Model independent extraction of the proton magnetic radius from electron scattering

ZACHARY EPSTEIN\textsuperscript{(a)}, GIL PAZ\textsuperscript{(b)}, JOYDEEP ROY\textsuperscript{(b)}

\textsuperscript{(a)} Department of Physics, University of Maryland, College Park, MD 20742, USA
\textsuperscript{(b)} Department of Physics and Astronomy
Wayne State University, Detroit, Michigan 48201, USA

Abstract

We combine constraints from analyticity with experimental electron-proton scattering data to determine the proton magnetic radius without model-dependent assumptions on the shape of the form factor. We also study the impact of including electron-neutron scattering data, and $\pi\pi \rightarrow N\bar{N}$ data. Using representative datasets we find for a cut of $Q^2 \leq 0.5$ GeV$^2$, $r_M^p = 0.91^{+0.03}_{-0.06} \pm 0.02$ fm using just proton scattering data; $r_M^n = 0.87^{+0.04}_{-0.05} \pm 0.01$ fm adding neutron data; and $r_M^{\pi\pi} = 0.87^{+0.02}_{-0.02}$ fm adding $\pi\pi$ data. We also extract the neutron magnetic radius from these data sets obtaining $r_M^n = 0.89^{+0.03}_{-0.03}$ fm from the combined proton, neutron, and $\pi\pi$ data.
1 Introduction

The first indication of the composite nature of the proton was the measurement of the magnetic moment of the proton by Frisch and Stern in 1933 \[1\]. As described by Otto Stern in his Nobel prize lecture, “The result of our measurement was very interesting. The magnetic moment of the proton turned out to be about 2.5 times larger than the theory predicted. Since the proton is a fundamental particle - all nuclei are built up from protons and neutrons - this result is of great importance. Up to now the theory is not able to explain the result quantitatively.” \[2\]. This statement is to some extent still true today. The response of the proton to electromagnetic field is described by two form factors, one “electric” \(G_E\) and one “magnetic” \(G_M\). The magnetic moment of the proton is just the value of \(G_M\) at zero 4-momentum transfer squared. Viewed as a Taylor series, the magnetic moment is the first in an infinite list of numbers needed to describe the response of the proton to a magnetic field. The next number would be the slope of the magnetic form factor at zero, which is related to the magnetic radius of the proton. For the electric form factor the value at zero is the total charge of the proton in units of \(e\), and the slope at zero defines the charge radius of the proton. The electric and magnetic radii of the proton are therefore as fundamental as the charge and magnetic moment of the proton. Currently, we cannot determine them accurately from theory, although lattice QCD is making progress on this issue, see for example \[3\]. We can measure them from experiment.

The determination of the charge radius of the proton has received considerable attention in the last few years as a result of the discrepancy between the extraction of the charge radius of the proton from muonic and regular hydrogen. The measurement reported by the CREMA collaboration in \[4\] has found \(r_E^p = 0.84184(67)\) fm, and more recently \[5\] \(r_E^p = 0.84087(39)\) fm. Both of these muonic hydrogen extractions are in conflict with the CODATA 2010 \[6\] value \(r_E^p = 0.87580(770)\) fm, based on only hydrogen and deuterium spectroscopic data. This discrepancy is often referred to as the “proton radius puzzle”.

The discrepancy has generated considerable debate. The discussion has focused on the one hand on recalculation of the theoretical input to the extraction of \(r_E^p\) from muonic hydrogen and on modifications of the theoretical calculation such as proton structure effects, e.g. \[7, 8, 9, 10, 11, 12, 13, 14, 15, 16, 17, 18, 19, 20, 21, 22, 23, 24, 25, 26, 27, 28, 29, 30, 31, 32, 33, 34, 35, 36, 37, 38, 39, 40, 41, 42, 43, 44, 45, 46, 47, 48, 49, 50, 51, 52, 53, 54, 55, 56, 57\], and on effects of new physics, e.g. \[58, 59, 60, 61, 62, 63, 64, 65, 66, 67, 68, 69, 70, 71, 72, 73, 74, 75\] on the other.

Apart from regular and muonic hydrogen, electron proton scattering data also allows to measure the charge radius of the proton. Many such extractions exist in the literature, using different data sets and functional forms. The main problem in robust extraction of the proton charge radius from the data is the need to reliably extrapolate the form factor to \(q^2 = 0\) in order to find its slope. Many of the existing extractions postulate a functional form for the form factor either explicitly, or implicitly by truncating a possibly general series expansion. Thus all of these extractions introduce model-dependence for the value of \(r_E^p\) which is very hard to assess.

The problem was solved in \[76\], which introduced a method of extraction that is free of such model dependence. The method, often called the “\(z\)-expansion” adapts an established tool in the study of meson form factors to the case of baryon form factors. The \(z\) expansion
relies on the known analytic properties of the electromagnetic form factors $G_E$ and $G_M$. They are analytic in the complex plane outside of a cut along the positive real $q^2$ axis that starts at $4m^2_\pi$ and extends to infinity. The location of the singularity also implies that the radius of convergence, if using a simple Taylor expansion for the form factors, is at most $4m^2_\pi$. Most of the data about the form factors is well above this value. But even if we use data that is strictly below it, it is questionable whether we can ignore higher terms in the Taylor expansion as it is often assumed. The $z$-expansion avoids this difficulty. By using the variable $z$ defined as

$$z(t, t_{\text{cut}}, t_0) = \frac{\sqrt{t_{\text{cut}} - t} - \sqrt{t_{\text{cut}} - t_0}}{\sqrt{t_{\text{cut}} - t} + \sqrt{t_{\text{cut}} - t_0}}$$

(1)

we can map the domain of the analyticity of the form factors onto the unit circle, see Figure 1.

For $G_E$ and $G_M$, $t_{\text{cut}} = 4m^2_\pi$. The free parameter $t_0$ determines the location of $z = 0$. Considered as a function of $z$, the form factor is analytic inside the unit circle and can be expressed as

$$G_{E,M}(q^2) = \sum_{k=0}^{\infty} a_k z(q^2)^k.$$

(2)

Intuitively, $z$ is the “right” variable in which to perform a Taylor expansion of the form factor. Unlike a Taylor expansion in $q^2$, the expansion is guaranteed to converge for $|z| < 1$. Since for finite negative $q^2$, $z$ is smaller than 1, this guarantees convergence for any $q^2$ measured in experiment. As an illustration to this intuitive picture, consider the proton magnetic form factor data tabulated in [77] and the neutron magnetic form factor data tabulated in [78, 79, 80, 81, 82, 83, 84]. Plotting the data points as a function of $Q^2 = -q^2$ for $0 < Q^2 \leq 1$ GeV$^2$, we see a considerable curvature, see Figure 2. If we plot the same data as a function of $z$ (using $t_{\text{cut}} = 4m^2_\pi$ and $t_0 = 0$) the data looks fairly linear. We can also easily estimate the slopes of the proton and neutron magnetic form factors. If we plot the normalized values of the form factors, i.e. the form factor values divided by their value at $q^2 = 0$ as a function of $z$, the slopes would be hard to distinguish. This implies that the magnetic radii of the proton and neutron are very similar. We will see later that this is indeed the case.

The magnetic radius of the proton is defined as $r_M^p \equiv \sqrt{\langle r^2 \rangle_M^p}$, where

$$\langle r^2 \rangle_M^p = \frac{6}{G_M^p(0)} \frac{d}{dq^2} G_M^p(q^2) \bigg|_{q^2=0}.$$ 

(3)

In 2010 the A1 collaboration reported a value of $r_M^p = 0.777(13)_{\text{stat.}}(9)_{\text{syst.}}(5)_{\text{model}}(2)_{\text{group}}$ fm [85]. This value is considerably lower than $r_M^p = 0.876 \pm 0.010 \pm 0.016$ fm extracted in
Figure 2: Proton (above the horizontal axes) and neutron (below the horizontal axes) magnetic form factor data as a function of $Q^2$ (left) and as a function of $z$ (right). Here we choose $t_0 = 0$ and use $t_{\text{cut}} = 4m_N^2$ in the definition of $z$, and plot data for $0 \leq Q^2 \leq 1.0 \text{ GeV}^2$.

[86] or $0.854 \pm 0.005$ fm extracted in [87], the two other extractions cited by the Particle Data Group (PDG) [88]. Are we facing also a magnetic radius puzzle?1

The purpose of this study is to apply the methods established in [76], to the extraction of the magnetic radius of the proton from scattering data. As in [76] we will use proton, neutron, and $\pi\pi$ scattering data to determine the magnetic radius of the proton from the reported measurement of the magnetic form factors of the proton and neutron. We will also determine the magnetic radius of the neutron.

The rest of the paper is organized as follows. In section 2 we discuss the analytic structure of the form factors and their constraints. In section 3 we extract the magnetic radius of the proton from proton, neutron, and $\pi\pi$ scattering data. In section 4 we extract the magnetic radius of the neutron from the same data. We present our conclusions in section 5.

2 Form factor constraints

The analytic structure of the form factors and their constraints were discussed in detail in [76]. Here we review some of the main ingredients needed for our analysis.

2.1 Form factor definitions

The Dirac and Pauli form factors, $F_1^N$ and $F_2^N$, respectively, are defined by [89, 90]

$$\langle N(p')|J^e_m|N(p)\rangle = \bar{u}(p') \left[ \gamma_\mu F_1^N(q^2) + \frac{i\sigma_{\mu\nu}}{2m_N} F_2^N(q^2)q^\nu \right] u(p),$$

(4)

[1]See the conclusions for values of $r_M^p$ not quoted by the PDG.
where \( q^2 = (p' - p)^2 = t \) and \( N \) stands for \( p \) or \( n \). The Sachs electric and magnetic form factors are related to the Dirac-Pauli basis by \( [91] \)

\[
G^N_E(t) = F_1^N(t) + \frac{t}{4m_N^2} F_2^N(t), \quad G^N_M(t) = F_1^N(t) + F_2^N(t).
\]

(5)

At \( t = 0 \) they are \( [88] \) \( G^p_E(0) = 1, G^p_M(0) = 0, G^p_M(0) = \mu_p = 2.793, G^p_M(0) = \mu_n = -1.913 \).

We define the isoscalar and isovector form factors as

\[
G^{(0)}_{M,E} = G^p_{M,E} + G^n_{M,E}, \quad G^{(1)}_{M,E} = G^p_{M,E} - G^n_{M,E},
\]

(6)

such that at \( t = 0 \) they are, \( G^{(0)}_{E}(0) = 1, G^{(1)}_{E}(0) = 1, G^{(0)}_{M}(0) = \mu_p + \mu_n, G^{(1)}_{M}(0) = \mu_p - \mu_n \).

Notice that \( G^{(0)}_{M,E} = 2G^{s,v}_{M,E}, G^{(1)}_{M,E} = 2G^{s,v}_{M,E} \) for \( G^{s,v}_{M,E} \) of \( [87] \).

2.2 Analytic structure

The electric and magnetic form factors are analytic functions of \( t \) outside of a cut that starts at the two-pion threshold \( t \geq 4m_{\pi}^2 \) on the real \( t \) axis. The scattering data lies on \( -Q_{\text{max}}^2 \leq t \leq 0 \), where \( Q_{\text{max}}^2 \) denotes the largest value of \( Q^2 \) in a given data set. The domain of analyticity can be mapped onto the unit disk via the conformal transformation \( \{1\} \). The mapping is shown in Figure 1. The maximal value of \( |z| \) depends on \( Q_{\text{max}}^2 \) and \( t_0 \). It is minimized for the choice \( t_{\text{opt}}^0 = t_{\text{cut}} \left(1 - \sqrt{1 + Q_{\text{max}}^2 / t_{\text{cut}}} \right) \) which is also the value used for Figure 1.

Since the values of the form factors at \( q^2 = 0 \) are well known, in the following we will use \( t_0 = 0 \). As discussed in \([76]\), the results do not depend on the choice of \( t_0 \). For this choice of \( t_0 \), the maximum value of \( |z| \) is 0.46, 0.58 for \( Q_{\text{max}}^2 = 0.5, 1.0 \) GeV\(^2\), respectively. The form factors can be expanded in a power series in \( z(q^2) \)

\[
G(q^2) = \sum_{k=0}^{\infty} a_k z(q^2)^k,
\]

(7)

where higher order terms are suppressed by powers of the maximum values of \( |z| \). The coefficients \( a_k \) are also bounded in size guaranteeing that the series converges.

The analytic structure implies the dispersion relation,

\[
G(t) = \frac{1}{\pi} \int_{t_{\text{cut}}}^{\infty} dt' \frac{\text{Im}G(t' + i0)}{t' - t}.
\]

(8)

Information about \( \text{Im}G \) over the cut can be translated into information about \( a_k \). As shown in \([76]\), we have

\[
a_0 = \frac{1}{\pi} \int_0^{\pi} d\theta \text{Re}G[t(\theta) + i0] = G(t_0),
\]

\[
a_k = -\frac{2}{\pi} \int_0^{\pi} d\theta \text{Im}G[t(\theta) + i0] \sin(k\theta) = \frac{2}{\pi} \int_{t_{\text{cut}}}^{\infty} dt \frac{\sqrt{t_{\text{cut}} - t_0}}{t - t_{\text{cut}}} \text{Im}G(t) \sin[k\theta(t)], \quad k \geq 1,
\]

(9)

where

\[
t = t_0 + \frac{2(t_{\text{cut}} - t_0)}{1 - \cos \theta} \equiv t(\theta).
\]

(10)
2.3 Bounds on the coefficients

In order to obtain a reliable and conservative extraction of the proton magnetic radius we need to establish appropriate bounds on the coefficients \( a_k \). In particular, it was shown in [76] that the bounds of \( |a_k| < 5 \) and \( |a_k| < 10 \) are very conservative for the electric form factor. We would like to determine similar bounds for the magnetic form factor.

2.3.1 Vector dominance ansatz

The first approach we use to estimate the size of \( a_k \) is the vector dominance ansatz, where the form factors are assumed to be dominated by vector meson exchange: \( \omega \) for \( I = 0 \), and \( \rho \) for \( I = 1 \) [92]. In particular, the imaginary part of the form factor is given by [93]

\[
\text{Im} G(t + i0) = \frac{N m_V^3 \Gamma_V}{(t - m_V^2)^2 + \Gamma_V^2 m_V^2} \theta(t - t_{\text{cut}}),
\]

where \( m_V \) and \( \Gamma_V \) are the mass and width of the vector meson and \( N \) is a normalization constant determined below. Also, \( t_{\text{cut}} = 9m_\pi^2 \) for \( I = 0 \) and \( t_{\text{cut}} = 4m_\pi^2 \) for \( I = 1 \). Using the dispersion relation [8] with (11) we find [94],

\[
G(t + i0) = \frac{N m_V^3 \Gamma_V}{\pi |b(t)|^2} \left[ \frac{1}{2} \log \left( \frac{|b(t_{\text{cut}})|^2}{|t_{\text{cut}} - t|^2} \right) + \frac{m_V^2 - t}{m_V \Gamma_V} \arg[b(t_{\text{cut}})] + i\pi \theta(t - t_{\text{cut}}) \right],
\]

where \( b(t) = t - m_V^2 + i\Gamma_V m_V \), and \( N \) is determined by the value of \( G(0) \).

This form allows us to calculate \( a_k \) explicitly from (9). Using \( m_\rho = 0.775 \text{ GeV} \), \( \Gamma_\rho = 0.149 \text{ GeV} \), \( m_\omega = 0.783 \text{ GeV} \) and \( \Gamma_\omega = 0.0085 \text{ GeV} \) [88], we have for \( I = 0 \): \( a_0 \approx 0.88 \), \( a_1 \approx 1.0 \), \( a_2 \approx 0.83 \), \( a_3 \approx -0.29 \), \( a_4 \approx -1.1 \). For \( I = 1 \), we have \( a_0 \approx 4.7 \), \( a_1 \approx 3.7 \), \( a_2 \approx 2.7 \), \( a_3 \approx 2.0 \), \( a_4 \approx -0.36 \). Also, using \( |\sin(k\theta)| \leq 1 \) allows us to obtain a \( k \)-independent bound on \( a_k \) for \( k \geq 1 \)

\[
\left| \frac{a_k}{a_0} \right| \leq \frac{2|N|}{|G_M(t_0)|} \text{Im} \left( \frac{-m_V^2}{b(t_{\text{cut}}) + \sqrt{(t_{\text{cut}} - t_0)b(t_{\text{cut}})}} \right).
\]

We find that \( |a_k/a_0| \leq 1.3 \) for \( I = 0 \) and \( |a_k/a_0| \leq 1.1 \) for \( I = 1 \). These results are very similar to those of [76]. An important difference from the electric case is that the magnetic form factors at \( q^2 = 0 \) are given by \( G_M^{(0)}(0) \approx 0.88 \) and \( G_M^{(1)}(0) \approx 4.7 \), compared to \( G_E^{(0)}(0) = 1 \). Since the vector dominance ansatz is normalized by the value at \( q^2 = 0 \), the coefficients \( a_k \) are proportional to this value. Thus we find that \( |a_k| \leq 1.1 \) for \( I = 0 \) and \( |a_k| \leq 5.1 \) for \( I = 1 \). We conclude that while \( |a_k| \leq 10 \) is a conservative estimate for this ansatz, a bound of 5, namely \( |a_k| \leq 5 \) is not conservative enough.

2.3.2 Explicit \( \pi\pi \) continuum

For the case of the magnetic isovector form factor the singularities that are closest to the cut arise from the two pion continuum. The imaginary part of \( G_M^{(1)}(0) \) close to the cut can
be described by the pion form factor $F_{\pi}(t)$ (normalized to $F_{\pi}(0) = 1$) and $f_{\perp}^{I}(t)$, a partial $\pi\pi \to N\bar{N}$ amplitude [95, 96, 97]:

$$\text{Im} G_{M}^{(1)}(t) = \sqrt{\frac{2}{t}} \left( \frac{t/4 - m_{\pi}^{2}}{t} \right)^{3/2} F_{\pi}(t)^{*} f_{\perp}^{I}(t).$$

(14)

Since $F_{\pi}(t)$ and $f_{\perp}^{I}(t)$ share the same phase [96], we will replace them in (14) by their absolute values [76]. The relation [14] holds only up to the four-pion threshold $t \leq 16 m_{\pi}^{2}$, but in order to estimate the bounds on the coefficients, we will extend (14) through the $\rho$ peak as in [76]. Values of $f_{\perp}^{I}(t)$ are taken from Table 2.4.6.1 of [97]. We interpolate their product with the prefactor in [14]. This interpolated function is multiplied by the values of $|F_{\pi}(t)|$ using the four $t$ values from [98] (0.101 to 0.178 GeV$^2$) and the 43 $t$ values from [99] (0.185 to 0.94 GeV$^2$). This gives us a discrete expression with 47 data points for $\text{Im} G_{M}^{(1)}(t)$ from 0.101GeV$^2$ to 0.94 GeV$^2$. We now use the experimental data up to $t = 0.8$ GeV$^2 \approx 40 m_{\pi}^{2}$ to calculate $a_k$ using (9). We find $a_0 \approx 7.9$, $a_1 \approx -5.5$, $a_2 \approx -6.1$, $a_3 \approx -2.9$, $a_4 \approx 1.1$. Using $|\sin(k \theta)| \leq 1$ gives $|a_k| \lesssim 7.2$ for $k \geq 1$. It is interesting to note that $a_k/a_0$ for these values is very similar to the analogous $a_k/a_0$ obtained for the isovector electric form factor in [76]. This can be traced to the fact that the shape of $f_{\perp}^{I}(t)$ is very similar to $f_{\perp}^{E}(t)$. This indicates that the main difference between the electric and magnetic form factors is their normalization.

### 2.3.3 Bounds on the $t \geq 4 m_{N}^2$

Above the two nucleon threshold one can use $e^{+}e^{-} \to N\bar{N}$ data to constrain the electric and magnetic form factor. In particular, the cross section is given by [100]

$$\sigma(t) = \frac{4 \pi \alpha^2}{3t} \sqrt{1 - \frac{4 m_{N}^2}{t}} \left( |G_{M}(t)|^2 + \frac{2 m_{N}^2}{t} |G_{E}(t)|^2 \right).$$

(15)

The contribution to $a_k$ from this region is given by [9]

$$\delta a_k = \frac{2}{\pi} \int_{4 m_{N}^2}^{\infty} \frac{dt}{t - t_0} \sqrt{\frac{t_{\text{cut}} - t_0}{t - t_{\text{cut}}}} \text{Im} G(t) \sin[k \theta(t)], \quad k \geq 1.$$

(16)

Since $|\text{Im} G(t) \sin[k \theta(t)]| \leq |G_{M}(t)| \leq \sqrt{|G_{M}(t)|^2 + \frac{2 m_{N}^2}{t} |G_{E}(t)|^2}$ we have

$$|\delta a_k| \leq \frac{2}{\pi} \int_{4 m_{N}^2}^{\infty} \frac{dt}{t - t_0} \sqrt{\frac{t_{\text{cut}} - t_0}{t - t_{\text{cut}}}} \sigma(t) \frac{3 t}{4 \pi \alpha^2} \left( 1 - \frac{4 m_{N}^2}{t} \right)^{-1/2}, \quad k \geq 1.$$

(17)

These bounds are valid for both the proton and neutron magnetic form factors. Using the $e^{+}e^{-} \to p\bar{p}$ data from [101], we perform the integral from $t = 4.0$ GeV$^2$ to 9.4 GeV$^2$ as a discrete sum, using the measured values of $\sigma(t)$ plus the 1σ error. We find $|\delta a_k^p| \leq 0.013$. The contribution above 9.4 GeV$^2$ can be conservatively estimated by assuming a constant value for the form factors. This gives a constant value of 0.031 for $\sqrt{|G_{M}(t)|^2 + 2 m_{p}^2 |G_{E}(t)|^2}/t$ above 9.4 GeV$^2$, leading to an additional 0.004. In total we have $|\delta a_k^p| \leq 0.013 + 0.004$. 

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Using the $e^+e^- \rightarrow n\bar{n}$ data from [102] for $t = 3.61$ to 5.95 GeV$^2$, we find $|\delta a_k^p| \leq 0.011$. Assuming a constant value, 0.32, for $\sqrt{|G_M^p(t)|^2 + 2m_n^2|G_E^p(t)|^2/t}$ above 5.95 GeV$^2$, leads to an additional 0.047, giving in total $|\delta a_k^p| \leq 0.011 + 0.047$.

These results are very similar to those of the electric form factor in [76], although for the electric form factor more stringent bounds were obtained. Compared to the bounds calculated above, these contributions are negligible. Our conclusion, as in [76], is that the contribution of the physical timelike region $t \geq 4m_N^2$ can be neglected.

2.3.4 Summary

All our studies point out that for the magnetic form factor the coefficients $a_k$ are smaller than 10. Since $a_0 = G^{(1)}(0) = \mu_p - \mu_n \approx 4.7$, a bound of 5 might be too stringent. In the following we will use a bounds of 10 and 15 instead of the bounds of 5 and 10 used in [76]. As we will see, even using a bound of 20 will not change the results in an appreciable way.

One could also argue that a bound on the ratio $|a_k/a_0| \leq 5,10$ is more appropriate. Since $a_0$ is known, this will translate to a bound of $|a_k| \leq 25,50$ in the $I = 1$ case. We prefer to use the more stringent bound of $|a_k| \leq 10,15$, but we will comment on the results when using these looser bounds.

It should be noted that for $t_0 = 0$, the magnetic radius depends only the coefficient of $z$. Writing $G_M^p(q^2) = \sum_{k=0}^\infty a_k z(q^2)^k$, where $z(q^2) \equiv z(q^2, 4m_\pi^2, 0)$, equation (3) implies that

$$r_M^p = \frac{hc}{2m_\pi c^2} \sqrt{-\frac{3a_1}{2\mu_p}},$$

where we are showing explicitly the factors of $h$ and $c$. A bound of 5, 10, 15, or 20, on $|a_k|$, implies also a bound of 1.2, 1.6, 2.0, and 2.3 fm on $r_M^p$. Writing $G_M^{(0)}(q^2) = \sum_{k=0}^\infty a_k^{(0)} z(q^2, 9m_\pi^2, 0)^k$ and $G_M^{(1)}(q^2) = \sum_{k=0}^\infty a_k^{(1)} z(q^2, 4m_\pi^2, 0)^k$ we have

$$r_M^p = \frac{hc}{2m_\pi c^2} \sqrt{-\frac{a_1^{(0)} + \frac{g}{3}a_1^{(1)}}{3\mu_p}}. \quad (19)$$

A bound of 5, 10, 15, or 20, on $|a_k^{(0,1)}|$, implies also a bound of 0.98, 1.4, 1.7, or 2.0 fm on $r_M^p$. For our default choice of bounds of 10 and 15 these values are much larger than the current range of values quoted by the PDG [88], roughly $0.7 - 0.9$ fm. Thus just the presence of our default bounds does not bias the extraction of the radius.

3 Extraction of the proton magnetic radius

3.1 Proton data

We extract the proton magnetic radius from the values of $G_M^p$, tabulated in [77]. We write the form factor as $G_M^p(q^2) = \sum_{k=0}^\infty a_k z(q^2)^k$, where $z(q^2) \equiv z(q^2, 4m_\pi^2, 0)$. We fit $k < k_{max}$

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parameters, where \( k_{\text{max}} = 2, \ldots, 12 \). We minimize the \( \chi^2 \) function

\[
\chi^2 = \sum_i (\text{data}_i - \text{theory}_i)^2 / (\sigma_i)^2,
\]

(20)

Where \( i \) ranges over the tabulated values of \([77]\) up to a given maximal value of \( Q^2 \), with \( Q^2 = 0.1, 0.2, \ldots, 1.2, 1.4, 1.6, 1.8 \) GeV\(^2\). As explained above, our default choice for the bounds on the coefficients is \( |a_k| < 10 \) and \( |a_k| < 15 \). The proton magnetic radius is obtained from \([3]\). The error bars are determined from the \( \Delta \chi^2 = 1 \) range. Usually, the \( \Delta \chi^2 = 1 \) range was determined from a numerical search algorithm. For some higher values of \( Q^2 \), the \( \chi^2(r^p_M) \) seems to have some discontinuities and in that case, the \( \Delta \chi^2 = 1 \) was extracted directly from \( \chi^2(r^p_M) \) curve. To ensure a conservative estimate of the error, we quote only one digit in the error bar.

The extracted values and the value of the minimum of \( \chi^2 \) do not vary with \( k_{\text{max}} \) for \( k_{\text{max}} > 4 \). In other words, the extracted values do not depend on the number of coefficients we fit. In the following we quote results with \( k_{\text{max}} = 8 \). The extracted values of the magnetic radius are very consistent over the range of \( Q^2 \). Thus for data with \( Q^2 \leq 0.5 \) GeV\(^2\), we have \( r^p_M = 0.91^{+0.03}_{-0.06} \) fm for a bound of 10 and \( r^p_M = 0.92^{+0.04}_{-0.07} \) fm for a bound of 15, while for \( Q^2 \leq 1.0 \) GeV\(^2\) we have \( r^p_M = 0.90^{+0.03}_{-0.07} \) fm for a bound of 10 and \( r^p_M = 0.91^{+0.04}_{-0.07} \) fm for a bound of 15.

We have studied the dependence of the extracted magnetic radius on the bounds on \( |a_k| \). If we use a bound of \( |a_k| < 20 \), the results above change to \( r^p_M = 0.93^{+0.03}_{-0.07} \) fm for \( Q^2 \leq 0.5 \) GeV\(^2\) and \( r^p_M = 0.91^{+0.04}_{-0.08} \) fm for \( Q^2 \leq 1.0 \) GeV\(^2\). These values are very similar to the ones obtained with \( |a_k| < 10 \) and \( |a_k| < 15 \). As discussed above we consider the bound \( |a_k| < 5 \) to be too stringent, but if we do use it we obtain \( r^p_M = 0.89^{+0.03}_{-0.05} \) fm for \( Q^2 \leq 0.5 \) GeV\(^2\) and \( r^p_M = 0.89^{+0.02}_{-0.05} \) fm for \( Q^2 \leq 1.0 \) GeV\(^2\), which are not statistically different from the results of our default bounds.

Another possible choice of bounds might be to bound \( |a_k/a_0| \). This is motivated by the fact that the vector dominance ansatz and the \( \pi-\pi \) data indicate that \( a_k/a_0 \) is similar for the electric and magnetic form factors. Thus we might choose \( |a_k/a_0| < 5, 10 \). We have checked the effect of these looser bounds on the extracted magnetic radius. For \( Q^2 \leq 0.5 \) GeV\(^2\), we have \( r^p_M = 0.92^{+0.03}_{-0.07} \) fm for a bound of \( |a_k/a_0| < 5 \) and \( r^p_M = 0.95^{+0.04}_{-0.08} \) fm for a bound of \( |a_k/a_0| < 10 \) while for \( Q^2 \leq 1.0 \) GeV\(^2\) we have \( r^p_M = 0.91^{+0.04}_{-0.08} \) fm for a bound of \( |a_k/a_0| < 5 \) and \( r^p_M = 0.92^{+0.05}_{-0.09} \) fm for a bound of \( |a_k/a_0| < 10 \). For the magnetic radius with \( t_0 = 0, a_0 = \mu_p \approx 2.8 \), so if we choose \( |a_k/a_0| < 5, 10 \) this translates to \( |a_k| < 14, 28 \) respectively. Comparing these results to the ones obtained above we notice a slight monotonic increase in the central value and the error bar with the loosening of the bound. The increase in the error bars is to be expected of course. Even with the looser bounds, the results we obtain are consistent with our default bounds.

Using our default bounds of \( |a_k| < 10 \) and \( |a_k| < 15 \), and using \( Q^2 \leq 0.5 \) GeV\(^2\) for concreteness we obtain \( r^p_M = 0.91^{+0.03}_{-0.06} \pm 0.02 \) fm. The first error is for a bound of 10 and the second error includes the maximum variation of the \( \Delta \chi^2 = 1 \) interval when we redo the fits with a bound of 15.
3.2 Proton and neutron data

Including neutron data allows us to separate the $I = 1$ and $I = 0$ isospin components of the proton magnetic form factor. Since for the $I = 0$ components $t_{\text{cut}} = 9m_n^2$, this increases the value of $t_{\text{cut}}$ and effectively increases the maximum value of $z$.

As before we use values of $G_M^n$ tabulated in [77]. For $G_M^p(Q^2)$ we use values published in [78, 79, 80, 81, 82, 83, 84]. We do not use the data reported in [106] and [107], as they were criticized for missing a systematic error, see section VIII of [83].

We form the $\chi^2$ as before and express $G_M^n$ and $G_M^p$ in terms of $G_M^{(0)}$ and $G_M^{(1)}$, see (6). We express $G_M^{(0)}$ as a power series in $z(t, 9m_n^2, 0)$ and $G_M^{(1)}$ as a power series in $z(t, 4m_n^2, 0)$, i.e.

$$G_M^{(0)}(t) = \sum_k a_k^{(0)} z^k(t, t_{\text{cut}} = 9m_n^2, 0) \quad (21)$$

$$G_M^{(1)}(t) = \sum_k a_k^{(1)} z^k(t, t_{\text{cut}} = 4m_n^2, 0) \quad (22)$$

As for the proton data alone, the extracted values of the magnetic radius do not depend on the number of the parameters we fit. The values are very consistent over the range of $Q^2$. Thus for data with $Q^2 \leq 0.5$ GeV$^2$, we have $r_M^p = 0.87^{+0.04}_{-0.05}$ fm for a bound of 10 and $r_M^p = 0.87^{+0.05}_{-0.05}$ fm for a bound of 15, while for $Q^2 \leq 1.0$ GeV$^2$ we have $r_M^p = 0.87^{+0.03}_{-0.05}$ fm for a bound of 10 and $r_M^p = 0.88^{+0.04}_{-0.05}$ fm for a bound of 15. These values are consistent with the values extracted from the proton data alone.

We have studied the dependence of the extracted magnetic radius on the bounds on $|a_k|$. If we use a bound of $|a_k| < 20$, the results above change to $r_M^p = 0.88^{+0.04}_{-0.06}$ fm for $Q^2 \leq 0.5$ GeV$^2$ and $r_M^p = 0.88^{+0.05}_{-0.06}$ fm for $Q^2 \leq 1.0$ GeV$^2$. These values are very similar to the ones obtained with $|a_k| < 10$ and $|a_k| < 15$. If we use the bound $|a_k| < 5$, we obtain $r_M^p = 0.87^{+0.02}_{-0.02}$ fm for $Q^2 \leq 0.5$ GeV$^2$ and $r_M^p = 0.87^{+0.02}_{-0.02}$ fm for $Q^2 \leq 1.0$ GeV$^2$. The central values are consistent with our default bounds, but the error bars are substantially smaller. This is to be expected since this bound is too stringent.

As explained above, another possible choice of bounds is $|a_k/a_0| < 5, 10$. For $Q^2 \leq 0.5$ GeV$^2$, we have in this case $r_M^p = 0.88^{+0.05}_{-0.06}$ fm for a bound of $|a_k/a_0| < 5$ and $r_M^p = 0.91^{+0.05}_{-0.07}$ fm for a bound of $|a_k/a_0| < 10$. For $Q^2 \leq 1.0$ GeV$^2$ we have $r_M^p = 0.89^{+0.04}_{-0.07}$ fm for a bound of $|a_k/a_0| < 5$ and $r_M^p = 0.90^{+0.05}_{-0.09}$ fm for a bound of $|a_k/a_0| < 10$. Since $a_0^{(0)} = \mu_p + \mu_n \approx 0.88$, $a_0^{(1)} = \mu_p - \mu_n \approx 4.7$, $|a_0^{(0)}/a_0^{(0)}| < 5$ implies $|a_0^{(0)}| < 4.4$ and $|a_0^{(1)}/a_0^{(0)}| < 5$ implies $|a_0^{(1)}| < 23.5$. Similarly $|a_k^{(0)}/a_0^{(0)}| < 10$ implies $|a_k^{(0)}| < 8.8$ and $|a_k^{(1)}/a_0^{(0)}| < 10$ implies $|a_k^{(1)}| < 47$. Comparing these results to the ones obtained above we notice again a monotonic increase in the central value and the error bar with the loosening of the bound. The increase in the error bars is to be expected of course. Even with the looser bounds, the results we obtain are consistent.

Using our default bounds of $|a_k| < 10$ and $|a_k| < 15$, and using $Q^2 \leq 0.5$ GeV$^2$ for concreteness we obtain $r_M^p = 0.87^{+0.04}_{-0.05} \pm 0.01$ fm.

\[\text{References:}\]
\[\text{103, 104.}\]
\[\text{105.}\]

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\[^{2}\text{See}\] contain the final results that supersedes the previous publications [103, 104]. For [83], the data is tabulated in [105].

\[^{3}\text{If we include these additional data points we obtain similar values of the magnetic radius but with much larger values of } \chi^2.\]
3.3 Proton, neutron, and $\pi\pi$ data

Between the two-pion and four-pion threshold the only state that can contribute to the imaginary part of the magnetic isovector form factor is that of two pions. Since we have information about $\text{Im} G_M^{(1)}(t)$ in this region, see \cite{14}, we can use it to raise the effective threshold for the isovector form factor from $t_{\text{cut}} = 4m_\pi^2$ to $t_{\text{cut}} = 16m_\pi^2$. We do that by fitting \cite{76}

$$G_M^{(1)}(t) = G_{\text{cut}}(t) + \sum_k a_k^{(1)}z^k(t, t_{\text{cut}} = 16m_\pi^2, 0). \quad (23)$$

$G_{\text{cut}}(t)$ is calculated using \cite{8} from the discrete expression for $\text{Im} G_M^{(1)}(t)$ described in section 2.3.2. As in \cite{76} we consider two cases for $G_{\text{cut}}(t)$. The first is generated by the values of $\text{Im} G_M^{(1)}(t)$ in the range $4m_\pi^2 < t < 16m_\pi^2$, and the second by the values of $\text{Im} G_M^{(1)}(t)$ in the range $4m_\pi^2 < t < 40m_\pi^2$. The second choice amounts to modeling the $\pi\pi$ continuum $16m_\pi^2 < t < 40m_\pi^2$ by $\text{Im} G_M^{(1)}(t)$ of \cite{14}. As explained in \cite{76}, this does not introduce model dependence since the difference between the true continuum and $G_{\text{cut}}(t)$ will be accounted for by the parameters in the $z$ expansion, as we do not change the value of $t_{\text{cut}} = 16m_\pi^2$.

In \cite{76} it was found that the second choice of $G_{\text{cut}}(t)$ led to a smaller size of the coefficients in the $z$ expansion of the isovector form factor. We would like to check if that holds true also in the magnetic case. We fit the same proton and neutron data for $Q^2_{\text{max}} = 1$ GeV$^2$, $t_0 = 0$, $k_{\text{max}} = 8$ and a bound of 15 on the coefficients using (23). For the first choice of $G_{\text{cut}}(t)$ we find the first two coefficients of the isoscalar form factor to be $-2_{-0.3}^{+0.5}, 3_{-2}^{+2}$ and the first two coefficients of the vector form factor to be $-13.5(3), 13_{-3}^{+6}$ (The value of $13_{-3}^{+6}$ was obtained by applying a bound of 15 on all the coefficients with the exception of the second one, which is left unbounded). For the second choice of $G_{\text{cut}}(t)$ we find the first two coefficients of the isoscalar form factor are not changed while the first two coefficients of the vector form factor are $2.6_{-0.5}^{+0.4}, 5_{-4}^{+5}$. As in the electric form factor case, we have a reduction in the size of the isovector coefficients when using the second form. We will therefore adopt that as our default choice. As we will see below, the value of the magnetic radius does not change if we use the first form of $G_{\text{cut}}(t)$.

We can understand the large size of the isovector coefficients when using $G_{\text{cut}}(t)$ calculated from $\text{Im} G_M^{(1)}(t)$ in the range $4m_\pi^2 < t < 16m_\pi^2$. From equations (6) and (23), the proton magnetic radius is given by

$$r_M^p = \frac{\hbar c}{2m_ec^2} \sqrt{\frac{1}{\mu_p} \left(-\frac{1}{3}a_1^{(0)} - \frac{3}{16}a_1^{(1)} + 4m_\pi^2c^4G'_{\text{cut}}(0)\right)}, \quad (24)$$

where $G'_{\text{cut}}(0)$ is obtained from \cite{8},

$$G'_{\text{cut}}(0) = \frac{1}{\pi} \int_{4m_\pi^2} dt' \frac{\text{Im} G(t' + i0)}{(t')^2}. \quad (25)$$

Since $\text{Im} G_M^{(1)}(t)$ from \cite{14} is positive in the relevant region, as we increase the upper limit in (25), $G'_{\text{cut}}(0)$ increases. Therefore $G'_{\text{cut}}(0)$ calculated from $\text{Im} G_M^{(1)}(t)$ in the range $4m_\pi^2 < t < 16m_\pi^2$ is larger than $G'_{\text{cut}}(0)$ obtained from the nucleon data. As a result, we see that $r_M^p$ is larger for the second choice of $G_{\text{cut}}(t)$ than for the first choice. This is consistent with the fact that the isovector form factor is smaller than the isoscalar form factor in the magnetic case. As a result, the magnetic radius is smaller for the second choice of $G_{\text{cut}}(t)$ than for the first choice.

We can also understand the difference in the size of the isovector coefficients when using the first and second choices of $G_{\text{cut}}(t)$ by examining the behavior of $G'_{\text{cut}}(0)$ as a function of the upper limit in (25). As the upper limit increases, $G'_{\text{cut}}(0)$ increases, which leads to a reduction in the size of the isovector coefficients. This is consistent with the fact that the isovector form factor is smaller than the isoscalar form factor in the magnetic case. As a result, the magnetic radius is smaller for the second choice of $G_{\text{cut}}(t)$ than for the first choice.
16m_{\pi}^2$ is smaller than $G'_{\text{cut}}(0)$ calculated from Im $G^{(1)}_M(t)$ in the range $4m_{\pi}^2 < t < 40m_{\pi}^2$ and as a result $|a_{1}^{(1)}|$ must be larger to maintain the same size of $r^p_M$ preferred by the data. In fact, since we can calculate $G'_{\text{cut}}(0)$, if we assume $r^p_M \approx 0.87$ fm and use $a_{1}^{(0)} \approx -2$, we can calculate and find $a_{1}^{(1)} \approx -13$ in the first case and $a_{1}^{(1)} \approx 3$ in the second case. These are the values we obtained above.

Using (23) we extract the magnetic radius. The extracted values of the magnetic radius do not depend on the number of the parameters we fit. The values are very consistent over the range of $Q^2$. Thus for data with $Q^2 \leq 0.5$ GeV$^2$, we have $r^p_M = 0.871^{+0.011}_{-0.015}$ fm for a bound of 10 and $r^p_M = 0.873^{+0.012}_{-0.016}$ fm for a bound of 15, while for $Q^2 \leq 1.0$ GeV$^2$ we have $r^p_M = 0.874^{+0.008}_{-0.015}$ fm for a bound of 10 and $r^p_M = 0.874^{+0.012}_{-0.014}$ fm for a bound of 15. These values are consistent with the values extracted above.

We have studied the dependence of the radius on the bounds on the coefficients. If we use a bound of 20, we have $r^p_M = 0.876^{+0.012}_{-0.018}$ for $Q^2 \leq 0.5$ GeV$^2$ and $r^p_M = 0.875^{+0.013}_{-0.016}$ for $Q^2 \leq 1.0$ GeV$^2$. These values are very similar to the ones we obtain with a bound of 10 and 15. If we use the too-stringent bound of 5 we obtain $r^p_M = 0.867^{+0.010}_{-0.013}$ for $Q^2 \leq 0.5$ GeV$^2$ and $r^p_M = 0.867^{+0.006}_{-0.008}$ for $Q^2 \leq 1.0$ GeV$^2$. These values are consistent, but the error bars are smaller.

Another possible choice of bounds is $|a_k/a_0| < 5, 10$. For $Q^2 \leq 0.5$ GeV$^2$, we find $r^p_M = 0.867^{+0.013}_{-0.015}$ fm for a bound of $|a_k/a_0| < 5$ and $r^p_M = 0.869^{+0.013}_{-0.015}$ fm for a bound of $|a_k/a_0| < 10$. For $Q^2 \leq 1.0$ GeV$^2$, we find $r^p_M = 0.867^{+0.008}_{-0.009}$ fm for a bound of $|a_k/a_0| < 5$ and $r^p_M = 0.873^{+0.009}_{-0.014}$ fm for a bound of $|a_k/a_0| < 10$. All these results are consistent with our default choices.

The decrease in the error bar when including the $\pi\pi$ data arises from the increase in the value of $t_{\text{cut}}$ from $4m_{\pi}^2$ to $16m_{\pi}^2$ for the isovector form factor. If we use (23) but with $t_{\text{cut}} = 4m_{\pi}^2$ we obtain results that are almost identical to the fits using the proton and neutron data alone. As another check of our results, we fit the data using (23), but with $G_{\text{cut}}(t)$ calculated using Im $G^{(1)}_M(t)$ in the range $4m_{\pi}^2 < t < 16m_{\pi}^2$. As discussed above, we use only a bound of 15 in this case. For $Q^2 \leq 0.5$ GeV$^2$ we find $r^p_M = 0.873^{+0.011}_{-0.010}$, and for $Q^2 \leq 1.0$ GeV$^2$ we find $r^p_M = 0.873^{+0.012}_{-0.012}$. These values are very close to the ones we obtained with the use of the default form of $G_{\text{cut}}(t)$.

The expression for Im $G^{(1)}_M(t)$ depends on $f^1(t)$. The tabulation of $f^1(t)$ in [97] does not quote any error. In [26] an error of 30% was used as a representative uncertainty. If we assume a 30% increase for $f^1(t)$ and hence for $G_{\text{cut}}(t)$ we obtain for $Q^2 \leq 0.5$ GeV$^2$ and a bound of 10, $r^p_M = 0.872^{+0.013}_{-0.015}$. If we assume a 30% decrease for $G_{\text{cut}}(t)$ we obtain for $Q^2 \leq 0.5$ GeV$^2$ and a bound of 10, $r^p_M = 0.867^{+0.010}_{-0.015}$.

In summary, all our checks produce consistent results for $r^p_M$. Using our default choices for the bounds and $G_{\text{cut}}(t)$, and using $Q^2 \leq 0.5$ GeV$^2$ for concreteness we obtain $r^p_M = 0.87^{+0.02}_{-0.02}$ fm. Our conservative error estimate includes the variation of the bounds and of $G_{\text{cut}}(t)$ where we choose to quote only one digit in our error estimate.
4 Extraction of the neutron magnetic radius

The data we have used to extract the magnetic radius of the proton can be used also to extract the magnetic radius of the neutron. The magnetic radius of the neutron is defined as \( r_M^n \equiv \sqrt{\langle r^2 \rangle_M} \), where

\[
\langle r^2 \rangle_M^n = \left. \frac{6}{G_M^n(0)} \frac{d}{dq^2} G_M^n(q^2) \right|_{q^2=0}.
\] (26)

We extract the neutron magnetic radius from the neutron, neutron and proton, and neutron, proton, and \( \pi\pi \) data sets. We follow the same default choices described above. In particular we will use a bound of 10 and 15 on the coefficients of the \( z \) expansion.

4.1 Neutron data

Using the neutron form factor data reported in [78, 79, 80, 81, 82, 83, 84] we fit \( G_M^n(q^2) = \sum_{k=0}^\infty a_k z(q^2)^k \) by minimizing the \( \chi^2 \) function of (20). For a cut \( Q^2 \leq 0.5 \text{ GeV}^2 \) we find \( r_M^n = 0.74^{+0.13}_{-0.06} \) fm for a bound of 10 and \( r_M^n = 0.65^{+0.21}_{-0.07} \) fm for a bound of 15. For a cut of \( Q^2 \leq 1.0 \text{ GeV}^2 \) we find \( r_M^n = 0.77^{+0.17}_{-0.09} \) fm for a bound of 10 and \( r_M^n = 0.74^{+0.20}_{-0.11} \) fm for a bound of 15. Obviously the errors bar for \( r_M^n \) extracted from the neutron data are much large than for \( r_M^p \). We prefer to quote only one digit in our error bar. We therefore determine \( r_M^n = 0.7^{+0.2}_{-0.1} \) fm from neutron data alone. Comparing to \( r_M^p = 0.91^{+0.03}_{-0.02} \) fm obtained from proton data alone, we find that \( r_M^n \) and \( r_M^p \) are consistent within errors.

4.2 Neutron and proton data

Adding the proton form factor data from [77] allows us to separate the isospin components. The magnetic radius of the neutron is given by an equation similar to (19)

\[
r_M^n = \frac{\hbar c}{2m_\pi c^2} \sqrt{\frac{-a_1^{(0)} + \frac{9}{4}a_1^{(1)}}{3\mu_n}}.
\] (27)

We fit the isoscalar and the isovector form factors as described before. For a cut \( Q^2 \leq 0.5 \text{ GeV}^2 \) we find \( r_M^n = 0.89^{+0.06}_{-0.09} \) fm for a bound of 10 and \( r_M^n = 0.88^{+0.08}_{-0.06} \) fm for a bound of 15. For a cut of \( Q^2 \leq 1.0 \text{ GeV}^2 \) we find \( r_M^n = 0.88^{+0.06}_{-0.08} \) fm for a bound of 10 and \( r_M^n = 0.89^{+0.07}_{-0.10} \) fm for a bound of 15. Again the errors bar for \( r_M^n \) are about twice as large as those for \( r_M^p \) from the same data set. Quoting only one digit we determine \( r_M^n = 0.9^{+0.1}_{-0.1} \) fm from neutron and proton data. Comparing to \( r_M^p = 0.87^{+0.03}_{-0.05} \) ± 0.02 fm obtained from the same proton and neutron data, we find that \( r_M^n \) and \( r_M^p \) are consistent within errors.

4.3 Neutron, proton, and \( \pi\pi \) data

Adding the \( \pi\pi \) data as described in the previous section leads to a reduction in the error bars. For a cut \( Q^2 \leq 0.5 \text{ GeV}^2 \) we find \( r_M^n = 0.89^{+0.03}_{-0.06} \) fm for a bound of 10 and \( r_M^n = 0.89^{+0.03}_{-0.03} \) fm.
for a bound of 15. If we take a 30\% variation of $f_1(t)$ as described above, we get values of $r_M^n$ within this range. For a cut of $Q^2 \leq 1.0$ GeV$^2$ we find $r_M^n = 0.88^{+0.03}_{-0.01}$ fm for a bound of 10 and $r_M^n = 0.88^{+0.03}_{-0.02}$ fm for a bound of 15. As before the errors bar for $r_M^n$ are about twice as large as those for $r_M^p$ from the same data set. Quoting only one digit for the error bar we determine $r_M^n = 0.89^{+0.03}_{-0.03}$ fm from neutron, proton, and $\pi\pi$ data. Comparing to $r_M^p = 0.87^{+0.02}_{-0.02}$ fm obtained from the same data set, we find that $r_M^n$ and $r_M^p$ are consistent within errors.

5 Conclusions

The recent large discrepancy in the extraction of the charge radius of the proton from spectroscopic measurements of regular and muonic hydrogen has motivated the reexamination of the extraction of nucleon radii from scattering data. Since the first measurement of the “size” of the proton \cite{108} almost 60 years ago, there have been many extractions of the charge radius of the proton. These were based on different data sets and postulated different functional forms for the form factors. These various extractions do not agree with each other. Even when using the same data sets, different functional forms can lead to different values of the charge radius of the proton.

A fundamental problem of many of these extractions is that they do not take into account the known analytic structure of the form factors. Therefore, it is unlikely that an arbitrary functional form will be consistent with this structure. This analytic structure constrains the form factors but does not determine it completely. Since the form factors are non-perturbative functions, one would like to incorporate the analytic structure while maintaining the flexibility of the functional form. The so-called “z-expansion” described in the introduction achieves both of these goals. It automatically incorporates the analytic structure and allows for flexible functional forms. It is therefore not surprising that the z-expansion has become a standard tool in analyzing meson form factors, see for example section 8.3.1 of \cite{109}.

To the best of our knowledge the first application of the z-expansion to baryon form factor was done in \cite{76}. That paper also has shown the need to impose some constraints on the coefficients of the z-expansion in order to have a result that is independent of the number of parameters. For meson form factors such as $B \to \pi$, constraints that bound the sum of the squares of the coefficients can be obtained from unitarity\footnote{See \cite{110} for a discussion of these unitarity bounds.}. For the nucleon form factors such constraints are less useful since there is a large distance between the two-pion threshold where the singularity begins, and the two-nucleon threshold where the unitarity bounds can be applied. The studies of \cite{76} have shown that a uniform bound on the coefficients can be applied. The methods of \cite{76} were later used in \cite{94} for a model-independent extraction of the axial mass parameter of the nucleon from neutrino-nucleon scattering data.

We have applied the same methods in this paper to extract the magnetic radius of the proton from scattering data in a model independent way. While not as severe as the proton charge radius problem, various extractions in recent years, e.g the ones cited by the PDG \cite{88}, are not consistent with each other. The goal of our study was to try and resolve these discrepancies.
We first studied the bounds on the coefficients of the z-expansion. In [76] bound of 5 and 10 were used. Since the value of the isovector magnetic form factor at zero momentum transfer is about 4.7, a bound of 5 on the coefficients might be too stringent. Our studies have shown that this is indeed the case, but bounds of 10 and 15 are conservative enough for the coefficients of the magnetic form factor. An alternative option is to use a bound of 5 and 10 on the ratio \( |a_k/a_0| \). Fitting the data using each of these prescriptions gives consistent results. Our default choice is to use the bound of 10 and 15.

We have extracted the magnetic radius of the proton from three data sets. The first contains values of proton magnetic form factor data tabulated in [77]. The second contains the proton data and the neutron magnetic form factor data tabulated in [78, 79, 80, 81, 82, 83, 84]. The third contains the proton and neutron data and the two-pion continuum data constructed from pion form factor data and a \( \pi\pi \rightarrow N\bar{N} \) partial amplitude using (14). In all the cases we use the listed data and do not apply any corrections. For each data set the extracted values quoted by the PDG are larger than these extraction, which is not unusual when using model-independent methods. For the first two values the first error is for a bound of 10 and the second error includes the maximum variation of the \( \Delta x^2 = 1 \) interval. The error on the third value combines both, as well as errors on the continuum contribution as discussed in section 3. In all cases we choose to quote one digit in our error bar. As expected the error decrease as we include more data, but the main effect is the change in the value of \( t_{\text{cut}} \). Using proton data alone we have \( t_{\text{cut}} = 4m_z^2 \). Adding the neutron data allows to set \( t_{\text{cut}} = 9m_z^2 \) for the isoscalar magnetic form factor. Adding the two-pion continuum allows to set \( t_{\text{cut}} = 16m_z^2 \) for the isovector magnetic form factor. The increase in \( t_{\text{cut}} \) leads to a decrease in the maximum value of \( |z| \) and a therefore for a smaller error.

Comparing our third value of the magnetic radius of the proton, \( r_M^p = 0.87 +0.02 \) fm, to the values quoted by the PDG [85], we find that they are more consistent with \( r_M^p = 0.876 \pm 0.010 \pm 0.016 \) fm extracted in [86] and \( r_M^p = 0.854 \pm 0.005 \) fm extracted in [87], than the A1 collaboration value of \( r_M^p = 0.777(13)_{\text{stat}}(9)_{\text{syst}}(5)_{\text{model}}(2)_{\text{group}} \) fm [88]. Our error bars are larger than these extraction, which is not unusual when using model-independent methods [76, 94]. Other extractions of the proton magnetic radius from scattering data that were not quoted by the PDG are \( r_M^p = 0.855 \pm 0.035 \) fm [94], \( r_M^p = 0.867 \pm 0.020 \) fm [97], and \( r_M^p = 0.86 +0.02 -0.03 \) fm [113]. Our results are consistent with these values too.

The same data can be used also for a model independent extraction of the neutron magnetic radius. Taking \( Q^2 \leq 0.5 \) GeV\(^2\) and fits with eight parameters for concreteness we find that for the neutron data set \( r_M^n = 0.7 +0.1 \) fm, for the proton and neutron data set \( r_M^n = 0.9 +0.1 \) fm, and for the proton, neutron, and \( \pi\pi \) data set \( r_M^n = 0.89 +0.03 \) fm. The last value can be compared to the value quoted by the PDG. \( r_M^n = 0.862^{+0.009}_{-0.008} \) fm [87]. Our results are consistent but our error bars are larger which can again be attributed to the use of model-independent methods.

It is interesting to note that the magnetic radius of the neutron is consistent within errors with the magnetic radius of the proton. In fact the magnetic radius of the proton is also consistent within errors with the value of the charge radius of the proton, \( r_E^p = 0.871 \pm 0.009 \pm 0.002 \pm 0.002 \) fm extracted using the same model-independent methods in [76]. We
will not interpret these results here, but using model-independent methods is essential in establishing these facts.

Our study shows the utility and robustness of the $z$-expansion in model-independent extraction of fundamental properties of nucleons such as the electric, magnetic, and axial radii. It would be interesting to apply the same methods to newer data sets such as that of the A1 collaboration and to include also polarization data.

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References

[1] R. Frisch and O. Stern, Zeitschrift für Physik 85, 4 (1933).

[2] http://www.nobelprize.org/nobel_prizes/physics/laureates/1943/stern-lecture.pdf

[3] J. R. Green, J. W. Negele, A. V. Pochinsky, S. N. Syritsyn, M. Engelhardt and S. Krieg, arXiv:1404.4029 [hep-lat].

[4] R. Pohl et al., Nature 466, 213 (2010).

[5] A. Antognini, F. Nez, K. Schuhmann, F. D. Amaro, FrancoisBiraben, J. M. R. Cardoso, D. S. Covita and A. Dax et al., Science 339, 417 (2013).

[6] P. J. Mohr, B. N. Taylor and D. B. Newell, Rev. Mod. Phys. 84, 1527 (2012) arXiv:1203.5425 [physics.atom-ph].

[7] F. Garcia Daza, N. G. Kelkar and M. Nowakowski, J. Phys. G 39, 035103 (2012) arXiv:1008.4384 [hep-ph].

[8] U. D. Jentschura, Annals Phys. 326, 500 (2011) arXiv:1011.5275 [hep-ph].

[9] R. J. Hill, G. Paz, Phys. Rev. Lett. 107, 160402 (2011), arXiv:1103.4617 [hep-ph]

[10] J. D. Carroll, A. W. Thomas, J. Rafelski and G. A. Miller, Phys. Rev. A 84, 012506 (2011) [arXiv:1104.2971][physics.atom-ph].

[11] M. I. Eides, Phys. Rev. A 85, 034503 (2012) arXiv:1201.2979 [physics.atom-ph].

[12] N. G. Kelkar, F. G. Daza and M. Nowakowski, Nucl. Phys. B 864, 382 (2012) arXiv:1203.0581 [hep-ph].

[13] E. Borie, Annals Phys. 327, 733 (2012).
[14] A. Antognini, F. Kottmann, F. Biraben, P. Indelicato, F. Nez and R. Pohl, Annals Phys. 331, 127 (2013) [arXiv:1208.2637 [physics.atom-ph]].

[15] P. Indelicato, Phys. Rev. A 87, no. 2, 022501 (2013) [arXiv:1210.5828 [physics.atom-ph]].

[16] K. M. Graczyk, Phys. Rev. C 88, 065205 (2013) [arXiv:1306.5991 [hep-ph]].

[17] D. -Y. Chen and Y. -B. Dong, Phys. Rev. C 87, no. 4, 045209 (2013).

[18] M. M. Giannini and E. Santopinto, arXiv:1311.0319 [hep-ph].

[19] E. Y. Y. Korzinin, V. G. Ivanov and S. G. Karshenboim, Phys. Rev. D 88, no. 12, 125019 (2013) [arXiv:1311.5784 [physics.atom-ph]].

[20] P. Indelicato, P. J. Mohr and J. Sapirstein, arXiv:1402.0439 [quant-ph].

[21] R. N. Faustov, A. P. Martynenko, G. A. Martynenko and V. V. Sorokin, Phys. Lett. B 733, 354 (2014) [arXiv:1402.5825 [hep-ph]].

[22] S. G. Karshenboim, arXiv:1405.6039 [hep-ph].

[23] T. Friedmann, Eur. Phys. J. C 73, 2298 (2013) [arXiv:0910.2229 [hep-ph]].

[24] T. Friedmann, Eur. Phys. J. C 73, 2299 (2013) [arXiv:0910.2231 [hep-ph]].

[25] A. De Rujula, Phys. Lett. B 693, 555 (2010) [arXiv:1008.3861 [hep-ph]].

[26] I. C. Cloet and G. A. Miller, Phys. Rev. C 83, 012201 (2011) [arXiv:1008.4345 [hep-ph]].

[27] M. Vanderhaeghen and T. Walcher, Nucl. Phys. News 21, 14 (2011) [arXiv:1008.4225 [hep-ph]].

[28] A. De Rujula, arXiv:1008.4546 [hep-ph].

[29] A. Kholmetskii, O. Missevitch and T. Yarman, arXiv:1010.2845 [physics.atom-ph].

[30] A. De Rujula, Phys. Lett. B 697, 26 (2011) [arXiv:1010.3421 [hep-ph]].

[31] M. O. Distler, J. C. Bernauer and T. Walcher, Phys. Lett. B 696, 343 (2011) [arXiv:1011.1861 [nucl-th]].

[32] G. A. Miller, A. W. Thomas, J. D. Carroll and J. Rafelski, Phys. Rev. A 84, 020101 (2011) [arXiv:1101.4073 [physics.atom-ph]].

[33] C. E. Carlson and M. Vanderhaeghen, Phys. Rev. A 84, 020102 (2011) [arXiv:1101.5965 [hep-ph]].

[34] A. Pineda, arXiv:1108.1263 [hep-ph].

[35] B. Y. Wu and C. W. Kao, arXiv:1108.2968 [hep-ph].
[36] C. E. Carlson and M. Vanderhaeghen, arXiv:1109.3779 [physics.atom-ph].

[37] A. L. Kholmetsky, O. V. Missevitch and T. Yarman, Eur. Phys. J. Plus 127, 44 (2012).

[38] J. -P. Karr and L. Hilico, Phys. Rev. Lett. 109, 103401 (2012) arXiv:1205.0633 [physics.atom-ph].

[39] M. C. Birse and J. A. McGovern, Eur. Phys. J. A 48, 120 (2012) arXiv:1206.3030 [hep-ph].

[40] G. A. Miller, A. W. Thomas and J. D. Carroll, Phys. Rev. C 86, 065201 (2012) arXiv:1207.0549 [nucl-th].

[41] G. A. Miller, Phys. Lett. B 718, 1078 (2013) arXiv:1209.4667 [nucl-th].

[42] O. W. Greenberg and S. Cowen, Phys. Rev. A 87, 042516 (2013) arXiv:1211.1619 [quant-ph].

[43] M. Gorchtein, F. J. Llanes-Estrada and A. P. Szczepaniak, Phys. Rev. A 87, 052501 (2013) arXiv:1302.2807 [nucl-th].

[44] T. Mart and A. Sulaksono, Phys. Rev. C 87, 025807 (2013) arXiv:1302.6012 [nucl-th].

[45] P. J. Mohr, J. Griffith and J. Sapirstein, Phys. Rev. A 87, no. 5, 052511 (2013) arXiv:1304.2076 [hep-ph].

[46] D. Robson, arXiv:1305.4552 [nucl-th].

[47] J. M. Alarcon, V. Lensky and V. Pascalutsa, Eur. Phys. J. C 74, 2852 (2014) arXiv:1312.1219 [hep-ph].

[48] E. J. Downie, W. J. Briscoe, R. Gilman and G. Ron, Phys. Rev. C 88, no. 5, 059801 (2013).

[49] U. D. Jentschura, Phys. Rev. A 88, 062514 (2013) arXiv:1401.3666 [physics.atom-ph].

[50] M. I. Eides, arXiv:1402.5860 [hep-ph].

[51] C. Peset and A. Pineda, arXiv:1403.3408 [hep-ph].

[52] R. K. Gainutdinov and A. A. Mutygullina, Bull. Russ. Acad. Sci. Phys. 78, 189 (2014) [Izv. Ross. Akad. Nauk Ser. Fiz. 78, 289292 (2014)].

[53] O. Tomalak and M. Vanderhaeghen, Phys. Rev. D 90, 013006 (2014) arXiv:1405.1600 [hep-ph].

[54] K. Pachucki and K. A. Meissner, arXiv:1405.6582 [hep-ph].

[55] S. D. Glazek, arXiv:1406.0127 [hep-th].
[56] M. Gorchtein, arXiv:1406.1612 [nucl-th].

[57] C. Peset and A. Pineda, arXiv:1406.4524 [hep-ph].

[58] J. Jaeckel and S. Roy, Phys. Rev. D 82, 125020 (2010) arXiv:1008.3536 [hep-ph].

[59] S. K. Kauffmann, Prespace. J. 1, 1295 (2010) arXiv:1009.3584 [physics.gen-ph].

[60] P. Brax and C. Burrage, Phys. Rev. D 83, 035020 (2011) arXiv:1010.5108 [hep-ph].

[61] V. Barger, C. -W. Chiang, W. -Y. Keung and D. Marfatia, Phys. Rev. Lett. 106, 153001 (2011) arXiv:1011.3519 [hep-ph].

[62] D. Tucker-Smith and I. Yavin, Phys. Rev. D 83, 101702 (2011) arXiv:1011.4922 [hep-ph].

[63] B. Batell, D. McKeen and M. Pospelov, Phys. Rev. Lett. 107, 011803 (2011) arXiv:1103.0721 [hep-ph].

[64] J. I. Rivas, A. Camacho and E. Goeklue, Phys. Rev. D 84, 055024 (2011) arXiv:1105.6345 [gr-qc].

[65] V. Barger, C. -W. Chiang, W. -Y. Keung and D. Marfatia, Phys. Rev. Lett. 108, 081802 (2012) arXiv:1109.6652 [hep-ph].

[66] C. E. Carlson and B. C. Rislow, Phys. Rev. D 86, 035013 (2012) arXiv:1206.3587 [hep-ph].

[67] L. -B. Wang and W. -T. Ni, Mod. Phys. Lett. A 28, 1350094 (2013) arXiv:1303.4885 [hep-ph].

[68] Z. Li and X. Chen, arXiv:1303.5146 [hep-ph].

[69] M. Moumni and A. BenSlama, Int. J. Mod. Phys. A 28, 1350139 (2013) arXiv:1305.3508 [hep-ph].

[70] C. E. Carlson and B. C. Rislow, Phys. Rev. D 89, 035003 (2014) arXiv:1310.2786 [hep-ph].

[71] W. -F. Chang, J. N. Ng and J. M. S. Wu, Phys. Lett. B 730, 347 (2014) arXiv:1310.6513 [hep-ph].

[72] R. Onofrio, Europhys. Lett. 104, 20002 (2013) arXiv:1312.3469 [hep-ph].

[73] S. G. Karshenboim, D. McKeen and M. Pospelov, arXiv:1401.6154 [hep-ph].

[74] W. Ubachs, W. Vassen, E. J. S. and and K. S. E. Eikema, Annalen Phys. 525, A113 (2013).

[75] P. Brax and C. Burrage, arXiv:1407.2376 [hep-ph].
[76] R. J. Hill and G. Paz, Phys. Rev. D 82, 113005 (2010) [arXiv:1008.4619 [hep-ph]].

[77] J. Arrington, W. Melnitchouk and J. A. Tjon, Phys. Rev. C 76, 035205 (2007) [arXiv:0707.1861 [nucl-ex]].

[78] A. Lung, L. M. Stuart, P. E. Bosted, L. Andivahis, J. Alster, R. G. Arnold, C. C. Chang and F. S. Dietrich et al., Phys. Rev. Lett. 70, 718 (1993).

[79] H. Gao, J. Arrington, E. J. Beise, B. Bray, R. W. Carr, B. W. Filippone, A. Lung and R. D. McKeown et al., Phys. Rev. C 50, 546 (1994).

[80] H. Anklin, D. Fritschi, J. Jourdan, M. Loppacher, G. Masson, I. Sick, E. E. W. Bruins and F. C. P. Joosse et al., Phys. Lett. B 336, 313 (1994).

[81] H. Anklin, L. J. deBever, K. I. Blomqvist, W. U. Boeglin, R. Bohm, M. Distler, R. Edelhoff and J. Friedrich et al., Phys. Lett. B 428, 248 (1998).

[82] G. Kubon, H. Anklin, P. Bartsch, D. Baumann, W. U. Boeglin, K. Bohinc, R. Bohm and M. O. Distler et al., Phys. Lett. B 524, 26 (2002) [nucl-ex/0107016].

[83] B. Anderson et al. [Jefferson Lab E95-001 Collaboration], Phys. Rev. C 75, 034003 (2007) [nucl-ex/0605006].

[84] J. Lachniet et al. [CLAS Collaboration], Phys. Rev. Lett. 102, 192001 (2009) [arXiv:0811.1716 [nucl-ex]].

[85] J. C. Bernauer et al. [A1 Collaboration], Phys. Rev. Lett. 105, 242001 (2010) [arXiv:1007.5076 [nucl-ex]].

[86] D. Borisyuk, Nucl. Phys. A 843, 59 (2010) [arXiv:0911.4091 [hep-ph]].

[87] M. A. Belushkin, H. W. Hammer and U. G. Meissner, Phys. Rev. C 75, 035202 (2007) [arXiv:hep-ph/0608337].

[88] J. Beringer et al. [Particle Data Group Collaboration], Phys. Rev. D 86, 010001 (2012).

[89] L. L. Foldy, Phys. Rev. 87, 688 (1952).

[90] G. Salzmann, Phys. Rev. 99, 973 (1955).

[91] F. J. Ernst, R. G. Sachs and K. C. Wali, Phys. Rev. 119, 1105 (1960).

[92] G. Hohler and E. Pietarinen, Nucl. Phys. B 95, 210 (1975).

[93] J. Schwinger, Annals Phys. 9, 169-193 (1960).

[94] B. Bhattacharya, R. J. Hill and G. Paz, Phys. Rev. D 84, 073006 (2011) [arXiv:1108.0423 [hep-ph]].

[95] P. Federbush, M. L. Goldberger and S. B. Treiman, Phys. Rev. 112, 642 (1958).
[96] W. R. Frazer and J. R. Fulco, Phys. Rev. 117, 1609 (1960).

[97] G. Höhler, Pion-nucleon scattering, in: H. Schopper (editor), Landolt-Börnstein database, Volume 9, subvolume b, part 1, Springer-Verlag, Berlin, 1983. [http://www.springermaterials.com/navigation/]

[98] S. R. Amendolia et al., Phys. Lett. B 138, 454 (1984).

[99] M. N. Achasov et al., J. Exp. Theor. Phys. 101, 1053 (2005) [Zh. Eksp. Teor. Fiz. 101, 1201 (2005)] [arXiv:hep-ex/0506076].

[100] N. Cabibbo and R. Gatto, Phys. Rev. 124, 1577 (1961).

[101] M. Ablikim et al. [BES Collaboration], Phys. Lett. B 630, 14 (2005) [arXiv:hep-ex/0506059].

[102] A. Antonelli et al., Nucl. Phys. B 517, 3 (1998).

[103] W. Xu, D. Dutta, F. Xiong, B. Anderson, L. Auberbach, T. Averett, W. Bertozzi and T. Black et al., Phys. Rev. Lett. 85, 2900 (2000) [nucl-ex/0008003].

[104] W. Xu et al. [Jefferson Lab E95-001 Collaboration], Phys. Rev. C 67, 012201 (2003) [nucl-ex/0208007].

[105] http://clasweb.jlab.org/cgi-bin/clasdb/msm.cgi?eid=111&mid=1&data=on

[106] P. Markowitz, J. M. Finn, B. D. Anderson, H. Arenhovel, A. R. Baldwin, D. Barkhuff, K. B. Beard and W. Bertozzi et al., Phys. Rev. C 48, 5 (1993).

[107] E. E. W. Bruins, T. S. Bauer, H. W. den Bok, C. P. Duif, W. C. van Hoek, D. J. J. de Lange, A. Misiejuk and Z. Papandreou et al., Phys. Rev. Lett. 75, 21 (1995).

[108] R. W. Mcallister and R. Hofstadter, Phys. Rev. 102, 851 (1956).

[109] S. Aoki, Y. Aoki, C. Bernard, T. Blum, G. Colangelo, M. Della Morte, S. Drr and A. X. El Khadra et al., arXiv:1310.8555 [hep-lat].

[110] T. Becher and R. J. Hill, Phys. Lett. B 633, 61 (2006) [hep-ph/0509090].

[111] I. Sick, Prog. Part. Nucl. Phys. 55, 440 (2005).

[112] X. Zhan, K. Allada, D. S. Armstrong, J. Arrington, W. Bertozzi, W. Boeglin, J. -P. Chen and K. Chirapatpimol et al., Phys. Lett. B 705, 59 (2011) [arXiv:1102.0318 [nucl-ex]].

[113] I. T. Lorenz, H. -W. Hammer and U. -G. Meissner, Eur. Phys. J. A 48, 151 (2012) [arXiv:1205.6628 [hep-ph]].