first two and last two indices. An easy computation shows that
\begin{equation}
B_{\text{euc}}(H) = (-2H_{k,jkl} + H_{j,klk})x^j dx^l.
\end{equation}
We also need the following basic lemma.

**Lemma 3.1.** Let $v$ be a symmetric 2-tensor with $\|v\|_{g_{\text{cal}}} = O(r^\delta)$ for some $\delta \in (2 - 2n, 0)$. Then there exists $u$ with $\|u\|_{g_{\text{cal}}} = O(r^{2+\delta})$ and $P_{g_{\text{cal}}}(u) = v$ if and only if
\begin{equation}
\langle v, o \rangle_{L^2(g_{\text{cal}})} = 0.
\end{equation}

**Proof.** This follows from standard theory of elliptic operators on weighted spaces (see [Bar86] for example). \qed

Next, we have the following proposition.

**Proposition 3.2.** Given $(H, \Lambda)$ with $H$ a $\Gamma_n$-invariant symmetric 2-tensor on $\mathbb{R}^{2n}$ satisfying
\begin{align}
\frac{1}{2}\Delta_{\text{euc}} H &= \Lambda g_{\text{euc}} \\
B_{\text{euc}} H &= 0,
\end{align}
there exists a solution $(h, \lambda)$ to
\begin{align}
P_{g_{\text{cal}}}(h) &= \Lambda g_{\text{cal}} + \lambda o \\
h &= H + O(r^{-2n+2+\epsilon}) \\
B_{g_{\text{cal}}}(h) &= 0,
\end{align}
if and only if
\begin{equation}
\lambda = -\frac{1}{\|o\|_{L^2}^2} \lim_{r \to \infty} \int_{S_r/\Gamma_n} \frac{n+1}{r} \langle H, o \rangle dS_r/\Gamma_n.
\end{equation}
**Proof.** Let \((h, \lambda)\) be a solution to (3.6). Let \(R\) be a fixed positive real number and consider the following cut-off function

\[
\chi(x) = \begin{cases} 
1 & |x| > 2R \\
0 & |x| < R 
\end{cases}
\]

Let \(h' = h - \chi H\), then \(h' = O(r^{-2n+2+\epsilon})\). We have

\[
P_{\text{gcal}}(h') = -P_{\text{gcal}}(\chi H) + \Lambda_{\text{gcal}} + \lambda o.
\]

Next, we claim that \(\Lambda_{\text{gcal}} - P_{\text{gcal}}(\chi H) = O(r^{-2n})\). To see this,

\[
P_{\text{gcal}}(\chi H) = -\frac{1}{2} \Delta(\chi H) - \delta \cdot R(H) = \Lambda_{\text{geuc}} + O(r^{-2n}),
\]

where we used that \(H_{ikl} = \delta_{kl}\). Using the fact that \(o\) is traceless,

\[
\langle P_{\text{gcal}}(\chi H) - \Lambda_{\text{gcal}}, o \rangle = \langle P_{\text{gcal}}(\chi H), o \rangle.
\]

Integrating by parts yields

\[
\int_{r \leq r_0} \langle P_{\text{gcal}}(\chi H), o \rangle dV_{\text{gcal}} = \frac{1}{2} \int_{r = r_0} \left( \langle \nabla H, \nabla o \rangle - \langle \nabla (\chi H), o \rangle \right) dS_{r_0/\Gamma_n}.
\]

Notice that we have the following asymptotic relations

\[
\nabla n o = \frac{-2n}{r} o + O(r^{-2n-2})
\]

\[
\nabla n (\chi H) = \frac{2}{r} H + O(1),
\]

as \(r \to \infty\), so taking the limit as \(r_0 \to \infty\) we obtain

\[
\int_X \langle P_{g_{\text{cal}}}(\chi H), o \rangle dV_{g_{\text{cal}}} = -\lim_{r \to \infty} \int_{S_r/\Gamma_n} \frac{n+1}{r} \langle H, o \rangle dS_{S_r/\Gamma_n}.
\]

Now applying Lemma 3.1 we should have

\[
\langle -P_{g_{\text{cal}}}(\chi H) + \Lambda_{g_{\text{cal}}} + \lambda o, o \rangle = 0
\]

which implies

\[
\lambda = -\frac{1}{||o||_{L^2}^2} \lim_{r \to \infty} \int_{S_r/\Gamma_n} \frac{n+1}{r} \langle H, o \rangle dS_{S_r/\Gamma_n}.
\]

Conversely, assume that (3.7) is satisfied. This implies that (3.16) is satisfied and by Lemma 3.1 there exist a solution \(h'\) for the following problem:

\[
-P_{g_{\text{cal}}}(\chi H) + \Lambda_{g_{\text{cal}}} + \lambda o = P(h').
\]

So \(h = h' + \chi H\) is the desired solution. Next, we show that \(h\) is in Bianchi gauge. Since the Calabi metric is Ricci-flat, we have

\[
0 = B_{g_{\text{cal}}} P_{g_{\text{cal}}} h = \frac{1}{2} \nabla^* \nabla B_{g_{\text{cal}}} h.
\]

Also,

\[
h = H + O(r^{-2n+2+\epsilon}) \implies B_{g_{\text{cal}}} h = O(r^{-2n+1+\epsilon}).
\]
Let $\omega = B_{g_{cal}} h$. Then integrating by parts implies

$$0 = \int_{B_r/\Gamma_n} \langle \omega, \Delta \omega \rangle dV_{g_{cal}} = \int_{B_r/\Gamma_n} |\nabla \omega|^2 dV_{g_{cal}} - \int_{S_r/\Gamma_n} \langle \omega, \nabla \tilde{r} \omega \rangle dS,$$

and taking the limit as $r \to \infty$ shows that $\nabla \omega = 0$ and the later implies that $\omega = 0$. □

3.1. Reformulation of the obstruction. We begin with some integral identities.

**Lemma 3.3.** On $\mathbb{R}^{2n}$ with coordinates $(x^1, \ldots, x^{2n})$ we have the following:

$$\int_{S^{2n-1}} x^i x^j dS = \frac{\omega_{2n-1}}{2n} \delta_{ij},$$

$$\int_{S^{2n-1}} x^i x^j x^k x^l dS = \frac{\omega_{2n-1}}{2n(2n+2)} (\delta_{kl} \delta_{ij} + \delta_{ki} \delta_{lj} + \delta_{kj} \delta_{il}).$$

**Proof.** This is proved, for example, in [Bre08]. □

Let $g$ be a Riemannian metric on $\mathbb{R}^{2n}$. The Taylor expansion of the metric in normal coordinates has the form

$$g_{kl}(z) = \delta_{kl} - \frac{1}{3} R_{ikjl} z^i z^j + O(|z|^3)$$

as $|z| \to 0$. The quadratic term is not in Bianchi gauge, so we use the following lemma.

**Lemma 3.4.** Assume that $g_0$ is Einstein of dimension $2n$, with $\text{Ric}(g_0) = \Lambda g_0$. Then there exists a coordinate system near the origin such that

$$g_{kl}(z) = \delta_{kl}(0) + H_{ijkl} z^i z^j + O(|z|^3),$$

as $|x| \to 0$, where

$$H_{ijkl} z^i z^j = -\frac{1}{3} R_{ikjl} z^i z^j - \frac{\Lambda}{3(n+1)} (\delta_{kl}|z|^2 + 2z^k z^l).$$

**Proof.** It is easy to verify that

$$B_{\text{euc}}(-\frac{1}{3} R_{ikjl} z^i z^j) = -\frac{2}{3} \Lambda z^i,$$

and that

$$B_{\text{euc}}(\delta_{ij}|z|^2 + 2z^i z^j) = -(2n + 2) z^i.$$

Consequently,

$$B_{\text{euc}}(H_{ijkl} z^i z^j) = 0.$$

Finally, we note that

$$\delta_{\text{euc}}^*(|z|^2 z^i) = \delta_{ij}|z|^2 + 2z^i z^j,$$

so the added term is a Lie derivative of a cubic vector field. A standard argument, see for example [AV12, Section 7], implies that there is a coordinate system with the stated property. □
**Proposition 3.5.** Let $H = H_{ijkl}x^i x^j dx^k dx^l$ be the symmetric 2-tensor with

\begin{equation}
H_{ijkl} = -\frac{1}{3} R_{ikjl} - \frac{\Lambda}{3(n+1)} (\delta_{ij}\delta_{kl} + 2\delta_{ik}\delta_{jl}).
\end{equation}

Then the constant $\lambda$ in Proposition 3.2 is given by

\begin{equation}
\lambda = \frac{1}{\|o\|^2} \left(2 - \frac{n}{2} \Lambda - \frac{1}{8}\langle R(\omega), \omega \rangle \right).
\end{equation}

**Proof.** Recall from (2.41) that

\begin{equation}
o = \frac{1}{1 + r^{2n}} \left(-g_{\text{cat}} + n(dr \otimes dr) + n(\bar{\theta} \otimes \bar{\theta}) \right),
\end{equation}

so in ALE coordinates, we have

\begin{equation}
o = \frac{1}{r^{2n}} \left(-g_{\text{euc}} + n(dr \otimes dr) + nr^2(\theta \otimes \theta) \right) + O(r^{-2n-1})
\end{equation}
as $r \to \infty$. Defining

\begin{equation}
o_{\text{euc}} = \frac{1}{r^{2n}} \left(-g_{\text{euc}} + n(dr \otimes dr) + nr^2(\theta \otimes \theta) \right),
\end{equation}

we have

\begin{equation}
\lim_{r \to \infty} \int_{S_r/\Gamma_n} \frac{n + 1}{r} \langle H, o \rangle dS_{S_r/\Gamma_n} = \int_{|x|=1} \langle n + 1, H, o_{\text{euc}} \rangle dS_{S_r/\Gamma_n}.
\end{equation}
The inner product is

\begin{equation}
\langle H, o_{\text{euc}} \rangle = -\langle H, g_{\text{euc}} \rangle + n\langle H, dr \otimes dr \rangle + n\langle H, r^2 \theta \otimes \theta \rangle.
\end{equation}

We lift up to $\mathbb{R}^{2n}$, and next we calculate each term in (3.37); the first term is

\begin{equation}
\langle H, g_{\text{euc}} \rangle = H_{ijkl} x^i x^j,
\end{equation}

so by (3.22),

\begin{equation}
\int_{|x|=1} \langle H, g_{\text{euc}} \rangle dS = \frac{\omega^{2n-1}}{2n} H_{ikk}.
\end{equation}

Also, since $r^2 = (x^i)^2$, we have

\begin{equation}
rdr = x^i dx^i
\end{equation}

\begin{equation}
r^2 dr \otimes dr = x^i x^j dx^k dx^l.
\end{equation}

Letting $J$ denote the standard complex structure on $\mathbb{R}^{2n}$, then $\theta = -J(\frac{dr}{r})$, so we have

\begin{equation}
(r\theta) \otimes (r\theta) = x^p x^q J^p_k J^q_l dx^k dx^l.
\end{equation}

Clearly,

\begin{equation}
r^2 \langle \theta \otimes \theta, H \rangle = x^i x^j x^p x^q J^p_k J^q_l H_{ijkl},
\end{equation}

\begin{equation}
r^2 \langle dr \otimes dr, H \rangle = x^i x^j x^k x^l H_{ijkl}.
\end{equation}
Using (3.23), this implies that
\[
\int_{|x|=1} \langle H, dr \otimes dr \rangle dS = \int_{|x|=1} H_{ijkl} x^i x^j x^k x^l dS
\]
(3.45)
\[
= \frac{\omega_{2n-1}}{2n(2n+2)} (\delta_{ij} \delta_{kl} + \delta_{ik} \delta_{jl} + \delta_{il} \delta_{jk}) H_{ijkl}
\]
\[
= \frac{\omega_{2n-1}}{2n(2n+2)} (H_{ikik} + H_{ikik} + H_{ikki}).
\]

Also, (3.23) implies
\[
\int_{|x|=1} \langle H, \theta \otimes \theta \rangle dS = \int_{|x|=1} H_{ijkl} x^i x^j x^p x^q J^p_k J^q_l dS
\]
(3.46)
\[
= \frac{\omega_{2n-1}}{2n(2n+2)} (\delta_{ij} \delta_{pq} + \delta_{ip} \delta_{jq} + \delta_{iq} \delta_{jp}) J^p_k J^q_l H_{ijkl}
\]
\[
= \frac{\omega_{2n-1}}{2n(2n+2)} (J^p_k J^q_l H_{ikil} + J^p_k J^q_l H_{ipkl} + J^p_k J^q_l H_{pqkl}).
\]

Putting together (3.39), (3.45), and (3.46), we obtain
(3.47)
\[
\int_{|x|=1} (n+1) \langle H, o_{euc} \rangle dS = -\frac{\omega_{2n-1}}{2n} (n+1) H_{kkll} + \frac{\omega_{2n-1}}{4} (H_{ikik} + H_{ikik} + H_{ikki})
\]
\[
+ \frac{\omega_{2n-1}}{4} (J^p_k J^q_l H_{ikil} + J^p_k J^q_l H_{ipkl} + J^p_k J^q_l H_{pqkl}).
\]

Next we compute the terms involving $H$ appearing in (3.47). First,
(3.48)
\[
H_{kkll} = -\frac{1}{3} R_{kkll} - \frac{\Lambda}{3(n+1)} (\delta_{kk} \delta_{ll} + 2 \delta_{kl} \delta_{kl})
\]
\[
= -\frac{1}{3} \Lambda (2n) - \frac{\Lambda}{3(n+1)} ((2n)^2 + 4n) = -2n \Lambda.
\]

The next two terms are
(3.49)
\[
H_{ikik} = -\frac{1}{3} R_{ikik} - \frac{\Lambda}{3(n+1)} (\delta_{ik} \delta_{ik} + 2 \delta_{ii} \delta_{kk}) = -\frac{\Lambda}{3(n+1)} (2n + 8n^2),
\]
and
(3.50)
\[
H_{ikki} = -\frac{1}{3} R_{ikki} - \frac{\Lambda}{3(n+1)} (\delta_{ik} \delta_{ik} + 2 \delta_{ik} \delta_{ik})
\]
\[
= \frac{1}{3} 2n \Lambda - \frac{\Lambda}{3(n+1)} 6n = \frac{2n \Lambda}{3} - \frac{2n \Lambda}{n + 1}.
\]

The first term involving $J$ is
(3.51)
\[
J^p_k J^q_l H_{ikl} = J^p_k J^q_l \left( -\frac{1}{3} R_{ikil} - \frac{\Lambda}{3(n+1)} (\delta_{ii} \delta_{kl} + 2 \delta_{ik} \delta_{il}) \right)
\]
\[
= -\frac{1}{3} \Lambda (2n) - \frac{\Lambda}{3(n+1)} ((2n)^2 + 4n) = -2n \Lambda.
\]
The next term is

\[ J^p_k J^q_l H_{pqkl} = J^p_k J^q_l \left( -\frac{1}{3} R_{pkql} - \frac{\Lambda}{3(n+1)} (\delta_{pq} \delta_{kl} + 2 \delta_{pk} \delta_{ql}) \right) \]

\[ = -\frac{1}{3} \langle R(\omega), \omega \rangle - \frac{2n\Lambda}{3(n+1)}. \]

(3.52)

For the last term, we compute

\[ J^p_k J^q_l H_{qpkl} = J^p_k J^q_l \left( -\frac{1}{3} R_{qkpl} - \frac{\Lambda}{3(n+1)} (\delta_{qp} \delta_{kl} + 2 \delta_{qk} \delta_{pl}) \right) \]

\[ = -\frac{1}{3} J^p_k J^q_l R_{qkpl} + \frac{2n\Lambda}{3(n+1)}. \]

(3.53)

Note that using the Bianchi identity, we have

\[ 0 = J^p_k J^q_l (R_{qkpl} + R_{kpql} + R_{pqkl}) \]

\[ = J^p_k J^q_l R_{qkpl} - J^k_p J^q_l R_{kpql} - J^p_k J^q_l R_{pqkl} \]

\[ = J^p_k J^q_l R_{qkpl} - J^k_p J^q_l R_{kpql} + J^p_k J^q_l R_{pqkl}. \]

(3.54)

The first and third terms in the last line of (3.54) are easily seen to be equal, thus we have

\[ J^p_k J^q_l R_{qkpl} = \frac{1}{2} \langle R(\omega), \omega \rangle. \]

(3.55)

Plugging (3.55) into (3.53), we obtain

\[ J^p_k J^q_l H_{qpkl} = -\frac{1}{6} \langle R(\omega), \omega \rangle + \frac{2n\Lambda}{3(n+1)}. \]

(3.56)

We now insert (3.48), (3.49), (3.50), (3.51), (3.52), and (3.56) into (3.47) to obtain

\[ \int_{|x|=1} (n+1) \langle H, o_{euc} \rangle dS = -\frac{\omega_{2n-1}}{2n} (n+1) (-2n\Lambda) \]

\[ + \frac{\omega_{2n-1}}{4} \left( -2n\Lambda - \frac{\Lambda}{3(n+1)} (2n + 8n^2) + \frac{2n\Lambda}{3} - \frac{2n\Lambda}{n+1} \right) \]

\[ + \frac{\omega_{2n-1}}{4} \left( -2n\Lambda - \frac{1}{3} \langle R(\omega), \omega \rangle - \frac{2n\Lambda}{3(n+1)} - \frac{1}{6} \langle R(\omega), \omega \rangle + \frac{2n\Lambda}{3(n+1)} \right). \]

(3.57)

This simplifies to

\[ \int_{|x|=1} (n+1) \langle H, o \rangle dS = \omega_{2n-1} \left( \frac{2-n}{2} \Lambda - \frac{1}{8} \langle R(\omega), \omega \rangle \right). \]

(3.58)

Finally, since we are integrating over \( S^{2n-1} / \Gamma \), where \( \Gamma \) is of order \( n \), the area of the unit sphere is \( \omega_{2n-1} / n \), which accounts for the extra factor of \( \frac{1}{n} \) appearing in (3.32). □
4. The refined approximate metric

The approximate metric described in the introduction will be referred to as the “naive” approximate metric. In this section, we will construct a “refined” approximate metric, which will be necessary to compute the obstruction. The construction closely follows that in [Biq13] with minor modifications. Recall that the Einstein orbifold is denoted by \((M_0, g_0)\) and \((X, g_{\text{cal}})\) denotes the Calabi metric. Also, recall that \(\{x^i\}, i = 1 \ldots 2n\), are ALE coordinates on \(X\), and that \(g_0\) satisfies
\[
Ric(g_0) = \Lambda g_0.
\]

Near \(p_0\), we use the coordinate system \(\{z^i\}, i = 1 \ldots 2n\), from Lemma 3.4, where \(g_0\) takes the form
\[
g_0 = g_{\text{euc}} + H + o(|z|^2),
\]
with
\[
H = H_{ijkl} z^i z^j dz^k dz^l,
\]
given by (3.26).

On the Calabi metric, using Proposition 3.2, there exists a solution \((h, \lambda)\) for the following problem
\[
P_{g_{\text{cal}}}Ric(h) = \Lambda g_{\text{cal}} + \lambda o, \quad h = H + O(r^{2-2n+\epsilon}),
\]
\[
B_{g_{\text{cal}}}(h) = 0,
\]
for any \(\epsilon > 0\), where \(\lambda\) is given by (3.32), as \(r \to \infty\). Define
\[
h_t = g_{\text{cal}} + th.
\]
Consider the Calabi metric on \(\mathbb{C}^n - \{0\}/\Gamma_n\) and define the following two regions
\[
DZ_t \equiv \{ \frac{1}{2} t^{-\frac{1}{4}} < r < 2t^{-\frac{1}{4}} \}
\]
\[
X_t \equiv \{ r < 2t^{-\frac{1}{4}} \},
\]
(DZ is short for damage zone). Now define the following function on \(\mathbb{C}^n - \{0\}/\Gamma_n\)
\[
\rho = \begin{cases} r & \text{when } r > 2 \\ 1 & \text{when } r \leq 1 \end{cases}
\]

**Lemma 4.1.** For \(t\) sufficiently small, \(h_t\) is a Riemannian metric on \(X_t\) and the following estimate holds on this region
\[
|\nabla^k (Ric_{h_t} - t\Lambda h_t - t\lambda o)|_{g_{\text{cal}}} \leq c_k t^2 \rho^{2-k},
\]
for some constants \(c_k\) and any \(k \geq 0\).
Proof. Since $h$ has quadratic leading term and $g_{\text{cal}}$ is ALE, there exist a constant $c$ so that $|h|_{g_{\text{cal}}}(p) \leq c \rho^2(p)$ for any $p \in X^t$. It follows that
\[
|th|_{g_{\text{cal}}}(p) \leq ct \rho^2(p) \leq 2ct^2.
\]
The right hand side is independent of $p$ so $h_t$ is a Riemannian metric on $X^t$ for small $t$. Recall that
\[
d_{g_{\text{cal}}}Ric(h) = \Lambda g_{\text{cal}} + \lambda o
\]
so we have
\[
Ric(g_{\text{cal}} + th) = Ric(g_{\text{cal}}) + td_{g_{\text{cal}}}Ric(h) + Q(th),
\]
where $Q$ is a nonlinear term. For any Riemannian metric $g$, and symmetric tensor $h$ small, recall the well-known expansion
\[
Rm_{g+h} = Rm_g + (g + h)^{-1} \nabla^2 h + (g + h)^{-2} \nabla h \ast \nabla h,
\]
where $\ast$ denotes various tensor contractions (see, for example [GV16, Section 3]). Contracting again, we obtain
\[
Ric_{g+h} = (g + h)^{-1} \nabla^2 h + (g + h)^{-2} \nabla h \ast \nabla h.
\]
It follows that for $t$ sufficiently small
\[
|Q(th)| \leq c_1 t^2(\|Rm_{g_{\text{cal}}}|h|^2 + |h|\|\nabla^2 h| + |\nabla h|^2).
\]
Using (4.4), it follows that
\[
|Ric(h_t) - t\Delta h_t - \lambda o| = c_1 t^2(\|Rm_{g_{\text{cal}}}|h|^2 + |h|\|\nabla^2 h| + |\nabla h|^2)
\]
Note that
\[
|\nabla^k h| \leq c_k \rho^{2-k},
\]
so the lemma for $k = 0$ follows. For $k > 0$ the proof is similar, and is omitted. \qed

Recalling the attaching map $\phi_t$ defined in (1.10), consider
\[
M^t = M_0^t \cup_{\phi_t} X^t
\]
where
\[
M_0^t = M \setminus B(p_0, (1/2)t^{1/4}).
\]
To define the refined approximate metric on $M^t$, let $\chi$ be as in (1.11),
\[
g_t = \begin{cases} 
g_0 & M \setminus B(p_0, 2t^{1/4}) \\
(1 - \chi_t(r))\phi_t^*g + \chi_t(r)th_t & (1/2)t^{-1/4} < r < 2t^{-1/4}
\end{cases}
\]
\[
th_t & r < (1/2)t^{-1/4}
\]
We have the following estimation.
Lemma 4.2. For each integer $k \geq 0$, there exist a constant $c_k$ so that $g_t$ satisfies the following estimate on $X^t$

\begin{equation}
|\nabla^k (Ric_{g_t} - \Lambda g_t - t\lambda \chi_t(\rho) o)|_{g_{cal}} \leq c_k t^2 \rho^{2-k},
\end{equation}

where the covariant derivatives and norm are taken with respect to $g_{cal}$.

Proof. If $r < \frac{1}{2} t^{-\frac{1}{4}}$ then $g_t = th_t$ and by Lemma 4.1 we are done. If

\begin{equation}
(1/2)t^{-1/4} < r < 2t^{-1/4}
\end{equation}

then $g_t = (1 - \chi_t(r))\phi_t^* g_0 + \chi_t(r)th_t$. Recall that

\begin{align*}
\phi_t^* g_0 &= tg_{euc} + t^2 H + tO(t^2 r^3) \\
th_t &= tg_{euc} + t^2 H + tO(r^{-2n}) + tO(r^{2-2n+\epsilon}).
\end{align*}

Since the quadratic terms agree, we have the estimate

\begin{equation}
g_t = tg_{euc} + t^2 H + tO(t^3 \rho^3).
\end{equation}

Writing

\begin{equation}
\theta = g_t - tg_{euc},
\end{equation}

expanding the Ricci tensor near $t \cdot g_{euc}$ we have

\begin{equation}
Ric_{g_t} = d_{t-euc}Ric(\theta) + Q_{euc}(\theta),
\end{equation}

where $Q_{euc}(\theta)$ satisfies the estimate

\begin{equation}
t^2 |Q_{euc}(\theta)| \leq c(|\theta||\nabla^2 \theta| + |\nabla \theta|^2),
\end{equation}

for some constant $c$, where the norms are with respect to $g_{euc}$. Note that \textcolor{red}{(4.22)} implies that

\begin{equation}
t\lambda \chi_t(\rho) o = O(t^2 \rho^2).
\end{equation}

From \textcolor{red}{(4.4)} it follows that

\begin{equation}
d_{t-euc}Ric(\theta) = t\Lambda g_{euc} + O(t^2 \rho^2),
\end{equation}

where $O$ is with respect to the Euclidean metric. Finally we obtain

\begin{equation}
Ric_{g_t} - \Lambda g_t - t\lambda \chi_t(\rho) o = O(t^2 \rho^2),
\end{equation}

so the lemma for $k = 0$ follows. For general $k$ the proof is similar and is omitted. \qed
5. Solution of the Einstein equation modulo obstructions

5.1. Function Spaces. In this section we define the function spaces that we will need in order to apply the implicit function theorem. Let $E$ be a tensor bundle over $X^t$ and $\delta_0 \in \mathbb{R}$ be a weight. We define $C^{k,\alpha}_{\delta_0}(X^t, E)$ to be

\[
|s|_\alpha \equiv \sup_{d(x,y) < \text{dist}} \frac{|s(x) - s(y)|}{d(x,y)^\alpha}
\]

\[
\|s\|_{C^{k,\alpha}_{\delta_0}} \equiv \sum_{i=0}^k (\sup_{\delta_0^n} |\nabla^i s|) + |\rho^{\delta_0+k+\alpha} \nabla^k s|_\alpha.
\]

Note $s(x)$ and $s(y)$ in (5.1) are compared by means of parallel transport. Let $i_{p_0}$ denote the injectivity radius at the point $p_0$. Define the following function

\[
\tilde{z} = \begin{cases} |z| & \text{for } |z| < \frac{1}{2} i_{p_0} \\ 1 & \text{on } M_0 - B(p_0, z_{p_0}) \end{cases}.
\]

We can define a similar norm $C^{k,\alpha}_{\delta_0}(M_0, E)$ for tensor bundles over $M_0$

\[
\|s\|_{C^{k,\alpha}_{\delta_0}} \equiv \sum_{i=0}^k (\sup z^{\delta_0+i} |\nabla^i s|) + |\tilde{z}^{\delta_0+k+\alpha} \nabla^k s|_\alpha.
\]

When $(M_0, \partial M_0, g_0, x)$ is AH, for another weight $\delta_\infty \in \mathbb{R}$, we define $C^{k,\alpha}_{\delta_0, \delta_\infty}(M_0, E)$ by the norm

\[
\|s\|_{C^{k,\alpha}_{\delta_0, \delta_\infty}} \equiv \|\tilde{x}^{-\delta_\infty} s\|_{C^{k,\alpha}_{\delta_0}},
\]

where $\tilde{x}$ is defined in (1.20). We also need to define function spaces $C^{k,\alpha}_{\delta_0, \delta_\infty; t}$ on the topological desingularization $M^t$. Consider the following norms

\[
\|s\|_{C^{k,\alpha}_{\delta_0, \delta_\infty; t}} \equiv t^{\frac{\delta_0}{2}} (\sup w^{\delta_0+i} |\nabla^i (\tilde{x}^{-\delta_\infty} s)|) + w^{\delta_0+k+\alpha} \nabla^k (\tilde{x}^{-\delta_\infty} s)|_\alpha,
\]

where $l = r - s$ for an $(r, s)$ tensor $E$ and $w$ is the weight function and the covariant derivatives are with respect to $g_t$ in last expression. It is easy to see that these norms are uniformly equivalent when $t$ is sufficiently small, so we will use them interchangeably, and omit the prime notation. The exponent of $t$ in (5.6) follows because

\[
t^{\frac{\delta_0}{2}} |s|_g = |s|_g,
\]

and since $\rho$ is comparable to $\tilde{z}^{\sqrt{t}}$ in the damage zone.

In the following, we fix $\delta_0 > 0$ and $\delta_\infty > 0$ sufficiently small.
5.2. Lyapunov-Schmidt reduction. Define the space of obstructions
\[ \mathcal{O}_t \subset C^{\alpha}_{2,\infty;\delta; t}(M^t) \]
to be the one-dimensional span of \( \chi(t) o \). Define the following operator
\[ \Phi_{g_t}(g) = \text{Ric}(g) - \Lambda g + \delta^* B_{g_t} g. \]
The essential ingredient to the gluing argument is the following implicit function theorem [Biq13].

**Lemma 5.1.** Let \( \Phi : E \to F \) be a differentiable map between Banach spaces and
\[ Q = \Phi - \Phi(0) - d_0 \Phi. \]
Assume there exist positive constants \( q, r_0, c \) such that
\[ \|Q(x) - Q(y)\| \leq q \|x - y\| \left( \|x\| + \|y\| \right) \quad x, y \in B(0, r_0) \]
\[ d_0 \Phi \text{ is an isomorphism and } \left\| (d_0 \Phi)^{-1} \right\| \leq c. \]
If \( r \leq \min(r_0, \frac{1}{2qc}) \) and \( \|\Phi(0)\| \leq \frac{r^2}{2c} \) then \( \Phi(x) = 0 \) has a unique solution in \( B(0, r) \).

Let \( v \in \mathbb{R} \) and consider the following metric on \( M^t \)
\[ g_{t,v} = g_t + tv \chi t o. \]
Notice that
\[ \| t \chi t o \|_{C^{2,\alpha}_{\delta;0,\infty; t}} = O(t^{\frac{\delta}{2}}). \]

**Proposition 5.2.** Assume \((M_0, g_0)\) is a non-degenerate Einstein orbifold with one singular point \( p_0 \) with orbifold group \( \Gamma_n \) (\( M_0 \) can either be compact or AHE). Then there exist \( \eta, \epsilon > 0 \) such that for any \( t < \eta \) and \( v \in \mathbb{R} \) with \( |v| \leq \epsilon \), there exist a unique metric \( \hat{g}_{t,v} \) solving the following equation
\[ \Phi(\hat{g}_{t,v}) \in \mathcal{O}_t, \]
where \( \hat{g}_{t,v} \) satisfies the following
\[ \| \hat{g}_{t,v} - g_t \|_{C^{2,\alpha}_{\delta;0,\infty; t}} \leq ct^{\frac{\delta}{2}}, \]
\[ \int_{\mathbb{P}^{n-1}} \left\langle \frac{1}{t} (\hat{g}_{t,v} - g_{t,v}), o \right\rangle_{\text{cal}} dV_{\text{cal}} = v. \]
Also, if we write \( \Phi(\hat{g}_{t,v}) = \lambda(t)(\chi_t o) \), then we have the following estimation
\[ \| \hat{g}_{t,v} - g_{t,v} \|_{C^{2,\alpha}_{\delta;0,\infty; t}} \leq ct^{1+\frac{\delta}{2}}, \]
\[ \lambda(t) = t\lambda + O(t^{2-\frac{\delta}{4}}) \]
as \( t \to 0 \).
Proof. The following proof is very similar to [Biq13, Proposition 9.1]. However, since scalings of various terms are different in higher dimensions, we give a fairly detailed outline here, omitting some computations. To begin, define the following operator

\[
\Phi : t^{-\frac{s_0}{2}} \tilde{C}^{2,\alpha}_{\delta_0,\delta_\infty,t} \oplus \mathbb{R} \rightarrow t^{-\frac{s_0}{2}} C^{\alpha}_{\delta_0+2,\delta_\infty,t}
\]

(5.21)

\[
\Phi(h, x) = \text{Ric}_{g_{t,v,h}} - \Lambda g_{t,v,h} + \delta^*_{g_{t,v,h}} B_{g_{t,v,h}} - x \chi o
\]

(5.22)

where \(g_{t,v,h} = g_{t,v} + h\) and \(\tilde{C}^{2,\alpha}_{\delta_0,\delta_\infty,t}\) is space of all \(C^{2,\alpha}_{\delta_0,\delta_\infty,t}\) tensors like \(h\) which satisfies:

\[
\int_{\mathbb{R}^n} \left\langle \frac{1}{t}(h - g_t), o \right\rangle \, dV_{cal} = v.
\]

(5.23)

For simplicity of notation, in the remainder of the proof we will omit the variable \(v\).

We want to apply Lemma 5.1 for \(E = t^{-\frac{s_0}{2}} \tilde{C}^{2,\alpha}_{\delta_0,\delta_\infty,t} \oplus \mathbb{R}\) and \(F = t^{-\frac{s_0}{2}} C^{\alpha}_{\delta_0+2,\delta_\infty,t}\). From Lemma 4.4 it follows that

\[
\|\Phi(0,0)\|_{C^{\alpha}_{\delta_0+2,\delta_\infty,t}} = O(t^{1+\frac{s_0}{4}}).
\]

(5.24)

The linearization of \(\Phi\) at \((0,0)\) is given by

\[
d(0,0)\Phi(h,x) = P_{t,v,h} - x \chi o + \frac{1}{2} (\text{Ric}_{g_{t,v}} \circ h + h \circ \text{Ric}_{g_{t,v}} - 2 \Lambda h).
\]

(5.25)

Also the quadratic terms are given by

\[
Q(h,x) = \Phi(h,x) - \Phi(0,0) - d_{(0,0)}\Phi(h,x).
\]

(5.26)

Using the expansion (4.14), and a similar expansion for the gauge term (see [GV16, Section 3]), it is not hard to see that there exist constants \(r_0\) and \(q\) so that (5.12) is satisfied. We need to show that \(d_{(0,0)}\Phi\) is an isomorphism with bounded inverse. Equivalently, we need to show that there exists a constant \(c > 0\) so that

\[
t^{-\frac{s_0}{2}} \|h\|_{C^{2,\alpha}_{\delta_0,\delta_\infty,t}} + |x| \leq c t^{-\frac{s_0}{2}} \|d_{(0,0)}\Phi\|_{C^{2,\alpha}_{\delta_0+2,\delta_\infty,t}}.
\]

(5.27)

Assume this is not true, then there exist a sequence \(t_i \to 0\) and \((h_i, x_i)\) such that

\[
t_i^{-\frac{s_0}{2}} \|h_i\|_{C^{2,\alpha}_{\delta_0,\delta_\infty,t_i}} + |x_i| = 1,
\]

(5.28)

but

\[
\|d_{(0,0)}\Phi(h_i, x_i)\|_{C^{2,\alpha}_{\delta_0+2,\delta_\infty,t_i}} \to 0.
\]

(5.29)

Let us first consider that case that \(x_i \to 0\). Then without loss of generality we may assume that

\[
t_i^{-\frac{s_0}{2}} \|h_i\|_{C^{\alpha}_{\delta_0,\delta_\infty,t_i}} \to 1,
\]

(5.30)

and

\[
t_i^{-\frac{s_0}{2}} \|d_{(0,0)}\Phi(h_i, x_i)\|_{C^{\alpha}_{\delta_0+2,\delta_\infty,t_i}} \to 0.
\]

(5.31)
To get a contradiction, we briefly outline the standard blow-up argument. Let \( p_i \in M_t \) be a sequence of points such that norm is attained. By passing to a subsequence we can assume that \( p_i \to p \). There are three cases for \( p \).

- If \( p \in M_0 \) then by standard elliptic theory we can obtain a limit \( h_\infty \) satisfying \( \|h_\infty\|_{C_{\delta_0}^{2,\alpha}} = 1 \) and \( P_{g_0} h_\infty = 0 \). But this is impossible because we assumed that \( M_0 \) is non-degenerate.

- If \( p \in X \) then we can obtain a limit \( h_\infty \) with \( \|h_\infty\|_{C_{\delta_0}^{2,\alpha}} = 1 \) and \( P_{g_{cal}} h_\infty = 0 \). From Theorem 2.2, \( h_\infty = c \cdot o \), for some constant \( c \). But then (5.18) implies that \( c = 0 \), a contradiction.

- If \( p \in \mathbb{C}^n / \Gamma_n \) then we can obtain a limit \( h_\infty \) with \( \|h_\infty\|_{C_{\delta_0}^{2,\alpha}} = 1 \) and \( \nabla^* \nabla h_\infty = 0 \), but this is impossible since the indicial roots are the integers and \( \delta_0 \) was chosen small but non-zero.

Next, we show that \( x_i \to 0 \). To prove this, we need the following estimates,

\[
\langle u, \chi_{t_o} \rangle_{L^2(g_{t_i})} \leq C t_i^{n-1} \frac{\delta_0}{2} \|u\|_{C_{\delta_0,\delta_\infty}^{2,\alpha}},
\]

\[
\langle P_{h}, \chi_{t_o} \rangle_{L^2(g_{t_i})} \leq C t_i^{n-1} \frac{\delta_0}{2} \|h\|_{C_{\delta_0,\delta_\infty}^{2,\alpha}},
\]

\[
\langle \chi_{t_o}, \chi_{t_o} \rangle_{L^2(g_{t_i})} = t_i^{n-2} \left( \langle o, o \rangle_{L^2(g_{t_i})} + O(t_i^{3n}) \right),
\]

for some constant \( C > 0 \). The proofs of these are straightforward and are omitted.

Using the above identities and the fact that \( t_i^{\frac{\delta_0}{2}} \|h_i\|_{C_{\delta_0,\delta_\infty}^{2,\alpha}} \leq 1 \) and estimating

\[
\langle d_{(0,0)} \Phi(h_{t_i}, x_{t_i}), \chi_{t_o} \rangle_{L^2(g_{t_i})},
\]

it follows that

\[
x_i t_i^{n-2} \left( c_1 + c_2 t_i^{\frac{3n}{2}} \right) = c_3 t_i^{n-1}
\]

for some constants \( c_1, c_2, \) and \( c_3 \) with both \( c_1 \) and \( c_2 \) non-zero. Consequently, we must have \( x_i \to 0 \) as \( i \to \infty \).

It remains to prove the expansion for \( \lambda(t) \). From Lemma 5.1 for each \( t \) sufficiently small, we obtain a solution \( (h_t, x_t) \) of \( \Phi(h_t, x_t) = 0 \). Defining \( \hat{g}_t = g_t + h_t \), we have \( \Phi_{g_t}(\hat{g}_t) = \lambda(t) (\chi_{t_o}) \). Notice that

\[
\langle \Phi_{g_t}(g_t + h_t), \chi_{t_o} \rangle_{L^2(g_{t_i})} = \lambda(t) \langle \chi_{t_o}, \chi_{t_o} \rangle_{L^2(g_{t_i})}.
\]

On the other hand

\[
\Phi_{g_t}(\hat{g}_t) = \Phi_{g_t}(g_t) + d_{g_t} \Phi_{g_t}(h_t) + Q_{g_t}(h_t),
\]

which implies that

\[
\langle \Phi_{g_t}(\hat{g}_t), \chi_{t_o} \rangle_{L^2(g_{t_i})} = \langle \Phi_{g_t}(g_t), \chi_{t_o} \rangle_{L^2(g_{t_i})} + \langle d_{g_t} \Phi_{g_t}(h_t), \chi_{t_o} \rangle_{L^2(g_{t_i})} + \langle Q_{g_t}(h_t), \chi_{t_o} \rangle_{L^2(g_{t_i})}.
\]
Using (5.32), Lemma 4.2, and the expansion (4.14), respectively, we can estimate each of the above terms by

\[ \langle \Phi_{g_t}(g_t), \chi_t o \rangle_{L^2(g_{cal})} = \lambda t^{n-1} \| o \|^2_{L^2(g_{cal})} + O(t^n), \]

(5.40)

\[ \langle d_{g_t} \Phi(h_t), \chi_t o \rangle_{L^2(g_t)} = O(t^{-\frac{d}{2}}), \]

(5.41)

\[ \langle Q_{g_t}(h_t), \chi_t o \rangle_{L^2(g_t)} = O(t^{n+1-\frac{d}{2}}), \]

(5.42)

as \( t \to 0 \), and we obtain

\[ \lambda(t) t^{n-2} \| o \|^2_{L^2(g_{cal})} = t^{n-2}(\lambda t \| o \|^2_{L^2(g_{cal})} + O(t^{2-\frac{d}{2}})), \]

(5.43)

which implies

\[ \lambda(t) = t \lambda + O(t^{2-\frac{d}{2}}), \]

(5.44)

as \( t \to 0 \).

Proof of Theorem 1.1. We first assume that \( g_0 \) is non-degenerate. Observe that the refined approximate metric is close to the naïve approximate metric in the \( C^{2,\alpha}_{\delta,0} \) norm, for \( t \) sufficiently small. From identical arguments as in [Biq13, Section 8], for any metric \( g_E \) satisfying (1.16), there exists a diffeomorphism \( \phi : M^t \to M^t \) so that \( B_{g_t}(g_t - \phi^* g_E) = 0 \), and \( \phi^* g_E \) remains close to \( g_t \) in the \( C^{2,\alpha}_{\delta,0} \) norm. The metric \( \phi^* g_E \) would then give a zero of the mapping \( \Phi_t \), but then (5.20) would then imply that \( \lambda = 0 \), a contradiction.

If \( g_0 \) is not non-degenerate, then there is a finite dimensional space of infinitesimal Einstein deformations on \((M_0, g_0)\), of dimension \( k \geq 0 \). The proof of Proposition 5.2 can be modified to yield a mapping

\[ \psi_t : B_1 \times (-\epsilon, \epsilon) \to B_1 \times (-\epsilon, \epsilon), \]

(5.45)

where \( B_1 \) is a small ball in \( \mathbb{R}^k \), whose zero set is in one-to-one correspondence with the zeroes of \( \Phi_{g_t} \) (see also the argument in [GV16, Section 11]). An almost identical argument as in the proof of Proposition 5.2 shows that the mapping \( \psi \) admits an expansion

\[ \psi(a_1, \ldots, a_k, v) = (\psi_1, \ldots, \psi_k, t \lambda + O(t^{2-\frac{d}{2}})), \]

(5.46)

as \( t \to 0 \), so the existence of an Einstein metric in the stated neighborhood still implies that \( \lambda = 0 \).

6. Analysis on asymptotically hyperbolic manifolds

The goal of this section is to study the asymptotic behavior of \( L^2 \)-solutions of some elliptic operators on an AHE manifold \((M^m, \partial M, g, x)\). We consider the following operators:

1. \( P_0 = \Delta_H - 2Ric \), acting on functions,

2. \( P_1 = B\delta^* = \frac{1}{2}(\nabla^* \nabla - Ric) = \frac{1}{2}(\Delta_H - 2Ric) \), acting on 1-forms,
(3) \( P_2 = \Delta_H - 2\text{Ric} \), acting on 2-forms,

(4) \( P = \frac{1}{2} \nabla^* \nabla - \ddot{R} \), acting on traceless symmetric 2-tensors.

We also need the following decompositions, noting that we will only consider \( C^\infty \) sections of bundles, so notations of regularity will be suppressed.

For \( \Omega^1(M) \mid_{\partial M} \), we have the decomposition

\[
\Omega^1(M) \mid_{\partial M} = \Omega^1_t(M) \mid_{\partial M} \oplus \Omega^1_n(M) \mid_{\partial M},
\]

where \( \Omega^1_t(M) \) is the space of one-forms on \( M \) which vanish on normal vectors (with respect to \( g \)) when restricted to the boundary and \( \Omega^1_n(M) \) is the space one-forms on \( M \) which vanish on tangential vectors when restricted to the boundary.

Next, for \( \Omega^2(M) \mid_{\partial M} \) we have the following,

\[
\Omega^2(M) \mid_{\partial M} = \Omega^2_t(M) \mid_{\partial M} \oplus (\Omega^1_n(M) \mid_{\partial M} \wedge \Omega^1_t(M) \mid_{\partial M}),
\]

where \( \Omega^2_t(M) \) is space of two-forms on \( M \) which satisfy \( \partial_x \omega = 0 \) on the boundary, where \( \partial_x \) is dual of \( dx \) with respect to \( \bar{g} \).

For \( S^2_0(M) \mid_{\partial M} \) we have the following decomposition

\[
S^2_0(M) \mid_{\partial M} = V_0 \oplus V_1 \oplus V_2,
\]

where

\[
V_0 \equiv \{ h \in S^2_0(M) \mid_{\partial M} \mid \text{tr}_g h = 0, \partial_x \omega h = 0 \},
\]

\[
V_1 \equiv \{ h \in S^2_0(M) \mid_{\partial M} \mid h = v \odot dx, v \in T^* \partial M \},
\]

\[
V_2 \equiv \{ h \in S^2_0(M) \mid_{\partial M} \mid h \text{ is a multiple of } ((m - 1)dx \odot dx - \bar{g}) \},
\]

where \( \odot \) denotes the symmetric product. Finally, let \( V_0, V_1, V_2 \) be those traceless symmetric tensors on \( M \) such that they lie in \( V_0, V_1, V_2 \) when restricted to the boundary, respectively.

An important property of these operators that they are uniformly degenerate at the boundary, and there are a finite set of numbers, called \( \text{indicial roots} \) which characterize invertibility of the associated indicial operators \cite{Lee06, Maz91a}. The following proposition gives the indicial roots of these operators, and will be used throughout the remainder of the paper. For simplicity, we state the result for solutions defined on all of \( M \), but note that the same conclusions hold only assuming that the solutions are defined in a neighborhood of the boundary of \( M \).

**Proposition 6.1.** Let \( (M^n, \partial M, g, x) \) be an AH manifold.

- The indicial roots associated with \( P_0 \) are given by

\[
\delta^+_0, P_0, \delta^-_0, P_0 = \frac{m - 1 \pm \sqrt{(m - 1)^2 + 8(m - 1)}}{2}.
\]
If $P_0(f) = 0$ for some function $f \in L^2(M)$ then $f = f_\infty x^{\delta_0, P_0} + o(x^{\delta_0, P_0})$ as $x \to 0$, where $f_\infty : \overline{M} \to \mathbb{R}$, and if $f_\infty = 0$ on $\partial M$, then $f \equiv 0$.

- The indicial roots associated with $P_1$ are given by

$$
\delta_{0,P_1}^+, \delta_{0,P_1}^- = m, -1
$$

$$
\delta_{1,P_1}^+, \delta_{1,P_1}^- = \frac{m - 1 \pm \sqrt{(m - 1)^2 + 8(m - 1)}}{2},
$$

where the roots $\delta_{0,P_1}^+$ correspond to sections of $\Omega_1^1 \overline{M}$ and $\delta_{1,P_1}^+$ corresponds to sections $\Omega_1^1 \overline{M}$. Moreover, if $\alpha \in \Omega_1^1(M)$ is in $L^2$ and solves $P_1(\alpha) = 0$, then we have the following expansion as $x \to 0$

$$
\alpha = \alpha_t x^{\delta_{0,P_1}^+, P_1} + o(x^{\delta_{0,P_1}^+, P_1}),
$$

where $x\alpha_t \in \Omega_1^1 \overline{M}$ and if $x\alpha_t \equiv 0$ on $\partial M$ then

$$
\alpha = \alpha_n x^{\delta_{1,P_1}^+, P_1} + o(x^{\delta_{1,P_1}^+, P_1}),
$$

where $x\alpha_n \in \Omega_1^1 \overline{M}$ and if $x\alpha_n \equiv 0$ on $\partial M$ then $\alpha \equiv 0$.

- The indicial roots associated with $P_2$ are given by

$$
\delta_{0,P_2}^+, \delta_{0,P_2}^- = \frac{m - 1 \pm \sqrt{(m - 1)^2 + 16}}{2}
$$

$$
\delta_{1,P_2}^+, \delta_{1,P_2}^- = m, -1,
$$

where $\delta_{0,P_2}^+$ corresponds to sections of $\Omega_2^1 \overline{M}$ and $\delta_{1,P_2}^+$ corresponds to sections of $(\Omega_1^1 \overline{M} \land \Omega_2^1 \overline{M})$. Moreover, if $\omega \in \Omega_2^1(M)$ is in $L^2$ and satisfies $P_2(\omega) = 0$, then we have the following expansion as $x \to 0$

$$
\omega = \omega_t x^{\delta_{0,P_2}^+, P_2} + o(x^{\delta_{0,P_2}^+, P_2}),
$$

where $x^2 \omega_t \in \Omega_2^1 \overline{M}$ and if $x^2 \omega_t \equiv 0$ on $\partial M$ then

$$
\omega = \omega_n x^{\delta_{1,P_2}^+, P_2} + o(x^{\delta_{1,P_2}^+, P_2}),
$$

where $x^2 \omega_n \in (\Omega_1^1 \overline{M} \land \Omega_2^1 \overline{M})$ and if $x^2 \omega_n \equiv 0$ on $\partial M$ then $\omega \equiv 0$.

- The indicial roots associated with $P$ are given by

$$
\delta_{0,P}^+, \delta_{0,P}^- = m - 1, 0
$$

$$
\delta_{1,P}^+, \delta_{1,P}^- = m, -1
$$

$$
\delta_{2,P}^+, \delta_{2,P}^- = \frac{m - 1 \pm \sqrt{(m - 1)^2 + 8(m - 1)}}{2},
$$

where $\delta_{0,P}^+$ corresponds to sections of $\Omega_2^1 \overline{M}$ and $\delta_{1,P}^+$ corresponds to sections of $(\Omega_1^1 \overline{M} \land \Omega_2^1 \overline{M})$. Moreover, if $\omega \in \Omega_2^1(M)$ is in $L^2$ and solves $P(\omega) = 0$, then we have the following expansion as $x \to 0$

$$
\omega = \omega_t x^{\delta_{0,P}^+, P} + o(x^{\delta_{0,P}^+, P}),
$$

where $x^2 \omega_t \in \Omega_2^1 \overline{M}$ and if $x^2 \omega_t \equiv 0$ on $\partial M$ then

$$
\omega = \omega_n x^{\delta_{1,P}^+, P} + o(x^{\delta_{1,P}^+, P}),
$$

where $x^2 \omega_n \in (\Omega_1^1 \overline{M} \land \Omega_2^1 \overline{M})$ and if $x^2 \omega_n \equiv 0$ on $\partial M$ then $\omega \equiv 0$. 


where \( \delta^+_0, \delta^+_1, \delta^+_2 \) corresponds to sections of \( V_0, V_1, V_2 \), respectively. Moreover, if \( h \in S^2_0(T^*M) \) is in \( L^2 \) and satisfies \( Ph = 0 \), then we have the following expansion as \( x \to 0 \)

\[
(6.13) \quad h = \sigma_0 x^{\delta^+_0} + o(x^{\delta^+_0})
\]

where \( x^2 \sigma_0 \in V_0 \) and if \( x^2 \sigma_0 \equiv 0 \) on \( \partial M \) then

\[
(6.14) \quad h = \sigma_1 x^{\delta^+_1} + o(x^{\delta^+_1}),
\]

where \( x^2 \sigma_1 \in V_1 \) and if \( x^2 \sigma_1 \equiv 0 \) on \( \partial M \) then

\[
(6.15) \quad h = \sigma_2 x^{\delta^+_2} + o(x^{\delta^+_2})
\]

where \( x^2 \sigma_2 \in V_2 \) and if \( x^2 \sigma_2 \equiv 0 \) on \( \partial M \) then \( h \equiv 0 \).

Proof. The proof for existence of an expansion around infinity for uniformly degenerate elliptic operators can be found in [Maz88, Maz91a]. The vanishing results follow from a unique continuation principle due to Mazzeo [Maz91b, Maz91a]. Since we were not able to find calculations of the indicial roots for all of these operators in one place in the literature, we will give an outline of the calculations below (see [Maz88] for calculations involving the Hodge Laplacian, and [GL91] for calculation of the indicial roots in case of \( P \)). We need the following two lemmas.

Lemma 6.2. Let \( (M, \partial M, g, x) \) be an asymptotically hyperbolic manifold. Let \( \hat{g} \) be a representative of the boundary conformal class. Then there exists a unique defining function \( x \) (defined in a collar neighborhood of \( \partial M \)) such that

\[
(6.16) \quad g = \sinh^{-2}(x)(dx^2 + g_x),
\]

where \( g_x \) is an \( x \)-dependent metric on a slice, and \( g_0 = \hat{g} \).

Proof. This is proved, for example, in [AD98]  \( \square \)

The above lemma reduces the calculation of indicial roots to calculations on the hyperbolic ball. To this end, consider the hyperbolic metric in normal coordinates

\[
(6.17) \quad g_H = dr^2 + \sinh^2(r)g_{S^{m-1}},
\]

where \( g_{S^{m-1}} \) is the standard metric on \( S^{m-1} \).

Lemma 6.3. Let \( f(r, \theta) \) be a function in polar coordinates, then

\[
(6.18) \quad \Delta f(r, \theta) = \frac{1}{\sinh^2(r)} \Delta_{S^{m-1}} f(r, \theta) + (m-1) \frac{\cosh(r)}{\sinh(r)} f'(r, \theta) + f''(r, \theta).
\]

Proof. Let \( p \in \mathbb{R}^m \), and let \( \nabla \) denotes the covariant derivative with respect to the metric \( g_H \) and let \( \nabla \) denote the covariant derivative on the \( (m - 1) \)-sphere with respect to the metric \( \sinh(r)^2 g_{S^{m-1}} \). We construct a local orthonormal frame \( e_0 = \partial_r, e_1, ..., e_{m-1} \) as follows. Let \( \partial_r = e_0 \) denote the unit radial vector field. Choose \( e_1, ..., e_m \) such that

\[
(6.19) \quad \nabla_{e_i} e_j(p) = 0
\]
and extend $e_1, ..., e_m$ by parallel transport along the radial direction. We obtain a local orthonormal frame around $p$ with following properties

\begin{align}
\nabla_{e_i} e_j(p) &= 0 \quad i, j \in \{1, ..., m - 1\} \\
\nabla_{e_0} e_i &= 0 \quad i \in \{0, ..., m - 1\}.
\end{align}

It follows that

\begin{equation}
\Delta f = \nabla^2 f(e_i, e_i) + \nabla^2 f(\partial_r, \partial_r).
\end{equation}

Also, notice that

\begin{equation}
\nabla^2 f(e_i, e_i) = \nabla_{e_i} \nabla_{e_i} f - \nabla_{e_i} e_i f;
\end{equation}

\begin{equation}
\nabla_{\partial_r} \partial_r = 0 \implies \nabla^2 f(\partial_r, \partial_r) = \nabla_{\partial_r} \nabla_{\partial_r} f,
\end{equation}

and

\begin{equation}
\nabla_{e_i} e_i(p) = 0 \implies \nabla_{e_i} e_i = \langle \nabla_{e_i} e_i, \partial_r \rangle \partial_r.
\end{equation}

Next, recall that

\begin{equation}
\Gamma^k_{ij} = \frac{1}{2} g^{kl} (\partial_i g_{jl} + \partial_j g_{il} - \partial_l g_{ij}).
\end{equation}

It follows that

\begin{equation}
\langle \nabla_{e_i} e_i, \partial_r \rangle = -\frac{\cosh(r)}{\sinh(r)}.
\end{equation}

Putting everything together we obtain (6.18). □

Next, we calculate the indicial roots for the above operators.

- Indicial roots for $P_0$.

From formula (6.18), we see that the terms with spatial derivatives are strictly lower order terms. Therefore, without loss of generality, we may assume that $f(r)$ is a radial function. Applying (6.18), we obtain

\begin{align}
\nabla_{e_i} e_i &= -(m - 1) \coth(r) f' - f'' + 2(m - 1) \\
\nabla_{\partial_r} \partial_r &= -(m - 1) f' - f'' + 2(m - 1) + O(1 - \coth(r)) f'.
\end{align}

Letting $f(r) = e^{-\delta r}$ we obtain the quadratic formula

\begin{equation}
(m - 1) \delta - \delta^2 + 2(m - 1) = 0.
\end{equation}

- Indicial roots for $P_1 = B\delta^*$.

Let $p \in \mathbb{R}^m$, and $e_0 = \partial_r, e_1, ..., e_{m-1}$ be a local orthonormal frame as in (6.20) and (6.21), and let $e^0, ..., e^{m-1}$ denote the dual frame. As in the previous case, terms in the operator with spatial derivatives will contribute lower order terms, so without loss
of generality, we may assume that $\alpha = \alpha_0 e^0 + \alpha_i e^i$ is a 1-form with radial coefficients. It follows that

$$
\begin{align*}
\nabla_i e^0 &= \coth(r) e^i, \\
\nabla^* \nabla \alpha &= -\alpha'' e^0 - \alpha' e^i - (m - 1) \coth(r) (\alpha'_0 e^0 + \alpha'_i e^i) + \mathcal{D} \alpha
\end{align*}
$$

where $\mathcal{D} = -\sum_{i=1}^{m-1} \nabla_i \nabla_i \alpha$ satisfies the following

$$
\begin{align*}
\mathcal{D}(e^0) &= (m - 1)e^0 + O(1 - \coth^2(r)) \\
\mathcal{D}(e^j) &= e^j + O(1 - \coth^2(r)) \quad (j \neq 0),
\end{align*}
$$

as $r \to \infty$. Switching to the disc model of hyperbolic space, recall the decomposition

$$
\Omega^{1,1} \mathbb{D}^m \big|_{\partial \mathbb{D}^m} = \Omega^{1,1}_r \mathbb{D}^m \big|_{\partial \mathbb{D}^m} \oplus \Omega^{1,1}_n \mathbb{D}^m \big|_{\partial \mathbb{D}^m},
$$

and note that

$$
(6.33) \quad Ric(\alpha) = -(m - 1)\alpha.
$$

We search for $s = f E_\mu$ where $E_\mu$ is an eigenform for $\mathcal{D} - Ric$. We obtain the following ODEs corresponding to $2(m - 1)$ and $m$ respectively

$$
(6.34) \quad \begin{align*}
-f'' - (m - 1)f' + 2(m - 1)f &= 0, \\
f'' - (m - 1)f' + mf &= 0,
\end{align*}
$$

and letting $f(r) = e^{-\delta r}$, we obtain the following quadratic formulas

$$
(6.35) \quad \begin{align*}
-\delta^2 + (m - 1)\delta + 2(m - 1) &= 0, \\
-\delta^2 + (m - 1)\delta + m &= 0.
\end{align*}
$$

- Indicial roots for $P_2$.

By the Weitzenböck formula on two-forms and since hyperbolic space has constant sectional curvature equal to $-1$, we have

$$
(6.36) \quad (\Delta_H \omega)_{ij} = (\nabla^* \nabla \omega)_{ij} + 2(2 - m)\omega_{ij},
$$

and it follows that $P_2(\omega) = (\nabla^* \nabla \omega) + 2\omega$. Let $p \in \mathbb{R}^m$. Let $e_0 = \partial_r, e_1, ..., e_{m-1}$ be a local orthonormal frame as in (6.20) and (6.21) and let $e^0, ..., e^{m-1}$ denote the dual frame. It follows that

$$
(6.37) \quad \begin{align*}
\nabla_i e^0 &= \coth(r)e^i, \\
\nabla_j e^i &= -\coth(r)\delta_{ij} e^0.
\end{align*}
$$

Again, without loss of generality, we may assume that $\omega = \omega_0 e^0 \wedge e^i + \omega_i e^i \wedge e^j$ is a 2-form with radial coefficients. We obtain

$$
(6.38) \quad \begin{align*}
\nabla^* \nabla \omega &= -\omega''_{0i} e^0 \wedge e^i - \omega''_{ij} e^i \wedge e^j - (m - 1) \coth(r) (\omega'_0 e^0 \wedge e^i + \omega'_i e^i \wedge e^j) \\
&\quad + \mathcal{D} \omega
\end{align*}
$$
where \( D\omega = -\sum_{i=1}^{m-1} \nabla_i \nabla_i \omega \) satisfies the following

\[
\begin{align*}
(6.39) \quad D(e^i \wedge e^j) &= 2e^i \wedge e^j + O(1 - \coth^2(r)) \\
(6.40) \quad D(e^i \wedge e^0) &= (m - 2)(e^i \wedge e^0) + O(1 - \coth^2(r)),
\end{align*}
\]

as \( r \to \infty \). Switching to the disc model, we have the decomposition

\[
\begin{align*}
\Omega^2 \mathbb{D}^m |_{\partial \mathbb{D}^m} &= \Omega^2 \mathbb{D}^m |_{\partial \mathbb{D}^m} \oplus (\Omega^1_n \mathbb{D}^m |_{\partial \mathbb{D}^m} \wedge \Omega^1_i \mathbb{D}^m |_{\partial \mathbb{D}^m}).
\end{align*}
\]

We obtain the following ODE for the roots on two-forms

\[
\begin{align*}
(6.42) \quad -f'' - (m - 1)f' + 4f &= 0, \\
(6.43) \quad -f'' - (m - 1)f' + mf &= 0,
\end{align*}
\]

and letting \( f(r) = e^{-\delta r} \) yields the following quadratic formulas

\[
\begin{align*}
(6.44) \quad -\delta^2 + (m - 1)\delta + 4 &= 0, \\
(6.45) \quad -\delta^2 + (m - 1)\delta + m &= 0.
\end{align*}
\]

- Indicial roots for \( P = \frac{1}{2} \nabla^* \nabla - \tilde{R} \).

Let \( p \in \mathbb{R}^m \), and let \( e_0 = \partial_r, e_1, ..., e_{m-1} \) be a local orthonormal frame as in (6.20) and (6.21), and let \( e^0, ..., e^{m-1} \) denote the dual frame. Again, without loss of generality, we may assume that \( h \) is a symmetric two-tensor with radial coefficients. We can write \( h \) in above frame as follows

\[
(6.46) \quad h = h_{00} e^0 \otimes e^0 + h_{0i} e^0 \otimes e^i + h_{ij} e^i \otimes e^j,
\]

and we obtain the following expression for the Laplacian

\[
\begin{align*}
\nabla^* \nabla h &= -h''_{00} e^0 \otimes e^0 - h''_{0i} e^0 \otimes e^i - h''_{ij} e^i \otimes e^j \\
&\quad - (m - 1) \coth(r) (h'_{00} e^0 \otimes e^0 + h'_{0i} e^0 \otimes e^i + h'_{ij} e^i \otimes e^j) + Dh,
\end{align*}
\]

where \( Dh = -\sum_{i=1}^{m-1} \nabla_i \nabla_i h \). Next, we compute eigenvalues of \( D \). First, note that

\[
\begin{align*}
\nabla_i (e^0 \otimes e^0) &= \coth(r) (e^0 \otimes e^i + e^i \otimes e^0) \\
\nabla_i \nabla_i e^0 \otimes e^0 &= 2 \coth^2(r) e^i \otimes e^i - 2 \coth^2(r) (e^0 \otimes e^0),
\end{align*}
\]

and

\[
\begin{align*}
D(e^0 \otimes e^0) &= \sum_{i=1}^{m-1} -2(e^i \otimes e^i - e^0 \otimes e^0) + O(1 - \coth^2(r)) \\
D(e^j \otimes e^j) &= 2(e^j \otimes e^j - e^0 \otimes e^0) + O(1 - \coth^2(r)) \\
D(e^0 \otimes e^i + e^i \otimes e^0) &= (m + 2)(e^0 \otimes e^i + e^i \otimes e^0) + O(1 - \coth^2(r)) \\
D(e^j \otimes e^k + e^k \otimes e^j) &= 2(e^j \otimes e^k + e^k \otimes e^j) + O(1 - \coth^2(r)),
\end{align*}
\]
as \( r \to \infty \). Next, we have the following for \( \hat{R} \)

\[
\begin{align*}
\hat{R}(e^p \otimes e^q + e^q \otimes e^p) &= (e^p \otimes e^q + e^q \otimes e^p) \text{ when } p \neq q \\
\hat{R}(e^p \otimes e^p) &= -e^0 \otimes e^0 - \ldots - e^p \otimes e^p - \ldots - e^{m-1} \otimes e^{m-1}.
\end{align*}
\]

We obtain

\[
\begin{align*}
\mathcal{C}(e^i \otimes e^j + e^j \otimes e^i) &= 0 + O(1 - \coth^2(r)) \quad (j > 0, k > 0) \\
\mathcal{C}(e^j \otimes e^j - e^k \otimes e^k) &= 0 + O(1 - \coth^2(r)) \\
\mathcal{C}(e^0 \otimes e^k + e^k \otimes e^0) &= \frac{m}{2} (e^0 \otimes e^k + e^k \otimes e^0) + O(1 - \coth^2(r)) \\
\mathcal{C}(e^j \otimes e^j) &= e^1 \otimes e^1 + \ldots + e^{m-1} \otimes e^{m-1} + O(1 - \coth^2(r)) \\
\mathcal{C}(e^0 \otimes e^0) &= (m - 1) e^0 \otimes e^0 + O(1 - \coth^2(r)) \\
\mathcal{C}(g) &= (m - 1) g + O(1 - \coth^2(r)) \\
\mathcal{C}(me^0 \otimes e^0 - g) &= (m - 1)(me^0 \otimes e^0 - g) + O(1 - \coth^2(r))
\end{align*}
\]

where \( \mathcal{C} = \frac{1}{2} \mathcal{D} - \hat{R} \). In the disc model, recall the decomposition of the space of traceless symmetric two tensors at the boundary

\[
\begin{align*}
V_0 &\equiv \{ h \in S_0^2(\mathbb{D}^m) \mid _{\partial \mathbb{D}^m} tr \hat{h} = 0, _{\partial \mathbb{D}^m} \hat{h} = 0 \}, \\
V_1 &\equiv \{ h \in S_0^2(\mathbb{D}^m) \mid _{\partial \mathbb{D}^m} h = v \otimes dx, v \in T^* S^{m-1} \}, \\
V_2 &\equiv \{ h \in S_0^2(\mathbb{D}^m) \mid _{\partial \mathbb{D}^m} h \text{ is a multiple of } ((m - 1)dx \otimes dx - \bar{g}) \},
\end{align*}
\]

where \( _{\partial \mathbb{D}^m} \) is the dual of \( dx \) with respect to \( \bar{g} \) and \( V_0, V_1, V_2 \) are eigenspaces of \( \mathcal{C} \) corresponding to \( 0, \frac{m}{2}, m - 1 \), respectively. Next, we search for \( s = fE_\mu \) where \( E_\mu \) is an eigentensor with eigenvalue \( \mu \) of \( \mathcal{C} \), which gives us three types indicial roots around infinity corresponding to the following ODEs for elements in \( V_0, V_1 \) and \( V_2 \) respectively

\[
\begin{align*}
-\frac{1}{2} f'' - \frac{(m - 1)}{2} f' &= 0 \\
-\frac{1}{2} f'' - \frac{(m - 1)}{2} f' + \frac{m}{2} f &= 0 \\
-\frac{1}{2} f'' - \frac{(m - 1)}{2} f' + (m - 1) f &= 0.
\end{align*}
\]

Letting \( f(r) = e^{-\delta r} \) we obtain the quadratic formulas

\[
\begin{align*}
- \delta^2 + (m - 1)\delta &= 0 \\
- \delta^2 + (m - 1)\delta + m &= 0 \\
- \delta^2 + (m - 1)\delta + 2(m - 1) &= 0.
\end{align*}
\]
The reader will notice that there are several coincidences between the indicial roots in Proposition 6.1. This is not an accident, due to the following relations between the various operators (which also contains some formulas which will be useful later).

**Lemma 6.4.** Let \((M^n, g)\) be an Einstein manifold such that \(\text{Ric}_g = \Lambda g\) then we have the following relations for \(f : M \to \mathbb{R}, \omega \in \Omega^1(M),\) and \(h \in \Gamma(S^2(T^*M))\).

1. If \(P_0(f) = 0\) then \(P_1(df) = 0\).
2. \(P(f \cdot g) = \frac{1}{2} P_0(f) \cdot g\).
3. If \(P_1(\omega) = 0\) then \(P_0(\delta \omega) = 0\) and \(P_2(d\omega) = 0\).
4. If \(P h = 0\) then \(P_1(Bh) = 0\) and \(P_0(\delta Bh) = 0\).
5. \(P \delta^* = \delta^* B \delta^*\).

**Proof.** The Bochner formula on 1-forms implies that

\[
P_1 = \frac{1}{2} (\nabla^* \nabla - \text{Ric}),
\]

\[
P_1 = \frac{1}{2} (\Delta_H - 2\text{Ric}).
\]

Since \(\delta\) and \(d\) commute with the Hodge Laplacian, the first and third relations follow. The second relation is obvious, since \(g\) is parallel.

Next, we recall that

\[
B(d_g \text{Ric}) = \Lambda B
\]

\[
d_g \text{Ric} (\delta^*) = \Lambda \delta^*,
\]

and

\[
d_g \text{Ric}(h) = P(h) - \delta^* B(h) + \Lambda h.
\]

Applying \(B\) to both sides of (6.59), we have

\[
B \delta^* B(h) = 0,
\]

which implies \(P_1(B(h)) = 0\). This proves the fourth relation. For the last relation we have

\[
P \delta^* = d_g \text{Ric}(\delta^*) + \delta^* B \delta^* - \Lambda \delta^* = \delta^* B \delta^*.
\]

\[\square\]

7. From germs to boundary tensors

To complete the proof of Theorem 1.2 we will need to find global solutions of \(Ph = 0\) on \(M\) with prescribed leading order quadratic asymptotics at the point \(p_0\), and such that \(h\) is bounded (in the interior metric) on approach to conformal infinity. The argument is based on the four-dimensional case considered by Biquard in [Biq13], but there are some key differences. We have the following definition.

**Definition 7.1.** Let \((M^m, \partial M, g, x)\) be an AH orbifold with one singular point \(p_0\). Let \(F_k\) denote the space of harmonic polynomials of degree \(k\), in normal coordinates at \(p_0\). We refer to elements of \(F_k \otimes S^2_0(T^*_{p_0}M)\) as \(k\)-germs at \(p_0\).
We next need a few basic lemmas.

**Lemma 7.2.** Let \((M^m, \partial M, x, g)\) be an AH manifold and let \(h\) be a symmetric two tensor defined in neighborhood of the boundary. Then there exist a one-form \(\alpha\) defined in neighborhood of the boundary such that \(s = h + \delta^*\alpha\) is in radial gauge (i.e. \(\partial_x s = 0\)). Furthermore, if \(|h|_g = o(x^{m-1})\), then \(\alpha\) can be chosen to satisfy \(|\alpha|_g = o(x^{m-1})\) as \(x \to 0\).

**Proof.** Using the boundary defining function \(x\), we identify a neighborhood of the boundary with \([0, \epsilon) \times \partial M\). Fix any point \((\epsilon_0, p) \in [0, \epsilon) \times \partial M\). We specify a local orthonormal frame on a slice as follows: choose \(\{\epsilon_0, \ldots, \epsilon_{m-1}\}\) at \((\epsilon_0, q)\), for \(q\) near \(p\) so that \(\epsilon_0\) is normal to the slice \(\{\epsilon_0\} \times \partial M\), and \(\{\epsilon_1, \ldots, \epsilon_{m-1}\}\) is a local orthonormal frame on the tangent space of the slice \(\{\epsilon_0\} \times \partial M\). Extend this locally by parallel transport along integral curves of \(\epsilon_0\) to an orthonormal frame in a neighborhood of the chosen point. Let \(\{e^0, \ldots, e^{m-1}\}\) denote the dual frame.

This frame has the following property

\[
\nabla e_i e^i = 0
\]

for \(i \in \{0, \ldots, m - 1\}\). Next, we write \(h\) in the above frame as follows

\[
h = \sum_{i=1}^{m-1} h_{\alpha i}(e^0 \otimes e^i + e^i \otimes e^0) + \sum_{i,j=1}^{m-1} h_{ij}(e^i \otimes e^j + e^j \otimes e^i) + h_{00} e^0 \otimes e^0,
\]

where \(h_{00}, h_{\alpha i}, h_{ij}\) are functions defined near the chosen point. Next, for a 1-form of the form \(\alpha = \sum_{i=0}^{m-1} \alpha_0 e^i\), we compute

\[
\nabla (\alpha_0 e^i) = d\alpha_0 \otimes e^i + \alpha_0 \omega_{ij} \otimes e^j
\]

and

\[
\delta^* \alpha = \frac{1}{2} \left( d\alpha_0 \otimes e^i + \alpha_0 \omega_{ij} \otimes e^j \right).
\]

We obtain

\[
e_0 \lrcorner (\delta^* \alpha) = \frac{1}{2} \left( d\alpha_0 (e_0) e^i + d\alpha_0 (e_0) e^0 + \alpha_0 \omega_{ij} (e_0) e^j + \alpha_0 \omega_{i0} \right)
\]

\[
= \frac{1}{2} \left( (d\alpha_0 (e_0)) e^i + d\alpha_0 (e_0) + \alpha_0 \omega_{i0} \right),
\]

where \(\omega_{ij}\) are the connection 1-forms satisfying \(\nabla e^j = \omega_{ji} \otimes e^i\), and we also used (7.1). We also have

\[
e_0 \lrcorner h = \sum_{i=1}^{m-1} h_{\alpha i} e^i + h_{00} e^0.
\]

Choosing \(\alpha_0\) to be a radial function, we first solve the equation \(d\alpha_0 (e_0) = h_{\alpha i}\) for functions \(\alpha_0\) when \(i \geq 1\), which we can solve by integrating along integral curves of \(e_0\). Then we solve the equation \(2\dot{\alpha}_0 + \alpha_0 \omega_{i0} = h_{00}\) for \(\alpha_0\). This defines \(\alpha\) only locally but we can extend it to a global form in neighborhood of the boundary by using a partition of unity argument. The statement about decay rates follows easily. \(\square\)
We also need the following analytic continuation result for analytic Riemannian manifolds.

**Lemma 7.3.** Let \((M^m, \partial M, g, x)\) be AHE and let \(P(s) = 0\) in some open set \(V \subset M\) such that \(s = \delta^* \alpha\) for some one-form \(\alpha\) defined in some open ball \(U \subset V\) then \(\alpha\) can be analytically continued to a vector field \(\tilde{\alpha}\) defined on any simply connected domain \(U \subset U_1 \subset V\) such that \(s = \delta^* \tilde{\alpha}\). Furthermore, if \(V\) contains a punctured ball around \(p_0\), and \(s = O(|z|^{\gamma})\) for \(\gamma \neq -1\) where \(|z|\) is distance to the point \(p_0\), then \(\tilde{\alpha} = O(|z|^{\gamma+1})\) as \(|z| \to 0\).

**Proof.** Since Einstein metrics are real analytic in harmonic coordinates one can use the infinitesimal version of analytic continuation technique (see [KN96 Corollary 6.4]) to extend \(\alpha\) to a larger domain. We refer to [And08 Lemma 2.6] for a complete proof. Finally, an examination of the proof shows that the solution is obtained by integration, and the last statement follows from this (the details are omitted). \(\square\)

Another crucial result we will need is the following unique continuation theorem due to Biquard.

**Theorem 7.4 (Biquard [Biq08]).** Let \((M^m, \partial M, g, x)\) be AHE. If \(h \in \Gamma(S^2(T^*U))\), where \(U\) is a neighborhood of \(\partial M\), is a solution of \(d_g \text{Ric}(h) = \Lambda h\), and \(h\) is in radial gauge (i.e. \(\partial_x h = 0\)), such that \(|h|_g = o(x^{m-1})\) as \(x \to 0\), then \(h \equiv 0\).

We next define a mapping on 0-germs

\[
S_0^\infty : S_0^2(T^*_p M_0) \to V_0,
\]

where \(V_0\) was defined in (6.4). Given a traceless symmetric matrix \(a_{ij}\), then there is a unique solution \(\tilde{h} \in \Gamma(S_0^2(T^*_p M_0))\) of \(P(\tilde{h}) = 0\) satisfying

\[
\tilde{h} = \frac{a_{ij}}{|z|^{m-2}}dz^i \otimes dz^j + O(|z|^{3-m})
\]

as \(|z| \to 0\), and such that \(\tilde{h} \in L^2(M_0 \setminus B(p_0, \epsilon))\). The solution is unique since \(M_0\) is assumed to be non-degenerate. From Proposition 6.1 \(\tilde{h}\) admits an expansion

\[
\tilde{h} = \sigma_0 x^{m-1} + o(x^{m-1}),
\]

with \(x^2 \sigma_0 \in \nabla_0\), as \(x \to 0\). Since \(\tilde{h}\) is unique, the mapping

\[
S_0^\infty (a_{ij} dz^i \otimes dz^j) \equiv x^2 \sigma_0|_{\partial M_0}
\]

is well-defined. Next, let \(E_0 = \text{Im}(S_0^\infty)\). In a similar fashion we can define a mapping on 1-germs

\[
S_0^\infty : F_1 \otimes S_0^2(T^*_p M_0) \to E_0^\perp \subset V_0,
\]

with respect to the \(L^2\) inner product on \(\partial M_0\), defined by the following. Given

\[
a_{ijk} z^i dz^j \otimes dz^k \in F_1 \otimes S_0^2(T^*_p M_0)
\]
then there is a solution $\tilde{h} \in \Gamma(S_0^2(T^*M_0))$ of $P(\tilde{h}) = 0$ satisfying

$$\tilde{h} = \frac{a_{ijkl}z^iz^jdz^k \odot dz^l}{|z|^m} + O(|z|^{2-m})$$

as $|z| \to 0$, and such that $\tilde{h} \in L^2(M_0 \setminus B(p_0, \epsilon))$. Again, from Proposition 6.1 there is an expansion

$$\tilde{h} = \sigma_0x^{m-1} + o(x^{m-1}),$$

where $x^2\sigma_0 \in \nabla_0$ as $x \to 0$. Since $\tilde{h}$ is only unique up to the solutions from the $S_0^\infty$ construction above, the mapping

$$S_1^\infty(a_{ijk}z^iz^jdz^k \odot dz^l) = x^2\sigma_0|\partial M_0$$

is well-defined modulo the image of $S_0^\infty$ and therefore gives a well-defined mapping to $E_0^\perp$. Letting $E_1 = Im(S_1^\infty)$, in a similar fashion, we may define a mapping on 2-germs

$$S_2^\infty : S_0^2(\mathbb{R}^m) \otimes S_0^2(T^*_pM_0) \to (E_0 \oplus E_1)^\perp.$$  

by the following. Given

$$a_{ijkl}z^iz^jdz^k \odot dz^l \in F_2 \otimes S_0^2(T^*_pM_0),$$

then there is a solution $\tilde{h} \in \Gamma(S_0^2(T^*M_0))$ of $P(\tilde{h}) = 0$ satisfying

$$\tilde{h} = \frac{a_{ijkl}z^iz^jdz^k \odot dz^l}{|z|^{m+2}} + O(|z|^{1-m})$$

as $|z| \to 0$, and such that $\tilde{h} \in L^2(M_0 \setminus B(p_0, \epsilon))$. Again, from Proposition 6.1 there is an expansion

$$\tilde{h} = \sigma_0x^{m-1} + o(x^{m-1}),$$

where $x^2\sigma_0 \in \nabla_0$ as $x \to 0$. Since $\tilde{h}$ is only unique up to the solutions from the $S_0^\infty$ and $S_1^\infty$ constructions above, the mapping

$$S_2^\infty(a_{ijkl}) = x^2\sigma_0|\partial M_0$$

is well-defined modulo the image of $S_0^\infty$ and $S_1^\infty$ and therefore gives a well-defined mapping to $(E_0 \oplus E_1)^\perp$.

**Proposition 7.5.** Let $(M_0^m, \partial M_0, g, x)$ be an AHE orbifold with one singular point $p_0$ of type $\Gamma_n (m = 2n)$ and non-degenerate. Then we have the following:

(1) $S_0^\infty$ is injective.

(2) $S_1^\infty$ is injective.

(3) The kernel of $S_2^\infty$ is the space of 2-germs of the form $|z|^{m+2}\mathcal{K}_{euc}(\frac{\alpha_1}{|z|^m})$, where $\alpha_1$ is a one-form with linear coefficients$^2$

$^2$ The reader should compare this statement with [Biq13, Lemma 10.1], from which our conclusion differs.
Proof. In any of these cases, let $s$ be the solution from the above construction which is blowing up at the point, and $L^2$ in a neighborhood of infinity, so that top $L^2$ weight vanishes. By Proposition 6.1, $s$ admits an expansion

$$s = \sigma_1 x^m + o(x^m)$$

(7.21)

where $x^2 \sigma_1 \in V_1$, as $x \to 0$. A straightforward calculation shows that $B s = o(x^m)$ as $x \to 0$. By Lemma 6.4, $P_1(B s) = 0$, so by Proposition 6.1, we must have

$$B s = \alpha_n x^\delta + o(x^\delta)$$

(7.22)

where $\alpha_n \in \Omega_1^1(M_0)$, as $x \to 0$. Then, again by the Lemma 6.4 and Proposition 6.1 and the fact that $\delta_1 > \delta_2$, we have that $d B s \equiv 0$. We need the following lemma.

Lemma 7.6. The function $f = \delta B s$ satisfies

$$B s = \frac{1}{\Lambda} df,$$

(7.23)

$$P_0(f) = 0.$$  

(7.24)

Proof. Since $tr_g s = 0$ and $d B s = 0$, using Lemma 6.4 we have

$$df = d \delta s = (\Delta_H - \delta d) \delta s = \Lambda \delta s - \delta d \delta s = \Lambda \delta s,$$

(7.25)

and (7.23) follows. Lemma 6.4 directly implies (7.24) since $P s = 0$. \hfill \Box

Letting $\tilde{f} = \frac{2f}{(2-m)\Lambda}$, from Lemma 7.6 we have that

$$P(s + \tilde{f} g) = 0,$$

(7.26)

$$B(s + \tilde{f} g) = 0.$$  

(7.27)

Consequently,

$$d_g Ric(s + \tilde{f} g) = \Lambda (s + \tilde{f} g),$$

(7.28)

$$s + \tilde{f} g = o(x^{m-1}),$$

(7.29)

as $x \to 0$.

Then by Theorem 7.4 and Lemma 7.2 it follows that

$$s + \tilde{f} g = \delta^*(\tilde{\alpha})$$

(7.30)

for some one-form $\alpha$ defined in a neighborhood of the boundary. Since $(M_0, g)$ is a non-compact and complete Riemannian manifold there exists a geodesic ray $\gamma$ starting from $p_0$ to infinity. Now applying Lemma 7.3 inductively on $\gamma$ we can analytically continue $\alpha$ to a one-form $\tilde{\alpha}$ defined in a neighborhood of $\gamma$ such that

$$s + \tilde{f} g = \delta^*(\tilde{\alpha}).$$

(7.31)

Equivalently,

$$s = \mathcal{K}(\tilde{\alpha}),$$

(7.32)

for a 1-form $\tilde{\alpha}$ satisfying

$$P_1(\tilde{\alpha}) = 0.$$  

(7.33)
In the case of $S_0^\infty$, let $s$ be the solution of $Ps = 0$ whose leading term is $|z|^{2-m}h_0$ at the point, where $h_0$ is a 0-germ at $p_0$, and $s = o(x^{m-1})$ as $x \to 0$. Since $s = O(|z|^{-m+2})$ around $p_0$ and (7.32) is satisfied, by Lemma 7.3, $\tilde{\alpha} = O(|z|^{-m+3})$. The Bochner formula implies that there is no nontrivial $L^2$ kernel for $P_1$. Since $P_1\tilde{\alpha} = 0$, this implies that $\tilde{\alpha} = 0$ by a standard removable singularity theorem for solutions of $P_1$, and therefore $h_0 = 0$.

In case of $S_1^\infty$, given any 1-germ $h_1$ at $p_0$ let $s$ be the solution of $Ps = 0$ whose leading term is $|z|^{-m}h_1$ at $p_0$, and $x^2s|_{\partial M_0} \in E_0^{\perp}$. If $h_1 \in \text{Ker}(S_1^\infty)$ then $s = O(x^{m-1})$ as $x \to 0$. Since $s = O(|z|^{-m+1})$ as $|z| \to 0$ and (7.32) is satisfied, by Lemma 7.3, $\tilde{\alpha} = O(|z|^{-m+2})$. This implies that $\tilde{\alpha} = 0$ by the removable singularity theorem, since there is no solution with that leading term which is also invariant under group action, and therefore $h_1 = 0$.

In case of $S_2^\infty$, given any 2-germ $h_2$ at $p_0$ let $s$ be the solution of $Ps = 0$ whose leading term is $|z|^{2-m}h_2$ at $p_0$, and $x^2s|_{\partial M_0} \in (E_0 \oplus E_1)^{\perp}$. If $h_2 \in \text{Ker}(S_2^\infty)$ then $s = O(x^{m-2})$ as $x \to 0$. Since $s = O(|z|^{-m+1})$ as $|z| \to 0$ and (7.32) is satisfied by Lemma 7.3, $\tilde{\alpha} = O(|z|^{-m+2})$. This implies that

\begin{equation}
\alpha_1 = \frac{\alpha_1}{|z|^m} + O(|z|^{-m+2}),
\end{equation}

where $\alpha_1$ is a one-form with linear coefficients. Then we have

\begin{equation}
s = \mathcal{K}(\tilde{\alpha}) = \mathcal{K}_{\text{euc}}(\frac{\alpha_1}{|z|^m}) + O(|z|^{-m+1})
\end{equation}

as $|z| \to 0$. For the converse, since there is no nontrivial $L^2$ kernel for $P_1$, given a linear form $\alpha_1 = a_{ij}z^i dz^j$, there exists a unique solution $\tilde{\alpha}_1$ of $P_1(\tilde{\alpha}_1) = 0$ which is in $L^2(M_0 \setminus B(p_0, \epsilon))$ satisfying

\begin{equation}
\tilde{\alpha}_1 = |z|^{2-m}a_{ij}z^i dz^j + O(|z|^{2-m}),
\end{equation}

as $|z| \to 0$. By Proposition 6.1, $\tilde{\alpha}_1$ admits an expansion

\begin{equation}
\tilde{\alpha}_1 = \alpha_1 x^{\delta^+_p, p_1} + o(x^{\delta^+_p, p_1}),
\end{equation}

where $x\alpha_1 \in \Omega^1_{\text{cl}} M_0$, as $x \to 0$. From Proposition 6.4, it follows easily that $h = \mathcal{K}\tilde{\alpha}_1$ is a solution of $P h = 0$, and thus the 2-germ $|z|^{m+2}\mathcal{K}_{\text{euc}}(\frac{\alpha_1}{|z|^m})$, is in the kernel of $S_2^\infty$, where $\alpha_1$ is any one-form with linear coefficients. \hfill \Box

8. From boundary tensors to germs

The following lemma describes a “duality” between boundary tensors and germs. We state the following

Lemma 8.1. Let $(M^m, \partial M, g, x)$ be AH with one orbifold point $p_0$ of type $\Gamma_n$. Let $\sigma_k, \tau_k$ be $k$-germs at $p_0$ and Let $s$ and $t$ be two solutions for $P$ such that

\begin{equation}
\begin{aligned}
s &= \sigma_k |z|^{2-m-2k} + o(z^{2-m-k}) \quad \text{as } |z| \to 0 \\
s &= \sigma_\infty x^{\delta^+_p, p} + o(x^{\delta^+_p, p}) \quad \text{as } x \to 0,
\end{aligned}
\end{equation}

\begin{equation}
\begin{aligned}
t &= \tau_k |z|^{2-m-2k} + o(z^{2-m-k}) \quad \text{as } |z| \to 0 \\
t &= \tau_\infty x^{\delta^+_p, p} + o(x^{\delta^+_p, p}) \quad \text{as } x \to 0.
\end{aligned}
\end{equation}
where \( x^2 \sigma_\infty \in \overline{V}_0 \), and
\[
\begin{align*}
t &= \tau_k + o(|z|^k) \text{ as } |z| \to 0 \\
t &= \tau_\infty x^{\delta_{0,p}} + o(x^{\delta_{0,p}}) \text{ as } x \to 0,
\end{align*}
\]
(8.2) where \( x^2 \tau_\infty \in \overline{V}_0 \). For \( k = 0 \) we have
\[
(m - 2) \frac{\omega_{m-1}}{n} \langle \langle \sigma_0, \tau_0 \rangle \rangle = (m - 1) \langle x^2 \sigma_\infty, x^2 \tau_\infty \rangle,
\]
(8.3) for \( k = 1 \) we have
\[
\frac{\omega_{m-1}}{n} \langle \langle \sigma_1, \tau_1 \rangle \rangle = (m - 1) \langle x^2 \sigma_\infty, x^2 \tau_\infty \rangle,
\]
(8.4) and for \( k = 2 \) we have
\[
\frac{2 \omega_{m-1}}{mn} \langle \langle \sigma_2, \tau_2 \rangle \rangle = (m - 1) \langle x^2 \sigma_\infty, x^2 \tau_\infty \rangle,
\]
(8.5) where the left inner product is the germ inner product at \( p_0 \) (see the proof for the definition), and the right inner product is the \( L^2 \) inner product on the boundary.

Proof. Let \( M_{0,\epsilon} = M_0 \setminus (B(p_0, \epsilon) \cup \partial M_0 \times [0, \epsilon]) \). Integrating by parts on \( M_{0,\epsilon} \) implies that
\[
\int_{M_{0,\epsilon}} (\langle Ps, t \rangle - \langle s, Pt \rangle) dV_g = \int_{\partial M_{0,\epsilon}} (\langle -\nabla \tilde{n}s, t \rangle + \langle s, \nabla \tilde{n}t \rangle) dS_g.
\]
(8.6) The interior boundary integral is given by
\[
\lim_{\epsilon \to 0} \int_{\partial B(p_0, \epsilon)} (\langle -\nabla \tilde{n}s, t \rangle + \langle s, \nabla \tilde{n}t \rangle) dS_g = (2 - m - 2k) \int_{S_{m-1}/\Gamma_n} \langle \sigma_k, \tau_k \rangle dS_{euc},
\]
(8.7) where the inner product denotes contraction on only the indices of the germs corresponding to \( S_0^2(T_{p_0}^*(M_0)) \).

Recall from Proposition 6.1 that \( \delta_{0,p} = m - 1 \) and \( \delta_{0,p} = 0 \). Since \( \sqrt{\det g} = x^{1-m} \sqrt{\det g} \) when we restrict both metrics to \( \partial M_0 \times \{ \epsilon \} \), it follows that the outer boundary integral is given by
\[
\lim_{\epsilon \to 0} \int_{\partial M_0 \times \{ \epsilon \}} (\langle -\nabla \tilde{n}s, t \rangle + \langle s, \nabla \tilde{n}t \rangle) dS_g = (m - 1) \int_{\partial M_0} \langle x^2 \sigma_\infty, x^2 \tau_\infty \rangle dS_{euc}.
\]
(8.8) If \( k = 0 \) then then working in a coordinate chart around \( p_0 \) with \( g_{ij}(p_0) = \delta_{ij} \) we have
\[
\sigma_0 = H_{kl} dz^k \otimes dz^l,
\]
(8.9)
\[
\tau_0 = K_{kl} dz^k \otimes dz^l,
\]
(8.10) and (8.3) follows immediately, with \( \langle \langle \sigma_0, \tau_0 \rangle \rangle = \sum_{k,l} H_{kl} K_{kl} \).

If \( k = 1 \), then
\[
\sigma_1 = H_{jkl} z^j dz^k \otimes dz^l,
\]
(8.11)
\[
\tau_1 = K_{jkl} z^j dz^k \otimes dz^l,
\]
(8.12) and (8.4) follows from (8.22) with \( \langle \langle \sigma_1, \tau_1 \rangle \rangle = \sum_{j,k,l} H_{jkl} K_{jkl} \).
Finally, if \( k = 2 \), then
\[
\sigma_2 = H_{ijkl}z^i z^j dz^k \otimes dz^l,
\]
\[
\tau_2 = K_{ijkl}z^i z^j dz^k \otimes dz^l,
\]
and (8.14) follows from (8.23) with \( \langle \sigma_2, \tau_2 \rangle = \sum_{i,j,k,l} H_{ijkl} K_{ijkl} \).

Next, we explain how to construct a map which sends boundary tensors to germs at \( p_0 \). Identify a neighborhood of the boundary with \( \partial M_0 \times [0, \epsilon) \). Let \( \varphi \) be a cutoff function which is 1 on \( \partial M_0 \times [0, \frac{\epsilon}{2}) \) and is zero on \( M_0 \setminus (\partial M_0 \times [0, \epsilon)) \). Given \( \sigma_0 \in V_0 \), consider the tensor \( \varphi x^{-2} \sigma_0 \). By solving a finite number of terms of a power series, we can then find \( s \) of the form \( s = \varphi x^{-2} \sigma_0 + \tilde{s} \), where \( \tilde{s} \) are lower order terms supported in a neighborhood of the boundary, such that \( Ps \) is in \( L^2 \) in a neighborhood of the boundary. (For this power series method, see for example [AG11] for the Hodge Laplacian acting on forms; the same technique applies to solutions of \( P \).) Then, using the non-degeneracy assumption, we can find a solution of \( Ps = 0 \) defined on all of \( M_0 \), which is in \( L^2 \) on any compact subset of \( M_0 \), so that
\[
s = x^{-2} \sigma_0 + o(1)
\]
as \( x \to 0 \). Note that by non-degeneracy, the solution \( s \) is unique. The solution \( s \) admits the following expansion around \( p_0 \)
\[
s = h_0 + o(1) \quad \text{as} \quad |z| \to 0,
\]
where \( h_0 \) is a 0-germ at \( p_0 \). Define a mapping
\[
S^0_\infty : V_0 \to S^2(T^*_p M_0), \quad \text{by}
\]
\[
S^0_\infty (\sigma_0) \equiv h_0.
\]
This mapping enjoys the following property.

**Lemma 8.2.** The mapping \( S^0_\infty : V_0 \to S^2(T^*_p M_0) \) is surjective.

*Proof.* Assume \( S^0_\infty \) is not surjective. Then there exist a 0-germ \( h_0 \) such that \( h_0 \) is orthogonal to the image of \( S^0_\infty \). Let \( \tilde{h} \) be the solution of \( P\tilde{h} = 0 \) with following leading terms
\[
\tilde{h} = h_0 |z|^{2-m} + o(|z|^{2-m}) \quad \text{as} \quad |z| \to 0,
\]
\[
\tilde{h} = S^0_\infty (h) x^{\delta^+_p} + o(x^{\delta^+_p}) \quad \text{as} \quad x \to 0.
\]
Also for a boundary tensor \( \tau \in V_0 \), let \( \tilde{\tau} \) be the solution of \( P\tilde{\tau} = 0 \) with following leading terms
\[
\tilde{\tau} = S^0_\infty (\tau) + o(1) \quad \text{as} \quad |z| \to 0,
\]
\[
\tilde{\tau} = x^{-2} \tau + o(1) \quad \text{as} \quad x \to 0.
\]

By applying Lemma 8.1 to \( \tilde{\tau} \) and \( \tilde{h} \) we see that
\[
\langle h_0, S^0_\infty \tau \rangle = c \langle S^0_\infty h_0, \tau \rangle,
\]
for every \( \tau \in V_0 \), where \( c \neq 0 \), so \( S^0_\infty (h) = 0 \). This implies that \( h_0 = 0 \) by Proposition 7.5 (1), which is a contradiction. \( \square \)
We next define a mapping from $E^\perp_0 \subset V_0$ (the orthogonal complement in $L^2$) to 1-germs by the following. Given $\sigma_0 \in E^\perp_0$, as above let $s$ be the solution of $Ps = 0$ defined on all of $M_0$, which is in $L^2$ on any compact subset of $M_0$, and satisfies
\begin{equation}
\tag{8.22}
s = x^{-2}\sigma_0 + o(1),
\end{equation}
as $x \to 0$. By Lemma 8.1, $s$ admits an expansion
\begin{equation}
\tag{8.23}
s = h_1 + o(|z|),
\end{equation}
as $|z| \to 0$, where $h_1$ is a 1-germ at $p_0$. We then define the mapping
\begin{equation}
\tag{8.24}
S^1_{\infty} : E^\perp_0 \to F_1 \otimes S^2(T^*_{p_0}M_0), \text{ by }
\end{equation}
\begin{equation}
\tag{8.25}
S^1_{\infty}(\sigma_0) \equiv h_1.
\end{equation}
This mapping is also onto.

**Lemma 8.3.** The mapping $S^1_{\infty} : E^\perp_0 \to F_1 \otimes S^2(T^*_{p_0}M_0)$ is surjective.

**Proof.** By Proposition 7.5 (2), $S^1_{\infty}$ is injective, so by the same argument in the proof of Lemma 8.2, it follows that $S^1_{\infty}$ is surjective. \]

We next define a mapping from $(E_0 \oplus E_1)^\perp \subset V_0$ (the orthogonal complement in $L^2$) to 2-germs by the following. Given $\sigma_0 \in (E_0 \oplus E_1)^\perp$, as above let $s$ be the solution of $Ps = 0$ defined on all of $M_0$, which is in $L^2$ on any compact subset of $M_0$, and satisfies
\begin{equation}
\tag{8.26}
s = x^{-2}\sigma_0 + o(1),
\end{equation}
as $x \to 0$. By Lemma 8.1, $s$ admits an expansion
\begin{equation}
\tag{8.27}
s = h_2 + o(|z|^2),
\end{equation}
as $|z| \to 0$, where $h_2$ is a 2-germ at $p_0$.

We then define the mapping
\begin{equation}
\tag{8.28}
S^2_{\infty} : (E_0 \oplus E_1)^\perp \to F_2 \otimes S^2(T^*_{p_0}M_0), \text{ by }
\end{equation}
\begin{equation}
\tag{8.29}
S^2_{\infty}(\sigma_0) \equiv h_2.
\end{equation}
The following Lemma says that $S^2_{\infty}$ is not surjective, and also gives a complete characterization of its image.

**Lemma 8.4.** Let $T = \text{Im}(S^2_{\infty})$ then
\begin{equation}
\tag{8.29}
T = S^\perp,
\end{equation}
where the orthogonal complement is with respect to the 2-germ inner product, and
\begin{equation}
\tag{8.30}
S = \left\{ |z|^{m+2}\mathcal{K}_\text{euc}\left(\frac{\alpha}{|z|^m}\right) \mid \alpha \text{ is a one-form with linear coefficients} \right\}.
\end{equation}

\footnote{The reader should compare to [Biq13, Lemma 11.1], where it is claimed that this mapping is surjective.}
Proof. Let \( h_2 \in T \) so \( h_2 = S^2_\infty(\tau) \) and let \( \tilde{h} \) be the solution of \( P\tilde{h} = 0 \) with following leading terms

\[
\begin{align*}
(8.31) & \quad \tilde{h} = h_2 + o(|z|^2) \quad \text{as} \quad |z| \to 0, \\
(8.32) & \quad \tilde{h} = x^{-2}\tau + o(1) \quad \text{as} \quad x \to 0.
\end{align*}
\]

For any \( t_2 \in S \), Let \( \tilde{t} \) be the solution of \( P\tilde{t} = 0 \) with the following leading terms

\[
\begin{align*}
(8.33) & \quad \tilde{t} = t_2|z|^{-2-m} + o(|z|^{-m}) \quad \text{as} \quad |z| \to 0, \\
(8.34) & \quad \tilde{t} = S^2_\infty(t_2)x^{-2}x^{m-1} + o(x^{m-1}) \quad \text{as} \quad x \to 0.
\end{align*}
\]

Applying Lemma 8.1 to \( \tilde{h} \) and \( \tilde{t} \) we obtain

\[
(8.35) \quad \langle h_2, t_2 \rangle = \langle S^2_\infty(\tau), t_2 \rangle = c\langle \tau, S^2_\infty(t_2) \rangle = 0,
\]

for some constant \( c \neq 0 \), so by Proposition 7.3, it follows that \( T \subset S^\perp \).

Next, assume that \( S^2_\infty \) is not surjective onto \( S^\perp \). Then there exists a nontrivial \( s_2 \in S^\perp \cap T^\perp \). Let \( \tilde{s} \) be the solution of \( P\tilde{s} = 0 \) with the following leading terms

\[
\begin{align*}
(8.36) & \quad \tilde{s} = s_2|z|^{-2-m} + o(|z|^{-m}) \quad \text{as} \quad |z| \to 0, \\
(8.37) & \quad \tilde{s} = S^2_\infty(s_2)x^{-2}x^{m-1} + o(x^{m-1}) \quad \text{as} \quad x \to 0.
\end{align*}
\]

Also, for any boundary tensor \( \tau \in (E_0 \oplus E_1)^\perp \), let \( \tilde{t} \) be the solution of \( P \) with following leading terms

\[
\begin{align*}
(8.38) & \quad \tilde{t} = S^2_\infty(\tau) + o(|z|^2) \quad \text{as} \quad |z| \to 0, \\
(8.39) & \quad \tilde{t} = x^{-2}\tau + o(1) \quad \text{as} \quad x \to 0.
\end{align*}
\]

Applying Lemma 8.1 to \( \tilde{t} \) and \( \tilde{s} \) we obtain

\[
(8.40) \quad \langle S^2_\infty(s_2), \tau \rangle = c\langle s_2, S^2_\infty(\tau) \rangle = 0,
\]

which implies that \( S^2_\infty(s_2) = 0 \) so \( s_2 \in S \) by Proposition 7.3 (3), which gives a contradiction. \( \square \)

Next, we give an explicit description of the elements of \( S \).

**Lemma 8.5.** Let \( \alpha = \frac{\alpha_i}{|z|^m} \) where \( \alpha_1 = a_{ij}z_idz^j \) is a one-form with linear coefficients. Then

\[
(8.41) \quad K_{euc}(\alpha)_{ij} = \frac{1}{2} \left( a_{ij} + a_{ji} \right) + \frac{mz_i z_j a_{pi}}{|z|^{m+2}} \quad \text{where} \quad K_{euc}(\alpha) = |z|^{-m-2}K_{ijkl}z^i z^j dz^i dz^j,
\]

Furthermore, we may write

\[
(8.42) \quad K_{euc}(\alpha) = |z|^{-m-2}K_{ijkl}z^i z^j dz^i dz^j,
\]

where

\[
K_{ijkl} = \frac{1}{2} \left( a_{ij} + a_{ji} \right) \delta_{kl} - \frac{m}{2} \left( \delta_{ik} a_{lj} + \delta_{il} a_{kj} \right) - \frac{m}{2} \left( \delta_{kj} a_{li} + \delta_{lj} a_{ki} \right)
\]

\[
+ \frac{-tr(a) \delta_{kl} \delta_{ij}}{m} + \frac{\delta_{ij}}{2} (a_{lk} + a_{kl}).
\]
Proof. Let \( \alpha_1 = a_{ij} x^i dx^j \) be a one-form with linear coefficients. Then we have

\[
(\delta^*_{\text{euc}} \alpha)_{ij} = \frac{1}{2}(\partial_i \alpha_j + \partial_j \alpha_i)
\]

(8.44)

Using \( \partial_i(|z|^{-m}) = -\frac{mz_i}{|z|^{m+2}} \) we obtain

\[
(\delta^*_{\text{euc}} \alpha)_{ij} = \frac{1}{2}\left(\frac{a_{ij} + a_{ji}}{|z|^m} + \frac{-mz_i z_p a_{pj} - m z_j z_k a_{ki}}{|z|^{m+2}}\right).
\]

(8.45)

Also,

\[
\delta_{\text{euc}}(\alpha) = -\partial_k X_k = -\partial_k \left(\frac{a_{pk} z_p}{|z|^m}\right) = -\frac{\text{tr}(a)}{|z|^m} + \frac{m z_k z_p a_{pk}}{|z|^{m+2}}.
\]

The formula \(8.41\) follows directly from these computations, and \(8.42\) follows easily from this.

\[\square\]

Lemma 8.6. The tensor \( r^2 g_{\text{euc}} = P_{ijkl} z^i z^j dz^k dz^l \), where

\[
P_{ijkl} = \delta_{ij} \delta_{kl}.
\]

The tensor \( r^2 dr \otimes dr = Q_{ijkl} z^i z^j dz^k dz^l \), where

\[
Q_{ijkl} = \frac{1}{2} (\delta_{ik} \delta_{jl} + \delta_{il} \delta_{jk})
\]

The tensor \( r^4 \theta \otimes \theta = A_{ijkl} z^i z^j dz^k dz^l \), where

\[
A_{ijkl} = \frac{1}{2} (J^k_i J^l_j + J^l_i J^k_j).
\]

Proof. The proof is immediate.

\[\square\]

Lemma 8.7. We have the following identities:

\[
\langle \langle Q, K \rangle \rangle = \frac{2 - m^2 - m}{2} \text{tr}(a),
\]

(8.50)

\[
\langle \langle A, K \rangle \rangle = \frac{m + 2}{2} \text{tr}(a),
\]

(8.51)

\[
\langle \langle K, P \rangle \rangle = 0.
\]

(8.52)

Proof. We compute

\[
\langle \langle Q, K \rangle \rangle = \frac{1}{2} \left((a_{ij} + a_{ji}) \delta_{ij} - \frac{m}{2} (\delta_{i} a_{jj} + \delta_{i} a_{ij}) - \frac{m}{2} (\delta_{ij} a_{ji} + \delta_{ij} a_{ii})\right)
\]

\[
+ \frac{-\text{tr}(a) \delta_{ij} \delta_{ij}}{m} + \frac{\delta_{ij}}{2} (a_{ji} + a_{ij})
\]

\[
= \frac{1}{2} \left(2 \text{tr}(a) - m^2 \text{tr}(a) - m \text{tr}(a)\right) = \frac{2 - m^2 - m}{2} \text{tr}(a).
\]

(8.53)
This proves the first identity. For the second identity, we have

\[
\langle \langle A, K \rangle \rangle = \frac{1}{2} \left( (a_{ij} + a_{ji}) \delta_{kl} J_i^k J_j^l - \frac{m}{2} (J_i^k J_j^l \delta_{ij} a_{kl} + J_i^k J_j^l \delta_{ik} a_{kj}) \right. \\
- \frac{m}{2} (J_i^k J_j^l \delta_{kj} a_{li} + J_i^k J_j^l \delta_{lj} a_{ki}) \left. + \frac{-tr(a) \delta_{kl} \delta_{ij} J_i^k J_j^l}{m} + J_i^k J_j^l \delta_{ij} \left( \frac{a_{ik} + a_{kl}}{2} \right) \right) \\
= \frac{1}{2} \left( 2tr(a) + mtr(a) \right) - tr(a) + \frac{m + 2}{2} tr(a).
\]

Finally, the last identity follows from the fact that \( K \) is traceless on the first and last two pairs of indices. \( \Box \)

**Lemma 8.8.** We have the following identities:

\[
B_{\text{euc}}(r^2 dr \otimes dr) = (-2n)r dr, \\
B_{\text{euc}}(r^4 \theta \otimes \theta) = 2r dr, \\
B_{\text{euc}}(r^2 g_{\text{euc}}) = 2(n - 1)r dr.
\]

**Proof.** Recall that if \( H = H_{ijkl} x^i dx^j \otimes dx^k \otimes dx^l \) then

\[
B_{\text{euc}}(H) = (-2H_{kijkl} + H_{jikkk}) x^i dx^j.
\]

It follows that

\[
B_{\text{euc}}(r^2 dr \otimes dr) = (-2) \frac{-\delta_{kk} \delta_{ij} + \delta_{kl} \delta_{jk}}{2} + \frac{\delta_{jk} \delta_{ik} + \delta_{kj} \delta_{ik}}{2}) x^j dx^l \\
= (-2n) \sum_i x^i dx^i.
\]

Next,

\[
B_{\text{euc}}(r^2 \theta \otimes \theta) = \left( -2 \frac{J_i^k J_j^l + J_i^l J_j^k}{2} + \frac{J_i^k J_j^l + J_i^l J_j^k}{2} \right) x^j dx^l = 2 \sum_i x^i dx^i.
\]

The formula \((8.57)\) follows easily. \( \Box \)

We next construct a special symmetric 2-tensor \( k_1 \) on the Calabi manifold.

**Proposition 8.9.** Let \( \sigma_2 \) denote the quadratic 2-tensor

\[
\sigma_2 = r^2 (-dr \otimes dr - (2n - 1)r^2 \theta \otimes \theta + g_{\text{euc}})
\]

\[
= S_{ijkl} x^i x^j dx^k dx^l.
\]

Then

\[
tr_{\text{euc}}(\sigma_2) = 0, \\
B_{\text{euc}}(\sigma_2) = 0, \\
\Delta_{\text{euc}}(\sigma_2) = 0,
\]

and

\[
\langle \langle S, K \rangle \rangle = 0,
\]
for all $K_{ijkl}$ of the form in (8.43). Furthermore, there exist a constant $c \neq 0$ and a tensor $k_1$ defined on the Calabi manifold with the following properties

\begin{equation}
  k_1 = c \cdot \sigma_2 + O(r^{-2n+2+\epsilon})
\end{equation}

(8.66)

\begin{align*}
  P_{g_{\text{cal}}}(k_1) &= o \\
  B_{g_{\text{cal}}}(k_1) &= 0,
\end{align*}

as $r \to \infty$.

**Proof.** For equation (8.62), we have

\begin{equation}
  \text{tr}_{\text{euc}}(\sigma_2) = r^2(-1 - (2n - 1) + 2n) = 0.
\end{equation}

(8.67)

For equation (8.63), using Lemma 8.8, we have

\begin{equation}
  B_{\text{euc}}(\sigma_2) = (2n - 2(2n - 1) + 2(n - 1))rdr = 0.
\end{equation}

(8.68)

Next, it follows from Lemma 8.6 that $S_{ijkl} = S_{klij}$, so (8.62) implies (8.64). To see (8.65), Lemma (8.7) implies that

\begin{equation}
  \langle S, K \rangle = \left( (n + 1)(1 - 2n) - (1 - 2n^2 - n) \right) tr(a) = 0.
\end{equation}

(8.69)

Next, let $(H, \Lambda) = (\sigma_2, 0)$ in Proposition 3.2 to find a solution of

\begin{align*}
  P_{g_{\text{cal}}}(k_1) &= \tilde{\lambda}o \\
  k_1 &= c \cdot \sigma_2 + O(r^{-2n+2+\epsilon})
\end{align*}

(8.70)

\begin{align*}
  B_{g_{\text{cal}}}(h) &= 0,
\end{align*}

where $\tilde{\lambda}$ is given by

\begin{equation}
  \tilde{\lambda} = -\frac{c}{\|o\|_2^2} \lim_{r \to \infty} \int_{S_r/T_n} \frac{n + 1}{r} \langle \sigma_2, o \rangle dS_{S_r/T_n}.
\end{equation}

(8.71)

The leading term in the integrand is given by

\begin{equation}
  \frac{n + 1}{r} \langle \sigma_2, o_{\text{euc}} \rangle = c \cdot 2n(1 - n^2)r^{1-m},
\end{equation}

(8.72)

so choosing the scaling constant $c$ to make $\tilde{\lambda} = 1$ finishes the proof. \qed

The following result is the key ingredient to our existence theorem in the AHE case, which will be proved in Section 9.

**Proposition 8.10.** Let $\sigma = c \cdot \sigma_2 = c \cdot S_{ijkl} z^i z^j d^k z^l$ be as Proposition 8.9. Then there exist a solution $\tilde{k}$ of $P_{g_0}(\tilde{k}) = 0$ defined on $M_0$ such that

\begin{align*}
  \tilde{k} &= \sigma + O(|z|) \text{ as } |z| \to 0 \\
  \tilde{k} &= k_{\infty} + o(1) \text{ as } x \to 0,
\end{align*}

(8.73) \quad (8.74)

where $x^2 k_{\infty} \in \nabla_0$.

**Proof.** The proof is immediate from Lemma 8.3 if we let $\sigma$ be the leading term of the tensor $k_1$ found in Proposition 8.9. \qed
9. Completion of proofs

In this section, we complete the proofs of Theorems 1.2 and 1.3. First, for \( u \in \mathbb{R} \), we define

\[
h_{t,u} = g_{cal} + th + utk_1,
\]

where \( h \) was defined in (4.4), and \( k_1 \) was defined in Proposition 8.9.

Lemma 9.1. For \( t \) and \( u \) sufficiently small, \( h_{t,u} \) is a Riemannian metric on \( X^t \). Moreover, for each integer \( \ell \geq 0 \), there exists a constant \( c_\ell \) so that \( h_{t,u} \) satisfies the following estimate on \( X^t \)

\[
|\nabla^\ell (\text{Ric}_{h_{t,u}} - t\Lambda h_{t,u} - t(\lambda + u)\chi_t(p)\sigma)|_{g_{cal}} \leq c_\ell t^2(1 + u^2 + t^2)\rho^{2-\ell}.
\]

Proof. Since \( h \) and \( k_1 \) both have quadratic leading terms, and \( g_{cal} \) is ALE, there exist constants \( c \) and \( c_1 \) so that

\[
|h|_{g_{cal}}(p) \leq c\rho^2(p),
\]

\[
|k_1|_{g_{cal}}(p) \leq c_1 \rho^2(p),
\]

for any \( p \in X^t \). Similarly, there exists a constant \( c_2 \) so that

\[
|\sigma|_{g_{cal}}(p) \leq c_2 \rho^{-2n}(p),
\]

for any \( p \in X^t \). It follows that

\[
|th|_{g_{cal}}(p) \leq c\rho^2(p) \leq \frac{1}{2}ct^{\frac{1}{2}}
\]

\[
|tuk_1|_{g_{cal}}(p) \leq c_1 u\rho^2(p) \leq \frac{1}{2}c_1 ut^{\frac{1}{2}}
\]

\[
|tuo|_{g_{cal}}(p) \leq c_2 u\rho^2(p) \leq \frac{1}{2}c_2 ut^{\frac{1}{2}}
\]

The right hand side is independent of \( p \) so \( h_{t,u} \) is a Riemannian metric for sufficiently small \( t \) and \( u \). Recall that

\[
d_{g_{cal}}\text{Ric}(h) = \Lambda g_{cal} + \lambda \sigma
\]

\[
d_{g_{cal}}\text{Ric}(\sigma) = 0
\]

\[
d_{g_{cal}}\text{Ric}(k_1) = 0,
\]

so we have

\[
\text{Ric}(g_{cal} + th + tuk_1) = \text{Ric}(g_{cal}) + d_{g_{cal}}\text{Ric}(th + tuk_1) + Q(th + tuk_1),
\]

where, using (4.14), \( Q \) is a nonlinear term satisfying

\[
|Q(th + tuk_1)| \leq b\left( |Rm_{g_{cal}}||th + tuk_1|^2 + |th + tuk_1||\nabla^2 (th + tuk_1)| + |\nabla (th + tuk_1)|^2 \right),
\]
for some constant $b$ and for $t$ and $u$ sufficiently small, and where the norms and covariant derivatives are measured with respect to the Calabi metric. Note that
\begin{align}
|\nabla^\ell h| &\leq c_\ell \rho^{2-\ell}, \\
|\nabla^\ell o| &\leq c_\ell \rho^{-2n-\ell}, \\
|\nabla^\ell k_1| &\leq c_\ell \rho^{2-\ell},
\end{align}
so it follows that
\begin{equation}
|Q(th + tu_k_1)| \leq b_1 t^2 (1 + t^2 + u^2) \rho^2
\end{equation}
and consequently
\begin{equation}
|Ric_{ht,u} - t\Lambda h_{t,u} - t\lambda o - tu o| \leq c_0 t^2 (1 + t^2 + u^2) \rho^2,
\end{equation}
for some constant $c_0$, so the lemma for $\ell = 0$ follows. The proof for $\ell > 0$ is similar and is omitted.

We next define
\begin{equation}
k = \begin{cases}
\tilde{k} & M_0 \setminus B(p_0, 2t^{1/4}) \\
(1 - \chi_{t}(r))\phi_1^* \tilde{k} + \chi_{t}(r)t^2 k_1 & (1/2)t^{-1/4} < r < 2t^{-1/4}, \\
t^2 k_1 & r < (1/2)t^{-1/4}
\end{cases}
\end{equation}
where $\chi_t$ is the cutoff function we defined in (1.11), and $\tilde{k}$ was defined in Proposition 8.10. For $u \in \mathbb{R}$, consider the metric on $M^t$ defined by $g_{t,u} = g_t + uk$, where $g_t$ is the refined metric defined in (4.20). Note that $g_{t,u}$ has the following expression
\begin{equation}
g_{t,u} = \begin{cases}
g_0 + uk & M_0 \setminus B(p_0, 2t^{1/4}) \\
(1 - \chi_{t}(r))\phi_1^*(g_0 + uk) + \chi_{t}(r)th_{t,u} & (1/2)t^{-1/4} < r < 2t^{-1/4}, \\
\chi_{t}(r)th_{t,u} & r < (1/2)t^{-1/4}
\end{cases}
\end{equation}
This metric satisfies the following modification of Lemma 4.2.

**Lemma 9.2.** For $t$ and $u$ sufficiently small, $g_{t,u}$ is a Riemannian metric on $M^t$. Moreover, for each integer $\ell \geq 0$, there exist a constant $c_\ell$ so that $g_{t,u}$ satisfies the following estimate on $X^t$
\begin{equation}
|\nabla^\ell (Ric_{g_{t,u}} - \Lambda g_{t,u} - t(\lambda + u)\chi_t(\rho) o)_{\text{scal}}| \leq c_\ell t^2 (1 + t^2 + u^2) \rho^{2-\ell}.
\end{equation}

**Proof.** If $r < (1/2)t^{-1/4}$, then the proof follows from the Lemma 9.1 by scaling. If
\begin{equation}
(1/2)t^{-1/4} < r < 2t^{-1/4},
\end{equation}
then
\begin{equation}
g_{t,u} = (1 - \chi_{t}(r))(\phi_1^*(uk + g_0)) + \chi_{t}(r)(th_t + t^2 uk_1).
\end{equation}
We have that
\begin{align}
\phi_1^* g_0 &= tg_{euc} + t^2 H + O(t^2 \rho^2), \\
th_t &= tg_{euc} + t^2 H + tO(r^{-2n}) + t O(r^{2-2n+\epsilon}),
\end{align}
and
\begin{align}
\phi^*_t(u\tilde{k}) &= t^2(u\sigma) + ut^2O(r^{-2n+2+\epsilon}) \\
t^2u_k &= t^2u(\sigma) + t^2O(r^{-2n+2+\epsilon}),
\end{align}
as \(t \to 0\). Since the quadratic terms agree in each of these, we have the estimate
\begin{align}
g_{t,u} = tg_{\text{euc}} + t^2H + t^2u\sigma + tO(t^3\rho^3),
\end{align}
as \(t \to 0\). Writing
\begin{align}
\theta = g_t - tg_{\text{euc}},
\end{align}
expanding the Ricci tensor near \(t \cdot g_{\text{euc}}\) we have
\begin{align}
\text{Ric}_{g_{t,u}} = d_{\text{euc}}\text{Ric}(\theta) + Q(\theta),
\end{align}
where \(Q(\theta)\) satisfies the estimate
\begin{align}
t^2|Q(\theta)| \leq c(|\theta||\nabla^2\theta| + |\nabla\theta|^2),
\end{align}
for some constant \(c\), where the norms are with respect to \(g_{\text{euc}}\). More precisely,
\begin{align}
\theta = t^2H + t^2u\sigma + tO(t^3\rho^3)
\end{align}
implies that
\begin{align}
|\nabla^2\theta| |\theta| &= O(t^4(1+u^2+t^2)\rho^2), \\
|\nabla\theta|^2 &= O(t^4(1+u^2+t^2)\rho^2).
\end{align}
From (4.4) it follows that
\begin{align}
d_{\text{euc}}\text{Ric}(\theta) = t\Lambda g_{\text{euc}} + O(t^2(1+u^2+t^2)\rho^2),
\end{align}
where \(O\) is with respect to the Euclidean metric. Note that (9.22) implies that
\begin{align}
t\lambda\chi_t(\rho) = O(t^2\rho^2).
\end{align}
Finally we obtain
\begin{align}
\text{Ric}_{g_{t,u}} - \Lambda g_{t,u} - t(\lambda + u)\chi_t(\rho) = O(t^2(1+t^2+u^2)\rho^2),
\end{align}
so the lemma for \(\ell = 0\) follows. For \(\ell > 0\), the proof is similar and is omitted. \(\Box\)

**Proof of Theorem 1.2.** We apply Proposition 5.2, keeping track of the dependence on the variable \(u\). From Proposition 5.2 we obtain a solution of
\begin{align}
\Phi(\hat{g}_{t,u,v}) = \lambda(t,u)(\chi_t\rho).
\end{align}
Lemma 9.2 implies that \(\lambda(t,u)\) is differentiable in the \(u\) variable, and if \(\lambda = 0\) then
\begin{align}
\lambda(t,u) = tu + O(t^2\rho^2).
\end{align}
By the classical implicit function theorem, for each \(t\) sufficiently small, there exists \(u(t)\) such that \(\lambda(t,u(t)) = 0\). For \(t\) sufficiently small, \(\text{Ric}_{g_{t,u,v}}\) is strictly negative, and a standard argument (see [Big06]) then implies that \(\hat{g}_{t,u,v}\) is Einstein. \(\Box\)
Proof of Theorem 1.3. First assume that $g_0$ is non-degenerate. Observe that the refined approximate metric is close to the naïve approximate metric in the $C^{2,\alpha}_{\delta_0,\delta_\infty}$ weighted norm, for $t$ sufficiently small. As above, from the results in [Biq13, Section 8], for any metric $g_E$ satisfying (1.16), then there exists a diffeomorphism $\varphi : M^t \to M^t$ so that $B_{g_t}(g_t - \varphi^* g_E) = 0$, and $\varphi^* g_E$ remains close to $g_t$ in the $C^{2,\alpha}_{\delta_0,\delta_\infty}$ weighted norm. The metric $\varphi^* g_E$ would then give a zero of the mapping $\Phi_t$, but then (5.20) would then imply that $\lambda = 0$, a contradiction.

If $g_0$ is not non-degenerate, then if $\delta_\infty$ is chosen sufficiently small, the linearized operator is Fredholm (see [Lee06]), so there is a finite-dimensional space of $L^2$ infinitesimal Einstein deformations on $(M_0, g_0)$. The remainder of the proof is almost identical to the proof of Theorem 1.1 given above.  

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